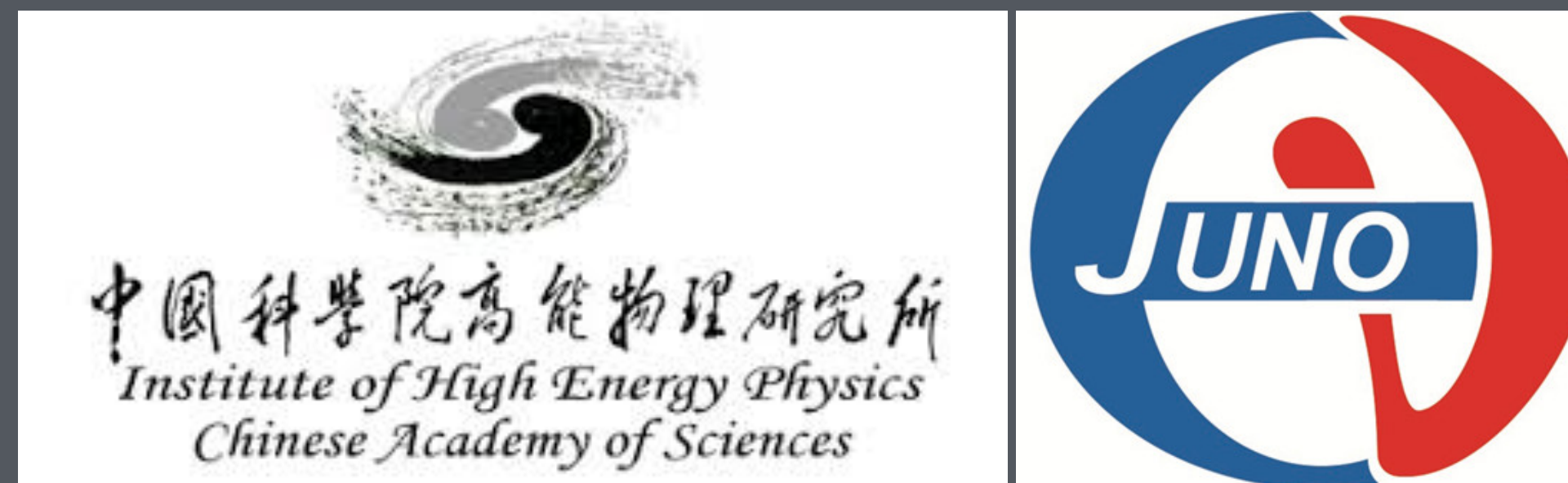


OVERVIEW OF MACHINE LEARNING APPLICATIONS IN JUNO

WUMING LUO (IHEP)
ON BEHALF OF JUNO
ICHEP 2024@PRAGE



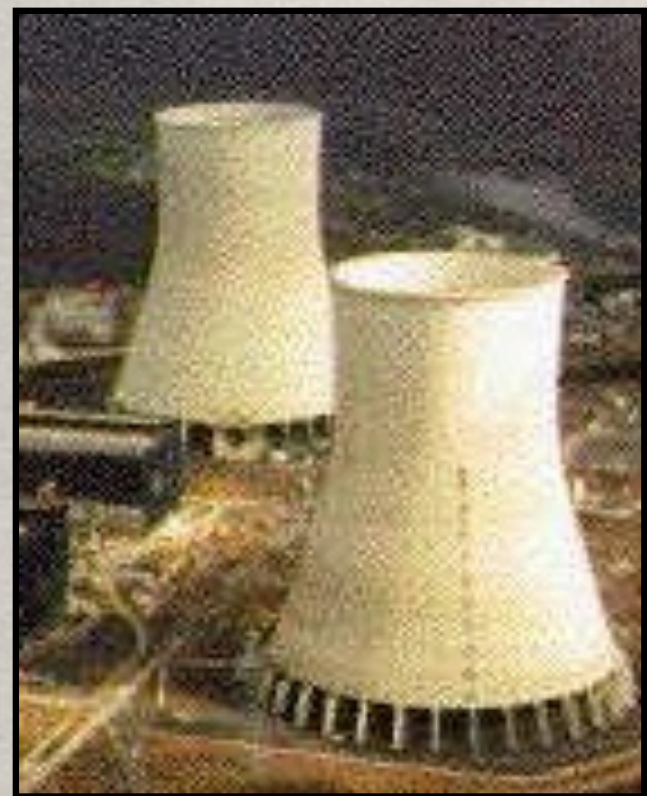
OUTLINE

- ❖ Introduction of JUNO
- ❖ Applications of Deep Learning in MeV region
 - ❖ PMT waveform reco for reactor anti-neutrinos
 - ❖ Vertex/Energy reco for reactor anti-neutrinos
- ❖ Applications of Deep Learning in GeV region
 - ❖ Directionality for atmospheric neutrinos
 - ❖ Particle Identification for atmospheric neutrinos
- ❖ Summary



JUNO

- ❖ Jiangmen Underground Neutrino Observatory(JUNO):
- ❖ Determine the neutrino mass ordering
- ❖ Measure neutrino oscillation parameters to sub-percent level
- ❖ SuperNova, Solar, Atm. Geo. etc



Wuming Luo

	DETECTOR TARGET MASS	ENERGY RESOLUTION
KamLAND	1000 t	6%@1MeV
D. Chooz	8+22 t	8%@1MeV
RENO	16 t	
Daya Bay	20 t	
Borexino	300 t	5%@1MeV
JUNO	20000 t	3%@1MeV

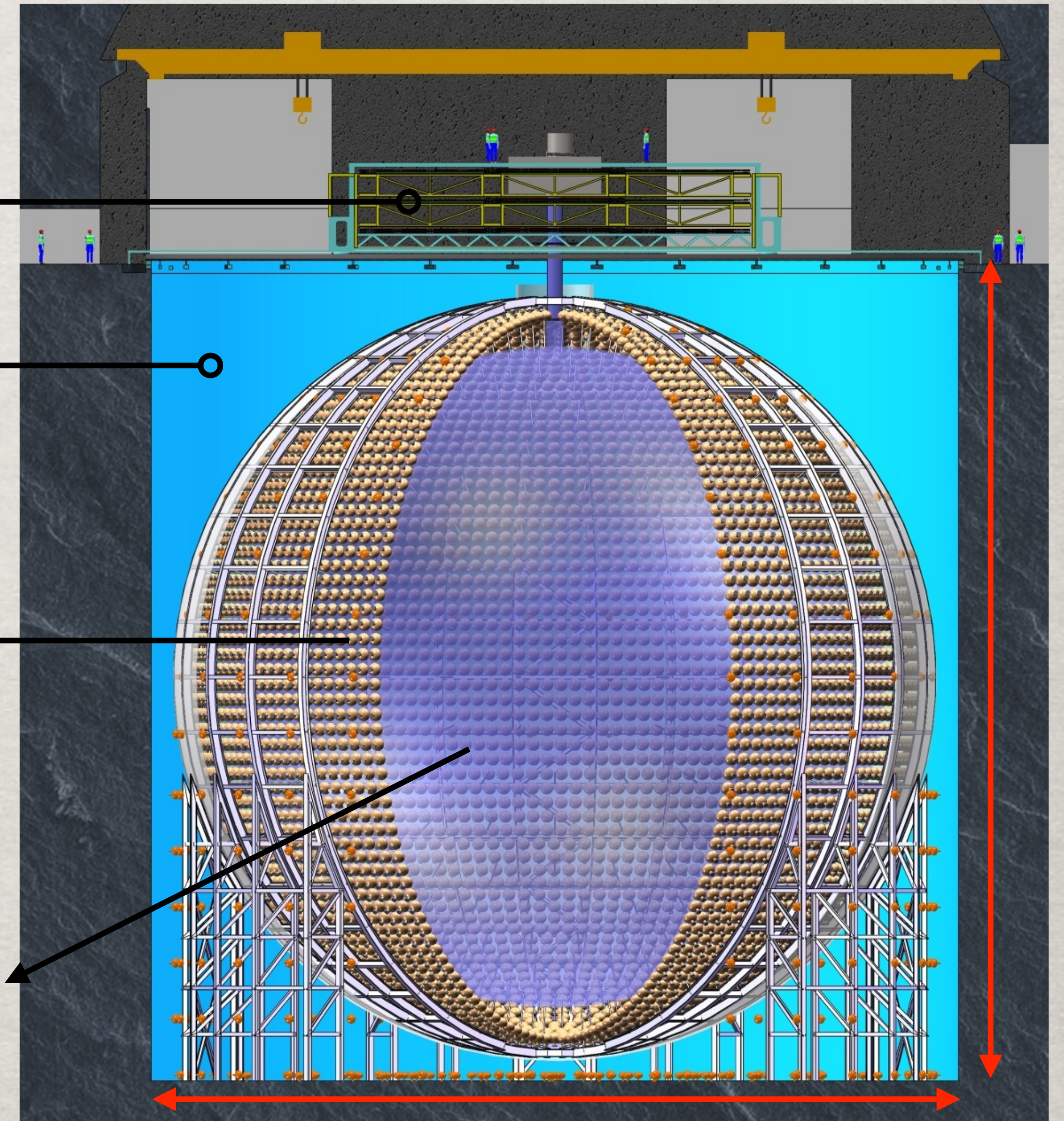
Top Tracker

Water Pool

Central Detector

~17,612 20" PMTs
 + ~25,600 3" PMTs
 + ~78% coverage

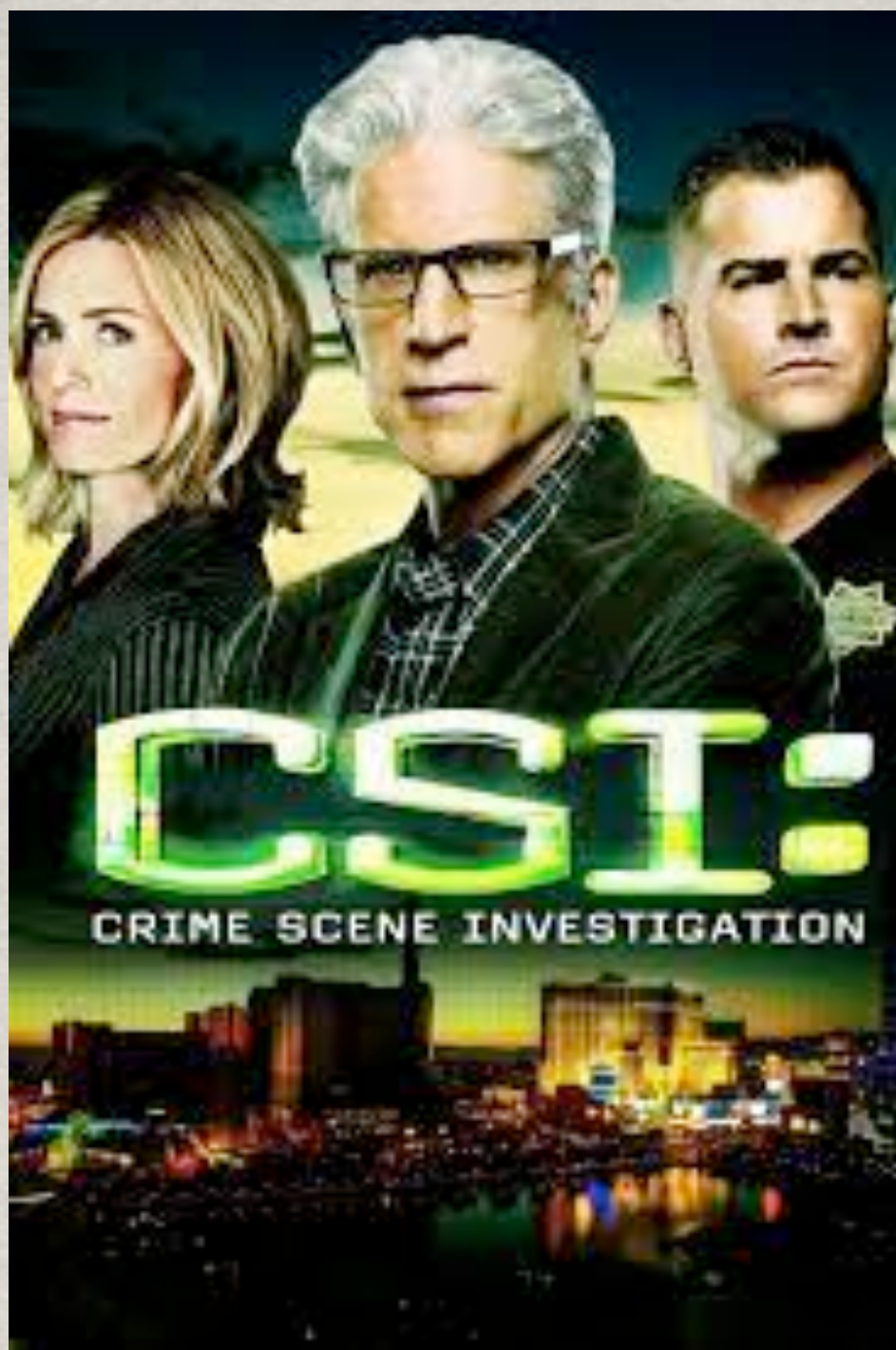
Liquid Scintillator
 20kton



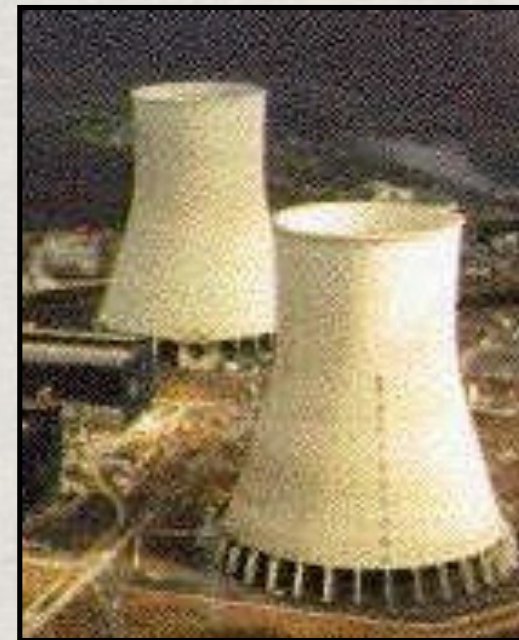
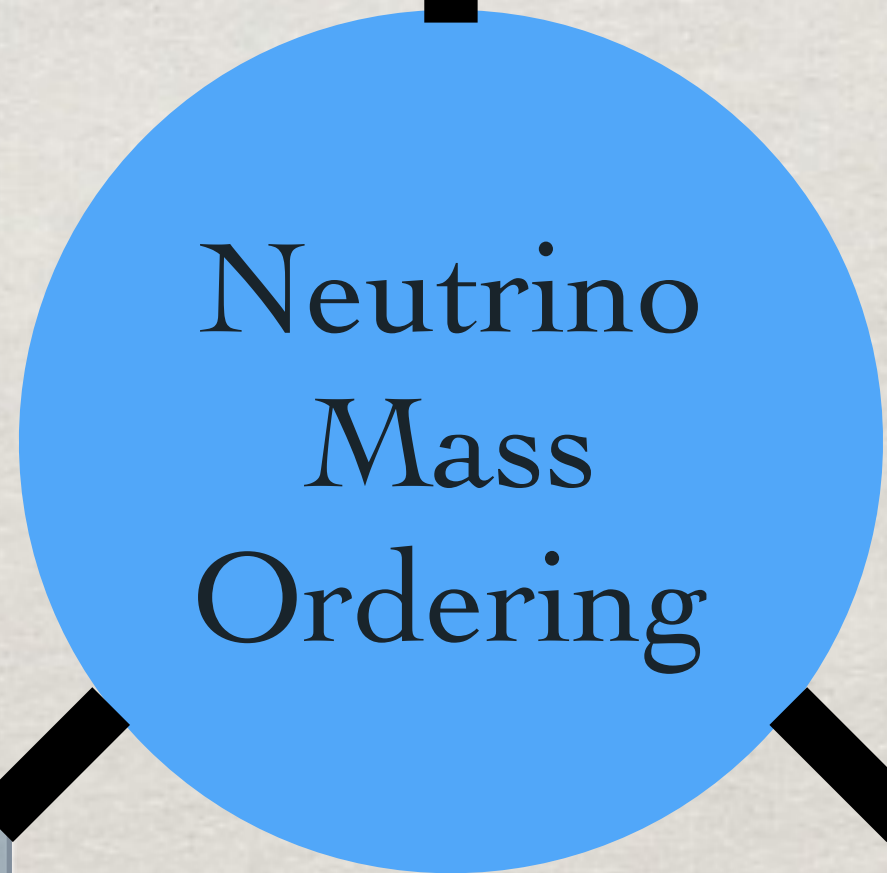
$\phi: 43.5\text{m}$

Depth: 44m





RECO IN JUNO



reactor $\bar{\nu}$
MeV

PMT waveform

energy

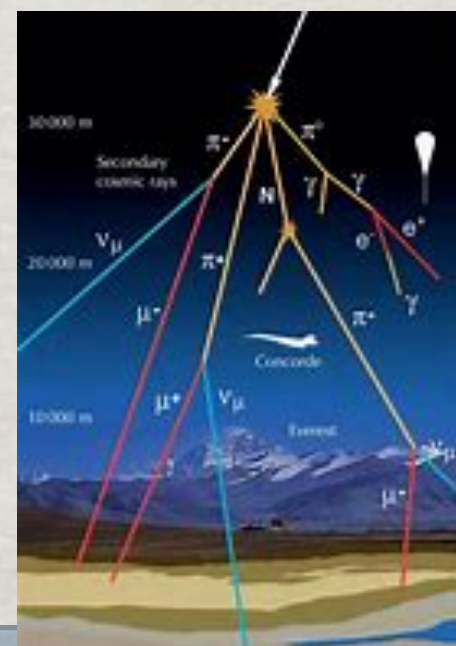
vertex

non-uniformity

PMT dark noise

PMT charge smear

^{14}C pileUp



atm. ν
GeV

cosmic μ
GeV

muon classification

multi-detector combined reco

track

shower vertex

Direction

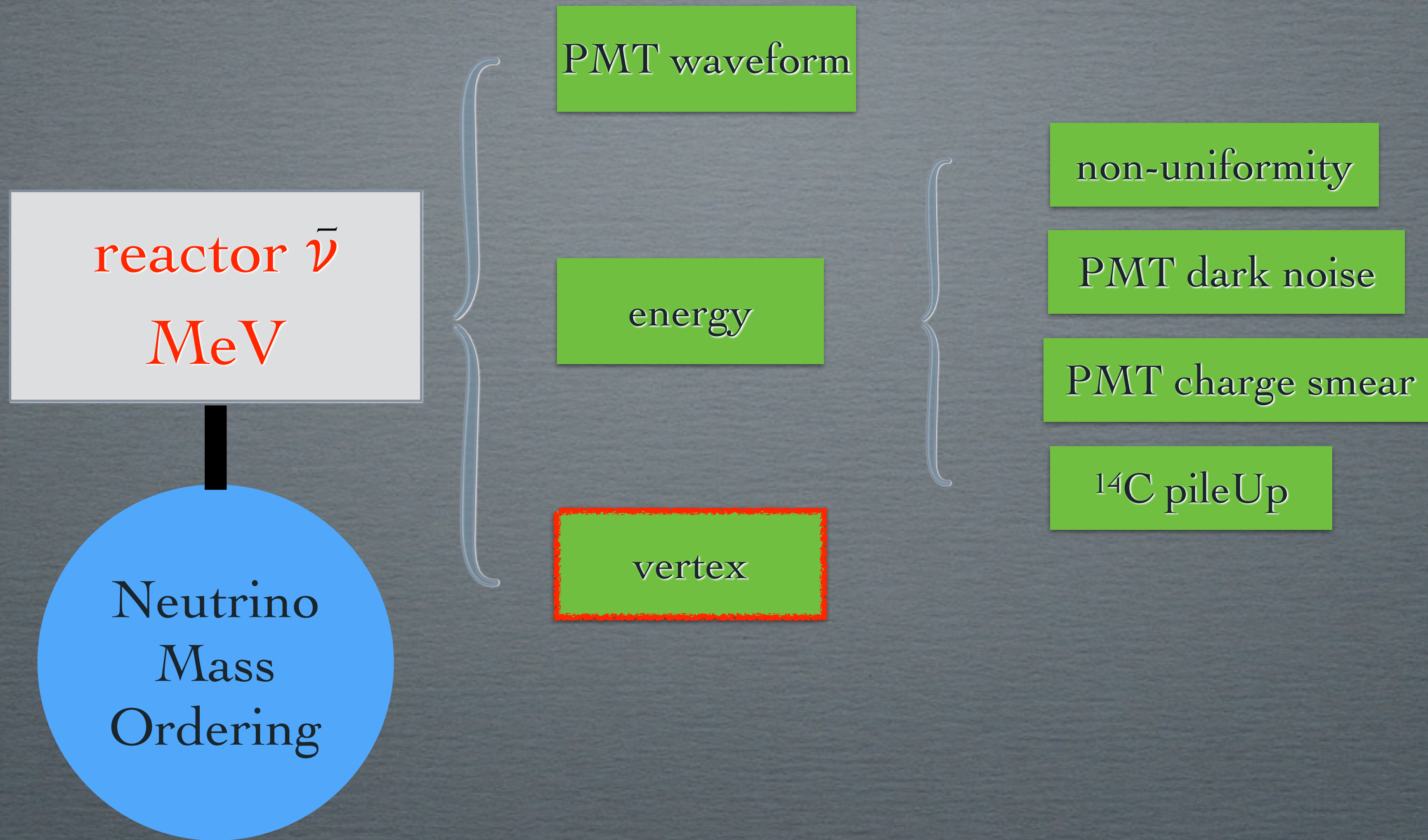
Particle ID

Energy

ν_{μ} vs ν_e vs NC

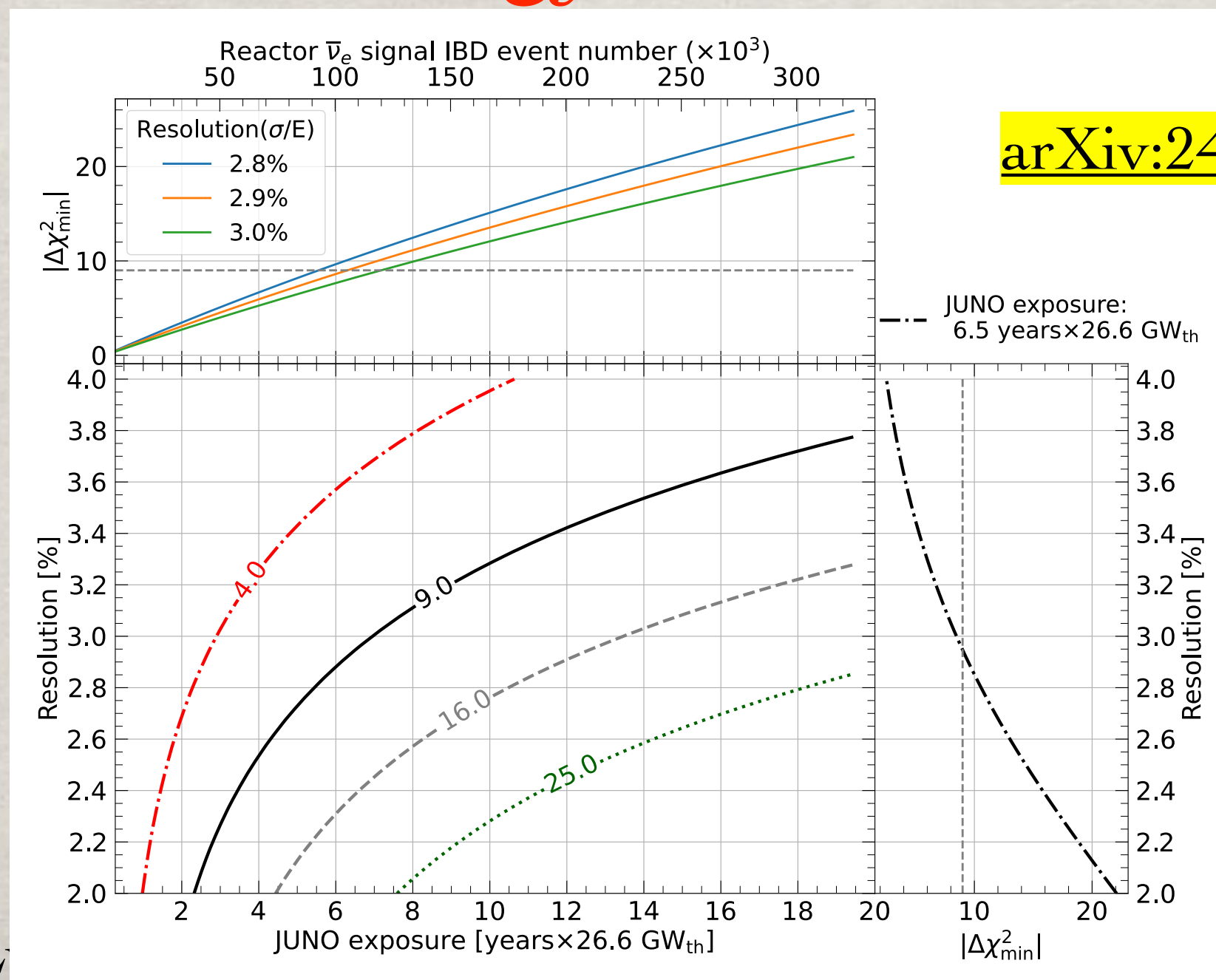
ν vs $\bar{\nu}$



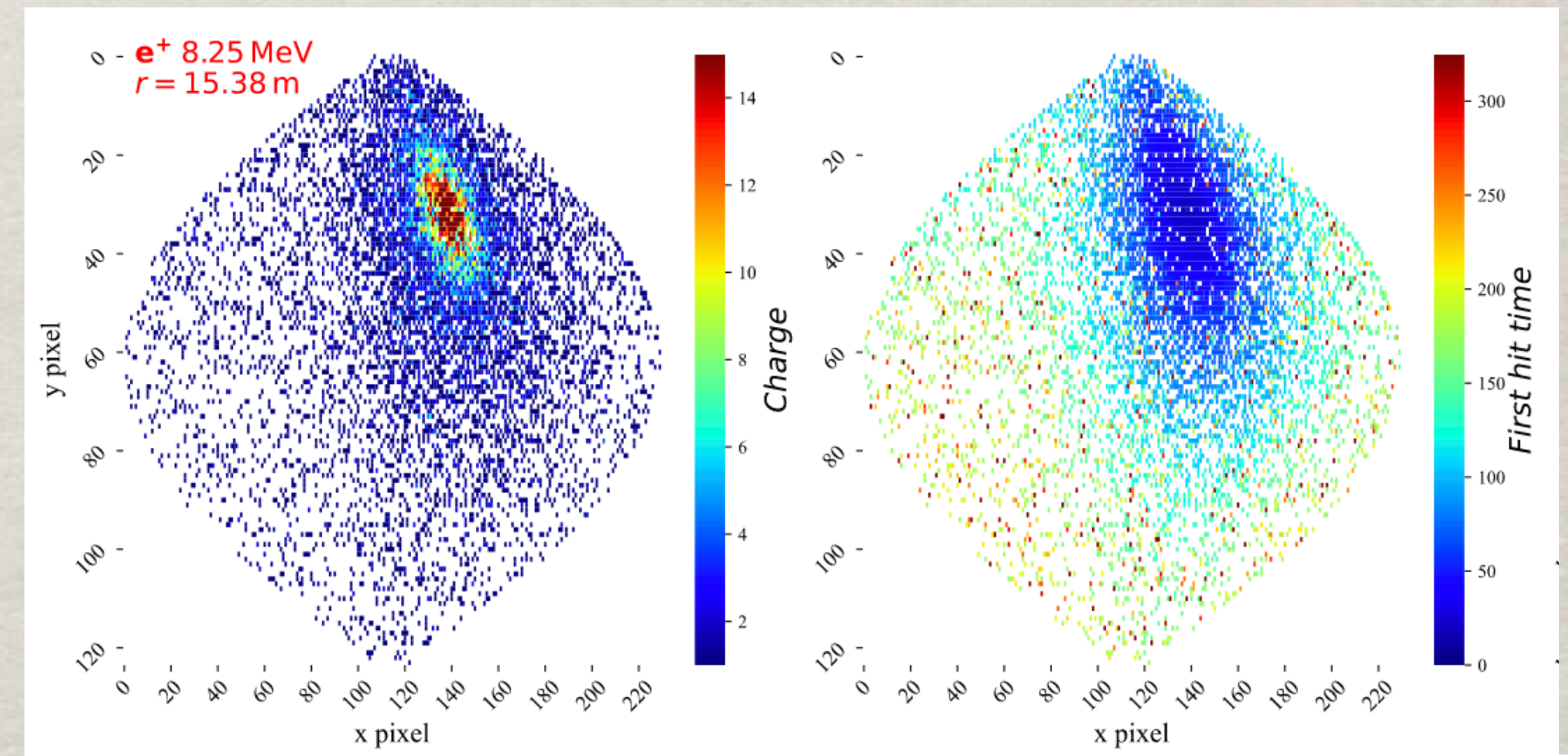


IMPORTANCE AND PRINCIPLES

- ❖ Large number of PMTs $O(10^4)$ installed on a sphere
- ❖ each PMT as a pixel \rightarrow JUNO as a Camera
- ❖ ensemble of PMTs charge/time form an image
- ❖ Image is highly vertex and energy dependent
- ❖ Vertex/energy reconstruction \leftrightarrow Image recognition



