



# QCD corrections to Coulomb potential at vanishing transverse momenta

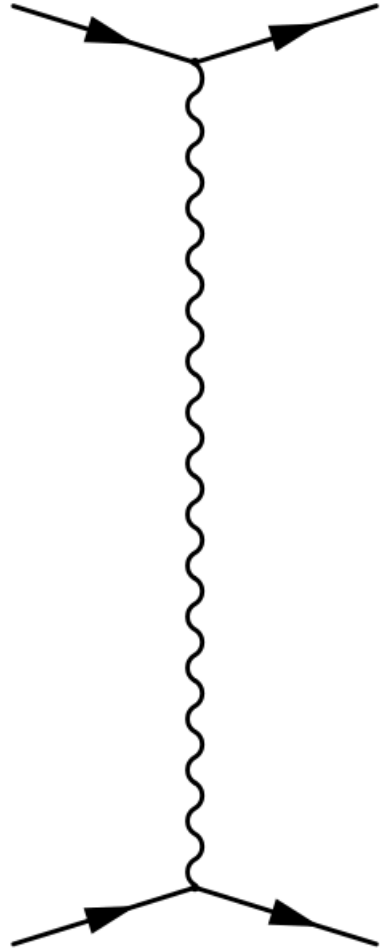
Martin Spousta

Charles University

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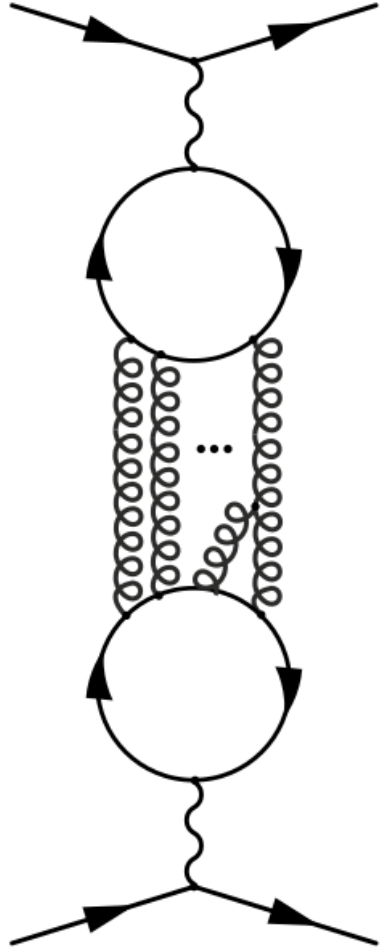
# Introduction



- Tree-level amplitude leading to Coulomb potential in  $q^2 \rightarrow 0$  limit.



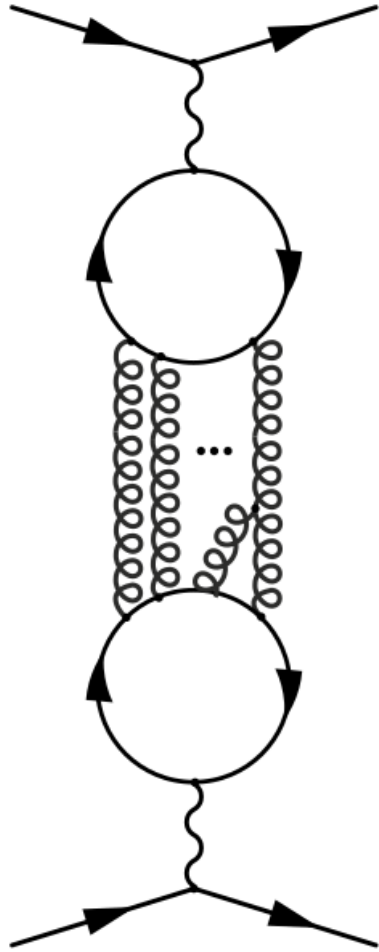
# Introduction



- Tree-level amplitude leading to Coulomb potential in  $q^2 \rightarrow 0$  limit.
- How will the QCD corrections to the photon propagator affect the behavior in  $q^2 \rightarrow 0$  limit?



