Disappearing tracks at a 10 TeV Muon Collider

IMCC physics studies meeting 07/09/2023

Federico Meloni (DESY)



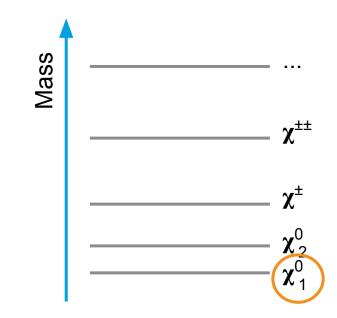




Co-funded by the European Union

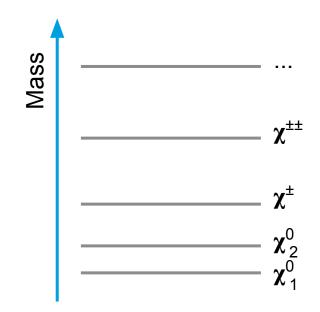
Start from the simplest interpretation of dark matter: it is the **thermal relic of at least a new stable neutral particle**.

• The SM is extended with *n*-tuplets that predict such neutral state χ^0



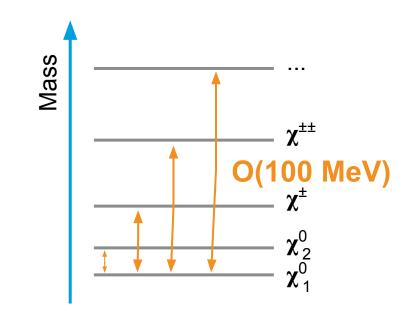
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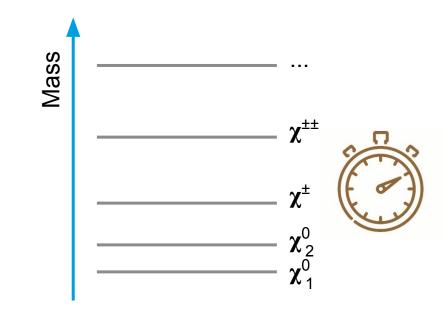
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- Heavier states can be long-lived

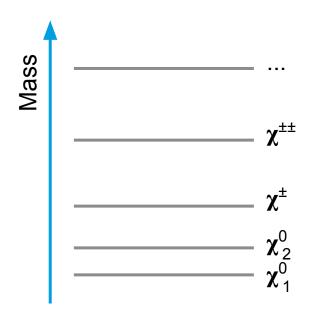


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Experimental signatures:

- Displaced vertices
- Kinked tracks
- Displaced tracks
- Disappearing tracks

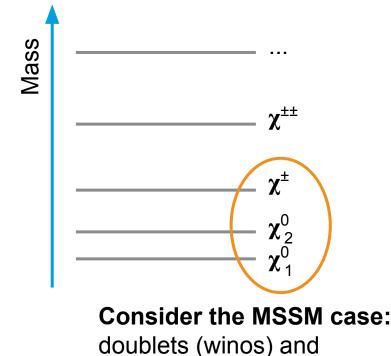


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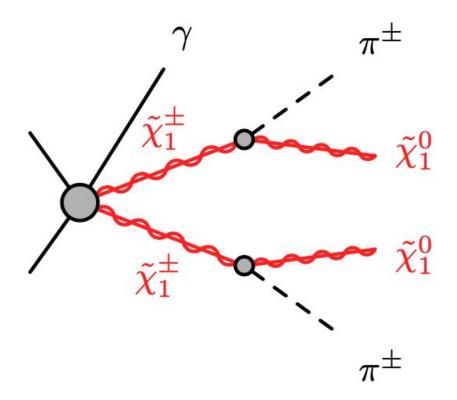
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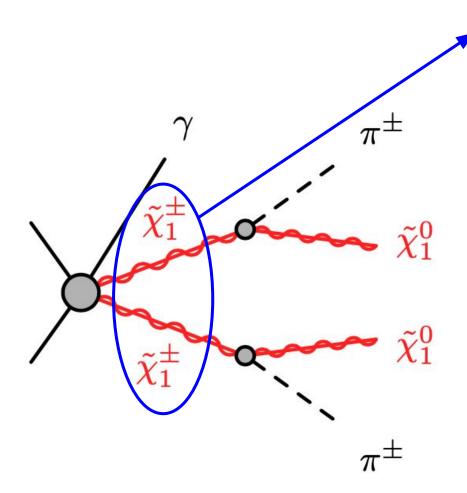


triplets (higgsinos)