arXiv:2405.18008

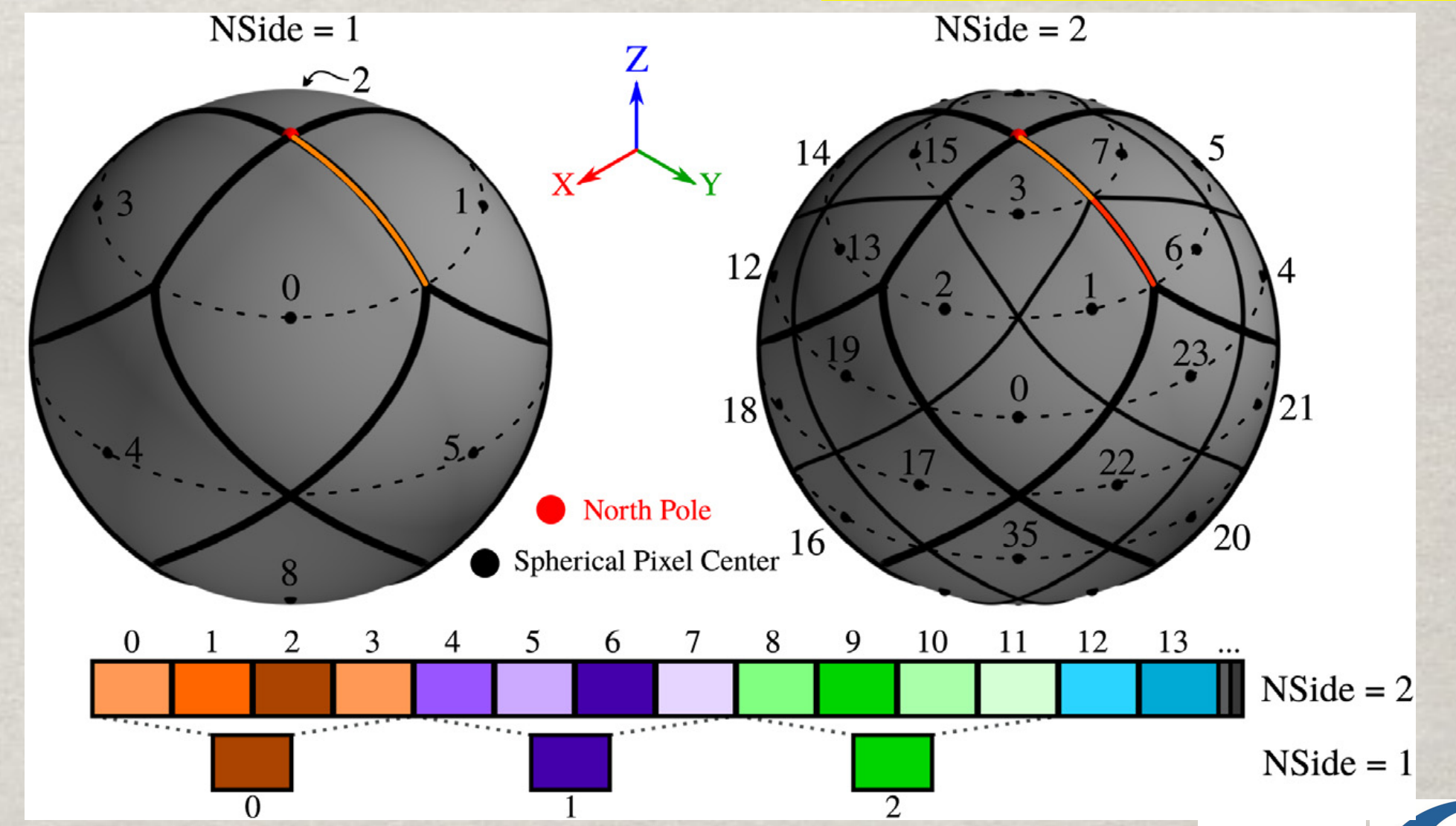
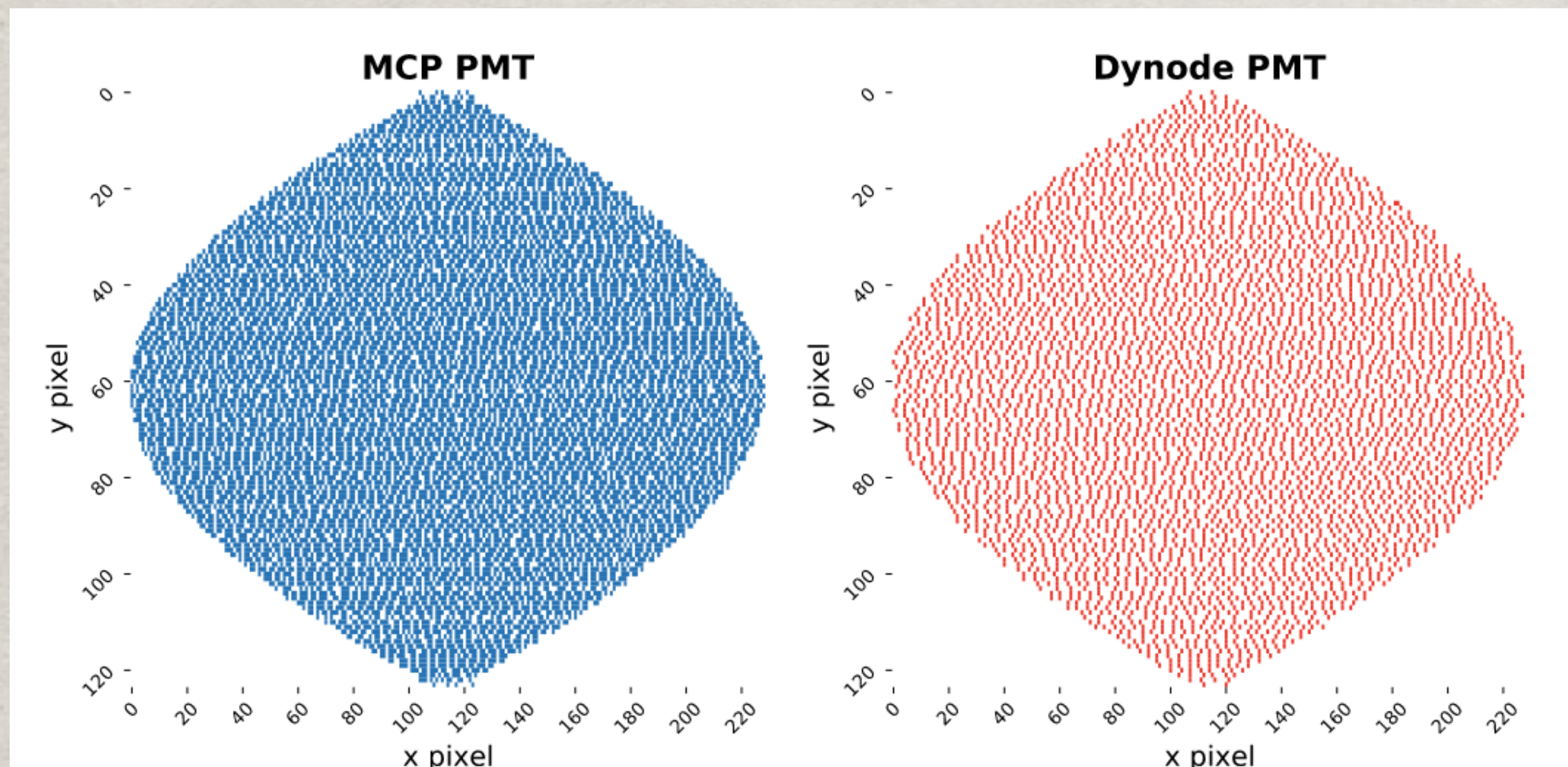


Nucl.Sci.Tech. 33 (2022) 7, 93

INPUTS

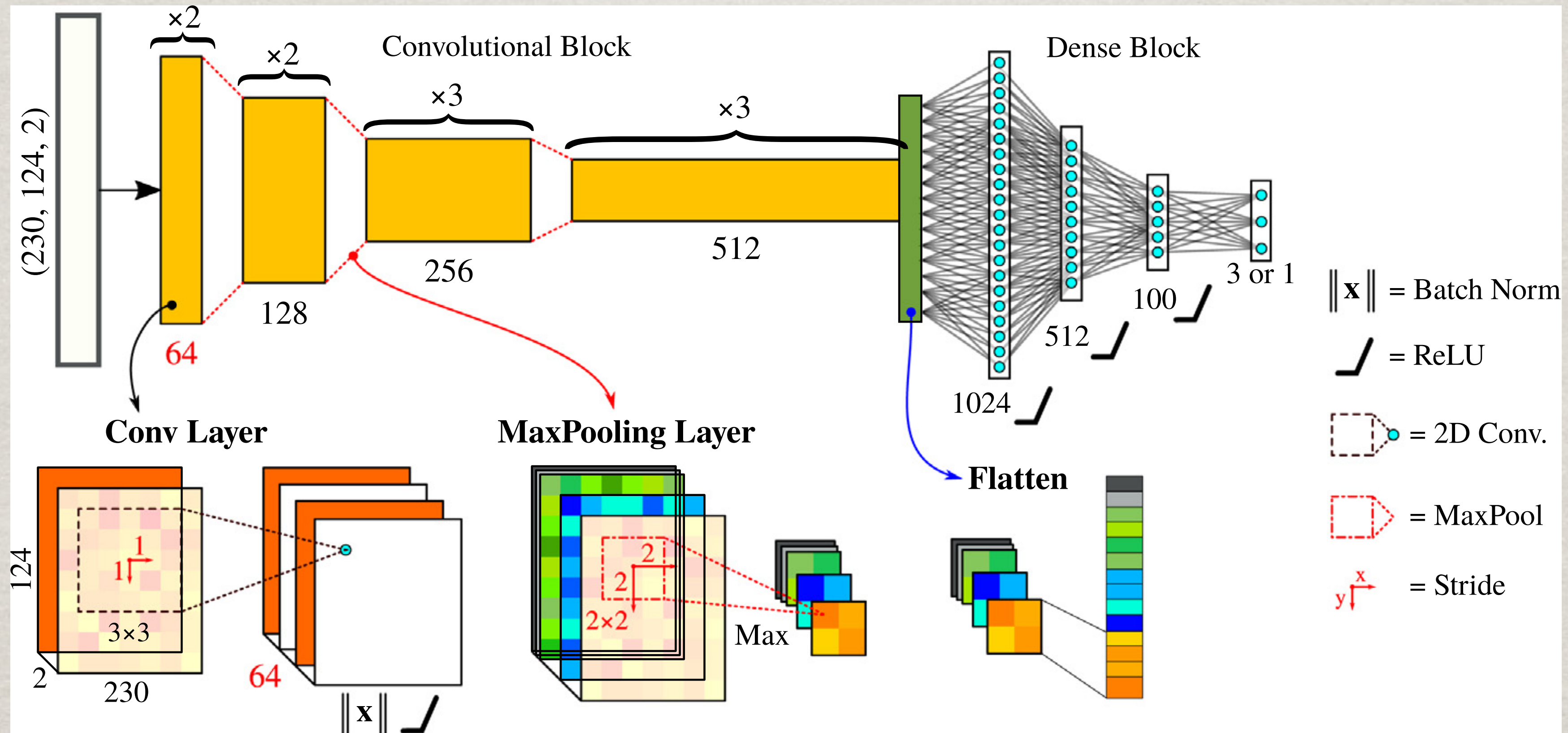
- ❖ Large number of PMTs $O(10^4)$ installed on a sphere
- ❖ Method 1: projection to 2D plane \rightarrow Plane CNN
- ❖ Method 2: HEALPix \rightarrow plane/spherical CNN
- ❖ Method 3: 3D models such as pointNet++/Transformer

NIMA 1010 (2021) 165527



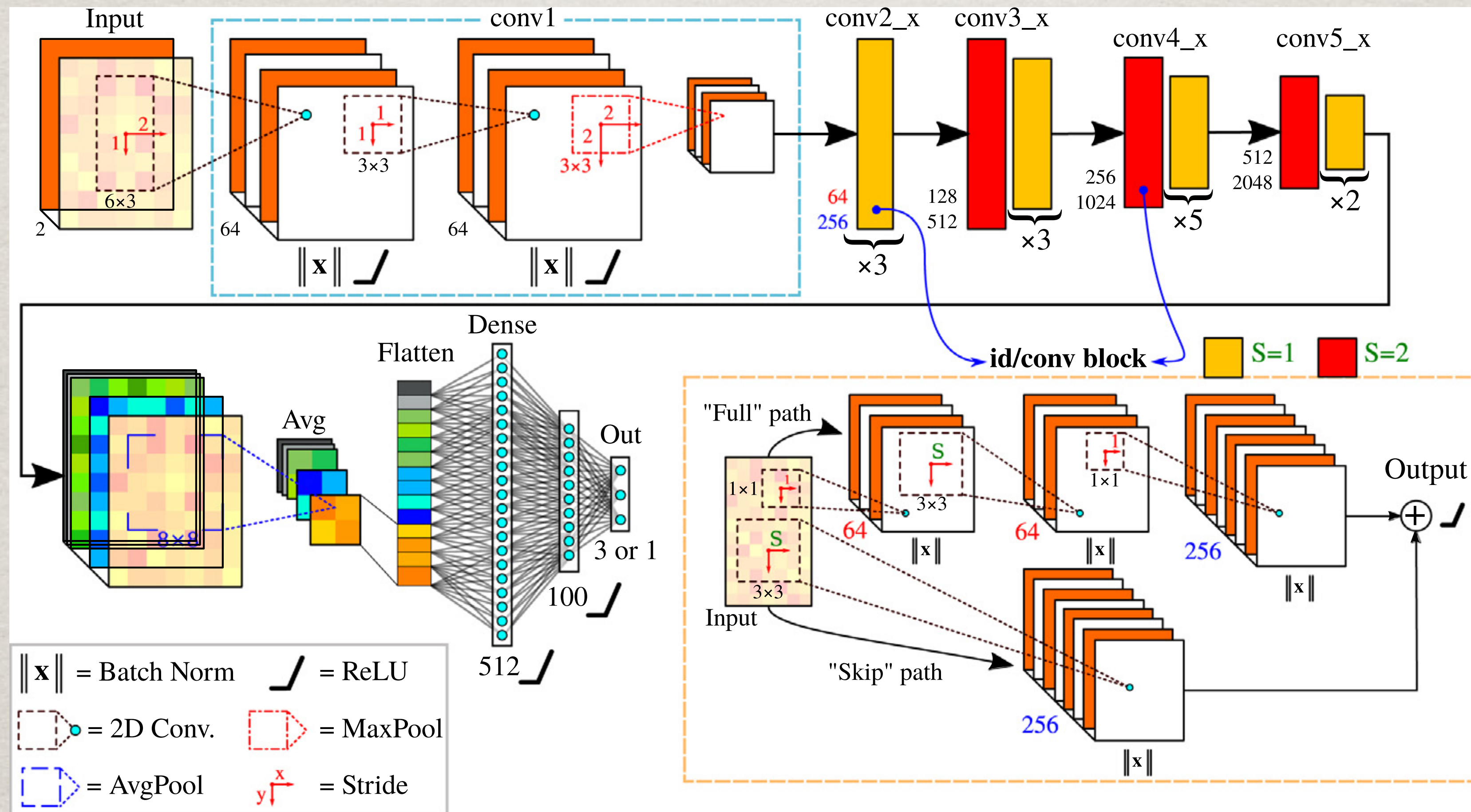
MODELS: VGG-J

NIMA 1010 (2021) 165527



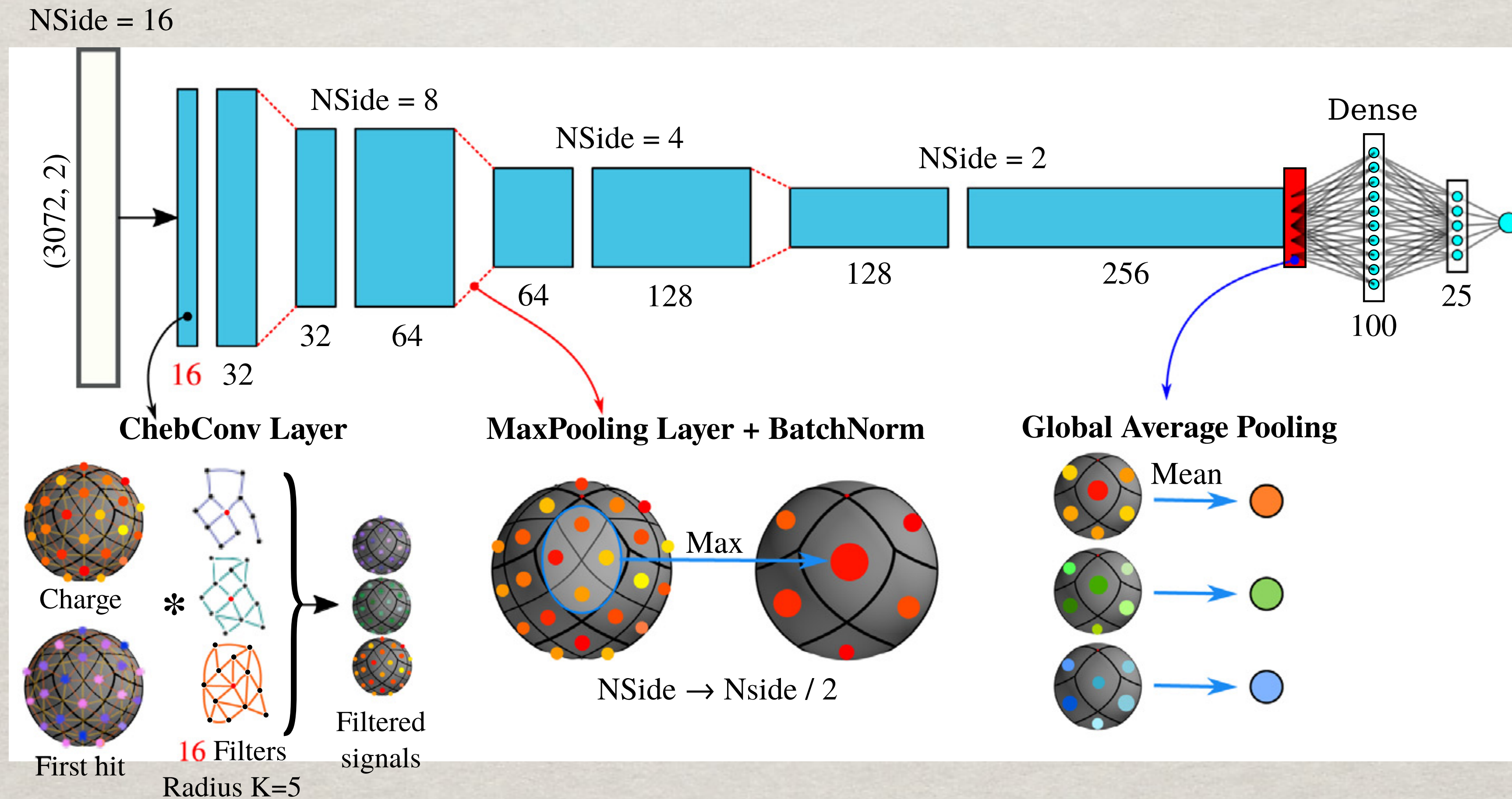
MODELS: RESNET-J

NIMA 1010 (2021) 165527



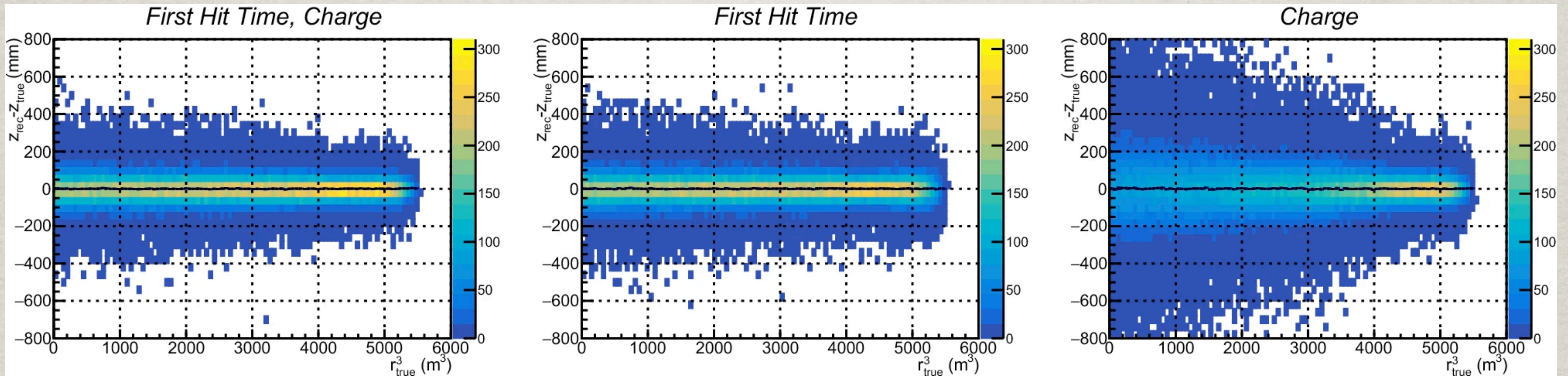
MODELS: GNN-J

NIMA 1010 (2021) 165527

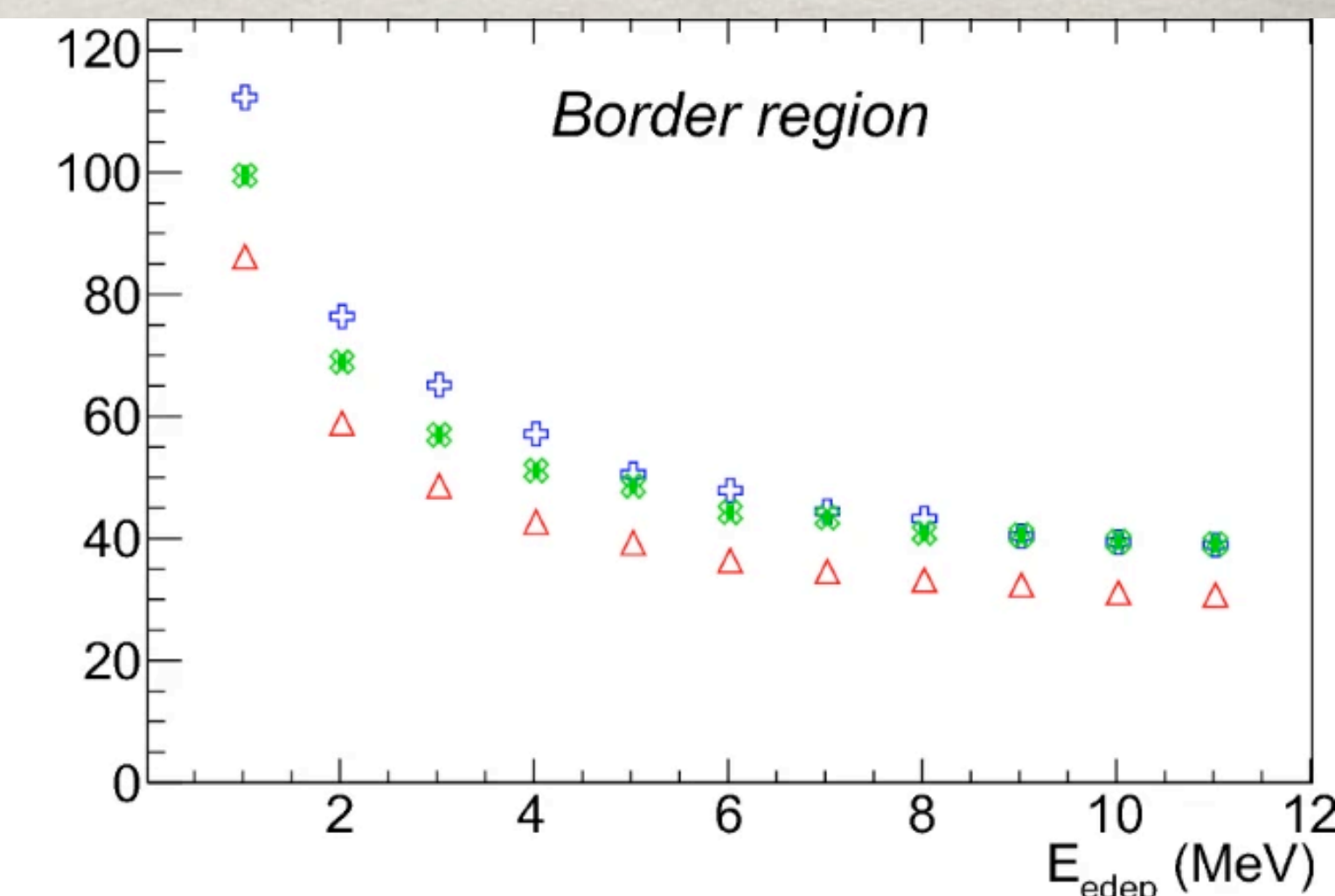
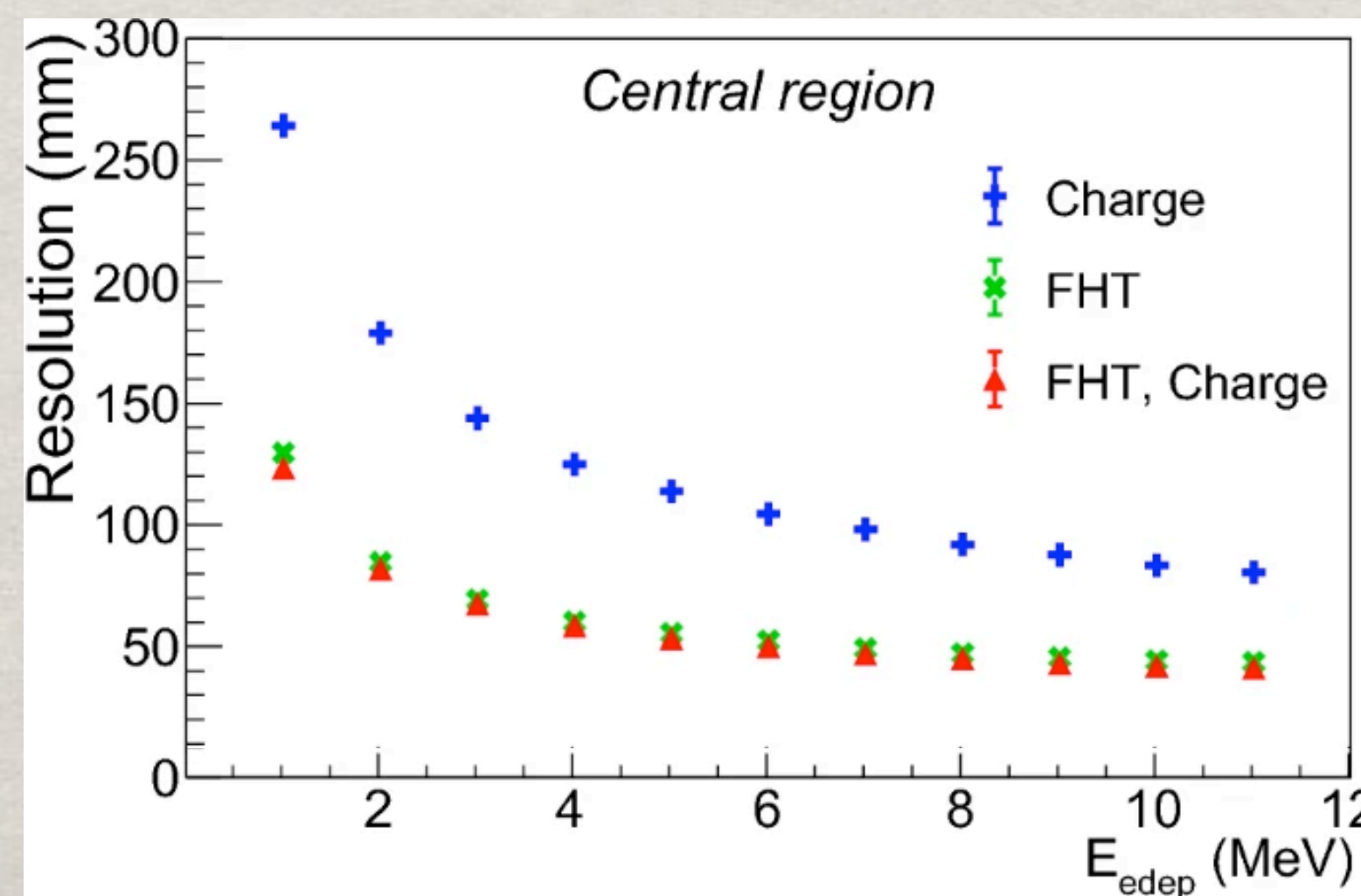


