

How is CPV in leptogenesis connected to CPV in ν -oscillations? —— the first general + explicit link in the type-I seesaw ——

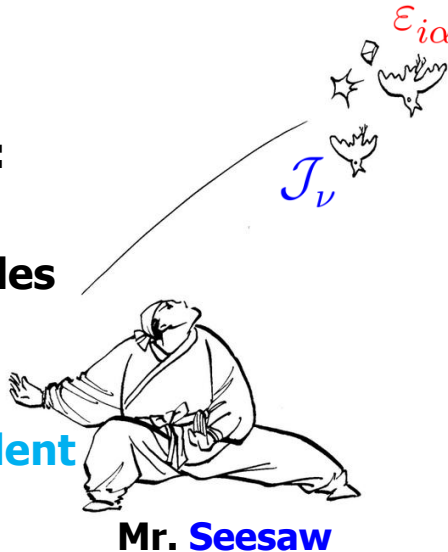
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Theorists:

Type-I seesaw mechanism:

- 3** heavy neutrino masses
- 9** active-sterile mixing angles
- 6** CP-violating phases

Hundreds of **model-dependent** connections on the market



Experimentalists:

Low-energy measurements:

- 3** light neutrino masses
- 3** active mixing angles
- 3** CP-violating phases

The **1st model-independent** link (**ZZX**, 2406.01142)

Seesaw: high gains + low costs

◆ **Fundamentals** of the electroweak SM structure → **reasons** for zero ν -mass:

- ◆ The Lorentz invariance
- ◆ Local $SU(2)_L \times U(1)_Y$ gauge symmetries
- ◆ The Higgs mechanism
- ◆ Renormalizability (no $d \geq 5$ operators)

- Plus *economical* particle content:
- No right-handed neutrino fields
 - Only one Higgs doublet

Go beyond the $d=4$ operators

$$m_1 = m_2 = m_3 = 0$$



SMEFT

The unique **d=5** operator for ν -mass and **Majorana** nature

$$\mathcal{O}_w = \frac{\bar{\ell}_L \widetilde{H} \widetilde{H}^T \ell_L^c}{\Lambda}$$

Steven Weinberg
1979

R-handed neutrinos + **Majorana** character

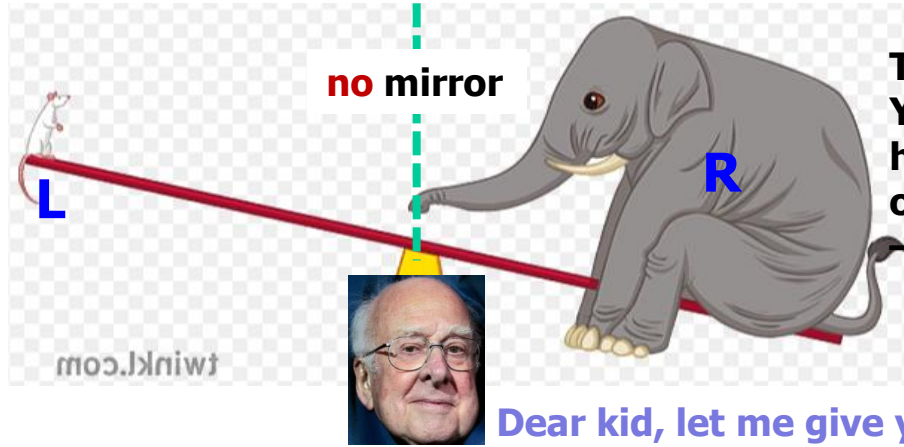
◆ The most convincing **UV** mechanism = **canonical seesaw** fully consistent with the SMEFT spirit, most natural/economical extension of the SM. *Bonus: Leptogenesis, SO(10) GUT-friendly...*

Integrate out the heavy dof

Three key issues of the seesaw

- ◆ **Right-handed** neutrino fields are **not** the **mirror** counterparts of the **left-handed** ones

It is said that I was **weightless** at birth, and it was you who fed me up a bit.



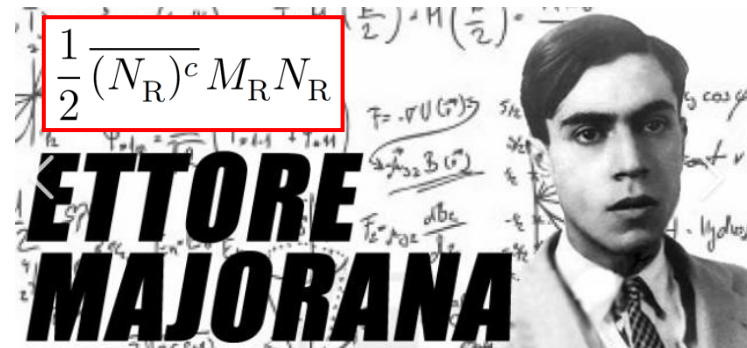
That is true, my love!
You may call me heavy neutrinos or heavy neutral leptons — both sound strange!

Dear kid, let me give you mass on behalf of God

- ◆ **Yukawa interactions** — the **Higgs** fields play a crucial role, as they do in generating masses for the charged fermions in the SM.

- ◆ The **Majorana** nature of massive neutrinos: N and N^c may have **self-interactions**, respecting all the fundamental symmetries of the SM.

Gell-Mann's totalitarian principle (1956)
Everything not forbidden is compulsory!



It works both before and after SSB

The **seesaw** mechanism formally works far above the **Fermi** scale, **before SSB** (ZZX, 2301.10461):

$$\begin{aligned}
 -\mathcal{L}_{\text{lepton}} &= \bar{\ell}_L Y_l H l_R + \bar{\ell}_L Y_\nu \tilde{H} N_R + \frac{1}{2} \overline{(N_R)^c} M_R N_R + \text{h.c.} \\
 &= \bar{\ell}_L Y_l l_R \phi^0 + \frac{1}{2} \overline{[\nu_L \quad (N_R)^c]} \begin{pmatrix} \mathbf{0} & Y_\nu \phi^{0*} \\ Y_\nu^T \phi^{0*} & M_R \end{pmatrix} \begin{bmatrix} (\nu_L)^c \\ N_R \end{bmatrix} + \bar{\nu}_L Y_l l_R \phi^+ - \bar{\ell}_L Y_\nu N_R \phi^- + \text{h.c.}
 \end{aligned}$$

The **basis transformation** related to the origin of active **Majorana** neutrino masses even **before SSB**:

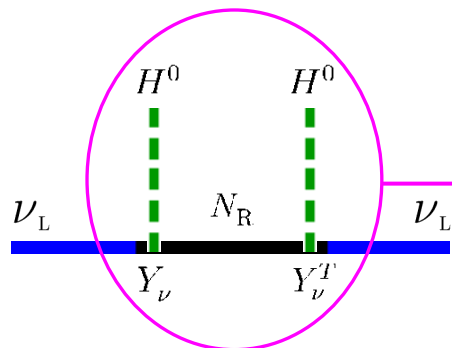
$$\mathbb{U}^\dagger \begin{pmatrix} \mathbf{0} & Y_\nu \phi^{0*} \\ Y_\nu^T \phi^{0*} & M_R \end{pmatrix} \mathbb{U}^* = \begin{pmatrix} D_\nu & \mathbf{0} \\ \mathbf{0} & D_N \end{pmatrix}$$

working masses: $\begin{cases} D_\nu \equiv \text{Diag}\{m_1, m_2, m_3\} & \text{light} \\ D_N \equiv \text{Diag}\{M_4, M_5, M_6\} & \text{heavy} \end{cases}$

SSB

$$\begin{pmatrix} \mathbf{0} & M_D \\ M_D^T & M_R \end{pmatrix}$$

6 × 6 mass matrix



Integrating out the heavy degrees of freedom:

$$-\mathcal{L}_{\text{mass}} = \frac{1}{2} \bar{\nu}_L M_\nu \nu_L^c + \text{h.c.} \quad M_\nu \simeq -Y_\nu \frac{\langle H \rangle^2}{M_R} Y_\nu^T$$

Consistent with the dim-5 Weinberg operator!

If you can untie **Weinberg's** knot, you will find new heavy **Majorana** neutrinos at a superhigh scale.