An example from supersymmetry



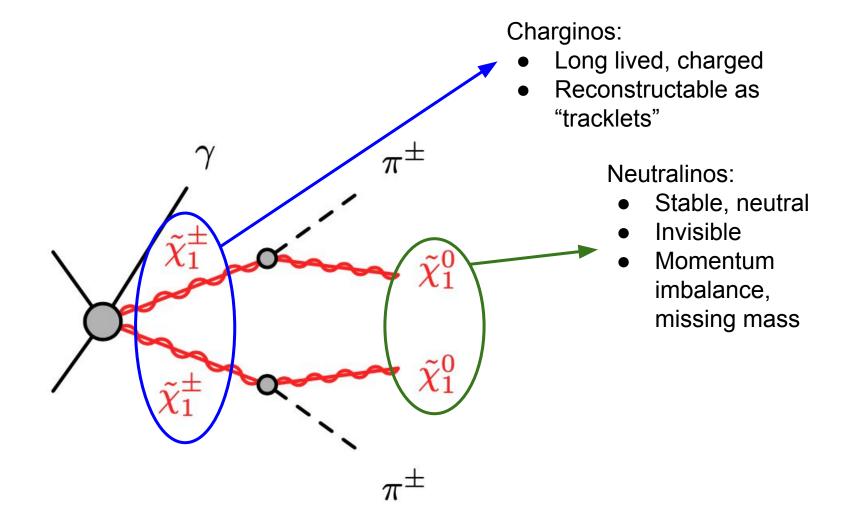
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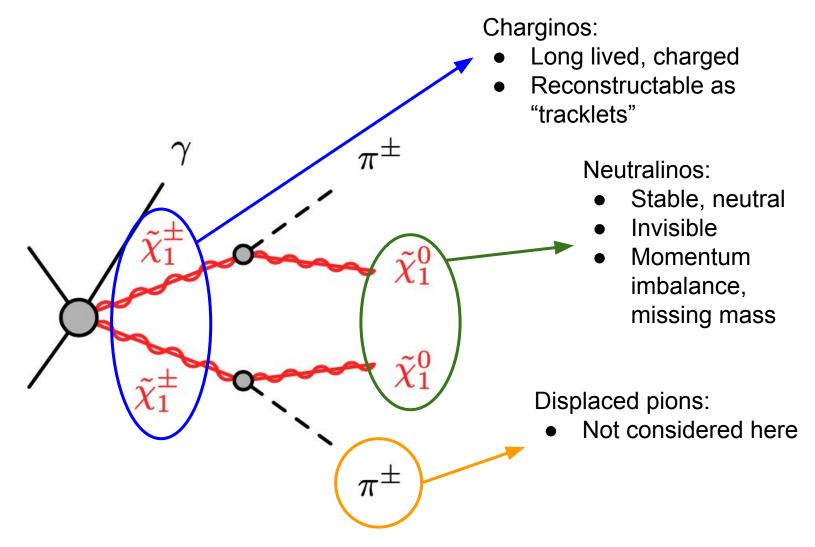
Charginos:

- Long lived, charged
- Reconstructable as "tracklets"

An example from supersymmetry



An example from supersymmetry



 π

 π^{\pm}

An example from supersymmetry

ISR/FSR:

- "Trigger" the event
- Momentum imbalance

Charginos:

- Long lived, charged
- Reconstructable as "tracklets"

Neutralinos:

- Stable, neutral
- Invisible
- Momentum imbalance, missing mass

Displaced pions:

• Not considered here

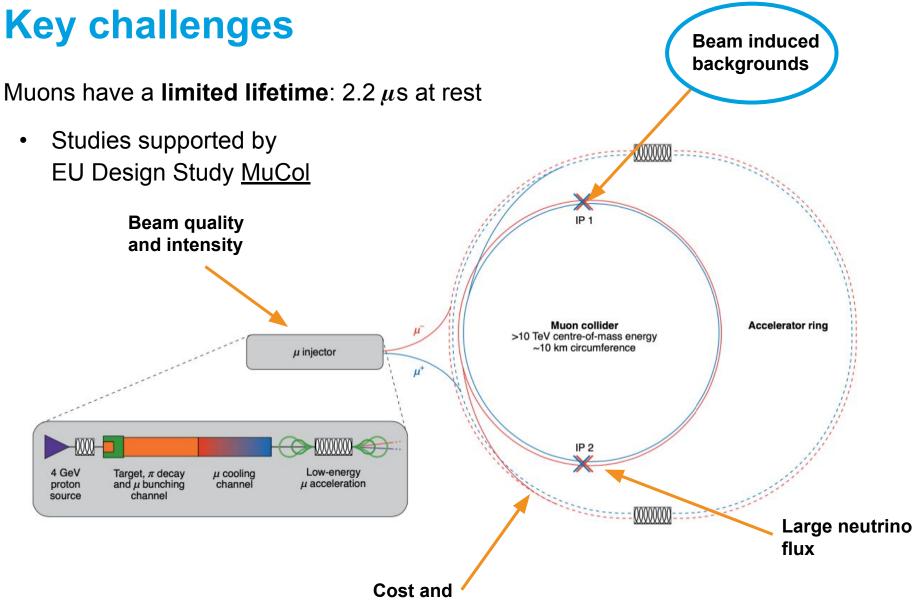
Where to look for these?

