Beth-Uhlenbeck Approach to Quark-Hadron Matter and Chemical Freeze-out in Heavy-Ion Collisions



David Blaschke (HZDR/CASUS, IFT UWr)

www.casus.science





2019/33/B/ST9/03059: Neutron stars: birth, structure and mergers 2021/43/P/ST2/03319: Bayesian analysis of the dense matter equation of state





Z HELMHOLTZ Centre for Environmental Research



Followit Ministry of Education and Research



CHAFT











QCD Phase Diagram





Landscape of our investigations



Figure from T. Kojo arXiv:1912.05326 [nucl-th]

QCD Phase Diagram





Landscape of our investigations



Gluons ↔ Vector mesons Quarks ↔ Baryons Goldstones ↔ Pseudoscalar mesons

Quark-Hadron Duality?

Mutual influence of Order parameters for χSB and CSC

From: T. Kojo, "QCD equations of state in quark-hadron continuity", Universe 4 (2018) 42

- T. Schaefer & F. Wilczek, Phys. Rev. Lett. 82 (1999) 3956
- C. Wetterich, Phys. Lett. B 462 (1999) 164
- T. Hatsuda, M. Tachibana, T. Yamamoto & G. Baym, Phys. Rev. Lett. 97 (2006) 122001

Contents





Introduction

- New research triangle Wroclaw Görlitz Dresden/Rossendorf: UWr CASUS & DZA HZDR
- Landscape of investigations: QCD Phase Diagram

Towards a unified approach to quark-nuclear matter

- Generalized Φ -derivable approach with clusters; cluster virial expansion
- Hadrons (mesons, baryons, multiquark states) as clusters in quark matter Mott dissociation of clusters
- Beth-Uhlenbeck approach to thermodynamics of quark-hadron matter
- Chemical freeze-out as "inverse" of the Mott effect for hadrons (χ SB) and nuclear clusters (Pauli blocking)

Relativistic density functionals for quark matter with confinement

- Density functional for warm, dense quark matter; chiral symmetry breaking and color superconductivity
- Quark confinement as density functional → effective Nambu model with density-dependent couplings
- Phase transition construction and hybrid neutron star properties

Unified EOS for quark-hadron matter



Cluster virial expansion & Beth-Uhlenbeck EoS





Clustering aspects in the QCD phase diagram



From: N.-U. Bastian, D.B., et al., Universe 4 (2018) 67; arxiv:1804.10178



 Φ -derivable approach to cluster virial expansion

$$\Omega = \sum_{l=1}^{A} \Omega_{l} = \sum_{l=1}^{A} \left\{ c_{l} \left[\operatorname{Tr} \ln \left(-G_{l}^{-1} \right) + \operatorname{Tr} \left(\Sigma_{l} \ G_{l} \right) \right] + \sum_{\substack{i,j \\ i+j=l}} \Phi[G_{i}, G_{j}, G_{i+j}] \right\} ,$$

$$G_A^{-1} = G_A^{(0)^{-1}} - \Sigma_A, \ \Sigma_A(1 \dots A, 1' \dots A', z_A) = \frac{\delta \Phi}{\delta G_A(1 \dots A, 1' \dots A', z_A)}$$

Stationarity of the thermodynamical potential is implied

$$\frac{\delta\Omega}{\delta G_A(1\ldots A, 1'\ldots A', z_A)} = 0 \; .$$

Cluster virial expansion follows for this Φ -functional



Figure: The Φ functional for A-particle correlations with bipartitions A = i + j.



Green's function and T-matrix, separable approx.



The T_A matrix fulfills the Bethe-Salpeter equation in ladder approximation

$$T_{i+j}(1,2,\ldots,A;1',2',\ldots,A';z) = V_{i+j} + V_{i+j}G_{i+j}^{(0)}T_{i+j}$$

which in the separable approximation for the interaction potential,

$$V_{i+j} = \Gamma_{i+j}(1, 2, \ldots, i; i+1, i+2, \ldots, i+j)\Gamma_{i+j}(1', 2', \ldots, i'; (i+1)', (i+2)', \ldots, (i+j)'),$$

leads to the closed expression for the T_A matrix

$$T_{i+j}(1,2,\ldots,i+j;1',2',\ldots,(i+j)';z) = V_{i+j}\{1-\prod_{i+j}\}^{-1},\$$

with the generalized polarization function

$$\Pi_{i+j} = \operatorname{Tr}\left\{ \Gamma_{i+j} G_i^{(0)} \Gamma_{i+j} G_j^{(0)} \right\}$$

