

北京航空航天大学
BEIHANG UNIVERSITY

Cosmological implications of the radiative electroweak symmetry breaking theories

Ke-Pan Xie

Beihang University

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[2408.03649 \[PRD\]](#), [2304.00908 \[PRD\]](#), [2206.04691 \[JHEP\]](#)

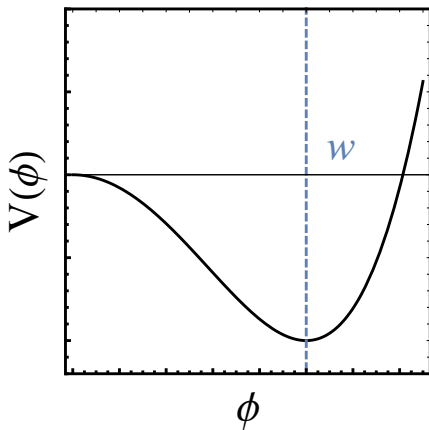
Patterns of spontaneous symmetry breaking

Conventional setup

$$V(\phi) = \frac{\mu^2}{2}\phi^2 + \frac{\lambda}{4}\phi^4$$

with $\mu^2 < 0$

$$\langle \phi \rangle = w = \sqrt{-\mu^2/\lambda};$$



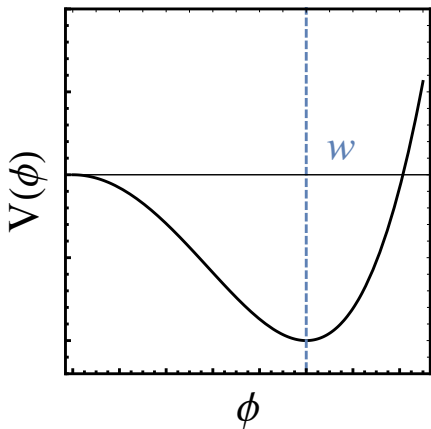
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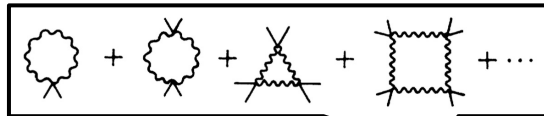
$$\langle \phi \rangle = w = \sqrt{-\mu^2/\lambda};$$



Radiative symmetry breaking [this talk]

$V(\phi) = \lambda\phi^4/4$ at tree-level; at **one-loop** level

$$V_1(\phi) = \frac{B}{4}\phi^4 \left(\log \frac{\phi}{w} - \frac{1}{4} \right)$$



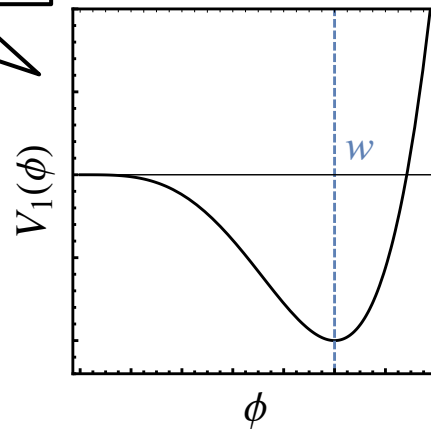
Scale & conformal anomaly

Dimensional transmutation

S. Coleman & E. Weinberg PRD 7 (1973) 1888-1910, citation 5000+

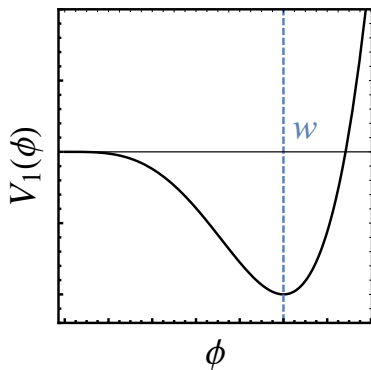
No bare ϕ^2 -term, hopefully solving hierarchy problem

Bardeen, FERMILAB-CONF-95-391-T;
Meissner *et al*, PLB 660 (2008) 260-266; etc



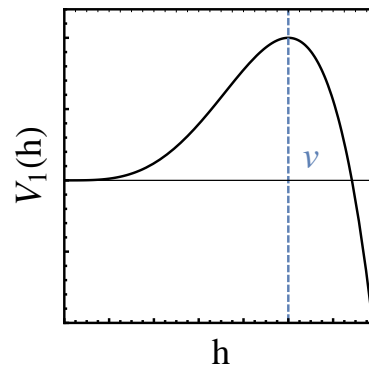
Radiative electroweak symmetry breaking?

However **not** directly applied to the Standard Model **Higgs!**



If identifying ϕ as h

$$V_h(h) \sim \frac{B_h}{4} h^4 \left(\log \frac{h}{v} - \frac{1}{4} \right)$$



$$B_h \propto (9g^4 - 48y_t^4) < 0$$

EW bosons top quark

Unstable electroweak vacuum!



Radiative electroweak symmetry breaking!

A realistic framework: new scalar ϕ , tree-level

$$V_0(h, \phi) = \frac{\lambda_s}{4} \phi^4 - \frac{\lambda'}{4} \phi^2 h^2 + \frac{\lambda_h}{4} h^4$$

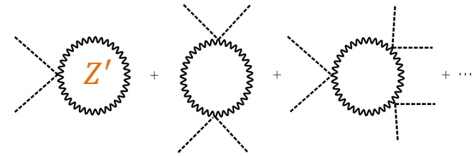
- Classically scale-invariant & conformal

Along the **new physics** direction

$$V_\phi(\phi) \approx \frac{B}{4} \phi^4 \left(\log \frac{\phi}{w} - \frac{1}{4} \right)$$

For example:

