

Double Beta Decay of ^{48}Ca with NEMO3 Calibration Software Development for SuperNEMO

Benjamin Richards UCL

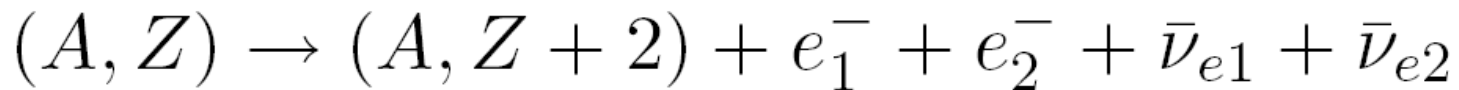
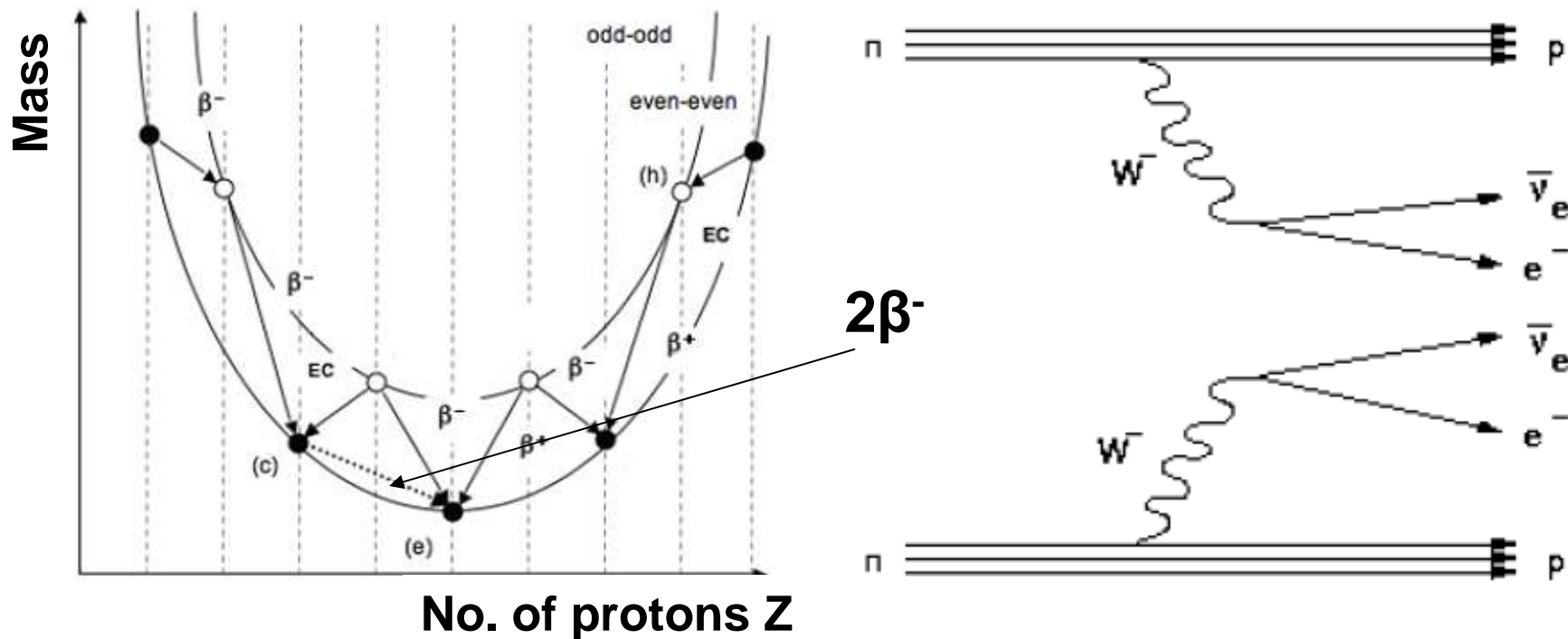
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Overview

- Description of double beta decay and its implications
- NEMO3 ^{48}Ca analysis work
- SuperNEMO software developments leading to calibration studies

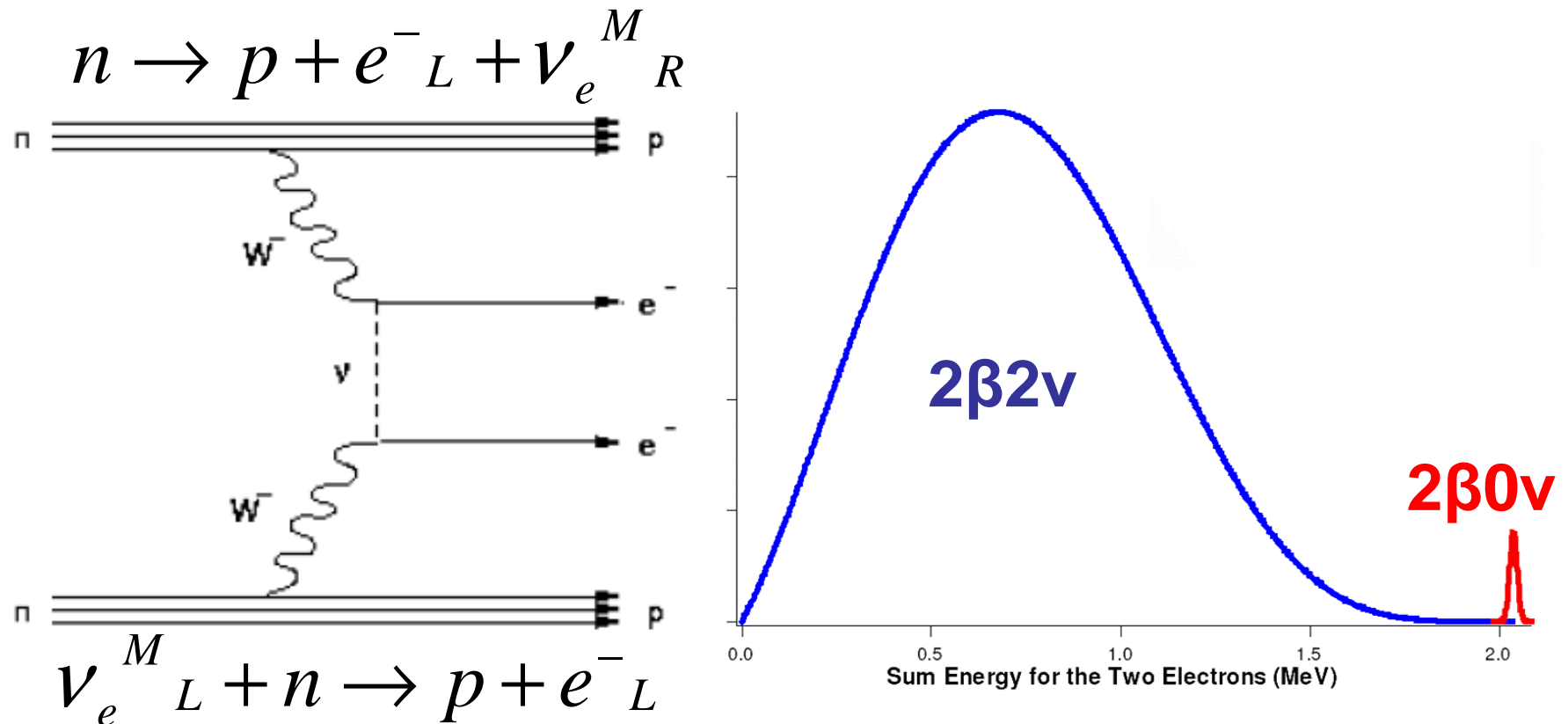
Double Beta Decay

For an atom with even A (mass no.) both even-even and odd-odd configurations can occur leading to the two parabolic distributions shown



$$(T_{1/2}^{2\nu})^{-1} = G^{2\nu} |M_{2\nu}|^2$$

Neutrinoless Double Beta Decay



$$(A, Z) \rightarrow (A, Z + 2) + e^-_1 + e^-_2$$

Mass proportional to inverse half-life

$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu} |M_{0\nu}|^2 \langle m_\nu \rangle^2$$

Neutrinoless $\beta\beta$ implications

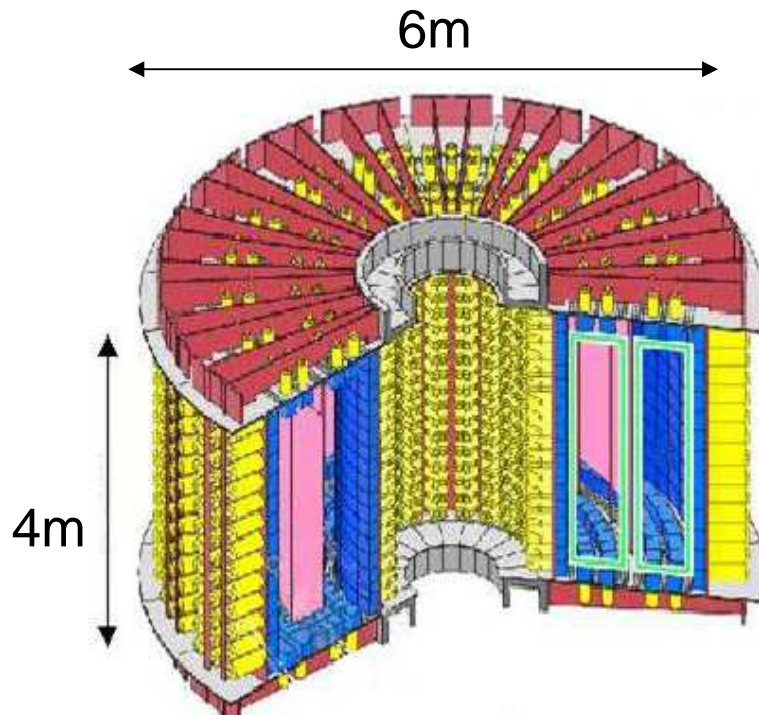
- Lepton number violating process
(**new physics**)
- Provides probe of **absolute mass** scale of neutrino
- Probes fundamental nature of neutrino
(**Dirac / Majorana**)
- Evidence for hierarchy (**normal / inverted**)
- Many GUTs require Majorana nature.

