## Overview of HNL sensitivities at protoDUNE

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Based on arXiv:2304.06765

In collaboration with Pilar Coloma, Jacobo López-Pavón and Laura Molina-Bueno

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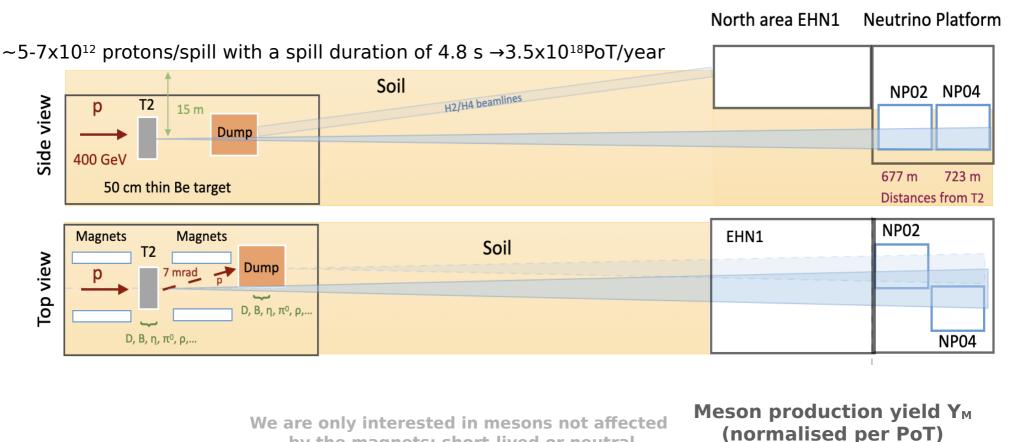






# **Experimental set-up**

#### **Experimental set-up: T2 target**



#### by the magnets: short-lived or neutral $\pi^0$ $\eta'$ D $D_s$ $\eta$ 400 GeV protons $0.05 | 4.8 \cdot 10^{-4} | 1.4 \cdot 10^{-4} | 7.4 \cdot 10^{-6}$ $4.03 \, | \, 0.46 \, |$ $\phi$ $J/\psi$ B $\omega$ $\rho$ $0.54 | 0.53 | 0.019 | 4.4 \cdot 10^{-5} | 1.2 \cdot 10^{-7} | 2.3 \cdot 10^{-8}$

Distributions obtained from Pythia

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## HNL

#### **HNL: Production**

$$L \supset -\frac{m_W}{v} \bar{N} U_{\alpha 4}^* \gamma^{\mu} l_{L\alpha} W_{\mu}^+ - \frac{m_Z}{\sqrt{2}v} \bar{N} U_{\alpha 4}^* \gamma^{\mu} \nu_{L\alpha} Z_{\mu}$$

We consider the simplified phenomenological benchmarks of one HNL mixing with one SM neutrino of a given flavour

 $U_{e4}$ 

 $U_{\mu 4}$ 

 $U_{\tau 4}$ 

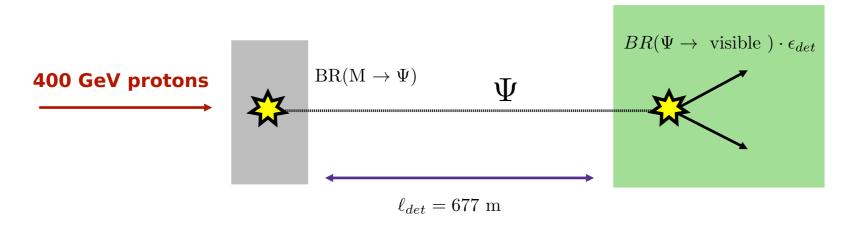
on [			
힣	Parent	2-body decay	3-body decay
We don't have pions and kaons	$\pi^+ \rightarrow$	$e^+N_4$	_
ive pic		$\mu^+ N_4$	
n't ha	$K^+ \rightarrow$	$e^+N_4$	$\pi^0 e^+ N_4$
We do		$\mu^+ N_4$	$\pi^0 \mu^+ N_4$
	$\tau^- \to$	$\pi^- N_4$	$e^-\overline{\nu}N_4$
		$ ho^- N_4$	$\mu^-\overline{\nu}N_4$

