

Phenomenology of Dirac Scotogenic Model

Sushant Yadav

(In Collaboration with Rahul Srivastava and Salvador Centelles Chuliá)

Department of Physics, IISER Bhopal, India

[arXiv:2206.11903](https://arxiv.org/abs/2206.11903)



Outline

- Introduction (SM and its Shortcomings)
- Scotogenic Model
- Dirac Scotogenic Model
- Dark Matter
- Higgs Vacuum Stability
- CDF II W Anomaly
- Conclusion

The Standard Model (SM) and its Shortcomings

- A quantum field theory which provides a very good understanding of three of the four fundamental forces: Strong, Weak, and EM.
- It is based on the "Gauge Principle". $SU(3)_C \times SU(2)_L \times U(1)_Y$
- It contains 3 generations of leptons and quarks with Gauge bosons and Higgs boson.
- SM stands as a remarkably successful theoretical framework, providing a comprehensive description of all known particles and their interactions with very great accuracy.

The Standard Model (SM) and its Shortcomings

- A quantum field theory which provides a very good understanding of three of the four fundamental forces: Strong, Weak, and EM.
- It is based on the "Gauge Principle". $SU(3)_C \times SU(2)_L \times U(1)_Y$
- It contains 3 generations of leptons and quarks with Gauge bosons and Higgs boson.
- SM stands as a remarkably successful theoretical framework, providing a comprehensive description of all known particles and their interactions with very great accuracy.
- There are various shortcomings of SM: Neutrino Mass, Dark Matter (DM), Matter-antimatter asymmetry, CDF-II W boson mass anomaly, Muon's anomalous magnetic dipole moment, Vacuum stability problem etc.
- Two important shortcomings are: [Neutrino Mass and Dark Matter](#)

The Standard Model (SM) and its Shortcomings

- A quantum field theory which provides a very good understanding of three of the four fundamental forces: Strong, Weak, and EM.
- It is based on the "Gauge Principle". $SU(3)_C \times SU(2)_L \times U(1)_Y$
- It contains 3 generations of leptons and quarks with Gauge bosons and Higgs boson.
- SM stands as a remarkably successful theoretical framework, providing a comprehensive description of all known particles and their interactions with very great accuracy.
- There are various shortcomings of SM: Neutrino Mass, Dark Matter (DM), Matter-antimatter asymmetry, CDF-II W boson mass anomaly, Muon's anomalous magnetic dipole moment, Vacuum stability problem etc.
- Two important shortcomings are: **Neutrino Mass and Dark Matter**
- **Neutrino Mass:**
 - ① The SM predicts neutrinos to be massless.
 - ② Neutrino oscillation experiments [Super-Kamiokande:1998kpq] predict masses for neutrinos.
 - ③ At least two neutrinos are massive.
- **Dark Matter:**
 - ① Discrepancy in galactic rotation curve \Rightarrow One possible solution is DM.
 - ② Flat galactic rotation curves seem to suggest that each galaxy is surrounded by significant amounts of non-visible matter known as dark matter.
 - ③ There is no candidate for dark matter within SM.

Scotogenic Model

Scotogenic Model: Minimal extension of SM (Proposed by Ernest Ma in 2006) [\[Ma:2006km\]](#).

- It provides tiny neutrino mass and dark matter stability simultaneously within the same framework.

Scotogenic Model

Scotogenic Model: Minimal extension of SM (Proposed by Ernest Ma in 2006) [\[Ma:2006km\]](#).

- It provides tiny neutrino mass and dark matter stability simultaneously within the same framework.
- Light neutrino masses are generated via the one-loop radiative seesaw mechanism.

Scotogenic Model

Scotogenic Model: Minimal extension of SM (Proposed by Ernest Ma in 2006) [Ma:2006km].

- It provides tiny neutrino mass and dark matter stability simultaneously within the same framework.
- Light neutrino masses are generated via the one-loop radiative seesaw mechanism.
- Two newly added BSM fields: Scalar doublet η and Fermion singlet N .
- A new symmetry Z_2 : new fields are odd under Z_2 and SM fields are even under Z_2 .

Scotogenic Model

Scotogenic Model: Minimal extension of SM (Proposed by Ernest Ma in 2006) [Ma:2006km].

