

NOvA - FNAL E929

Run: 46043 / 46

Event: 6183 / NuMI

UTC Sun Mar 26, 2023

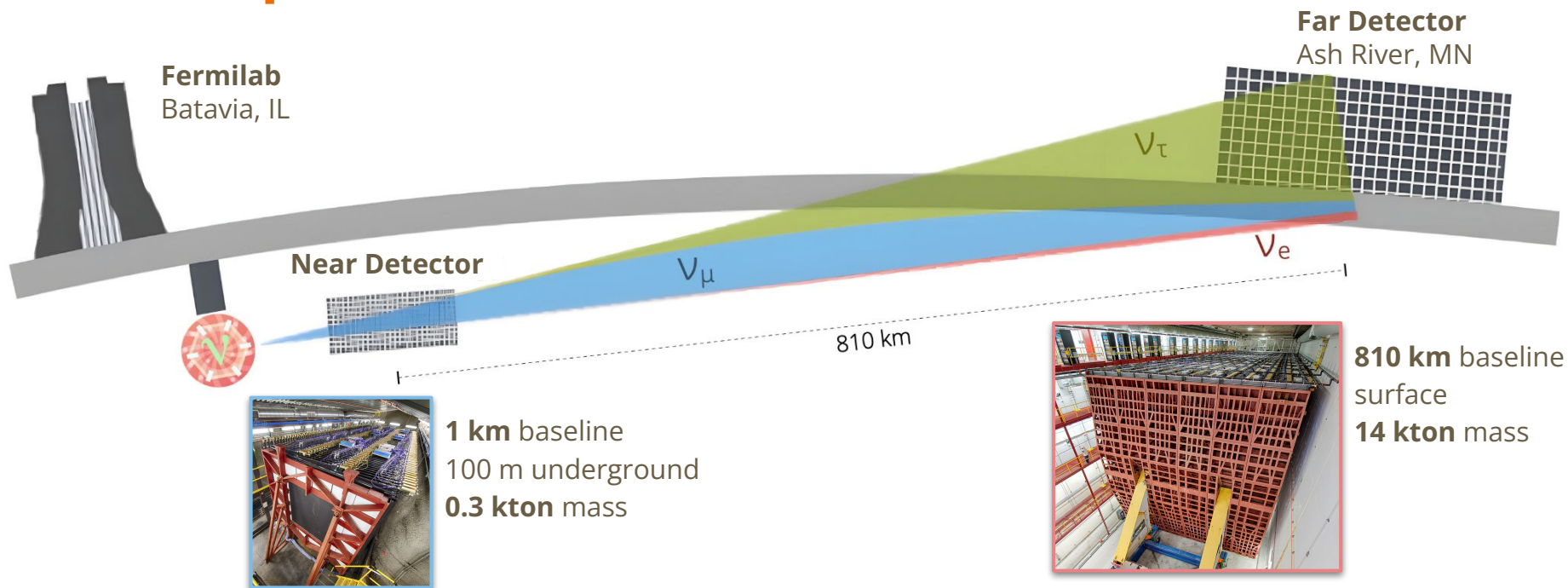
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42nd International Conference on High Energy Physics  
July 18th, 2024 - Prague, Czech Republic

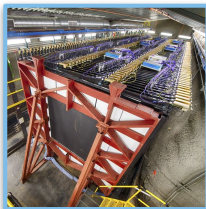
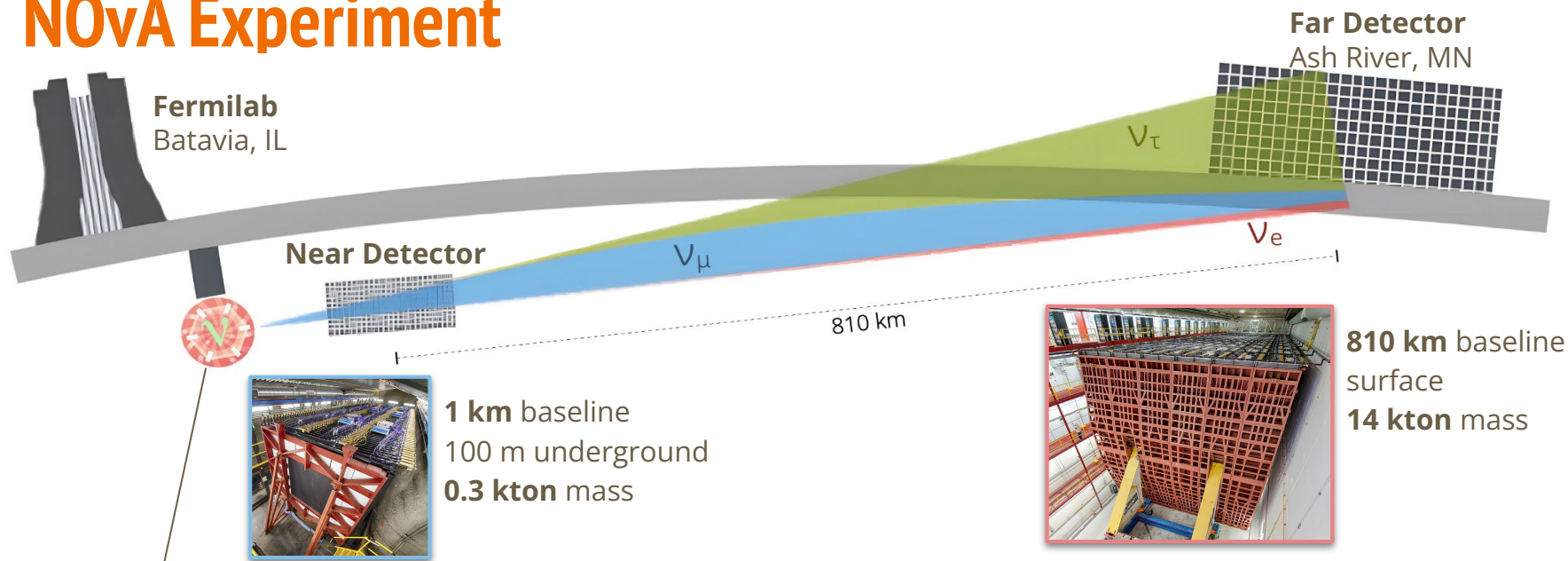


Caltech

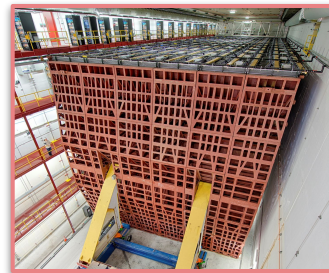
# NOvA Experiment



# NOvA Experiment



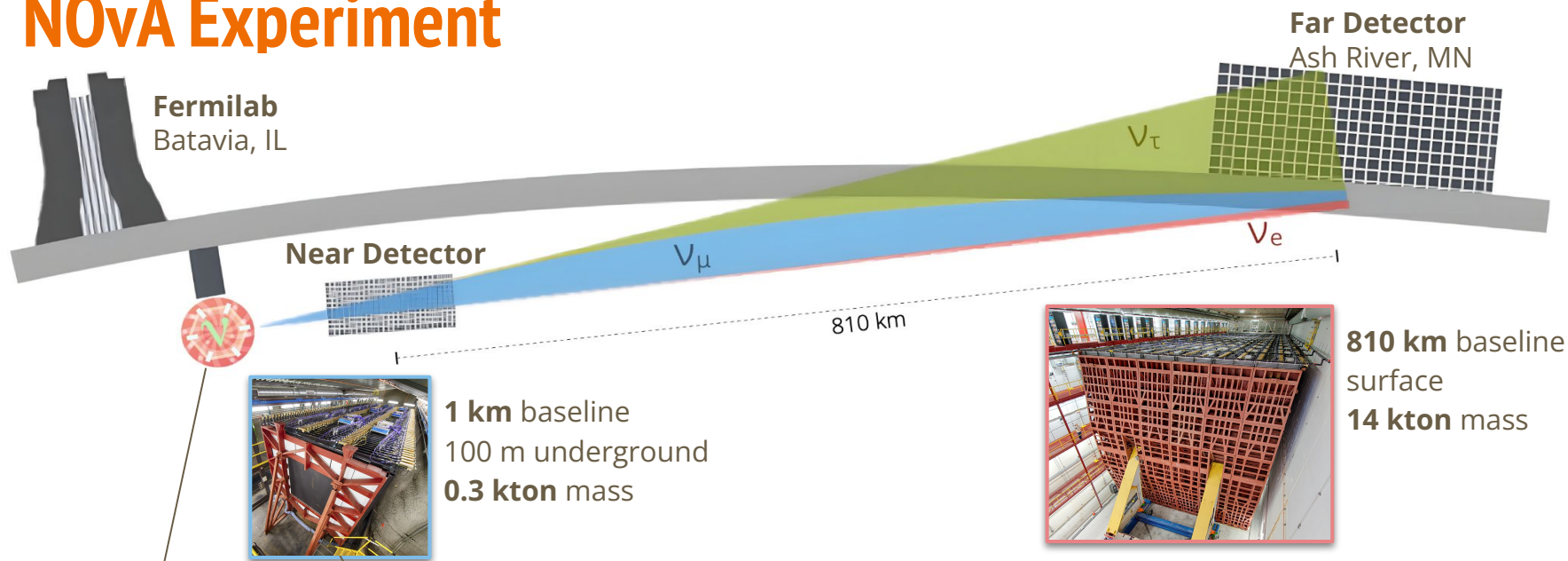
**1 km** baseline  
100 m underground  
**0.3 kton** mass



**810 km** baseline  
surface  
**14 kton** mass

**High intensity,  
 $\nu_\mu$ -pure beam.**  
Run in neutrino and  
antineutrino mode

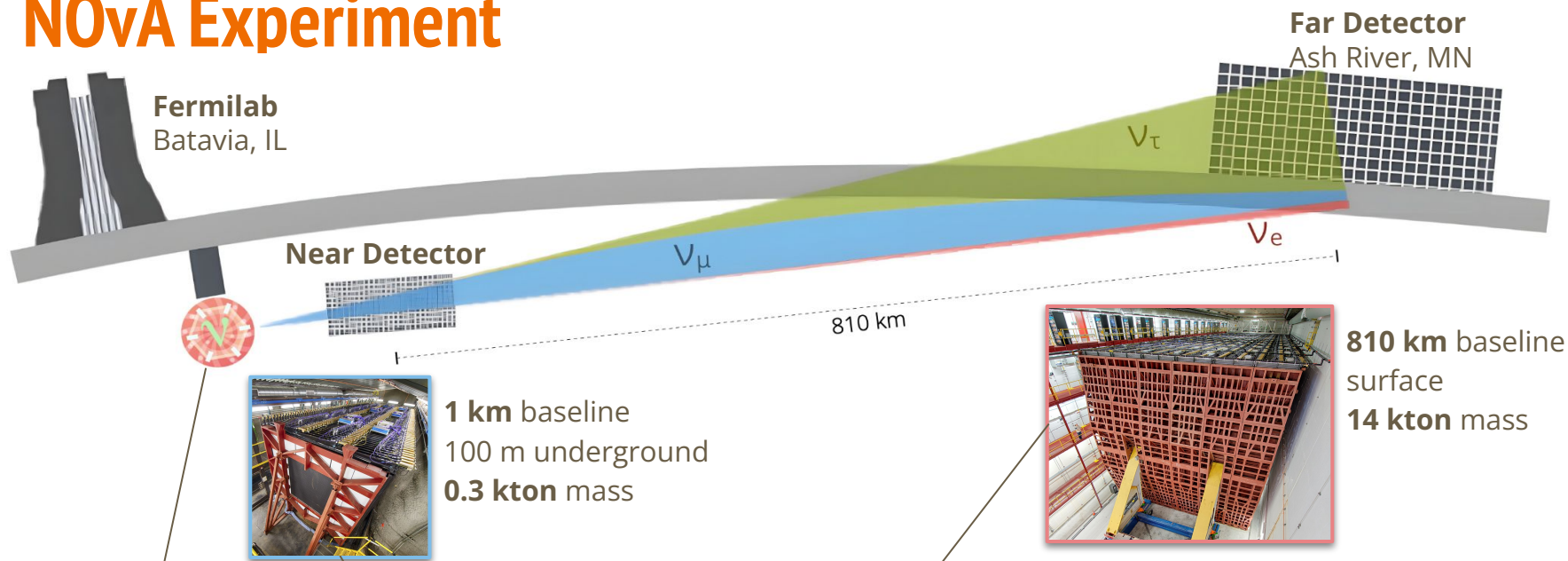
# NOvA Experiment



**High intensity,  
 $\nu_\mu$ -pure beam.**  
Run in neutrino and  
antineutrino mode

Use high-statistics in  
near detector to  
**constrain systematic  
uncertainties**

