

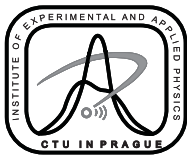
Magnetic Monopoles with Internal Structure

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The idea:

- Spectrum of certain gauge field theories contain magnetic monopoles ('t Hooft, Polyakov, 1974)
- In some of the theories the monopole solutions can be found even analytically
- Alas, number of these theories is rather limited. . .
- Our goal:
 - Enlarge the number of these theories (by resorting to effective field theories)
 - Perhaps the new exact solution will have interesting features. . .?

't Hooft–Polyakov monopole @ Georgi–Glashow model

(a short reminder)

Georgi–Glashow model

- Simplest model that allows for magnetic monopoles
⇒ Prototypical, if not archetypal, example
- Spontaneously broken $SU(2)$ theory with adjoint (real triplet) scalar:

$$\mathcal{L} = \frac{1}{2}(D^\mu \phi)^2 - \frac{1}{4g^2}(\mathbf{F}^{\mu\nu})^2 - \frac{1}{4}\lambda(\phi^2 - v^2)^2$$

where

$$\begin{aligned}D^\mu \phi &= \partial^\mu \phi + \mathbf{A}^\mu \times \phi \\ \mathbf{F}^{\mu\nu} &= \partial^\mu \mathbf{A}^\nu - \partial^\nu \mathbf{A}^\mu + \mathbf{A}^\mu \times \mathbf{A}^\nu\end{aligned}$$

- In the vacuum ($\phi^2 = v^2$) the $SU(2)$ is broken down to the “electromagnetic” $U(1)$
- Spectrum:
 - Perturbative: Photon, W^\pm , Higgs boson
 - Non-perturbative: Magnetic monopole. . .

The monopole solution

Static, spherically symmetric classical solution with magnetic charge

- The solution is guaranteed to exist (and be stable) on topological reasons
 - Since $\pi_2(SU(2)/U(1)) = \mathbb{Z} \neq \{1\}$
- Asymptotically, the solution approaches the vacuum and:
 - $SU(2)$ is broken down to the electromagnetic $U(1)$
 - Electromagnetic field can be defined
 - Magnetic field happens to be non-vanishing!
 - \Rightarrow magnetic charge $q_m = 4\pi$

Technically:

- The monopole solution is found using the “Hedgehog” Ansatz:

$$\phi_a = v \frac{x_a}{r} H(r), \quad \mathbf{A}_a^i = \frac{1}{r^2} \varepsilon_{iab} x_b (K(r) - 1), \quad \mathbf{A}_a^0 = 0$$

- The form factors $H(r)$, $K(r)$ are subject to 2nd-order ODEs
- \Rightarrow Difficult! \Rightarrow Solution can be found only numerically
- However, there is an exception. . .

The BPS limit:

- Limit of vanishing potential: $\lambda \rightarrow 0$
- But the boundary condition $\phi^2(\infty) = v^2$ is kept!

⇒ EOMs are just 1st-order!

- Analytical solution exists ($\rho \equiv vgr$):

$$K = \frac{\rho}{\sinh \rho}, \quad H = \coth \rho - \frac{1}{\rho}$$

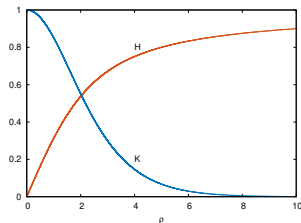
- Mass of the monopole:

$$M = \frac{4\pi v}{g}$$

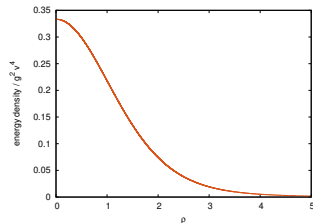
The lesson: The BPS limit is good – it makes the life **easier and happier!**

⇒ We will use it also in the following

Solution $K(\rho)$, $H(\rho)$:



Energy density $\mathcal{E}(\rho)$:



Effective extension

The idea

Recall the Georgi–Glashow model:

$$\mathcal{L} = \frac{1}{2}(D^\mu \phi)^2 - \frac{1}{4g^2}(\mathbf{F}^{\mu\nu})^2 - V(\phi^2)$$

The Lagrangian is a linear combination of

$$(D^\mu \phi)^2 \quad \text{and} \quad (\mathbf{F}^{\mu\nu})^2$$

→ The idea:

- Consider also the terms

$$(\phi \cdot D^\mu \phi)^2 \quad \text{and} \quad (\phi \cdot \mathbf{F}^{\mu\nu})^2$$

- Make the coefficients of this linear combination ϕ -dependent

In other words: We consider the most general model quadratic in $D^\mu \phi$ and $\mathbf{F}^{\mu\nu}$

⇒ Our generalized model:

$$\mathcal{L} = \frac{v^2}{2} \left[f_1^2 \left(\frac{(D^\mu \phi)^2}{\phi^2} - \frac{(\phi \cdot D^\mu \phi)^2}{\phi^4} \right) + f_3^2 \frac{(\phi \cdot D^\mu \phi)^2}{\phi^4} \right] \\ - \frac{1}{4g^2} \left[f_2^2 \left((\mathbf{F}^{\mu\nu})^2 - \frac{(\phi \cdot \mathbf{F}^{\mu\nu})^2}{\phi^2} \right) + f_4^2 \frac{(\phi \cdot \mathbf{F}^{\mu\nu})^2}{\phi^2} \right] - V(\phi^2)$$

- f_i^2 : Positive, dimensionless and gauge-invariant functions of ϕ
 - Technically: $\phi = vH\mathbf{n}$ (where $\mathbf{n}^2 = 1$) $\Rightarrow f_i^2 = f_i^2(H)$
- In fact, we consider a whole class of models:
 - Each set of the functions f_i^2 defines a particular theory
 - E.g., for $f_{1,3}^2 = H^2$ and $f_{2,4}^2 = 1$ we recover the Georgi–Glashow model
- Important: We modified only the interactions
 - ⇒ Topology remains the same as in the Georgi–Glashow model
 - ⇒ Monopole solutions are again present!
- (Unimportant: Negative powers of ϕ and the projector-like structures are just for convenience – see next slides...)

The BPS limit

How to achieve the BPS limit?

- Georgi–Glashow model: Only $V \rightarrow 0$
- The generalized model: Also $V \rightarrow 0$, but f_i must satisfy

$$f_3 f_4 = H \frac{d}{dH} (f_1 f_2)$$

Then the Bogomolny bound (lower energy bound of a static configuration) is saturated, when

- The EOM of the 1st-order is satisfied:

$$D^i \phi = \frac{H}{g} \left[\frac{f_2}{f_1} \left(\mathbf{B}^i - \frac{\phi \cdot \mathbf{B}^i}{\phi^2} \phi \right) + \frac{f_4}{f_3} \frac{\phi \cdot \mathbf{B}^i}{\phi^2} \phi \right] \quad (\text{where } \mathbf{B}^i \equiv -\frac{1}{2} \epsilon^{ijk} \mathbf{F}^{jk})$$

- By definition, if this eq. is satisfied, the energy density is a total derivative:

$$\mathcal{E} = \partial^i \left(\frac{f_1 f_2}{gH} \phi \cdot \mathbf{B}^i \right)$$

The BPS limit: Spherical symmetry

- Recall the spherically-symmetric “hedgehog” Ansatz:

$$\phi_a = v \frac{x_a}{r} H(r), \quad \mathbf{A}_a^i = \frac{1}{r^2} \varepsilon_{iab} x_b (K(r) - 1), \quad \mathbf{A}_a^0 = 0$$

- The EOMs follow as ($\rho \equiv vgr$):

$$\begin{aligned} \partial_\rho(\log K) &= -\frac{f_1}{f_2} \\ \partial_\rho(\log H) &= \frac{1 - K^2}{\rho^2} \frac{f_4}{f_3} \end{aligned}$$

- 1st order \Rightarrow There is a hope that, perhaps, they could be easily solvable...

Two classes of exact solutions

Indeed! There are (at least) two classes when the EOMs can be solved analytically:

1 If $\frac{f_3}{f_4} = H \frac{d}{dH} \left(\frac{f_1}{f_2} \right)$:

$$K = \frac{\rho}{\sinh \rho}$$
$$H = \left(\frac{f_1}{f_2} \right)^{-1} (\kappa) \quad \text{where } \kappa \equiv \coth \rho - \frac{1}{\rho}$$

→ The “t Hooft–Polyakov” class

2 If $f_1 = f_2$:

$$K = \xi \exp(-\rho)$$
$$H = \left(\int \frac{dH}{H} \frac{f_3}{f_4} \right)^{-1} (1 - \lambda) \quad \text{where } \lambda \equiv \frac{1}{\rho} - \xi^2 \left[\frac{e^{-2\rho}}{\rho} + 2\text{Ei}(-2\rho) \right]$$

