

Using an Effective Charges Method to Find $\alpha_s(M_Z)$

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- 2 Event Shape Observables
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Scheme dependence problem:

- Fixed-order results depend on **renormalisation scheme** (RS) and scale, μ .
- Standard method:
 - Use \overline{MS} , set $\mu = Q$ (e.g. com energy) and vary by factor of 2.
 - Large **theoretical error**.

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- Standard method:
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 - Large **theoretical error**.
- Effective Charges Method (ECH):
 - Don't have to argue for a particular value of μ .
 - Only use μ in intermediate steps of derivation. Final result is μ **independent**.
 - Links variables to **universal Λ_{QCD}** .

Consider a generic, dimensionless QCD quantity $\mathcal{R}(Q)$. By dimensional analysis:

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Invert:

$$\frac{\Lambda}{Q} = f^{-1}(\mathcal{R}(Q)) \quad (1)$$

Want to find f^{-1} .

Take $\mathcal{R}(Q)$ to be an **effective charge**:

$$\mathcal{R}(Q) = a(1 + r_1 a + r_2 a^2 + \dots) \quad (2)$$

where $a = \alpha_s(\mu)/\pi$

QCD β -function equation:

$$\frac{\partial a}{\partial \ln \mu} = \beta(a) = -ba^2(1 + ca + c_2a^2 + \dots) \quad (3)$$

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For $\mathcal{R}(Q)$:

$$\frac{d\mathcal{R}}{d \ln Q} = \rho(\mathcal{R}(Q)) = -b\mathcal{R}^2(1 + c\mathcal{R} + \rho_2\mathcal{R}^2 + \rho_3\mathcal{R}^3 + \dots) \quad (4)$$

where $\rho_2 = c_2 + r_2 - r_1c - r_1^2$ is **RS-invariant**.

Exponentiate and rearrange to get:

$$\Lambda_{\mathcal{R}} = Q \mathcal{F}(\mathcal{R}(Q)) \mathcal{G}(\mathcal{R}(Q)) \quad (5)$$

where $\mathcal{F}(\mathcal{R}(Q)) = \exp\left[-\frac{1}{b\mathcal{R}} - \frac{c}{b} \ln \frac{c\mathcal{R}}{1+c\mathcal{R}}\right]$.

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$\Lambda_{\mathcal{R}}$ is **scheme independent** but specific to the effective charge. Relate to $\Lambda_{\overline{MS}}$ which is **universal**.

Can move between Λ 's in different schemes **exactly**.

$$\Lambda_{\overline{MS}} = \Lambda_{\mathcal{R}} \exp\left(\frac{-r}{b}\right) \left(\frac{2c}{b}\right)^{\frac{c}{b}} \quad (6)$$

where $r = r_1^{\overline{MS}}(Q)$.

Can convert $\Lambda_{\overline{MS}}$ to α_s using the β -function eqn at an appropriate order.

Event Shape (ES) Observables

Using data from LEP (DELPHI, OPAL, L3) and PETRA (JADE).

Moments of: $1-T$, C , B_W , B_T , Y_3 , ρ_E

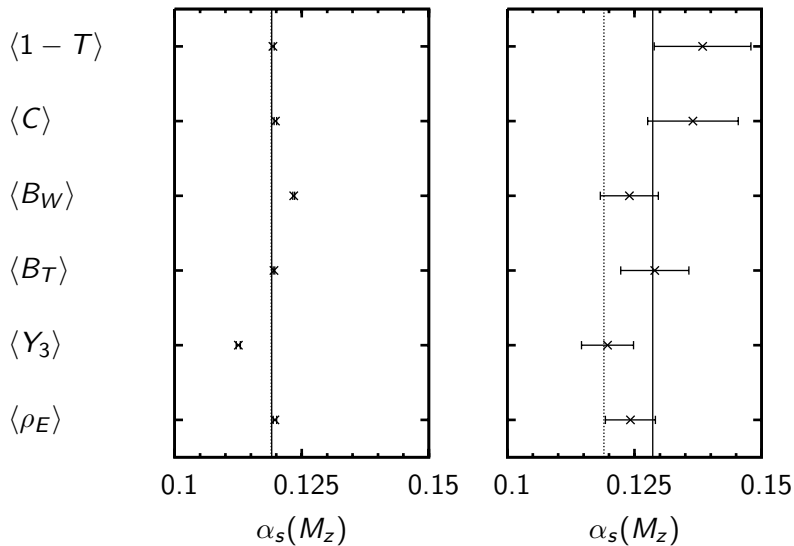
ES moments have PT expansion:

$$\langle y^n \rangle = \left(\frac{\alpha_s(Q)}{2\pi} \right) \bar{\mathcal{A}}_{y,n} + \left(\frac{\alpha_s(Q)}{2\pi} \right)^2 \bar{\mathcal{B}}_{y,n} + \left(\frac{\alpha_s(Q)}{2\pi} \right)^3 \bar{\mathcal{C}}_{y,n} + \dots \quad (7)$$

DELPHI [DELPHI coll., Feb 2003, hep-ex/0307048] have looked at means ($n=1$) with NLO ECH. Very promising results.

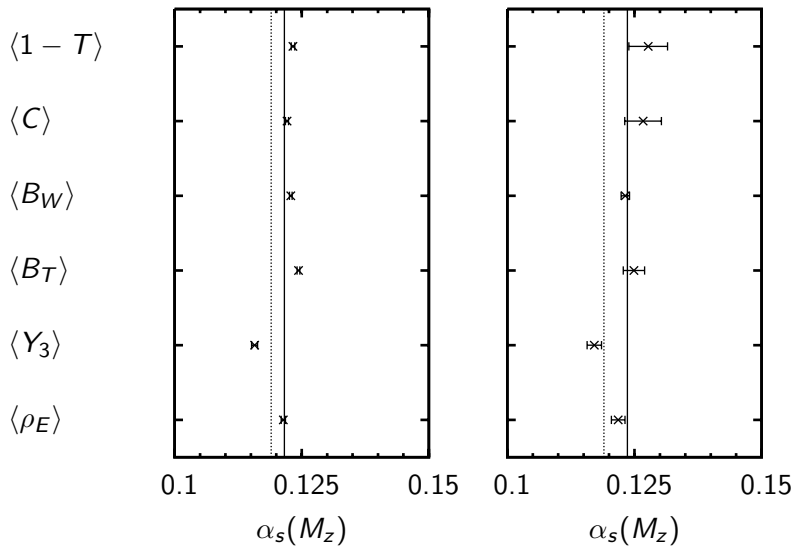
Results: Means

NLO ECH and $\overline{\text{MS}}$ PT results for the means:



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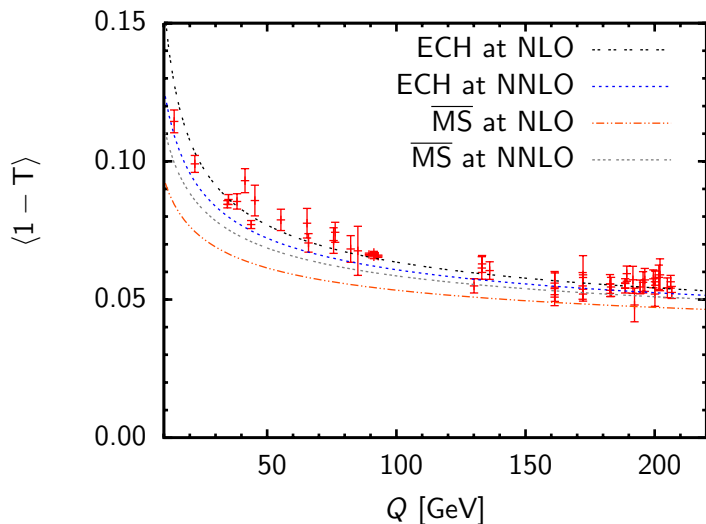
Results: Plots approximating the data

Look at a different way of implementing ECH method:

- Take a fixed value of $\alpha_s(M_Z) = 0.1190$ [Baikov, Chetyrkin, Kühn; Jan 2008]
- Work **backwards** through ECH method to get an approximation to data.
- Can compare with $\overline{\text{MS}}$ results for fixed α_s using (7).

