

Short Baseline Neutrino Oscillations

Žarko Pavlović

Los Alamos National Laboratory

Standard Model & Neutrino Oscillations

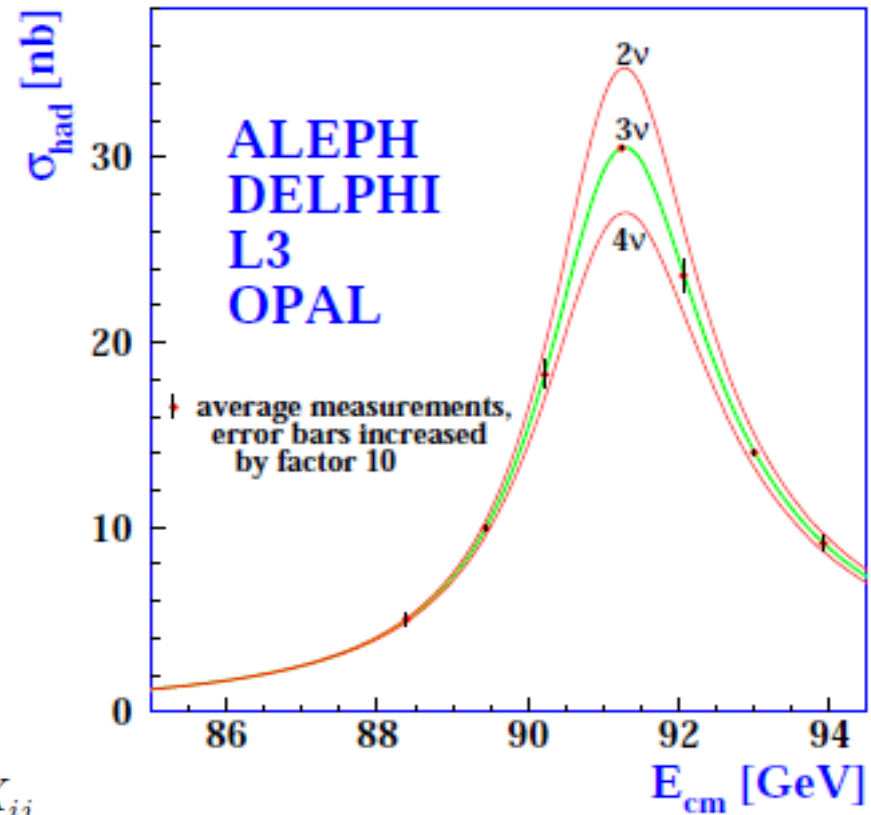
- 3 neutrinos
 - Initially assumed massless
- Mixing matrix:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

- Oscillation Probability:

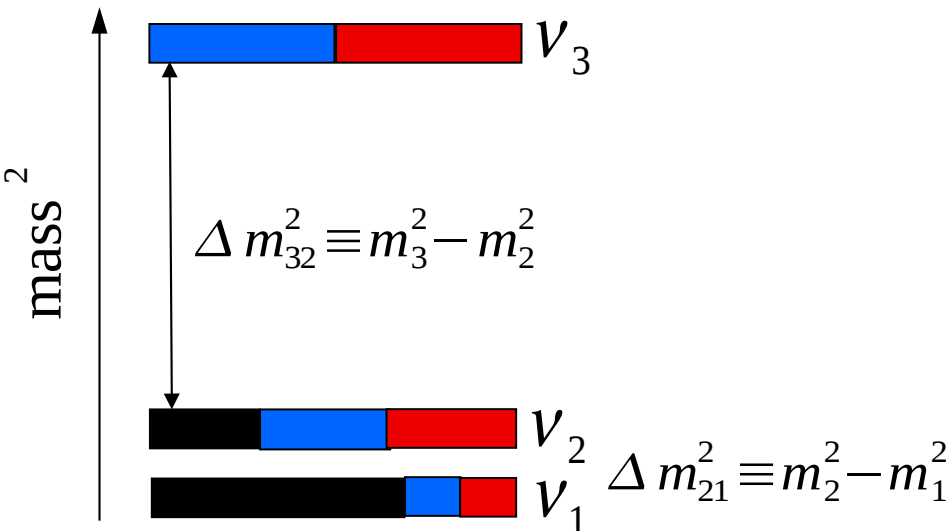
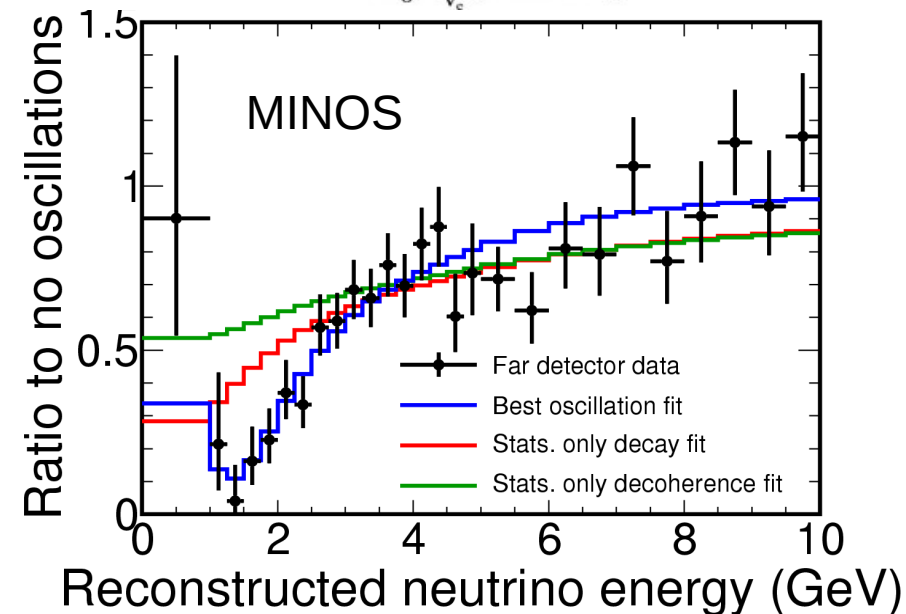
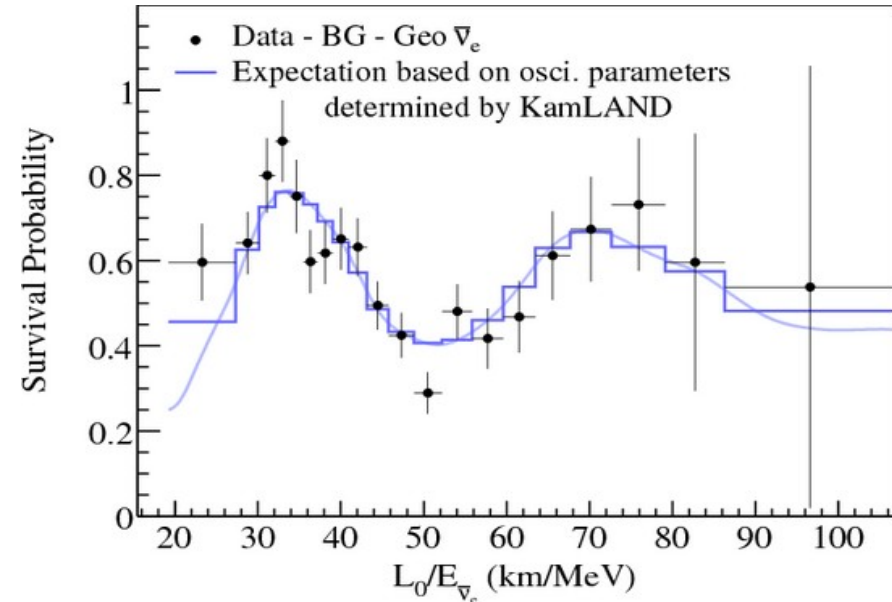
$$P_{\alpha\beta} = |\langle \nu_\beta | \nu_\alpha(t) \rangle|^2 = \delta_{\alpha\beta} - 4 \sum_{i < j}^n \text{Re}[U_{\alpha i} U_{\beta i}^* U_{\alpha j}^* U_{\beta j}] \sin^2 X_{ij} + 2 \sum_{i < j}^n \text{Im}[U_{\alpha i} U_{\beta i}^* U_{\alpha j}^* U_{\beta j}] \sin 2X_{ij}$$

$$X_{ij} = \frac{(m_i^2 - m_j^2)L}{4E} = 1.27 \frac{\Delta m_{ij}^2}{\text{eV}^2} \frac{L/E}{\text{m/MeV}}$$



Neutrino Oscillations

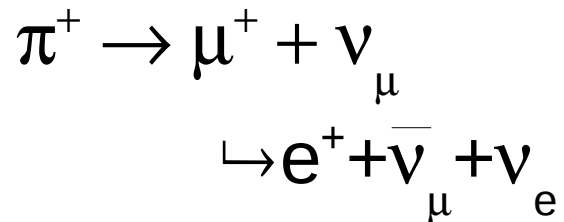
- Lot of experimental evidence
- L/E dependence
- Precise measurement of atmospheric and solar Δm^2



LSND

- Evidence for oscillations at higher Δm^2 than atmospheric and solar

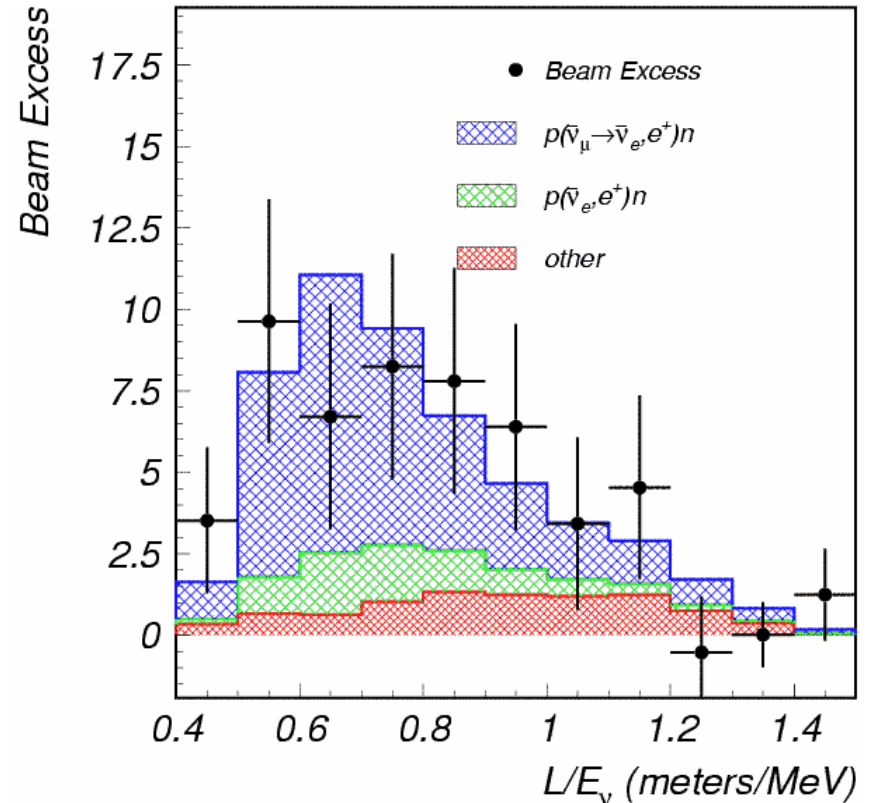
- Stopped pion beam



- Excess of $\bar{\nu}_e$ in $\bar{\nu}_{\mu}$ beam

- $\bar{\nu}_e$ signature: Cherenkov light from e^+ with delayed n-capture

- Excess = $87.9 \pm 22.4 \pm 6$ (3.8σ)

