## <span id="page-0-0"></span>Probing the photon emissivity of thermal QCD matter on the lattice

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Lattice@CERN2024 workshop, CERN, 8 July 2024





Harvey Meyer [Photon emissivity of QCD matter](#page-37-0)

This talk is mainly based on the following publications:

- 1. 2001.03368 (PRD): Rate of photon production in the quark-gluon plasma from lattice QCD, Marco Cè, Tim Harris, HM, Aman Steinberg, Arianna Toniato.
- 2. 2205.02821 (PRD): Photon emissivity of the quark-gluon plasma: A lattice QCD analysis of the transverse channel, Marco Cè, Tim Harris, Ardit Krasniqi, HM, Csaba Török.
- 3. 2309.09884 (PRD): Probing the photon emissivity of the quark-gluon plasma without an inverse problem in lattice QCD, Marco Cè, Tim Harris, Ardit Krasniqi, HM, Csaba Török.

#### Brief advertisement for our recent preprint:

4. 2407.01657: Hot QCD matter around the chiral crossover: A lattice study with  $O(a)$ -improved Wilson fermions, Ardit Krasniqi, Marco Cè, Renwick Hudspith, HM.

### Motivation (I): heavy-ion collision phenomenology

- $\triangleright$  direct photons (those not produced via hadronic decays) encode information about the environment in which they were created
- If for  $1 \text{GeV} \leq p_T \leq 2 \text{GeV}$ : expect a dominant contribution from (quasi-)thermal photons: quark-gluon plasma and hadronic phase

ALICE [2308.16704]

 $\blacktriangleright$  Used measurement of  $e^+e^$ pairs with  $m_{ee} < 30$  MeV: fraction of direct/inclusive dielectrons is the same as for photons.



### Photon emissivity and thermal vector spectral functions

$$
\rho^{\mu\nu}(\mathcal{K}) = \int d^4x \, e^{i\mathcal{K}\cdot x} \, \frac{1}{Z} \sum_n e^{-E_n/T} \langle n | [j^\mu(x), j^\nu(0)] | n \rangle
$$

▶ Rate of dilepton production per unit volume plasma: [McLerran, Toimela 1985]

$$
d\Gamma_{\ell^+\ell^-}(\mathcal{K}) = \alpha^2 \frac{d^4 \mathcal{K}}{6\pi^3 \mathcal{K}^2} \frac{-\rho^{\mu}{}_{\mu}(\mathcal{K})}{e^{\beta \mathcal{K}^0} - 1} \qquad (\mathcal{K}^2 \equiv \omega^2 - k^2)
$$

 $\triangleright$  Rate of photon production per unit volume plasma:

$$
d\Gamma_{\gamma}(\mathbf{k}) = \alpha \frac{d^3k}{4\pi^2 k} \frac{-\rho^{\mu}{}_{\mu}(k,\mathbf{k})}{e^{\beta k} - 1}.
$$

$$
\omega = 0
$$
\n
$$
\omega = k, \ \mathcal{K}^2 < 0
$$
\n
$$
\omega = k
$$

Predictions for  $\sigma(\omega=k)=-\frac{1}{2}$  $\frac{1}{2}\rho^{\mu}{}_{\mu}(k,\bm{k})$  in non-Abelian plasmas:



 $\blacktriangleright$   $\sigma(\omega)$  vanishes in the vacuum;

- ightharpoonup at  $\omega \neq 0$  it vanishes for thermal, non-interacting quarks;
- $\blacktriangleright$  ideal probe of the medium!

Arnold, Moore Yaffe JHEP 11 (2001) 057 and JHEP 12 (2001) 009. AdS/CFT: Caron-Huot et al. JHEP 12 (2006) 015.

