

# Exotic Di-Photon Resonances at the ATLAS Experiment

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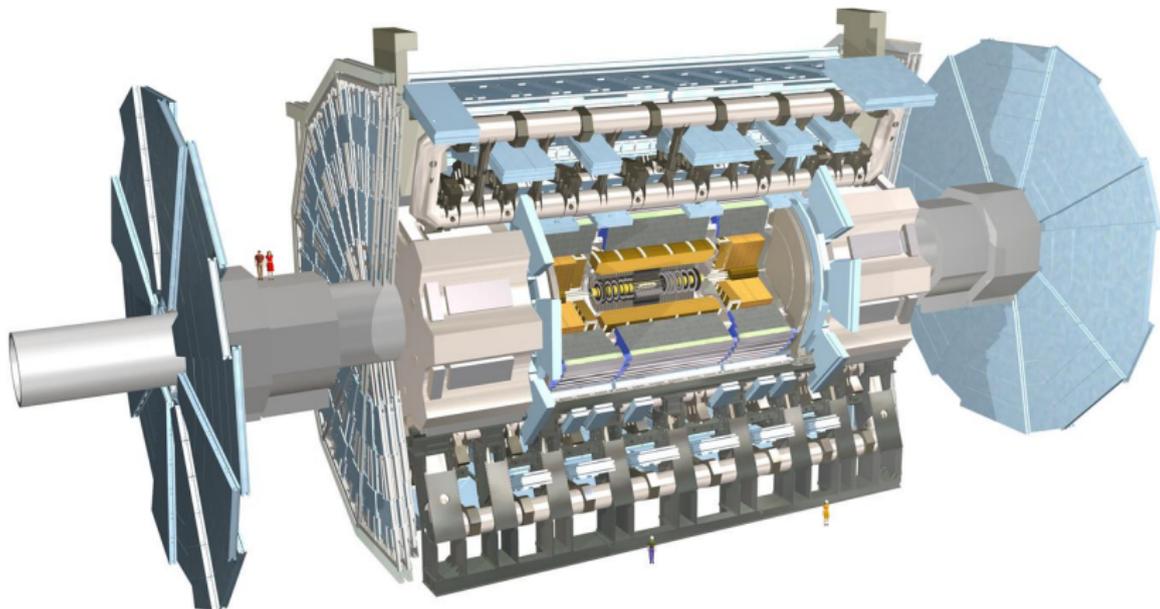
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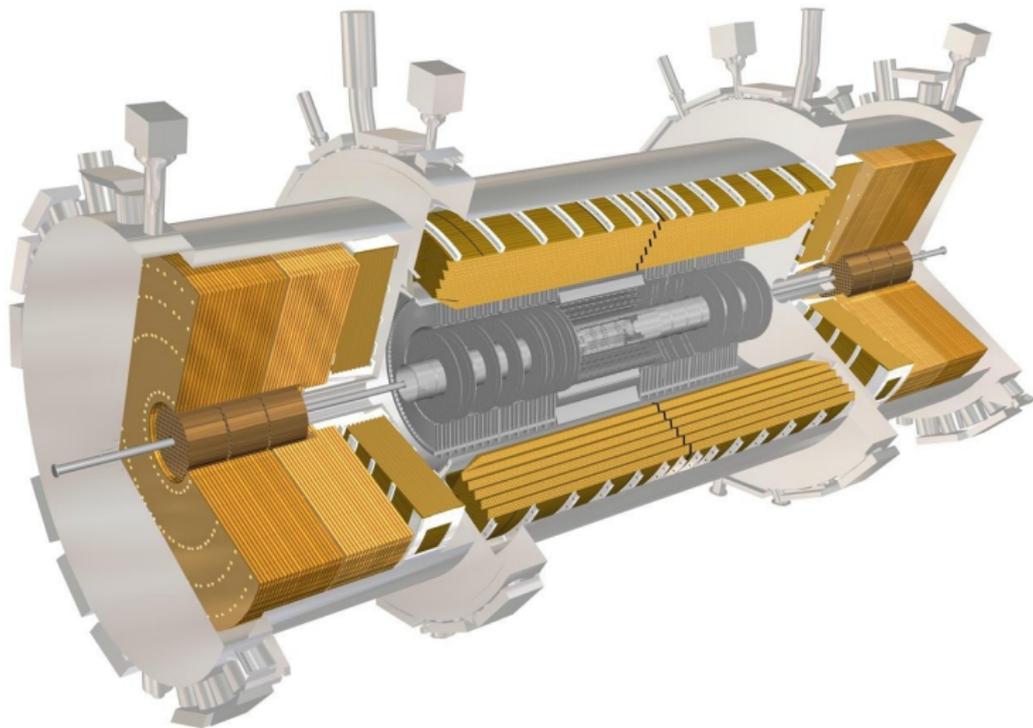
# ATLAS

