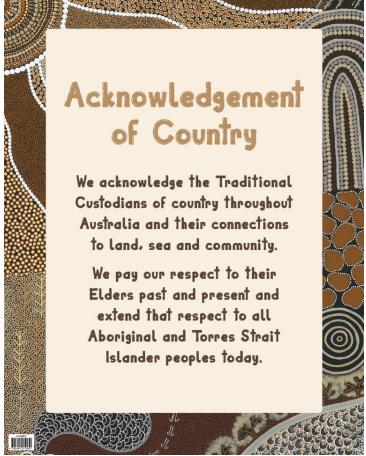
# Scattering Amplitudes, double copies and equivalence theorems in Compactifed Warped Extra-Dimensions

# Dipan Sengupta

R. Sekhar Chivukula (UCSD), Dennis Foren (UCSD-> Wolfram Computers), Kirtimaan A. Mohan (MSU), Elizabeth H. Simmons (UCSD), X. Wang (UCSD -> INFN, Rome) and J. A. Gill (Adelaide -> UNSW), George Sanmayan (Adelaide) + Anthony Williams (Adelaide)



Phys.Rev.D 100 (2019) 11, 115033 Phys.Rev.D 101 (2020) 7, 075013 Phys.Rev.D 101 (2020) 5, 055013 Phys.Rev.D 103 (2021) 9, 095024 Phys. Rev. D 107, (2023) 03505, Phys. Rev. D 109 (2024), 1 Phys. Rev. D 109 (2024), 7

### With





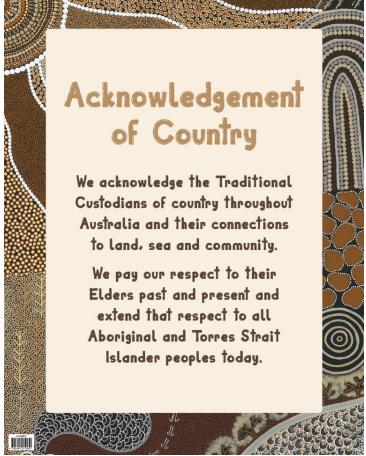




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### With









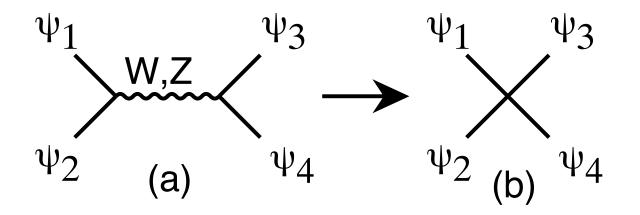
# **Effective Field Theories**

A systematic expansion in powers of momentum to extract the validity of Quantum field theories.

# Effective Field Theories

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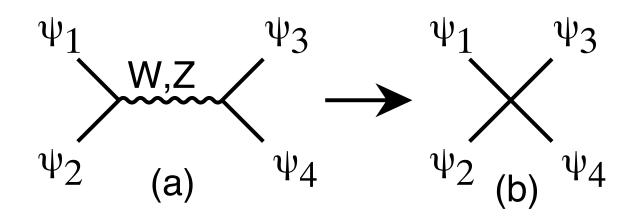
Four Fermi Theory



# **Effective Field Theories**

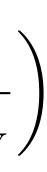
A systematic expansion in powers of momentum to extract the validity of Quantum field theories.

Four Fermi Theory



 $\underbrace{ \begin{pmatrix} \psi_3 \\ \psi_4 \end{pmatrix}}_{\psi_4} \underbrace{ \psi_1 \\ \psi_2 \end{pmatrix}_{(b)} \psi_4}^{\psi_3} \qquad i\mathcal{A} = \left( -i\frac{e}{\sin\theta_w} \right)^2 J_-^{\mu} J_+^{\nu} \frac{-ig_{\mu\nu}}{q^2 - M_W^2} = -i\frac{e^2}{\sin^2\theta_w M_W^2} J_-^{\mu} J_{\mu+} + O\left(\frac{q^2}{M_W^2}\right)$ 

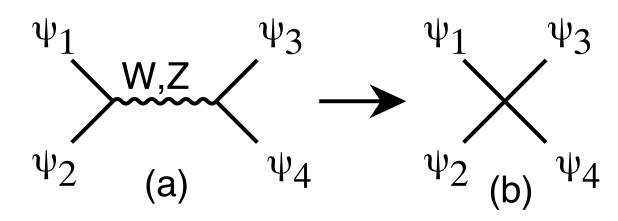
EWSB and Unitarity



# **Effective Field Theories**

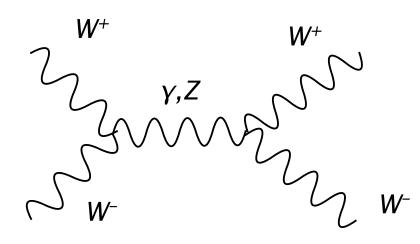
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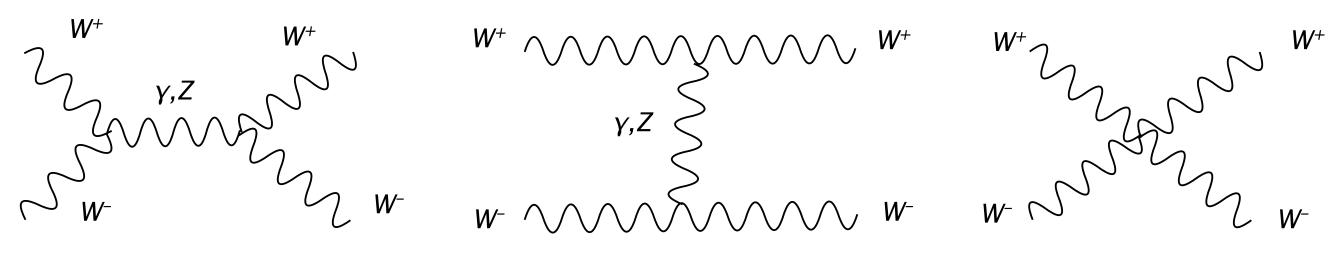
### Four Fermi Theory

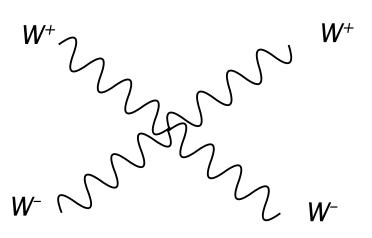


$$i\mathcal{A} = \left(-i\frac{e}{\sin\theta_w}\right)$$

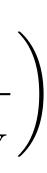
# EWSB and Unitarity







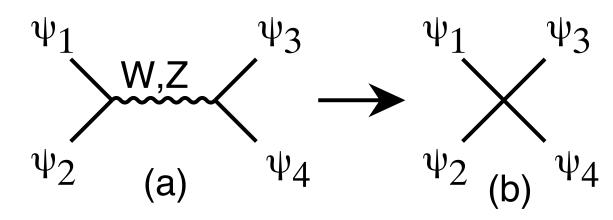
 $\frac{1}{v}\int J_{-}^{\mu}J_{+}^{\nu}\frac{-ig_{\mu\nu}}{q^{2}-M_{W}^{2}} = -i\frac{e^{2}}{\sin^{2}\theta_{w}M_{W}^{2}}J_{-}^{\mu}J_{\mu+} + O\left(\frac{q^{2}}{M_{W}^{2}}\right)$ 



# **Effective Field Theories**

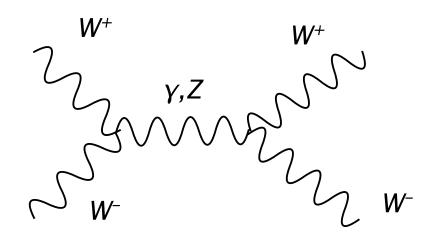
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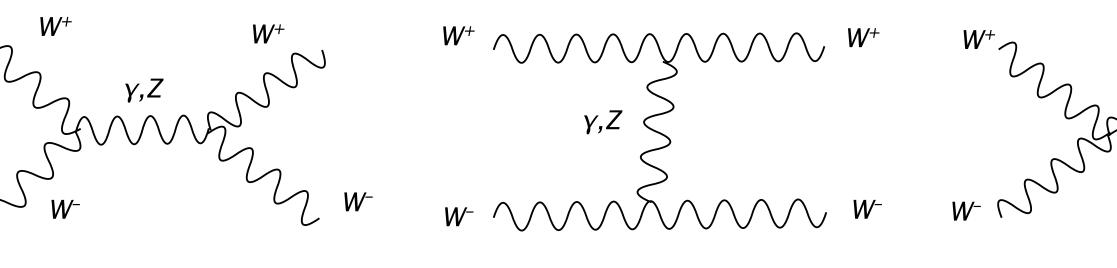
### Four Fermi Theory

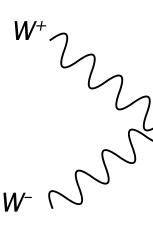


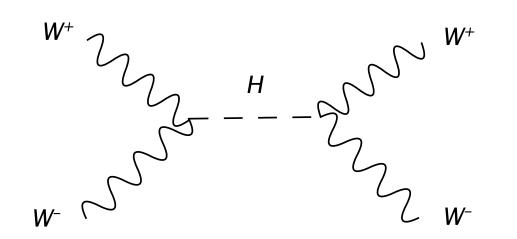
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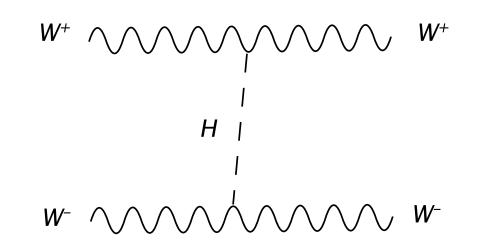
# EWSB and Unitarity







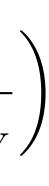


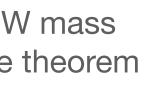


$$\int^{2} J^{\mu}_{-} J^{\nu}_{+} \frac{-ig_{\mu\nu}}{q^{2} - M_{W}^{2}} = -i \frac{e^{2}}{\sin^{2} \theta_{w} M_{W}^{2}} J^{\mu}_{-} J_{\mu+} + O\left(\frac{q^{2}}{M_{W}^{2}}\right)$$

$$\begin{split} & \stackrel{W^{*}}{\longrightarrow} \qquad \epsilon_{L}^{\mu} = k^{\mu}/M_{W} + O(M_{W}^{2}/E_{W}^{2}) \\ & \stackrel{W^{*}}{\longrightarrow} \qquad \epsilon_{L}^{W^{+}} \cdot \epsilon_{L}^{W^{-}} \approx \frac{k_{W^{+}} \cdot k_{W^{-}}}{m_{W}^{2}} = \frac{s}{2m_{W}^{2}} \qquad \begin{array}{c} \text{EFT breaks down at } V_{W^{*}} \\ & \text{Goldstone Equivalence} \end{array} \end{split}$$

The Higgs kicks in and cancels the divergence to a constant





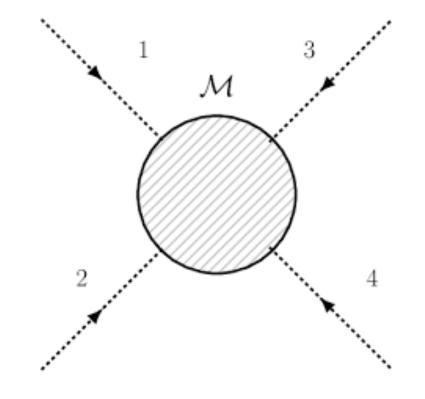
### **Power Counting in Momenta**

$$S = \int \mathrm{d}^{\mathsf{d}} x \, \mathscr{L}(x) \qquad [\mathscr{L}(x)] = \mathsf{d} \qquad \mathscr{L}(x) = \sum_{i} c_{i} O_{i}$$

$$S = \int \mathrm{d}^{\mathsf{d}} x \, \bar{\psi} \, i \partial \!\!\!/ \psi \qquad S = \int \mathrm{d}^{\mathsf{d}} x \, \frac{1}{2} \partial_{\mu} \phi \, \partial^{\mu} \phi$$

$$[\psi] = \frac{1}{2}(\mathsf{d} - 1) \qquad [\phi] = \frac{1}{2}(\mathsf{d} - 2)$$

### A generic EFT in D dimensions



A scattering amplitude with some Momentum scale

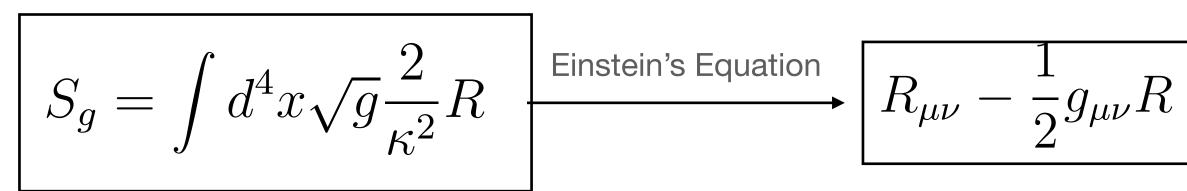
 $\mathbf{P}_i(x)$ 

 $\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\mathcal{D} \leq 4} + \frac{\mathcal{L}_5}{\Lambda} + \frac{\mathcal{L}_6}{\Lambda^2} + \dots$ 

$$\mathscr{A} \sim \left(\frac{p}{\Lambda}\right)^{\mathscr{D} - \mathsf{d}}$$

See for example, Manohar, Tasi lectures 2022

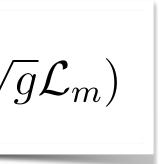
### The Classical Action for Gravity



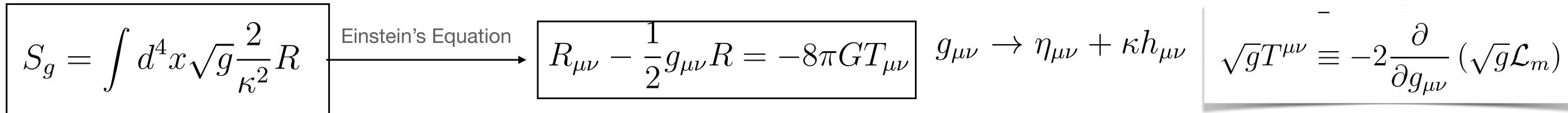
**Conserved Source** 

$$= -8\pi G T_{\mu\nu}$$

$$\sqrt{g}T^{\mu\nu} \equiv -2\frac{\partial}{\partial g_{\mu\nu}} \left(\sqrt{g}\right)$$



### The Classical Action for Gravity

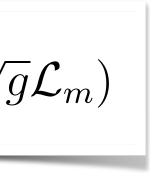


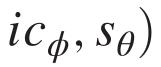
The Ricci is a two derivative object

Einstein Hilbert Action is a dimension 6 operator with a cut-off MPI

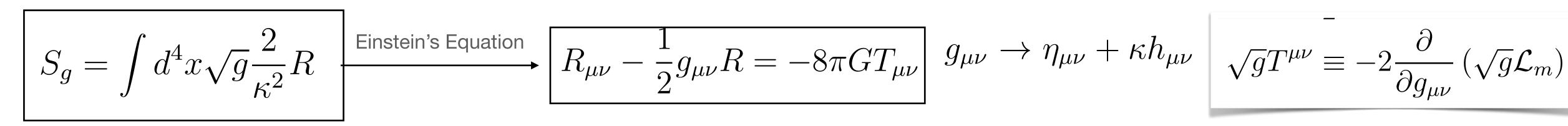
**Conserved Source** 

 $\begin{aligned} \epsilon^{\mu\nu}_{\pm 2} &= \epsilon^{\mu}_{\pm 1} \epsilon^{\nu}_{\pm 1} \\ \epsilon^{\mu}_{\pm 1} &= \pm \frac{e^{\pm i\phi}}{\sqrt{2}} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi}, -c_{\theta}s_{\phi} \mp ic_{\phi}, s_{\theta} \right) \end{aligned}$ 



