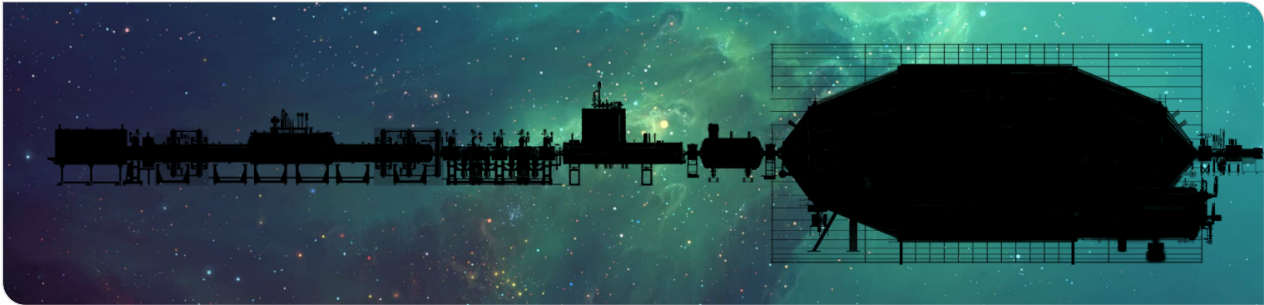


# Probing Physics Beyond the Neutrino Mass at the KATRIN Experiment

42<sup>nd</sup> International Conference on High Energy Physics

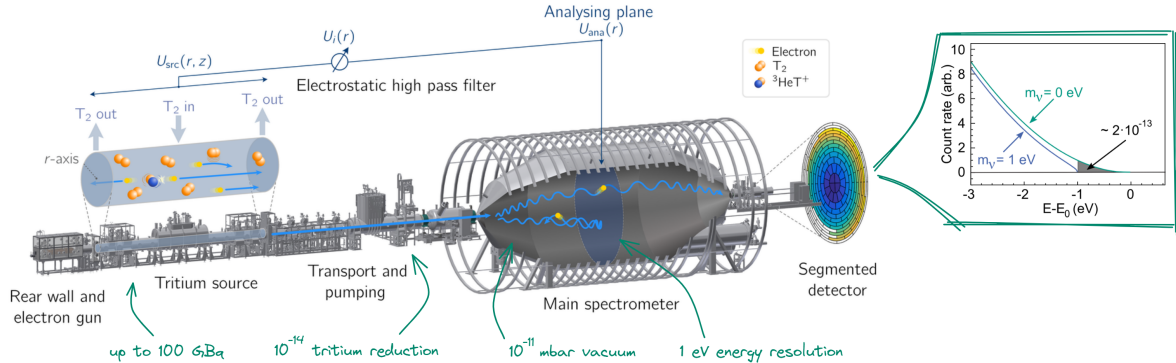
Caroline Fengler for the KATRIN Collaboration | July 20<sup>th</sup>, 2024



# The KATRIN Experiment

70 m long set-up: a gaseous tritium source & high resolution MAC-E filter

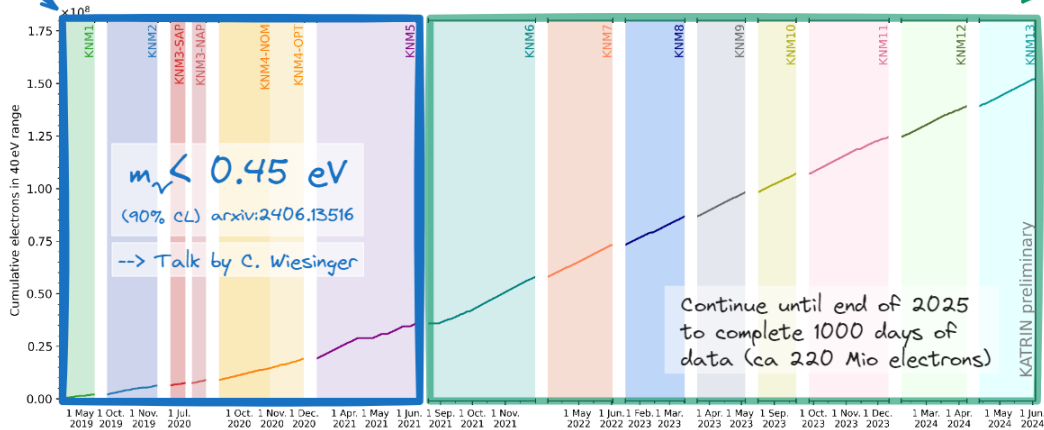
Continuous  $\beta$ -spectrum: measurement of effective mass  $m(\nu_e)$  based on kinematic parameters & energy conservation