A full parameterization of seesaw

A **block parametrization** of **active-sterile** flavor mixing in the seesaw framework:

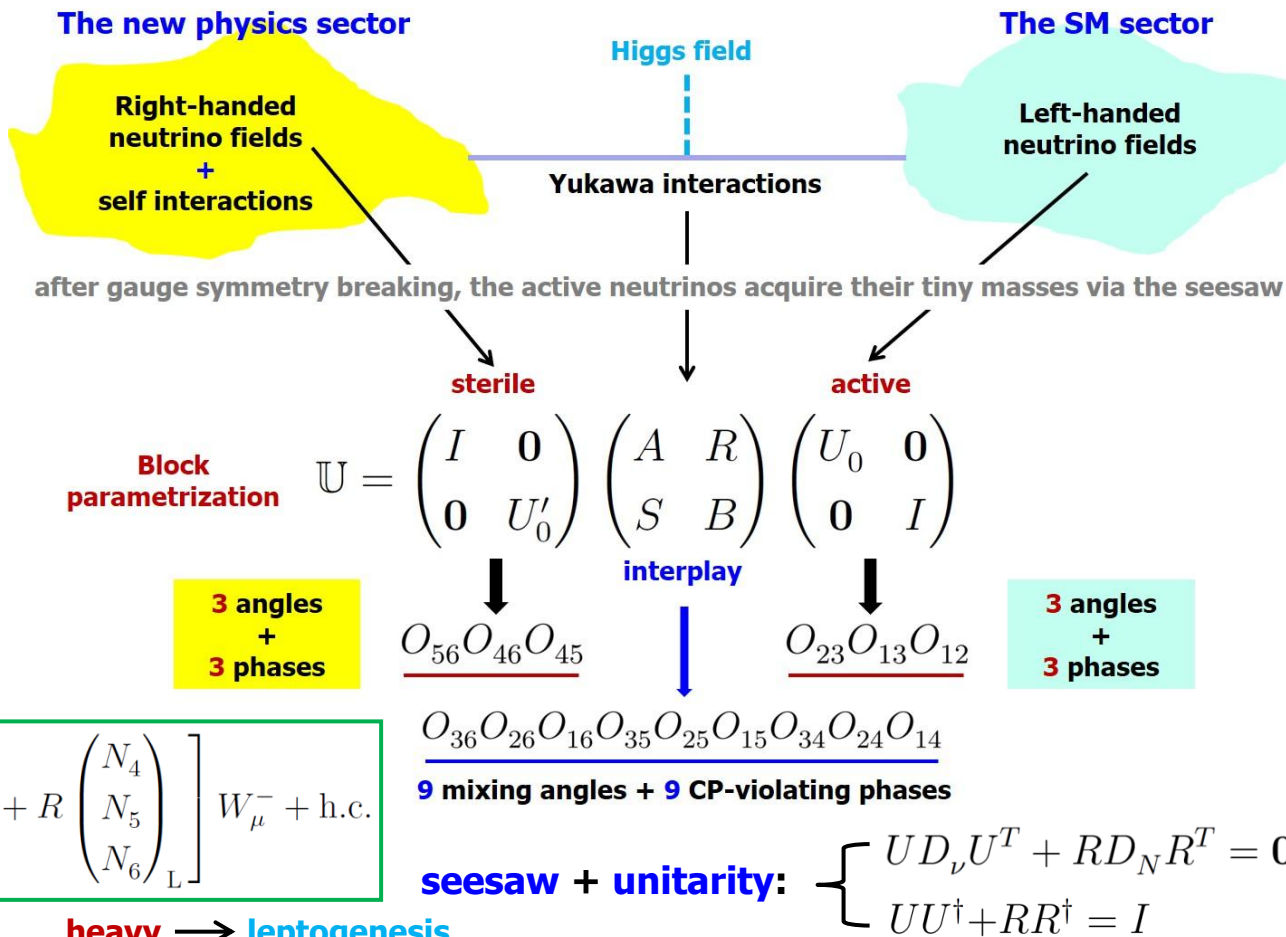
- ◆ reflects salient features of the seesaw dynamics
- ◆ offers **generic** + **explicit** expressions of observables using the **Euler-like** angles and phases (**ZZX**, 1110.0083)

The weak **charged-current** interactions of leptons:

$U = AU_0$: the **PMNS** matrix;
 R : an analogue for heavy.

$$-\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} \overline{(e \ \mu \ \tau)_L} \gamma^\mu \left[U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}_L + R \begin{pmatrix} N_4 \\ N_5 \\ N_6 \end{pmatrix}_L \right] W_\mu^- + \text{h.c.}$$

oscillations ← **light** **heavy** → leptogenesis



Original vs derivational seesaw parameters

5

$$U_0 = \begin{pmatrix} c_{12}c_{13} & \hat{s}_{12}^*c_{13} & \hat{s}_{13}^* \\ -\hat{s}_{12}c_{23} - c_{12}\hat{s}_{13}\hat{s}_{23}^* & c_{12}c_{23} - \hat{s}_{12}^*\hat{s}_{13}\hat{s}_{23}^* & c_{13}\hat{s}_{23}^* \\ \hat{s}_{12}\hat{s}_{23} - c_{12}\hat{s}_{13}c_{23} & -c_{12}\hat{s}_{23} - \hat{s}_{12}^*\hat{s}_{13}c_{23} & c_{13}c_{23} \end{pmatrix} \text{derivational parameters!}$$

$$c_{ij} \equiv \cos \theta_{ij}$$

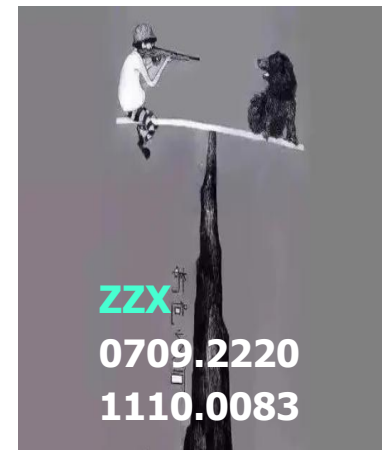
$$\hat{s}_{ij} \equiv e^{i\delta_{ij}} \sin \theta_{ij} \text{ (for } 1 \leq i < j \leq 6 \text{)}$$

$$A = \begin{pmatrix} c_{14}c_{15}c_{16} & 0 & 0 \\ -c_{14}c_{15}\hat{s}_{16}\hat{s}_{26}^* - c_{14}\hat{s}_{15}\hat{s}_{25}^*c_{26} & c_{24}c_{25}c_{26} & 0 \\ -\hat{s}_{14}\hat{s}_{24}^*c_{25}c_{26} & & 0 \\ -c_{14}c_{15}\hat{s}_{16}c_{26}\hat{s}_{36}^* + c_{14}\hat{s}_{15}\hat{s}_{25}^*\hat{s}_{26}\hat{s}_{36}^* & -c_{24}c_{25}\hat{s}_{26}\hat{s}_{36}^* - c_{24}\hat{s}_{25}\hat{s}_{35}^*c_{36} & c_{34}c_{35}c_{36} \\ -c_{14}\hat{s}_{15}c_{25}\hat{s}_{35}^*c_{36} + \hat{s}_{14}\hat{s}_{24}^*c_{25}\hat{s}_{26}\hat{s}_{36}^* & -\hat{s}_{24}\hat{s}_{34}^*c_{35}c_{36} & \\ +\hat{s}_{14}\hat{s}_{24}^*\hat{s}_{25}\hat{s}_{35}^*c_{36} - \hat{s}_{14}c_{24}\hat{s}_{34}^*c_{35}c_{36} & & \end{pmatrix}$$

$$R = \begin{pmatrix} \hat{s}_{14}^*c_{15}c_{16} & \hat{s}_{15}^*c_{16} & \hat{s}_{16}^* \\ -\hat{s}_{14}^*c_{15}\hat{s}_{16}\hat{s}_{26}^* - \hat{s}_{14}^*\hat{s}_{15}\hat{s}_{25}^*c_{26} & -\hat{s}_{15}^*\hat{s}_{16}\hat{s}_{26}^* + c_{15}\hat{s}_{25}^*c_{26} & c_{16}\hat{s}_{26}^* \\ +c_{14}\hat{s}_{24}^*c_{25}c_{26} & & \\ -\hat{s}_{14}^*c_{15}\hat{s}_{16}c_{26}\hat{s}_{36}^* + \hat{s}_{14}^*\hat{s}_{15}\hat{s}_{25}^*\hat{s}_{26}\hat{s}_{36}^* & -\hat{s}_{15}^*\hat{s}_{16}c_{26}\hat{s}_{36}^* - c_{15}\hat{s}_{25}^*\hat{s}_{26}\hat{s}_{36}^* & c_{16}c_{26}\hat{s}_{36}^* \\ -\hat{s}_{14}^*\hat{s}_{15}c_{25}\hat{s}_{35}^*c_{36} - c_{14}\hat{s}_{24}^*c_{25}\hat{s}_{26}\hat{s}_{36}^* & +c_{15}c_{25}\hat{s}_{35}^*c_{36} & \\ -c_{14}\hat{s}_{24}^*\hat{s}_{25}\hat{s}_{35}^*c_{36} + c_{14}c_{24}\hat{s}_{34}^*c_{35}c_{36} & & \end{pmatrix}$$

