

LFV muon decays with Neutrino NSI

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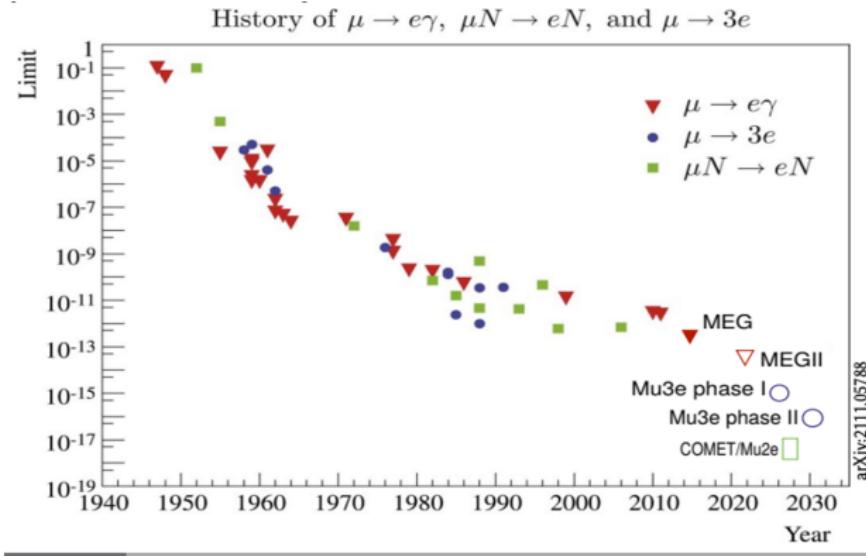


मार्गीय प्रौद्योगिकी संस्थान हैदराबाद
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LFV muon decays



- Precision and BSM physics
- Neutrino NSI and muon LFV

cLFV

- Neutrino Oscillation indicate LFV
- LFV in charged lepton not observed yet
- In the SM, $\text{Br}(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} |\sum_i U_{\beta i}^* U_{\alpha i} \frac{m_{\nu_i}^2}{M_W^2}|^2 \leq 10^{-54}$
- But, $\text{Br}(b \rightarrow s\gamma) \approx (3.36 \pm 0.23) \times 10^{-4}$



- NP can enhance $\text{Br}(\mu \rightarrow e\gamma)$ by few orders
- cLFV are very clean probes - unambiguous sign of new physics
- Implications of neutrino NSI on muon LFV

Neutrinos and Open Questions

Neutrino: Key to New Physics

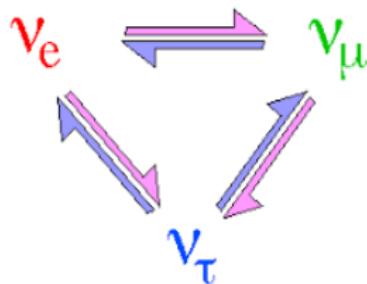
To decipher signature(s) of BSM Physics

Open Questions

- CP violation in lepton sector ?
- Majorana or Dirac ?
- Absolute mass of neutrinos?
- Mass ordering: sign of (Δm_{13}^2) ?
- $\theta_{23} > \pi/4, \theta_{23} < \pi/4, \theta_{23} = \pi/4$?
- Sterile neutrino(s) ?

Neutrino Oscillations

- Neutrino oscillations provide pathway to Physics beyond the standard model.



- Three neutrino flavor eigenstates (ν_e , ν_μ , ν_τ) are unitary linear combinations of three neutrinos mass eigenstates (ν_1 , ν_2 , ν_3) with masses m_1 , m_2 , $m_3 \rightarrow$ Neutrino mixing
- standard parameterization for PMNS matrix:

$$U_{PMNS} = U_{23}(\theta_{23}) U_{13}(\theta_{13}, \delta_{cp}) U_{12}(\theta_{12})$$

CP Violation

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Controls CP Violation

$$U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

- strength of CP violation is parameterized by the Jarlskog invariant:

$$J_{CP}^{PMNS} = \sin \theta_{12} \cos \theta_{12} \sin \theta_{13} \cos^2 \theta_{13} \sin \theta_{23} \cos \theta_{23} \sin \delta_{CP}$$

$J_{CKM} \approx 3 \times 10^{-5}$ (PDG)
[arxiv:0308040 (Lepton Photon 2003) using $\gamma \approx 70^\circ$]

- Using the recent results of nuFit v5.1, in lepton sector:

$$J_{PMNS} \approx 0.034 \cdot \sin \delta_{CP}$$

CPV in Lepton Sector

- CPV in lepton sector is essential
- CPV can be measured in oscillation experiment $P(\nu_\alpha \rightarrow \nu_\beta)$
- Comparing neutrino probability with anti-neutrino probability
- So for CP Violation in neutrino mixing matrix

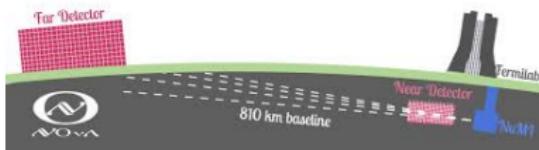
$$P(\nu_\alpha \rightarrow \nu_\beta) \neq P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta)$$

- In this discussion, we will use $P(\nu_\mu \rightarrow \nu_e)$ as oscillation channel.

