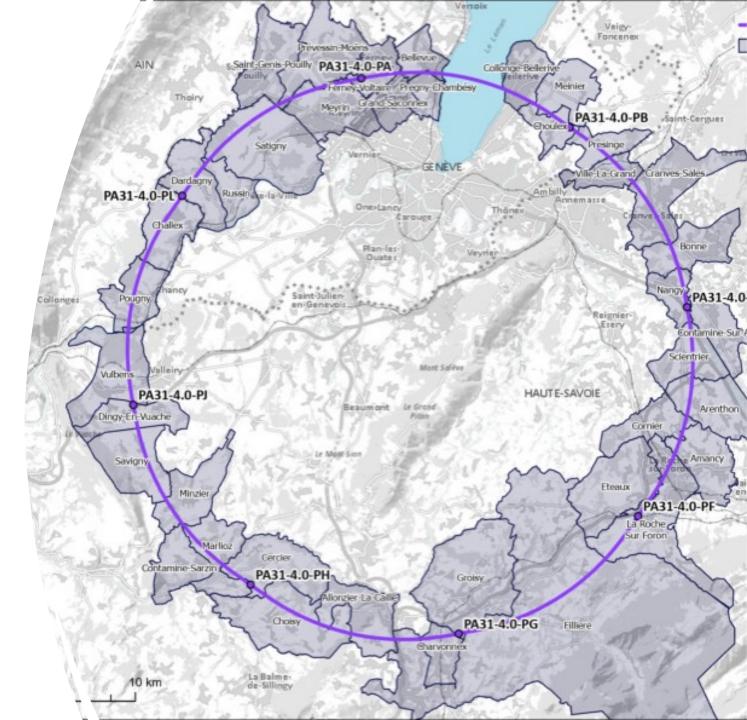
# BSM in FCC

*Zeynep Demiragli* (*Boston University*)

FCC Week 2024, San Francisco



Zeynep Demiragli – FCC Week 2024, San Francisco

# Where the wild things are?



Illustrator: Maurice Sendak

FCC-ee opportunity: Precision and Exploration

- > Explore indirectly using precision measurements.
- > Explore directly & discover: ALPs, dark photons, HNLs,
  - exotic Higgs any of which could be feebly interacting particles

# Where the wild things are?





Illustrator: Maurice Sendak

FCC-ee opportunity: Precision and Exploration

> Explore indirectly using precision measurements.

> Explore directly & discover: ALPs, dark photons, HNLs,

exotic Higgs any of which could be feebly interacting particles

# Where the wild things are?





Illustrator: Maurice Sendak

FCC-ee opportunity: Precision and Exploration

> Explore indirectly using precision measurements.

> Explore directly & discover: ALPs, dark photons, HNLs,

exotic Higgs any of which could be feebly interacting particles

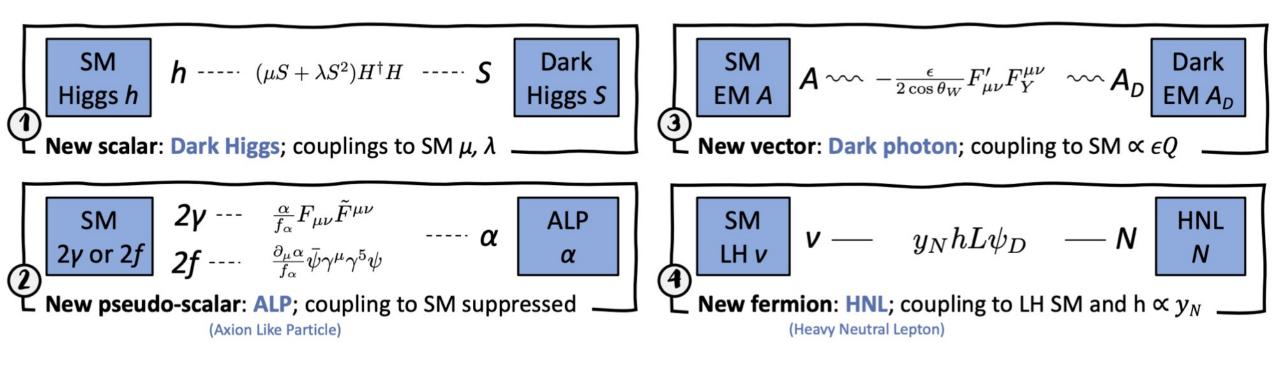
Working point	Z, years 1-2	Z, later	WW, years 1-2	WW, later	ZH	$t\overline{t}$	
$\sqrt{s} \; (\text{GeV})$	88, 91, 94		157, 163		240	340 - 350	365
Lumi/IP $(10^{34}  \mathrm{cm}^{-2} \mathrm{s}^{-1})$	70	140	10	20	5.0	0.75	1.20
Lumi/year $(ab^{-1})$	34	68	4.8	9.6	2.4	0.36	0.58
Run time (year)	2	2	2	0	3	1	4
					$1.45 \times 10^6 \mathrm{ZH}$	$1.9 \times 10^{-1}$	$b^{6} t \overline{t}$
Number of events	$6 \times 10^{1}$	$^{2}$ Z	$2.4 imes10^8\mathrm{WW}$		+	+330k	ZH
					45k WW $\rightarrow$ H	$+80 \mathrm{kWW}$	$/ \rightarrow H$

### Personal Favorite Motivator: Dark Sector



While the dynamics of the dark sector could be complicated...

to observe a dark sector, we need a portal interaction:



New Physics could be light and feebly interacting with SM

# Personal Favorite Motivator: Dark Sector



Feebly interacting Particles: From Energy Frontier to Intensity Frontier

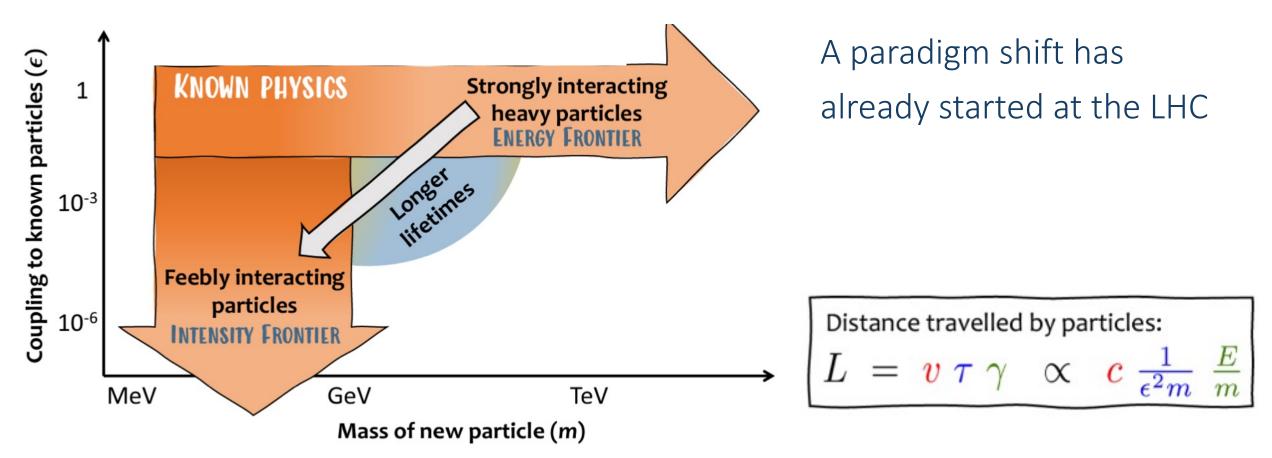


Diagram credit to A. Sfyrla

# Personal Favorite Motivator: Dark Sector



Feebly interacting Particles: From Energy Frontier to Intensity Frontier

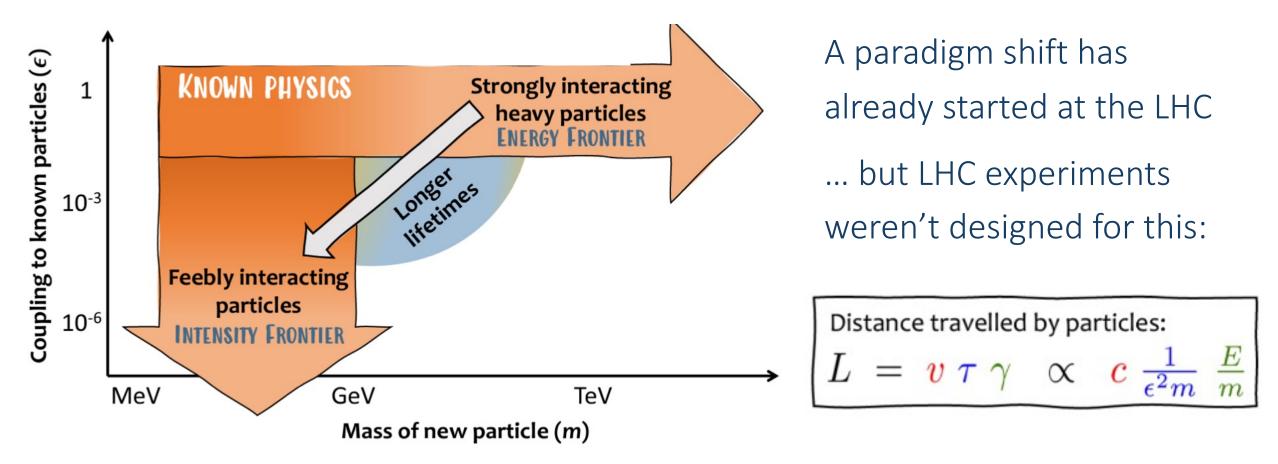


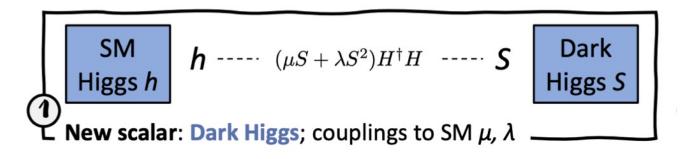
Diagram credit to A. Sfyrla

See L. Skinnari's talk for detector requirements!

# Scalar Portal: Dark Higgs, BSM Higgs



8

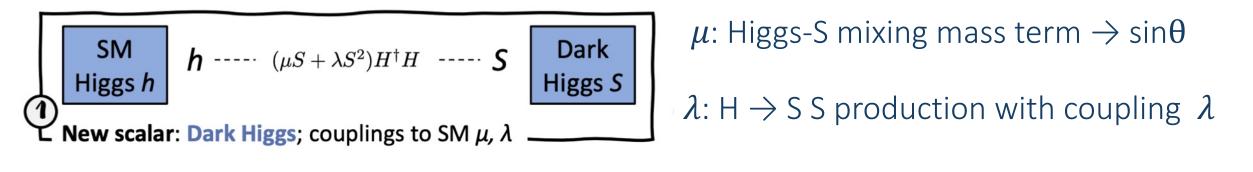


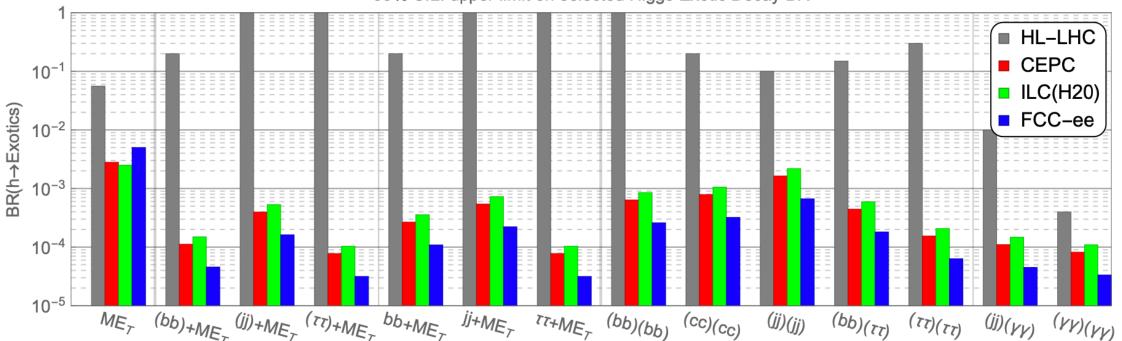
 $\mu$ : Higgs-S mixing mass term  $\rightarrow \sin \theta$ 

The mixing implies that the SM couplings of the Higgs deviate from their SM values

# Scalar Portal: Dark Higgs, BSM Higgs







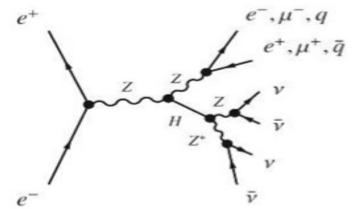
95% C.L. upper limit on selected Higgs Exotic Decay BR

Zeynep Demiragli – FCC Week 2024, San Francisco

https://arxiv.org/pdf/1612.09284

In SM H  $\rightarrow$  ZZ  $\rightarrow$  (vv<sup>-</sup>)(vv<sup>-</sup>) process has a BR of about 10<sup>-3</sup>

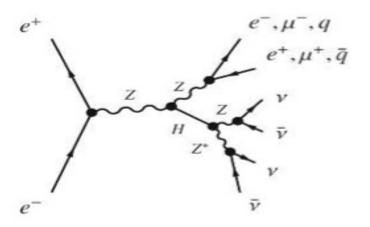
- ≻Current ATLAS/CMS sensitivity: ~10-15% BR observed
- ➢Inv BR would be significantly enhanced in presence of New Physics



