

CryoEDM: building a simulation framework for ultra-cold neutrons

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- CryoEDM
 - What are UCN?
 - Simulating UCN in CryoEDM

Why are we interested in EDMs anyway?

- The existence of a permanent EDM in fundamental particles would be a violation of Time reversal symmetry and therefore CP symmetry.

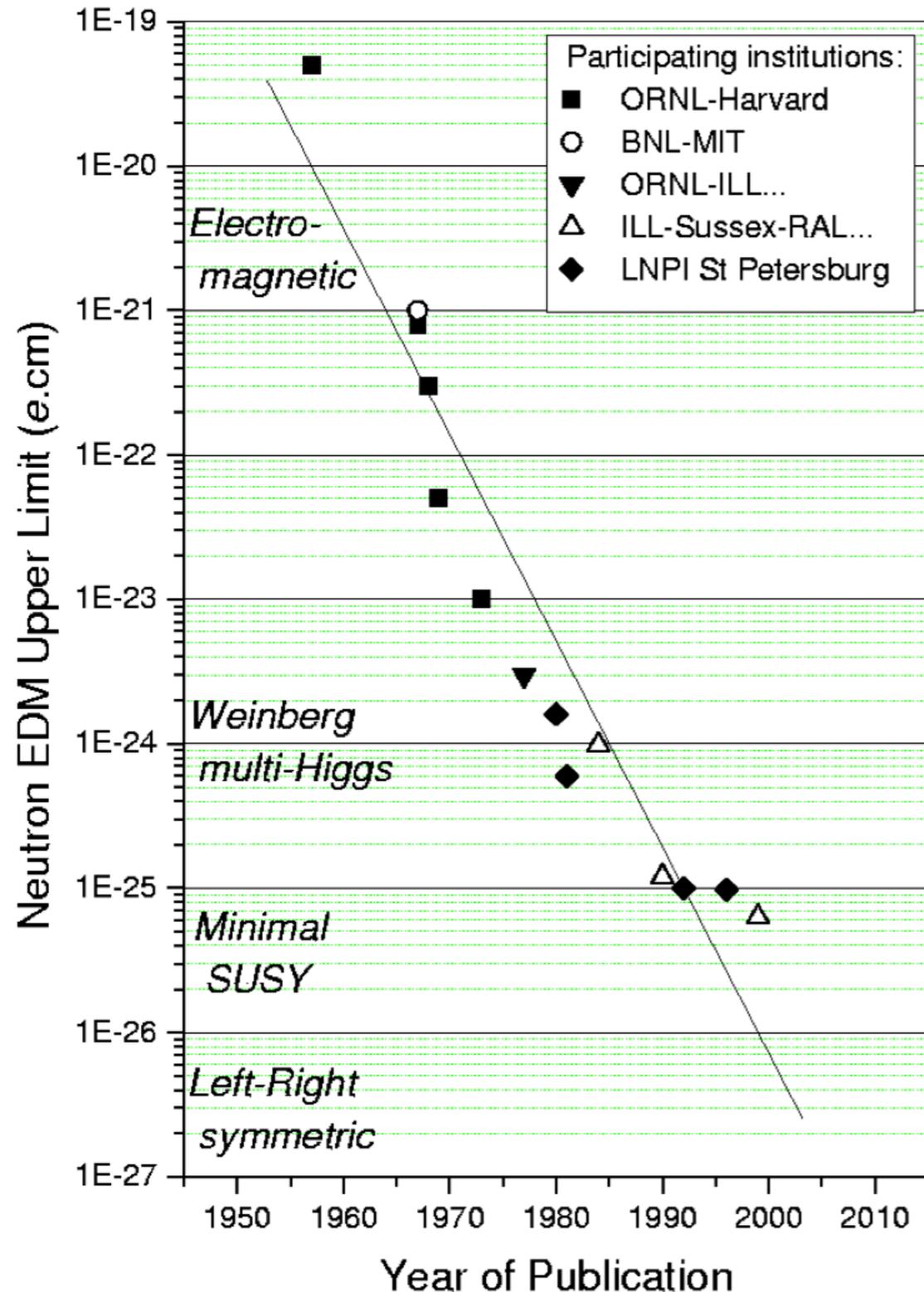
$$\vec{d} = d\vec{J}$$

$$\hat{H} = \underbrace{\vec{d} \cdot \vec{E}}$$

Asymmetric under T , P and CP

- CP violation is thought to be one of the key ingredients necessary for the predominance of matter over anti-matter in the universe.
 - ➔ Sakharov's necessary conditions for Baryogenesis
- However, sources of CP-violation in the standard model are generally thought to be **insufficient to generate presently observed baryon asymmetry**.

Neutron EDM Measurements



- Many different theoretical models predict new sources of CP-violation (such as the many from supersymmetry) and therefore **give predictions for the size of the nEDM.**
- Thus, searches for a nEDM place experimental limits on many theories of beyond-standard-model physics

Sussex-RAL 2006 Best Limit:

$$d_n < 2.9 \times 10^{-26} \text{ ecm}$$

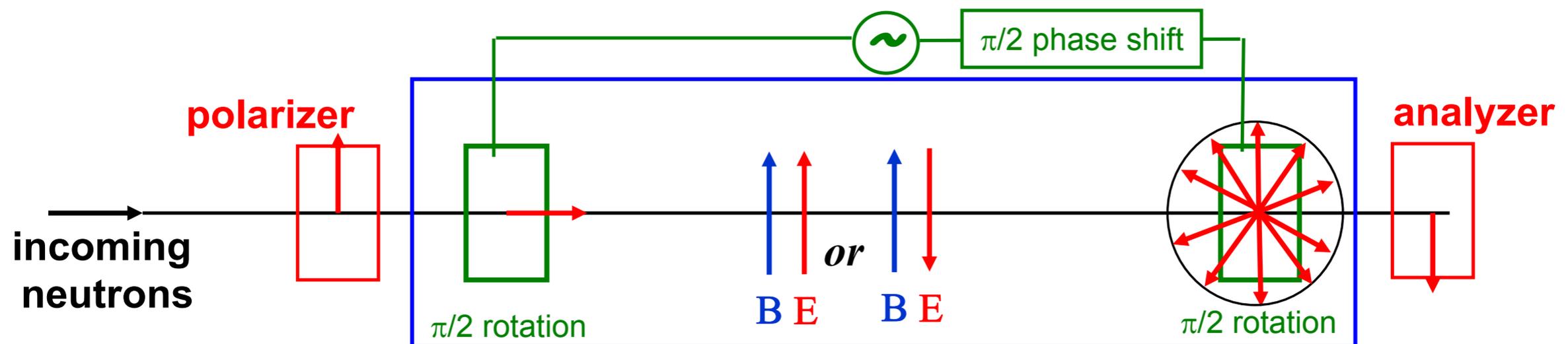
Measurement Principle

- Measure changes in neutron's Larmor precession frequency due to a potential EDM's interaction with an applied Electric field.

