

## Axion-like particle and dark sector search at BESIII

Xiaoxuan Ding Peking University On behalf of the BESIII collaboration

July 18th, Prague

42<sup>M</sup> INTERNATIONAL CONFERENCE ON HIGH ENERGY PHYSICS 18-24 July 2024 • Standard Model (SM) is incredibly successful in describing the properties of visible matter

 $\checkmark$  Ieptons, photon, Z,  $W^{\pm}$ , Higgs, quarks, mesons, hadrons



#### Matter-antimatter asymmetry



#### Galactic rotation curve



arXiv:astro-ph/0403324

#### PRL 131, 161802(2023)

There is nothing new to be discovered in physics now. 68% Dark Energy Dark Matter 27% Velocity (km s<sup>-1</sup>) О 20.000 30,000 40,000 Distance (light years) Visible Matter of Elem ntary Particles Dark Matter (Territors) Interactions / force carriers naus 63.2 MeM 38 DAVID area deve 115.11 GeV/P H C Ĵ up charm top gluon higgs nA 7 Mehile d s b down strange bottom photon O STO MeVIC 0.00 MeVA CMA Device e μ Physics BSM must exist! Z boson electron muon tau LEPTONS \*LB eVIC 17 Metallo V electron neutrino muon neutrino tau neutrino W boson



### **Dark sector**

- A collection of particles that are not charged directly under the SM strong, weak, or electromagnetic forces.
  - May interact with SM particle through portal interactions, not through gravitational effect not.
- ➡ Dark sector particles
- Axion: QCD axion and axion like particles
- Muonphilic vector or scalar
- **Dark photon**  $\gamma'$  : Massive or massless
- SUSY, dark Higgs, heavy neutrinos, dark fermion



If their mass are in the MeV-GeV range  $\rightarrow$  Accessible at BEPCII and BESIII experiment.



**Beijing Spectrometer III Detector** 

## **BESIII Experiment**

### **Beijing Electron–Positron Collider II**



NIMA 614, 345 (2010) CPC 44, 040001 (2020)





What is ALP	
O QCD axion ( <i>a</i> )	PRL. 40, 223 (1978) PRL. 40, 279 (1978)
• Predicted by the Peccei-Quinn (PQ) s	ution to the strong CP problem, could also be a dark matter candidate
<b>O</b> Axion-like particles (ALPs)	
• Have the same quantum numbers as <b>masses and couplings.</b>	e QCD axion, but could have <b>arbitrary</b> PLB 753,482 (2016)
• The ALP-photon coupling $g_{a\gamma\gamma}$ is mo	y discussed $\rightarrow$ ALP decays to two photon $e^{-}$ $\gamma$
• Experimental bounds on $g_{a\gamma\gamma}$ with $m_{e^+e^-}$ colliders	range of $MeV/c^2 - GeV/c^2$ mainly from
	Resonant ALP production
$\frac{1}{\sqrt{J/\psi}}$	s searched via $\gamma a, a \rightarrow \gamma \gamma$
Using 2.7B $\psi(2S)$ Data, in 0.165 < $m_a$ < 2.84 GeV/ $c^2$	Using 10B $J/\psi$ Data, in $0.18 < m_a < 2.85 \text{ GeV}/c^2$ JHEP. 06, 091 (2019)
PLB 838(2023) 137698	arXiv:2404.04640 Accepted by PRD(L)
dingxx@stu.pku.edu.cn	

ALP search

Dark Sector

Summary

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Introduction

BEPCII & BESIII





dingxx@stu.pku.edu.cn

0.5

0<sup>L</sup>

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 $10^{-3}$ 

 $10^{-2}$ 

 $10^{-1}$ 

 $m_a (GeV/c^2)$ 

2.5

3

2

1.5

 $m_a (GeV/c^2)$ 

10



date for  $0.18 < m_a < 2.85 \text{ GeV}$ 

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10<sup>1</sup>

BESIII (1/w)

10<sup>0</sup>

 $10^{-4}$ 

10-3

Beam dump

 $10^{-1}$ 

*m*<sub>a</sub> (GeV/*c*<sup>2</sup>)