Parent	2-body decay	3-body decay
$D^+ \rightarrow$	$e^+N_4$	$e^{+}\overline{K^{0}}N_{4}$
	$\mu^+ N_4$	$\mu^+ \overline{K^0} N_4$
	$ au^+ N_4$	
$D_s^+ \rightarrow$	$e^+N_4$	_
	$\mu^+ N_4$	
	$ au^+ N_4$	

#### (normalised per PoT)

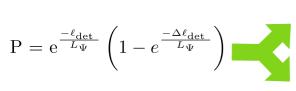
#### New Physics: Decay in flight inside the detector

### **Detector(NP02) Liquid Argon TPC**



$$N_{dec}^{M} = N_{PoT} Y_{M} BR(M \to \Psi) \int dS \int dE_{\Psi} \mathcal{P} \left( c \tau_{\Psi} / m_{\Psi}, E_{\Psi}, \Omega_{\Psi} \right) \frac{dn^{M \to \Psi}}{dE_{\Psi} dS}$$

 $N_{det} = N_{dec}^{M} \cdot BR(\Psi \rightarrow \text{visible}) \cdot \epsilon_{det}$ 

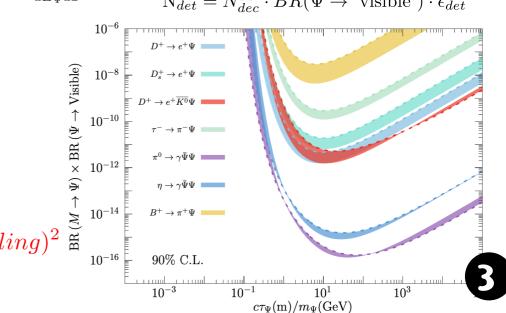


#### Large couplings

$$e^{-rac{\ell_{
m det}}{L_{\Psi}}}$$

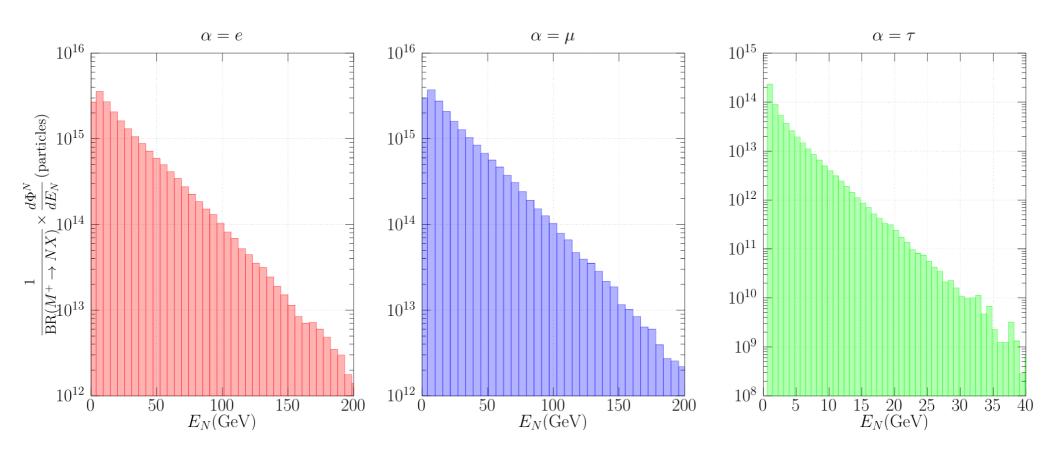
$$1 - e^{-\frac{\Delta \ell_{\text{det}}}{L_{\Psi}}} \propto (coupling)$$

