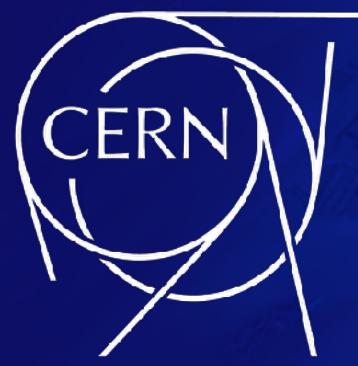


Enabling BSM Studies with Neutrino Data

Joachim Kopp (CERN & JGU Mainz)
NuXTract Workshop • CERN • 2–6 October 2023



Neutrino Physics Beyond the Standard Model

**Sterile Neutrino
Oscillations**

Heavy Neutral Lepton Decay

**Non-standard
Neutrino Interactions**

Axion Portal



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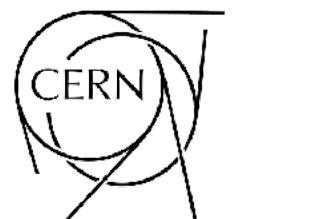
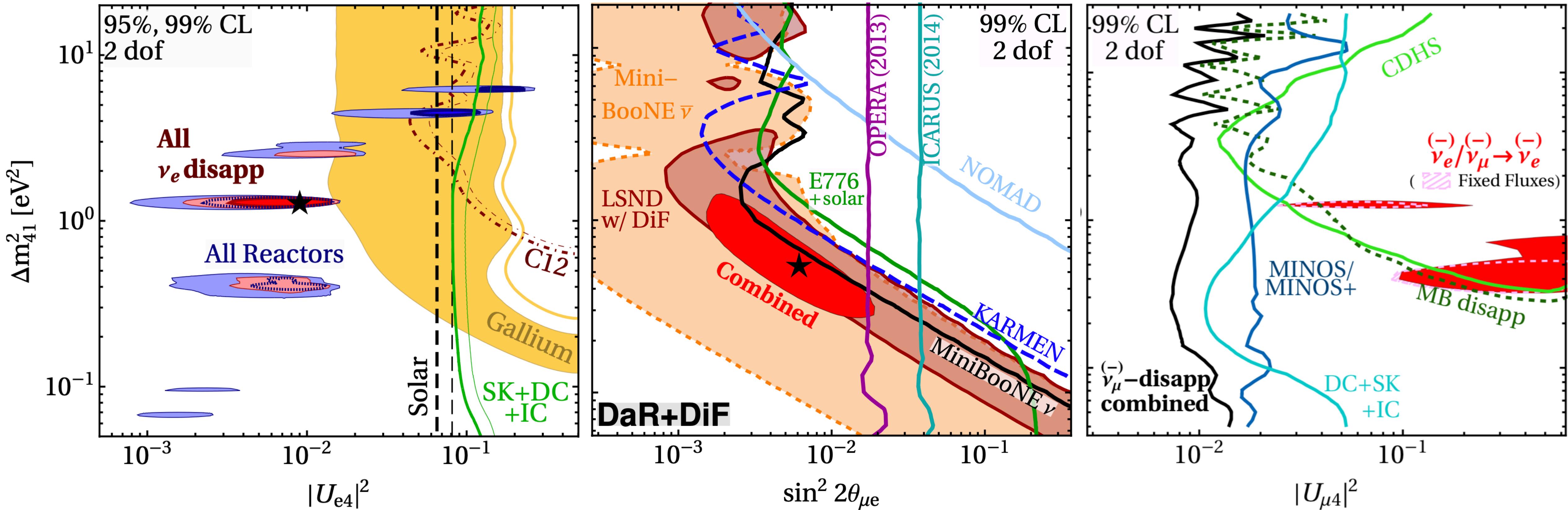
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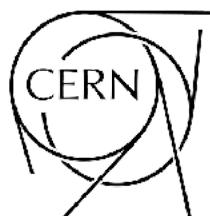
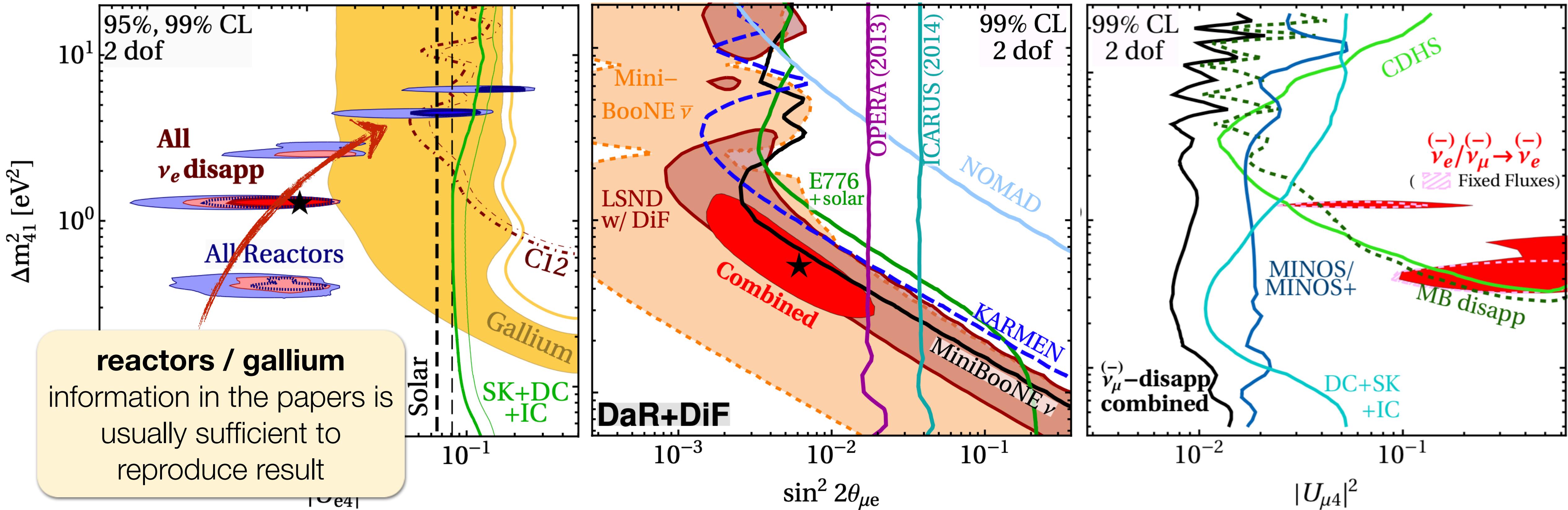
Sterile Neutrino Oscillations – Enabling Global Fits

