

THE CYGNO EXPERIMENT

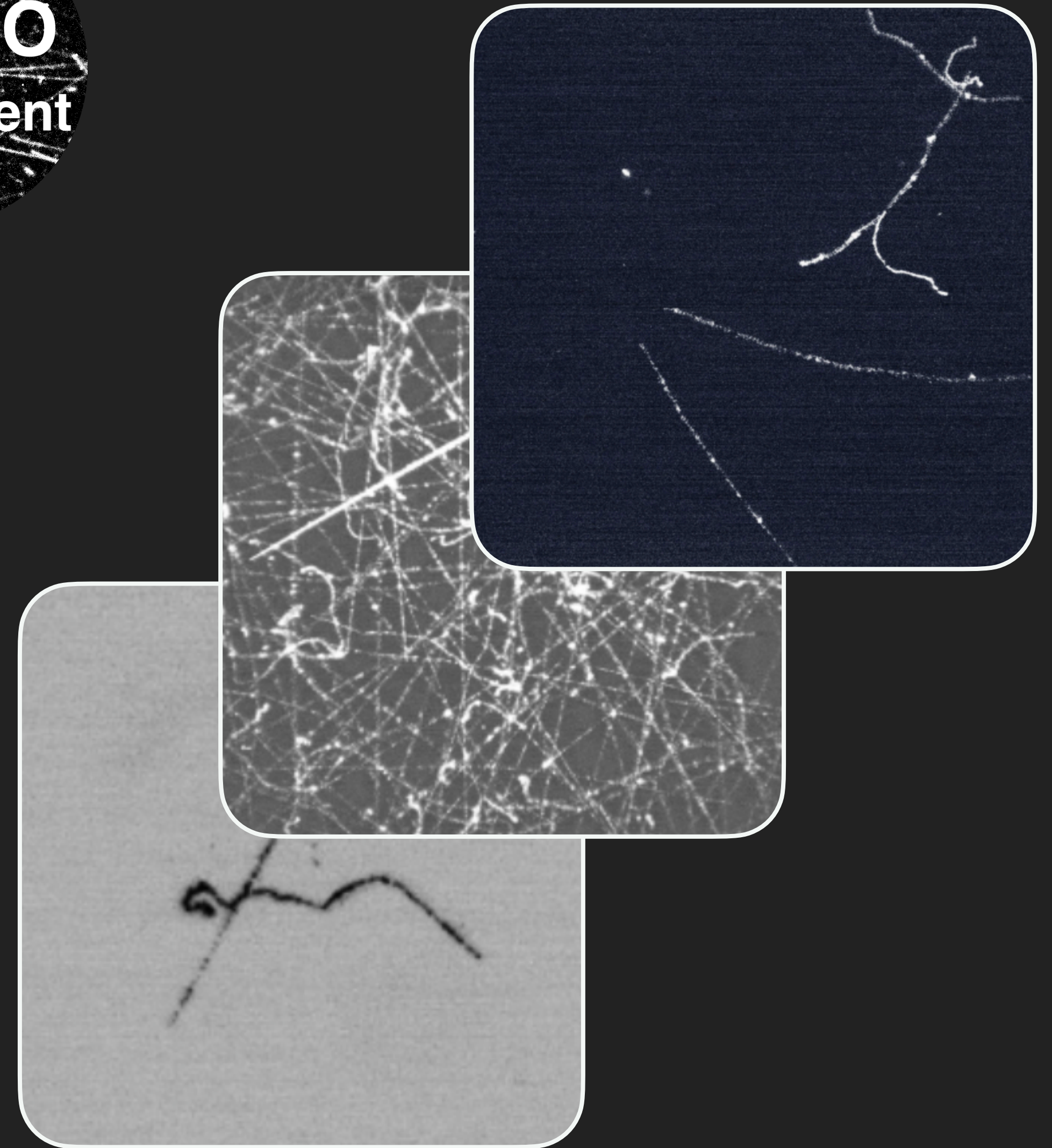


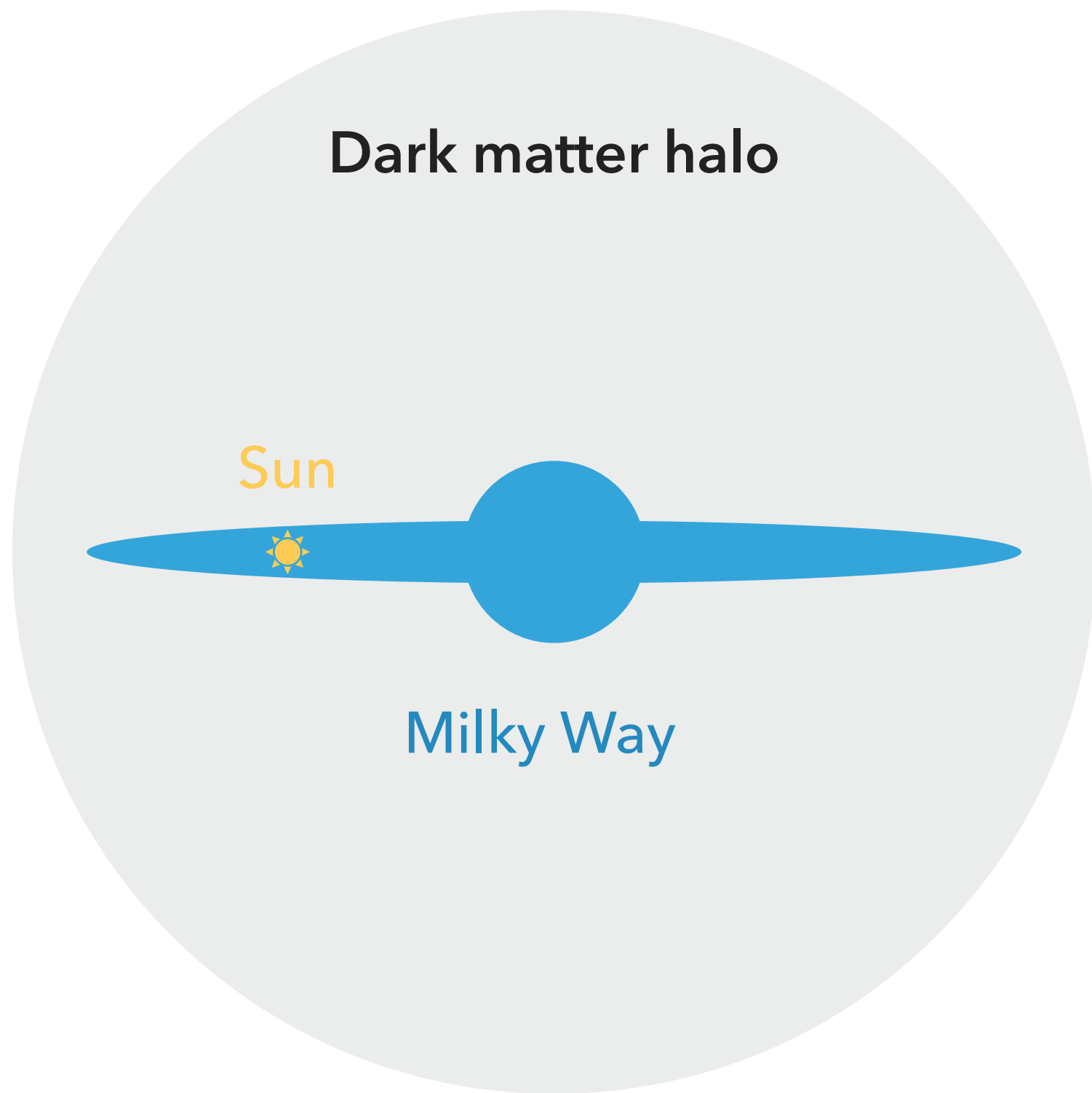
MELBA D'ASTOLFO

On behalf of the CYGNO collaboration

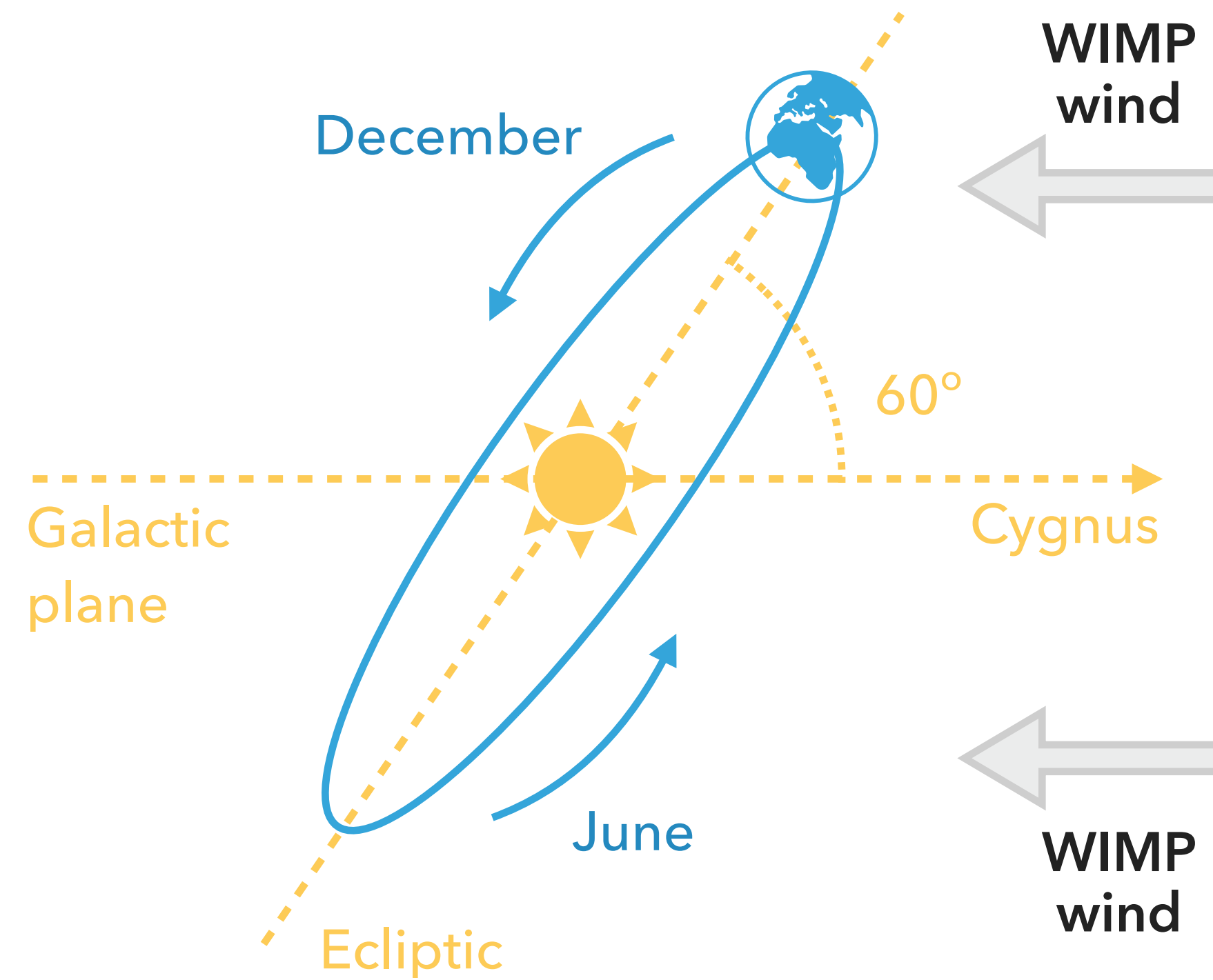
Gran Sasso Science Institute

melba.dastolfo@gssi.it

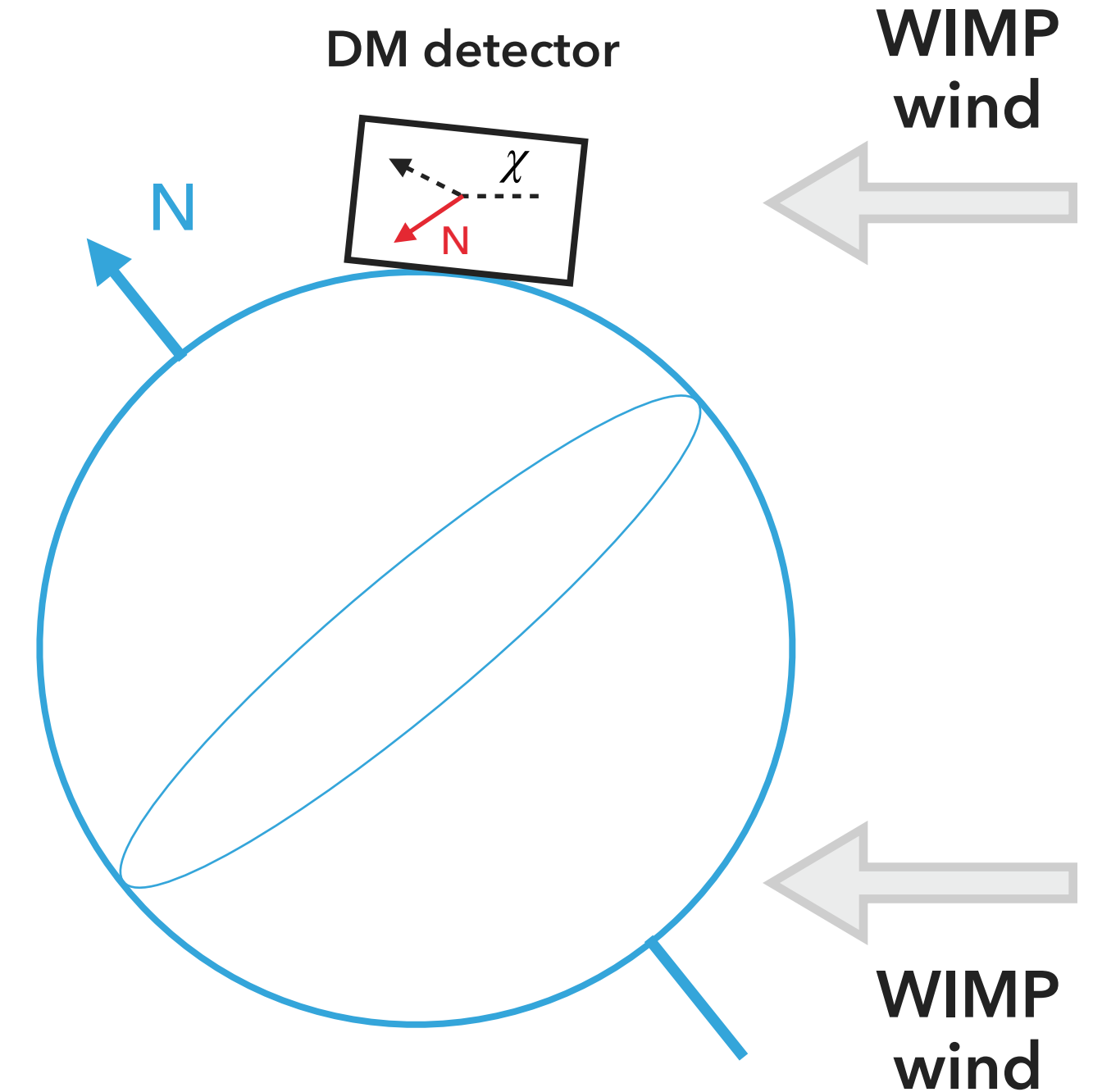




Our galaxy



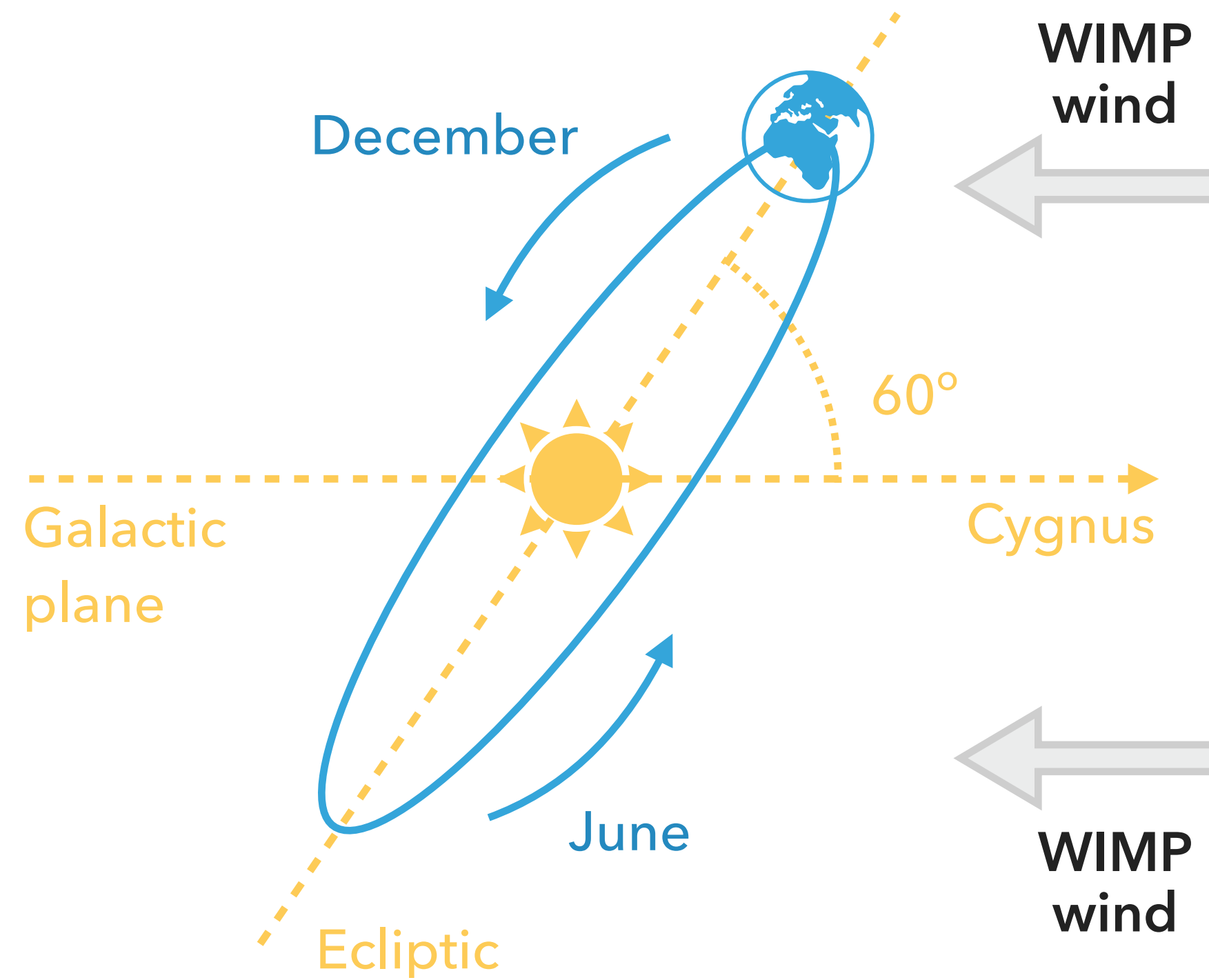
Solar System



Earth



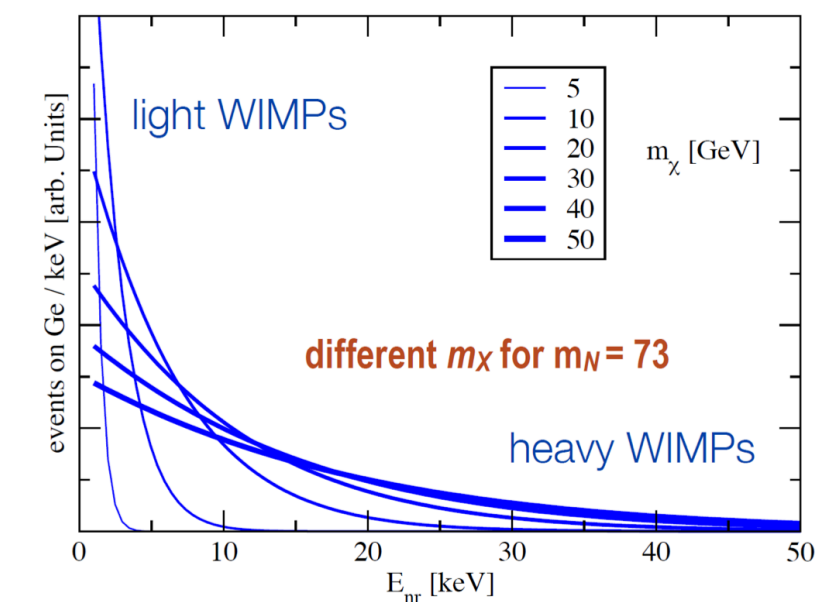
Our galaxy



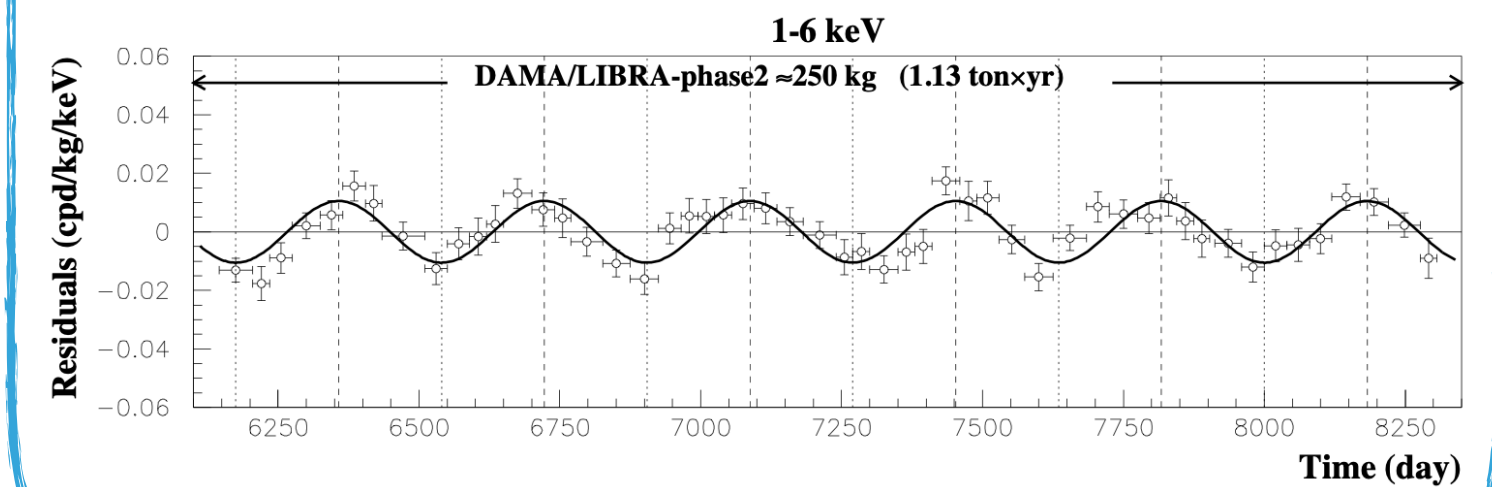
Solar System

Main physical observables

⚡ ENERGY DEPOSIT