**Small couplings** 



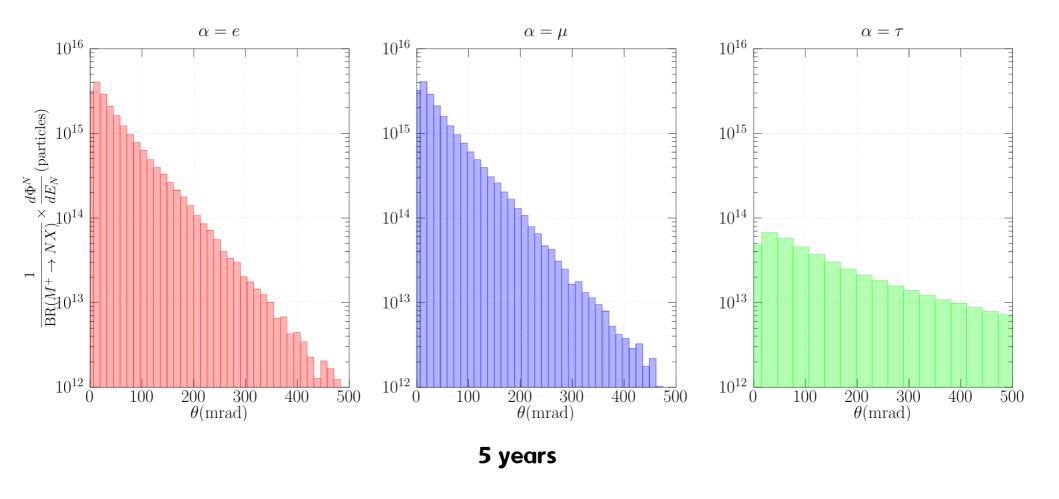
Salvador Urrea (IFIC)

#### **HNL: Fluxes**



5 years

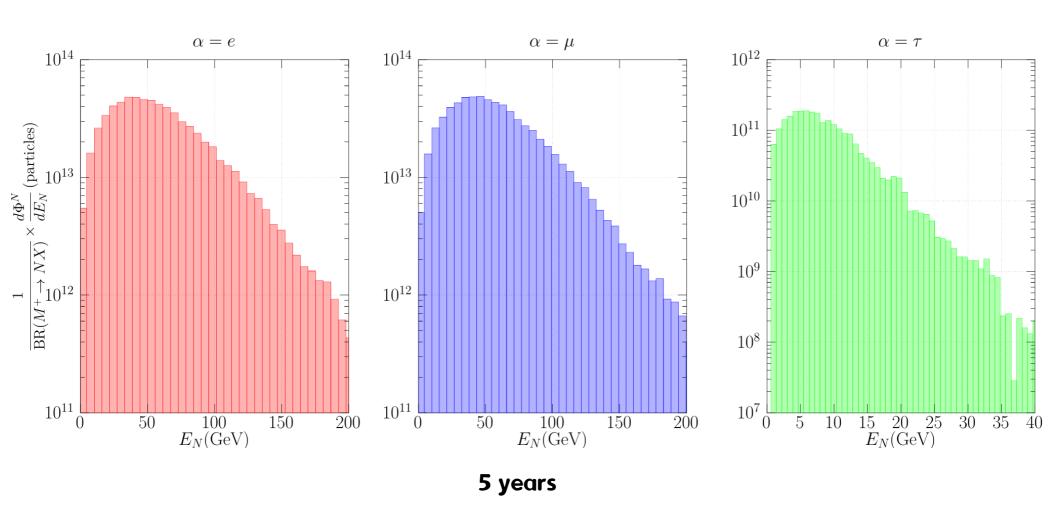
#### **HNL: Fluxes**



 $\theta = \text{Angle of the HNL}$  with respect to the forward direction

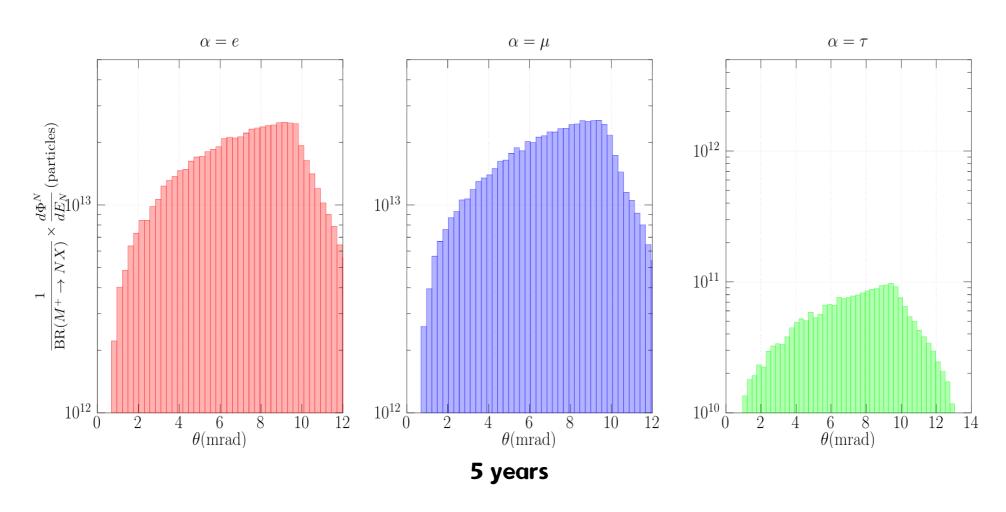
### **HNL: Fluxes with detector acceptance**

