



Q-Pix:

pixel-based charge readout for kton scale LArTPC

Shion Kubota for Q-Pix Consortium
Harvard University / The University of Manchester

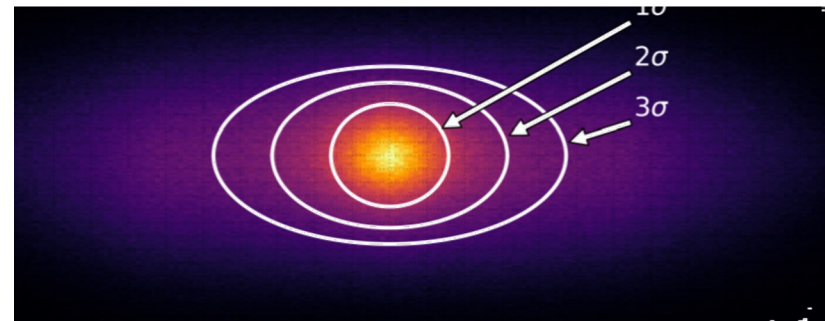
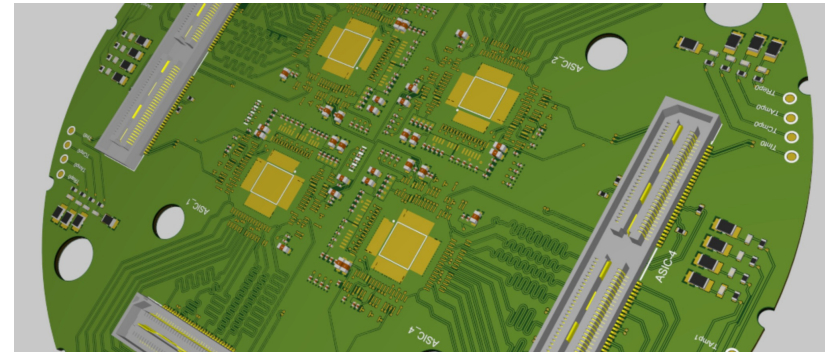


Contents

1. LArTPC
 - a) How it works
 - b) Examples

2. Q-Pix
 - a) Pixel-based readout technology
 - b) How it works with toy example
 - c) Q-Pix demonstration with commercial parts

3. Physics studies with Q-Pix
 - a) Beam studies
 - b) Supernova studies
 - c) Solar studies





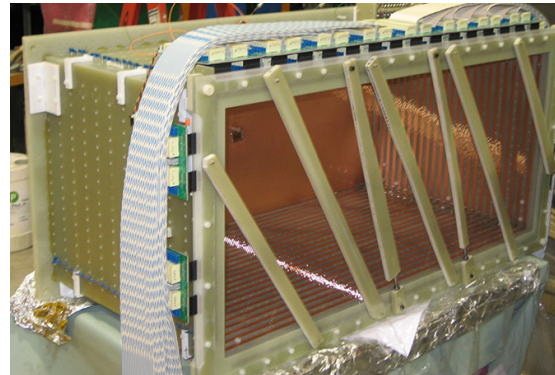
Liquid Argon Time Projection Chamber (LArTPC)

Traditional wire-based LArTPC has been used in many different experiments

DUNE



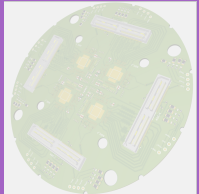
MicroBooNE



ArgoNeut

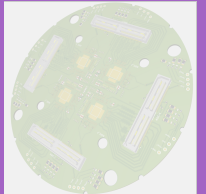
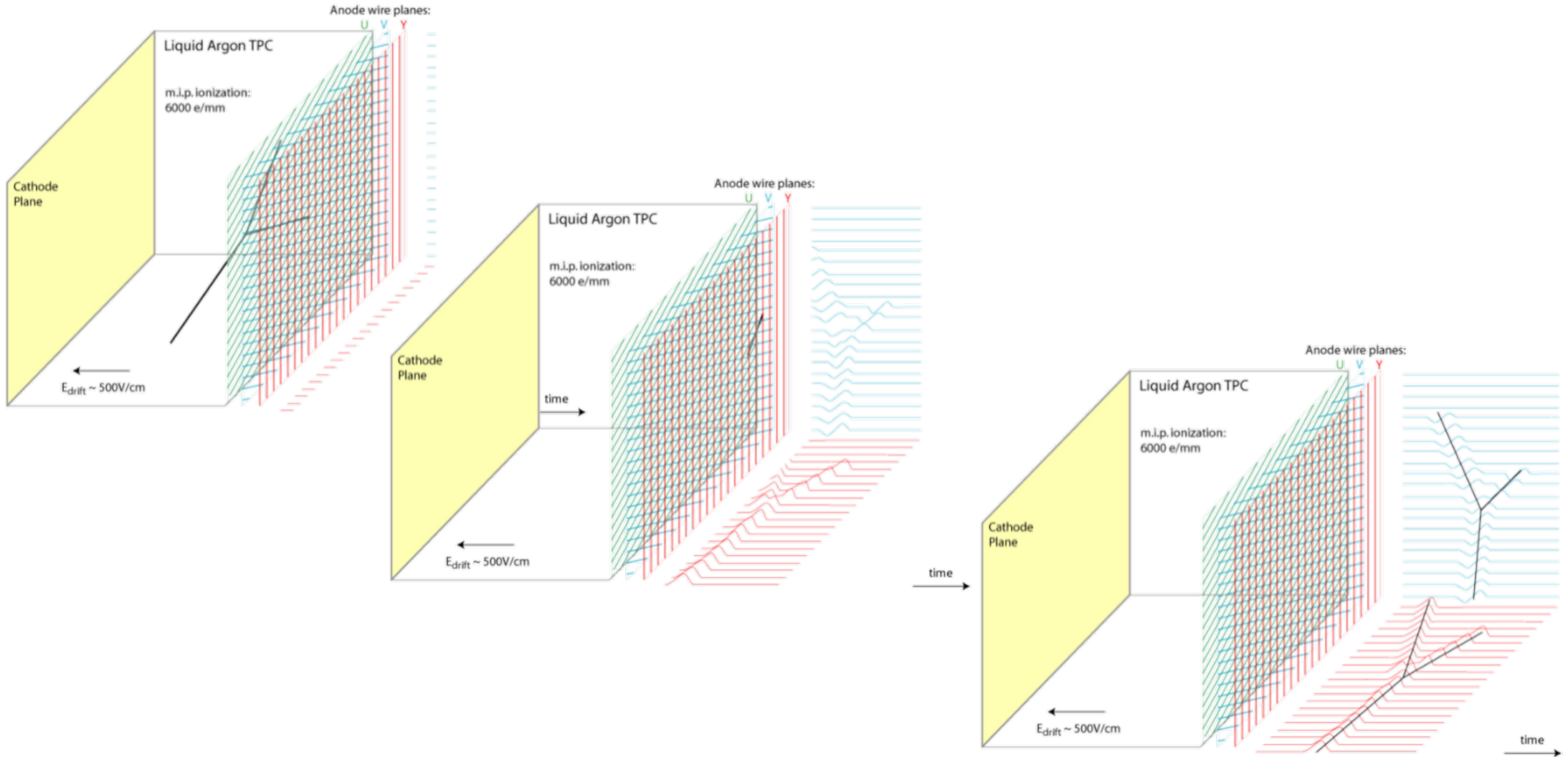


ICARUS



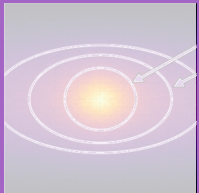
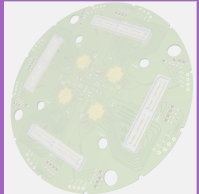
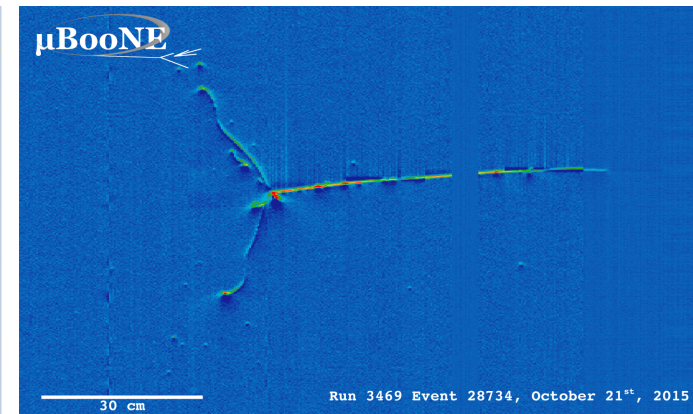
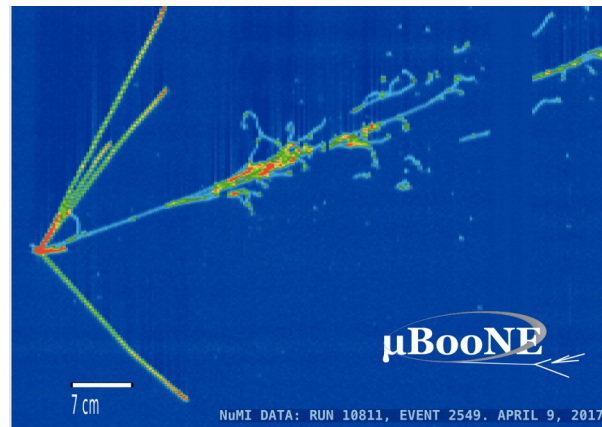
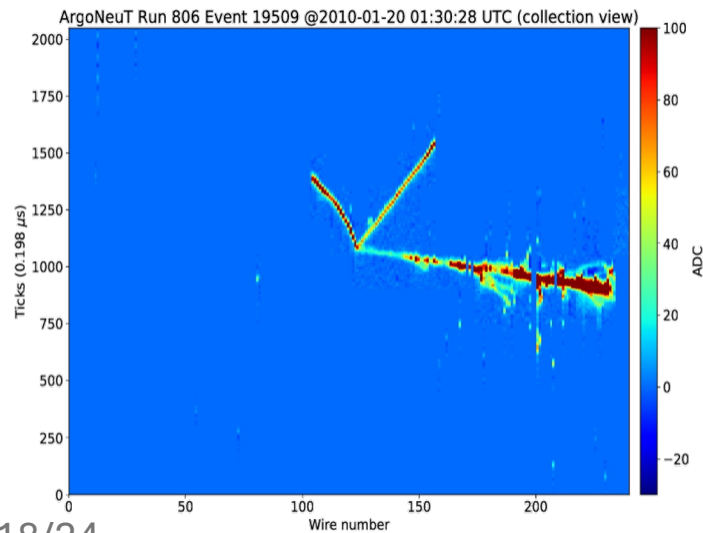
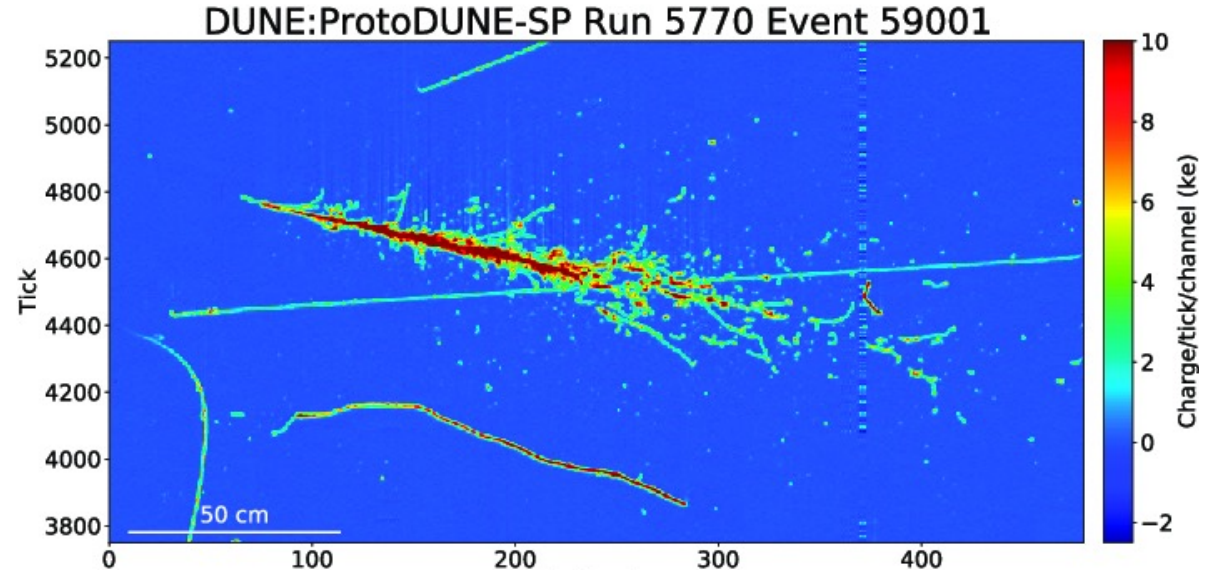
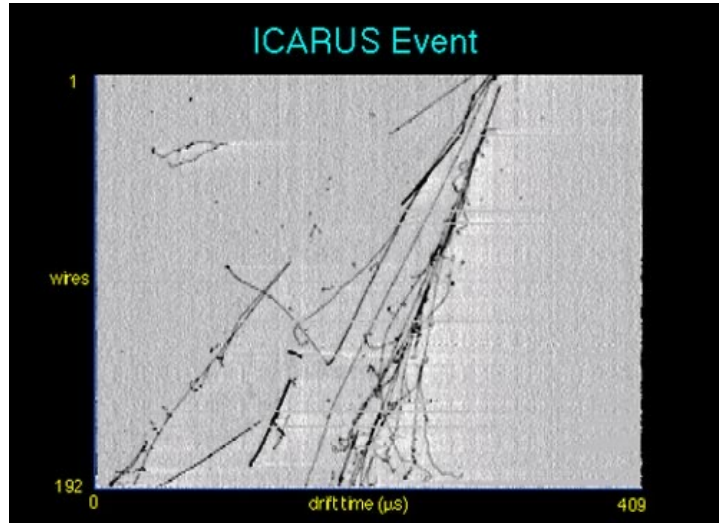


Liquid Argon Time Projection Chamber (LArTPC)





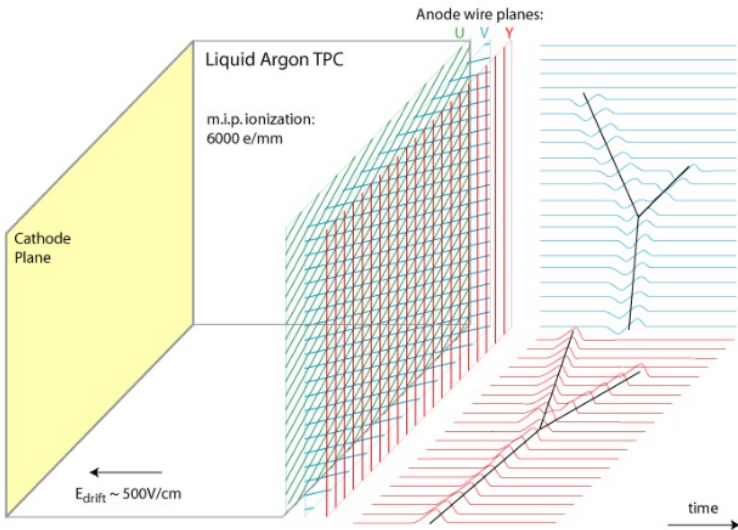
Liquid Argon Time Projection Chamber (LArTPC)



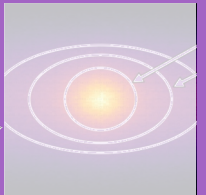
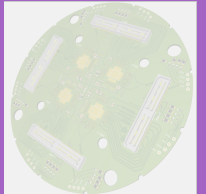
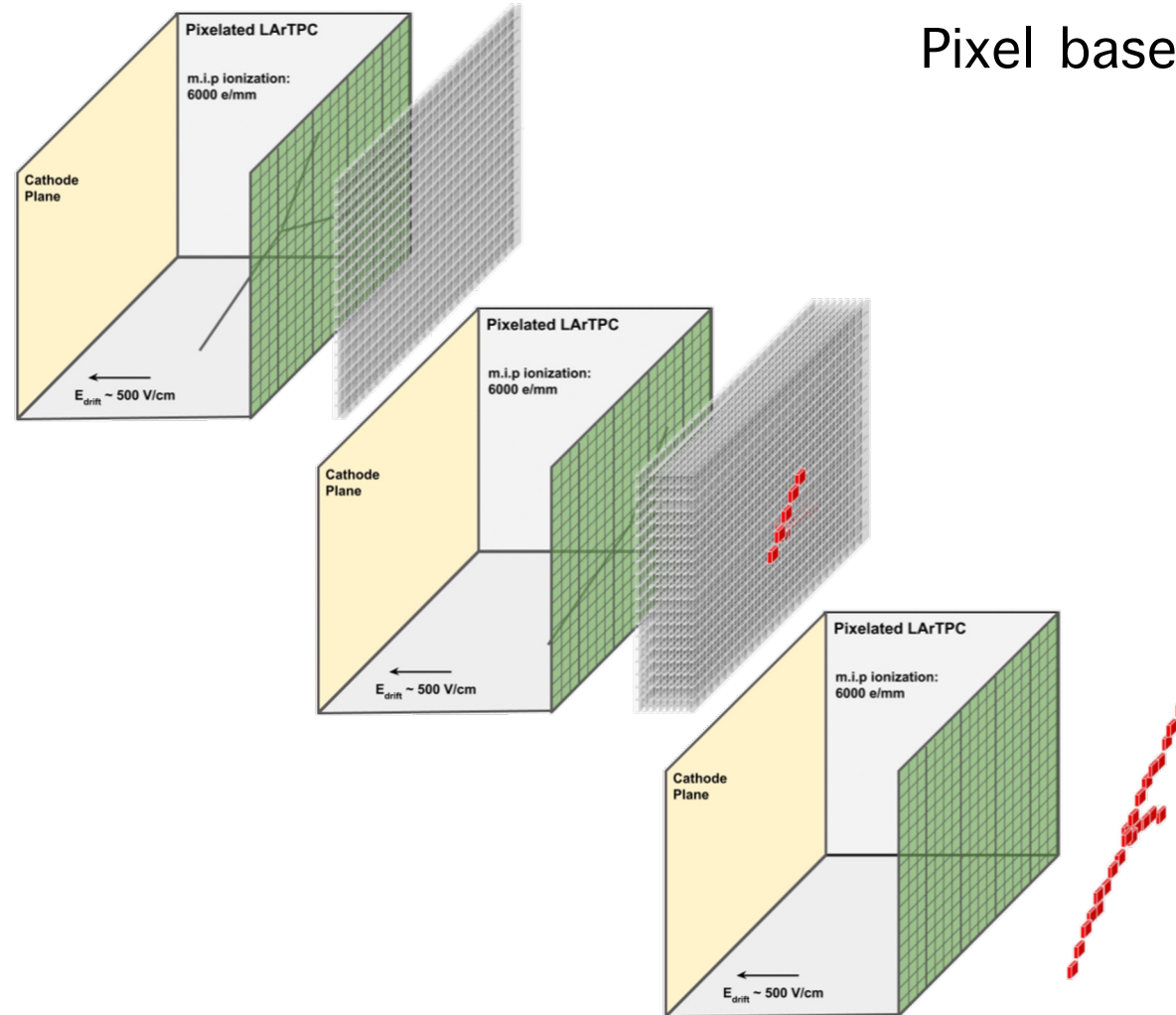


Wire vs Pixelated LArTPC

Wire based



Pixel based





Pixelated LArTPC – why not!

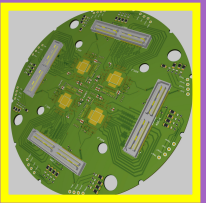
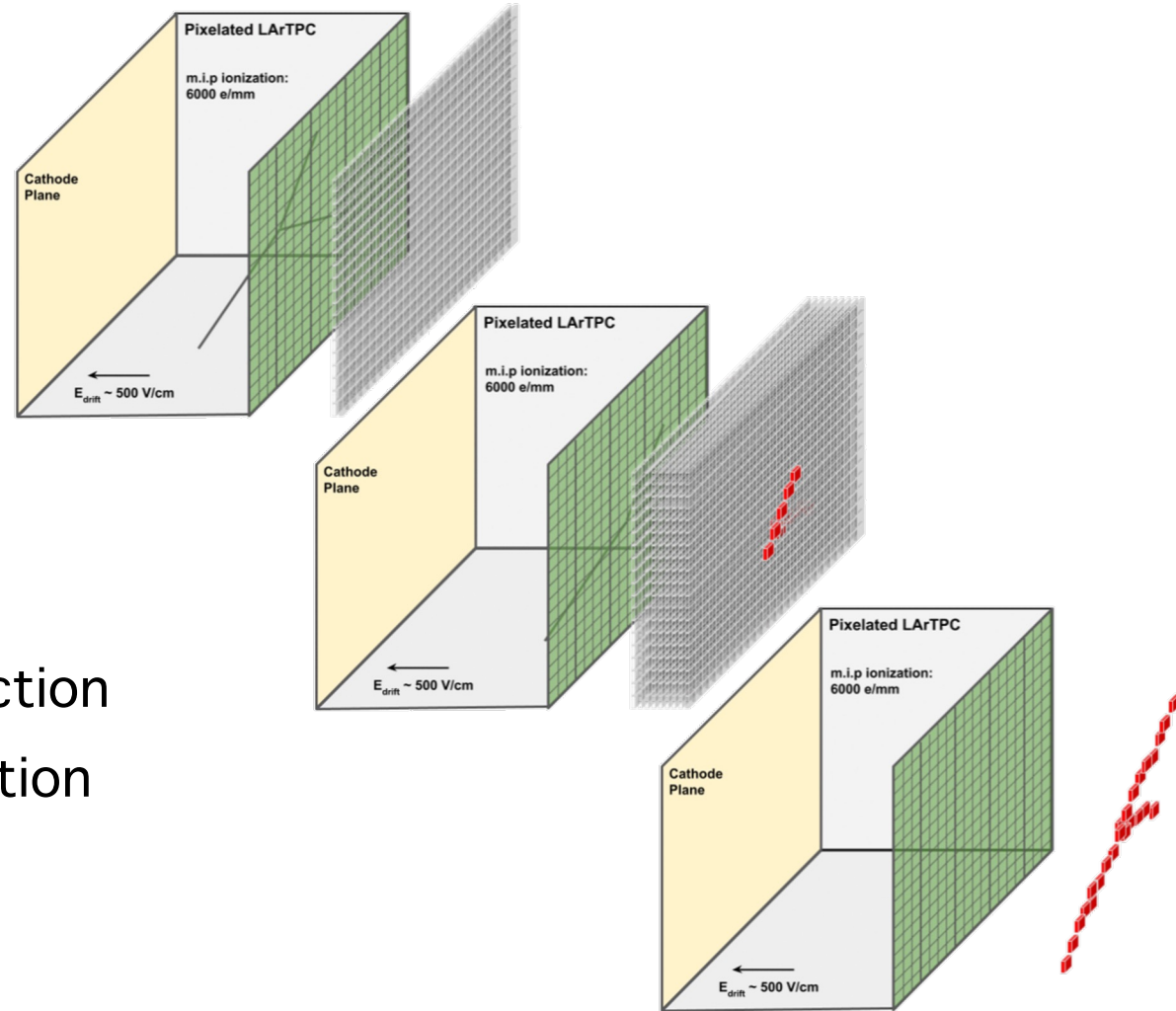
Challenge:

Great increase in

- 1) the number of channels
- 2) the amount of data

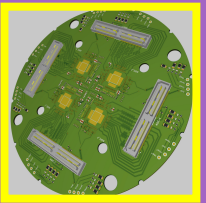
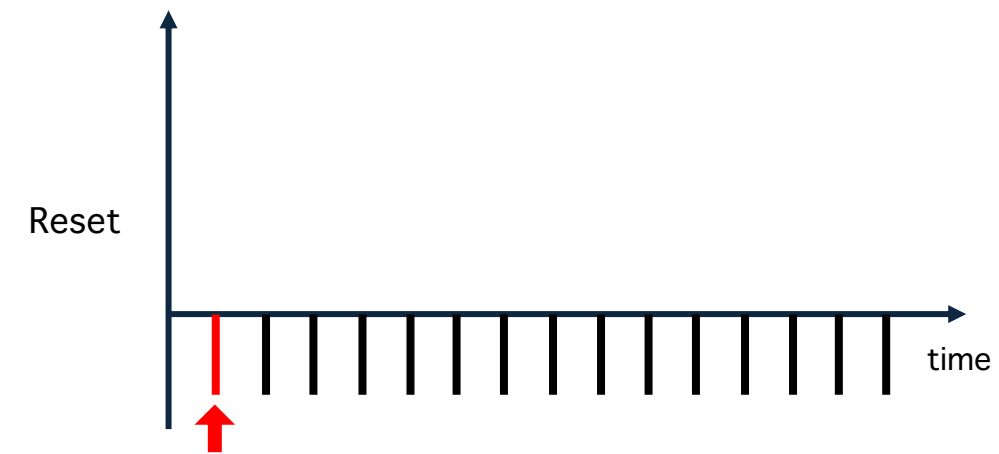
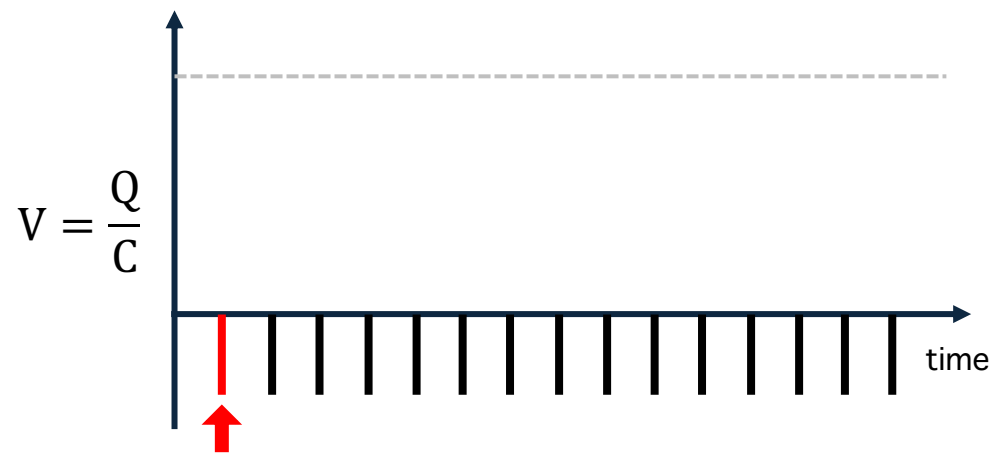
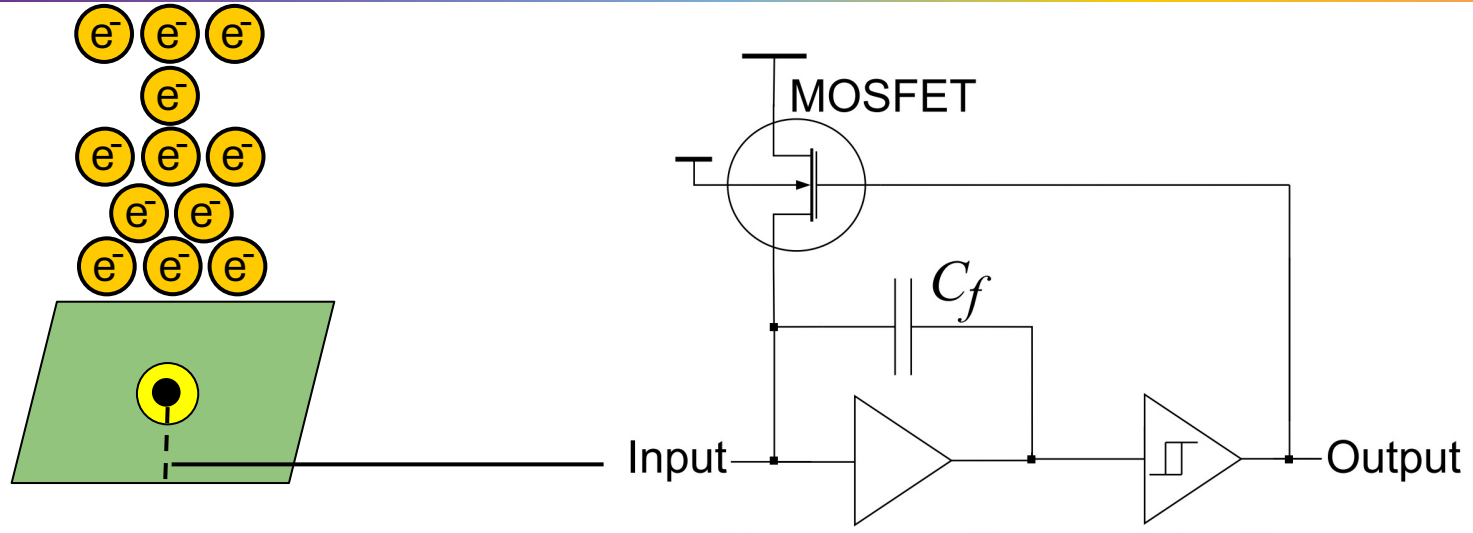
Solutions:

- 1) Electronic principle of least action
- 2) New way to quantize information



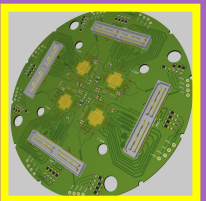
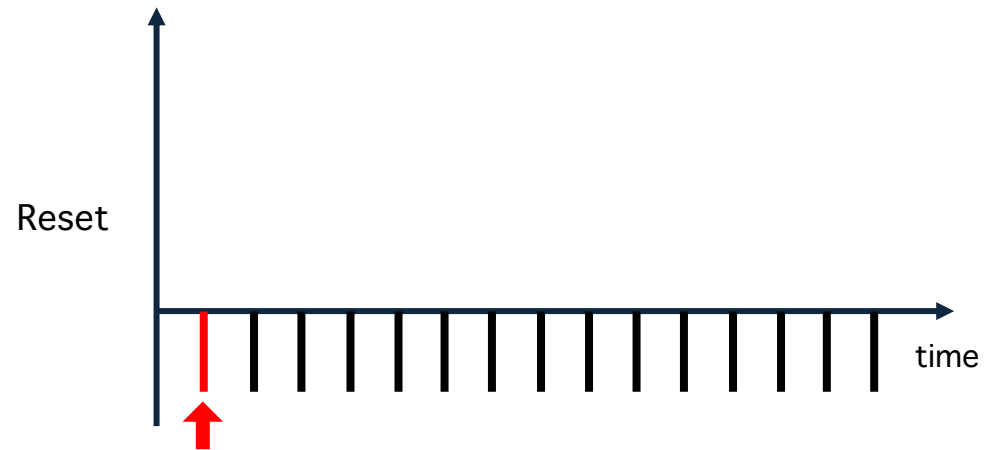
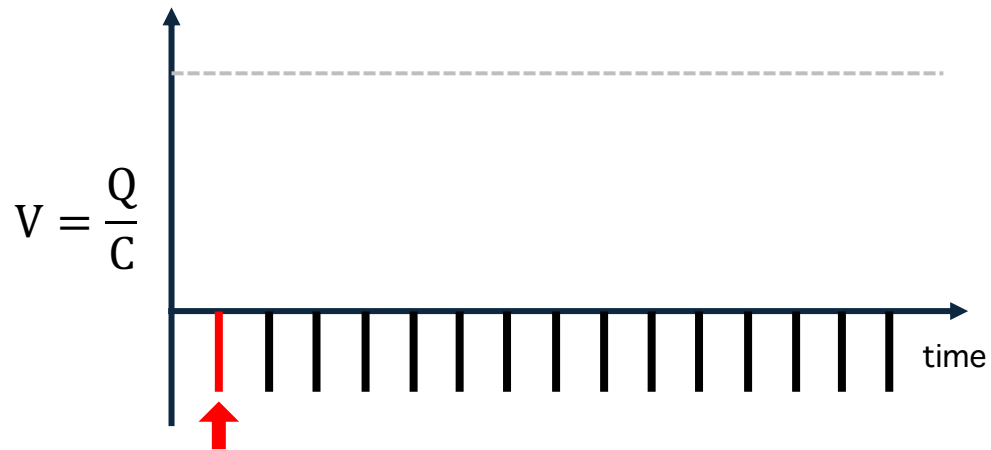
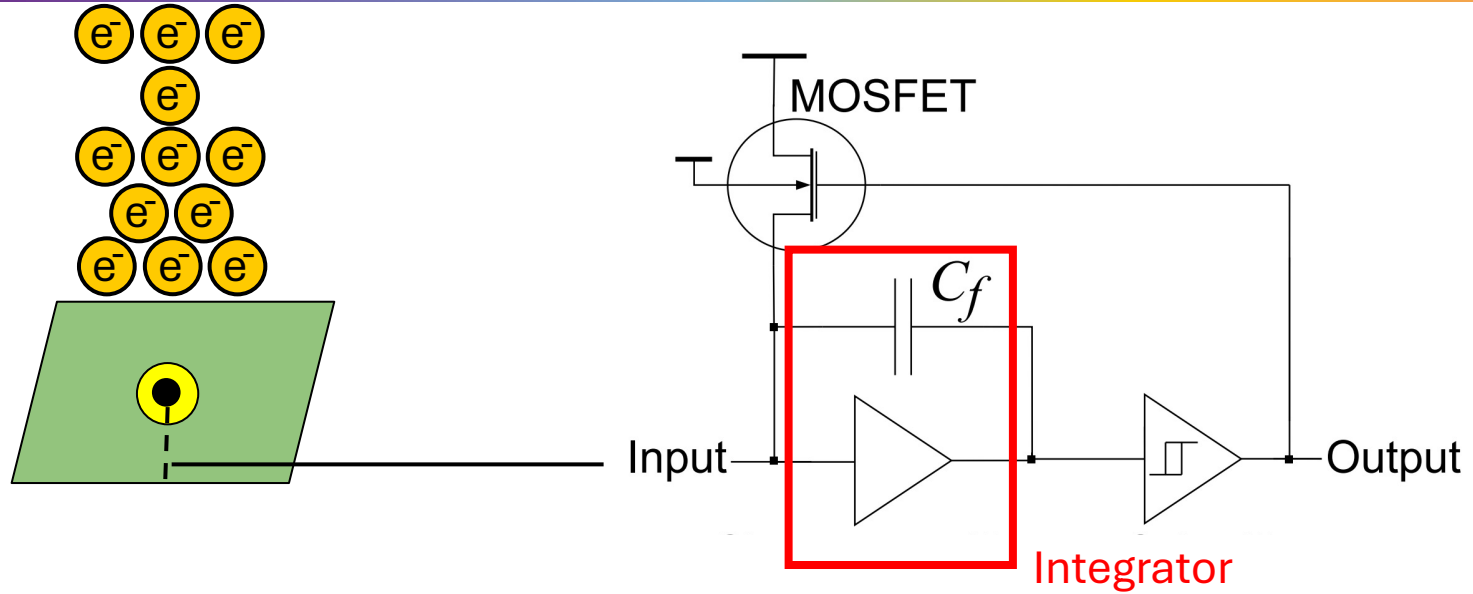


Toy Example



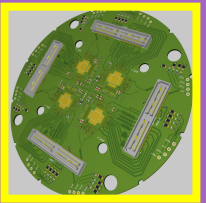
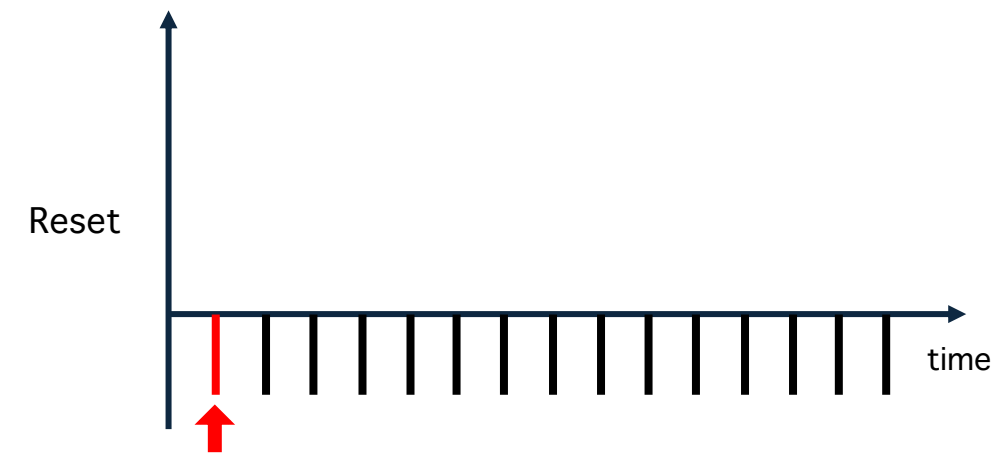
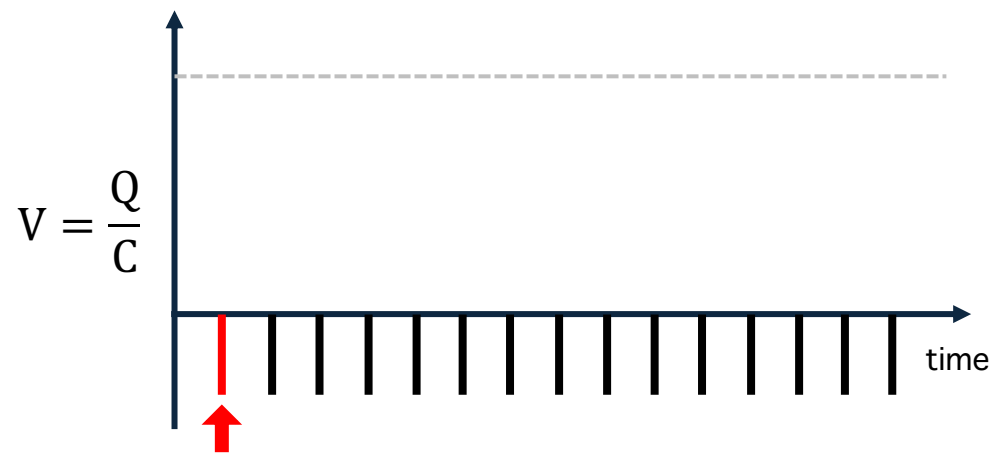
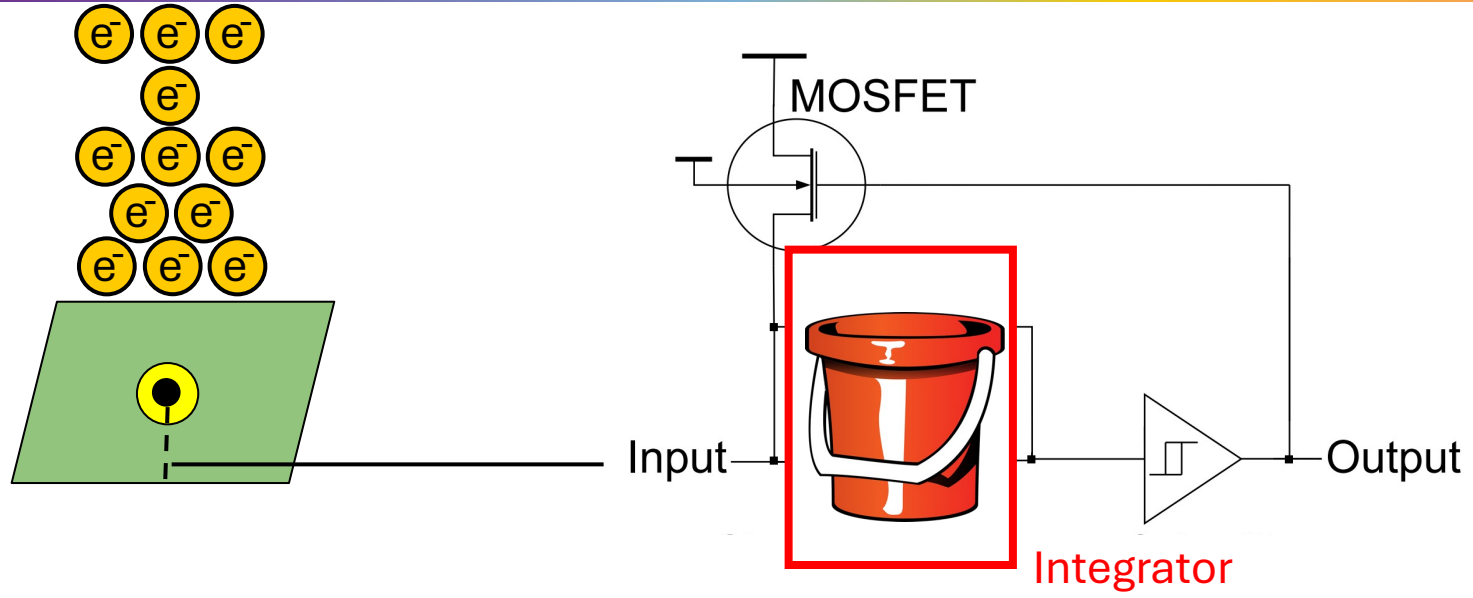


Toy Example



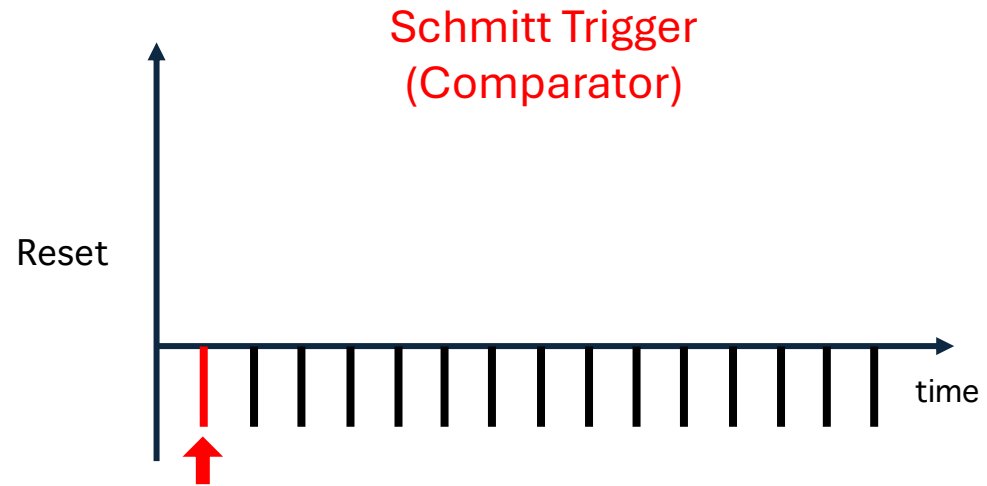
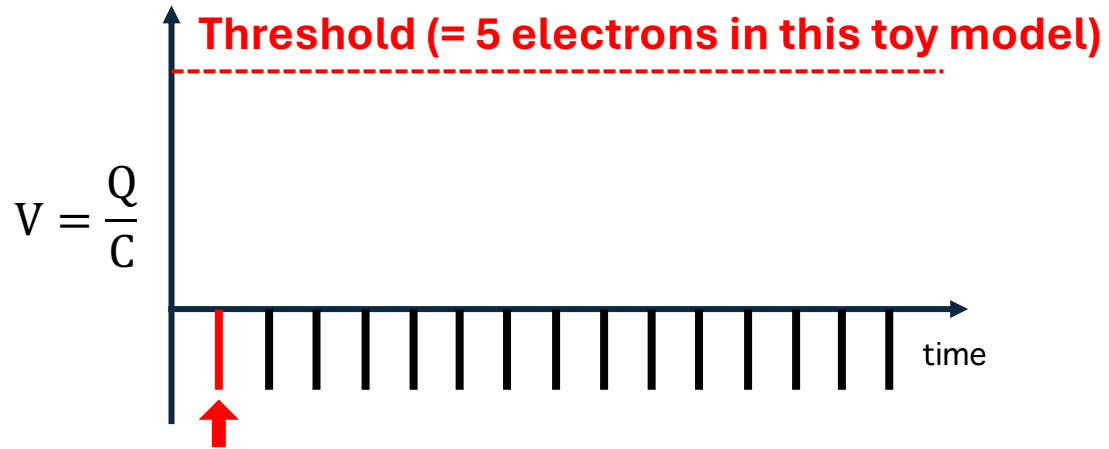
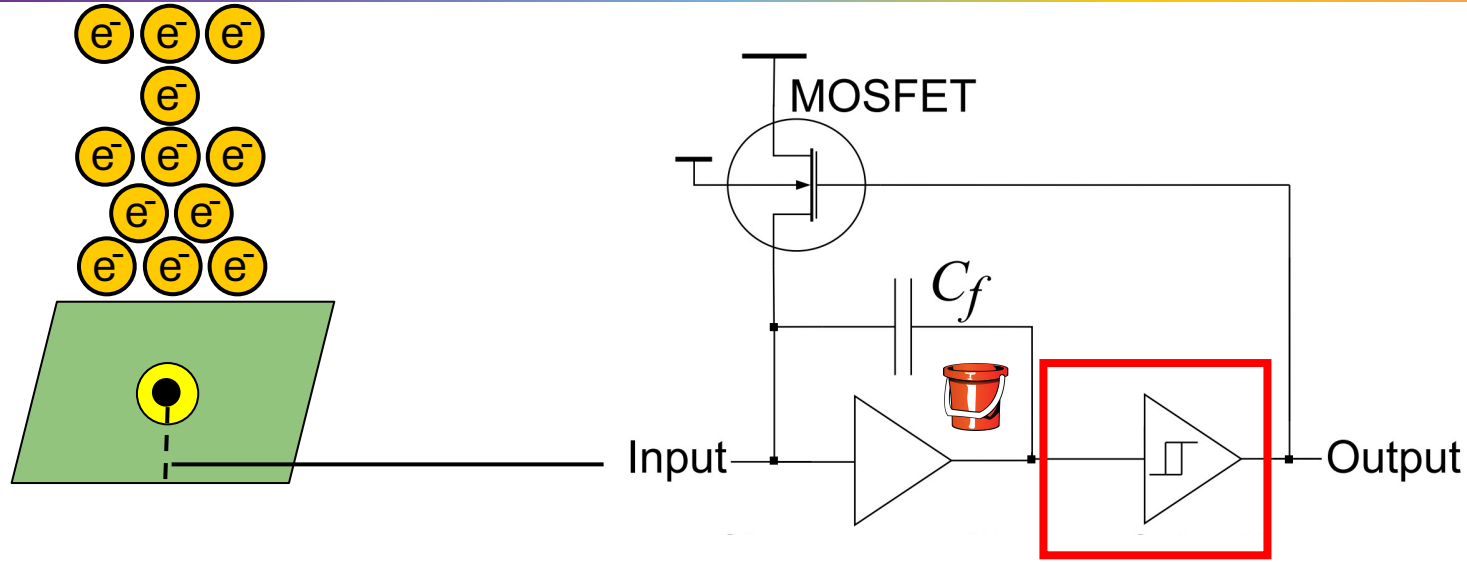