Long Baseline Experiments: NO ν A and DUNE

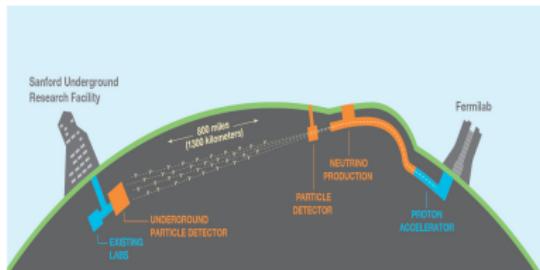


- Detect neutrinos in Fermilab's NuMI beam
- 14 mrad off-axis, $E \approx 2$ GeV
- Active liquid scintillator calorimeter
- Baseline $\rightarrow 810$ Km
- Two Detectors:
 - Near detector $\rightarrow 0.3$ kt
 - Far Detector $\rightarrow 14$ kt



DUNE

- proposed future superbeam experiment at Fermilab
- Liquid Argon (LAr) detector of mass 40 kt
- Baseline $\rightarrow 1300$ Km
- Far detector \rightarrow Homestake mine in South Dakota.



Long Baseline Experiments: T2K and T2HK

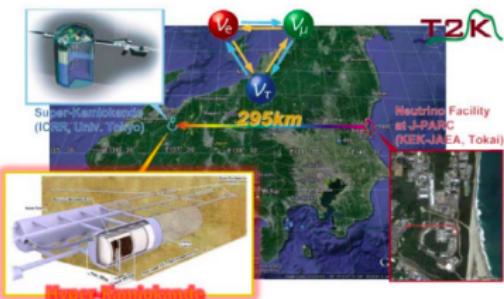


- Detect neutrinos in JPARC beam
- 43 mrad off-axis, $E \approx 0.65$ GeV
- water Cherenkov Detector
- Baseline \rightarrow 295 Km
- Two Detectors:
 - Near Detector \rightarrow ND280, 280 metres from the target
 - Far Detector \rightarrow (Super K), 295 km from the target in Tokai.



T2HK

- Upgraded version of T2K
- fiducial mass will be increased by about twenty times
- will contain two 187 kt third generation Water Cherenkov detectors
- Baseline \rightarrow 295 Km

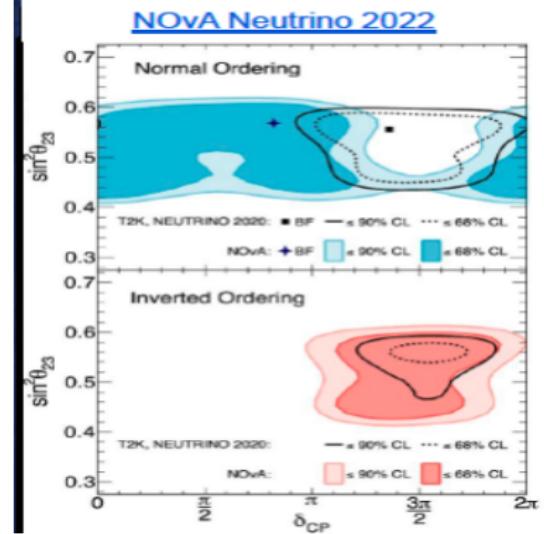
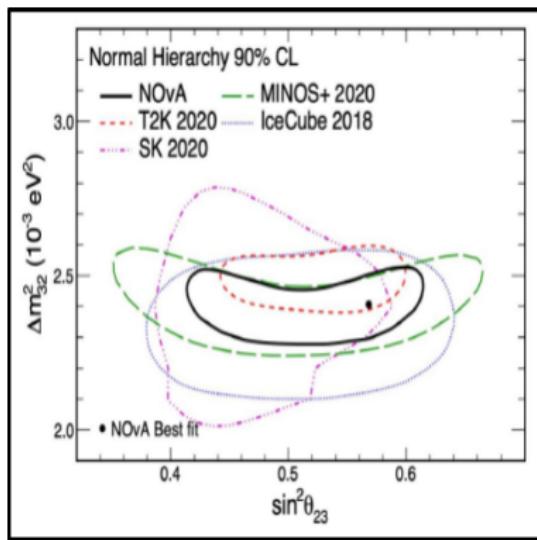


Exploring New Physics

- The main difference between NO ν A-T2K as well as DUNE-T2HK is the baseline and matter density, apart from energy.
- Neutrinos at NO ν A and DUNE experience stronger matter effects than T2K and T2HK
- New physics signature could probably be inferred from this exercise
- **Non-standard Interactions (NSI) → LBL CP Sensitivity**