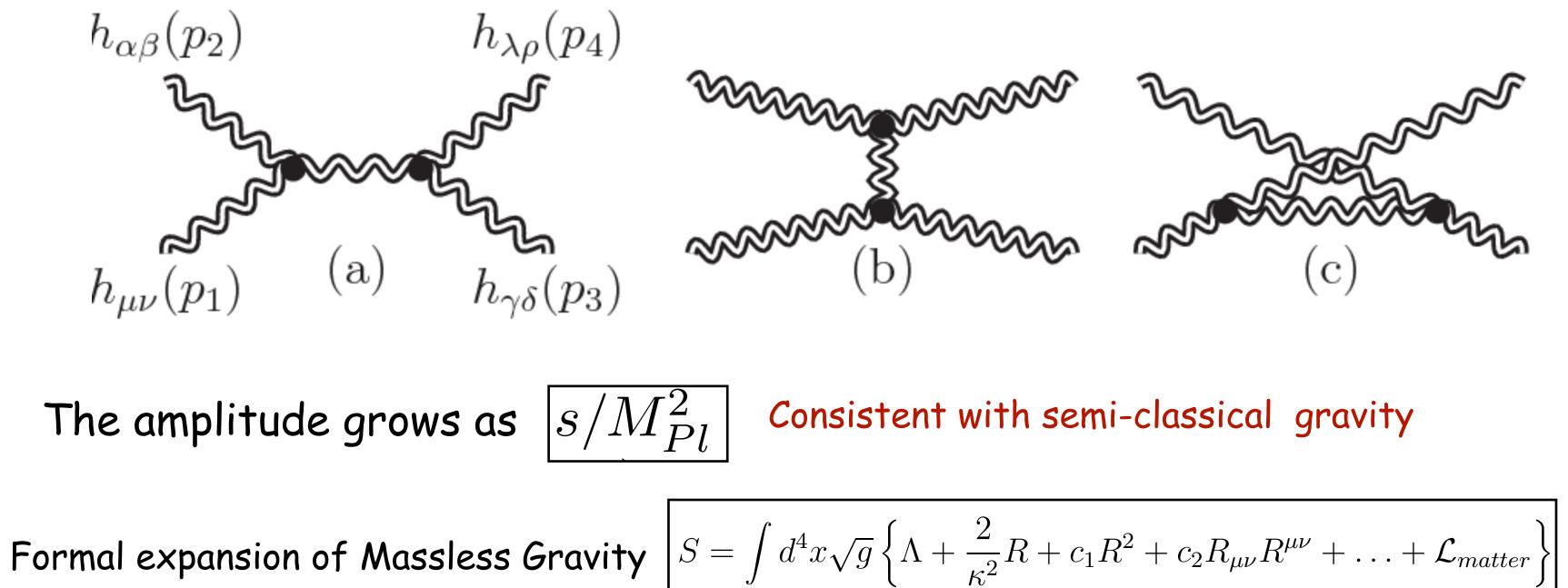


### The Classical Action for Gravity

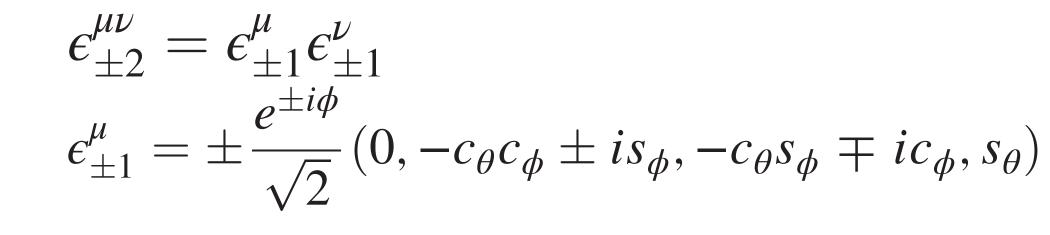


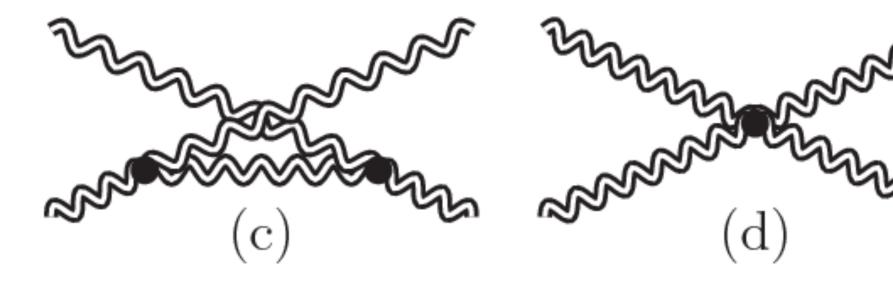
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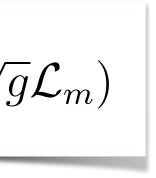


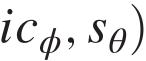
**Conserved Source** 





Consistent with semi-classical gravity





Diffeomorphism

 $h_{\mu\nu} \to h_{\mu\nu} + \partial_{(\mu}\xi_{\nu)}$ 

Equivalent to transformation for Gauge Theories

D.O.F counting in d dimensions for the massless graviton

$$d(d+1)/2 - 2d = d(d-3)/2$$

D.O.F counting in d dimensions for the massless gauge boson



$$A_{\mu} \to A_{\mu} + \partial_{\mu}\xi$$

Diffeomorphism

$$h_{\mu\nu} \to h_{\mu\nu} + \partial_{(\mu}\xi_{\nu)}$$

Equivalent to transformation for Gauge Theories

D.O.F counting in d dimensions for the massless graviton

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D.O.F counting in d dimensions for the massless gauge boson

Let's think of a Massive Photon

$$\mathcal{L}_{\rm Proca} = -\frac{1}{4}F_{\mu\nu}^2 - \frac{1}{2}m^2 A_{\mu}A^{\mu}$$

$$\begin{array}{ll} \mbox{Propagator} & \frac{-i}{p^2 + m^2} \left( \eta_{\mu\nu} + \frac{p_{\mu}p_{\nu}}{m^2} \right) \\ \\ \mbox{Stuckelberg Trick} & A_{\mu} \rightarrow A_{\mu} + \partial_{\mu}\phi & \boxed{-\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2}m^2A_{\mu}A_{\mu}} \\ \\ \mbox{Propagators} & \frac{-i\eta_{\mu\nu}}{p^2 + m^2}, & \frac{-i}{p^2 + m^2} \end{array}$$

$$A_{\mu} \to A_{\mu} + \partial_{\mu}\xi$$

Mass term explicitly breaks the gauge redundancy

$$\delta A_{\mu} = \partial_{\mu} \Lambda$$
$$\delta A_{\mu} = \partial_{\mu} \Lambda$$
$$\delta \phi = -m\Lambda$$



Let's think of a Massive Graviton

$$S_G = \int d^4x \sqrt{g}R + m^2((h_{\mu\nu})^2 - h^2) + \kappa h_{\mu\nu}T^{\mu\nu}$$
 Fierz-Pauli Theory:



Let's think of a Massive Graviton

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 Fierz-Pauli Theory:

Stuckelberg

$$h_{\mu\nu} \to h_{\mu\nu} + \partial_{\mu}A_{\nu} + \partial_{\nu}A_{\mu}$$

rescale  $A_{\mu} 
ightarrow rac{1}{m} A_{\mu}, \ \phi 
ightarrow rac{1}{m^2} \phi$  Assume source is conserved , vanishing

$$S = \int d^{D}x \ \mathcal{L}_{m=0}(h') - \frac{1}{2} F_{\mu\nu} F^{\mu\nu} - 2 \frac{D-1}{D-2} \partial_{\mu} \phi \partial^{\mu} \phi + \kappa h'_{\mu\nu} T^{\mu\nu} + \frac{2}{D-2} \kappa \phi T$$

- Scalar couples to the trace of the stress energy tensor and does not decouple. 1.
- Behaves like a Scalar-Tensor/Brans-Dicke Theory. Affects the Newtonian Potential 2.
- vanDam-Veltman-Zakharov Discontinuity. M-> 0 limit not smooth under Stuckleberg 3.

Gravity as an Effective Field Theory : Diffeomorphism and Mass Terms

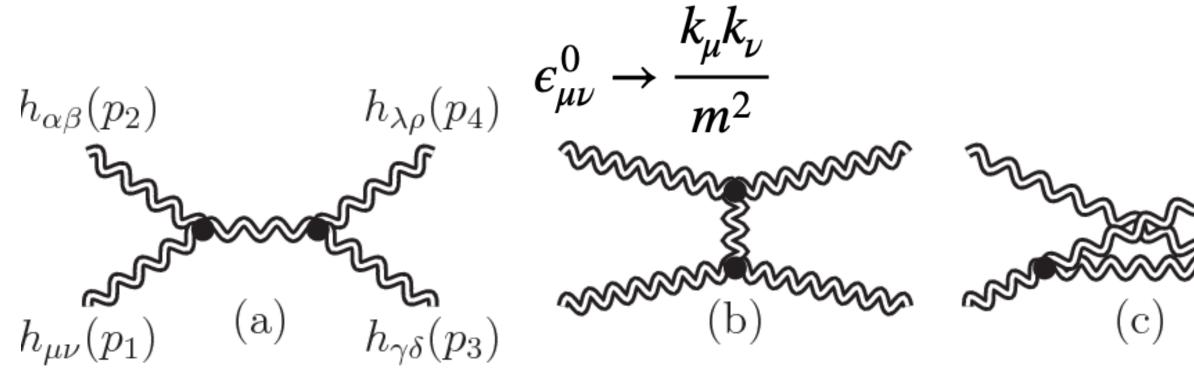
$$A_{\mu} \to A_{\mu} + \partial_{\mu}\phi$$

5 propagating degrees of freedom 2 transverse + 3 longitudinal

 $\partial_{\nu}T^{\mu\nu}$ 







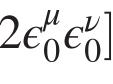
$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

$$\epsilon_{\pm 1}^{\mu\nu} = \frac{1}{\sqrt{2}} [\epsilon_{\pm 1}^{\mu} \epsilon_{0}^{\nu} + \epsilon_{0}^{\mu} \epsilon_{\pm 1}^{\nu}],$$

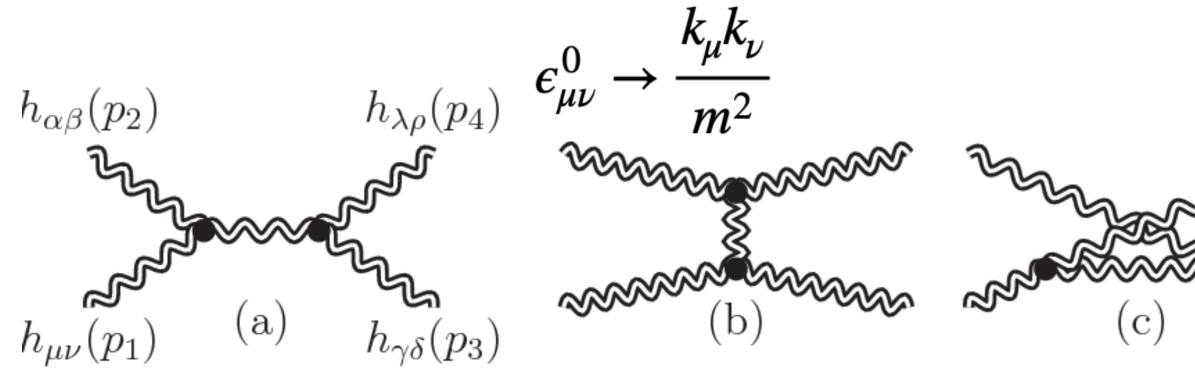
 $\epsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left[ \epsilon_{+1}^{\mu} \epsilon_{-1}^{\nu} + \epsilon_{-1}^{\mu} \epsilon_{+1}^{\nu} + 2\epsilon_{0}^{\mu} \epsilon_{0}^{\nu} \right]$ 

$$\epsilon^{\mu}_{\pm 1} = \pm \frac{e^{\pm i\phi}}{\sqrt{2}} \left(0, -c_{\theta}c_{\phi} \pm is_{\phi}, -c_{\theta}c_{\phi} \pm is_{\phi}\right)$$

$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$



### $-c_{\theta}s_{\phi} \exists$



Power Counting

- 1. Each external polarization grows as s/m<sup>2</sup>
- 2. Each vertex grows as s
- 3. The propagator grows as 1/s

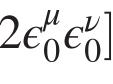
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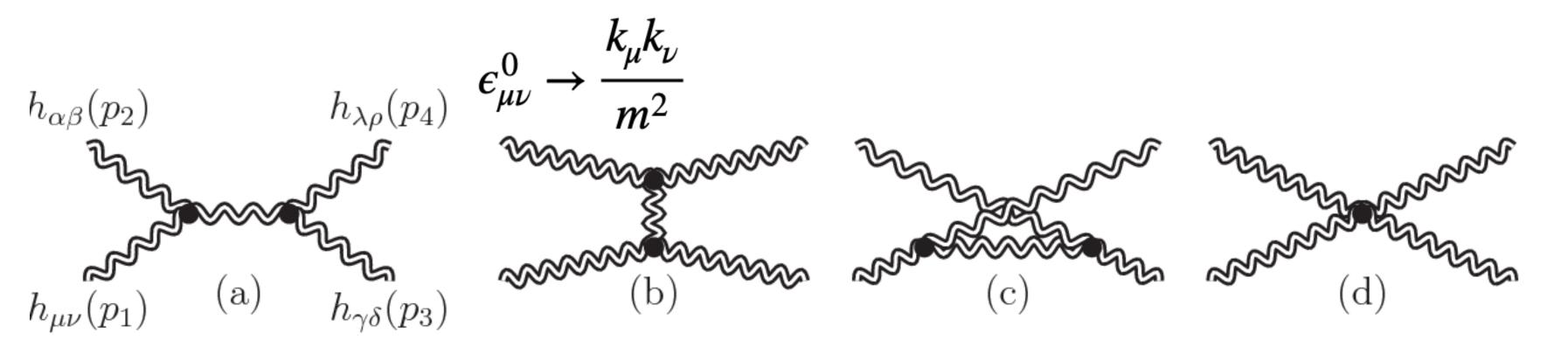
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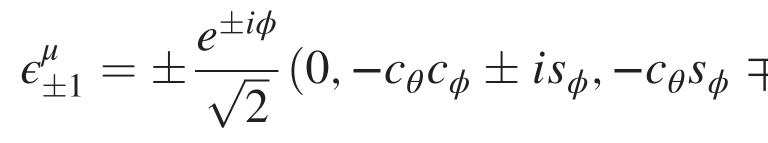
Amplitude grows as 
$$\frac{s^5}{m^8 M_{pl}^2}$$
 Discontinuity as m->

Unitarity is violated at a scale  $\Lambda_5 = (M_{pl}m^4)^{1/5} \ll$ 

$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

$$\epsilon_{\pm 1}^{\mu\nu} = \frac{1}{\sqrt{2}} [\epsilon_{\pm 1}^{\mu} \epsilon_{0}^{\nu} + \epsilon_{0}^{\mu} \epsilon_{\pm 1}^{\nu}],$$

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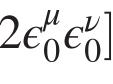


$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$

0. Does not reduce to nstein-Hilbert action

$$M_{pl}$$

De-Rham, Gabadadze, Tolley (2010) Cheung and Remen (2017) Bonifacio, Rosen, Hinterbichler (2019) Georgi, Arkani-Hamed, Schwartz (2001) Schwartz (2003)



## Most general potential

$$S = \frac{1}{2\kappa^2} \int d^D x \left[ \left( \sqrt{-g}R \right) - \sqrt{-g} \frac{1}{4} m^2 V(g,h) \right]$$

Non-linear massive gravity

### Most general potential

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$$\begin{split} V_2(g,h) &= \langle h^2 \rangle - \langle h \rangle^2, \\ V_3(g,h) &= +c_1 \langle h^3 \rangle + c_2 \langle h^2 \rangle \langle h \rangle + c_3 \langle h \rangle^3, \\ V_4(g,h) &= +d_1 \langle h^4 \rangle + d_2 \langle h^3 \rangle \langle h \rangle + d_3 \langle h^2 \rangle^2 + d_4 \langle h^2 \rangle \langle h \rangle^2 + d_5 \langle h \rangle^4, \\ V_5(g,h) &= +f_1 \langle h^5 \rangle + f_2 \langle h^4 \rangle \langle h \rangle + f_3 \langle h^3 \rangle \langle h \rangle^2 + f_4 \langle h^3 \rangle \langle h^2 \rangle + f_5 \langle h^2 \rangle^2 \langle h \\ &+ f_6 \langle h^2 \rangle \langle h \rangle^3 + f_7 \langle h \rangle^5, \end{split}$$

Tune coefficients to raise the scale, avoid ghosts

### Most general potential

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$$V_{3}(g,h) = +c_{1} \langle h^{3} \rangle + c_{2} \langle h^{2} \rangle \langle h \rangle + c_{3} \langle h \rangle^{3},$$
  

$$V_{4}(g,h) = +d_{1} \langle h^{4} \rangle + d_{2} \langle h^{3} \rangle \langle h \rangle + d_{3} \langle h^{2} \rangle^{2} + d_{4} \langle h^{2} \rangle$$
  

$$V_{5}(g,h) = +f_{1} \langle h^{5} \rangle + f_{2} \langle h^{4} \rangle \langle h \rangle + f_{3} \langle h^{3} \rangle \langle h \rangle^{2} + f_{4} \langle h^{2} \rangle + f_{6} \langle h^{2} \rangle \langle h \rangle^{3} + f_{7} \langle h \rangle^{5},$$

### Tune coefficients to raise the scale, avoid ghosts

$$\begin{split} c_1 &= 2c_3 + \frac{1}{2}, & c_2 &= -3c_3 - \frac{1}{2}, \\ d_1 &= -6d_5 + \frac{3}{2}c_3 + \frac{5}{16}, & d_2 &= 8d_5 - \frac{3}{2}c_3 - \frac{1}{4}, \\ d_3 &= 3d_5 - \frac{3}{4}c_3 - \frac{1}{16}, & d_4 &= -6d_5 + \frac{3}{4}c_3, \end{split} \qquad \begin{array}{l} \text{De-Rahm,} \\ \text{Cheung and} \\ \text{Bonifiacion,} \end{array}$$

 $\langle h \rangle^2 + d_5 \langle h \rangle^4,$  $\langle h^3 \rangle \langle h^2 \rangle + f_5 \langle h^2 \rangle^2 \langle h \rangle$ 

Gabadadze, Tolley 2011, ... nd Remen, 2020,.. , Rosen, Hinterbichler, 2021, ...

