



Neutrino Mass & Mass Ordering with Cosmic Gravitational Focusing

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SFG, Pedro Pasquini, Liang Tan, **JCAP 05 (2024) 108** [arXiv:2312.16972]
SFG & Liang Tan [arXiv:2409.11115]



Pedro Pasquini



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UNSW, Sydney
December 13, 2024

李改道研究所
Tsung-Dao Lee Institute

Georg G. Raffelt

Stars as Laboratories for Fundamental Physics

The Astrophysics of Neutrinos, Axions, and Other
Weakly Interacting Particles

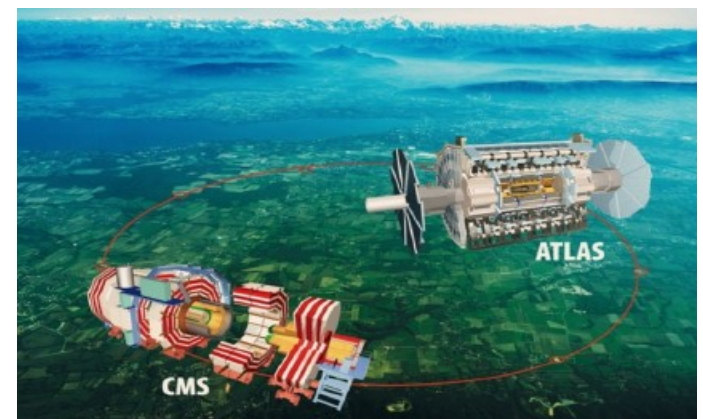
In the standard model, neutrinos have been assigned the most minimal properties compatible with experimental data: zero mass, zero charge, zero dipole moments, zero decay rate, zero almost everything.

Importance of Neutrino Masses

- **Higgs boson** \Rightarrow electroweak symmetry breaking & mass. $\sim O(100)\text{GeV}$
- **Chiral symmetry breaking** \Rightarrow majority of mass.
- The world seems not affected by the tiny neutrino mass?
 - Neutrino mass \Rightarrow Mixing
 - 3 Neutrino \Rightarrow possible **CP violation**
 - CP violation \Rightarrow **Leptogenesis**
 - \Rightarrow **Matter-Antimatter Asymmetry**
 - There is something left in the Universe.
 - **EW Baryogenesis** is not enough.



Daya Bay @ **March 8, 2012**



LHC @ **July 4, 2012**

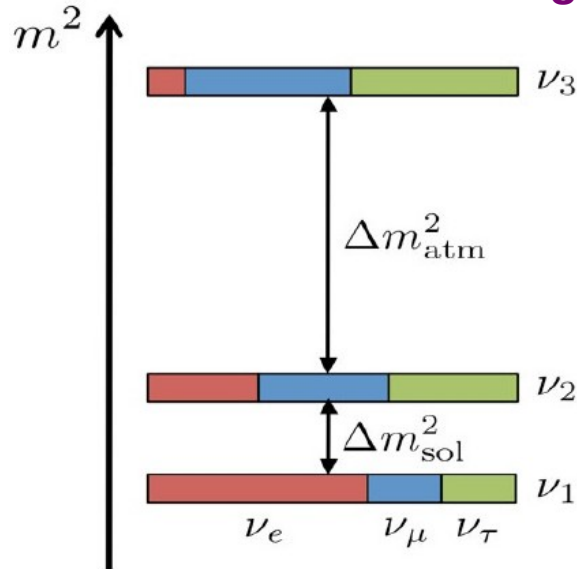
Global Fit of Oscillation Parameters

(for NO)	-1σ	Best Value	$+1\sigma$
$\Delta m_s^2 \equiv \Delta m_{21}^2$ (10^{-5}eV^2)	7.30	7.50	7.72
$ \Delta m_a^2 \equiv \Delta m_{31}^2 $ (10^{-3}eV^2)	2.52	2.56	2.59
$\sin^2 \theta_s$ ($\theta_s \equiv \theta_{12}$)	0.302 (33.3°)	0.318 (34.3°)	0.334 (35.3°)
$\sin^2 \theta_a$ ($\theta_a \equiv \theta_{23}$)	0.544 (47.54°)	0.566 (48.79°)	0.582 (49.72°)
$\sin^2 \theta_r$ ($\theta_r \equiv \theta_{13}$)	0.02147 (8.43°)	0.02225 (8.58°)	0.02280 (8.69°)
δ_D	191°	216°	257°
δ_{Mi}	??	??	??

Salas, Forero, Gariazzo, Martinez-Mirave, Mena, Ternes, Tortola & Valle, [arXiv:2006.11237]

ν Mass Ordering

Normal Ordering (正序)



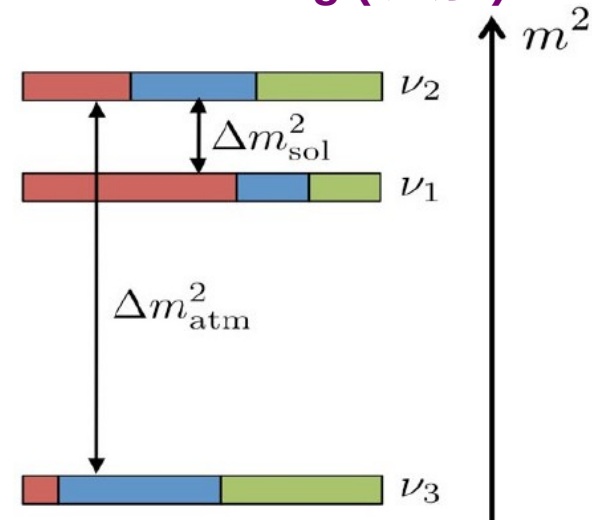
$$m_1 \lesssim m_2 < m_3$$

m_1

$$m_2 = \sqrt{m_1^2 + \Delta m_s^2}$$

$$m_3 = \sqrt{m_1^2 + \Delta m_a^2}$$

Inverted Ordering (反序)



$$m_3 < m_1 \lesssim m_2$$

$$m_1 = \sqrt{m_3^2 + \Delta m_a^2}$$

$$m_2 = \sqrt{m_3^2 + \Delta m_a^2 + \Delta m_s^2}$$

m_3

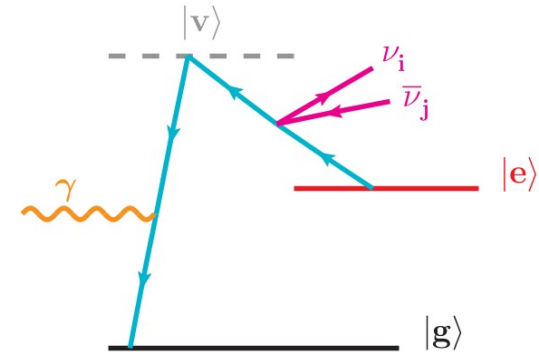
江门中微子实验 (JUNO)

However, ν oscillation experiment cannot measure the absolute mass!

ν Oscillation

Beta Decay

Radiative Emission of ν Pairs



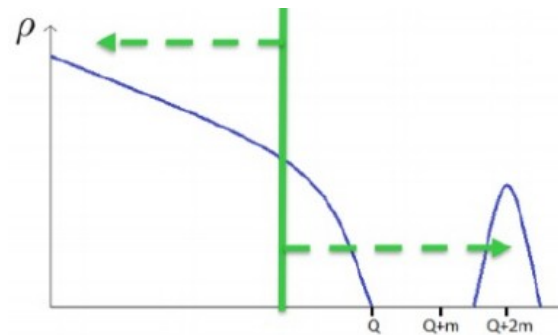
SFG, Pedro Pasquini, **JHEP 12 (2023) 083** [arXiv:2306.12953]
Phys.Lett.B 841 (2023) 137911 [arXiv:2206.11717]
Eur.Phys.J.C 82 (2022) 3, 208 [arXiv:2110.03510]