Dentler et al. 2018



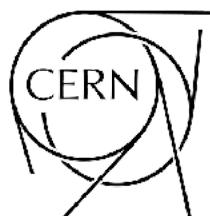
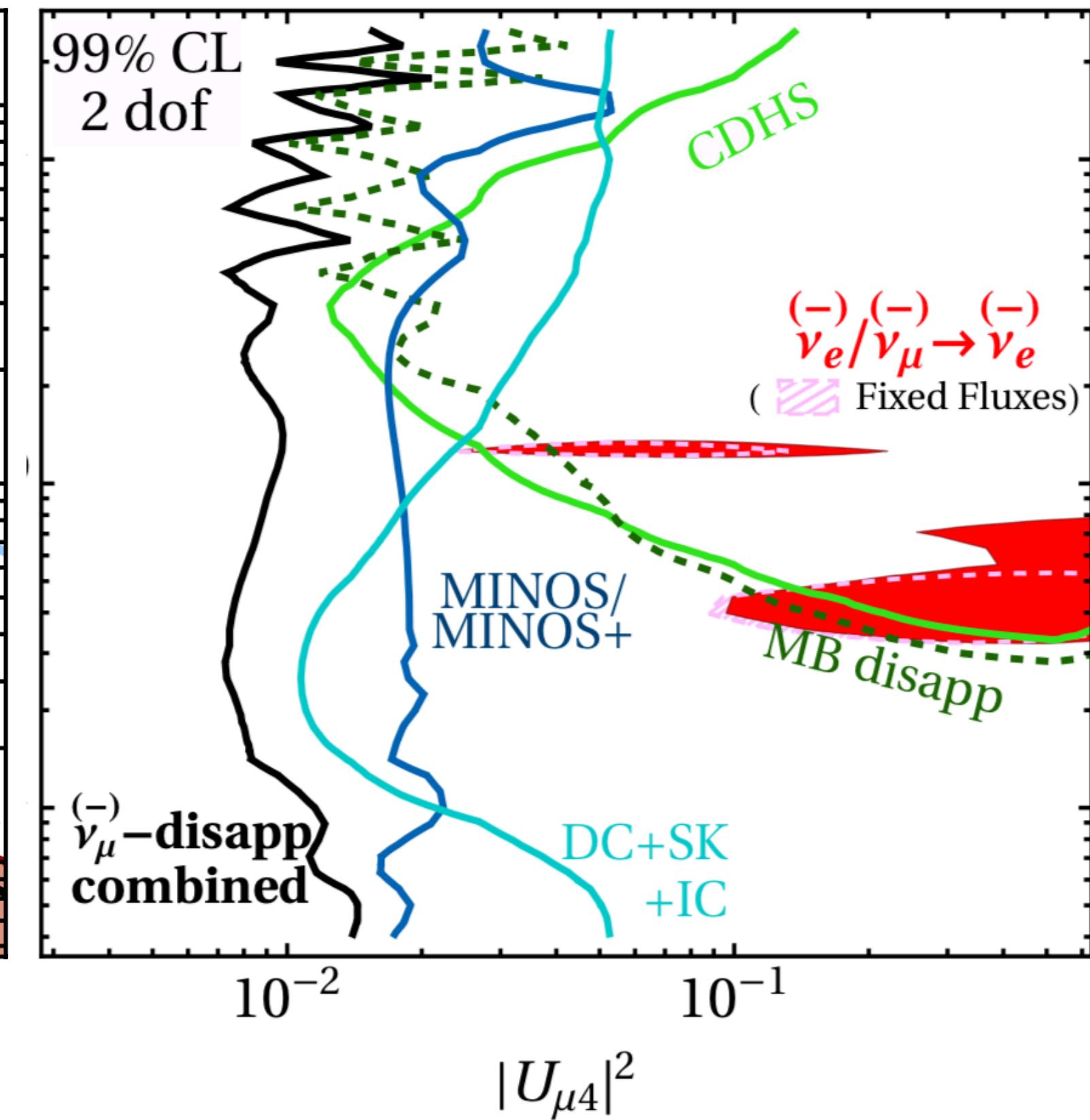
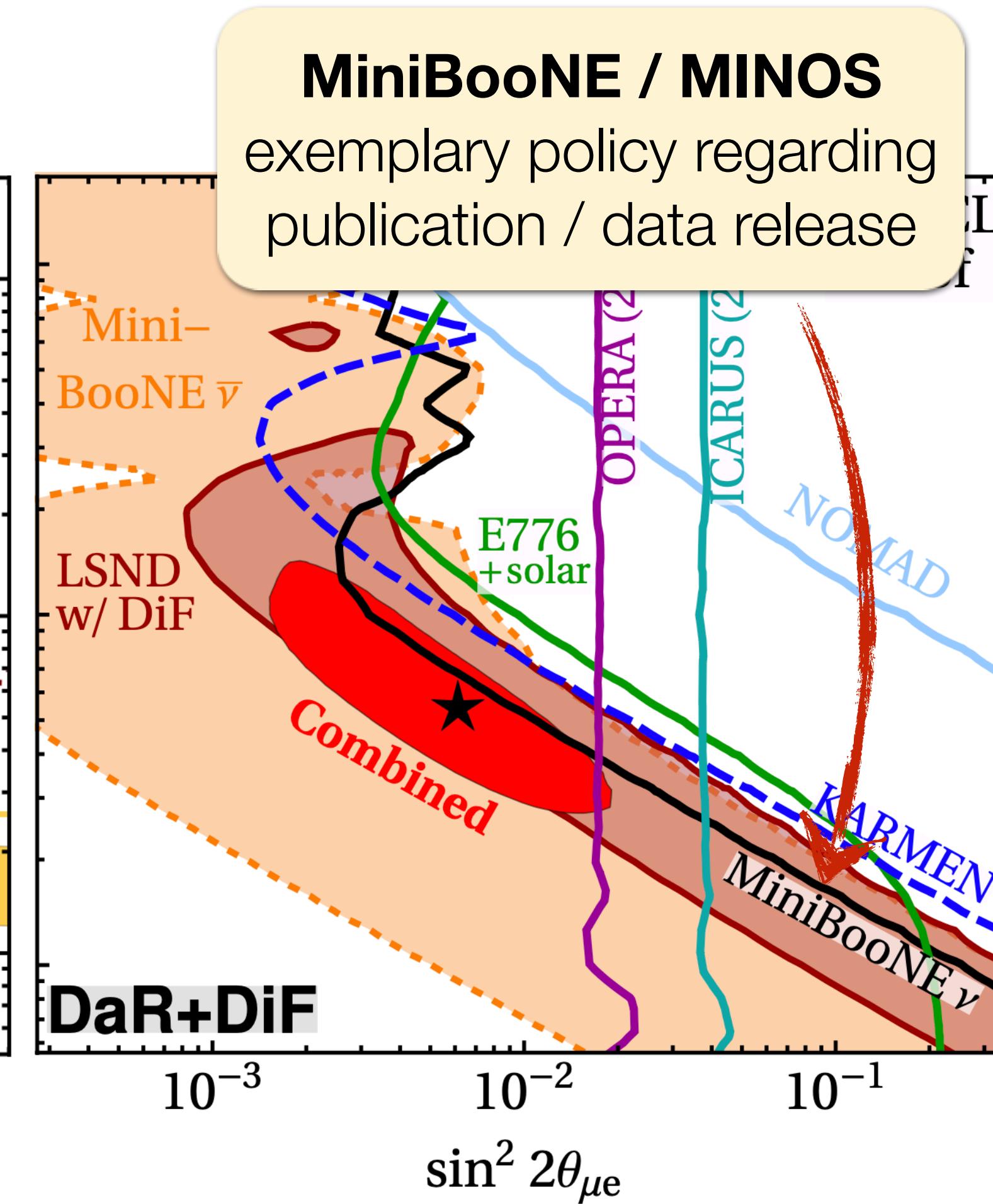
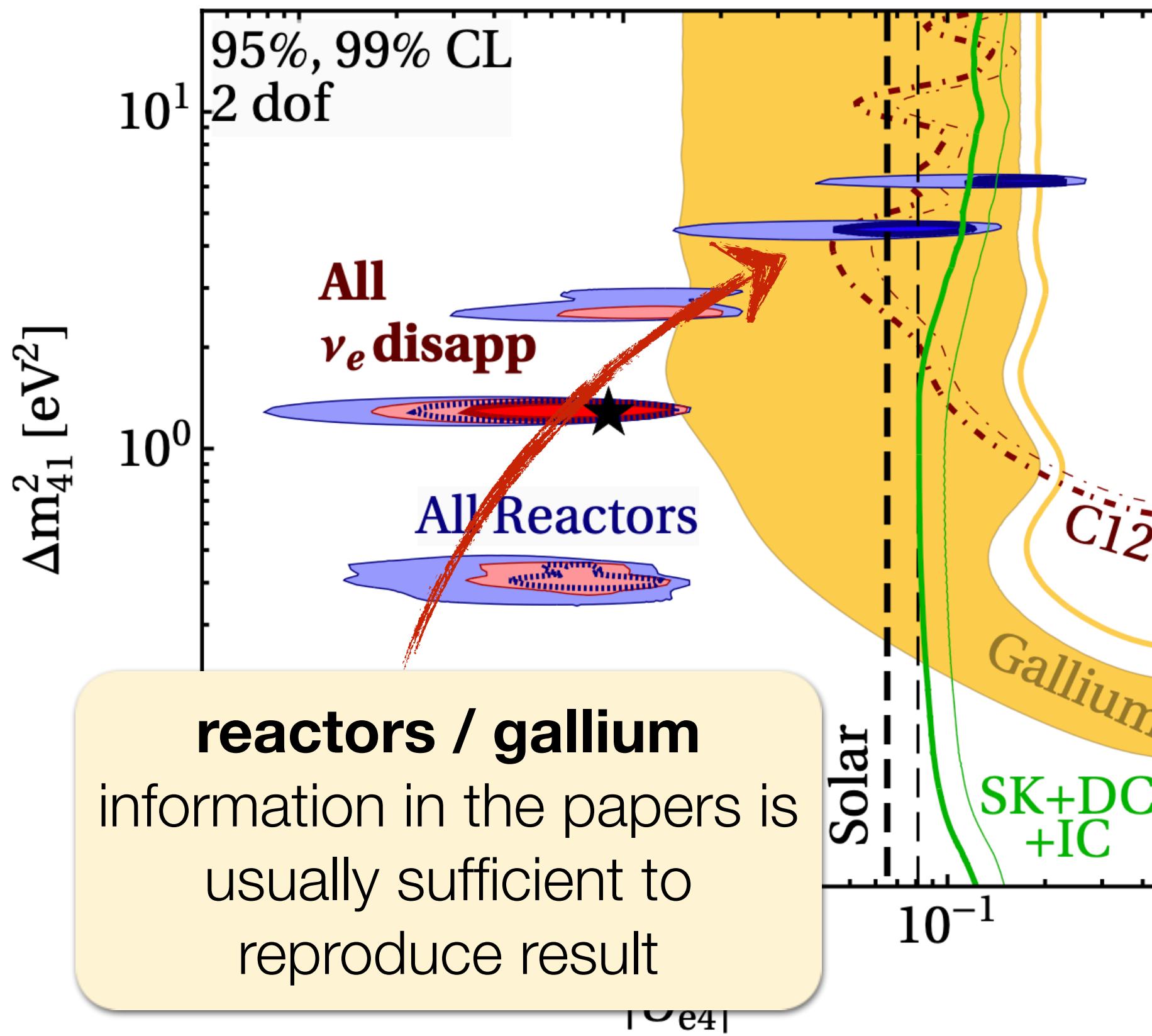
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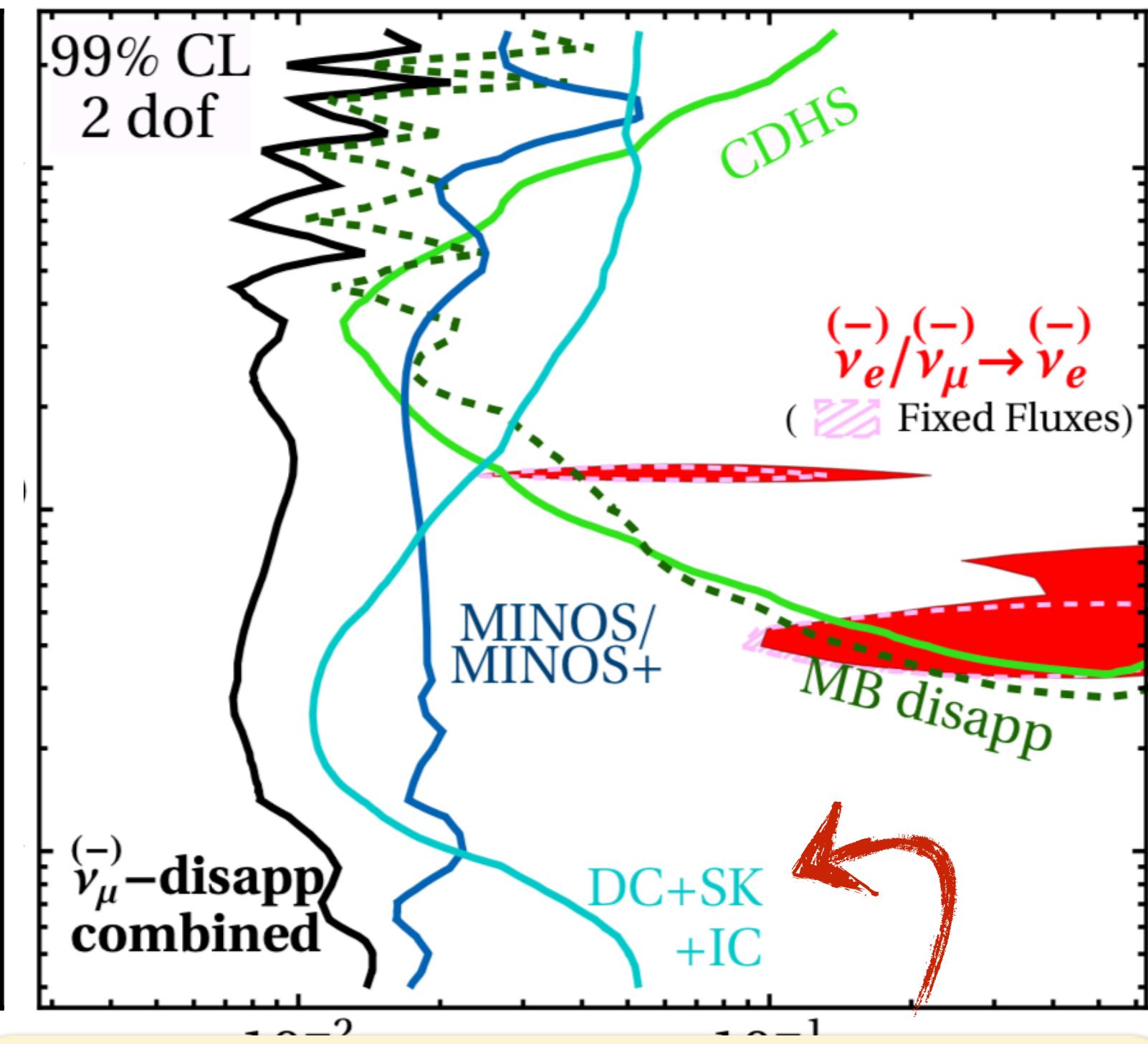
