

On an active muon veto system for the COSINUS experiment

Mariano Cababie



TECHNISCHE
UNIVERSITÄT
WIEN

HEPHY / TU Wien
@ ALPS 2024-DM

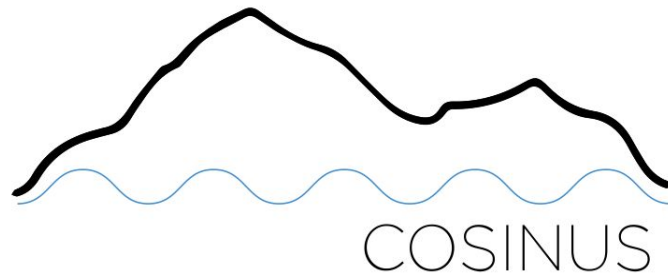


HEPHY
INSTITUTE OF
HIGH ENERGY PHYSICS



COSINUS

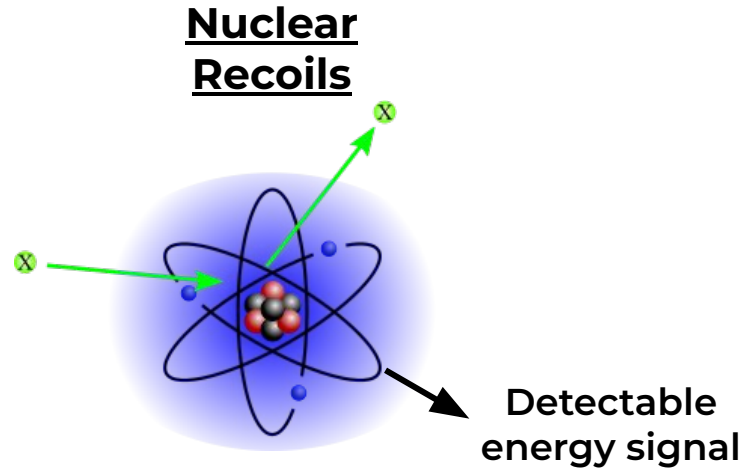
Cryogenic **O**bservatory for **S**ignals seen in
Next Generation **U**nderground **S**earches