### Most general potential

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$$V_{4}(g,h) = +d_{1} \langle h^{4} \rangle + d_{2} \langle h^{3} \rangle \langle h \rangle + d_{3} \langle h^{2} \rangle^{2} + d_{4} \langle h^{2} \rangle$$
  

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$$+f_{6} \langle h^{2} \rangle \langle h \rangle^{3} + f_{7} \langle h \rangle^{5},$$

### Tune coefficients to raise the scale, avoid ghosts

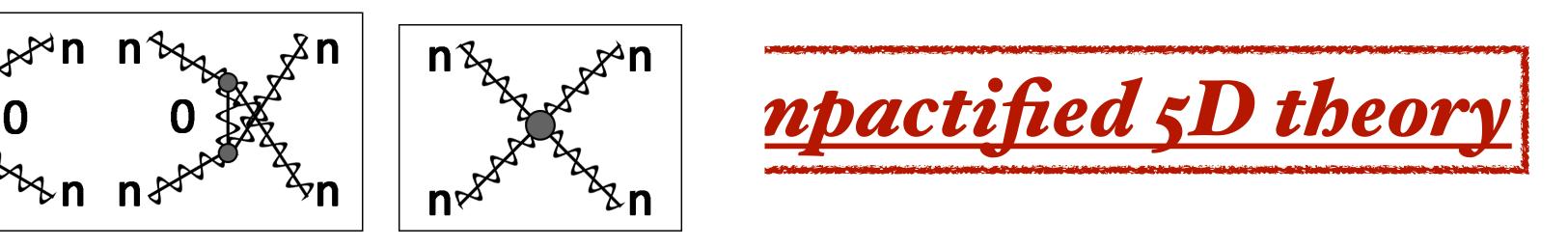
$c_1 = 2c_3 + \frac{1}{2},$	$c_2 = -3c_3 - \frac{1}{2},$	
$d_1 = -6d_5 + \frac{3}{2}c_3 + \frac{5}{16},$	$d_2 = 8d_5 - \frac{3}{2}c_3 - \frac{1}{4},$	De-Rahm, Gal Cheung and F
$d_3 = 3d_5 - \frac{3}{4}c_3 - \frac{1}{16},$	$d_4 = -6d_5 + \frac{3}{4}c_3,$	Bonifiacio, Ro

De-Rahm, Gabadadze, Tolley 2011, ... Cheung and Remen, 2020,.. Bonifiacio, Rosen, Hinterbichler, 2021, ...

# Cut-off scale raised to $\ \Lambda_3$ Can show that cut-off can't be raised above $\ \Lambda_3$

Realizations of this set up : dRGT gravity, Bi/Multigravity

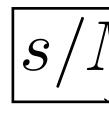
 $\langle h \rangle^2 + d_5 \langle h \rangle^4,$  $\langle h^3 \rangle \langle h^2 \rangle + f_5 \langle h^2 \rangle^2 \langle h \rangle$ 



 $\mathcal{S}_{5D}^{EH} \propto \frac{1}{M_5^3} \int d^5 x \sqrt{-g} R_{5D}$ 

High energy growth

Coupled channel analysis



 $s^{3/2}$  /



**A.** Flat Extra dimension compactified on a torus **B.** The Randall Sundrum Model (ADS)

### 5D diffeomorphism with a 5D Planck mass

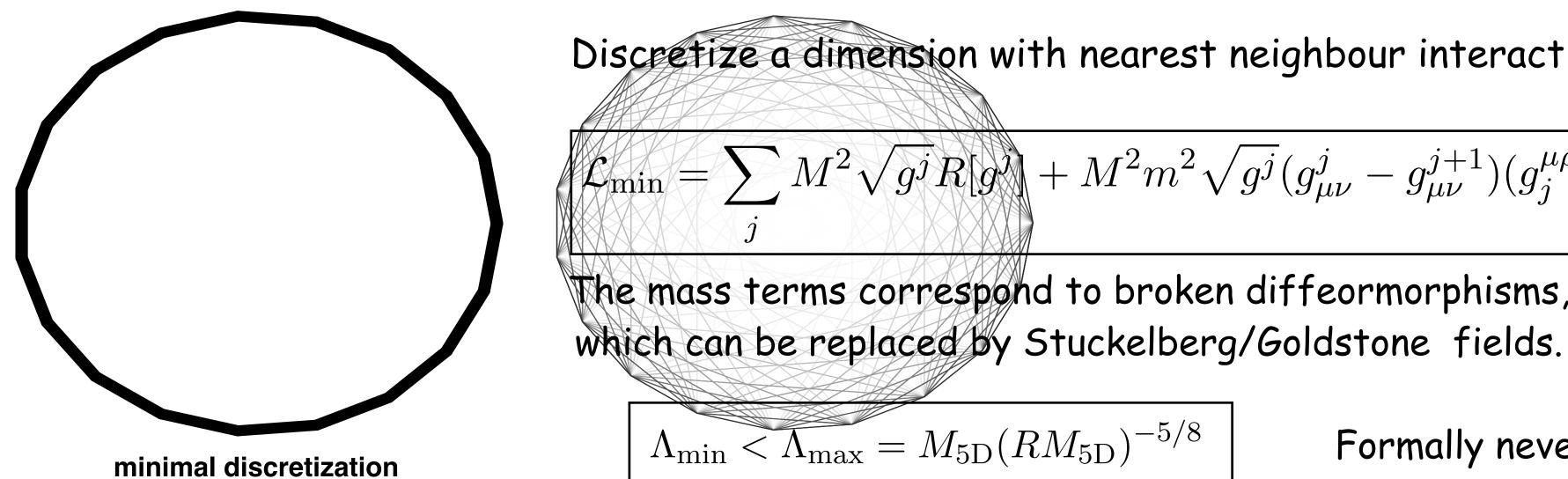
<u>Compactification (IR phenomenon) should not change the high energy (UV) behavior,</u>





# Understanding the problem : Geometrical Deconstruction of Dimensions

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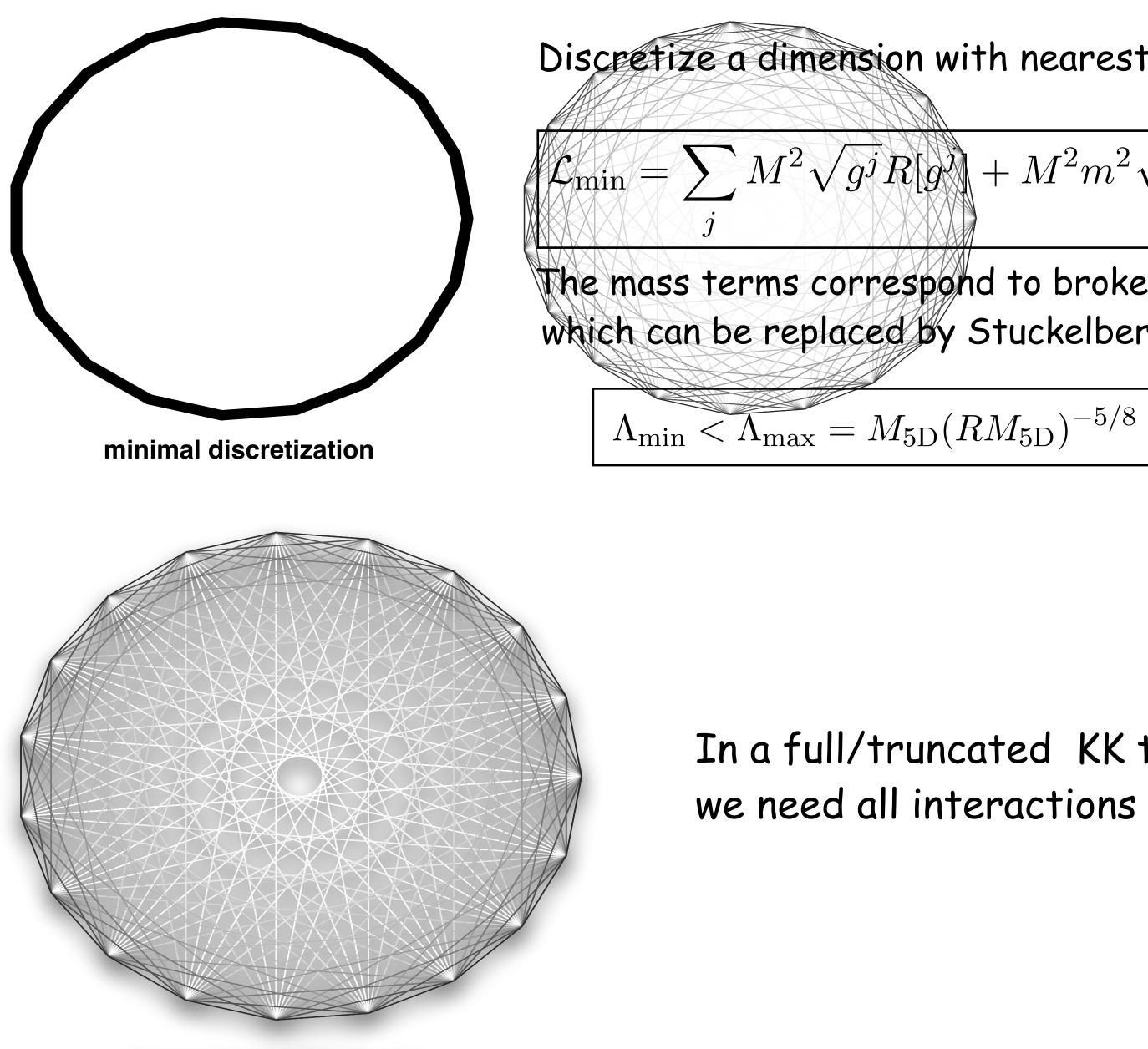


Discretize a dimension with nearest neighbour interaction

$$+ M^2 m^2 \sqrt{g^j} (g^j_{\mu\nu} - g^{j+1}_{\mu\nu}) (g^{\mu\rho}_j g^{\nu\sigma}_j - g^{\mu\nu}_j g^{\rho\sigma}_j) (g^j_{\rho\sigma} - g^{j+1}_{\rho\sigma})$$
ond to broken diffeormorphisms,  $\Lambda_{\min} = (Nm_1^4 M_{\rm Pl})^{1/5}$ 

Formally never recovers the full 5D

# Understanding the problem : Geometrical Deconstruction of Dimensions



truncated KK theory

Discretize a dimension with nearest neighbour interaction

$$+ M^2 m^2 \sqrt{g^j} (g^j_{\mu\nu} - g^{j+1}_{\mu\nu}) (g^{\mu\rho}_j g^{\nu\sigma}_j - g^{\mu\nu}_j g^{\rho\sigma}_j) (g^j_{\rho\sigma} - g^{j+1}_{\rho\sigma})$$
ond to broken diffeormorphisms,  $\Lambda_{\min} = (Nm_1^4 M_{\rm Pl})^{1/5}$ 
Stuckelberg/Goldstone fields.

Formally never recovers the full 5D

In a full/truncated KK theory, we need all interactions and not just nearest neighbour ones

> Arkani-Hamed, Georgi, Schwartz 2002 Arkani-Hamed, Schwartz 2003 Schwartz 2003





# **Compact Extra Dimensions : A primer**

Anstaz- The Fundamental Theory is 5 or D dimensional

$$S_{4+n} = -M_*^{n+2} \int d^{4+n} x \sqrt{g^{(4+n)}} R^{(4+n)} \qquad M_{Pl}^2 = M_*^{n+2} V_{(n)} = M_*^{n+2} (2\pi r)^n$$

Fluctuations in 5D  $g_{MN}(x, x^5) = \begin{pmatrix} g_{\mu\nu} & g_{\mu5} \\ g_{5\mu} & g_{55} \end{pmatrix} = \begin{pmatrix} \eta_{\mu} & \eta_{55} \end{pmatrix}$ 

**5D** gauge invariance  $h_{MN} \rightarrow h_{MN} + \delta h_{MN} = h_{MN} + \partial_N \epsilon_M + \partial_M \epsilon_N$ 

- There exists a number of n new spacial compact dimensions. For instance a simple manifold could be just  $\mathcal{M}_4 \times T^n$ .
- The fundamental Planck scale of the theory is very low  $M_D \sim \text{TeV}$ .
- The SM degrees of freedom are localized on a 3D-brane stretching along the 3 noncompact space dimensions.

$$\frac{1}{r} = M_* \left(\frac{M_*}{M_{Pl}}\right)^{\frac{2}{n}} = (1\text{TeV})10^{-\frac{32}{n}}$$

 $r \sim 2 \cdot 10^{-1}$ 

$$ds^2 = (G_{MN})dx^M$$

$$\begin{array}{ccc} h_{\mu\nu} + h_{\mu\nu} & h_{\mu5} \\ h_{\mu5} & 1 + h_{55} \end{array} \right)$$

$$^{-17}10^{\frac{32}{n}}$$
 cm  $n = 1$   $r = 2 \cdot 10^{15}$  cm



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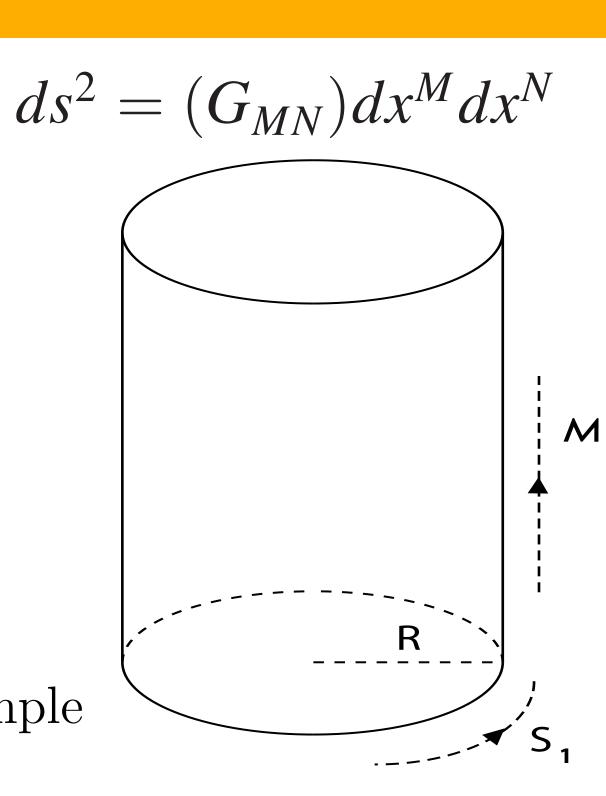
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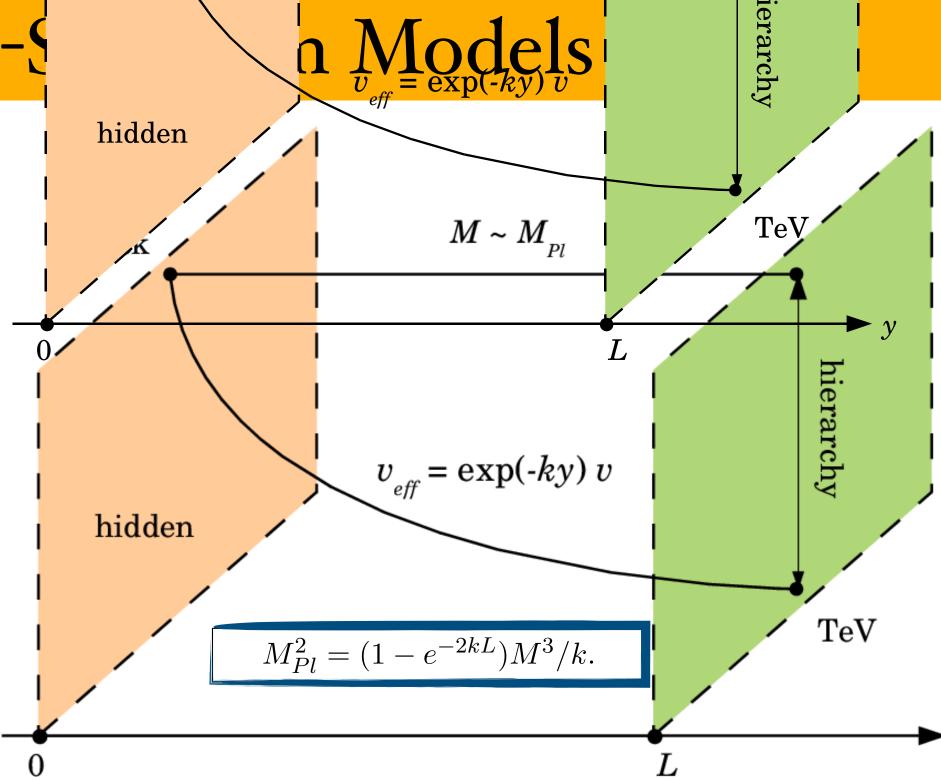
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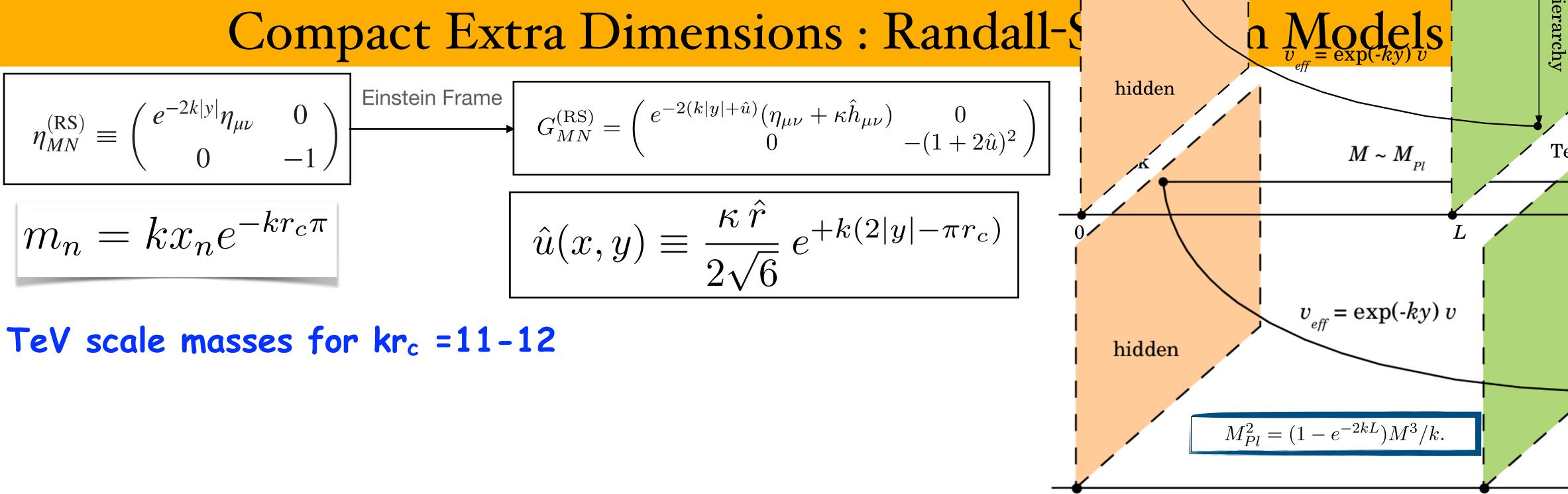