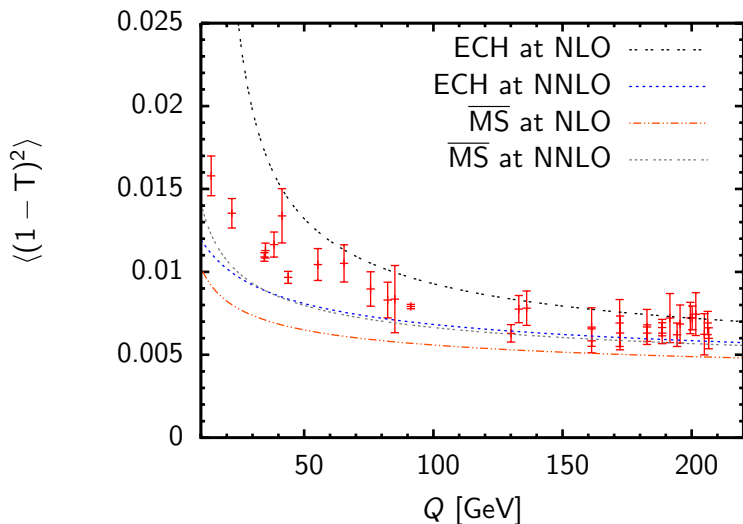
ECH plots: $\langle 1 - T \rangle$

ECH and $\overline{\text{MS}}$ curves at NLO and NNLO against data for $\langle 1 - T \rangle$:



ECH plots: $\langle(1 - T)^2\rangle$

ECH and $\overline{\text{MS}}$ curves at NLO and NNLO against data for $\langle(1 - T)^2\rangle$:



Looking at $\Lambda_{\mathcal{R}}$

Look at values of $\Lambda_{\mathcal{R}}$ to help see where the perturbative regime is:

$$\begin{aligned}\langle 1 - T \rangle : & \quad \Lambda_{\mathcal{R}} = 4 \text{ GeV} \\ \langle (1 - T)^2 \rangle : & \quad \Lambda_{\mathcal{R}} = 24 \text{ GeV}\end{aligned}$$

Have data over the range 14.0-206.6 GeV.

Conclusions:

- ECH at NLO works well for ES means - better than \overline{MS}
- ECH at NNLO often works less well than NLO
- Some higher moments are not suited to this method

In progress:

- Inclusion of non-perturbative power corrections

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Thanks for listening!

Back ups: ES definitions

Thrust:

$$T = \max \frac{\sum_i |\mathbf{p}_i \cdot \mathbf{n}_T|}{\sum_i |\mathbf{p}_i|}$$

C-parameter - defined by the eigenvalues of the linear momentum tensor:

$$\theta_{m,n} = \frac{\sum_i p_i^m p_i^n / |\mathbf{p}_i|}{\sum_i |\mathbf{p}_i|}$$
$$C = 3(\lambda_1 \lambda_2 + \lambda_2 \lambda_3 + \lambda_3 \lambda_1)$$

Wide jet broadening - define jet broadening in two hemispheres divided by plane perpendicular to \mathbf{n}_T :

$$B_x = \frac{\sum_{\mathbf{p}_k \in H_x} |\mathbf{p}_k \times \mathbf{n}_T|}{2 \sum_{\mathbf{p}_k \in H_x} |\mathbf{p}_k|}$$
$$B_W = \max(B_a, B_b)$$

Back ups: ES definitions cont.

Total jet broadening - also defined by B_x :

$$B_T = B_a + B_b$$

Three-to-two jet transition parameter in the Durham algorithm:

$$y_{ij}^D = \frac{2 \min(E_i^2, E_j^2)(1 - \cos \theta_{ij})}{E_{\text{vis}}^2}$$

Heavy jet mass in the E-scheme - hemisphere observable:

$$\rho = \max \frac{M_x^2}{E_{\text{vis}}^2} = \max \frac{1}{E_{\text{vis}}^2} \left(\sum_{\mathbf{p}_k \in \mathbf{H}_x} p_k \right)^2 \quad (8)$$

The use of the E-scheme corrects for the effect of hadron masses. In the E-scheme the four-momentum $p = (\mathbf{p}, E)$ is replaced by $(\hat{p}E, E)$ where \hat{p} is a unit vector in the \mathbf{p} -direction. The definition in the E-scheme is identical to (8) for massless hadrons.