A general purpose particle detector



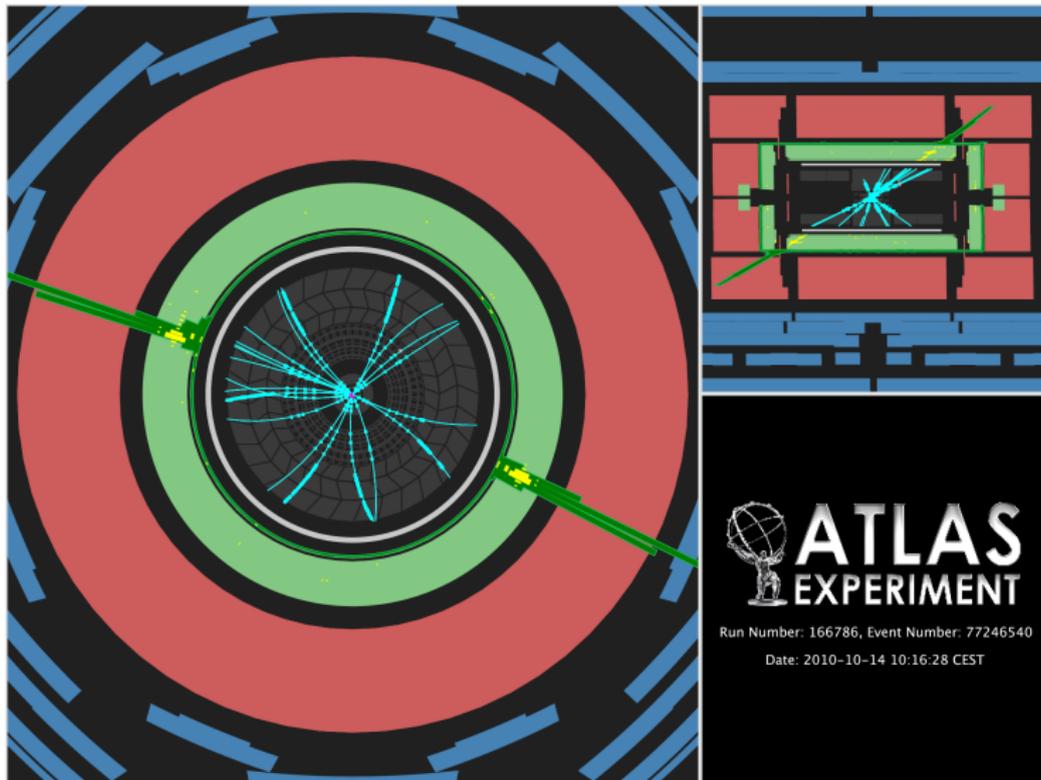
By now, a familiar image.

# The EM Calorimeter

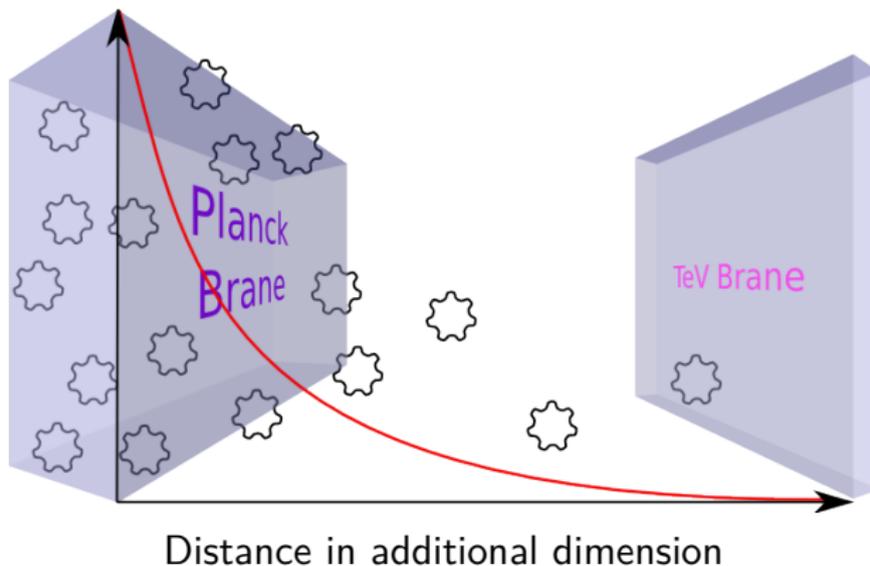


Provides good resolution and efficiency out to high energies.  
Barrel:  $|\eta| < 2.5$ , inner layer granularity  $\Delta\eta \times \Delta\phi = 0.05 \times 0.05$

