

# Conference Summary and Highlights

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ICHEP 2024

## A wealth of new results!

Soshi Asai - "Lets enjoy our physics!"

281 posters, 918 parallel talks, 40 plenary talks.

1388 Participants

## Selected Highlights

**Sincere apologies to those whose superb result(s) are not presented!**

Thank you to all the speakers for their availability to discuss their results!

**A breathtaking overview of High Energy Physics**

**Marumi Kado**

Max Planck Institute for Physics

# The General Framework

Raymond Volkas

Two main outcomes of the LHC: **The discovery of the Higgs boson and nothing else (so far)!**

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + i \bar{\Psi} \not{D} \Psi + h.c.$$

Simplicity, governed by symmetries only 3 (EW) and 2 (QCD) parameters!

$$+ \bar{\Psi}_i y_{ij} \Psi_j \phi + h.c. + \frac{c_{ij}}{\Lambda} L_i L_j H H ?$$
$$+ |D_\mu \phi|^2 - V(\phi) + \Lambda^4 ?$$

Not governed by symmetries and with **26 parameters set by “hand” of experiments!**

## Open problems

### Hierarchies

- Gauge Hierarchy and Naturalness
- Flavour hierarchy including neutrino masses

### The strong CP problem

$$\theta \frac{\alpha_s}{8\pi} F_{\mu\nu}^A \tilde{F}^{A\mu\nu} \quad \theta < 10^{-10} \quad \text{From neutron electric dipole moment}$$

### The existence of Dark Matter (new field?)

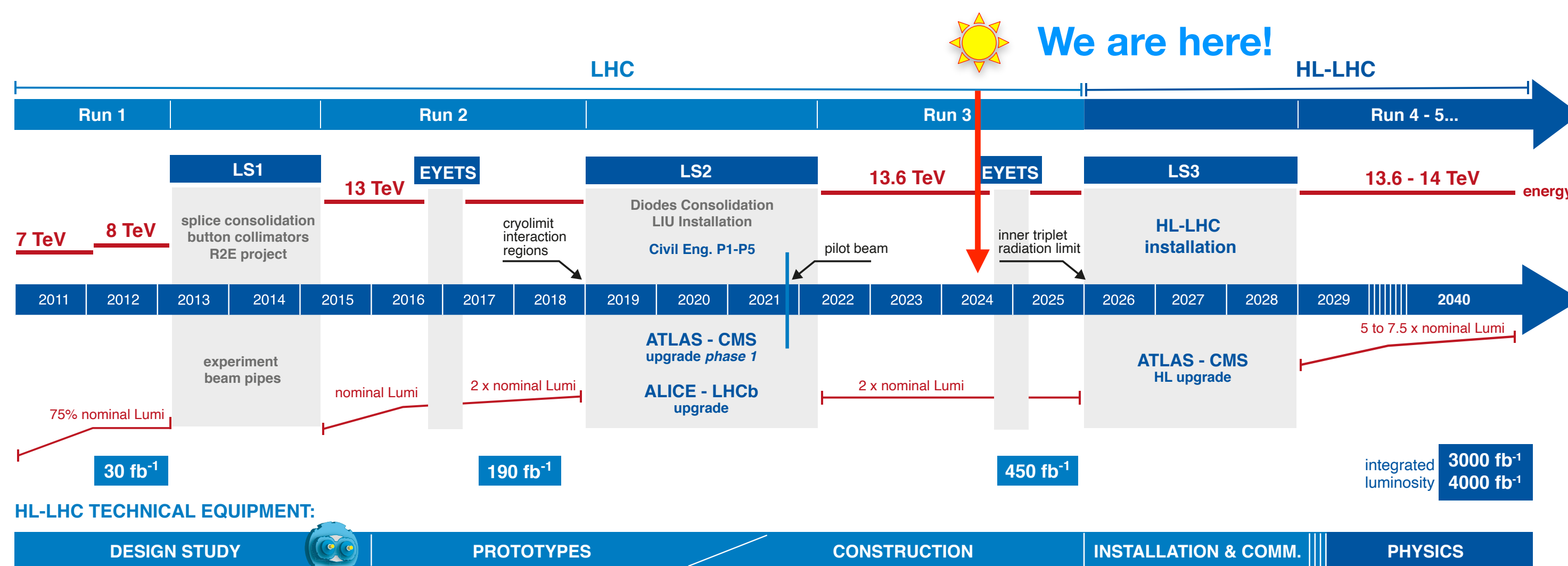
### The nature of Dark energy

## Open questions

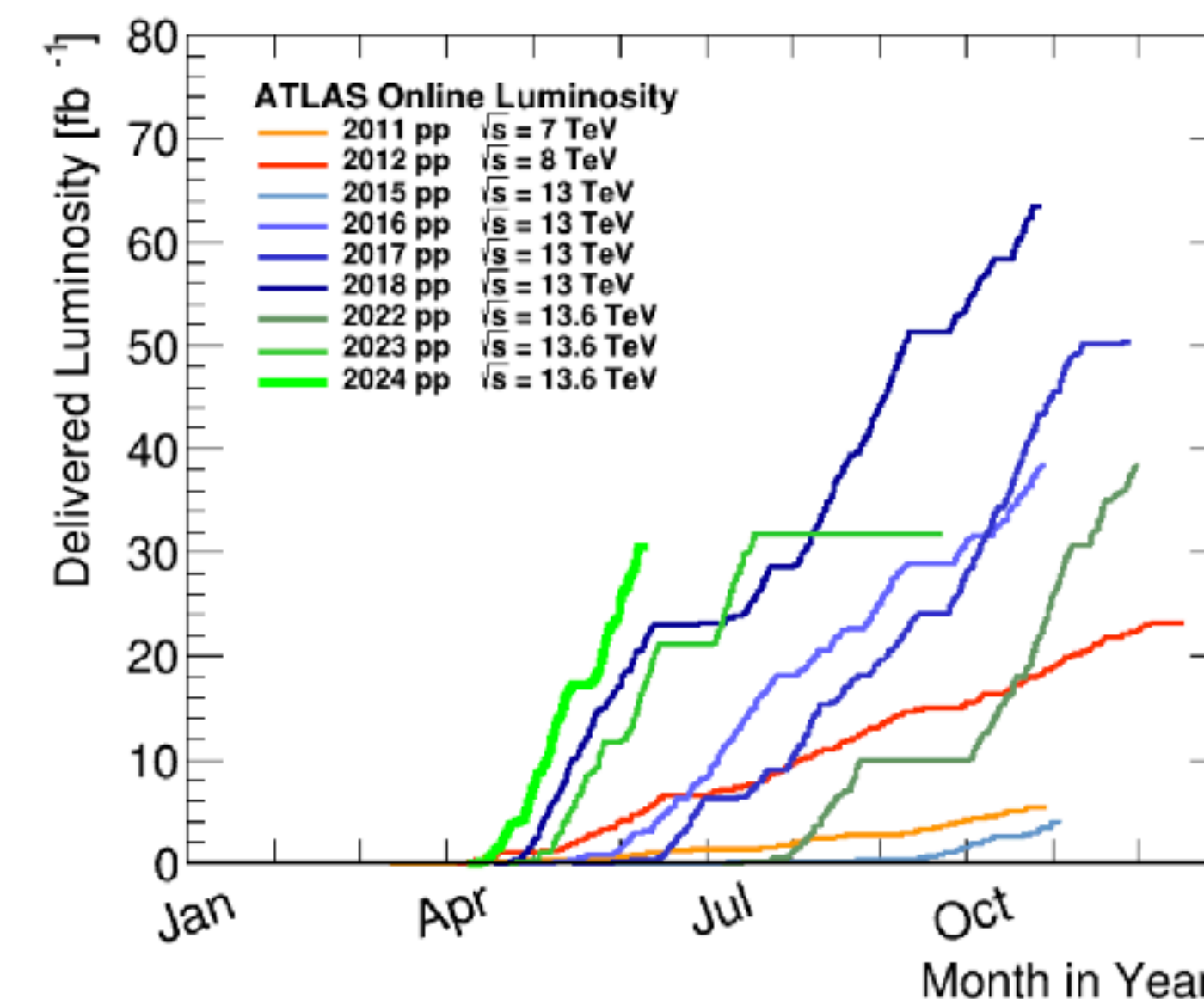
- What is the origin of the asymmetry between matter and anti-matter in the universe?
- What are the properties of QCD confinement?
- Why do electrons have precisely the same charge as the protons?

# Main Colliders - Experimental Landscape

- LHC is refining its results on a **clean and well calibrated dataset of  $\sim 140 \text{ fb}^{-1}$**  at 13 TeV the Run 3 nearly equaling the Run 2 dataset  $\sim 120 \text{ fb}^{-1}$  at 13.6 TeV.



**2024 - High availability operation, Full mastery of considerable inherent operational risks**



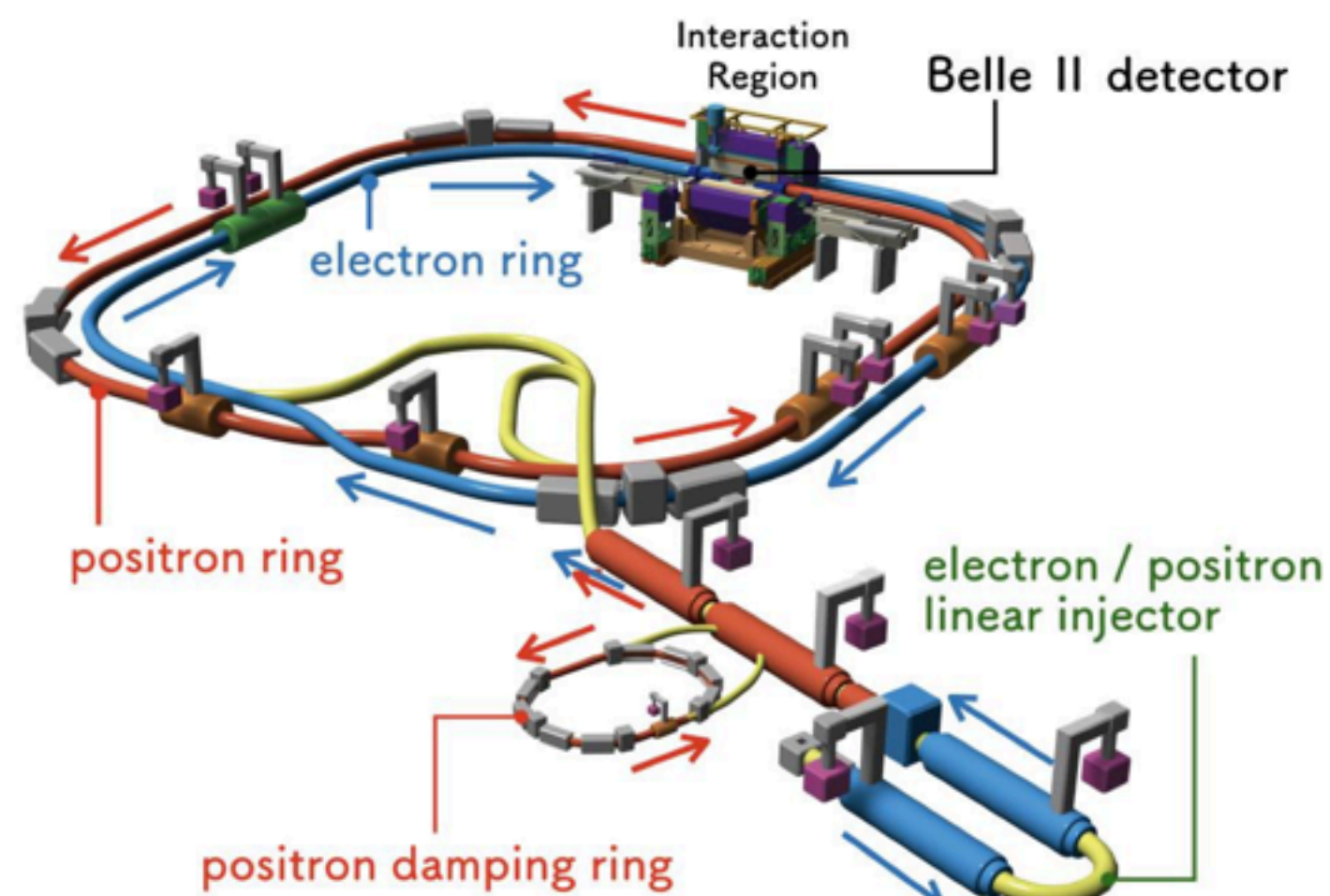
- Currently about half way into entire LHC operations schedule, **Major upgrades leading** to the High Luminosity during the third long shutdown now on the horizon! (See backup for more details)

**Approximately x10 Luminosity delivered (in terms of results x20)**

# Main Colliders - Experimental Landscape

Sergei Nagaitsev

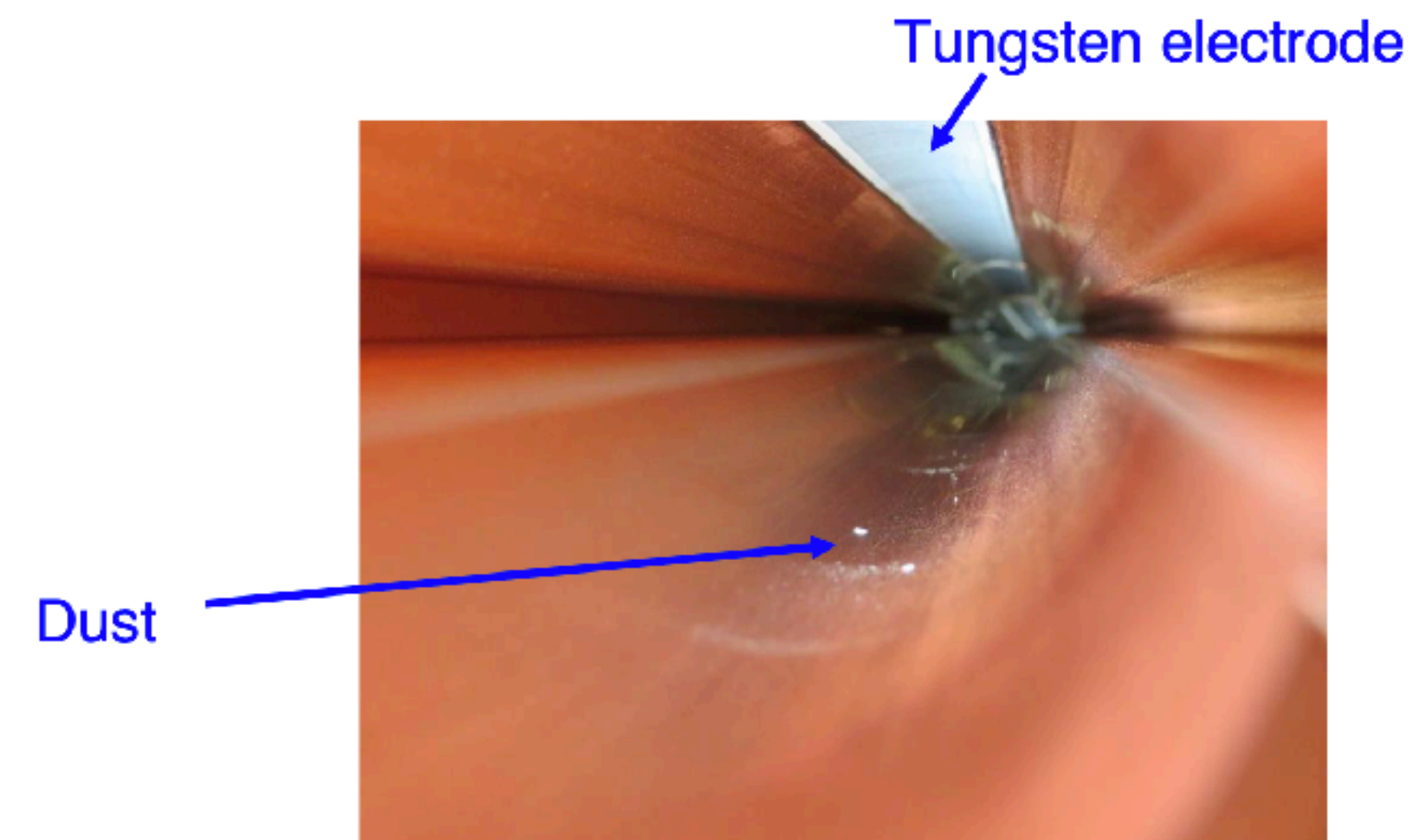
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- Super KEK-B and Belle II world's highest instantaneous luminosity ( $4.7 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$ ), difficult year with a focus on understanding **Sudden Beam Losses** (Belle II is currently running with VTX off).

## Sergei also discussed other facilities:

- RHIC AA (and future EIC, NICA)
- High intensity facilities Fermilab, JPARC, PSI, TRIUMF, CERN PS and SPS
- VEPP and BEPC continuing to bring very useful data (discussed in this talk).
- DAΦNE facility still brings physics potential (PADME) and progress in accelerator R&D!



Countermeasures to the SBL will be implemented during the 2024 summer maintenance period.

# The Higgs Sector - Pillars of Higgs physics

Nicolas Berger, Matthew McCullough

The pillars of Higgs physics:

$\frac{2m_V^2}{v}$

$|\mathcal{D}_\mu \phi|^2$

This term could not exist without a vev

$\frac{m_f}{v}$

$\bar{\Psi}_i y_{ij} \Psi_j \phi + h.c.$

$\frac{3m_H^2}{v}$

$\frac{3m_H^2}{v^2}$

$V(\phi)$

**Message 1** (Matthew McCullough): It is of utmost importance to measure the most precisely measured coupling (hZZ) to probe the Higgs compositeness.

Unambiguous proof of the existence of the Higgs condensate!

	Current	HL-LHC	FCC (ee)
$\kappa_{W,Z}$	6%	1.5%, 1.7 %	0.4%, 0.2 %

**Most precisely known Higgs coupling** tells us how elementary the Higgs boson is!

**The Higgs could well be a pNGB as the pion!**

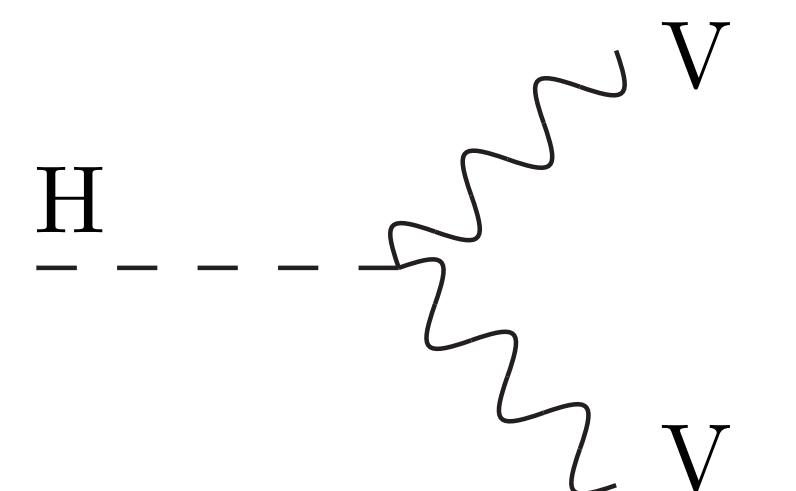
**Message 2** (Matthew McCullough): Precision in Higgs physics is key.

**Precision at HL-LHC is limited by TH (HO, PDFs and  $\alpha_s$ )**

# The Higgs Sector - Pillars of Higgs physics

Nicolas Berger, Matthew McCullough, Sakura Schafer-Nameki

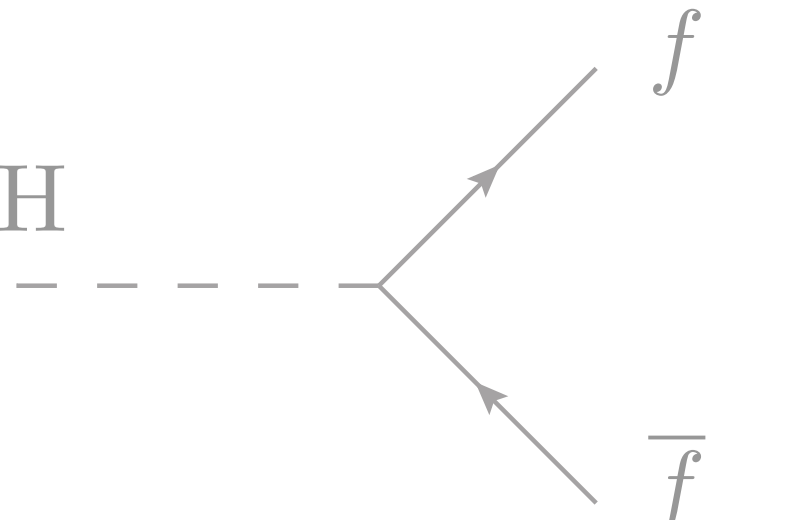
The pillars of Higgs physics:



$\frac{2m_V^2}{v}$

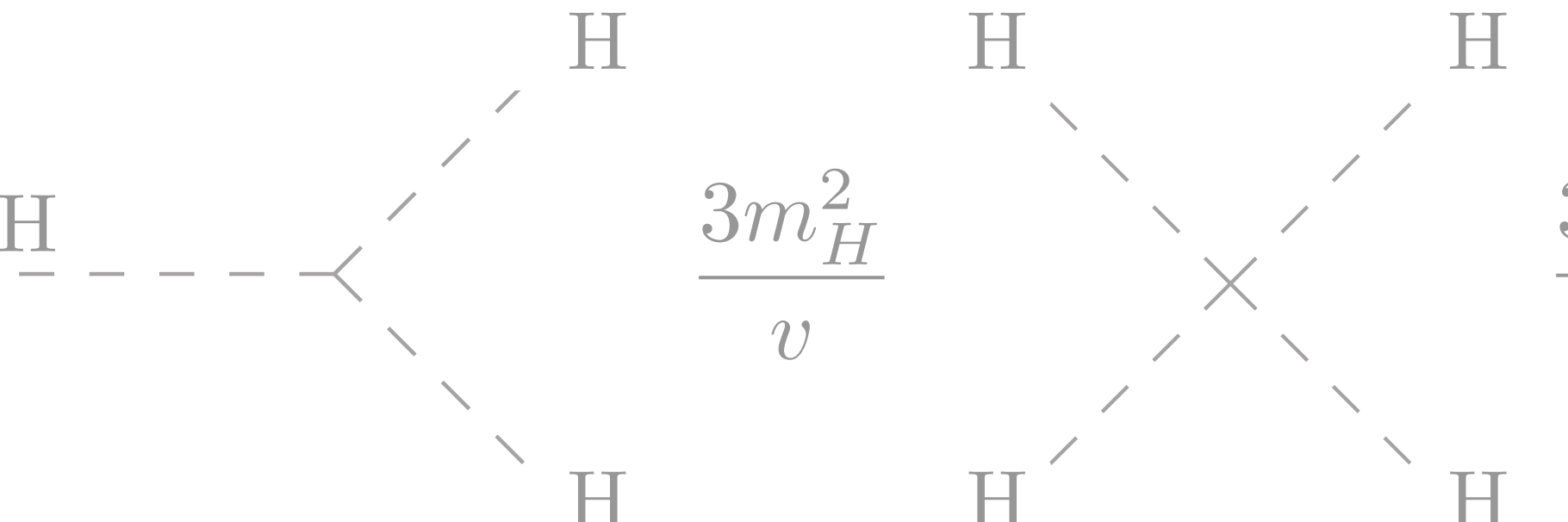
$|\partial_\mu \phi|^2$

This term could not exist without a vev



$\frac{m_f}{v}$

$\bar{\Psi}_i \gamma_{ij} \Psi_j \phi + h.c.$



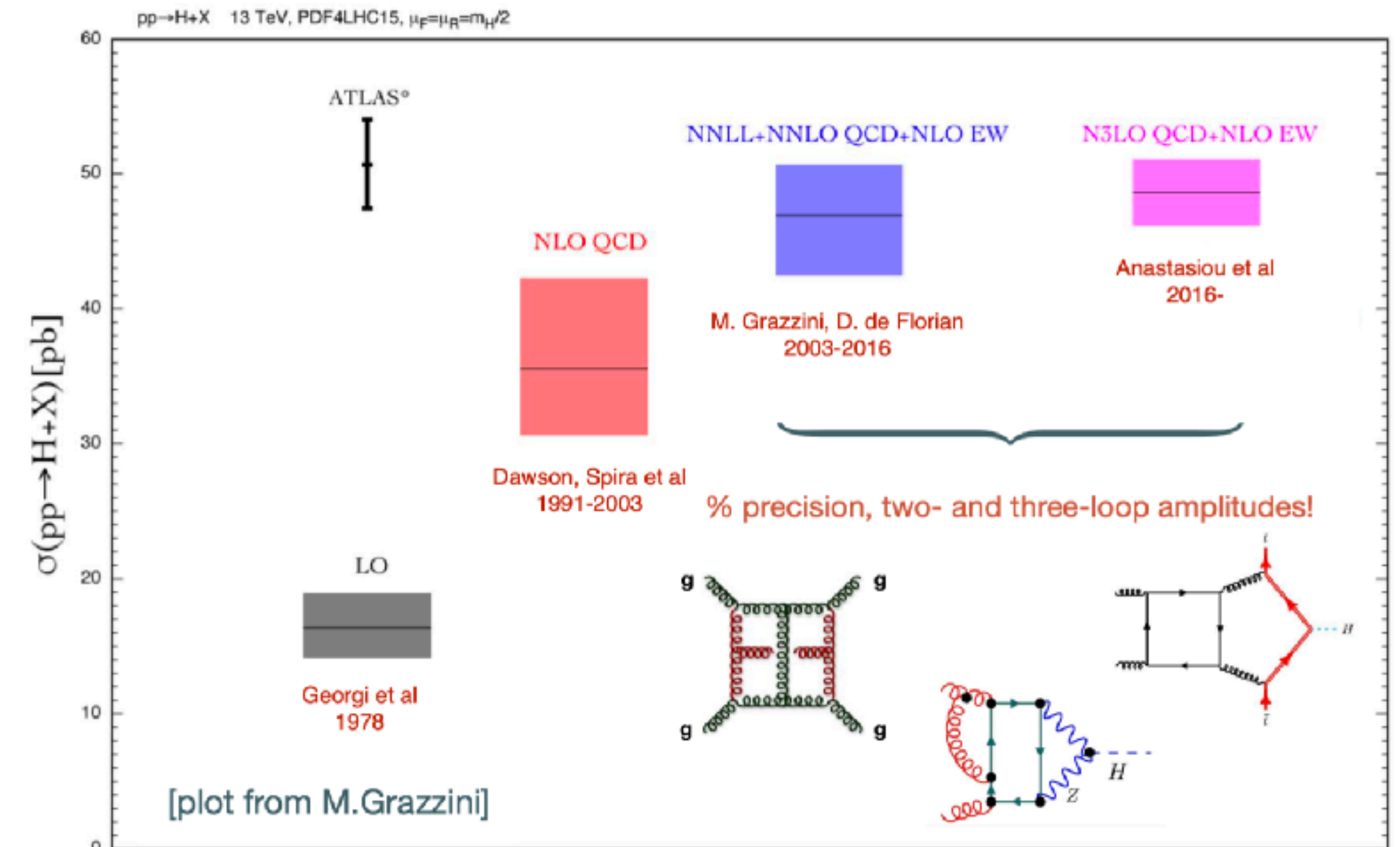
$\frac{3m_H^2}{v}$

$\frac{3m_H^2}{v^2}$

$V(\phi)$

**Message 1** (Matthew McCullough): It is of utmost importance to measure the most precisely measured coupling (hZZ) to probe the Higgs compositeness.

**Message 2** (Sakura Schafer-Nameki): Much progress thanks to formal theory: **amplitudes!**



# The Higgs Sector - Pillars of Higgs physics

Nicolas Berger, Matthew McCullough, Maurizio Pierini

The pillars of Higgs physics:

Diagram showing a Higgs boson (H) decaying into two photons ( $\gamma$ ) via a loop of a vector boson (V). The coupling is given by  $\frac{2m_V^2}{v}$ .

Handwritten note:  $|\mathcal{D}_\mu \phi|^2$

This term could not exist without a vev

Diagram showing a Higgs boson (H) decaying into a fermion (f) and an antifermion ( $\bar{f}$ ). The coupling is given by  $\frac{m_f}{v}$ .

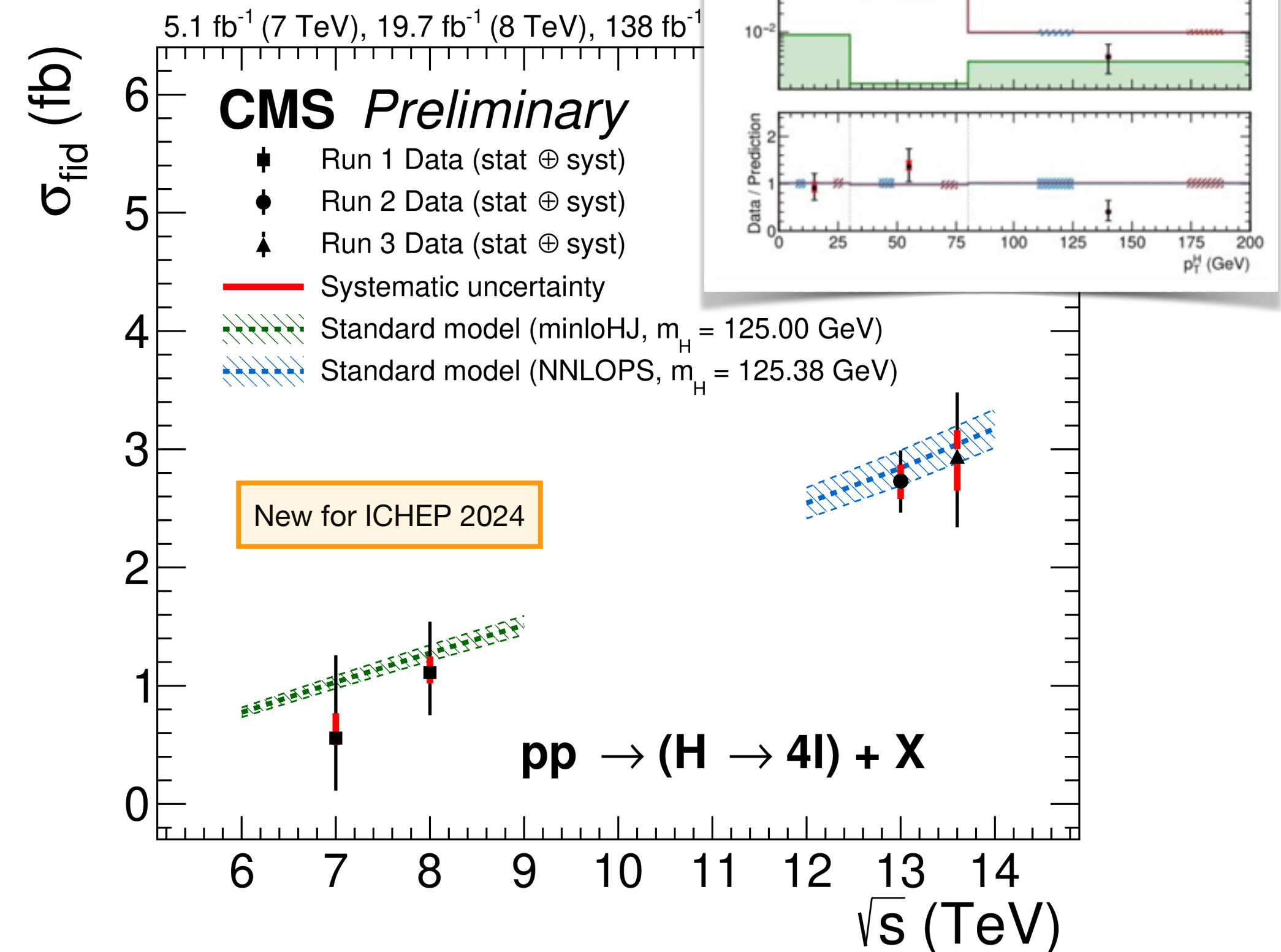
Handwritten note:  $\bar{\Psi}_i y_{ij} \Psi_j \phi + h.c.$

Diagrams showing Higgs boson (H) production and decay into four Higgs bosons (H) via s-channel and t-channel. The coupling is given by  $\frac{3m_H^2}{v}$  and  $\frac{3m_H^2}{v^2}$ .

Handwritten note:  $V(\phi)$

**Message 1** (Matthew McCullough): It is of utmost importance to measure the most precisely measured coupling (hZZ) to probe the Higgs compositeness.

**New measurements at 13.6 TeV (Run 3)**



# The Higgs Sector - Pillars of Higgs physics

Nicolas Berger, Matthew McCullough

The pillars of Higgs physics:

$\frac{2m_V^2}{v}$

$|\mathcal{D}_\mu \phi|^2$

This term could not exist without a vev

$\frac{m_f}{v}$

$\bar{\Psi}_i y_{ij} \Psi_j \phi + h.c.$

$\frac{3m_H^2}{v}$

$\frac{3m_H^2}{v^2}$

$V(\phi)$

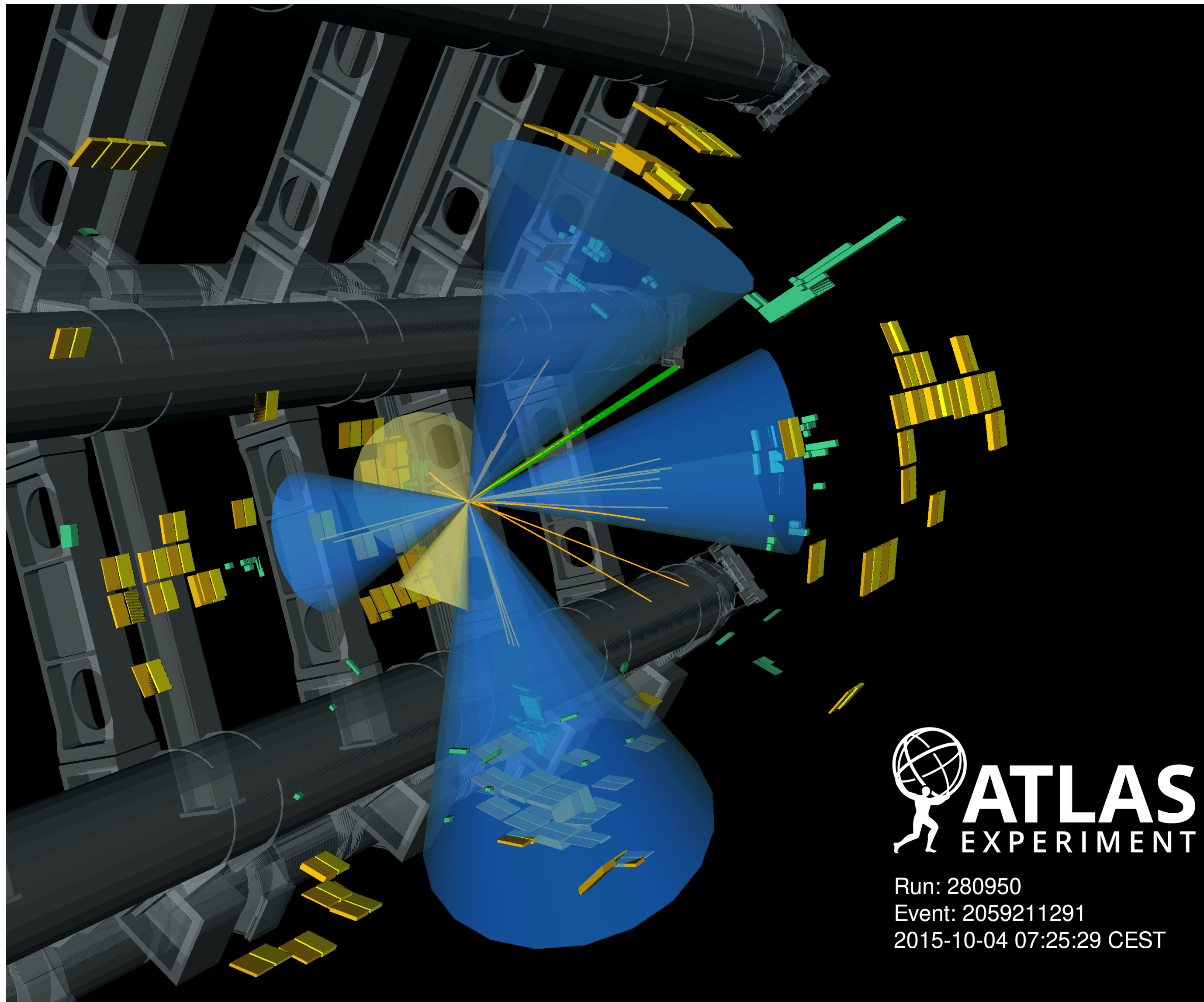
	Current	HL-LHC	FCC (ee)	
$\kappa_{W,Z}$	6%	1.5%, 1.7 %	0.4%, 0.2 %	
$\kappa_t$	11%	3.4%	-	
$\kappa_b$	11%	3.7%	0.7%	
$\kappa_\tau$	8%	1.9%	0.7%	
$\kappa_\mu$	20%	4.3%	8.9%*	
		HL-LHC	FCC (ee)	FCC (hh)



# Top Yukawa Coupling at the LHC

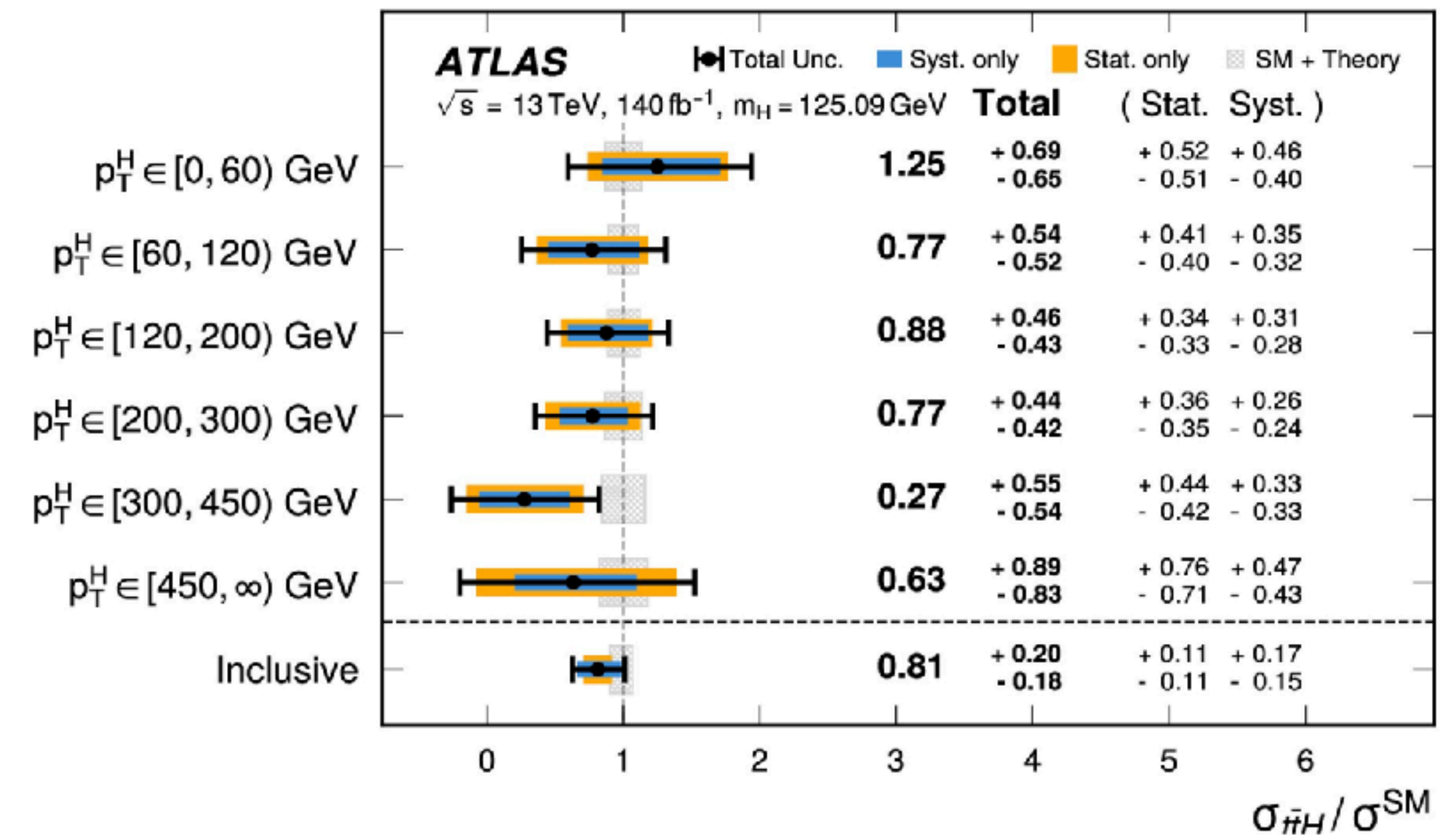
Nicolas Berger, Monica Dunford

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## ttH news from ATLAS!

**Very complex final state** Main challenge: tt+bb background



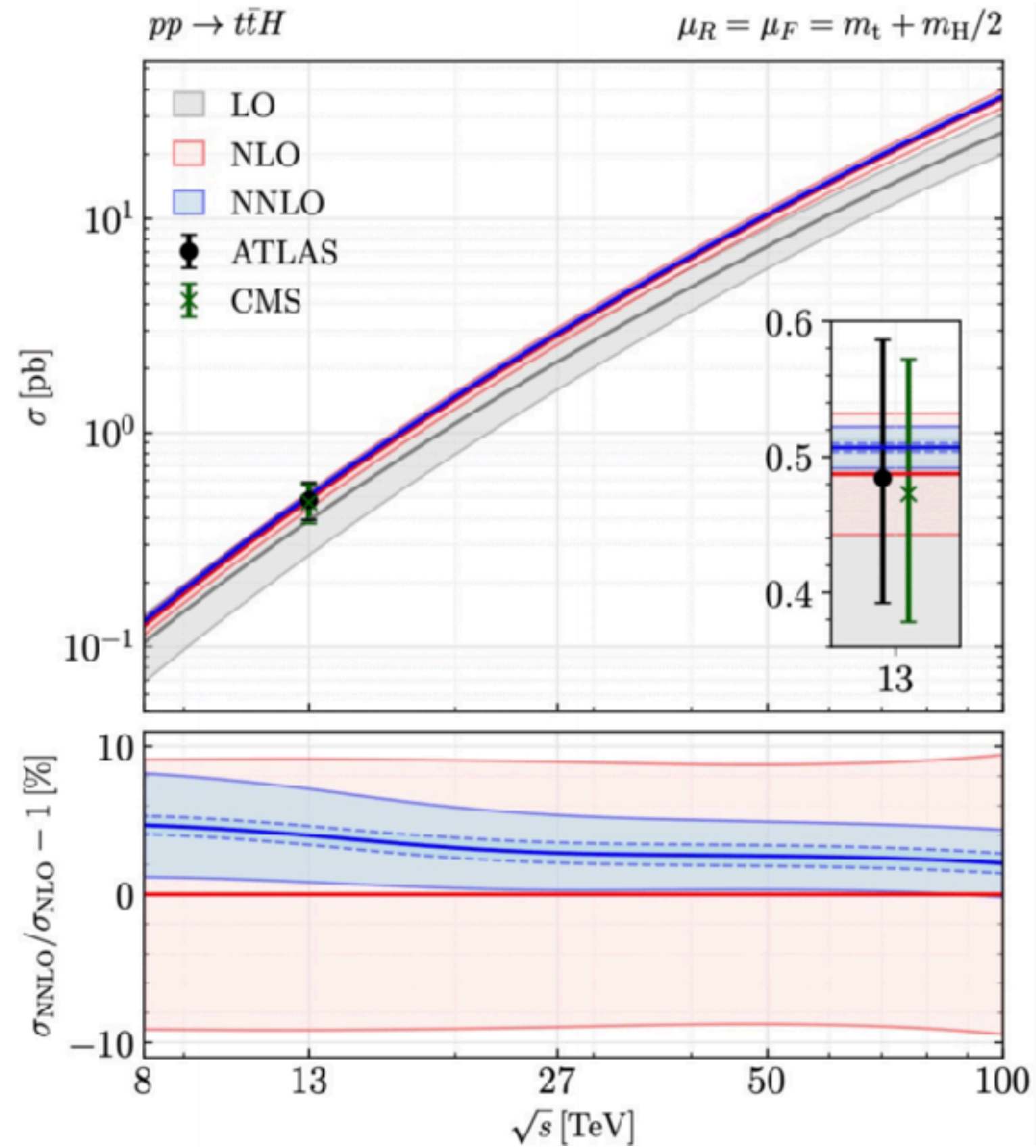
**Best single ttH measurement!**  
Overall uncertainty improved by factor 1.8, 4.6 $\sigma$

# Top Yukawa Coupling at the LHC

Nicolas Berger, Matthew McCullough, Monica Dunford, Francesco di Bello

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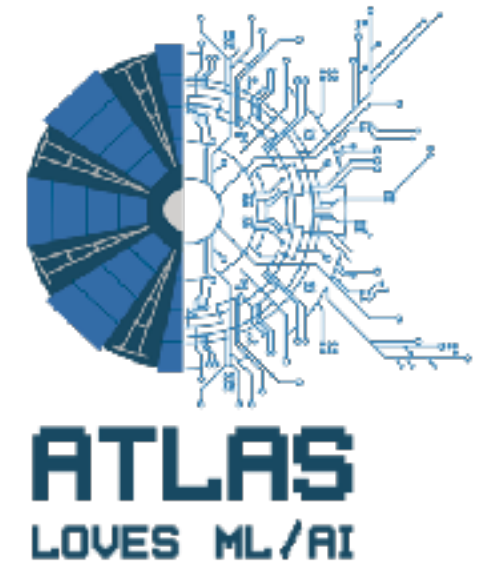
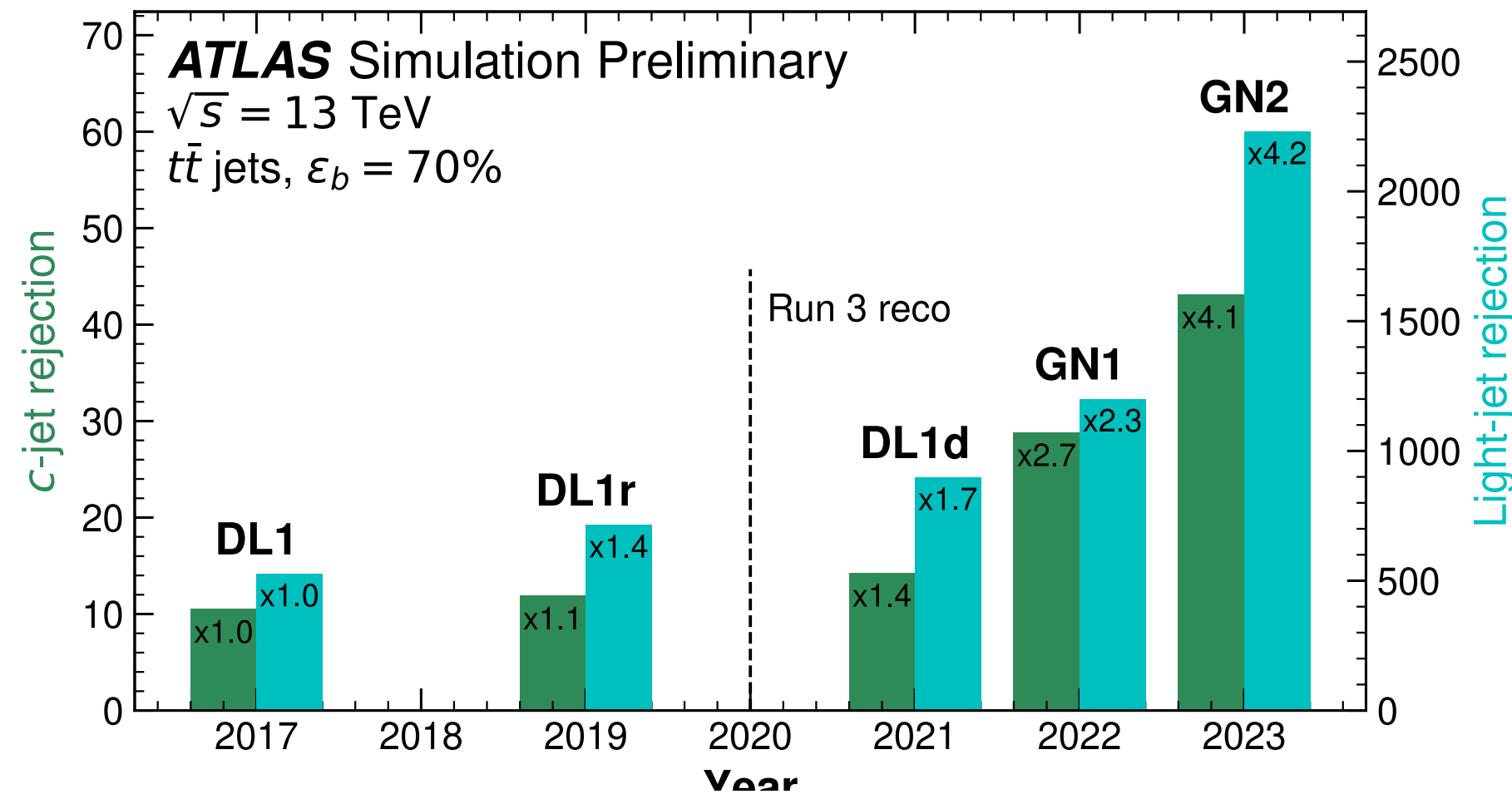
## Recent example from Mathew (ttH)



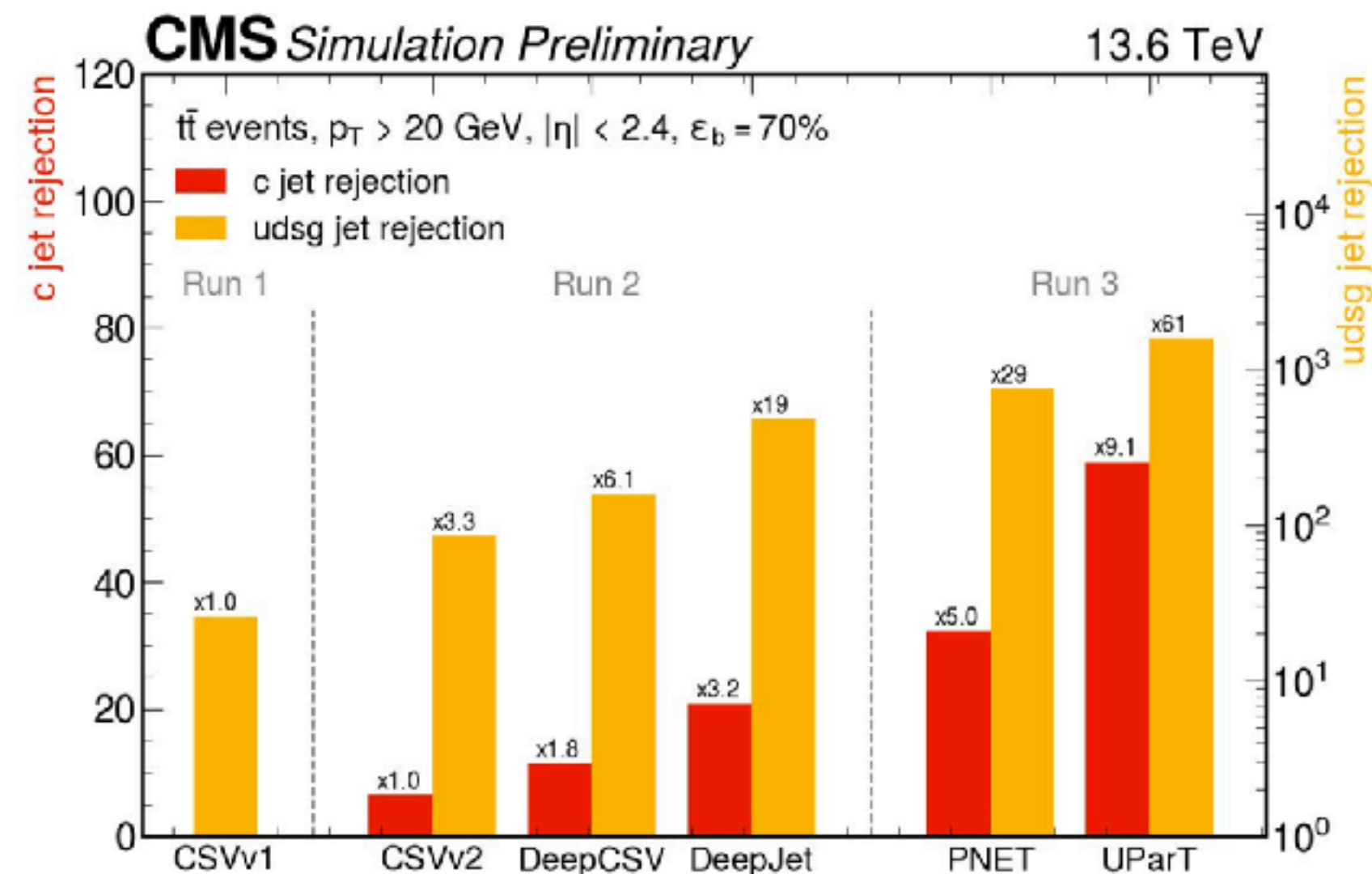
**Message 2 (again)** (Matthew McCullough): Precision in Higgs physics is key.

## ttH from ATLAS

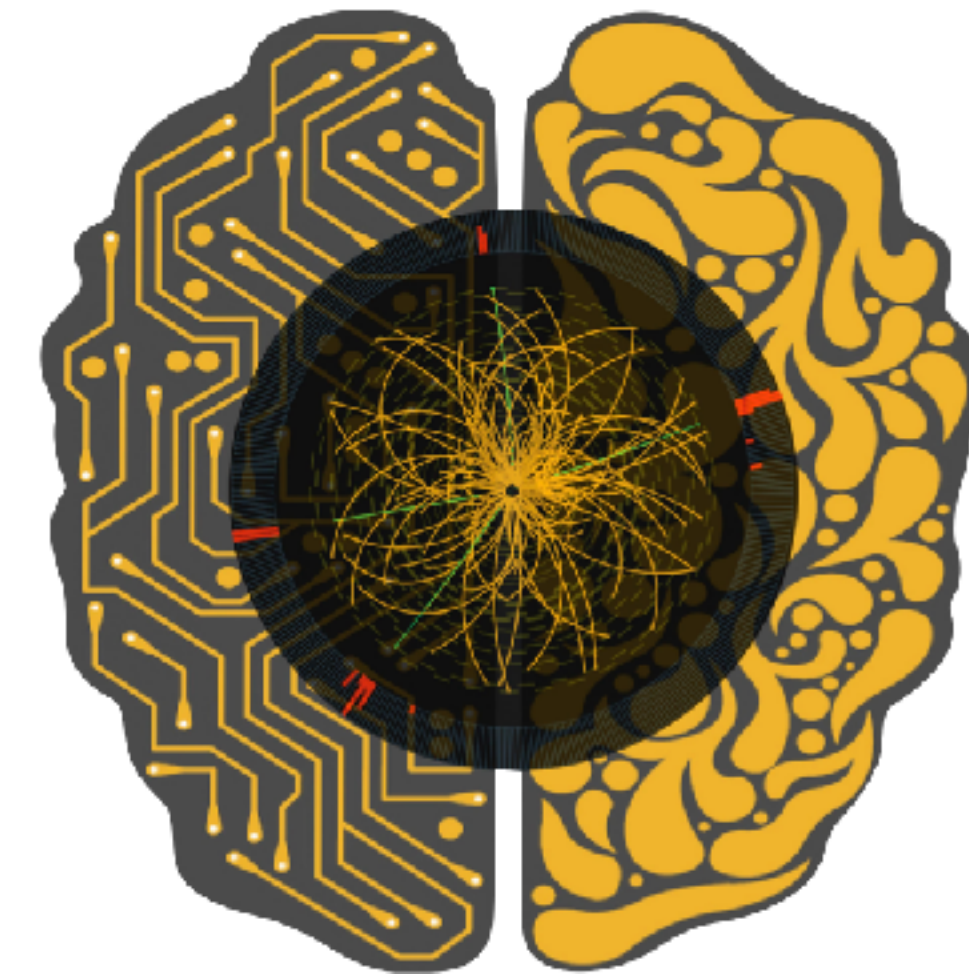
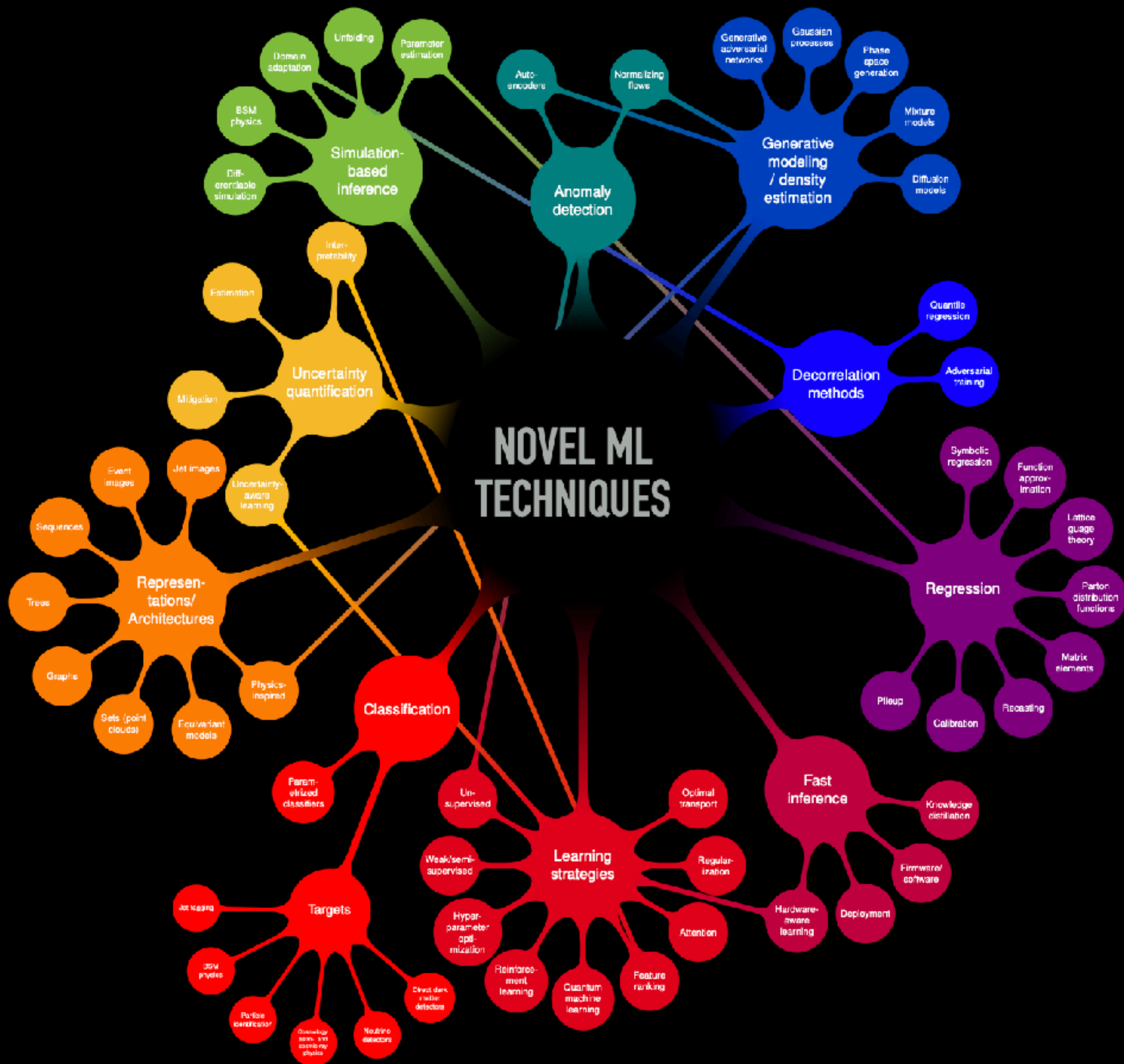
AI in HEP reconstruction has a significant impact!



The most shown plot at this conference!



There are 4 b-quark jets in the  $t\bar{t}H(bb)$  event topology!



Array of ML opportunities beyond classification and regression, in simulation, unfolding, anomaly detection, etc.

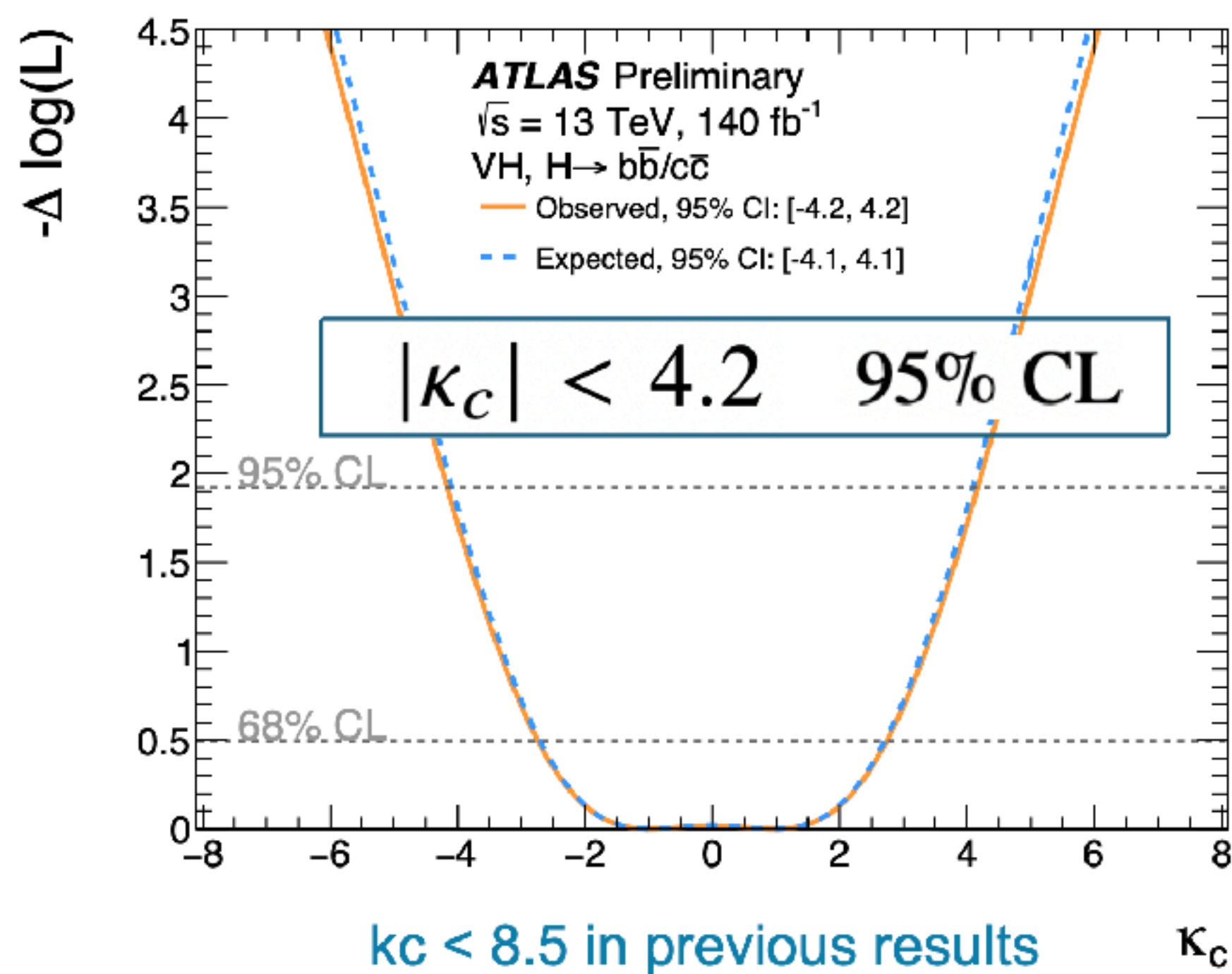
New ideas also have to be concerned with **robustness and interpretability**

# Yukawa Coupling to Charm at the LHC

Nicolas Berger, Monica Dunford, Francesco di Bello

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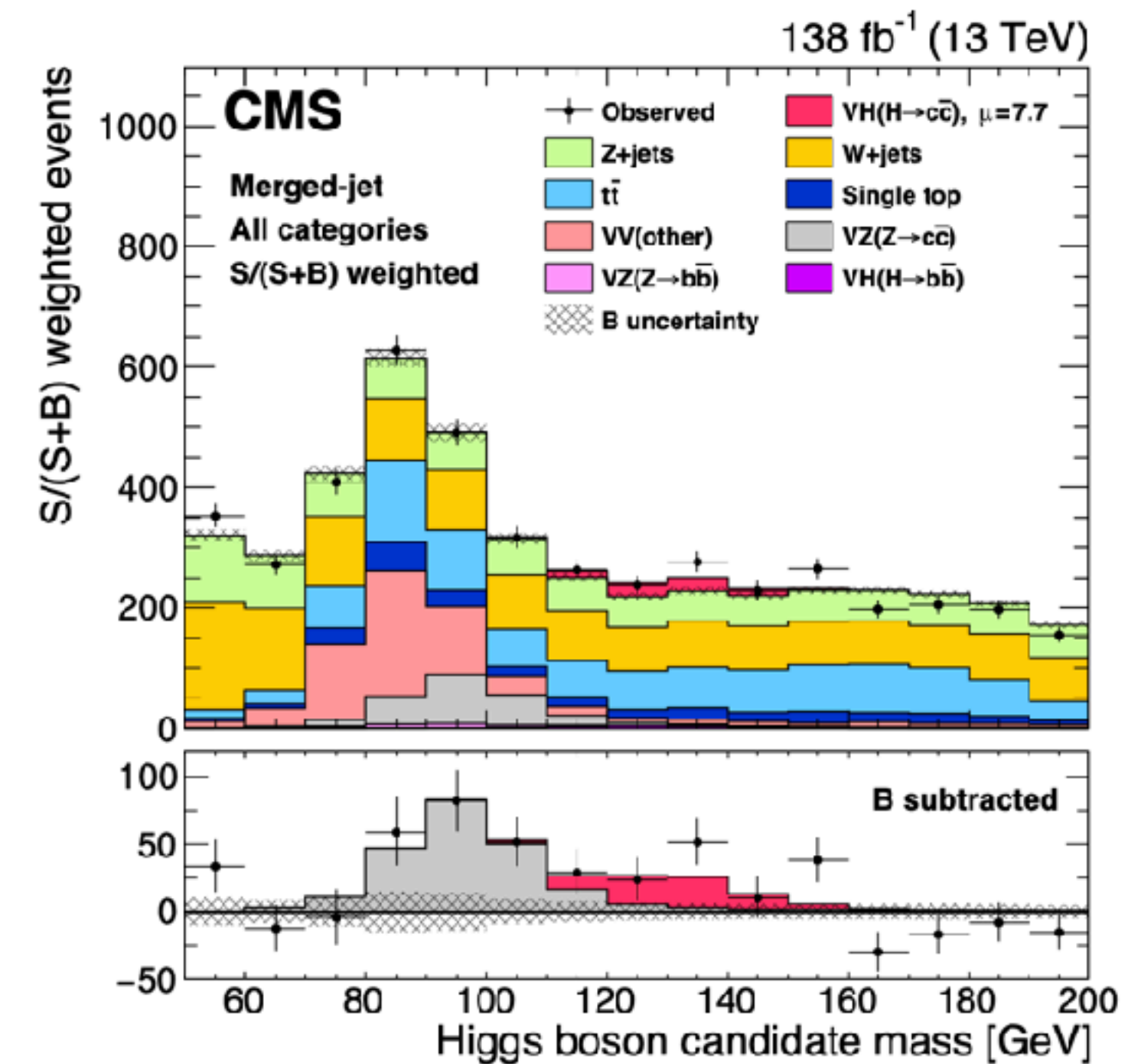
Refined analysis of Run 2 data with now Graph NN charm tagging!



$$\mu_{VH}^{cc} = 1.0^{+5.4}_{-5.2} = 1.0^{+4.0}_{-3.9} (\text{stat.})^{+3.6}_{-3.5} (\text{syst.}).$$

Improvement by a factor of 2 w.r.t. previous result

Use of state-of-the-art ML techniques [Particle Net](#) uses Dynamic Graph CNN



Constraints on charm Yukawa  $1.1 < \kappa_c < 5.5$

Yields a precision on  $\kappa_c$  of  $\sim 40\%$  per experiment at HL-LHC

New perspective at the LHC!

# The Higgs Sector - Pillars of Higgs physics

Nicolas Berger, Francesco di Bello

The pillars of Higgs physics:

$\frac{m_f}{v}$

$\bar{\Psi}_i y_{ij} \Psi_j \phi + h.c.$

$\frac{2m_V^2}{v}$

$|D_\mu \phi|^2$

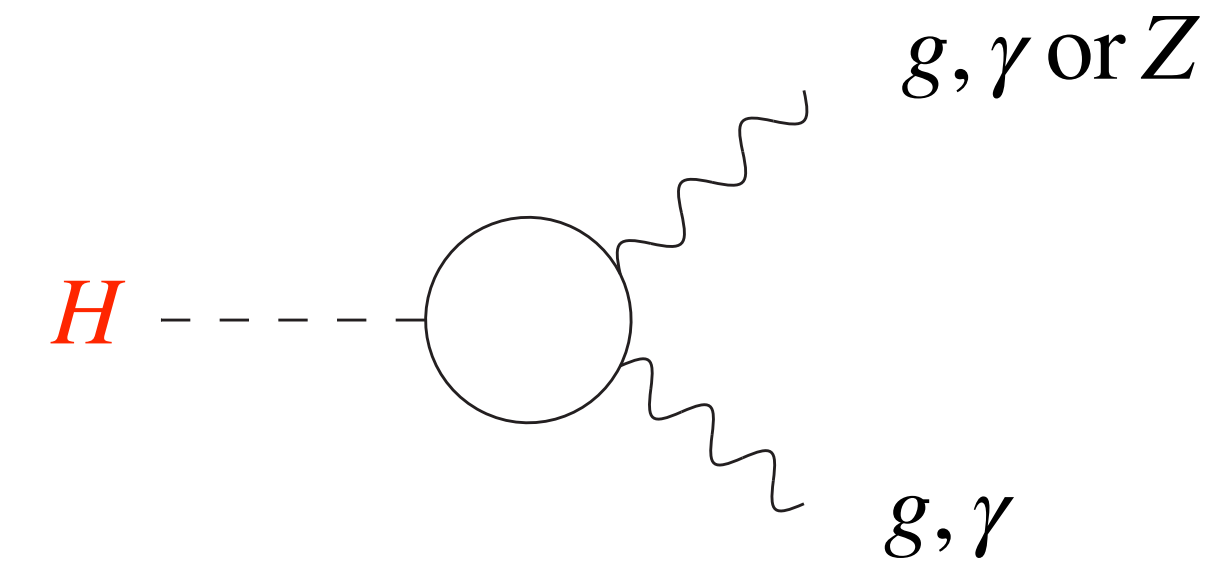
This term could not exist without a vev

$\frac{3m_H^2}{v}$

$\frac{3m_H^2}{v^2}$

$V(\phi)$

Probing new particles through loops in production and decays!



	Current	HL-LHC	FCC (ee)
$\kappa_\gamma$	6%	1.8%	3.9%
$\kappa_g$	7%	2.5%	1%
$\kappa_{Z\gamma}$	30%	9.8%	

# The Higgs Sector - Pillars of Higgs physics

Cristina Mantilla-Suarez, Dilla Maria Portillo Quintero

The pillars of Higgs physics:

$H \rightarrow f \bar{f}$       $\frac{m_f}{v}$       $\bar{\Psi}_i y_{ij} \Psi_j \phi + h.c.$

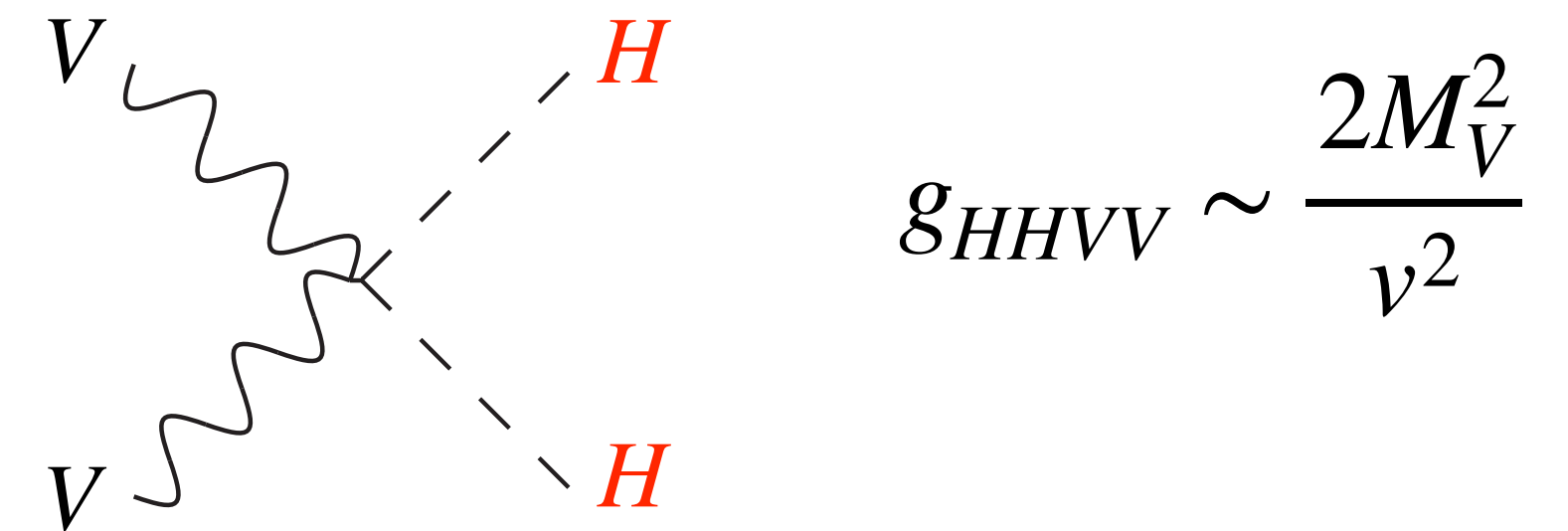
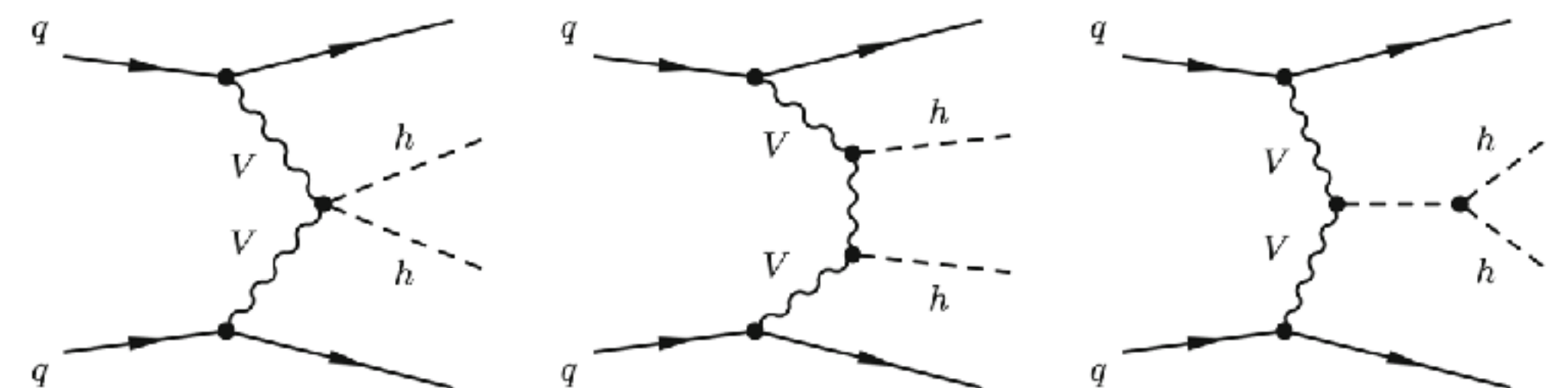
$H \rightarrow V V$       $\frac{2m_V^2}{v}$       $|D_\mu \phi|^2$      This term could not exist without a vev

$H \rightarrow H H$       $\frac{3m_H^2}{v}$       $H H \rightarrow H H$       $\frac{3m_H^2}{v^2}$       $V(\phi)$

## Non vanishing di-Higgs (to VV) coupling!

Without observing HH production

Done in VBF(HH) production with a significant negative interference with



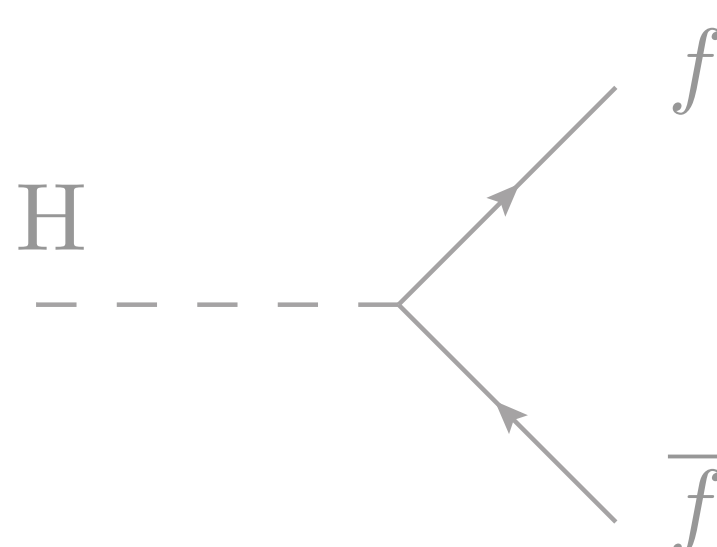
$k_{2V} \in [0.67, 1.38]$

CMS result (ATLAS similar)

# The Higgs boson self coupling!

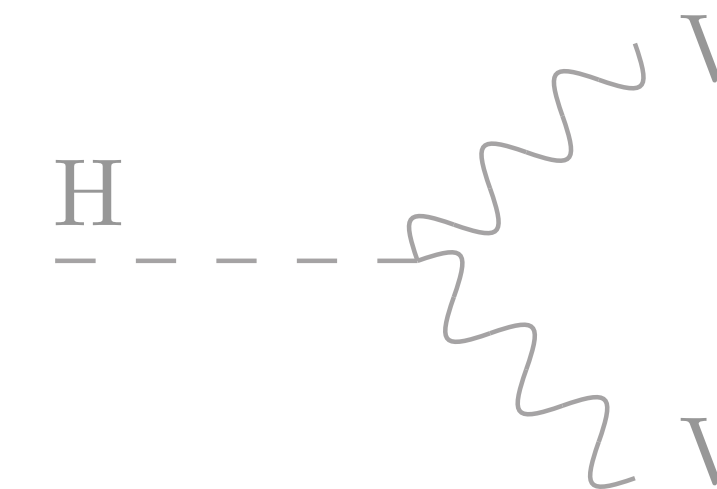
Nicolas Berger, Matthew McCullough

The pillars of Higgs physics:



A Feynman diagram showing a Higgs boson (H) decaying into a fermion (f) and an antifermion ( $\bar{f}$ ). The coupling is given by  $\frac{m_f}{v}$ .