The one-frequency free i-particle Green's function is defined by the (i - 1)-fold Matsubara sum



Useful relationships for many-particle functions

$$G_{i+j}^{(0)} = G_{i+j}^{(0)}(1,2,\ldots,i+j;\Omega_{i+j}) = \sum_{\Omega_i} G_i^{(0)}(1,2,\ldots,i;\Omega_i) G_j^{(0)}(i+1,i+2,\ldots,i+j;\Omega_j) .$$

Another set of useful relationships follows from the fact that in the ladder approximation both, the full two-cluster (i + j particle) T matrix and the corresponding Greens' function

$$G_{i+j} = G_{i+j}^{(0)} \left\{ 1 - \Pi_{i+j} \right\}^{-1}$$
(1)

have similar analytic properties determined by the i + j cluster polarization loop integral and are related by the identity

$$T_{i+j}G_{i+j}^{(0)} = V_{i+j}G_{i+j} .$$
(2)

which is straightforwardly proven by multiplying Equation for the T_{i+j} - matrix with $G_{i+j}^{(0)}$ and using Equation (1). Since these two equivalent expressions in Equation (2) are at the same time equivalent to the two-cluster irreducible Φ functional these functional relations follow

$$T_{i+j} = \delta \Phi / \delta G_{i+j}^{(0)} ,$$

$$V_{i+j} = \delta \Phi / \delta G_{i+j} .$$



Generalized Beth-Uhlenbeck EOS from Φ -deriv.

Consider the partial density of the A-particle state defined as

$$n_{A}(T,\mu) = -\frac{\partial\Omega_{A}}{\partial\mu} = -\frac{\partial}{\partial\mu}d_{A}\int\frac{d^{3}q}{(2\pi)^{3}}\int\frac{d\omega}{2\pi}\left[\ln\left(-G_{A}^{-1}\right) + \operatorname{Tr}\left(\Sigma_{A} \ G_{A}\right)\right] + \sum_{\substack{i,j\\i+j=A}}\Phi[G_{i},G_{j},G_{i+j}]$$
spectral representation for $F(\omega)$ and Matsubara summation

Using spectral representation for $F(\omega)$ and Matsubara summation

$$F(iz_n) = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{\mathrm{Im}F(\omega)}{\omega - iz_n}, \quad \sum_{z_n} \frac{c_A}{\omega - iz_n} = f_A(\omega) = \frac{1}{\exp[(\omega - \mu)/T] - (-1)^A}$$

with the relation $\partial f_A(\omega)/\partial \mu = -\partial f_A(\omega)/\partial \omega$ we get for Equation (3) now

$$n_{A}(T,\mu) = -d_{A}\int \frac{d^{3}q}{(2\pi)^{3}}\int \frac{d\omega}{2\pi}f_{A}(\omega)\frac{\partial}{\partial\omega}\left[\operatorname{Im}\ln\left(-G_{A}^{-1}\right) + \operatorname{Im}\left(\Sigma_{A} \ G_{A}\right)\right] + \sum_{\substack{i,j\\i+j=A}} \frac{\partial\Phi[G_{i},G_{j},G_{A}]}{\partial\mu},$$

where a partial integration over ω has been performed For two-loop diagrams of the sunset type holds a cancellation³ which generalize here for cluster states

$$d_A \int \frac{d^3 q}{(2\pi)^3} \int \frac{d\omega}{2\pi} f_A(\omega) \frac{\partial}{\partial \omega} \left(\operatorname{Re} \Sigma_A \operatorname{Im} G_A \right) - \sum_{\substack{i,j \\ i+j=A}} \frac{\partial \Phi[G_i, G_j, G_A]}{\partial \mu} = 0 \ .$$