# Highest invariant mass $\gamma\gamma$ pair ATLAS saw in 2010



# Brane worlds - a solution to the hierarchy problem?

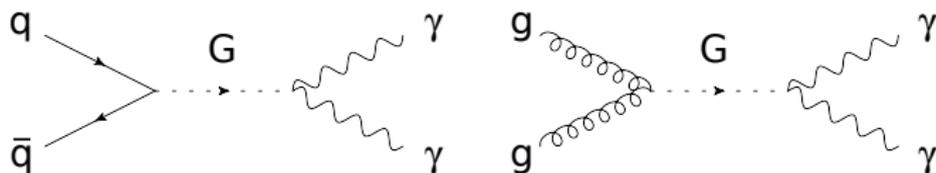


- The Randall Sundrum model
- Could solve the Hierarchy problem, explaining why Gravity is weak compared with the other forces.
- Components of momentum in the additional dimension would manifest themselves as mass in ours.

## Interesting features of $G \rightarrow \gamma\gamma$

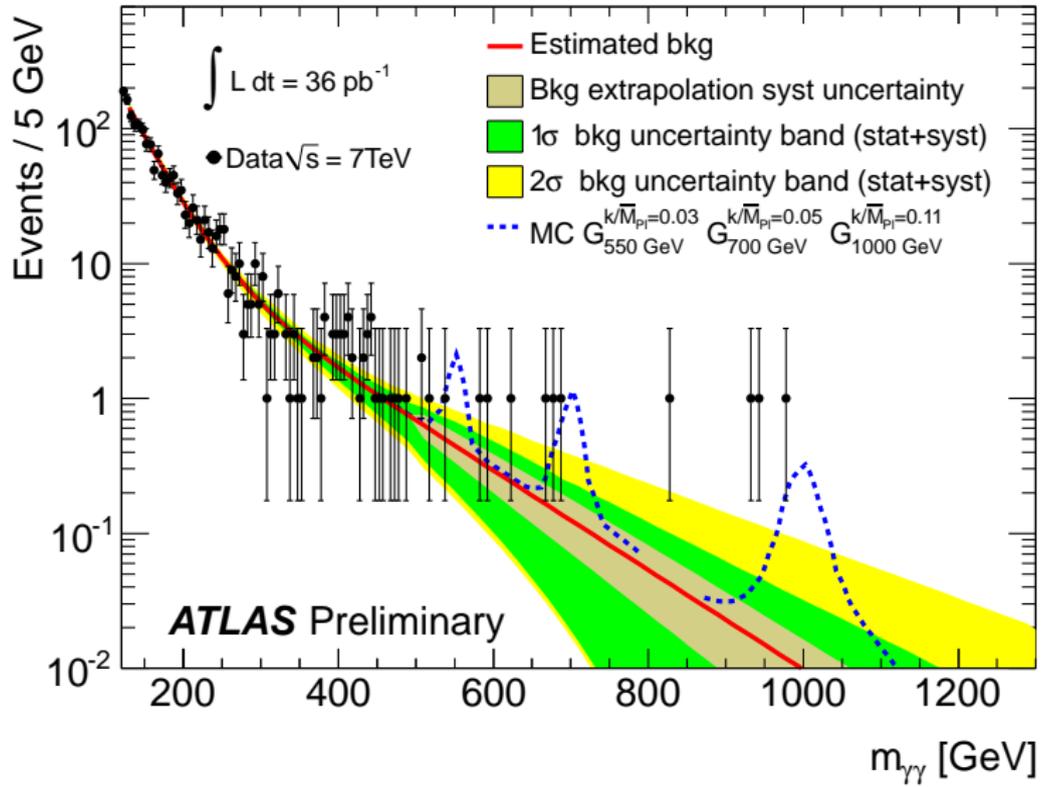
Through Kaluza Klein, the  $G$  would have a number of massive states, of which we might observe the first.

