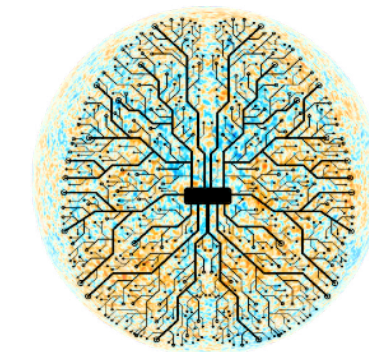
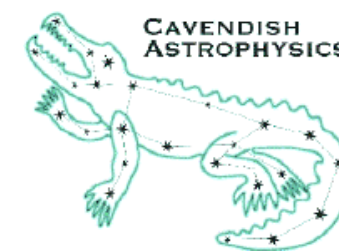


Calibrating Tension Statistics with Neural Ratio Estimators

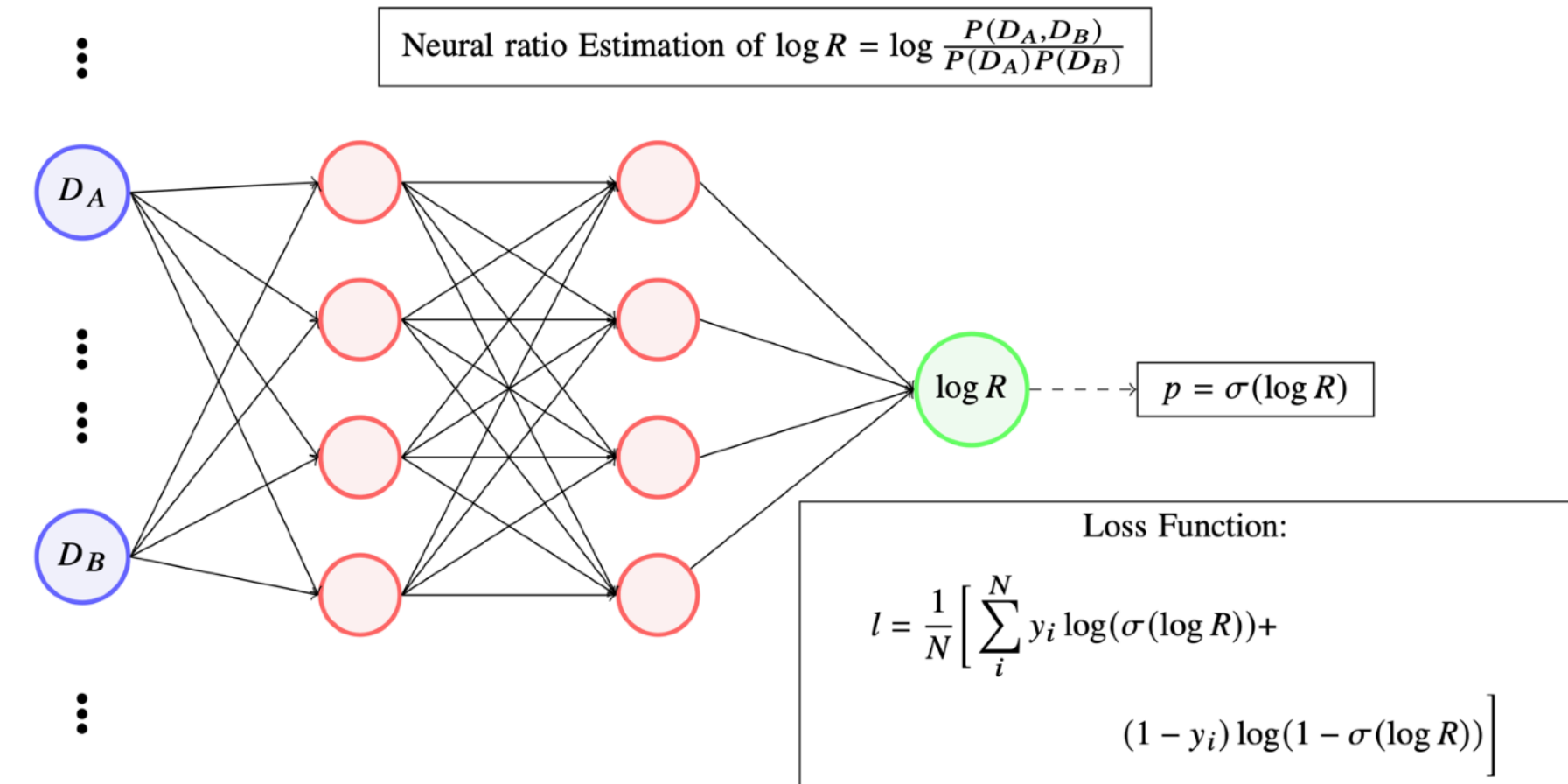
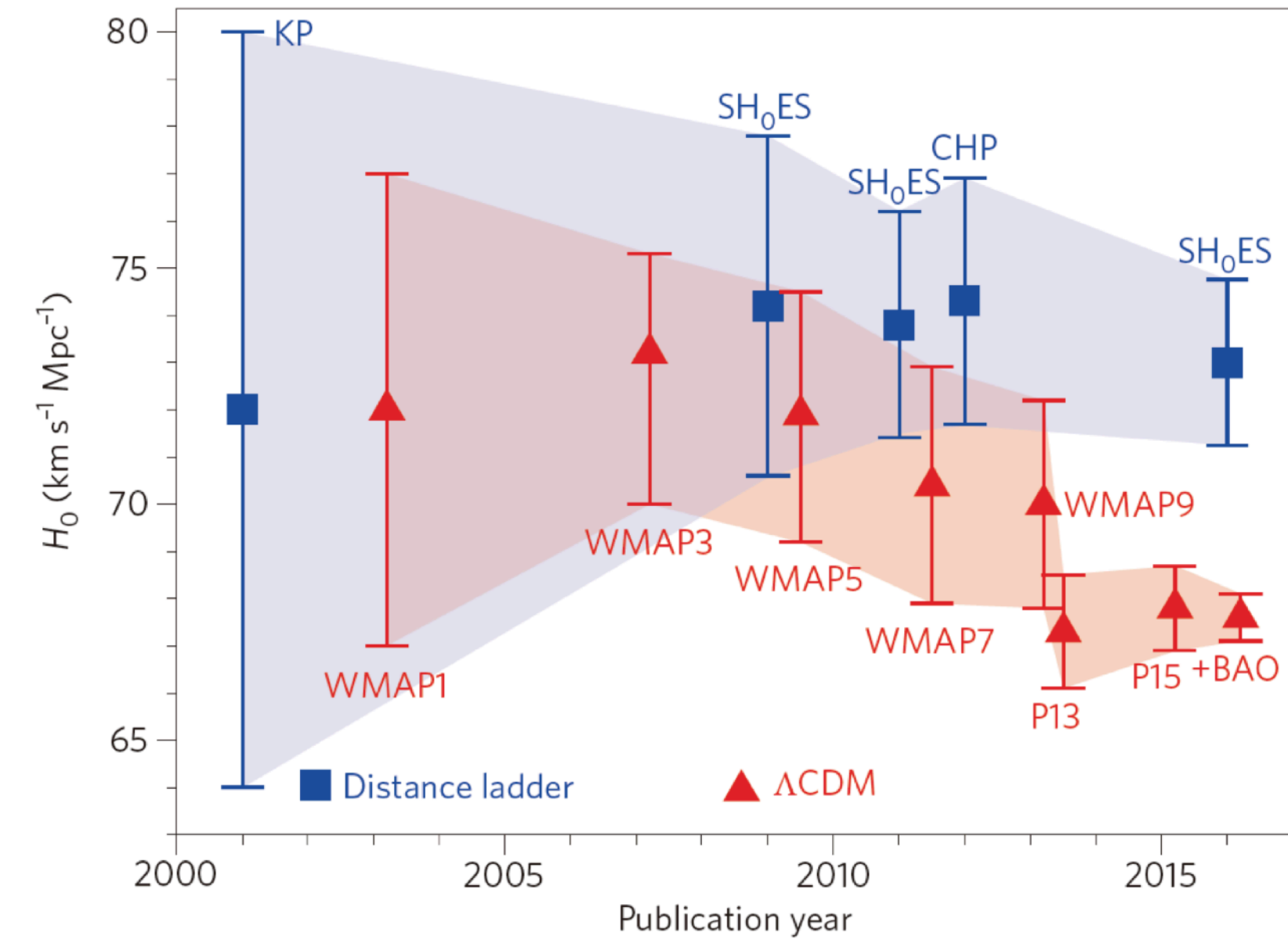
Harry Bevins

with Thomas Gessey-Jones and Will Handley
University of Cambridge



Calibrating Tensions

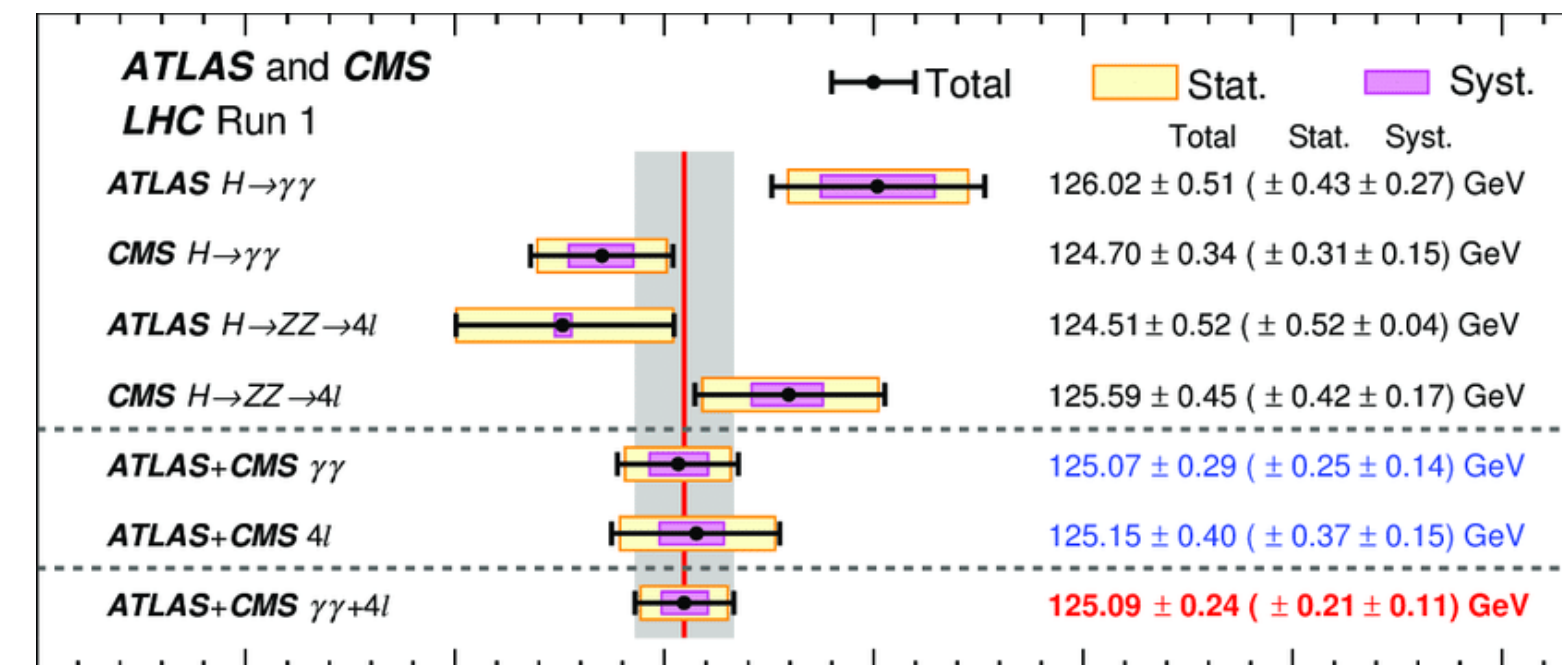
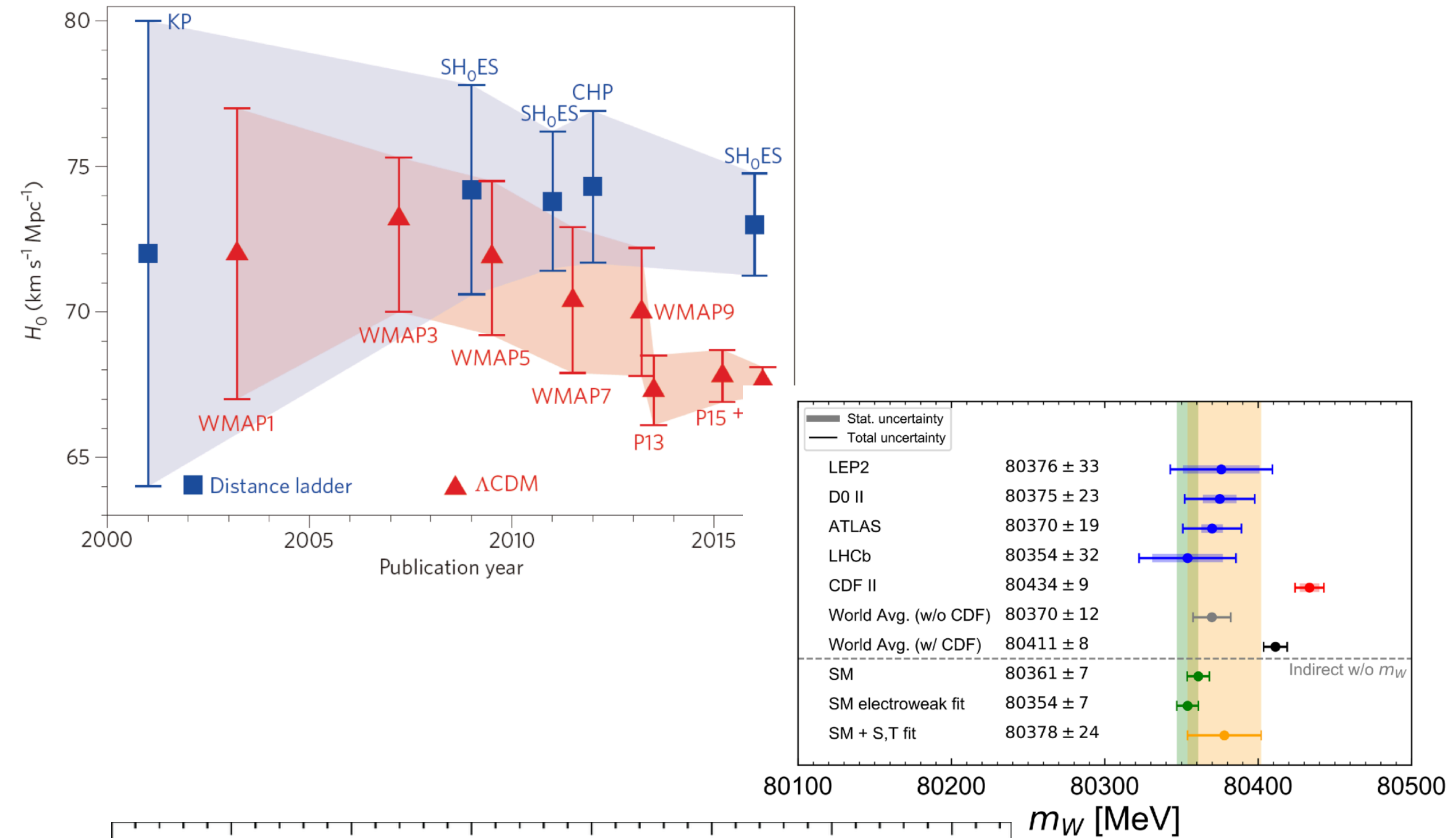
1. Why are we interested and how do we measure tension?
2. Calibrating with Neural Ratio Estimation
3. Demonstrations



