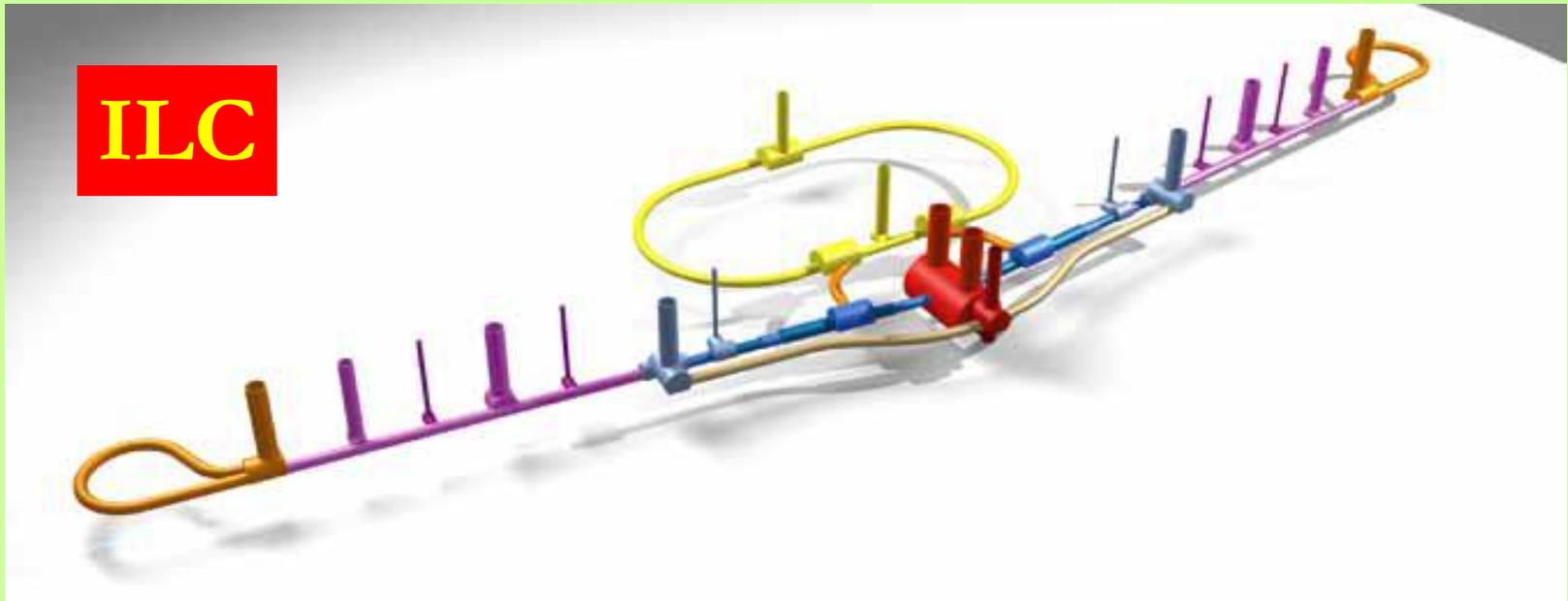


# Future Opportunities in Particle Physics



Barry Barish  
Caltech

*DPF Meeting 13-Aug-2011*

# *What is Particle Physics?*

## The definition from Wikipedia:

- “Particle physics is a branch of physics that studies the elementary subatomic constituents of matter and radiation, and their interactions. The field is also called high energy physics, because many elementary particles do not occur under ambient conditions on Earth. They can only be created artificially during high energy collisions with other particles in particle accelerators.”
- “Particle physics has evolved out of its parent field of nuclear physics and is typically still taught in close association with it. Scientific research in this area has produced a long list of particles.”

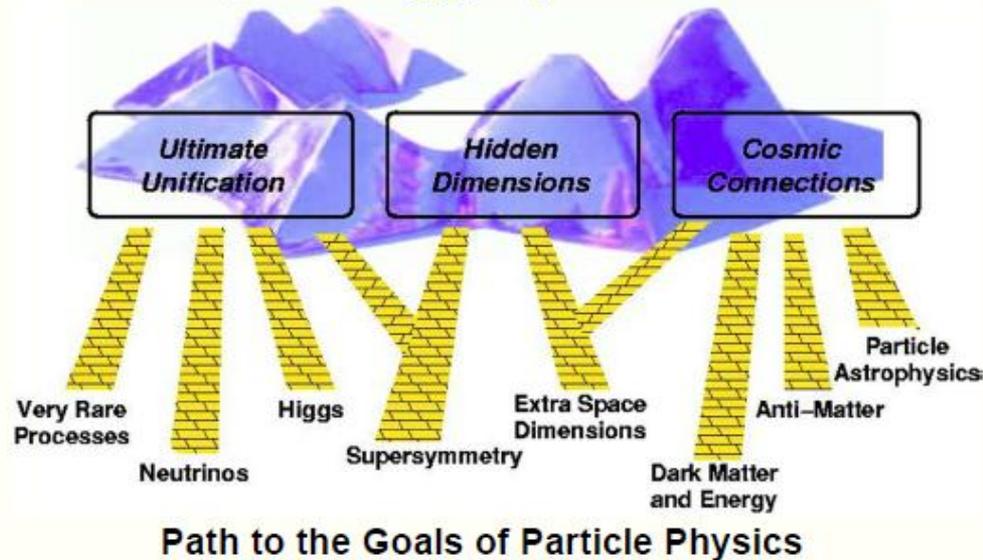
# *What is Particle Physics?*

## The definition from DoE Office of Science:

- “High Energy Physics explores the most fundamental questions about the nature of the universe. The Office of High Energy Physics supports a program focused on three frontiers of scientific discovery. At the *energy frontier*, powerful accelerators investigate the constituents and architecture of the universe. At the *intensity frontier*, astronomically large amounts of particles and highly sensitive detectors offer a second, unique pathway to investigate rare events in nature. At the *cosmic frontier*, natural sources of particles from space reveal the nature of the universe. Together these three interrelated discovery frontiers create a complete picture, advancing Department of Energy *missions* through the development of key cutting-edge technologies and the training of future generations of scientists.”

# *What is Particle Physics?*

## Matter, Energy, Space and Time



*From each of these goals flows a diverse research program that will be carried out in partnership with colleagues across the globe.*

The definition from Bagger-Barish (2001)

# *Particle Physics: an Inquiry Based Science*

1. How can we solve the mystery of dark energy?
2. Are there extra dimensions of space?
3. Do all the forces become one?
4. Why are there so many kinds of particles?
5. What is dark matter?

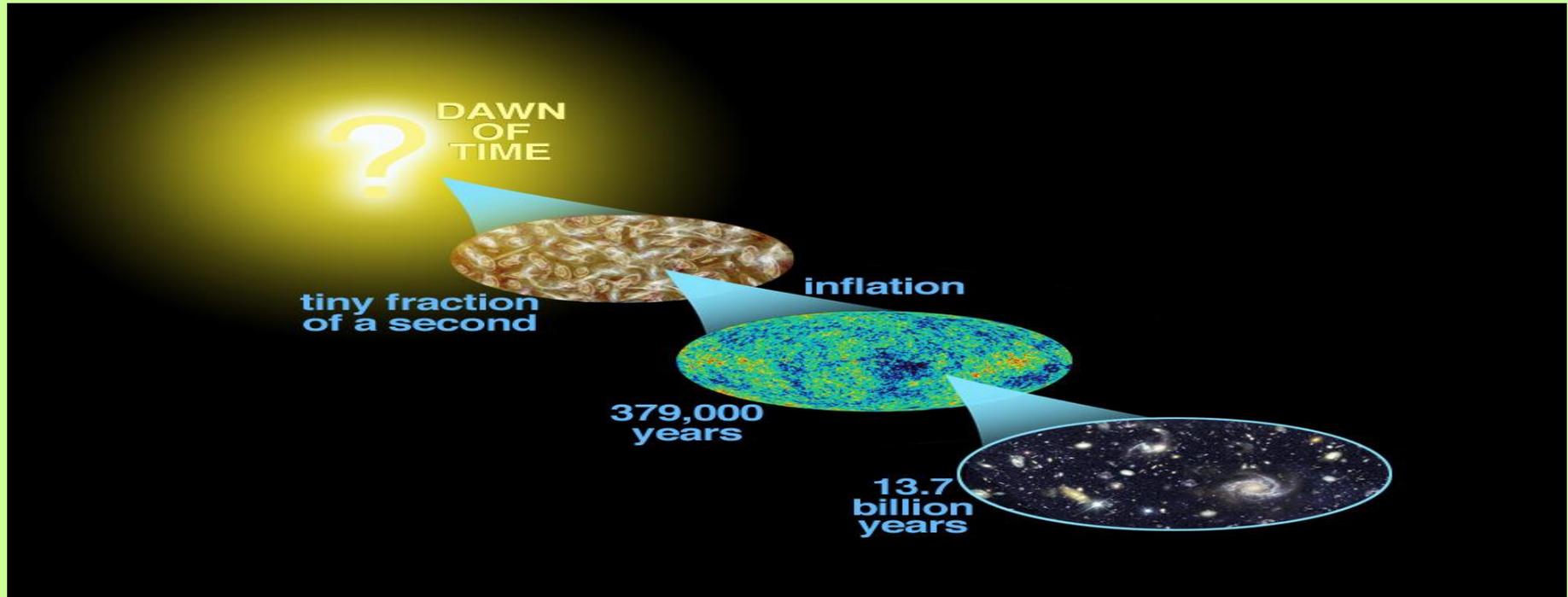
How can we make it in the laboratory?

6. What are neutrinos telling us?
7. How did the universe come to be?
8. What happened to the antimatter?
9. Are there undiscovered principles of nature:

New symmetries, new physical laws?

