# A data-driven method to estimate the $\bar{p}$ background in Mu2e

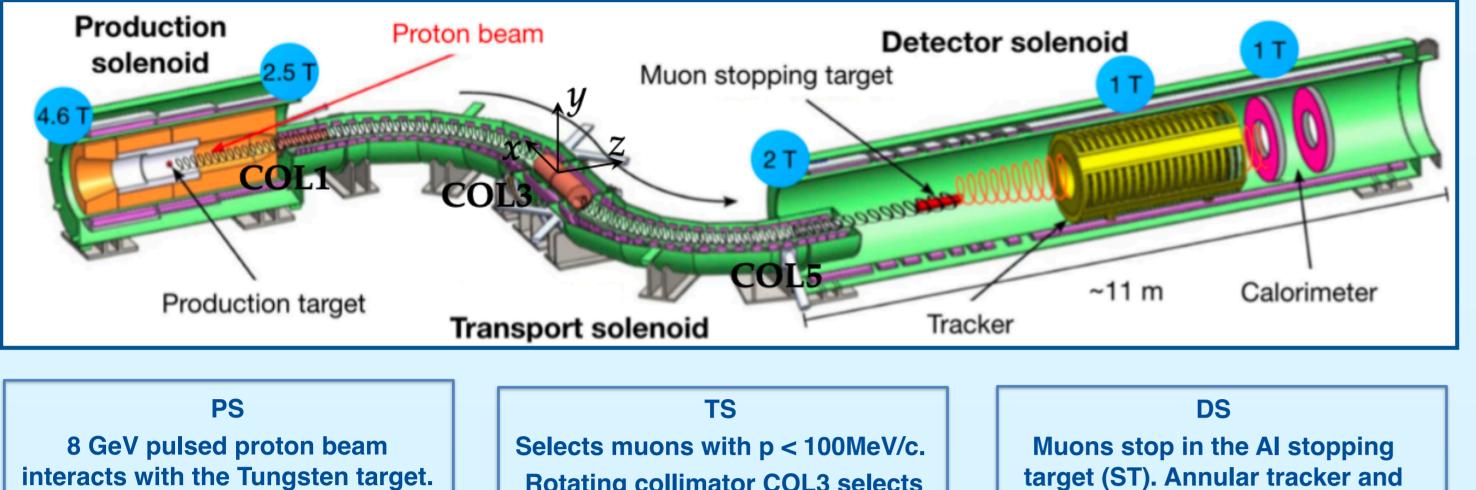
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### **Mu2e: An Overview**

Search for neutrinoless, coherent conversion  $\mu^- N \rightarrow e^- N$  in the field of an Al nucleus by measuring,

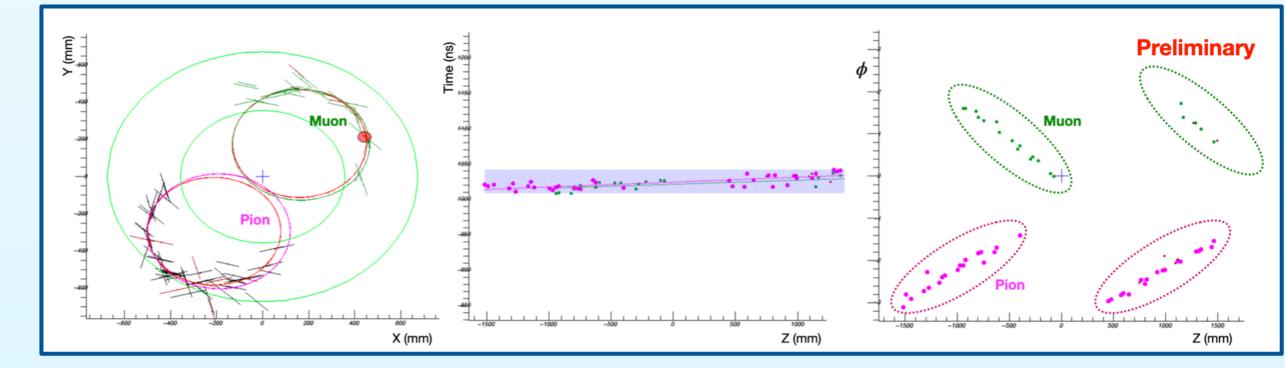
$$R_{\mu e} = \frac{\Gamma(\mu^- + N(Z, A) \rightarrow e^- + N(Z, A))}{\Gamma(\mu^- + N(Z, A) \rightarrow \nu_\mu + N(Z - 1, A))}$$

Signal: Monochromatic, 104.97 MeV/c e<sup>-</sup>.