Why are we interested?

Why are we interested in tension?

- Important to be able to independently observe and confirm experimental results
- When two experiments give different results we call this a tension
- Understanding where tension comes from can lead us to new physics and a better understanding of our instruments

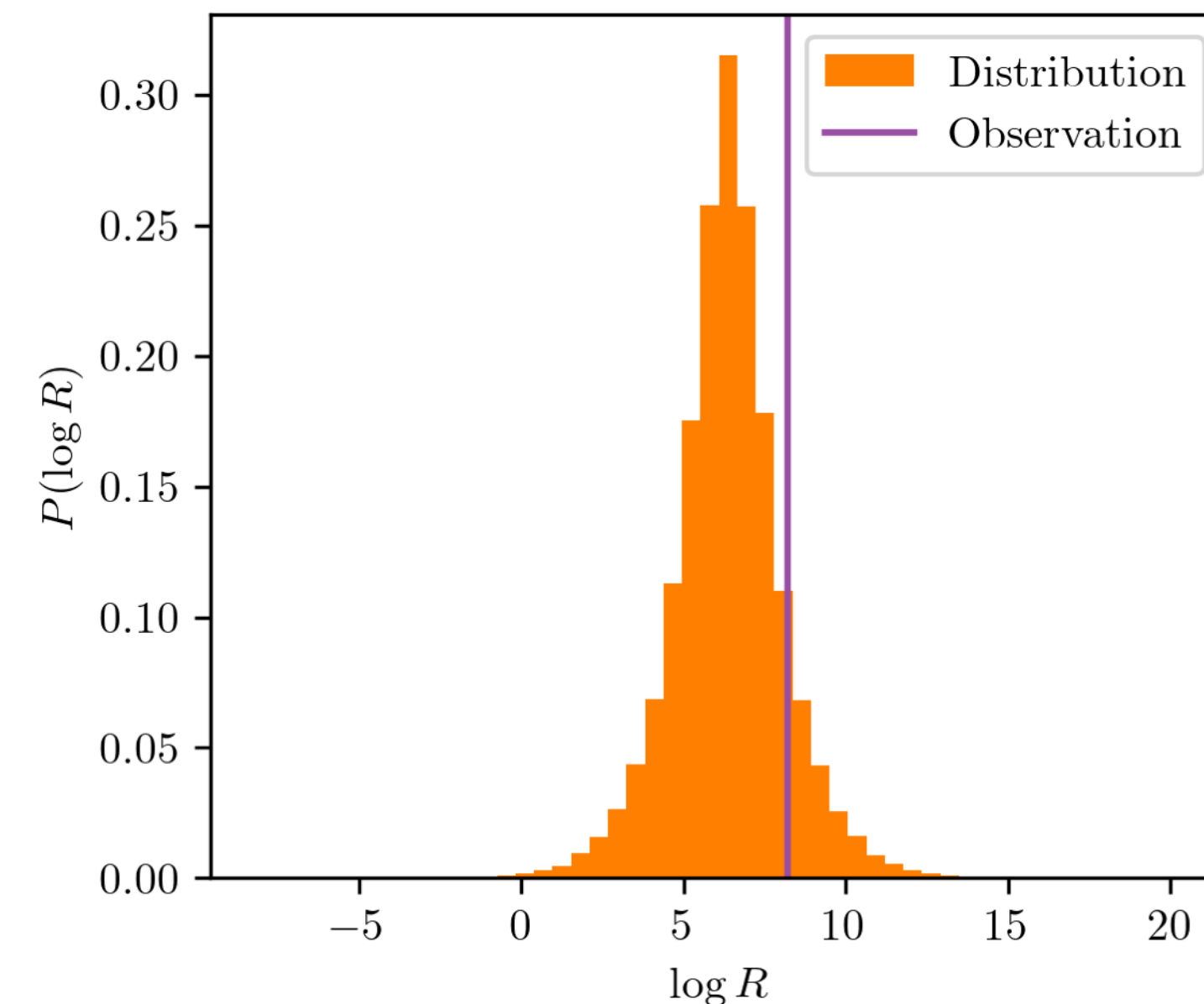
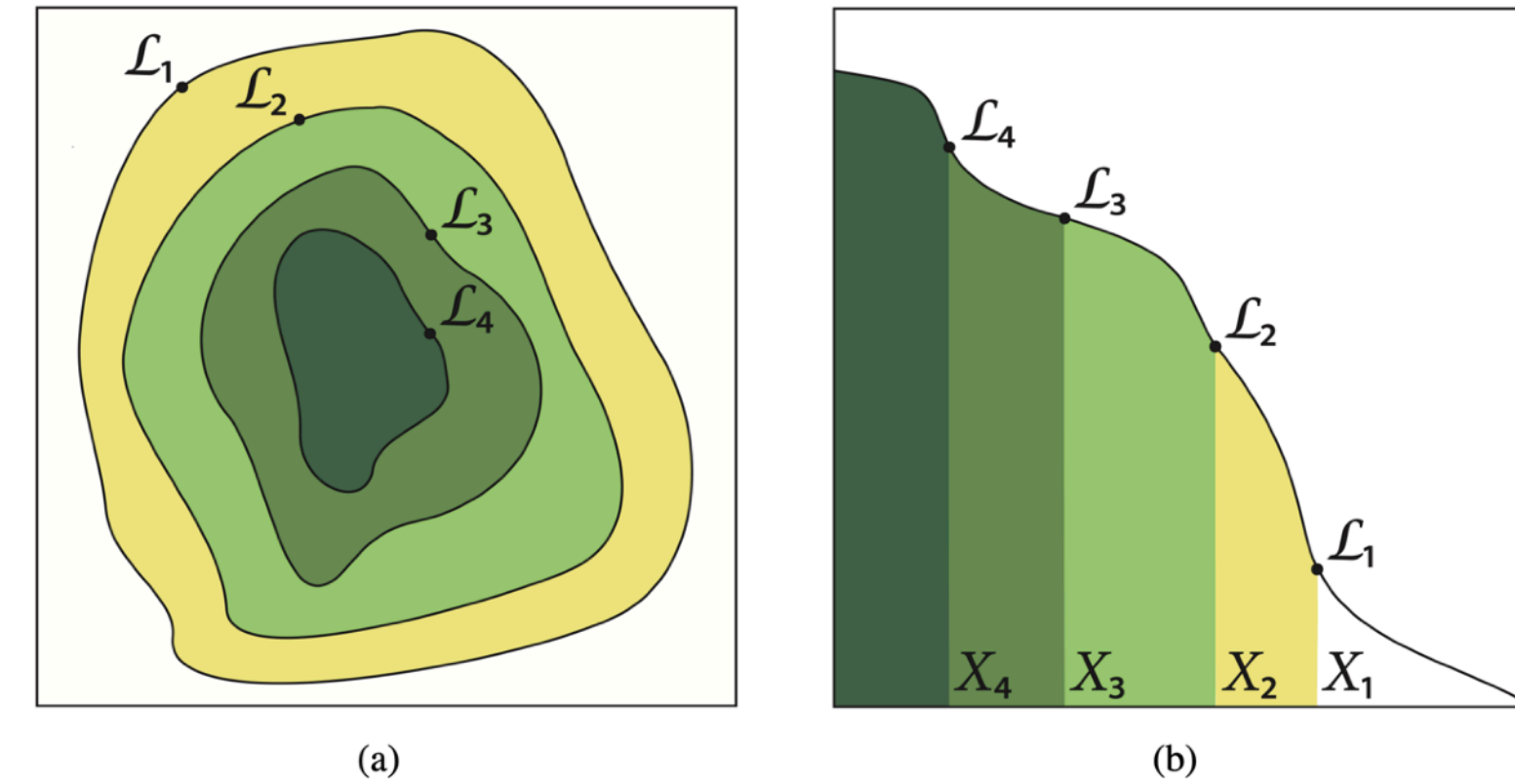


Measuring tension

- Parameter differences, goodness of fit degradation, suspiciousness (see 2012.09554 for a review)
- Here, interested in evidence ratio