# NOvA Experiment



Fermilab  
Batavia, IL

Far Detector  
Ash River, MN

Near Detector

$\nu_\mu$

$\nu_\tau$

$\nu_e$

810 km

1 km baseline  
100 m underground  
0.3 kton mass

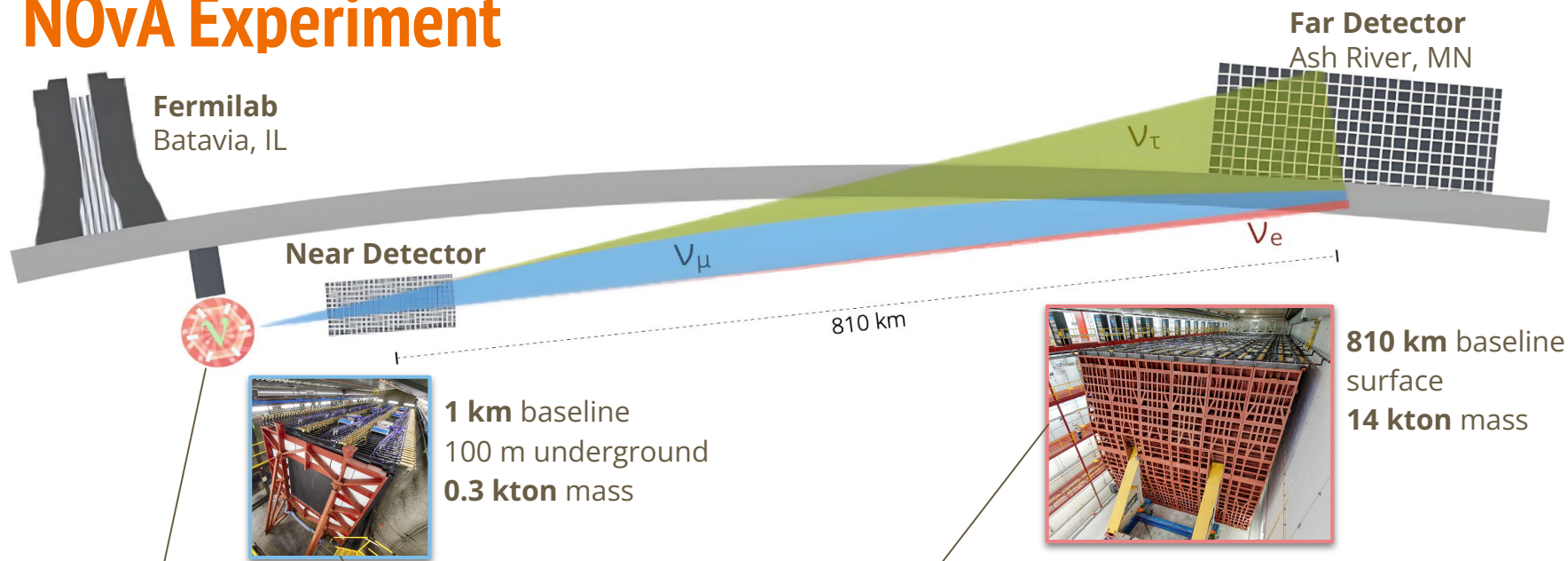
810 km baseline  
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**High intensity,  
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Measure four inclusive  
channels at far detector:  
 **$\bar{\nu}_\mu/\nu_\mu$  disappearance  
and  $\bar{\nu}_e/\nu_e$  appearance**

# NOvA Experiment



**High intensity,  $\nu_\mu$ -pure beam.**  
Run in neutrino and antineutrino mode

Use high-statistics in near detector to **constrain systematic uncertainties**

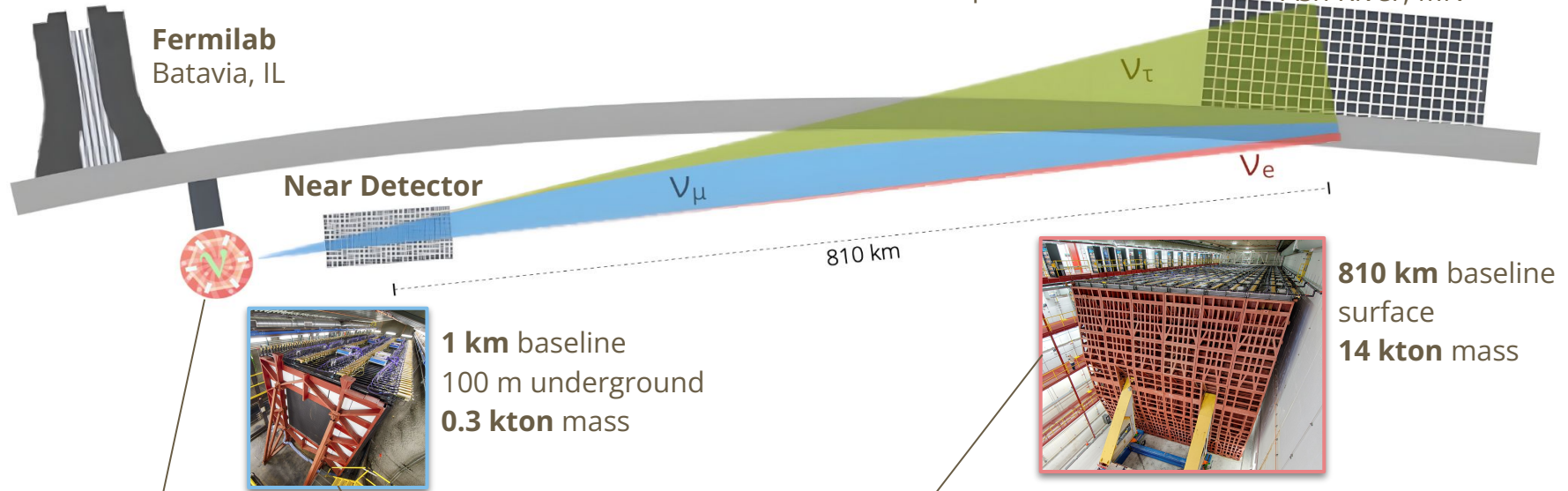
Measure four inclusive channels at far detector:  
 **$\bar{\nu}_\mu/\nu_\mu$  disappearance**  
and  **$\bar{\nu}_e/\nu_e$  appearance**

**Fit oscillation parameters**  
( $\Delta m_{32}^2, \sin^2\theta_{23}, \sin^2\theta_{13}, \delta_{CP}$ )  
to observed FD data and simulated predictions

# NOvA Experiment

NOvA addresses many compelling questions:

- What is the neutrino mass hierarchy?
- Is there CP symmetry violation in neutrinos?
- Is the 3-flavor model complete?



Fermilab  
Batavia, IL

Near Detector

Far Detector  
Ash River, MN

810 km

1 km baseline  
100 m underground  
0.3 kton mass

810 km baseline  
surface  
14 kton mass

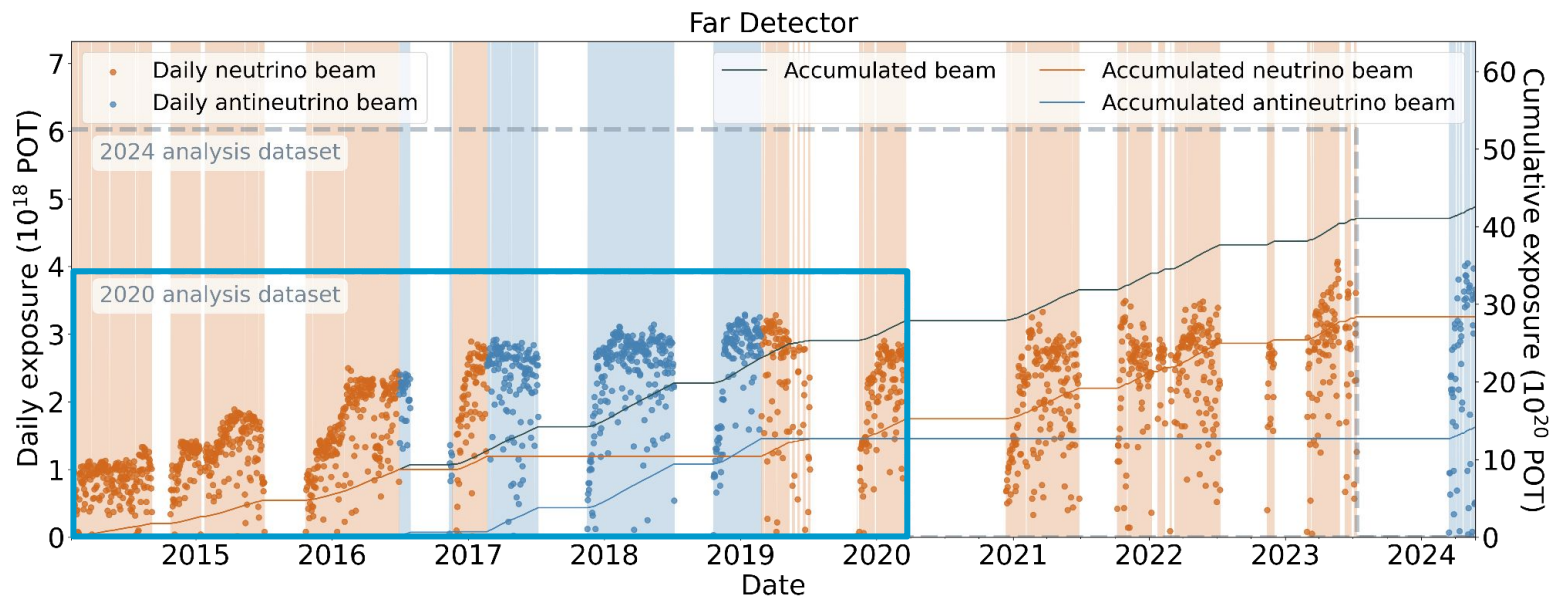
**High intensity,  $\nu_\mu$ -pure beam.**  
Run in neutrino and antineutrino mode

Use high-statistics in near detector to **constrain systematic uncertainties**

Measure four inclusive channels at far detector:  
 **$\bar{\nu}_\mu/\nu_\mu$  disappearance**  
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**Fit oscillation parameters**  
( $\Delta m_{32}^2$ ,  $\sin^2\theta_{23}$ ,  $\sin^2\theta_{13}$ ,  $\delta_{CP}$ )  
to observed FD data and simulated predictions

# New 2024 Oscillation Result



## 2020 data set:

Frequentist: [Phys. Rev D 106, 032004 \(2022\)](#)

Sterile ( $\nu$ -mode only): [FNAL JETP \(2022\)](#)

Bayesian: [Phys. Rev. D 110, 012005 \(2024\)](#)

NSI: [arXiv:2403.07266 \(2024\)](#)

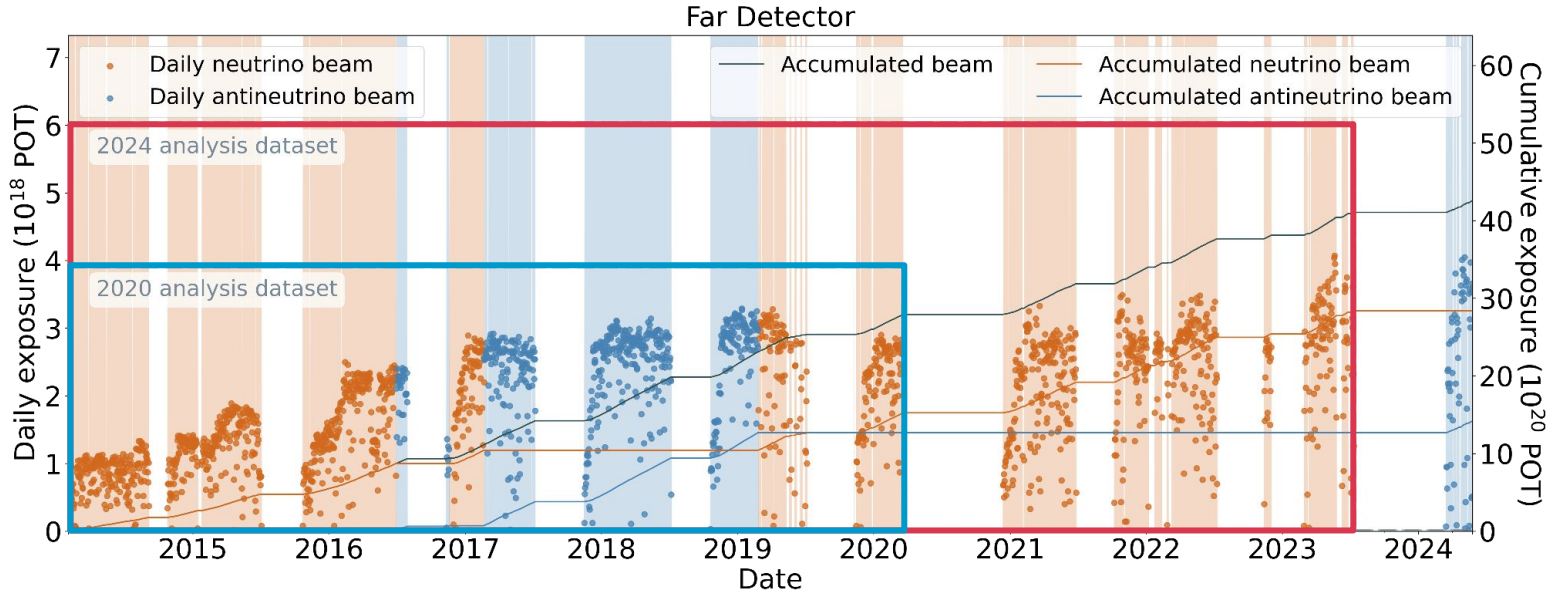
NOvA-T2K joint fit: [following talk](#) (T. Nosek)





# New 2024 Oscillation Result

Celebrating new 1.018MW NuMI beam power record.  
See poster from K. Yonehara: [Achievement in Beam Power Records for the NOvA Target System](#)



**New 2024 data set: Double the neutrino mode data since 2020!**

## 2020 data set:

Frequentist: [Phys. Rev D 106, 032004 \(2022\)](#)

Sterile ( $\nu$ -mode only): [FNAL JETP \(2022\)](#)

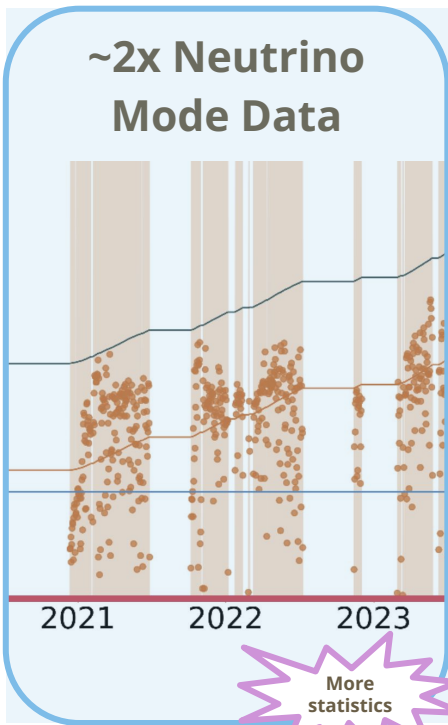
Bayesian: [Phys. Rev. D 110, 012005 \(2024\)](#)

NSI: [arXiv:2403.07266 \(2024\)](#)

NOvA-T2K joint fit: [following talk](#) (T. Nosek)