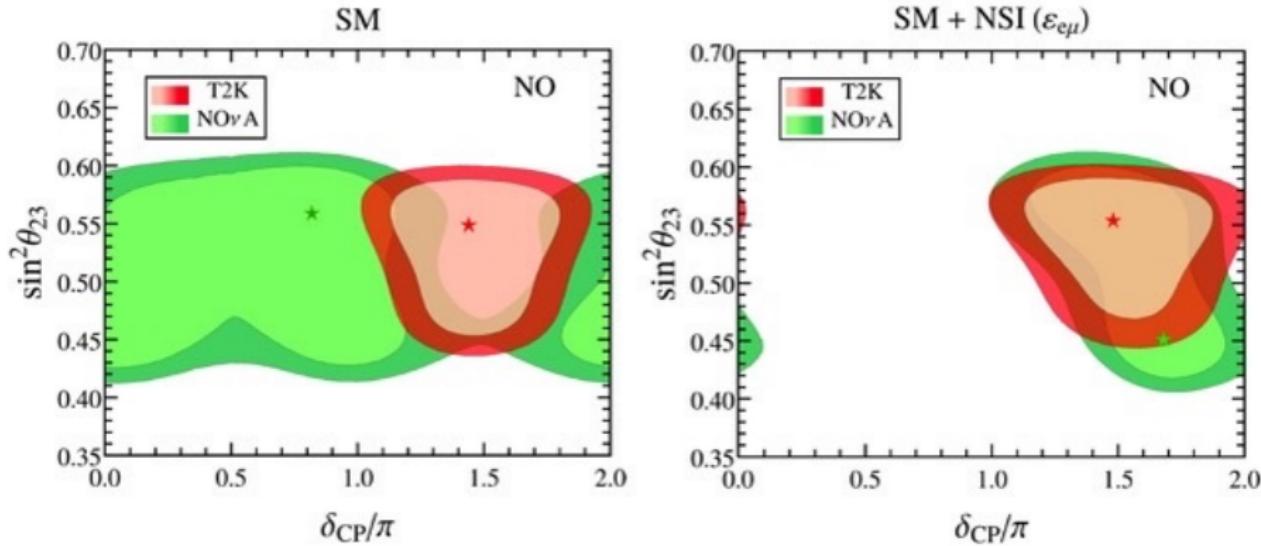
B Brahma, A Giri EPJ C 82, 1145 (2022)
[2302.09592, 2306.05258]

Some results



- The best fit value for θ_{23} in the higher octant and different values of δ_{CP} by NOvA for NO and IO.

LBL ν -CP Tension!!



PRL,126, 051802 (2021)

PRL 126, 051801 (2021)

Neutrino Non-Standard Interactions

- NSI can be characterized by dimension-six four-fermion operators of the form:

$$\mathcal{L}_{NSI} = -2\sqrt{2}G_F \sum_{\alpha,\beta,f,P} \epsilon_{\alpha\beta}^{f,P} [\bar{\nu}_\alpha \gamma^\mu \nu_\beta] [\bar{f} \gamma_\mu f] \quad (1)$$

- The neutrino propagation Hamiltonian in the presence of matter, NSI, can be expressed as

$$H_{Eff} = \frac{1}{2E} \left[U_{PMNS} \begin{bmatrix} 0 & 0 & 0 \\ 0 & \delta m_{21}^2 & 0 \\ 0 & 0 & \delta m_{31}^2 \end{bmatrix} U_{PMNS}^\dagger + V \right] \quad (2)$$

where,

$$V = 2\sqrt{2}G_F N_e E \begin{bmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} e^{i\phi_{e\mu}} & \epsilon_{e\tau} e^{i\phi_{e\tau}} \\ \epsilon_{\mu e} e^{-i\phi_{e\mu}} & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} e^{i\phi_{\mu\tau}} \\ \epsilon_{\tau e} e^{-i\phi_{e\tau}} & \epsilon_{\tau\mu} e^{-i\phi_{\mu\tau}} & \epsilon_{\tau\tau} \end{bmatrix}$$

where, $\epsilon_{\alpha\beta} e^{(i\phi_{\alpha\beta})} \equiv \sum_{f=e,u,d} (\epsilon_{\alpha\beta}^{fL} + \epsilon_{\alpha\beta}^{fR}) \frac{N_f}{N_e}$

Probability

- In the presence of NSI from $e\mu$ and $e\tau$ sectors, the probability can be expressed as the sum of terms *:

$$P_{\mu e} = P_{SM} + P_{\epsilon_{e\mu}} + P_{\epsilon_{e\tau}} + P_{Int} + h.o.$$

where,

$$P_0 = 4s_{13}^2 s_{23}^2 f^2 + 8s_{13}s_{23}s_{12}c_{12}c_{23}rfg \cos(\Delta + \delta_{CP}) + 4r^2 s_{12}^2 c_{12}^2 c_{23}^2 g^2$$

- P_0 denotes the SM probability expression

where,

$$f \equiv \frac{\sin[(1-\hat{A})\Delta]}{1-\hat{A}}, \quad g \equiv \frac{\sin \hat{A}\Delta}{\hat{A}}, \quad \hat{A} = \frac{2\sqrt{2}G_F N_e E}{\Delta m_{31}^2}, \quad \Delta = \frac{\Delta m_{31}^2 L}{4E}, \quad r = \frac{\Delta m_{21}^2}{\Delta m_{31}^2}$$

(*Phys.Rev.D77:013007,2008, JHEP 0903:114,2009, JHEP 0904:033,2009,
Phys.Rev.D93,093016(2016))

Probability

$$\begin{aligned} P_{\epsilon_{e\mu}} &= 4\hat{A}\epsilon_{e\mu}[xf^2s_{23}^2\cos(\Psi_{e\mu}) + xfgc_{23}^2\cos(\Delta + \Psi_{e\mu}) + yg^2c_{23}^2\cos\phi_{e\mu} \\ &+ ygfs_{23}^2\cos(\Delta - \phi_{e\mu})] + 4\hat{A}^2\epsilon_{e\mu}^2[f^2s_{23}^4 + g^2c_{23}^4 + 2fgs_{23}^2c_{23}^2\cos\Delta] \end{aligned}$$