- It provides tiny neutrino mass and dark matter stability simultaneously within the same framework.
- Light neutrino masses are generated via the one-loop radiative seesaw mechanism.
- Two newly added BSM fields: Scalar doublet η and Fermion singlet N .
- A new symmetry Z_2 : new fields are odd under Z_2 and SM fields are even under Z_2 .
- Two possible DM candidates:
 - 1 Neutral Scalar η^0
 - 2 Neutral Fermion N

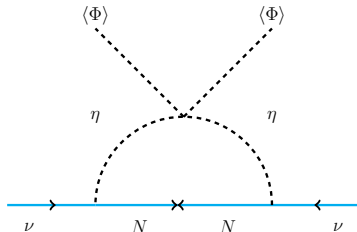


Figure 1: Leading neutrino mass generation diagram.

Dirac Scotogenic Model

Dirac Scotogenic Model:

- It is theoretical framework [\[Bonilla:2018ynb\]](#) to obtain stable dark matter along with naturally small Dirac neutrino masses generated at the loop level.
- This is achieved through the symmetry breaking of the global $U(1)_{B-L}$ symmetry already present in the SM.

Dirac Scotogenic Model

Dirac Scotogenic Model:

- It is theoretical framework [\[Bonilla:2018ynb\]](#) to obtain stable dark matter along with naturally small Dirac neutrino masses generated at the loop level.
- This is achieved through the symmetry breaking of the global $U(1)_{B-L}$ symmetry already present in the SM.
- Dirac/Majorana nature of neutrinos is intimately connected with the $U(1)_{B-L}$ symmetry of the SM and its possible breaking pattern.

$U(1)_{B-L} \rightarrow \mathcal{Z}_m \equiv \mathcal{Z}_{2n+1}$ with $n \in \mathbb{Z}^+$ \Rightarrow neutrinos are Dirac particles

$U(1)_{B-L} \rightarrow \mathcal{Z}_m \equiv \mathcal{Z}_{2n}$ with $n \in \mathbb{Z}^+$ \Rightarrow neutrinos can be Dirac or Majorana

$$\text{Lepton doublet } L_i \begin{cases} \sim \omega^n & \text{under } \mathcal{Z}_{2n} & \Rightarrow \text{Dirac neutrinos} \\ \sim \omega^n & \text{under } \mathcal{Z}_{2n} & \Rightarrow \text{Majorana neutrinos} \end{cases} \quad (1)$$

where $\omega^{2n} = 1$ or $\omega = \exp \frac{2\pi i}{2n}$.

Dirac Scotogenic Model

Dirac Scotogenic Model:

- It is theoretical framework [Bonilla:2018ynb] to obtain stable dark matter along with naturally small Dirac neutrino masses generated at the loop level.
- This is achieved through the symmetry breaking of the global $U(1)_{B-L}$ symmetry already present in the SM.
- Dirac/Majorana nature of neutrinos is intimately connected with the $U(1)_{B-L}$ symmetry of the SM and its possible breaking pattern.

$U(1)_{B-L} \rightarrow \mathcal{Z}_m \equiv \mathcal{Z}_{2n+1}$ with $n \in \mathbb{Z}^+$ \Rightarrow neutrinos are Dirac particles

$U(1)_{B-L} \rightarrow \mathcal{Z}_m \equiv \mathcal{Z}_{2n}$ with $n \in \mathbb{Z}^+$ \Rightarrow neutrinos can be Dirac or Majorana

$$\text{Lepton doublet } L_i \begin{cases} \sim \omega^n & \text{under } \mathcal{Z}_{2n} & \Rightarrow \text{Dirac neutrinos} \\ \sim \omega^n & \text{under } \mathcal{Z}_{2n} & \Rightarrow \text{Majorana neutrinos} \end{cases} \quad (1)$$

where $\omega^{2n} = 1$ or $\omega = \exp \frac{2\pi i}{2n}$.

- This $U(1)_{B-L}$ is multipurpose:
 - ① Protect the stability of DM:
By forbidding terms that lead to decay or mixing of dark sector particles to the SM particles.

Dirac Scotogenic Model

Dirac Scotogenic Model:

- It is theoretical framework [Bonilla:2018ynb] to obtain stable dark matter along with naturally small Dirac neutrino masses generated at the loop level.
- This is achieved through the symmetry breaking of the global $U(1)_{B-L}$ symmetry already present in the SM.
- Dirac/Majorana nature of neutrinos is intimately connected with the $U(1)_{B-L}$ symmetry of the SM and its possible breaking pattern.