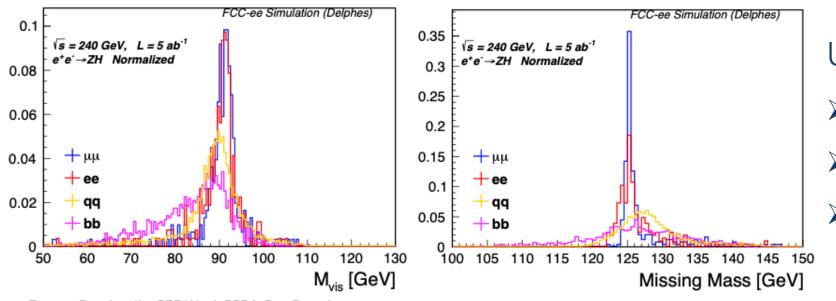
In SM H  $\rightarrow$  ZZ  $\rightarrow$  (vv<sup>-</sup>)(vv<sup>-</sup>) process has a BR of about 10<sup>-3</sup>

≻Current ATLAS/CMS sensitivity: ~10-15% BR observed

≻Inv BR would be significantly enhanced in presence of New Physics



Signal ( $H \rightarrow$ inv )	Energy	Luminosity	Channels	Backgrounds
ZH	240 GeV	5 ab <sup>-1</sup>	ee,μμ,bb,qq	Z, ZZ, ZH, WW



Using Delphes simulation

➢ qq resolution worse than ee/mm

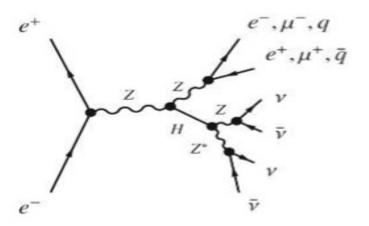
 $> M_{miss}$  resolution best for mm

Mass resolution for bb worst due to neutrinos in b hadron decays

In SM H  $\rightarrow$  ZZ  $\rightarrow$  (vv<sup>-</sup>)(vv<sup>-</sup>) process has a BR of about 10<sup>-3</sup>

≻Current ATLAS/CMS sensitivity: ~10-15% BR observed

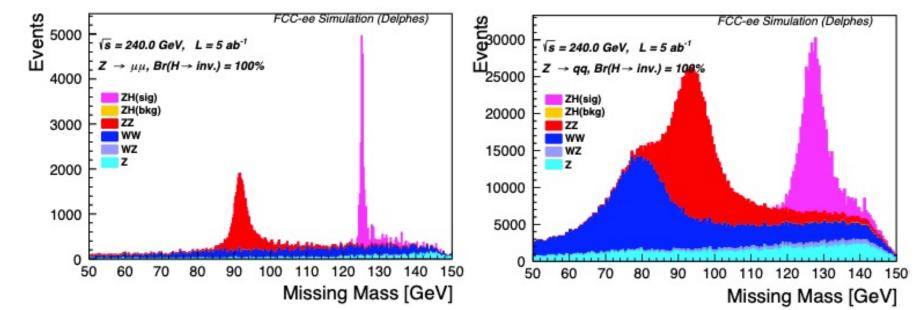
≻Inv BR would be significantly enhanced in presence of New Physics



Signal ( H $ ightarrow$ inv )	Energy	Luminosity	Channels	Backgrounds
ZH	240 GeV	5 ab <sup>-1</sup>	ee,μμ,bb,qq	Z, ZZ, ZH, WW

Additional requirements for Z(bb) channel, to cope with worse resolution of bb system

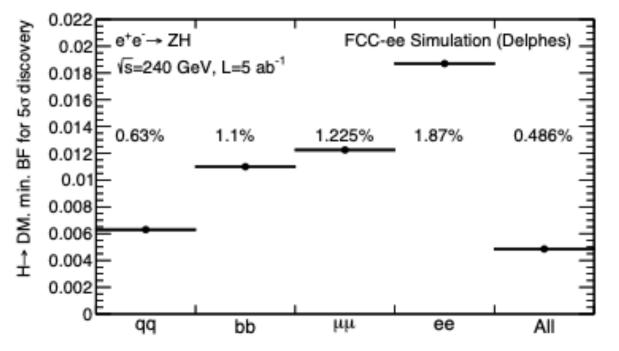
Use distribution of  $M_{\rm miss}$  in likelihood fit.



Zeynep Demiragli – FCC Week 2024, San Francisco

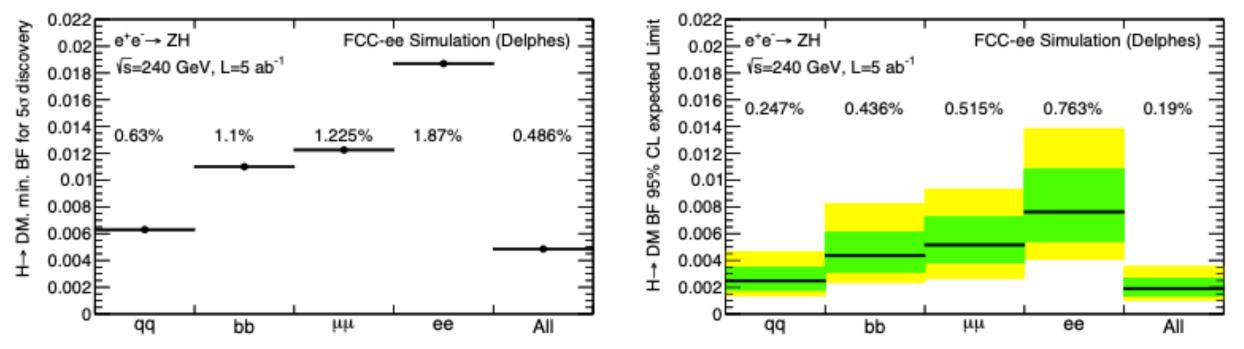
#### Reach SM precision of $\simeq 0.1\%$

