

On the speed of sound in dense two colour QCD

A Quick Update

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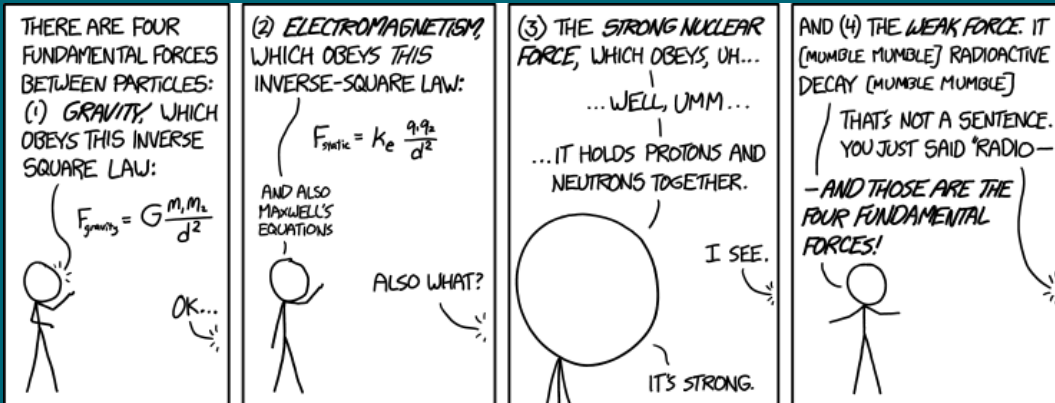
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ITPQFTM+ 2023

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- 1 What is Lattice QCD?
 - 2 Scale Setting
 - 3 Lattice Action
 - 4 Why Two Colour QCD (QC₂D)?
 - 5 Thermodynamic Results
 - 6 Neutron Star Speed of Sound
 - 7 Conclusion
- Acronyms78



"Of these four forces, there's one we don't really understand." "Is it the weak force or the strong—" "It's gravity."

Euclidean QCD Lagrangian given by

$$\mathcal{L}_{QCD} = \frac{1}{4} \text{Tr} G_{\mu\nu} G_{\mu\nu} + \bar{\psi} [\gamma_\mu (\partial_\mu + ig_s A_\mu) + m] \psi \quad (1)$$

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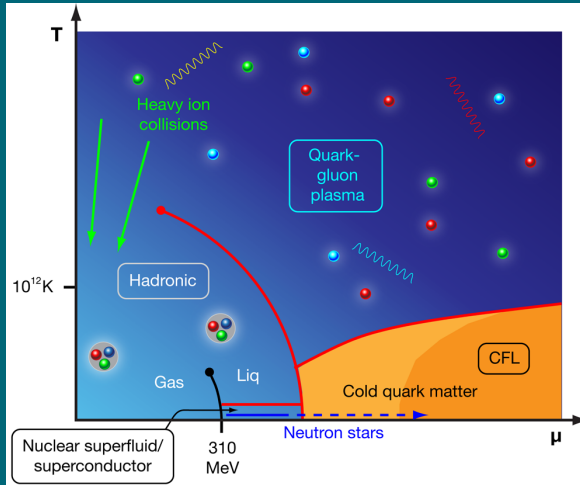
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- For $g_s \ll 1$ can use perturbation theory (very high temperature/very short distances)
- For $g_s \sim 1$ or non-perturbative phenomena must use non-perturbative approach



A sketch of the QCD Phase Diagram

- Want to solve the path integral

$$\langle \mathcal{O} \rangle = \frac{1}{Z} \int \mathcal{D}[\Phi] \mathcal{O}[\Phi] e^{-S[\Phi]} \quad (2)$$

- Space-time itself is discretised
- 4-Dimensional Lattice with periodic boundary conditions.
- Fermions lie on vertices, gauge bosons on the links