Neutrino Ettore Majorana Observatory (NEMO3)

Situated in the Fréjus Underground
Laboratory (LSM) in Modane (Savoie, France)

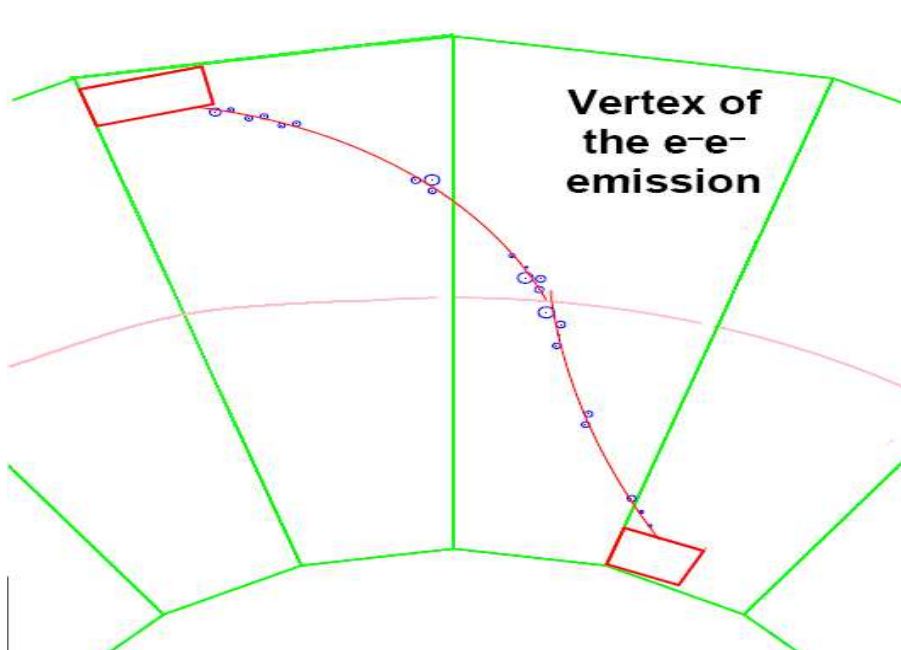
Feb 2003 - Jan 2011

NEMO3



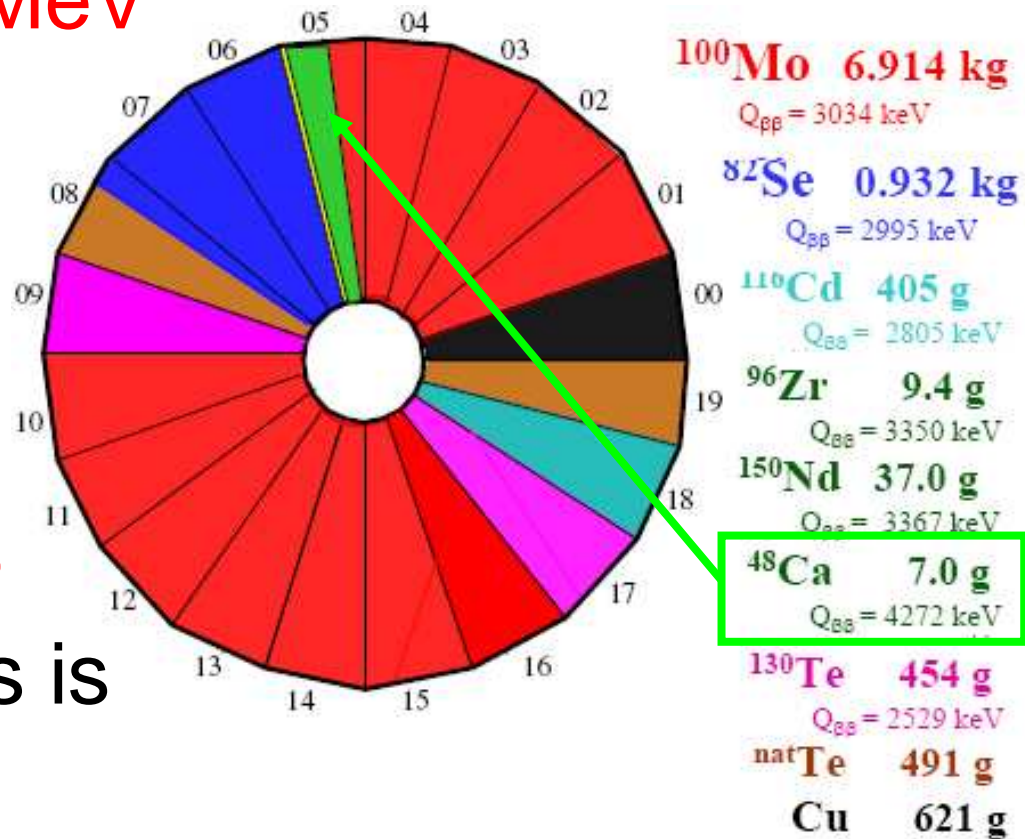
NEMO3 Detector

- Geiger drift cell tracker
- Polystyrene calorimeter
- **10kg** of source in total
- 7 different sources
- Magnetic field
- Passive shielding
- **0.3 eV** mass sensitivity
- Energy resolution ~ **14% FWHM** at 1MeV



^{48}Ca Analysis

- Highest $Q_{\beta\beta} = 4.27\text{MeV}$
- Natural abundance 0.187%
- Enriched using electromagnetic separation $\sim 39.9\%$
- Active isotope mass is 6.99g.



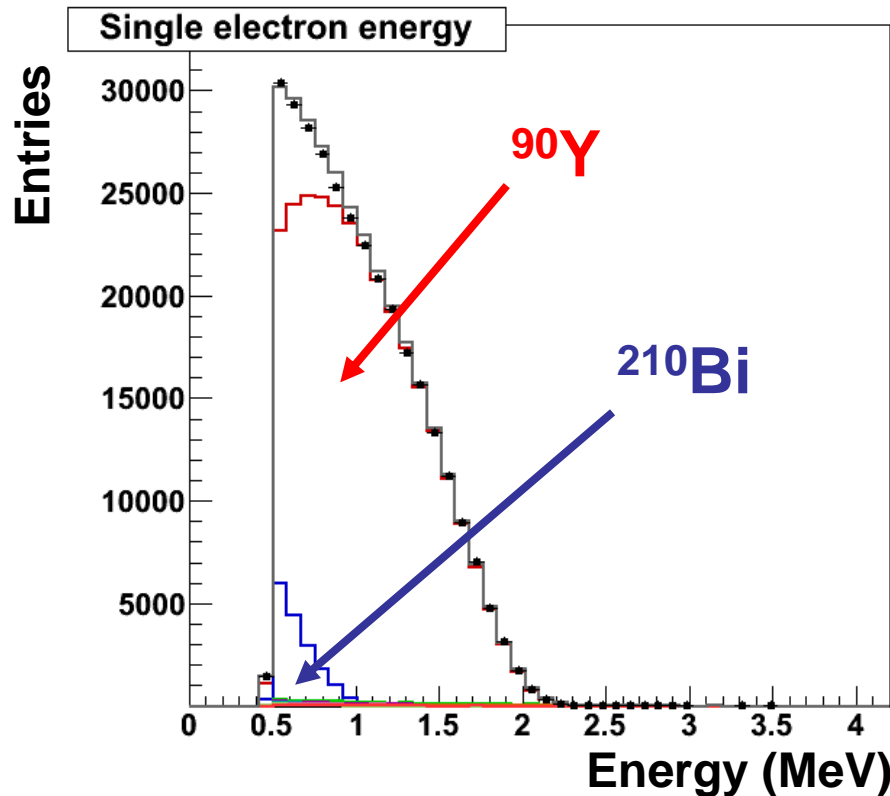
Sources of Backgrounds

- Natural impurities in PMTs and construction materials
- Internal impurities in the source foils
- Radon gas emanation leading to Bi and Pb deposits on wires and surfaces of foils

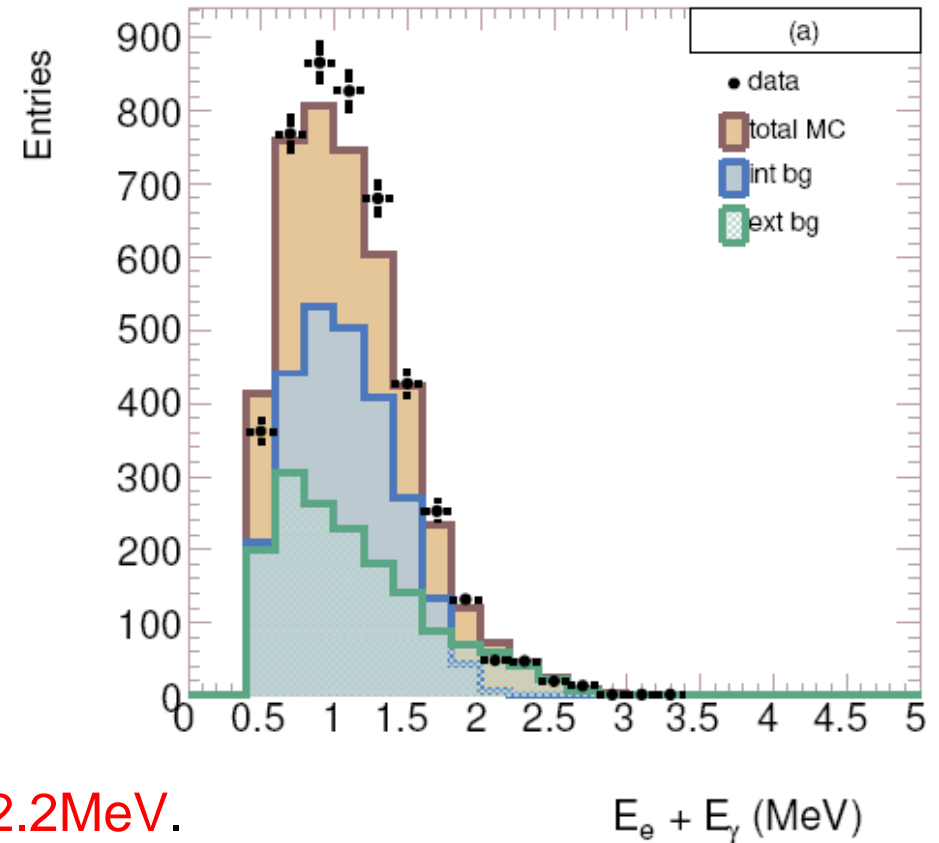
Background Component	Activity (Bq)	Efficiency to mimic 2e
^{228}Ac pmt	515	$<5.8 \times 10^{-10}$
^{208}Tl pmt	41.6	$<1 \times 10^{-8}$
^{214}Bi pmt	374	1.8×10^{-8}
^{40}K pmt	954	1×10^{-11}
^{60}Co iframe	50.7	2.5×10^{-11}
^{214}Bi swire	$(84 \pm 6) \times 10^{-3}$	$< 1 \times 10^{-6}$
^{214}Pb swire	$(84 \pm 6) \times 10^{-3}$	1.4×10^{-7}
^{208}Tl swire	2.8×10^{-3}	1.2×10^{-8}
^{210}Bi swire	5.05 ± 1	1×10^{-9}
^{214}Bi sfoil	$(8.5 \pm 1.9) \times 10^{-3}$	1.1×10^{-6}
^{214}Pb sfoil	$(8.5 \pm 1.9) \times 10^{-3}$	1.3×10^{-6}
^{210}Bi sfoil	$(17.4 \pm 5) \times 10^{-3}$	9×10^{-9}