$$\hbar\omega_0 = -2\mu_n|\mathbf{B}_0| \mp 2d_n|\mathbf{E}_0|$$

- Changing **B** field and **E** field from parallel to anti-parallel results in,

$$\delta\omega_0 = -\frac{4d_n E_0}{\hbar}$$



EDM Statistical sensitivity

$$\sigma_{edm} \propto \frac{\hbar}{\alpha ET \sqrt{N_n}}$$

α = polarisation efficiency

E = applied E field

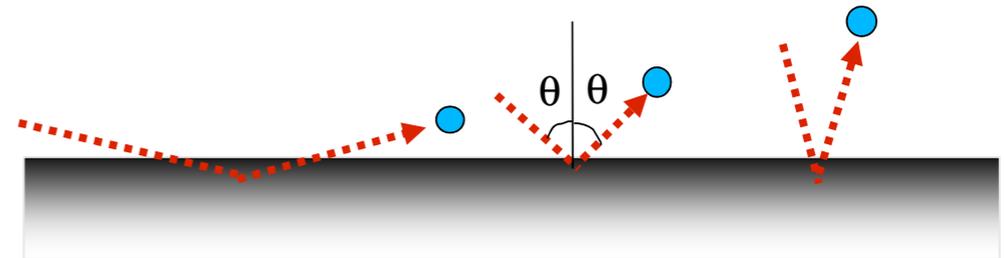
T = observation time

N = total number of neutrons detected

- Previous Sussex/RAL nEDM experiment was eventually limited by systematics - (Geometric phase effect)
- Measurement is highly dependent on a highly uniform magnetic field with variations of order $\approx 1 \text{ nT/m}$ (Earth's field $\approx 30 \mu\text{T}$)

Ultra-Cold Neutrons

- Ultra-cold neutrons are typically of energy: $E \lesssim 200 \times 10^{-9} eV$
- Typical UCN velocity: $v \leq 5 \text{ ms}^{-1}$
- Wavelength: $\lambda \geq 400 \text{ \AA}$



- Defining feature - **reflects from material boundary at all angles**
- UCN can be stored in material bottles for **hundreds of seconds**, limited by:
 - Capture by material's nuclei
 - Inelastic up-scattering by thermal vibrations.
- UCN are also heavily restricted by gravity!

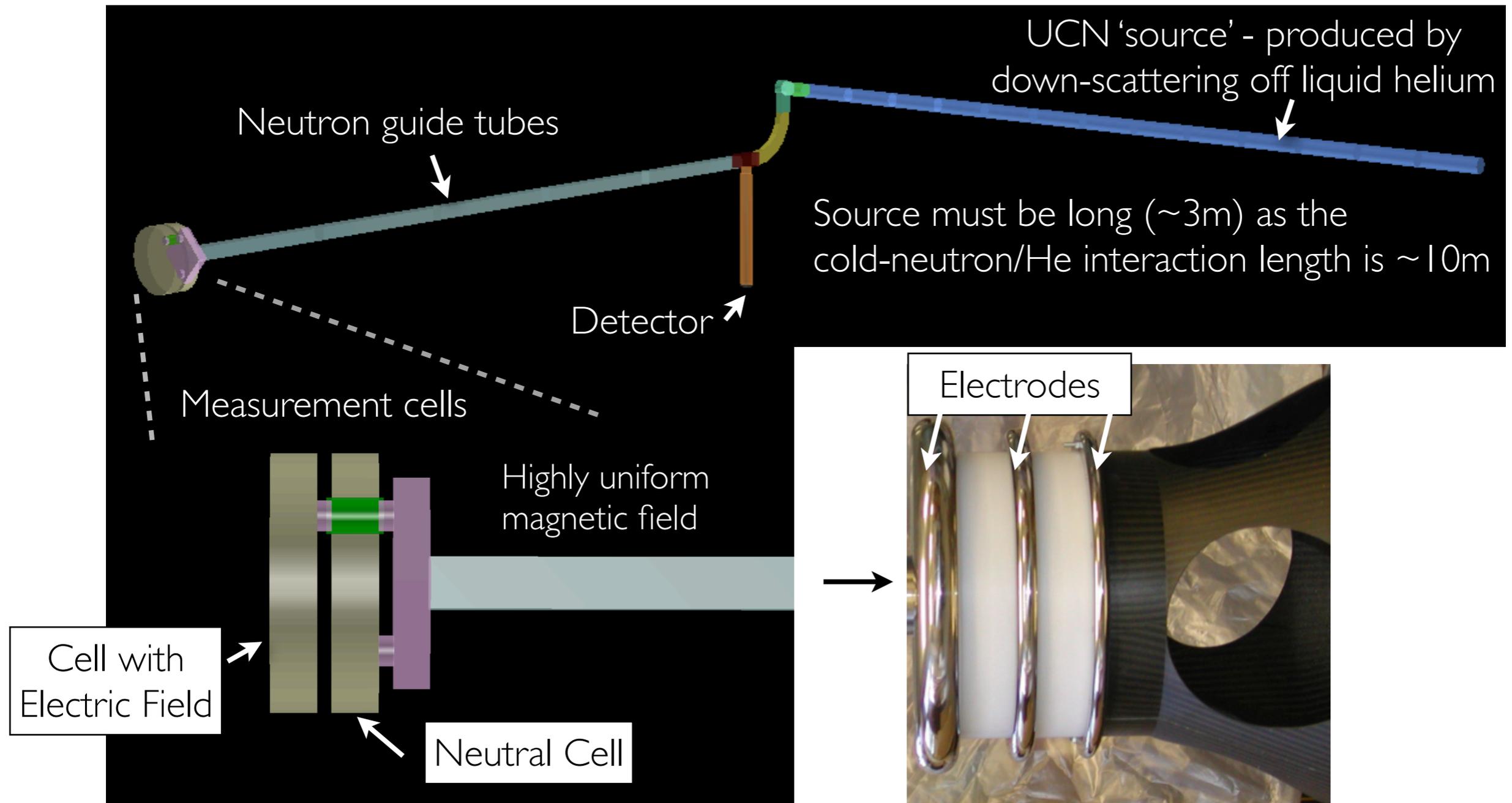
$$V_g = m_n g h = (939 \text{ MeV}/c^2 * 9.81 \text{ m/s}^2) h \approx (100 \text{ neV/m}) * h$$