$$\text{where } \Psi_{e\mu} = \phi_{e\mu} + \delta_{CP}$$

$$\begin{aligned} P_{\epsilon_{e\tau}} &= 4\hat{A}\epsilon_{e\tau}[xf^2s_{23}c_{23}\cos(\Psi_{e\tau}) - xfgs_{23}c_{23}\cos(\Delta + \Psi_{e\tau}) \\ &- yg^2s_{23}c_{23}\cos\phi_{e\tau} + ygfs_{23}c_{23}f\cos(\Delta - \phi_{e\tau})] \\ &+ 4\hat{A}^2\epsilon_{e\tau}^2s_{23}^2c_{23}^2[g^2 + f^2 - 2fg\cos\Delta] \end{aligned}$$

$$\text{where } \Psi_{e\tau} = \phi_{e\tau} + \delta_{CP}$$

$$\begin{aligned} P_{Int} &= 8\hat{A}^2c_{23}s_{23}\epsilon_{e\mu}\epsilon_{e\tau}[g^2c_{23}^2 + f^2s_{23}^2 + 2fgc_{23}^2\cos(\phi_{e\mu} - \phi_{e\tau})\cos\Delta \\ &- fg\cos(\Delta - \phi_{e\mu} + \phi_{e\tau})] \end{aligned}$$

Non-standard Interactions

- The flavor changing parameter of NSI:

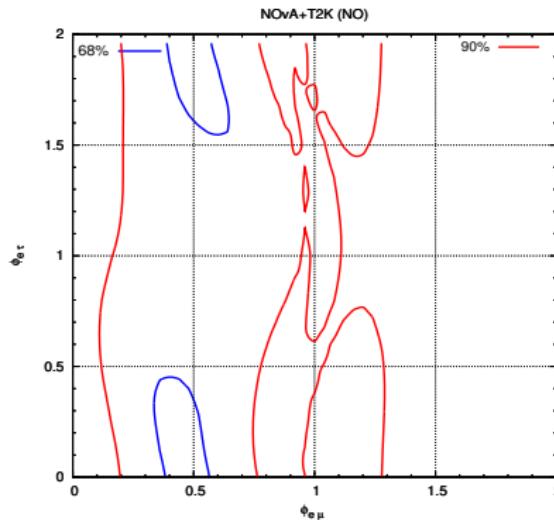
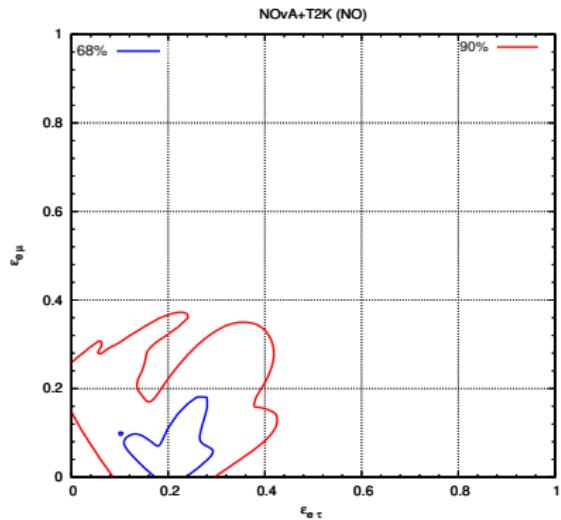
$$|\epsilon_{e\mu}|e^{i\phi_{e\mu}}, |\epsilon_{e\tau}|e^{i\phi_{e\tau}}, |\epsilon_{\mu\tau}|e^{i\phi_{\mu\tau}}$$

- In this work, we consider only the propagation NSI.
- Will discuss the effect of NSI ranges on sensitivity as well as oscillation probability plots for DUNE and T2HK.
- Use GLoBES and its additional public tools to deal with non-standard interactions *.

(*Comp.Phys.Comm, 167 (2005) 195; Comp. Phys. Comm, 177 (2007) 432;
<https://www.mpi-hd.mpg.de/personalhomes/globes/tools/snu-1.0.pdf> (2010).)

Dual NSI, $\epsilon_{e\mu}$ and $\epsilon_{e\tau}$ Sector

- Allowed regions in the plane spanned by NSI coupling $\epsilon_{e\mu}$ and $\epsilon_{e\tau}$ (left) and NSI coupling phase $\phi_{e\mu}$ and $\phi_{e\tau}$ (right) determined by the combination of T2K and NOvA for NO.
- The allowed regions at the 68% and 90% C.L.



Analysis Details

NSI Range

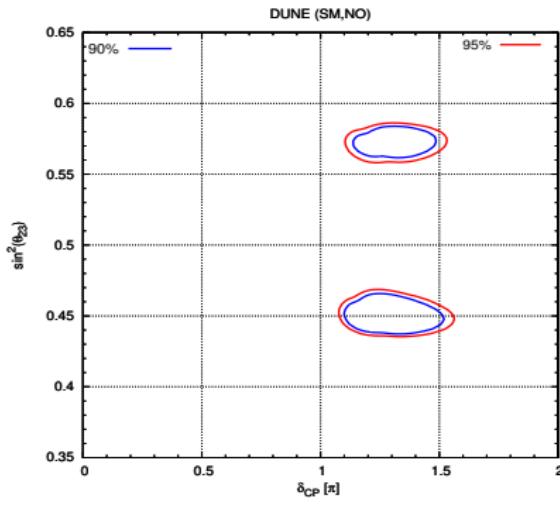
From the allowed region plots in the previous slides, the best-fit points are:

Mass Ordering	$ \epsilon_{e\mu} $	$ \epsilon_{e\tau} $
NO	0.22	0.06
IO	0.04	0.2
Mass ordering	$\phi_{e\mu}/\pi$	$\phi_{e\tau}/\pi$
NO	0.48	1.88
IO	1.24	1.87