The **original seesaw** parameters in **A+R**:

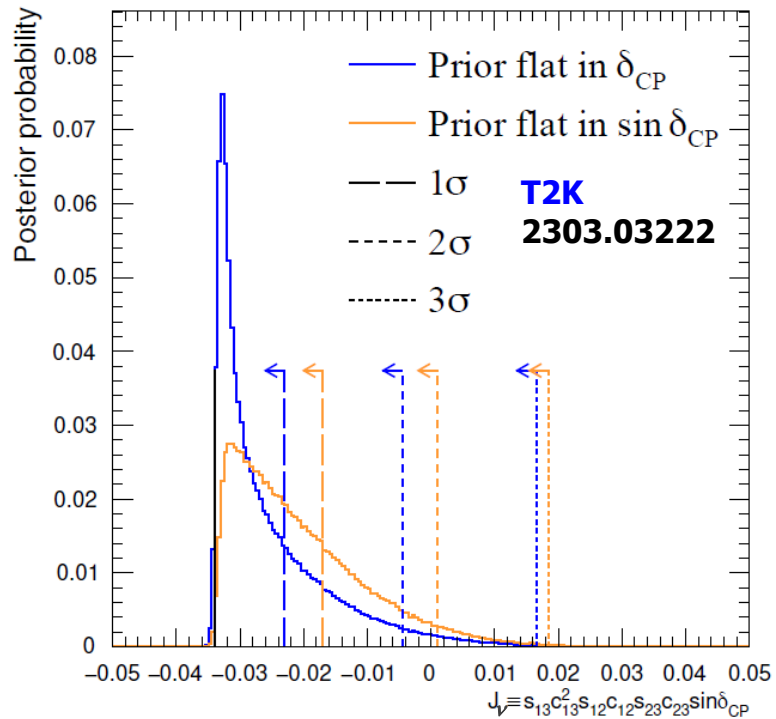
9 angles + 6 phases



You may calculate **everything** that can in principle be measured, in terms of **18** seesaw parameters.

Two kinds of fundamental CP violation

◆ **3-flavor ν -oscillations are established and a 2σ hint for CPV is achieved.**



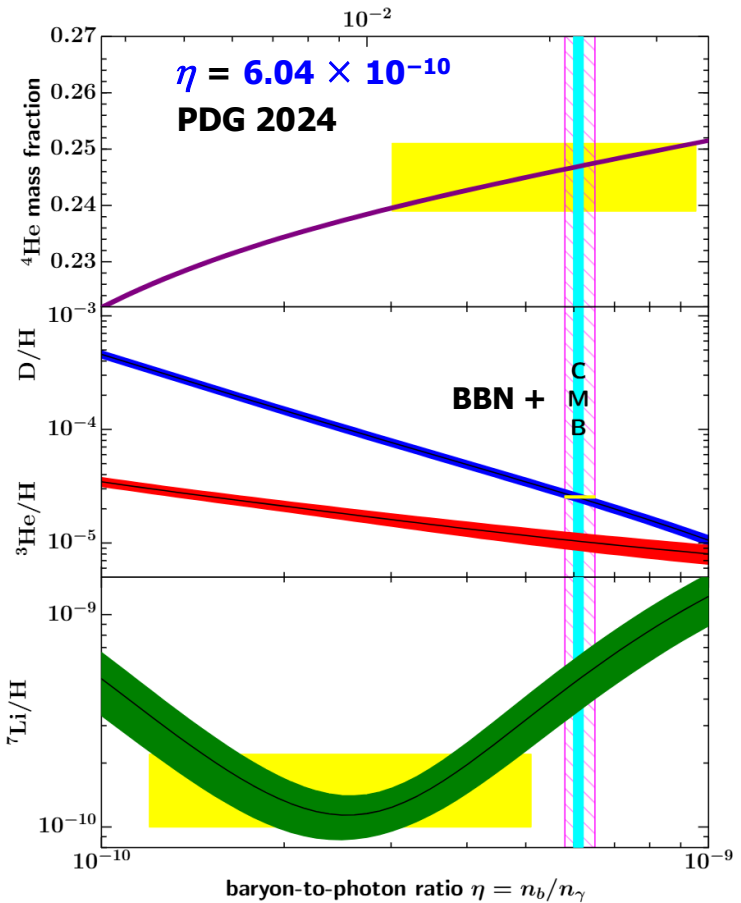
direct connection via seesaw + Leptogenesis ?

NO
W. Buchmüller
M. Plümacher:
hep-ph/9608308

YES
ZZX, 2406.01142



◆ **Cosmic CPV is already established.**



$$P(\nu_\mu \rightarrow \nu_e) = -4 \sum_{i < j} \left(\mathcal{R}_{ij} \sin^2 \frac{\Delta_{ji} L}{4E} \right) - 8 \mathcal{J}_\nu \prod_{i < j} \sin \frac{\Delta_{ji} L}{4E}$$

A bridge between light and heavy

- ◆ The exact seesaw formula — a bridge between the **original** and **derivational** flavor parameters:

$$UD_\nu U^T + RD_N R^T = 0 \longrightarrow M_\nu \equiv U_0 D_\nu U_0^T = (iA^{-1}R) D_N (iA^{-1}R)^T$$

Degrees of freedom (**mass** + **mixing angle** + **CPV phase**): **3 + 3 + 3 (derivational)** ← **3 + 9 + 6 (original)**

- ◆ The **Jarlskog** invariant of CP violation in **ν -oscillations**:

$$\mathcal{J}_\nu \equiv \text{Im} \left[(U_0)_{\alpha i} (U_0)_{\beta i'} (U_0)_{\alpha i'}^* (U_0)_{\beta i}^* \right]$$

(α, β) run cyclically over (e, μ, τ) , (i, i') run cyclically over $(1, 2, 3)$

$$\begin{cases} D_\nu \equiv \text{Diag}\{m_1, m_2, m_3\} \\ D_N \equiv \text{Diag}\{M_4, M_5, M_6\} \\ \Delta_{ii'} \equiv m_i^2 - m_{i'}^2 \end{cases}$$

- ◆ On the one hand, we use the **light degrees of freedom** to get the relation (**ZZX**, 2306.02362)

$$\mathcal{J}_\nu = \frac{\text{Im} \left[(M_\nu M_\nu^\dagger)_{e\mu} (M_\nu M_\nu^\dagger)_{\mu\tau} (M_\nu M_\nu^\dagger)_{\tau e} \right]}{\Delta_{21} \Delta_{31} \Delta_{32}} \longleftarrow \text{already measured}$$

- ◆ On the other hand, we use the **original seesaw-related parameters** to calculate the same quantity *in the leading order approximation* of $A^{-1}R$, because the **non-unitarity** of U characterized by $R \neq 0$ has been well constrained by precision measurements (**M. Blennow** et al, 2306.01040)

$$\left. \vphantom{\text{On the other hand, we use the original seesaw-related parameters}} \right\} A^{-1}R \simeq \begin{pmatrix} \hat{S}_{14}^* & \hat{S}_{15}^* & \hat{S}_{16}^* \\ \hat{S}_{24}^* & \hat{S}_{25}^* & \hat{S}_{26}^* \\ \hat{S}_{34}^* & \hat{S}_{35}^* & \hat{S}_{36}^* \end{pmatrix}$$