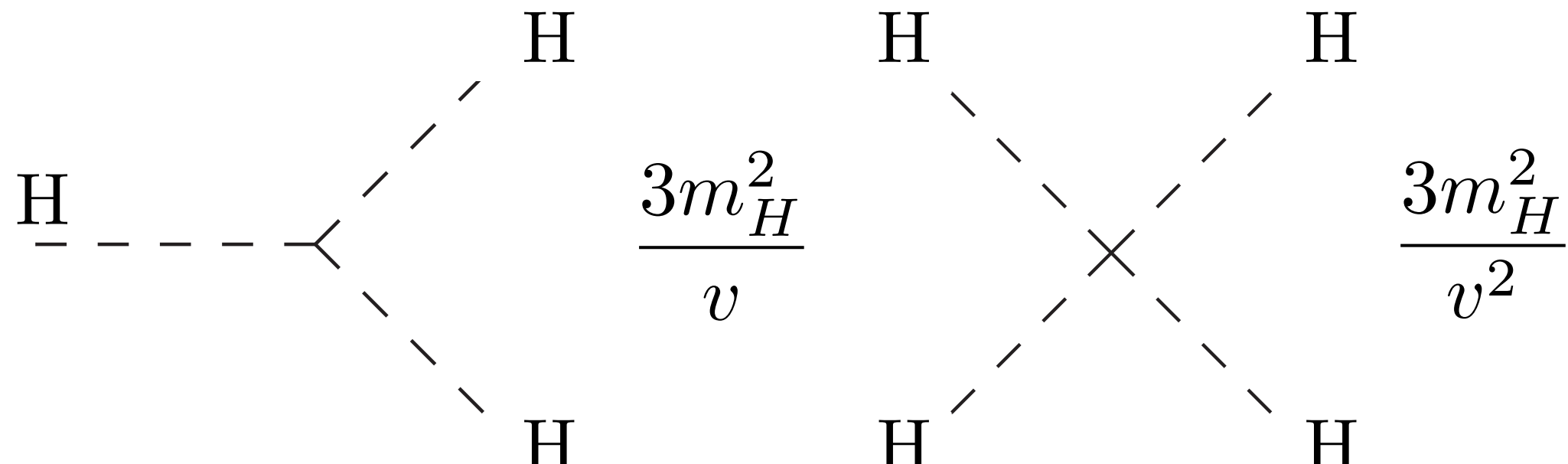
$\bar{\Psi}_i y_{ij} \Psi_j \phi + h.c.$



A Feynman diagram showing a Higgs boson (H) decaying into two vector bosons (V). The coupling is given by  $\frac{2m_V^2}{v}$ .

$|\mathcal{D}_\mu \phi|^2$

This term could not exist without a vev



Two Feynman diagrams showing Higgs boson self-couplings. The first diagram shows a Higgs boson (H) decaying into two Higgs bosons (H) with a coupling of  $\frac{3m_H^2}{v}$ . The second diagram shows a Higgs boson (H) decaying into two Higgs bosons (H) with a coupling of  $\frac{3m_H^2}{v^2}$ .

$V(\phi)$

Despite the fact that in “Vanilla SUSY and vanilla composite models it is difficult to have large deviations in trilinear w.r.t. vector boson coupling”

*Matthew McCullough*

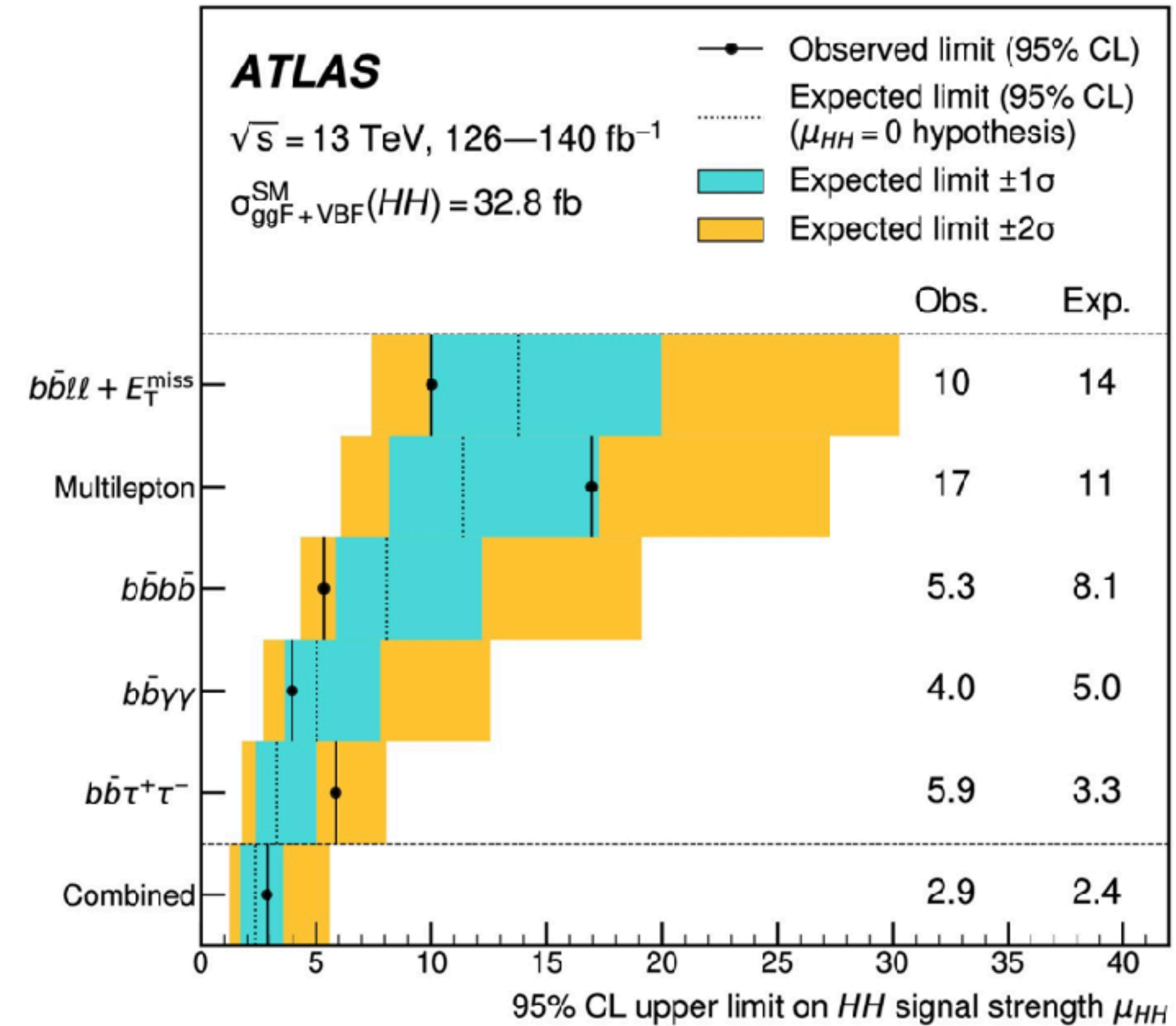
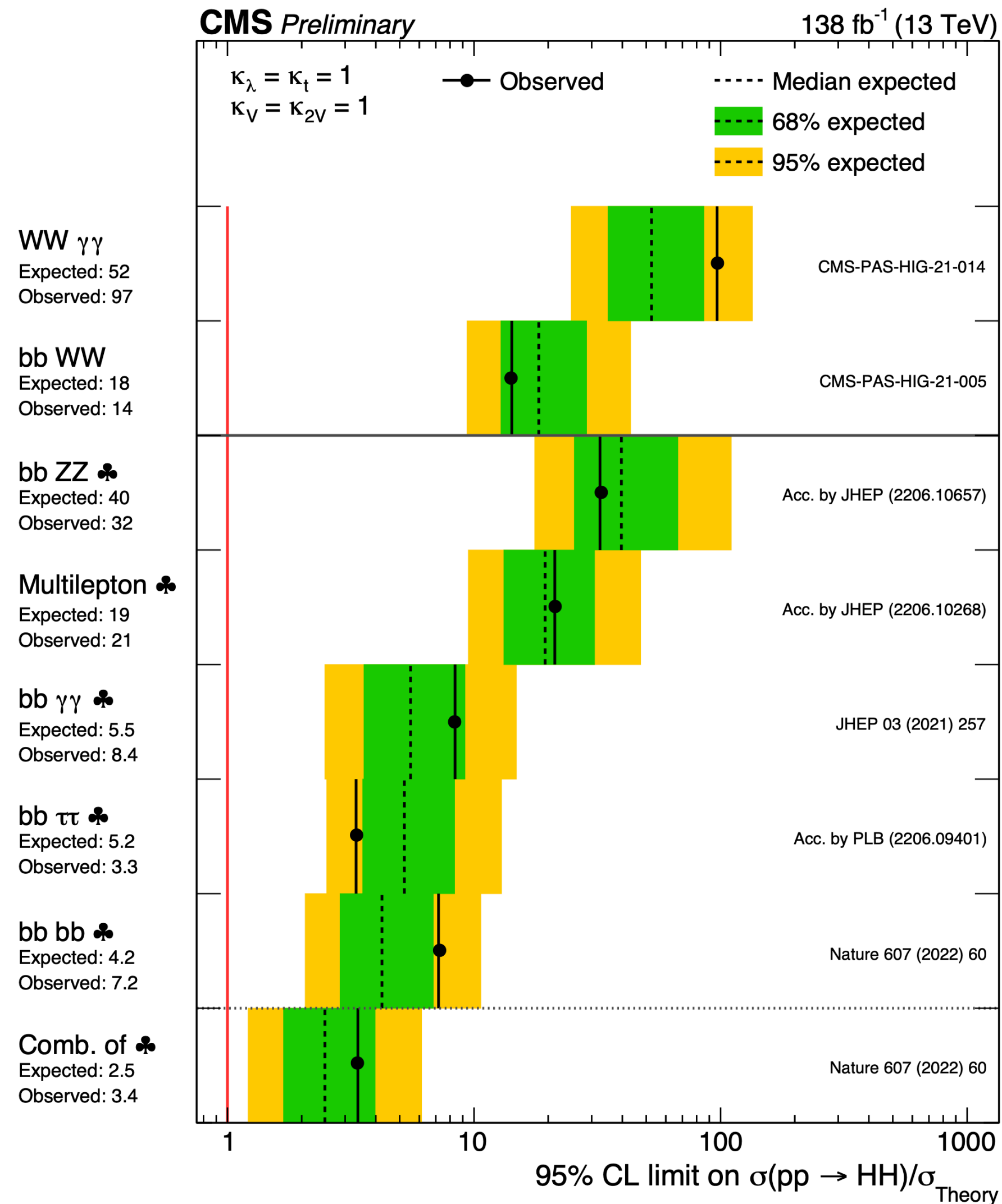
**Message 3** (Matthew McCullough - as well as Georg Weiglein in parallel session): Large trilinear deviations are possible while deviations of the Higgs to Z coupling remain small.

“Arguably the most important of them all!”

# Higgs Self Coupling and HH Production

Nicolas Berger, Matthew McCullough

“Arguably the most important of them all!”



Observed limits start deviating from expectation!!

Both experiments have  $\sim 1\sigma$  sensitivity to a signal (Obs. ATLAS  $0.4\sigma$  and CMS  $\sim 1\sigma$ ) with Run 2!!

Naive comb. ATLAS-CMS sensitivity with Run 3 close  $2.5\sigma$  with improvements (and as much data as possible) **aim at  $3\sigma$**

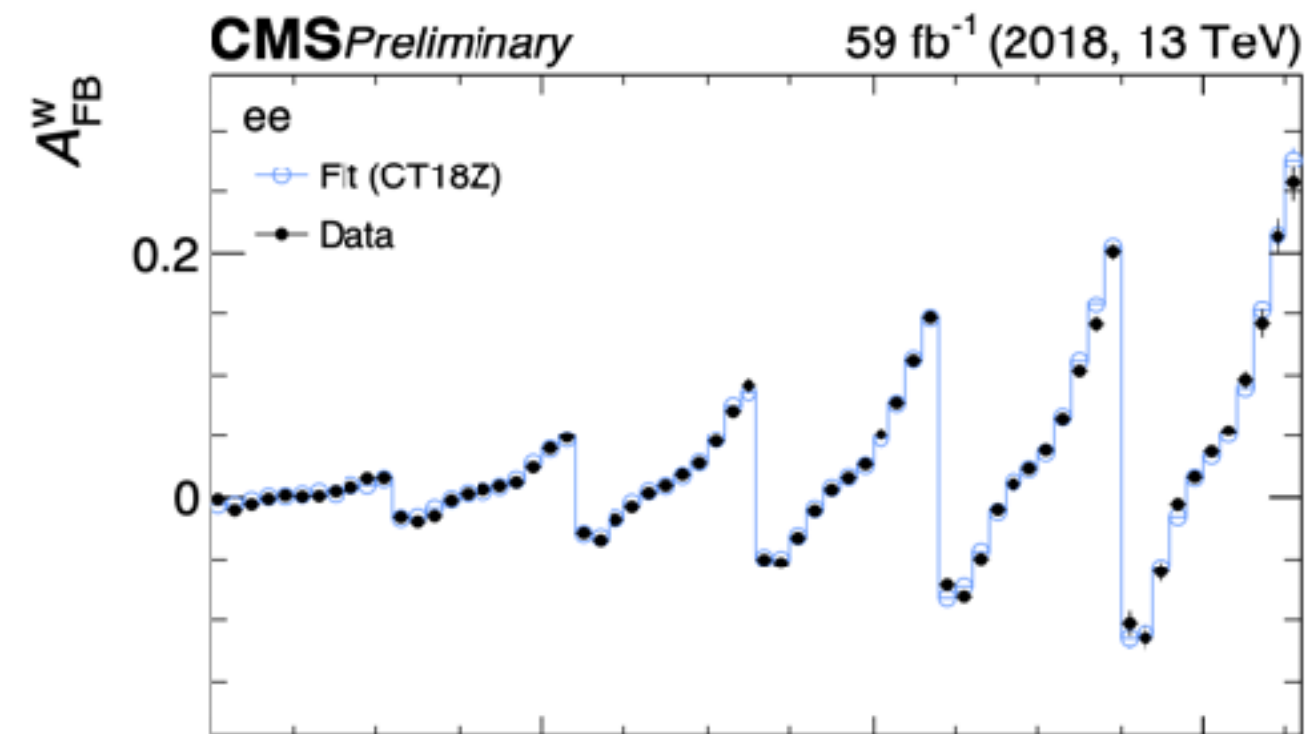


# The Gauge Sectors - LHC at Precision Frontier

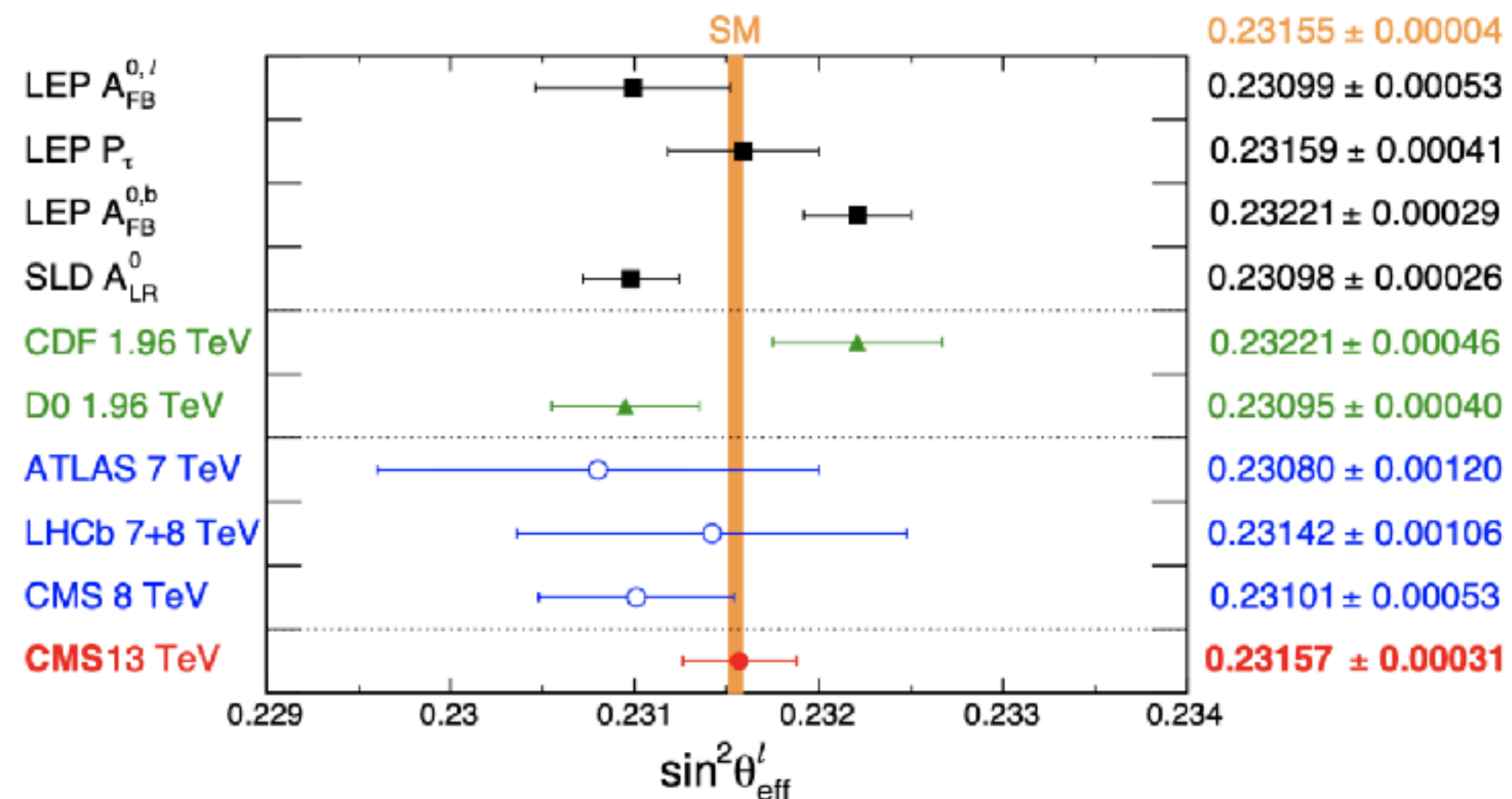
Maurizio Pierini, Yasmine Amhis

## New measurements of $\sin^2 \theta_W$ by CMS and LHCb through $A_{FB}$

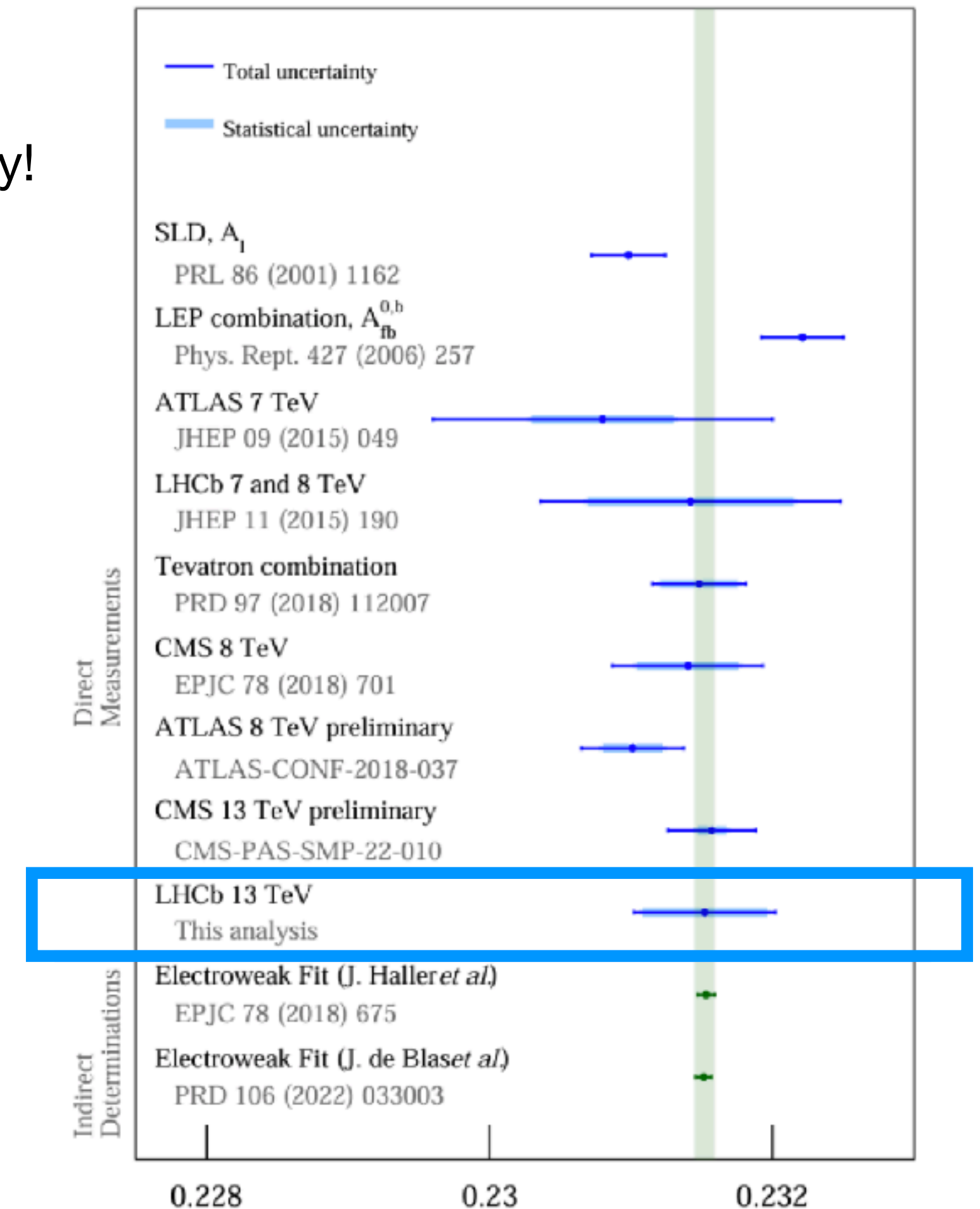
$A_{FB}$  in pp collisions is a tricky question, the forward region direction is given by the valence quark i.e. the system boost direction.



Precision comparable to the most precise single measurements at LEP  $A_{FB}^b$  and SLD  $A_{LR}$  determination



CERN Seminar yesterday!  
[Link](#)



$$\sin^2 \theta_{eff}^{\ell} = 0.23152 \pm 0.00044 \pm 0.00005 \pm 0.00022$$

# Precise Determination of $\alpha_s$ using $Z \rightarrow \ell^+ \ell^-$

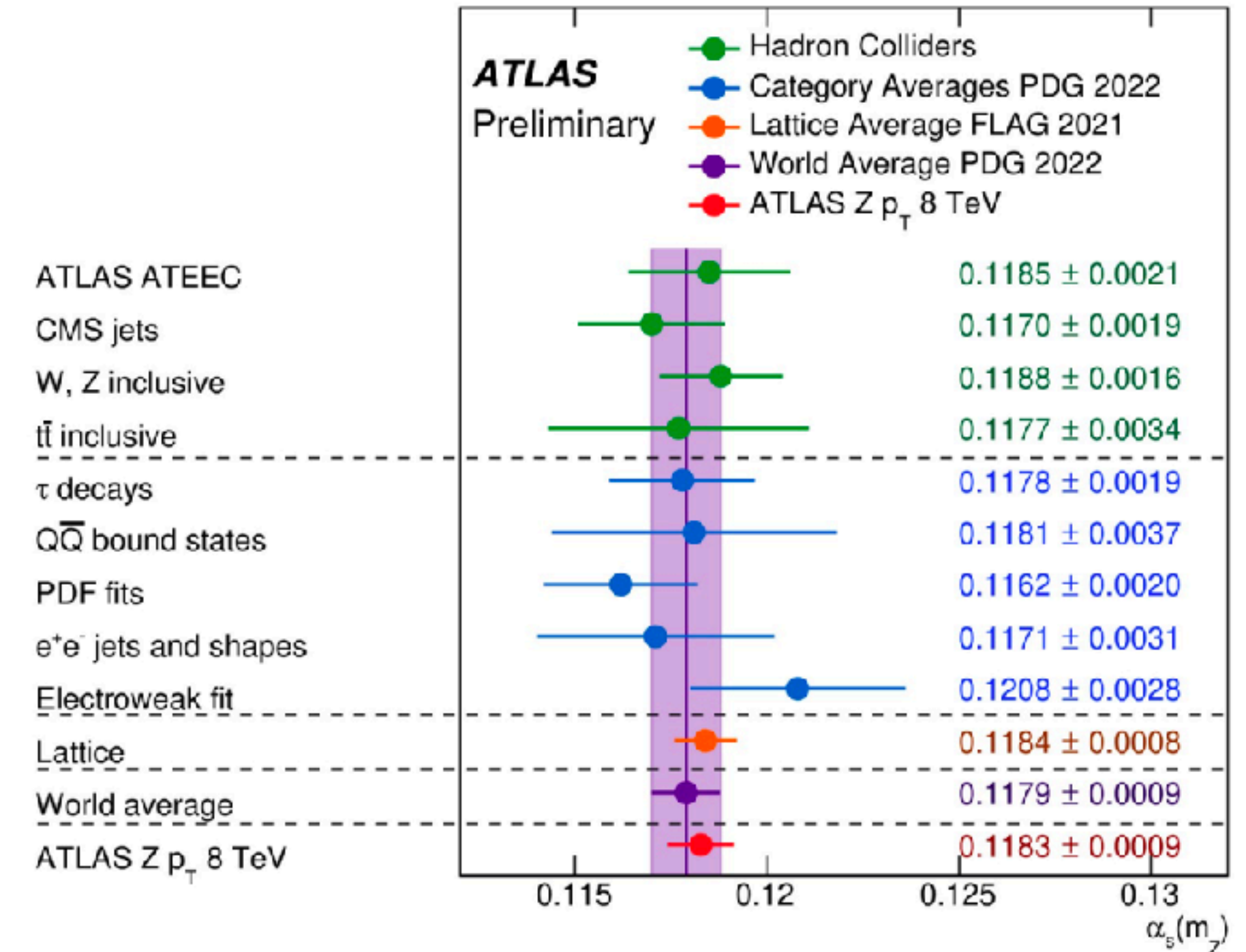
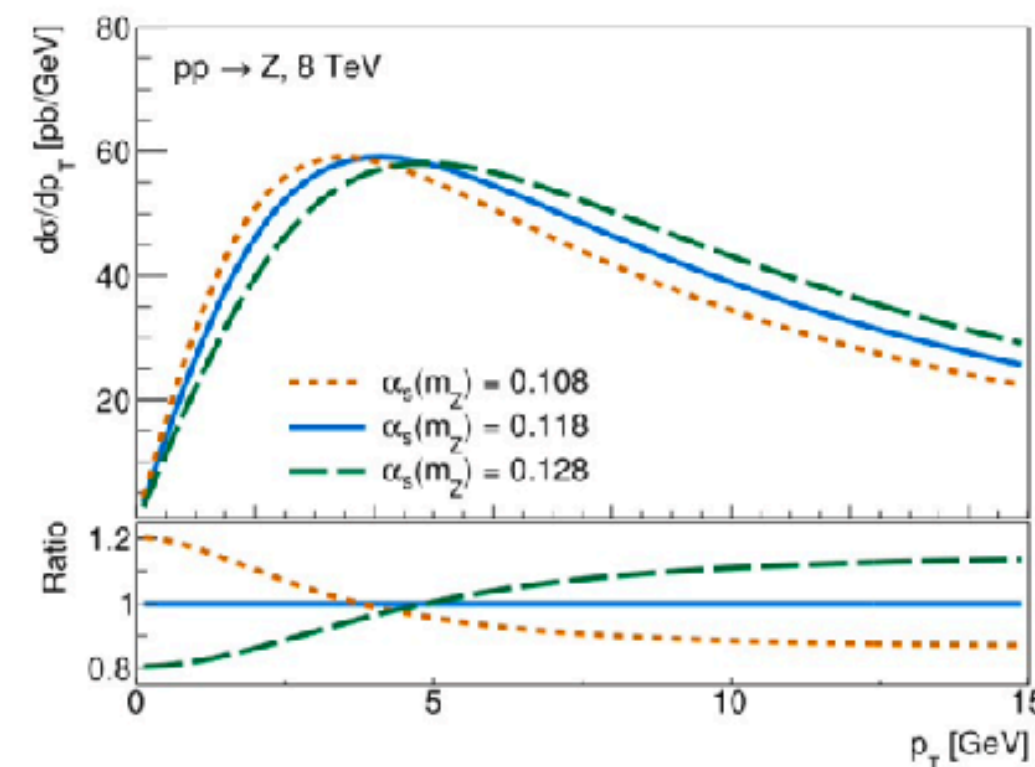
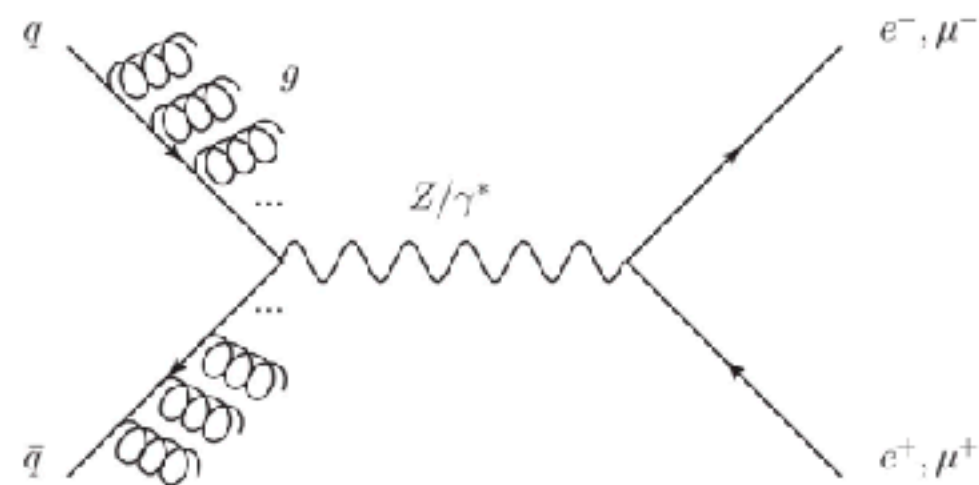
## Conditions for a competitive measurement of $\alpha_s$

- 1) is proportional to the strong coupling constant;
- 2) can be predicted theoretically with a percent precision (NNLO and higher);
- 3) is independent (nearly independent) of poorly-known parton distribution functions;
- 4) refers to low(er) region of hard momentum region;
- 5) does not suffer from unknown non-perturbative effects.

**“Inclusive Z transverse momentum seems to fit the bill”!**

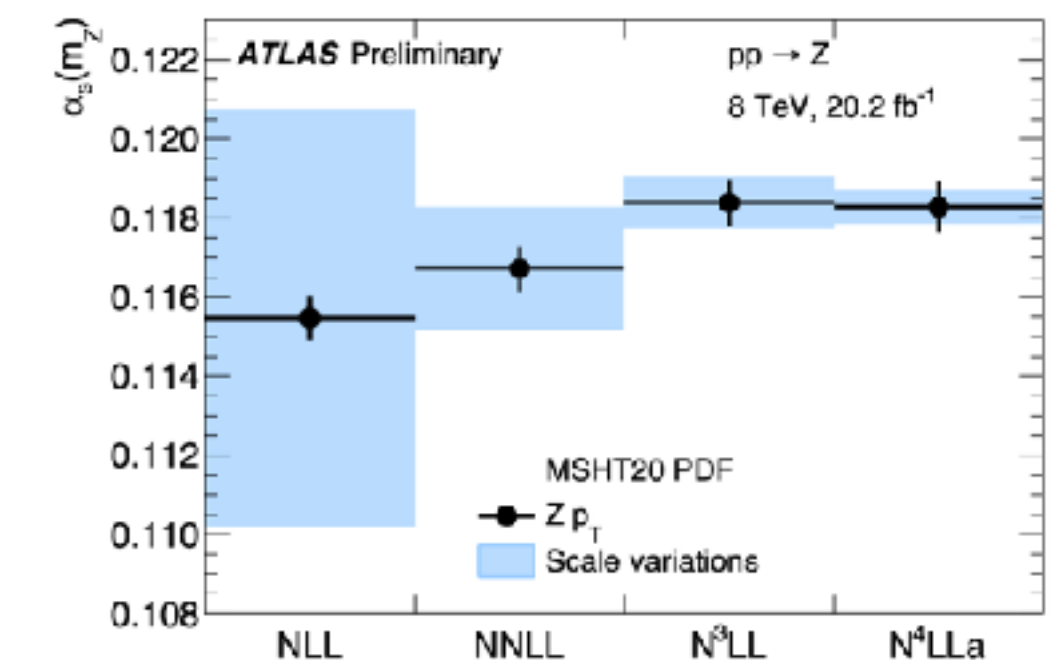
$$\frac{d\sigma_Z}{\sigma_Z dp_\perp} \sim \frac{\alpha_s(p_\perp)}{2\pi p_\perp} \ln \frac{M_Z}{p_\perp}$$

Using **Sudakov peak** in  $p_T$ , based on **resummed calculations**



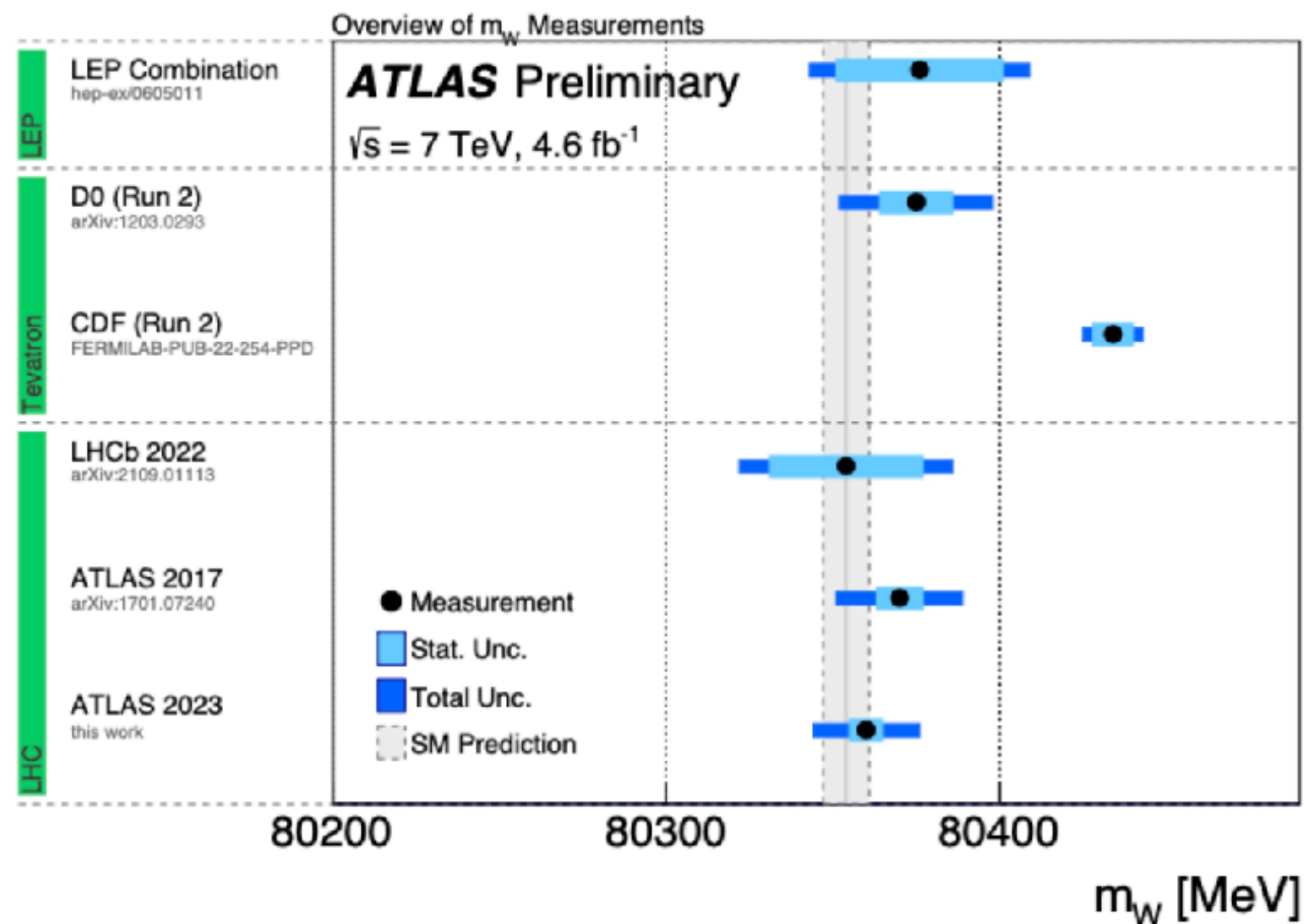
**Precision on par with lattice QCD and world average!**

**Such precision would not be possible without precise TH predictions!**



# W Mass Puzzle

## Measurements at LHC from ATLAS and LHCb



The measurement relies on the ratio of W/Z pT (as noted by Kirill) non trivial QCD and EW corrections can modify this spectrum and bring correction of up to ~20 MeV

Before discussing the tension of the CDF measurement with the SM, need to address the tension between measurements!

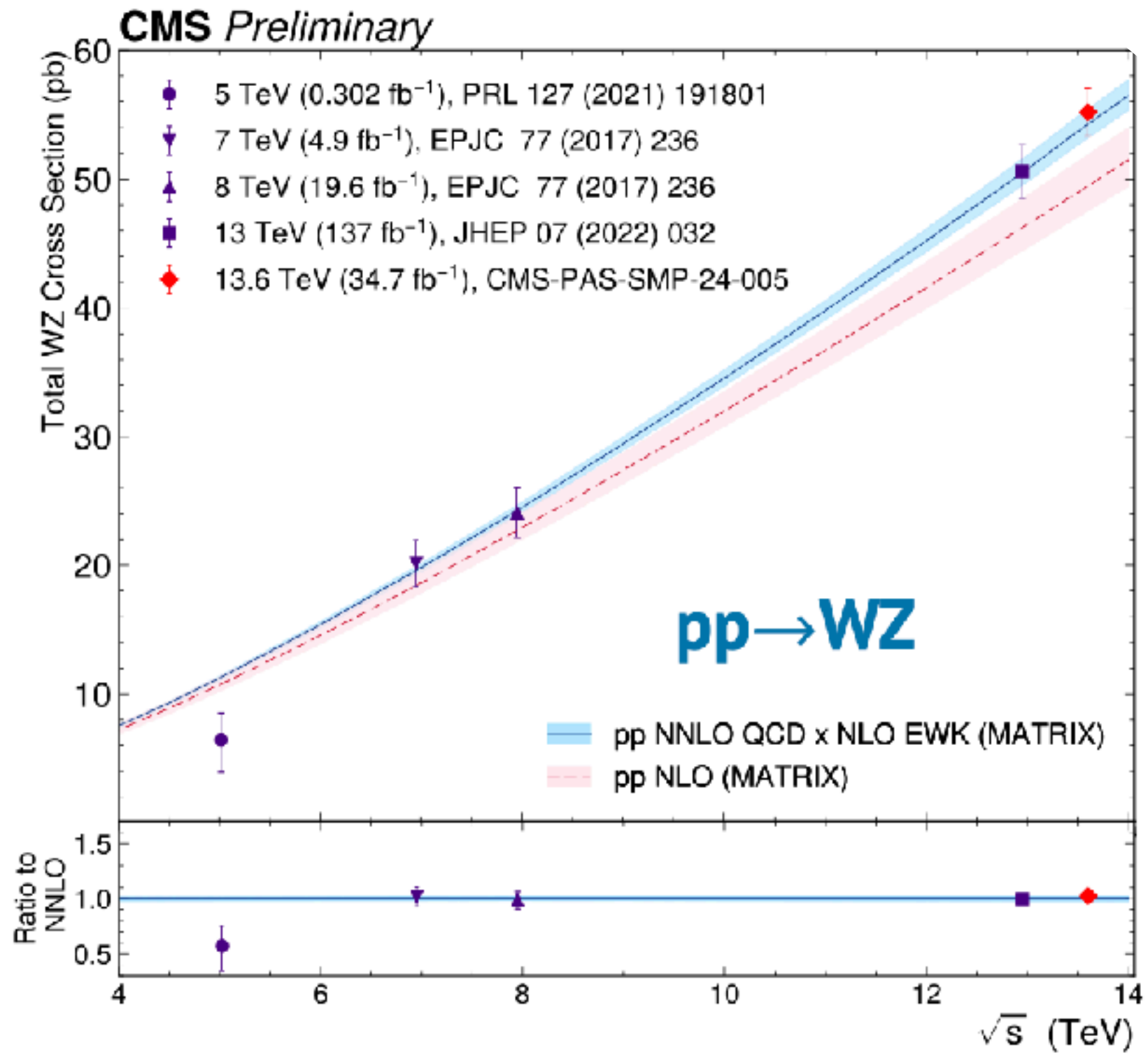
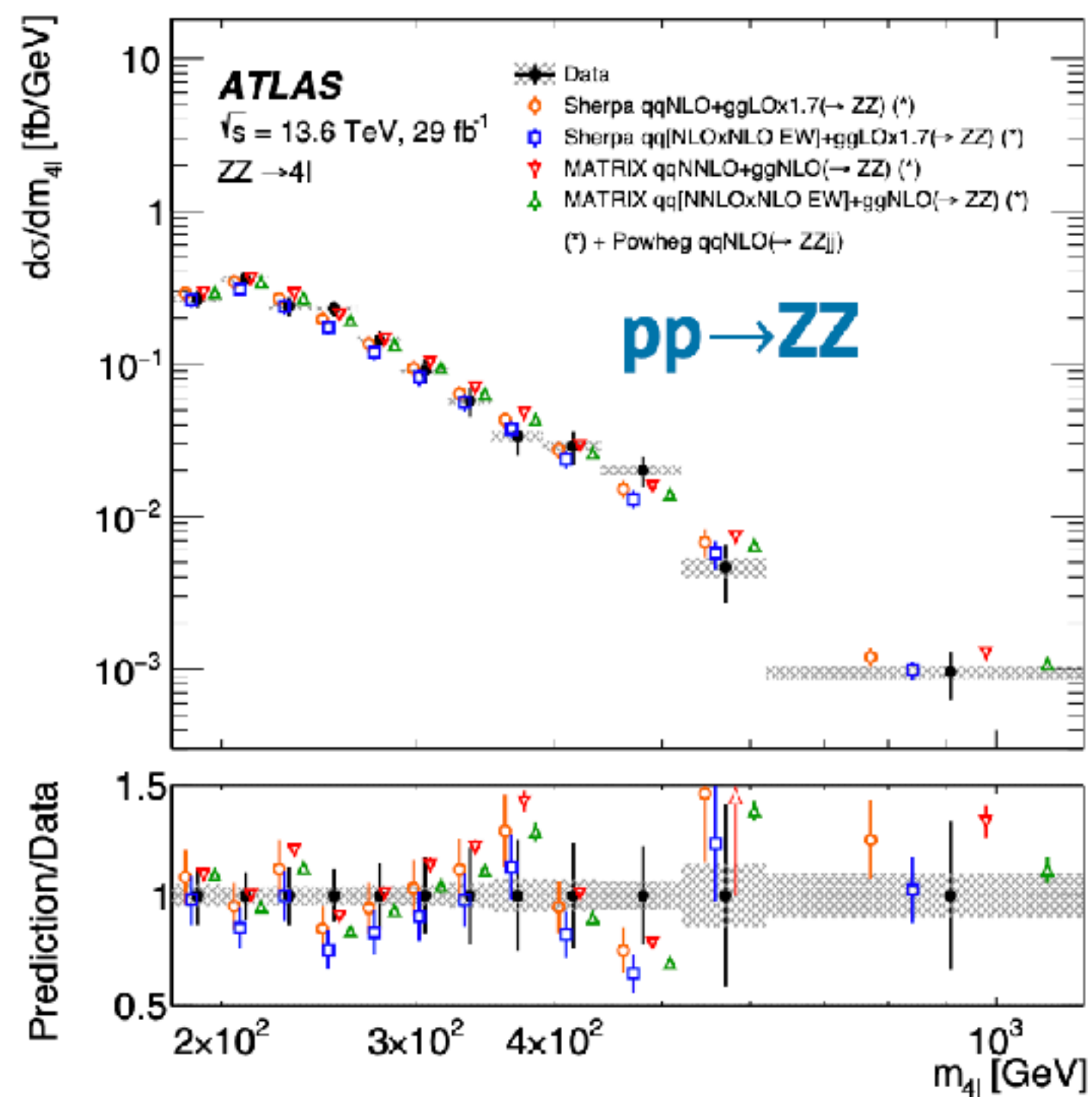
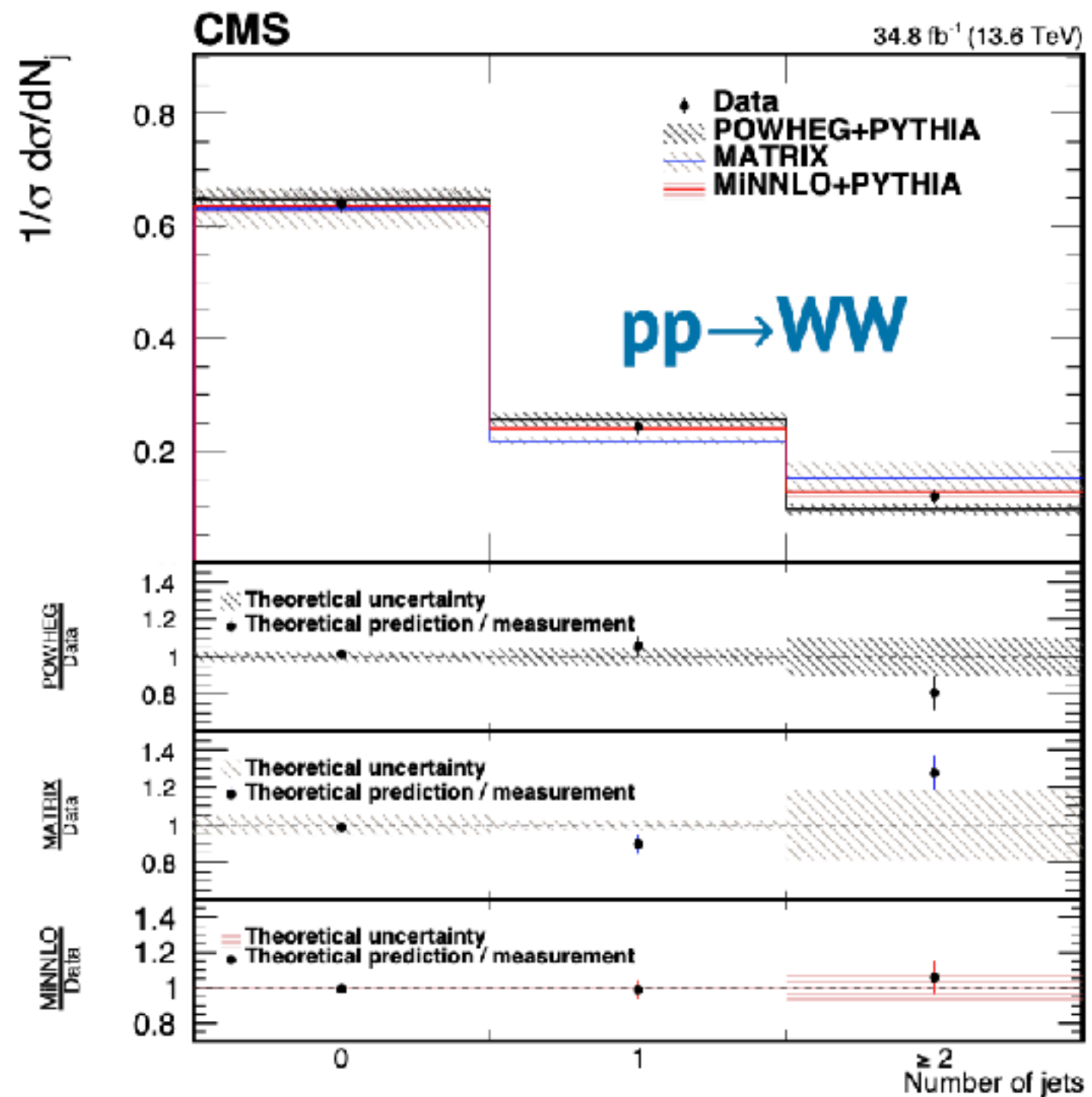
The tension between ATLAS and CDF W mass is of  $4\sigma$

Significant evidence of measurement systematic bias!

$$m_W = 80360 \pm 5_{(\text{stat.})} \pm 15_{(\text{syst.})} = 80360 \pm 16 \text{ MeV}$$

# First Measurements of di-boson production at Run 3

Monica Dunford, Philip Sommer



First measurements of the diboson processes at Run 3 (with 2022 data)

# The Quark Sector - Top Entanglement

Maurizio Pierini, Didar Dobur

In top pair production at the LHC, top quarks are **not produced polarised**, however a **spin correlation** exists.



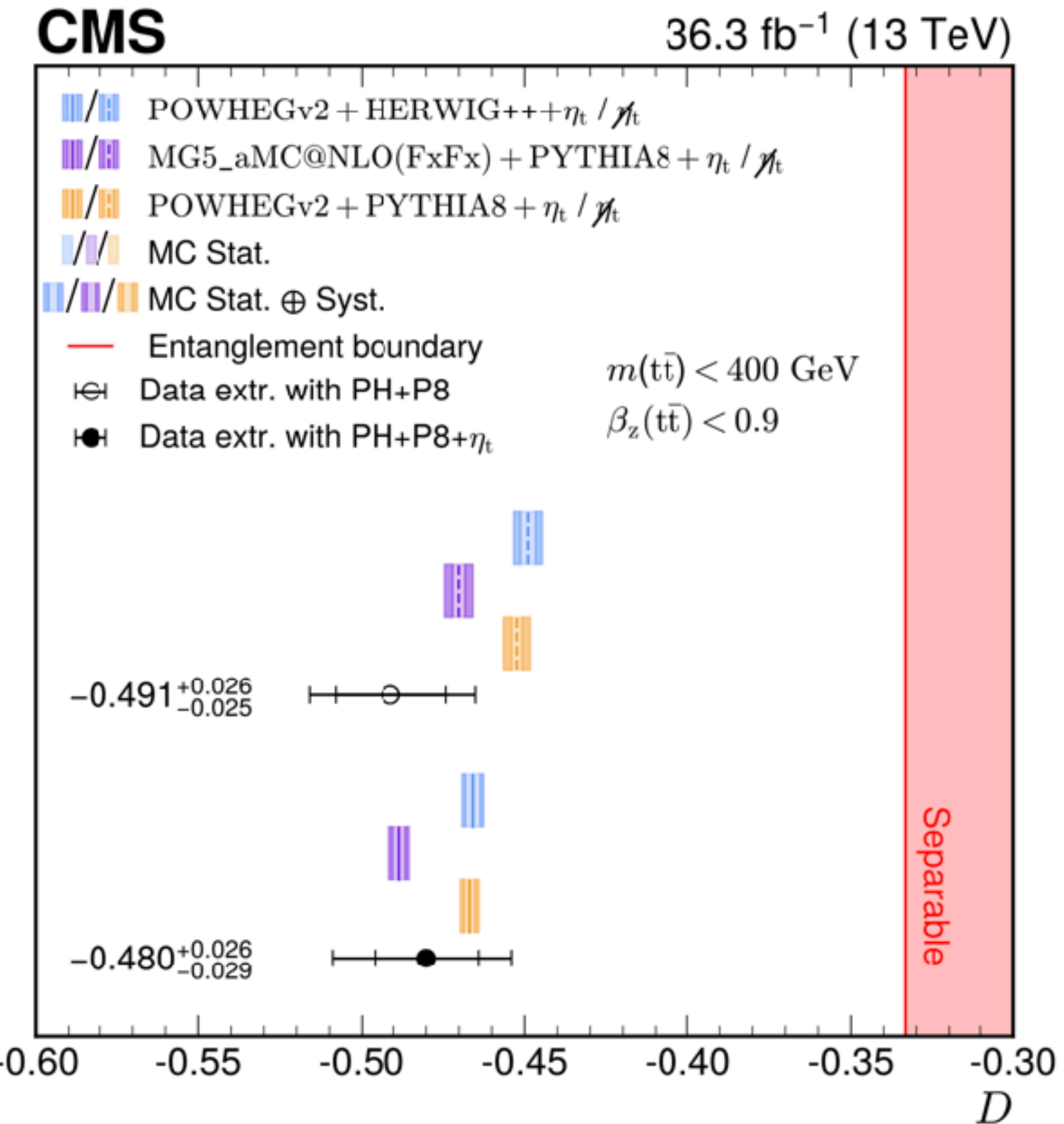
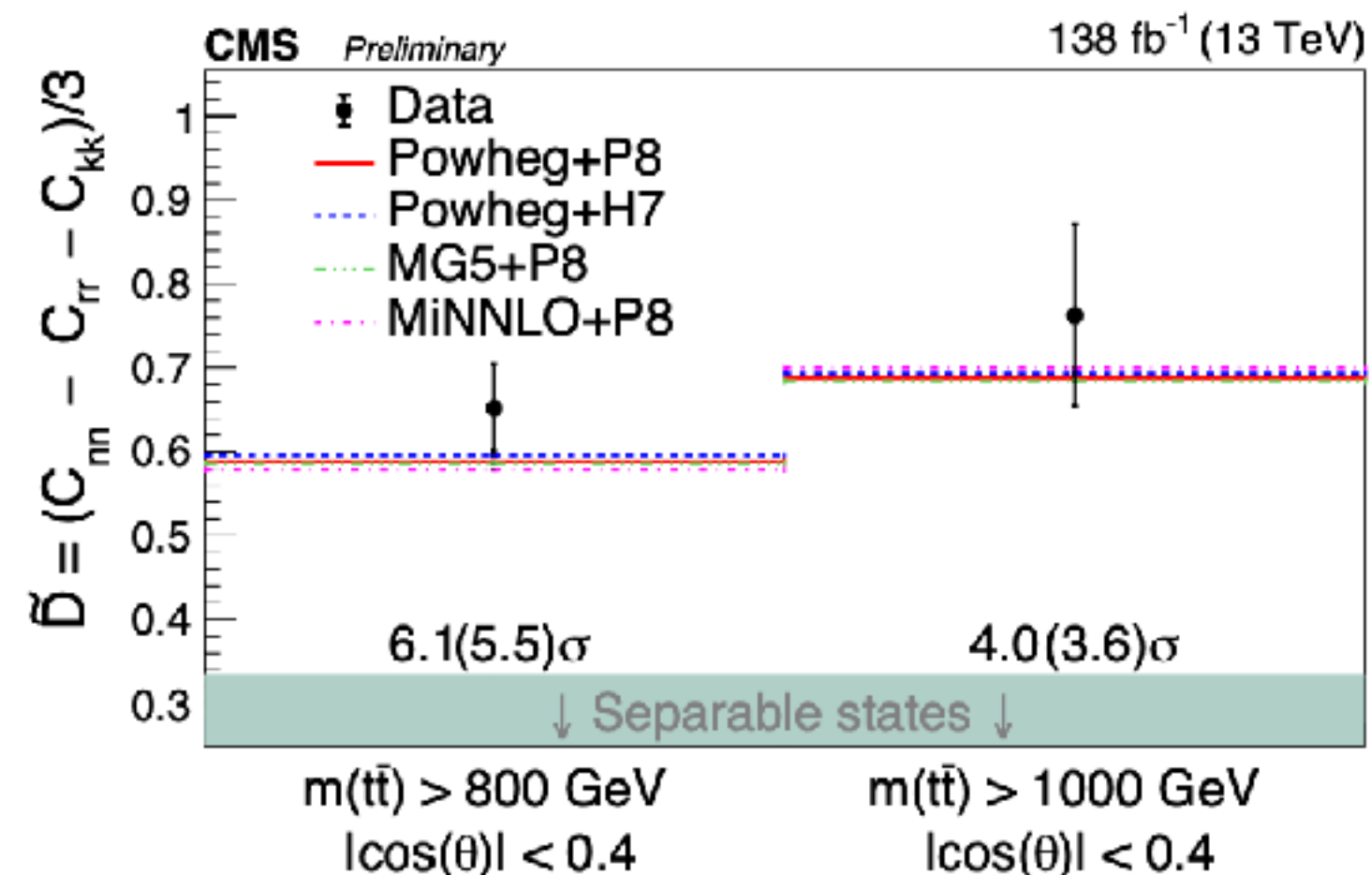
At threshold the  $gg \rightarrow t\bar{t}$  production is dominated by the “singlet” spin configuration, **which is a pure, superposed and maximally entangled Bell state**:

$$\frac{1}{\sqrt{2}} (|\uparrow\downarrow\rangle - |\downarrow\uparrow\rangle)$$

From the measurement of the spin density matrix we can probe whether this correlation is of quantum nature or not!

Initially measured near threshold where it is easier! CMS went beyond with:

- At production threshold in  $t\bar{t} \rightarrow b\ell\nu b\ell\nu$  events
- At high  $m_{t\bar{t}}$  with  $t\bar{t} \rightarrow b\ell\nu bq\bar{q}$  events, (phase space dominated 90% by space-like events)

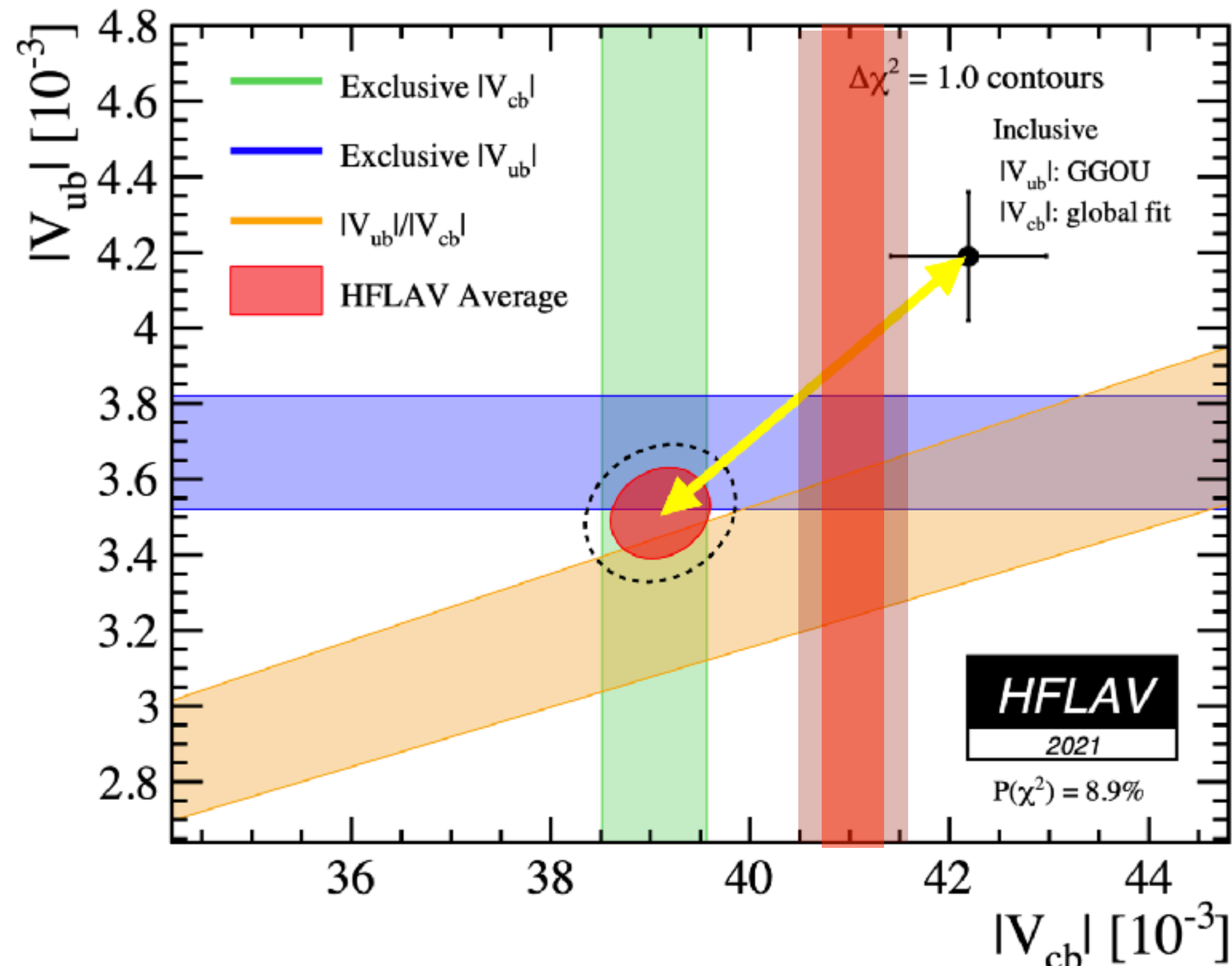


Very important elements (space-like) to go beyond entanglement towards the **violation of Bell Inequalities!** (With higher sensitivity)

# The $V_{ub}$ and $V_{cb}$ Puzzle

## New measurements from Belle II and Babar

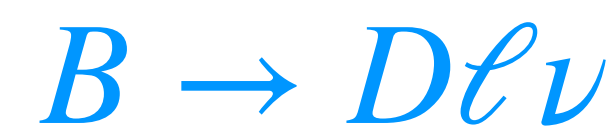
$|V_{cb}|$  and  $|V_{ub}|$  discrepancy  $\sim 3\sigma$  between exclusive and inclusive (have different TH uncertainties)!



Limiting factor in precision flavour physics!

Inclusive extraction from  $B \rightarrow X_c \ell \nu$ , where  $X_c$  goes to anything ([Bernlochner et al.](#) and [Bordone et al.](#)).

New exclusive  $|V_{cb}|$  results from BaBar and Belle II using fully differential information for the first time!



$$|V_{cb}| = (41.1 \pm 1.2) \times 10^{-3}$$



$$|V_{cb}| = (41.0 \pm 0.7) \times 10^{-3}$$

“Perhaps on the right path to resolve this puzzle”

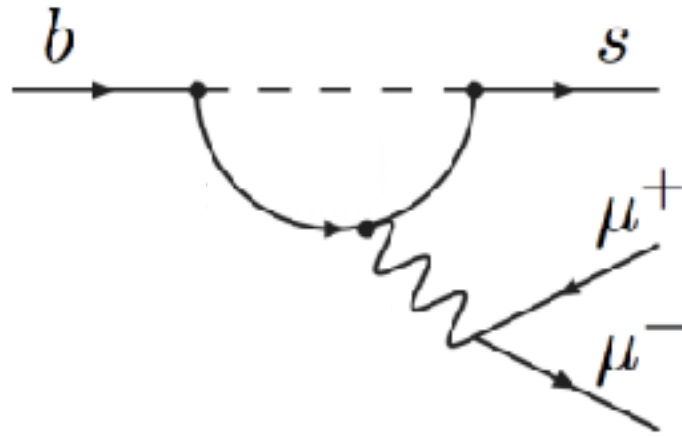
# Lepton Flavour Universality

Jim Libby, Yasmine Amhis, Maurizio Pierini

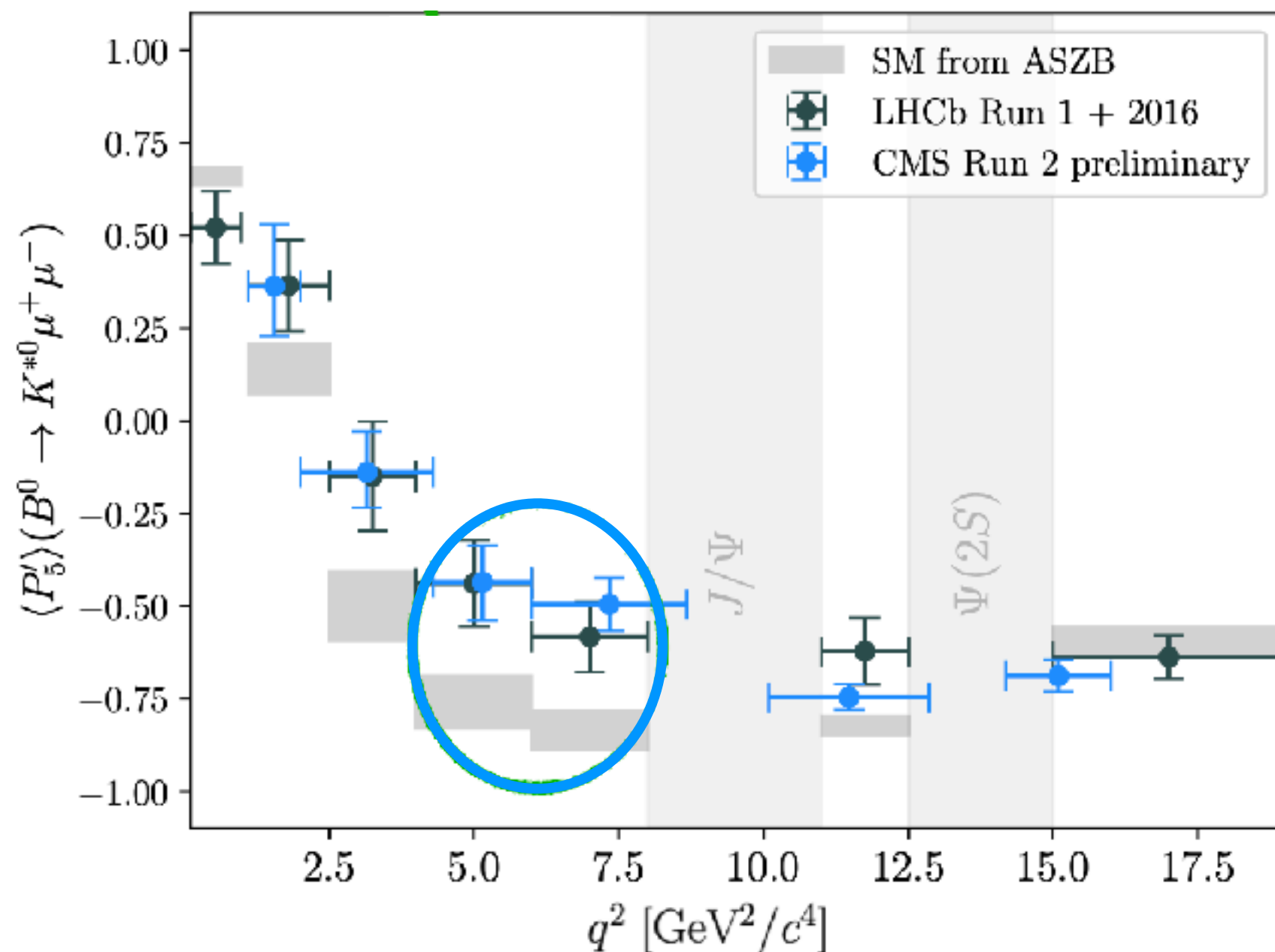
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## RK\* from LHCb cancelled end of 2022

Still a 2-3 Standard Deviation in the angular distribution and absolute branching fractions of  $B^0 \rightarrow K^{0*} \mu^+ \mu^-$  ([paper](#))



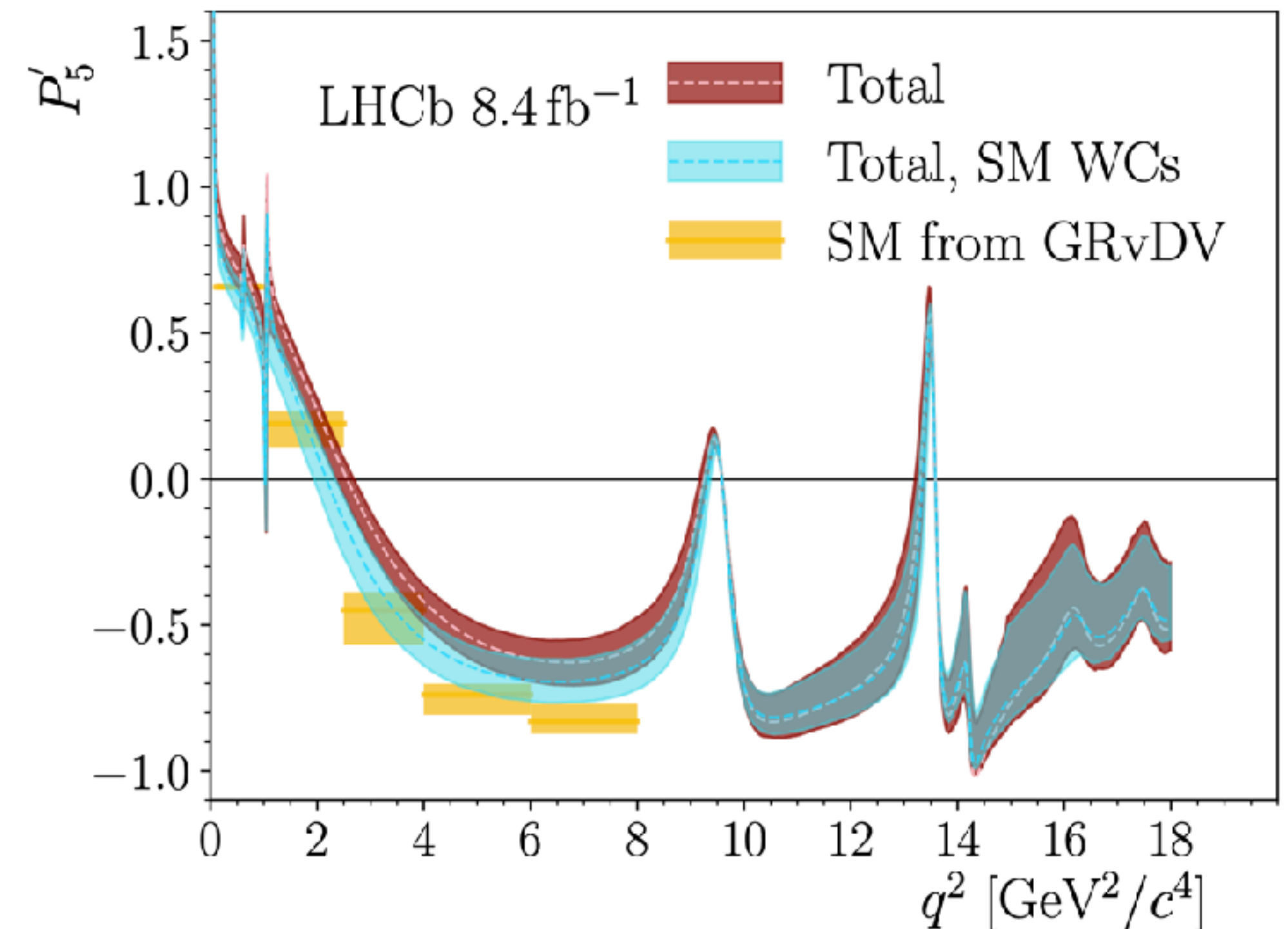
## $B^0 \rightarrow K^{0*} \mu^+ \mu^-$ also from CMS



$P'_5$  angular observable essentially free from form factor uncertainties.

## Comprehensive analysis of local and nonlocal amplitudes $B^0 \rightarrow K^{0*} \mu^+ \mu^-$ ([paper](#))

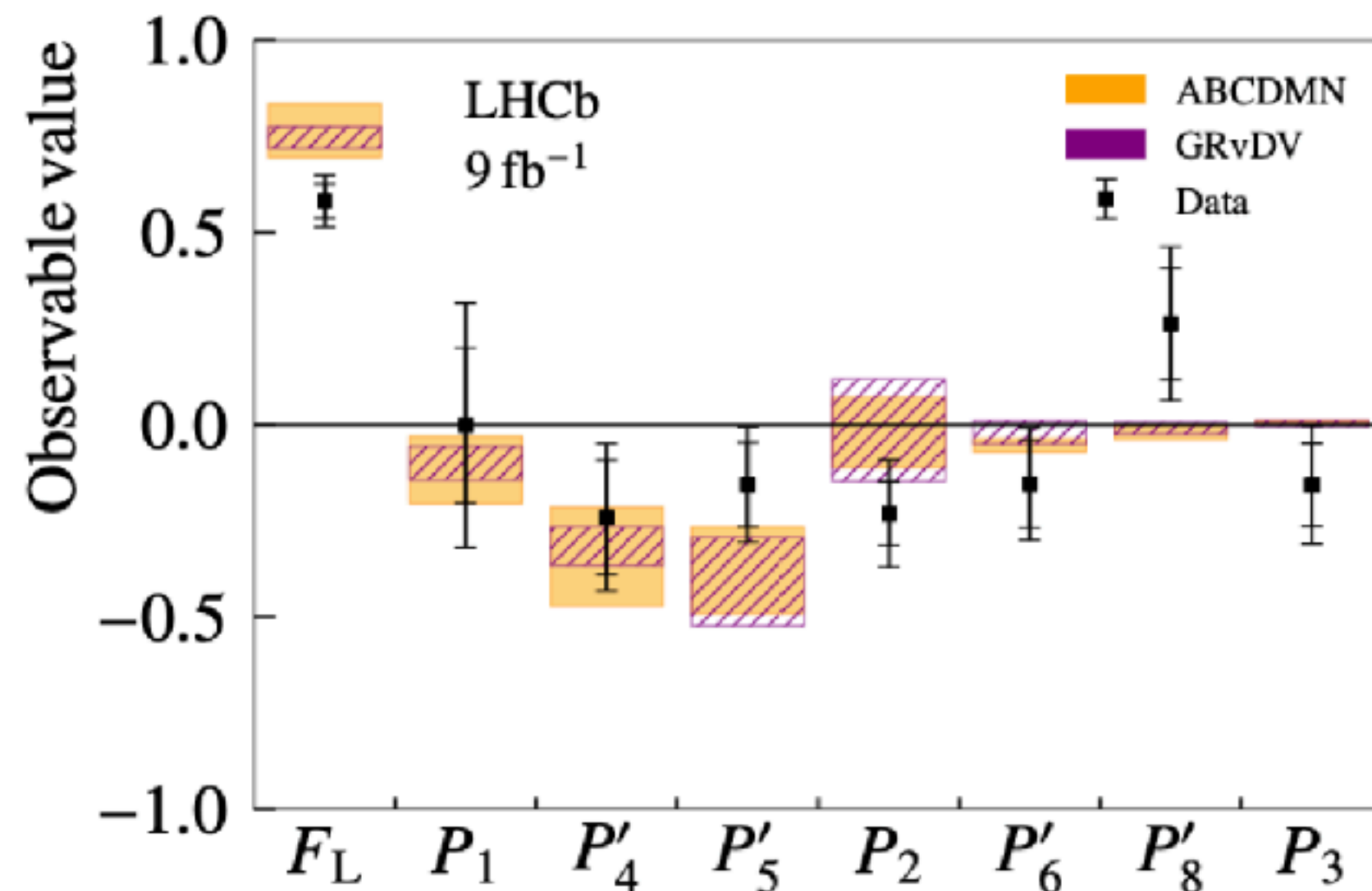
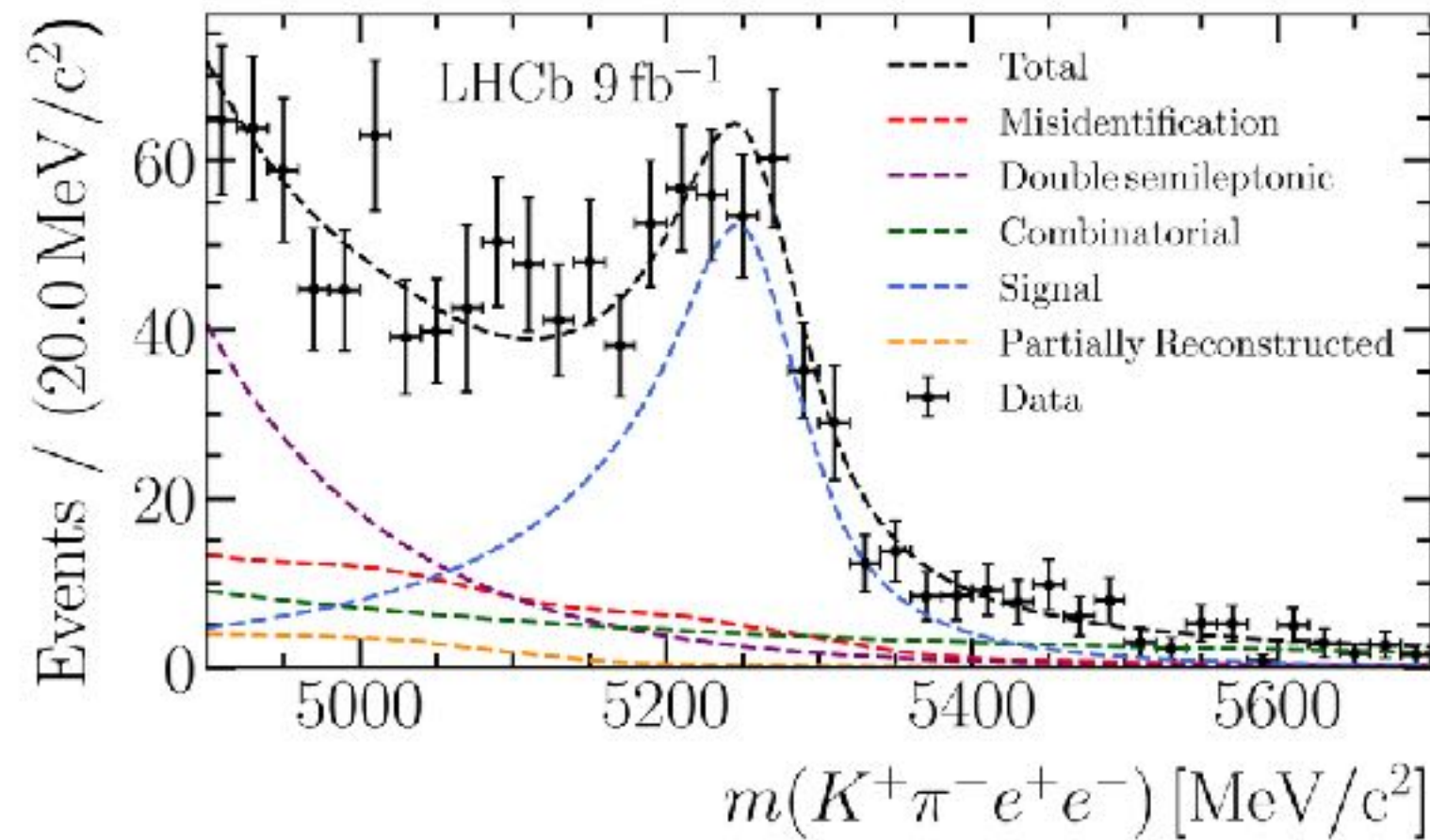
Careful long-distance contributions weaken these tensions are not considered



Tension now down to  $2.1\sigma$

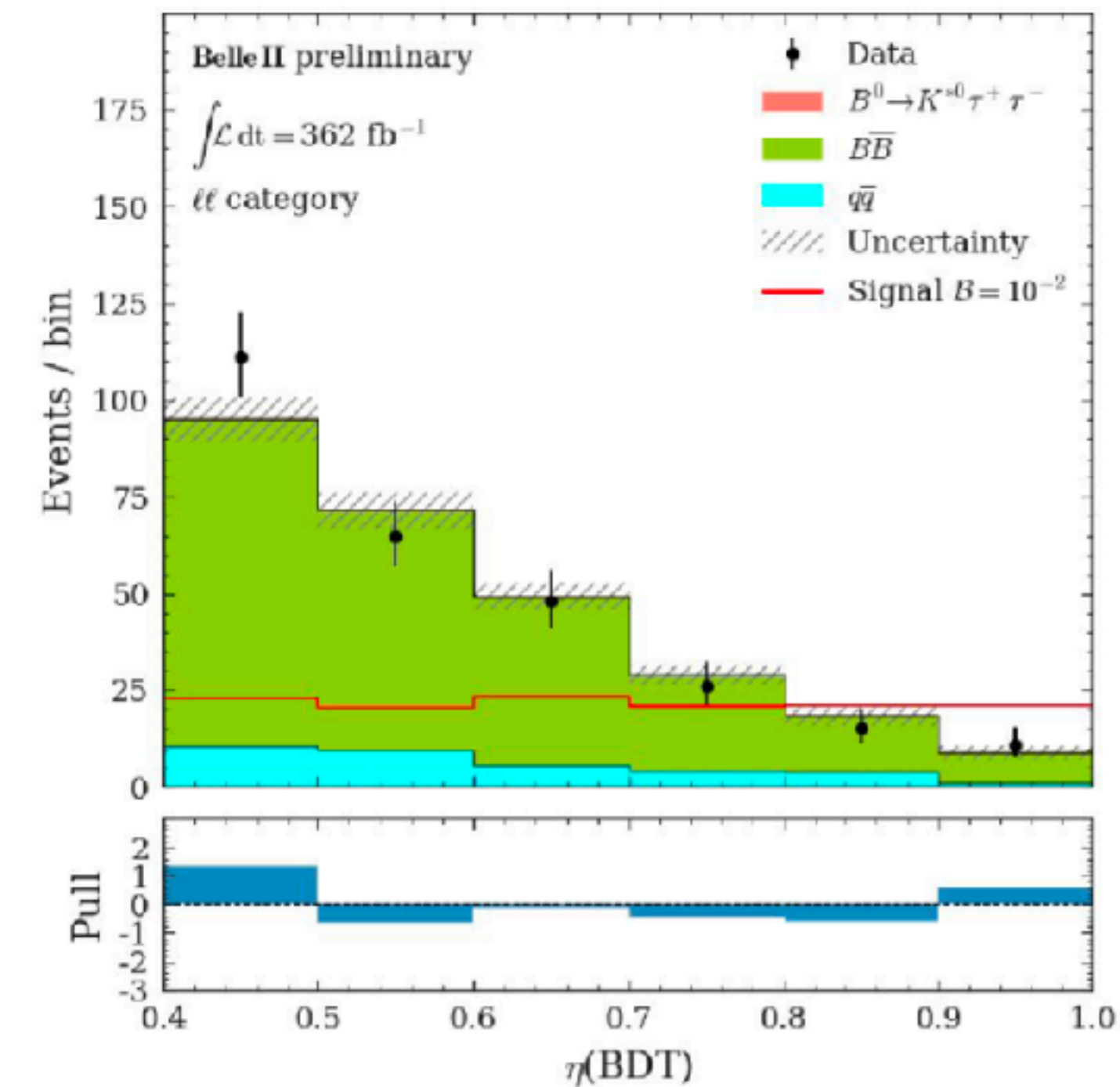
# Lepton Flavour Universality

## New $B^0 \rightarrow K^{0*} e^+ e^-$ measurement from LHCb



## New $B^0 \rightarrow K^{0*} \tau^+ \tau^-$ measurement from Belle II

Analysis particularly sensitive to new physics affecting the  $B^0 \rightarrow K^{0*} \mu^+ \mu^-$  decay!



$$\mathcal{B}(B^0 \rightarrow K^{*0} \tau^+ \tau^-) < 1.73 \times 10^{-3}$$

Limit twice improved over the Belle result!