Measuring Backgrounds in ^{48}Ca

1e channel



1e1y channel

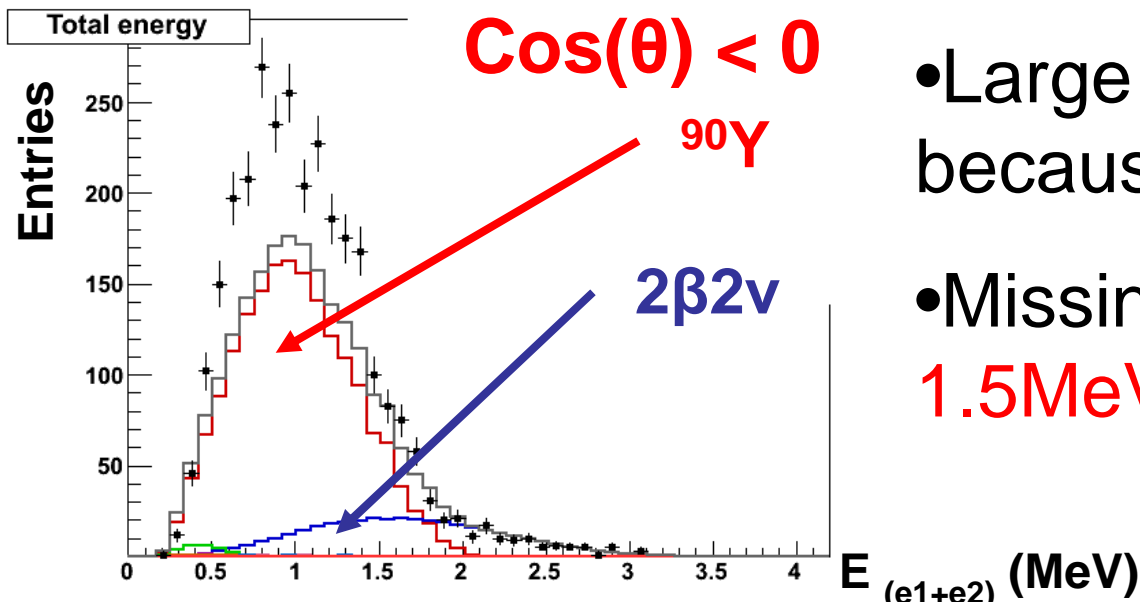
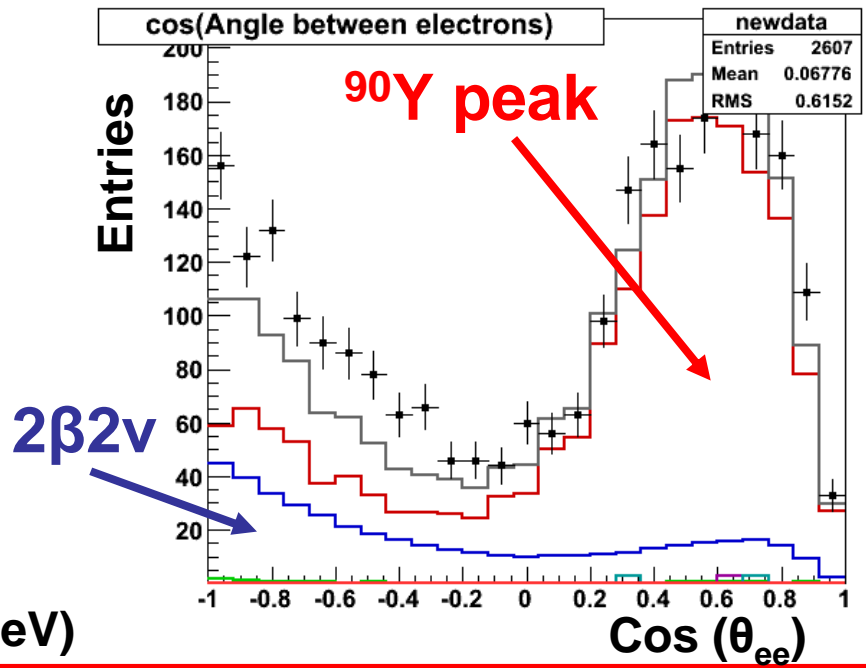
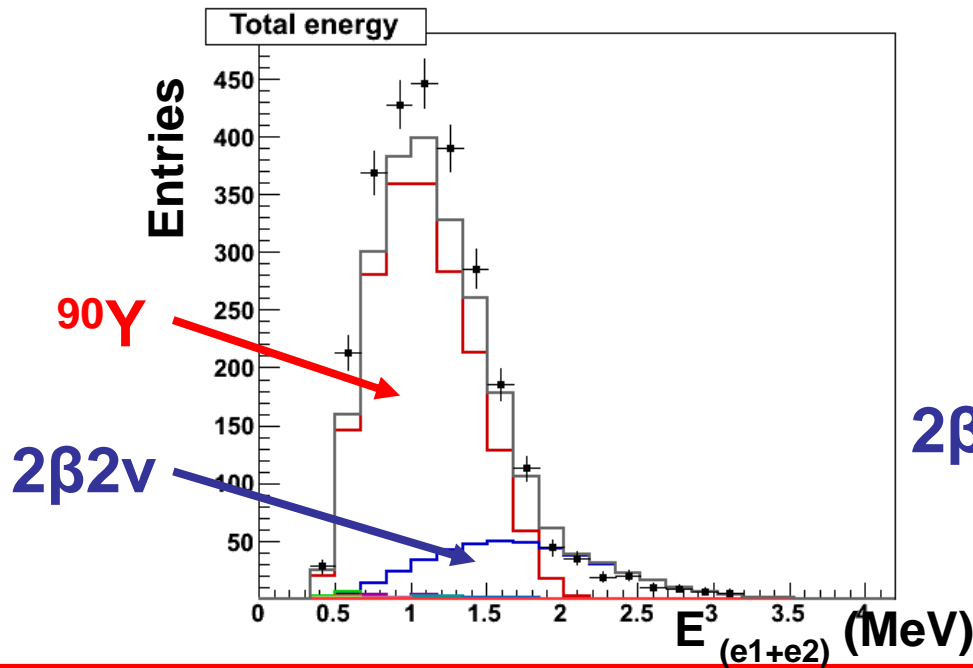


- ^{90}Y is the dominant background $Q_\beta = 2.2\text{MeV}$.

- Comes from ^{90}Sr decay half-life 28.8 years

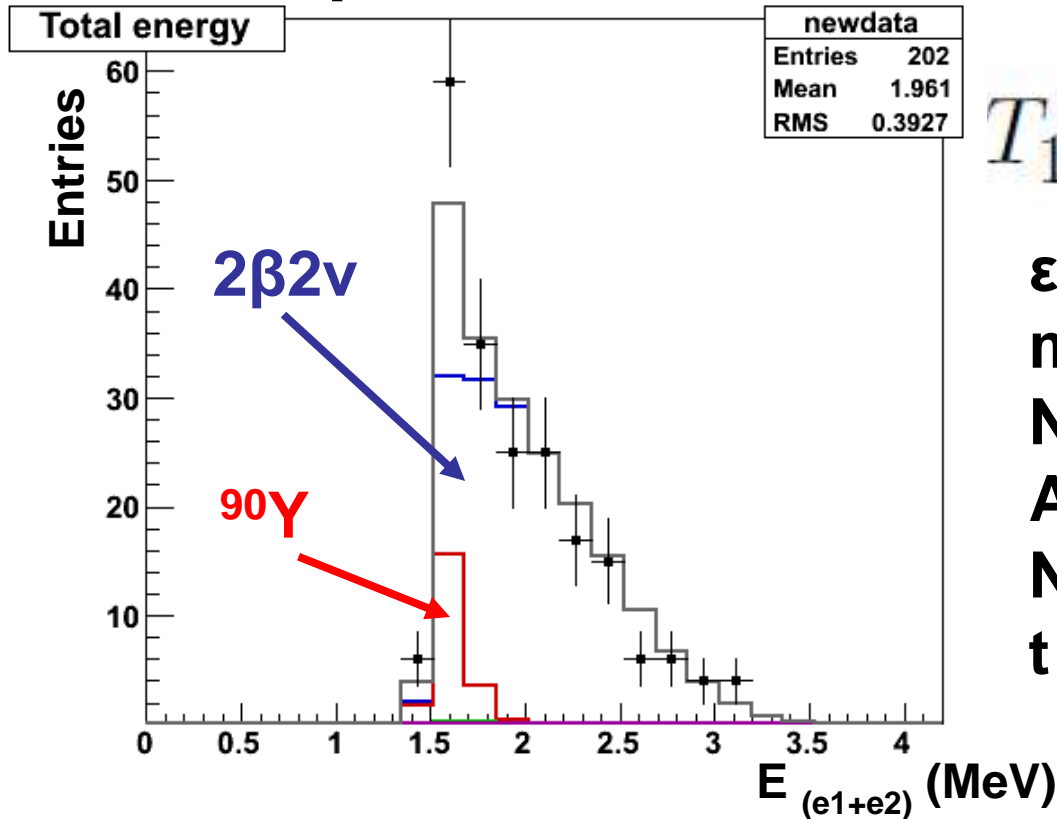
- Evidence of a **missing background** in the 1e1y channel and 2e

^{48}Ca $2e$ spectra



- Large peak on $\text{Cos}(\theta)$ because of ^{90}Y
- Missing background **below 1.5 MeV** consistent with $1e1\gamma$

2β2ν Half-life Measurement



$$T_{1/2} = \varepsilon \frac{m_i N_A}{A_r N_{dec}} \ln(2) t.$$

$$\varepsilon = 0.03$$

$$m_i = 6.99\text{g}$$

$$N_A = 6.022 \times 10^{23}$$

$$A_r = 48$$

$$N_{dec} = 180$$

$$t = 4.03 \text{ yr}$$

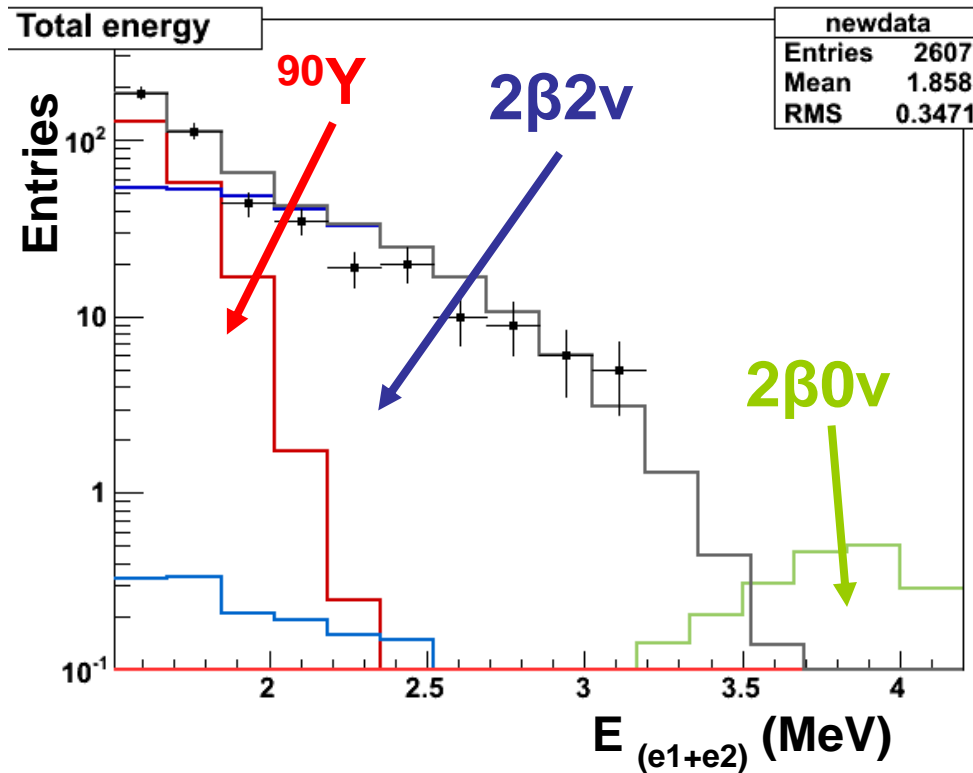
$$T_{1/2} = 4.09^{+0.26}_{-0.24} (\text{stat.}) \pm 0.29 (\text{syst.}) \times 10^{19} \text{ yr}$$

World's most precise measurement

Calculation of matrix element $M^{2\nu} = 0.0238 \pm 0.0015$

Theoretical calculations = 0.026 & 0.028

Neutrinoless limit



$$T_{1/2} = \varepsilon \frac{m_i N_A}{A_r N_{dec}} \ln(2) t.$$

$$\varepsilon = 0.189$$

$$m_i = 6.99 \text{g}$$

$$N_A = 6.022 \times 10^{23}$$

$$A_r = 48$$

$$N_{dec} = 2.30$$

$$t = 4.25 \text{ yr}$$

Energy window [3.36MeV – 4.2MeV] Chosen by optimising signal/background

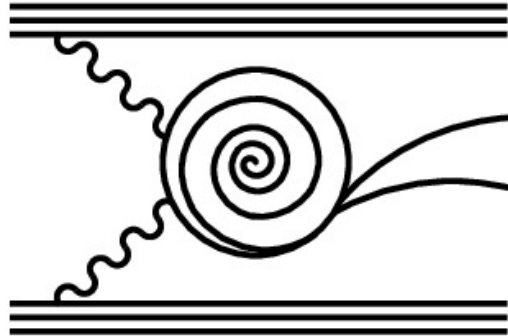
$$T_{1/2} > 2.12 \times 10^{22} \text{ yr at 90\% CL}$$

Compared to our previous result, 1.3×10^{22} yr

But latest ELEGANT-VI result is 5.8×10^{22} yr PRC (2008)