- $U(1)_X$ with g_X
- $B = 3g_X^4/8\pi^2$



Then $\phi \rightarrow w$, and along the Higgs direction

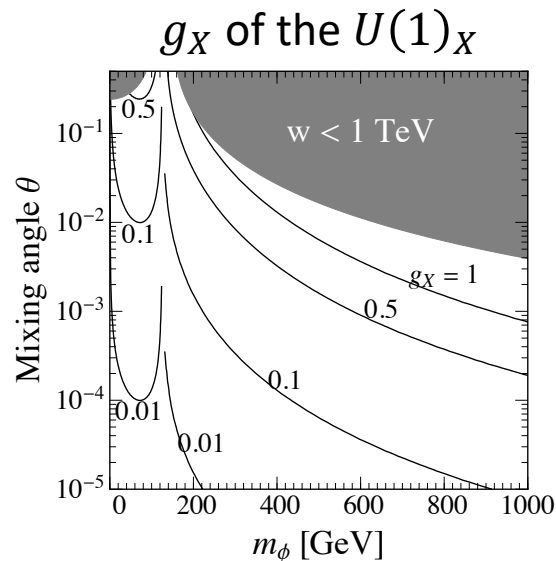
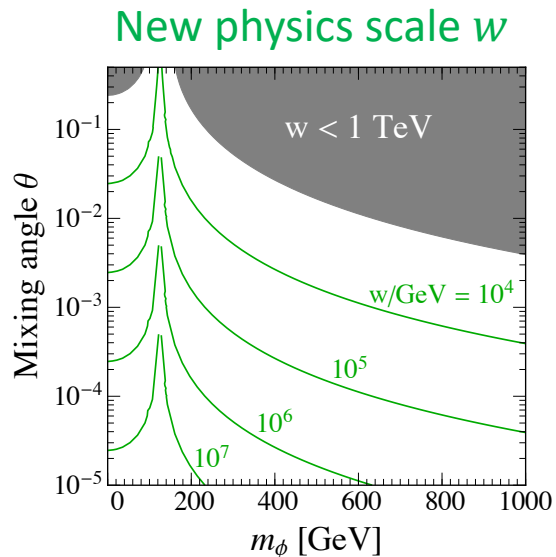
$$V_h(h) \approx \left(-\frac{\lambda'}{4} \phi^2 h^2 + \frac{\lambda_h}{4} h^4 \right) \rightarrow \left(\frac{\mu_h^2}{2} h^2 + \frac{\lambda_h}{4} h^4 \right)$$

$\mu_h^2 < 0 \rightarrow$ EWSB! [Hempfling, PLB 379, 153 \(1996\)](#); [Iso et al, PLB 676, 81 \(2009\)](#); [Chun et al, PLB 725, 158 \(2013\)](#), etc

Vacuum structure decomposition

Only **2** free parameters [Liu and KPX, PRD 110 (2024) 11, 115001]

- Use (m_ϕ, θ) as input, all other parameters derived **rigorously**

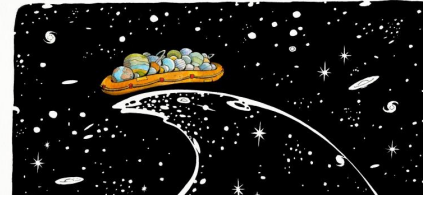
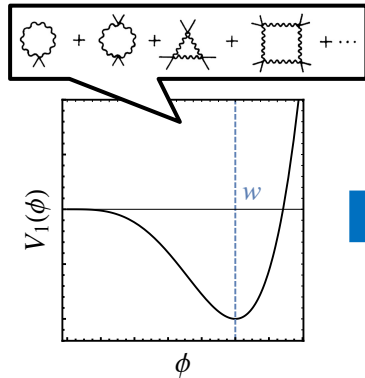


ϕ cannot be degenerate with the Higgs boson, $m_\phi \neq 125$ GeV

$w \gg v_{EW}$ needed for sequential symmetry breaking [Chataignier *et al*, JHEP 08 (2018) 083]

Cosmological implications

If the scale is dynamically generated,
then the early Universe must be **boiling**



Gravitational waves



(Novel) dark matter

then more interesting consequences...

(Novel) origin of matter, baryogenesis

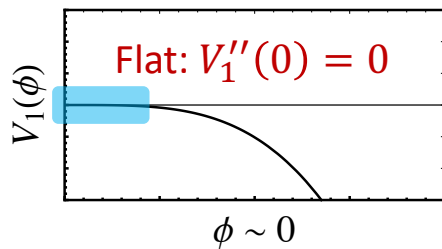


Why the Universe was boiling

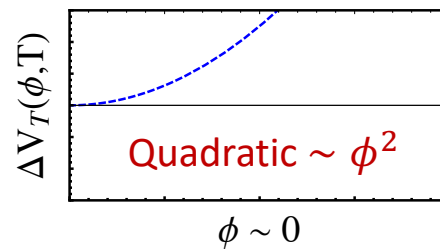
Early Universe → temperature corrections

Near the field origin,

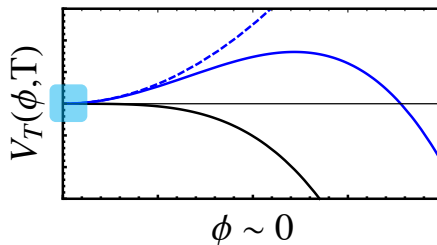
$$V_T \approx \underbrace{\frac{B}{4} \phi^4 \left(\log \frac{\phi}{w} - \frac{1}{4} \right) - \frac{\lambda'}{4} h^2 \phi^2 + \frac{\lambda_h}{4} h^4}_{V_1(\phi)} + \underbrace{\frac{c_\phi T^2}{2} \phi^2 + \frac{c_h T^2}{2} h^2}_{\Delta V_T(\phi, T)}$$



+

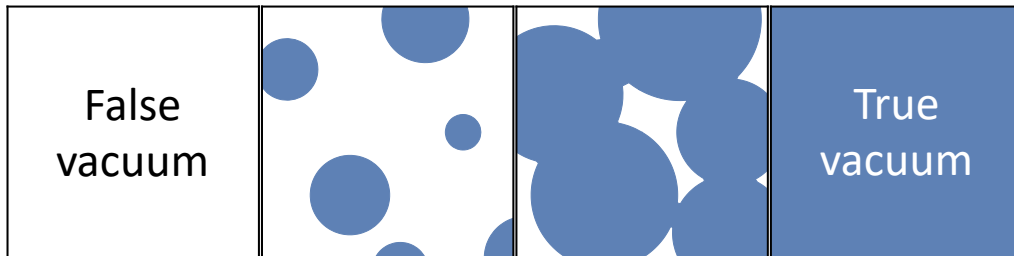
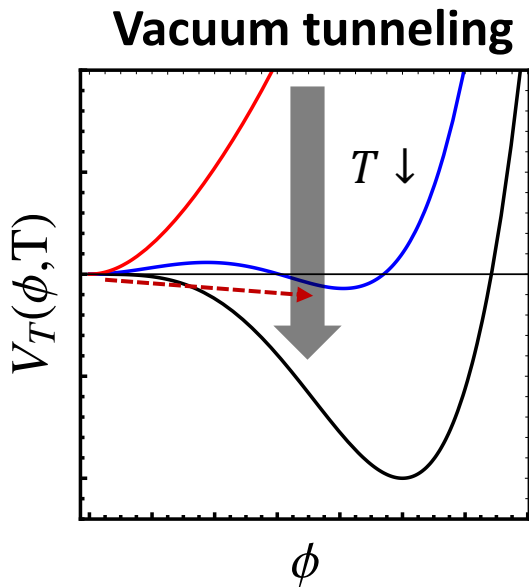


Local minimum exists for $T > 0$
The Universe is trapped in the
false vacuum!



