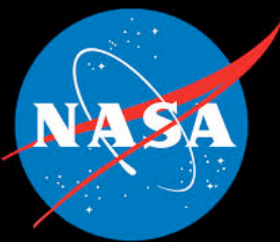


# The Swampland and Neutrino Physics



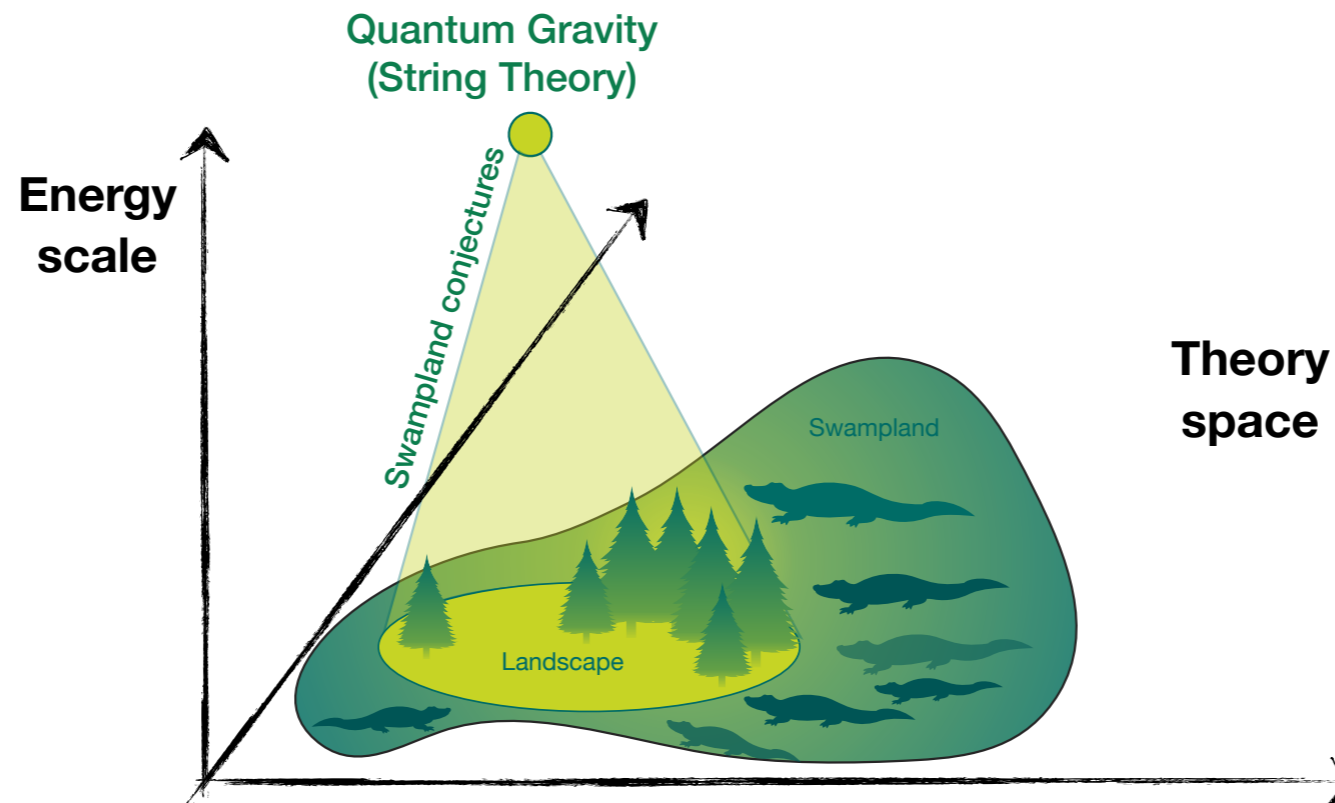
**Luis Anchordoqui**  
**CUNY**





# Swampland Program

- Not all effective field theories can consistently coupled to gravity
  - **anomaly cancellation is not sufficient**
  - **consistent ultraviolet completion can bring non-trivial constraints**
- “good” theories live in landscape and “bad” ones live in swampland