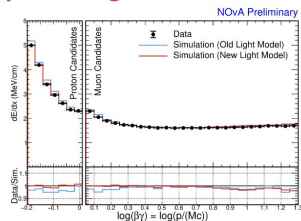
# Improvements to Oscillation Analysis

## ~2x Neutrino Mode Data

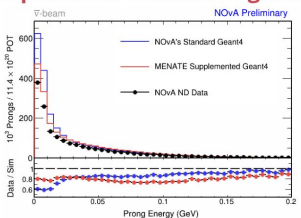


## Detector Characterization

### Improved Light Production Model

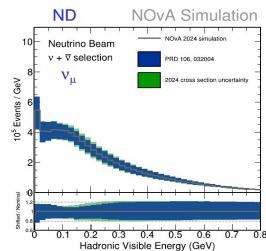


### Improved n-C Scattering Model



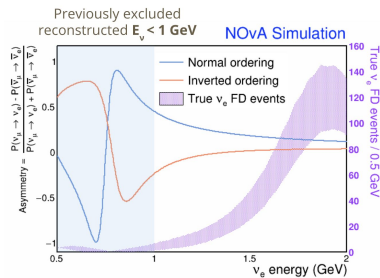
## Cross Section Modeling

### Additional Systematic Uncertainties for Pion Production



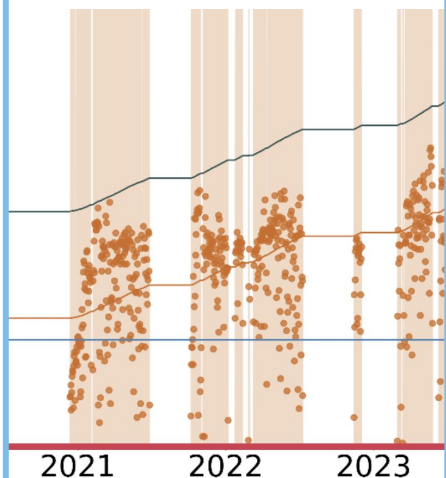
## Low Energy $\nu_e$ Sample

### New Selection to Enhance Sensitivity



# Improvements to Oscillation Analysis

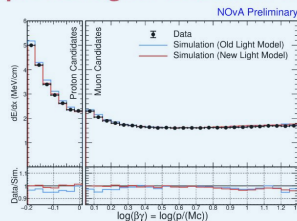
## ~2x Neutrino Mode Data



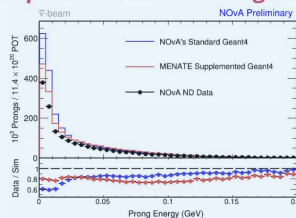
More statistics

## Detector Characterization

### Improved Light Production Model

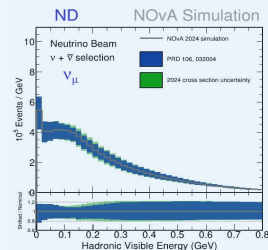


### Improved n-C Scattering Model



## Cross Section Modeling

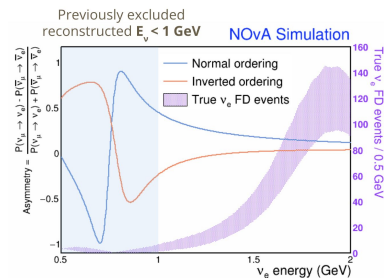
### Additional Systematic Uncertainties for Pion Production



Simulation improvements

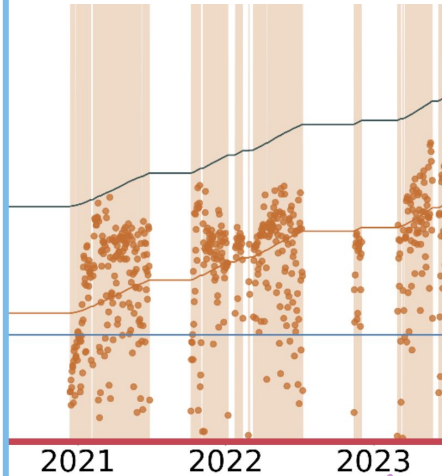
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# Improvements to Oscillation Analysis

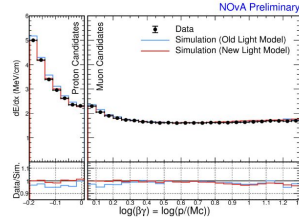
## ~2x Neutrino Mode Data



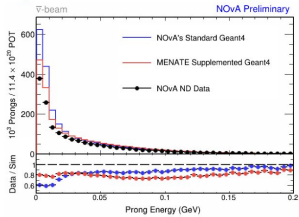
More statistics

## Detector Characterization

### Improved Light Production Model

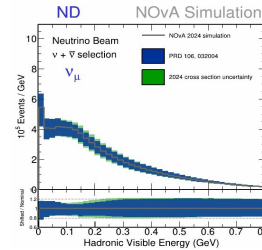


### Improved n-C Scattering Model



## Cross Section Modeling

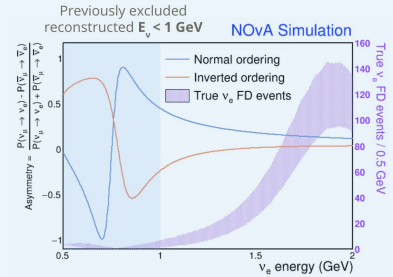
### Additional Systematic Uncertainties for Pion Production



Simulation improvements

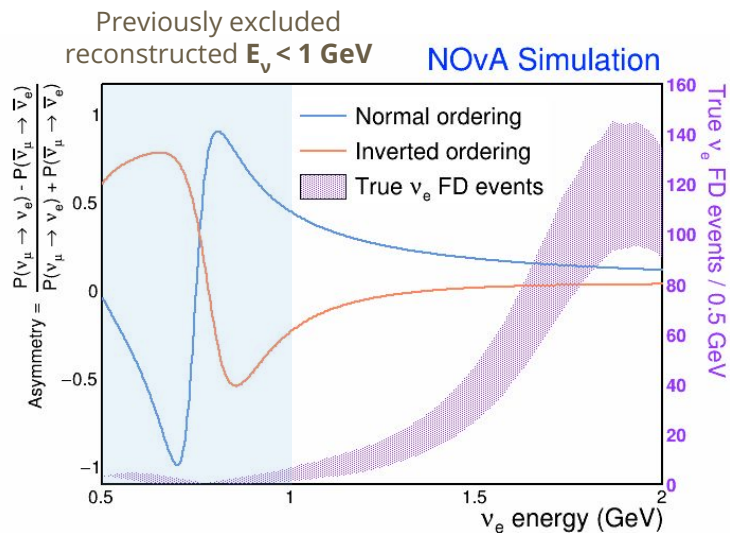
## Low Energy $\nu_e$ Sample

### New Selection to Enhance Sensitivity



New selection

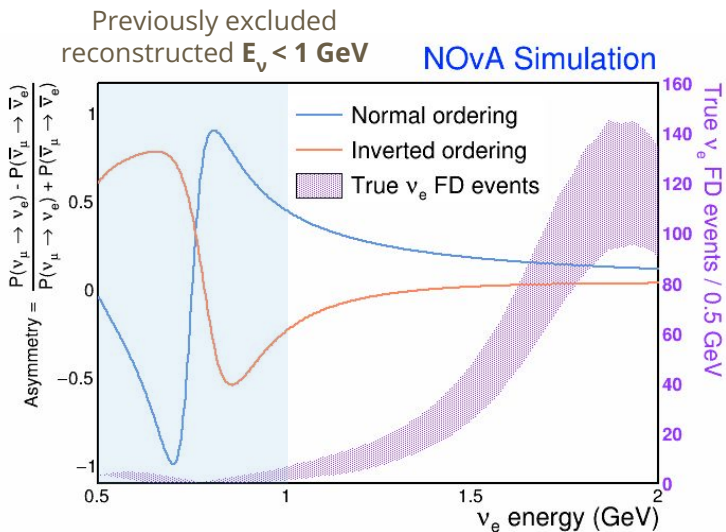
# Low Energy $\nu_e$ Sample



Low expected event rate but maximum  
ordering sensitivity from  $\nu_e - \bar{\nu}_e$  asymmetry  
at lower  $E_\nu$

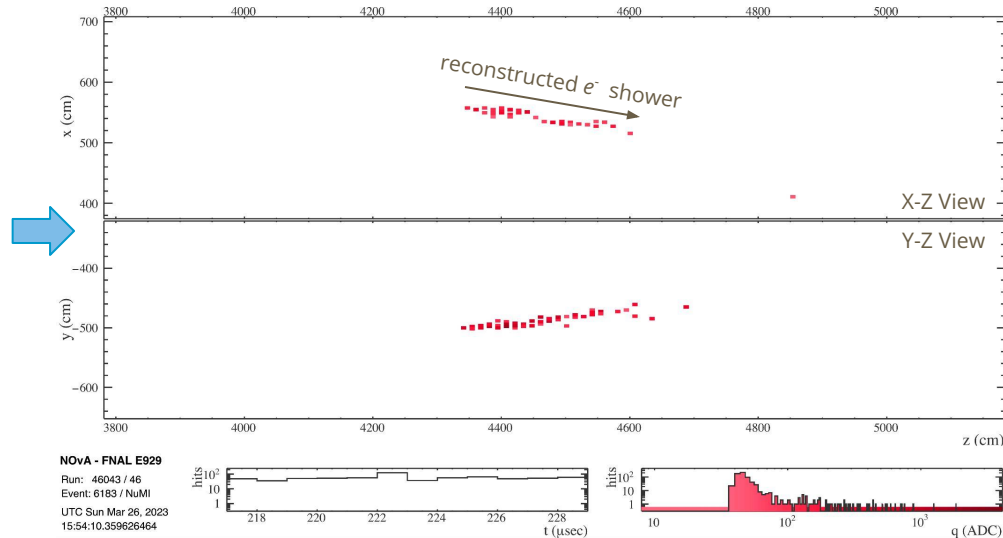
# Low Energy $\nu_e$ Sample

Neutrino-mode data event selected in the far detector by the low energy  $\nu_e$  2024 analysis.



Low expected event rate but maximum ordering sensitivity from  $\nu_e - \bar{\nu}_e$  asymmetry at lower  $E_\nu$

The reconstructed energy is 1.4 GeV.



Designed new selection to retain more  $\nu_e$  candidates in the  $0.5 \text{ GeV} \leq \text{reconstructed } E_\nu \leq 1.5 \text{ GeV}$  range using BDT to reject backgrounds

# Results with 2024 Data Set



NOvA - FNAL E929

Run: 46043 / 46

Event: 6183 / NuMI

UTC Sun Mar 26, 2023

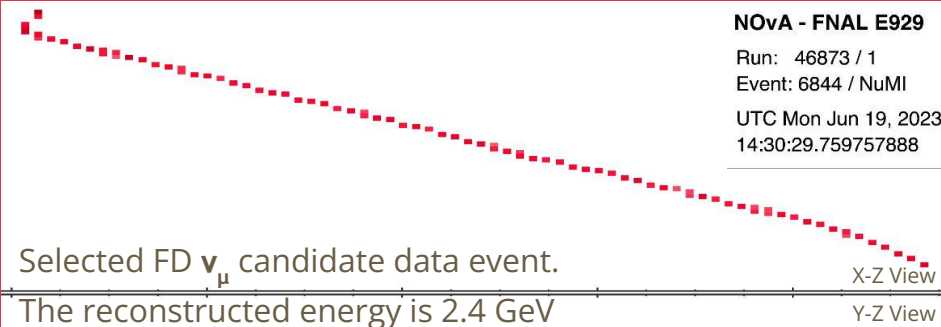
15:54:10.359626464

Selected FD  $\nu_e$  candidate data event.

The reconstructed energy is 1.4 GeV

X-Z View

Y-Z View



NOvA - FNAL E929

Run: 46873 / 1

Event: 6844 / NuMI

UTC Mon Jun 19, 2023

14:30:29.759757888

Selected FD  $\nu_\mu$  candidate data event.

The reconstructed energy is 2.4 GeV

X-Z View

Y-Z View

reconstructed

$p$  track

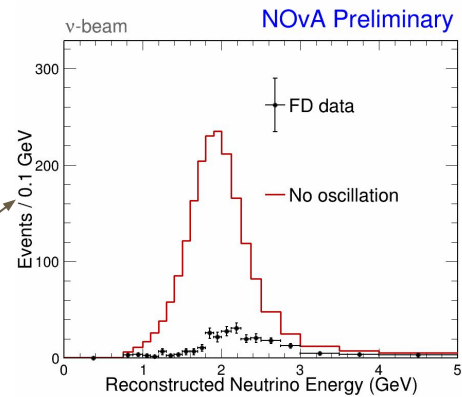
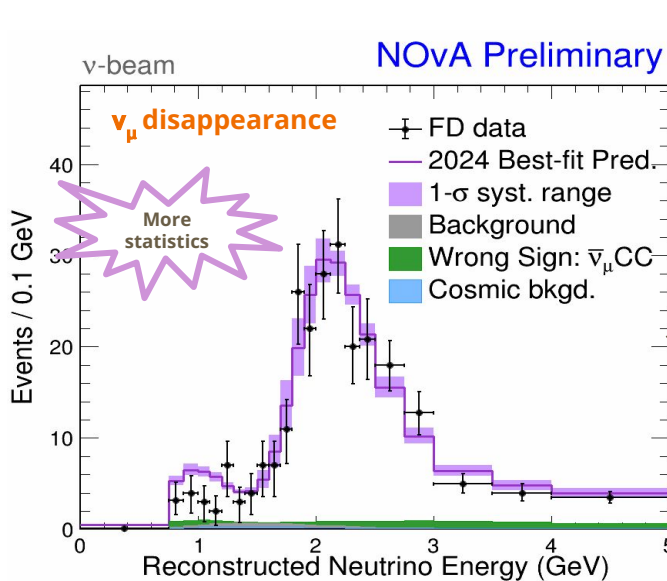


reconstructed  $\mu^-$  track





# Far Detector Muon (Anti)Neutrinos

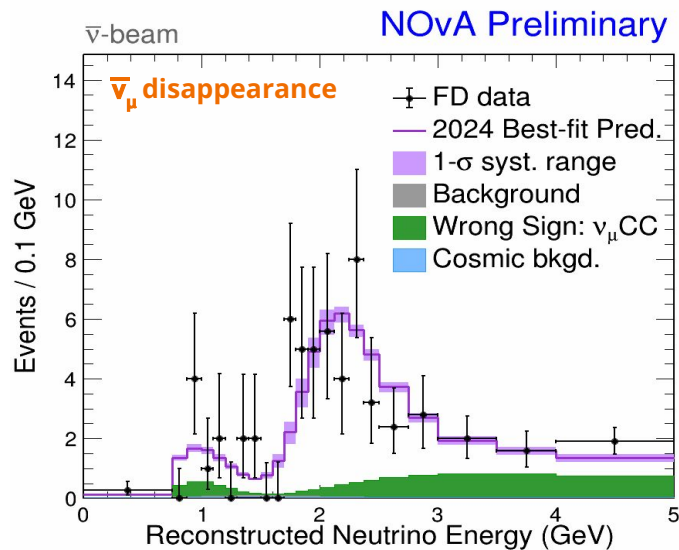
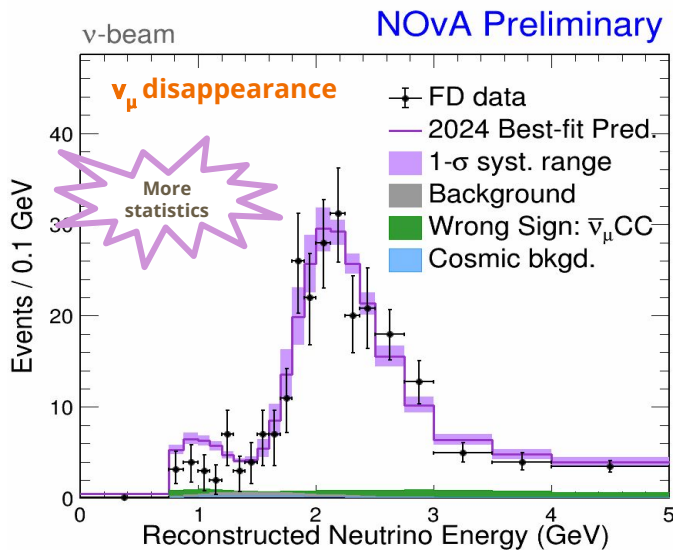