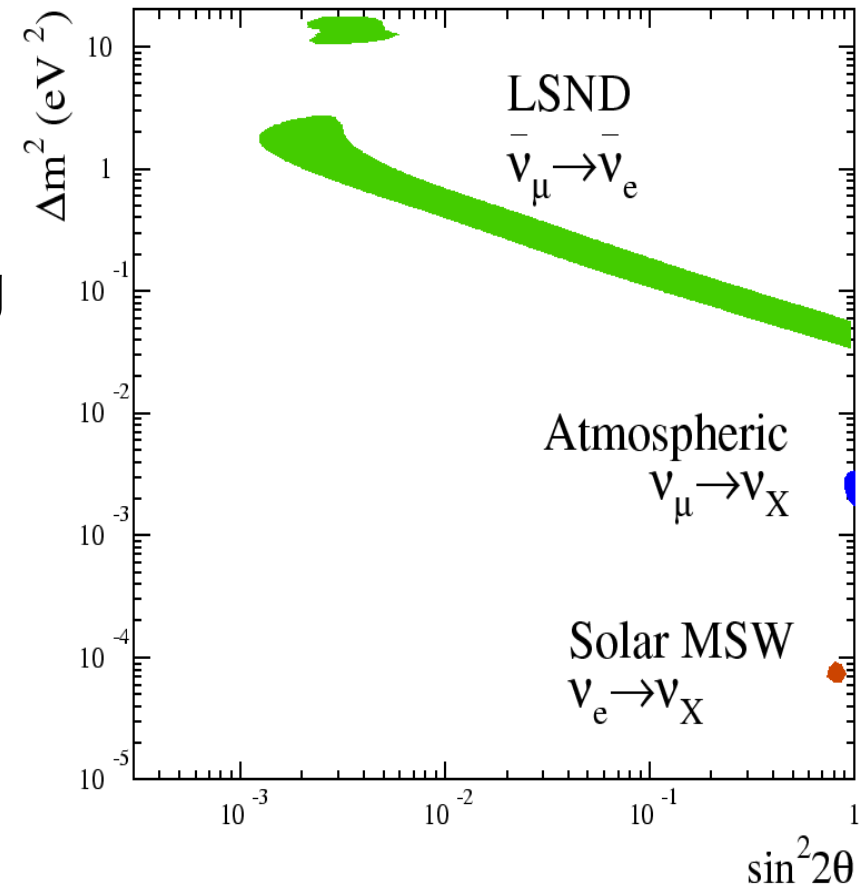
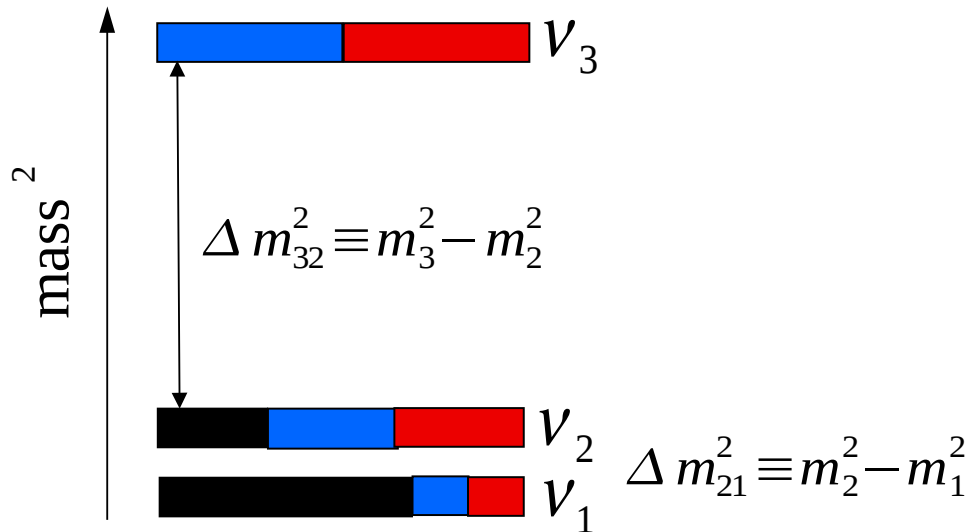


LSND signal

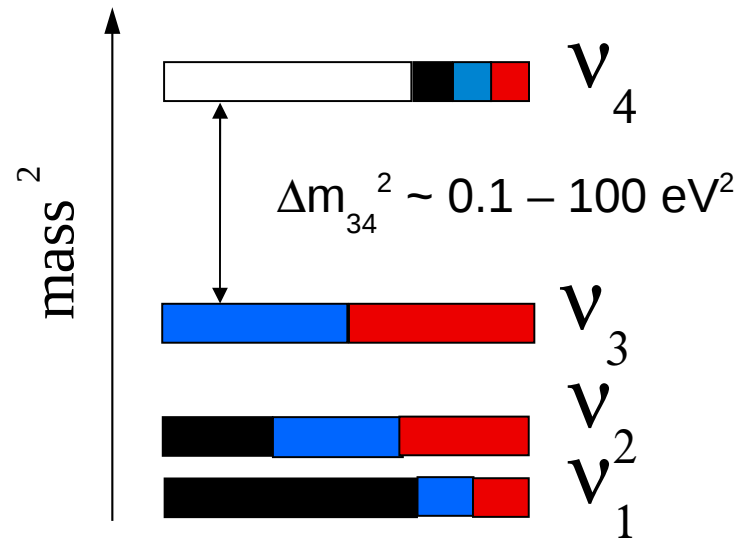
- Assuming two neutrino oscillations

$$\begin{aligned}
 P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) &= \sin^2(2\theta) \sin^2\left(\frac{1.27 L \Delta m^2}{E}\right) \\
 &= 0.245 \pm 0.067 \pm 0.045 \%
 \end{aligned}$$

- Can't reconcile LSND result with atmospheric and solar neutrino using only 3 Standard Model neutrinos – only two independent mass splittings



Sterile neutrinos



- 3 active neutrinos + 1 sterile neutrino
- Sterile neutrino has no Standard Model interactions
- Active neutrinos can oscillate into sterile
- 3 parameters relevant for short baseline exp.: Δm_{41}^2 ,

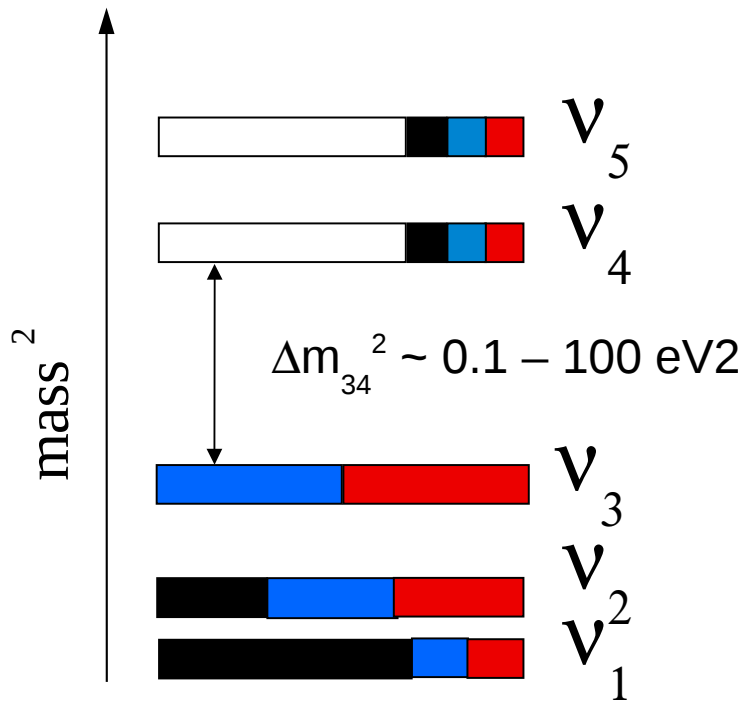
$$|U_{e4}| \text{ and } |U_{\mu 4}|$$

$$P(\nu_\mu \rightarrow \nu_e) = 4|U_{e4}|^2|U_{\mu 4}|^2 \sin^2(1.27 \Delta m_{41}^2 L/E)$$

$$P(\nu_e \rightarrow \nu_e) = 1 - 4|U_{e4}|^2(1 - |U_{e4}|^2) \sin^2(1.27 \Delta m_{41}^2 L/E)$$

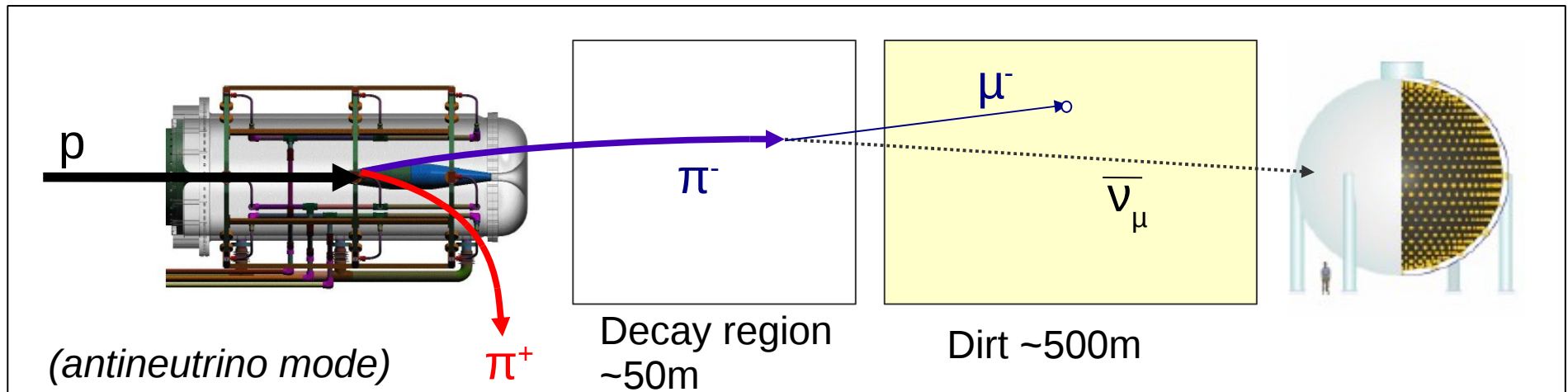
$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - 4|U_{\mu 4}|^2(1 - |U_{\mu 4}|^2) \sin^2(1.27 \Delta m_{41}^2 L/E)$$

More sterile neutrinos



- Next minimal extension 3+2 models
- Favored by fits to world data
- Model allows CP violation
 - $\nu_\mu \rightarrow \nu_e \neq \bar{\nu}_\mu \rightarrow \bar{\nu}_e$

MiniBooNE experiment



- Similar L/E as LSND
 - MiniBooNE $\sim 500\text{m}/\sim 500\text{MeV}$
 - LSND $\sim 30\text{m}/\sim 30\text{MeV}$
- Horn focused neutrino beam ($p+\text{Be}$)
 - Horn polarity \rightarrow neutrino or anti-neutrino mode
- 800t mineral oil Cherenkov detector