Simulating Ultra-Cold Neutrons

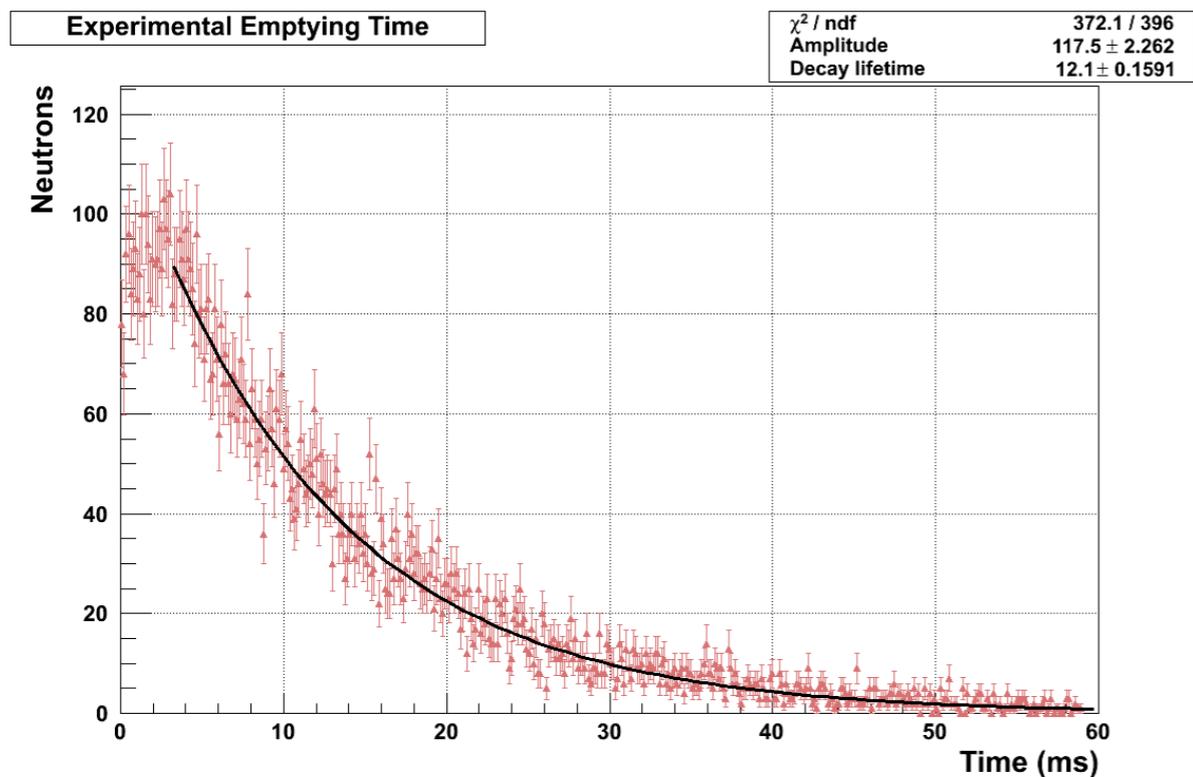
- UCN are very different to typical particle physics simulations (more akin to a ideal Gas + Quantum Mechanics)
- They also require their own special treatment when it comes to simulating them.
 - Highly gravitationally bound (parabolic trajectories)
 - UCN wall interactions (reflection, absorption, up-scattering, diffuse reflection)
 - Magnetic field interactions (neutron magnetic moment, spin-polarisation)
 - Electric field interactions ($\vec{B}' = \frac{1}{\gamma} \frac{\vec{v} \times \vec{E}}{c^2}$ relativistic motion-induced magnetic field)
- No 'off-the-shelf' simulation package in this area, so started from ground up, designed specifically for UCN physics.