Clear oscillation signature as compared to no-oscillation prediction

We observe 384  $\nu_\mu$  and 106  $\bar{\nu}_\mu$  candidates in the FD. The central value and uncertainties are constrained by selections in the high-statistics, functionally identical ND via extrapolation.

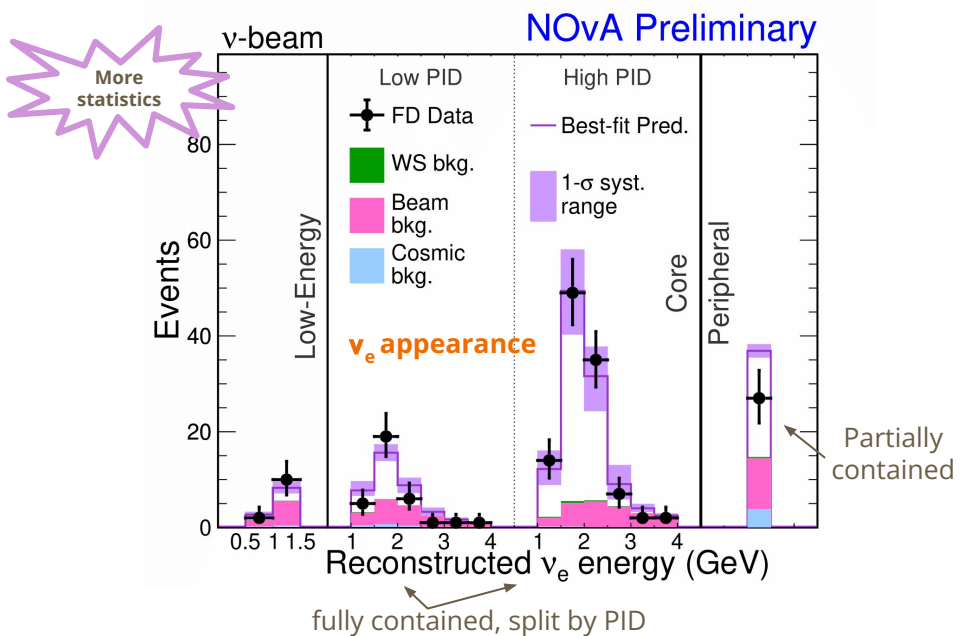


# Far Detector Muon (Anti)Neutrinos



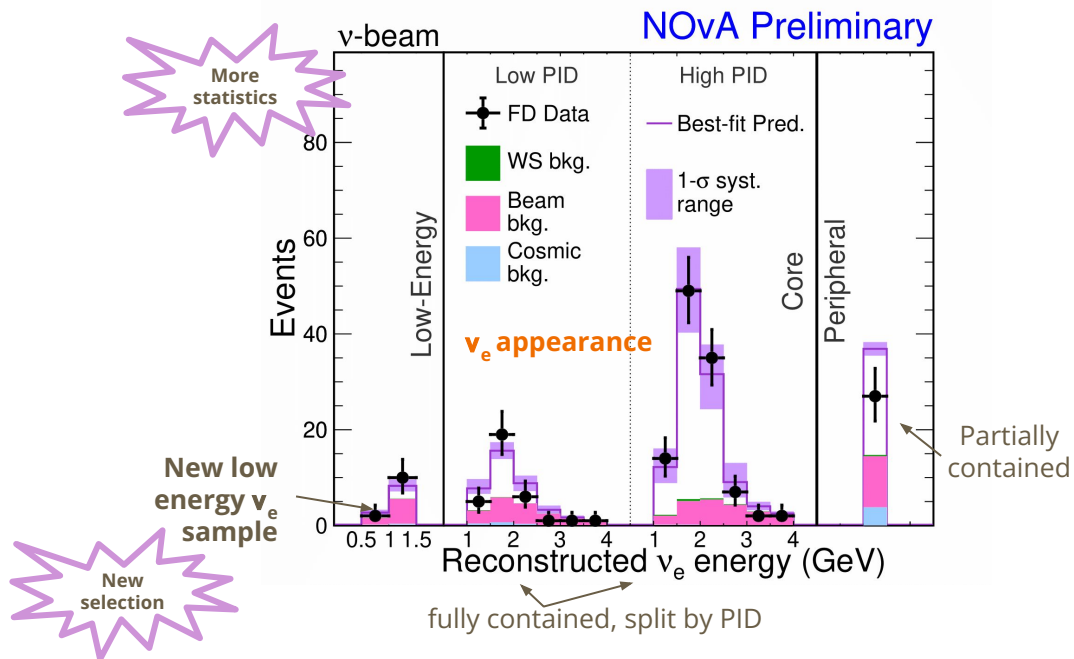
We observe 384  $\nu_\mu$  and 106  $\bar{\nu}_\mu$  candidates in the FD. The central value and uncertainties are constrained by selections in the high-statistics, functionally identical ND via extrapolation.

# Far Detector Electron (Anti)Neutrinos



**We observe 181  $\nu_e$  and 32  $\bar{\nu}_e$  candidates in the FD.** The central value and uncertainties are similarly constrained by near-far detector extrapolation.

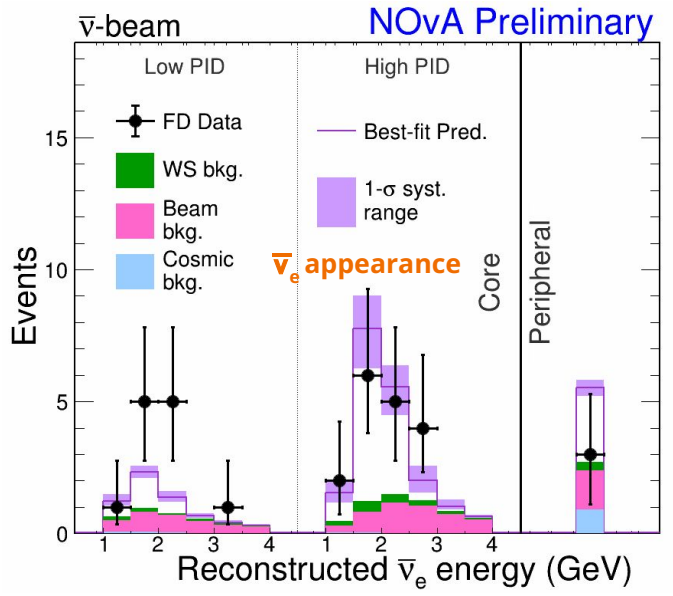
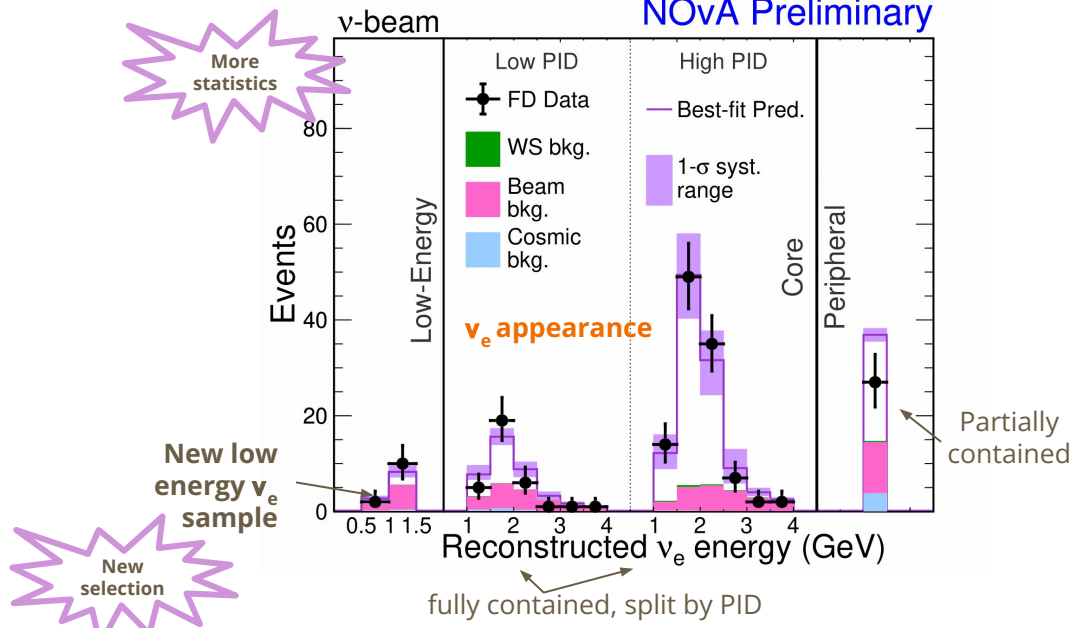
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Simulation improvements

# Far Detector Electron (Anti)Neutrinos

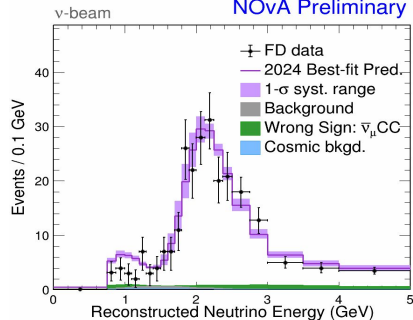


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# Fit to Oscillation Parameters

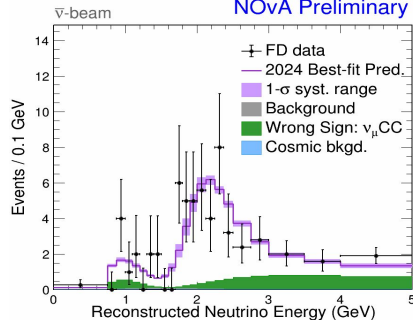
## $\nu_\mu$ disappearance

NOvA Preliminary



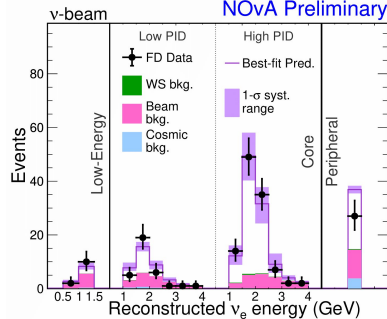
## $\bar{\nu}_\mu$ disappearance

NOvA Preliminary



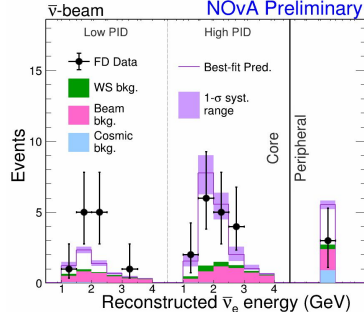
## $\nu_e$ appearance

NOvA Preliminary



## $\bar{\nu}_e$ appearance

NOvA Preliminary

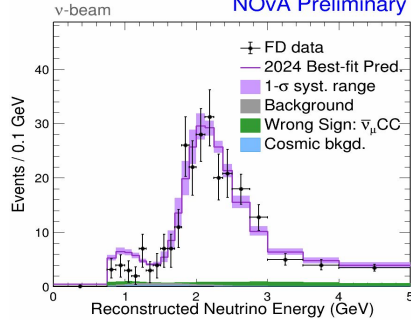


Fit to  $\Delta m^2_{32}$ ,  $\sin^2\theta_{23}$ ,  $\sin^2 2\theta_{13}$ ,  $\delta_{CP}$

# Fit to Oscillation Parameters

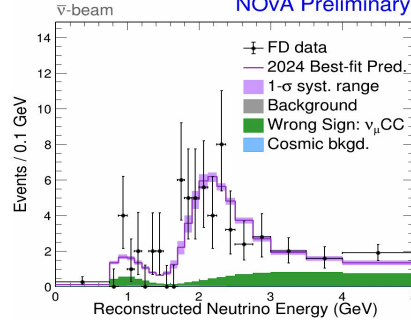
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NOvA Preliminary



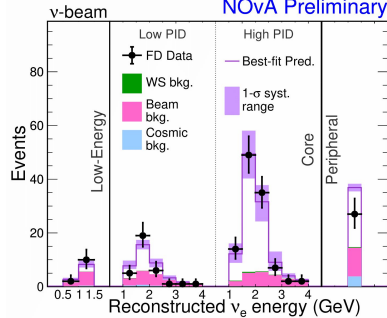
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NOvA Preliminary