#### **HNL** intersecting the detector



#### **HNL: Fluxes with detector acceptance**

#### **HNL** intersecting the detector

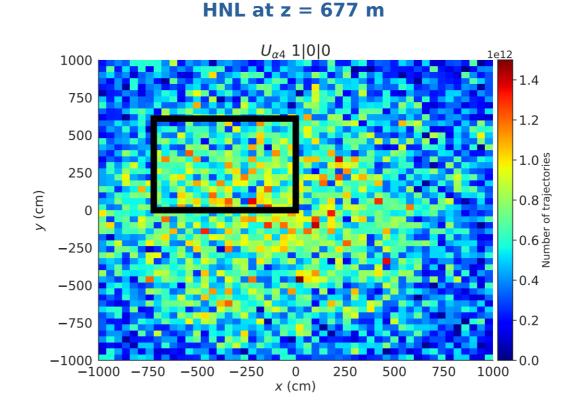


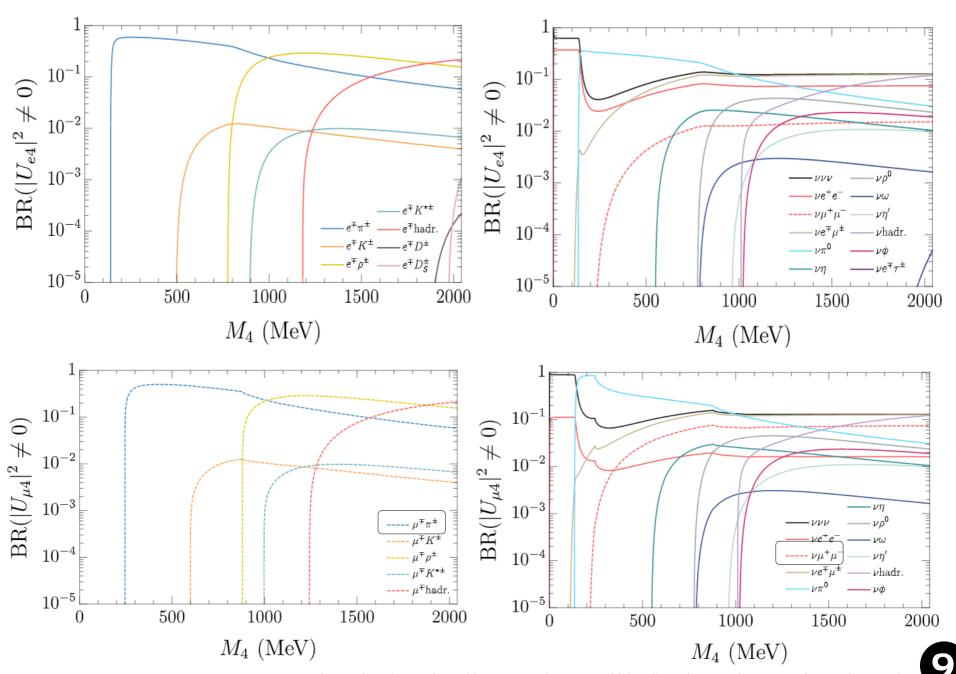
 $\theta = \mbox{Angle}$  of the HNL with respect to the forward direction Salvador Urrea (IFIC)