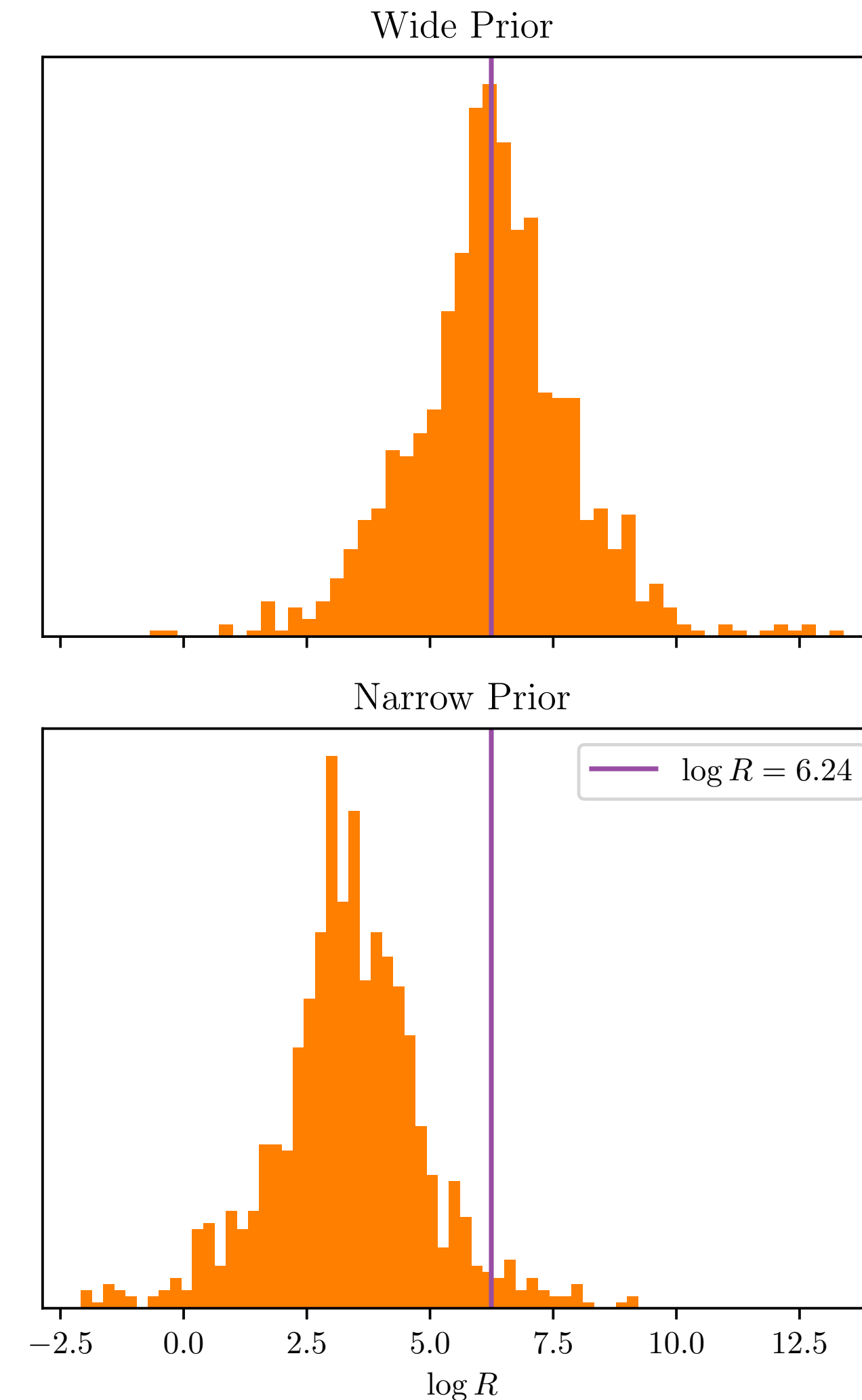
$$R = \frac{P(D_A, D_B)}{P(D_A)P(D_B)} = \frac{Z_{AB}}{Z_A Z_B}$$

- For any pair of experiments, model and prior there is a distribution of in concordance R values

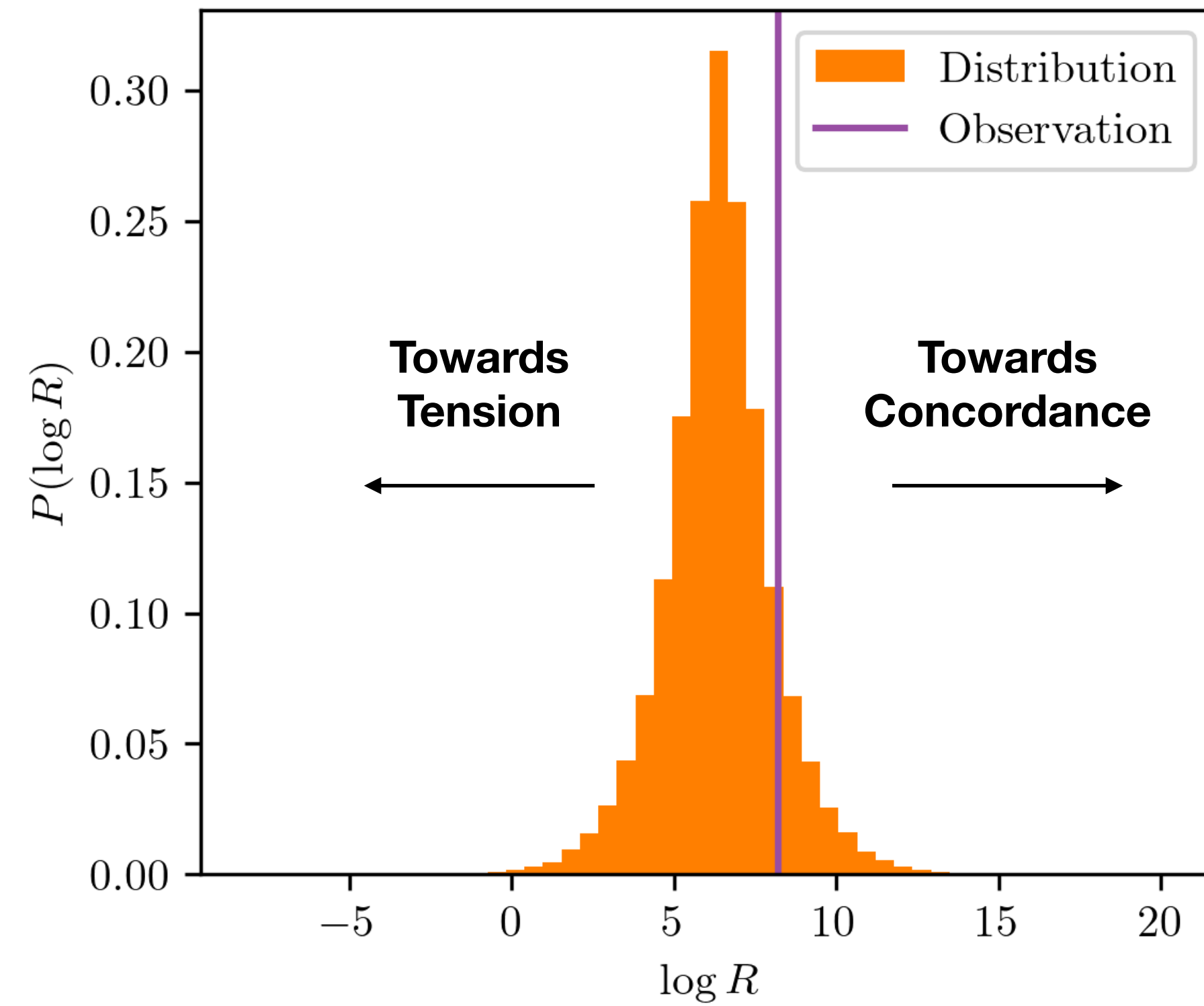
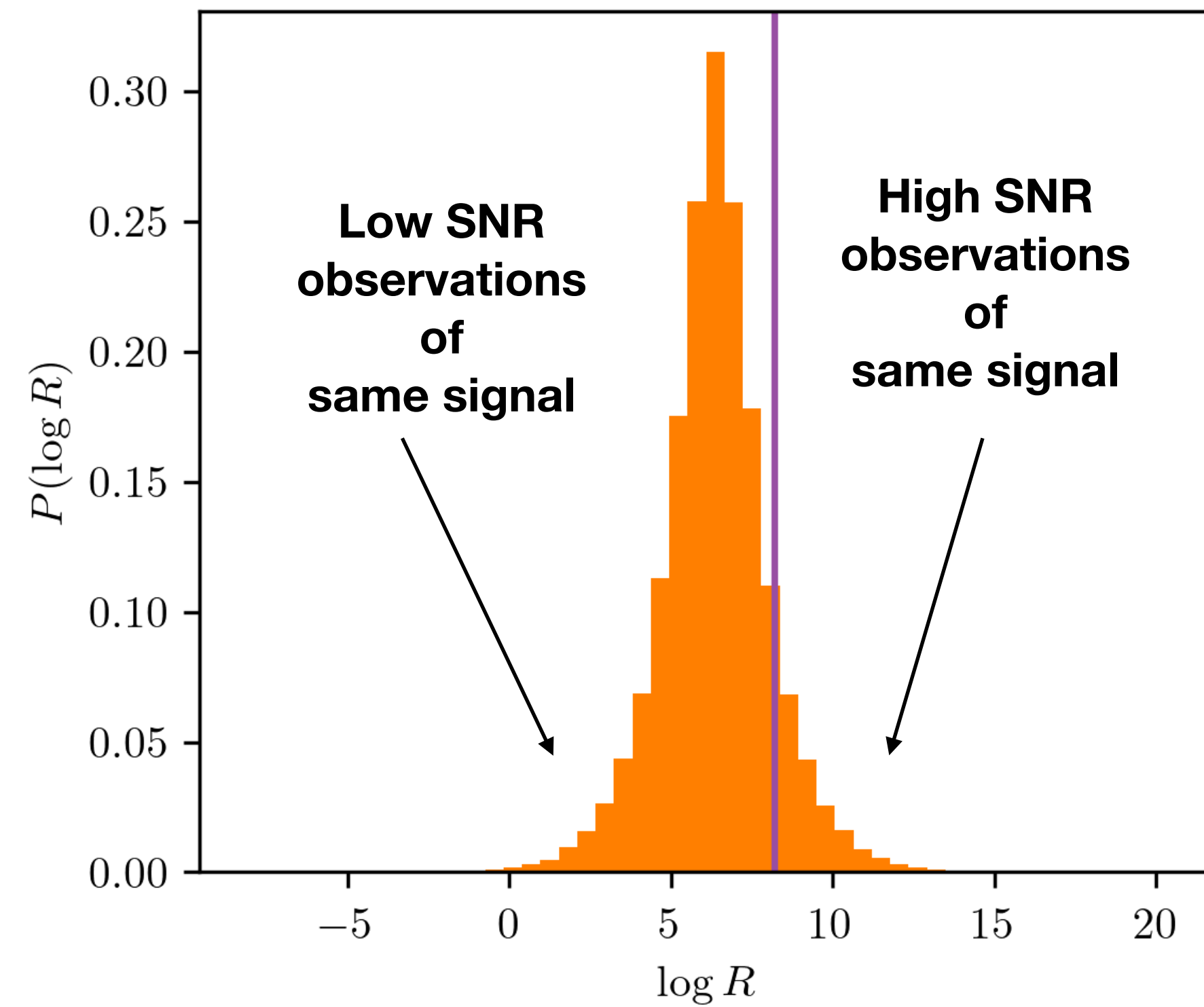


Measuring tension

- The fractional increase in our confidence in one experiment given data from another
- Dimensionally consistent and parameterisation invariant
- But prior dependent and hard to interpret
 - $R \gg 1 \rightarrow$ in concordance
 - $R \ll 1 \rightarrow$ in tension



Measuring tension



Calibrating with Neural Ratio Estimation

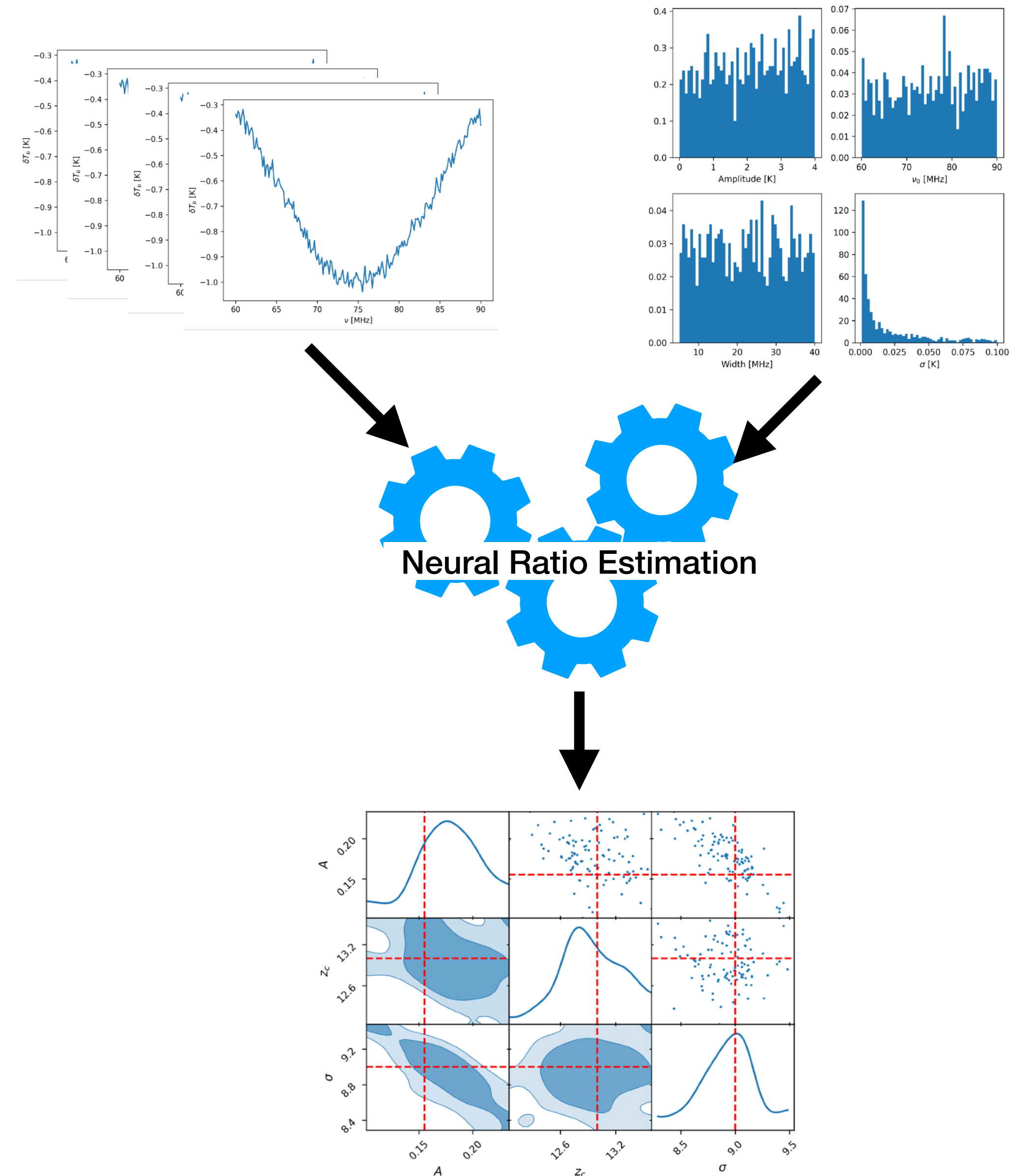
Neural Ratio Estimation

- Essentially just classifiers
- Take in two inputs A and B and estimate the probability that they are drawn from joint distribution vs disjoint