### A one-parameter family of expressions for the photon emissivity

 $\blacktriangleright$  at finite temperature, there are two independent,  $\mathsf{O}(3)$  invariant components:  $(k\equiv |\bm{k}|, \quad \hat{k}^i=k^i/k)$ 

$$
\rho_L(\omega,k) \equiv (\hat{k}^i \hat{k}^j \rho^{ij} - \rho^{00}), \qquad \rho_T(\omega,k) \equiv \frac{1}{2} (\delta^{ij} - \hat{k}^i \hat{k}^j) \rho^{ij}.
$$

► current conservation:  $\omega^2 \rho^{00}(\omega, k) = k^i k^j \rho^{ij}(\omega, k)$  implies that  $\rho_L$  vanishes at lightlike kinematics,  $\mathcal{K}^2=0.$ 

Introduce  $(\lambda \in \mathbb{R})$ 

$$
\rho(\omega, k, \lambda) = 2\rho_T + \lambda \rho_L \qquad \stackrel{\lambda=1}{\equiv} -\rho^{\mu}{}_{\mu} .
$$

 $\blacktriangleright$  Photon rate can be written  $(\forall \lambda)$ 

$$
d\Gamma_{\gamma}(\mathbf{k}) = \alpha \frac{d^3k}{4\pi^2 k} \frac{\rho(k,k,\lambda)}{e^{\beta k} - 1}.
$$

### Choosing  $\lambda$ : weak and strong coupling spectral fcts



**Spatial momentum**  $k = \pi T$ : (see hep-th/0607237 and 1310.0164)

- $\rho_T$  is positive-definite and free of the diffusion pole
- $\triangleright$  ( $\rho_T \rho_L$ ) vanishes in the vacuum, is strongly suppressed at large  $\omega$ and obeys a superconvergent sum-rule.
- At  $\omega = k$ , the two channels should be equal: non-trivial consistency check for lattice-based calculations!

### Lattice QCD and vector correlators

Imaginary-time path-integral representation of QFT (Matsubara formalism).

Imaginary-time **vector correlators**  $({\gamma^{\mu}, \gamma^{\nu}}) = 2g^{\mu\nu} = 2diag(1, -1, -1, -1)$ ),

$$
G^{\mu\nu}(x_0,\mathbf{k}) = \int d^3x \; e^{-i\mathbf{k}\cdot\mathbf{x}} \operatorname{Tr}\left\{ \frac{e^{-\beta H}}{Z(\beta)} j^{\mu}(x) \, j^{\nu}(0) \right\}, \qquad j^{\mu} = \sum_f Q_f \, \bar{\psi}_f \gamma^{\mu} \psi_f
$$

**Spectral representation** ( $u$  is a real four-vector):  $\rightsquigarrow$  inverse problem

$$
u_{\mu}G^{\mu\nu}u_{\nu}(x_0,\mathbf{k}) = \int_0^{\infty} \frac{d\omega}{2\pi} \underbrace{\frac{(u_{\mu}\rho^{\mu\nu}u_{\nu})(\omega,\mathbf{k})}{\sinh(\beta\omega/2)}}_{\geq 0} \cosh[\omega(\beta/2 - x_0)].
$$

### Parameters of the lattice calculations

- $N_f = 2$  flavours of dynamical  $O(a)$  improved Wilson fermions with Wilson gauge action; ensembles generated with the openQCDv1.6 code.
- $\triangleright$   $T \simeq 254$  MeV,  $L = 4/T \simeq 3.1$  fm;  $m_{\pi}(T = 0) \simeq 270$  MeV.
- $\blacktriangleright$  Isovector current correlator is computed.



See 2001.03368 ( $\rho_T - \rho_L$ ), and 2205.02821 ( $\rho_T$ ).