Supernova ν Time Delay

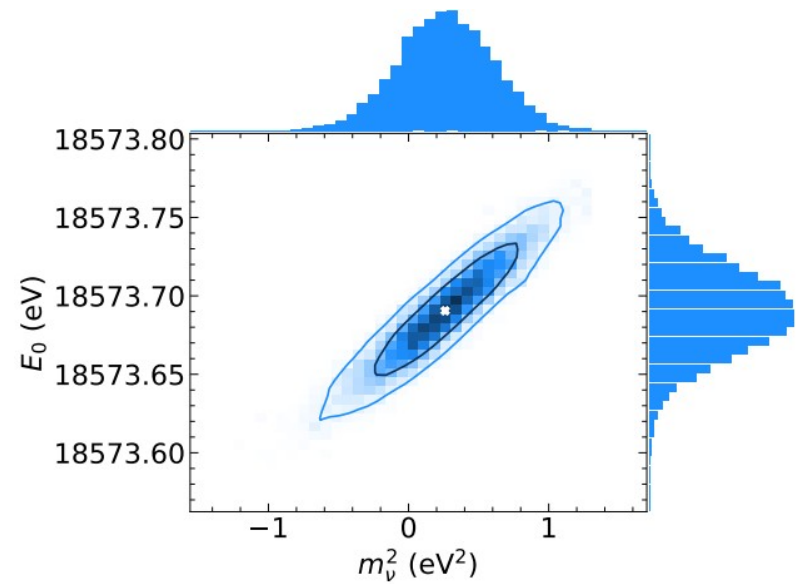
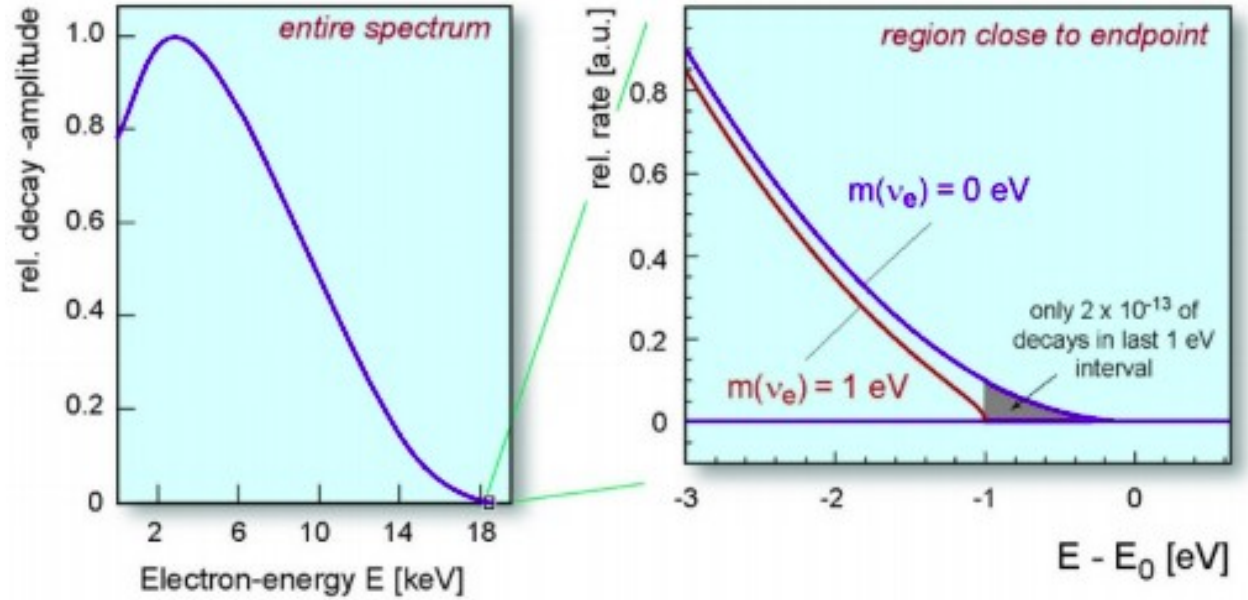
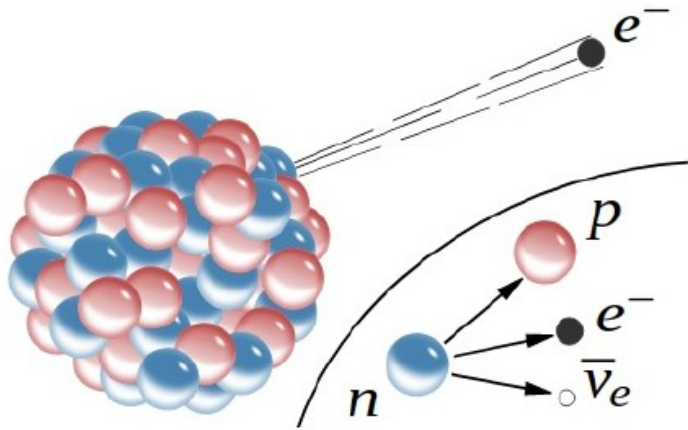
SFG, Chui-Fan Kong, Alexei Smirnov, **PRL 133 (2024) 121802** [arXiv:2404.17352]

CvB Detection

CMB & LSS

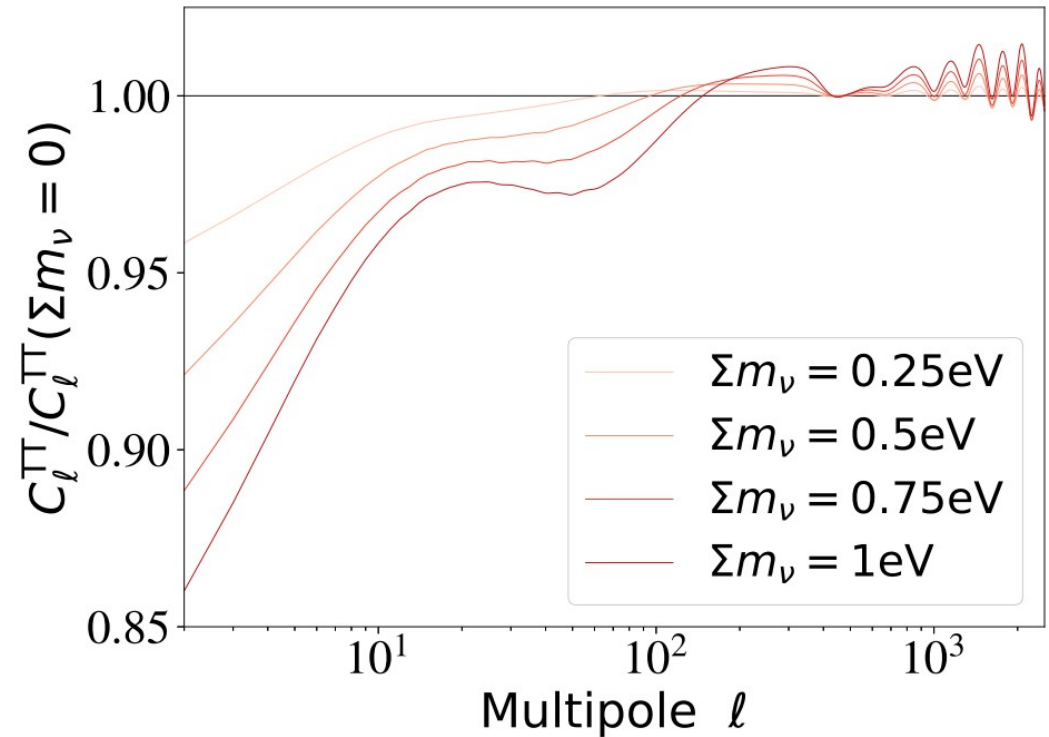
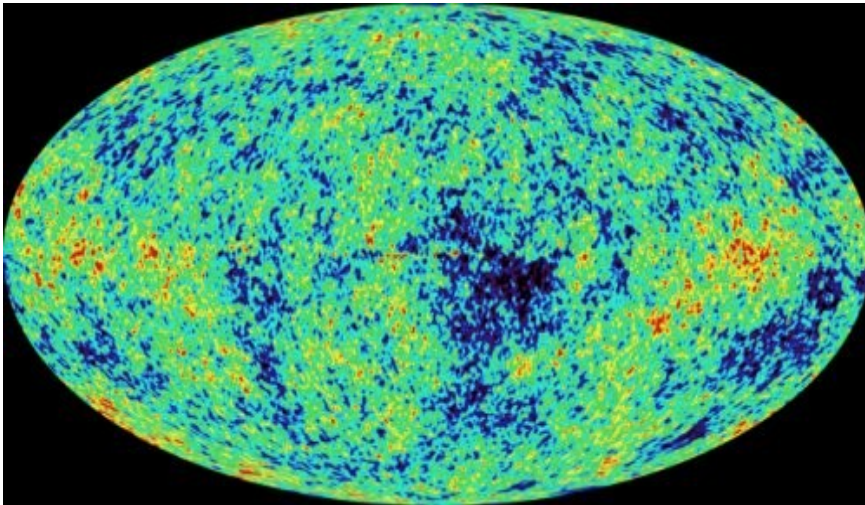


Neutrino Mass @ β Decay



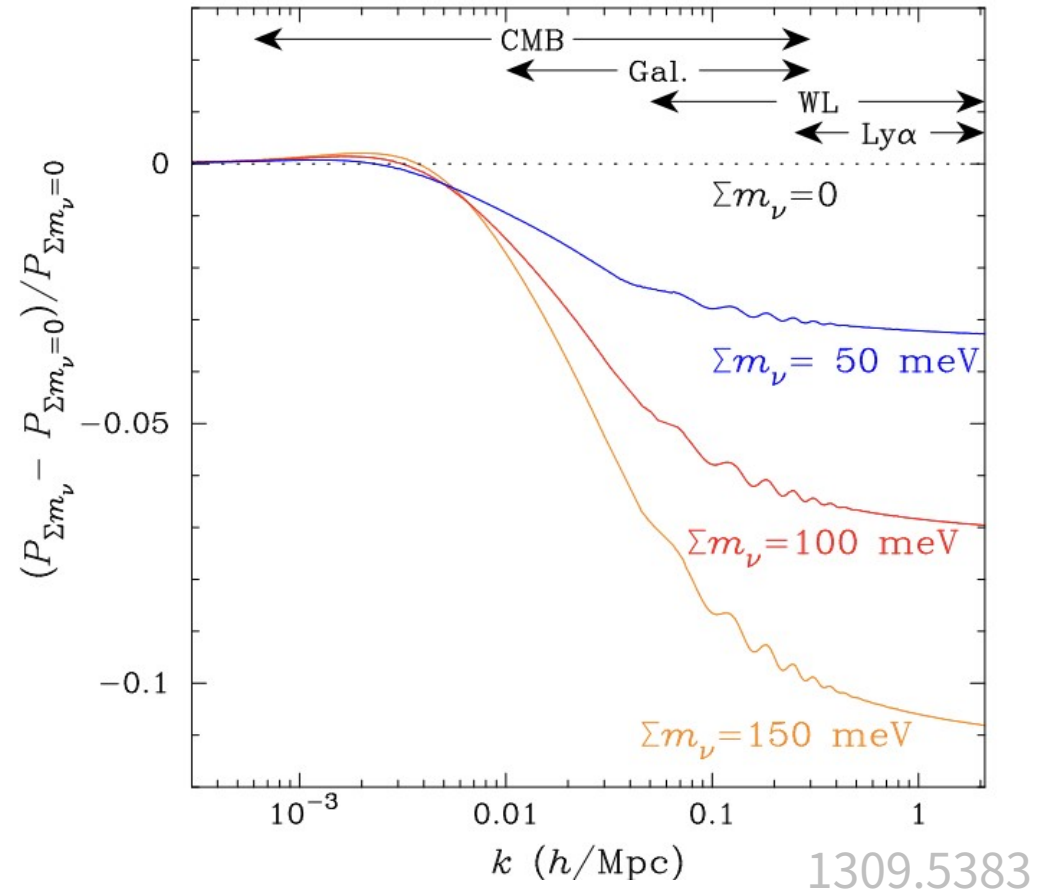
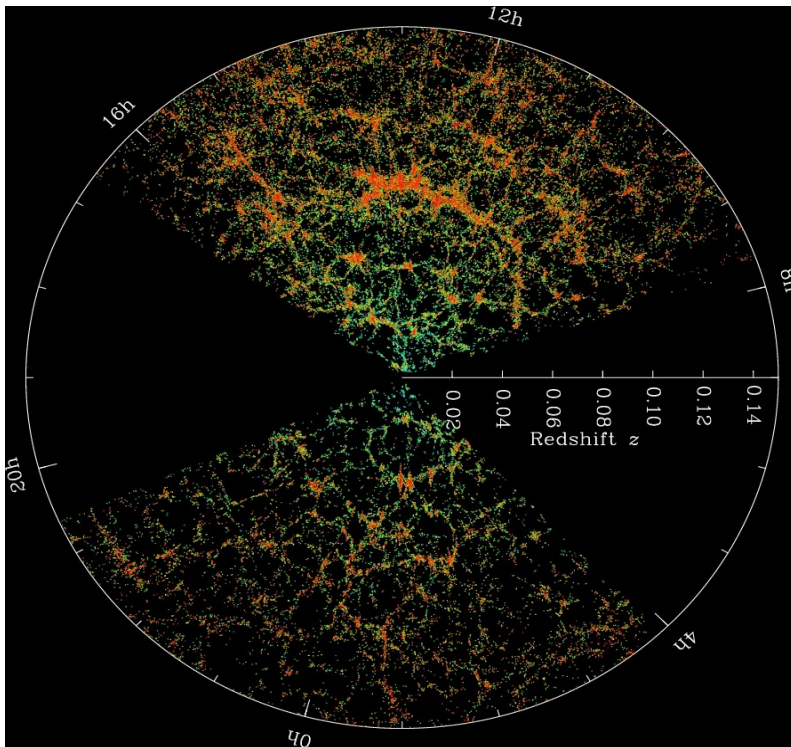
Nature Phys. 18 (2022) 2, 160-166