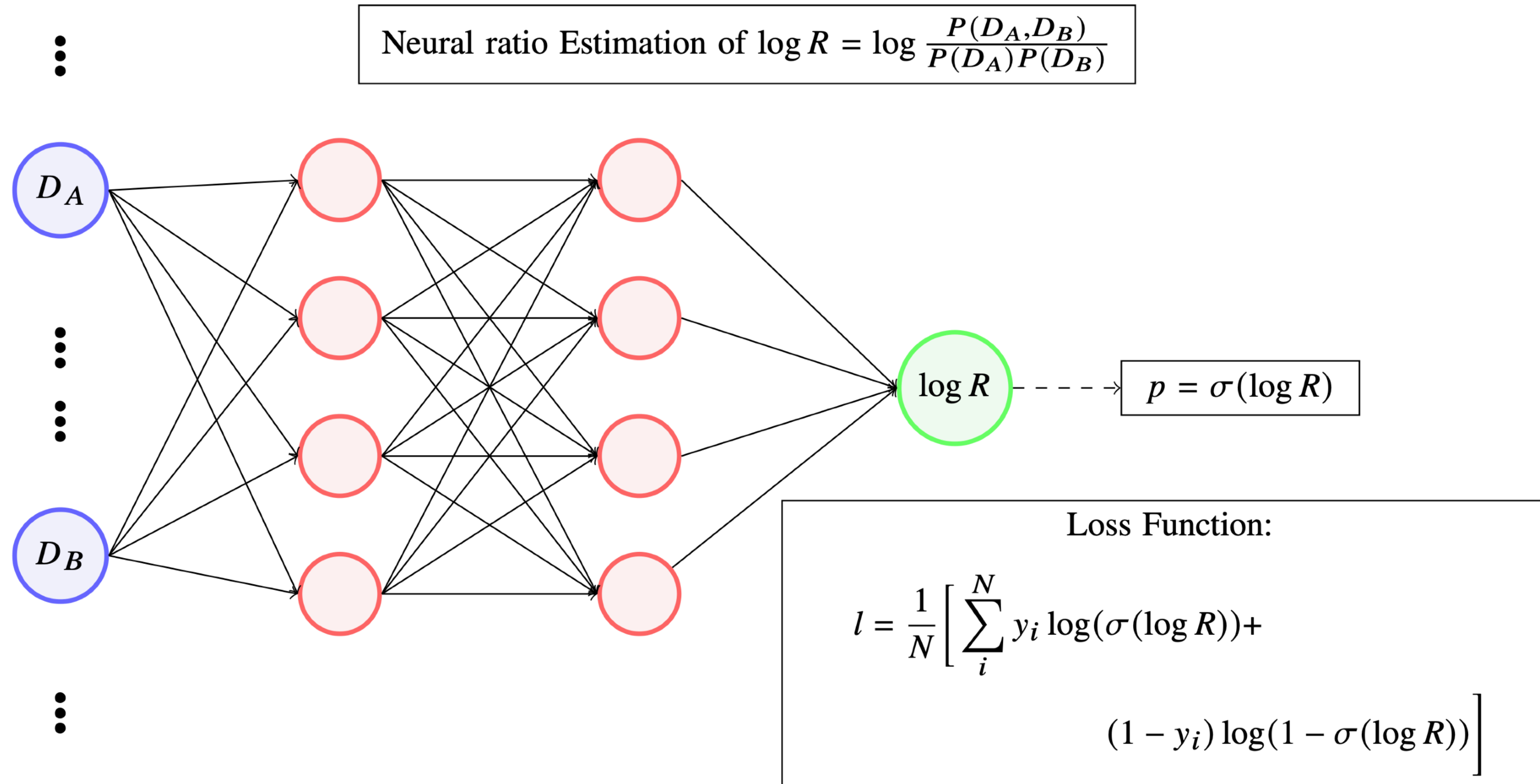
$$r = \frac{P(A, B)}{P(A)P(B)}$$

- Used for parameter inference

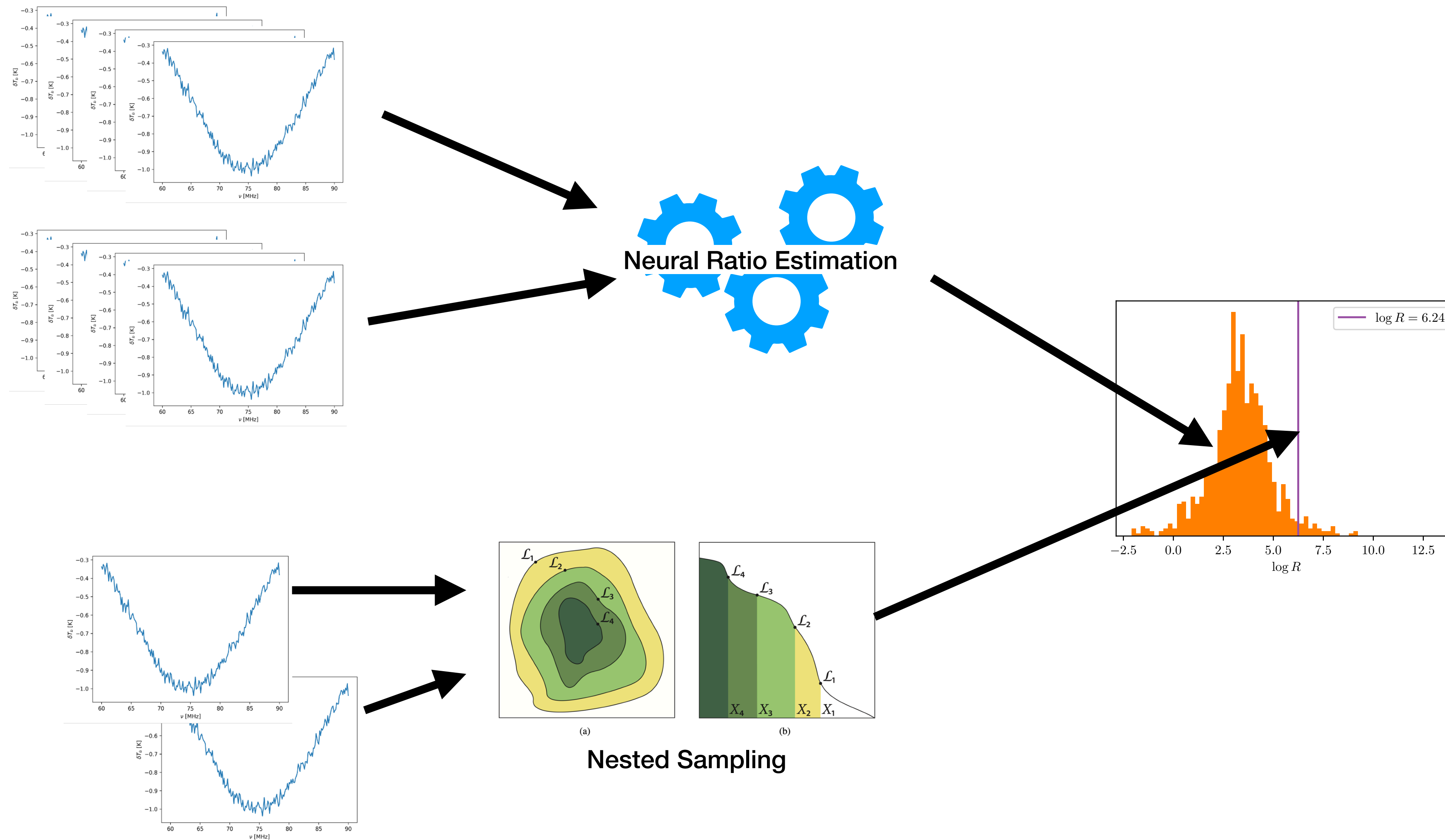
$$r = \frac{P(D, \theta)}{P(D)P(\theta)} = \frac{P(D | \theta)}{P(D)} = \frac{L(\theta)}{Z}$$



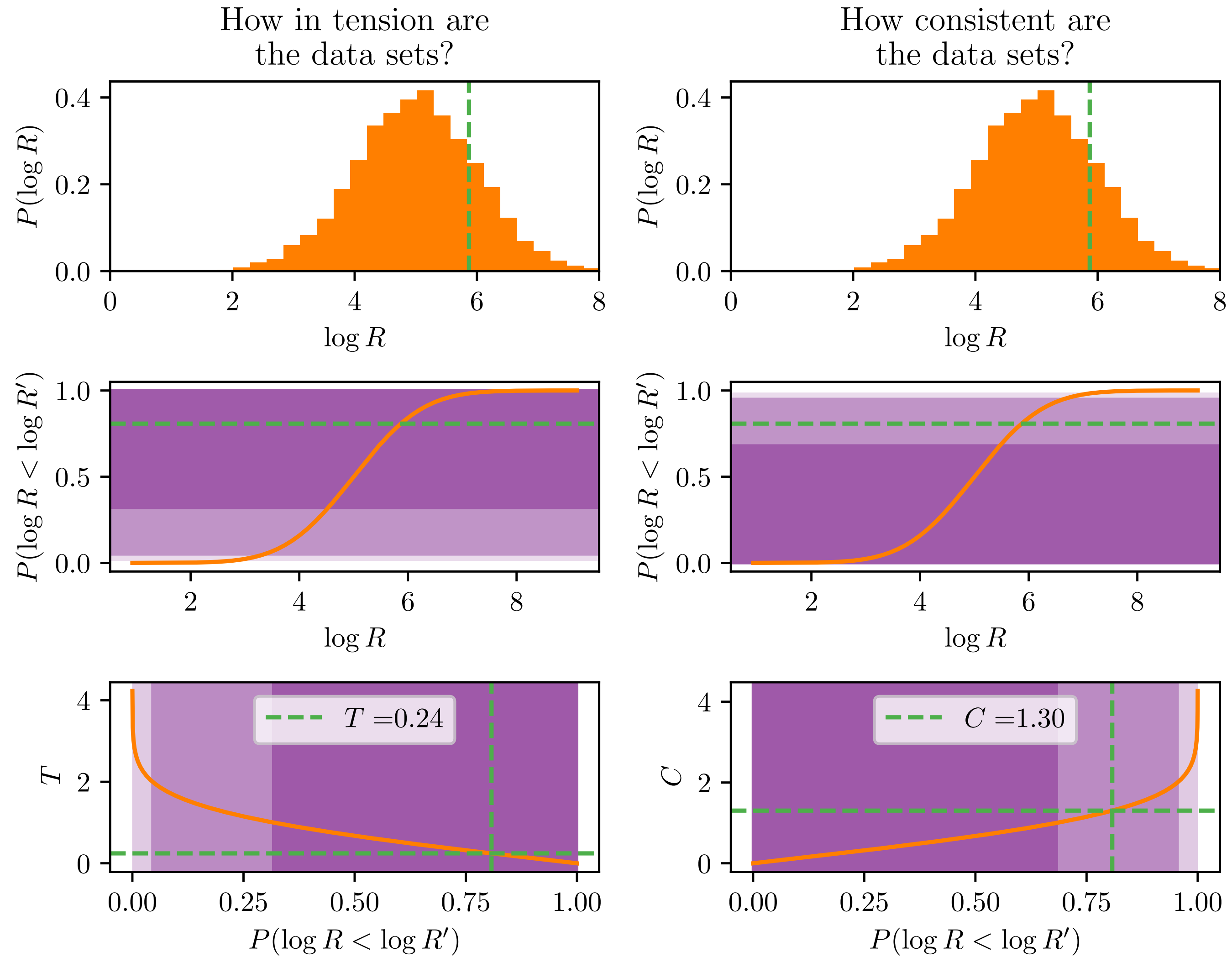
R with NREs



Direct predictions or calibration?