## Continuum extrapolation of the  $\rho_T - \rho_L$  Euclidean correlator



## Analysis of the  $(\rho_T - \rho_L)$  channel

'Hydrodynamics' prediction at small  $\omega, k$ : with D the diffusion coefficient,

$$
\rho(\omega, k, -2)/\omega \approx \frac{4\chi_s \, Dk^2}{\omega^2 + (Dk^2)^2} \qquad \omega, k \ll D^{-1}.
$$

From the operator-product expansion:

$$
\rho(\omega, k, \lambda = -2) \stackrel{\omega \to \infty}{\sim} k^2/\omega^4 \; : \qquad \int_0^\infty d\omega \, \omega \, \rho(\omega, k, -2) = 0.
$$

 $\rightsquigarrow$  5-parameter ansatz:

$$
\rho(\omega, k, -2) = \frac{A(1 + B\omega^2) \tanh(\omega\beta/2)}{[(\omega - \omega_0)^2 + b^2][( \omega + \omega_0)^2 + b^2][\omega^2 + a^2]}.
$$

Analysis strategy: always determine  $B$  so as to satisfy the sum rule; scan over all other parameters to determine the  $\chi^2$  landscape.

For an early Backus-Gilbert analysis, see Harris, Steinberg, Brandt, Francis, HM 1710.07050.



← spec.funct. poorly constrained at small  $k \simeq 0.4$  GeV

> Representative spectral functions describing the lattice  $2(T - L)$  correlator

←− spec.funct. better constrained at  $k \approx 2\pi T \simeq 1.6$  GeV

### Final result of analysis of the  $(\rho_T - \rho_L)$  channel

$$
D_{\text{eff}}(k) \equiv \frac{\rho(\omega = k, k, \lambda)}{4\chi_s k}, \qquad \chi_s = \beta \int d^3x \langle V_0(x) V_0(0) \rangle.
$$



Cè et al. 2205.02821 (PRD).

### Comparison of NLO+LPM spectral function to  $T - L$  lattice data



- $\blacktriangleright$  Continuum-extrapolated quenched data from  $N_t = 20, 24, 30$ .
- $\blacktriangleright$  General behaviour reproduced, but differences are visible.

2403.11647 Ali, Bala, Francis, Jackson, Kaczmarek, Turnwald, Ueding, Wink

# The transverse channel

Cè et al. 2205.02821 (PRD).

## Taking the continuum limit of  $G_T(\tau, k)$



Three different discretisations, joint continuum extrapolation.

Use treelevel improvement.

### Euclidean correlator: lattice vs. NLO prediction



- $\triangleright$  NLO prediction lies a few percent higher than the lattice data.
- $\blacktriangleright$  (For this comparison we set  $\chi_s^{\text{AdS/CFT}} = \frac{N_c^2 T^2}{8}$  $\frac{1}{8}T^2 = \frac{9T^2}{8}$  $\frac{1}{8}$ .)

### (I) Backus-Gilbert spectral function



$$
\frac{\bar{\rho}(\omega,k)}{f(\omega)} = \int_0^\infty d\omega' \, \Delta(\omega,\omega') \, \frac{\rho(\omega')}{f(\omega')} = \sum_i g_i(\omega) G_{\rm T}(\tau_i).
$$

Choice made here:  $f(\omega)=\frac{\omega^2}{\tanh(\beta\omega/2)}$  in order to make  $\frac{\rho(\omega')}{f(\omega')}$  a 'slowly varying function', since the BG method is exact if  $\frac{\rho(\omega')}{f(\omega')}$  is constant.

### (II) Fit ansätze for the spectral functions

$$
\rho(\omega) = \rho_{\text{fit}}(\omega)(1 - \Theta(\omega, \omega_0, \Delta)) + \rho_{\text{pert}}(\omega)\Theta(\omega, \omega_0, \Delta)
$$

with  $\omega_0 \approx 2.5$  GeV the matching frequency,

$$
\Theta(\omega, \omega_0, \Delta) = (1 + \tanh[(\omega - \omega_0)/\Delta])/2
$$

a smooth step function and  $\rho_{\rm pert}(\omega)$  from [Jackson, Laine 1910.09567].