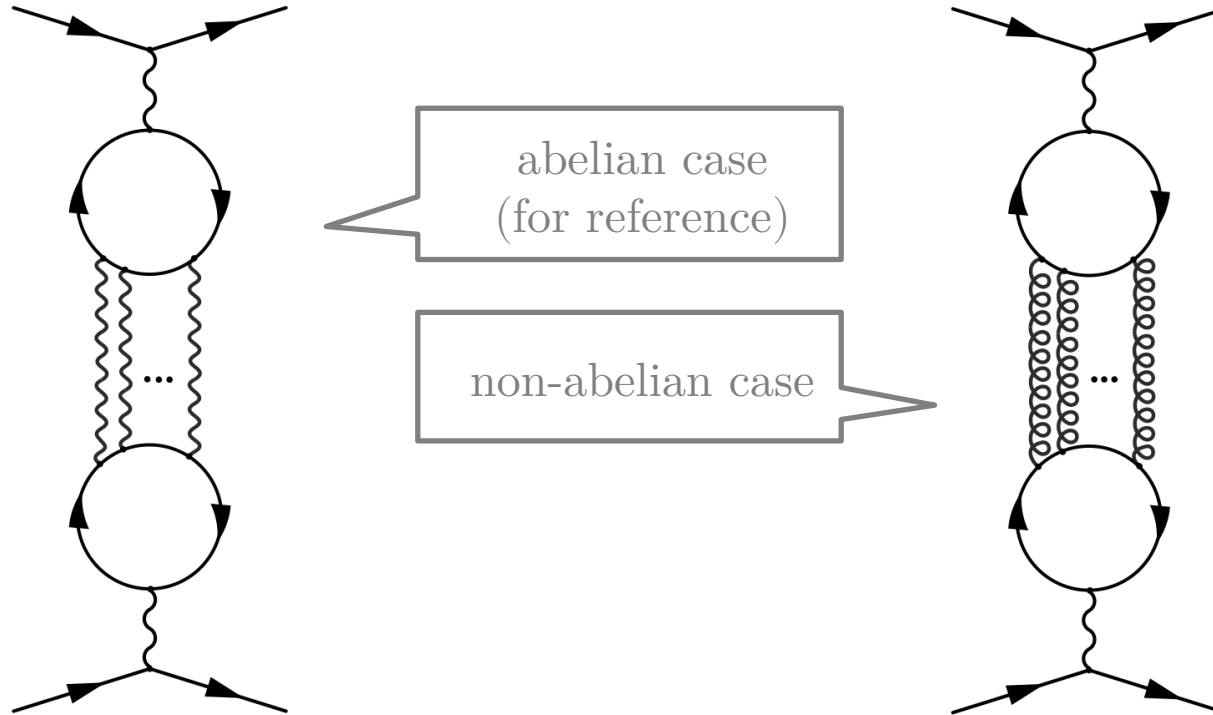
# Introduction



- Tree-level amplitude leading to Coulomb potential in  $q^2 \rightarrow 0$  limit.
- How will the QCD corrections to the photon propagator affect the behavior in  $q^2 \rightarrow 0$  limit?
- For  $q^2 \gg 0$ , these are local, diverging leading e.g. to running coupling.
- For  $q^2 \rightarrow 0$ , inter-loop gluon exchanges should be an important source of confining forces [[arXiv:1604.08082](https://arxiv.org/abs/1604.08082)].
- In many calculations [[arXiv:0709.2877](https://arxiv.org/abs/0709.2877), [arXiv:1408.5409](https://arxiv.org/abs/1408.5409), ...],  $\alpha_s(q^2 \rightarrow 0) < 1 \Rightarrow$  IR behavior can be investigated perturbatively.
- First, investigate amplitudes with 2 loops + multi-gluon exchanges in  $q^2 \rightarrow 0$  limit in a systematic way.



# General structure of amplitude



$$i\mathcal{M}^{\alpha\beta} = (i\sqrt{a})^2 \frac{-ig^{\alpha\mu}}{q^2 + i\epsilon} Y_{\mu\nu}(q) \frac{-ig^{\nu\beta}}{q^2 + i\epsilon},$$

$$Y_{\mu\nu}(q) = \int \prod_i du_i \underbrace{A_{\mu, \mu_1 \dots \mu_{n-1}}(u_i, q)}_{\text{n-point quark loop}} \underbrace{B^{\mu_1 \dots \mu_{n-1}, \nu_1 \dots \nu_{n'-1}}(u_i, q)}_{\text{bosonic exchanges}} \underbrace{A'_{\nu, \nu_1 \dots \nu_{n'-1}}(u_i, q)}_{\text{n'-point quark loop}}$$