*from the Quantum Universe*

# The Nature of Particle Physics



*“There are more things in heaven and earth,  
Horatio, than are dreamt of in your philosophy”  
(Hamlet, I.5)*

# Many Experimental Probes

- Neutrinos

- » Particle physics and astrophysics using a weakly interacting probe



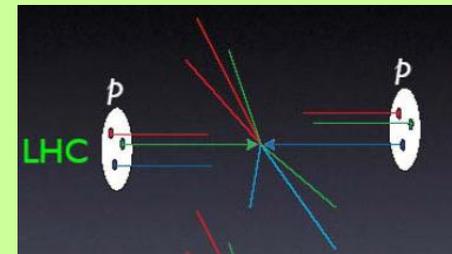
- Particle Astrophysics/Cosmology

- » Dark Matter; Cosmic Microwave, etc



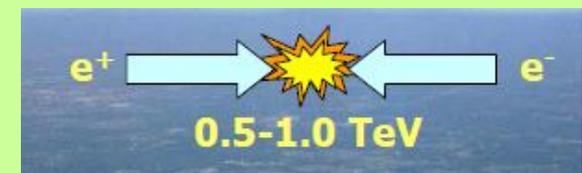
- High Energy pp Colliders

- » Opening up a new energy frontier (~ 1 TeV scale)



- High Energy Lepton Colliders

- » Precision Physics at the new energy frontier



# and more Experimental Probes

- Precision Physics

- » SuperB factories
- » Fermilab:  $g-2$ ;  $\mu \rightarrow e$ ; etc.



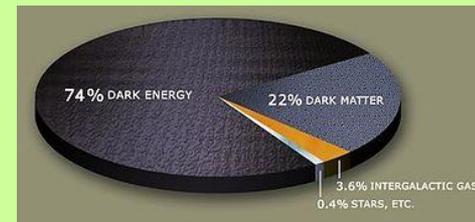
- Gravity

- » Gravitational waves; quantum gravity



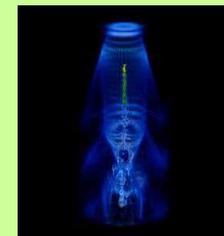
- Dark Energy

- » Space missions



- New Generations of Accelerators

- » Plasma/laser wakefield acceleration



# Addressing the Questions

- Neutrinos

- » Particle physics and astrophysics using a weakly interacting probe



# Neutrinos – Broad field; Specialized Expts

- **Neutrino Properties**

- » Why are neutrino masses so small and what is their mass scale?
- » Separation and ordering of neutrino masses?
- » Are neutrinos their own antiparticles?

- **Neutrinos in Astrophysics**

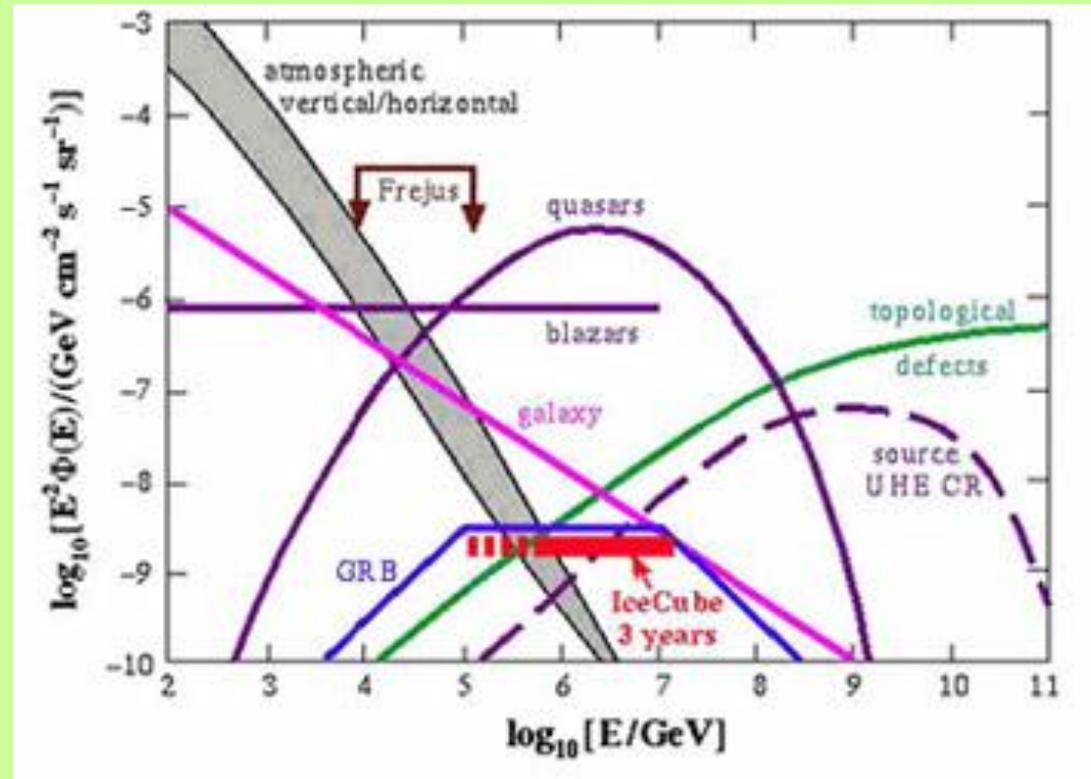
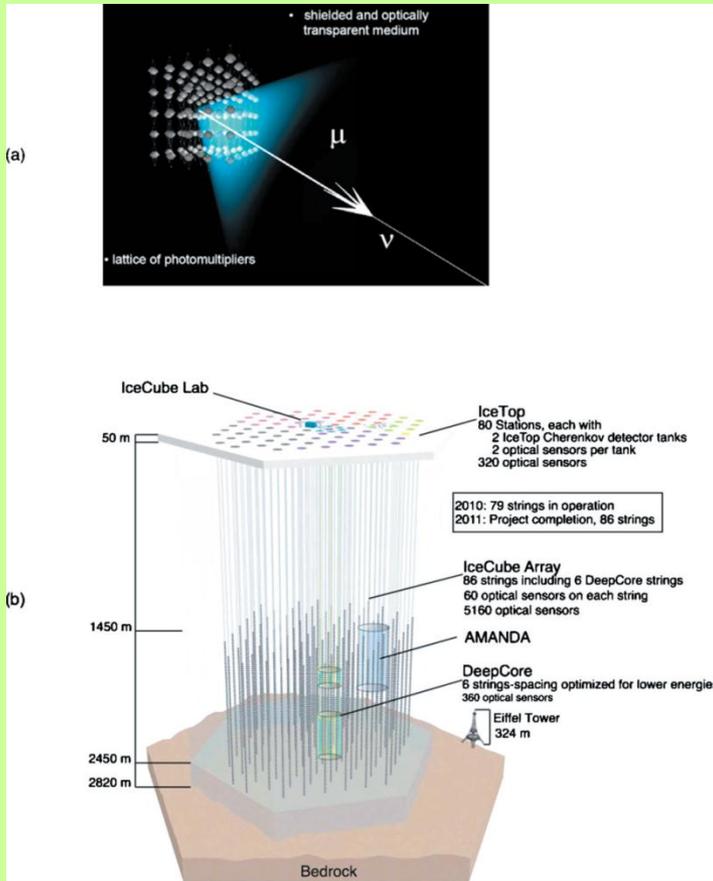
- » Astrophysical sources of neutrinos, and their contribution to the dark matter?

- **Neutrinos and Fundamental Particle Physics**

- » CP violation in neutrinos, leptogenesis, possible role in the early universe and in understanding the particle antiparticle asymmetry in nature?

# Ice Cube: Neutrino Astronomy

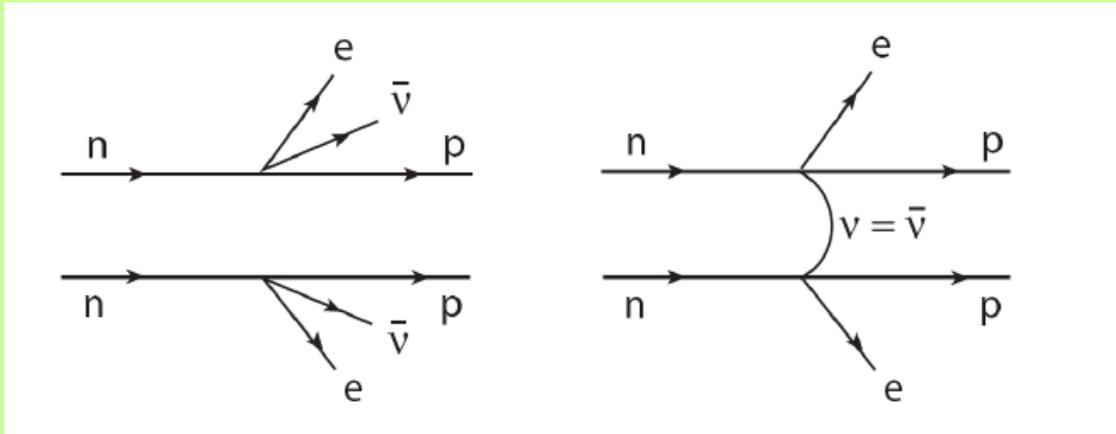
- Neutrino Astrophysics – Investigating astrophysical sources emitting ultra high energy neutrinos



South Pole

# Is the neutrino its own antiparticle?

## ■ Neutrinoless Double Beta Decay



## ■ If neutrino is a Majorana particle (own antiparticle)

- » Determine neutrino mass scale
- » Importance for Grand Unification
- » Lepton number violation
- » Implications for leptogenesis

# Neutrinoless Double Beta Decay

- **Experimental Challenges – What technique?**
  - » Isotopes decaying to monochromatic electrons over continuum background. No obvious best isotope? Therefore, several large-scale experiments are likely needed.
  - » Which technique scales to the needed mass (Xenon, Germanium, ???)
- **Experimental Challenges – Scale required**
  - » Present experiments – 10 kg  $\rightarrow$  sensitivity to  $m_\nu < 1$  eV
  - » Need  $\sim 1$  ton detectors to reach atmospheric  $m_\nu \sim 50$  meV
  - » Need  $\sim 50$  tons to reach solar  $m_\nu \sim 1$  meV scale
- **Other issues: Low background environment; costs; etc**

# Experimental Probes

- Neutrinos

- » Particle physics and astrophysics using a weakly interacting probe



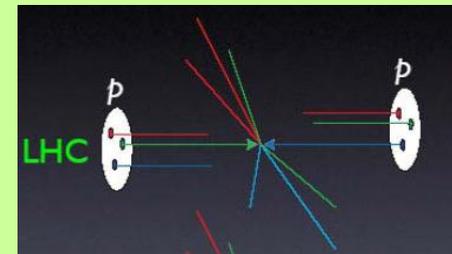
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- » Dark Matter; Cosmic Microwave, etc