VERTEX: CHARGE VS TIME

Nucl.Sci.Tech. 33 (2022) 7, 93

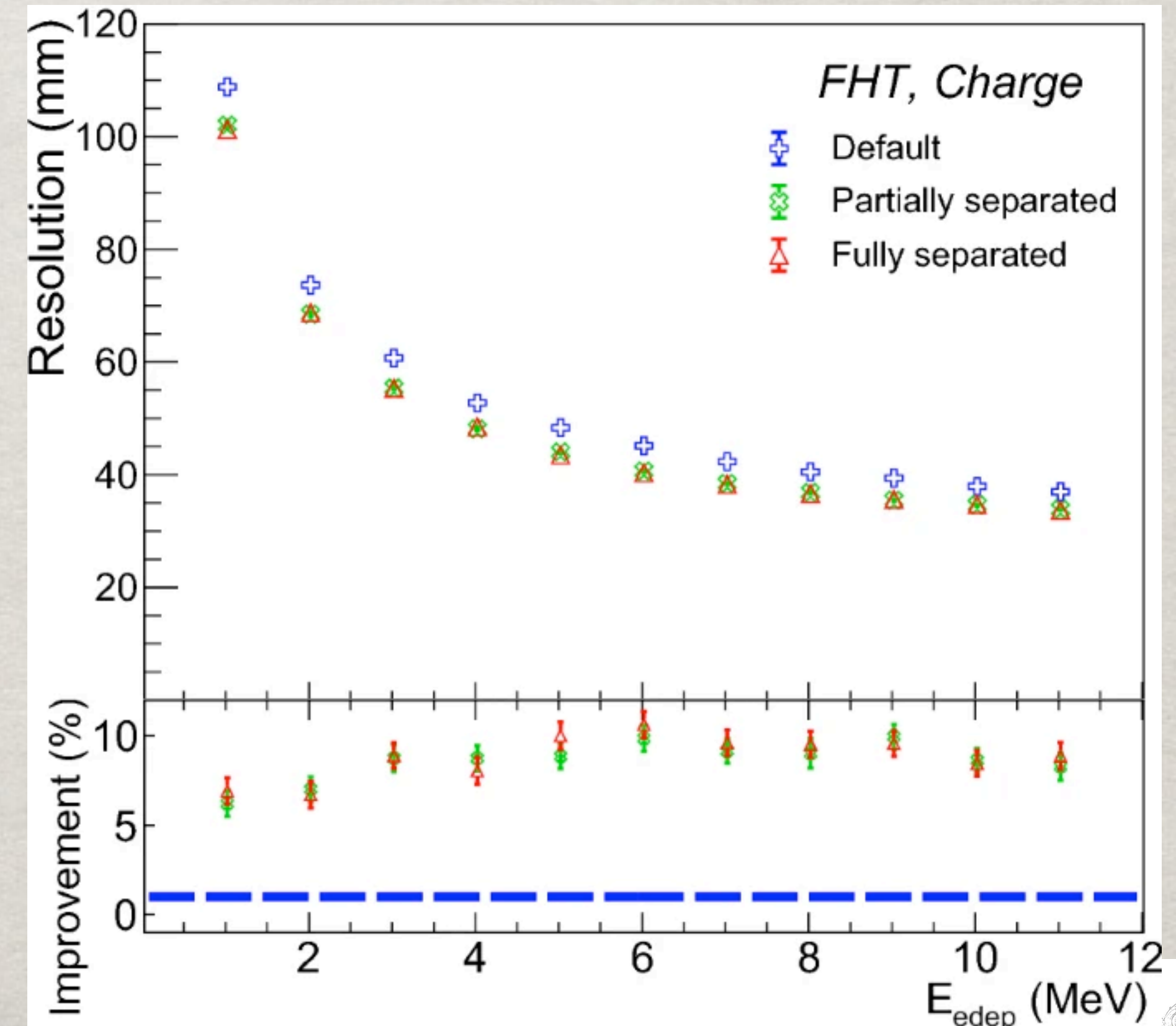
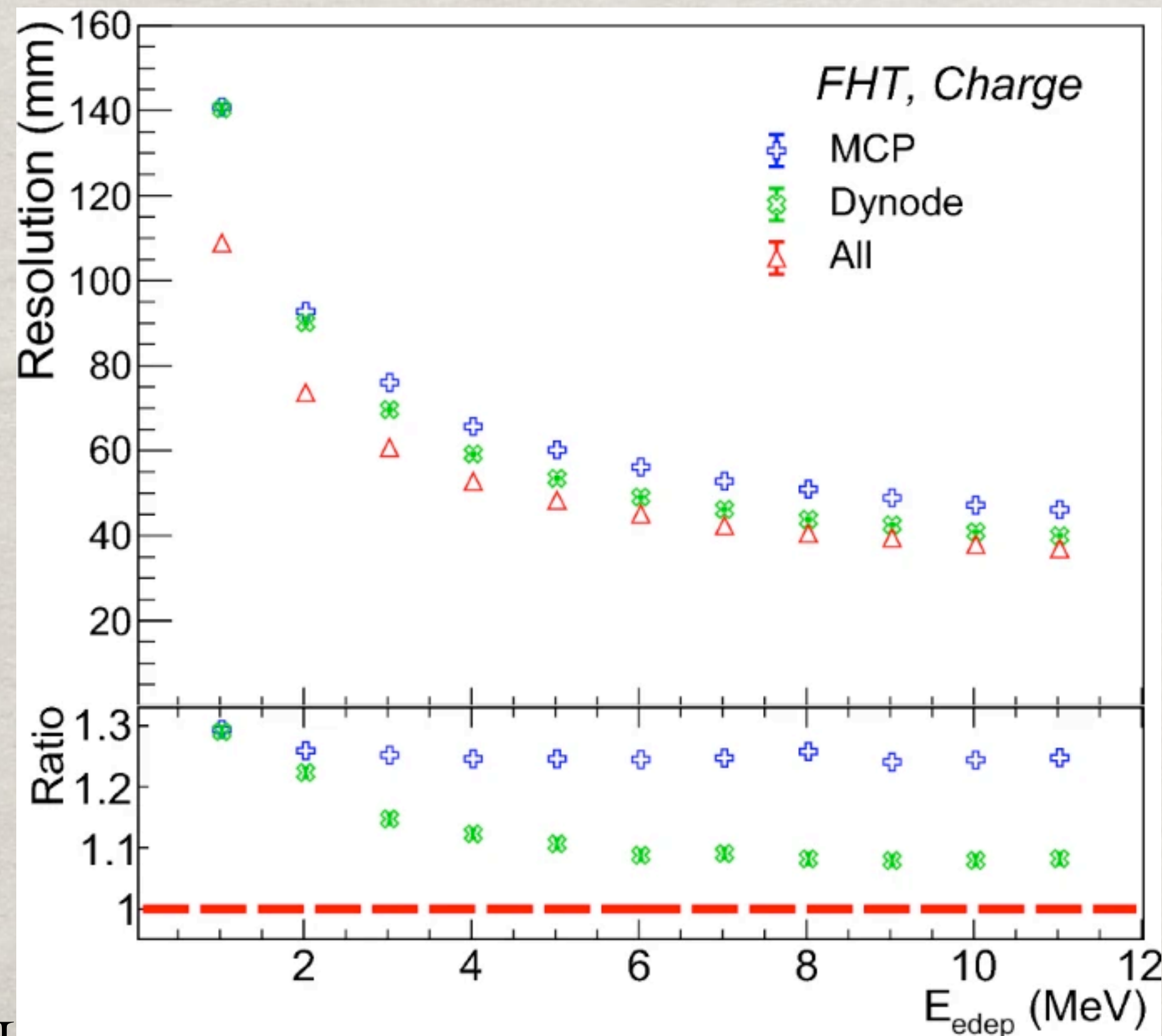


❖ Charge also helps constrain the Vertex, especially in the border region



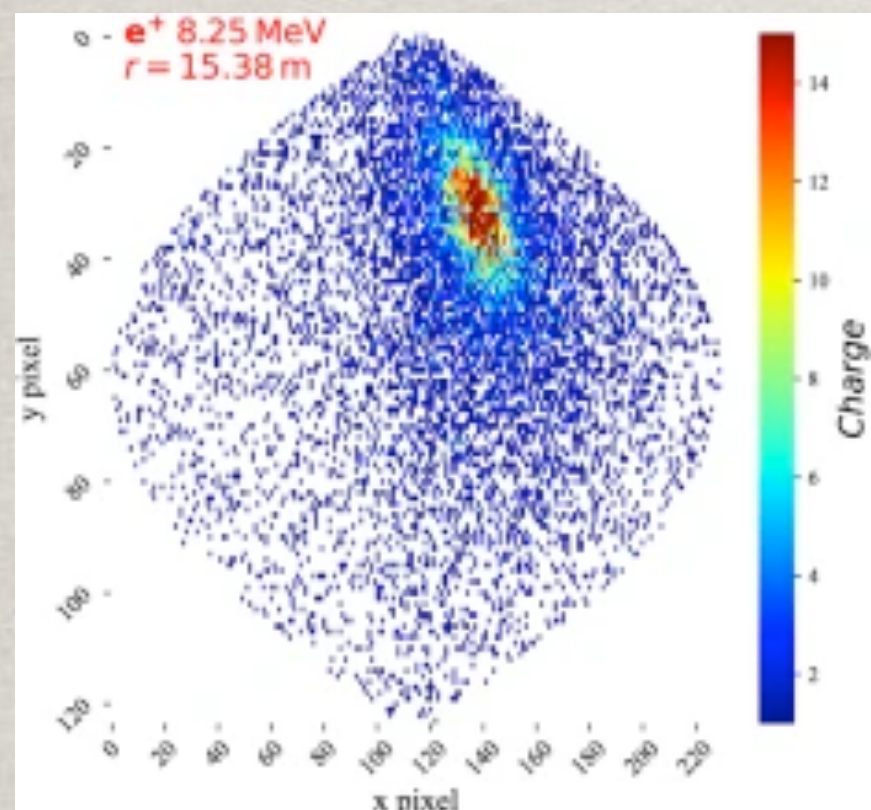
VERTEX: MCP VS DYNODE

- ❖ Comparison between MCP PMTs and Dynode PMTs Nucl.Sci.Tech. 33 (2022) 7, 93
- ❖ Number: 12612 vs 5000; Time resolution(σ_{tts}): 12 ns vs 2.8 ns
- ❖ Separating charge/time info by PMT type gives the best performance

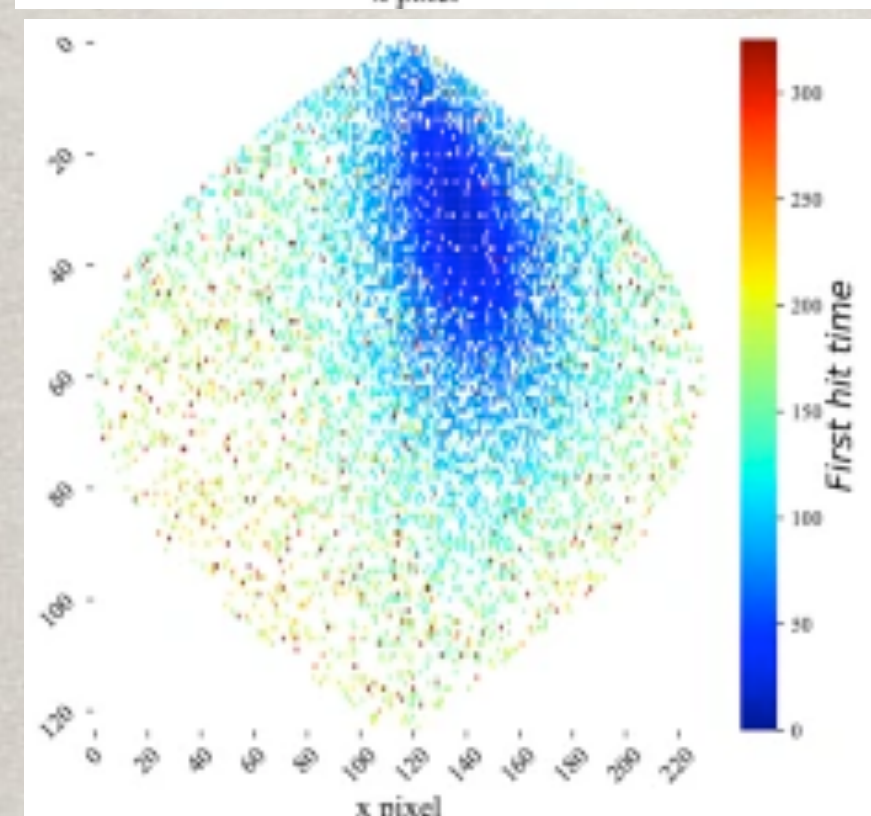


VERTEX: ADDITION OF 2ND HIT

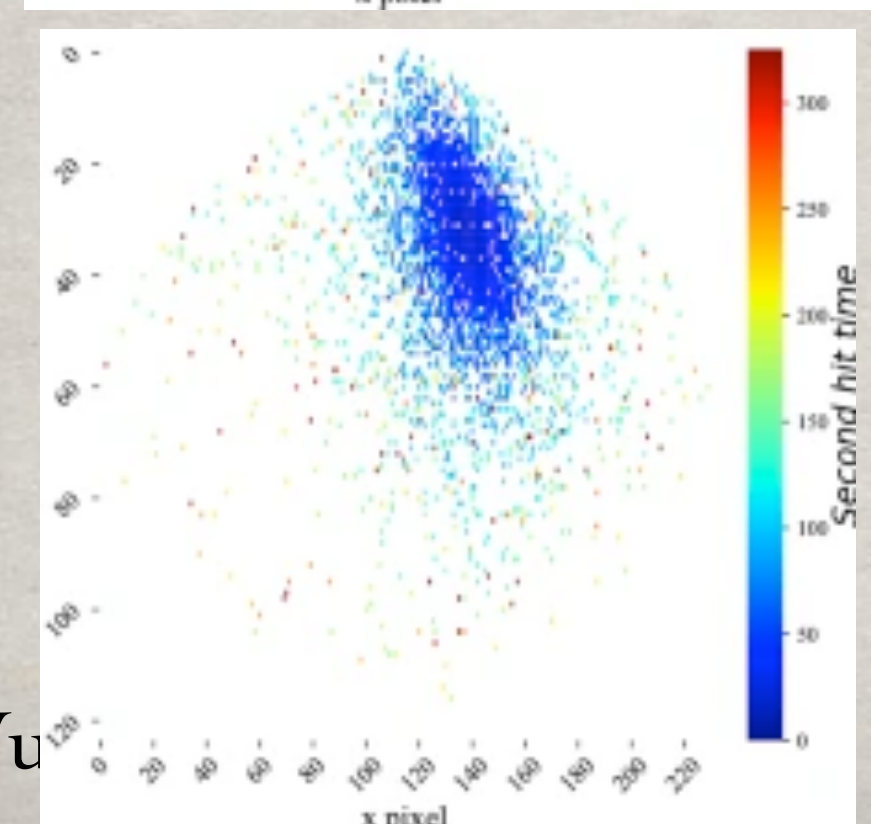
Nucl.Sci.Tech. 33 (2022) 7, 93



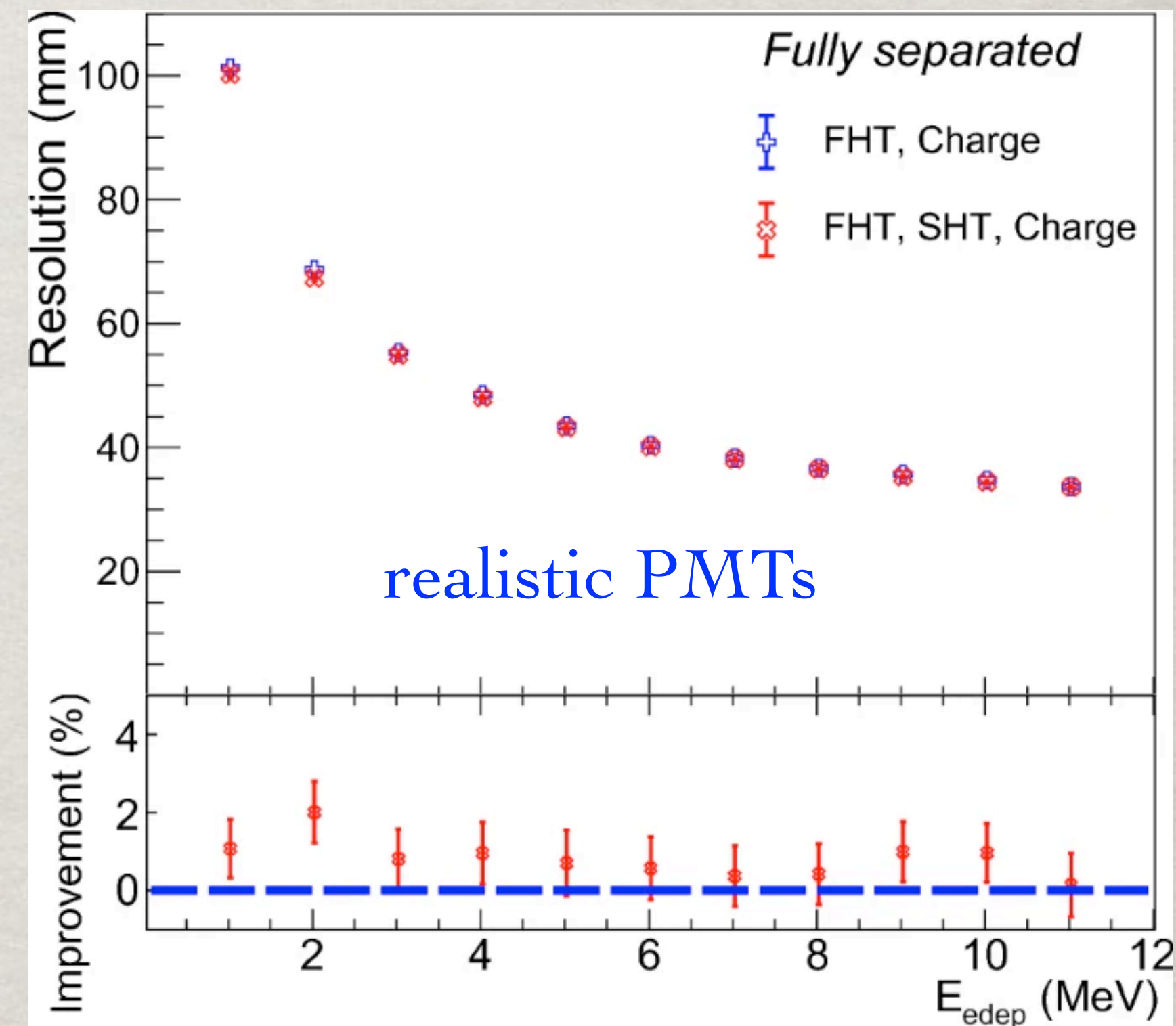
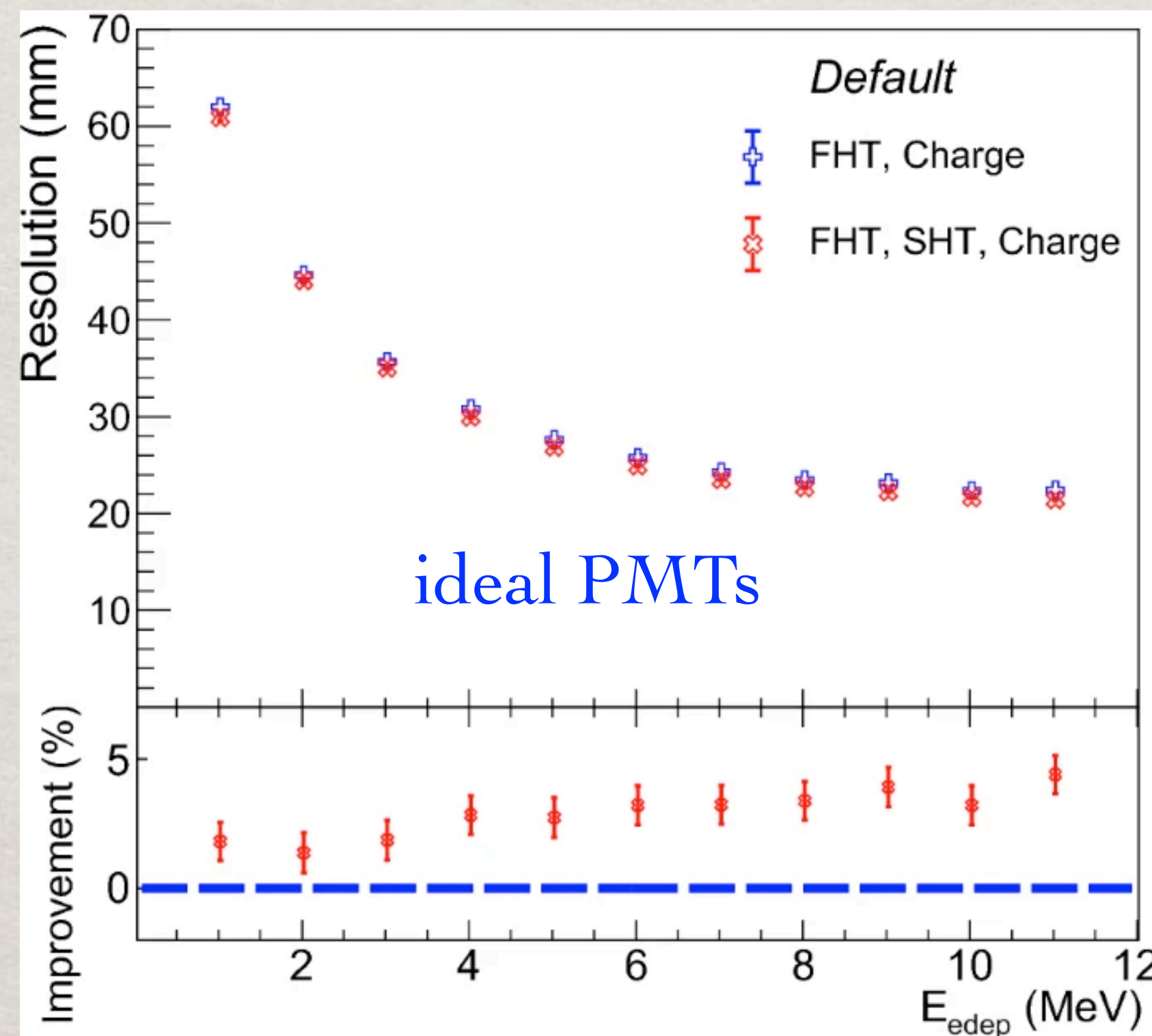
Charge



FHT:
t(1st Hit)



SHT:
t(2nd Hit)



- ❖ Later hits are also useful in principle
- ❖ PMT Time resolution is the key

ENERGY: BDT/FCDNN

Eur.Phys.J.C 82 (2022) 11, 1021

- ❖ Inputs: feature engineering
- ❖ Models: BDT or FCDNN
- ❖ Optimization:
 - ❖ hyper-parameters
 - ❖ feature selection

Notation	Description
AccumCharge	Total accumulated charge
nPMTs	Number of fired PMTs
$x_{cc}, y_{cc}, z_{cc}, R_{cc}, \theta_{cc}, \phi_{cc}, J_{cc}, \rho_{cc}, \gamma_z^{cc}, \gamma_y^{cc}, \gamma_x^{cc}$	Center of charge
$pe_{n\%}, pe_{mean}, pe_{std}, pe_{skew}, pe_{kurtosis}$	Charge distribution
$x_{cht}, y_{cht}, z_{cht}, R_{cht}, \theta_{cht}, \phi_{cht}, J_{cht}, \rho_{cht}, \gamma_z^{cht}, \gamma_y^{cht}, \gamma_x^{cht}$	Center of FHT
$ht_{n\%}, ht_{n_{i+1}\% - n_i\%}, ht_{mean}, ht_{std}, ht_{skew}, ht_{kurtosis}$	FHT distribution