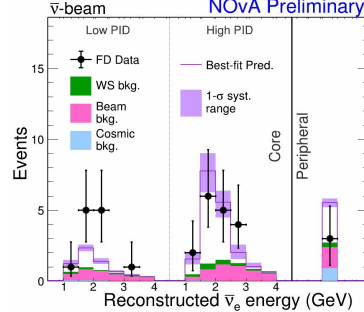
## $\nu_e$ appearance

NOvA Preliminary



## $\bar{\nu}_e$ appearance

NOvA Preliminary



Fit to  $\Delta m^2_{32}$ ,  $\sin^2\theta_{23}$ ,  $\sin^22\theta_{13}$ ,  $\delta_{CP}$

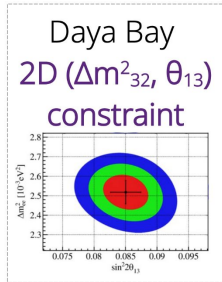
Consider three  $\theta_{13}$  possibilities:

$\theta_{13}$  unconstrained  
(NOvA only)

or

Daya Bay  
1D  $\theta_{13}$  constraint  
 $\sin^2 2\theta_{13} = 0.0851 \pm 0.0024$

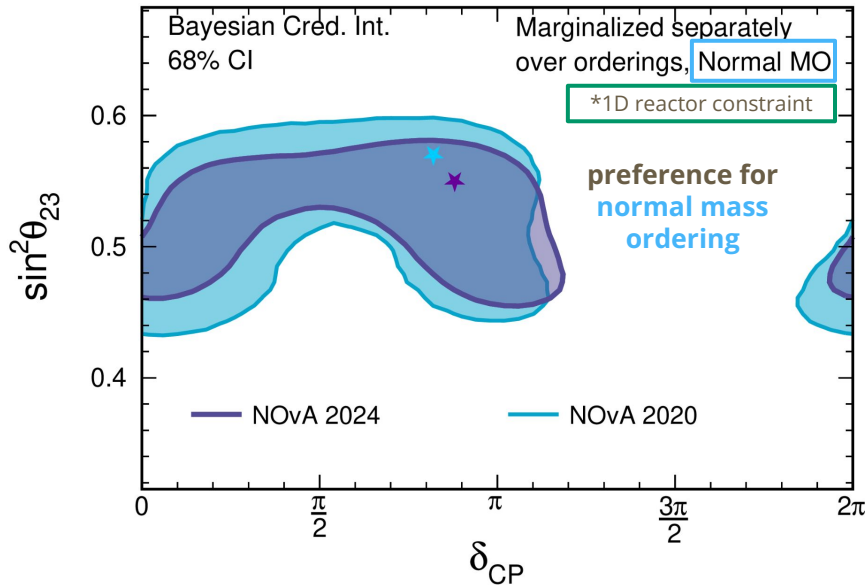
or



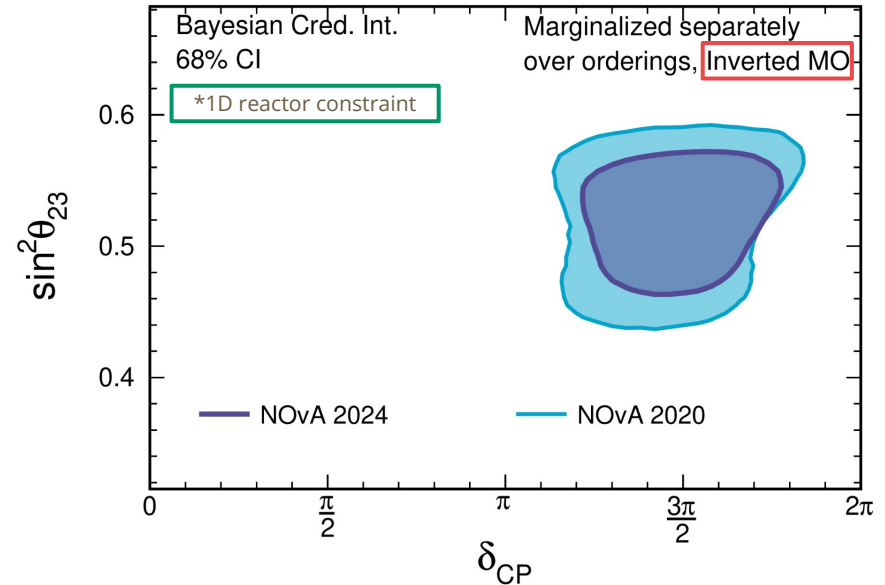
# Comparison 2020 and 2024 Results

See strong consistency between 2020 and 2024 results, with the improved constraint in ~same regions

NOvA Preliminary



NOvA Preliminary



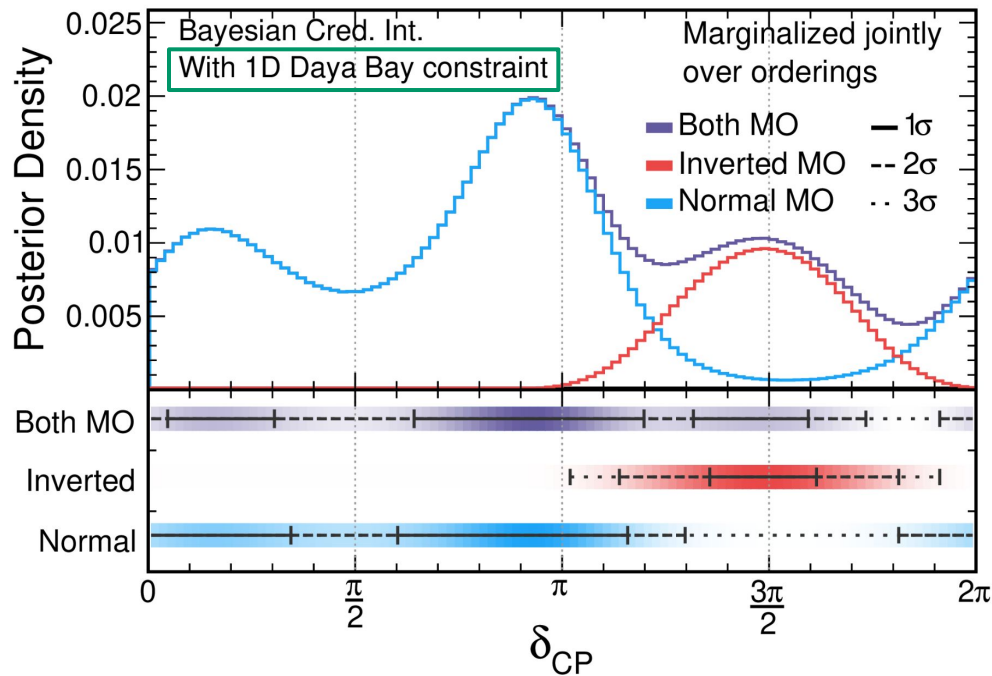
# 2024 Results: $\delta_{CP}$

Statements on  $\delta_{CP}$  dependent on mass ordering determination

CP-conserving points favored in **normal ordering**

CP-conservation outside  $3\sigma$  interval in **inverted ordering**, preference around maximal violation at  $3\pi/2$

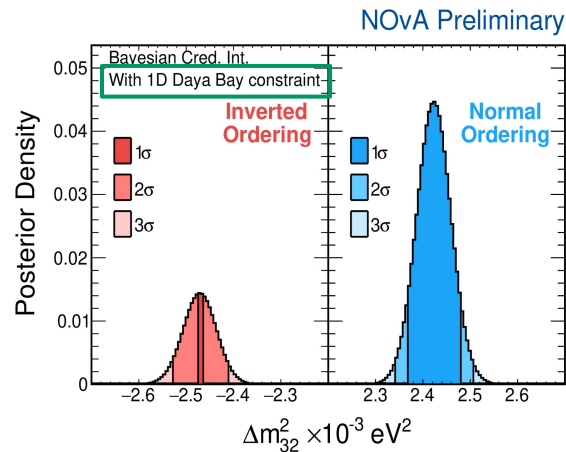
NOvA Preliminary





# 2024 Results: $\Delta m_{32}^2$

Small preference for **normal mass ordering**

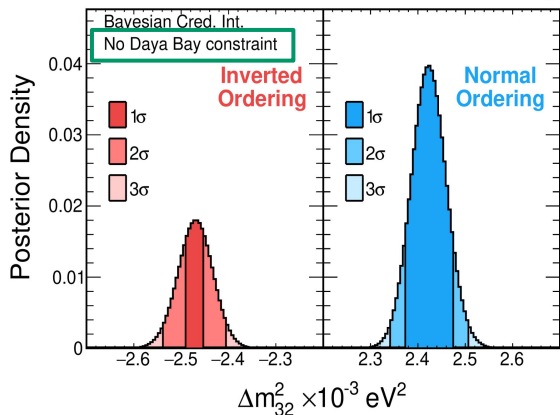


Daya Bay 2023 1D $\sin^2 2\theta_{13}$		
	Probability	Bayes Factor
Normal MO preference	<b>76%</b>	3.2

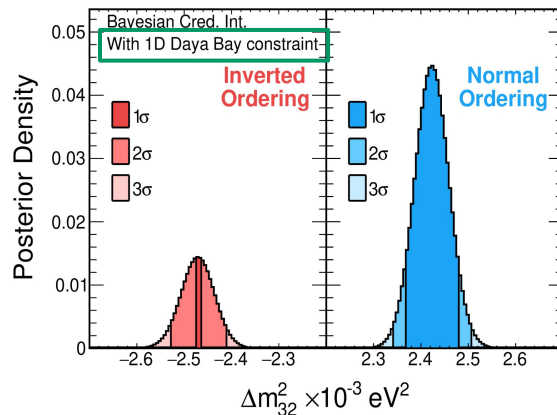
# 2024 Results: $\Delta m_{32}^2$

Preference for **normal mass ordering** is enhanced by choice of reactor constraint

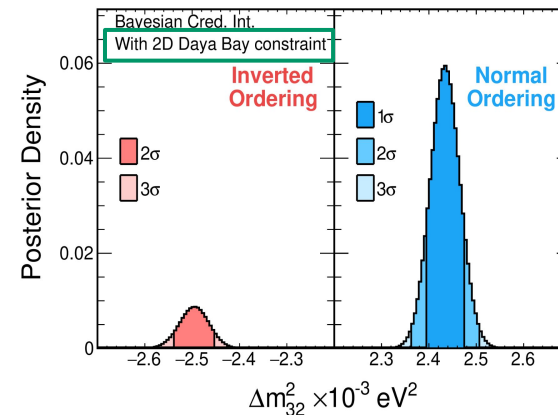
NOvA Preliminary



NOvA Preliminary



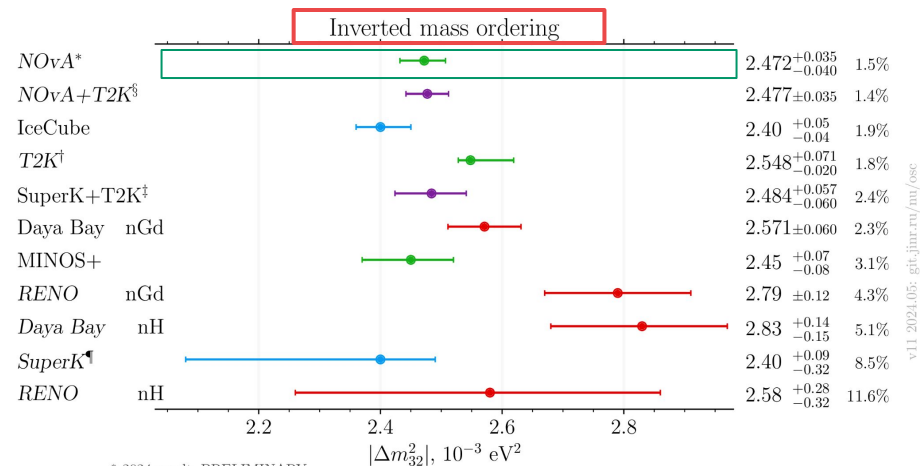
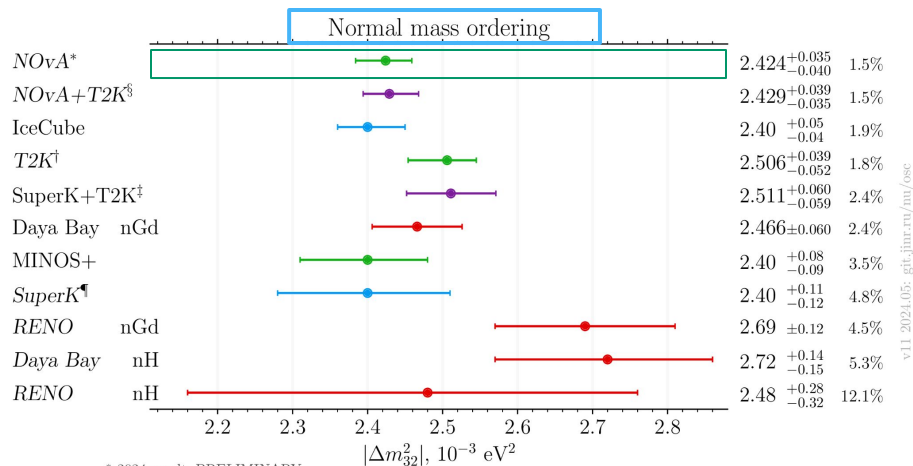
NOvA Preliminary



	No constraint		Daya Bay 2023 1D $\sin^2 2\theta_{13}$		Daya Bay 2023 2D ( $\sin^2 2\theta_{13}, \Delta m_{32}^2$ )	
	Probability	Bayes Factor	Probability	Bayes Factor	Probability	Bayes Factor
Normal MO preference	<b>69%</b>	2.2	<b>76%</b>	3.2	<b>87%</b>	6.8



# 2024 Results: $\Delta m_{32}^2$



\* 2024 result, PRELIMINARY  
 Preliminary  
 Published † based on 2020 ana.  
 ‡ Neutrino-2022 result

§ SKI-V result, arXiv:2311.05105  
 ‡ based on SK IV and T2K 2020, arXiv:2405.12488