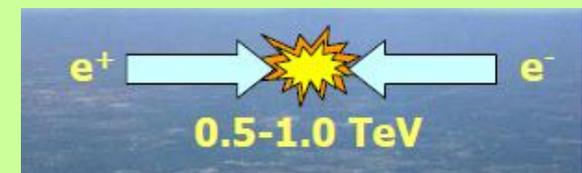
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- » Precision Physics at the new energy frontier

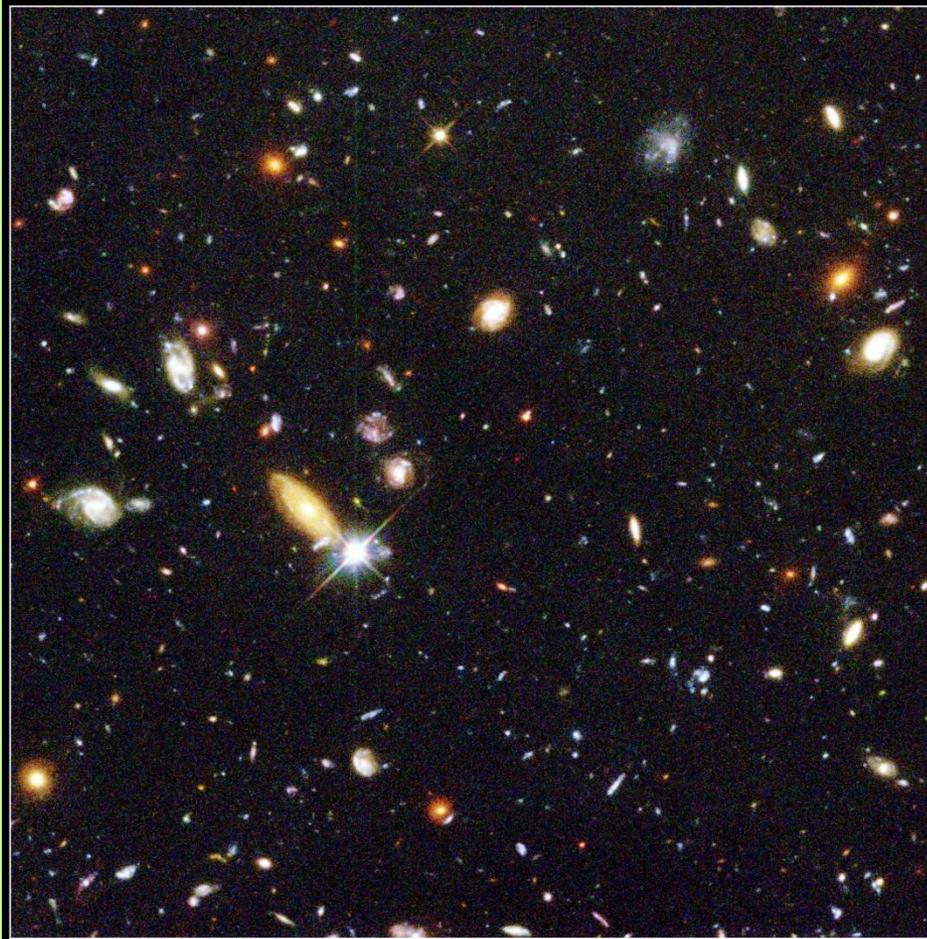


# Approaches

- Particle Astrophysics/Cosmology
  - » Dark Matter; Cosmic Microwave, etc



# Dark Matter



**Hubble Deep Field**  
Hubble Space Telescope · WFPC2

PRC96-01a · ST ScI OPO · January 15, 1995 · R. Williams (ST ScI), NASA

What don't we see?

Dark Matter

Neutrinos

Dark Energy

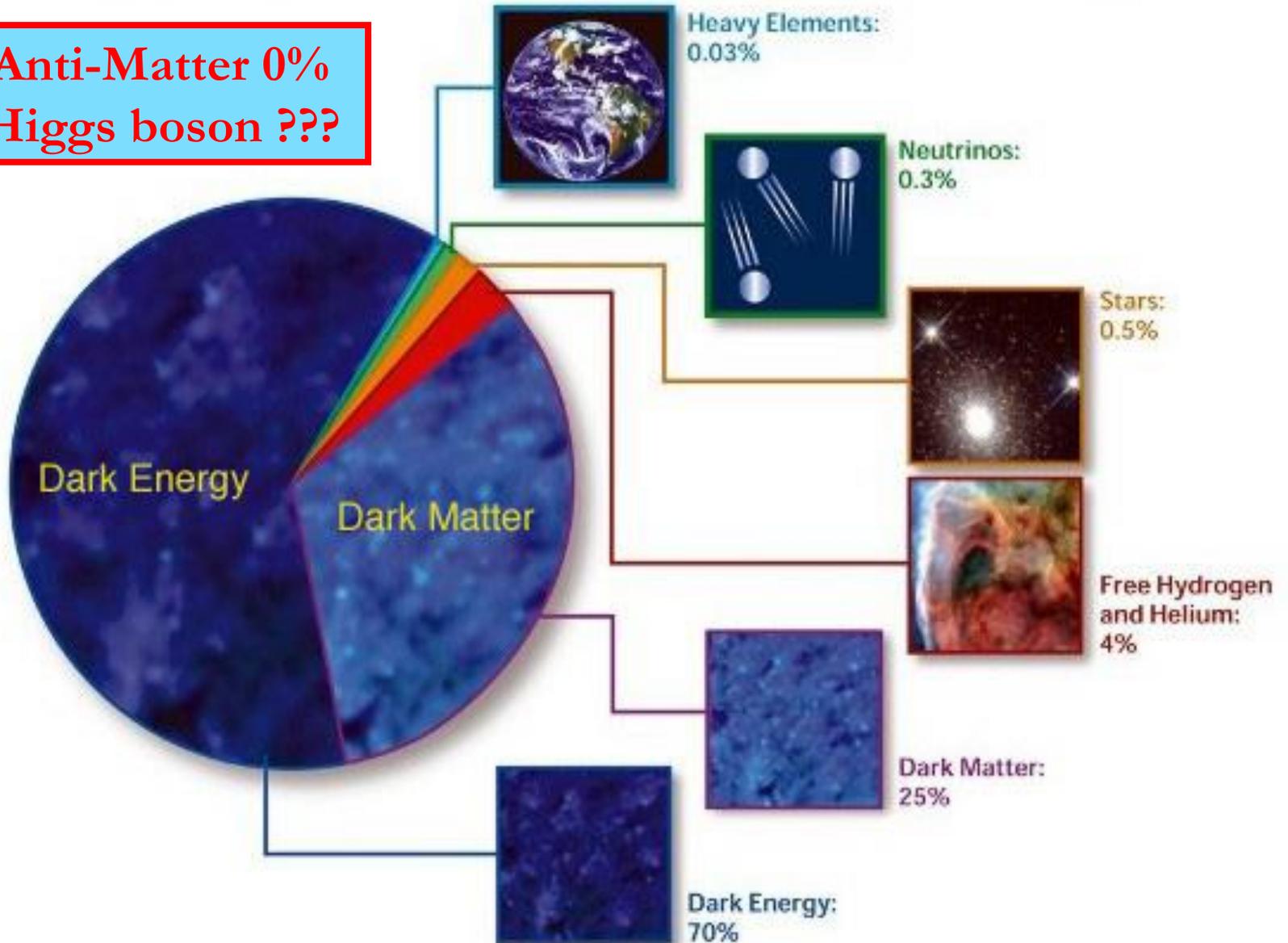
...

Higgs Bosons !

Antimatter !!

# The Energy Budget of the Universe

Anti-Matter 0%  
Higgs boson ???



# Dark Matter

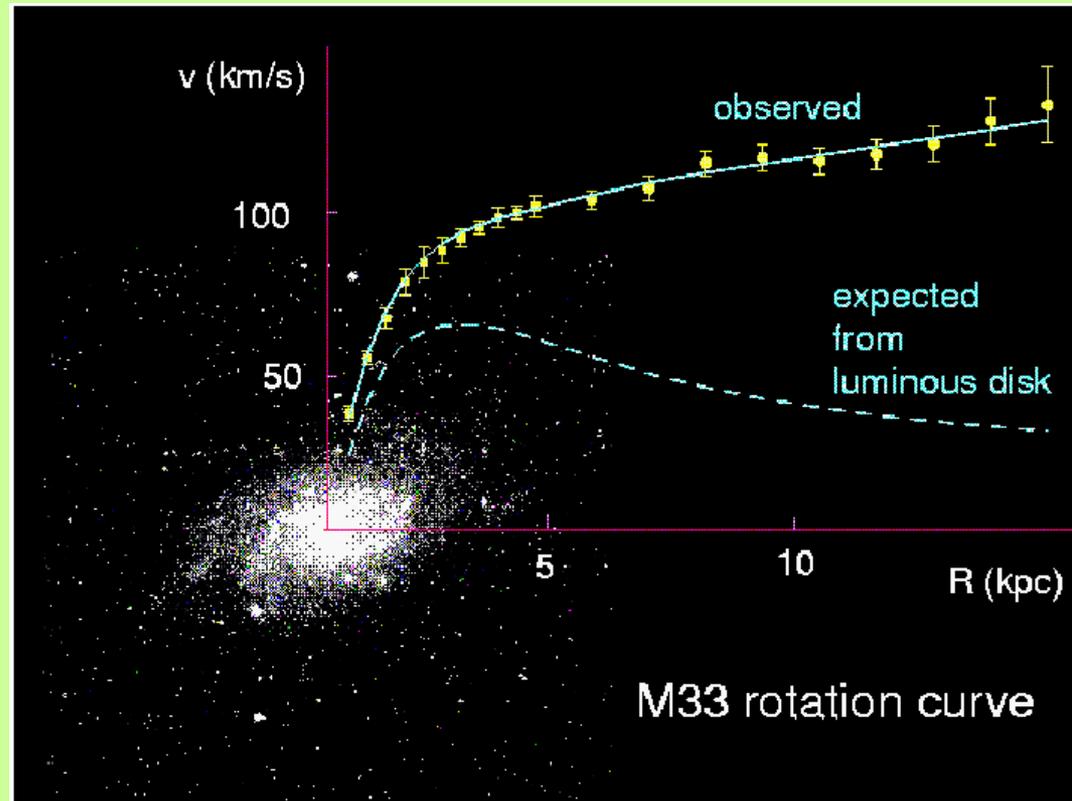
## *the evidence*

From the Kepler's law,  $v_{circ} = \sqrt{\frac{GM(r)}{r}}$  for  $r$  much larger than the luminous terms, you should have  $v \propto r^{-1/2}$  However, Instead, it is flat or rises slightly.