 $10^{-2}$ 

Introduction	BEPCII & BESIII	ALP search	Dar	k Sector	Summary
( <b>Other</b> )		√Search •Ma •Mu	for fully invisus ssive dark pho onphilic scala	sible decays oton or $X_0$ or vector $X_1$	This talk
Dark sector search	tor search		$\eta,\eta',\omega,\phi$	PRD 105, L071101 (2022) PRD 98, 032001 (2018) PRD 87, 012009 (2013)	
and the second		<b>√</b> Other s	earches with	invisible signatures	PLB 852 ,138614 (2024)
		•FC	NC process:		PRD 105, L071102 (2022) PRD 105, 106, 072008 (2022)
		$\Sigma^+$	$\rightarrow p + \text{ invisit}$	ole, $\Lambda_c^+ \to p + \text{ invis}$	ible, $D^0 \to \pi^0 \nu \bar{\nu}$
		• $J/\psi$	$\gamma \rightarrow \gamma + invisit$	ble	PRD 101, 112005 (2020)
		<b>√</b> Visible	e dark photon	searches	
		t a l		. 1+1-	

•  $e^+e^- \rightarrow \gamma \gamma', \gamma' \rightarrow l^+l^-$ •  $J/\psi \rightarrow \eta \gamma', \gamma' \rightarrow e^+e^-$ 

PLB 774, 252(2017) PRD 99, 012006 (2019)



Introduction	BEPCII & BESIII	ALP search	Dark Sector	
What is Dark p	bhoton $\gamma'$			
O An extra Abelian g	gauge group $U(1)_D$ :		Magging of the group of	

PLB, 196 (1986)

- A minimal extension to SM, causing the associated spin-one boson the dark photon
- The dark photon has a kinetic mixing with SM photon through

with a kinetic mixing parameter  $\varepsilon \sim 10^{-3}$  (empirical, very small)

 $\rightarrow \frac{1}{2} \mathcal{E} F'_{\mu\nu} F^{\mu\nu}$ 

- *ɛ*: controls the coupling strength
- $F'_{\mu\nu}$ : field strengths of the dark photon

• Massive  $\gamma'$ , if the symmetry is spontaneously broken

Summary

• Massless  $\gamma'$ , if the symmetry is unbroken





Introduction	BEPCII & BESIII	ALP search	Dark Sector	Summary
Muonphilic scala	<b>r</b> $X_0$ or vector $X_1$			
$\rightarrow$ Similar to the previo	us dark photon, an extra <i>U</i> (1	) group is added as a mini	mal JHEP10(2020)207 Mod. Phys. Lett. A 06, 527 (1991 PRD 43, R22 (1991)	).
extension to the SM				
• $U(1)_{L_{\mu}-L_{\tau}}$ model : a	new massive vector boson	$X_1$ or scalar boson $X_0$ only	$\mathcal{L}_{\mu}^{ ext{scalar}} = -g_0  X$	$X_0 \ \overline{\mu} \ \mu,$
couple to the second	nd or third generations of l	<b>eptons</b> $(\mu, \nu_{\mu}, \tau, \nu_{\tau})$ with the	$\mathcal{L}_{\mu}^{\rm vector} = -g_1 X$	$X_{1\alpha} \overline{\mu} \gamma^{lpha} \mu.$
coupling strength g	' <sub>1</sub> , <i>8</i> ' <sub>0</sub>	Eur. Phys. J. C 81, 861	(2021).	γ γ
• The light muonphi	lic scalar or vector particle	es can contribute to the m		
anomalous magne	tic moment and explain the $q_2'$	$(g-2)_{\mu}$ anomaly $g'_{\mu}$	$\mu^-$	$\mu^-$
с —	$\gamma \mu^{-}$	$c \qquad \gamma \qquad \mu^{-}$	Can be accessible	e via $J/\psi  ightarrow \mu^+ \mu^- X_{0,1}$
			with $X_{0,1}$ invisible	at BESIII
(4	a) $\searrow \mu^+$	(b)	$\searrow \mu^+$	
dingxx@stu.pku.e	du.en	ICHEP 2024		14

Introduction	BEPCII & BESIII	ALP search	Dark Sector	Summary

### Three cases of muonphilic particles



Introduction	BEPCII & BESIII	ALP search	Dark Sector	Summary

- Search for a muonphilic scalar  $X_0$  or vector  $X_1$  via  $J/\psi \rightarrow \mu^+\mu^- +$  invisible decays <u>PRD 109, L031102 (2024)</u>
  - **O** Data samples: **9B**  $J/\psi$  events

• Signal extraction: A series of unbinned maximum likelihood fits are performed to  $M_{\text{recoil}}^2(\mu^+\mu^-)$  or  $M_{\text{recoil}}(\mu^+\mu^-)$ 

$$\begin{split} M_{recoil}^2(\mu^+\mu^-) &= \\ (p_J - p_{\mu^+} - p_{\mu^-})^2 \end{split}$$