Toy Example

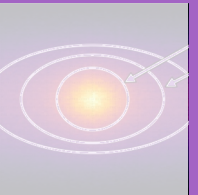
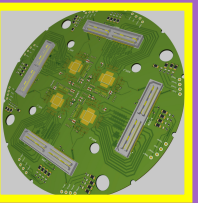




Toy Example

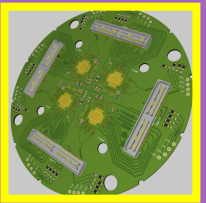
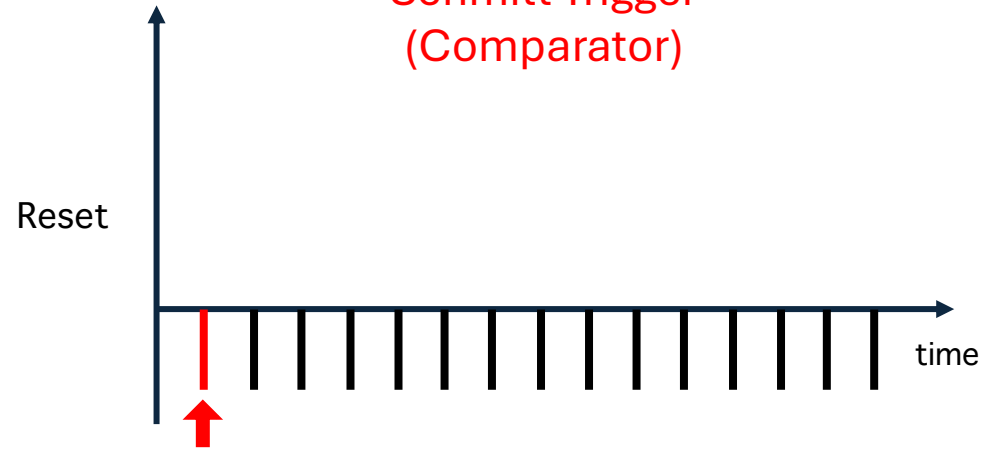
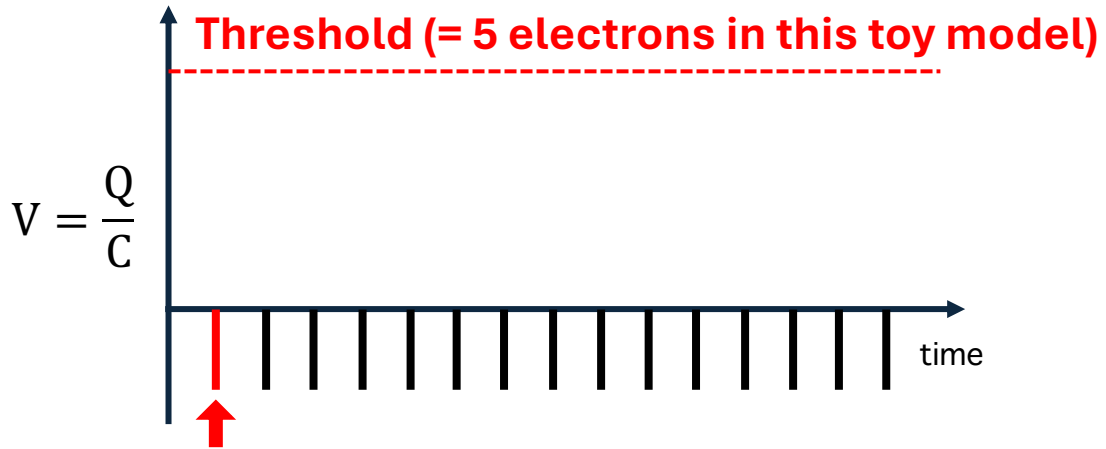
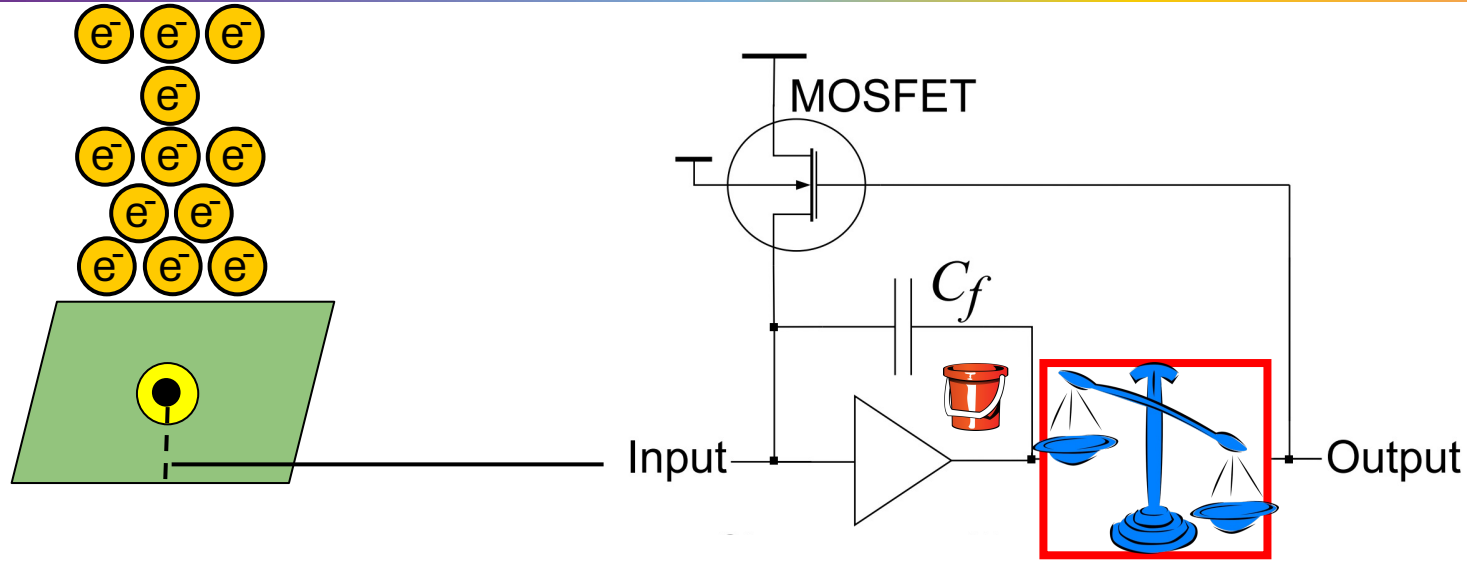


Schmitt Trigger
(Comparator)



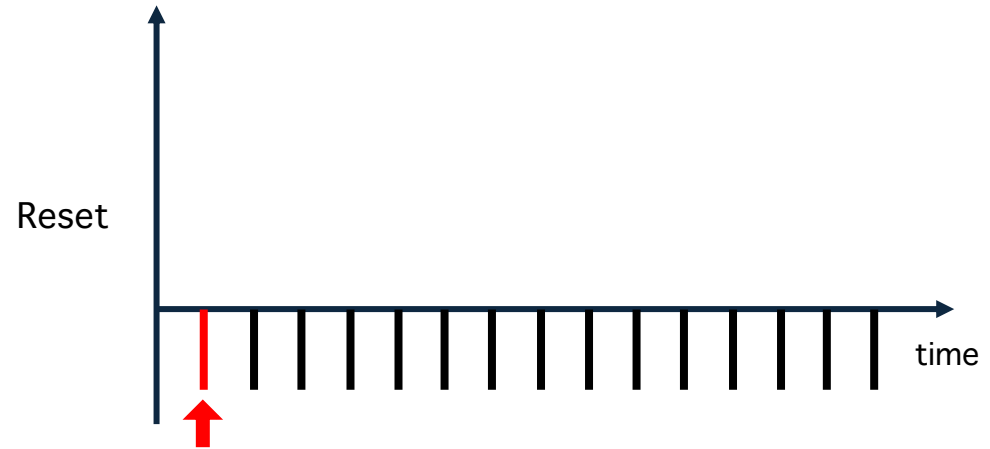
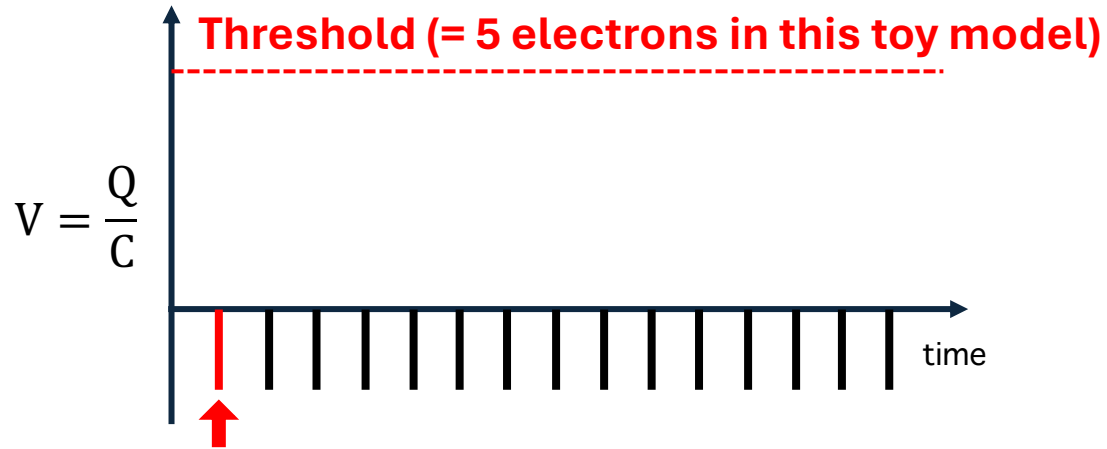
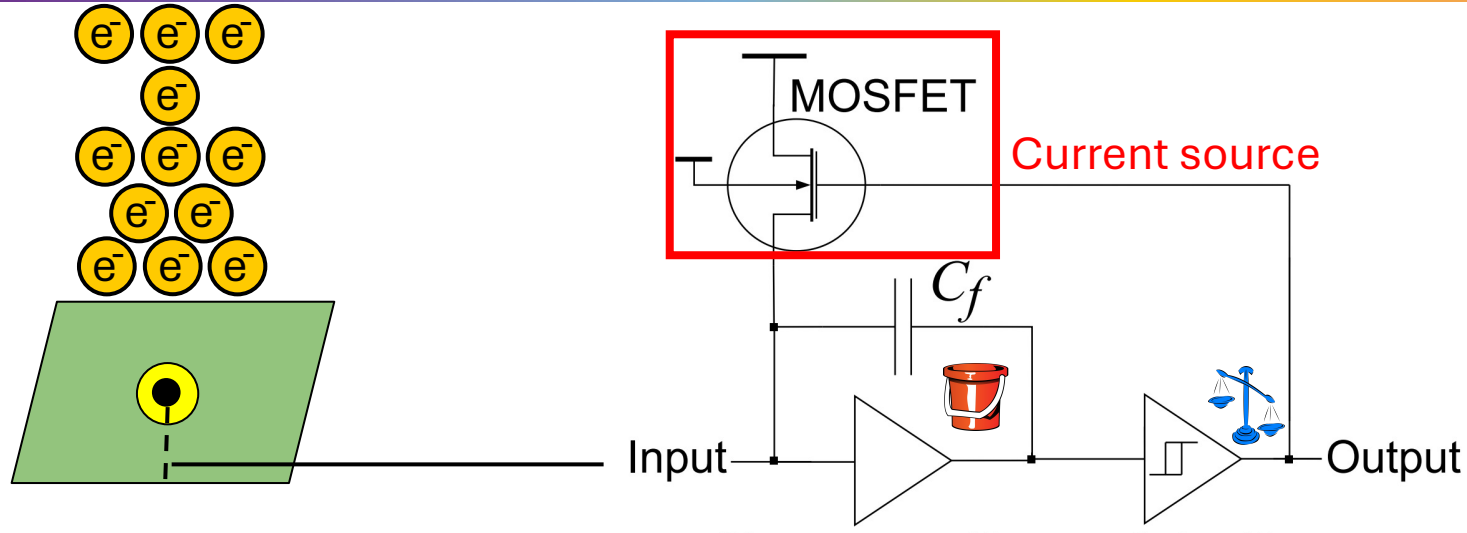


Toy Example

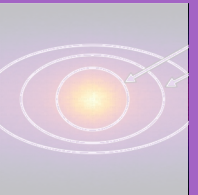
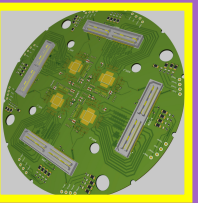




Toy Example

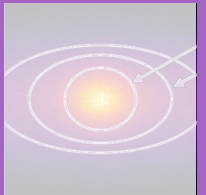
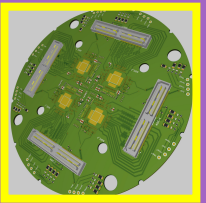
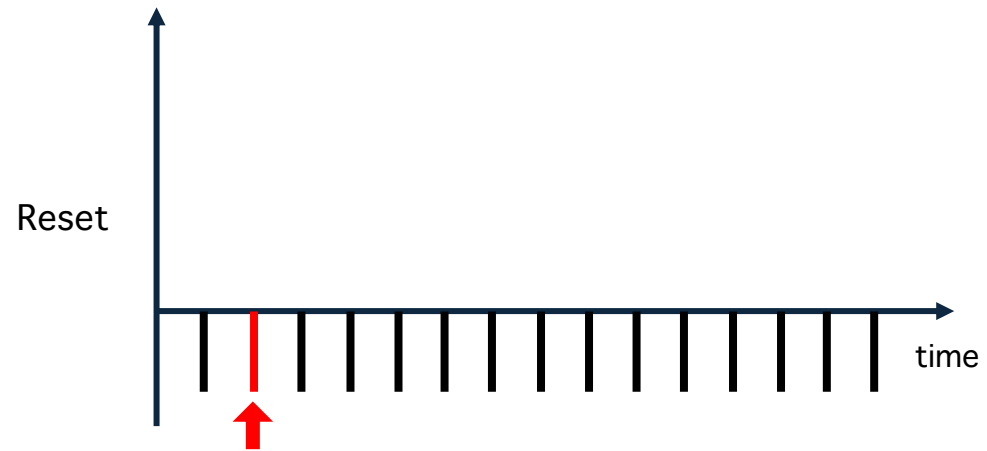
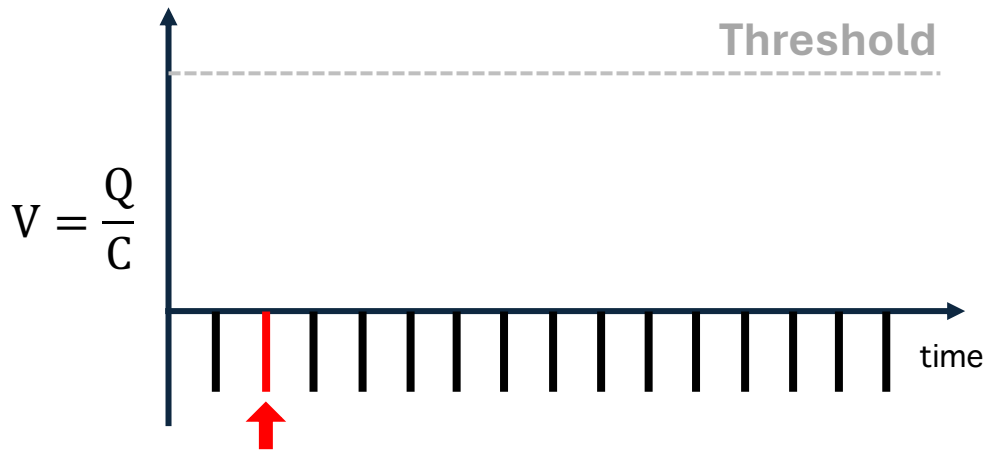
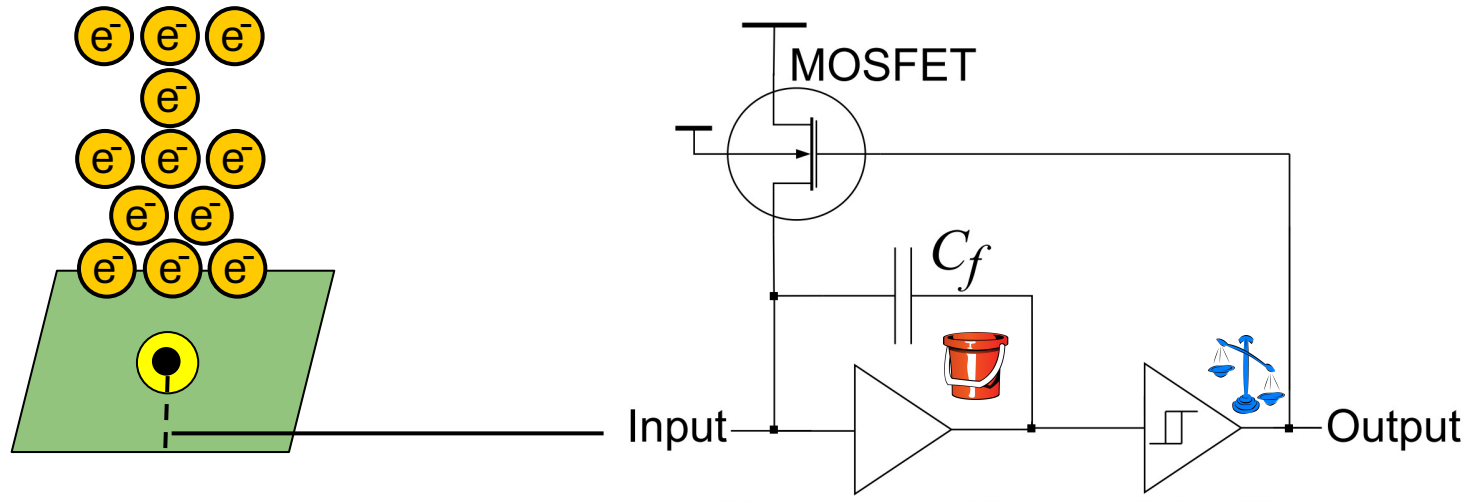


※ the reset happens for 5 electrons



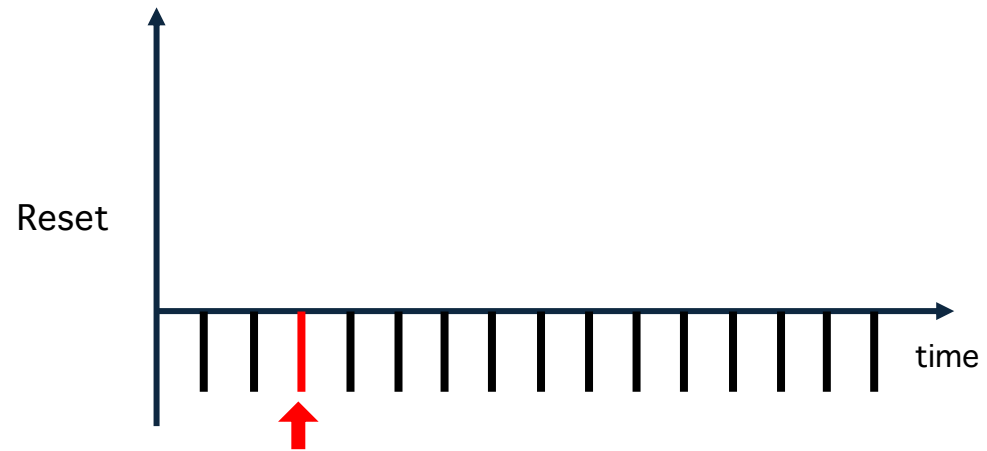
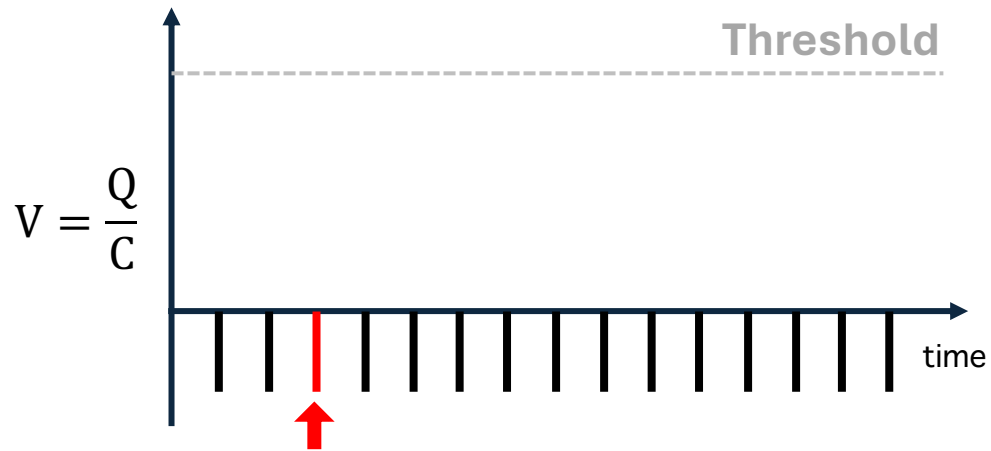
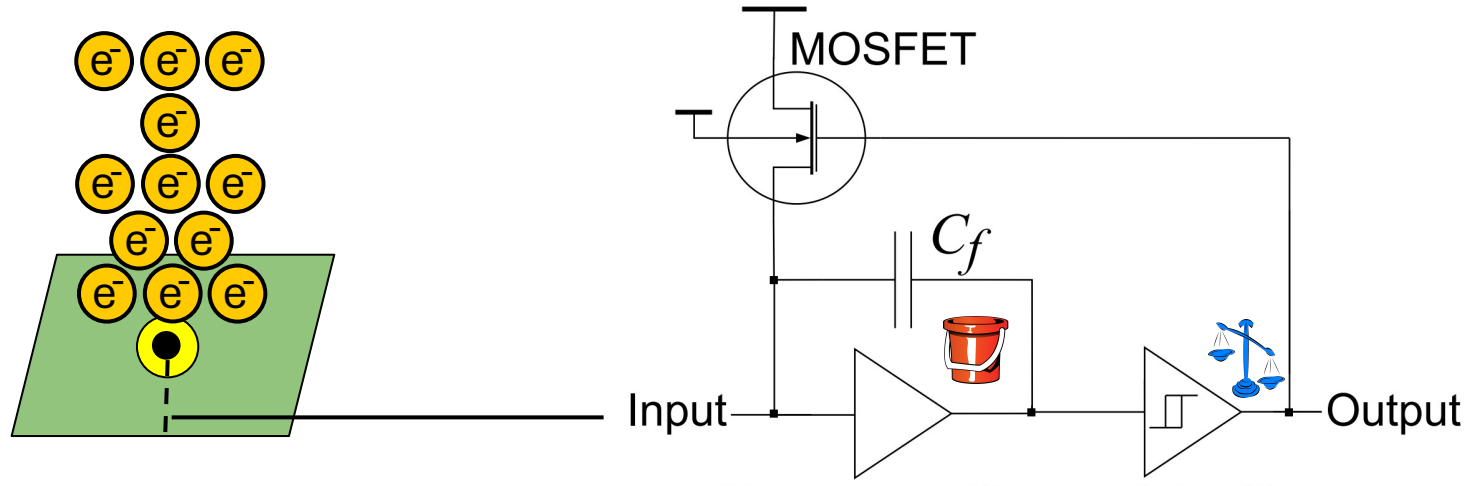


Toy Example

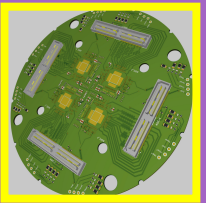




Toy Example

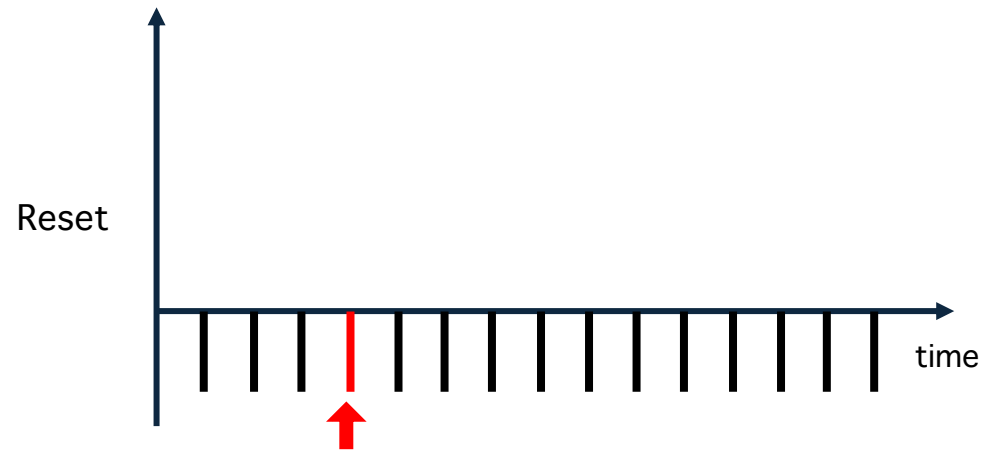
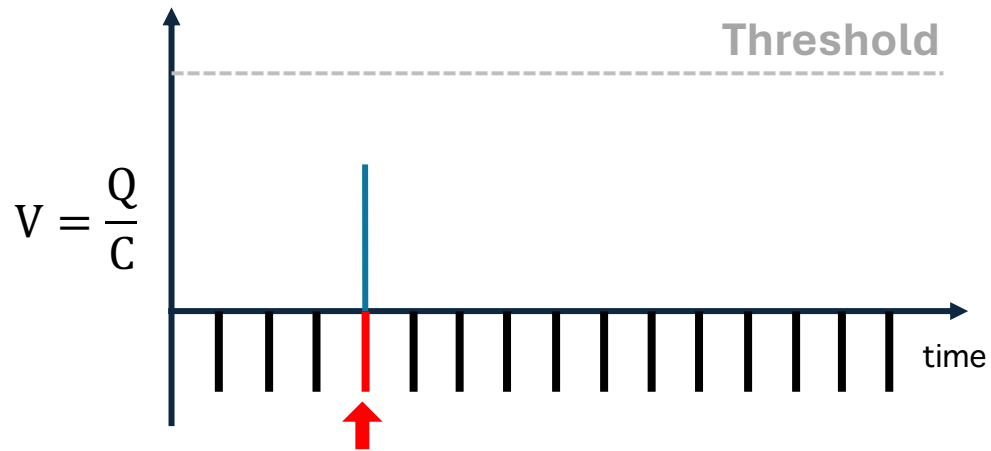
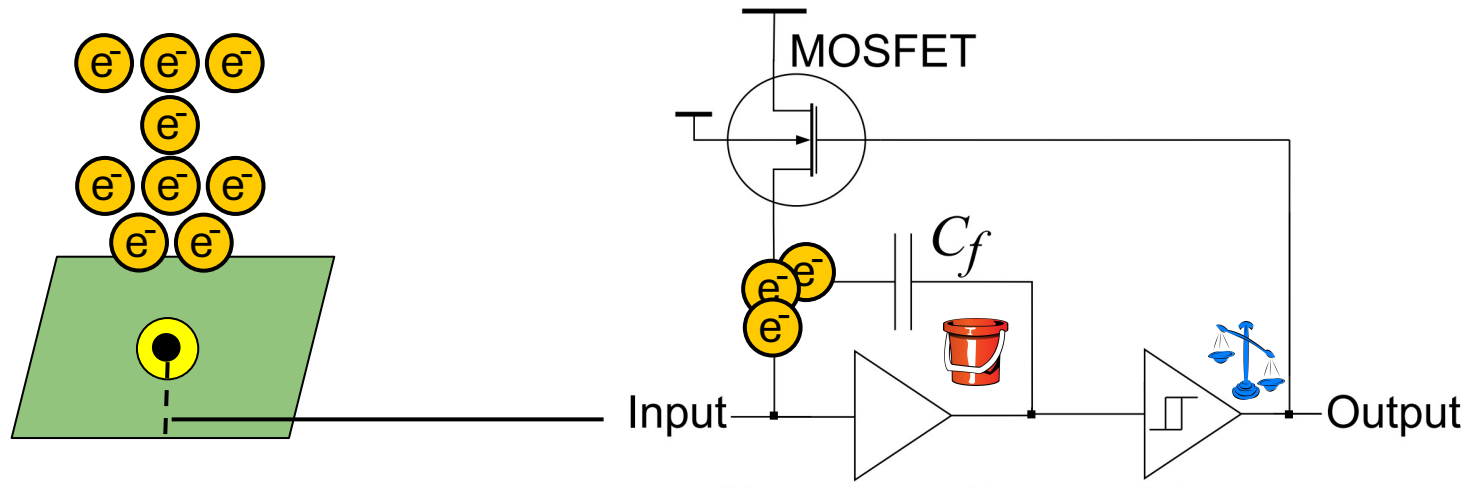


※ the reset happens for 5 electrons

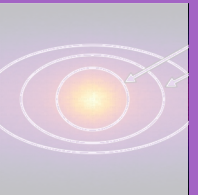
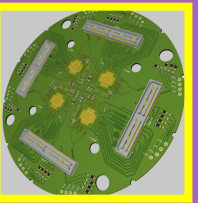




Toy Example

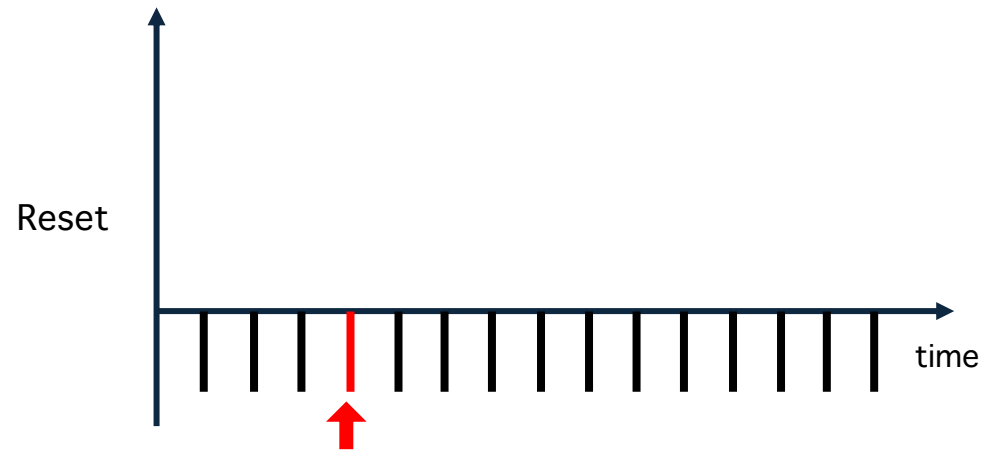
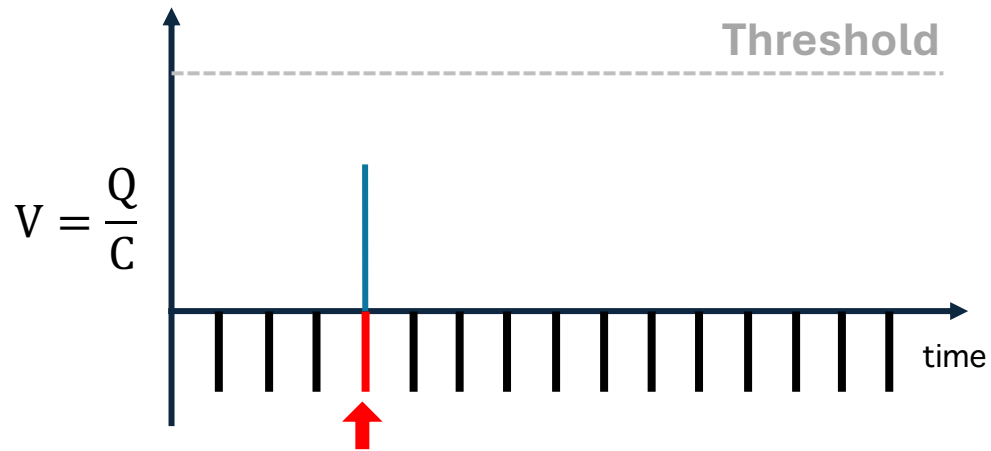
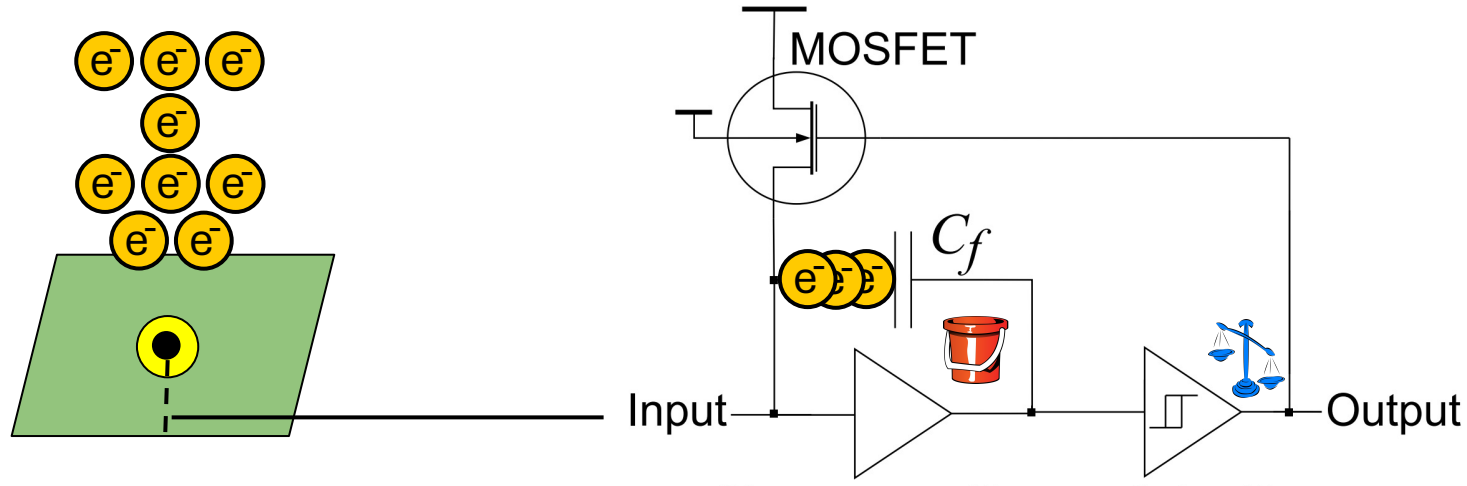


※ the reset happens for 5 electrons

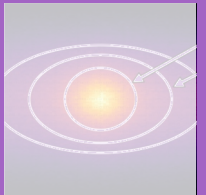
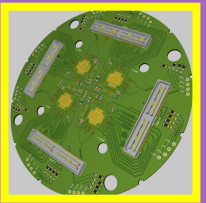




Toy Example

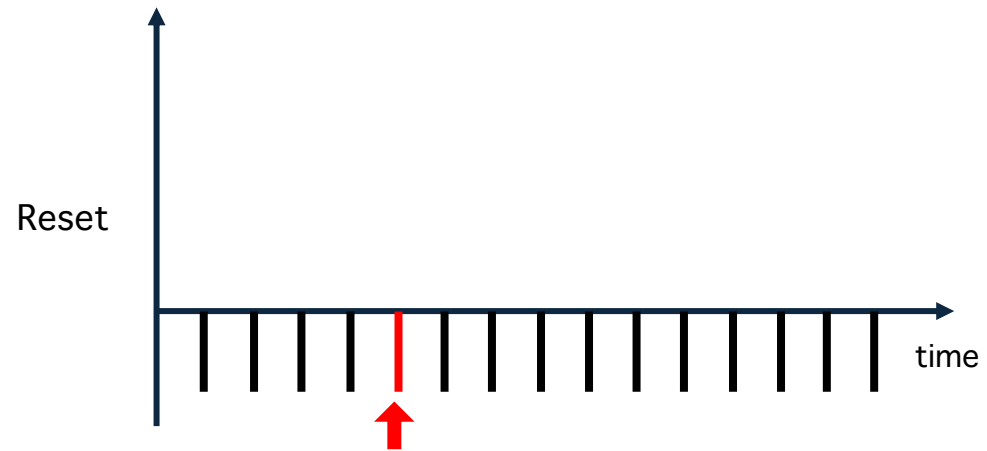
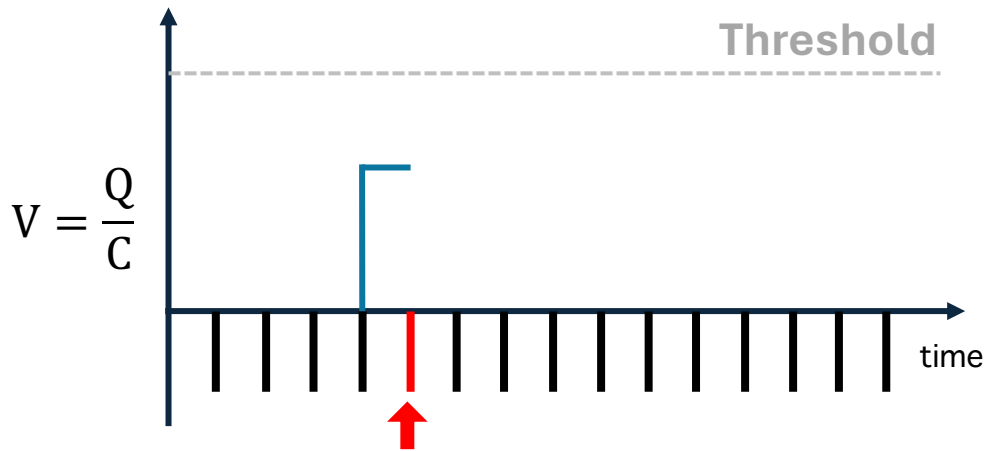
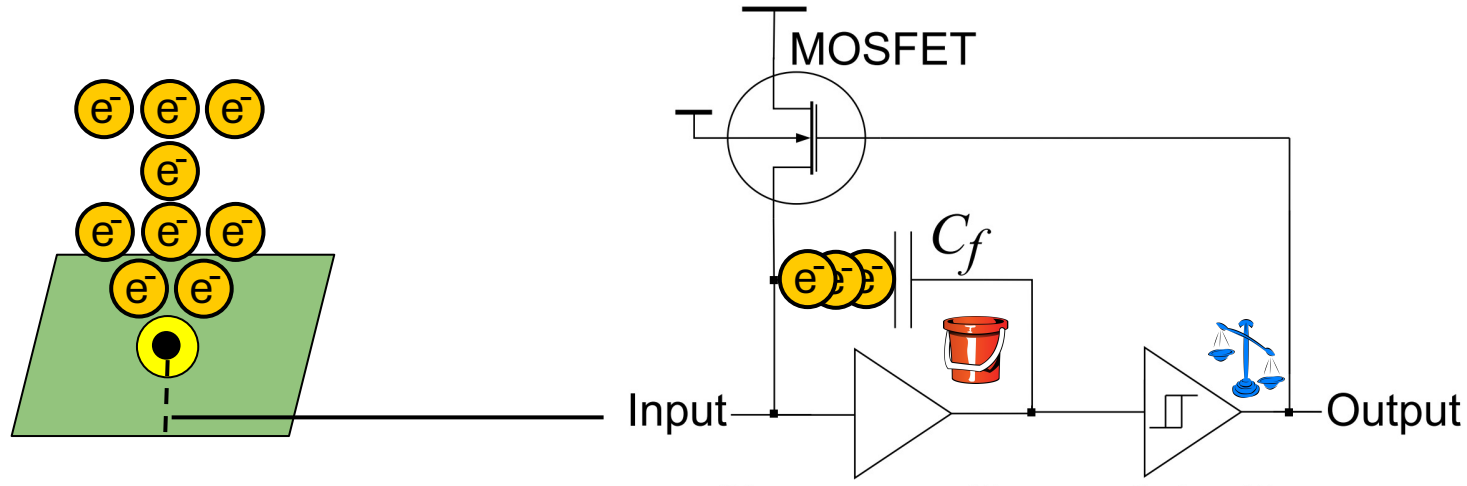