This is the most direct evidence for dark matter.

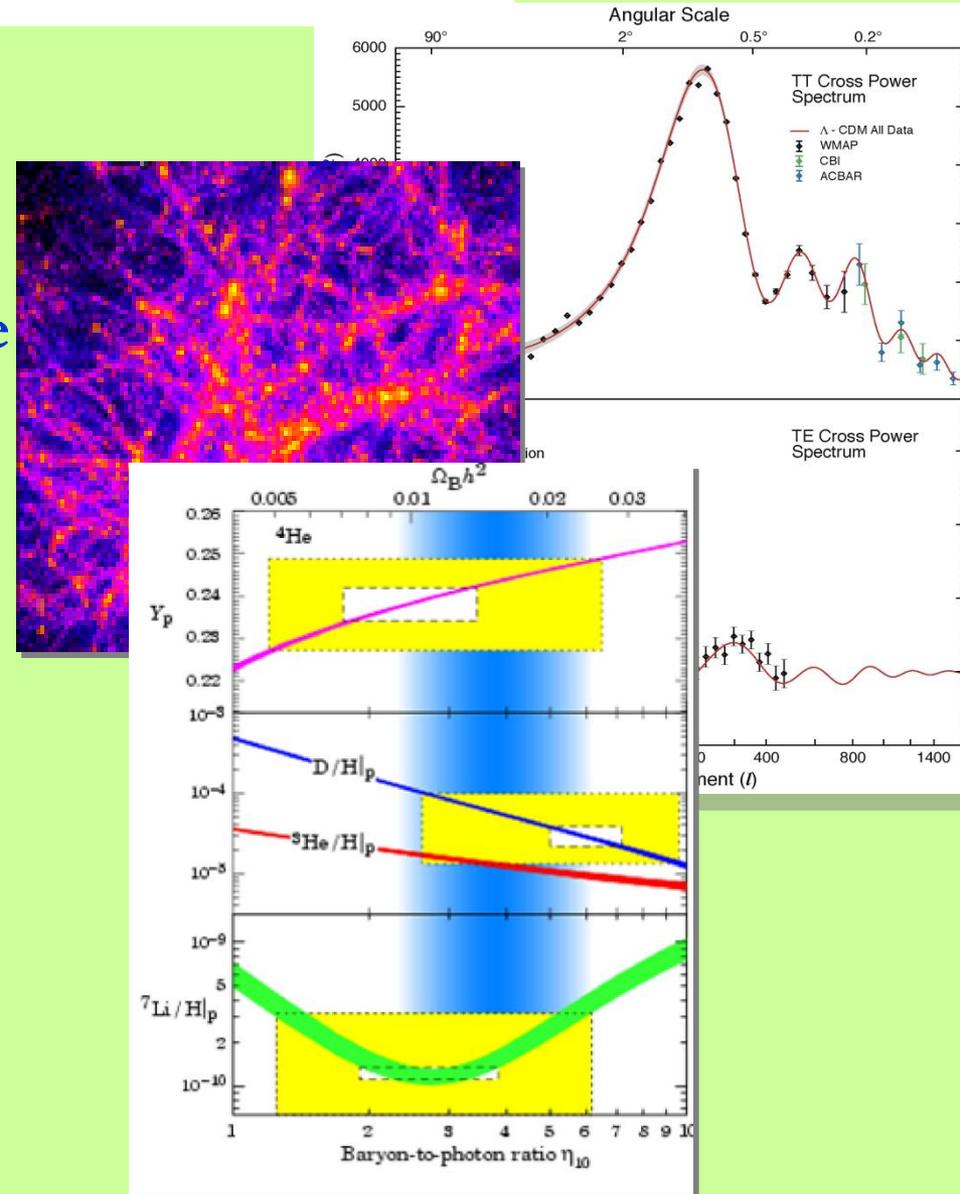
There are many complementary measurements at all scales

Corbelli & Salucci (2000);  
Bergstrom (2000)



# Other Dark Matter Evidence

- Evidence from a wide range of astrophysical observations including rotation curves, CMB, lensing, clusters, BBN, SN1a, large scale structure
- Each observes dark matter through its gravitational influence
- Still no (reliable) observations of dark matter's electroweak interactions (or other non-gravitational interactions)
- Still no (reliable) indications of dark matter's particle nature



# Dark Matter Particle Candidates

Axions, Neutralinos, Gravitinos, Axinos, Kaluza-Klein Photons, Kaluza-Klein Neutrinos, Heavy Fourth Generation Neutrinos, Mirror Photons, Mirror Nuclei, Stable States in Little Higgs Theories, WIMPzillas, Cryptons, Sterile Neutrinos, Sneutrinos, Light Scalars, Q-Balls, D-Matter, Brane World Dark Matter, Primordial Black Holes, ...

**EVIDENCE STRONGLY FAVORS NON-BARYONIC COLD DARK MATTER**

# Leading Dark Matter Candidate

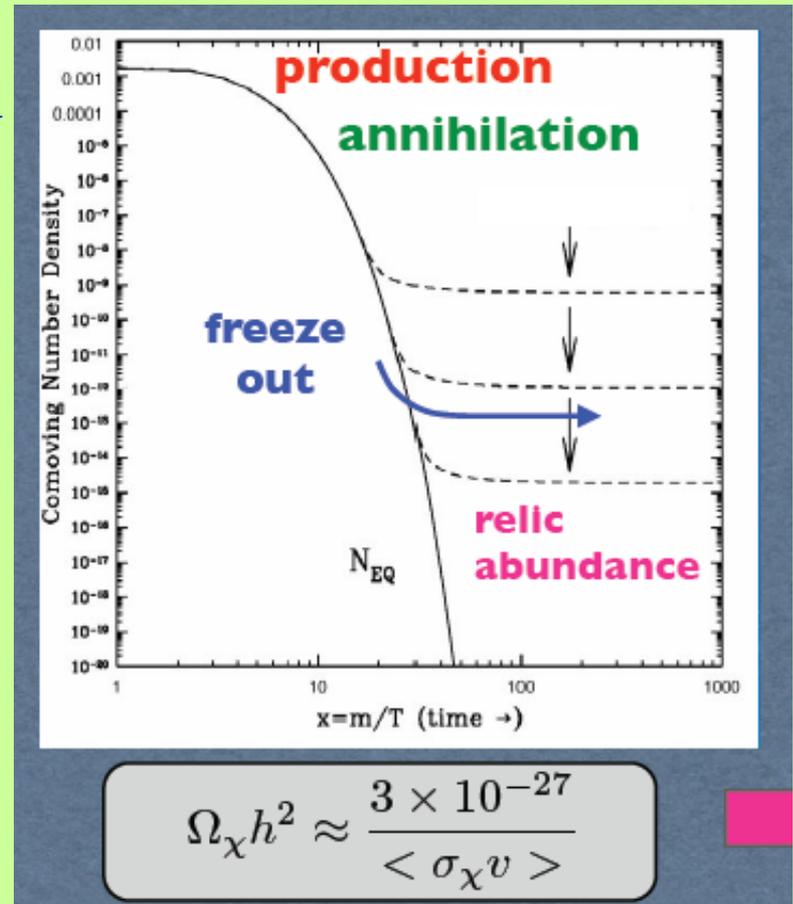
## *Weakly Interacting Massive Particles (WIMPs)*

Weakly interacting particles produced thermally in the early universe

Large mass compared to standard particles.

Due to their large mass, they are relatively slow moving and therefore “cold dark matter.”

Leading candidate – “Supersymmetric Particles”



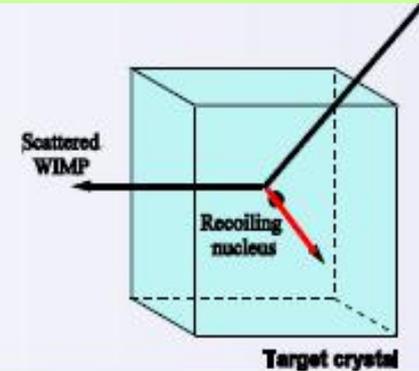
Supersymmetric dark matter would solve one of biggest problems in astrophysics and particle physics at the same time !

# Searching for Dark Matter

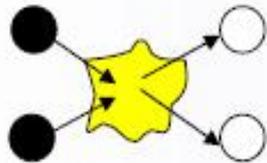
## Underground

- Direct Detection:

Look for the elastic scattering of dark matter with nuclei



## In Space



- Neutrinos
- Photons
- Electron-Positron, Antimatter

- Indirect Detection:

Look for the annihilation products

## On Accelerators

Look for signals of new physics



# Experimental Probes

- Neutrinos

- » Particle physics and astrophysics using a weakly interacting probe



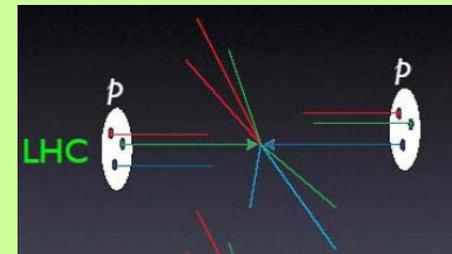
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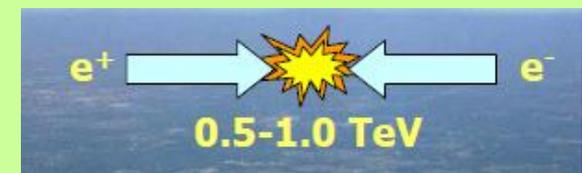
- High Energy pp Colliders

- » Opening up a new energy frontier (~ 1 TeV scale)