Using generalized optical theorems we can show that $(G_A = |G_A| \exp(i\delta_A))$

$$\frac{\partial}{\partial \omega} \left[\operatorname{Im} \ln \left(-G_A^{-1} \right) + \operatorname{Im} \Sigma_A \operatorname{Re} G_A \right] = 2 \operatorname{Im} \left[G_A \operatorname{Im} \Sigma_A \frac{\partial}{\partial \omega} G_A^* \operatorname{Im} \Sigma_A \right] = -2 \sin^2 \delta_A \frac{\partial \delta_A}{\partial \omega} .$$

The density in the form of a generalized Beth-Uhlenbeck EoS follows

$$n(T,\mu) = \sum_{i=1}^{A} n_i(T,\mu) = \sum_{i=1}^{A} d_i \int \frac{d^3q}{(2\pi)^3} \int \frac{d\omega}{2\pi} f_i(\omega) 2\sin^2 \delta_i \frac{\partial \delta_i}{\partial \omega} .$$

³B. Vanderheyden & G. Baym, J. Stat. Phys. (1998), J.-P. Blaizot et al., PRD (2001)

Unified approach to quark-nuclear matter Example: deuterons in nuclear matter



The Φ -derivable thermodynamical potential for the nucleon-deuteron system reads

$$\Omega = -\mathrm{Tr} \{ \ln(-G_1) \} - \mathrm{Tr} \{ \Sigma_1 G_1 \} + \mathrm{Tr} \{ \ln(-G_2) \} + \mathrm{Tr} \{ \Sigma_2 G_2 \} + \Phi[G_1, G_2] ,$$

where the full propagators obey the Dyson-Schwinger equations

$$G_1^{-1}(1,z) = z - E_1(p_1) - \Sigma_1(1,z); \quad G_2^{-1}(12,1'2',z) = z - E_1(p_1) - E_2(p_2) - \Sigma_2(12,1'2',z),$$

with selfenergies and
$$\Phi$$
 functional

$$\Sigma_1(1,1') = \frac{\delta\Phi}{\delta G_1(1,1')}; \quad \Sigma_2(12,1'2',z) = \frac{\delta\Phi}{\delta G_2(12,1'2',z)}, \Phi = \bigoplus,$$

fulfilling stationarity of the thermodynamic potential $\partial\Omega/\partial G_1 = \partial\Omega/\partial G_2 = 0$. For the density we obtain the cluster virial expansion

$$n = -rac{1}{V}rac{\partial\Omega}{\partial\mu} = n_{
m qu}(\mu,T) + 2n_{
m corr}(\mu,T) \; ,$$

with the correlation density in the generalized Beth-Uhlenbeck form

$$n_{\rm corr} = \int \frac{dE}{2\pi} g(E) 2\sin^2 \delta(E) \frac{d\delta(E)}{dE}$$
.



Cluster virial expansion for quark-hadron matter

$$\Omega = \sum_{i=Q,M,D,B} c_i \left[\operatorname{Tr} \ln \left(-G_i^{-1} \right) + \operatorname{Tr} \left(\Sigma_i \ G_i \right) \right] + \Phi \left[G_Q, G_M, G_D, G_B \right] ,$$



When Φ functional for the system is given by 2-loop diagrams holds

$$n = -\frac{\partial\Omega}{\partial\mu} = \sum_{a} a n_{a}(T,\mu)$$
$$= \sum_{a} a d_{a} \int \frac{d\omega}{\pi} \int \frac{d^{3}q}{(2\pi)^{3}} \left\{ f_{\phi}^{(a),+} - \left[f_{\phi}^{(a),-} \right]^{*} \right\} 2 \sin^{2} \delta_{a}(\omega,q) \frac{\partial \delta_{a}(\omega,q)}{\partial \omega} ,$$

Analogous for the entropy density $s = -\partial \Omega / \partial T$.