$U(1)_{B-L} \rightarrow \mathcal{Z}_m \equiv \mathcal{Z}_{2n+1}$ with $n \in \mathbb{Z}^+ \Rightarrow$ neutrinos are Dirac particles

$U(1)_{B-L} \rightarrow \mathcal{Z}_m \equiv \mathcal{Z}_{2n}$ with $n \in \mathbb{Z}^+ \Rightarrow$ neutrinos can be Dirac or Majorana

$$\text{Lepton doublet } L_i \begin{cases} \approx \omega^n & \text{under } \mathcal{Z}_{2n} \Rightarrow \text{Dirac neutrinos} \\ \sim \omega^n & \text{under } \mathcal{Z}_{2n} \Rightarrow \text{Majorana neutrinos} \end{cases} \quad (1)$$

where $\omega^{2n} = 1$ or $\omega = \exp \frac{2\pi i}{2n}$.

- This $U(1)_{B-L}$ is multipurpose:
 - 1 Protect the stability of DM:
By forbidding terms that lead to decay or mixing of dark sector particles to the SM particles.
 - 2 Protect the Dirac nature of the neutrinos and the smallness of neutrino masses by forbidding the tree-level coupling with the Higgs field.

Dirac Scotogenic Model

Dirac Scotogenic Model:

- It is theoretical framework [Bonilla:2018ynb] to obtain stable dark matter along with naturally small Dirac neutrino masses generated at the loop level.
- This is achieved through the symmetry breaking of the global $U(1)_{B-L}$ symmetry already present in the SM.
- Dirac/Majorana nature of neutrinos is intimately connected with the $U(1)_{B-L}$ symmetry of the SM and its possible breaking pattern.

$U(1)_{B-L} \rightarrow \mathcal{Z}_m \equiv \mathcal{Z}_{2n+1}$ with $n \in \mathbb{Z}^+$ \Rightarrow neutrinos are Dirac particles

$U(1)_{B-L} \rightarrow \mathcal{Z}_m \equiv \mathcal{Z}_{2n}$ with $n \in \mathbb{Z}^+$ \Rightarrow neutrinos can be Dirac or Majorana

$$\text{Lepton doublet } L_i \begin{cases} \approx \omega^n & \text{under } \mathcal{Z}_{2n} & \Rightarrow \text{Dirac neutrinos} \\ \sim \omega^n & \text{under } \mathcal{Z}_{2n} & \Rightarrow \text{Majorana neutrinos} \end{cases} \quad (1)$$

where $\omega^{2n} = 1$ or $\omega = \exp \frac{2\pi i}{2n}$.

- This $U(1)_{B-L}$ is multipurpose:
 - 1 Protect the stability of DM:
By forbidding terms that lead to decay or mixing of dark sector particles to the SM particles.
 - 2 Protect the Dirac nature of the neutrinos and the smallness of neutrino masses by forbidding the tree-level coupling with the Higgs field.
 - 3 It also predicts that the lightest neutrino is massless.

The Model Setup

- $(-4, -4, 5)$ Chiral solutions to $U(1)_{B-L}$ anomaly cancellation conditions (forbidding the tree-level neutrino Yukawa couplings).

	Fields	$SU(2)_L \otimes U(1)_Y$	$U(1)_{B-L}$	Z_6
Fermions	L_i	$(\mathbf{2}, -1/2)$	-1	ω^4
	ν_{R_i}	$(\mathbf{1}, 0)$	$(-4, -4, 5)$	$(\omega^4, \omega^4, \omega^4)$
	N_{L_i}	$(\mathbf{1}, 0)$	$-1/2$	ω^5
	N_{R_i}	$(\mathbf{1}, 0)$	$-1/2$	ω^5
Scalars	H	$(\mathbf{2}, 1/2)$	0	1
	η	$(\mathbf{2}, 1/2)$	$1/2$	ω
	ξ	$(\mathbf{1}, 0)$	$7/2$	ω

Table 1: Charge assignment for all the fields.

- 0 B-L charge for Higgs is zero to preserve Yukawa terms for fermions that give mass to them.

Breaking Pattern of $U(1)_{B-L}$ Symmetry

- The $U(1)_{B-L} \rightarrow Z_6$ breaking happens because of the presence of the soft term

$$(\kappa \eta^\dagger H \xi + h.c.)$$

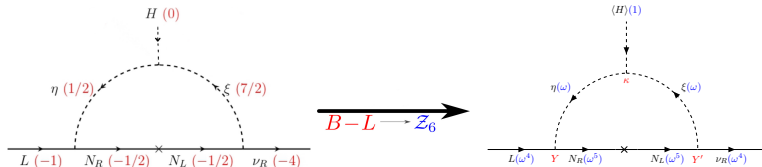


Figure 2: Charge assignment and symmetry breaking pattern for $U(1)_{B-L} \rightarrow Z_6$.

