

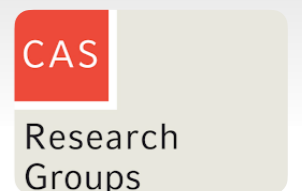
Disentangling tensions from systematics with CLONES

(Constrained LOcal & NEsting Environment Simulations)

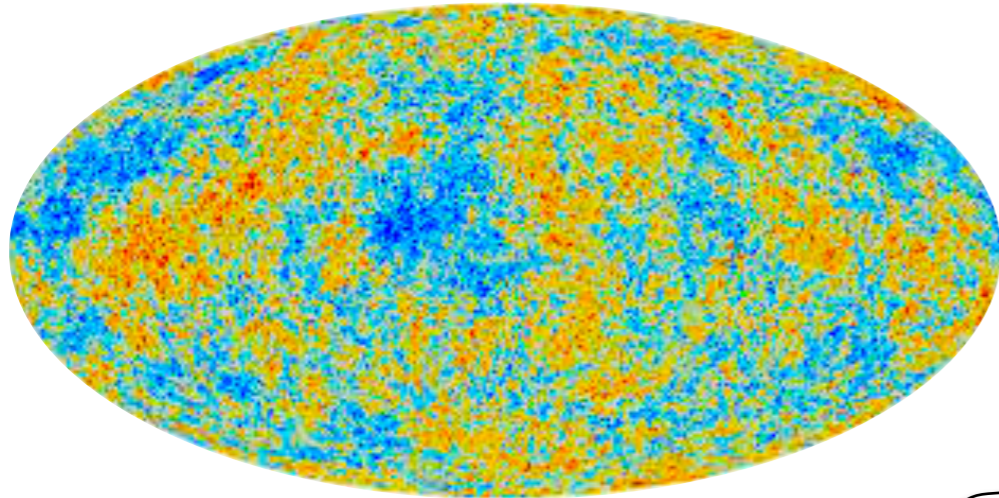
Jenny Sorce
and many collaborators

Researcher at CRIStAL, Lille & Associate Researcher at IAS, Orsay &
Guest researcher at AIP, Potsdam & CAS fellow at LMU, Munich

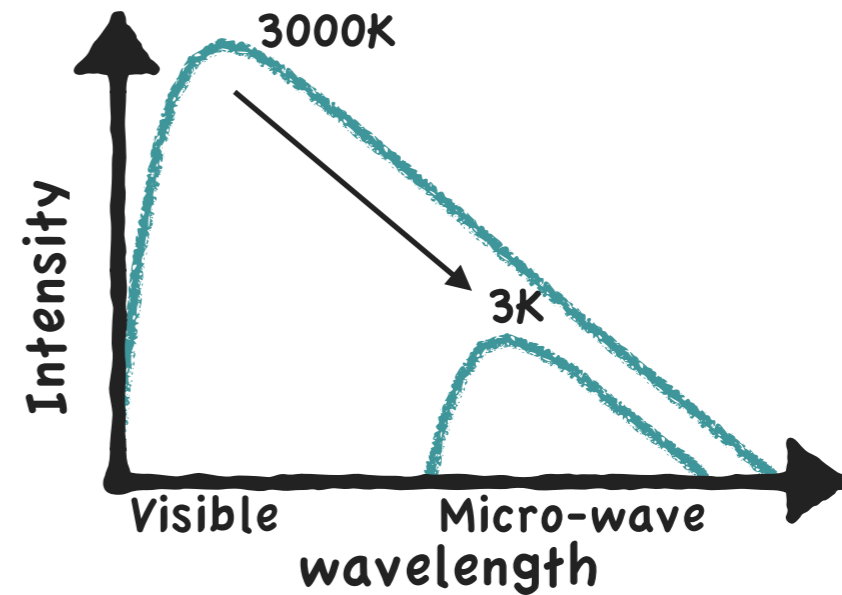
Tensions in Cosmology, Corfu - September 7th, 2023



Cosmic Microwave Background

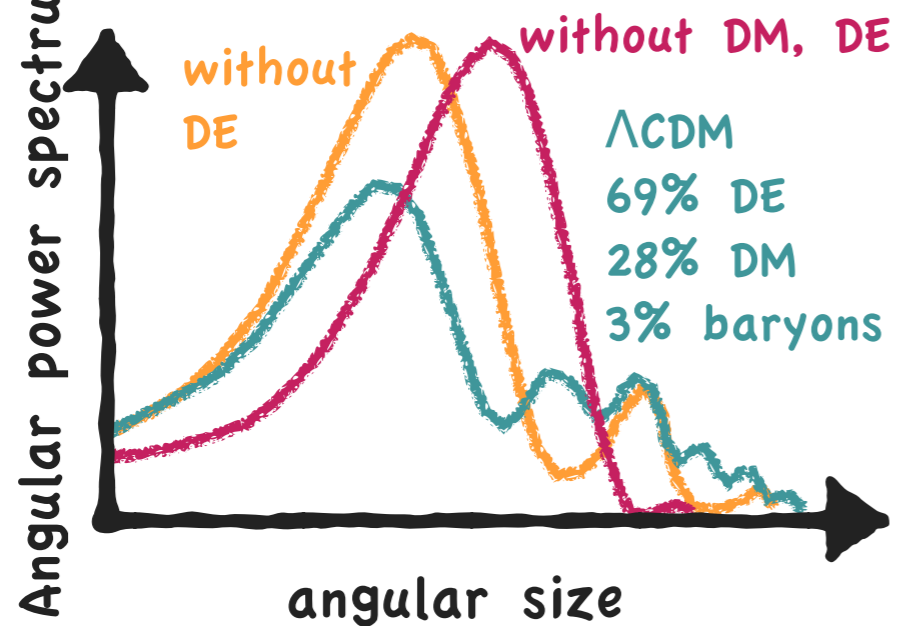


Universe expansion

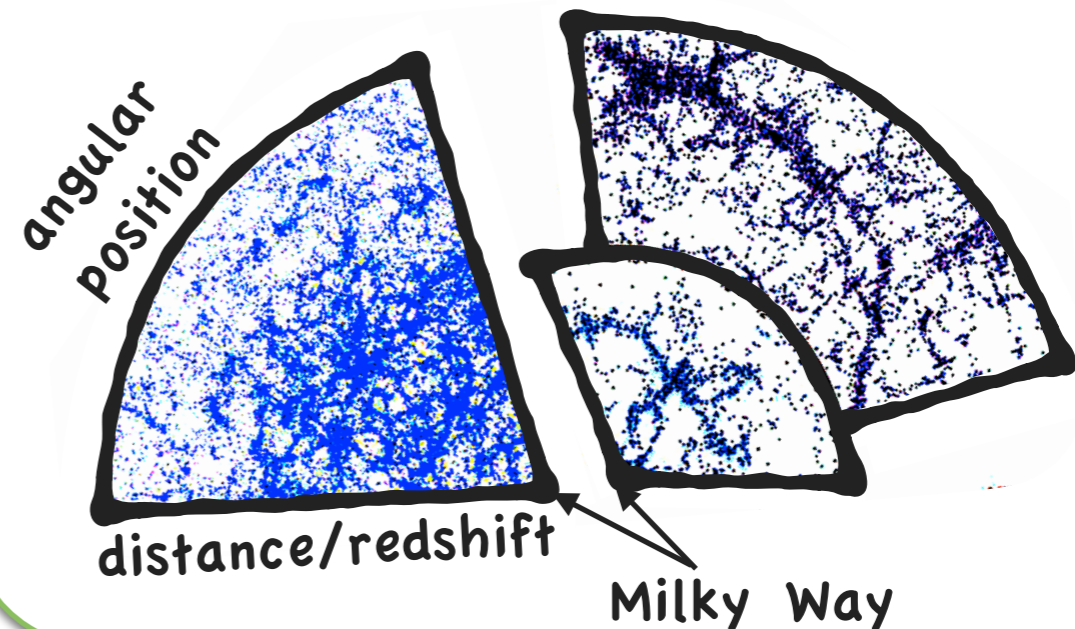


Λ CDM

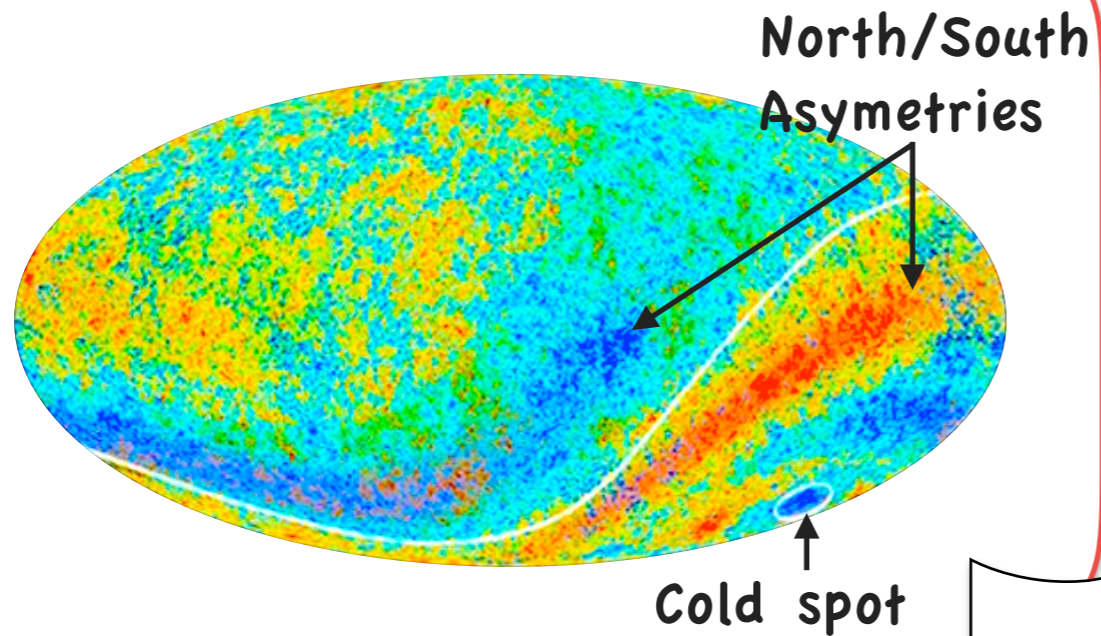
Universe content



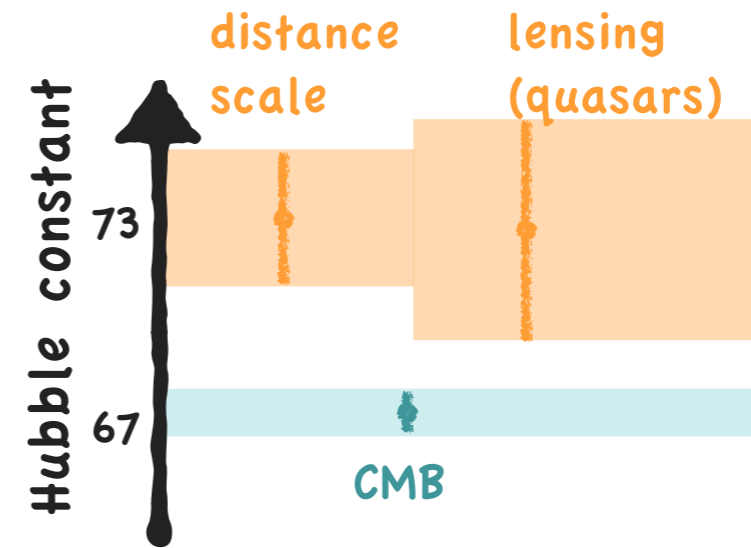
Cosmic Web and galaxies



Anomalies in the CMB

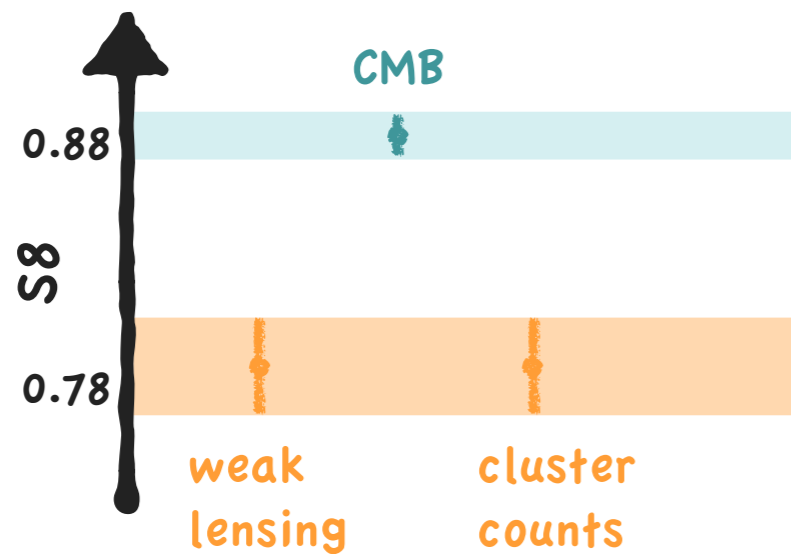


Universe expansion rate (H_0)

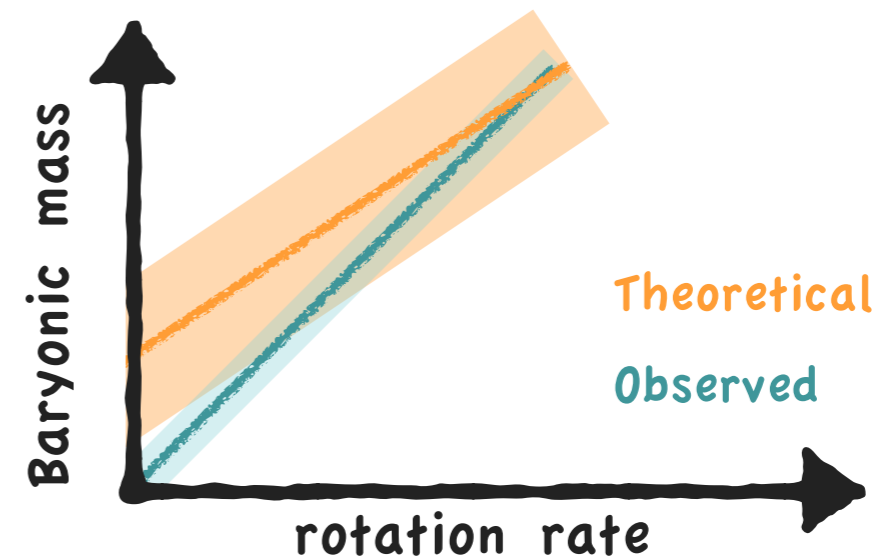


~~**Λ CDM**~~

S8 (σ_8, Ω_m)

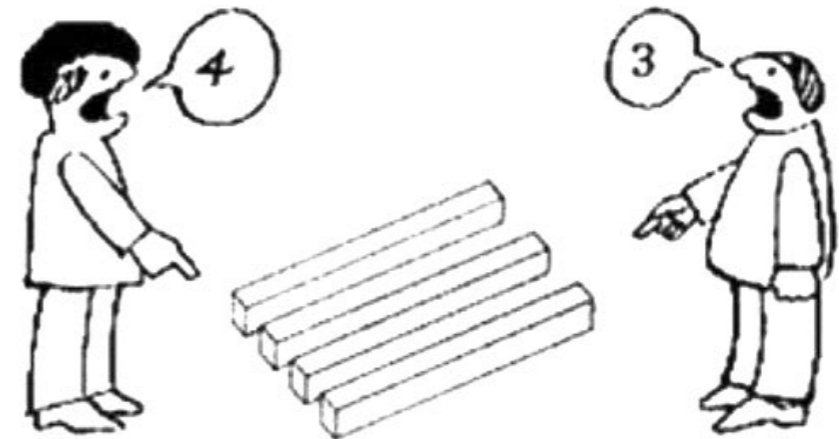


Galaxy properties



New physics or biases/ systematics?

