Singlet-Doublet Fermionic Dark Matter with Dirac Neutrino Mass, (g-2) and $\Delta N_{\rm eff}$

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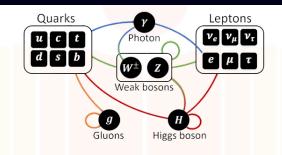
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Overview of the talk

- Introduction
- Oirac Neutrino Mass
- The Model (Singlet-Doublet Fermionic Dark Matter)
- Anomalous Magnetic Moment
- Oirect Detection
- $\bullet \Delta N_{\rm eff}$
- Summary

Introduction

What We Have..



Neutrino Physics



- Neutrino Mass
- Type of the Neutrino Field (Majorana/Dirac).
 - Number of Neutrino Species $(\mathrm{N_{eff}}=2.99^{+0.34}_{-0.33} \text{ at } 2\sigma \text{ confidence level})[1807.06209]$

Dark Matter

- Non-luminous, Non-baryonic matter
- Evidences from Astrophysical and Cosmologica observations
- $\Omega_{\rm DM} h^2 \simeq 0.12 \pm 0.001 \ [1807.06209]$



• N_{eff} is the effective degrees of freedom, other than photon(γ), present in the Universe.

$$N_{\rm eff} = \left(\frac{11}{4}\right)^{4/3} \sum_{\rm boson} \left(\frac{\rho_{rad} - \rho_{\gamma}}{\rho_{\gamma}}\right) + \sum_{\rm fermion} \left(\frac{\rho_{rad} - \rho_{\gamma}}{\rho_{\gamma}}\right)$$

- 3 flavours of active neutrinos present in SM contributes only $N_{\nu}=3.045$ (hep-ph/0506164)
- Any deviation of SM contribution from $N_{\rm eff}$, can be thought of as from extra relativistic particles/radiation.

$$\Delta N_{\rm eff} = \left(\frac{\rho_{\nu_R}}{\rho_{\nu_L}}\right)_{T_D(\nu_L)}$$

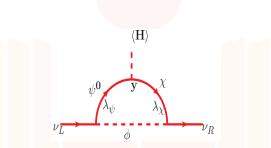
• Any extra relativistic particle may affect the matter-radiation equality and Hubble expansion rate. And hence, different imprints on CMB spectra.

Assumption: There are as many neutrinos as anti-neutrinos.

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Neutrino Mass Generation

Dirac Neutrino



Effective Langrangian term for Dirac Neutrino Mass, $\mathcal{L} \supset \overline{L}_L \widetilde{H} \nu_R$ $\Rightarrow M_{\nu} \overline{\nu}_L \nu_R$

• Do not forget about oscillation parameters $\Delta m_{sol}^2 = 7.54 \pm 0.18 \times 10^{-5} \text{ eV}^2,$ $\Delta m_{atm}^2 = 2.43 \pm 0.03 \times 10^{-3} \text{ eV}^2(\text{N0})[\text{PTET 2022}]$ and cosmological mass bound, $\sum m_{\nu_i} \leq 0.12 \text{ eV}$ [1807.06209].

Gauge	Fermion Fields				Scalar
Group	L	Ψ	χ	ν_R	ϕ
$SU(2)_L$	2	2	1	1	1
$U(1)_Y$	-1	-1	0	0	0
Z ₃	ω	ω^2	ω^2	ω^2	1
Z ₂	+1	-1	-1	+1	-1

We have considered 3 generations of Siglet- Doublet fermion to generate 3 non-zero neutrino mass states.

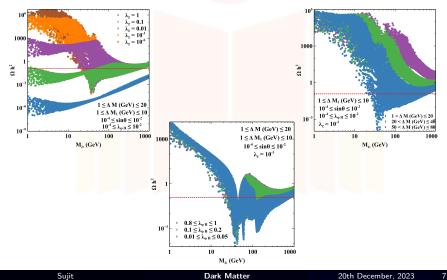
Lagrangian

$$\mathcal{L} \supset i\overline{\Psi}\gamma^{\mu}D_{\mu}\Psi + i\overline{\chi}\gamma^{\mu}\partial_{\mu}\chi - M_{\Psi}\overline{\Psi}\Psi - M_{\chi}\overline{\chi}\chi - y\overline{\Psi}\widetilde{H}\chi - \lambda_{\Psi}\overline{\ell}\Phi\Psi - \lambda_{\chi}\overline{\nu_{R}}\Phi\chi + h.c.$$

