# Can the QCD axion feed a DE component?

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#### The axion defining interaction

 $\mathscr{L}_{a} = \frac{\alpha_{s}}{8\pi} \left( \frac{a(x)}{F} + \bar{\theta} \right) G\tilde{G} + \mathscr{L} \left( \partial_{\mu} a(x), \psi, \varphi, A_{\mu} \right) + \left[ \delta \mathscr{L}_{\text{eff}}(a(x), \ldots) \right]$ 

 $a \rightarrow a + \text{const.}$ 

 $\partial_{\mu}a \rightarrow \partial_{\mu}a$  invariant

Absent or suppressed  $\Lambda_{\text{eff}} \sim m_P \& d \ge 10$ 

# The axion defining interaction

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 $a \rightarrow a + \text{const.}$ 

- 1.  $\theta$  is removed via a shift of the axion field  $a \to a \overline{\theta} F$
- 2. Minimum of the vacuum energy occurs for  $\langle a(x) \rangle \rightarrow 0$ : solves strong CP problem
- 3. The  $a \ G \tilde{G}$  interaction generates a mass term:

$$F^2 m_a^2 = i \int d^4 x \left\langle \frac{\alpha_s}{8\pi} G \tilde{G}(x) \frac{\alpha_s}{8\pi} G \tilde{G}(0) \right\rangle$$

 $\partial_{\mu}a(x), \psi, \varphi, A_{\mu} + [\delta \mathscr{L}_{eff}(a(x), \ldots)]$ 

 $\partial_{\mu}a \rightarrow \partial_{\mu}a$  invariant

Absent or suppressed  $\Lambda_{\rm eff} \sim m_P \& d \ge 10$ 

)  $\geq \chi \leftarrow$  "Topological susceptibility"



In a hot plasma, at T >> T<sub>c</sub>, free color charges screen the correlator:  $\chi = 0$ 

- At T < T<sub>c</sub> color charges are confined in SU(3) singlets, no screening:  $\chi = (160 \text{ MeV})^4$

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 $\mathcal{X} = \mathcal{X}(\mathsf{T})$  $\Rightarrow$ 

What is the T dependance?

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 $\mathcal{X} = \mathcal{X}(\mathsf{T})$ =>

DIGA (lowest order):  $n = \beta_0 - n_f - 4 = \frac{11}{3}N + \frac{1}{3}n_f - 4$  n = 8 (QCD) n~668 **IILM** (more appropriate for  $T \sim T_{osc}$ ):

[Interacting inst. liquid model: Shellard & Wanz, 2010]

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What is the T dependance?  $m_a^2(T) \sim T^{-n}$  [n ~ n(T)]



#### Effective mass, lattice calculations

#### Lattice QCD: we can compute axion mass

$$m_a^2 f_a^2 = \chi(T)$$



At high T (no mesons) we can analytically compute potential (DIGA)

$$V(\theta) = -\chi(T)\cos\theta$$

#### Particles with varying mass: Effective Equation of State



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No. Not enough energy density:  $\rho_b \lesssim \Lambda_b^4 < T_0^4 \sim \rho_{rad} \ll \rho_{DE}$ 

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- If a dominates puniverse, => <u>acceleration</u> already for n> 2
- Could a PNGB b(x), coupled to a "dark" gauge group  $G_b$  that is undergoing a confining PT <u>now</u> ( $\Lambda_b < T_0$ ), explain the DE?







#### Take $G_a \times G_b$ , $G_a = SU(3)_{QCD}$ ; $G_b = SU(3)$ or SU(2); $\Lambda_b \leftrightarrow \Lambda_a$





#### $\mathscr{L}_{V} \sim \bar{\psi}_{I} \psi_{R} \Phi_{1} + \bar{\chi}_{I} \chi_{R} \Phi_{2} \psi \sim (1,3), \ \chi \sim (3,3)$

Take  $G_a \times G_b$ ,  $G_a = SU(3)_{QCD}$ ;  $G_b = SU(3)$  or SU(2);  $\Lambda_b \leftrightarrow \Lambda_a$ 

$$\rightarrow \quad \bar{\psi}_L \psi_R v_1 e^{i\frac{a_1}{v_1}} + \bar{\chi}_L \chi_R v_2 e^{i\frac{a_2}{v_2}}$$







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 $\mathscr{L}_V \sim \bar{\psi}_I \psi_R \Phi_1 + \bar{\chi}_I \chi_R \Phi_2$  $\psi \sim (1,3), \ \chi \sim (3,3)$ 

This generates the potenti  $V = \Lambda_a^4 \left[ 1 - \cos\left(\frac{\varphi_a}{r}\right) \right] + \Lambda_b^4 \left[ 1 - \frac{\varphi_a}{r} \right] + \Lambda_b^4 \left[ 1 - \frac{\varphi_a}$ 

$$S; G_b = SU(3) \text{ or } SU(2); \Lambda_b \leftrightarrow \Lambda_a$$

$$\rightarrow \quad \bar{\psi}_L \psi_R v_1 e^{i\frac{a_1}{v_1}} + \bar{\chi}_L \chi_R v_2 e^{i\frac{a_2}{v_2}}$$

ial:  

$$F, F' \propto v_2, \ f \propto v_1$$

$$COS\left(\frac{\varphi_a}{F'} + \frac{\varphi_b}{f}\right)$$

$$\left(\begin{array}{c}\varphi_a\\\varphi_b\end{array}\right) = \begin{pmatrix}cos\beta & sin\beta\\-sin\beta & cos\beta\end{pmatrix}\begin{pmatrix}a_1\\a_2\end{pmatrix}$$