20

Unified approach to quark-nuclear matter Cluster virial expansion for quark-hadron matter



The cluster decomposition of the thermodynamic potential is given as

$$\Omega_{ ext{total}}(T,\mu,\phi,ar{\phi}) = \Omega_{ ext{PNJL}}(T,\mu,\phi,ar{\phi}) + \Omega_{ ext{pert}}(T,\mu,\phi,ar{\phi}) + \Omega_{ ext{MHRG}}(T,\mu,\phi,ar{\phi}),$$

where the first two terms describe the quark and gluon degrees of freedom via the mean-field thermodynamic potential for quark matter in a gluon background field \mathcal{U}

$$\Omega_{PNJL}(T,\mu,\phi,\bar{\phi}) = \Omega_Q(T,\mu,\phi,\bar{\phi}) + \mathcal{U}(T,\phi,\bar{\phi})$$

with a perturbative correction $\Omega_{pert}(T, \mu, \phi, \overline{\phi})$.

00

The Mott-Hadron-Resonance-Gas (MHRG) part for the multi-quark clusters is

$$\Omega_{MHRG}(\mathcal{T},\mu,\phi,\bar{\phi}) = \sum_{i=M,B,\dots} \Omega_i(\mathcal{T},\mu,\phi,\bar{\phi}),$$

where the multi-quark states are described by the GBU formula:

$$m = -\frac{\partial\Omega}{\partial\mu} = \sum_{a} a n_{a}(T,\mu)$$
$$= \sum_{a} a d_{a} \int \frac{d\omega}{\pi} \int \frac{d^{3}q}{(2\pi)^{3}} \left\{ f_{\phi}^{(a),+} - \left[f_{\phi}^{(a),-} \right]^{*} \right\} 2 \sin^{2} \delta_{a}(\omega,q) \frac{\partial \delta_{a}(\omega,q)}{\partial \omega} ,$$

where d_i is the degeneracy factor, a is the number of valence quarks in the cluster an $f_{\phi}^{(a),+}$, $\left[f_{\phi}^{(a),-}\right]^*$ are the Polyakov-loop modified distribution functions. Analogous for the entropy density $s = -\partial \Omega / \partial T$.

Unified approach to quark-nuclear matter Polyakov-loop modified distribution functions

For multiquark clusters with net number a of valence quarks holds

$$f_{\phi}^{(a),\pm} \stackrel{(a \text{ even})}{=} \frac{(\phi - 2\bar{\phi}y_{a}^{\pm})y_{a}^{\pm} + y_{a}^{\pm 3}}{1 - 3(\phi - \bar{\phi}y_{a}^{\pm})y_{a}^{\pm} - y_{a}^{\pm 3}},$$

$$f_{\phi}^{(a),\pm} \stackrel{(a \text{ odd})}{=} \frac{(\bar{\phi} + 2\phi y_{a}^{\pm})y_{a}^{\pm} + y_{a}^{\pm 3}}{1 + 3(\bar{\phi} + \phi y_{a}^{\pm})y_{a}^{\pm} + y_{a}^{\pm 3}},$$

where $y_a^{\pm} = e^{-(E_p \mp a\mu)/T}$ and $E_p = \sqrt{\vec{p}^2 + M^2}$. It is instructive to consider the two limits $\phi = \bar{\phi} = 1$ (deconfinement)

$$f_{\phi=1}^{(a=0,2,4,\ldots),\pm} = \frac{y_a^{\pm}}{1-y_a^{\pm}}, \quad f_{\phi=1}^{(a=1,3,5,\ldots),\pm} = \frac{y_a^{\pm}}{1+y_a^{\pm}},$$

and $\phi = \bar{\phi} = 0$ (confinement),

$$f_{\phi=0}^{(a=0,2,4,\ldots),\pm} = \frac{y_a^{\pm^3}}{1-y_a^{\pm^3}}, \ f_{\phi=0}^{(a=1,3,5,\ldots),\pm} = \frac{y_a^{\pm^3}}{1+y_a^{\pm^3}}.$$



Unified approach to quark-hadron matter



Inputs: mass spectrum & phase shifts (models)



Unified approach to quark-hadron matter Inputs: mass spectrum (Particle Data Tables)