Direct dark matter detection

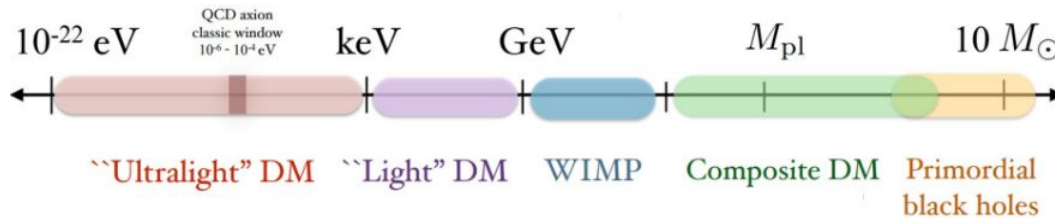
Dark matter

- ❖ No charge
- ❖ No self-interaction
- ❖ Cold
- ❖ Stable



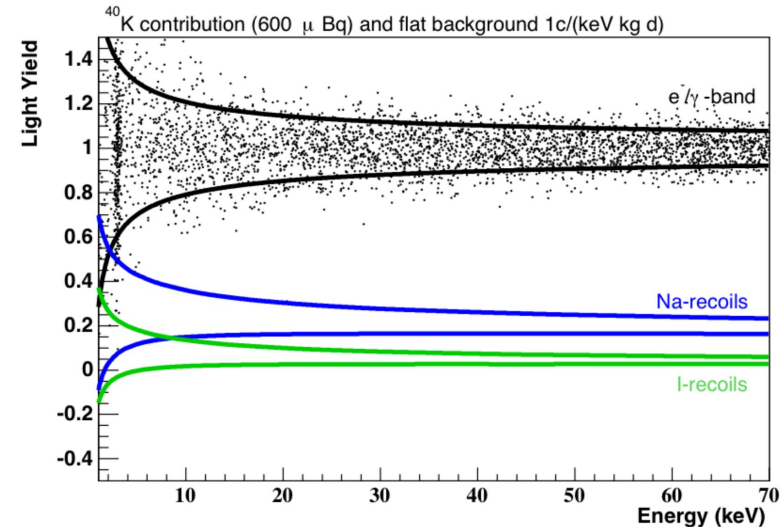
WIMP

- ❖ All above and...
- ❖ $O(\text{GeV})$ mass

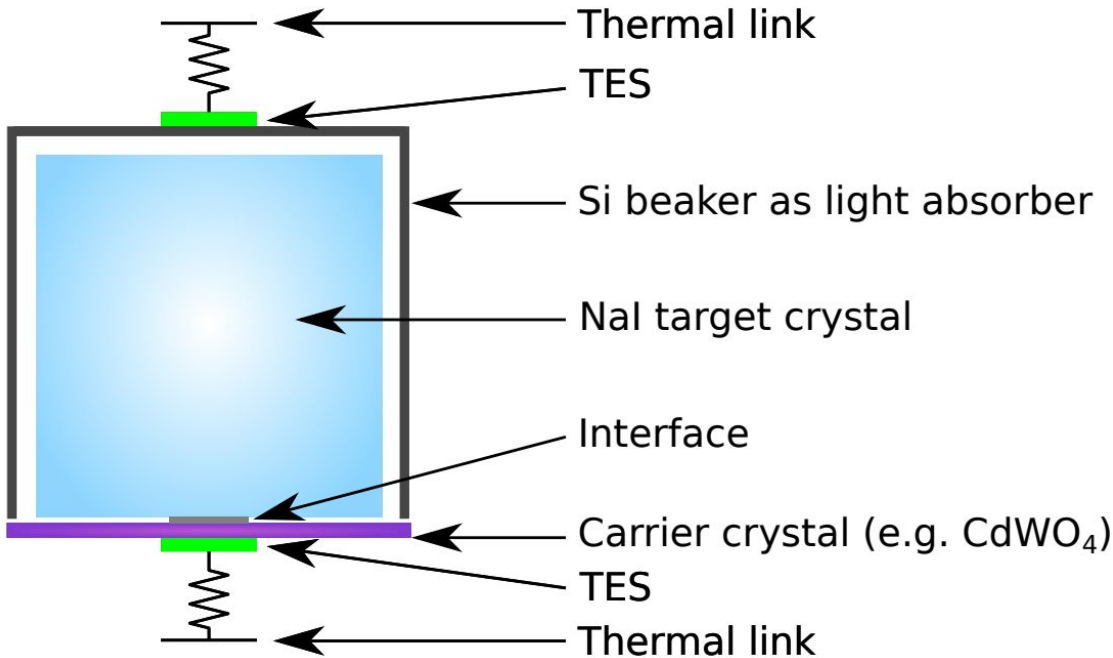


Cryogenic detectors

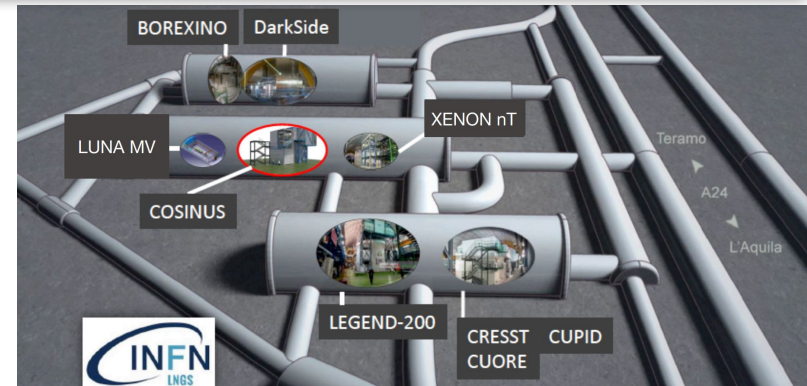
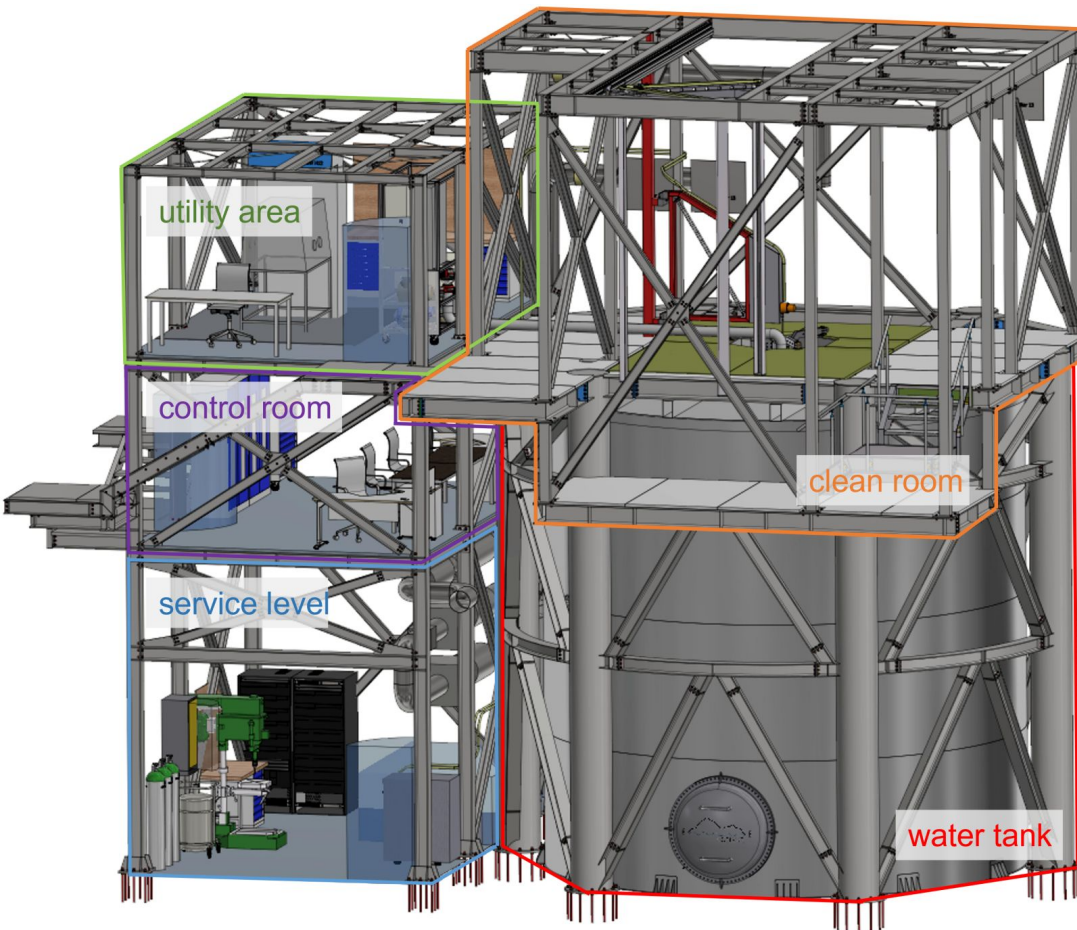
Particle discrimination (NaI)