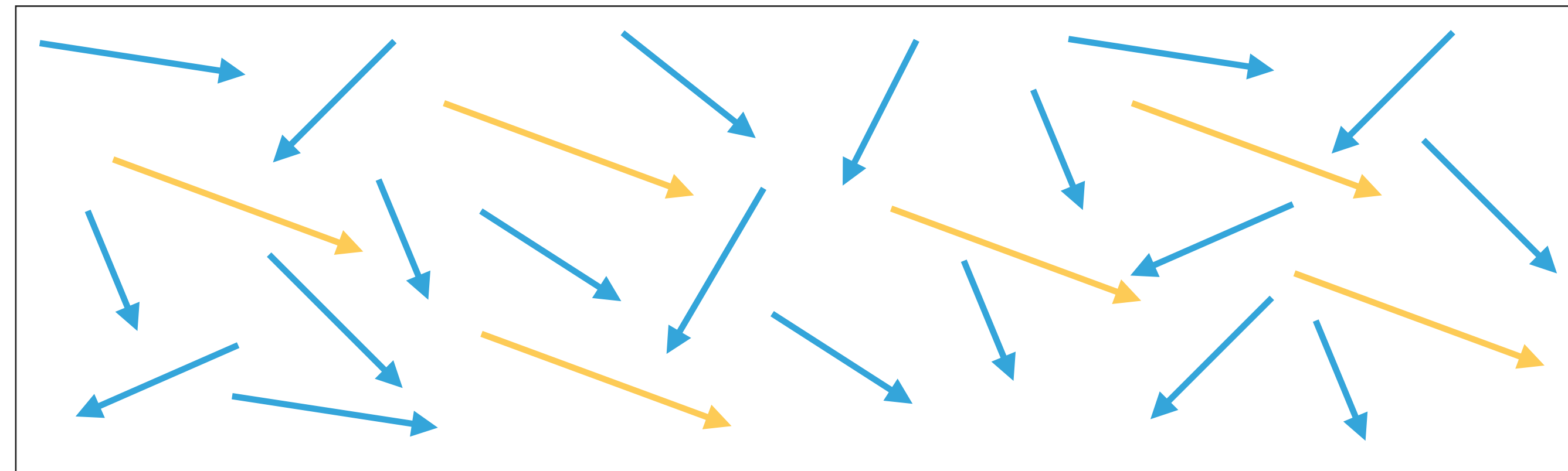
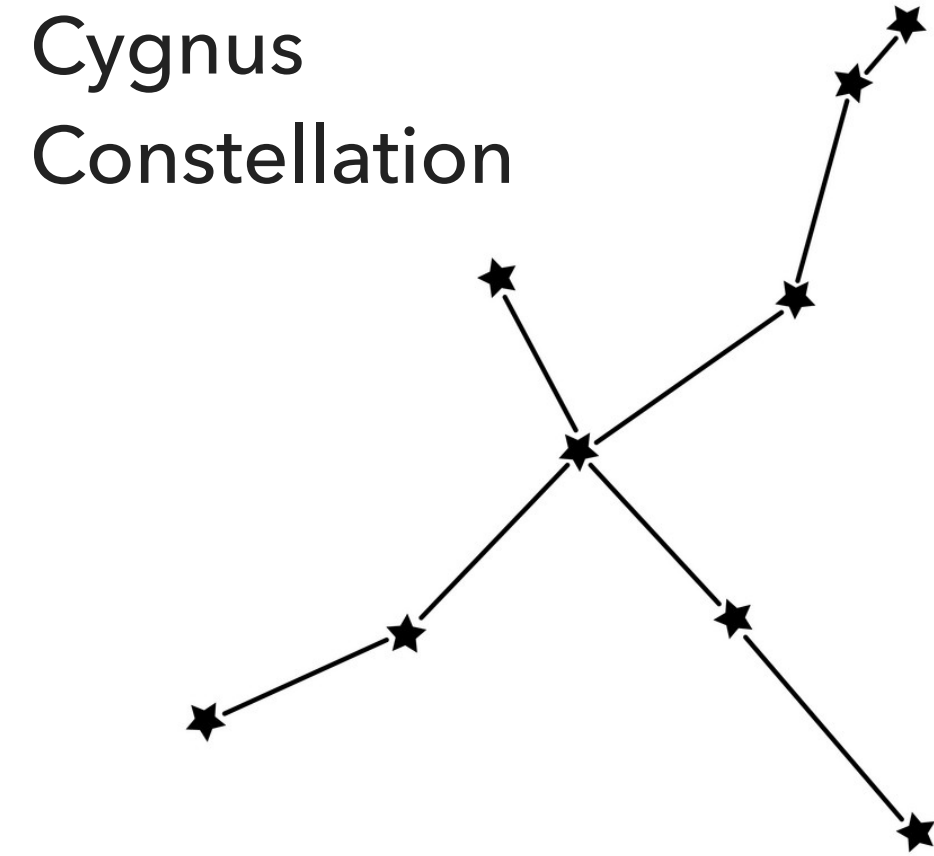


🕒 ANNUAL MODULATION



Bernabei et al., arXiv:1805.10486

Earth

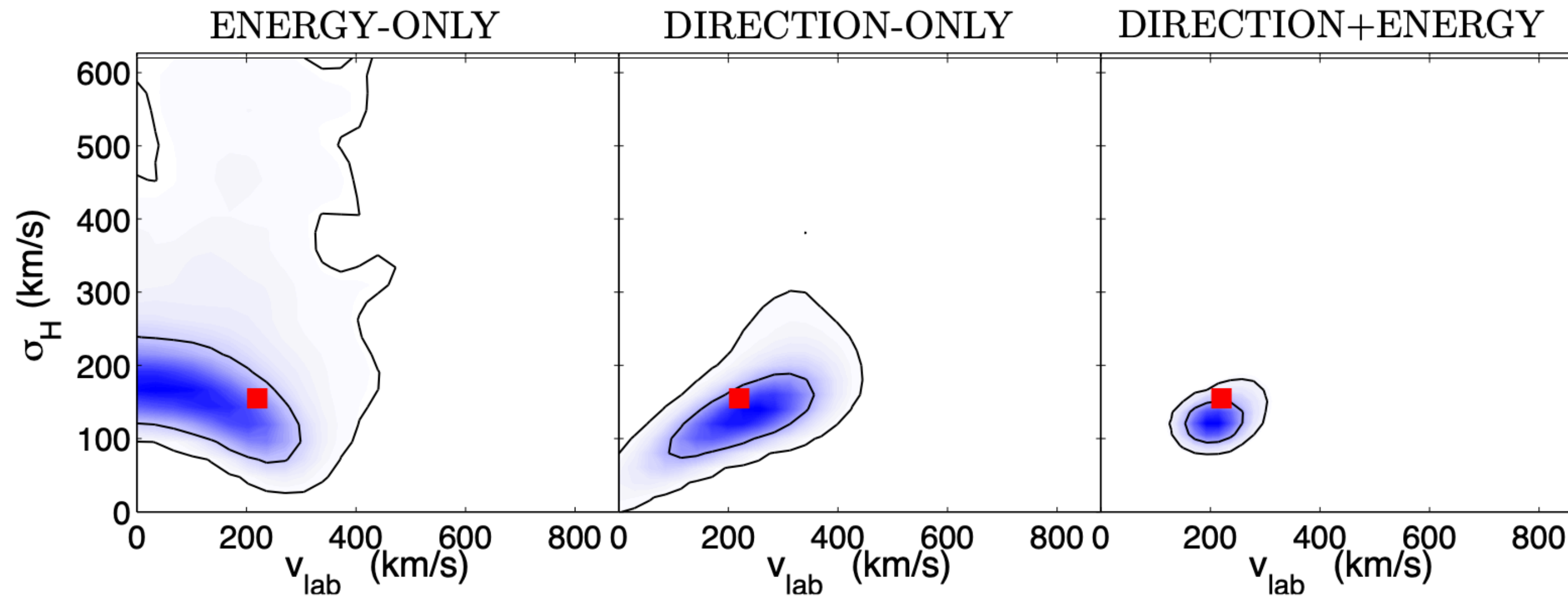


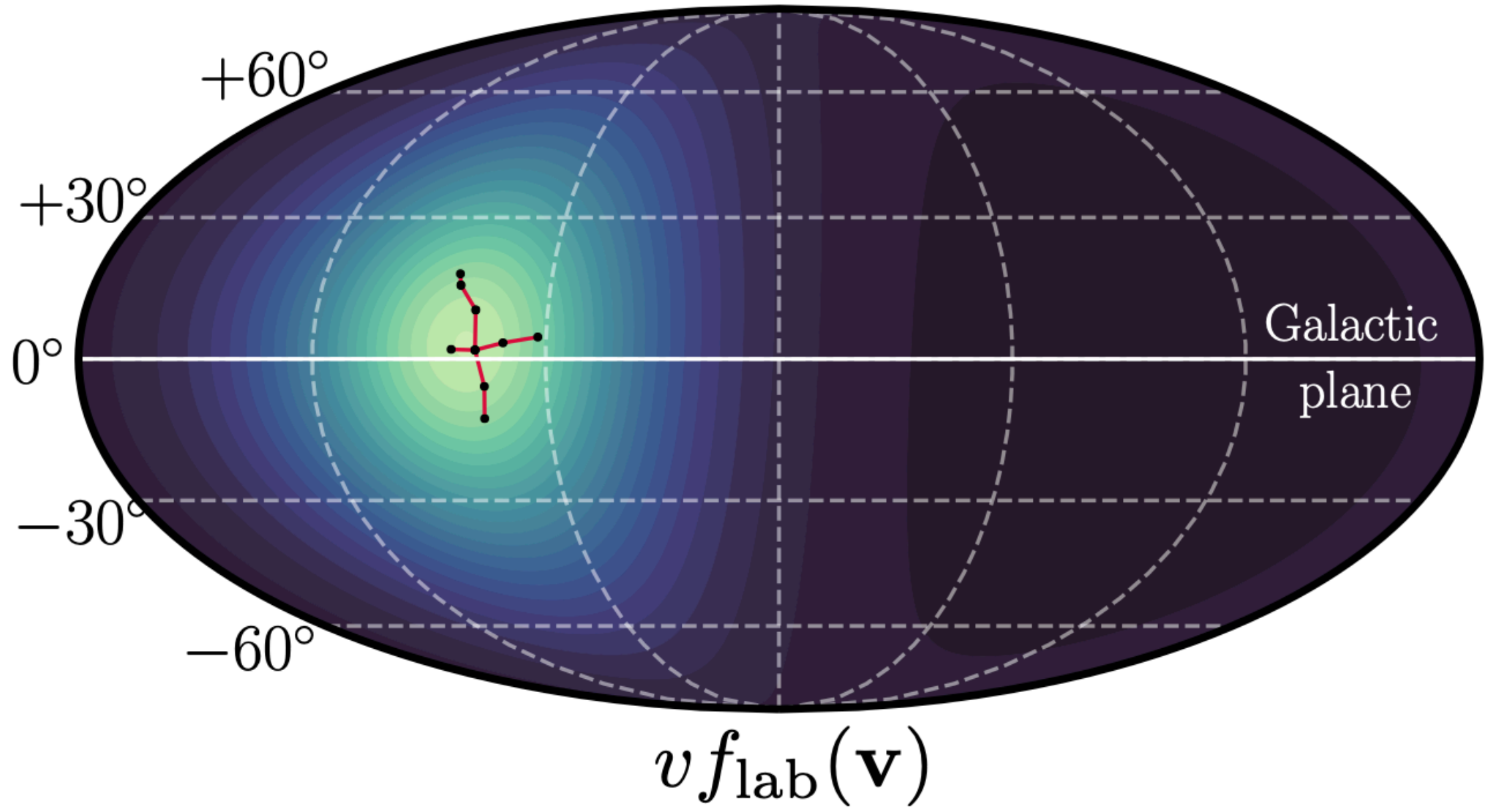
Signal
Background

Directional detection aims to measure **both the energy and the direction** of the recoiling nuclei

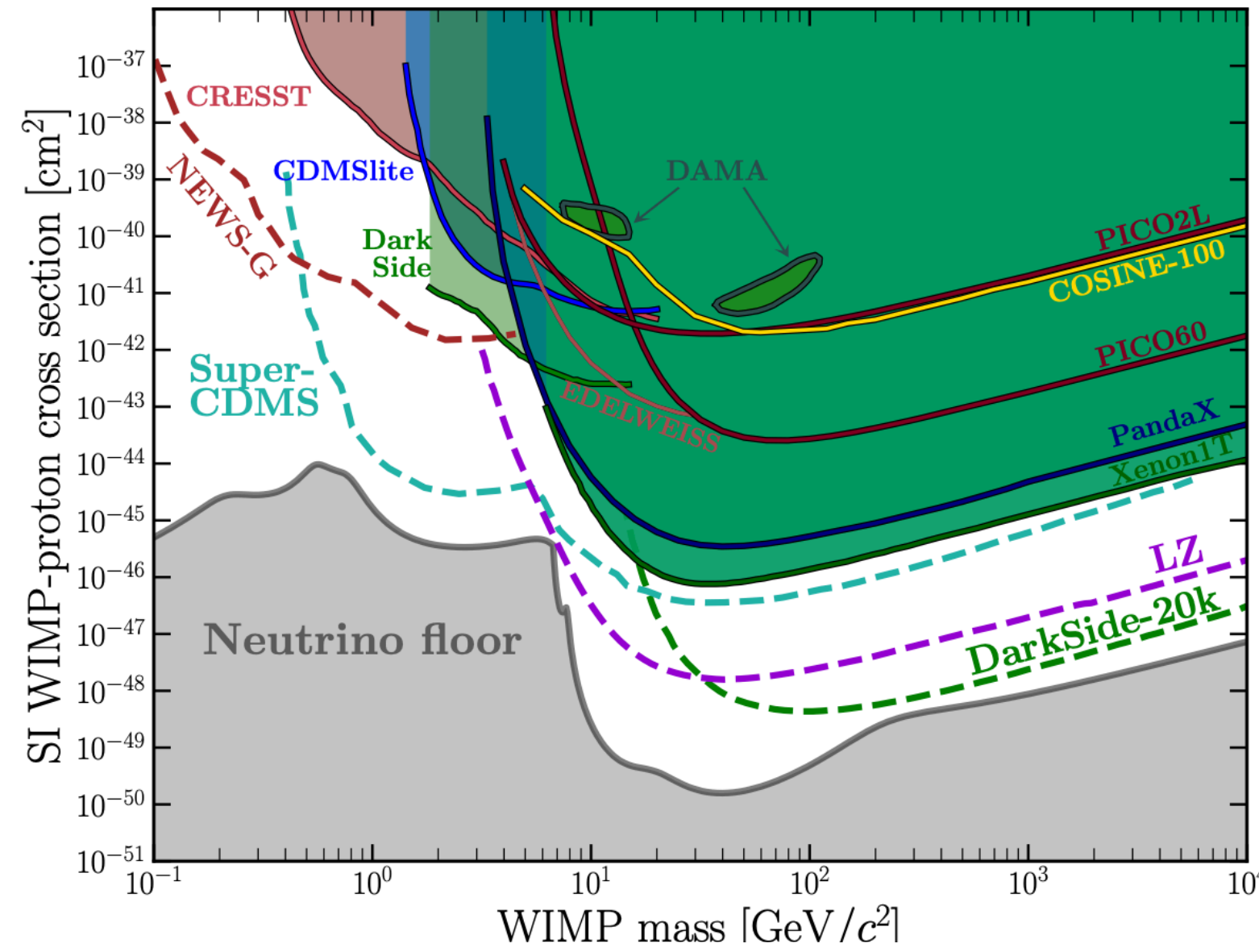
ASSUMPTIONS

- $m_\chi = 50$ GeV
- MIMAC-like experiment



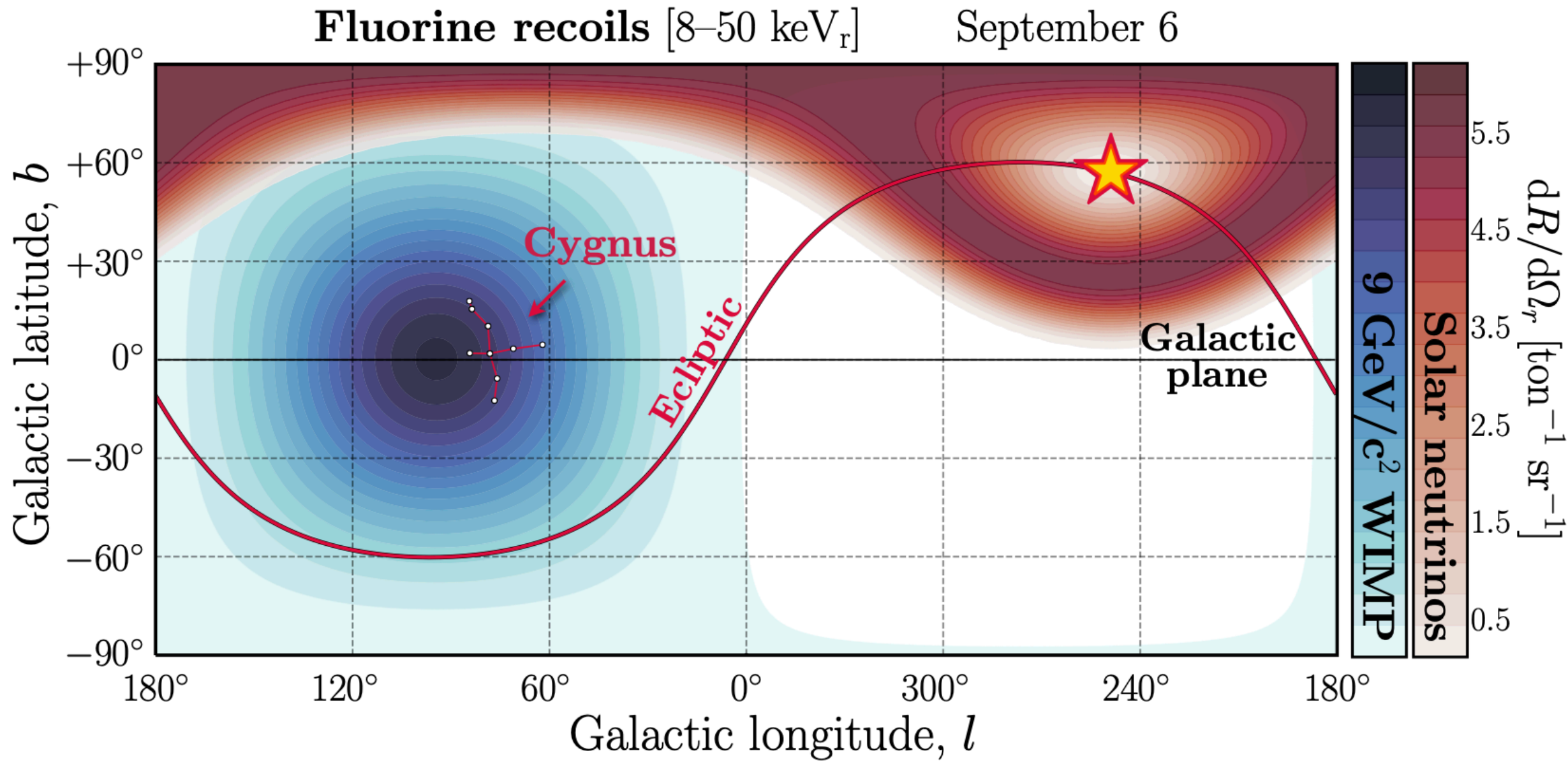
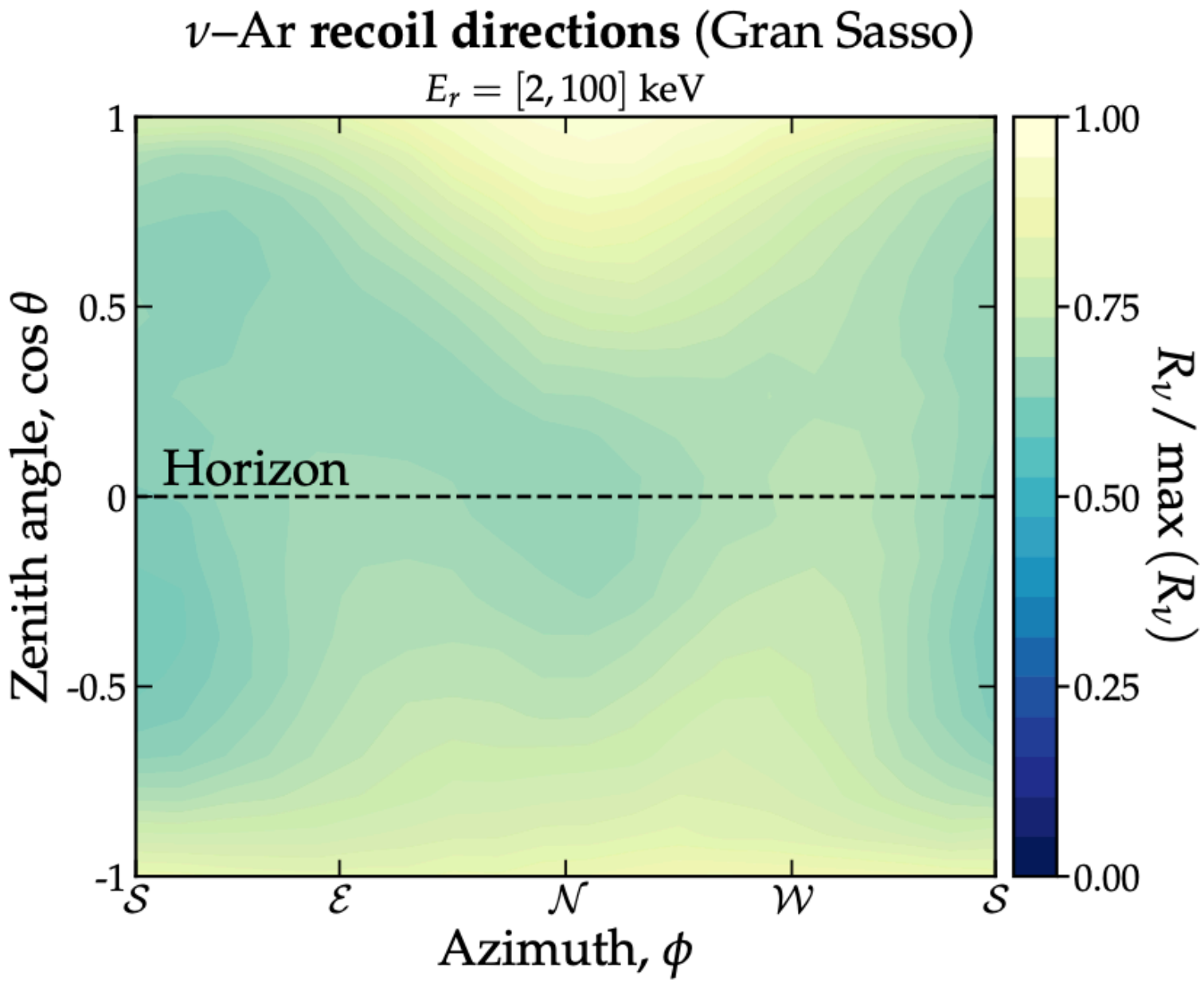


