

From particle emission at accelerator facilities to axion properties in neutron stars: results and implications of the latest femtoscopic studies involving pions by ALICE

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on behalf of the ALICE Collaboration

Technical University of Munich

19<sup>th</sup> of June 2024 ICHEP 2024, Prague





- Inside of neutron stars is still an open topic
- Hyperons as one of the most discussed scenarios (YN & YNN)



#### **Neutron stars and QCD axions**



Impact of QCD axions on the equation of state (EoS)
 → Can lead to stiffer EoS

Reuven Balkin et al., JHEP 2020, 221 (2020) Reuven Balkin et al., arXiv 2307.14418



#### **Neutron stars and QCD axions**

- Impact of QCD axions on the equation of state (EoS)
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**Goal:** Study of  $pp\pi^{\pm}$  interactions using femtoscopy in small colliding systems

 $\rightarrow$  Access dynamics of pions with few nucleons



 $p_1$ 



 $r \sim 1 - 2 \, \text{fm}$ 

## On today's menu





#### **Two-particle femtoscopy**



- Observed femtoscopic correlation functions depend on
  - final-state interaction  $\left|\psi(\vec{k}^*,\vec{r}^*)\right|^2$
  - particle emitting source  $S(\vec{r}^*)$
  - $\rightarrow$  Talk by <u>M. Korwieser (19<sup>th</sup> of July 17:00)</u>
- The study of the interactions needs a well constrained source!
- Two-body  $p\pi^{\pm}$  interaction known from scattering experiments

L. Fabbietti, V. Mantovani Sarti and O. Vazquez Doce, Annu. Rev. Nucl. Part. Sci. (2021) 71:377-402



$$C(k^*) = \int S(\vec{r}^*) |\psi(\vec{k}^*, \vec{r}^*)|^2 d^3 \vec{r}^* = \mathcal{N} \frac{N_{\text{same}}(k^*)}{N_{\text{mixed}}(k^*)}$$

### Two-particle correlation function of $p\pi^+$





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# Fitting of data of $p\pi^+$



$$C_{\text{total}} = \mathcal{N} \times C_{\text{bckg}} \times \left[ \lambda_{\text{Gen}} C_0 + (1 - \lambda_{\text{Gen}}) \right] + \left[ N_{\Delta} P S(p_{\text{T}}, T) \times Sill(M_{\Delta}, \Gamma_{\Delta}) \right]$$

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## **Background contribution**

• Correlated background due to "mini-jet" contribution from hadronization process





# **Background contribution**

- Correlated background due to "mini-jet" contribution from hadronization process
- Background modelled with MC simulations using Pythia:
  - Obtain MC correlation function for pairs with common and non-common partonic origin (ancestors) separately
  - Use common  $C_{\rm c}$  and non-common  $C_{\rm nc}$  as templates to build the background
  - $C_{\text{bckg}} = \mathcal{N} \times [w_c C_c + (1 w_c) C_{\text{nc}}]$







$$C_{\text{total}} = \mathcal{N} \times C_{\text{bckg}} \times \left[ \lambda_{\text{Gen}} C_0 + (1 - \lambda_{\text{Gen}}) \right] + \left[ N_{\Delta} P S(p_{\text{T}}, T) \times Sill(M_{\Delta}, \Gamma_{\Delta}) \right]$$

• Background  $C_{bckg}$  via MC templates, controlled by  $w_c$ 

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$$C_{\text{total}} = \mathcal{N} \times C_{\text{bckg}} \times [\lambda_{\text{Gen}} C_0 + (1 - \lambda_{\text{Gen}})] + N_{\Delta} PS(p_T, T) \times Sill(M_{\Delta}, \Gamma_{\Delta})$$

- Background  $C_{bckg}$  via MC templates, controlled by  $w_c$
- Interaction C<sub>0</sub>(r<sub>core</sub>) Coulomb + strong interaction (fixed from scattering lengths)

M. Hoferichter et al., Phys.Rept. 625 (2016) 1–88 M. Hennebach et al., EPJA 50 (2014) 12, 190 M. Hoferichter et al., Phys.Rept. 625 (2016) 1-88.