Where
$$D_{\mu} = \partial_{\mu} - g_1 \frac{\tau_i}{2} W^i_{\mu} - g_2 \frac{\gamma}{2} B_{\mu}, \Psi = (\psi^0 \ \psi^-)^T$$

$$\begin{pmatrix} \psi^0 \\ \chi \end{pmatrix} = U \begin{pmatrix} \chi_1 \\ \chi_2 \end{pmatrix} = \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \chi_1 \\ \chi_2 \end{pmatrix}, \tag{1}$$

Parameter space is spanned by $(\lambda_{\chi}, \lambda_{\psi}, sin\theta, M_{\chi_2}, \Delta M = M_{\chi_1} - M_{\chi_2}, \Delta M_1 = M_{\phi} - M_{\chi_2})$



Dark Matter

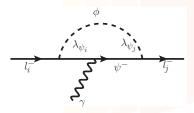
Anomalous Magnetic Moment

$$\vec{\mu_{\mu}} = g \frac{e}{2m_{\mu}} \vec{s}$$

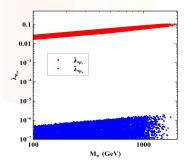
g-value is basically the sum of internal degrres of freedom and effects from quantum corrections.

$$\Delta a_{\mu} = a_{\mu}^{\exp} - a_{\mu}^{SM} = 249(48) \times 10^{-11} [2308.06230]$$

Where a_{μ}^{SM} includes all quantum corrections from SM.



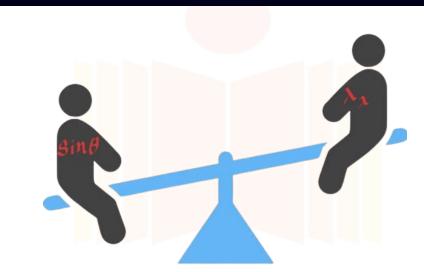
Loop diagram for (g-2) and LFV



Allowed parameter space from $(g-2)_{\mu}$ and LFV $(\mu^-
ightarrow \gamma e^-)$

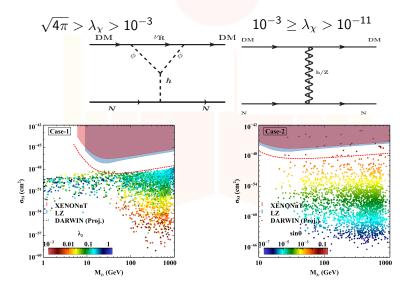
Dark Matter

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Trade-off between λ_{χ} and $\sin \theta$.

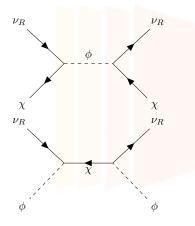
Direct Detection



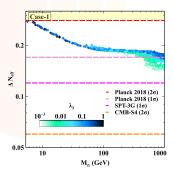
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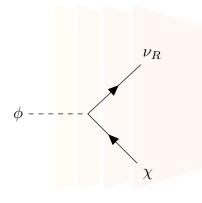
Dark Matter

$\begin{array}{c} \textbf{Model} \\ \Delta \textit{N}_{\rm eff} \mbox{ (Thermal)} \end{array}$

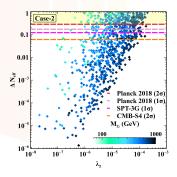


 $\sqrt{4\pi} > \lambda_{\chi} > 10^{-3}$ Thermal Production of ν_R





 $10^{-3} \ge \lambda_{\chi} > 10^{-11}$ Non-thermal Production of ν_R



- We incorporated Dirac neutrino mass in Singlet-Doublet fermionic DM model.
- Analysed our model parameters with muon (g-2) and LFV bounds.
- The extended model gives more freedom on singlet-doublet mass splitting than the usual SD model to satisfy Relic Density.
- Both thermal and non-thermal production of RHN has been done for $\Delta N_{\rm eff}$ analysis.

Thank You !!!