If dark matter is explained by a single particle, we expect this particle to be heavy

Higgsino ~ 1.1 TeV Wino ~ 2.7 TeV

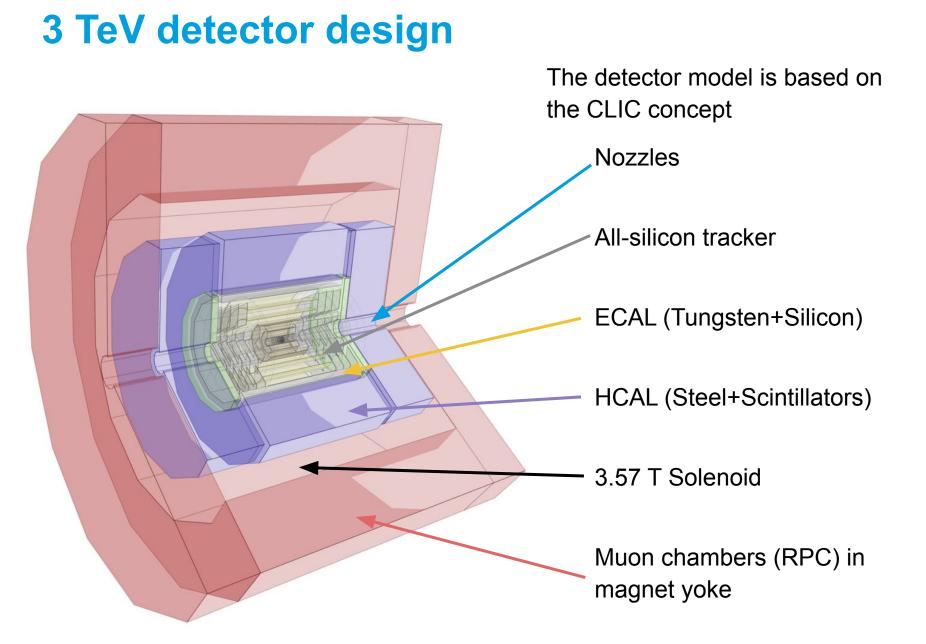
A muon collider is an excellent machine to look for these

Had a first look with full simulation in R. Capdevilla, FM, R. Simoniello, J, Zurita [2102.11292]



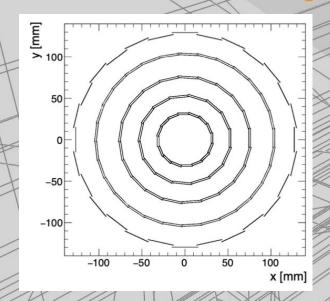
power

2303.08533



Signal event

No beam-induced backgrounds



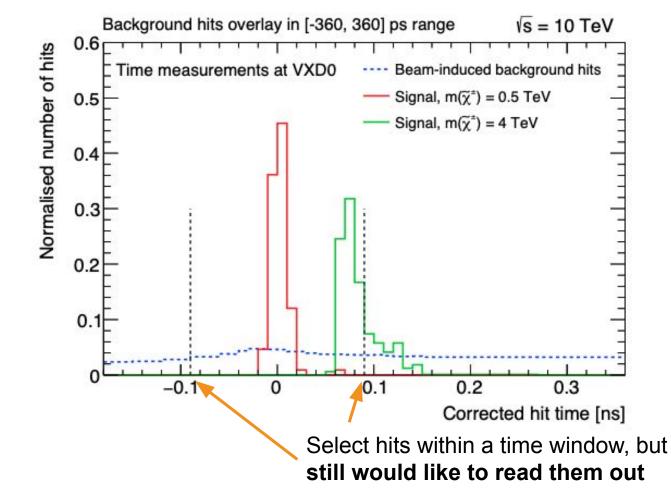
Signal event

Beam-induced backgrounds from MARS

BIB rejection: timing

Exploit particle arrival times to reduce BIB

• Correct for time of flight Corrected time = $t_{measured} - \frac{|r|}{c}$



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3 TeV detector

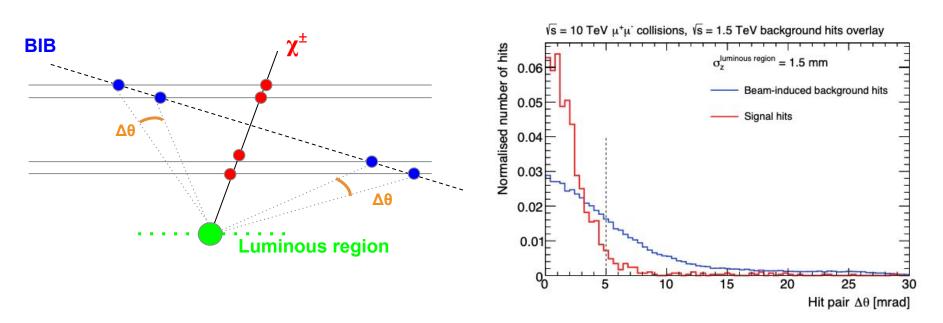
1.5 TeV BIB overlay Extrapolated to 10 TeV

BIB rejection: stubs

3 TeV detector 1.5 TeV BIB overlay Extrapolated to 10 TeV

The layout of the vertex detector can be exploited to reject hits from BIB particles

- Look for pairs of hits in neighbouring double-layers forming "stubs" that point back to the luminous region
- Work ongoing to apply a similar approach at the cluster level