- High Energy Lepton Colliders

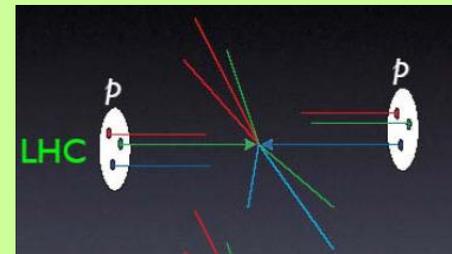
- » Precision Physics at the new energy frontier



# Addressing the Questions

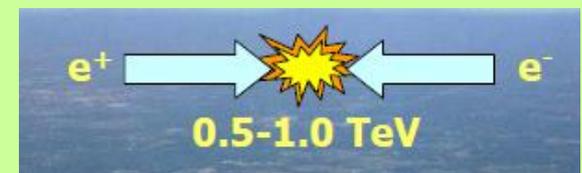
- **High Energy pp Colliders**

- Opening up a new energy frontier  
( $\sim 1$  TeV scale)



- **High Energy  $e^+e^-$  Colliders**

- Precision Physics at the new energy frontier



# Exploring the Terascale

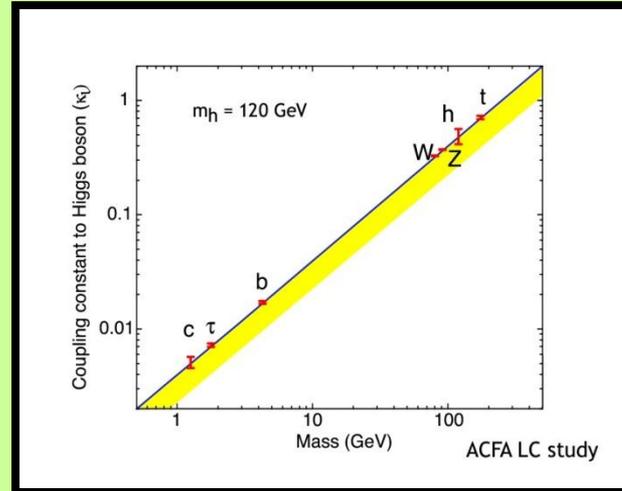
## *the tools*

- **The LHC**
  - » It will lead the way and has large reach
  - » Quark-quark, quark-gluon and gluon-gluon collisions at 0.5 - 5 TeV
  - » Broadband initial state
- **A Lepton Collider (e.g. ILC or ?)**
  - » A second view with high precision
  - » Electron-positron collisions with fixed energies
  - » Well defined initial state
- **Together, these are our tools for the terascale**

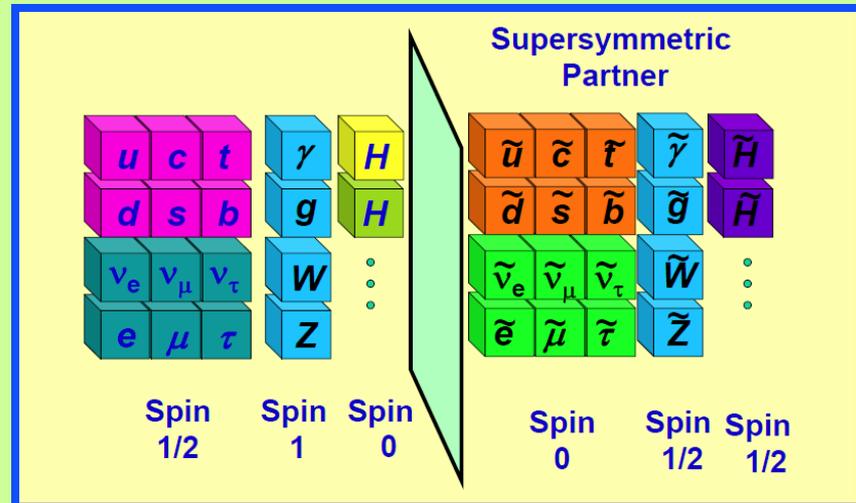
# LHC

*The machine and detectors work very well and results are emerging*

- Higgs



- Supersymmetry

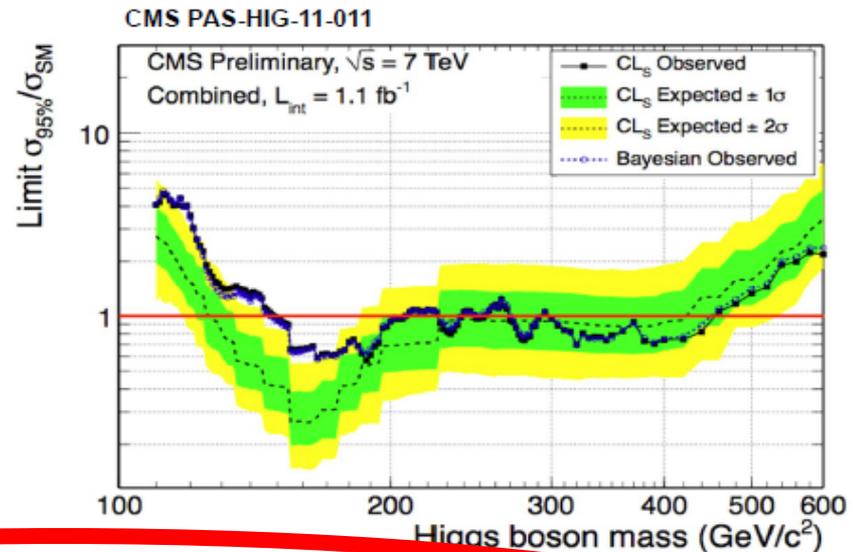
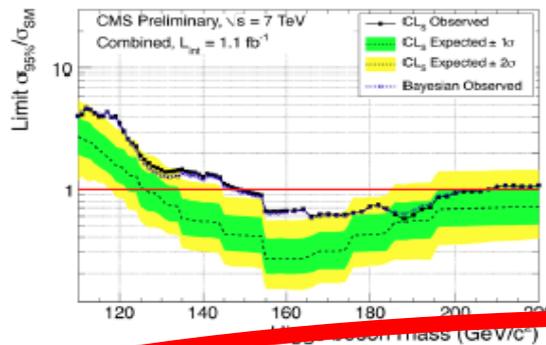


# LHC/CMS: Higgs Search



## SM Higgs exclusion limits

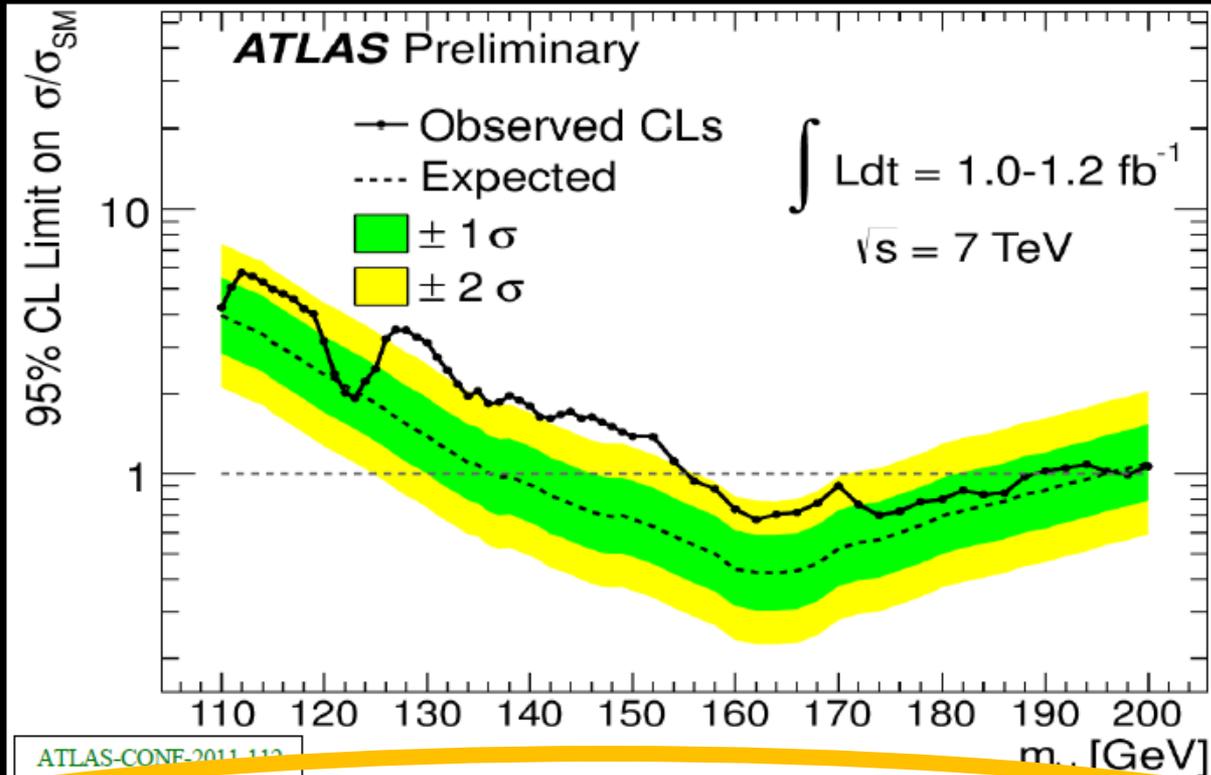
**Expected exclusion: 127-420 GeV**  
**Observed exclusion: 149-206 GeV**  
**+ 300-440 GeV + parts in between**



In the low mass part (114-149 GeV) we see a couple of interesting regions showing excesses larger than  $3\sigma$  (local significance without correction for LEE effects). Further study with the new data we are collecting will hopefully tell us if we are seeing a background fluctuation or a first sign of the Higgs boson.