# Compact Extra Dimensions : Randall-S



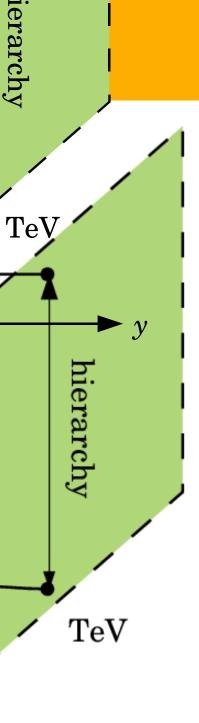
### Randall-Sundrum 1999



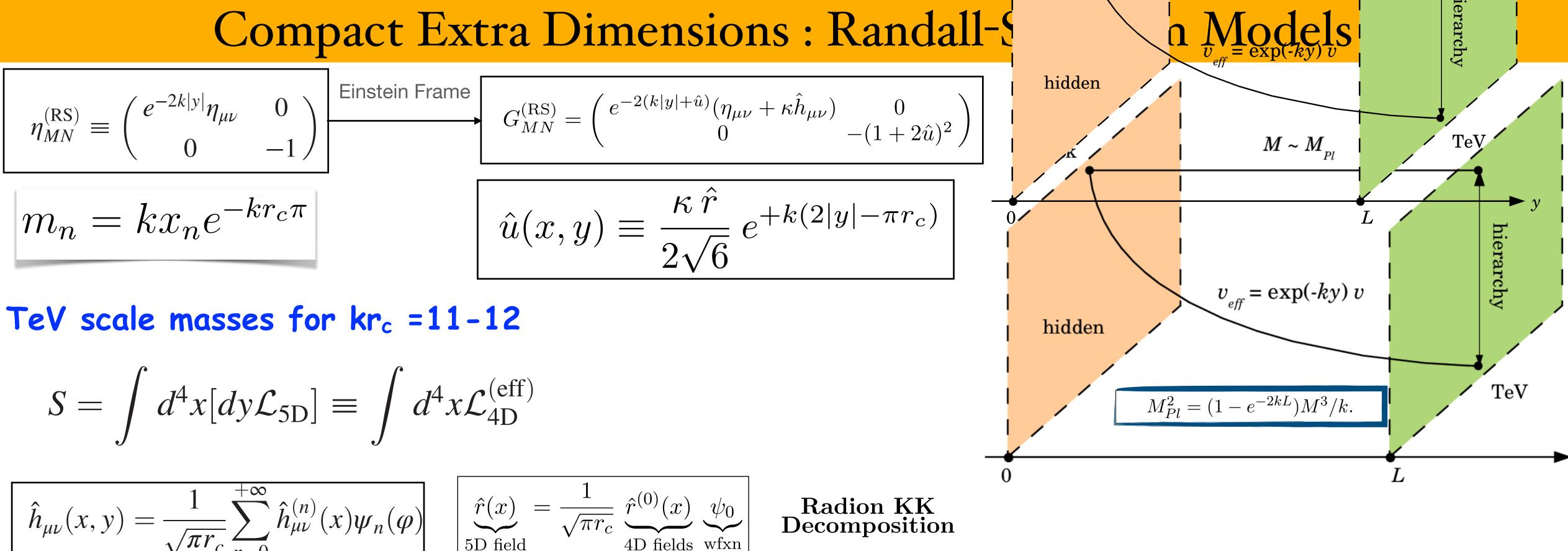


L

Randall-Sundrum 1999







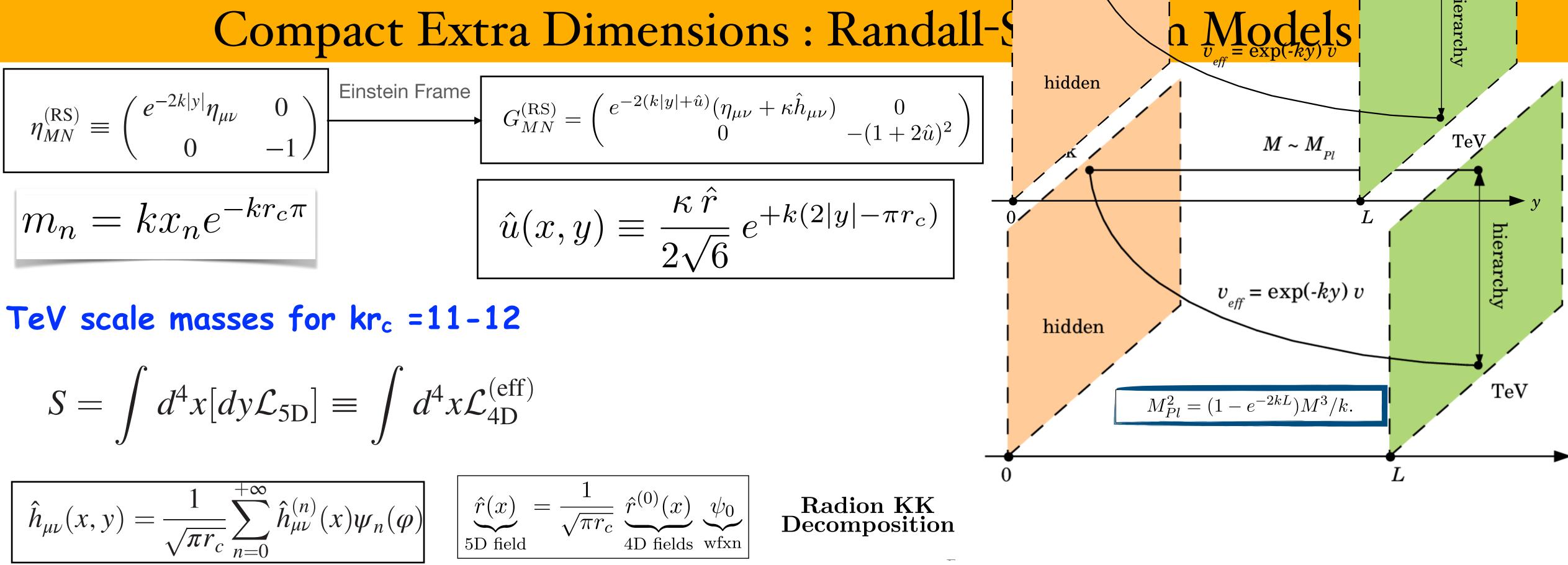
$$S = \int d^4x [dy \mathcal{L}_{5D}] \equiv \int d^4x \mathcal{L}_{4D}^{(\text{eff})}$$

$$\hat{h}_{\mu\nu}(x,y) = \frac{1}{\sqrt{\pi r_c}} \sum_{n=0}^{+\infty} \hat{h}_{\mu\nu}^{(n)}(x) \psi_n(\varphi)$$

$\hat{\alpha}(\alpha)$		1	$\hat{\alpha}(0)(\infty)$	
$\hat{r}(x)$	—	$\sqrt{\pi r_c}$	$\underbrace{\hat{r}^{(0)}(x)}_{}$	$\psi_0$
5D field		v	4D fields	wfxr

Randall-Sundrum 1999



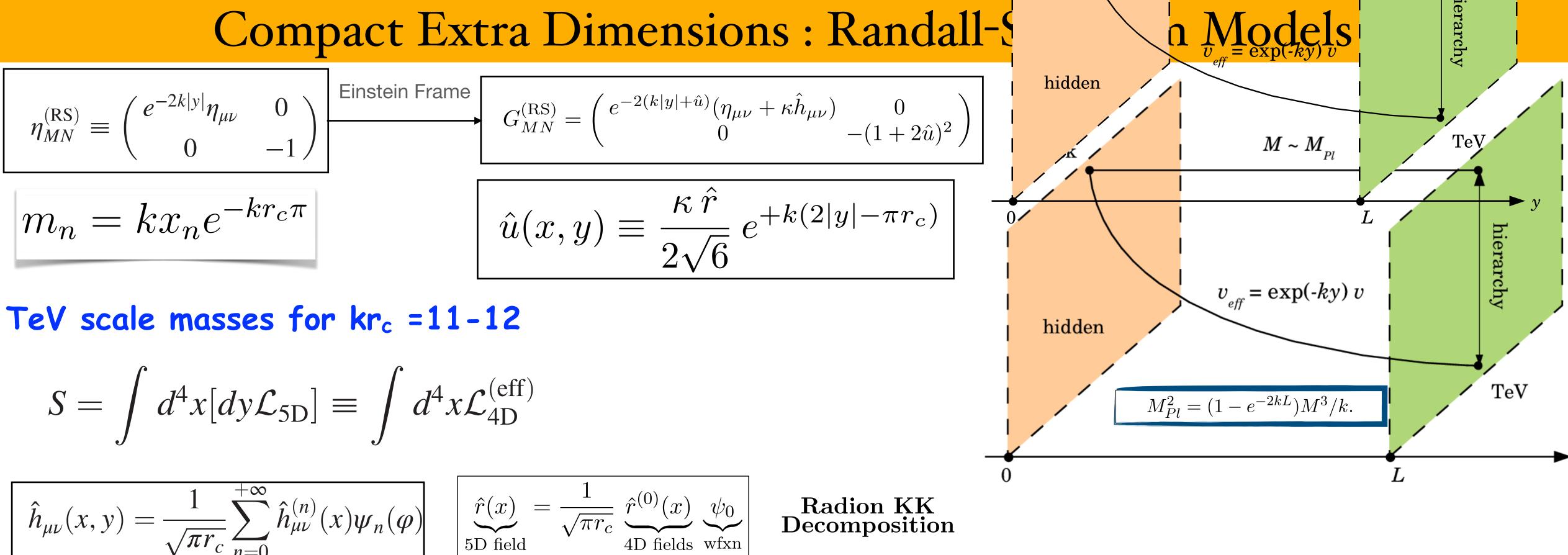


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• is orthonormal+complete with discrete spectrum  $\mu_n$ 

Randall-Sundrum 1999





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• is orthonormal+complete with discrete spectrum  $\mu_n$ 

Orthonormality

$$-\frac{d}{dy}\left[e^{-4k|y|}\frac{d\psi_n}{dy}\right] = m_n^2 e^{-2k|y|}\psi_n$$

$$\frac{1}{\pi r_c} \int_{-\pi r_e}^{+\pi r_e} dy e^{-2k|y|} \psi$$

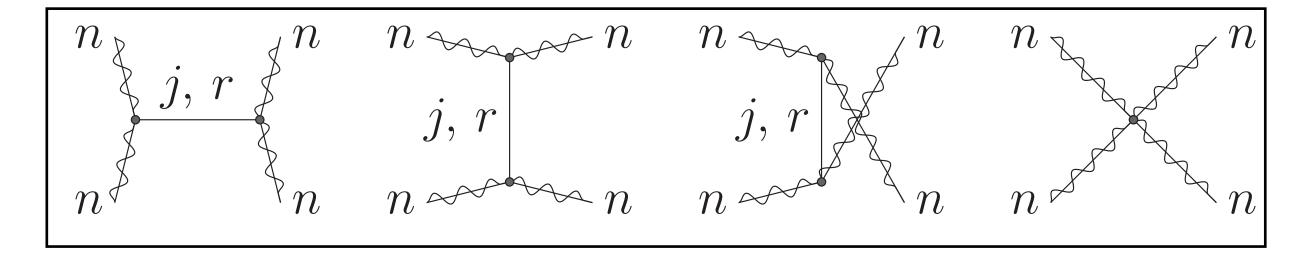
Completeness

 $\psi_m(y)\psi_n(y) = \delta_{mn} \left| \frac{1}{\pi r_c} e^{-2k|y|} \sum_i \psi_j(y)\psi_j(y') = \delta(y-y') \right|$ 

# Randall-Sundrum 1999



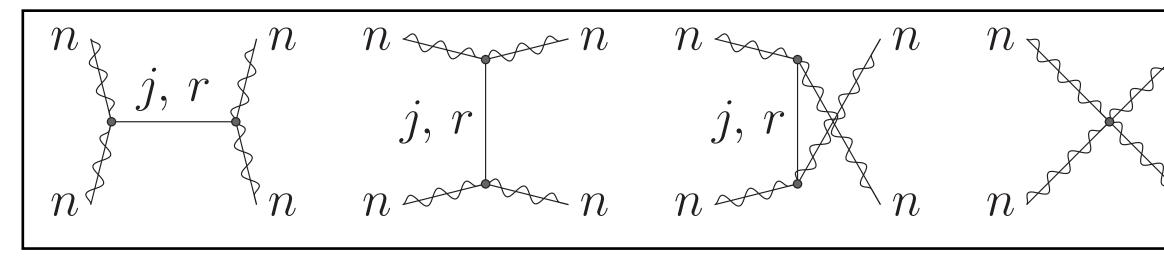
# An Elastic scattering process in compactified theories

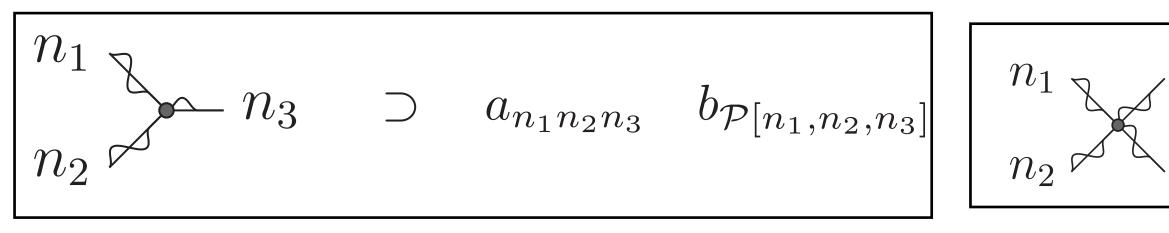




$$R_{5\mathrm{D}} = \tilde{G}^{MN} R_{MN}$$

# An Elastic scattering process in compactified theories





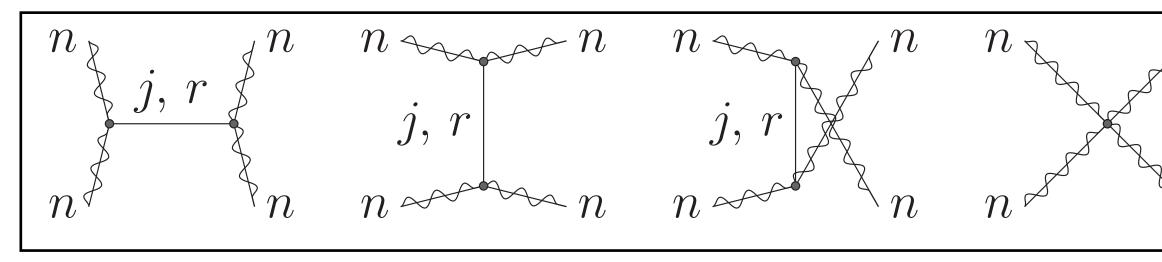
$$R_{5D} = \tilde{G}^{MN} R_{MN}$$
Expansion of Ricci gives two different  
types of coupling structures
$$n_{3}$$