Could discover NP with H  $\rightarrow$  inv above SM background with BF ~ 0.5%



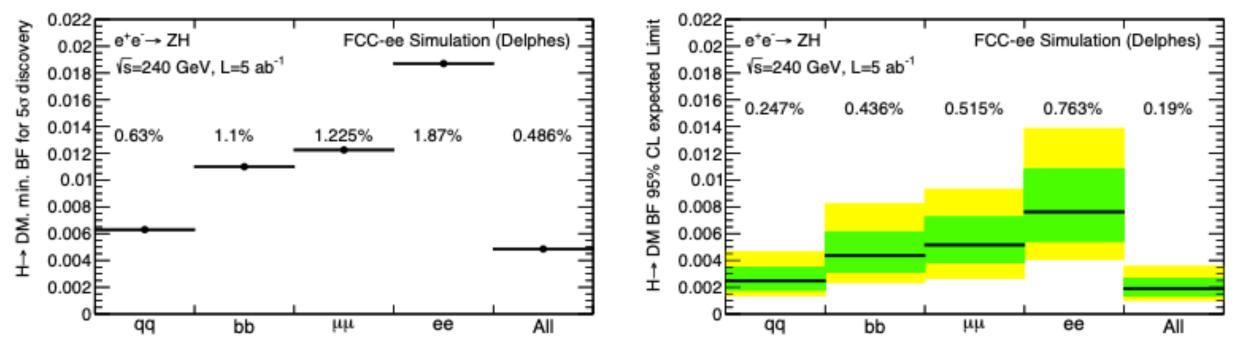
#### Reach SM precision of ≈ 0.1%

Could discover NP with H  $\rightarrow$  inv above SM background with BF ~ 0.5%



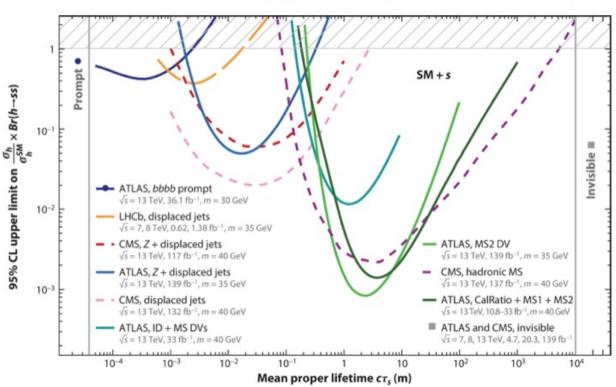
#### Reach SM precision of ≈ 0.1%

Could discover NP with H  $\rightarrow$  inv above SM background with BF ~ 0.5%



Recent work (FCC MIT Workshop) compares CLD full sim and CLD & IDEA Delphes fast sim.

- Efficiency is ~identical for IDEA and CLD fast simulations!
- Electron eff is worse for full sim than for fast sim & Muon eff is very similar for full & fast sim.



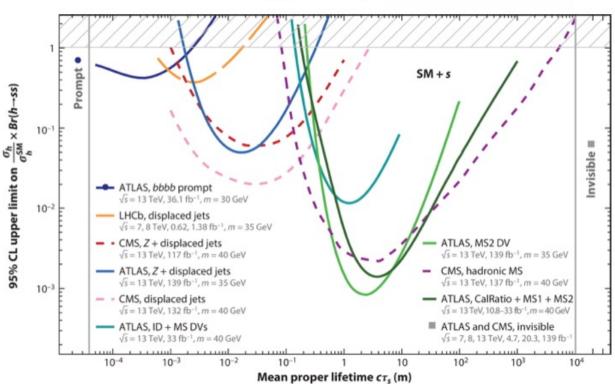
Review: Exotic Higgs Decays arXiv:2111.12751

For sufficiently small mixing, the scalar can be long-lived:  $c\tau \sim$  meters if  $\theta < 1e-6$ 

 $ZH \rightarrow Z (ee/mm) + H \rightarrow s s$ 

BOSTO

UNIVERSITY



Review: Exotic Higgs Decays arXiv:2111.12751

For sufficiently small mixing, the scalar can be long-lived:  $c\tau \sim$  meters if  $\theta < 1e-6$ 

$$ZH \rightarrow Z (ee/mm) + H \rightarrow s s \rightarrow 4b$$

#### Full chain:

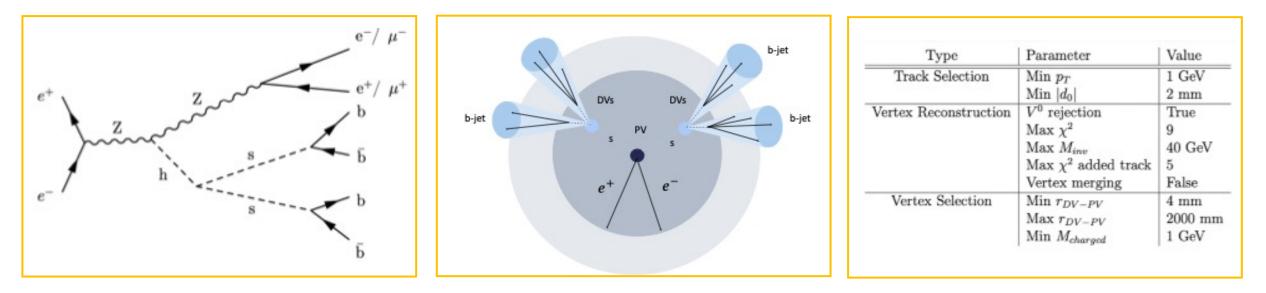
MadGraph v3.4.1 (for parton level) + Pythia8 (parton shower / hadronization) + Delphes ( winter2023 IDEA)

Signal ( $H \rightarrow$ inv )	Energy	Luminosity	Channels	Backgrounds
ZH	240 GeV	5 ab <sup>-1</sup>	4b	WW, ZZ, ZH