Therefore a mass limit of $m_{0\nu} < 18 \text{ eV}$

supernemo



collaboration



Calibration Software Development for SuperNEMO

Detector description

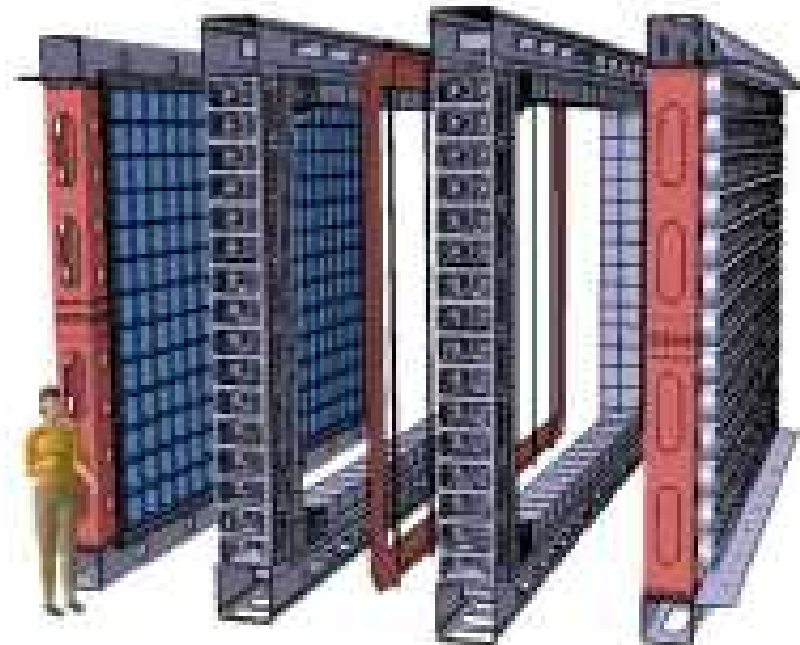
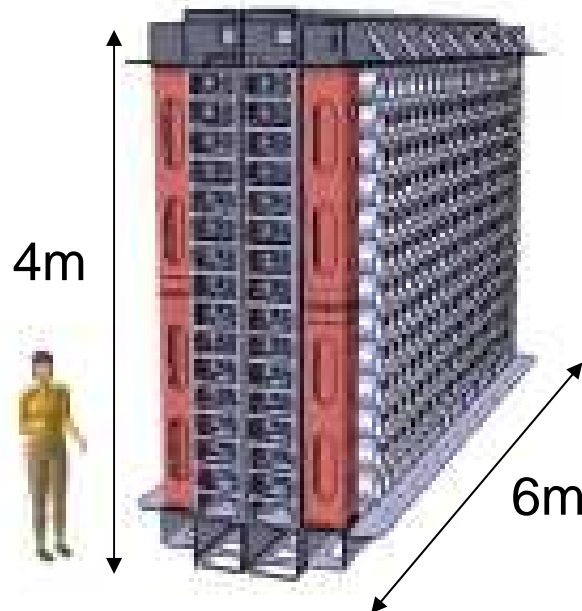
SuperNEMO

- Geiger drift cell tracker (improved eff)
- Plastic Bc404 calorimeter
- **100kg** of ^{82}Se source
- Modular design (~ 20 modules)
- **50-80 meV** sensitivity
- Energy resolution < **7% FWHM** at 1Mev

Nemo3

polystyrene
10kg (multi source)

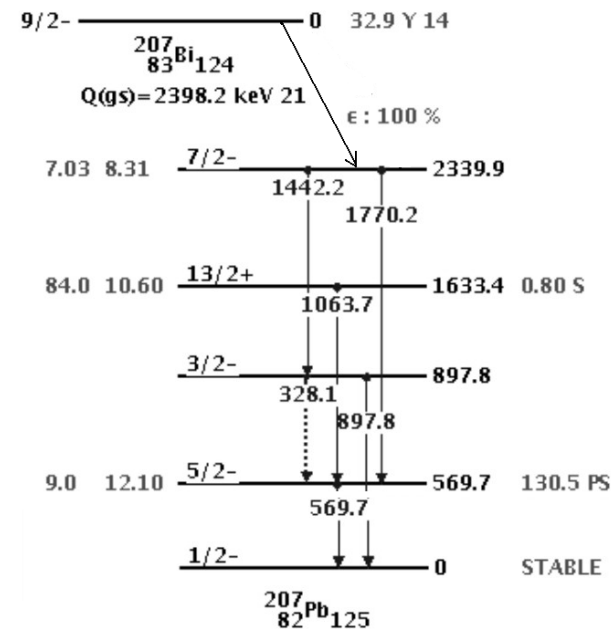
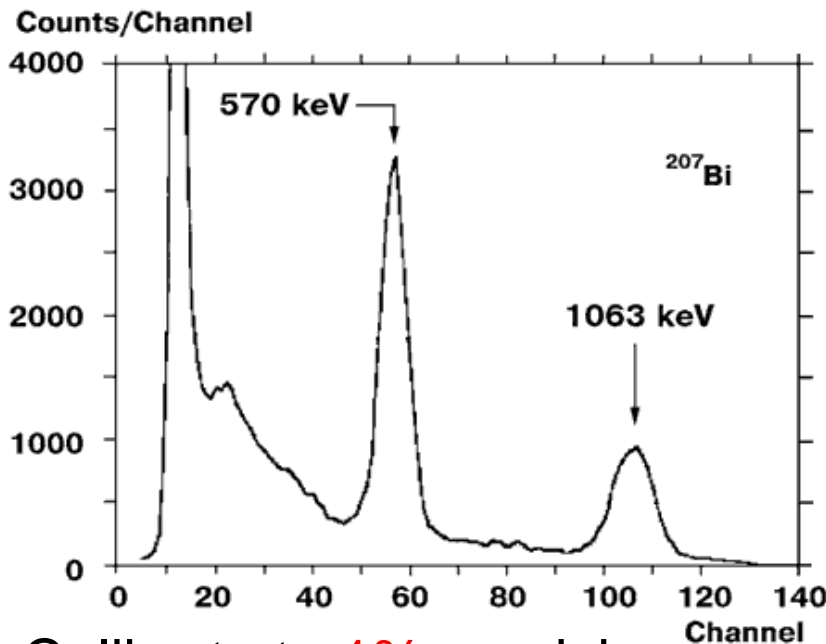
0.3eV
14%



Energy Calibration Constraints

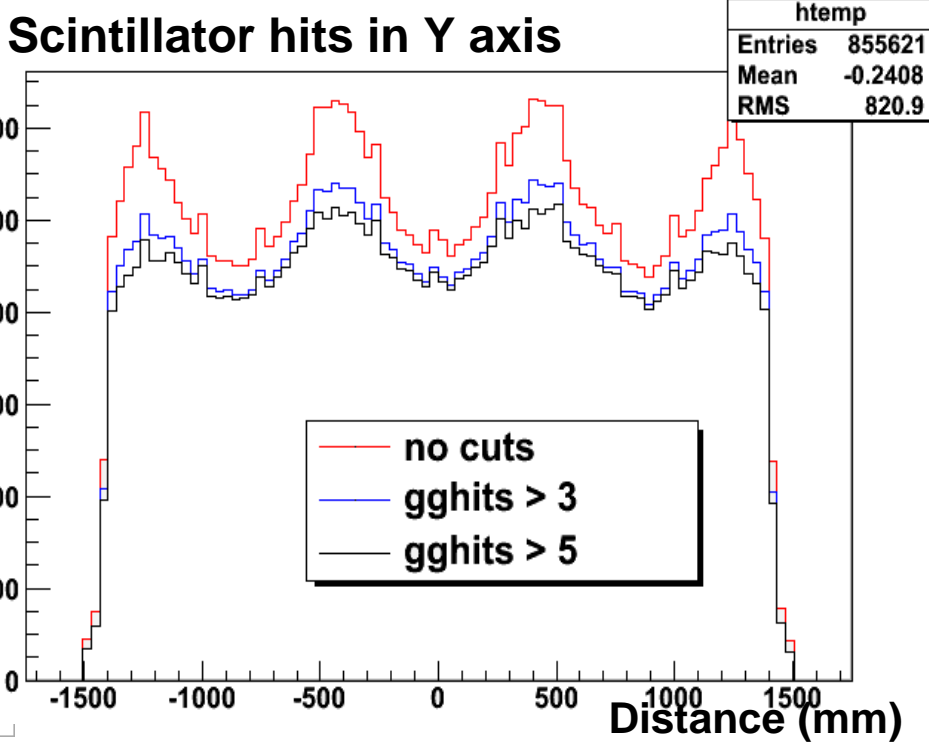
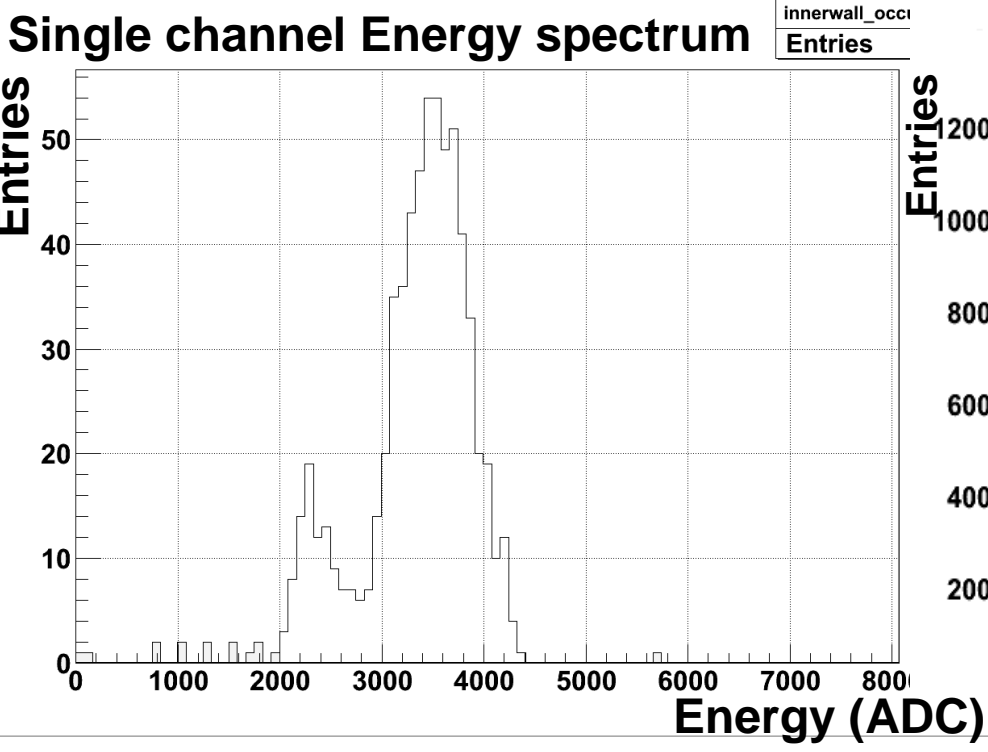
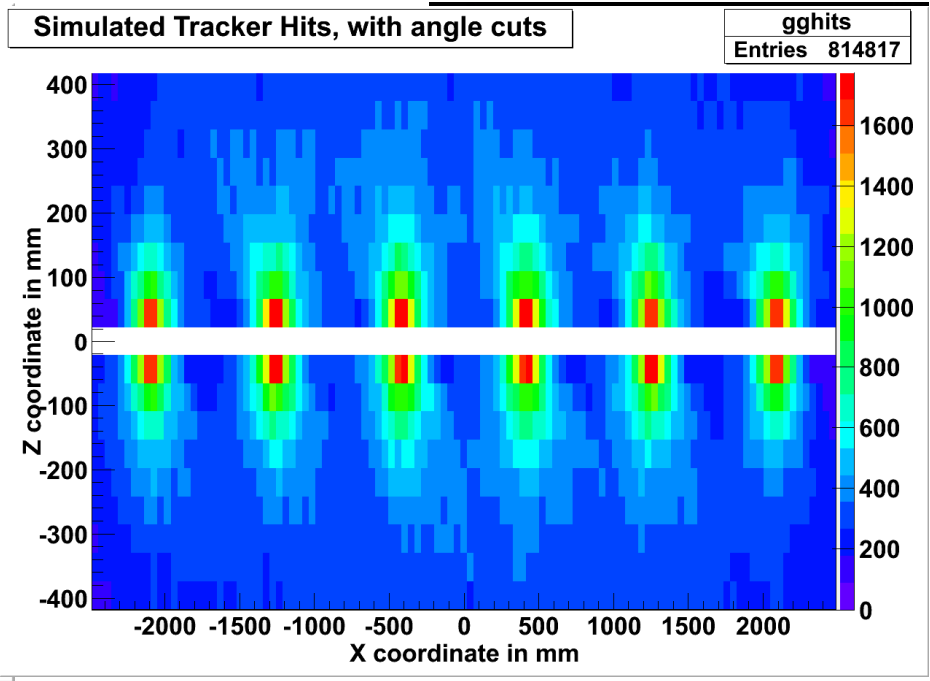
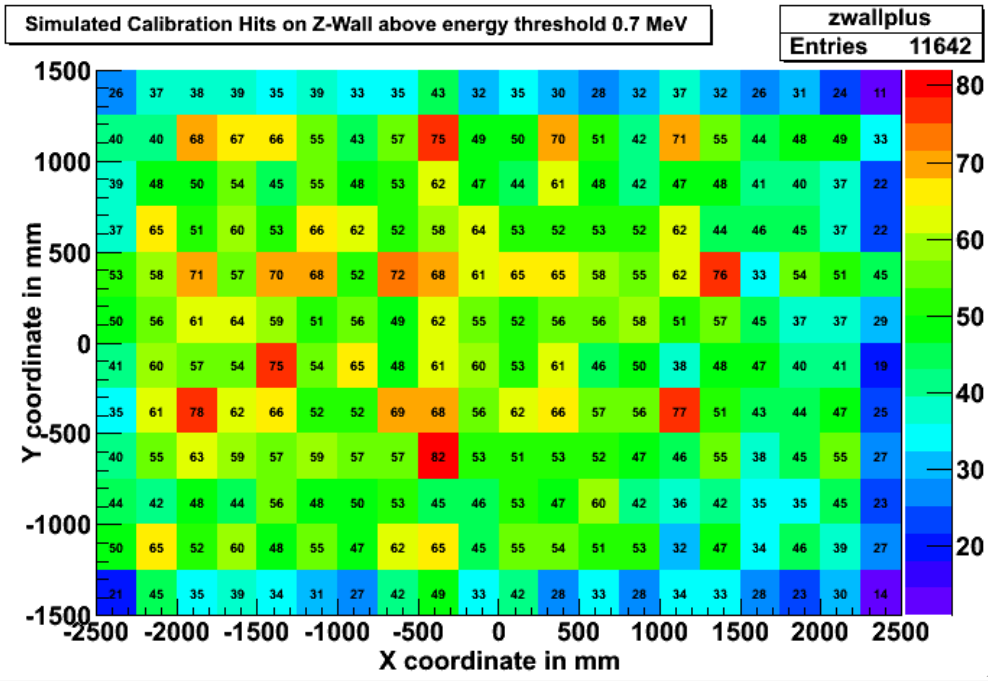
Calibration of ADC with energy scale needed periodically