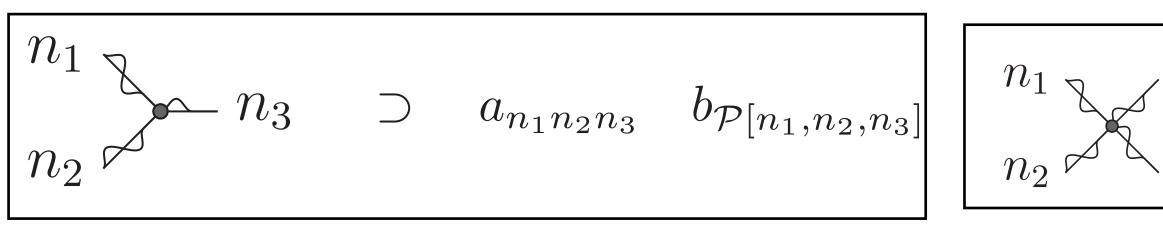
$$n_{4}$$

$$n_{1} \rightarrow a_{n_{1}n_{2}n_{3}n_{4}} \quad b_{\mathcal{P}[n_{1},n_{2},n_{3},n_{4}]}$$

$$n_{1} \rightarrow c \rightarrow b_{n_{1}n_{2}r}$$

# An Elastic scattering process in compactified theories





Define  $\mathcal{L}_{h^H r^R}^{(\text{RS})} \equiv \text{all terms in } \mathcal{L}_{5D}^{(\text{RS})}$  with *H* graviton fields and *R* radion fields. By construction, each term in this set is either

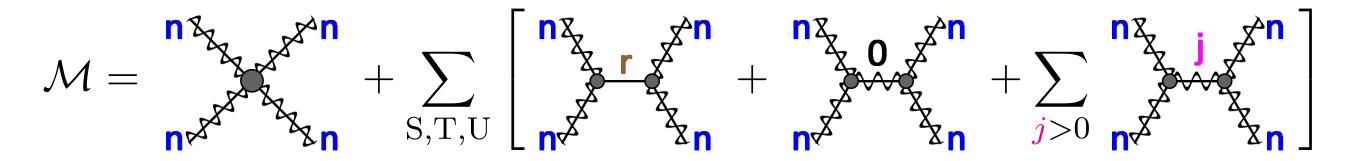
- A-Type: has two spatial derivatives  $\partial_{\mu}\partial_{\nu}$ , or
- **B-Type:** has two extra-dimensional derivatives  $\partial_u^2$

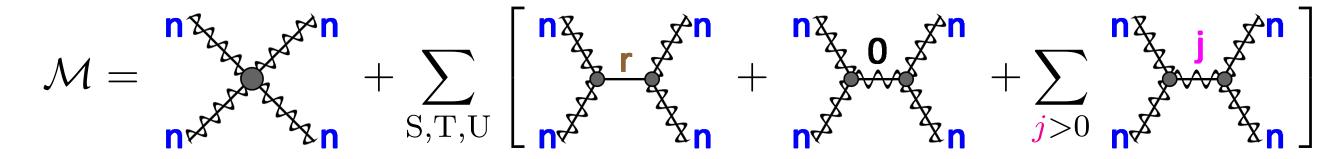
$$\begin{aligned} \mathcal{L}_{h^{H}r^{R}}^{(\mathrm{RS})} &= \mathcal{L}_{A:h^{H}r^{R}}^{(\mathrm{RS})} + \mathcal{L}_{B:h^{H}r^{R}}^{(\mathrm{RS})} \\ &= \kappa^{(H+R-2)} \left[ e^{k[2(R-1)|y| - R\pi r_{c}]} \ \overline{\mathcal{L}}_{A:h^{H}r^{R}}^{(\mathrm{RS})} + e^{i t_{A}} \right] \end{aligned}$$

$$R_{5D} = \tilde{G}^{MN} R_{MN}$$
Expansion of Ricci gives two different  
types of coupling structures
$$n_{3}$$

$$\sum a_{n_{1}n_{2}n_{3}n_{4}} b_{\mathcal{P}[n_{1},n_{2},n_{3},n_{4}]} \qquad \boxed{n_{1}}_{n_{2}} \stackrel{r}{\sim} \sum b_{n_{1}n_{2}r}$$

 $e^{k[2(R-2)|y|-R\pi r_c]} \overline{\mathcal{L}}_{B:h^H r^R}^{(RS)}$ 





$$\mathcal{M}(s,\cos\theta) \equiv \sum_{k=-\infty}^{+5} \overline{\mathcal{M}}^{(k)}(\cos\theta) \cdot s^k$$

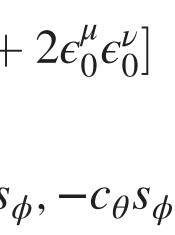
$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

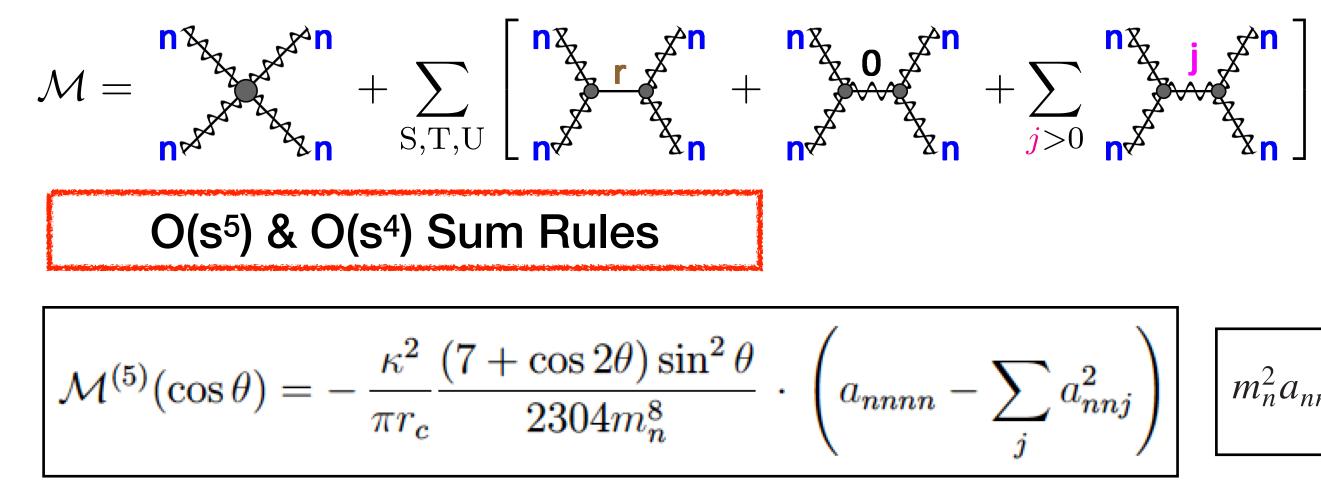
$$\epsilon_{\pm 1}^{\mu\nu} = \frac{1}{\sqrt{2}} \left[ \epsilon_{\pm 1}^{\mu} \epsilon_{0}^{\nu} + \epsilon_{0}^{\mu} \epsilon_{\pm 1}^{\nu} \right],$$

$$\epsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left[ \epsilon_{+1}^{\mu} \epsilon_{-1}^{\nu} + \epsilon_{-1}^{\mu} \epsilon_{+1}^{\nu} + \right]$$

$$\epsilon_{\pm 1}^{\mu} = \pm \frac{e^{\pm i\phi}}{\sqrt{2}} \left(0, -c_{\theta}c_{\phi} \pm is\right)$$

$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$





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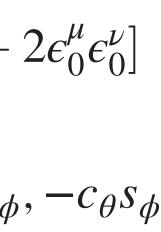
$$m_n^2 a_{nnnn} = \frac{3}{4} \sum_j m_j^2 a_{nnj}^2$$

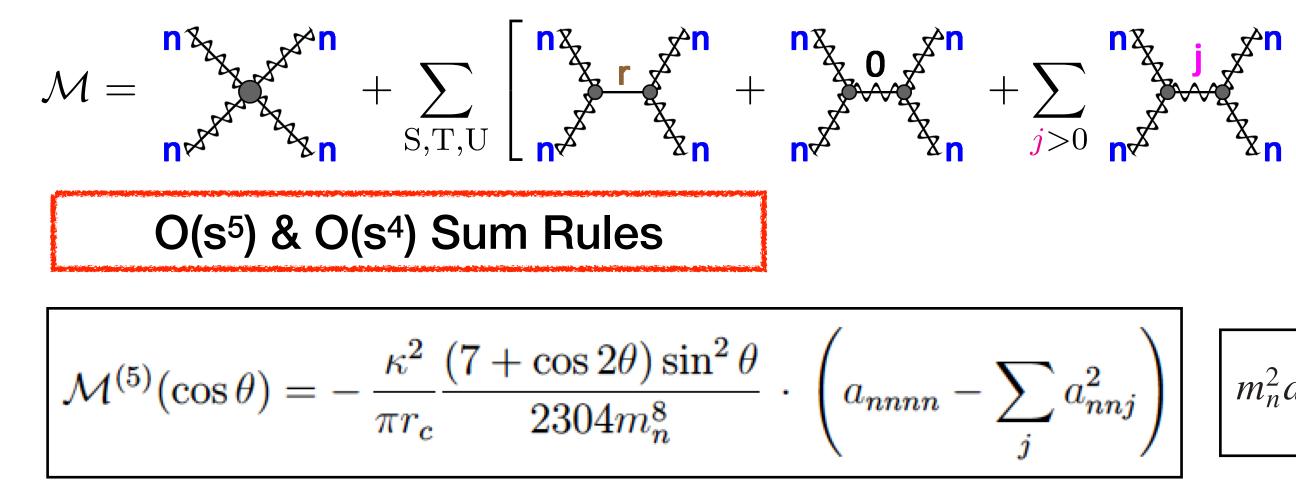
$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

$$\epsilon_{\pm 1}^{\mu\nu} = \frac{1}{\sqrt{2}} [\epsilon_{\pm 1}^{\mu} \epsilon_{0}^{\nu} + \epsilon_{0}^{\mu} \epsilon_{\pm 1}^{\nu}],$$

$$\frac{1}{\alpha_{nnj}^{2}} \qquad \epsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left[ \epsilon_{\pm 1}^{\mu} \epsilon_{-1}^{\nu} + \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu} + 2\epsilon_{0}^{\mu} \epsilon_{0}^{\nu} \right] \\
\epsilon_{\pm 1}^{\mu} = \pm \frac{e^{\pm i\phi}}{\sqrt{2}} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi}, -c_{\theta}s_{\phi} + is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left( 0, -c_{\theta}c_{\phi}$$

$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$





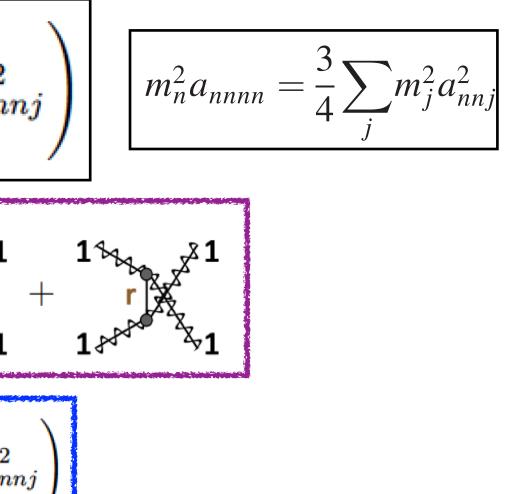
$$\mathcal{M}^{(4)}(\cos\theta) = \frac{\kappa^2}{\pi r_c} \frac{(7+\cos 2\theta)^2}{27648m_n^8} \cdot \left(4m_n^2 a_{nnnn} - 3\sum_j m_j^2 a_n^2\right)$$
$$\mathcal{M}^{(4)}(\cos\theta) = \frac{\kappa^2}{\pi r_c} \frac{\sin^2\theta}{3456m_n^8} \cdot \left(-108\frac{b_{nnr}^2}{r_c^4} + 12m_n^4 a_{nn0}^2 - 16m_n^4 a_{nnnn} + 15\sum_j m_j^4 a_n^2\right)$$

$$\mathcal{M}(s,\cos\theta) \equiv \sum_{k=-\infty}^{+5} \overline{\mathcal{M}}^{(k)}(\cos\theta) \cdot s^k$$

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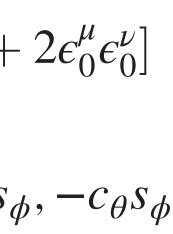
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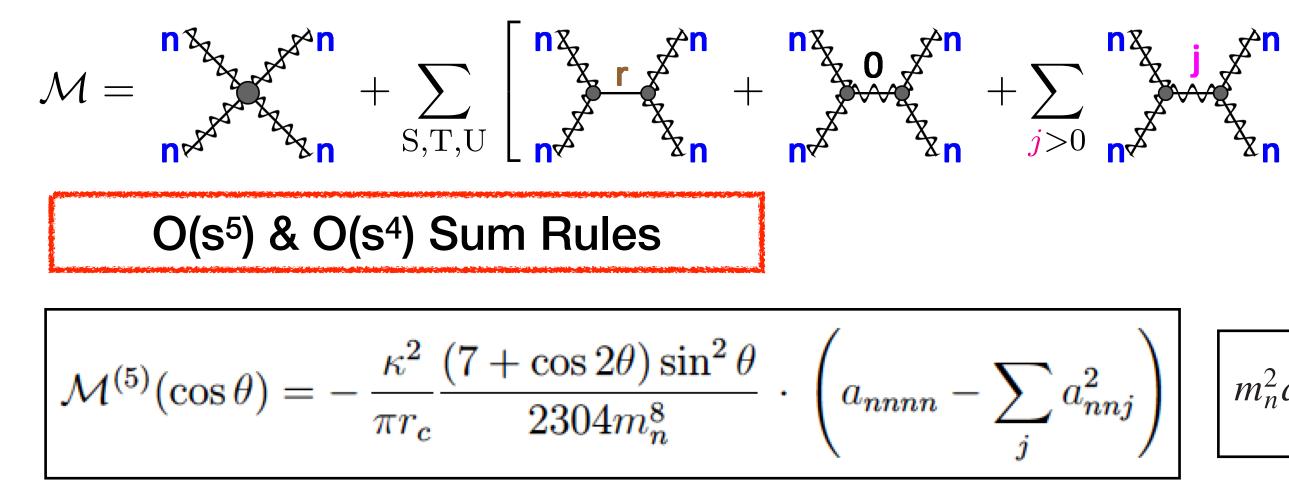


$$\epsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left[ \epsilon_{+1}^{\mu} \epsilon_{-1}^{\nu} + \epsilon_{-1}^{\mu} \epsilon_{+1}^{\nu} + \frac{i\phi}{\sqrt{6}} \right]$$

$$\epsilon_{\pm 1}^{\mu} = \pm \frac{e^{\pm i\varphi}}{\sqrt{2}} \left(0, -c_{\theta}c_{\phi} \pm is\right)$$

$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$





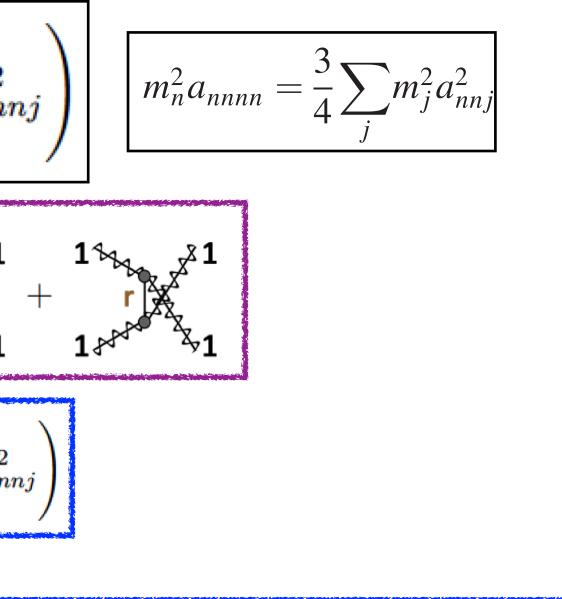
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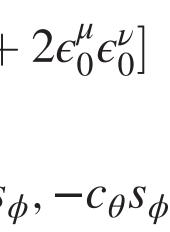


$$m_j^6 a_{nnj}^2 = 5 m_n^2 \sum_j m_j^4 a_{nnj}^2 - rac{16}{3} m_n^6 a_{nnnn}$$

$$\epsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left[ \epsilon_{+1}^{\mu} \epsilon_{-1}^{\nu} + \epsilon_{-1}^{\mu} \epsilon_{+1}^{\nu} + \frac{1}{\sqrt{6}} \right]$$

$$\epsilon_{\pm 1}^{\mu} = \pm \frac{e^{\pm i\phi}}{\sqrt{2}} \left(0, -c_{\theta}c_{\phi} \pm is\right)$$