Simple time clustering alone is insufficient for  $\bar{p}$  annihilation events as the tracks are mostly simultaneous in time.

However, hits from different particles could be well separated in  $\phi = tan^{-1}(y/x)$ .



y vs x, time v/s z and  $\phi$  v/s z views for an example  $p\bar{p}$  annihilation at the ST event.

Mu2e Run I will operate in low intensity mode:  $1.6 \times 10^7$ protons/pulse,  $\sim 2.5 \times 10^4$  $\mu^{-}/pulse$  in ST. For high intensity mode, the

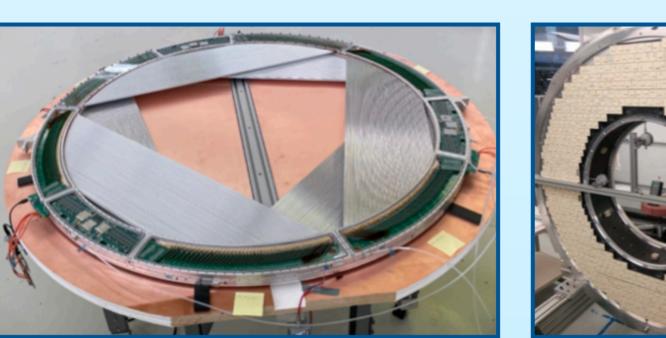
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Mostly produces pions.

**Rotating collimator COL3 selects**  $\mu^-$  or  $\mu^+$  beam.

The tracker consists of 18 stations with 1152 straws per station. The straws are filled with 80%:20% Ar :  $CO_2$  mixture.

The calorimeter has 2 disks covering radii 37 cm-66 cm. Each disk has 674 pure CsI crystals.



calorimeter to detect the

conversion  $e^-$  (CE).

The expected Run I 5 $\sigma$  discovery sensitivity is  $R_{\mu e} = 1.2 \times 10^{-15}$ . If no signal, the expected upper limit is  $R_{\mu e} < 6.2 \times 10^{-16}$  at 90% CL.

The estimated  $\overline{p}$  background for Run 1 is  $0.01 \pm 0.003(stat) \pm 0.010(syst)^*$ . The systematic error is dominated by the uncertainty on the  $\overline{p}$  production cross section.

### $\bar{p}$ background in Mu2e

 $\overline{p}$ s are produced in the pW interactions in the PS.

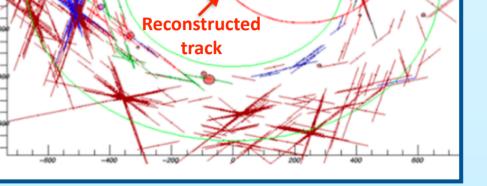
 $p\overline{p}$  annihilation at ST can produce  $e^-$ s by  $\pi^0 \to \gamma\gamma$  decays followed by  $\gamma$ conversions and  $\pi^- \rightarrow \mu^- \overline{\nu}$  decays followed by the  $\mu^-$  decays.

It cannot be suppressed by the time window cut because  $\overline{p}$ s are much slower than other beam particles.

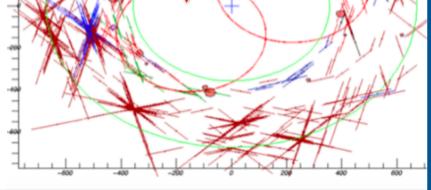
Absorber elements placed at entrance and centre of the TS to suppress the  $\overline{p}$ s.

corresponding numbers are  $\times 2.5$  higher.

