

# Cauchy Horizon Stability in the Self Similar LTB Spacetime

Emily Duffy

Dublin City University, Ireland    Supported by the Irish Research Council for Science, Engineering and Technology

# Outline

- 1 Introduction
- 2 The LTB Spacetime
- 3 Cauchy Horizon Behaviour
- 4 Conclusions

# The Self-Similar LTB Spacetime

- Inhomogeneous dust collapsing into a singularity.
- Self-similarity can be imposed by requiring the spacetime to admit a homothetic Killing vector field  $\vec{\xi}$  such that

$$\mathcal{L}_{\vec{\xi}}g_{\mu\nu} = 2g_{\mu\nu}. \quad (1)$$

- The resulting line-element is given by

$$ds^2 = e^{2p}(-dz^2 + (e^\nu(z) - z^2)dp^2 - 2zdpdz + S^2(z)d\Omega^2), \quad (2)$$

where  $z = -t/r$  and  $p = \ln(r)$ .

- We get a naked singularity for  $\lambda \in (0, 0.09)$ , where  $m(r) = \lambda r$  is the Misner-Sharp mass.
- Gerlach-Sengupta formalism - gives gauge invariant linear perturbations for the stress-energy and the metric using a spherical decomposition. Even parity only.

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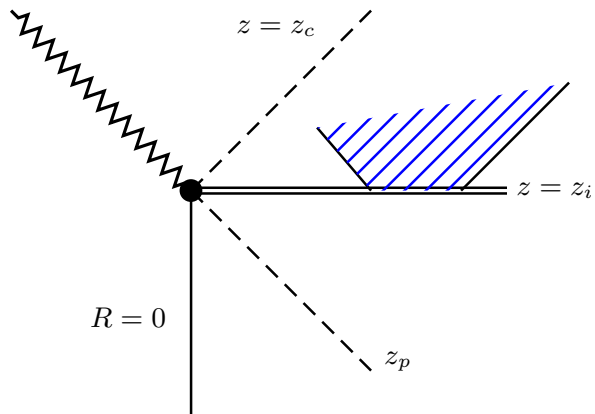
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# The Self-Similar LTB Spacetime



# First Order Reduction

- The 5 dimensional symmetric hyperbolic system

$$\tau \frac{\partial \vec{u}}{\partial \tau} + A(\tau) \frac{\partial \vec{u}}{\partial \rho} + B(\tau) \vec{u} = \vec{\Sigma}(\tau, \rho). \quad (3)$$

## Theorem 1: Existence and Uniqueness

The IVP consisting of the system (3) along with the initial data

$$\vec{u}|_{\tau_1} = \vec{f}(\rho), \quad (4)$$

where  $\vec{f} \in C_0^\infty(\mathbb{R}, \mathbb{R}^5)$ , possesses a unique solution  $\vec{u}(\tau, \rho)$ ,  $\vec{u} \in C^\infty(\mathbb{R} \times (0, \tau_1], \mathbb{R}^5)$ . For all  $\tau \in (0, \tau_1]$ ,  $\vec{u}(\tau, \cdot) : \mathbb{R} \rightarrow \mathbb{R}^5$  has compact support.

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# Behaviour of the Averaged Perturbation

- We introduce  $\hat{u} := \int_{\mathbb{R}} \vec{u} dp$  which obeys

$$\tau \frac{d\hat{u}}{d\tau} = -B(\tau)\hat{u} + \hat{\Sigma}(\tau). \quad (5)$$

## Theorem 2: Behaviour of $\hat{u}$

Let  $\vec{u}$  be a solution of (3) obeying theorem 1. Then the behaviour of  $\hat{u}$  is given by

$$\hat{u}_i \text{ is } O(1) \text{ as } \tau \rightarrow 0, \quad \hat{u}_5 \text{ is } O(\tau^{-b}) \text{ as } \tau \rightarrow 0 \quad (6)$$

for  $i = 1, \dots, 4$ , where  $b$  is the only non-zero eigenvalue of the matrix  $B(\tau = 0)$ , after it is put in Jordan canonical form.  $b > 0$ .