Are we a neutral observer? Are our surveys free of cosmic variance?



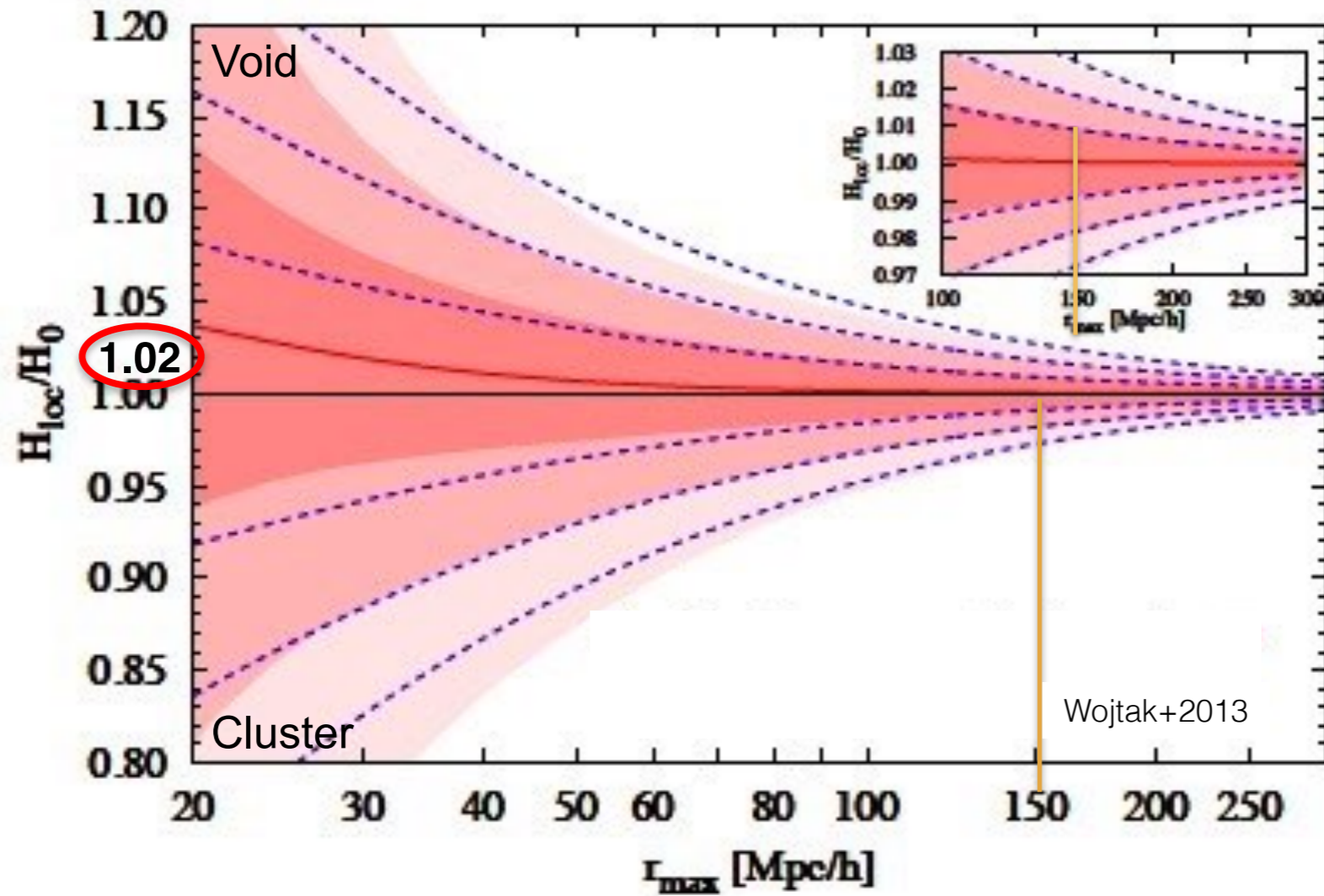
Are we comparing apple-to-apple?
Do we understand enough?



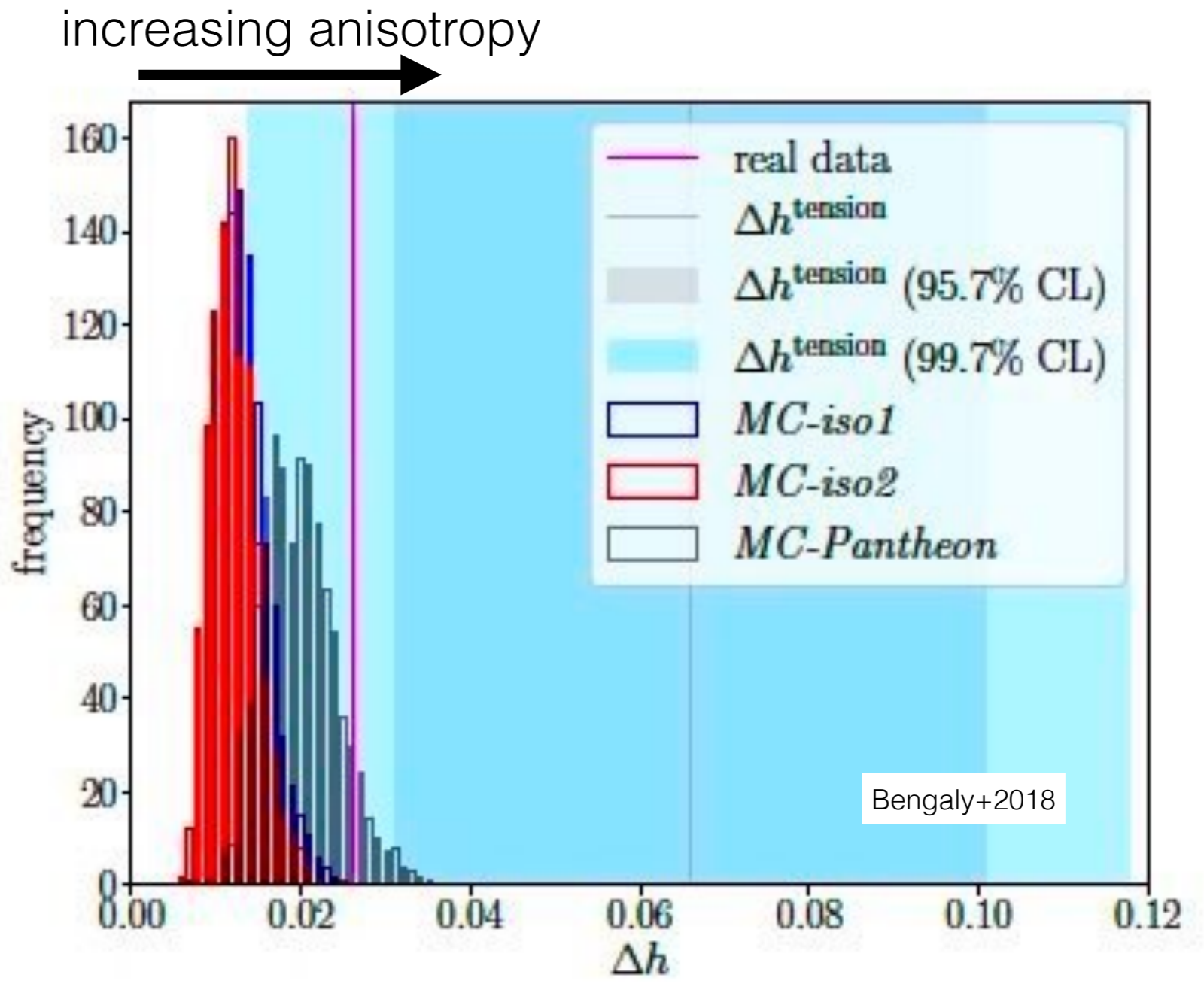
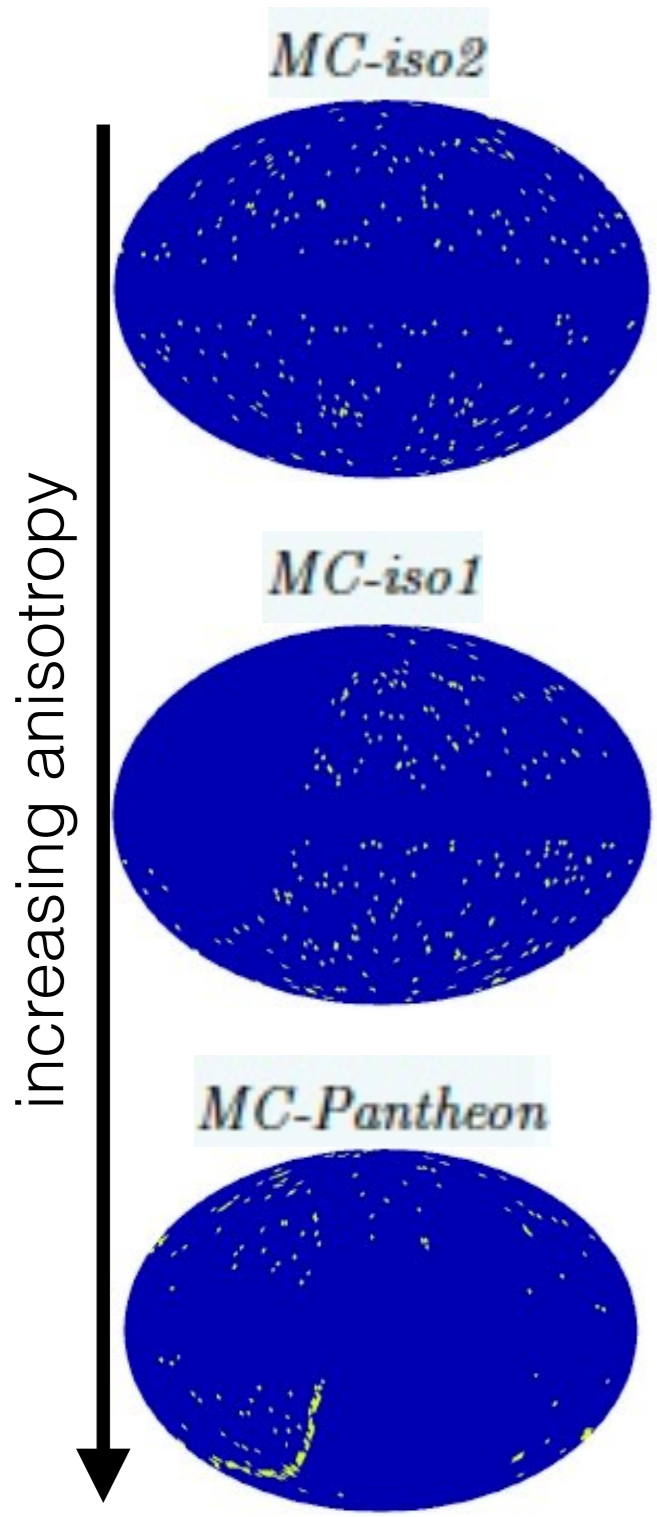
What about foreground and
evolution effects?



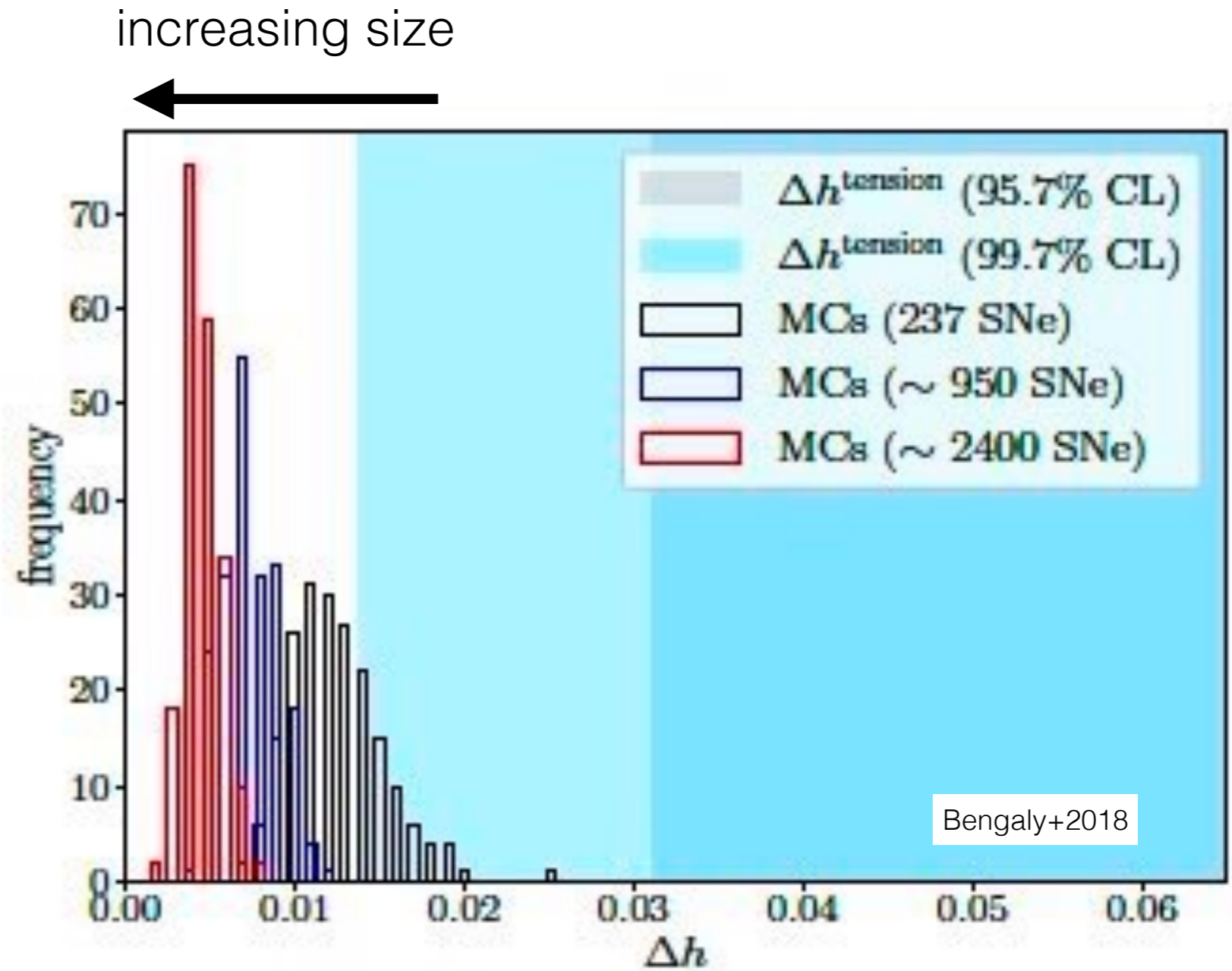
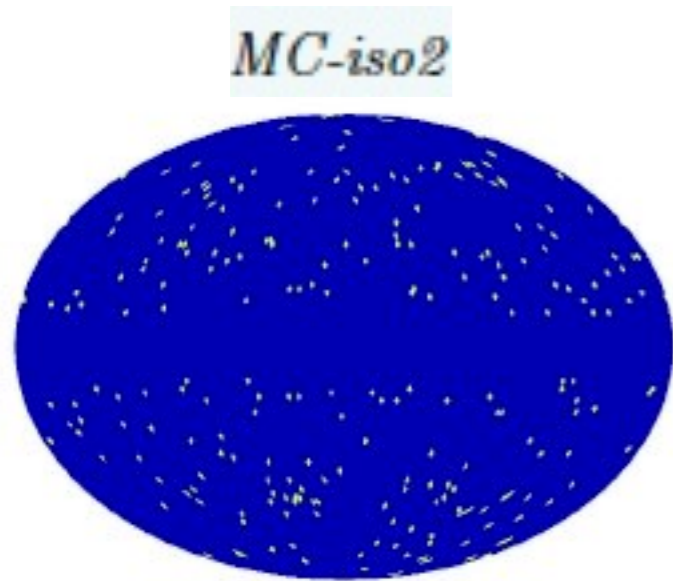
As many effects on values as environments