# Lepton Flavour Universality

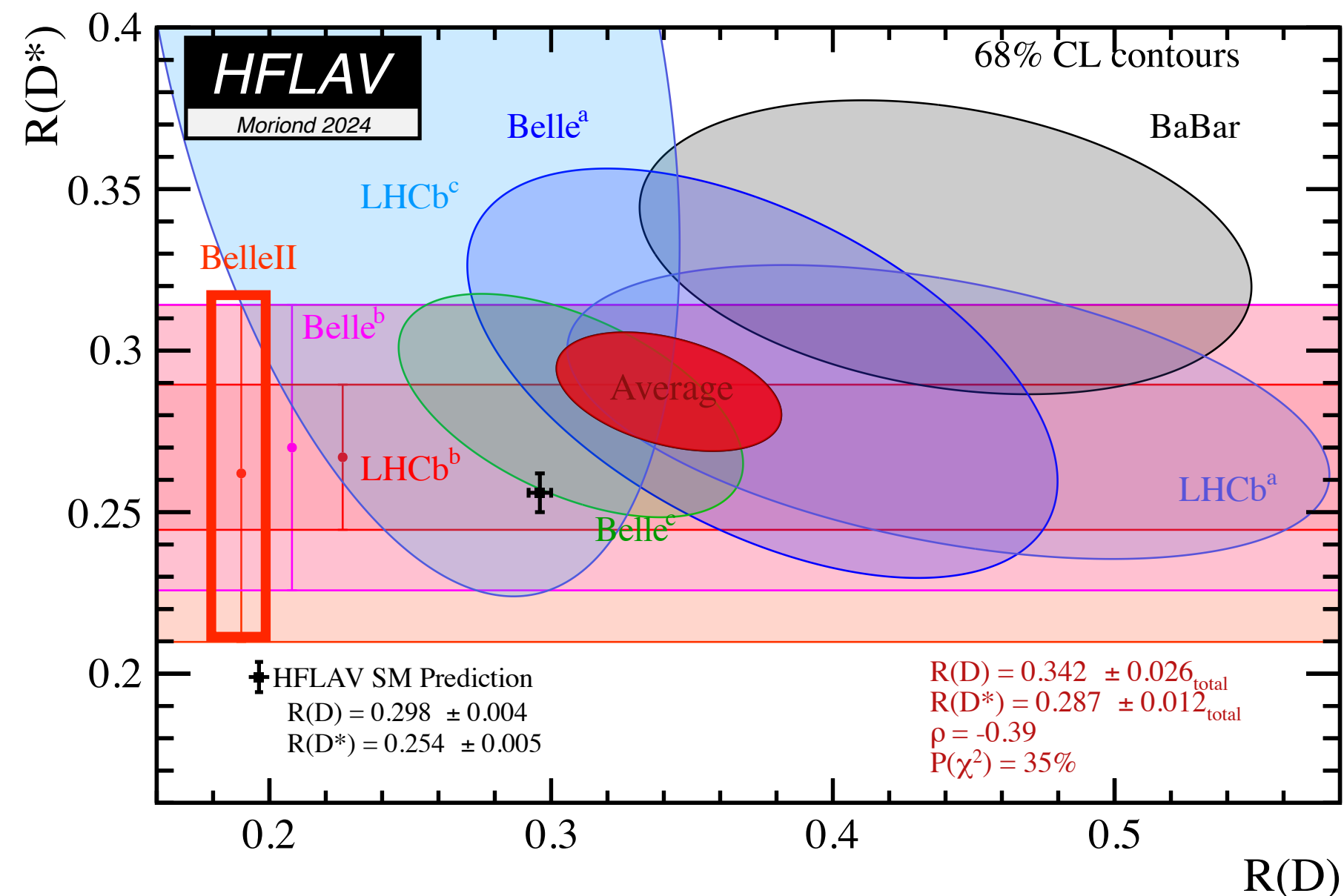
Jim Libby, Alakabha Datta, Monica Dunford

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## First Belle II $R_{D^*}$ measurement!

Both TH and EXP clean!

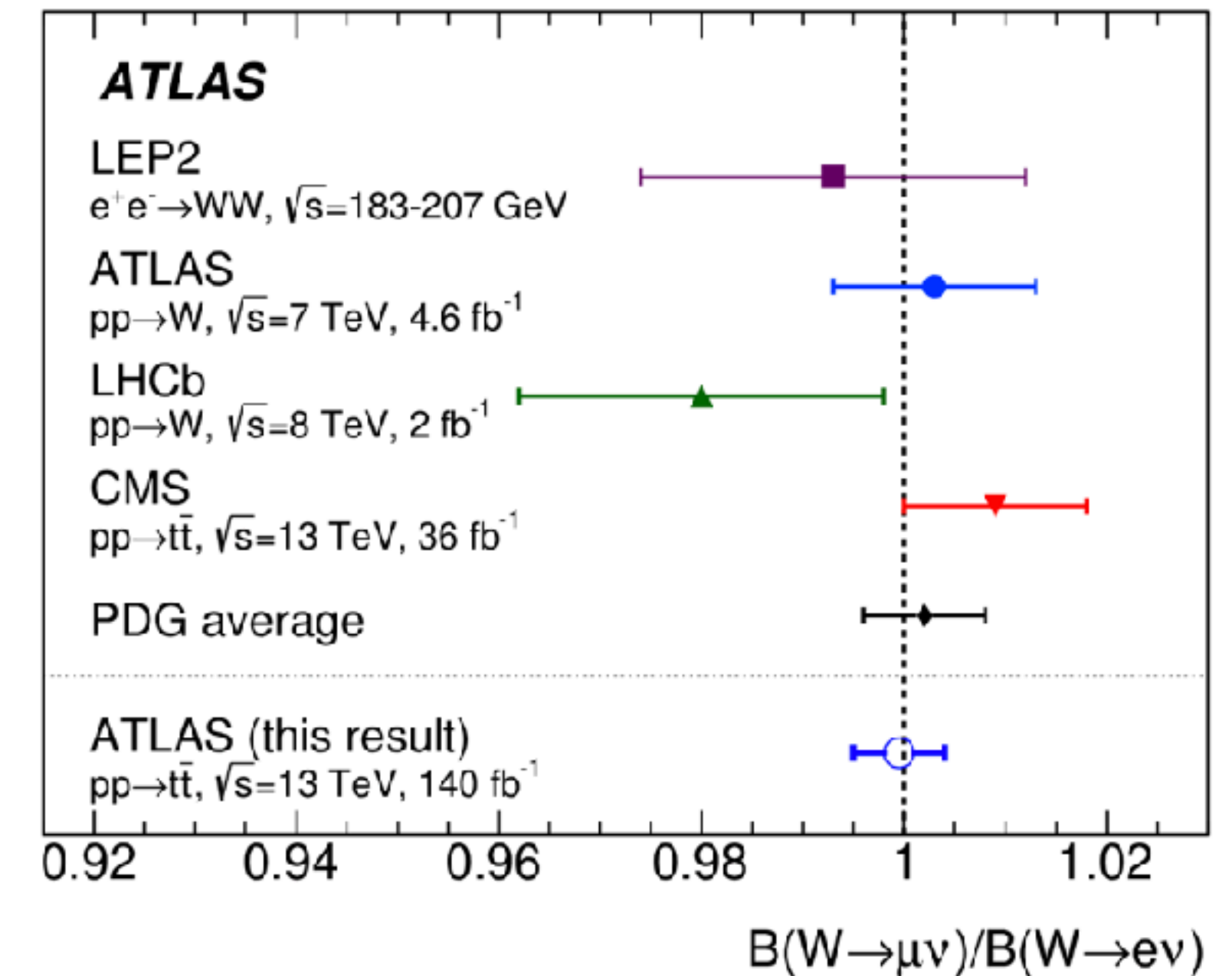
$$\mathcal{R}(D^{(*)}) = \frac{\mathcal{B}(B^0 \rightarrow D^{(*)-} \tau^+ \nu_\tau)}{\mathcal{B}(B^0 \rightarrow D^{(*)-} \mu^+ \nu_\mu)}$$



$$R_D^* = 0.26 \pm 0.04^{+0.04}_{-0.03}$$

- Systematic uncertainty related mainly to size of control samples
- Comparable precision to equivalent Belle result with 1/4 the sample

## Lepton universality measurements for “on-shell” W bosons in top decays



ATLAS result more precise than current world average

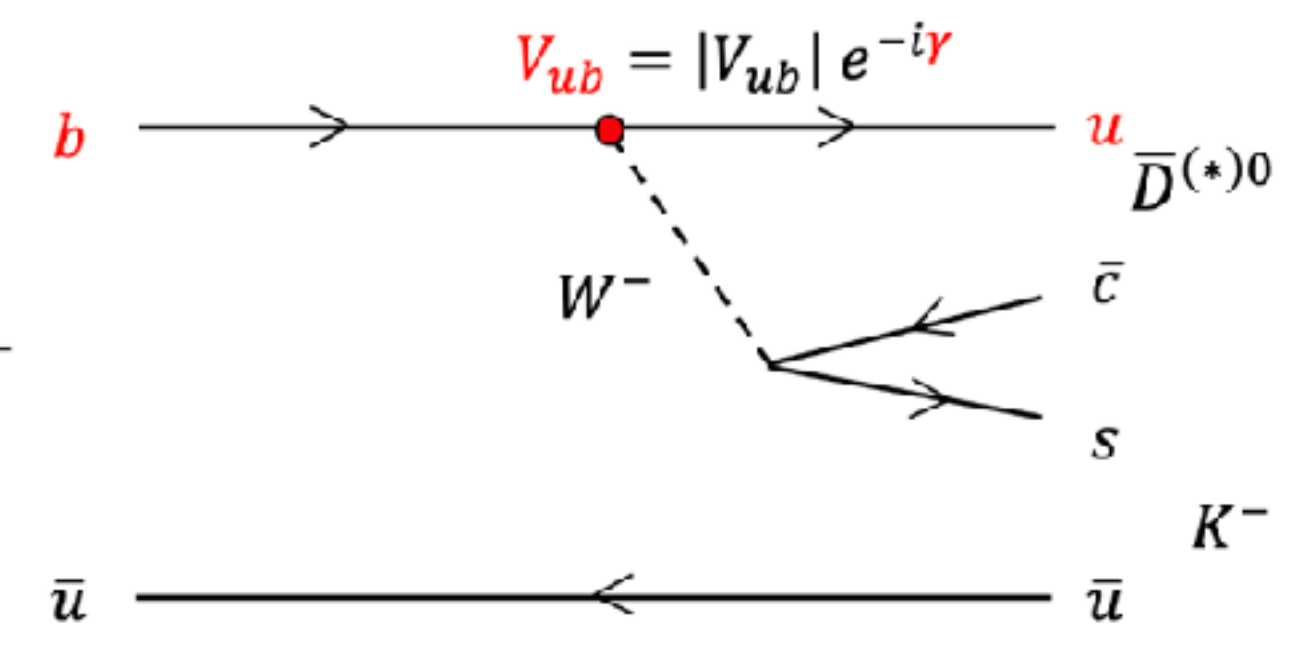
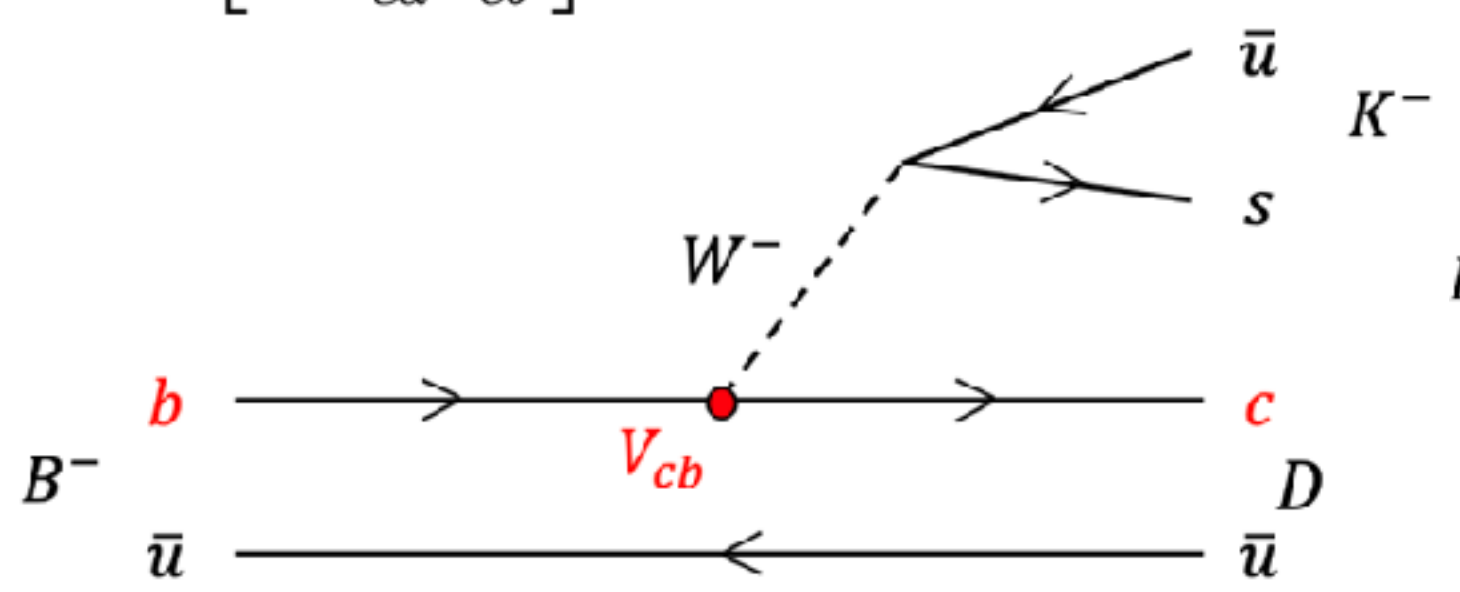
Also with W decays from tops CMS measures  $|V_{cs}|$  !

$$R_W^c = \frac{\mathcal{B}(W \rightarrow cq)}{\mathcal{B}(W \rightarrow ud)} \quad |V_{cs}| = 0.959 \pm 0.021$$

# Latest CKM $\gamma$ News (Belle II - LHCb)

Recent Measurements of  $\gamma$  in the golden channel  $B^\pm \rightarrow DK^\pm$

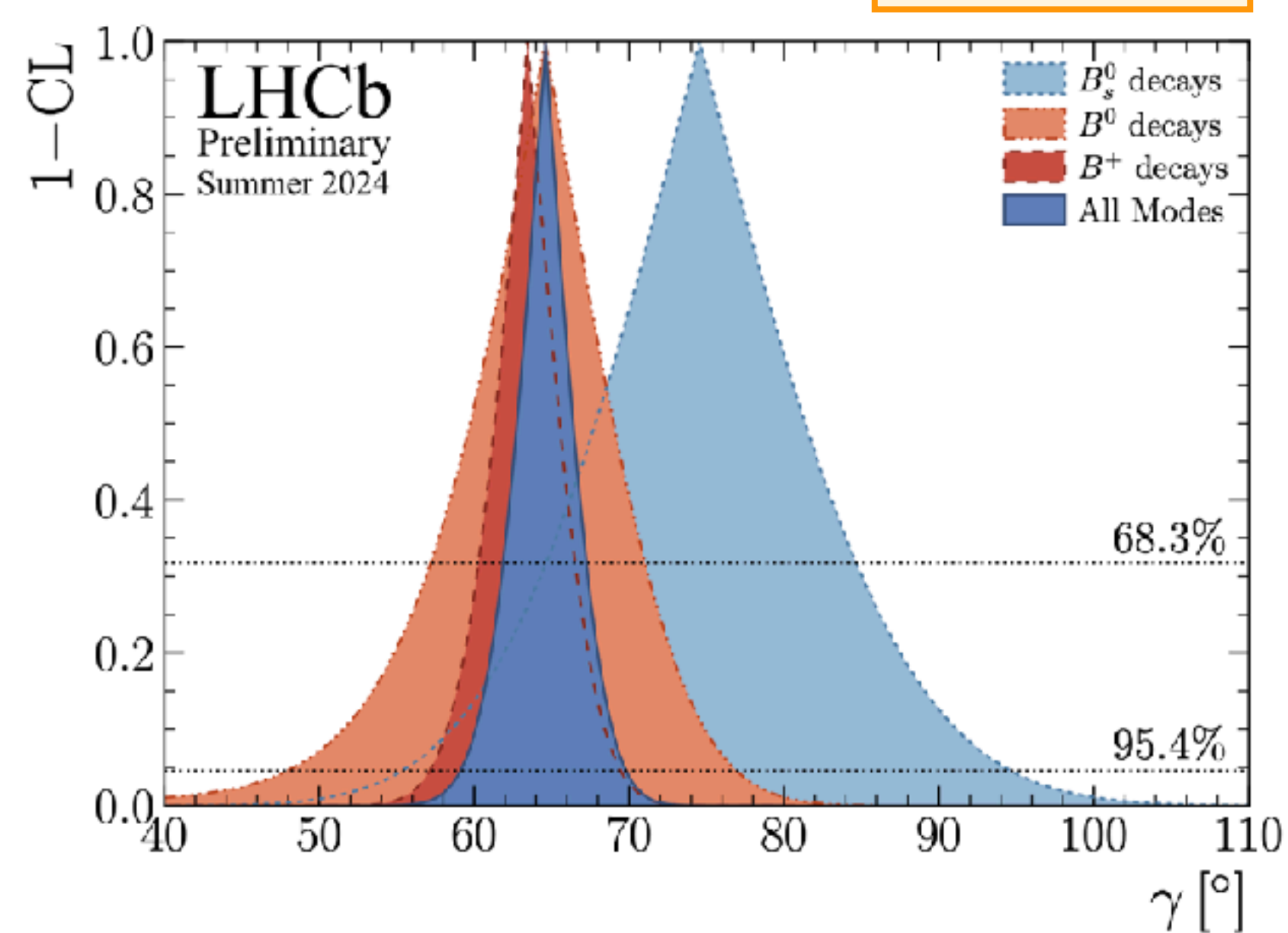
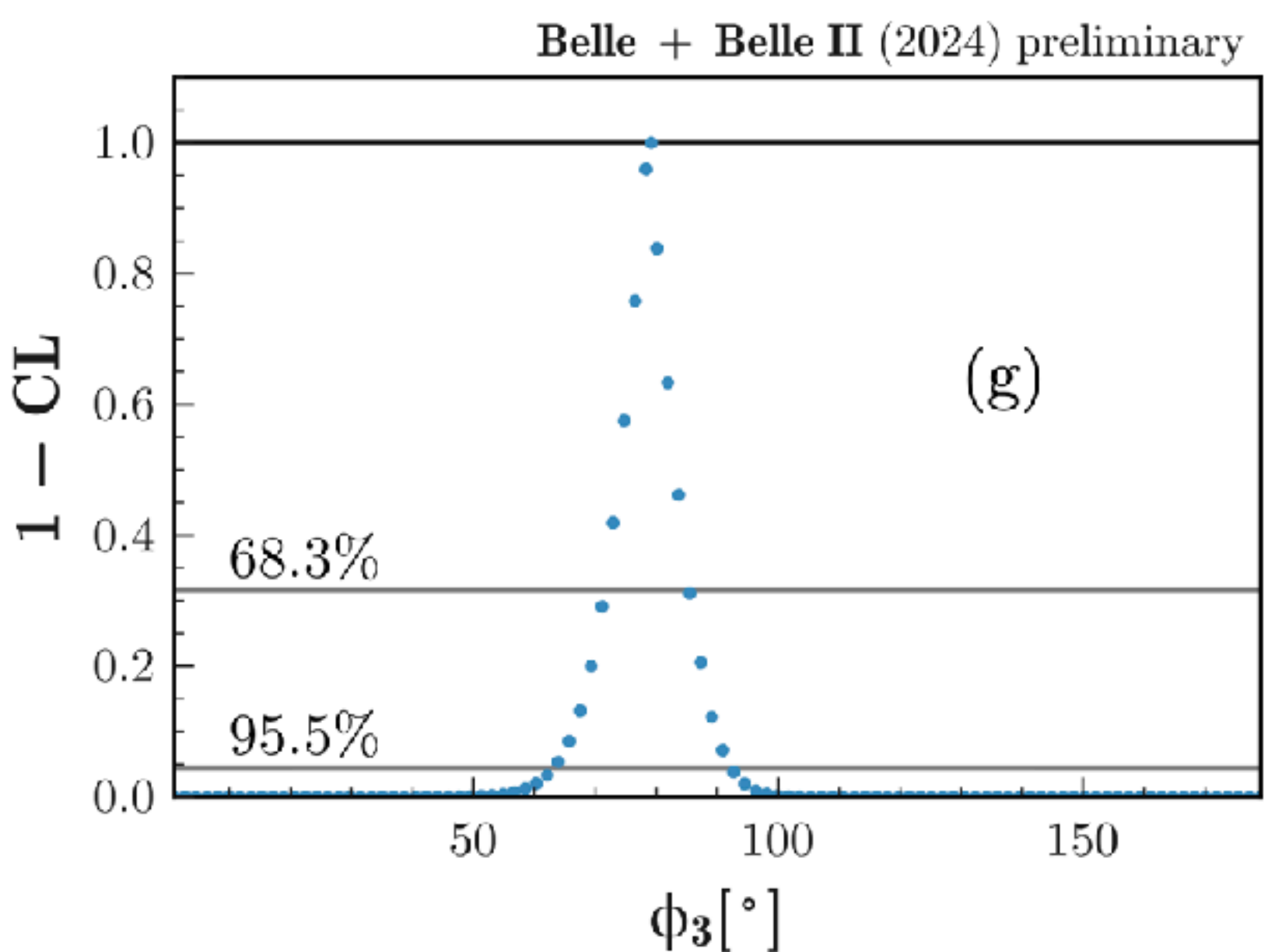
$$\gamma = \arg \left[ -\frac{V_{ud}V_{ub}^*}{V_{cd}V_{cb}^*} \right]$$



Lack of Lattice QCD needs makes it a **“pristine observable”** in flavour physics!

Charm input from BESIII/CLEO is critical

New for ICHEP 2024



Combination from Belle II

$$\gamma = (78.6^{+7.2}_{-7.3})^\circ$$

Combination from LHCb!

$$\gamma = (64.7 \pm 2.8)^\circ$$

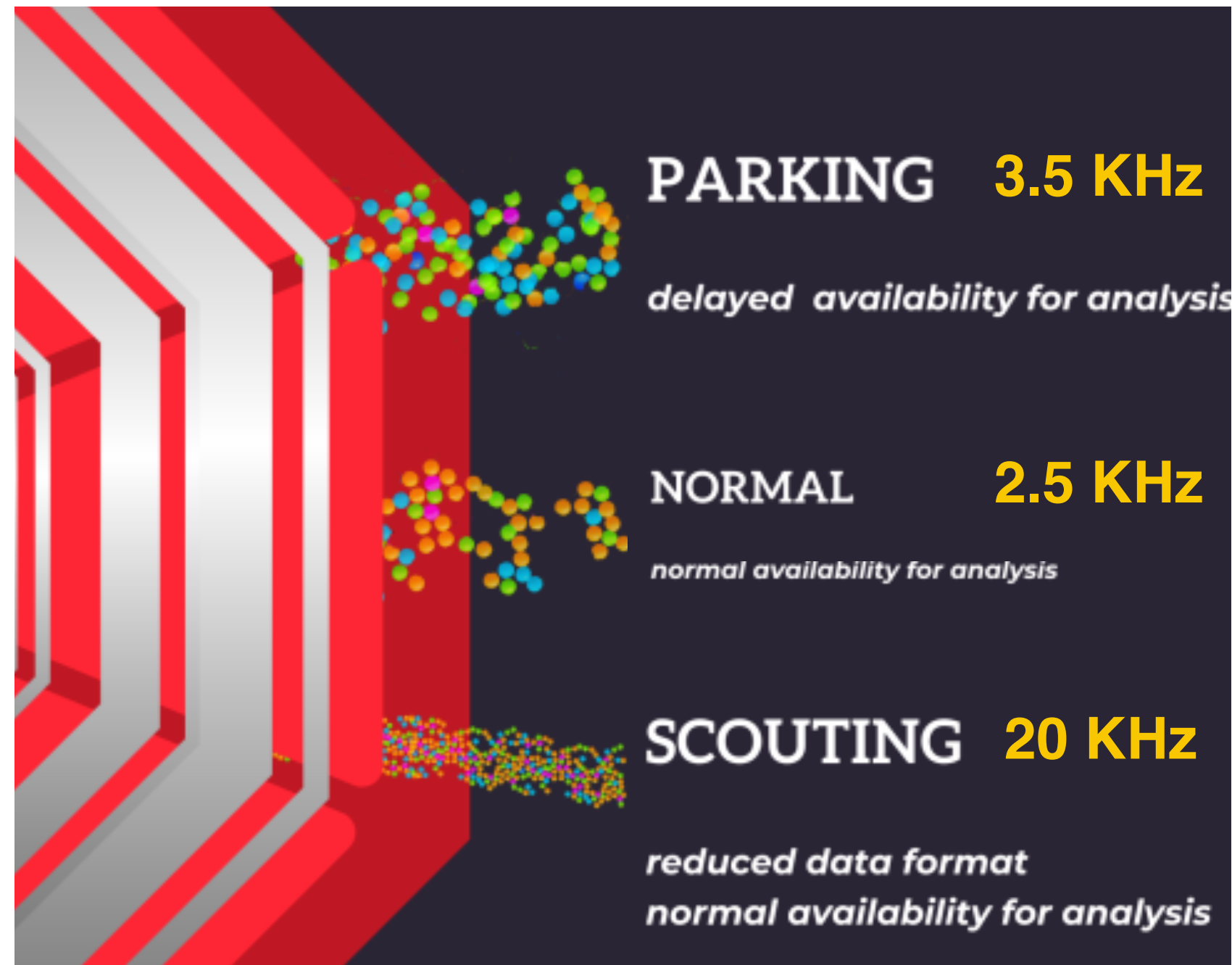
**Measurement from LHCb has surpassed the target goal for Run 2!!**

From CKM fitter  $\gamma = (66.3^{+0.7}_{-1.9})^\circ$

# Scouting, Parking and Trigger Improvements

Maurizio Pierini, David Rohr, Javier Duarte

[CMS-EXO-23-007](#)

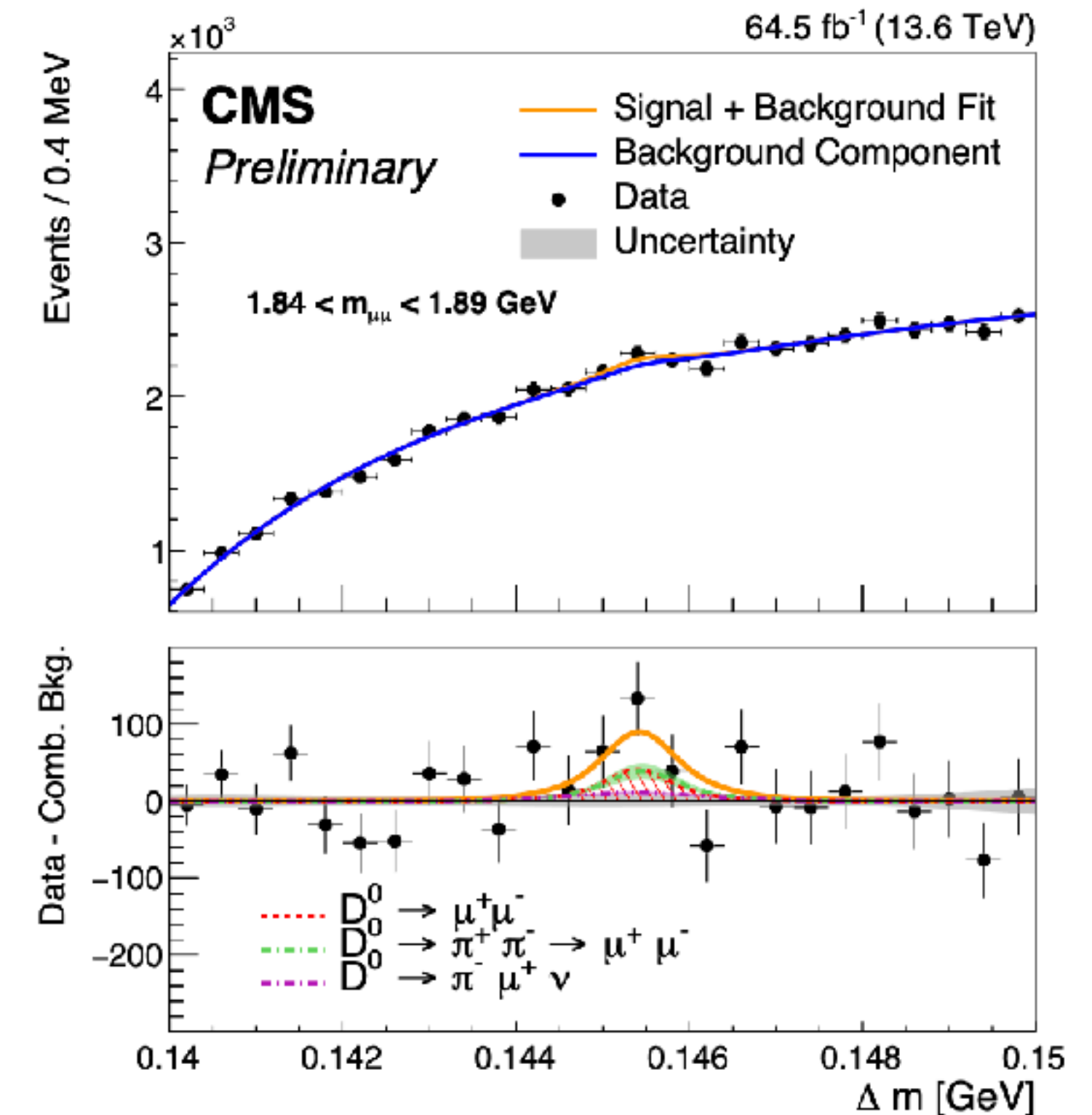


## First Results using Run3 Parking

Search for  $D^0 \rightarrow \mu^+ \mu^-$  - improved by 35% over previous best limit

Further improvements foreseen!

[CMS-BPH-23-008](#)



Large scale project to support the developments (with powerful hardware and software) to bring experiments to the next level with efficient data flows and structures, with ambitious and large ML models e.g. GNN tracking!

See [talk](#) by D. Rohr

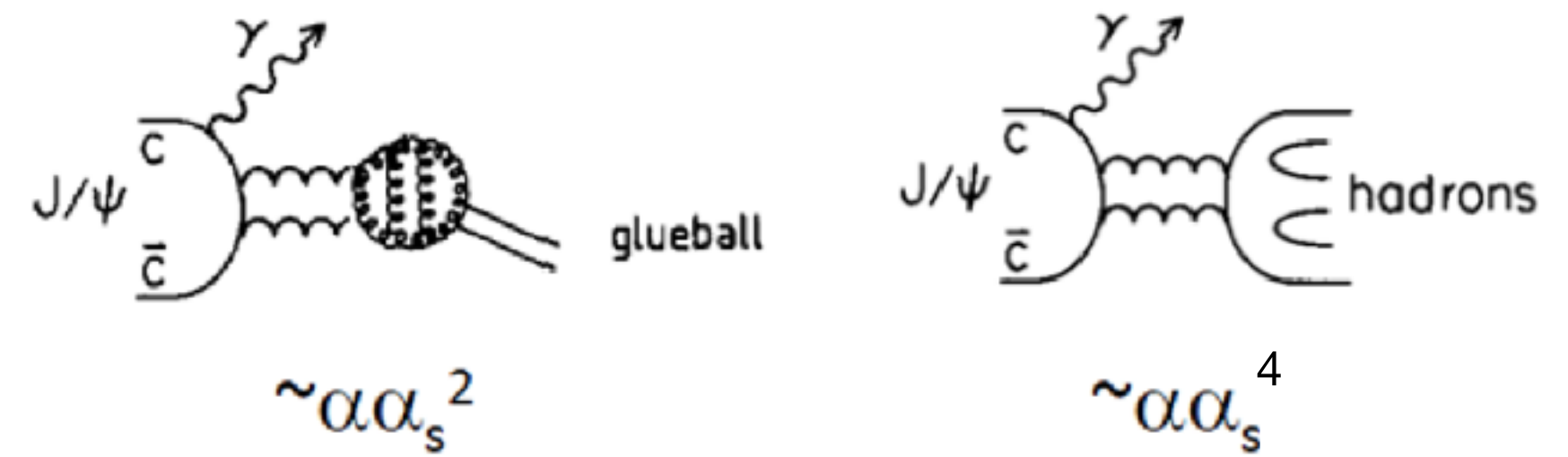
# Focus on New States

## Observation of $\chi_c(3872)$ radiative decays at LHCb!

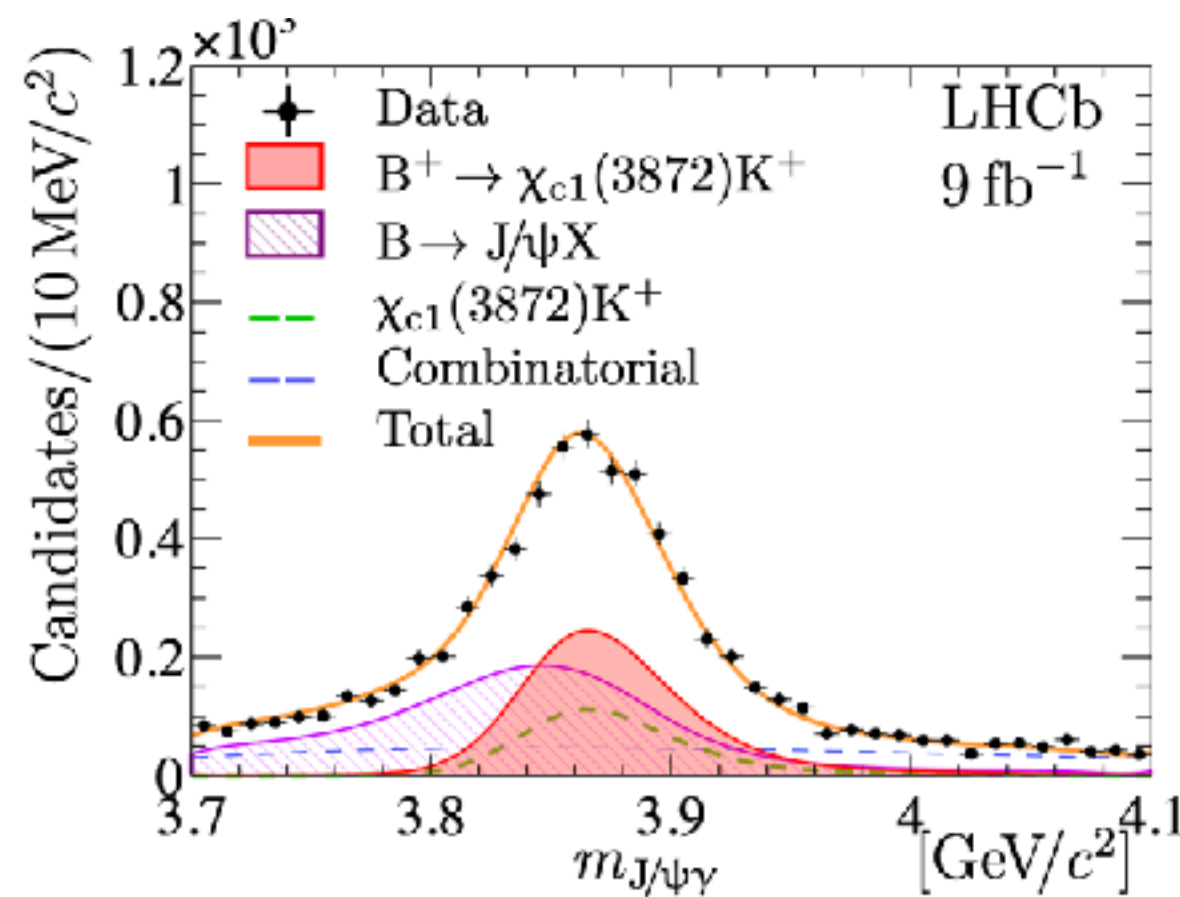
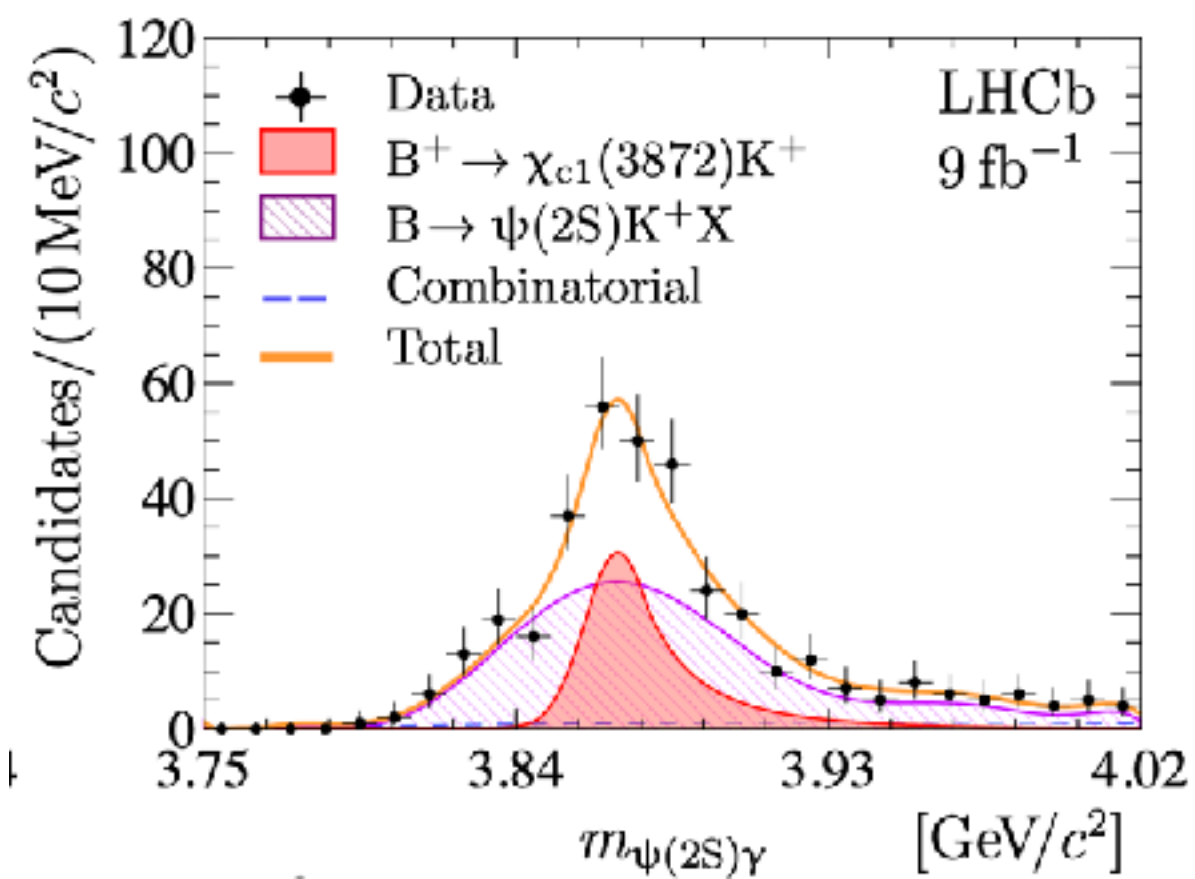
$\chi_c(3872)$  discovered 20 years ago, could it be D meson molecular state?

Now, the LHCb collaboration is closer to finding out what it is made up of!

## Discovery of a Glueball-like particle X(2370) at BESIII

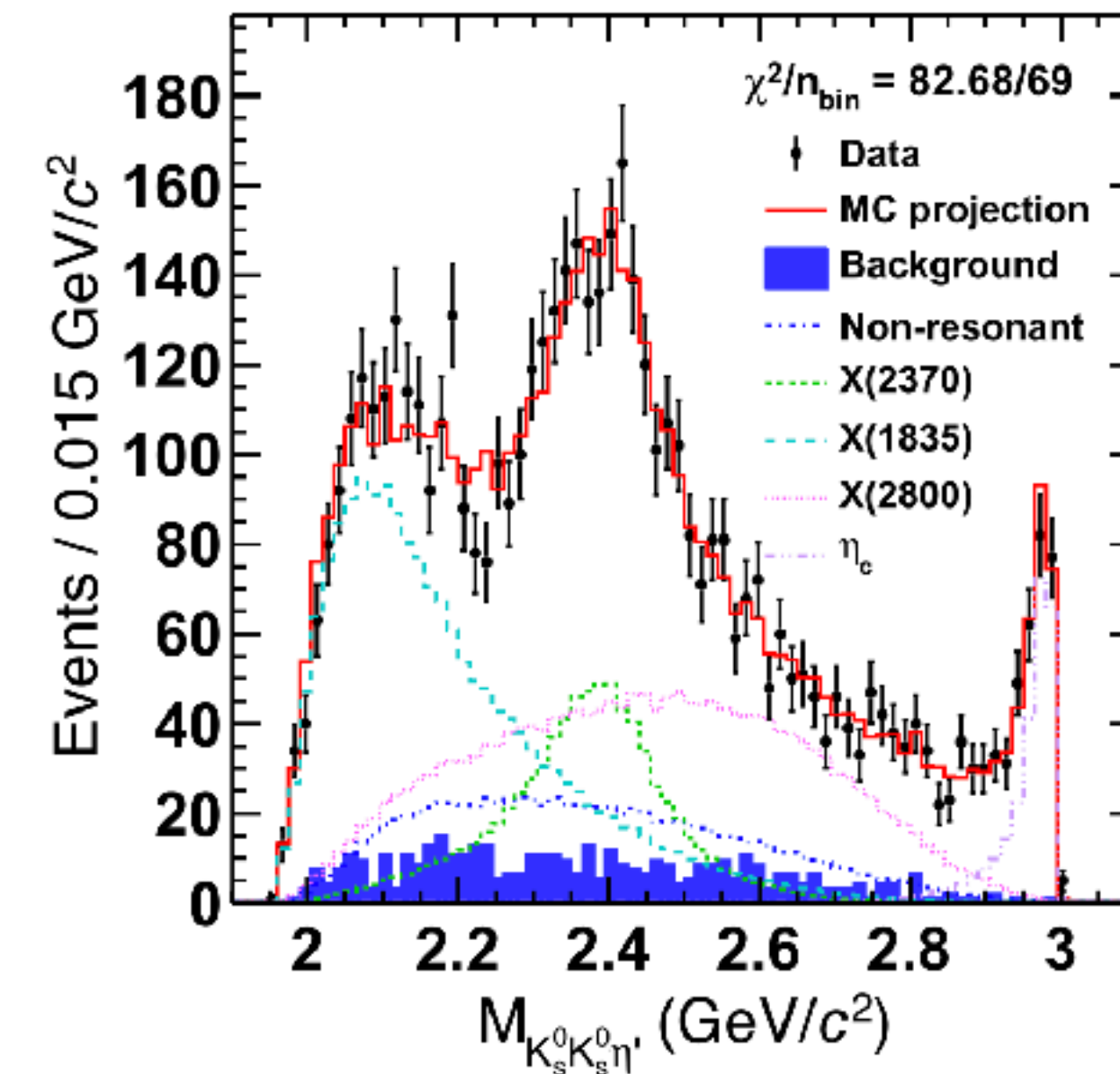


Radiative  $J/\psi$  decays are gluon rich!



$$\mathcal{R}_{\psi\gamma} = \frac{\mathcal{B}_{B^+ \rightarrow (\chi_{c1}(3872) \rightarrow \psi(2S)\gamma) K^+}}{\mathcal{B}_{B^+ \rightarrow (\chi_{c1}(3872) \rightarrow J/\psi\gamma) K^+}} = 1.67 \pm 0.21 \pm 0.12 \pm 0.04.$$

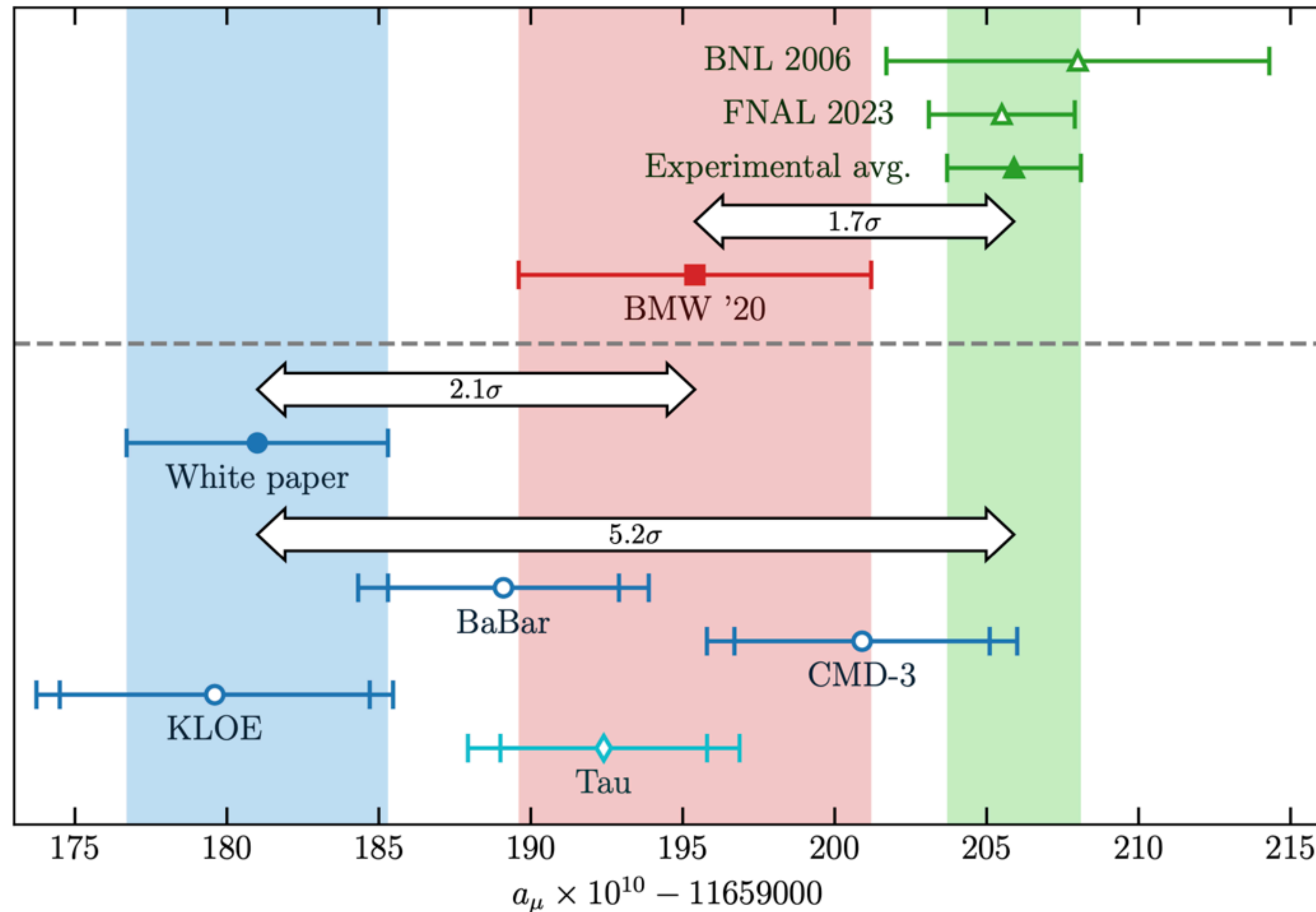
**Strong indication of a sizeable charmonium or tetraquark compact component of the X(3872)!**



Candidate for elusive lightest pseudoscalar glueball predicted by LQCD

# The Charged Lepton Sector

Several important news from the front of  $(g_\mu - 2)$  predictions from BaBar and Lattice in conjunction with data!



Scanning  $e^+e^- \rightarrow \pi^+ \pi^-$  ECM = 0.32-2 GeV

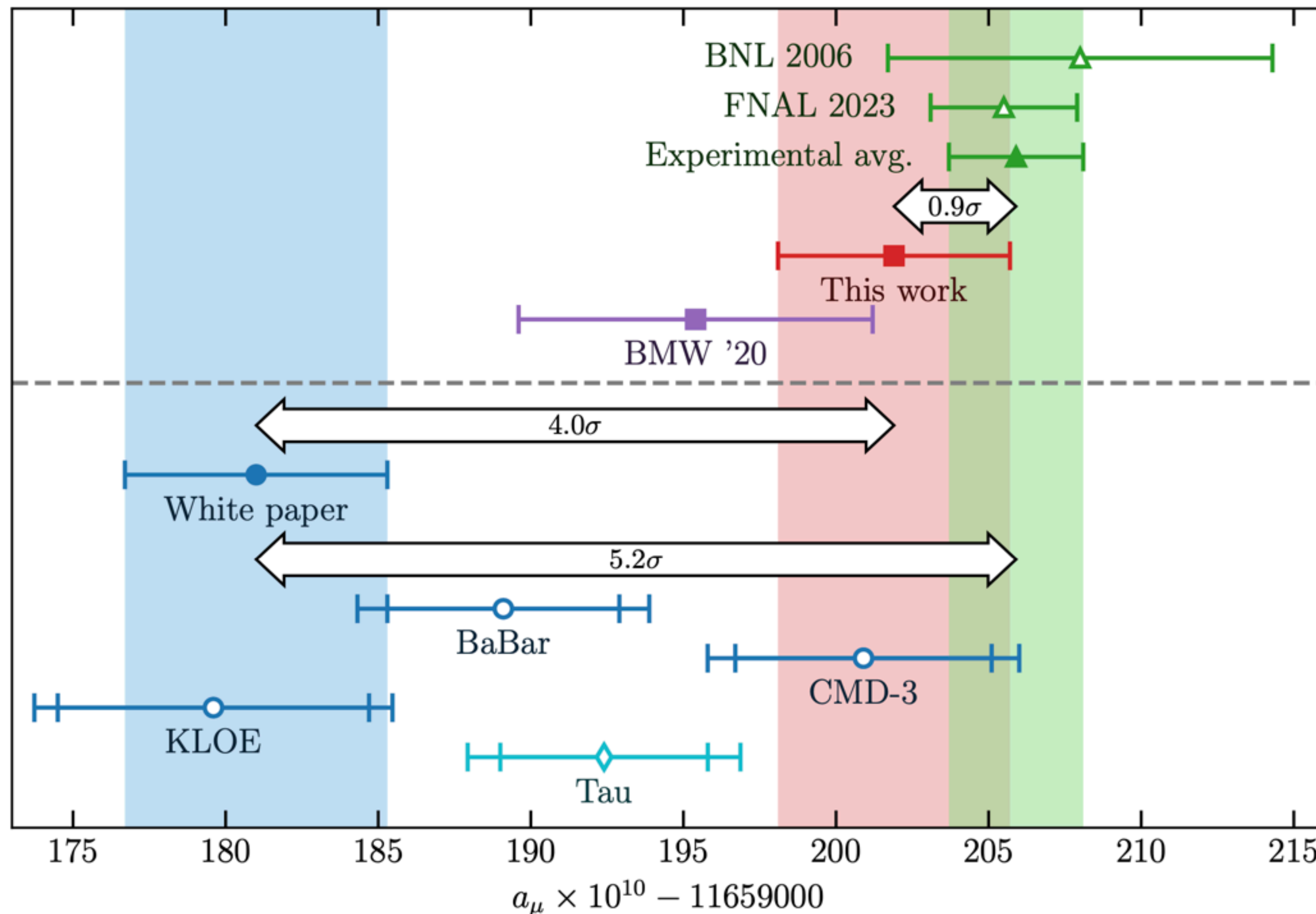
**CMD-3** at VEPP-2000  $e^+e^-$  collider

Better detector performance Larger statistics (x30 CMD-2)

# Muon Anomalous Magnetic Moment

Toru Lijima, Bogdan Malaescu, Zoltan Fodor

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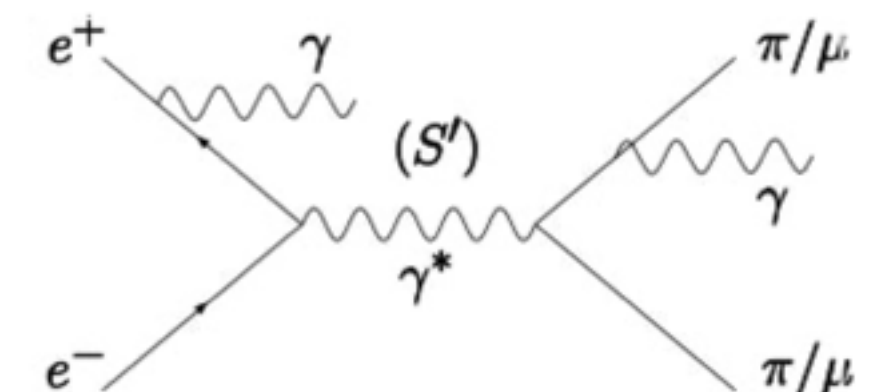
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**CMD-3** at VEPP-2000  $e^+e^-$  collider

Better detector performance Larger statistics (x30 CMD-2)

New BMW result including finer lattice and long distance effects from  $e^+e^-$  data!

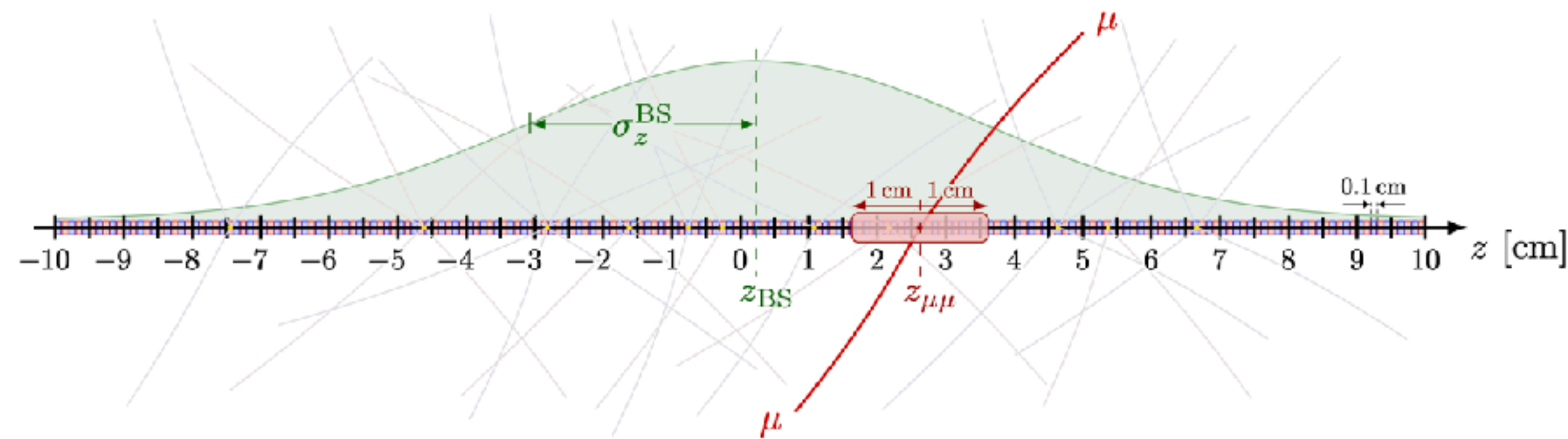
New BaBar studies of higher order radiation and impact on the vacuum polarisation predictions of  $(g-2)$ !

**NNLO Radiative corrections need to be better understood and accounted in analyses!**



# Tau Anomalous Magnetic Moments

Toru Lijima, Maurizio Pierini



Beautiful analysis selecting isolated low multiplicity vertices, sensitive to photo-production of tau pairs!

Large gain in sensitivity! Only ~3 times larger than the Schwinger term (QED part)

However still almost 3-4 orders of magnitude above sensitive corrections e.g. EW!

$$\begin{aligned}
 a_\tau^{\text{QED}} &= 1.1732 \times 10^{-3}, \\
 a_\tau^{\text{had}} &= 3.2(4) \times 10^{-6}, \\
 a_\tau^{\text{EW}} &= 4.7 \times 10^{-7}.
 \end{aligned}$$

## CMS

138 fb<sup>-1</sup> (13 TeV)

• Observed — 68% CL — 95% CL

**OPAL**  
 $ee \rightarrow Z \rightarrow \tau\tau\gamma$   
 PLB 434 (1998) 188

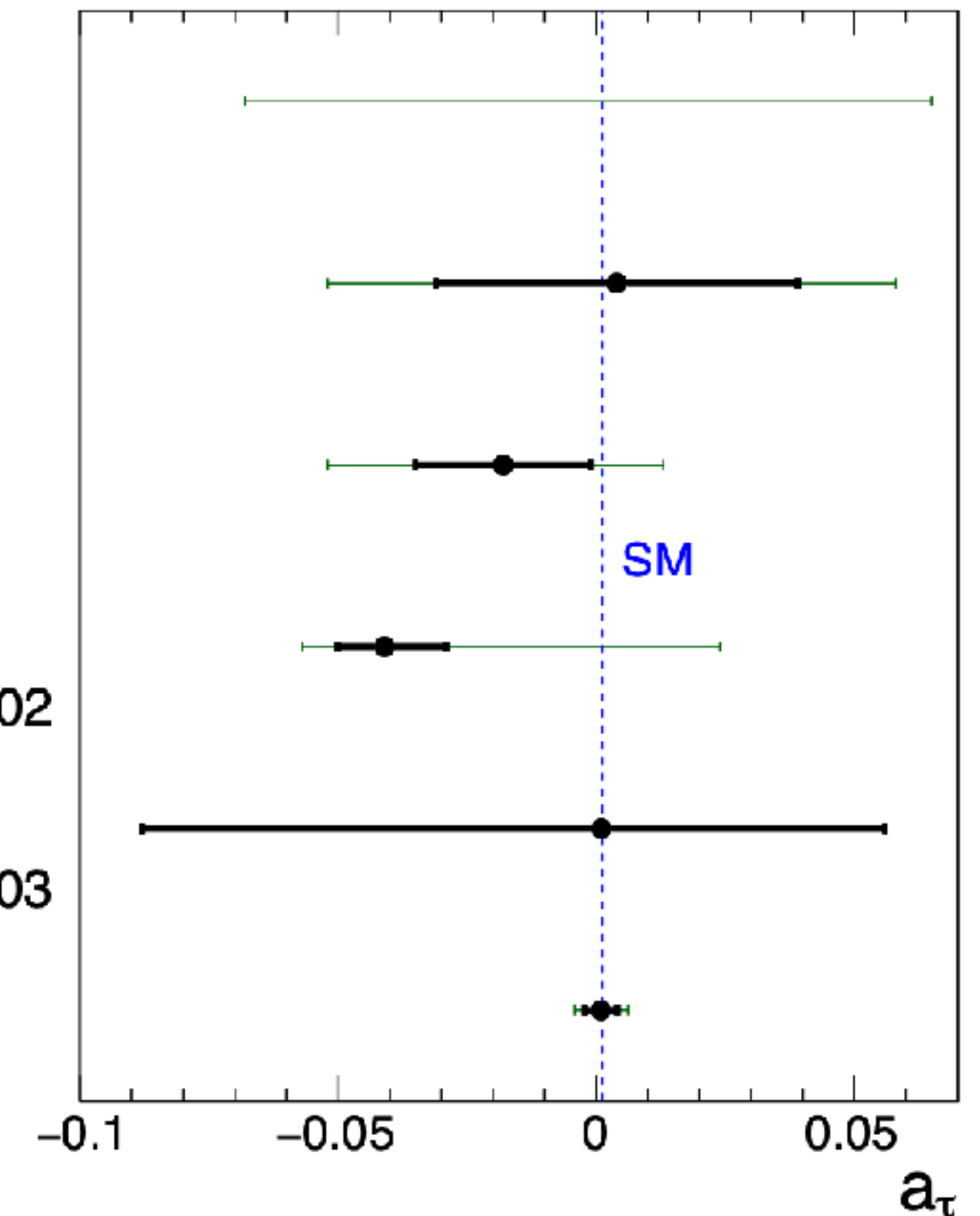
**L3**  
 $ee \rightarrow Z \rightarrow \tau\tau\gamma$   
 PLB 434 (1998) 169

**DELPHI**  
 $\gamma\gamma \rightarrow \tau\tau$  ( $\gamma$  from e)  
 EPJC 35 (2004) 159

**ATLAS**  
 $\gamma\gamma \rightarrow \tau\tau$  ( $\gamma$  from Pb)  
 PRL 131 (2023) 151802

**CMS**  
 $\gamma\gamma \rightarrow \tau\tau$  ( $\gamma$  from Pb)  
 PRL 131 (2023) 151803

**CMS**  
 $\gamma\gamma \rightarrow \tau\tau$  ( $\gamma$  from p)  
 This result



$$a_\tau = \frac{g_\tau}{2} - 1 = 0.0009^{+0.0032}_{-0.0031}$$

# The Neutrino Sector - Entering Precision Era

## Accelerator Neutrino Oscillations

The current two main players  $\nu_\mu$ -beam experiments!

### NOvA

#### Off Axis

Fermilab to Ash River

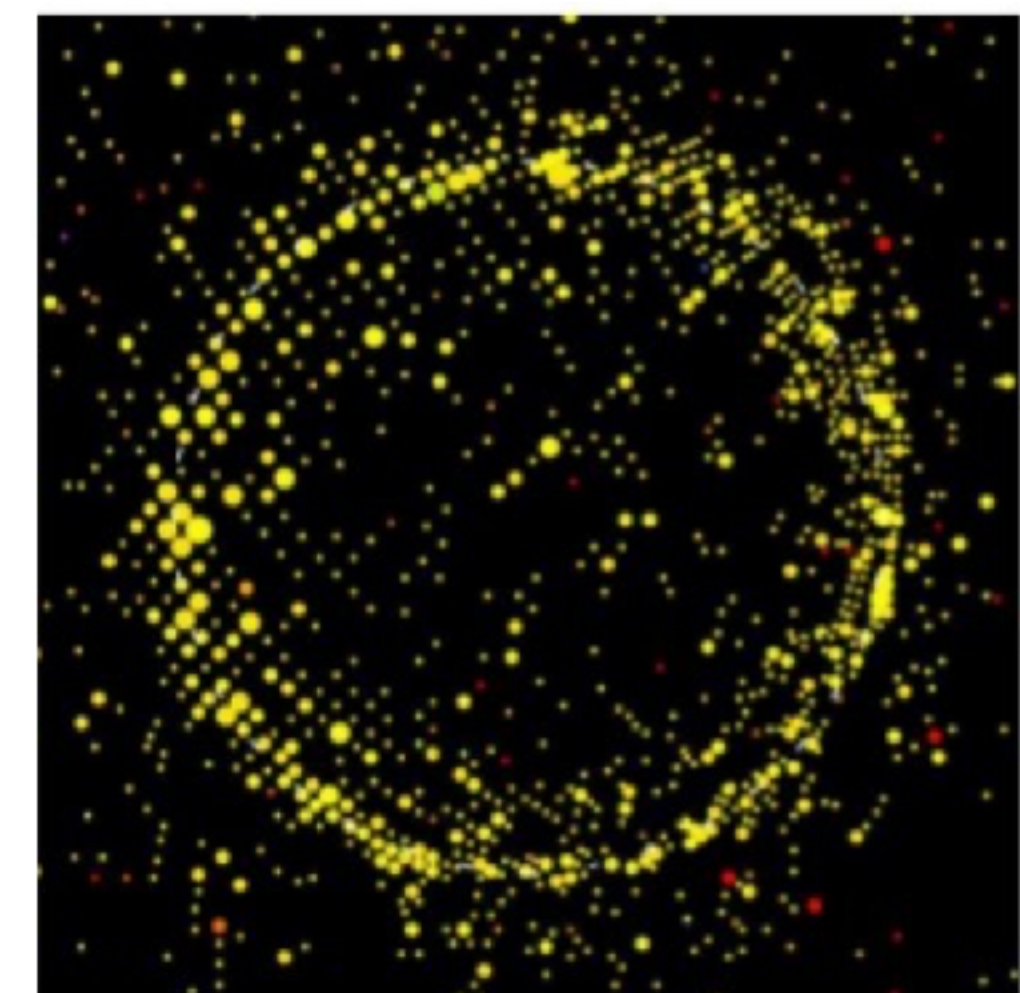
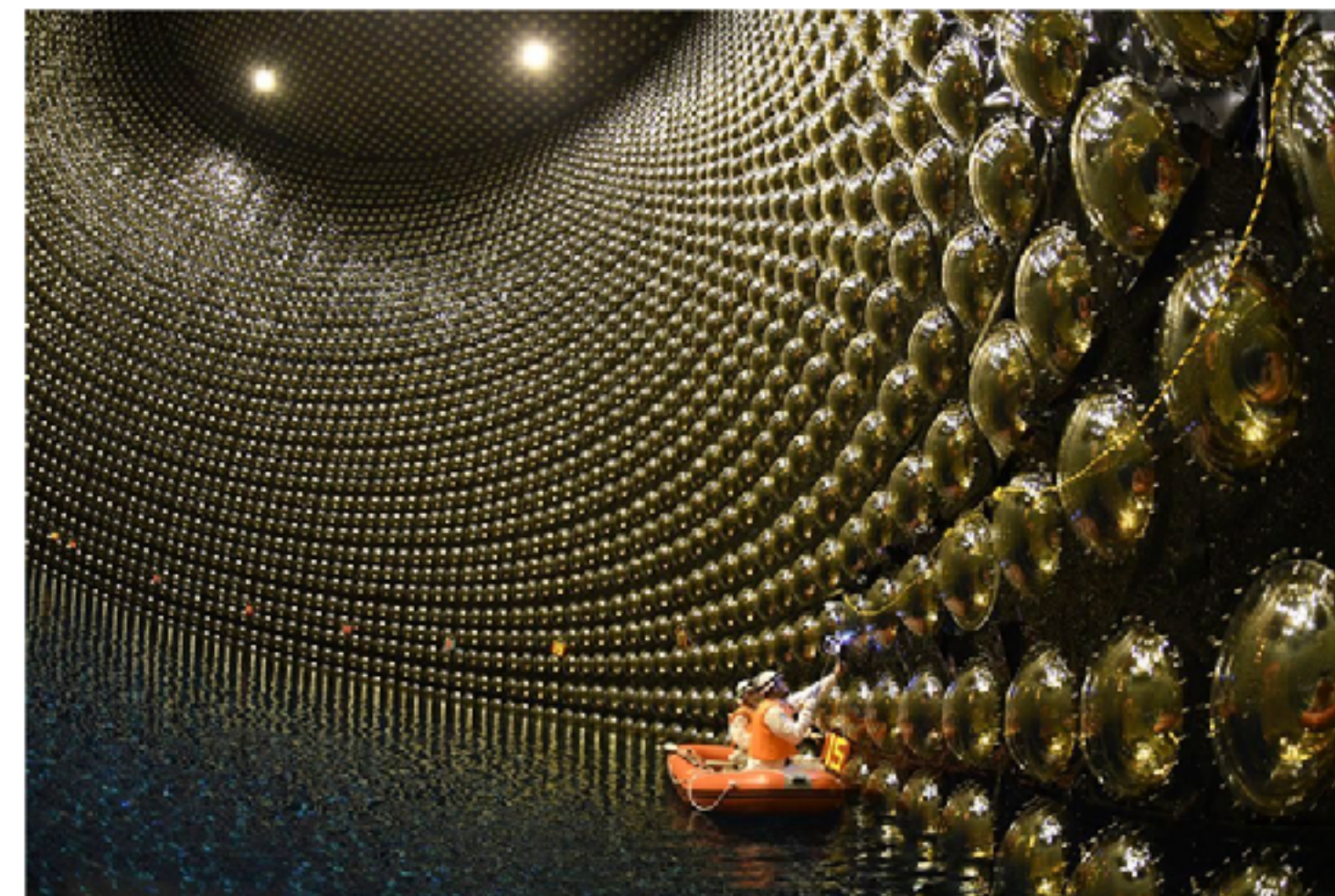
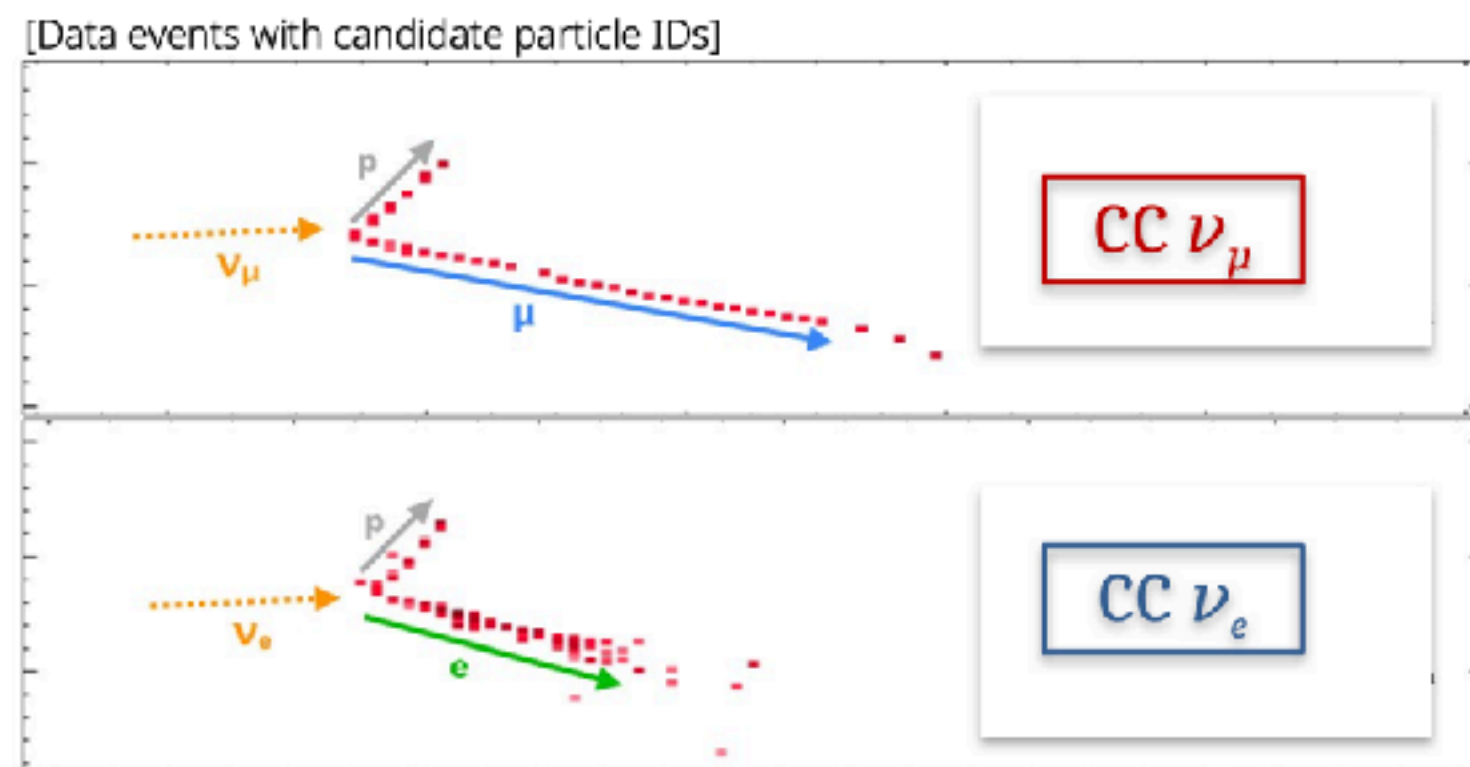
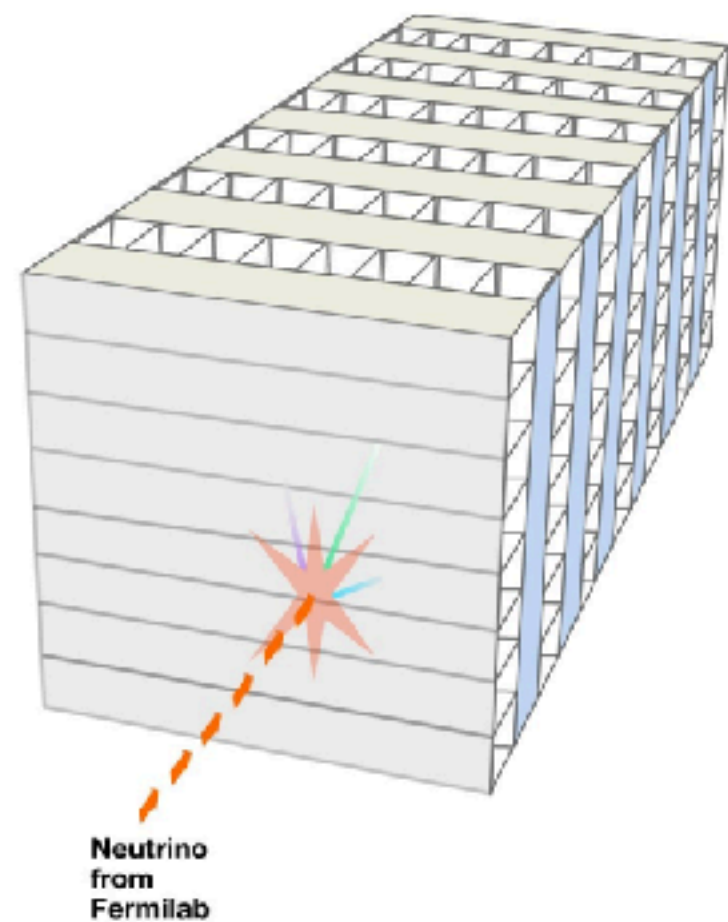
Improved sensitivity to mass ordering!  
810 km/GeV - E 2 GeV -  $0.8^\circ$  off-axis

### T2K

Slightly off axis (J-PARC to Super K)

490 km/GeV - E 0.6 GeV -  $2.5^\circ$  off-axis

Sensitive to  $\nu_\mu$  disappearance and  $\nu_e$  appearance!