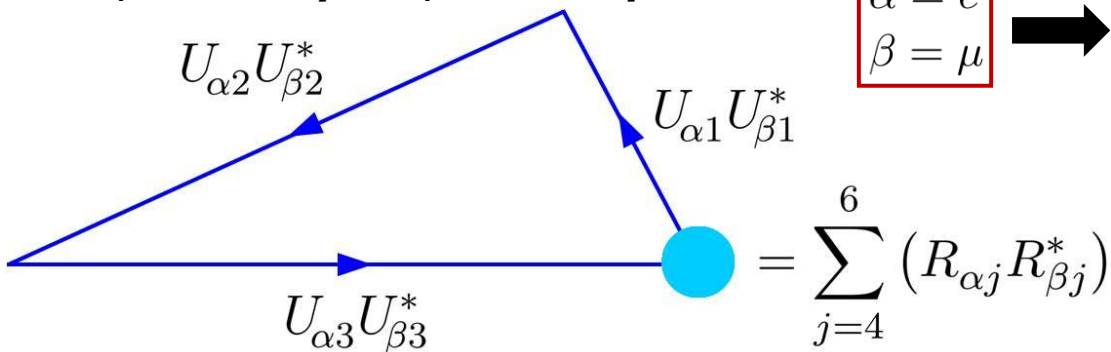
Is this approximation really safe?

◆ Of course, one may use the **non-unitary PMNS** matrix $U = AU_0$ to define the more general *Jarlskog* invariants to describe CP violation in neutrino oscillations. But one can show that their **leading terms** are the same, coming from the **unitarity limit** (ZZX, 1110.0083):

$$\mathcal{J}_{\alpha\beta}^{ii'} = \mathcal{J}_\nu + \text{corrections}$$

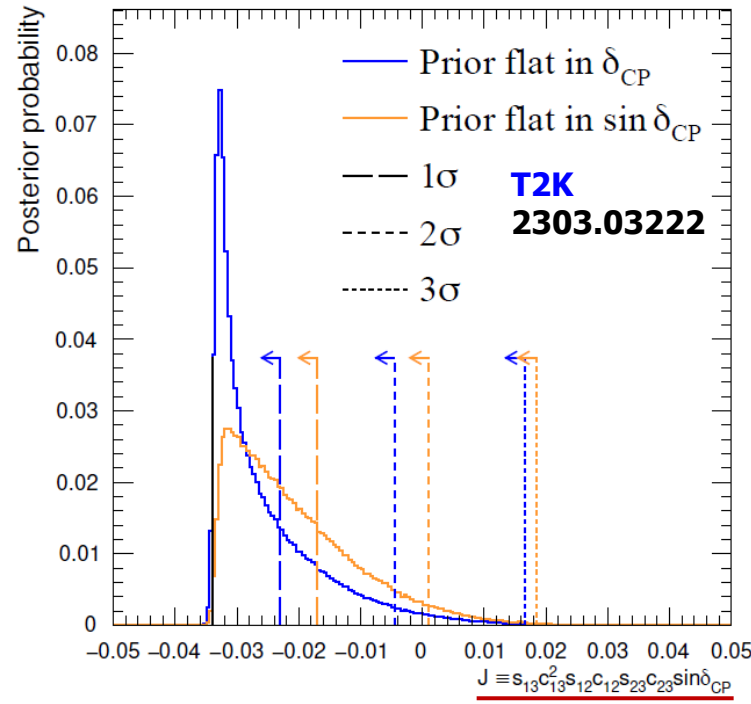
\uparrow $\leq 1\%$ \uparrow $< 0.01\%$

◆ Yes, absolutely safe, at least by 2044!



$$\mathcal{J}_{\alpha\beta}^{ii'} \equiv \text{Im} (U_{\alpha i} U_{\beta i'} U_{\alpha i'}^* U_{\beta i}^*)$$

(α, β) run cyclically over (e, μ, τ)
 (i, i') run cyclically over $(1, 2, 3)$



How many terms to be calculated?

9

◆ Let us classify the analytical results in terms of the products of **heavy Majorana neutrino masses**.

$$\text{Im} \left[(M_\nu M_\nu^\dagger)_{e\mu} (M_\nu M_\nu^\dagger)_{\mu\tau} (M_\nu M_\nu^\dagger)_{\tau e} \right]$$

CPV from 2-family or (and) 3-family interferences

Term 6 0 0: M_4^6, M_5^6, M_6^6

Term 5 1 0: $M_4^5 M_5, M_4^5 M_6, M_5^5 M_4, M_5^5 M_6, M_6^5 M_4, M_6^5 M_5$

Term 4 2 0: $M_4^4 M_5^2, M_4^4 M_6^2, M_5^4 M_4^2, M_5^4 M_6^2, M_6^4 M_4^2, M_6^4 M_5^2$

Term 3 3 0: $M_4^3 M_5^3, M_4^3 M_6^3, M_5^3 M_6^3$

Term 4 1 1: $M_4^4 M_5 M_6, M_5^4 M_4 M_6, M_6^4 M_4 M_5$

Term 3 2 1: $M_4^3 M_5^2 M_6, M_4^3 M_6^2 M_5, M_5^3 M_4^2 M_6, M_5^3 M_6^2 M_4, M_6^3 M_4^2 M_5, M_6^3 M_5^2 M_4$

Term 2 2 2: $M_4^2 M_5^2 M_6^2$

◆ There are totally **6** independent **original CP-violating phases** in the canonical seesaw mechanism, measuring the **inter-family interference effects** in all processes of heavy and light Majorana neutrinos.

$$I_{jk} \equiv \sum_{i=1}^3 \hat{s}_{ij}^* \hat{s}_{ik} = I_{kj}^*, \quad (j, k = 4, 5, 6)$$

$$(M_\nu M_\nu^\dagger)_{e\mu} = \sum_{j=4}^6 \sum_{k=4}^6 M_j M_k I_{jk} \hat{s}_{1j}^* \hat{s}_{2k}$$

×

$$(M_\nu M_\nu^\dagger)_{\mu\tau} = \sum_{j=4}^6 \sum_{k=4}^6 M_j M_k I_{jk} \hat{s}_{2j}^* \hat{s}_{3k}$$

×

$$(M_\nu M_\nu^\dagger)_{\tau e} = \sum_{j=4}^6 \sum_{k=4}^6 M_j M_k I_{jk} \hat{s}_{3j}^* \hat{s}_{1k}$$

✓

✓

$$\begin{cases} \alpha_i \equiv \delta_{i4} - \delta_{i5} \\ \beta_i \equiv \delta_{i5} - \delta_{i6} \\ \gamma_i \equiv \delta_{i6} - \delta_{i4} \end{cases} \quad \alpha_i + \beta_i + \gamma_i = 0 \quad (\text{for } i = 1, 2, 3)$$

◆ The **general** and **explicit** expression of the **Jarlskog invariant** in the seesaw mechanism is a linear combination of the above **5** terms:

$$\mathcal{J}_\nu = \frac{T_{33} + T_{42} + T_{411} + T_{321} + T_{222}}{\Delta_{21}\Delta_{31}\Delta_{32}}$$

◆ The two **2-family interference** terms are:

$$T_{33} = \underbrace{M_4^3 M_5^3}_{\text{blue}} \left(I_{44} I_{55} - |I_{45}|^2 \right) \left[\sum_{i=1}^3 s_{i4} s_{i5} s_{i'4} s_{i'5} \left(s_{i4}^2 s_{i''5}^2 + s_{i''4}^2 s_{i'5}^2 - s_{i'4}^2 s_{i''5}^2 - s_{i''4}^2 s_{i'5}^2 \right) \sin(\alpha_i + \alpha_{i'}) \right. \\ \left. - \sum_{i=1}^3 s_{i4}^2 s_{i5}^2 \left(s_{i'4}^2 s_{i''5}^2 - s_{i''4}^2 s_{i'5}^2 \right) \sin 2\alpha_i \right]$$

i, i' and i'' run cyclically over 1, 2 and 3

+ term{4 → 5, 5 → 6; α_i → β_i} + term{4 → 6, 5 → 4; α_i → γ_i}

$$T_{42} = \underbrace{M_4^2 M_5^2}_{\text{blue}} \left(I_{44} I_{55} - |I_{45}|^2 \right) \left[\sum_{i=1}^3 s_{i4} s_{i5} s_{i'4} s_{i'5} \left(\underbrace{M_4^2 I_{44}}_{\text{blue}} s_{i''4}^2 - \underbrace{M_5^2 I_{55}}_{\text{blue}} s_{i''5}^2 \right) \sin(\alpha_i - \alpha_{i'}) \right]$$