$$\text{Light yield} = \frac{\text{Energy in Light}}{\text{Total Energy}}$$

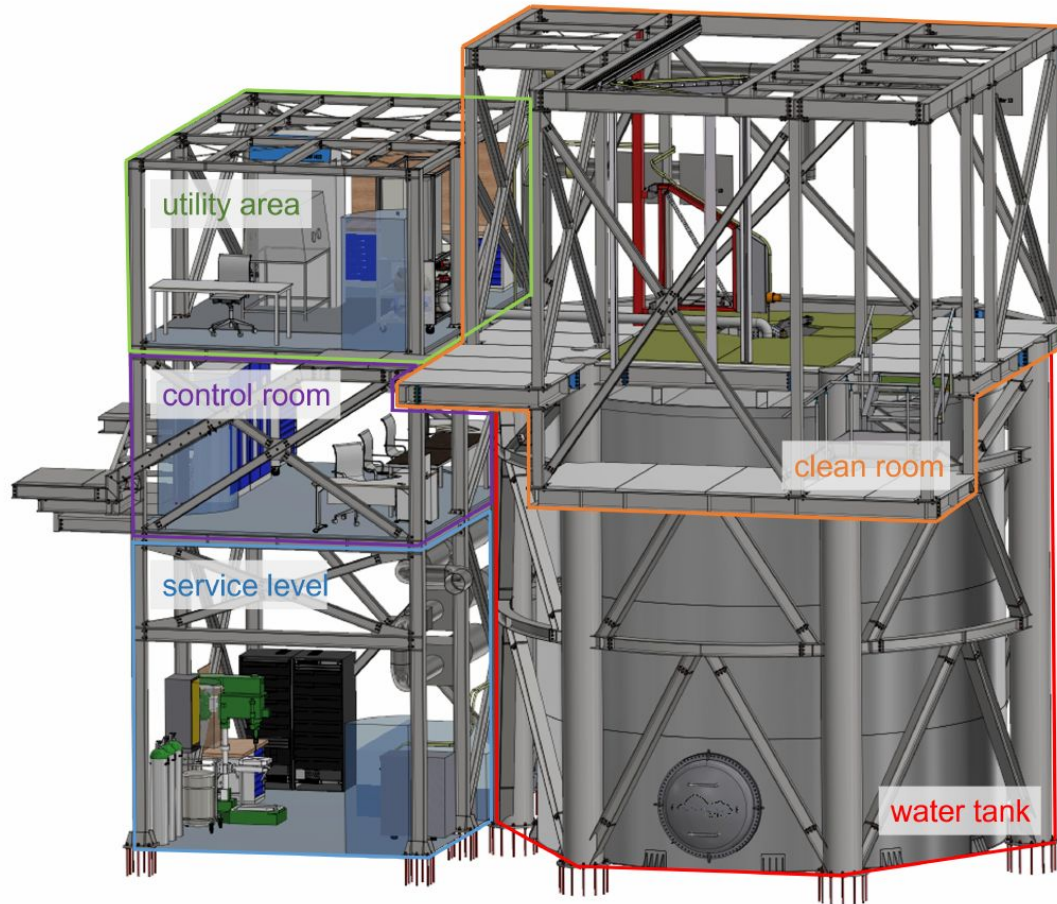


The COSINUS experiment site



- ◆ Clear division in multiple working areas
- ◆ Instrumented water tank for muon veto, surrounding cryostat.
- ◆ Cleanroom facility above cryostat/water tank for easy access and sample manipulation

The COSINUS experiment site

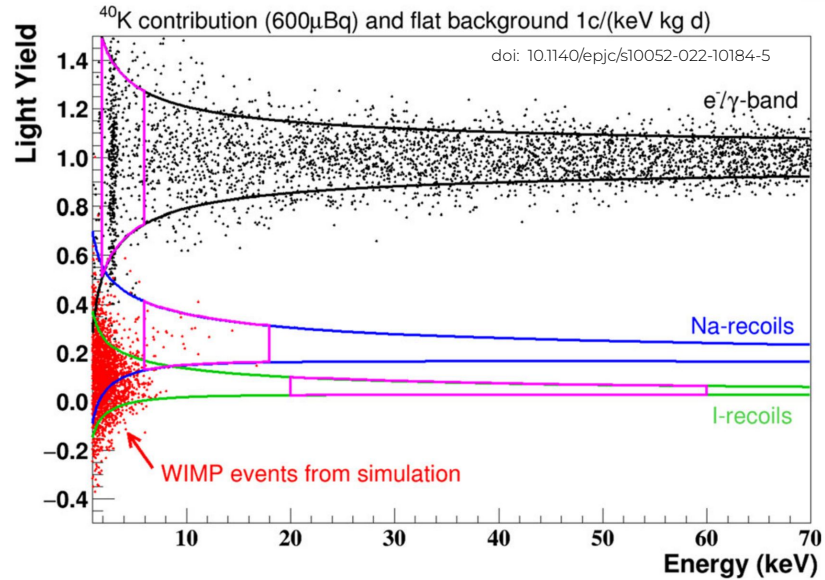


The COSINUS experiment site



The COSINUS experiment site

Danger!

