# Dark matter complementarity in Snowmass

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### What is Snowmass?



Particle physics community planning exercise organized by the Division of Particles and Fields (DPF) of the American Physical Society

#### Snowmass goals:

- work on/collect new scientific studies, mostly concerning on future directions for the field
- engage the community and junior scientists
- to prepare a collective vision for the next decade of US particle physics

#### Organized in 10 Frontiers

- Accelerator, Cosmic, Community Engagement, Computing, Energy, Instrumentation, Neutrino, Rare processes and Precision Measurements, Theory, Underground + Snowmass Young (for early career input)
- ▶ Each Frontier has 5-10 Topical Groups (eg. dark matter!)
- The vision is taken by the Particle Physics Project Prioritization Panel (P5) as input to a 10-year strategic plan
  - P5 science drivers from Snowmass 2013: Higgs boson as tool for discovery, pursue physics associated to neutrino mass, identify the new physics of dark matter, understand dark energy and inflation, explore the unknown

The P5 website: https://www.usparticlephysics.org/p5/ The P5 process, slides by chair H. Murayama

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- ▶ Empirical evidence of DM from astrophysical observations at different scales
  - interacts gravitationally, long lived and neutral
  - no information about its nature (what DM could be?)
  - only measured quantitative property is its mass abundance
    - very large set of possible DM masses can account for observed final relic density



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### Snowmass: what is the road to discover DM?

arXiv:2210.01770



#### Over the next decade, need of unified strategy to maximize discovery potential

- many compelling theoretical targets from simple/minimal models will be accessible via planned and proposed experiments, colliders, and observatories across experimental frontiers
- understanding and harnessing complementarity between techniques and technologies is essential
- if DM discovered, complementarity needed to triangulate fundamental nature of DM
- if DM is not found, we will still gain important information on its properties

#### What is the strategy ahead to allow DM discovery?

- <u>build a portfolio of experiments of different scales</u>: experiments at all scales needed to cover very broad range of theoretically motivated parameter space
- <u>leverage us expertise in international projects</u>: coordination and cooperation across borders is critical. Build a strong US-based program and pursue opportunities for US expertise to play key role in international projects
- <u>Provide support to further strengthen the theory program</u>: essential to make connections between experimental frontiers and guide new approaches to understand and detect DM
- <u>Support inter-disciplinary collaborations and increase in the research budget: essential to enable new ideas,</u> new technologies, and new analyses

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# Why complementarity?

Dífferent DM search approaches and complementary necessary to investigate DM across Frontiers:



Experiments can simultaneously discover DM by detecting its relic components or producing it

- relic DM probes give insight into DM halo properties and provide connection of new particles discovery with cosmological DM, DM production probes give insight on DM interactions and on dark-sector beyond SM
- Different experimental approaches have unique parameter-space coverage and DM properties sensitivity
  - for eg. DM detection experiments are sensitive to non-relativistic DM signals and production probes explore the physics of DM in relativistic regime
  - terrestrial experiments are sensitive to rare DM interaction processes; cosmological and astrophysical probes have unique sensitivity to properties like lifetime, DM self-interaction, etc ...
- ▶ DM experiments can be co-located and/or profit from the same or similar technological infrastructure
  - wider exploration of DM with a more efficient use of shared resources

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# Inputs and opinions from people outside of Snowmass?

In the following wishlist of inputs from the community for different frontier

### Goal:

- Inform European community of efforts on dark matter complementarity within Snowmass
- are there more recommendations that the European community would like to bring forward with respect to the ones already introduced in the whitepaper?

Cosmic: direct and Indirect searches

- ▶ Focus: particles DM with mass ev-1019 Gev
  - <u>direct detection</u>: delving deep into heavy DM, with spin-independent reaching neutrino floor and increasing sensitivity in scenarios where the neutrino floor remains distant, searching wide with new techniques to probe DM in MeV and lower ranges
  - <u>indirect detection</u>: delving deep in TeV range with model-agnostic probes of minimal thermal relic scenario, searching wide in MeV range with new probe of low-energy cosmic-ray anti-deuterons
- ▶ Roadmap:
  - Diverse and continuous direct and indirect detection experiments portfolio at multiple scale
  - Moderate- and large-scale experiments allow to delve deep in GeV-TeV range, small-scale experiments provides ability to test expanded models range
  - Support theory, simulations, calibration, bkg modeling and complementary astrophysical measurements
  - Rgp towards improved detector technologies are fundamental and crucial