Gluon fields:

$$A_\mu(x) \in \mathfrak{su}(N) \quad (3)$$

$$U_\mu(n) \in SU(N) \quad (4)$$

Gauge Action:

$$S_g[A] = -\frac{1}{2g^2} \int d^4x G_{\mu\nu}^a G_{\mu\nu}^a \quad (5)$$

$$S_g[U] = -\frac{a^4}{2g^2} \sum_{n \in \Lambda} \text{Tr} G_{\mu\nu}(n) G^{\mu\nu}(n) \quad (6)$$

This discretisation is not unique

Fermion Action:

$$S_F [\psi, \bar{\psi}, A] = \sum_{f=1}^{N_f} \int d^4x \bar{\psi}^{(f)}(x) \left(\gamma_\mu (\partial_\mu + iA_\mu(x)) + m^{(f)} \right) \psi^{(f)}(x) \quad (7)$$

$$S_F[\psi, \bar{\psi}, U] = \alpha^4 \sum_{n \in \Lambda} \sum_{f=1}^{N_f} \bar{\psi}^{(f)}(n) \left(\sum_{\mu=1}^4 \gamma_\mu \frac{U_\mu(n) \psi_\mu^{(f)}(n + \hat{\mu}) - U_\mu^\dagger(n - \hat{\mu}) \psi_\mu^{(f)}(n - \hat{\mu})}{2a} + m \psi^{(f)}(n) \right) \quad (8)$$

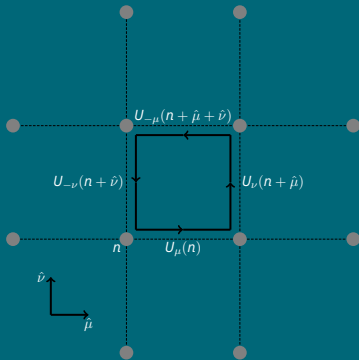


Figure: A Plaquette.

- Bosonic observables represented by closed loop on the lattice (Wilson loop)
- Closed loop in temporal extent (Polyakov loop) is order parameter for deconfinement
- Simplest closed loop is a 1×1 square called a *plaquette*

$$\begin{aligned}
 U_{\mu\nu}(n) = & \\
 & U_{\mu}(n)U_{\nu}(n + \hat{\mu})U_{-\mu}(n + \hat{\mu} + \hat{\nu})U_{-\nu}(n + \hat{\nu}) = \\
 & U_{\mu}(n)U_{\nu}(n + \hat{\mu})U_{\mu}^{\dagger}(n + \hat{\nu})U_{\nu}^{\dagger}(n) \quad (9)
 \end{aligned}$$

- Gauge configurations produced with probability weight

$$e^{-S[U]} = \det M[U] e^{-S_g[U]} \quad (10)$$

using Hybrid Monte Carlo (HMC). $M[U]$ fermion matrix.

- Metropolis algorithm is inefficient. Updates one site per step.
- updating all sites naïvely gives very large changes to action
- HMC involves a global update in a fictitious time τ defined by *Hamiltonian* instead.

Generating Configurations

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Fun Fact!

For an exact calculation, $\delta H = 0$ due to energy conservation so all fields would be accepted.

- Observables on the lattice are dimensionless, must set the scale to give a physical meaning.
- Choose an observable and use it to fix scale.
- We use string tension $\sigma = (440 \text{ MeV})^2$ and compare using static quark potential.

Ensemble	β	κ	a (fm)	$\frac{m_\pi}{m_\rho}$	m_π (MeV)
V. Coarse	1.7	0.178	0.233	0.779(4)	688(11)
Coarse	1.9	0.1680	0.178(6)	0.805(9)	717(25)
Light	1.7	0.1810	0.189(4)	0.61(5)	638(33)
Fine	2.1	0.1577	0.138(6)	0.810(7)	637(28)
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Table: Table of lattice parameters, lattice spacings and pion masses

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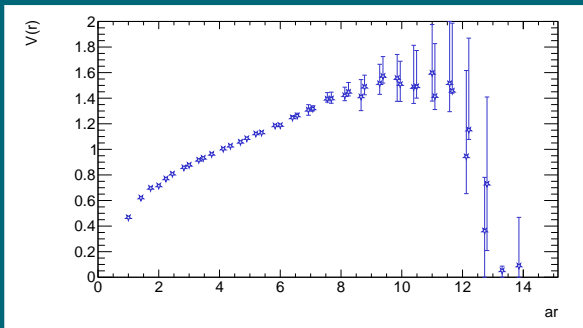


Figure: Static Quark Potential for $\beta = 2.0$, $\kappa = 0.1642$ on a $16^3 \times 32$ lattice.