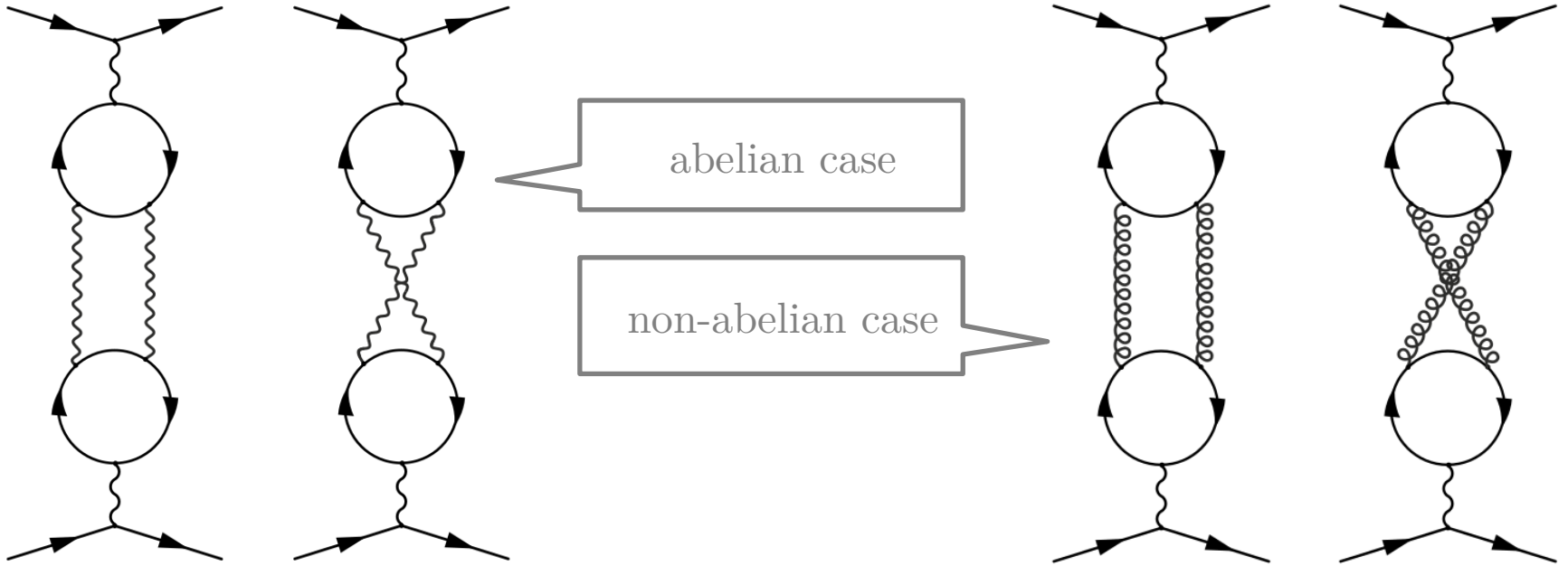
n-point  
quark loop

bosonic exchanges

n'-point  
quark loop



# $n=n'=3$ contribution



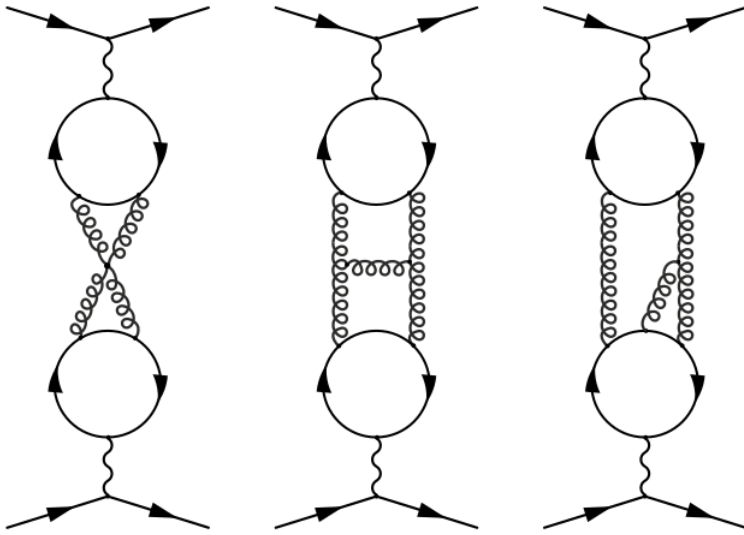
$$Y_{\mu\nu} = \alpha_{(s)}^2 \int d^4 u_1 \int d^4 p \operatorname{Tr} \left\{ \gamma_\mu \frac{i}{\not{p} + \not{q} - m} \gamma_{\mu_1} \frac{i}{\not{p} + \not{u}_1 + \not{q} - m} \gamma_{\mu_2} \frac{i}{\not{p} - m} \right\} \frac{-ig_{\mu_1 \nu_1}}{(u_1 + q)^2 - M^2} \frac{-ig_{\mu_2 \nu_2}}{u_1^2 - M^2} \int d^4 p' \operatorname{Tr} \left\{ \gamma_\nu \frac{i}{\not{p}' - m} \gamma_{\nu_2} \frac{i}{\not{p}' + \not{u}_1 + \not{q} - m} \gamma_{\nu_1} \frac{i}{\not{p}' + \not{q} - m} \right\},$$

$= 0$   $\begin{cases} \rightarrow$  Furry's theorem  
 $\rightarrow$  If  $u_i^2 = M^2 = 0$  (abelian case)

$\neq 0 \rightarrow$  If  $u_i^2 = M^2 \neq 0$  (non-abelian case, IR gluon mass [[arXiv:2201.09747](https://arxiv.org/abs/2201.09747)])



# $n=3, n'=3,4$ contribution



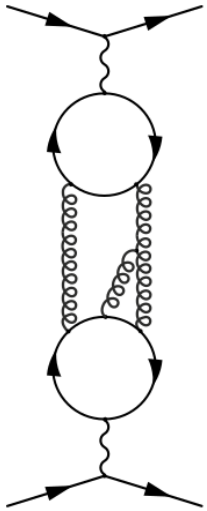
non-abelian only

- Calculations performed using Package-X + tailored Python interface
- Same limiting results as for  $n=n'=3$   
=> discuss the structure of solution for one particular amplitude



# n=3, n'=4 contribution

$$Y_{\mu\nu}(q) = \int \prod_i du_i A_{\mu, \mu_1 \dots \mu_{n-1}}(u_i, q) B^{\mu_1 \dots \mu_{n-1}, \nu_1 \dots \nu_{n'-1}}(u_i, q) A'_{\nu, \nu_1 \dots \nu_{n'-1}}(u_i, q)$$



- After performing the contraction due to bosonic fields – several thousands terms.
- Some terms contain scalar product  $u_i \cdot u_j$  inside PV coefficients.
- Currently possible to calculate only part of the amplitude with no scalar products:

$$Y_{\mu\nu}^{\text{part}} = g_{\mu\nu} \alpha_s^{\frac{5}{2}} \left[ \sum_{j=0}^4 c_{0j} \log^j \frac{\mu^2}{M^2} + \frac{1}{\epsilon} \sum_{j=0}^3 c_{1j} \log^j \frac{\mu^2}{M^2} + \frac{1}{\epsilon^2} \sum_{j=0}^2 c_{2j} \log^j \frac{\mu^2}{M^2} + \mathcal{O}(\epsilon) \right]$$

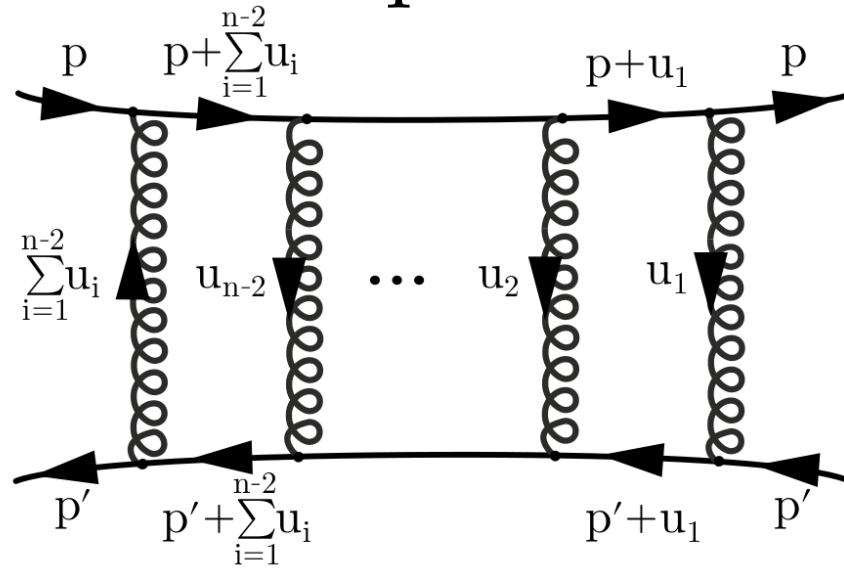
$$c_{ij} = c_{ij}(m^2, M^2, \log(m^2/M^2), \text{PV}(m, M))$$