Calibration of R



Examples

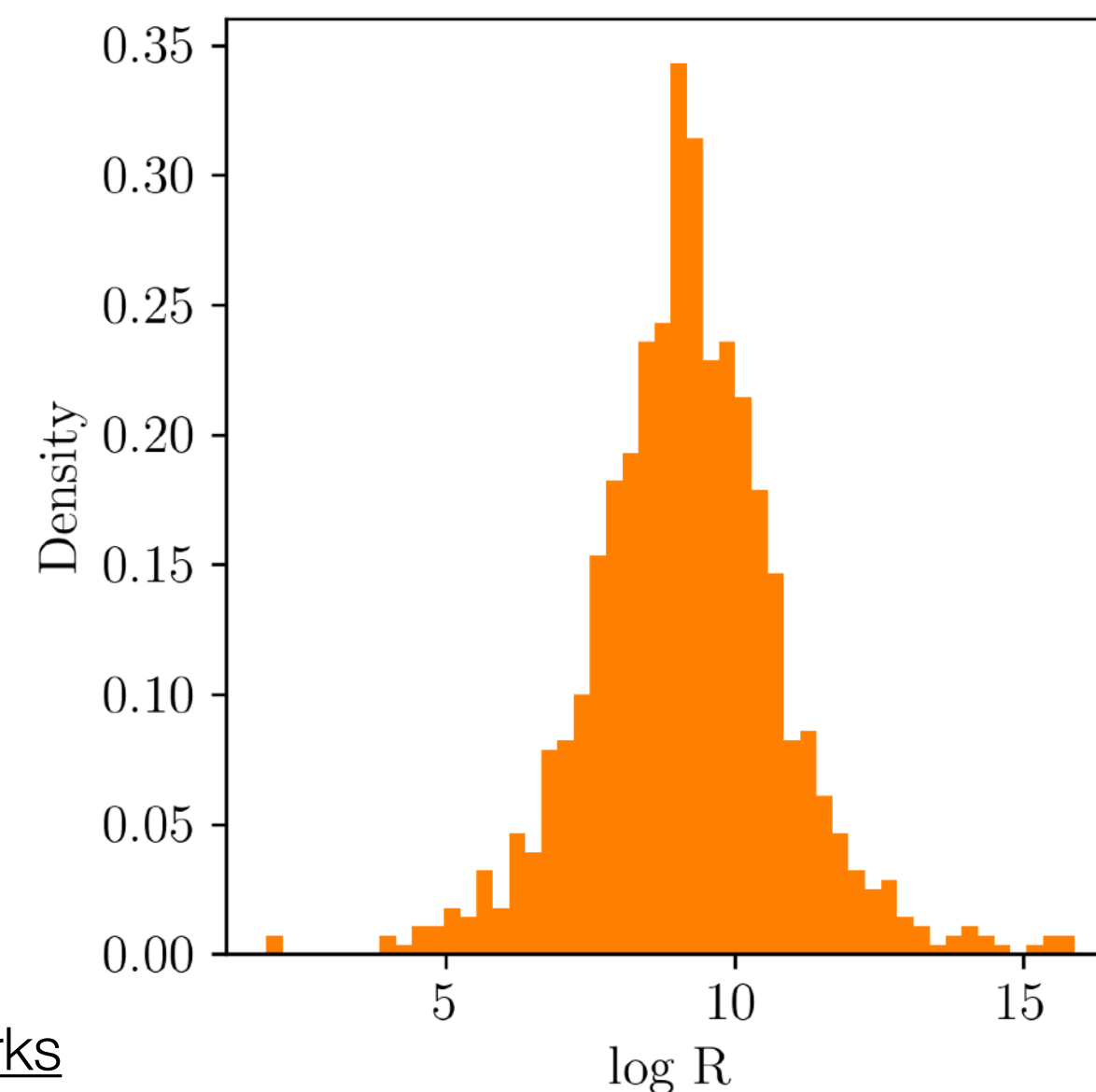
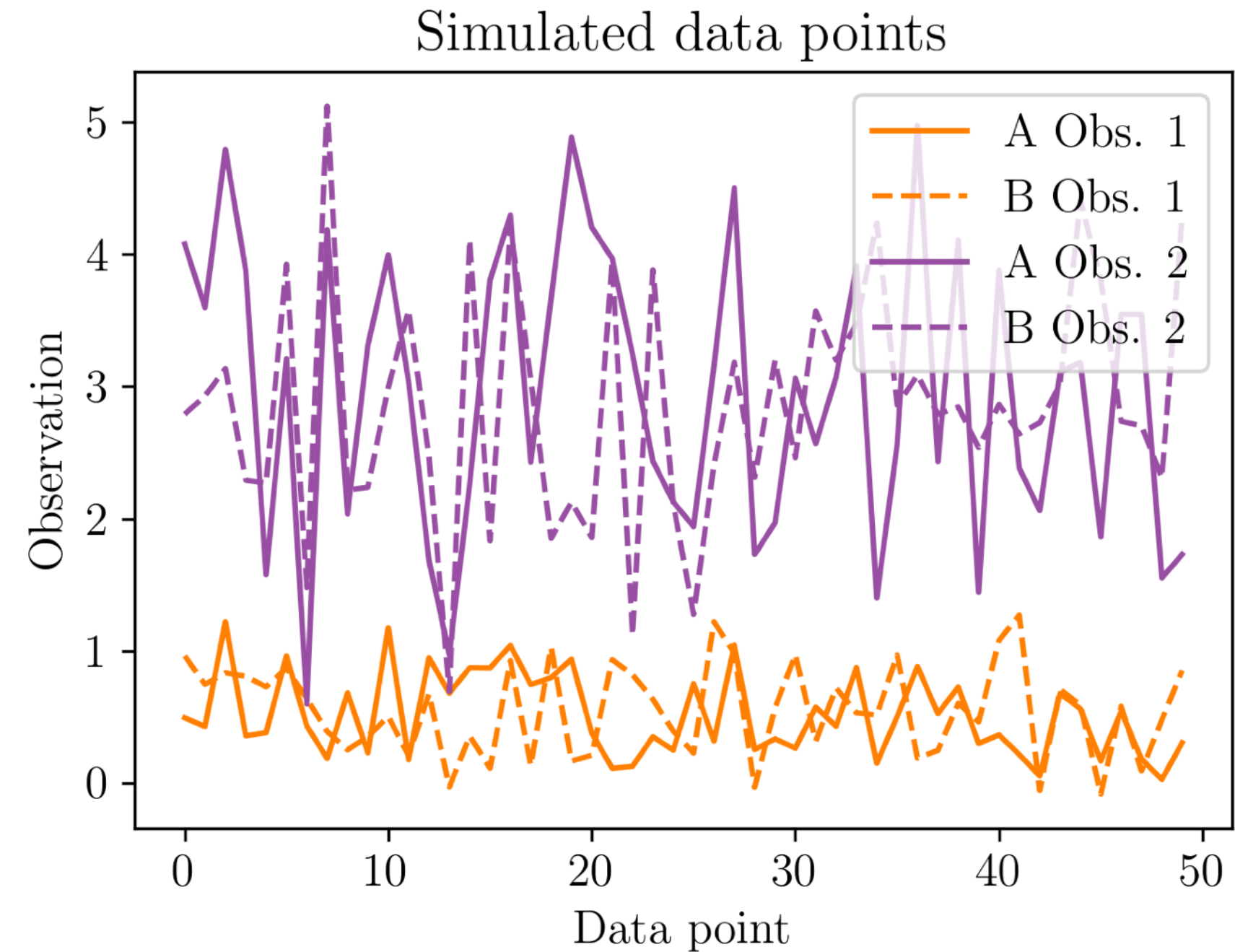
Analytic Example: Set Up

- Define a linear model

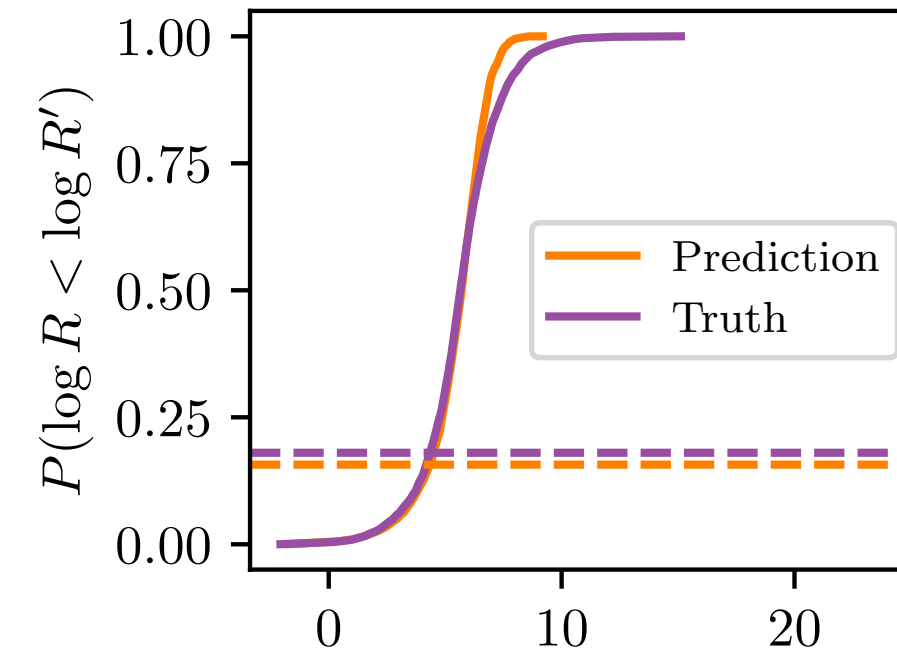
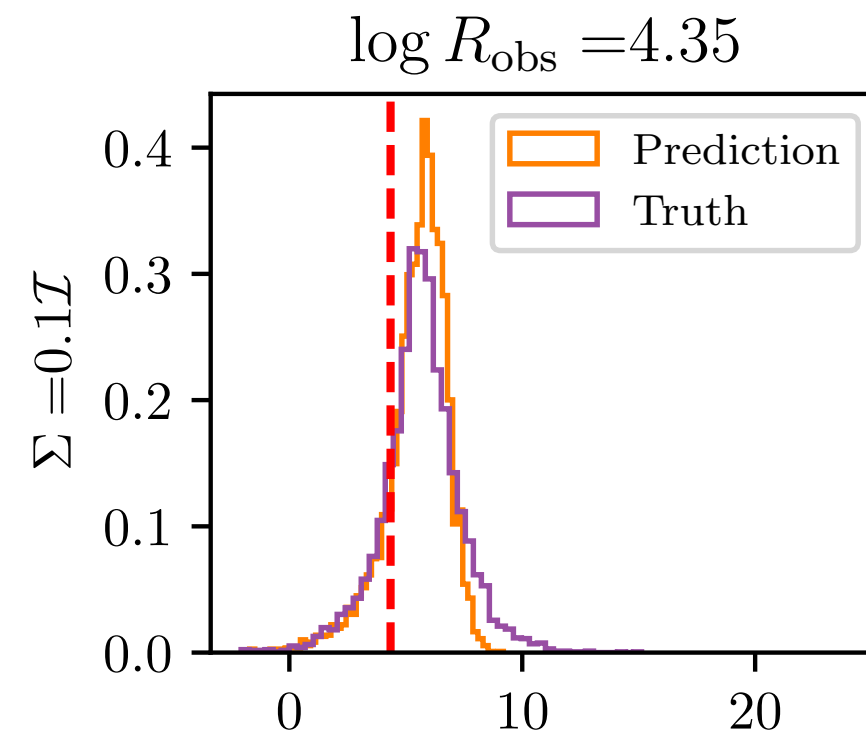
$$D_A = M_A \theta + m_A \pm \sqrt{C_A}$$