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- 1 Produce gauge configs for chose of β and κ .
- 2 Calculate pseudoscalar/vector correlators and extract mass ratio.
- 3 Evaluate Wilson Loops/Lines numerically
- 4 Fit Wilson lines to form

$$W(r, \tau) = Ae^{-V(r)\tau} \quad (11)$$

with fit parameters A and $V(r)$

- 5 Fit $V(r)$ to the Cornell Form in lattice units

$$aV(r) = aV_0 + \frac{\alpha}{\left(\frac{r}{a}\right)} + a^2\sigma\frac{r}{a} \quad (12)$$

Momentum Space Massless Free Fermion Propagator

$$\tilde{D}(p)^{-1} \Big|_{m=0} = \frac{i \sum_{\mu=1}^4 \gamma_{\mu} p_{\mu}}{p^2} \quad (13)$$

$$\tilde{D}(p)^{-1} \Big|_{m=0} = \frac{ia^{-1} \sum_{\mu=1}^4 \gamma_{\mu} \sin(p_{\mu}a)}{a^{-2} \sum_{\mu=1}^4 \sin^2(p_{\mu}a)} \quad (14)$$

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- Who ordered those!?!
This is the infamous fermion doubling problem. Techniques for removing these doublers include

- Wilson Fermions (increase mass of doublers so they decouple from the theory)
- (Rooted) Staggered Fermions (mixes Dirac and space-time indices)
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- Want translational invariance, hermiticity and locality; you get an equal number of left and right handed fermions. No-Go Theorem!²
- This work uses Wilson Fermions, which explicitly break chiral symmetry.

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Wilson Fermions

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- Add term to momentum space Dirac operator that is
 - 1 0 if $p_\mu = 0$
 - 2 Non-zero if $p_\mu = \frac{\pi}{a}$

- Naïve Dirac operator given by

$$\tilde{D}(p) = m + \frac{i}{a} \sum_{\mu=1}^4 \gamma_{\mu} \sin(p_{\mu}a) \quad (15)$$

- Instead consider

$$\tilde{D}(p) = m + \frac{i}{a} \sum_{\mu=1}^4 \gamma_{\mu} \sin(p_{\mu}a) + \frac{1}{a} \sum_{\mu=1}^4 (1 - \cos(p_{\mu}a)) \quad (16)$$

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- In $a \rightarrow 0$ limit these extra masses are infinite so decouple
- Chiral symmetry is *explicitly* broken by the Wilson term
- Get order $\mathcal{O}(a)$ discretisation errors

- But we can do better!!!
- Let

$$Q_{\mu\nu} = U_{\mu,\nu}(n) + U_{\mu,-\nu}(n) + U_{-\mu,-\nu}(n) + U_{-\mu,\nu}(n) \quad (17)$$

Wilson Fermions

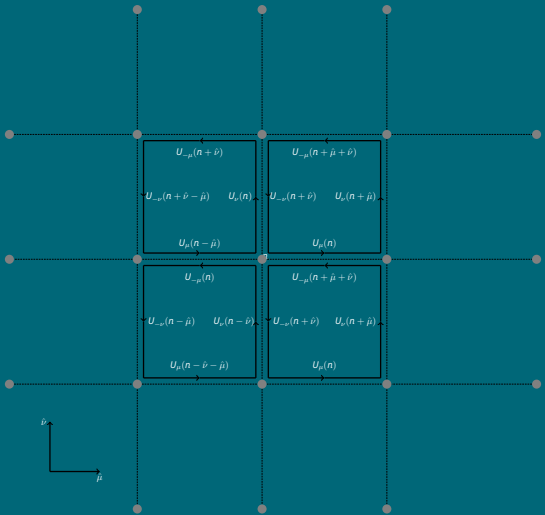


Figure: Clover action

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- Then the *Symanzik Improved* action of the form