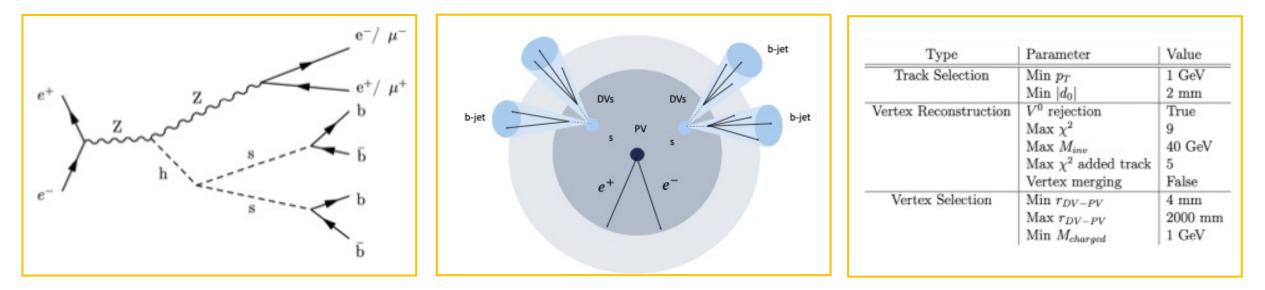
BOST

UNIVERSIT

#### Pre-selection: 2 opposite sign, same flavor leptons (invariant mass within Z) + 2 displaced vertex



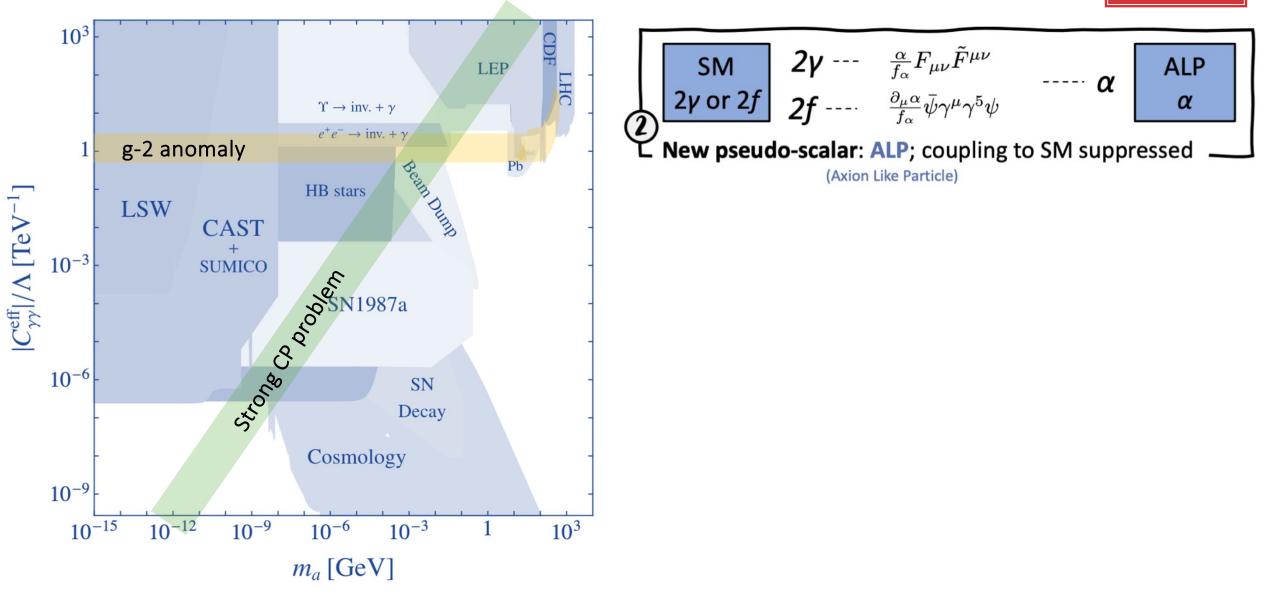
#### Pre-selection: 2 opposite sign, same flavor leptons (invariant mass within Z) + 2 displaced vertex



#### Zero background search => 3 events is 95% CL exclusion

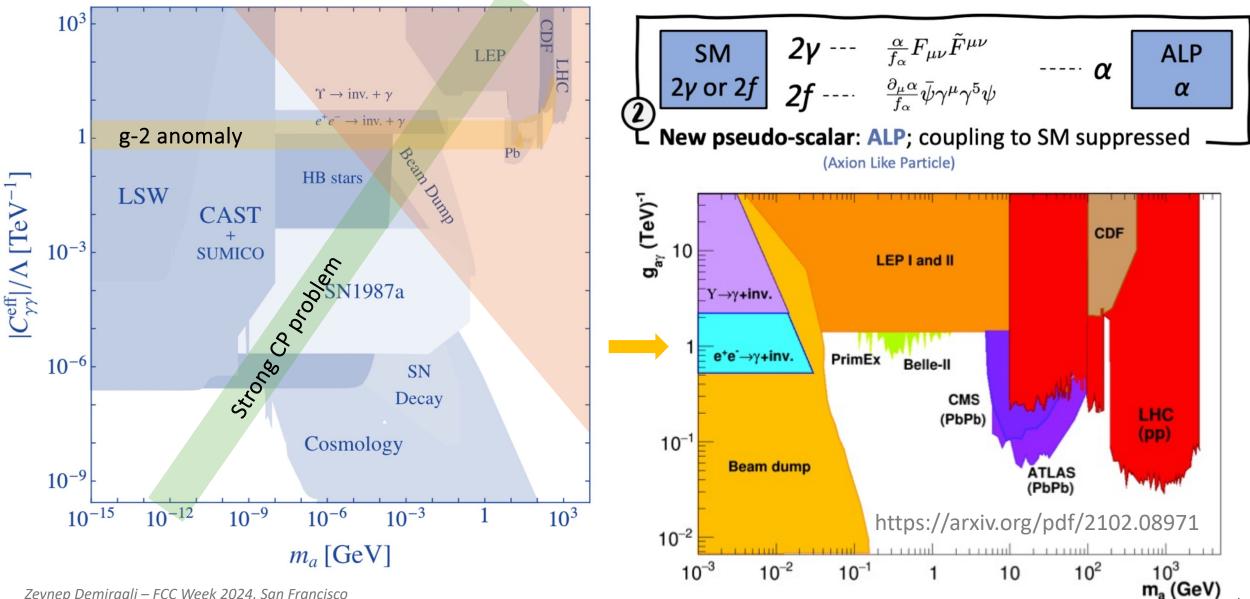
$m_s, \sin  heta \mid c au \; [ m mm]$	${ m BR}(h  o ss)$	Before selection	Pre-selection	$70 < m_{ll} < 110~{\rm GeV}$	$n\_DVs \geq 2$
20 GeV, 1e-5         3.4           20 GeV, 1e-6         341.7           20 GeV, 1e-7         34167.0           60 GeV, 1e-5         0.9           60 GeV, 1e-6         87.7           60 GeV, 1e-7         8769.1	$\begin{array}{c} 6.98 \times 10^{-4} \\ 6.98 \times 10^{-4} \\ 6.98 \times 10^{-4} \\ 2.06 \times 10^{-4} \\ 2.06 \times 10^{-4} \\ 2.06 \times 10^{-4} \\ 2.06 \times 10^{-4} \end{array}$	$\begin{array}{c} 55.2 \pm 0.552 \\ 55.2 \pm 0.552 \\ 55.2 \pm 0.552 \\ 16.32 \pm 0.163 \\ 16.32 \pm 0.163 \\ 16.32 \pm 0.163 \end{array}$	$52.84 \pm 0.538 \\52.44 \pm 0.538 \\52.38 \pm 0.540 \\15.62 \pm 0.127 \\15.62 \pm 0.196 \\15.52 \pm 0.159$	$\begin{array}{c} 49.02 \pm 0.520 \\ 49.02 \pm 0.521 \\ 49.68 \pm 0.524 \\ 14.59 \pm 0.154 \\ 14.61 \pm 0.196 \\ 14.62 \pm 0.155 \end{array}$	$5.0 \pm 0.166$ $37.1 \pm 0.453$ $0.8 \pm 0.067$ $0.0033 \pm 0.0023$ $10.96 \pm 0.167$ $6.49 \pm 0.103$