Resonant production at LHC:



- Model has two free parameters,  $M_G$  and  $k/\overline{M_{pl}}$ 
  - Resonance width ( $\Gamma$ ) and cross section ( $\sigma$ )  $\propto k/\overline{M_{pl}}$
- $\gamma\gamma$  channel is particularly interesting because the branching ratio is enhanced relative to leptons.
- Backgrounds:
  - Irreducible: SM diphoton
  - Reducible: fakes from rare QCD and multi-jet configurations
- Both fall rapidly with increasing mass. If the Graviton existed, it would appear strongly above the background.

# Mass spectrum



# Analysis Strategy

## Outline

Searching for a narrow resonance on a smoothly falling background, therefore it is necessary to know the expected number of signal and background events. Once we have these, we can apply  $CL_s$  to obtain frequentist limits.

*“What is the maximum  $\sigma \times Br$  that this process is allowed to have, given our data, with 95% confidence?”*

# Analysis Strategy

## Selection

- Standard ATLAS  $E\gamma$  data, object quality and jet cleaning cuts
- Primary vertex requirement
- $\eta$  outside crack region and inside barrel
- $p_T > 25$  GeV
- Loose  $\gamma$  identification cuts
  - (greater mass reach, preserved background shape, limit doesn't get worse)
- Diphoton pair with  $M_{\gamma\gamma} > 120$  GeV.

1,650 candidates after selections

# Analysis Strategy

## Background determination

- Background determination
    - Not possible use Monte-Carlo for QCD background - it is known not to be modelled very well and would require simulating our detector for more events than is feasible
    - Therefore it is necessary to use data-driven methods
- 1 Pick a control region, in which  $G$  is ruled out by previous experiments (e.g.  $M_{\gamma\gamma} < 500$  GeV)
  - 2 Model with a double exponential and fit to the data
  - 3 Extrapolate
  - 4 Compute uncertainties for anything which can affect the shape of the extrapolated curve.

# Analysis Strategy

Uncertainties which affect background extrapolation

Systematic uncertainties on the shape are important, for example:

- Statistical uncertainty on fit
- NLO corrections
- Choice of fit window
- Detector response
  - Energy scale
  - Resolution
  - Identification
  - Efficiency

.. and their extrapolation to high  $M_{\gamma\gamma}$ .

# Analysis Strategy

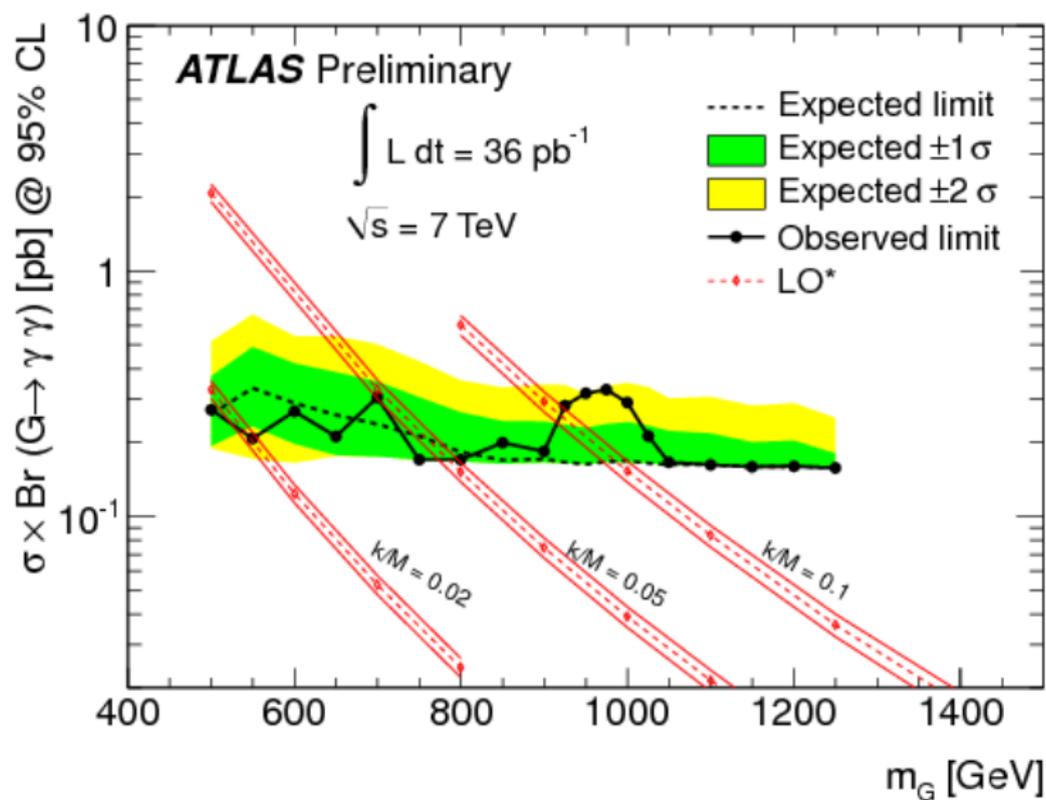
## Search and limit setting

Using the modified frequentist method in the spirit of  $CL_S$ , perform many pseudo-experiments with a test statistic for each one.