Simulating Ultra-Cold Neutrons

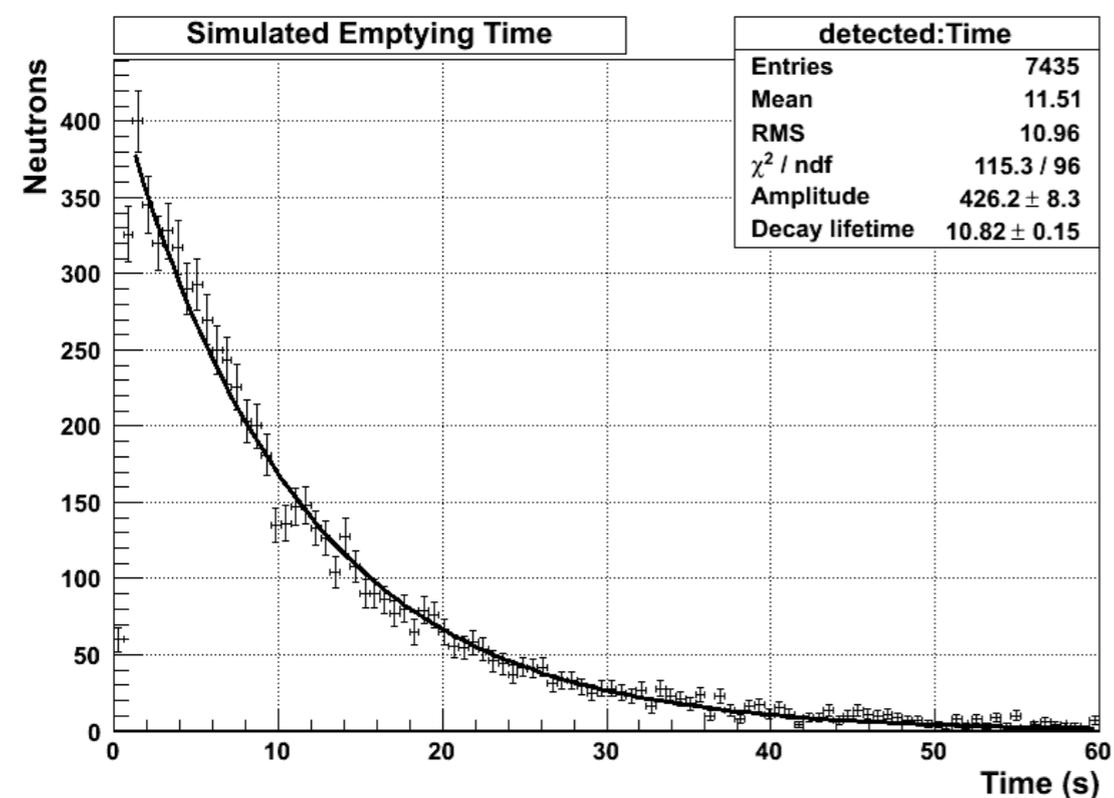
- Have developed a complete model of the neutron-facing parts of CryoEDM apparatus.



Simulating Ultra-Cold Neutrons



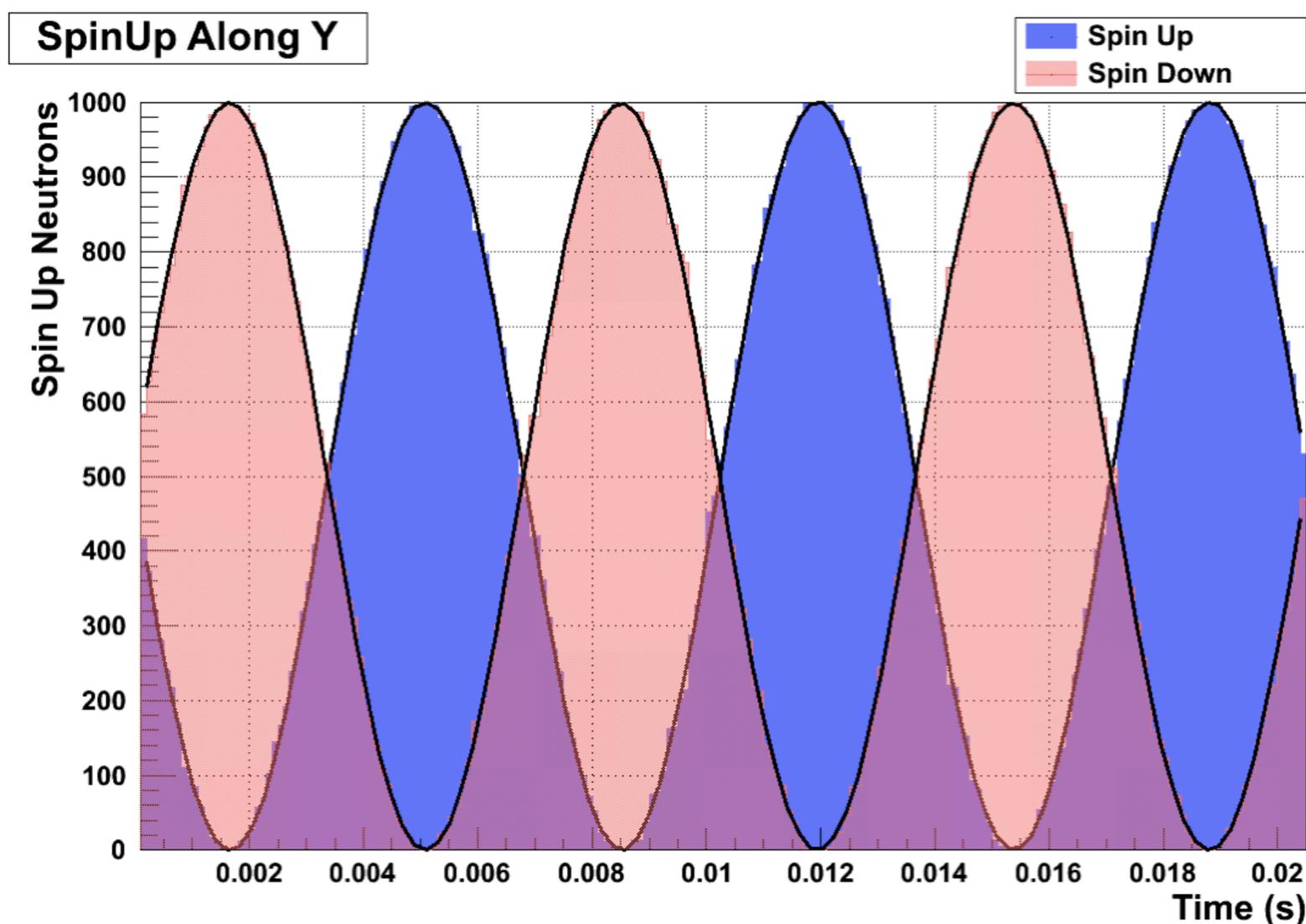
- Simulation has been tested, and is now in use on CryoEDM
 - Used to understand flow of UCN 'gas' through apparatus
 - Modelling factors that affect storage lifetime of UCN



- Able to show agreement with CryoEDM data on UCN flow through experiment

Simulating Ultra-Cold Neutrons

- We also model the interaction of neutron spin with the various static and time-varying magnetic fields present