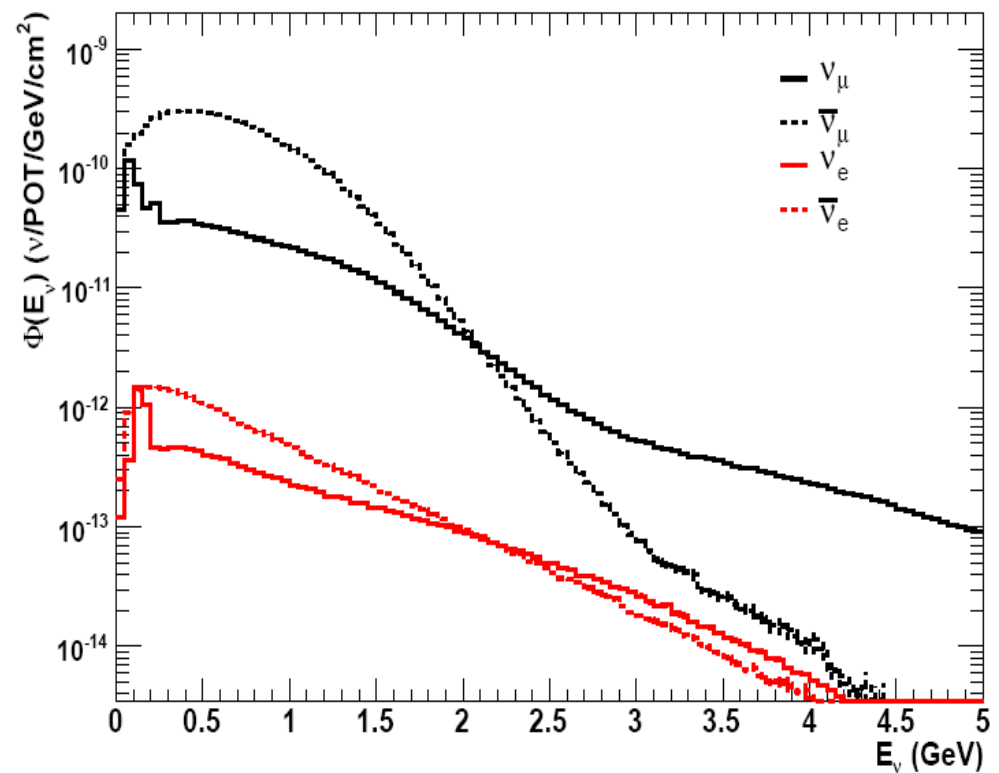
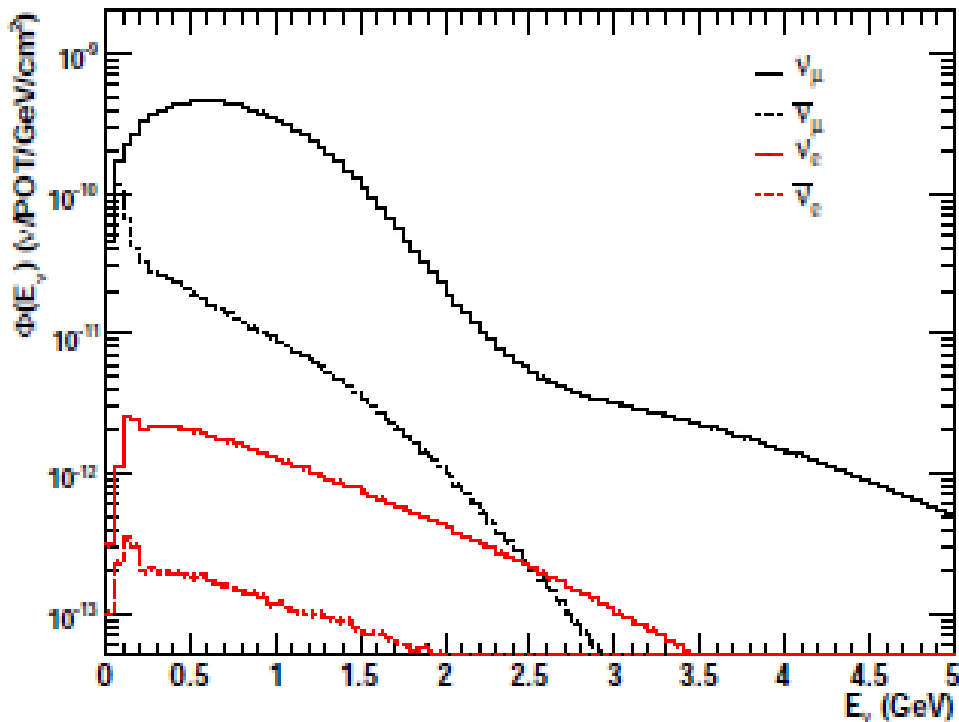
Predicted neutrino flux

- Neutrino mode

| | |
|-----------------------|-------|
| ν_μ | 93.6% |
| $\bar{\nu}_\mu$ | 5.8% |
| $\nu_e + \bar{\nu}_e$ | 0.6% |

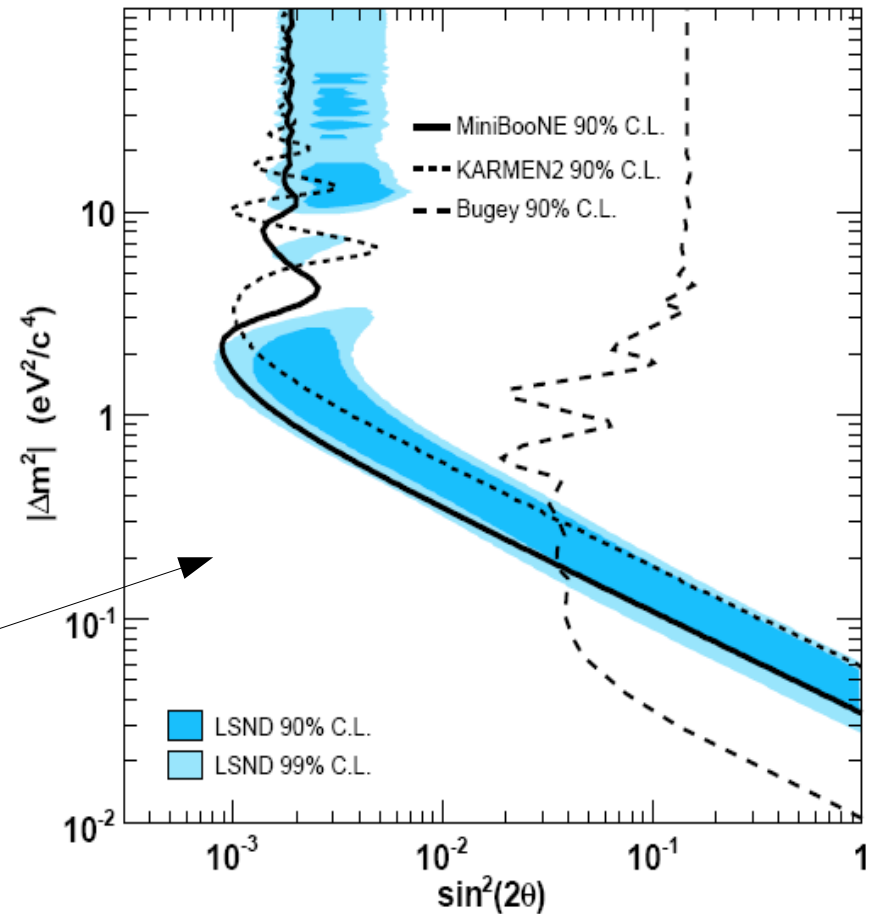
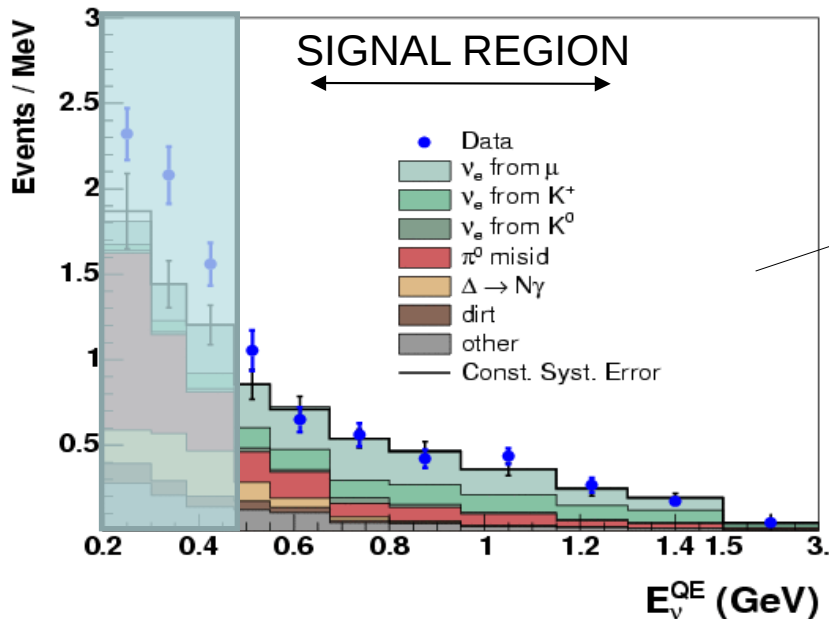
- Anti-neutrino mode

| | |
|-----------------------|-------|
| ν_μ | 15.7% |
| $\bar{\nu}_\mu$ | 83.7% |
| $\nu_e + \bar{\nu}_e$ | 0.6% |



MiniBooNE neutrino result

- 6.5e20 Protons On Target (POT)
- No excess of events in fit region ($E > 475$ MeV)
- Disfavors LSND 2 ν oscillation explanation for neutrinos (assuming no CP or CPT violation)



Phys. Rev. Lett. 98, 231801 (2007)

MiniBooNE neutrino result

Excess of events observed at low energy:

$$128.8 \pm 20.4 \pm 38.3 \text{ (} 3.0\sigma \text{)}$$

Shape not consistent with 2ν oscillations

Magnitude consistent with LSND

Anomaly Mediated Neutrino-Photon Interactions at Finite Baryon Density:
Jeffrey A. Harvey, Christopher T. Hill, & Richard J. Hill, arXiv:0708.1281

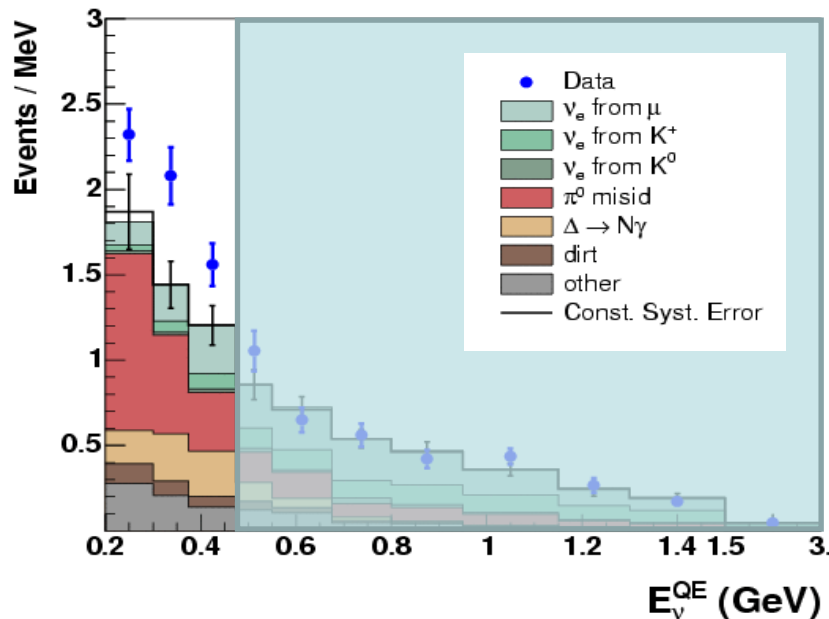
CP-Violation 3+2 Model: Maltoni & Schwetz, arXiv:0705.0107; T. Goldman, G. J. Stephenson Jr., B. H. J. McKellar, Phys. Rev. D75 (2007) 091301.

Extra Dimensions 3+1 Model: Pas, Pakvasa, & Weiler, Phys. Rev. D72 (2005) 095017

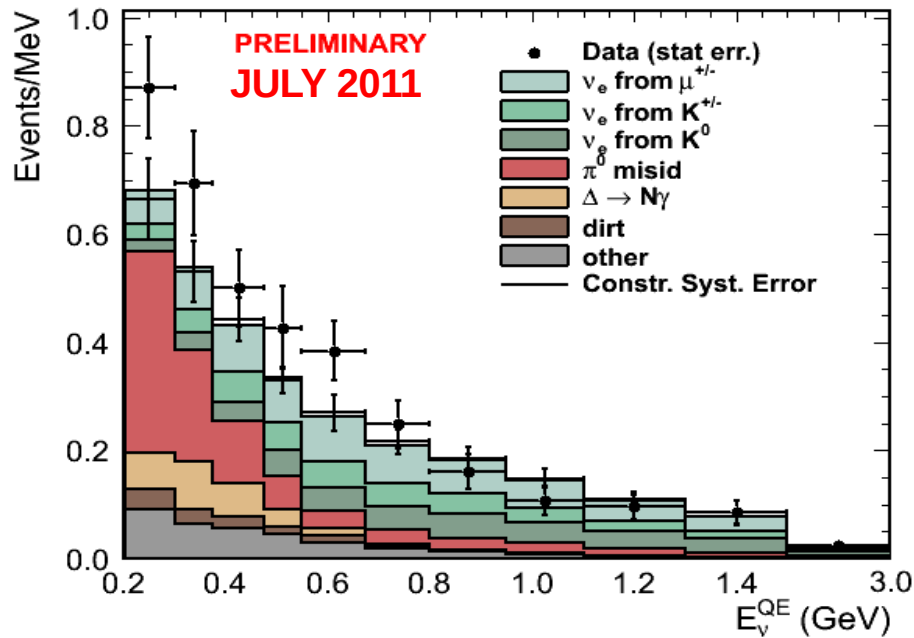
Lorentz Violation: Katori, Kostelecky, & Tayloe, Phys. Rev. D74 (2006) 105009

CPT Violation 3+1 Model: Barger, Marfatia, & Whisnant, Phys. Lett. B576 (2003) 303

New Gauge Boson with Sterile Neutrinos:
Ann E. Nelson & Jonathan Walsh, arXiv:0711.1363