- Remarkably  $Y_{\mu\nu}^{\text{part}} \neq 0$  for  $M^2 \neq 0$ . This unexpected **IR finite behavior** can be generalized ...





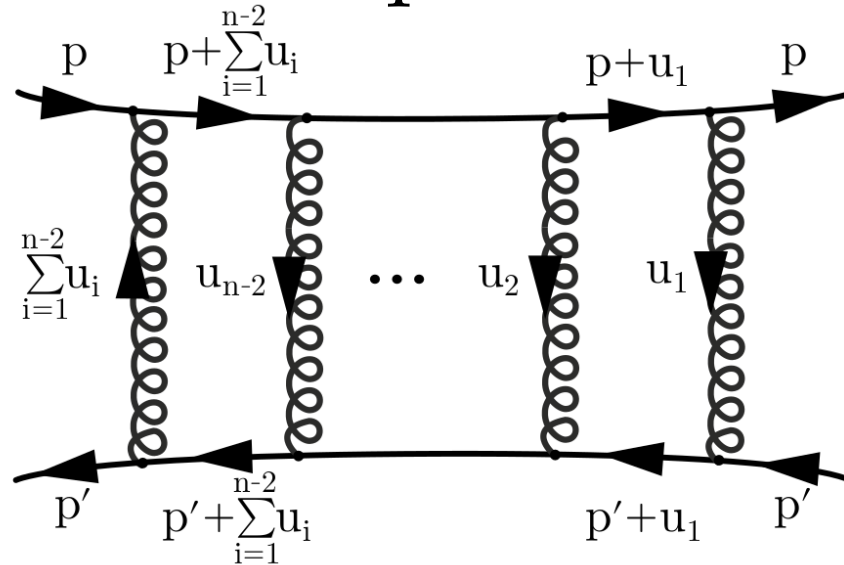
# Generalized partial solution



... for any n-even



# Generalized partial solution



... for any n-even

$$Y_{\mu\nu}^{\text{part}} \sim g_{\mu\nu} \alpha_s^{n-1} [M^2(1 + \log M^2 + \log^2 M^2) + \mathcal{O}(M^4)] \equiv g_{\mu\nu} f_M$$

... that is  $Y_{\mu\nu}^{\text{part}} \neq 0$  for  $M^2 \neq 0$ . Full solution for any n-even is then:

$$Y_{\mu\nu} = \left( g_{\mu\nu} - \frac{q_\mu q_\nu}{q^2} \right) \left( \underbrace{f_M(\alpha_s, M)}_{\text{finite}} + \underbrace{f_\Delta(\alpha_s, m, M, q)}_{\text{analytical function of } q} + \underbrace{f_N(\alpha_s, m, M, q)}_{\text{non-analytical function of } q} \right)$$



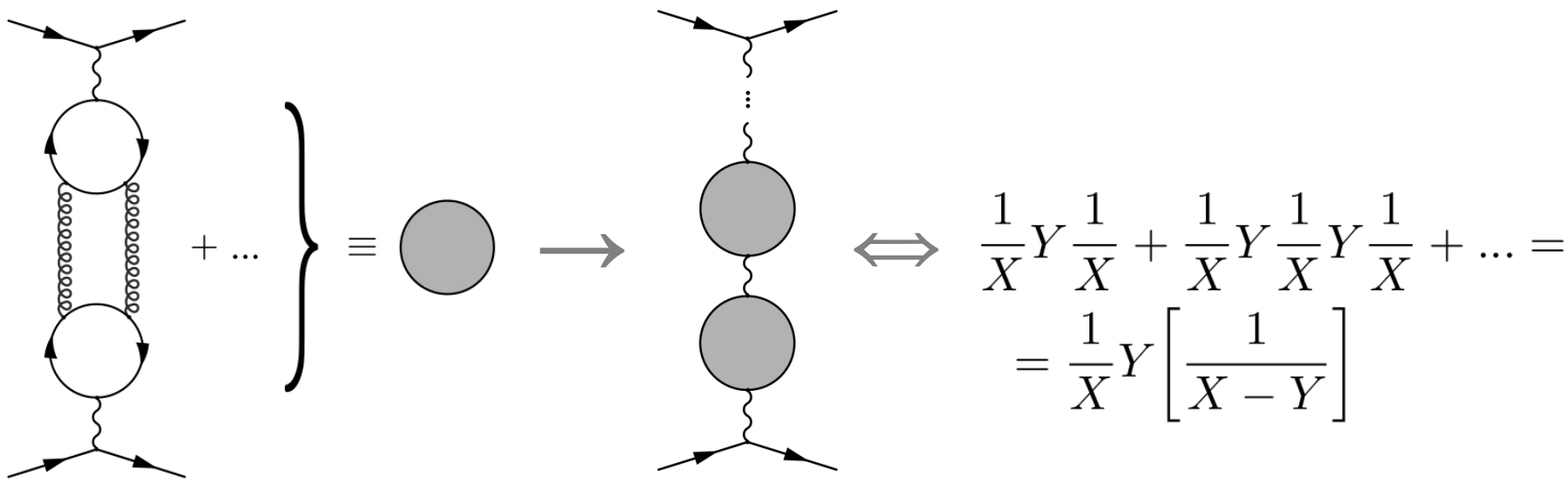
# Full solution: resummation

- Need to **resum** these contributions to get full solution.
- We empirically know that these loop interactions give rise to confining forces in the  $q^2 \rightarrow 0$  limit which are different than Coulomb interaction  $\Rightarrow$  resum them **separately** from free photon propagator in  $q^2 \rightarrow 0$  limit.
- Presence of **finite solution** allows to perform this resummation.



