Coannihilation and Scotogenic Fermionic Dark Matter

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Motivation

- Neutrino mass generation and Dark matter are two important unsolved issues in Particle Physics.
- Cosmological bound of $\sum_{i} m_{\nu_i} < 0.09$ eV [Di Valentino et al. 2106.15267] makes usual seesaw schemes (I & III) hardly accessible to collider experiments.
- In the case of scotogenic models neutrino masses are generated radiatively and hence loop-suppressed; additionally, neutrino masses are symmetry-protected.
- Dark mediated neutrino mass generation is a very interesting idea. Imposition
 of Z₂ symmetry stabilizes the DM. [Ma hep-ph/0601225, Tao hep-ph/9603309]
- Singlet-triplet scotogenic model is one such model. [Hirsch et al. 1307.8134]
- The DM in this model can be bosonic as well as fermionic.
- The scalar DM for this model has been studied in great detail. It resembles the inert-2HDM scenario. [Diaz et al. 1612.06569, Avila et al. 1910.08422]
- Triplet-like fermionic DM in this model is phenomenologically not very interesting; one can go up to mass of 2.5 TeV only.
- Singlet-like fermionic DM is more rich in phenomenology.

Valle, Rojas, Hirsch, Vicente, Restrepo, De Romari, ... 1307.8134, 1603.05685, 1605.01915, 1612.06569, 1907.11938, 1910.08422 ...

Singlet-Triplet Model

Particle Content:

	Sta	ndard Mo	odel	New S	Scalars	New Fermions			
	L	e ^c	Φ	Ω	η	Σ	F		
Multiplicity	3	3	1	1	1	1	1		
U(1) _Y	-1/2	1	1/2	0	1/2	0	0		
<i>SU</i> (2) _L	2	1	2	3	2	3	1		
\mathcal{Z}_2	+	+	+	+	_	-	-		

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Yukawa:

$$\mathcal{L} = Y^{\alpha\beta}\bar{L}_{\alpha}\Phi e_{\beta} + \boxed{Y_{F}^{\alpha}\bar{L}_{\alpha}\tilde{\eta}F^{c} + Y_{\Sigma}^{\alpha}\bar{L}_{\alpha}\Sigma^{c}\tilde{\eta}} + \boxed{Y_{\Omega}Tr[\bar{\Sigma}\Omega]F^{c}} + \boxed{\frac{M_{\Sigma}}{2}Tr[\bar{\Sigma}\Sigma^{c}] + \frac{M_{F}}{2}\bar{F}F^{c}} + \text{h.c.}$$

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Scalar Potential:

$$= -\mu_{\phi}^{2} \left(\Phi^{\dagger} \Phi \right) + \mu_{\eta}^{2} \left(\eta^{\dagger} \eta \right) - \frac{1}{2} \mu_{\Omega}^{2} \operatorname{Tr}(\Omega^{\dagger} \Omega) + \mu_{\Omega}^{\phi} \left(\Phi^{\dagger} \Omega \Phi \right) + \mu_{\Omega}^{\eta} \left(\eta^{\dagger} \Omega \eta \right) \right)$$

$$+ \left[\frac{1}{2} \lambda_{1} (\Phi^{\dagger} \Phi)^{2} + \frac{1}{2} \lambda_{2} (\eta^{\dagger} \eta)^{2} + \lambda_{3} (\Phi^{\dagger} \Phi) (\eta^{\dagger} \eta) + \lambda_{4} (\Phi^{\dagger} \eta) (\eta^{\dagger} \Phi) + \frac{1}{2} \lambda_{5} \left[(\Phi^{\dagger} \eta)^{2} + (\eta^{\dagger} \Phi)^{2} \right] \right]$$

$$+ \left[\frac{1}{2} \lambda_{\Omega}^{\phi} (\Phi^{\dagger} \Phi) \operatorname{Tr}(\Omega^{\dagger} \Omega) + \frac{1}{4} \lambda_{\Omega}^{\Omega} \left[\operatorname{Tr}(\Omega^{\dagger} \Omega) \right]^{2} + \frac{1}{2} \lambda_{\Omega}^{\eta} (\eta^{\dagger} \eta) \operatorname{Tr}(\Omega^{\dagger} \Omega) \right]$$