Vafa '05

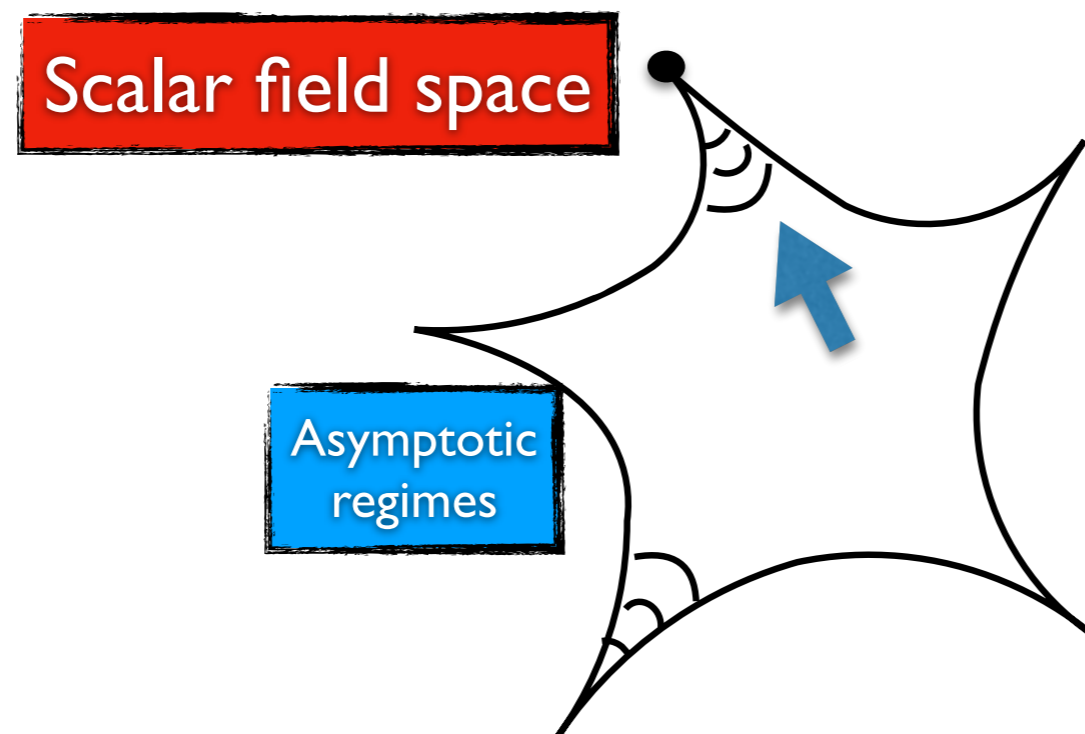
- In theory space ➡ frontier discerning good theories from those downgraded to swampland is drawn by family of conjectures classifying properties that theory must call for/avoid to enable consistent completion into QG
  - **criteria => conjectures**
  - supported by arguments based on string theory and black-hole physics**
- Swampland conjectures provide bridge from quantum gravity to astrophysics, cosmology, and particle physics

# Where Do We Live in String Landscape?

## Distance Conjecture

- At large distance in field space  $\phi \Rightarrow$  tower of exponentially light states  
 $m \sim e^{-\alpha\phi}$  with  $\alpha \sim \mathcal{O}(1)$  parameter in Planck units Ooguri-Vafa '06
- Kaluza-Klein tower  $\blacktriangleright$  decompactification of  $d$  extra dimensions  
 $\phi = \ln R \blacktriangleright m \sim 1/R$
- Species scale where gravity becomes strong  $\blacktriangleright \Lambda_{\text{QG}} = (m^d M_p^2)^{1/(d+2)}$

Dvali '07



- Smallness of some physical parameters might signal a large distance corner in the string landscape of vacua
- Such parameters can be  $\blacktriangleright$  scales of dark energy and neutrino masses

# Dark Dimension Proposal for Dark Energy

- AdS distance conjecture  $\phi = -\ln |\Lambda|$  Lüst-Palti-Vafa '19
- Extension to dS  $\blacktriangleright m \sim \frac{1}{\lambda} \left( \frac{\Lambda}{M_p^4} \right)^a M_p \Rightarrow 1/4 \leq a \leq 1/2$  Montero-Vafa-Valenzuela '22
- $a \leq 1/2$  unitarity bound  $m_{\text{spin}-2}^2 \geq 2H^2 \sim \Lambda$  Higuchi '87
- $a \geq 1/4$  estimate of 1-loop contribution  $\Lambda \gtrsim m^4$