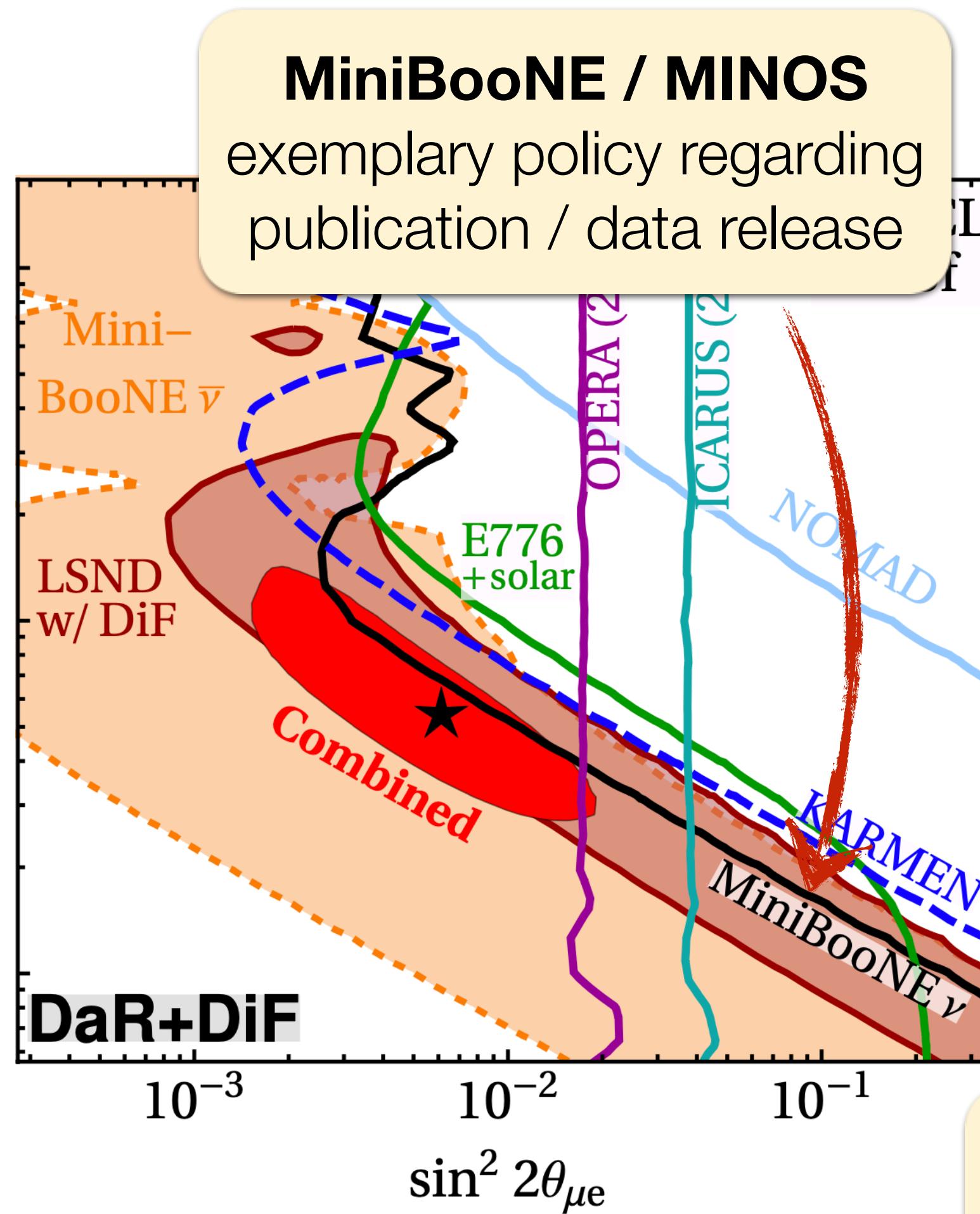
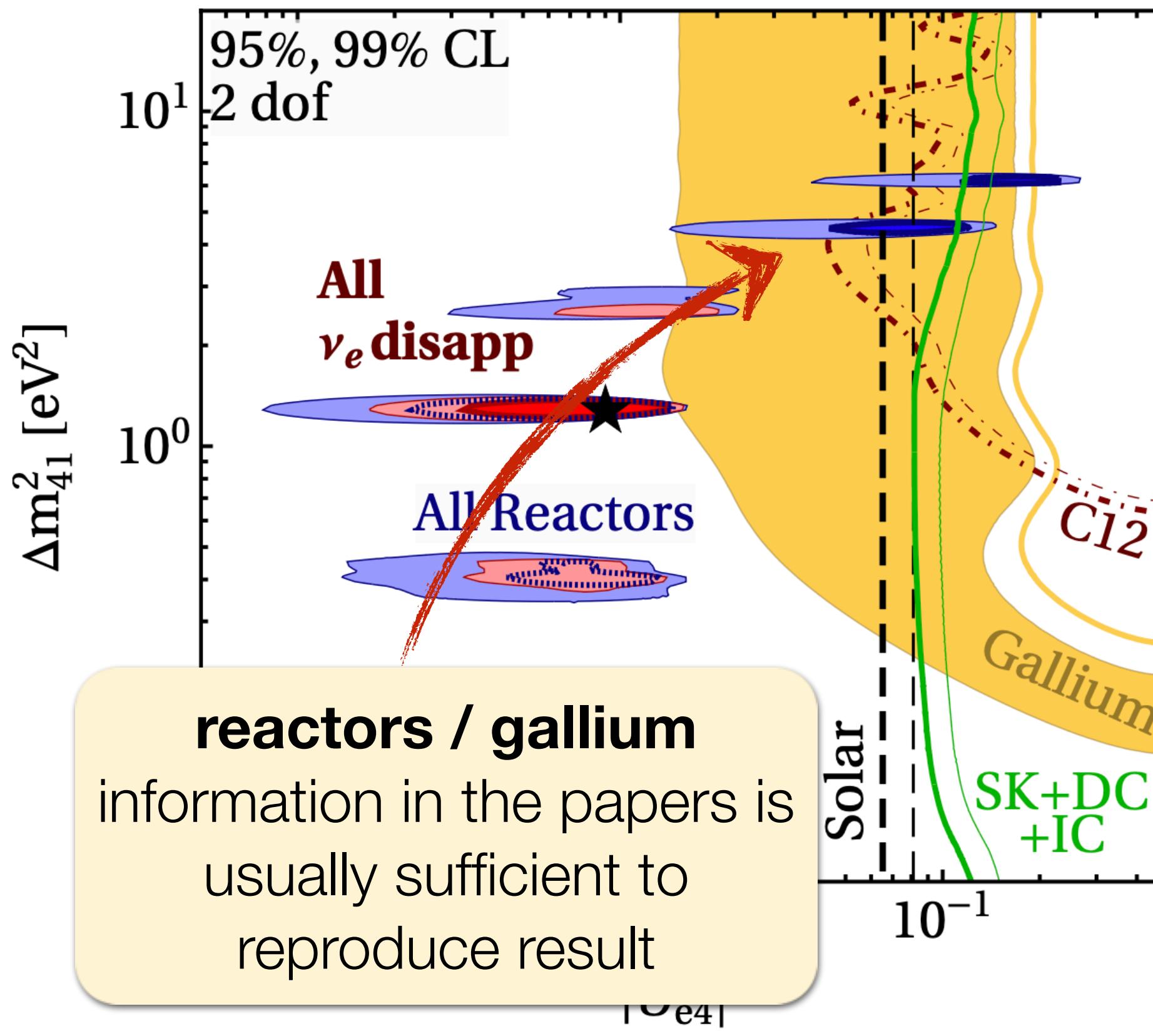
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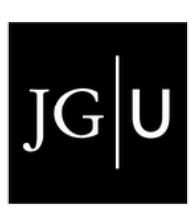
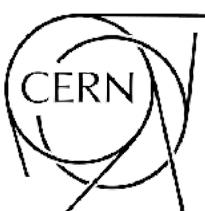