$\nu_e$ -like



# Accelerator Neutrino Oscillations

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The current two main players  $\nu_\mu$ -beam experiments!

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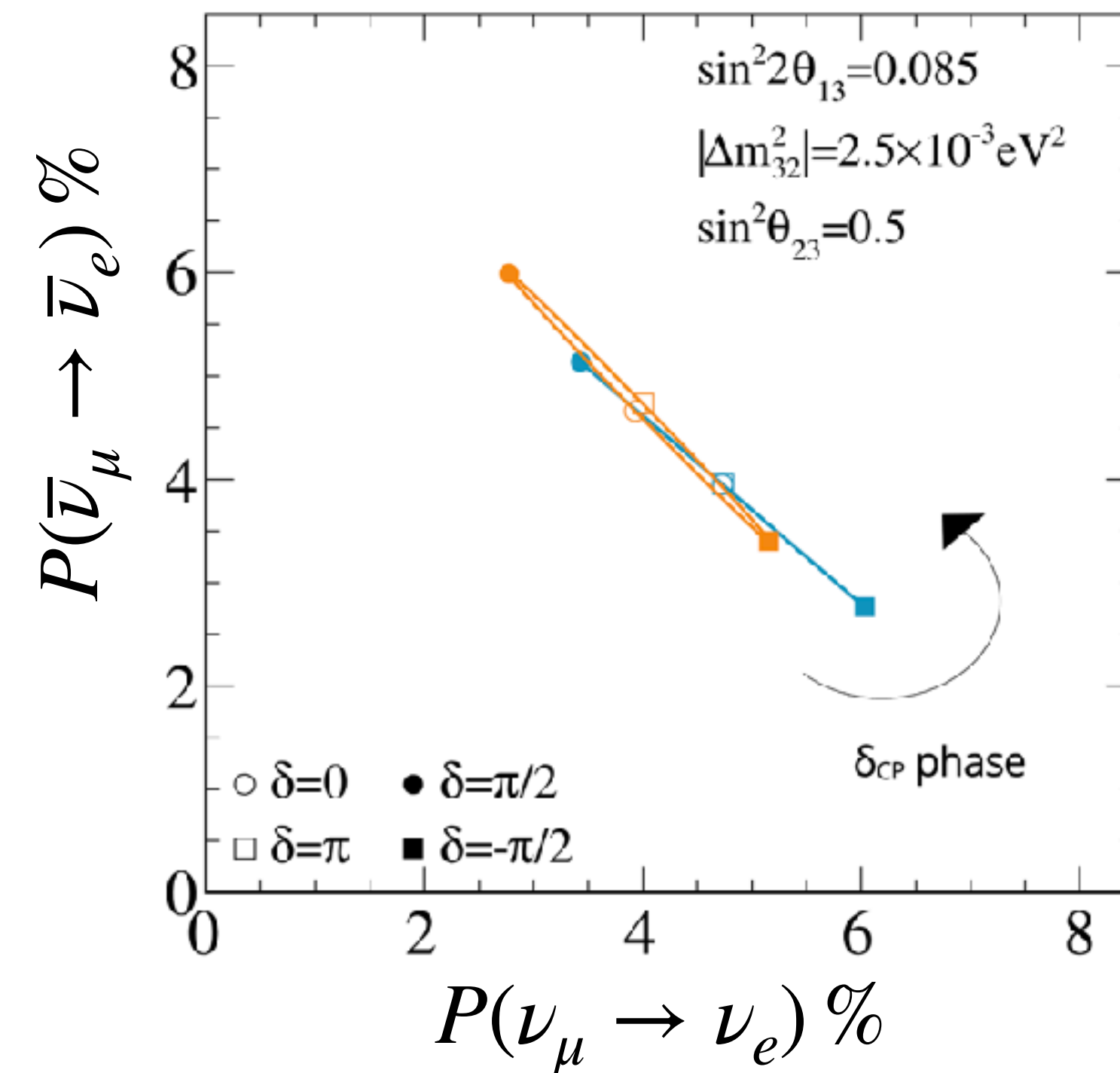
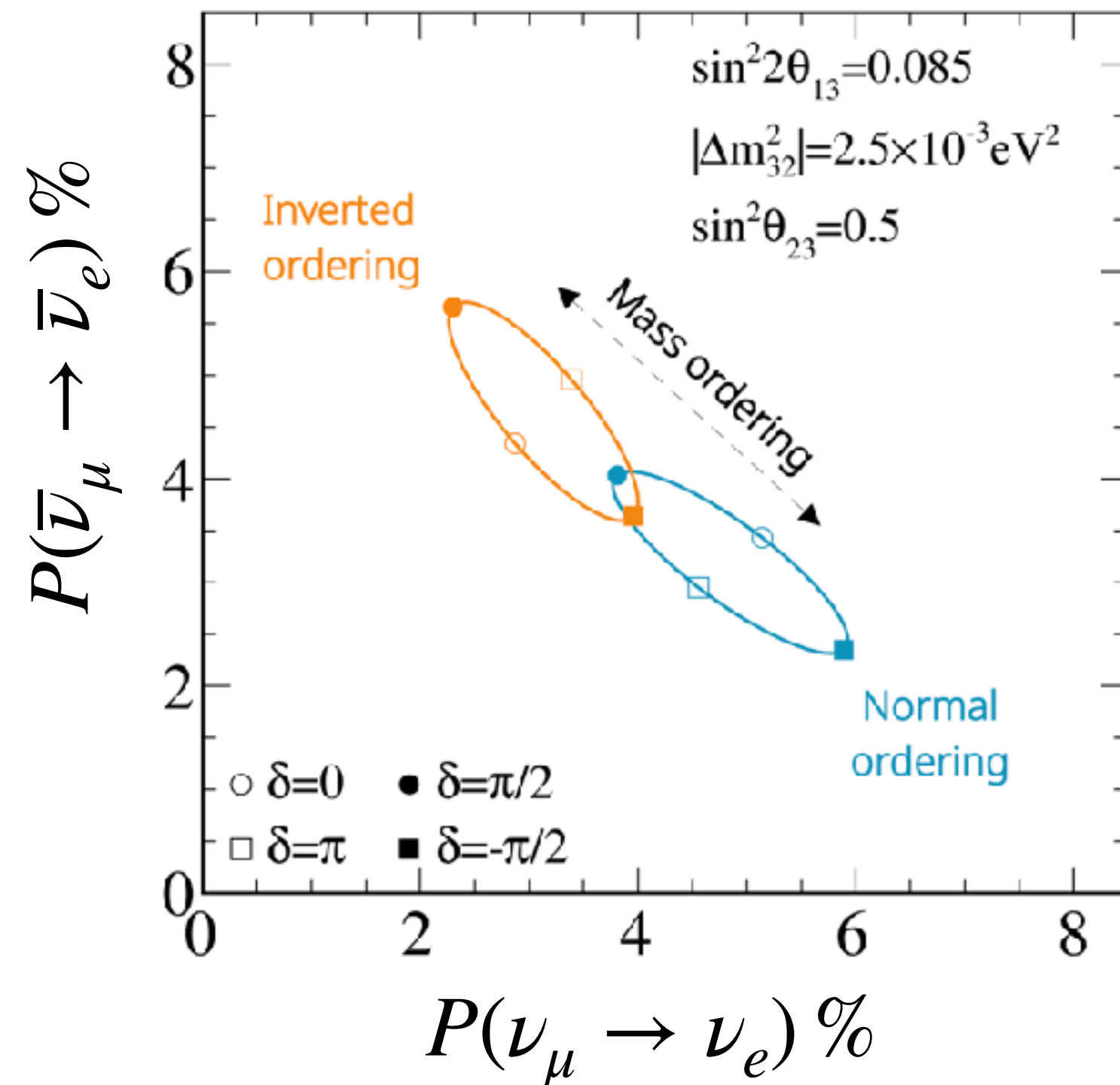
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# Accelerator Neutrino Oscillations

## Accelerator Neutrino Oscillations

The current two main players  $\nu_\mu$ -beam experiments!

New combination of the two experiments for this conference!

**NOvA**

**Off Axis**

Fermilab to Ash River

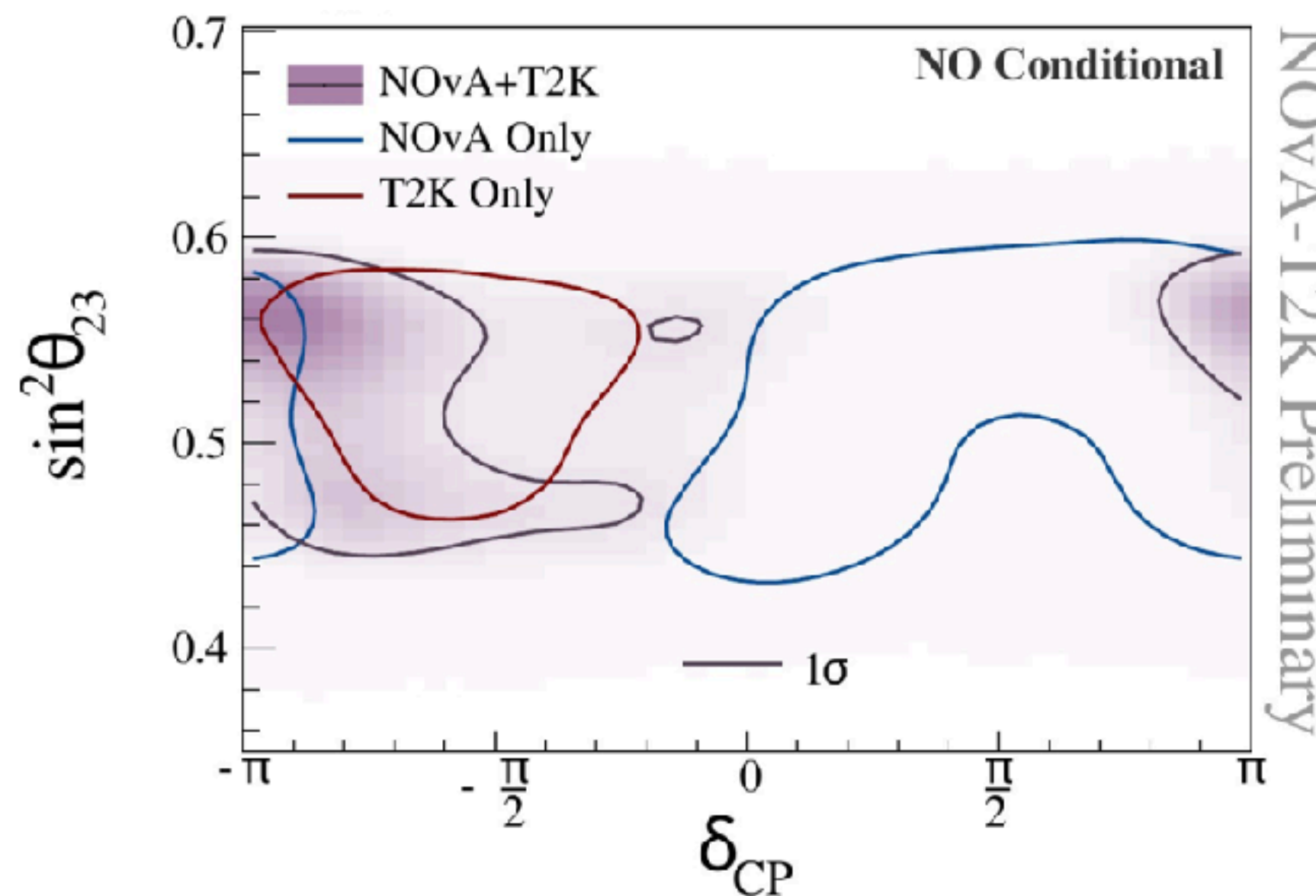
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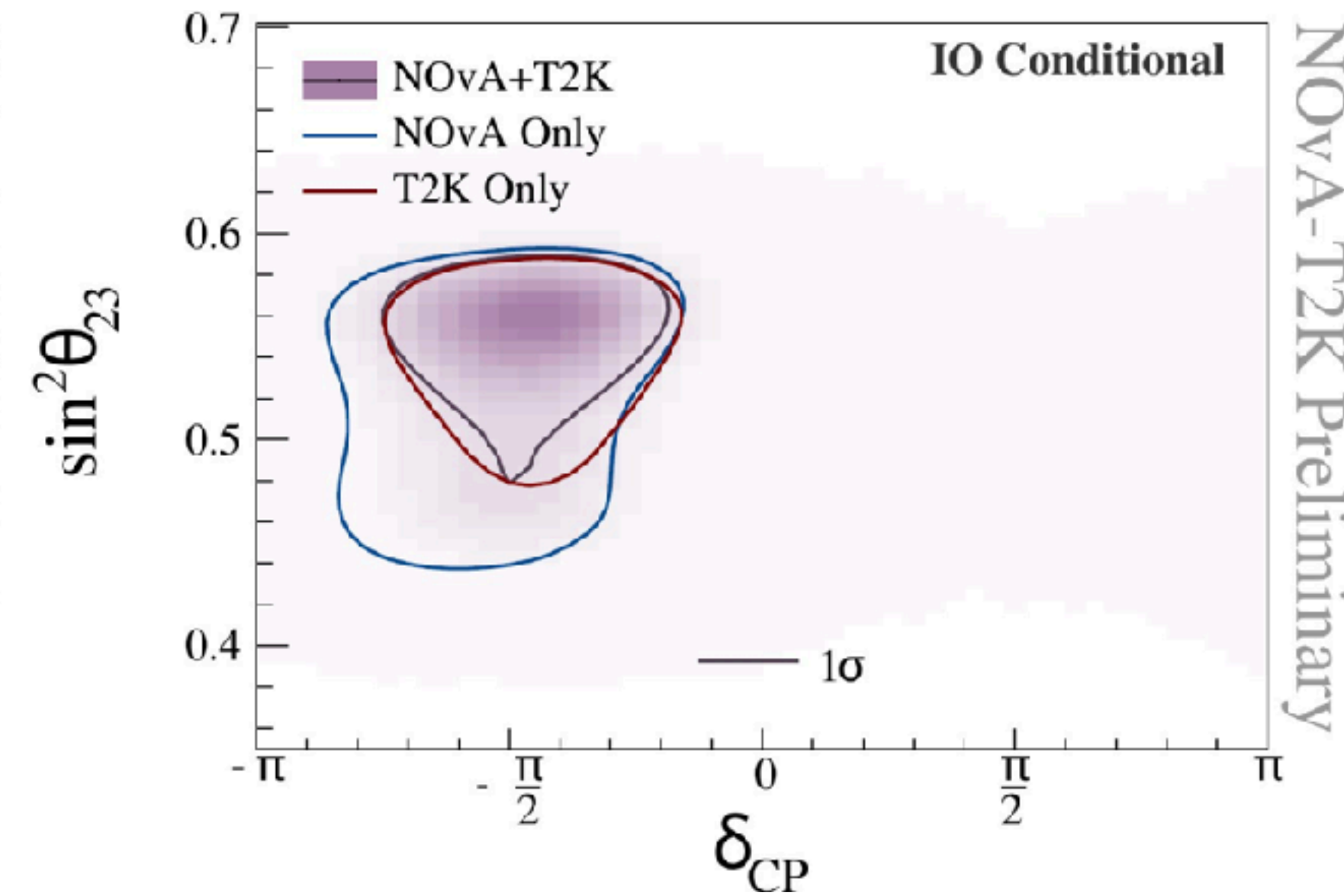
**T2K**

**Slightly off axis (J-PARC to Super K)**

490 km/GeV - E 0.6 GeV -  $2.5^\circ$  off-axis



Mild preference for Inverted Ordering  
but influenced by  $\theta_{13}$  constraint



CP-conserving points are *outside*  
 $3\sigma$  intervals in IO  
Expect CPV *if* ordering is inverted

# Accelerator Neutrino Prospects

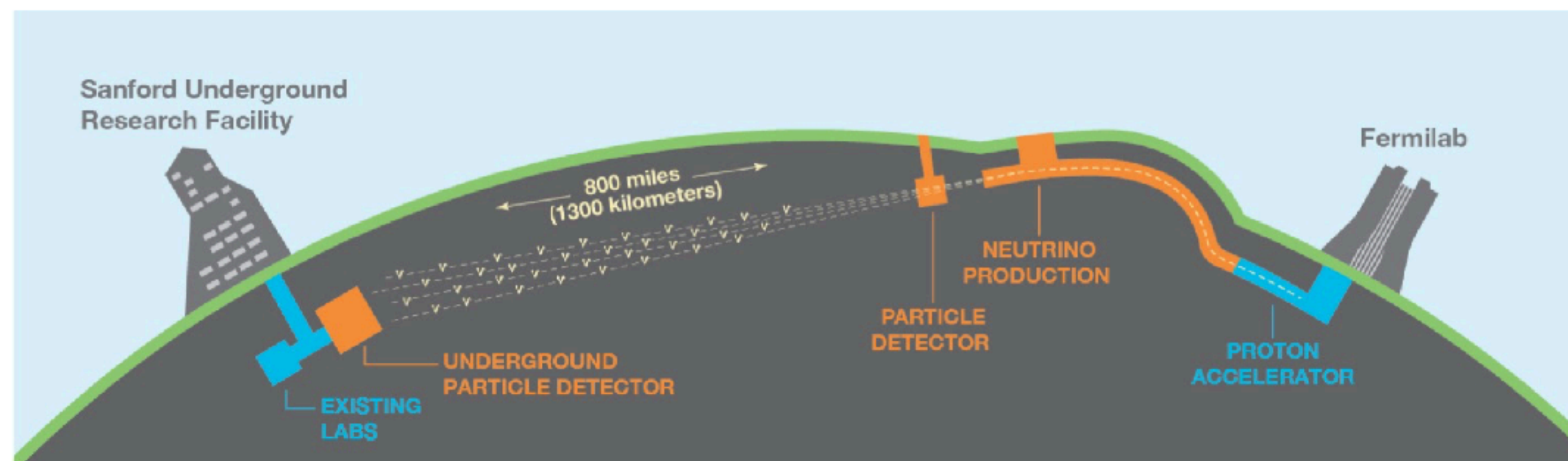
- Need definitive measurements! Two large next-generation projects are under preparation:

## DUNE:

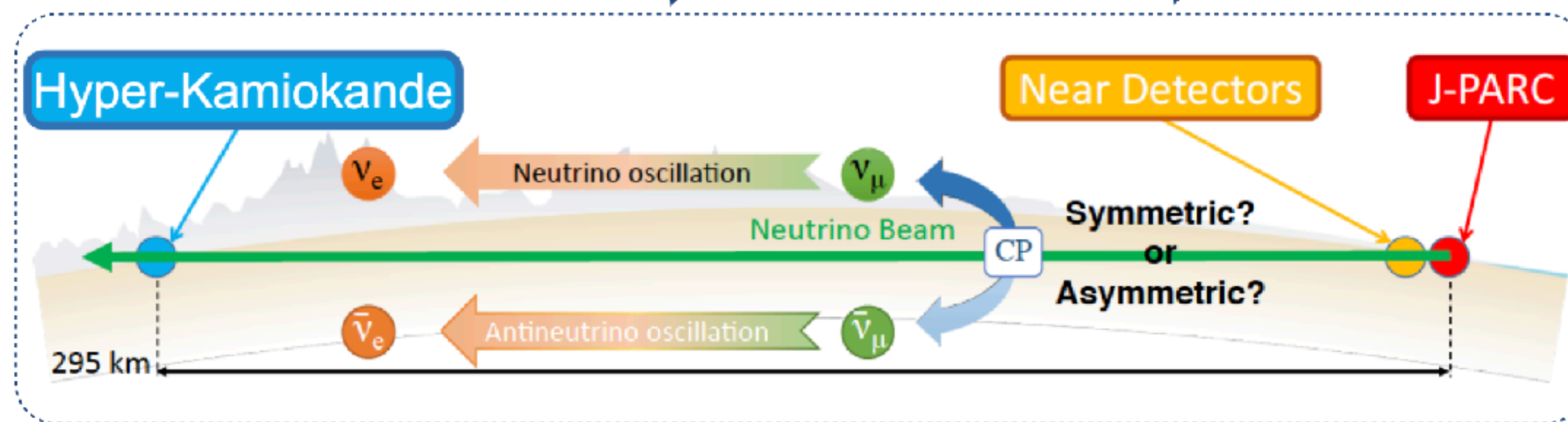
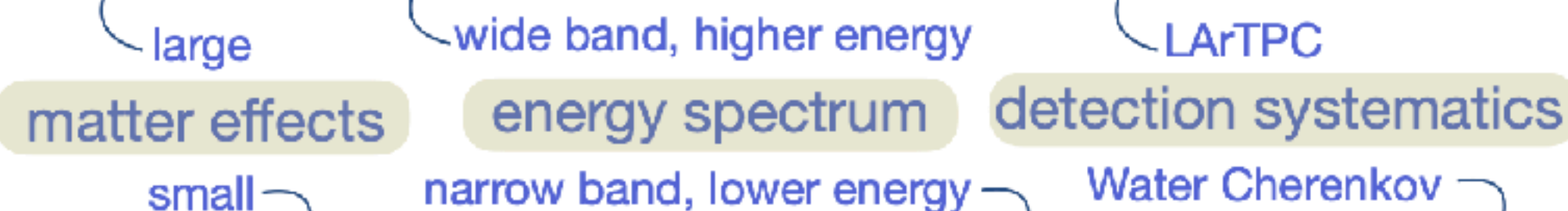
- > 2 MW beam
- Liquid-Argon TimeProjection Chamber (LArTPC) technology
- $\geq 40$  kton far detector fiducial mass
- First physics in ~2029

## Hyper-Kamiokande:

- 1.3 MW beam
- Water Cherenkov far detector
- 190 kton far detector fiducial mass
- First physics in ~2027



Large degree of complementarity:

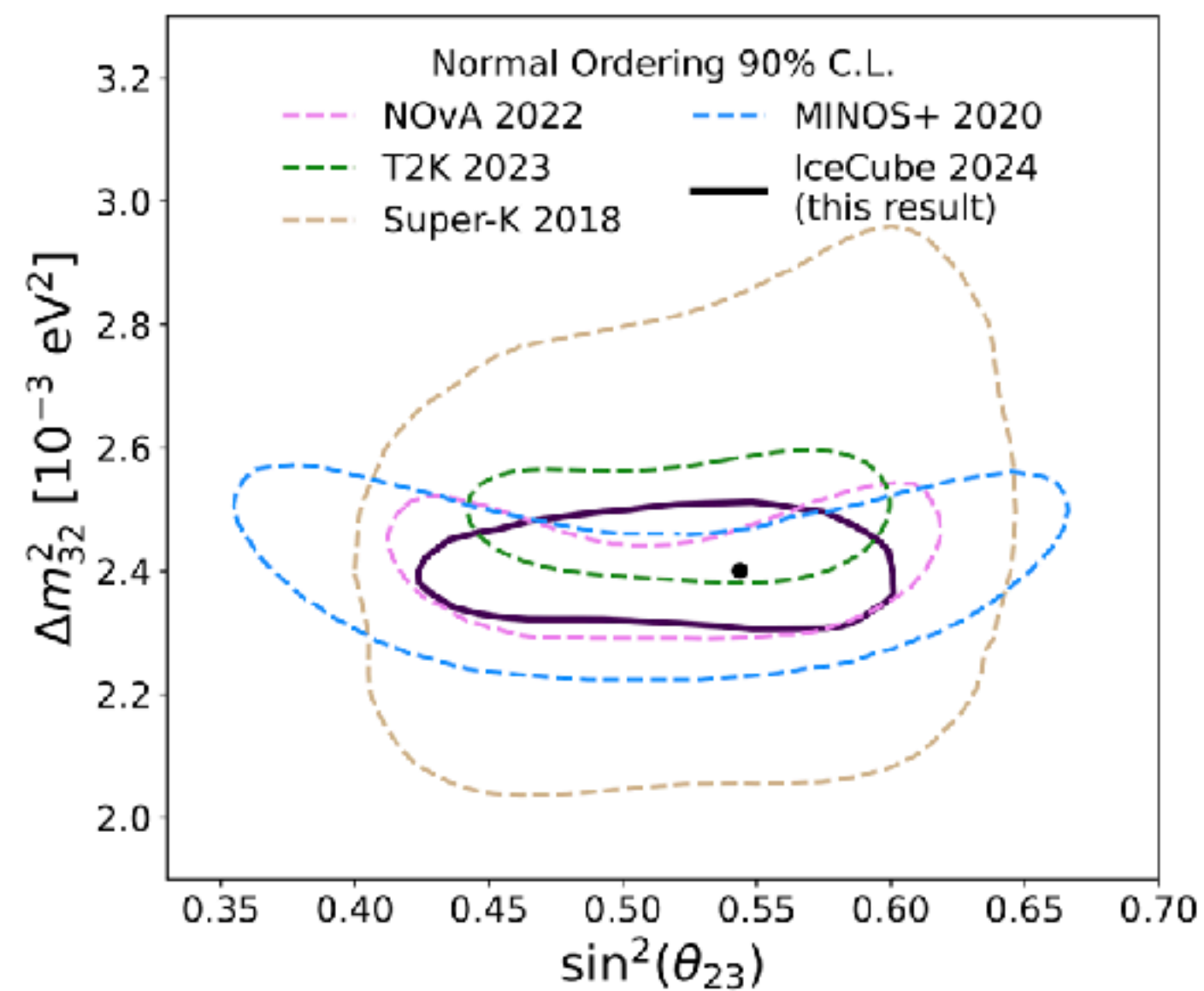


From T. Nakadaira's [talk](#) at ICHEP 2024

# Atmospheric Neutrinos

## ICECUBE and Deep Core

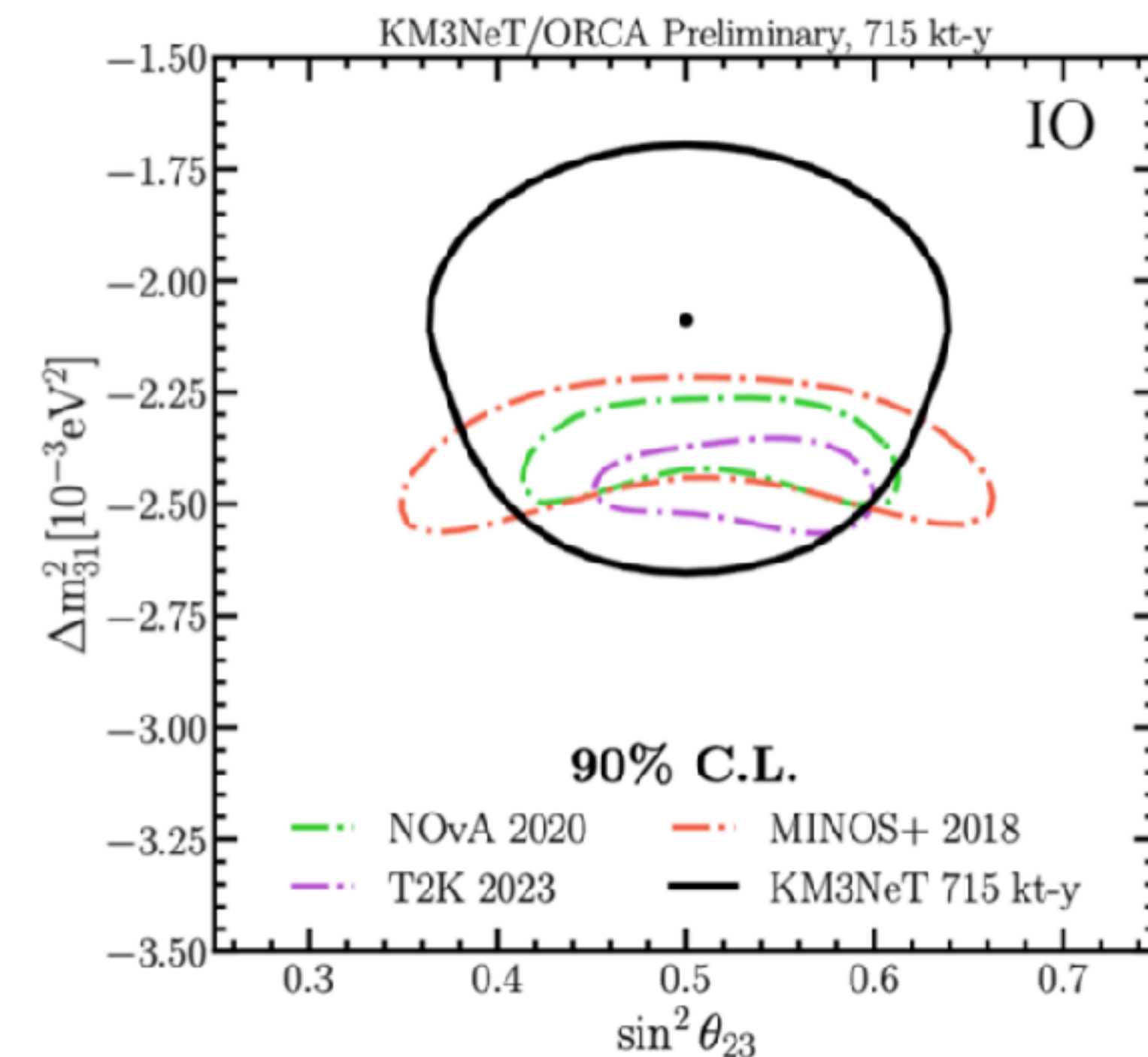
- 1km<sup>3</sup> of ice instrumented with strings of Digital Optical Modules (DOMs), each with a PMT
- DeepCore: 8 densely region at the center (threshold ~8 GeV)



- Comparable and compatible results between accelerator and atmospheric neutrinos.

## Km3net and ORCA

- 115 strings optimized for neutrino oscillation measurements
- Each DOM has 31 3-inch PMTs
- About 20% of DOMs already installed



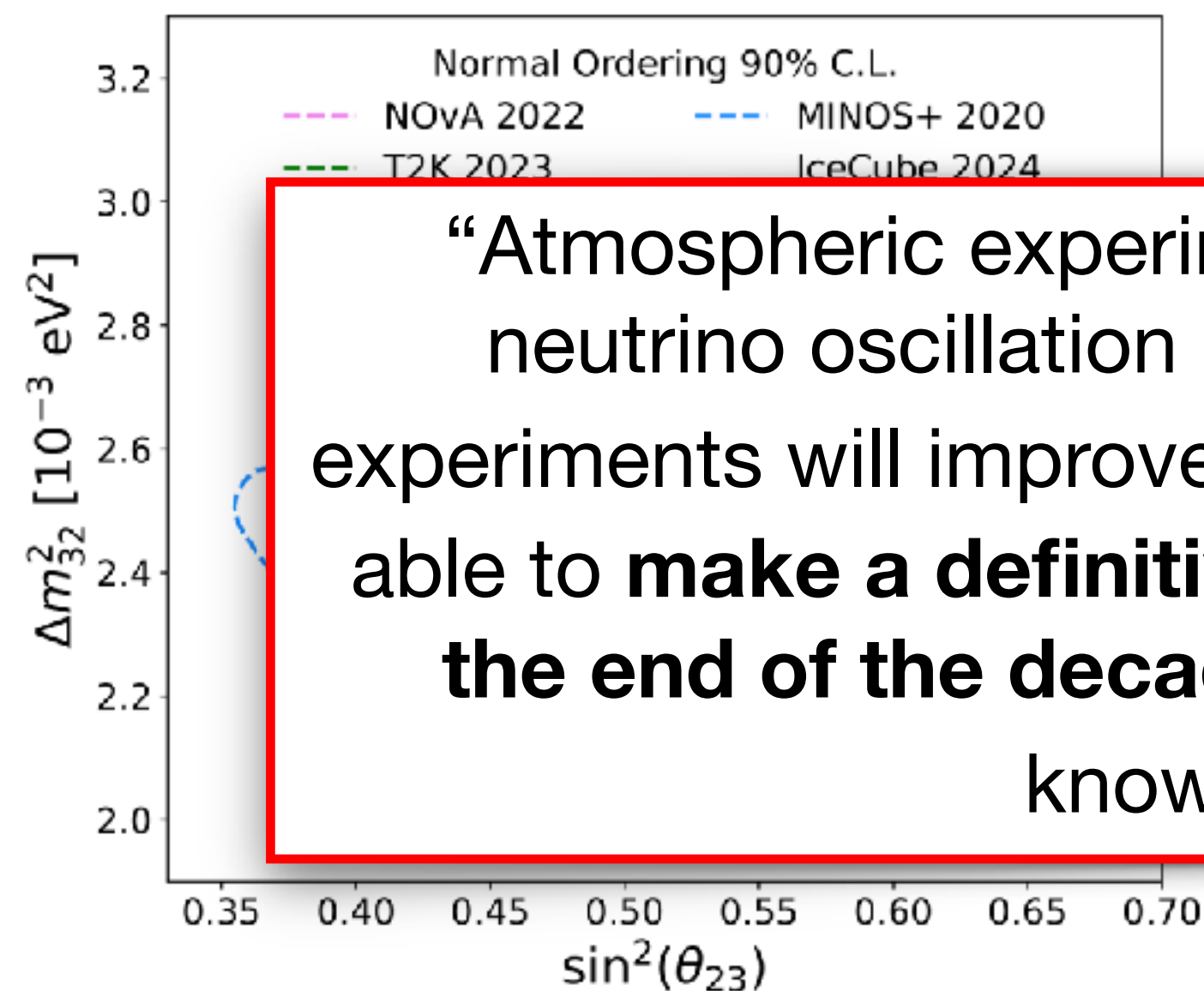
# Atmospheric Neutrinos

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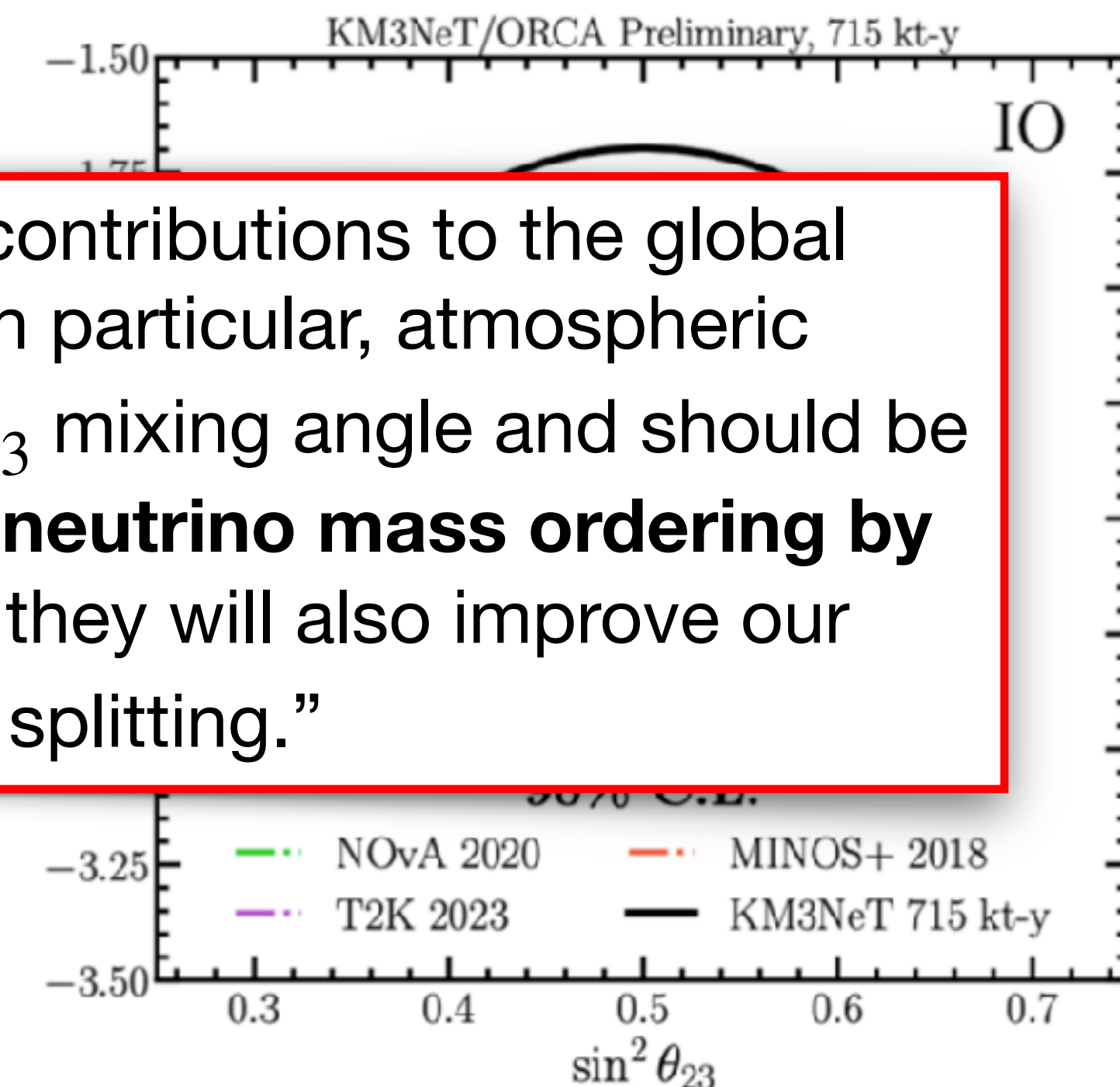
## Km3net and ORCA

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“Atmospheric experiments will make leading contributions to the global neutrino oscillation landscape before 2030. In particular, atmospheric experiments will improve our knowledge of the  $\theta_{23}$  mixing angle and should be able to **make a definitive measurement of the neutrino mass ordering by the end of the decade.** Together with JUNO, they will also improve our knowledge of the  $\Delta m_{32}^2$  mass splitting.”

- Comparable and compatible results between accelerator and atmospheric neutrinos.



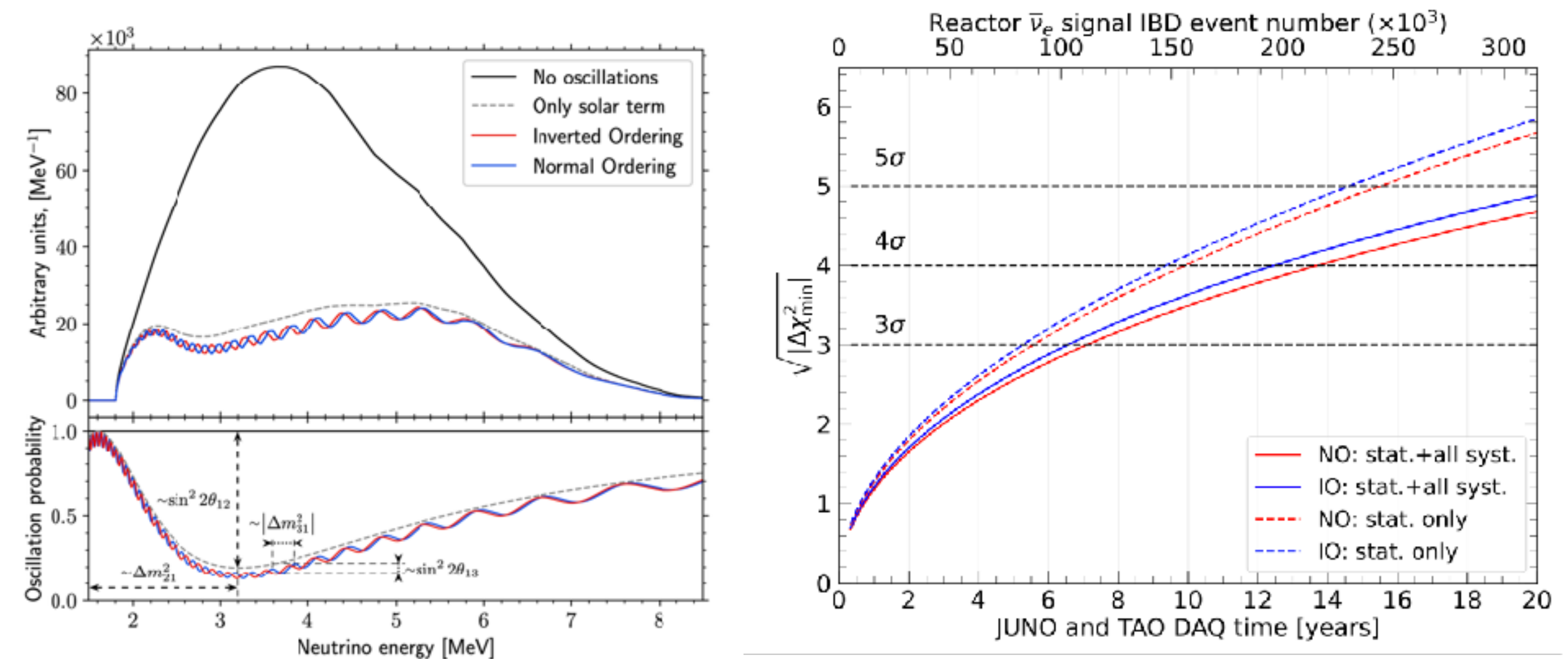
# Reactor Oscillations

## JUNO Medium baseline experiment

20-kton Liquid Scintillator neutrino observatory located in Southern China

Mass hierarchy from the electron anti-neutrino disappearance pattern through the interference effect of quasi-vacuum oscillation of reactor antineutrinos

JUNO reactor neutrino oscillation analysis alone provides a median  $3\sigma$  sensitivity to NMO in 6.5 years!



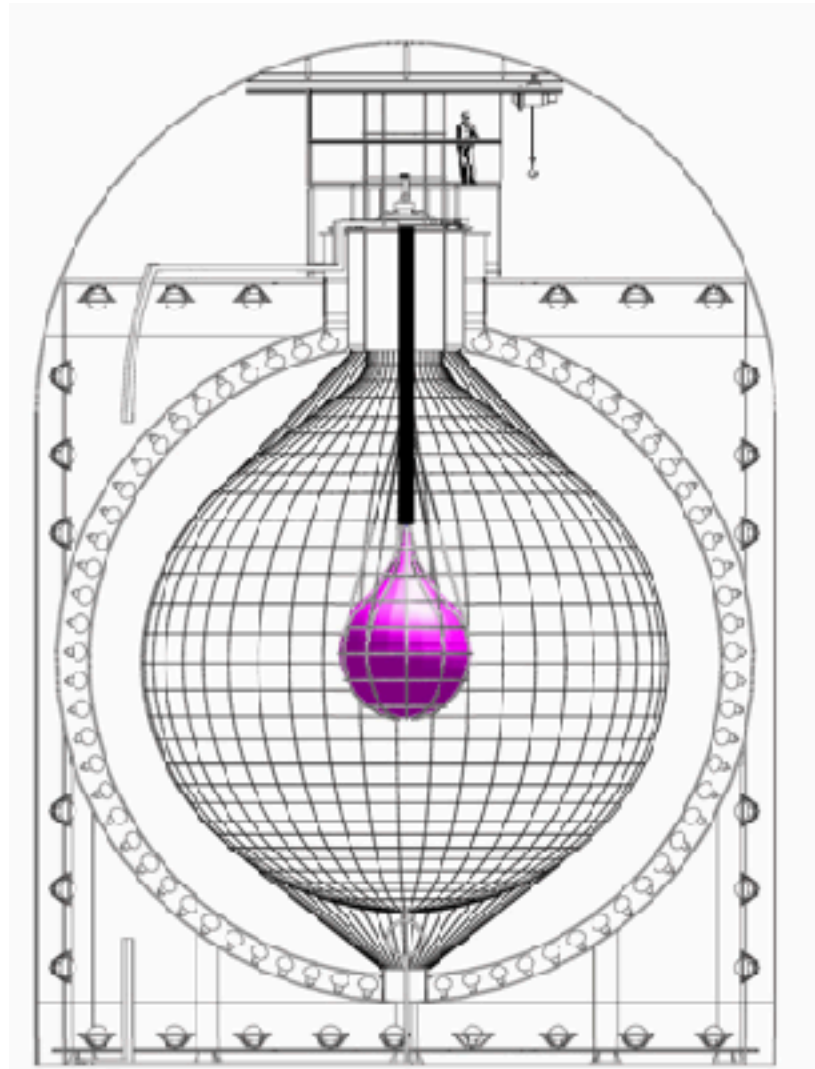
# Neutrinoless double beta decay

Susanne Mertens

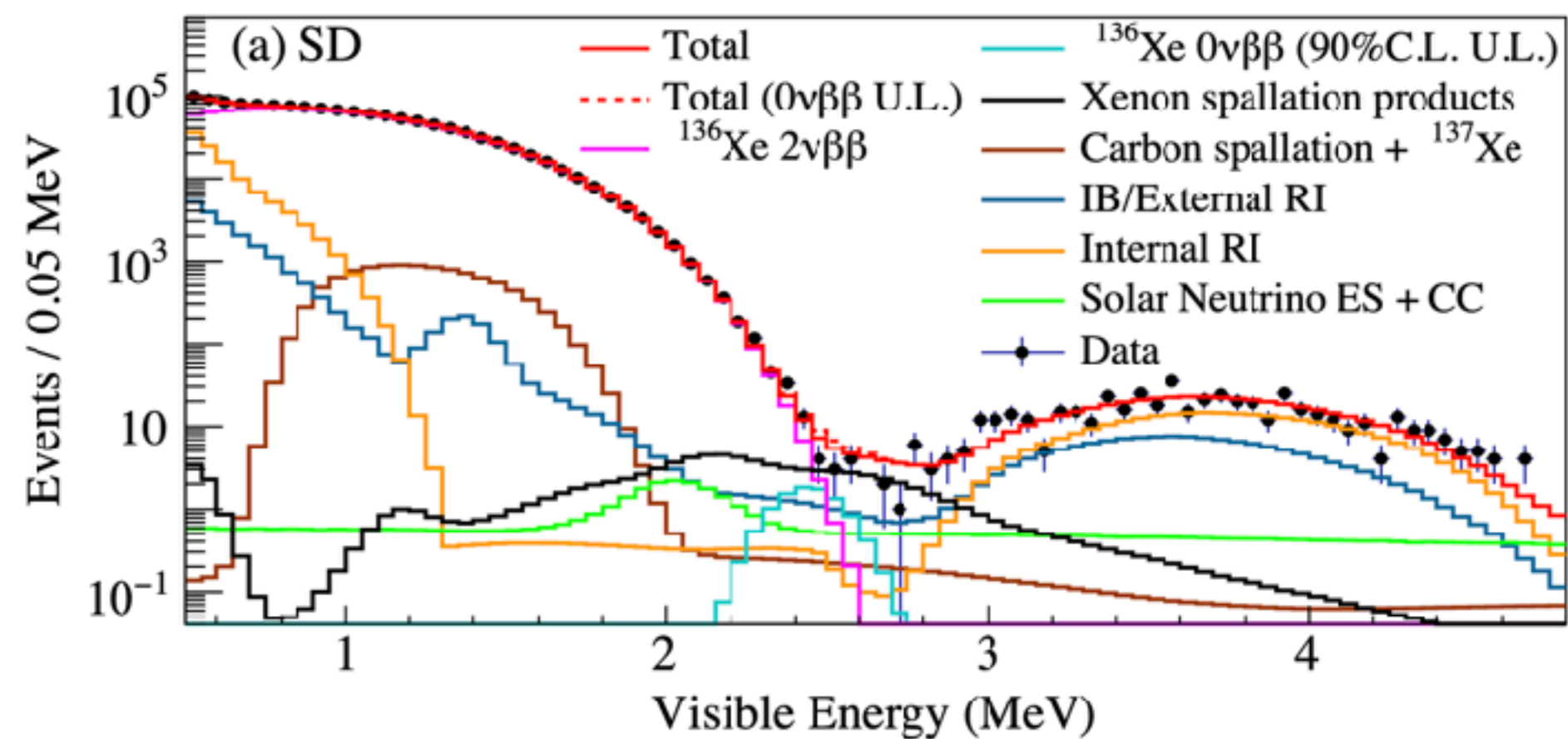
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## KamLandZen

Mini-balloon Radius = 1.90 m  
Xenon mass = 750 kg  
Data taking started in 2019



The largest number of  $\beta\beta$  nuclei. Low BG by distillation and filtration of both Liquid Scintillator and Xenon.



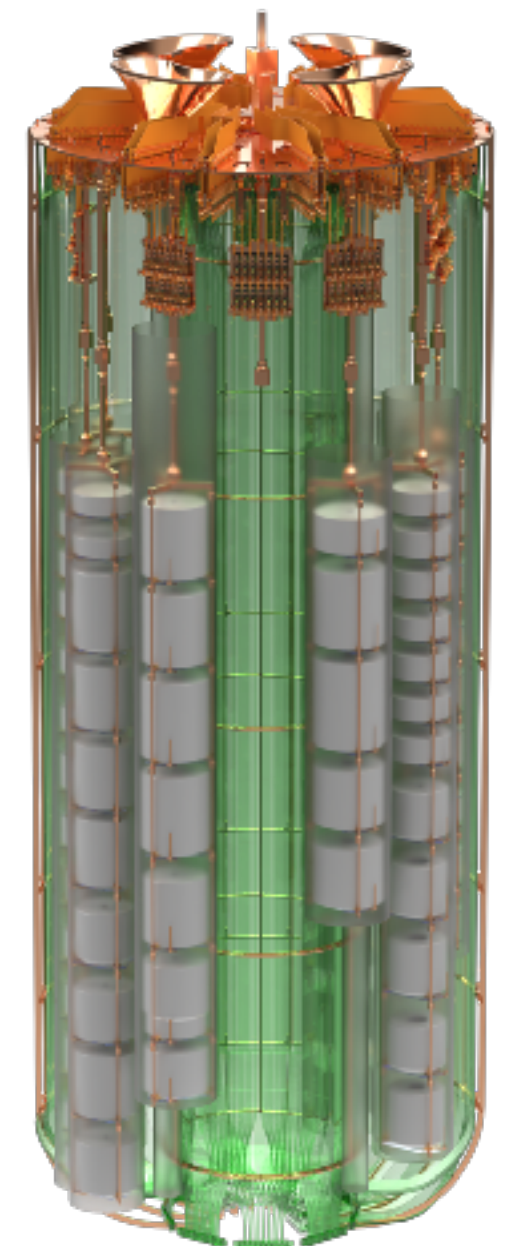
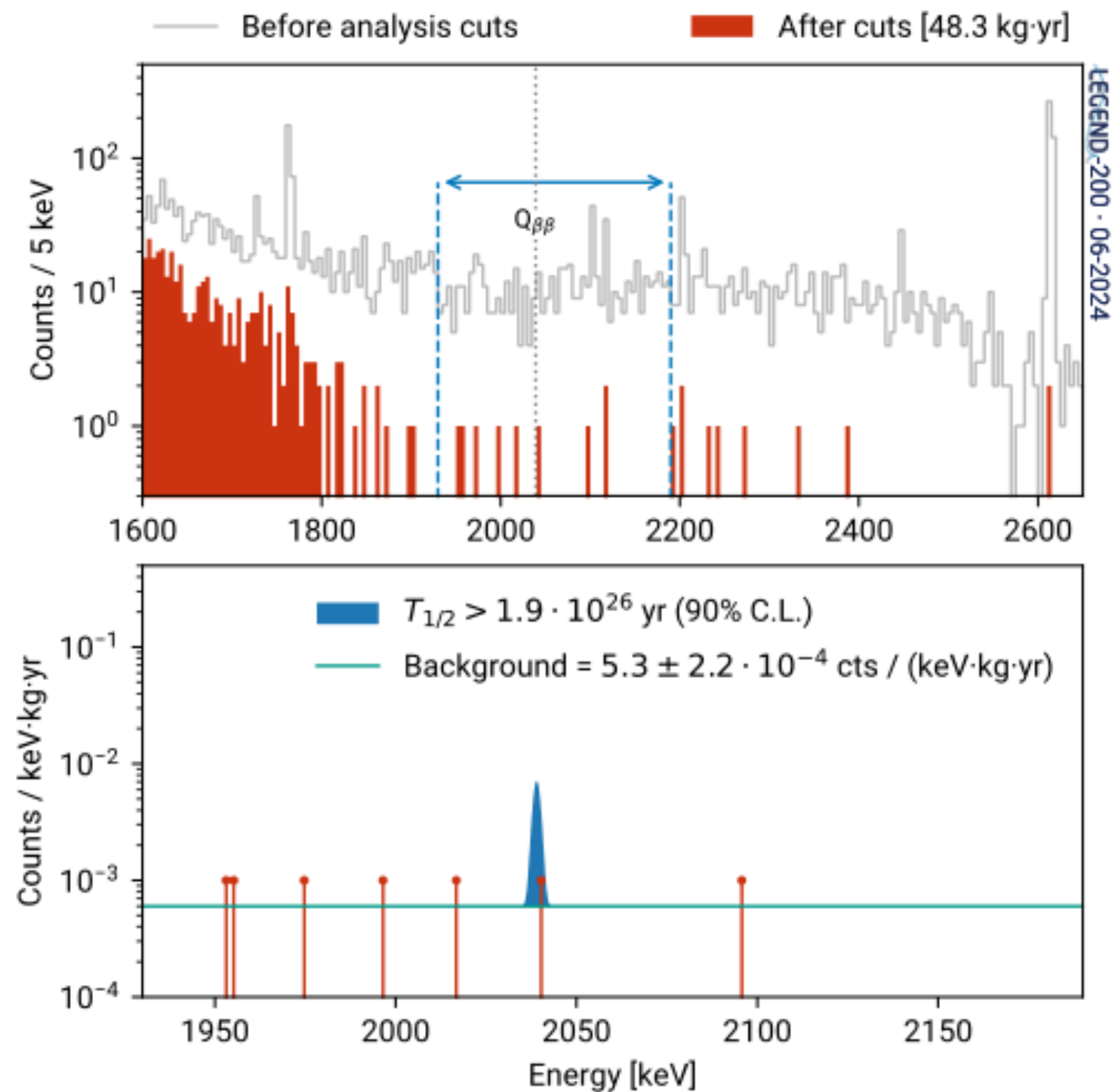
Limiting factor is the muon spallation of Xenon background.

$$\langle m_{\beta\beta} \rangle < 28 - 122 \text{ meV}$$

$$\text{KamLAND-ZEN-1T: } \langle m_{\beta\beta} \rangle < \sim 20 \text{ meV}$$

## Legend-200

Germanium Semiconductor, with enrichment to > 90% in  $^{76}\text{Ge}$  ( $Q_{\beta\beta}=2039 \text{ keV}$  Excellent energy resolution (0.1 % FWHM @  $Q_{\beta\beta}$ )



$$\langle m_{\beta\beta} \rangle < 75 - 178 \text{ meV}$$

Legend 1000

$$\langle m_{\beta\beta} \rangle < \sim 20 \text{ meV}$$

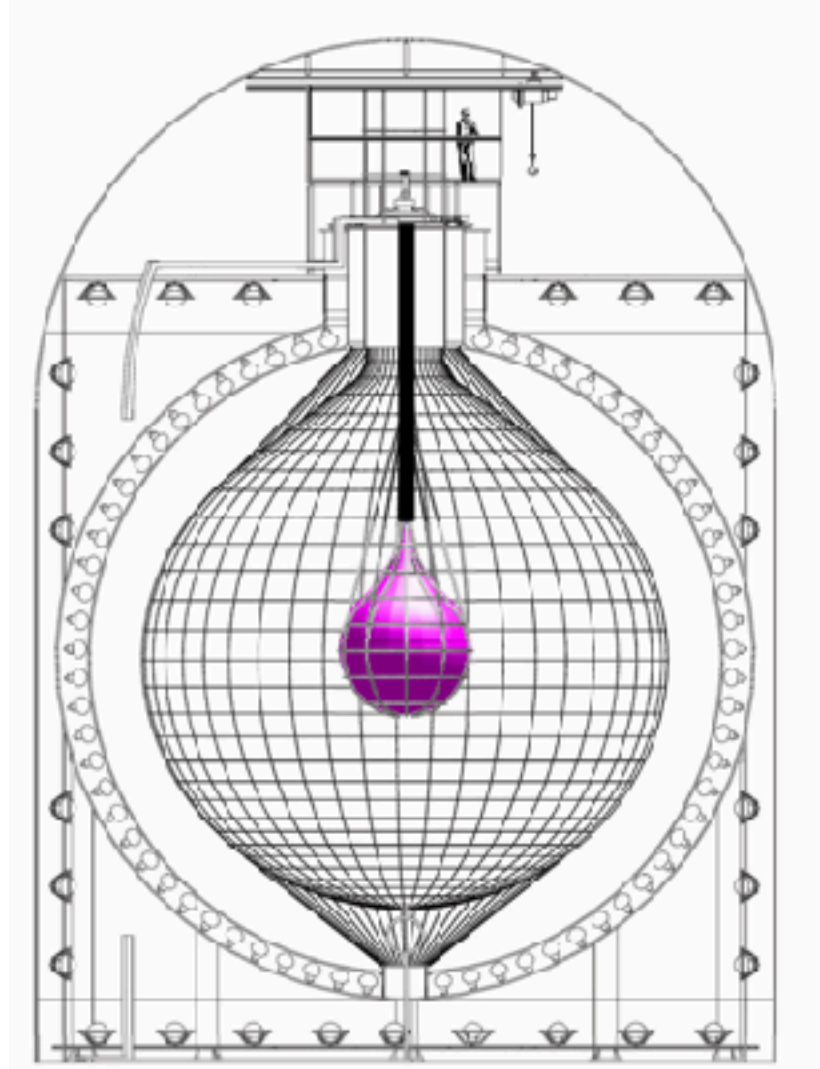
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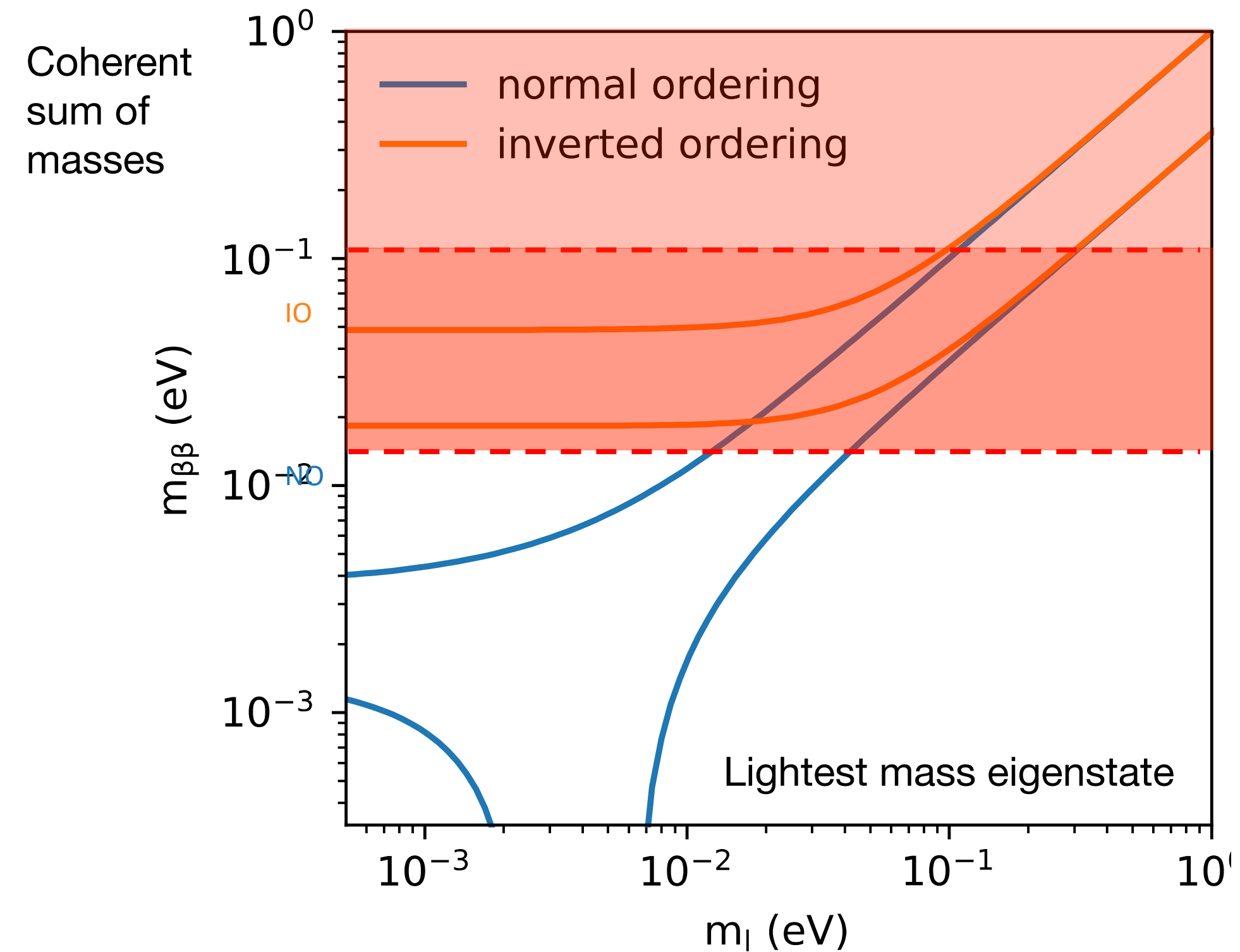
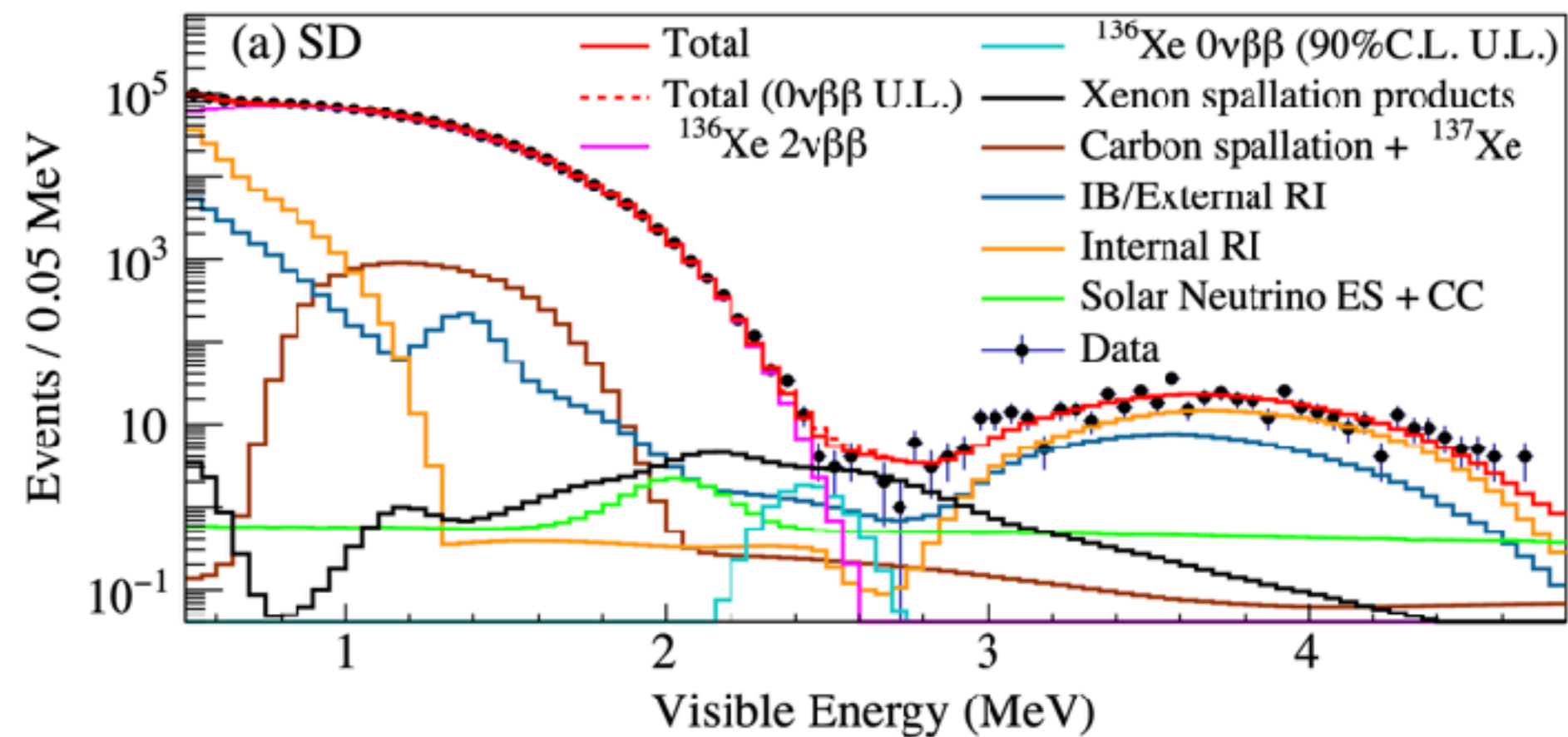
40

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 Data taking started in 2019



The largest number of  $\beta\beta$  nuclei. Low BG by distillation and filtration of both Liquid Scintillator and Xenon.



Next Generation experiments Legend-1000, KamLAND-ZEN-1T and nEXO will cover the inverted ordering!

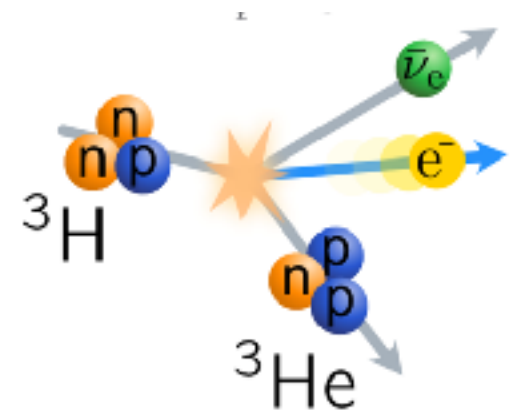
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# Absolute Neutrino Mass

Susanne Mertens

## KATRIN Experiment



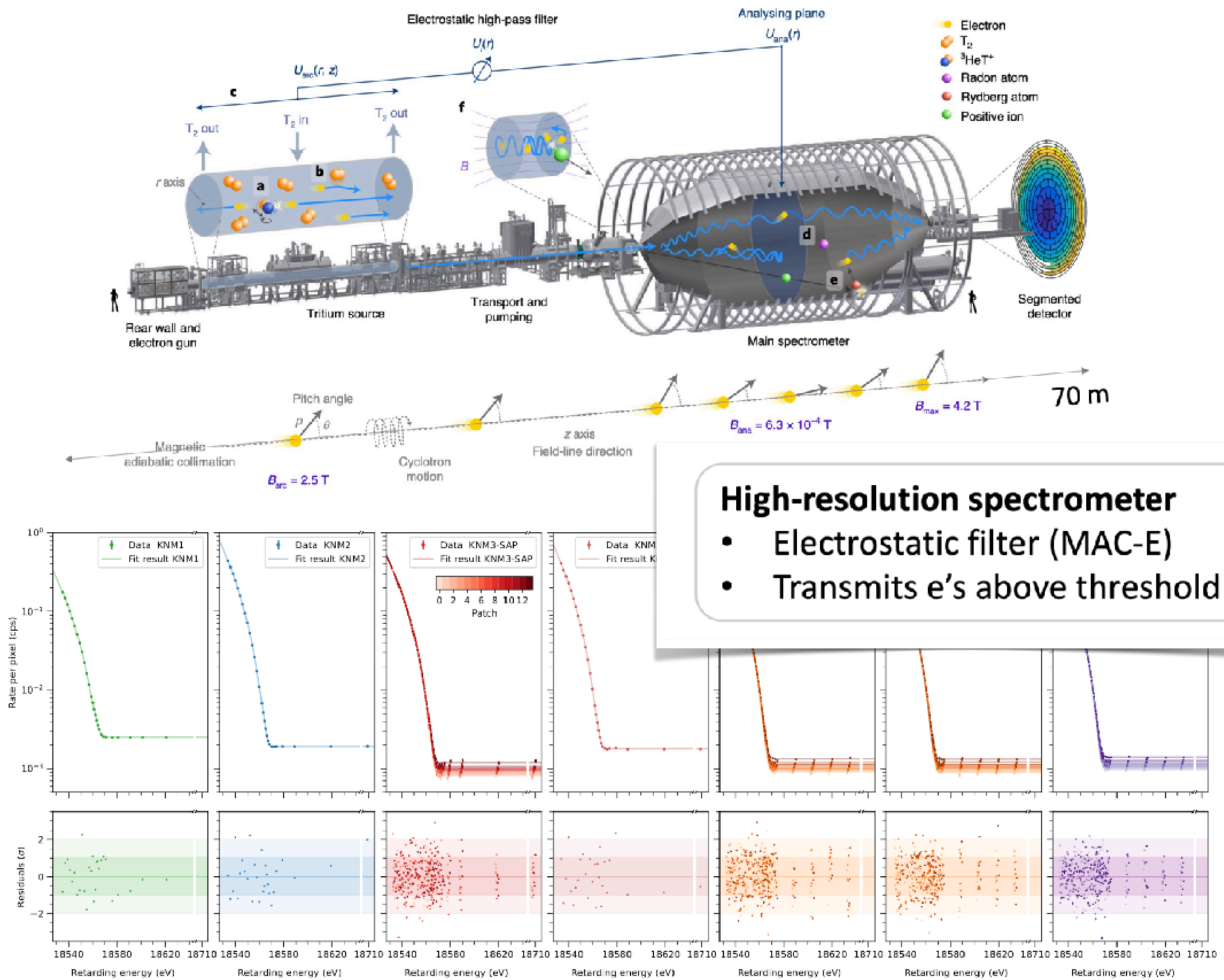
**High-activity tritium source**

- 30  $\mu\text{g}$  of gaseous  $\text{T}_2$
- $10^{11}$   $\text{T}_2$  decays/s

- Best fit:  
 $m_\nu^2 = (-0.14^{+0.13}_{-0.15}) \text{eV}^2$  (stat. dom.)
- New limit:  
 $m_\nu < 0.45 \text{ eV}$  (90% CL)  
 Neutrino-24 (2024), arXiv:2406.13516 (2024)

### Final goal (in 2026):

- $< 0.3 \text{ eV}$  sensitivity

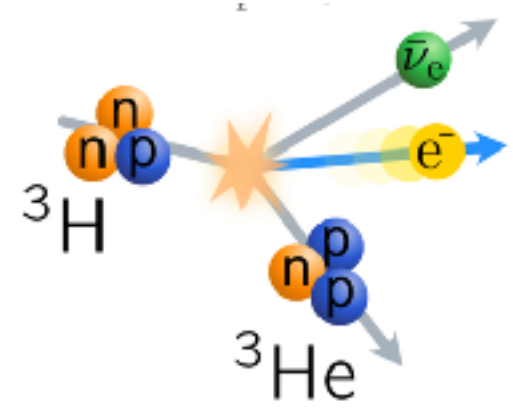


**High-resolution spectrometer**

- Electrostatic filter (MAC-E)
- Transmits e's above threshold

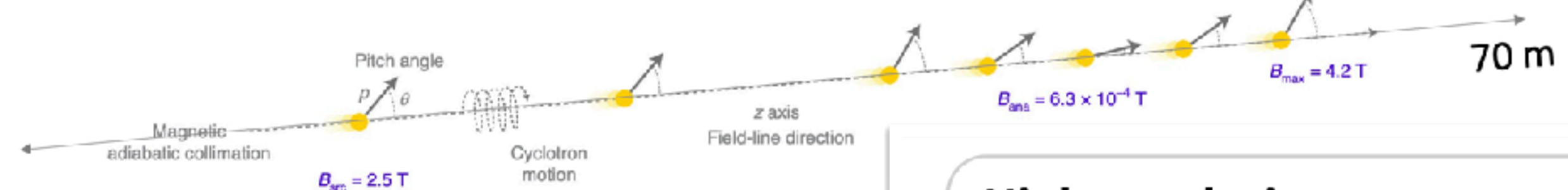
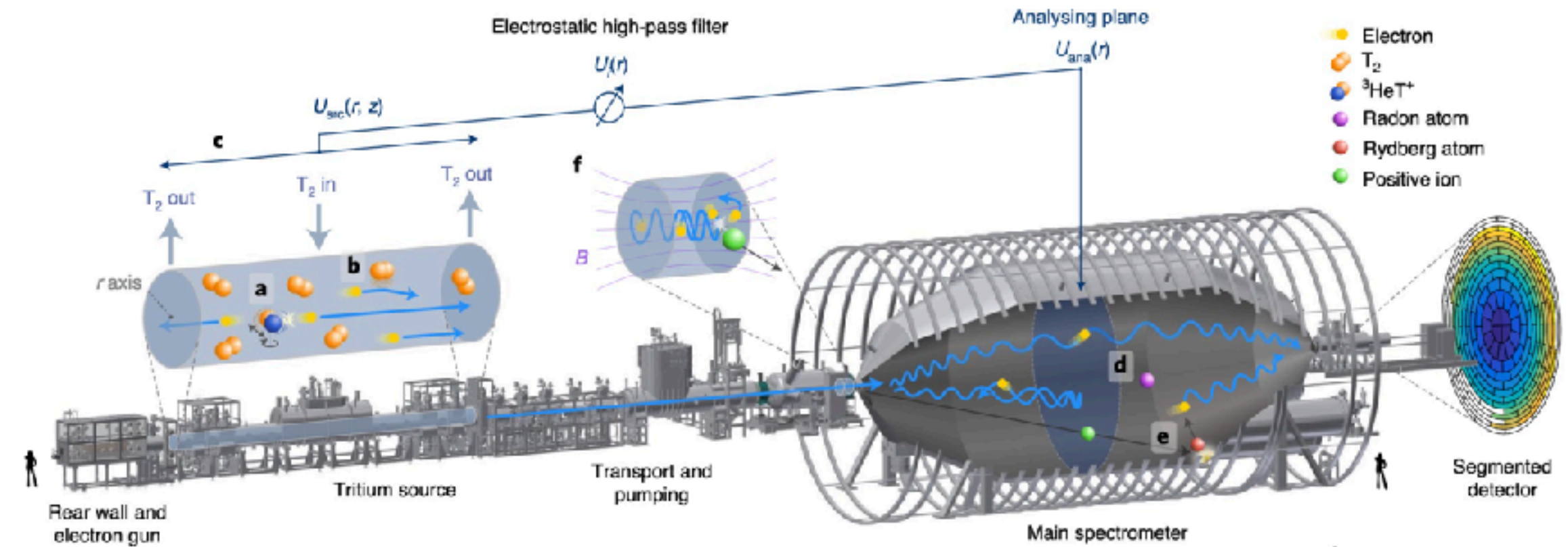
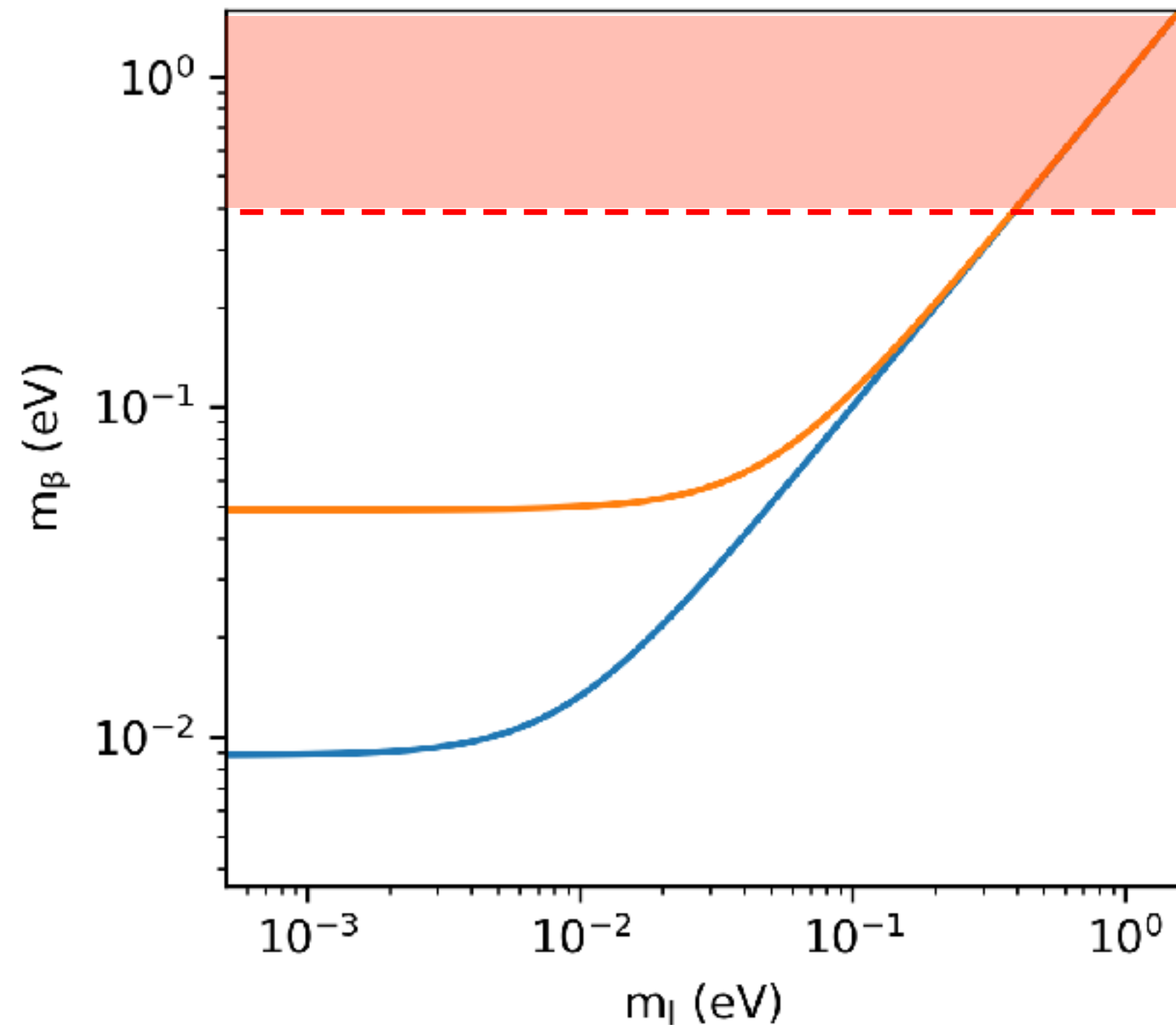
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## KATRIN Experiment



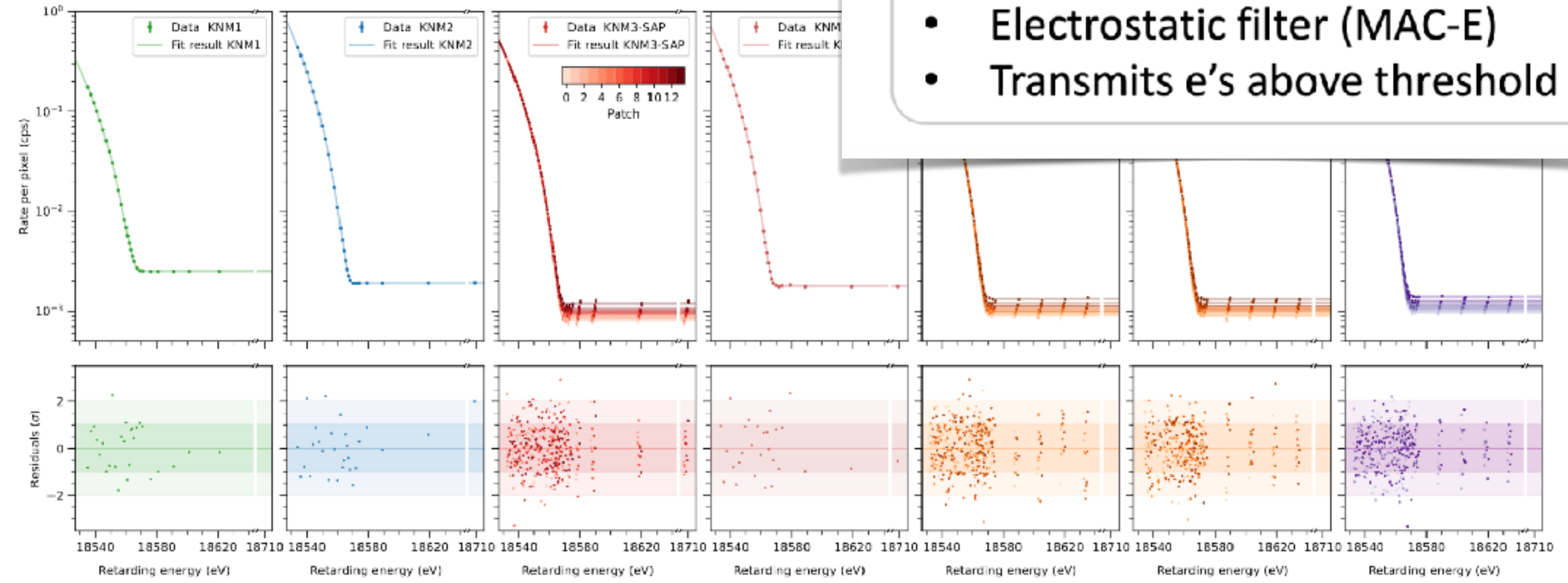
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- $10^{11}$   $\text{T}_2$  decays/s



**High-resolution spectrometer**

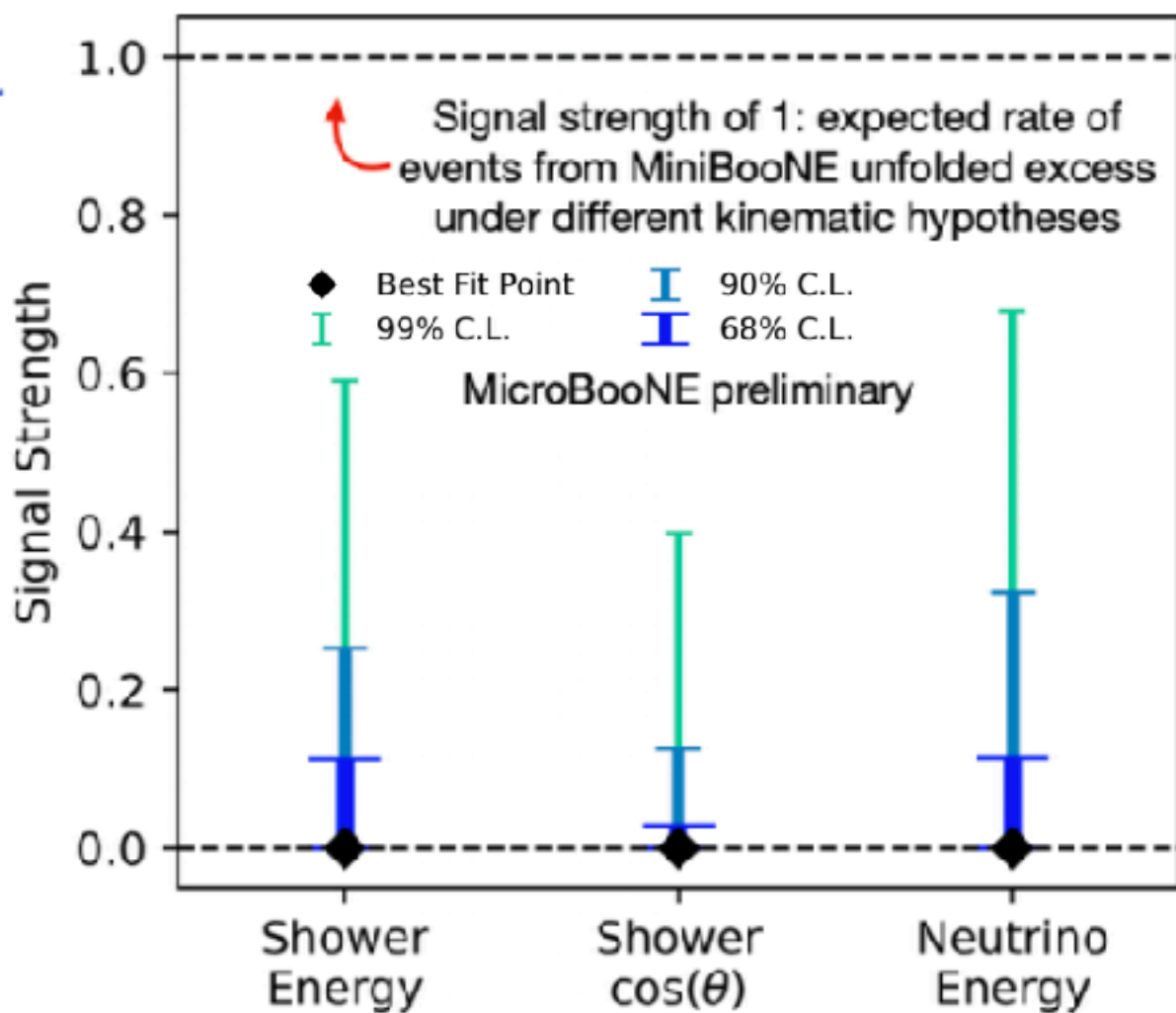
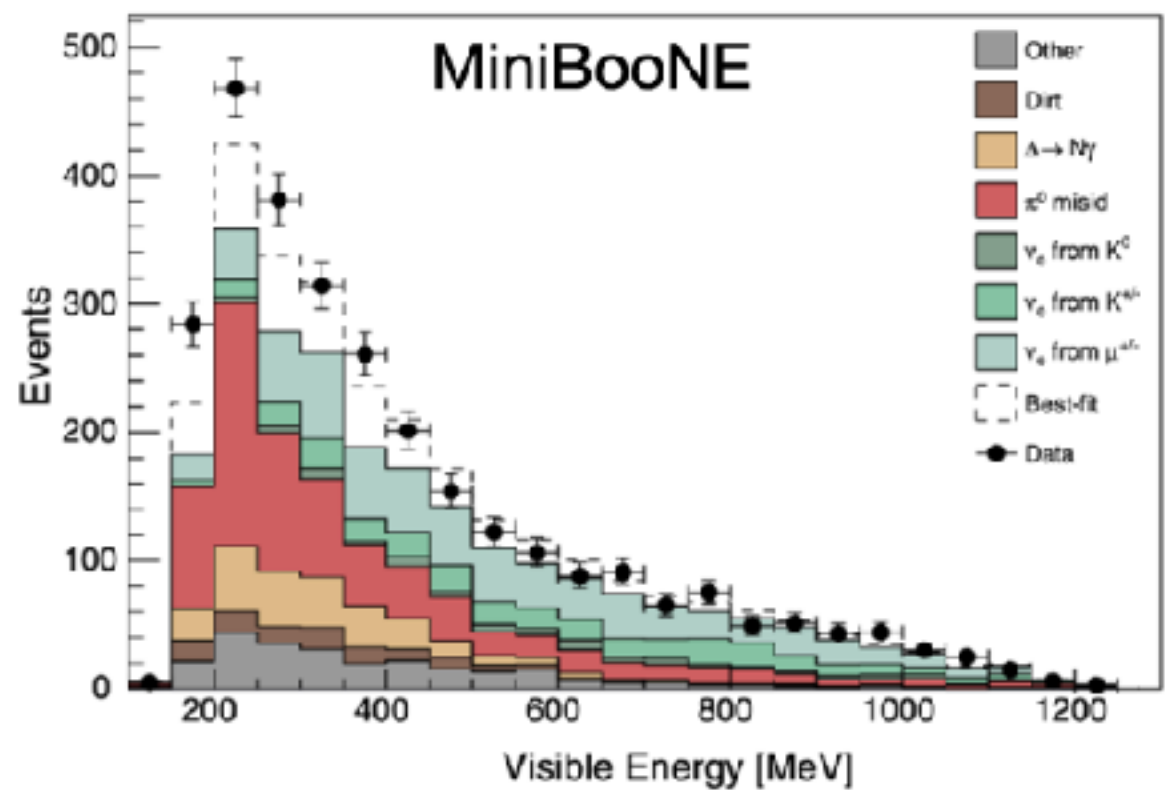
- Electrostatic filter (MAC-E)
- Transmits e's above threshold



# Neutrino Anomalies

## LSND/MiniBooNE

LSND/MiniBooNE observed  $6\sigma$  excess of electron (anti)neutrinos in muon (anti)neutrino beam!