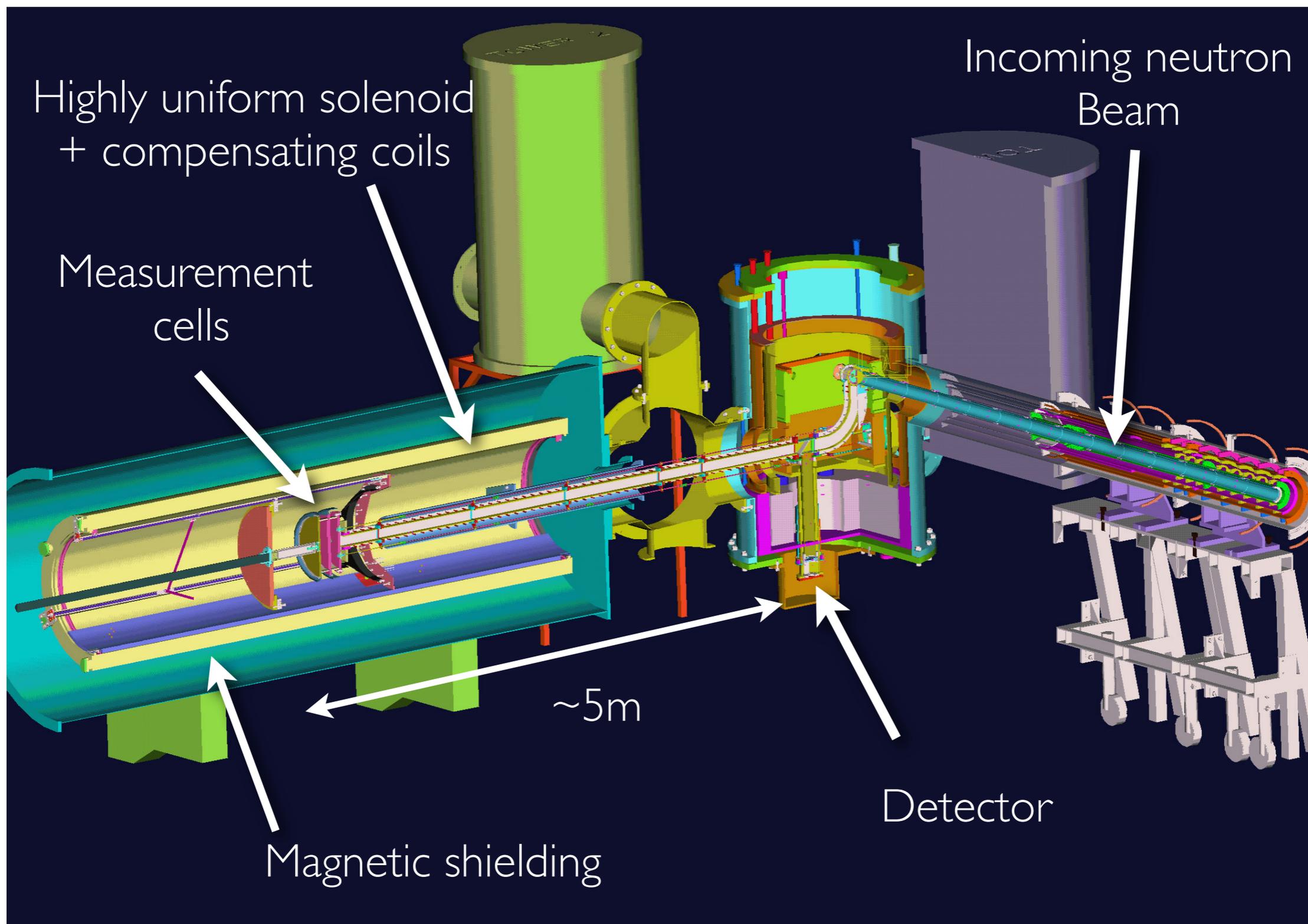


- Simulation is used to model effect of measured magnetic inhomogeneities on the characteristic depolarisation time.
- Also models systematics like $v \times E$ effect.

Summary

- Have created and tested a completely new monte-carlo simulation package for the physics of ultra cold neutrons.
- Models particle motion under gravity, utilising a new parabolic-path-finding algorithm, that supports arbitrary, complex geometries.
- We simulate neutron spin interaction under any kind of static or time-varying magnetic field (including the relativistically induced ' $v \times E$ ' magnetic field)
- Simulation shows good agreement with initial experimental data and is now used by CryoEDM in understanding the physics.

CryoEDM



1.4 Generating Matter-Antimatter Asymmetry

A.D. Sakharov, JETP Lett. 5, 24, (1967)

1. Just after the Big Bang, ($t < 10^{-6}\text{s}$), matter and antimatter were in thermal equilibrium ($T \gg 1\text{GeV}$) - balance between baryons and anti-baryons.
2. At some point, a symmetry-breaking process caused a slight imbalance between baryons and anti-baryons.
3. When the universe cooled to somewhere below ($T \sim 1\text{GeV}$), all the anti-baryons annihilated leaving a few baryons and lots of photons.
4. The photons are now microwaves in the CMB from subsequent expansion.