New $\bar{\nu}_e$ appearance results



- Updated analysis with 8.58E20 POT - ~50% more data
- Nearly same analysis as before
 - New constraint on ν_e flux from K^+ decays using SciBooNE + fit to K^+ production global data (1105.2871, accepted by Phys. Rev. D)
- Compared to last analysis:
 - excess of events in 200-475 MeV is more significant
 - excess somewhat reduced in 475-1250 MeV region, but still consistent with LSND

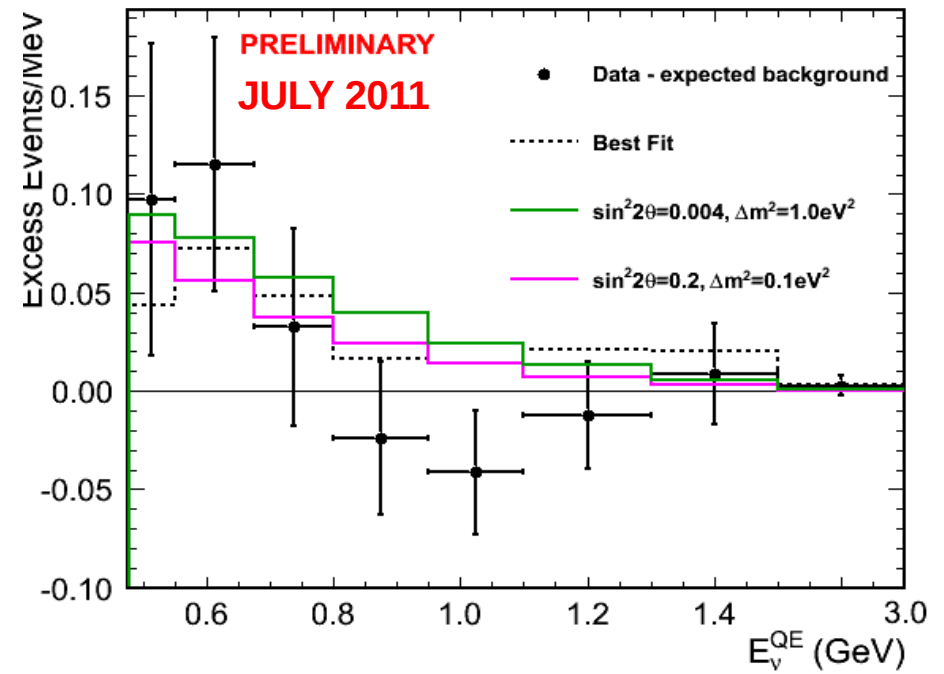
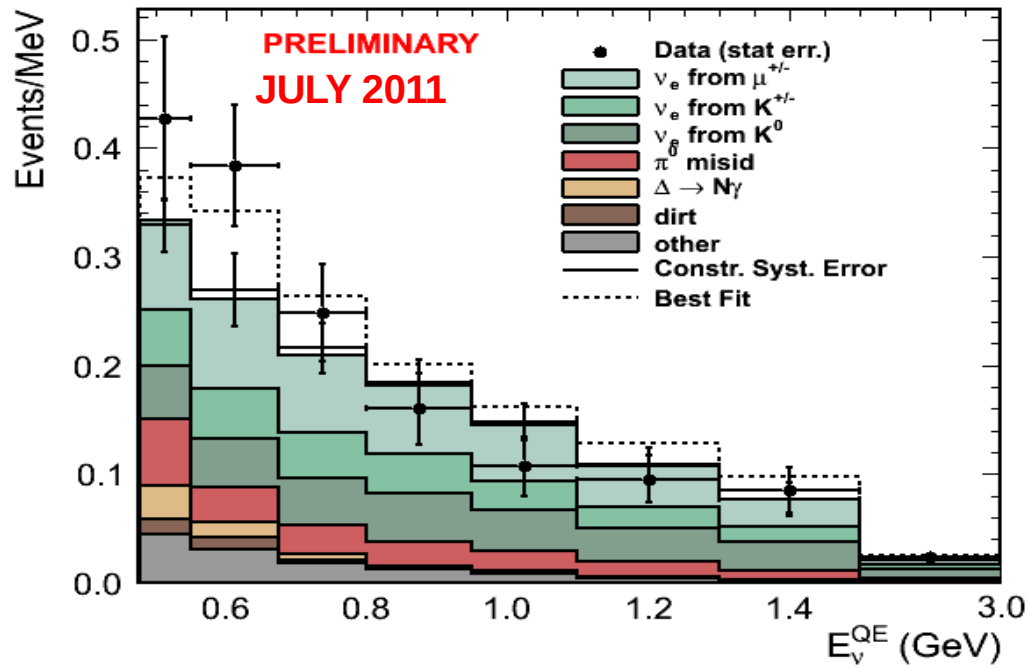
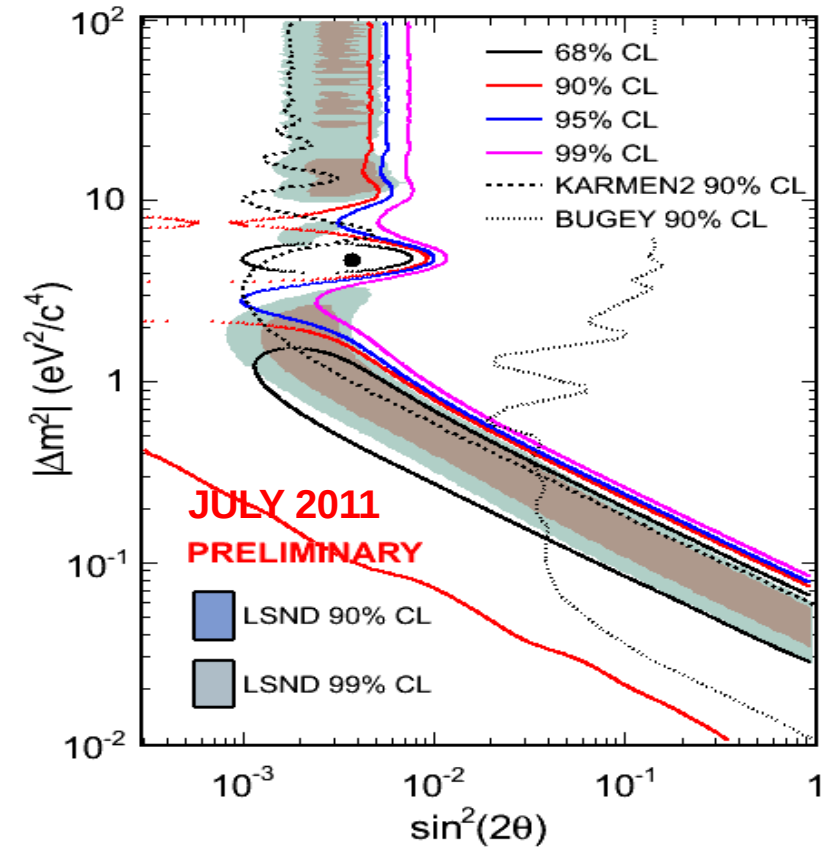
| | 200-475MeV | 475-1250MeV |
|-------------------------------------|------------|-------------|
| Data | 189 | 168 |
| MC | 150.4±18.5 | 151.7±19.4 |
| Excess | 38.6±18.5 | 16.3±19.4 |
| LSND Best Fit | 11.4 | 33.3 |
| Expectation from ν low E excess | 17 | 0 |
| LSND+Low E | 28.4 | 33.3 |

Fit $E > 475 \text{ MeV}$

- 8.58E20 POT
- $E > 475$ is neutrino mode fit region
- Oscillations favored over background only hypotheses at 91.4% CL
- Probability:

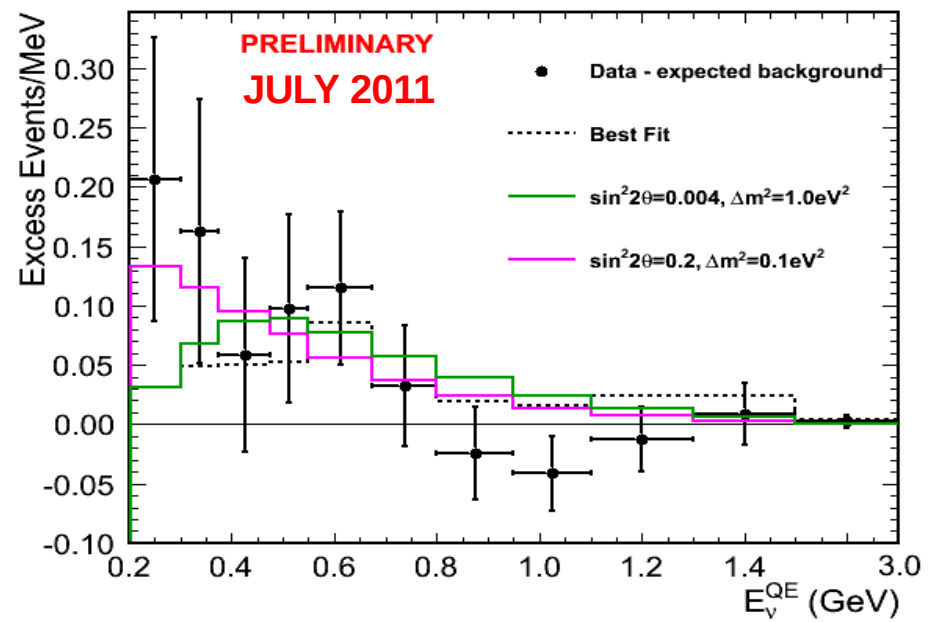
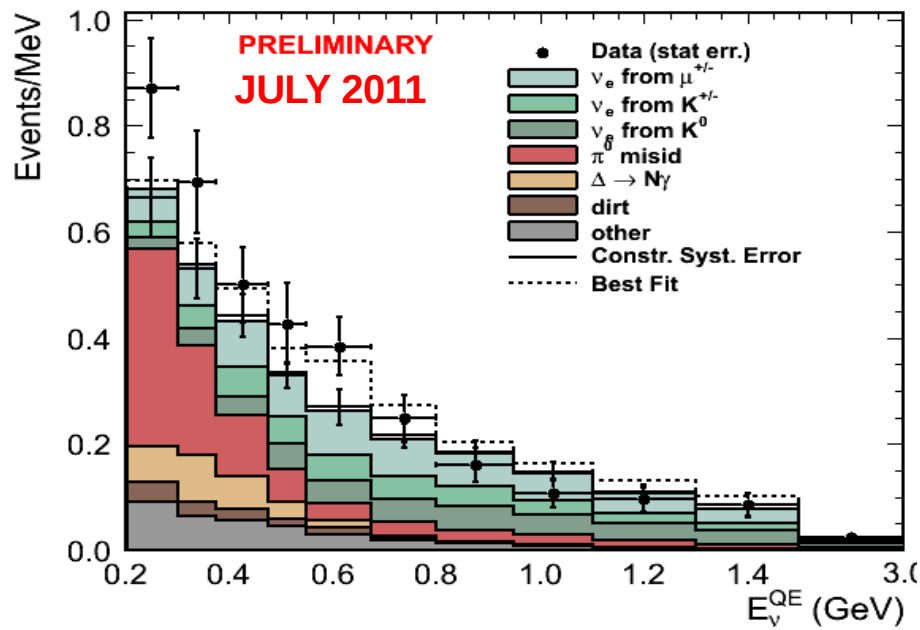
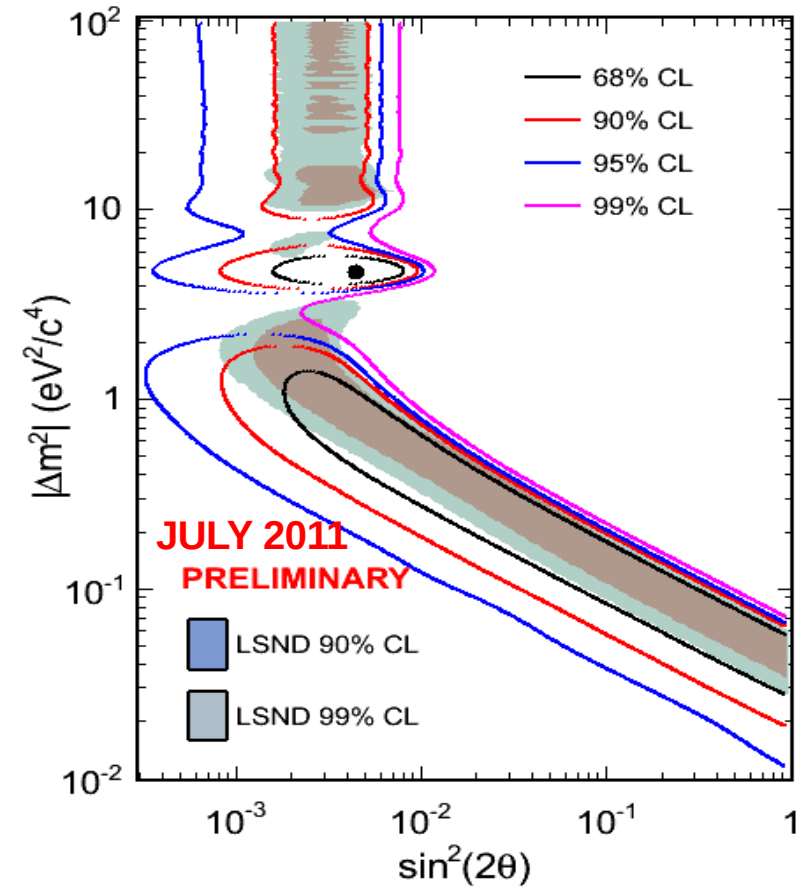
$$P(\text{null}) = 14.9\%$$

