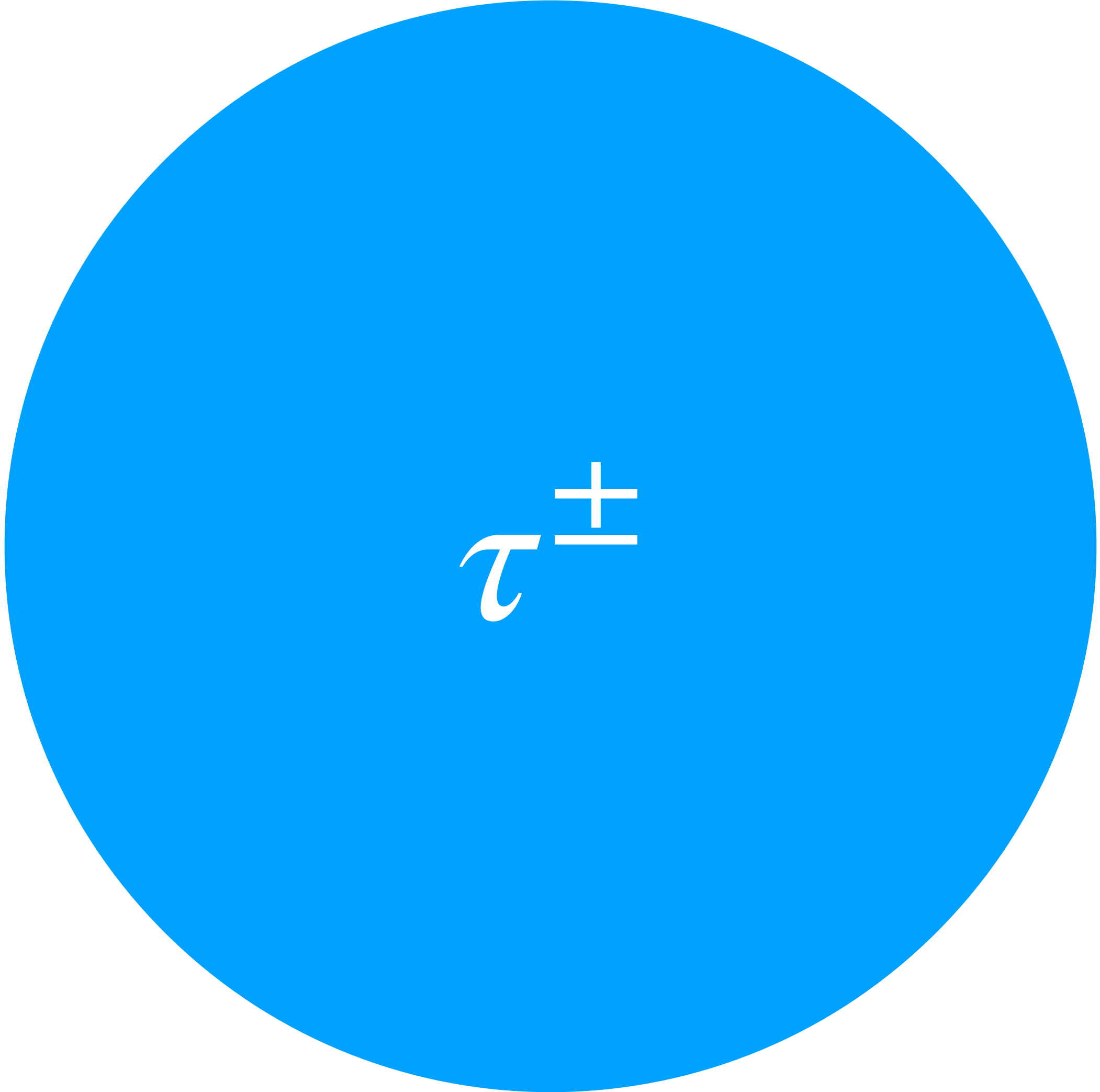


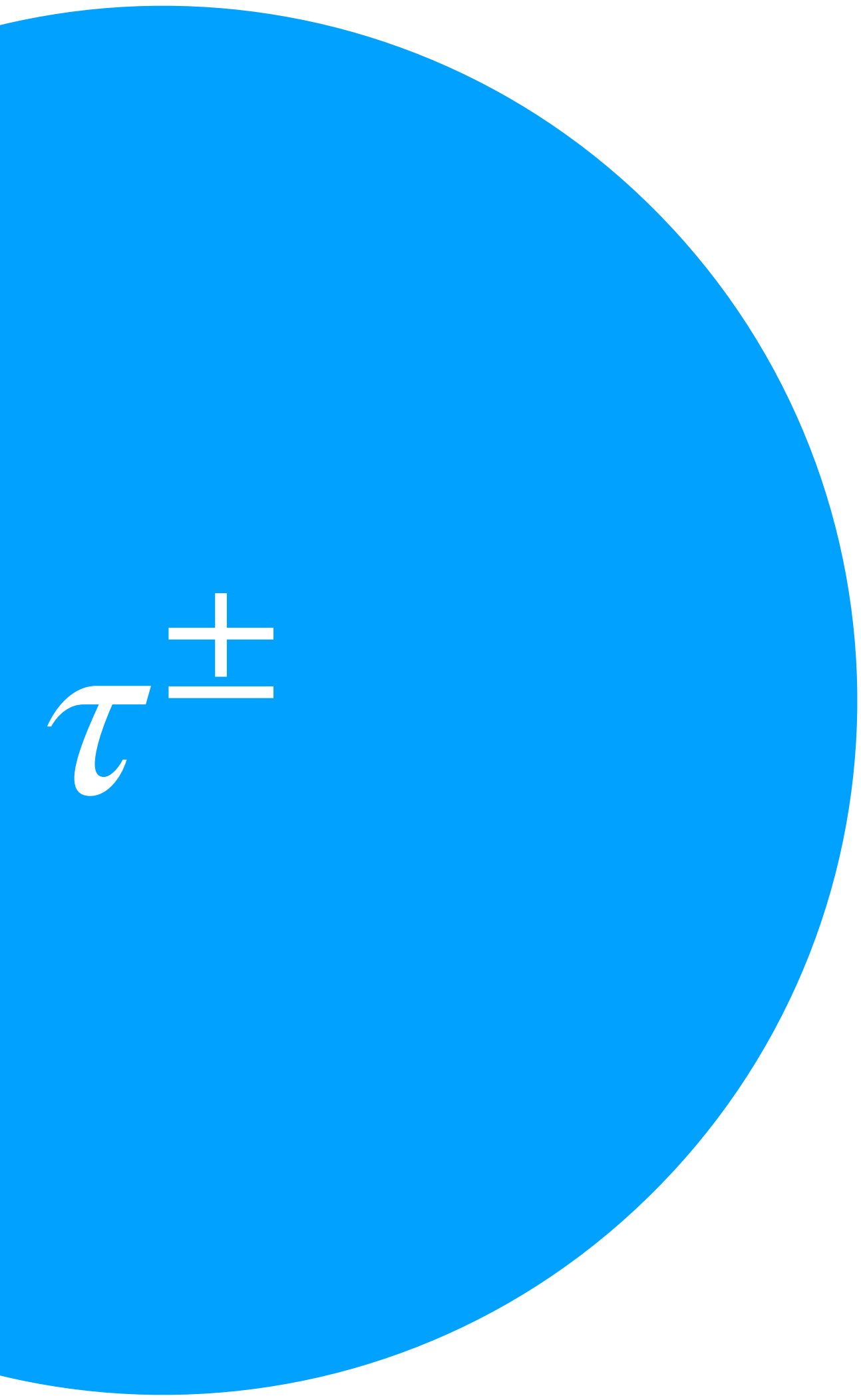
Neutrino & Dark Matter Connections

Kevin J. Kelly, Texas A&M University

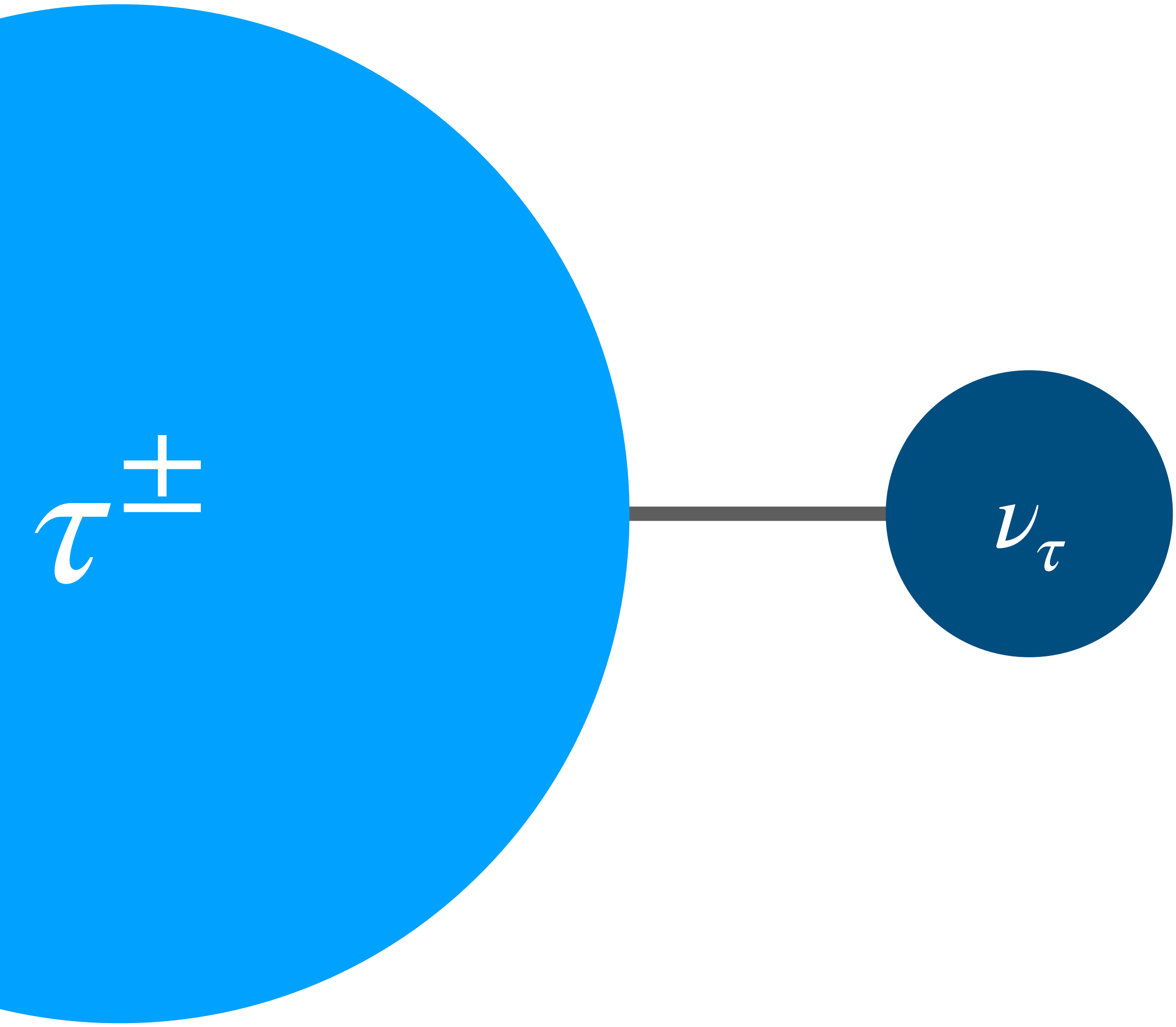
TAU2023 — 4-8 December, 2023

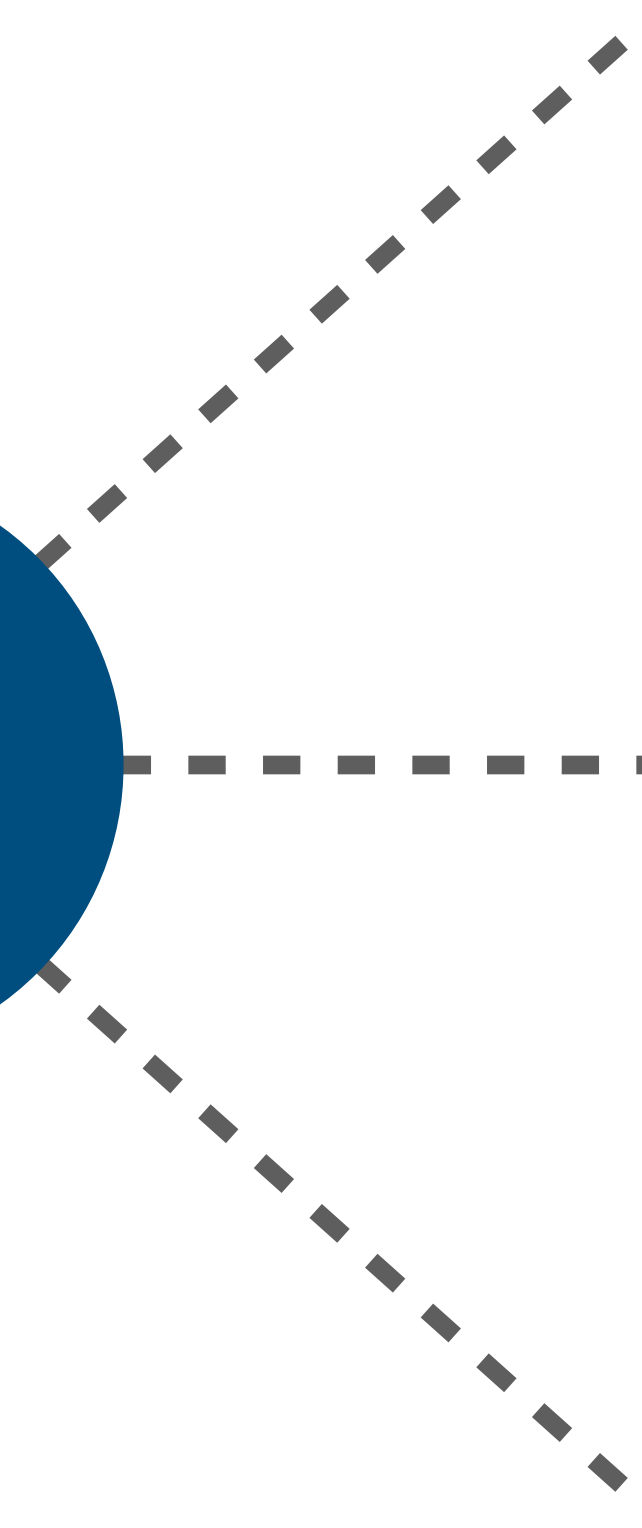
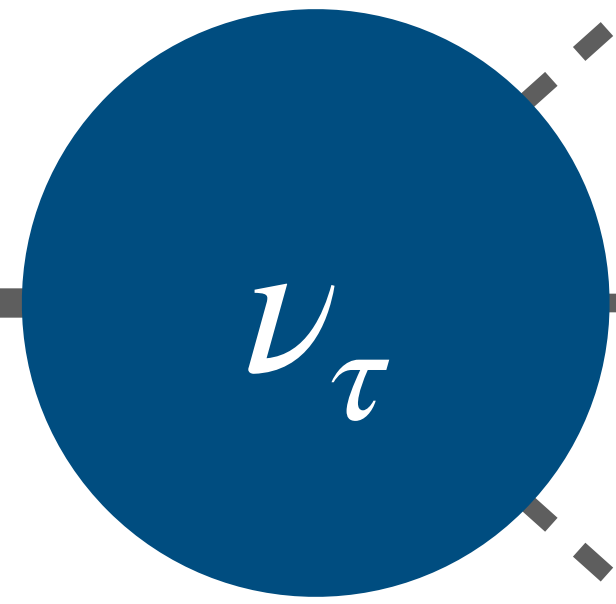
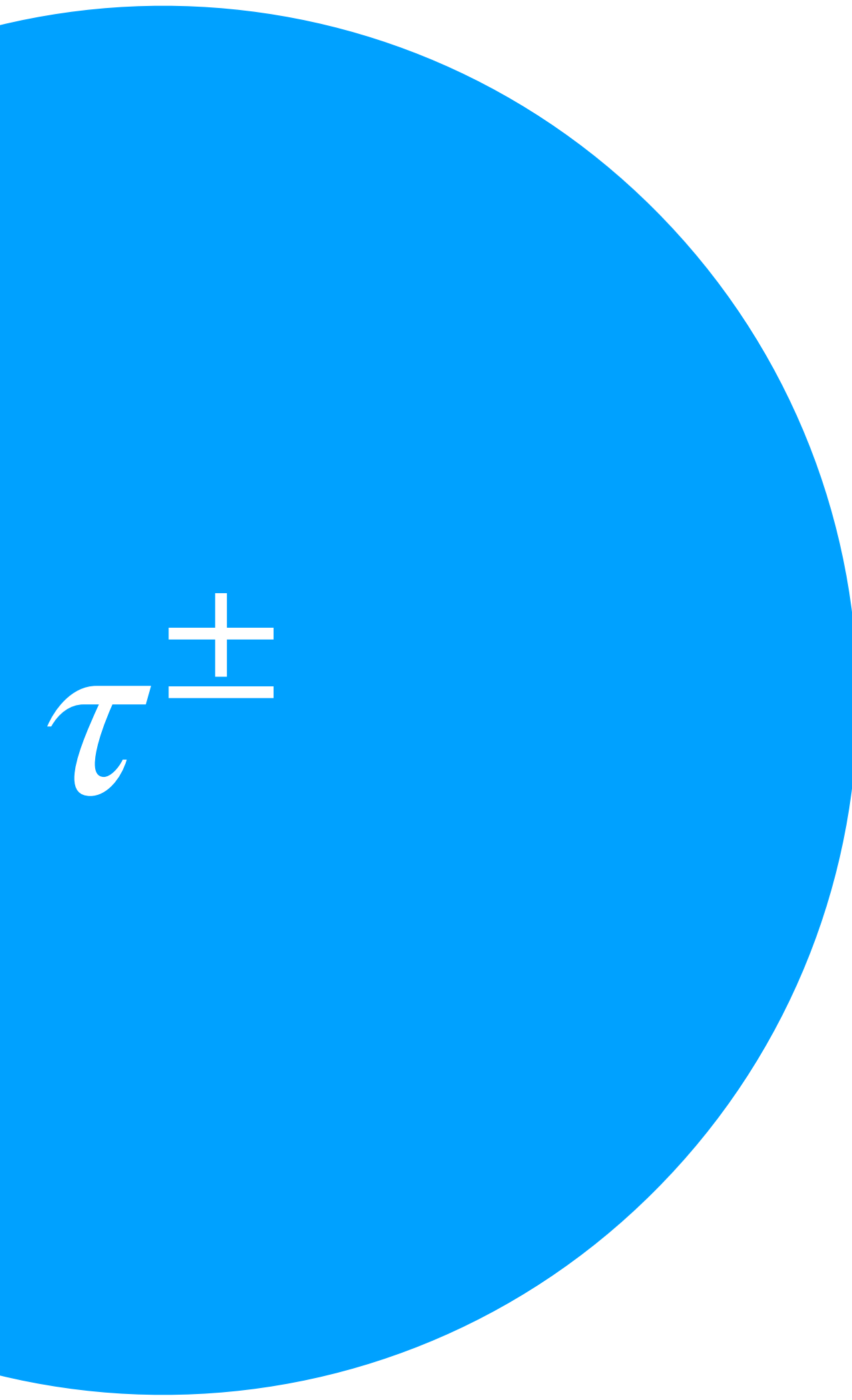
kjkelly@tamu.edu