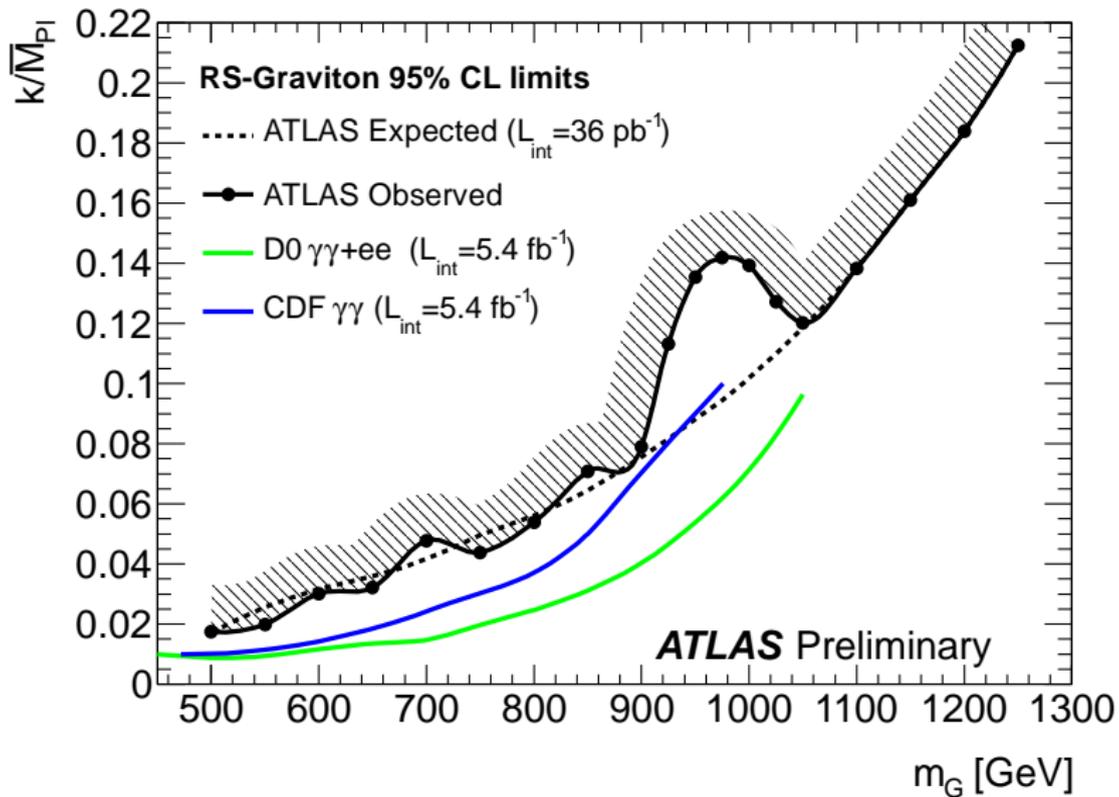
For each experiment:

- Simulate signal and background in multiple mass windows
- Define likelihood  $\mathcal{L}$  as the probability to observe the data given just the background ( $B$ ), or the signal + background ( $S+B$ ).
- Log likelihood ratio:  $LLR = \ln \mathcal{L}(S+B) - \ln \mathcal{L}(B)$
- Take the sum of the log likelihood ratio across bins
- Perform many pseudo-experiments under both  $S+B$  and  $B$  hypotheses
- Scale the cross section to find the point at which  $CL_S = 0.05$

# CLs $\sigma \times \text{Br}$ Limits



# Graviton exclusion



## Conclusion and outlook

- ATLAS is rapidly catching up with DØ and CDF sensitivity for the Randall Sundrum model, obtaining similar limits with  $150\times$  less data
- Gravitons excluded for masses below 500 GeV
  - (1200 GeV for larger couplings)
- What next?
  - Combination with  $e^+e^-$  and  $\mu^+\mu^-$  channels
  - Looking forward to 2011 data!

Paper reference: ATLAS-CONF-2011-044