## Observations $\blacktriangleright$

$$\Lambda \sim 10^{-120} M_p^4 \text{ (No deviations from Newton's law } R < 30 \mu\text{m)} \Rightarrow a = 1/4$$

Consistent with string computations LAA-Antoniadis-Lüst-Lüst '23

astrophysical constraints  $\Rightarrow d = 1$  extra dimension

$$\Rightarrow \text{species scale (5d Planck mass)} \blacktriangleright \Lambda_{\text{QG}} \sim m^{1/3} M_p^{2/3} \sim 10^9 \text{ GeV}$$

$$10^{-4} \lesssim \lambda \lesssim 10^{-2}$$



# More Physics Implications of Dark Dimension

- Neutrino masses originate in 5D bulk brane interactions of the form

$$\mathcal{L} \supset h_{ij} \bar{L}_i \tilde{H} \Psi_j(y=0)$$

lepton doublets  
(localized on SM brane)

coupling constants

$\tilde{H} = -i\sigma_2 H^*$

3 bulk R-neutrinos

(evaluated at position of SM brane  $y = 0$  in 5<sup>th</sup>-dimension coordinate  $y$ )

- Expanding  $\Psi_j$  into modes canonically normalized leads to Yukawa 3x3 matrix suppressed by square root of bulk volume

$$Y_{ij} = \frac{h_{ij}}{\sqrt{\pi R M_s}} \sim h_{ij} \frac{M_s}{M_p} \quad M_s \lesssim \Lambda_{\text{QG}}$$

- Mesoscopic extra dimension produces suppression of 4-dimensional Yukawa couplings ➔ yielding naturally light Dirac neutrinos

- Recent analysis of  $\nu$ -oscillation data with 3 bulk neutrinos  $\Rightarrow$

$$m \gtrsim 2.5 \text{ eV} \quad \vee \quad R \lesssim 0.4 \mu\text{m} \Rightarrow \lambda \lesssim 10^{-3} \quad \wedge \quad \Lambda_{\text{QG}} \sim 10^9 \text{ GeV}$$

Forero-Giunti-Ternes-Tyagi '22

- Bound can be relaxed in presence of bulk  $\nu_R$ -neutrino masses

LAA-Antoniadis-Cunat '23

# Dark Energy, Gravitino, and KK Towers

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➤ Gravitino conjecture

Cribiori-Lüst-Scalisi '21

Castellano-Font-Herraez-Ibanez '21

$$\tilde{m} = \frac{1}{\lambda_{3/2}} \left( \frac{m_{3/2}}{M_p} \right)^n M_p \quad n > 0$$

➤ Combine gravitino conjecture with dark dimension proposal ➡ two possibilities

➤ One KK Tower

$$\tilde{m} = m \Rightarrow m_{3/2} \sim 1 \text{ eV}$$

➤ Two KK towers ➡ one for dark energy and one for SUSY breaking

$$\tilde{m} \neq m \Rightarrow m_{3/2} \sim 250 \text{ TeV}$$

LAA-Antoniadis-Cribiori-Lust-Scalisi '23



# Dark Dimension Hierarchy from Inflation

➤ Interesting possibility ➡ extra dimension expands with time

$R_0 \sim 1/M_s$  to  $R \sim \mu\text{m}$  requires  $\sim 42$  e-folds! LAA-Antoniadis-Lüst '22

$$ds_5^2 = a_5^2(-d\tau^2 + d\vec{x}^2 + R_0^2 dy^2) \quad R_0 \text{ ➡ initial size prior to inflation}$$

$$= \frac{ds_4^2}{R} + R^2 dy^2 \quad ; \quad ds_4^2 = a^2(-d\tau^2 + d\vec{x}^2) \quad \Rightarrow \quad a^2 = R^3$$

After 5d inflation of  $N = 42$  e-folds  $\Rightarrow$  63 e-folds in 4d with  $a = e^{3N/2}$