- This residual Z_6 symmetry simultaneously protects the Dirac nature of neutrinos and the stability of DM.

The Scalar Potential

The general form of the scalar potential is given by

$$\begin{aligned} V = & -\mu_H^2 H^\dagger H + \mu_\eta^2 \eta^\dagger \eta + \mu_\xi^2 \xi^* \xi + \frac{1}{2} \lambda_1 (H^\dagger H)^2 + \frac{1}{2} \lambda_2 (\eta^\dagger \eta)^2 + \frac{1}{2} \lambda_3 (\xi^* \xi)^2 \\ & + \lambda_4 (H^\dagger H)(\eta^\dagger \eta) + \lambda_6 (\eta^\dagger \eta)(\xi^* \xi) + \lambda_7 (H^\dagger \eta)(\eta^\dagger H) + \lambda_8 (H^\dagger H)(\xi^* \xi) \\ & + (\kappa \eta^\dagger H \xi + h.c.) \end{aligned} \quad (2)$$

The Scalar Potential

The general form of the scalar potential is given by

$$\begin{aligned} V = & -\mu_H^2 H^\dagger H + \mu_\eta^2 \eta^\dagger \eta + \mu_\xi^2 \xi^{*\dagger} \xi + \frac{1}{2} \lambda_1 (H^\dagger H)^2 + \frac{1}{2} \lambda_2 (\eta^\dagger \eta)^2 + \frac{1}{2} \lambda_3 (\xi^{*\dagger} \xi)^2 \\ & + \lambda_4 (H^\dagger H)(\eta^\dagger \eta) + \lambda_6 (\eta^\dagger \eta)(\xi^{*\dagger} \xi) + \lambda_7 (H^\dagger \eta)(\eta^\dagger H) + \lambda_8 (H^\dagger H)(\xi^{*\dagger} \xi) \\ & + (\kappa \eta^\dagger H \xi + h.c.) \end{aligned} \quad (2)$$

Bounded from below scalar potential, ensured by the following conditions [\[Kannike:2016fmd\]](#)

$$\begin{aligned} \lambda_1, \lambda_2, \lambda_3 &\geq 0; & \lambda_4 &> -\sqrt{\lambda_1 \lambda_2}, & \lambda_6 &> -\sqrt{\lambda_2 \lambda_3}, & \lambda_8 &> -\sqrt{\lambda_1 \lambda_3}, \\ \sqrt{\frac{\lambda_3}{2}} \lambda_4 + \sqrt{\frac{\lambda_1}{2}} \lambda_6 + \sqrt{\frac{\lambda_2}{2}} \lambda_8 + \sqrt{\frac{\lambda_1 \lambda_2 \lambda_3}{8}} \\ &> -\sqrt{(\lambda_4 + \sqrt{\lambda_1 \lambda_2})(\lambda_8 + \sqrt{\lambda_1 \lambda_3})(\lambda_6 + \sqrt{\lambda_2 \lambda_3})} \end{aligned} \quad (3)$$

Neutrino Mass

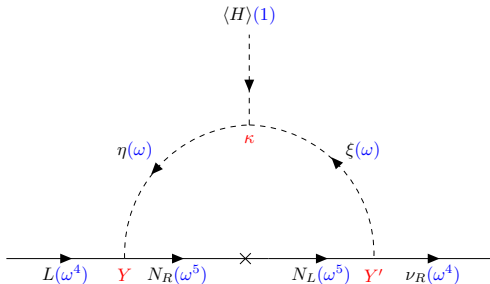


Figure 3: Leading neutrino mass generation diagram.

The relevant Yukawa Lagrangian for neutrino masses is given by

$$-\mathcal{L}_Y \supset Y_{il} \bar{L}_i \tilde{\eta} N_{Rl} + Y'_{li} \bar{N}_{Ll} \nu_{Ri} \xi + M_{lm} \bar{N}_{Rl} N_{Lm} + h.c. \quad (4)$$

We can calculate neutrino masses from the diagram Fig.3 as

$$(M_\nu)_{ij} = \frac{1}{16\pi^2} \sum_{k=1}^3 Y_{ik} Y'_{kj} \frac{\kappa \nu}{m_\xi^2 - m_\eta^2} M_k \sum_{l=1}^2 (-1)^l B_0(0, m_l^2, M_k^2). \quad (5)$$

Phenomenology

- We performed a detailed numerical scan for the model parameters with various experimental and theoretical constraints.
- We have implemented the model in SARAH-4.14.5 [[Staub:2015kfa](#)] and SPheno-4.0.5 [[Porod:2011nf](#)] to calculate all the vertices, mass matrices and tadpole equations.
- Thermal component to the DM relic abundance as well as the DM nucleon scattering cross sections are determined by micrOMEGAS-5.2.13 [[Belanger:2014vza](#)]