$$S_I = S_{\text{Wilson}} + c_{\text{sw}} a^5 \sum_{n \in \Lambda} \sum_{\mu < \nu} \psi(\bar{n}) \frac{1}{2} \sigma_{\mu\nu} \left(\frac{i}{8a^2} (Q_{\mu\nu}(n) - Q_{\nu\mu}(n)) \right) \psi(n) \quad (18)$$

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- This is a work in progress

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 - Cheat?

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so can use HMC

- Theory is *qualitatively similar* showing QCD properties such as deconfinement and chiral symmetry breaking.

QC₂D

DISCLAIMER!!!!

- We are **NOT** claiming that the strong force actually follows $SU(2)$ instead of $SU(3)$.
We are merely using $SU(2)$ to probe an otherwise inaccessible régime using lattice techniques, much like running 2-d instead of 3-d simulations.
- The individual quarks are still fermionic.

Things to keep in mind

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- Max 10 flavours to have asymptotic freedom (vs up-to 16 in three-colour QCD)
- Baryons are now quark-quark pairs (diquarks), so follow Bose-Einstein Statistics³
- BEC phase appears at high density
- Now get 5 massless Goldstone bosons in the chiral limit.

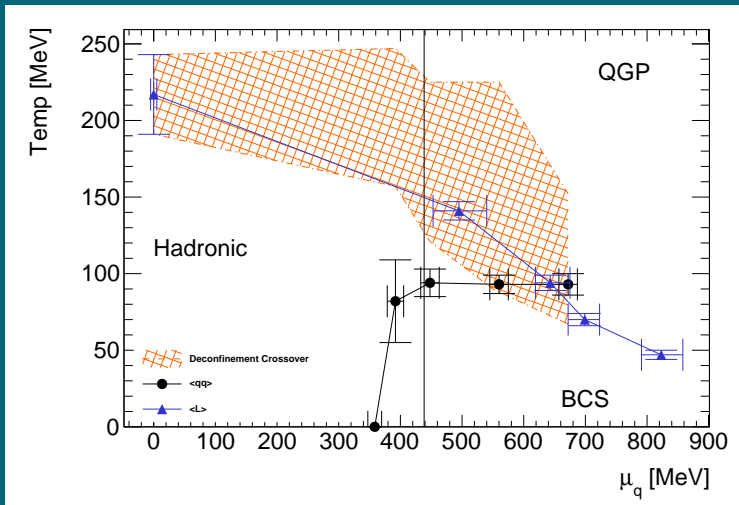
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- These results were first presented at XQCD 2022 and XV QCHS Conference⁴
- Using a spatial extent $N_s = 24$
- Conducted a temperature scan on coarse lattice at $a\mu = 0.400$
- Temperature is given by

$$T = \frac{1}{a_\tau N_\tau} \quad (19)$$

- Varying the number of sites along the time direction allows us to complete a temperature scan

⁴Lawlor et al. 2022.

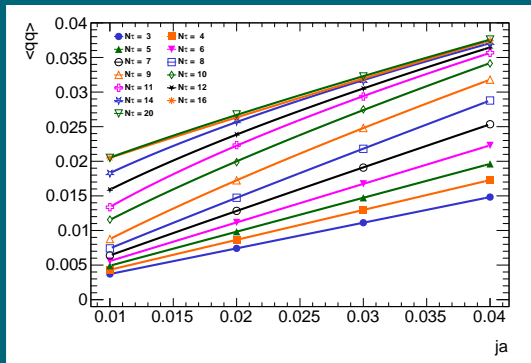


Phase diagram of QC₂D for $\frac{m_\pi}{m_\rho} = 0.80(1)$.