	$p\pi^+$	pπ <sup>-</sup>
Scattering Length	-0.125 fm	0.121 fm

# Fitting of Data $p\pi^+$



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M. Hoferichter et al., Phys.Rept. 625 (2016) 1–88 M. Hennebach et al., EPJA 50 (2014) 12, 190

• Resonance description: Sill distribution  $Sill(M_{\Delta}, \Gamma_{\Delta})$ ,  $M_{\Delta}$  fixed to 1215 MeV

F. Giacosa et al., EPJA 57 (2021) 12

•  $PS(p_T, T)$  phase-space factor

$$PS(p_{\mathrm{T}},T) \propto \frac{m}{\sqrt{m^2 + p_{\mathrm{T}}^2}} \times \exp\left(-\frac{\sqrt{m^2 + p_{\mathrm{T}}^2}}{T}\right)$$

• Fit between between 0 and 450 MeV in k\*

# Fitting of the $p\pi^+$ correlation function



 Fit procedure repeated for different pair transverse mass ranges

• 
$$m_{\rm T} = \sqrt{\overline{m}^2 + k_{\rm T}^2}$$
 and  
 $\vec{k}_{\rm T} = \frac{1}{2} [\vec{p}_{\rm T,1} + \vec{p}_{\rm T,2}]$ 



# Core radius scaling $p\pi^+$



- *r*<sub>core</sub>: size of emission source of **primordial** particles
- r<sub>core</sub> of pπ<sup>+</sup> follows common scaling of pp, pK<sup>+</sup>, π<sup>±</sup>π<sup>±</sup> in pp collisions

ALICE, PLB, 811:135849, 2020 ALICE, <u>arXiv:2311.14527</u>, EPJC in press

→ Common emission source for all hadrons



#### On today's menu

 $r \sim 1 - 2 \text{ fm}$ 

A



Emission at small distances for  $p\pi^{\pm}$  systems!

Common source scaling observed in mesonmeson, meson-baryon, and baryon-baryon systems → Common emission source for all hadrons



 $\overline{p_1}$ 

 $\overline{p_2}$ 

#### On today's menu



Emission at small distances for  $p\pi^{\pm}$  systems!

Common source scaling observed in mesonmeson, meson-baryon, and baryon-baryon systems → Common emission source for all hadrons



π

π

π

Do we see effects beyond pairwise interactions in  $pp\pi^{\pm}$ ?



 $r \sim 1 - 2 \, \text{fm}$ 

# **Three-particle femtoscopy**

• Pair relative momentum not applicable in three-body system  $\rightarrow$  Use Lorentz-invariant hyper-momentum  $Q_3$ 

$$Q_3 = \sqrt{-q_{12}^2 - q_{23}^2 - q_{13}^2}$$

ALICE, PRC 89 (2014) 2, 024911



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• Three-particle correlation functions:  $3 \rightarrow 3$  scattering processes





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 $p_2$ 

 $\overline{p_3}$ 

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• Three-particle correlation functions:  $3 \rightarrow 3$  scattering processes

$$C(Q_3) = \mathcal{N} \frac{N_{\text{same}}(Q_3)}{N_{\text{mixed}}(Q_3)}$$

- Challenge of isolating three-body effects
  - $\rightarrow$  Effects of two-body and potential three-body interactions in the system



 $p_1$ 

 $p_2$ 

 $p_3$ 



#### • Cumulant decomposition:

R. Kubo, Journal of the Physical Society of Japan 17 no. 7, (1962) 1100–1120





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Lower-order contributions estimated in a data-driven way using the same- and mixed-events distributions



# Three-particle correlation function of $pp\pi^+$

- Overall attractive effects in triplet correlation function
- Signal consisting of two-body and potential three-body effects