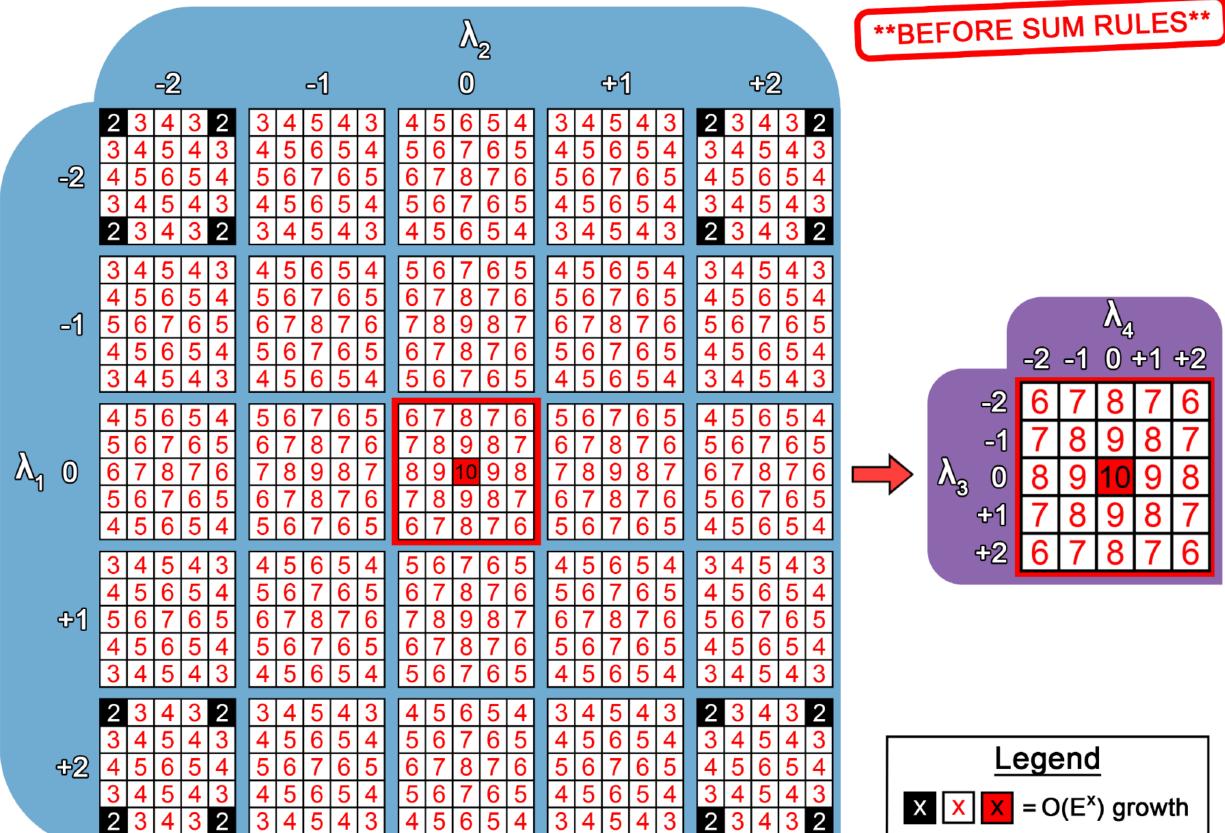
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Amplitudes grow as  $s/M_{Pl}^{2}$ : Consistent with gauge-invariance expectations

Elastic 2-to-2 KK Mode Scattering Matrix Elements in RS1 Fastest Energy Growth per Helicity Combination:  $(\lambda_1, \lambda_2) \rightarrow (\lambda_3, \lambda_4)$ 

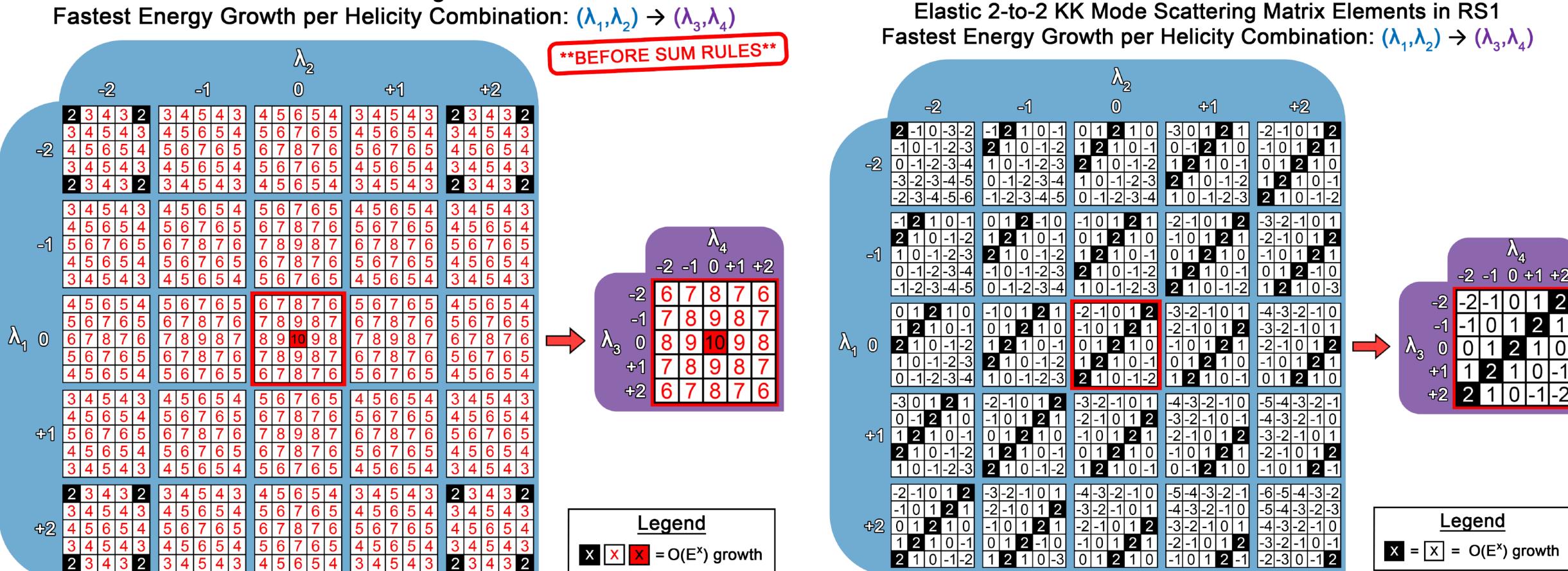


Phys.Rev.D 100 (2019) 11, 115033 Phys.Rev.D 101 (2020) 7, 075013 Phys.Rev.D 101 (2020) 5, 055013



# Amplitudes grow as $s/M_{Pl}^{2}$ : Consistent with gauge-invariance expectations

Elastic 2-to-2 KK Mode Scattering Matrix Elements in RS1



Phys.Rev.D 100 (2019) 11, 115033 Phys.Rev.D 101 (2020) 7, 075013 Phys.Rev.D 101 (2020) 5, 055013

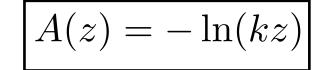
# Amplitudes grow as $s/M_{Pl}^{2}$ : Consistent with gauge-invariance expectations





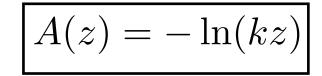
$$G_{MN} = e^{2A(z)} \begin{pmatrix} e^{-\kappa\varphi/\sqrt{6}}(\eta_{\mu\nu} + \kappa h_{\mu\nu}) & \frac{\kappa}{\sqrt{2}}A_{\mu} \\ \frac{\kappa}{\sqrt{2}}A_{\mu} & -\left(1 + \frac{\kappa}{\sqrt{6}}\varphi\right)^2 \end{pmatrix} dz \\ \hline h_{\mu\nu}(x^{\alpha}, z) = \sum_{n=0}^{\infty} h_{\mu\nu}^{(n)}(x^{\alpha})f^{(n)}(z) \\ A_{\mu}(x^{\alpha}, z) = \sum_{n=1}^{\infty} A_{\mu}^{(n)}(x^{\alpha})g^{(n)}(z) \\ \varphi(x^{\alpha}, z) = \sum_{n=0}^{\infty} \varphi^{(n)}(x)k^{(n)}(z) \\ \end{pmatrix}$$

 $ds^{2} = e^{2A(z)} (\eta_{\mu\nu} dx^{\mu} dx^{\nu} - dz^{2})$ 



$$G_{MN} = e^{2A(z)} \begin{pmatrix} e^{-\kappa\varphi/\sqrt{6}}(\eta_{\mu\nu} + \kappa h_{\mu\nu}) & \frac{\kappa}{\sqrt{2}}A_{\mu} \\ \frac{\kappa}{\sqrt{2}}A_{\mu} & -\left(1 + \frac{\kappa}{\sqrt{6}}\varphi\right)^{2} \end{pmatrix} \begin{bmatrix} d \\ d \\ h_{\mu\nu}(x^{\alpha}, z) = \sum_{n=0}^{\infty} h_{\mu\nu}^{(n)}(x^{\alpha})f^{(n)}(z) \\ A_{\mu}(x^{\alpha}, z) = \sum_{n=1}^{\infty} A_{\mu}^{(n)}(x^{\alpha})g^{(n)}(z) \\ \varphi(x^{\alpha}, z) = \sum_{n=0}^{\infty} \varphi^{(n)}(x)k^{(n)}(z) \end{bmatrix} \xrightarrow{\bullet} \text{Only 0 mode}$$

 $ds^{2} = e^{2A(z)} (\eta_{\mu\nu} dx^{\mu} dx^{\nu} - dz^{2})$ 



- K gravitons, O mode + tower of massive states
- I KK graviphotons/goldstones
- e/radion is physical, higher modes higgsed/goldstones

$$G_{MN} = e^{2A(z)} \begin{pmatrix} e^{-\kappa\varphi/\sqrt{6}}(\eta_{\mu\nu} + \kappa h_{\mu\nu}) & \frac{\kappa}{\sqrt{2}}A_{\mu} \\ \frac{\kappa}{\sqrt{2}}A_{\mu} & -\left(1 + \frac{\kappa}{\sqrt{6}}\varphi\right)^{2} \end{pmatrix} \begin{bmatrix} d_{\mu\nu}(x^{\alpha}, z) = \sum_{n=0}^{\infty} h_{\mu\nu}^{(n)}(x^{\alpha})f^{(n)}(z) \\ A_{\mu}(x^{\alpha}, z) = \sum_{n=1}^{\infty} A_{\mu}^{(n)}(x^{\alpha})g^{(n)}(z) \\ \varphi(x^{\alpha}, z) = \sum_{n=0}^{\infty} \varphi^{(n)}(x)k^{(n)}(z) \end{bmatrix} \xrightarrow{} Only 0 \text{ mode}$$

$$\mathcal{L} = \mathcal{L}_{EH} + \mathcal{L}_{CC} + \Delta \mathcal{L} + \mathcal{L}_{GF} + \mathcal{L}_{m}$$

$$\mathcal{L}_2 = \frac{1}{2} h^n_{\mu\nu} \mathcal{D}^{\mu\nu\rho\sigma}_h h^n_{\rho\sigma} + \frac{1}{2} A^n_\mu \mathcal{D}^{\mu\nu}_A A^n_\nu + \frac{1}{2} \varphi D_\varphi \varphi$$

$$\mathcal{L}_{\rm GF} = F^{\mu}F_{\mu} - F_5F_5,$$

$$F_{\mu}^n = -\left(\partial^{\nu}h_{\mu\nu}^n - \frac{1}{2}\partial_{\mu}h^n + \frac{1}{\sqrt{2}}m_nA_{\mu}^n\right),$$

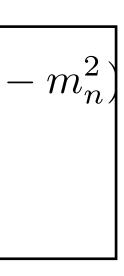
$$F_5^n = -\left(\frac{1}{2}m_nh^n - \frac{1}{\sqrt{2}}\partial^{\mu}A_{\mu}^n + \sqrt{\frac{3}{2}}m_n\varphi^n\right)$$

 $ds^{2} = e^{2A(z)} (\eta_{\mu\nu} dx^{\mu} dx^{\nu} - dz^{2})$ 

$$A(z) = -\ln(kz)$$

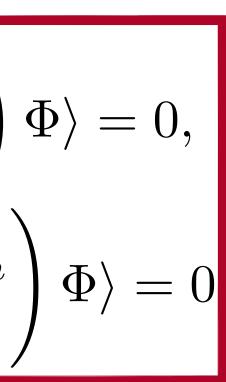
- K gravitons, O mode + tower of massive states
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- e/radion is physical, higher modes higgsed/goldstones

Gauge Fixing Lagrangian



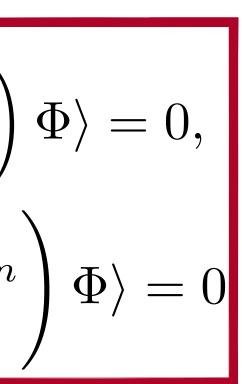
Ward Identities 
$$\begin{array}{l} \langle \mathbf{T}F_{\mu}(x)\Phi\rangle = \langle \mathbf{T}F_{5}(x)\Phi\rangle = 0 \\ \\ \langle \mathbf{T}\left(\partial^{\nu}(h_{\mu\nu}^{n} - \frac{1}{2}\eta_{\mu\nu}h^{n}) + \frac{1}{\sqrt{2}}m_{n}A_{\mu}^{n}\right) \\ \\ \langle \mathbf{T}\left(\frac{1}{2}m_{n}h^{n} - \frac{1}{\sqrt{2}}\partial^{\mu}A_{\mu}^{n} + \sqrt{\frac{3}{2}}m_{n}\varphi^{n}\right) \end{array}$$

Any External on-shell physical fields after LSZ amputation

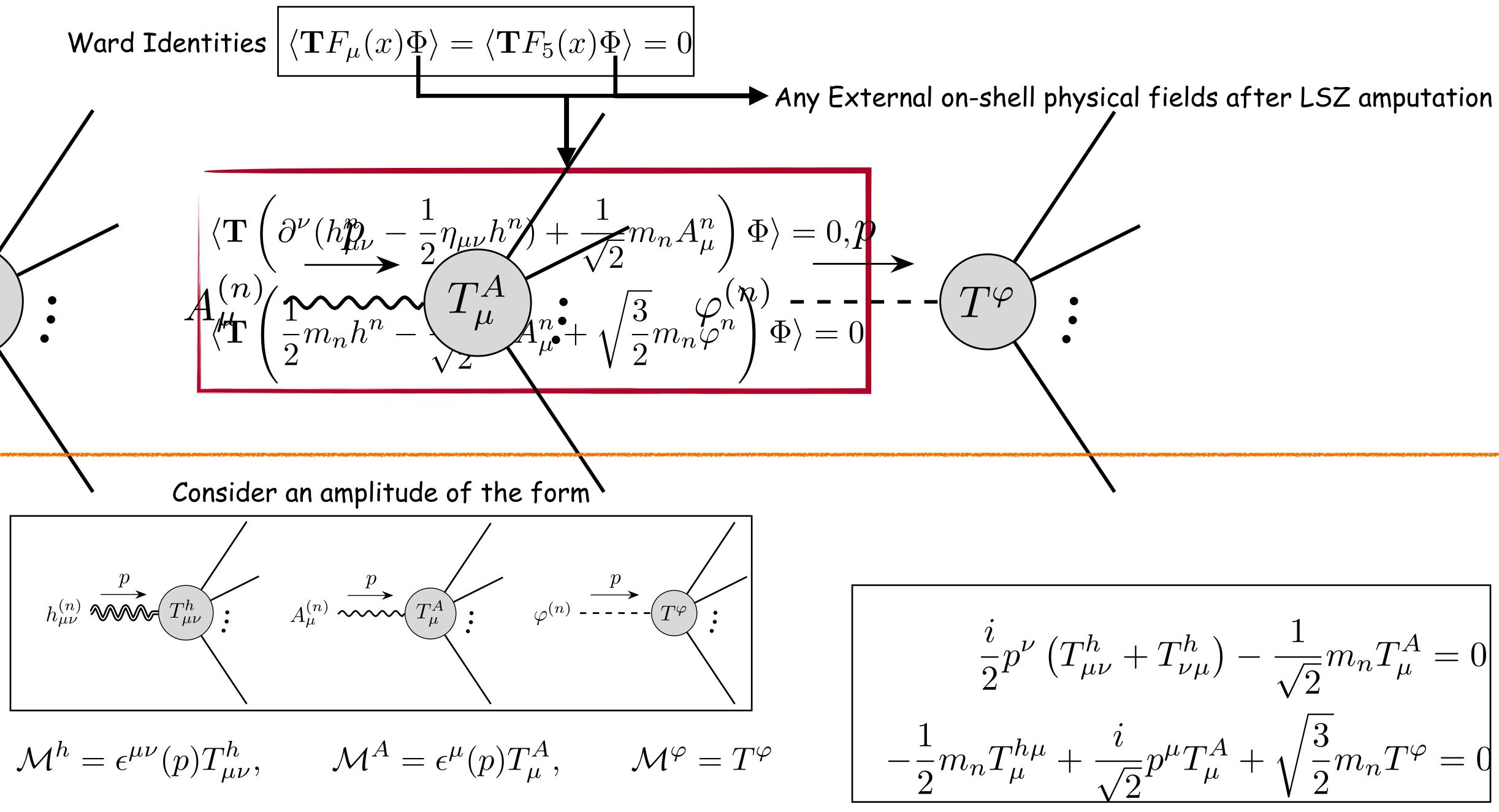


$$\begin{aligned} \text{Vard Identities} \quad & \left\langle \mathbf{T}F_{\mu}(x)\Phi \right\rangle = \langle \mathbf{T}F_{5}(x)\Phi \rangle = 0 \\ \\ & \left\langle \mathbf{T}\left(\partial^{\nu}(h_{\mu\nu}^{n} - \frac{1}{2}\eta_{\mu\nu}h^{n}) + \frac{1}{\sqrt{2}}m_{n}A_{\mu}^{n}\right) \\ \\ & \left\langle \mathbf{T}\left(\frac{1}{2}m_{n}h^{n} - \frac{1}{\sqrt{2}}\partial^{\mu}A_{\mu}^{n} + \sqrt{\frac{3}{2}}m_{n}\varphi^{n}\right) \\ \end{aligned} \end{aligned}$$