Massive neutrino decreasing CMB lensing power spectrum



$$C_l^{\kappa\kappa} \propto (\Omega_m h^2)^2 A_s \left(1 - 0.02 \frac{f_\nu}{4 \times 10^{-3}} \right) \quad f_\nu = \Omega_\nu / \Omega_m$$

Suppress matter power spectrum



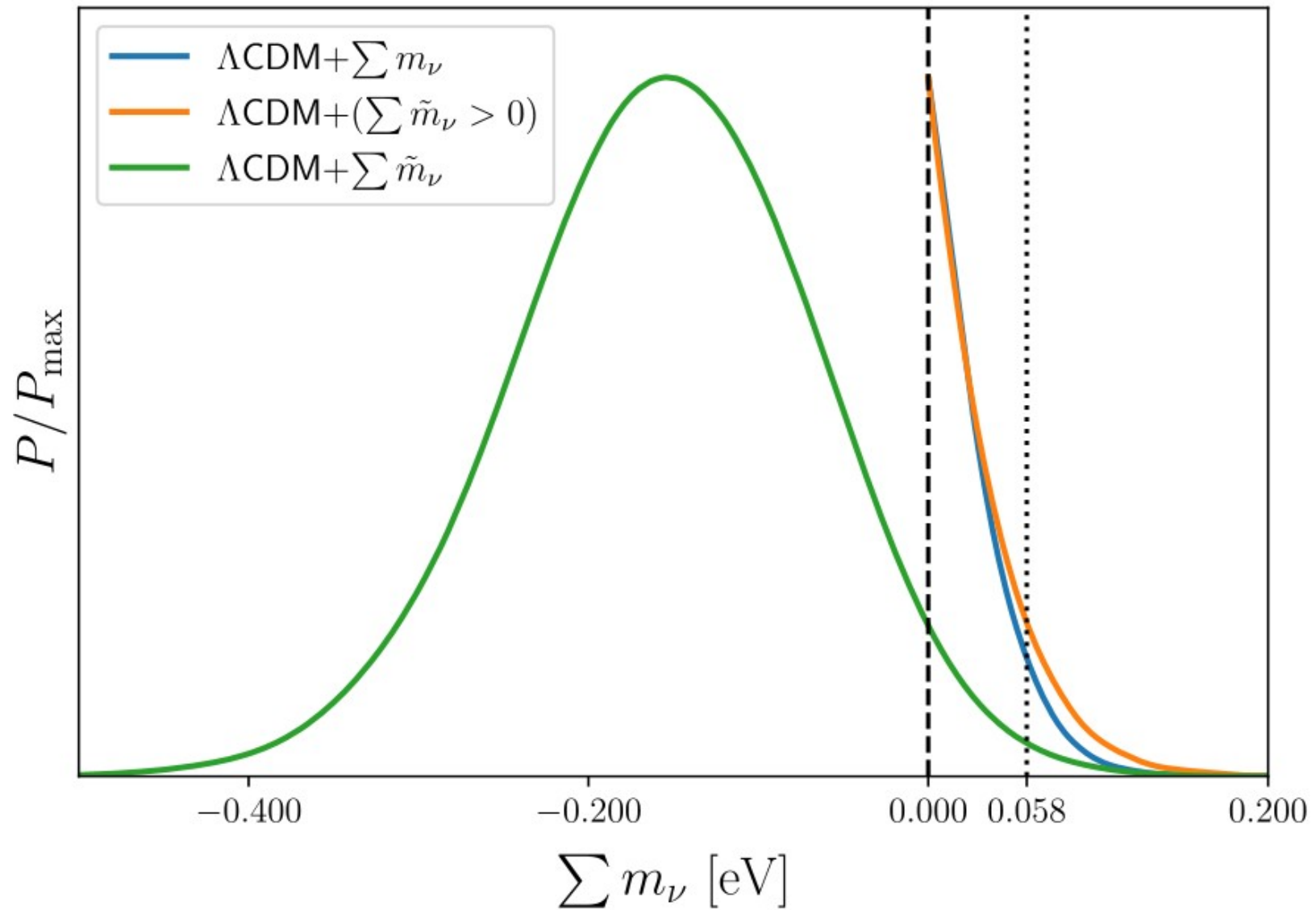
$$P(\Sigma m_\nu)(k \gg k_{\text{fs}}, z) \approx \left(1 - 2f_\nu - \frac{6}{5}f_\nu \log \frac{1+z_\nu}{1+z} \right) P(\Sigma m_\nu=0)(k \gg k_{\text{fs}}, z)$$

Cosmological Systematics

model / dataset	Ω_m	H_0 [km s ⁻¹ Mpc ⁻¹]	Σm_ν [eV]	N_{eff}
ΛCDM + Σm_ν				
DESI+CMB	0.3037 ± 0.0053	68.27 ± 0.42	< 0.072	—
ΛCDM + N_{eff}				
DESI+CMB	0.3058 ± 0.0060	68.3 ± 1.1	—	3.10 ± 0.17
wCDM + Σm_ν				
DESI+CMB	0.282 ± 0.013	$71.1^{+1.5}_{-1.8}$	< 0.123	—
DESI+CMB+Panth.	0.3081 ± 0.0067	67.81 ± 0.69	< 0.079	—
DESI+CMB+Union3	0.3090 ± 0.0082	67.72 ± 0.88	< 0.078	—
DESI+CMB+DESY5	0.3152 ± 0.0065	67.01 ± 0.64	< 0.073	—
wCDM + N_{eff}				
DESI+CMB	0.281 ± 0.013	$71.0^{+1.6}_{-1.8}$	—	2.97 ± 0.18
DESI+CMB+Panth.	0.3090 ± 0.0068	67.9 ± 1.1	—	3.07 ± 0.18
DESI+CMB+Union3	0.3097 ± 0.0084	67.8 ± 1.2	—	3.06 ± 0.18
DESI+CMB+DESY5	0.3163 ± 0.0067	67.2 ± 1.1	—	3.09 ± 0.18
$w_0 w_a$CDM + Σm_ν				
DESI+CMB	$0.344^{+0.032}_{-0.026}$	$64.7^{+2.1}_{-3.2}$	< 0.195	—
DESI+CMB+Panth.	0.3081 ± 0.0069	68.07 ± 0.72	< 0.155	—
DESI+CMB+Union3	0.3240 ± 0.0098	66.48 ± 0.94	< 0.185	—
DESI+CMB+DESY5	0.3165 ± 0.0069	67.22 ± 0.66	< 0.177	—

DESI [arXiv:2404.03002]

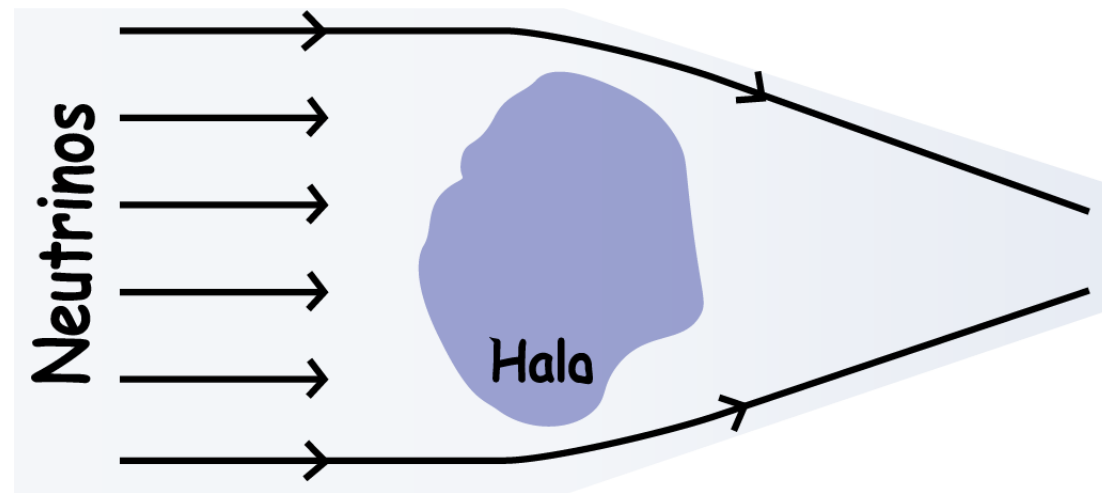
Results from DESI BAO + CMB



Craig, Green, Meyers & Rajendran [arXiv:2405.00836]