^{207}Bi used as a conversion electron source of **~ 1 MeV** electrons

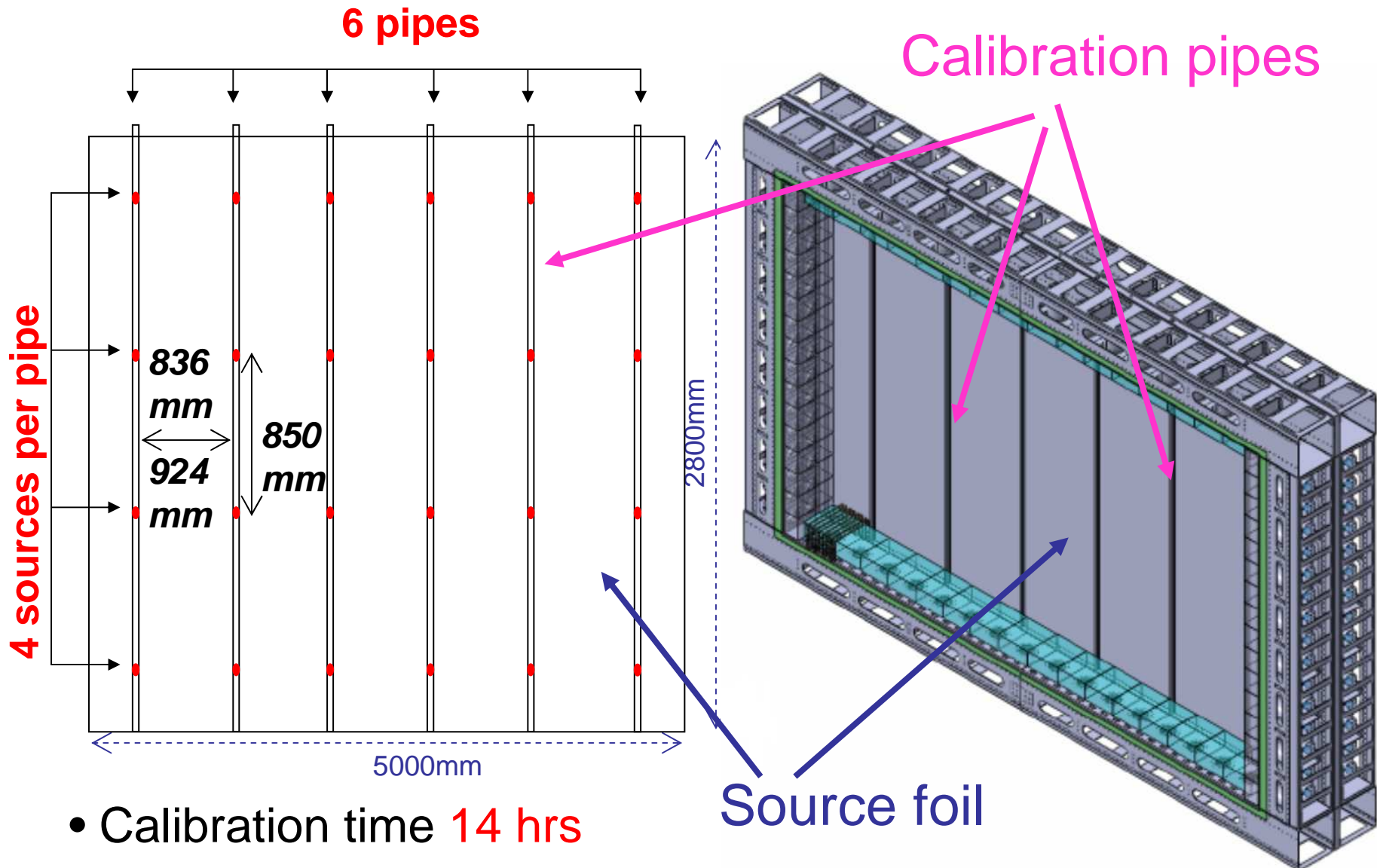


- Calibrate to **1%** precision
- Enough hits in each channel to reconstruct **1 MeV** peak
- Smallest area possible

- **Little downtime** as possible
- Without going over maximum₁₆ Geiger rate



Physical Model



- Calibration time **14 hrs**
- Geiger Rate **< 9 Hz**
- 6 x 4 configuration

Future Work

^{48}Ca Analysis:

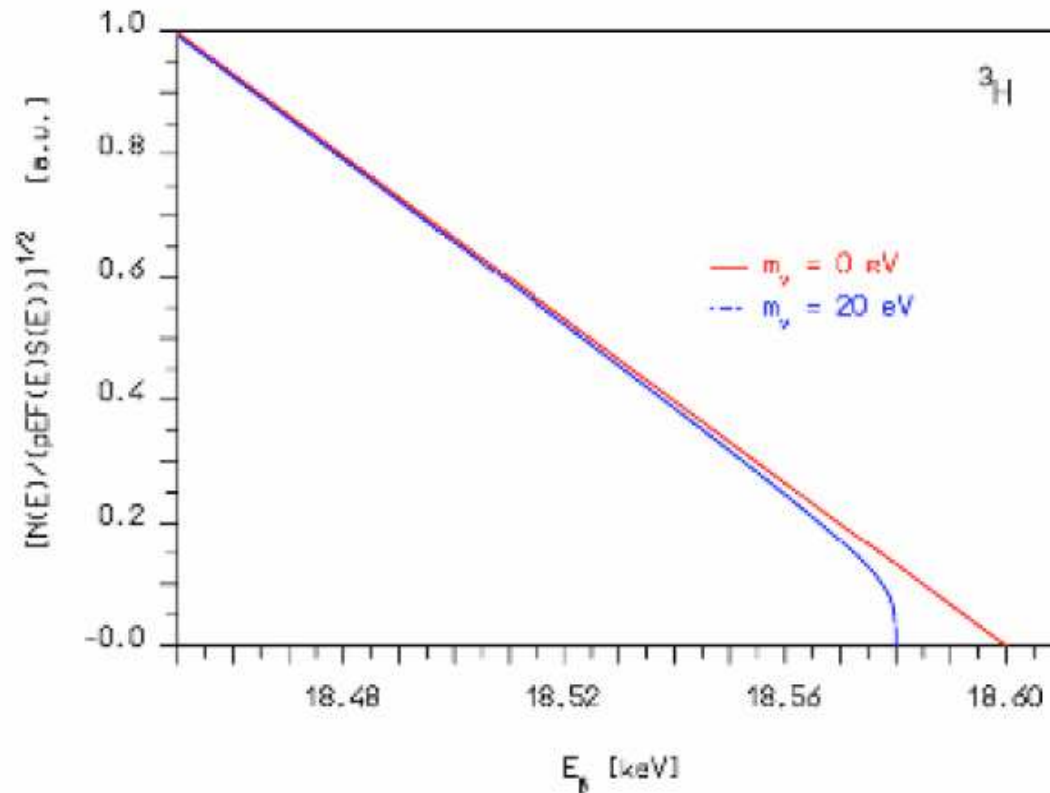
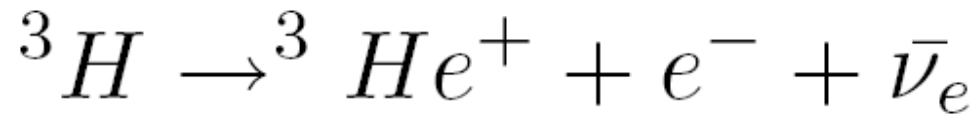
- Disentangle low energy background
- Consolidate all data for better precision $2\beta 2\nu$ result (**factor of 2**)
- Improve on $2\beta 0\nu$ limit

Software Development & Calibration:

- Extend simulation for ^{90}Sr (third calib point at 2.2MeV) and ^{60}Co (coincident gammas 1.1MeV for timing)
- Core software development
- Commissioning the tracker module