In this scenario the actual matter-antimatter asymmetry is really small, given by the ratio of baryons to CMB photons:

$$\frac{n_{\text{baryon}}}{n_{\gamma}} \approx 10^{-10}$$

1.5 Sakharov Conditions

- Sakharov identified three requirements for a process to generate the observed baryon asymmetry.
 1. *The process must violate baryon number conservation.*
 2. *There must be a period of non-thermal equilibrium*
 3. *There must be a process that violates CP symmetry.*
- **Can this CP/T-violation needed in the very early universe be understood better through an observable quantity today?**

2.2 Scattering From Fixed Nucleus

- At low energies can represent nuclear potential as a spherical square well.
- In this case, the solution for the scattered wave function may be expanded in eigenfunction of angular momentum:

$$\psi_s(\vec{r}) = \frac{f(\theta)}{r} e^{i\vec{k}\vec{r}}$$
$$f(\theta) = \frac{1}{2ik} \sum_l (2l+1) [e^{2i\delta_l(k)} - 1] P_l(\cos\theta)$$

- Since the cold neutron de Broglie wavelength is much larger than the range of the nuclear force, the scattering is predominantly s-wave, and

$$f(\theta) \simeq \text{constant} \equiv -a$$

which is known as the scattering length

2.3 Scattering Length

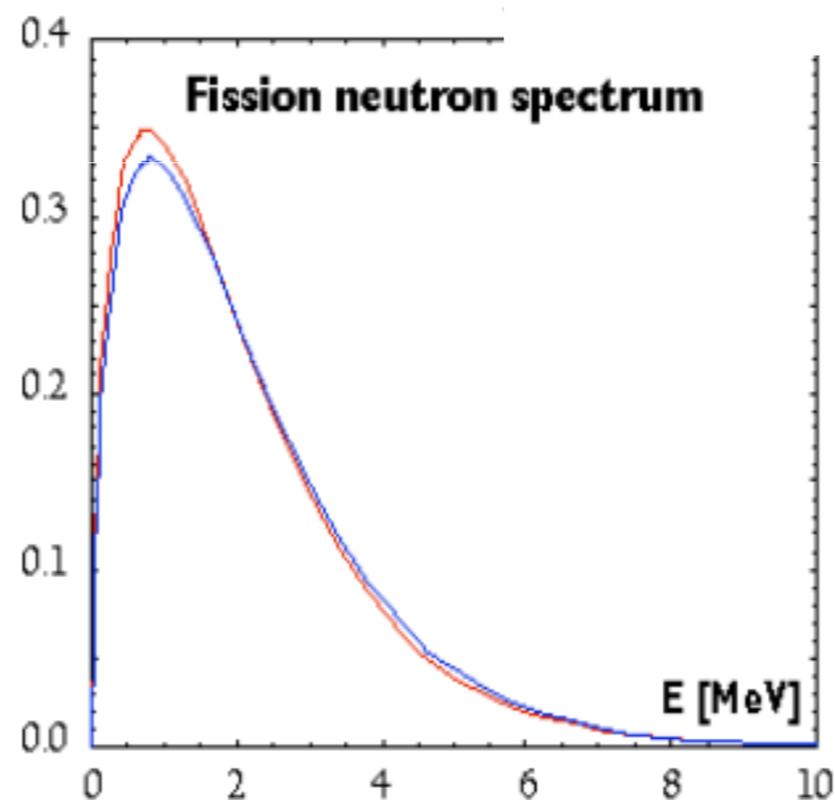
- The scattering length is actually determined experimentally and characterises the scattering cross section.

element	b (fm)
H	-3.74
Be	7.79
C	6.65
Al	3.45
Si	4.15
Ti	-3.44
Fe	9.45
Co	2.49
Ni/ ⁵⁸ Ni	10.3/14.4
Cu/ ⁶⁵ Cu	7.72/10.6
Cd	4.87-0.7 <i>i</i>

- Can be imaginary - implies strongly absorbing material. Real part can be positive or negative, with negative scattering length is also often absorbing.

2.6 UCN Production

- Basic challenge for Experiments using UCN:

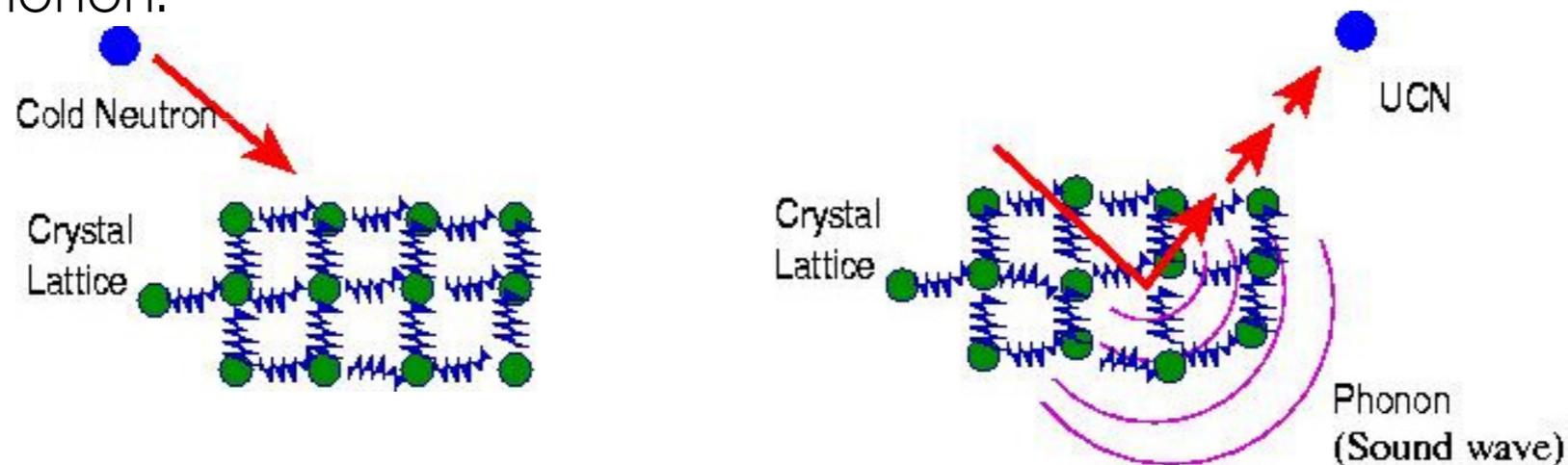


Typical UCN density in previous nEDM experiment $\sim 25 \text{ UCN cm}^{-3}$

How do we increase the proportion of UCN?