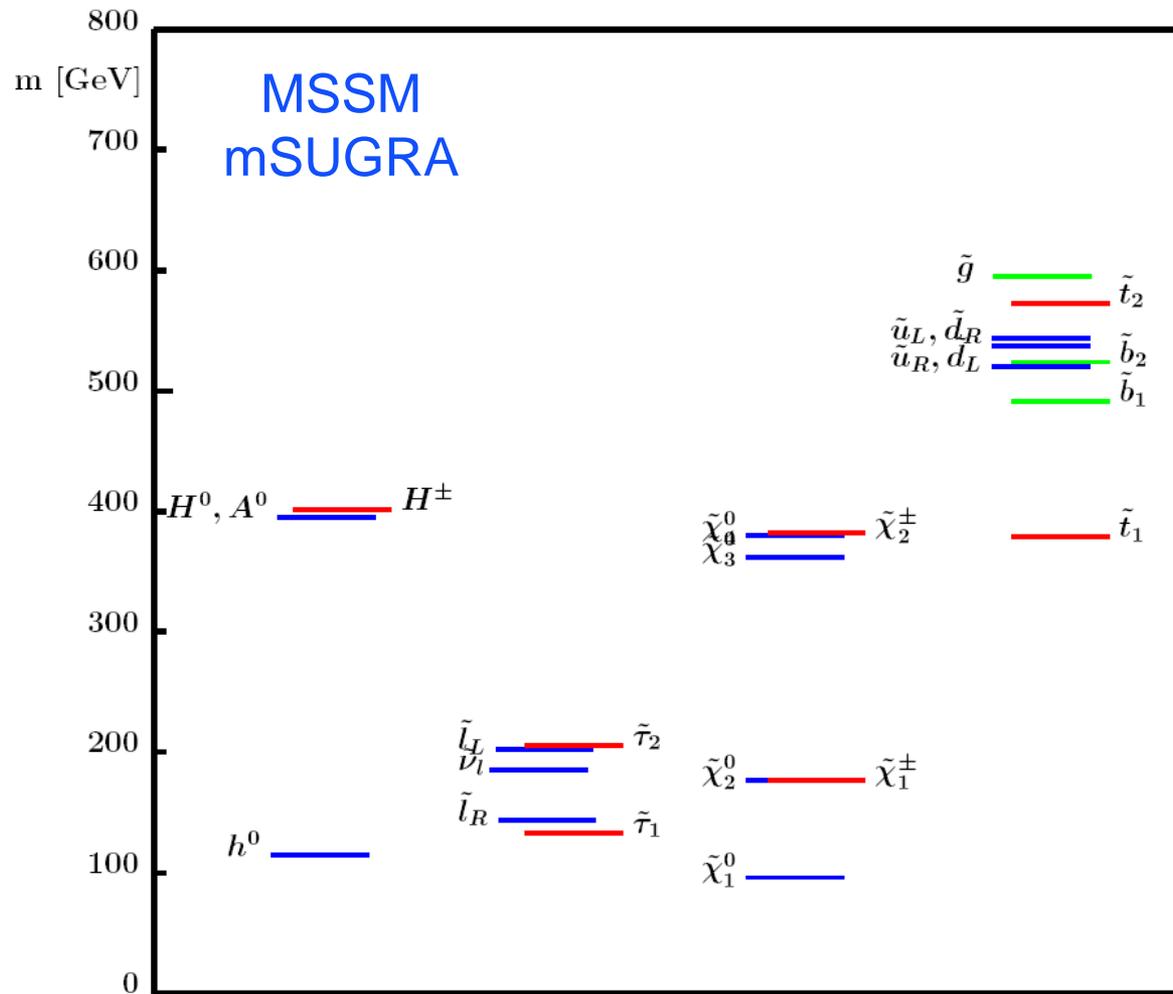
# LHC/ATLAS: Higgs Searches

## The Low Mass Region Is there more to say yet?



Limit on  $\sigma/\sigma_{\text{SM}}$  from data is generally rather poorer than median expectation if no Higgs  
Is it a fluctuation, or a background, or something else?

# Spectrum of Supersymmetric Particles



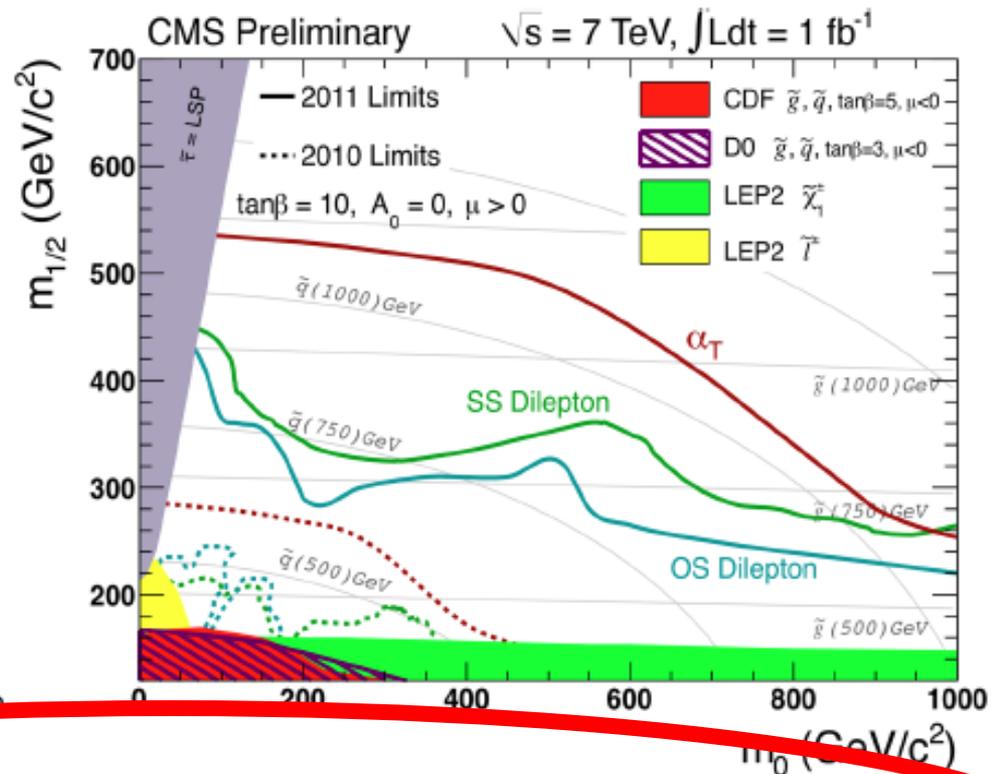
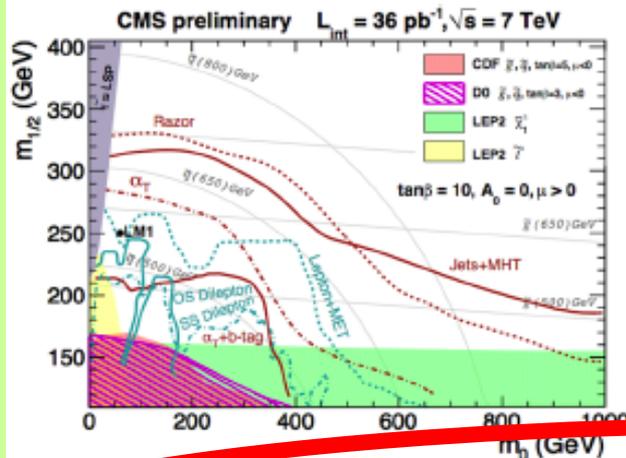
squarks and  
sgluons heavy  
yielding long  
decay chains  
ending with LSP  
neutrino

# LHC/CMS Supersymmetric Particles



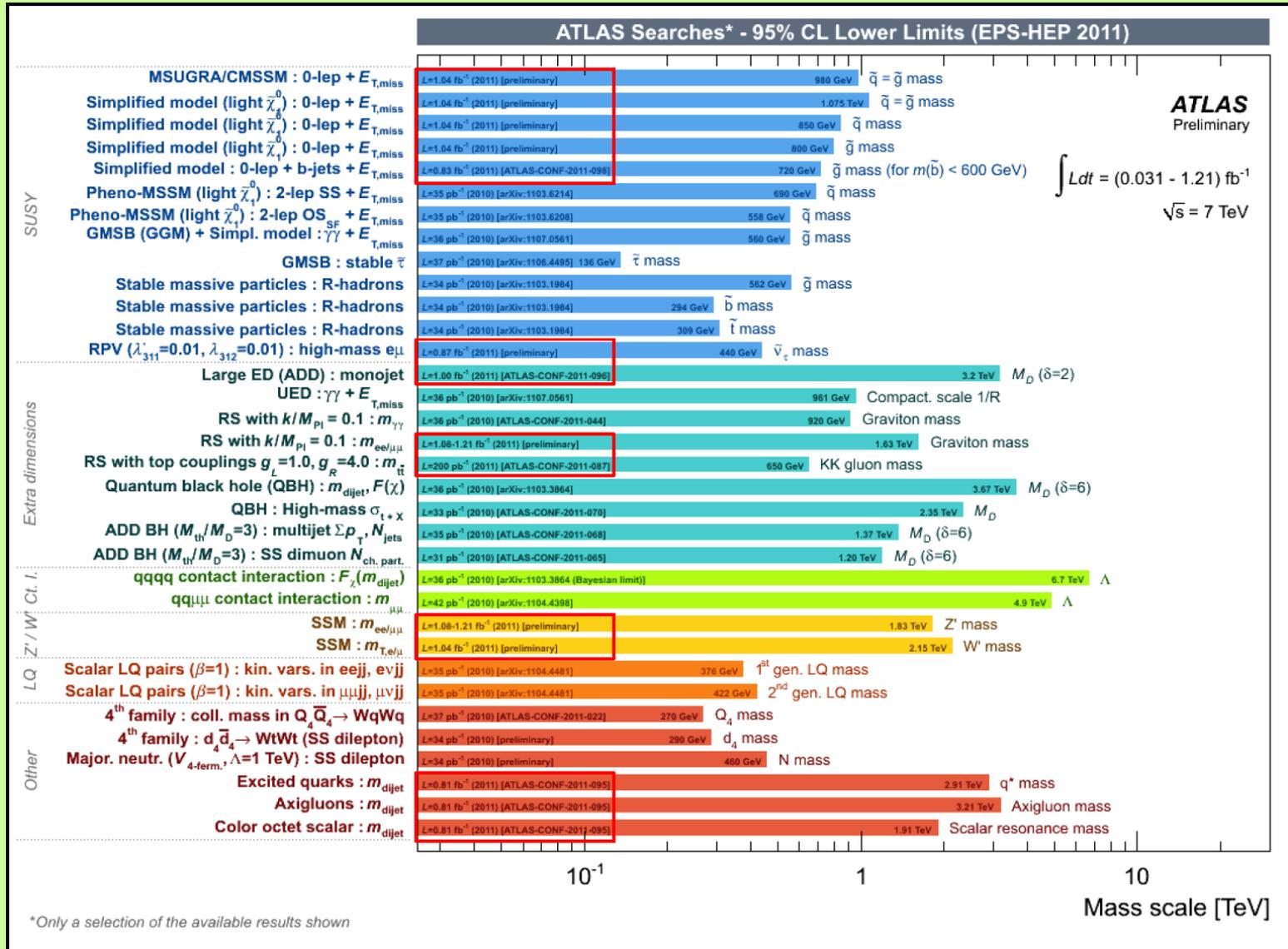
## Progress on SUSY

Results of the first three SUSY analyses completed on 2011 data ( $\alpha_T$ , Same Sign and Opposite Sign dileptons).