$$D_B = M_B \theta + m_B \pm \sqrt{C_B}$$

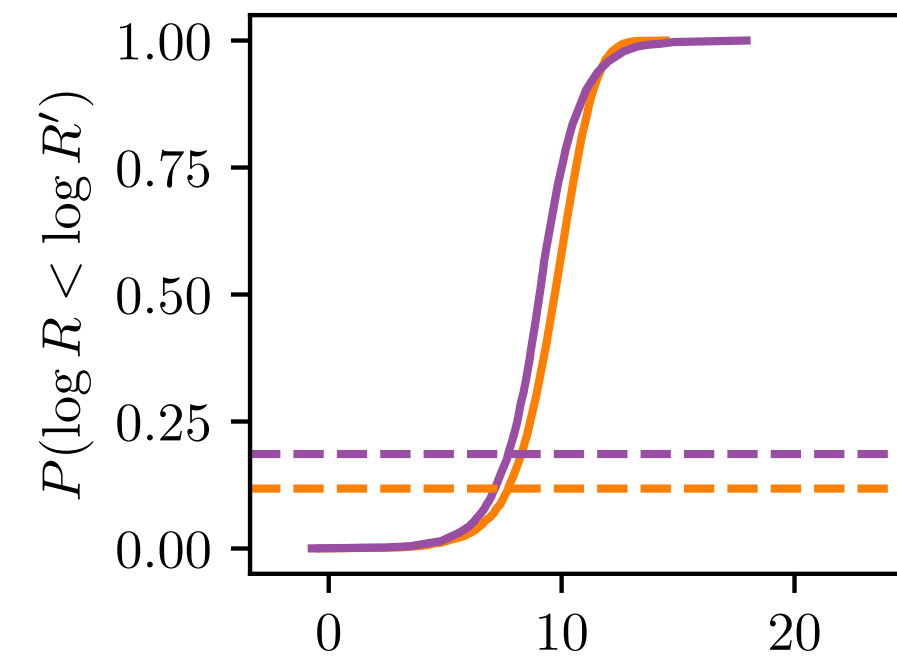
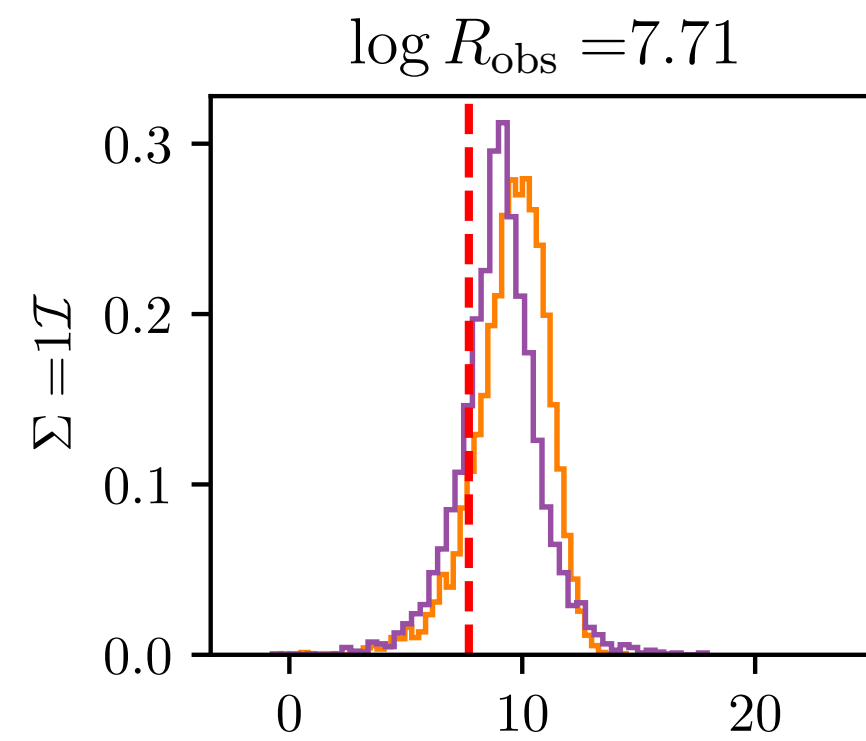
- $n_{dims} = 3, n_{data} = 50$
- Gaussian prior and likelihood
- Can analytically calculate $Z_A = P(D_A)$, Z_B and Z_{AB} and therefore get $\log R$
- Using lsbi package (<https://github.com/handley-lab/lsbi>)



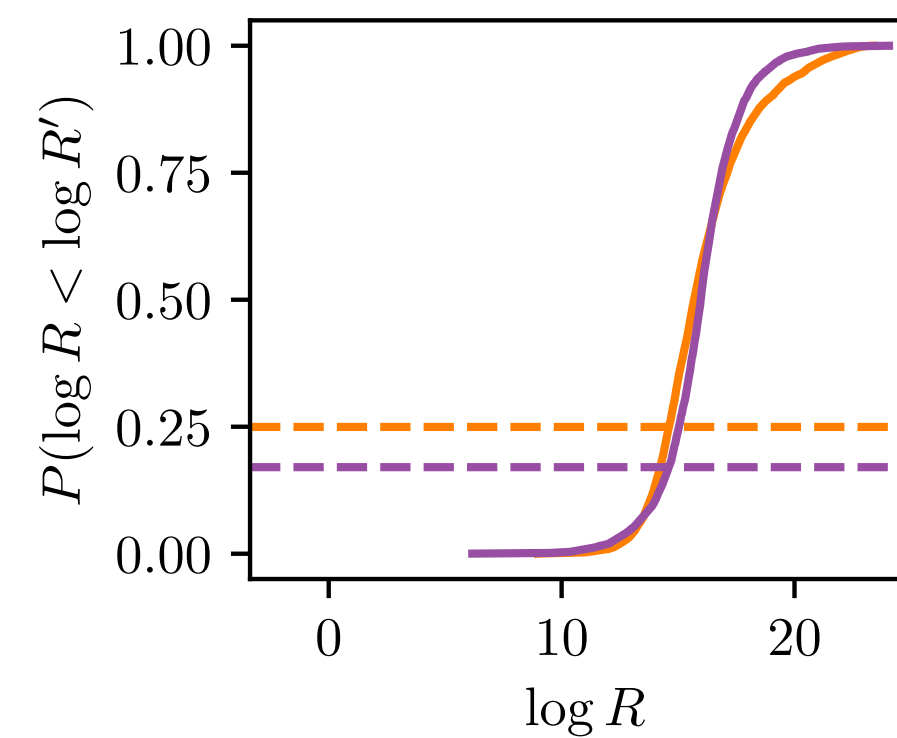
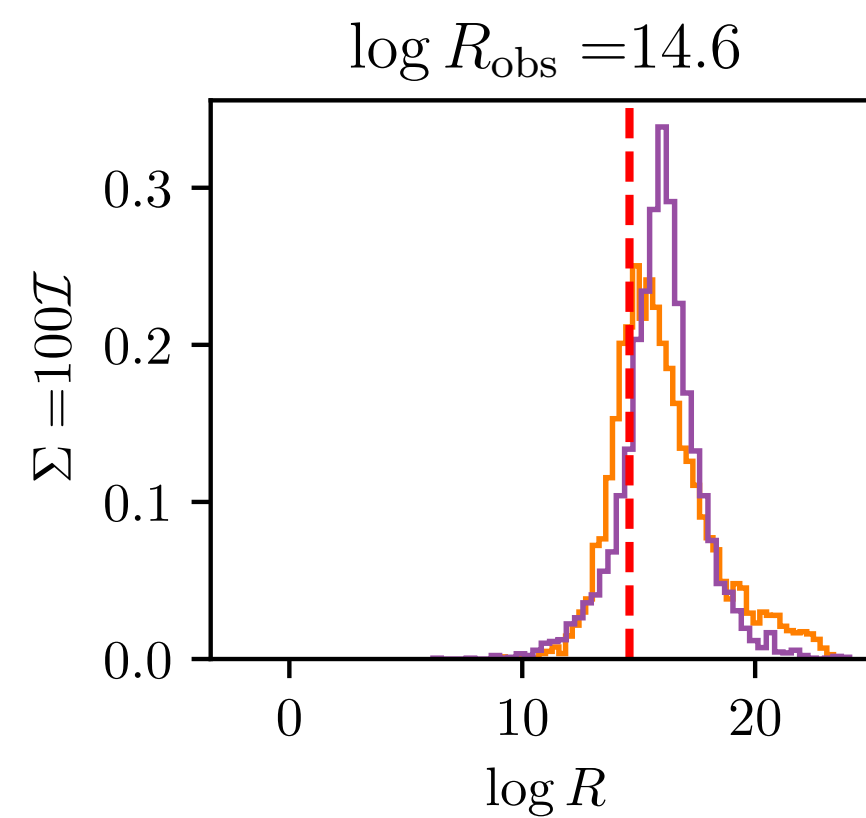
Analytic Example: Prior Dependence



	T	C
Truth	1.340	0.228
TENSIONNET	1.416	0.198



	T	C
Truth	1.323	0.235
TENSIONNET	1.563	0.148



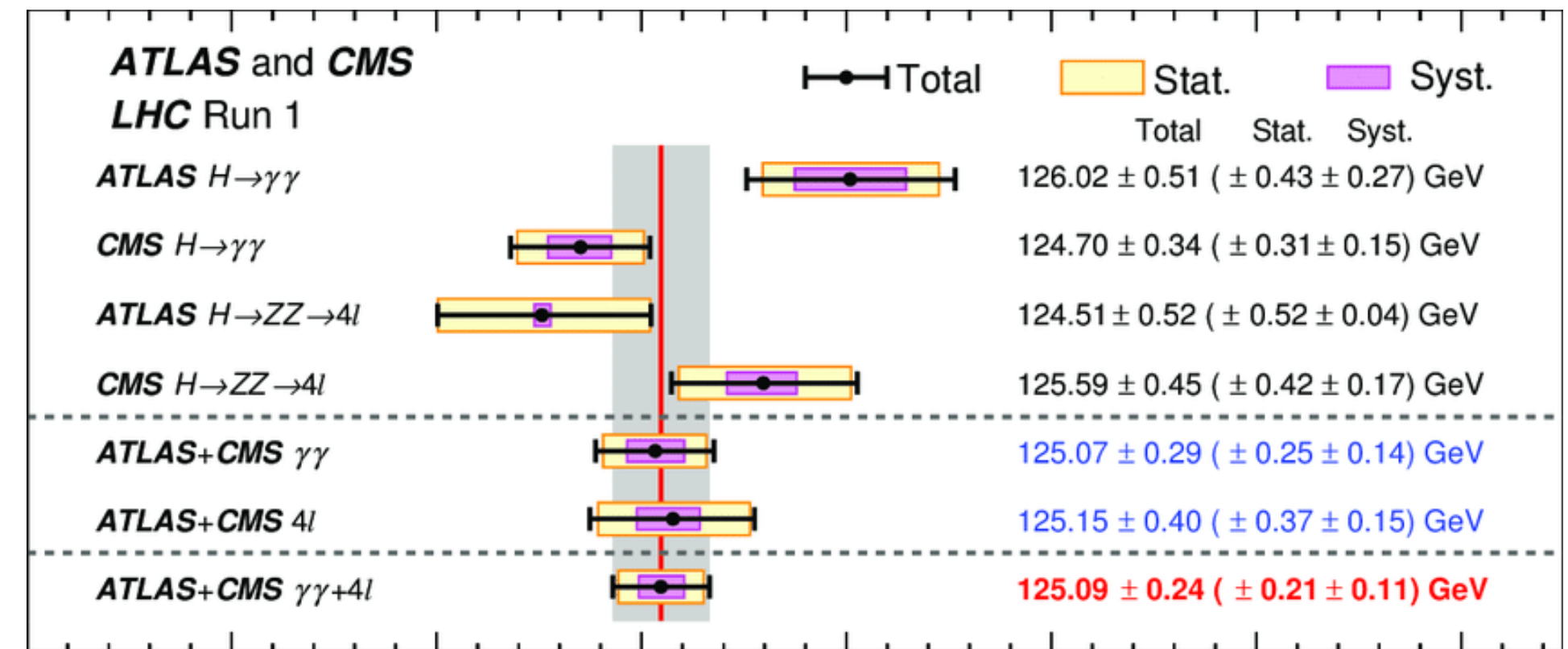
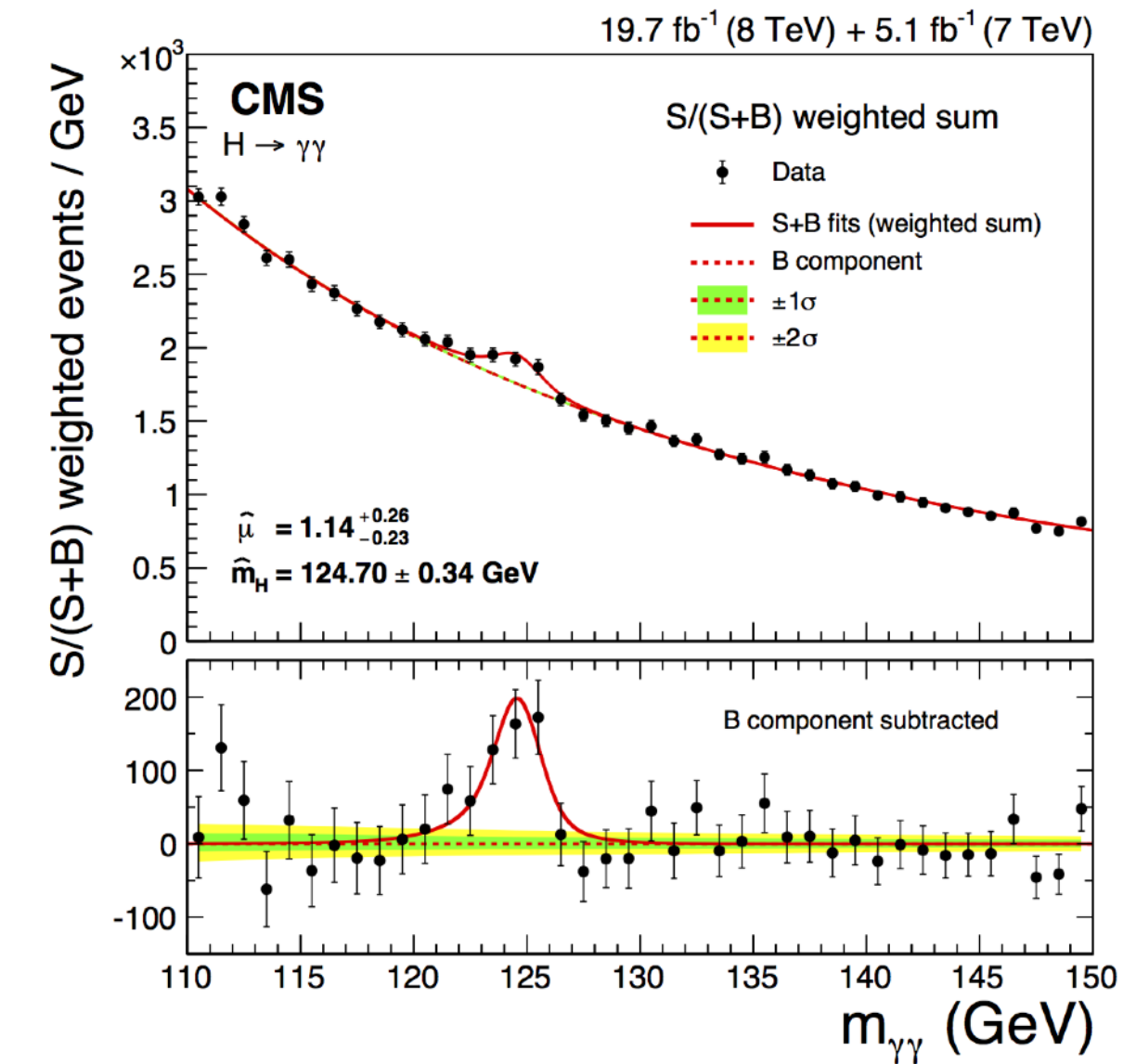
	T	C
Truth	1.371	0.215
TENSIONNET	1.151	0.318