$$\tau^{\pm}$$



$$\tau^{\pm}$$

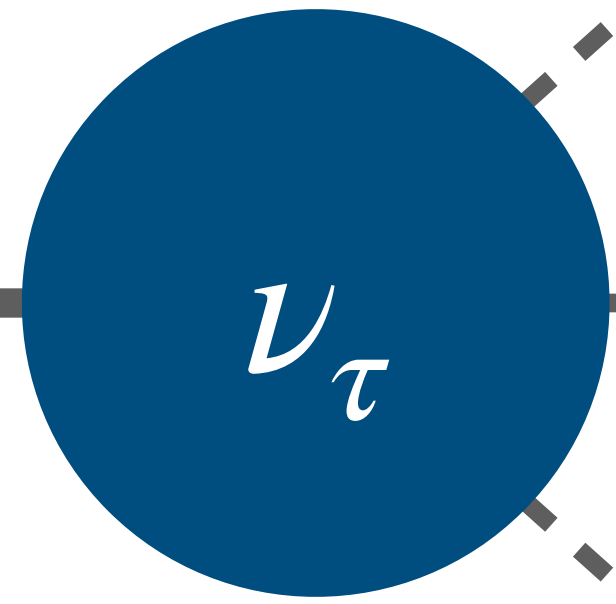
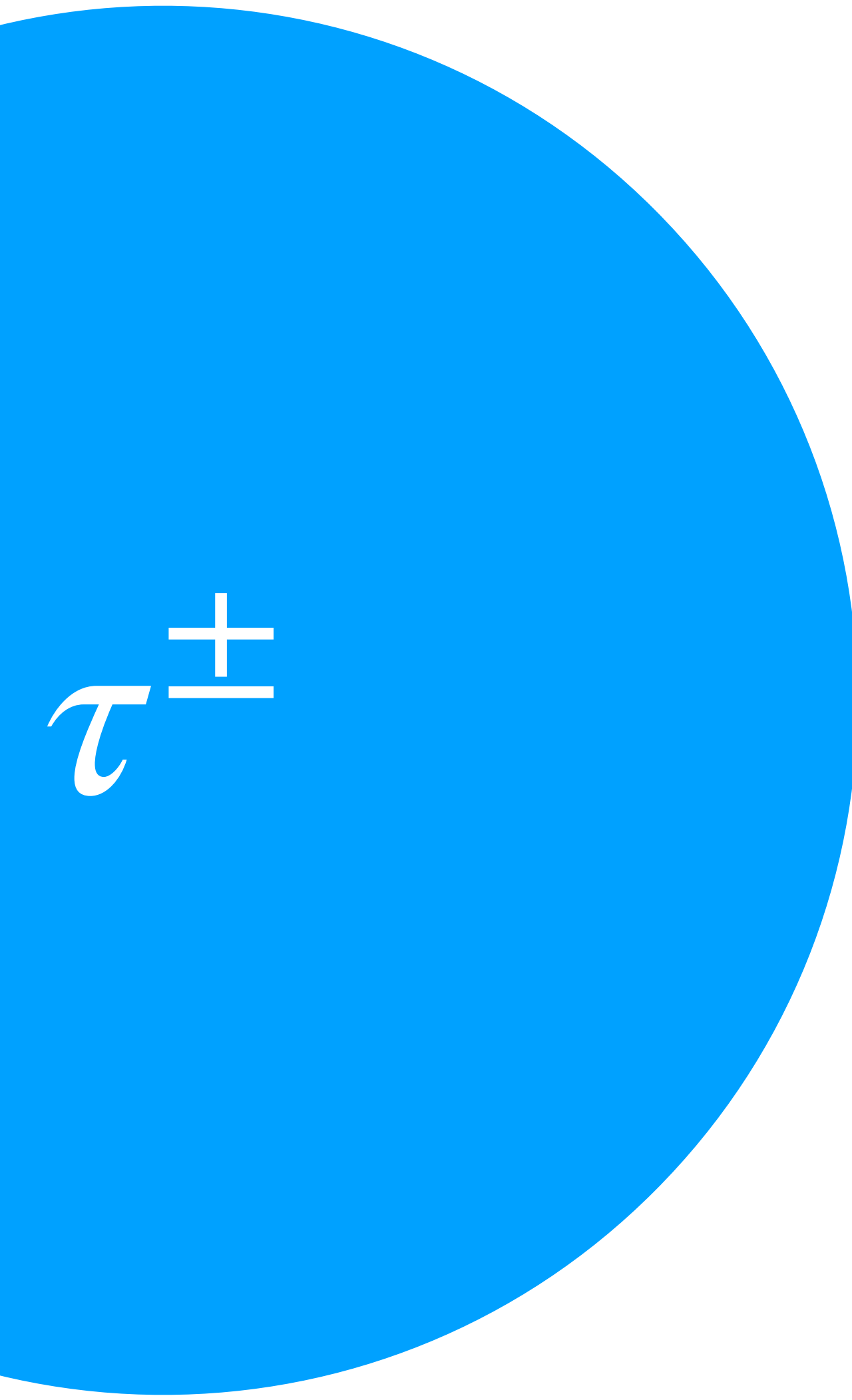




Precision understanding of Lepton Mixing, CP Violation, etc.

New Physics in Neutrino Oscillations

Connections to Dark Sector Physics



Precision understanding of Lepton Mixing, CP Violation, etc.

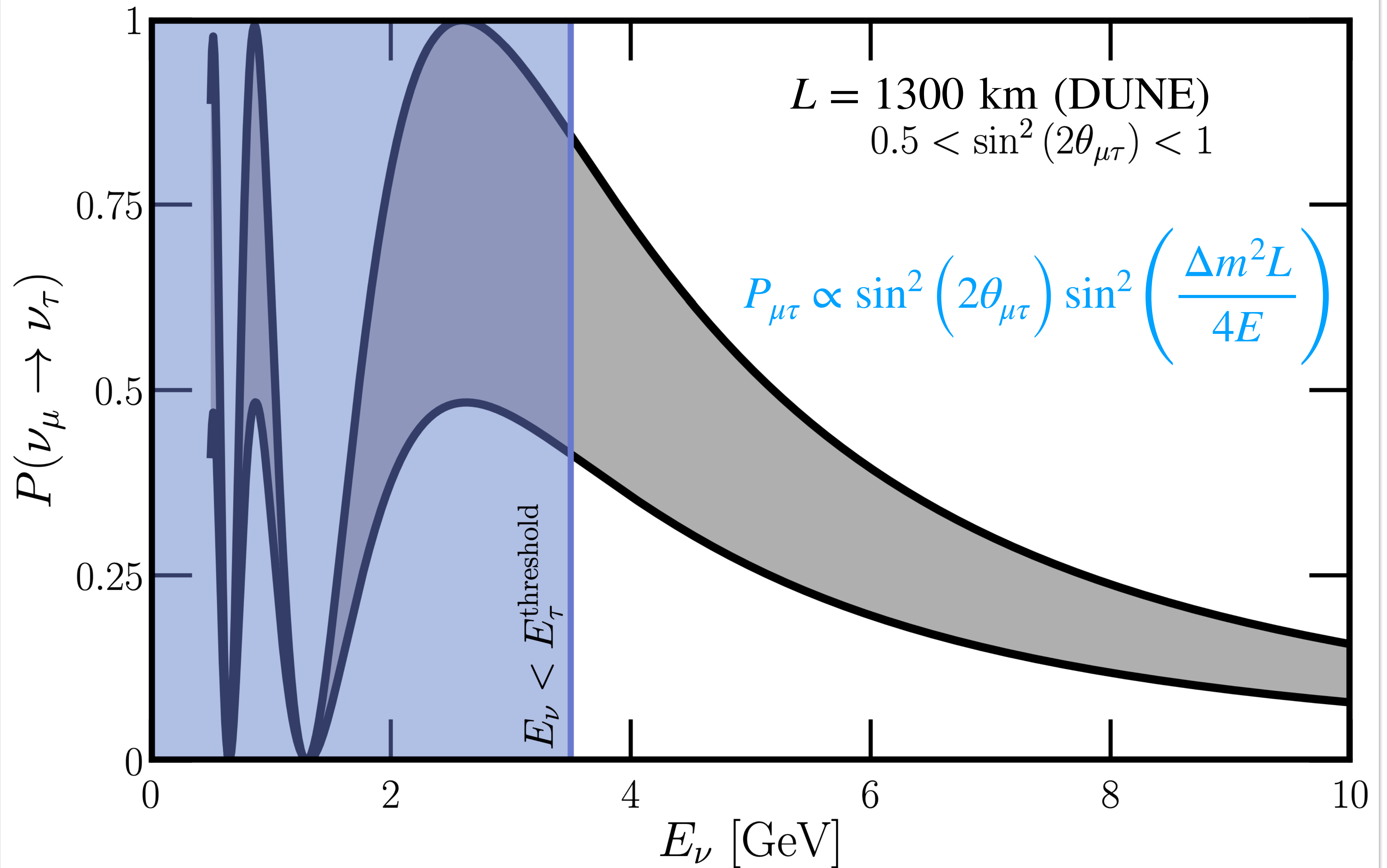


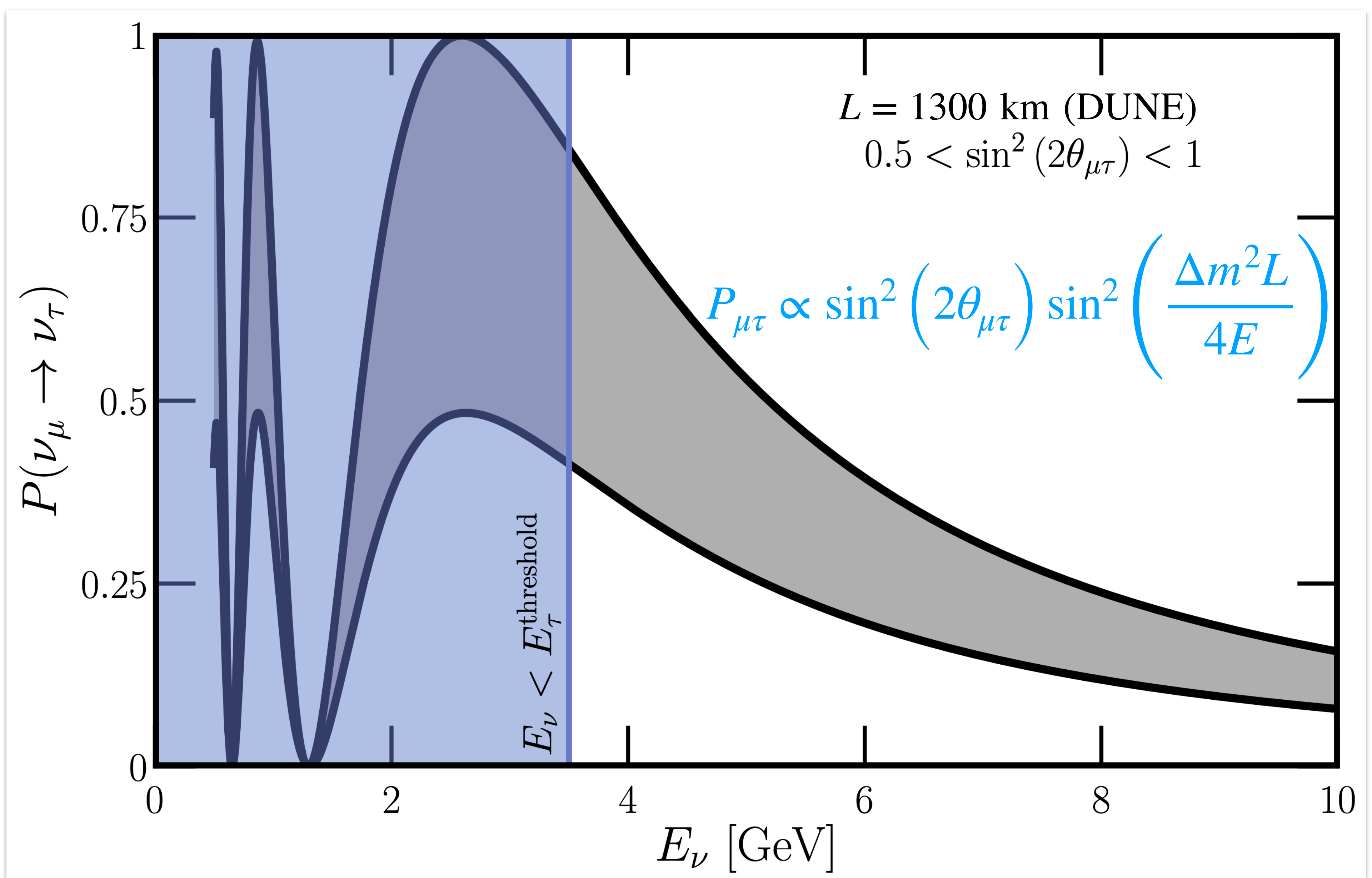
New Physics in Neutrino Oscillations
(A. Aurisano's talk after this!)



Connections to Dark Sector Physics

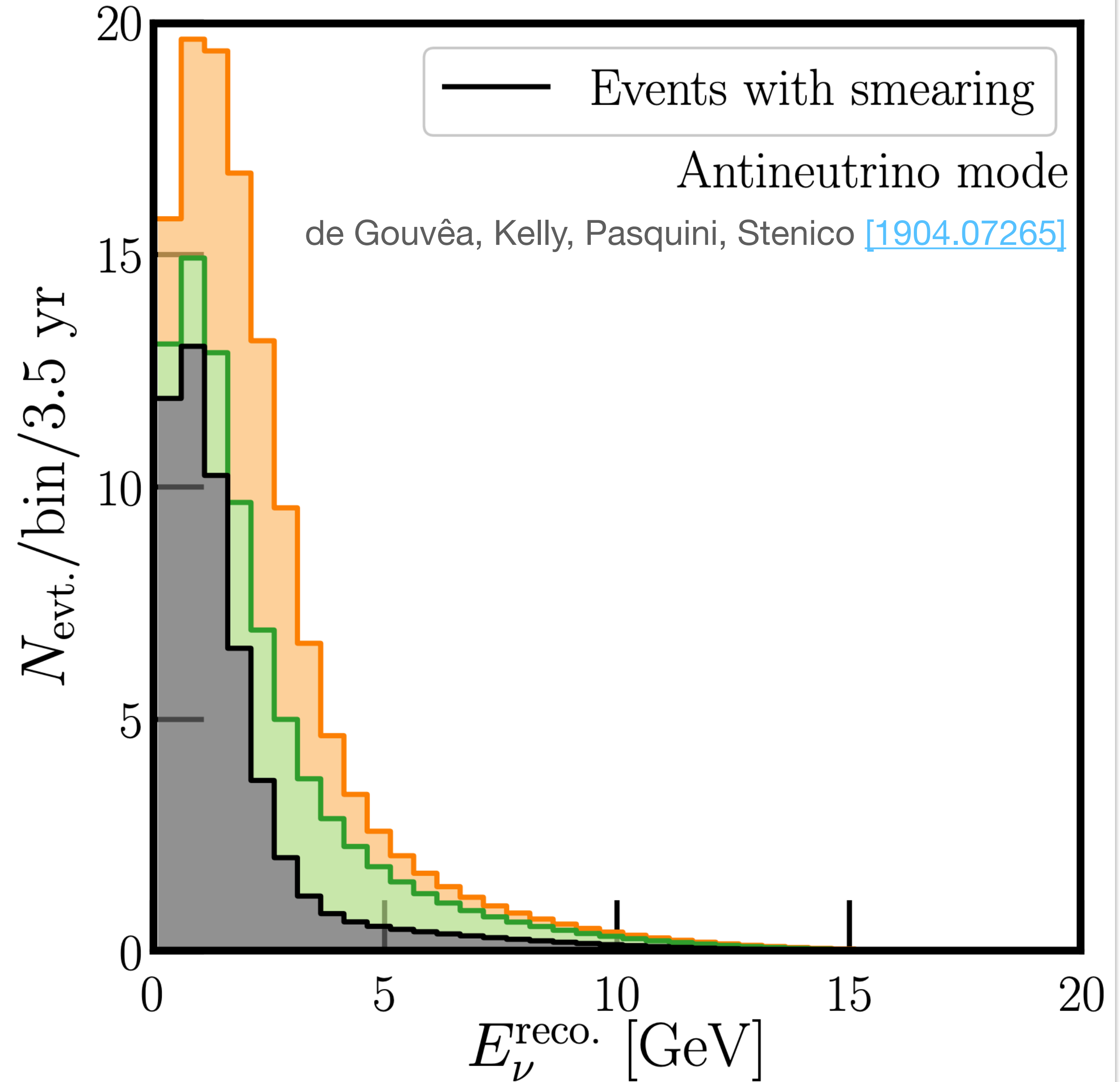
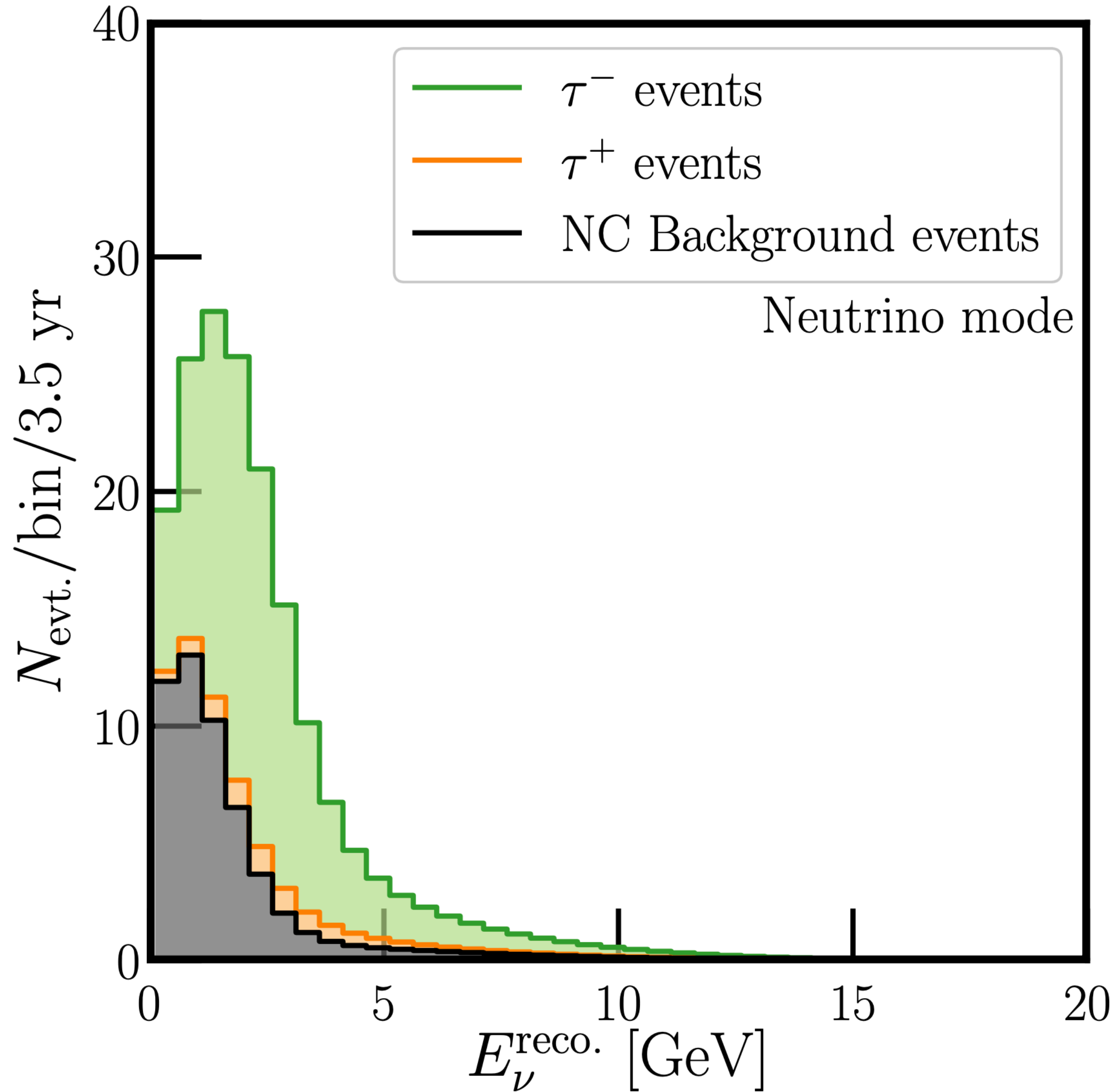
Tau Neutrinos for Three-Flavor Oscillation Physics