# Neutrino Mass Results

New release with 36 Mio electrons in region of interest

Data taking and analysis in progress with now more than 150 Mio electrons in region of interest

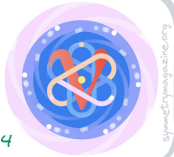


# Beyond the Neutrino Mass - New physics searches

## "Kink" search

for eV-scale sterile  $\nu$   
close to the endpoint

--> PRL 126 (2021) 091803  
& PRD 105 (2022) 072004

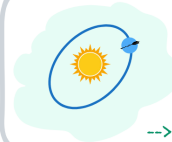


symmetrybreaking.org

## Line search

for capture of  
local cosmic relic  $\nu$

--> PRL 129 (2022) 011806



Search for Lorentz  
violation through  
sidereal modulation

--> PRD 107 (2023) 082005



## "Kink" search

for keV-scale sterile  $\nu$   
far from the endpoint

--> EPJ C 83 (2023) 763

European Commission

Search for shape  
distortions through  
exotic weak interactions

--> eg JHEP 01 (2019) 206



# Beyond the Neutrino Mass - New physics searches

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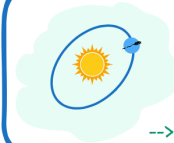
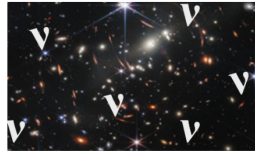


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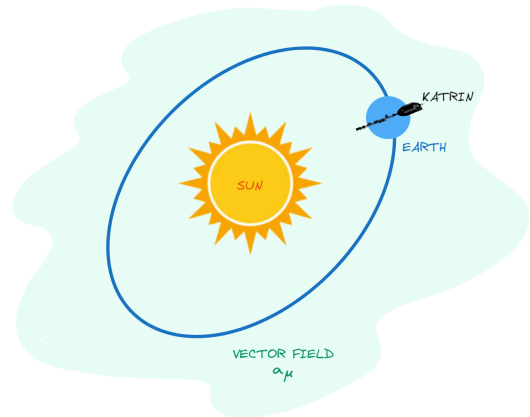


# Search for Lorentz Invariance Violations

- **Motivation:** BSM theories (String theory, loop quantum gravity and non-commutative QFT) suggest CPT and Lorentz invariance violation at high energies.
- **Constraints:** Neutrino oscillation, time-of-flight experiments, experiments using interaction processes (KATRIN)

$$\mathcal{L}_{SME}^a = -\bar{\psi}_w a^\mu \gamma_\mu \psi_w \quad w \in \{T, H, e, n\}$$

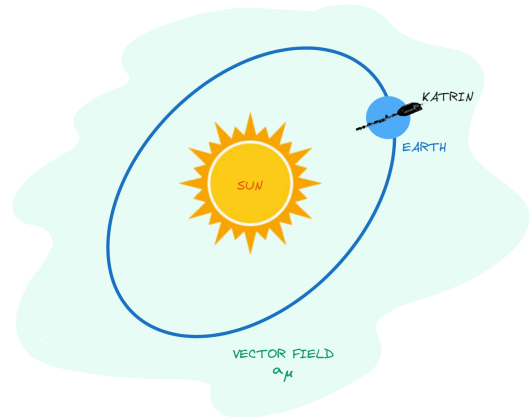
- Produces terms  $\propto a^\mu p_\mu = a_0 p_0 - a_i p_i$   
 → time-dependent & time-independent shift of  $E_0$