A) Polynomial ansatz:

$$
\frac{\rho_{\text{fit},1}(\omega)}{T^2} = \sum_{n=0}^{N_{\text{p}}-1} A_n \left(\frac{\omega}{\omega_0}\right)^{1+2n},
$$

B) Piecewise polynomial ansatz:

$$
\frac{\rho_{\text{fit},2}(\omega)}{T^2} = \begin{cases} A_0 \frac{\omega}{\omega_0} + A_1 \left(\frac{\omega}{\omega_0}\right)^3, & \text{if } \omega \leq k, \\ B_0 \frac{\omega}{\omega_0} + B_1 \left(\frac{\omega}{\omega_0}\right)^3, & \text{if } \omega > k. \end{cases}
$$

### Representative lattice-QCD results for the spectral functions



- $\triangleright$  Piecewise polynomial cannot 'decide' between having a min. or max. at  $\omega = k$ .
- $\triangleright$  Qualitatively, both the 'AdS/CFT' type and the 'NLO' type are compatible with the lattice data.

### Final result for the photon emissivity



- $\triangleright$  most fits give a larger photon emissivity than the weak-coupling prediction, but overall the lattice result is still compatible with it.
- $\triangleright$  the AdS/CFT prediction is also consistent with the lattice data for  $k \geq \pi T/2$ .

2205.02821 Török et al.

# Probing the photon emissivity without an inverse problem

Cè et al. 2309.09884 (PRD, Editor's Suggestion).

### A dispersion relation for a Euclidean correlator at zero virtuality

In Let  $\sigma(\omega) \equiv \rho_T(\omega, |\mathbf{k}| = \omega)$  be the spectral function proportional to the photon emissivity;

In let  $H_E(\omega_n) \equiv G_E(\omega_n, k = i\omega_n)$  the momentum-space Euclidean correlator with Matsubara frequency  $\omega_n = 2\pi T n$ and *imaginary* spatial momentum  $k = i\omega_n$ ;

 $\triangleright$  once-subtracted dispersion relation:  $(\sigma(\omega) \sim \omega^{1/2}$  at weak coupling)

$$
H_E(\omega_n) = -\frac{\omega_n^2}{\pi} \int_0^\infty \frac{d\omega}{\omega} \frac{\sigma(\omega)}{\omega^2 + \omega_n^2}
$$

In these energy-moments of  $\sigma(\omega)$  are directly accessible without involving an inverse problem.

HM, 1807.00781 (EPJC).

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### How different energy-moments probe  $\sigma(\omega)$

$$
H_E(\omega_n) - H_E(\omega_r) = \int_0^\infty \frac{d\omega}{\pi} \, \omega \left[ \frac{1}{\omega^2 + \omega_n^2} - \frac{1}{\omega^2 + \omega_r^2} \right] \sigma(\omega).
$$



 $\blacktriangleright$   $H_E(\omega_1)$  receives a sizeable contribution from the soft photons;

 $\blacktriangleright$   $H_E(\omega_2) - H_E(\omega_1)$  probes the emission of hard photons.

### Formulation on the lattice

Our standard representation: [Cè, ... HM, 2112.00450]

$$
H_E(\omega_n) = -\int_0^\beta dx_0 \int d^3x \left(e^{i\omega_n x_0} - e^{i\omega_n x_2}\right) e^{\omega_n x_3} \langle j_1(x) j_1(0) \rangle.
$$

The  $e^{i\omega_n x_2}$  term subtracts a static contribution which vanishes in the continuum (for the same reason the polarisation tensor component  $\Pi_{11}(q)$  vanishes at lightlike virtuality for  $q_1 = 0$ ).

- **In** This representation has the advantage that  $H_E(\omega_n)$  vanishes exactly in the vacuum even at finite lattice spacing;
- $\blacktriangleright$  as a consequence, cutoff effects at finite temperature are strongly reduced.

NB. similar task to computing hadronic vacuum polarisation in muon  $(q - 2)!$ 

**Computing**  $H_E(\omega_1)$ 



 $N_f = 2$ ,  $T = 254$  MeV,  $N_t = 16, 20, 24$ : Monte-Carlo chains were extended.