➤ Dark Dimension from 5D inflation

- connect weakness of gravity to size of observable universe
- scale invariant density fluctuations from 5D inflation
- radion stabilization

# Neutrino-Modulino Mixing

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➤ Modulino  $\mathcal{S}$  ➔ fermionic partner of radion

➤ Simple construct ➔ relevant light scale for SM singlets is  $m_{3/2}$

and dimensionless coupling constant with visible matter is  $\lambda_i = \alpha_i \frac{m_{3/2}}{M_p}$

➤ Modulino mass is generated by SUSY breaking as the coupling  $\lambda_i$

but @ 2nd order in  $m_{3/2}$  ➔  $m_4 \sim \beta \frac{m_{3/2}^2}{M_p}$

$$m_{3/2} \sim 250 \text{ TeV} \wedge \beta \sim 2 \Rightarrow m_4 \sim 50 \text{ eV}$$

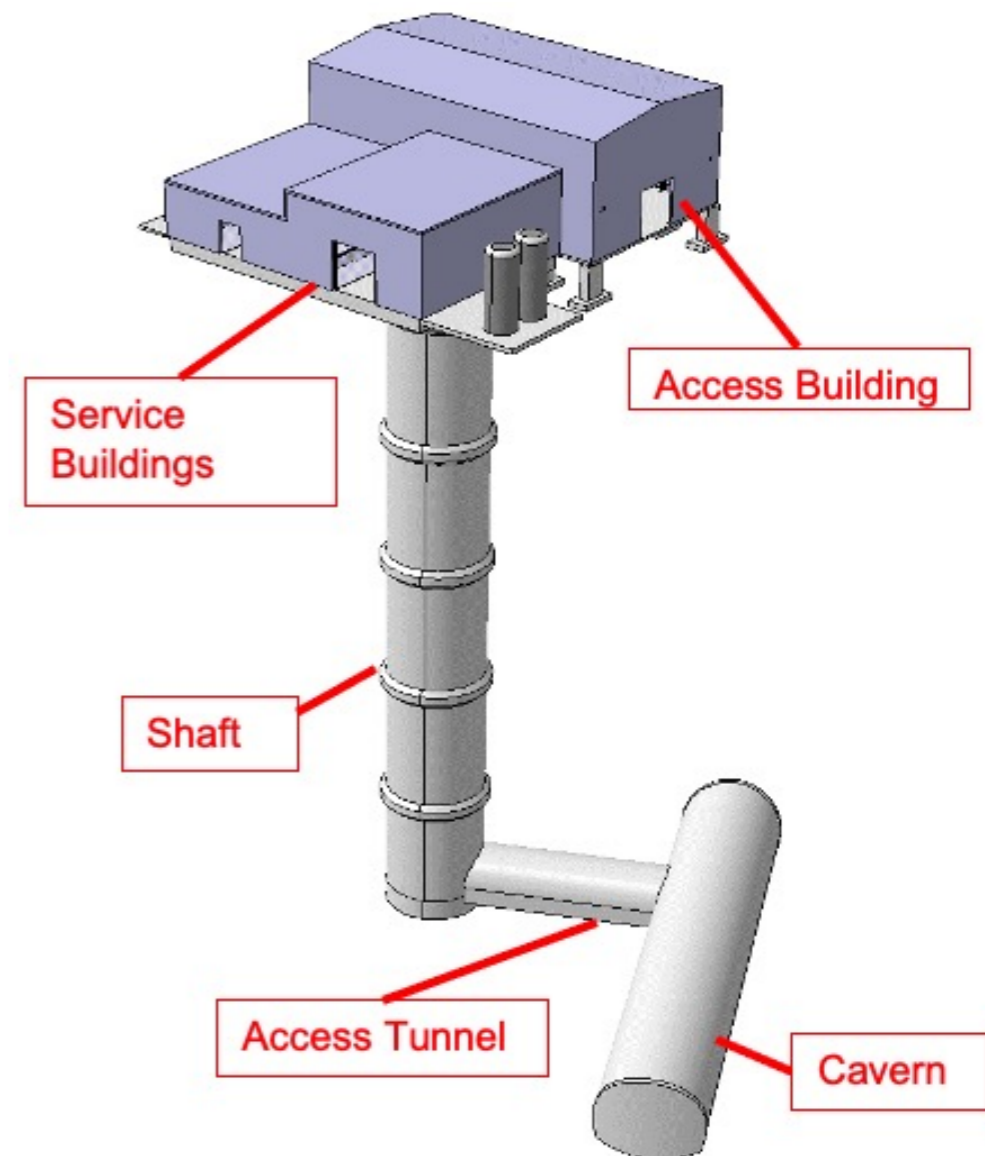
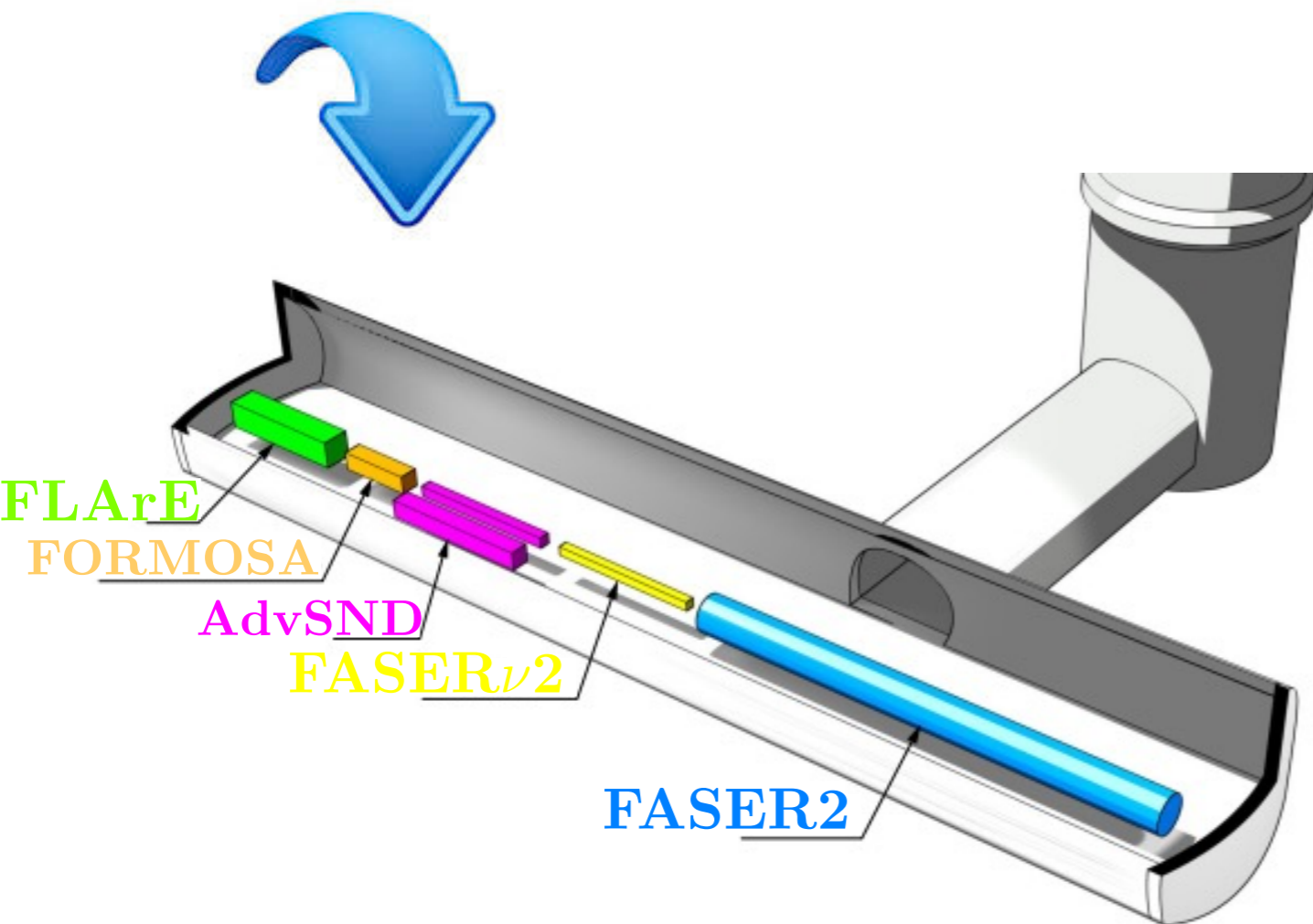
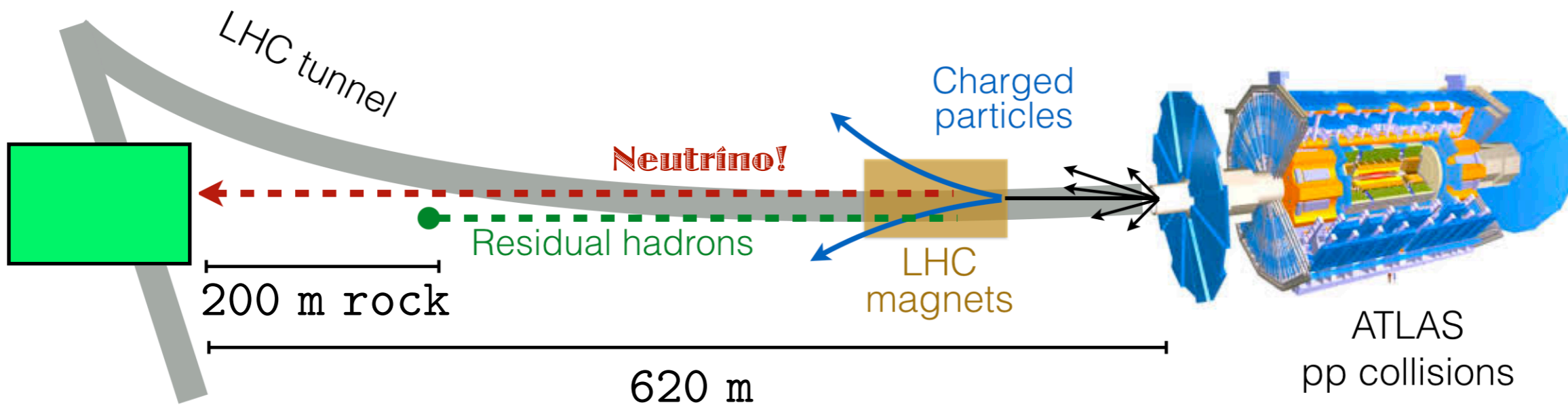
➤  $\mathcal{S}$  mixing with active SM neutrinos involves the electroweak symmetry breaking