Hadronisation corrections for event shape moments are expected to be additive:

$$\langle y^n \rangle = \langle y^n \rangle_{pt} + \langle y^n \rangle_{np} \quad (9)$$

Dispersive model for power corrections: accounts for NP behaviour at low energies by replacing α_s with an α_{eff} below an IR cutoff scale, μ_I . Done s.t. the integral of the coupling up to μ_I is finite:

$$\frac{1}{\mu_I} \int_0^{\mu_I} dQ \alpha_{eff}(Q^2) = \alpha_0(\mu_I) \quad (10)$$

Leads to:

$$\langle y \rangle_{np} = a_y P$$

Back ups: power corrections (cont.)

with:

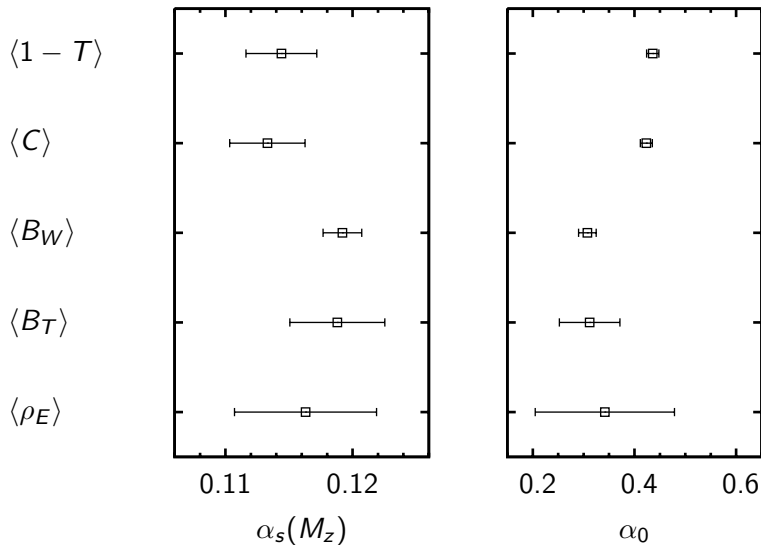
$$P = \frac{4C_F}{\pi^2} \mathcal{M} \left(\alpha_0 - \left[\alpha_s(\mu_R) + \frac{b}{\pi} \left(1 + \ln \frac{\mu_R}{\mu_I} + \frac{K}{2b} \right) \alpha_s^2(\mu_R) \right. \right. \\ \left. \left. + \left(4bc \left(1 + \ln \frac{\mu_R}{\mu_I} + \frac{L}{4bc} \right) + 8b^2 \left(1 + \ln \frac{\mu_R}{\mu_I} + \frac{K}{2b} \right) \right. \right. \right. \\ \left. \left. \left. + 4b^2 \ln \frac{\mu_R}{\mu_I} \left(\ln \frac{\mu_R}{\mu_I} + \frac{K}{b} \right) \right) \frac{\alpha_s^3(\mu_R)}{4\pi^2} \right] \right) \frac{\mu_I}{Q}$$

where \mathcal{M} is the Milan factor and:

$$K = \left(\frac{67}{18} - \frac{\pi^2}{6} \right) C_A - \frac{5}{9} N_F \\ L = C_A^2 \left(\frac{245}{24} - \frac{67}{9} \frac{\pi^2}{6} + \frac{11}{6} \zeta_3 + \frac{11}{5} \left(\frac{\pi^2}{6} \right)^2 \right) + C_F N_F \left(-\frac{55}{24} + 2\zeta_3 \right) \\ + C_A N_F \left(-\frac{209}{108} + \frac{10}{9} \frac{\pi^2}{6} - \frac{7}{3} \zeta_3 \right) - \frac{1}{27} N_F^2$$

Back ups: power corrections (cont.)