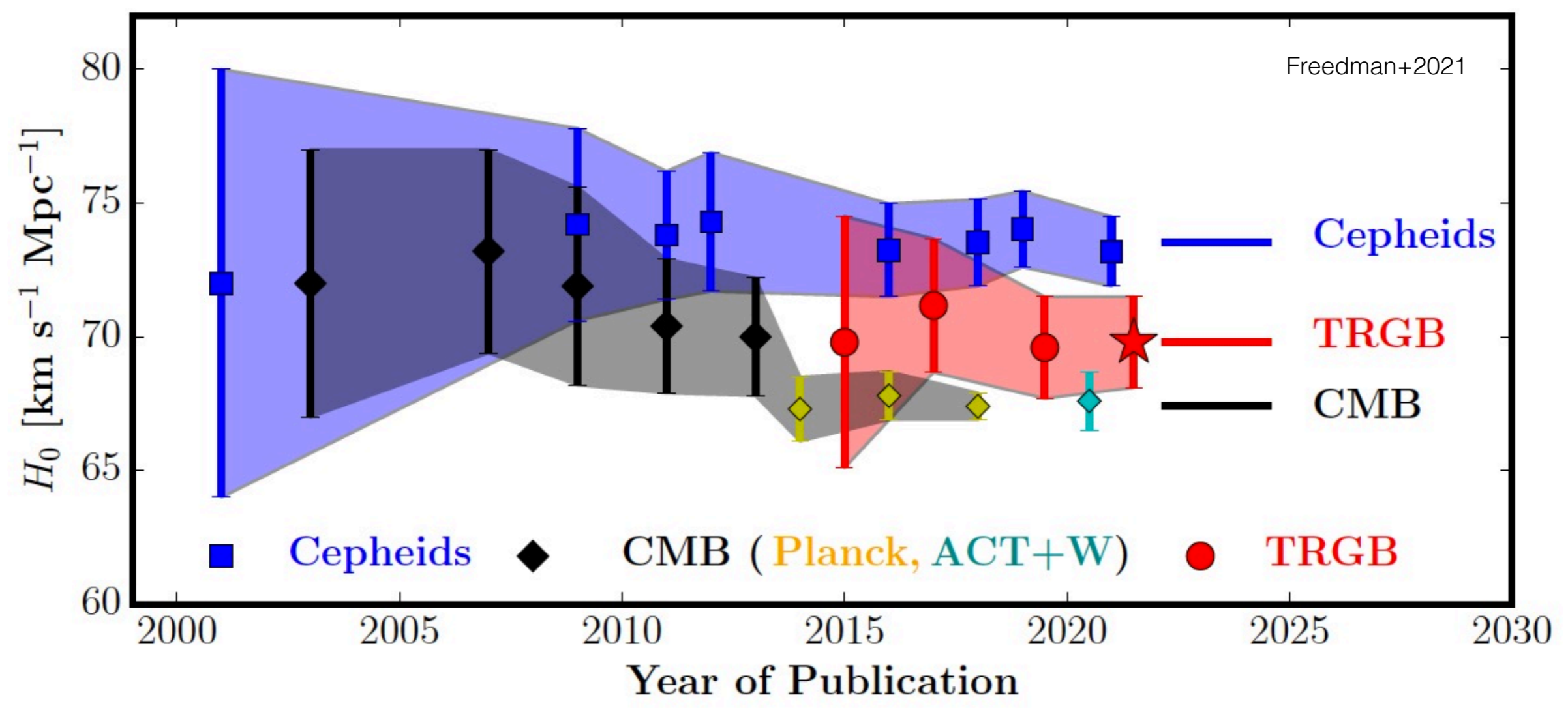
For an average environment: a 2% bias !



For an average survey: a 1-2% bias !



For a survey size divided by 10: a 1-2% bias !



For different calibrators: a 5% difference !

local density

For an average environment: **a 2% bias !**

survey anisotropy

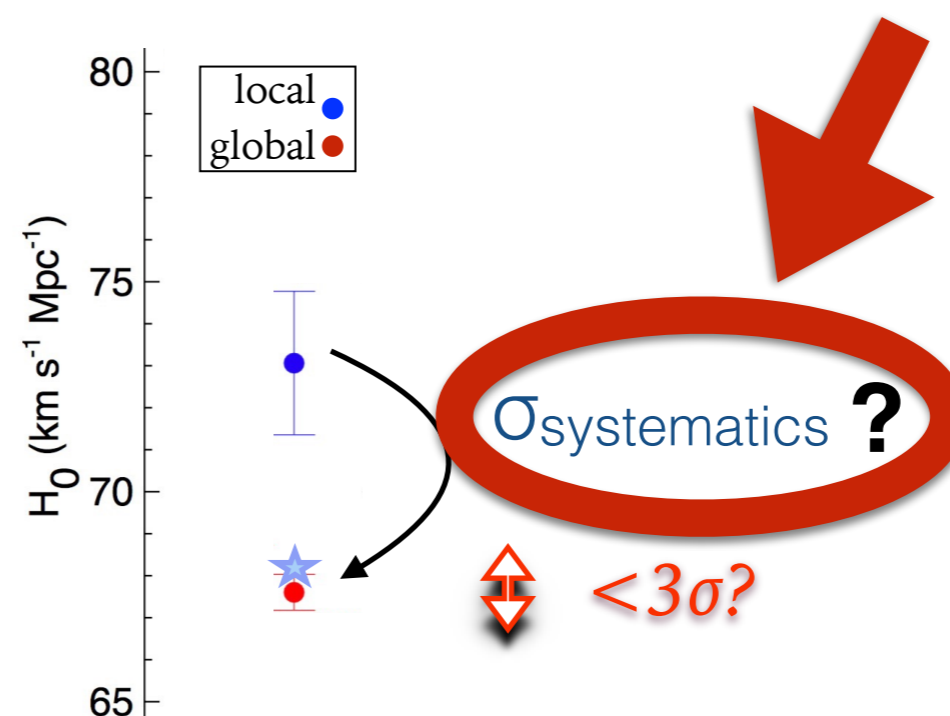
For an average survey: **a 1-2% bias !**

survey size

For a survey size divided by 10: **a 1-2% bias !**

calibrator nature

For different calibrators: **a 5% difference !**

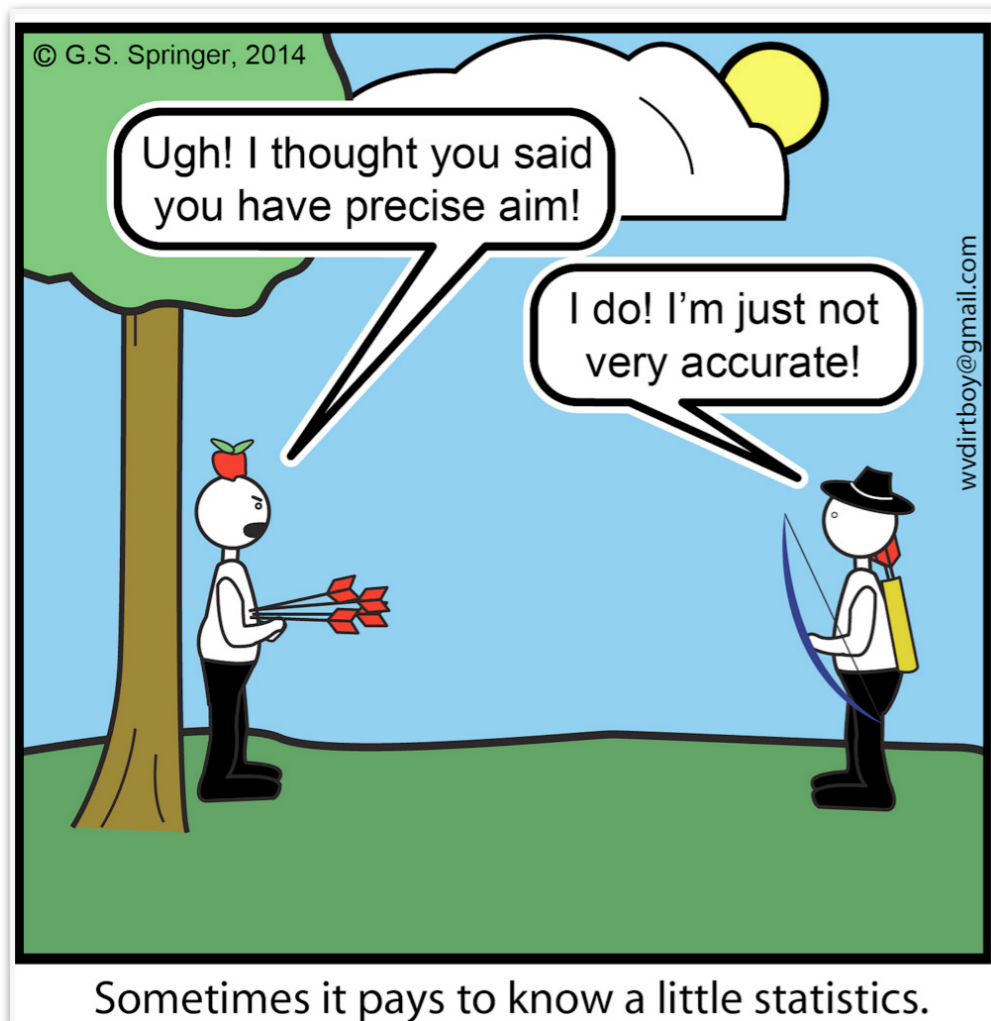


→ Λ CDM is not (yet) ruled out

$$H_0 = X \pm \sigma_{\text{measurement}} \pm \sigma_{\text{systematics}}$$

- nb measurements
- instruments/tools sensitivity
= precision

Standard cosmological simulations can give the total uncertainty but cannot disentangle tensions and systematics

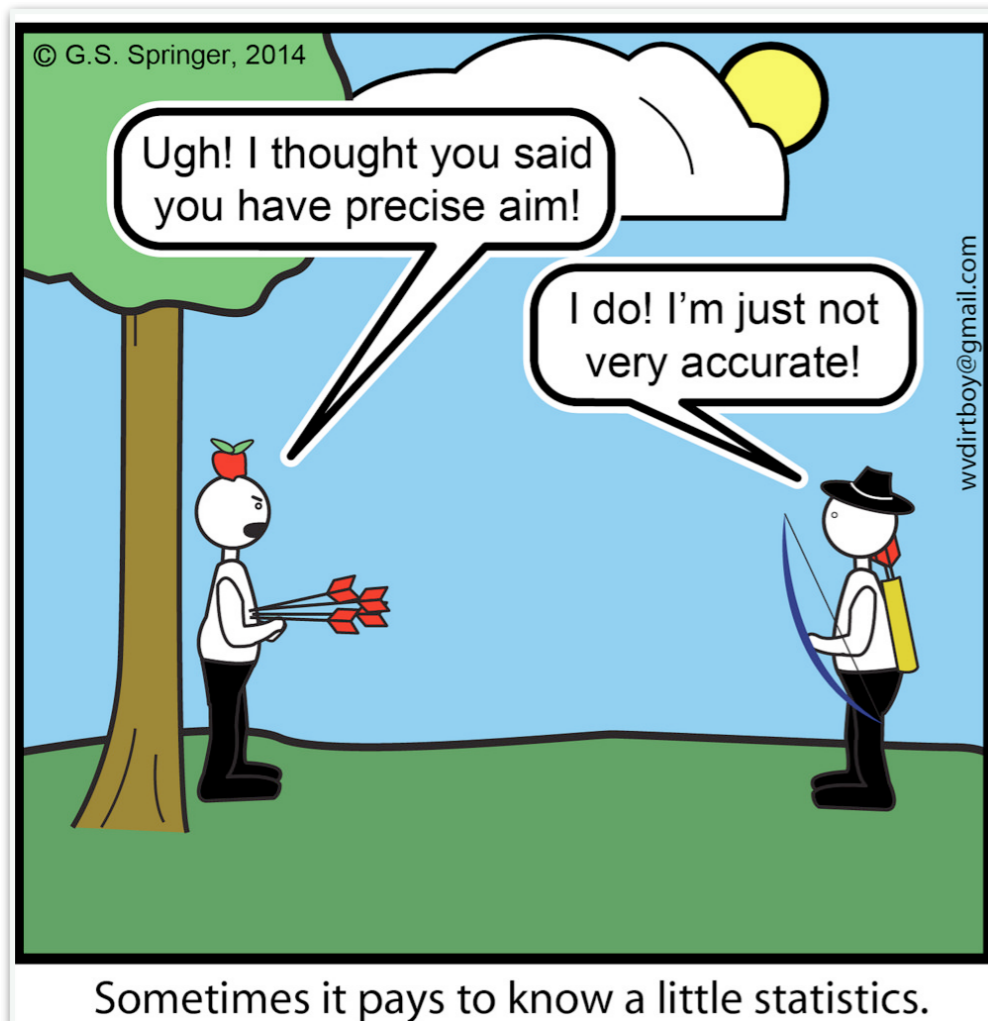


$$H_0 = X \pm \sigma_{\text{measure}} \pm \sigma_{\text{systematics}}$$

- nb measurements
- instruments/tools sensitivity = precision

Standard cosmological simulations can give the total uncertainty but cannot disentangle tensions and systematics

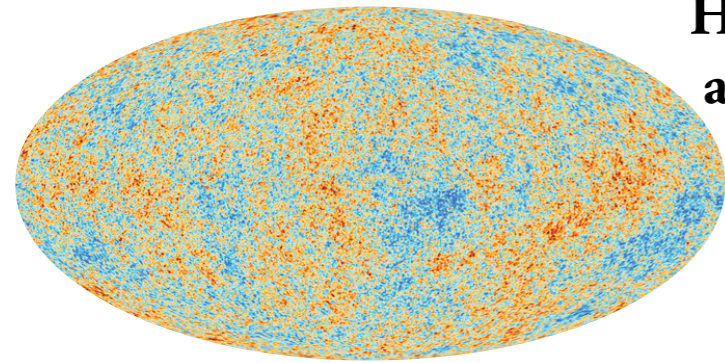
Constrained cosmological simulations can help disentangling real tensions from systematics (by reproducing our particular case)



Standard cosmological simulations

Initial conditions (ICs)

Part of the Universe at
13.7 light-Gyr
Photons received today
have been emitted when it
was ~380 000 yrs. old