Fields:

$$\Phi = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2} \phi^+ \\ v_{\phi} + \phi^0 + i G^0 \end{pmatrix}, \quad \Omega = \begin{pmatrix} \frac{v_{\Omega}}{\sqrt{2}} + \frac{\Omega^0}{\sqrt{2}} & \Omega^+ \\ \Omega^- & -\frac{v_{\Omega}}{\sqrt{2}} - \frac{\Omega^0}{\sqrt{2}} \end{pmatrix}, \quad \eta = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2} \eta^+ \\ \eta_R^0 + i \eta_I^0 \end{pmatrix}$$

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Neutral scalars:

$$\{\phi^{0},\Omega^{0}\} \longrightarrow \{h,H\}: \quad \mathcal{M}_{0}^{2} = \begin{pmatrix} \lambda_{1}v_{\phi}^{2} & \lambda_{\Omega}^{\phi}v_{\Omega}v_{\phi} - \frac{1}{\sqrt{2}}\,\mu_{\Omega}^{\phi}\,v_{\phi} \\ \lambda_{\Omega}^{\phi}v_{\Omega}v_{\phi} - \frac{1}{\sqrt{2}}\,\mu_{\Omega}^{\phi}\,v_{\phi} & 2\lambda_{\Omega}^{\Omega}v_{\Omega}^{2} + \frac{1}{2\sqrt{2}}\,\mu_{\Omega}^{\phi}\,\frac{v_{\phi}^{2}}{v_{\Omega}} \end{pmatrix}$$

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Charged scalars:

$$\{\phi^{\pm}, \Omega^{\pm}\} \rightarrow \{G^{\pm}, H^{\pm}\}: \mathcal{M}_{\pm}^{2} = \sqrt{2}\,\mu_{\Omega}^{\phi} \begin{pmatrix} v_{\Omega} & \frac{1}{2}\,v_{\phi} \\ \frac{1}{2}\,v_{\phi} & \frac{1}{4}\,\frac{v_{\phi}^{2}}{v_{\Omega}} \end{pmatrix} + g^{2}\,\xi_{W^{\pm}} \begin{pmatrix} \frac{1}{4}\,v_{\phi}^{2} & -\frac{1}{2}\,v_{\phi}\,v_{\Omega} \\ -\frac{1}{2}\,v_{\phi}\,v_{\Omega} & v_{\Omega}^{2} \end{pmatrix}$$

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\mathcal{Z}_2 -odd scalars:

$$m_{\eta_R^0}^2 = \mu_{\eta}^2 + \frac{1}{2}\lambda_{\Omega}^{\eta}v_{\Omega}^2 - \frac{1}{\sqrt{2}}\mu_{\Omega}^{\eta}v_{\Omega} + \frac{1}{2}(\lambda_3 + \lambda_4 + \lambda_5)v_{\phi}^2$$
$$m_{\eta_I^0}^2 = \mu_{\eta}^2 + \frac{1}{2}\lambda_{\Omega}^{\eta}v_{\Omega}^2 - \frac{1}{\sqrt{2}}\mu_{\Omega}^{\eta}v_{\Omega} + \frac{1}{2}(\lambda_3 + \lambda_4 - \lambda_5)v_{\phi}^2$$
$$m_{\eta^{\pm}}^2 = \mu_{\eta}^2 + \frac{1}{2}\lambda_{\Omega}^{\eta}v_{\Omega}^2 + \frac{1}{\sqrt{2}}\mu_{\Omega}^{\eta}v_{\Omega} + \frac{1}{2}\lambda_3v_{\phi}^2$$

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$$\Phi = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2} \phi^+ \\ v_{\phi} + \phi^0 + i G^0 \end{pmatrix}, \quad \Omega = \begin{pmatrix} \frac{v_{\Omega}}{\sqrt{2}} + \frac{\Omega^0}{\sqrt{2}} & \Omega^+ \\ \Omega^- & -\frac{v_{\Omega}}{\sqrt{2}} - \frac{\Omega^0}{\sqrt{2}} \end{pmatrix}, \quad \eta = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2} \eta^+ \\ \eta_R^0 + i \eta_I^0 \end{pmatrix}$$

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$$egin{aligned} & \lim \lambda_5 o 0 \ \implies & m_{\eta^0_R} = m_{\eta^0_I} \end{aligned}$$