Mesons

PDG	d_i	$M_{\rm PDG}$	M_i	$M_{\rm th}^{<}$	$M_{\rm th}^{>}$
mesons		[MeV]	[MeV]	[MeV]	[MeV]
π^+/π^0	3	140	140	1254	11.2
K^+/K^0	4	494	494	1397	129.6
η	1	548	878	1349	90.1
ρ^+/ρ^0	9	775	783	1254	11.2
ω	9	783	783	1254	11.2
K^{*+}/K^{*0}	12	895	806*)	2651	140.8
η'	1	960	878	1349	90.1
a_0	3	980	$1095^{*)}$	2508	22.4
f_0	1	980	1095 ^{*)}	2508	22.4
ϕ	3	1020	1069	1540	248
$\pi_2(1880)$	15	1895	1095 ^{*)}	2508	22.4
$f_2(1950)$	5	1944	$1095^{*)}$	2508	22.4
$a_4(2040)$	27	1996	$1095^{*)}$	2508	22.4
$f_2(2010)$	5	2011	$1095^{*)}$	2508	22.4
$f_4(2050)$	9	2018	1095* ⁾	2508	22.4
$K_4^*(2045)$	36	2045	1238* ⁾	2651	140.8
$\phi(2170)$	3	2175	$1381^{*)}$	2794	259.2
$f_2(2300)$	5	2297	1095* ⁾	2508	22.4
$f_2(2340)$	5	2339	$1095^{*)}$	2508	22.4

Baryons

PDG	d_i	$M_{\rm PDG}$	M_i	$M_{\mathrm{th},i}^{<}$	$M_{\text{th},i}^{>}$
baryons		[MeV]	[MeV]	[MeV]	[MeV]
n/p	4	939	939	1881	16.8
Λ	2	1116	1082	2024	135.2
Σ	6	1193	1082	2024	135.2
Δ	16	1232	$1251^{**)}$	3135	28
Ξ^0	2	1315	1225	2167	253.6
Ξ^{-}	2	1322	1225	2167	253.6
$\Sigma(1385)$	6	1385	1394 ^{**)}	3278	146.4
$\Lambda(1405)$	2	1405	$1394^{**)}$	3278	146.4
N(1440)	4	1440	$1251^{**)}$	3135	28
N(2195)	36	2220	1251 ^{**)}	3135	28
$\Sigma(2250)$	6	2250	$1394^{**)}$	3278	146.4
$\Omega^{-}(2250)$	2	2252	1680**)	3564	383.2
N(2250)	20	2275	$1251^{**)}$	3135	28
A(2350)	10	2350	$1394^{**)}$	3278	146.4
$\Delta(2420)$	48	2420	$1251^{**)}$	3135	28
N(2600)	24	2600	$1251^{**)}$	3135	28

... and colored clusters (model) !

Unified approach to quark-hadron matter Inputs for the phase shifts (models)





Step-up (SU) model \rightarrow Hadron Resonance Gas

Step-up-step-down model Step-up-continuum model \rightarrow Mott Hadron Resonance Gas (MHRG)

Unified approach to quark-hadron matter Results for Mott-Hadron Resonance Gas (MHRG)



Unified approach to quark-hadron matter Entropy for MHRG: role of the sin²-term

Unified approach to quark-hadron matter Results for the entropy density

Unified approach to quark-hadron matter

Results for pressure=thermodynamic potential

5

 $\mu_{\rm B}/T = 0$ Borsányi et al., PRL (2021) Borsányi et al., PLB (2014) The pressure is obtained by 1 integrating the entropy density over temperature 3 s(T) = dp(T) / dTp/T⁴ \rightarrow p(T) = $\int dT' s(T')$ 2 Excellent agreement with Lattice QCD thermodynamics ! 1 0 100 200 300 400 500 600 0 T [MeV]

Unified approach to quark-hadron matter

Results for the entropy density & composition

Abrupt hadronisation (change in the composition) at the chiral crossover transition with Tc=156 MeV \rightarrow important for understanding chemical freeze-out in ultrarelativistic heavy-ion collisions.

Unified approach to quark-hadron matter Chiral condensate & baryon susceptibilities

1,25 Borsanyi et al., JHEP (2010) Bazavov et al., PRD (2017) $V_{excl} = 0$ $V_{\text{excl}} = V_0 \text{th}((T_V - T)/\delta_V) + V_0$ $V_0 = 2.9 \text{ fm}^{-3}$ BHRG 0.75 $T_v = 163 \text{ MeV}$ $\delta_{\rm V} = 13 \, {\rm MeV}$ 0.5 0,75 X^B/X² 0,5 0 hadrons ored clusters 0,25 -0.25free quark gas -0.50└ 100 150 200 250 100 150 200 250 T [MeV] T [MeV] Smooth crossover for the chiral condensate, Ratio of baryon susceptibilities shows: Despite abrupt change in the composition!