We tested the reconstruction procedure with datasets containing  $p\overline{p}$  annihilation events mixed with low and high intensity backgrounds as well.



**Default Reconstruction** 



**CHEP 2024** 

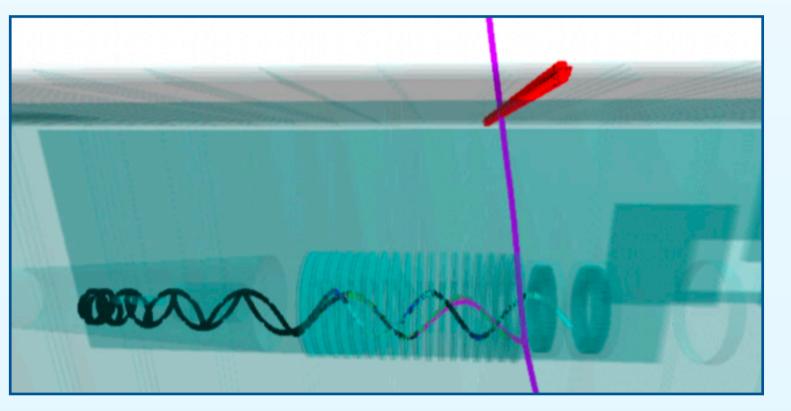
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**New Reconstruction** 

## **Contribution of cosmic rays to multi-track events**

A cosmic ray veto (CRV) system built from scintillator counters surrounds the DS to identify the cosmic rays.

- The multi-track events from cosmic rays are: 1) Muons trapped in the magnetic bottle structure of the DS or interacting with calorimeter producing  $e^{-}/e^{+}$ , which then travel up towards the ST and back to the calorimeter end.
- 2) Muons interacting with ST producing  $e^{-}/e^{+}$ .



Cosmic  $\mu^-$  hits the CRV scintillation bars (shown in red), performs helical motion in the DS back and forth.

Thus, most cosmic multi-track events are made of an upstream and downstream moving leg of the same particle while  $\bar{p}$  annihilation at the ST gives multiple particle tracks moving

 $p\overline{p}$  annihilation at ST can give multi- tracks final state with p  $\sim$  100 MeV/c for each track at much higher rate than signal-like  $e^{-}$ . From GEANT4 simulations,

$$\frac{N_{e^-perMeV}(110 > p > 90MeV/c)}{N_{multi-track}(p \ge 80MeV/c)} \approx \frac{1}{50}$$



### **Mu2e event reconstruction**

Mu2e event reconstruction is optimised for single  $e^{-}$  track events.

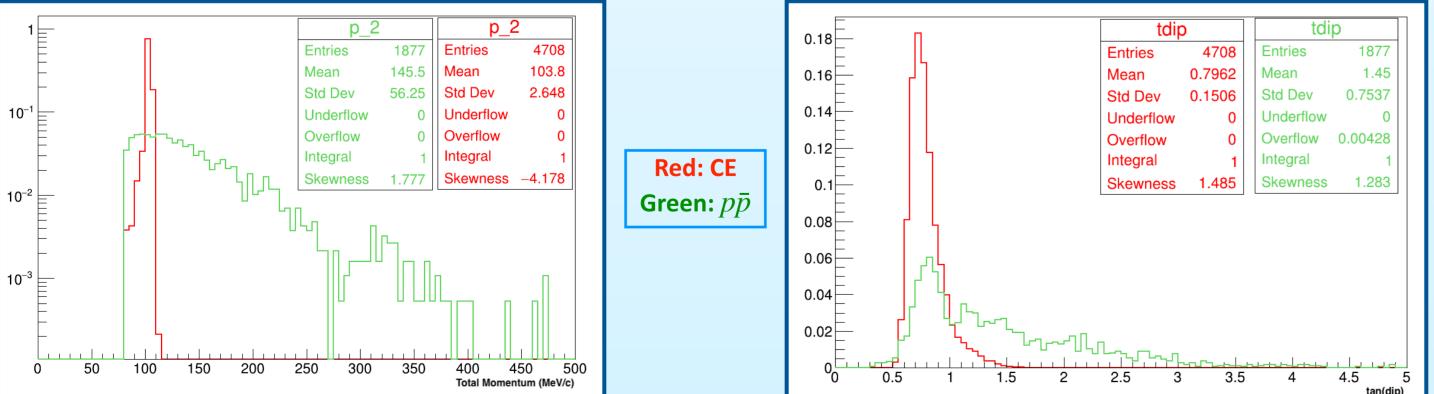
From MC studies, 90% of hits in an event are from low energy  $e^-, e^+, p_-$ 