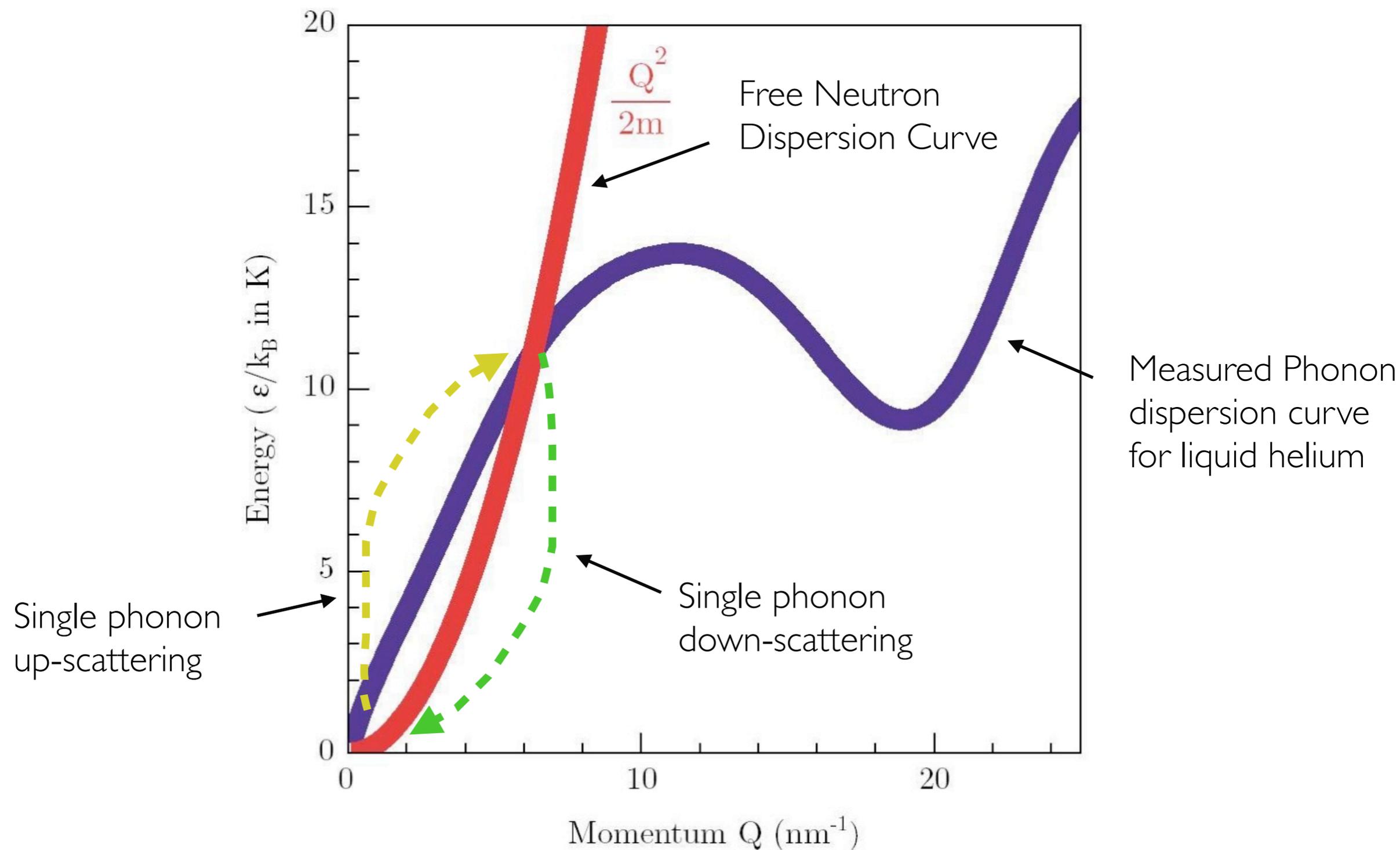
2.7 Super-Thermal UCN Production

- Pendlebury and Golub (Phys. Lett., 53, 1975) showed that a system such as superfluid Liquid ^4He can significantly increase the proportion of UCN, since it is an entirely coherent scatterer and has a zero neutron absorption cross-section.
- Cold neutrons, $\lambda \sim 9\text{\AA}$ can down-scatter into the UCN energy regime by emission of a single phonon.



- Super-fluid liquid helium, cooled to temperatures such as $\sim 0.5\text{K}$ will therefore have very few 9\AA phonons available to facilitate the reverse process - single phonon UCN up-scattering.

2.8 UCN Production in Liquid Helium



2.9 UCN Production in Liquid Helium

- Measurements of super-thermal production of UCN by neutron EDM group at Sussex/RAL found:

Measured UCN Production rate: $(0.91 \pm 0.13) \text{ cm}^{-3} \text{ s}^{-1}$

Theoretical single-phonon UCN production rate: $(1.19 \pm 0.18) \text{ cm}^{-3} \text{ s}^{-1}$

C.A.Baker et al., Phys.Lett. A308 67-74 (2002)

- Other modes of down/up-scattering, involving multiple-phonon emission/absorption, and can be particularly significant in relation to UCN up-scattering.

3.0 Experiments using UCN

- Many other research areas with an impact into BSM physics, are using UCN to increase their precision:
 - Neutron Lifetime (P. Huffman; FNSS 09)
 - Neutron Beta-decay correlations (S. Baessler; FNSS 09)
 - ‘Exotic’ neutron decay - $n \rightarrow \bar{n}$ oscillations
 - Tests of quantisation of gravity with UCN (H. Abele; FNSS 09)

2.2 Fermi Potential

- This potential in fact, **defines**, the UCN energy regime, since at this energy the neutrons will be totally internally reflected:

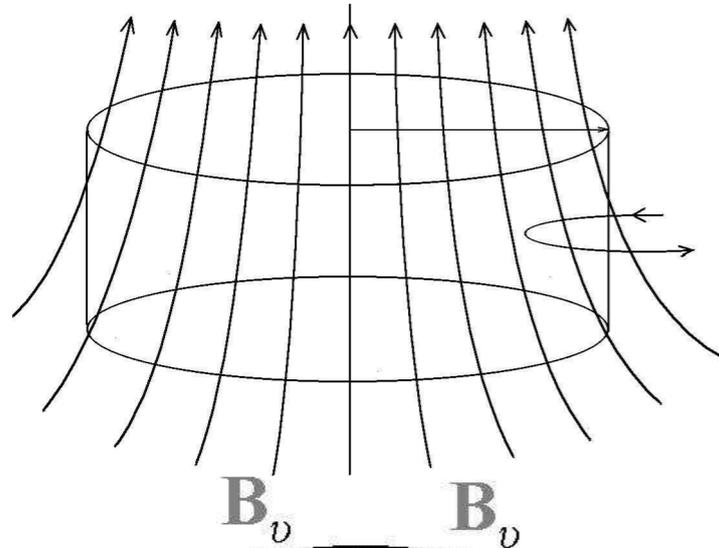
$$E_{\text{UCN}} < V_{\text{Boundary}}$$