+ term{4 → 5, 5 → 6; α_i → β_i} + term{4 → 6, 5 → 4; α_i → γ_i}

Switching off the **3rd** heavy neutrino species "**6**", we can immediately arrive at the results in the **minimal seesaw** case (**ZZX**, 2306.02362):



9 combinations of **3** original CPV phases

◆ The simplest **3-family interference** term is obtained as follows:

$$\begin{aligned}
 T_{411} = & \underline{M_4^4 M_5 M_6} I_{44} s_{14} s_{24} s_{34} \left[-s_{14} s_{24} s_{34} \sum_{i=1}^3 s_{i5}^2 \left[\underline{s_{i'6}^2 \sin 2(\alpha_i + \gamma_{i'})} - \underline{s_{i''6}^2 \sin 2(\alpha_i + \gamma_{i''})} \right] \right. \\
 & + \sum_{i=1}^3 s_{i4} (s_{i'4}^2 - s_{i''4}^2) \left[\underline{s_{i5}^2 s_{i'6} s_{i''6} \sin(2\alpha_i + \gamma_{i'} + \gamma_{i''})} - \underline{s_{i6}^2 s_{i'5} s_{i''5} \sin(\alpha_{i'} + \alpha_{i''} + 2\gamma_i)} \right] \\
 & + \sum_{i=1}^3 s_{i4} s_{i'5} s_{i'6} (s_{i4}^2 + s_{i''4}^2) \left[\underline{s_{i'5} s_{i''6} \sin(\alpha_{i'} - \beta_{i'} + \gamma_{i''})} - \underline{s_{i'6} s_{i''5} \sin(\alpha_{i''} - \beta_{i'} + \gamma_{i'})} \right] \\
 & + \sum_{i=1}^3 s_{i4} s_{i''5} s_{i''6} (s_{i4}^2 + s_{i'4}^2) \left[\underline{s_{i'5} s_{i''6} \sin(\alpha_{i'} - \beta_{i''} + \gamma_{i''})} - \underline{s_{i'6} s_{i''5} \sin(\alpha_{i''} - \beta_{i''} + \gamma_{i'})} \right] \\
 & \left. + \sum_{i=1}^3 2s_{i4} s_{i5} s_{i6} (s_{i'4}^2 + s_{i''4}^2) \left[\underline{s_{i'5} s_{i''6} \sin(\alpha_{i'} - \beta_i + \gamma_{i''})} - \underline{s_{i'6} s_{i''5} \sin(\alpha_{i''} - \beta_i + \gamma_{i'})} \right] \right]
 \end{aligned}$$

+ term $\{ (4, 5, 6) \rightarrow (5, 4, 6); (\alpha_i, \beta_i, \gamma_i) \rightarrow -(\alpha_i, \gamma_i, \beta_i) \}$

+ term $\{ (4, 5, 6) \rightarrow (6, 5, 4); (\alpha_i, \beta_i, \gamma_i) \rightarrow -(\beta_i, \alpha_i, \gamma_i) \}$

The terms $T_{321} + T_{222}$ are very complicated and can be found in **ZZX, 2406.01142**.

Counting the phase combinations (1)

◆ After a very tedious survey of all terms of the **Jarlskog invariant**, we find **240** linear combinations of the **6 original seesaw phase parameters** — **72** of them:

$$\sin \alpha_1, \quad \sin \alpha_2, \quad \sin \alpha_3; \quad \sin \beta_1, \quad \sin \beta_2, \quad \sin \beta_3; \quad \sin \gamma_1, \quad \sin \gamma_2, \quad \sin \gamma_3;$$

$$\sin 2\alpha_1, \quad \sin 2\alpha_2, \quad \sin 2\alpha_3; \quad \sin 2\beta_1, \quad \sin 2\beta_2, \quad \sin 2\beta_3; \quad \sin 2\gamma_1, \quad \sin 2\gamma_2, \quad \sin 2\gamma_3;$$

$$\sin(\alpha_1 + \alpha_2), \quad \sin(\alpha_2 + \alpha_3), \quad \sin(\alpha_3 + \alpha_1); \quad \sin(\beta_1 + \beta_2), \quad \sin(\beta_2 + \beta_3), \quad \sin(\beta_3 + \beta_1);$$

$$\sin(\gamma_1 + \gamma_2), \quad \sin(\gamma_2 + \gamma_3), \quad \sin(\gamma_3 + \gamma_1);$$

$$\sin(\alpha_1 - \alpha_2), \quad \sin(\alpha_2 - \alpha_3), \quad \sin(\alpha_3 - \alpha_1); \quad \sin(\beta_1 - \beta_2), \quad \sin(\beta_2 - \beta_3), \quad \sin(\beta_3 - \beta_1);$$

$$\sin(\gamma_1 - \gamma_2), \quad \sin(\gamma_2 - \gamma_3), \quad \sin(\gamma_3 - \gamma_1);$$

$$\sin(\alpha_1 + \beta_2), \quad \sin(\alpha_1 + \beta_3), \quad \sin(\alpha_2 + \beta_1), \quad \sin(\alpha_2 + \beta_3), \quad \sin(\alpha_3 + \beta_1), \quad \sin(\alpha_3 + \beta_2);$$

$$\sin(\alpha_1 + \gamma_2), \quad \sin(\alpha_1 + \gamma_3), \quad \sin(\alpha_2 + \gamma_1), \quad \sin(\alpha_2 + \gamma_3), \quad \sin(\alpha_3 + \gamma_1), \quad \sin(\alpha_3 + \gamma_2);$$

$$\sin(\beta_1 + \gamma_2), \quad \sin(\beta_1 + \gamma_3), \quad \sin(\beta_2 + \gamma_1), \quad \sin(\beta_2 + \gamma_3), \quad \sin(\beta_3 + \gamma_1), \quad \sin(\beta_3 + \gamma_2);$$

$$\sin 2(\alpha_1 + \beta_2), \quad \sin 2(\alpha_1 + \beta_3), \quad \sin 2(\alpha_2 + \beta_1), \quad \sin 2(\alpha_2 + \beta_3), \quad \sin 2(\alpha_3 + \beta_1), \quad \sin 2(\alpha_3 + \beta_2);$$

$$\sin 2(\alpha_1 + \gamma_2), \quad \sin 2(\alpha_1 + \gamma_3), \quad \sin 2(\alpha_2 + \gamma_1), \quad \sin 2(\alpha_2 + \gamma_3), \quad \sin 2(\alpha_3 + \gamma_1), \quad \sin 2(\alpha_3 + \gamma_2);$$

$$\sin 2(\beta_1 + \gamma_2), \quad \sin 2(\beta_1 + \gamma_3), \quad \sin 2(\beta_2 + \gamma_1), \quad \sin 2(\beta_2 + \gamma_3), \quad \sin 2(\beta_3 + \gamma_1), \quad \sin 2(\beta_3 + \gamma_2);$$