\* 2024 result, PRELIMINARY  
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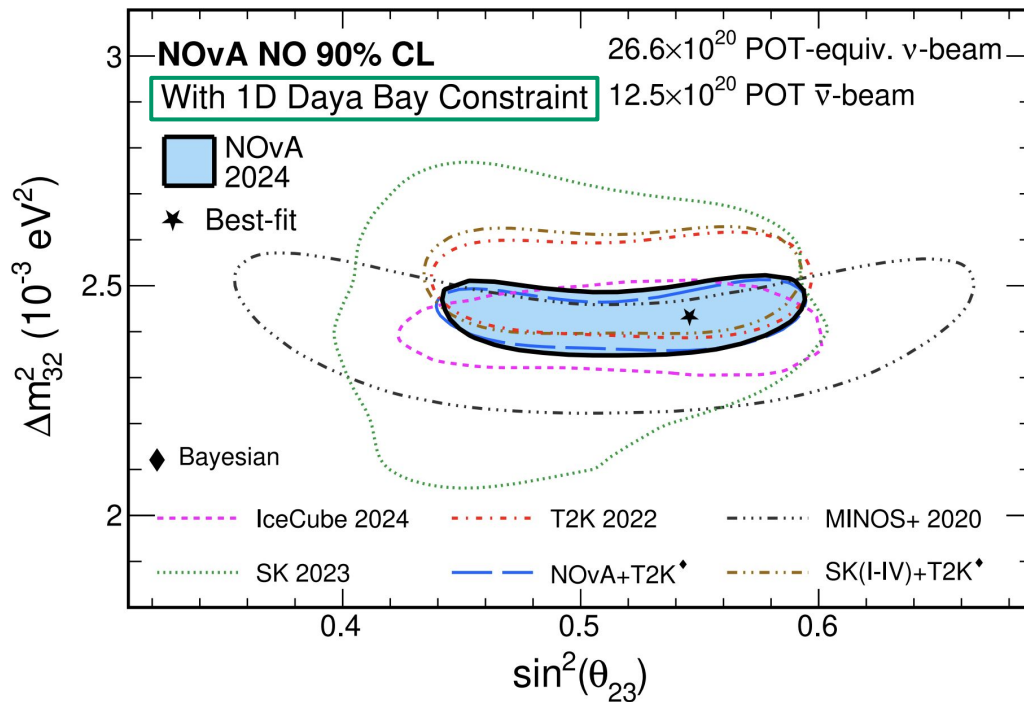
NOvA-only  $|\Delta m_{32}^2|$  result has world-leading precision of  $\sim 1.5\%$ , best single-experiment measurement.

# 2024 Results: $\sin^2\theta_{23}$

In the  $\nu_2 - \nu_3$  sector, NOvA measurements are consistent with accelerator, atmospheric, and joint results

IceCube 2024: [arXiv:2405.02163](https://arxiv.org/abs/2405.02163) T2K 2022: [10.5281/zenodo.6683821](https://arxiv.org/abs/10.5281/zenodo.6683821)  
MINOS+ 2020: [Phys. Rev. Lett. 125, 131802](https://arxiv.org/abs/125.131802)  
SK 2023: [Phys. Rev. D109, 072014](https://arxiv.org/abs/109.072014)  
NOvA+T2K 2024: [KEK IPNS seminar, FNAL JETP seminar](https://arxiv.org/abs/2405.12488)  
T2K+SK 2024: [arXiv:2405.12488](https://arxiv.org/abs/2405.12488)

NOvA Preliminary



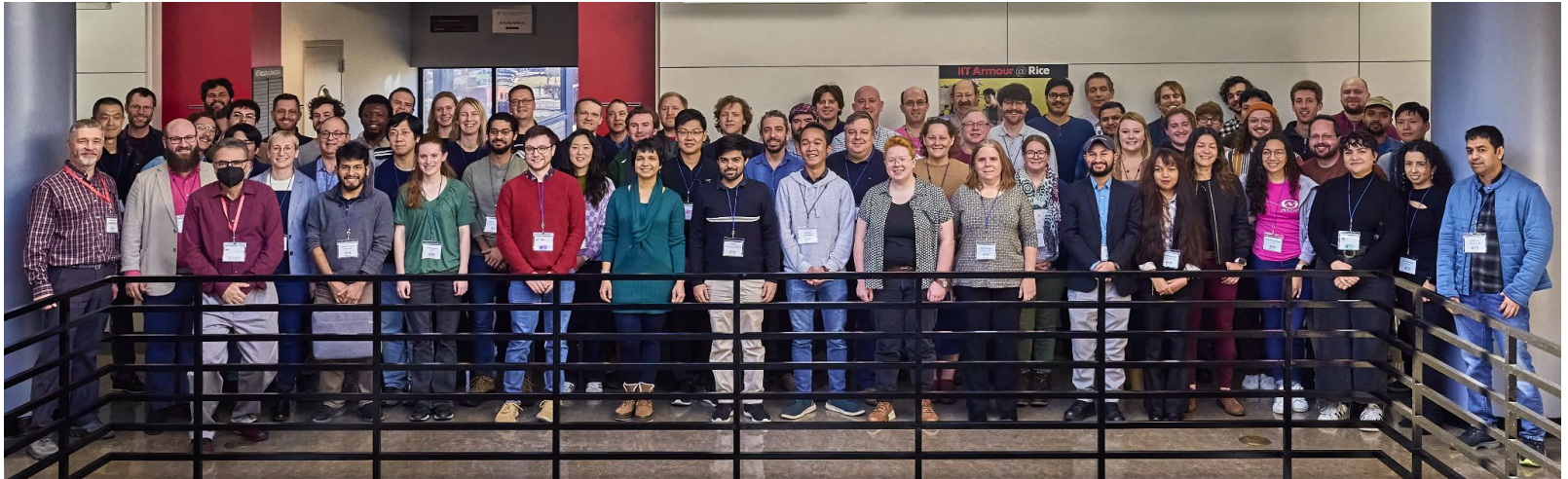
# Summary

More NOvA Talks at ICHEP:

- [Triple Differential Muon Antineutrino Charge Current Inclusive Cross Section Measurement in NOvA](#) ( P. Singh)
- [Deep Learning Event Reconstruction at NOvA](#) (W. Wu)

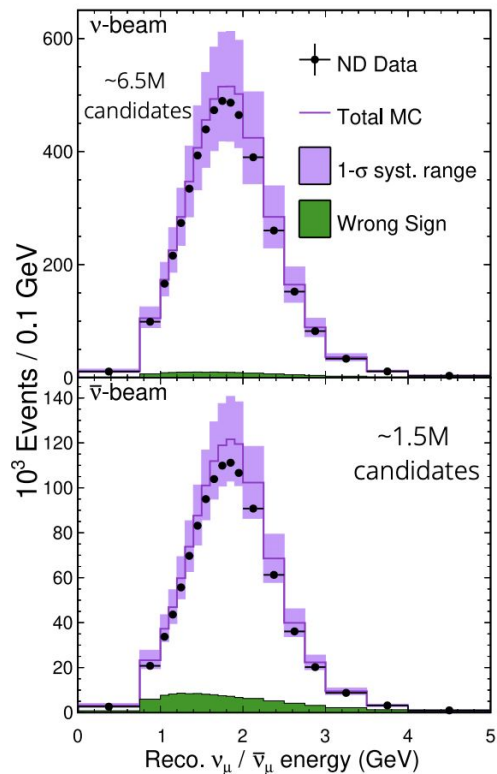
- **The new 2024 analysis is a major update to NOvA's oscillation result**
  - Doubled neutrino-mode dataset with 10 years of neutrino & antineutrino data
  - Improvements to the detector characterization, cross section uncertainties, and a new low-energy  $\nu_e$  sample
- **Most precise single-experiment measurement of  $\Delta m_{32}^2$  (1.5%)**
- Statements on **CP violation strongly coupled to mass-hierarchy determination**
- Strong synergy with reactor measurements
- The best is yet to come!
  - **Goal to double the anti-neutrino mode data in the final data set**
  - Test beam results could address some of the largest systematic uncertainties in NOvA
  - Broader program including sterile searches, NSI, cross section measurements, cosmic ray physics, exotics... and more!

Thanks!

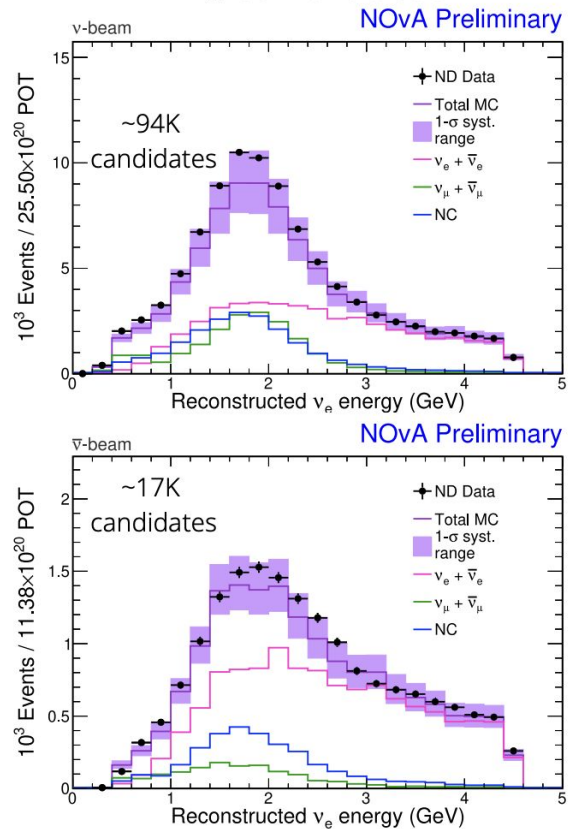


# ND data distributions

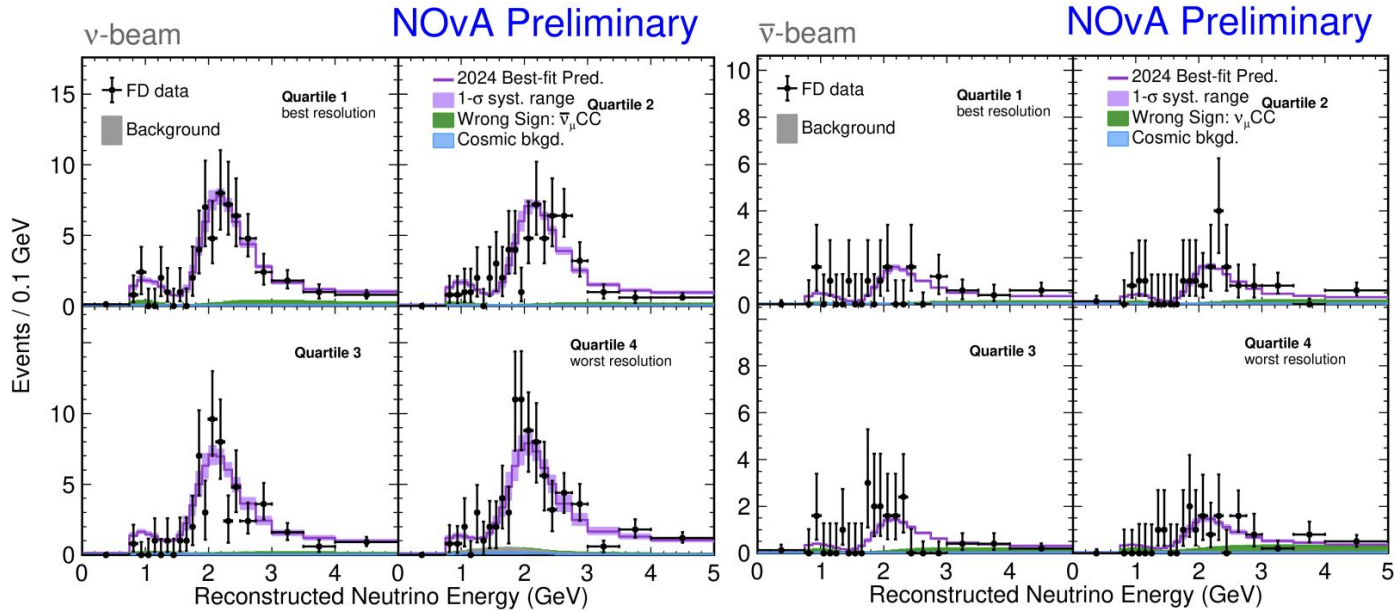
## $\nu_\mu$ candidates



## $\nu_e$ candidates



# FD data: $\nu_\mu$ in $E_{\text{had}}/E_\nu$ quartiles



Extrapolation procedure is performed in  
 $|\mathbf{p}_t|$  subpopulations of  $E_{\text{had}}/E_\nu$  quartiles  
Resolutions range from Q1 6.5% (5.4%) to Q4 12.6% (11.2%) in  $\nu$  ( $\bar{\nu}$ ) mode

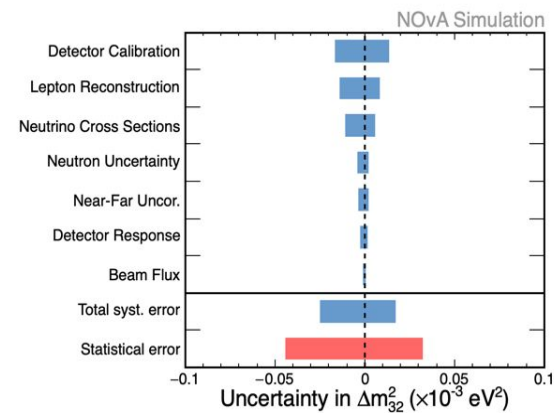
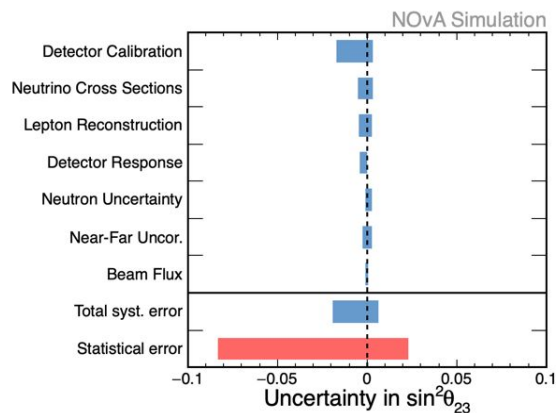
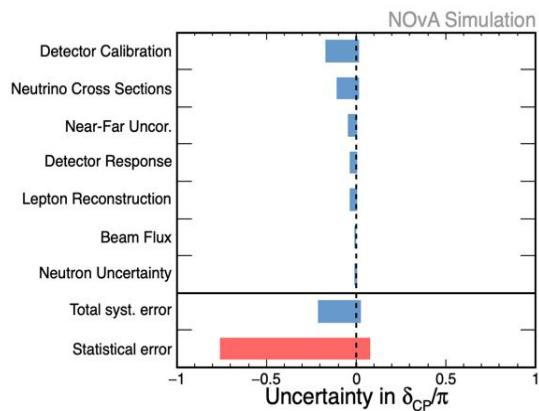