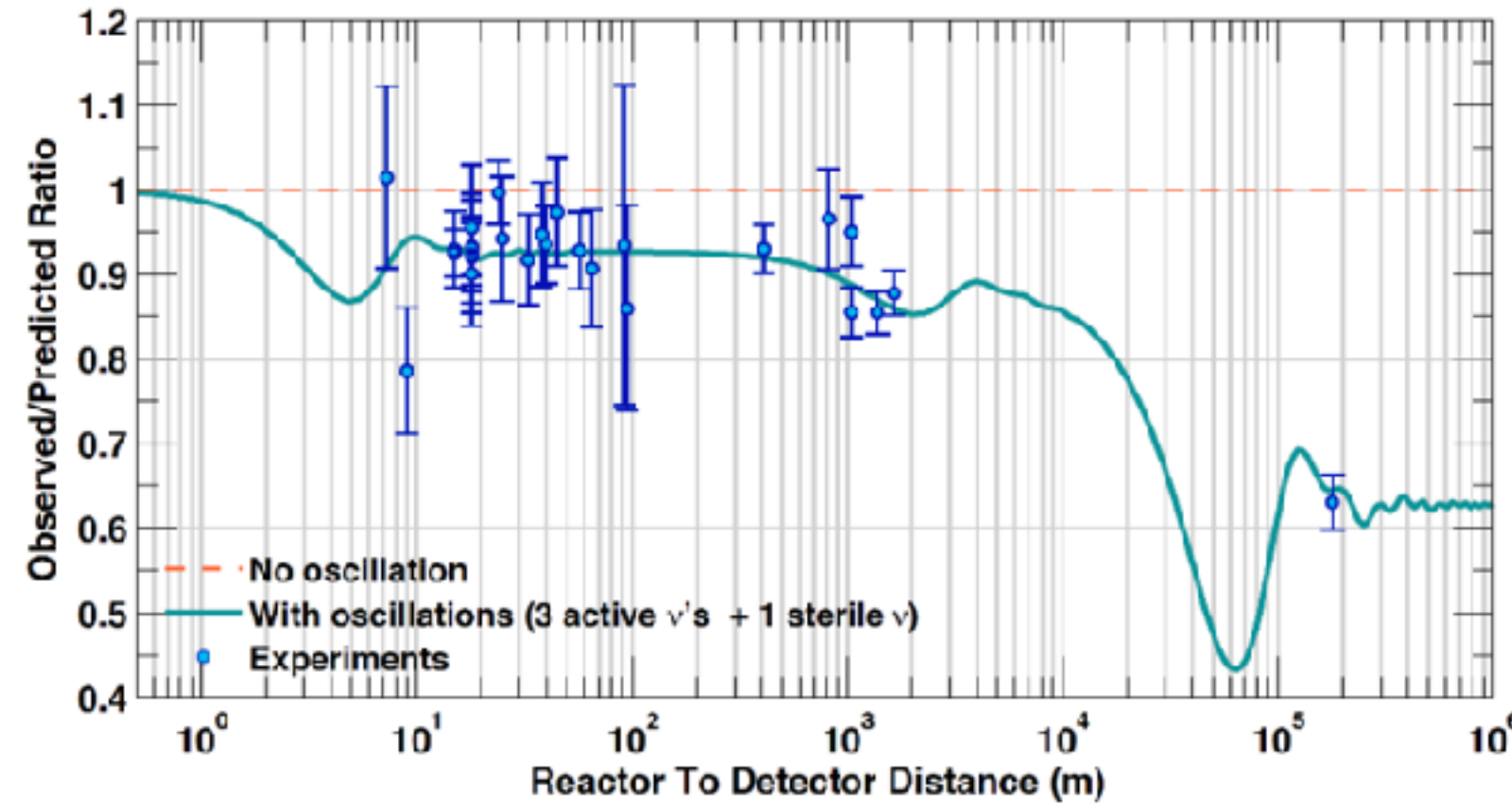


Excess not seen by MicroBooNE experiment

**SBN program** at Fermilab to completely settle the question!

## Reactor anti-neutrino deficit

6% deficit of reactor anti-neutrinos  $\sim 3\sigma$  with respect to flux prediction models.

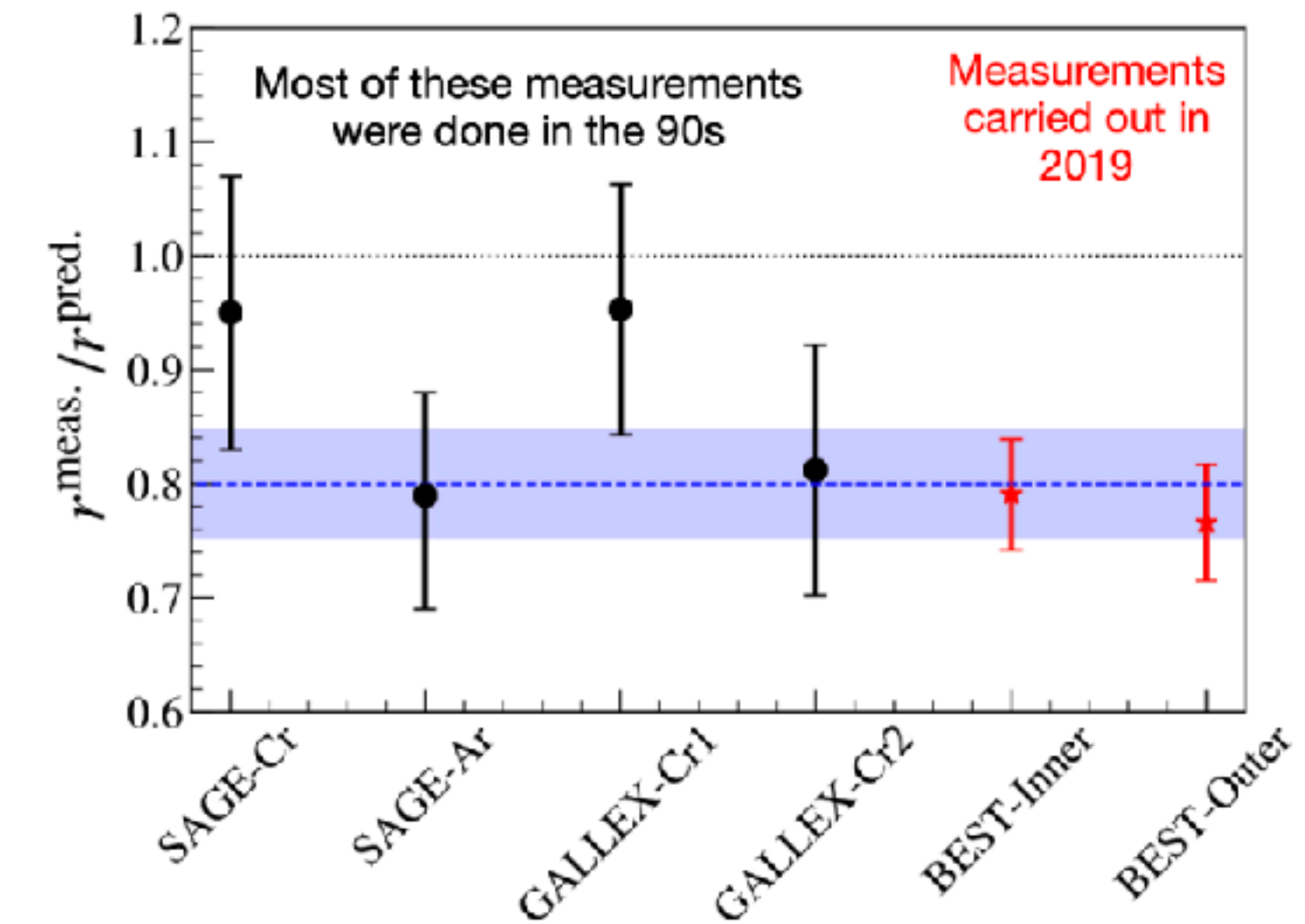


Recent fission data suggests that the flux were overestimated... **by about the right amount!**

Also essentially excluded by short baseline experiments searching for sterile neutrinos: DANSS, NEOS, PROSPECT, SoLid, STEREO.

## Gallium anomaly

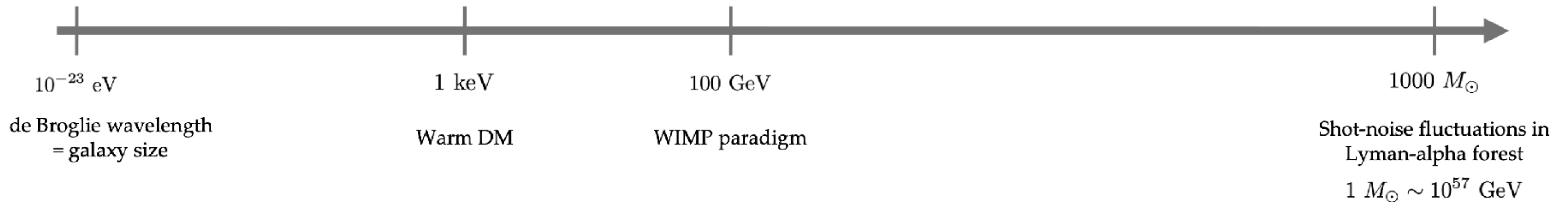
Seen by several experiments (including recent)



Also excluded by short baseline experiments.

# Dark Matter and a Possible Dark Sectors

The Range of Possibilities is **Stunning!**



## A clear roadmap!

Challenge #1: Fully Cover Electroweakino DM, e.g. with Cherenkov Telescopes

Challenge #2: Search for WIMPs to the Neutrino Background in Direct Detection

Challenge #3: Build out the suite of axion searches

Challenge #4: Build out the suite of accelerator searches—**high energy and intensity**—for hidden sectors

Challenge #5: Cover the abundance-driven light DM models in direct detection

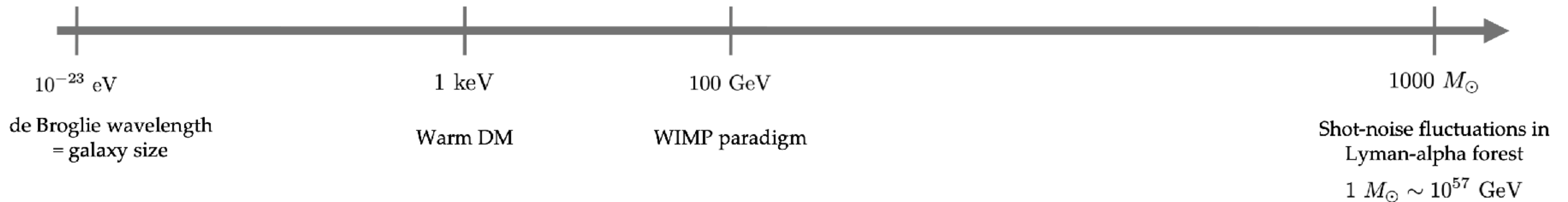
Challenge #6: Observe the Dark Matter Power on Small Scales

...

**Jocelyn and Stefania have covered (#2 - #5) !**

# The Road Map

The Range of Possibilities is **Stunning!**



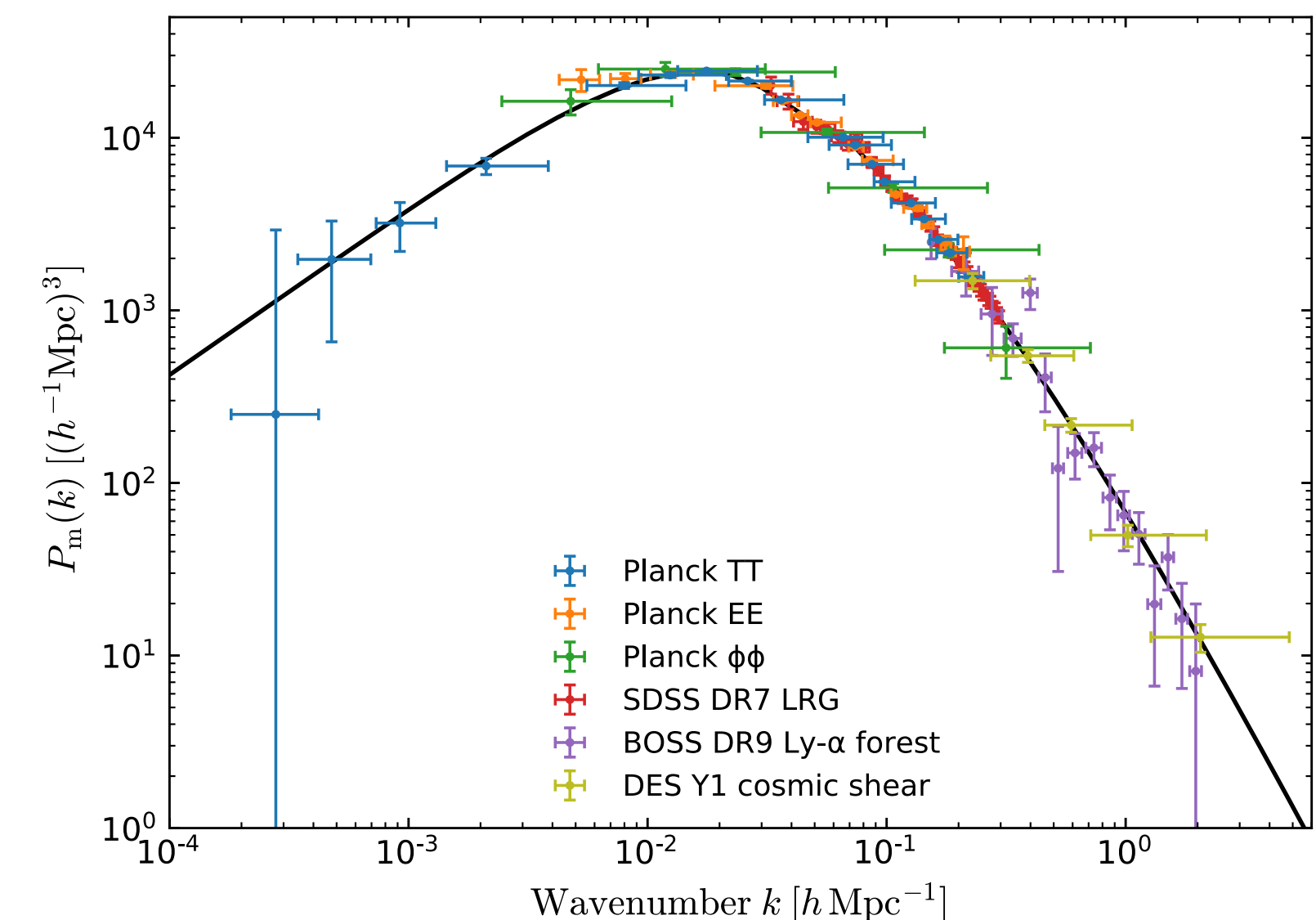
Challenge #1: Fully Cover Electroweakino DM, e.g. with Cherenkov Telescopes

- LHC is very sensitive to strongly produced SUSY particles, less so for pure electroweakinos
- Cherenkov telescopes can cover to larger masses!

Challenge #6: Observe the Dark Matter Power on Small Scales

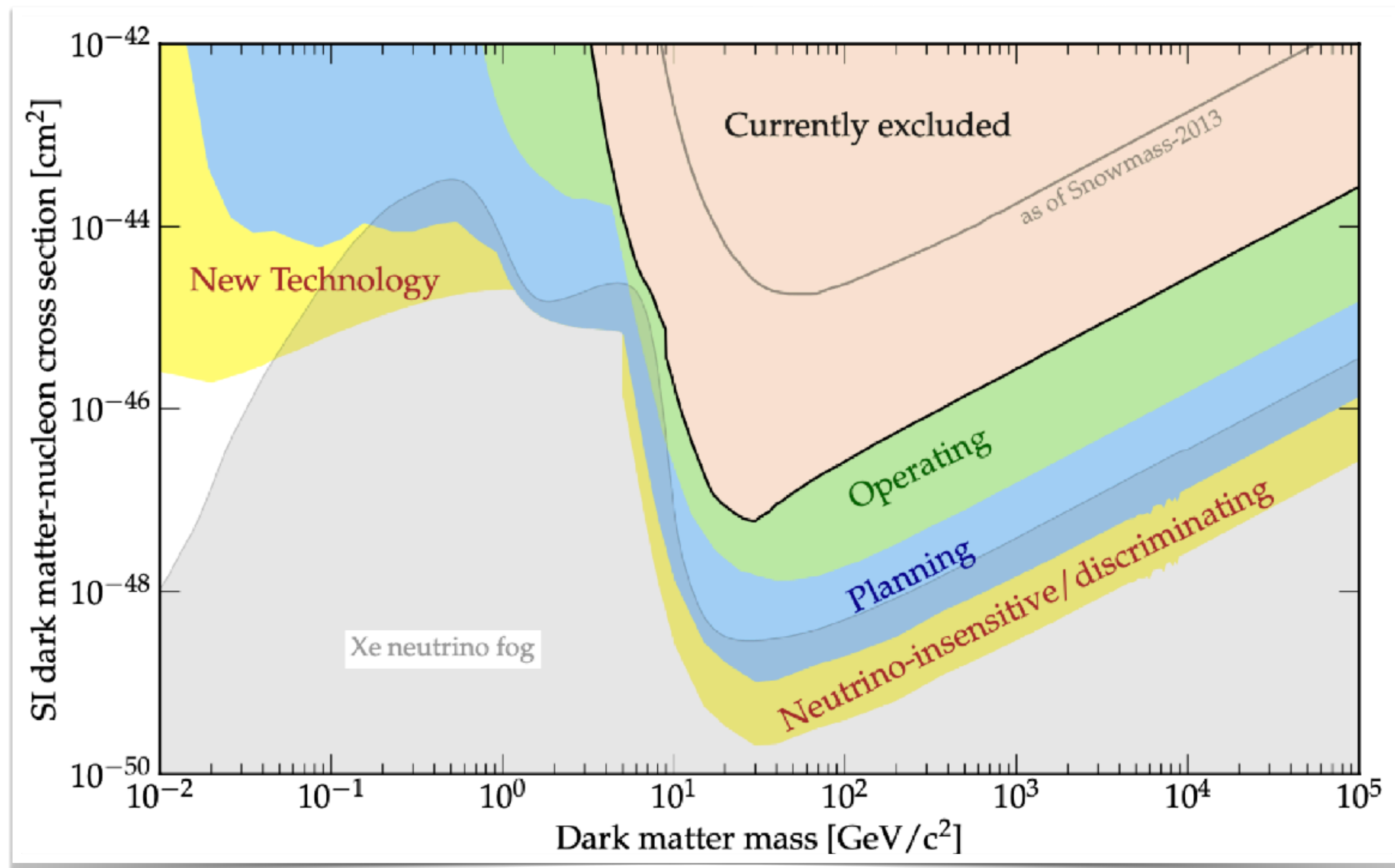
Theories as **PBH, axions,** predict different behavior than  $\Lambda_{\text{CDM}}$  at sub-halo scales, requiring new measurements of small scale!

(e.g. measuring the changes in the metric due to transiting DM substructure using Pulsar Timing Arrays)



# Reaching the Neutrino Fog

## Challenge #2: Search for WIMPs to the Neutrino Background in Direct Detection



### - Currently operating detectors:

- Xenon nT, PandaX-4T, LZ-7T
- Argon DarkSide

### - Next generation:

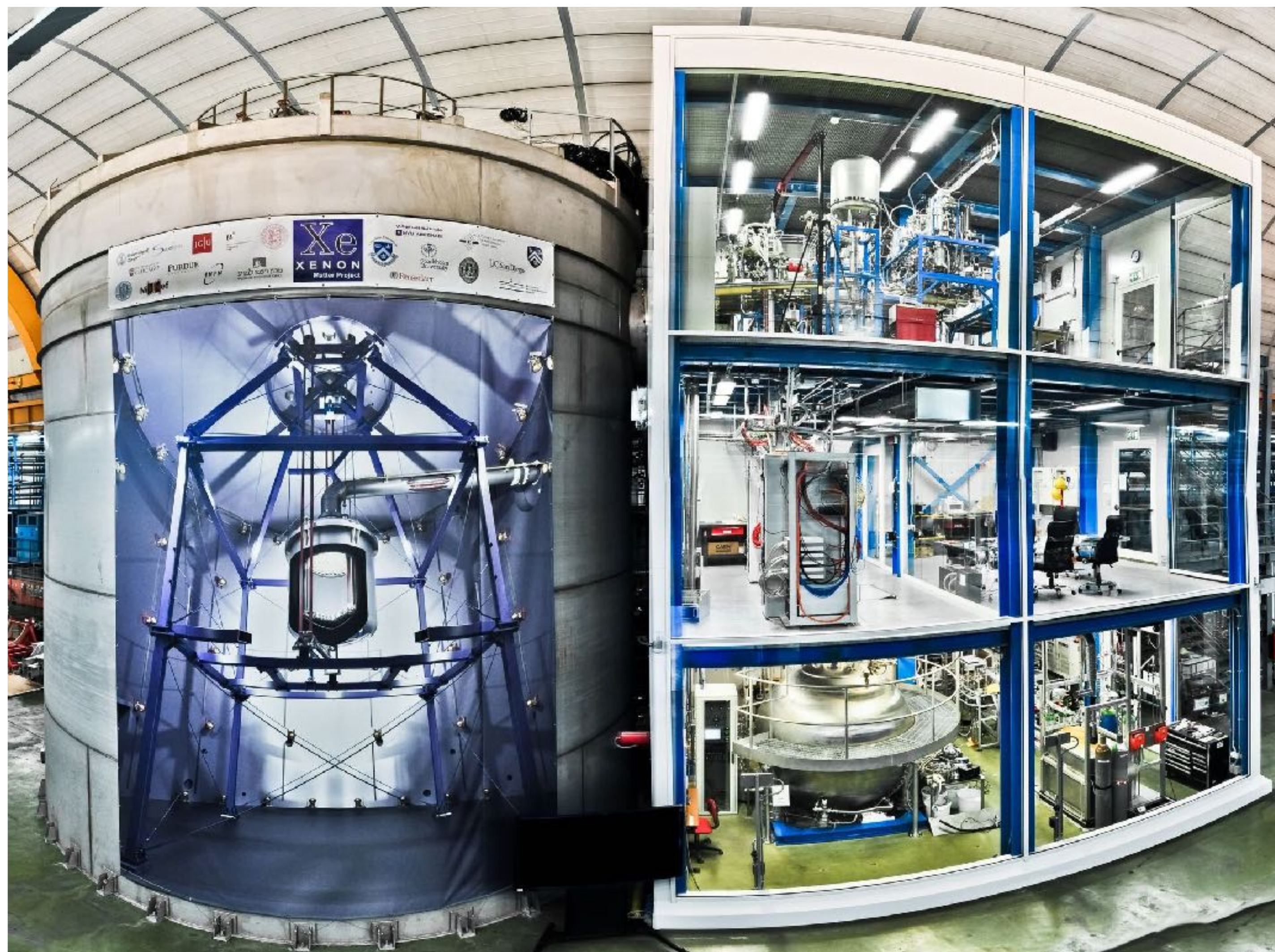
- PandaX-20T
- Darwin XLZD 40-60T

# Dual Phase Xenon TPC Experiments

Jocelyn Monroe, Di Gangi

47

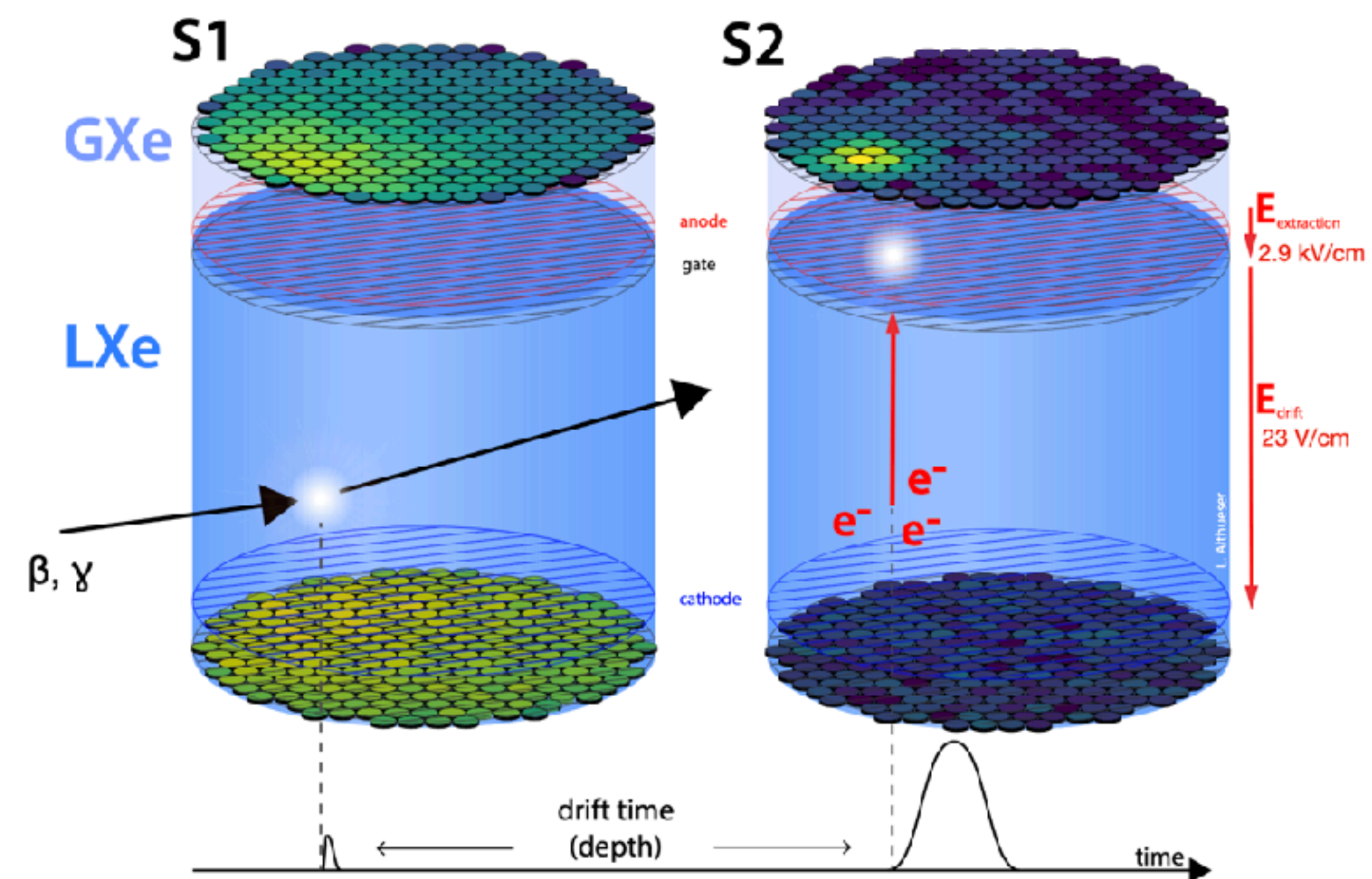
Challenge #2: Search for WIMPs to the Neutrino Background in Direct Detection



## Xenon nT

Liquid Xenon Scintillation

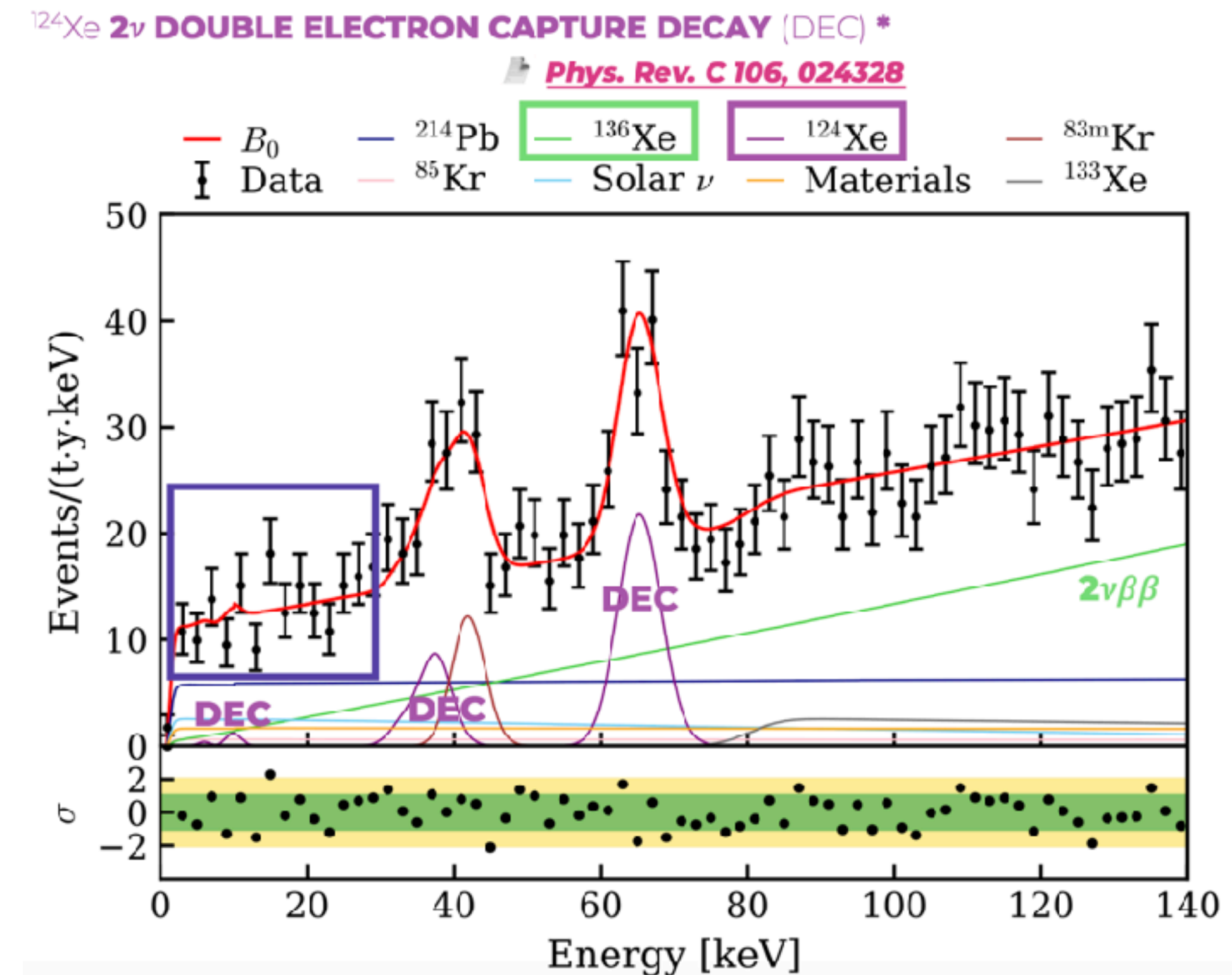
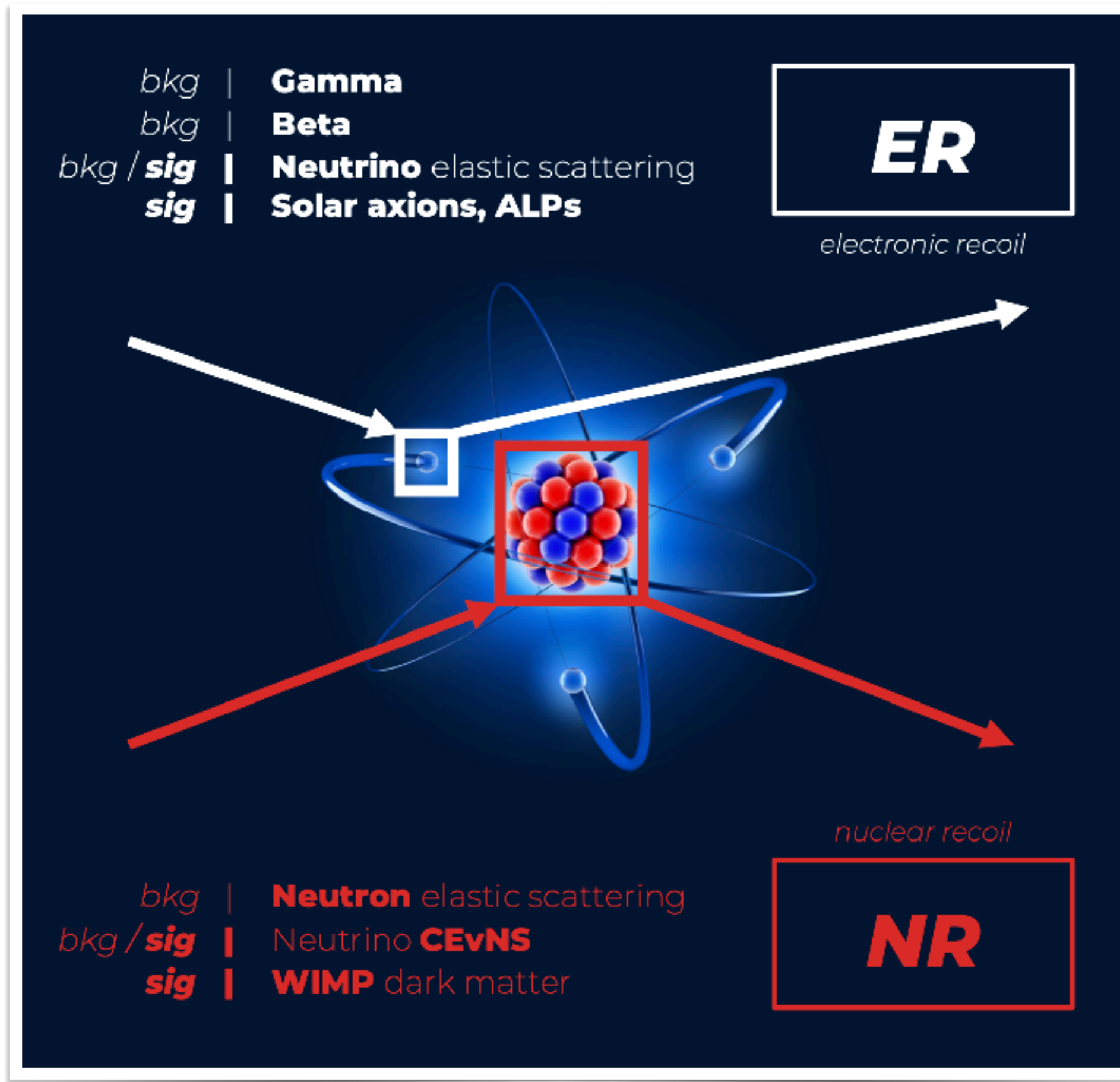
Ionisation channel (Gaseous Xe scintillation)



# Electron Recoils

## Challenge #2: Search for WIMPs to the Neutrino Background in Direct Detection

## Electron Recoil in Xenon

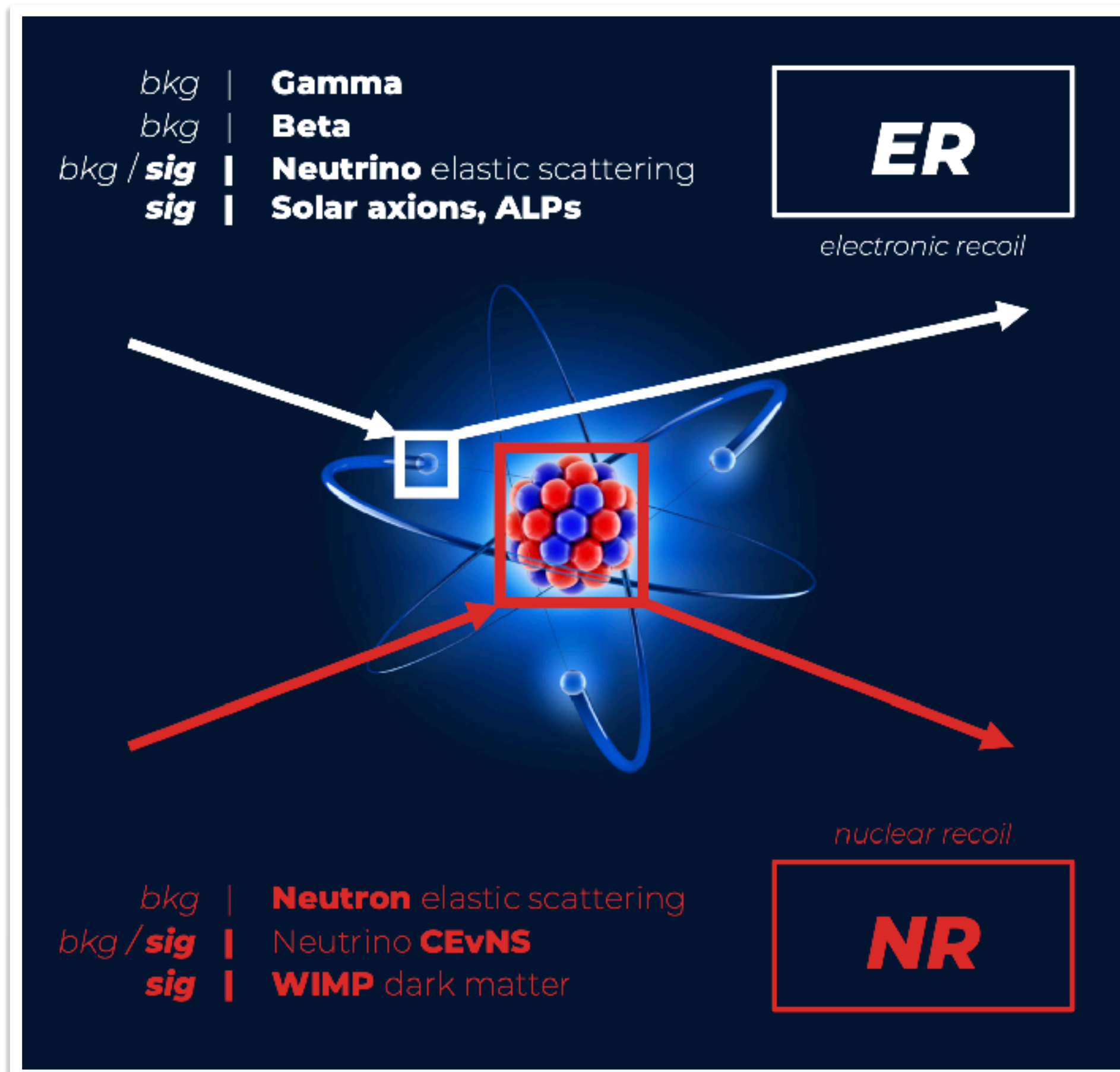


- Observation of DEC: Two of the orbital **electrons** are captured **protons** in the nucleus emission of 2 neutrinos
- Solar neutrinos background subdominant (backup PandaX-4T)

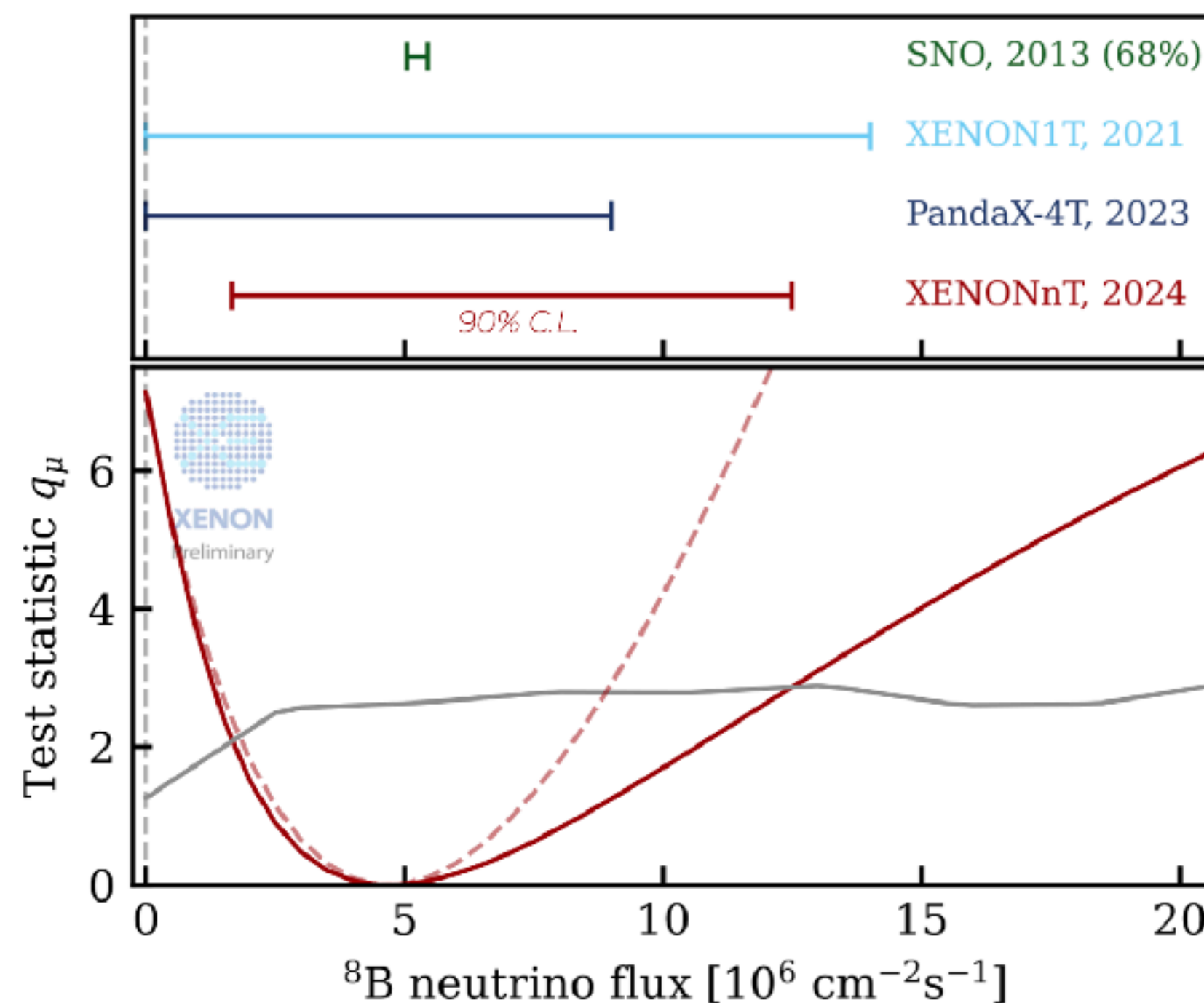


# Reaching the Neutrino Fog

## Challenge #2: Search for WIMPs to the Neutrino Background in Direct Detection



## Nuclear Recoils from $^8B$ solar neutrinos **CEvNS** in Xenon!



**2.73  $\sigma$**   
**SIGNIFICANCE**

**37 Events observed** with 26 bkg and 11 signal expected!

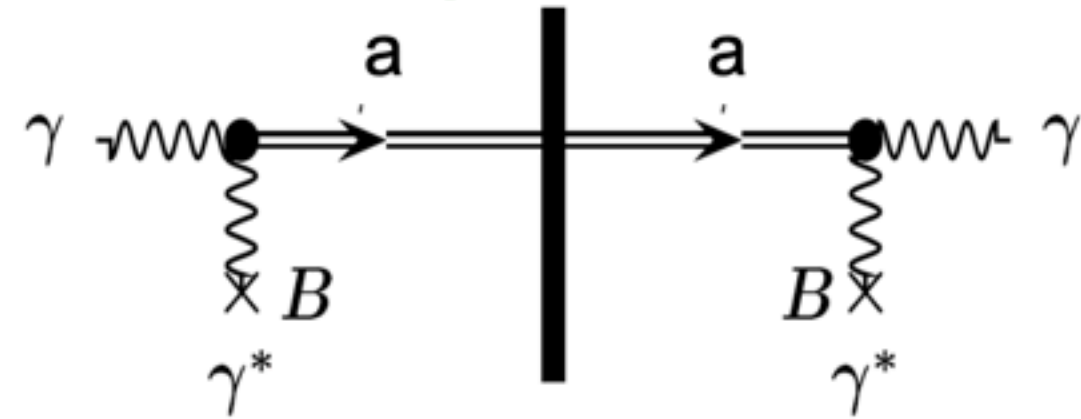
**Great achievement, but now need to learn how to fight it!**

# Axions an Ambitious Program

Challenge #3: Build out the suite of axion searches

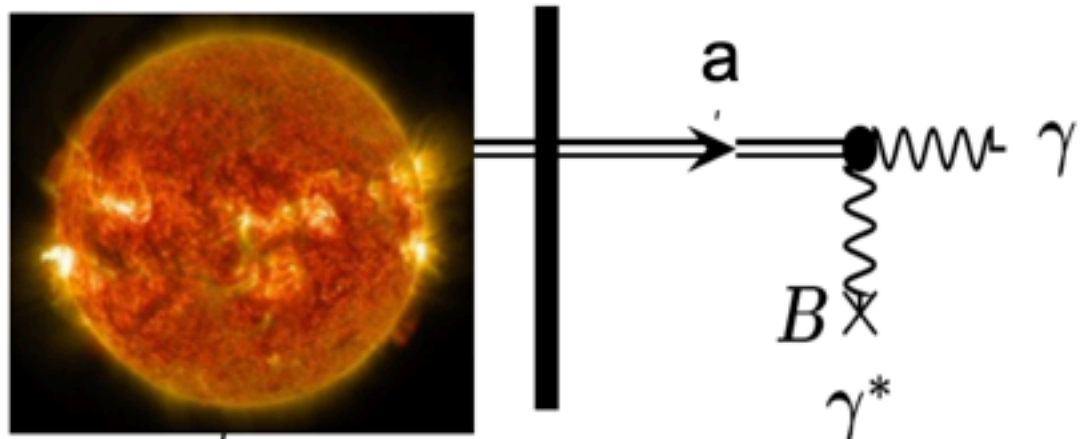
Light shining through a wall (LSW)

[Anselm 85; van Bibber 87]



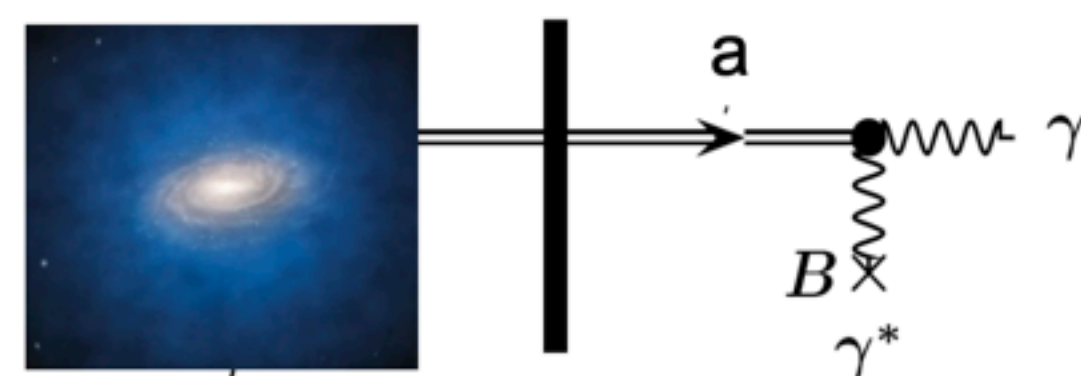
Helioscope: Sun shining through a wall

[Sikivie 83]

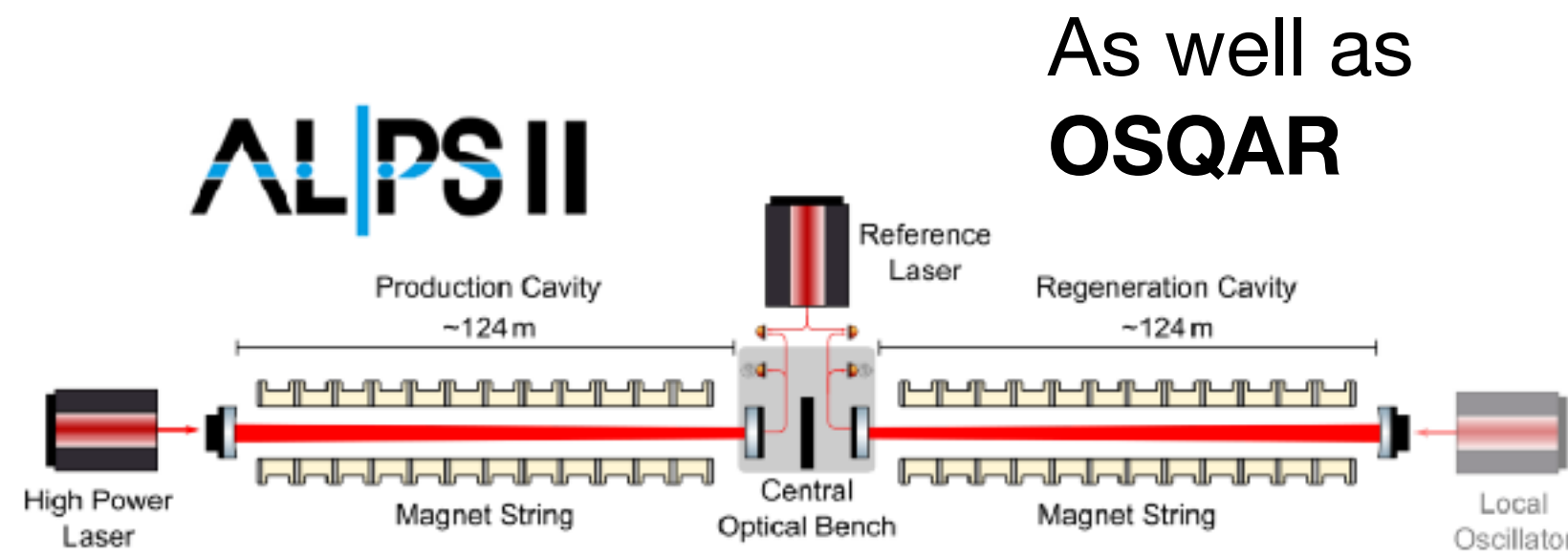


Haloscope: DM shining through a wall

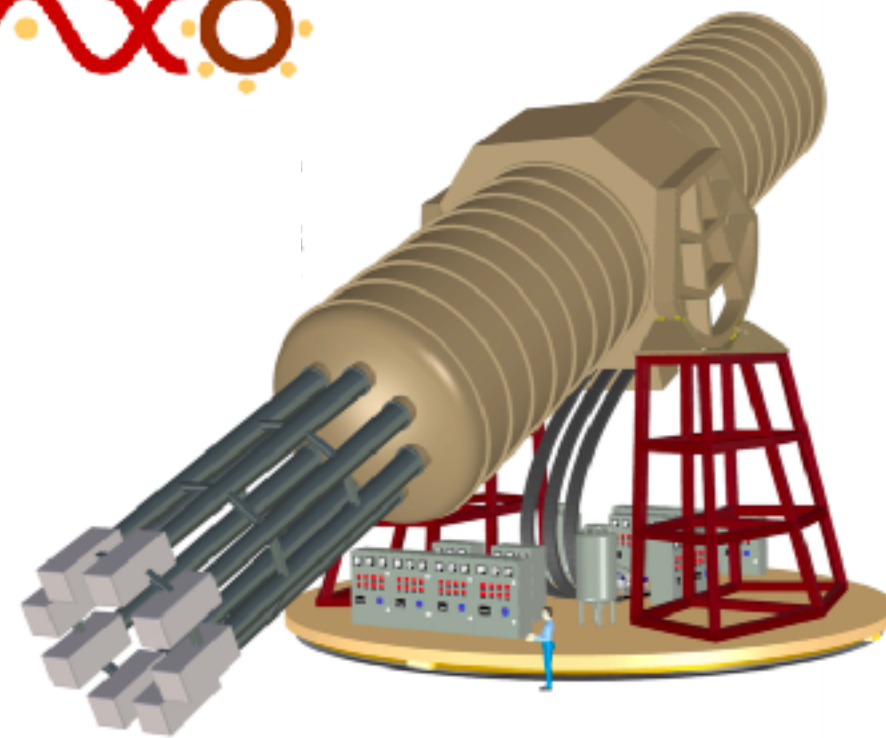
[Sikivie 83]



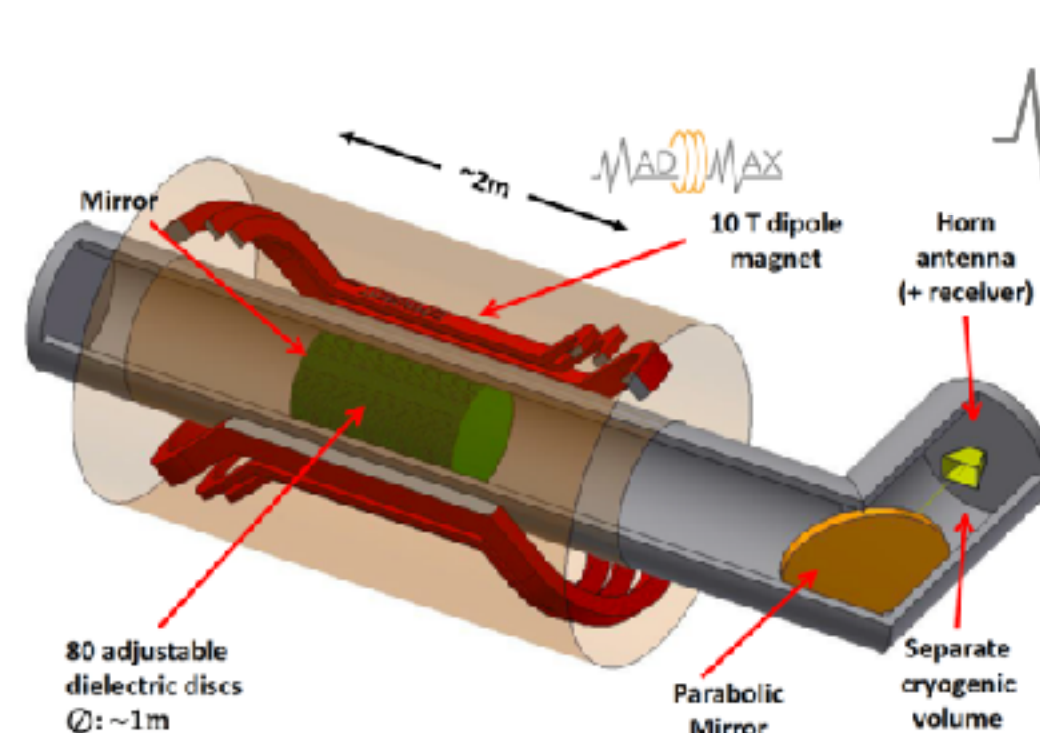
[Lindner]



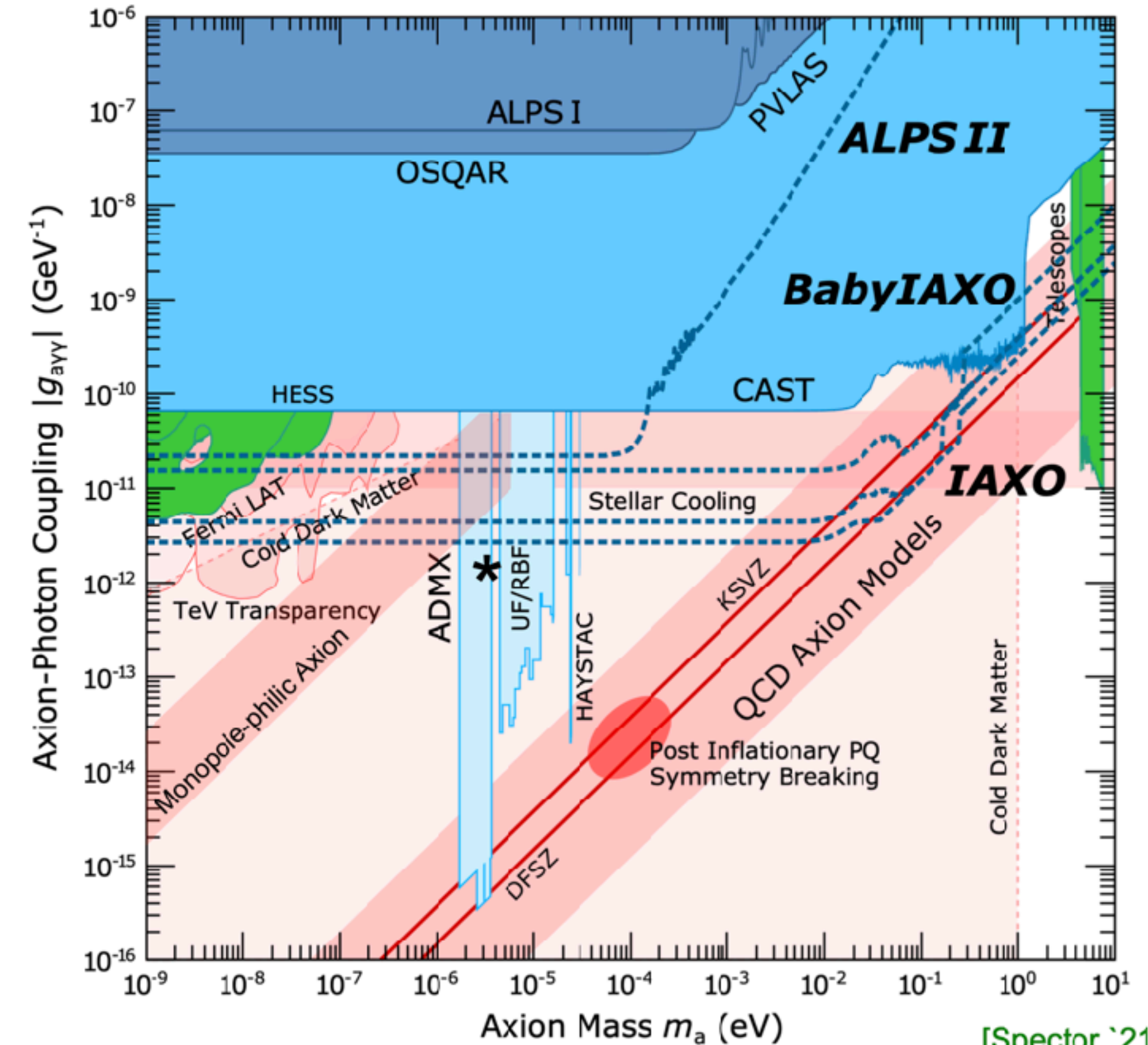
As well as  
**OSQAR**



As well as  
**SUMICO,**  
and **CAST**



As well as **ADMX,**  
**HAYSTAC, QUAX, CAST-**  
**CAPP, RADES**



[Spector '21]

# Dark Sector, Intensity Frontier and Meson Factories

Stefania Gori, Kathryn Zurek, Livia Soffi

51

Challenge #4: Build out the suite of accelerator searches—**high energy and intensity**—for hidden sectors

There are numerous gaps in the experimental exploration of dark sectors, both at the **electroweak-TeV** scale and at the **sub-GeV** scale.

NP particles can have a width that is suppressed by small mass splitting, multi-body final states, high NP scale small coupling, etc.

**Sub-GeV** new physics motivates searches at high intensity facilities and Meson factories!

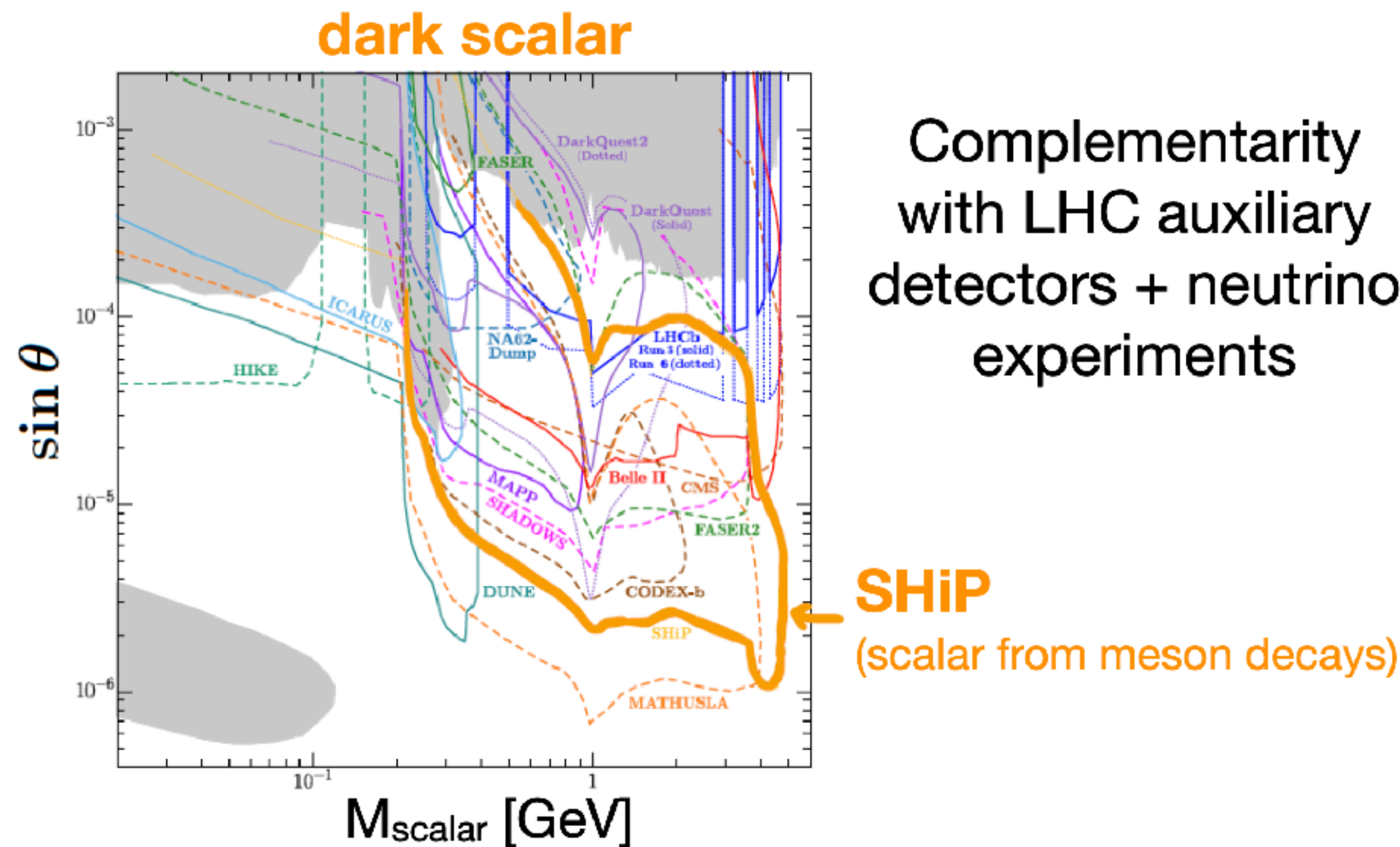
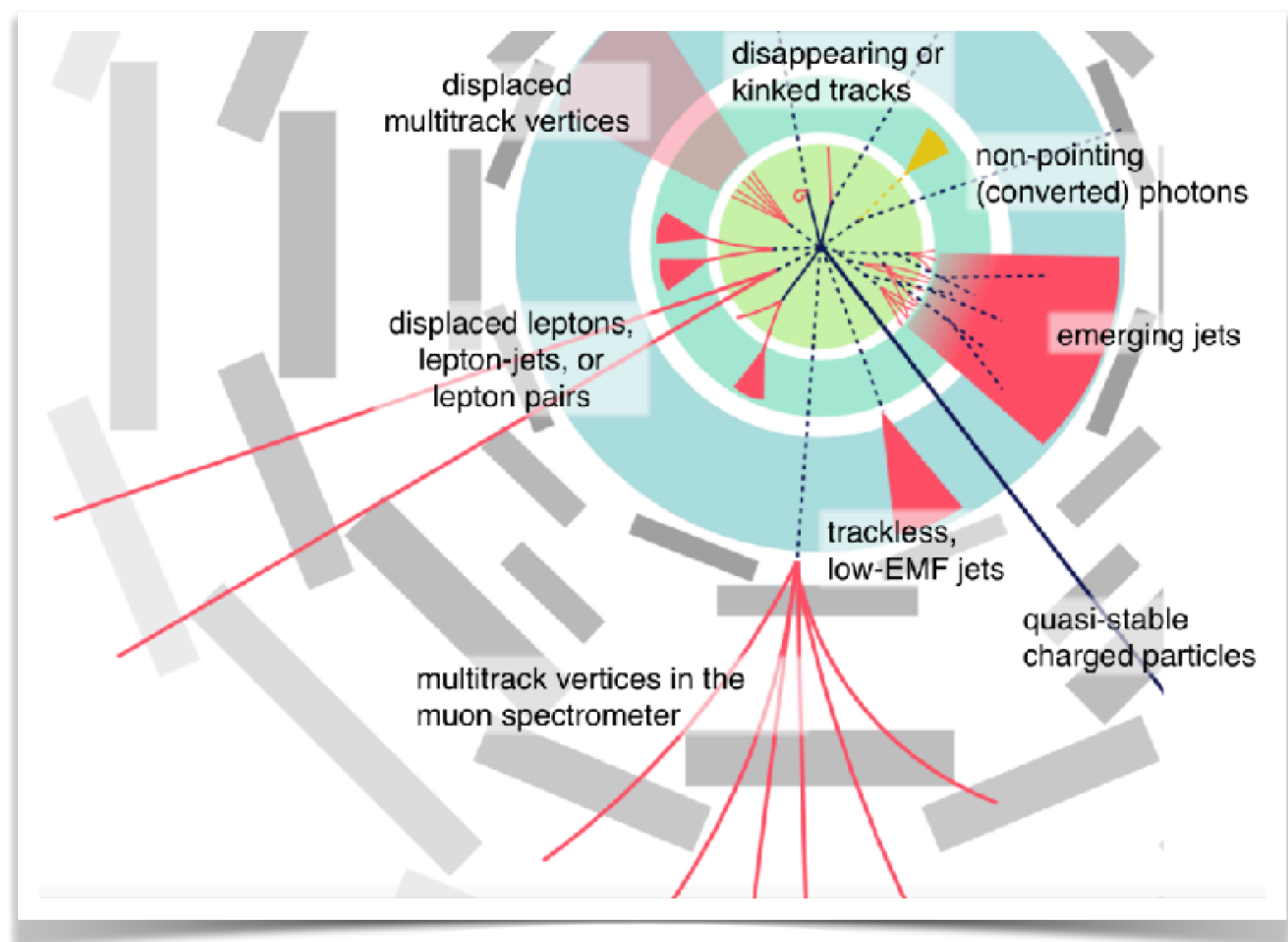


Image from H. Russel

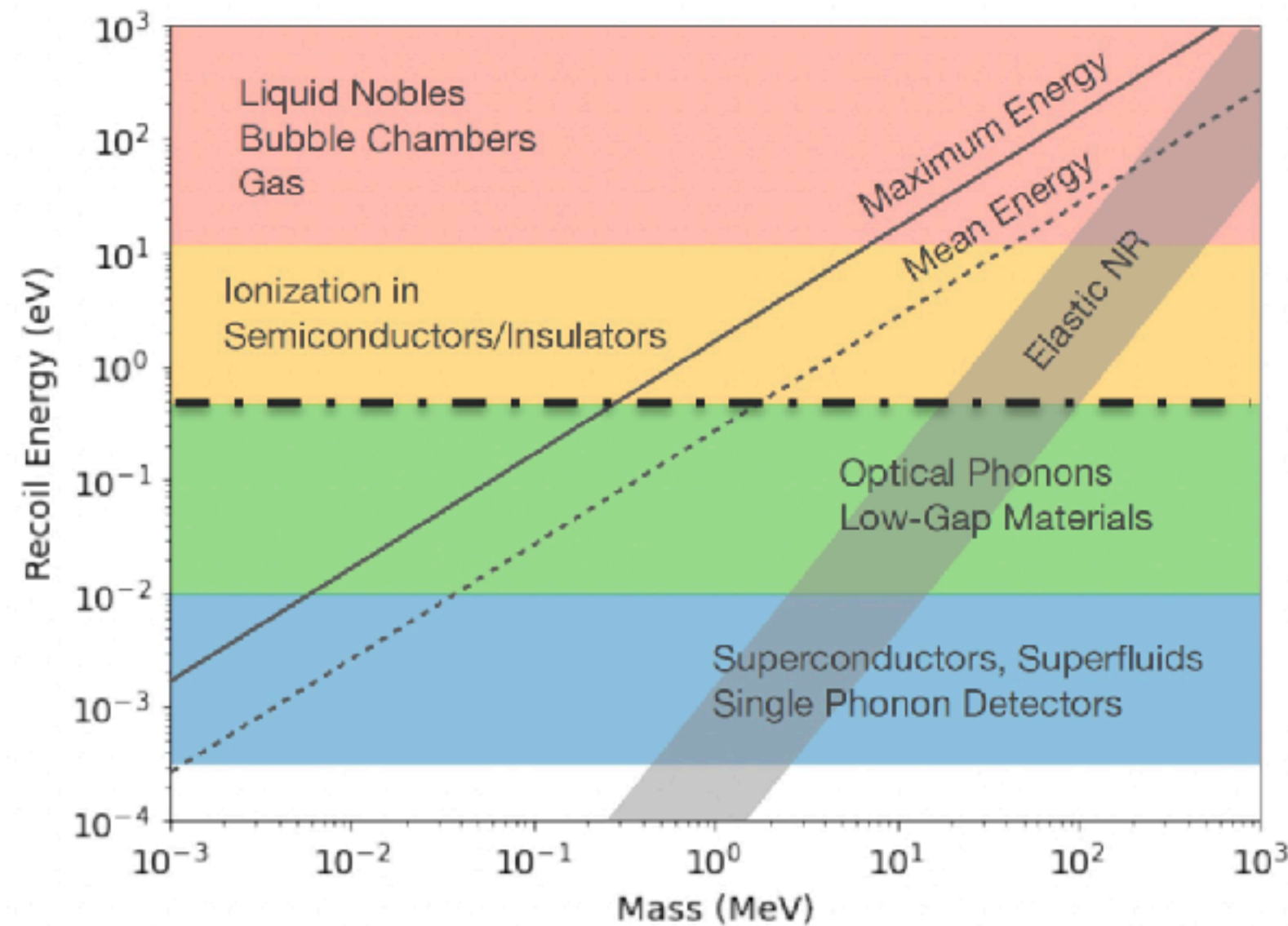
# Dark Sector, Intensity Frontier and Meson Factories

Stefania Gori, Kathryn Zurek, Jocelyn Monroe, Michael Doser

Challenge #5: Cover the abundance-driven light DM models in direct detection

Hidden sector DM: huge range of techniques... **Models motivate from relic abundance** can be tested w/1 kg-yr exposures!