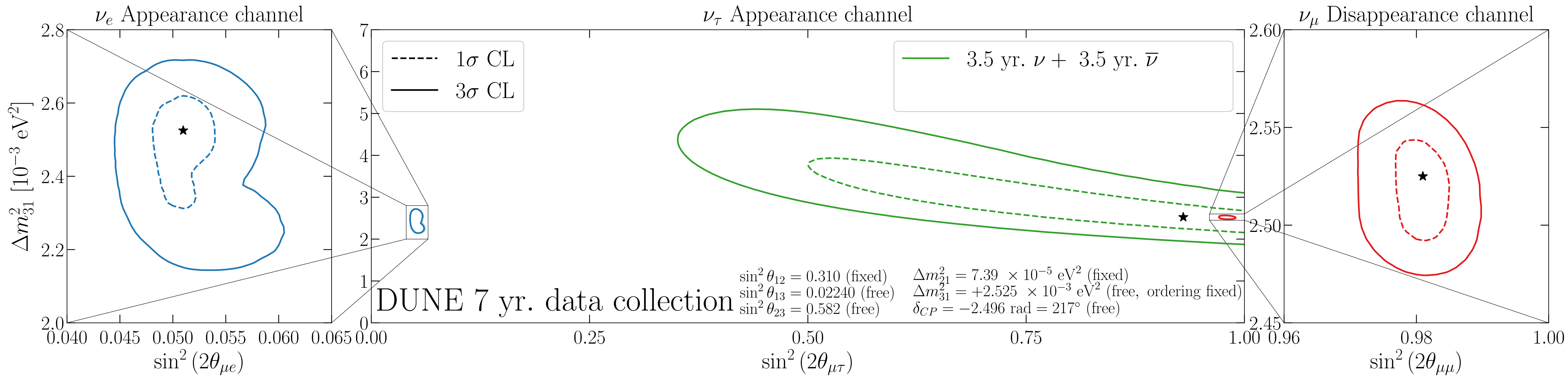




Three-flavor: $\sin^2(2\theta_{\mu\tau}) = 4 |U_{\mu 3}|^2 |U_{\tau 3}|^2 = 4 |U_{\mu 3}|^2 (1 - |U_{\mu 3}|^2 - |U_{e 3}|^2)$

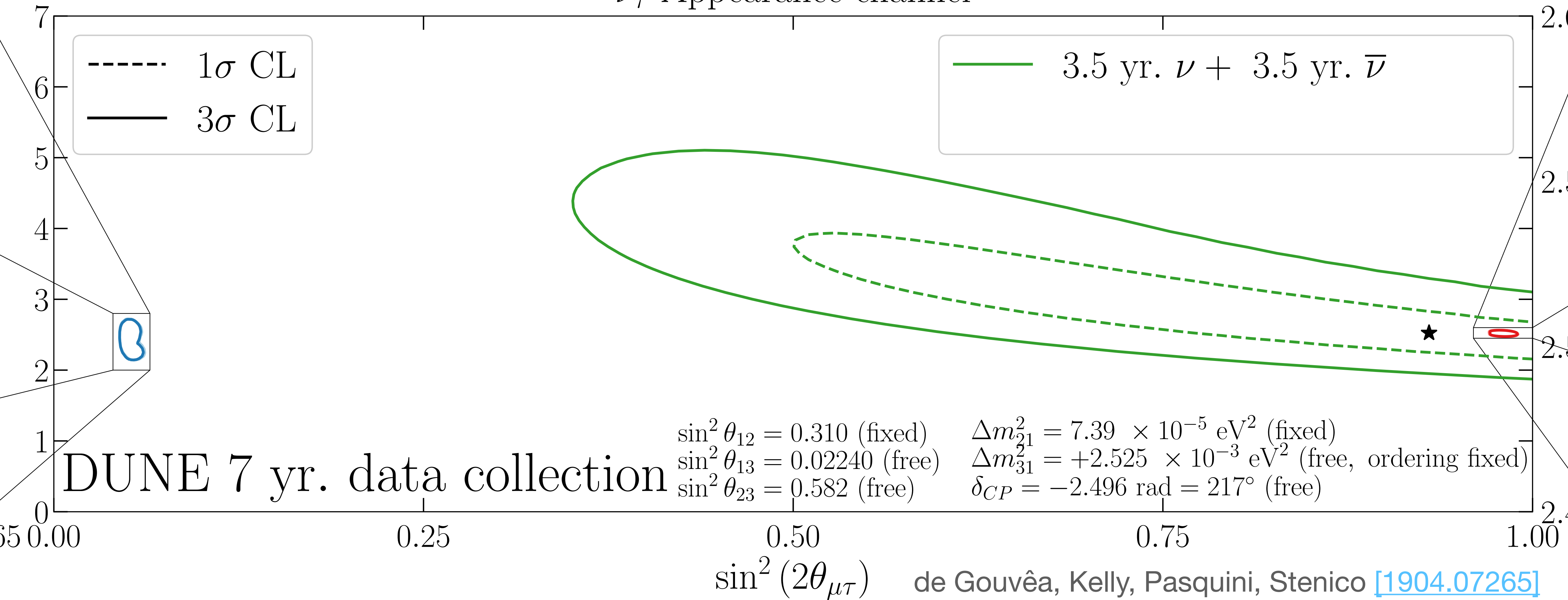
DUNE & Tau-Neutrino Appearance





de Gouvêa, Kelly, Pasquini, Stenico [\[1904.07265\]](#)

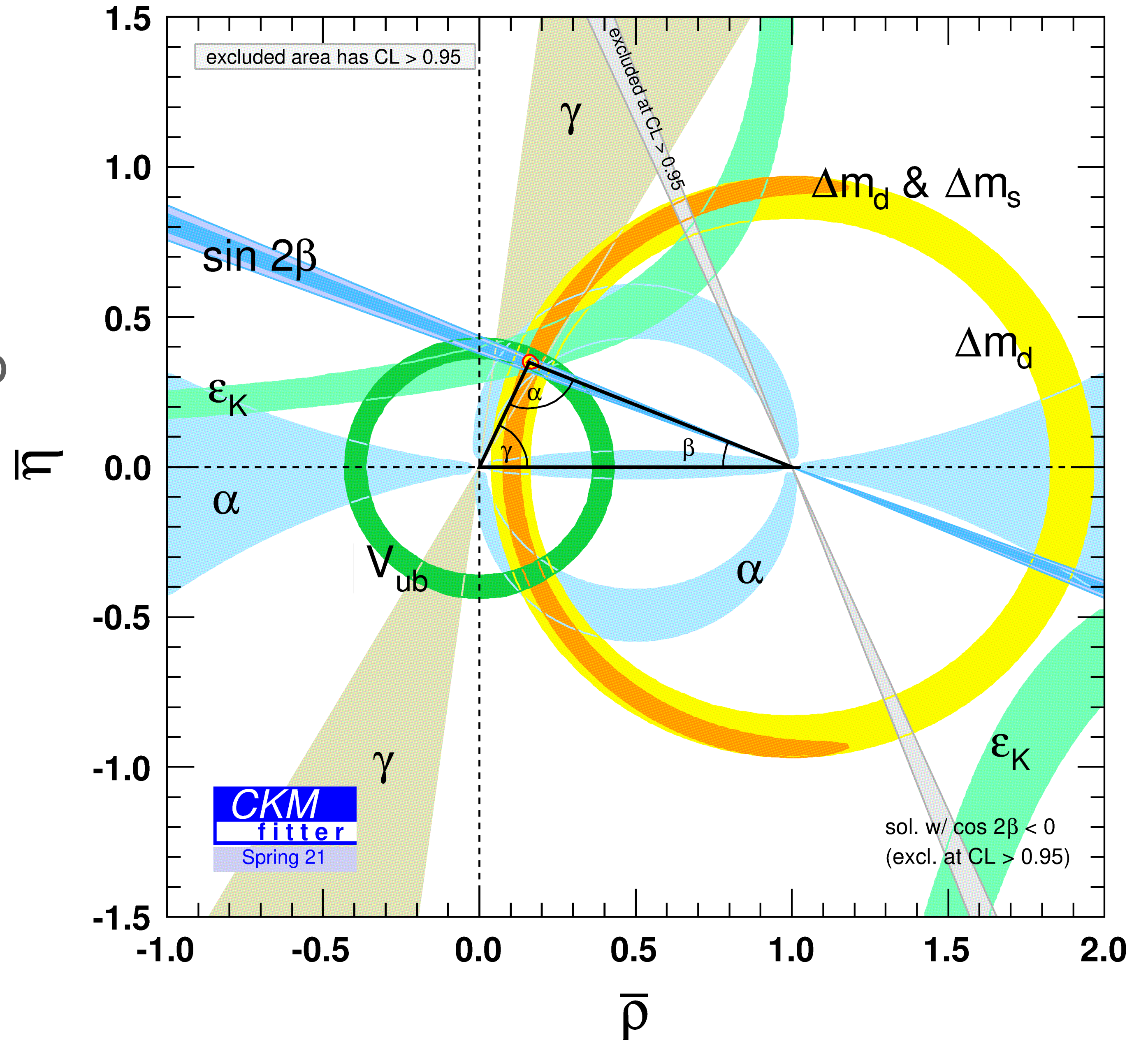
ν_τ Appearance channel

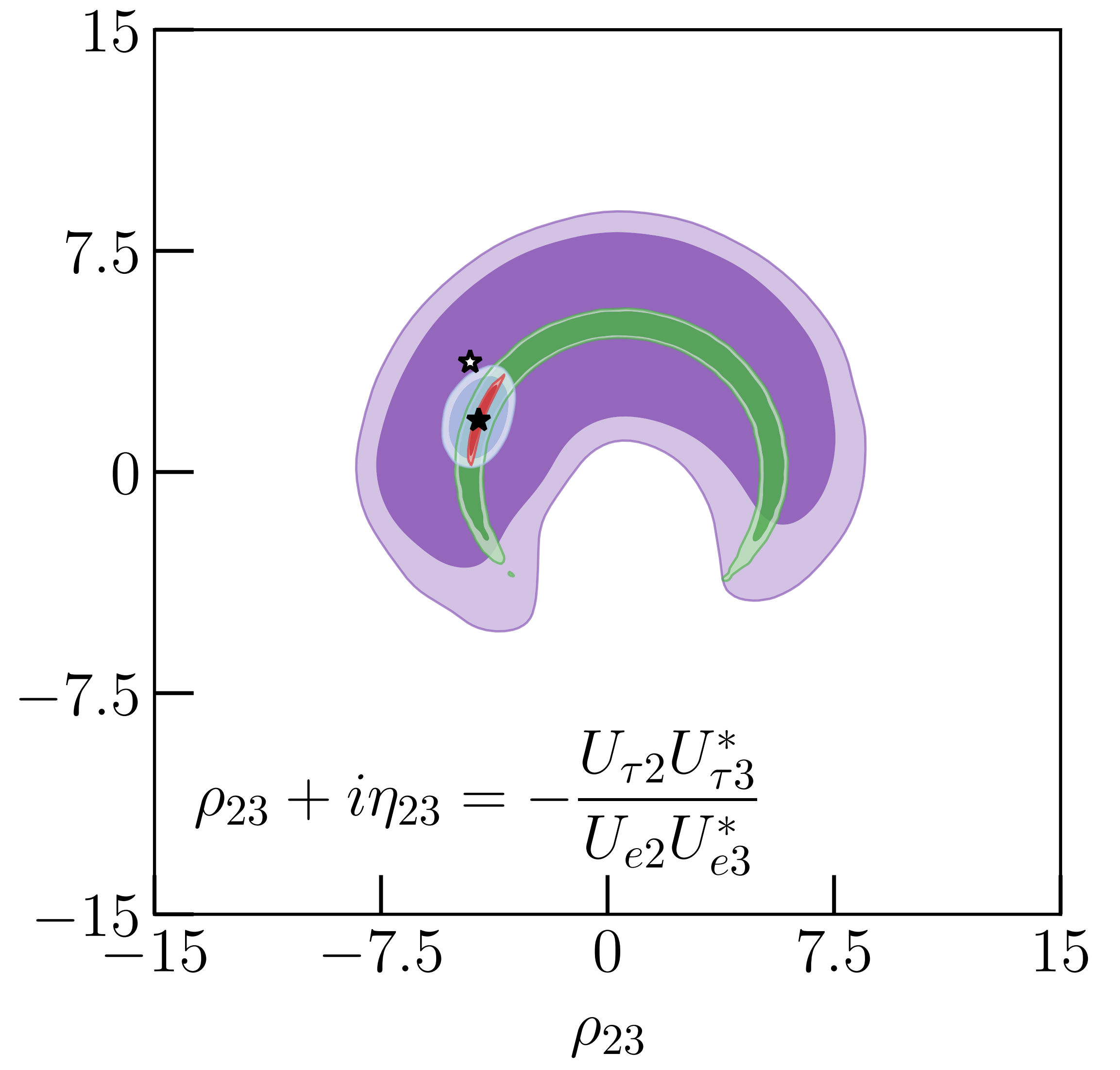
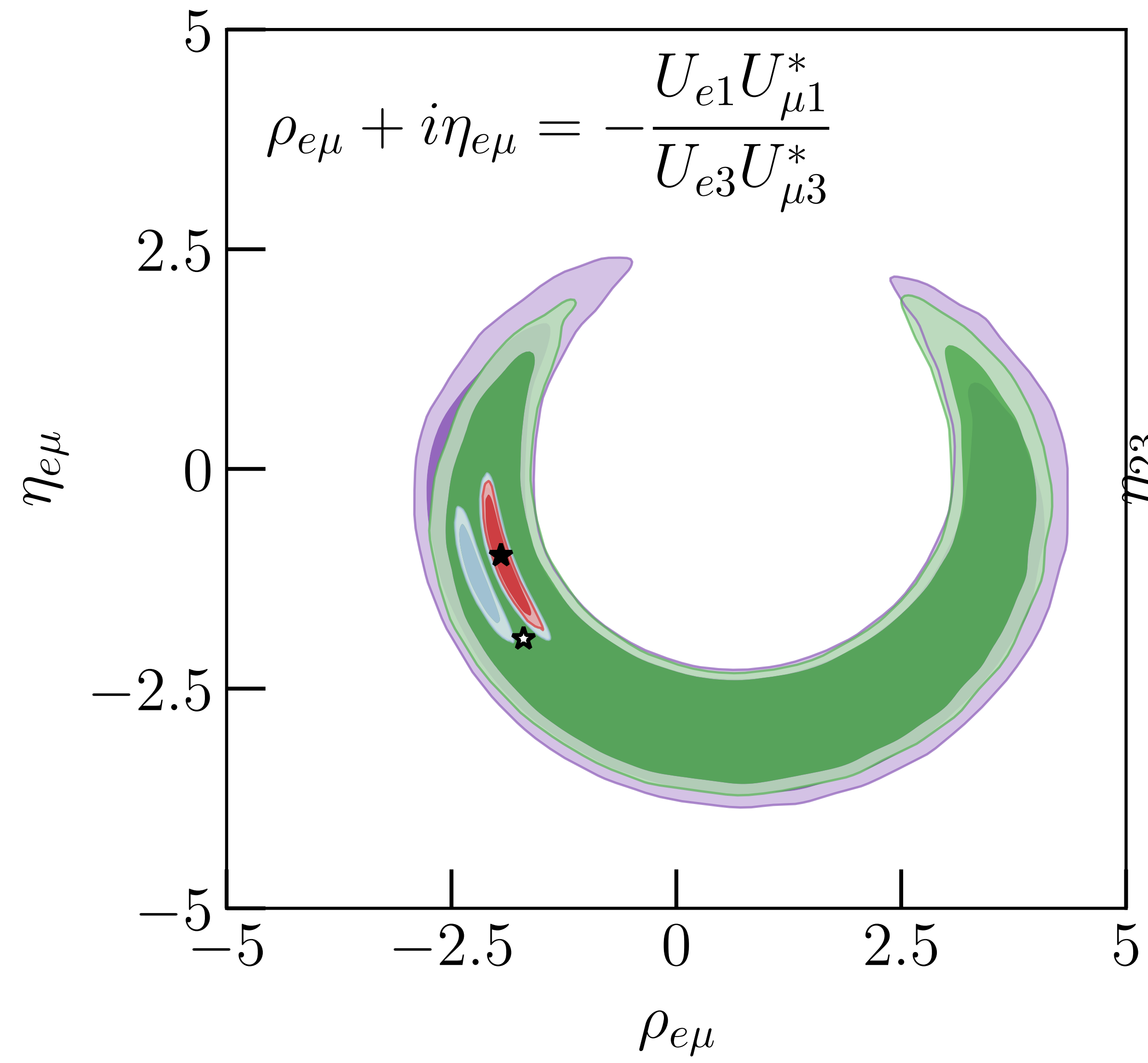


Why Stress-test?