Provide a means of discovering dark matter



Extend searches for WIMPs below the neutrino floor

Capability to discriminate against neutrinos





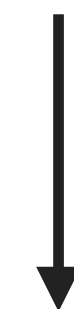
Gaseous TPC

- He:CF₄ (60:40)
- Room temperature
- Atm pressure



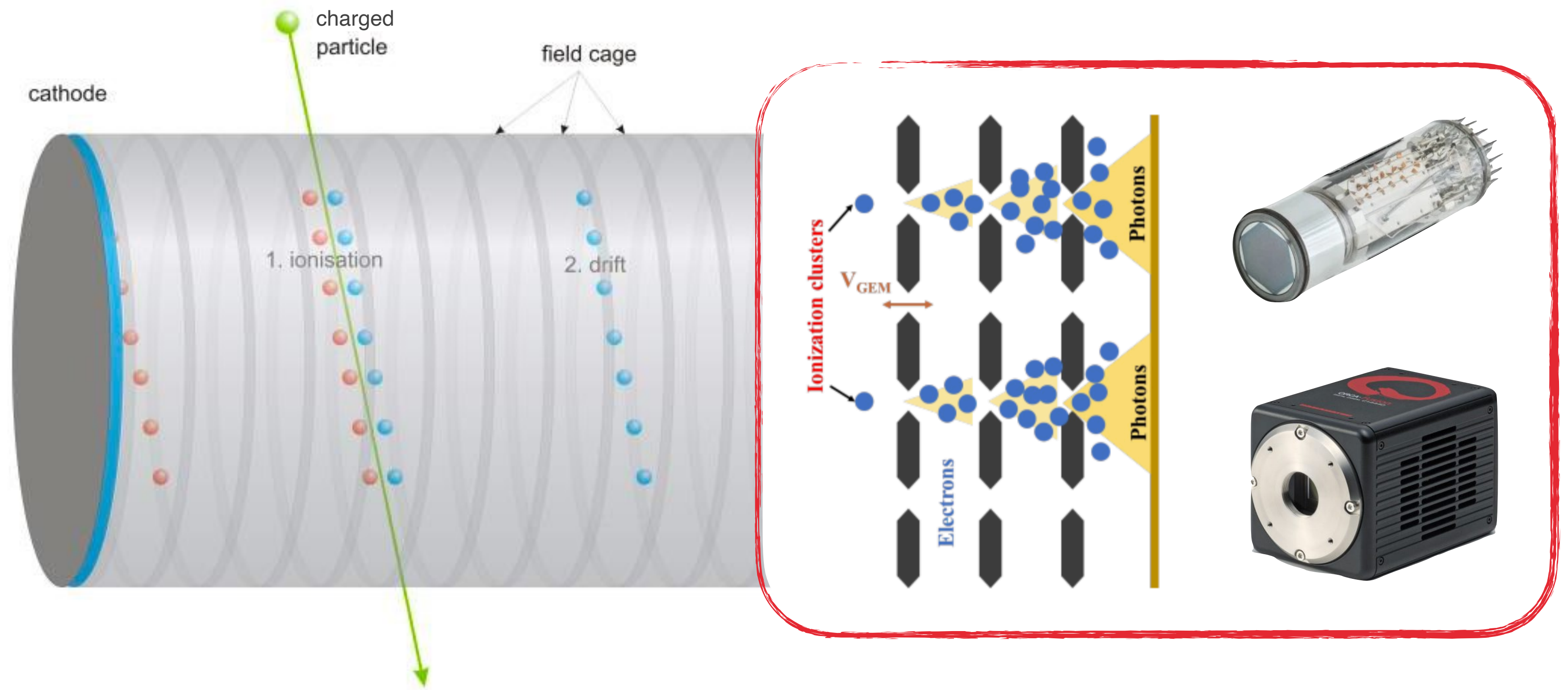
Triple GEM

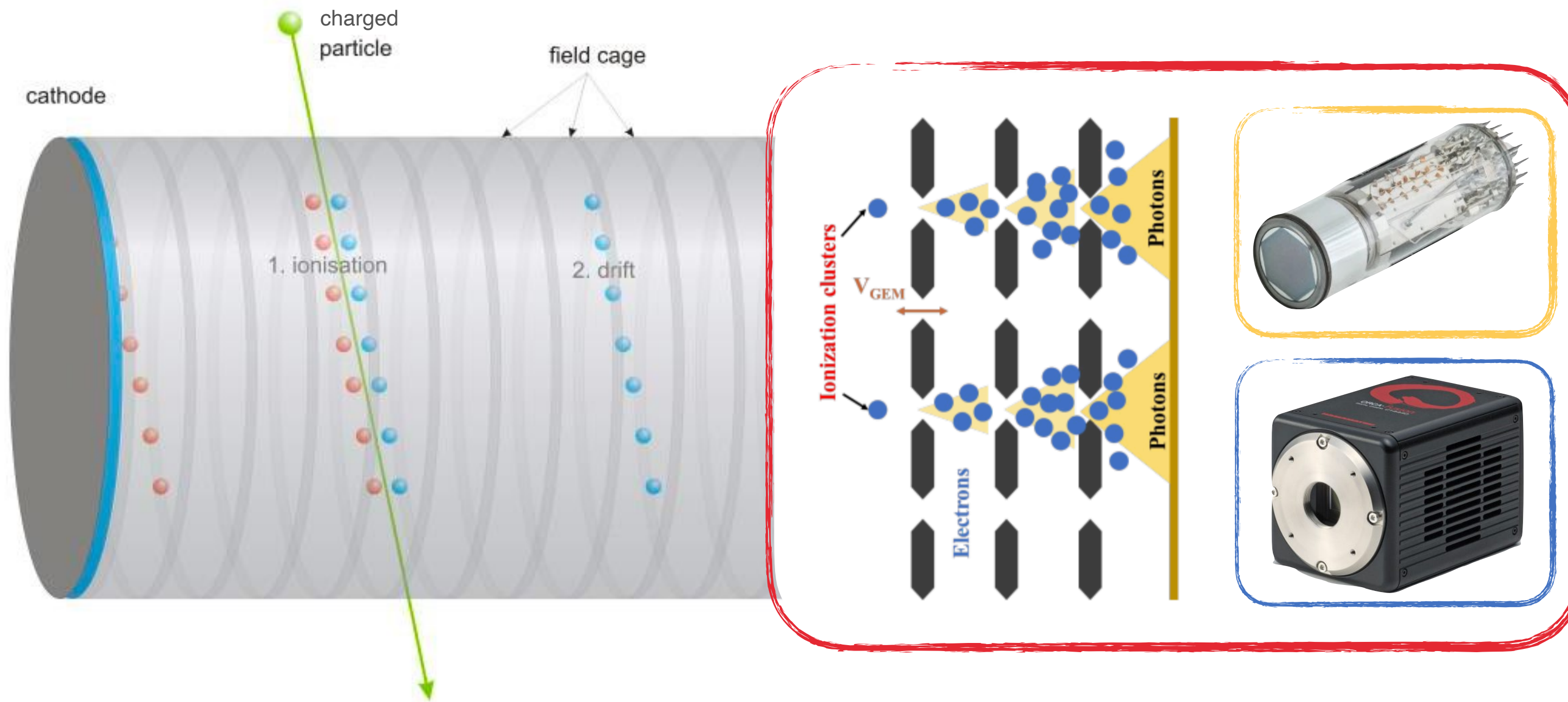
Charge multiplication



Camera + PMT

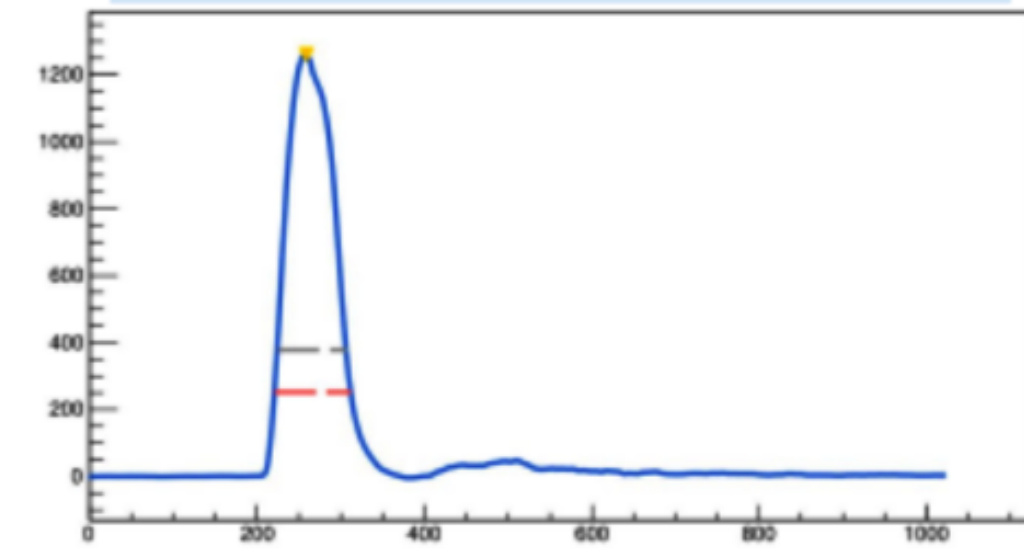
Light from gas scintillation during electron avalanche



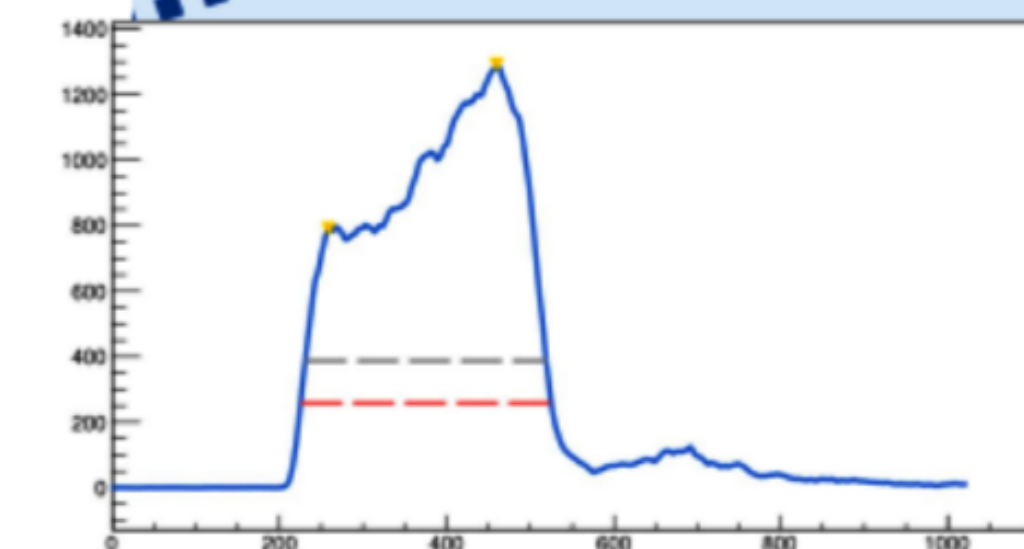


Integrated z + Energy

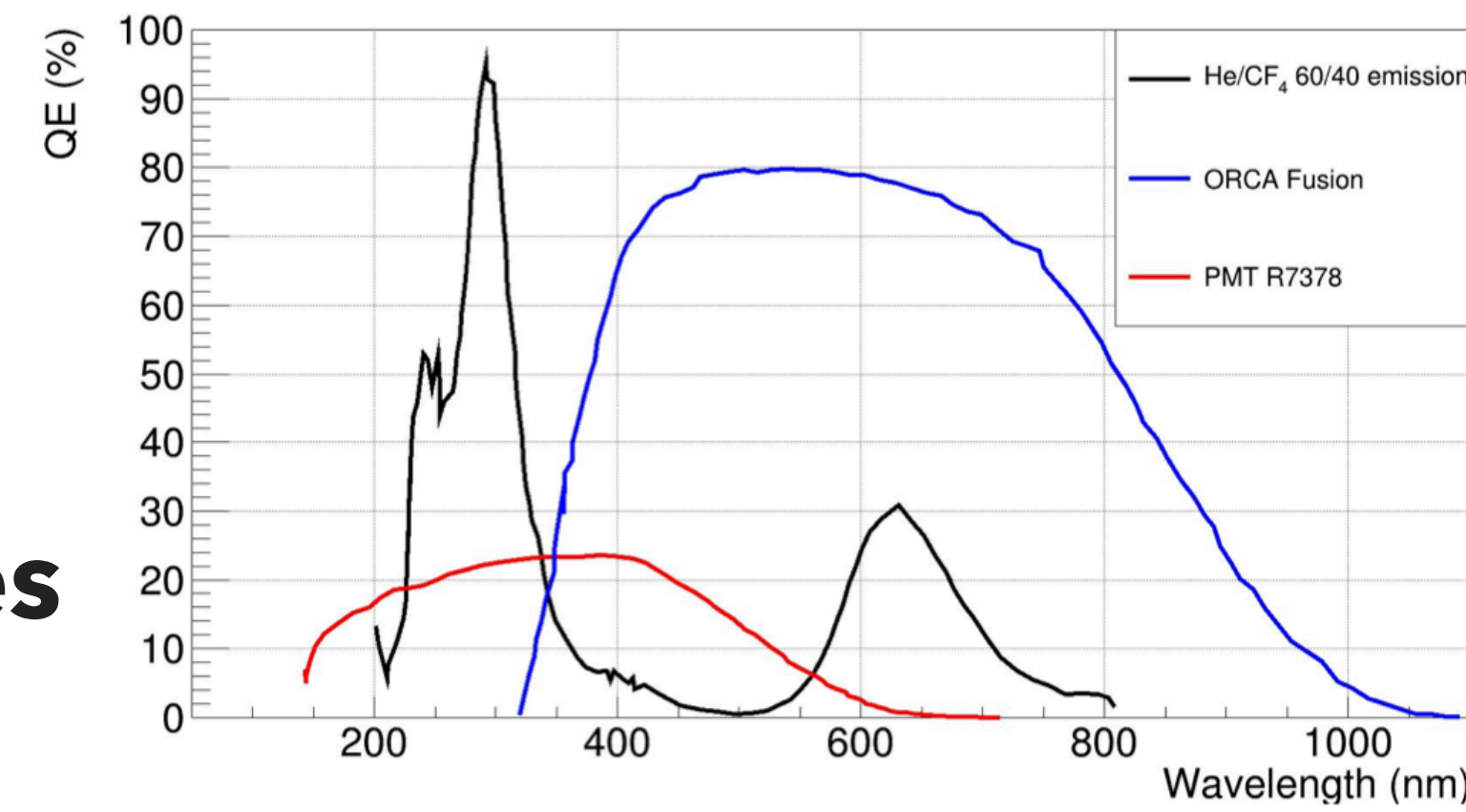
Straight track



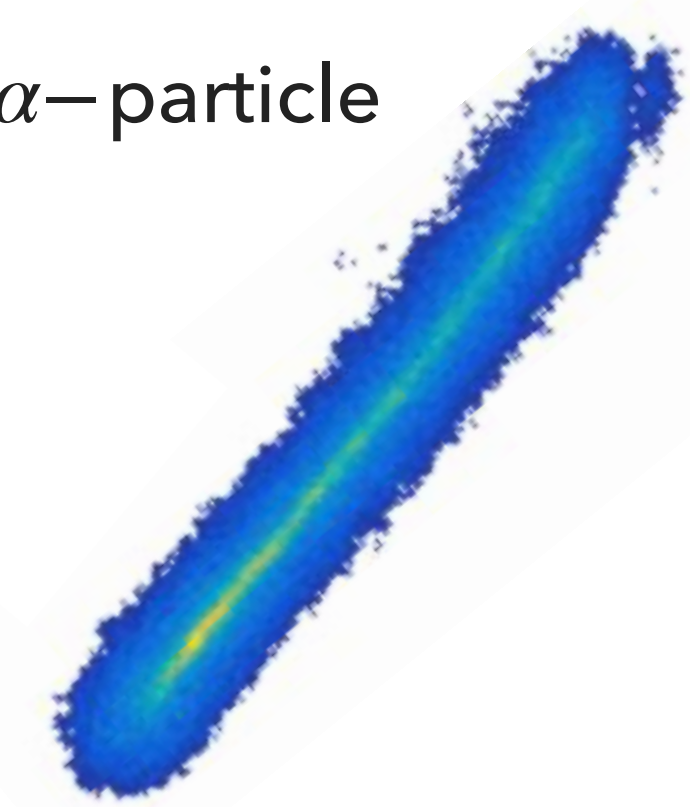
Tilted track



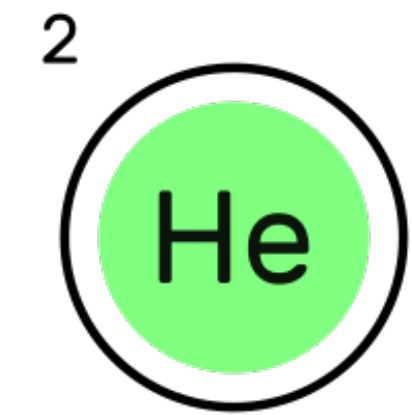
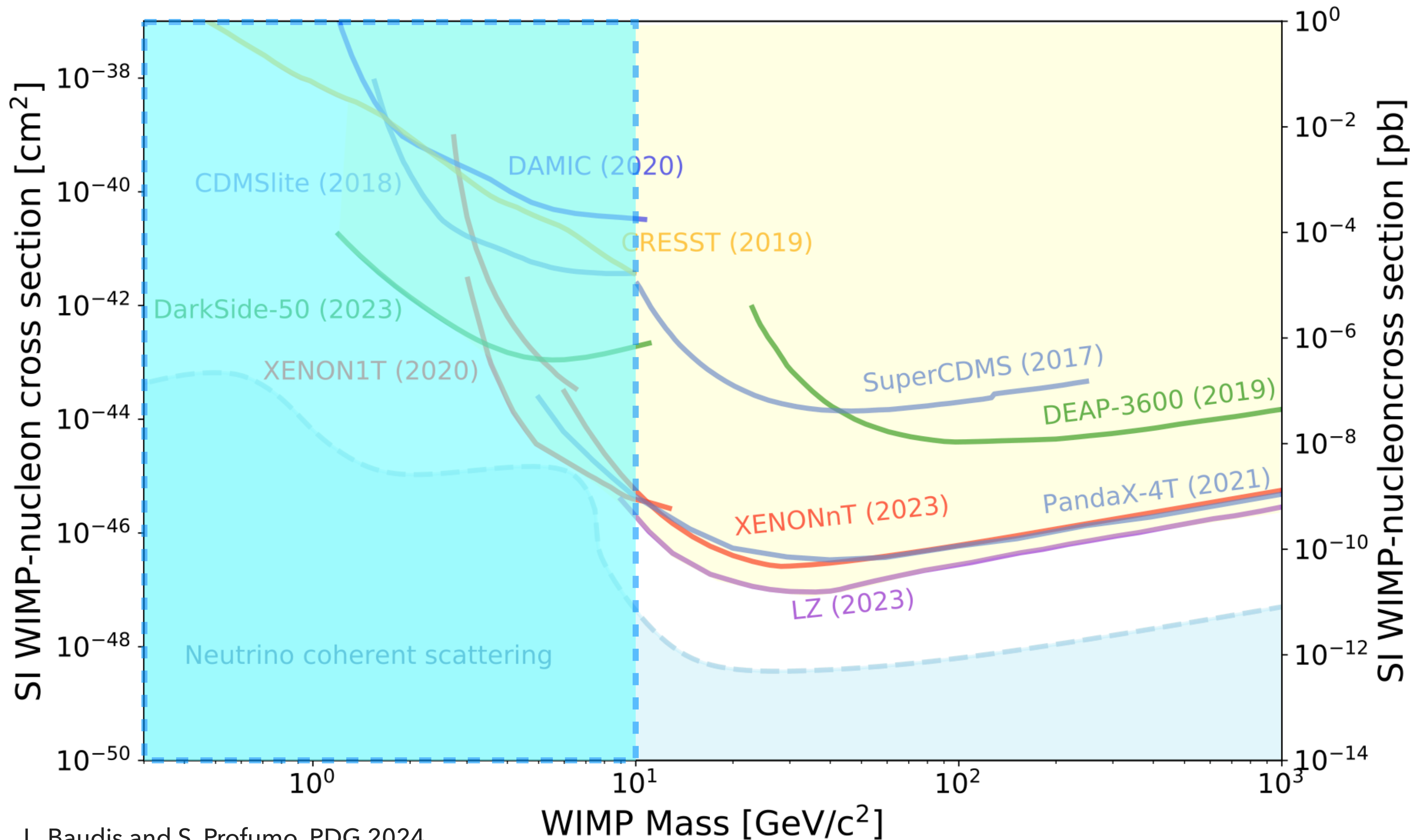
Energy + X & Y coordinates



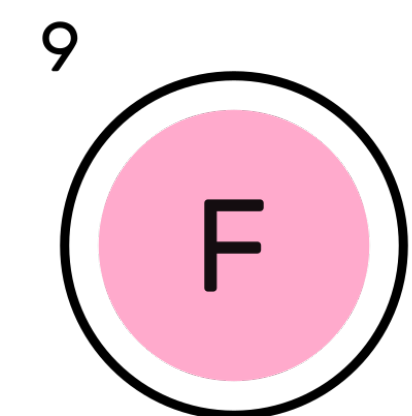
α -particle



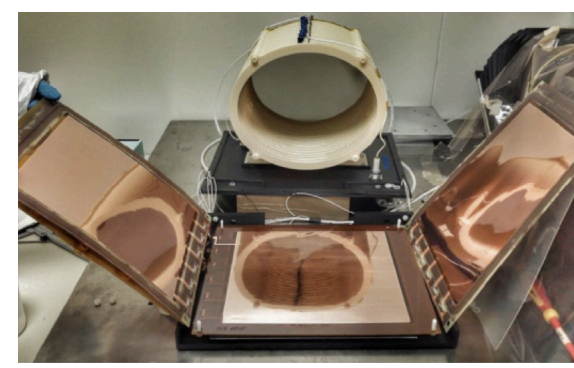
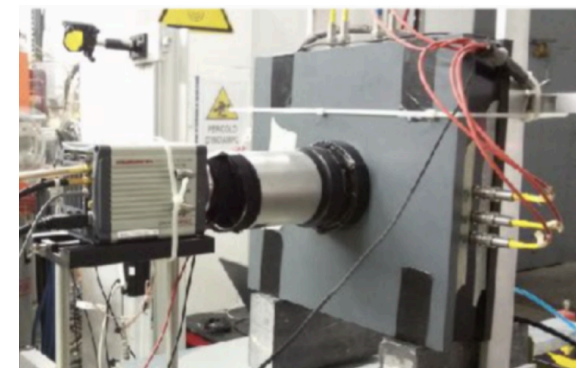
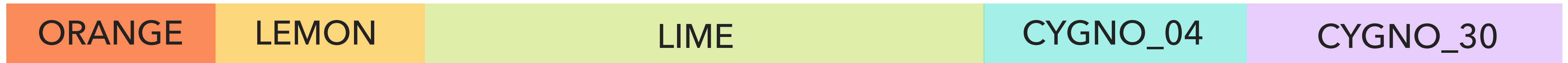
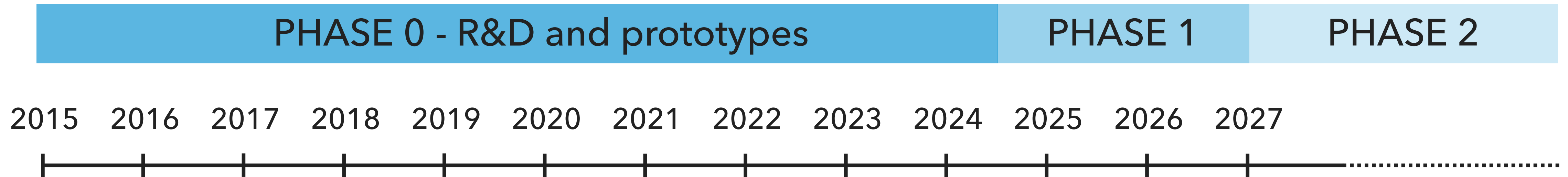
CYGNO Dark Matter exploration region