Homogeneous
and Isotropic
Universe

→

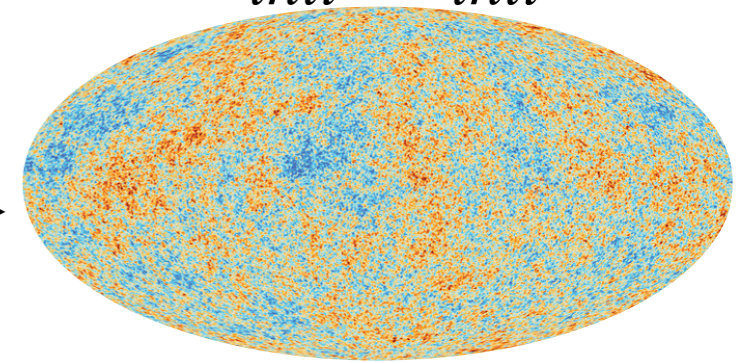
$$P(k)$$

Gaussian
initial density
field

→

$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

→

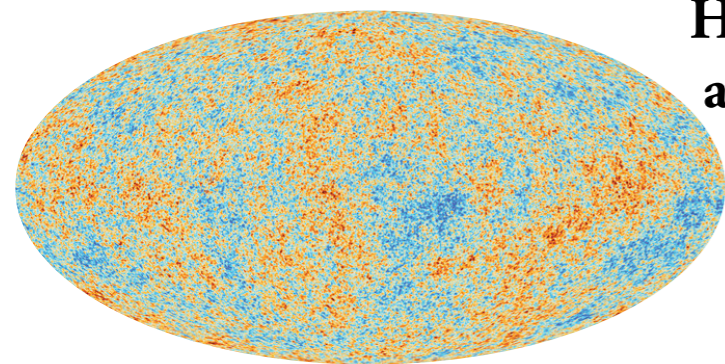


initial conditions of
a random patch of
the Universe
 $\{\delta_{init}, v_{init}\}$

Standard cosmological simulations

Initial conditions (ICs)

Part of the Universe at
13.7 light-Gyr
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Homogeneous
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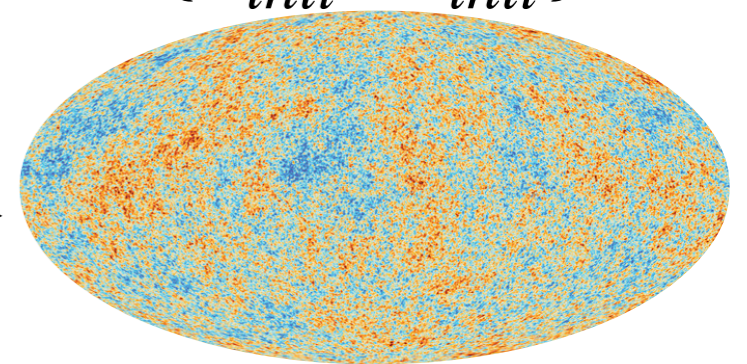
$$P(k)$$

Gaussian
initial density
field

→

$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

→



initial conditions of
a random patch of
the Universe
 $\{\delta_{init}, v_{init}\}$

Linear perturbation
theory (Euler+
Continuity+Poisson)

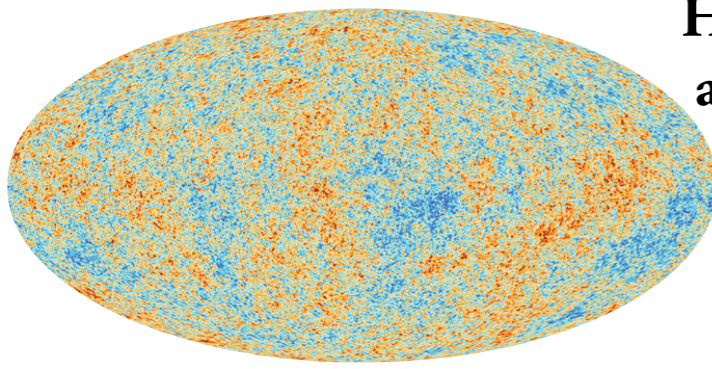
Why only δ ? → $\nabla \cdot v = -\dot{a}f\delta$

NB: only divergent (no tidal) but periodic boundaries

Standard cosmological simulations

Initial conditions (ICs)

Part of the Universe at
13.7 light-Gyr
Photons received today
have been emitted when it
was ~380 000 yrs. old



Homogeneous
and Isotropic
Universe

→

$$P(k)$$

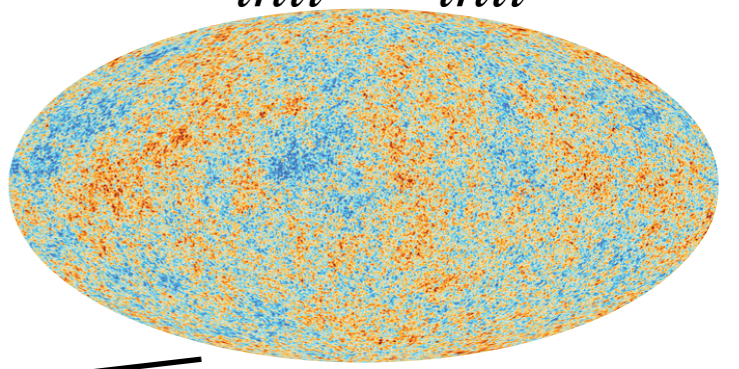
Gaussian
initial density
field

→

$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

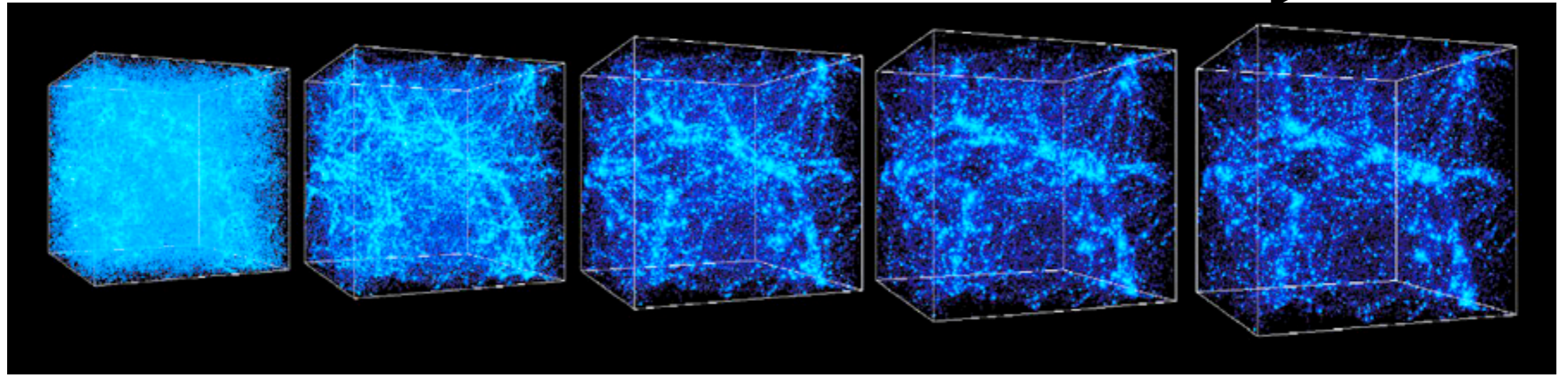
→

initial conditions of
a random patch of
the Universe
 $\{\delta_{init}, v_{init}\}$



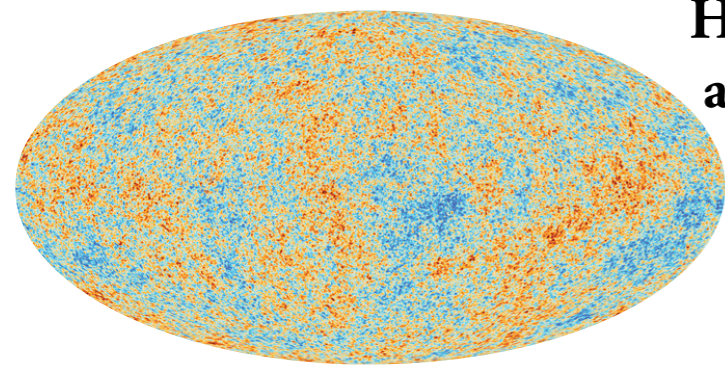
Evolution

Linear perturbation
theory + "kick"



Constrained cosmological simulations

Part of the Universe at
13.7 light-Gyr
Photons received today
have been emitted when it
was $\sim 380\,000$ yrs. old



Homogeneous
and Isotropic
Universe

→

$$P(k)$$

Gaussian
initial density
field

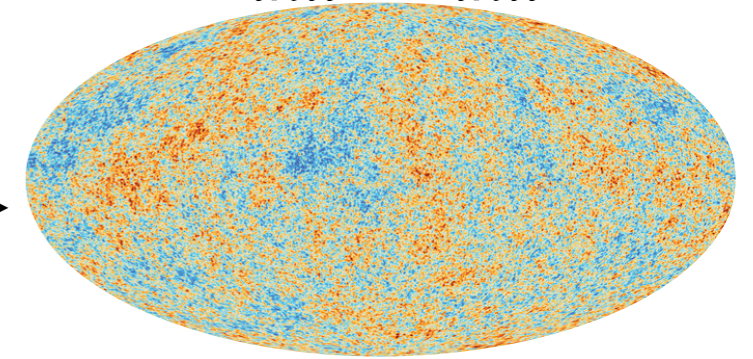
→

$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

?

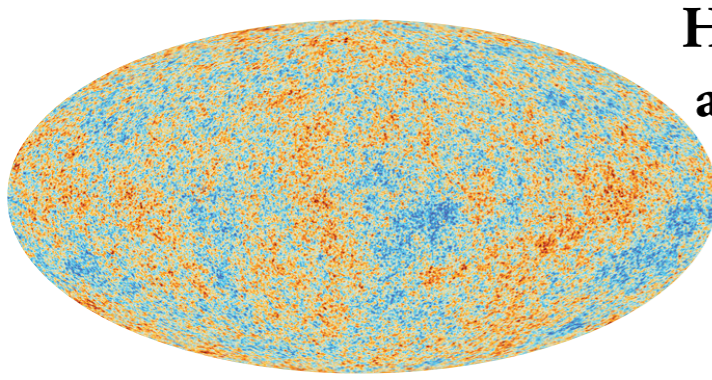
initial conditions of
the local Universe

$$\{\delta_{init}, v_{init}\}$$



Constrained cosmological simulations

Part of the Universe at
13.7 light-Gyr
Photons received today
have been emitted when it
was ~380 000 yrs. old



Homogeneous
and Isotropic
Universe

$$\longrightarrow P(k)$$

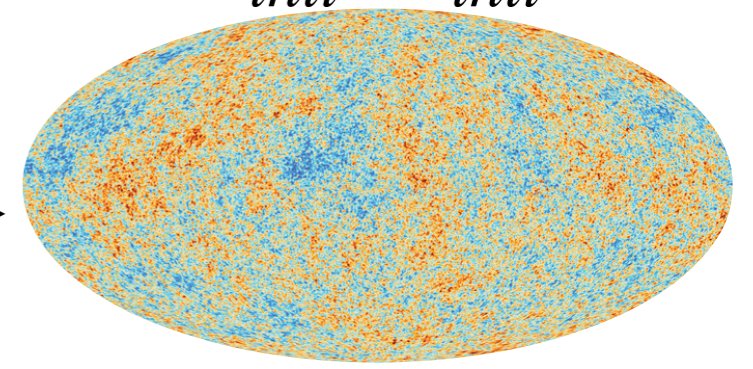
Gaussian
initial density
field

$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

?

initial conditions of
the local Universe

$$\{\delta_{init}, v_{init}\}$$



Type of constraints

Redshift



Λ CDM

Peculiar
velocity

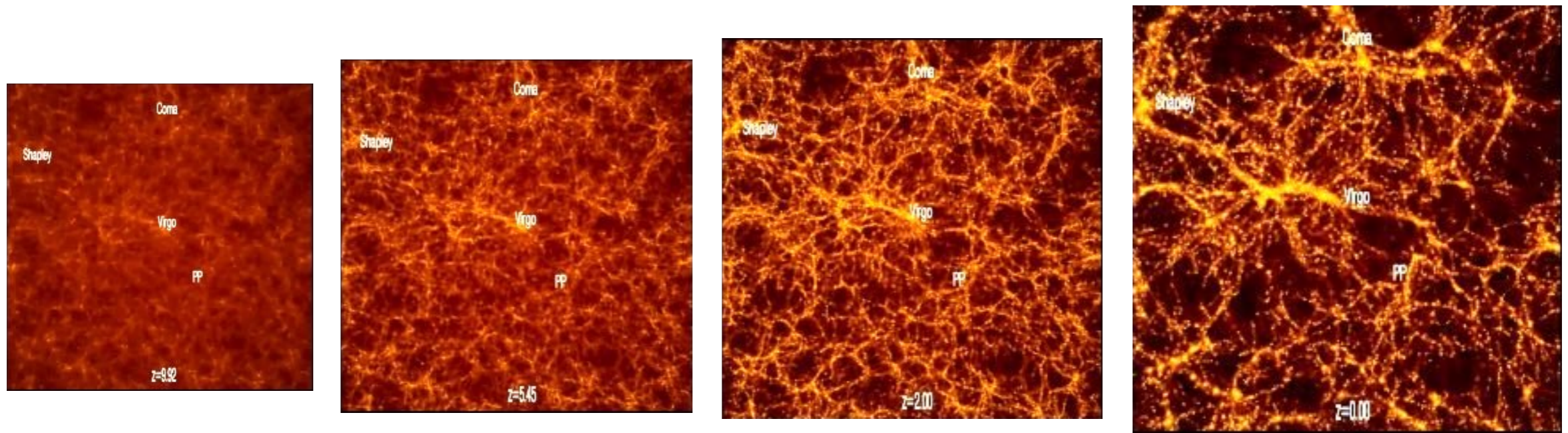


Λ CDM

NB: both with pros and cons!