## Result for  $H_E(\omega_1)$ : comparison to strong/weak-coupling predictions



 $\blacktriangleright$   $-H_E(\omega_1)$  turns out to be smaller than the value predicted from the AMY spectral function  $\sigma(\omega)$ ; is it due to its peak around  $\omega = 0$ ?

 $\triangleright$  for  $H_E(\omega_2)$ : better control of the large  $x_3$  regime needed.

Cè et al, 2309.09884 (PRD).

## **Conclusion**

- $\triangleright$  Photon rate: first lattice calculation in dynamical QCD with continuum limit.
- $\blacktriangleright$  The transverse and the transverse-minus-longitudinal channels show small but statistically significant deviations from the state-of-the-art weak-coupling calculations at  $T = 254$  MeV;
- $\blacktriangleright$  Treatment of the inverse problem with Backus-Gilbert method or with an explicit fit ansatz for  $\sigma(\omega)$  leads to consistent results for the photon emissivity for  $\pi T \leq \omega \leq 2\pi T$ .
- $\blacktriangleright$  Those results are still compatible with weak-coupling prediction and AdS/CFT.
- $\blacktriangleright$  Dispersion relation at fixed photon virtuality  $q^2=0$ :  $H_E(\omega_1)$  is lower than the value derived from the weak-coupling spectral function  $\sigma(\omega)$ ; the perturbative uncertainty is now the limiting factor.

### Thermal modification of vector spectral function [Krasniqi et al. 2407.01657]

- **D** Quasi physical-mass  $N_f = 2 + 1$  simulations,  $m_\pi(T = 0) = 130$  MeV; parameters of CLS ensemble E250,  $L = 6.1$  fm,  $a \simeq 0.064$  fm.
- 3 ensembles:  $N_t = 24, 20, 16$  corresponding to temperatures 128, 154 and 192 MeV.
- ightharpoonup among other observables, we studied (at  $p = 0$ ; isovector vector and axial-vector channels)

$$
G(x_0, \mathbf{p}, T) - \sum_{n \in \mathbb{Z}} G(|x_0 + n\beta|, \mathbf{p}, 0)
$$
  
= 
$$
\int_0^\infty d\omega \left( \underbrace{\rho(\omega, \mathbf{p}, T) - \rho(\omega, \mathbf{p}, 0)}_{\equiv \Delta \rho(\omega, \mathbf{p}, T)} \right) \frac{\cosh[\omega(\frac{\beta}{2} - x_0)]}{\sinh[\frac{\beta}{2}\omega]}.
$$

Used Backus-Gilbert method with rescaling function  $f(\omega) = \omega$  to constrain  $\Delta \rho(\omega, \mathbf{p}, T)$ .

NB.  $\Delta \rho(\omega, \mathbf{p}, T) \stackrel{\omega \to \infty}{\sim} 1/\omega^2$ .

### Reminder: vacuum spectral functions from  $\tau$  decays



Fig. from Davier, Höcker, Zhang, doi:10.1103/RevModPhys.78.1043

## Results of BG analysis



Harvey Meyer [Photon emissivity of QCD matter](#page-0-0)

### Spectral sum rules

$$
\int_0^\infty \frac{d\omega}{\omega} \Delta \rho_V(\omega, \mathbf{0}, T) = 0.
$$
 [1107.4388]

**I** generically, a transport peak appears in the spectral function at  $\omega = 0$ , and therefore the contribution of the  $\rho$  meson must be reduced in order to satisfy the sum rule [Brandt, Francis, Jäger, HM 1512.07249]

$$
\int_0^\infty \frac{d\omega}{\omega} \Delta \rho_A(\omega, \mathbf{0}, T) = G_A(\omega_n = 0, \mathbf{p} = 0, T) - G_A(0, \mathbf{0}, 0)
$$
  