# Search for Lorentz Invariance Violations

## Time-dependent & time-independent shift of $E_0$ :

- **Rotation of Earth:** relative direction of KATRIN acceptance angle changes w.r.t Lorentz-violating vector  $a^\mu$
- **LIV-signature:** Measured endpoint energy  $E_0$  oscillates with sidereal time (23 h 56 min 4 s)  
→ Sensitive to  $|(a_{of}^{(3)})_{11}|$
- **LIV-signature:** Global shift of measured endpoint energy  $E_0$   
→ Sensitive to  $|(a_{of}^{(3)})_{00}|$  and  $|(a_{of}^{(3)})_{10}|$



# Results from first campaign

## Results:

KATRIN Coll., PRD **107** (2023) 082005

- No significant oscillation of  $E_0$  observed

### First upper limit:

$$\left| \left( a_{of}^{(3)} \right)_{11} \right| < 3.7 \times 10^{-6} \text{ GeV (90 \% CL)}$$

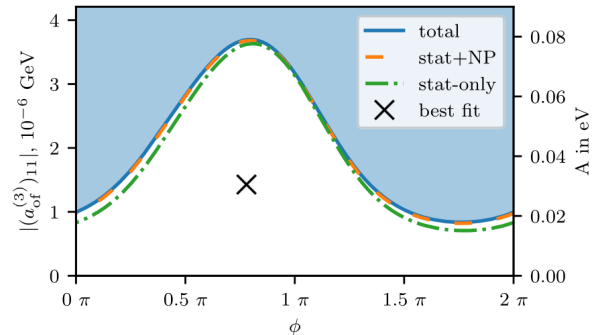
- No significant shift of  $E_0$  observed

### Improved upper limits:

$$\left| \left( a_{of}^{(3)} \right)_{00} \right| < 3.0 \times 10^{-8} \text{ GeV (90 \% CL)}$$

$$\left| \left( a_{of}^{(3)} \right)_{10} \right| < 6.4 \times 10^{-4} \text{ GeV (90 \% CL)}$$

$$A = \sqrt{\frac{3}{2\pi} \left| \left( a_{of}^{(3)} \right)_{11} \right|} \sqrt{B^2 \cos^2 \chi \cos^2 \xi + (\beta_{rot} - B \sin \xi)^2}$$



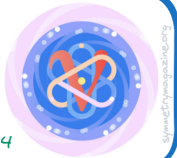


# Beyond the Neutrino Mass - New physics searches

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--> PRL 126 (2021) 091803  
& PRD 105 (2022) 072004

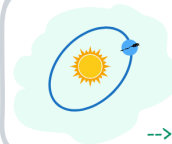


sponsored by the KIT

## Line search

for capture of  
local cosmic relic  $\nu$

--> PRL 129 (2022) 011806



Search for Lorentz  
violation through  
sidereal modulation

--> PRD 107 (2023) 082005



## "Kink" search

for keV-scale sterile  $\nu$   
far from the endpoint

--> EPJ C 83 (2023) 763

sponsored by the KIT

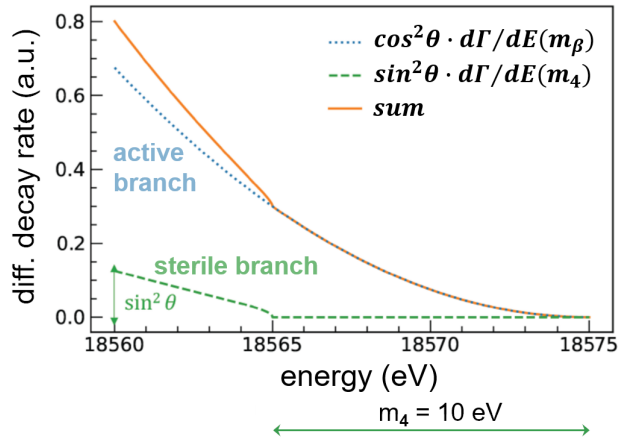
Search for shape  
distortions through  
exotic weak interactions