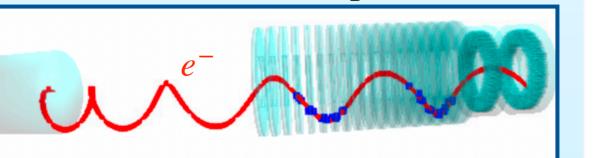
They are flagged background prior to reconstruction.

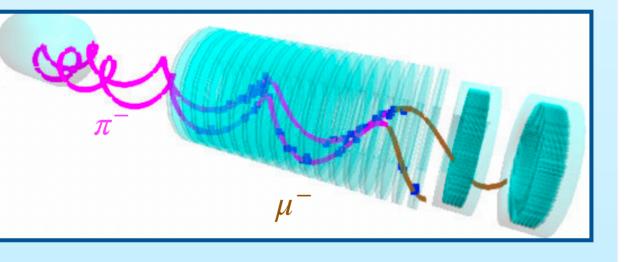
Assuming hits by same particle have close reconstructed times, they are time clustered. *TimeClusters* are input for pattern recognition which search for 3-D *Helices*.

Finally, the reconstructed track parameters are determined by the Kalman fit.

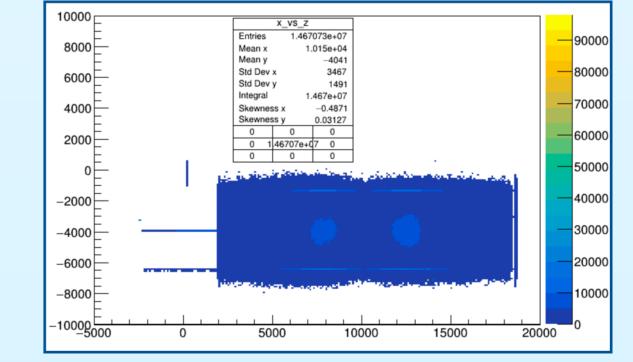
1		p	p_2		p_2	
		Entries	1877	Entries	4708	
-		Mean	145.5	Mean	103.8	
_	4	Std Dev	56.25	Std Dev	2.648	

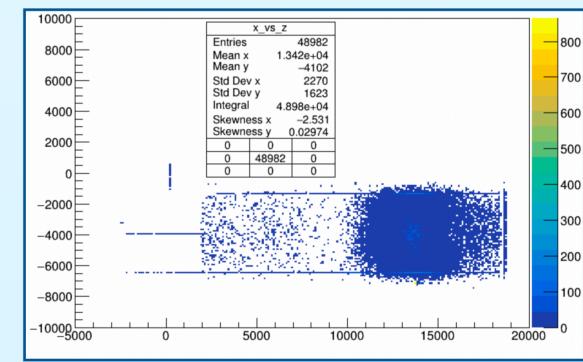






downstream from the ST.