## Theorem 3: $L^q$ Divergence

$\lim_{\tau \rightarrow 0} \|\vec{u}\|_q = \infty$  for  $1 \leq q < \infty$ .

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## Behaviour of $\vec{x}$ on the Cauchy horizon

Corollary 1: Bound on  $\vec{x} = \tau^b \vec{u}$

$\vec{x} := \tau^b \vec{u}$  is uniformly bounded in the range  $\tau \in (0, \tau^*)$  for some  $\tau^* > 0$ .

That is

$$|\vec{x}| = |\tau^b \vec{u}| \leq \lambda + \epsilon \tau^{2b-1} + O(\tau^{2b}), \quad (7)$$

where  $\lambda$  and  $\epsilon$  are constants.

Note that  $\hat{x}$  is non-zero on the Cauchy horizon. Define

$$\vec{x}^{(n)} := \vec{x}(\tau^{(n)}, \rho), \quad (8)$$

where  $\tau^{(n)}$  is a sequence of  $\tau$ -values such that as  $n \rightarrow \infty$ ,  $\tau^{(n)} \rightarrow 0$ .

Theorem 4:  $\vec{x}$  exists on the Cauchy horizon

$\vec{x}_{ch} := \lim_{\tau \rightarrow 0} \vec{x}(\tau, \rho)$  exists on the Cauchy horizon.

Theorem 5:  $\vec{x} \in L^1$

$\vec{x}(\cdot, \rho) \in L^1(\mathbb{R}, \mathbb{R}^5)$  for all fixed  $\tau$ .

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# Behaviour of $\vec{x}$ on the Cauchy horizon

## Lemma 1: Existence of a Dominated Subsequence

Define  $\vec{x}^{(n)}$  as in (8). Then there exists a subsequence

$$\vec{x}^{(n_m)} := \vec{x}(\tau^{(n_m)}, p) \quad (9)$$

which converges to  $\vec{x}(0, p)$  as  $n_m \rightarrow \infty$ , such that  $\vec{x}^{(n_m)}$  is dominated, that is, there exists some  $h(p) \in L^1(\mathbb{R})$  such that

$$|\vec{x}^{(n_m)}| \leq h(p) \quad \forall m. \quad (10)$$

Theorem 6:  $\vec{x}$  is non-zero at  $\tau = 0$

Define  $\vec{x}^{(n)}$  as in (8). Then there exists some interval  $(a, b) \subset \mathbb{R}$  such that  $\vec{x}(0, p) \neq 0$  within this interval.

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# Physical Interpretation of Results

- The perturbed Weyl scalars are given by

$$\delta\Psi_0 = \frac{Q}{2r^2} \bar{l}^A \bar{l}^B k_{AB}, \quad (11)$$

$$\delta\Psi_4 = \frac{Q^*}{2r^2} \bar{n}^A \bar{n}^B k_{AB}, \quad (12)$$

and the scalar  $\delta P_{-1}$  is given by  $\delta P_{-1} = |\delta\Psi_0 \delta\Psi_4|^{1/2}$ .

## Theorem 7: Perturbed Weyl Scalars

Given a solution  $\vec{u}$  to the same IVP, the perturbed Weyl scalars  $\delta\Psi_0$  and  $\delta\Psi_4$ , as well as the scalar  $\delta P_{-1}$ , generically diverge on the Cauchy horizon.

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# Conclusions

- Unsolved mystery: How to handle the problem of scaling?
- Results so far show that the even parity perturbations diverge on the Cauchy horizon for all  $l$ .
- In the odd parity case, perturbations were found to remain finite on the Cauchy horizon for all  $l$ .
- Apply these methods to the self-similar perfect fluid spacetime.

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