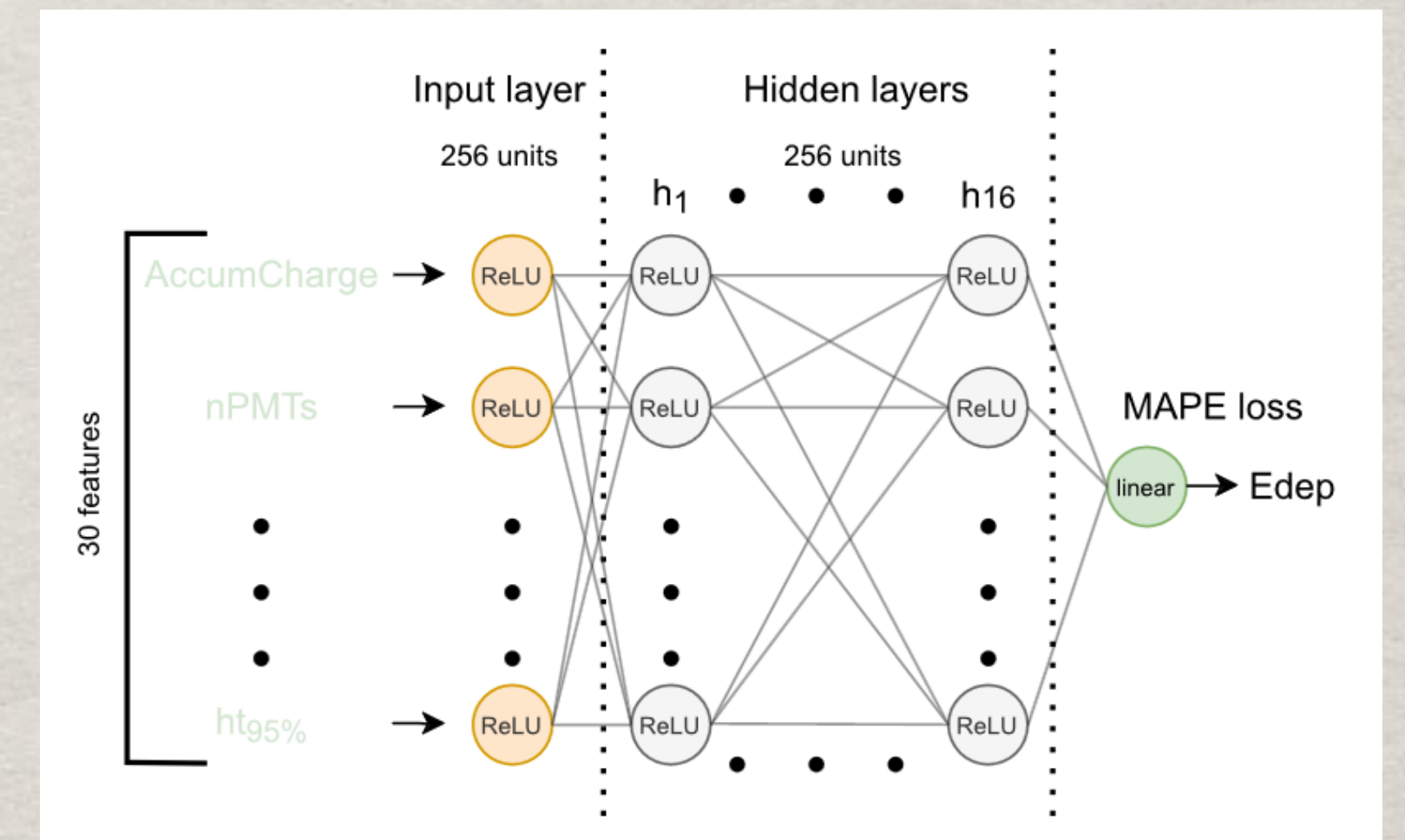
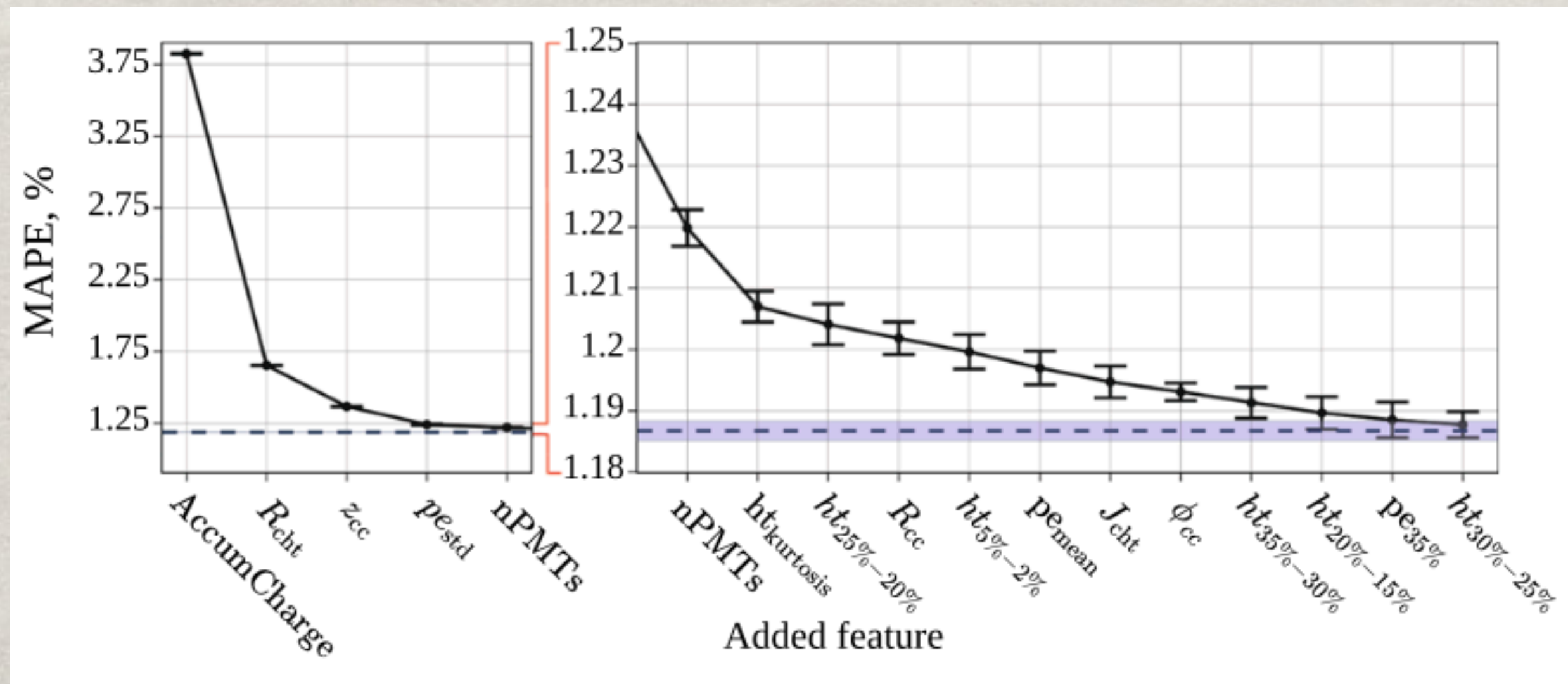


Fig. 12 Selected FCDNN architecture and its main selected hyperparameters



PMT WAVEFORM PHOTON COUNTING

W. Luo@Neutrino2024

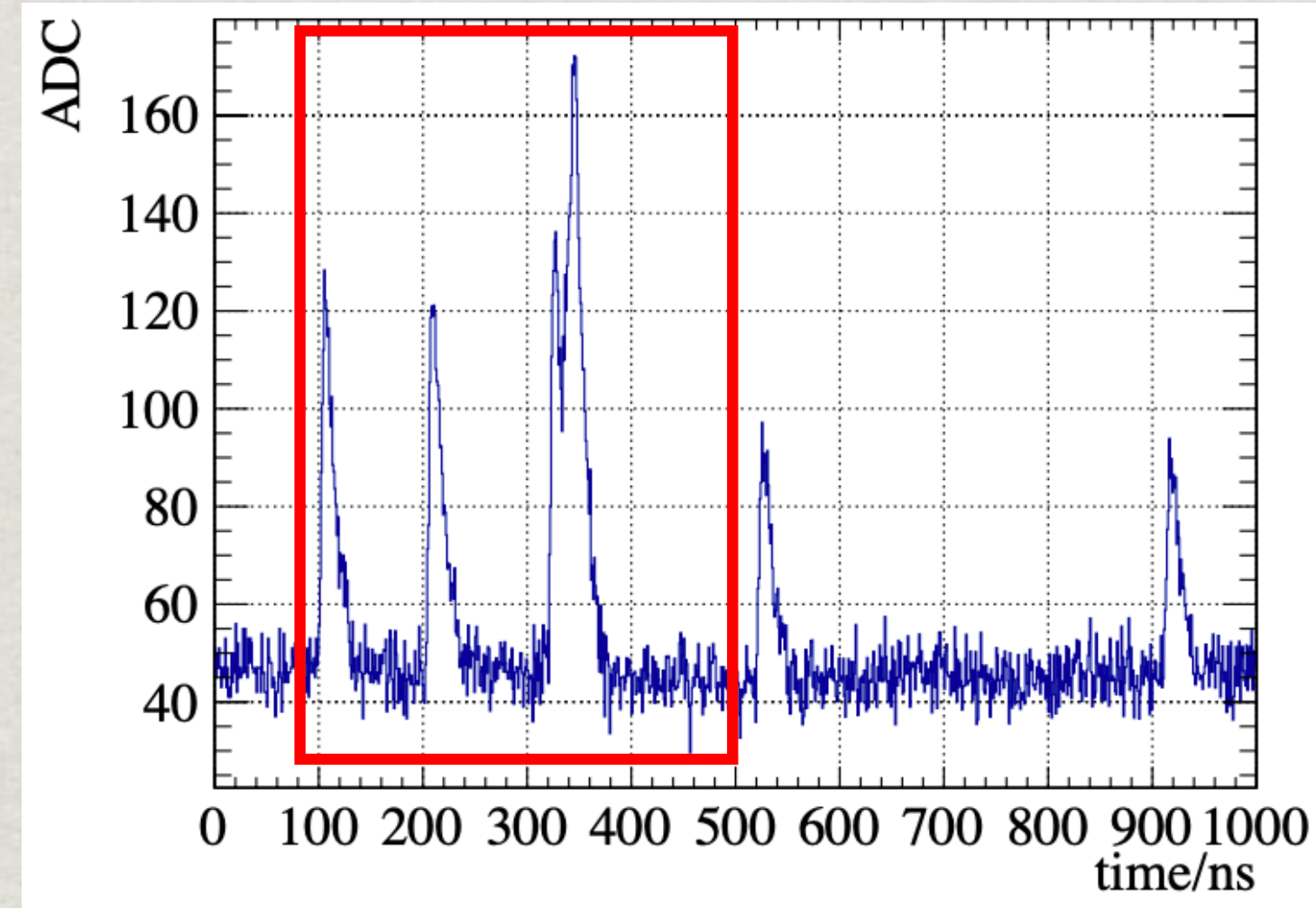
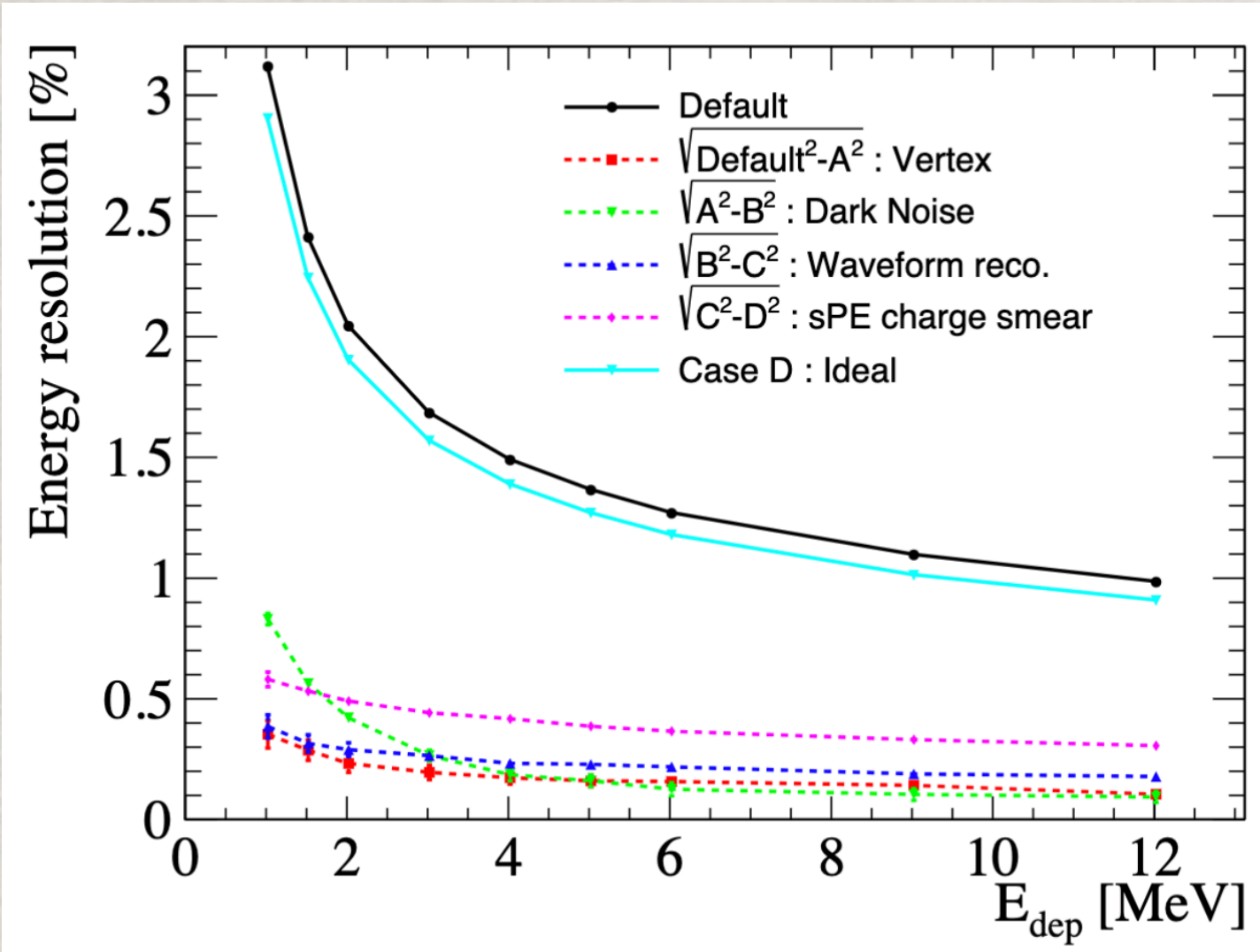
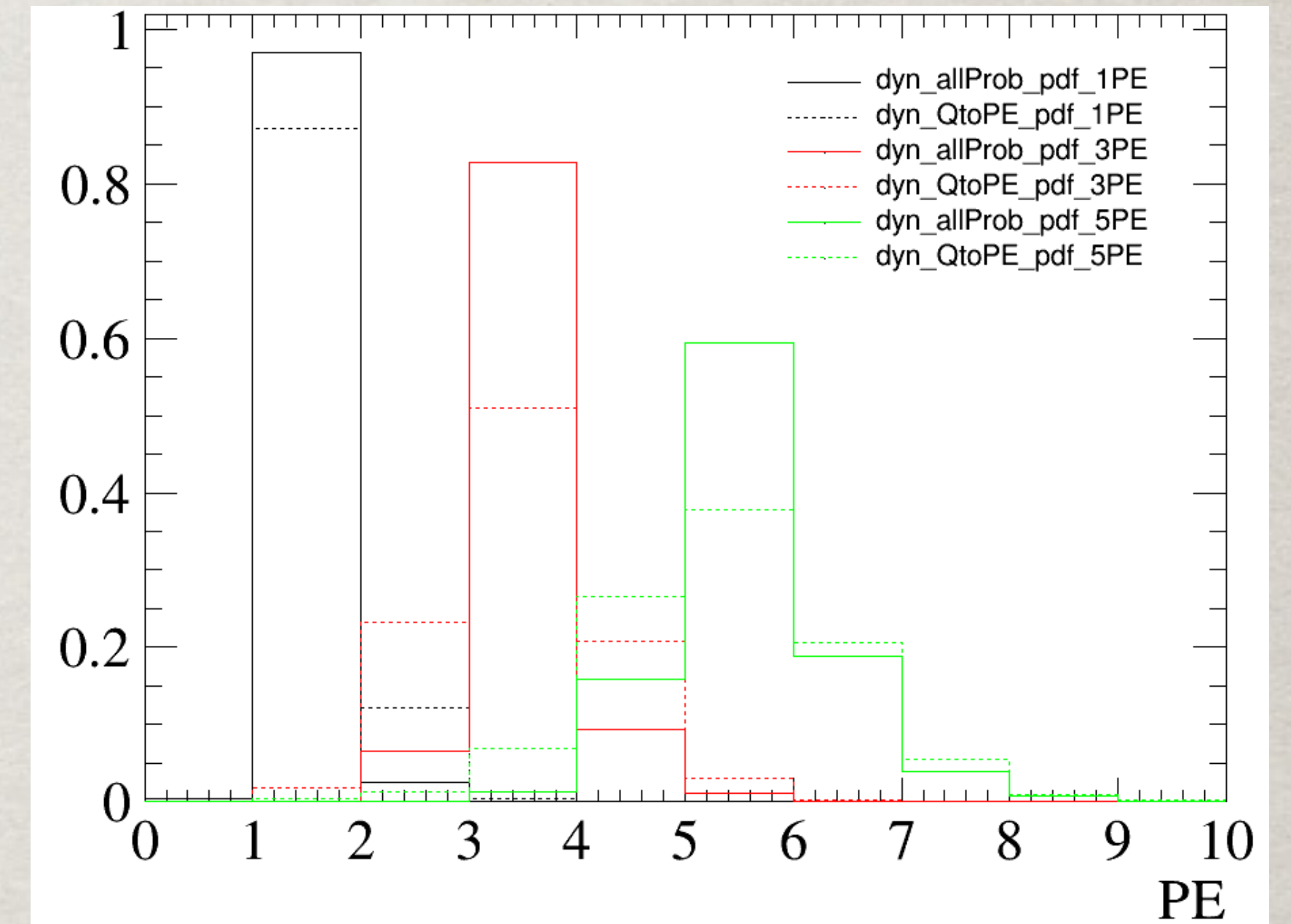
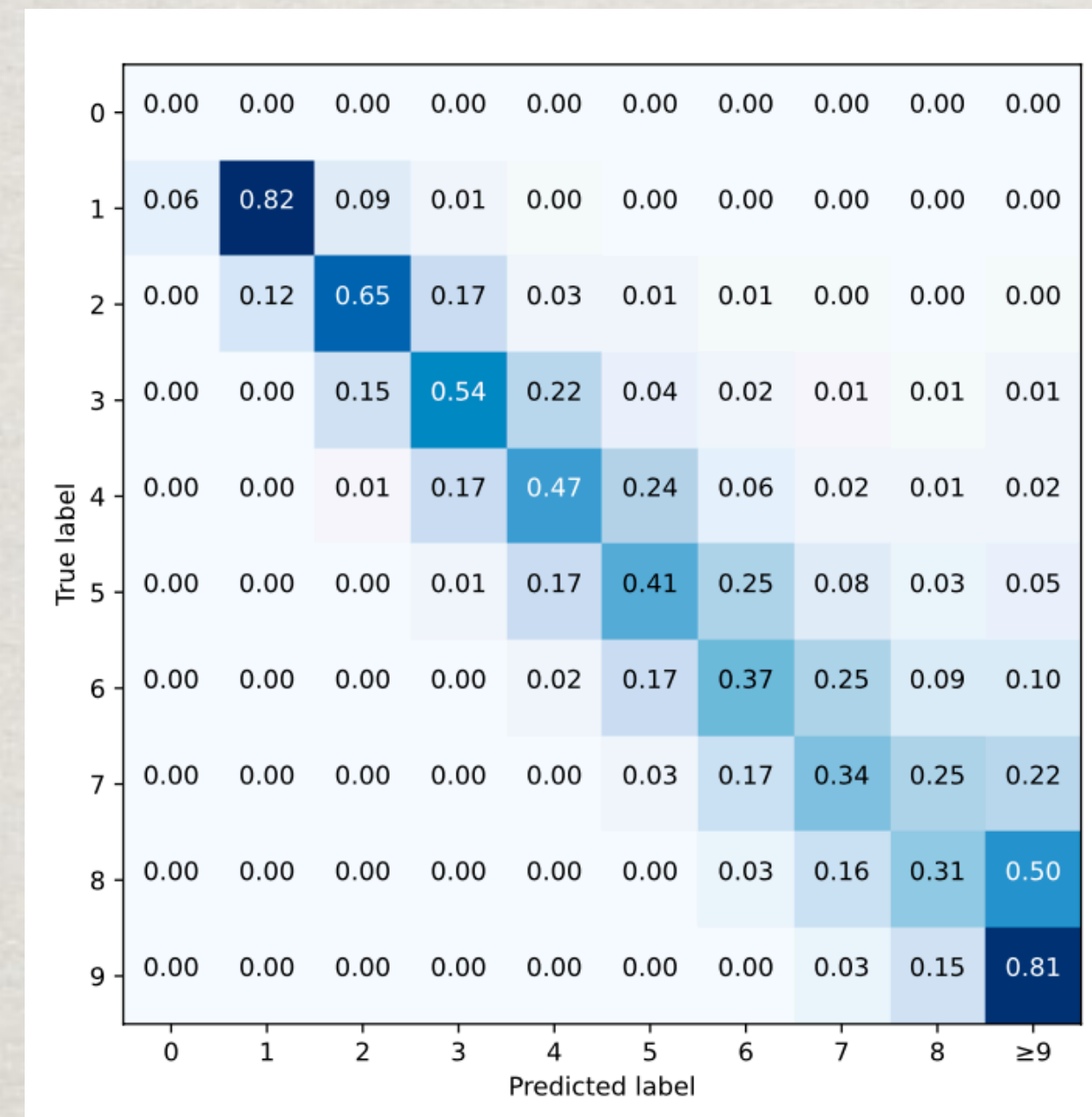
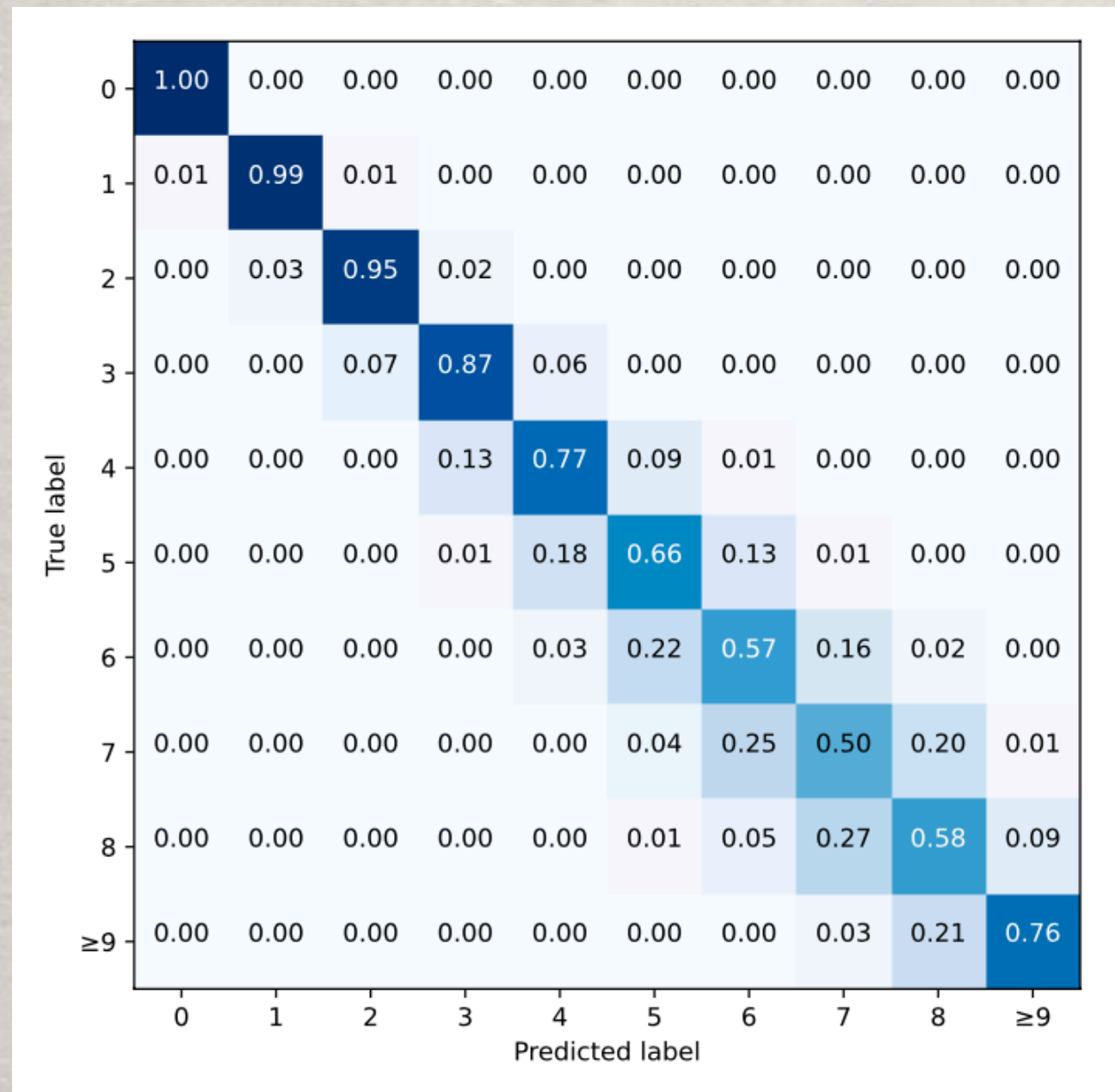


Table 2: Modified RawNet architecture. For convolutional layers, numbers inside parentheses refer to filter length, stride size, and number of filters. For gated recurrent unit (GRU) and fully-connected layers, numbers inside the parentheses indicate the number of nodes.

Layer	Input	Output shape
Strided -conv	Conv(3,3,128) BN LeakyReLU	(128, 140)
Res block	$\left\{ \begin{array}{l} \text{Conv}(3,1,128) \\ \text{BN} \\ \text{LeakyReLU} \\ \text{Conv}(3,1,128) \\ \text{BN} \\ \text{LeakyReLU} \\ \text{MaxPool}(3) \end{array} \right\} \times 2$	(128, 46)
Res block	$\left\{ \begin{array}{l} \text{Conv}(3,1,256) \\ \text{BN} \\ \text{LeakyReLU} \\ \text{Conv}(3,1,256) \\ \text{BN} \\ \text{LeakyReLU} \\ \text{MaxPool}(3) \end{array} \right\} \times 2$	(256, 1)
GRU	GRU(1024)	(1024,)
Speaker embedding	FC(128)	(128,)
Output	FC(10)	(10,)

- ❖ Input: pre-processed PMT waveform within 420ns signal window
- ❖ Model: Customized RawNet
- ❖ Output: $\{p_k\}$ the probability for predicting ($k=0, 1, \dots, \geq 9$) PEs

PHOTON COUNTING PERFORMANCE



- ❖ Left: Confusion matrix of RawNet
 - ❖ 99% (95%, 87%) accuracy for 1PE (2PEs, 3PEs)
 - ❖ Accuracy decreases rapidly as nPEs increases
- ❖ Right: Confusion matrix based on charge classification
 - ❖ The accuracy is markedly inferior to that of RawNet

[W. Luo@Neutrino2024](mailto:W.Luo@Neutrino2024)



ENERGY RECONSTRUCTION

Note	Algo. Name	Observable	Likelihood: $\kappa \leq K_T$	Likelihood: $\kappa > K_T$
Charge, Time, Maximum Likelihood Estimation	QTMLE (reference)	q (charge)	$\mathcal{L}(q_i \mu_i) = \sum_{k=1}^{+\infty} P_Q(q_i k)P(k, \mu_i)$	
PE, Time, Maximum Likelihood Estimation	PETMLE (ideal)	k (true PEs)	$\mathcal{L}(k_i \mu_i) = P(k_i, \mu_i)$	
Charge, Probability, Time, Maximum Likelihood Estimation	QPTMLE (realistic)	$\{p_k\}$, q	$\mathcal{L}(\{p_k^i\} \mu_i) = \sum_{k=0}^9 R_{K_T k} p_k^i P(k, \mu_i)$	
Charge, PE, Time, Maximum Likelihood Estimation	QPETMLE (100% accuracy)	$k(p_k=1)$, q	$\mathcal{L}(k_i \mu_i) = P(k_i, \mu_i)$	$\mathcal{L}(q_i \mu_i) = \sum_{k=1}^{+\infty} P_Q(q_i k)P(k, \mu_i)$
Charge, Confusion Matrix, Time, Maximum Likelihood Estimation	QCTMLE	κ ($p_{\kappa}:\text{max}$), q	$) = \sum_{k=0}^9 C_{k\kappa_i} \times P(k, \mu_i)$	

where μ_i is the expected nPEs for the i-th PMT, $P(k, \mu_i)$ is just the Poisson probability of observing k p.e. given μ_i and $P_Q(q_i|k)$ is the charge pdf for k p.e..