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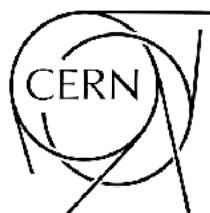
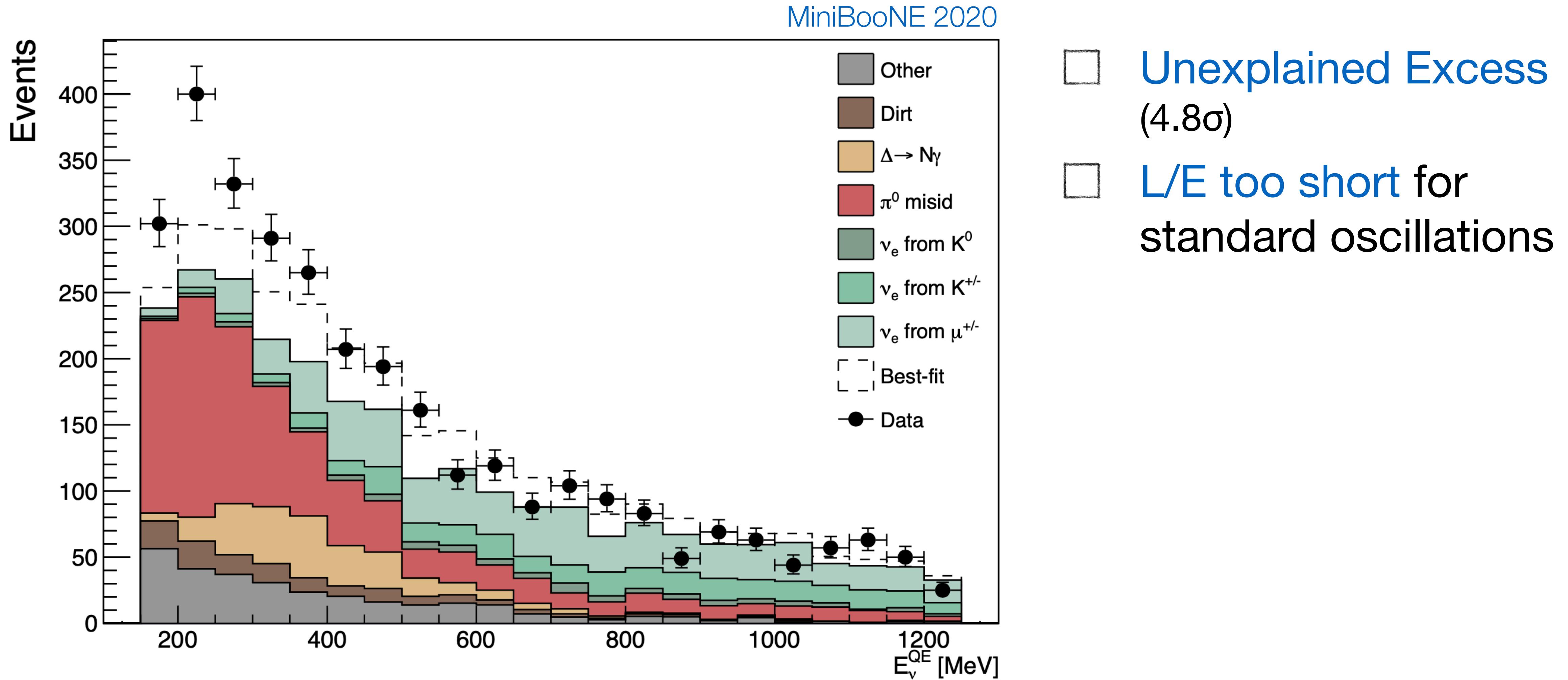
Dentler et al. 2018



SuperKamiokande
totally impossible to reproduce without investing years of work
(or collaborating with Michele Maltoni who has invested years of work)