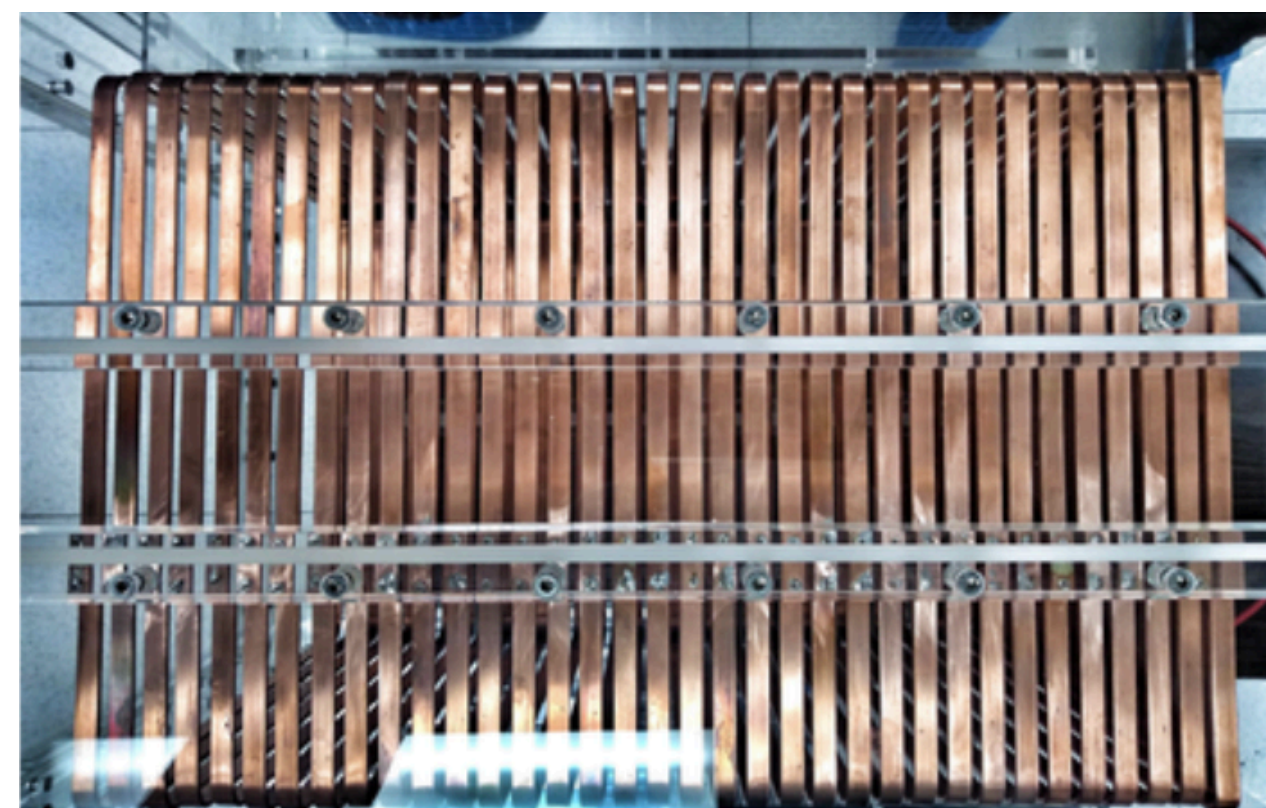
- **Light** target for low mass WIMPs
- Sensitive to **SI** couplings



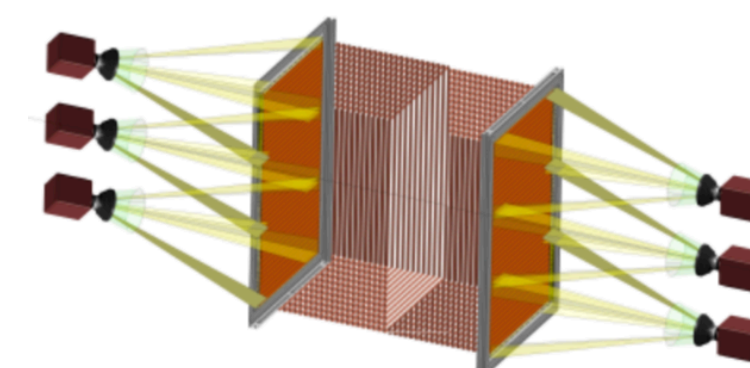
- **Heavier** target for intermediate masses
- Sensitive to **SD** couplings



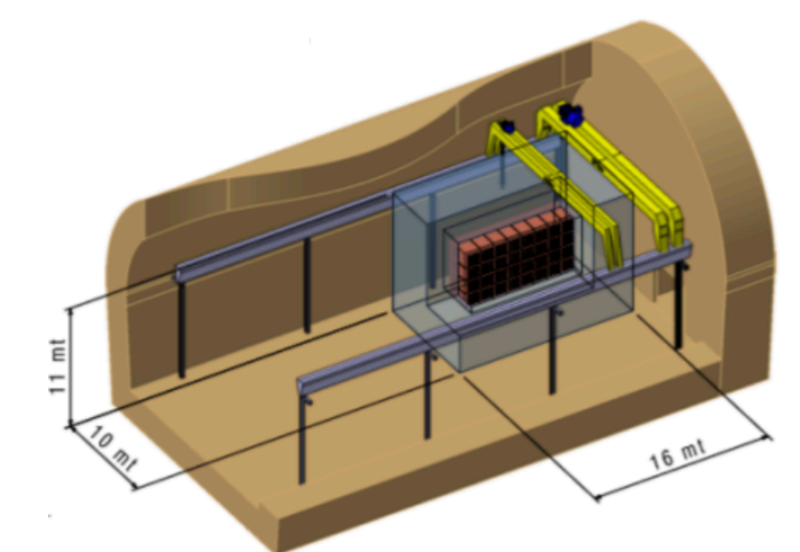
- 1 cm drift
- 3D printing
- 20 cm drift



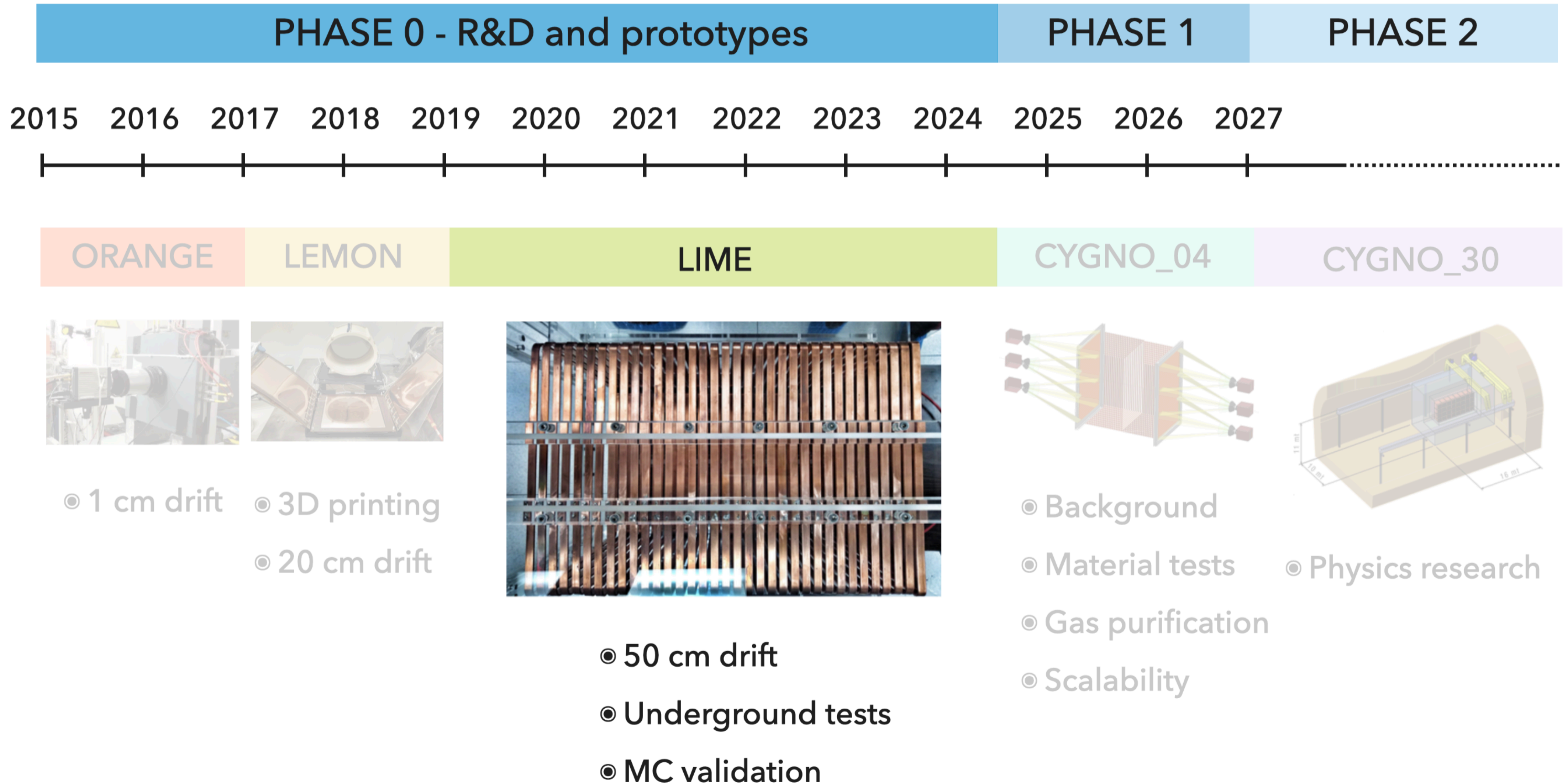
- 50 cm drift
- Underground tests
- MC validation

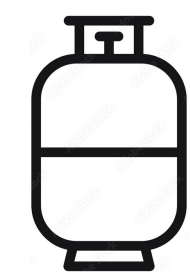
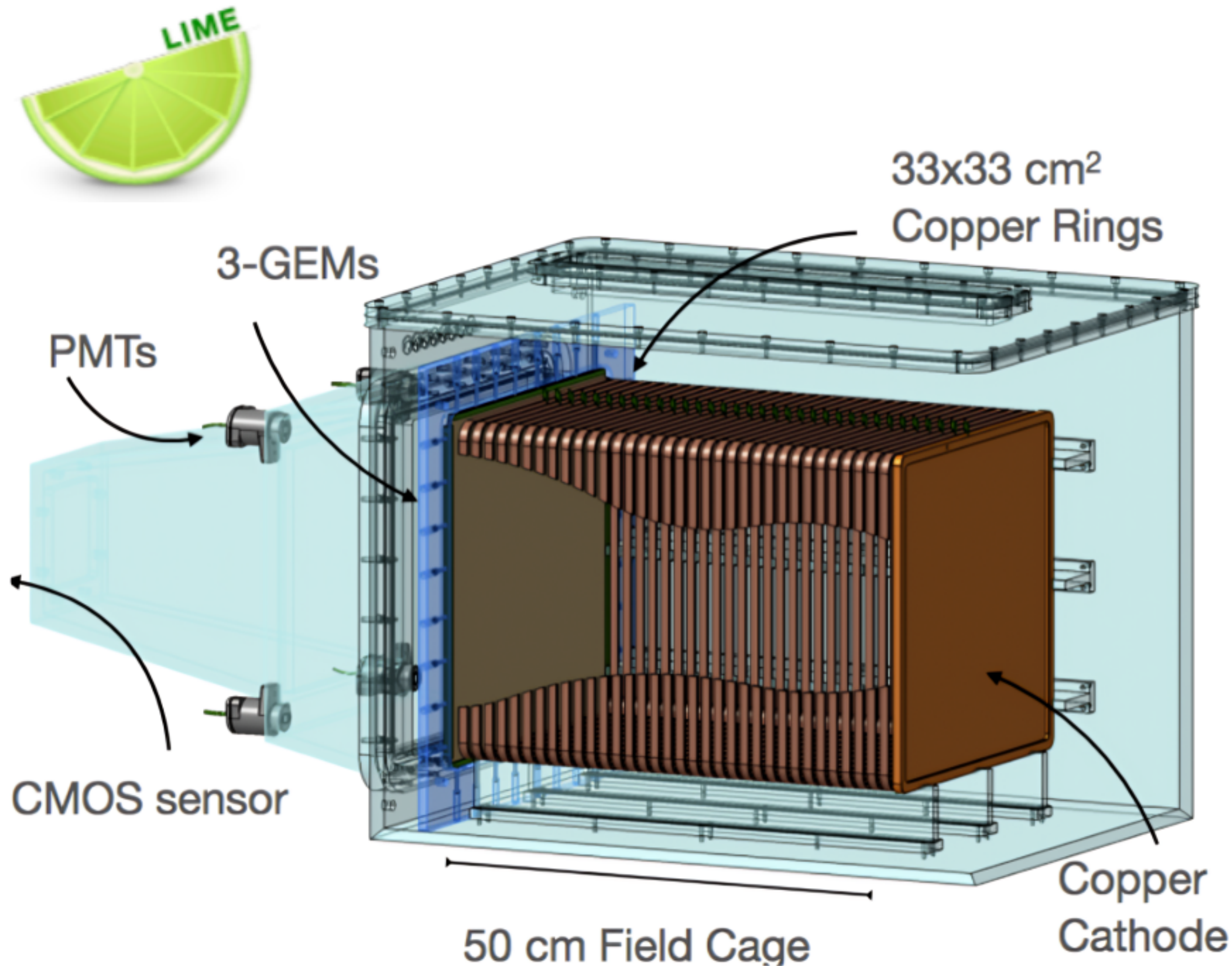


- Background
- Material tests
- Gas purification
- Scalability

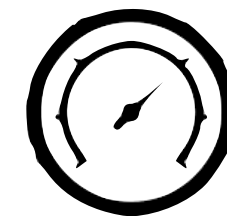


- Physics research





50 L active volume of He:CF₄ (60:40)



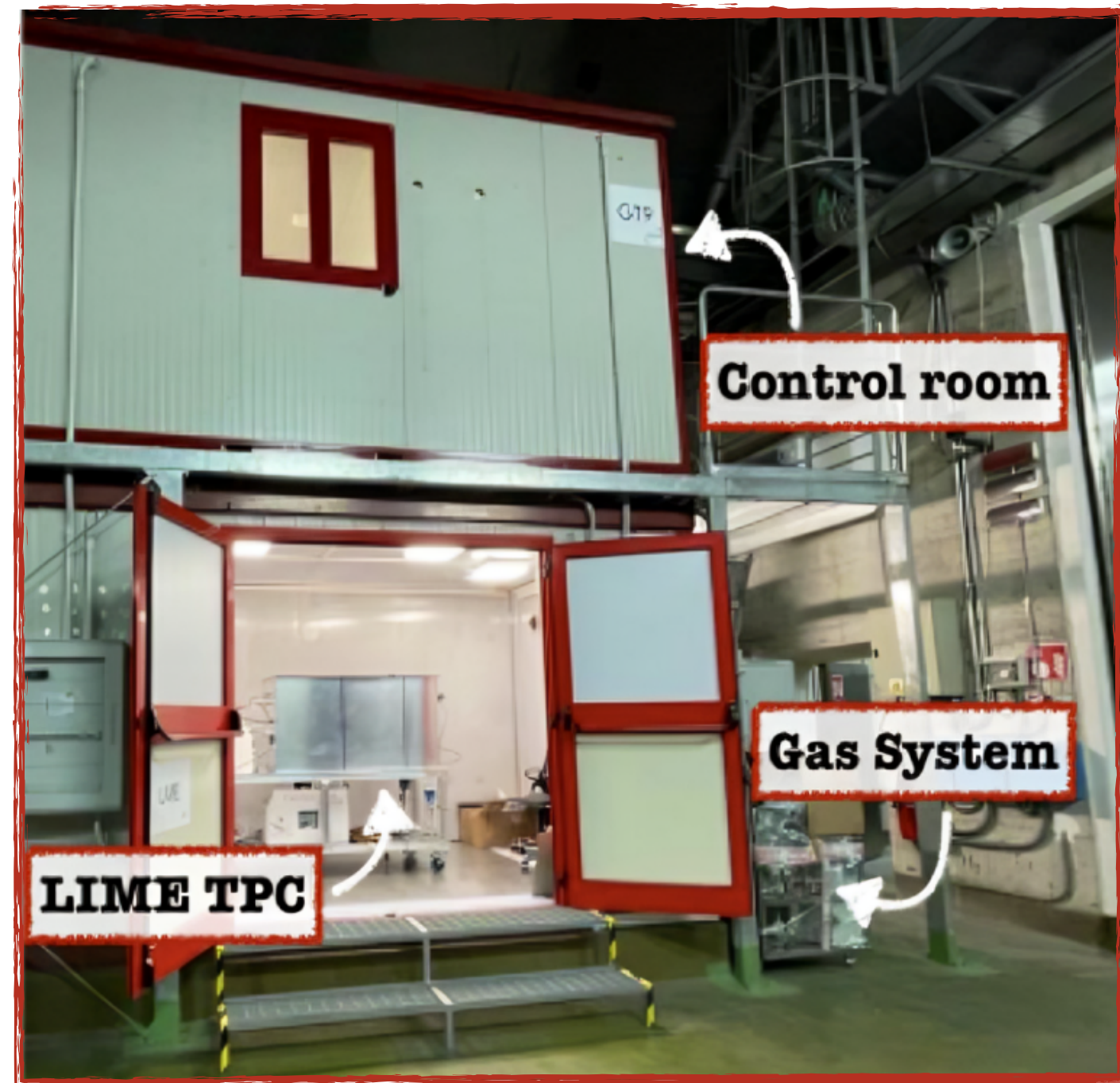
Atmospheric pressure and room temperature



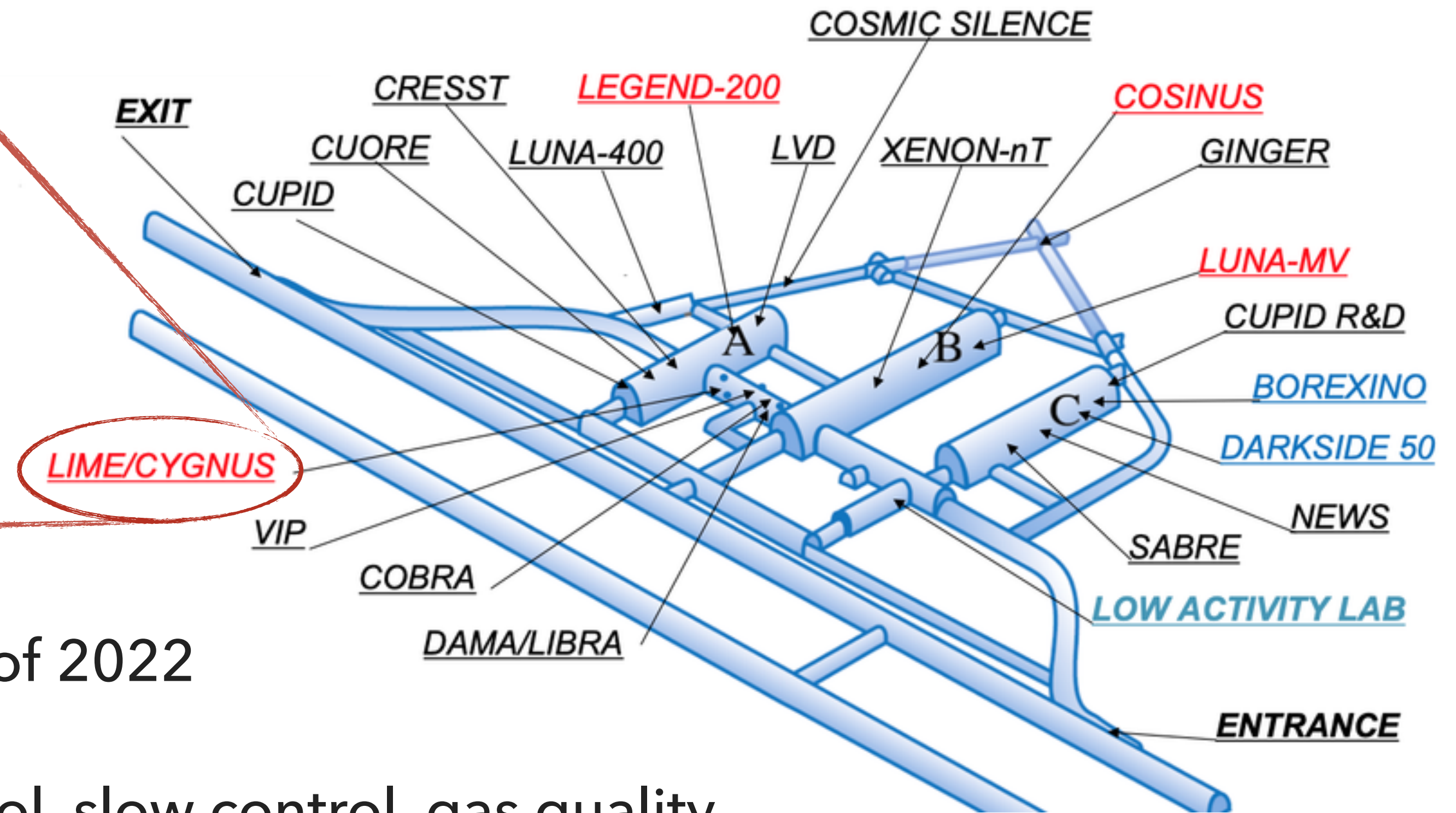
Triple 33x33 cm² GEM stack for amplification



Optical readout → 4 PMTs + 1 sCMOS camera (ORCA fusion)



- Running
- Construction/Commissioning
- Decommissioning



Placed underground at the beginning of 2022



Several initial tests: DAQ, remote control, slow control, gas quality, detector operation optimization