4

Fermion mixing & Neutrino mass

Fermion mixing:

$$\{F, \Sigma^{0}\} \rightarrow \{\chi_{1}^{0}, \chi_{2}^{0}\}: \quad \mathcal{M}_{\chi} = \begin{pmatrix} M_{F} & Y_{\Omega}v_{\Omega} \\ Y_{\Omega}v_{\Omega} & M_{\Sigma} \end{pmatrix} \implies V \cdot \mathcal{M}_{\chi} \cdot V^{T} = \operatorname{diag}(m_{\chi_{1}^{0}}, m_{\chi_{2}^{0}})$$
$$\boxed{\operatorname{tan}(2\theta) = \frac{2 Y_{\Omega}v_{\Omega}}{M_{\Sigma} - M_{F}}} \qquad \boxed{m_{\chi_{1,2}^{0}} = \frac{1}{2} \left[(M_{\Sigma} + M_{F}) \mp \sqrt{(M_{\Sigma} - M_{F})^{2} + 4Y_{\Omega}^{2}v_{\Omega}^{2}} \right]}$$

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Neutrino mass:



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Neutrino mass:

Neutrino mass matrix: $\mathcal{M}_{\nu} = Y_{\nu} \cdot \mathcal{F} \cdot Y_{\nu}^{T}$

$$\begin{split} \mathbf{Y}_{\nu} &= \begin{pmatrix} \mathbf{Y}_{E}^{1} & \mathbf{Y}_{Z}^{1} / \sqrt{2} \\ \mathbf{Y}_{F}^{2} & \mathbf{Y}_{Z}^{2} / \sqrt{2} \\ \mathbf{Y}_{F}^{2} & \mathbf{Y}_{Z}^{2} / \sqrt{2} \end{pmatrix} \cdot \mathbf{V}^{T}(\theta), \qquad \mathcal{F} = \begin{pmatrix} \frac{\mathcal{I}_{1}}{32\pi^{2}} & 0 \\ 0 & \frac{\mathcal{I}_{2}}{32\pi^{2}} \end{pmatrix}, \\ \mathcal{I}_{j} &= m_{\chi_{j}^{0}} \begin{bmatrix} \frac{\ln \left(m_{\chi_{j}^{0}}^{2} / m_{\eta_{R}^{0}}^{2}\right)}{\left(m_{\chi_{j}^{0}}^{2} / m_{\eta_{R}^{0}}^{2}\right) - 1} - \frac{\ln \left(m_{\chi_{j}^{0}}^{2} / m_{\eta_{I}^{0}}^{2}\right)}{\left(m_{\chi_{j}^{0}}^{2} / m_{\eta_{I}^{0}}^{2}\right) - 1} \end{bmatrix} \\ \hline \mathbf{Y}_{\nu} &= \mathbf{U} \cdot (\widetilde{\mathcal{M}}_{\nu})^{1/2} \cdot \rho \cdot (\mathcal{F})^{-1/2} \qquad \text{with} \quad \mathcal{U}^{\dagger} \mathcal{M}_{\nu} \mathcal{U}^{*} = \widetilde{\mathcal{M}}_{\nu} \ , \\ \rho^{\mathrm{NO}} &= \begin{pmatrix} 0 & 0 \\ \cos \omega & \sin \omega \\ -\sin \omega & \cos \omega \end{pmatrix} \quad \text{and} \quad \rho^{\mathrm{IO}} = \begin{pmatrix} \cos \omega & \sin \omega \\ -\sin \omega & \cos \omega \\ 0 & 0 \end{pmatrix}. \end{split}$$



 $\lim \lambda_5 \to 0 \implies \mathcal{I}_j \propto \lambda_5 \ m_{\chi_j^0} \longrightarrow \text{Smaller } \lambda_5 \text{ and } m_{\chi_j^0} \implies \text{ bigger } Y_{F,\Sigma}$