Constrained cosmological simulations ▶ e.g. CLONES



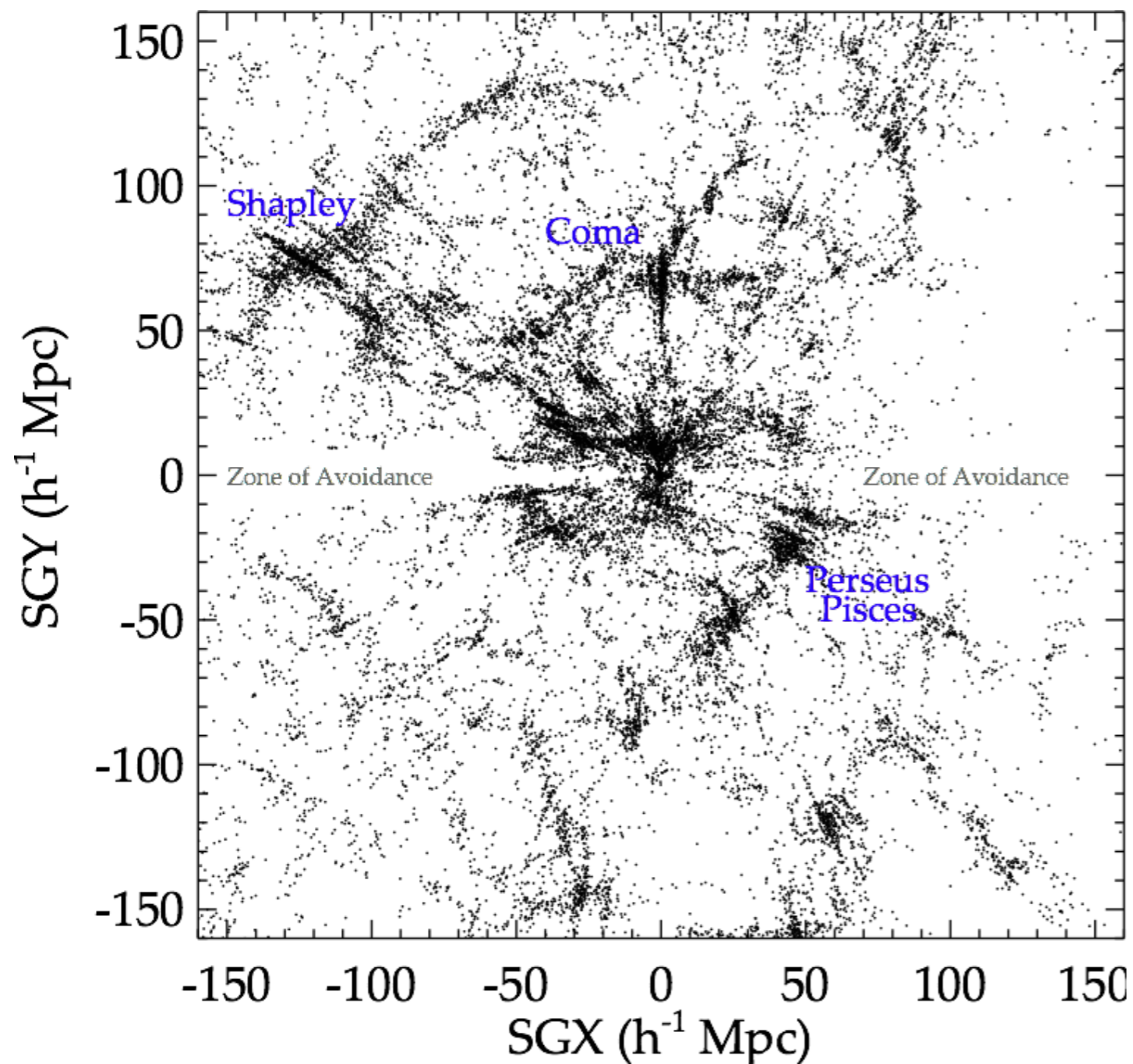
Evolution

Sorce+2016

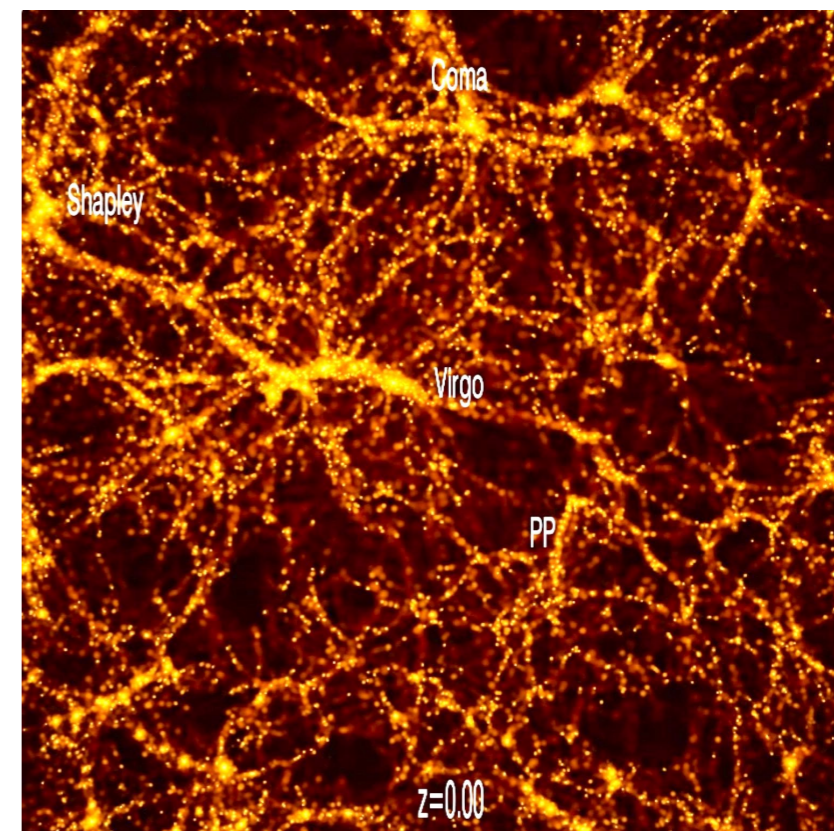
Sorce2018



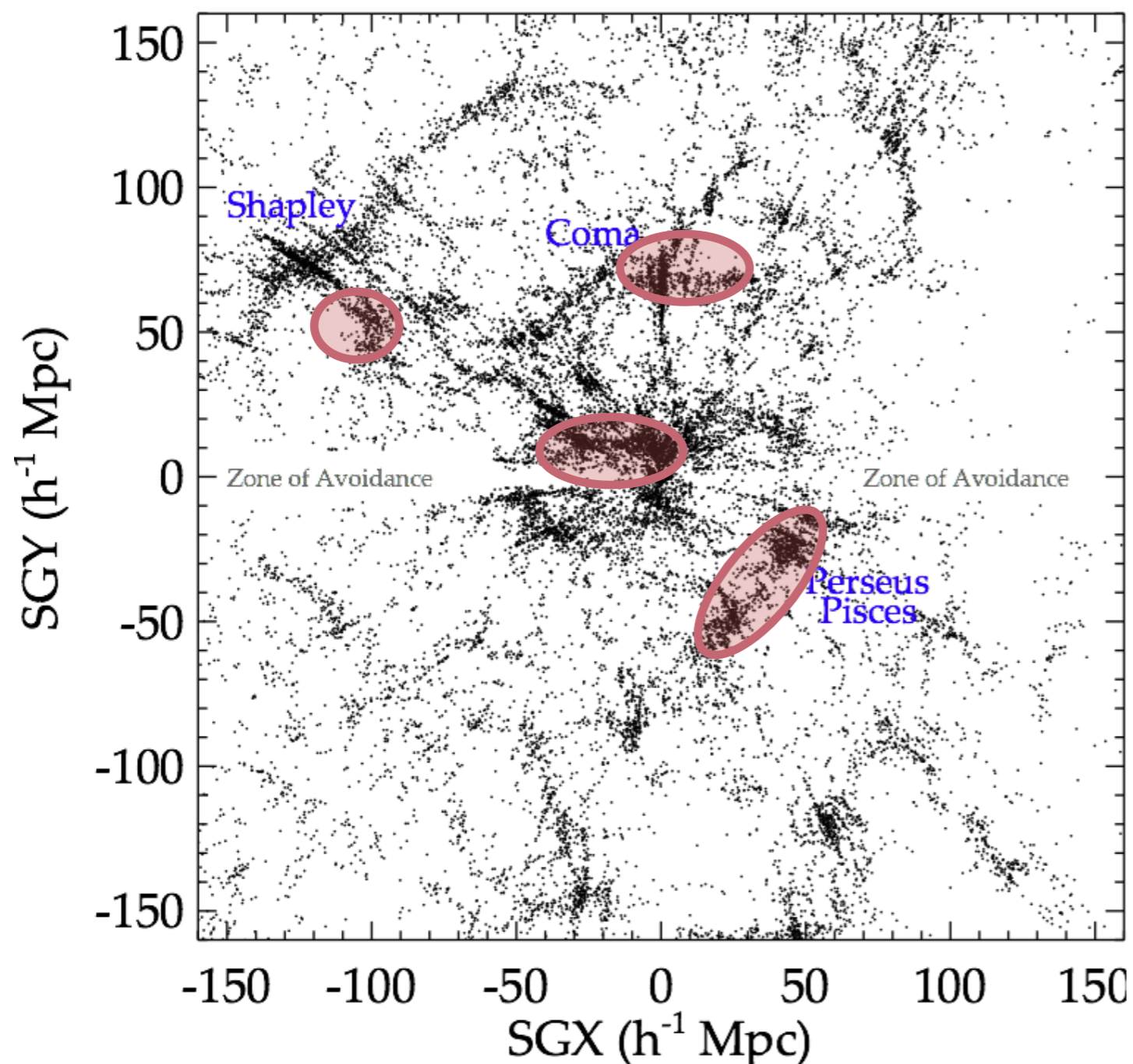
CLONES = Constrained LOcal & Nesting Environment Simulations



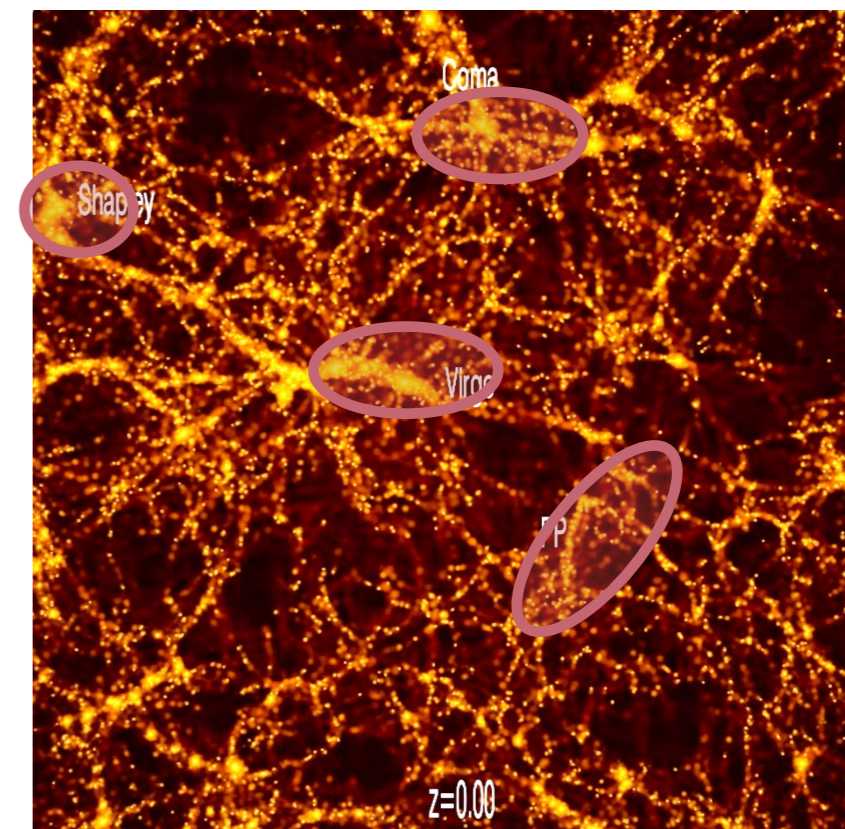
Note the fingers of gods



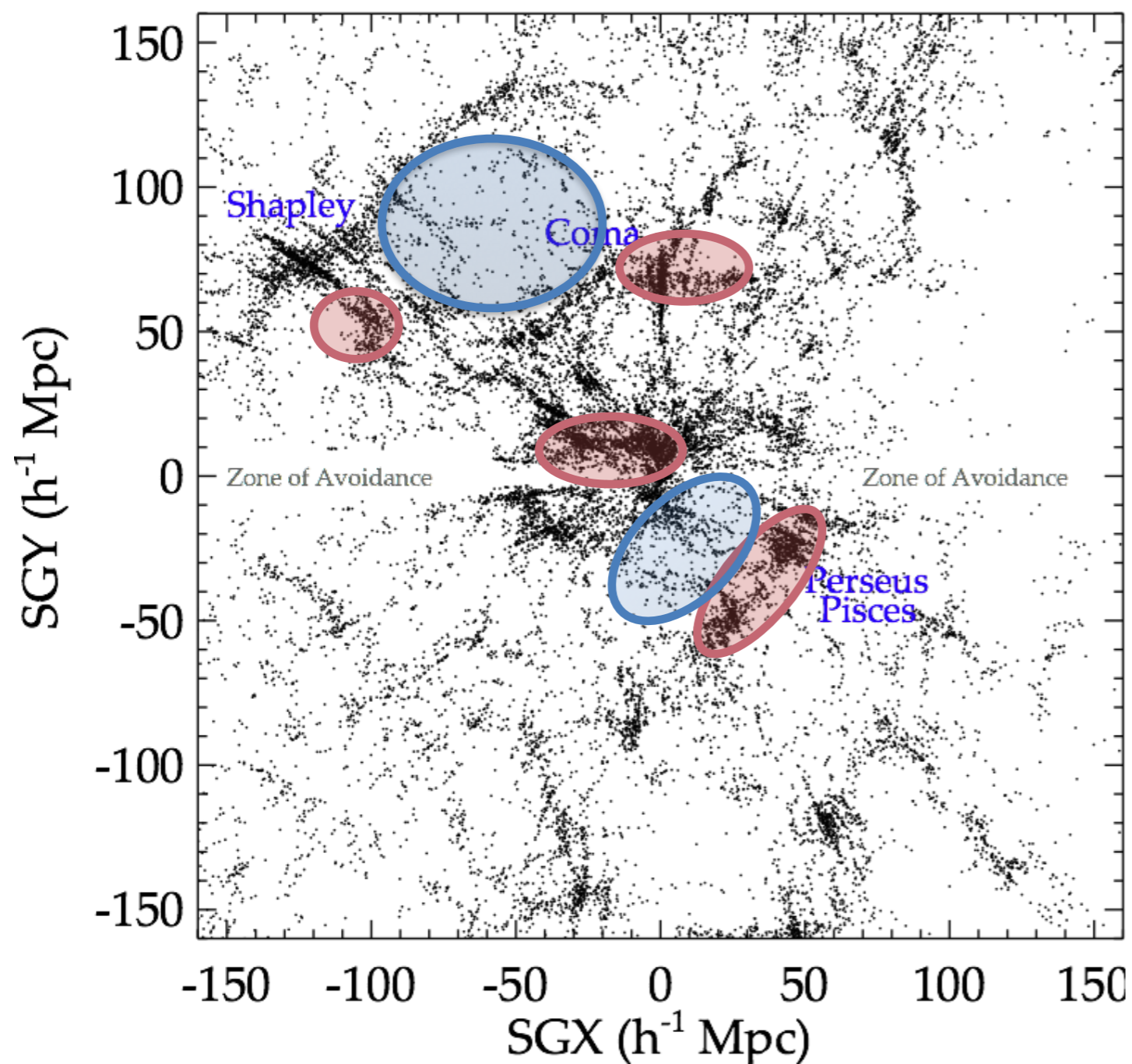
500 Mpc/h, 1024^3 particles,
DM only, Planck cosmology