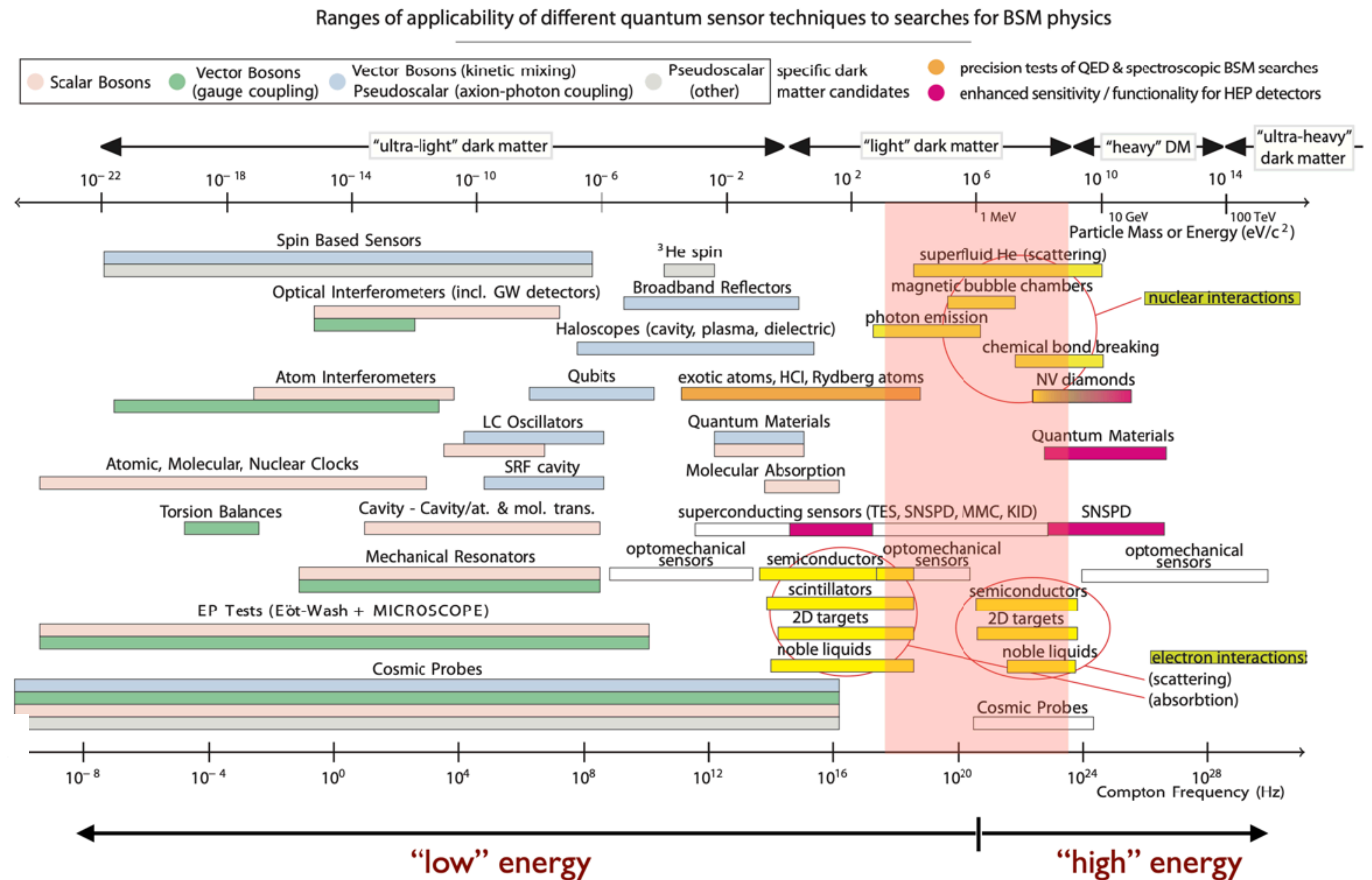
Quantum sensors can go to sub-eV recoil energies (a blossoming field!)



$\Delta E \sim 1 \text{ eV}$   
e.g. Si, Ge, GaAs, diamond,  
Quantum Dots, organic  
scintillators...

$\Delta E \sim 10 - 100 \text{ meV}$   
e.g. GaAs, sapphire, Dirac  
materials, doped s/c, ...

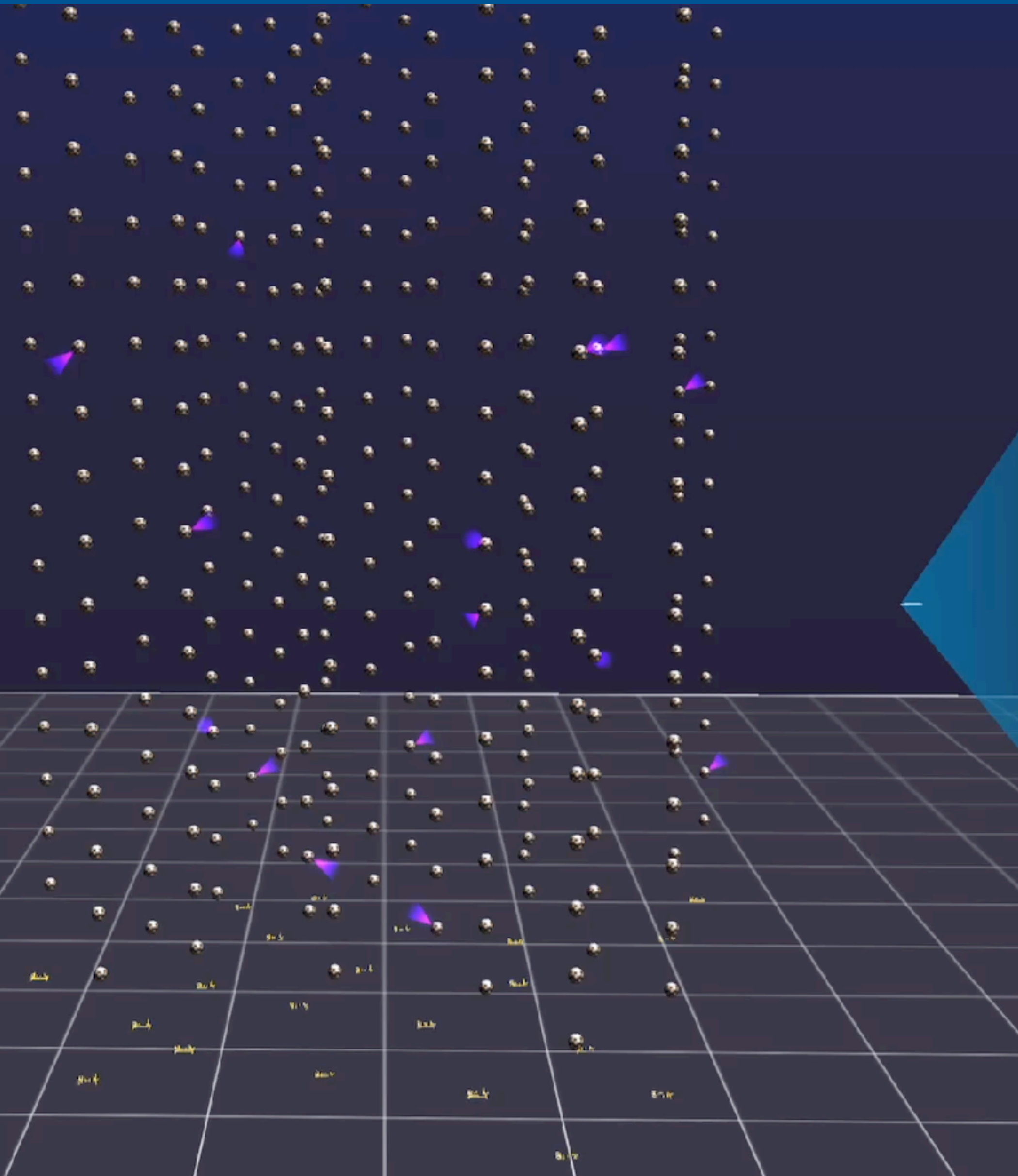
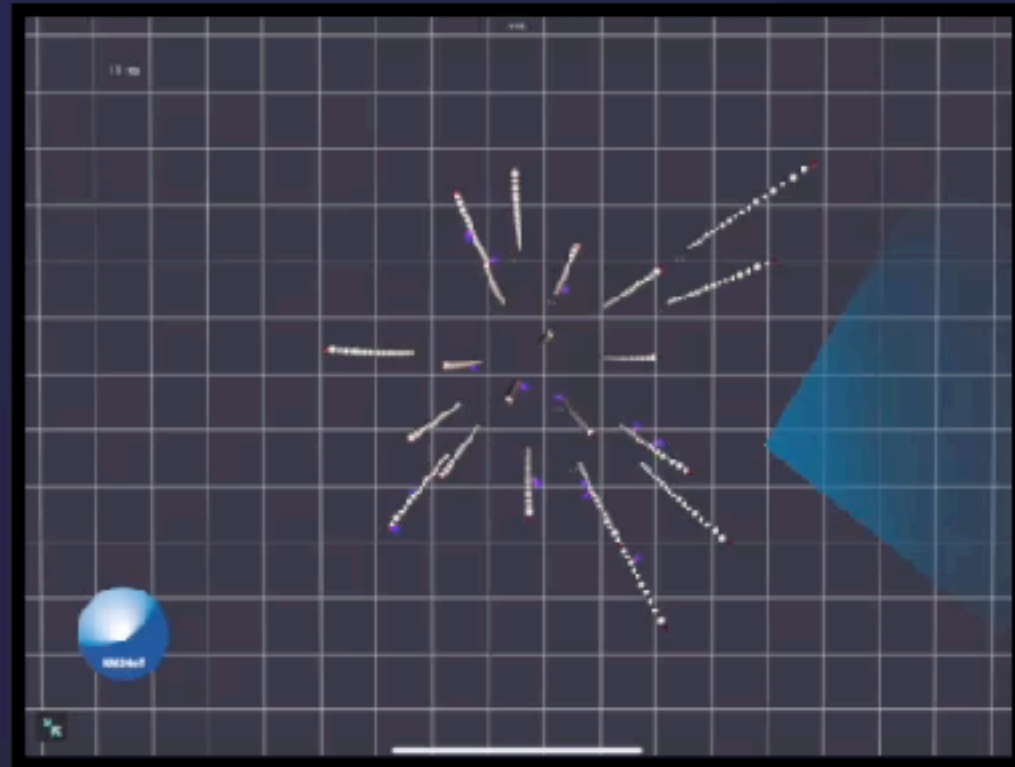
$\Delta E \sim 1 \text{ meV}$   
e.g. superfluids,  
superconductors



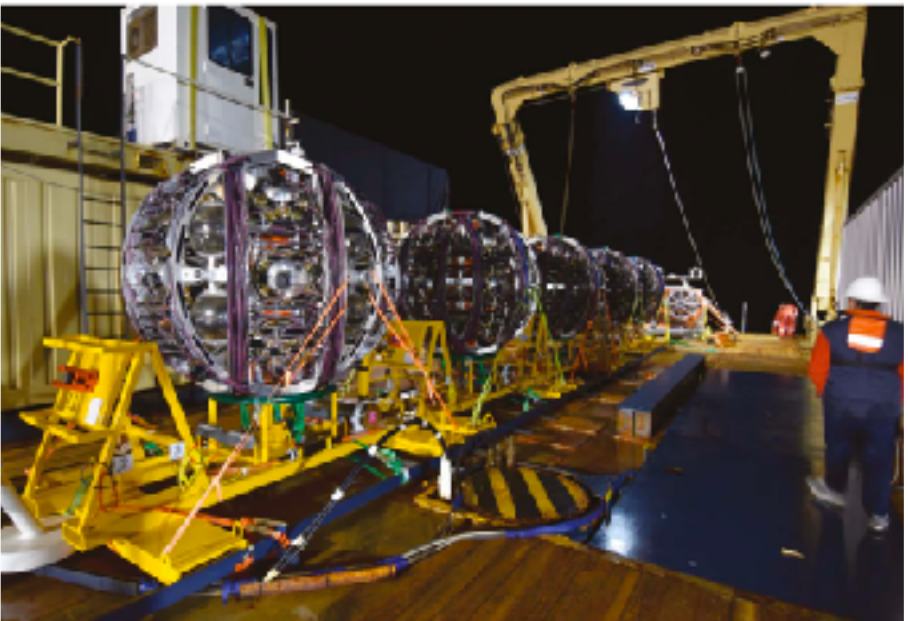
# Multi-Messenger Astrophysics

Paschal Coyle, Vitor de Souza, Lu Lu

53



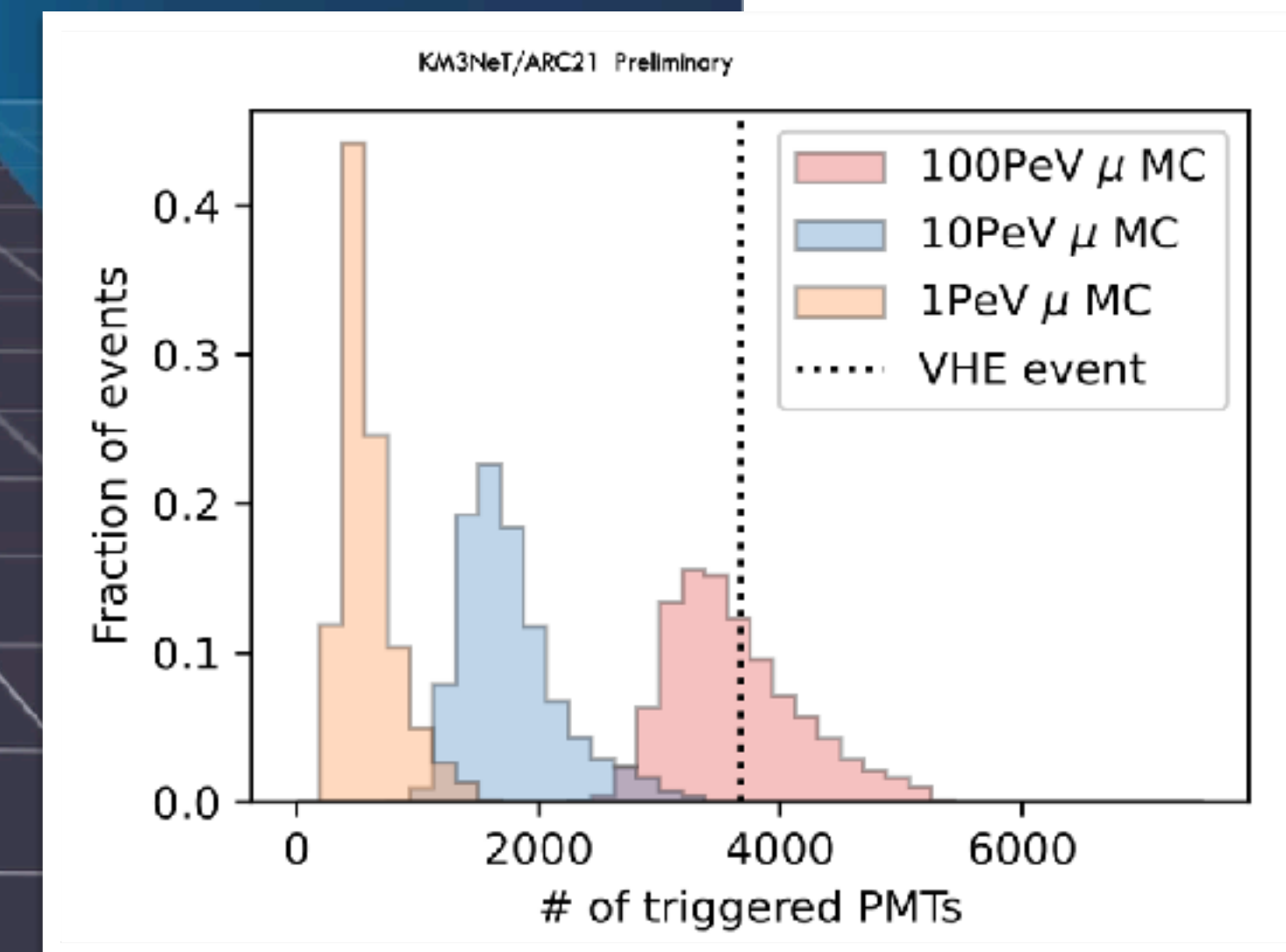
**NEWS IN FOCUS**



Five ARCA light detectors on board a ship, ready for deployment.

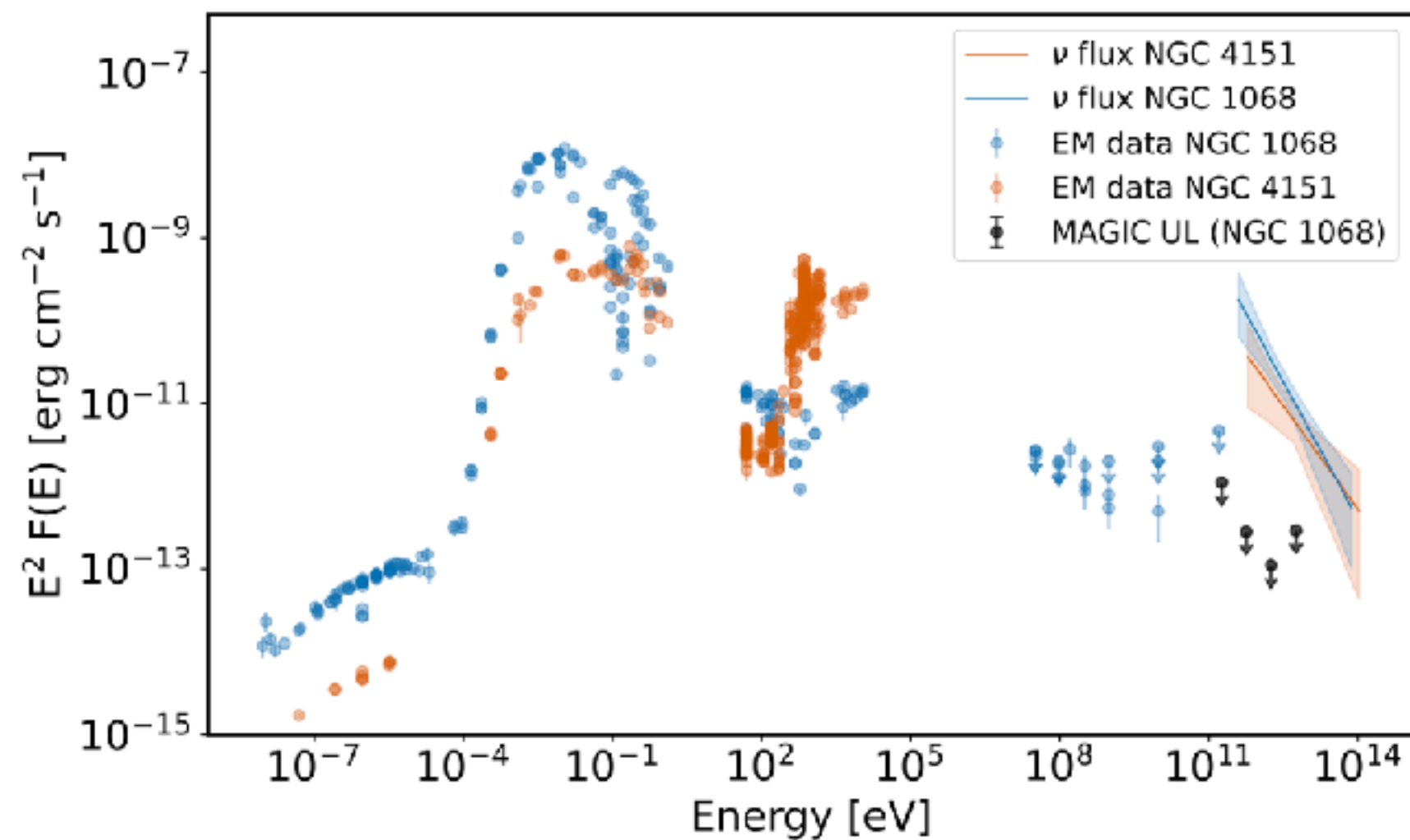
### 'FANTASTIC' NEUTRINO COULD BE MOST ENERGETIC EVER FOUND

Ultra-high-energy particle spotted by deep-sea detectors could point to a massive cosmic event



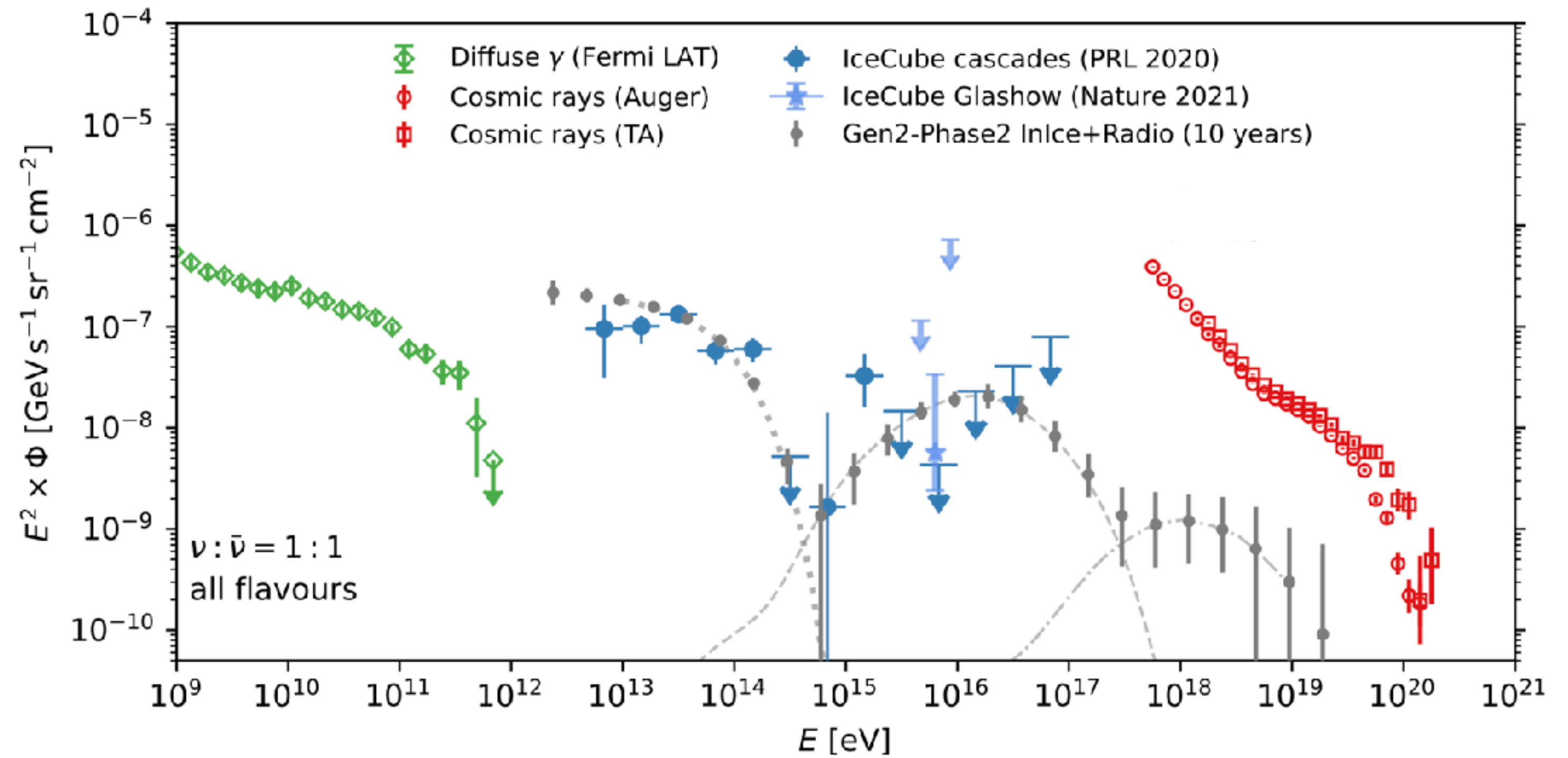
# Multi-Messenger Astrophysics

Observation by IceCube of NGC1068!



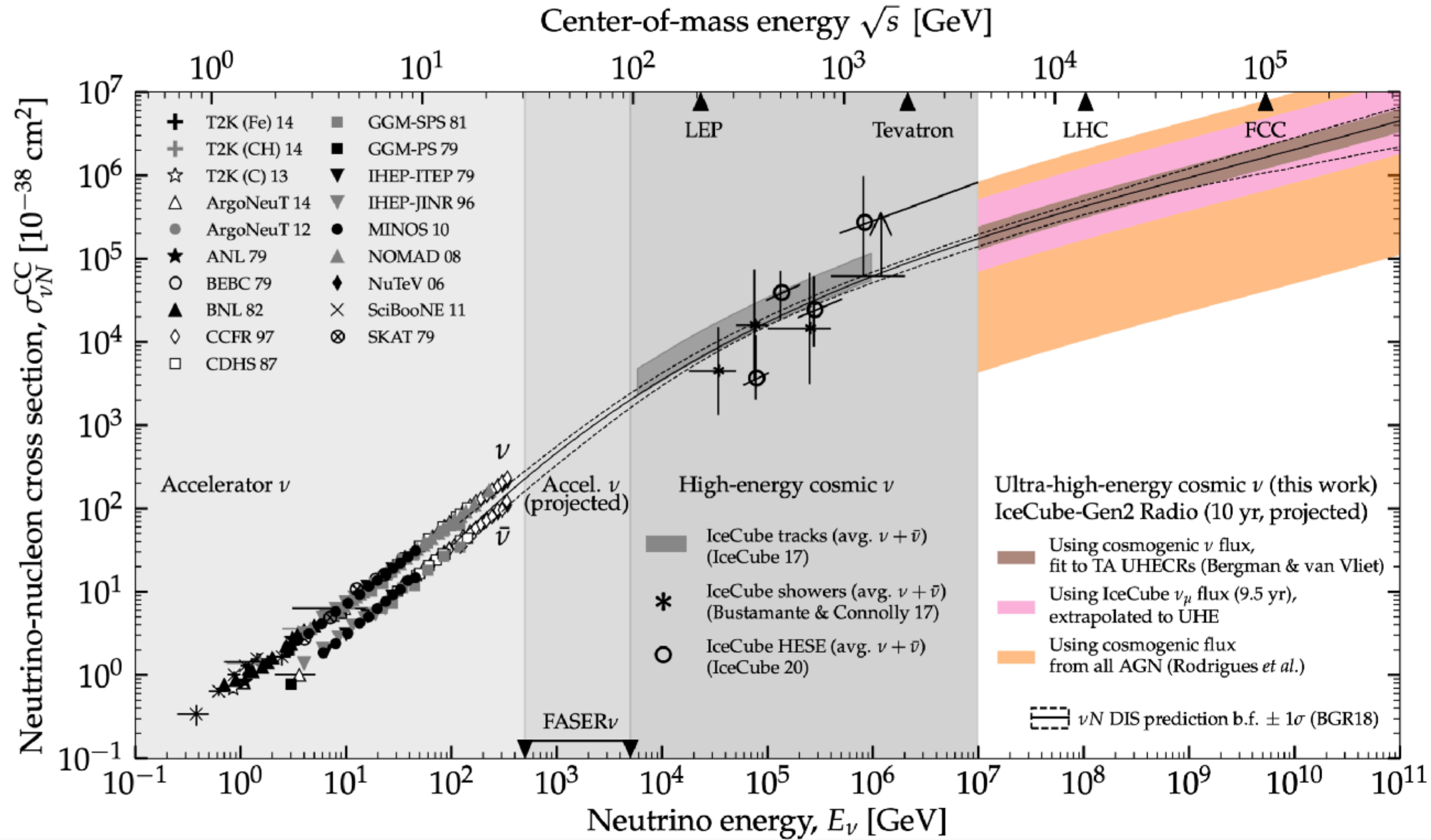
In 2022 ESO found a cloud of cosmic dust (obscuring/attenuating  $\gamma$  rays) at the centre of NGC1068 hiding a supermassive black hole.

Evidence for diffuse astrophysical neutrinos: “10 PeV Neutrinos a gateway to  $10^{20}$  eV cosmic rays!”



IceCube Gen2 in the race towards Ultra High Energy neutrinos with ARA, Trinity 18, GRAND, and POEMMA projects!

# The birth of Collider Neutrinos (at the LHC)



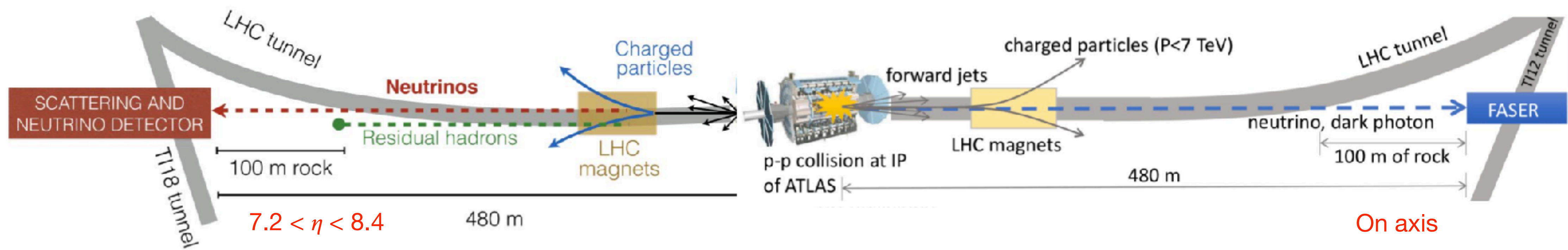
Accelerator neutrinos, up to O(100 GeV)

Cosmic neutrinos

# The birth of Collider Neutrinos (at the LHC)

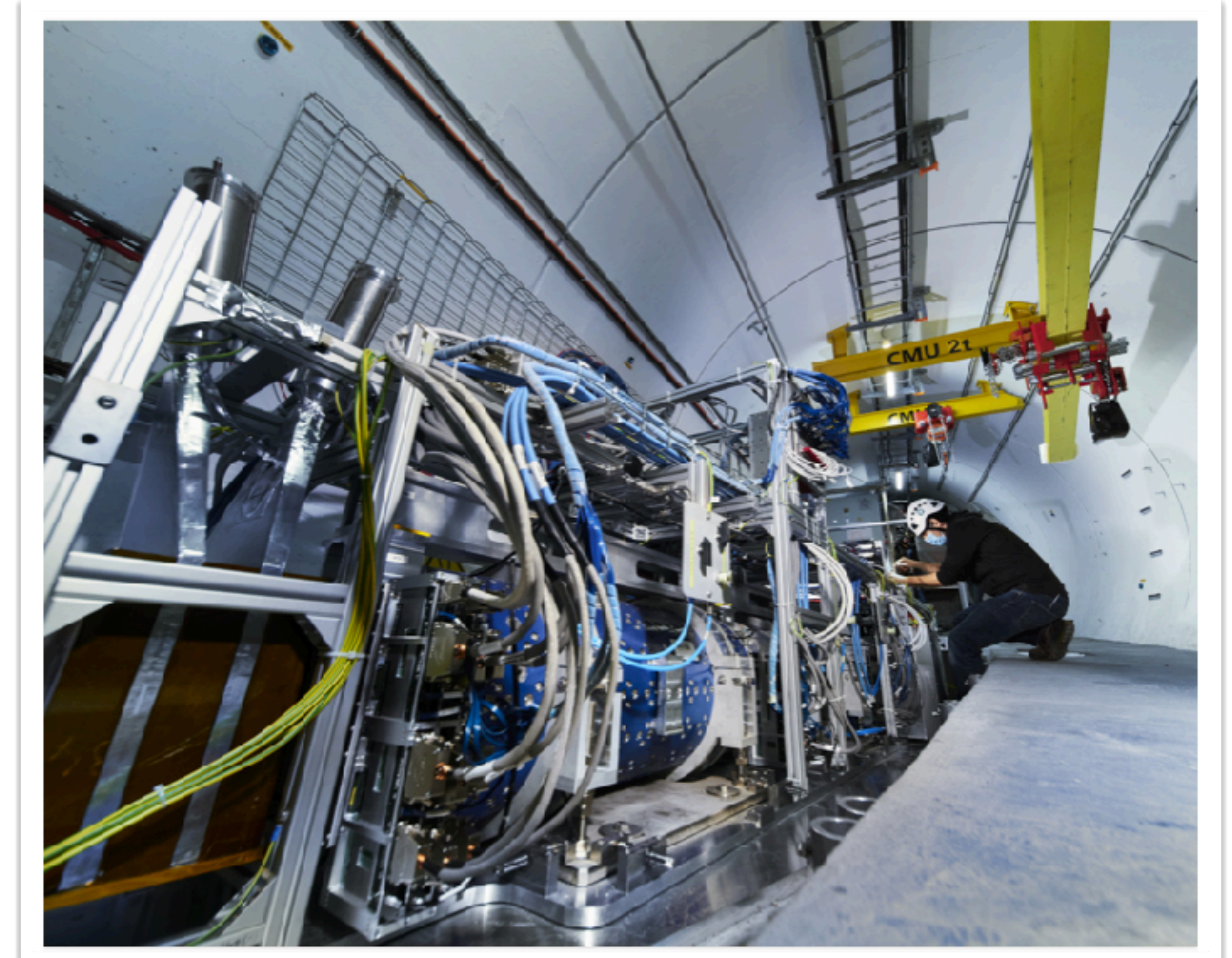
Dimitrievsky, Biswas

56



**SND**

**Faser-v**



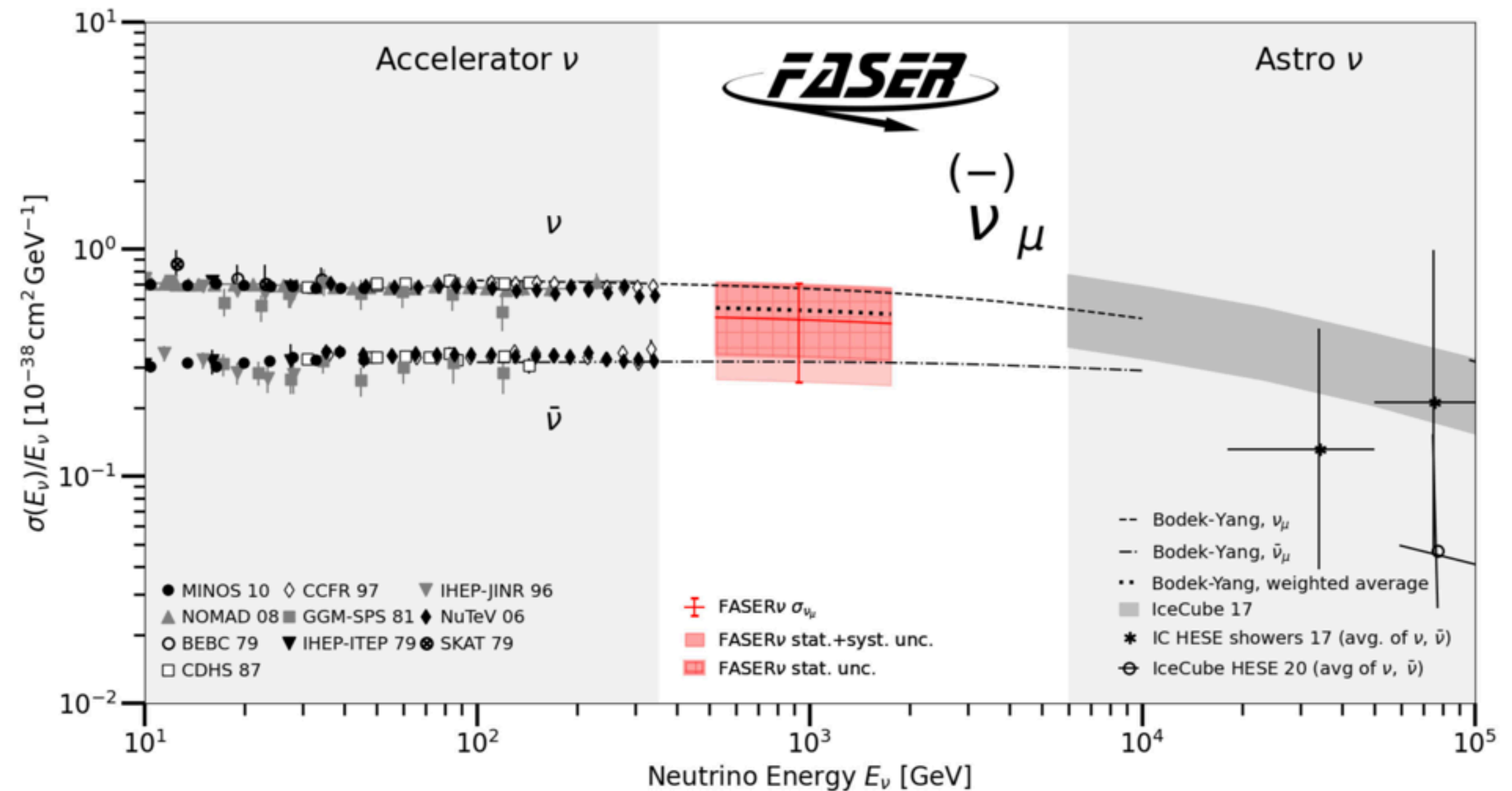
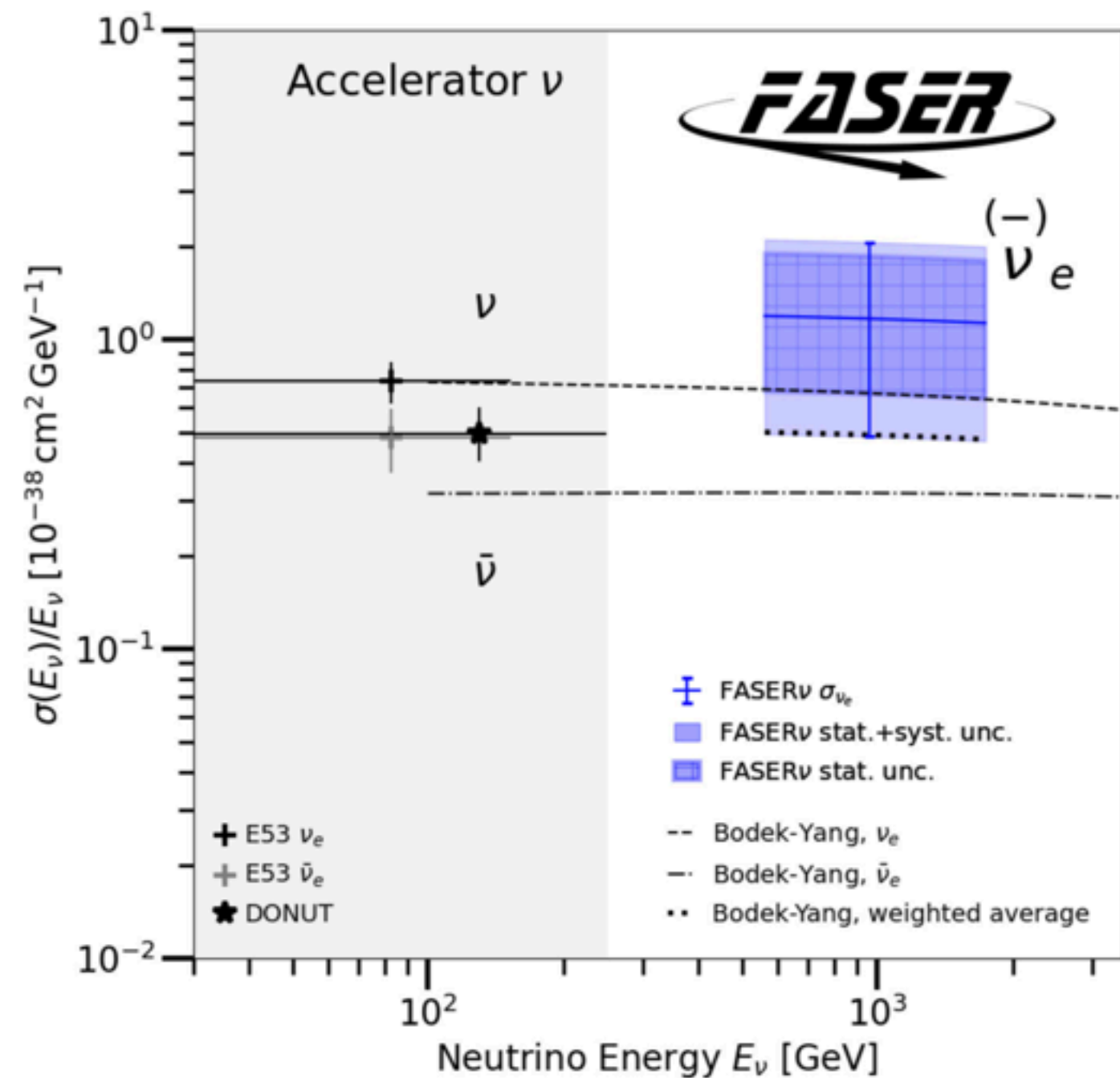
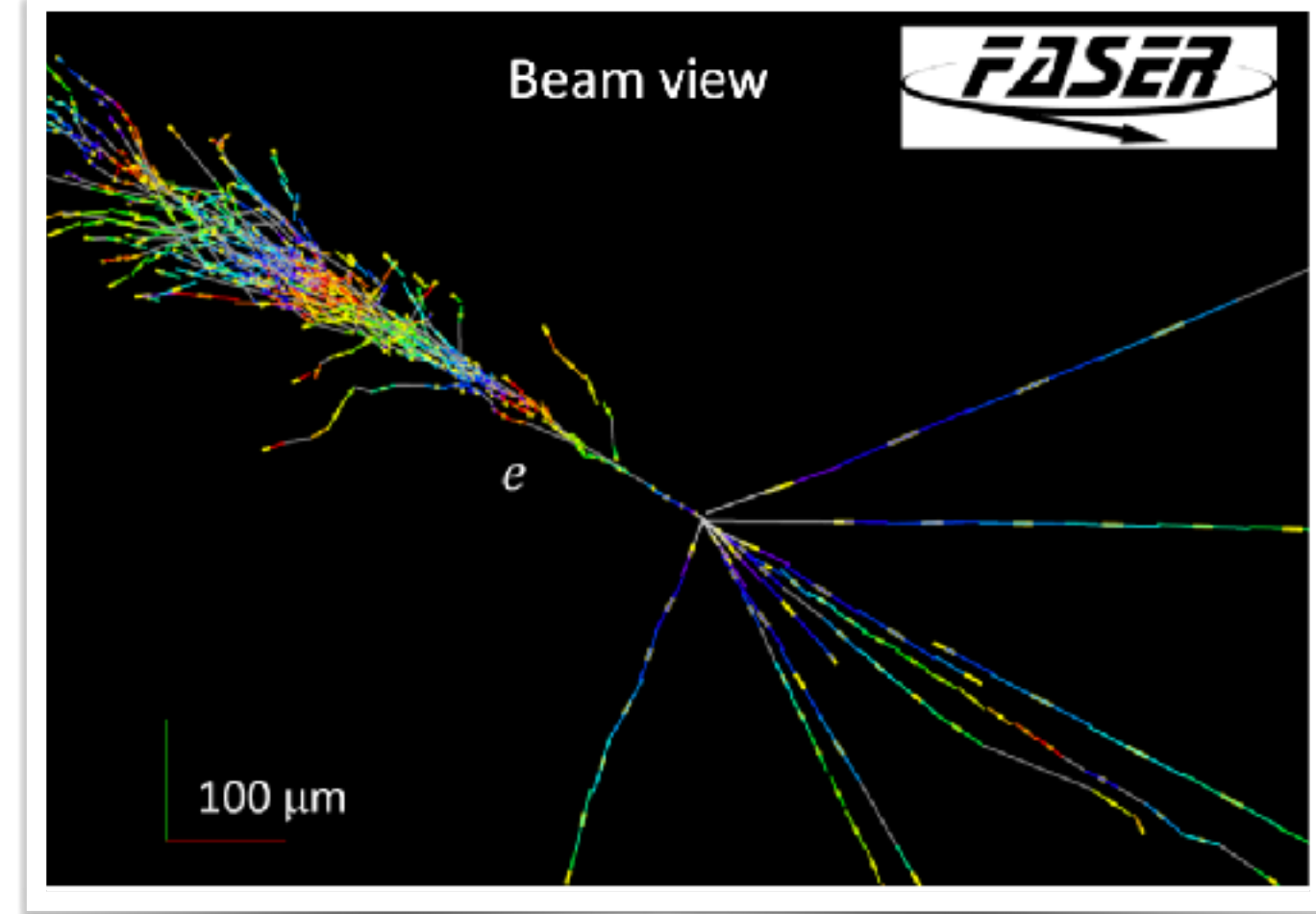
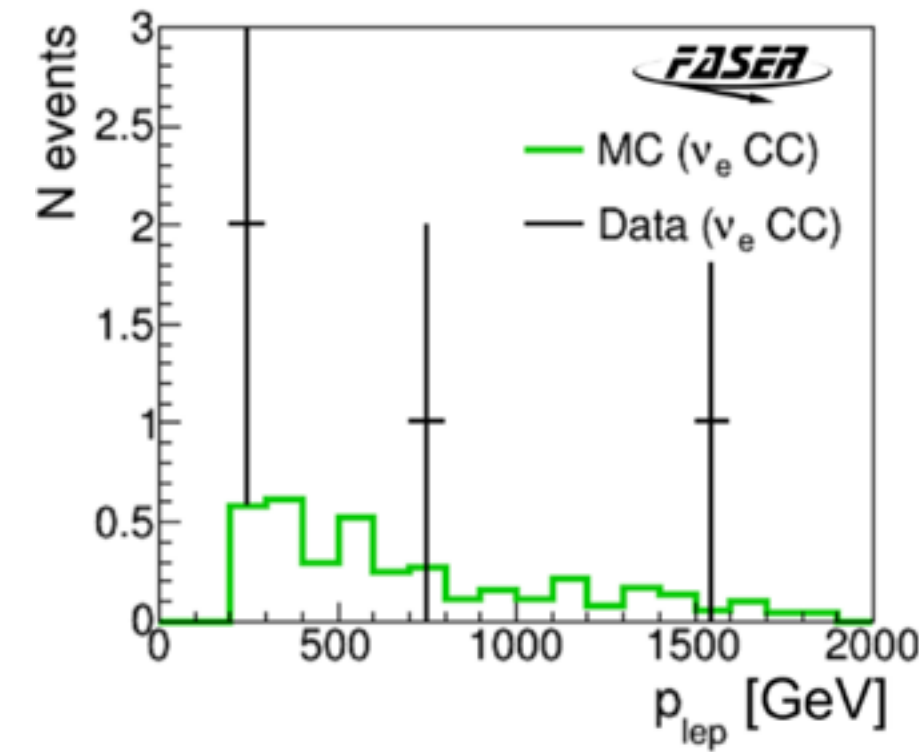


# The birth of Collider Neutrinos (at the LHC)

## Faser-ν

Results with emulsions (produced in Japan, sent to CERN and processed again in Japan)!

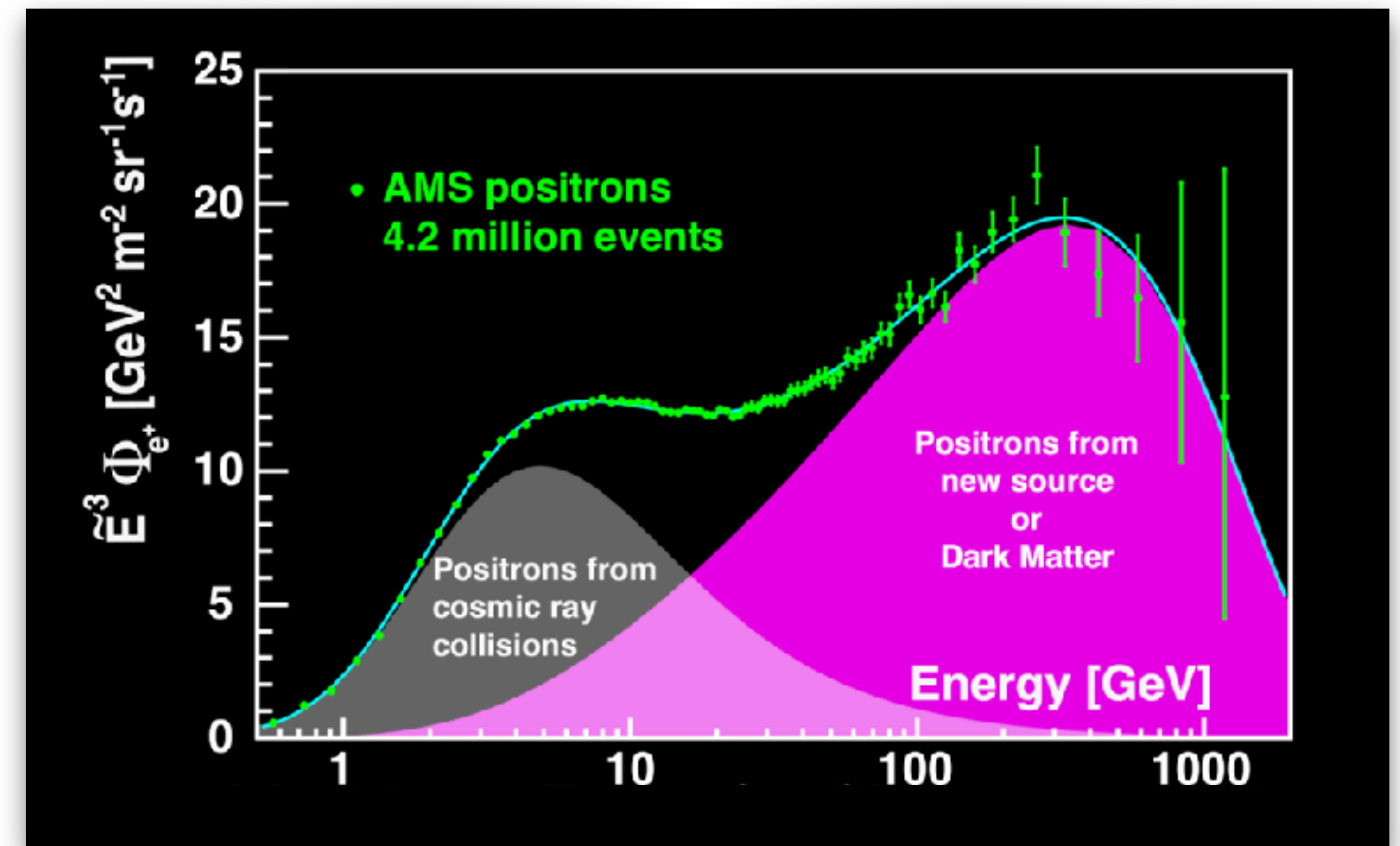
Results with emulsions (produced in Japan, sent to CERN and processed again in Japan)!



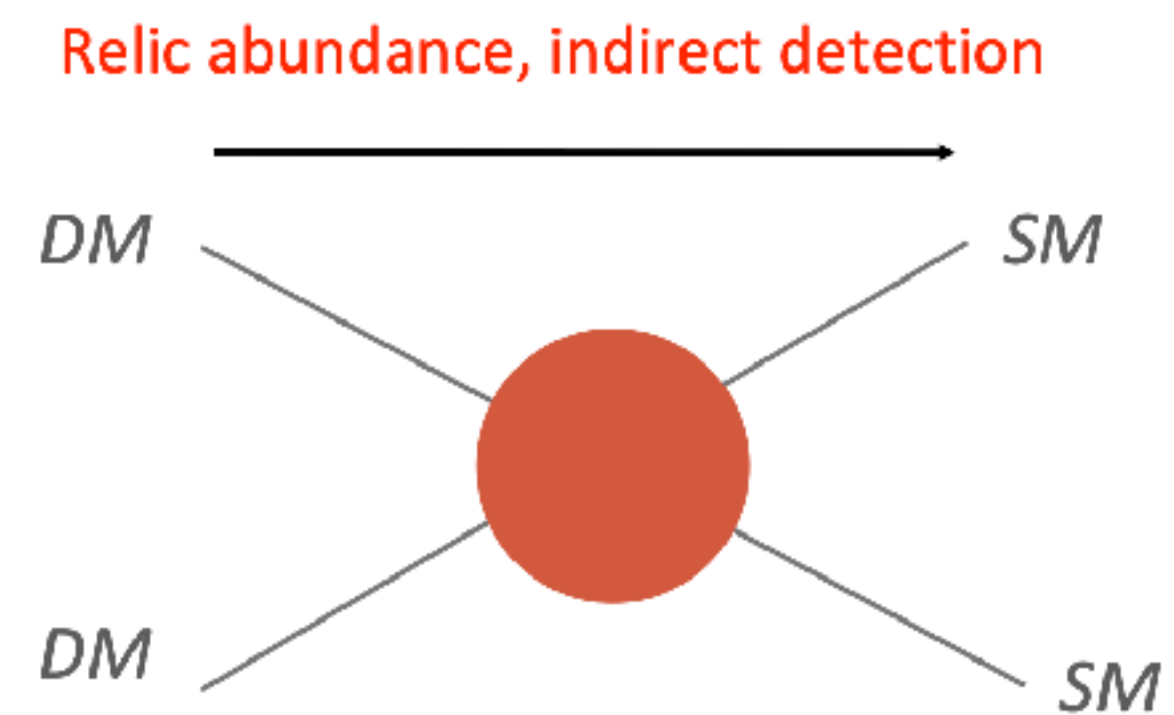
# AMS Positron Data



Latest data with complete dataset 2 times larger than previous release!



Interesting to note that there is a rather **sharp cut-off** indicative of mass range for DM candidate mass (in the DM annihilation hypothesis)



# LHCb on ${}^3\overline{\text{He}}$ production in $\Lambda_b$ decays

Yasmine Amhis

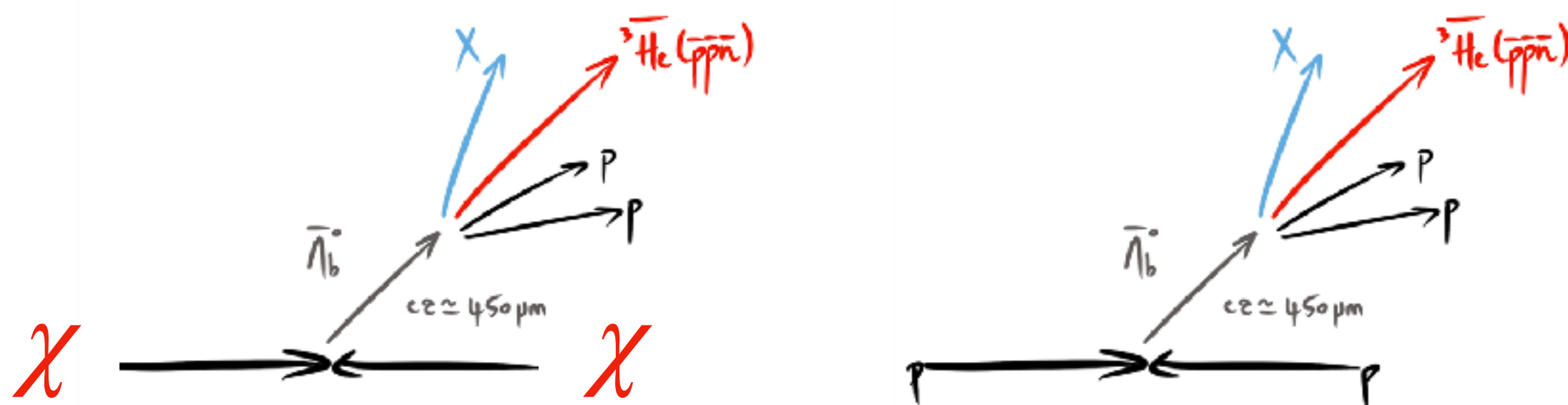
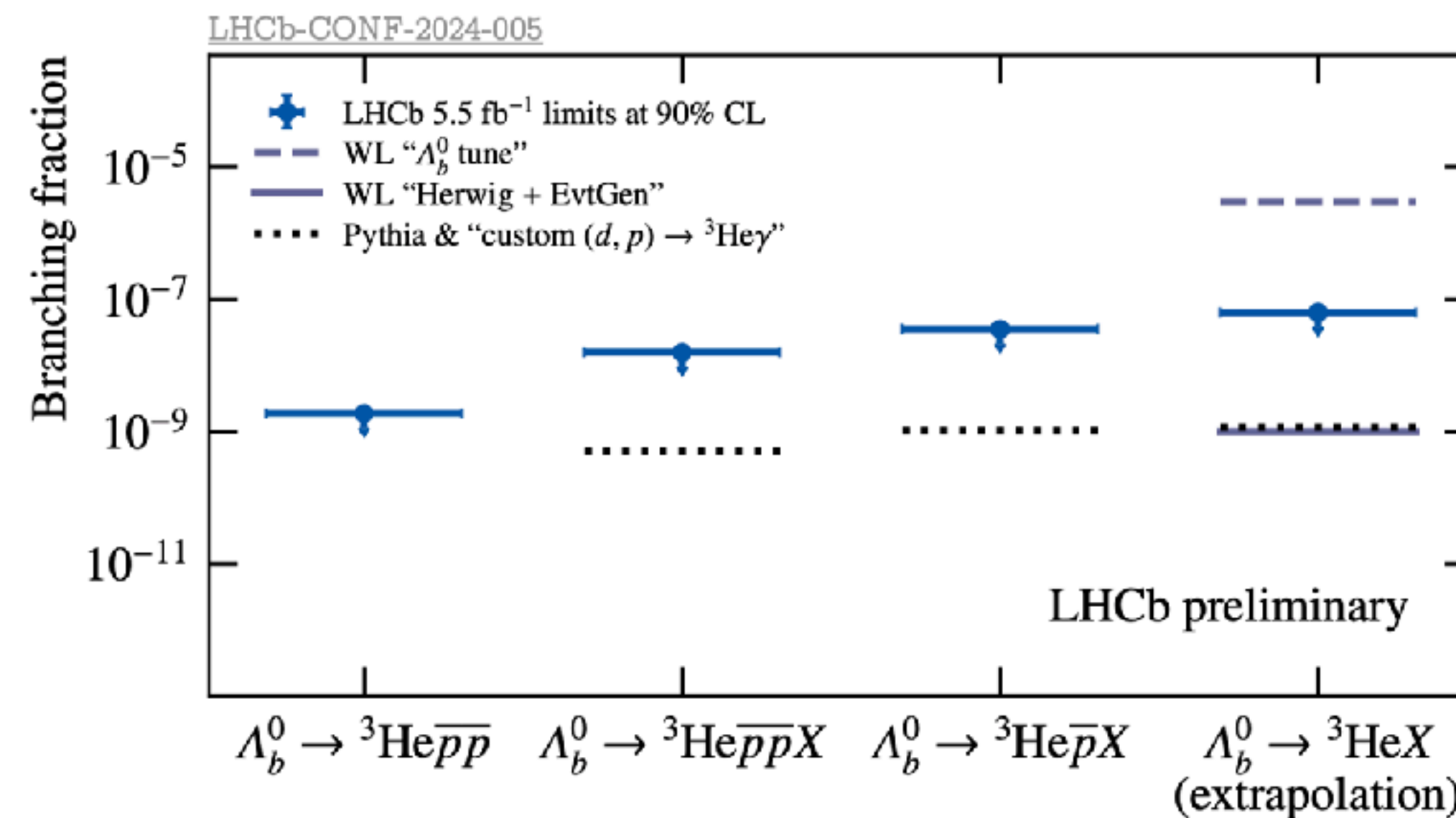
59



Anti-helium in Cosmic Rays could be a signature of physics beyond the standard model (e.g. **DM annihilation**)!

Interesting scenario where anti- $\Lambda_b$  are produced in Dark Matter scenario annihilation.

He3 identified with correlated measurements of charge between VELO and Silicon Strips Tracker



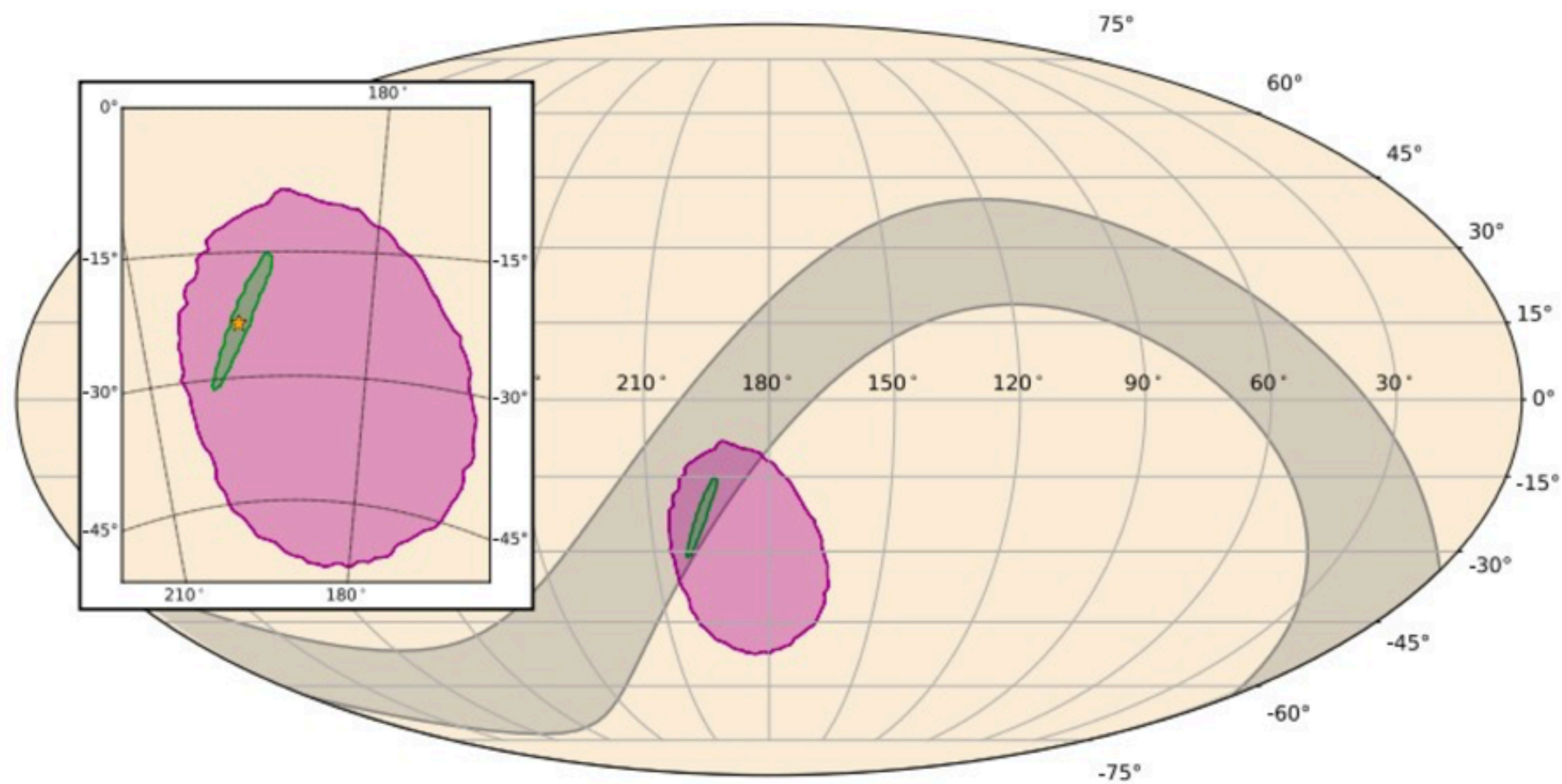
Important to understand the production of  $\Lambda_b, {}^3\overline{\text{He}}$  in cosmic rays

# Multi-Messenger Astrophysics

Vitor de Souza, Samya Nissanke

60

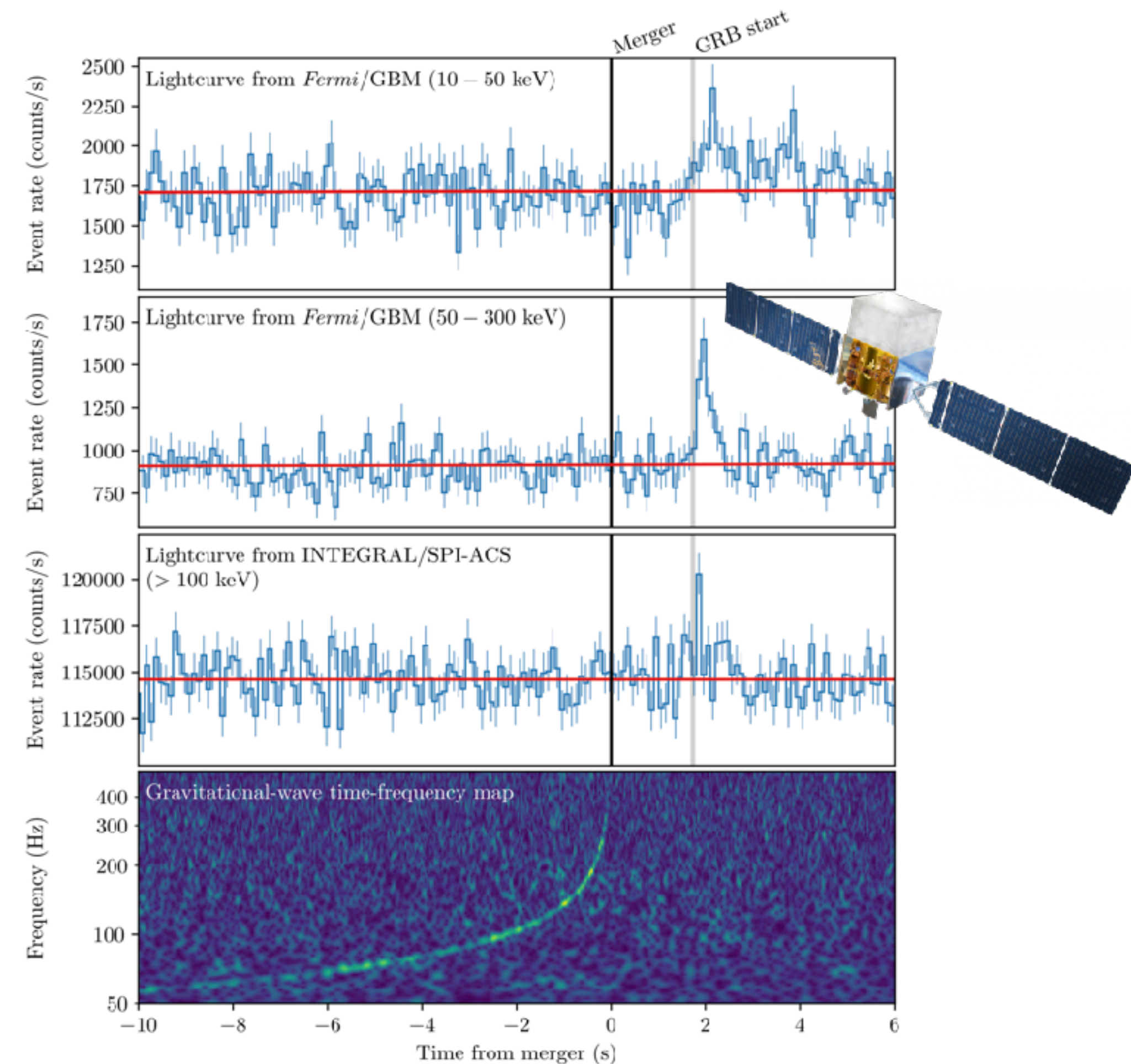
Gravitational Waves and Gamma-rays from a Binary Neutron Star Merger: GW170817 and GRB 170817A



Signal from LIGO-Hanford, LIGO- Livingston data, Fermi Gam. Ray Burst monitor (NASA) and INTEGRAL (ESA)

Test of speed of gravity and light and equivalence principle

$$-3 \times 10^{-15} \leq \frac{v_{\text{gw}}}{v_{\text{em}}} - 1 \leq 7 \times 10^{-16}$$

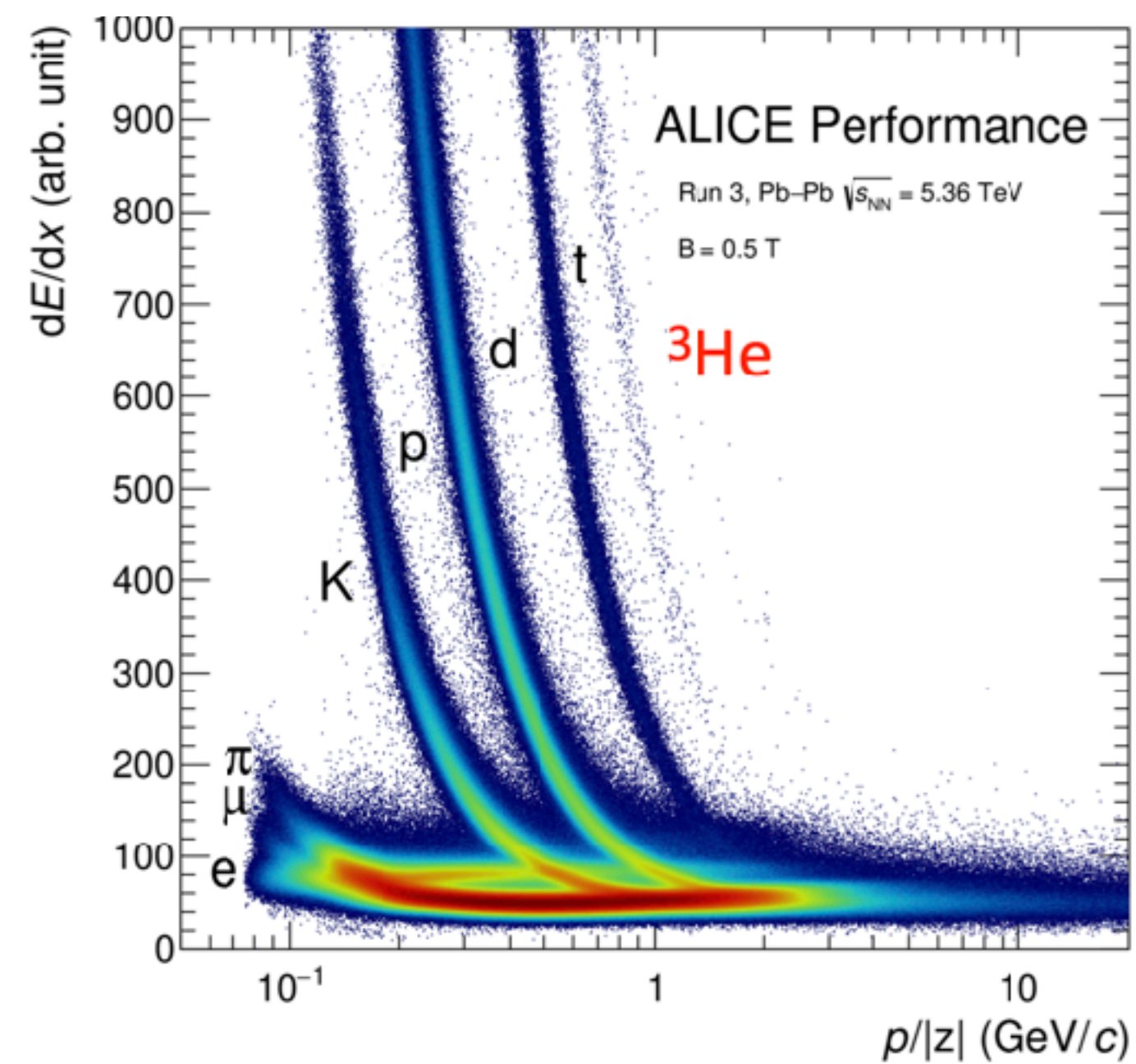


Signal from LIGO-Hanford, LIGO- Livingston data, Fermi (NASA) and INTEGRAL (ESA)

Also allows to constrain neutron star Equation of State!