$$P(\text{Best Fit}) = 35.5\%$$

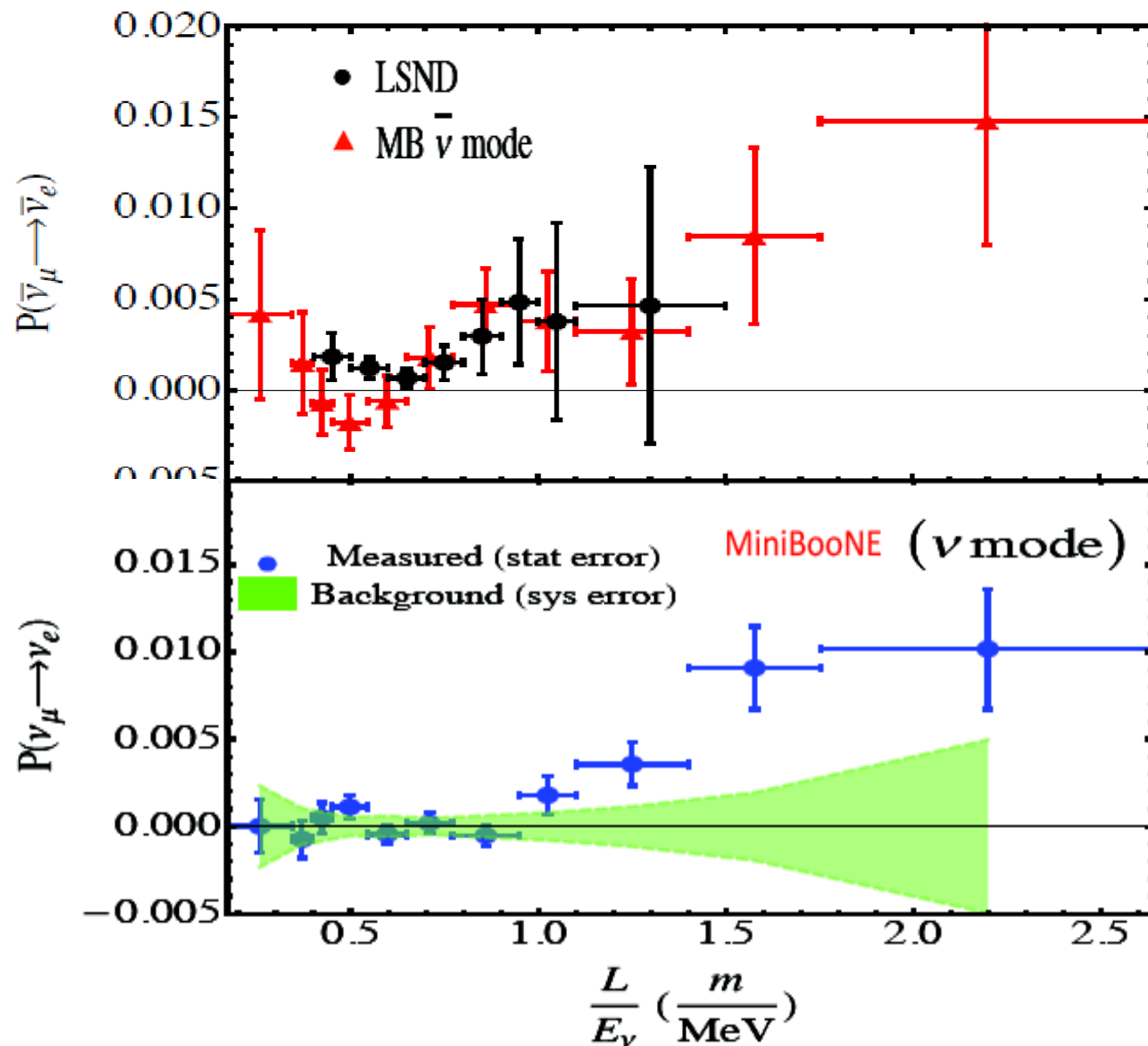


$E > 200 \text{ MeV}$

- $E < 475 \text{ MeV}$:
- Larger background & systematics
- Oscillations favored over background only hypotheses at 97.6% CL (model dependent)
- No assumption made about low energy excess
- $p(\text{null}) = 10.1\%$
 $p(\text{BF}) = 50.7\%$

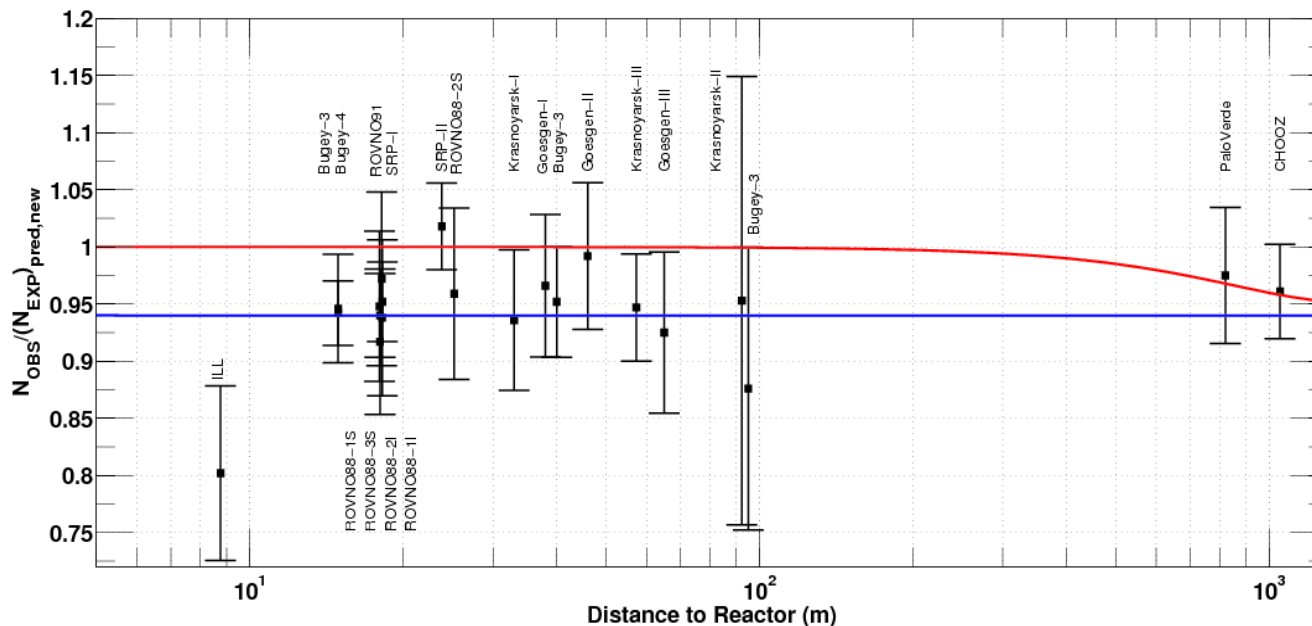


LSND vs MB direct comparison



Reactor antineutrino anomaly

- Recent re-evaluation of reactor fluxes $\rightarrow +3\%$
- Observed/predicted event rate = 0.943 ± 0.023
- Deviation from unity at 98.6% CL



Gallium Anomaly

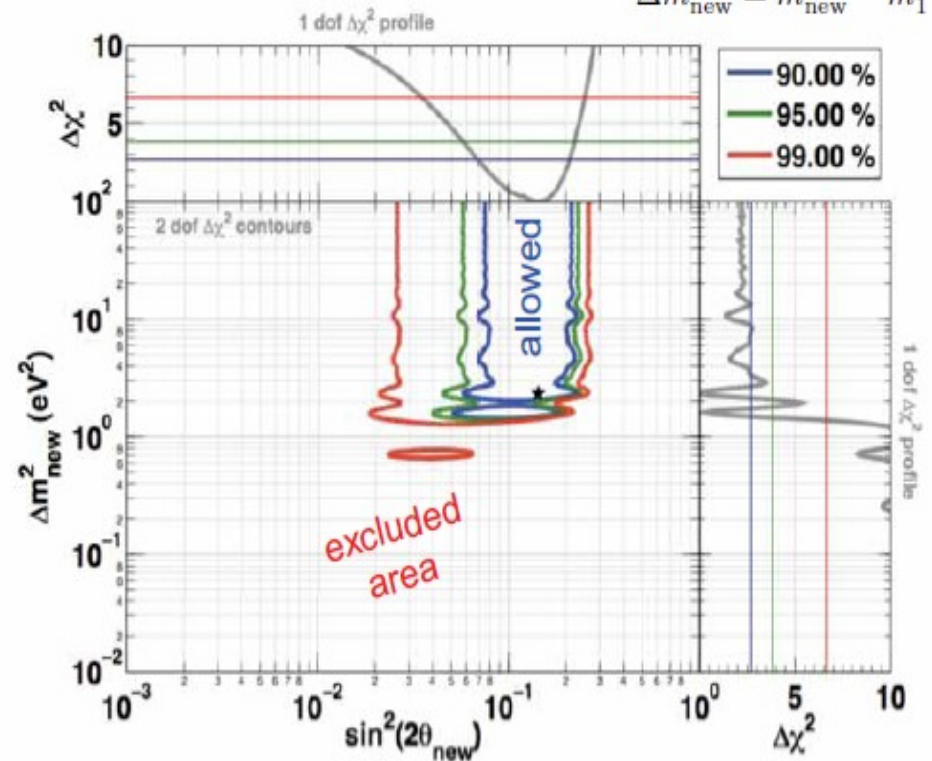
- GALLEX and SAGE calibration runs with intense MCi sources (ν_e)
- Neutrinos detected through radiochemical counting of Ge nuclei: ${}^{71}\text{Ga} + \nu_e \rightarrow {}^{71}\text{Ge} + e^-$
 - 2 runs at GALLEX with ${}^{51}\text{Cr}$ source ($\sim 750\text{keV}$)
 - 1 run at SAGE with ${}^{51}\text{Cr}$ source
 - 1 run at SAGE with ${}^{37}\text{Ar}$ source ($\sim 810\text{keV}$)
- All runs observed deficit of neutrino interactions compared to the expected activity
 - $R = \text{meas}/\text{pred} = 0.86 \pm 0.06$

Sterile neutrinos?