Tracklet reconstruction

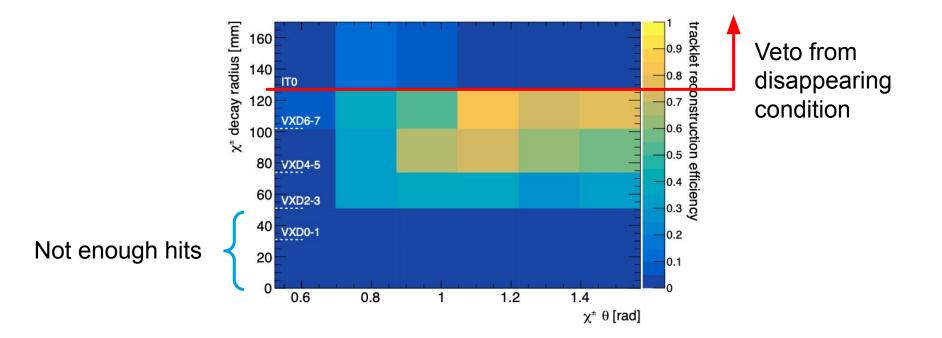
After BIB rejection cuts

3 TeV detector 1.5 TeV BIB overlay Extrapolated to 10 TeV

Impose a "disappearing condition" (hit veto) at the first layer of the IT (12.7 cm)

Efficiencies evaluated with truth matching to χ^{\pm}

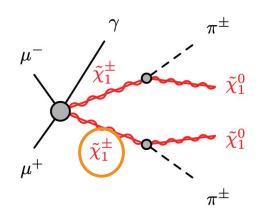
- Evaluated vs the χ^{\pm} decay radius and polar angle θ

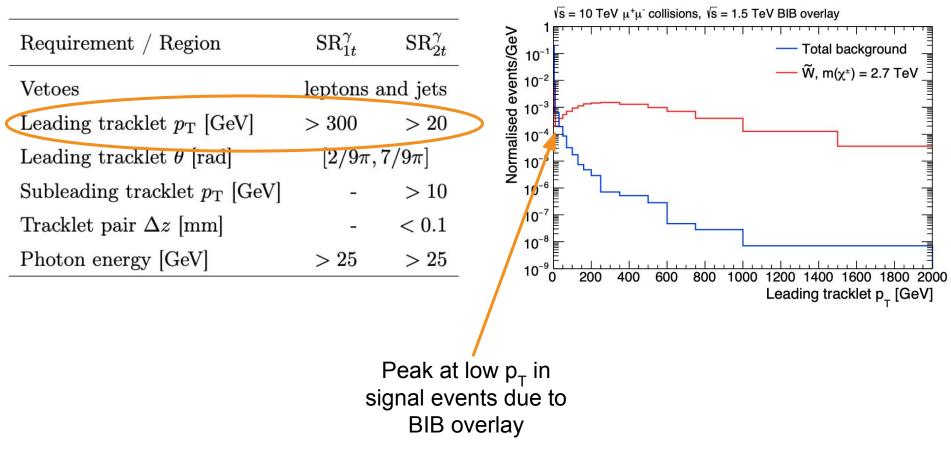


Event selection

Relatively simple event selection:

• Tracklet p_{T} (single most important quantity)