◆ Of **240** linear combinations of the **6 original seesaw phase parameters** — **60** of them:

$$\sin(2\alpha_1 - \beta_2), \quad \sin(2\alpha_1 - \beta_3), \quad \sin(2\alpha_2 - \beta_1), \quad \sin(2\alpha_2 - \beta_3), \quad \sin(2\alpha_3 - \beta_1), \quad \sin(2\alpha_3 - \beta_2);$$

$$\sin(2\alpha_1 - \gamma_2), \quad \sin(2\alpha_1 - \gamma_3), \quad \sin(2\alpha_2 - \gamma_1), \quad \sin(2\alpha_2 - \gamma_3), \quad \sin(2\alpha_3 - \gamma_1), \quad \sin(2\alpha_3 - \gamma_2);$$

$$\sin(2\beta_1 - \alpha_2), \quad \sin(2\beta_1 - \alpha_3), \quad \sin(2\beta_2 - \alpha_1), \quad \sin(2\beta_2 - \alpha_3), \quad \sin(2\beta_3 - \alpha_1), \quad \sin(2\beta_3 - \alpha_2);$$

$$\sin(2\beta_1 - \gamma_2), \quad \sin(2\beta_1 - \gamma_3), \quad \sin(2\beta_2 - \gamma_1), \quad \sin(2\beta_2 - \gamma_3), \quad \sin(2\beta_3 - \gamma_1), \quad \sin(2\beta_3 - \gamma_2);$$

$$\sin(2\gamma_1 - \alpha_2), \quad \sin(2\gamma_1 - \alpha_3), \quad \sin(2\gamma_2 - \alpha_1), \quad \sin(2\gamma_2 - \alpha_3), \quad \sin(2\gamma_3 - \alpha_1), \quad \sin(2\gamma_3 - \alpha_2);$$

$$\sin(2\gamma_1 - \beta_2), \quad \sin(2\gamma_1 - \beta_3), \quad \sin(2\gamma_2 - \beta_1), \quad \sin(2\gamma_2 - \beta_3), \quad \sin(2\gamma_3 - \beta_1), \quad \sin(2\gamma_3 - \beta_2);$$

$$\sin(\alpha_1 + \alpha_2 + 2\beta_3), \quad \sin(\alpha_1 + \alpha_2 + 2\gamma_3), \quad \sin(\alpha_1 + \alpha_3 + 2\beta_2), \quad \sin(\alpha_1 + \alpha_3 + 2\gamma_2),$$

$$\sin(\alpha_2 + \alpha_3 + 2\beta_1), \quad \sin(\alpha_2 + \alpha_3 + 2\gamma_1); \quad \sin(\beta_1 + \beta_2 + 2\alpha_3), \quad \sin(\beta_1 + \beta_2 + 2\gamma_3),$$

$$\sin(\beta_1 + \beta_3 + 2\alpha_2), \quad \sin(\beta_1 + \beta_3 + 2\gamma_2), \quad \sin(\beta_2 + \beta_3 + 2\alpha_1), \quad \sin(\beta_2 + \beta_3 + 2\gamma_1);$$

$$\sin(\gamma_1 + \gamma_2 + 2\alpha_3), \quad \sin(\gamma_1 + \gamma_2 + 2\beta_3), \quad \sin(\gamma_1 + \gamma_3 + 2\alpha_2), \quad \sin(\gamma_1 + \gamma_3 + 2\beta_2),$$

$$\sin(\gamma_2 + \gamma_3 + 2\alpha_1), \quad \sin(\gamma_2 + \gamma_3 + 2\beta_1);$$

$$\sin(\alpha_1 + \beta_2 + \gamma_3), \quad \sin(\alpha_1 + \beta_3 + \gamma_2), \quad \sin(\alpha_2 + \beta_1 + \gamma_3), \quad \sin(\alpha_2 + \beta_3 + \gamma_1),$$

$$\sin(\alpha_3 + \beta_1 + \gamma_2), \quad \sin(\alpha_3 + \beta_2 + \gamma_1).$$

Counting the phase combinations (3)

◆ Of **240** linear combinations of the **6 original seesaw phase parameters** — **54** of them:

$$\begin{aligned} & \sin(\alpha_1 + \alpha_2 - \beta_1), \quad \sin(\alpha_1 + \alpha_2 - \beta_2), \quad \sin(\alpha_1 + \alpha_2 - \beta_3), \quad \sin(\alpha_1 + \alpha_3 - \beta_1), \\ & \sin(\alpha_1 + \alpha_3 - \beta_2), \quad \sin(\alpha_1 + \alpha_3 - \beta_3), \quad \sin(\alpha_2 + \alpha_3 - \beta_1), \quad \sin(\alpha_2 + \alpha_3 - \beta_2), \\ & \sin(\alpha_2 + \alpha_3 - \beta_3); \quad \sin(\alpha_1 + \alpha_2 - \gamma_1), \quad \sin(\alpha_1 + \alpha_2 - \gamma_2), \quad \sin(\alpha_1 + \alpha_2 - \gamma_3), \\ & \sin(\alpha_1 + \alpha_3 - \gamma_1), \quad \sin(\alpha_1 + \alpha_3 - \gamma_2), \quad \sin(\alpha_1 + \alpha_3 - \gamma_3), \quad \sin(\alpha_2 + \alpha_3 - \gamma_1), \\ & \sin(\alpha_2 + \alpha_3 - \gamma_2), \quad \sin(\alpha_2 + \alpha_3 - \gamma_3); \quad \sin(\beta_1 + \beta_2 - \alpha_1), \quad \sin(\beta_1 + \beta_2 - \alpha_2), \\ & \sin(\beta_1 + \beta_2 - \alpha_3), \quad \sin(\beta_1 + \beta_3 - \alpha_1), \quad \sin(\beta_1 + \beta_3 - \alpha_2), \quad \sin(\beta_1 + \beta_3 - \alpha_3), \\ & \sin(\beta_2 + \beta_3 - \alpha_1), \quad \sin(\beta_2 + \beta_3 - \alpha_2), \quad \sin(\beta_2 + \beta_3 - \alpha_3); \quad \sin(\beta_1 + \beta_2 - \gamma_1), \\ & \sin(\beta_1 + \beta_2 - \gamma_2), \quad \sin(\beta_1 + \beta_2 - \gamma_3), \quad \sin(\beta_1 + \beta_3 - \gamma_1), \quad \sin(\beta_1 + \beta_3 - \gamma_2), \\ & \sin(\beta_1 + \beta_3 - \gamma_3), \quad \sin(\beta_2 + \beta_3 - \gamma_1), \quad \sin(\beta_2 + \beta_3 - \gamma_2), \quad \sin(\beta_2 + \beta_3 - \gamma_3); \\ & \sin(\gamma_1 + \gamma_2 - \alpha_1), \quad \sin(\gamma_1 + \gamma_2 - \alpha_2), \quad \sin(\gamma_1 + \gamma_2 - \alpha_3), \quad \sin(\gamma_1 + \gamma_3 - \alpha_1), \\ & \sin(\gamma_1 + \gamma_3 - \alpha_2), \quad \sin(\gamma_1 + \gamma_3 - \alpha_3), \quad \sin(\gamma_2 + \gamma_3 - \alpha_1), \quad \sin(\gamma_2 + \gamma_3 - \alpha_2), \\ & \sin(\gamma_2 + \gamma_3 - \alpha_3); \quad \sin(\gamma_1 + \gamma_2 - \beta_1), \quad \sin(\gamma_1 + \gamma_2 - \beta_2), \quad \sin(\gamma_1 + \gamma_2 - \beta_3), \\ & \sin(\gamma_1 + \gamma_3 - \beta_1), \quad \sin(\gamma_1 + \gamma_3 - \beta_2), \quad \sin(\gamma_1 + \gamma_3 - \beta_3), \quad \sin(\gamma_2 + \gamma_3 - \beta_1), \\ & \sin(\gamma_2 + \gamma_3 - \beta_2), \quad \sin(\gamma_2 + \gamma_3 - \beta_3); \end{aligned}$$