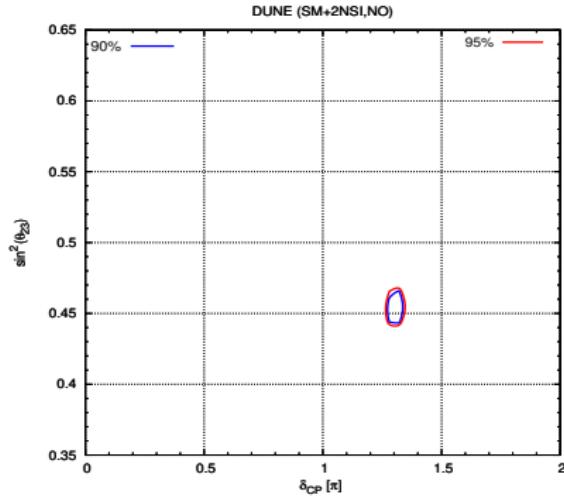
- In SM Plots the standard parameters θ_{13} is marginalized
- In SM+NSI plots, along with θ_{13} the NSI magnitudes ($|\epsilon_{e\mu}|$, $|\epsilon_{e\tau}|$) as well as phase ($\phi_{e\mu}, \phi_{e\tau}$) are marginalized
- The plots display the allowed regions at the 68% and 95% level

DUNE Sensitivity with dual NSI inclusion

SM, NO



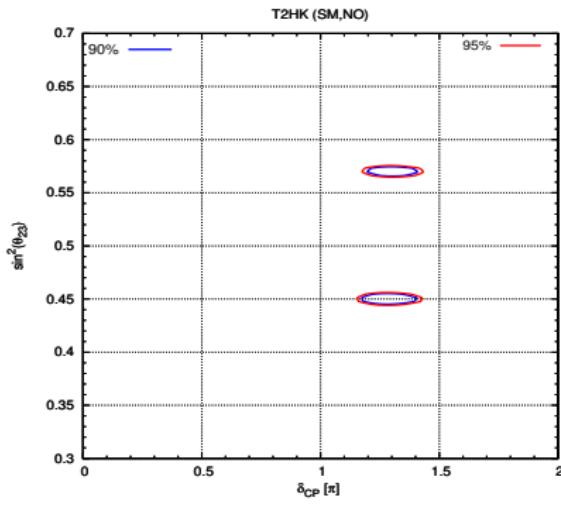
SM+ dual NSI, $\epsilon_{e\mu}$ and $\epsilon_{e\tau}$ Sector, NO



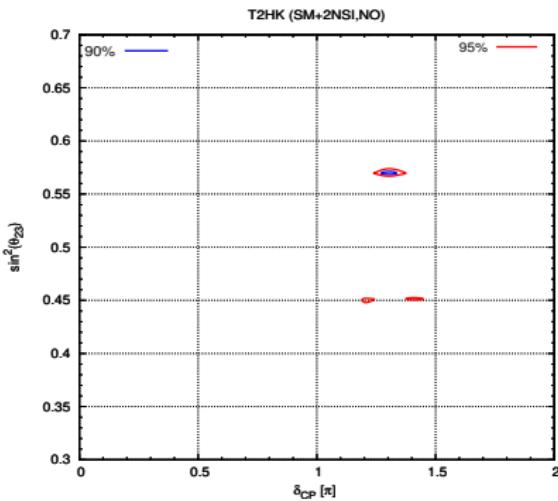
- With the inclusion of dual NSI from $e - \mu$ and $e - \tau$ sector, the allowed region corresponding to the higher octant in DUNE vanishes.

T2HK Sensitivity with dual NSI inclusion

SM, NO



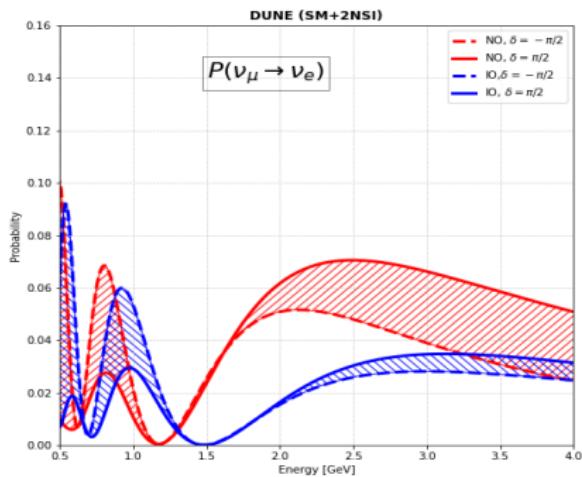
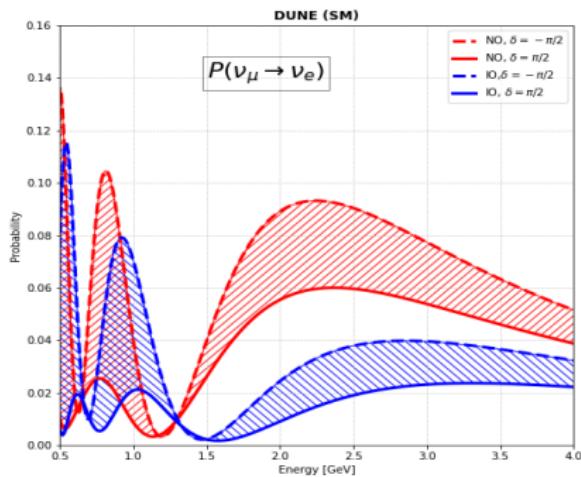
SM+dual NSI, $\epsilon_{e\mu}$ and $\epsilon_{e\tau}$ Sector,
NO



- With the inclusion of dual NSI from $e - \mu$ and $e - \tau$ sector, the allowed region corresponding to both the octants does not vanish completely.