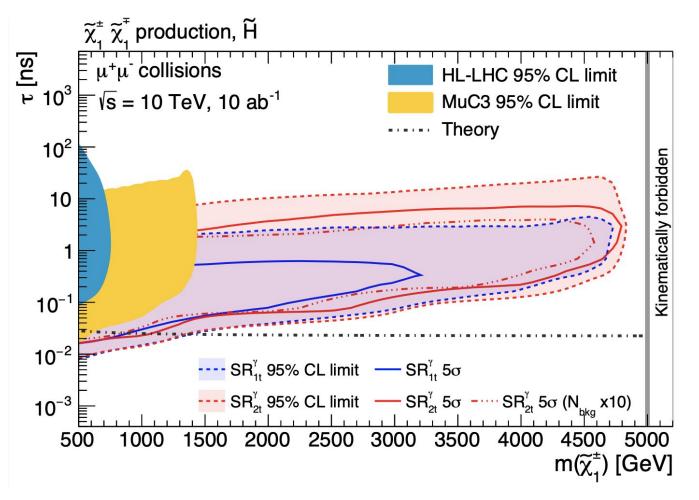




Expected sensitivity

3 TeV detector 1.5 TeV BIB overlay Extrapolated to 10 TeV

Pure higgsino models at MuC 10

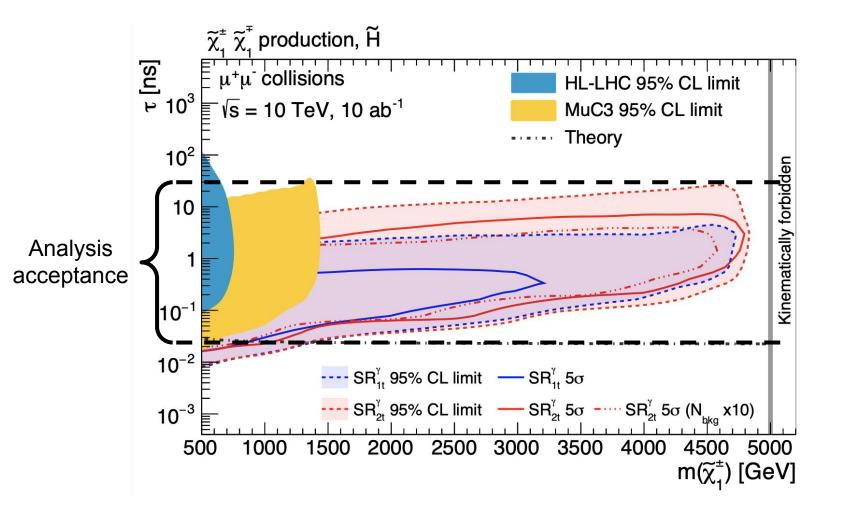


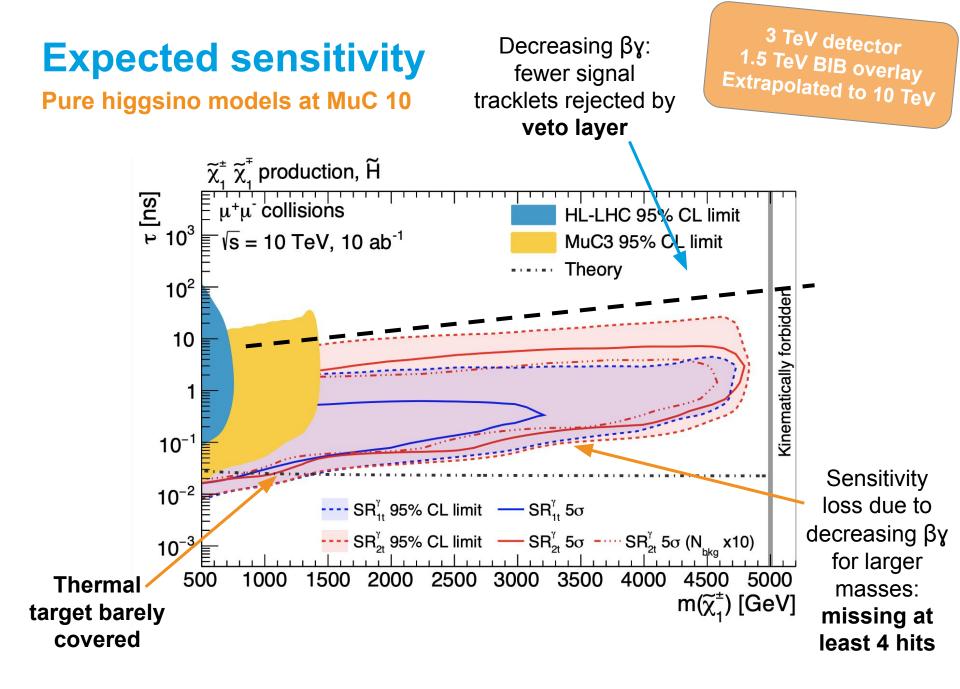
See also detailed comparison with fast sim results from Han, Liu, Wang, Wang [2009.11287, 2203.07351] in MuC Forum report [2209.01318]

Expected sensitivity

3 TeV detector 1.5 TeV BIB overlay Extrapolated to 10 TeV

Pure higgsino models at MuC 10





A new detector for 10 TeV

Detector model based on ongoing work D. Calzolari, K. DiPetrillo, R. Hillman, T. Holmes, S. Jindariani, B. Johnson, L. Lee, T. Madlener, FM, I. Ojalvo, P. Pani, S. Pagan Griso, K. Pedro, R. Powers, B. Rosser, L. Rozanov, A. Vendrasco, J. Zhang

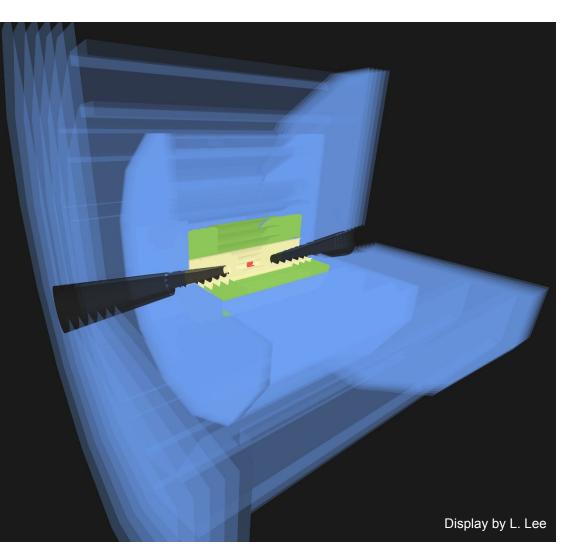
The detectors need to be ready to **measure both TeV-scale particles as well as GeV-scale**

Detector sizes need to grow

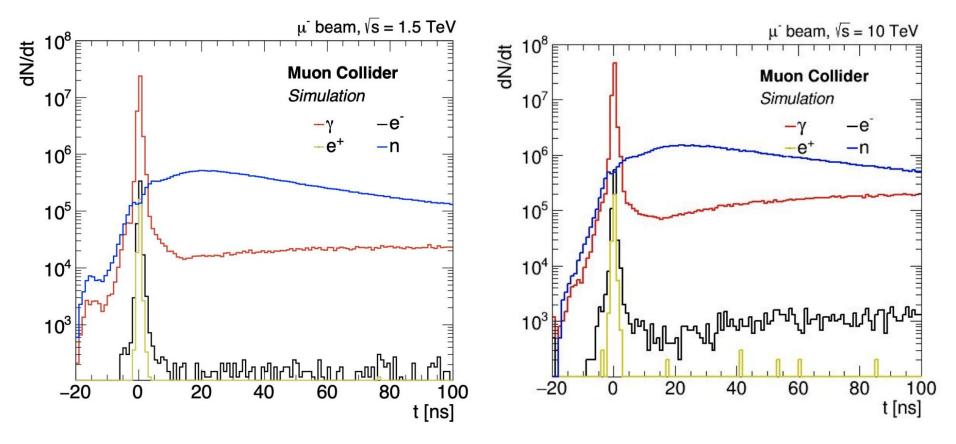
- Thicker calorimeters
- Bigger trackers with high
 precision in more places

Physics benchmarks are key to guide the detector a design

Can a revised detector preserve, or even improve, the sensitivity to winos and higgsinos?