5

Constraints

Bounded below: i) $\lambda_1 \ge 0$, ii) $\lambda_2 \ge 0$, iii) $\lambda_0^{\Omega} \ge 0$, iv) $\lambda_3 + \sqrt{\lambda_1 \lambda_2} \ge 0$, v) $\lambda_3 + \lambda_4 - |\lambda_5| + \sqrt{\lambda_1 \lambda_2} \ge 0$, vi) $\lambda_0^{\phi} + \sqrt{2\lambda_1 \lambda_0^{\Omega}} \ge 0$, vii) $\lambda_0^{\eta} + \sqrt{2\lambda_2 \lambda_0^{\Omega}} \ge 0$, viii) $\sqrt{2\lambda_1\lambda_2\lambda_0^{\Omega}} + \lambda_3\sqrt{2\lambda_0^{\Omega}} + \lambda_0^{\phi}\sqrt{\lambda_2} + \lambda_0^{\eta}\sqrt{\lambda_1}$ $+\sqrt{\left(\lambda_{3}+\sqrt{\lambda_{1}\lambda_{2}}
ight)\left(\lambda_{\Omega}^{\phi}+\sqrt{2\lambda_{1}\lambda_{\Omega}^{\Omega}}
ight)\left(\lambda_{\Omega}^{\eta}+\sqrt{2\lambda_{2}\lambda_{\Omega}^{\Omega}}
ight)}\geq0$ [Merle et al. 1603.05685, Kannike 1205.3781] **Perturbativity**: $\lambda_i < 4\pi$ and $Y_i < \sqrt{4\pi}$ \mathcal{Z}_2 symmetry: $\mu_0^\eta \leq \mathcal{O}(1 \text{ TeV})$ [Merle et al. 1603.05685] **EWPO**: $S = -0.02 \pm 0.10$, $T = 0.03 \pm 0.12$, $U = 0.01 \pm 0.11$ $\rho_{exp} = 1.00038 \pm 0.00020 \longrightarrow \rho = 1 + (4 v_{\Omega}^2 / v_{\phi}^2) \implies v_{\Omega} \lesssim 4 \text{ GeV} (3\sigma)$ Neutrino oscillation: $\sin^2 \theta_{12} = 0.304^{+0.012}_{-0.012}, \quad \sin^2 \theta_{23} = 0.537^{+0.016}_{-0.020}, \quad \sin^2 \theta_{13} = 0.0022^{+0.00062}_{-0.00063}$

 $\Delta m_{21}^2 = 7.42^{+0.20}_{-0.21} \times 10^{-5} \text{ eV}^2, \quad \text{`m}_{31}^2 = 2.517^{+0.026}_{-0.028} \times 10^{-3} \text{ eV}^2, \quad \delta_{\text{CP}} = 197^{\circ + 27^{\circ}}_{-24^{\circ}}$

[de Salas et al. 2006.11237]

Constraints

<u>Direct searches</u>: $m_H > 150 \text{ GeV}$ [LEP & LHC], $M_{\Sigma} > 100 \text{ GeV}$ (from Σ^+) [LEP],

$m_{\eta^+} > 70 \text{ GeV}$ [OPAL, hep-ph/0703056].

Though there exist some bounds on $\eta^0_{R,I}$ [0810.3924] from netralino search at LEP, they are not very constraining.

Bounds from LEP and LHC on charged scalar and neutral leptons are not directly applicable to H^+ , η^+ and $\chi^0_{1,2}$.

LHC bound on triplet fermion is not directly applicable to Σ .

LHC searches on slepton decaying to lepton and massless neutralino might put some bound on m_{H^+} . So, we choose $m_{H^+} > 400$ GeV (conservatively [1908.08215]).

Decay width: $\mathcal{B}(h \to inv) < 13\%$, $\delta \Gamma_Z < 5 \text{ MeV} (2\sigma)$, $\delta \Gamma_W < 90 \text{ MeV} (2\sigma)$ [PDG]

<u>cLFV</u>: $\mathcal{B}(\mu \to e\gamma) < 4.2 \times 10^{-13}$ [MEG], $\mathcal{B}(\mu \to 3e) < 1.0 \times 10^{-12}$ [SINDRUM] $\mathcal{C}(\mu, Au \to e, Au) < 7.0 \times 10^{-13}$ [SINDRUMI] **Relic density**: $\Omega h^2 = 0.120 \pm 0.001$ [Planck]

DM direct detection (Strongest):

0.1 GeV - 4 GeV: DarkSide-50 [2302.01830]

4 GeV - 10 GeV: XENON1T ⁸B [2012.02846] and PandaX-4T ⁸B [2207.04883]

10 GeV - 10 TeV: LZ [2207.03764]

XENON1T [1805.12562], XENONnT [2303.14729], PandaX-4T [2107.13438].