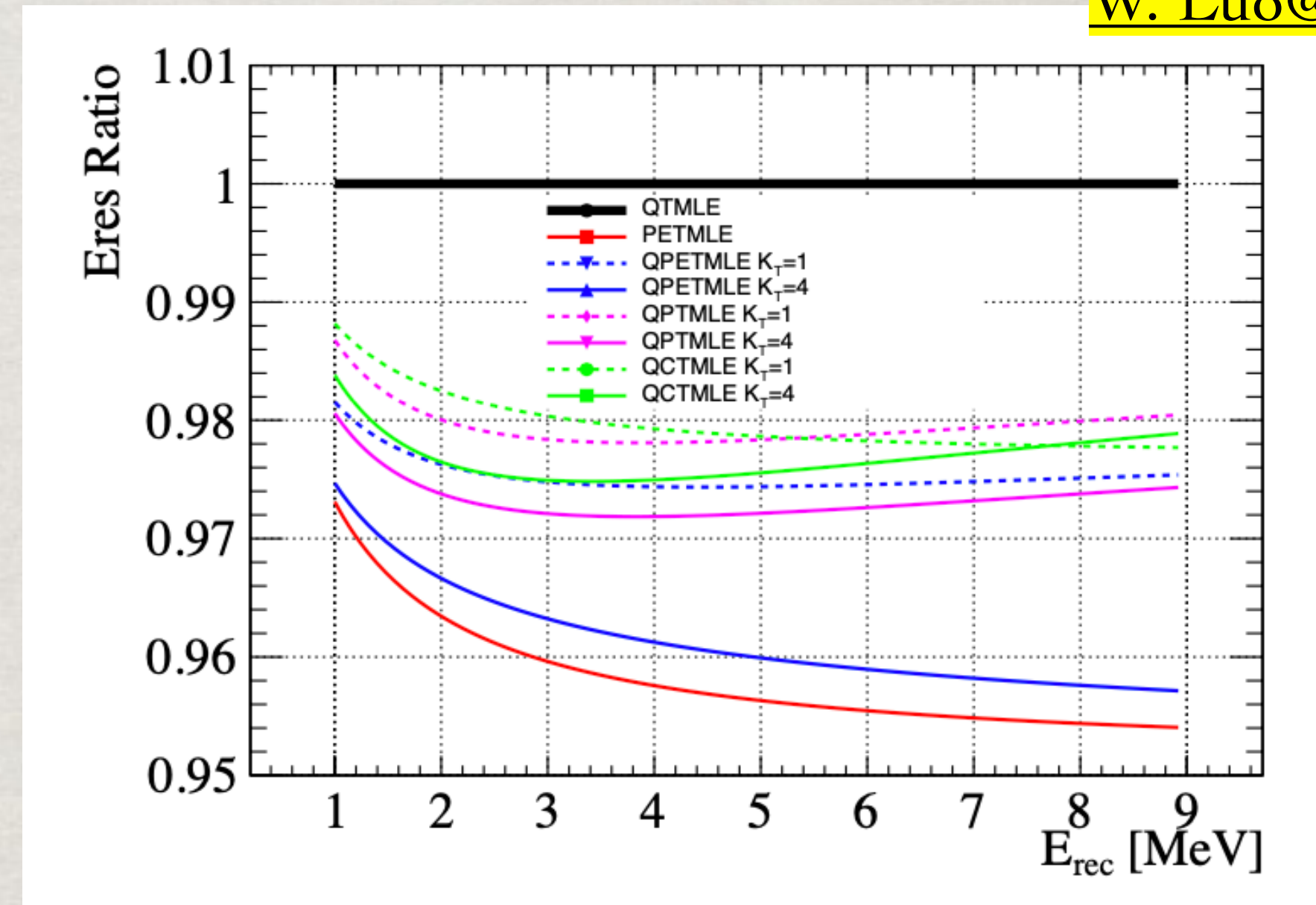
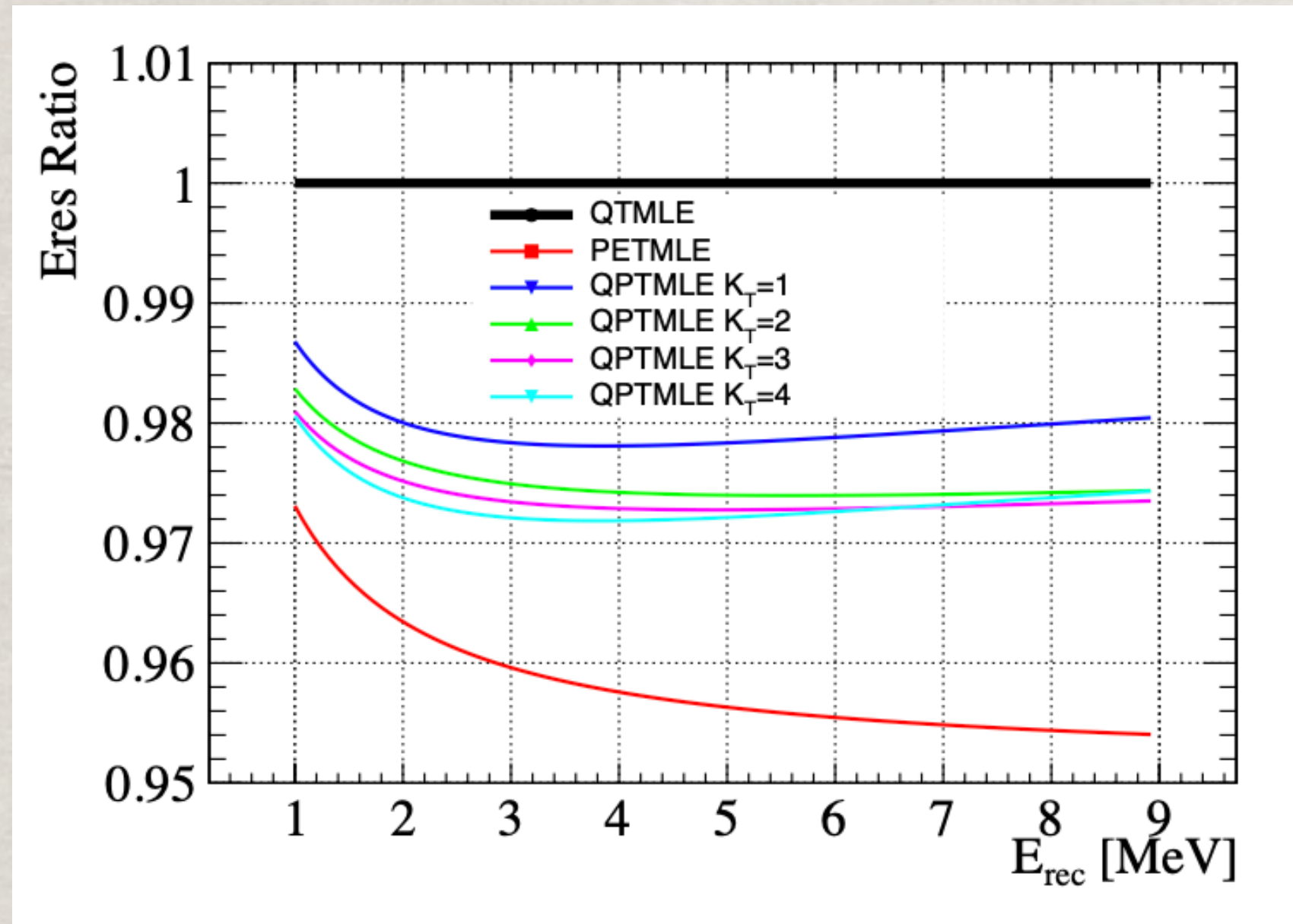
$$R_{K_T k} = \sum_{\kappa=0}^{K_T} C_{k\kappa}$$

confusion matrix $C_{kk'}$



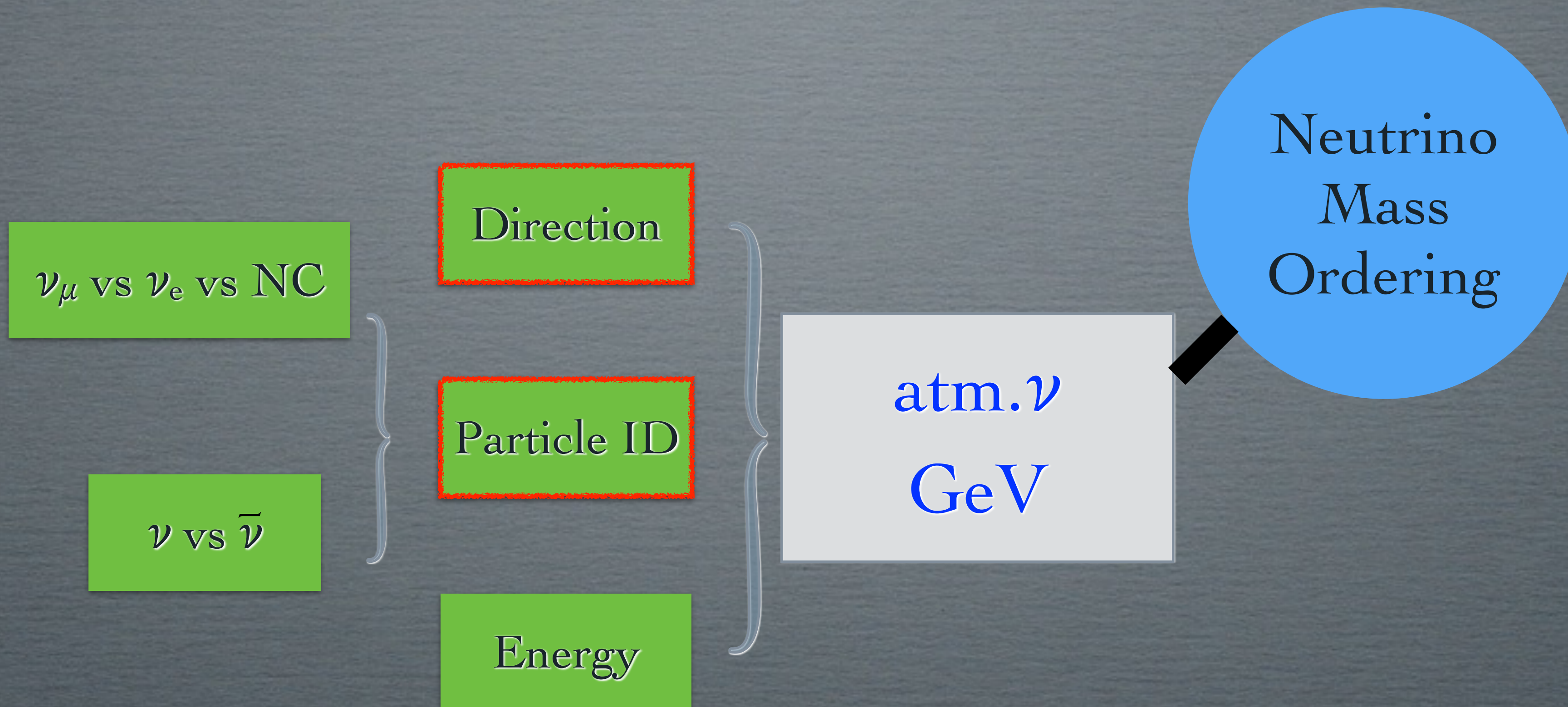
RELATIVE IMPROVEMENT

W. Luo@Neutrino2024



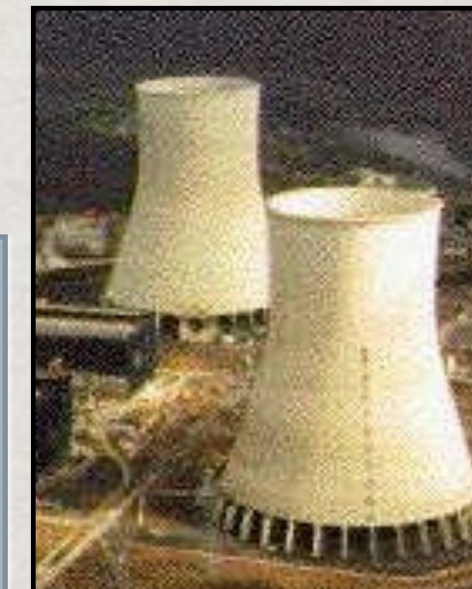
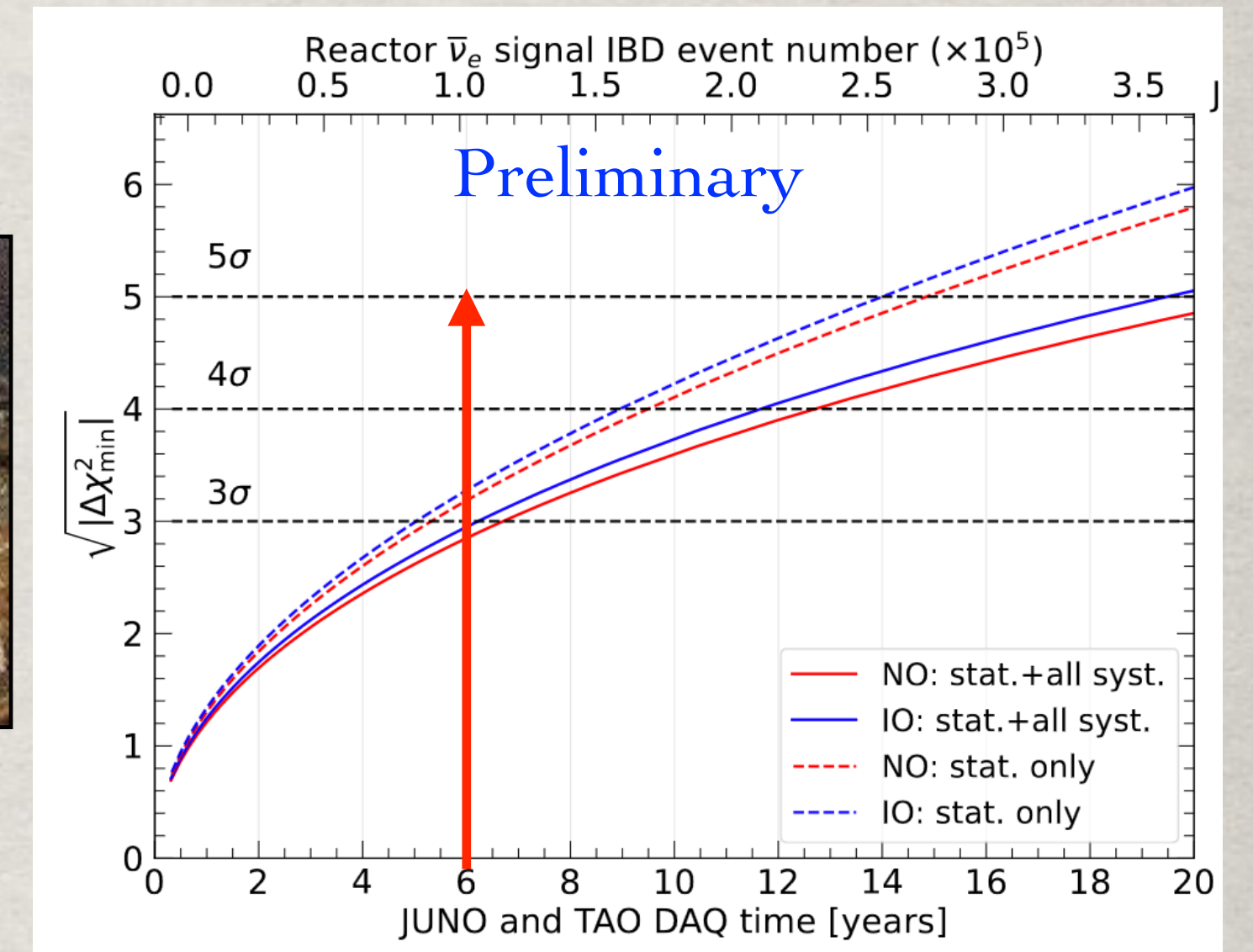
- ❖ Using the photon counting information for PMTs with ($\kappa \leq K_T$) PEs can improve the energy resolution
- ❖ The improvement becomes smaller as K_T increases due to the dropping accuracy for high PEs
- ❖ Additional checks were done to validate the results





JUNO NMO SYNERGY

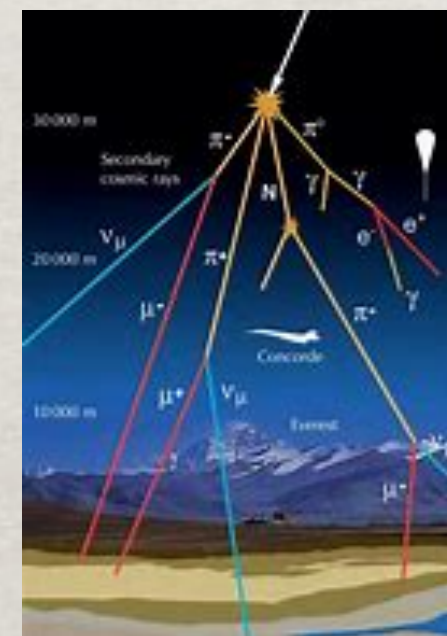
- ❖ NMO @ 6years $\Delta\chi^2$: Reactor(~ 9), atm.(~ 1.96), $|\Delta m^2_{ee}|$ (4|1.5% or 9|1%)
- ❖ 1.96 of atm. was estimated with assumptions
- ❖ Can we do better than Yellow Book?



Yellow Book assumptions	
Event Selection $\nu_e/\bar{\nu}_e$	$E_{vis} > 1\text{GeV}$ $Y_{vis} = E_h/E_{vis} < 0.5$
Directionality	$\sigma_{\theta\mu} = 1^\circ$ $\sigma_{\theta\nu} = 10^\circ$ CC-e vs CC- μ vs NC: 100% eff.
Classification	ν vs $\bar{\nu}$: Ne, Y_{vis} ,
Energy	$\sigma_{E_{vis}} = 1\%/ \sqrt{E}$

reactor ν
MeV

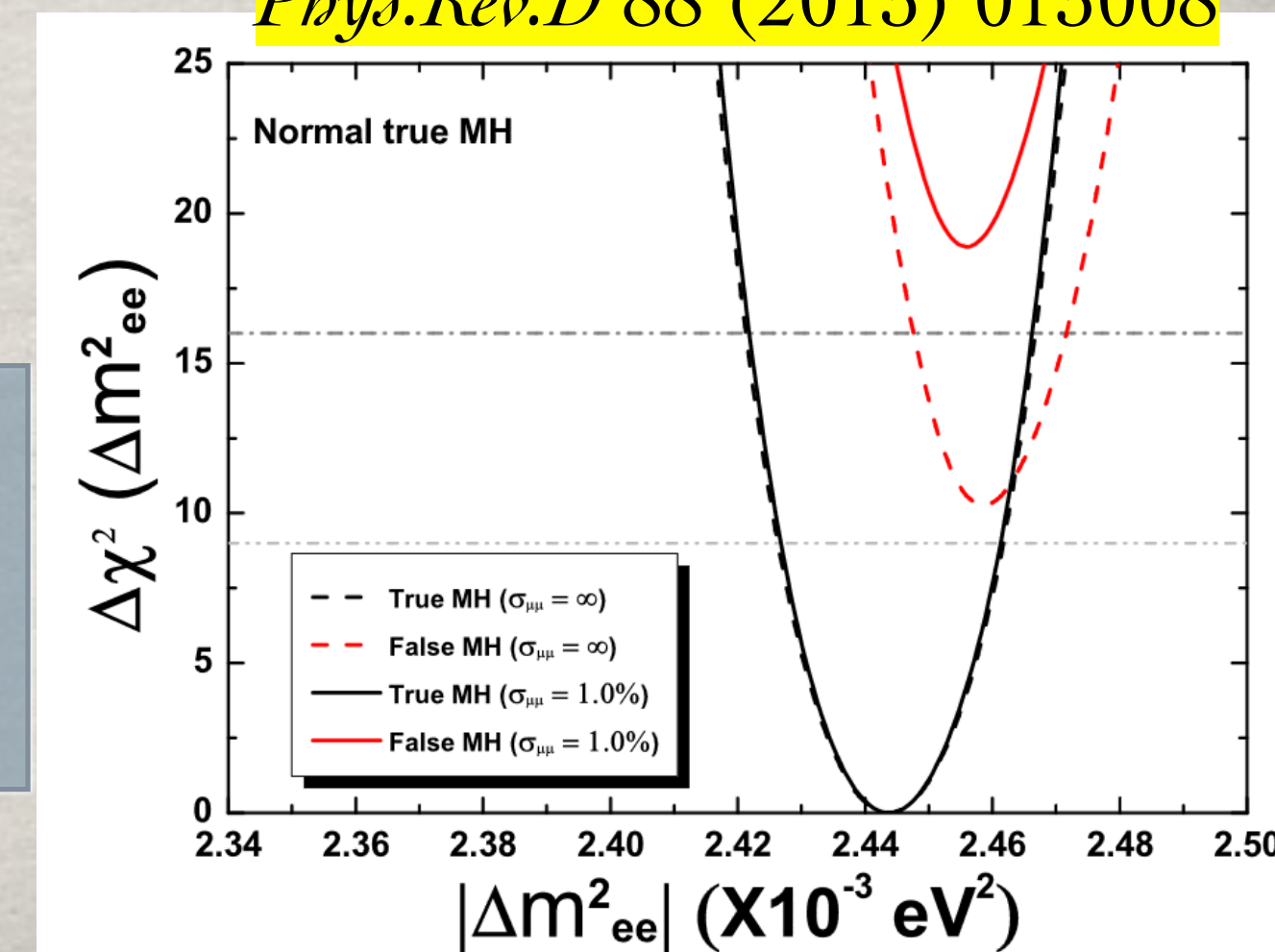
Neutrino
Mass
Ordering



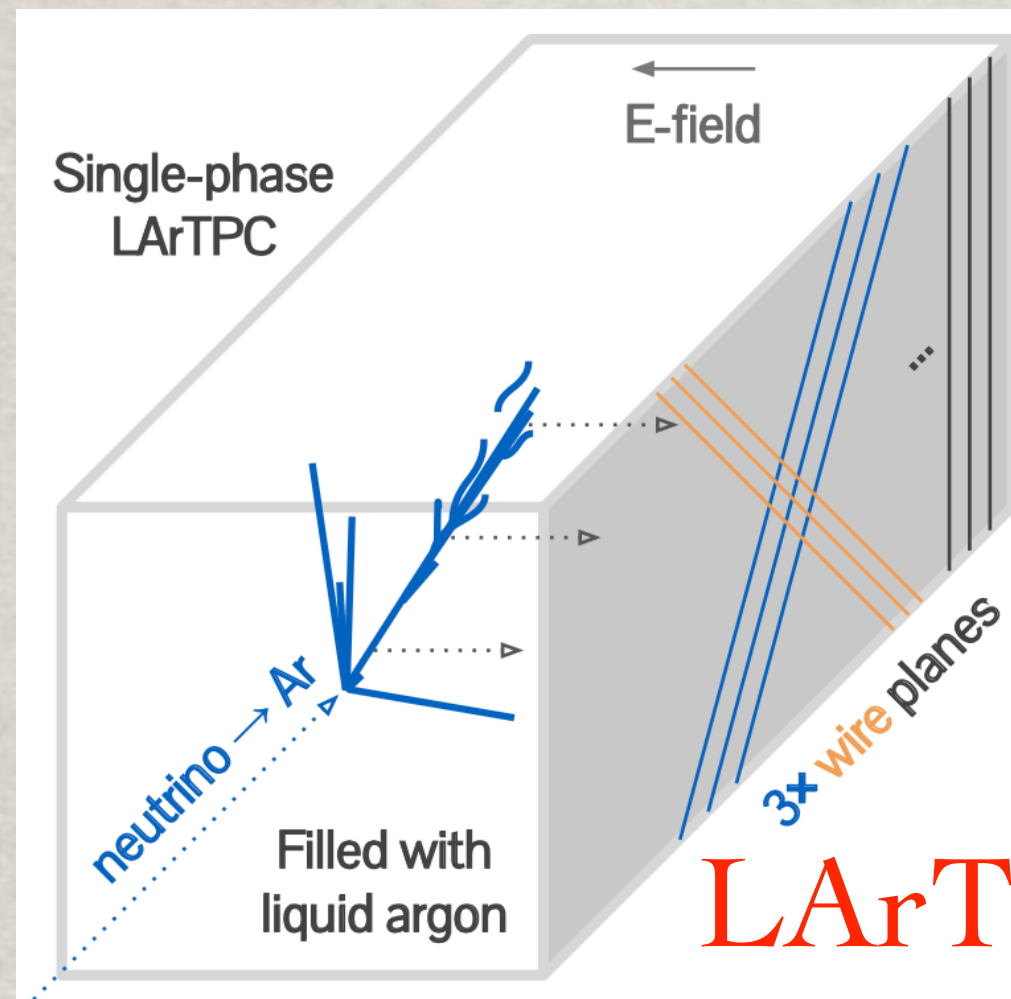
atm. ν
GeV

$|\Delta m^2_{\alpha\alpha}|$

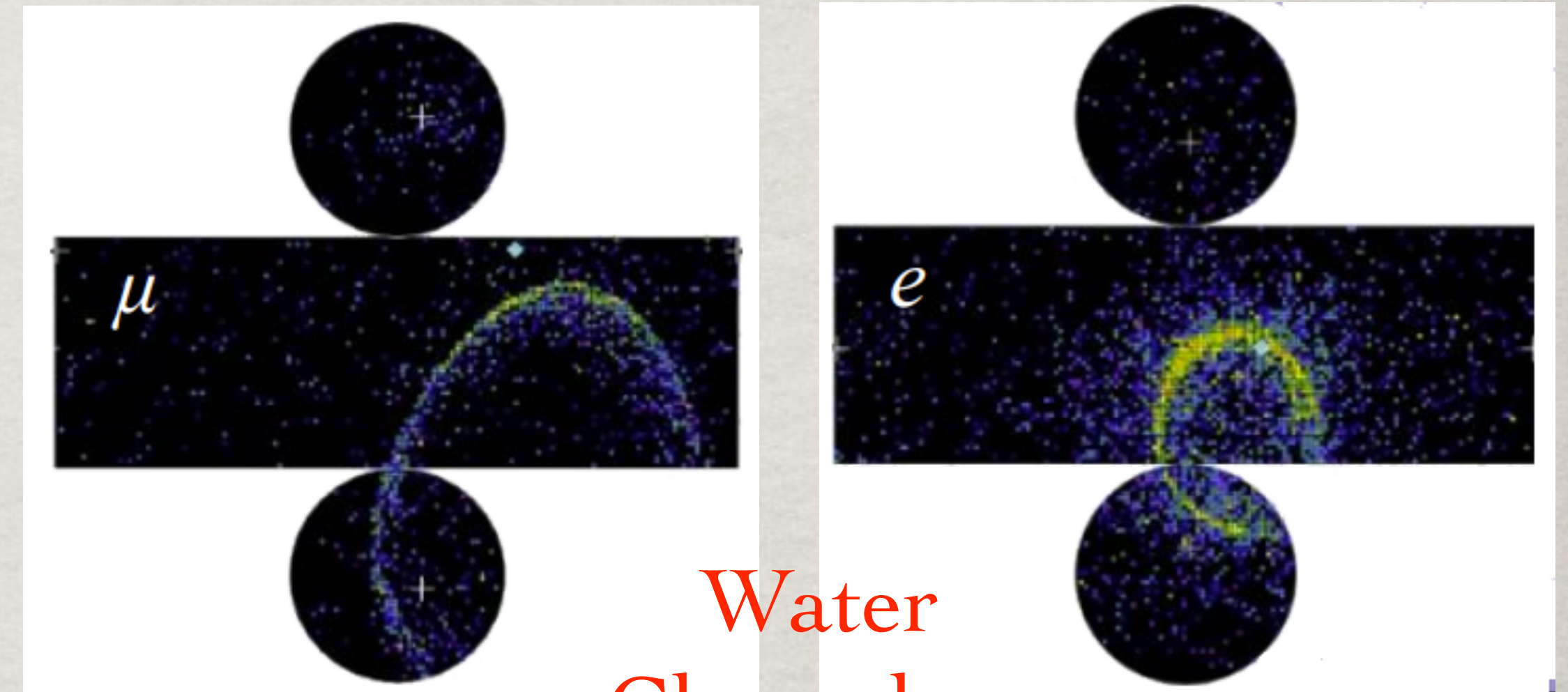
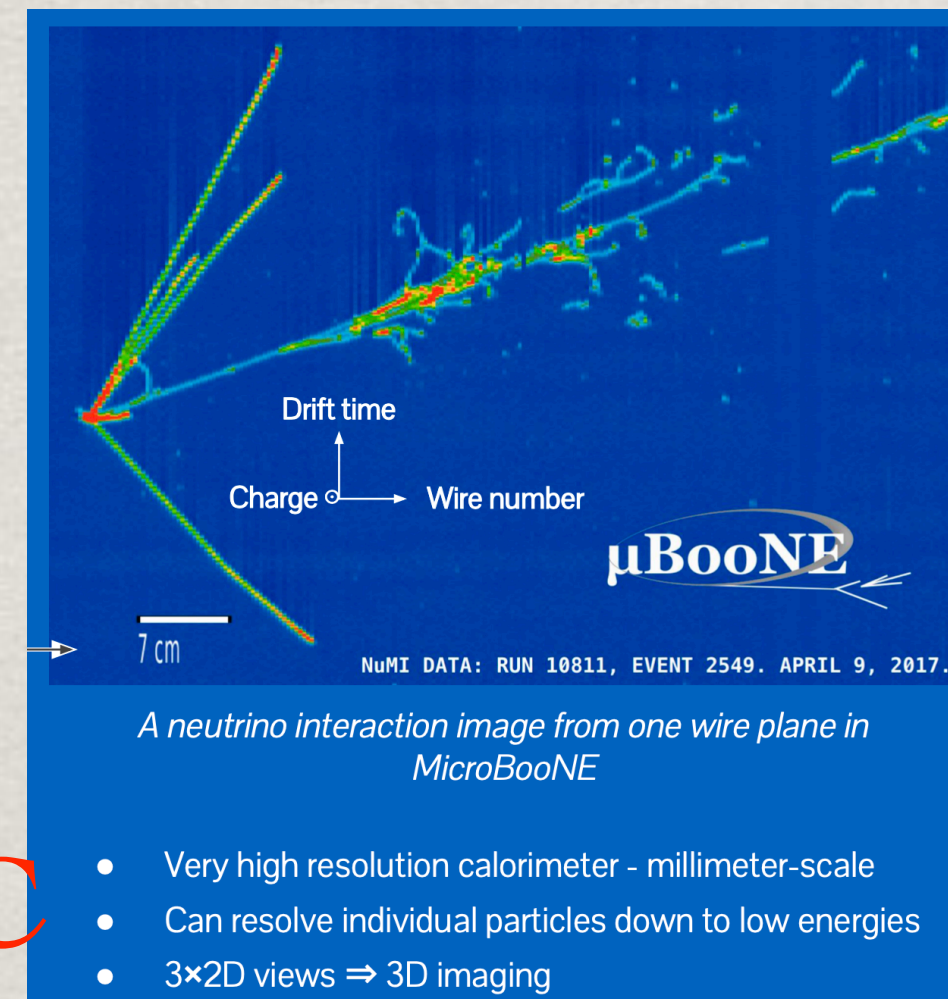
Yufeng Li et al.
Phys.Rev.D 88 (2013) 013008