- The maximum local significance is  $2.5\sigma$ at  $M(X_{0,1}) = 720 \text{ MeV}/c^2$
- **O** No evidence for  $J/\psi \rightarrow \mu^+\mu^- + X_{0,1}$ signals

• Low mass region, with  $M(X_{0,1}) = 120 \text{MeV}/c^2$ 



High mass region, with  $M(X_{0,1}) = 720 \text{MeV}/c^2$ 





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### Summary

→BESIII plays an active role in dark sector and axion-like particle search, with many first searches or best limits

### Axion-like particles (ALPs) search

Search for an **axion-like** particle in radiative  $J/\psi$  decays

✓ Using 2.7B  $\psi$ (3686) data PLB 838 137698 (2023)

✓ Using 10B  $J/\psi$  data

Accepted by PRD(L)



### Dark sector search

PRD 109, L031102 (2024)

✓ Search for a **muonphilic** scalar  $X_0$  or vector  $X_1$  via  $J/\psi \rightarrow \mu^+\mu^-$ + invisible =>>

✓ Search for Massive dark photon with  $e^+e^- \rightarrow \gamma \gamma'$  PLB 839 (2023) 137785

Stringent limits on the **coupling**  $g'_{0,1}$  are set



Introduction

**BEPCII & BESIII** 

ALP search



### Future

- BESIII has recently collected 20 fb<sup>-1</sup> of  $\psi(3770)$  data sample
  - ➡More conclusive results are ongoing!

Enlightening the dark, coming is the future!

### Thanks for your attention!

## Back UP







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- → A massless dark photon  $\gamma'$  could induce FCNC process through higher dimensional operators
- **Data samples**: 4.5 fb<sup>-1</sup>  $e^+e^-$  annihilation data at  $\sqrt{s} = 4.6 \sim 4.7$  GeV
- Strategy: Double-Tag technique
- **o** Results:
  - $\checkmark$  No signal is observed

✓ 90% C.L. upper limit on BF is set  $\mathscr{B}(\Lambda_c^+ \to p\gamma') < 8 \times 10^{-5}$ 



# $\begin{array}{c} c & & & u \\ \Lambda_c^+ & u & & & u \\ d & & & & d \end{array}$

10 hadronic decay modes



Introduction	BEPCII & BESIII	ALP search	Dark Sector	Summary
$\circ$ Mativation on	Dark sector			

- The dark sector could be light and communicate with the visible sector through a feeble portal interaction
  - The dark sector models can be classified based on the mediator particle or "portal"



• Beyond these, certain anomaly gauged  $U(1)_{L_{\mu}-L_{\tau}}$  model



Introduction	BEPCII & BESIII	ALP search	Dark Sector	Summary
Search for an	axion-like parti	cle in radiative	$J/\psi$ decays	PLB 838 137698 (2023)
O Data samples:	2.7B $\psi$ (3686) events			γ . [ <sup>2</sup>
O Strategy:			$\psi(2S)$	$ \begin{array}{c} \overline{c} \\ \overline{c} \\ \overline{c} \\ \overline{d} \\ J/\psi \end{array} $
Search for $J/\psi \rightarrow$	$\gamma a, a \rightarrow \gamma \gamma$ with $\psi$ (3686	$) \rightarrow \pi^+ \pi^- J/\psi$ decays	c c	
Search range:	$0.165 < m_a < 2.84 \text{ GeV}/c^2$	2		in a start of the
• <i>a</i> : negligible	decay width and lifetime	▶ decay width $\Gamma_a = g_a$	$m_{a\gamma\gamma}^2 m_a^3/64\pi$	$m_{L/k}^2$ ( $m^2$ ) <sup>3</sup>
• $\psi(3686)$ deca	ay		$\mathcal{B}(J/\psi  ightarrow \gamma a)$ =	$= \frac{J/\psi}{32\pi\alpha_{\rm em}} g_{a\gamma\gamma}^2 \left(1 - \frac{m_a}{m_{J/\psi}^2}\right) \mathcal{B}(J/\psi \to e^+ e^-),$
<ul><li>☆ preclude</li><li>QED bac</li></ul>	the pollution from non-resolution $e^+e^- \rightarrow \gamma\gamma(\gamma)$	onant production, avoid	large $(\overline{z_{2}})^{10^5}$	$ \begin{array}{c} \bullet  \text{Data} \qquad \qquad J/\psi \rightarrow \gamma \pi^0 \pi^0 \\ & \swarrow  J/\psi \rightarrow \gamma \pi^0  \swarrow  J/\psi \rightarrow \gamma \eta_c \\ & \qquad \qquad J/\psi \rightarrow \gamma \eta  \swarrow  J/\psi \rightarrow \gamma \gamma \gamma \\ & \qquad \qquad$
• Three $\gamma \gamma$ com fits on $M_{\gamma\gamma}$	binations per event, perform	m unbinned maximum-li	kelihood $50^{\circ}$ 10 <sup>4</sup>	