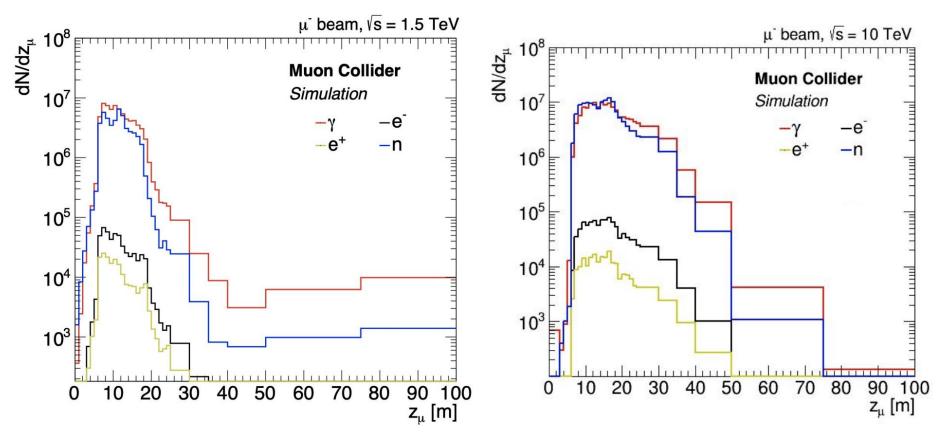
Moving to 10 TeV BIB simulation: time



Simulation based on FLUKA (thanks to Daniele Calzolari and the MDI group!)

- 5 TeV muon and anti-muon beams
- Sampled 110 metres of decays in the collider lattice

Moving to 10 TeV BIB simulation: μ decay z

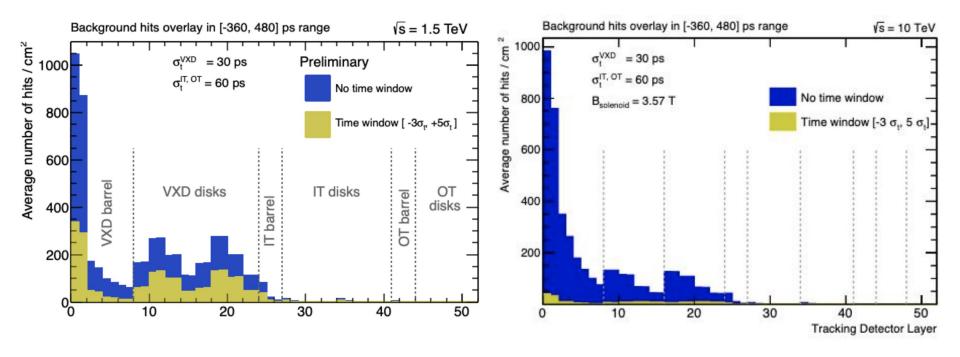


Lack of very-long tail comes from better modelling in FLUKA (now including synchrotron radiation)

• See more in Daniele's <u>talk</u>

See also <u>Nazar Bartosik's</u> <u>10 TeV BIB study</u>

Occupancies: tracker



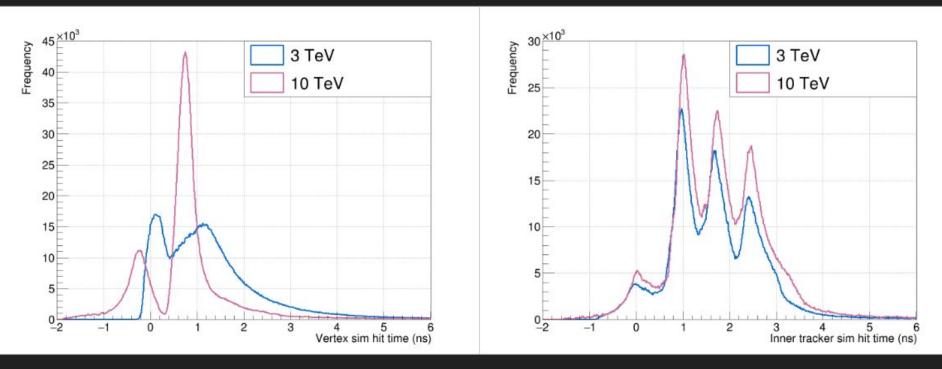
Increasing solenoid field to 5T has an effect of the order of 10% pre-timing Work in progress on:

- Different BIB structure in endcaps
- Timing selections seem much more effective than at 3 TeV