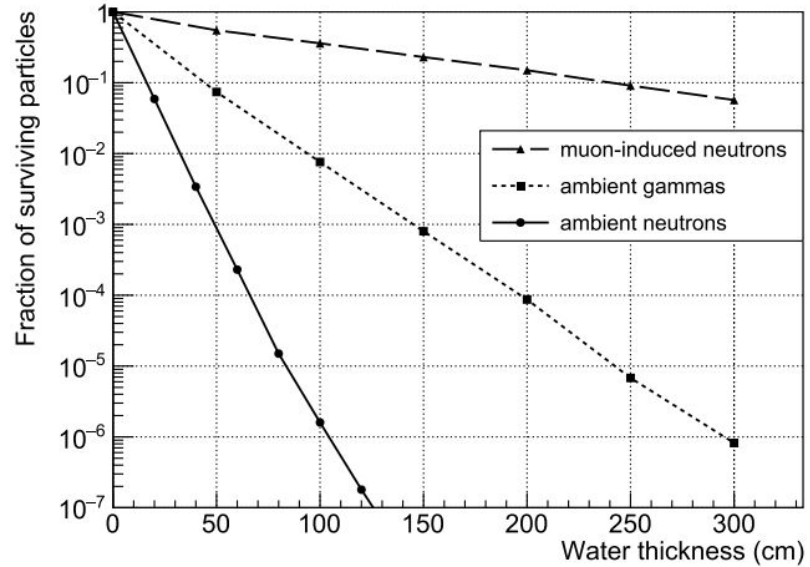


Neutrons!



The COSINUS experiment site

- ❖ Water tank as **passive** shielding
- ❖ Ambient and radiogenic neutron simulated at 1 neutron per kg . year for NaI detector



The COSINUS muon veto system

Publications on muon veto system

- ❖ **Passive shielding:** *“Simulation-based design study for the passive shielding of the COSINUS dark matter experiment”*, 2022, doi: 10.1140/epjc/s10052-022-10184-5
- ❖ **Active shielding:** *“Water Cherenkov muon veto for the COSINUS experiment: design and simulation optimization”*, on peer-review process.

The COSINUS muon veto system

Publications on muon veto system

- ❖ **Passive shielding:** *“Simulation-based design study for the passive shielding of the COSINUS dark matter experiment”*, 2022, doi: 10.1140/epjc/s10052-022-10184-5



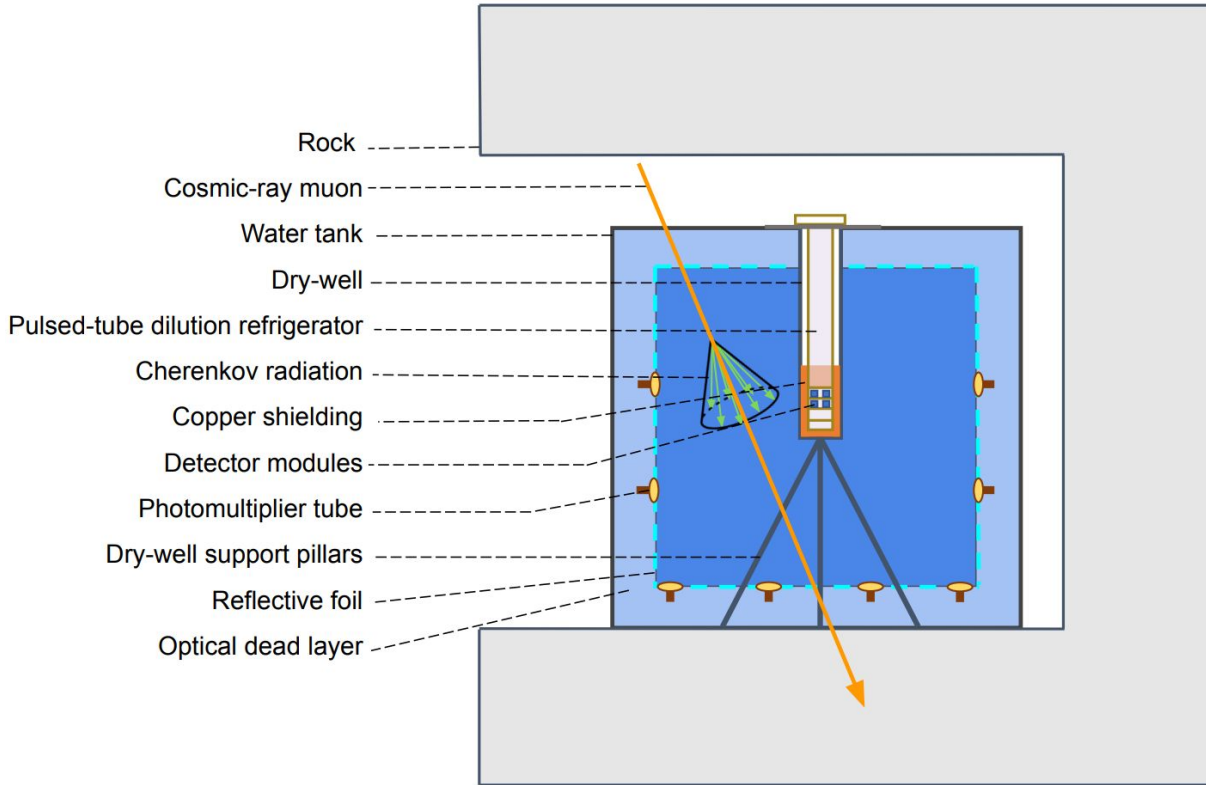
“...muon-induced spallation and subsequent hadronic showers result in a further emission of neutrons. A muon veto to tag such events is therefore crucial.”