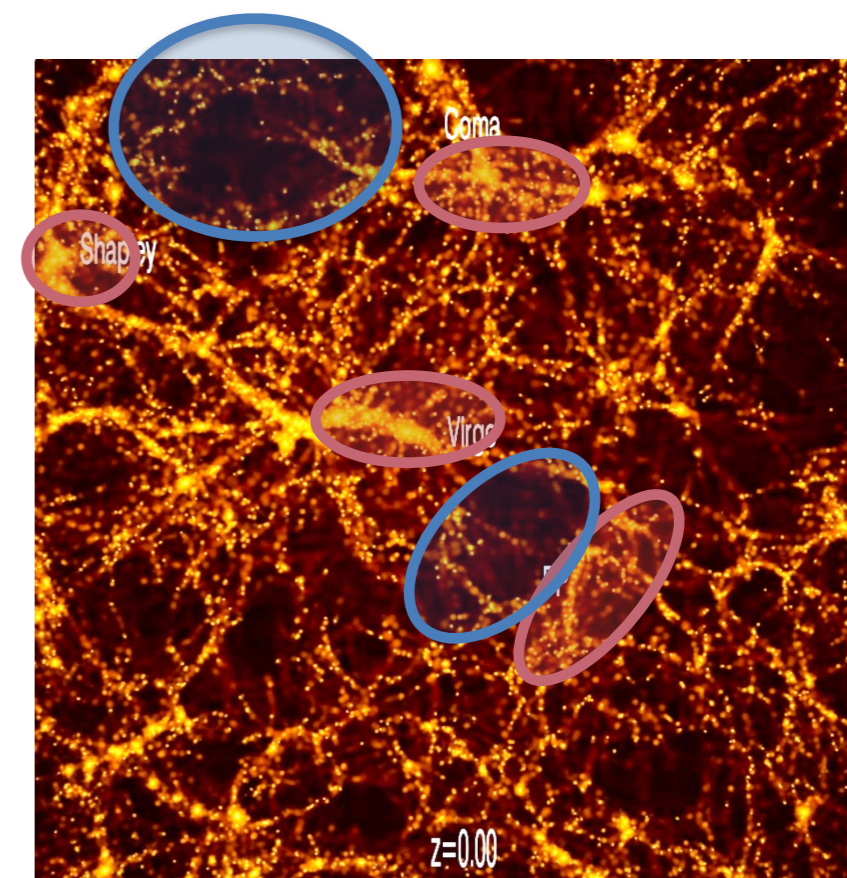
Note the fingers of gods



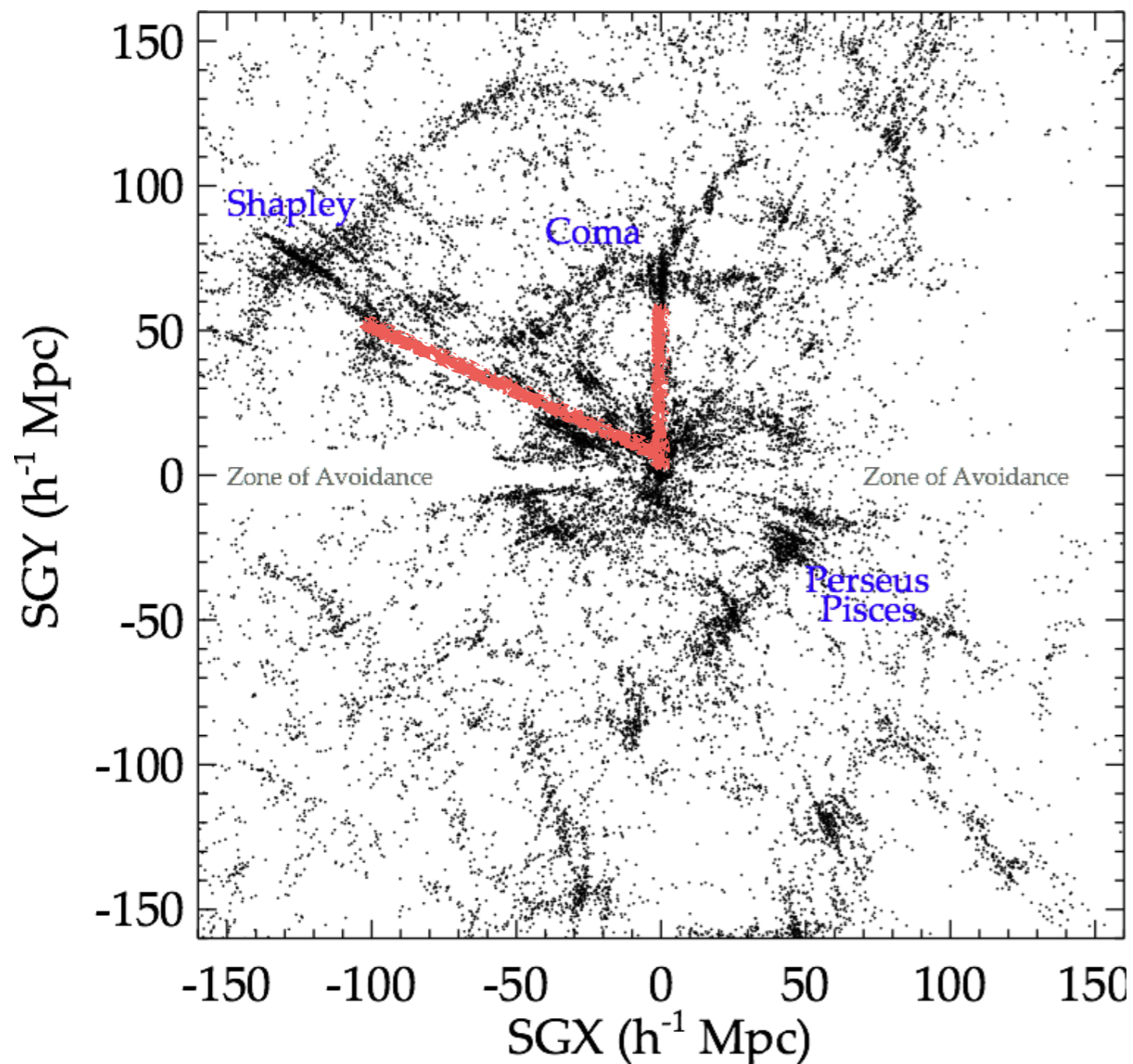
500 Mpc/h, 1024^3 particles, DM only, Planck cosmology



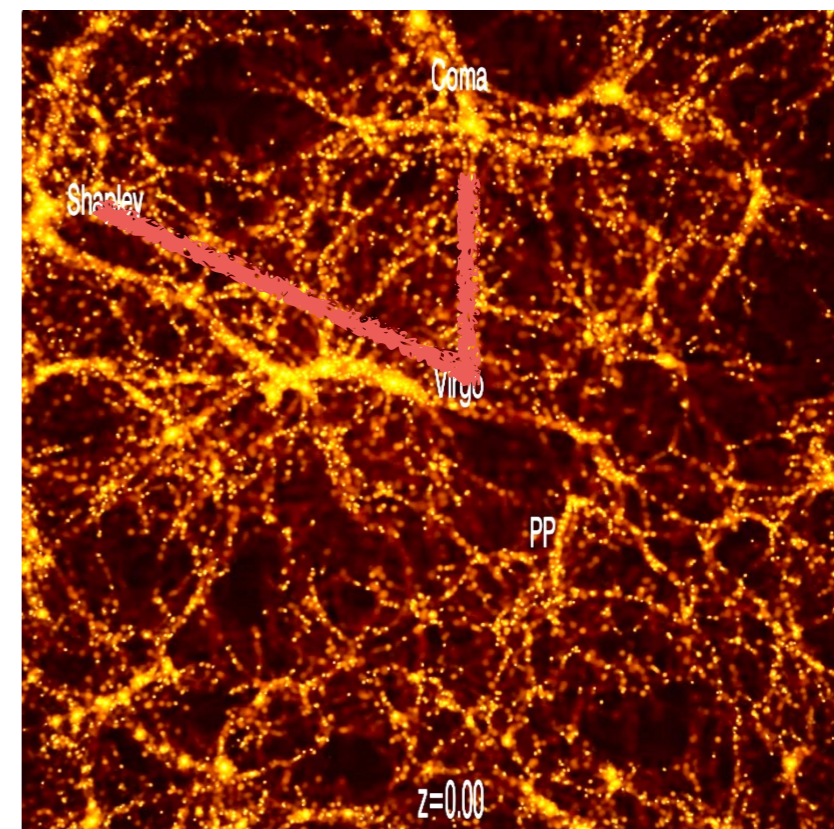
Note the fingers of gods



500 Mpc/h, 1024^3 particles,
DM only, Planck cosmology

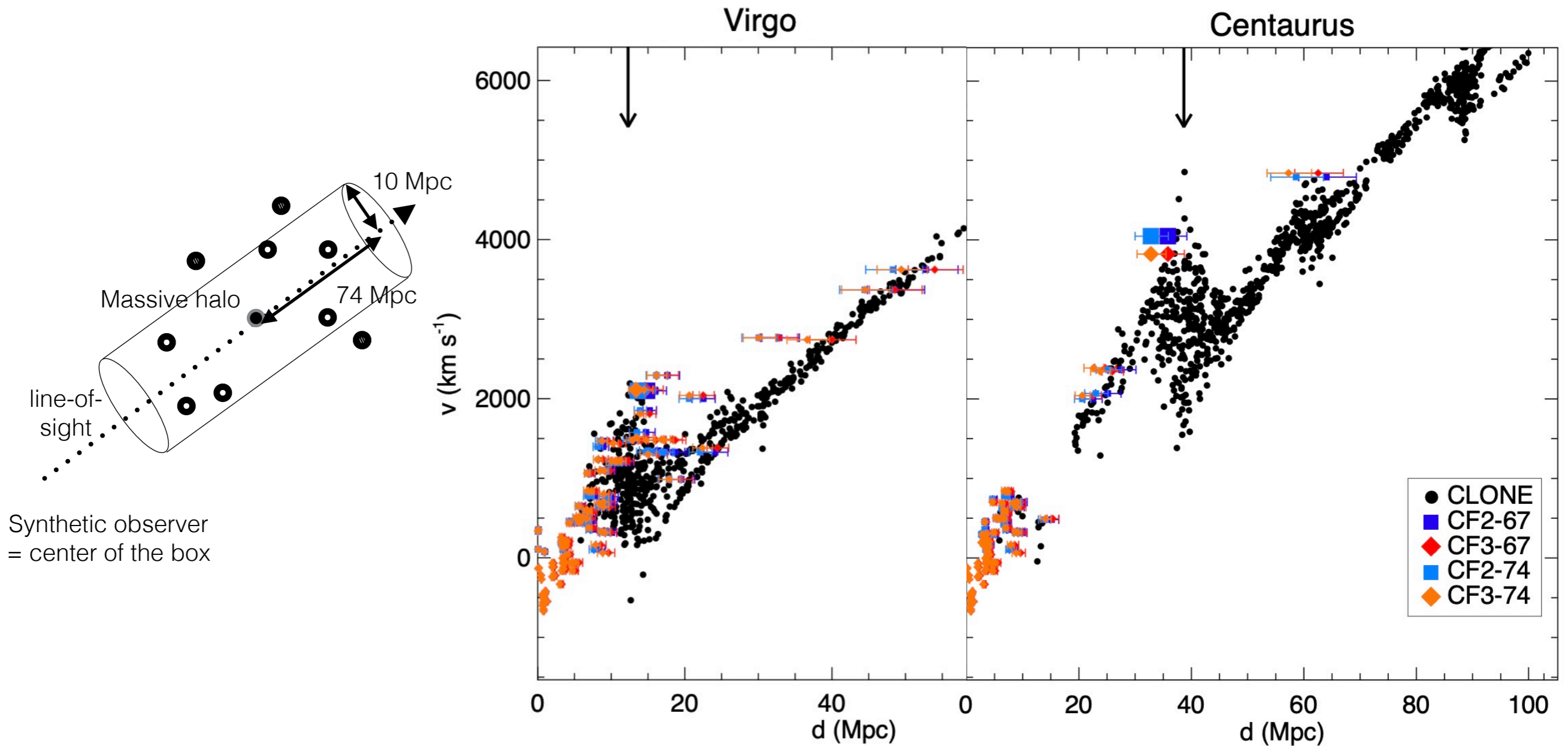


Note the fingers of gods

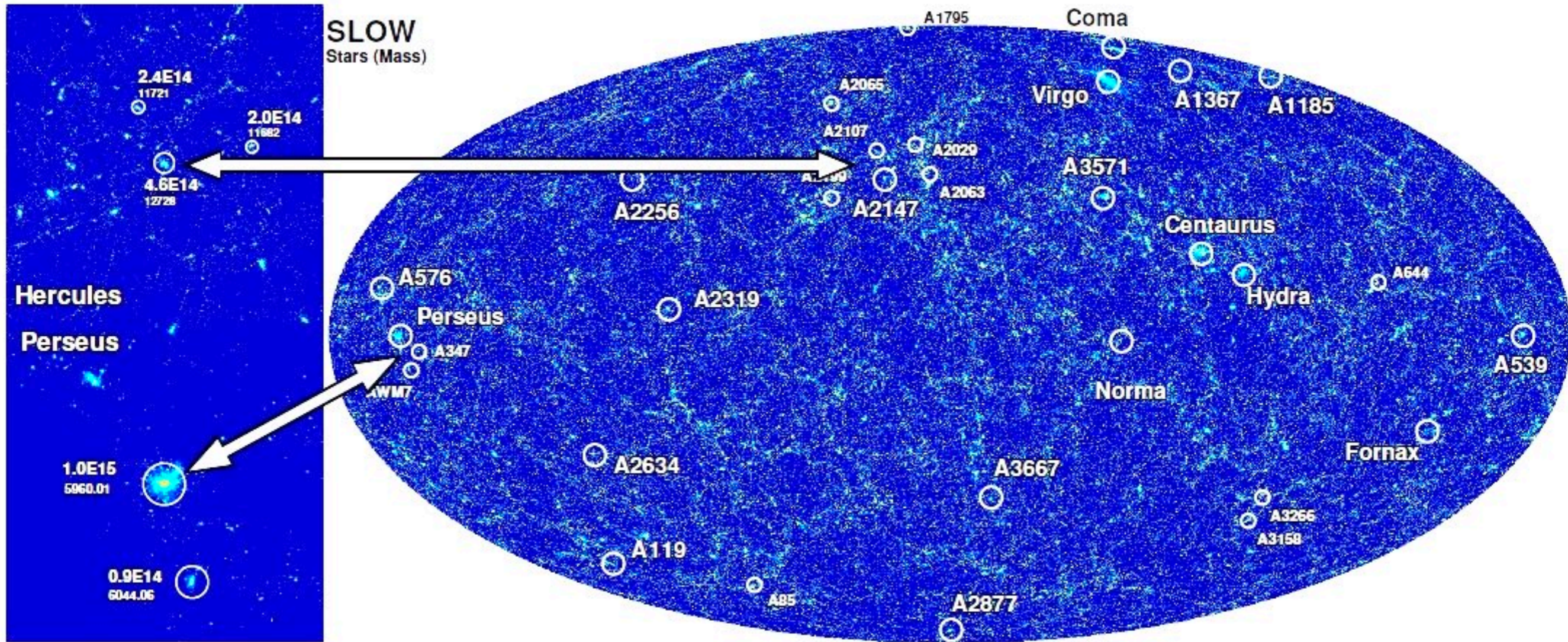


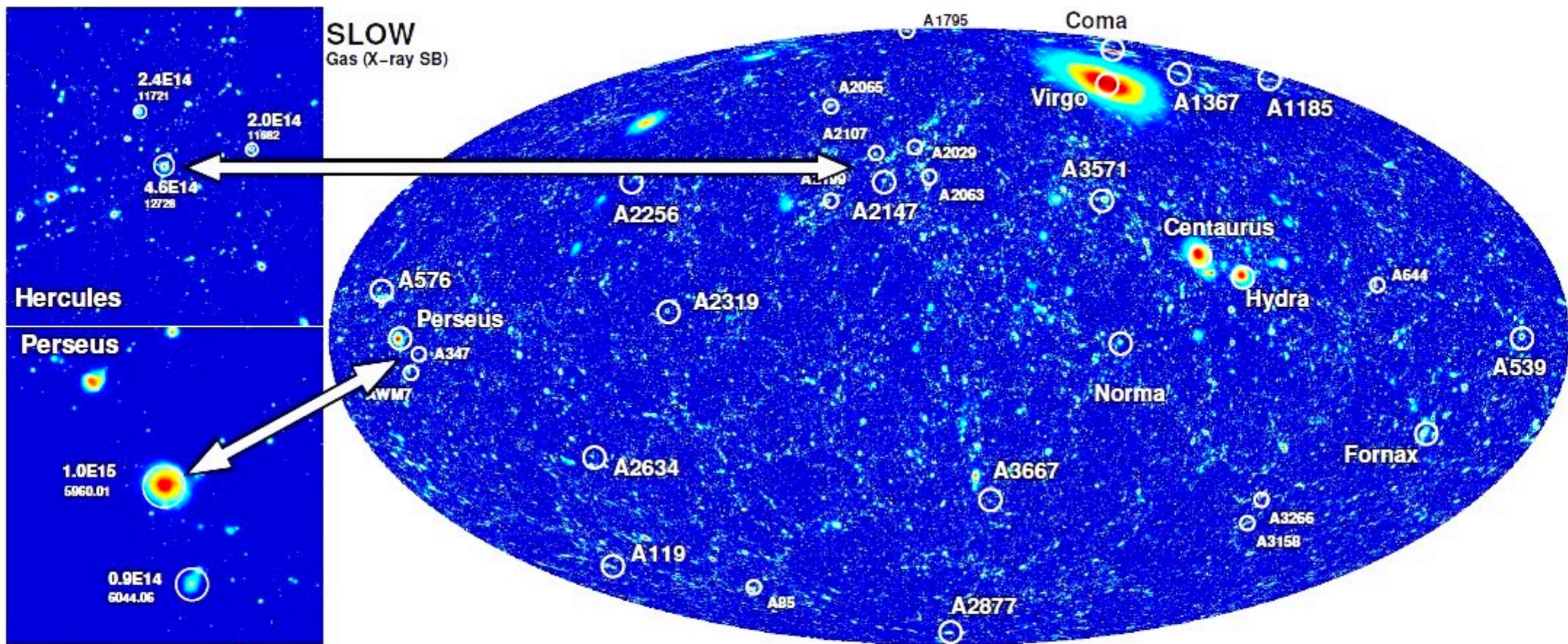
500 Mpc/h, 1024^3 particles,
DM only, Planck cosmology