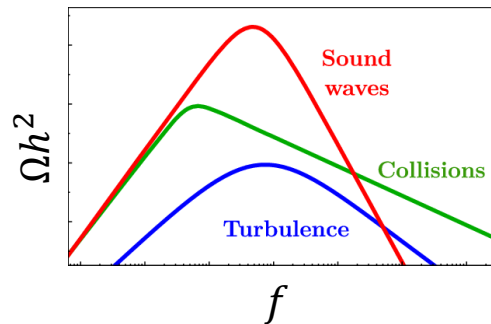
Why the Universe was boiling

A first-order phase transition



Stochastic GWs:

1. Bubble collisions
2. Sound waves
3. Turbulence

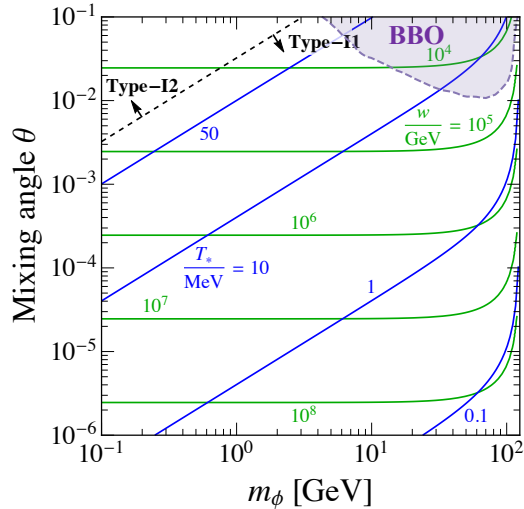


Just for an intuitive explanation; full one-loop calculation included in paper

*See also talks of [Huai-Ke Guo](#), [Alan S. Cornell](#), [Tomasz Dutka](#), [Joshua Cesca](#), and [Haipeng An](#)

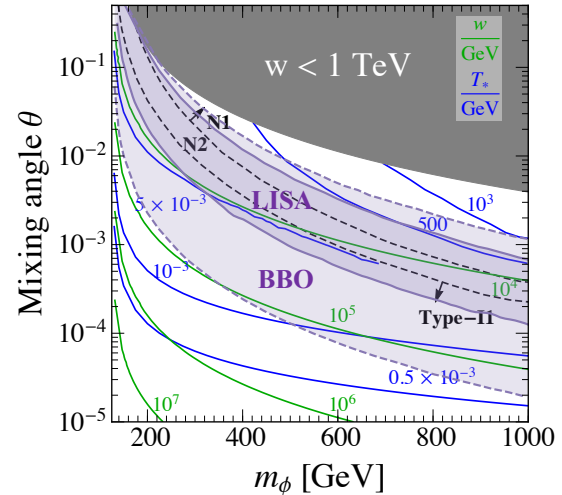
FOPT and gravitational waves

Only two free parameters: new scalar boson mass m_ϕ , mixing angle θ
 Fully thermal one-loop calculation [Liu and KPX, PRD 110 (2024) 11, 115001]



$\leftarrow m_\phi < 125 \text{ GeV}$

$m_\phi > 125 \text{ GeV} \rightarrow$



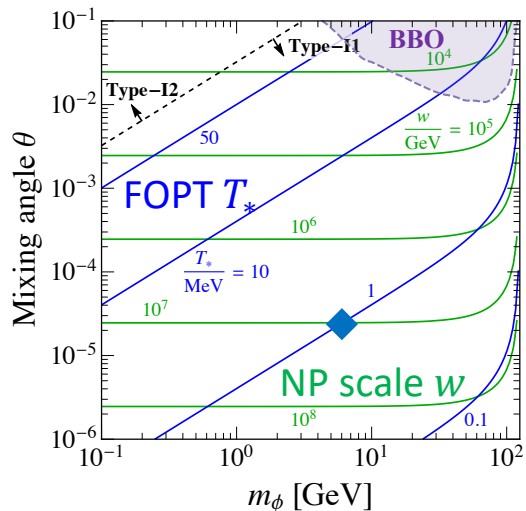
Blue: FOPT temperature T_*

Green: new physics scale w

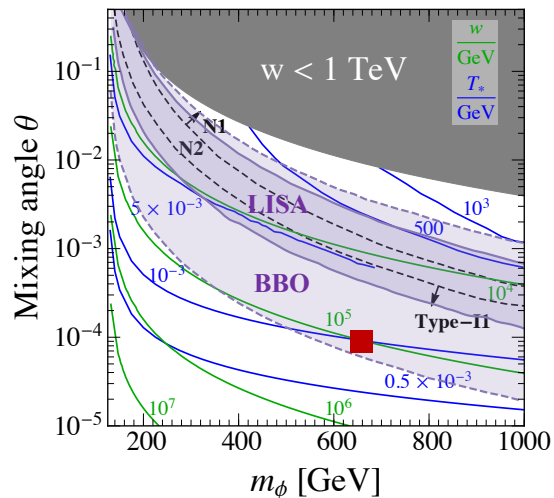
Purple: LISA (similar for TianQin, Taiji) and BBO gravitational wave detection

FOPT and gravitational waves

Only two free parameters: new scalar boson mass m_ϕ , mixing angle θ
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$T_* \ll w$:
 Supercooling!

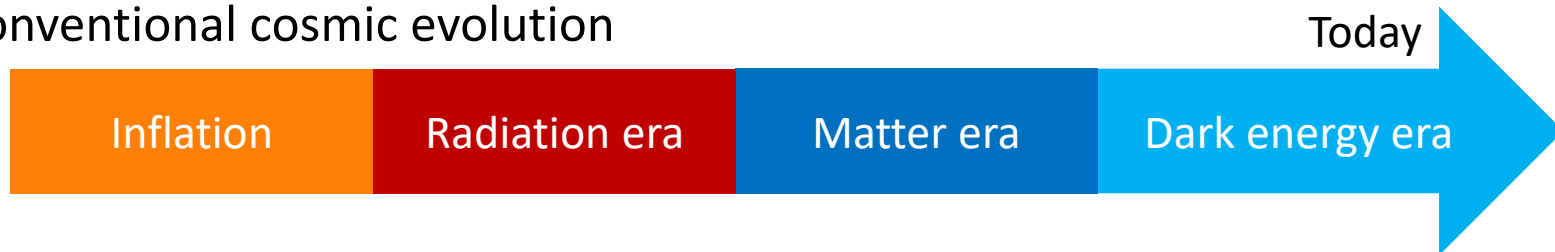


◆ $w = 10^7 \text{ GeV}, T_* = 1 \text{ MeV},$
 $T_{\text{rh}} \sim 1 \text{ TeV}$