- ❖ **Active shielding:** *“Water Cherenkov muon veto for the COSINUS experiment: design and simulation optimization”*, on peer-review process.



“The active veto will reduce the cosmogenic neutron background rate to $0.11 \pm 0.02 \text{ cts} \cdot \text{kg}^{-1} \cdot \text{year}^{-1}$ ”

The COSINUS muon veto system



- ❖ Cherenkov light can come from cosmogenic muons (95%) or showers of those muons (5%).
- ❖ MUSUN code used for generating muons, ImpCRESST used for physics simulation
- ❖ 7x7 cylindrical water tank with 28 PMTs. Internal surface covered with reflective foil

Goals and challenges

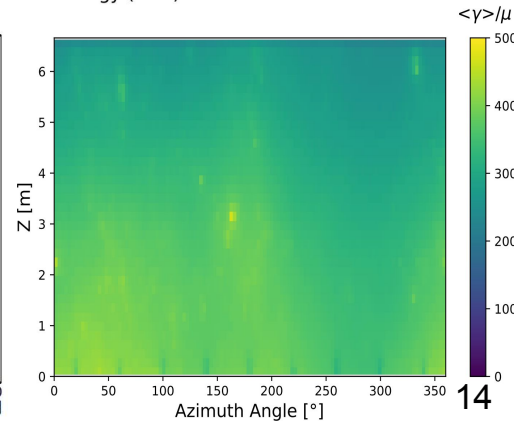
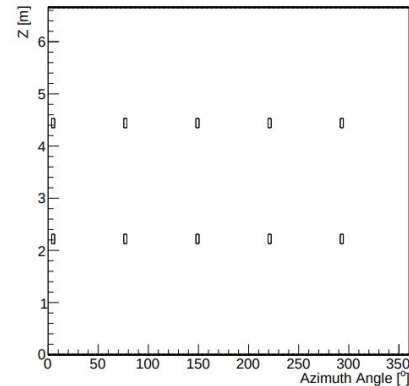
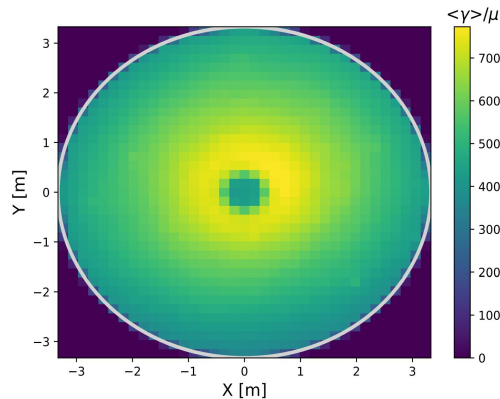
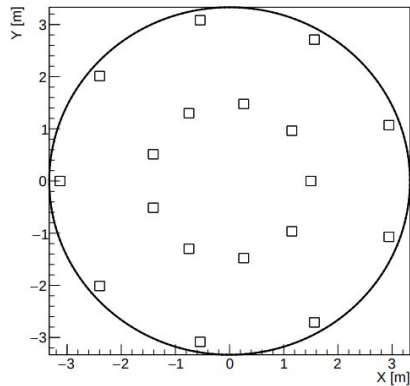
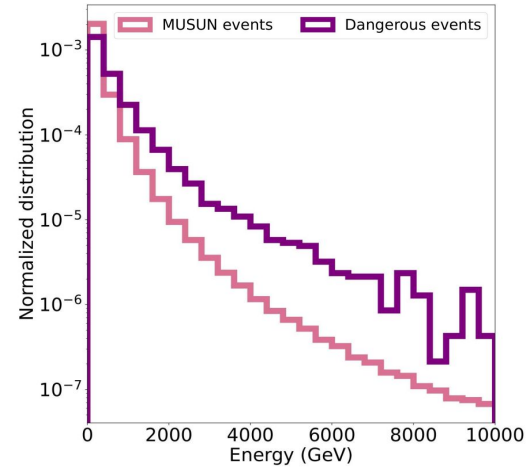


Hamamatsu R5912-30

- ❖ Effectively detect (and veto) muons that generate neutrons
- ❖ Consider the ambient gamma background ($\sim 10^5$ γ s per second through water tank)
- ❖ Take dark counts ($\sim 1200/6000$ Hz) into account