# Full solution: resummation

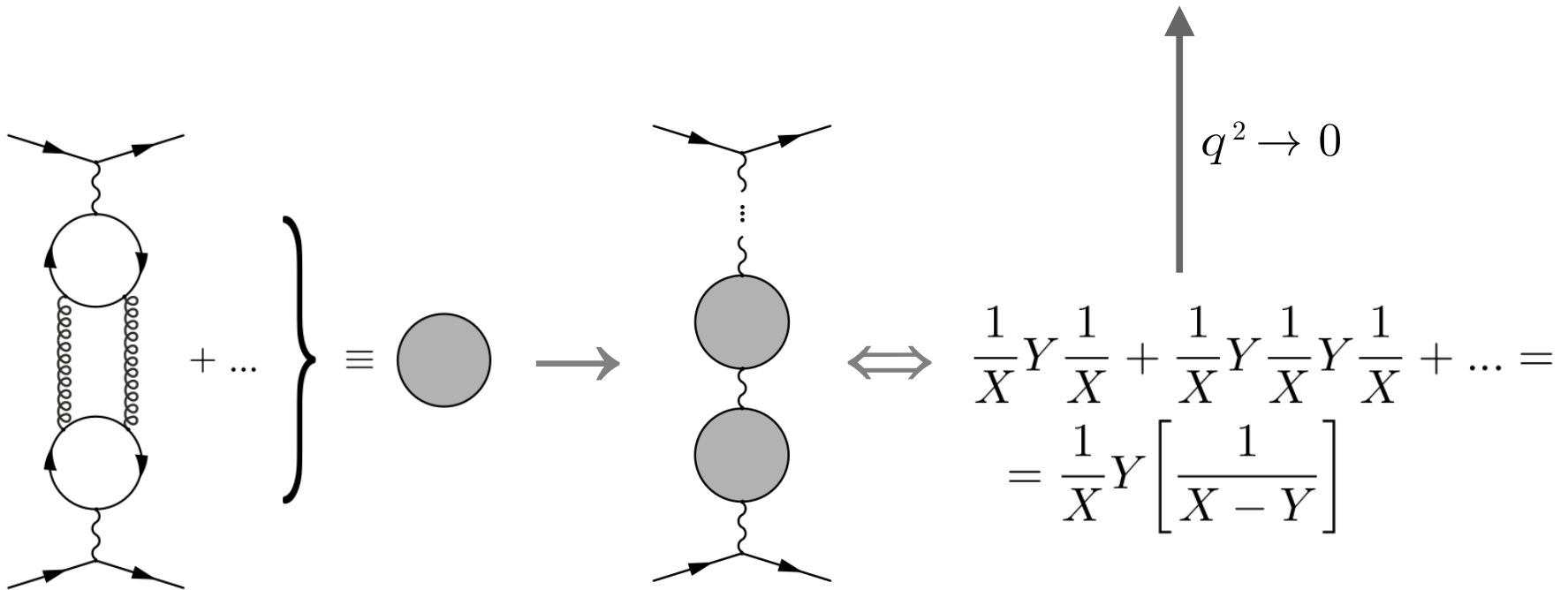
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# Full solution: resummation

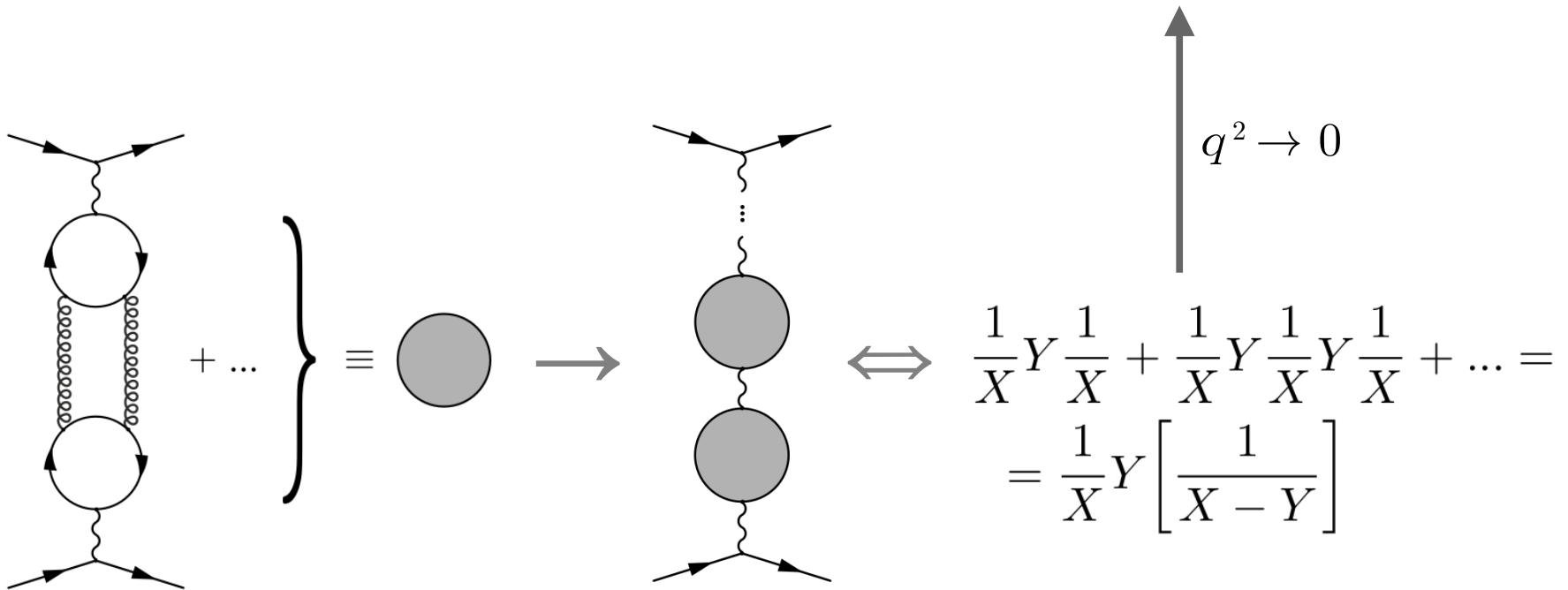
$$i\mathcal{M}^{\alpha\beta} = -g^{\alpha\beta} \left( \frac{ia}{q^2} - \frac{1}{f_M(\alpha_s, M) + f_{\Lambda,0}(\alpha_s, m, M) + f_N(\alpha_s, m, M, q)} + \mathcal{O}(q) \right)$$