Cosmic ν Fluid (CvF) vs DM Halo

Gravitational attraction between DM halo & CvF

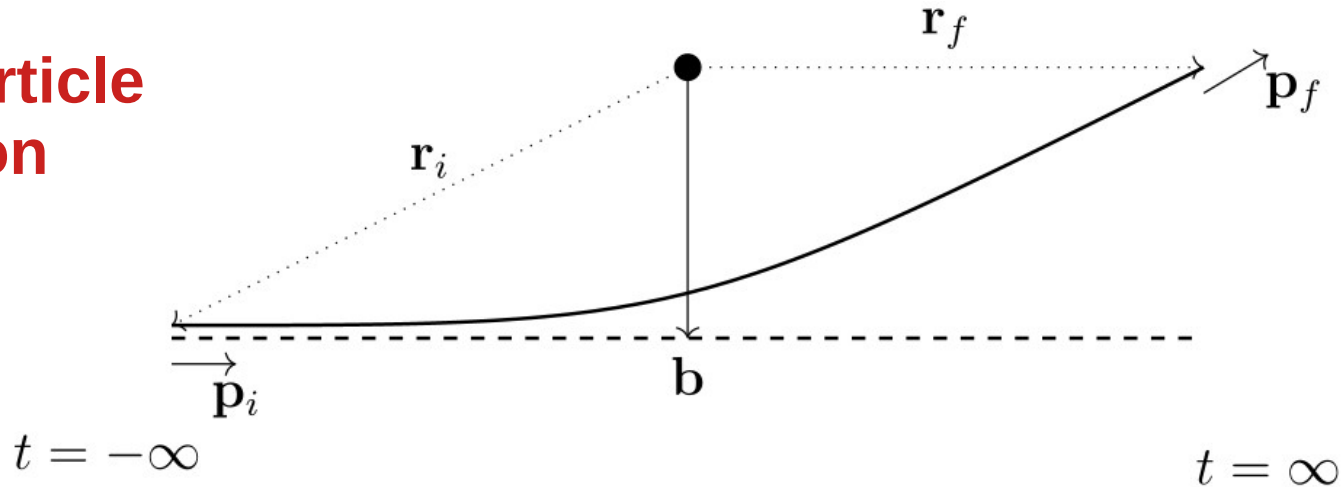


CvF

DM
Halo

vs Stone in Water Flow

Single-particle description



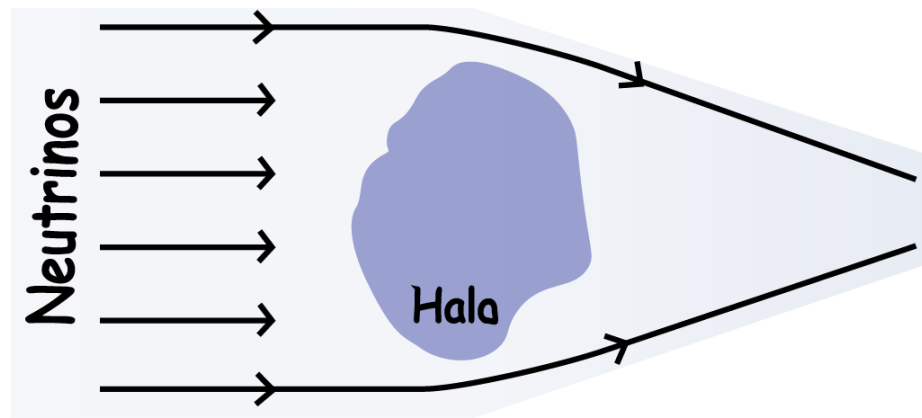
$$ds^2 \equiv - \left(1 - \frac{2GM}{r} \right) dt^2 + \left(1 - \frac{2GM}{r} \right)^{-1} dr^2 + r^2 d\Omega^2$$

$$\Delta\phi = 2|\mathbf{b}| \int_{r_{\min}}^{\infty} \frac{dr}{r^2} \left[1 + \frac{2b_{90}}{r} - \frac{|\mathbf{b}|^2}{r^2} + \frac{2GM|\mathbf{b}|^2}{r^3} \right]^{-1/2}$$

$$b_{90} \equiv GMm_{\nu}^2/|\mathbf{p}_i|^2 \quad r_{\min}^3 + 2b_{90}r_{\min}^2 - |\mathbf{b}|^2r_{\min} + 2GM|\mathbf{b}|^2 = 0$$

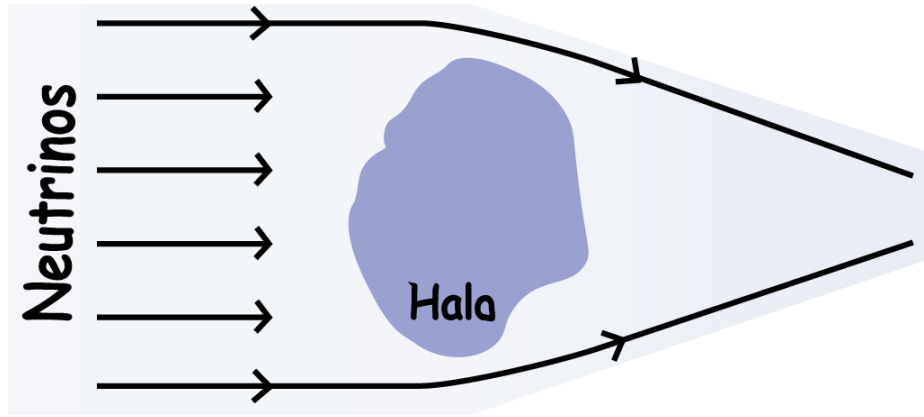
SFG, Pedro Pasquini, Liang Tan, **JCAP 05 (2024) 108** [arXiv:2312.16972]

Single-particle description



$$\Delta\phi \approx \pi + 2\frac{GM}{|\mathbf{b}|} \left(\frac{m_\nu^2}{|\mathbf{p}_i|^2} + 2 \right) \quad f_\nu(\mathbf{p}_i, \mathbf{v}) \approx \frac{2}{e^{|\mathbf{p}_i - E_{\mathbf{p}_i} \mathbf{v}/c|/T} + 1}$$

$$\Delta\mathbf{p}^\parallel \equiv \mathbf{p}_f^\parallel - \mathbf{p}_i = (-\cos \Delta\phi - 1) \mathbf{p}_i \approx -\frac{2G^2 M^2}{|\mathbf{b}|^2} \left(\frac{m_\nu^2}{|\mathbf{p}_i|^2} + 2 \right)^2 \mathbf{p}_i$$



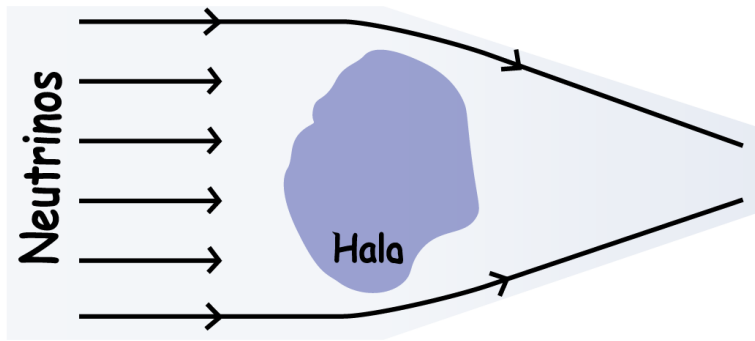
**Boltzmann
description**

$$f_\nu(\mathbf{x}, \mathbf{p}) \equiv \bar{f}_\nu(\mathbf{p}) + \delta f_\nu(\mathbf{x}, \mathbf{p})$$

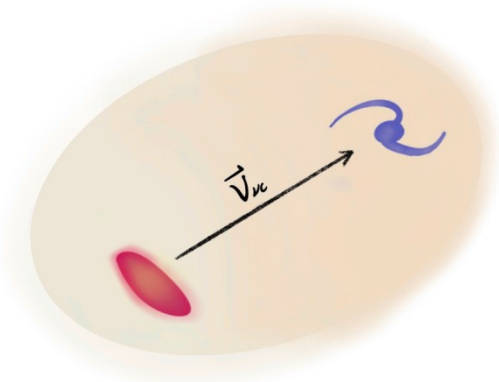
$$\left\{ \partial_t + \frac{\mathbf{p} \cdot \nabla_{\mathbf{x}}}{aE_{\mathbf{p}}} - \left[(H + \dot{\Phi})\mathbf{p} + \frac{E_{\mathbf{p}}}{a} \nabla_{\mathbf{x}} \Psi - \frac{|\mathbf{p}|^2 \nabla_{\mathbf{x}} \Phi - \mathbf{p}(\mathbf{p} \cdot \nabla_{\mathbf{x}} \Phi)}{aE_{\mathbf{p}}} \right] \cdot \nabla_{\mathbf{p}} \right\} f_\nu(\mathbf{x}, \mathbf{p}) = 0$$

$$\delta \tilde{f}_\nu(\mathbf{k}, \mathbf{p}) = \tilde{\Psi}(\mathbf{k}) \left(\frac{m_\nu^2 + 2\mathbf{p}^2}{\mathbf{p} \cdot \mathbf{k}} \mathbf{k} - \mathbf{p} \right) \cdot \nabla_{\mathbf{p}} \bar{f}_\nu(\mathbf{p})$$

$$\text{Im}[\delta \tilde{\rho}_\nu] = - \frac{(\mathbf{v}_{\nu c} \cdot \hat{\mathbf{k}}) \tilde{\Psi}}{4\pi} \int d|\mathbf{p}'| (m_\nu^4 + 3m_\nu^2 |\mathbf{p}'|^2 + 2|\mathbf{p}'|^4) \frac{d\bar{f}_\nu}{d|\mathbf{p}'|} \Theta(|\mathbf{p}'| - E_{\mathbf{p}'} |\mathbf{v}_{\nu c} \cdot \hat{\mathbf{k}}|)$$



$$\delta_m = \delta_{m0} + \sum_{i=1}^3 \frac{\delta\rho_{\nu i}}{\rho_m}$$



Density enhancement downwind!