--> eg JHEP 01 (2019) 206



# Light Sterile Neutrinos

- **Motivation:** Multiple anomalies in the oscillation data, could be explained by  $\geq 1$  eV sterile neutrino
- **Analysis:** Add sterile  $\beta$ -spectrum with sterile mass  $m_4$  and active-to-sterile mixing  $\sin^2 \theta$  to active neutrino  $\beta$ -spectrum



# Light Sterile Neutrinos

$$\sin^2 2\theta = 4 |U_{e4}|^2 (1 - |U_{e4}|^2)$$

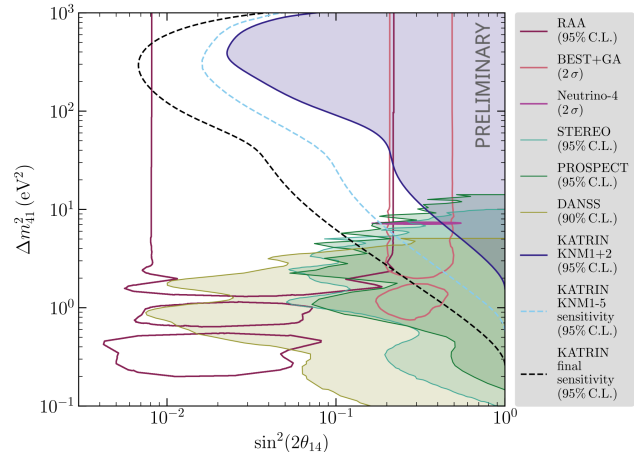
## Results of campaigns 1+2:

KATRIN Coll., Phys. Rev. D,  
10.1103/PhysRevD.105.072004

- No significant sterile neutrino signal observed in first two measurement campaigns
- Excluded large  $\Delta m_{41}^2$  solutions of reactor and gallium anomalies

## Projection for campaigns 1-5:

- Significant improvement in sensitivity
- Able to test last part of the Gallium Anomaly (GA) not excluded by short baseline oscillation experiments, and Neutrino-4 result.

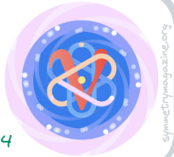


# Beyond the Neutrino Mass - New physics searches

## "Kink" search

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--> PRL 126 (2021) 091803  
& PRD 105 (2022) 072004

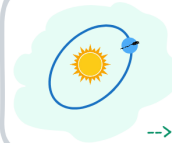


symmetrybreaking.org

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European Commission



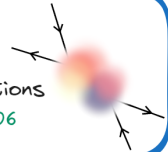
## "Kink" search

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far from the endpoint

--> EPJ C 83 (2023) 763

Search for shape  
distortions through  
exotic weak interactions

--> eg JHEP 01 (2019) 206



# Theory of General Neutrino Interactions (GNI)

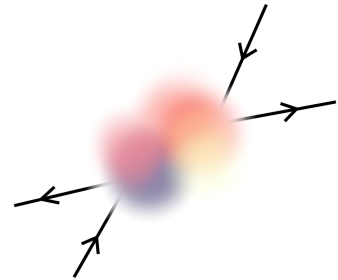
- Generalisation of neutrino Non-Standard Interactions (NSI)
- Considers scalar, pseudoscalar, vector, axial vector or tensor interactions of neutrinos with fermions:

$$\mathcal{L}_{GNI}^{NC} = -\frac{G_F}{\sqrt{2}} \sum_{j=1}^{10} \tilde{\epsilon}_{j,f}^{(\sim)} (\bar{\nu} O_j \nu) (\bar{f} O'_j f)$$

$$\mathcal{L}_{GNI}^{CC} = -\frac{G_F V_{Y\delta}}{\sqrt{2}} \sum_{j=1}^{10} \tilde{\epsilon}_{j,ud}^{(\sim)} (\bar{e} O_j \nu) (\bar{u} O'_j d) + h.c.$$