# FD Selections

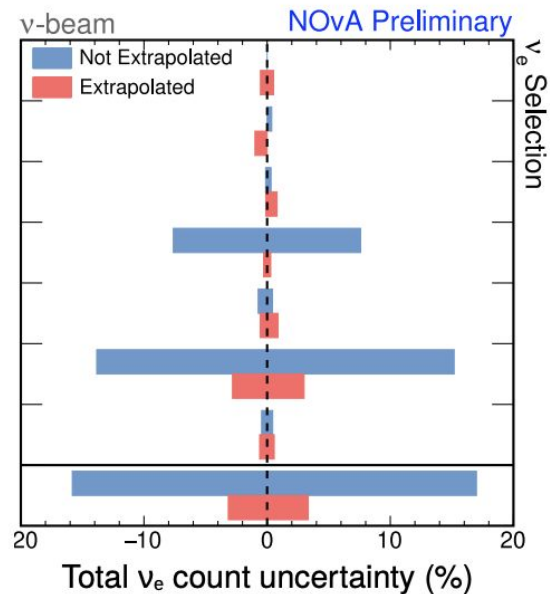
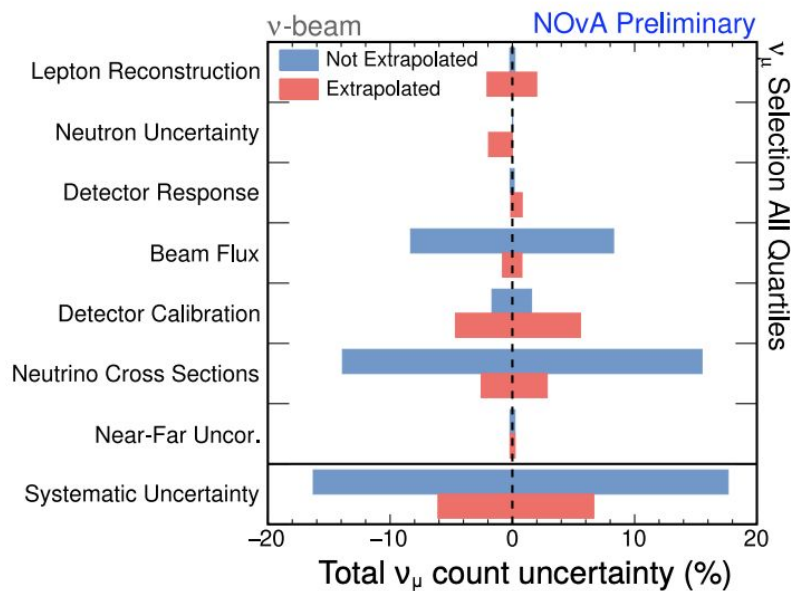
	Neutrino Beam			Anti-neutrino Beam	
	$\nu_\mu$	$\nu_e$	LowE $\nu_e$	$\bar{\nu}_\mu$	$\bar{\nu}_e$
$\nu_\mu \rightarrow \nu_\mu$	372.7	4.0	0.3	24.7	0.2
$\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$	24.7	0.1	0.0	71.8	0.2
$\nu_\mu \rightarrow \nu_e$	0.4	121.6	2.9	0.0	2.1
$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	0.0	1.7	0.1	0.0	18.2
Beam $\nu_e + \bar{\nu}_e$	0.1	26.3	0.8	0.0	6.5
NC	5.5	16.3	5.0	0.8	2.0
Cosmic	4.6	5.7	0.5	0.7	1.1
Others	1.5	0.8	0.1	0.2	0.1
Signal	398.2	121.6	2.9	96.7	18.2
Background	11.3	54.9	6.8	1.7	12.2
Best fit	409.5	176.5	9.7	98.4	30.4
Observed	384.0	169.0	12.0	106.0	32.0

**Table:** The Ana2024 observed and predicted  $\nu_\mu$  ( $\bar{\nu}_\mu$ ) disappearance and  $\nu_e$  ( $\bar{\nu}_e$ ) appearance events at the Far Detector including the lowE  $\nu_e$  events. The predicted number of events are scaled to the Ana2024 NOUO best-fit values.

# Uncertainties on Oscillation Parameters



# Systematic Uncertainties

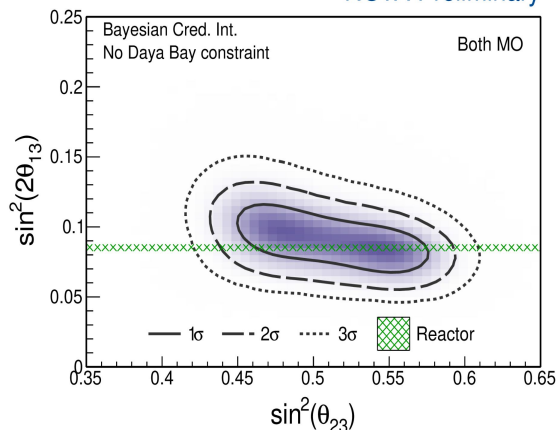


# 2024 Results: $\theta_{23}$

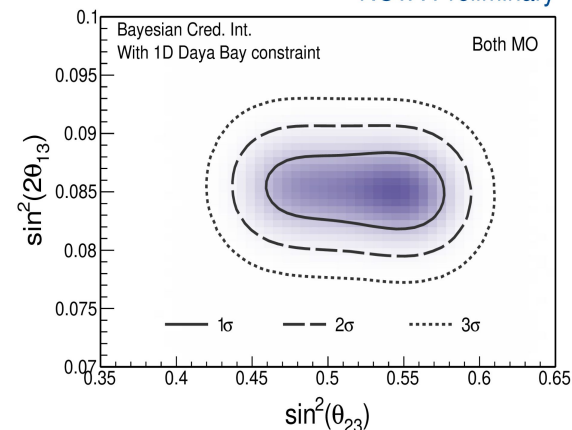
Mild preference for upper octant (69% probability) that emerges from applying the reactor constraint

Maximal mixing is allowed at  $<1\sigma$

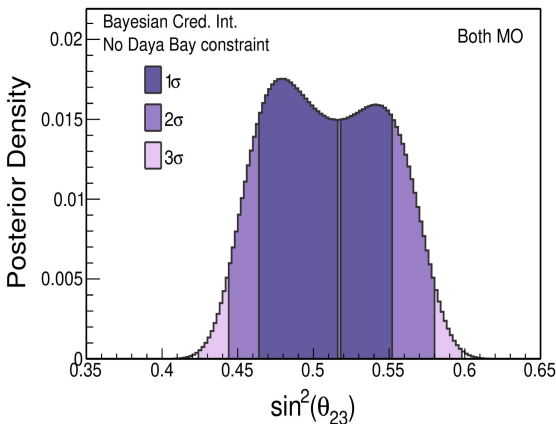
NOvA Preliminary



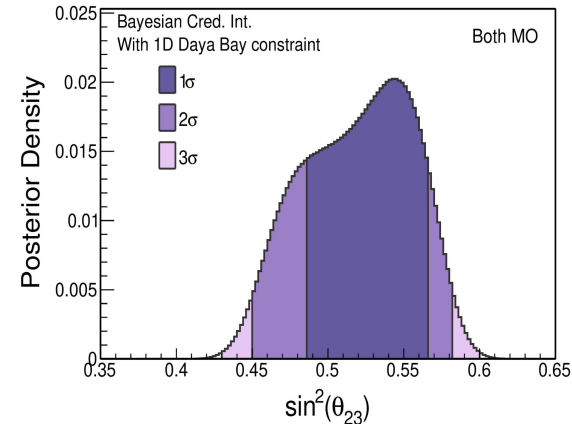
NOvA Preliminary



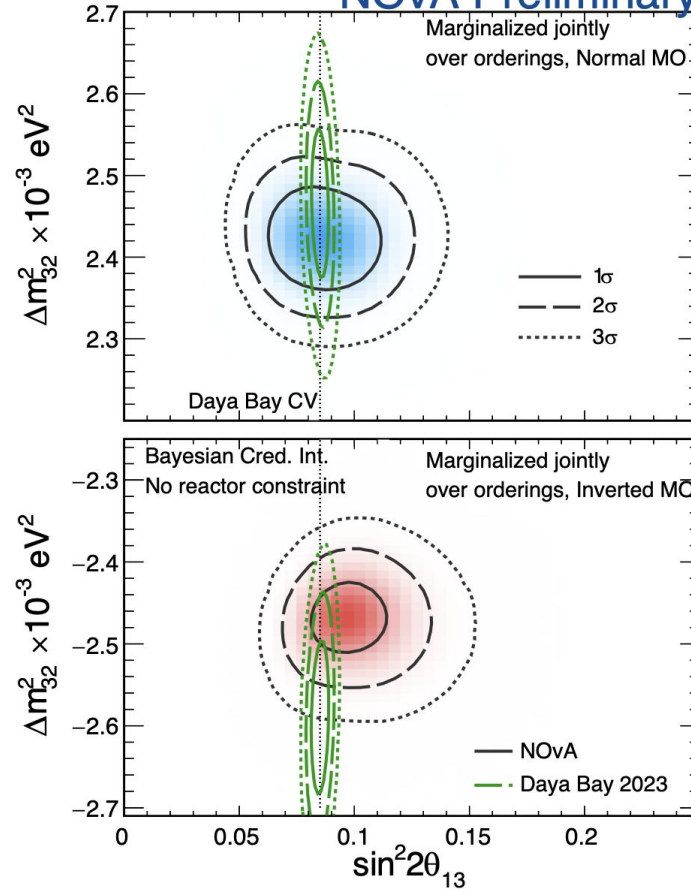
NOvA Preliminary



NOvA Preliminary

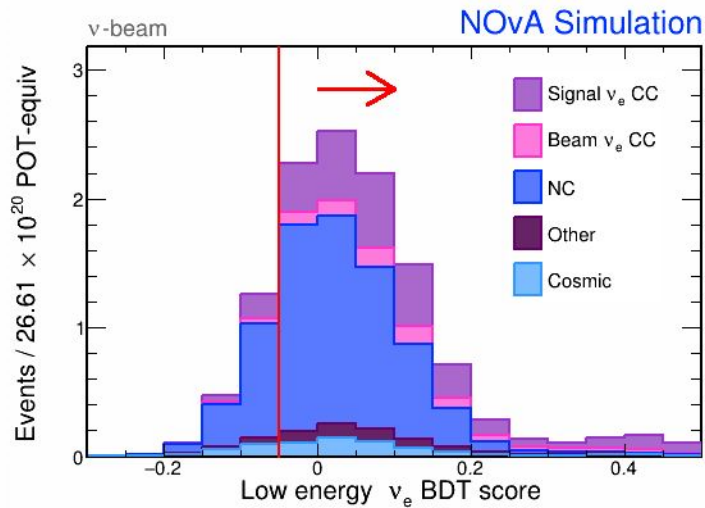
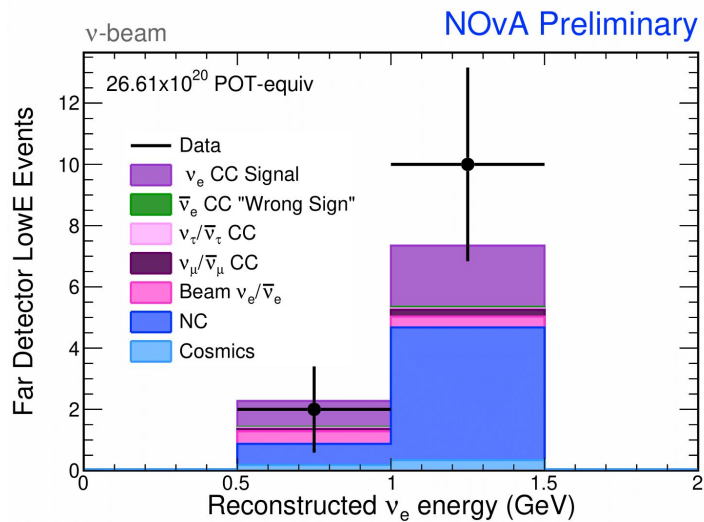


# 2D Daya Bay



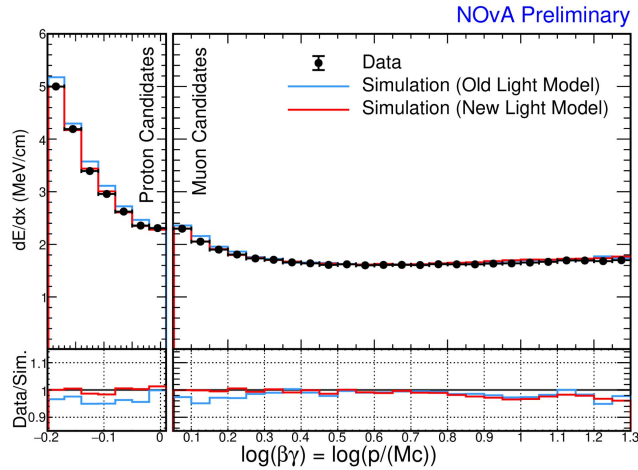
Daya Bay  
2D ( $\Delta m_{32}^2, \theta_{13}$ )  
constraint  
(PRL 130, 161802)