<https://github.com/htjb/tension-networks>

<https://arxiv.org/abs/2407.15478>

Bump Hunting

- Imagine two experiments recording excess events via some channel at a similar mass
- Obvious that the experiments are observing the same signal
- Want to quantify how well they agree
- For example run 1 measurements of Higgs Boson mass at ATLAS and CMS

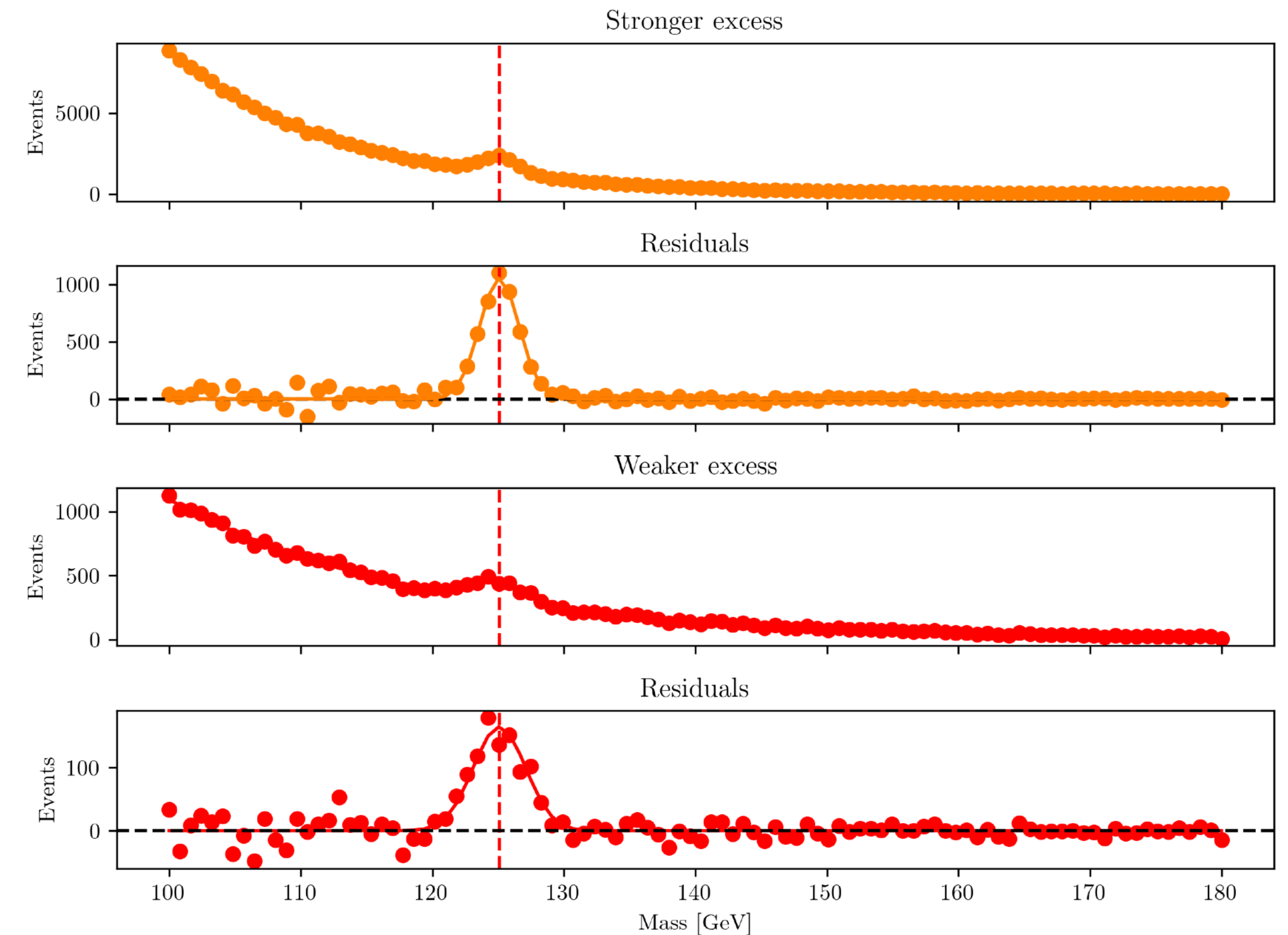


Bump Hunting

- Modelling background with

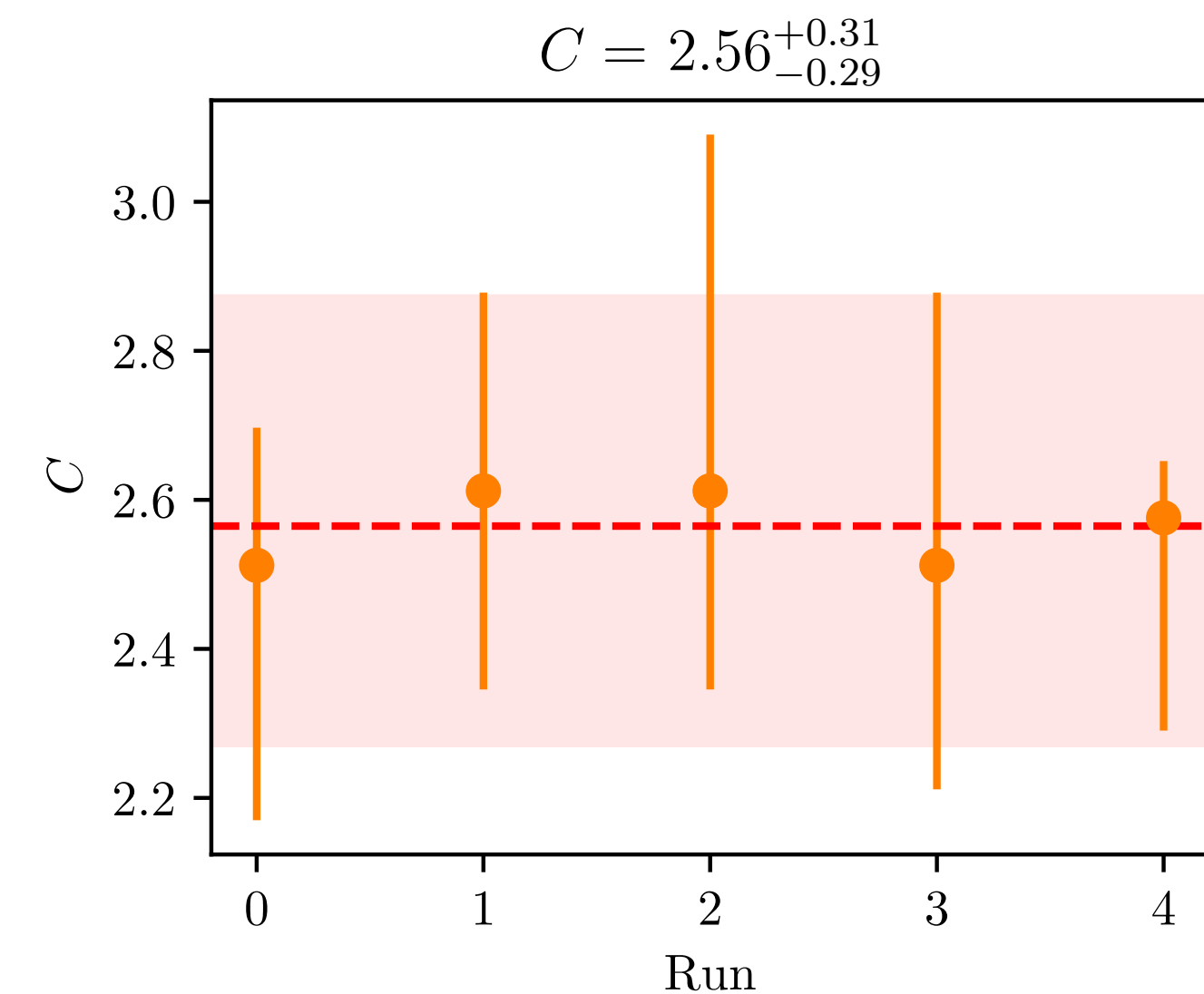
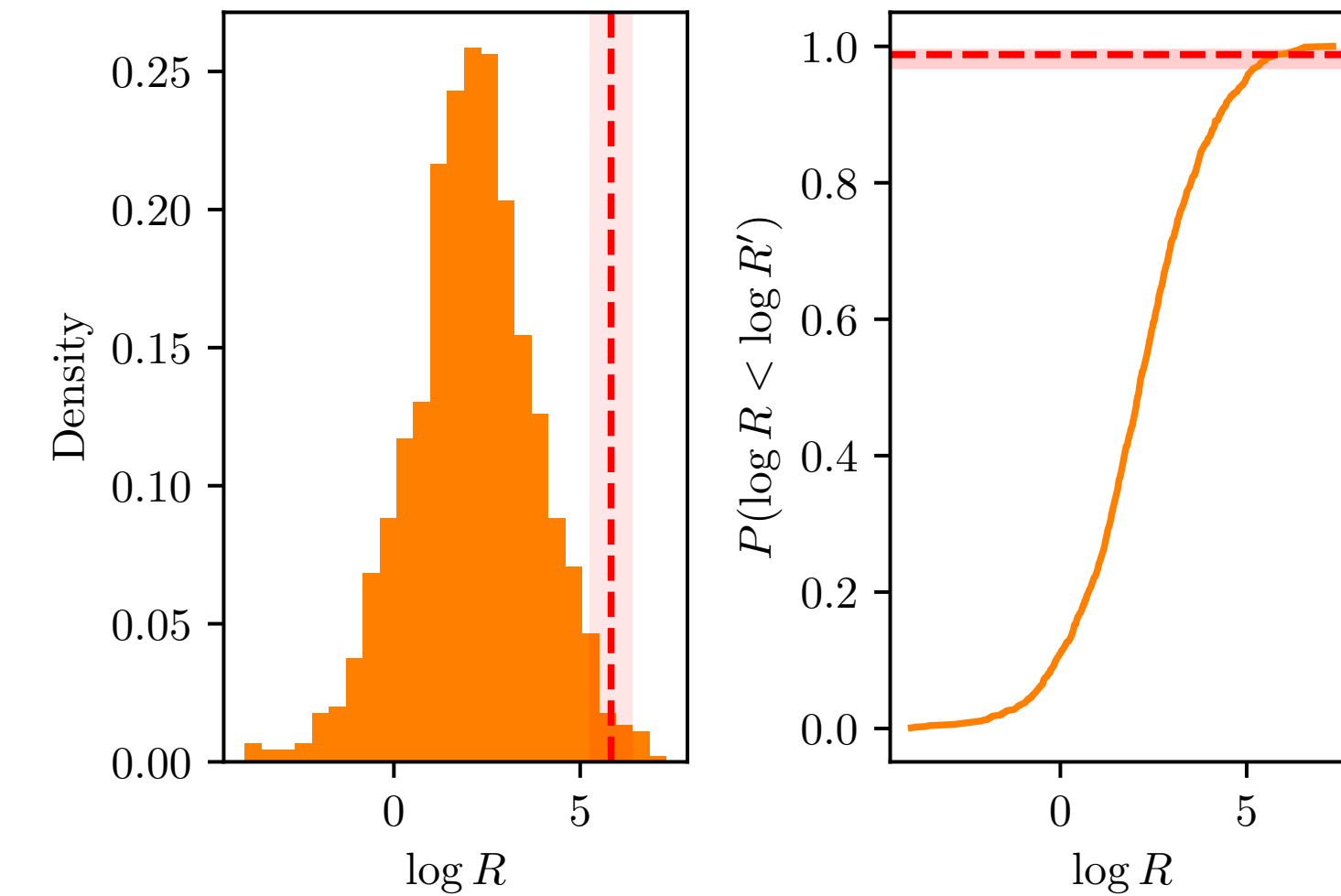
$$B(E) = \sum_{i=0}^{N=2} \theta_{1,i} \exp(-\theta_{2,i}E)$$