Effectiveness of timing selections

Vertex Barrel Sim Hit Timing

Inner Tracker Barrel



- Disparity in VXB hit timing affects the efficiency of timing cut

Different timing structure being investigated

- No obvious bugs found so far, looking at dependency on machine lattice
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Table: courtesy of Sergo Jindariani

Data rates and power

	Upper timing cut (ns)	Module size (cm²)	Maximum hits/cm²	Reduction using cluster shapes	Data payload per module (Gbps)	Transmission power per module (W)	Total Transmission Power (W)
VXD barrel L1/L2	15	10	3000	x2	50	0.5	25
VXD barrel L1/L2	1	10	2400	x2	35	0.3	20
VXD barrel L3-8	15	10	1400	-	50	0.5	100
VXD disks	15	10	700	-	21	0.2	60
IT barrel	15	50	170	-	24	0.2	360
IT disks	15	50	60	-	9	0.09	100
OT barrel	15	100	10	-	3	0.03	170
OT disks	15	100	5	-	1.5	0.15	30

L1/L2 rates still require tighter time window (~1ns)

L3/L8 now less comfortable

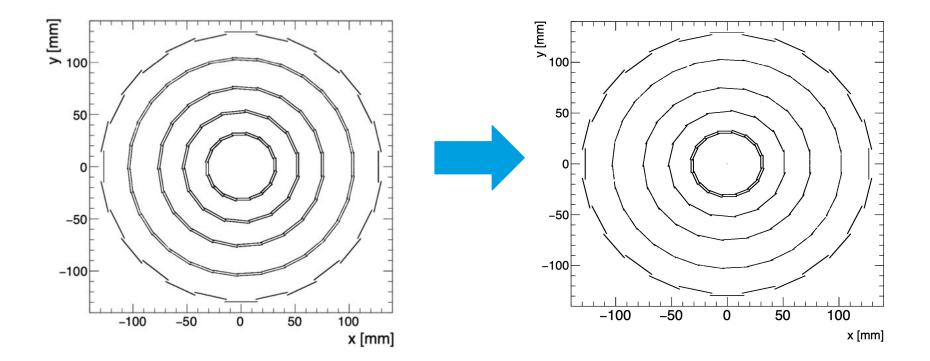
• Total L3/L8 power could be halved by removing the double layers

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Updating to a new tracker

These considerations, together with a greatly improved tracking software (now based on the ACTS library) made the double layers possibly redundant

• Barrel region of vertex detector revised keeping only one double layer pair



Updated detection efficiency

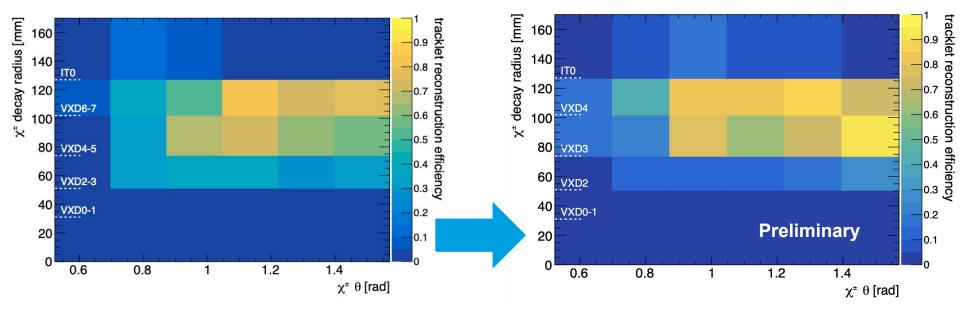
10 TeV detector Preliminary 10 TeV BIB overlay

New tracker layout and tracking algorithm

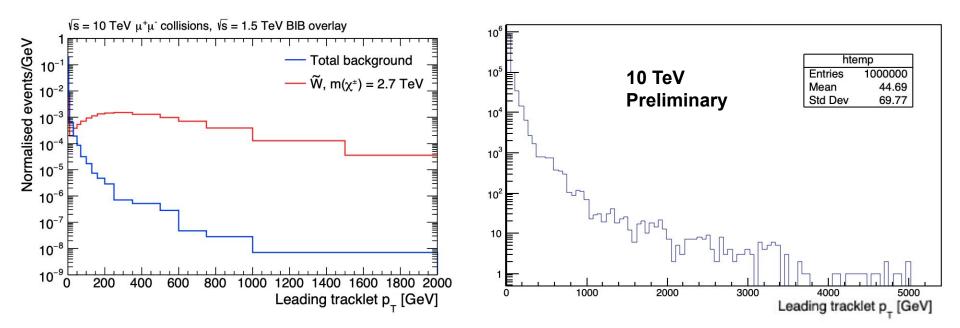
Kept all tracklet quality requirements as before except:

 $N_{hit} \ge 3$ hits No requirement on stubs

Found similar detection efficiency, and greatly reduced fake tracklet rate



Updated fake rates



Reconstructed fake tracklet p_{τ} spectrum similar

Probability of finding a fake tracklet per event lowered by ~1/10

• Estimates being refined with higher MC statistics

Expected sensitivity

Pure higgsino models at MuC 10

DESY.