Tour de force results on quark mixing from the CKMfitter group

Closure tests allow for honest evaluation of our models — is the “three massive neutrinos” paradigm good or not?



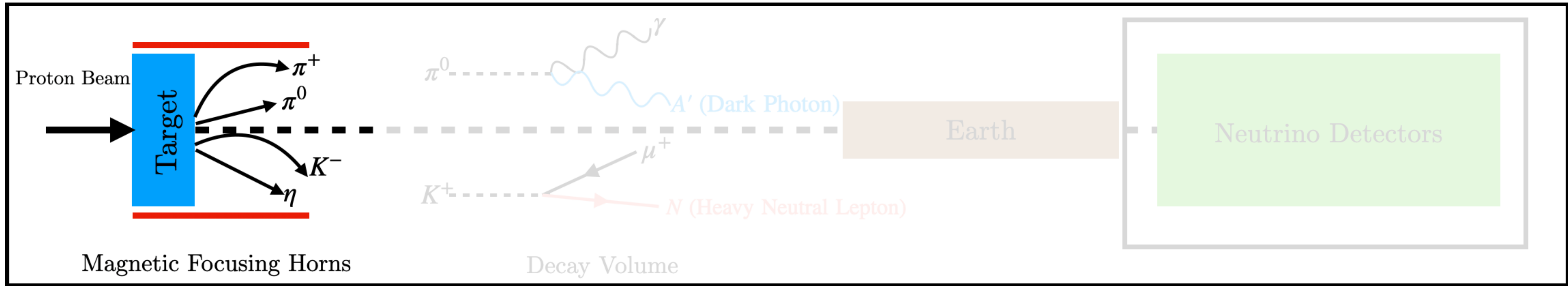


Analogue to the Lepton Sector

Ellis, Kelly, Li [\[2004.13719\]](#)

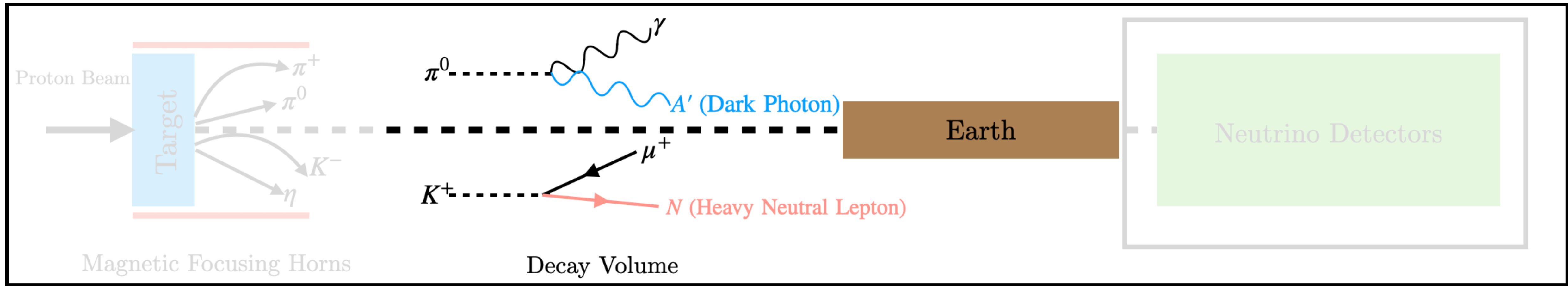
Tau Neutrinos and Dark Sectors

Neutrino Facilities as Dark Sector Machines



1) Charged and Neutral Mesons are produced in the high-energy/high-intensity proton collisions.

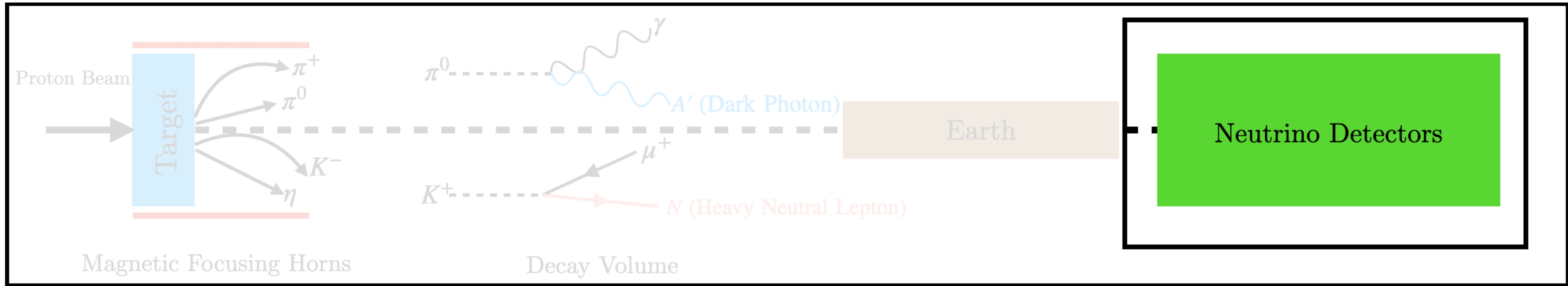
Neutrino Facilities as Dark Sector Machines



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Neutrino Facilities as Dark Sector Machines

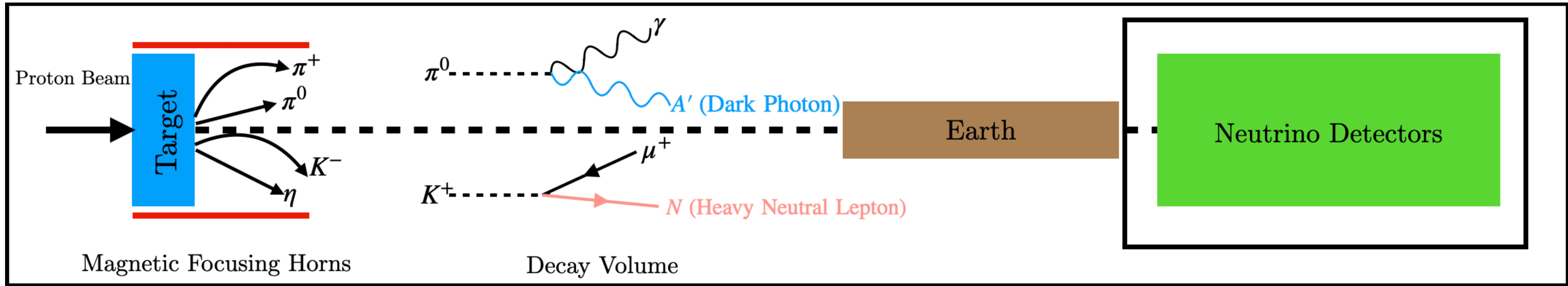


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Neutrino Facilities as Dark Sector Machines



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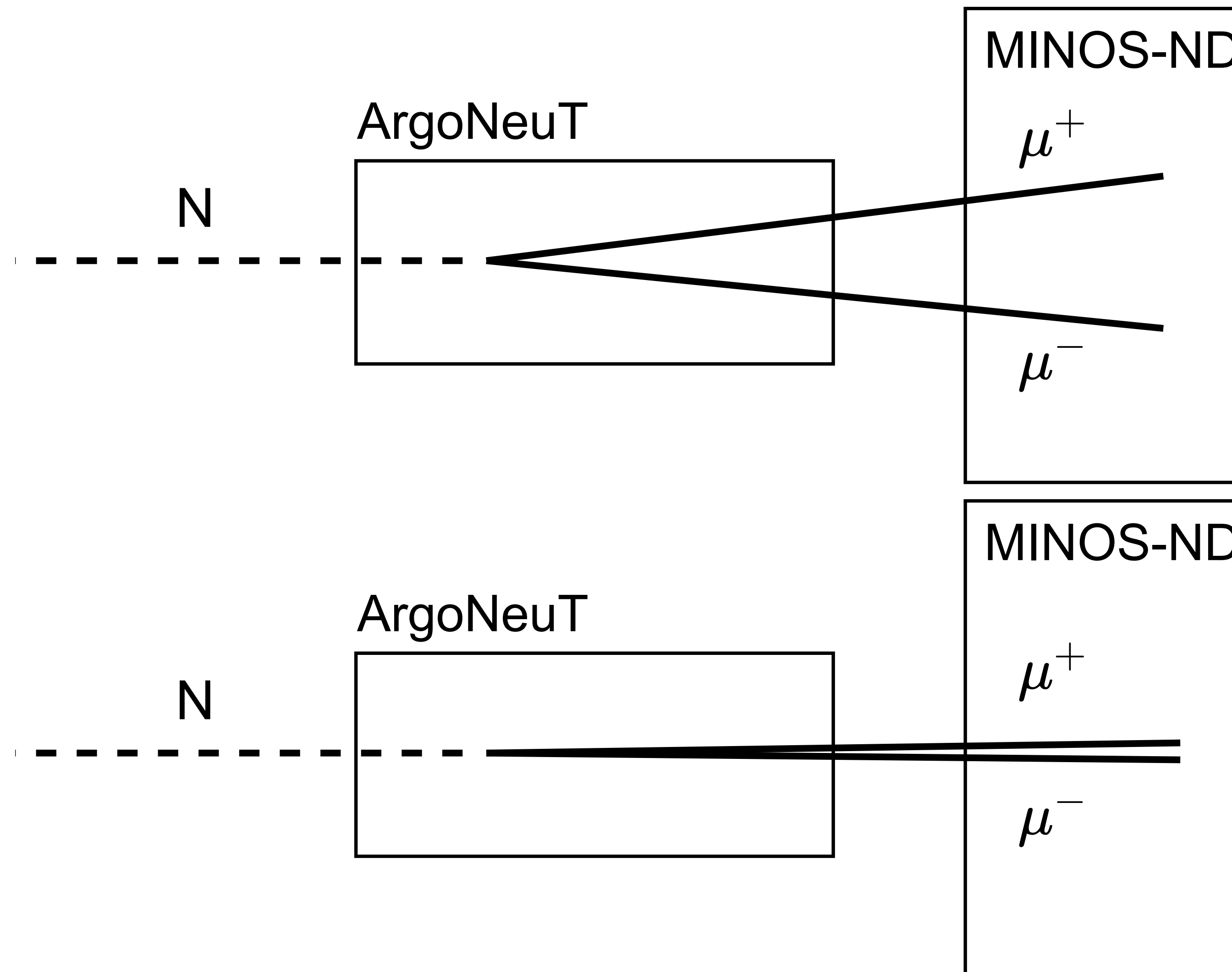
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ArgoNeuT

Tabletop ton-scale liquid-argon detector at Fermilab that took data in the NuMI beamline (120 GeV protons)

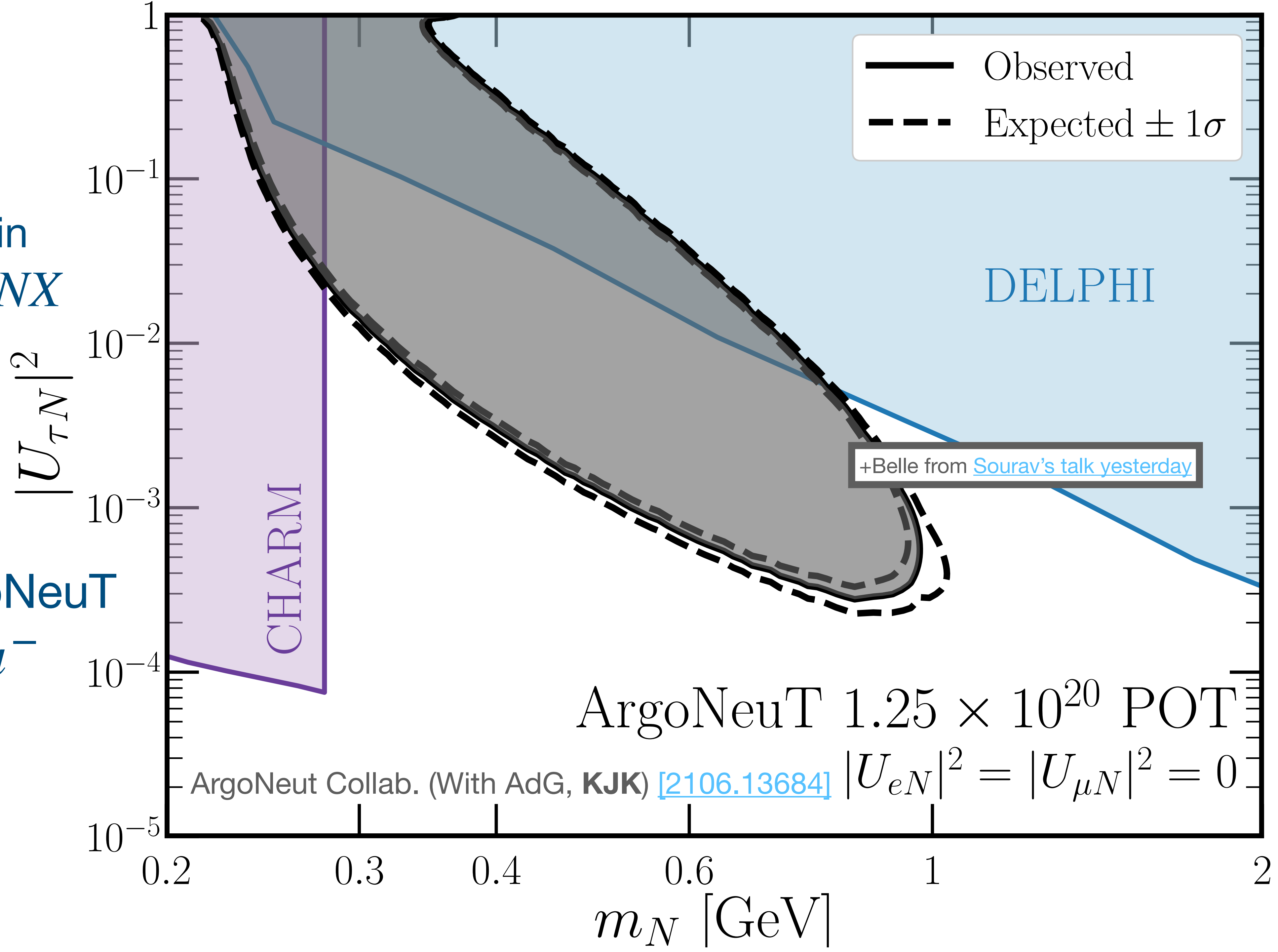
HNL production possible through D-meson production: $D^\pm \rightarrow \tau^\pm, \tau^\pm \rightarrow N$