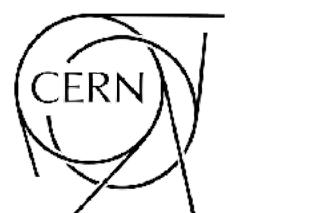
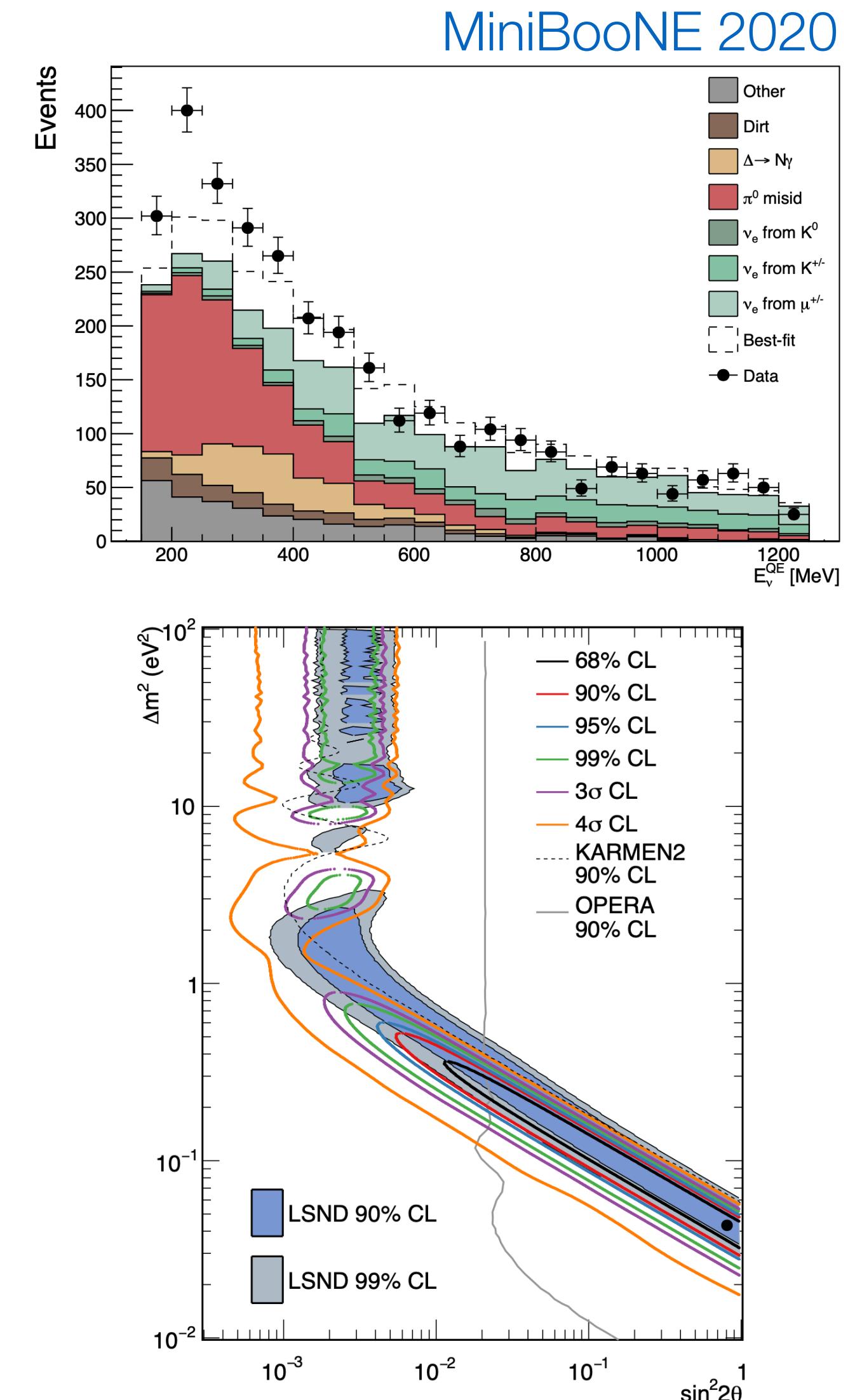
Sterile Neutrino Oscillations – MiniBooNE



Usefulness of MiniBooNE's Data Releases

MiniBooNE's oscillation data releases feature

- data shown in plots in machine-readable form
- flux tables
- efficiencies
- MC event samples (E_{true} vs E_{reco})
 - can be reweighted to match any BSM oscillation scenario
 - can be used to extract detector / analysis response function
- Covariance matrices
- Detailed instructions on how to reproduce the official results
- A very approachable collaboration
 - real interest in outside input
 - ability to share information without too much red tape



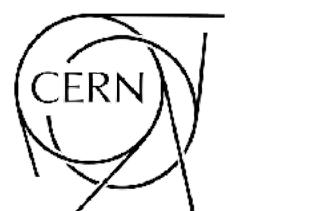
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Non-Standard Neutrino Interactions

- new dimension-6 operators
 - NC: new oscillation effects
 - CC: new production / detection effects

$$\mathcal{L}_{\text{NSI,NC}} = \sum_{f,\alpha,\beta} 2\sqrt{2}G_F \varepsilon_{\alpha\beta}^{f,P} (\bar{\nu}_\alpha \gamma_\mu P_L \nu_\beta) (\bar{f} \gamma^\mu P f) + \text{h.c.}$$

$$\mathcal{L}_{\text{NSI,CC}} = \sum_{f,f',\alpha,\beta} 2\sqrt{2}G_F \varepsilon_{\alpha\beta}^{ff',P} (\bar{\nu}_\alpha \gamma_\mu P_L \ell_\beta) (\bar{f}' \gamma^\mu P f) + \text{h.c.}$$

- Production / propagation / detection effects should be considered **together** (EFT needs **SU(2)** invariant UV completion, e.g. SMEFT)

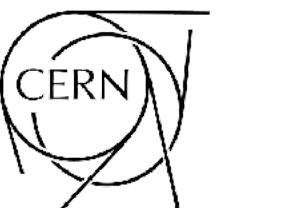
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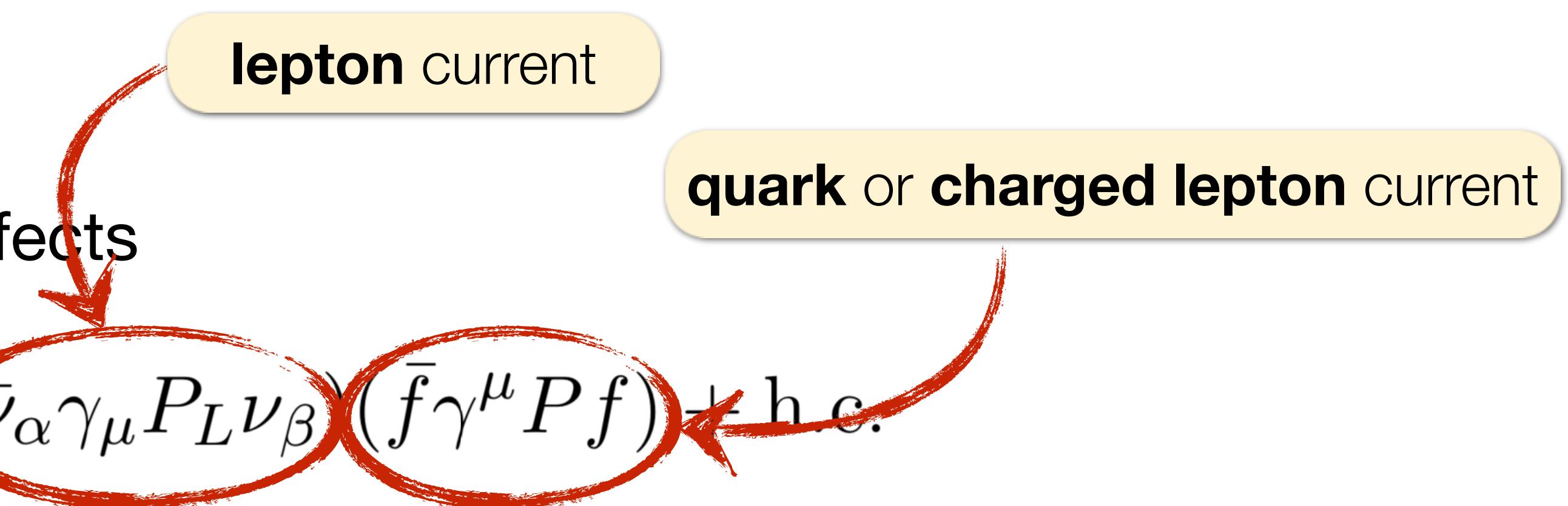
JG|U

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Non-Standard Neutrino Interactions

- ☐ new dimensionless coefficients
- strength of the new interaction
- relative to SM weak interactions

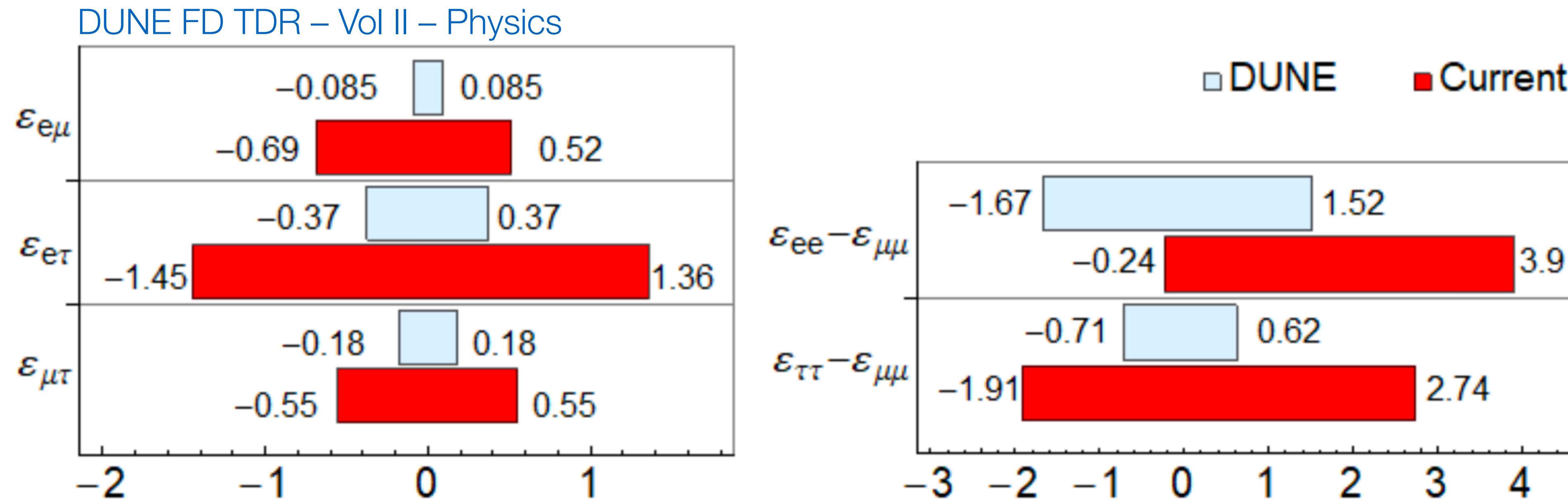
interaction effects

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- ☐ Production / propagation / detection effects should be considered **together** (EFT needs **SU(2)** invariant UV completion, e.g. SMEFT)