- Reactor data and GALLEX/SAGE
- Data consistent with sterile neutrino oscillations
- Null disfavored at 99.8%
- High Δm^2 excluded by limits on $\nu_{e\mu}$ disappearance using KARMEN and LSND (*arxiv:1106.5552*)

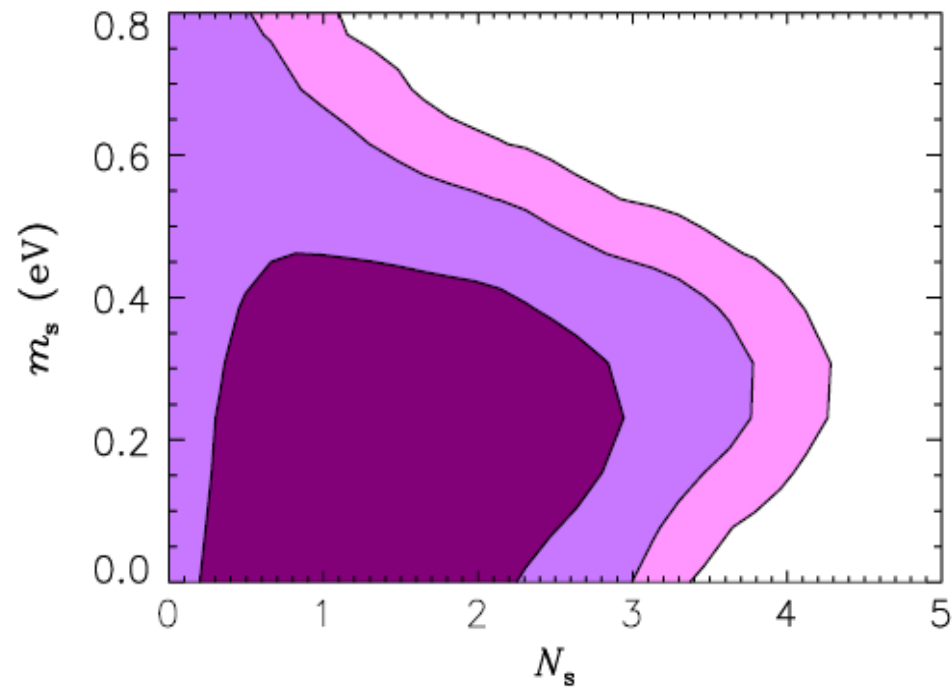
$$\sin^2(2\theta) = 0.14 \pm 0.07$$

$$\Delta m^2 > 1.5 \text{eV}^2 @ 99\% \text{ CL}$$



Cosmology

- Data consistent with extra sterile neutrinos
- N_s = number of thermalized sterile neutrinos



Phys.Rev.Lett. 105, 181301 (2010)

3+N models require large $\bar{\nu}_\mu$ disappearance

- In general:

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) < \frac{1}{4} P(\bar{\nu}_\mu \rightarrow \bar{\nu}_x) P(\bar{\nu}_e \rightarrow \bar{\nu}_x)$$

- From reactor experiments:

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_x) < 8\%$$

- From LSND/MiniBooNE:

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \sim 0.25\%$$

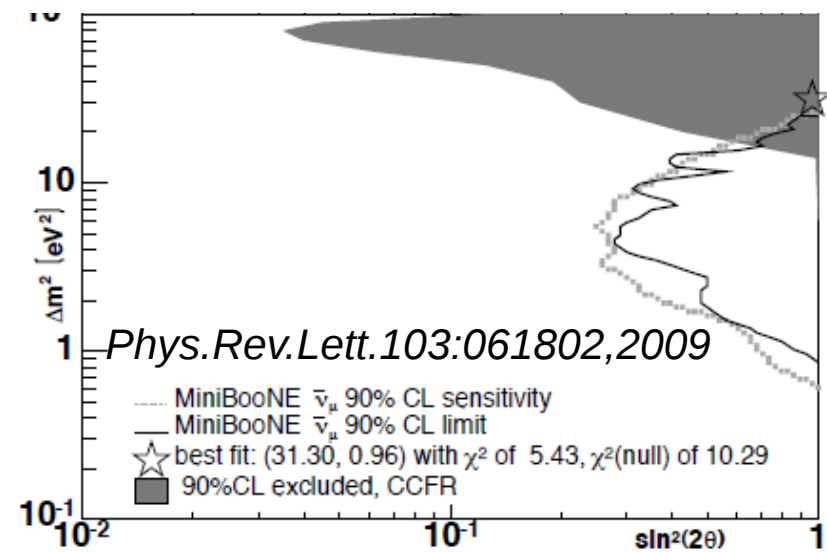
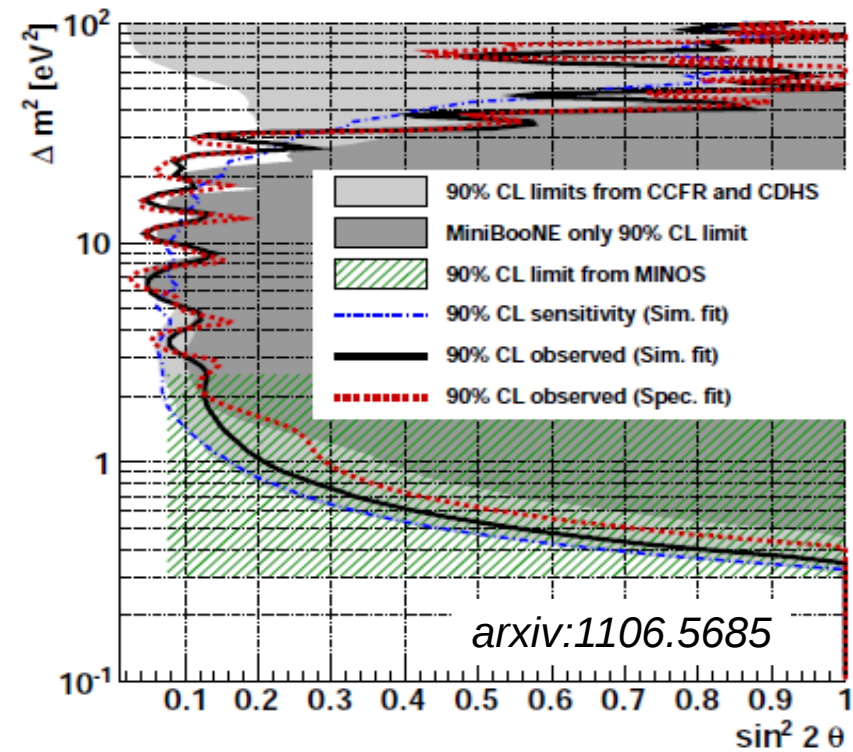
- Therefore:

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_x) > 10\%$$

*Assuming light neutrinos are mostly active and sterile neutrinos are heavy

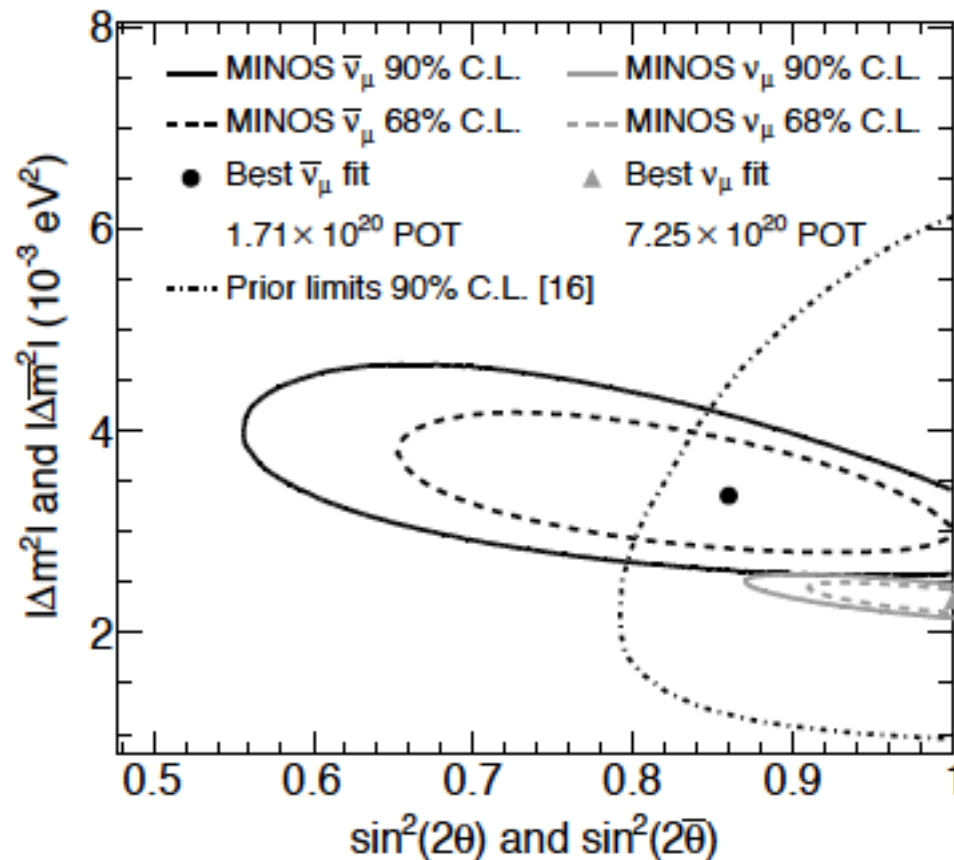
ν_{μ} disappearance

- Provides a constraint on ν_e appearance
- Combined SciBooNE-MiniBooNE analysis in neutrino mode
- MiniBooNE only in anti-neutrino mode



MINOS ν_{μ} vs $\bar{\nu}_{\mu}$

- Hint of NSI or CPT violation?



Future outlook

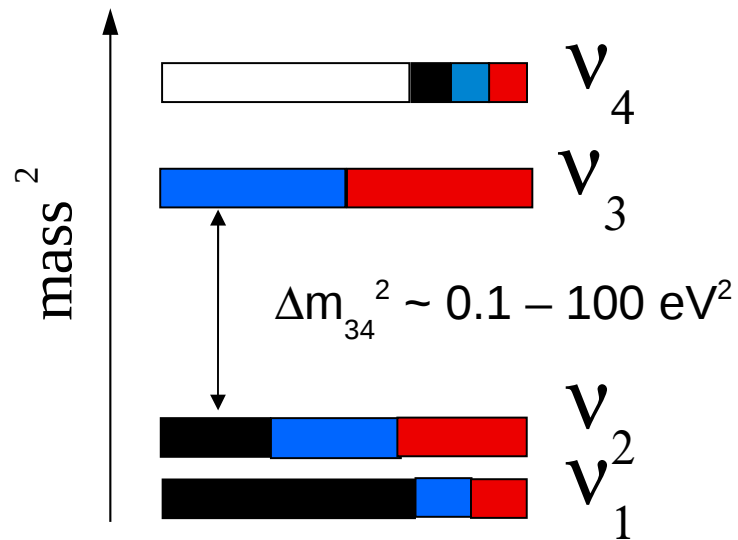
- MiniBooNE – more antineutrino data, approved to run through spring 2012
- Joint MiniBooNE/SciBooNE $\bar{\nu}_\mu$ disappearance
- MicroBooNE - resolve the low energy excess
- MINOS+ - sterile neutrinos, NSI, ...
- BooNE
- Stopped pion source exp. (OscSNS,...)
- Icarus at CERN-PS

Conclusion

- MiniBooNE anti neutrino data consistent with LSND and $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ oscillations at $\Delta m^2 \sim 1 \text{eV}^2$
- Few other hints of sterile neutrinos – reactor anomaly, gallium anomaly, cosmology
- Very active topic:
 - Workshop on Sterile Neutrinos and on the Reactor (anti)-Neutrino Anomaly, TUM, Garching, Feb 8 2011
 - Beyond3nu, Gran Saso, May 3-4 2011
 - Short Baseline Neutrino Workshop, Fermilab, May 12-14 2011
 - Sterile Neutrinos At The Crossroads, Virginia Tech, Sep 26-28 2011