◆ Of **240** linear combinations of the **6 original seesaw phase parameters** — **54** of them:

$$\begin{aligned}
 & \sin(\alpha_1 + \beta_2 - \gamma_1), \quad \sin(\alpha_1 + \beta_2 - \gamma_2), \quad \sin(\alpha_1 + \beta_2 - \gamma_3), \quad \sin(\alpha_1 + \beta_3 - \gamma_1), \\
 & \sin(\alpha_1 + \beta_3 - \gamma_2), \quad \sin(\alpha_1 + \beta_3 - \gamma_3), \quad \sin(\alpha_2 + \beta_1 - \gamma_1), \quad \sin(\alpha_2 + \beta_1 - \gamma_2), \\
 & \sin(\alpha_2 + \beta_1 - \gamma_3), \quad \sin(\alpha_2 + \beta_3 - \gamma_1), \quad \sin(\alpha_2 + \beta_3 - \gamma_2), \quad \sin(\alpha_2 + \beta_3 - \gamma_3), \\
 & \sin(\alpha_3 + \beta_1 - \gamma_1), \quad \sin(\alpha_3 + \beta_1 - \gamma_2), \quad \sin(\alpha_3 + \beta_1 - \gamma_3), \quad \sin(\alpha_3 + \beta_2 - \gamma_1), \\
 & \sin(\alpha_3 + \beta_2 - \gamma_2), \quad \sin(\alpha_3 + \beta_2 - \gamma_3); \quad \sin(\alpha_1 - \beta_1 + \gamma_2), \quad \sin(\alpha_1 - \beta_1 + \gamma_3), \\
 & \sin(\alpha_1 - \beta_2 + \gamma_2), \quad \sin(\alpha_1 - \beta_2 + \gamma_3), \quad \sin(\alpha_1 - \beta_3 + \gamma_2), \quad \sin(\alpha_1 - \beta_3 + \gamma_3), \\
 & \sin(\alpha_2 - \beta_1 + \gamma_1), \quad \sin(\alpha_2 - \beta_1 + \gamma_3), \quad \sin(\alpha_2 - \beta_2 + \gamma_1), \quad \sin(\alpha_2 - \beta_2 + \gamma_3), \\
 & \sin(\alpha_2 - \beta_3 + \gamma_1), \quad \sin(\alpha_2 - \beta_3 + \gamma_3), \quad \sin(\alpha_3 - \beta_1 + \gamma_1), \quad \sin(\alpha_3 - \beta_1 + \gamma_2), \\
 & \sin(\alpha_3 - \beta_2 + \gamma_1), \quad \sin(\alpha_3 - \beta_2 + \gamma_2), \quad \sin(\alpha_3 - \beta_3 + \gamma_1), \quad \sin(\alpha_3 - \beta_3 + \gamma_2); \\
 & \sin(\alpha_1 - \beta_1 - \gamma_2), \quad \sin(\alpha_1 - \beta_1 - \gamma_3), \quad \sin(\alpha_1 - \beta_2 - \gamma_1), \quad \sin(\alpha_1 - \beta_2 - \gamma_3), \\
 & \sin(\alpha_1 - \beta_3 - \gamma_1), \quad \sin(\alpha_1 - \beta_3 - \gamma_2), \quad \sin(\alpha_2 - \beta_1 - \gamma_2), \quad \sin(\alpha_2 - \beta_1 - \gamma_3), \\
 & \sin(\alpha_2 - \beta_2 - \gamma_1), \quad \sin(\alpha_2 - \beta_2 - \gamma_3), \quad \sin(\alpha_2 - \beta_3 - \gamma_1), \quad \sin(\alpha_2 - \beta_3 - \gamma_2), \\
 & \sin(\alpha_3 - \beta_1 - \gamma_2), \quad \sin(\alpha_3 - \beta_1 - \gamma_3), \quad \sin(\alpha_3 - \beta_2 - \gamma_1), \quad \sin(\alpha_3 - \beta_2 - \gamma_3), \\
 & \sin(\alpha_3 - \beta_3 - \gamma_1), \quad \sin(\alpha_3 - \beta_3 - \gamma_2).
 \end{aligned}$$

- ◆ The flavor-dependent CP-violating asymmetries in **LNV** decays of heavy Majorana neutrinos:

$$\varepsilon_{j\alpha} \equiv \frac{\Gamma(N_j \rightarrow \ell_\alpha + H) - \Gamma(N_j \rightarrow \bar{\ell}_\alpha + \bar{H})}{\sum_\alpha [\Gamma(N_j \rightarrow \ell_\alpha + H) + \Gamma(N_j \rightarrow \bar{\ell}_\alpha + \bar{H})]}$$

$$\simeq \frac{1}{8\pi \langle H \rangle^2 \sum_\beta |R_{\beta j}|^2} \sum_{k=4}^6 \left\{ M_k^2 \text{Im} \left[(R_{\alpha j}^* R_{\alpha k}) \sum_\beta \left[(R_{\beta j}^* R_{\beta k}) \xi(x_{kj}) + (R_{\beta j} R_{\beta k}^*) \zeta(x_{kj}) \right] \right] \right\}$$

- ◆ **Baryogenesis via leptogenesis** in the early Universe:

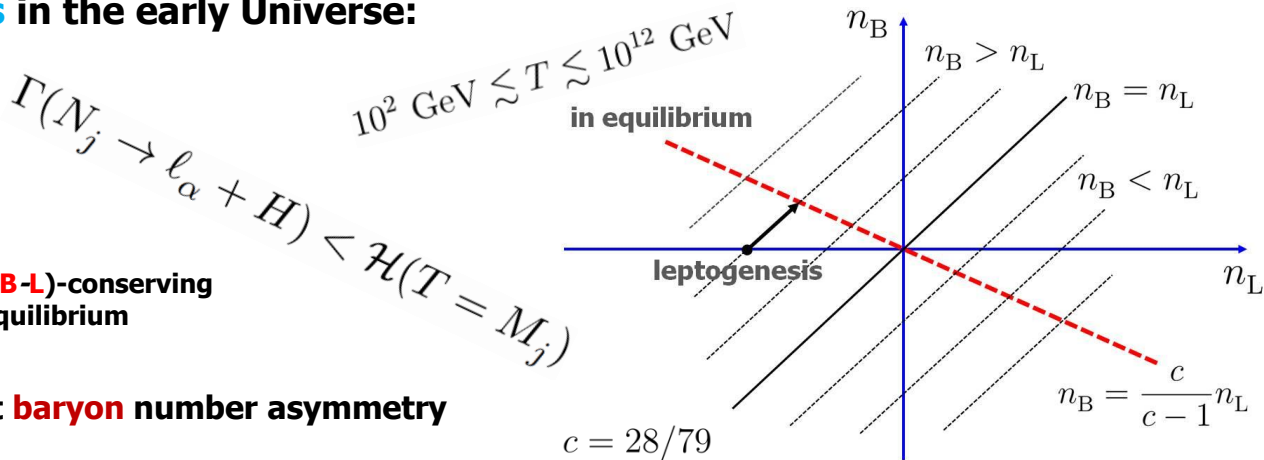
A net **lepton** number asymmetry

$$Y_L \equiv \frac{n_L - n_{\bar{L}}}{s} = \frac{1}{g_*} \sum_{j,\alpha} \kappa_{j\alpha} \varepsilon_{j\alpha}$$

sphaleron-induced (**B-L**)-conserving process in thermal equilibrium

$$Y_B \equiv \frac{n_B - n_{\bar{B}}}{s} = -c Y_L$$

A net **baryon** number asymmetry



How many phase combinations (1)

- ◆ A **flavor-dependent** CP-violating asymmetry of the first ($j = 4$) heavy Majorana neutrino decays:

$$\begin{aligned} \varepsilon_{4e} = \frac{1}{8\pi\langle H \rangle^2 I_{44}} & \left\{ M_5^2 s_{14} s_{15} \left[\xi(x_{54}) \sum_{i=1}^3 s_{i4} s_{i5} \sin(\alpha_1 + \alpha_i) + \zeta(x_{54}) \left[s_{24} s_{25} \sin(\alpha_1 - \alpha_2) \right. \right. \right. \\ & \left. \left. \left. + s_{34} s_{35} \sin(\alpha_1 - \alpha_3) \right] \right] - M_6^2 s_{14} s_{16} \left[\xi(x_{64}) \sum_{i=1}^3 s_{i4} s_{i6} \sin(\gamma_1 + \gamma_i) \right. \right. \\ & \left. \left. \left. + \zeta(x_{64}) \left[s_{24} s_{26} \sin(\gamma_1 - \gamma_2) + s_{34} s_{36} \sin(\gamma_1 - \gamma_3) \right] \right] \right\} \end{aligned}$$

The formulas for "5" and "6" can be similarly written out.

$$\begin{aligned} \varepsilon_{4\mu} = \frac{1}{8\pi\langle H \rangle^2 I_{44}} & \left\{ M_5^2 s_{24} s_{25} \left[\xi(x_{54}) \sum_{i=1}^3 s_{i4} s_{i5} \sin(\alpha_2 + \alpha_i) + \zeta(x_{54}) \left[s_{14} s_{15} \sin(\alpha_2 - \alpha_1) \right. \right. \right. \\ & \left. \left. \left. + s_{34} s_{35} \sin(\alpha_2 - \alpha_3) \right] \right] - M_6^2 s_{24} s_{26} \left[\xi(x_{64}) \sum_{i=1}^3 s_{i4} s_{i6} \sin(\gamma_2 + \gamma_i) \right. \right. \\ & \left. \left. \left. + \zeta(x_{64}) \left[s_{14} s_{16} \sin(\gamma_2 - \gamma_1) + s_{34} s_{36} \sin(\gamma_2 - \gamma_3) \right] \right] \right\} \end{aligned}$$

The formulas for "5" and "6" can be similarly written out.