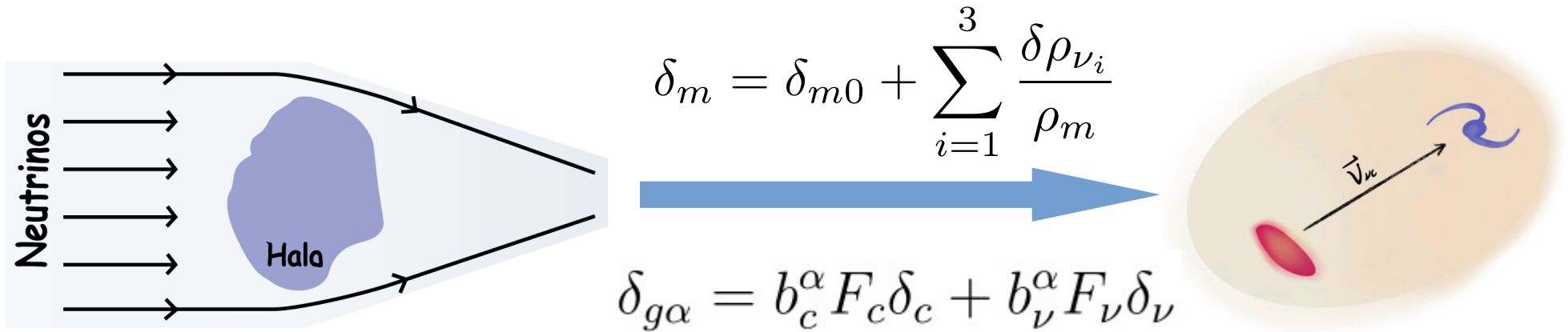
$$\delta\rho_\nu(-\mathbf{x}) = -\delta\rho_\nu(\mathbf{x})$$

$$\tilde{\delta}_m \equiv \tilde{\delta}_{m0}(1 + i\tilde{\phi})$$

$$[\tilde{A}(\mathbf{k})]^* = \int d\mathbf{x} e^{i\mathbf{k}\cdot\mathbf{x}} A(\mathbf{x}) = \int d\mathbf{x} e^{-i\mathbf{k}\cdot\mathbf{x}} A(-\mathbf{x})$$

$$= - \int d\mathbf{x} e^{-i\mathbf{k}\cdot\mathbf{x}} A(\mathbf{x}) = -\tilde{A}(\mathbf{k})$$

SFG, Pedro Pasquini, Liang Tan [arXiv:2312.16972]

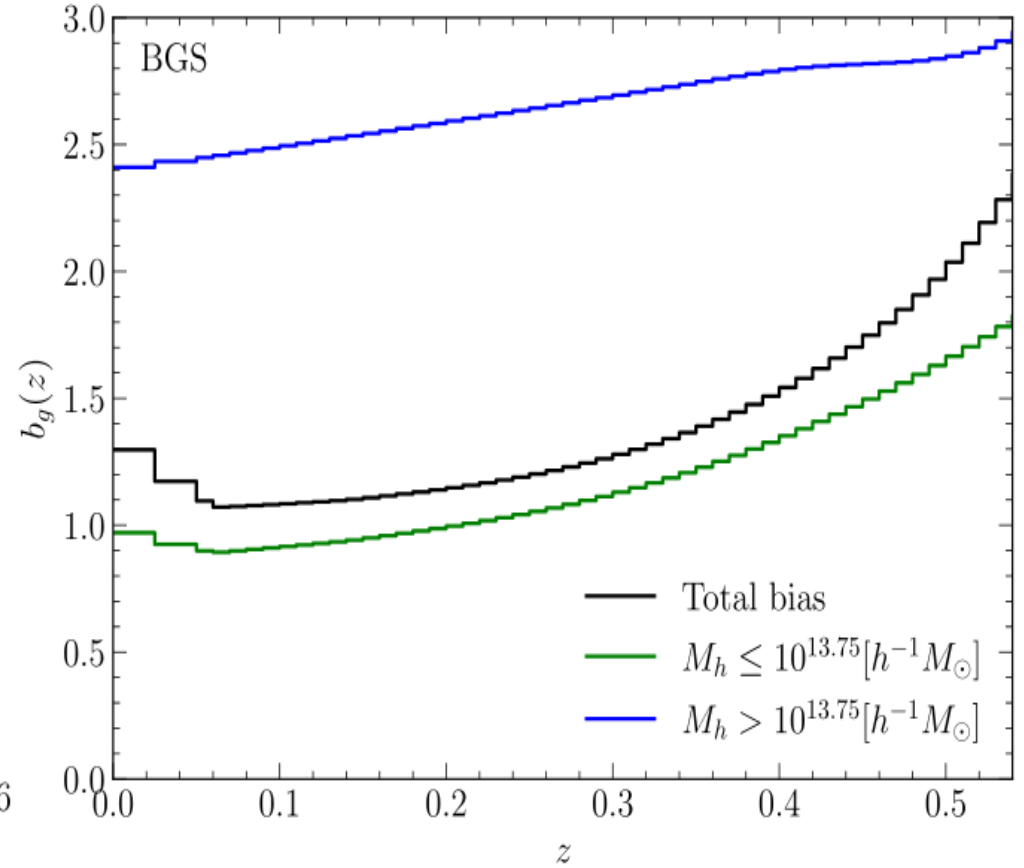
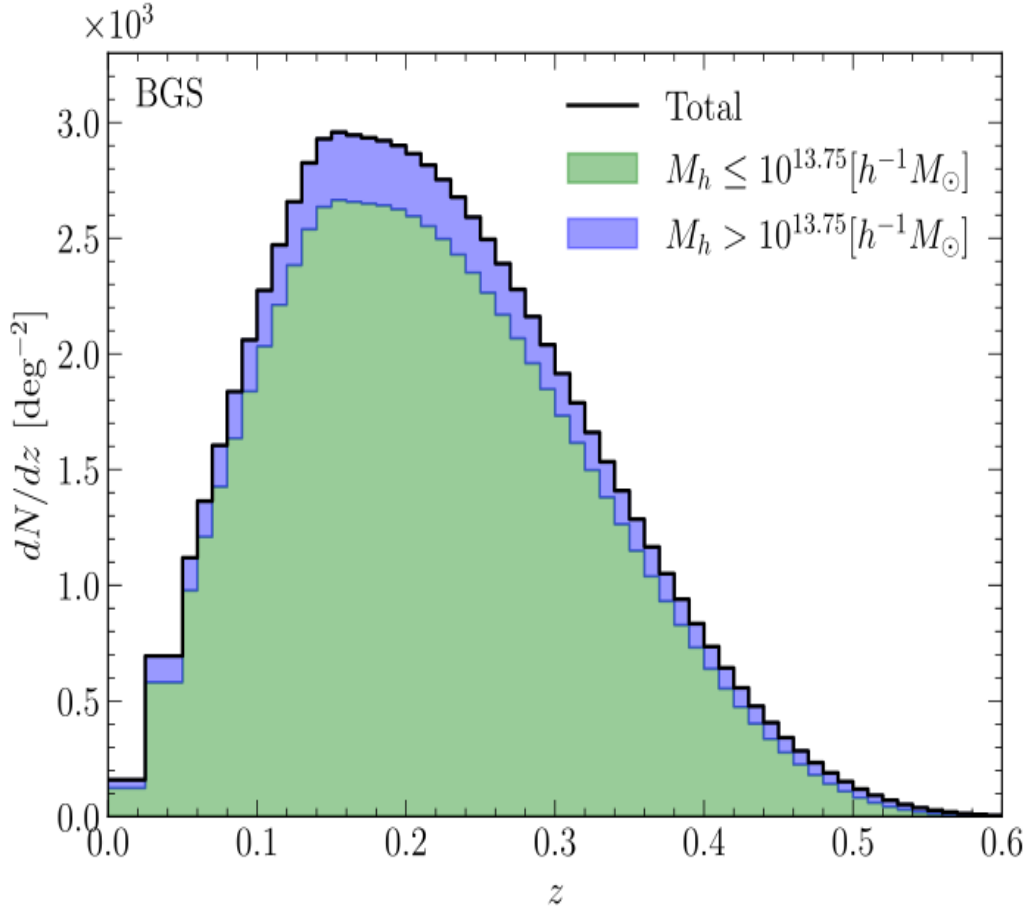


$$\delta_{g\alpha, \text{RSD}}(\mathbf{x}) \equiv \delta_{g\alpha}(\mathbf{x}) - \frac{\partial}{\partial x} \left(\frac{\mathbf{u}_m \cdot \hat{\mathbf{x}}}{aH} \right)$$

$$\text{Im}[\tilde{\delta}_{g\alpha, \text{RSD}} \tilde{\delta}_{g\beta, \text{RSD}}^*] = -i\Delta b \left[\mu_{\mathbf{k}}^2 \frac{\dot{\tilde{\phi}}}{H} + (f\mu_{\mathbf{k}}^2 + 1)\tilde{\phi} \right] \tilde{\delta}_{m0}^2$$

