# Electromagnetic effects on charged particles

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- 1) Prologue;
- 2) Do never agree with your boss ;
- 3) EM effects on charged particles ;
- 4) Space-time evolution of the system ;
- 5) EM effects in small systems ;
- 6) UPC's ?
- 7) <u>No</u> epilogue.

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by I. Sputowska

Collision energy in the c.m.s. (centerof-mass system), per one pair of colliding nucleons.

## 1) Prologue



• Charged spectators generate electromagnetic fields.

- These modify charged pion spectra in the final state.
- We use this effect as a new source of information on the space-time evolution of the system.



Please note: advantages of fixed-target w.r.t. collider experiments:

- typically better coverage of kinematically available phase-space (p<sub>x</sub>,p<sub>y</sub>,p<sub>z</sub>) : "forward" hemisphere of the collision.
- Full coverage of low transverse momentum starting from p<sub>T</sub>=0 ;

 $(p_T = \sqrt{p_x^2 + p_y^2})$ er to develop

- Easier to develop (add new subdetectors);
- Cheaper (?)





A historical (?) question:

- what is a heavy-ion collision?

A simple consequence of nucleonnucleon processes? "New" physics? Both?

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 $p_L$  $X_{E}$ beam (c.m.s.)

unless present Ы this room

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## $\pi^+/\pi^-$ ratios in Pb+Pb @ $\sqrt{s}_{NN}$ = 17.3 GeV

- simple superposition of proton and neutron collisions?



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 $\pi^+/\pi^-$ 

A.Trzcińska et al., PRL87, 2001 R.Schmidt et al., PRC67, 2003 S.Mizutori et al., PRC61, 2000





## Hypothesis no. 1: the "neutron halo" ?

P.Pawłowski, A.Szczurek, PRC70, 2004



- Analysis of collision geometry: b = 10.9 ± 0.5 fm
- Not possible to obtain 75% n, 25 % p





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XF

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 $X_F$ 





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WA98: NPA 663 (2000) 725

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# 4) **Space-time evolution of** forward pion production

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#### Plan:

- Formulation of a simple model;
- Validation with exp. data on rapidity distributions;
- Application to EM effects. PRC102 (2020) 014901

PRC 95 (2017) 024908 PRC 99 (2019) 024908 (\*)

(\*) Yes, both papers have page no. 024908

l<sub>E</sub> [fm] σ 3 2 1 0 1.5 0.5  $\cap$ beam

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π

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 The pion rapidity distribution from one fire streak in Pb+Pb collisions is similar to the pion rapidity distribution in **p+p** reactions :



**Ap/up** 

20 È

15 F

(a)

NA49,

<sup>60</sup> 50 40

30

(b)

Model, b = 9.72 f

NA49, C4

π, Pb+Pb, √s\_<sub>NN</sub>= 17.3 GeV

(c)

NA49, C2

vp/up

40 Ē

NA 49, C3

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**Ap/up** 

20

15 F

(a)

NA49,

30

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Model, b = 9.72 f

NA49, C4

π̄, Pb+Pb, √s̄<sub>NN</sub>= 17.3 GeV

(c)

NA49, C2

NA 49, C3

# Application to EM effects

Longitudinal evolution of the system → from our model ;







- <u>Fragmentation</u> (expansion) of the spectator charge  $\rightarrow$  included ;
- Isospin (p/n) effects between  $\pi^+$  and  $\pi^- \rightarrow$  included  $\rightarrow$  PRC 99 (2019) 024908 ;
- <u>Azimuthal anisotropies</u> (flow), <u>vorticity</u>, <u>transverse expansion</u>  $\rightarrow$  included optionally ;
- The pion creation time  $\tau$  (taken in the fire streak c.m.s.)  $\rightarrow$  taken as free parameter.



0.0

0.2

X<sub>E</sub>

0.3

0.4

0.5

0.1

0.0

-0.1



**First quantitative description** of the electromagnetic distortion of  $\pi$ +/ $\pi$ - ratios in Pb+Pb collisions at 158 GeV/nucleon beam energy ( $\sqrt{s_{NN}}$ =17.3 GeV) [PRC102 (2020) 014901].

- → Reasonable agreement with experimental data for  $x_F \ge 0.1$ ;
- ➔ Inclusion of spectator expansion improves the description of exp. data ;
- → Short pion creation times (  $0.5 < \tau < 2$  fm/c, to be compared with ~ 5.5 fm/c at y=0).



- EM effects induced by spectators bring new information on the space-time evolution of the process of (fast) pion production;
- We obtained this information and used it ;
- Results look reasonable.
- → Small systems ;
- → UPC (gamma-gamma).

 $\pi^{*}$ π π 4 d<sub>E</sub> I 3 2 1 0 1.5 0.5  $\mathbf{O}$ oeam

[fm]

#### NA61/SHINE experiment

#### SHINE = SPS Heavy Ion and Neutrino Experiment





 $\sqrt{s_{NN}}$  = 5.1–17.3(27.4) GeV



#### Strong interactions

- study the onset of deconfinement
- search for the critical point

Adapted from Antoni Marcinek, QM22

5) Small systems

#### Cosmic rays and neutrinos





 $\sqrt{s_{_{NN}}}$ = 8.8 GeV



NA61/SHINE collaboration, MESON2021, 18 May 2021

See also e.g.: A. Marcinek for NA61/SHINE, QM22

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## Modelling EM effects in new Ar+Sc data



A. Rybicki, A. Szczurek, Phys. Rev. C 75, 054903 (2007)



 Non-expanding spectator system cannot describe data (contrary to Pb+Pb, see A. Rybicki et al., APPB 46,737 (2015))

need significant expansion velocity  $\beta$  of the charge cloud

 Optimal description: charge cloud moves slower than spectator system → presence of participant charge?

Antoni Marcinek (IFJ PAN)

APPB 50 (2019), 1127

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needed.

Xe+La?



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colliding nuclei

# 6) Comment: how about gamma-gamma processes in ultra-peripheral collisions?

- Think about **Pb+Pb**  $(\gamma \ \gamma) \rightarrow e^+e^-$ ;
- What happens to leptons once created ?
- Subject of long discussions (here in Kraków);
- Conceptual difficulties ;
- Never measured.





 $R_1$ 

h



# 6) No epilogue

The electromagnetic (EM) fields resulting from the presence of charged spectators induce distortions in

spectra of charged particles, and result in charge splitting of directed flow.

These spectator-induced EM effects can be used to study the space-time evolution of particle production.

Such studies have shown that in high energy heavy-ion collisions, faster pions are produced closer to the spectator system. They also provided an independent estimate for the time of pion creation, at y=0.

A first quantitative description of the EM distortion of charge ratios ( $\pi$ +/ $\pi$ -) of fast pions produced in Pb+Pb collisions at  $\sqrt{s_{NN}}$ =17.3 GeV has been obtained. This gives an indication of significantly shorter pion production time scales (shorter proper times  $\tau$ ) w.r.t. what was obtained at central rapidity.

**First ever measurements** of these effects in **small systems** at the CERN SPS are now available from NA61/SHINE.

Thank you !

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## **Extra slides**

√s<sub>NN</sub>= 17.3 GeV

 p+A collisions with our model from Sec. 4.



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√s<sub>мм</sub>= 17.3 GeV

 p+A collisions with our model from Sec. 4.



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