# Two-particle contributions of $pp\pi^+$





# Two-body contributions of $pp\pi^+$





#### Three-body effects in $pp\pi^{\pm}$





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ALI-PREL-576418

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# Three-body effects in $pp\pi^\pm$

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- In both cases cumulant compatible with zero for large Q<sub>3</sub>
   → No three-body effects
- Three-body effects for small Q<sub>3</sub>< 200 MeV/c</li>
  - Repulsion for  $\mathrm{pp}\pi^+$
  - Attraction for  $pp\pi^-$



# What to take home



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# Backup

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- Final product of supernova explosions
- Very compact objects:
  - M ≈ 1-2 M<sub>☉</sub>
  - R ≈ 10-15 km (~ size of Munich area!)





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- Very compact objects:
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- Very dense and rather cold objects:
  - extreme densities of several  $\rho_0$
  - $T_{max} \sim few MeV$







- Final product of supernova explosions
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  - R ≈ 10-15 km (~ size of Munich area!)
- Very dense and rather cold objects:
  - extreme densities of several  $\rho_0$
  - $T_{max} \sim few MeV$
- Physics of neutron stars driven by their equation of state (EoS)





• EoS dependent on the particle composition and the possible interactions between them


## **Neutron Stars**



- EoS dependent on the particle composition and the possible interactions between them
- EoS linked to masses and radii of neutron stars via TOV equations



#### **Neutron Stars**



- EoS linked to masses and radii of neutron stars via TOV equations
- Pure neutron matter (PNM) supports heavy neutron stars of  $\rm 2M_{\odot}$



Adapted from D. Lonardoni et al., PRL 114, 092301 (2015)



- Chemical potential µ = m + Fermi energy
- Fermi energy increases with density
  - $\rightarrow \mu_n = \mu_{\Lambda}$ : conversion into baryons with strangeness (hyperons)







• High baryonic densities allow for the existence of strange particles, e.g. Λ hyperons



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- However: EoS softens with appearance of  $\Lambda$  hyperons
  - $\rightarrow$  cannot support heavy neutron stars



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- High baryonic densities allow for the existence of strange particles, e.g. Λ hyperons
- However: EoS softens with appearance of  $\Lambda$  hyperons
  - $\rightarrow$  cannot support heavy neutron stars
- Three-body interactions such as  $\Lambda NN$  play an important role



Adapted from D. Lonardoni et al., PRL 114, 092301 (2015)

# Three-body hadronic interactions with ALICE



 Three-hadron interactions of *ppp* & *pp*Λ are currently studied in ALICE with femtoscopy

ALICE, Eur. Phys. J. A 59 (2023) 145

Results of ppp from Run2 have first comparisons with theory

A. Kievsky, E. Garrido, M. Viviani, L.E. Marcucci, L. Serksnyte, R. Del Grande, <u>arXiv:2310.10428</u>, accepted by PRC

Calculations of ppΛ are currently on the way



A. Kievsky, E. Garrido, M. Viviani, L.E. Marcucci, L. Serksnyte, R. Del Grande, <u>arXiv:2310.10428</u>, accepted by PRC

#### **Other Three-Body Studies**





## **Other Three-Body Studies**





# **ALICE - A Large Ion Collider Experiment**

- pp at  $\sqrt{s}$  = 13 TeV
- 10<sup>9</sup> high-multiplicity (HM) events (Run 2)
- Direct detection of charged particles (protons, kaons, pions, deuterons)
- Very good PID capabilities of the detector resulting in very pure samples (protons ~ 98%, pions 99%)



# **Two-body femtoscopy**



L. Fabbietti and V. Mantovani Sarti and O. Vazquez Doce, Study of the Strong Interaction Among Hadrons with Correlations at the LHC, Annu. Rev. Nucl. Part. Sci. (2021) 71:377-402