Within the constrained SSM models we are crossing the border of excluding gluinos and squarks up to 1TeV and beyond. The air is getting thin for constrained SUSY. More conclusive results after summer.

# LHC/ATLAS Supersymmetric Particles



# more Experimental Probes

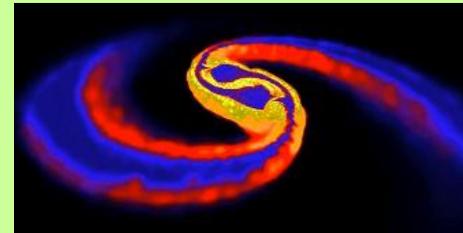
## ■ Precision Physics

- » SuperB factories
- » Fermilab:  $g-2$ ;  $\mu \rightarrow e$ ; etc.



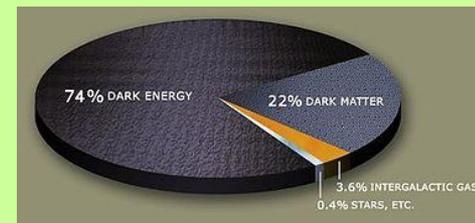
## ■ Gravity

- » Gravitational waves; quantum gravity



## ■ Dark Energy

- » Space mission (Europe)



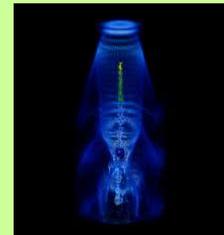
## ■ New Generations of Accelerators

- » Plasma/laser wakefield acceleration



# Addressing the Questions

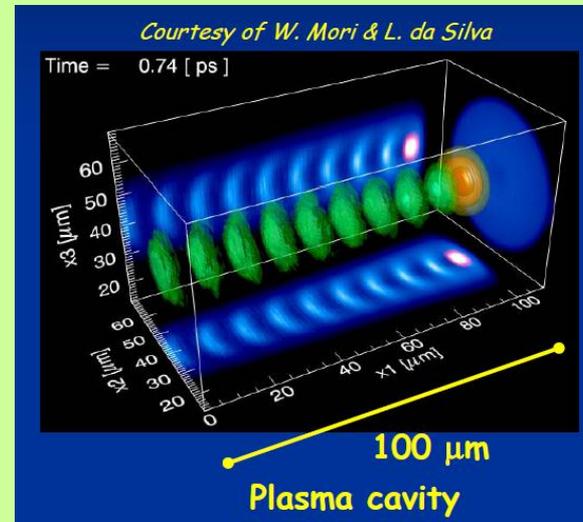
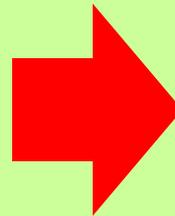
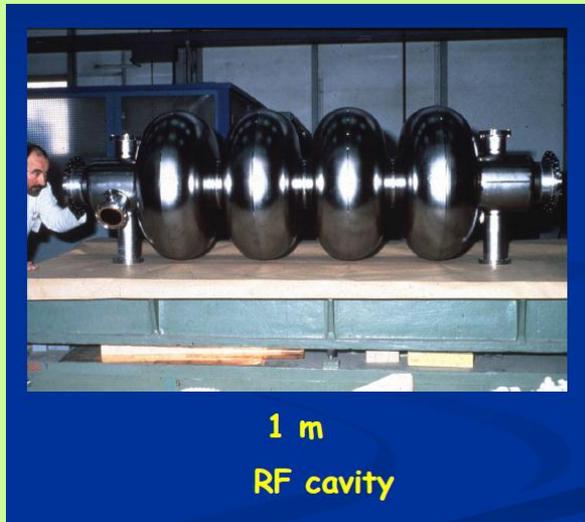
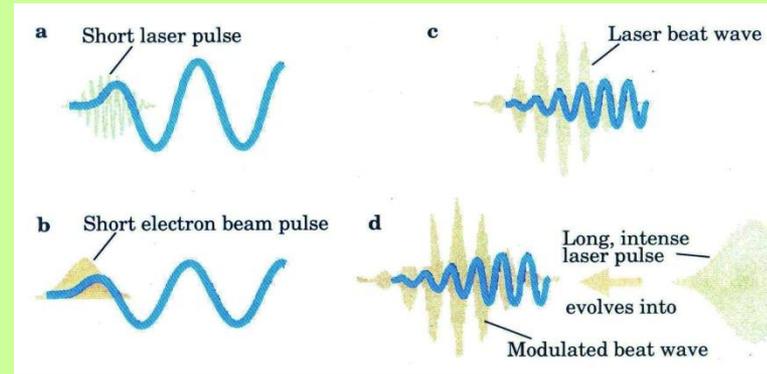
- **New Generations of Accelerators**
  - » **Plasma/laser wakefield acceleration**



# Future of Accelerator Based Particle Physics?

Future of Accelerators (eliminate materials!)