#### **HNL: Fluxes**

#### • Wide HNL beam

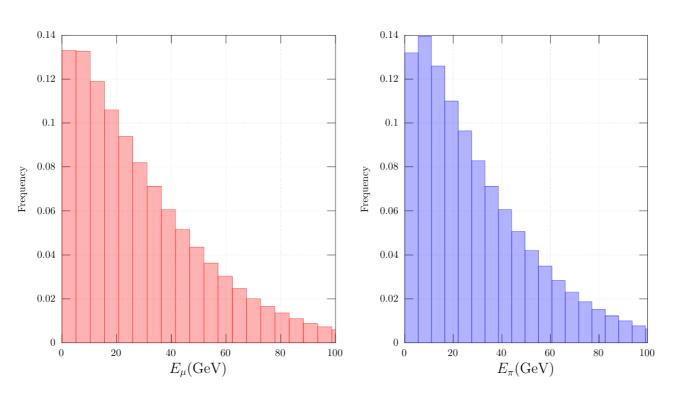
- Small changes in the geometry will not significantly change the results
- Any of the two ProtoDUNE detectors can be used

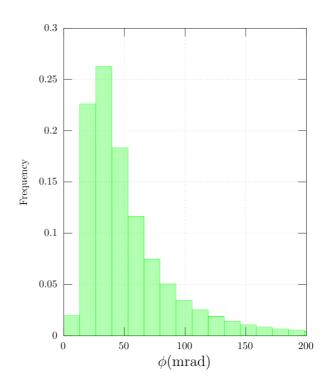




#### Signal: two body decay

$$N_4 \longrightarrow \pi^+ \mu^-$$



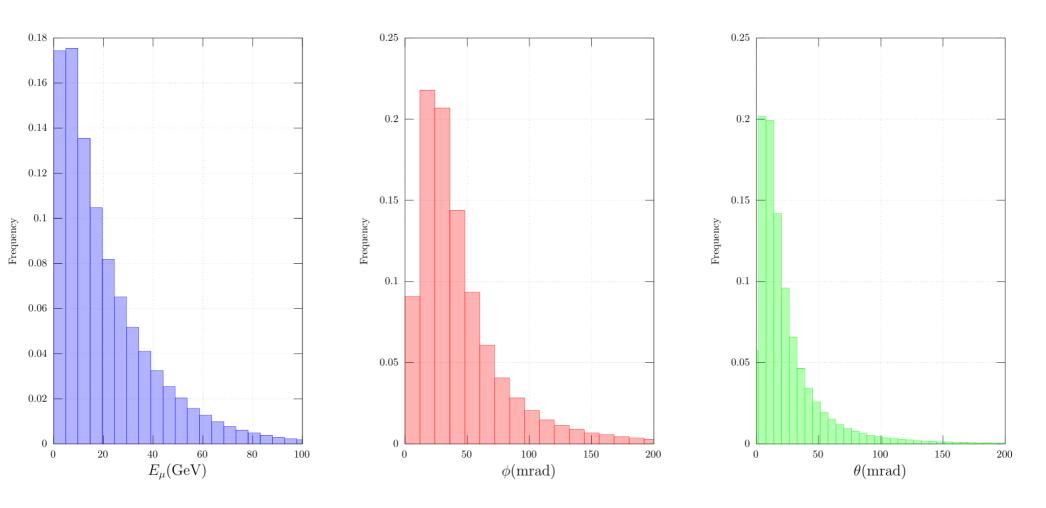