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Change in the **sign** wrt  
Coulomb interaction

Exact  $1/r$  behavior of resulting potential:

- due to the presence of  $f_M$ ,  
 $f_N$  can have any  $q^2$  behavior  
and any  $\mu^2$  dependence
- no  $1/r^k$  corrections to  $1/r$

Can be generalized (backup), but sufficient to see for one class of amplitudes.



# Sign of the interaction

	Yukawa		electromagnetic			
	$ff$	$f\bar{f}$	$ff$	$f\bar{f}$	$f_1^- f_2^+$	$f_1^- \bar{f}_2^-$
fermion current	+1	-1	+1	+1	+1	+1
intermediate particle	+1	+1	-1	-1	-1	-1
normal product	-1	+1	-1	+1	-1	-1
charge	+1	+1	+1	+1	-1	+1
$(-1) \cdot \text{total}$	+1	+1	-1	+1	+1	-1

- Sign of Coulomb interaction is **complex** (!)





# Sign of the interaction

	Yukawa		“residual”			
	$ff$	$f\bar{f}$	$ff$	$f\bar{f}$	$f_1^- f_2^+$	$f_1^- \bar{f}_2^-$
fermion current	+1	-1	+1	+1	+1	+1
intermediate particle	+1	+1	-1	-1	-1	-1
normal product	-1	+1	-1	+1	-1	-1
charge	+1	+1	+1	+1	+1	+1
summation			-1	-1	-1	-1
$(-1) \cdot \text{total}$	+1	+1	+1	-1	+1	+1

- Sign of Coulomb interaction is **complex** (!)
- Probing **strongly interacting degrees of freedom** in fermion’s wave function  
 $\Rightarrow$  electromagnetic charge is irrelevant.



# Sign of the interaction

	Yukawa		“residual”			
	$ff$	$f\bar{f}$	$ff$	$f\bar{f}$	$f_1^- f_2^+$	$f_1^- \bar{f}_2^-$
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intermediate particle	+1	+1	-1	-1	-1	-1
normal product	-1	+1	-1	+1	-1	-1
charge	+1	+1	+1	+1	+1	+1
summation			-1	-1	-1	-1
$(-1) \cdot \text{total}$	+1	+1	+1	-1	+1	+1

- Sign of Coulomb interaction is **complex** (!)
- Probing **strongly interacting degrees of freedom** in fermion’s wave function  $\Rightarrow$  electromagnetic charge is irrelevant.
- Total sign of the “QCD residual interaction”: **attractive except for fermion-anti-fermion** which is repulsive.



# Strength of the interaction

- Strength given by the hadronic part of elmg. coupling:

$$\alpha(q^2) = \frac{\alpha_0}{1 - \Delta\alpha(q^2)} = \alpha_0 [1 + \Delta\alpha(q^2)_{\text{had}} + \Delta\alpha(q^2)_{\text{lep}} + \Delta\alpha(q^2)_{\text{top}} + \mathcal{O}(\Delta\alpha^2)]$$

$$\alpha_0 \Delta\alpha(q^2)_{\text{had}} \equiv a(q^2)$$

- Magnitude of  $a$  can be estimated using **alphaQED** package [[arXiv:1905.05078](https://arxiv.org/abs/1905.05078)].
- The ratio of  $a$  to fine structure constant  $\alpha_0$  is:

$$a / \alpha_0 = 10^{-20} \quad \dots \text{ for eV scale}$$

$$a / \alpha_0 = 10^{-26} \quad \dots \text{ for meV scale}$$

(i.e. not far from the ratio of gravity to elmg. interaction let's say for  $e^-$ )



# Where it can act

- Terrestrial experiments have no sensitivity to measure it (20 orders of magnitude smaller than Coulomb force).
- It can play a role in **large-scale astrophysical objects** due a presence of QED plasma in thermal equilibrium, e.g.
  - Intracluster medium (**ICM**) in clusters of galaxies
  - Warm ionized medium (**WIM**) in interstellar space