• Exclude mass intervals around  $\pi^0$ ,  $\eta$ ,  $\eta'$  peaks when extracting the signal



3

2.5

2

Introduction	BEPCII & BESIII	ALP search	Dark Sector	Summary

### Search for an axion-like particle in radiative $J/\psi$ decays

PLB	838	(2023)	) 137698
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Chinksmule Constant and maint

### Table 1

The $M_{\gamma\gamma}$ fit intervals to	$m_{\alpha}$ points.
$m_a$ points (GeV/ $c^2$ )	$M_{\gamma\gamma}$ fit intervals (GeV/ $c^2$ )
0.165 - 0.35	0.06 - 0.45
0.35 - 0.75	0.25 - 0.85
0.75 - 1.20	0.65 - 1.30
1.20 - 2.84	$(m_a - 0.2) - (m_a + 0.2)$

674 hypothesis

#### arXiv:2404.04640

TABLE I. The fit intervals of  $m_{\gamma\gamma}$  for various  $m_a$  points.

$m_a$ range	$m_{\gamma\gamma}$ fit interval	Polynomial
$({ m GeV}/c^2)$	$({ m GeV}/c^2)$	function order
0.180 - 0.420	0.16, 0.46	$4^{th}$
0.421 - 0.490	0.39, 0.51	$5^{th}$
0.610 - 0.880	0.59, 0.90	$5^{th}$
1.020 - 1.099	1.00, 1.20	$5^{th}$
1.100 - 2.770	$m_a - 0.10, m_a + 0.10$	$3^{\rm rd}$
2.772 - 2.850	2.70, 2.88	$4^{th}$



The largest value of upward local significance is determined to be 3.5 $\sigma$  at ma = 2.786 GeV/  $c^2$ 



Introduction	BEPCII & BESIII	ALP search	Dark Sector	Summary
Search for N	Massive dark p	ohoton with e	$^+e^- \rightarrow \gamma \gamma'$	PLB 839 (2023) 137785

$$\sigma(e^+e^- \to \gamma\gamma') = \frac{2\pi\alpha^2}{s}\epsilon^2 \left(1 - \frac{m_{\gamma'}^2}{s}\right) \times \left(\left(1 + \frac{2\frac{m_{\gamma'}^2}{s}}{\left(1 - \frac{m_{\gamma'}^2}{s}\right)^2}\right) \log \frac{(1 + \cos\theta_c)^2}{(1 + \cos\theta_c)^2} - 2\cos\theta_c\right) \quad \begin{array}{l} \cos\theta_c = 0.6 \text{ is the } \cos\theta_c = 0.6 \text{ is th$$

Search for single photon signals in  $1.3 < E(\gamma) < 1.8$  GeV corresponding to  $1.5 < m_{\gamma'} < 2.9$  GeV

Low E(γ) region → low trigger efficiency & high background level
 High E(γ) region → saturation of the EMC electronics

dingxx@stu.pku.edu.cn

Introduction	BEPCII & BESIII	ALP search	Dark Sector	Summary			
Search for a muonphilic scalar $X_0$ or vector $X_1$ via $J/\psi \rightarrow \mu^+ \mu^- +$ invisible decays							