# Two-body femtoscopy

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Repulsion

k\*

200

400

L. Fabbietti and V. Mantovani Sarti and O. Vazquez Doce, Study of the Strong Interaction Among Hadrons with Correlations at the LHC, Annu. Rev. Nucl. Part. Sci. (2021) 71:377-402



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# **Two-body Femtoscopy**



- Observed femtoscopic correlation functions depend on
  - Final-state interaction  $\left|\psi(\vec{k}^*, \vec{r}^*)\right|^2$
  - Particle emitting source  $S(\vec{r}^*)$
- The study of the interactions, especially also of exotic pairs, need a well constrained source!



$$C(k^*) = \int S(\vec{r}^*) |\psi(\vec{k}^*, \vec{r}^*)|^2 d^3 \vec{r}^* = \mathcal{N} \frac{N_{same}(k^*)}{N_{mixed}(k^*)}$$

## **The Source for Baryon Pairs**





- Known interaction allows the extraction of the source size
- Source modelled as a gaussian

$$S(r) = \frac{1}{(4\pi r_{eff}^2)^{3/2}} \exp\left(-\frac{r^2}{4r_{eff}^2}\right)$$

• *r<sub>eff</sub>* as free fit parameter



## A Common Baryon Source?

• Effective source size for different transverse mass bins, with

 $m_{\rm T} = \sqrt{\overline{m}^2 + k_T^2}$ and  $\vec{k}_{\rm T} = \frac{1}{2} [\vec{p}_{{\rm T},1} + \vec{p}_{{\rm T},2}]$ 

• Different effective source size for pp and  $p\Lambda$  pairs!





Proper source modelling

Resonance contributions fixed from statistical hadronization model and EPOS

• Effective source not taking short-lived resonances into account









- Source distribution of particles from Gaussian core (r<sub>core</sub>) and decay of short-lived particles
- Different effective source size for ppand  $p\Lambda$  pairs **BUT**
- Common core source of primordial baryons!



## A Common Baryon Hadron Source!



ALICE, arXiv:2311.14527

ALICE pp  $\sqrt{s} = 13 \text{ TeV}$ 

High-mult. (0-0.17% INEL>0)

- Particle emission source studied with
  - pp ALICE, PLB, 811 135849, 2020
  - ALICE, arXiv:2311.14527 •  $pK^+$
  - $\pi^{\pm}\pi^{\pm}$  ALICE, <u>arXiv:2311.14527</u>
- Common source for all hadrons!
- Scaling allows to extract the source size of particle pairs with unknown interaction

 $\rightarrow$  Possibility to study interaction for exotic pairs (strange and charm sector)







# Fitting of data $p\pi^+$



 $C_{\text{total}} = \mathcal{N} \times C_{\text{bckg}} \times [\lambda_{\text{Gen}} C_0 + (1 - \lambda_{\text{Gen}})] + N_{\Delta} PS(p_{\text{T}}, T) \times Sill(M_{\Delta}, \Gamma_{\Delta})$ 

- Background  $C_{bckg}$  via MC templates, controlled by  $w_c$
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M. Hoferichter et al, Phys.Rept. 625 (2016) 1-88 M. Hennebach et al, Eur.Phys.J.A 50 (2014) 12, 190

Sill distribution Sill(M<sub>Δ</sub>, Γ<sub>Δ</sub>), M<sub>Δ</sub> fixed to 1215 MeV
F. Giacosa et al, Eur.Phys.J.A 57 (2021) 12

•  $PS(p_T, T)$  phase-space factor

$$PS(p_{\mathrm{T}},T) \propto \frac{m}{\sqrt{m^2 + p_{\mathrm{T}}^2}} \times \exp\left(-\frac{\sqrt{m^2 + p_{\mathrm{T}}^2}}{T}\right)$$

Fit between between 0 and 450 MeV in k\*



- Overall normalisation N
- $W_c$
- r<sub>core</sub>
- Scaling of  $\Delta^{++} N_{\Delta}$
- T (kinetic decoupling temp.)
- Width of  $\Delta^{++}$

