- Magnitude of residual force in a Debye volume of plasma:

$$\Phi_{\text{deb}}(r) = \frac{1}{4\pi\epsilon_0} \frac{e}{r} \exp\left(-\frac{\sqrt{2}r}{\lambda_D}\right) \rightarrow F_{\text{res}}(\lambda_D) = 2n_i e \int^{\lambda_D} \frac{\partial\Phi_{\text{deb}}(r)}{\partial r} d^3x \Big|_{\alpha_0 \rightarrow a(q^2)}$$

- **Comparison to gravity** within Debye volume:

$$F_{\text{res}} \sim 10^{-40} \text{ GeV/fm} \quad F_G \sim 10^{-61} \text{ GeV/fm} \quad \dots \text{ for ICM}$$

$$F_{\text{res}} \sim 10^{-46} \text{ GeV/fm} \quad F_G \sim 10^{-62} \text{ GeV/fm} \quad \dots \text{ for WIM}$$



# Summary + speculation

- Basic properties of residual interaction:
  - Potential with exact  $1/r$  behavior.
  - **Attractive** for all fermions except for fermion-anti-fermion.
  - **Universal** – each particle has a hadronic part in its wave function (including e.g. neutrino).
  - **Not localized** to a single point but has an intrinsic spacial structure for  $q^2 \rightarrow 0$  (c.f. polarization tensor).
  - **Small interaction strength** (not completely far from that of gravity).
  - Linear **dependence on mass** (cf. normalization of bispinors – backup).
- Speculation: Can this be a new candidate for an **emergent gravity**?
  - Only if virtual photons are universally present unaffected by multipole structure of charged particles.
  - If not, still important for large scale astrophysical objects: ICM, WIM, etc. and possibly a part of DM puzzle.



# More information

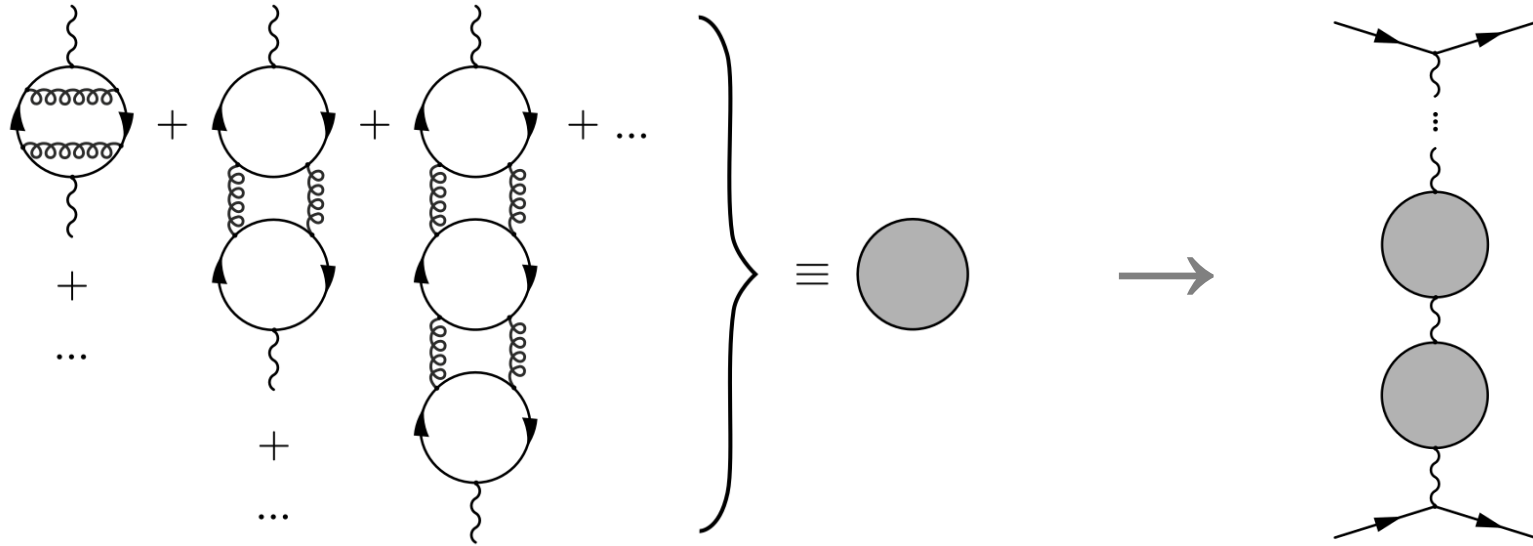
- For details see [EPJP 139 \(2024\) 5, 374](#) + publication in preparation
- Little more in the backup

# Backup slides





# Generalization



$$\frac{1}{X} Y \frac{1}{X} + \frac{1}{X} Y \frac{1}{X} Y \frac{1}{X} + \dots = \frac{1}{X} Y \left[ \frac{1}{X - Y} \right]$$
$$\frac{y}{q^2(q^2 - y)} = -\frac{1}{q^2} + \frac{1}{q^2 - y} = -\frac{1}{q^2} + \frac{1}{q^2 - y} = -\frac{1}{q^2} - \frac{1}{y} (1 + \mathcal{O}(q^2))$$

- Why **not to** resum them to the photon propagator? Because we empirically know that they (or their subclass) give rise to confining forces in the  $q^2 \rightarrow 0$  limit (which are different than Coulomb force).