※ the reset happens for 5 electrons

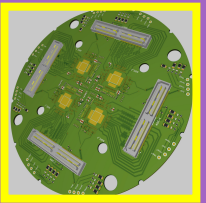




Toy Example

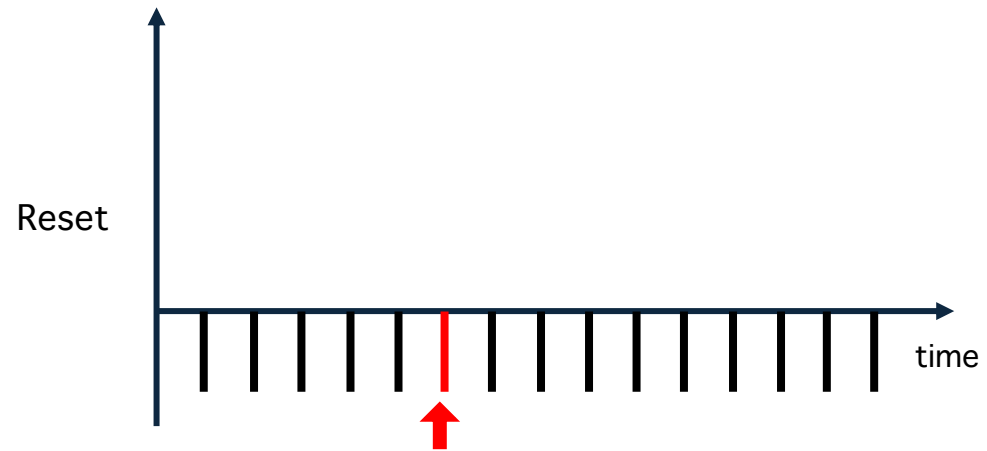
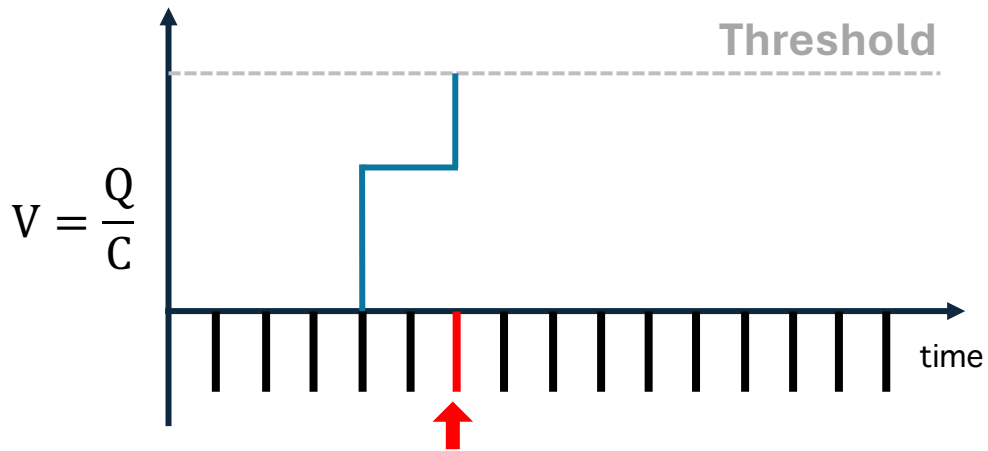
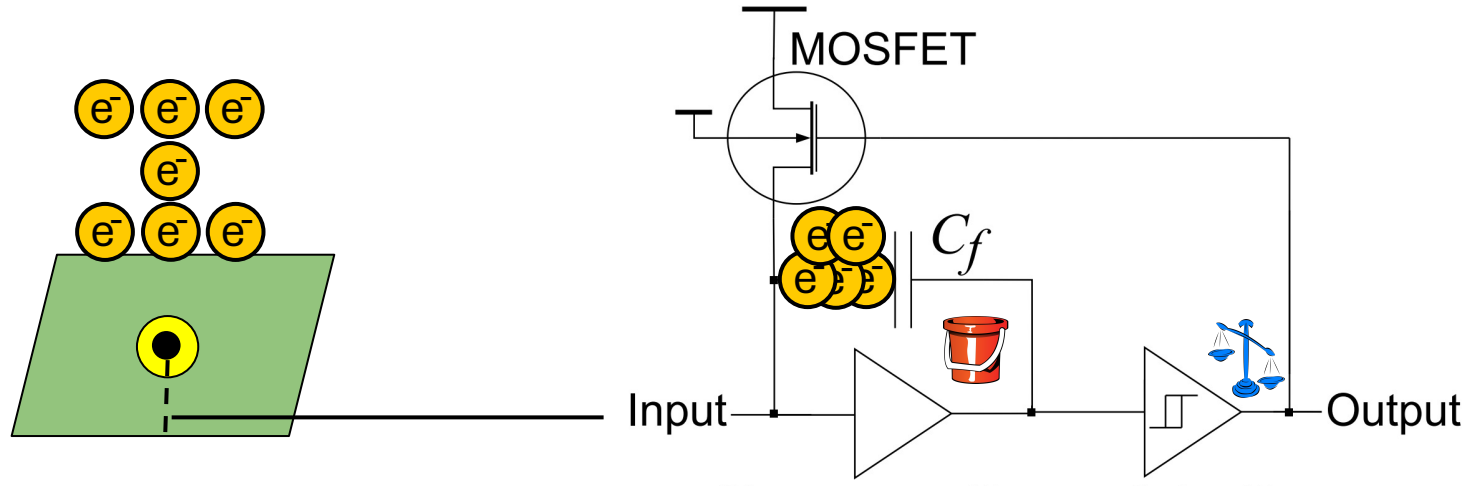


※ the reset happens for 5 electrons

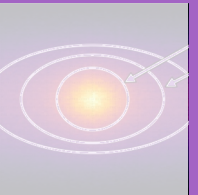
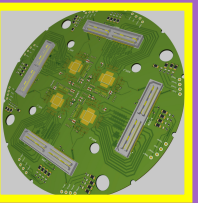




Toy Example

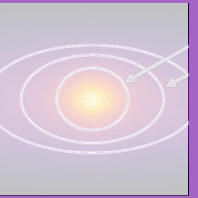
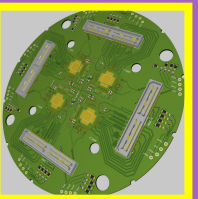
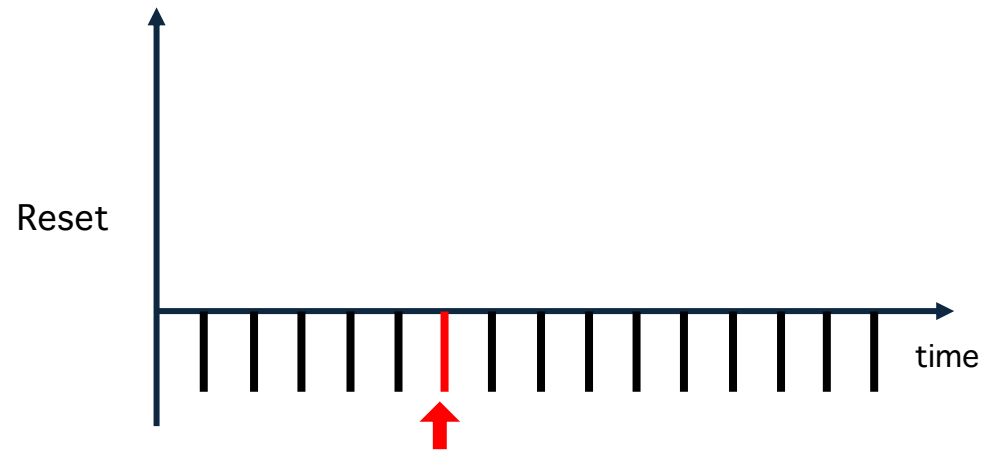
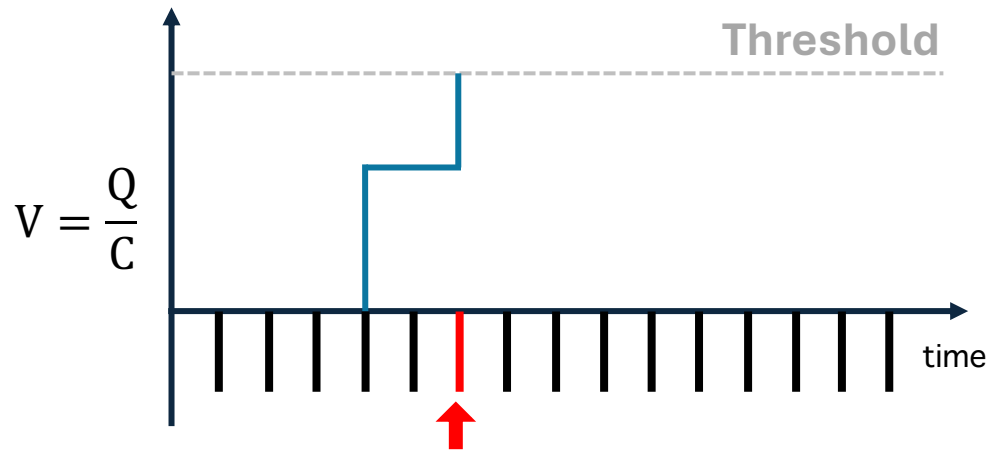
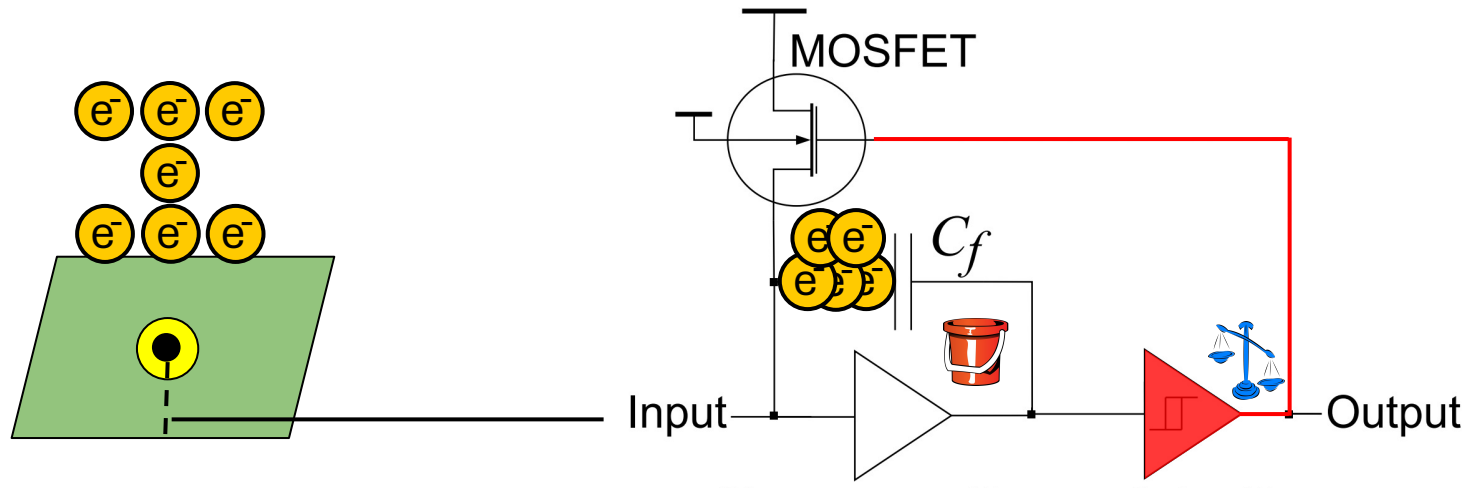


※ the reset happens for 5 electrons



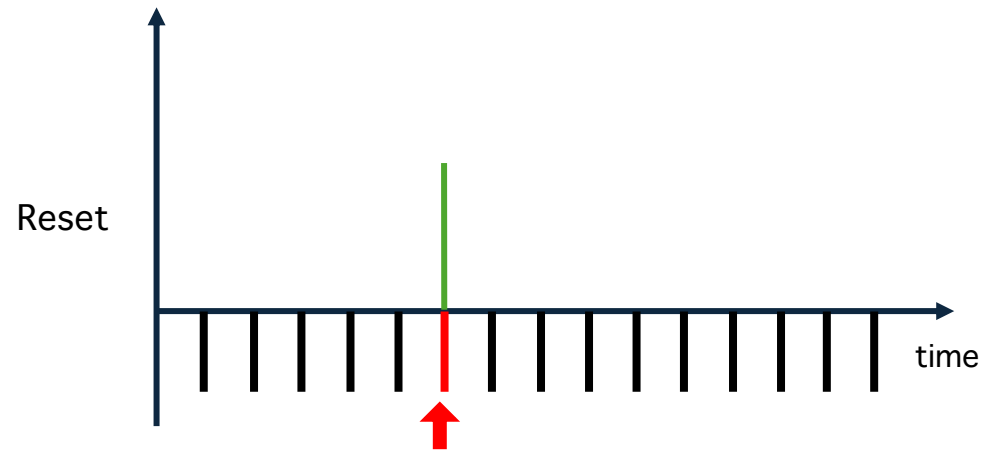
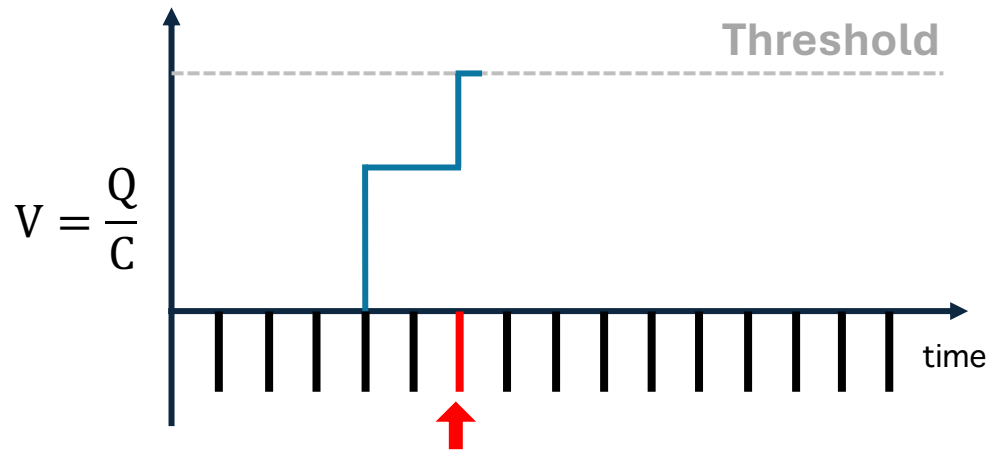
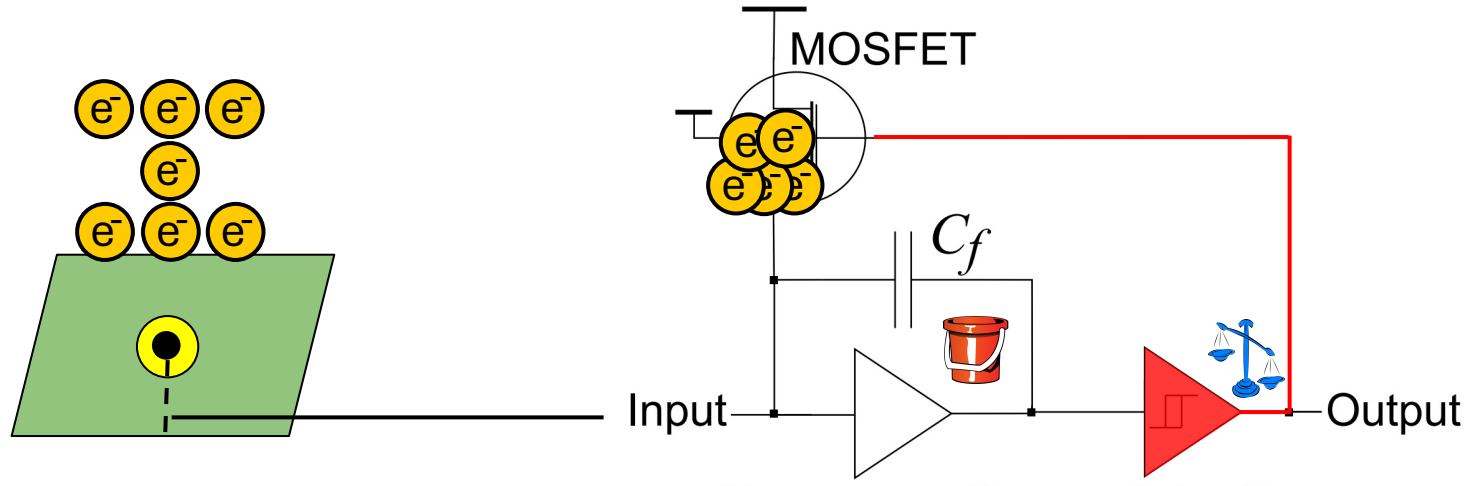


Toy Example

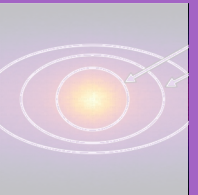
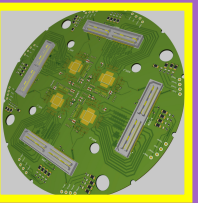




Toy Example

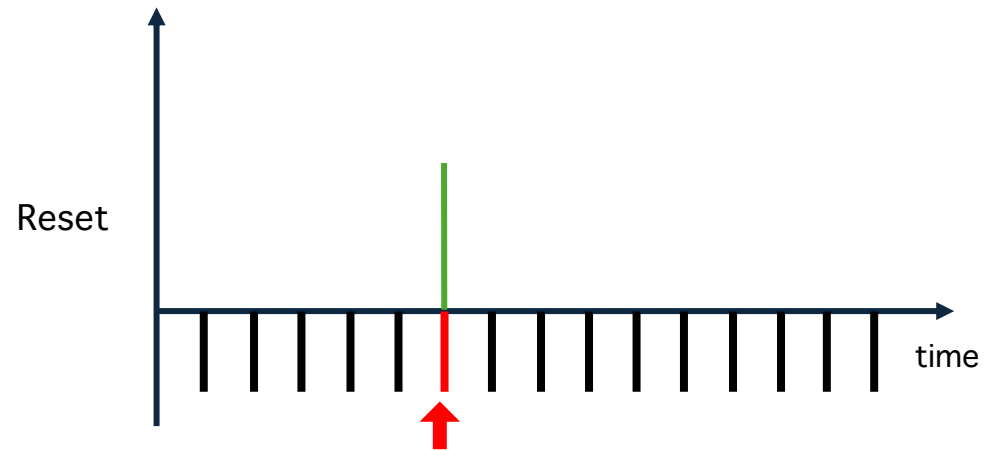
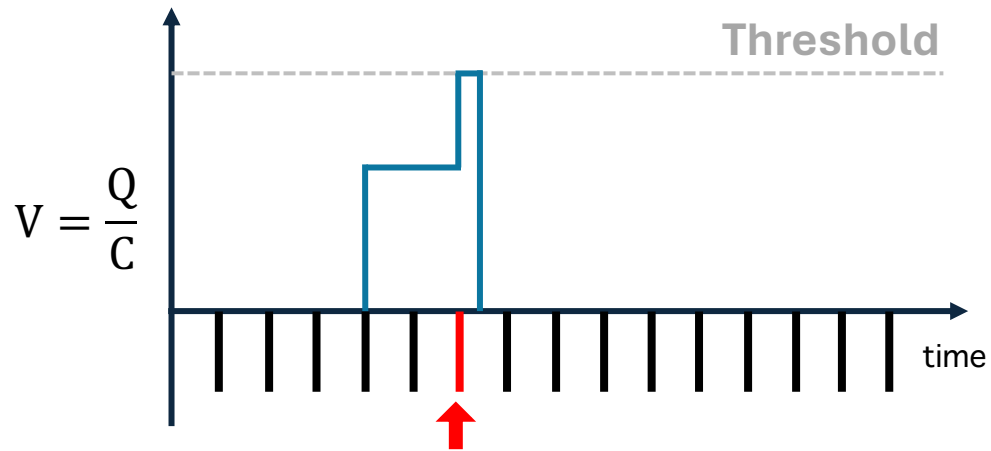
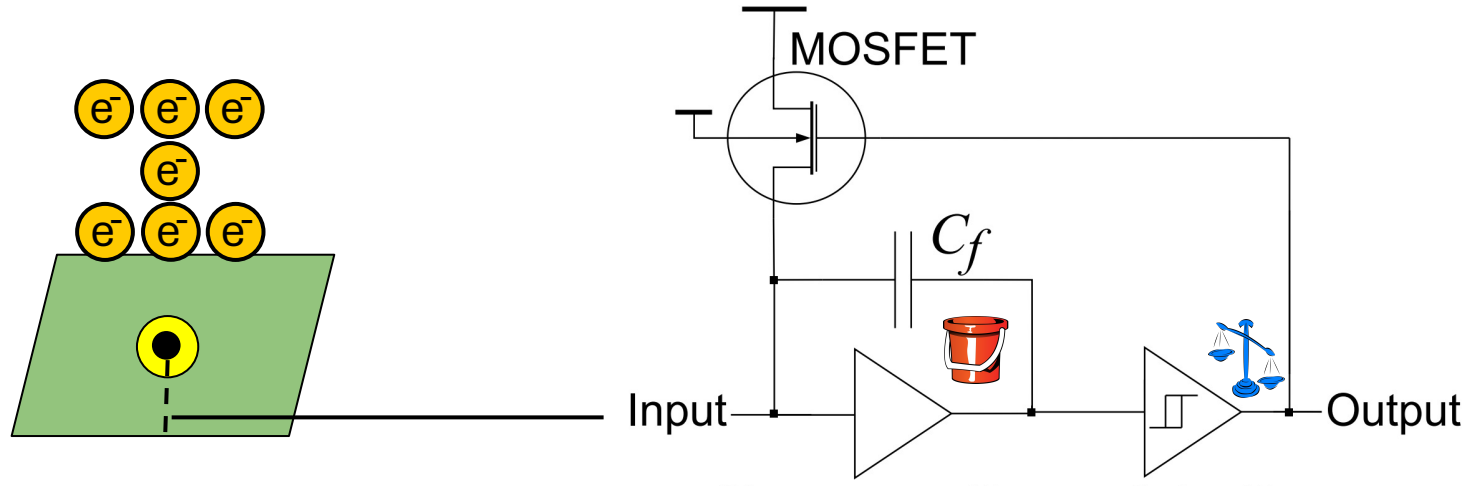


※ the reset happens for 5 electrons

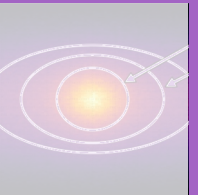
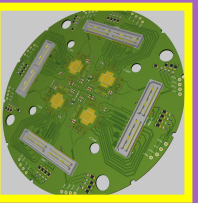




Toy Example

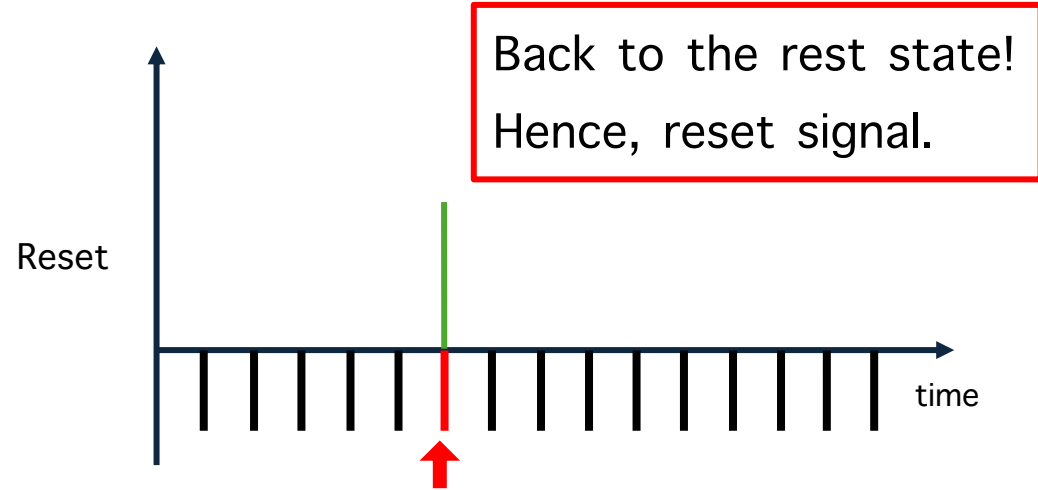
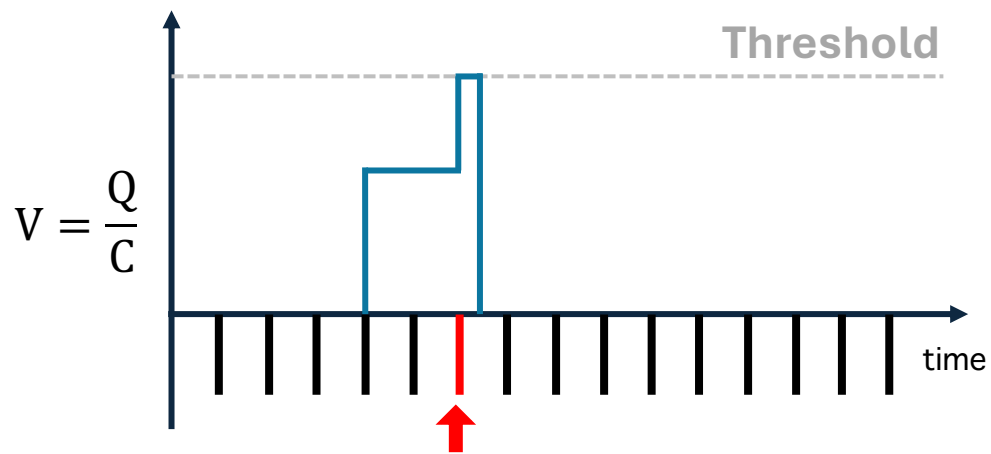
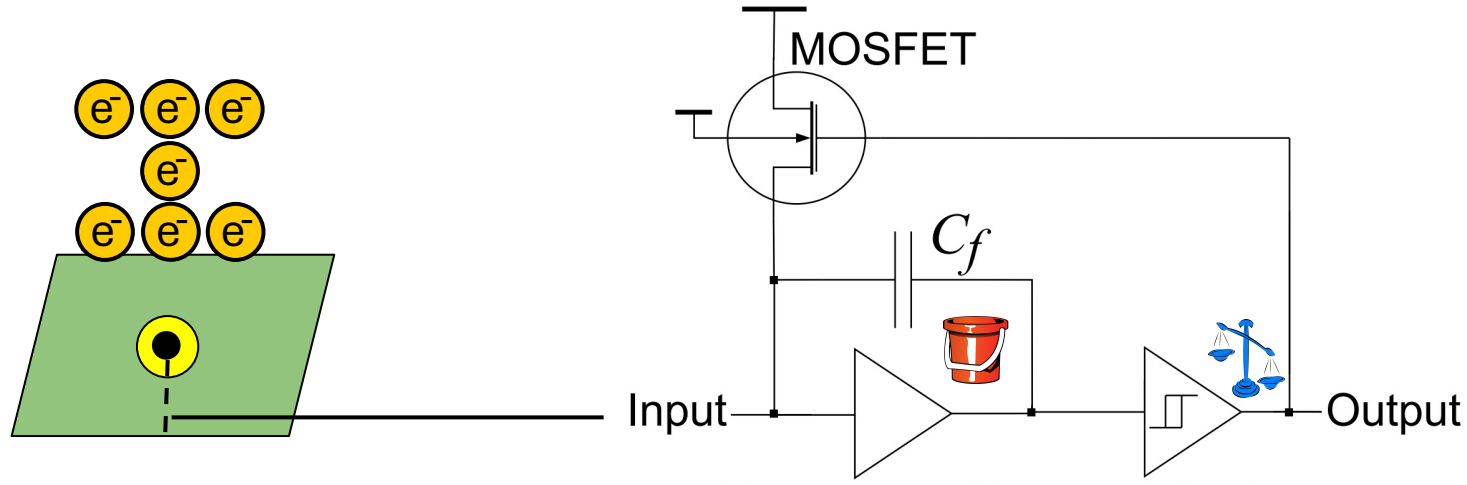


※ the reset happens for 5 electrons

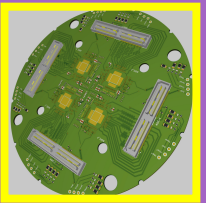




Toy Example

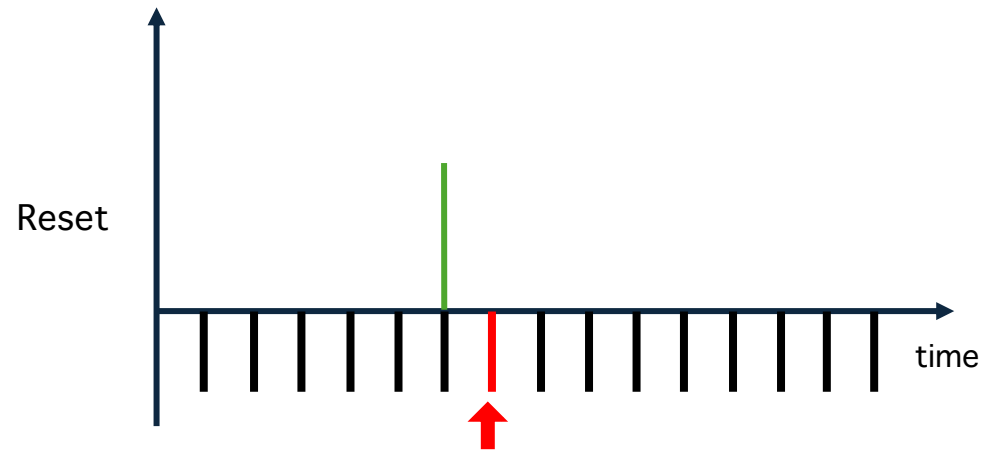
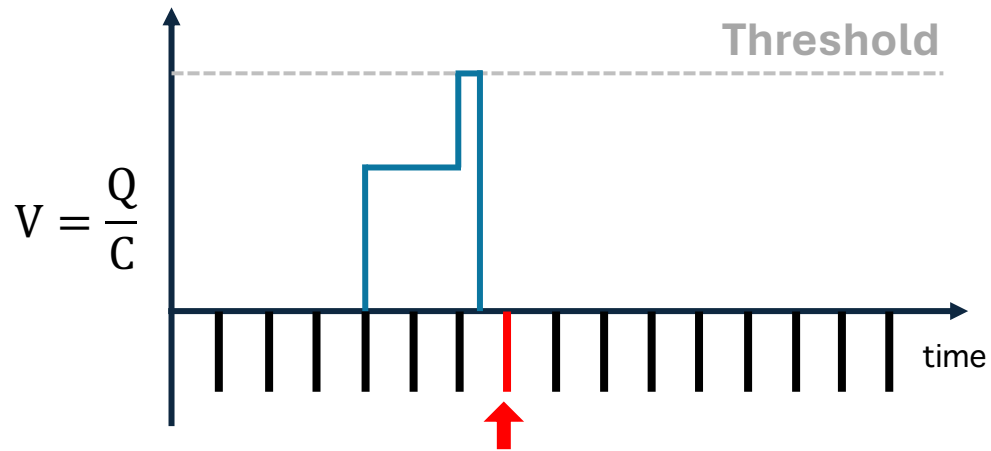
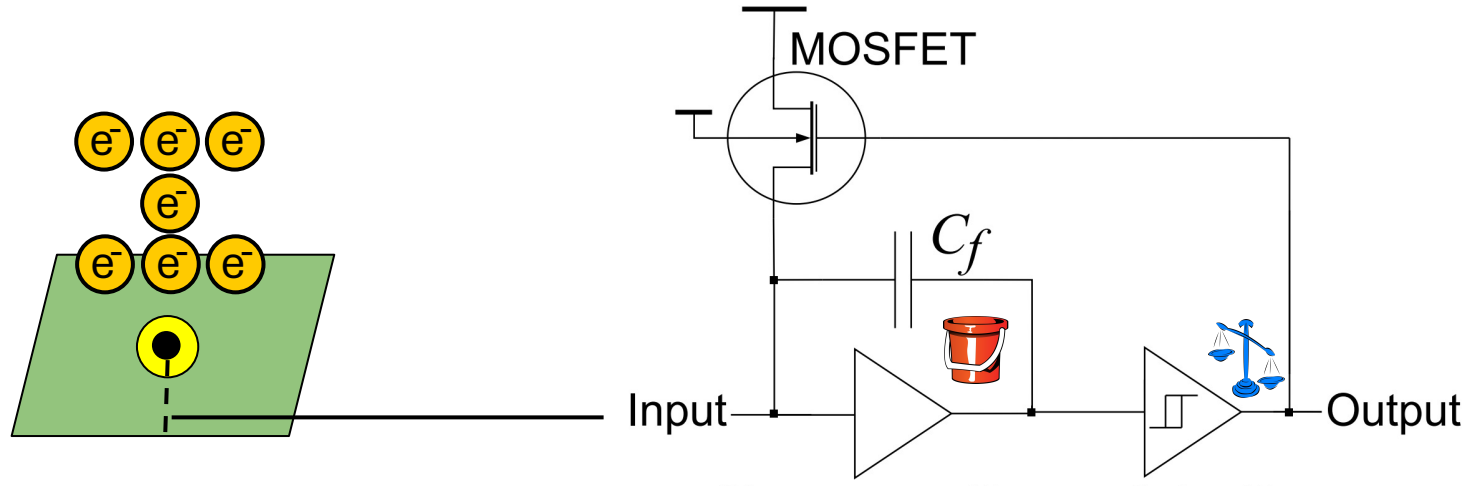


※ the reset happens for 5 electrons

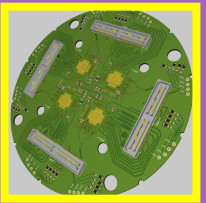




Toy Example

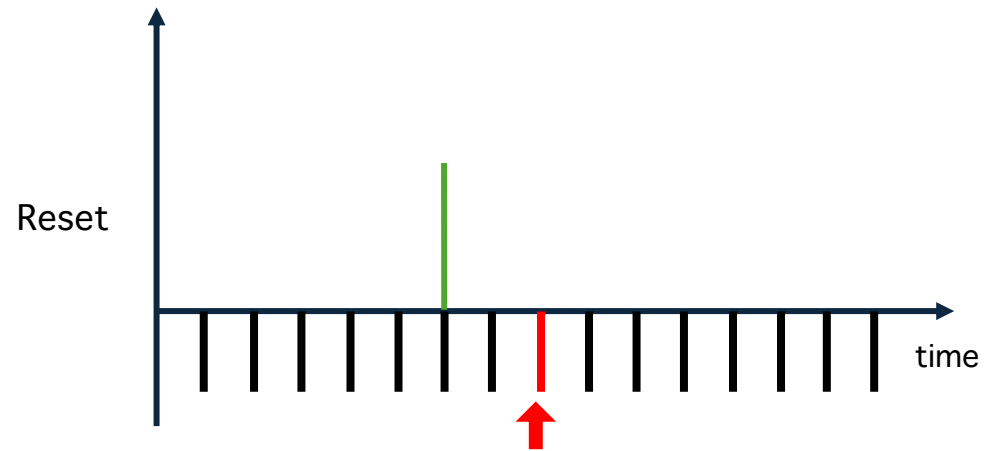
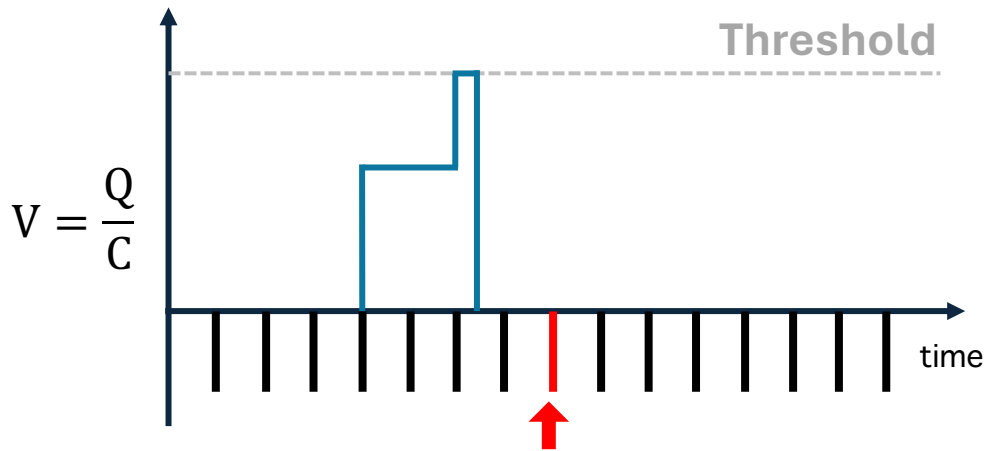
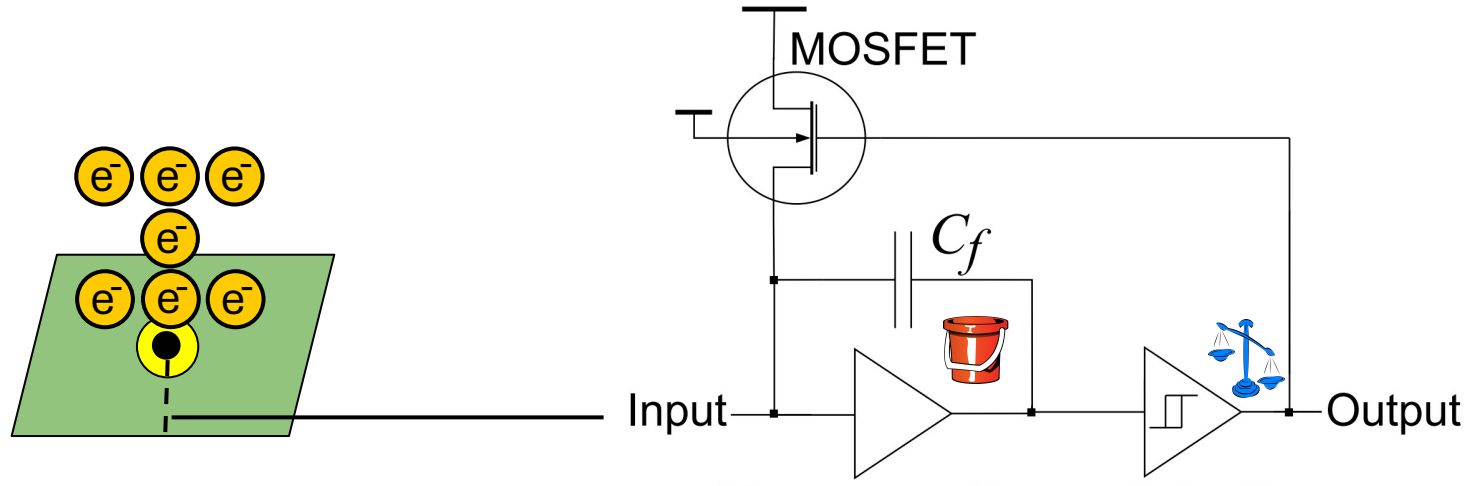


※ the reset happens for 5 electrons

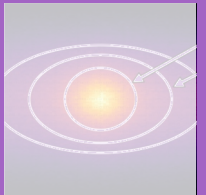
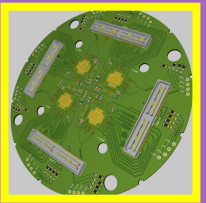




Toy Example

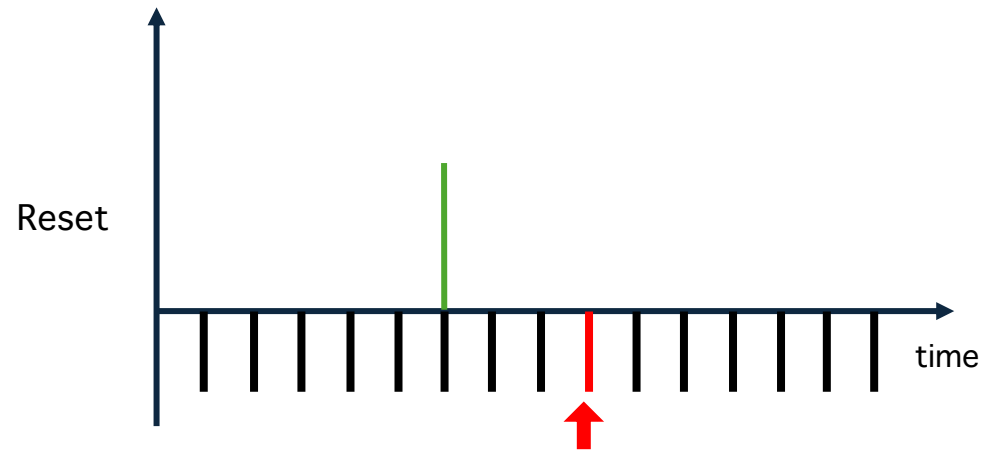
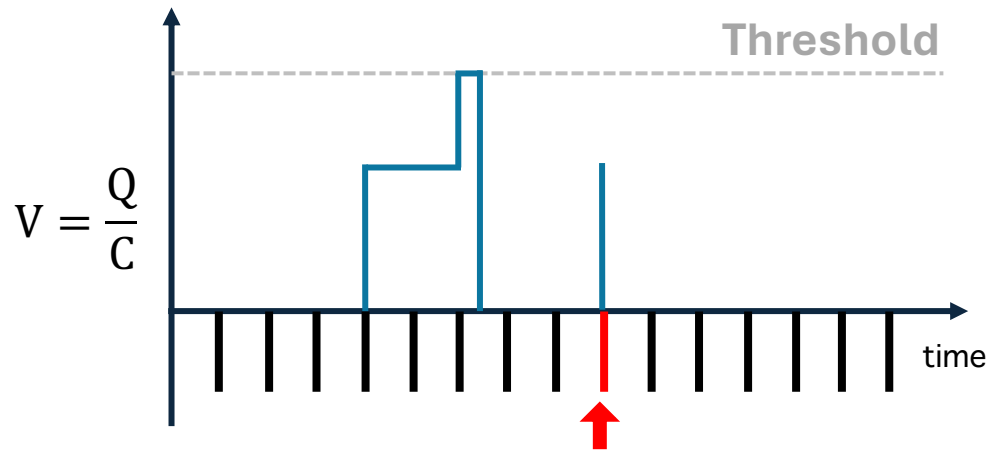
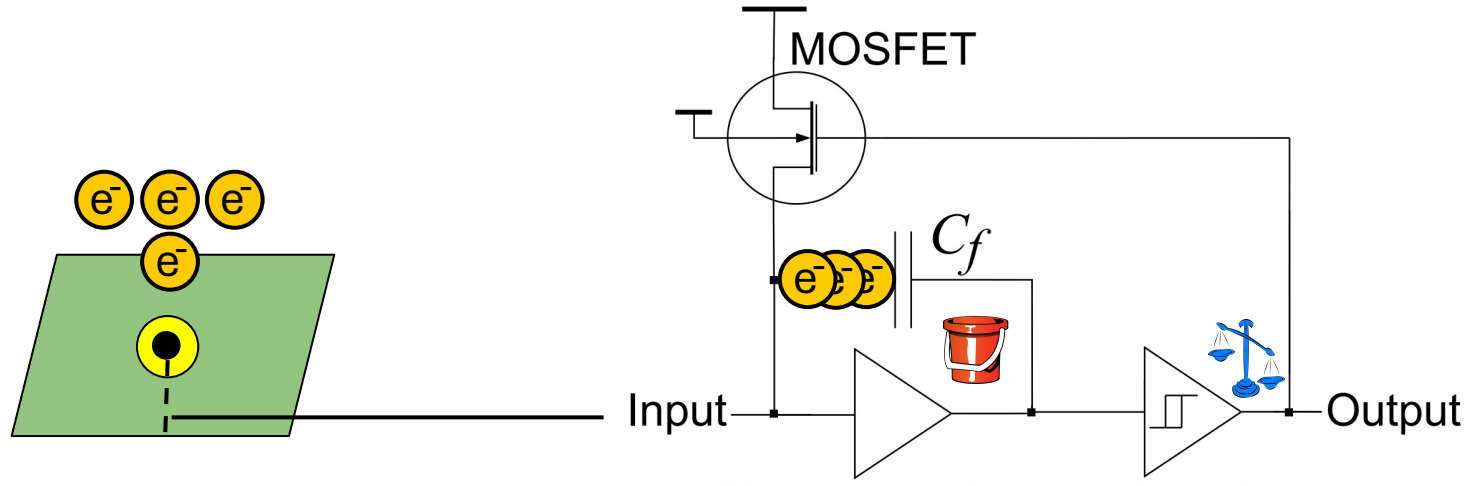


※ the reset happens for 5 electrons

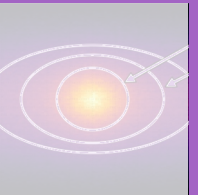
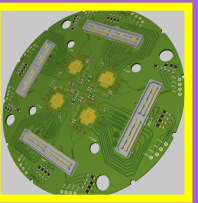