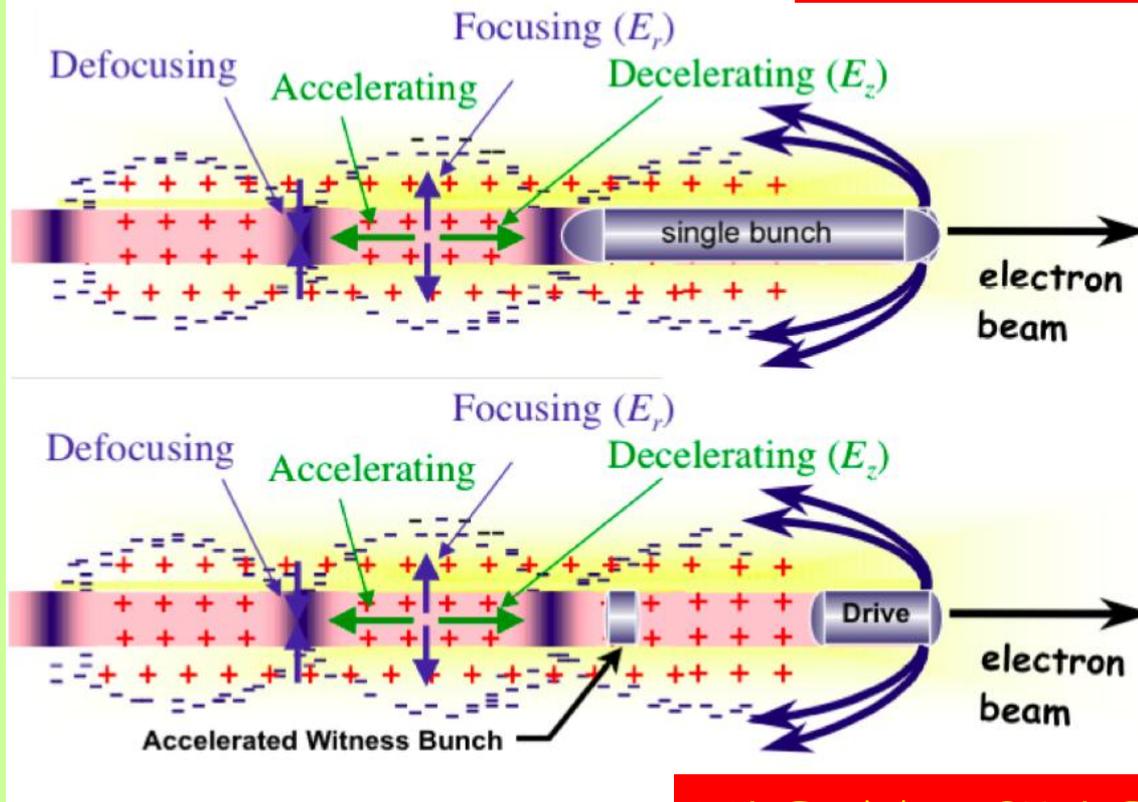
Plasma/Laser  
Wakefield Acceleration



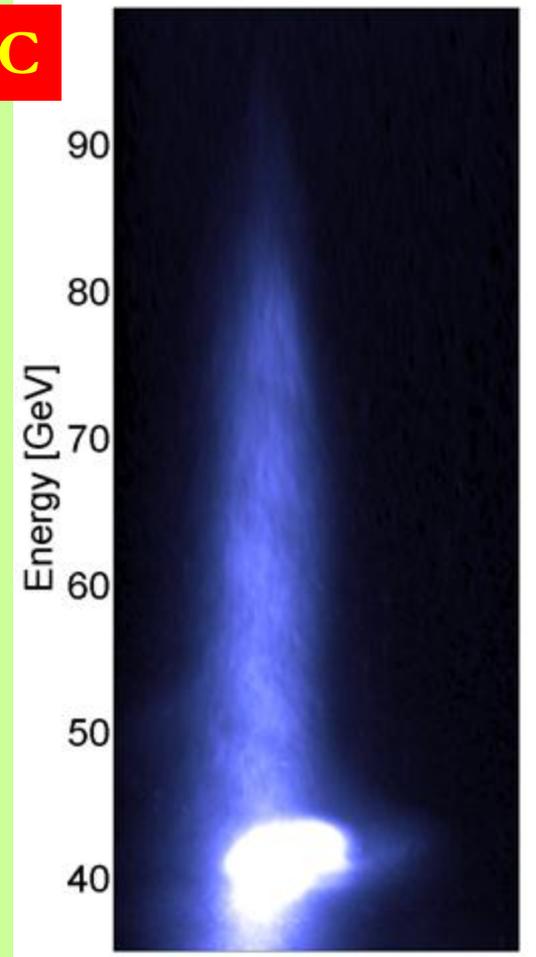
# Compact Acceleration

*50 GeV/meter has been achieved*

**FFTB at SLAC**

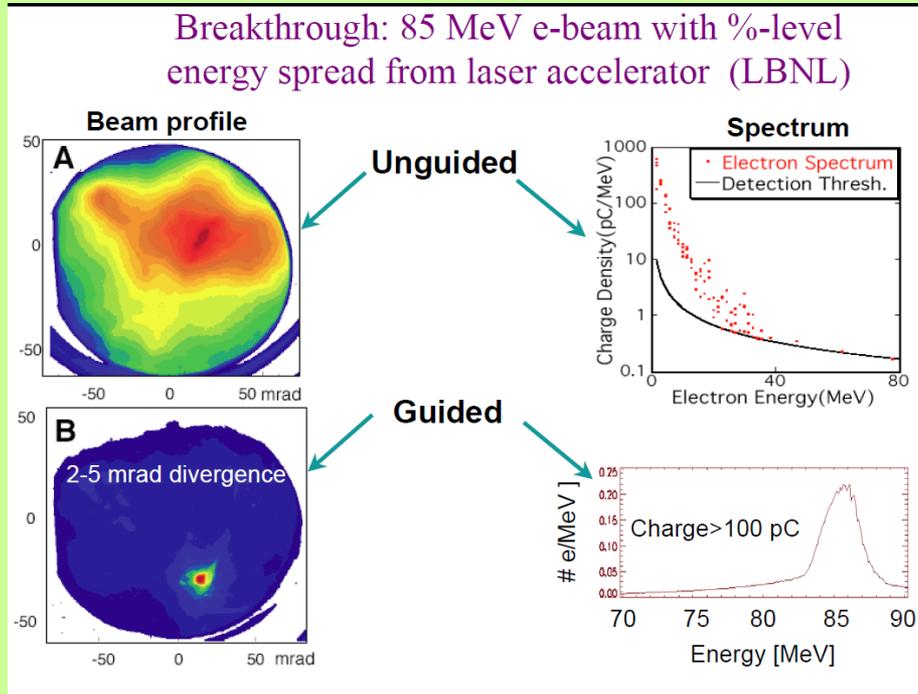


**FACET at SLAC**

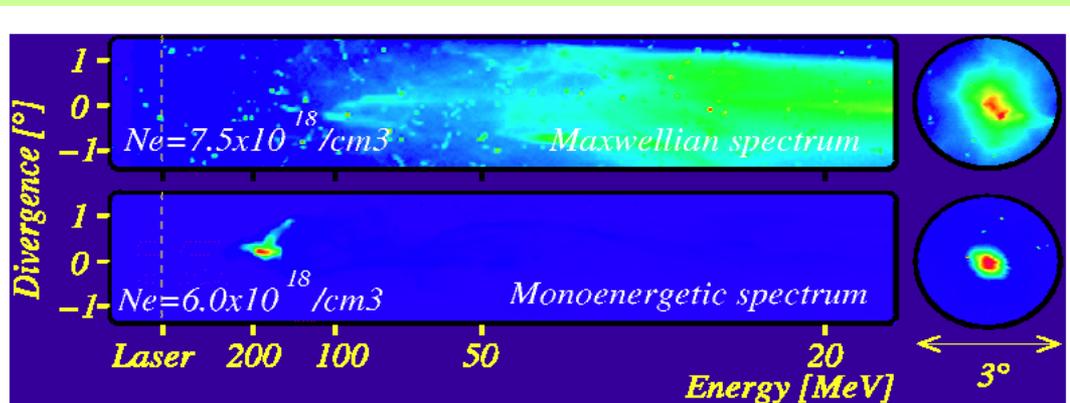


# Controlling the beams

LBNL



Reducing energy spread to ~ percent level



Reducing angular divergence (< 1 degree)

# Advanced Accelerator R&D projects

## **FACET**

### **Location**

**SLAC National Accelerator Laboratory**

**Creates electron wake using  
an electron beam in a plasma**

**Based on previous experiments that  
doubled the energies of a few  
electrons—from 42 billion to 85 billion  
electron volts—in 84 centimeters**

### **Stimulus funds**

**\$14.5 million**

**Next step/midterm goal:  
is to create tighter electron bunches  
and accelerate them from 23 billion to  
46 billion electronvolts in 40  
centimeters**

## **BELLA**

### **Location**

**Lawrence Berkeley National Laboratory**

**Creates electron wake using  
a laser beam in a plasma**

**Based on previous experiments that  
accelerated tightly packed electron  
bunches—the kind needed for physics  
experiments—from zero to 1 billion  
electron volts in 3 centimeters**

### **Stimulus funds**

**\$20 million**

**Next step/midterm goal:  
is to accelerate already-tight electron  
bunches to higher energies, from zero to  
10 billion electronvolts in 80  
centimeters**

# Lots of fundamental questions

1. How can we solve the mystery of dark energy?
2. Are there extra dimensions of space?
3. Do all the forces become one?
4. Why are there so many kinds of particles?
5. What is dark matter?

How can we make it in the laboratory?

6. What are neutrinos telling us?
7. How did the universe come to be?
8. What happened to the antimatter?
9. Are there undiscovered principles of nature:

New symmetries, new physical laws?

*from the Quantum Universe*

# Many Opportunities

- Neutrinos

- » Particle physics and astrophysics using a weakly interacting probe



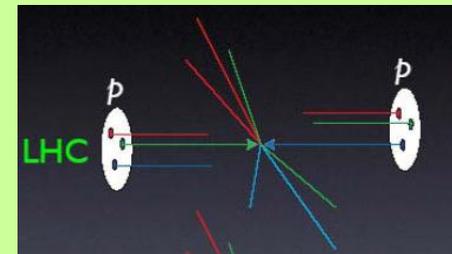
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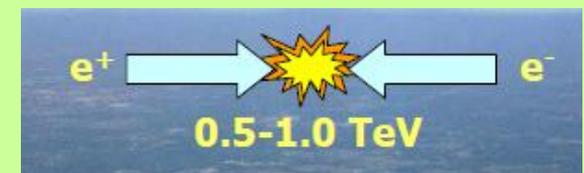
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- » Opening up a new energy frontier (~ 1 TeV scale)



- High Energy Lepton Colliders

- » Precision Physics at the new energy frontier



# and more Opportunities

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- » Fermilab:  $g-2$ ;  $\mu \rightarrow e$ ; etc.



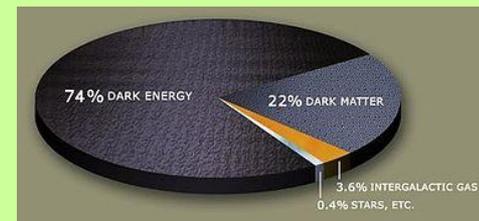
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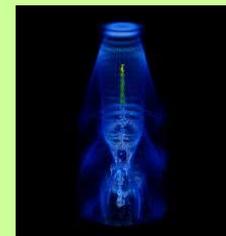
- Dark Matter, Dark Energy

- » Accelerators; underground; space missions

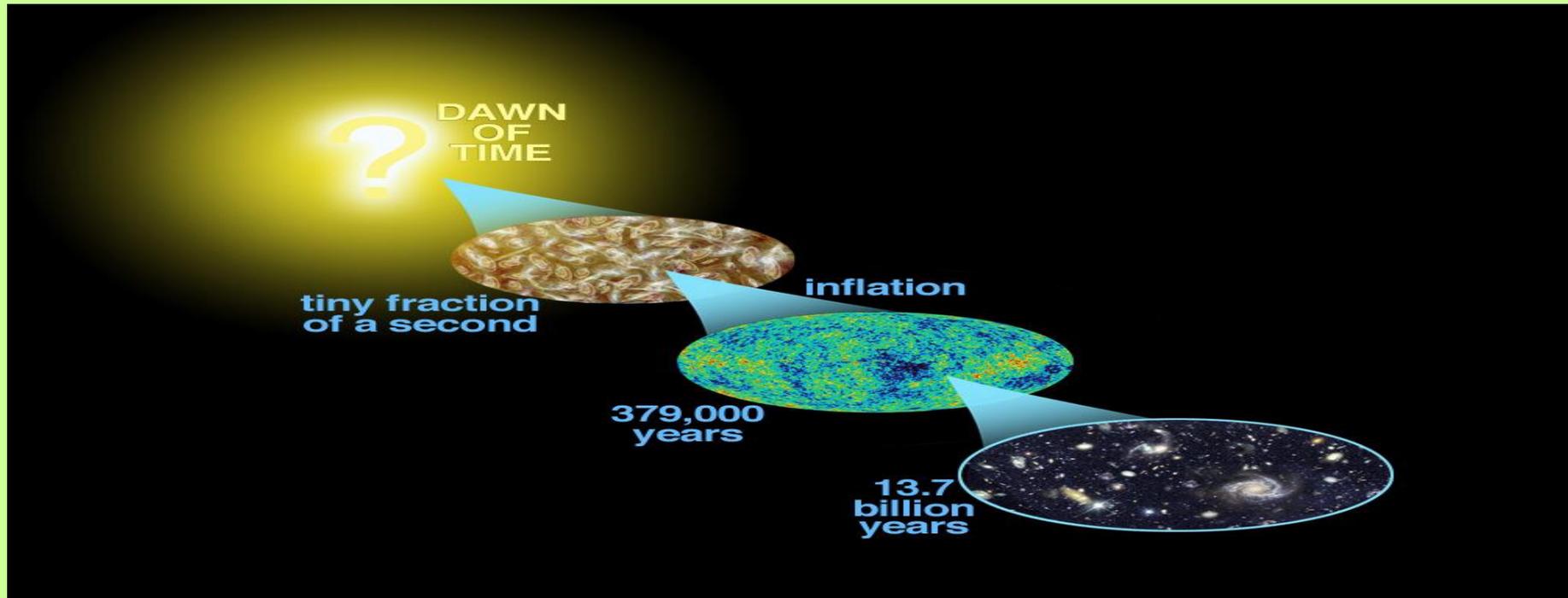


- New Generations of Accelerators

- » Plasma/laser wakefield acceleration



What will we discover by **DPF2021** or **DPF 2031** ??



*“There are more things in heaven and earth,  
Horatio, than are dreamt of in your philosophy”  
(Hamlet, I.5)*