ECH and power corrections at NLO:



Integrate up (4):

$$\begin{aligned} b \ln \frac{Q}{\Lambda_{\mathcal{R}}} &= \int_0^{\mathcal{R}(Q)} \frac{b dx}{\rho(x)} + \kappa \\ &= \int_0^{\mathcal{R}(Q)} \frac{b dx}{\rho(x)} + \int_0^{\infty} \frac{dx}{x^2(1+cx)} \end{aligned}$$

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 &= F(\mathcal{R}) + G(\mathcal{R}) \\
 &= \frac{1}{\mathcal{R}} + c \ln \left[\frac{c\mathcal{R}}{1+c\mathcal{R}} \right] + G(\mathcal{R}) \tag{11}
 \end{aligned}$$

Result for $G(\mathcal{R})$ depends on the order you work to. At NLO $G(\mathcal{R}) = 0$. At NNLO there is an analytic result.

ES moments have PT expansion:

$$\begin{aligned} \langle y^n \rangle = & \left(\frac{\alpha_s(\mu_R)}{2\pi} \right) \bar{\mathcal{A}}_{y,n} + \left(\frac{\alpha_s(\mu_R)}{2\pi} \right)^2 \left(\bar{\mathcal{B}}_{y,n} + \bar{\mathcal{A}}_{y,n} b \ln \frac{\mu_R^2}{Q^2} \right) \\ & + \left(\frac{\alpha_s(\mu_R)}{2\pi} \right)^3 \left(\bar{\mathcal{C}}_{y,n} + 2\bar{\mathcal{B}}_{y,n} b \ln \frac{\mu_R^2}{Q^2} + \bar{\mathcal{A}}_{y,n} \left(b^2 \ln^2 \frac{\mu_R^2}{Q^2} + 2bc \ln \frac{\mu_R^2}{Q^2} \right) \right) \\ & + \dots \end{aligned} \quad (12)$$

Results: $\langle(1 - T)^n\rangle$

NLO and NNLO ECH results for $\langle(1 - T)^n\rangle$:

$\langle 1 - T \rangle$

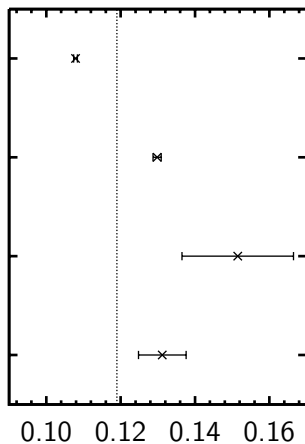
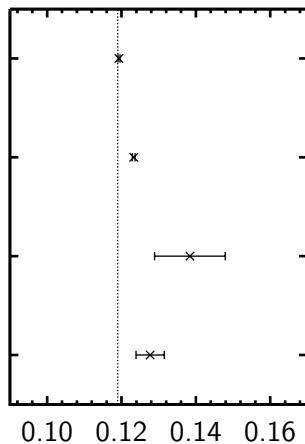
$\langle(1 - T)^2\rangle$

ECH, NLO

ECH, NNLO

\overline{MS} , NLO

\overline{MS} , NNLO



Asymptotic Series

Look at PT expansions of ESs. These are asymptotic series.

Definition of Asymptotic Series

$$f(g) = \sum_{n=0}^{\infty} f_n g^n$$

is asymptotic if the series diverges for all $g \neq 0$ and satisfies the bound:

$$\left| f(g) - \sum_{n=0}^N f_n g^n \right| \leq f_{N+1} |g|^{N+1}$$

So error is **bounded by the size of the first term you neglect**.

Asymptotic Series

So look at PT expansion for $\rho(\mathcal{R}) = -b\mathcal{R}^2(1 + c\mathcal{R} + \rho_2\mathcal{R}^2 + \dots)$:

$\langle 1 - T \rangle$:

- $\rho(\mathcal{R}) = -b\mathcal{R}^2(1 + 0.0785 - 0.0494)$

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$\langle 1 - T \rangle$:

- $\rho(\mathcal{R}) = -b\mathcal{R}^2(1 + 0.0785 - 0.0494)$

$\langle (1 - T)^2 \rangle$:

- $\rho(\mathcal{R}) = -b\mathcal{R}^2(1 + 0.1049 - 0.3227)$