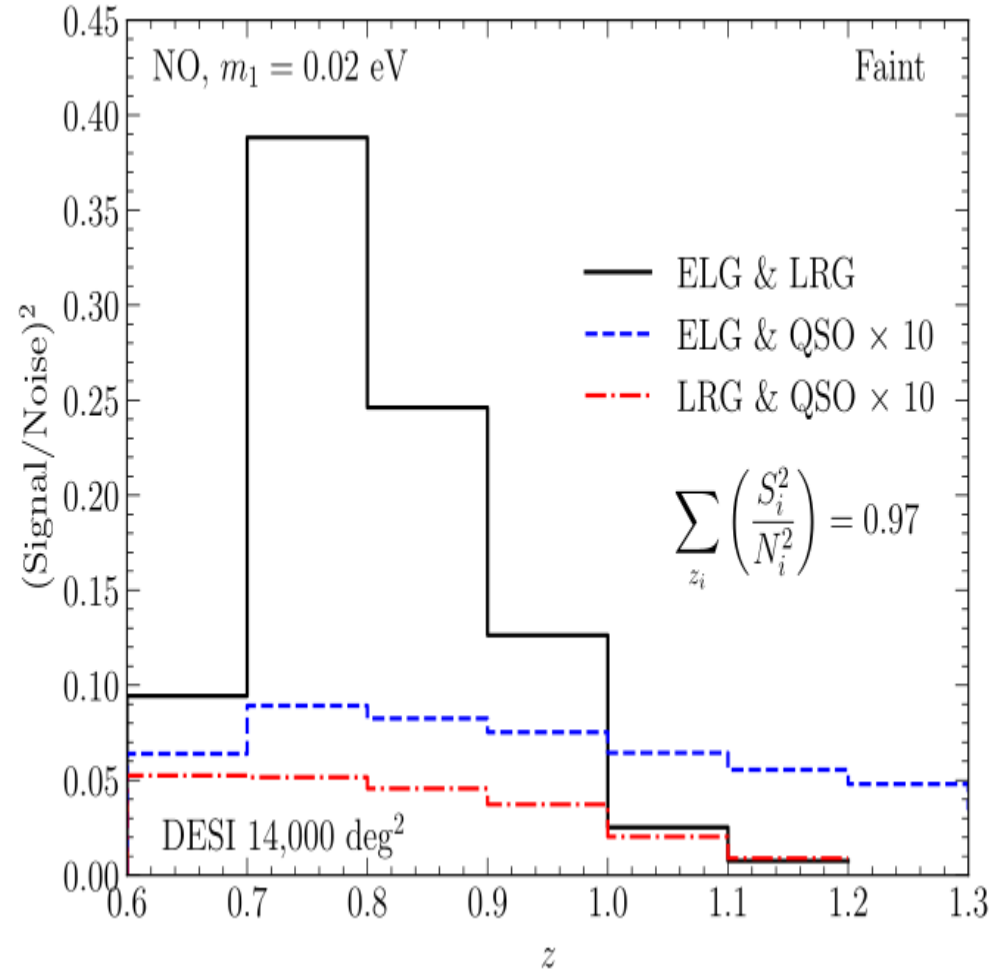
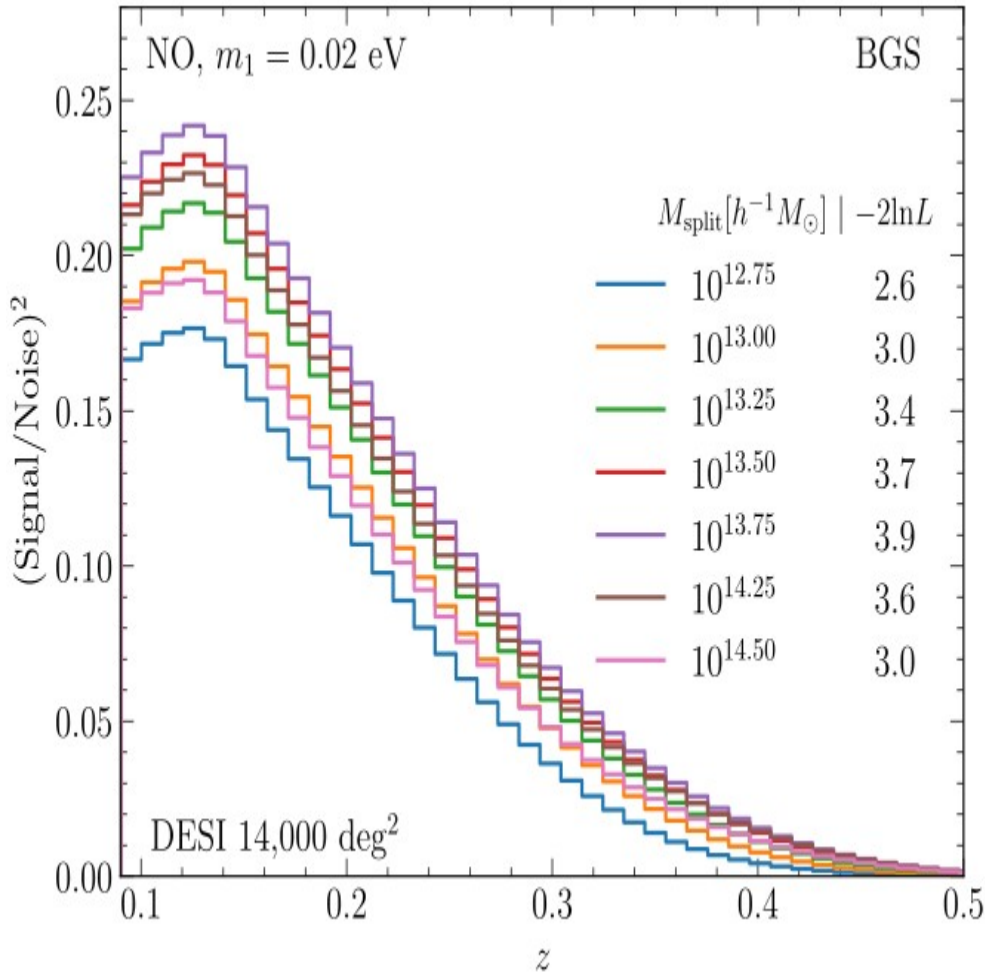
Galaxies with different bias