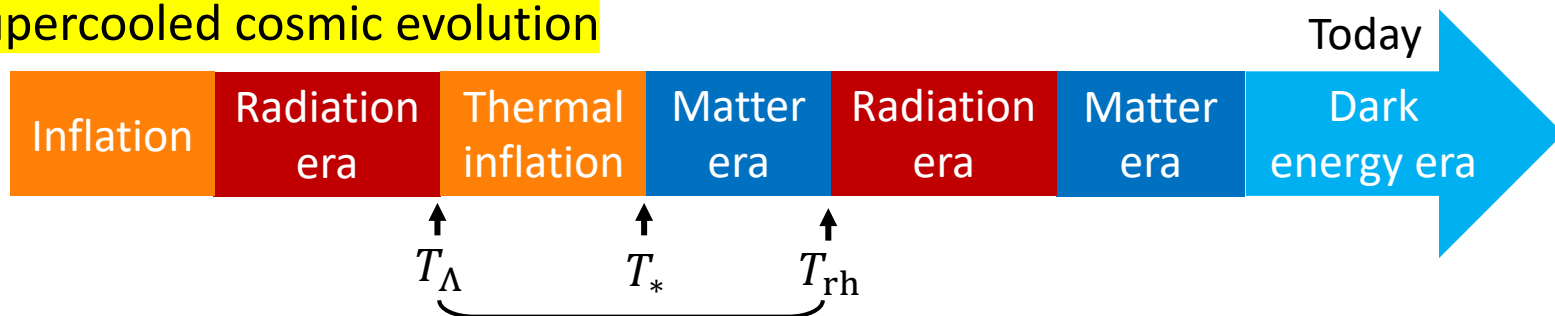
■ $w = 10^5 \text{ GeV}, T_* = 1 \text{ MeV},$
 $T_{\text{rh}} \sim 3 \text{ TeV}$

Exotic cosmic history

Conventional cosmic evolution



Supercooled cosmic evolution

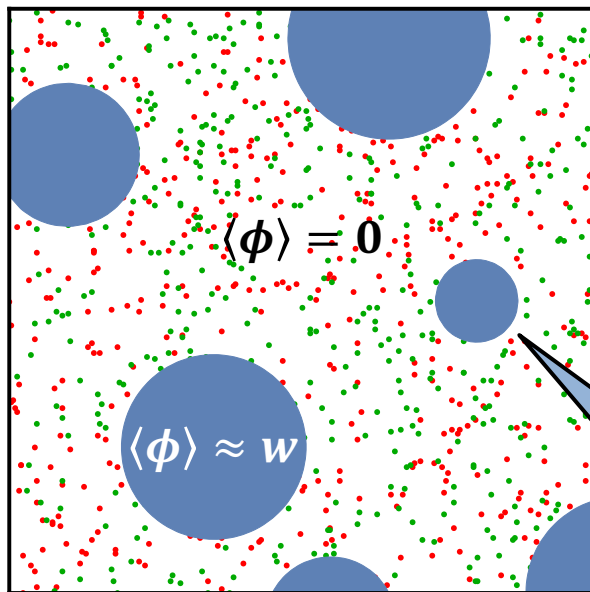


Rich implications on dark matter and baryogenesis mechanisms

Impact on elementary particle physics

Possible interactions (*scale-invariant* at tree-level)

$$\mathcal{L}_{\text{int}} \supset \bar{\chi} i \gamma^\mu \partial_\mu \chi - y_\chi \phi \bar{\chi} \chi + \partial_\mu S^\dagger \partial^\mu S - \frac{\lambda_S}{2} \phi^2 S^\dagger S$$



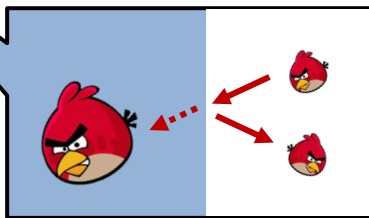
Outside the bubble:

- Particles massless

Mass discontinuity

Inside the bubble:

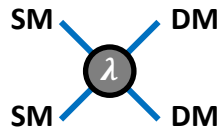
- $m_\chi \approx y_\chi w$ and $m_S \approx \sqrt{\lambda_S/2} w$



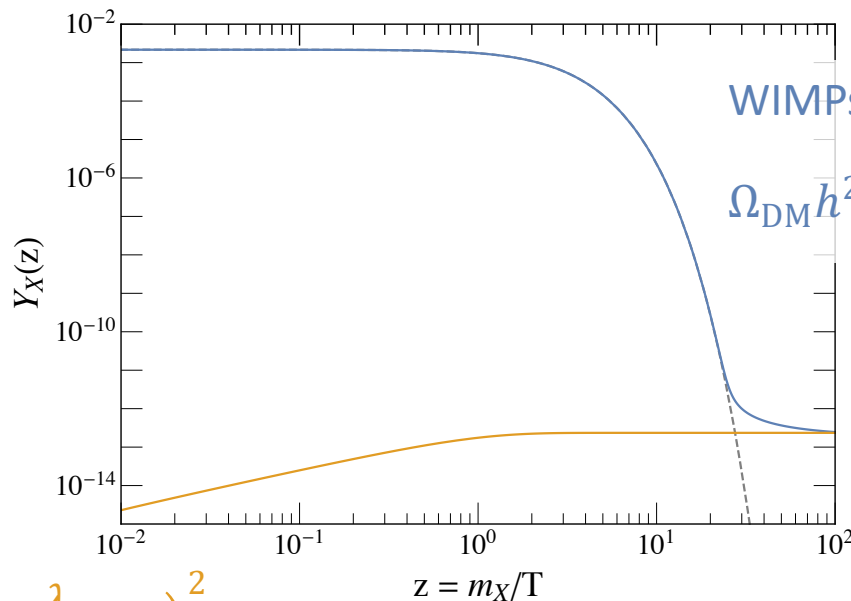
Particles scatter with the bubble wall

Conventional dark matter scenarios

2 → 2 scattering process between SM & DM
Smooth cosmic evolution



$$\langle \sigma v_{\text{rel}} \rangle \sim \frac{\lambda^2}{32\pi m_{\text{DM}}^2}$$



WIMPs: freeze-out

$$\Omega_{\text{DM}} h^2 \sim 0.1 \left(\frac{0.5}{\lambda} \right)^2 \left(\frac{m_X}{\text{TeV}} \right)^2$$