 $\phi =$ Opening angle of the muon-pion pair

$$m_N = 1 \text{GeV}, U_{\mu 4} = 3 \cdot 10^{-8}$$

#### Signal: three body decay

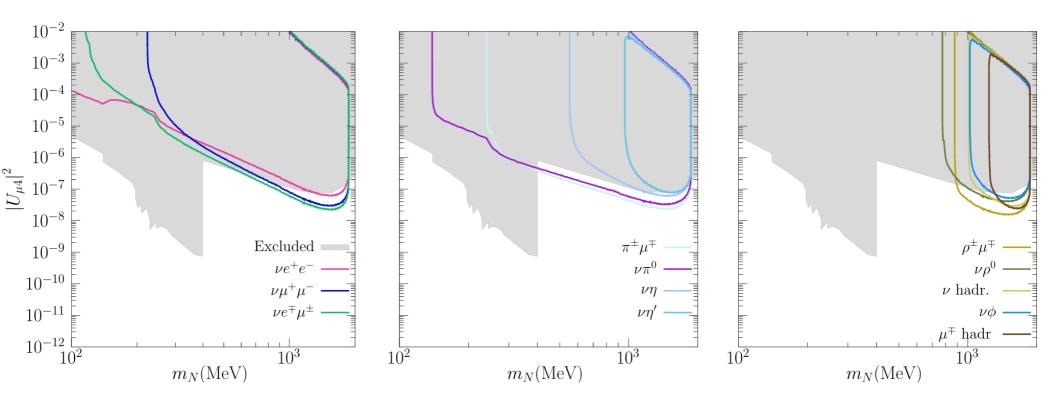
$$N_4 \longrightarrow \nu \mu^+ \mu^-$$



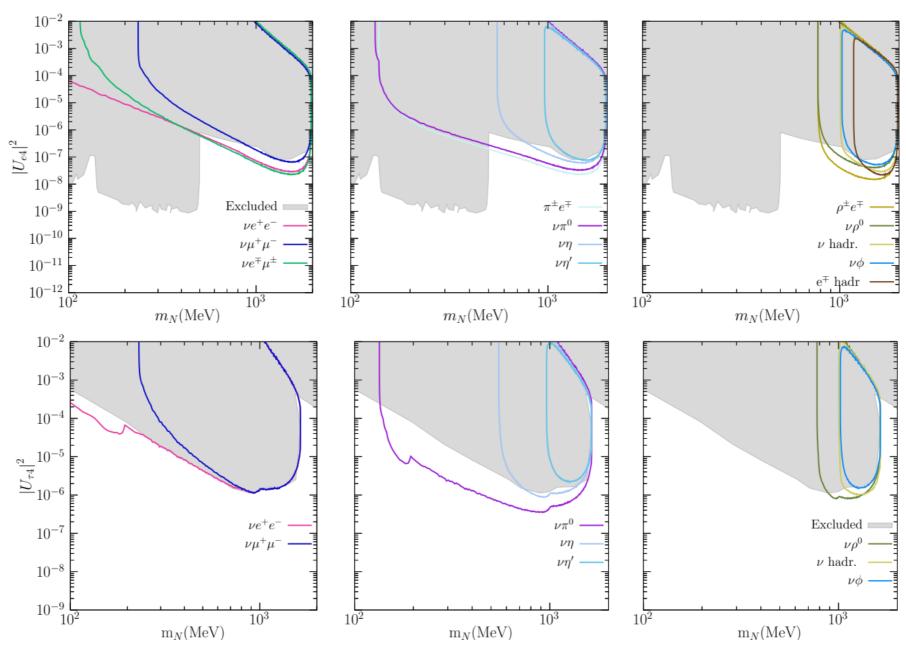
 $\theta = \text{Angle of the muon pair with respect to the forward direction} \qquad m_N = 1 \text{GeV}, U_{\mu 4} = 3 \cdot 10^{-8}$ 