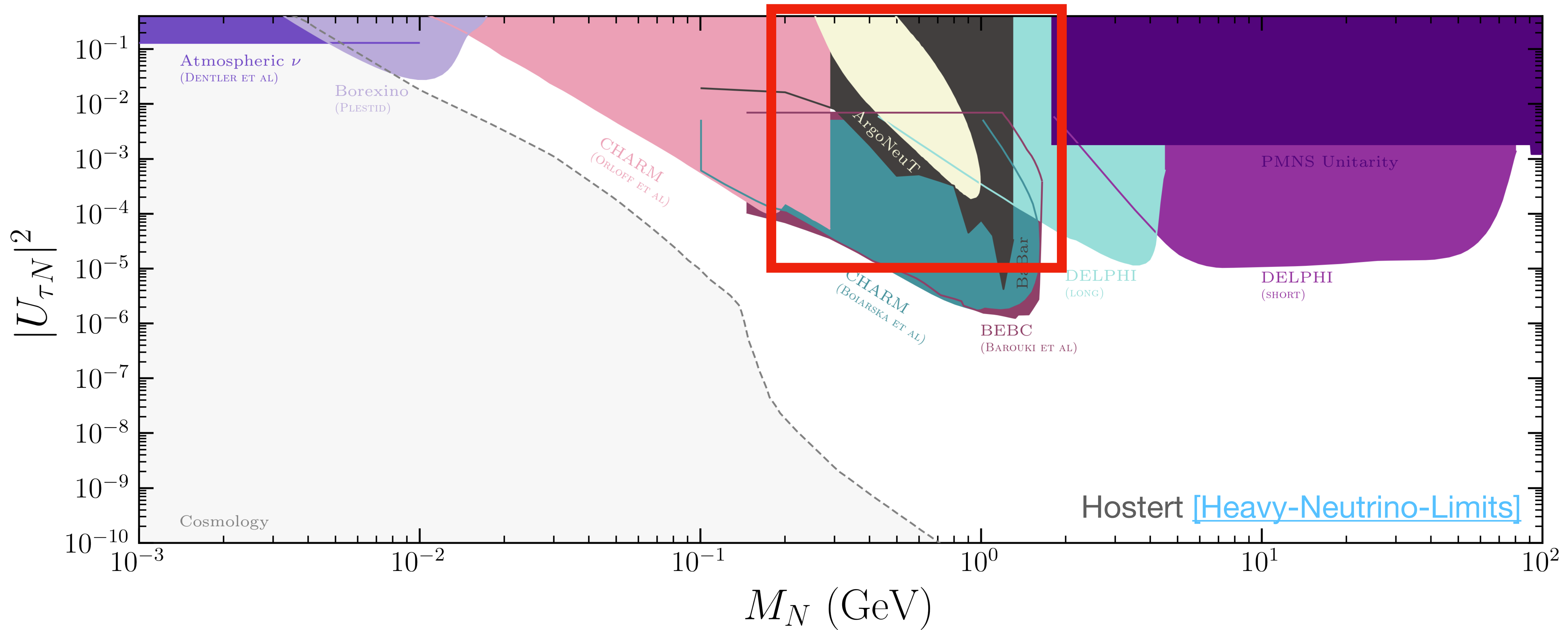


ArgoNeuT

HNLs produced in
 $D^\pm \rightarrow \tau^\pm, \tau^\pm \rightarrow NX$

Decay in/near ArgoNeuT
via $N \rightarrow \nu_\tau \mu^+ \mu^-$

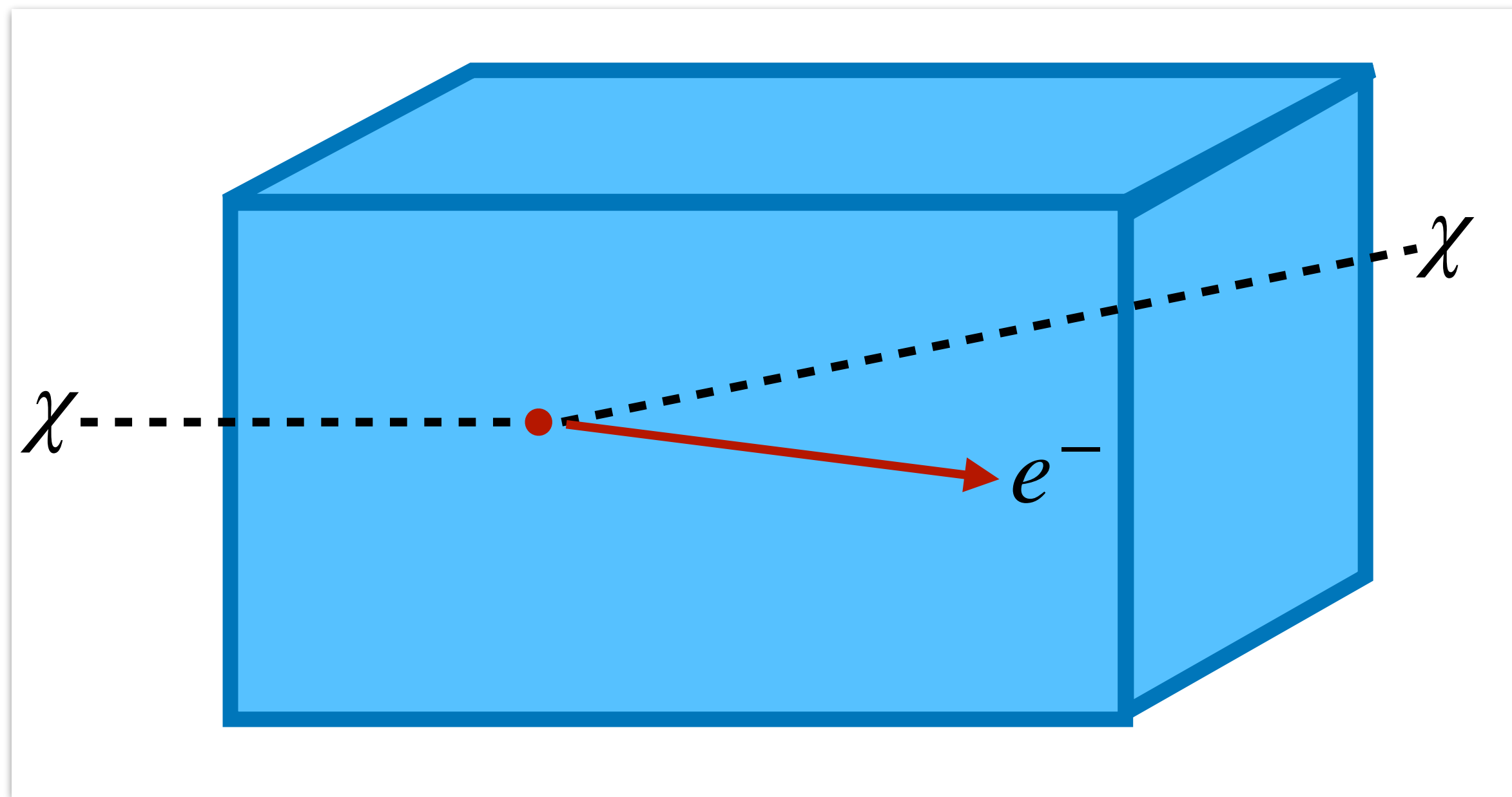




Full Disclosure...

Complementarity of Neutrino Detectors

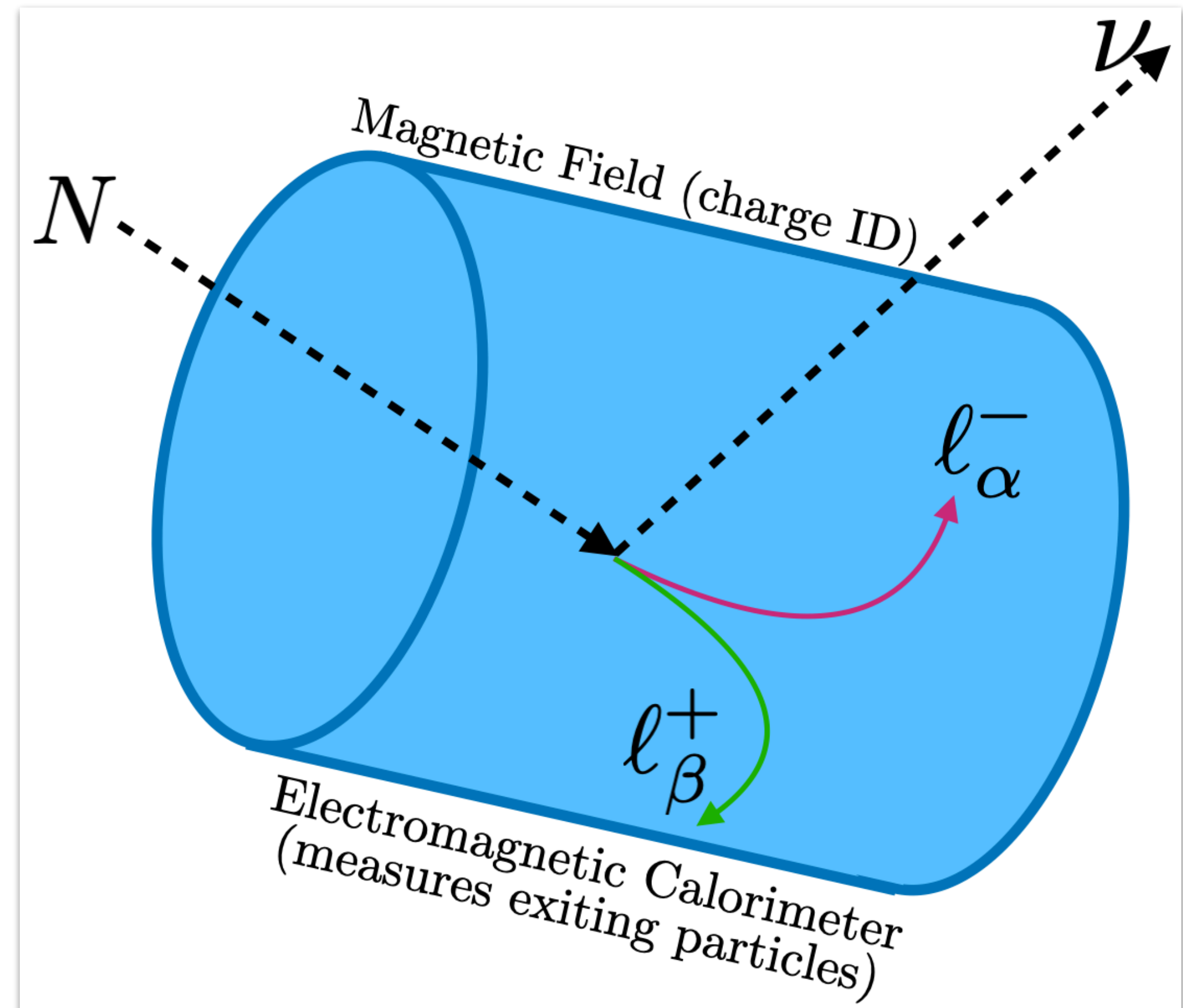
Liquid Detectors (SBND, ICARUS, etc.)



Large mass for rare-particle scattering

Excellent particle ID, energy resolution, etc.

Gaseous Detectors (DUNE NDGAr)