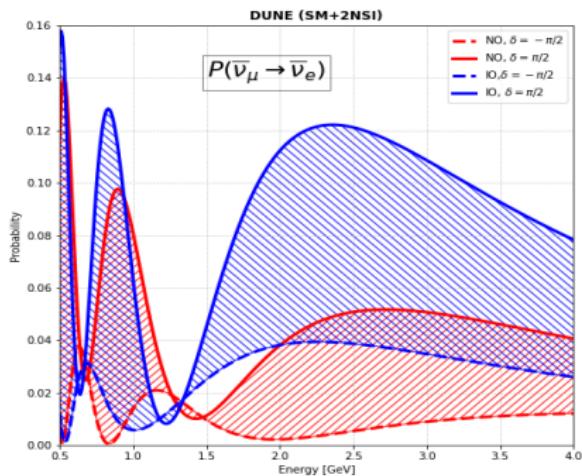
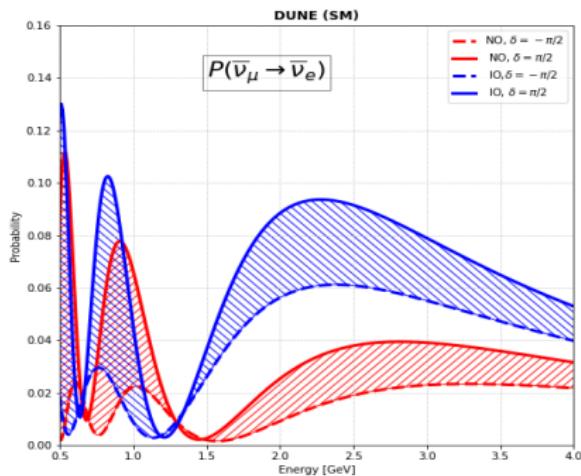
Probability, $P(\nu_\mu \rightarrow \nu_e)$ (DUNE)

- For the SM scenario, we see a good separation between NO-IO for both $\delta_{CP} = 90^\circ$ as well as $\delta_{CP} = -90^\circ$.
- For SM and dual NSI scenario, we still have some separation between NO-IO for $\delta_{CP} = -90^\circ$ in mid energy region, and they gradually merges around 4 GeV.



Probability, $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$ (DUNE)

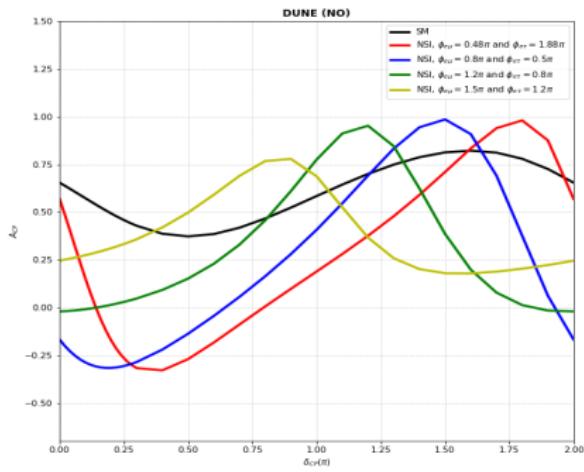
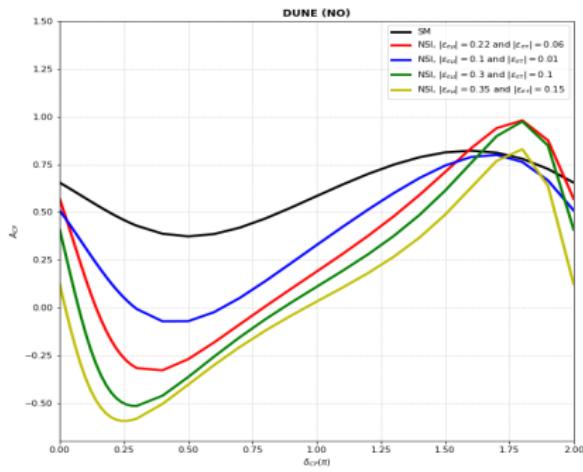
- For the SM scenario, we see a good separation between NO-IO for both $\delta_{CP} = 90^\circ$ as well as $\delta_{CP} = -90^\circ$.
- For SM and dual NSI scenario, the separation between NO-IO for $\delta_{CP} = 90^\circ$ becomes more than in the SM case. Compared with the SM case, the NO-IO separation decreases for $\delta_{CP} = -90^\circ$.