Coherent neutrino-nucleon scattering \rightarrow Neutrino floor [Billard et al. 2104.07634]

Parameters from Lagrangian:

Complex Yukawa	Real Yukawa	Scalar mass terms	Scalar quartic	Fermionic
couplings	couplings	and trilinear couplings	couplings	masses
$\begin{array}{c c} Y_{F}^{1}, \ Y_{F}^{2}, \ Y_{F}^{3}, \\ Y_{\Sigma}^{1}, \ Y_{\Sigma}^{2}, \ Y_{\Sigma}^{3} \end{array}$	Y _Ω	$\mu_{\phi}, \mu_{\eta}, \mu_{\Omega}, \mu_{\Omega}^{\phi}, \mu_{\Omega}^{\eta}$	$\begin{array}{c}\lambda_{1},\lambda_{2}\lambda_{3},\lambda_{4},\lambda_{5},\\\lambda_{\Omega}^{\phi},\lambda_{\Omega}^{\eta},\lambda_{\Omega}^{\Omega}\end{array}$	M_F, M_{Σ}

Define:
$$\Delta m_{\Sigma F} = M_{\Sigma} - M_F$$
, $\Delta m_{\eta^+ F} = m_{\eta^+} - M_F$ and $\Delta m_{\eta^0_{\eta^+}}^2 = m_{\eta^0_{\eta^-}}^2 - m_{\eta^+}^2$

After Reduction:

There remain 16 independent parameters (No Majorana phase).

BP0:

M _F (GeV)	$\Delta m_{\Sigma F}$ (GeV)	Δm_{η^+F} (GeV)	$\frac{\Delta m_{\eta_l^0 \eta^+}^2}{(\text{GeV}^2)}$	μ_{Ω}^{η} (GeV)	ν _Ω (GeV)	Y_{Ω}	Re(ω)	$Im(\omega)$	λ_1	λ_2	λ_3	λ_5	λ^{ϕ}_{Ω}	$\lambda_{\Omega}^{\Omega}$	λ_{Ω}^{η}
[1, 1000]	200	500	1000	400	4.0	2.0	π/4	$\pi/4$	0.2626	0.5	0.5	10-8	0.5	0.5	0.5

More on cLFV



More on cLFV



DM annihilation (No coannihilation)



Dependencies:

 $\begin{array}{l} \label{eq:scalar} \text{Scalar mixing and } m_{H^0} \colon \{\lambda_1, \lambda^\phi_\Omega, v_\Omega\} \\ \text{Fermion mixing: } \{v_\Omega, Y_\Omega, \, \Delta m_{\Sigma F}\} \\ \text{Couplings of SM and } \mathcal{Z}_2 \text{-odd leptons through } \eta \colon Y_{F,\Sigma} \longrightarrow \{\lambda_5, \operatorname{Im}(\omega), M_F\} \\ \text{Masses of } \mathcal{Z}_2 \text{-odd particles: } \{M_F, \, \Delta m_{\Sigma F}(\operatorname{slightly}), \, \Delta m_{\eta^+ F}(\operatorname{slightly})\} \end{array}$

Relic vs m_{DM} (BP0)



- 1st dip at $m_{DM} = 62.5 \text{ GeV} (m_h/2)$: $\chi_1^0 \chi_1^0 \rightarrow h^0 \rightarrow SM SM$
- 2nd dip at $m_{DM} = 200$ GeV $(m_H/2)$: $\chi_1^0 \chi_1^0 \rightarrow H^0 \rightarrow SM SM$
- 3rd dip near $m_{DM} = 250 \text{ GeV} \{ (m_{W^{\pm}} + m_{H^{\pm}})/2, (m_{h^0} + m_{H^0})/2 \}$:

Opening of $\chi_1^0 \chi_1^0 o W^{\pm} H^{\mp}$ and $\chi_1^0 \chi_1^0 o h^0 H^0$

• 4th dip near $m_{DM} = 400 \text{ GeV} (m_{H^0}, m_{H^{\pm}})$:

Opening of $\chi_1^0 \chi_1^0 \to H^0 H^0$ and $\chi_1^0 \chi_1^0 \to H^+ H^-$

Fermion-fermion coannihilation



Fermion-scalar coannihilation



Direct detection



$$\sum_{DM-N}^{SI} \approx \frac{\mu_{red}^2}{\pi} \left[\frac{Y_{\Omega} f_N m_N}{2v} \sin 2\theta \sin 2\beta \left(\frac{1}{m_{h^0}^2} - \frac{1}{m_{H^0}^2} \right) \right]^2$$

$$m_N: \text{ Nucleon mass}$$

$$\begin{array}{l} m_{N}: \text{Nucleon mass}\\ \overline{h}_{N}: \text{Nucleon form factor} \approx 0.3\\ \mu_{\text{red}}: \text{ Reduced mass} = (m_{\chi_{1}^{0}} m_{N})/(m_{\chi_{1}^{0}} + m_{N})\\ \overline{h}_{N}: \text{ Fermion mixing angle} \longrightarrow \tan(2\theta) = \frac{2 Y_{\Omega} v_{\Omega}}{M_{\Sigma} - M_{F}}\\ \overline{h}_{N}: \text{ Scalar mixing angle} \longrightarrow \tan 2\beta = \frac{4 v_{\Omega} v_{\phi} (\mu_{\Omega}^{\phi} - \sqrt{2} \lambda_{\phi}^{\Omega} v_{\Omega})}{4 \sqrt{2} \lambda_{\Omega}^{\Omega} v_{\Omega}^{3} - 2 \sqrt{2} \lambda_{1} v_{\Omega} v_{\phi}^{2} + \mu_{\Omega}^{\phi} v_{\phi}^{2}} \end{array}$$

Dependencies:
$$\sigma_{\mathbf{DM}-\mathbf{N}}^{SI} \longrightarrow \{\lambda_1, \lambda_{\Omega}^{\phi}, v_{\Omega}, Y_{\Omega}, \Delta m_{\Sigma F}\}$$
 (but not on $m_{\chi_1^0}$)

Cases	ν _Ω (GeV)	M _F (GeV)	$\Delta m_{\Sigma F}$ (GeV)	Δm_{η^+F} (GeV)	$\Delta m^2_{\eta^0_l \eta^+}$ (GeV ²)	μ_{Ω}^{η} (GeV)	YΩ	$Re(\omega)$	$Im(\omega)$	λ_1	λ_2	λ_3	λ_5	λ^{ϕ}_{Ω}	$\lambda_{\Omega}^{\Omega}$	λ_{Ω}^{η}
Scenario-I																
BP1			[100, 500]	[100, 500]	1000											
BP_1^{FF}	4.0	[3, 10000]	[1, 50]	[100, 500]	1000	400	[0.1, 3.5]	$[-\pi,\pi]$	$[-2\pi, 2\pi]$	0.2626	0.5	0.5	$[10^{-9}, 0.5]$	0.5	0.5	0.5
BP_1^{FS}			[100, 500]	[1, 30]	[1, 1000]											
Scenario-II																
BP ₂			[100, 500]	[100, 500]	1000	1			1	1.46						
BP ₂ ^{FF}	1.5	[3, 10000]	[1, 50]	[100, 500]	1000	400	[0.1, 3.5]	$[-\pi,\pi]$	$[-2\pi, 2\pi]$	0.2626	0.5	0.5	[10 ⁻⁹ , 0.5]	0.5	0.5	0.5
BP ₂ ^{FS}			[100, 500]	[1, 30]	[1, 1000]											

No Superscript: No coannihilation Superscript FF: Fermion-fermion coannihilation Superscript FS: Fermion-Scalar coannihilation

 $m_{H^0} = \{400 \text{ GeV} (\text{Scenario-I}), 1100 \text{ GeV} (\text{Scenario-II})\}$

Scanning: Relic



17

Scanning: Relic



17

Scanning: Direct detection

Scenario-I (High v_{Ω})



✓ This model cannot accommodate light fermionic dark matter (both singlet-like and triplet-like) below 62.5 GeV.

If Without any coannihilation feasible parameter-space can be obtained only for $m_{DM} > m_H$ (apart from discrete resonance at $m_{DM} = m_H/2$)

If Fermion-fermion coannihilation provides feasible parameter-space for $m_{DM} > 100$ GeV whereas fermion-scalar coannihilation does the same job from $m_{DM} > 62.5$ GeV.

 \measuredangle While fermion-scalar coannihilation is the most promising scenario for higher values of v_{Ω} , fermion-fermion coannihilation is the most promising case for lower values of v_{Ω} .