Any External on-shell physical fields after LSZ amputation







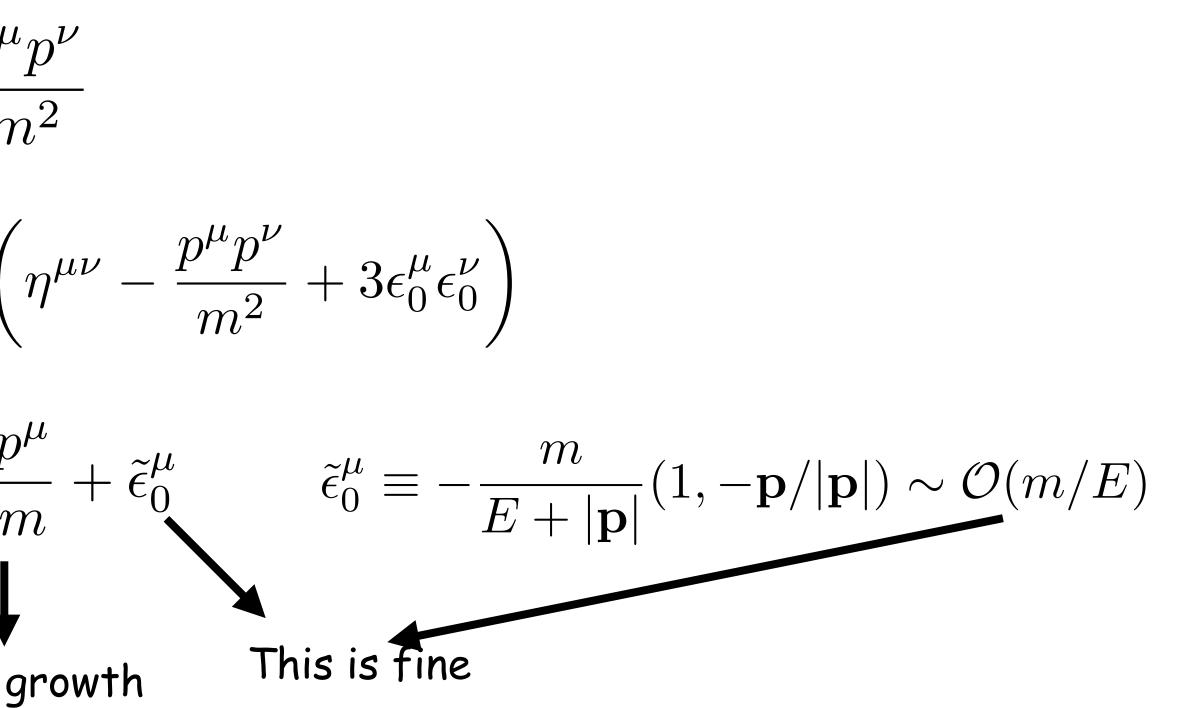


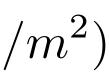


Consider the longitudinal polarizations of the KK graviton

$$\begin{split} \varepsilon_{0}^{\mu\nu} &= \frac{1}{\sqrt{6}} \left( \varepsilon_{+}^{\mu} \varepsilon_{-}^{\nu} + \varepsilon_{-}^{\mu} \varepsilon_{+}^{\nu} + 2\varepsilon_{0}^{\mu} \varepsilon_{0}^{\nu} \right) & \text{The bad high end} \\ \text{Jsing the polarization sum} & \sum_{\lambda=\pm,0} \varepsilon_{\lambda}^{\mu} \varepsilon_{\lambda}^{\nu*} = -\eta_{\mu\nu} + \frac{p^{\mu}}{m} \\ \text{Rewrite the longitudinal polarizations as} & \varepsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left( \\ \text{And separate out the bad high energy behaviour} & \varepsilon_{0}^{\mu} \equiv \frac{p}{n} \\ & \downarrow \\ \text{Bad High Energy g} \\ \varepsilon_{0}^{\mu\nu} &= \tilde{\varepsilon}_{0}^{\mu\nu} + \frac{1}{\sqrt{6}} \left( \eta^{\mu\nu} + 2 \frac{p^{\mu} p^{\nu}}{m^{2}} + 3 \frac{p^{\mu} \tilde{\varepsilon}_{0}^{\nu} + p^{\nu} \tilde{\varepsilon}_{0}^{\mu}}{m} \right) \\ & \mathcal{O} \left( \frac{m^{2}}{E^{2}} \right) \end{split}$$

energy growth part comes from the last term  $~~\epsilon_0^{\mu
u}\sim {\cal O}(E^2/m^2)$  .





Use Ward Identities to write the scattering amplitude as

$$T^{h}_{\mu\nu}\epsilon^{\mu\nu}_{0} = T^{\varphi} + \mathcal{O}(s^{0})$$

KK Goldstone Theorem :

Scattering Amplitude of longitudinally polarized KK graviton=Scalar Goldstones in the High Energy Limit

$$T^h_{\mu\nu}\epsilon^{\mu\nu}_{\pm 1} = -iT^A_\mu\epsilon^\mu_\pm + \mathcal{O}(s^0)$$

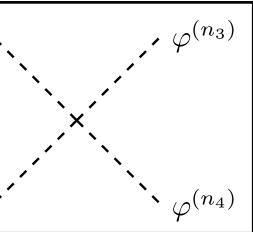
 $\mathcal{O}(m^2/E^2)$ 

Similarly for Helicity 1 one states

$$\begin{split} & \underbrace{\varphi_{(n_2)}^{(n_2)} \qquad \varphi_{(n_3)}^{(n_4)} \qquad \sum_{S^{(n_3)}}^{(n_3)} \qquad \varphi_{(n_4)}^{(n_4)} \qquad \sum_{S^{(n_2)}}^{(n_2)} \qquad \varphi_{(n_4)}^{(n_4)} \qquad \sum_{S^{(n_2)}}^{(n_2)} \qquad \varphi_{(n_4)}^{(n_4)} \qquad \sum_{S^{(n_2)}}^{(n_4)} \qquad \sum_{S^{(n_2)}}^{(n_4)} \qquad \sum_{S^{(n_2)}}^{(n_4)} \qquad \sum_{S^{(n_2)}}^{(n_4)} \qquad \sum_{S^{(n_4)}}^{(n_4)} \qquad \sum_{S^{(n_4)}}^{(n_$$

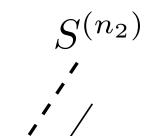
$$\mathcal{M}\left[S^{(n_1)}S^{(n_2)} \to \varphi^{(n_3)}\varphi^{(n_4)}\right] = \frac{\kappa^2 s}{32} (1 - \cos 2\theta) \left\langle k^{(n)}k \right\rangle$$

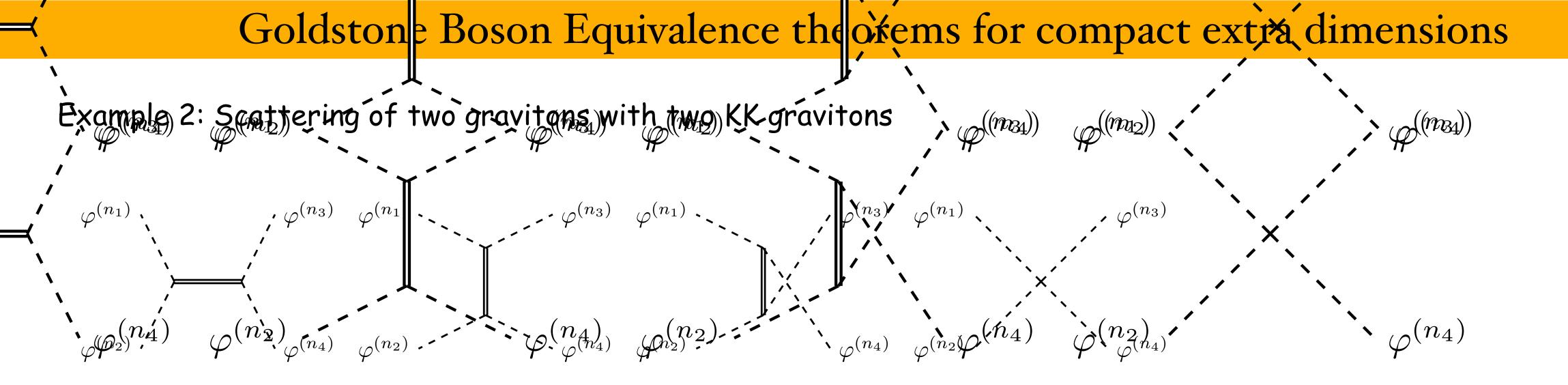
t extra dimensions  $\varphi^{(n_4)}$ 



$$\mathcal{L}_m = \sqrt{G} \left( \frac{1}{2} G^{MN} \partial_M S \partial_N S - \frac{1}{2} M_S^2 S^2 \right)$$

 $k^{(n)}f^{(n)}_Sf^{(n)}_S
angle + \mathcal{O}(s^0)$  Agrees with Unitary gauge calculation





Expectation from Goldstone Equivalence Theorem

$$\mathcal{M}\left[h_L^{(n_1)}h_L^{(n_2)} \to h_L^{(n_3)}h_L^{(n_4)}\right] = \mathcal{M}\left[\varphi^{(n_1)}\varphi^{(n_2)} \to \varphi^{(n_3)}\varphi^{(n_4)}\right] + \mathcal{O}(s^0)$$

Agrees with Unitary gauge calculation

$$\mathcal{M}\left[\varphi^{(n_1)}\varphi^{(n_2)} \to \varphi^{(n_3)}\varphi^{(n_4)}\right] = \frac{\kappa^2 s}{64} \frac{(\cos 2\theta + 7)^2}{\sin^2 \theta} \left\langle k^{(n)} k^{(n)} k^{(n)} k^{(n)} \right\rangle + \mathcal{O}(s^0)$$

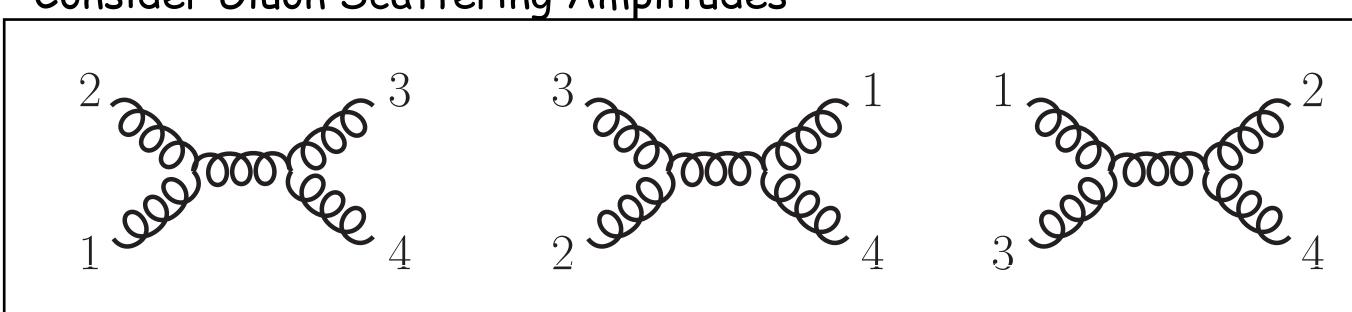
Chivukula, Gill, Mohan, DS, Simmons Wang et al, Phys. Rev. D 109 (2024), 1 Phys. Rev. D 109 (2024), 7

# A Short Summary of Double Copy: BCJ double copy : Bern-Carrasco-Johannson 2010



# A Short Summary of Double Copy: BCJ double copy : Bern-Carrasco-Johannson 2010

Consider Gluon Scattering Amplitudes



$$n_{s} = -\frac{1}{2} \left\{ \left[ (\varepsilon_{1} \cdot \varepsilon_{2}) p_{1}^{\mu} + 2(\varepsilon_{1} \cdot p_{2}) \varepsilon_{2}^{\mu} - (1 \leftrightarrow 2) \right] \left[ (\varepsilon_{3} \cdot \varepsilon_{4}) p_{3\mu} + 2(\varepsilon_{3} \cdot p_{4}) \varepsilon_{4} + s \left[ (\varepsilon_{1} \cdot \varepsilon_{3}) (\varepsilon_{2} \cdot \varepsilon_{4}) - (\varepsilon_{1} \cdot \varepsilon_{4}) (\varepsilon_{2} \cdot \varepsilon_{3}) \right] \right\},$$

The Contact Interaction is factored into the above definition by suitable kinematic reshuffling

$$i\mathcal{A}_{4}^{\text{tree}} = g^{2} \left( \frac{n_{s}c_{s}}{s} + \frac{n_{t}c_{t}}{t} + \frac{n_{u}c_{u}}{u} \right)$$

$$c_{s} = -2f^{a_{1}a_{2}b}f^{ba_{3}a_{4}} \qquad c_{t}n_{t} = c_{s}n_{s}\big|_{1\to 2\to 3\to 1}$$

$$c_{u}n_{u} = c_{s}n_{s}\big|_{1\to 3\to 2\to 1}$$

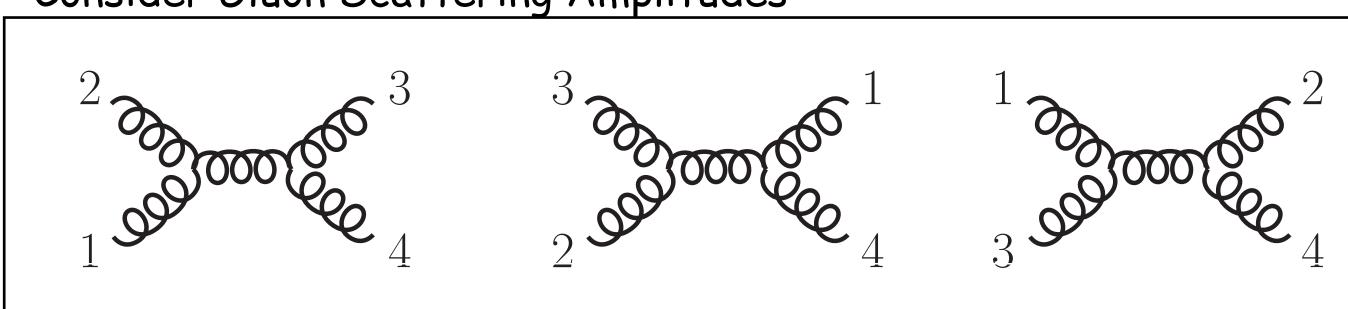
$$c_{4\mu} - (3 \leftrightarrow 4)\big]$$



# A Short Summary of Double Copy: BCJ double copy : Bern-Carrasco-Johannson 2010

 $c_s =$ 

Consider Gluon Scattering Amplitudes



$$n_{s} = -\frac{1}{2} \left\{ \left[ (\varepsilon_{1} \cdot \varepsilon_{2}) p_{1}^{\mu} + 2(\varepsilon_{1} \cdot p_{2}) \varepsilon_{2}^{\mu} - (1 \leftrightarrow 2) \right] \left[ (\varepsilon_{3} \cdot \varepsilon_{4}) p_{3\mu} + 2(\varepsilon_{3} \cdot p_{4}) \varepsilon_{4\mu} - (3 \leftrightarrow 4) \right] + s \left[ (\varepsilon_{1} \cdot \varepsilon_{3}) (\varepsilon_{2} \cdot \varepsilon_{4}) - (\varepsilon_{1} \cdot \varepsilon_{4}) (\varepsilon_{2} \cdot \varepsilon_{3}) \right] \right\},$$

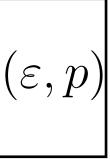
$$(1)$$

The Contact Interaction is factored into the above definition by suitable kinematic reshuffling Gauge Invariance demands that the amplitude must vanish under polarization to momentum replacement