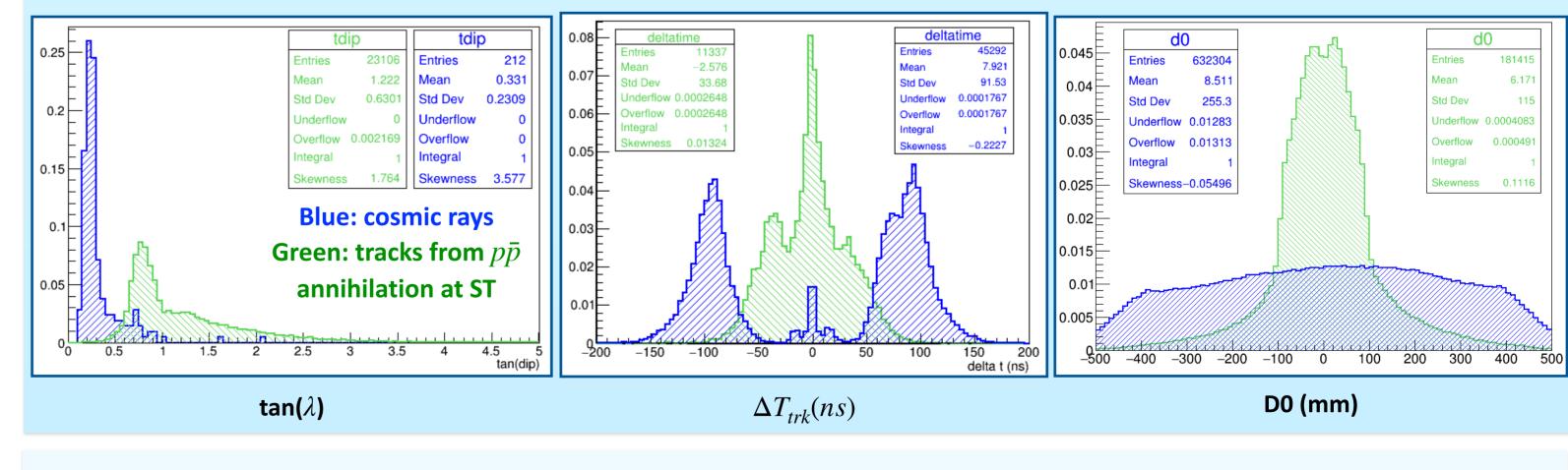
### x v/s z view for all cosmic ray events

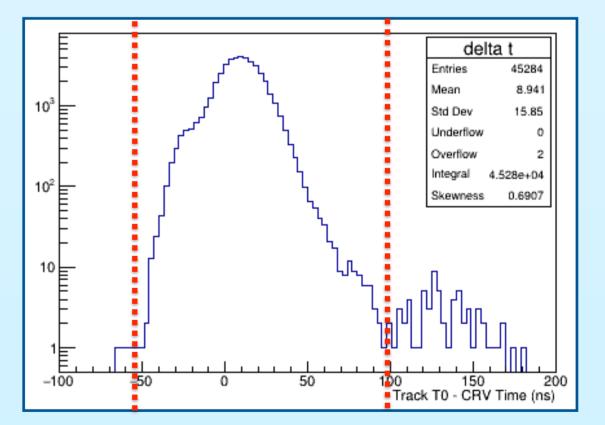
x v/s z view for multi-track cosmic ray events

About 99.98% of these multi-track events can be vetoed using the signal from the CRV. Cosmic event candidates are identified by the timing window

$$-50 < \Delta T_{CRV} < 100 ns$$
 where  $\Delta T_{CRV} = T_0 - T_{CRV}$ .

For events with no matched CRV signal, we have identified track parameters: pitch( $tan(\lambda)$ ), impact parameter (D0) that can be used to distinguish tracks from cosmic muons from  $\bar{p}$  background events.





The default algorithms to flag background hits and form *TimeClusters* use an ANN trained for efficient CE search, which removes a large fraction of pion and muon hits. This reduces the efficiency of reconstructing tracks from  $p\bar{p}$  annihilation significantly. We developed new algorithms, without any ANN, highly efficient for a wide spectrum of particle topologies.

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### Conclusion

We developed a novel data-driven approach to constrain the  $\overline{p}$  background in Mu2e.

Tested the reconstruction with pure  $p\overline{p}$  annihilation events and with  $p\overline{p}$  annihilation events mixed with low and high intensity backgrounds.

We reconstruct about 42% of the multi-track events from  $p\bar{p}$  annihilation at the ST where each track has a total momentum > 80 MeV/c and makes at least 20 hits in the tracker.

Compared to the default reconstruction, the number of multi-track events increased by  $\times$  2.1 times.

Currently, we are working on improving the reconstruction further and getting a final estimate on the  $\bar{p}$  background.