- Assume that GNI arise from heavy New Physics  $\rightarrow$  Map low energy GNI operators onto dim 6 SM(N)EFT terms.

$$\mathcal{L}_{EFT}(\phi) = \mathcal{L}_{SM}(\phi) + \sum_{n \geq 5} \frac{1}{\Lambda^{n-4}} C_i^{(n)} O_i^{(n)}(\phi)$$



$\rightarrow$  Enables broad search for New Physics through precision measurements.

Bischer and Rodejohann, Nucl. Phys. B,  
10.1016/j.nuclphysb.2019.114746

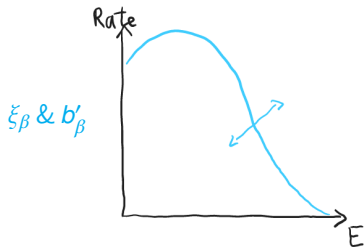
$$\frac{d\Gamma_{\text{GNI}}}{dE} = \frac{d\Gamma_{\text{SM}}}{dE} \sum_{k=\beta, N} \sqrt{(E_0 - E)^2 - m_k^2} \cdot \xi_k \left[ 1 - b'_k \frac{m_k}{E_0 - E} \right] \Theta(E_0 - m_k - E)$$

- Total differential decay rate for **light active neutrinos** and **additional heavier neutrinos**
- Dimensionless coefficients  $\xi_k$  and  $b'_k$  defined in terms of factors  $\epsilon$ ,  $\hat{\epsilon}$ ,  $U_{e4}$  and nuclear form factors  $g_V$ ,  $g_S$ ,  $g_T$  and  $g_A$ .
- Recover SM for  $\xi_N = b'_k = 0$ .

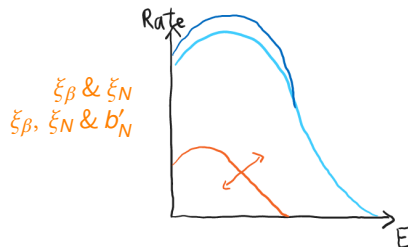
Bischer and Rodejohann, Nucl. Phys. B, 10.1016/j.nuclphysb.2019.114746

$$\frac{d\Gamma_{\text{GNI}}}{dE} = \frac{d\Gamma_{\text{SM}}}{dE} \sum_{k=\beta, N} \sqrt{(E_0 - E)^2 - m_k^2} \cdot \xi_k \left[ 1 - b'_k \frac{m_k}{E_0 - E} \right] \Theta(E_0 - m_k - E)$$

Only light active neutrinos



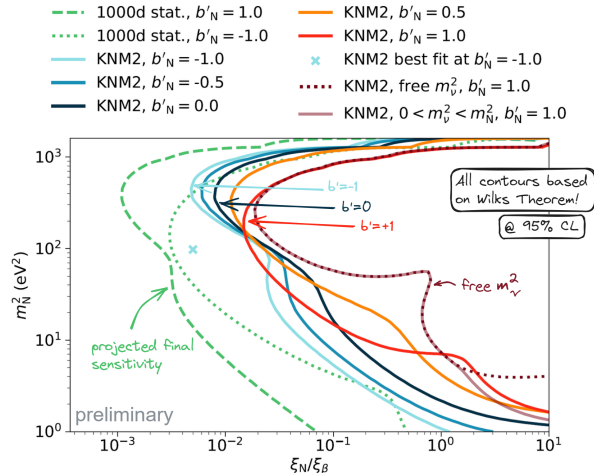
With additional heavier neutrinos



# Sensitivity of GNI for **additional heavier neutrinos**

- Able to search for GNI with KATRIN.
- No significant signal found in the KNM2 data set.
- Studied change in contour shape when allowing  $m_\nu^2 \neq 0$ .
- Sensitivity is dominated by statistics.  
→ Further significant improvements expected for final data set.
- Able to probe more specific scenarios, such as single types of interactions, right-handed W bosons, Leptoquarks and charged Higgs.