- Potential can be positive or negative, with negative signifying an absorbing material

Material	Effective potential (units of 10^{-8} eV) (equation (2.12))	Theoretical loss factor, $f = W/V \times 10^4$ (equation (3.8))
Ni (spin up)	27	2.5
Ni (spin down)	21	
Be	25	0.1 (300 K)
C	18	0.1
Cu	17	3.1
SiO ₂	11	—
Mg	6.1	0.2
Al	5.6	0.5
Perspex	3.3	—
Polyethylene	-1.1	—
H ₂ O	-1.7	—

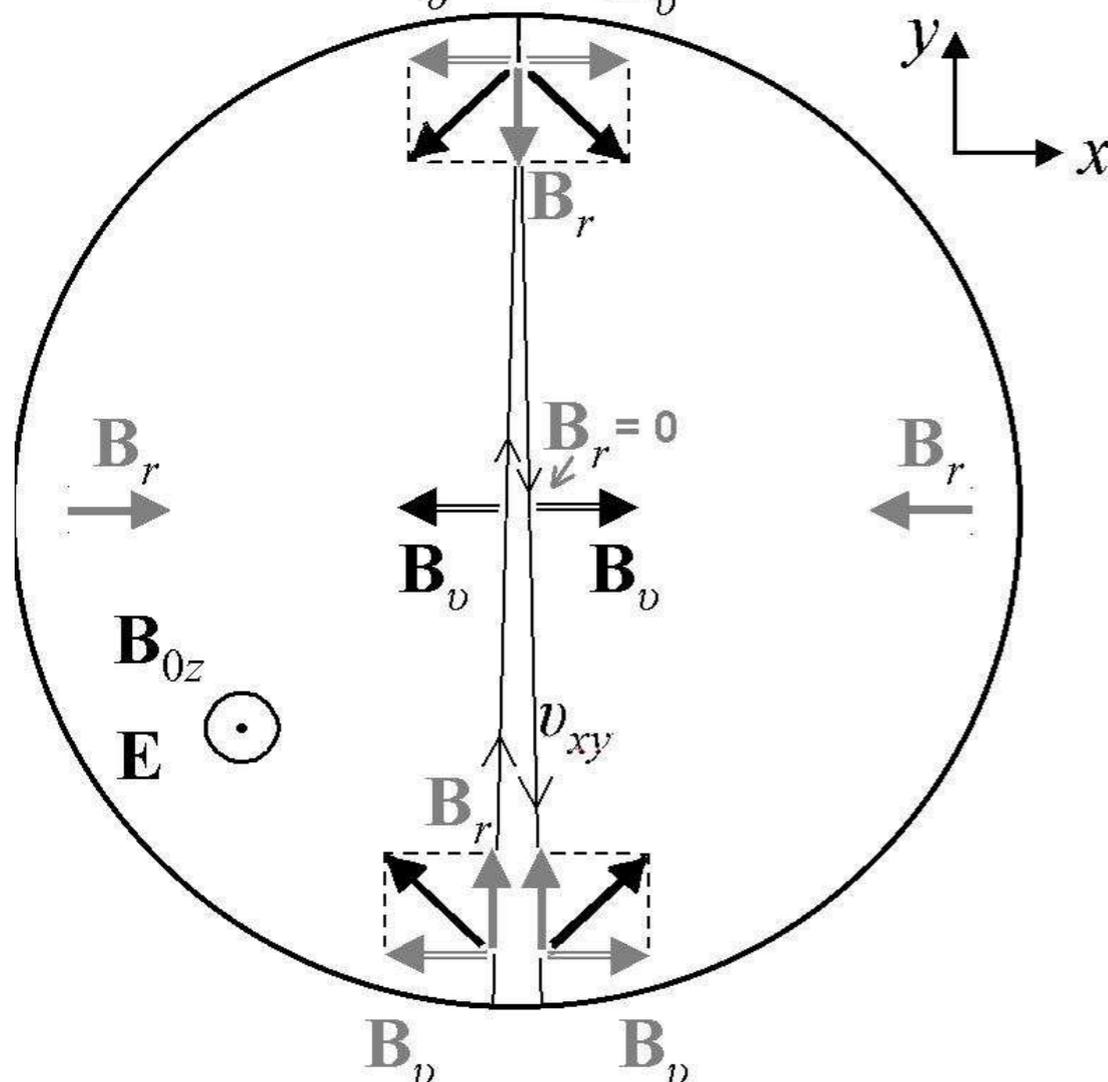
Geometric Phase

* J.M.Pendlebury et al. Phys. Rev.A 70, (2004)



- Combination of two factors:
 - Small inhomogeneities in the magnetic field
- $$\partial B_{0z} / \partial z \Rightarrow B_r \propto r$$
- Relativistic motion-induced field,

$$\vec{B}' = \frac{1}{\gamma} \frac{\vec{v} \times \vec{E}}{c^2}$$



- Two effects combine to produce an additional rotating magnetic field in x-y plane, shifting the neutron precession frequency.
- Shift is proportional to \vec{E} , just like a neutron EDM.
- Contribution to neutron EDM is roughly,

$$d_n \propto \frac{v_{\perp} |\partial B / \partial z|}{B_0^2}$$

Ultra-Cold Neutrons

- UCN wavelength is large compared to the atomic spacing in a solid and therefore interact macroscopically with the medium.
- The surface of a material can be approximated as a potential step.
- UCN can be considered as I-D wave mechanics coming to life