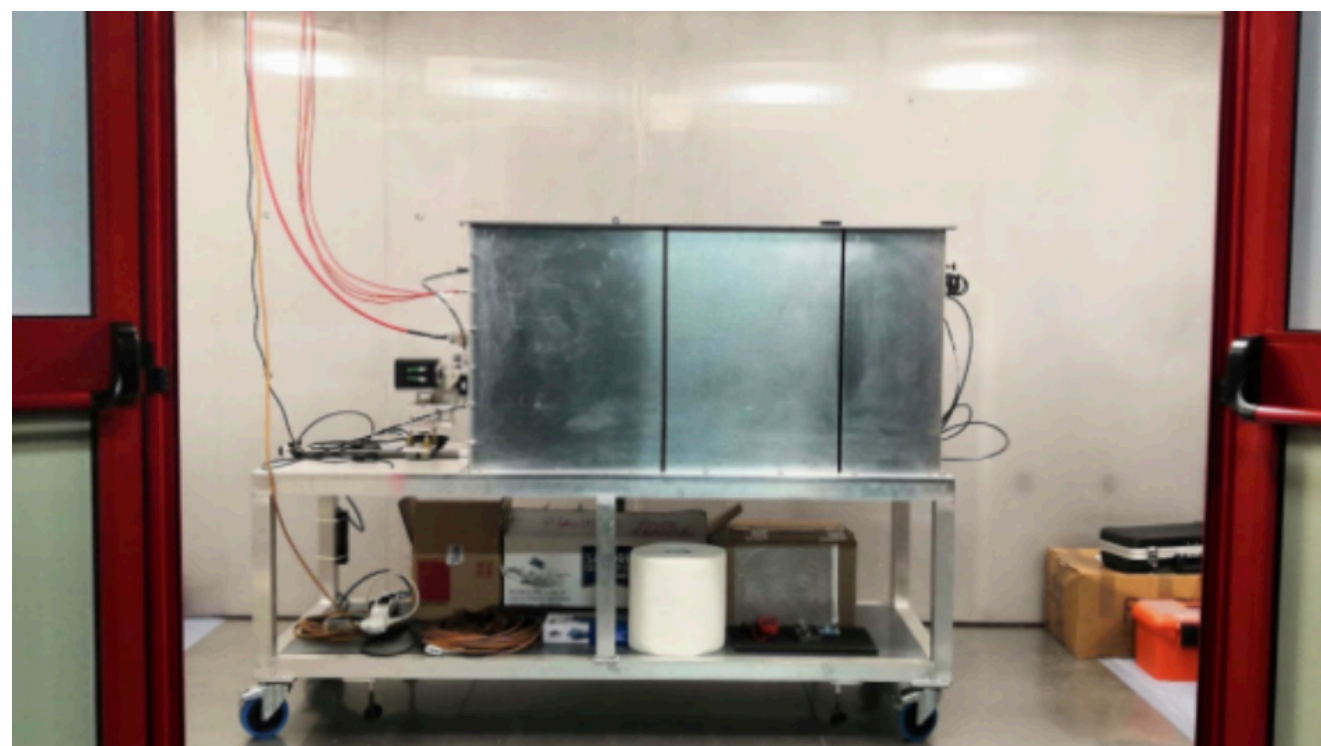


Multiple radioactive source runs: ^{55}Fe , ^{133}Ba , ^{152}Eu , AmBe

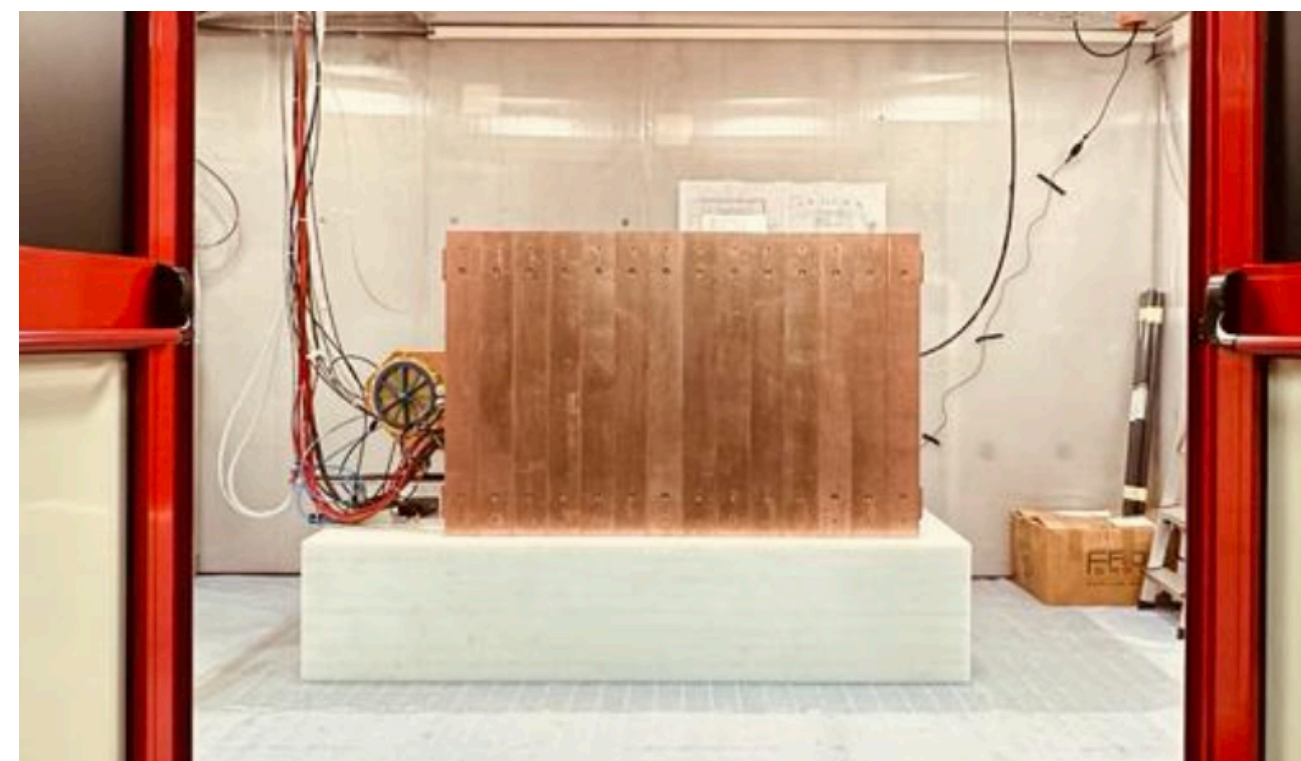
Phase	Shielding	GEM V [V]	# Pictures	Live time [s]	Rate PMTs [Hz]	Fresh gas flux [L/h]
RUN 1	None	420	2,86E+05	1,76E+05	30	10
RUN 2	4 cm Cu	440	2,98E+05	1,91E+05	3.5	20
RUN 3	10 cm Cu	440	1,72E+05	1,91E+05	1.6	20
RUN 4	+ 40 cm H ₂ O			Under analysis		

RUN 1

Background dominated by the external sources

***RUN 2-3***

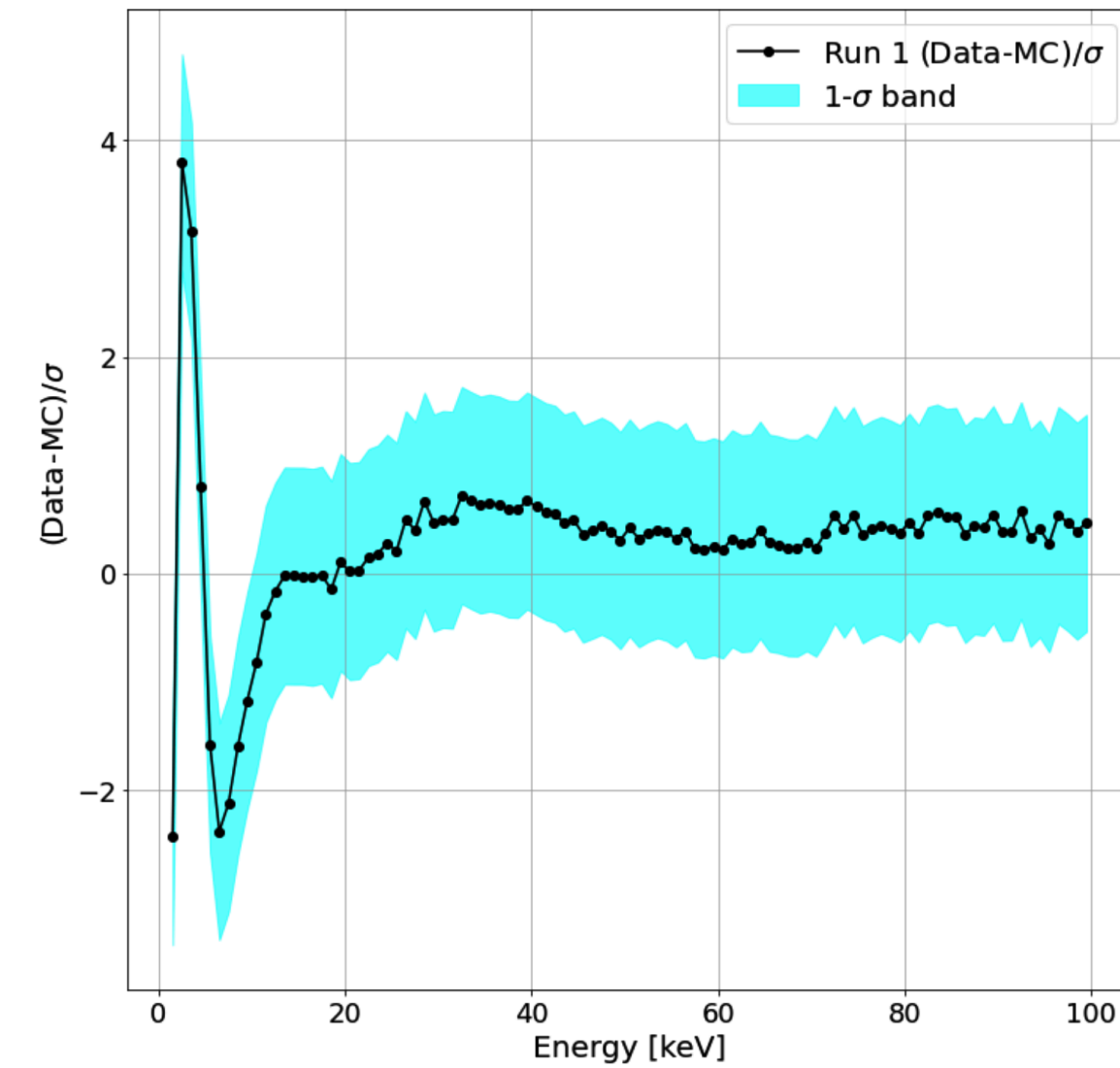
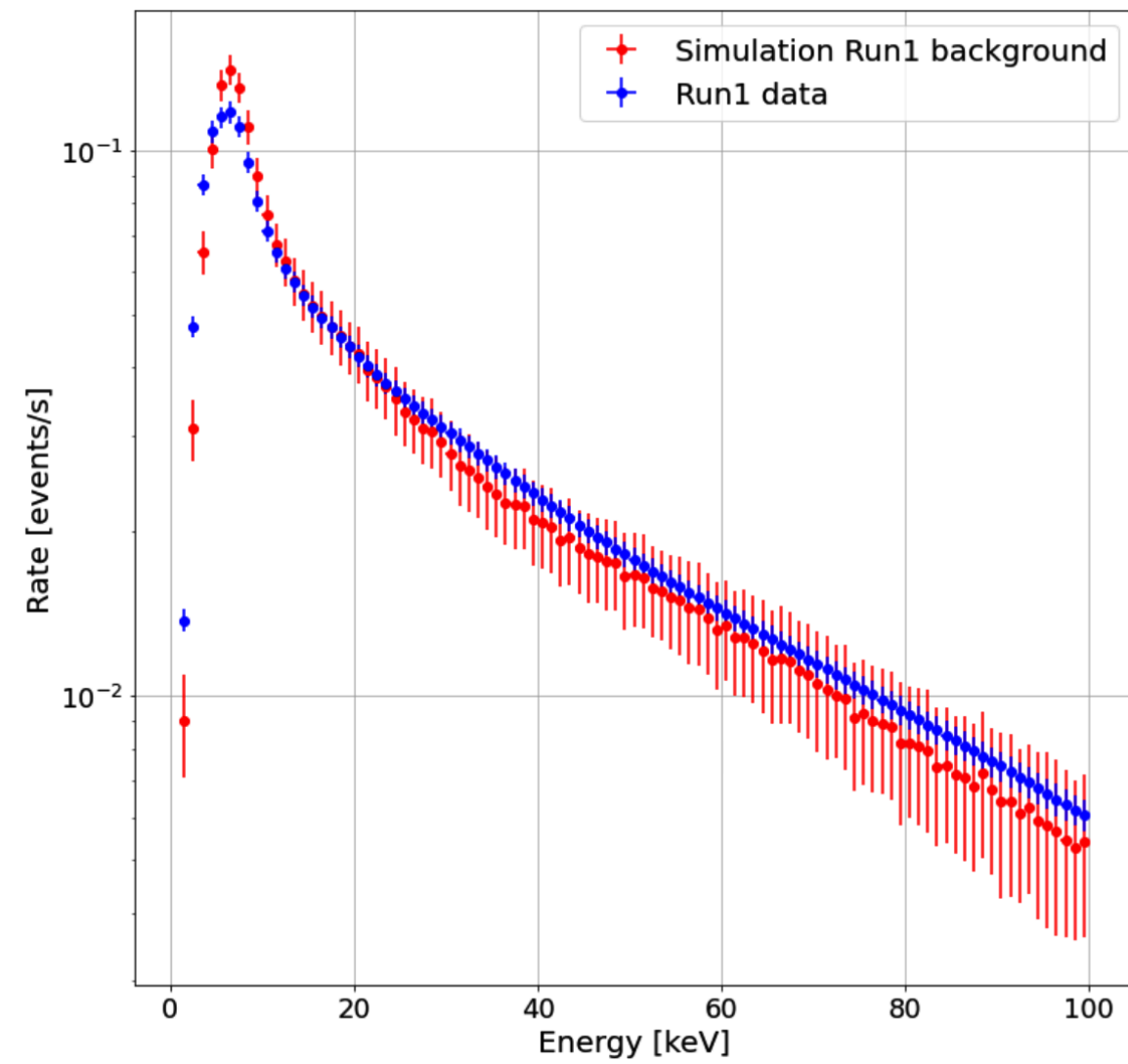
External background suppressed by a factor ~40-600

***RUN 4***

Residual external neutron background suppression

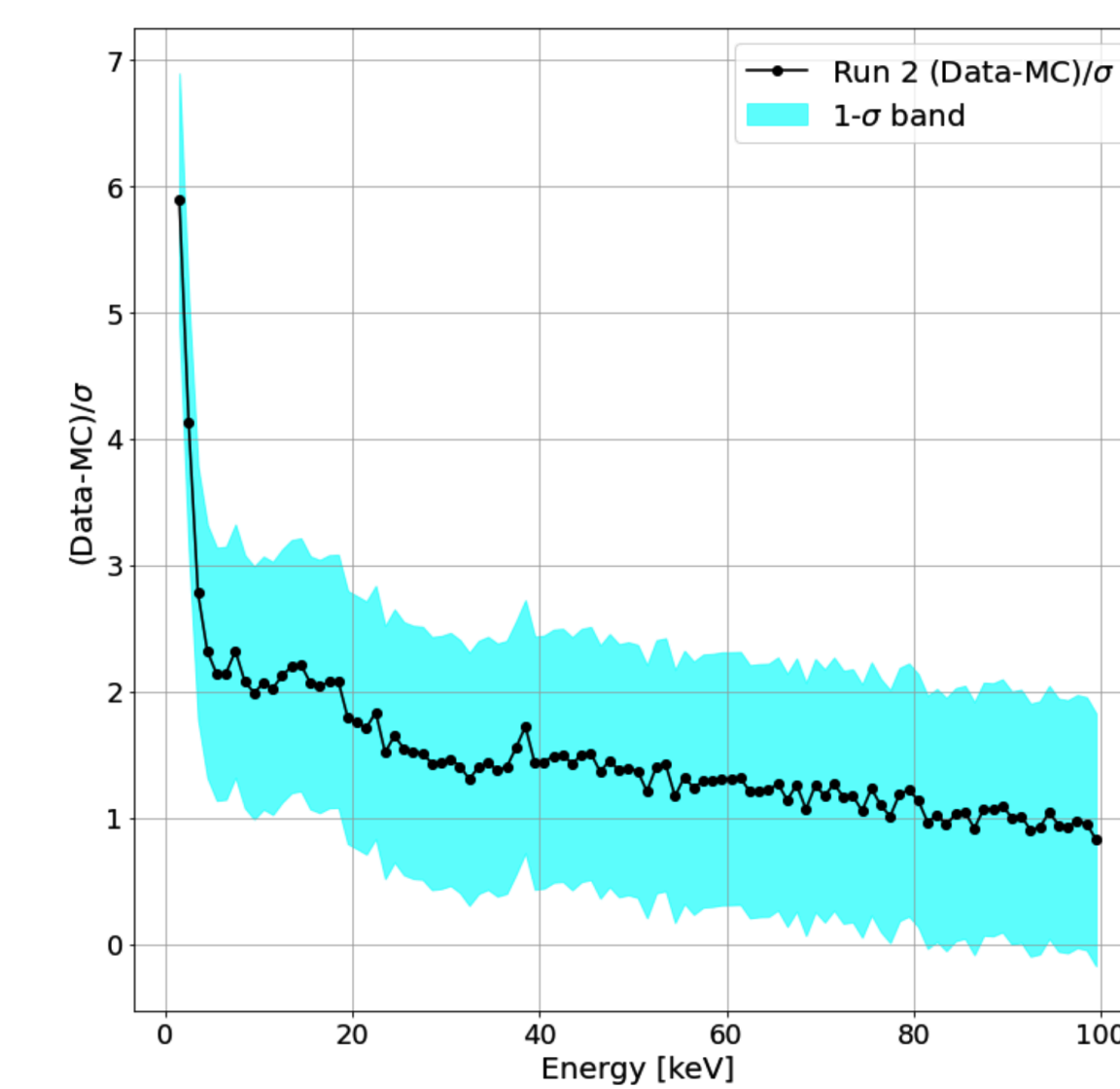
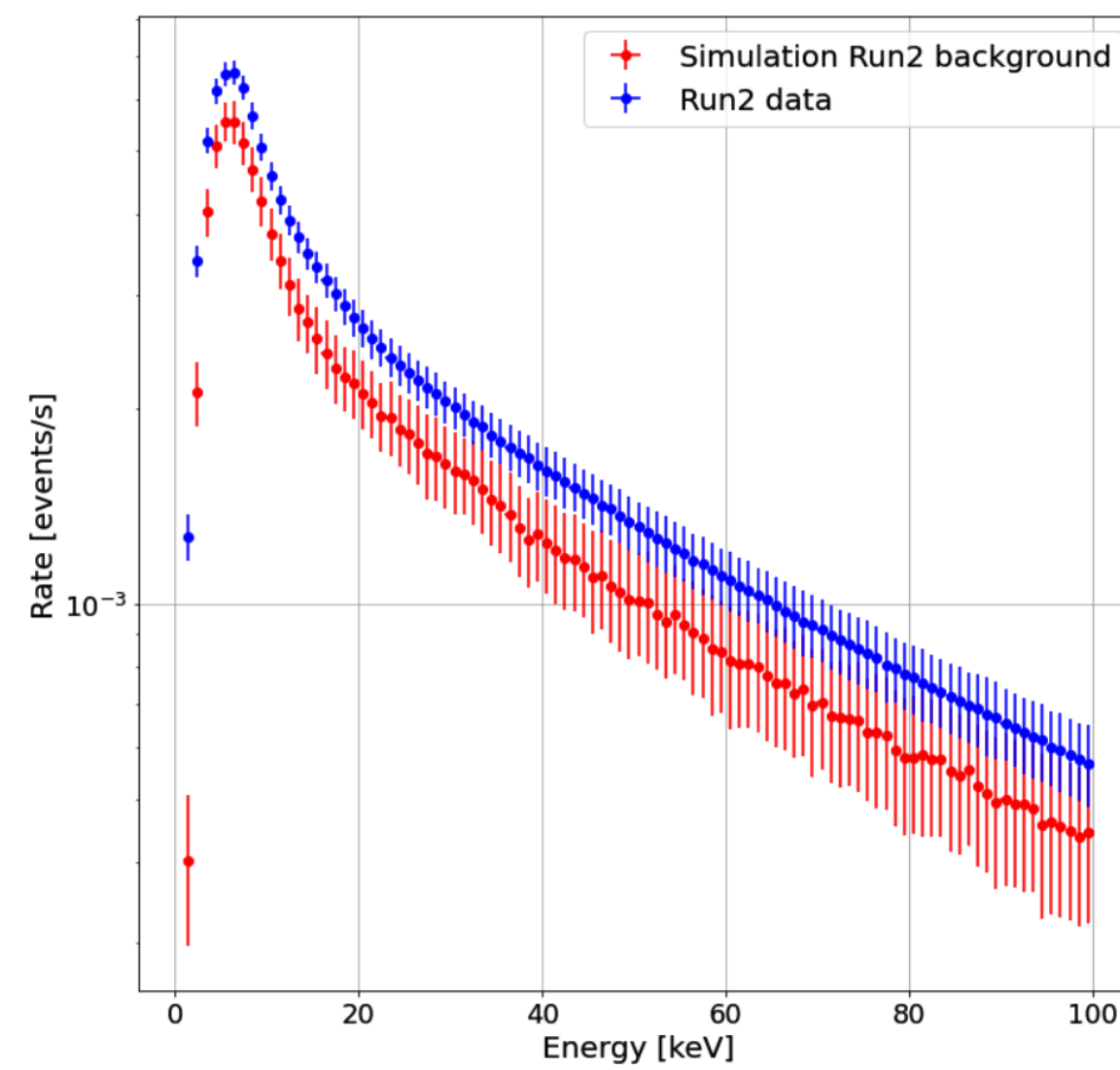


R
U
N
1



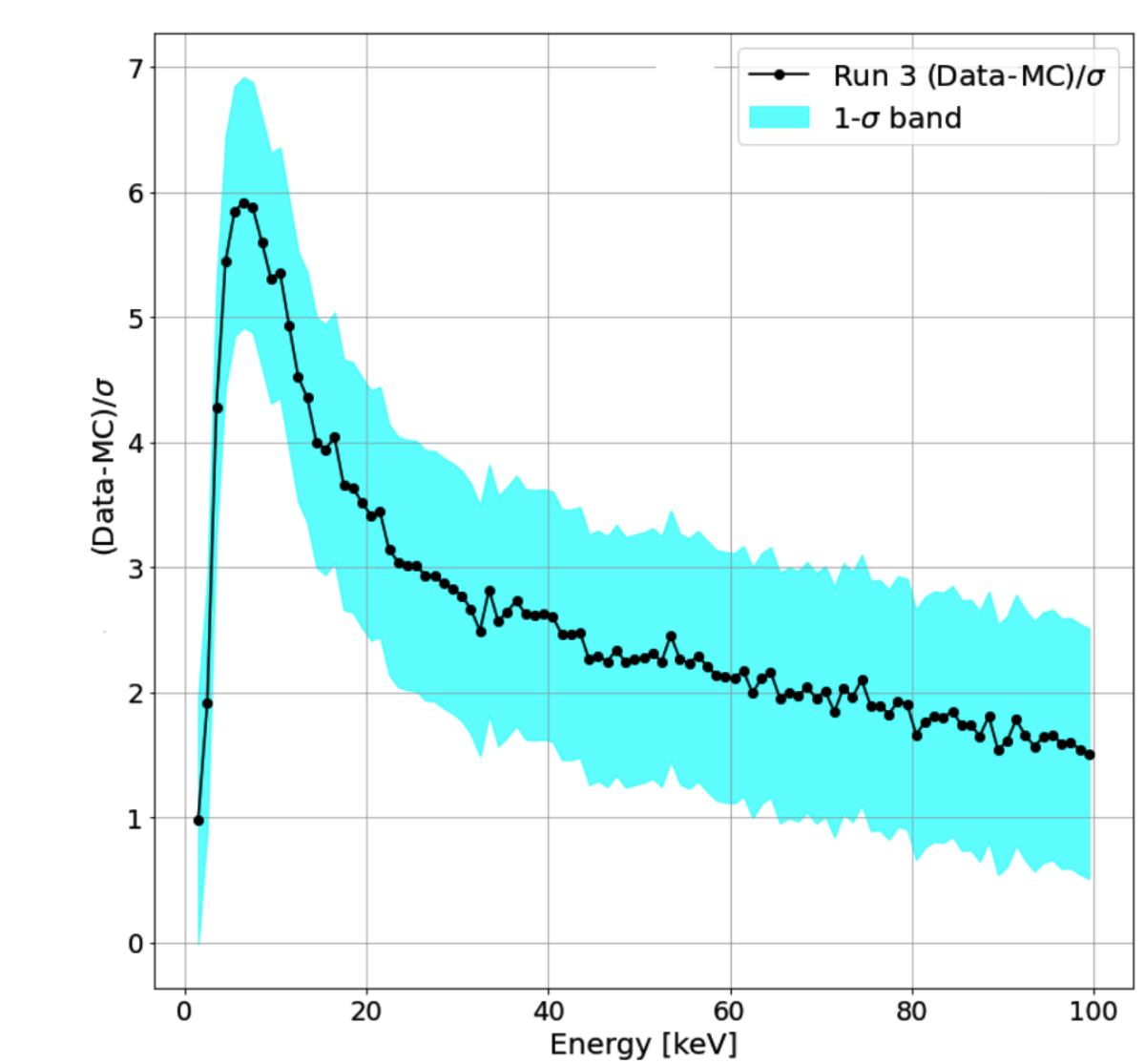
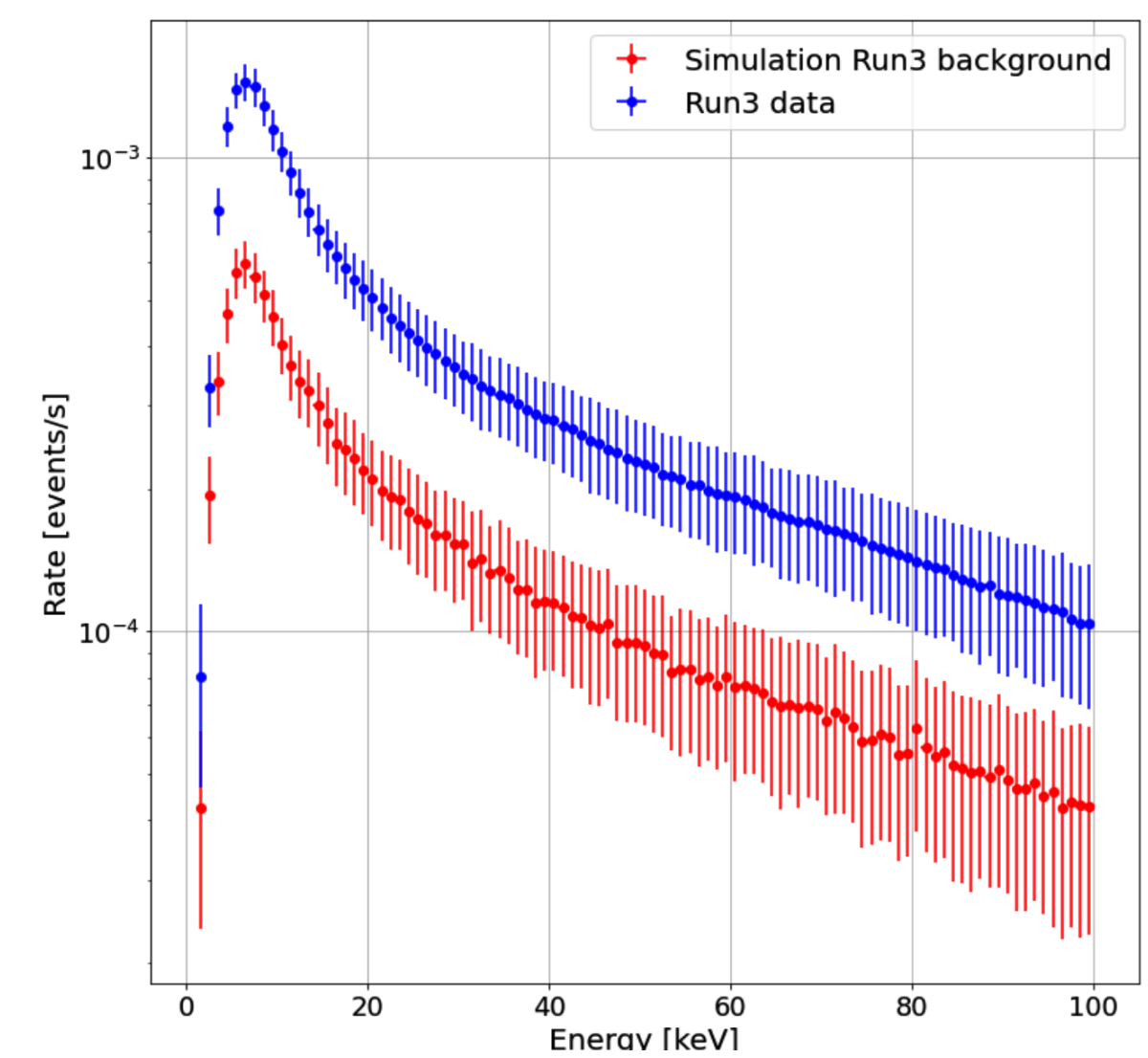
Agreement → capability of simulating external background