Full release of GNI analysis will follow soon.  
Stay tuned!





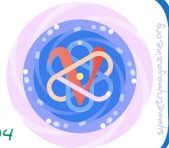
# Summary & Outlook

New publication in preparation, able to test critical parts of the parameter space

"Kink" search

for eV-scale sterile  $\nu$   
close to the endpoint

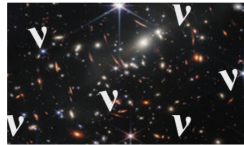
--> PRL 126 (2021) 091803  
& PRD 105 (2022) 072004



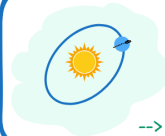
Line search

for capture of  
local cosmic relic  $\nu$

--> PRL 129 (2022) 011806



Analysis constrained new  
parameter and improved others



Search for Lorentz  
violation through  
sidereal modulation

--> PRD 107 (2023) 082005

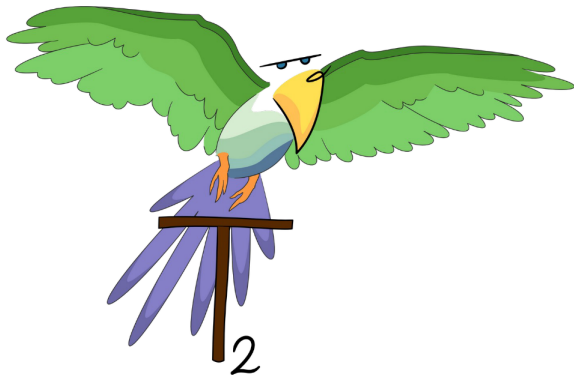
Search for shape  
distortions through  
exotic weak interactions

--> eg JHEP 01 (2019) 206



Glimpse at first results with KATRIN,  
full analysis in publication process

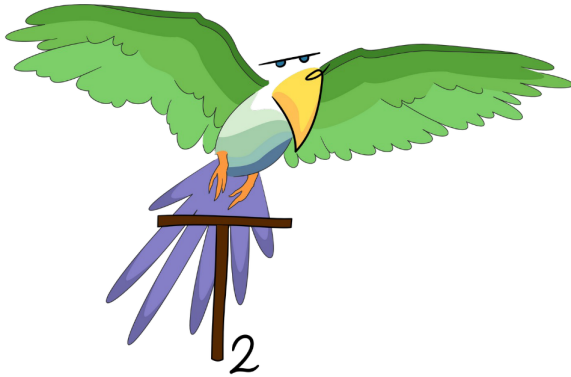
# Thank you for your attention!



We acknowledge the support of Helmholtz Association (HGF), Ministry for Education and Research BMBF (05A23PMA, 05A23PX2, 05A23VK2, and 05A23WO6), the doctoral school KSETA at KIT, Helmholtz Initiative and Networking Fund (grant agreement W2/W3-118), Max Planck Research Group (MaxPlanck@TUM), and Deutsche Forschungsgemeinschaft DFG (GRK 2149 and SFB-1258 and under Germany's Excellence Strategy EXC 2094 – 390783311) in Germany; Ministry of Education, Youth and Sport (CANAM-LM2015056, LTT19005) in the Czech Republic; Istituto Nazionale di Fisica Nucleare (INFN) in Italy; the National Science, Research and Innovation Fund via the Program Management Unit for Human Resources & Institutional Development, Research and Innovation (grant B37G660014) in Thailand; and the Department of Energy through Awards DE-FG02-97ER41020, DE-FG02-94ER40818, DE-SC0004036, DE-FG02-97ER41033, DE-FG02-97ER41041, DE-SC0011091 and DE-SC0019304 and the Federal Prime Agreement DE-AC02-05CH11231 in the United States. This project has received funding from the European Research Council (ERC) under the European Union Horizon 2020 research and innovation programme (grant agreement No. 852845). We thank the computing cluster support at the Institute for Astroparticle Physics at Karlsruhe Institute of Technology, Max Planck Computing and Data Facility (MPCDF), and the National Energy Research Scientific Computing Center (NERSC) at Lawrence Berkeley National Laboratory.