Backup slides

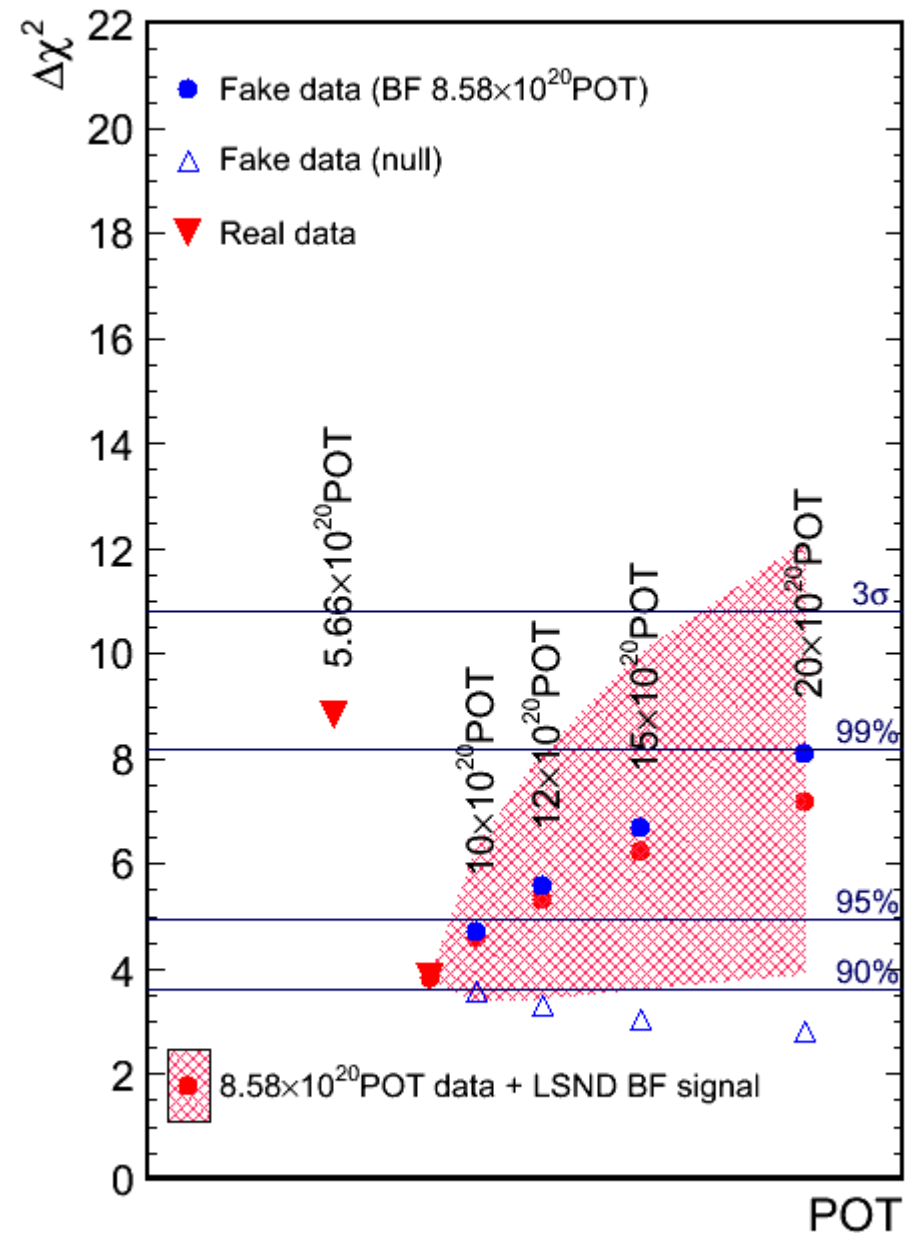
2+2



- Within 2+2 model Sterile neutrino participates in either solar or atmospheric neutrino oscillations (or both)
 - Experiments measuring solar and atmospheric Δm^2 disfavor oscillations to pure sterile neutrinos
- => 2+2 is strongly disfavored

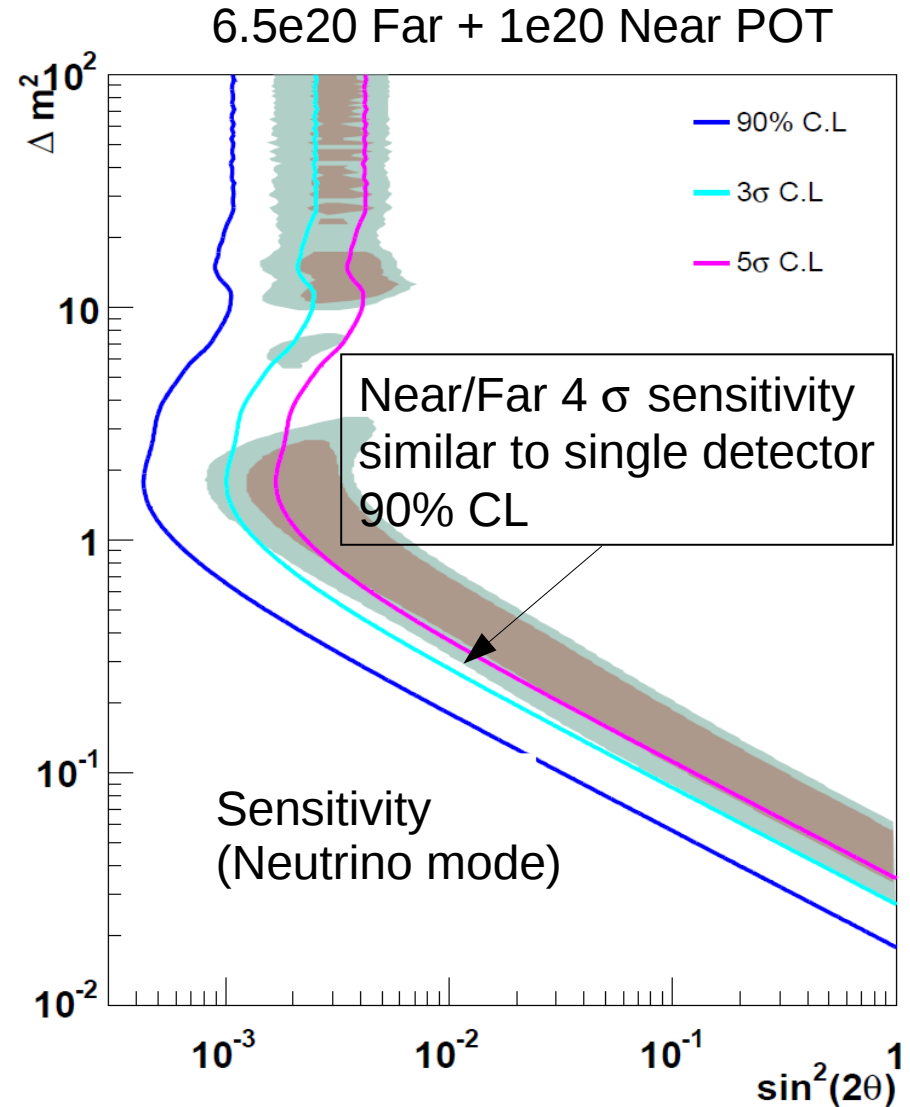
Future sensitivity

- Potential exclusion of null point assuming best fit signal
- Combined analysis of ν_e and $\bar{\nu}_e$



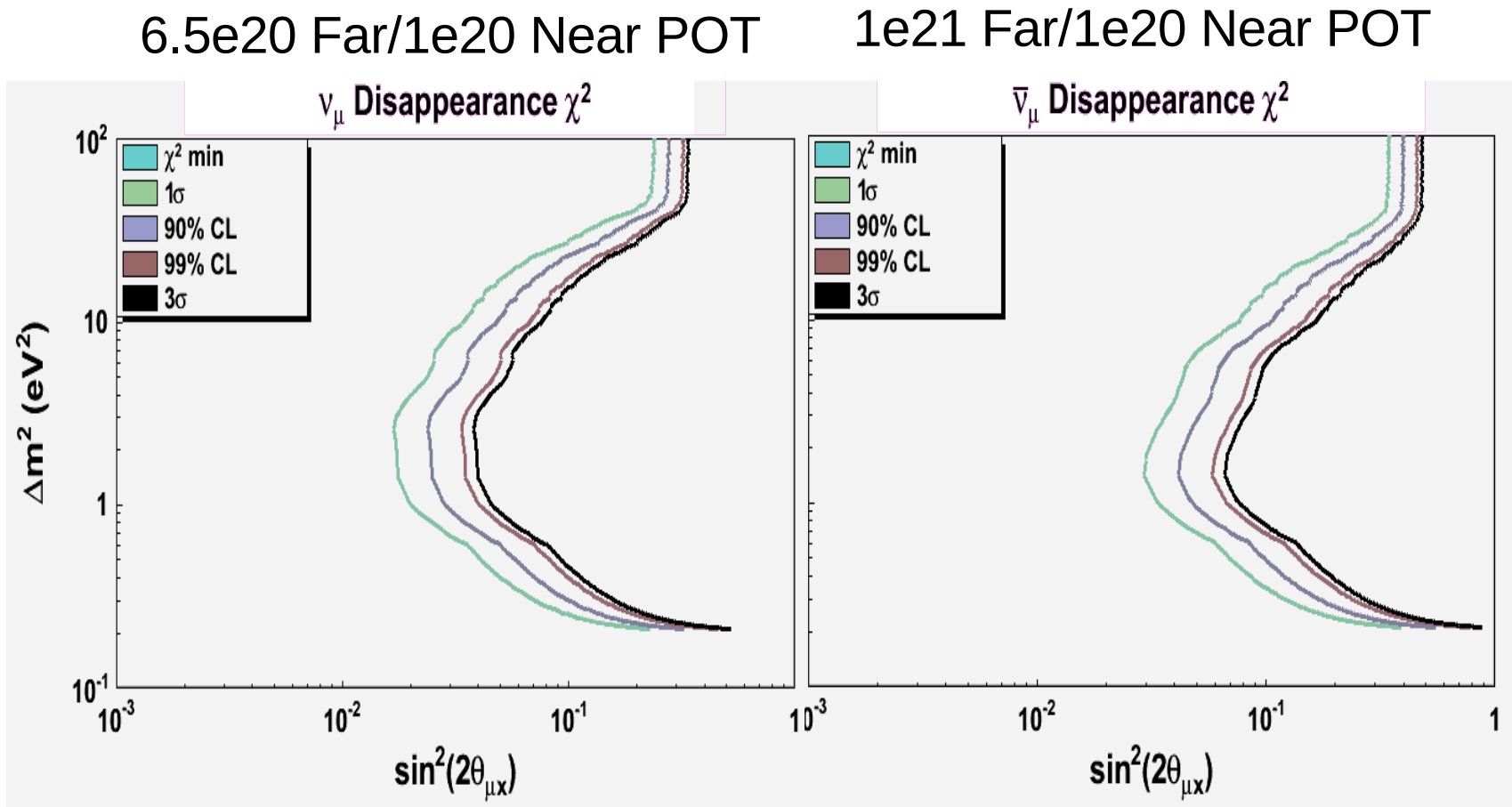
BooNE

- MiniBooNE like detector at 200m
- Flux, cross section and optical model errors cancel in 200m/500m ratio analysis
- Present neutrino low energy excess is 6 sigma statistical; 3 sigma when include systematics
- Study L/E dependence
- Gain statistics quickly, already have far detector data



BooNE

- Better sensitivity to ν_μ ($\bar{\nu}_\mu$) disappearance
- Look for CPT violation ($\nu_\mu \rightarrow \nu_\mu \neq \bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$)