R
U
N
2



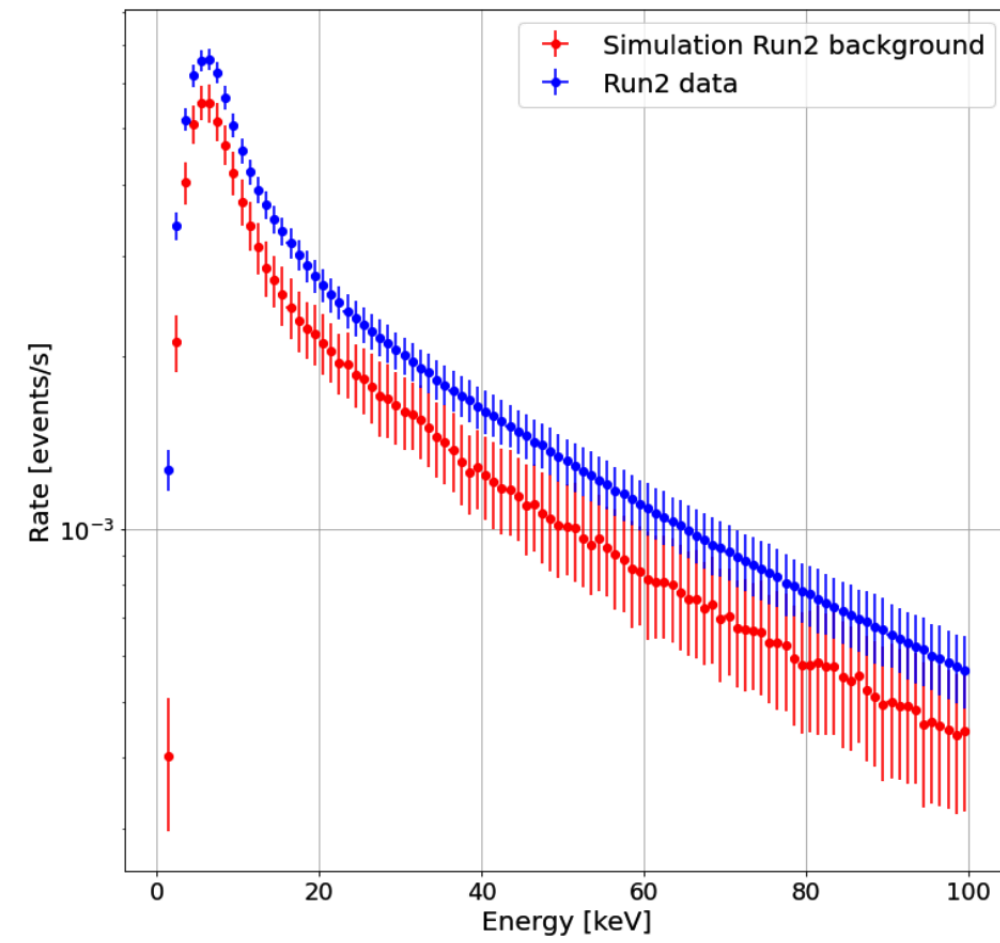
Difference ±22% → internal component missing

R
U
N
3



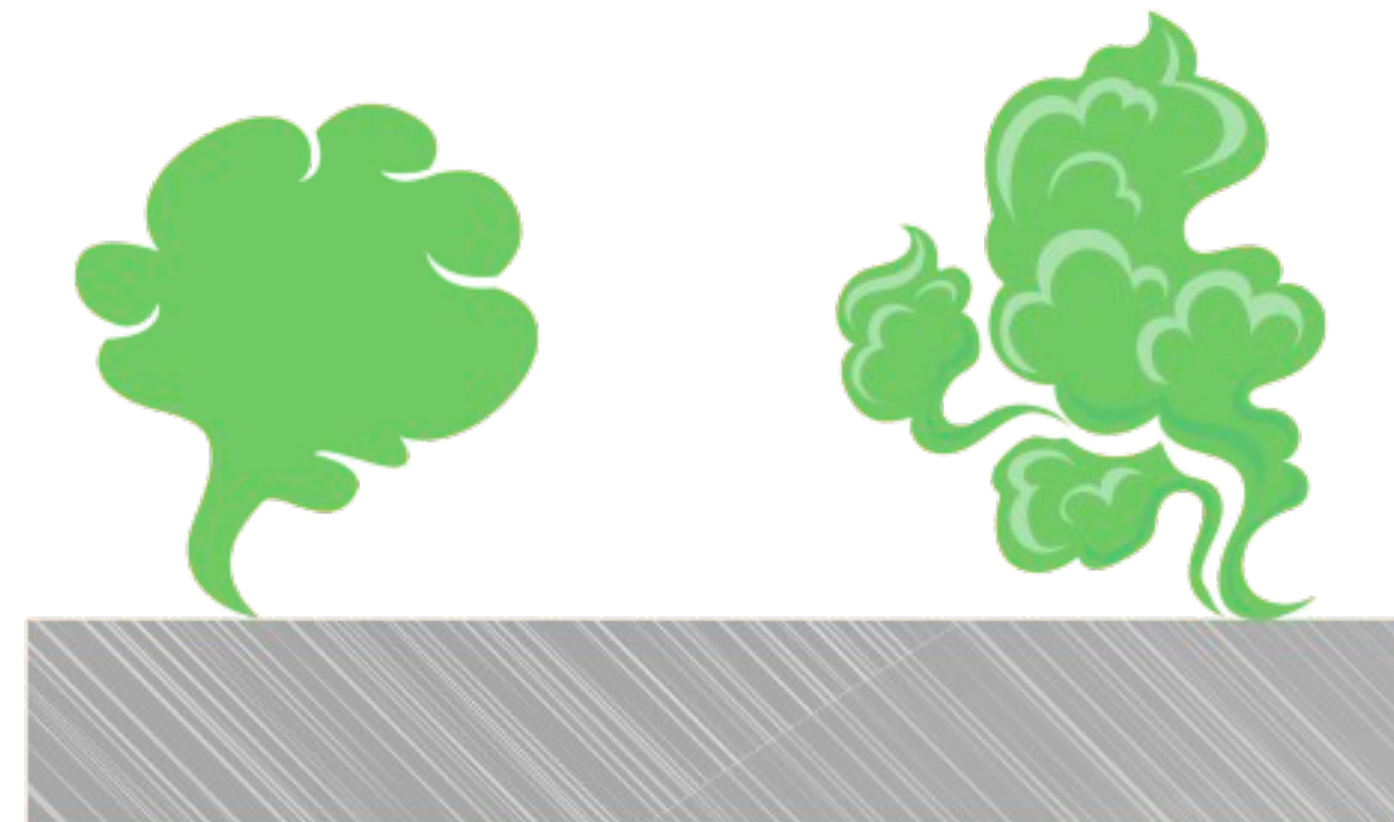
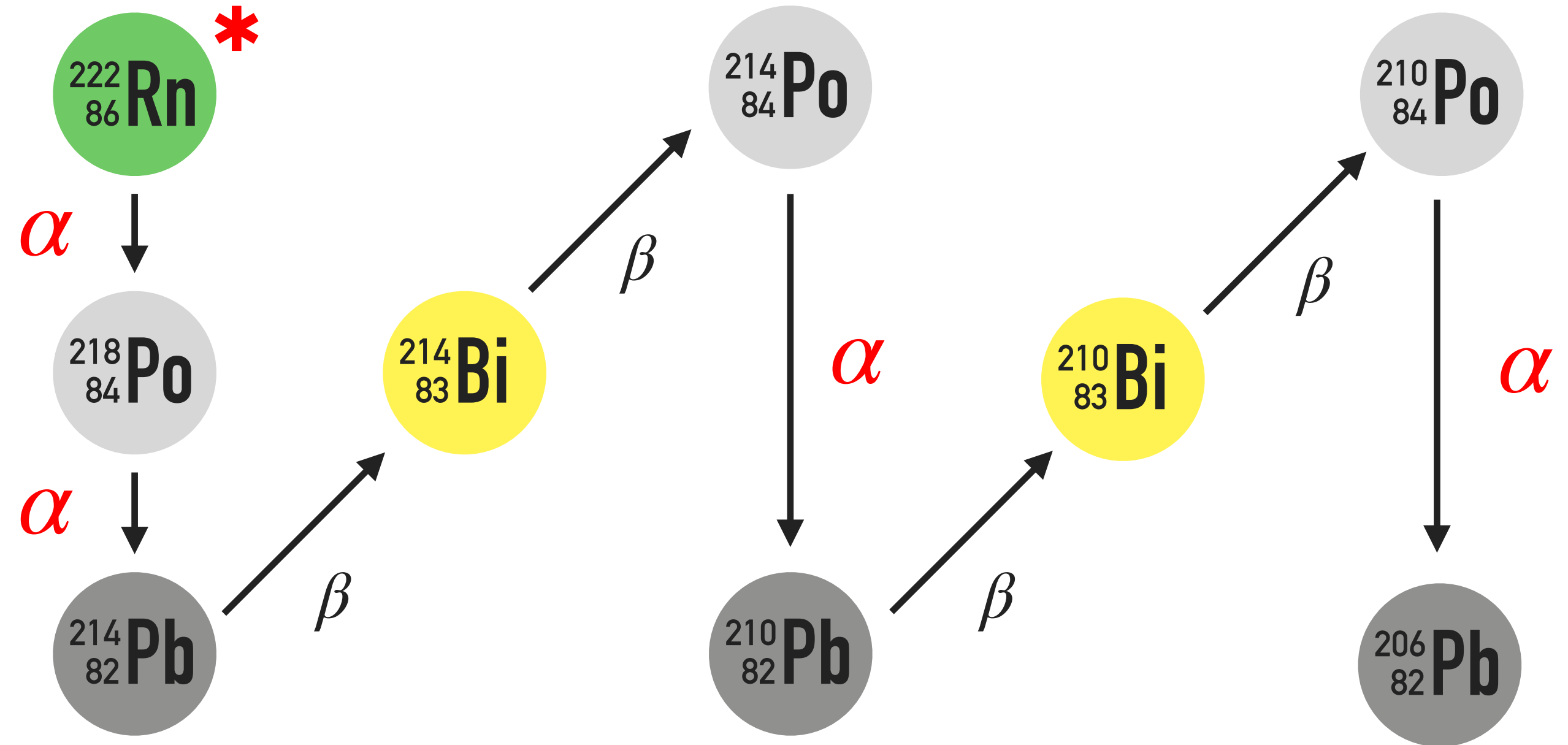
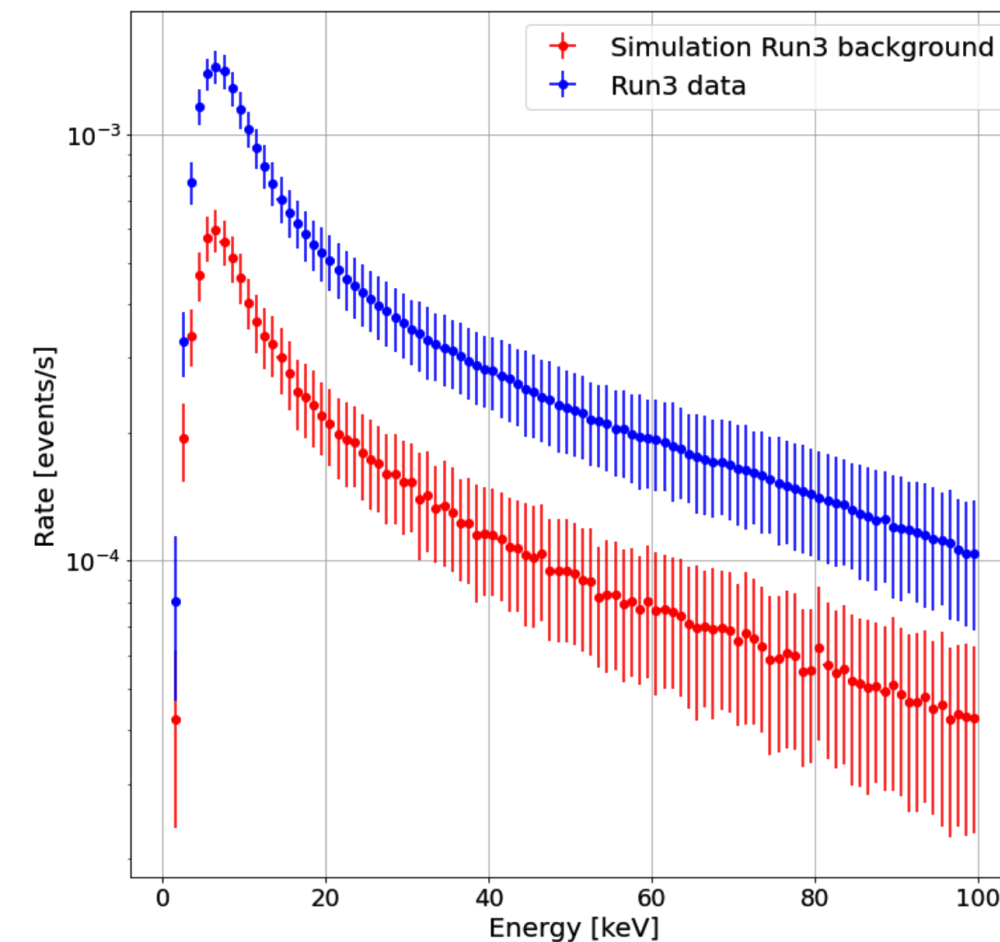
Difference ±60% → internal component missing