Toy Example

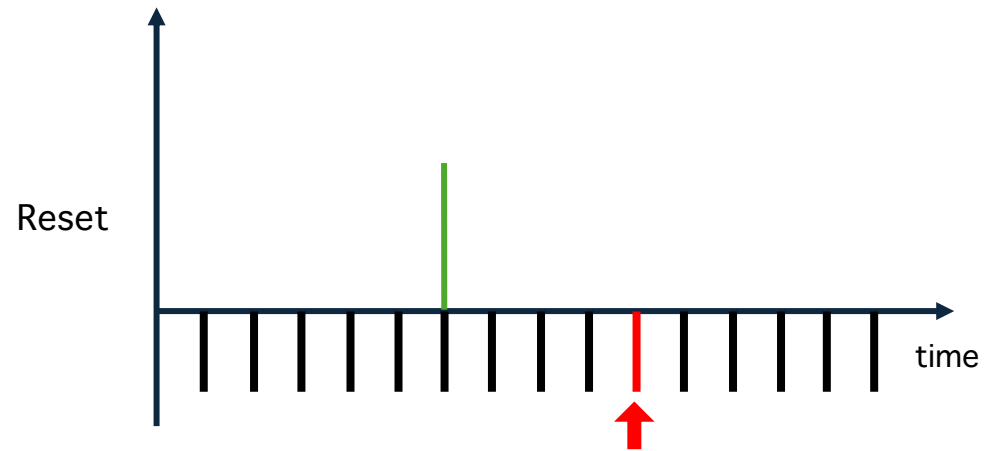
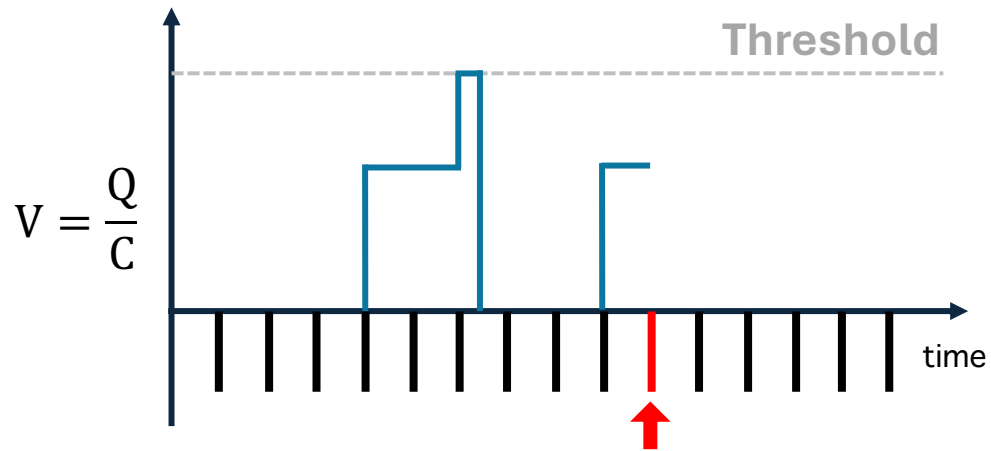
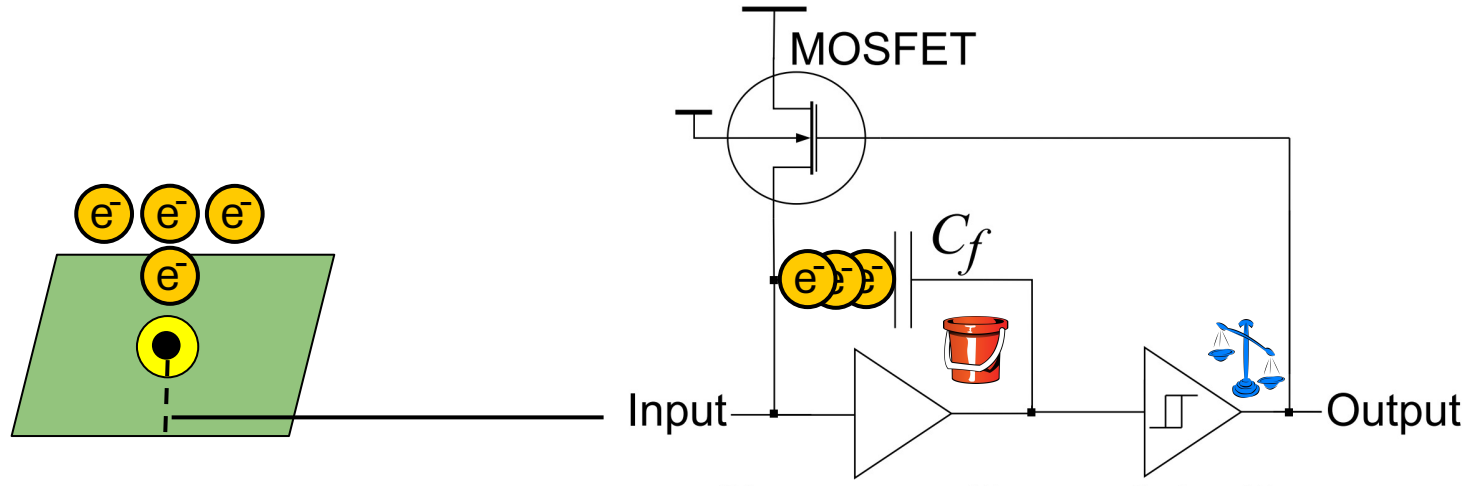


※ the reset happens for 5 electrons

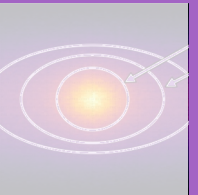
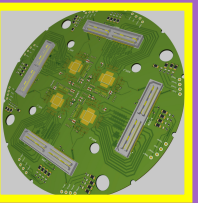




Toy Example

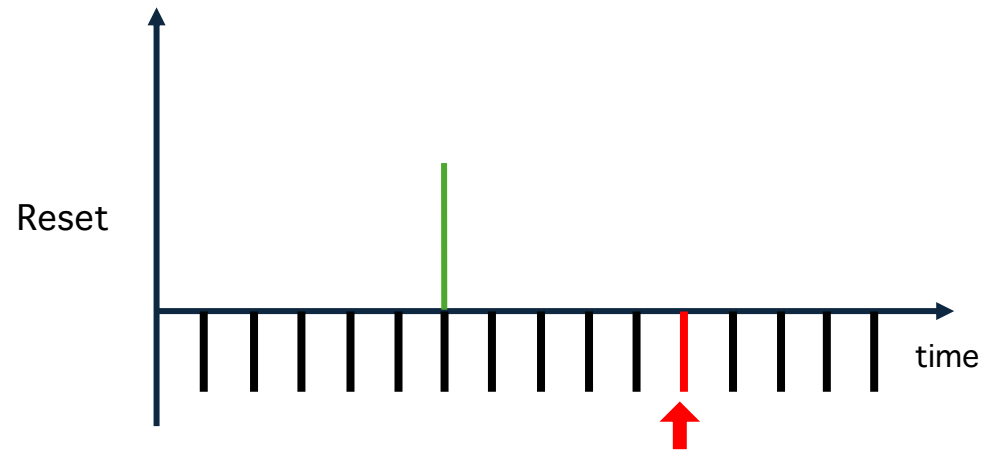
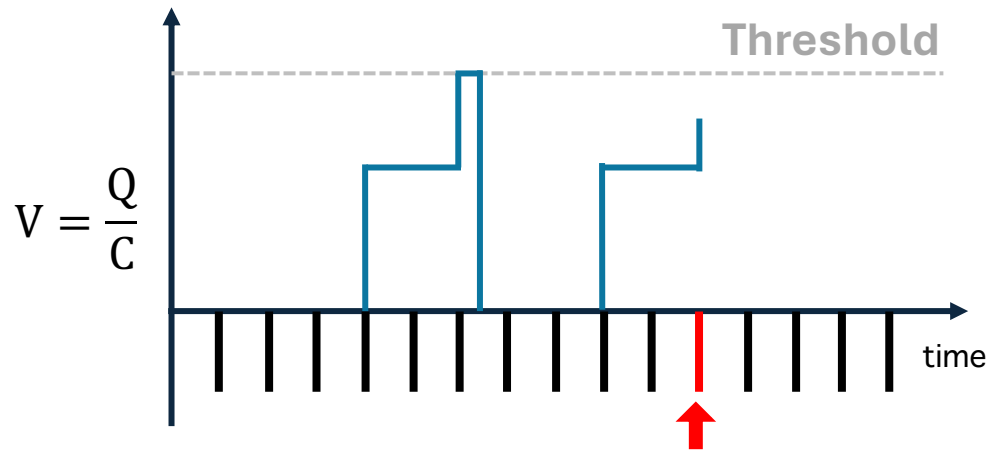
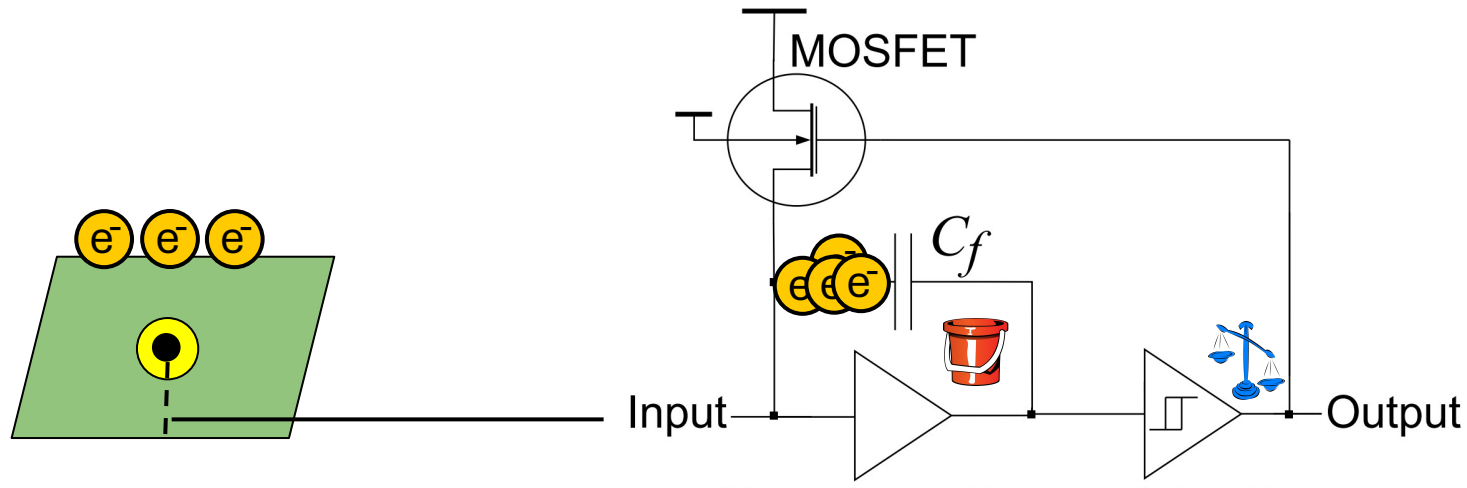


※ the reset happens for 5 electrons

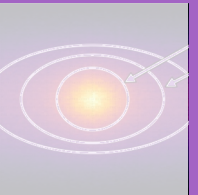
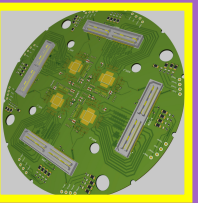




Toy Example

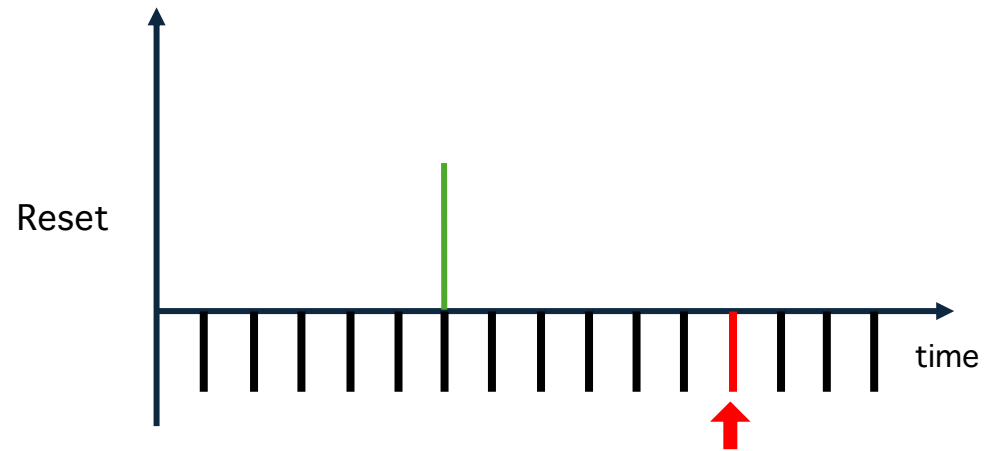
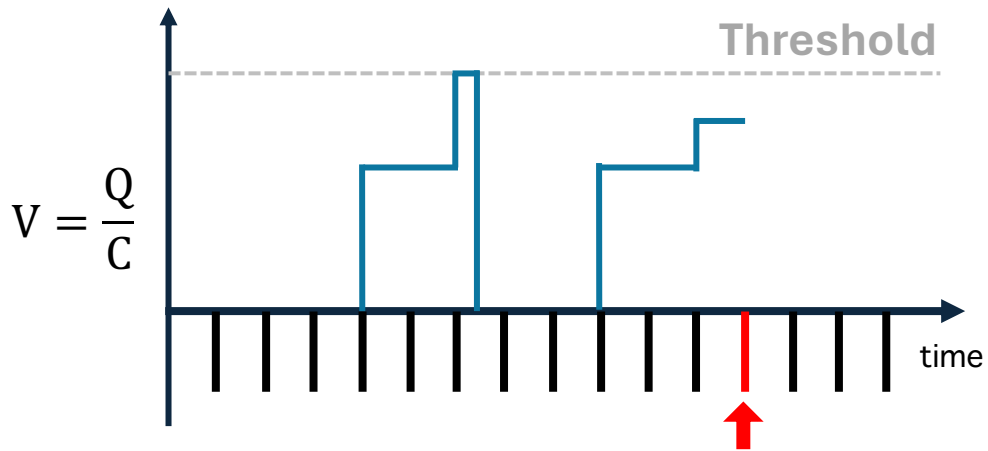
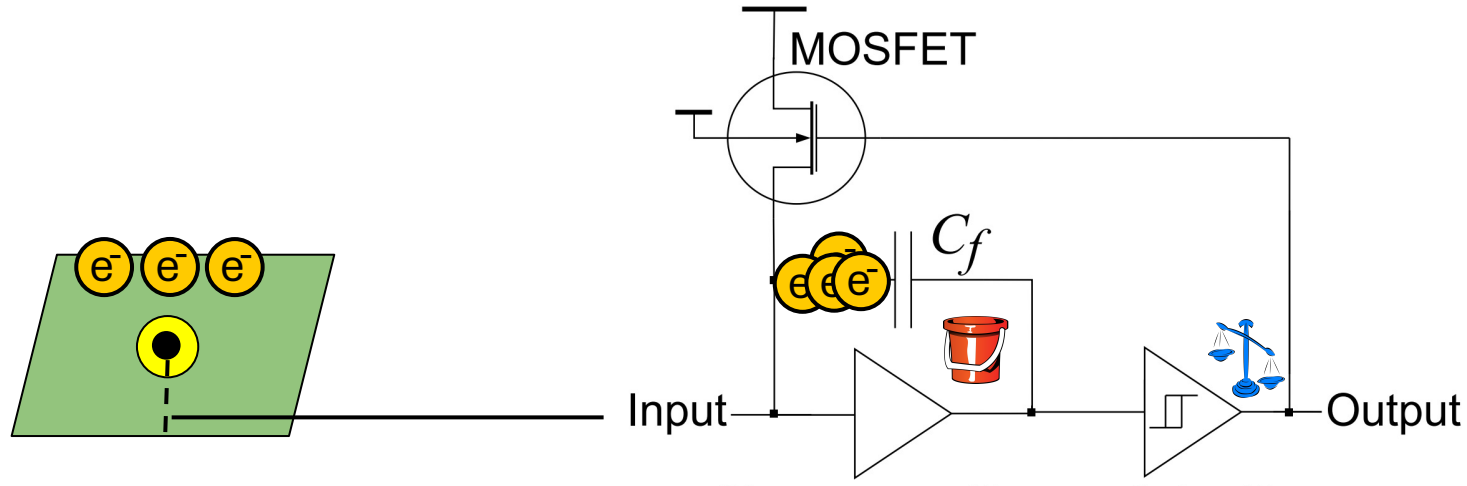


※ the reset happens for 5 electrons

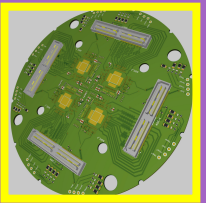




Toy Example

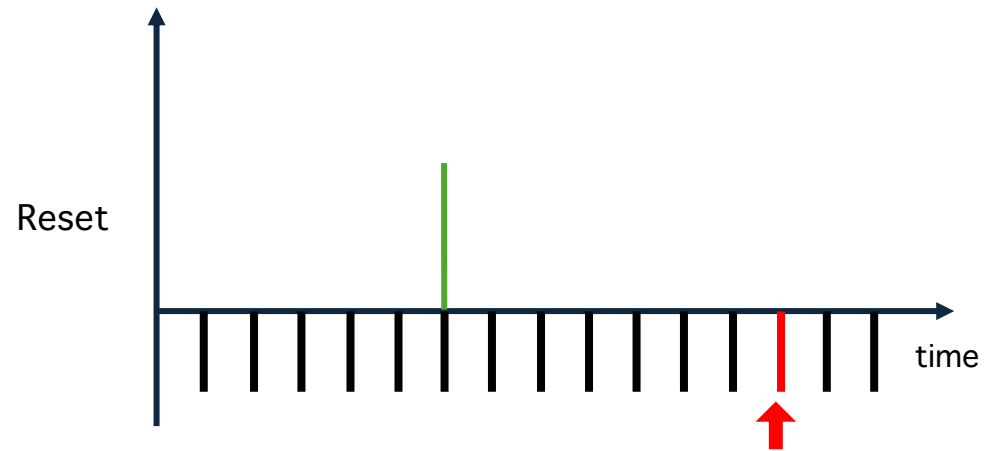
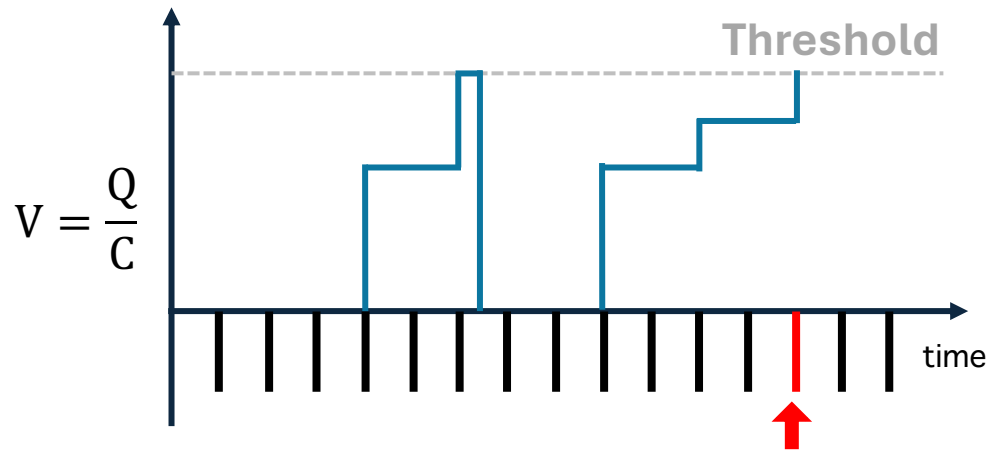
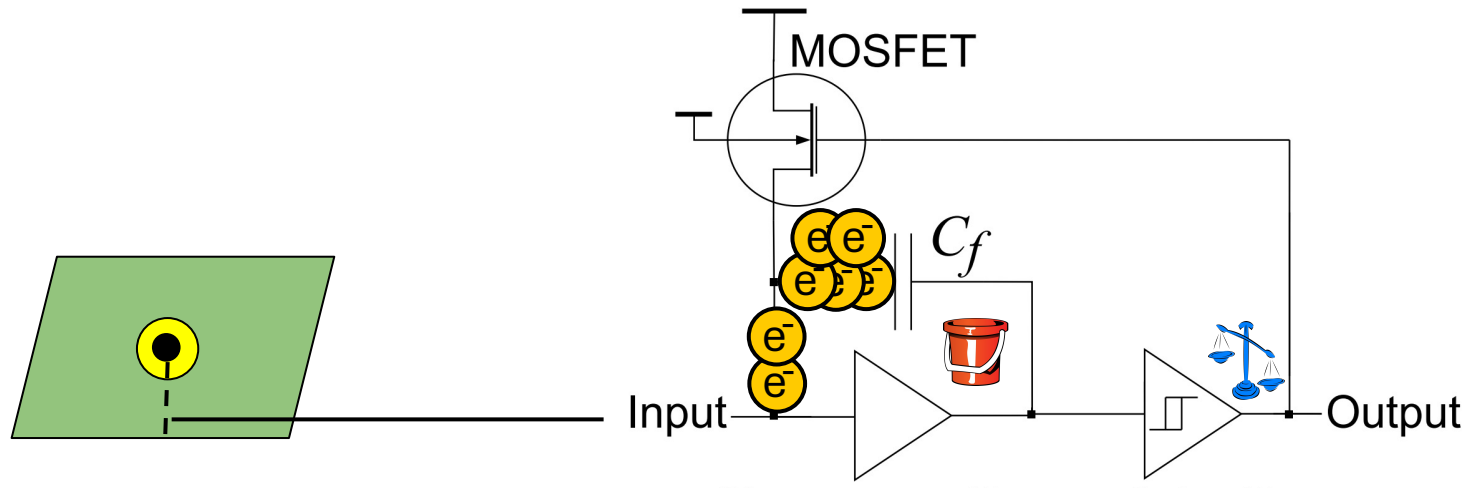


※ the reset happens for 5 electrons

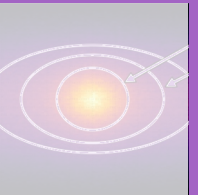
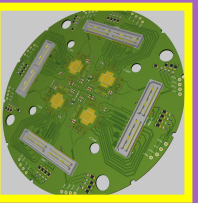




Toy Example

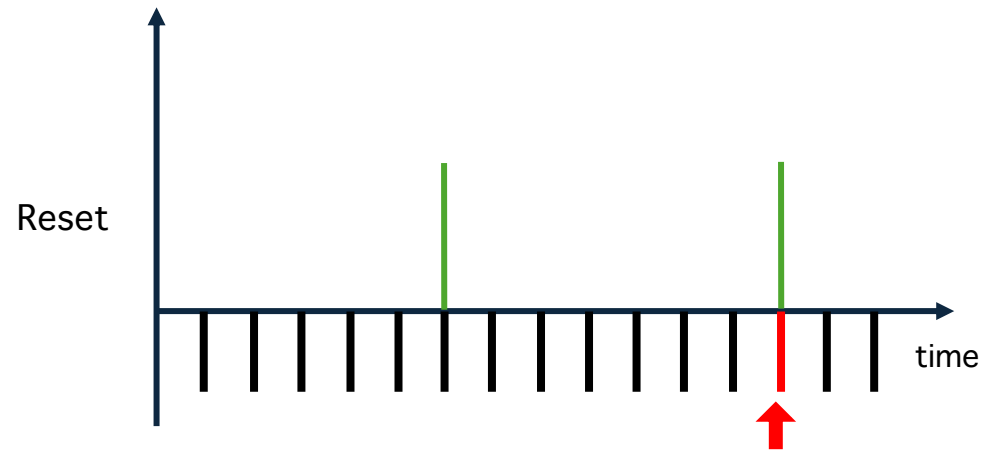
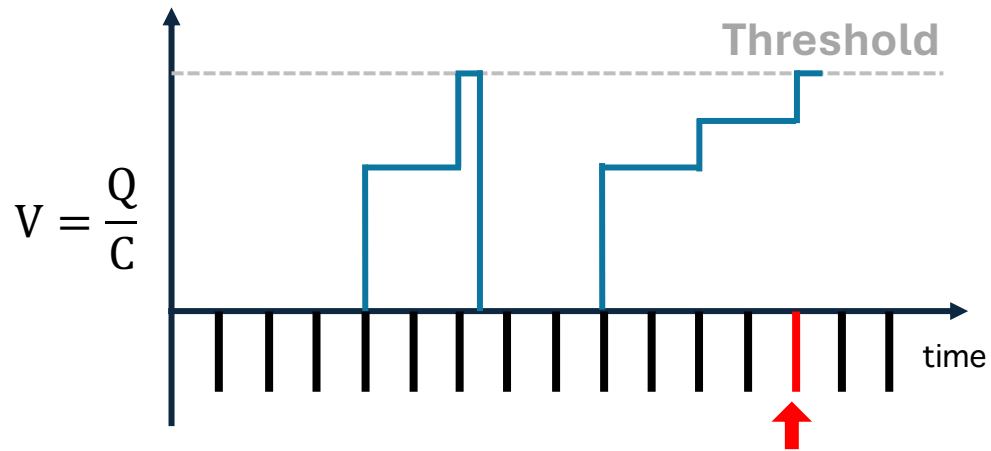
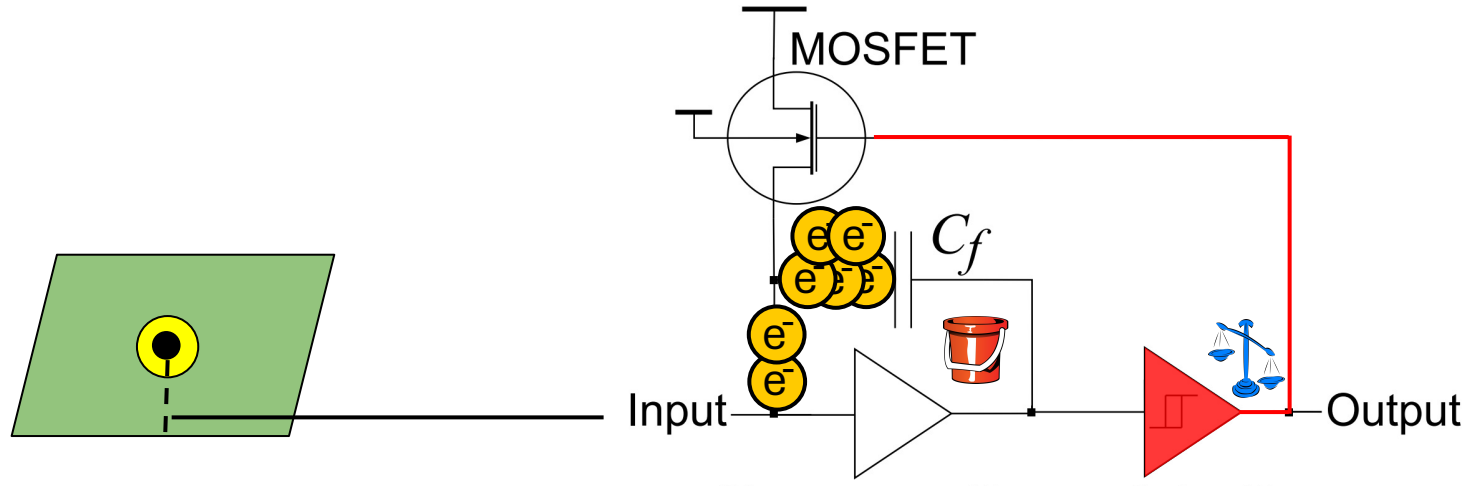


※ the reset happens for 5 electrons

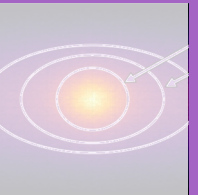
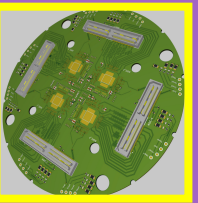




Toy Example

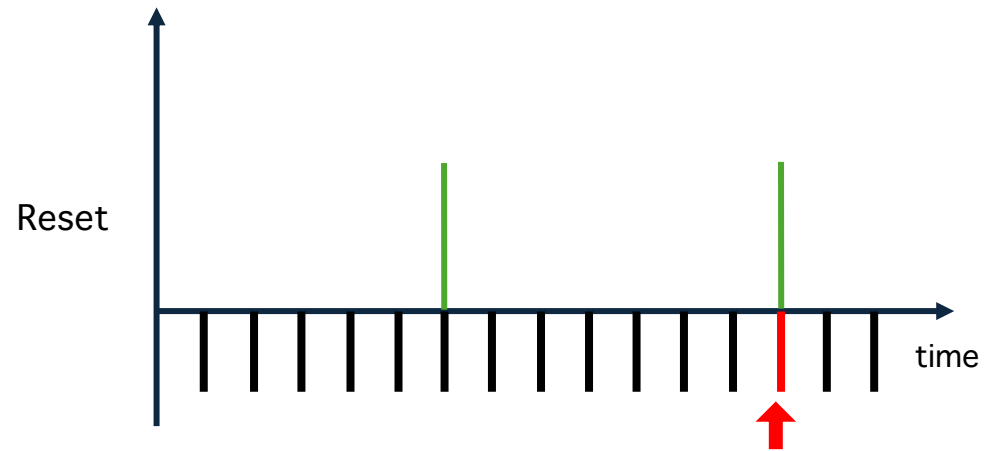
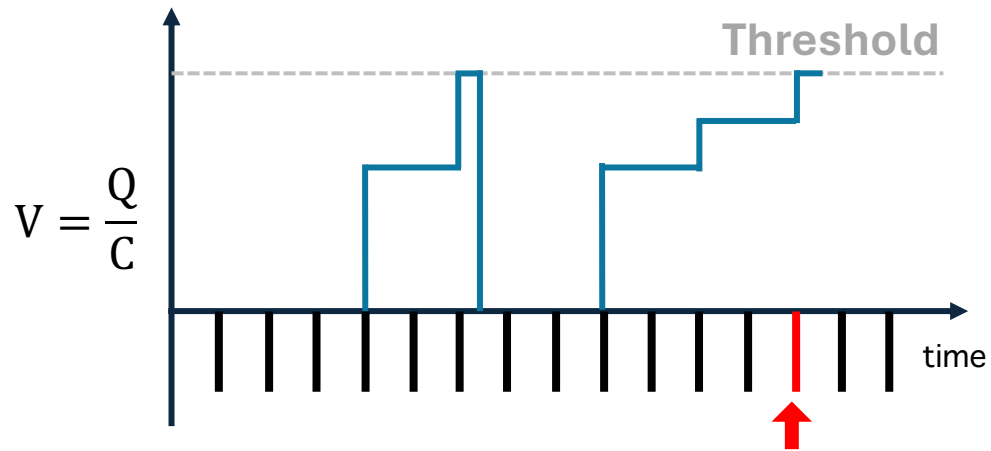
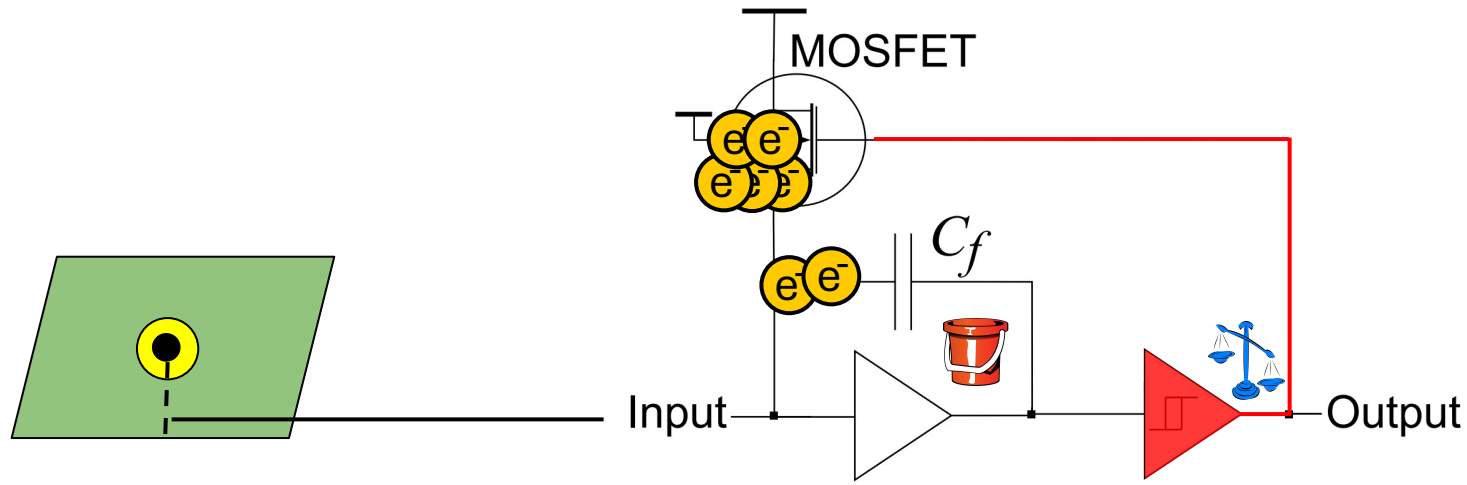


※ the reset happens for 5 electrons

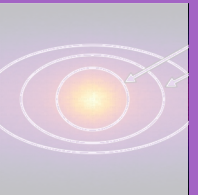
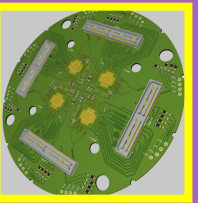




Toy Example

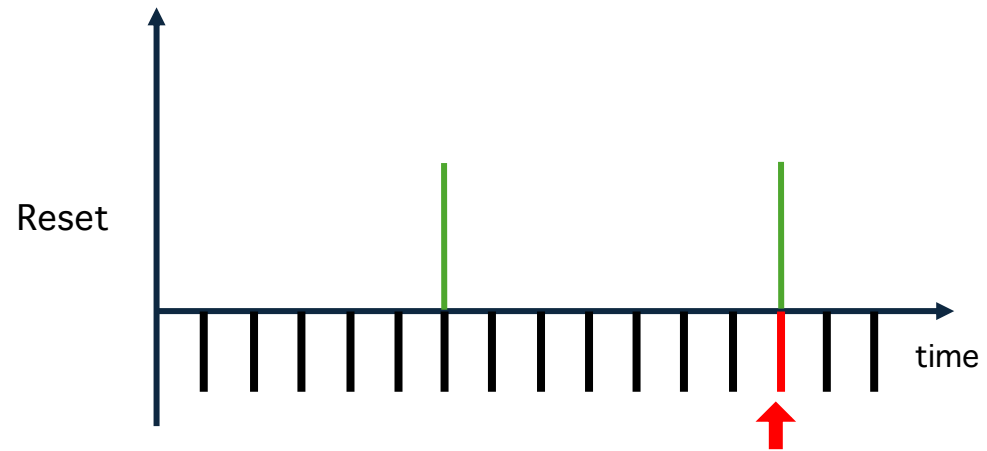
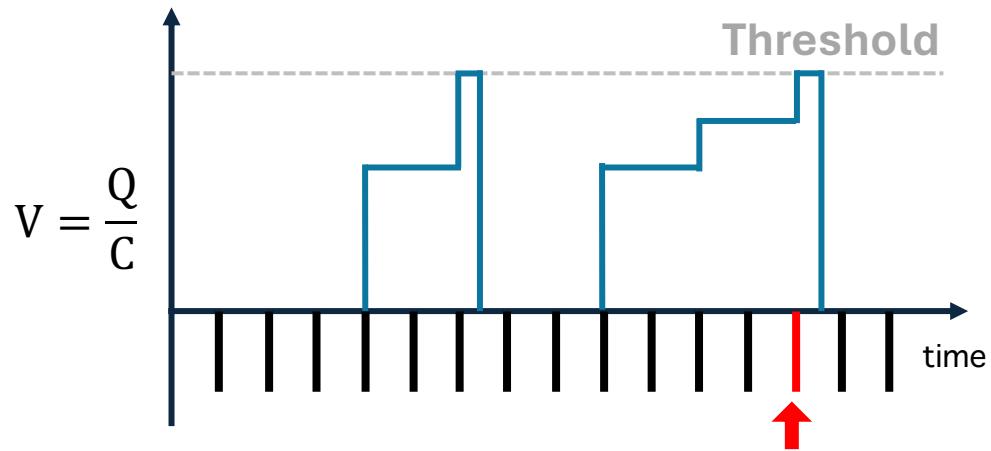
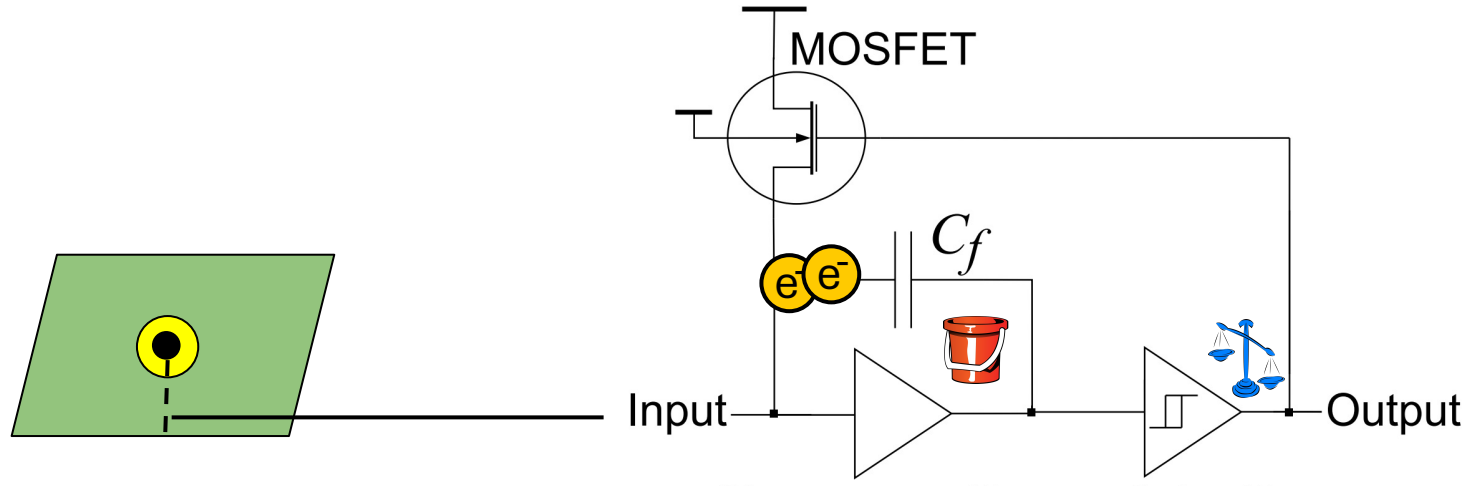


※ the reset happens for 5 electrons

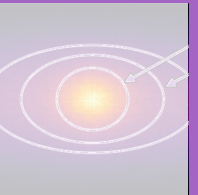
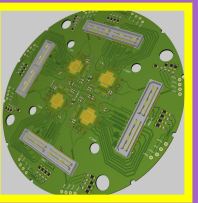




Toy Example

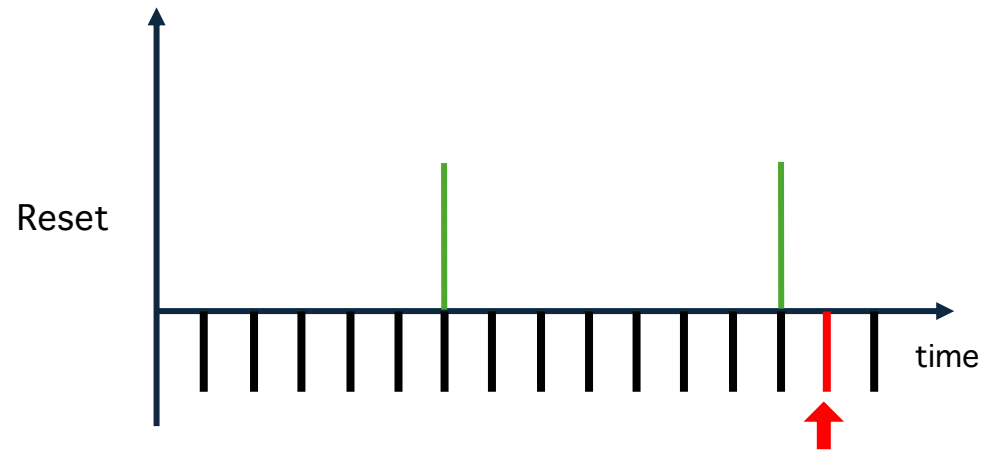
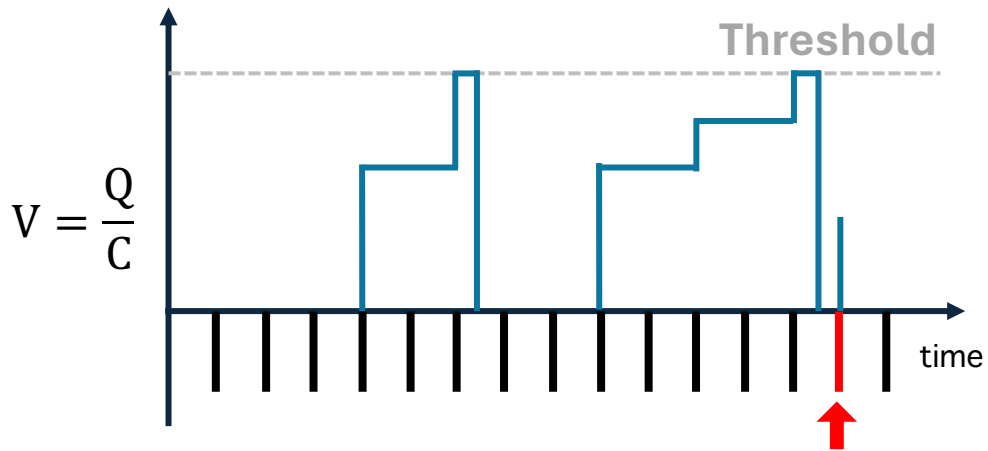
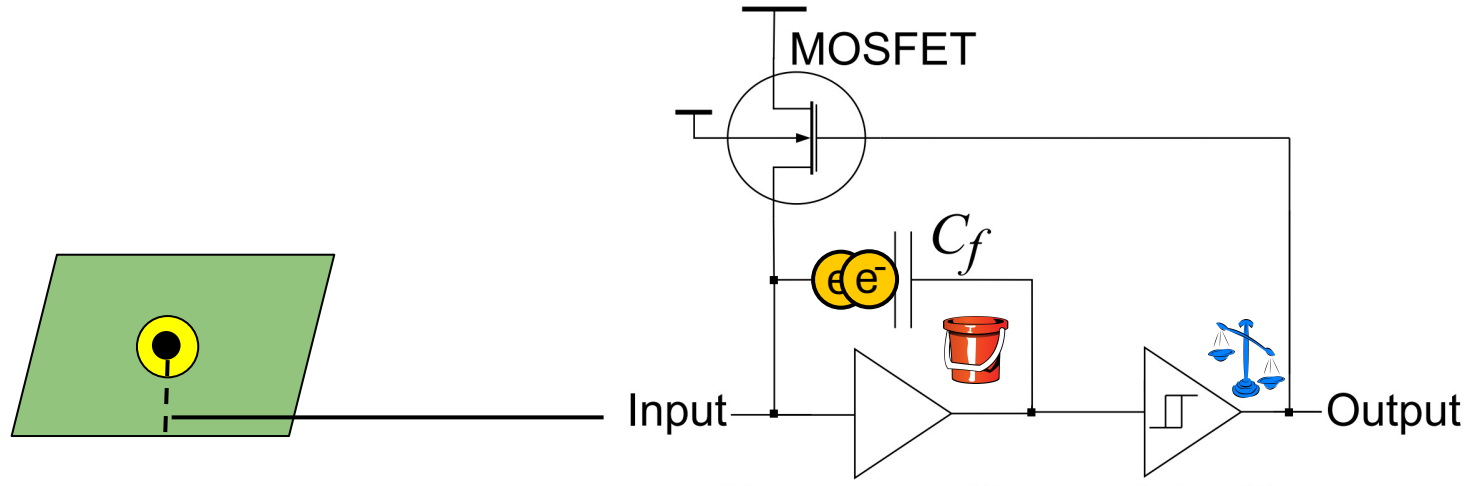


※ the reset happens for 5 electrons

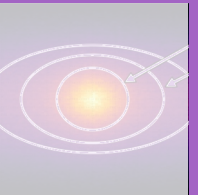
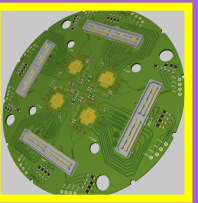




Toy Example



※ the reset happens for 5 electrons



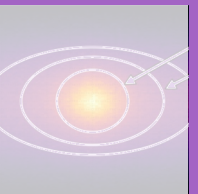
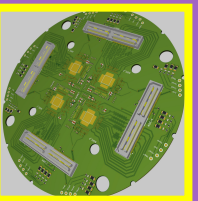
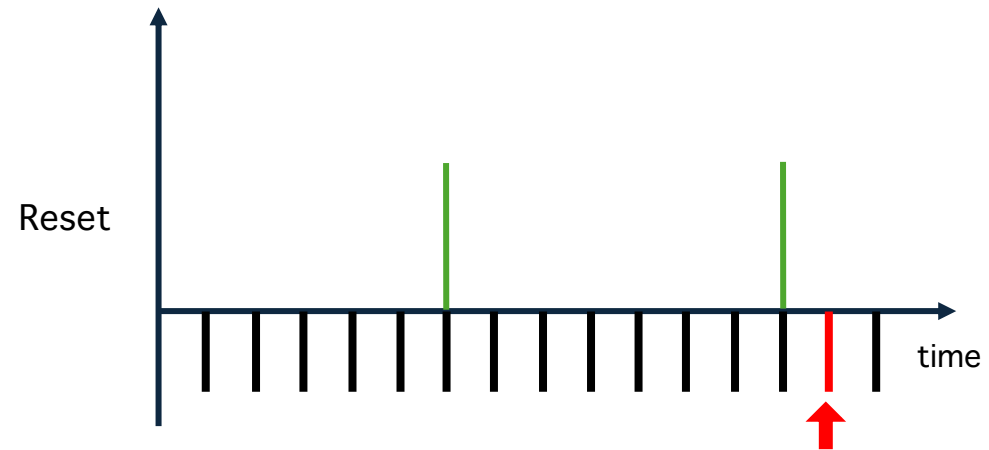
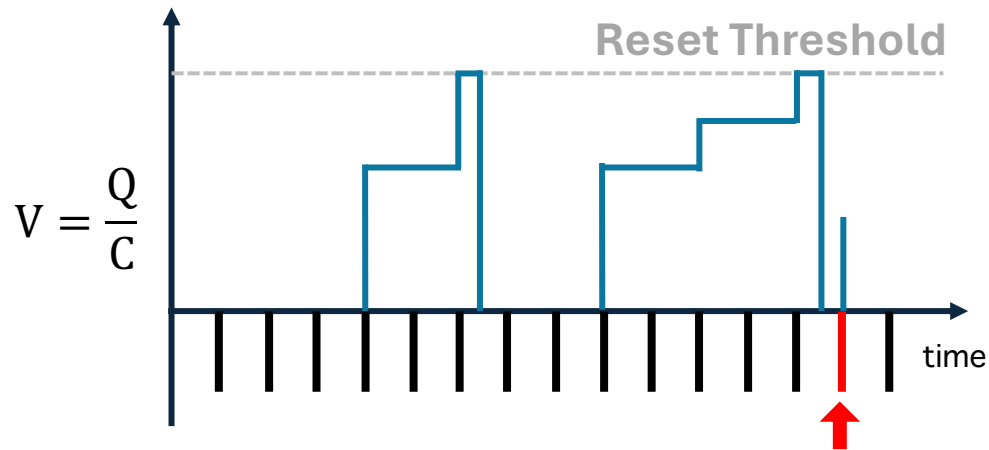


Toy Example

※ the reset happens for 5 electrons

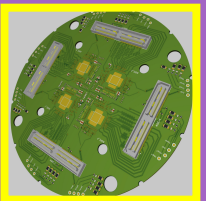
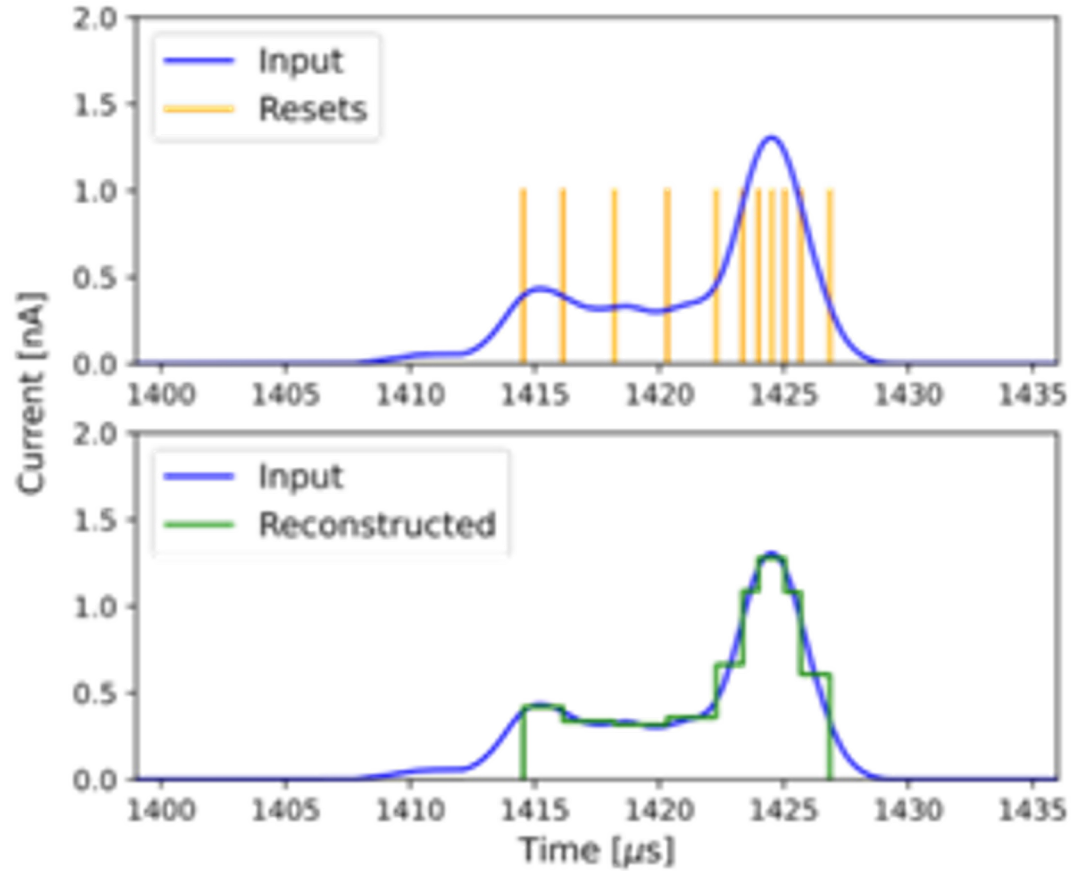
We are measuring
how long it took (Δt) to accumulate fixed amount of Q

Instead of recording the entire waveform,
we only need to record time stamps!