Phenomenology

- We performed a detailed numerical scan for the model parameters with various experimental and theoretical constraints.
- We have implemented the model in SARAH-4.14.5 [[Staub:2015kfa](#)] and SPheno-4.0.5 [[Porod:2011nf](#)] to calculate all the vertices, mass matrices and tadpole equations.
- Thermal component to the DM relic abundance as well as the DM nucleon scattering cross sections are determined by micrOMEGAS-5.2.13 [[Belanger:2014vza](#)]

we have also imposed the following additional conditions when generating the allowed points:

- Neutrino oscillation parameters.
- Bounded from below scalar potential, ensured by the vacuum stability constraints.
- Perturbativity of Yukawas and quartic couplings.
- If η^0 is the DM particle its mass must be smaller than the charged counterpart η^+ . This implies $\lambda_7 < 0$ in the small mixing limit.
- Finally, we impose the LEP constraint on the light-neutral component of a doublet. This limit is actually simply $m_{\eta R} + m_{\eta I} > m_Z$ which in our case translates to $m_{\eta^0} > m_Z/2 \approx 45.6$ GeV.

Doublet DM Case:

- Relic density computation and direct detection prospects for doublet DM case involve the exchange of a Higgs or Z boson.
- Our analysis shows that magenta points cover three mass regions in the relic density plot:
 - ① The low mass region from 10 GeV to around 30 GeV (Ruled out by LEP constraints).
 - ② The medium mass region from 58 GeV to around 122 GeV.
 - ③ The high mass region from 200 GeV to around 4.8 TeV.

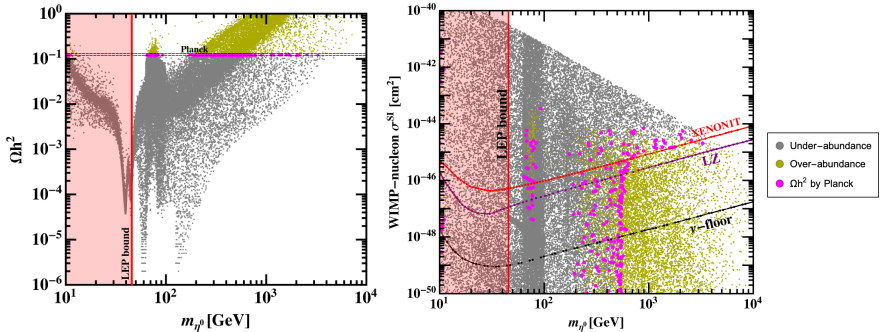


Figure 4: **Left:** Relic density plot for η^0 dominated DM. **Right:** Spin-independent η^0 WIMP-nucleon cross section for the η^0 dominated DM candidate case.

Dark Matter

Singlet DM Case:

- Relic density computation and direct detection prospects involve just a Higgs portal in the singlet case.
- Magenta points cover a broad mass region up to around 5 TeV that satisfy all theoretical and experimental constraints.

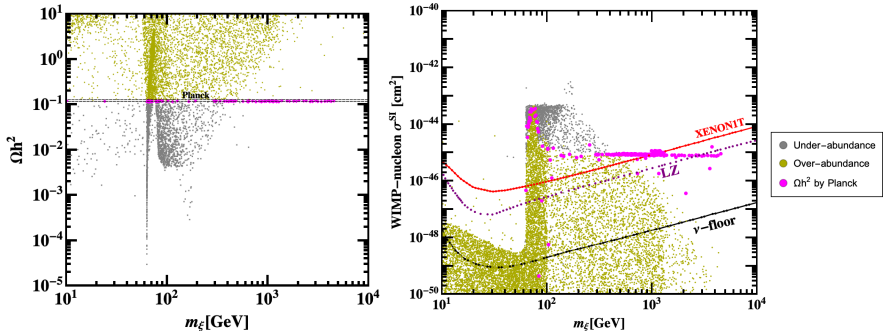


Figure 5: **Left:** Relic density vs singlet DM mass. **Right:** Spin-independent WIMP-nucleon cross section for the ξ dominated DM candidate vs DM mass.

Fermionic DM Case:

- Fermionic DM candidate satisfies all the constraints up to around 2 TeV.
- In the relic density computation for fermionic DM, coannihilation channels between dark fermion and doublet scalar become important if the relative dark fermion-scalar mass difference is below 10 GeV.