Muon veto simulation

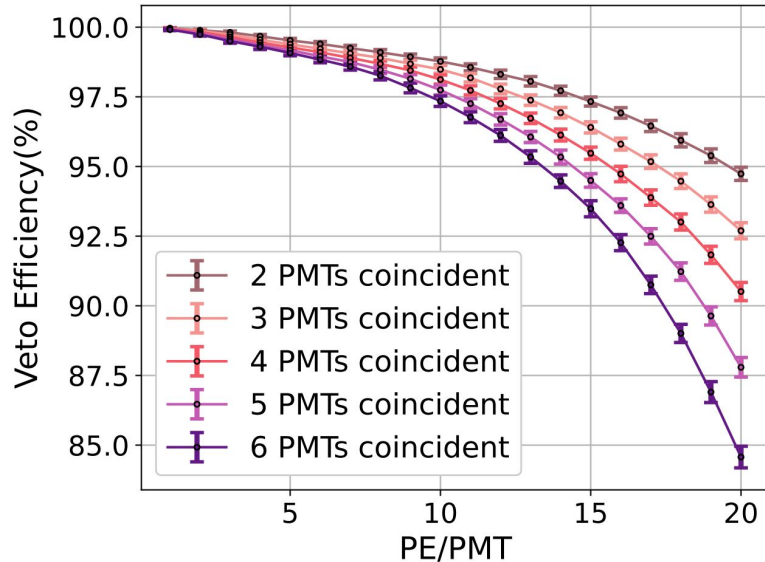
- ❖ 30×10^6 muons simulated (~ 13 years) \rightarrow 11682 ($\sim 0.04\%$) tagged as dangerous (i.e. muon-induced neutrons enter the detector volume)
- ❖ Higher energy events tend to be more dangerous (negligible angular resolution)
- ❖ Illumination map can be reconstructed:



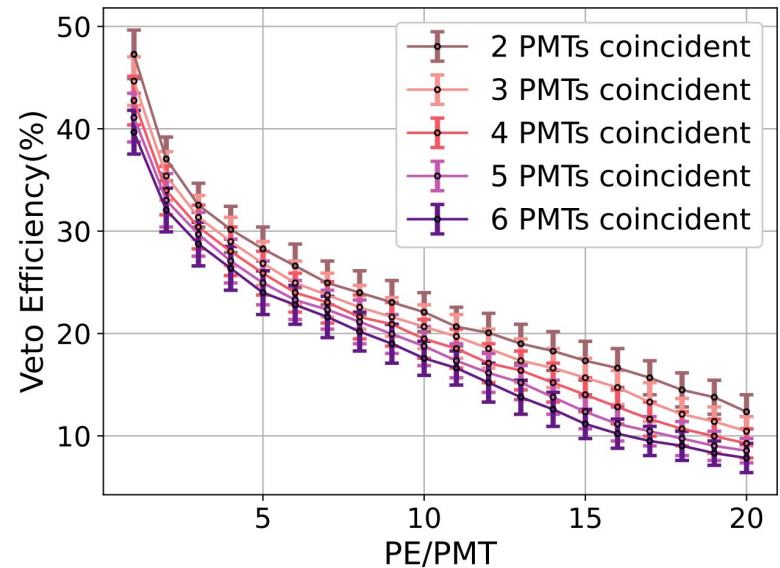
Muon veto efficiency

- ❖ Efficiency can be **optimized** varying two parameters: **threshold** and **coincidence**
- ❖ Ideally, threshold will be set at 1 PE, what about coincidence?

Muon events



Shower events



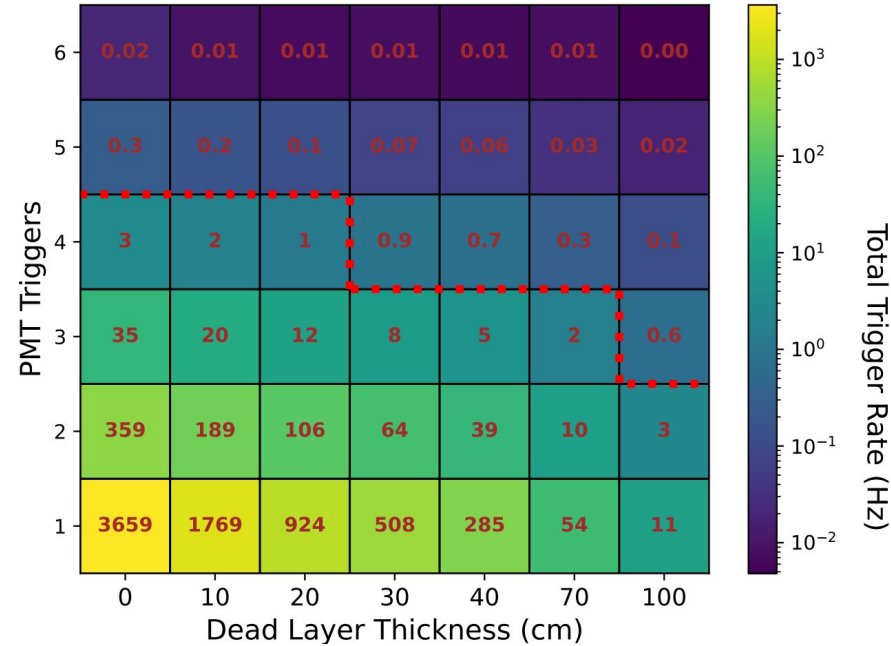
Ambient gamma background

- ❖ Gamma photons sub-products can also generate Cherenkov radiation
- ❖ Utilizing an optical dead layer may drastically reduce this background
- ❖ The probability p of having a spurious background event in coincidence with a signal event in our detector during a period ΔT is given by

$$p(\Delta T) = 1 - e^{-R_B \Delta T}$$

where R_B is the rate of the background event.

For a $\Delta T = 10\text{ms}$ and $p < 0.01$, R_B should be $< 1\text{Hz}$.