 $\phi =$ Opening angle of the muon pair

### **HNL:** Decays into visible channels

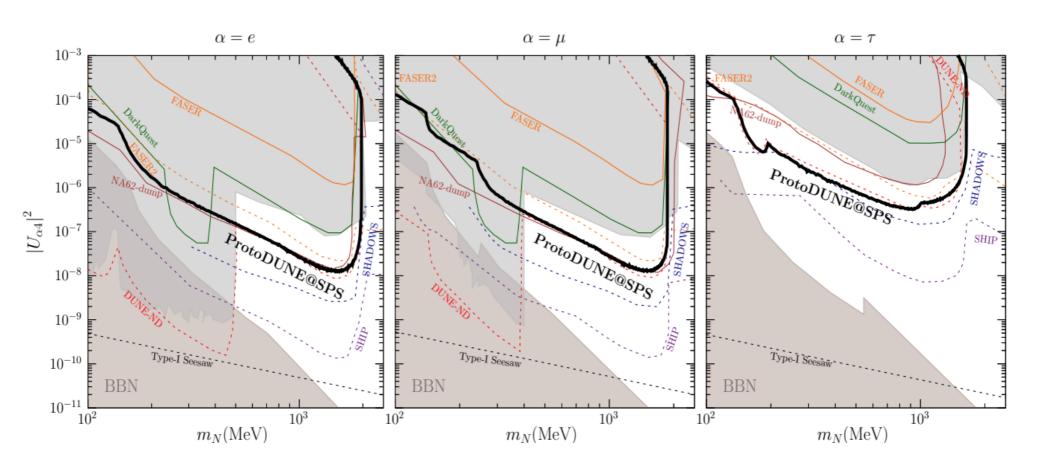


### **HNL:** Decays into visible channels



#### HNL: Decays into visible channels (combination)

We consider the following channels  $N \to \nu e e, \nu \mu \mu, \nu e \mu, e \pi, \mu \pi$  and  $\nu \pi^0$ 



## Thank you











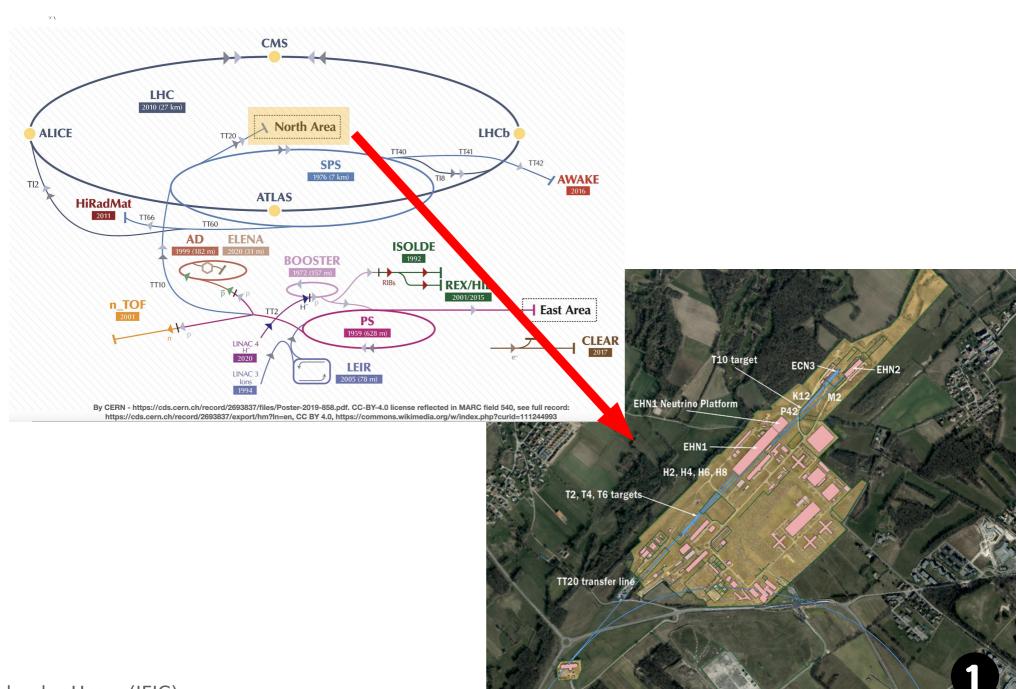






# Back-up

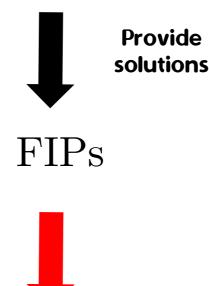
### **Experimental set-up: Extracted beam lines**



#### **Motivation**

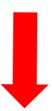
## Open problems in Particle Physics

Origin of neutrino masses, Baryon asymmetry of the universe and the origin of dark matter



They come in many forms

Vector (*Dark Photon*), Scalar (*Dark Higgs*), Fermion (*Heavy neutral lepton*), Pseudo-scalar (*Axion*)



Both the interaction strengths with SM particles and the masses of the FIPs range over many orders of magnitude.