Non-Standard Neutrino Interactions



- For quark flavor-universal V–A interactions (as in the SM):
 - straightforward implementation in terms of modified oscillation probabilities
- For interactions with different Lorentz / flavour structures:
 - modified meson decays \rightarrow need production MC events to re-decay mesons

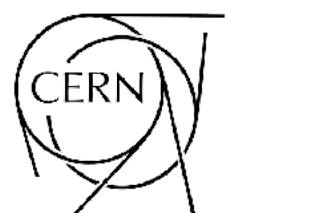
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Heavy Neutral Leptons

- Just another word for **sterile neutrino**, typically used if mass \sim MeV–GeV
- coupled via the neutrino portal

$$\mathcal{L} \supset y \bar{L} (i\sigma^2 H^*) N$$

- leads to mixing between ν and N
 - ⇒ any process that makes ν in the SM can also make N
(suppressed by a mixing angle)
 - ⇒ meson decays
 - ⇒ need **ability to re-decay mesons**



Heavy Neutral Leptons

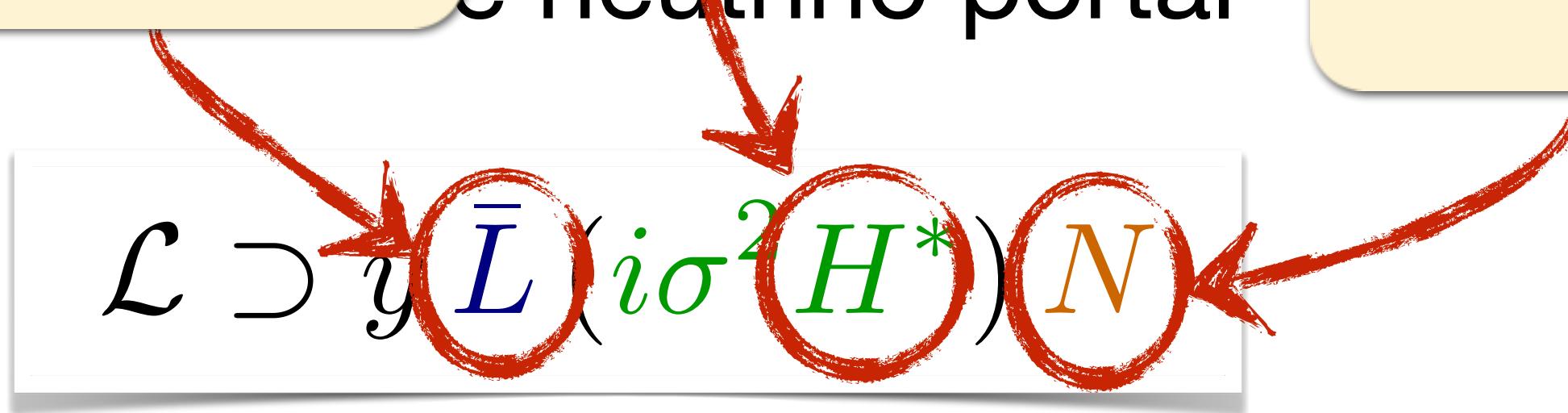
- Just another word for **sterile neutrino**.

typically used in SM Higgs doublet – G

- SM lepton doublet \bar{L} **neutrino portal**

Heavy Neutral lepton

(singlet fermion = sterile ν
= right-handed ν)



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⇒ meson decays

⇒ need **ability to re-decay mesons**

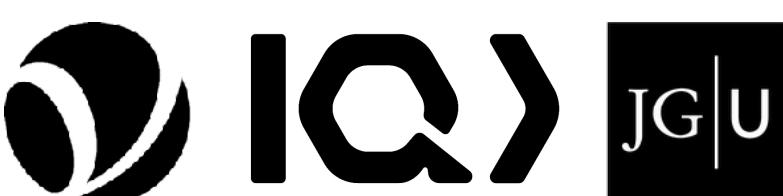
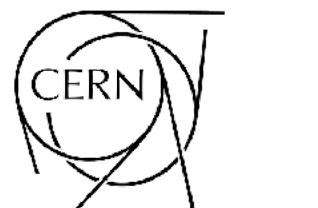
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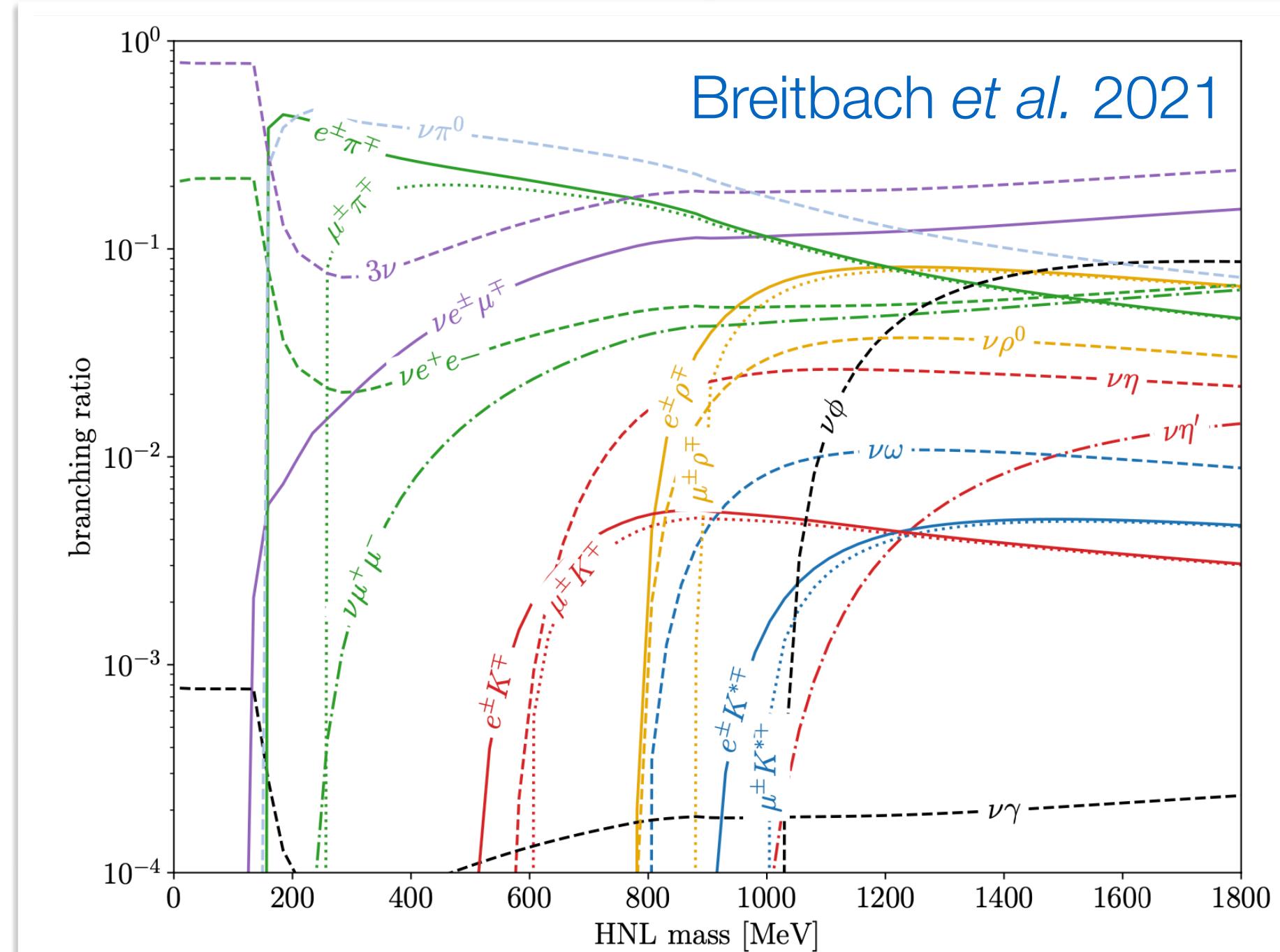
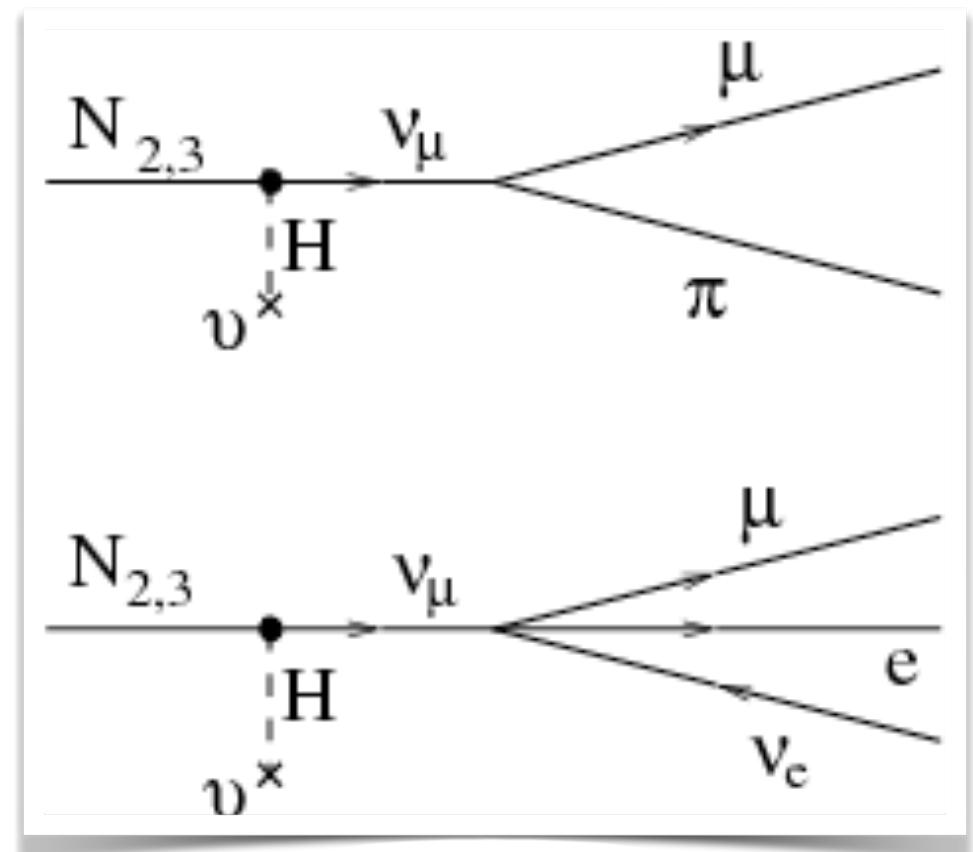
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 - ➡ need **ability to re-decay mesons**