# $\alpha_s(0)$ I.

$\alpha_s(0)$	IR behavior	Definition	RS	Gauge	Quarks	Framework	References
3.2	freezes	gh.-gl. vertex	MOM	Landau	Yes	Lattice	[331]
4.45	freezes	gh.-gl. vertex	MOM	Landau	No	Lattice	[329]
● 0.5	freezes	pQCD form	MOM	Indep.	No	Lattice	[329]
● $0.29 \pm 0.03$	freezes	Q-Q pot.	$V$	Indep.	No	Lattice	[212]
● 0.40	freezes	Q-Q pot.	$V$	Indep.	Yes	Lattice	[212]
$3.3 \pm 0.7$	freezes	gh.-gl. vertex	$(\tau)$	Landau	No	Lattice $\Lambda = 1.1$	[237]
● 0	vanishes	gh.-gl. vertex	MOM	Landau	No	Lattice	see caption
● 0	vanishes	gh.-gl. vertex	MOM	Landau	Yes	Lattice	[264,276]
$2.5 \pm 0.5$	freezes	gh.-gl., q.-gl. v.	MOM	Land. Coul.	Yes	Lattice	[258]
● 0.52	freezes	Q-Q pot.	$V$	Indep.	Yes	Lat. pert. stoch. th.	[338]
diverges	$e^{m/Q}$	Schrod. func.	$\overline{MS}$	–	Yes	Lattice	[341]
2.97	freezes	gh.-gl. vertex	–	indep.	Yes	Stoch. Quant.	[244,348]
$\infty$	$1/Q^2$	pQCD form	$V$	Coulomb	Yes	Variational ap.	[163]
● 0.71	freezes	pQCD form	$\overline{MS}$	Indep.	Yes	Pheno.	[222]
$\infty$	$1/Q^2$	pQCD form	$V$	Indep.	Yes	Pheno.	[294]
● 0.58	freezes	pQCD form	$V$	Indep.	Yes	Pheno.	[221],
$\infty$	$1/Q^2$	pQCD form	$\overline{MS}$	Indep.	Yes	Pheno.	[226]
● $0.42 \pm 0.03$	(Av. IR val.)	pQCD form	–	Indep.	Yes	Pheno.	[206]
● $\simeq 0.6$	freezes	pQCD form	–	Indep.	Yes	Pheno.	[206]

A. Deur et al., Prog. Part. Nucl. Phys. 90 (2016) 1–74 [ [arXiv:1604.08082](https://arxiv.org/abs/1604.08082) ]

**23 out of 39 papers** in the review suggest  $0 \leq \alpha_s(0) < 1$ .



# $\alpha_s(0)$ II.

$\alpha_s(0)$	IR behavior	Definition	RS	Gauge	Quarks	Framework	References
●0.78	mono. incr.	pQCD form	$\overline{MS}$	Indep.	Yes	Pheno.	[374]
●0.60	freezes	pQCD form	$(\overline{MS})$	Indep.	Yes	Pheno. $\Lambda = 0.2$	[219]
●0.82	freezes	pQCD form	$(\overline{MS})$	Indep.	Yes	Pheno.	[292]
diverges	$1/Q^2$	pQCD form	$V$	Indep.	Yes	Pheno.	[224]
●0.83	freezes	pQCD form	–	Indep.	Yes	Pheno.	[220]
●0.83	freezes	$\alpha_R$	Indep.	Indep.	Yes	Opt. Pert. Theo.	[388]
1.77	freezes	gh.-gl. vertex	$\overline{MS}$ , MOM	Landau	Yes	Gribov–Zwanziger	[351,214]
$7.7 \pm 2$	freezes	–	–	Landau	No	FRG	[343]
2.97	freezes	gh.-gl. vertex	–	Landau	No	FRG	[257]
$\sim 4$	freezes	–	Indep.	Indep.	Yes	Analytic Ap. $\Lambda = 0.56$	[166]
1.25	mono. decr.	–	$(\overline{MS})$	Indep.	Yes	Analytic Ap. $\Lambda = 0.32$	[362,359]
diverges	$1/Q^2$	–	Indep.	Indep.	Yes	Analytic Approach	[368,377]
●0.47	freezes	–	Indep.	Indep.	Yes	Analytic Approach	[378]
●0.37–0.60	freezes	pQCD form	$\overline{MS}$	Indep.	No	BPT	[218,386]
●0.81	freezes	pQCD form	$V$	Indep.	Yes	BPT	[45]
●0.97 $\pm$ 0.44	freezes	pQCD form	$V$	Indep.	Yes	BPT	[45,387]
●0.50 $\pm$ 0.08	freezes	pQCD form	$\overline{MS}$	indep.	Yes	duality.	[389]
●0	vanishes	gh.-gl. vertex	–	–	No	$\phi^4$ -YM mapping	[390]
●0	vanishes	pQCD form	$\overline{MS}$	indep.	No	Bogoliubov comp.	[170]
●0	vanishes	pQCD form	$\overline{MS}$	Landau	Yes	Curci–Ferrari Model	[392]

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# Normalization

- **Coulomb case.** Full amplitude, including also initial and final state description

$$i\mathcal{M} = \frac{-ie^2}{4\pi\epsilon_0\hbar c} \frac{1}{|\mathbf{q}|^2} (2m\xi'^{\dagger}\xi)_p (2m\xi'^{\dagger}\xi)_k$$

- Non-relativistic Born approximation:

$$\langle p' | iT | p \rangle = -i\tilde{\mathcal{V}}(\mathbf{q}) (2\pi) \delta(E_{p'} - E_p)$$

- **Mass factors** (and  $\hbar c$ ) in relativistic amplitude **dropped manually** to match non-relativistic conventions.



# Open questions

- Can implementing such attractive force into **astrophysical simulations** help with interpretation of the data? Could it be a part of the dark matter puzzle?
- Can we prove that this force is not an **emergent gravity**?
- Can we improve on the **magnitude of hadronic part** of  $\alpha_0$  at (m)eV scales to have better estimates strength of the force (e.g. using lattice QCD)?



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