Many different types of experiments are needed



When the interaction stregth is sufficiently large and the mass ranges from  $10^{-2}$  GeV to 10 GeV, it can be accessed by accelerator-based experiments

ProtoDUNE run as a Fixed-target experiment

## Summary

 The excellent imaging capabilities, the large fiducial volume and the convenient location with respect to the T2 target of the ProtoDUNE detectors make them ideal to search for weakly interacting massive particles in Beyond Standard Model scenarios, such as long lived unstable particles and stable particles. In particular HNL and millicharged particles

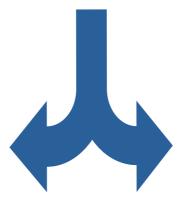
## **Outlook**

- A dedicated analysis is required to determine the expected backgrounds and efficiencies for the different detection channels consider.
- The development of a dedicated new trigger is needed for this type of searches
- Other models of new physics can be explored: ALPs, light dark matter, etc.

# **New Physics**

### **New Physics: Type of searches**

New particles produced in meson decays



Long-lived

(HNL, ALPs, dark photon,...)

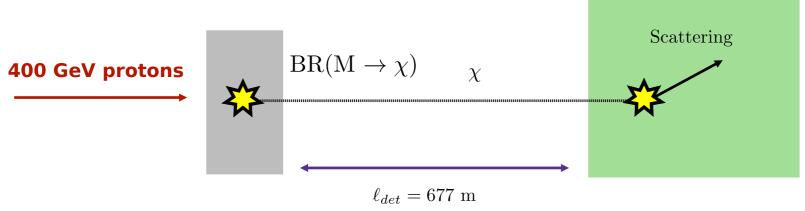
Decay in flight inside the detector Modify cross sections

Very long-lived (Stable)

(Milicharged particles,...)

#### **New Physics: stable particles**

# **Detector(NP02) Liquid Argon TPC**



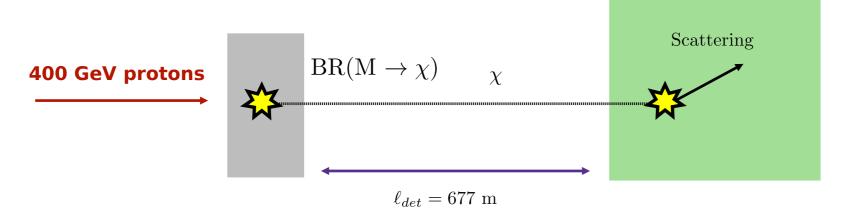
$$\langle \sigma \rangle = \frac{1}{\Phi^{\chi}} \int_{0}^{\infty} \int_{T^{\min}}^{T^{\max}} \frac{d\sigma}{dT} (E_{\chi}, \{X\}) \frac{d\Phi^{\chi}}{dE_{\chi}} dT dE_{\chi} \qquad \qquad N_{ev} = \epsilon_{\det} N_{trg} \langle \sigma \rangle \Phi^{\chi} N_{PoT},$$

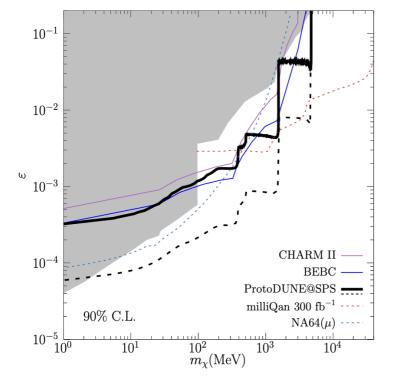
$$10^{-6} \int_{10^{-7}}^{10^{-6}} \frac{d\sigma}{dT} (E_{\chi}, \{X\}) \frac{d\Phi^{\chi}}{dE_{\chi}} dT dE_{\chi} \qquad \qquad N_{ev} = \epsilon_{\det} N_{trg} \langle \sigma \rangle \Phi^{\chi} N_{PoT},$$

$$10^{-2} \int_{0}^{\infty} \frac{10^{-8}}{10^{-10}} \frac{10^{-8}}{10^{-10}} \int_{0}^{\infty} \frac{10^{-10}}{10^{-10}} \int_{0}^{\infty$$

#### Millicharged particles

# **Detector(NP02) Liquid Argon TPC**





$$N_{ev} = \epsilon_{det} N_{trg} \langle \sigma \rangle \Phi^{\chi} N_{PoT},$$

$$\sigma \sim \varepsilon^2 \left( \frac{30 \text{MeV}}{T_{\text{min}}} \right) 10^{-26} \text{ cm}^{-2},$$