- **Very generic extension of SM**
 - leftovers of **GUT multiplets?**
- **Useful phenomenological tool**
 - ν masses (seesaw mechanism)
 - cosmic baryon asymmetry (leptogenesis)
 $m \gg 100$ GeV (thermal)
or $m < 100$ GeV (ARS))
 - dark matter ($m \sim$ keV)
 - mediator to a dark sector



Heavy Neutral Leptons

- decay inside the detector into a variety of modes
 - very hard to extract without dedicated searches by the experiments
 - best chance would probably be with **event-by-event information** on
 - ▶ **particle IDs** (with probability of correctness)
 - ▶ **4-momenta** (with errors)
 - systematic uncertainties?
- backgrounds from SM neutrino interactions



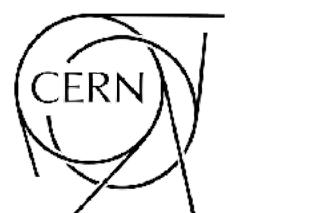
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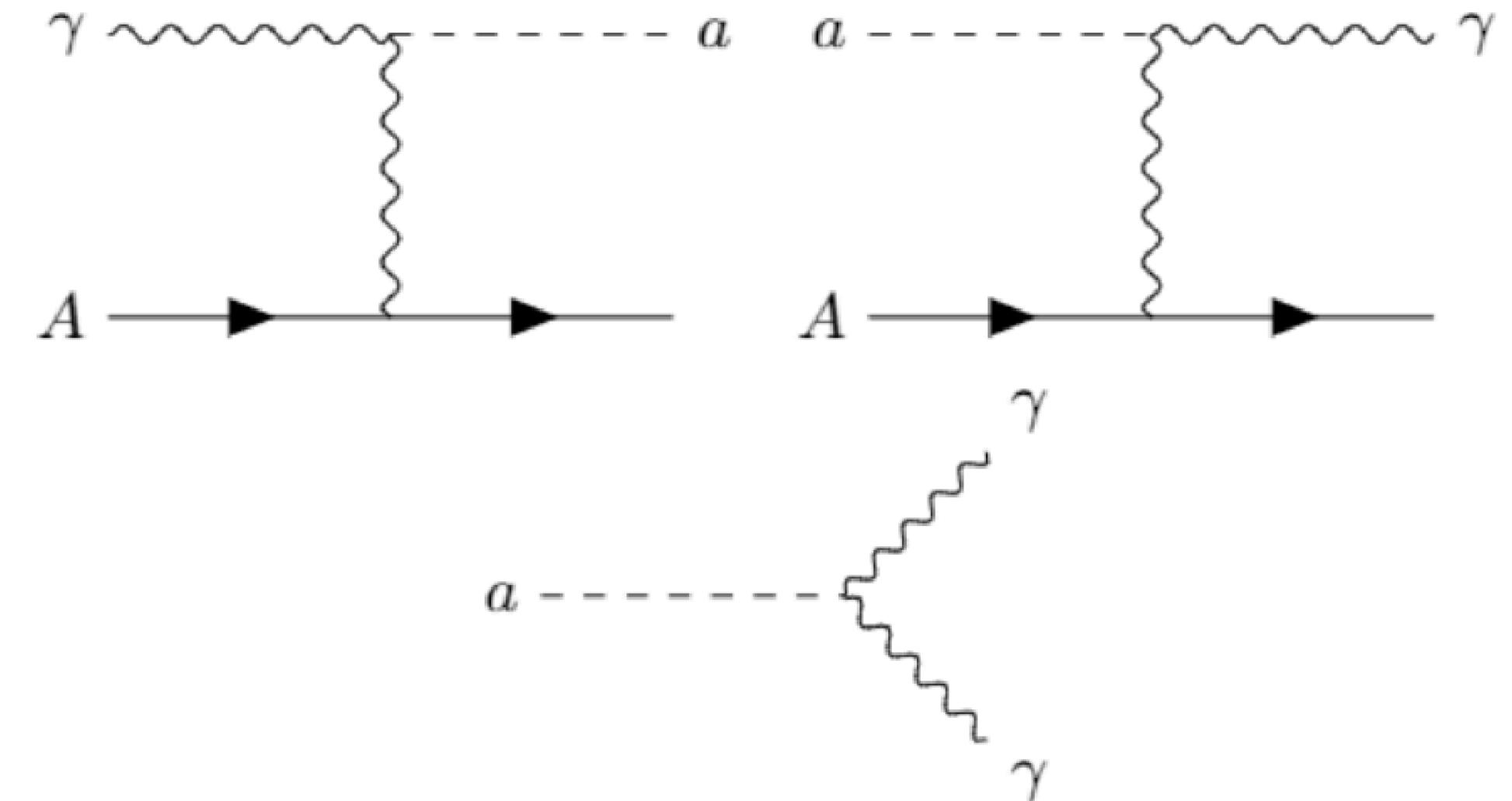
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Axion-Like Particles