CP Asymmetry:DUNE

- Baseline = 1300 Km, Energy = 2.6 GeV

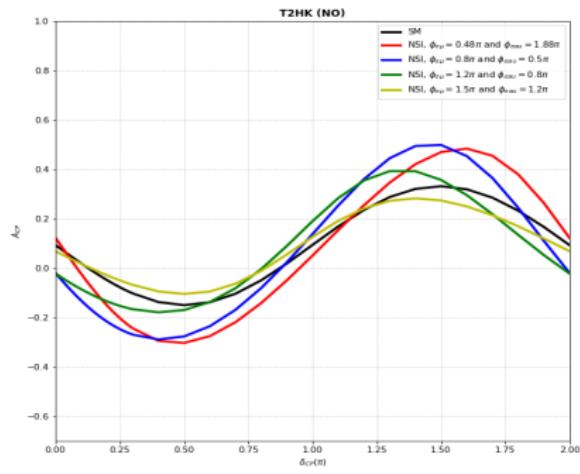
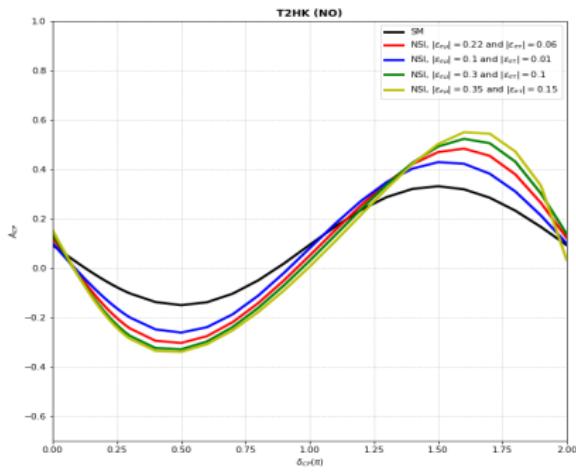
$$A_{CP} = \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}$$



CP Asymmetry:T2HK

- Baseline = 295 Km, Energy = 0.6 GeV

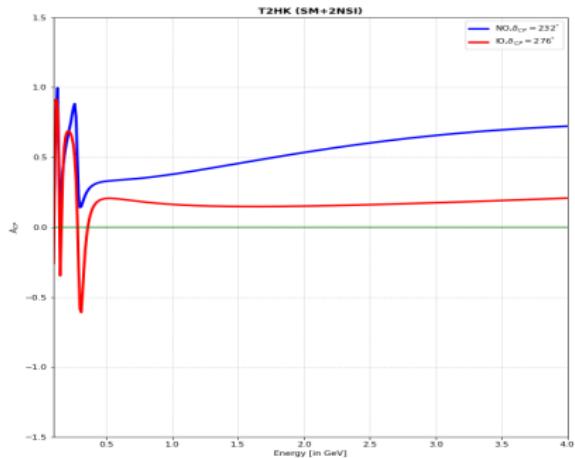
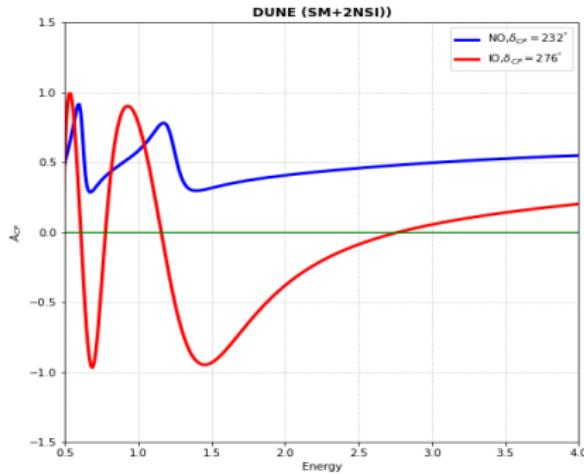
$$A_{CP} = \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}$$



CP Asymmetry versus Energy

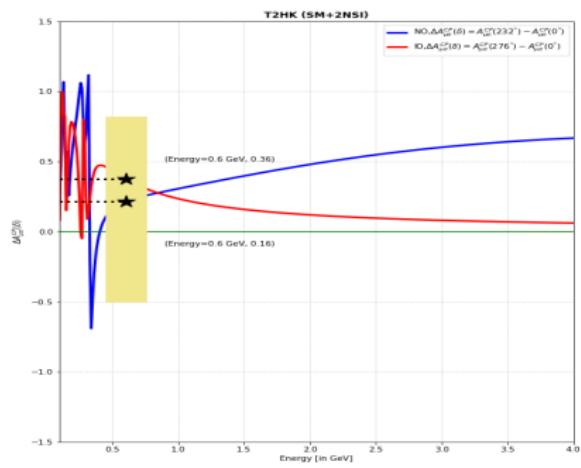
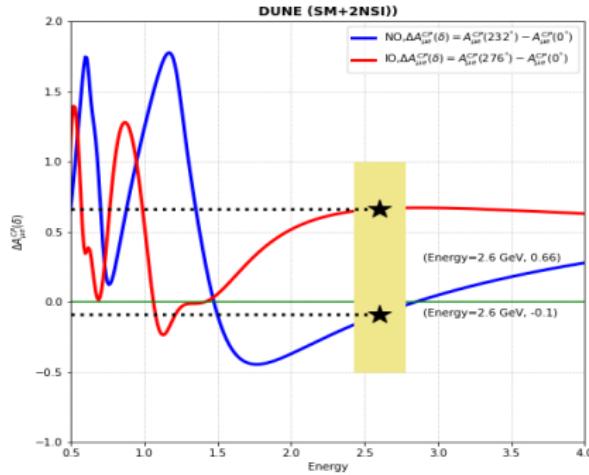
- For DUNE: Baseline = 1300 Km, $\delta_{CP} = 232^\circ$ (NO) and 272° (IO)
- For T2HK: Baseline = 295 Km, $\delta_{CP} = 232^\circ$ (NO) and 272° (IO)