Decay width

$$\begin{split} |\mathcal{M}_{\mu\mu\chi_{0}}|^{2} &= \left(\frac{2}{3}e^{2}g_{0}\frac{f_{J}}{m_{J}}\right)^{2}\frac{-8}{3\,m_{J}^{2}(m_{J}-2\,E_{-})^{2}(-2\,E_{-}-2\,E_{X}+m_{J})^{2}} \left(-4\,m_{\mu}^{2}\left(4\,E_{-}^{2}\left(m_{X}^{2}-2\,E_{X}\,m_{J}\right)\right)\right) \\ &+E_{-}\left(-8\,E_{X}^{2}\,m_{J}+4\,E_{X}\left(m_{X}^{2}+2\,m_{J}^{2}\right)-4\,m_{X}^{2}\,m_{J}\right)-E_{X}^{2}\left(m_{X}^{2}-6\,m_{J}^{2}\right)-2\,E_{X}\,m_{J}\left(m_{X}^{2}+m_{J}^{2}\right)+m_{X}^{2}m_{J}^{2}\right) \\ &+4\,E_{-}^{2}\left(2\,E_{X}^{2}\,m_{J}^{2}+m_{X}^{2}\,m_{J}(m_{J}-2\,E_{X})+m_{X}^{4}\right) \\ &+4\,E_{-}\left(2\,E_{X}^{3}\,m_{J}^{2}-2\,E_{X}^{2}\,m_{J}\left(m_{X}^{2}+m_{J}^{2}\right)+E_{X}\left(m_{X}^{4}+3\,m_{X}^{2}\,m_{J}^{2}\right)-m_{X}^{2}\,m_{J}\left(m_{X}^{2}+m_{J}^{2}\right)\right) \\ &-16\,E_{X}^{2}\,m_{\mu}^{4}+m_{J}\left(-4\,E_{X}^{3}\,m_{J}^{2}+2\,E_{X}^{2}\left(3\,m_{X}^{2}\,m_{J}+m_{J}^{3}\right)-2\,E_{X}\left(m_{X}^{4}+2\,m_{X}^{2}\,m_{J}^{2}\right)+m_{X}^{2}\,m_{J}\left(m_{X}^{2}+m_{J}^{2}\right)\right) \end{split}$$

where  $E_{-}$ , the energy of  $\mu^{-}$  and  $E_X$ , the energy of  $X_0$  are measured in the rest frame of  $J/\psi$ .

$$\Gamma_{\mu\mu X_{0,1}} = \int_{E_X^{min}}^{E_X^{max}} \int_{E_-^{min}}^{E_-^{max}} \frac{|\mathscr{M}_{\mu\mu X_{0,1}}|^2}{64\pi^3 m_J} dE_- dE_X,$$

$$\begin{split} |\mathcal{M}_{\mu\mu X_{1}}|^{2} &= \left(\frac{2}{3}e^{2} g_{1} \frac{f_{J}}{m_{J}}\right)^{2} \frac{-16}{3 m_{J}^{2}(m_{J}-2 E_{-})^{2}(-2 E_{-}-2 E_{X}+m_{J})^{2}} \left(16 E_{-}^{4} m_{J}^{2}+32 E_{-}^{3} m_{J}^{2}(E_{X}-m_{J}) + 2m_{\mu}^{2} \left(4 E_{-}^{2} \left(m_{J}(m_{J}-2 E_{X})+m_{X}^{2}\right)-4 E_{-} \left(2 E_{X}^{2} m_{J}-E_{X} \left(m_{X}^{2}+3 m_{J}^{2}\right)+m_{J} \left(m_{X}^{2}+m_{J}^{2}\right)\right) + 2 E_{X}^{2} \left(m_{X}^{2}+3 m_{J}^{2}\right)-2 E_{X} m_{J} \left(m_{X}^{2}+2 m_{J}^{2}\right)+m_{J}^{2} \left(m_{X}^{2}+m_{J}^{2}\right)\right) + 4 E_{-}^{2} \left(m_{J}^{2} \left(6 E_{X}^{2}-14 E_{X} m_{J}+7 m_{J}^{2}\right)+m_{X}^{2} m_{J} (3 m_{J}-2 E_{X})+m_{X}^{4}\right) + 4 E_{-} \left(2 E_{X}^{3} m_{J}^{2}-2 E_{X}^{2} m_{J} \left(m_{X}^{2}+4 m_{J}^{2}\right)+E_{X} \left(m_{X}^{4}+5 m_{X}^{2} m_{J}^{2}+9 m_{J}^{4}\right)-m_{J} \left(m_{X}^{4}+3 m_{X}^{2} m_{J}^{2}+3 m_{J}^{4}\right)\right) + 8 E_{X}^{2} m_{\mu}^{4}+m_{J} \left(-4 E_{X}^{3} m_{J}^{2}+2 E_{X}^{2} \left(3 m_{X}^{2} m_{J}+5 m_{J}^{3}\right)-2 E_{X} \left(m_{X}^{2}+2 m_{J}^{2}\right)^{2}+m_{J} \left(m_{X}^{4}+3 m_{X}^{2} m_{J}^{2}+2 m_{J}^{4}\right)\right)\right), \end{split}$$

where  $E_{-}$ , the energy of  $\mu^{-}$  and  $E_X$ , the energy of  $X_0$  are measured in the rest frame of  $J/\psi$ .