=  $f_\pi^2(T = 0) - f_\pi^2(T) + O(m_q^2)$ 

- $\blacktriangleright$   $f_{\pi}(T)$ , the decay constant of the 'static screening pion', falls off monotonously with increasing temperature; it is  $O(m_a)$  in the chirally restored phase.
- $\blacktriangleright$  Thus the RHS

## Ensembles used for computing  $H_E(\omega_1)$  and  $H_E(\omega_2)$



2309.09884

### Motivation: some heavy-ion collision phenomenology

- $\triangleright$  direct photons (those not produced via hadronic decays) encode information about the environment in which they were created
- $\triangleright$  for  $1 \text{GeV} \leq p_T \leq 2 \text{GeV}$ : expect a sizeable contribution from thermal photons: quark-gluon plasma and hadronic phase
- RHIC,  $\sqrt{s_{NN}} = 200$  GeV:
	- **IF PHENIX measurement of**  $p_T < 3$  GeV photons shows clear excess over  $N_{\text{coll}}$ -scaled pp measurement [PRL 104, 132301 (2010)]
	- ▶ PHENIX: large photon anisotropy wrt reaction plane [PRL 109, 122302 (2012)]
	- ▶ STAR: photon yield  $\sim$  3 times smaller than PHENIX: unresolved tension
	- $\blacktriangleright$  PHENIX: higher-statistics study [PRC 109, 044912 (2024)]
- ► LHC,  $\sqrt{s_{NN}} = 2760$  GeV: ALICE measured photon yield [PLB 754, 235 (2016)].
- Find recently, ALICE measured dileptons at  $\sqrt{s_{NN}} = 5020$  GeV, in particular determining direct photon yield [2308.16704]

Phenomenology: Gale et al. [2106.11216]. Review G. David [1907.08893].

### Lattice papers on the photon rate

- ▶ Karsch, Laermann, Petreczky, Stickan, Wetzorke 2002; S. Gupta 2004; Aarts, Allton, Foley, Hands, Kim 2007: quenched calculations,  $k = 0$ .
- **IF** hep-lat/0610061 (LAT06): Aarts, Allton, Foley, Hands: quenched,  $k \neq 0$
- ▶ 1012.4963 (PRD): Ding, Francis, Kaczmarek, Karsch, Laermann, Soeldner, quenched calculation with continuum limit,  $k = 0$ .
- 1212.4200 (JHEP): Brandt, Francis, HM, Wittig:  $N_f = 2$ ,  $N_t = 16$ ,  $k = 0$ ,  $m_{\pi} = 270$ ,  $T = 250$ MeV.
- $\triangleright$  1307.6763 (PRL), 1412.6411 (JHEP): Aarts, Allton, Amato, Giudice, Hands, Skullerud:  $N_f = 2 + 1$ ,  $k = 0$ , anisotropic, fixed-scale temperature scan,  $m_\pi = 384$  MeV
- ▶ 1512.07249 (PRD): Brandt, Francis, Jäger, HM,  $N_f = 2$ ,  $k = 0$ ,  $N_t = 12 \rightarrow 24$ ,  $m_{\pi} = 270$ , fixed-scale scan across the phase transition.
- ▶ 1604.07544 (PRD): Ghiglieri, Kaczmarek, Laine, F. Meyer: quenched calculation with continuum limit,  $k \neq 0$ .
- 2001.03368 (PRD): Cè, Harris, HM, Steinberg, Toniato,  $N_f = 2$  calculation with continuum limit at  $T = 250$ MeV,  $k \neq 0$ ; transv-minus-longitud. channel.
- $\triangleright$  2205.02821 (PRD): idem, but transverse channel.
- ▶ 2309.09884 (PRD): Cè, Harris, Krasniqi, HM, Török, calculation of energy-moments at zero virtuality.
- $\triangleright$  2403.11647: Ali, ... Francis, Kaczmarek et al.: analysis of transv-minus-longitud. channel.