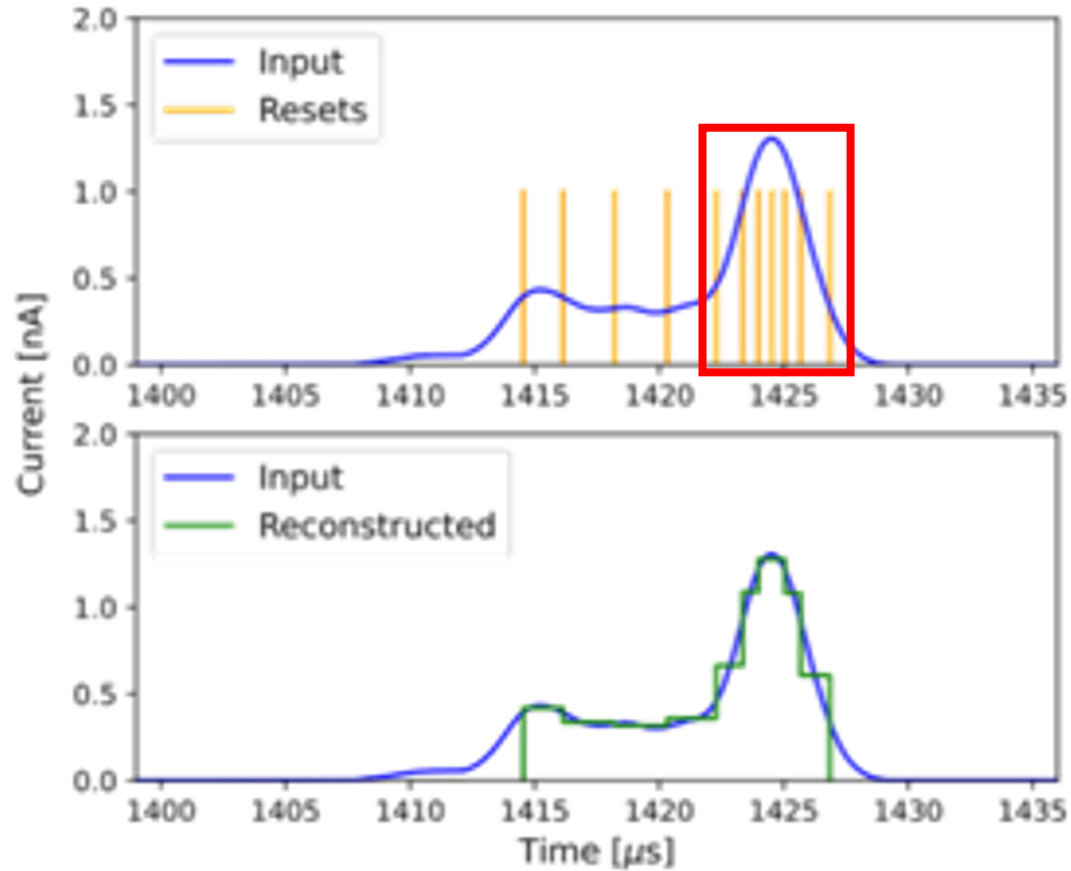


Solution : Pixelization

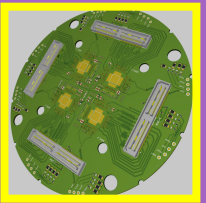




Solution : Pixelization

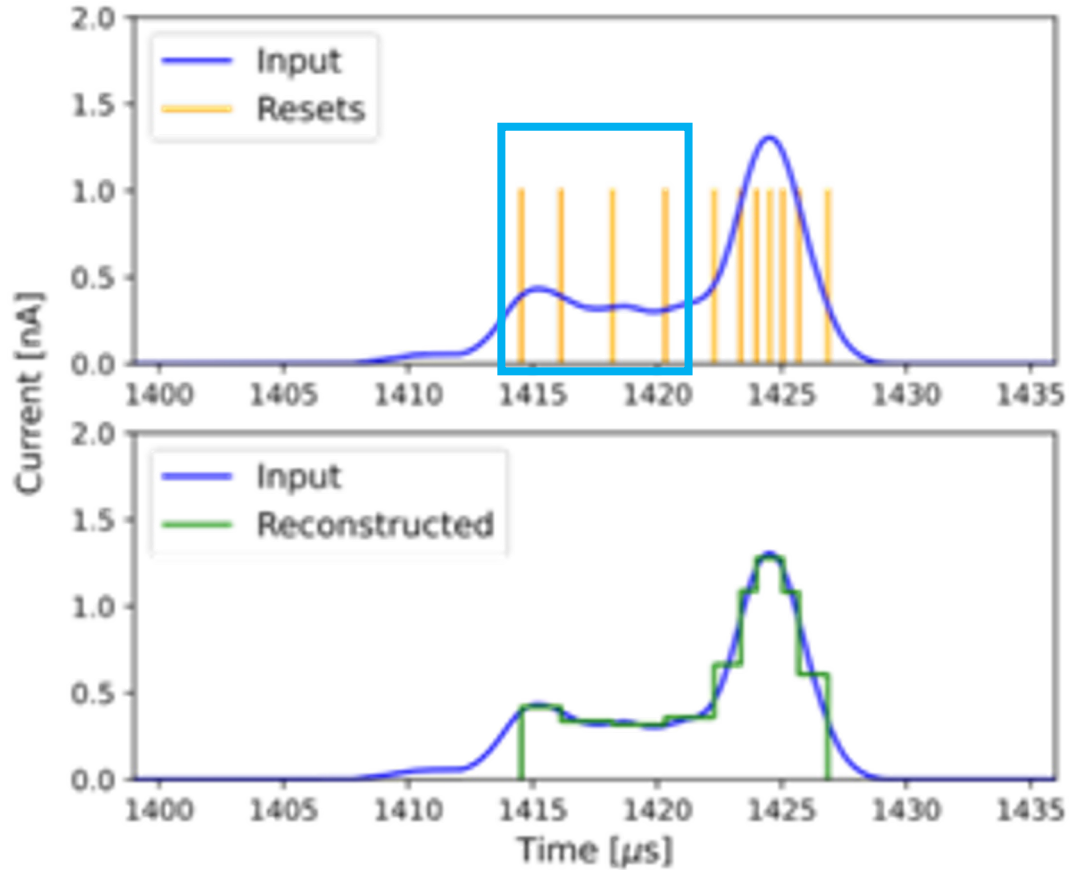


Denser area
↓
took pixel **less** time to get
unit amount of charge
↓
More current

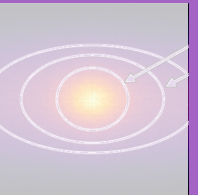
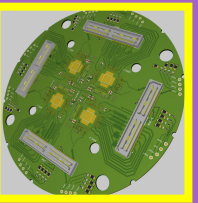




Solution : Pixelization

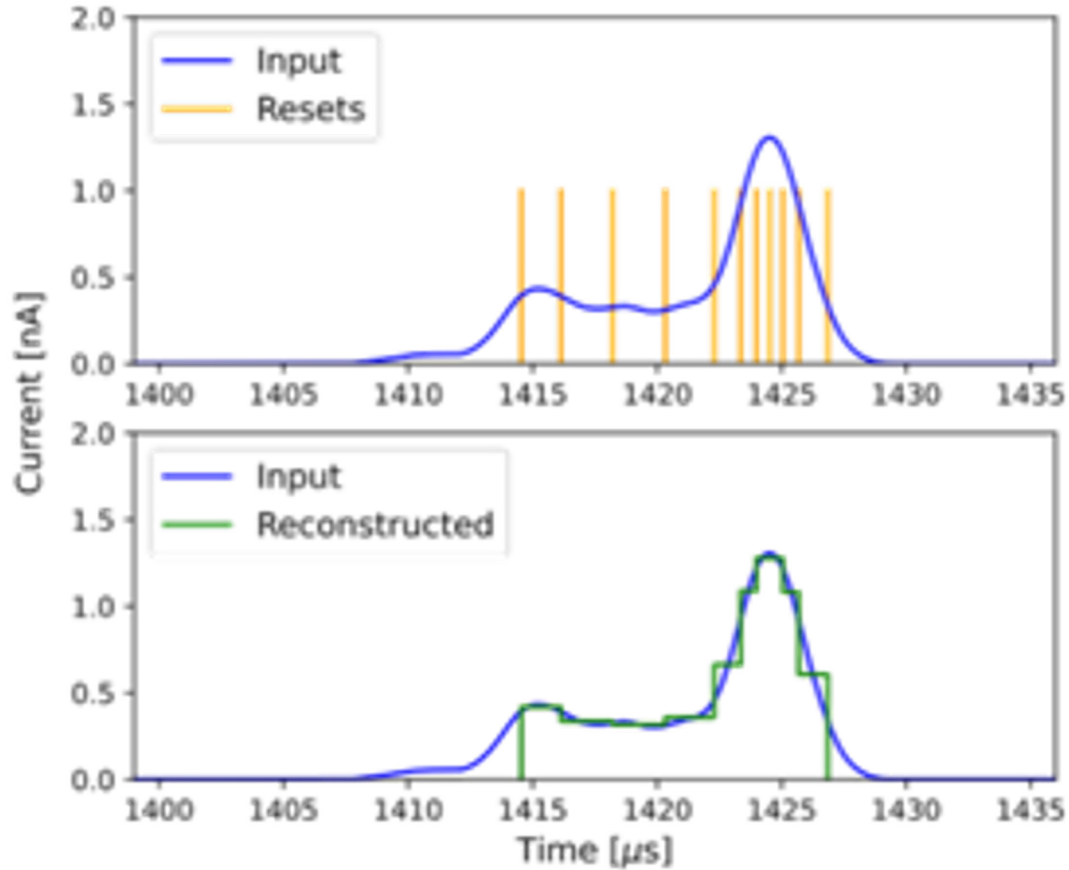


Sparse area
↓
took pixel **more** time to get
unit amount of charge
↓
Less current





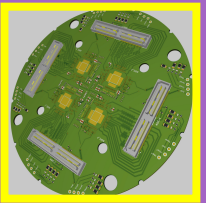
Solution : Pixelization



With this method,
the data rate is **10⁶ times less**
than the traditional wire readout!

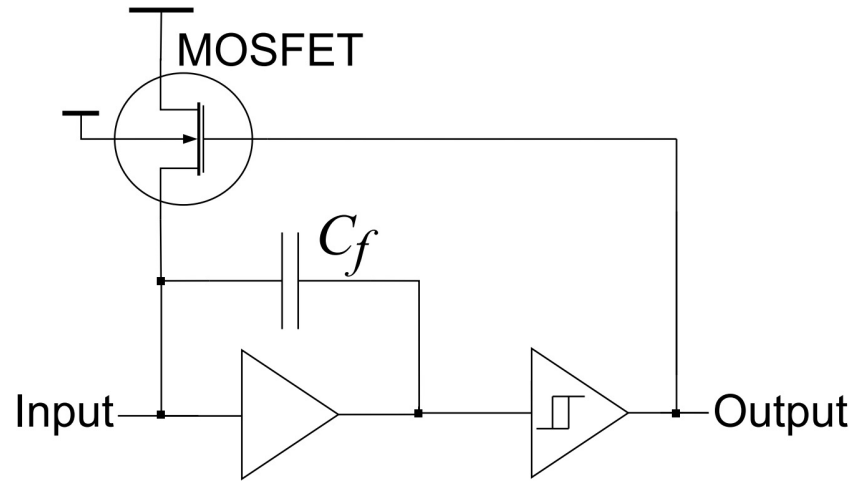
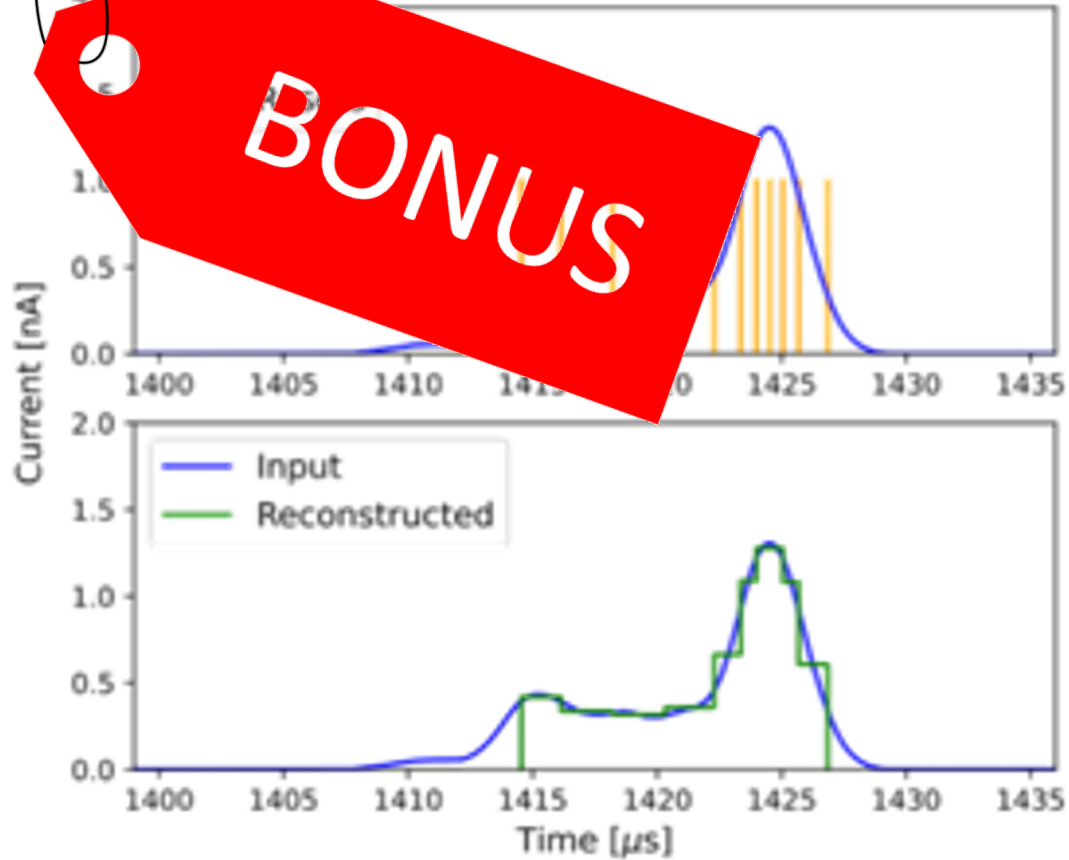


Q-Pix can offer a direct solution
to the data-rate problem that
pixelized kton-scale LArTPC has

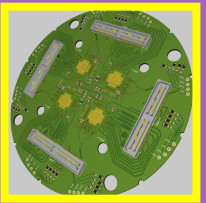




Solution : Pixelization

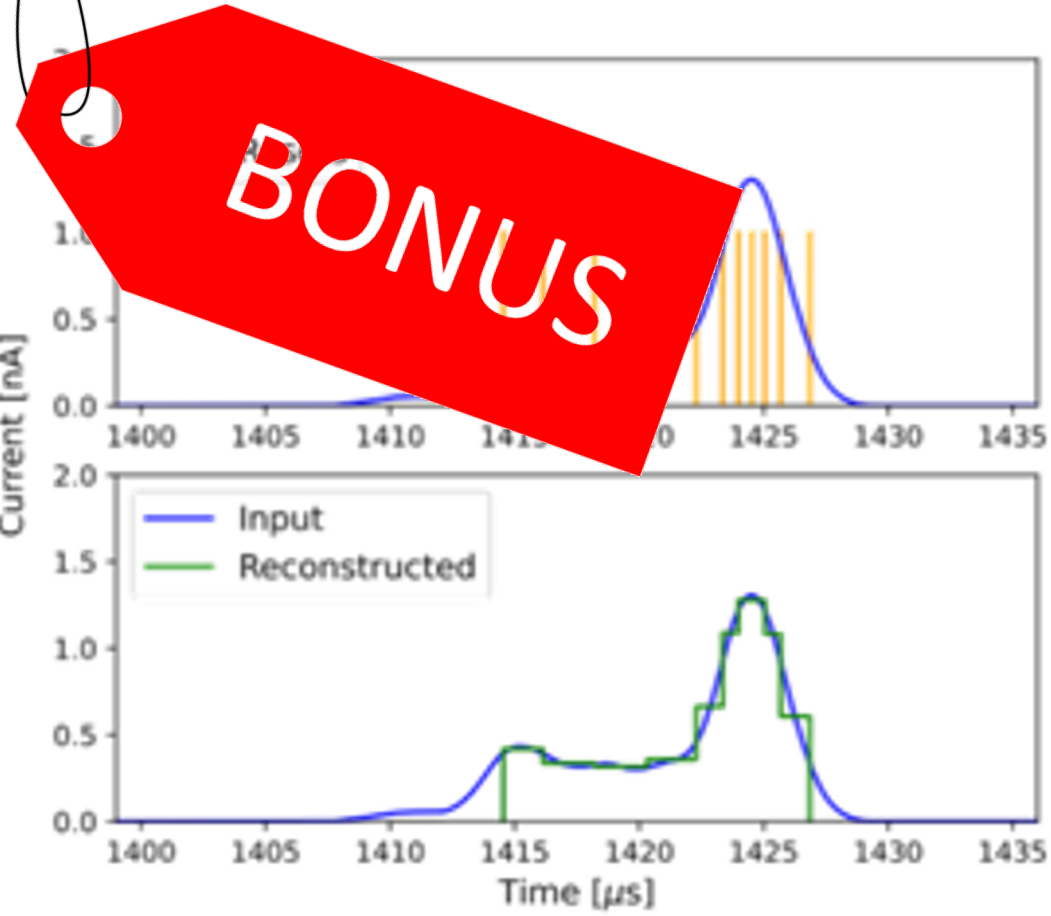


Moreover, this can be seen as
one-bit Sigma-Delta modulator

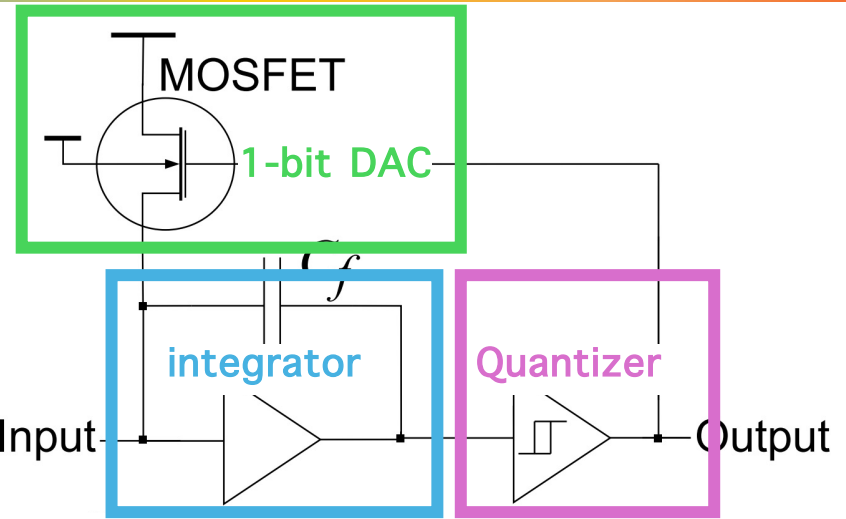




Solution : Pixelization



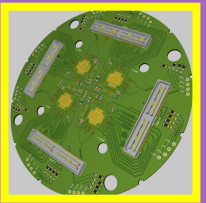
BONUS



Moreover, this can be seen as

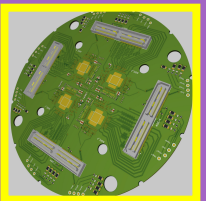
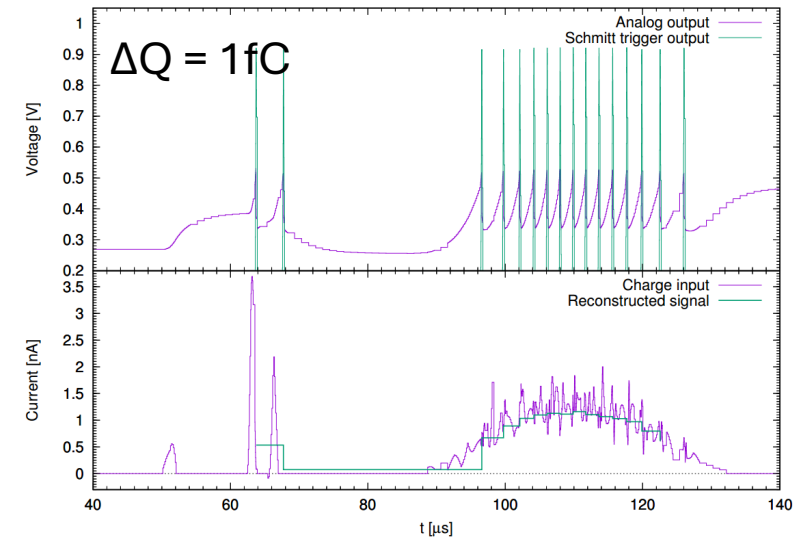
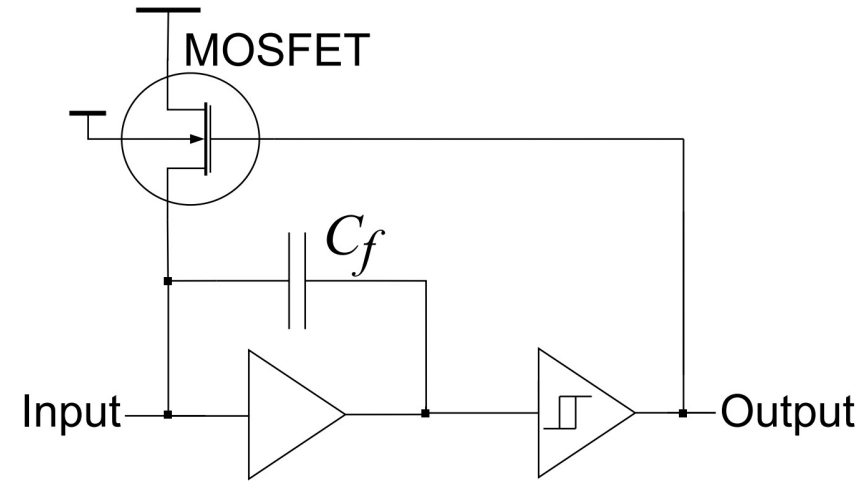
one-bit Sigma-Delta modulator

- Greater flexibility in data type
- More efficient data analysis
- More applications



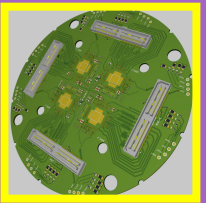
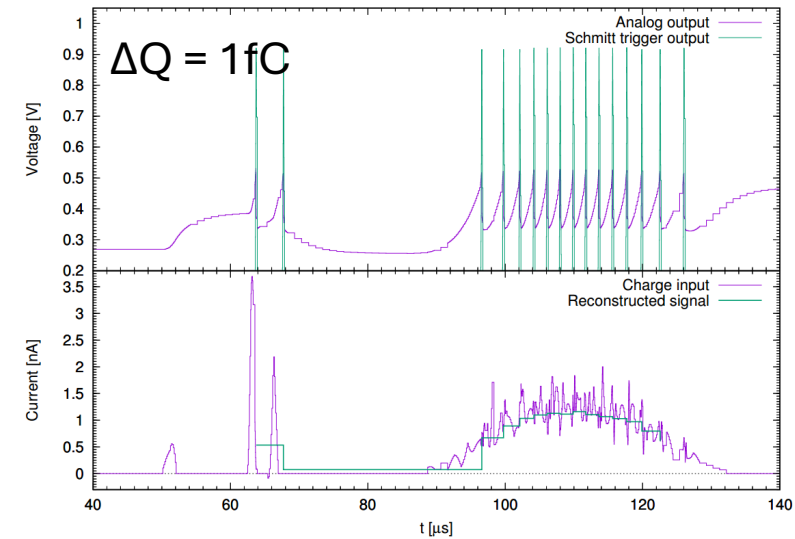
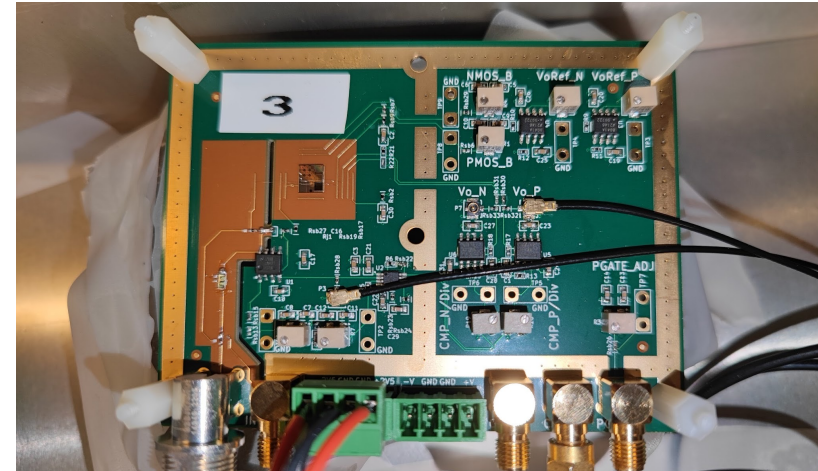


Demonstrating Q-Pix with commercial components



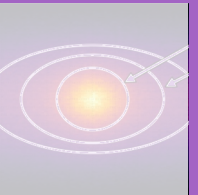
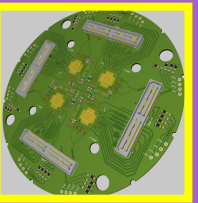
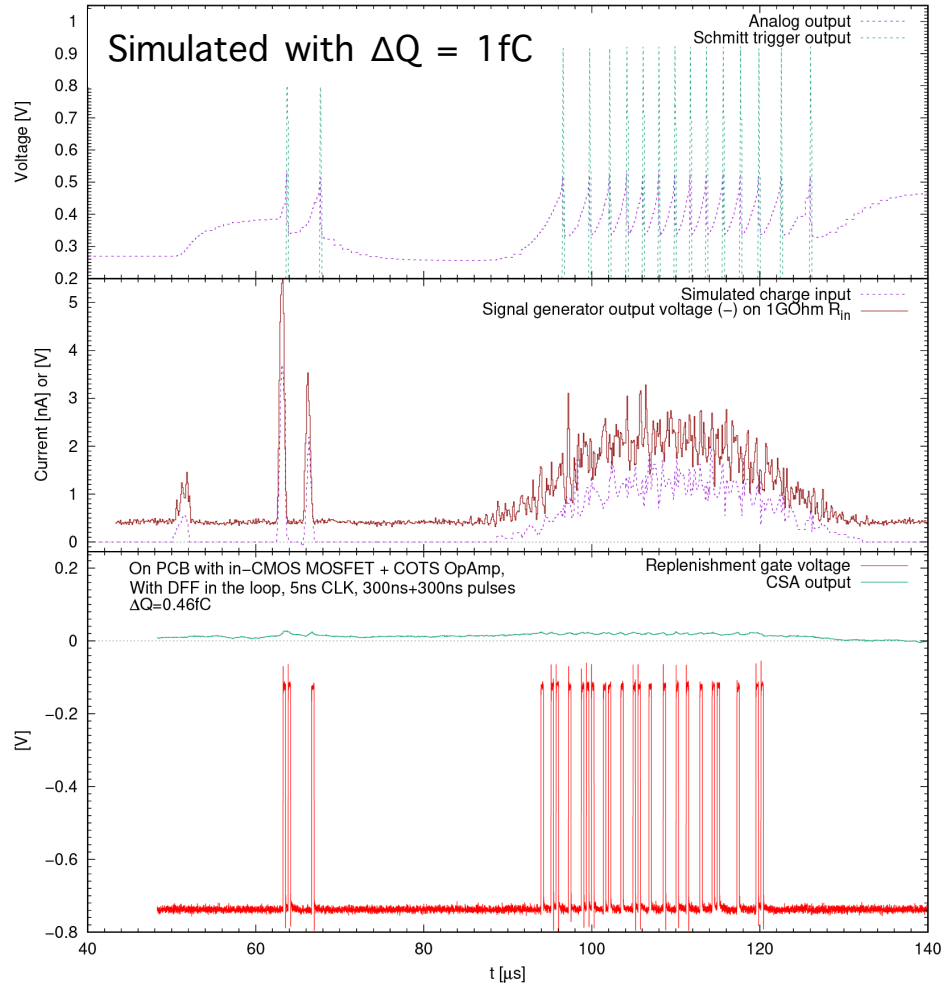
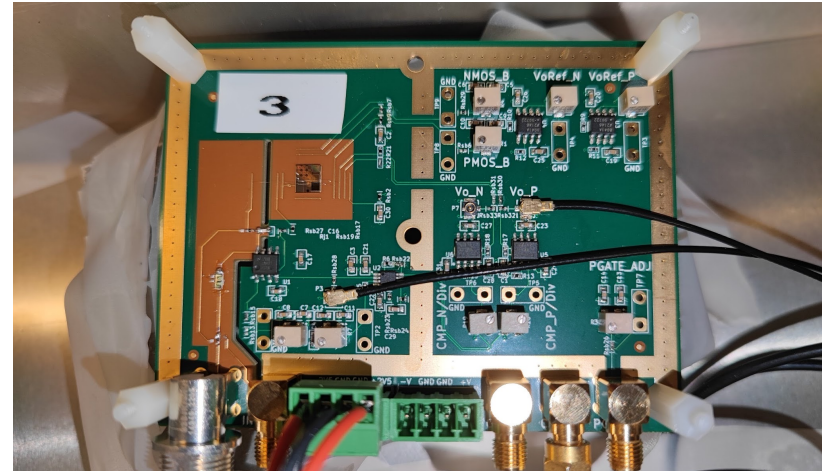


Demonstrating Q-Pix with commercial components





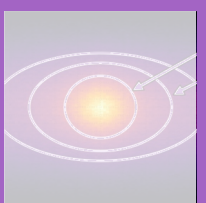
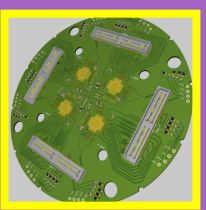
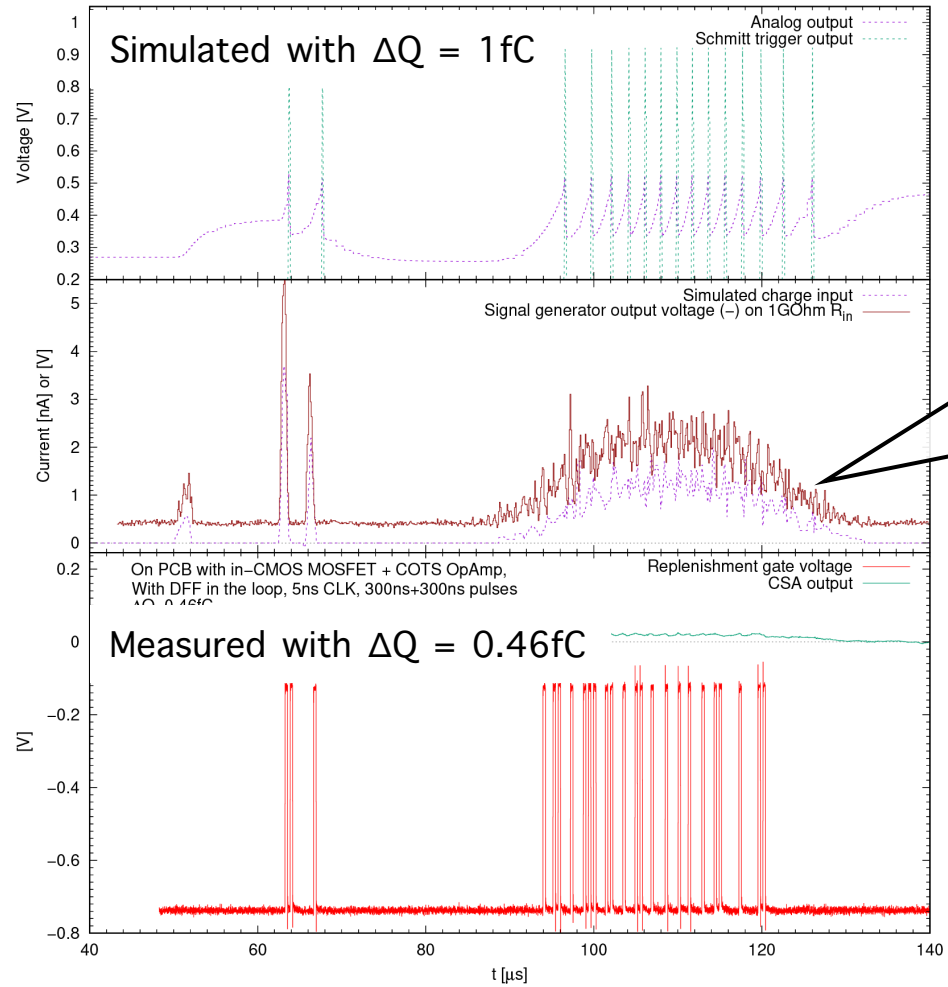
Demonstrating Q-Pix with commercial components





Demonstrating Q-Pix with commercial components

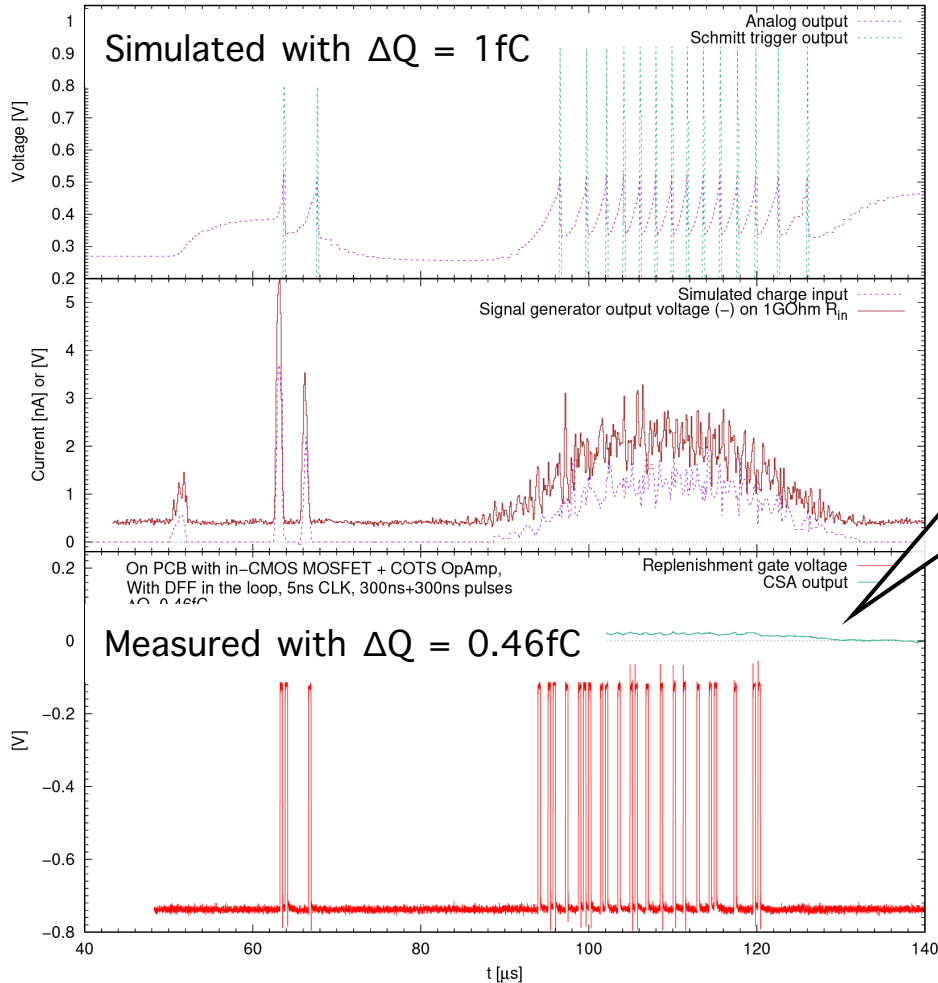
1. We generated the charge input same as the signal we used for simulation.





Demonstrating Q-Pix with commercial components

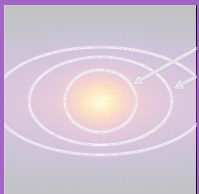
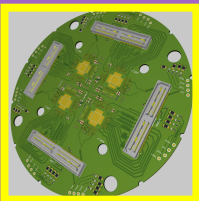
1. We generated the charge input same as the signal we used for simulation.
2. Fed it to the Q-Pix front-end made with commercial components
3. We could get better results than simulation!



0.46 fC enables ~50 KeV signal detection!