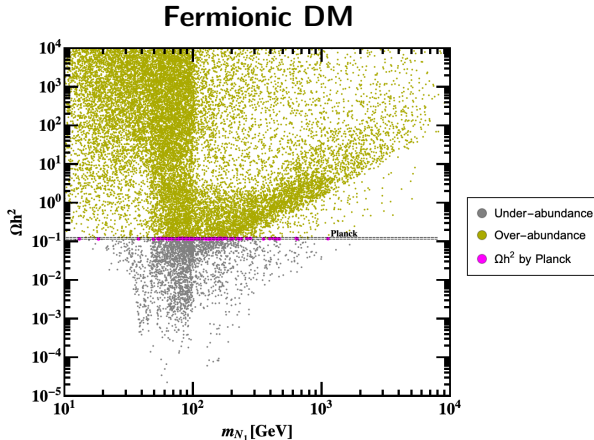


Figure 6: Relic density vs fermionic DM mass.

Higgs Vacuum in Standard Model

- In SM, there exists a problem with the stability of the electroweak vacuum since the electroweak vacuum becomes unstable at large scale ($\sim 10^{10}$ GeV).
- At this elevated scale, the quartic coupling of the SM Higgs, denoted as λ_{HH} , undergoes a transition to a negative value as dictated by the evolution of the renormalization group equations (RGE).

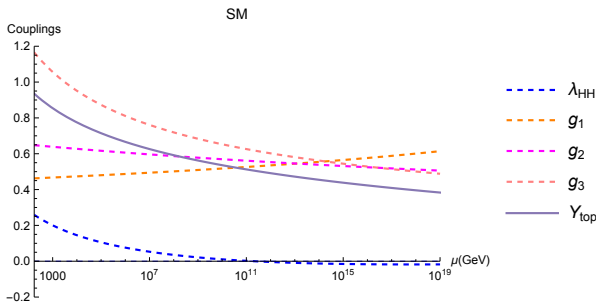


Figure 7: The RG evolution of the SM gauge couplings g_1 , g_2 , g_3 , the top quark Yukawa coupling Y_{top} and the quartic Higgs boson self-coupling λ_{HH} in the Standard Model.

Vacuum Stability in Dirac Scotogenic Model

- The beta functions for various gauge, quartic and Yukawa couplings in the model are evaluated up to the two-loop level.

Vacuum Stability in Dirac Scotogenic Model

- The beta functions for various gauge, quartic and Yukawa couplings in the model are evaluated up to the two-loop level.
- In our analysis, we observe a notable dependence of the quartic Higgs self-coupling (λ_{HH}) on various interaction couplings, namely $\lambda_{H\eta}$, $\lambda'_{H\eta}$ and $\lambda_{H\xi}$ denoted by λ_4 , λ_7 and λ_8 , respectively, within the scalar potential.

Vacuum Stability in Dirac Scotogenic Model

- The beta functions for various gauge, quartic and Yukawa couplings in the model are evaluated up to the two-loop level.
- In our analysis, we observe a notable dependence of the quartic Higgs self-coupling (λ_{HH}) on various interaction couplings, namely $\lambda_{H\eta}$, $\lambda'_{H\eta}$ and $\lambda_{H\xi}$ denoted by λ_4 , λ_7 and λ_8 , respectively, within the scalar potential.
- As we explore the parameter space, we find that the values of these couplings within the range of 0.15 to 0.50 yield significant corrections to λ_{HH} .

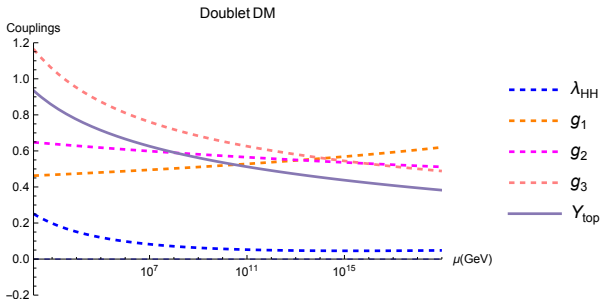


Figure 8: The RG evolution of the SM gauge couplings g_1 , g_2 , g_3 , the top quark Yukawa coupling Y_{top} and the quartic Higgs boson self-coupling λ_{HH} in the Dirac Scotogenic Model.