# Backup

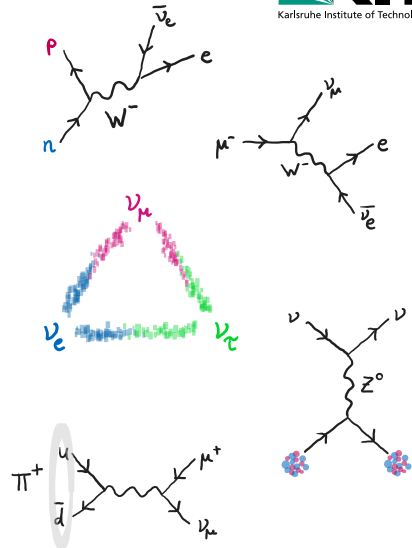


# Search for General Neutrino Interactions

## ■ Possible interaction channels:

- Neutrino oscillation
- LFV in  $\mu^-$ - and  $\tau$ -decays
- Neutrino scatterings, e.g.  $CE\nu NS$
- $\pi$ -decay
- $\beta$ -decay

- Different interaction channels are sensitive to different combinations of  $\epsilon_i$  in GNI Lagrangian.



# Flavour Space Tensor

$j$	$\epsilon_j$	$O_j$	$O'_j$
1	$\epsilon_L$	$\gamma_\mu(1 - \gamma^5)$	$\gamma^\mu(1 - \gamma^5)$
2	$\tilde{\epsilon}_L$	$\gamma_\mu(1 + \gamma^5)$	$\gamma^\mu(1 - \gamma^5)$
3	$\epsilon_R$	$\gamma_\mu(1 - \gamma^5)$	$\gamma^\mu(1 + \gamma^5)$
4	$\tilde{\epsilon}_R$	$\gamma_\mu(1 + \gamma^5)$	$\gamma^\mu(1 + \gamma^5)$
5	$\epsilon_S$	$(1 - \gamma^5)$	1
6	$\tilde{\epsilon}_S$	$(1 + \gamma^5)$	1
7	$-\epsilon_P$	$(1 - \gamma^5)$	$\gamma^5$
8	$-\tilde{\epsilon}_P$	$(1 + \gamma^5)$	$\gamma^5$
9	$\epsilon_T$	$\sigma_{\mu\nu}(1 - \gamma^5)$	$\sigma^{\mu\nu}(1 - \gamma^5)$
10	$\tilde{\epsilon}_T$	$\sigma_{\mu\nu}(1 + \gamma^5)$	$\sigma^{\mu\nu}(1 + \gamma^5)$

$$\mathcal{L}_{GNI}^{CC} = -\frac{G_F V_{\gamma\delta}}{\sqrt{2}} \sum_{j=1}^{10} \left( \overset{(\sim)}{\epsilon}_{j,ud} \right)^{\alpha\beta\gamma\delta} (\bar{e}_\alpha O_j \nu_\beta) (\bar{u}_\gamma O'_j d_\delta) + h.c.$$

- $\epsilon_{L/R}$ : Coupling for left-/right-handed vector-like interactions
- $\epsilon_S$ : Coupling for scalar interactions
- $\epsilon_P$ : Coupling for pseudo-scalar interactions
- $\epsilon_T$ : Coupling for tensor-like interactions

# Effect of GNI parameters on $\beta$ -spectrum

- Effect of neutrino mass and GNI on the tritium  $\beta$ -spectrum.
- $b'_N$  enhances/diminishes kink-like structure due to additional heavier mass state contribution to the spectrum.