$$\Delta b \equiv b_c^\alpha - b_c^\beta$$



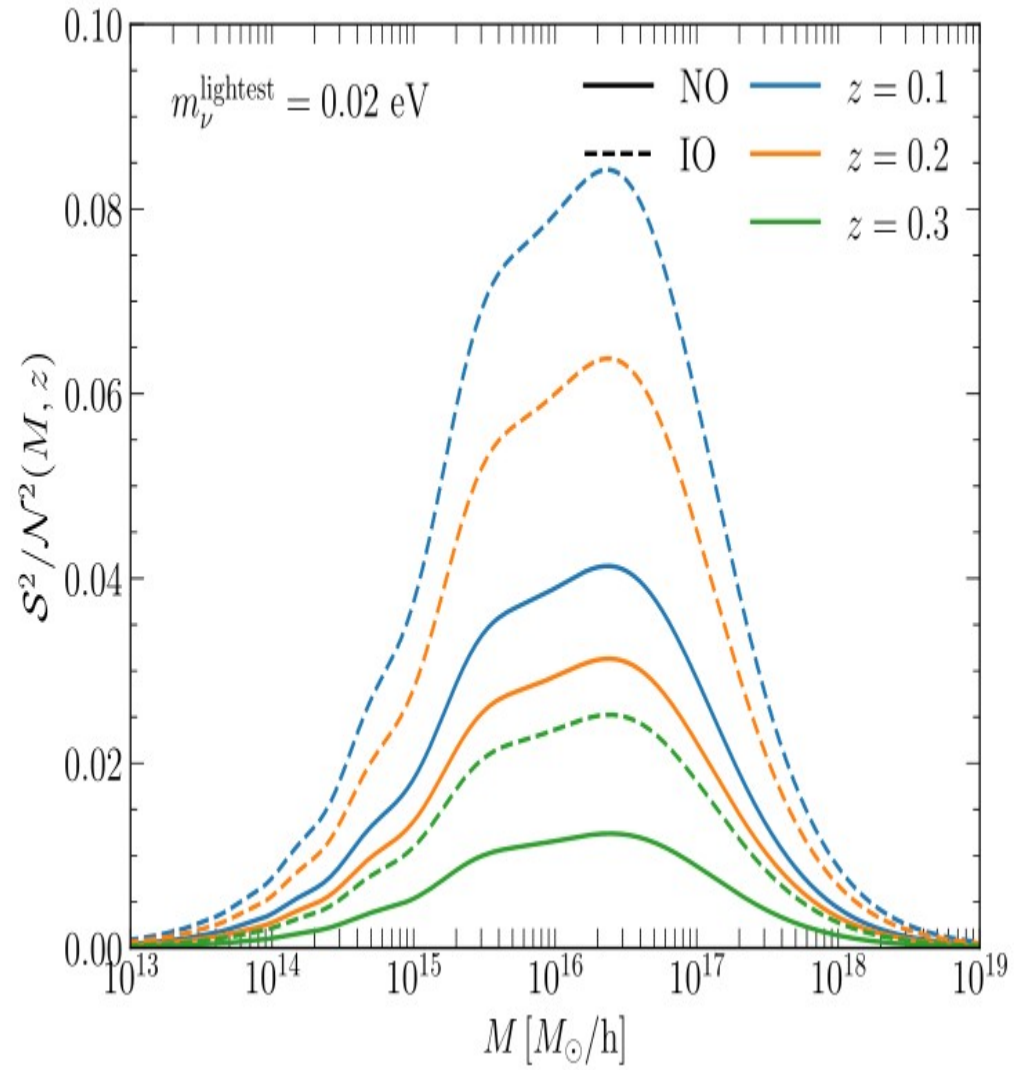
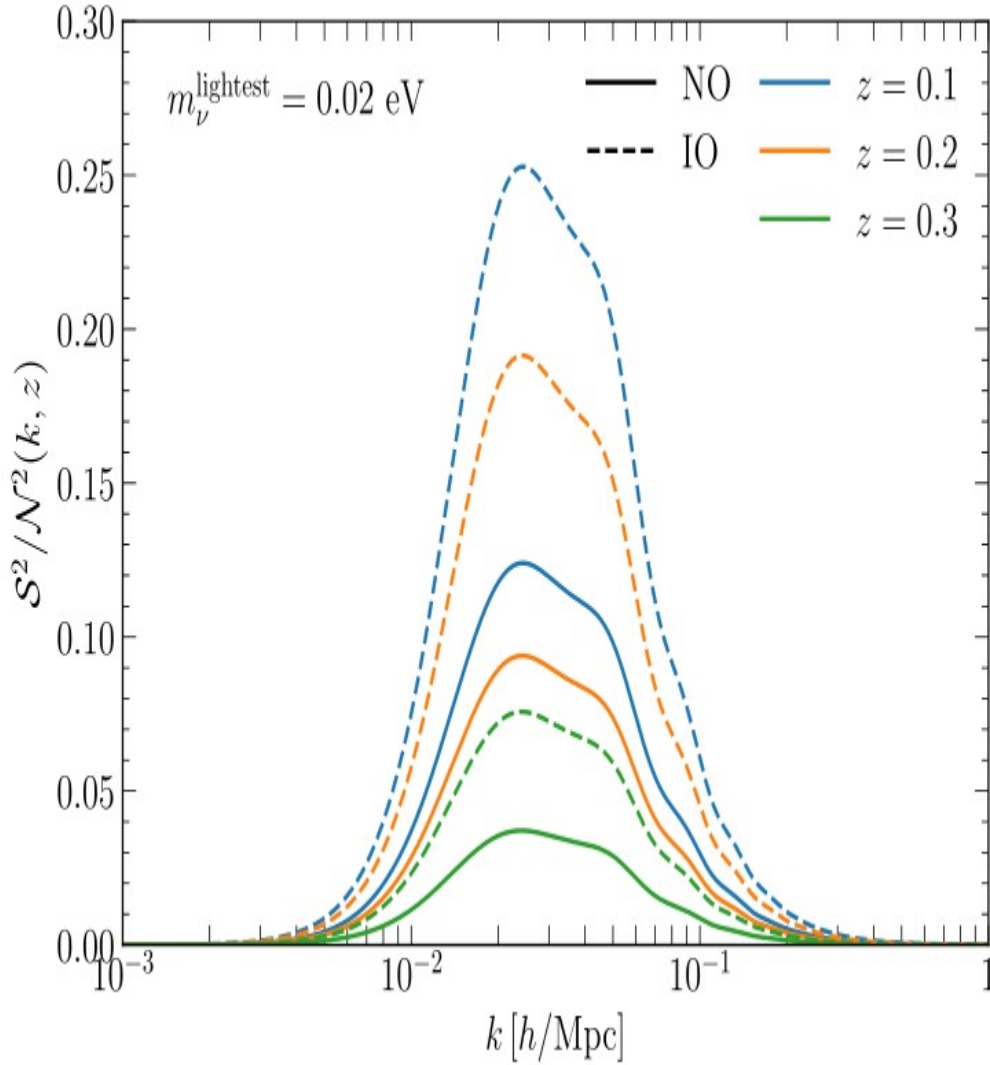
$$n_g(z) \equiv \int d \ln M_h \frac{dn(z)}{d \ln M_h} \langle N(M_h) \rangle \quad b_g(z) \equiv \frac{1}{n_g} \int d \ln M_h \frac{dn(z)}{d \ln M_h} \langle N(M_h) \rangle b_h(M_h, z)$$

SFG, Pedro Pasquini, Liang Tan, **JCAP 05 (2024) 108** [arXiv:2312.16972]



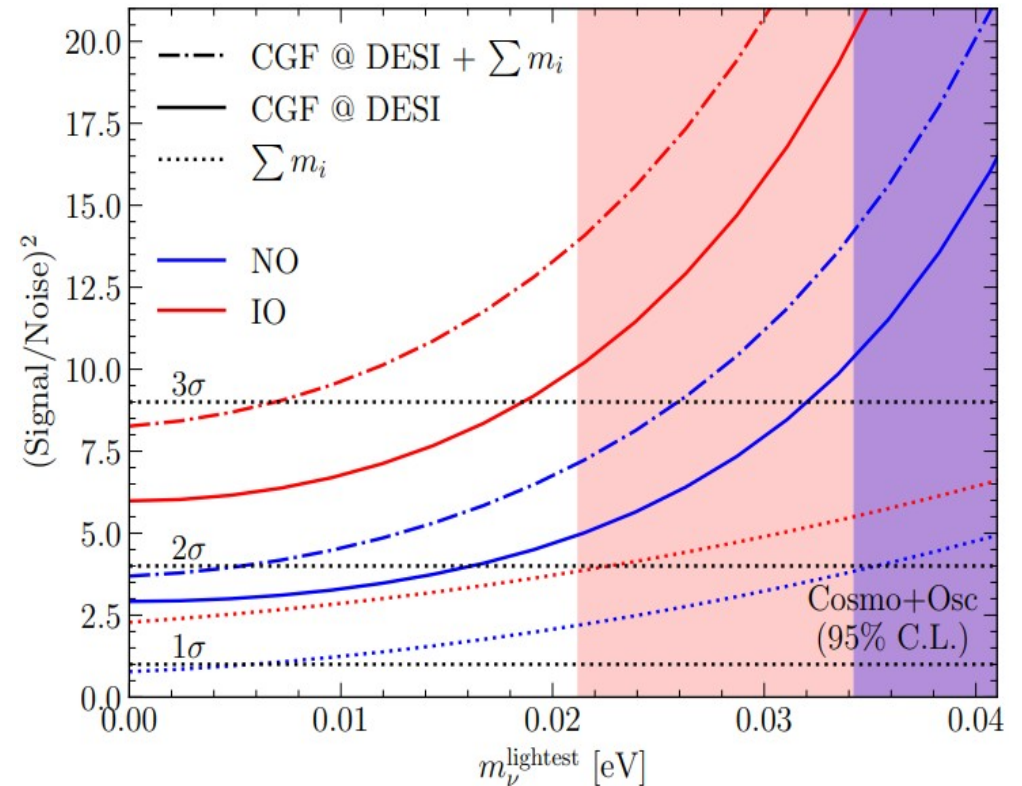
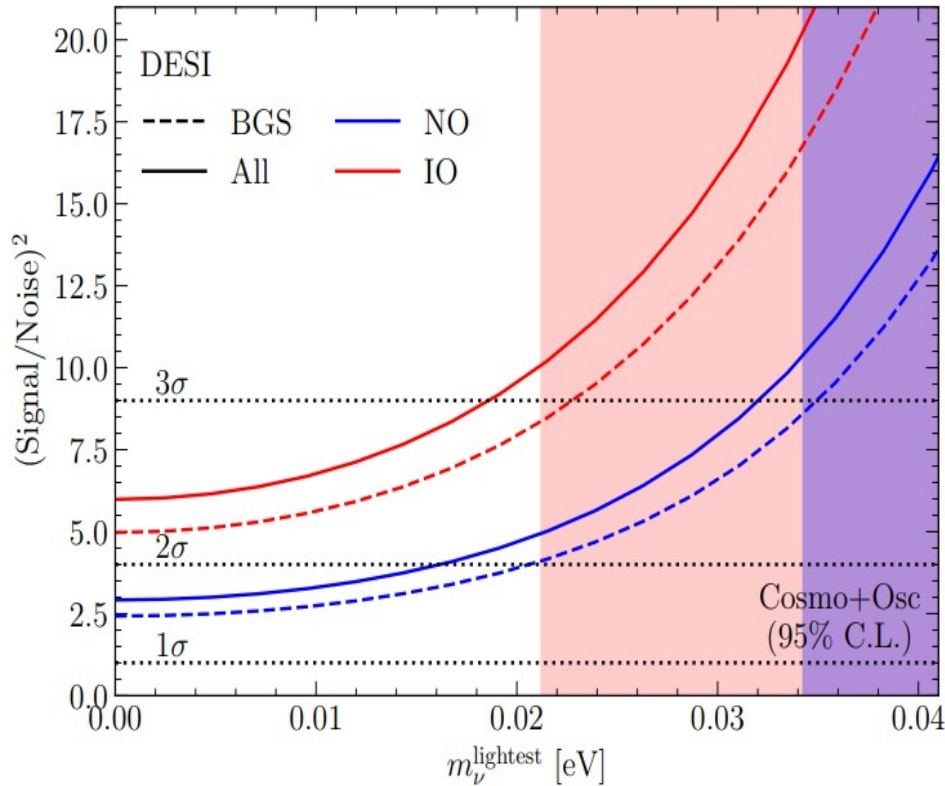
SFG, Pedro Pasquini, Liang Tan, **JCAP 05 (2024) 108** [arXiv:2312.16972]