## Pseudo-scalar Portal: <u>Axion-like Particles</u>



JNIVERSIT

## Pseudo-scalar Portal: Axion-like Particles



BOSTO

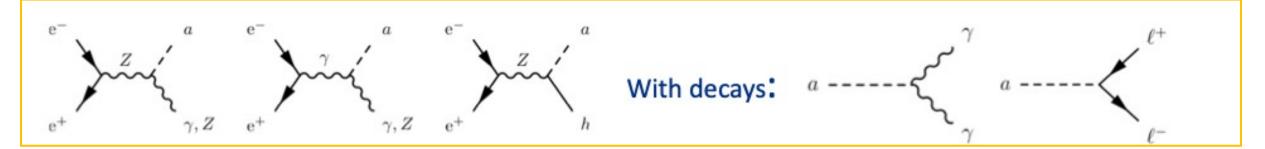
JNIVERSIT

Zeynep Demiragli – FCC Week 2024, San Francisco

# Case Study: <u>Axion-like Particles</u>



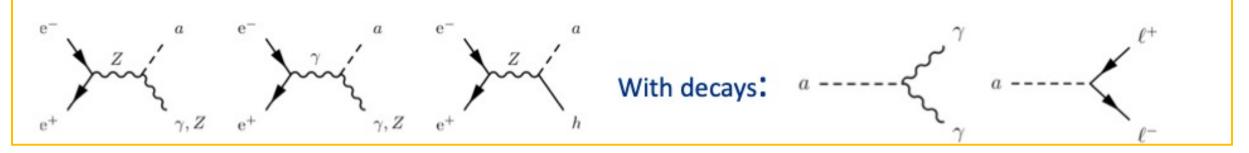
#### "Standard" approach:



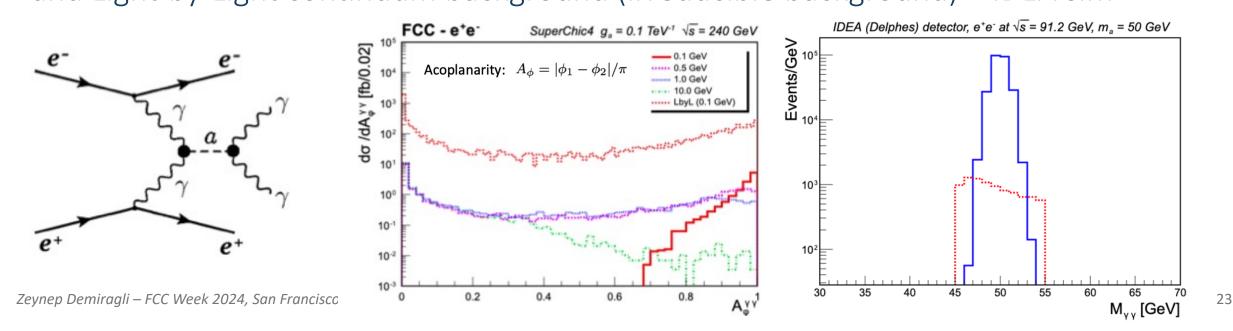
# Case Study: <u>Axion-like Particles</u>



#### "Standard" approach:

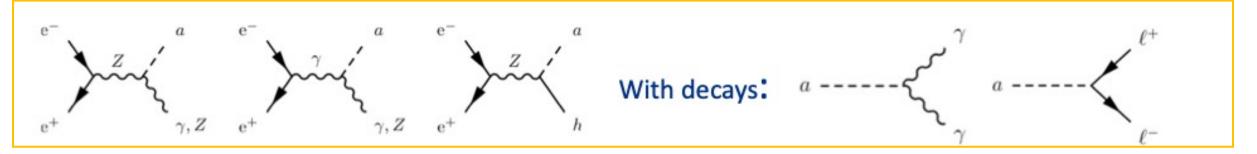


<u>ALP production via photon-photon fusion:</u> Using SC4 MC generator for the ALP signal and Light-by-Light continuum background (irreducible background) + IDEA sim



# Case Study: <u>Axion-like Particles</u>

"Standard" approach:



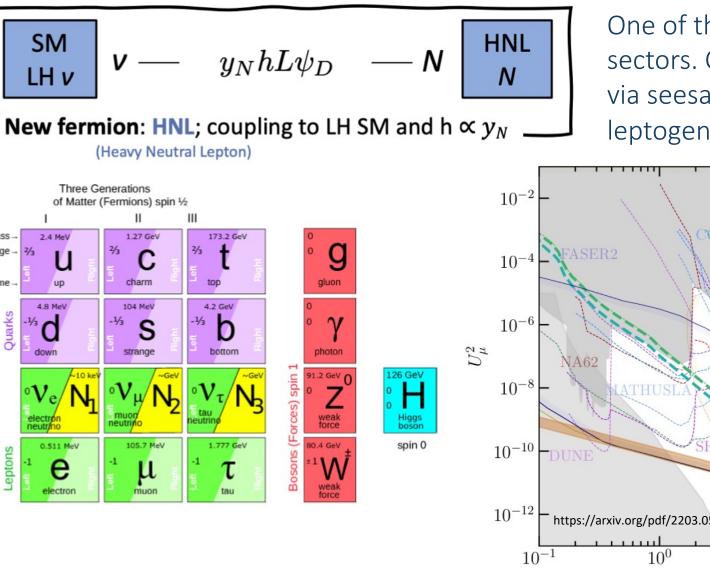
ightarrow γγ → a extends current LHC limits for m<sub>a</sub> = 5 - 350GeV by 2(O) magnitude

 $rightarrow e^+ e^- 
ightarrow Z 
ightarrow \gamma a$  extends current LHC limits for  $m_a$  = 0.1 − 90 GeV by 3(0) magnitude

For low ALP mass, sophisticated detectors & techniques are needed to isolate the overlapping photons

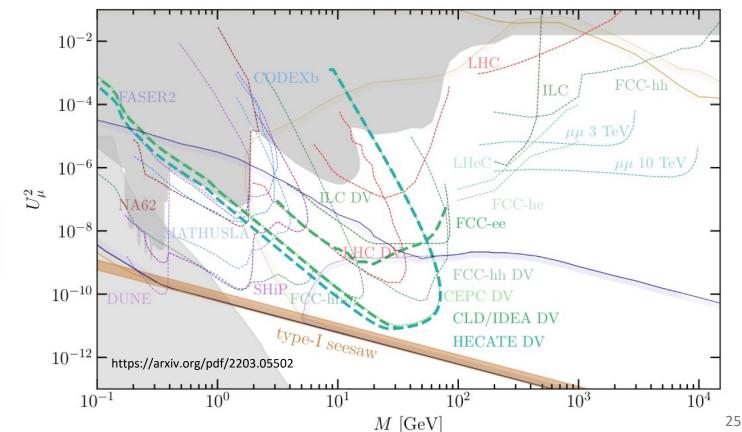
g<sub>aγ</sub> (TeV)<sup>-1</sup> 10 BaBar LEP I and II CDF CMS (PbPb) LHC ATLAS (pp)  $10^{-1}$ Beam Dump 10<sup>-2</sup>  $\gamma \gamma \rightarrow a$ FCCee (365 GeV) **FCCee** 10<sup>-3</sup> (240 GeV) **FCCee** FCCee (160 GeV)  $(e^+e^-\rightarrow \gamma a)$  $10^{-4}$ **FCCee** (91 GeV)  $10^{-5}$ 10<sup>-2</sup> 10<sup>-3</sup>  $10^{-1}$ 10<sup>2</sup> 10<sup>3</sup> 10 1 m<sub>a</sub> (GeV)