Extra Slides

Tritium β Decay

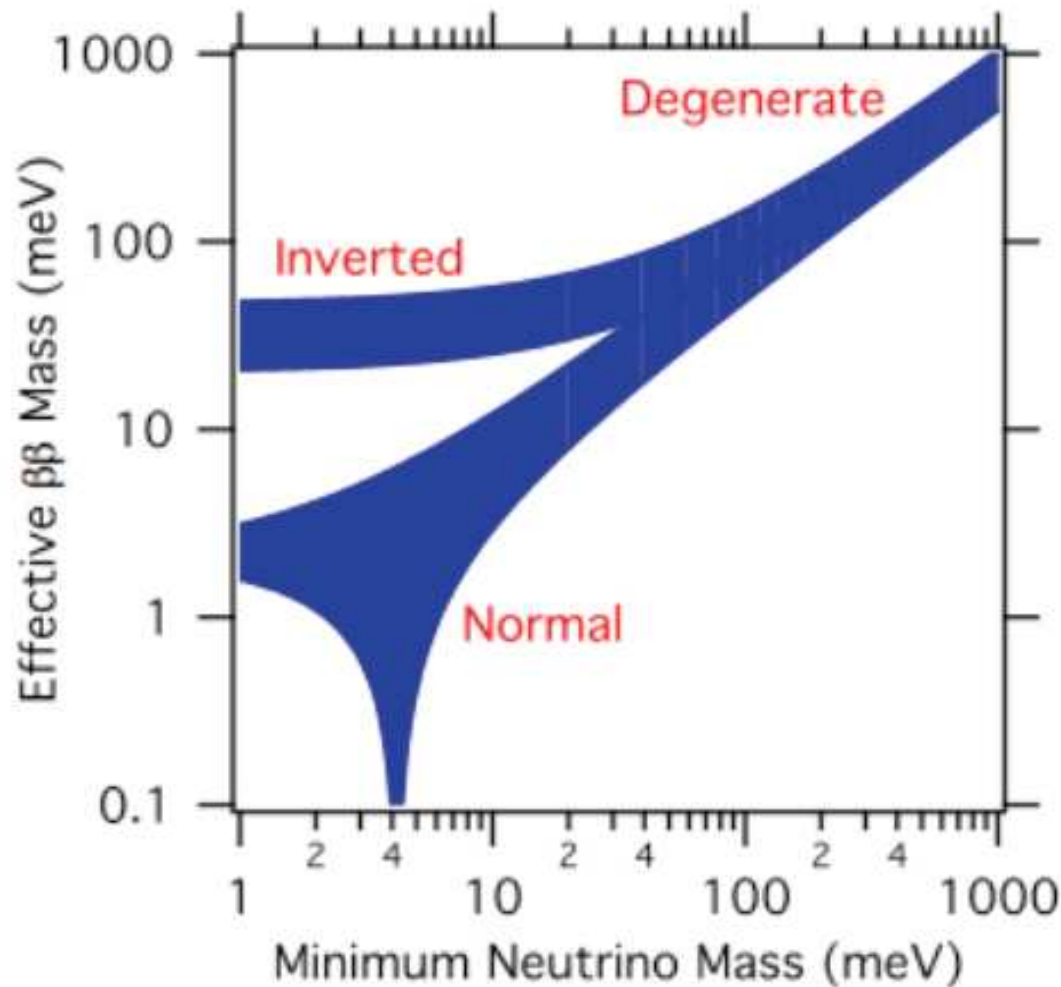


$m_{\bar{\nu}_e} < 2.2 \text{ eV}$
At 2σ

Super allowed so the matrix element does not affect the shape of the spectrum

$$K(E) = \sqrt{\frac{N(E)}{Fp_e E}} \propto (E_0 - E) \left[1 - \left(\frac{m_{\bar{\nu}_e} c^2}{E_0 - E} \right)^2 \right]^{1/4}$$

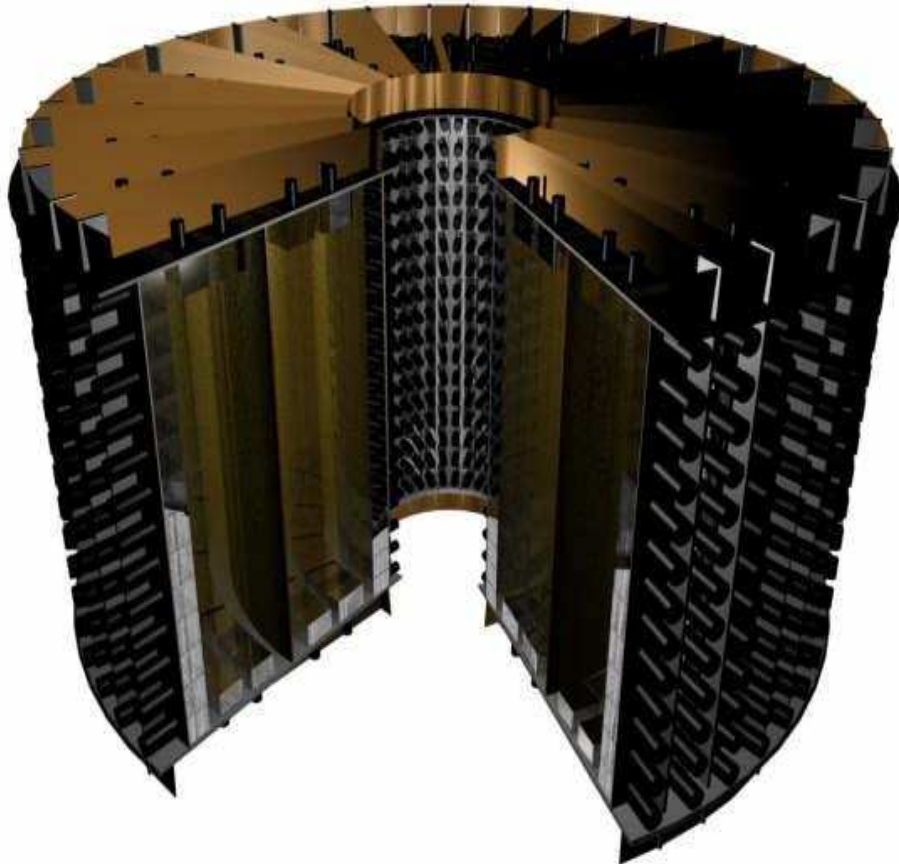
Evidence for hierarchy



$$m_{ee} \equiv \sum_i U_{ei}^2 m_i \equiv m_1 |U_{e1}|^2 e^{i\alpha_1} + m_2 |U_{e2}|^2 e^{i\alpha_2} + m_3 |U_{e3}|^2 e^{-2i\delta}$$

NEMO3 Detector

- Tracker + Calorimeter
- 10Kg of source
- Magnetic field
- Passive shielding
- 4800 MWE
- 0.3 eV sensitivity
- Energy resolution ~ 14% FWHM at 1MeV

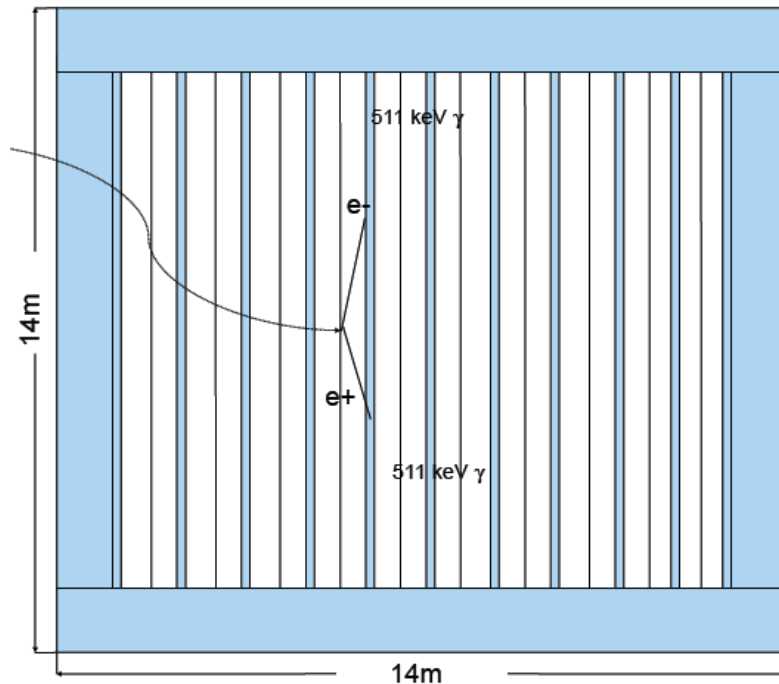


- Divided into 20 individual segments
- Source foils 3.1m in diameter, 2.5m high, 30-60mg/cm² thick and of different isotopes
- Either side of the foil are concentric tracking volumes composed of 6180 drift cells
- The drift cells are 270 cm long and contain an ethyl alcohol and helium gas mixture
- Surrounding this is the calorimeter made up of 1940 polystyrene scintillator blocks of 10cm
- A magnetic field parallel to the foil access of 25G is generated by an external solenoid.
- There is an external shield of 20cm iron that reduces gamma rays and thermal neutrons
- Outside this there is a borated water shield thermalize fast neutrons and capture thermal neutrons

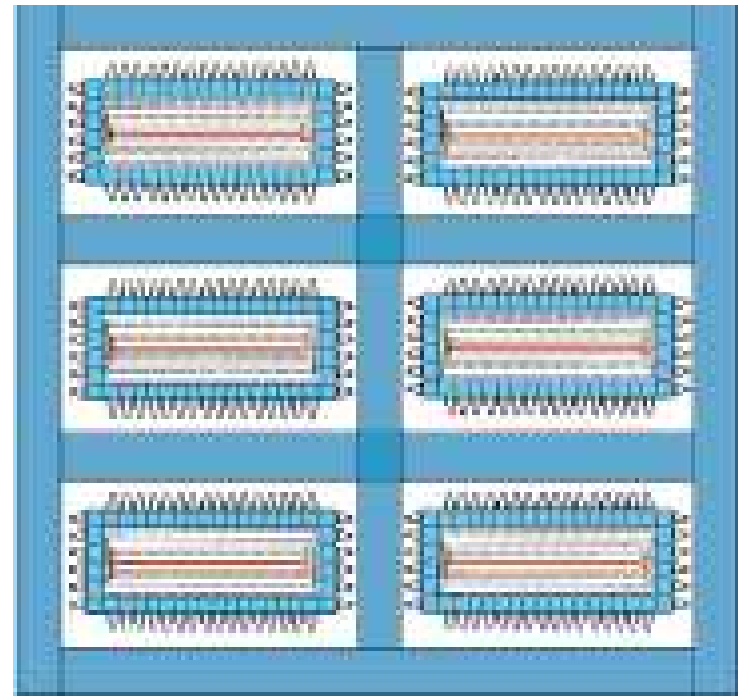
SuperNEMO

- Tracker + Calorimeter
- 100Kg of source
- modular design (~ 20 modules)
- 4800 MWE
- 50-80 meV sensitivity
- Energy resolution < 7% FWHM at 1Mev

