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Xueli Luo*, Stefan Hanke, Katharina Battes, Christian Day Karlsruhe Institute of Technology, Institute for Technical Physics, 76021 Karlsruhe, Germany

The Einstein Telescope (ET) is a design concept for a thirdgeneration European gravitational wave detector that will be about 10 times more sensitive than today's instruments. It is designed as an equilateral triangle with 10 km long arms, 200 to 300 meters underneath the ground, and with detectors being located in each corner. Any two adjacent arms comprise two independent interferometers, of which one will detect lowfrequency gravitational wave signals (ET-LF), while the other will be optimized for operation at higher frequencies (ET-HF). Fig. 1 The proposed Einstein Telescope

Monte Carlo simulation of the thermal radiation heat load to the cryogenic mirror and vacuum

system of the Einstein Telescope

Conclusions and outlook

- Systematic Monte Carlo simulations are carried out to obtain the pressure profiles along the beamline of ET-LF, and a design configuration of the cyropumps of ET-LF (Fig. 6) is suggested.
- With this configuration of the cyropumps, the pressures in the tower and in the arm are decoupled and the requirements regarding to pressure and frost mitigation aspects can be fulfilled.
- Thermal radiation heat loads for the corresponding cryogenic circuits of the system at different temperatures are obtained with radiation exchange factor (REF) method.
- In this method, all 55 surfaces in the model are radiating and absorbing according their temperatures and emissivity, and the matrix (55 x 55) of REF is obtained from the simulation.
- **ProVac3D, a novel and versatile Test Particle Monte Carlo (TPMC) simulation code developed by ourselves, is used both in vacuum simulation and in heat load simulation.**
- To guaranty the simulation precision, at least 10¹² test particles are simulated with 2000 CPUs of a super computer in parallel.
- Ideally, the sum of the heat loads should be zero since the whole system is an enclosure. In simulation, the ratio of the heat loads to the sum of the thermal radiation of all surfaces in the system is less than 5×10^{-10} , which shows a very high simulation precision.
- Together with the assessment of the vacuum and the radiation heat load of ET-HF which is going on in a similar way, important information to the cryogenic plant can be provided.

In order to reduce seismic noise, thermal noise and other systematic noise, the beamline pipes require ultra-high or high vacuum conditions. In addition, the mirrors of ET-LF will be cooled to cryogenic temperatures below 20 K. Obviously, the mirror at this temperature will adsorb gases as a frost layer, degrading its optical properties and increasing laser power absorption.

In this presentation, systematic Monte Carlo simulations are used to (1) assist to develop the vacuum pumping systems, based on cryopumps; (2) obtain the radiation heat loads for the corresponding cryogenic supply and for the mirror.

gravitational wave observatory

- **Hydrogen pumping needed only on** the left side for the adjacent tower flow
- \blacksquare 1 m left H₂ pumping sufficient
- **Pressure around mirror well below** requirement
- Water frost formation rate drives pump design for 80 K
- One monolayer water ice deposition time is \sim 2 years
- **Tight thermal budget of mirror at** constant T

Water ice build-up ~2 years for 1 ML Fig. 6 Design configuration of the cryopumps of ET-LF (not in scale)

Expected: ε =0.5

Fig. 3 Outgassing behaviour of mild steel (left) measured in our OMA setup (right)

 Beam pipe vacuum is mainly dependent to the outgassing, and tower vacuum is almost decoupled from the beam pipe by cryogenic pumping.

View factor method Simulation quick Hard to solve a matrix equation sometimes (Errors from sparse matrix)

$\dot{Q}_i = -\sigma \varepsilon_i A_i T_i^4 + \sum_j^N \sigma \varepsilon_j A_j T_i^4 R(j \to i),$

where σ is Stefan-Boltzmann constant, the matrix R is the results of 55 independent simulations, and its component $R(j\rightarrow i)$ represents the absorption probability of the ith surface when the jth surface is the source. This equation is easy to understand: the first term is the energy emitted from one surface, and the second term is the energy this surface obtained, coming from all surfaces. Finally, one surface will emit energy when the sum value is negative, and will get energy (thermal radiation heat loads) when it is positive.

Table 2 Simulation results of the heat load of ET-LF

(1) 80 K pump unshielded, and the heat load to it will be ~50-60% due to a passive shielding between pump and beam pipe.

(2) Mirror of ɛ=0.01 (99% reflectivity), and the heat load to it is almost proportional to ɛ;

Water frost on the mirror:

(3) Factor 12 of the simulated value for entire ET-LF (4 regions per interferometer in a triangle configuration with 3 interferometers)

Note: Losses, thermal conductivity by supports and all details are not yet considered.