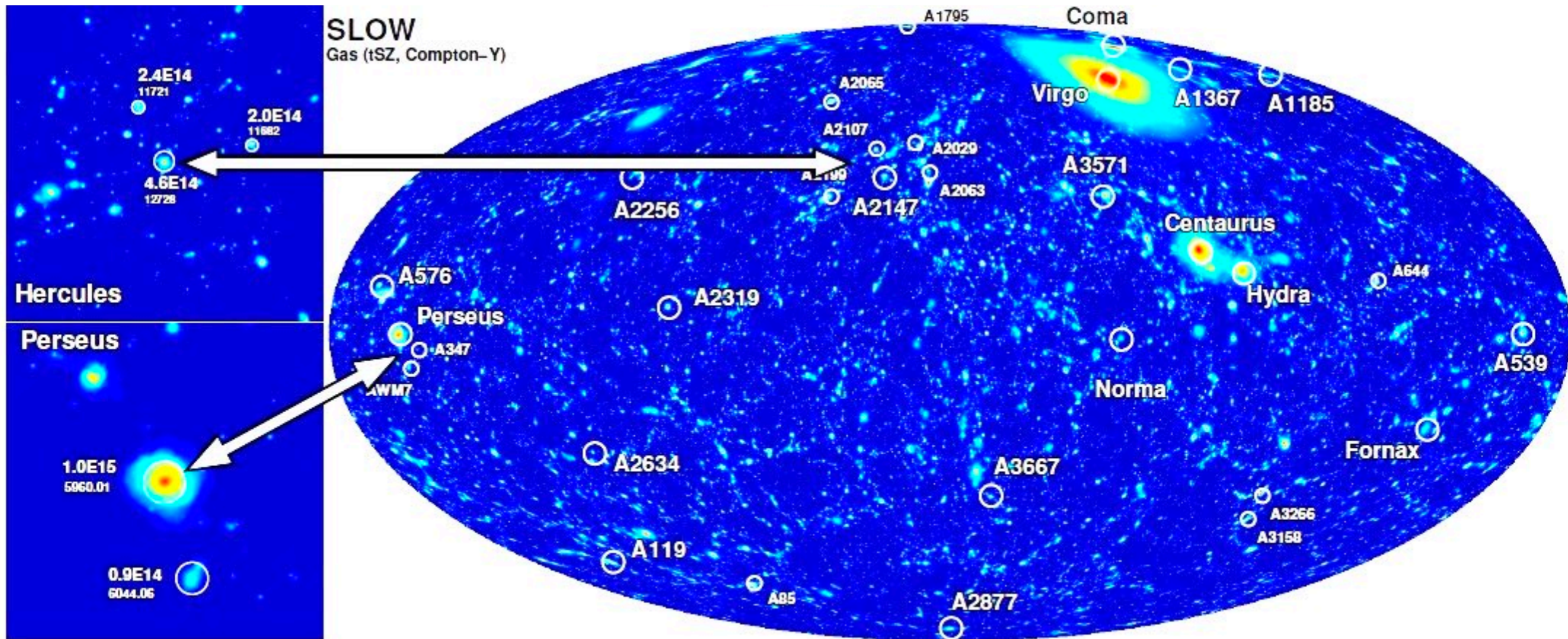
Velocity wave signatures in the Hubble diagram

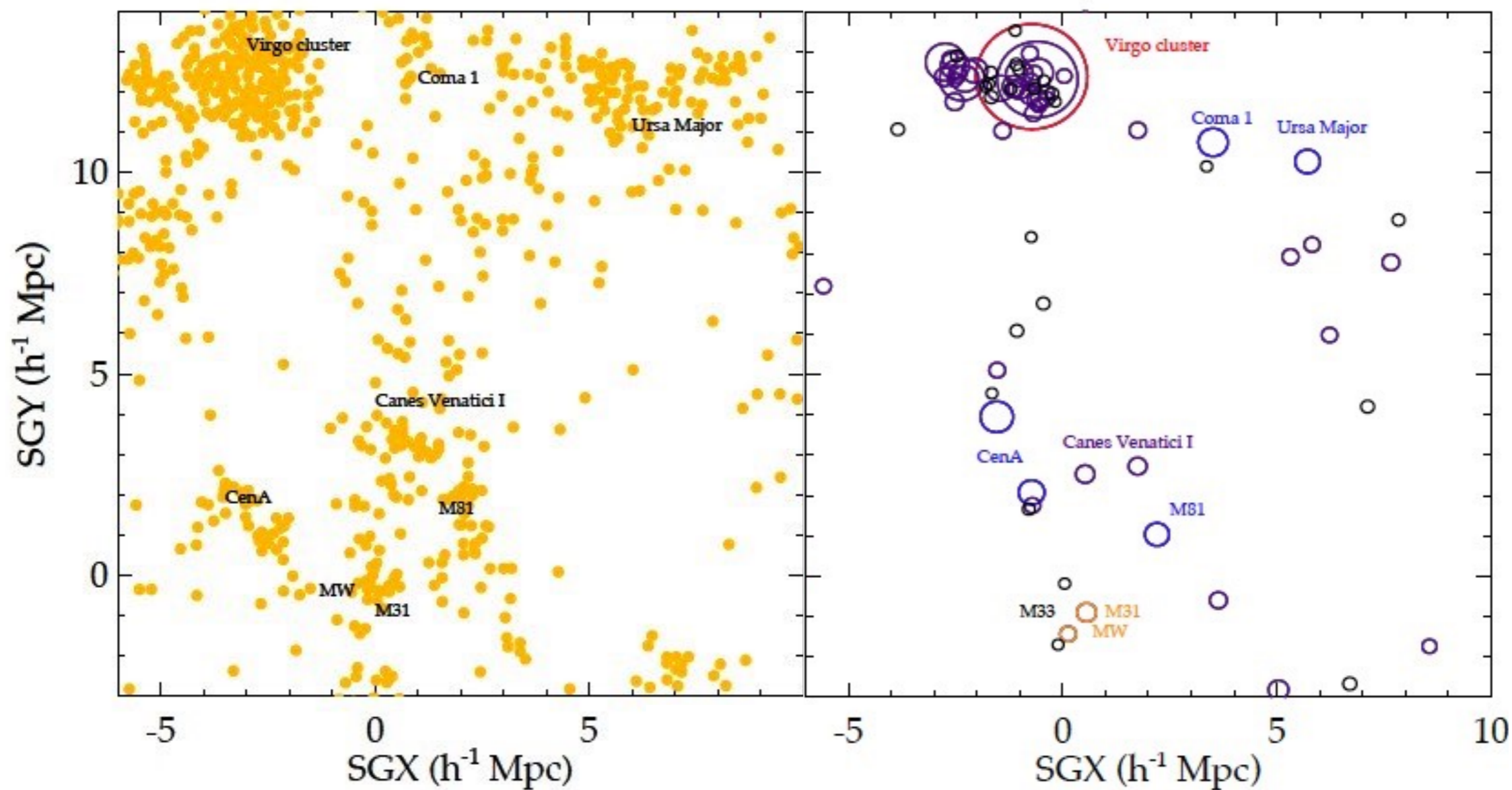


500 Mpc/h, 2048³ particles, DM only, Planck cosmology





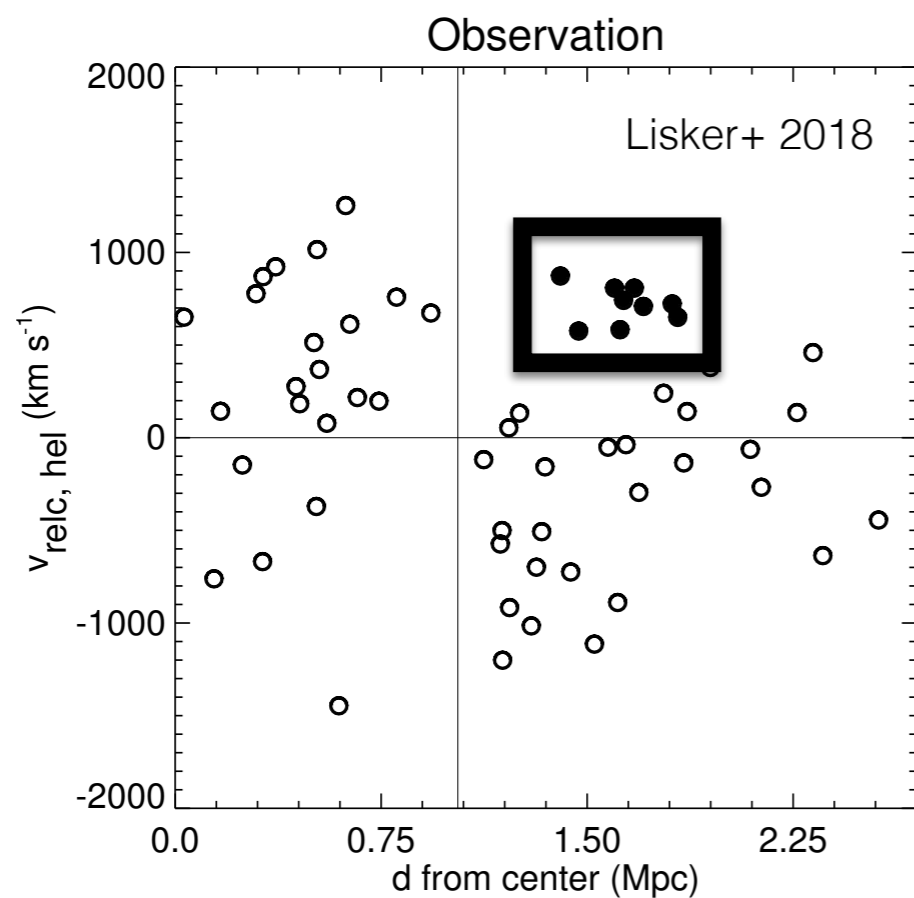




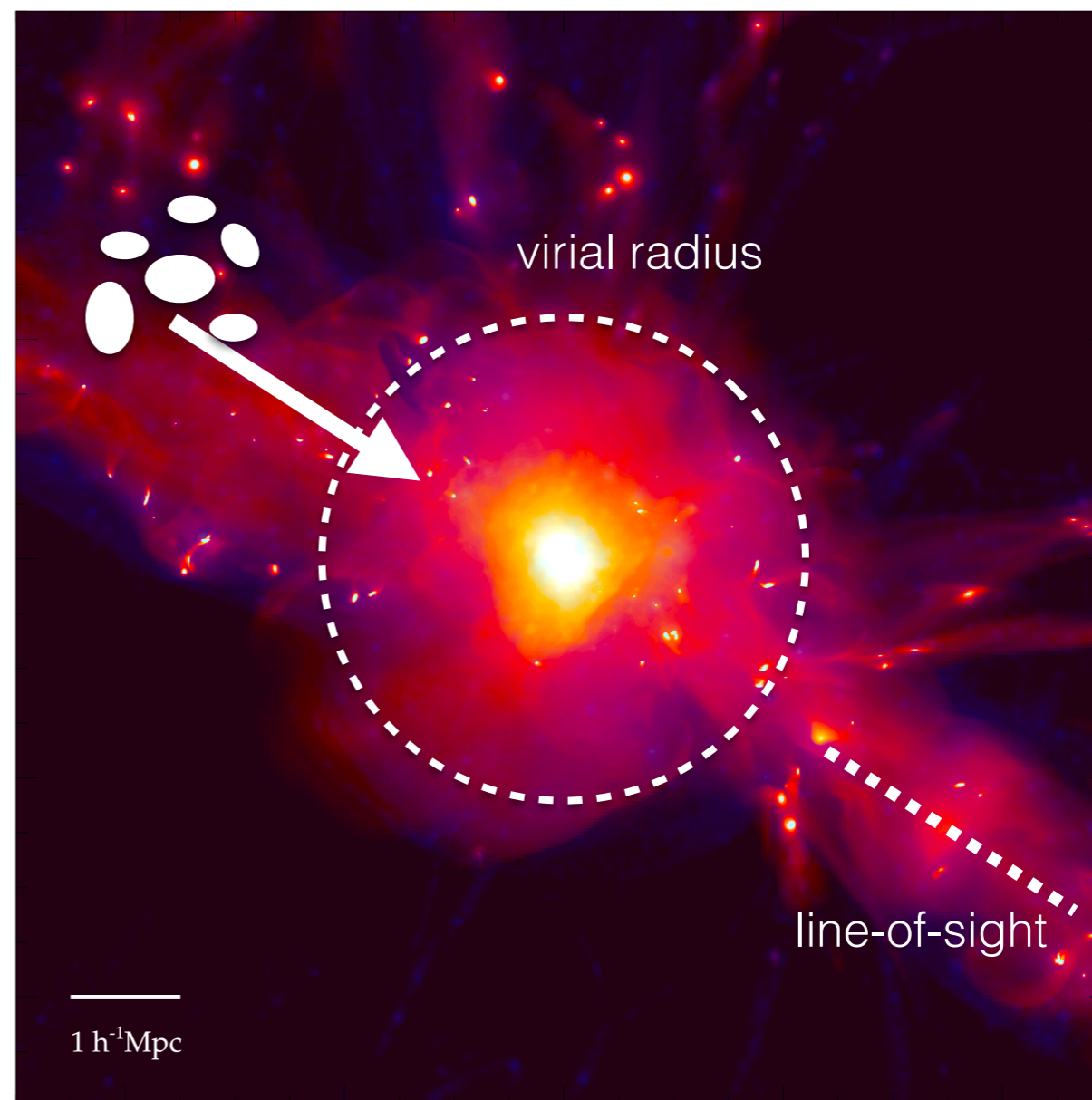
64 Mpc/h, 2048^3 particles, DM
only, Planck cosmology