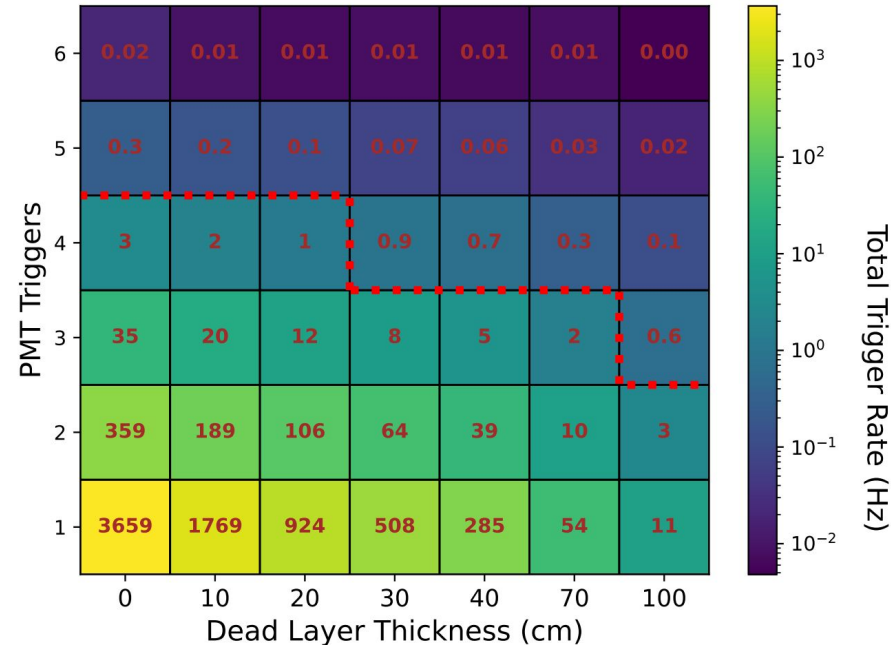
Ambient gamma background

- ❖ Gamma photons sub-products can also generate Cherenkov radiation
- ❖ Utilizing an optical dead layer may drastically reduce this background
- ❖ The probability p of having a spurious background event in coincidence with a signal event in our detector during a period ΔT is given by

$$p(\Delta T) = 1 - e^{-R_B \Delta T}$$

where R_B is the rate of the background event.

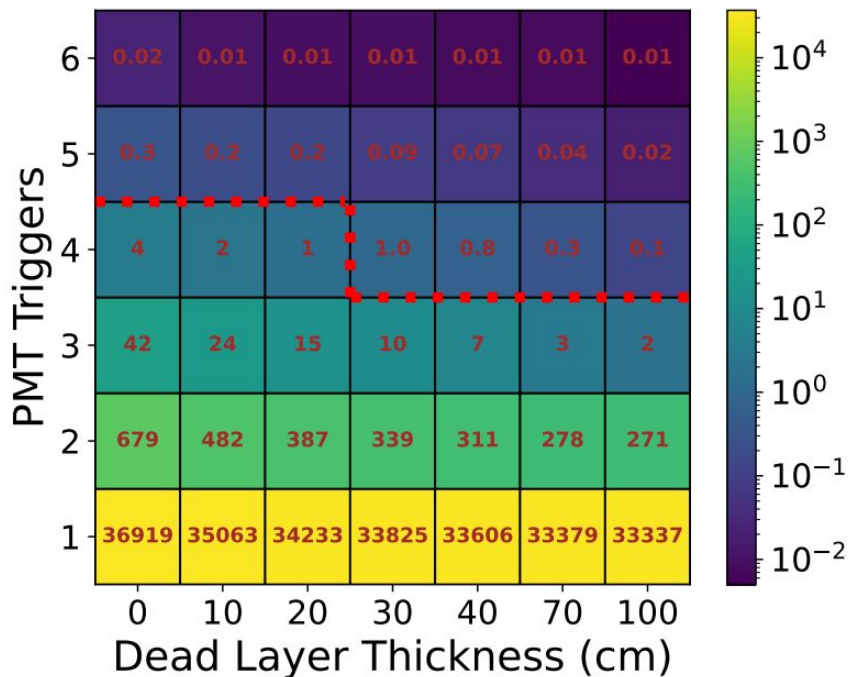
For a $\Delta T = 10\text{ms}$ and $p < 0.01$, R_B should be $< 1\text{Hz}$.



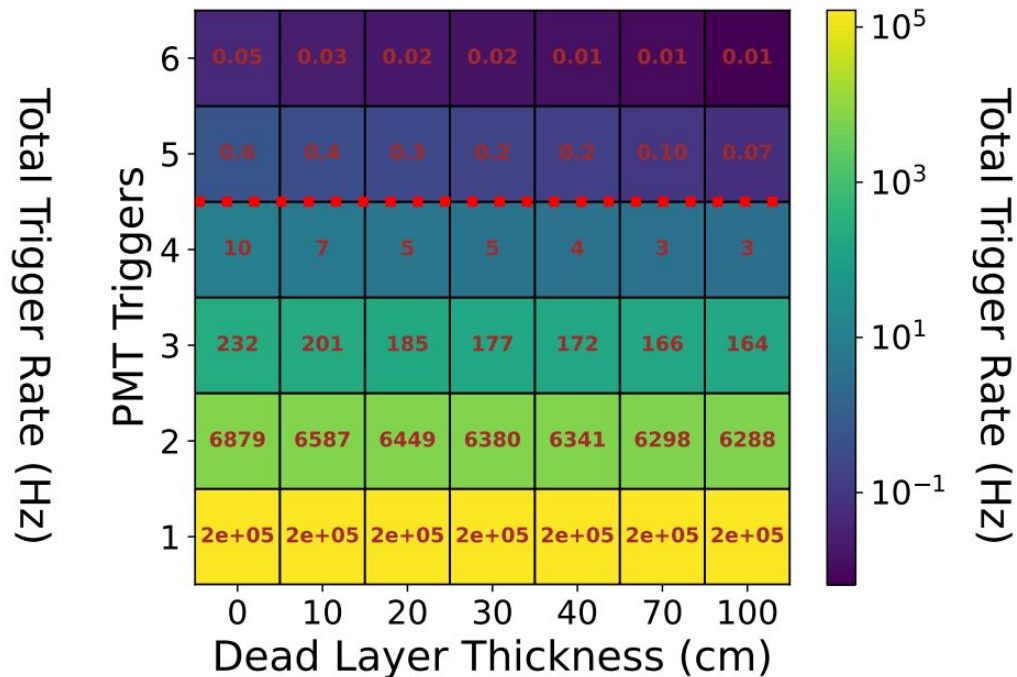
Dead layer not recommended!