- ◆ A **flavor-dependent** CP-violating asymmetry of the first ($j = 4$) heavy Majorana neutrino decays:

$$\varepsilon_{4\tau} = \frac{1}{8\pi\langle H \rangle^2 I_{44}} \left\{ M_5^2 s_{34} s_{35} \left[\xi(x_{54}) \sum_{i=1}^3 s_{i4} s_{i5} \sin(\alpha_3 + \alpha_i) + \zeta(x_{54}) \left[s_{14} s_{15} \sin(\alpha_3 - \alpha_1) + s_{24} s_{25} \sin(\alpha_3 - \alpha_2) \right] \right] - M_6^2 s_{34} s_{36} \left[\xi(x_{64}) \sum_{i=1}^3 s_{i4} s_{i6} \sin(\gamma_3 + \gamma_i) + \zeta(x_{64}) \left[s_{14} s_{16} \sin(\gamma_3 - \gamma_1) + s_{24} s_{26} \sin(\gamma_3 - \gamma_2) \right] \right] \right\}$$

The formulas for "5" and "6" can be similarly written out.

- ◆ The **flavor-independent** CP-violating asymmetry $\varepsilon_4 \equiv \varepsilon_{4e} + \varepsilon_{4\mu} + \varepsilon_{4\tau}$, for example:

$$\varepsilon_4 = \frac{1}{8\pi\langle H \rangle^2 I_{44}} \left[\sum_{i=1}^3 \sum_{i'=1}^3 s_{i4} s_{i'4} \left[M_5^2 \xi(x_{54}) s_{i5} s_{i'5} \sin(\alpha_i + \alpha_{i'}) - M_6^2 \xi(x_{64}) s_{i6} s_{i'6} \sin(\gamma_i + \gamma_{i'}) \right] \right]$$

- ◆ Totally **27** linear combinations of the **6 original seesaw phase parameters** in CP violation of three heavy Majorana neutrino decays ($i, i' = 1, 2, 3$):

$$\sin(\alpha_i \pm \alpha_{i'}), \sin(\beta_i \pm \beta_{i'}), \sin(\gamma_i \pm \gamma_{i'})$$

- ◆ The analytical results obtained above imply that the *CP-violating asymmetries* can be expressed as a linear combinations of the **sines** of the **6 original seesaw phase parameters**:

$$\begin{aligned}\varepsilon_{j\alpha} &= \sum_{i=1}^3 (C'_{\alpha i} \sin \alpha_i + C'_{\beta i} \sin \beta_i) \\ \varepsilon_j &= \sum_{i=1}^3 (C''_{\alpha i} \sin \alpha_i + C''_{\beta i} \sin \beta_i)\end{aligned}$$

It's then straightforward to extract the coefficients from the formulas of CP violating asymmetries.

- ◆ In comparison, the achieved result of the **Jarlskog invariant** implies that it can also be expressed as a linear combinations of the **sines** of the **6 original seesaw phase parameters**:

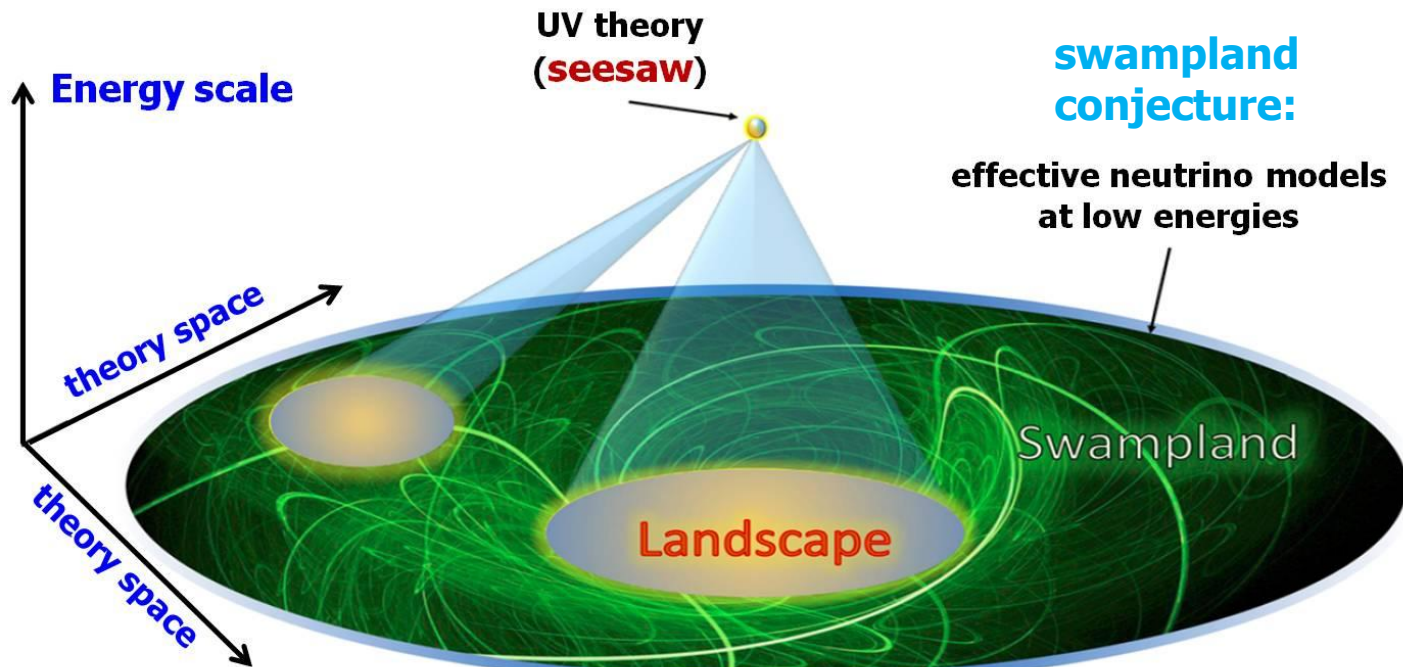
$$\mathcal{J}_\nu = \sum_{i=1}^3 (C_{\alpha i} \sin \alpha_i + C_{\beta i} \sin \beta_i)$$

It is straightforward to extract the **lengthy** coefficients from the $T_{33} + T_{42} + T_{411} + T_{321} + T_{222}$ terms, but the expressions are so complicated that they cannot be presented here.

◆ Extending the SM framework in a way that is as **natural** and **economical** as possible, we have argued that the **canonical seesaw mechanism** is most convincing to give mass to the active neutrinos. It is fully consistent with the spirit of **Weinberg's** EFT and thus should be located in **Vafa's landscape** of particle physics.



Cumrun Vafa 2005



- ◆ The **new era** of precision measurements, as characterized by **JUNO**, **DUNE** and **T2HK**, is coming. It is **high time** to experimentally test the **canonical seesaw** in a **systematical** and **model-independent** way at low energies.
- ◆ This becomes possible, with the help of a **complete Euler-like block** parametrization of the seesaw flavor structure, since it makes **analytical** calculations of all observables possible. The present talk give a **PoC** example by clarifying the **Buchmüller-Plümacher** claim. **For the first time**, we have shown that a **direct, explicit** and **model-independent** connection exists between CP violation at **high** and **low** energy scales.
- ◆ **A take-home message**: to really test the seesaw, you should calculate everything by using the **original seesaw parameters** instead of the **derivational ones** or a **mixture**.
- ◆ We are making progress in calculating all the other observables of **3 active neutrino oscillations** (in collaboration with **Jing-yu Zhu**).

THANK YOU