➤ Simplest effective operator  $\lambda_i \bar{L}_i \mathcal{S} H$  ➔  $m_{\nu \mathcal{S}} \sim m_{3/2} \langle H \rangle / M_p \ll m_4$

analog to 3 + 1 scheme



# Forward Physics Facility

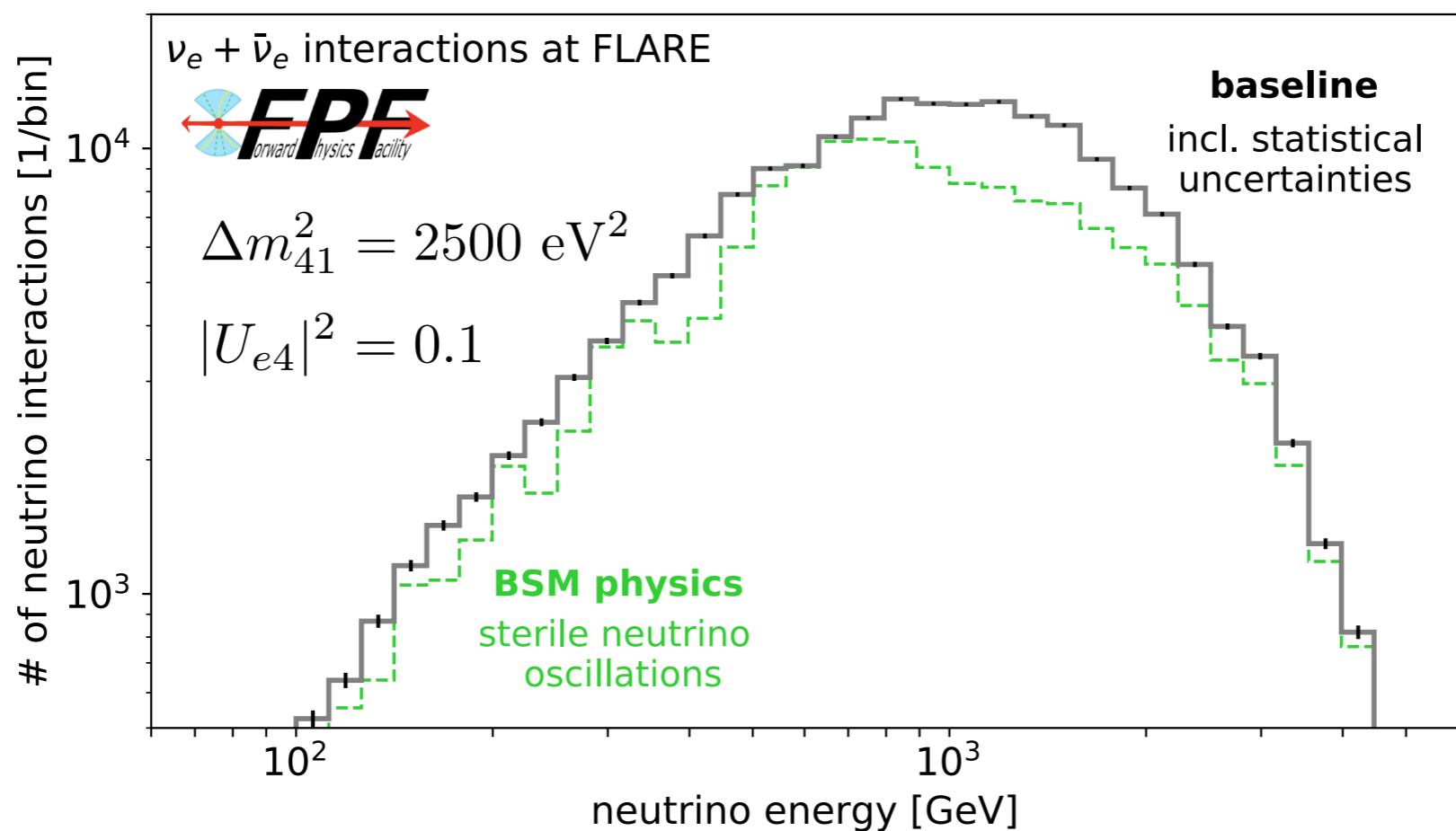


# Neutrino-Modulino Oscillations

- Survival probability ➡ 3 + 1 scheme

$$P(\nu_\alpha \rightarrow \nu_\alpha) = 1 - 4|U_{\alpha 4}|^2(1 - |U_{\alpha 4}|^2) \sin^2 \frac{\Delta m_{41}^2 L}{4E}$$

- FLArE sensitive to neutrino square mass difference satisfying  $\frac{\Delta m_{41}^2 L}{4E} = \frac{\pi}{2}$   
 for  $L \sim 620 \text{ m} \wedge E \sim \text{TeV} \Rightarrow \Delta m_{41}^2 \sim 2500 \text{ eV}^2$