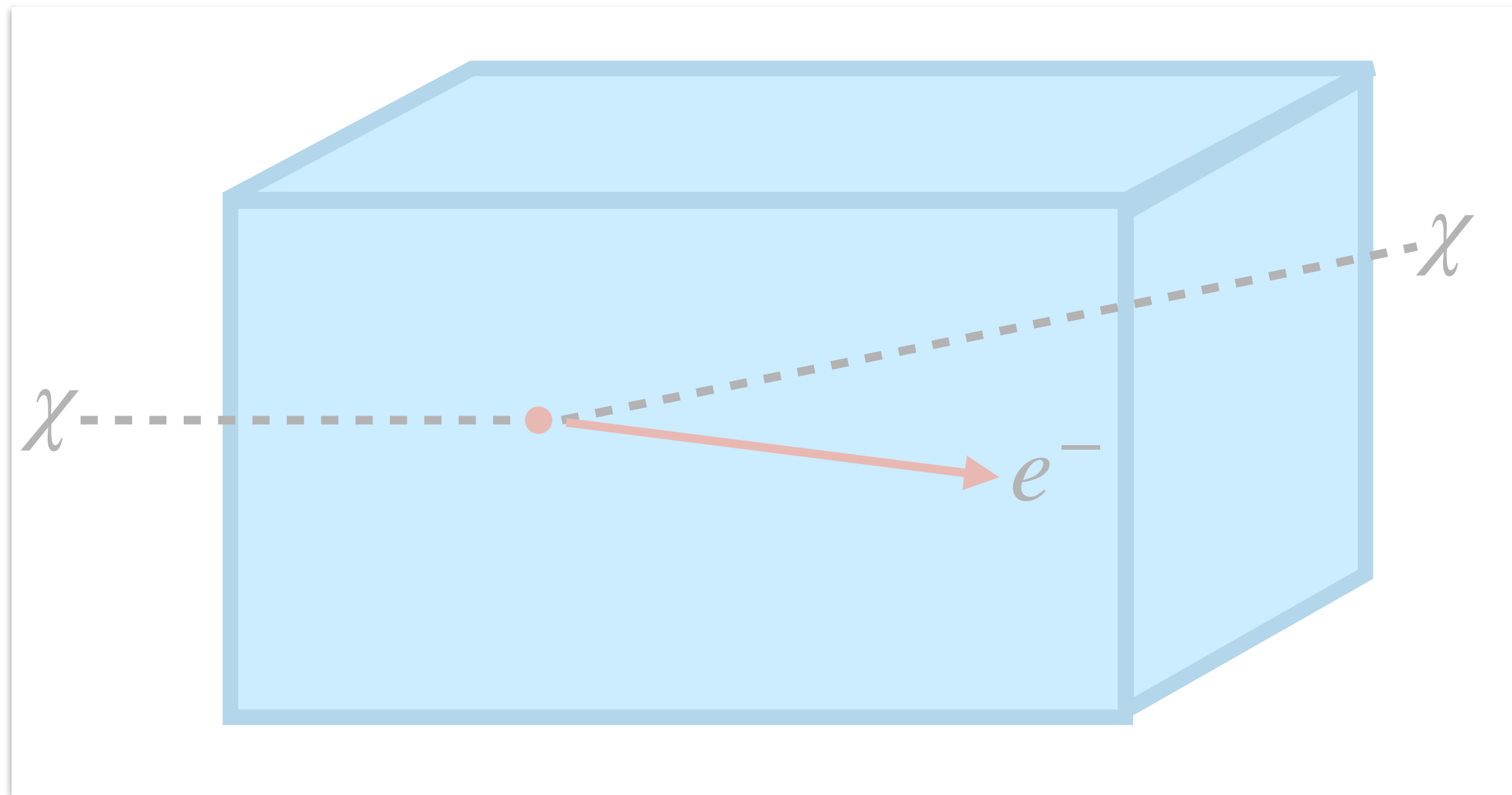


Decay Signal \propto Volume

Neutrino Scattering Backgrounds \propto Mass

Complementarity of Neutrino Detectors

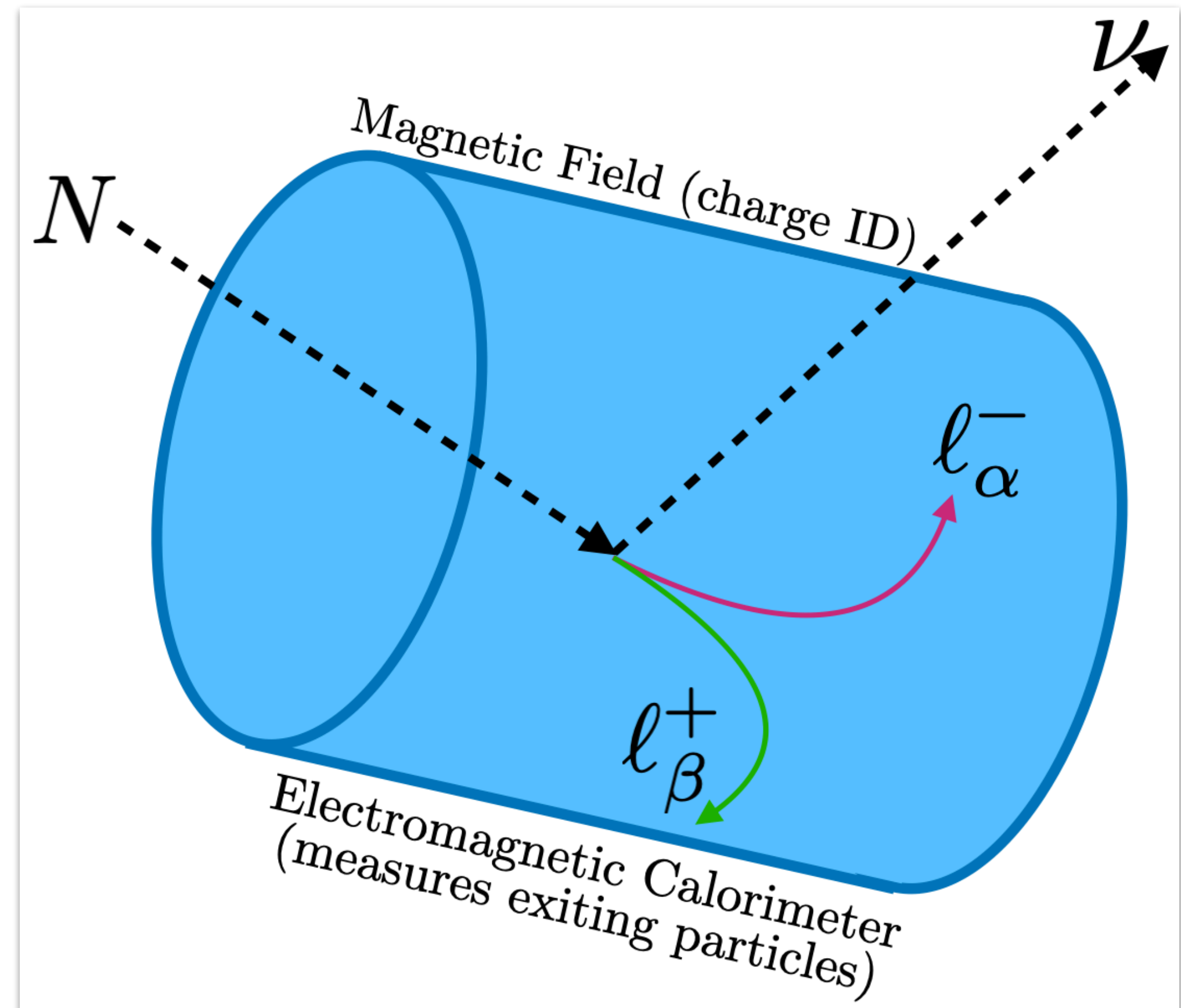
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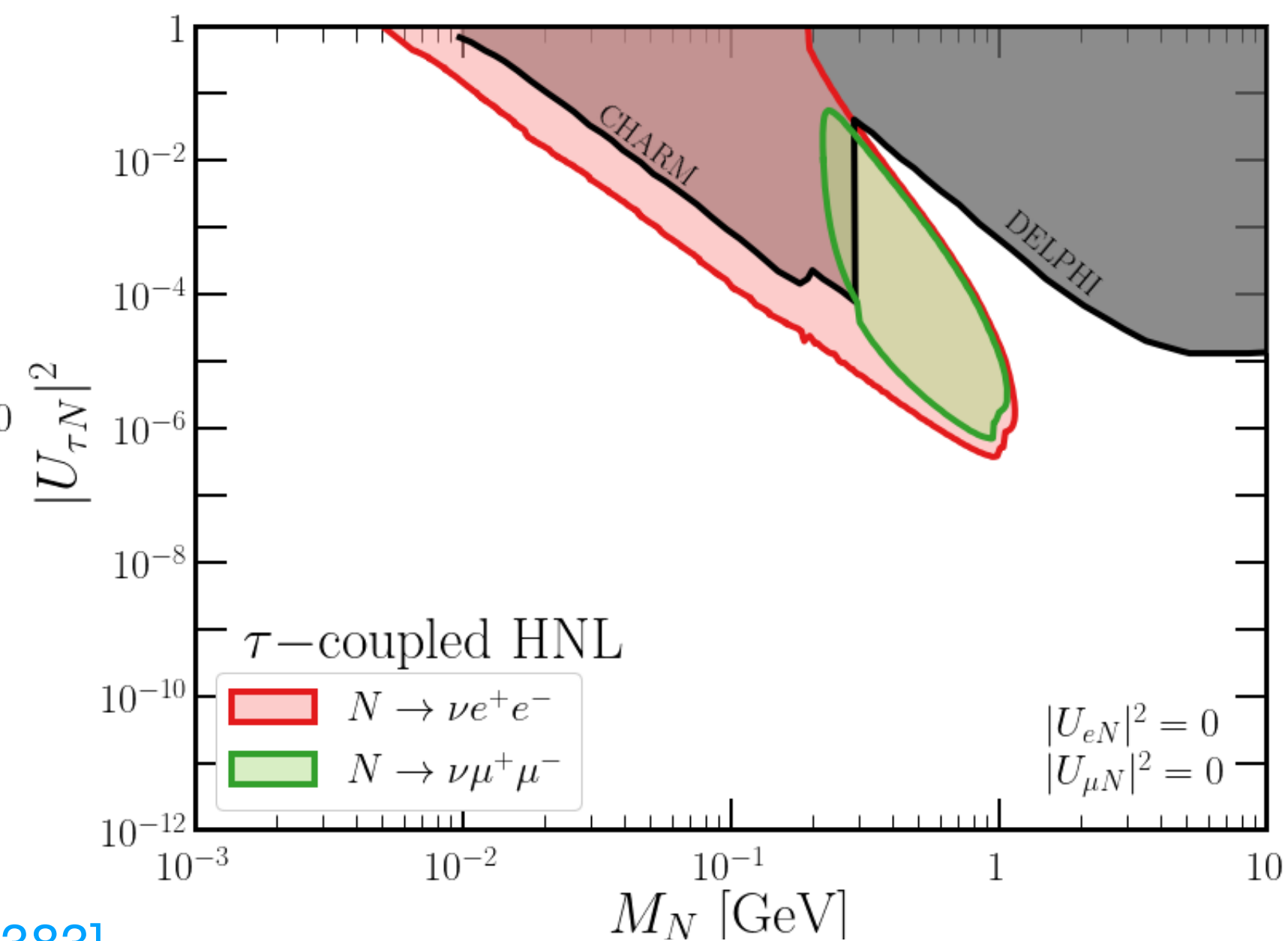
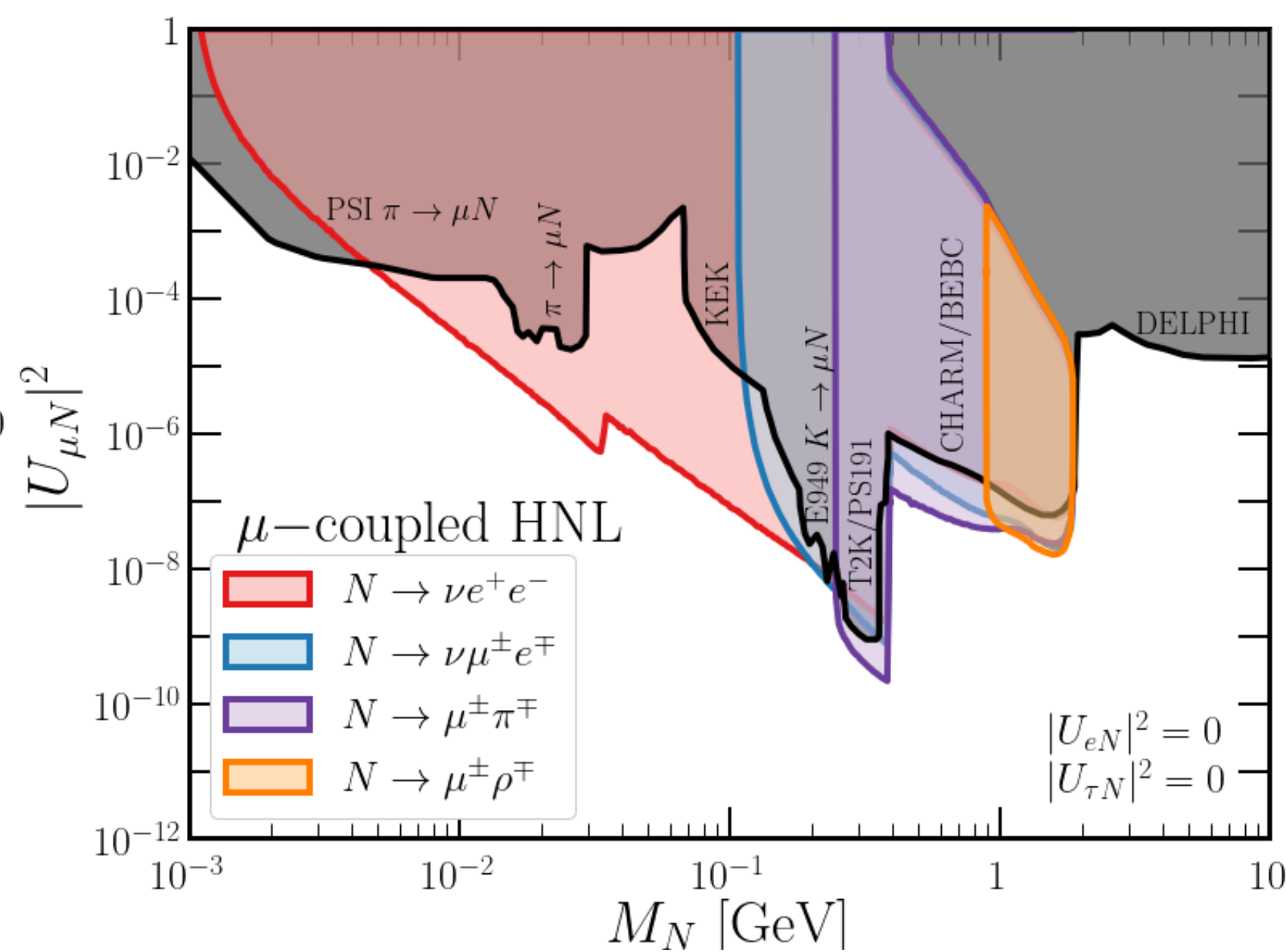
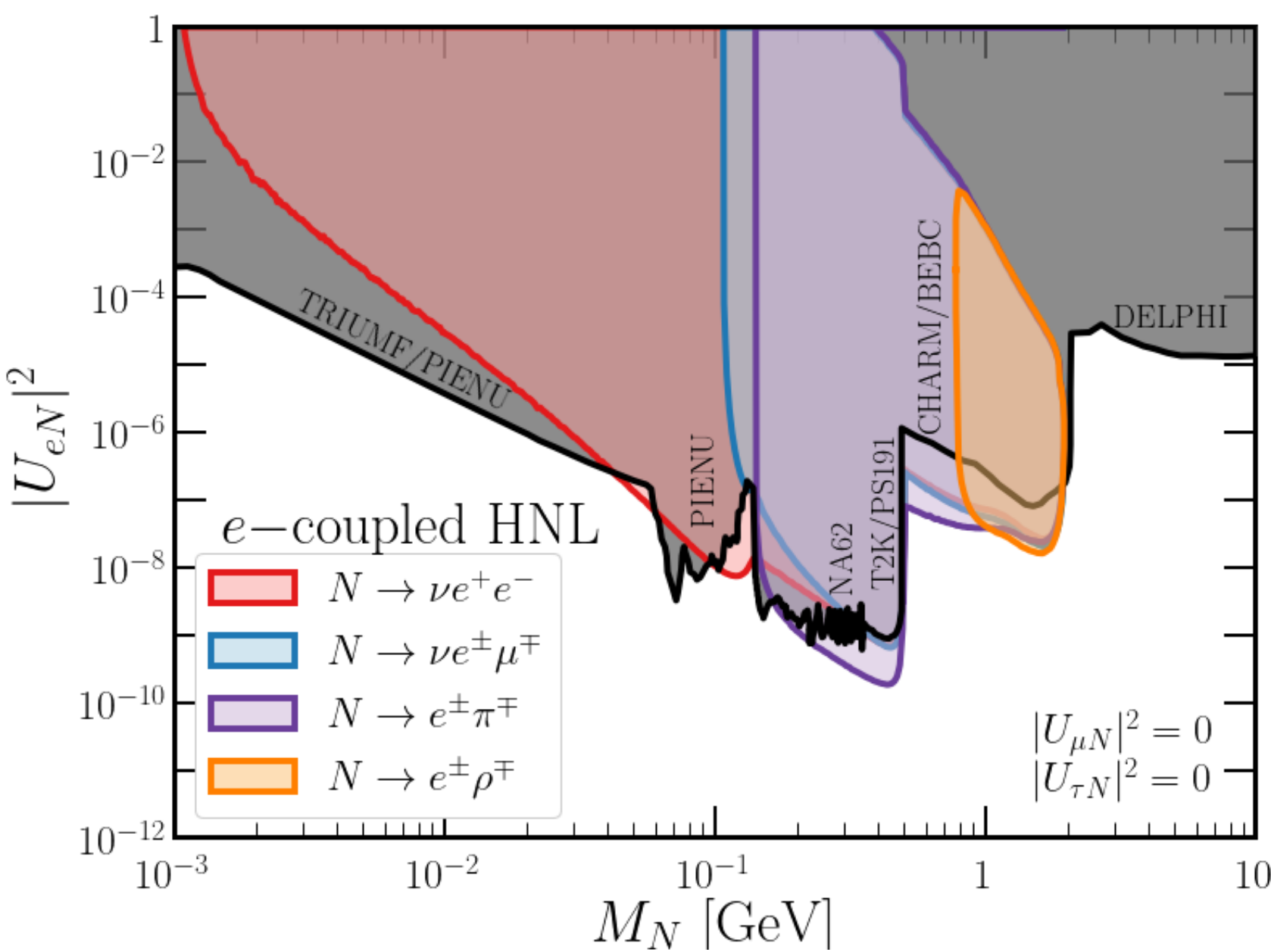


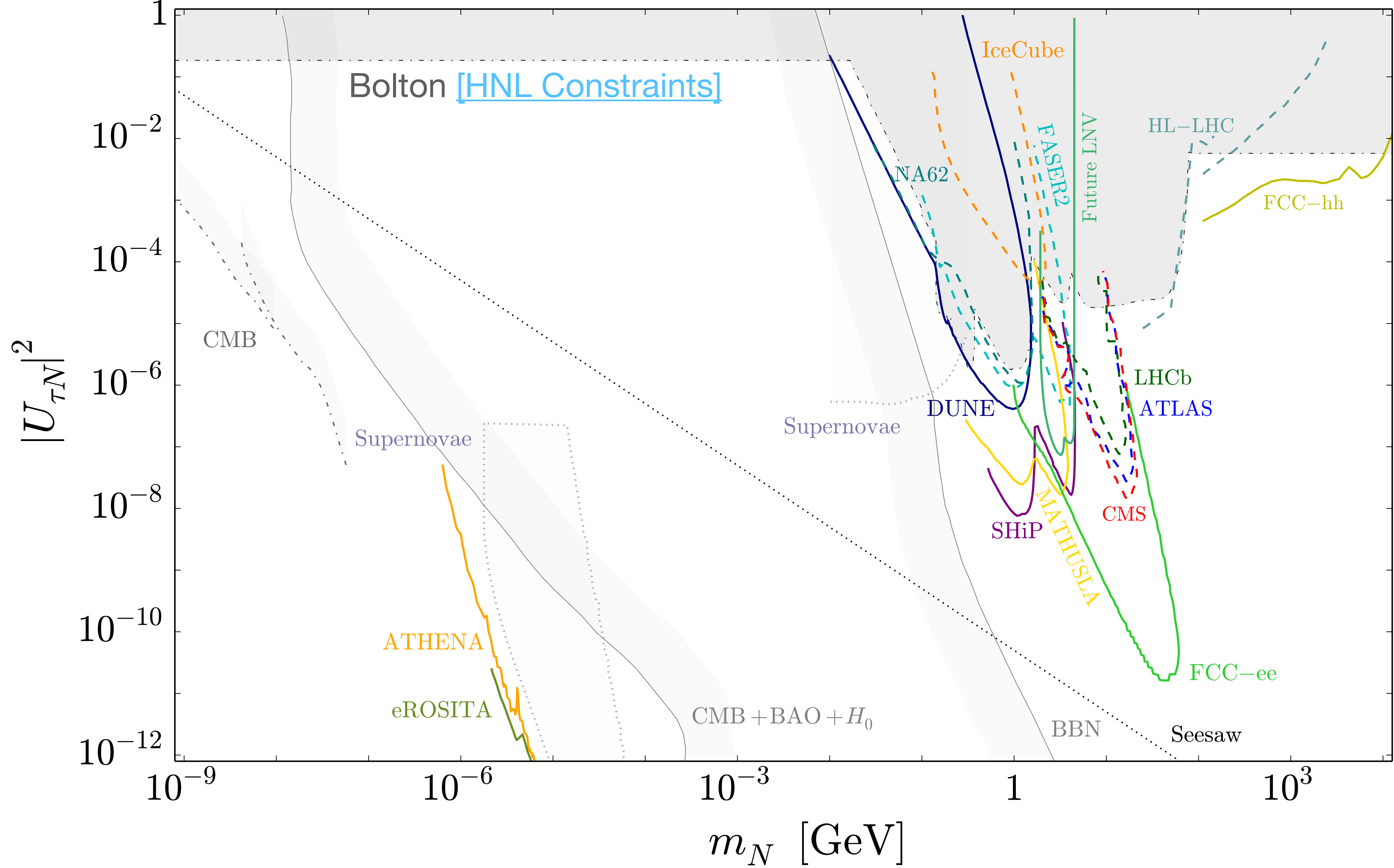
Decay Signal \propto Volume

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DUNE NDGA's HNL Capability

Berryman, de Gouvêa, Fox, Kayser, **KJK**, Raaf [\[1912.07622\]](https://arxiv.org/abs/1912.07622)





Looking Ahead...

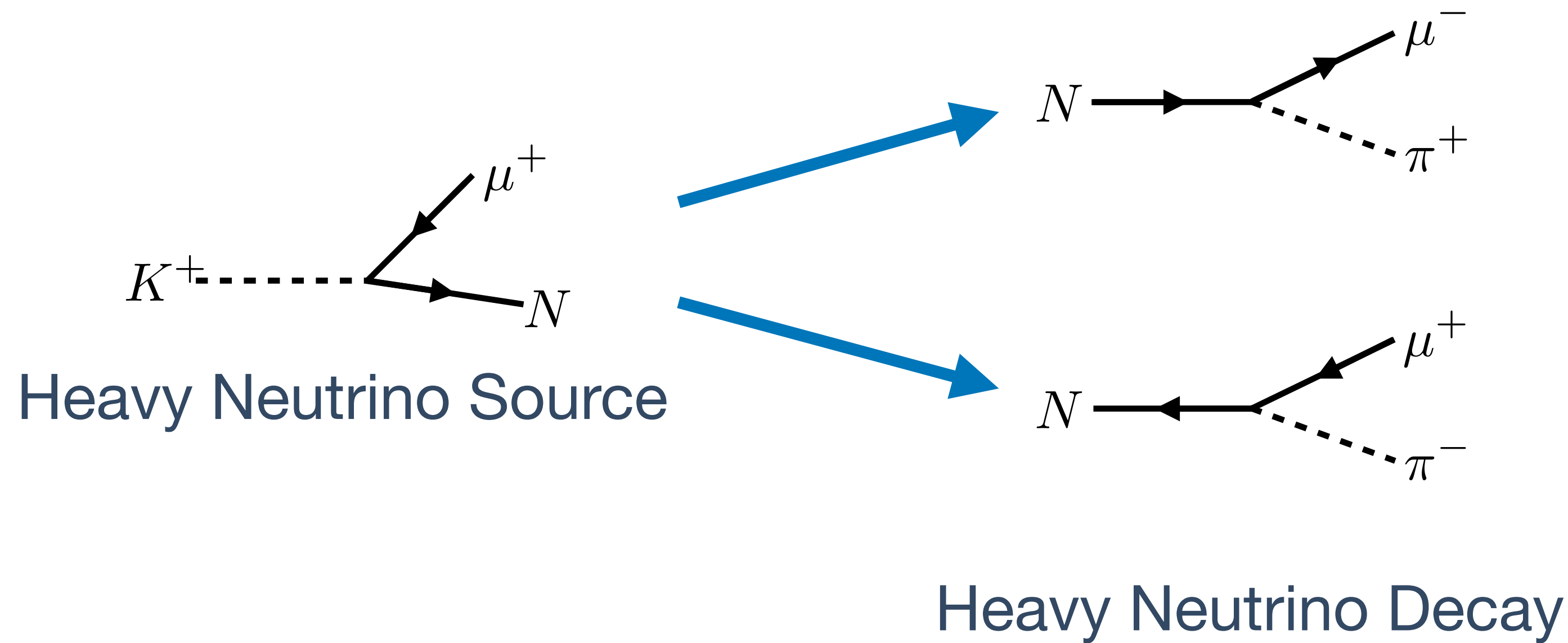
What do we do with a discovery?

Lepton-Number-Violation in a (Heavy) Neutrino Beam

Is the new particle a Dirac or Majorana Fermion?

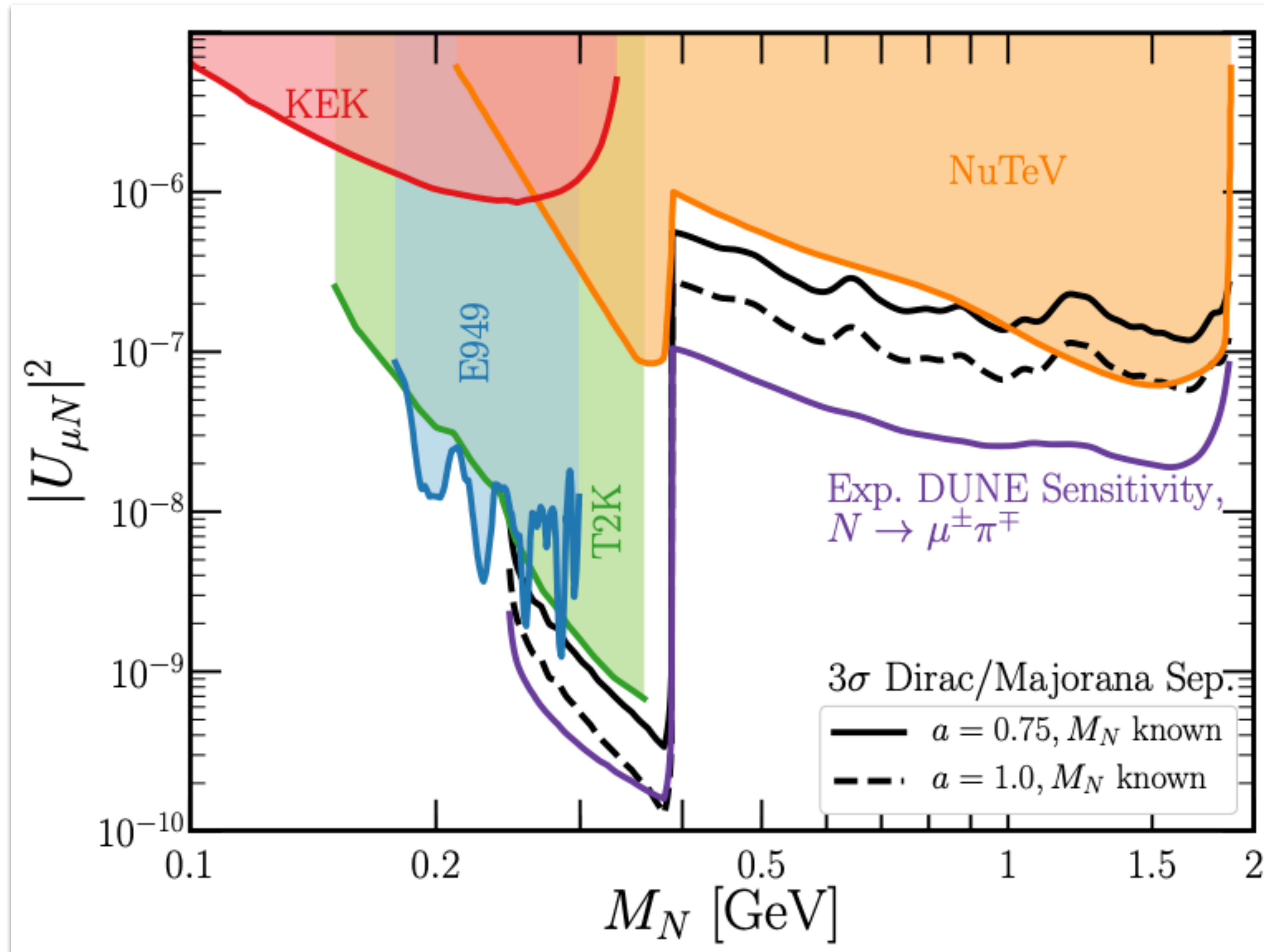


Do the new particle's interactions preserve or violate Lepton Number conservation?