CHALLENGES AND OPPORTUNITIES



LArTPC



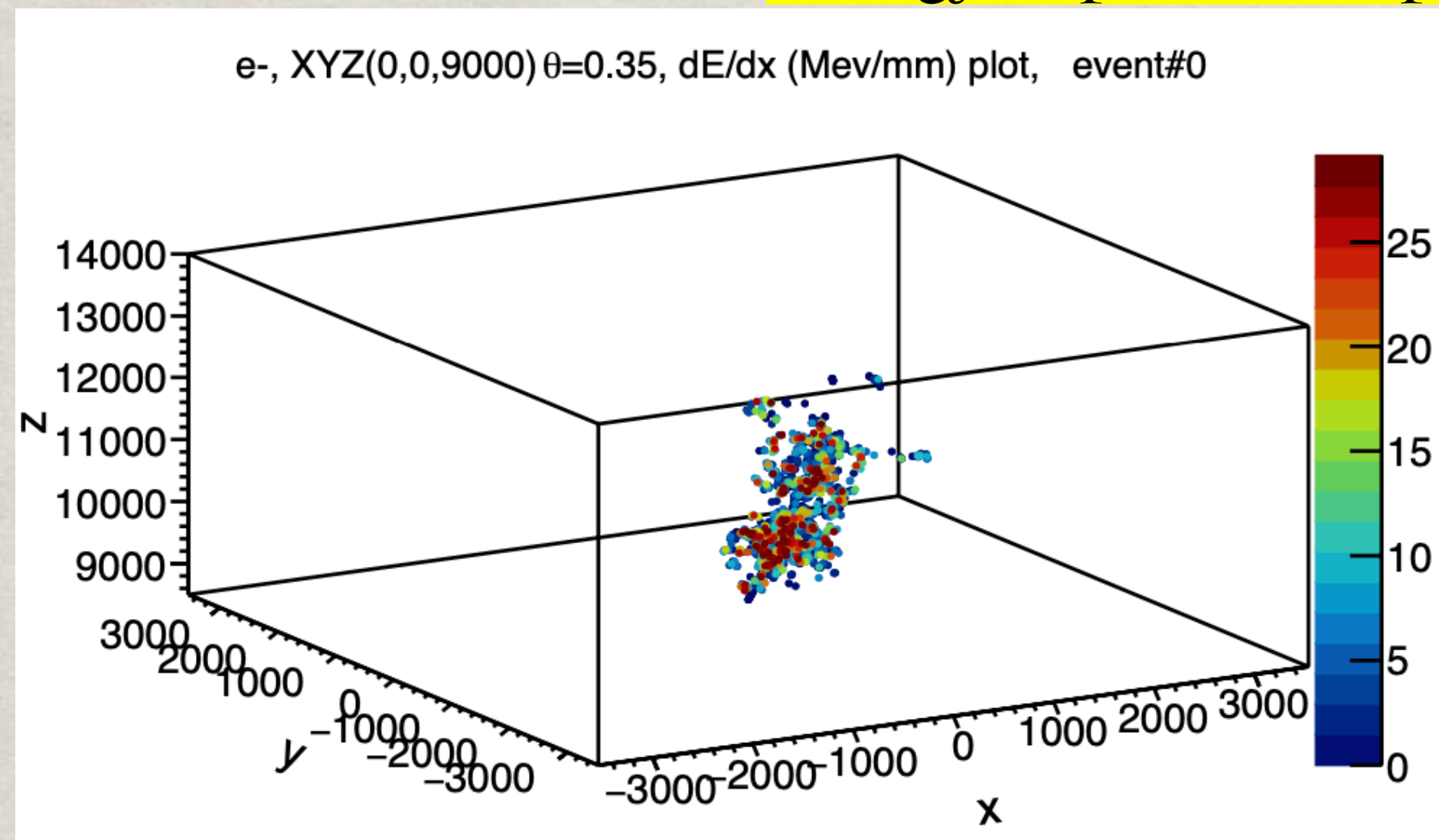
Water
Cherenkov

- ❖ Neither track information, nor Cherenkov rings for JUNO
- ❖ Advantages of JUNO: 1. large PMT coverage(78%), large volume; 2. excellent neutron tagging; 3. hadronic component visible in LS; 4. can measure distinctive isotopes

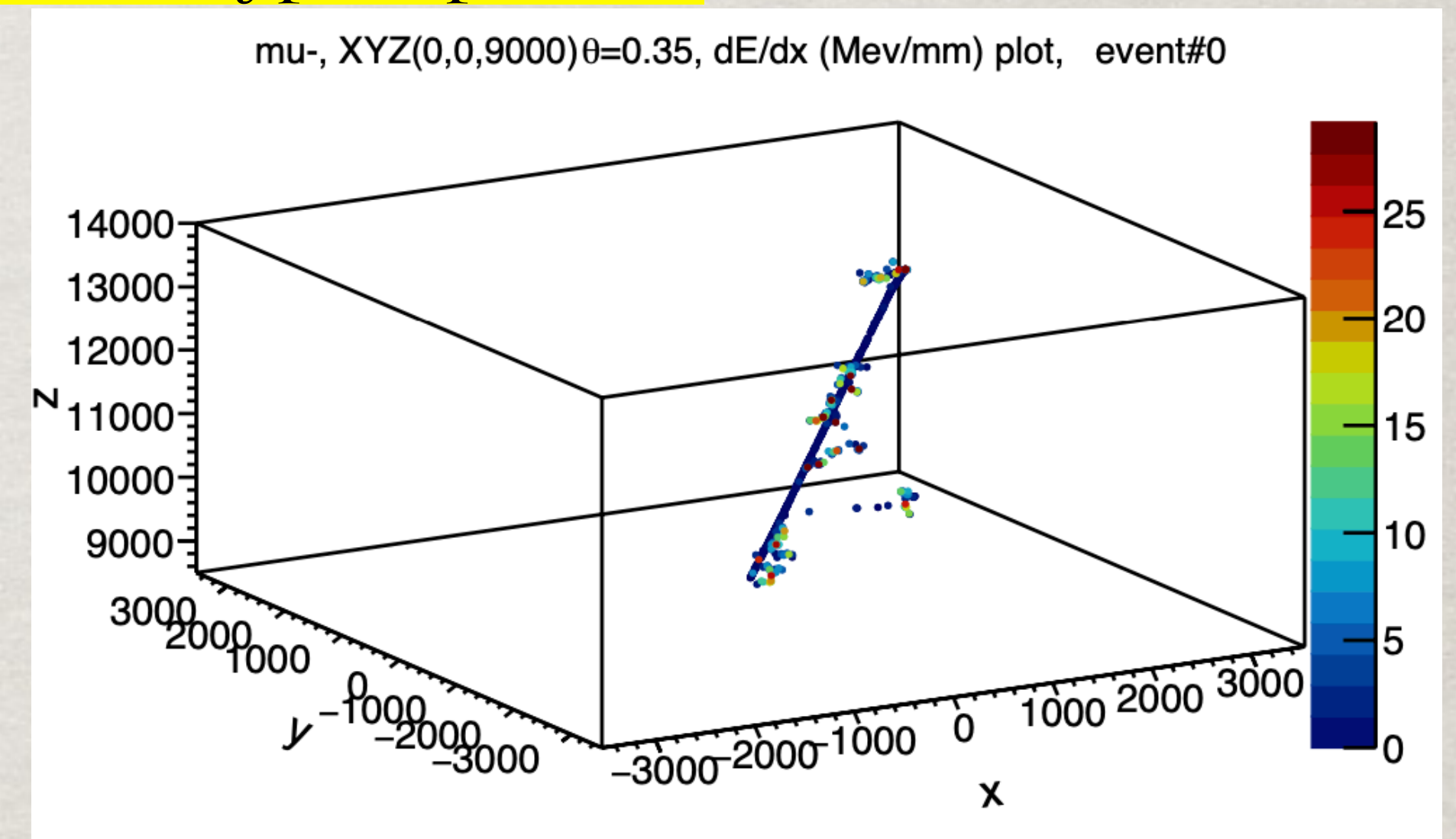
PARTICLE TOPOLOGY

Energy deposition topology in LS for different type of particles

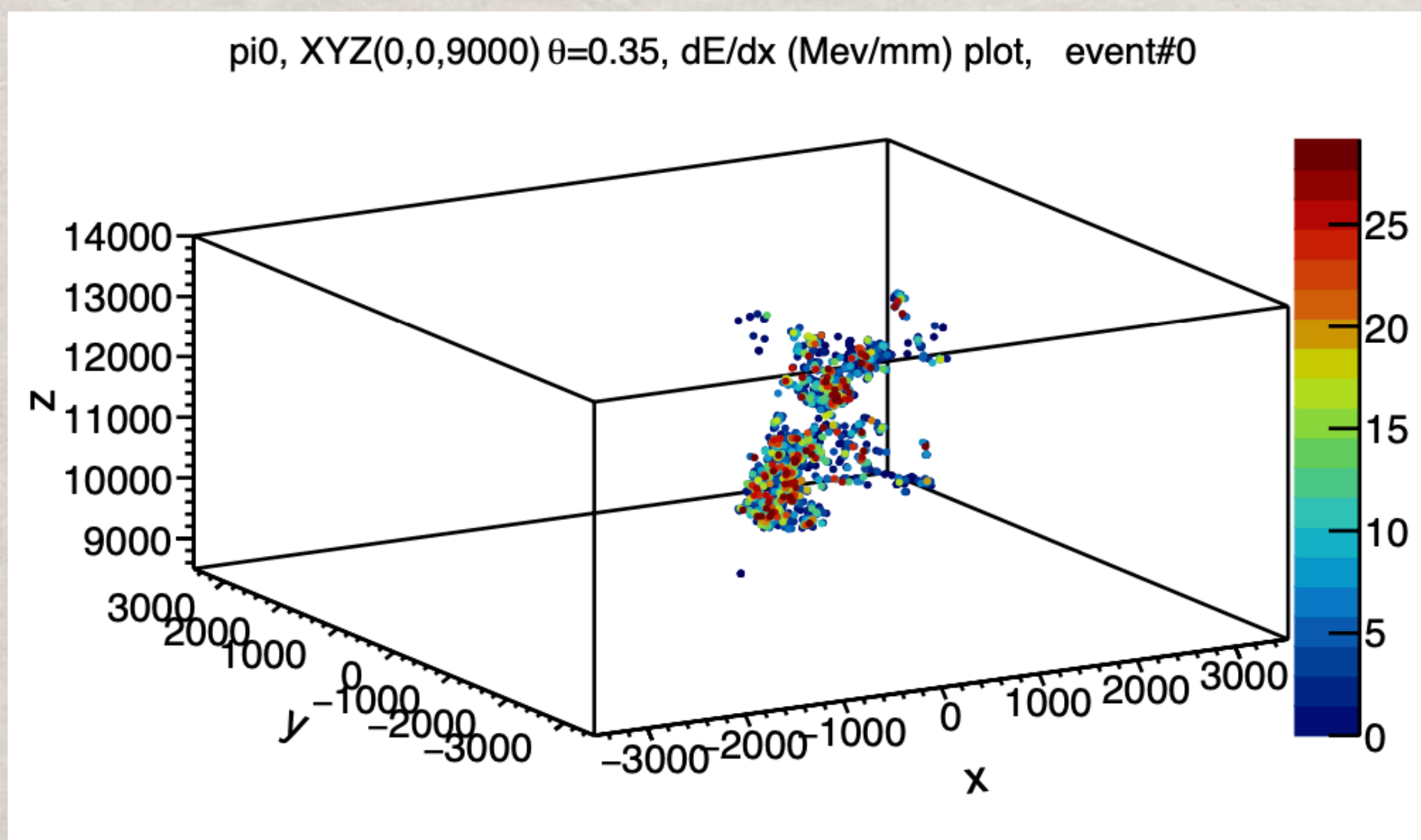
e^-



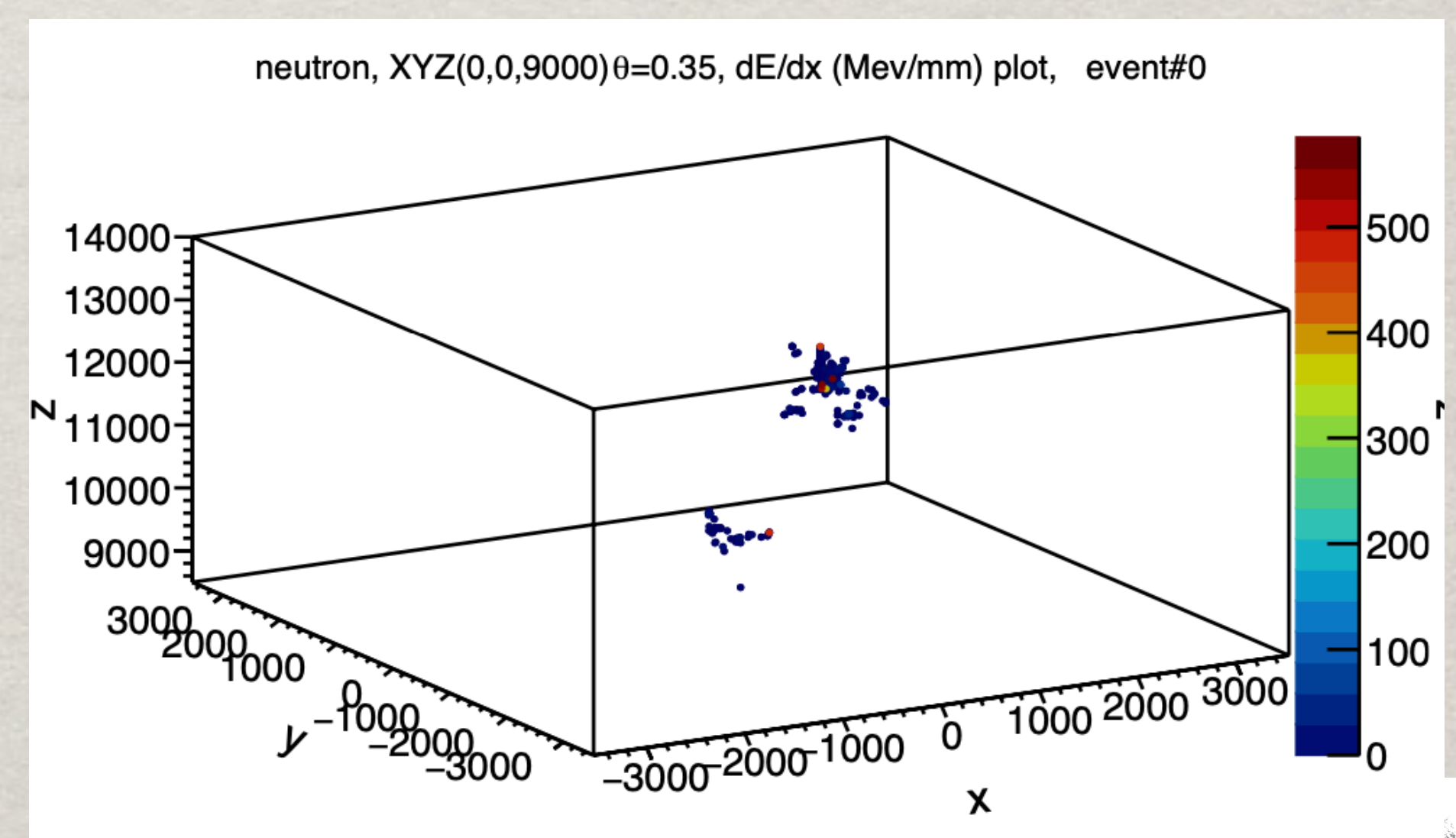
μ



π^0

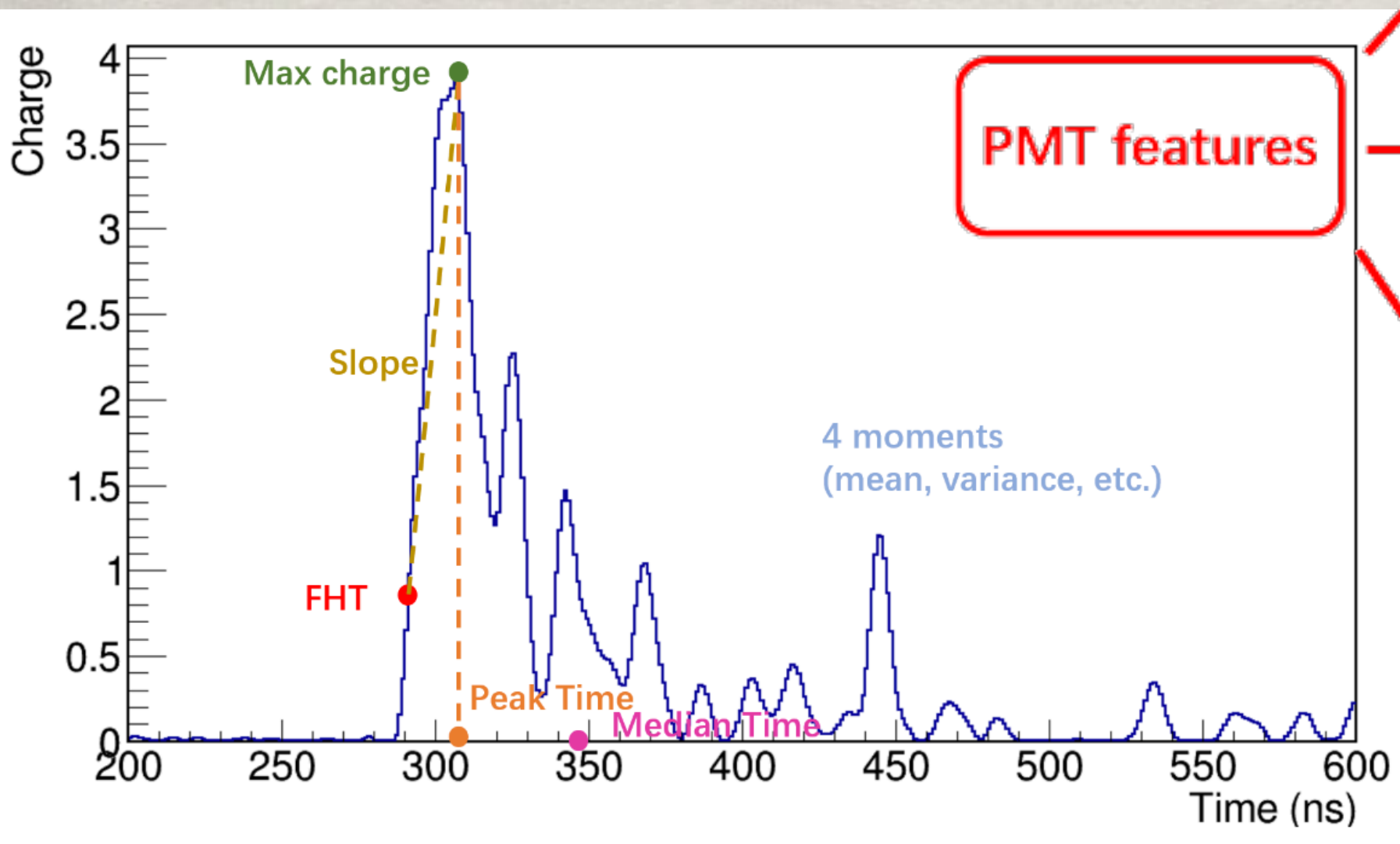
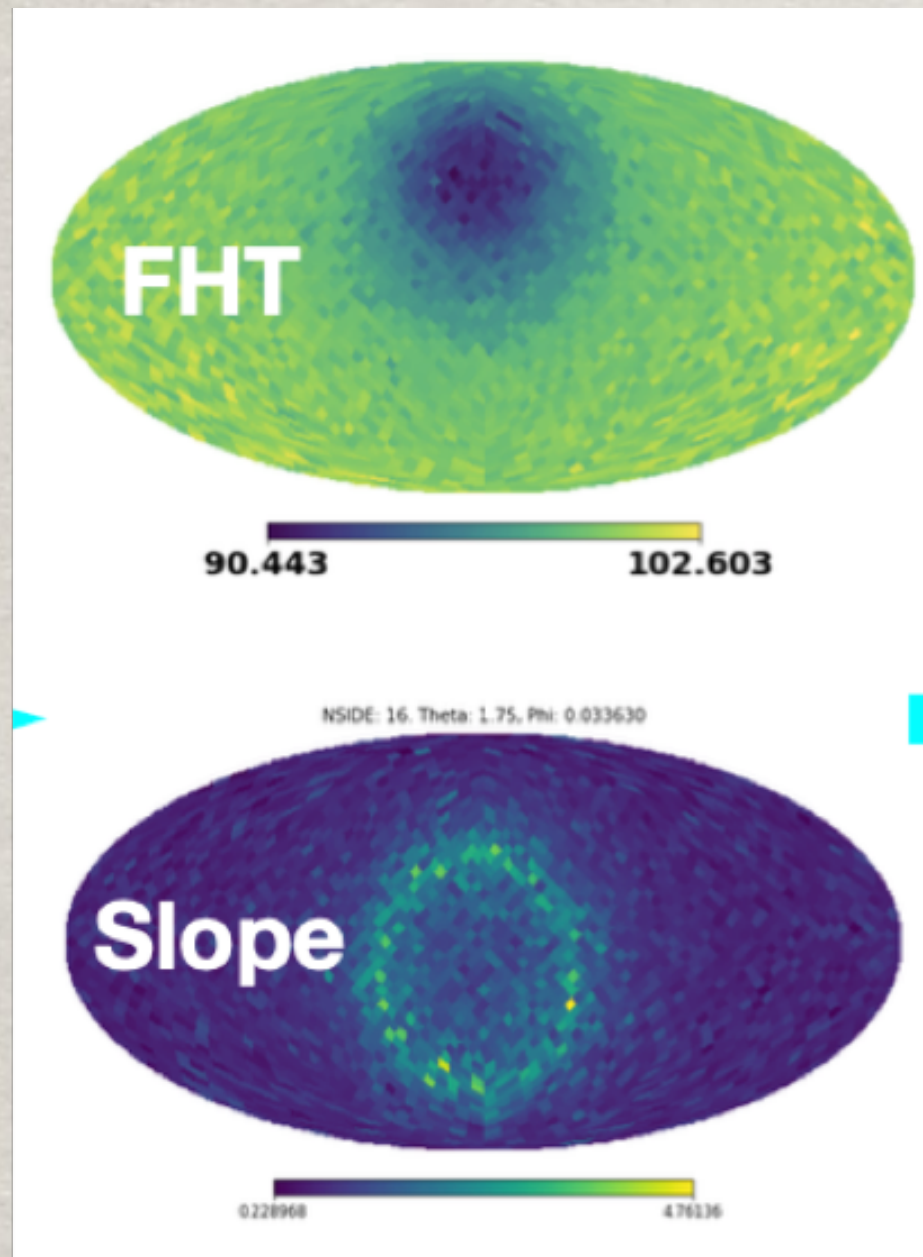


n

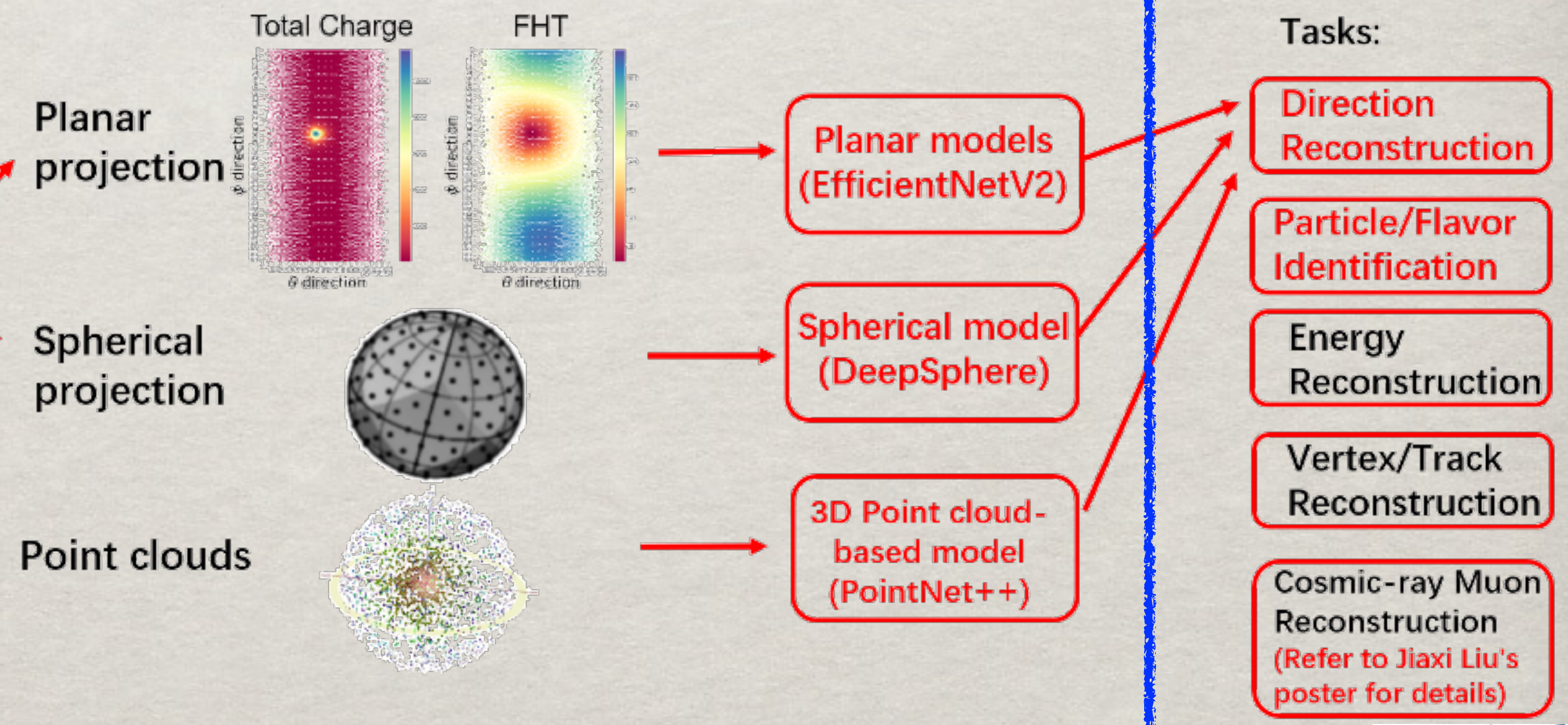


RECO/PID METHODOLOGY

- ❖ Step 1: feature extraction from PMT waveforms
- ❖ Step 2: model building
- ❖ Step 3: optimization and validation