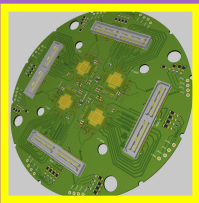
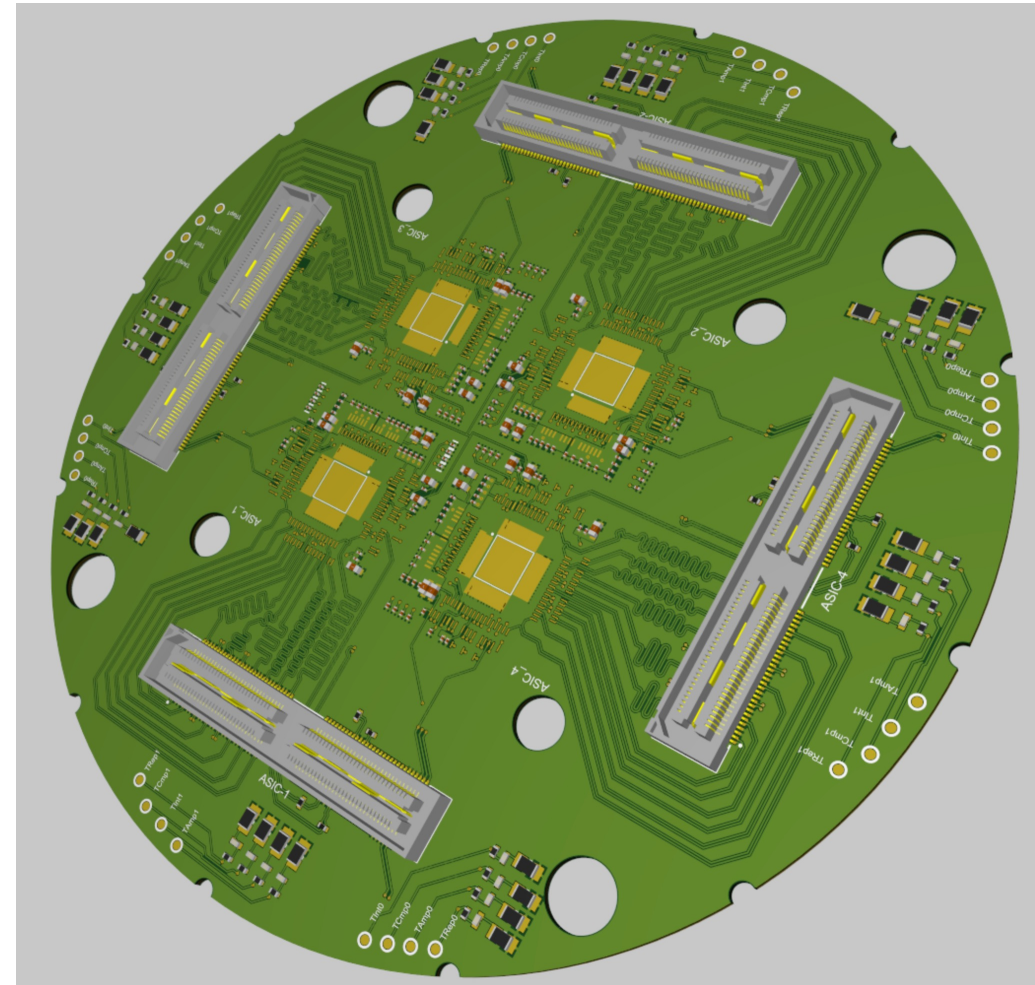
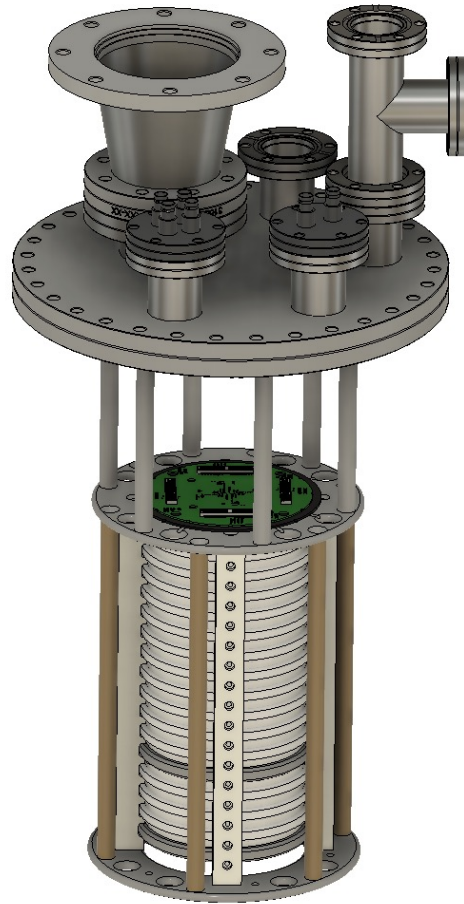
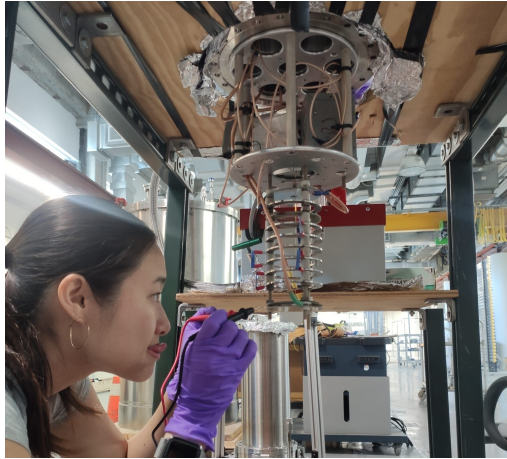
Demonstrating the Q-Pix front-end using discrete OpAmp and CMOS transistors

Peng Miao^a, Jonathan Asaadi^b, James B. R. Battat^d, Mikyung Han^a, Kevin Keefe^{e,b}, S. Kohani^e, Austin D. McDonald^{b,c}, David Nygren^b, Olivia Seidel^b, Yuan Mei^{a,b,*}





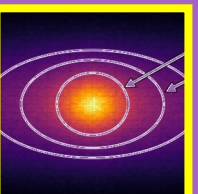
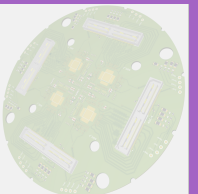
Future of Q-Pix





Physics with Q-Pix

'maximize the discovery potential of a kiloton scale LArTPC'





Physics with Q-Pix

‘maximize the discovery potential of a kiloton scale LArTPC’

Paper published from
JINST in 2020

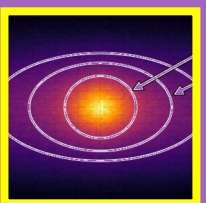
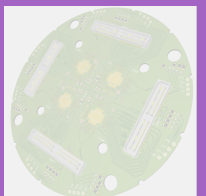
**Enhancing
beam event studies**

Paper published from
PRD in 2022

**Enhancing
supernovae studies**

Paper
In preparation

**Enabling
solar studies**





Physics with Q-Pix

‘maximize the discovery potential of a kiloton scale LArTPC’

Paper published from JINST in 2020

Enhancing beam event studies

Paper published from PRD in 2022

Enhancing supernovae studies

Paper In preparation

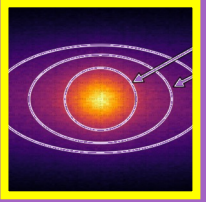
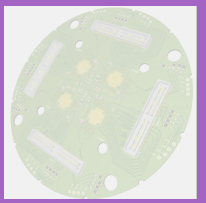
Enabling solar studies



~20GeV range : high energy neutrinos

~40MeV range : low energy neutrinos

~20MeV range : very low energy neutrinos





Beam physics (~20GeV)

Paper published from
JINST in 2020

**Enhancing
beam event studies**

Paper published from
PRD in 2022

**Enhancing
supernovae studies**

Paper
In preparation

**Enabling
solar studies**

Enhancing neutrino event reconstruction with pixel-based 3D readout for liquid argon time projection chambers

C. Adams,^{a,1} M. Del Tutto,^{b,c} J. Asaadi,^d M. Bernstein,^c E. Church,^e R. Guenette,^c
J.M. Rojas,^{c,2} H. Sullivan^d and A. Tripathi^d

^aArgonne National Laboratory, Lemont, IL, U.S.A.

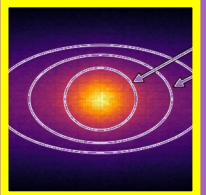
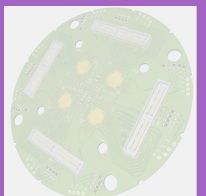
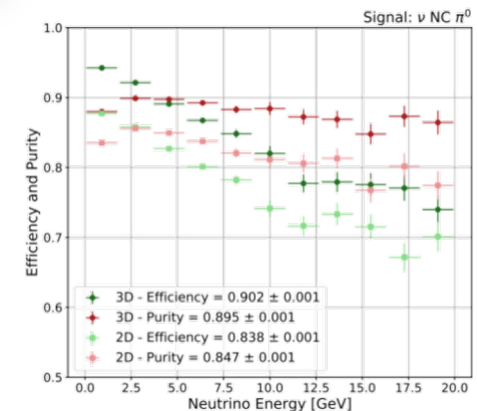
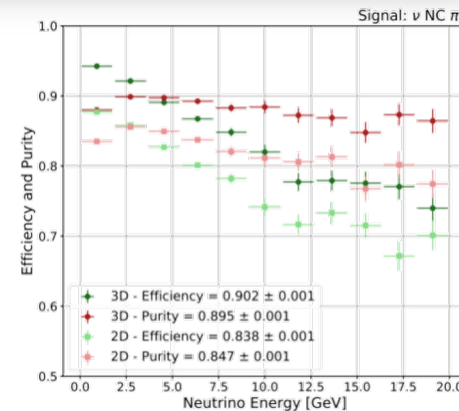
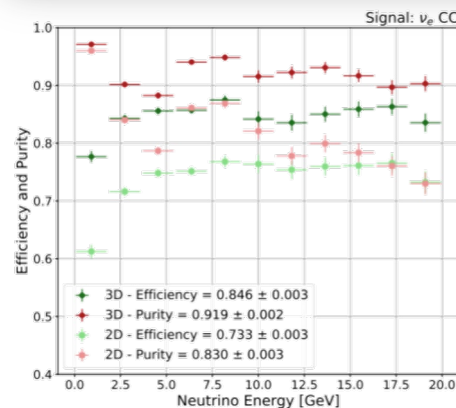
^bFermi National Accelerator Laboratory, Batavia, IL, U.S.A.

^cHarvard University, Cambridge, MA, U.S.A.

^dUniversity of Texas-Arlington, Arlington, TX, U.S.A.

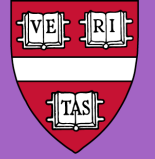
^ePacific Northwest National Laboratory, Richland, WA, U.S.A.

Category	Accuracy [%]	
	3D	2D
Neutrino Interaction	94	91
Proton Multiplicity	91	87
Charge Pion Presence	94	91
Neutral Pion Presence	95	94



Beam physics ($\sim 20\text{GeV}$)

Good tracking



Paper published from
JINST in 2020

**Enhancing
beam event studies**

Paper published from
PRD in 2022

**Enhancing
supernovae studies**

Paper
In preparation

**Enabling
solar studies**

Enhancing neutrino event reconstruction with pixel-based 3D readout for liquid argon time projection chambers

C. Adams,^{a,1} M. Del Tutto,^{b,c} J. Asaadi,^d M. Bernstein,^c E. Church,^e R. Guenette,^c
J.M. Rojas,^{c,2} H. Sullivan^d and A. Tripathi^d

^aArgonne National Laboratory, Lemont, IL, U.S.A.

^bFermi National Accelerator Laboratory, Batavia, IL, U.S.A.

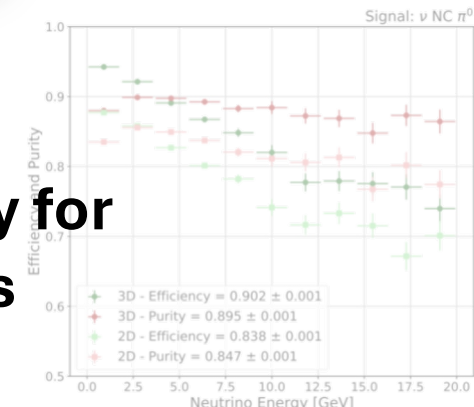
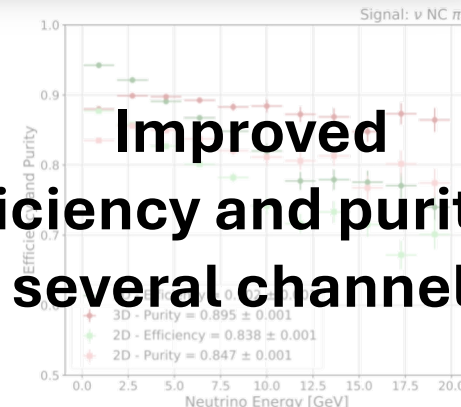
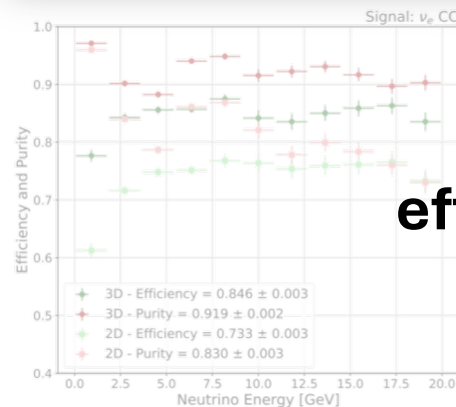
^cHarvard University, Cambridge, MA, U.S.A.

^dUniversity of Texas-Arlington, Arlington, TX, U.S.A.

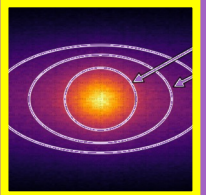
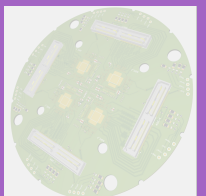
^ePacific Northwest National Laboratory, Richland, WA, U.S.A.

Enhanced
event reconstruction
accuracy

	2D	3D
Accuracy [%]	88	91
Track Momentum	91	97
Charge	91	91
Neutral Pos Process	95	94



**Improved
efficiency and purity for
several channels**



Supernova physics ($\sim 40\text{MeV}$)

Good tracking



Paper published from
JINST in 2020

**Enhancing
beam event studies**

Paper published from
PRD in 2022

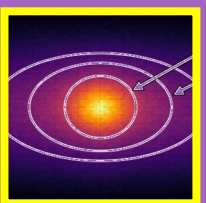
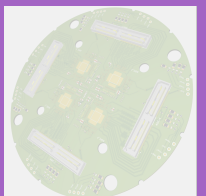
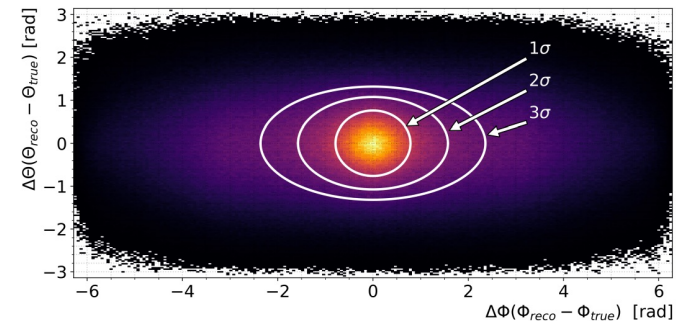
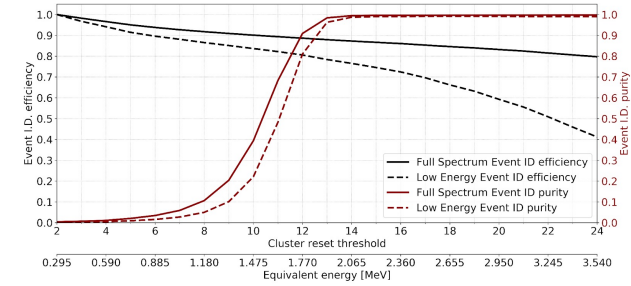
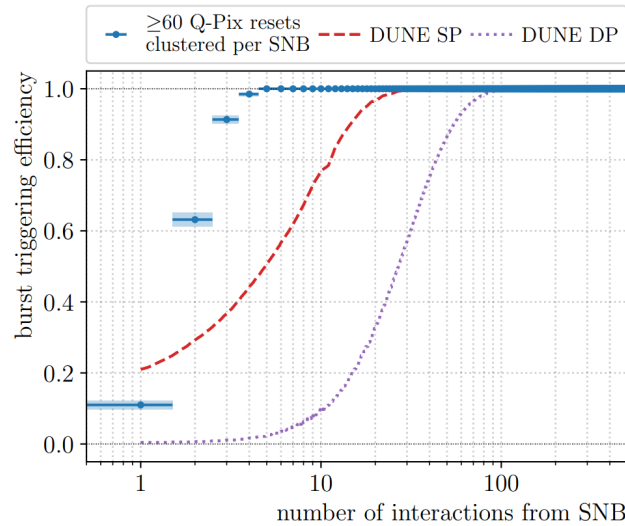
**Enhancing
supernovae studies**

Paper
In preparation

**Enabling
solar studies**

Enhanced low-energy supernova burst detection in large liquid argon time projection chambers enabled by Q-Pix (The Q-Pix Collaboration)

S. Kubota,¹ J. Ho,^{1,*} A.D. McDonald,^{2,1,†} N. Tata,¹ J. Asaadi,² R. Guenette,^{3,1} J.B.R. Battat,⁴
D. Braga,⁵ M. Demarteau,⁶ Z. Djurcic,⁷ M. Febraro,⁶ E. Gramellini,⁵ S. Kohani,⁸ C. Mauger,⁹ Y. Mei,¹⁰
F.M. Newcomer,⁹ K. Nishimura,⁸ D. Nygren,² R. Van Berg,⁹ G.S. Varner,⁸ and K. Woodworth⁵



Supernova physics ($\sim 40\text{MeV}$)

Good tracking
 Low energy threshold
 Better energy resolution



Paper published from JINST in 2020

Enhancing beam event studies

Paper published from PRD in 2022

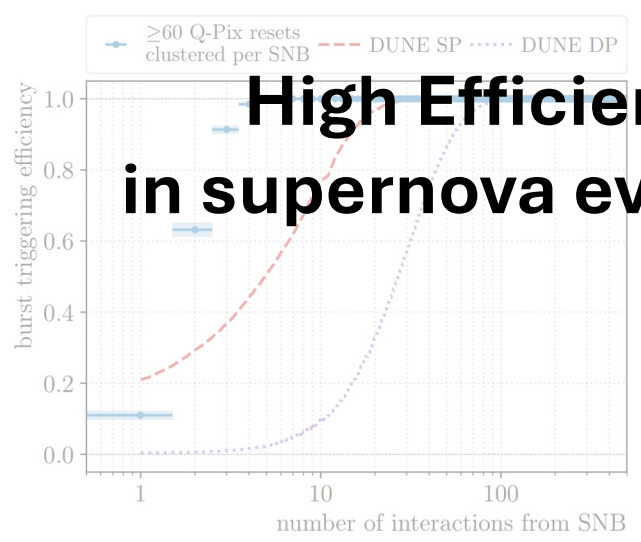
Enhancing supernovae studies

Paper In preparation

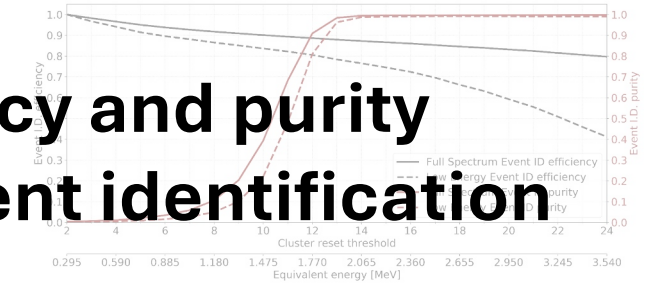
Enabling solar studies

Enhanced low-energy supernova burst detection in large liquid argon time projection chambers enabled by Q-Pix
 (The Q-Pix Collaboration)

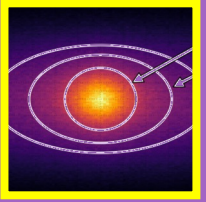
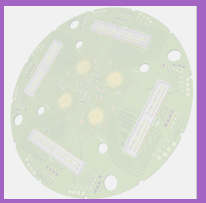
S. Kubota,¹ J. Ho,^{1,*} A.D. McDonald,^{2,1,†} N. Tata,¹ J. Asaadi,² R. Guenette,^{3,1} J.B.R. Battat,⁴ D. Braga,⁵ M. Demarteau,⁶ Z. Djurcic,⁷ M. Febraro,⁶ E. Gramellini,⁵ S. Kohani,⁸ C. Mauger,⁹ Y. Mei,¹⁰ F.M. Newcomer,⁹ K. Nishimura,⁸ D. Nygren,² R. Van Berg,⁹ G.S. Varner,⁸ and K. Woodworth⁵



High Efficiency and purity in supernova event identification



Supernova pointing



Solar physics ($\sim 20\text{MeV}$)

Good tracking
Low energy threshold
Better energy resolution



Solar neutrinos offer many opportunities

- Enhancement in oscillation parameters measurement
- Understanding solar models
- Potential discovery of hep neutrinos

→ Requires :

better tracking

lower energy threshold

better energy resolution

Paper published from
JINST in 2020

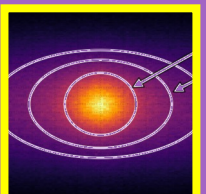
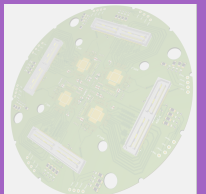
**Enhancing
beam event studies**

Paper published from
PRD in 2022

**Enhancing
supernovae studies**

Paper
In preparation

**Enabling
solar studies**



Solar physics ($\sim 20\text{MeV}$)

Good tracking
Low energy threshold
Better energy resolution



Solar neutrinos offer many opportunities

- Enhancement in oscillation parameters measurement
- Understanding solar models
- Potential discovery of hep neutrinos

→ Requires :

better tracking
lower energy threshold
better energy resolution

Exactly the
improvements
Q-Pix can offer!

Paper published from
JINST in 2020

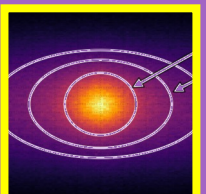
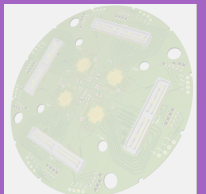
**Enhancing
beam event studies**

Paper published from
PRD in 2022

**Enhancing
supernovae studies**

Paper
In preparation

**Enabling
solar studies**



Solar physics ($\sim 20\text{MeV}$)

Good tracking
Low energy threshold
Better energy resolution



Solar neutrinos offer many opportunities

- Enhancement in oscillation parameters measurement
- Understanding solar models
- Potential discovery of hep neutrinos

→ Requires :

better tracking
lower energy threshold
better energy resolution

Exactly the
improvements
Q-Pix can offer!

→ Background control becomes crucial for continuous readouts
Paper on this is in preparation. Stay tuned! 😊

Paper published from
JINST in 2020

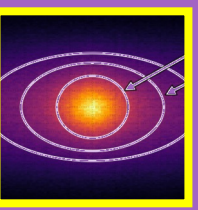
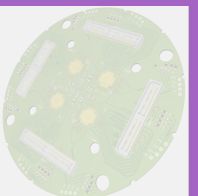
**Enhancing
beam event studies**

Paper published from
PRD in 2022

**Enhancing
supernovae studies**

Paper
In preparation

**Enabling
solar studies**



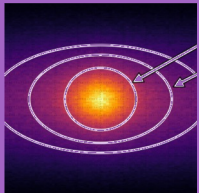
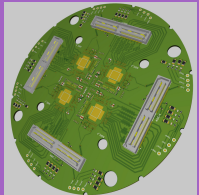


Summary

- Wire based LArTPC has been proven to be successful in many neutrino experiments but can be improved with pixel technology.
- Q-Pix is a new technology whose default state is ‘do-nothing’ – which is suited for large scale detector
- Q-Pix offers

Good tracking
Low energy threshold
Better energy resolution

 enabling low E neutrino studies.

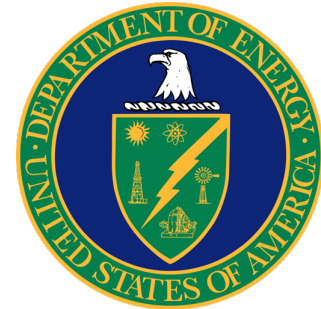




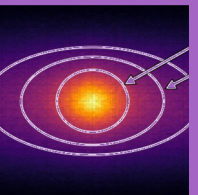
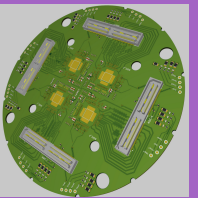
Thank you!



The University of Manchester



Q-Pix consortium would like to thank the DOE and the Gordon & Betty Moore Foundation for their support





Links to papers



Q-Pix: Pixel-scale Signal Capture for Kiloton Liquid Argon TPC Detectors: Time-to-Charge Waveform Capture, Local Clocks, Dynamic Networks

⇒ Q-Pix general info



Demonstrating the Q-Pix front-end using discrete OpAmp and CMOS transistors

⇒ Q-Pix commercial components result



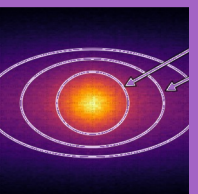
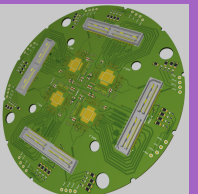
Enhancing Neutrino Event Reconstruction with Pixel-Based 3D Readout for Liquid Argon Time Projection Chambers

⇒ Beam physics with Q-Pix (High energy)



Enhanced low-energy supernova burst detection in large liquid argon time projection chambers enabled by Q-Pix

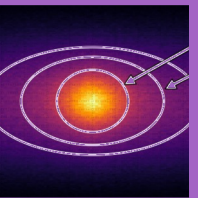
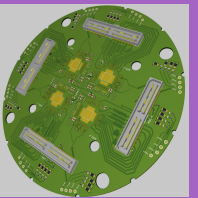
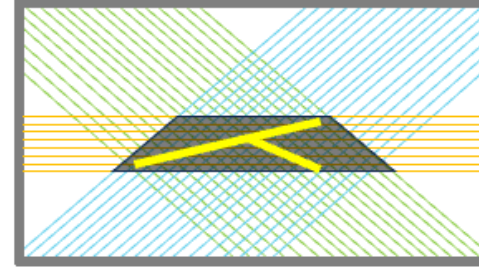
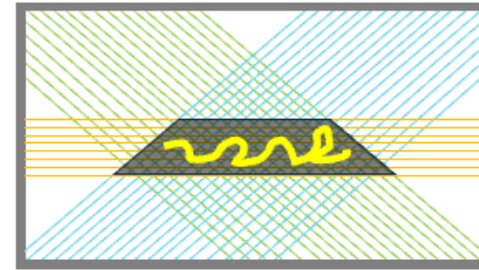
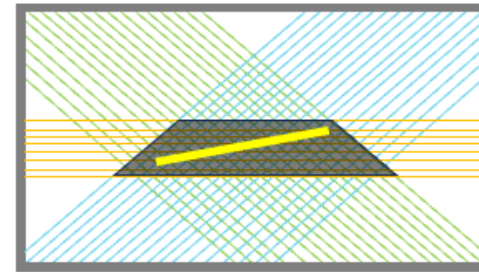
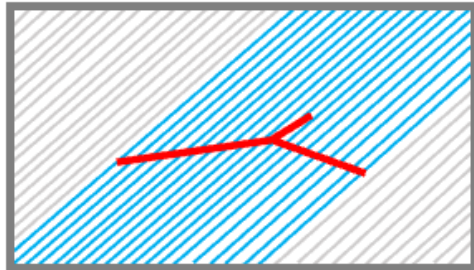
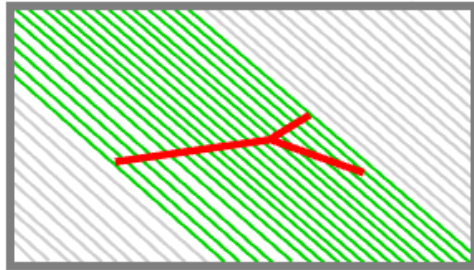
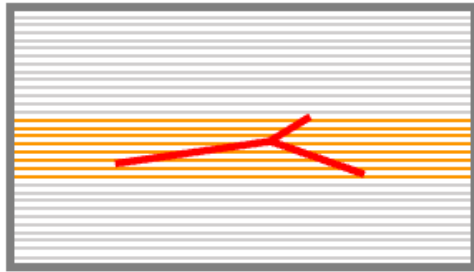
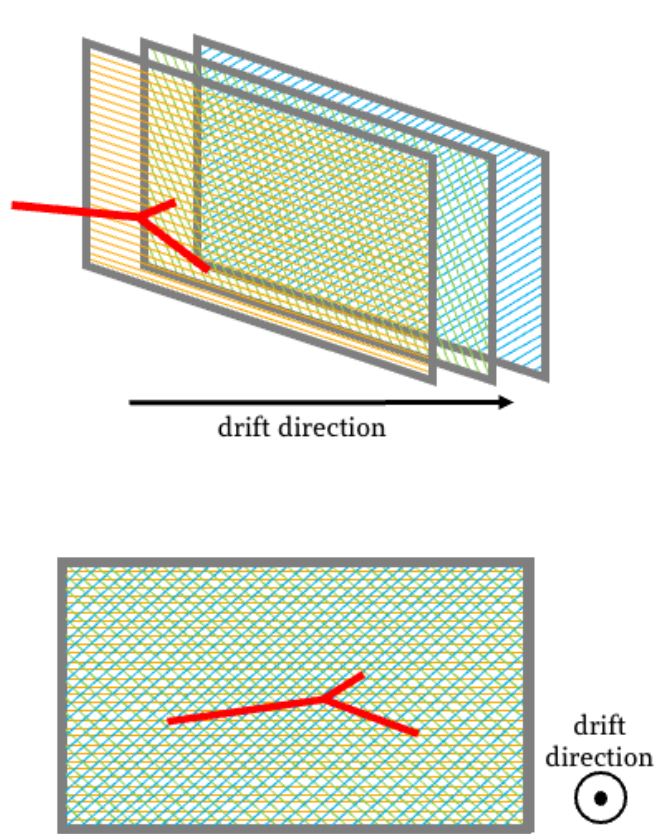
⇒ Supernovae neutrino physics with Q-Pix (Low energy)



Backup



Wire Ambiguity





Sigma Delta Modulator

Integrator

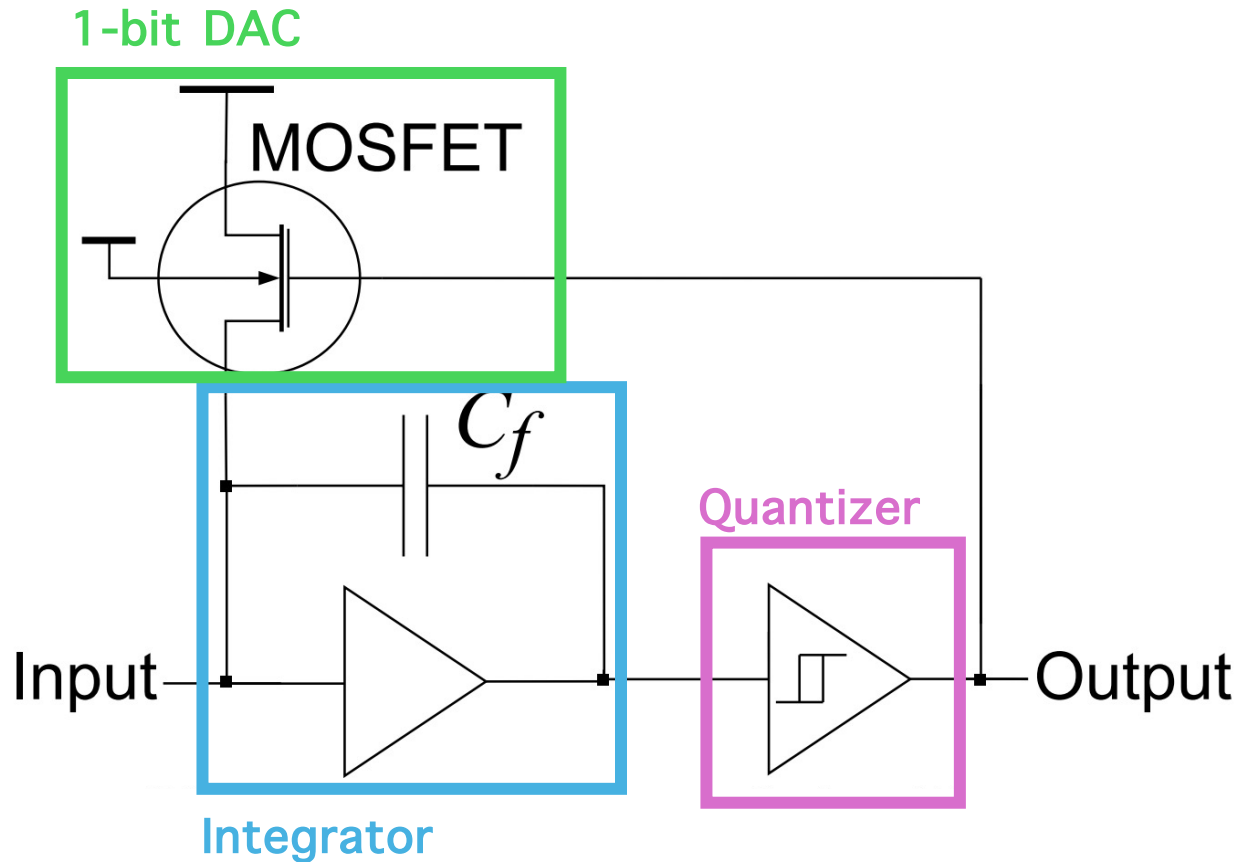
Sums the input signal over time, effectively smoothing it and making the system more tolerant to noise.

Quantizer

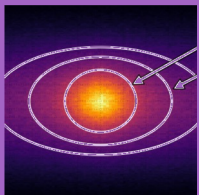
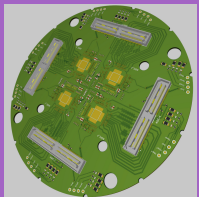
Converts the analog signal from the integrator into a digital signal. It is a single-bit quantizer with output of either 0 or 1.

DAC

Receives the output from the quantizer and converts the digital output back into an analog signal. It limits the amount of charge on C_f .



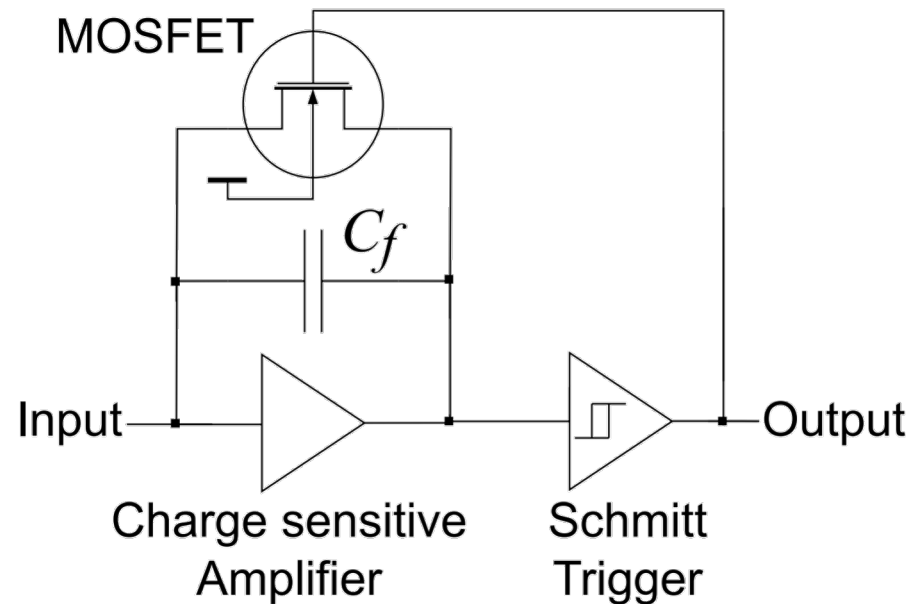
[Paper published on this](#)



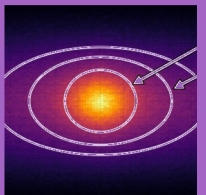
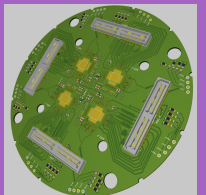
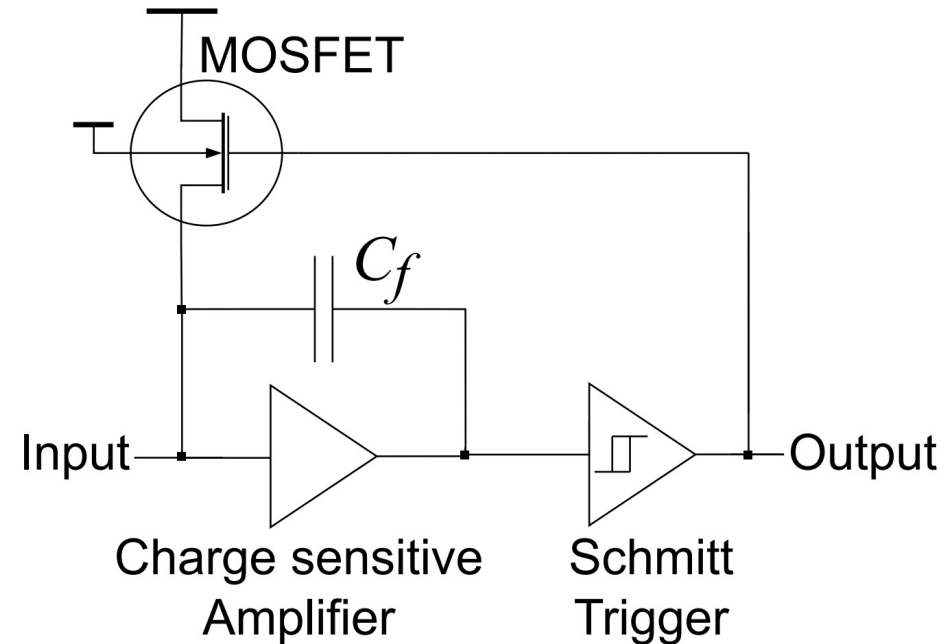


Reset vs Replenishment scheme

Reset



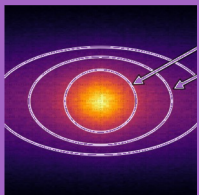
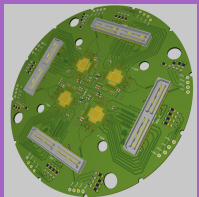
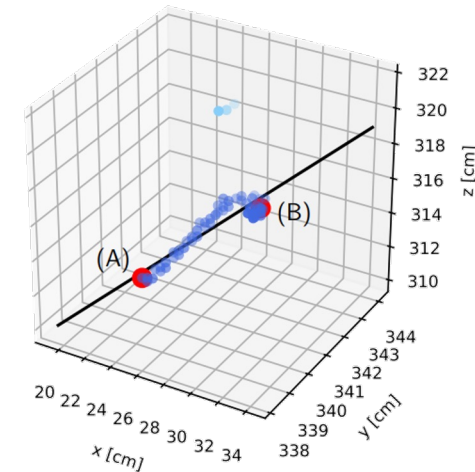
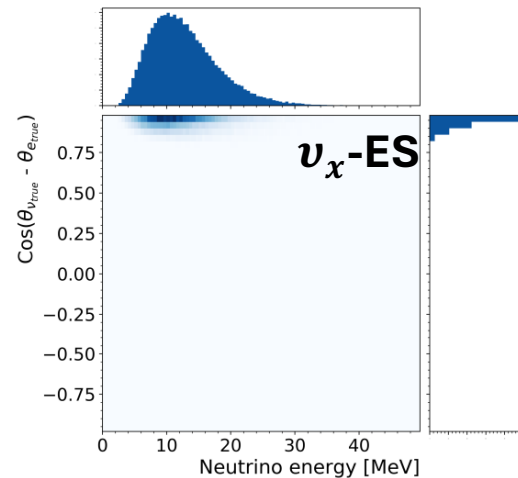
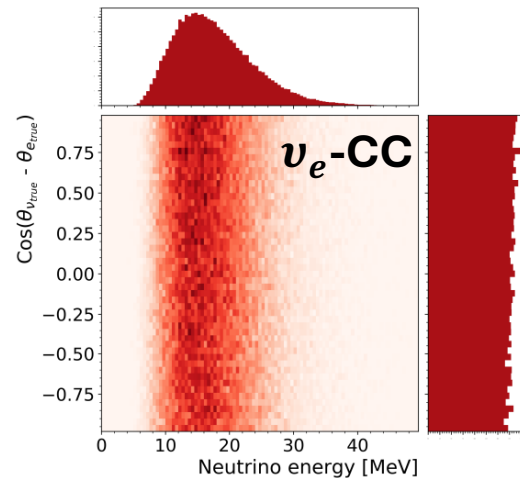
Replenishment





Supernova Pointing

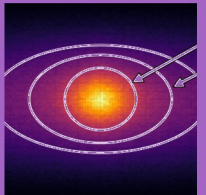
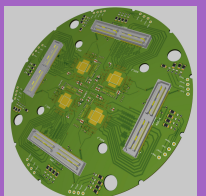
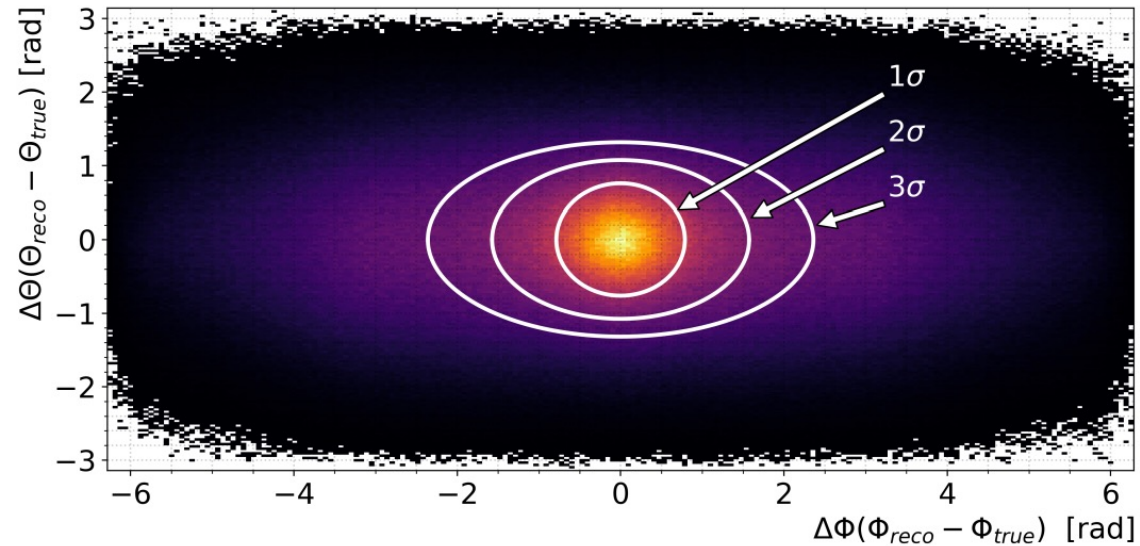
- The intrinsic 3D nature of the pixel data collected by Q-Pix allows us to get directional information from the identified supernova events.
- ~10% of all the events collected are neutrino-electron elastic scattering events (ν_x ES) and the rest are neutrino-charged current (ν_e -CC).
- ν_x -ES events preserve information about the direction of the neutrino
 - The direction of the neutrino tells us where in the sky the supernova burst occurs
 - This is a critical aspect of the identification of a SNB event for astronomers and particle physicists!





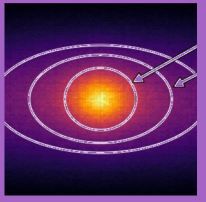
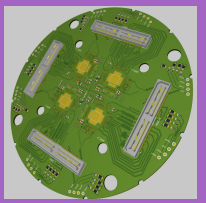
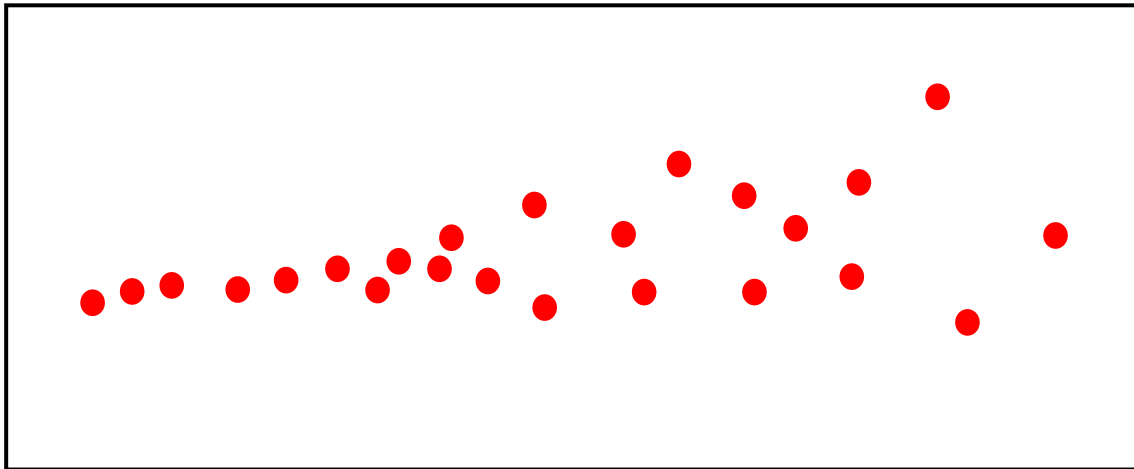
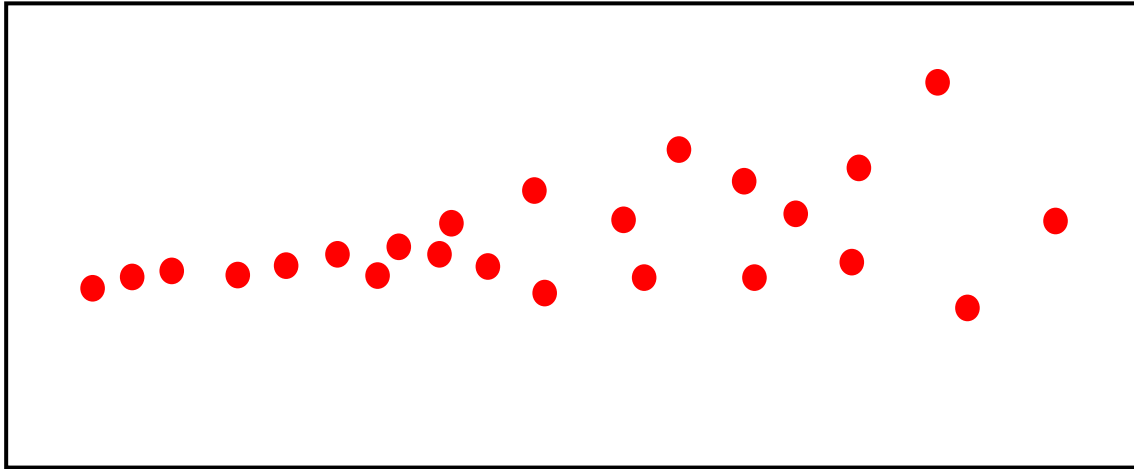
Supernova Pointing

- By reconstructing a direction vector for each neutrino interaction, we can come up with a hypothesis of where the SBN event occurred in the sky
 - We correctly identify the SBN direction within 20 degrees 80% of the time
 - The other 20% we have the direction wrong by 180 degrees
- Repeating this over 10,000 unique SBN events, we computed how confident we are with the direction with a 10 kTon Q-Pix module
 - 10 kpc supernova would be reconstructed within $\theta = 33^\circ$ and $\phi = 45^\circ$ at 1σ , and $\theta = 99^\circ$ and $\phi = 135^\circ$ at 3σ .