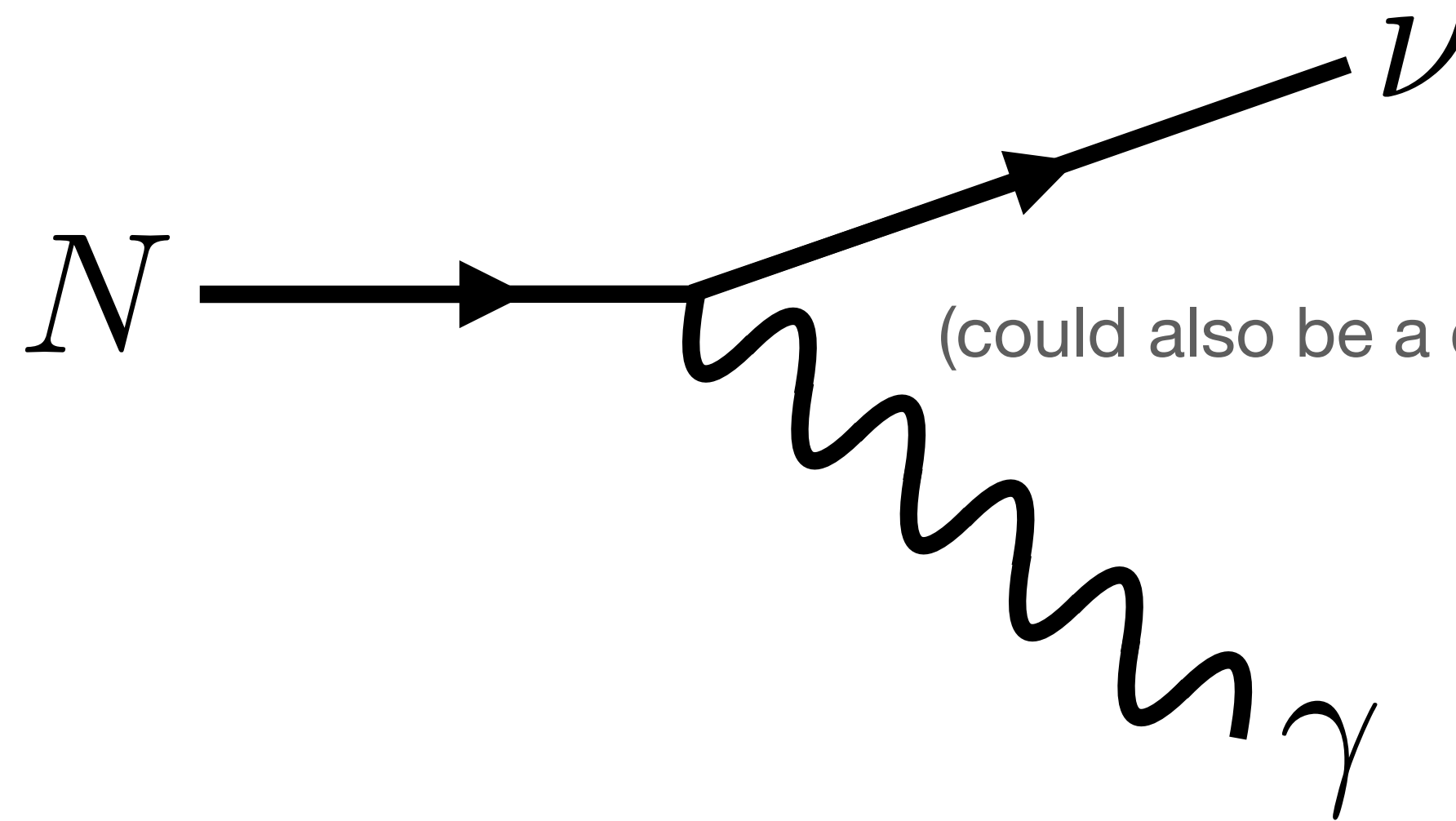
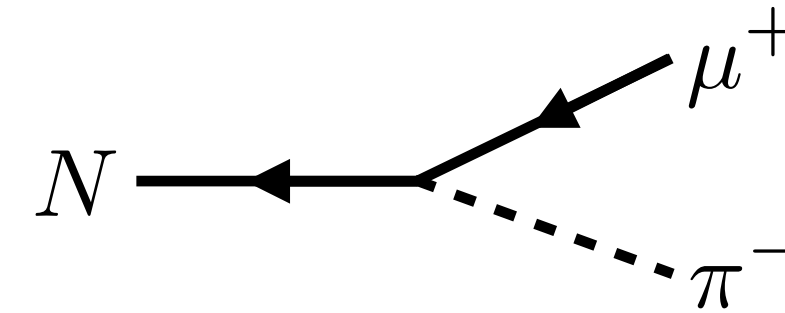
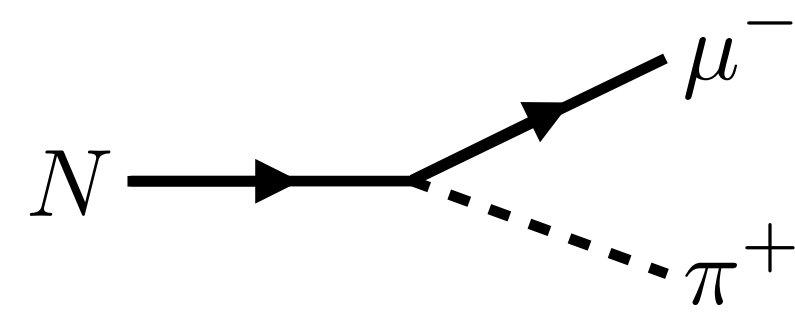
- Do these two chains occur with equal probability?

Next-Generation Prospects



What if we're not lucky?

- What if the HNL is lighter than the pion? Then there are no fully-visible final states to decay into, and Lepton Number can't be identified on an event-by-event basis.

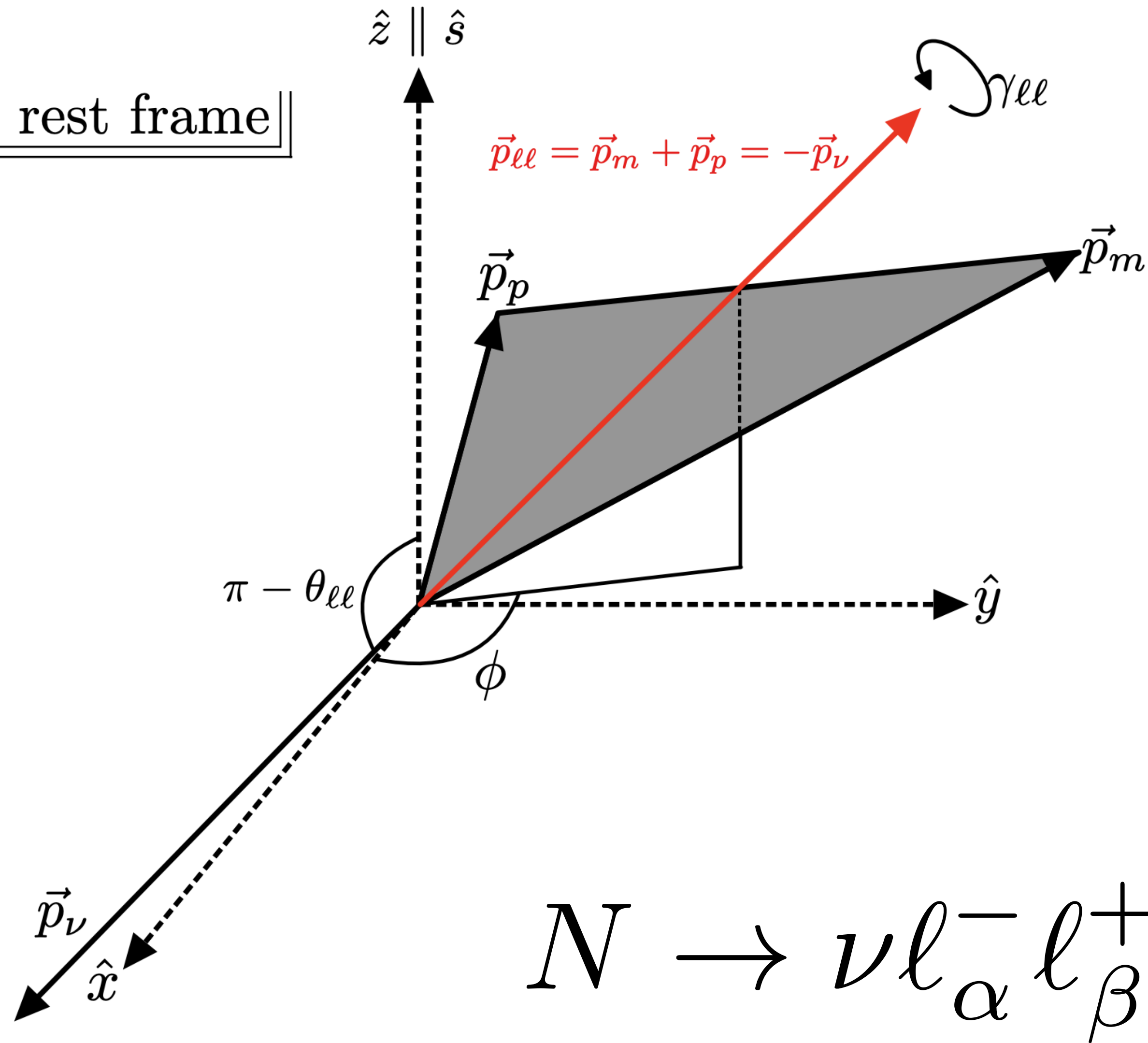


(could also be a charged-lepton pair instead of a photon, etc.)

Still, there are differences between Dirac/Majorana fermions: Measure the *distribution* of outgoing (visible) particles

Dirac vs. Majorana in Three-Body Decays

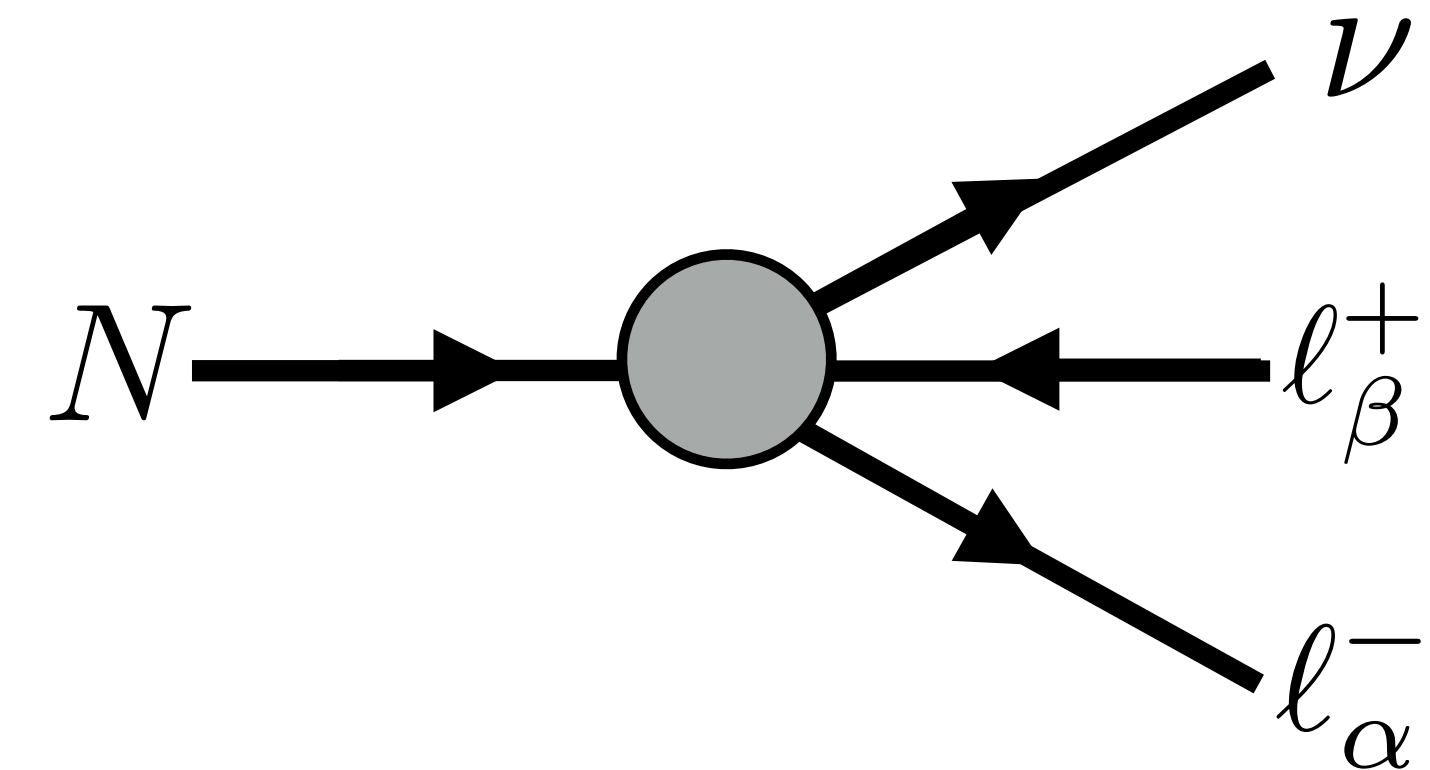
N rest frame



If N is a Majorana fermion, its decays are forward/backward symmetric if either

- The final-state charged leptons are identical (e.g. electron/positron pair).
- Whatever detection mechanism being used is charge-blind (can't distinguish electron from positron or muon from antimuon)

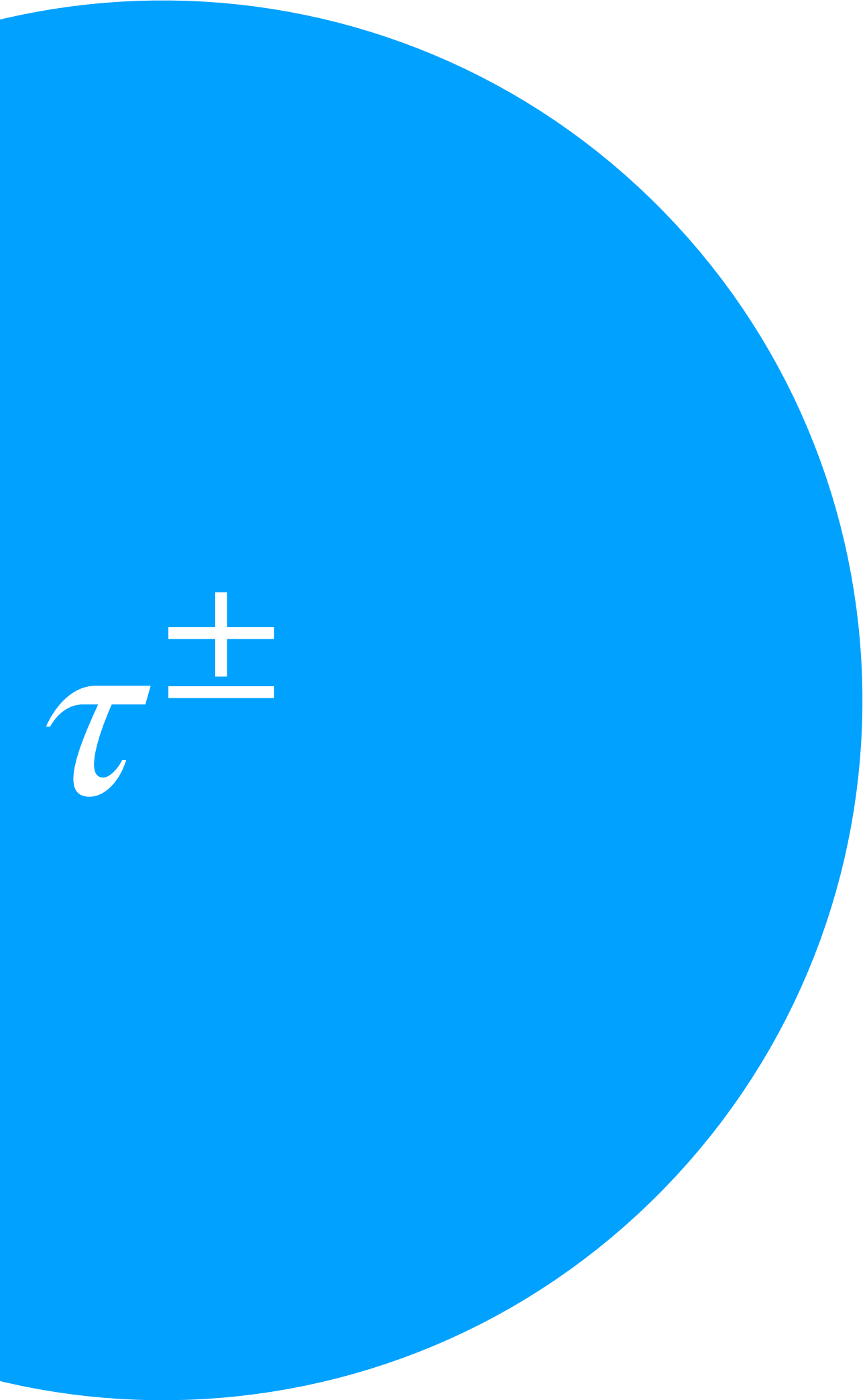
$$\mathcal{M}_1 = G_{NL} [\bar{u}_\nu \Gamma_N P_S u_N] [\bar{u}_\alpha \Gamma_L v_\beta]$$