- And the excess as a gaussian centred around a mass of 125 GeV
- Imagining two collider experiments observing excesses
- One with more events hence less noise and a greater confidence



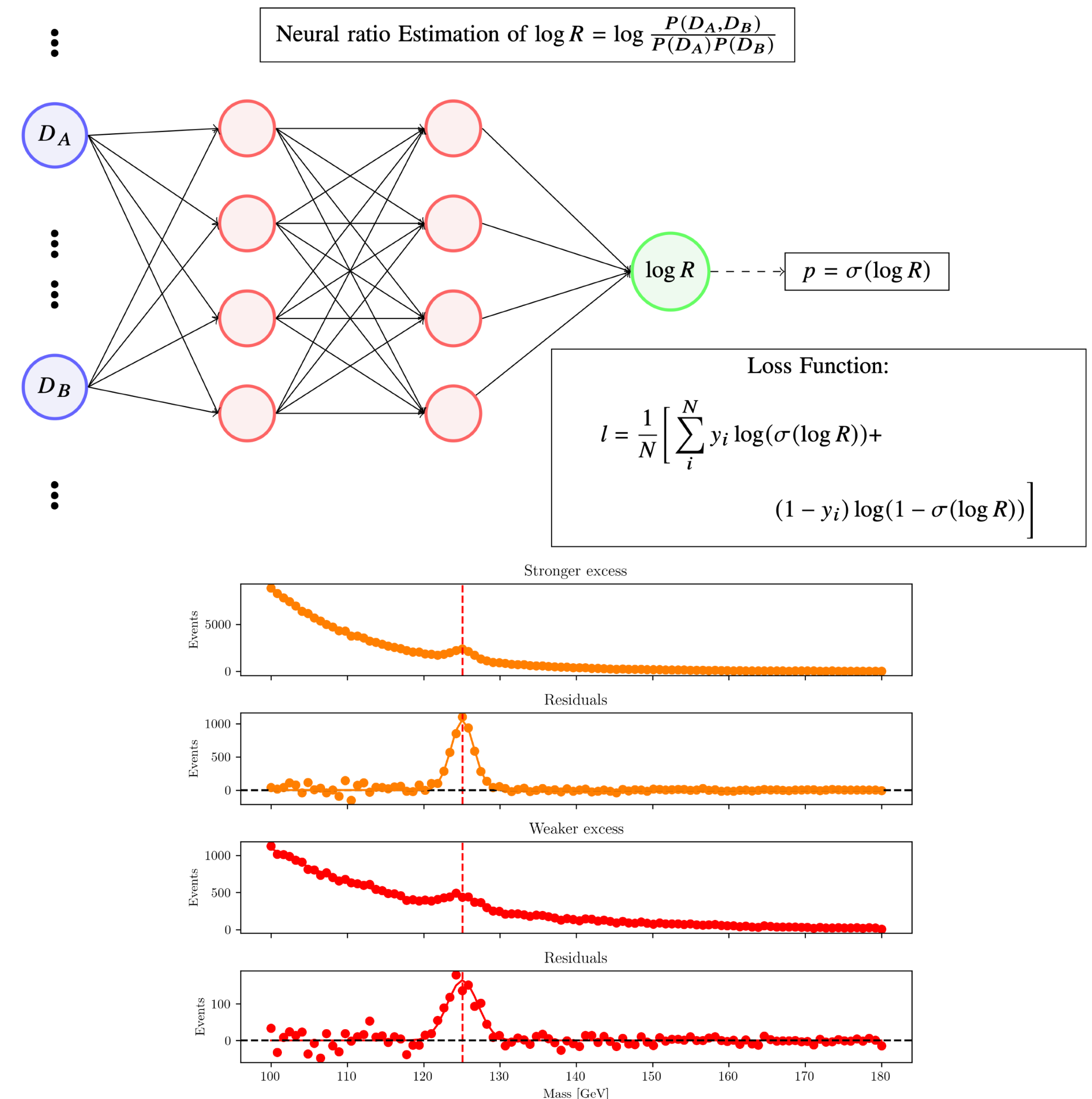
Bump Hunting

- Since I have an analytic model for the background and signal I can generate a range of simulations from a wide prior to train the NRE
- To calculate R_{obs} I use a product of Poisson distributions for my likelihood and the nested sampling algorithm
- Translating R_{obs} into units of σ concordance gives $C = 2.56^{+0.31}_{-0.29}$ averaged over five training runs



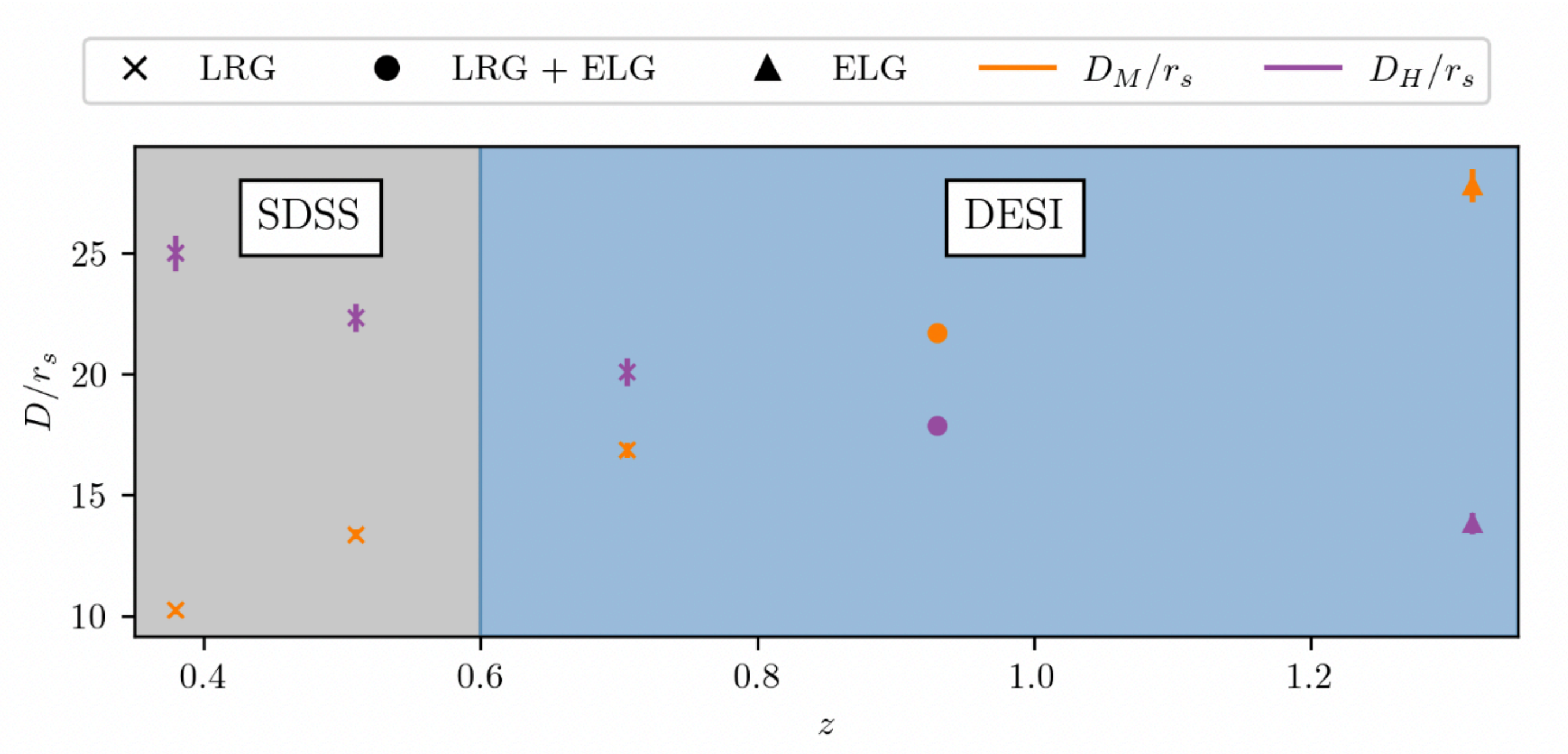
Conclusions

- Understanding tensions can help us identify new physics or instrumental systematics
- R statistic is an appropriately Bayesian choice
- We can use Neural Ratio Estimation to help us interpret the tension between different experiments
- Paper: arXiv:2407.15478
- Github: <https://github.com/htjb/tension-networks>

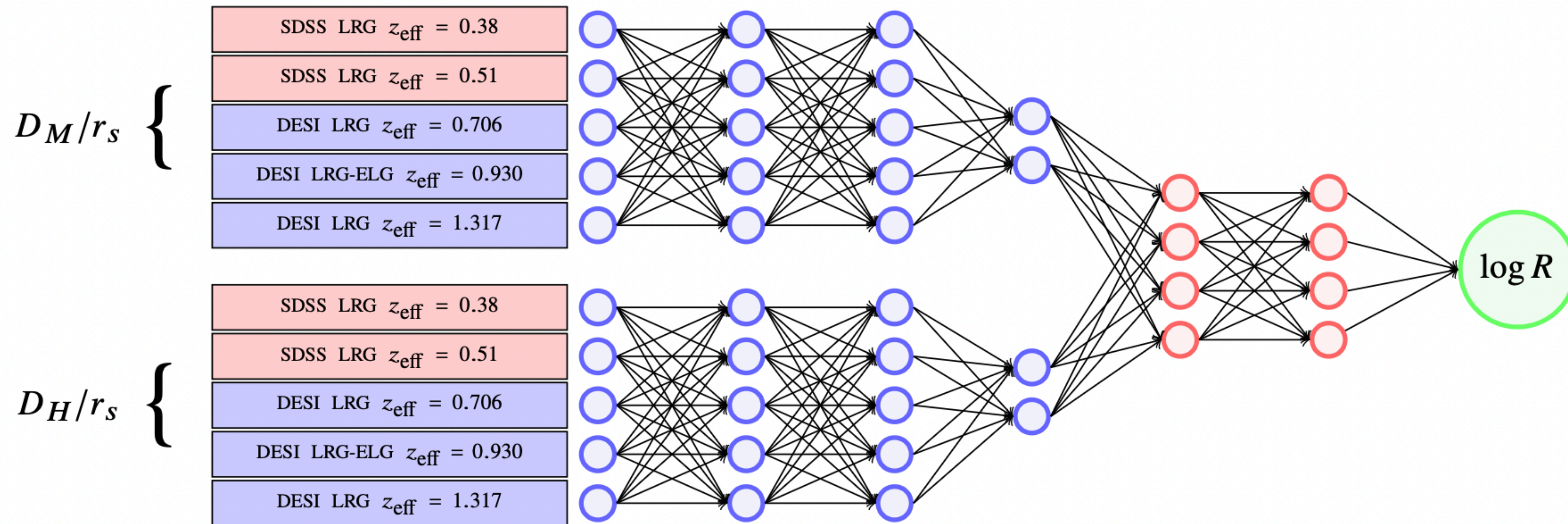


DESI + SDSS: Joint Data Set

- No existing correlated likelihood to evaluate a true R_{obs} with Nested Sampling
- Select different measurements from each survey to maximise the effective volume [e.g. 2404.03002]
- Focusing on LRG and ELG
- Add Quasars and $\text{Ly}\alpha$ in the future



DESI + SDSS: NRE Set Up



DESI + SDSS: Results

- We find $T = 1.22 \pm 0.20$