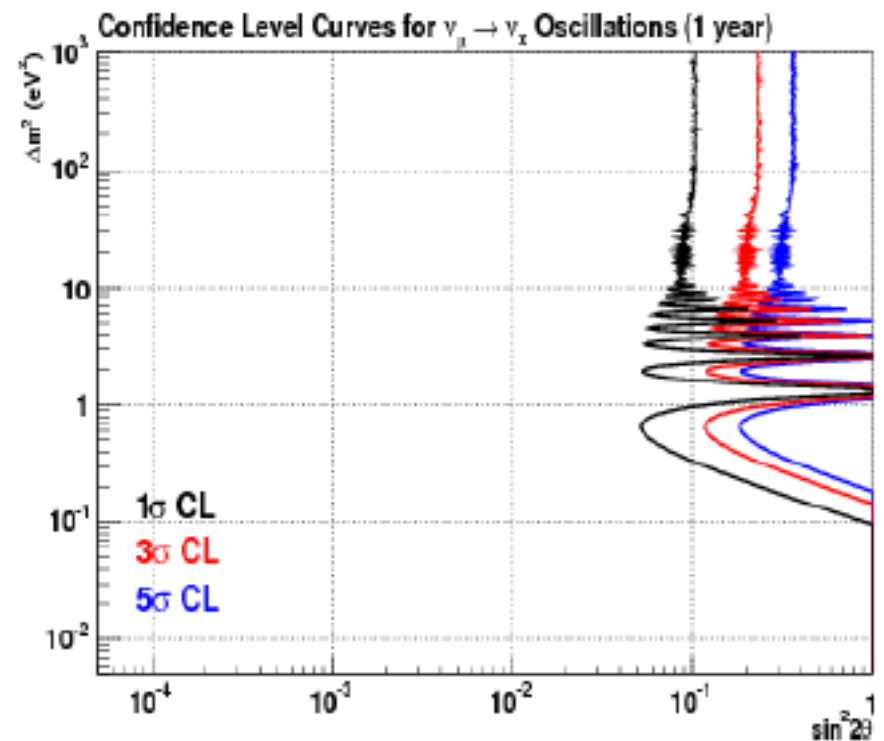
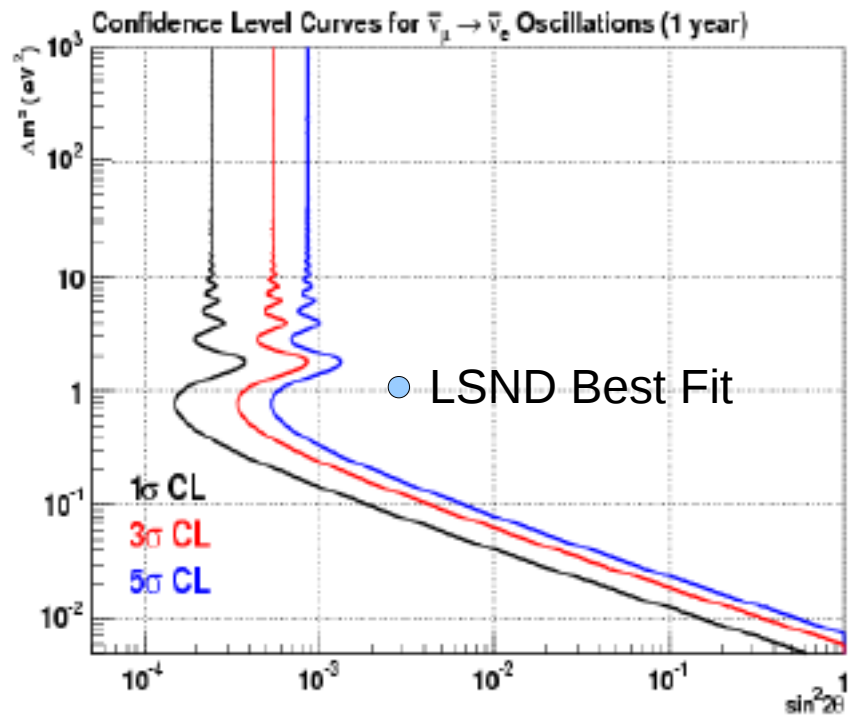
OscSNS

- Spallation neutron source at ORNL
- 1GeV protons on Hg target (1.4MW)
- Free source of neutrinos
- Well understood flux of neutrinos



OscSNS

- $\bar{\nu}_e$ appearance (left) and ν_μ disappearance sensitivity (right) for 1 year of running



LSND $\bar{\nu}_e$ Background Estimates

| Estimate | $\bar{\nu}_e/\bar{\nu}_\mu$ | $\bar{\nu}_e$ Bkgd | LSND Excess |
|-------------------|-----------------------------|--------------------|-----------------|
| LSND Paper | 0.086% | 19.5+-3.9 | 87.9+-22.4+-6.0 |
| Zhemchugov Poster | 0.071% | 16.1+-3.2 | 91.3+-22.4+-5.6 |
| Dydak Seminar | 0.116% | 26.3+-5.3 | 81.1+-22.4+-7.0 |

All $\bar{\nu}_e$ background estimates assume a 20% error. Note that the $\bar{\nu}_e/\bar{\nu}_\mu$ ratio determines the background!

LSND Paper: A. Aguilar et al., Phys. Rev. D 64, 112007 (2001); (uses **MCNP**)

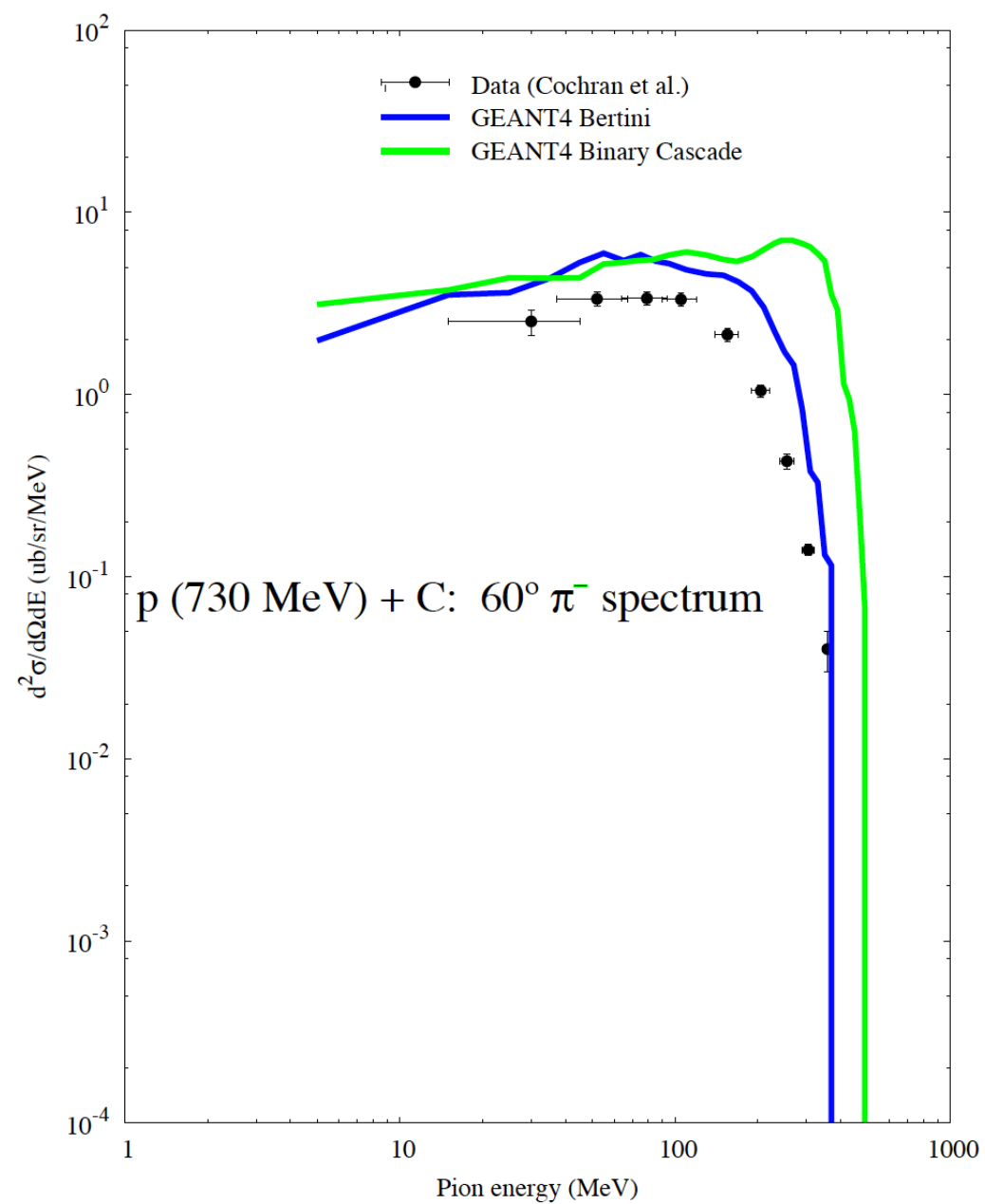
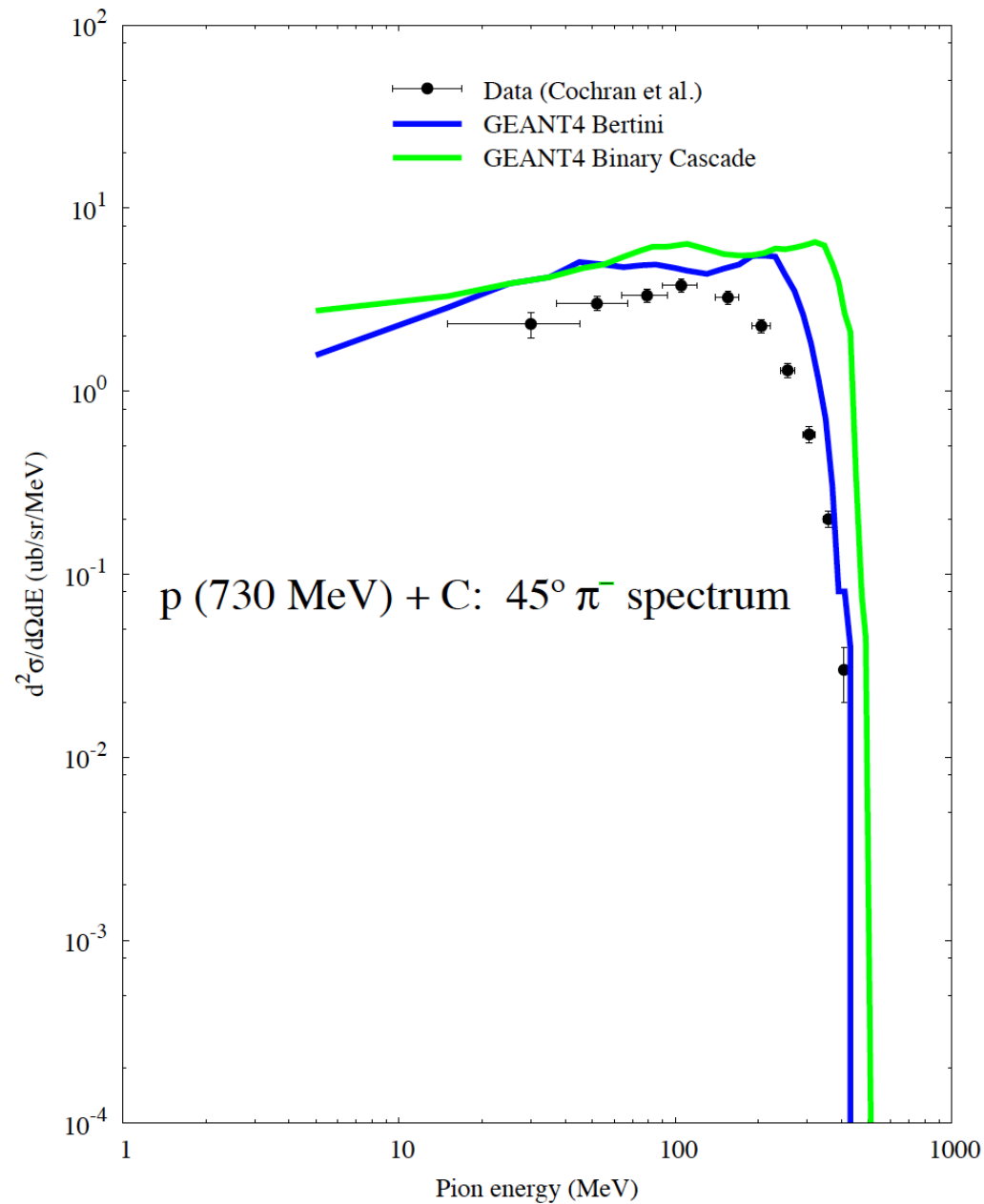
Zhemchugov Poster: **FLUKA** $\bar{\nu}_e/\bar{\nu}_\mu$ ratio presented at the ICHEP 2010 Conference, Paris

Dydak Seminar: **FLUKA** $\bar{\nu}_e/\bar{\nu}_\mu$ ratio presented at FNAL on January 14, 2011

Although the analysis of Zhemchugov, Dydak et al. is not fully understood or endorsed, their $\bar{\nu}_e/\bar{\nu}_\mu$ ratios agree reasonably well with the published LSND results.

Note that LSND measures the correct rate of $\bar{\nu}_\mu p \rightarrow \mu^+ n$ interactions, which confirms the π^- production and background estimates. Note also, that FLUKA & GEANT4 overestimate π^- production at ~ 800 MeV. Note that N_{gs} events are included in the LSND background estimate.

GEANT4 Overestimates π^- Production!



$\nu_e \text{ C} \rightarrow e^- N_{gs}$ Events Do Not Simulate
 $\bar{\nu}_e \text{ p} \rightarrow e^+ n$ Events!

For N_{gs} β decay to be considered a 2.2 MeV γ :
 $\Delta r < 2\text{m}$, $\Delta t < 500\mu\text{s}$, $19 < N_{\text{hits}} < 51$

The number of N_{gs} events with a β that satisfies this initial requirement is approximately: $(600)(1)(1/31.8)(0.05) \sim 1$ event.

The number of N_{gs} events with $R_\gamma > 10 \sim 0.1$ events.

This background is included in the LSND background estimate.