D.B., O. Ivanytskyi & G. Röpke, in preparation (2024)

Unified approach to quark-hadron matter

Applications to cool hybrid neutron star matter

N. Bastian and D.B., EPJA 57 (2021) 35

Chemical freeze-out in heavy-ion collisions

"inverse" of Mott dissociation by χSR or Pauli blocking

High temperature, low baryon density:

- \rightarrow χSB entails confinement of hadrons and clusters
- → chem. Freeze-out coincides with LQCD crossover

Intermediate temperature (30-100 MeV), high density:

- \rightarrow Pauli blocking destroys nuclear clusters
- \rightarrow Mott dissociation line for clusters (a) = freeze-out line

Chemical freeze-out in heavy-ion collisions

"inverse" of Mott dissociation by χSR or Pauli blocking

High temperature, low baryon density:

- \rightarrow χSB entails confinement of hadrons and clusters
- → chem. Freeze-out coincides with LQCD crossover

Intermediate temperature (30-100 MeV), high density:

- \rightarrow Pauli blocking destroys nuclear clusters
- \rightarrow Mott dissociation line for clusters (a) = freeze-out line

Relativistic density functional for quark matterWhat is new?O. Ivanytskyi & D.B., Phys. Rev. D 105 (2022) 114042

Interaction $\mathcal{U} = D_0 \left[(1+\alpha) \langle \overline{q}q \rangle_0^2 - (\overline{q}q)^2 - (\overline{q}i\vec{\tau}\gamma_5 q)^2 \right]^{\varkappa}$

Parameters

 D_0 - dimensionfull coupling, controls interaction strength α - dimensionless constant, controls vacuum quark mass

 $\langle \overline{q}q \rangle_0$ - χ -condensate in vacuum (introduced for the sake of convenience)

$$\begin{split} \varkappa &= 1/3 \\ & \Downarrow \\ \text{motivated by String Flip model} \\ & \mathcal{U}_{SFM} \propto \langle q^+ q \rangle^{2/3} \\ \Sigma_{SFM} &= \frac{\partial \mathcal{U}_{SFM}}{\partial \langle q^+ q \rangle} \propto \langle q^+ q \rangle^{-1/3} \propto \text{separation} \end{split}$$

Dimensionality

$$\begin{bmatrix} \mathcal{U} \end{bmatrix} = energy^4 \\ [\overline{q}q] = energy^3$$
 \Rightarrow $[D_0]_{\varkappa=1/3} = energy^2 = [string tension]$

self energy = string tension \times separation \Rightarrow confinement

 $\varkappa = 1$

Nambu-Jona-Lasinio model

Relativistic density functional for quark matter Expansion around mean fields

$$\mathcal{U} = \underbrace{\mathcal{U}_{MF}}_{0^{\text{th}} \text{ order}} + \underbrace{\left(\overline{q}q - \langle \overline{q}q \rangle\right) \Sigma_{S}}_{1^{\text{st}} \text{ order}} - \underbrace{G_{S}\left(\overline{q}q - \langle \overline{q}q \rangle\right)^{2} - G_{PS}\left(\overline{q}i\vec{\tau}\gamma_{5}q\right)^{2}}_{2^{\text{nd}} \text{ order}} + \dots$$
Mean-field scalar self-energy
$$\Sigma_{S} = \frac{\partial \mathcal{U}_{MF}}{\partial \langle \overline{q}q \rangle}$$
Effective medium dependent couplings
$$G_{S} = -\frac{1}{2} \frac{\partial^{2} \mathcal{U}_{MF}}{\partial \langle \overline{q}q \rangle^{2}}, \quad G_{PS} = -\frac{1}{6} \frac{\partial^{2} \mathcal{U}_{MF}}{\partial \langle \overline{q}i\vec{\tau}\gamma_{5}q \rangle^{2}}$$

Relativistic density functional for quark matter EOS and Mass-radius diagram for hybrid neutron stars

Observational constraints prefer early onset of deconfinement

Thanks to my collaborators:

T. Fischer, G. Röpke, A. Bauswein, <u>O. Ivanytskyi</u>, O. Vitiuk, N. Bastian, M. Cierniak, U. Shukla, S. Liebing, K. Maslov,

- A. Ayriyan,
- H. Grigorian,
- D.N. Voskresensky,
- M. Kaltenborn,
- G. Grunfeld,
- D. Alvarez-Castillo,
- B. Dönigus, ...