### Neutrino Portal: Heavy Neutral Leptons



One of the renormalizable portals to dark sectors. Could also address : Neutrino masses via seesaw mechanism, Baryogenesis via leptogenesis and oscillation anomalies...

BOSTON JNIVERSITY



Zeynep Demiragli – FCC Week 2024, San Francisco

4

mass

charge.

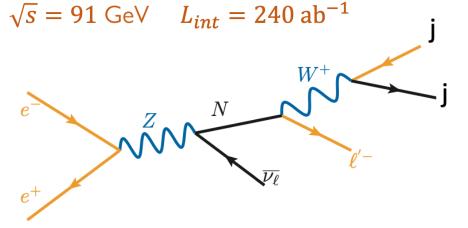
name

Quarks

Leptons

# Case Study: Prompt & Displaced HNL- µjj





High production rate. (~50% of the BR) Jets can be well separated or collimated.

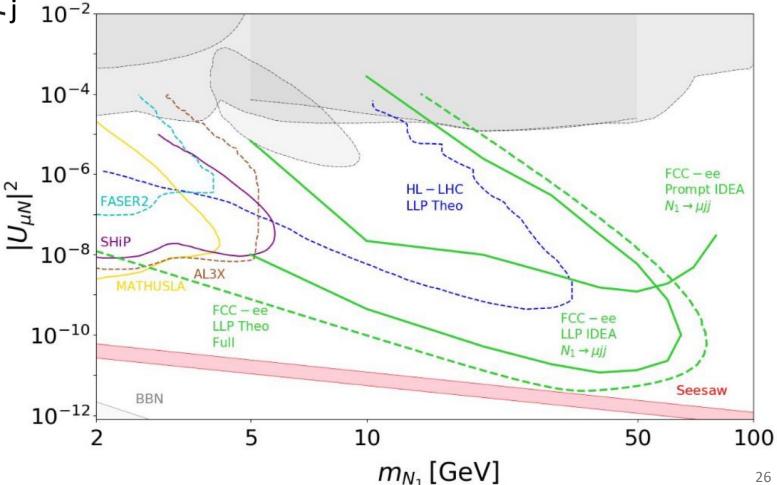
Primary backgrounds from Z :

 $\succ$  Z  $\rightarrow$  bb/cc/uds, Z  $\rightarrow$   $\mu\mu$  /  $\tau\tau$ 

 $rac{}{}e e \rightarrow \mu \nu q q \Rightarrow$  irreducible

Prompt selection: 1 muon, 1 or 2 jets, good PV

<u>Displaced selection</u>: Radial vertex position > 0.5 mm



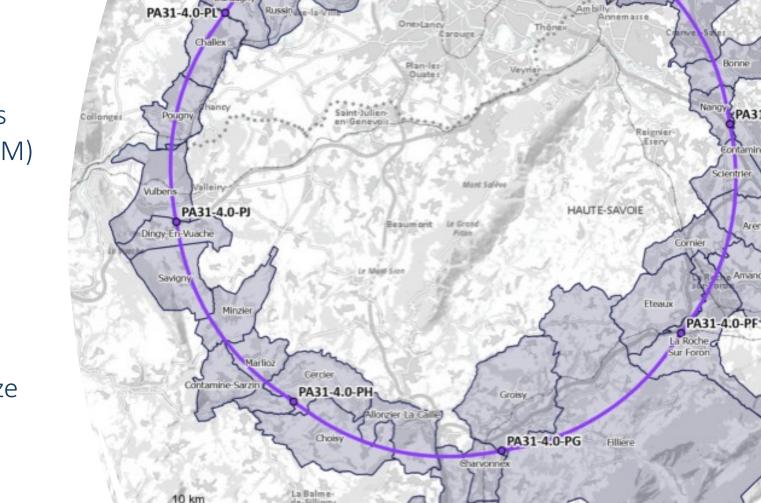
Zeynep Demiragli – FCC Week 2024, San Francisco

## FCC: A Gateway to Discovery

Today's Focus: Dark sector signals, including Heavy Neutral Leptons, Axionlike particles, and exotic Higgs decays.

Beyond the Dark Sector: The FCC offers numerous Beyond Standard Model (BSM) opportunities, with the <u>Z pole</u> and ZH pole runs being especially crucial. The FCC presents a large and unique phase space for exploration.

Let's design our detectors to maximize these opportunities!



PA31-4.0-PA

Foncene

PA31-4.0-PB

PA31-4.0

Arenthon

Contamine-Si

Meinie

Collonge-Belle

GE NEVE



### From <u>P. Azzi's slides</u>

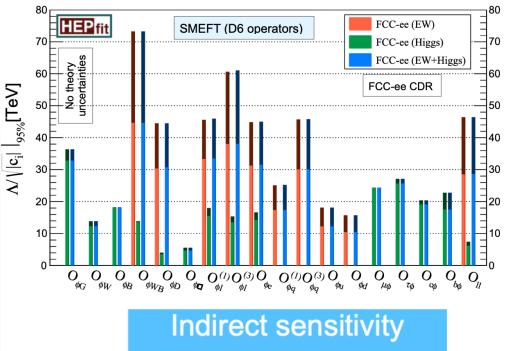
#### Indirect BSM sensitivity from EWPO

- Target: reduce syst. uncertainties to the level of statistical
- •Exquisite √s precision (100keV@Z, 300keV@WW)
- ~50 times better precision than LEP/LSD on EW precision observables