$$A_{CP} = \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}$$



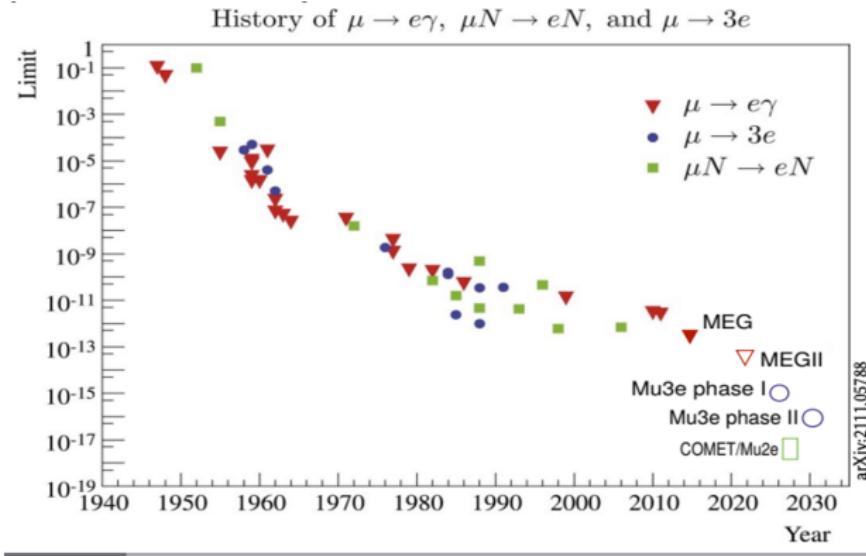
CP Asymmetry versus δ_{CP}

- For DUNE: Baseline = 1300 Km, Energy = 2.6 GeV
- For T2HK: Baseline = 295 Km, Energy = 0.6 GeV
- SM parameter δ_{CP} is varied from 0 to 2π



$$\Delta A_{\alpha\beta}^{CP}(\delta_{CP}) = A_{\alpha\beta}(\delta \neq 0) - A_{\alpha\beta}(\delta = 0)$$

LFV muon decays

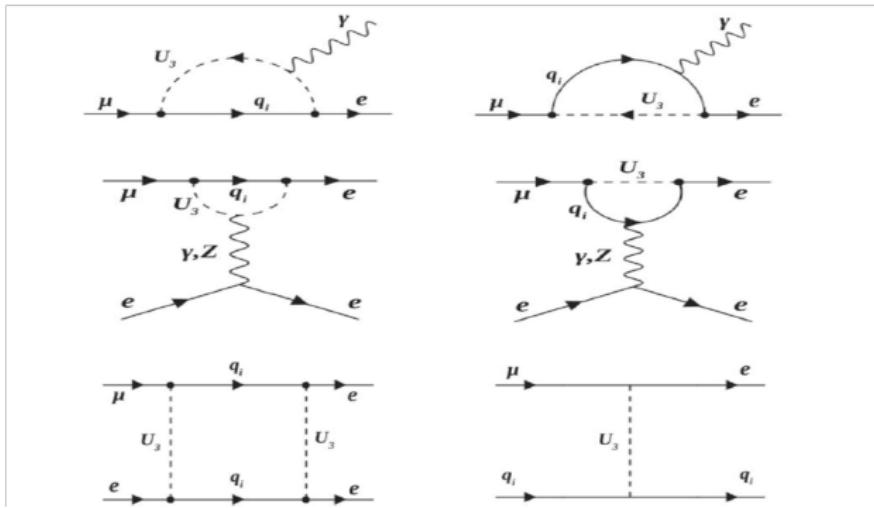


- Precision and BSM physics
- Neutrino NSI and muon LFV

Muon LFV with Leptoquark

- Some anomalies if the flavour and neutrino sector
- Leptoquarks: possible solution for simultaneous solution
- We consider U_3 vector leptoquark
- $\mathcal{L} \supset \chi_{ij}^{LL} \bar{Q}_L^{i,a} \gamma^\mu (\sigma_k \cdot U_{3,\mu}^k)^{ab} L_L^{j,b} + h.c.$
- And the effective four-fermion interaction
- $\mathcal{L}_{\text{eff}}^{\text{down}} = -\frac{2}{m_{LQ}^2} \lambda_{j\beta}^{LL} \lambda_{i\alpha}^{LL*} (\bar{d}^i \gamma_\mu P_L d^j) (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta)$
- $\mathcal{L}_{\text{eff}}^{\text{up}} = -\frac{2}{m_{LQ}^2} \lambda_{j\beta}^{LL} \lambda_{i\alpha}^{LL*} (\bar{u}^i \gamma_\mu P_L u^j) (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta)$
- vector leptoquark: $U_3(\bar{3}, 3, 2/3)$

Leptoquark contribution to LFV muon decays



Considering leptoquark mass 2 TeV

- $\mathcal{B}(\mu \rightarrow e\gamma) \approx 4.6 \times 10^{-18}$
- $\mathcal{B}(\mu \rightarrow eee) \approx 1.0 \times 10^{-20}$
- $\mathcal{B}(\mu \rightarrow e)_{Ti} \approx 6.8 \times 10^{-19}$

Summary

- With Dual NSI, allowed region for octant θ_{23} for DUNE and T2HK
- Striking differences in oscillation probabilities for ν channel in DUNE and T2HK, consequences for mass ordering
- CP asymmetry with NSI show significant differences in LBL Expts
- CP asymmetry vs. Energy show differences for DUNE and T2HK
- ΔA_{CP} vs. Energy for DUNE and T2HK exhibits sensitive pattern for NO and IO scenarios
- Muon LFV decays improved bounds

Thank You !!