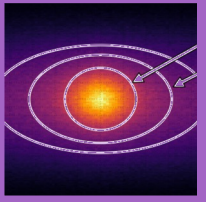
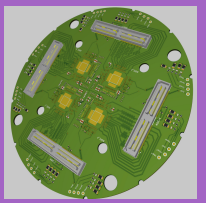
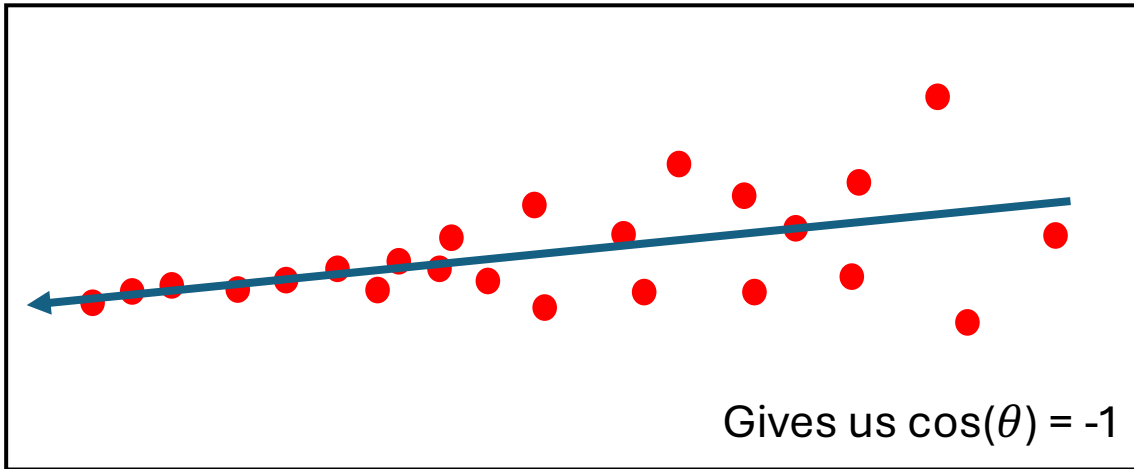
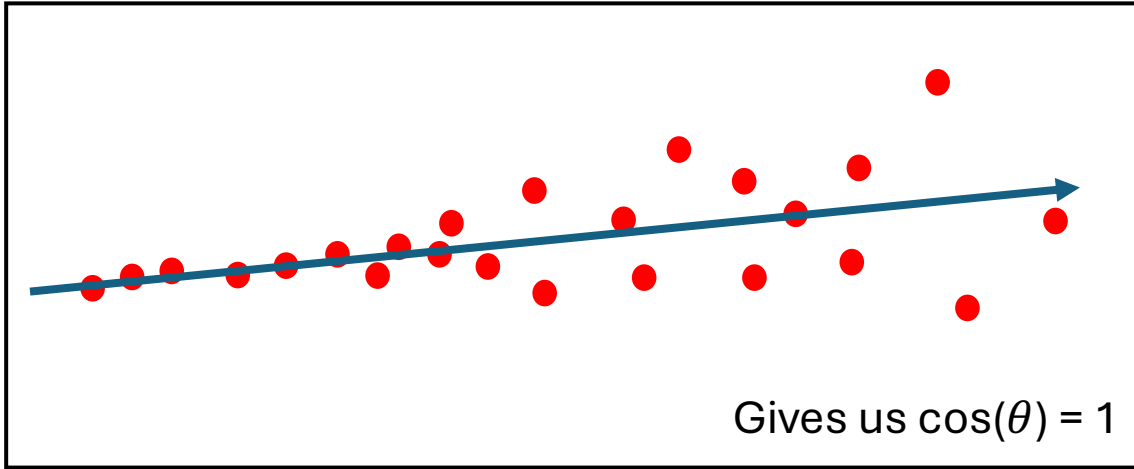


Directionality Flipping



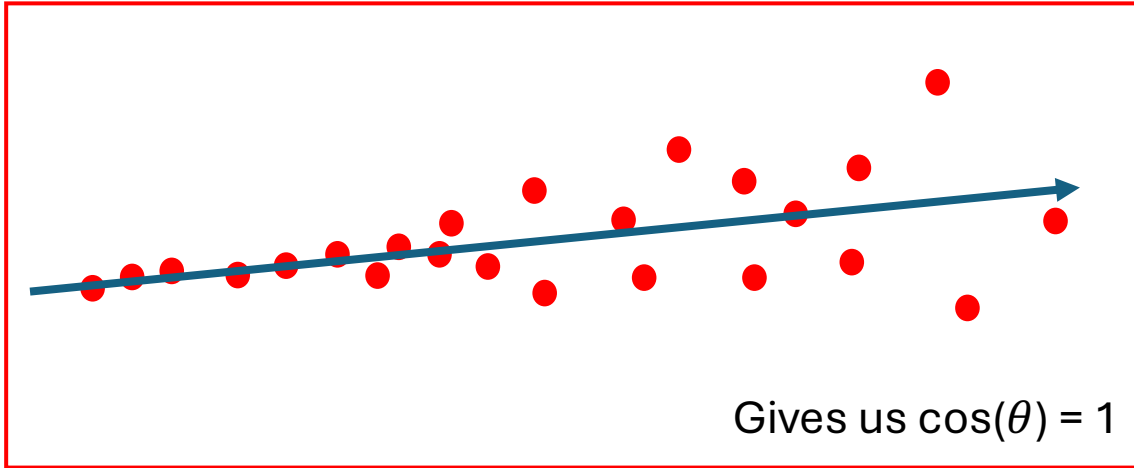


Directionality Flipping



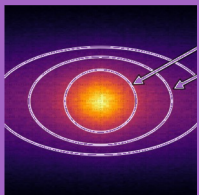
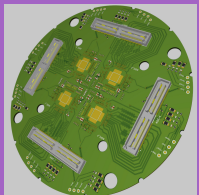
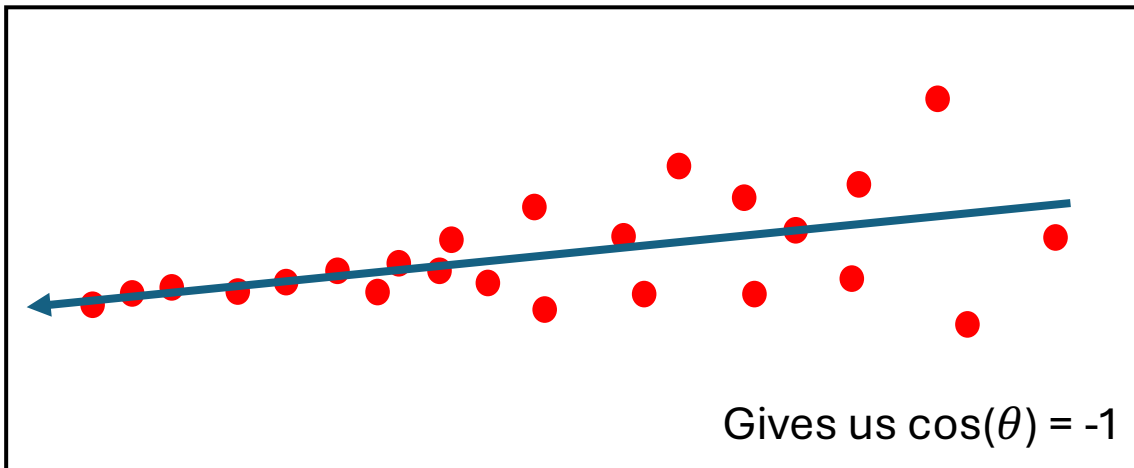


Directionality Flipping



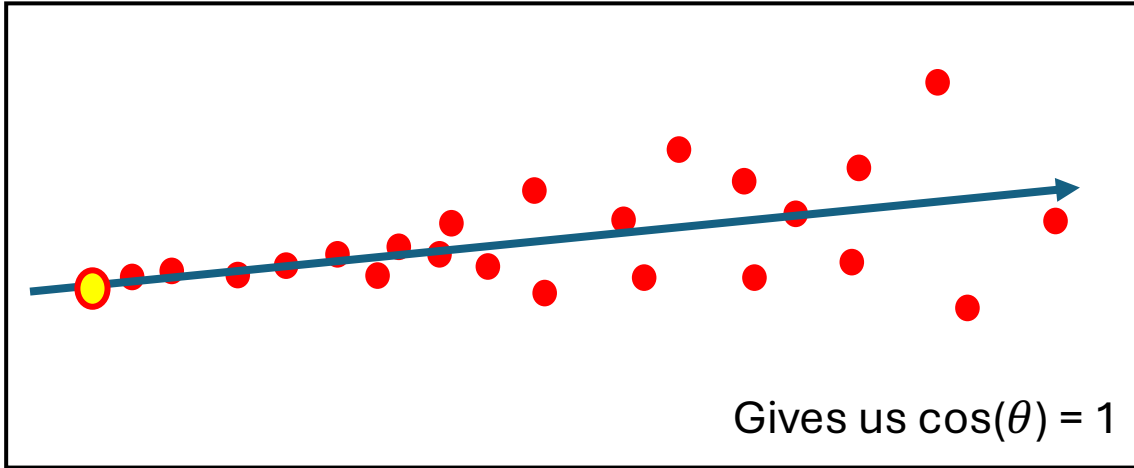
Electron scatters more often as it loses energy.

More scattered = less energy = end of the track
Less scattered = more energy = beginning of the track

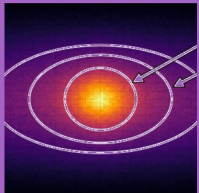
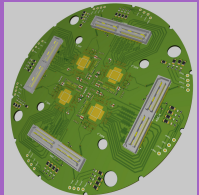
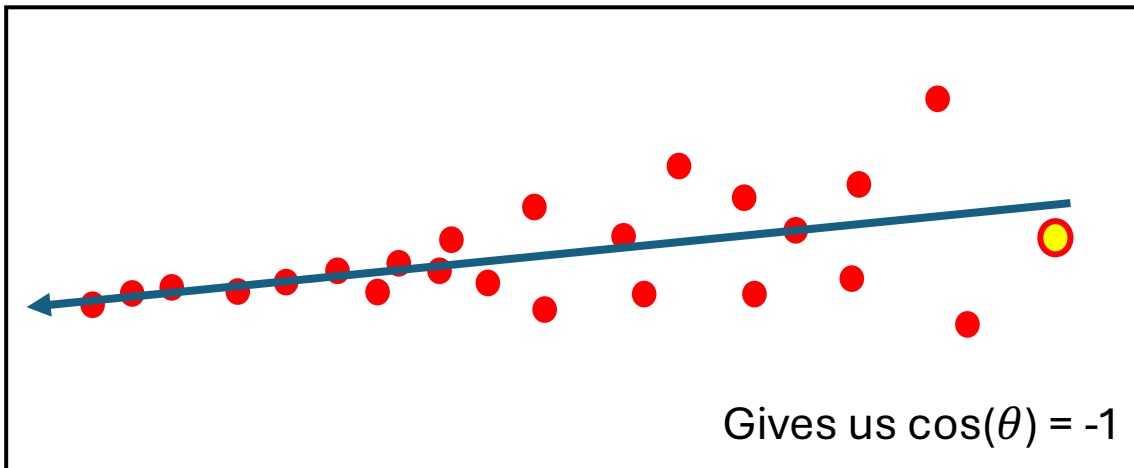




Directionality Flipping

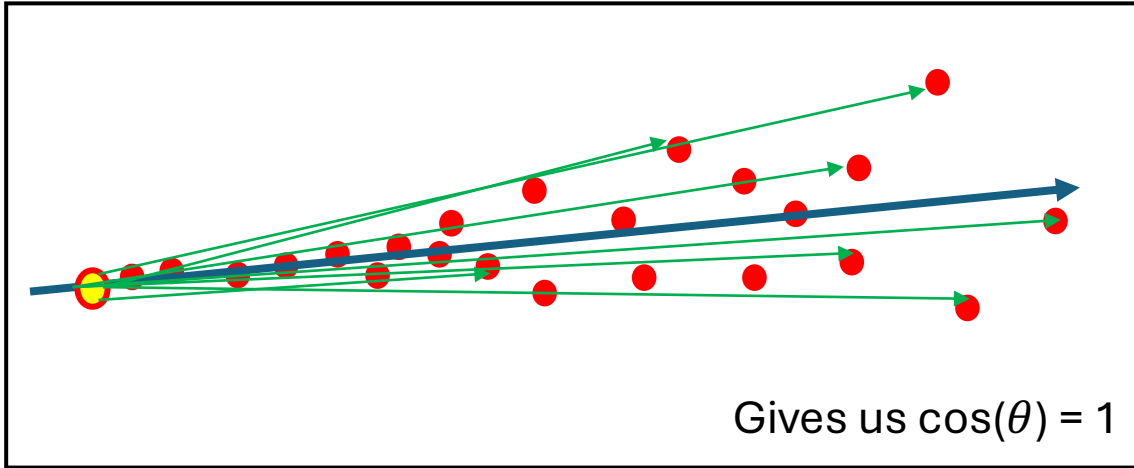


Pick data points at both ends



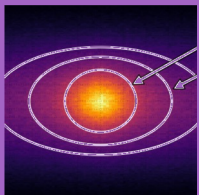
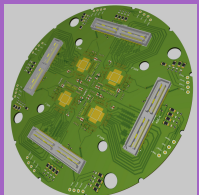
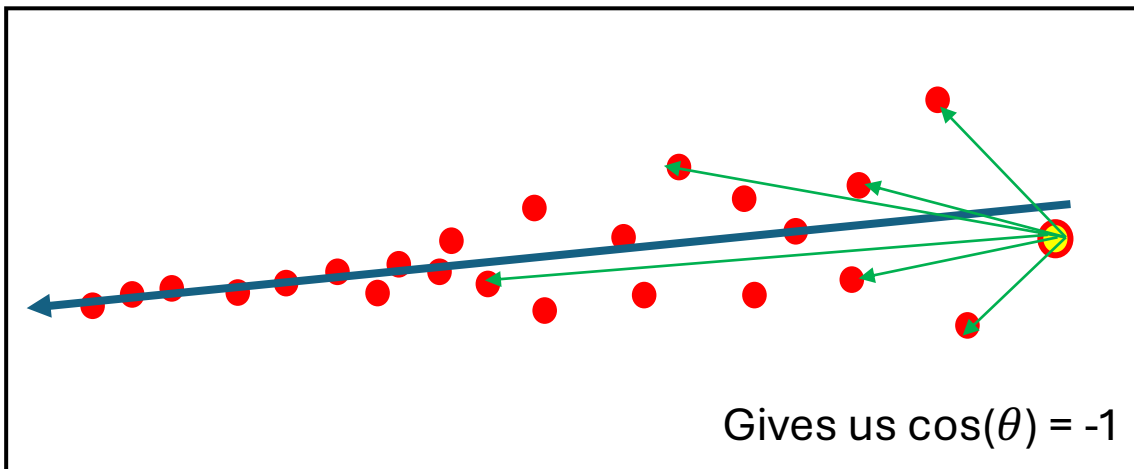


Directionality Flipping



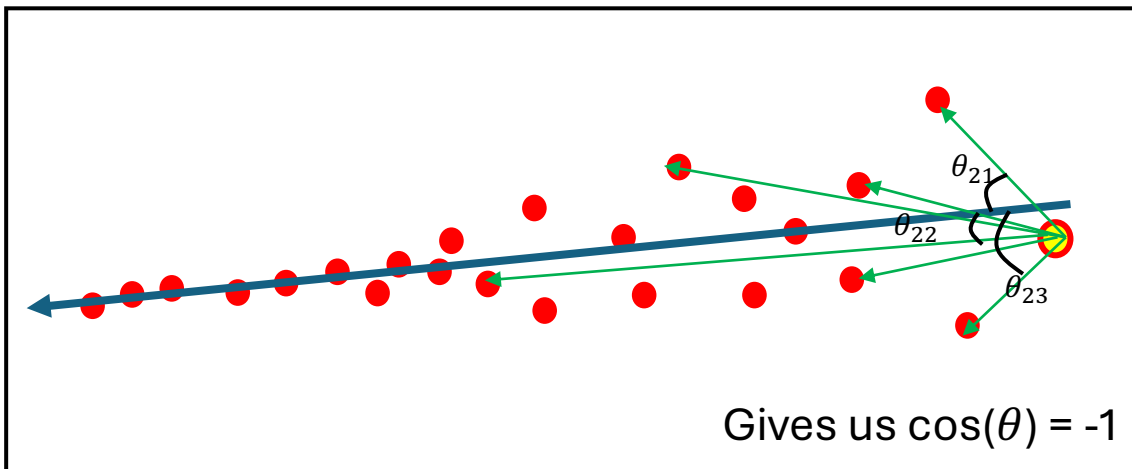
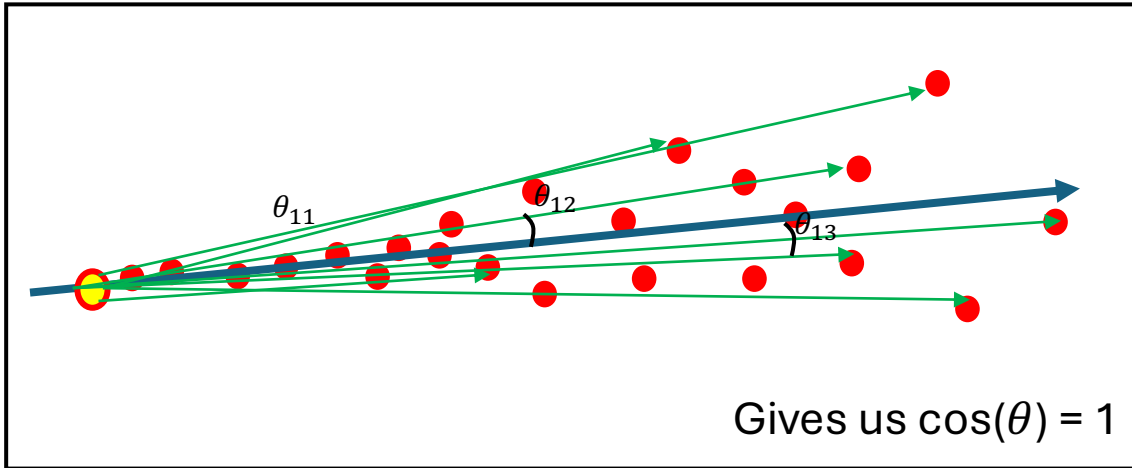
Pick data points at both ends

Draw line from two ends to every single data point





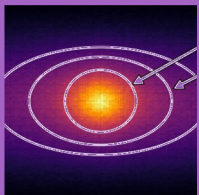
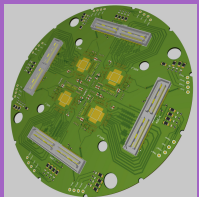
Directionality Flipping



Pick data points at both ends

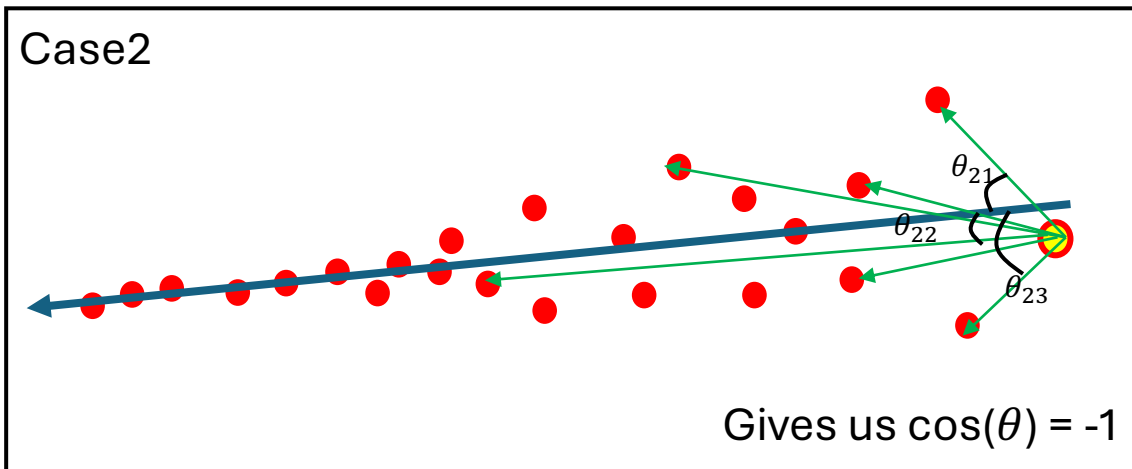
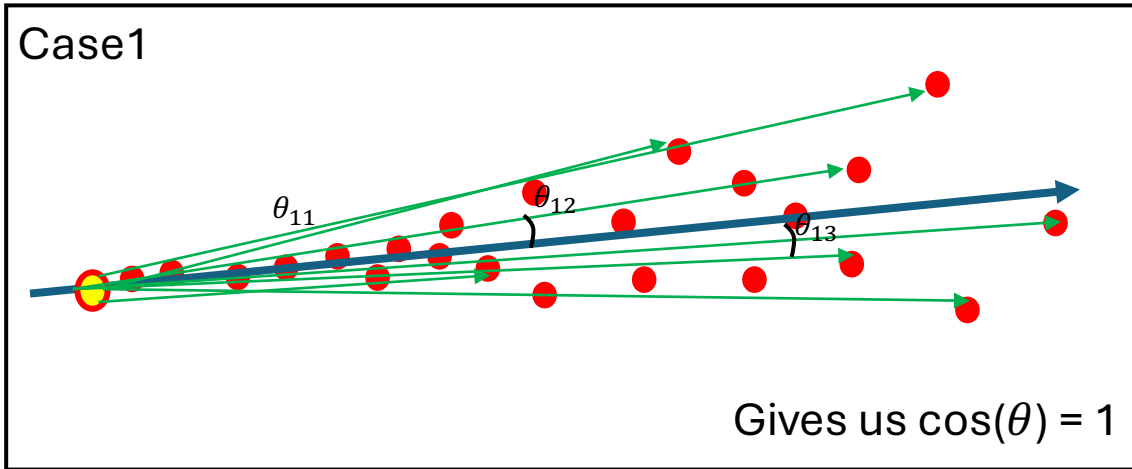
Draw line from two ends to every single data point

Calculate the cosine of angle between the lines and computed axis, and sum them





Directionality Flipping



Pick data points at both ends

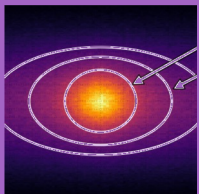
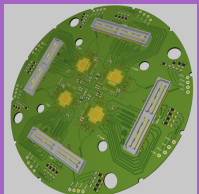
Draw line from two ends to every single data point

Calculate the cosine of angle between the lines and computed axis, and sum them

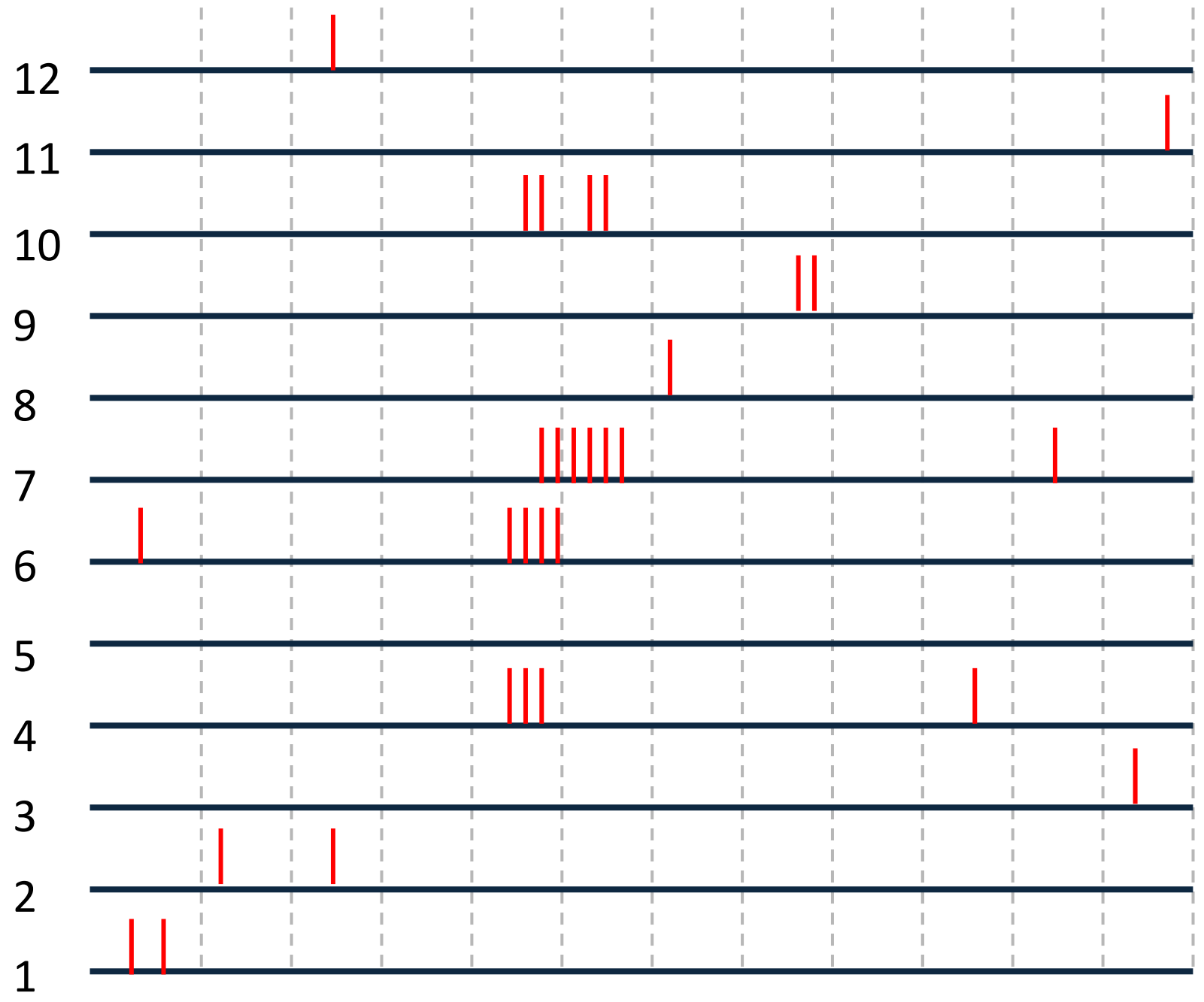
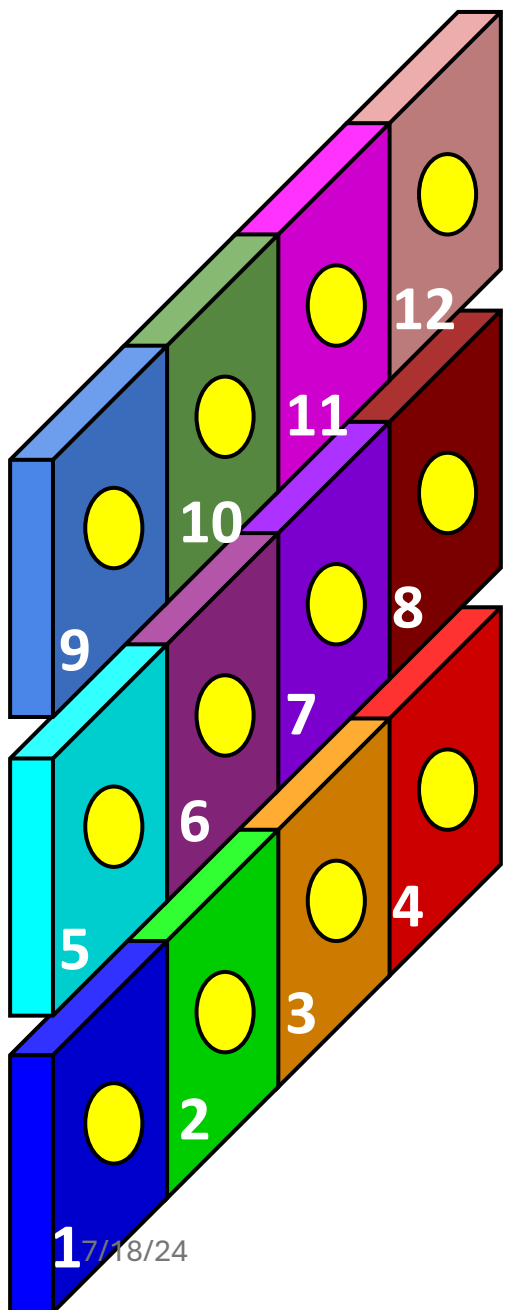
θ values are smaller \rightarrow $\cos(\theta)$ values are bigger
 $\text{sum_cos_case1} = \cos(\theta_{11}) + \cos(\theta_{12}) + \cos(\theta_{13}) \dots$

θ values are bigger \rightarrow $\cos(\theta)$ values are smaller
 $\text{sum_cos_case2} = \cos(\theta_{21}) + \cos(\theta_{22}) + \cos(\theta_{23}) \dots$

$\text{sum_cos_case1} > \text{sum_cos_case2}$

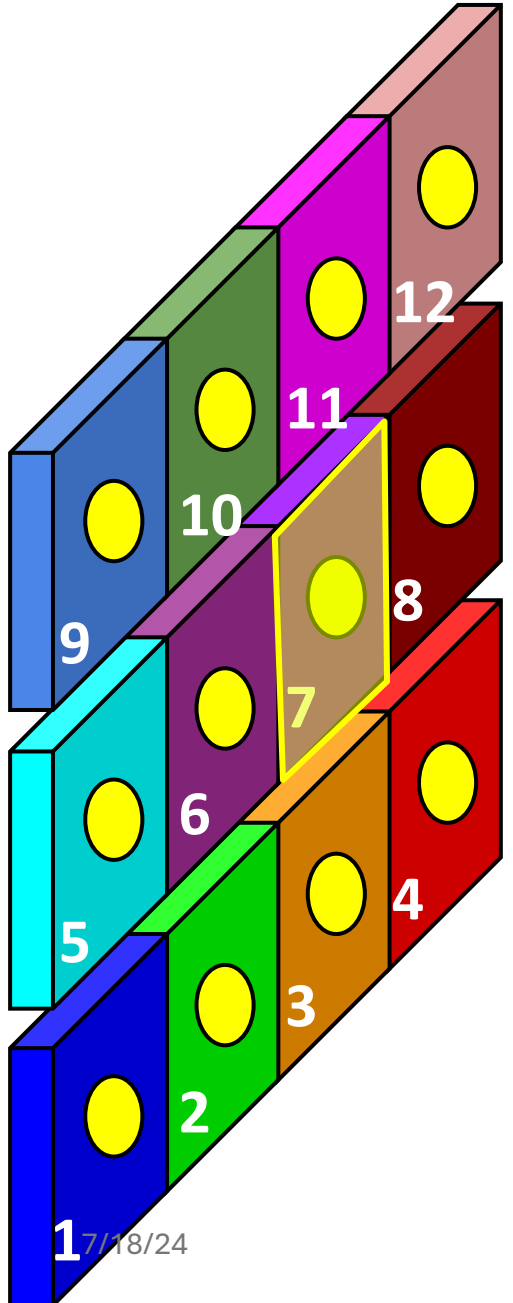


| = 1 RTD



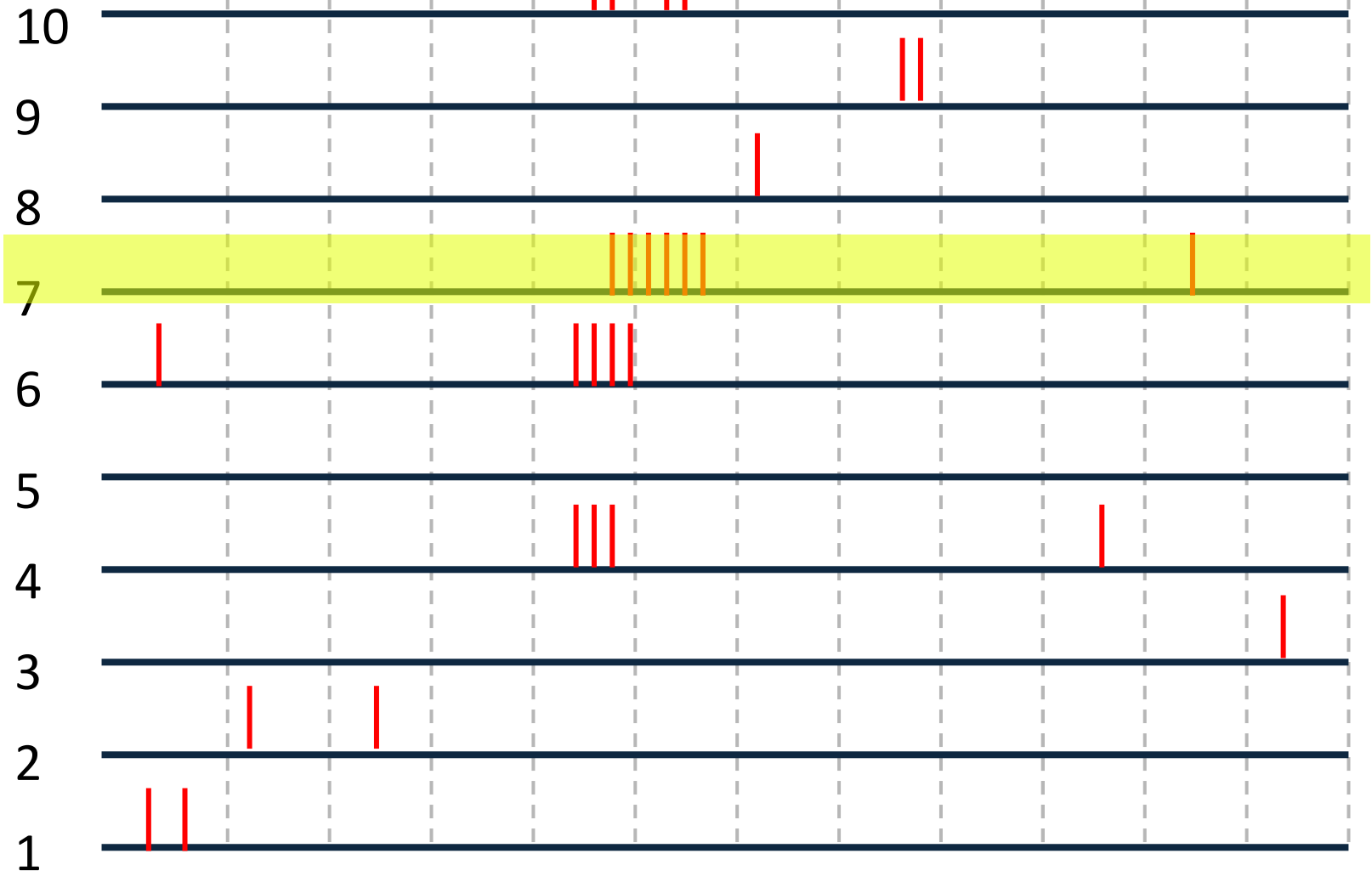
time

| = 1 RTD

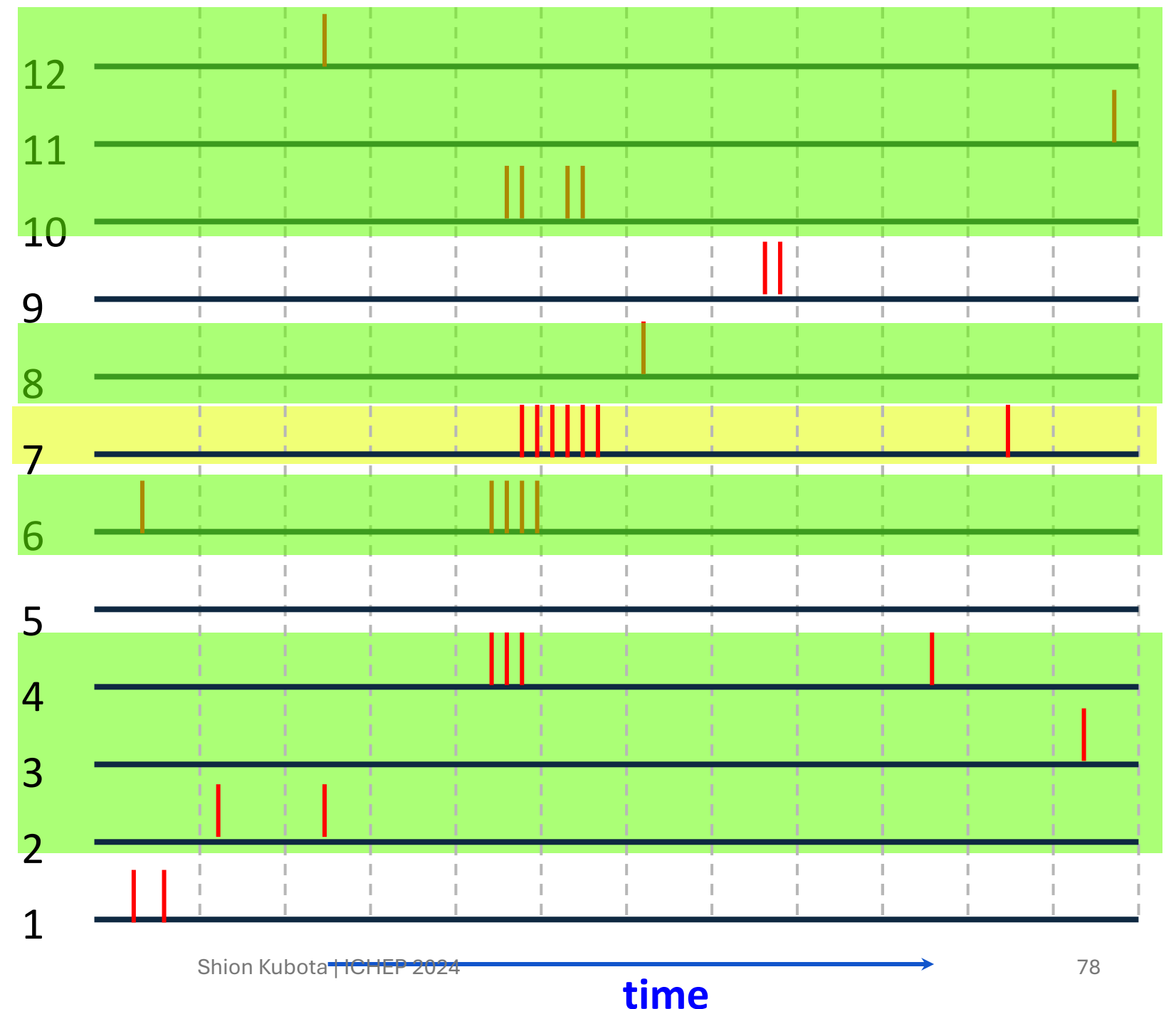
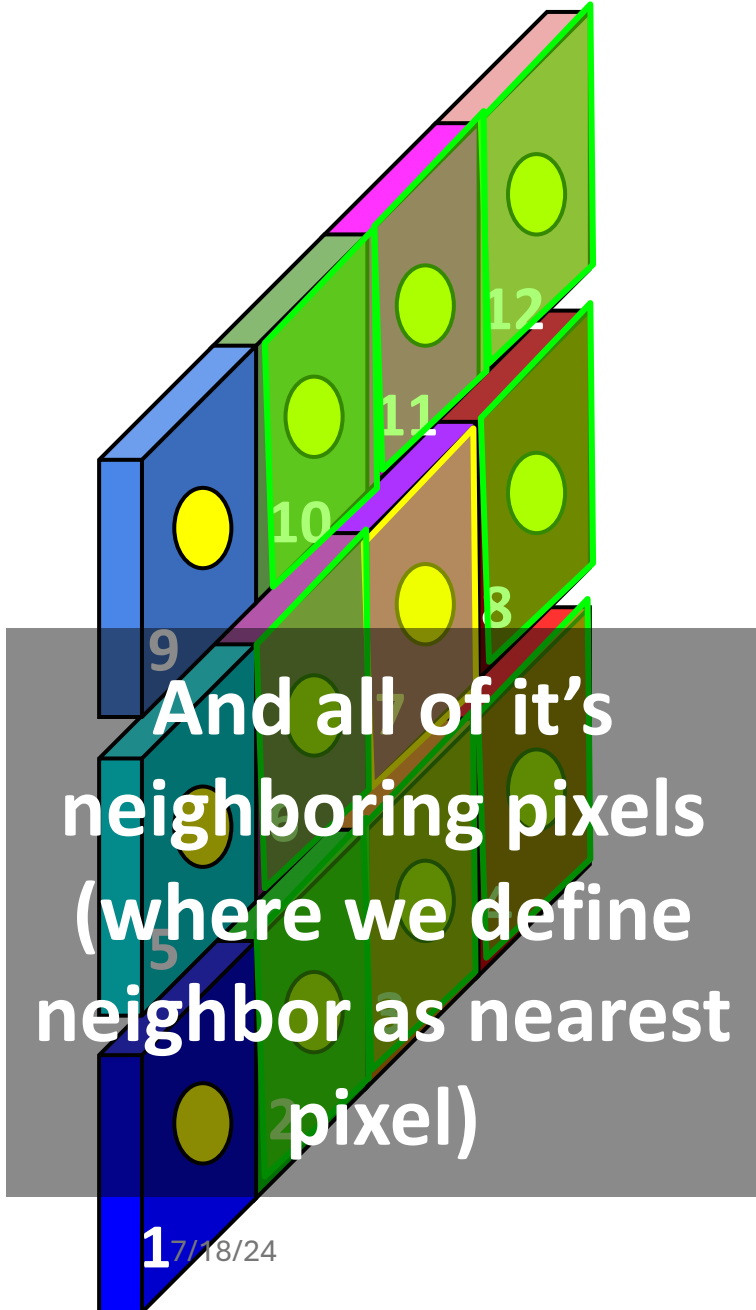