FIMPs: freeze-in

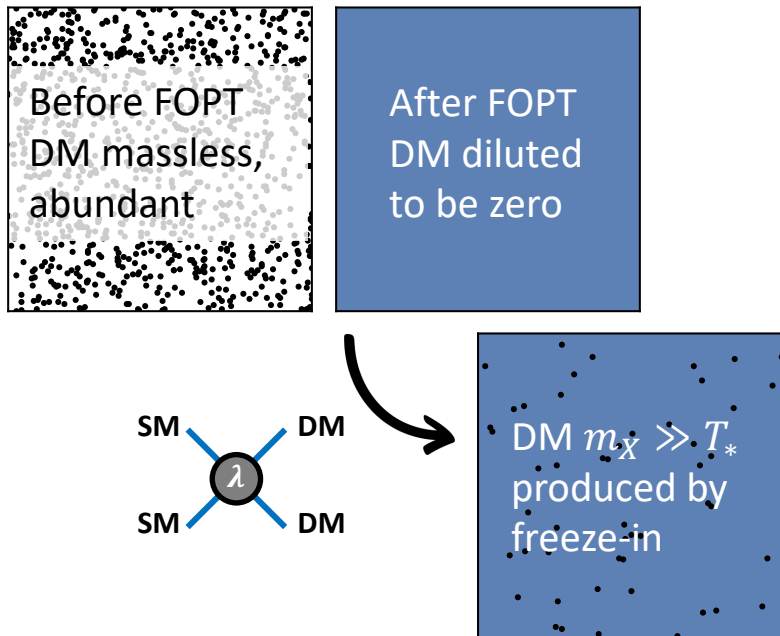
$$\Omega_{\text{DM}} h^2 \sim 0.1 \left(\frac{\lambda}{2.5 \times 10^{-11}} \right)^2$$

A new possibility arises in a supercooled Universe...

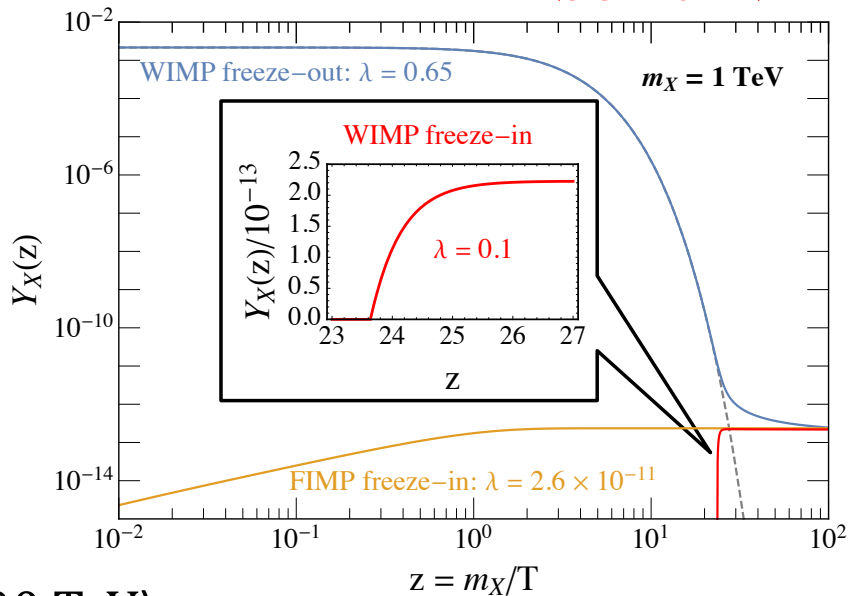
Dark matter in a supercooled Universe

Supercooled **WIMP**: **mass changes** after the FOPT, and experiences **freeze-in**

Wang and KPX, PRD 108 (2023) 5, 055035; Hambye *et al*, JHEP 08 (2018) 188; etc



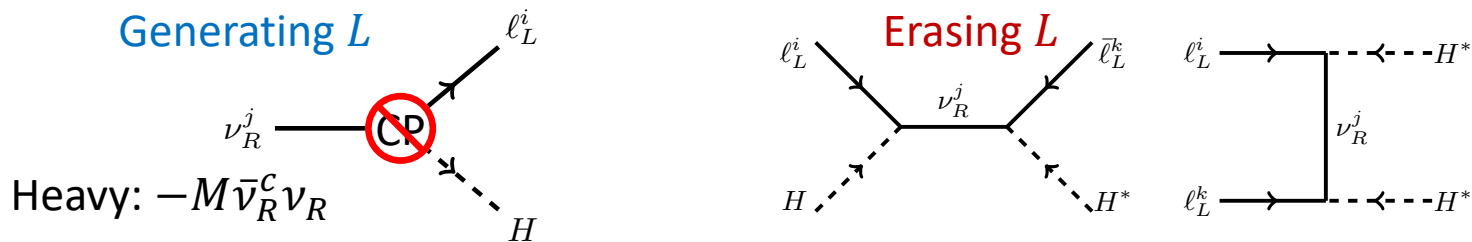
$$\Omega_{\text{DM}} h^2 \sim 0.1 (1 + 2z_2) \left(\frac{\lambda e^{-z_2}}{3.5 \times 10^{-11}} \right)^2$$



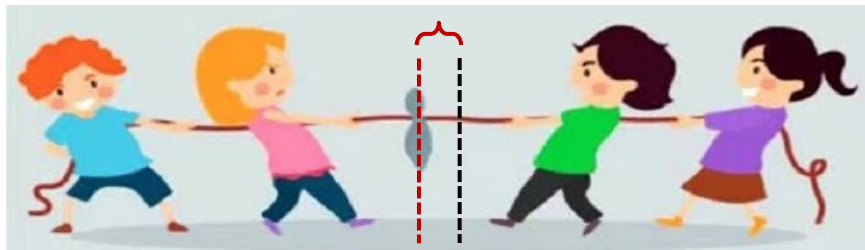
Could go **beyond** the GK bound ($\sim 100 \text{ TeV}$)