R
U
N
2

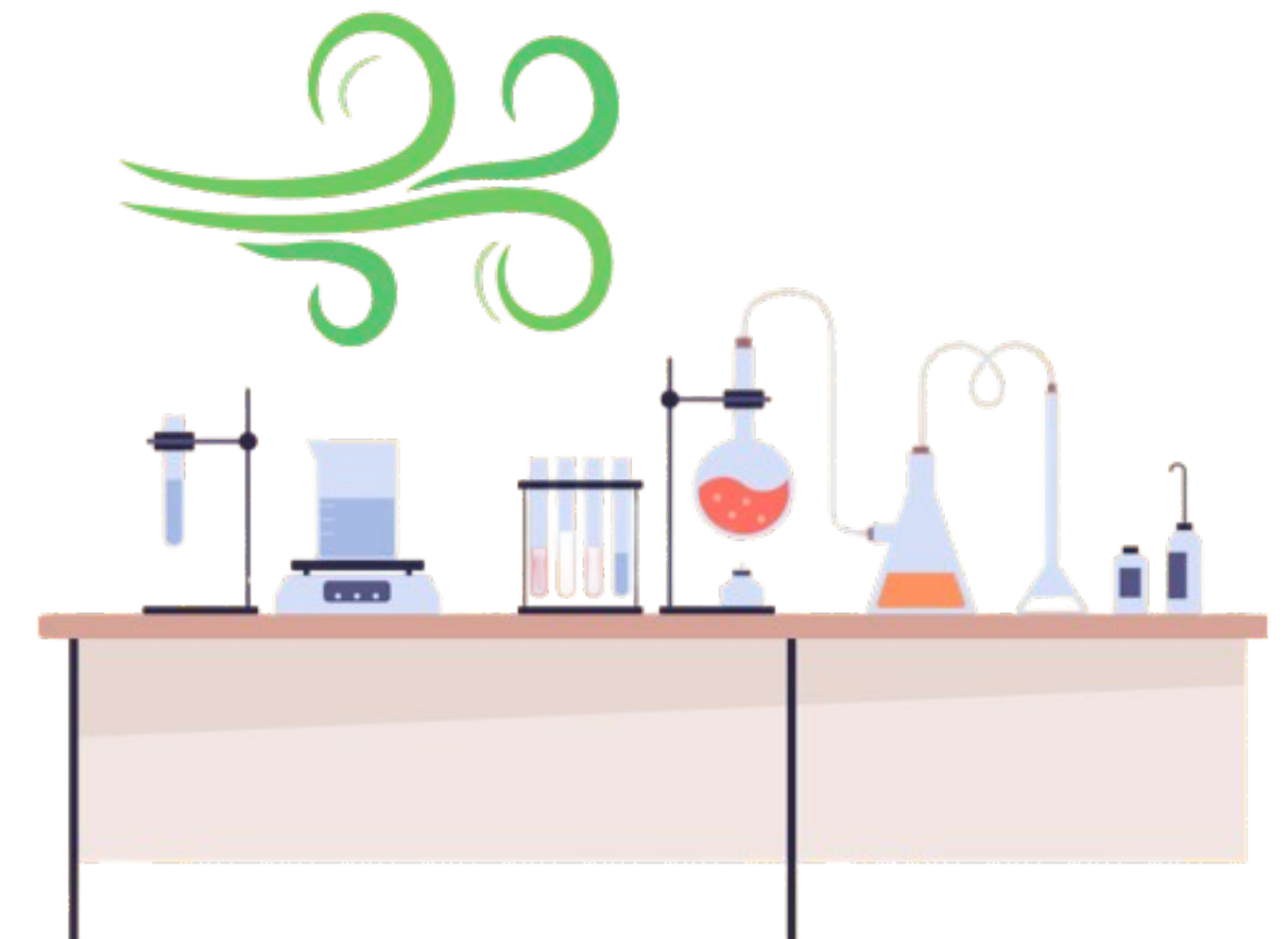


A deeper analysis suggests *Radon contamination* could explain the discrepancy

R
U
N
3

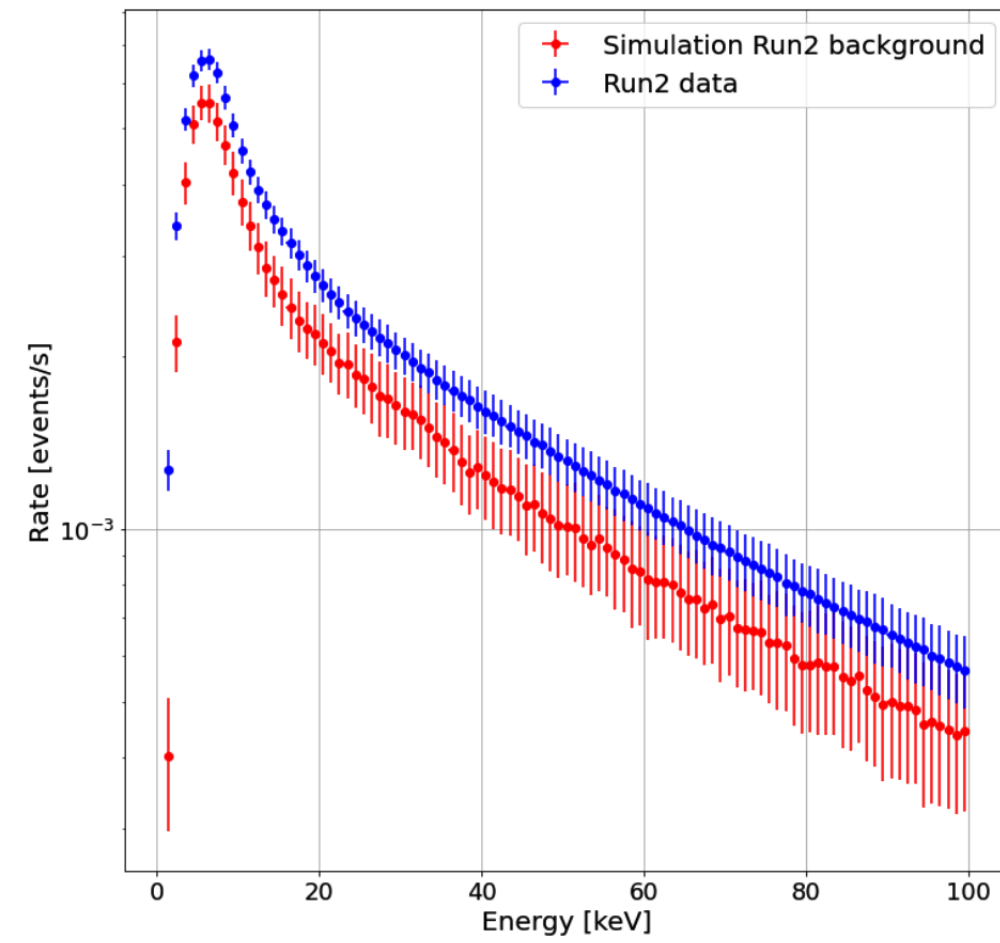


* RADON EMANATION



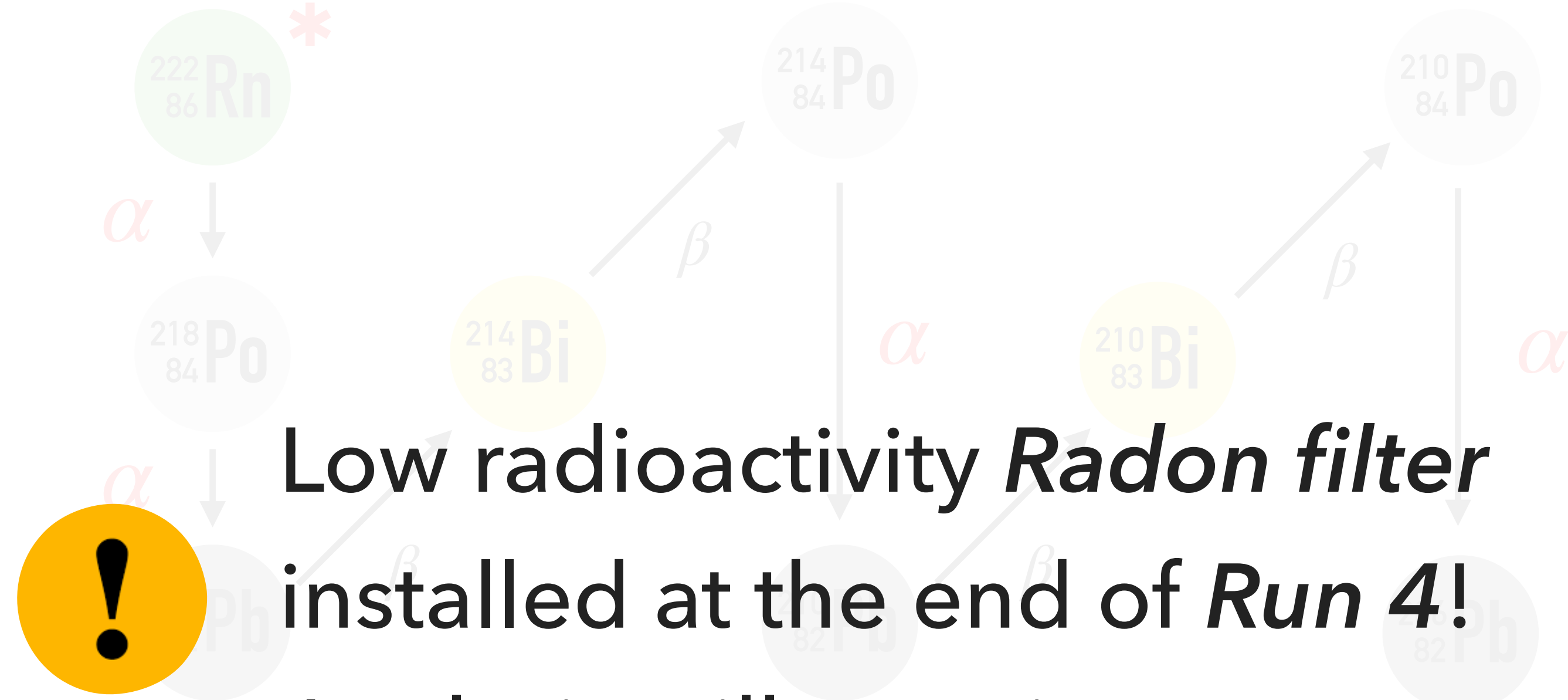
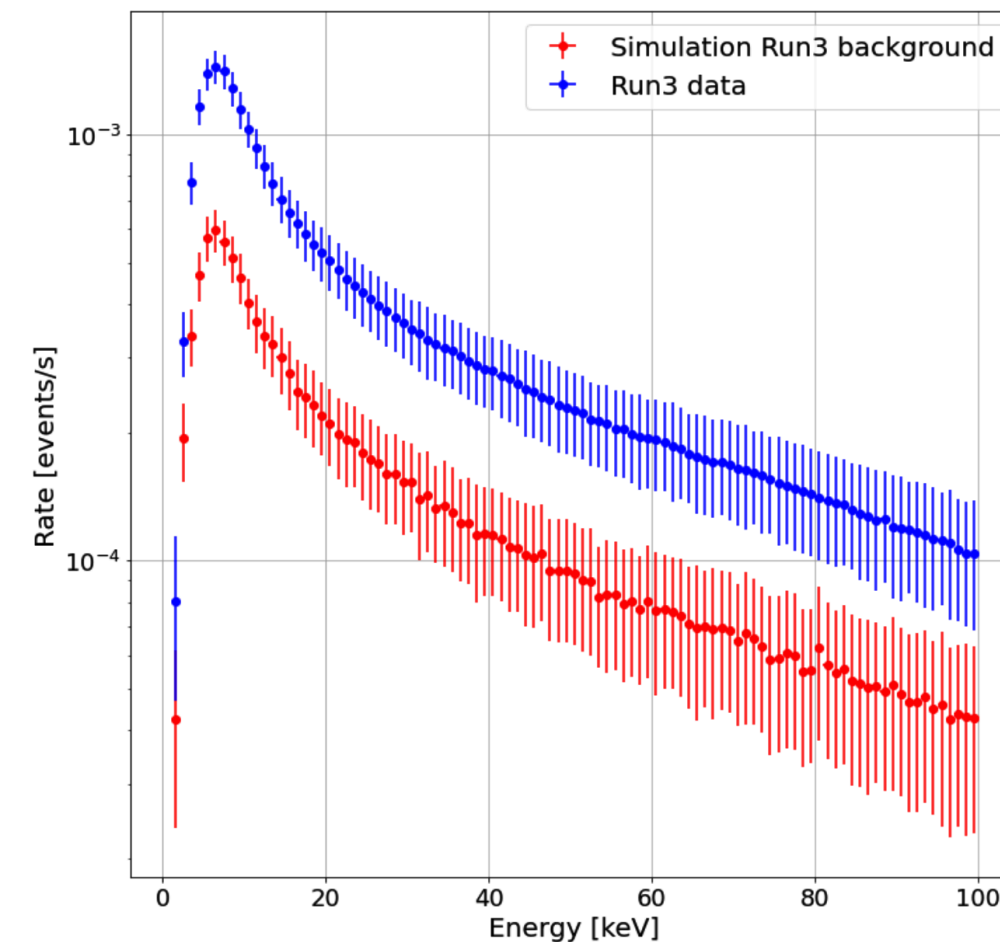
* UNDERGROUND LAB AIR

R
U
N
2



A deeper analysis suggests *Radon contamination* could explain the discrepancy

R
U
N
3



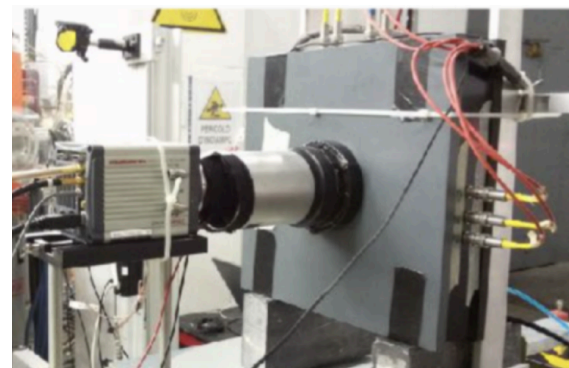
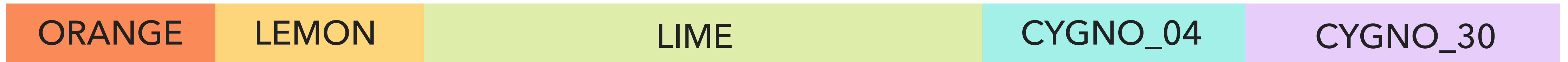
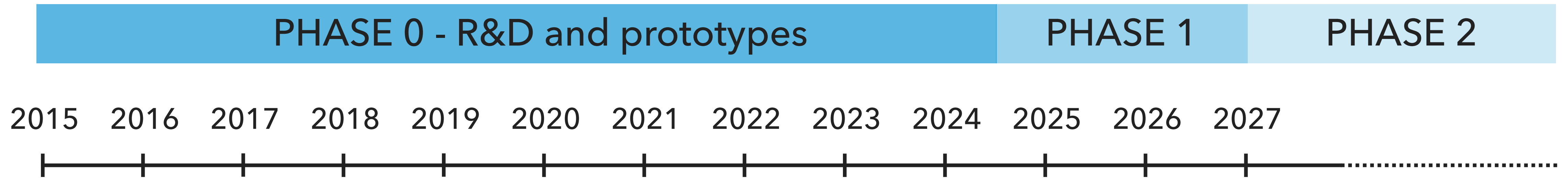
Low radioactivity *Radon filter* installed at the end of *Run 4*! Analysis still ongoing



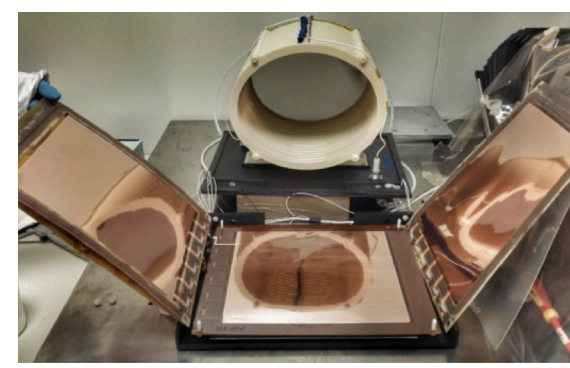
* RADON EMANATION



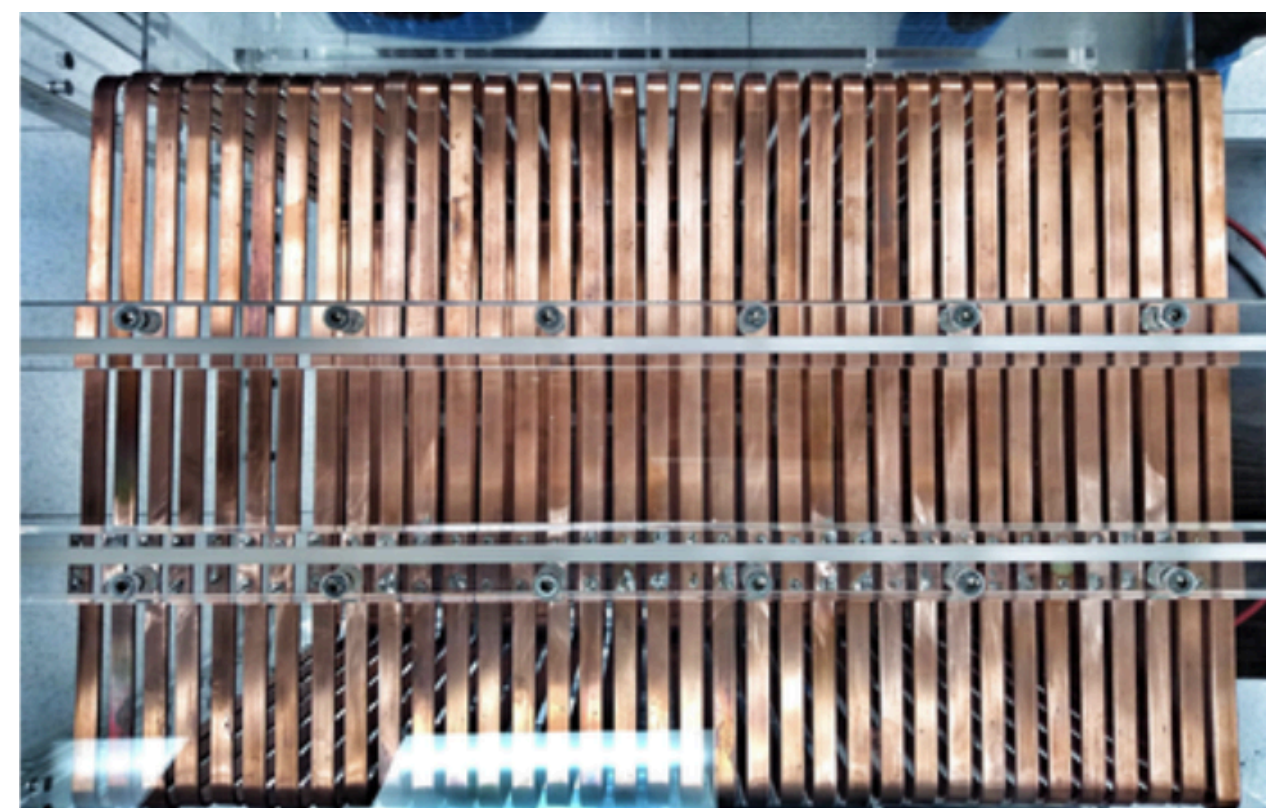
* UNDERGROUND LAB AIR



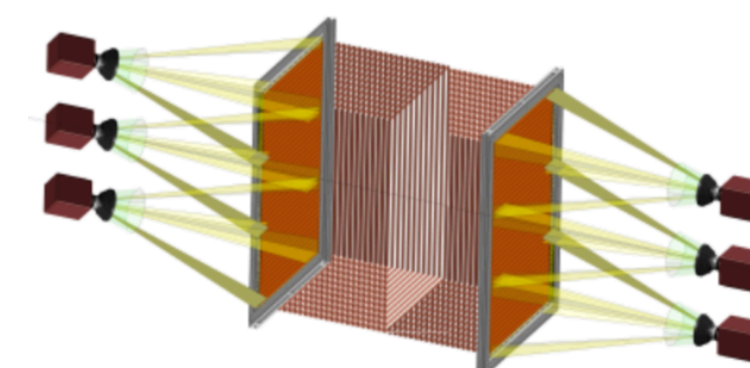
- 1 cm drift



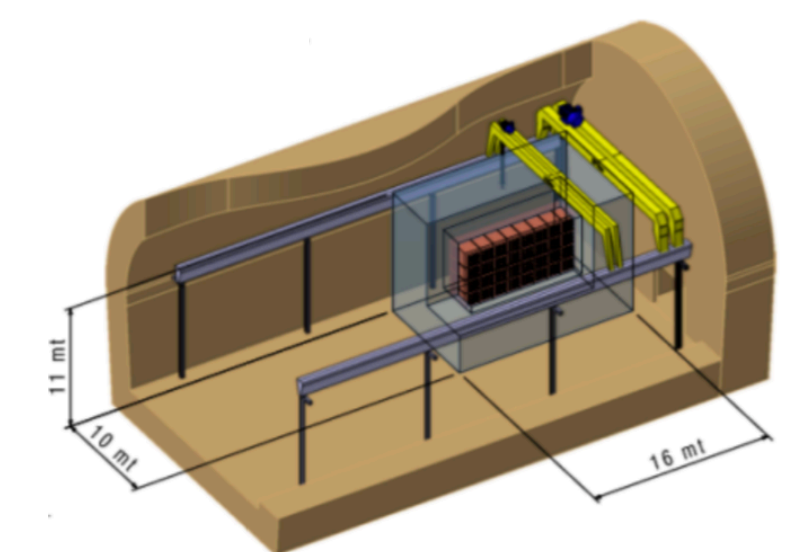
- 3D printing
- 20 cm drift



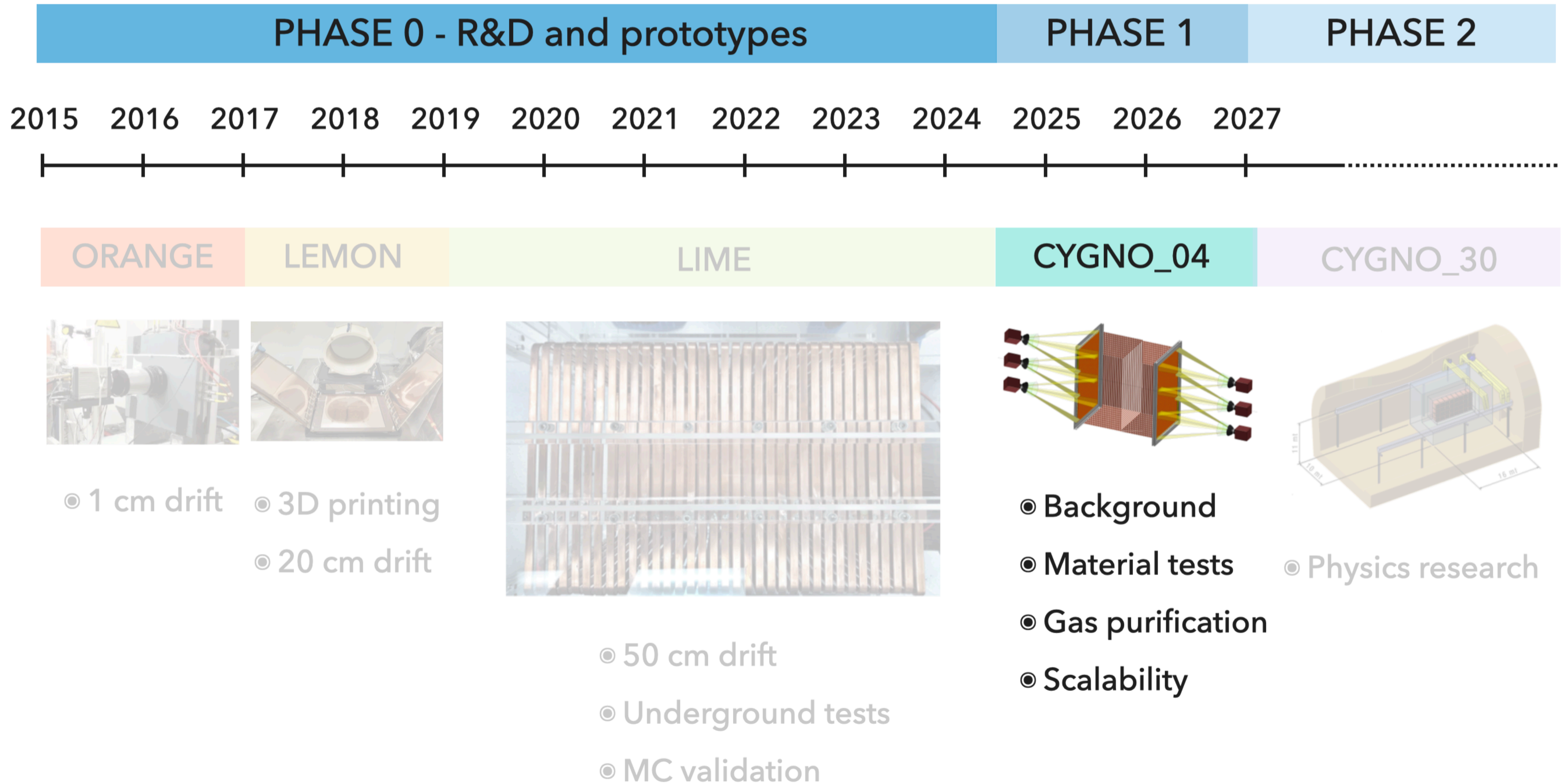
- 50 cm drift
- Underground tests
- MC validation