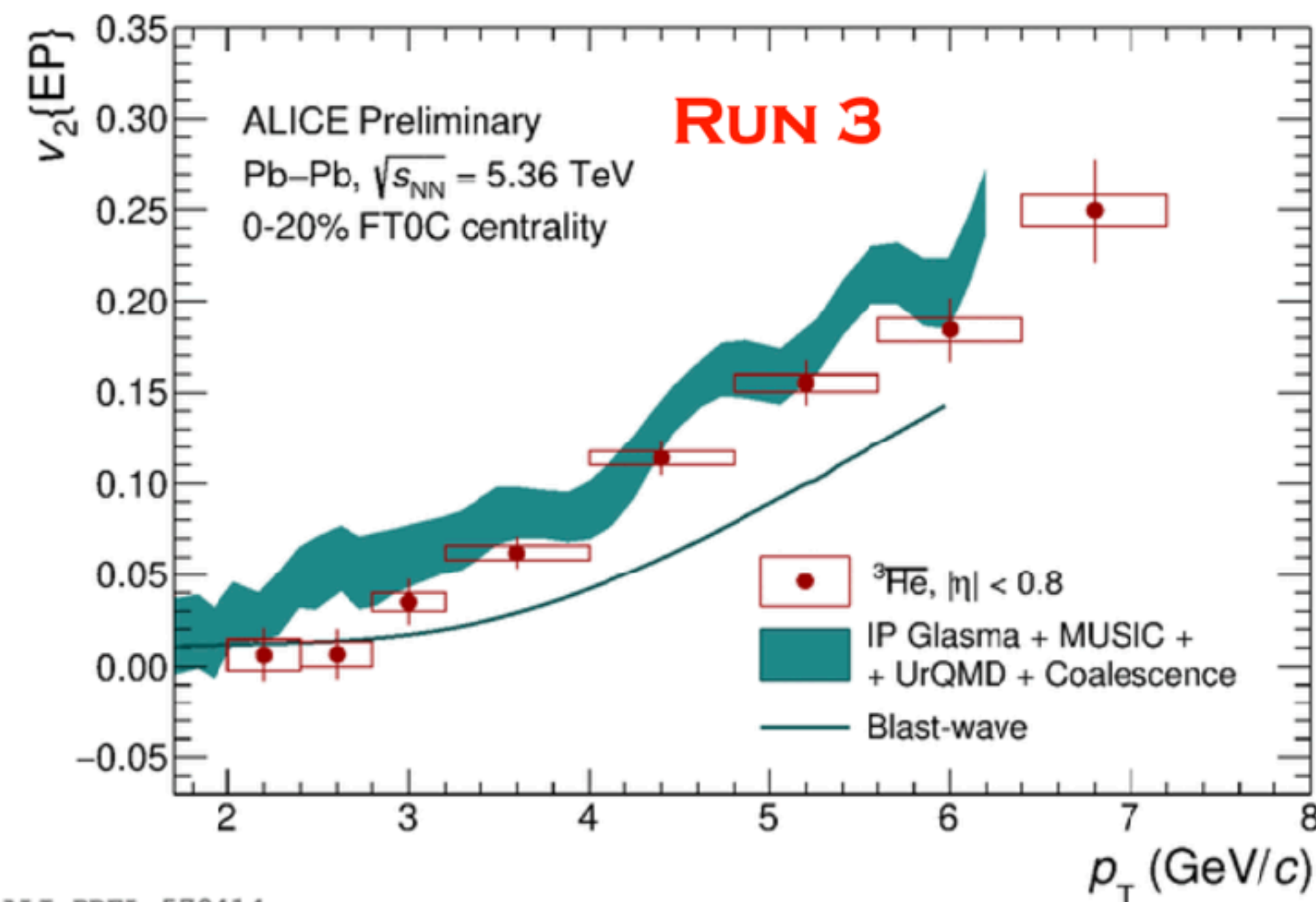
# Heavy Ion Physics

**Anti-Helium 3 production in heavy ion collisions, can give an interesting measure of the medium viscosity!**



ALI-PERF-529714

Helium nucleus ( $Z = 2$ ), clean PID, rare probe



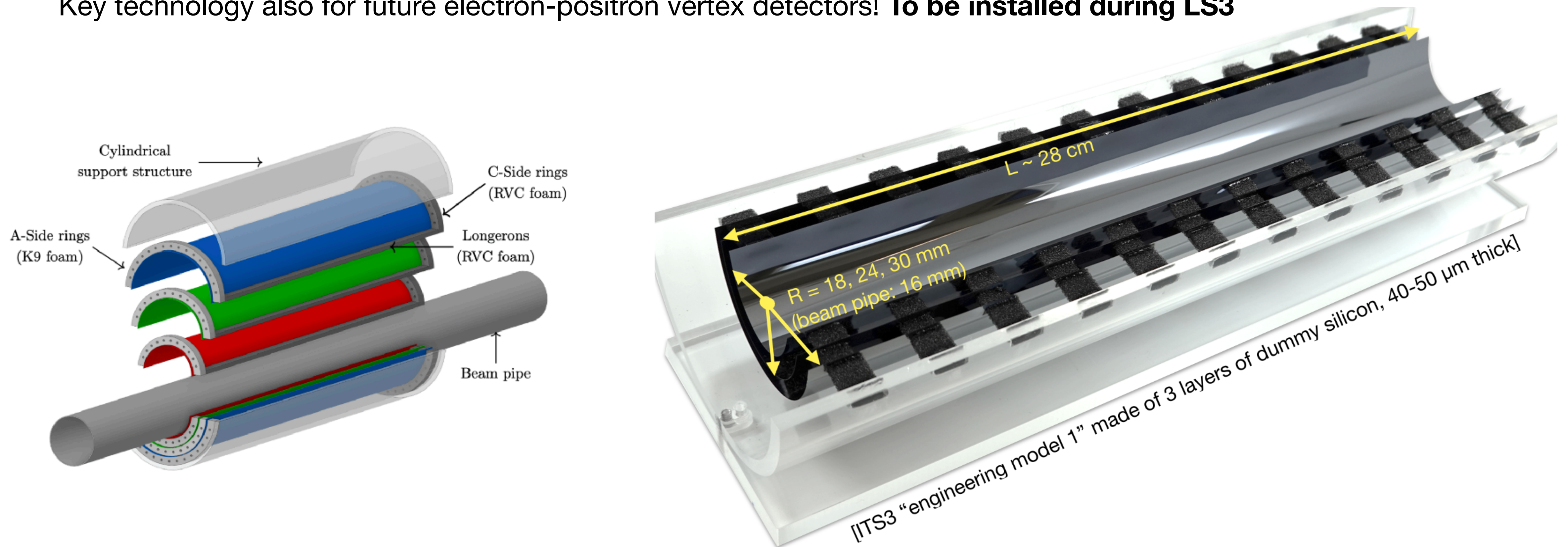
ALI-PREL-570414

Problem of **collective expansion** microscopic models



# Featuring ITS3 Upgrade for ALICE

Monolithic CMOS (MAPS - Monolithic Active Pixel Sensors) technology for the ITS3 upgrade of ALICE - Key technology also for future electron-positron vertex detectors! **To be installed during LS3**

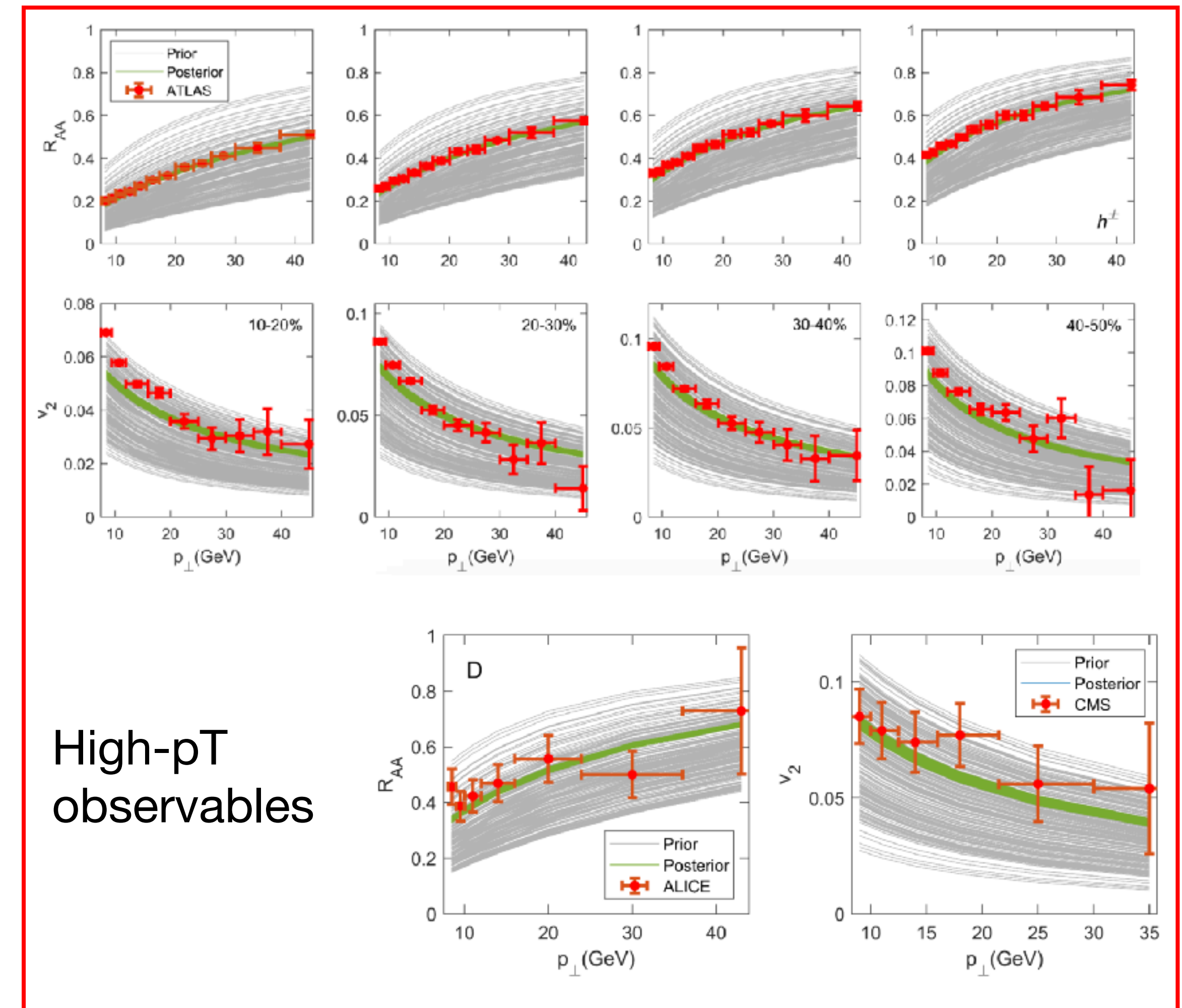
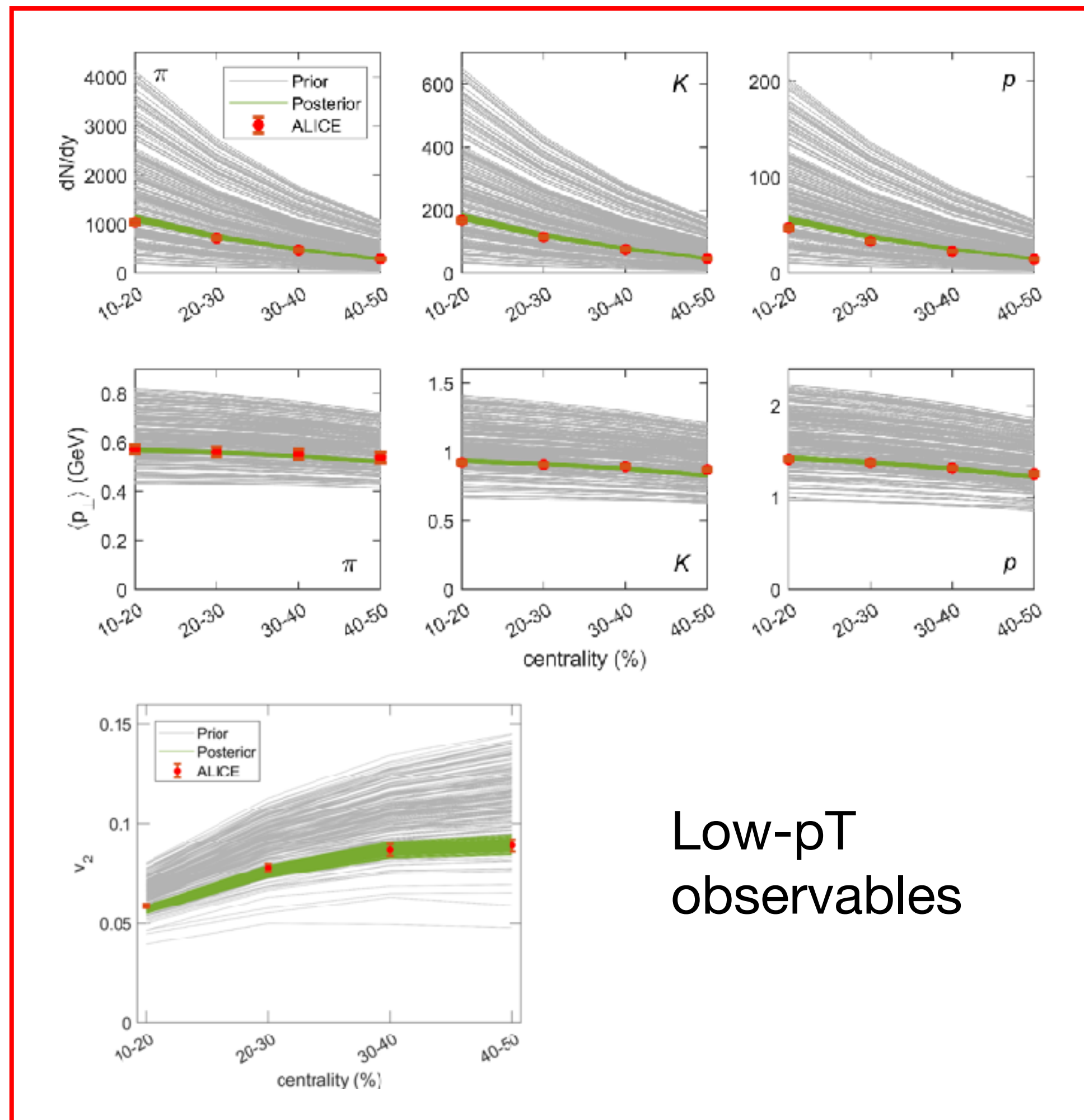


The design of the new vertex detector aims to reduce the material budget of the **first detection layer to an unprecedented minimum of 0.05%  $X_0$** , and to **get closer to the interaction point at a radial distance of 18 mm**.

Requires having in the active area only the thin MAPS silicon sensor (<50 $\mu$ m)

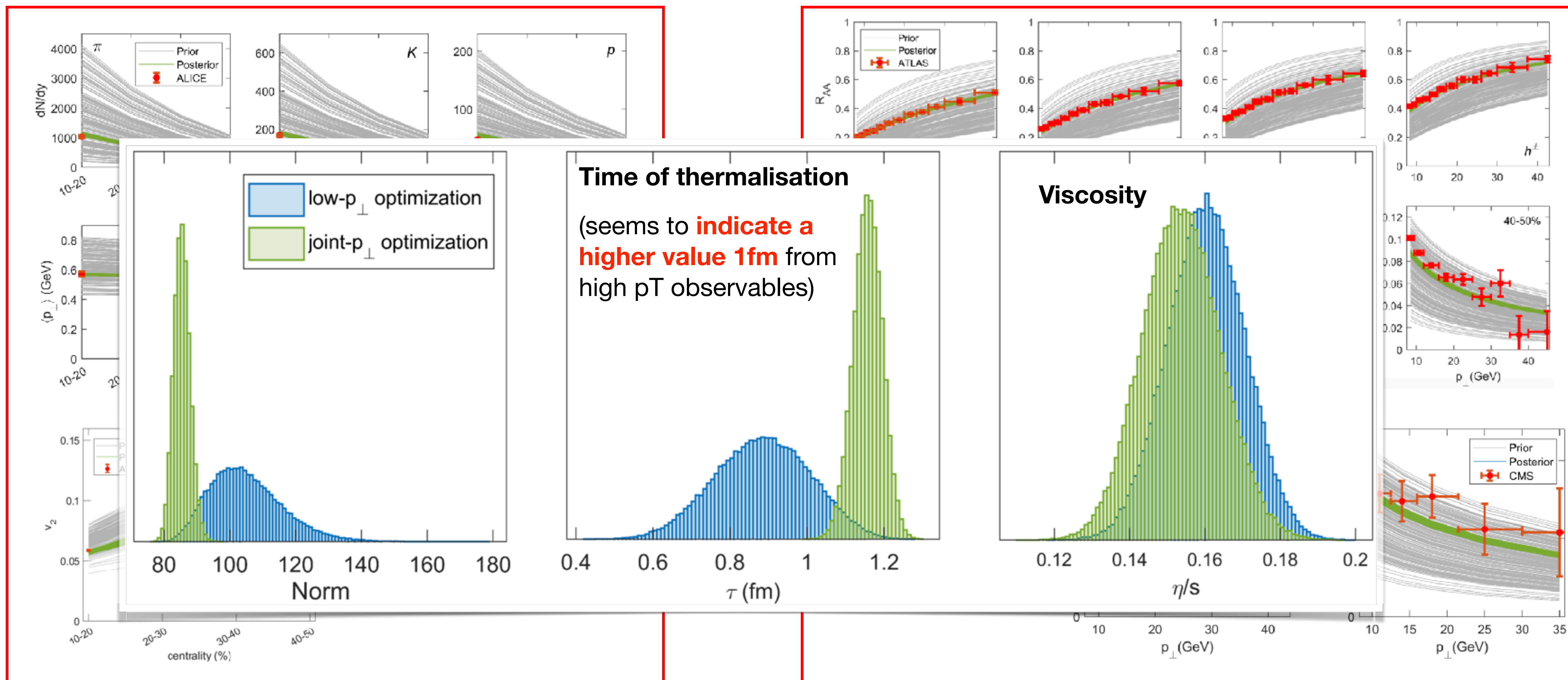
# Heavy Ion Physics

Development of a global framework to constrain QGP properties through both high-pT (Nuclear modification factors RAA) and low pT (flows) observables - DREENA!



# Heavy Ion Physics

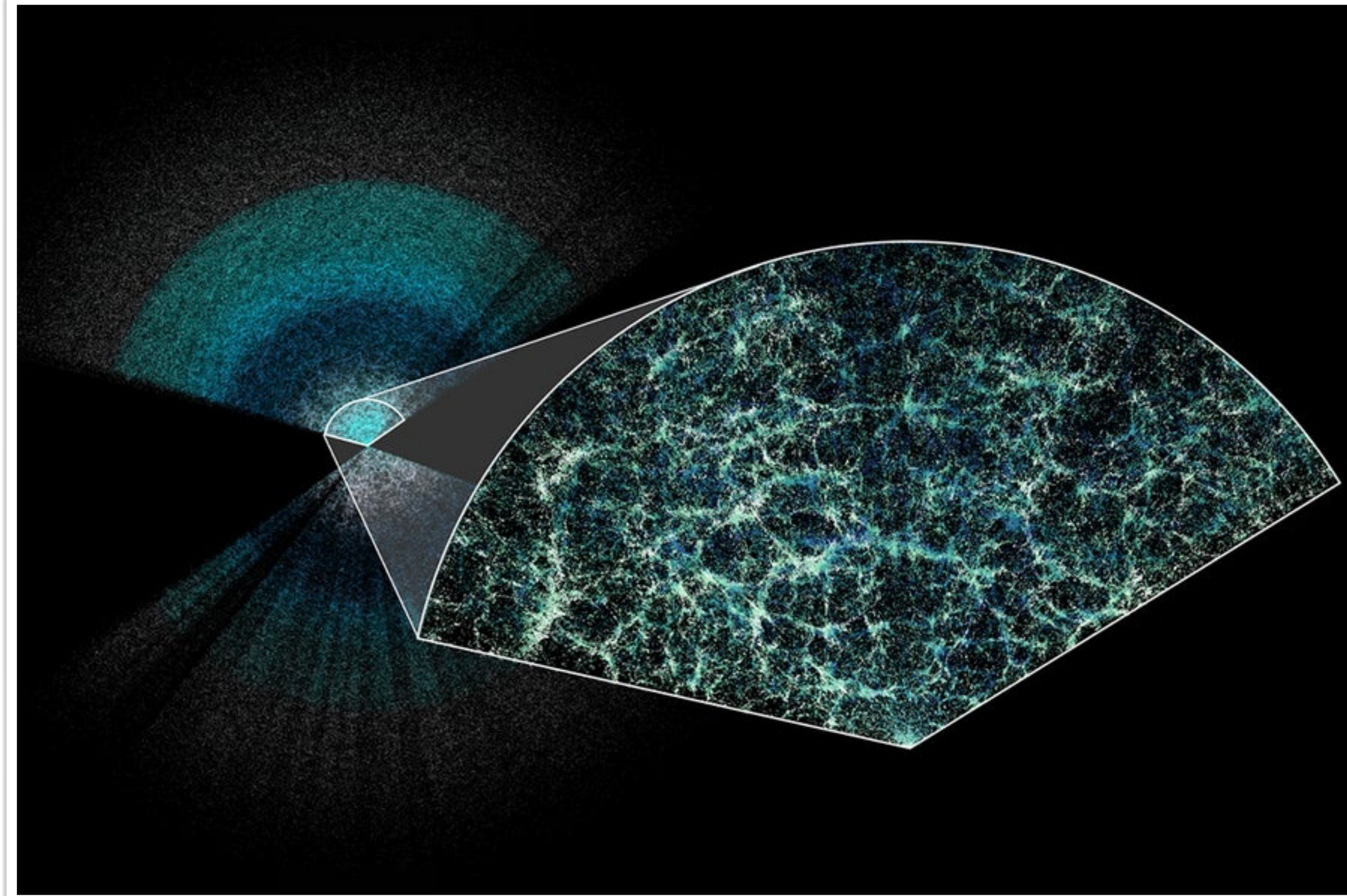
Development of a global framework to constrain QGP properties through both high- $p_T$  (Nuclear modification factors RAAs) and low  $p_T$  (flows) observables - DREENA!



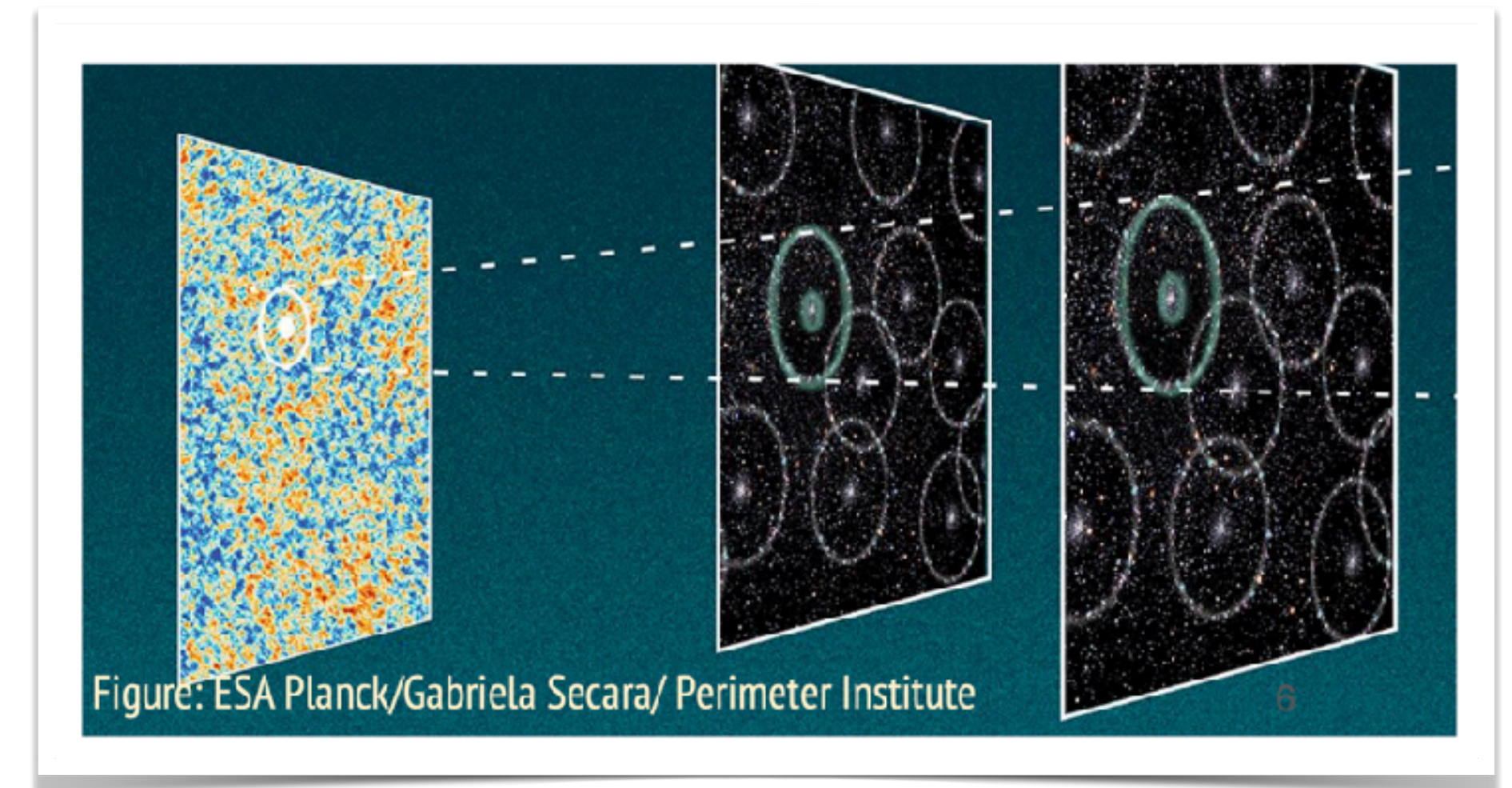


# Cosmology

**DESI** - Dark Energy Spectroscopic Instrument in operations since 2021 and first publication of Year 1 Results!  
In one year more specific data than all previous experiments!

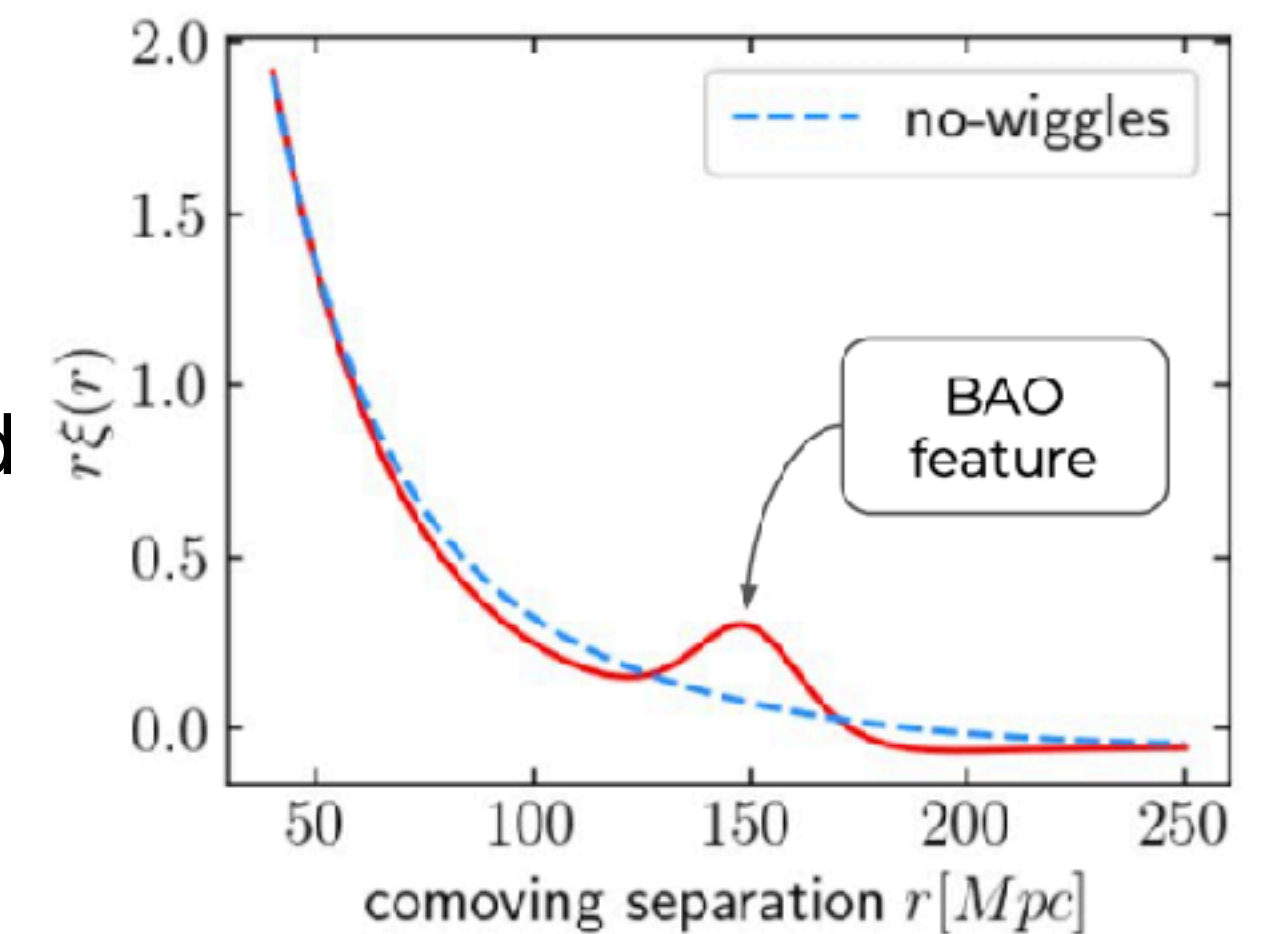


DESI 3D map of our universe to date.



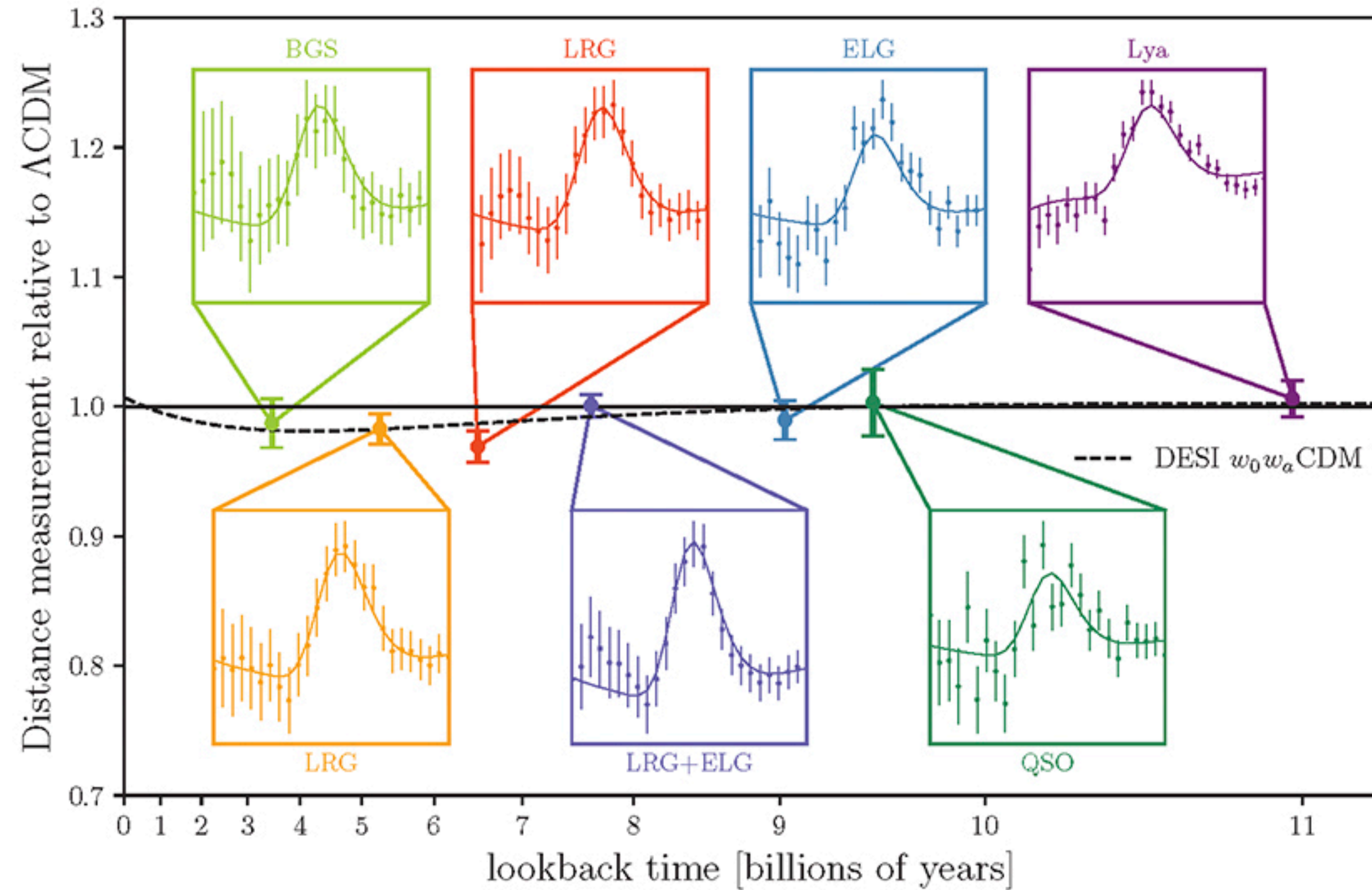
## Baryon acoustic oscillations

Early universe overdensities leave imprints (frozen when baryons and radiation decouple  $\sim 400,000$ y) in the distribution of galaxies.



# Cosmology

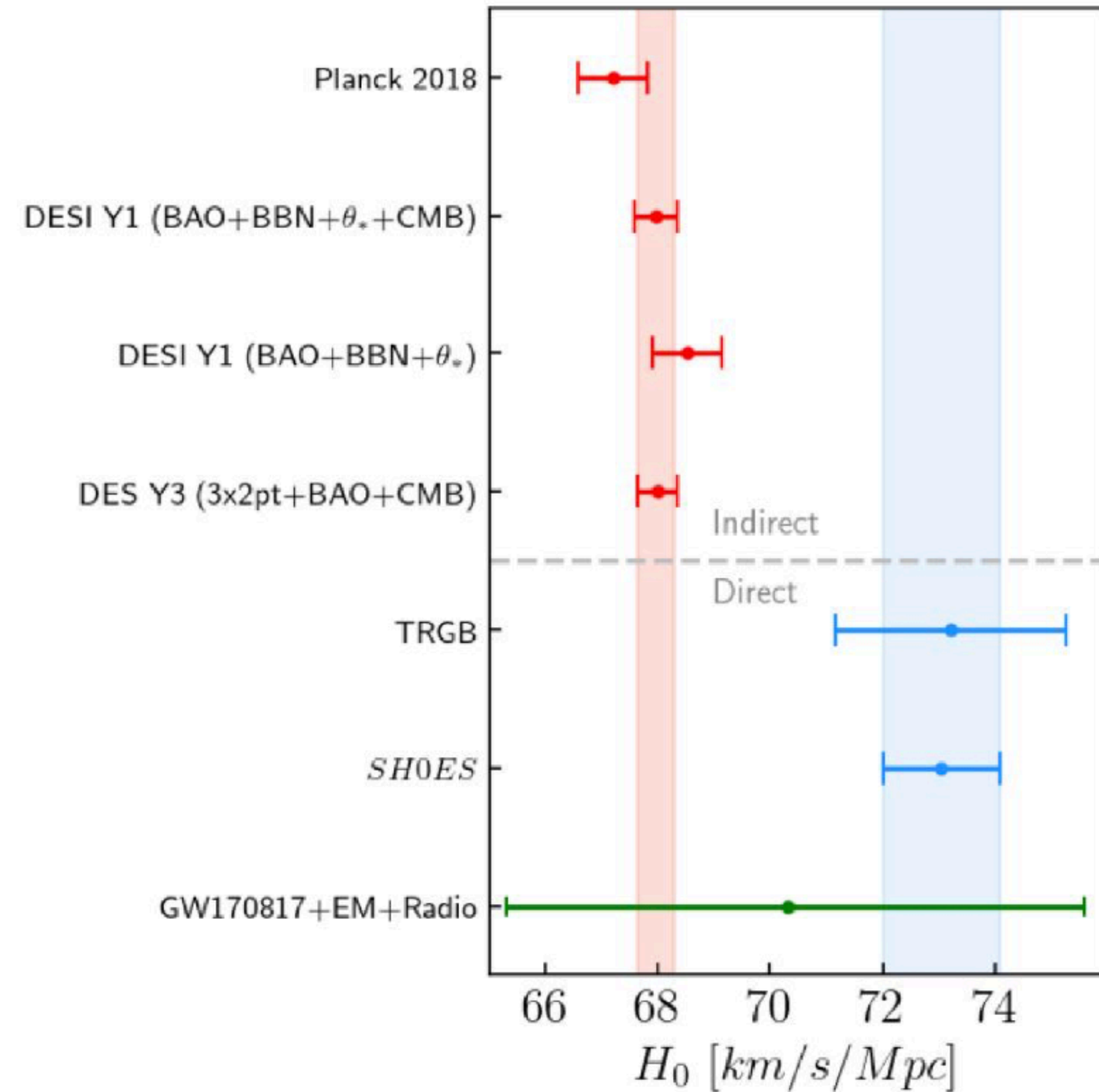
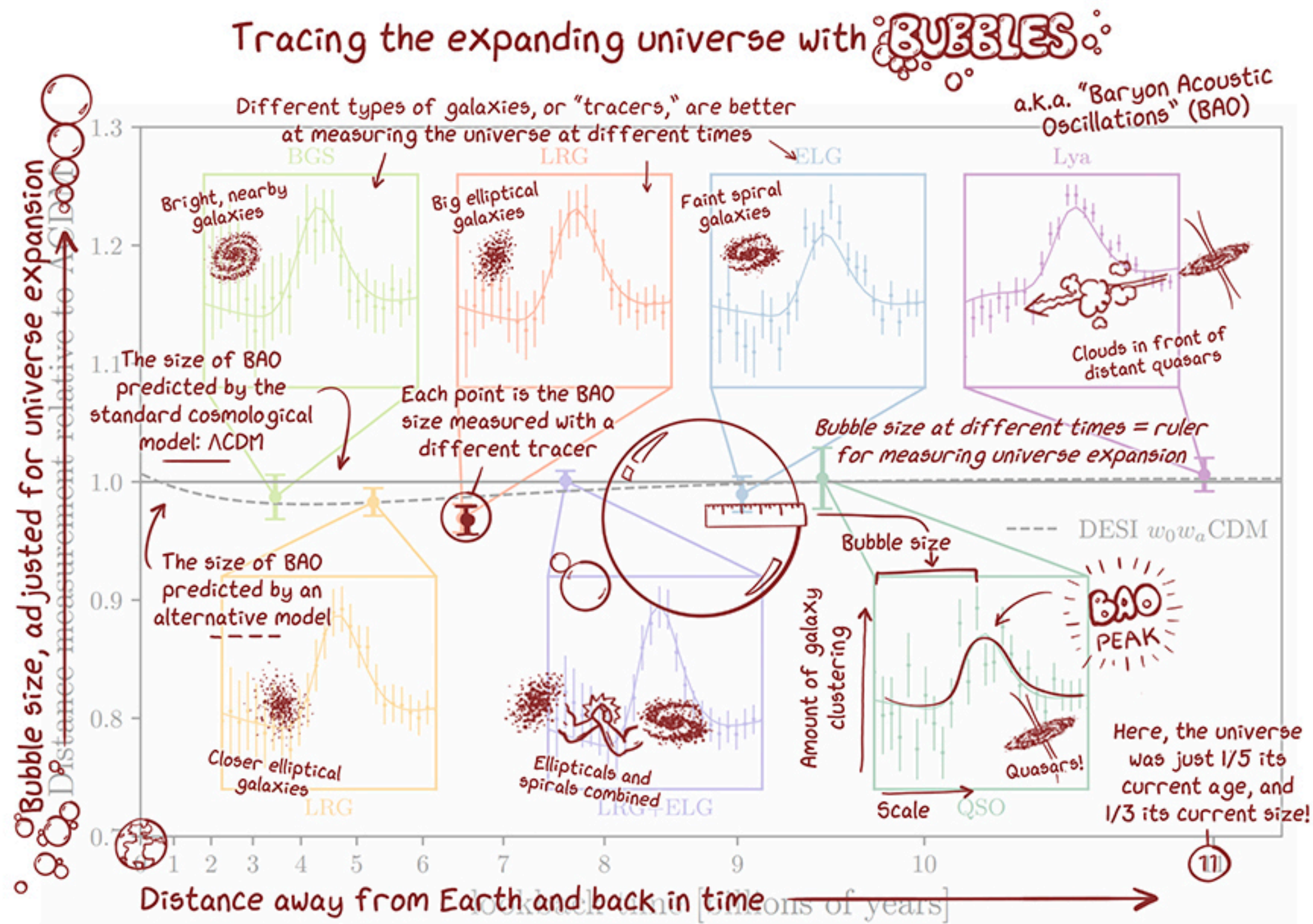
- 6 million spectra grouped in 7 different redshifts [ $0.4 < z < 4.2$ ] (LGR, BGS, QSO, Lya)
- Fully exploits the capabilities of DESI and increases sensitivity!



# Cosmology

- 6 million spectra grouped in 7 different redshifts [0.4 < z < 4.2] (LGR, BGS, QSO, Lya)
- Fully exploits the capabilities of DESI and increases sensitivity!

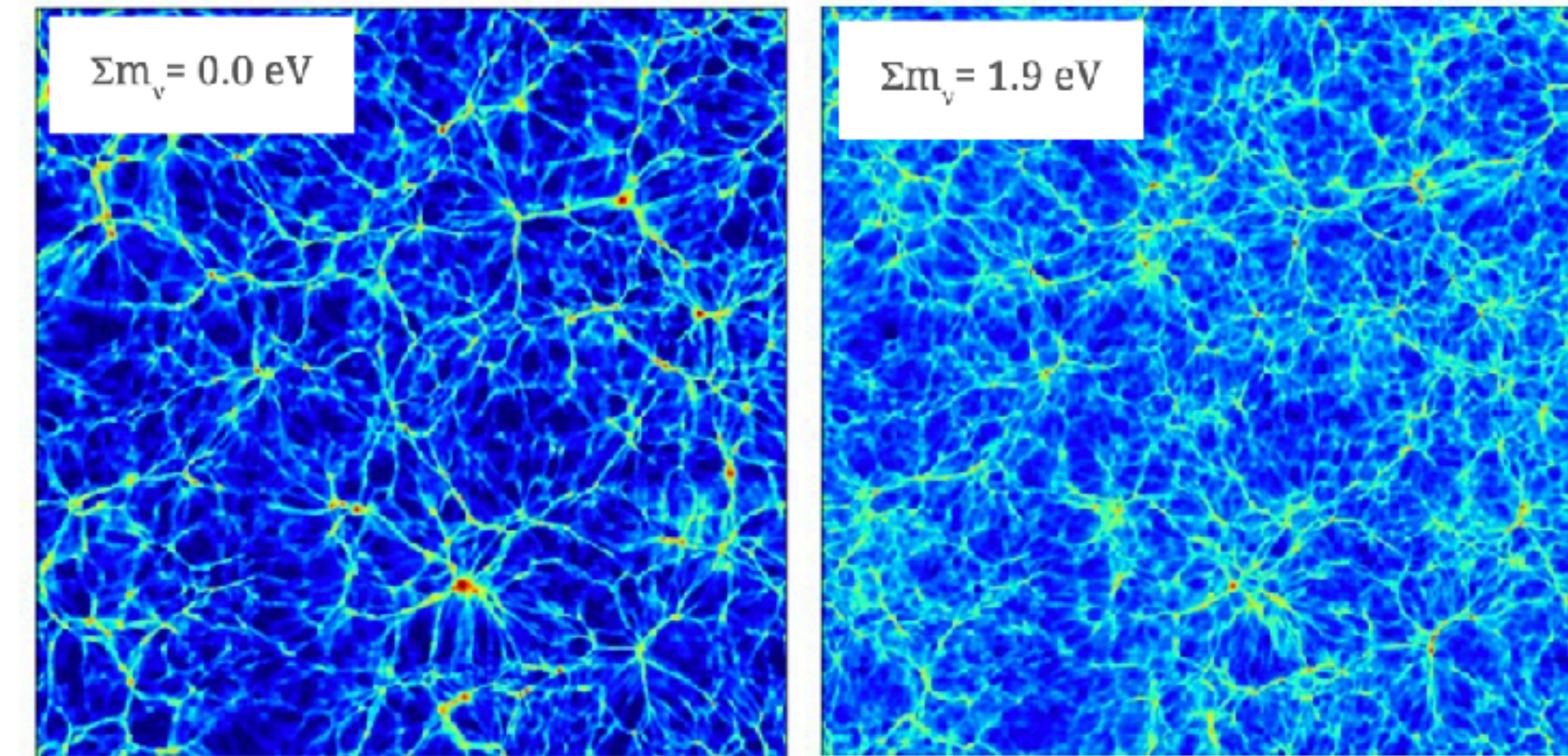
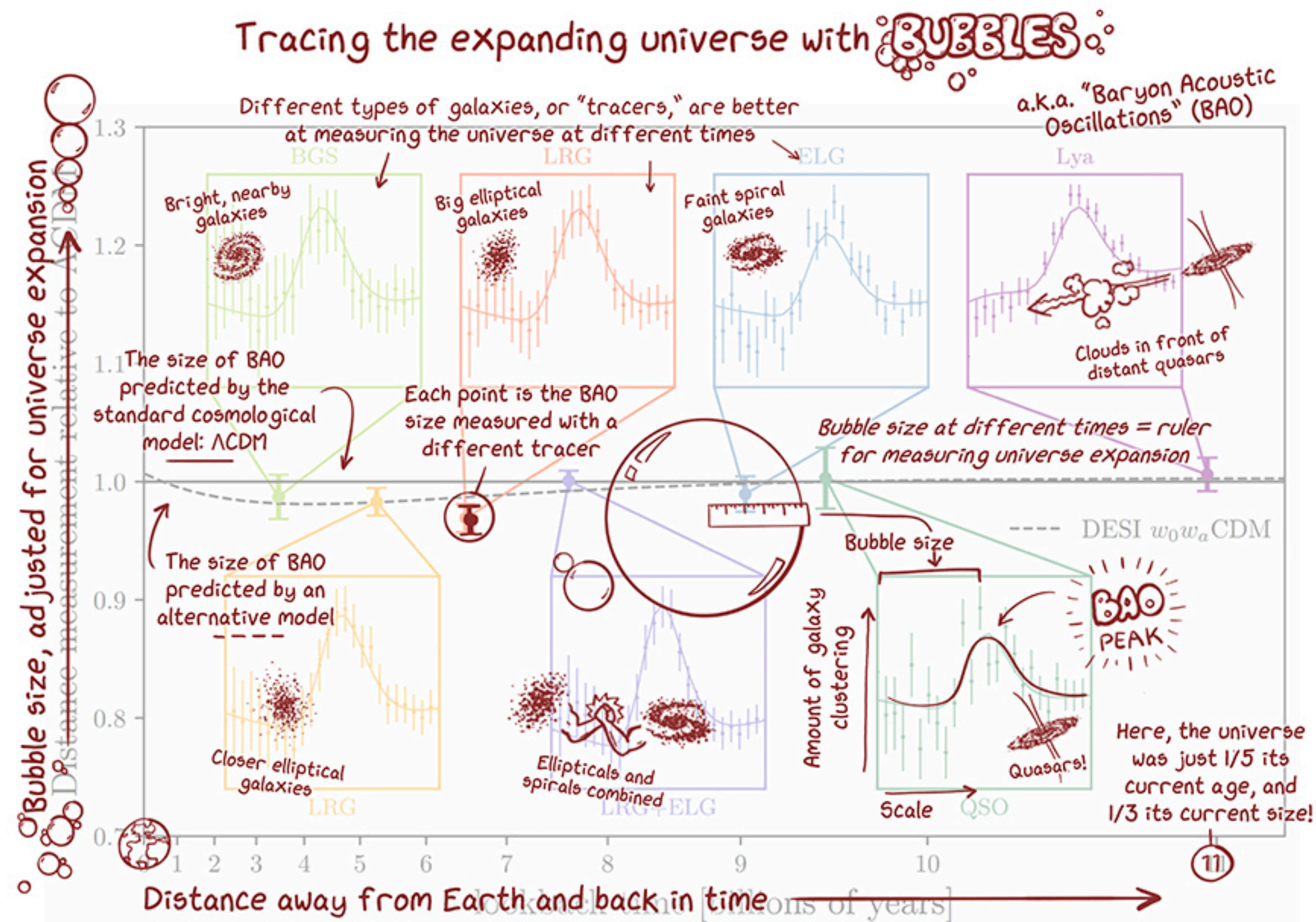
The Hubble tension does not significantly change with much improved Cosmological BAO measurements!



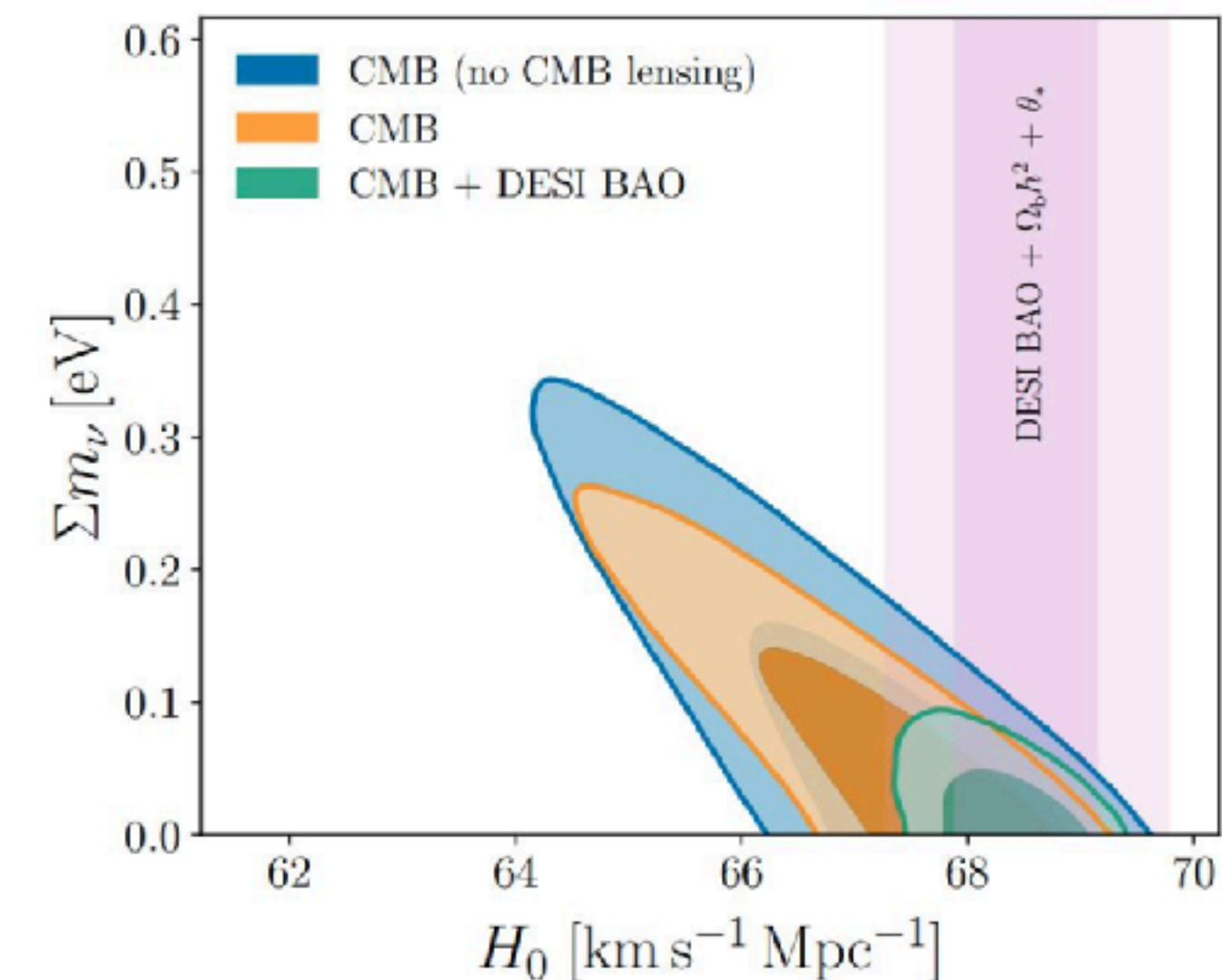
# Cosmology

- 6 million spectra grouped in 7 different redshifts [0.4 < z < 4.2] (LGR, BGS, QSO, Ly $\alpha$ )
- Fully exploits the capabilities of DESI and increases sensitivity!

The sum of neutrino masses affect how matter is clustering in the Universe

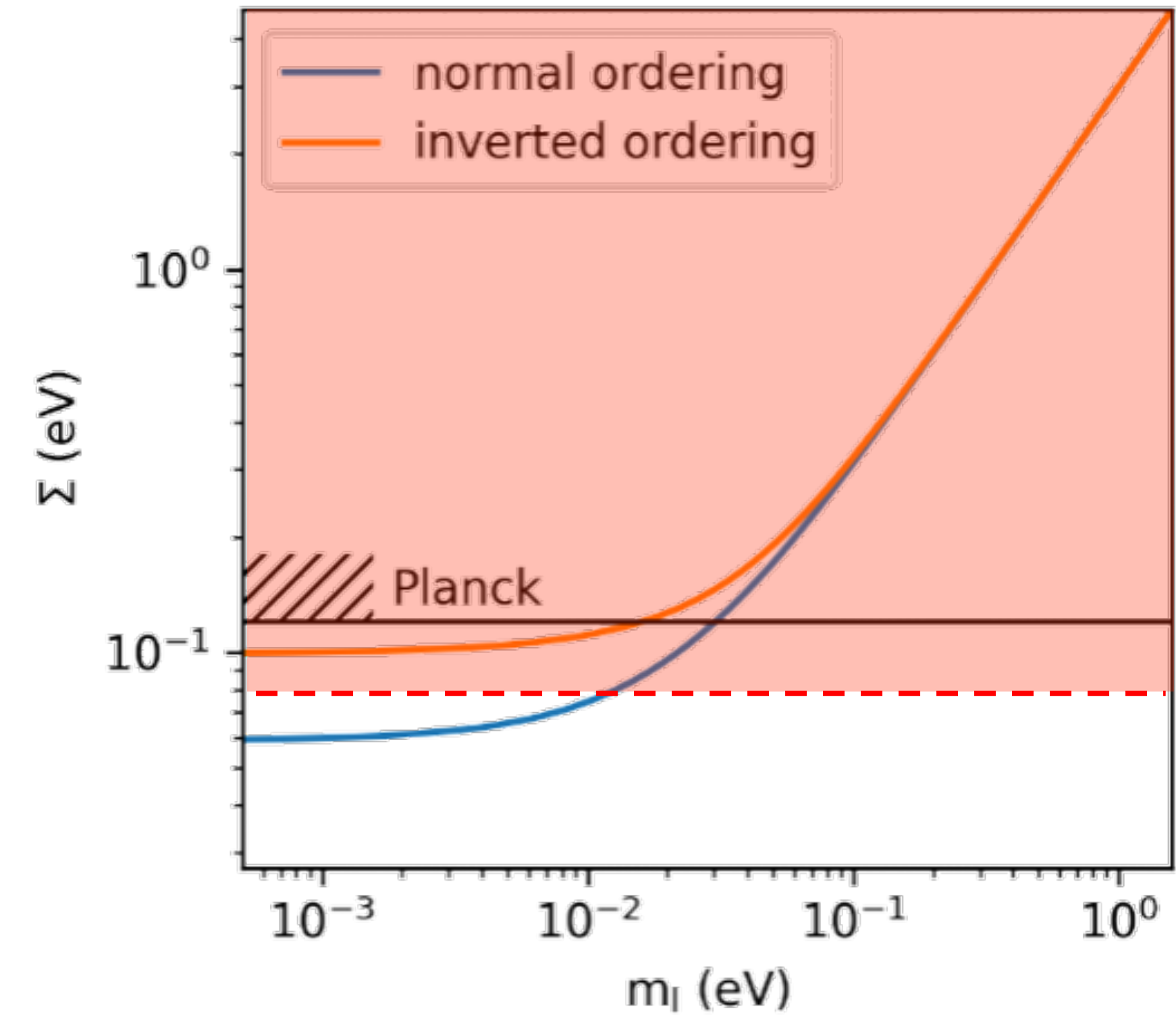
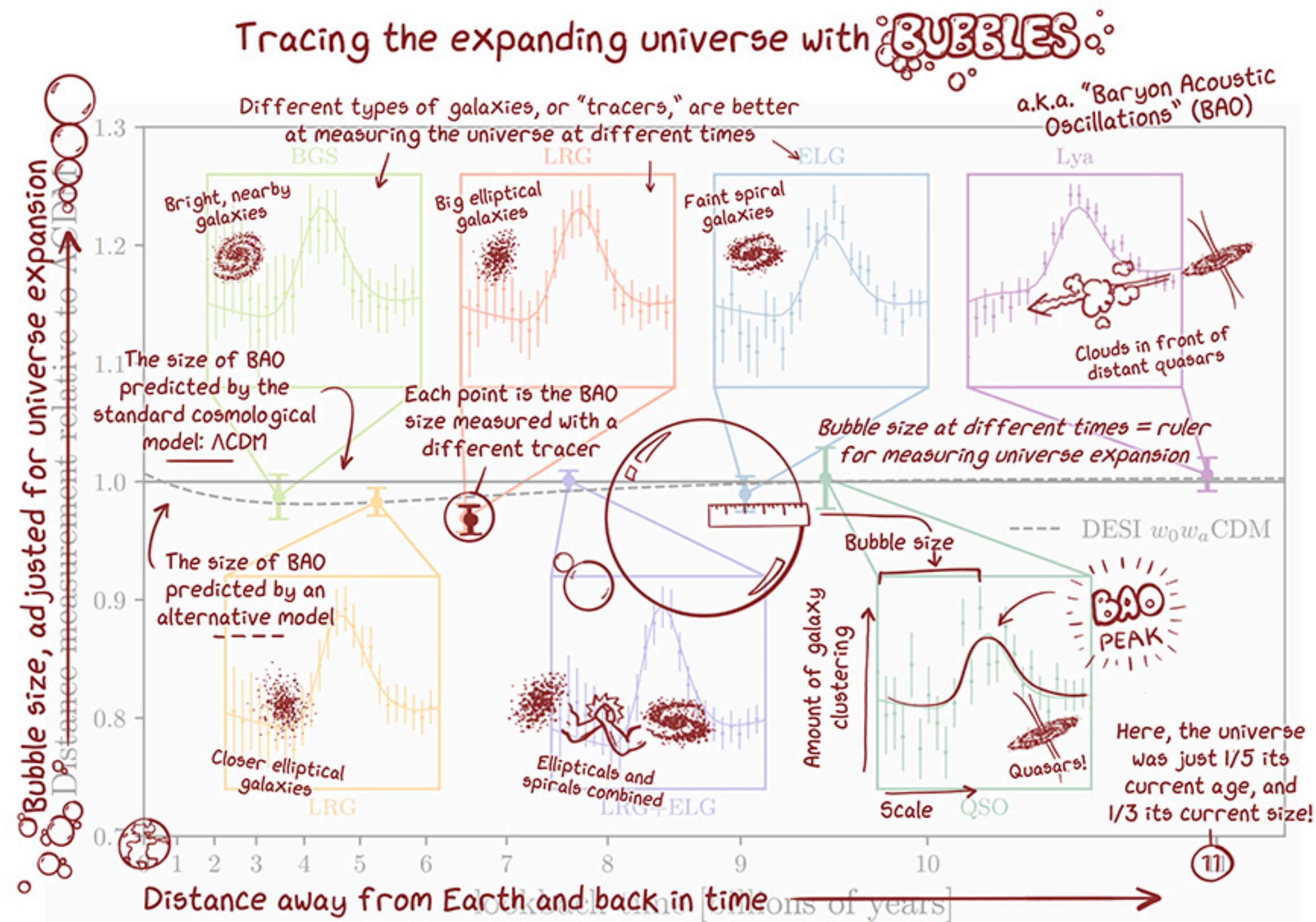


Argawal & Feldman, MNRAS 2011



# Cosmology

- 6 million spectra grouped in 7 different redshifts [0.4 < z < 4.2] (LGR, BGS, QSO, Ly $\alpha$ )
- Fully exploits the capabilities of DESI and increases sensitivity!



CMB + DESI (BAO) sets a limit at  $\Sigma m_\nu < 0.072$  eV

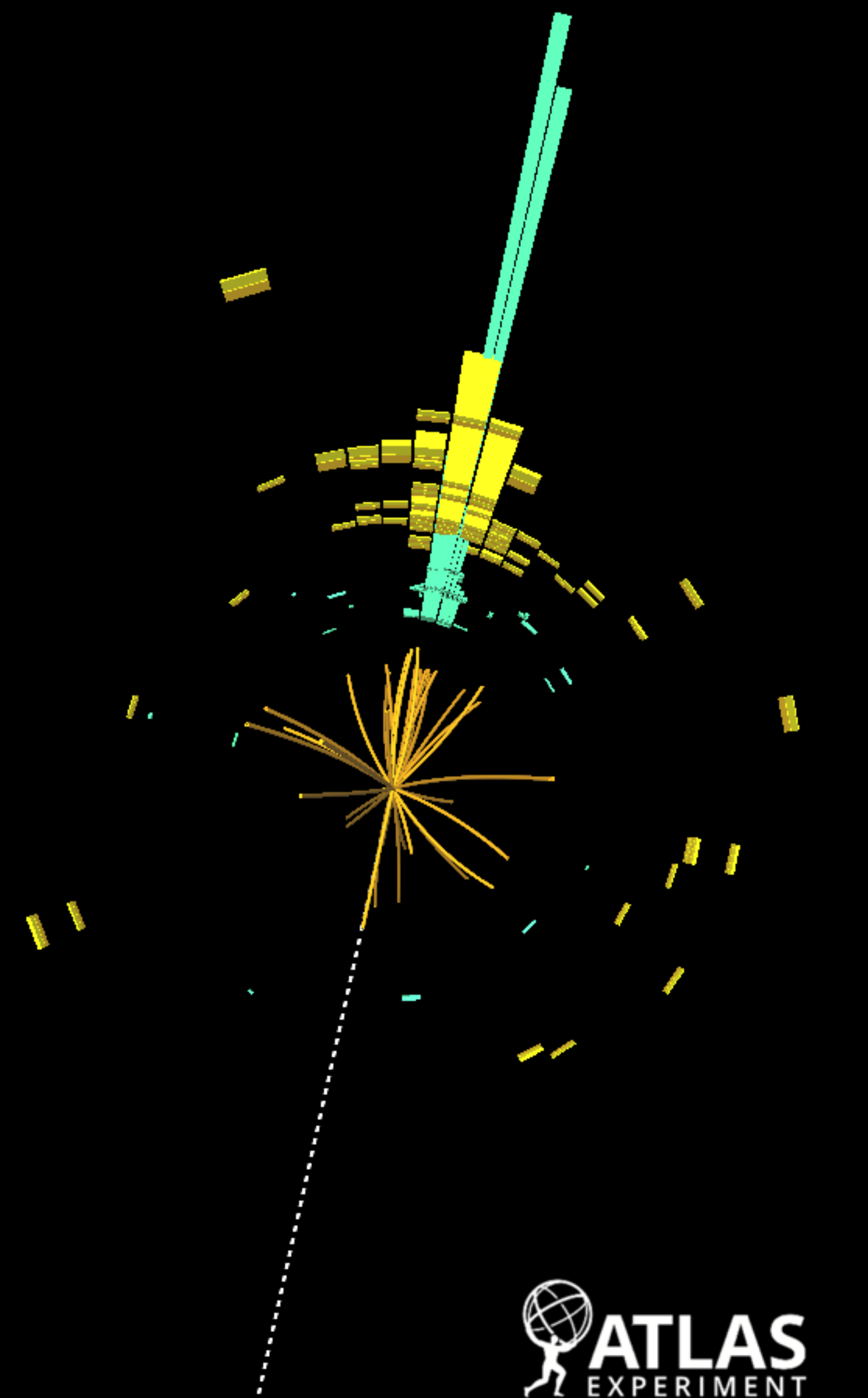
NO seems to be favoured, but this result is dependent on the cosmological model and priors adopted

# Direct Searches for new Physics at the LHC

Livia presented more than 20 new searches for new physics in extended higgs sectors, Dark sector, SUSY, Heavy fermions, and EFT. Impressive harvest of searches for new physics mostly at Run 2!

Leaving no stones unturned!

Reference	Topic	Experiment	Model	Explored energy range [GeV]
<a href="#">HDBS-2021-07</a>	$H \rightarrow aa \rightarrow bb\tau\tau$	ATLAS	Extended Higgs Sector	0 - 300
<a href="#">HDBS-2020-11</a>	$H^\pm \rightarrow cs$	ATLAS		0 - 300
<a href="#">HDBS-2023-19</a> <a href="#">HIG-24-002</a>	Combination of charged H $H \rightarrow ZZ \rightarrow 4l$	ATLAS CMS		0 - 3000
<a href="#">HIG-22-004</a>	$A \rightarrow Zh(\tau\tau)$	CMS		0 - 1200
<a href="#">SUS-24-001</a>	$\phi \rightarrow b\bar{b}$	CMS		0 - 1800
<a href="#">EXOT-2018-55</a>	Prompt Leptonjets	ATLAS	Dark Sector +ALPs	0 - 300
<a href="#">EXOT-2022-04</a>	Neutral LLP into displaced jets	ATLAS		0 - 1200 - displaced
<a href="#">SUS-23-004</a>	mono- $t$	CMS		0 - 1200 - dark matter
<a href="#">SUS-23-012</a>	mono- $h(\tau\tau)$	CMS		0 - 1200 - dark matter
<a href="#">SUS-23-018</a>	$H \rightarrow Za \rightarrow ll\chi\chi$	CMS		0 - 1800
<a href="#">SUS-24-004</a>	pMSSM	CMS	Supersymmetry	0 - 1200
<a href="#">SUS-23-003</a>	Compressed SUSY w/ RJR	CMS		0 - 300 - $\tilde{\Delta}_m$
<a href="#">ATLAS-2024-011</a>	Run3 displaced leptons	ATLAS		0 - 1200 - displaced
<a href="#">ATLAS-2024-008</a>	VLL $\rightarrow \tau b$	ATLAS	Heavy Fermions	0 - 1200
<a href="#">EXO-23-015</a>	VLL $\rightarrow \tau a(\gamma\gamma)$	CMS		0 - 1200 - displaced
<a href="#">B2G-22-005</a>	$t^* \rightarrow tg$	CMS		0 - 3000
<a href="#">EXO-23-010</a>	$ll + b - \text{jets, non - resonant}$	CMS	EFT	0 - 3000
<a href="#">EXO-24-007</a>	Low mass dijet+ISR	CMS	Z' Mediator	0 - 300
<a href="#">EXO-22-006</a>	$Z' \rightarrow \mu\mu + b - \text{jets, resonant}$	CMS		0 - 300

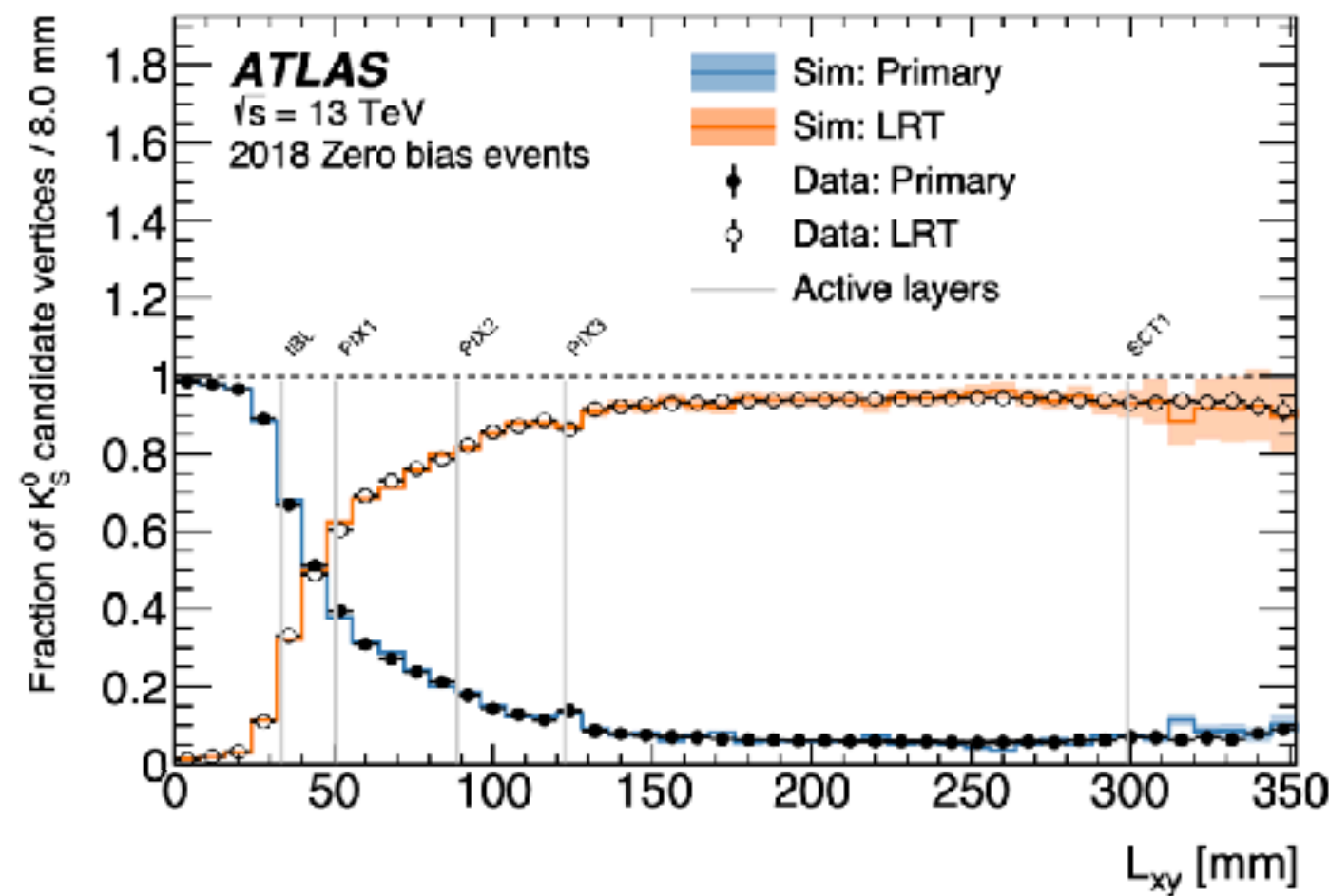


# Direct Searches at Colliders

Livia Soffi

## Run 3 search!

Improving reconstruction techniques e.g. ATLAS Large Radius Tracking at Run 3 and reprocessed Run 2!

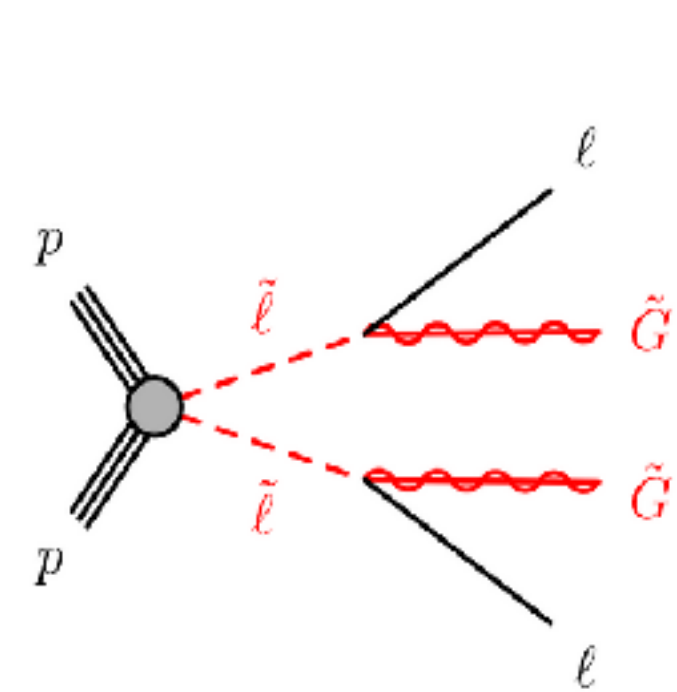


LRT performance tested with Ks reconstruction ([Paper](#))

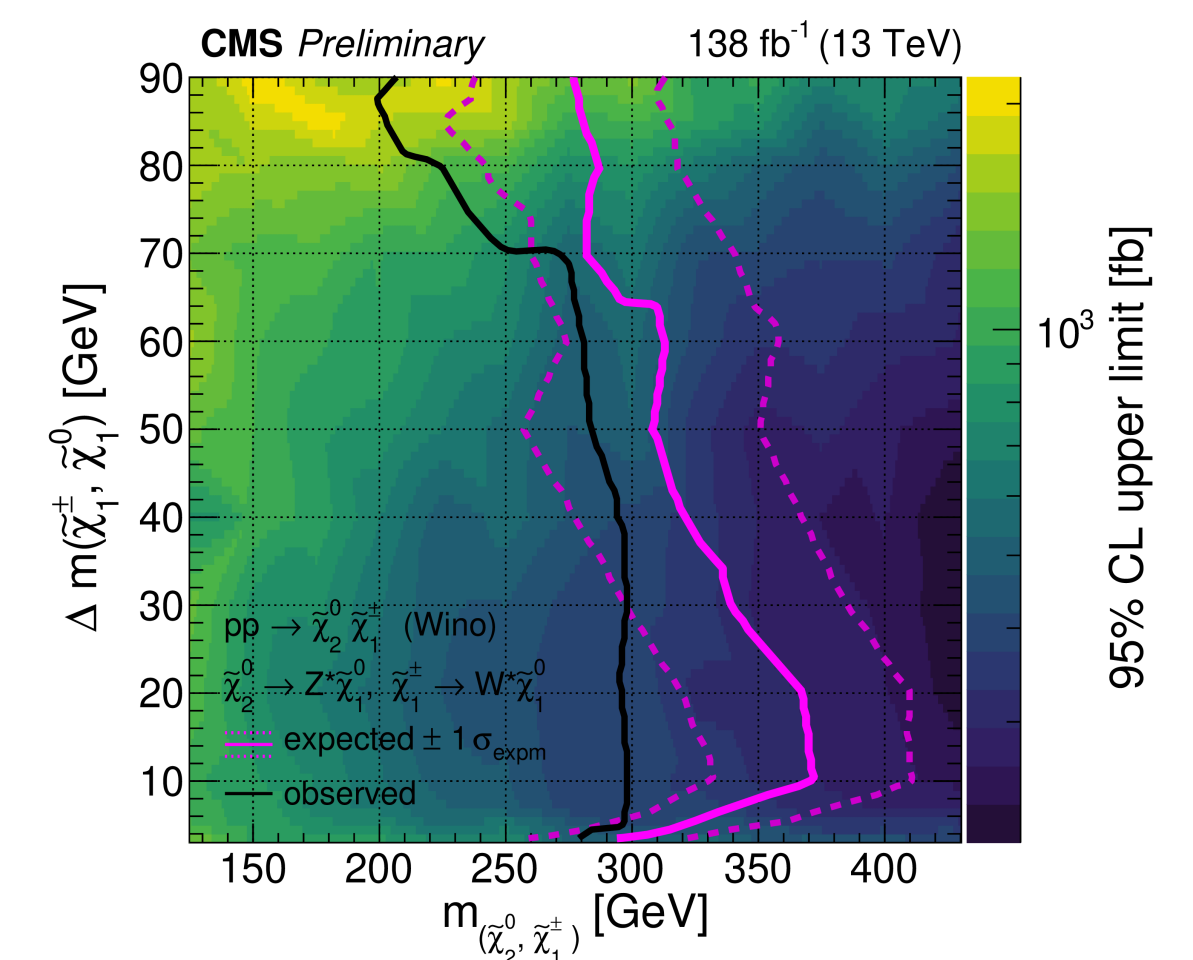
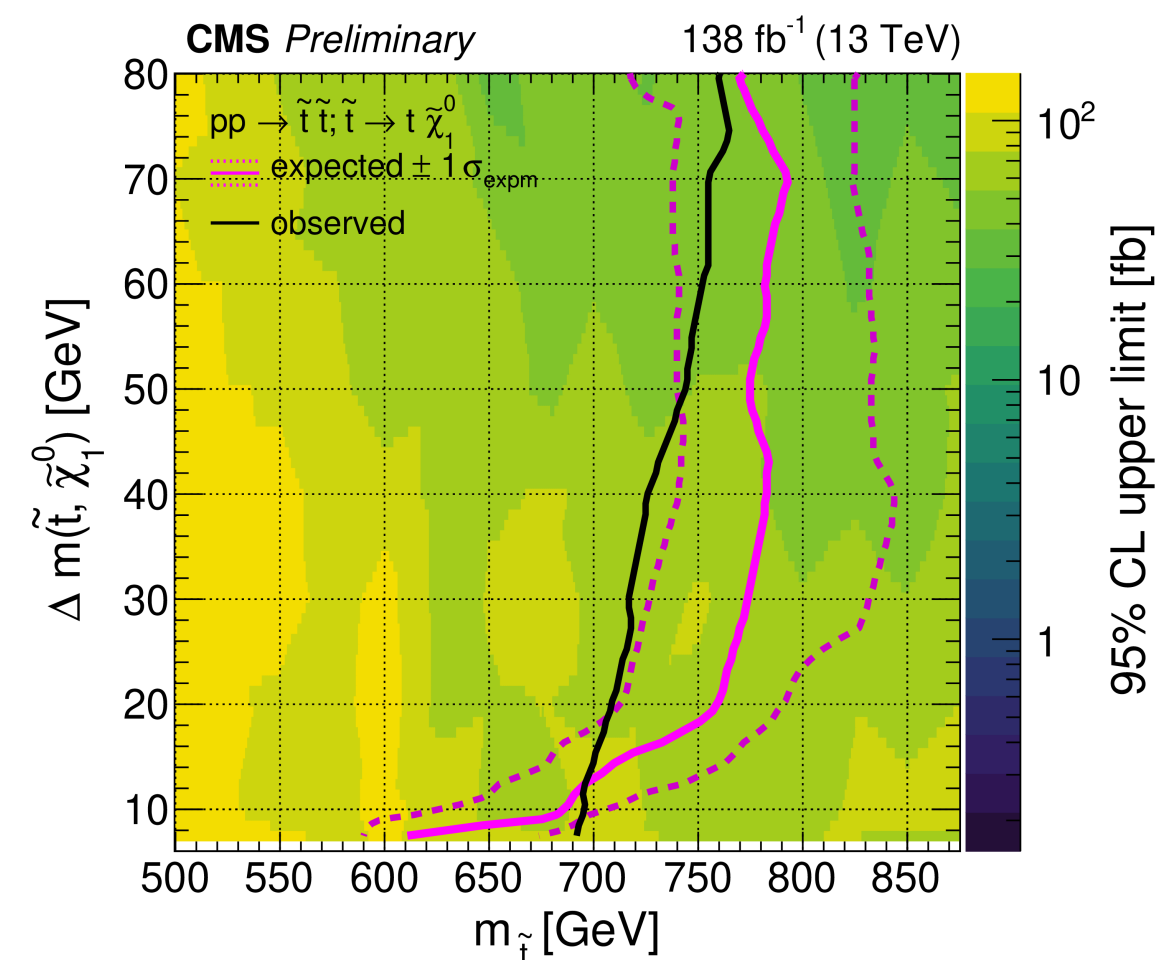
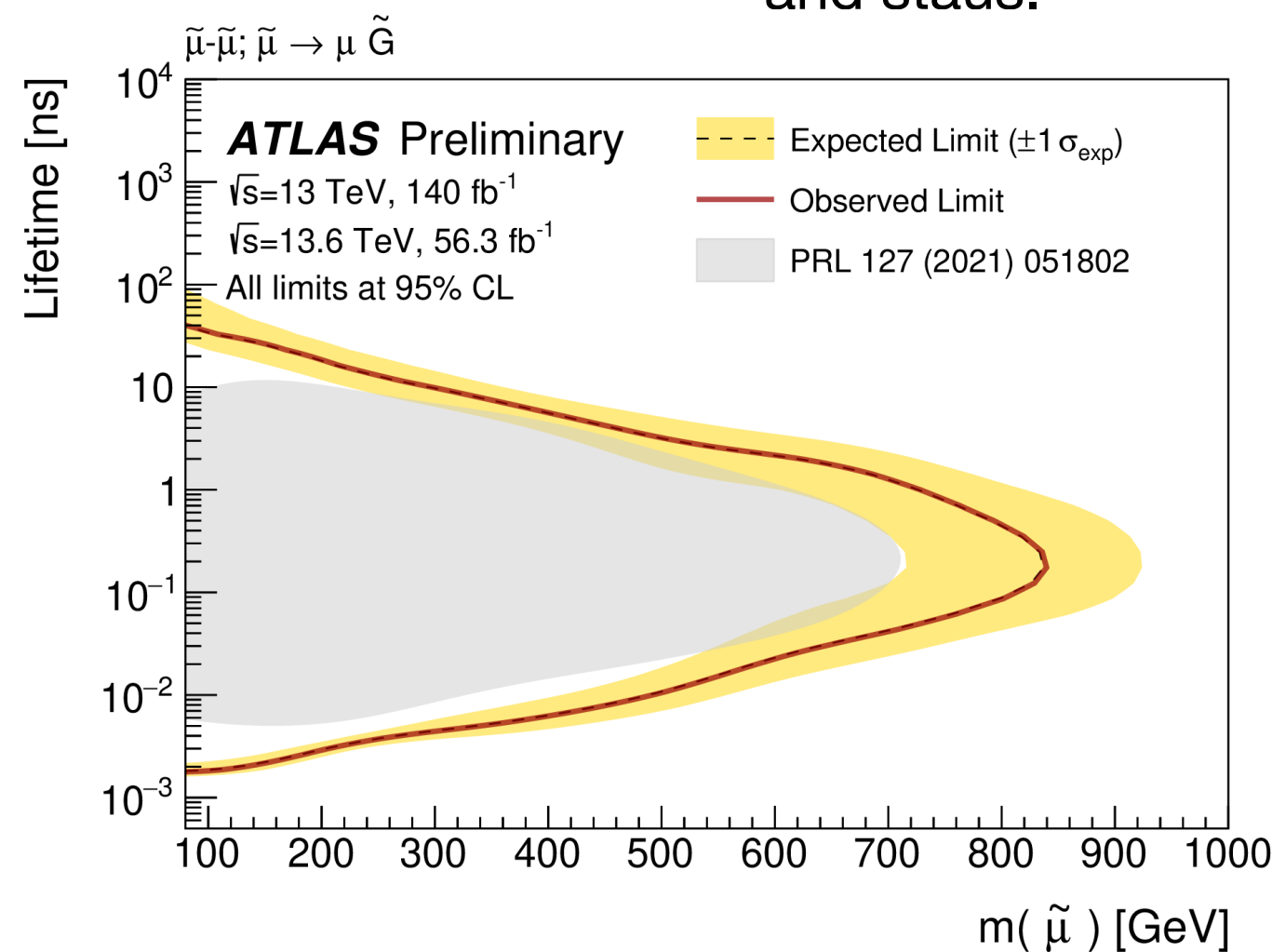
Search done for smuons, selectron and staus.

Searches for EWK production SUSY compressed scenarios

Wide range of signatures targeting electroweakinos, sleptons, and top squarks with focus on events with a high transverse momentum system from initial-state-radiation jets and significant missing transverse momentum.