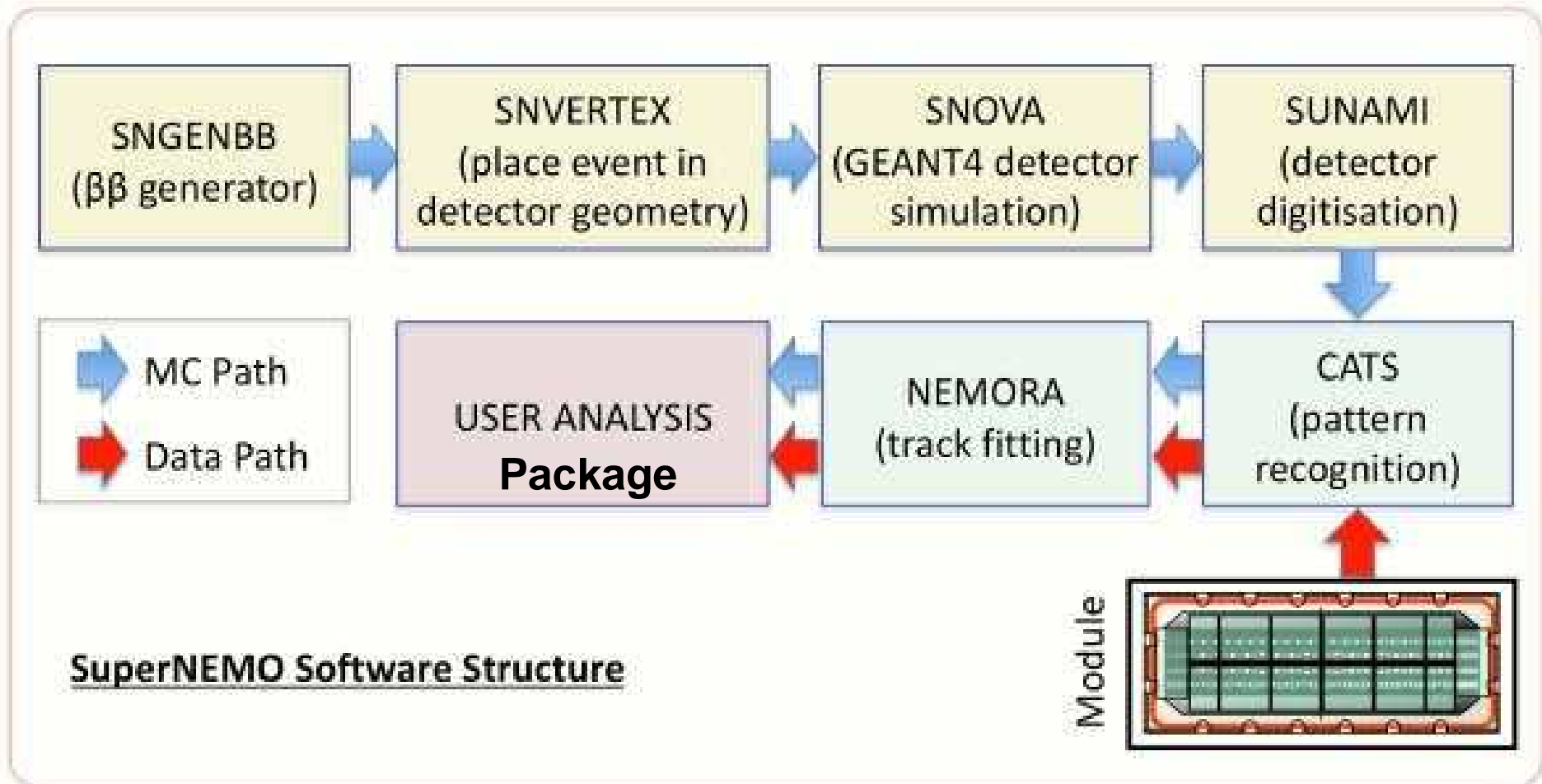
Bar design



Block design



Software description



Analysis Package

Provides:

- Full and easy access to both reconstructed, true and raw data
- Easy to use inbuilt data structure and root io
- Framework for applying cuts, event selection and plots
- Inbuilt analysis debugging
- Tools and functions for analysis

Simulation Results

n_T	n_S	x_T (mm)	y_T (mm)	No of events simulated	C_{\min} ($E > 0.7$ MeV)	T_{\max}	Cal time (hrs)	Tracke r max hit rate (Hz)*	Possible $\beta\beta$ Foil Sheet width (mm)
5	4	1012	850	10^6	5	2200	27.8	8.8	
6	4	748	850	10^6	5	1921	23.1	9.2	
6	4	836	850	9.5×10^5	8	1771	13.7	8.9	200
6	4	924	850	10^6	8	1840	14.5	8.8	180
6	4	924	900	10^6	6	1897	19.3	9.1	

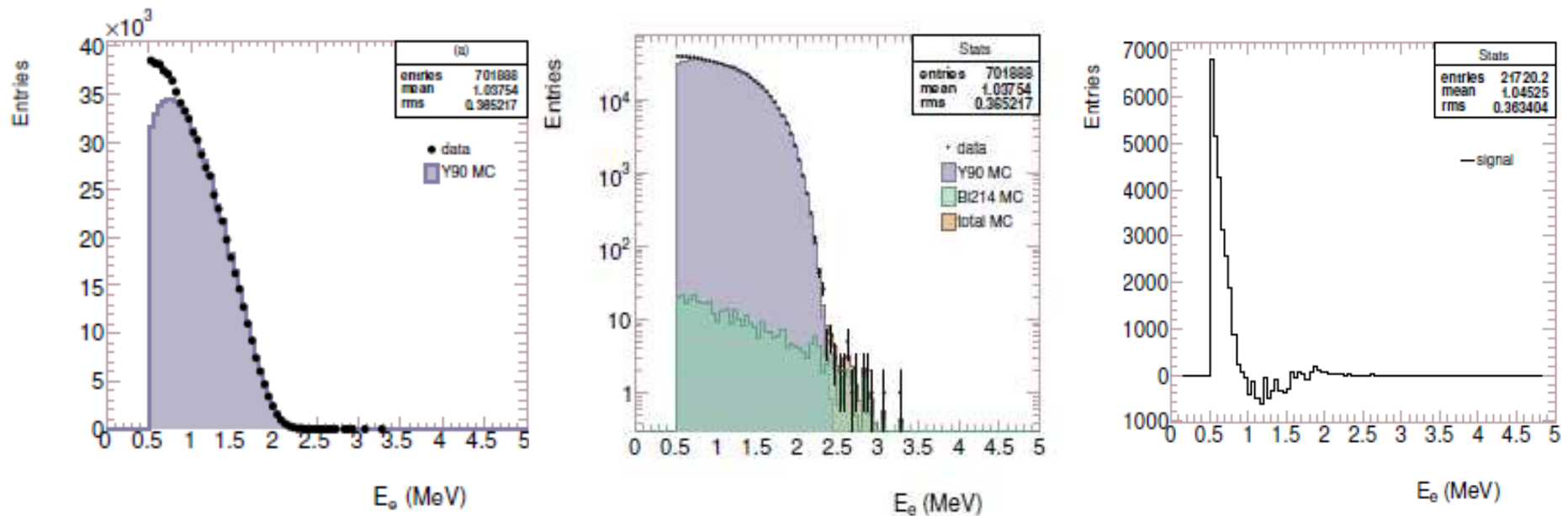
Backgrounds

- For time being using existing Internal background measurements
- Sources :
- HPGE measurements
- Nemo3 1e and 1e1y measurements

Background	Activity (Bq)
^{228}Ac	1.46×10^{-5}
^{212}Bi	1.46×10^{-5}
^{214}Bi	4.45×10^{-6}
$^{137}\text{Cs}^*$	4.95×10^{-5}
$^{152}\text{Eu}^*$	5.25×10^{-4}
^{40}K	8.00×10^{-4}
^{214}Pb	4.60×10^{-6}
^{208}Tl	5.25×10^{-6}
^{90}Y	0.03

The Y90 background

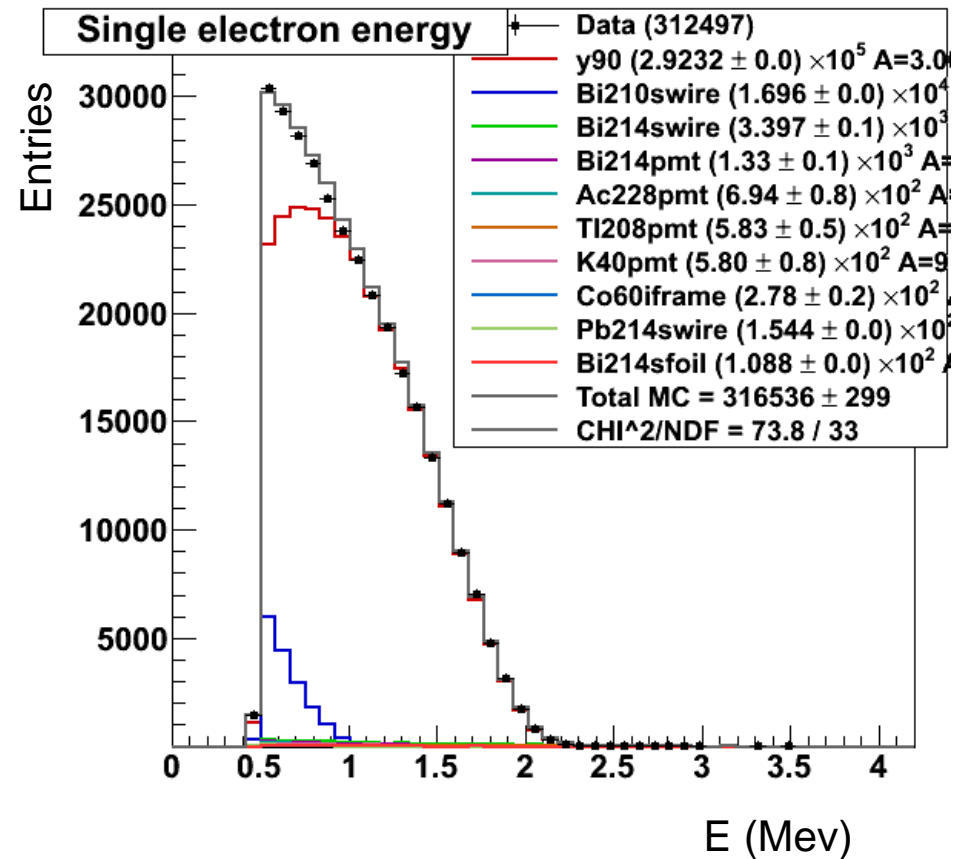
- The Y90 background component dominates
- It was measured using the 1e spectrum



- High energy deficit was fit by Bi214 Swire & Sfoil
- Low energy deficit not determined

Y90 Background

- Bi210 is a good fit for the low energy
- With an activity of 5.05Bq
- Can be out of equilibrium with Bi214 due to Pb210



Cuts and Efficiencies

Cuts used background efficiencies

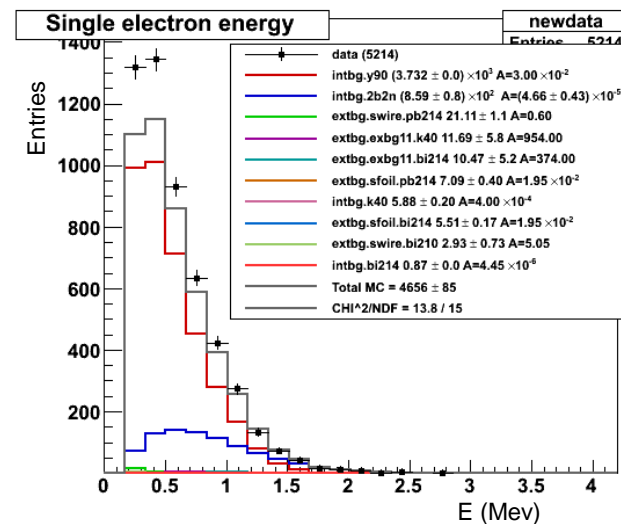
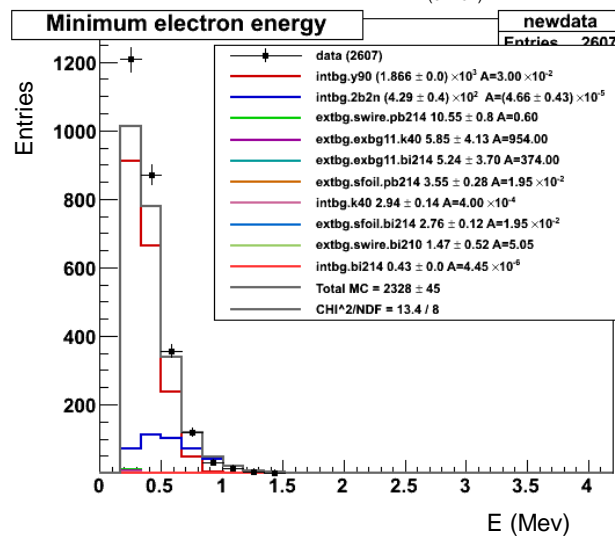
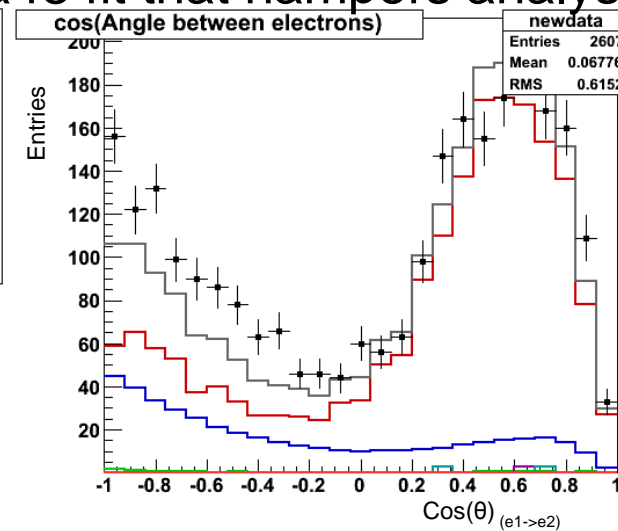
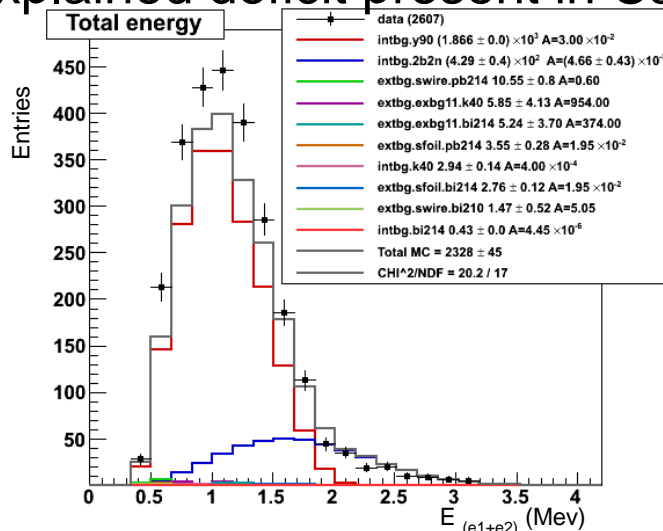
Largest