→ The “ ξ ” class

3 Perhaps some other(s)...

The “t Hooft–Polyakov” class

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If $\frac{f_3}{f_4} = H \frac{d}{dH} \left(\frac{f_1}{f_2} \right)$:

$$K = \frac{\rho}{\sinh \rho}$$
$$H = \left(\frac{f_1}{f_2} \right)^{-1} (\kappa) \quad \text{where } \kappa \equiv \coth \rho - \frac{1}{\rho}$$

- This class is a direct generalization of the 't Hooft–Polyakov solution (included here as a special case $f_{1,3} = H$ and $f_{2,4} = 1$)
- However, this class contains monopoles with some novel features
- Let's see an example...

An example: Power-function model

- The simplest example (two-parametric class of theories with $n \geq m$):

$$\begin{aligned} f_1^2 &= H^{n+m}, & f_3^2 &= nmH^{n+m}, \\ f_2^2 &= H^{n-m}, & f_4^2 &= \frac{n}{m}H^{n-m}. \end{aligned}$$

- The Lagrangian:

$$\begin{aligned} \mathcal{L} &= \frac{1}{2} \left(\frac{\phi}{v} \right)^{n+m-2} \left[(D^\mu \phi)^2 + (nm - 1) \frac{(\phi \cdot D^\mu \phi)^2}{\phi^2} \right] \\ &\quad - \frac{1}{4g^2} \left(\frac{\phi}{v} \right)^{n-m} \left[(\mathbf{F}^{\mu\nu})^2 + \left(\frac{n}{m} - 1 \right) \frac{(\phi \cdot \mathbf{F}^{\mu\nu})^2}{\phi^2} \right] \end{aligned}$$

- Solution:

$$K = \frac{\rho}{\sinh \rho}, \quad H = \sqrt[m]{\kappa} \quad \left(\text{where } \kappa \equiv \coth \rho - \frac{1}{\rho} \right)$$

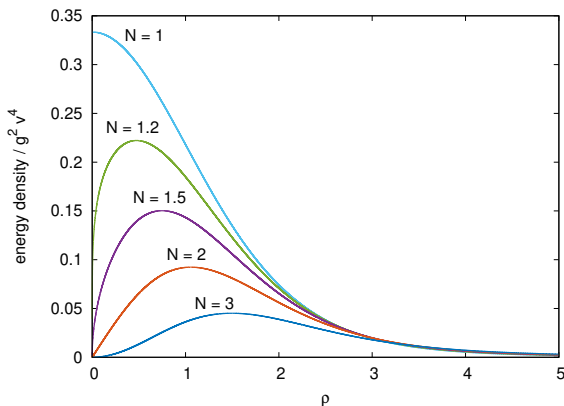
- Energy density:

$$\mathcal{E} = \kappa^N \left[2\kappa \frac{K^2}{\rho^2} + \frac{N}{\kappa} \frac{(1 - K^2)^2}{\rho^4} \right] v^4 g^2 \quad \text{where} \quad N \equiv \frac{n}{m}$$

Power-function model: Energy density

Where is the monopole's energy concentrated?

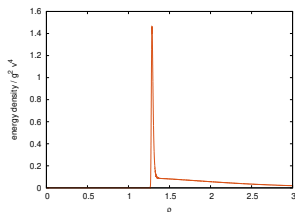
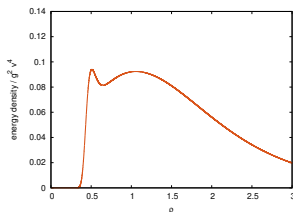
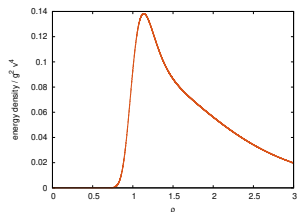
- For $N = 1$:
 - Most of the energy is in the center of the monopole
 - This is the 't Hooft–Polyakov monopole
- For $N > 1$:
 - Energy density vanishes in the center of the monopole!
 - The energy is stored in a *spherical shell* around the center of the monopole
 - \Rightarrow “Hollow” monopoles
 - Seems to be in fact quite typical behavior (also in other models...)