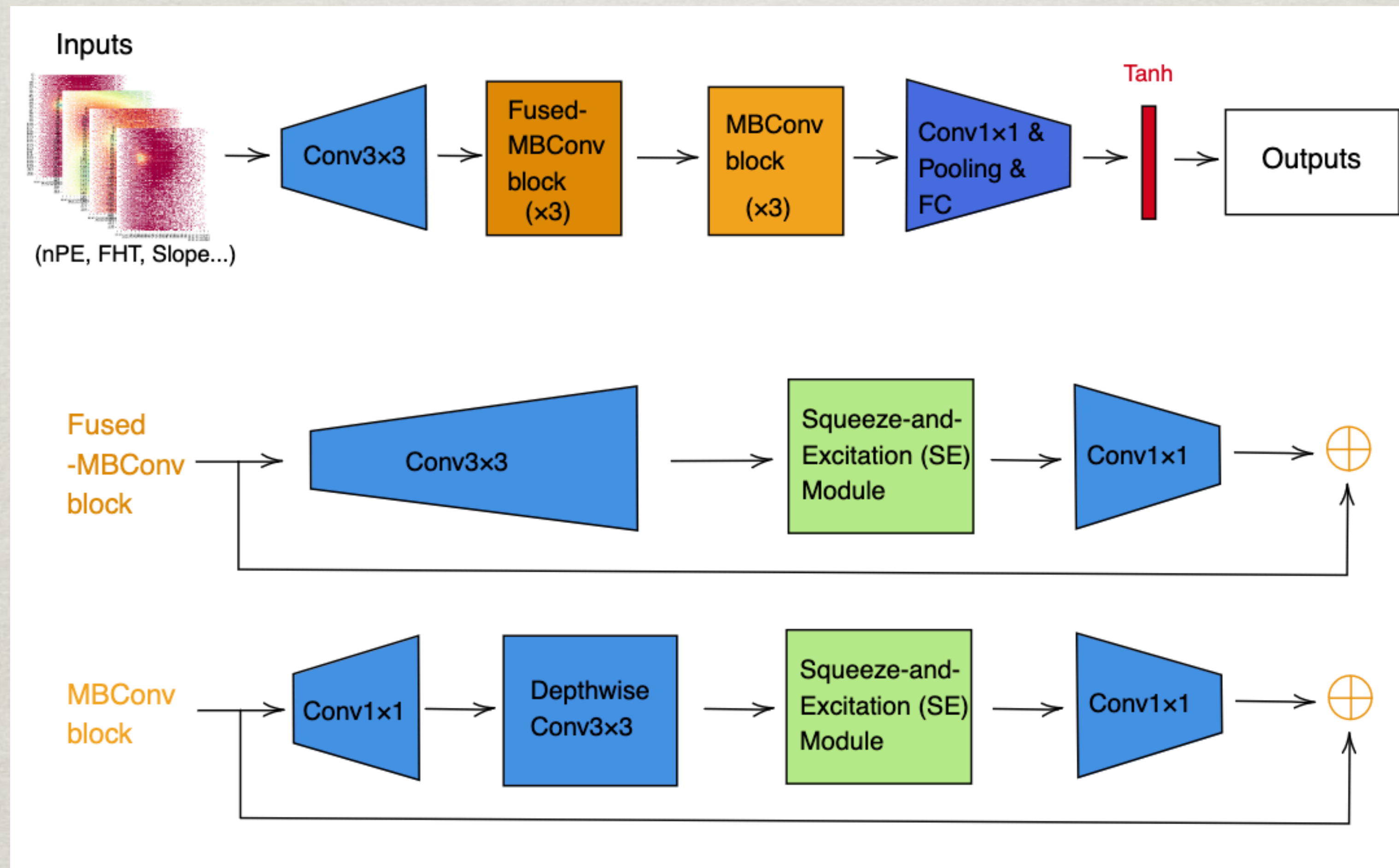
PMT features



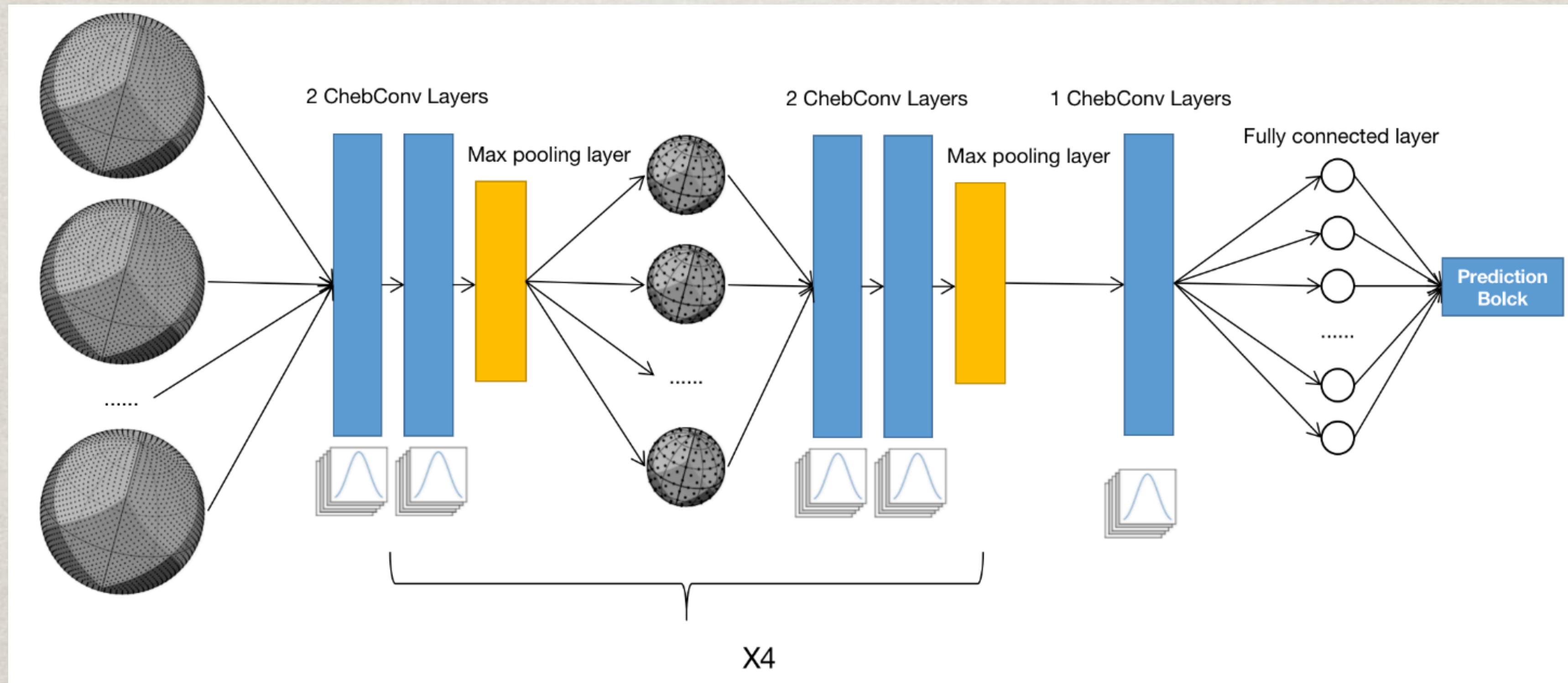
- Tasks:
- Direction Reconstruction
 - Particle/Flavor Identification
 - Energy Reconstruction
 - Vertex/Track Reconstruction
 - Cosmic-ray Muon Reconstruction (Refer to Jiaxi Liu's poster for details)



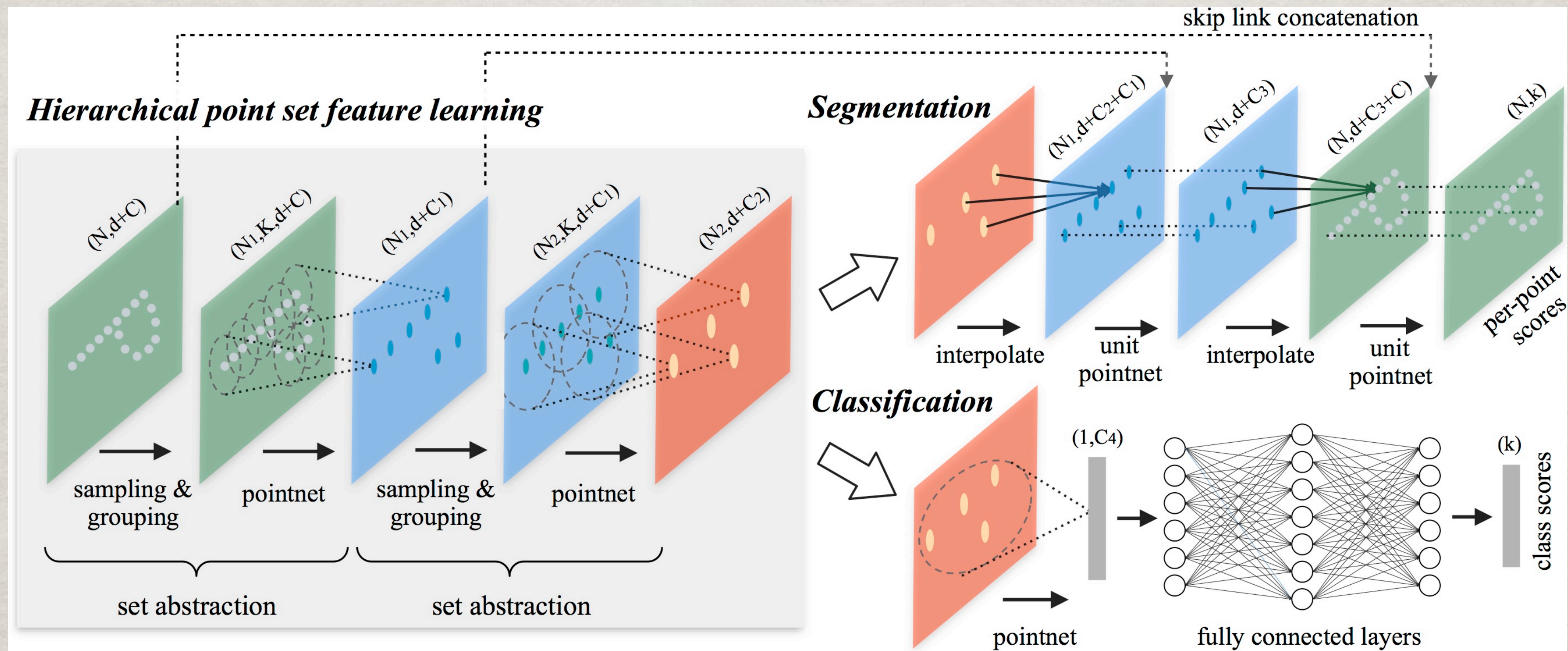
PLANE MODEL: EFFICIENTNETV2-S



SPHERICAL MODEL: DEEPSPHERE

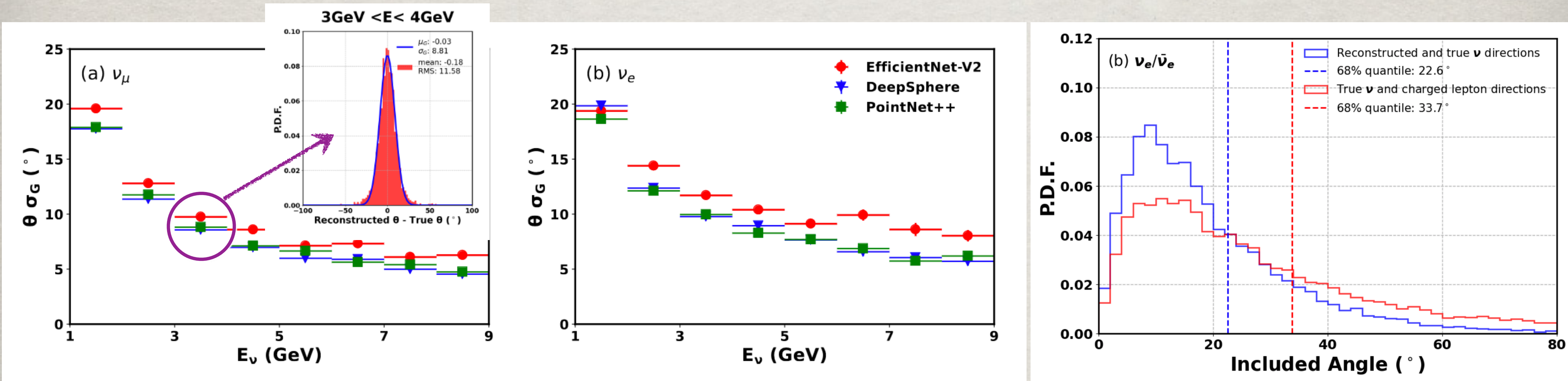


3D MODEL: POINTNET++



DIRECTIONALITY

Phys.Rev.D 109 (2024) 5, 052005



- ❖ Directly reconstruct the direction of ν instead of the charged lepton
- ❖ mitigate the intrinsic large uncertainty between the two
- ❖ hadronic component in LS also helps, advantageous w.r.t. Water Cerenkov
- ❖ Energy dependent Zenith Angle resolution, less than 10° for $E > 3\text{GeV}$

J. Phys. G: 43 (2016) 030401

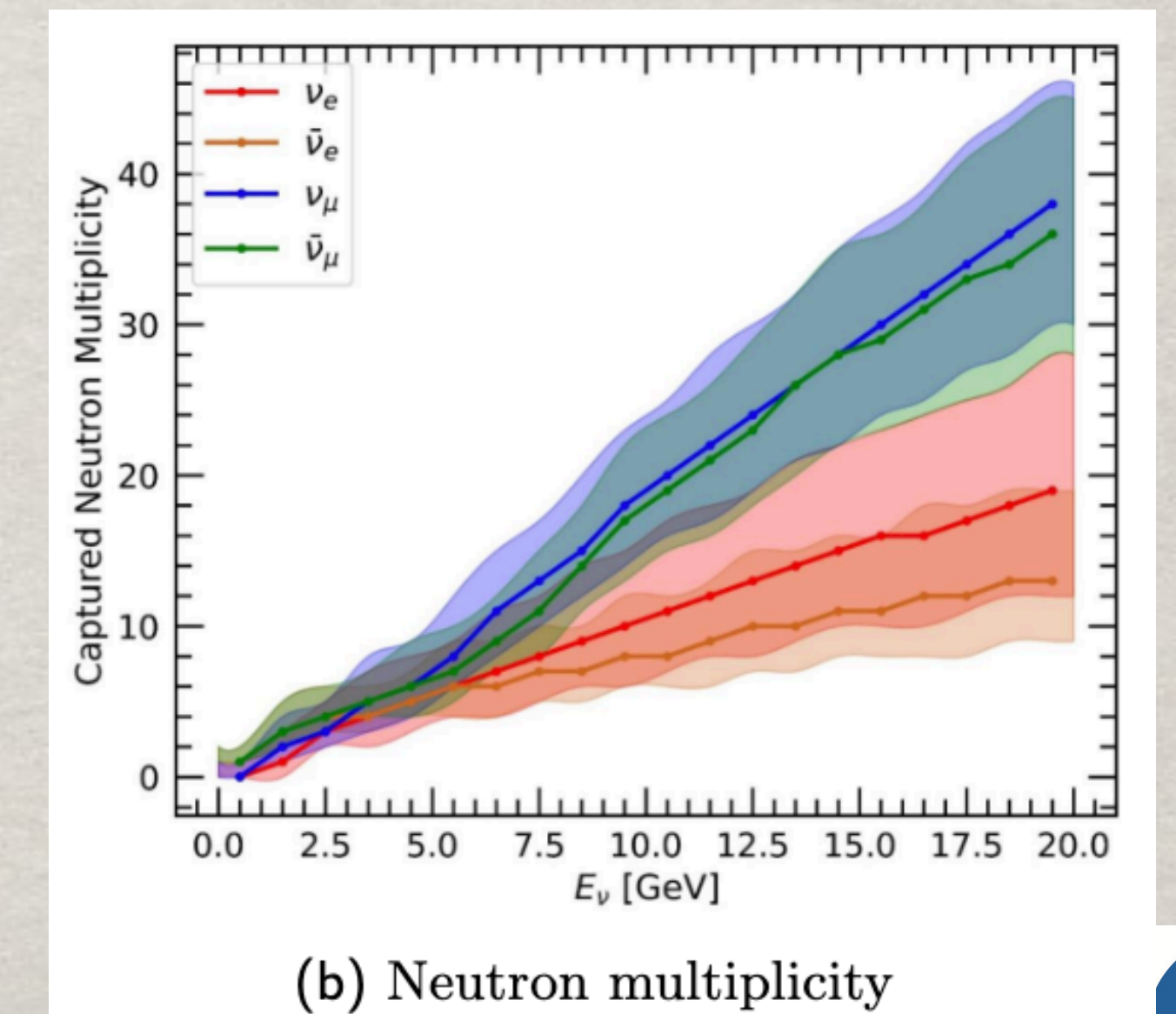
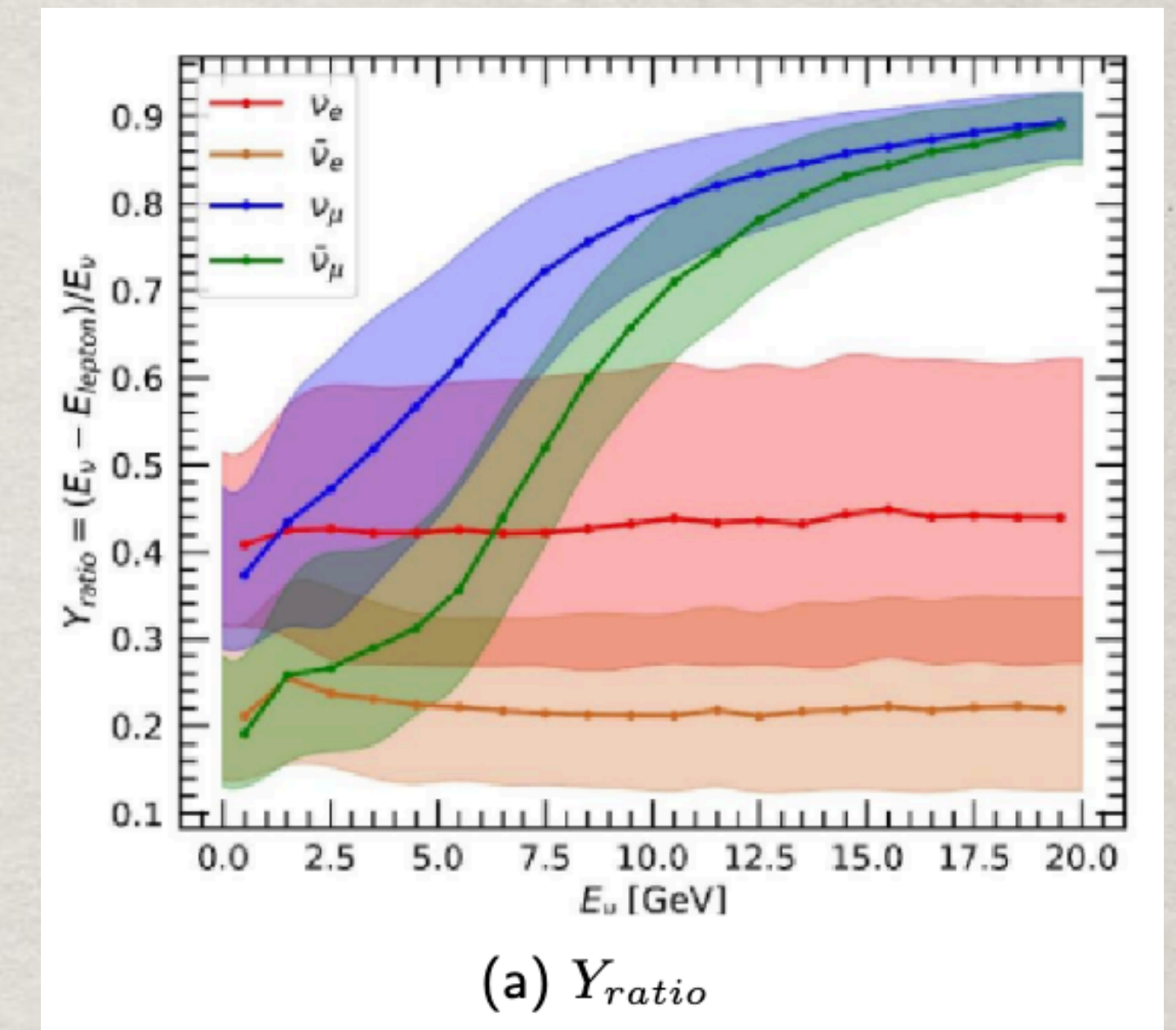
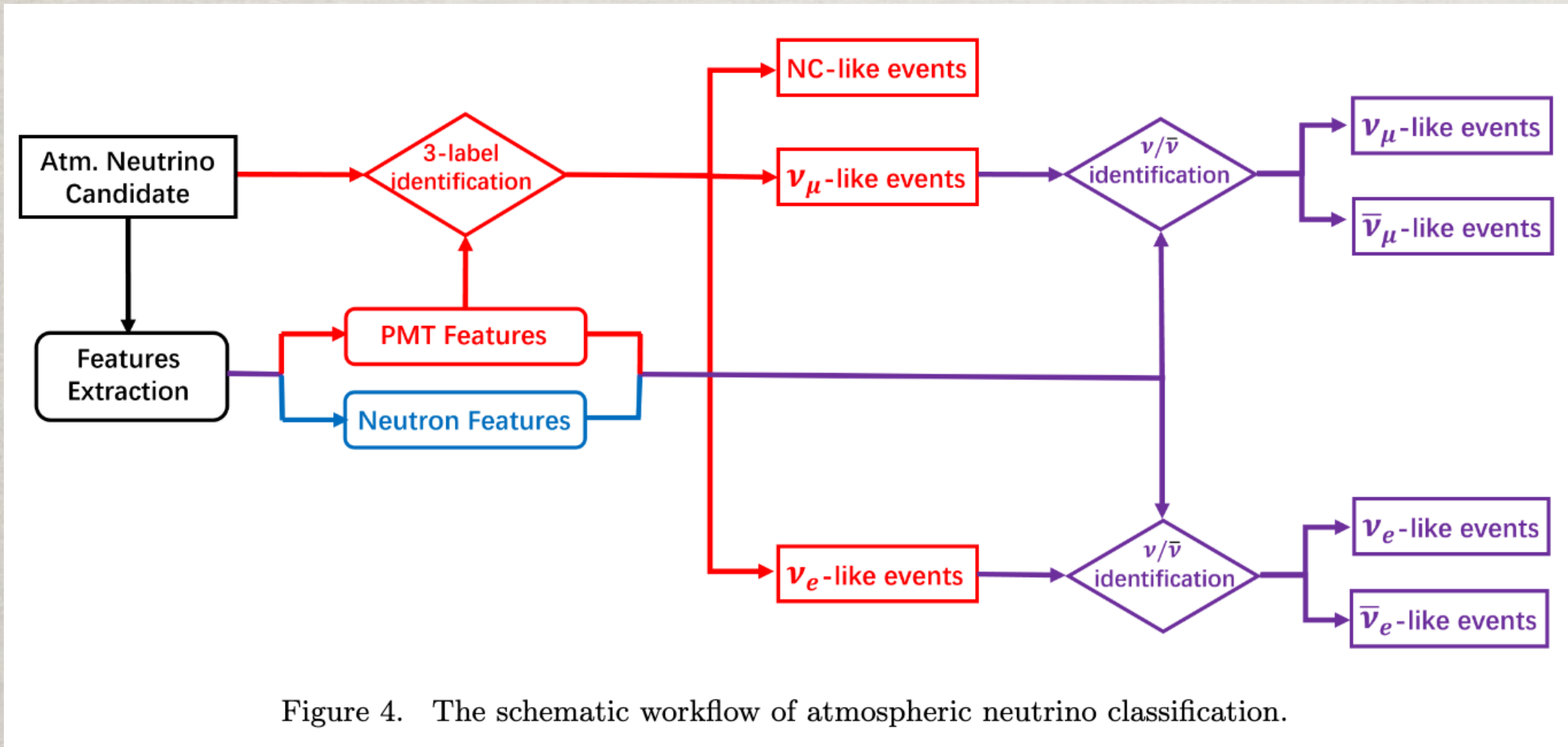
Yellow Book

$$\sigma_{\theta_\mu} = 1^\circ$$

$$\sigma_{\theta_\nu} = 10^\circ$$

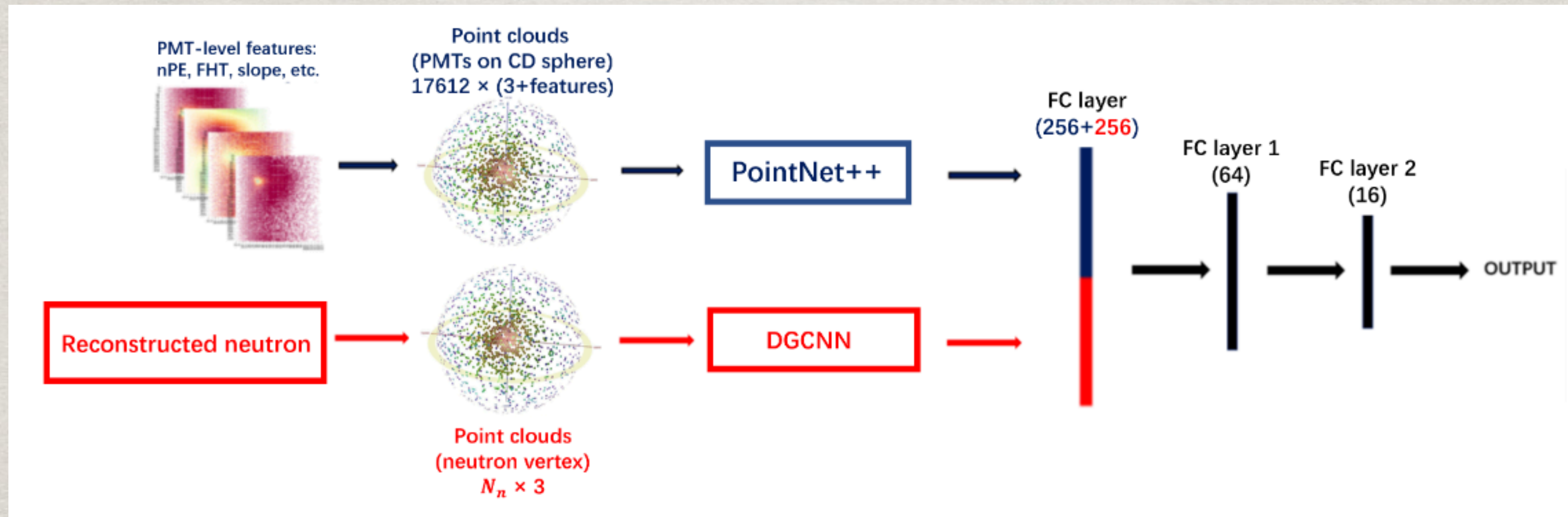


PID STRATEGY



- ❖ Both leptons&hadrons visible, different topology
- ❖ step1: CC-e/CC-mu/NC classification
- ❖ step2: $\bar{\nu}$ vs ν