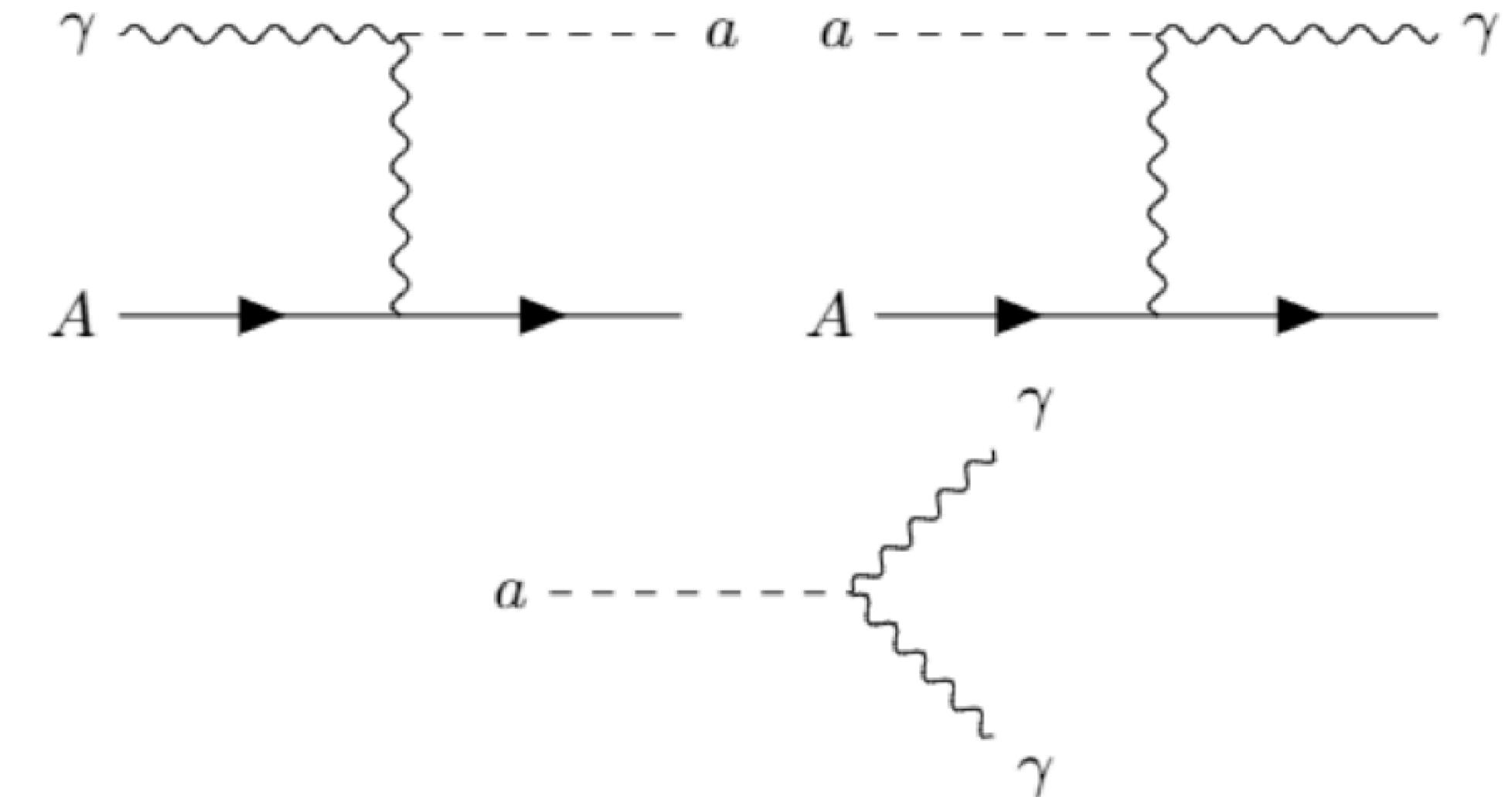
- a new pseudoscalar particle with a shift symmetry $a \rightarrow a + \text{const.}$
 - if **sub-eV**: possible solution to the **strong CP problem** (“axion”) → non-accelerator searches (astrophysics, quantum sensors)
 - frequent occurrence in **string compactifications**
 - **dark matter candidate**



e.g. Dent et al. 2020

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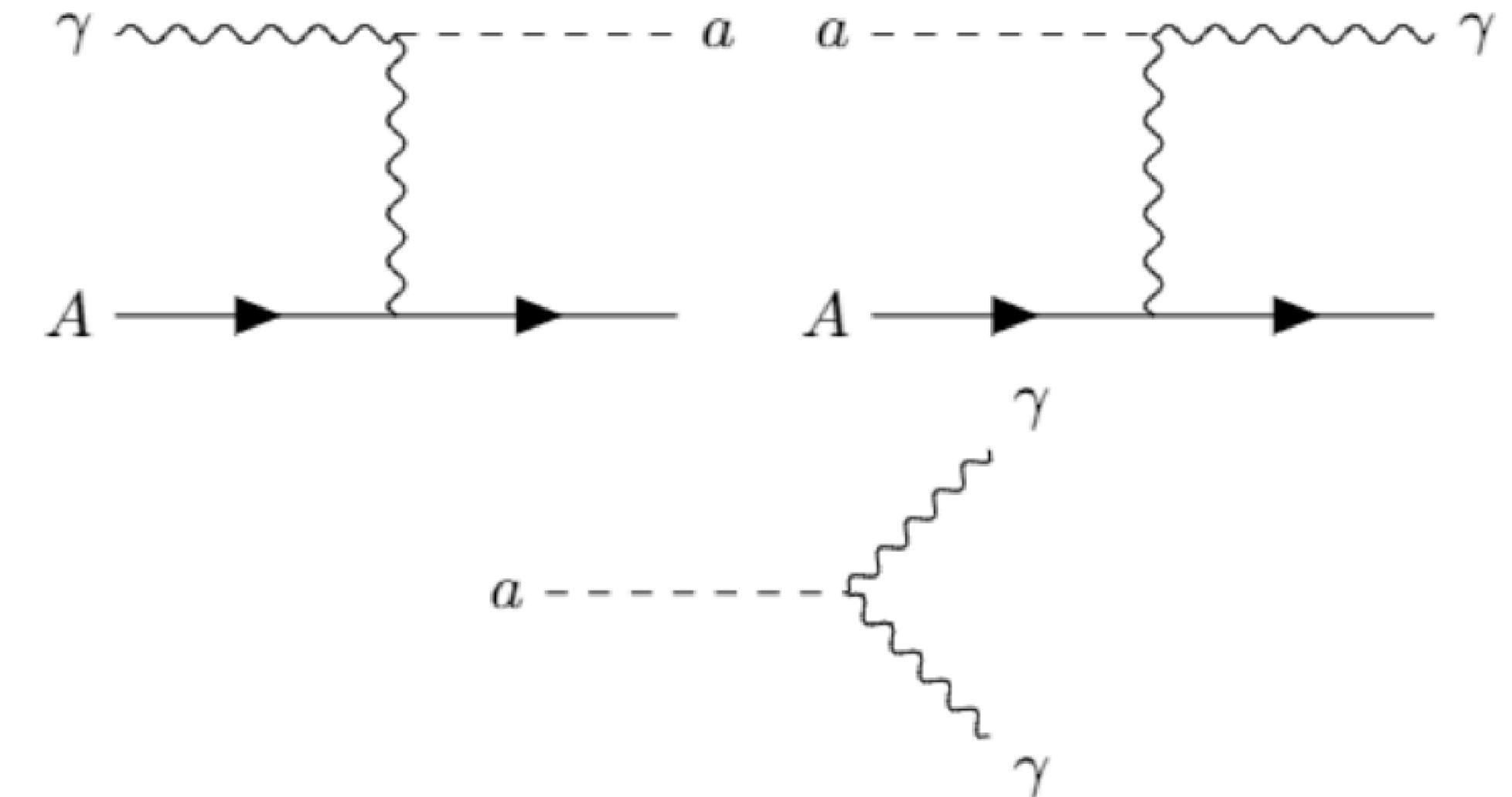
$$\mathcal{L}_a = \frac{1}{2} \partial_\mu a \partial^\mu a + \frac{\alpha_s}{4\pi f_a} a \text{tr} G^{\mu\nu} \tilde{G}_{\mu\nu} + \frac{s\alpha}{8\pi f_a} a F^{\mu\nu} \tilde{F}_{\mu\nu} + \mathcal{L}_a^{\text{int}} \left[\frac{\partial_\mu a}{f_a}; \psi \right]$$

Axion-Like Particles

- a new pseudoscalar particle with a shift symmetry $a \rightarrow a + \text{const.}$

- if **sub-eV**: possible solution to the **strong CP problem** (“axion”)
→ non-accelerator searches
(astrophysics, quantum sensors)
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- signatures with **one or two electromagnetic showers** very important
 - also relevant for e.g. dark photon decay to e^+e^-



e.g. Dent et al. 2020

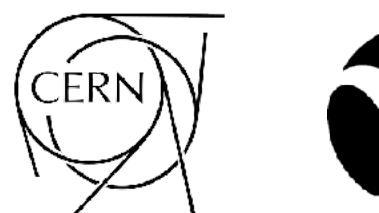
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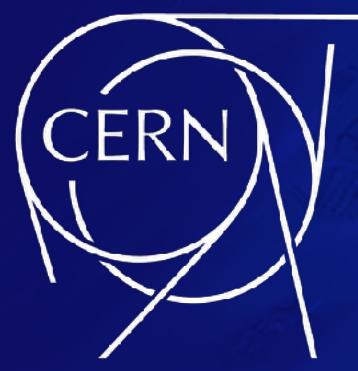
Wish List

- beam Monte Carlo events (with full decay history)
- detector Monte Carlo events (true energy \leftrightarrow reconstructed energy)
- covariance matrices
- efficiencies





Thank You!



QUANTUM
TECHNOLOGY
INITIATIVE