### About the life of the $\Delta^{++}$





### About the life of the $\Delta^{++}$





## Rescattering of the $\Delta^{++}$



- Study of kinetic mass shifts of  $\rho(770)$  and K\*(892) in Au+Au reactions at E<sub>beam</sub> = 1.23 AGeV with UrQMD
- Fitting of Data with PS x BW
- However: Temperature not fixed to chemical freezout but free parameter





## **Rescattering of the** $\Delta^{++}$

- Paper: Tom Reichert, Marcus Bleicher, <u>Nucl.Phys.A 1028 (2022) 122544</u>
- Study of kinetic mass shifts of  $\rho(770)$  and K\*(892) in Au+Au reactions at E<sub>beam</sub> = 1.23 AGeV with UrQMD
- Fitting of Data with PS x BW
- However: Temperature not fixed to chemical freezout but free parameter ("Kinetic Decoupling Temperature")
  - $\rightarrow$  good agreement between UrQMD and fit



### About the life of the $\Delta^{++}$





# Kinetic decoupling temperature $\Delta^{++}$

• Low "decoupling temperature" of about 25 MeV



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# Kinetic decoupling temperature $\Delta^{++}$



- Low "decoupling temperature" of about 25 MeV
- This does not mean that pp collisions are cold! (NeV)  $^{\rm (2E21)}_{\rm V^{+V}}$





# Kinetic decoupling temperature $\Delta^{++}$

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- Low "decoupling temperature" of about 25 MeV
- This does not mean that pp collisions are cold!
- We see a modification of the phase space of resonance
- $\rightarrow$  hadronic moshpit for the  $\Delta^{++}$



Thanks to Berkin and his AI friends!



# Width & kinetic decoupling temperature $\Delta^{++}$



- Width constant ~ 90 MeV
- Low "decoupling temperature"  $\rightarrow$  modification of the phase space of resonance


# Fitting of data $p\pi^-$



 $C_{\text{total}} = N \times C_{\text{bckg}} \times [\lambda_{\text{Gen}} C_0 + (1 - \lambda_{\text{Gen}})] + N_{\Delta} PS(p_{\text{T}}, T) \times Sill(M_{\Delta}, \Gamma_{\Delta}) + N_{\Lambda} Gaus(M_{\Lambda}, \Gamma_{\Lambda})$ 

- Background  $C_{bckg} = [1 + N_B(w_cC_c + (1 w_c)C_{NC} 1) + Sill(M_2, \Gamma_2) + Sill(M_3, \Gamma_3)]$
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$$PS(p_{\rm T},T) \propto \frac{m}{\sqrt{m^2 + p_{\rm T}^2}} \times \exp\left(-\frac{\sqrt{m^2 + p_{\rm T}^2}}{T}\right)$$

Fit between between 0 and 450 MeV in k\*

Free Parameters of the fit:

- Overall normalisation
- $w_c \& N_B$
- r<sub>core</sub>
- Scaling of  $\Delta^0 N_\Delta$
- T (kinetic decoupling temp.)
- Width of  $\Delta^0$
- Scaling of  $\Lambda$   $\textit{N}_{\Lambda}$
- Mass of  $\Lambda$
- Width of  $\Lambda$





## $p\pi^{-}$ - $m_{T}$ interval 2













## $p\pi^{-}$ - $m_{T}$ interval 5









# Width & kinetic decoupling temperature $\Delta^0$



- Width constant ~ 90 MeV
- Low "decoupling temperature"  $\rightarrow$  modification of the phase space of resonance



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ALI-PREL-576373





### Source



#### Source from resonance source model



r\* in fm

# Three-particle correlation function of $pp\pi^{-}$

- Overall attractive effects in triplet correlation function
- Signal consisting of two-body and potential three-body effects



# Two-particle contributions of $pp\pi^-$





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# Two-body contributions of $pp\pi^-$