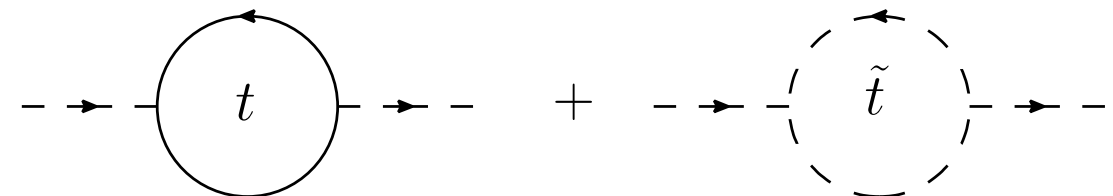
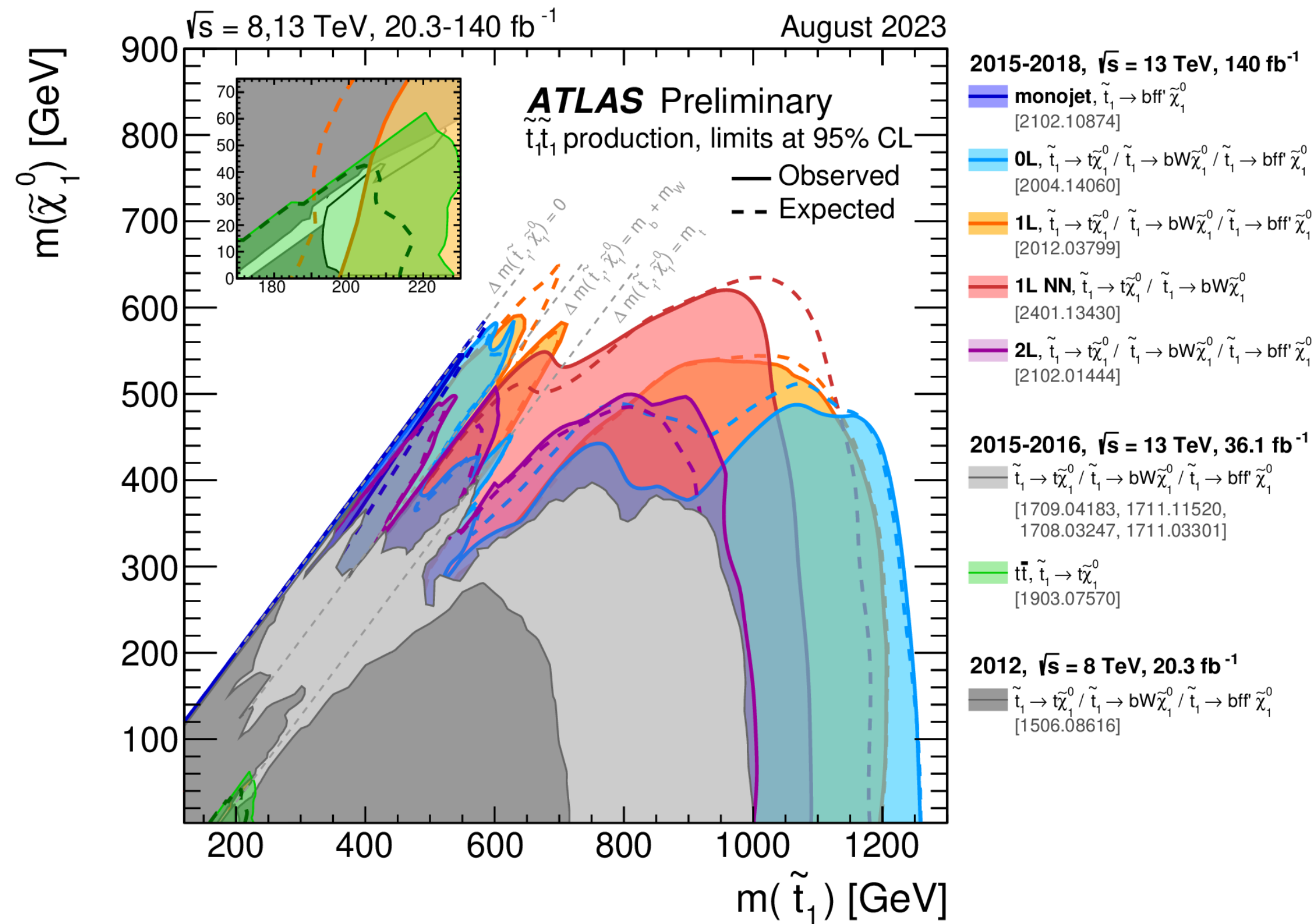


Long lifetime due to the small coupling to the low mass gravitino!



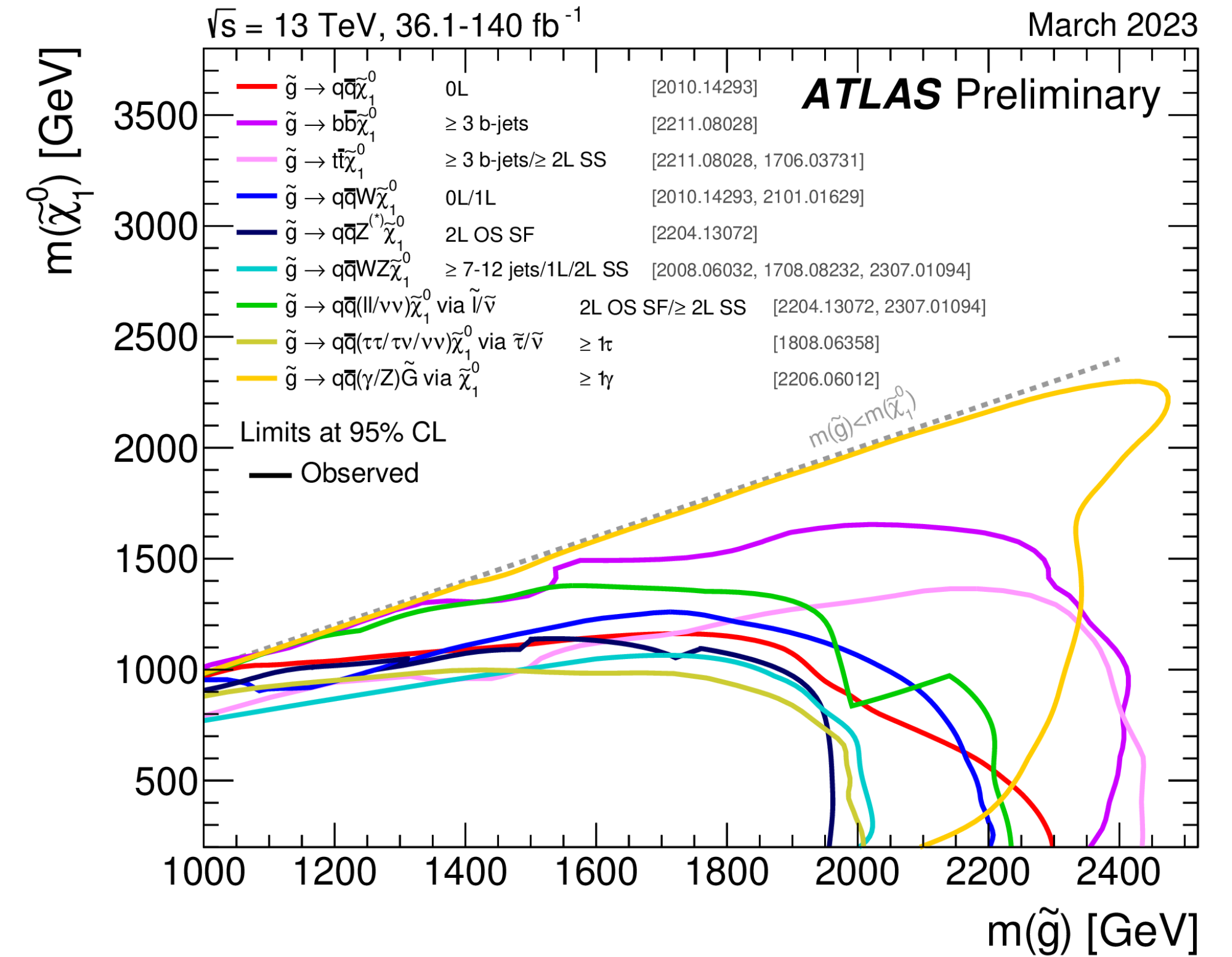
# Constraining Naturalness (SUSY) Scenarios

## Stop searches

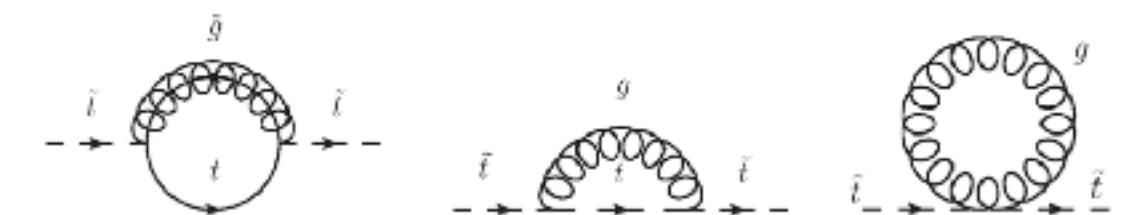


**Not so natural SUSY:** Stops > 1.2 TeV ~Tuning of factor **20**, but these exclusions are under specific conditions, and there are unexcluded corridors.

## Glino searches



**Stop also a scalar requires light gluinos** to be light enough: for gluinos > 2.4 TeV ~tuning of Factor of **30**





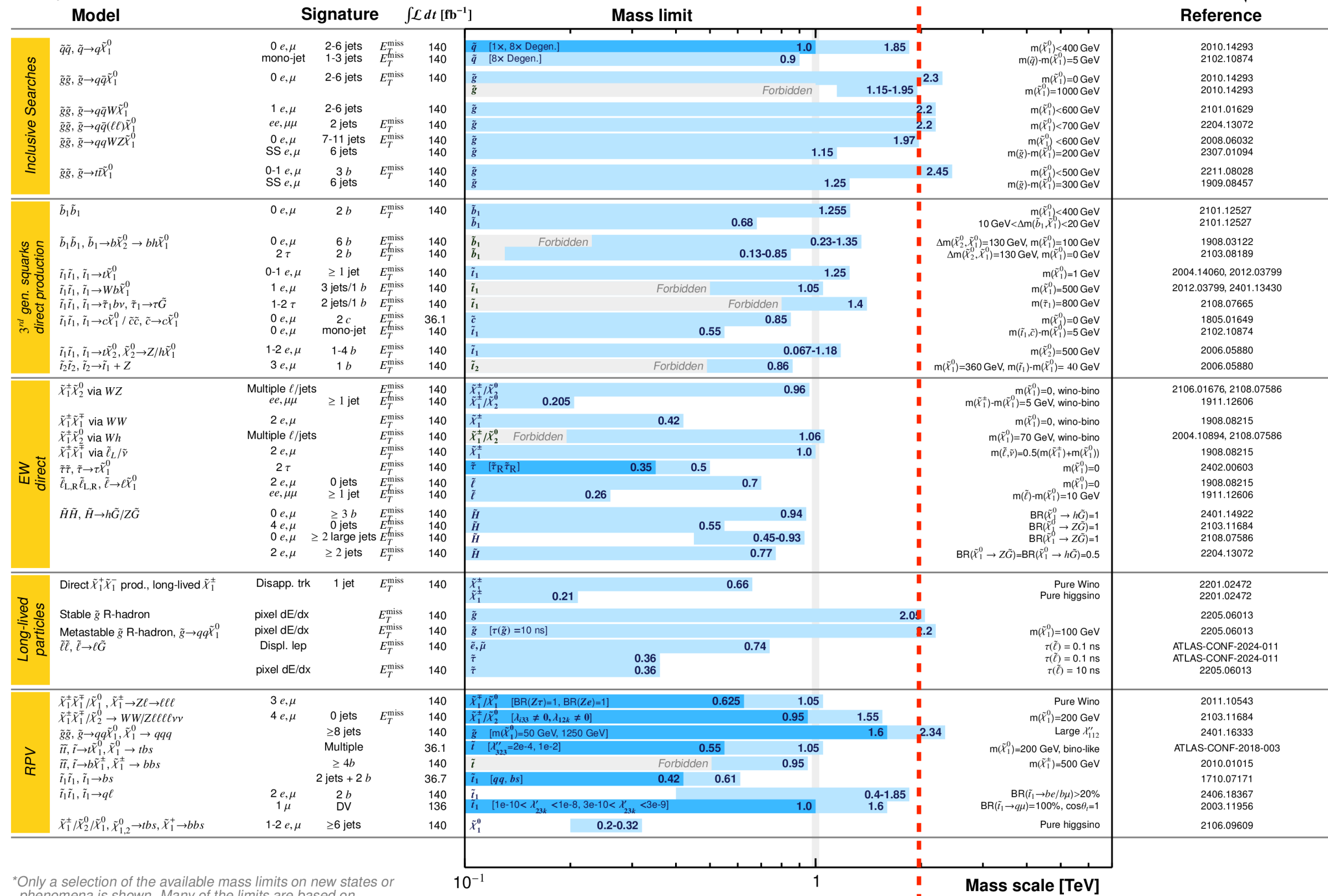
# Very Large Number of SUSY Searches

(in large variety of topologies and models)

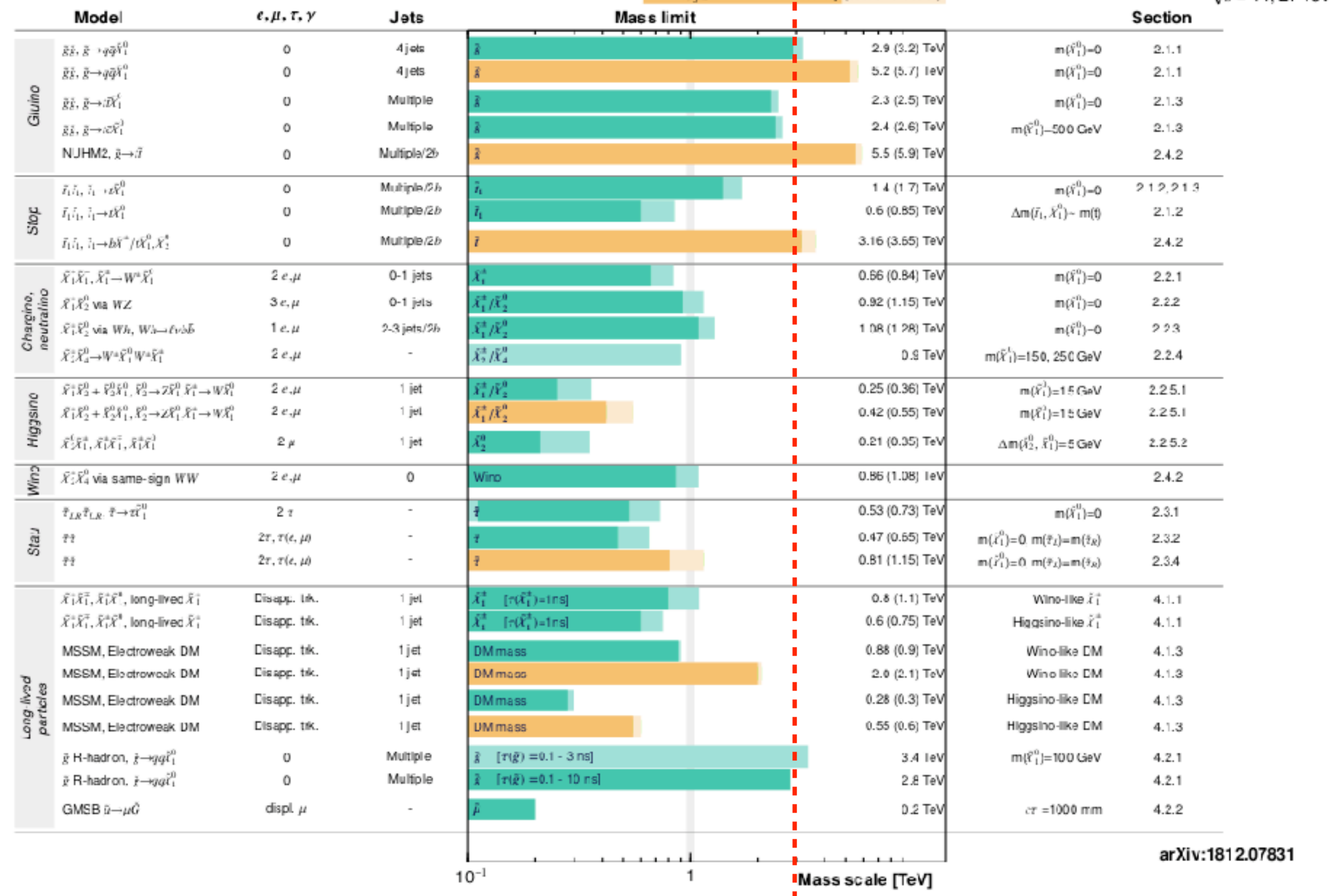
Livia Soffi

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ATLAS SUSY Searches\* - 95% CL Lower Limits  
July 2024



HL/HE-LHC SUSY Searches



\*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

arXiv:1812.07831

2 TeV

3 TeV

Example from ATLAS (similar for CMS)

HL-LHC YR  
1812.07831

# Very Large Number of SUSY Searches

(in large variety of topologies and models)

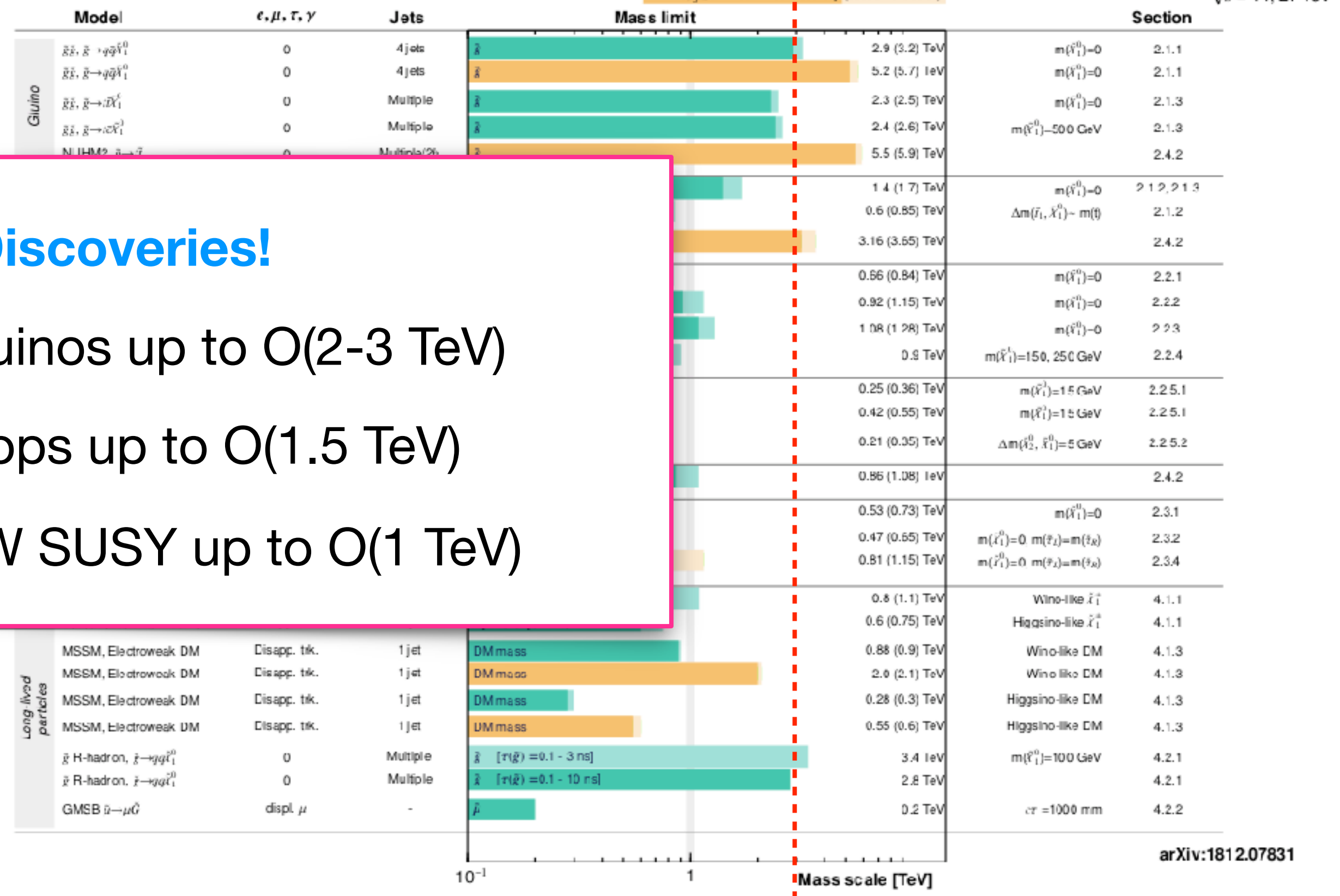
ATLAS SUSY Searches\* - 95% CL Lower Limits  
July 2024

Model	Signature	$\int L dt [fb^{-1}]$	Mass limit	Reference						
Inclusive Searches	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0 e, $\mu$	2-6 jets	$E_T^{miss}$	140	$\tilde{q}$ [1x, 8x Degen.]	1.0	1.85	$m(\tilde{\chi}_1^0) < 400$ GeV	2010.14293
	$\tilde{q}$	mono-jet	1-3 jets	$E_T^{miss}$	140	$\tilde{q}$ [8x Degen.]	0.9		$m(\tilde{q}) - m(\tilde{\chi}_1^0) = 5$ GeV	2102.10874
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0 e, $\mu$	2-6 jets	$E_T^{miss}$	140	$\tilde{g}$	Forbidden	2.3	$m(\tilde{\chi}_1^0) = 0$ GeV	2010.14293
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}W\tilde{\chi}_1^0$	1 e, $\mu$	2-6 jets		140	$\tilde{g}$		2.2	$m(\tilde{\chi}_1^0) < 600$ GeV	2101.01629
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(t\bar{t})\tilde{\chi}_1^0$	ee, $\mu\mu$	2 jets	$E_T^{miss}$	140	$\tilde{g}$		2.2	$m(\tilde{\chi}_1^0) < 700$ GeV	2204.13072
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}WZ\tilde{\chi}_1^0$	0 e, $\mu$	7-11 jets	$E_T^{miss}$	140	$\tilde{g}$		1.97	$m(\tilde{\chi}_1^0) < 600$ GeV	2008.06032
3 <sup>rd</sup> gen. squarks direct production	$\tilde{b}_1\tilde{b}_1$	0 e, $\mu$	2 b	$E_T^{miss}$	140	$\tilde{b}_1$	Forbidden		$m(\tilde{\chi}_1^0) = 200$ GeV	2307.01094
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0 \rightarrow b h\tilde{\chi}_1^0$	0 e, $\mu$	6 b	$E_T^{miss}$	140	$\tilde{b}_1$	Forbidden			
EW direct	$\tilde{\chi}_1^{\pm}\tilde{\chi}_2^0$ via WZ	Multiple $\ell$ /jets	$\geq 1$ jet	$E_T^{miss}$	140	$\tilde{\chi}_1^{\pm}/\tilde{\chi}_2^0$	0.205			
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}$ via WW	2 e, $\mu$		$E_T^{miss}$	140	$\tilde{\chi}_1^{\pm}/\tilde{\chi}_1^{\mp}$	Forbidden			
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_2^0$ via Wh	Multiple $\ell$ /jets		$E_T^{miss}$	140	$\tilde{\chi}_1^{\pm}/\tilde{\chi}_2^0$	Forbidden			
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}$ via $\tilde{\ell}_R/\nu$	2 e, $\mu$		$E_T^{miss}$	140	$\tilde{\chi}_1^{\pm}/\tilde{\chi}_1^{\mp}$	Forbidden			
	$\tilde{\tau}\tilde{\tau}, \tilde{\tau} \rightarrow \tau\tilde{\chi}_1^0$	2 $\tau$		$E_T^{miss}$	140	$\tilde{\tau}$ [FR $\tau$ R]	0.3			
	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 e, $\mu$	0 jets	$E_T^{miss}$	140	$\tilde{\ell}$	0.26			
Long-lived particles	Direct $\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}$ prod., long-lived $\tilde{\chi}_1^{\pm}$	Disapp. trk	1 jet	$E_T^{miss}$	140	$\tilde{\chi}_1^{\pm}$	0.21			
	Stable $\tilde{g}$ R-hadron	pixel dE/dx		$E_T^{miss}$	140	$\tilde{g}$			$\tau(\tilde{g}) = 0.1$ ns	ATLAS-CONF-2024-011
	Metastable $\tilde{g}$ R-hadron, $\tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	pixel dE/dx		$E_T^{miss}$	140	$\tilde{g}$ [ $\tau(\tilde{g}) = 10$ ns]			$\tau(\tilde{g}) = 0.1$ ns	ATLAS-CONF-2024-011
RPV	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}/\tilde{\chi}_1^0/\tilde{\chi}_2^0, \tilde{\chi}_1^{\pm} \rightarrow Z\ell\ell$	3 e, $\mu$		$E_T^{miss}$	140	$\tilde{\chi}_1^{\pm}/\tilde{\chi}_1^0/\tilde{\chi}_2^0$ [BR(Z $\tau$ )=1, BR(Z $e$ )=1]	0.625	1.05	Pure Wino	2011.10543
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}/\tilde{\chi}_1^0/\tilde{\chi}_2^0 \rightarrow WW/Z\ell\ell\nu\nu$	4 e, $\mu$	0 jets	$E_T^{miss}$	140	$\tilde{\chi}_1^{\pm}/\tilde{\chi}_1^0/\tilde{\chi}_2^0$ [ $A_{13} \neq 0, A_{124} \neq 0$ ]	0.95	1.55	$m(\tilde{\chi}_1^0) = 200$ GeV	2103.11684
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow q\tilde{q}q$	Multiple	$\geq 8$ jets	$E_T^{miss}$	140	$\tilde{g}$ [ $m(\tilde{\chi}_1^0) = 50$ GeV, 1250 GeV]		1.6	Large $A'_{12}$	2401.16333
	$\tilde{t}\tilde{t}, \tilde{t} \rightarrow b\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow tbs$	Multiple	$\geq 4b$	$E_T^{miss}$	36.1	$\tilde{t}$ [ $A'_{23} = 2e-4, 1e-2$ ]	0.55	1.05	$m(\tilde{\chi}_1^0) = 200$ GeV, bino-like	ATLAS-CONF-2018-003
	$\tilde{t}\tilde{t}, \tilde{t} \rightarrow b\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow bbs$	Multiple	$\geq 4b$	$E_T^{miss}$	140	$\tilde{t}$	Forbidden	0.95	$m(\tilde{\chi}_1^0) = 500$ GeV	2010.01015
	$\tilde{t}\tilde{t}, \tilde{t} \rightarrow bs$	2 jets + 2 b		$E_T^{miss}$	36.7	$\tilde{t}$ [ $qq, bs$ ]	0.42	0.61		1710.07171
RPV	$\tilde{t}\tilde{t}, \tilde{t} \rightarrow q\ell$	2 e, $\mu$	2 b	$E_T^{miss}$	140	$\tilde{t}$		0.4-1.85	BR( $\tilde{t} \rightarrow b\ell/h\nu$ ) > 20%	2406.18367
	$\tilde{t}\tilde{t}, \tilde{t} \rightarrow q\ell$	1 $\mu$	DV	$E_T^{miss}$	136	$\tilde{t}$		1.6	BR( $\tilde{t} \rightarrow q\mu$ ) = 100%, $\cos\theta = 1$	2003.11956
	$\tilde{\chi}_1^0/\tilde{\chi}_2^0/\tilde{\chi}_1^0/\tilde{\chi}_1^0 \rightarrow tbs, \tilde{\chi}_1^0 \rightarrow bbs$	1-2 e, $\mu$	$\geq 6$ jets		140	$\tilde{\chi}_1^0$	0.2-0.32		Pure higgsino	2106.09609

\*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

ATLAS Preliminary  
 $\sqrt{s} = 13$  TeV

HL/HE-LHC SUSY Searches



Simulation Preliminary  
 $\sqrt{s} = 14, 27$  TeV

2 TeV

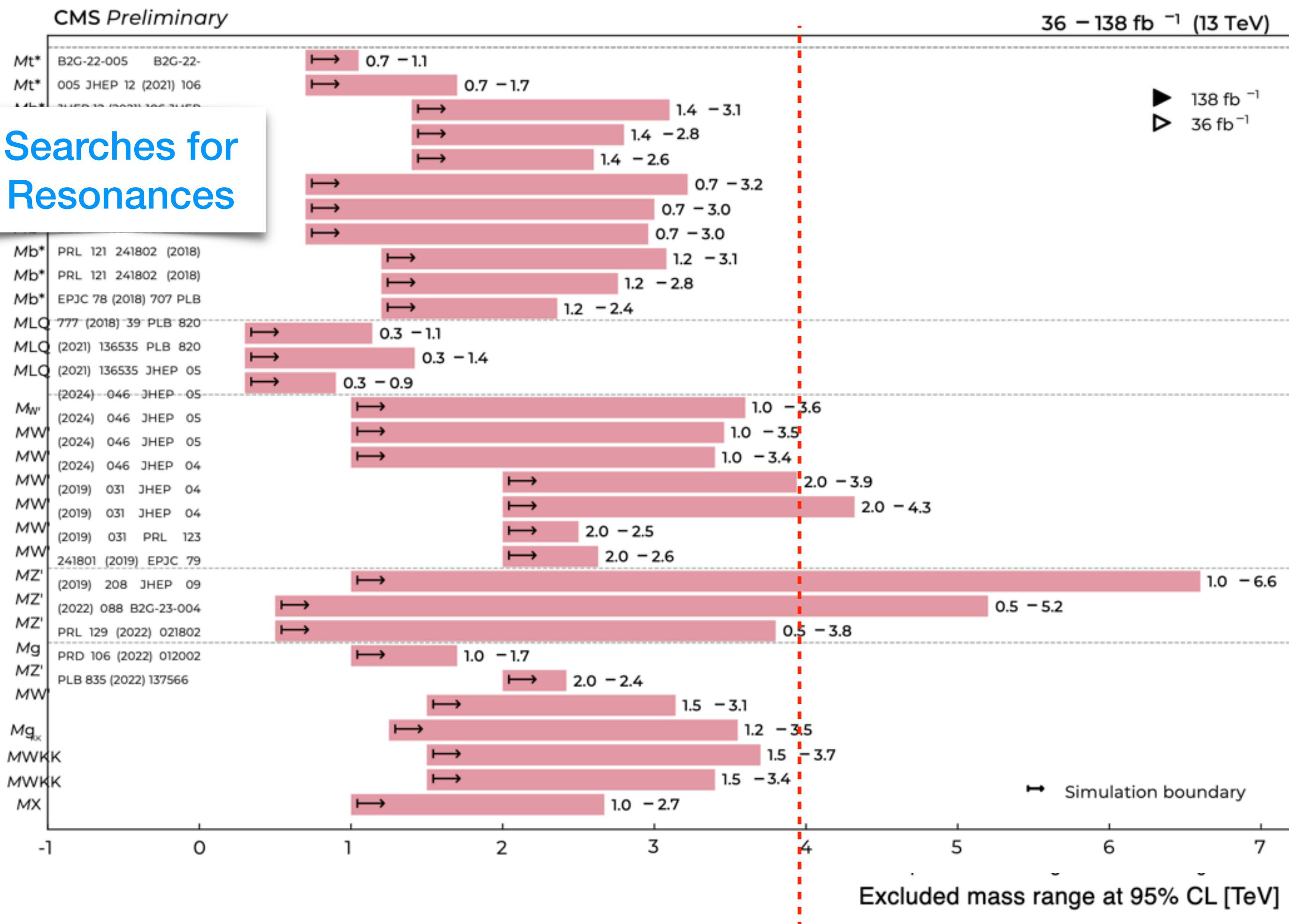
3 TeV

Example from ATLAS (similar for CMS)

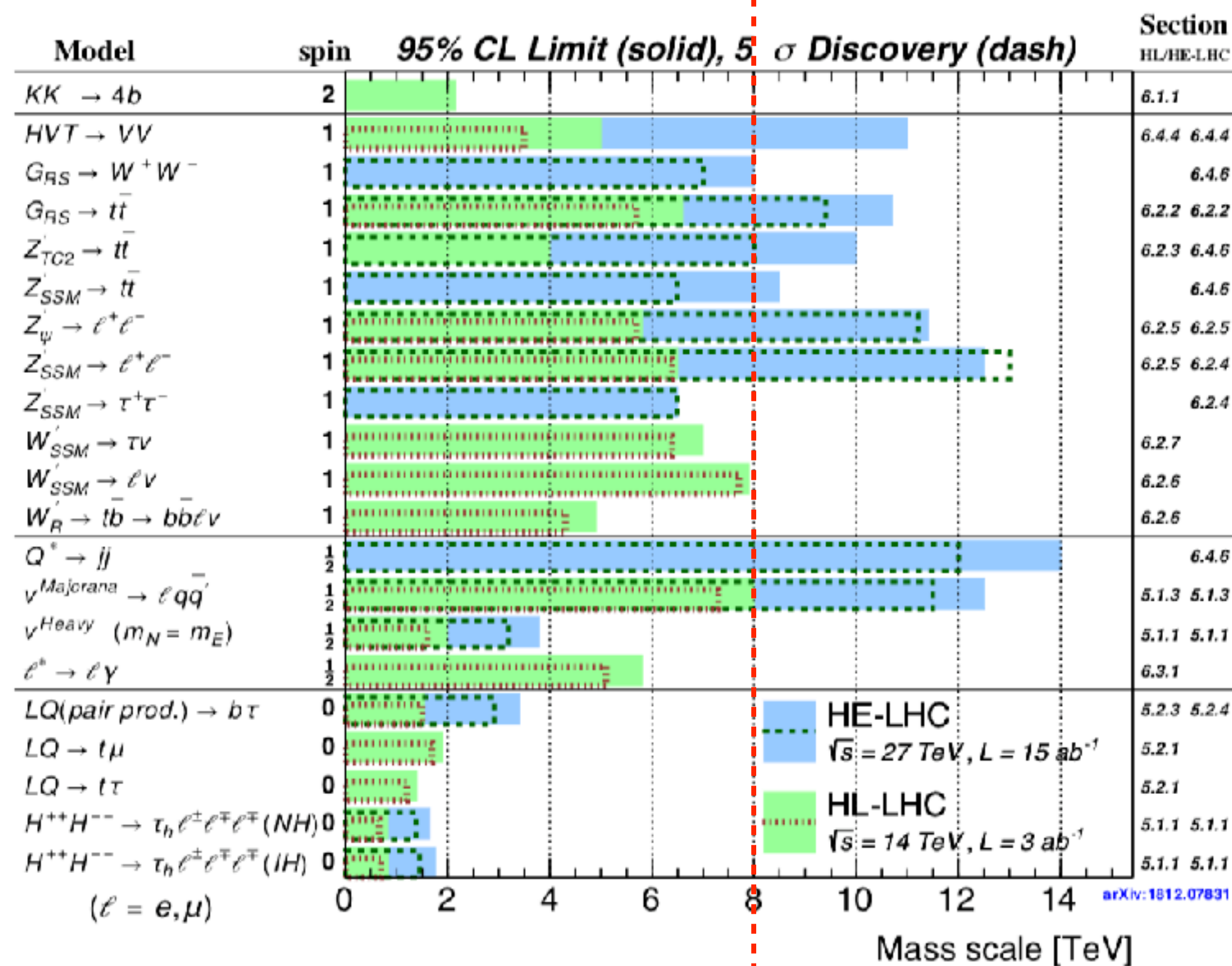
# Very Large Number of Searches

(in large variety of topologies and models)

Searches for Resonances



4 TeV



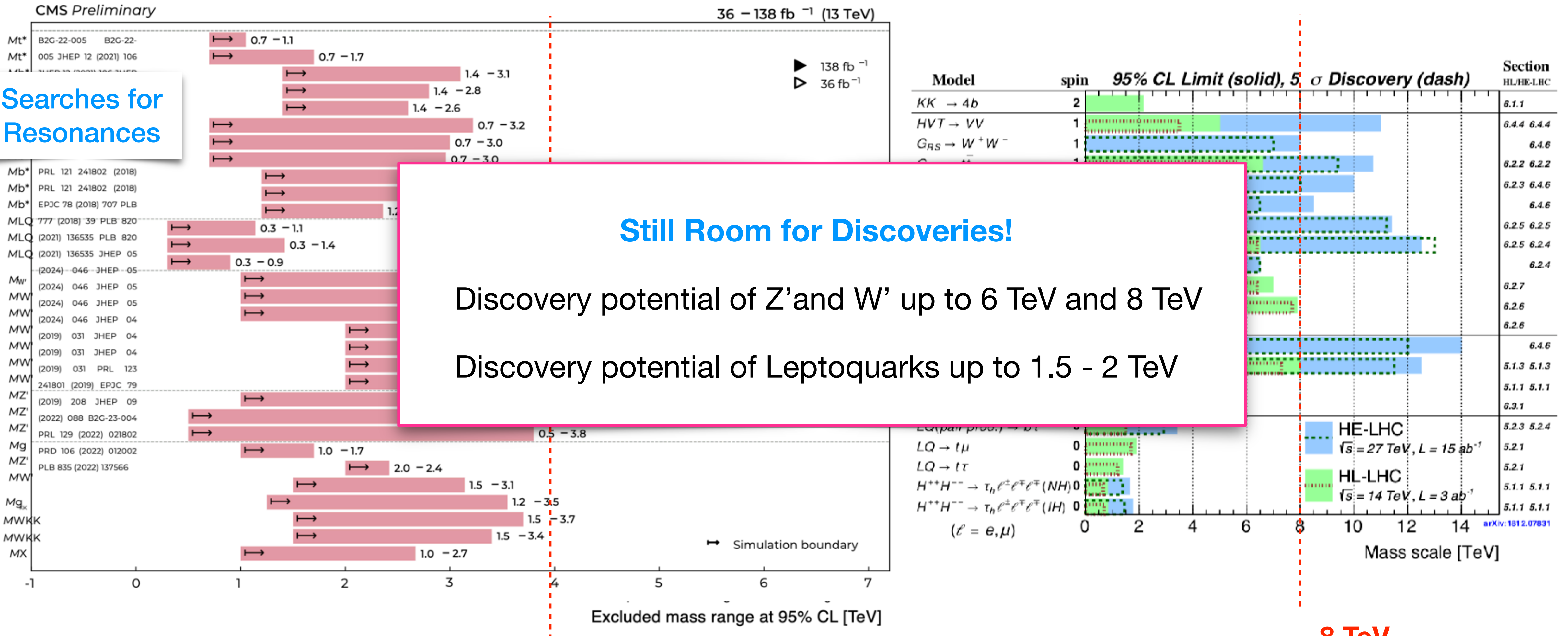
8 TeV

Example from CMS (similar for ATLAS) - latest plot in the backup!

# Very Large Number of Searches

(in large variety of topologies and models)

Searches for Resonances



4 TeV

8 TeV

Example from CMS (similar for ATLAS) - latest plot in the backup!



# A Scientific Mission for the 21st Century

Rende Steerenberg

## LHC Run 2

2014-2018 13 TeV  
100% to 2x Nom. Lumi, PU 40  
Int. Lumi. 190 fb<sup>-1</sup>

Higgs couplings to Fermions of the third generation (top, bottom and taus)!

## LS2

2018-2022  
Experiments Phase-I and accelerator upgrades

## HL-LHC (Runs 4-6)

2029-2041 13.6 - 14 TeV and 2x Nominal Luminosity, PU 140 - 200  
Int. Lumi. 3000 fb<sup>-1</sup>

di-Higgs boson production and Higgs self coupling and precision Higgs physics!

CLIC 380 GeV- 3 TeV

ILC 250 GeV - 1 TeV

Cool Copper Collider 250 - 550 GeV

2010

2020

2030

2040

2050

2060

2070

## LS1

2012-2014  
Consolidation of LHC interconnections

## LHC Run 1

2009-2012 7-8 TeV  
75% Nom. Lumi, PU 30-40  
Int. Lumi. 30 fb<sup>-1</sup>

Discovery of the Higgs Boson, measurements of Higgs Boson couplings to bosons (gluons, photons, W and Z)

## LS3

2026-2029 HL-LHC installation and major exp. upgrades

## LHC Run 3

2022-2026 13.6 TeV  
2x Nom. Lumi., PU 60  
Int. Lumi. 450 fb<sup>-1</sup>

Higgs couplings to Fermions of the second generation (muons) and more rare decays

FCC-ee 90 - 265 GeV

FCC-hh 100 TeV

CepC 90 - 240 GeV

SppC

Muon Collider

LHC

Ultimate Precision  $e^+e^-$

Ultimate Energy (pp,  $\mu^+\mu^-$ )

# Future Large Collider Projects



Hitoshi Murayama

**The field is facing a defining moment in its history!**

**P5 - Reveal the secrets of the Higgs boson and much more... ultimate EW precision factory!**

“We would not consider the theory of electromagnetism established if we had only verified the strength of electromagnetic forces to within 10% accuracy.” (Salam, Wang, Zanderighi, [Nature](#))

**P5 - A realistic path to a 10 TeV parton center-of-momentum (pCM) collider.**

Closing quotes from the panel discussion ([video](#))

**Yifang Wang:** “In the Future, not very far from now there will be a Higgs factory!”

**Lia Merging:** “Need to continue to be bold and ambitious and dream big!”

**Fabiola Gianotti:** “As we have seen at this conference the field is extremely vibrant and exciting!”



# Conference Summary and Highlights



thank you

Many thanks to all participants for the excellent discussions and great atmosphere!

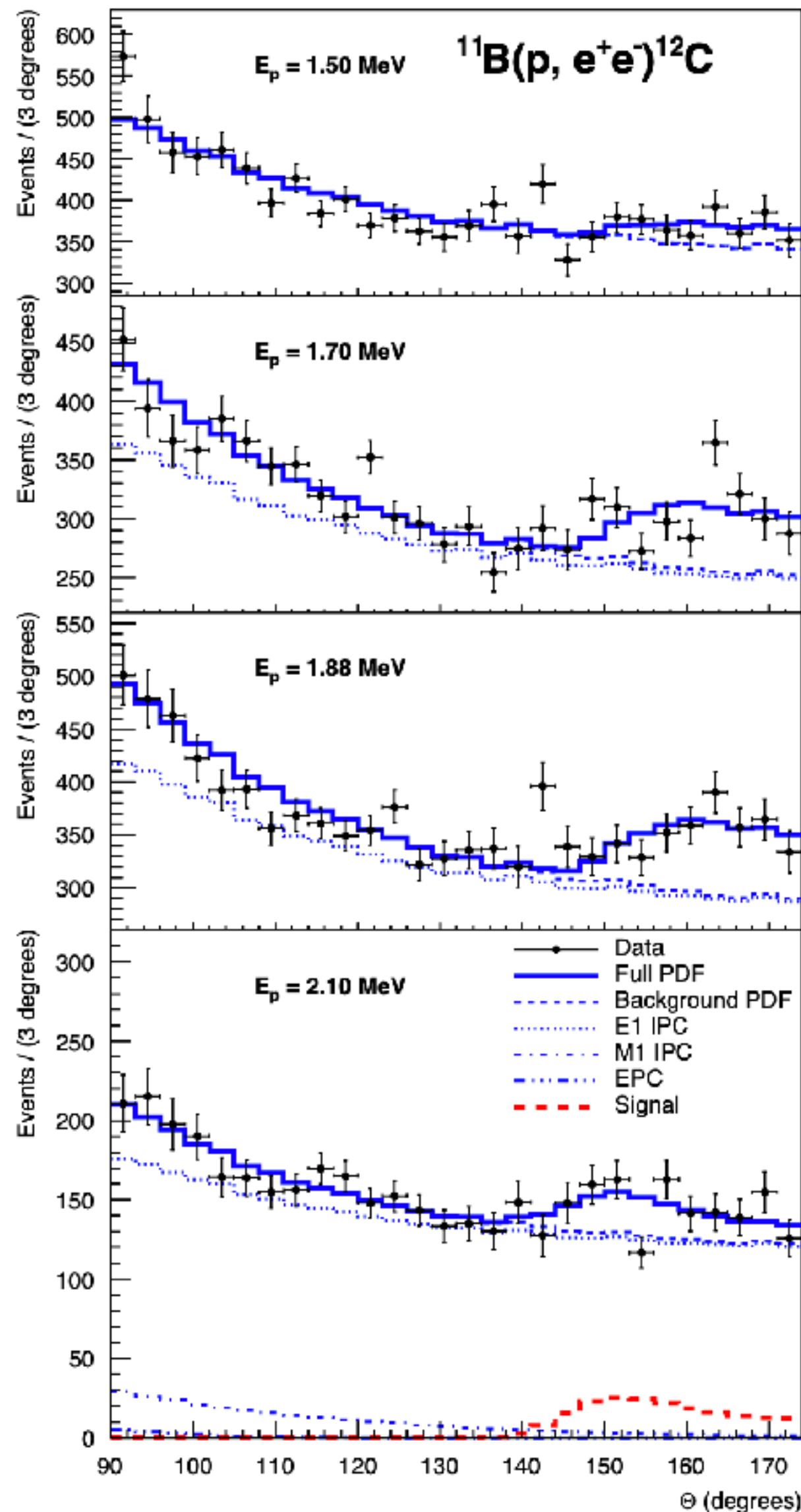
Many thanks and congratulations to all the speakers and collaborators for the splendid talks and results!

Many thanks to the organisers for the amazing conference in this amazing place! The organisation was perfect!

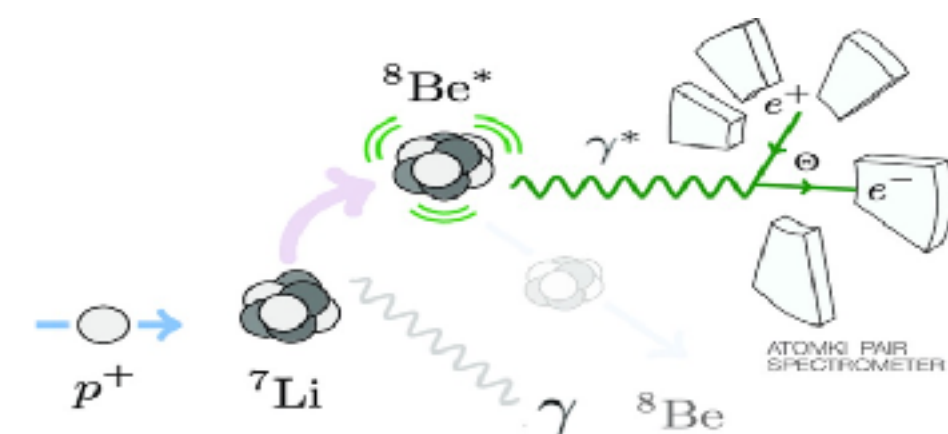


# Backup

# $^8\text{Be}$ Anomaly and PADME

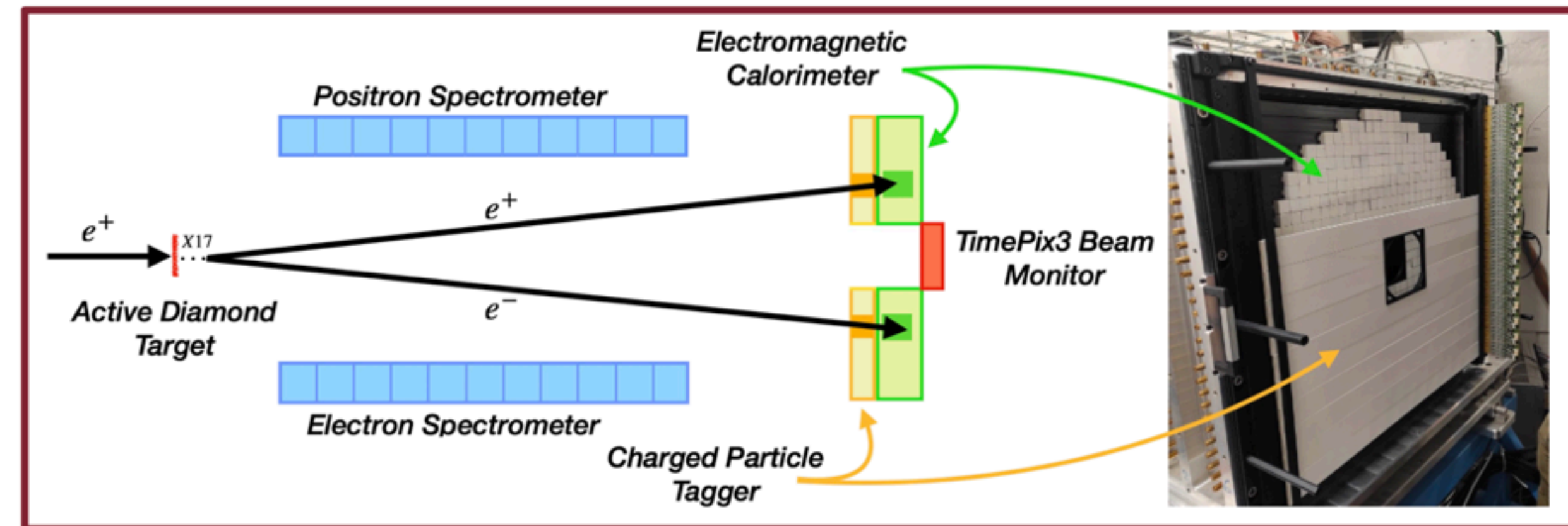


The ATOMKI institute observes the long standing  $^8\text{Be}$  anomaly, observed also in  $^4\text{He}$  and  $^{12}\text{C}$ , i.e. a significant excess compatible with new particle of 17 MeV mass.



DAΦNE Beam Test Facility is the only facility in the world with a positron beam at 282 MeV (yielding 17 MeV centre-of-mass collisions with fixed target electron!)

**PADME experiment** (Positron Annihilation into Dark Matter Experiment)



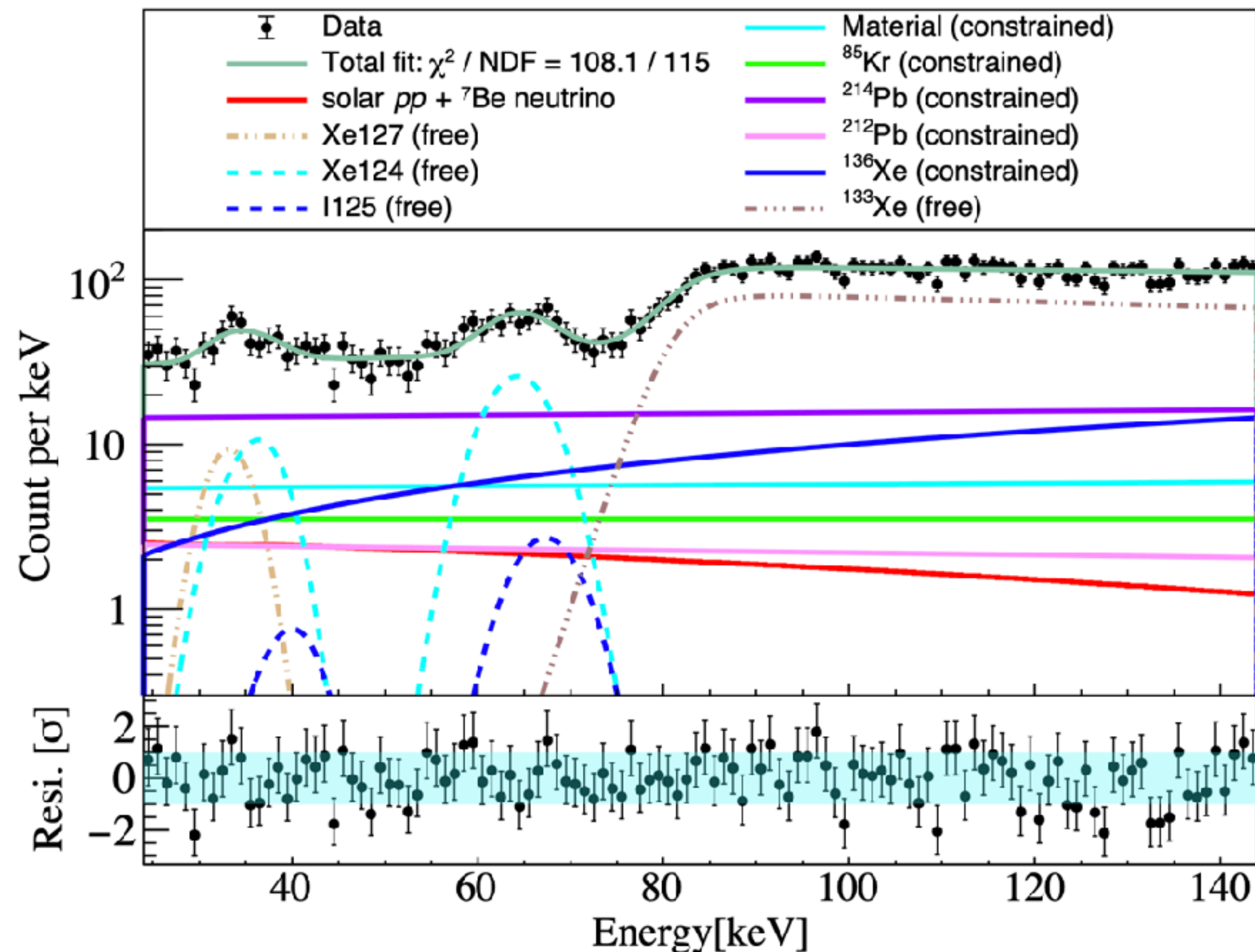
Run has finished and data analysis is ongoing, hoping to shine light on  $^8\text{Be}$

**Data not unblinded yet!**

# Solar Neutrinos in the Electron Recoils

Challenge #2: Search for WIMPs to the Neutrino Background in Direct Detection

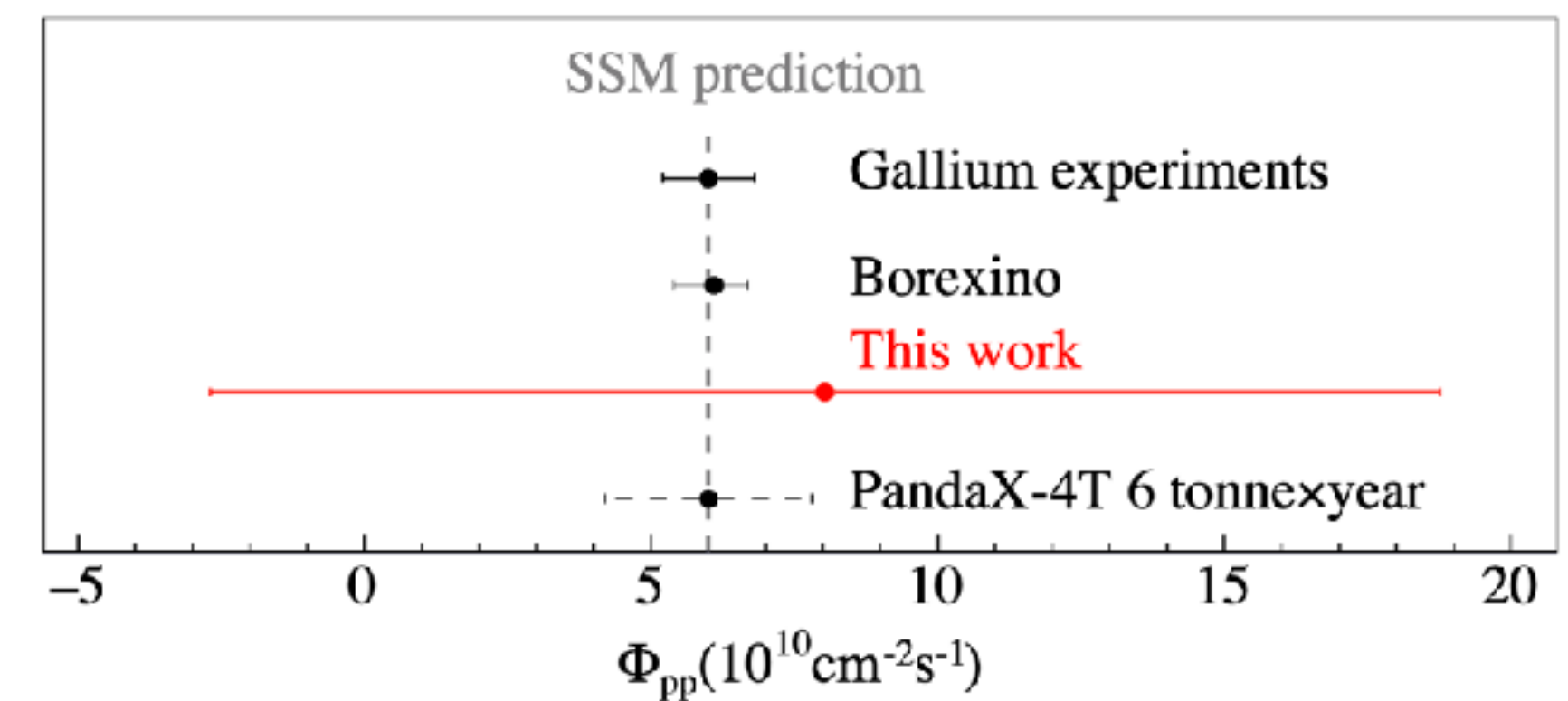
## Constraints on the $pp + {}^7\text{Be}$ neutrino flux from PandaX-4T



PandaX-4T is a dual phase liquid Xe TPC

$pp + {}^7\text{Be}$  neutrinos are a sub-dominant background in the electron recoil!

No evidence yet for solar neutrinos - **Yet!**



# LHC Machine Towards Major LS3 Upgrades

## NEW TECHNOLOGIES FOR THE HIGH-LUMINOSITY LHC

CERN [site](#)

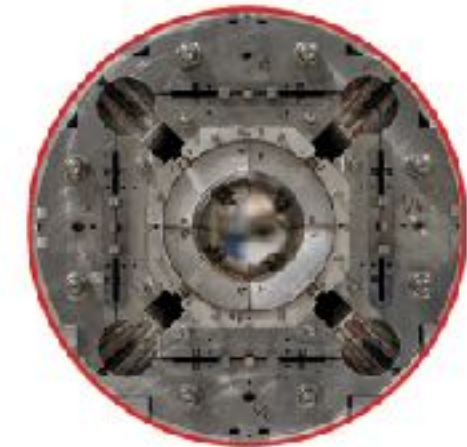


### CIVIL ENGINEERING

2 new 300-metre service tunnels and 2 shafts near ATLAS and CMS.

### "CRAB" CAVITIES

16 superconducting "crab" cavities for the ATLAS and CMS experiments to tilt the beams before collisions.



### FOCUSING MAGNETS

12 more powerful quadrupole magnets for the ATLAS and CMS experiments, designed to provide the final focusing of the beams before collisions.



### COLLIMATORS

15 to 20 additional collimators and replacement of 60 collimators with improved performance to reinforce machine protection.

### CRYSTAL COLLIMATORS

New crystal collimators in the IR7 cleaning insertion to improve cleaning efficiency during operation with ion beams.

ATLAS

ALICE

LHC TUNNEL

CMS

LHCb



### SUPERCONDUCTING LINKS

Electrical transmission lines based on a high-temperature superconductor to carry the very high DC currents to the magnets from the powering systems installed in the new service tunnels near ATLAS and CMS.

Front page of the CERN Courier says it all!



LS3 installation fully on track!

Nb<sub>3</sub>Sn series magnets manufactured at Fermilab arrived at CERN! See CERN [News](#).



# ATLAS and CMS Towards Major LS3 Upgrades

A new ATLAS for the high-luminosity era

18 January 2023 | By [Stelan Guinçon](#), [Christian Ohm](#), [Caterina Verrieri](#)

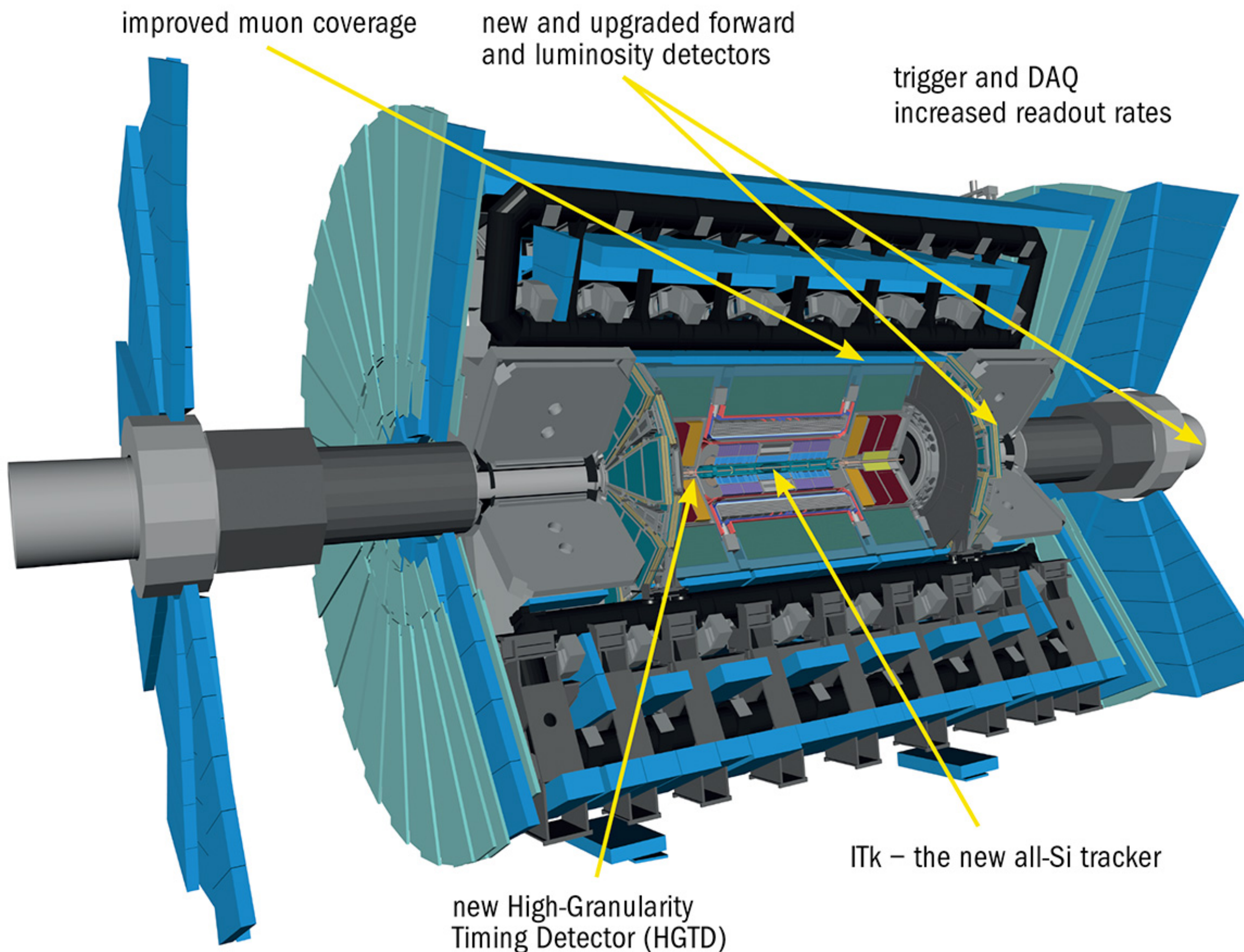
Feature [link](#)

ACCELERATORS | FEATURE

## CMS prepares for Phase II

9 January 2023

CERN Courier  
article [link](#)



**Trigger/HLT/DAQ**

- Track information in L1-Trigger
- L1-Trigger: 12.5 ms latency – output 750 kHz
- HLT output 7.5 kHz

**New Endcap Calorimeters**

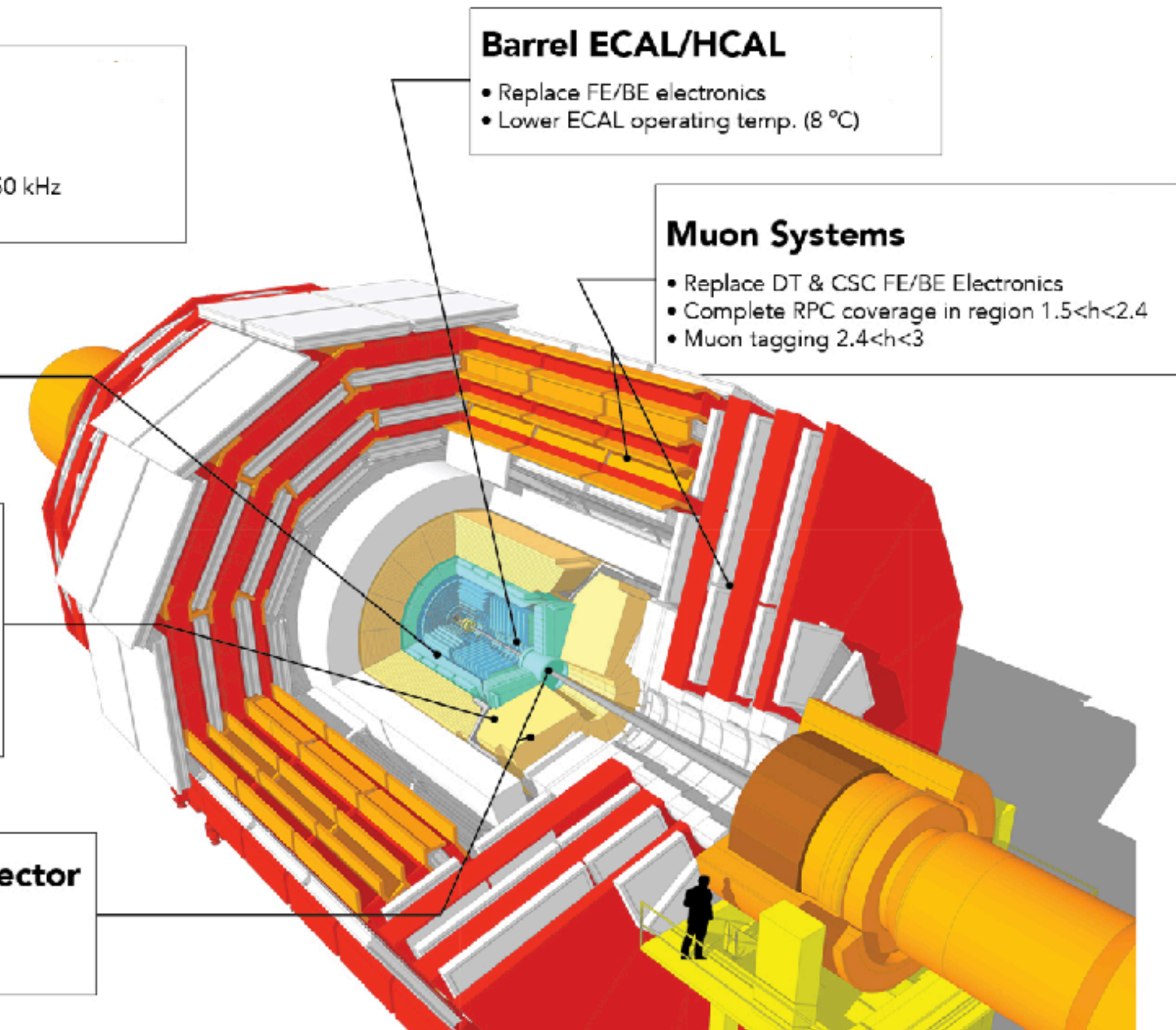
- Rad. tolerant – high granularity
- 3D capable

**New Tracker**

- Rad. tolerant – high granularity – significant less material
- 40 MHz selective readout ( $p_T > 2$  GeV) in Outer Tracker for L1-Trigger
- Extended coverage to  $h=4$

**MIP Precision Timing Detector**

- Barrel: Crystal +SiPM
- Endcap: Low Gain Avalanche Diodes



From [CLASSE](#) (Cornell)

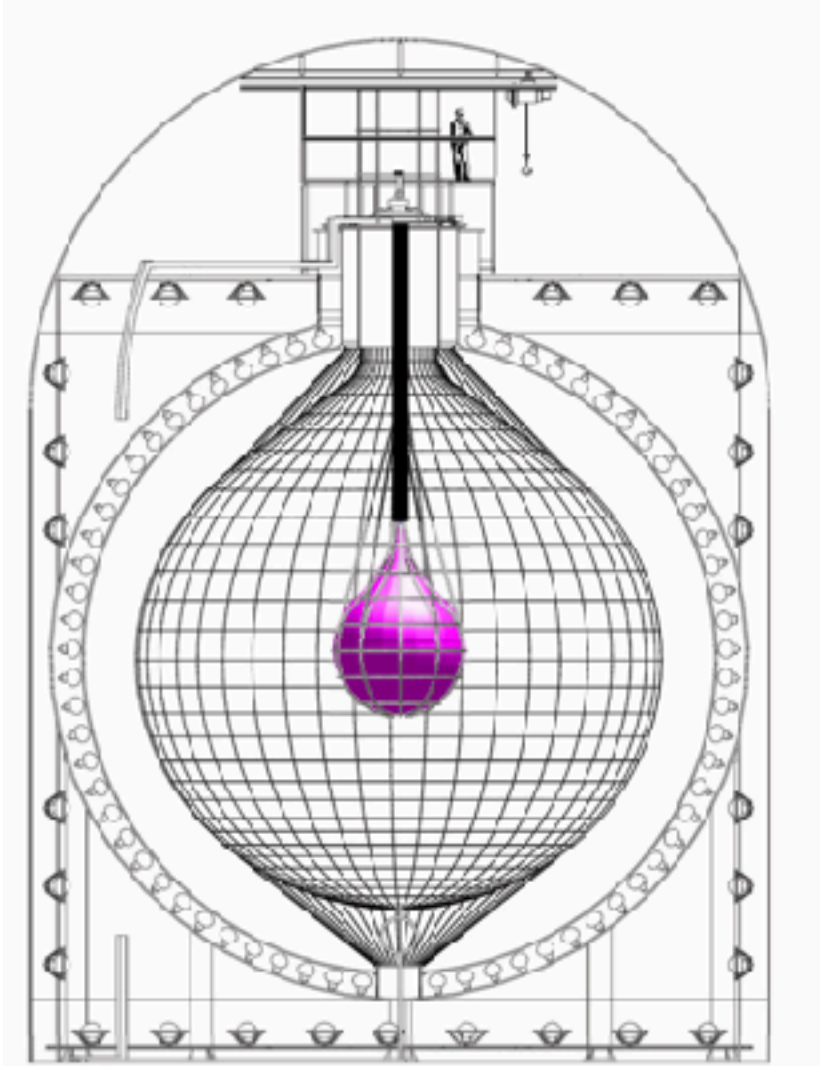
# Neutrinoless double beta decay

Susanne Mertens

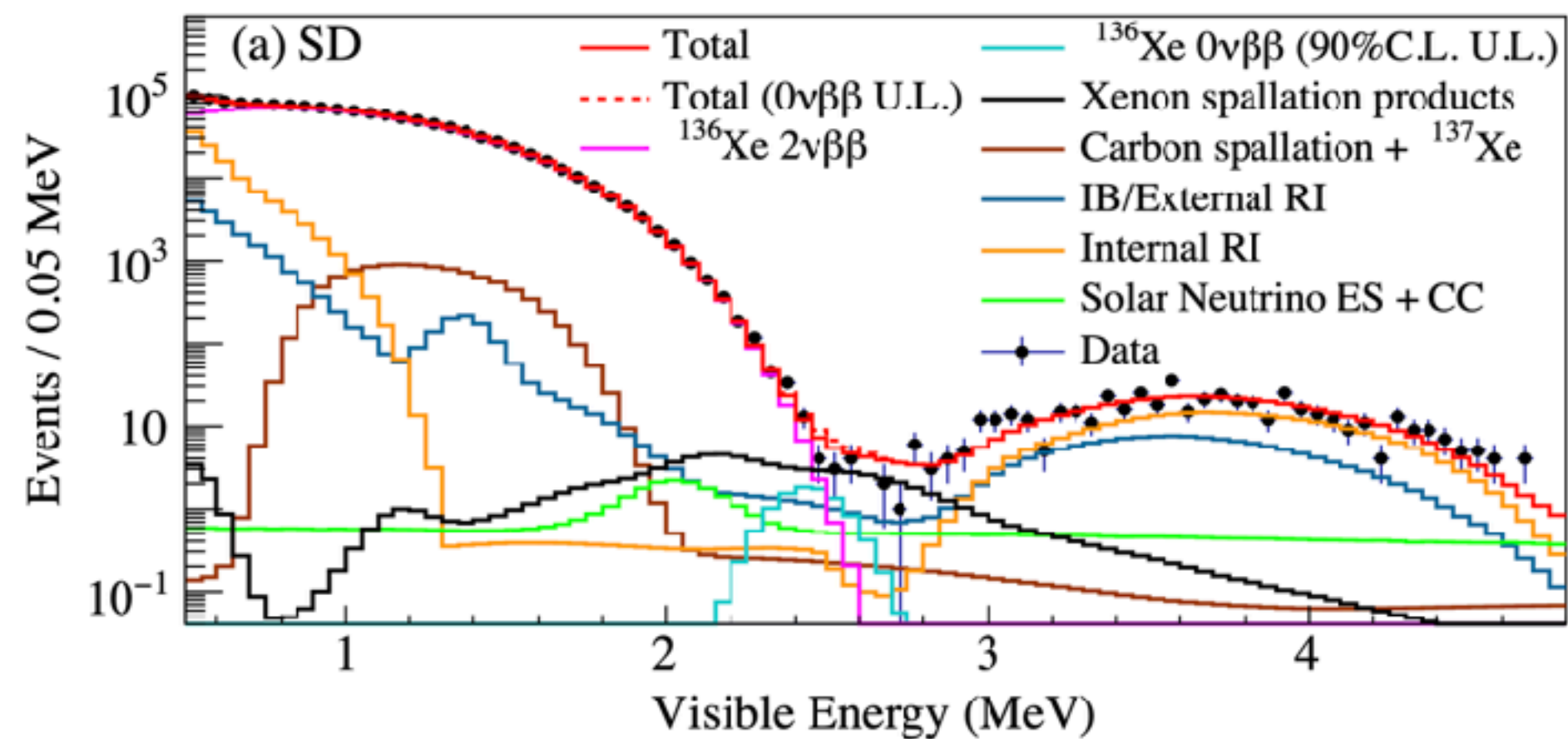
86

## KamLandZen

Mini-balloon Radius = 1.90 m  
Xenon mass = 750 kg  
Data taking started in 2019



The largest number of  $\beta\beta$  nuclei. Low BG by distillation and filtration of both Liquid Scintillator and Xenon.

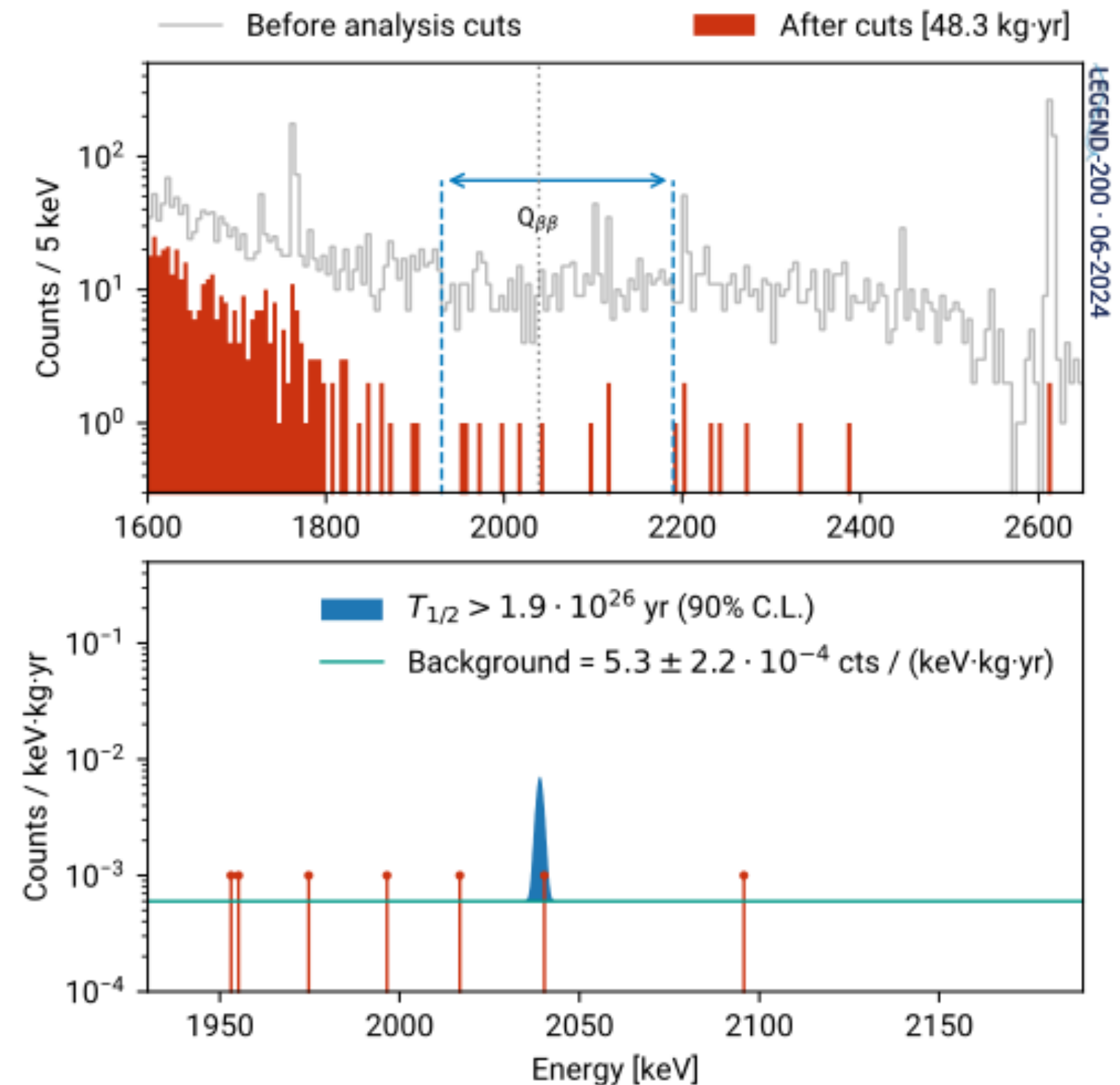


$$\langle m_{\beta\beta} \rangle < 28 - 122 \text{ meV}$$

$$\text{KamLAND-ZEN-1T: } \langle m_{\beta\beta} \rangle < \sim 20 \text{ meV}$$

## Legend-200

Germanium Semiconductor, with enrichment to > 90% in  $^{76}\text{Ge}$  ( $Q_{\beta\beta}=2039 \text{ keV}$ ). Excellent energy resolution (0.1 % FWHM @  $Q_{\beta\beta}$ )



$$\langle m_{\beta\beta} \rangle < 75 - 178 \text{ meV}$$

Legend 1000

$$\langle m_{\beta\beta} \rangle < \sim 20 \text{ meV}$$