PID ML INPUT & MODEL



- ❖ PMT features \rightarrow PointNet++ (x, y, z, feature_i...)
- ❖ Neutron candidates \rightarrow DGCNN (x, y, z)

PID PRELIMINARY PERFORMANCE

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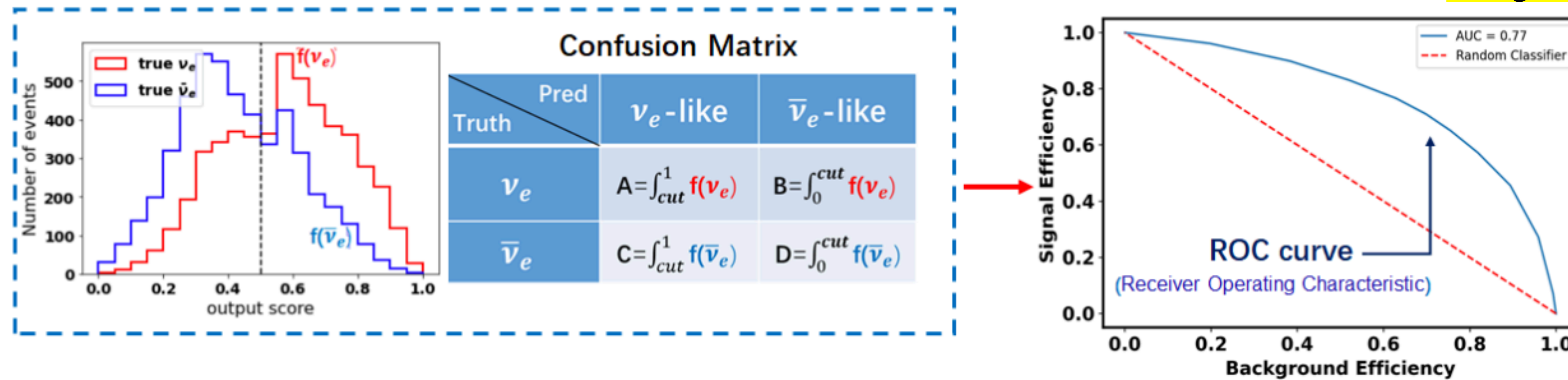
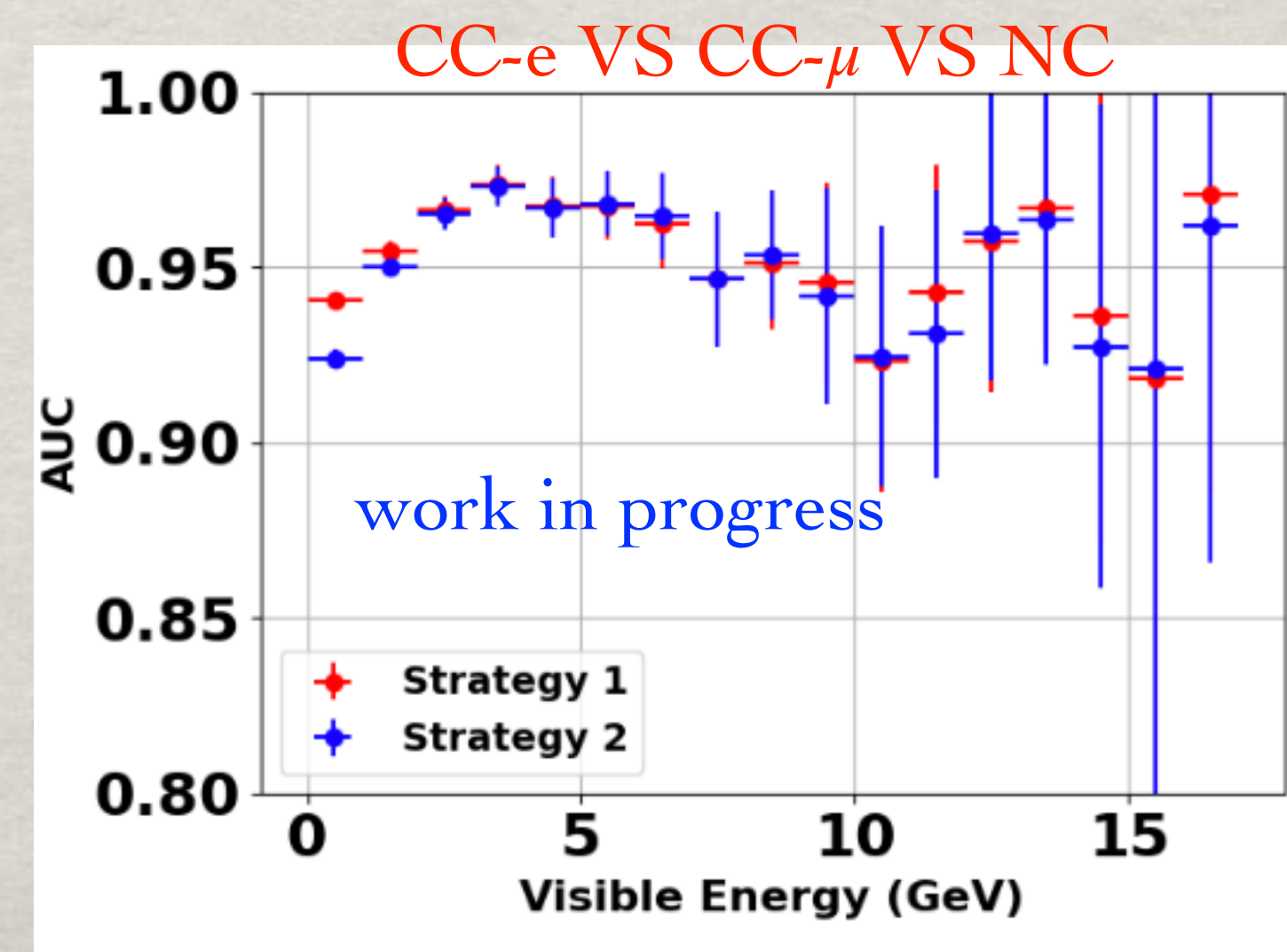
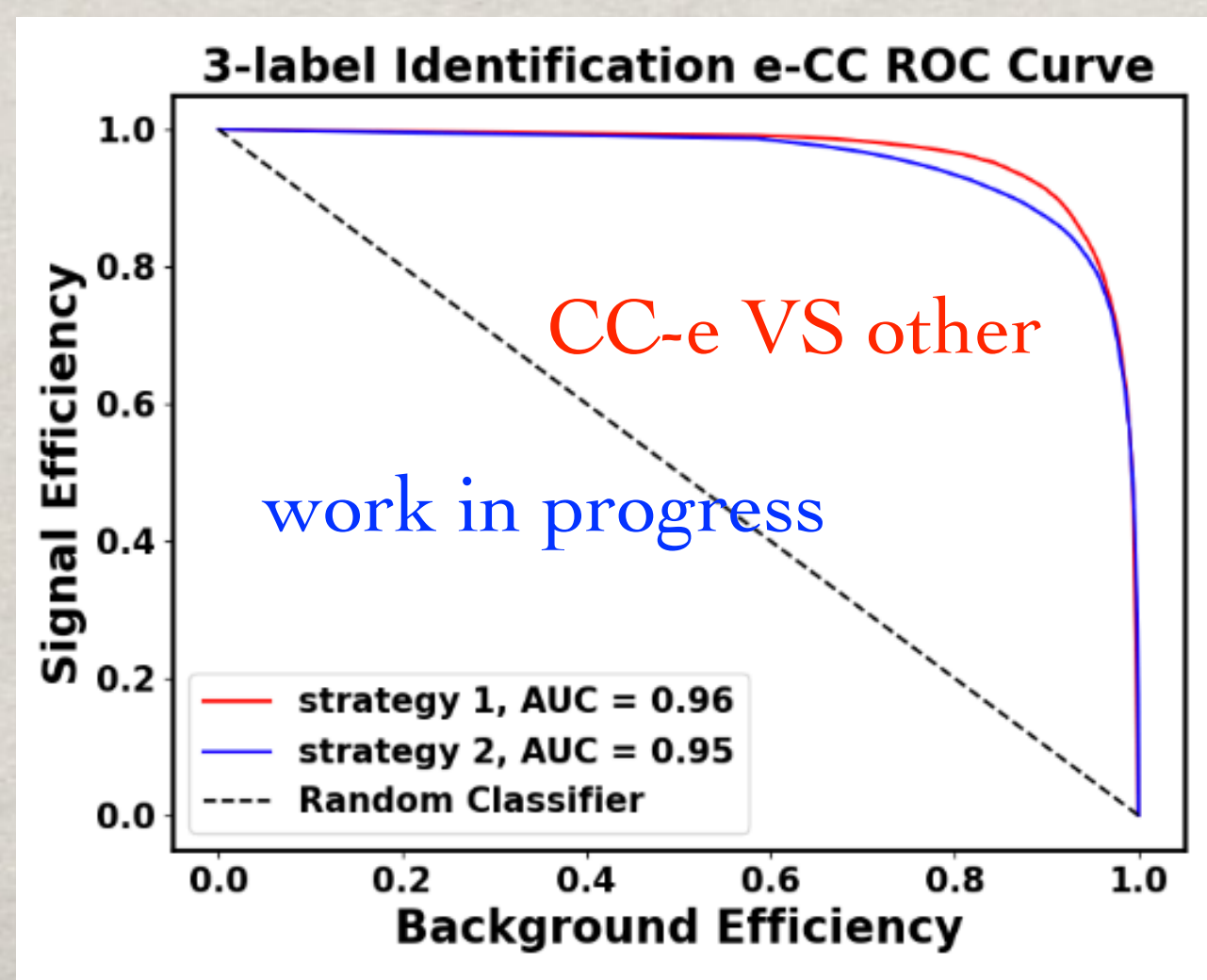


Fig. 8: Illustration of the AUC score using $\nu_e/\bar{\nu}_e$ classification as an example. The AUC score can be viewed as an optimisation of $\nu_e/\bar{\nu}_e$ efficiencies.



PID PRELIMINARY PERFORMANCE

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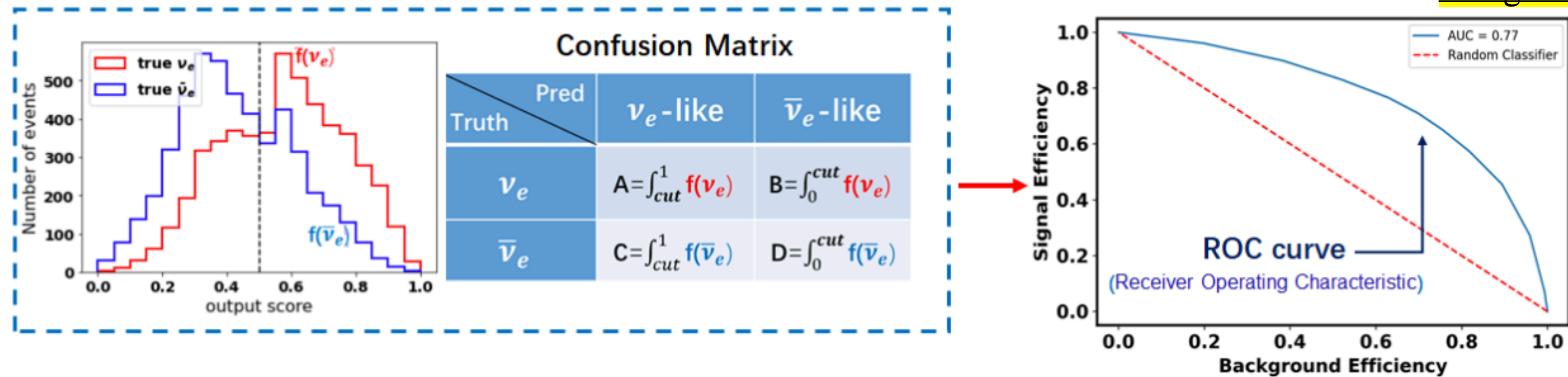
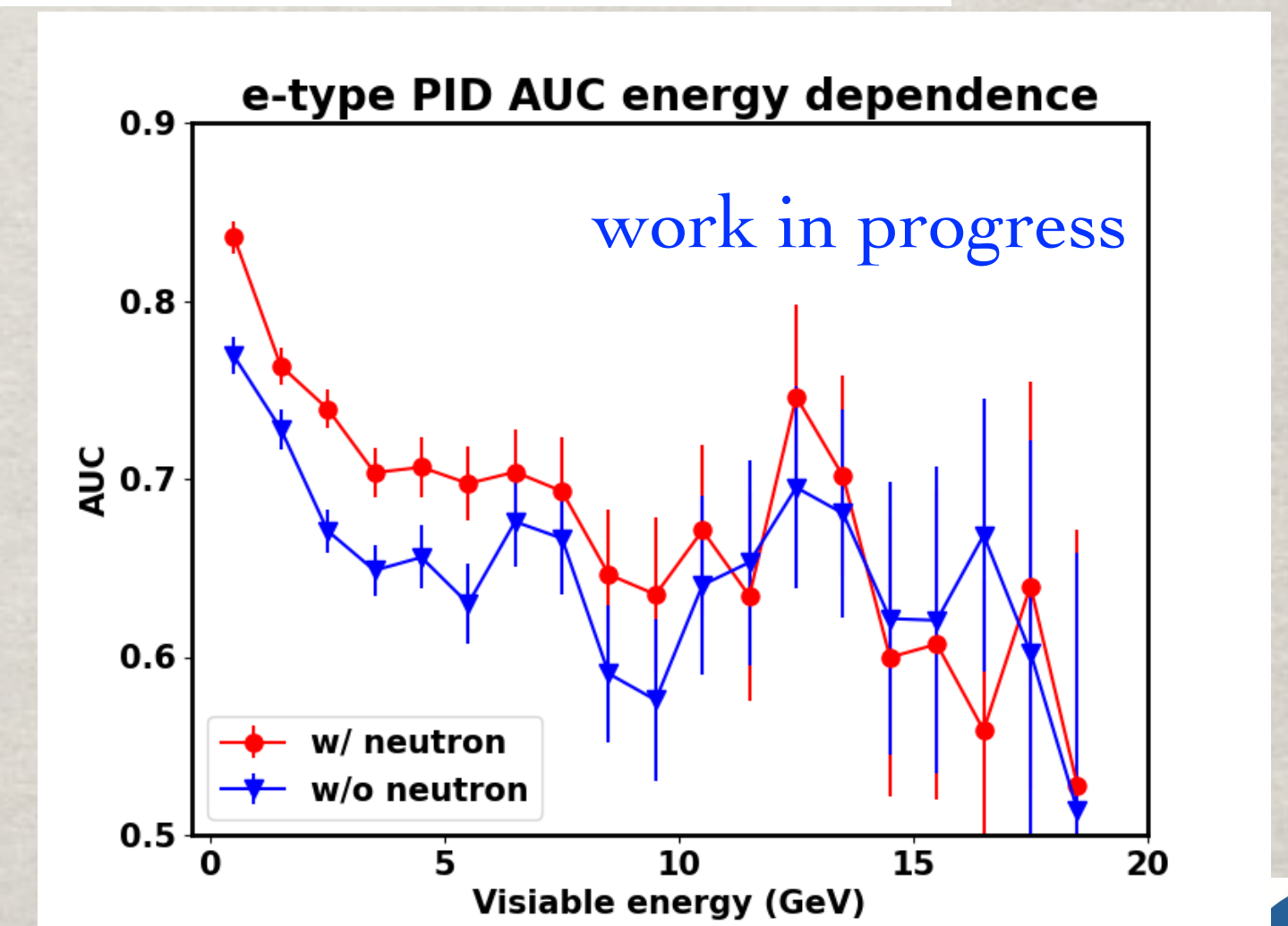
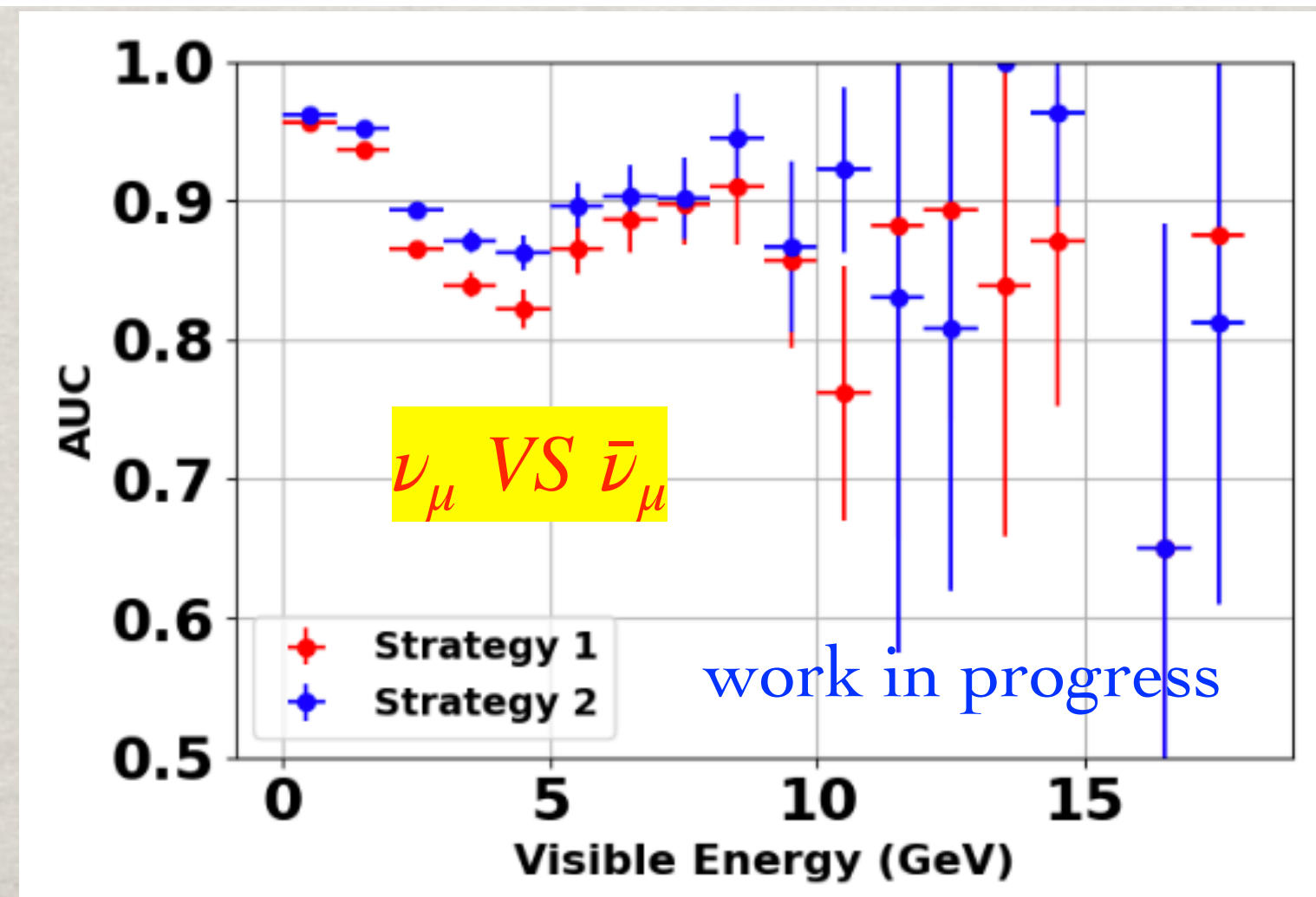
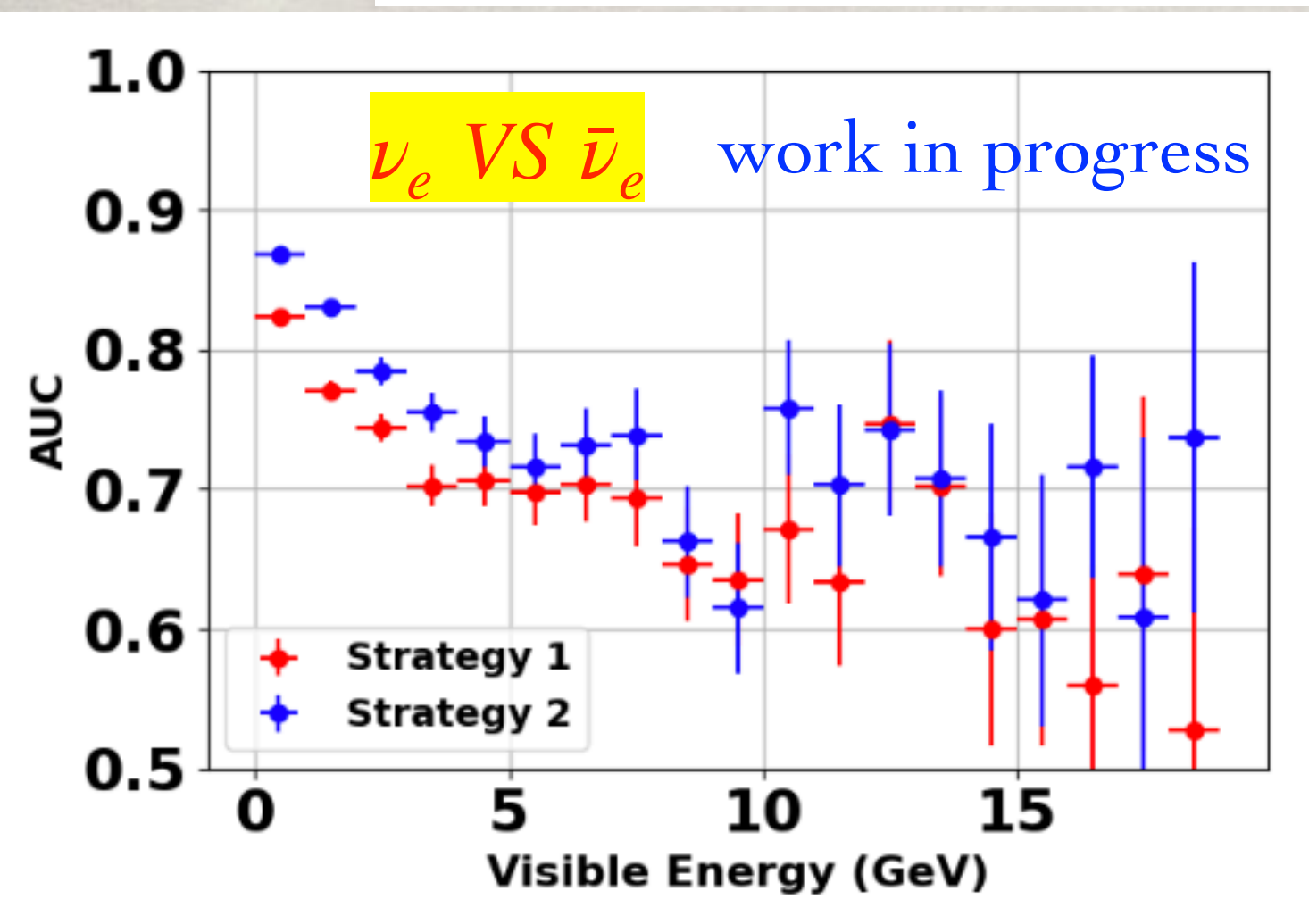


Fig. 8: Illustration of the AUC score using $\nu_e/\bar{\nu}_e$ classification as an example. The AUC score can be viewed as an optimisation of $\nu_e/\bar{\nu}_e$ efficiencies.



MORE INTERESTING TOPICS

- ❖ PMT de-noising, waveform reco
- ❖ ^{14}C pileUp identification
- ❖ Muon classification/combined reco
- ❖ Separation of Scintillation and Cherenkov photons?
- ❖ Multi-target reco?
- ❖ And more...



SUMMARY

- ❖ With $O(10^4)$ PMTs and 20kton LS, JUNO provides a perfect scenario for Machine Learning applications
- ❖ A few examples were presented
 - ❖ MeV reactor ν : PMT waveform/vertex/energy reco
 - ❖ GeV atm. ν : directionality reco and PID
- ❖ More interesting and challenging problems...
- ❖ ML can further enhance the capability of the detector, yielding faster, better results

THANK YOU!



WE ARE HIRING

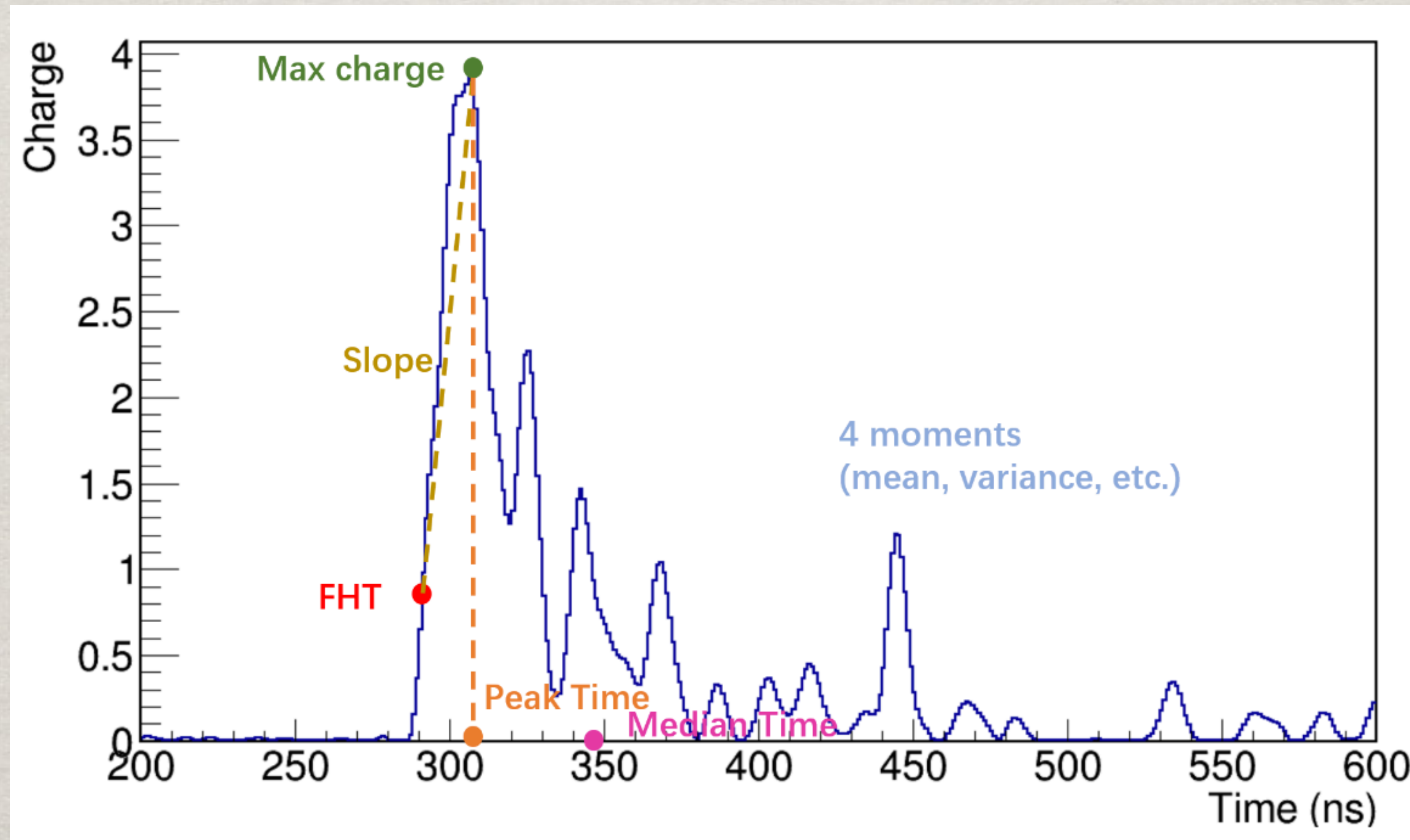
- ❖ Several postDoc openings at JUNO
- ❖ Commissioning, Reconstruction, Physics Analysis and more...
- ❖ Competitive salary!
- ❖ Email: luowm@ihep.ac.cn



BACK UP

INPUTS: PMT FEATURES

❖ Feature variables extracted from PMT waveforms



ATM. NEUTRINOS: DIRECTIONALITY