Wroclaw Group ...

University of Wroclaw, Institute of Theoretical Physics

Division: Theory of Elementary Particles - Collaborations

Polish-German WE-Heraeus Seminar & Max Born Symposium:

03.12. 06.12. 2023 Many-particle systems under extreme conditions

https://events.hifis.net/event/1076

Understanding complex systems with data

CASUS pushes the frontier of data-driven complex systems science

CASUS - The center for data-driven complex systems science

70 people, 20 countries, the brightest minds

Prof. Dr. Justin Calabrese Ecological Data Science

Dr. Attila Cangi Matter under Extreme Conditions

Sandia National Laboratories

CASUS

Dr. Artur Yakimovich Machine Learning for Infection and Disease

Dr. Ricardo Martinez-Garcia Dynamics of Complex Living Systems

PRINCETON UNIVERSITY

Dr. Michael Hecht Mathematical Foundations of Complex System Science

OF MOLECULAR CELL BIOLOGY AND GENETICS

CRTD

Dr. Weronika Schlechte-Welnicz SCULTETUS Center

Excellence recognized

ERC Starting Grant Dr. Tobias Dornheim

CASUS/UWr joint Professor Prof. Dr. David Blaschke **Professor University** of Arizona (Faculty) **Dr. Jesse Alston**

COLLEGE OF AGRICULTURE & LIFE SCIENCES Natural Resources & the Environment

Fullbright scholarship Dr. Debanjan Konar

UÌ

G

SPINISHS WITH

Federal Ministry of Education and Research

STAATSMINISTERIUM FÜR WISSENSCHAFT UND KUNST

New: The German Centre for Astrophysics (DZA)

Research Technology Digitization

"Science Creating Prospects for the Region!"

Scientific Commission: 13. July 2022 Structural and Transfer-Commission: 30. August 2022 Final decision (Approval): 29. September 2022

DZA

David Blaschke - Density functional approach to quark-hadron matter

Why in Saxony? Lusatia is a unique region for Astrophysics, Technology and Digitization

JWST results – primordial black holes !

Talk at University of Wroclaw by Günther Hasinger, Founding director of the German Centre for Astrophysics In Görlitz:

Key role plays the QCD hadronization transition !

Different peaks correspond to different particles created at the early universe phase transitions and the corresponding reduction in the sound velocity.

BH mass corresponds to the horizon size at each time.

Only requirement is enough fluctuation power in a volume fraction of 10⁻⁹ of the early Universe.

Carr, Clesse, García-Bellido 2019

David Blaschke - Density functional approach to quark-hadron matter

JWST results – primordial black holes !

LIGO-Virgo-KAGRA | Aaron Geller | Northwestern

Deutsches Zentrum für Astrophysik

Courtesv: Günther Hasinger (Karpacz 2024)

JWST results – primordial black holes !

1.25

1.00

0.75

0.50

0.25

Courtesv: Günther Hasinger (Karpacz 2024)

New constraints on PBH mass function

Original MACHO & OGLE microlensing constraints (Wyrzykowski, L., et al. 2011, solid). Reanalysis of the MACHO constraints on PBH in the light of the new Gaia MW rotation curve (Garcia-Bellido, J. & Hawkins, M., 2024, dashed) New 20-yr OGLE microlensing constraints (Mroz, P. et al., arXiv 2403.02386). Search for Subsolar-Mass Binaries in the First Half of Advanced LIGO's

and Advanced Virgo's Third Observing Run.

Just about fits!

Deutsches Zentrum für Astrophysik

Courtesv: Günther Hasinger (Karpacz 2024)