Simulated & Observed Virgo clusters

Lisker+2018: from observation, remnant of a group of $\sim 10\%$ m_{cluster} that infall 2-3 Gyr ago

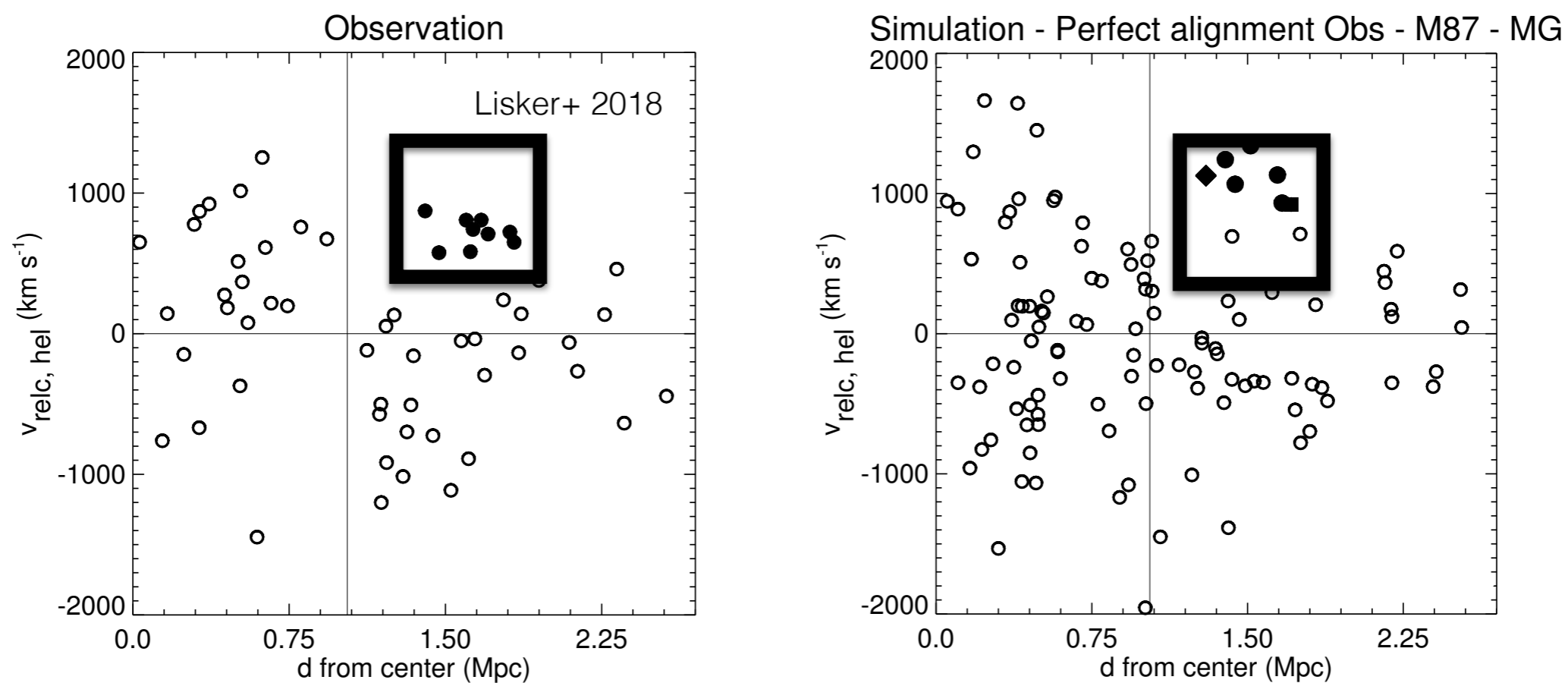


Group of galaxies that fell within the line-of-sight?



Simulated & Observed Virgo clusters

Lisker+2018: from observation, remnant of a group of $\sim 10\%$ m_{cluster} that infall 2-3 Gyr ago

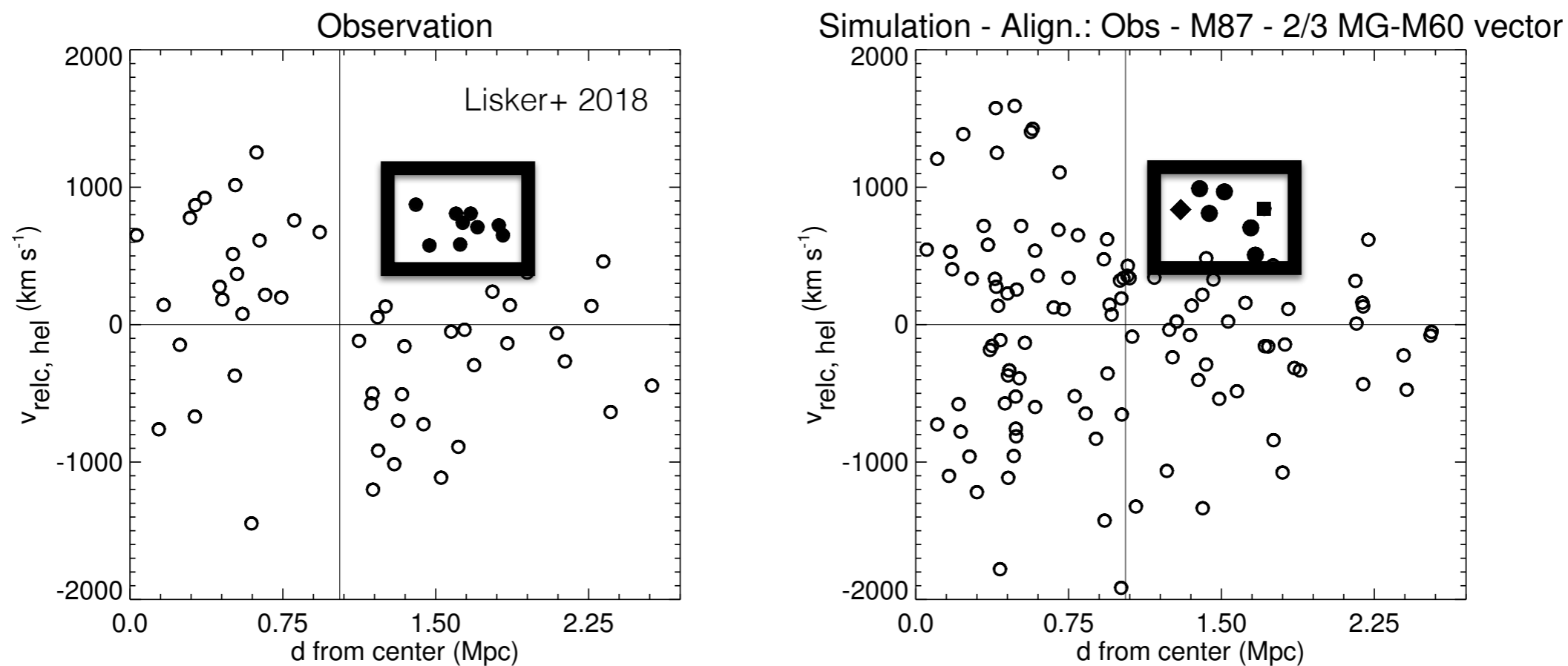


**Group of galaxies that fell
within the line-of-sight?**

Sorce+2021

Simulated & Observed Virgo clusters

Lisker+2018: from observation, remnant of a group of $\sim 10\%$ m_{cluster} that infall 2-3 Gyr ago

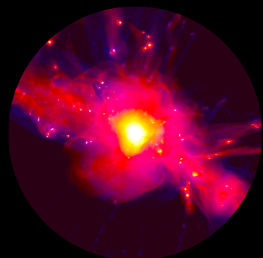


Group of galaxies that fell
quasi within the line-of-sight

Sorce+2021

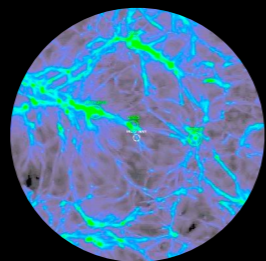
Agreement with observational predictions

Because they can reproduce what we observe
(quasi) **free of cosmic variance**, they can reduce
biases and help solve the question :
tensions or systematics ?



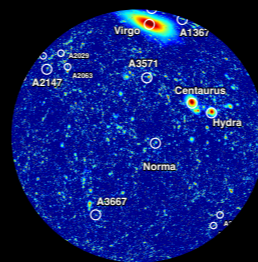
Virgo Cluster

(Sorce+2016, 2019, 2021, in prep., Olchanski & Sorce 2018, Lebeau+submitted.)

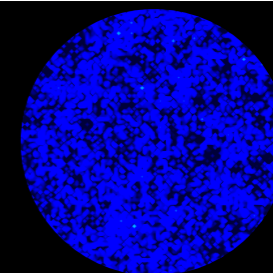


Cosmic Rays in the local Universe

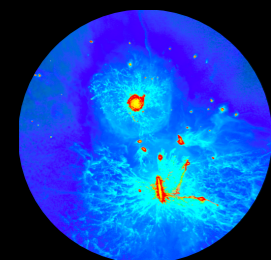
(Hackstein+2018, Boess+in prep.)



SLOW: local web
(Dolag, Sorce+2023)

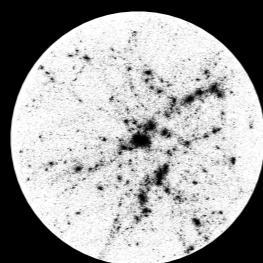


LOCALIZATION: local cluster signatures
(Sorce, Aghanim, Lebeau, Jung, Dolag)



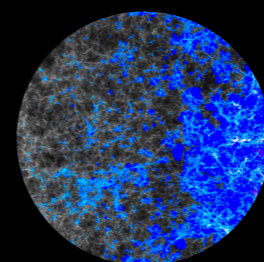
HESTIA: Local Group

Carlesi, Sorce+2016, Carlesi+2016, 2017, Libeskind+2020, Damle+2022, Newton+2022; Luis+2022, Dupuy+2022, Arora+2022, Khoperskov+2022a,b,c



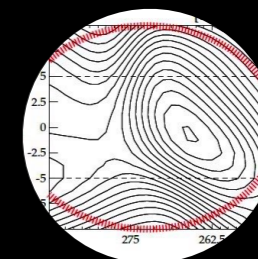
Coma connectivity

(Malavasi, Sorce, Dolag, Aghanim 2023)



CoDa: Reionization of the local Universe

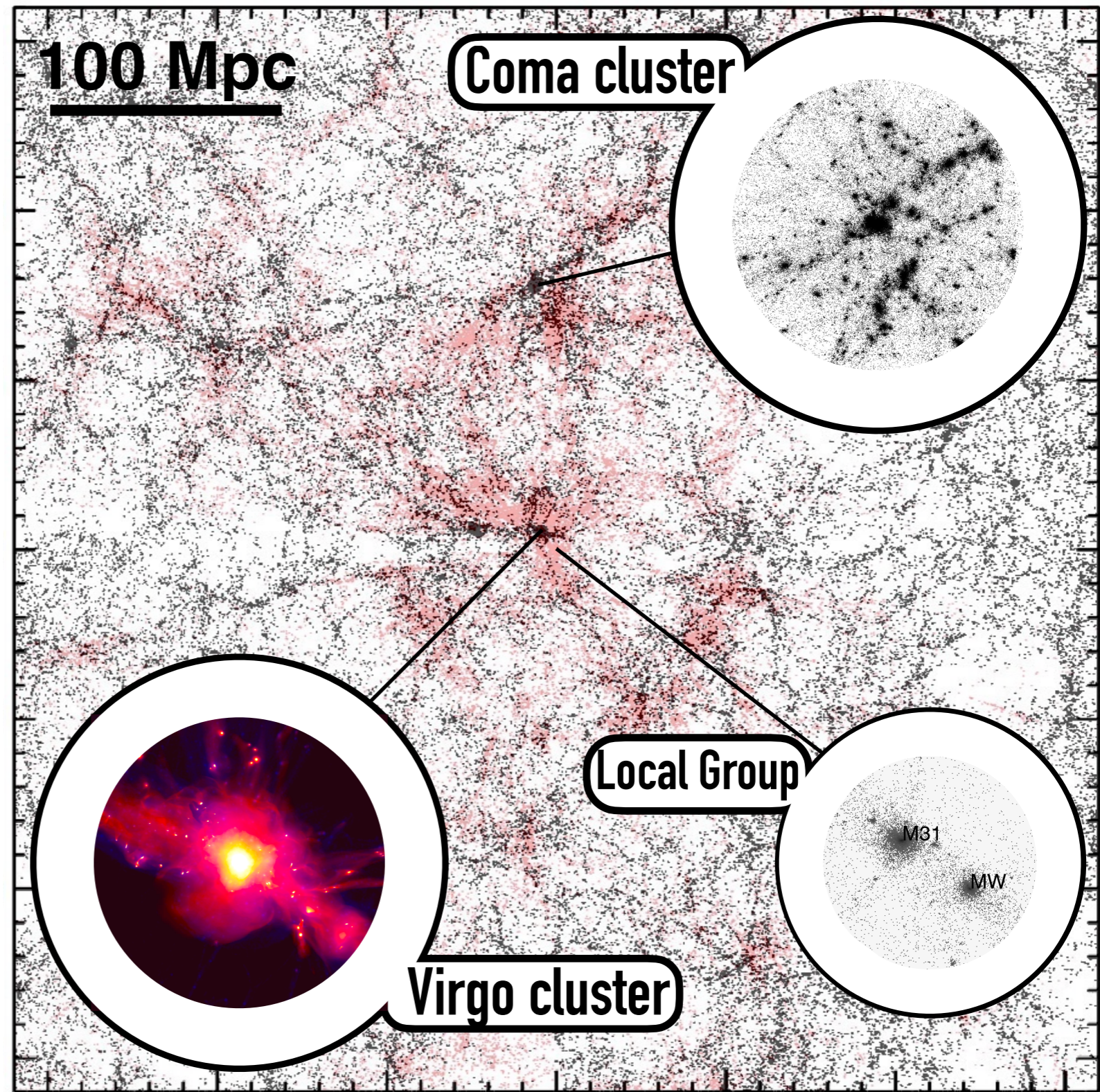
(Ocvirk+2020, Lewis+2020, Gronke+2021, Sorce+2022, Lewis+2022, Park+2022)



Zone of Avoidance
(Sorce+2017)

Conclusion

- **Standard** cosmological simulations give only the full uncertainty
- **Constrained** cosmological simulations can permit **disentangling tensions from systematics (by reducing biases)**
- **CLONES are constrained** cosmological simulations valid down to the cluster scales with induced smaller scales
- CLONES are **widely used** and **maybe you are the next users!**



**Thank you, Merci, Grazie,
Gracias, Danke, ευχαριστώ
Mahalo, 谢谢, ありがとう,
הודת, Obrigada, Dank u,
Tak, Cảm ơn, Dziękuję, 감사합니다
Kiitos, Aitäh, diolch, dankewol,
ಧನ್ಯವಾದಗಳು, ...***

* Missing your 'thanks' spelling? It means I did not get the chance to learn how to say it so far