Conventional leptogenesis

Heavy ν_R decays at $T \lesssim M$, suffers from thermal washout



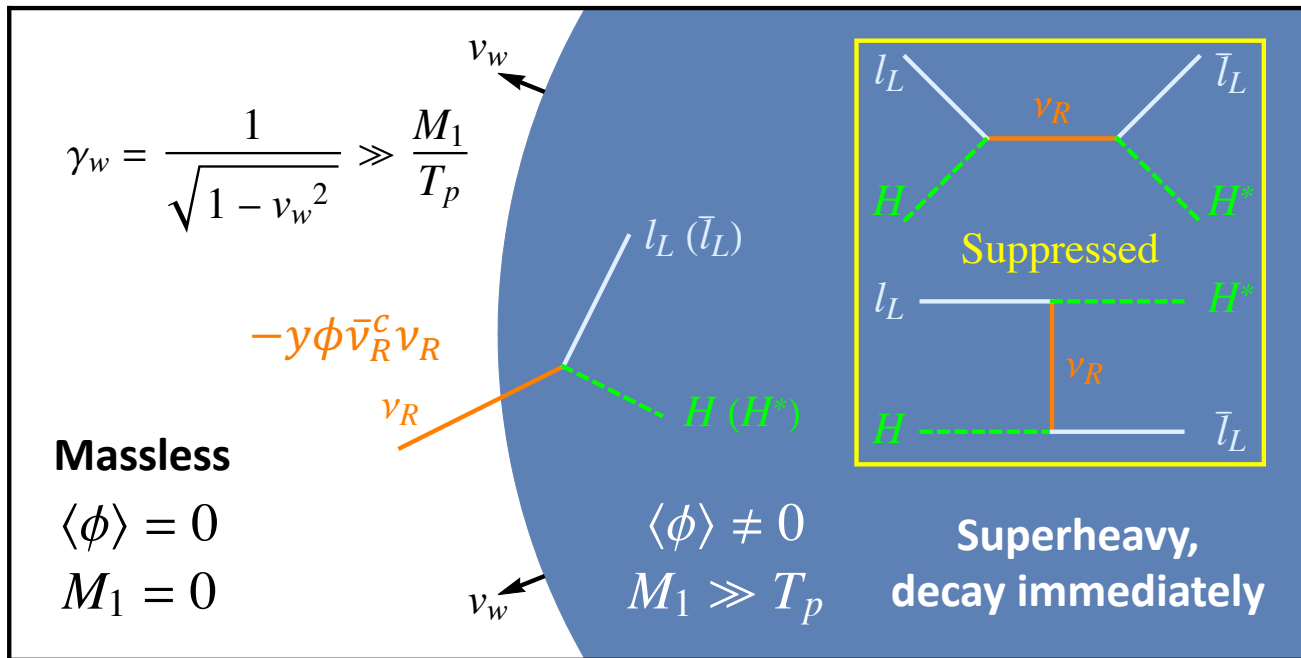
Only 1% survives



A new possibility arises in a supercooled Universe...

Leptogenesis in a supercooled Universe

Supercooled leptogenesis



- Leptogenesis **without** thermal washout

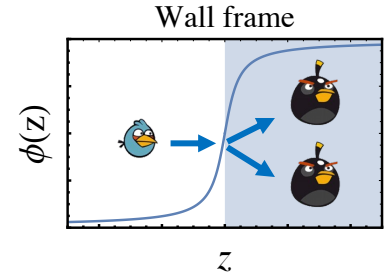
Huang and KPX, JHEP 09 (2022) 052; Chun *et al*, JHEP 09 (2023) 164; Baldes *et al*, PRD 104 (2021) 115029; etc

More implications

Heavy particle production on the wall if $\gamma_w T_* \gg m_S \gg T_*$

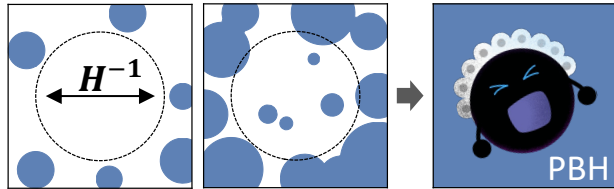
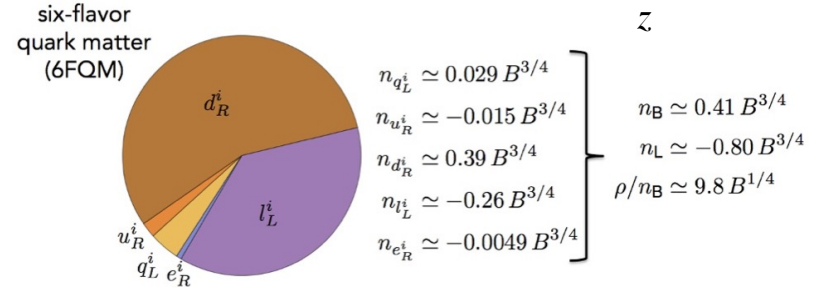
- Dark matter or baryogenesis

Azatov *et al*, JHEP 03 (2021) 288; JHEP 10 (2021) 043;
Baldes *et al*, PRD 104 (2021) 115029; Ai *et al*, 2406.20051; etc



Quark nugget soliton dark matter from a QCD FOPT induced by classically conformal theories

Bai *et al*, JHEP 06 (2018) 072; Witten, PRD 30 (1984) 272



Primordial black holes formation

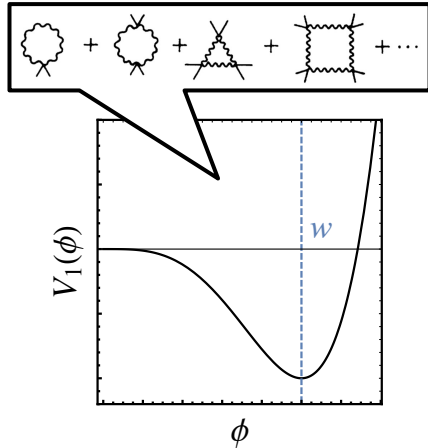
- Bubble collisions Hawking *et al*, PRD 26 (1982) 2681; Jung *et al*, 2110.04271; etc

- False vacuum "islands" collapse Liu, Bian, Cai, Guo, Wang, PRD 105, L021303; Gouttenoire, PLB 855, 138800 (2024); Ai *et al*, 2409.02175, etc

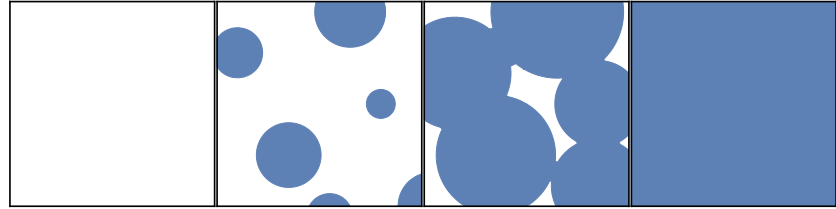
Conclusion & outlook

Simple assumption

Dynamical origin of the scale



Rich cosmological implications



FOPTs + detectable GWs

and maybe impacts on

- Cosmic evolution history
- Dark matter and/or baryogenesis
- Solitons or primordial black holes

A lot of future directions to explore!