1 7/18/24

For this example, we start by analyzing the RTD's on a particular pixel

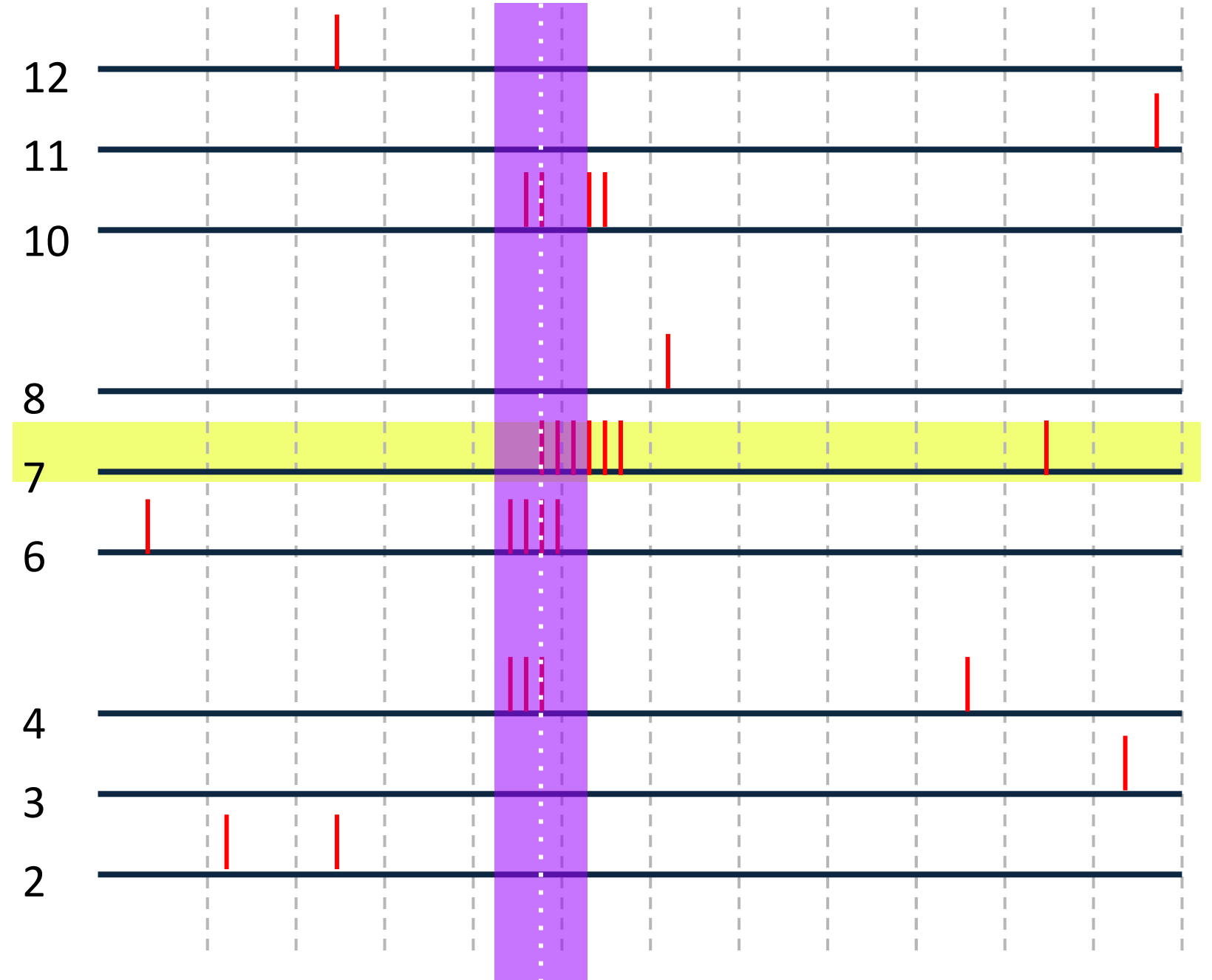
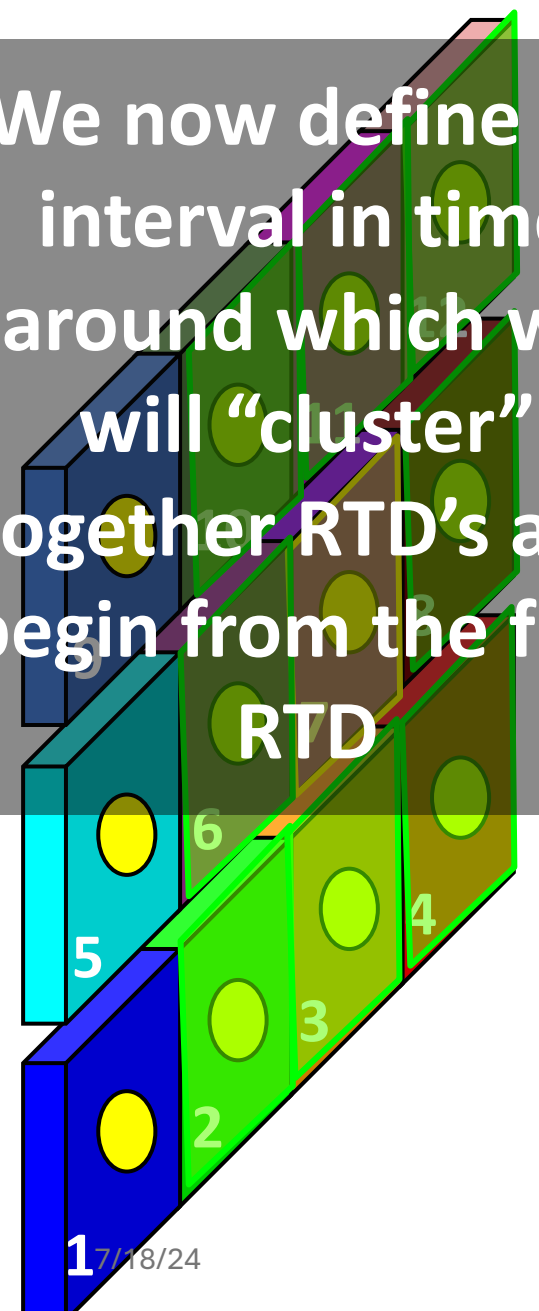


| = 1 RTD

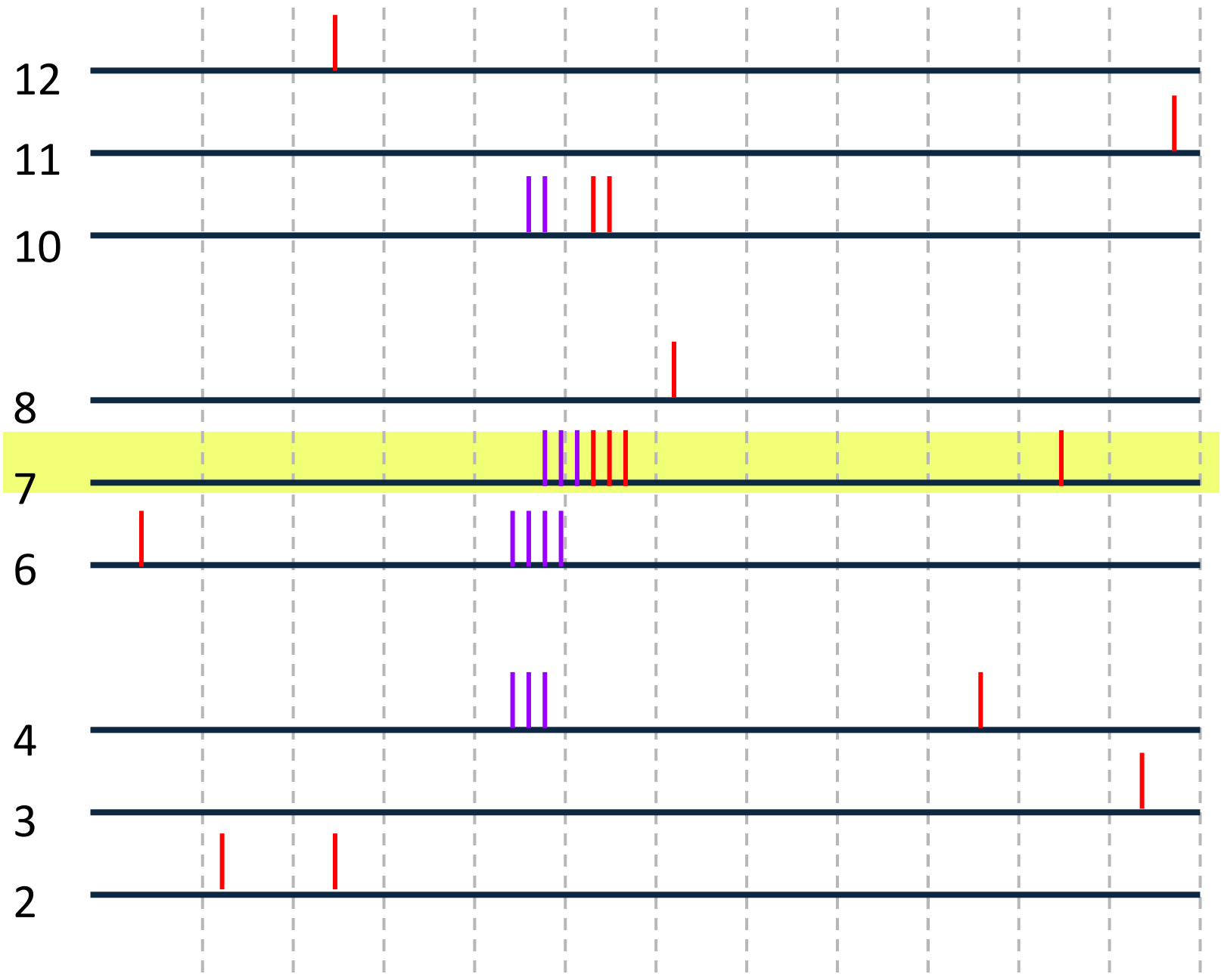
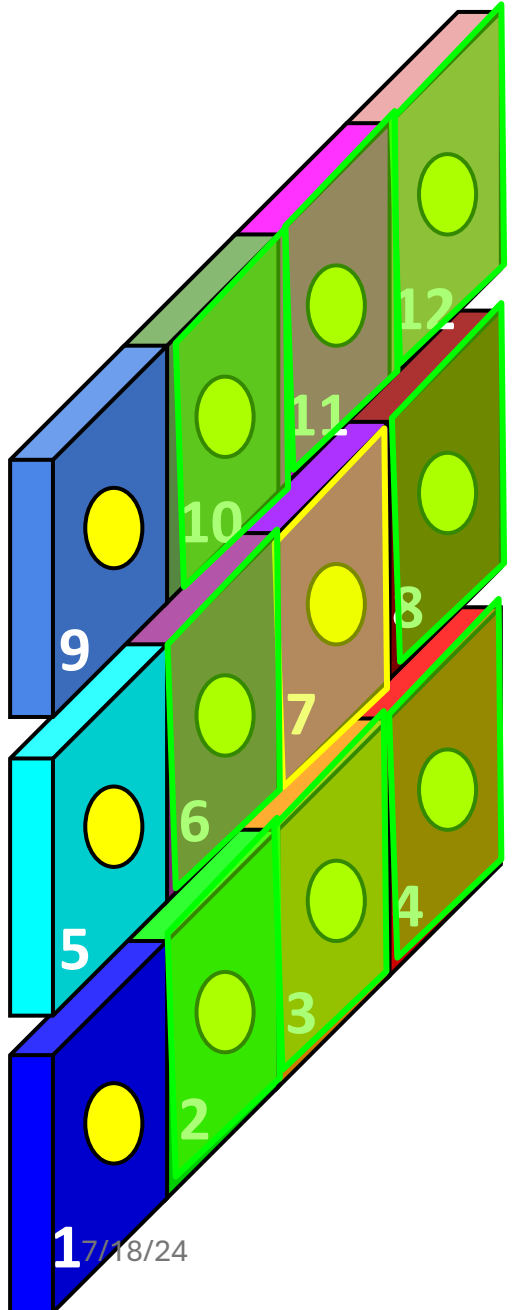


| = 1 RTD

We now define an interval in time around which we will "cluster" together RTD's and begin from the first RTD



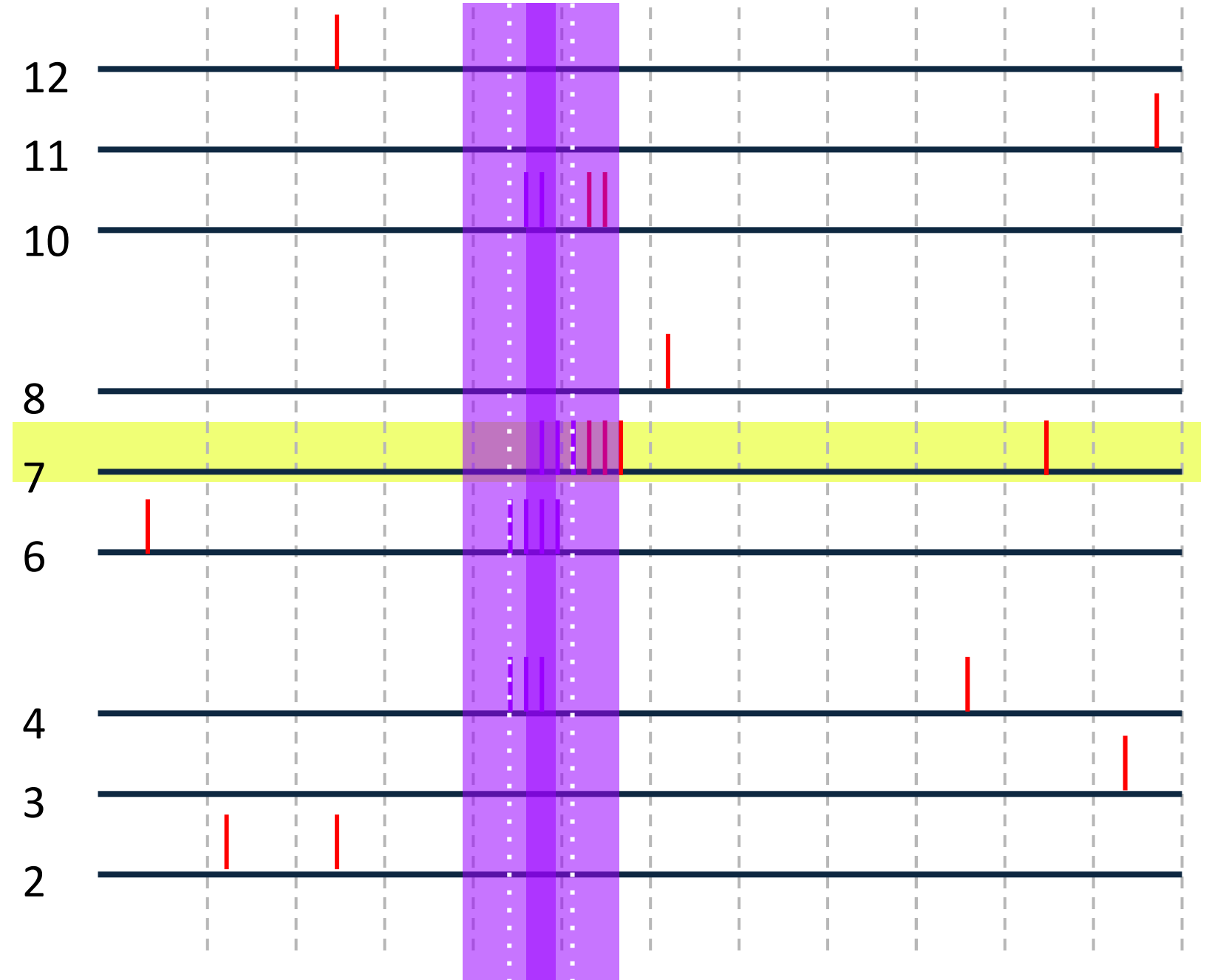
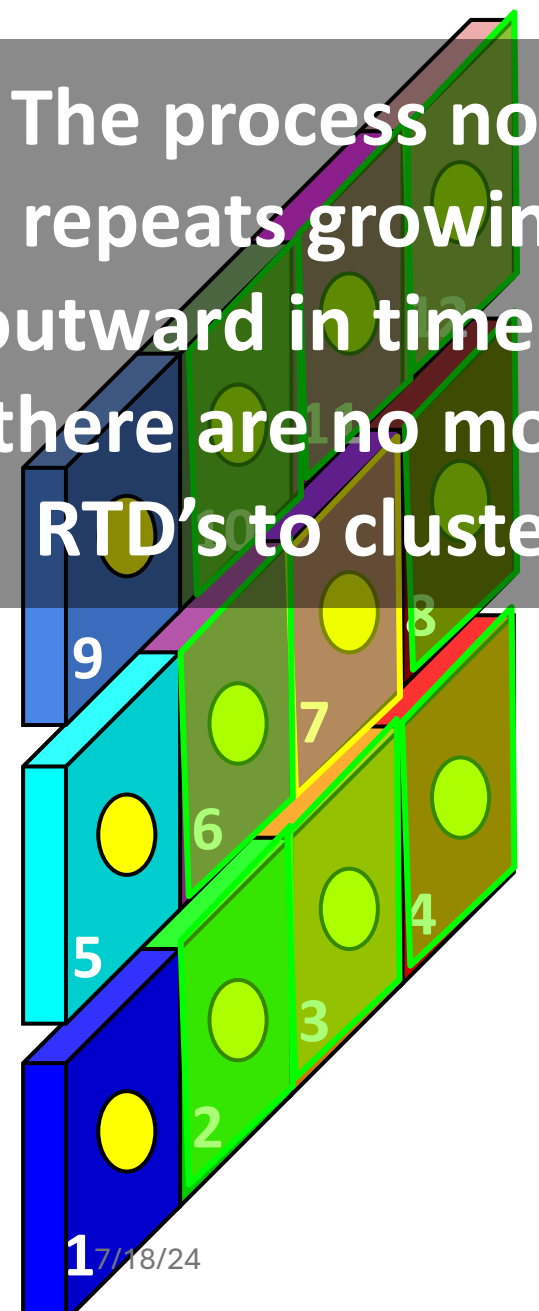
| = 1 RTD



time

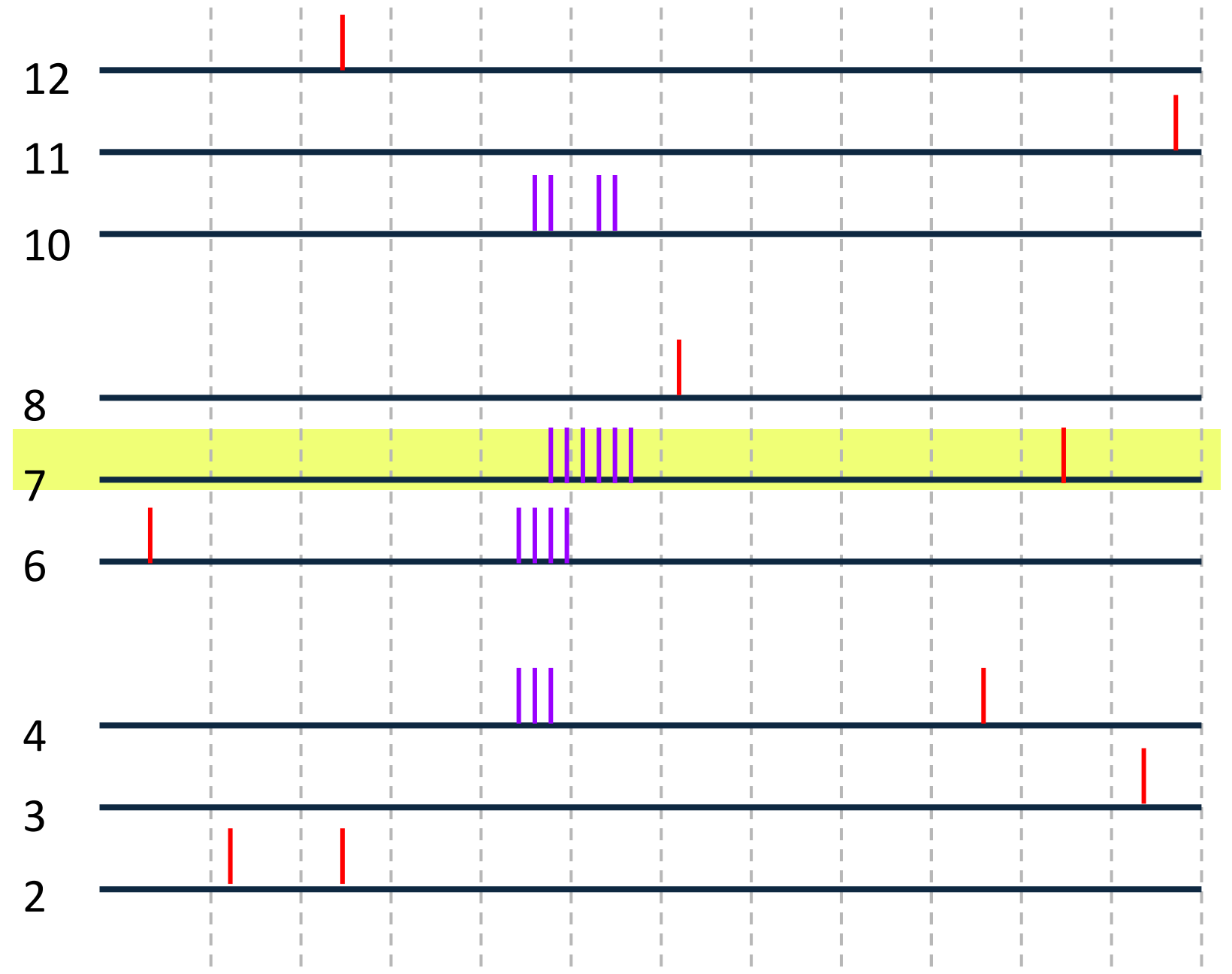
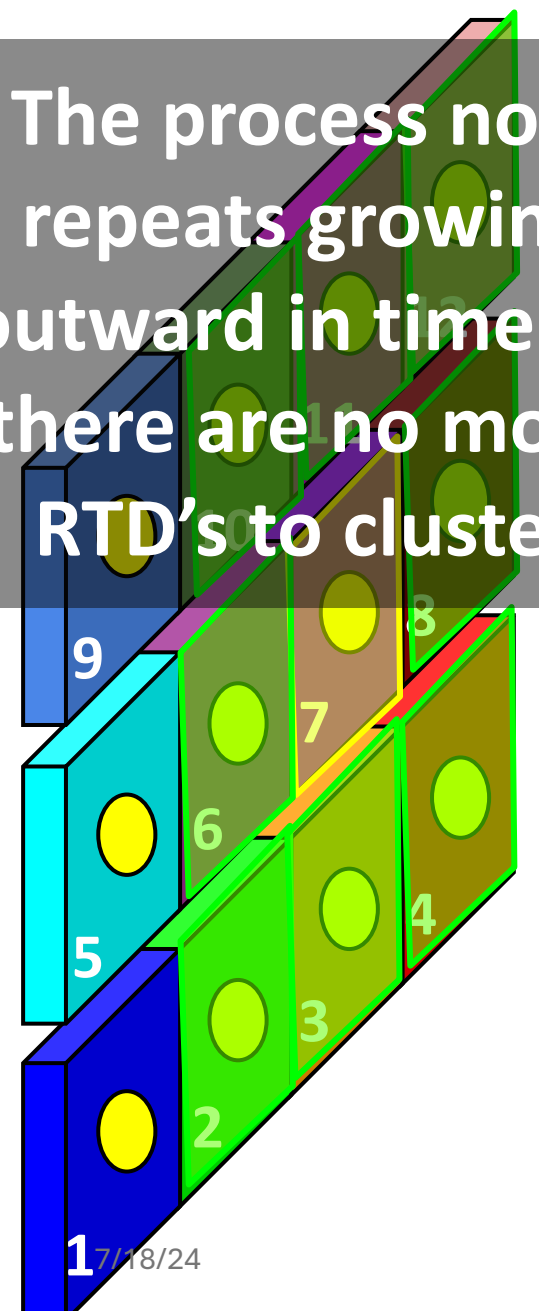
| = 1 RTD

The process now repeats growing outward in time till there are no more RTD's to cluster



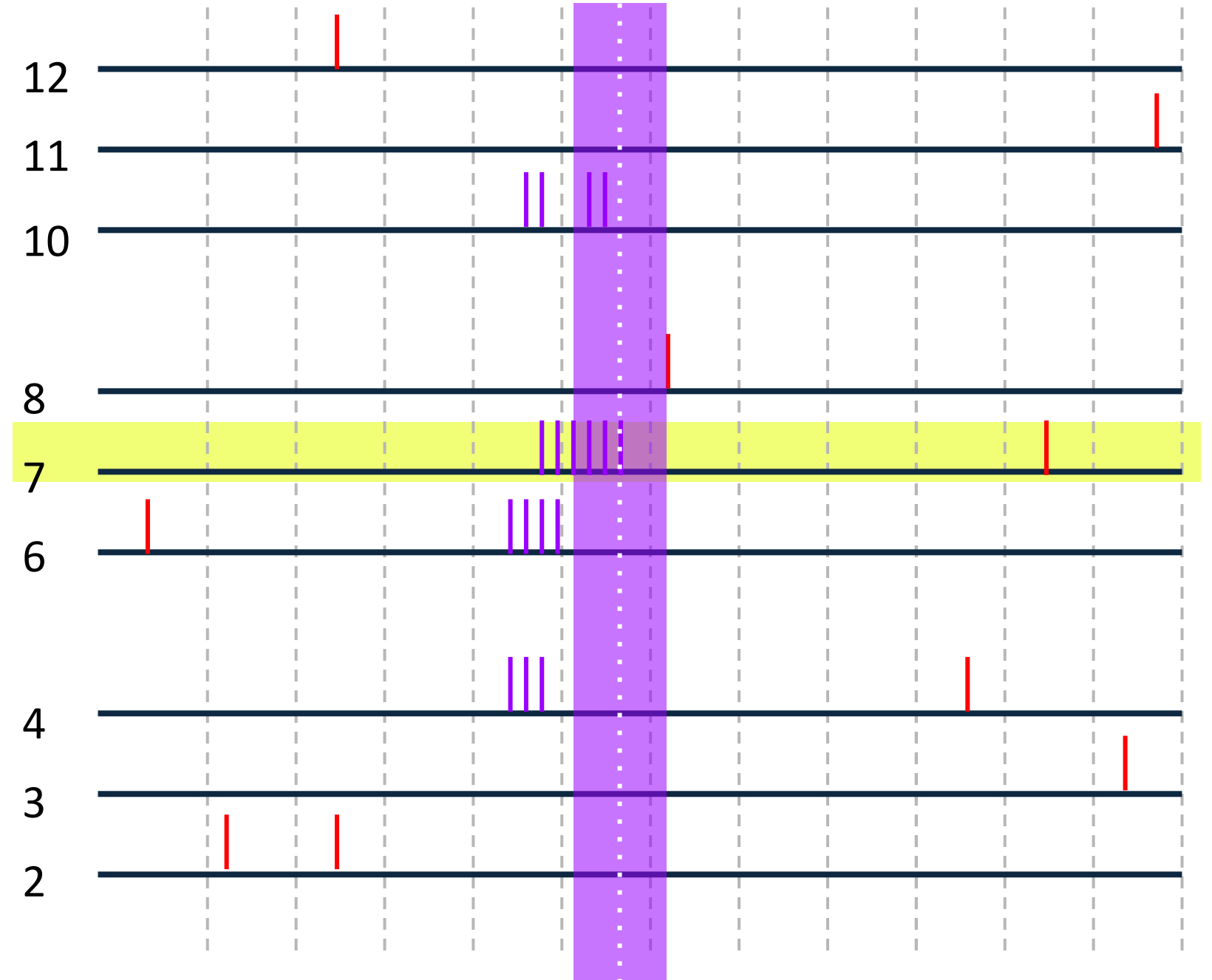
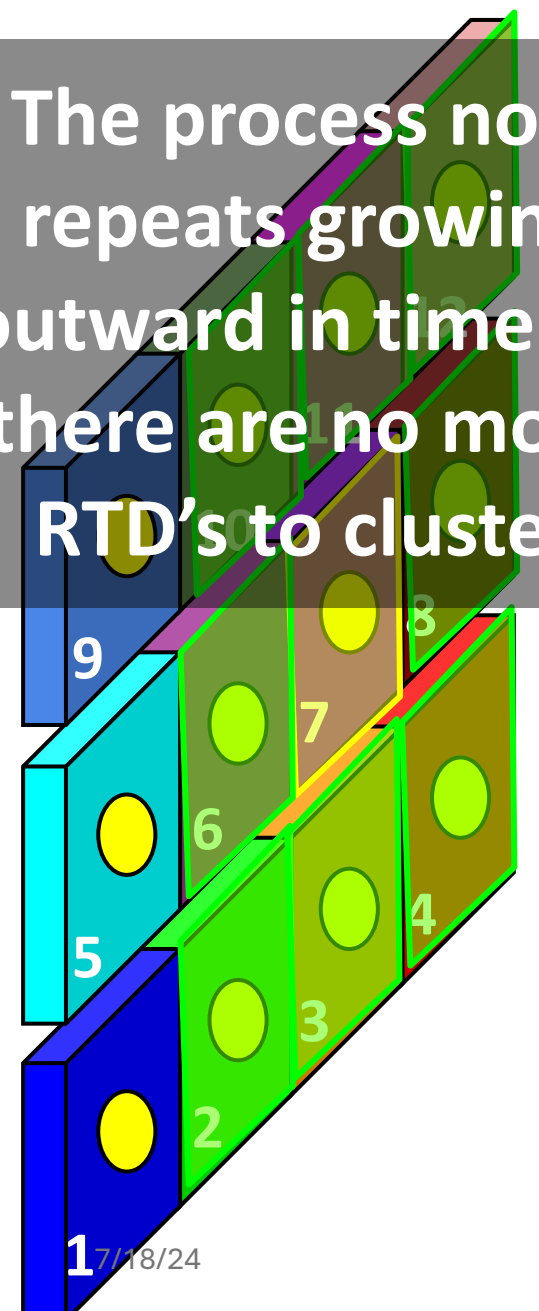
| = 1 RTD

The process now repeats growing outward in time till there are no more RTD's to cluster

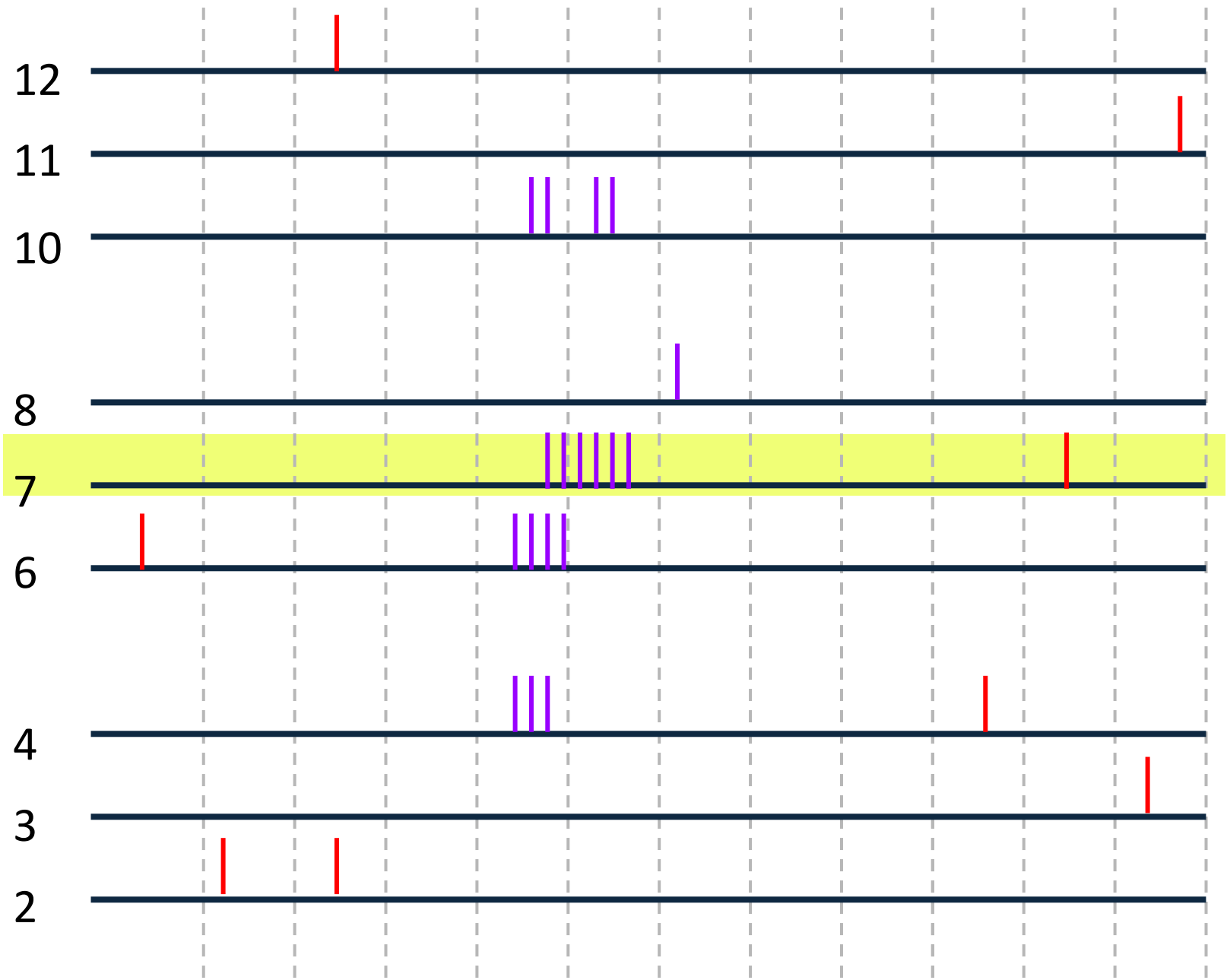
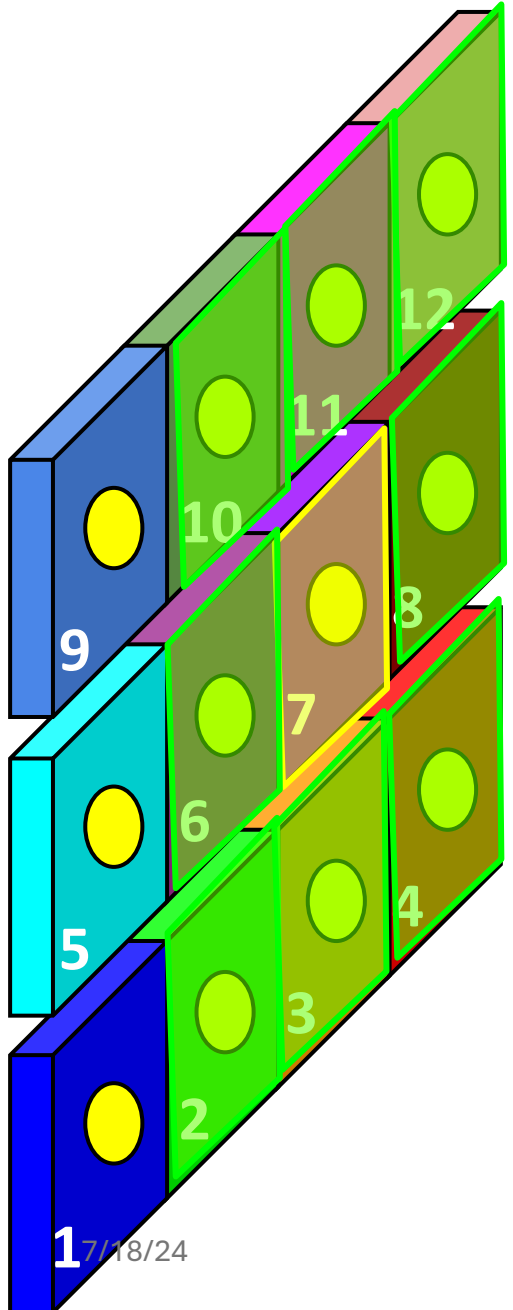


| = 1 RTD

The process now repeats growing outward in time till there are no more RTD's to cluster

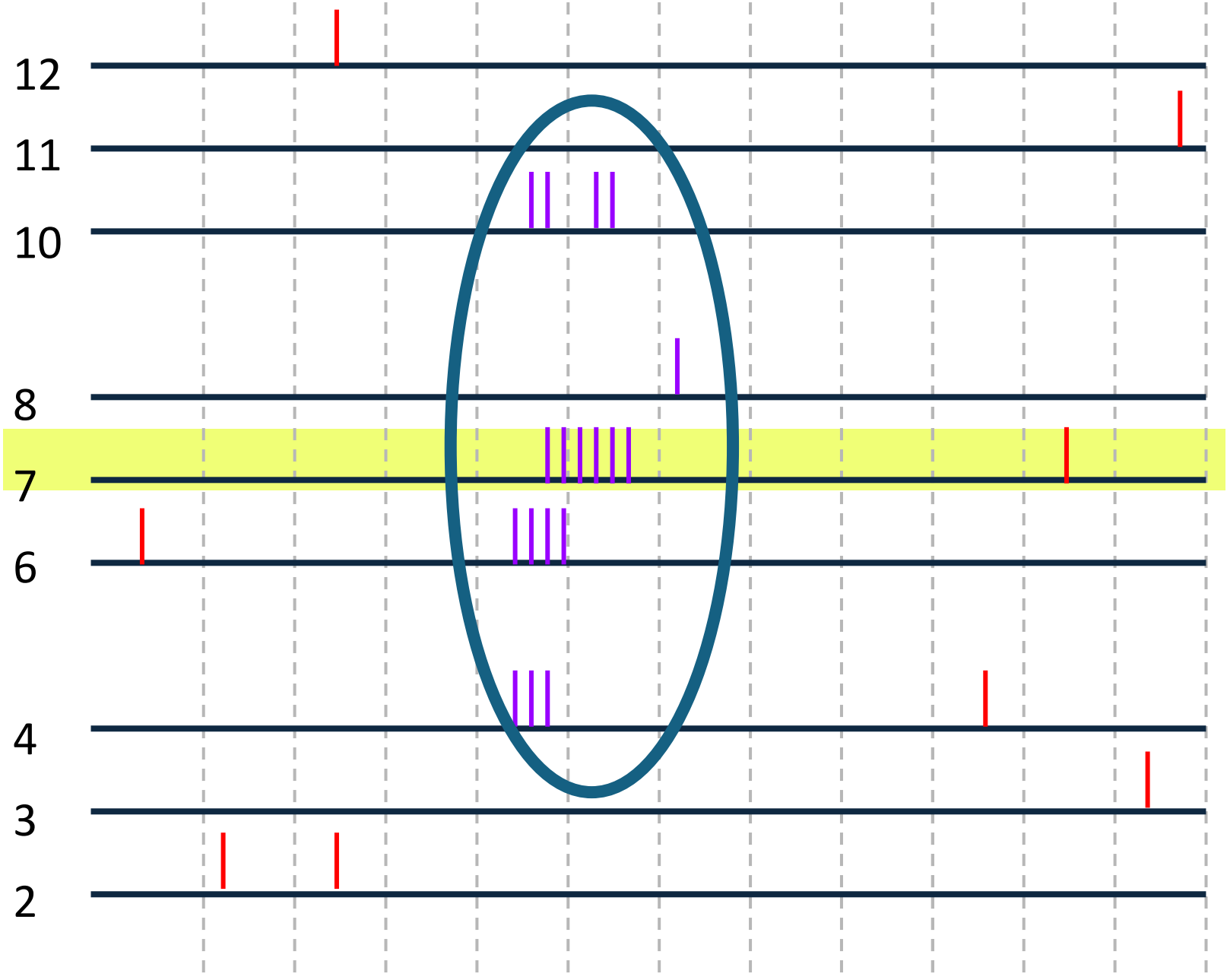
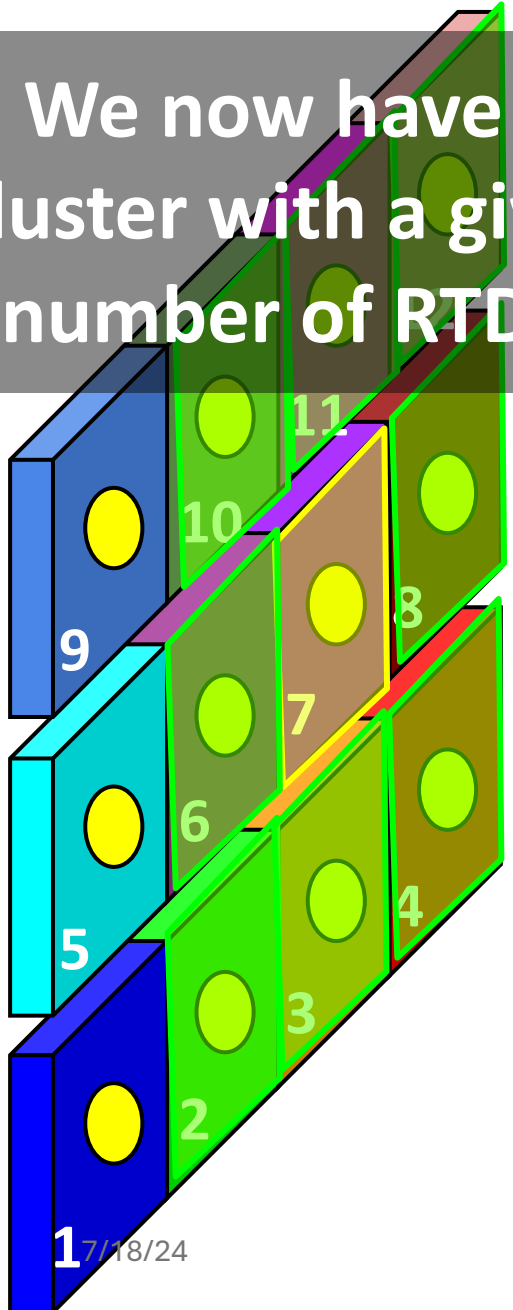


| = 1 RTD



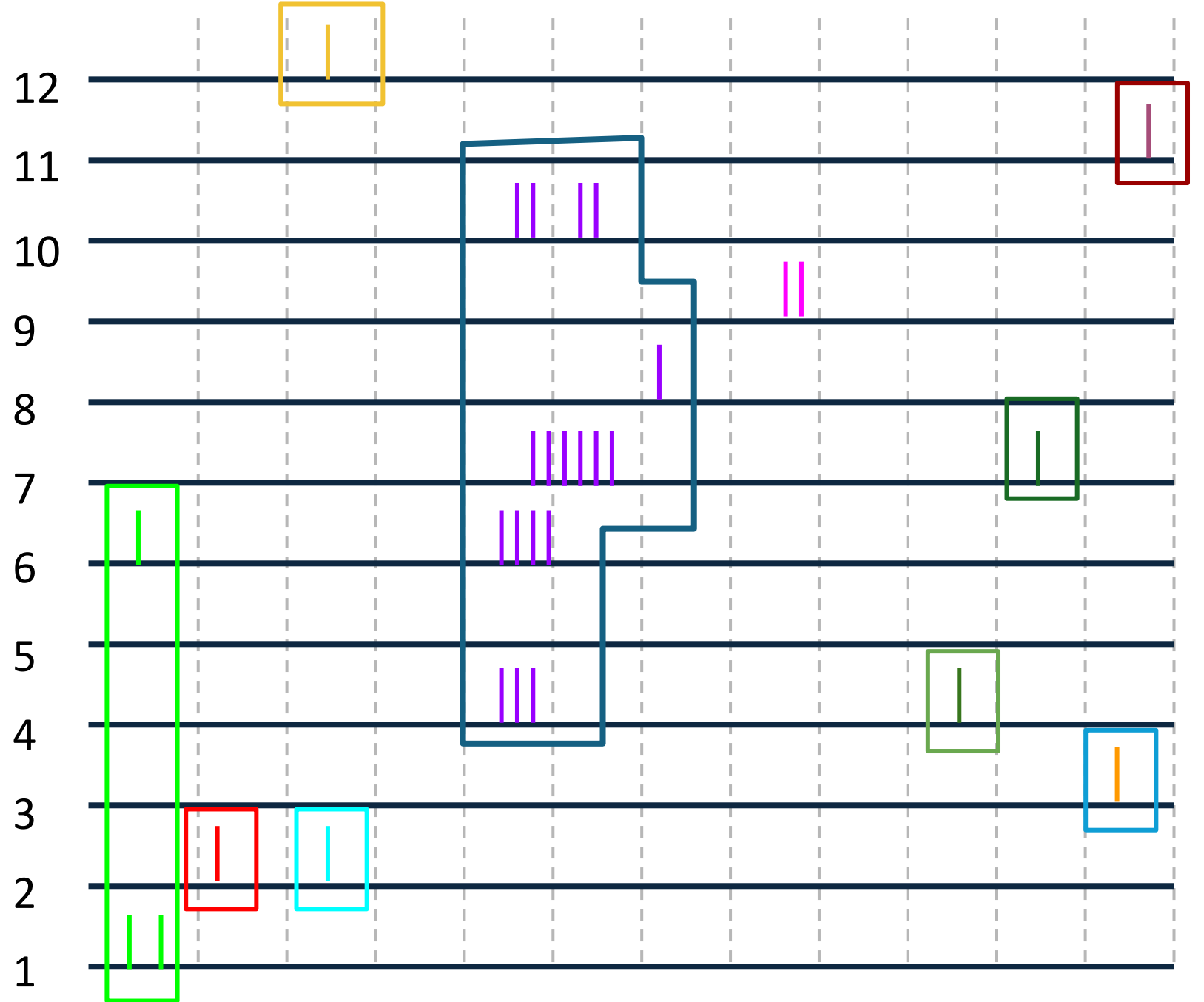
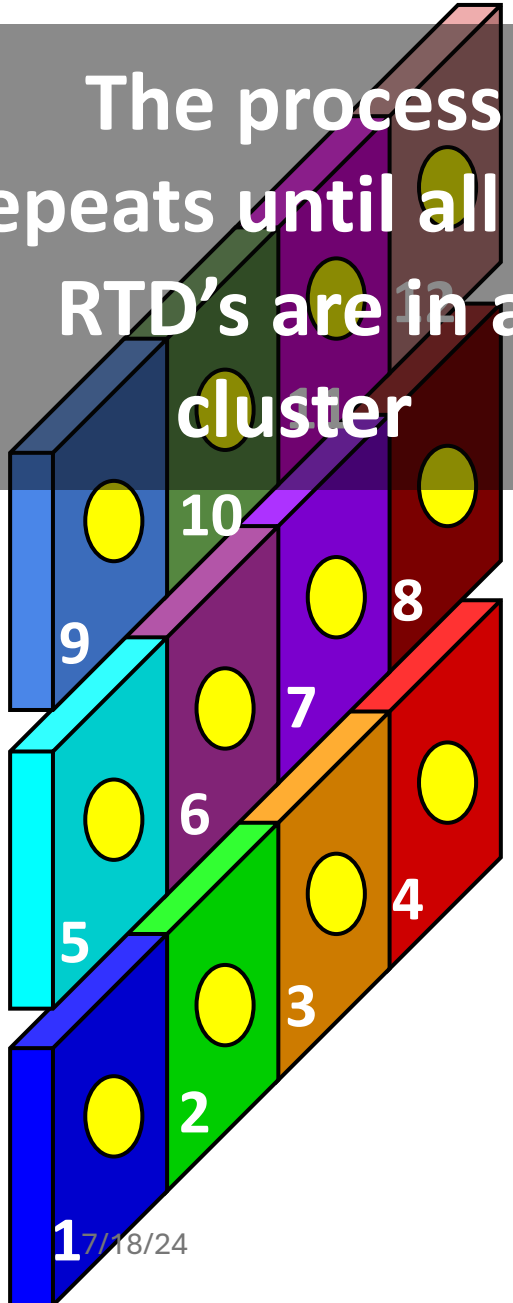
| = 1 RTD

We now have a cluster with a given number of RTD's



| = 1 RTD

The process repeats until all the RTD's are in a cluster





Solar neutrino chains