# Low E Sample



# Detector Characterization

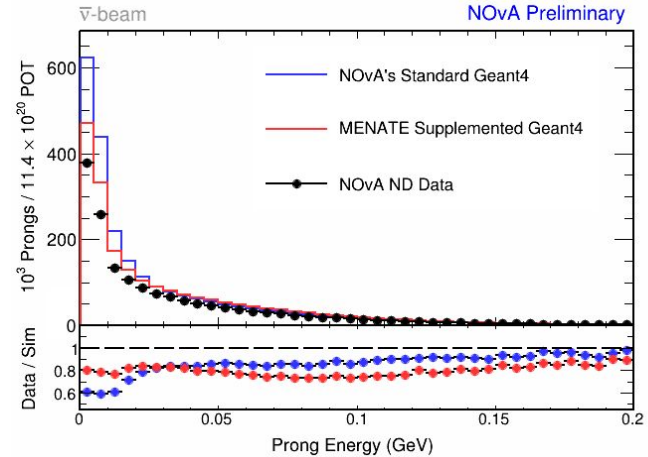
## Improved Light Production Model



Better modeling of **Cherenkov and Scintillation light in both ND and FD.**

Dedicated bench measurements & in situ stopping muon and proton tracks

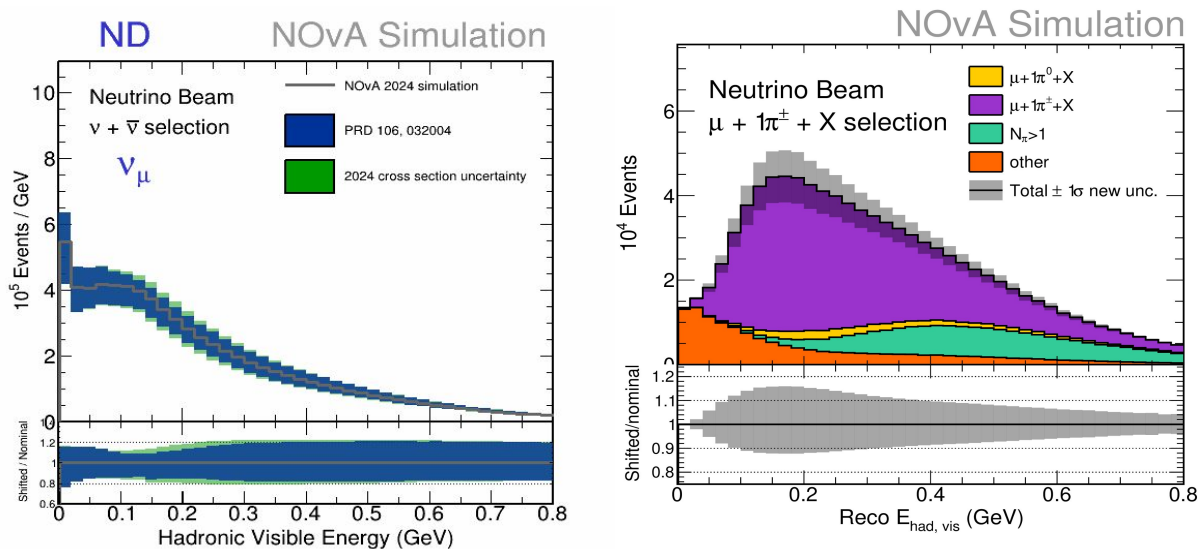
## Improved n-C Scattering Model



Difference between MENATE\_R\* and default Geant4.10.4 informs systematic uncertainty.

# Cross Section Modeling

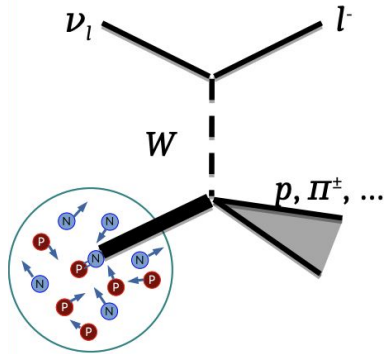
## Additional Systematic Uncertainties for Pion Production



**New cross section uncertainties on resonance and deep inelastic scattering interactions.** Increases uncertainty on pion production around peak hadronic visible energy; modest impact on overall CC Inclusive selections.

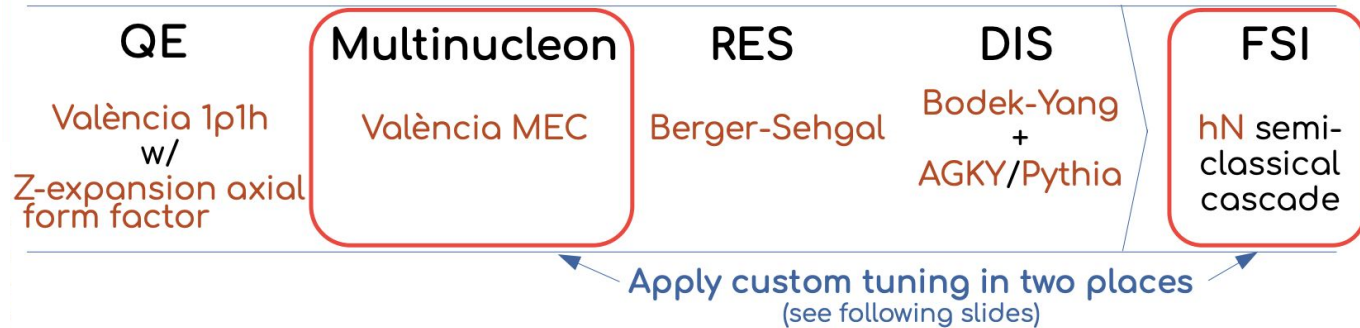


# Cross section model



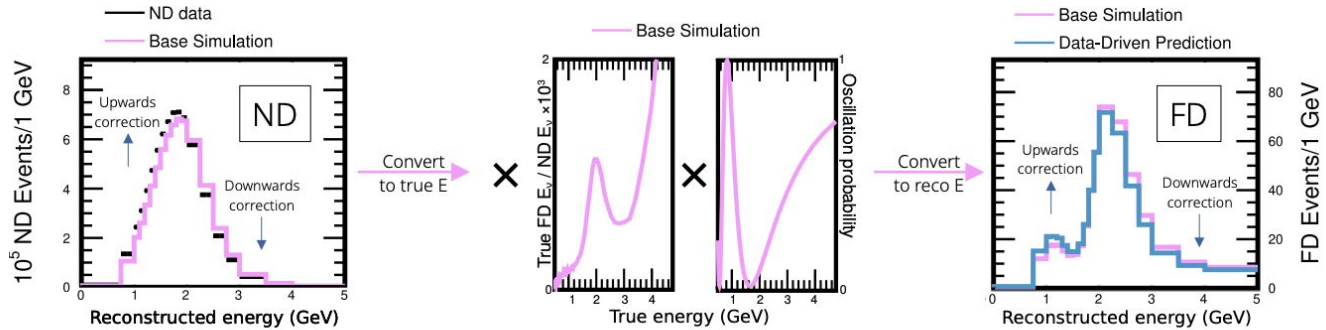
Base simulation: GENIE 3.0.6

- No stock comprehensive model configuration (CMC) agrees well with data
- We choose a “theory-driven” set of models\* and make *post hoc* adjustments to improve agreement



\* We call our model collection **N1810j\_00\_000**. It is built by starting with GENIE's **G1810b\_00\_000** and substituting the Z-expansion QE axial form factor for the dipole one. This combination was not available in the 3.0.6 release, but it may be available in future versions.

# Constraining predictions



Correcting ND simulation  
to agree with data in reco  $E_\nu$ ...

... via Far/Near transformation that  
comprises well understood effects  
(beam divergence, detector  
acceptance) + oscillations

... results in constrained  
FD  $E_\nu$  prediction highly correlated  
with ND correction

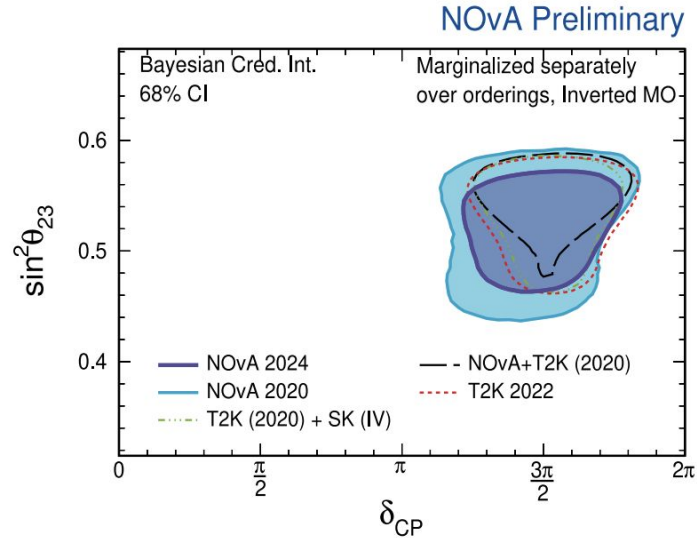
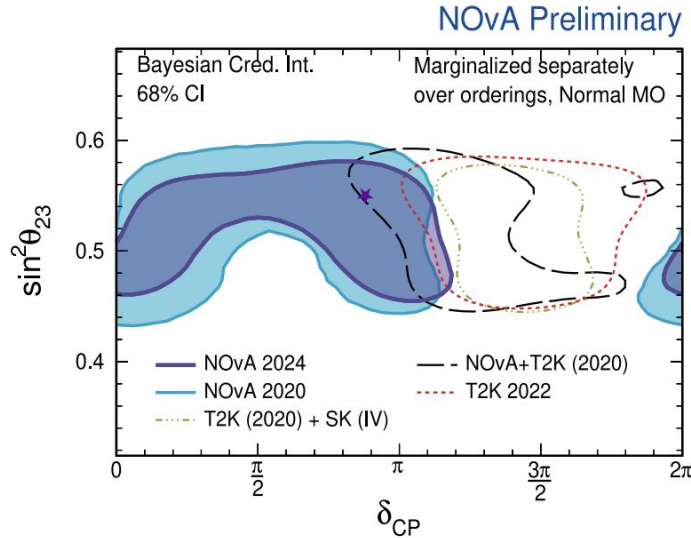
## Constrain nominal prediction and effect of systematic uncertainties

- Shift *all elements of sim.*, then redo constraint
- Since post-correction all variations forced to agree at ND, spread at FD is reduced
- Effects that are not shared between detectors unaffected, or increase in some cases (e.g.: calibration)

## Subdivide or use different samples to better account for ND/FD differences:

- $\nu_\mu$ : differences in resolution, acceptance  
subdivide by  $E_{had}/E_\nu \times |p_\tau|$  [12 bins]
- $\nu_e$  bknds: different oscillation behavior  
constrain vs' parent  $\pi$  and K (beam  $\nu_e$ );  
subdivide by Michel electron multiplicity ( $\nu_\mu$ , NC)

# Comparisons to NOvA+T2K

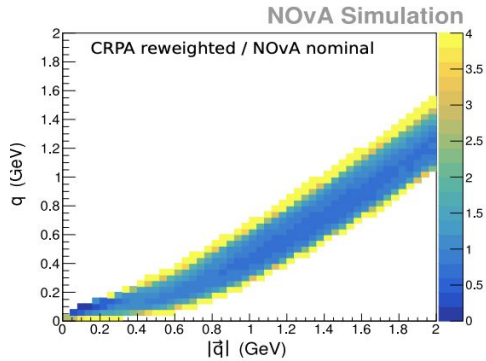


NOvA vs. other data favor different regions in NO,  
same region in IO

Note: results use different choices of reactor constraint

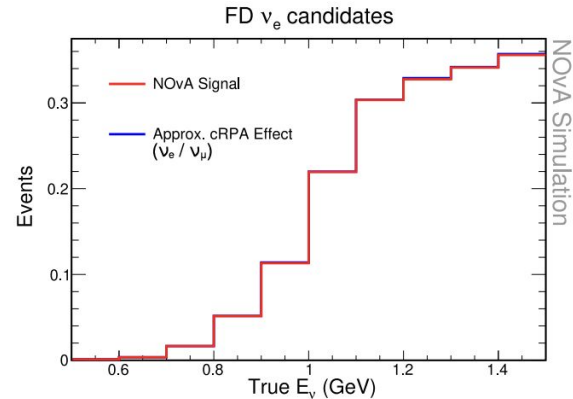
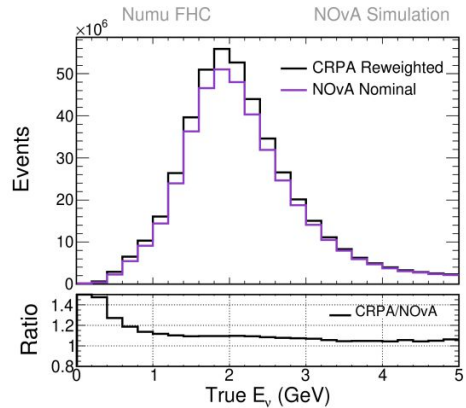
NOvA 2020: 2019 PDG avg  $\theta_{13}$     T2K: 2019 PDG avg  $\theta_{13}$   
 NOvA 2024: Daya Bay 2023 1D  $\theta_{13}$     NOvA+T2K: Daya Bay 2023 1D  $\theta_{13}$   
 T2K+SK: 2019 PDG avg  $\theta_{13}$

# Alternate CCQE model: cRPA (1)



- Continuum Random Phase Approximation (CRPA) CCQE model\*
  - Improved treatment of low-momentum-transfer nuclear dynamics
  - Opens additional phase space at edges of kinematically allowed region relative to base NOvA model (València)
  - Affects lowest neutrino energies most
- Studied impact on NOvA samples
  - Generated alternate sample using GENIE 3.4 implementation†
  - Reweighted NOvA events using ratio to base GENIE prediction
  - Found effect of  $\nu_e/\nu_\mu$  difference to be negligible on NOvA samples, so used weights for  $\nu_\mu(\bar{\nu}_\mu)$  everywhere

\*Phys. Rev. C92, 024606 †Phys. Rev. D106, 073001



# Alternate CCQE model: cRPA (2)

- Tested impact on NOvA analysis with fake data
  - Compared extrapolated CRPA prediction to extrapolated nominal
  - Spectral impact is well within extrapolated uncertainty budget
    - Impacts  $\nu_\mu$  quartile 1 the most, as expected (highest CCQE fraction)
    - ~Negligible impact on  $\nu_e$  samples (except in LowE, where stat uncertainty is enormous anyway)
  - Performed fits using extrapolated CRPA prediction as fake data
- Overall analysis impact is very small
  - $\Delta m^2_{32}$ : resulting bias  $\sim 0.1\%$  ( $\sim 7\%$  of  $1\sigma$  interval)
  - $\sin^2\theta_{23}$ : resulting bias  $\sim 0.4\%$