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  - provide information about our astrophysical environment
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  - provide only known way to directly study DM fundamental properties through gravity
  - sensitive to DM mass, lifetime, self-interaction, and other dark sector particles in extreme large scales and environments inaccessible to terrestrial experiments (wide range of parameter space)
  - ability to link results of terrestrial DM experiments to cosmological measurements
- ▶ Roadmap:
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  - Construction of future cosmology experiments critical for DM understanding. HEP community should invest strategically an support design, construction, operation of these facilities to maximize their sensitivity to DM
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Rare & precision: Dark Sectors at High Intensities

- Focus: particle DM with mass MeV-GeV
  - search wide, provide unique and unprecedented access to low-mass DM
  - powerful probe of thermal DM models
  - also probe of light SM-neutral dark-sector particles decaying into SM particles (long lifetimes) \_

#### Roadmap:

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- Broaden and invest in completion of "Basic Research Needs Dark Matter New Initiatives" (DMNI) program, small-scale experiments essential to explore sub-Gev thermal DM and dark-sector not accessible by large multipurpose detectors. CCM200 and LDMX selected projects, but need to identify dark-sector future ideas
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- ▶ Roadmap:
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  - Intermediate future: electron-positron collider, further explore if DM is produced via a SM or BSM mediator especially for models favoring lepton couplings
  - longer term: multi-Tev discovery machine, allow higher energy scales for dark-sector physics, will reach thermal milestone for thermal DM from 3 GeV complementing lower-mass Frontier experiments



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Neutrino: DM in neutrino

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  - indirectly probe or constrain DM by anomalous spectra of neutrino fluxes
  - *improved understanding of neutrino fluxes*, properties, cross sections can assist in understanding the bkg in DM searches with neutrino experiments and in other frontiers
- ▶ Roadmap:
  - Realization of DUNE in its full scope within the next decade. Will expand opportunities in DM and dark-sector searches through high intensity, high power proton beams and near/large-scale detectors
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  - define connections between different experimental programs and Frontiers
  - motivate specific experimental directions
- Instrumentation: new technologies often opening up new regions of parameter space for exploration
  - promote R&D into instrumentation, development of technical workforce, tools to share common knowledge
- Computing: essential for raw data collection, storage, data processing and analysis, results interpretation ...
  - build stronger partnerships with national supercomputing facilities, enhance industry collaborations on machine learning technique, support scalable software infrastructure tools across communities
  - essential to have efficient, well-maintained, and validated simulation infrastructure
  - promote widely-adopted data and software standards, in support of global analyses of experimental results
- Underground facilities: needed to host experiments (eg. to evade cosmic ray backgrounds)
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- Accelerator frontier: foundation of intensity-frontier and energy-frontier searches for DM and dark sectors
  - efficiently utilize existing and upcoming facilities to explore possible rare process experiments
  - explore technically and document accelerator possibilities for energy frontier to operate in 2040
- Community Engagement and Workforce Development:
  - material support to enhance engagement efforts at all levels of society, also as recruiting tool for HEP
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- ▶ Very large set of viable theoretical scenarios that can provide DM candidates
  - possibilities span enormous ranges in DM mass and interaction strength
  - <u>complementarity between DM searches pivotal</u> to enable discovery of DM fundamental nature and to allow triangulation of its properties (both in case of discovery or null results)



- Case study: WIMP DM lightest particle of multiplet interacting with SM through weak force (EWK SUSY)
  - DM Dirac (Higgsino) or Majorana(Wino) particle, m<sub>DM</sub> only free parameter fixed by relic density (TeV scale, but lighter for non-thermal production), cross-sections not fully probed by current experimental effort



direct detection: DM-nucleon interactions expected to be very rare, requiring large exposure for detection.

*Indirect detection:* signal from DM annihilation into photons or gauge bosons accessible by gamma-ray telescopes and other indirect detection experiments

colliders: can probe DM invisible signature, semi-visible decays of long-lived intermediate states (TeV scale), loop effects due to new particles (lower energies)

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  - eg. benchmark vector portal, dark vector boson (dark photon) mediator that mixes with SM photon



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- in portal models, both low and high mass particles are needed for theoretical consistency. Only a combination of Cosmic, Energy and RaregPrecision Frontier experiment can discover DM and associated particles
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Case study: interactions mediated by heavy neutral lepton (HNL) that mix with SM neutrinos

Complementarity: coverage of viable phase space

cosmic probes: constraints from suppression of DM structure for producing keV-mass sterile neutrinos with a "warm" initial momentum distribution

Indirect detection: search for X-ray lines originating from the decay of keV mass sterile neutrinos

neutrino experiments: different dependencies on the model behavior of sterile neutrinos in the early Universe

colliders: low mixing values and correct DM relic require other production mechanisms which involve heavy BSM particles that can be probed by Energy Frontier