Quantity	Current precision	FCC-ee stat. (syst.) precision	Required theory input	Available calc. in 2019	Needed theor $\operatorname{improvement}^{\dagger}$
$m_{ m Z} \ \Gamma_{ m Z} \ \sin^2  heta_{ m eff}^\ell$	$2.1 \mathrm{MeV}$ $2.3 \mathrm{MeV}$ $1.6  imes 10^{-4}$	$\begin{array}{l} 0.004  (0.1)  {\rm MeV} \\ 0.004  (0.025)  {\rm MeV} \\ 2(2.4) \times 10^{-6} \end{array}$	non-resonant $e^+e^- \rightarrow f\bar{f},$ initial-state radiation (ISR)	NLO, ISR logarithms up to 6th order	NNLO for $e^+e^- \rightarrow f\bar{f}$
$m_W$	$12{ m MeV}$	$0.25 (0.3) \mathrm{MeV}$	lineshape of $e^+e^- \rightarrow WW$ near threshold	NLO (ee $\rightarrow$ 4f or EFT framework)	NNLO for ee $\rightarrow$ WW, W $\rightarrow$ ff in EFT setup
HZZ coupling	—	0.2%	cross-sect. for $\rm e^+e^- \rightarrow ZH$	NLO + NNLO QCD	NNLO electroweak
$m_{ m top}$	$100{ m MeV}$	17 MeV	threshold scan $e^+e^- \rightarrow t\bar{t}$	$N^{3}LO$ QCD, NNLO EW, resummations up to NNLL	Matching fixe orders with resummations merging with MC, $\alpha_{s}$ (inpu

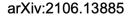
Need TH results to fully exploit Tera-7

 $^\dagger {\rm The}$  listed needed theory calculations constitute a minimum baseline; additional partial higher-order contributions may also be required.



to 70TeV-scale sector

connected to EW/Higgs



10

### M. Carena Slides

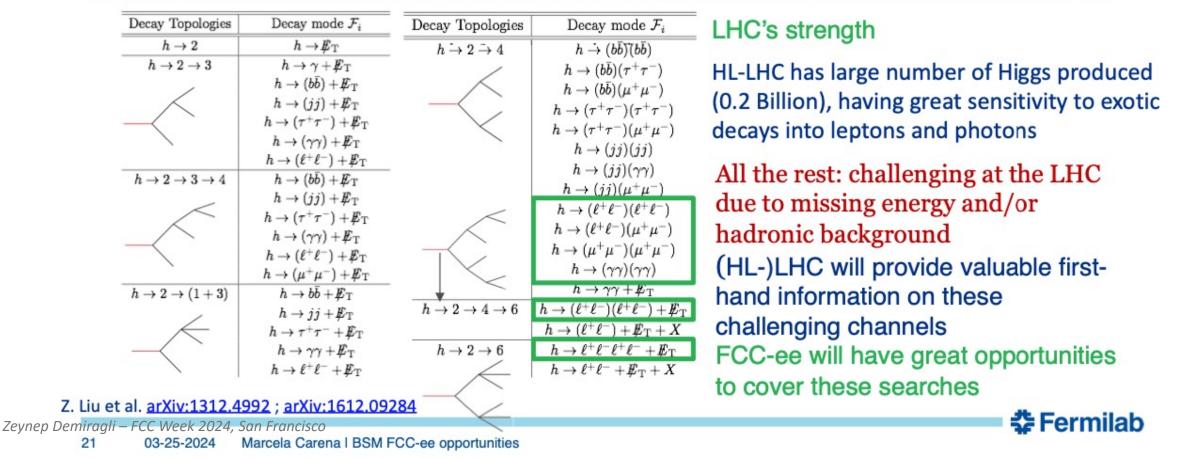
21

UNIVERSIT

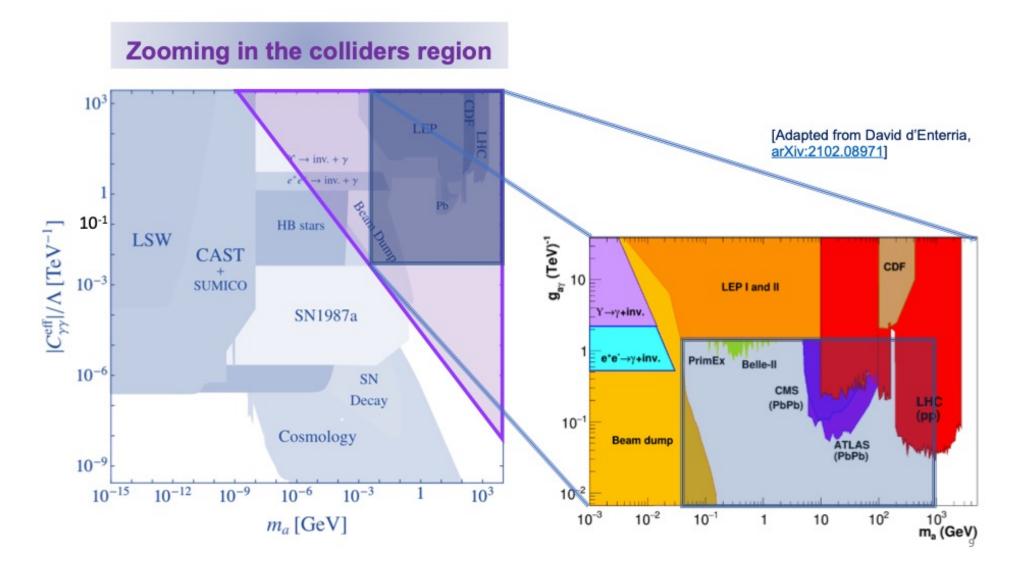
30

#### Higgs exotic decays: a rich variety of possibilities

- Focus on 2-body Higgs decays to BSM particles with subsequent decays to BSM or SM particles
- These processes are well-motivated by SM + Scalar singlets, 2HDMs (+ Scalar), SUSY models, ٠ gauge SM extensions (e.g. dark photons), SM + Fermion/s (e.g. Heavy Neutral leptons), etc.



#### Patricia R. Teles



#### HNL Full Cuts



1. Event Filter	2. Event Selection	3. Vertex selection
$\begin{array}{l} 1 { m ~muon} \ \geq 3 { m ~tracks} \ E_{\mu} \geq 3 { m ~GeV} \ E_{miss} \geq 5 { m ~GeV} \end{array}$	1 lepton (muon) Cuts on $p_{miss}$ , jets, $\mu$ and visible mass	$\begin{array}{l} N_{tracks} - N_{tracks}^{primary} < 5 \\ \chi^2_{vtx, primary} < 10 \end{array}$
4. Mass-dependent kin. selection	5a. Displacement: prompt	5b. Displacement: LL
$M_{\rm min}$ within $2 \times 10\% \sqrt{M}$	$r^{primary} > 0.5 \text{ mm}$	mrimaru

 $M_{vis}$  within  $2 \times 10\% \sqrt{M}$  $E_{miss}$  within  $2 \times 10\% \sqrt{p_{\nu}}$   $\begin{array}{l} r_{vert}^{primary} > 0.5 \ \mathrm{mm} \\ D_{0,\mu} < 8\sigma \ \mathrm{if} \ M_{N_1} > 70 \end{array}$ 

 $r_{vert}^{primary} < 0.5 \text{ mm}$