Superfluid Transition

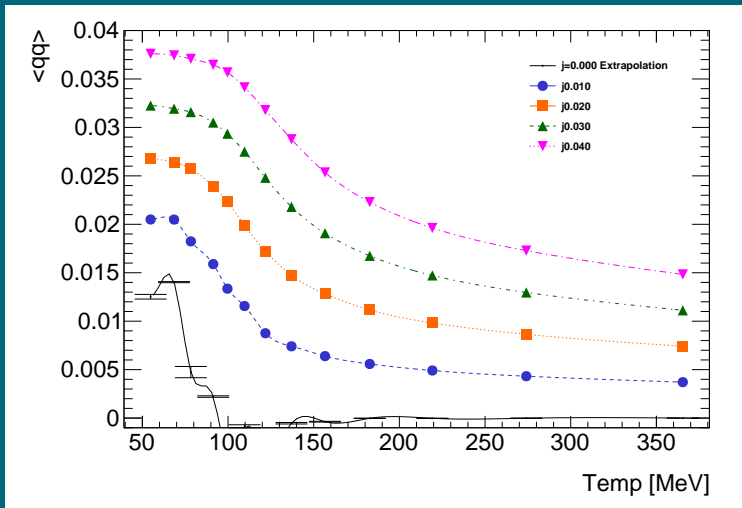
- Molecular dynamics requires inverting a large sparse matrix via conjugate gradient. ^a
- At non-zero baryon density, fermion matrix acquires non-zero density of very small eigenvalues, slowing down the computation
- Diquark source j lifts these eigenvalues, with "physical" results recovered by extrapolation of j to zero.

^ai.e. make supercomputer go brrr



Diquark condensate vs Diquark Source

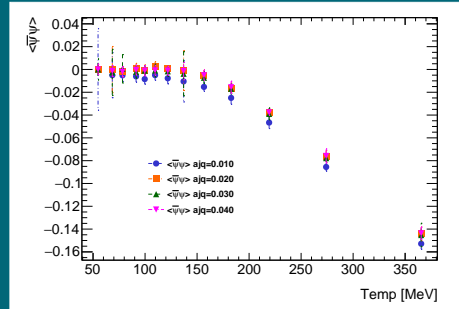
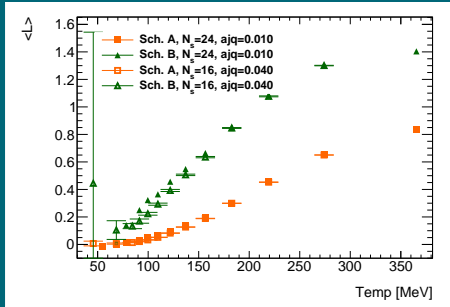
Superfluid Transition



Diquark condensate vs Temperature

- The superfluid phase transition occurs around $T \sim 100$ MeV.
- This indicates that the superfluid phase transition is indeed distinct from the deconfinement crossover.

Bosonic Observables



Renormalised Polyakov loop with two renormalisation schemes. These results are also compatible with earlier results on smaller volumes.

A: $L_R(N_\tau = 4, \mu = 0) = 0.5$

B: $L_R(N_\tau = 4, \mu = 0) = 1.0$

Unrenormalised but subtracted chiral condensate

The change in behaviour from constant to decreasing at $T \sim 150$ MeV suggests that the crossover coincides with the deconfinement crossover, not the superfluid transition.

Neutron Star (NS)

- Equation of State (EoS) of dense nuclear matter an unanswered question
- NS form natural labs for exploring dense nuclear matter
- Measurements of NS tidal deformabilities using Gravitational Wave (GW) indicates EoS soft at nuclear densities.
- Pulsar observations indicate stiff EoS at $M > 2M_{\odot}$. Non-monotonic C_s ?

	CEFT	Dense NM	Pert. QM	CFTs
c_s^2	$\ll 1$	$[0.8, 1]$	$\lesssim 1/3$	$1/3$

Table: C_s^2 predictions in four different limits.^a

^aAnnala et al. 2023.