#### <u>Case study</u>: wavelike DM

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- Complementarity: in case of discovery from direct detection experiments
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DM nature triangulate DM nature fundamental open (employing results from question in physics multiple approaches) Viable DM candidates yenergy frontier: high-energy collide Macroscopic DM Compact objects Interaction strength Asymmetric DM  $10^{-4}$ current: LHC CF, RF, E near term: HL-LHC WIMP hermal DM CF, E SIMP/Elder  $10^{-9}$ Scalar-vector light DM Dark sectors thermal milestone  $y = \varepsilon^2 \alpha \left(\frac{m_{\rm DM}}{m_{\rm med}}\right)^2$ Secluded dark sectors self-interactions, dark radiation, light relics.  $10^{-14}$ μeV feV eV MeV TeV  $10 \ M_{\odot}$ zeV 1 GeV 10 GeV100 GeV 100 MeVDark matter mass dark matter mass need to delve deep and search enormous theoretically wide across frontiers motivated space

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\* Complementarity across frontiers is at the hearth of the journey to DM discovery

- feedback on the roadmaps discussed?
- other complementarity cases that should be highlighted beyond those in Snowmass document was summarized in this talk?
- concrete suggestions from your community / experiment on how to make these roadmaps happen?

your

inpu





### How can we study DM interactions?

- ▶ Assuming DM-SM interactions enables different searches:
  - índírect detection, products from DM annihilation
  - dírect detection, nuclear recoils from DM-nuclei scattering
  - colliders, DM production
  - Complementarity essential: eg. info about lifetime in case of DM discovery at colliders (~10<sup>-7</sup>s), particle properties compared with cosmological constraints
- ▶ Example of complementarity for DM and dark-sector production

Mev-Gev:

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GeV-TeV:

- extensions of the SM at the GeV-TeV scale

 $\rightarrow$  high-energies

- dark sector mediators could be heavy, light, and/or long-lived
- \* rich connections and complementarity between type of experiments





## How can we study DM interactions?

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energy scale or mass

Important to identify common benchmarks to better exploit complementarity!

### Snowmass: what is the road to discover DM?

#### DM and visible matter have similar energy density - compelling evidence of non-gravitational interactions but ... many viable DM candidates and properties

- no single experimental technique can probe all possibilities
- over last decade, broaden pursuable DM candidates through new experiments/searches thanks to advances in technology, analysis techniques and theoretical models
- ▶ Over the next decade, need of unified strategy to maximize discovery potential
  - many compelling theoretical targets from simple/minimal models will be accessible via planned and proposed experiments, colliders, and observatories across experimental frontiers
  - understanding and harnessing complementarity between techniques and technologies is essential
  - if DM discovered, complementarity needed to triangulate fundamental nature of DM
  - if DM is not found, we will still gain important information on its properties

#### ▶ What is the strategy ahead to allow DM discovery?

- <u>build a portfolio of experiments of different scales</u>: experiments at all scales needed to cover very broad range of theoretically motivated parameter space
- <u>leverage US expertise in international projects</u>: coordination and cooperation across borders is critical. Build a strong US-based program and pursue opportunities for US expertise to play key role in international projects
- <u>Provide support to further strengthen the theory program</u>: essential to make connections between experimental frontiers and guide new approaches to understand and detect DM
- <u>Support inter-disciplinary collaborations and increase in the research budget:</u> essential to enable new ideas, new technologies, and new analyses

DM nature triangulate DM nature fundamental open (employing results from question in physics multiple approaches) Viable DM candidates yenergy frontier: high-energy col Macroscopic DM Compact objects Interaction strength Asymmetric DM  $10^{-4}$ current: I HC CF, RF, E ear term: HL-LHC WIMP nermal DM SIMP/Elder CF,  $10^{-9}$ Scalar-vector light DM Dark sectors thermal milestone  $y = \varepsilon^2 \alpha \left( \frac{m_{\rm DM}}{m_{\rm med}} \right)$ Secluded dark sectors self-interactions, dark radiation, light relics.  $10^{-14}$ μeV feV eV  $10 \ M_{\odot}$ zeV MeV TeV 1 GeV 10 GeV100 MeV. 100 GeV Dark matter mass dark matter mass need to delve deep and search enormous theoretically wide across frontiers motivated space

- incorporate multiple complementary approaches
- increase dedicated support for research, including funding for cross-collaborations
- funding to support a strong and vibrant theory program which motivate (new) DM searches, interpretation of results