 $\widetilde{\chi}_{1}^{\pm} \widetilde{\chi}_{1}^{\mp}$ production, \widetilde{H} [<u>su</u>] ບ 10³ $\begin{bmatrix} \mu^{+}\mu^{-} \text{ collisions} \end{bmatrix}$ HL-LHC 95% CL limit √s = 10 TeV, 10 ab⁻¹ ·-· Theory **Preliminary** 10² Kinematically forbidden 10 10^{-1} 10⁻² ---- SR_{1t}^{γ} 95% CL limit ---- SR_{1t}^{γ} 5 σ ----- SR_{1t}^{γ} 5 σ (N_{bkg} x10) ---- SR_{2t}^{γ} 95% CL limit ---- SR_{2t}^{γ} 5 σ ----- SR_{2t}^{γ} 5 σ (N_{bkg} x10) 10⁻⁵ Thermal target 500 1500 2000 2500 3000 3500 4000 1000 4500 5000 m(ĩ̃₁⁺) [GeV] covered by both selections

10 TeV detector Preliminary 10 TeV BIB overlay



Physics benchmarks are key in evaluating new detector design options

Dark matter is one of the most important topics of research in particle physics, and **disappearing tracks are the ultimate handle** to target minimal dark matter models

The preliminary update to a 10 TeV detector model and simulation seems to indicate the **initial estimates were conservative**

• Thermal targets possibly well within reach

Physics reach of a multi-TeV μ C relies on (among other things) **successful** detector design programme today

Thank you!

Contact

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Many thanks to S. Jindariani, D. Schulte, and M. Wing for inputs and useful discussions

The 12 challenges of a MuC

	Target	Status	Notes	Future work Refine design, including proton acceleration. Accumulation and compression of bunches.			
Pulse compression	1-3 ns	SPS does O(1) ns	Need higher intensity. O(30) ns loses only factor 2 in the produced muons.				
High-power targets	2 MW	2 MW	Available for neutrino and spallation neutrons. Aim for 4 MW to have margin.	Develop target design for 2 MW, O(1) ns bunches create larger thermal shocks. Prototype in 2030s.			
Capture solenoids	15 T	13 T	ITER central solenoid.	Study superconducting cables and validate cooling. Investigate HTS cables.			
Cooling solenoids	50 T	30-40 T	30 T leads to a factor 2 worse transverse emittance with respect to design.	Extend designs to the specs of the 6D cooling channel. Demonstrator.			
RF in magnetic field	>50 MV/m	65 MV/m	MUCOOL published results. Requires test in non-uniform B.	Design to the specs of 6D cooling. Demonstrator.			
6D cooling	10 ⁻⁶	0.9 (1 cell)	MICE result (no re-acceleration). Emittance exchange demonstrated at g-2.	Optimise with higher fields and gradients. Demonstrator.			
RCS dynamics	-		Simulation. 3 TeV lattice design in place.	Develop lattice design for a 10 TeV accelerator ring.			
Rapid cycling magnets	2 T/ms 2.5 T/ms Normal conducting magnets. 2 T peak 1.81 T peak HTS demonstrated 12 T/ms, 0.24 T peak.			Design and demonstration work. Optimise power management and re-use.			
Ring magnets aperture	20 T quads	12-15 T (Nb3Sn)	Need HTS or revise design to lower fields.	Design and develop larger aperture magnets, 12-16 T dipoles and 20 T HTS quads.			
Collider dynamics	-		3 TeV lattice in place with existing technology.	Develop lattice design for a 10 TeV collider.			
Neutrino radiation	10 μSv/year	123	3 TeV ok with 200 m deep tunnel. 10 TeV requires a mover system.	Study mechanical feasibility of the mover system impact on the accelerator and the beams.			
Detector shielding	Negligible	LHC-level	Simulation based on next-gen detectors.	Optimise detector concepts. Technology R&D.			

Muon collider target parameters

Parameter	Symbol	Unit	Target value		CLIC		
Centre-of-mass energy	$E_{\rm cm}$	TeV	3	10	14	3	1
Luminosity	L	$10^{34} {\rm cm}^{-2} {\rm s}^{-1}$	1.8	20	40	5.9	
Luminosity above $0.99 \times \sqrt{s}$	$\mathcal{L}_{0.01}$	$10^{34} {\rm cm}^{-2} {\rm s}^{-1}$	1.8	20	40	2 🔸	Ļ '
Collider circumference	C_{coll}	km	4.5	10	14		Beamstrahlung
Muons/bunch	N	10^{12}	2.2	1.8	1.8	0.0037	1
Repetition rate	f_r	Hz	5	5	5	50	
Beam power	$P_{\rm coll}$	MW	5.3	14.4	20	28	
Longitudinal emittance	ϵ_L	MeVm	7.5	7.5	7.5	0.2	
Transverse emittance	ϵ	$\mu { m m}$	25	25	25	660/20	
Number of bunches	n_b		1	1	1	312	
Number of IPs	$n_{ m IP}$		2	2	2	1	1
IP relative energy spread	δ_E	%	0.1	0.1	0.1	0.35	
IP bunch length	σ_z	mm	5	1.5	1.07	0.044	
IP beta-function	β	mm	5	1.5	1.07	1	
IP beam size	σ	$\mu { m m}$	3	0.9	0.63	0.04/0.001	

Based on extrapolation of the MAP parameters

 Plan to operate 5 years at each centre-of-mass energy (FCC-hh to operate for 25 years)