Table: Radius constraints for neutron stars for $\simeq 1.4M_{\odot}$ and $\simeq 2.1M_{\odot}$ NSs. taken from Minamikawa et al. 2023

	radius [km]	mass [M_{\odot}]
GW170817 (primary)	$11.9^{+1.4}_{-1.4}$	$1.46^{+0.12}_{-0.10}$
GW170817 (second)	$11.9^{+1.4}_{-1.4}$	$1.27^{+0.09}_{-0.09}$
J0030+0451 (NICER)	$13.02^{+1.24}_{-1.06}$	$1.44^{+0.15}_{-0.14}$
J0030+0451 (NICER)	$12.71^{+1.14}_{-1.19}$	$1.34^{+0.15}_{-0.16}$
PSR J0740+6620 (NICER)	$12.35^{+0.75}_{-0.75}$	$2.08^{+0.07}_{-0.07}$
PSR J0740+6620 (NICER)	$12.39^{+1.30}_{-0.98}$	$2.08^{+0.07}_{-0.07}$

- Recent results (Kojo and Suenaga 2022) indicate QC₂D is applicable NS studies.
- First principles calculation of C₅² using QC₂D
- Coarse lattice parameters. $12^3 \times 24 \Rightarrow T \approx 50 \text{ MeV}$
- Work in progress. No diquark extrapolation

Pressure

$$C_s^2 = \frac{dP}{d\varepsilon} \quad (20)$$

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- Energy Density ε has fermionic and baryonic component

$$\varepsilon_g = \frac{3\beta}{N_c} \text{Re} (\text{Tr} U_{ij} - \text{Tr} U_{i0}) \quad (21)$$

$$\varepsilon_q = \kappa \langle \bar{\psi} D_0 \psi \rangle \quad (22)$$

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- Two different approaches for pressure

Pressure

Integral Method

$$P(\mu) = \int_0^{\mu} n_q(\mu') d\mu' \quad (23)$$

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- $n_q(\mu)$ has form of cubic polynomial
- Fit $n_q(\mu)$, extract coefficients and integrate analytically.
- Currently use polynomial interpolation, cubic spline on the cards.

Integral Method

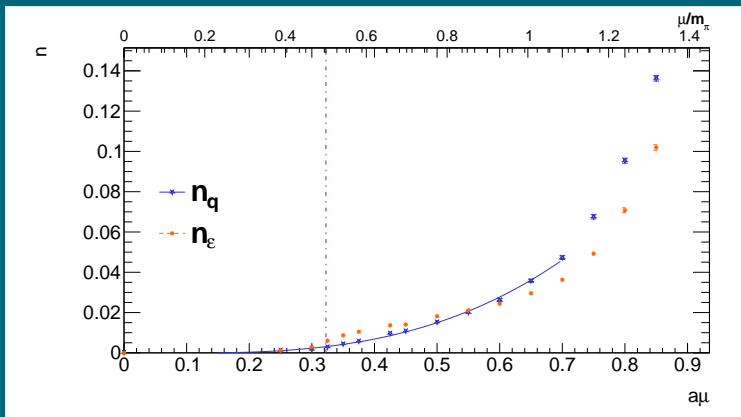


Figure: Quark and Energy Number Densities

$T_{\mu\mu}$

Derivative Method

Conformal Anomaly

$$T_{\mu\mu} = \varepsilon - 3p \tag{23}$$

$T_{\mu\mu}$

Derivative Method

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- Can find P from conformal anomaly

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Derivative Method

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- Can find P from conformal anomaly
- Calculate gluonic and fermionic components separately.

$$T_{\mu\mu}^g = -\frac{3\alpha}{N_c} \frac{\partial \beta}{\partial \alpha} \text{Re} (\text{Tr } U_{ij} + \text{Tr } U_{i0}) \quad (24)$$

$$T_{\mu\mu}^q = -\alpha \frac{\partial \kappa}{\partial \alpha} \kappa^{-1} (4N_f N_c - \langle \bar{\psi} \psi \rangle) \quad (25)$$

$T_{\mu\mu}$

Derivative Method

Conformal Anomaly

$$T_{\mu\mu} = \varepsilon - 3p \quad (23)$$

- Can find P from conformal anomaly
- Calculate gluonic and fermionic components separately.