Other models – other features

In other models there can be other novel features:

- Wider and more pronounced “cavity” in the center of “hollow monopole”
(The energy density falls off *exponentially* in the origin)
- Two or more local maxima of the energy density
(Spherical shell of energy around the monopole center is structured, with several “sub-shells”)
- Sharper and more pronounced energy shell



(All these density energy plots correspond to exact solutions)

The “ ξ ” class

The “ ξ ” class

If $f_1 = f_2$:

$$K = \xi \exp(-\rho)$$

$$H = \left(\int \frac{dH}{H} \frac{f_3}{f_4} \right)^{-1} (1 - \lambda) \quad \text{where } \lambda \equiv \frac{1}{\rho} - \xi^2 \left[\frac{e^{-2\rho}}{\rho} + 2\text{Ei}(-2\rho) \right]$$

(where $\text{Ei}(x) = -\int_{-x}^{\infty} \frac{e^{-t}}{t} dt$ is the *exponential integral*)

Notice the presence of a parameter ξ :

- Not present in the Lagrangian: Emerges as a constant of integration
- Regularity of energy density: $\xi \in (-1, 1)$ (otherwise unconstrained)

The parameter ξ is physical (i.e., *measurable*):

- It appears in a physical quantity: The energy density
- It controls the profile of the energy density
- However, it doesn't change the total energy: Monopoles with different ξ have the same mass

$\Rightarrow \xi$ is an internal degree of freedom (or moduli space parameter) of the monopole!

An example: Power-function model

A particular example (a representative of a larger family):

- Lagrangian:

$$\mathcal{L} = \frac{1}{2} \left(\frac{\phi}{v} \right)^6 \left\{ (D^\mu \phi)^2 + \left[16 \left(\frac{\phi}{v} \right)^{-2} - 1 \right] \frac{(\phi \cdot D^\mu \phi)^2}{\phi^2} \right\} \\ - \frac{1}{4g^2} \left(\frac{\phi}{v} \right)^8 \left\{ (\mathbf{F}^{\mu\nu})^2 + \left[4 \left(\frac{\phi}{v} \right)^2 - 1 \right] \frac{(\phi \cdot \mathbf{F}^{\mu\nu})^2}{\phi^2} \right\}$$

- Solution:

$$K = \xi e^{-\rho},$$

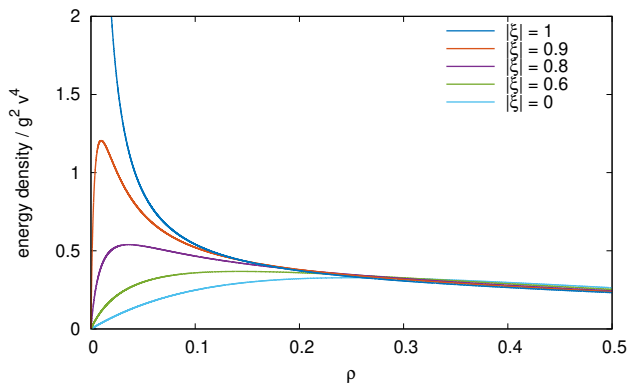
$$H = \left[1 + \frac{1 - \xi^2 e^{-2\rho}}{\rho} - 2\xi^2 \text{Ei}(-2\rho) \right]^{-1/2}$$

- Energy density:

$$\mathcal{E} = H^8 \left[2 \frac{K^2}{\rho^2} + 4H^2 \frac{(1 - K^2)^2}{\rho^4} \right] v^4 g^2$$

→ depends on ξ through K and H

The energy density



Indeed:

- The parameter ξ controls the “shape” of the monopole
- For $|\xi| \geq 1$ the energy density is singular

Summary & Outlook

Summary:

- Considered a class of effective $SU(2)$ models that admit monopole solutions
- Constructed the BPS limit
- Found exact monopole solutions with interesting properties
 - Interesting/weird/bizarre energy density profiles
 - New internal degree of freedom ξ : Controls the “shape” of the monopole, but doesn't change its total energy

Outlook (short term): Are there any deeper/physical reasons for...

- ... the existence of the hollow cavity in the monopoles?
- ... the occurrence of ξ ? Perhaps some symmetry?

Outlook (long term):

- SUSY? Dyons? Higher magnetic charges? Multi-monopole solutions? Dynamics? ...?

Reference: *Phys.Rev.D* 107 (2023) 12, 12, arXiv:2303.15602