#### Results for the effective diffusion coefficient  $D_{\text{eff}} \equiv \frac{\rho_T(\omega=k,k)}{2\chi_s k}$  $2\chi_s k$



- I left panel: comparison with results of  $(T L)$  channel analysis: piecewise-polynomial ansatz with max. at  $\omega = k$  is disfavoured for  $k \geq \pi T$ .
- right panel: forbidding a max. at the  $1\sigma$  level, predictivity is much stronger.
- polynomial ansatz favours values even larger than the  $\mathcal{N} = 4$  SYM prediction from AdS/CFT (hep-th/0607237).

### Representation through non-static screening masses

$$
\tilde{G}_E(\omega_r, x_3) = -2 \int_0^\beta dx_0 e^{i\omega_r x_0} \int dx_1 dx_2 \langle J_1(x)J_1(0) \rangle = \sum_n |A_n^{(r)}|^2 e^{-E_n^{(r)}|x_3|}
$$
\n
$$
\Rightarrow H_E(\omega_r) \equiv \int_0^\infty dx_3 \, \tilde{G}_E(\omega_r, x_3) e^{\omega_r x_3} = 2\omega_r^2 \sum_{n=1}^\infty |A_n|^2 \frac{1}{\zeta_n^{(r)} \zeta_n^{(r)} \zeta_n^{(r)} \zeta_n^{(r)}}.
$$

$$
\frac{H_E(\omega_r)}{=O(g^2)} = \int_{-\infty}^{\infty} dx_3 \, \sigma_E(\omega_r, x_3) \, e^{-2 \pi i \omega_r} \sum_{n=0}^{\infty} \underbrace{P_n^{(n)}(E_n^{(r)}(E_n^{(r)}(2-\omega_r^2))}_{=O(g^{-2})}.
$$

This helps explain the connection observed in [Brandt et al, 1404.2404] between non-static screening masses and the LPM-resummation contributions to the photon emission rate [Aurenche et al, hep-ph/0211036].

### <span id="page-37-0"></span>Sketch of the (standard) derivation of the dispersion relation

$$
G_R(\omega,k) = i(\delta_{il} - \frac{k_ik_l}{k^2}) \int d^4x \; e^{i\mathcal{K}\cdot x} \theta(x^0) \left\langle [j^i(x), j^l(0)] \right\rangle.
$$
 But

$$
[j^{\mu}(x), j^{\nu}(0)] = 0
$$
 for  $x^{2} < 0$ ,

 $\Rightarrow$  the retarded correlator  $H_R(\omega) \equiv G_R(\omega, k = \omega)$  at lightlike momentum is analytic for Im  $(\omega) > 0$ . Similarly, the advanced correlator  $H_A(\omega)$  is analytic for Im  $(\omega) < 0$ .

Define the function  $H(\omega) = \begin{cases} H_R(\omega) & \text{Im}(\omega) > 0 \\ H_R(\omega) & \text{Im}(\omega) < 0 \end{cases}$  $H_A(\omega)$  Im  $(\omega) > 0$ <br> $H_A(\omega)$  Im  $(\omega) < 0$ 

It is analytic everywhere, except for a discontinuity on the real axis:

$$
H(\omega + i\epsilon) - H(\omega - i\epsilon) = H_R(\omega) - H_A(\omega) = i\sigma(\omega),
$$

Write a Cauchy contour-integral representation (using two half-circles) of  $H(\omega)$ just above the real axis, where it coincides with  $H_R(\omega)$ :

$$
H_R(\omega) = H_R(\omega_r) + \int_{-\infty}^{\infty} \frac{d\omega'}{2\pi} \sigma(\omega') \left[ \frac{1}{\omega' - \omega - i\epsilon} - \frac{1}{\omega' - \omega_r - i\epsilon} \right].
$$

The dispersion relation for the Euclidean correlator follows from the observation  $G_E(\omega_n,k^2)=G_R(i\omega_n,k^2)$  $n > 0$ .