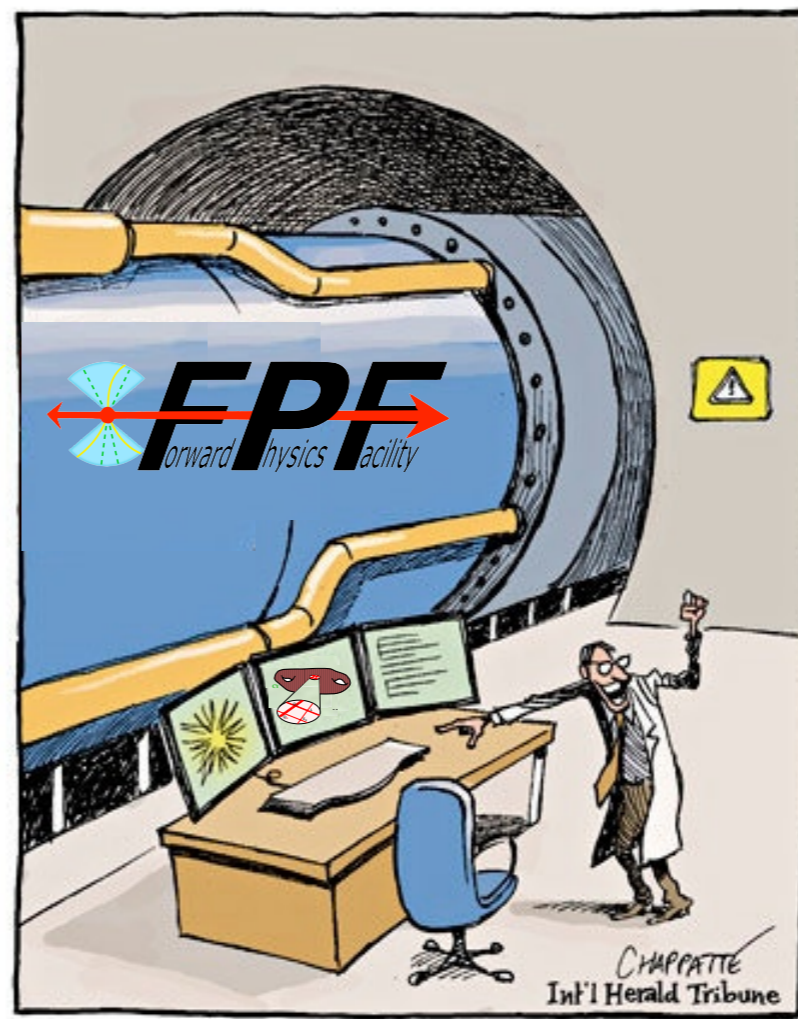
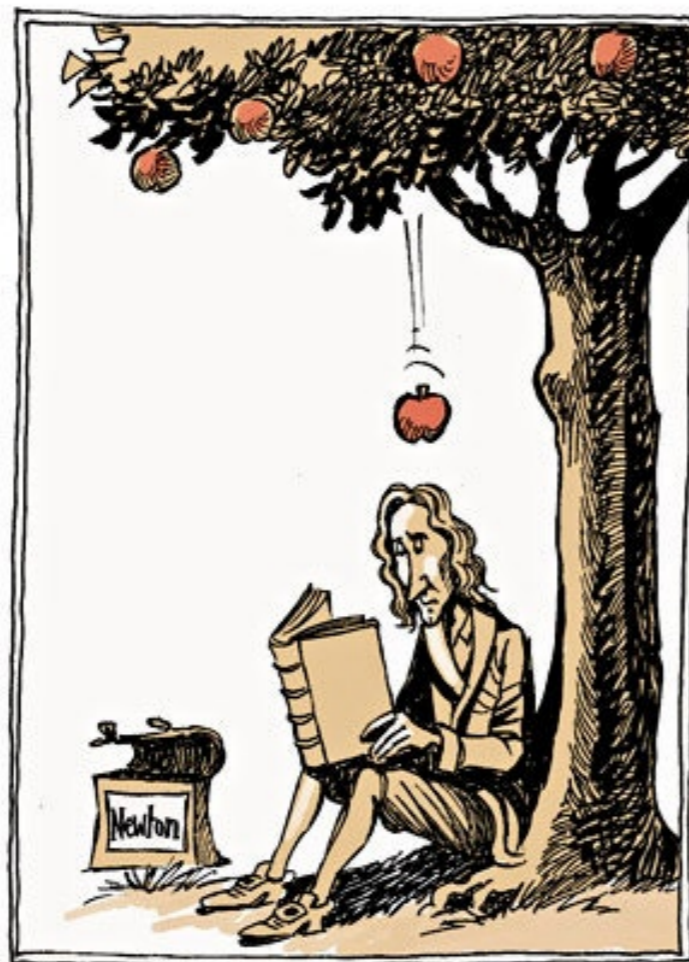
Study of FPF sensitivity to  $U_{\alpha 4}$  is underway ➡ Kling-Trojanowski-Makela to appear



# Take Home Message

- FPF experiments will be able to probe models with high scale SUSY breaking by searching for neutrino-modulino oscillations
- This highlights complementarity between ATLAS/CMS and FPF experiments in exploring the SUSY parameter space

## Collisions That Changed The World







*That's all Folks!*