Scintillator hits = 2
 Tracks = 2
 Associated tracks
 Track Charge < 0
 Reconstructed correctly
 $\Delta XY < 2\text{cm}$
 $\Delta Z < 7\text{cm}$
 Reconstructed to vertex on Ca foil
 Triggered blocks below threshold
 Scintillator hits > 200keV
 Track hits in first 2 layers near foil
 Track hits in last 2 layers near blocks
 Track length > 60 cm
 TOF $0.04 < \text{pint} < 1$
 TOF $0 < \text{pext} < 0.01$
 Isolated scintillator hits
 Delayed alpha single hit > $70\mu\text{s}$
 Delayed alpha grouped hit > $20\mu\text{s}$
 Cut on no associated fast hits
 Only one gamma > 200keV
 No gammas $150\text{keV} < E < 200\text{keV}$

<u>Background</u>	<u>Efficiency</u>
Bi214 sfoil	1.08E-006
Pb214 sfoil	1.34E-006
Bi210 swire	1.15E-009
Bi214 swire	1.00E-006
Pb214 swire	1.41E-007
iron	2.48E-010
Ac228 intbgr	1.98E-004
Bi214 intbgr	7.57E-004
Eu152 intbgr	1.08E-005
Tl208 intbgr	3.92E-004
Y90 intbgr	4.79E-004 ₃₁

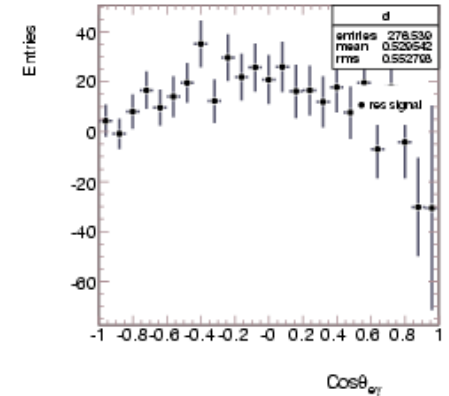
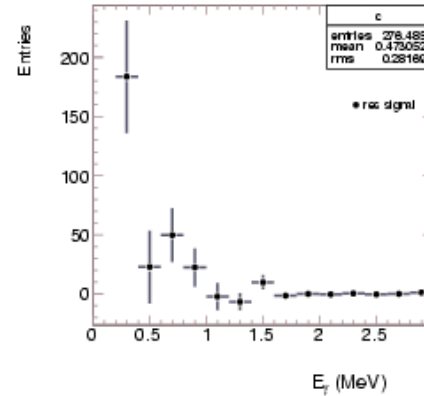
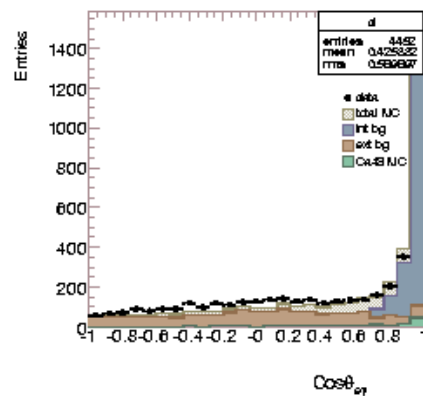
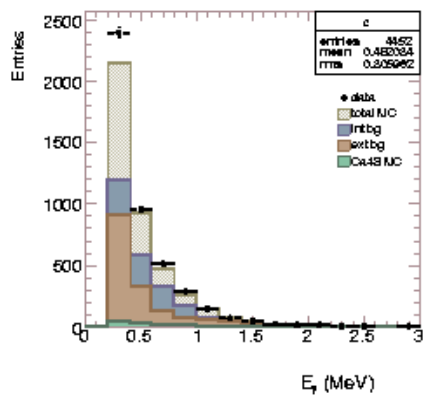
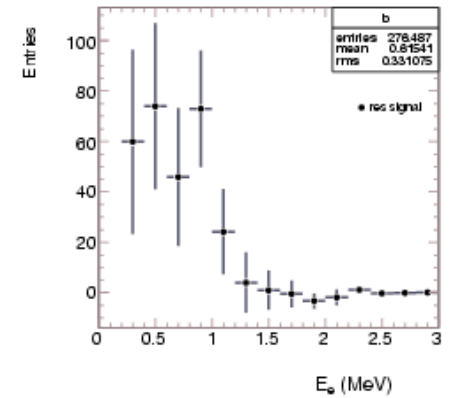
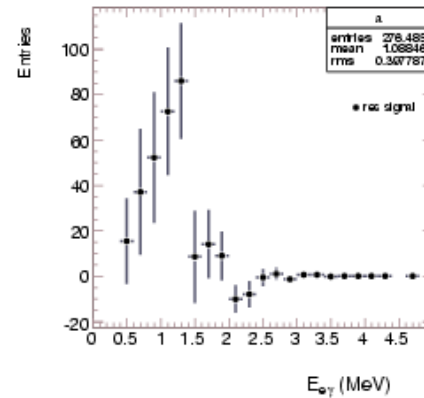
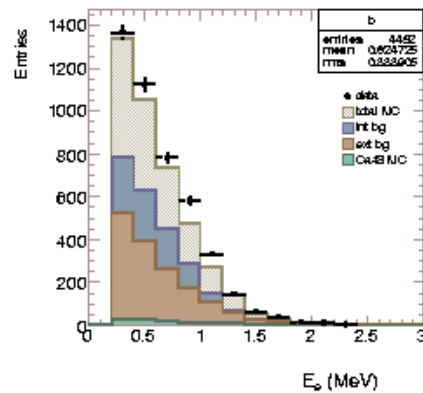
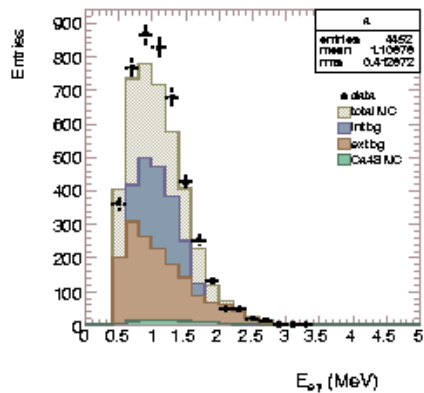
Issues with the Ca48 spectra

- The dominant background is of Y90. However, there is an apparent unexplained deficit present in Ca48 fit that hampers analysis.

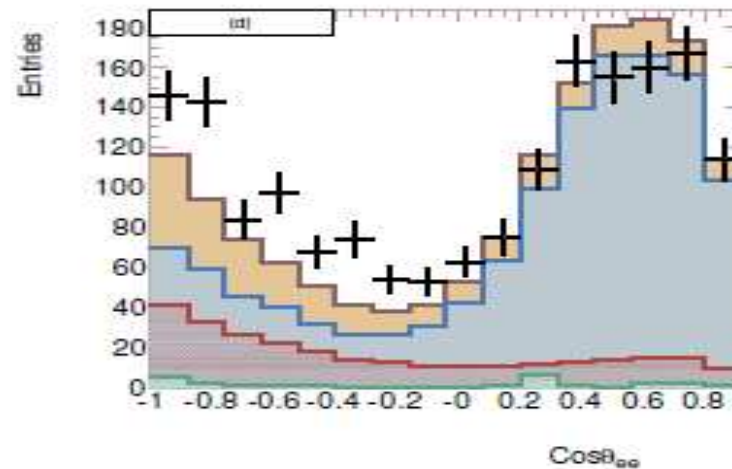


Issues with the Ca48 spectra

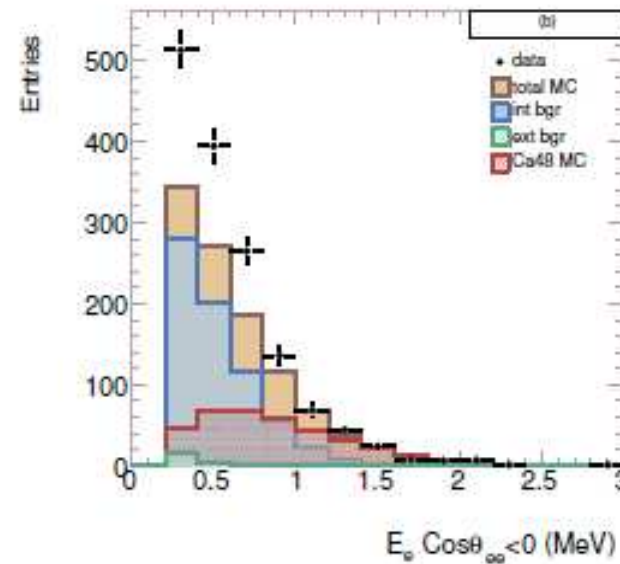
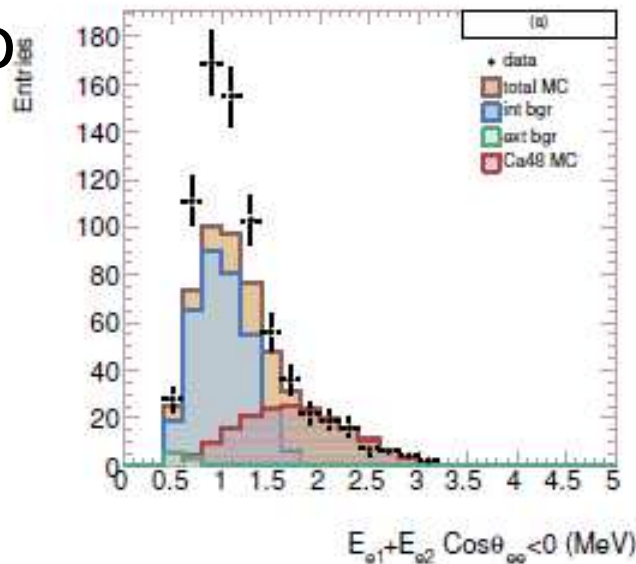
- Deficit also present in the $1e1\gamma$ spectra
- Backgrounds are well understood for $E_{(e1+e2)} > 1.5\text{MeV}$



Issues with the Ca48 spectra



- Cut $\cos(\theta) > 0$ to increase signal to background ratio



Results