 $\ddot{A} + 3H\dot{A} +$  $A = \begin{pmatrix} \varphi_a \\ \varphi_b \end{pmatrix}; \quad \mathcal{M}^2 = m_a^2 \begin{pmatrix} 1 & \epsilon r(T) \\ \epsilon r(T) & r(T) \end{pmatrix}$ 

$$\mathcal{M}^2 A = 0$$
  
$$(f); \qquad m_a = \frac{\Lambda_a^2}{F}, \quad r(T) = \frac{m_b^2(T)}{m_a}, \quad \epsilon = \frac{f}{F'}$$





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Assumption: at T=0  $m_b = \Lambda_b^2 / f > m_a$  [f<<F, i.e. v<sub>1</sub> << v<sub>2</sub>]



$$\ddot{A} + 3H\dot{A} + \mathscr{M}^{2}A = 0$$

$$A = \begin{pmatrix} \varphi_{a} \\ \varphi_{b} \end{pmatrix}; \quad \mathscr{M}^{2} = m_{a}^{2} \begin{pmatrix} 1 & \epsilon r(T) \\ \epsilon r(T) & r(T) \end{pmatrix}; \quad m_{a} = \frac{\Lambda_{a}^{2}}{F}, \quad r(T) = \frac{m_{b}^{2}(T)}{m_{a}}, \quad \epsilon = \frac{f}{F'}$$

#### Assumption: at T=0 $m_b = \Lambda$

This implies a Level Crossing  $m_b(T_{LC}) = m_a$  (width  $\Gamma_{LC} \sim 3\epsilon$ ) where QCD axions  $\varphi_a$  can partially convert into b-axions  $\varphi_b$ 

$$\frac{A_b^2}{f} > m_a \quad [f << F, i.e. v_1 << v_2]$$







t<sub>LC</sub>





Adiabatic  $m_{a} (\varepsilon t_{LC}) \gg 1$   $Plot: [\varepsilon t_{LC} m_{a} = 50]$ 





t<sub>LC</sub>





Adiabatic  $m_{a} (\varepsilon t_{LC}) \gg 1$   $Plot: [\varepsilon t_{LC} m_{a} = 50]$ 





*t*LC









![](_page_26_Picture_2.jpeg)

Severe Constraining Conditions

![](_page_27_Figure_2.jpeg)

 $f > T_{\rm LC} > T_{\rm DE} > T_0 > \Lambda_b$ 

![](_page_27_Figure_4.jpeg)

![](_page_27_Picture_5.jpeg)

Severe Constraining Conditions

$$m_b(T_{\rm LC}) \sim \frac{\Lambda_b^2}{f} \left(\frac{T_b}{T_{\rm LC}}\right)^3 = m_a = \frac{\Lambda_a^2}{F}$$

 $f > T_{\rm LC} > T_{\rm DE} > T_0 > \Lambda_b$ 

Which imply a pre-inflation scenario  $F \gtrsim 10^{14} \,\text{GeV}, \ [m_a \lesssim 6 \cdot 10^{-8} \,\text{eV}], \ \theta_a \lesssim 6\%$ 

![](_page_28_Figure_5.jpeg)

![](_page_28_Picture_6.jpeg)

Severe Constraining Conditions

$$\begin{split} m_b(T_{\rm LC}) \sim \frac{\Lambda_b^2}{f} \left(\frac{T_b}{T_{\rm LC}}\right)^3 &= m_a = \frac{\Lambda_a^2}{F} \\ f > T_{\rm LC} > T_{\rm DE} > T_0 > \Lambda_b \\ \text{Which imply a pre-inflation scenario} \\ F \gtrsim 10^{14} \,\text{GeV}, \quad [m_a \lesssim 6 \cdot 10^{-8} \,\text{eV}], \quad \theta_a \lesssim 6 \,\% \\ \text{And a non-adiabatic level crossing} \end{split}$$

$$\epsilon \sim 10^{-25} \left( \frac{\Lambda_b}{10^{-4} \text{eV}} \frac{160 \text{MeV}}{\Lambda_a} \right)$$

 $t_{\rm LC} = 10^9 \,\mathrm{yr}, \ [z_{\rm LC} \sim 5] \quad \Rightarrow \quad m_a t_{\rm LC} \lesssim 10^{25}$ 

2

![](_page_29_Figure_5.jpeg)

![](_page_29_Picture_6.jpeg)

Severe Constraining Conditions

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$$\epsilon \sim 10^{-25} \left( \frac{\Lambda_b}{10^{-4} \text{eV}} \frac{160 \text{MeV}}{\Lambda_a} \right)$$

![](_page_30_Figure_5.jpeg)

#### Conclusions

![](_page_31_Picture_1.jpeg)

![](_page_32_Picture_0.jpeg)

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![](_page_32_Picture_3.jpeg)

![](_page_32_Picture_4.jpeg)

![](_page_33_Picture_0.jpeg)

- It is consistent with different, possibly evolving, EoS (Quintessence, Λ, Phantom DE)

![](_page_33_Picture_6.jpeg)

![](_page_34_Picture_0.jpeg)

- It can shed light on the "why now ?" puzzle

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![](_page_34_Picture_8.jpeg)

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- Only viable for pre-inflationary axion scenarios.

- It is consistent with different, possibly evolving, EoS (Quintessence, Λ, Phantom DE)

- If the QCD axion constitutes the DM, there is not much freedom for model building.

![](_page_35_Figure_10.jpeg)

![](_page_35_Picture_11.jpeg)

![](_page_36_Picture_0.jpeg)

- It can shed light on the "why now ?" puzzle
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![](_page_36_Picture_5.jpeg)

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- If the QCD axion constitutes the DM, there is not much freedom for model building.

Thanks for your attention !

![](_page_36_Figure_12.jpeg)

![](_page_36_Picture_13.jpeg)