Dark current

- ❖ Dark current can be taken into account by adding it as a Poisson-like contribution over gamma background



1200Hz



6000Hz

Conclusions

- ❖ Effectively detect (and veto) muons that generate neutrons

5-fold PMT coincidence yield a efficiency veto of

→ **99.63(16)%** for muons and **44.4+-5.6 %** for showers
(while the first composes 95% of dangerous events!)



Hamamatsu R5912-30

- ❖ Consider the ambient gamma background($\sim 10^5$ γ s per second though water tank)

→ **5-fold PMT coincidence reduces rate <1Hz**

- ❖ Take dark counts ($\sim 1200/6000$ Hz) into account

→ **5-fold PMT coincidence reduces rate <1Hz**

Conclusions

- ❖ Effectively detect (and veto) muons that generate neutrons

5-fold PMT coincidence yield a efficiency veto of

→ **99.63(16)%** for muons and **44.4+-5.6 %** for showers
(while the first comprises 95% of dangerous events!)



Hamamatsu R5912-30

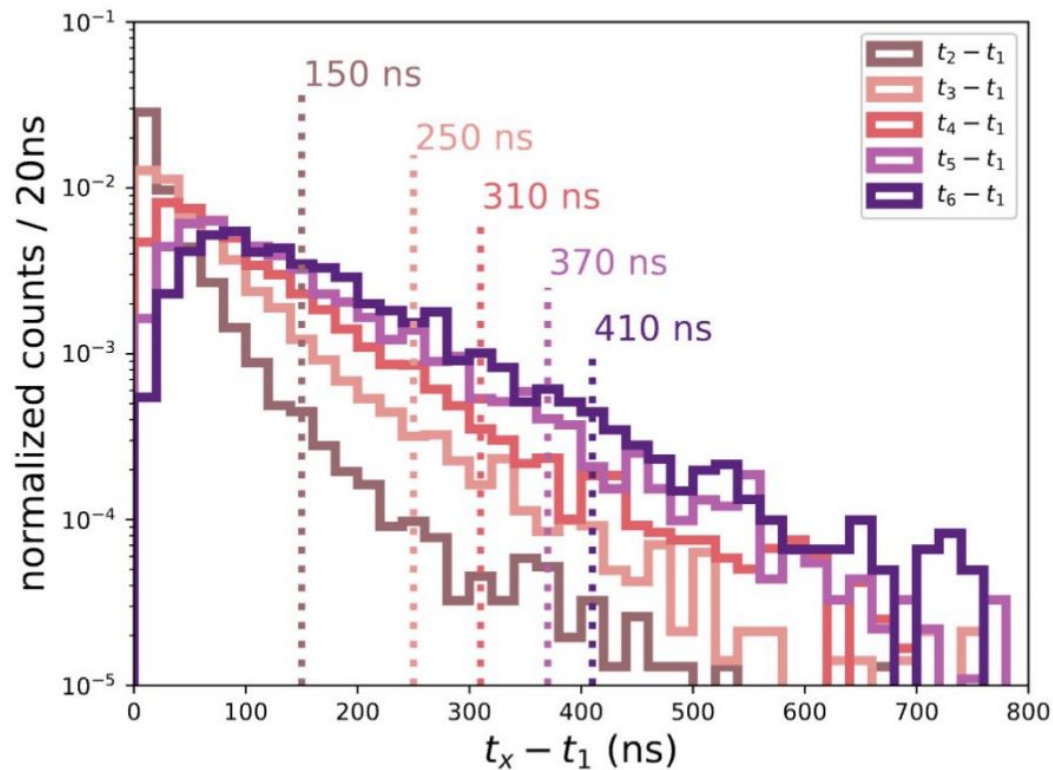
- ❖ Consider the ambient gamma background ($\sim 10^5$ γs per second though water tank)

→ **5-fold PMT coincidence reduces rate <1Hz**

- ❖ Take dark counts ($\sim 1200/6000$ Hz) into account

→ **5-fold PMT coincidence reduces rate <1Hz**

Muon-induced
neutron
background of
 0.11 ± 0.02
 $\text{cts} \cdot \text{kg}^{-1} \cdot \text{year}^{-1}$



- For a 6-fold PMT coincidence 95% of events will be tagged with a coincidence window of > 410 ns