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- Needs fewer configurations
- Karsch Coefficients need different lattice spacings.

$T_{\mu\mu}$

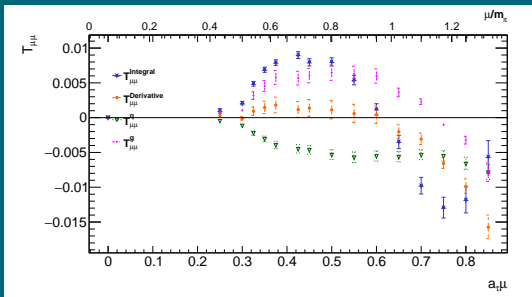


Figure: Unrenormalised Conformal Anomaly evaluated using the derivative and integral method, with the gluonic and fermionic components shown for the derivative method.

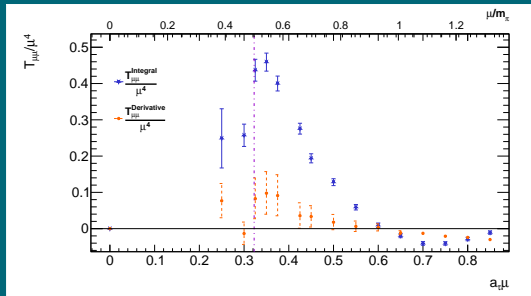


Figure: Renormalised Conformal Anomaly

c_s^2

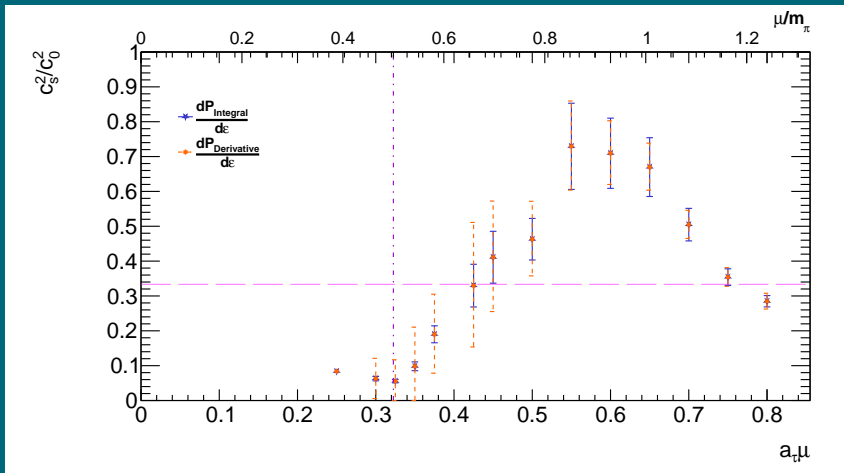


Figure: Speed of sound squared calculated using derivative and integral methods. Calculated using a $12^3 \times 24$ coarse lattice. Results are consistent with Iida and Ito 2022

Summary and bucket list

- The superfluid phase transition and deconfinement crossover are distinct
- Speed of sound exceeds the conformal limit. EoS stiff at intermediate densities?





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- The superfluid phase transition and deconfinement crossover are distinct
- Speed of sound exceeds the conformal limit. EoS stiff at intermediate densities?
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- Get the code to actually run correctly on GPUs
- EuroHPC benchmark access just awarded (LUMI-C/Vera).

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QC₂D Two Colour QCD

QCD Quantum Chromodynamics

HMC Hybrid Monte Carlo

RHIC Relativistic Heavy Ion Collisions

EoS Equation of State

NS Neutron Star

GW Gravitational Wave