CDF-II W anomaly

- In 2022, the CDF-II collaboration reported a 7σ excess on the mass of the W boson with respect to the SM prediction.
- In the Dirac Scotogenic model, the doublet dark scalar leads to radiative corrections to W boson mass. [\[CentellesChulia:2022vpz\] arXiv:2206.11903](#)

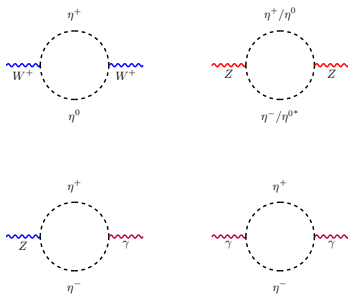


Figure 9: One loop polarization diagrams that contribute to the oblique S , T and U parameters.

Doublet Scalar DM

CDF-II mW analysis for doublet scalar DM:

- Doublet scalar mass is constrained to the medium mass region (58-86 GeV) after applying CDF-II mW constraints.

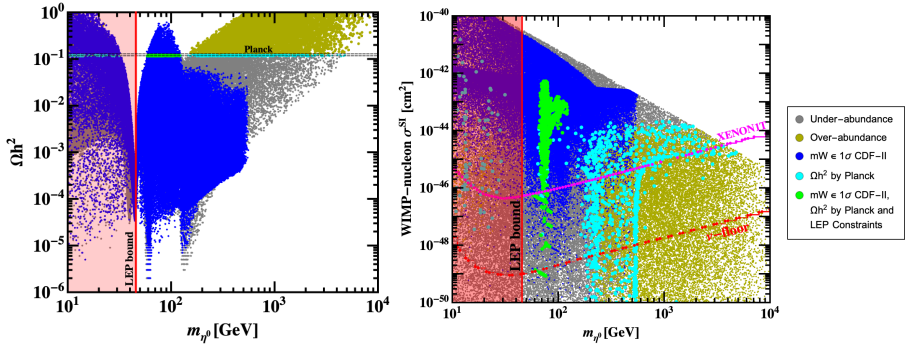


Figure 10: **Left:** Relic density plot for η^0 dominated DM. **Right:** Spin-independent WIMP-nucleon cross section for the η^0 dominated DM candidate case.

Singlet Scalar DM

CDF-II mW analysis for singlet scalar DM:

- Singlet scalar mass is constrained to up to around 500 GeV after applying CDF-II mW constraints.

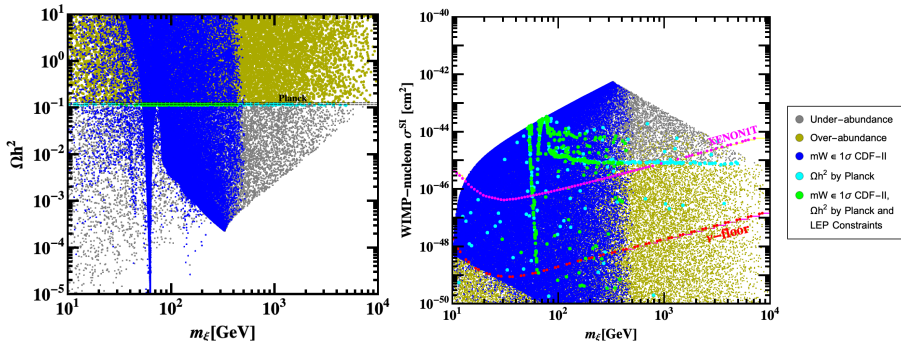


Figure 11: **Left:** Relic density vs singlet DM mass. **Right:** Spin-independent WIMP-nucleon cross section for the ξ dominated DM candidate vs DM mass.

Conclusion

- We have shown that the Dirac scotogenic model can reproduce the neutrino masses and mixing, the DM relic abundance and explain the CDF-II W boson mass anomaly while also ensuring the stability of electroweak vacuum up to the Planck energy scale.
- We find that the case of a mainly scalar doublet DM is constrained for the mass range of 58-86 GeV by the combination of the requirements that W boson mass remains within 1σ of the CDF-II measurement and the constraints coming from DM relic density, direct detection and invisible Z boson decays.
- We showed that if the singlet scalar is the DM candidate then all the above constraints are simultaneously satisfied along with W boson mass within 1σ range of the CDF-II measurement, where the singlet DM mass is constrained up to around 500 GeV.
- Fermionic DM is permissible within a mass range extending from 10 GeV to approximately 2000 GeV.
- Moreover, we find that we have to take some scalar couplings within the range of 0.15 to 0.50 to get enough correction to Higgs self-coupling so that electroweak vacuum will remain stable up to the Planck energy scale

Thank you for your attention!