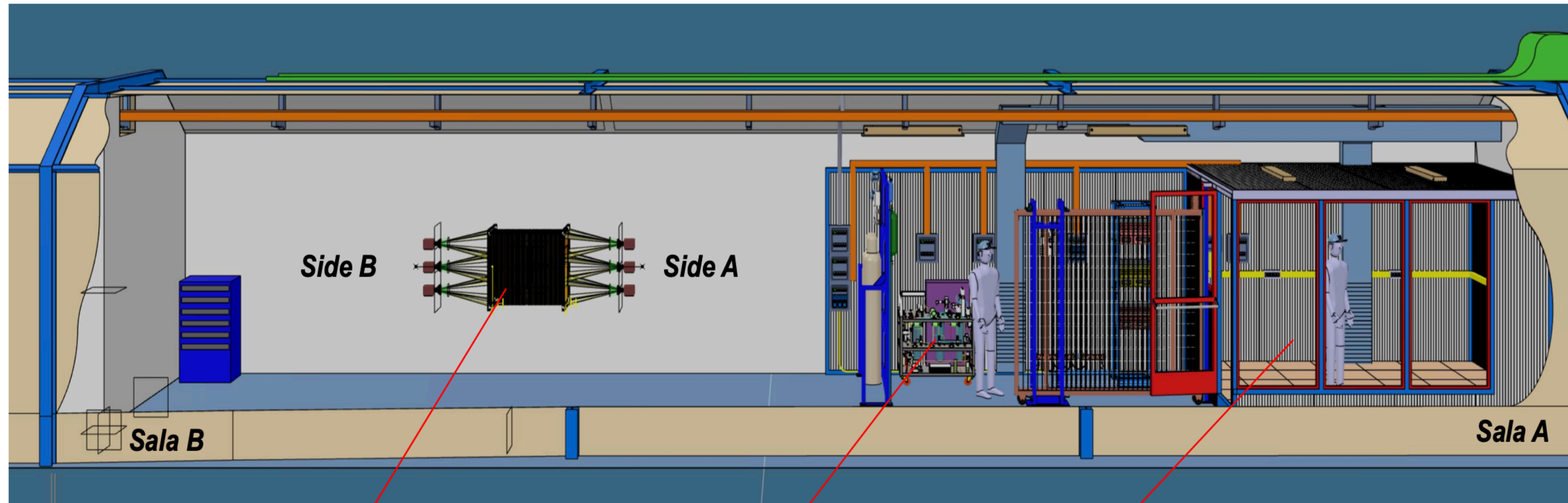


- Background
- Material tests
- Gas purification
- Scalability



- Physics research

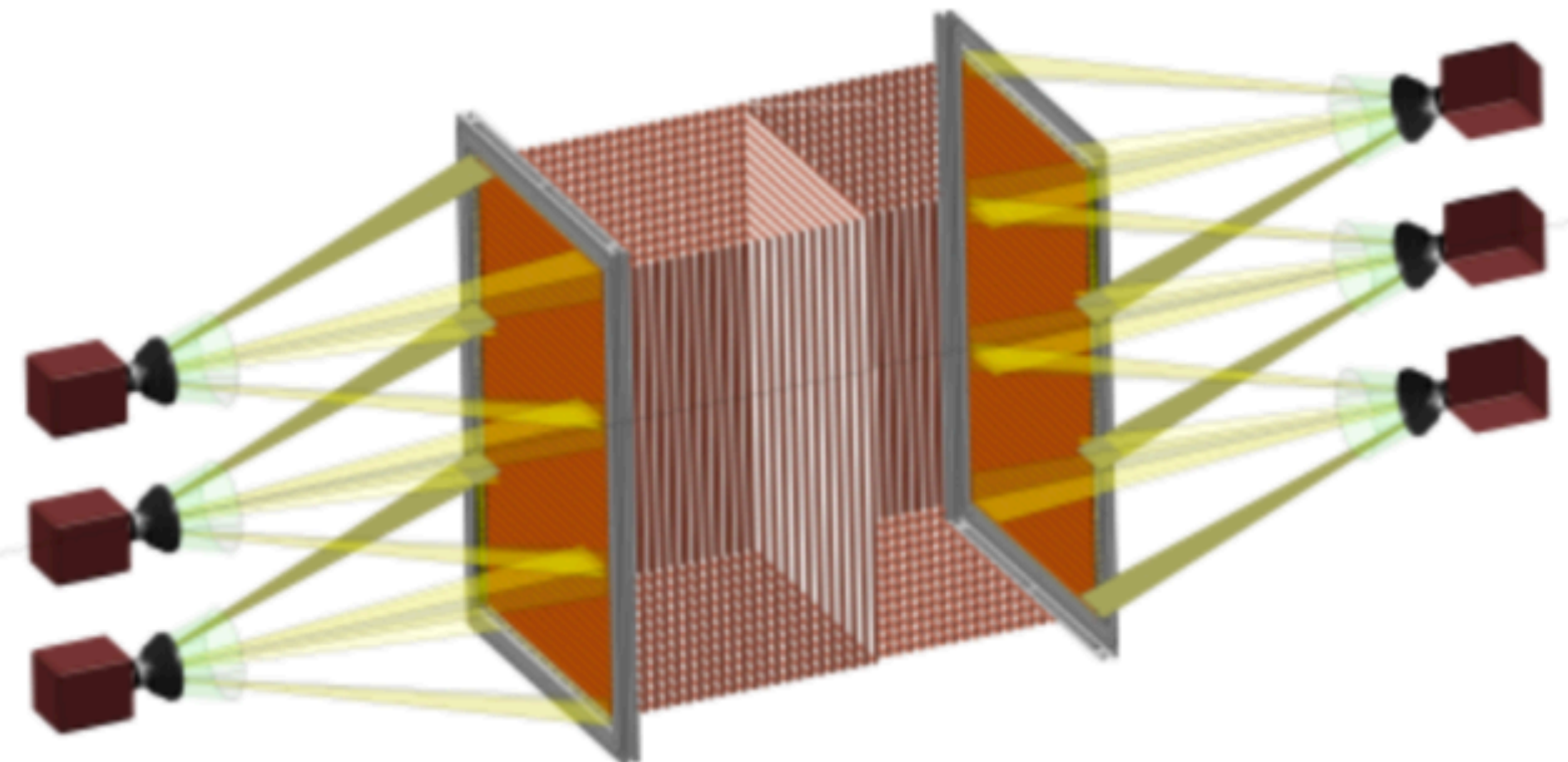




CYGNO_04

SERVICE AREA

CONTROL ROOM



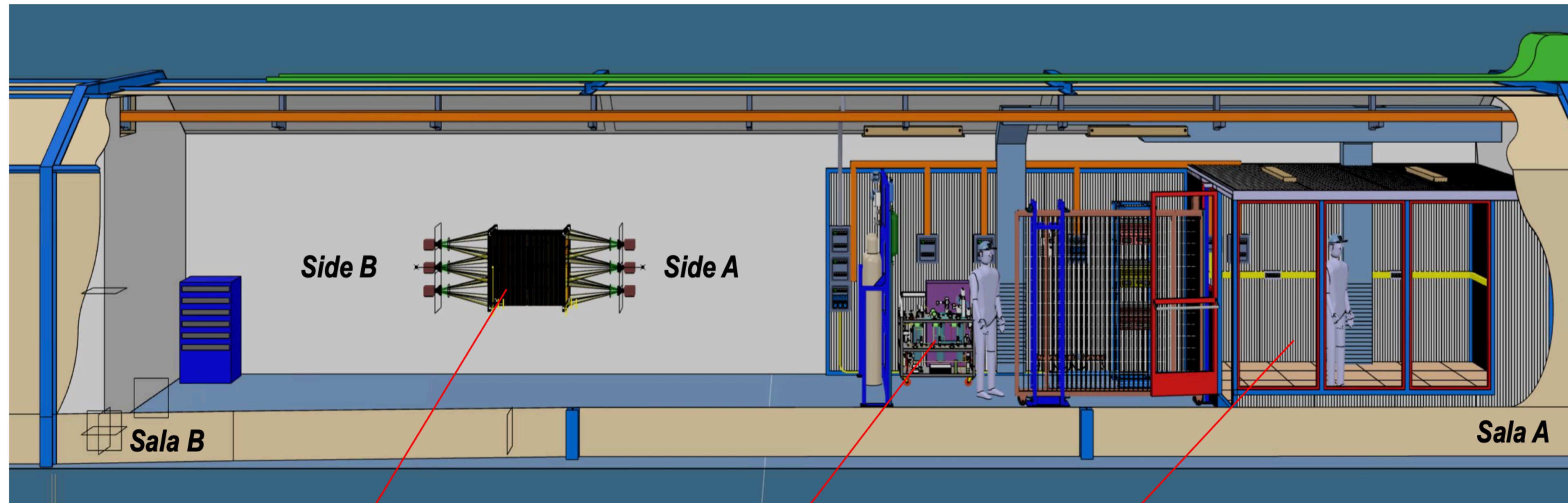
Back-to-back 0.4 m³ TPC, with central cathode



Triple 50 um GEMs amplification per side



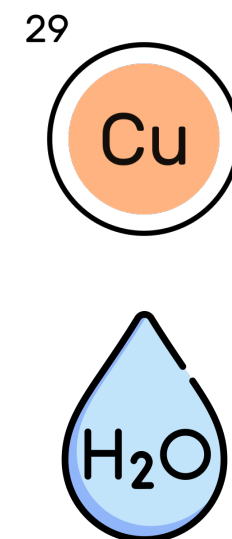
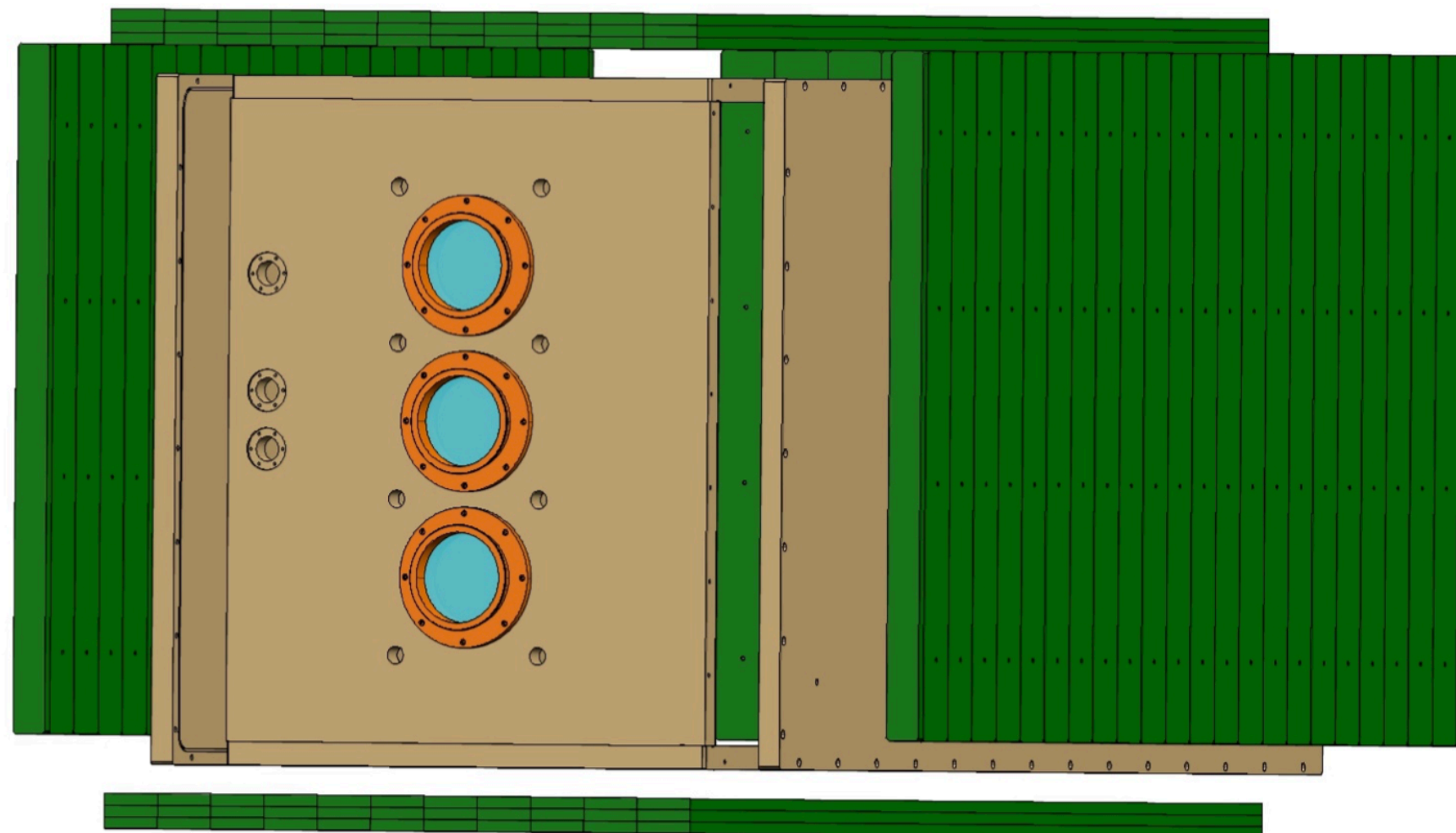
3 sCMOS ORCA Quest per side + PMTs (# TBD)



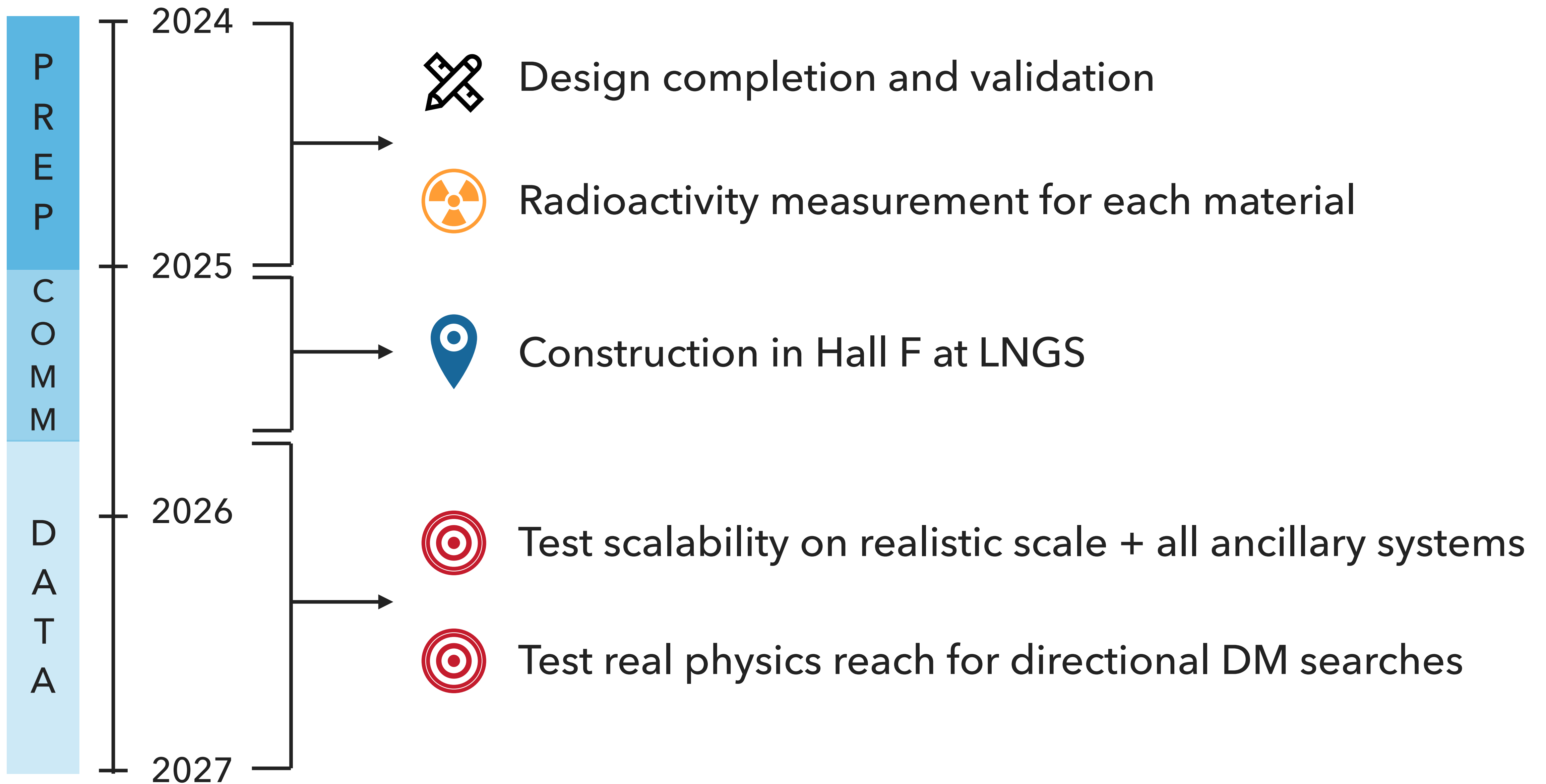
CYGNO_04

SERVICE AREA

CONTROL ROOM



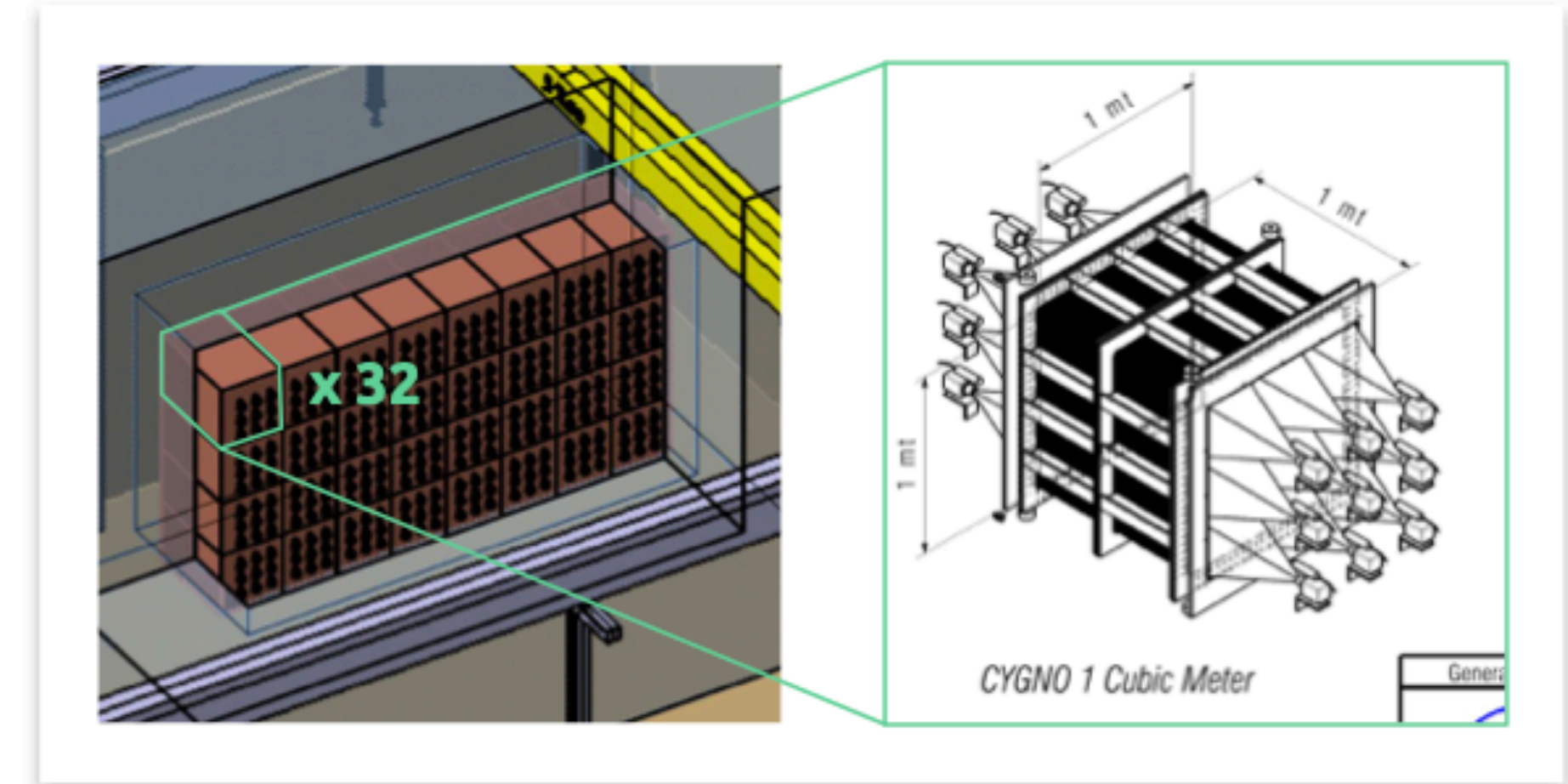
Projected shielding composed of
10 cm Cu + 100 cm H₂O



CYGNO-30 - Prospects



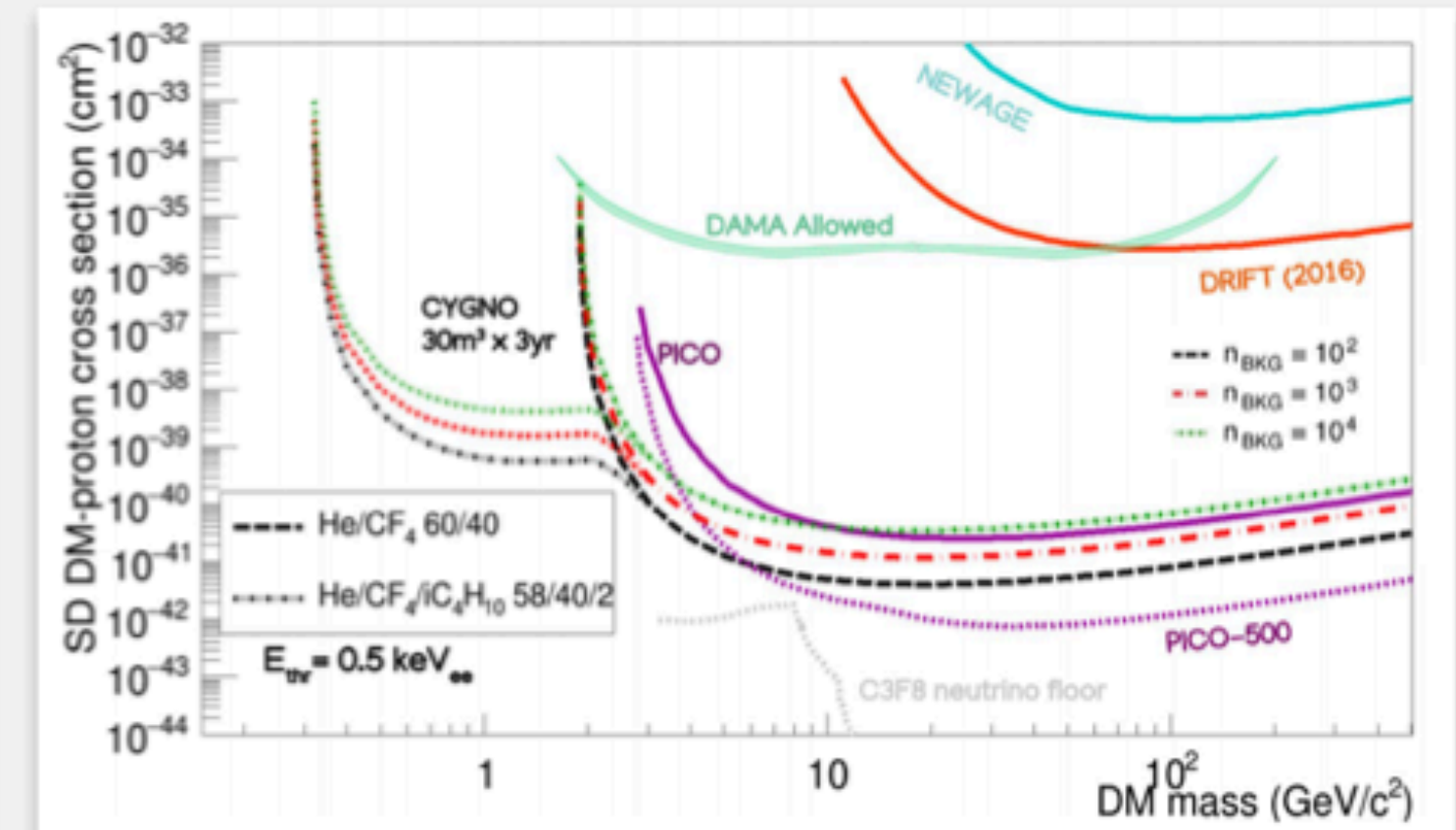
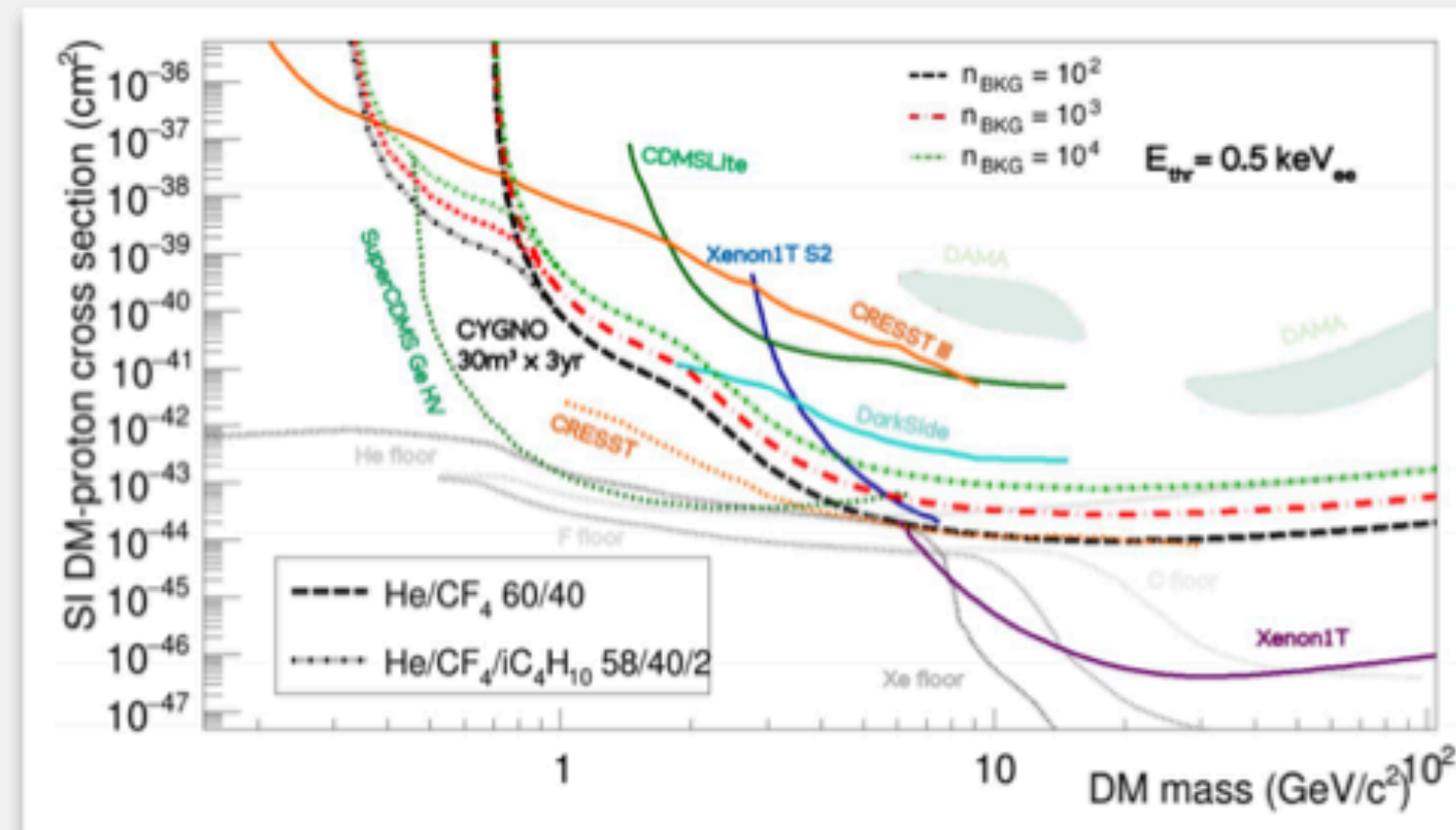
- **Low mass (0.5 - 10 GeV) directional DM searches**
- > 2027
- **30 - 100 m³** detector
- **0.5 - 1 keV_{ee}** energy threshold
- **30°** angular resolution



Expected **SI** and **SD** (90% CL)
interaction cross-section exclusion

Quenching factor simulated
with **SRIM** → Direct
measurement incoming!

He / (eventually H) allows us to
explore very low DM masses!





Energy response

Study of linearity and energy resolution (15-20%) performed with different X-ray sources

arXiv:2305.06168 [hep-ex]