Thank you!

Backup: structure of the potential

Why the potential is written as

$$V_1(h, \phi) \approx \frac{B}{4} \phi^4 \left(\log \frac{\phi}{w} - \frac{1}{4} \right) - \frac{\lambda'}{4} h^2 \phi^2 + \frac{\lambda_h}{4} h^4$$

In principle, both ϕ & h receive Coleman-Weinberg corrections

However if $w \gg v_{EW}$, then **sequential** SSB [Chataignier *et al*, JHEP 08 (2018) 083]

- ϕ -direction: loop correction dominates,

$$\frac{\lambda_\phi}{4} \phi^4 \sim \frac{B}{4} \phi^4 (\log \phi + C)$$

- h -direction: tree-level contribution dominates

$$\frac{\lambda_h}{4} h^4 \gg \frac{B_h}{4} h^4 (\log h + C)$$

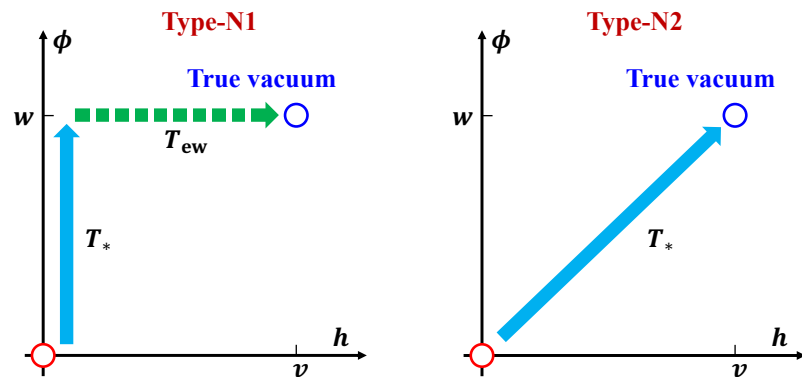
Therefore, along the ϕ -direction we adopt the log potential while along the h -direction we use the tree-level potential

Backup: thermal history in 2D field space

Normal pattern:

high scale \rightarrow low scale

- **N1**: ϕ -FOPT \rightarrow EW crossover
- **N2**: joint ϕ -EW-FOPT

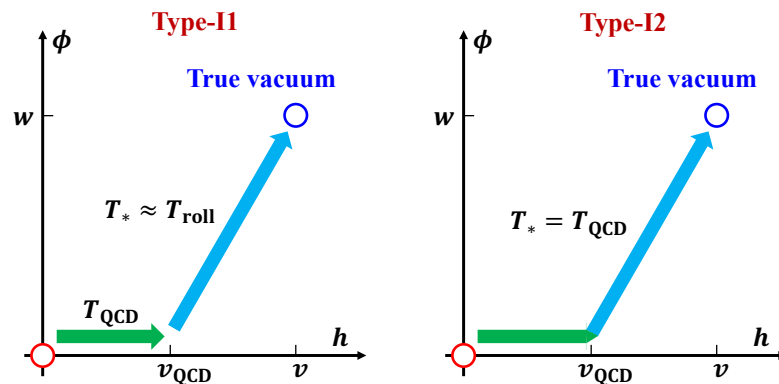


Detailed classification in Liu and KPX, PRD 110 (2024) 11, 115001

Inverted pattern:

low scale \rightarrow high scale

- **I1**: QCD-EW-FOPT \rightarrow ϕ -FOPT
- **I2**: joint QCD-EW- ϕ -FOPT



- **The inverted pattern**: QCD-FOPT triggers EWPT $h: 0 \mapsto v_{\text{QCD}} \approx 100 \text{ MeV}$ via $-y_t h \langle \bar{t} t \rangle$ Witten, NPB 177, 477 (1981); Iso *et al*, PRL 119 (2017) 14, 141301

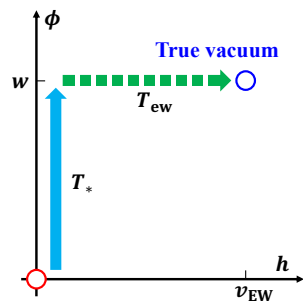
Backup: the normal pattern

ϕ -FOPT: $0 \mapsto w_*$ at T_* , then $V_T \rightarrow V_T|_{\phi \approx w} \approx \frac{c_h}{2} (T^2 - T_{ew}^2) h^2 + \frac{\lambda_h}{4} h^4$

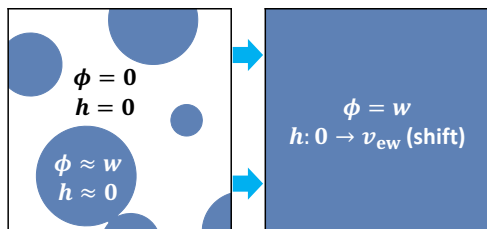
Depending on the sign of $(T_*^2 - T_{ew}^2)$, two sub-types

$$T_{ew} = \frac{m_h}{\sqrt{2c_h}} \sim 140 \text{ GeV}$$

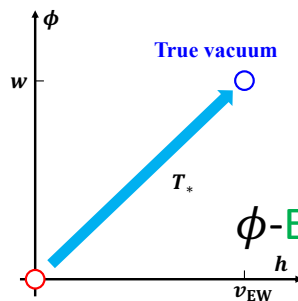
- Type-N1: $T_* > T_{ew}$



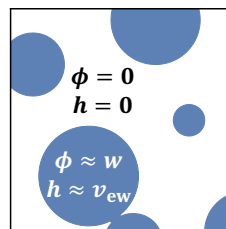
ϕ -FOPT
followed by
EW crossover



- Type-N2: $T_* < T_{ew}$



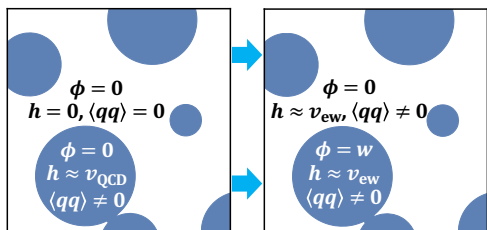
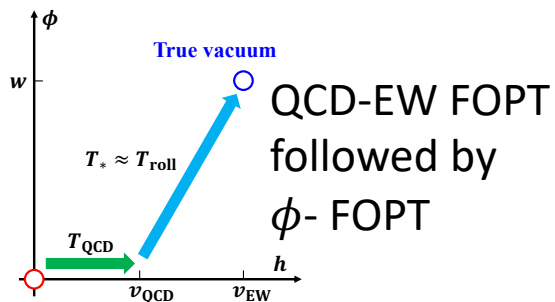
ϕ -EW FOPT