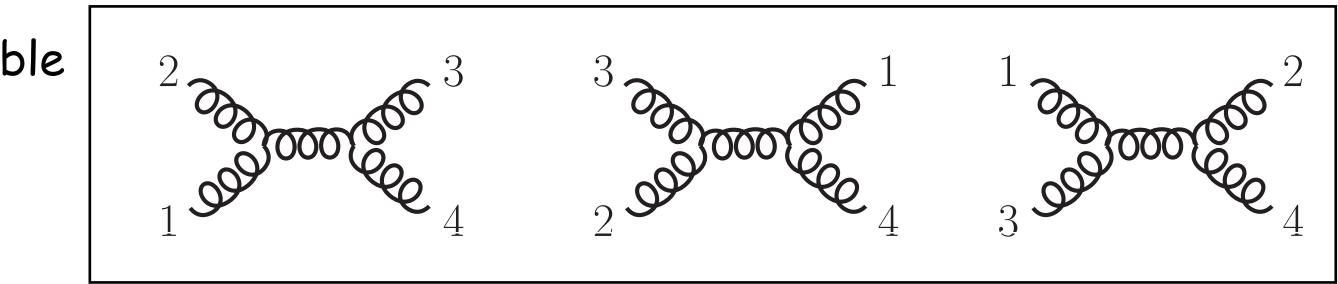
$$n_{s}|_{\varepsilon_{4} \to p_{4}} = -\frac{s}{2} \Big[ (\varepsilon_{1} \cdot \varepsilon_{2}) \big( (\varepsilon_{3} \cdot p_{2}) - (\varepsilon_{3} \cdot p_{1}) \big) + \operatorname{cyclic}(1, 2, 3) \Big] \equiv s \,\alpha(\varepsilon, p) \qquad \left[ \frac{n_{s}c_{s}}{s} + \frac{n_{t}c_{t}}{t} + \frac{n_{u}c_{u}}{u} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) \,\alpha(z_{s} + c_{t} + c_{u}) + c_{s}(z_{s} + c_{t} + c_{u}) \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) \,\alpha(z_{s} + c_{t} + c_{u}) \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) \,\alpha(z_{s} + c_{t} + c_{u}) \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) \,\alpha(z_{s} + c_{t} + c_{u}) \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) \,\alpha(z_{s} + c_{t} + c_{u}) \,\alpha(z_{s} + c_{t} + c_{u}) \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) \,\alpha(z_{s} + c_{t}) \,\alpha(z_{s} + c_{t$$

$$i\mathcal{A}_{4}^{\text{tree}} = g^{2} \left( \frac{n_{s}c_{s}}{s} + \frac{n_{t}c_{t}}{t} + \frac{n_{u}c_{u}}{u} \right)$$
$$c_{s} = -2f^{a_{1}a_{2}b}f^{ba_{3}a_{4}} \qquad c_{t}n_{t} = c_{s}n_{s} \big|_{1 \to 2 \to 3 \to 1}$$
$$c_{u}n_{u} = c_{s}n_{s} \big|_{1 \to 3 \to 2 \to 1}$$



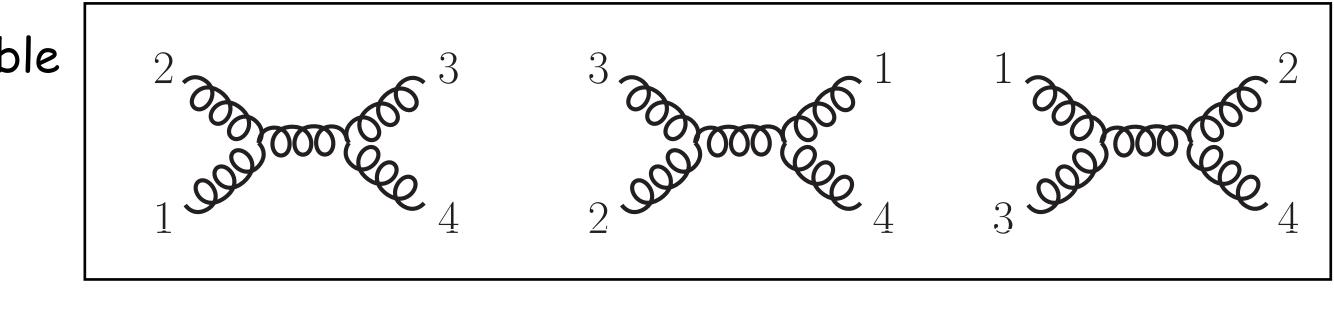


Color and Kinematic Factors are mutually Interchangable



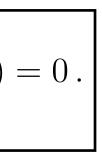
Color and Kinematic Factors are mutually Interchangable

$$i\mathcal{A}_{4}^{\text{tree}} = g^{2} \left( \frac{n_{s}c_{s}}{s} + \frac{n_{t}c_{t}}{t} + \frac{n_{u}c_{u}}{u} \right)$$
  
Massless Gravity amplitude



# Diffeomorphism

 $\left. \left| \frac{n_s^2}{t} + \frac{n_t^2}{t} + \frac{n_u^2}{u} \right|_{\varepsilon_4^{\mu\nu} \to p_4^{\mu} \varepsilon_4^{\nu} + p_4^{\nu} \varepsilon_4^{\mu}} = \right.$ Invariance  $= 2(n_s + n_t + n_u) \alpha(\varepsilon, p) = 0.$ 

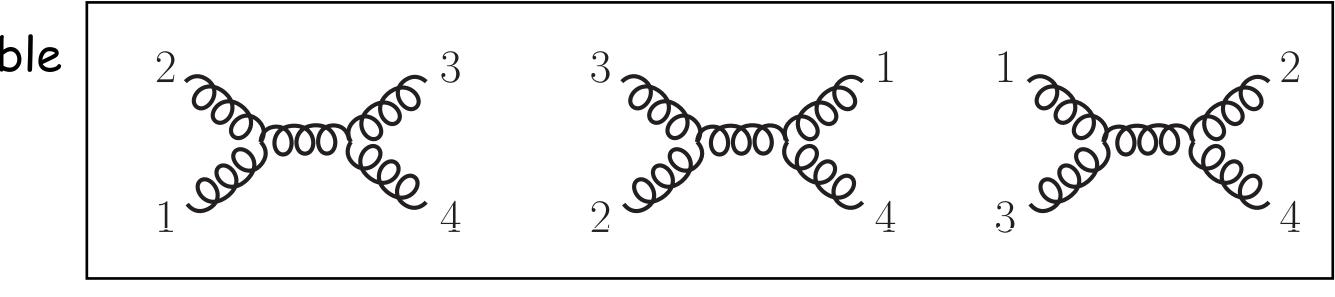


Color and Kinematic Factors are mutually Interchangable

$$i\mathcal{A}_{4}^{\text{tree}} = g^{2} \left( \frac{n_{s}c_{s}}{s} + \frac{n_{t}c_{t}}{t} + \frac{n_{u}c_{u}}{u} \right)$$

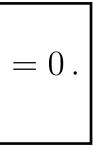
$$Massless Gravity amplitude$$

$$i\mathcal{A}_{4}^{\text{tree}}\Big|_{\substack{c_{i} \to \tilde{n}_{i} \\ g \to \kappa/2}} \equiv i\mathcal{M}_{4}^{\text{tree}} = \left(\frac{\kappa}{2}\right)^{2} \left(\frac{n_{s}^{2}}{s} + \frac{n_{t}^{2}}{t} + \frac{n_{u}^{2}}{u}\right)$$



# Diffeomorphism

# $\left|\frac{n_s^2}{t} + \frac{n_t^2}{t} + \frac{n_u^2}{u}\right|_{\varepsilon_4^{\mu\nu} \to p_4^{\mu} \varepsilon_4^{\nu} + p_4^{\nu} \varepsilon_4^{\mu}} =$ Invariance $= 2(n_s + n_t + n_u) \alpha(\varepsilon, p) = 0.$

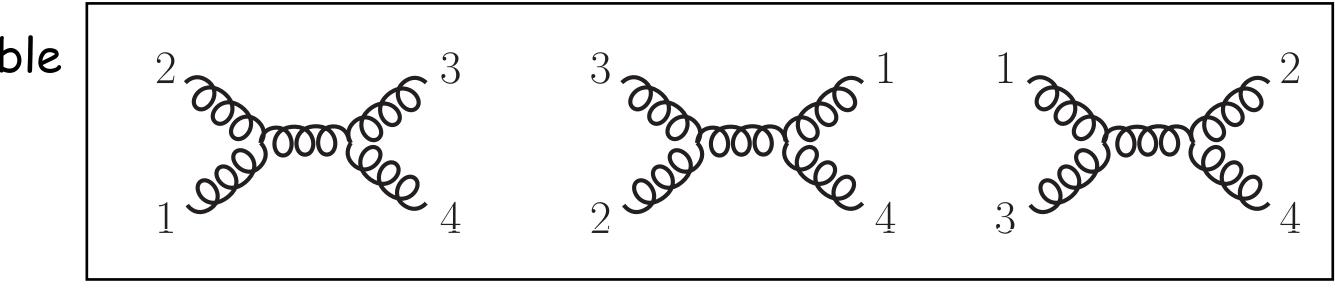


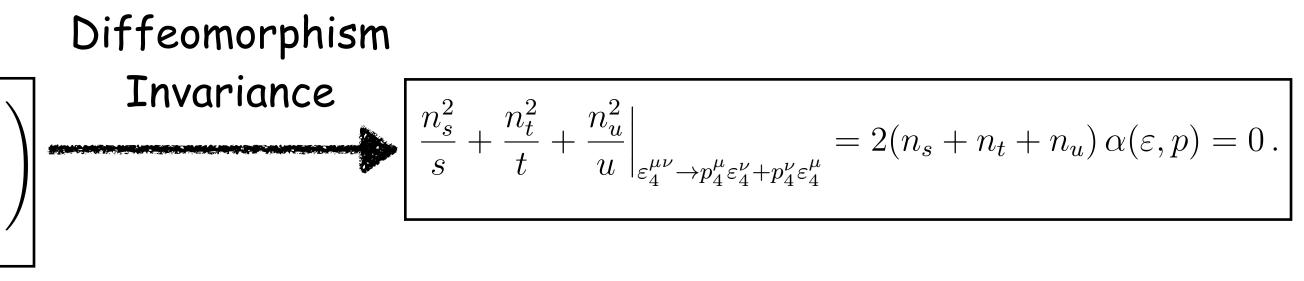
Color and Kinematic Factors are mutually Interchangable

$$i\mathcal{A}_{4}^{\text{tree}} = g^{2} \left( \frac{n_{s}c_{s}}{s} + \frac{n_{t}c_{t}}{t} + \frac{n_{u}c_{u}}{u} \right)$$

$$Massless Gravity amplitude$$

$$i\mathcal{A}_{4}^{\text{tree}}\Big|_{\substack{c_{i} \to \tilde{n}_{i} \\ g \to \kappa/2}} \equiv i\mathcal{M}_{4}^{\text{tree}} = \left(\frac{\kappa}{2}\right)^{2} \left(\frac{n_{s}^{2}}{s} + \frac{n_{t}^{2}}{t} + \frac{n_{u}^{2}}{u}\right)$$

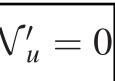




 $\mathcal{M}[h_{n_1}^L(k_1), ..., h_{n_N}^L(k_N), \Phi]$  $= \mathcal{M}[\phi_{n_1}(k_1), \ldots, \phi_{n_N}(k_N), \Phi]$ 

# Statement of Gravitational Equivalence Theorem

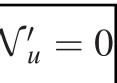
# $\mathcal{T}[A_{n_1}^{aL}A_{n_2}^{bL} \rightarrow A_{n_3}^{cL}A_{n_4}^{dL}] = \mathcal{T}[A_{n_1}^{a5}A_{n_2}^{b5} \rightarrow A_{n_3}^{c5}A_{n_4}^{d5}]$ Statement of 5D Gauge theory Equivalence Theorem $\begin{aligned} \mathcal{T}'[4A_L^n] &= g^2 \left( \frac{\mathcal{C}_s \mathcal{N}'_s}{s_0} + \frac{\mathcal{C}_t \mathcal{N}'_t}{t_0} + \frac{\mathcal{C}_u \mathcal{N}'_u}{u_0} \right) \end{aligned} \qquad \begin{array}{l} \text{Kinematic Jacobi} \\ \text{Identity} \\ \hline \mathcal{T}'[4A_5^n] &= g^2 \left( \frac{\mathcal{C}_s \widetilde{\mathcal{N}}'_s}{s_0} + \frac{\mathcal{C}_t \widetilde{\mathcal{N}}'_t}{t_0} + \frac{\mathcal{C}_u \widetilde{\mathcal{N}}'_u}{u_0} \right) \end{aligned} \qquad \begin{array}{l} \text{Kinematic Jacobi} \\ \text{Identity} \\ \hline \mathcal{N}'_s + \mathcal{N}'_t + \mathcal{N}'_u = 0 \end{aligned}$



 $\mathcal{M}[h_{n_1}^L(k_1), ..., h_{n_N}^L(k_N), \Phi]$  $= \mathcal{M}[\phi_{n_1}(k_1), \dots, \phi_{n_N}(k_N), \Phi]$ 

# Statement of Gravitational Equivalence Theorem

# $\mathcal{T}[A_{n_1}^{aL}A_{n_2}^{bL} \rightarrow A_{n_3}^{cL}A_{n_4}^{dL}] = \mathcal{T}[A_{n_1}^{a5}A_{n_2}^{b5} \rightarrow A_{n_3}^{c5}A_{n_4}^{d5}]$ Statement of 5D Gauge theory Equivalence Theorem $\left| \mathcal{T}'[4A_L^n] = g^2 \left( \frac{\mathcal{C}_s \mathcal{N}'_s}{s_0} + \frac{\mathcal{C}_t \mathcal{N}'_t}{t_0} + \frac{\mathcal{C}_u \mathcal{N}'_u}{u_0} \right) \right| \begin{array}{c} \text{Kinematic Jacobi} \\ \text{Identity} \\ \mathcal{N}'_s + \mathcal{N}'_t + \mathcal{N}'_u = 0 \end{array} \right|$ $\left| \widetilde{\mathcal{T}}'[4A_5^n] = g^2 \left( \frac{\mathcal{C}_s \widetilde{\mathcal{N}}'_s}{s_0} + \frac{\mathcal{C}_t \widetilde{\mathcal{N}}'_t}{t_0} + \frac{\mathcal{C}_u \widetilde{\mathcal{N}}'_u}{u_0} \right) \right|$



# Amplitudes/double copies and equivalence theorems

Color -> Kinematic replacement

$$(\mathcal{C}_s, \mathcal{C}_t, \mathcal{C}_u) \to (\mathcal{N}_s, \mathcal{N}_t, \mathcal{N}_u)$$

$$\begin{aligned} \overline{\mathcal{M}_0[h_L^n h_L^n \to h_L^n h_L^n]} &= c_0 g^2 \left[ \frac{(\mathcal{N}_s^0)^2}{s_0} + \frac{(\mathcal{N}_t^0)^2}{t_0} + \frac{(\mathcal{N}_u^0)^2}{u_0} \right] \\ \overline{\mathcal{M}}_0[\phi_n \phi_n \to \phi_n \phi_n] &= c_0 g^2 \left[ \frac{(\widetilde{\mathcal{N}}_s^0)^2}{s_0} + \frac{(\widetilde{\mathcal{N}}_t^0)^2}{t_0} + \frac{(\widetilde{\mathcal{N}}_u^0)^2}{u_0} \right] \end{aligned}$$

• Works for Compactified Torus on flat space-times -> Hang-He PRD 2022

• Works for a compactified spacetime with an ADS background -> Chivukula, DS, Gill, Wang et al, PRD 2023



- O Supersymmetric Structure of Compactified theories -> Chivukula, Simmons, Wang 2021
- O Scattering amplitudes for Moduli/radius stabilized geometries -> Chivukula, Foren, Mohan, DS, Simmons 2021,2023
- O Goldstone Equivalence Theorem with Matter couplings and generalized Poincare symmtries (Kac-Moody Algebras) -> Chivukula, Gill, Mohan, DS, Simmons, Wang 2023 (Josh Gill's Poster)
- O Scattering Amplitudes in models with curvature localized on D-branes/DGP model correspondence -> Chivukula, Mohan, DS< Simmons, Wang 2024
- O Scattering Amplitudes and matter couplings in massive gravity -> Gill, DS, Williams 2022 (Josh Gill's Poster)
- O Applications for KK portal dark matter models -> To appear soon

# Further works in this direction





- Compactified theories of extra dimensions -> No low energy cut-off
- Cancellations due to different diagrams reduce  $O(s^5)$  growth to O(s).
- No low energy cut-off for consistent models of stabilization
- Uncovered sum rules enforcing this cancellation
- Can show -> Analysis extends to matter on brane or bulk
- Consistent with literature on massive gravity.
- Pheno papers : Doing an unitarity analysis for DM models, ultralight radion as a candidate ...
- Theory papers : Spinor Helicity/Goldstone Equivalence calculation ?
- More connections with massive gravity community ...

# Conclusions

• Possible to double-copy a compactified gauge theory to compactified gravity for flat toroidal compactification







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