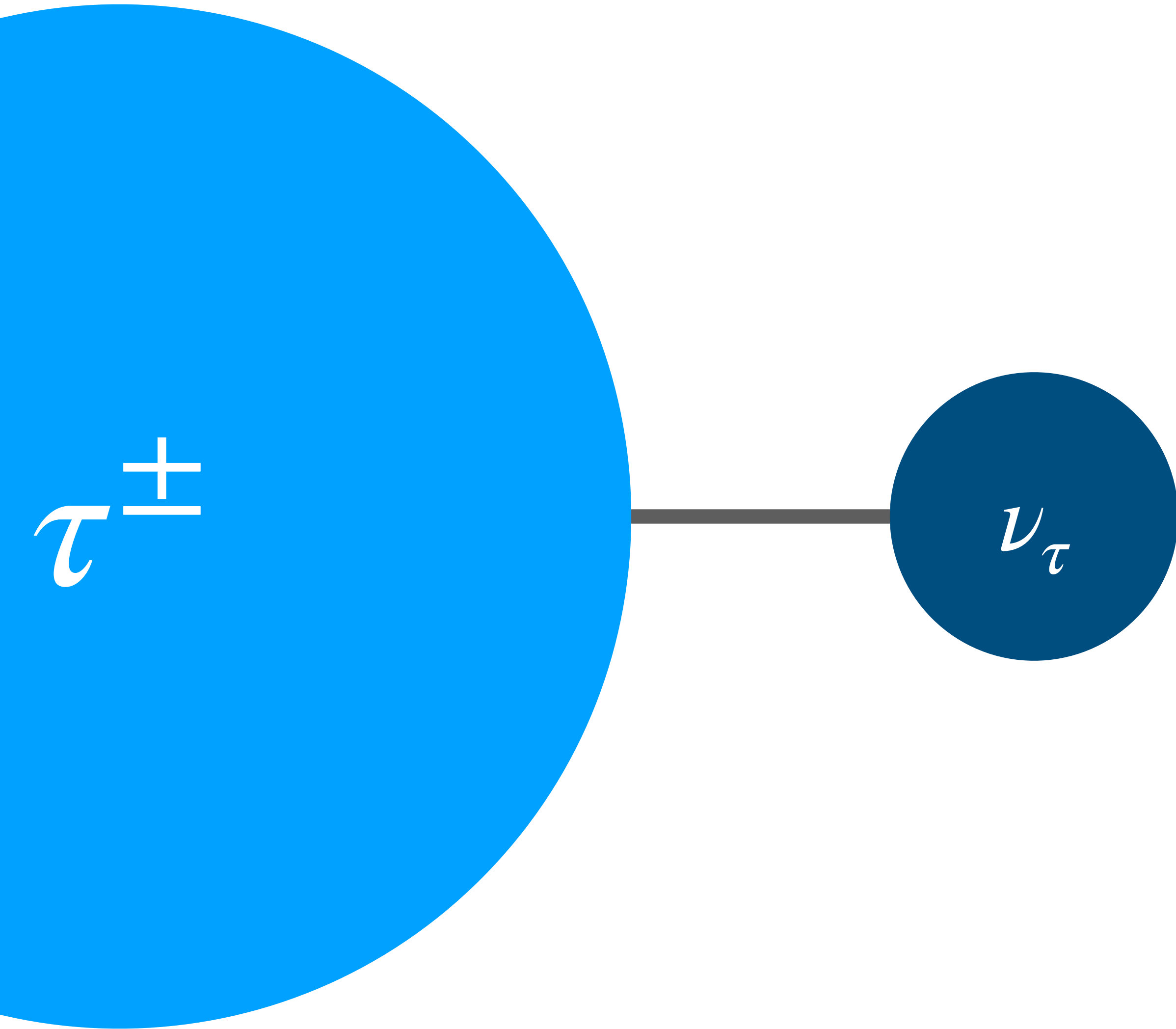
$$\Gamma_N, \Gamma_L \in \left\{ \mathbf{1}, \gamma^5, \gamma^\mu, \gamma^\mu \gamma^5, \sigma^{\mu\nu} = \frac{i}{2} [\gamma^\mu, \gamma^\nu] \right\}$$

de Gouvêa, Fox, Kayser, **KJK** [[2104.05719](#)]
 (Building off Balantekin, de Gouvêa, Kayser [[1808.10518](#)])

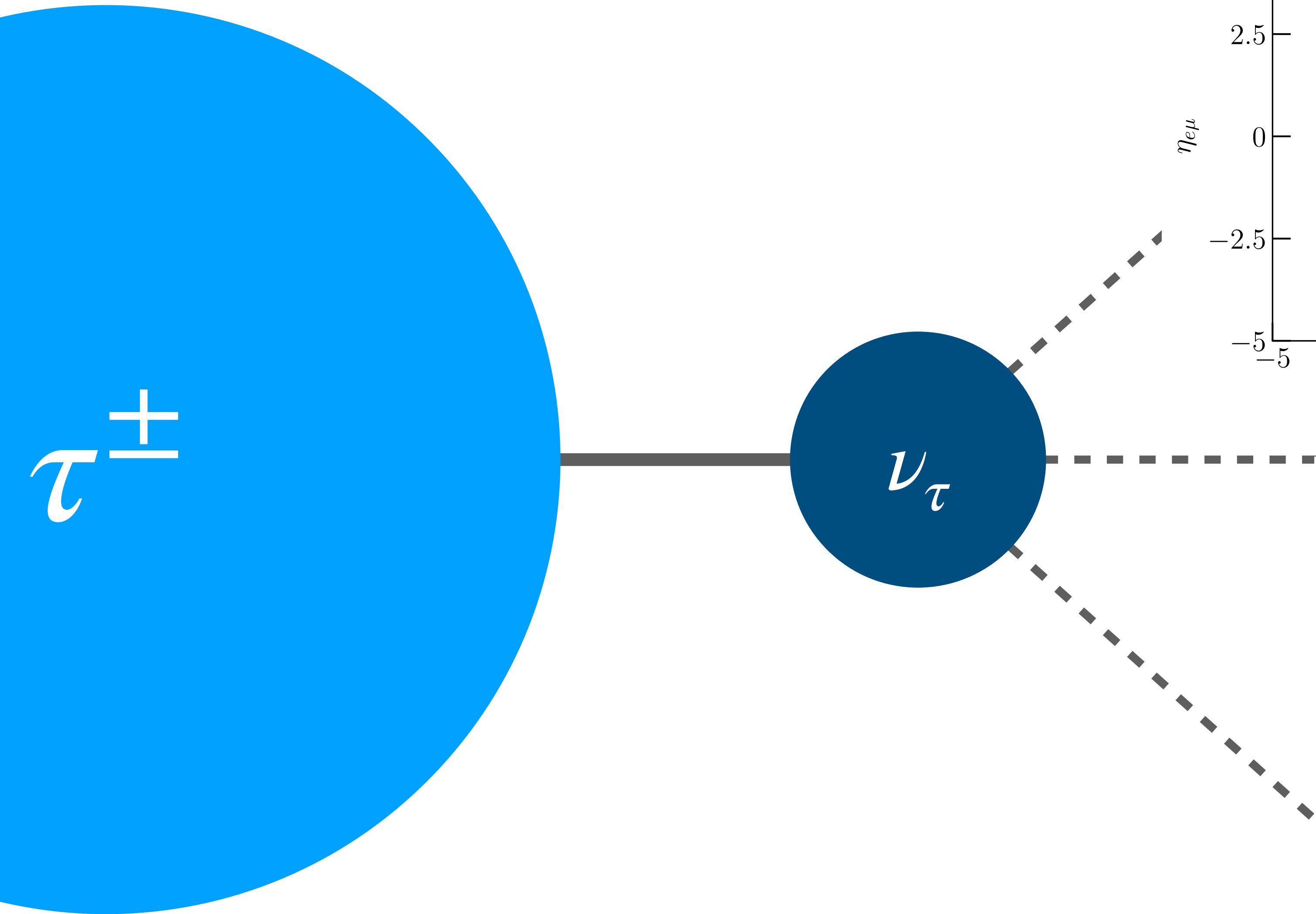
Takeaways


$$\tau^{\pm}$$

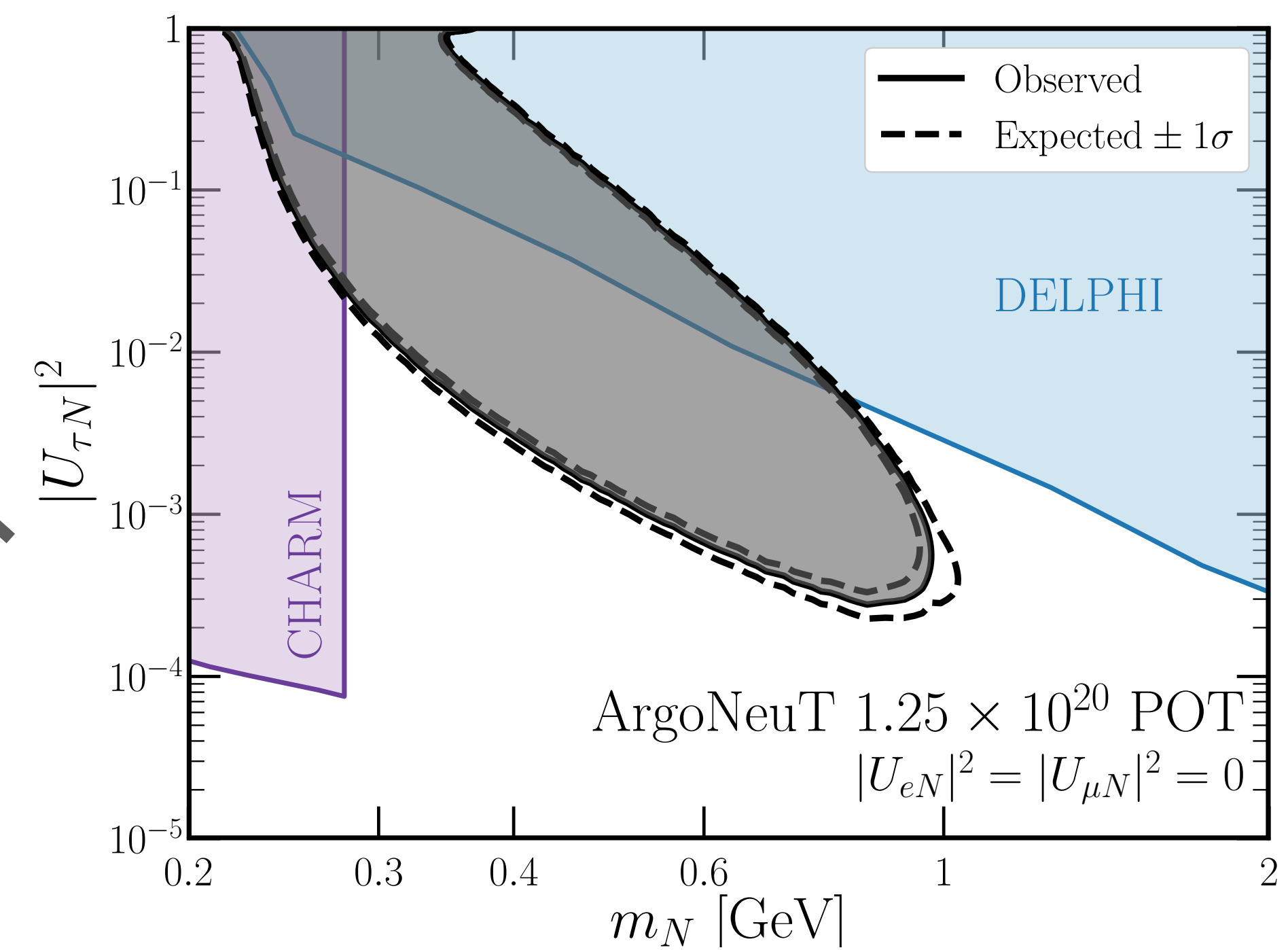
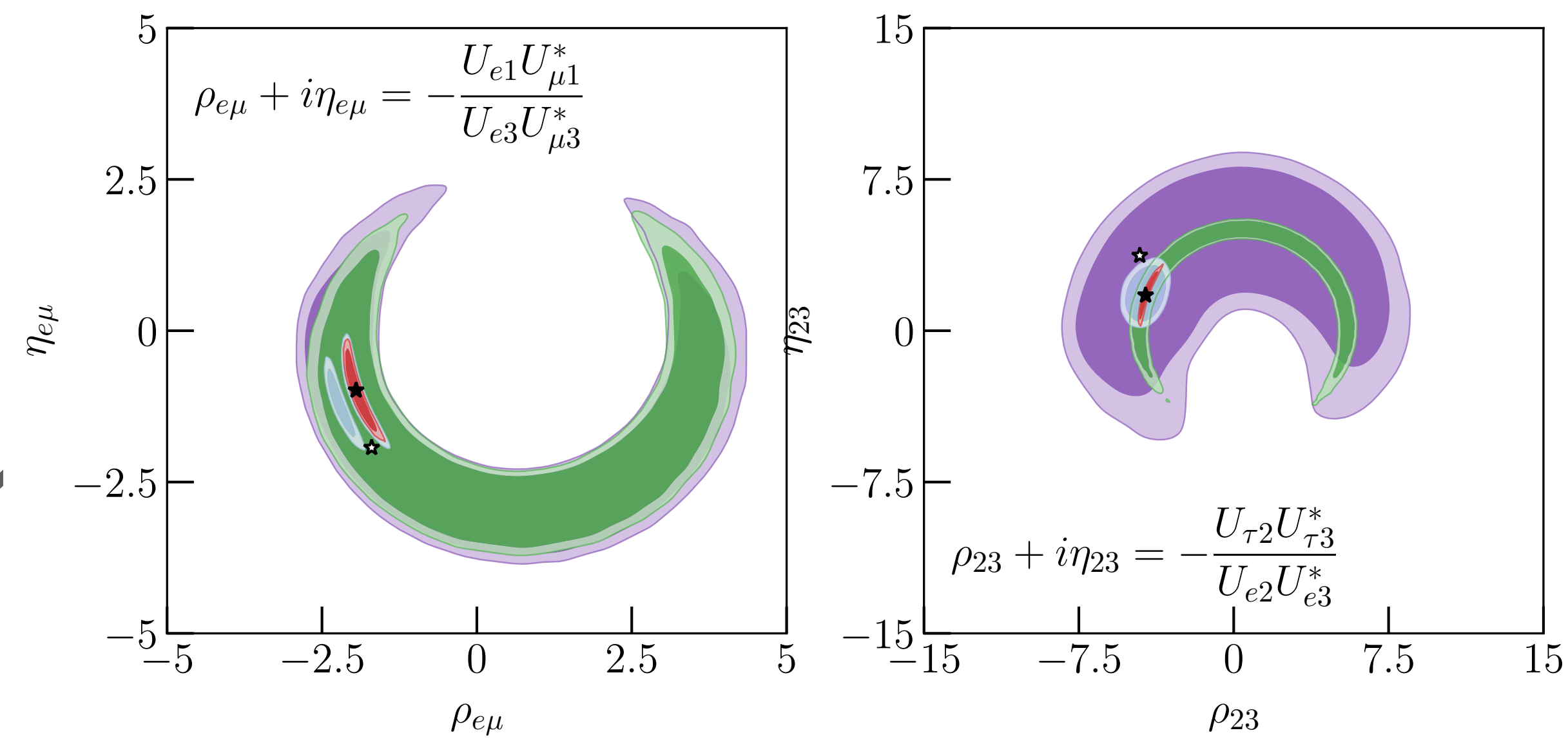
Takeaways



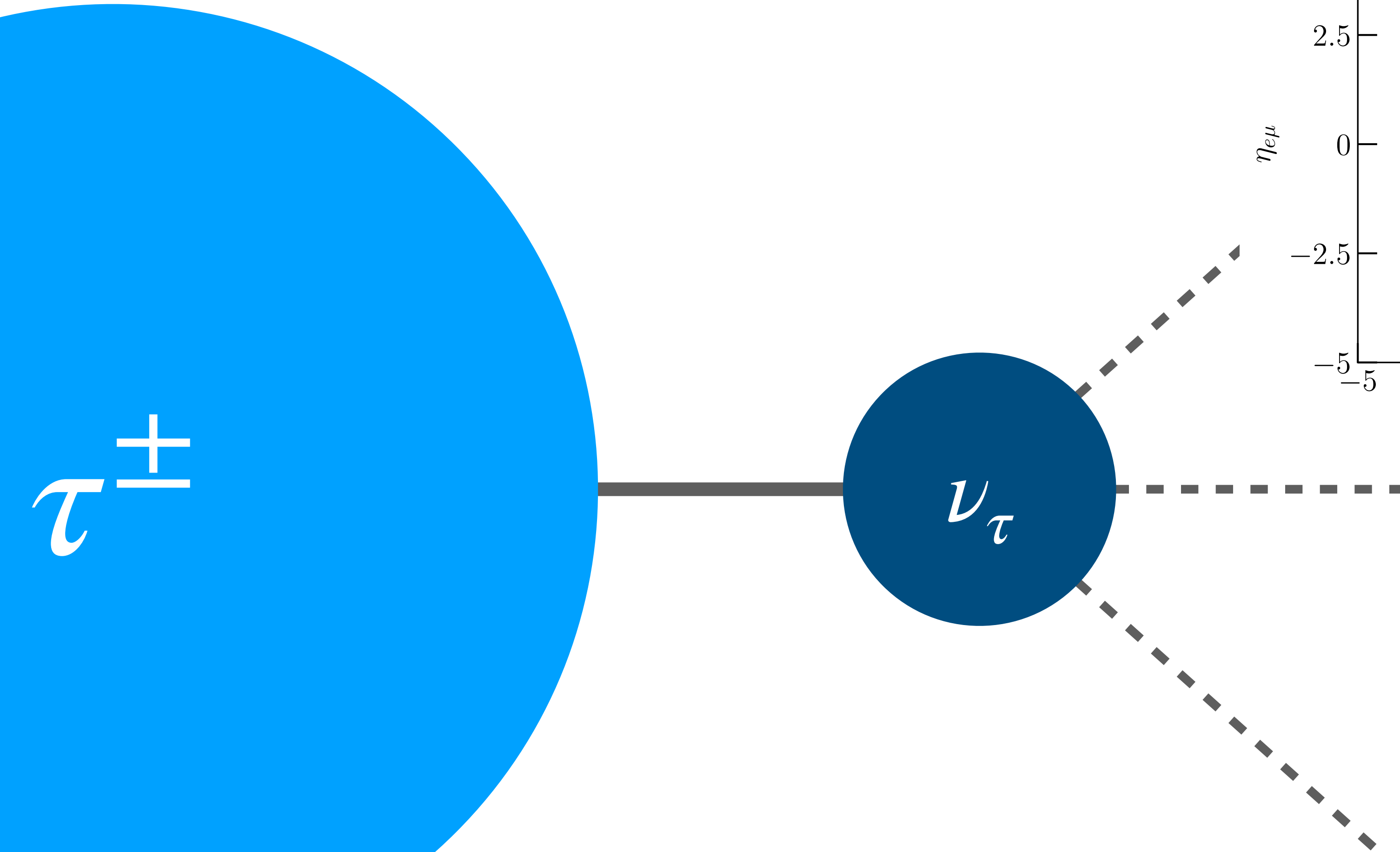
Takeaways



- All Current Data
- All Future Data
- All Current Data, Unitarity Assumed
- All Future Data, Unitarity Assumed



Takeaways



With a wealth of ν_τ and rich detectors now and in the coming decade, there's substantial new territory ahead. Stay tuned for results, and hopefully discoveries!

- All Current Data
- All Current Data, Unitarity Assumed
- All Future Data
- All Future Data, Unitarity Assumed