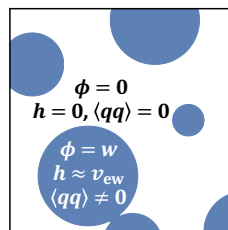
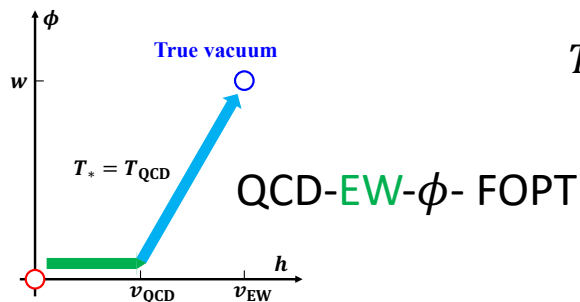
Backup: the inverted pattern

QCD FOPT $\langle \bar{q}q \rangle: 0 \mapsto \mathcal{O}(100 \text{ MeV})$ at T_{QCD} , **triggering** EWPT $h: 0 \mapsto v_{\text{QCD}}$ **via**
 $-y_t h \langle \bar{t}t \rangle$, $V_T \rightarrow V_T|_{h \approx v_{\text{QCD}}} \approx \frac{B}{4} \phi^4 \left(\log \frac{\phi}{w} - \frac{1}{4} \right) + \frac{c\phi}{2} (T^2 - T_{\text{roll}}^2) \phi^2$

- **Type-I1:** $T_{\text{OCD}} > T_{\text{roll}}$



- **Type-I2:** $T_{\text{OCD}} < T_{\text{roll}}$



$$T_{\text{QCD}} \approx 85 \text{ MeV}$$

$$v_{\text{QCD}} \approx 100 \text{ MeV}$$

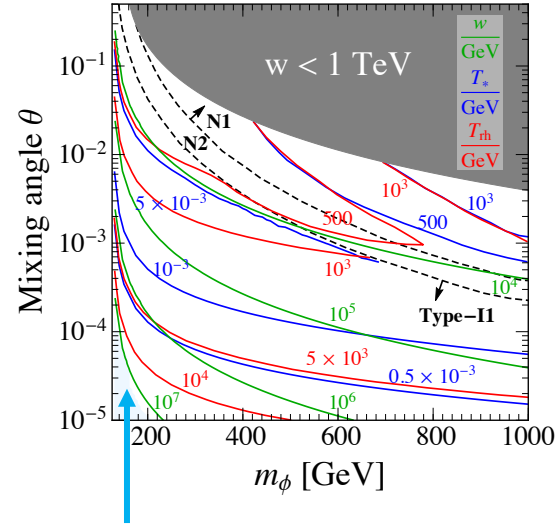
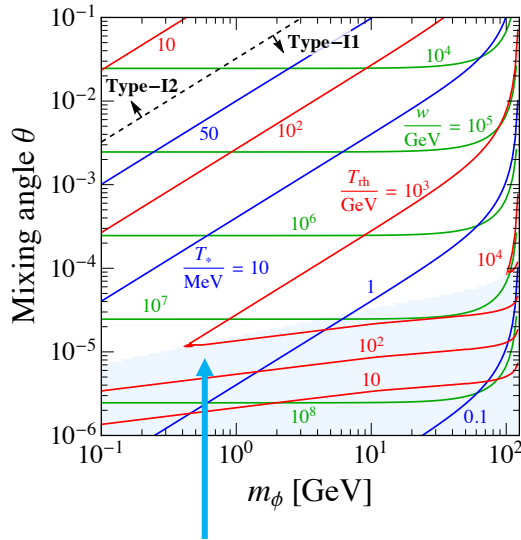
$$T_{\text{roll}} = \frac{m_h v_{\text{QCD}}}{\sqrt{2c\phi w}}$$

First proposed by Witten
 Witten, NPB 177, 477 (1981)

Backup: complete phase diagrams

The Universe is reheated to $T_{\text{rh}} \gtrsim T_*$ [Liu and KPX, PRD 110 (2024) 11, 115001]

- The EW symmetry might be **restored**, leading to a **second EWPT**

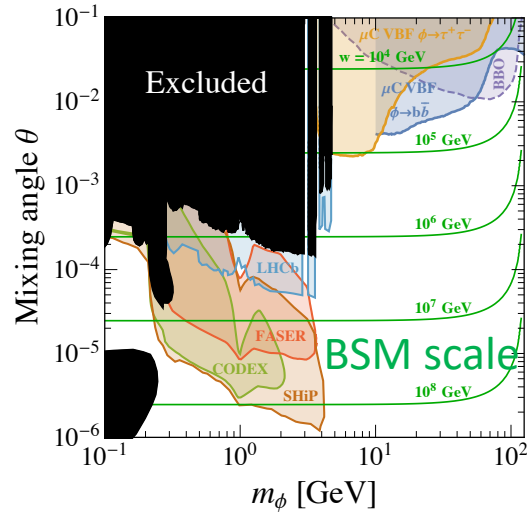


- If the ϕ decay rate is small, reheating is slow: early matter domination
- More detailed phase diagram can be plotted in principle

Backup: phenomenology

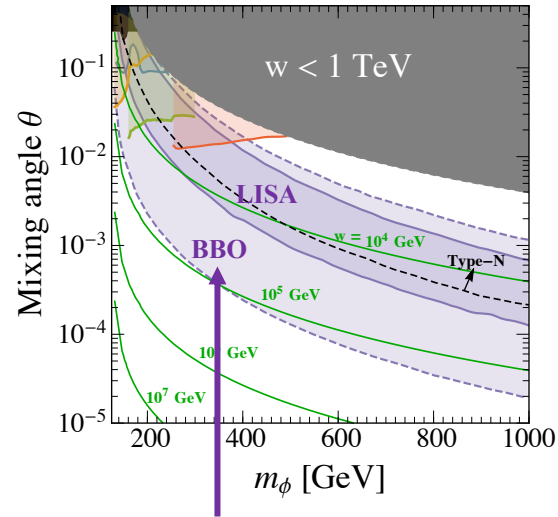
Collider search [Liu and KPX, PRD 110 (2024) 11, 115001]

- HL-LHC $\phi \rightarrow ZZ$; 10 TeV muon collider VBF $\phi \rightarrow b\bar{b}, VV, hh$
- Long-lived particle search



HL-LHC $\phi \rightarrow ZZ$

10 TeV MC VBF $\phi \rightarrow b\bar{b}, VV, hh$



Gravitational wave search: LISA (also TianQin, Taiji) and BBO