Dirac Scoto Mass Spectrum

Fleshing out the $SU(2)_L$ components of the scalars, we can write the doublets as

$$H = \begin{pmatrix} H^+ \\ H^0 \end{pmatrix}, \quad \eta = \begin{pmatrix} \eta^+ \\ \eta^0 \end{pmatrix} \quad (6)$$

$$H^0 = \frac{1}{\sqrt{2}}(v + h + iA), \quad \eta^0 = \frac{1}{\sqrt{2}}(\eta_R + i\eta_I), \quad \xi = \frac{1}{\sqrt{2}}(\xi_R + i\xi_I) . \quad (7)$$

We can now compute the masses of the physical scalar states after symmetry breaking

$$m_h^2 = \lambda_1 v^2, \quad (8)$$

$$m_{\eta^\pm}^2 = \mu_\eta^2 + \frac{\lambda_4}{2} v^2, \quad (9)$$

The real part of ξ will mix with the real part of η^0 and similarly the imaginary part of ξ will mix with the imaginary part of η^0 with the same mixing matrix.

$$m_{(\xi_R, \eta_R)}^2 = m_{(\xi_I, \eta_I)}^2 = \begin{pmatrix} \mu_\xi^2 + \lambda_8 \frac{v^2}{2} & \kappa \frac{v}{\sqrt{2}} \\ \kappa \frac{v}{\sqrt{2}} & \mu_\eta^2 + (\lambda_4 + \lambda_7) \frac{v^2}{2} \end{pmatrix} \quad (10)$$

Mass Spectrum

We can compute the mixing angle

$$\tan 2\theta = \frac{\sqrt{2}\kappa v}{(\mu_\xi^2 - \mu_\eta^2) + (\lambda_8 - \lambda_4 - \lambda_7)\frac{v^2}{2}}, \quad (11)$$

and the mass eigenstates for the real/imaginary part of neutral scalars η^0 and ξ are given by

$$m_{1R}^2 = m_{1I}^2 = \left(\mu_\xi^2 + \lambda_8 \frac{v^2}{2}\right) \cos^2 \theta + \left(\mu_\eta^2 + (\lambda_4 + \lambda_7) \frac{v^2}{2}\right) \sin^2 \theta - 2\kappa v \sin \theta \cos \theta = m_\xi^2 \quad (12)$$

$$m_{2R}^2 = m_{2I}^2 = \left(\mu_\xi^2 + \lambda_8 \frac{v^2}{2}\right) \sin^2 \theta + \left(\mu_\eta^2 + (\lambda_4 + \lambda_7) \frac{v^2}{2}\right) \cos^2 \theta + 2\kappa v \sin \theta \cos \theta = m_{\eta^0}^2 \quad (13)$$

W mass and the S,T,U parameters

We can calculate the BSM contributions to S , T and U of the model as

$$S = \frac{1}{12\pi} \log \frac{m_{\eta^0}^2}{m_{\eta^+}^2} \quad (14)$$

$$T = \frac{G_F}{4\sqrt{2}\pi^2\alpha_{em}} \left(\frac{m_{\eta^0}^2 + m_{\eta^+}^2}{2} - \frac{m_{\eta^0}^2 m_{\eta^+}^2}{m_{\eta^+}^2 - m_{\eta^0}^2} \log \frac{m_{\eta^+}^2}{m_{\eta^0}^2} \right) \quad (15)$$

$$U = \frac{1}{12\pi} \left(\frac{(m_{\eta^0}^2 + m_{\eta^+}^2) (m_{\eta^0}^4 - 4m_{\eta^0}^2 m_{\eta^+}^2 + m_{\eta^+}^4) \log \left(\frac{m_{\eta^+}^2}{m_{\eta^0}^2} \right)}{(m_{\eta^+}^2 - m_{\eta^0}^2)^3} - \frac{5m_{\eta^0}^4 - 22m_{\eta^0}^2 m_{\eta^+}^2 + 5m_{\eta^+}^4}{3(m_{\eta^+}^2 - m_{\eta^0}^2)^2} \right)$$

In terms of the oblique S , T and U parameters, the corrections to the W boson mass are given by

$$m_W^2 = m_W^{2(\text{SM})} + \frac{\alpha_{em} \cos^2 \theta_w}{\cos^2 \theta_w - \sin^2 \theta_w} m_Z^2 \left[-\frac{1}{2} S + \cos^2 \theta_w T + \frac{(\cos^2 \theta_w - \sin^2 \theta_w)}{4 \sin^2 \theta_w} U \right] \quad (16)$$

where θ_w is the weak angle, α_{em} is the fine-structure constant and $m_W^{(\text{SM})}$ is the Standard Model prediction for m_W .