Redshift Distribution



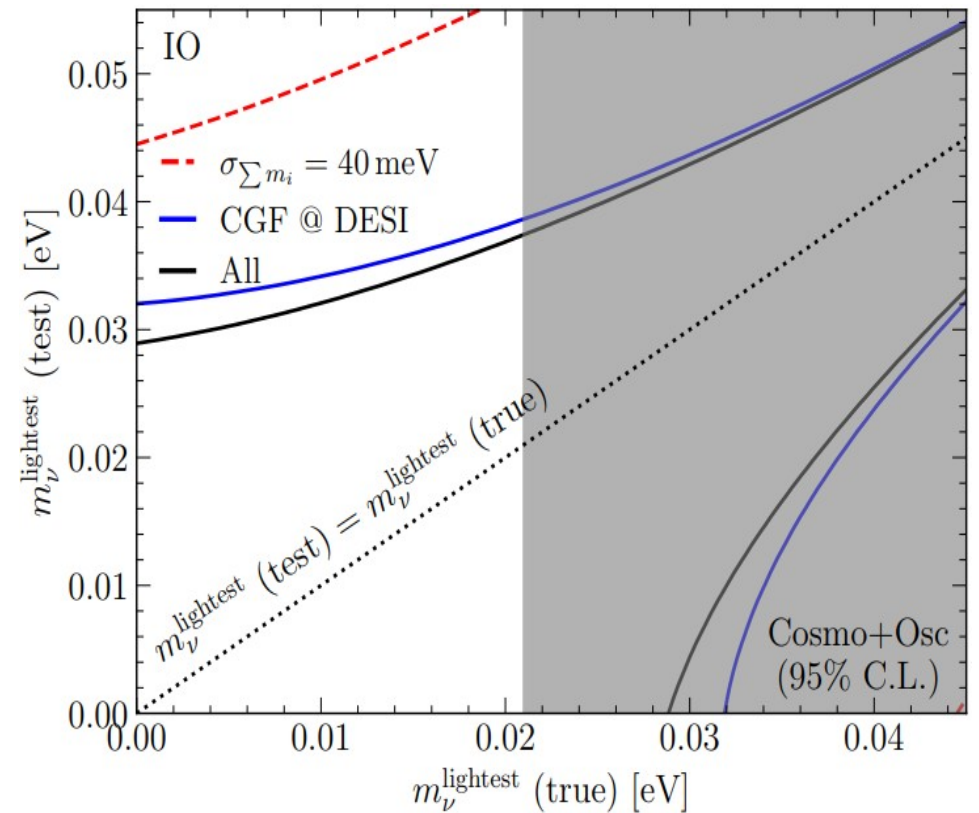
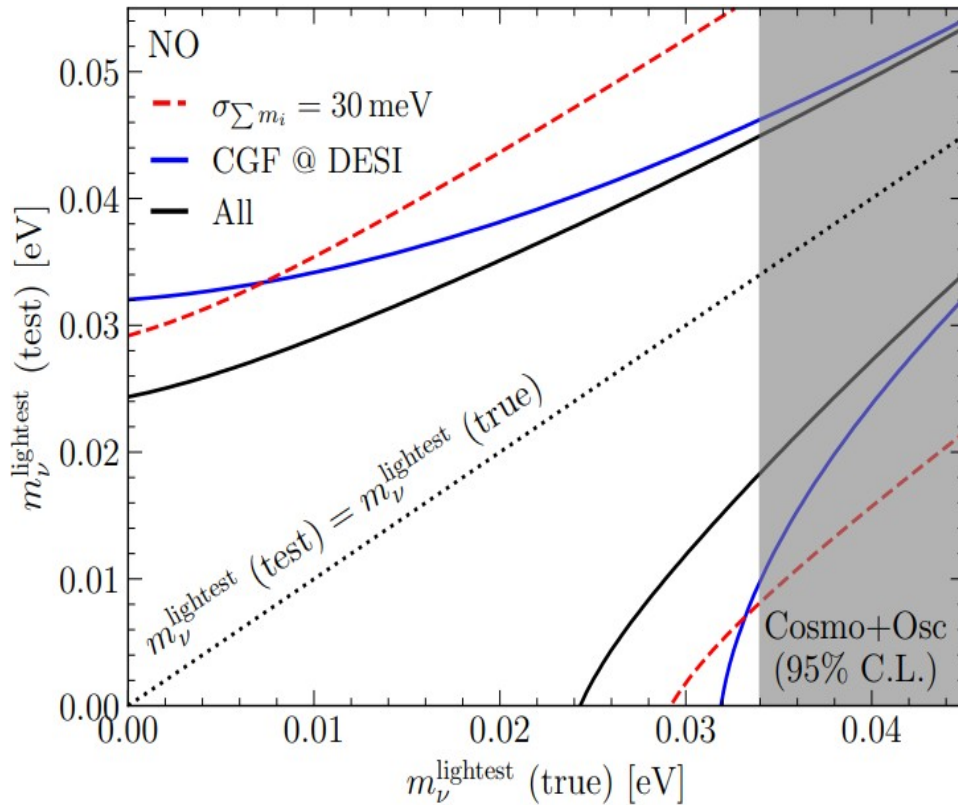
SFG, Pedro Pasquini, Liang Tan, **JCAP 05 (2024) 108** [arXiv:2312.16972]

$$\text{Im} \left[\tilde{\delta}_{g\alpha} \left(\tilde{\delta}_{g\beta} \right)^* \right] \propto \left(\mathbf{v}_{\nu c} \cdot \hat{\mathbf{k}} \right) \left(f_0 m_\nu^4 + f_1 m_\nu^2 T^2 + f_2 T^4 \right)$$



SFG, Pedro Pasquini, Liang Tan, **JCAP 05 (2024) 108** [arXiv:2312.16972]

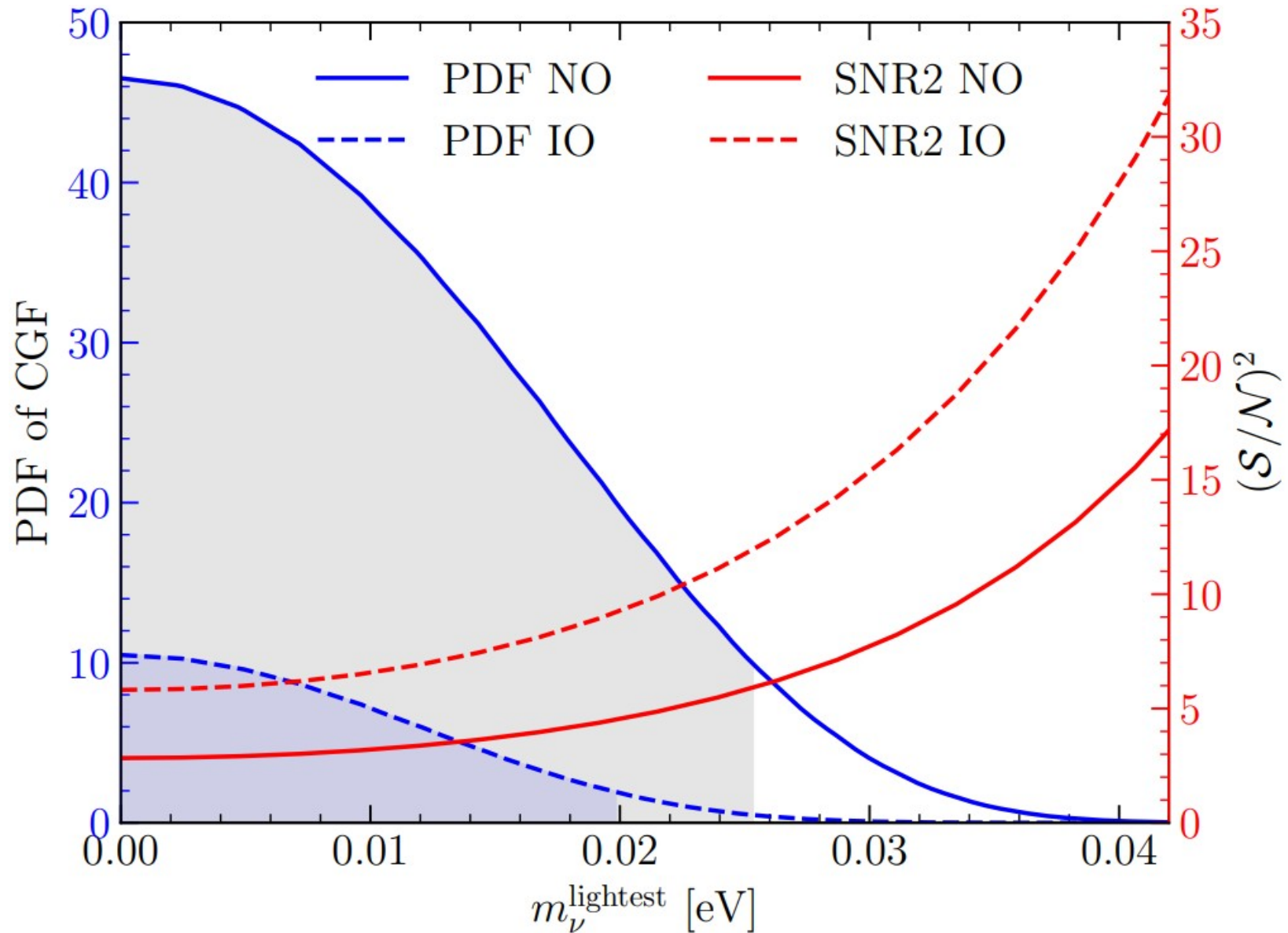
$$\text{Im} \left[\tilde{\delta}_{g\alpha} \left(\tilde{\delta}_{g\beta} \right)^* \right] \propto \left(\mathbf{v}_{\nu c} \cdot \hat{\mathbf{k}} \right) \left(f_0 m_\nu^4 + f_1 m_\nu^2 T^2 + f_2 T^4 \right)$$



DESI, Euclid, Subaru PFS, **CSST**

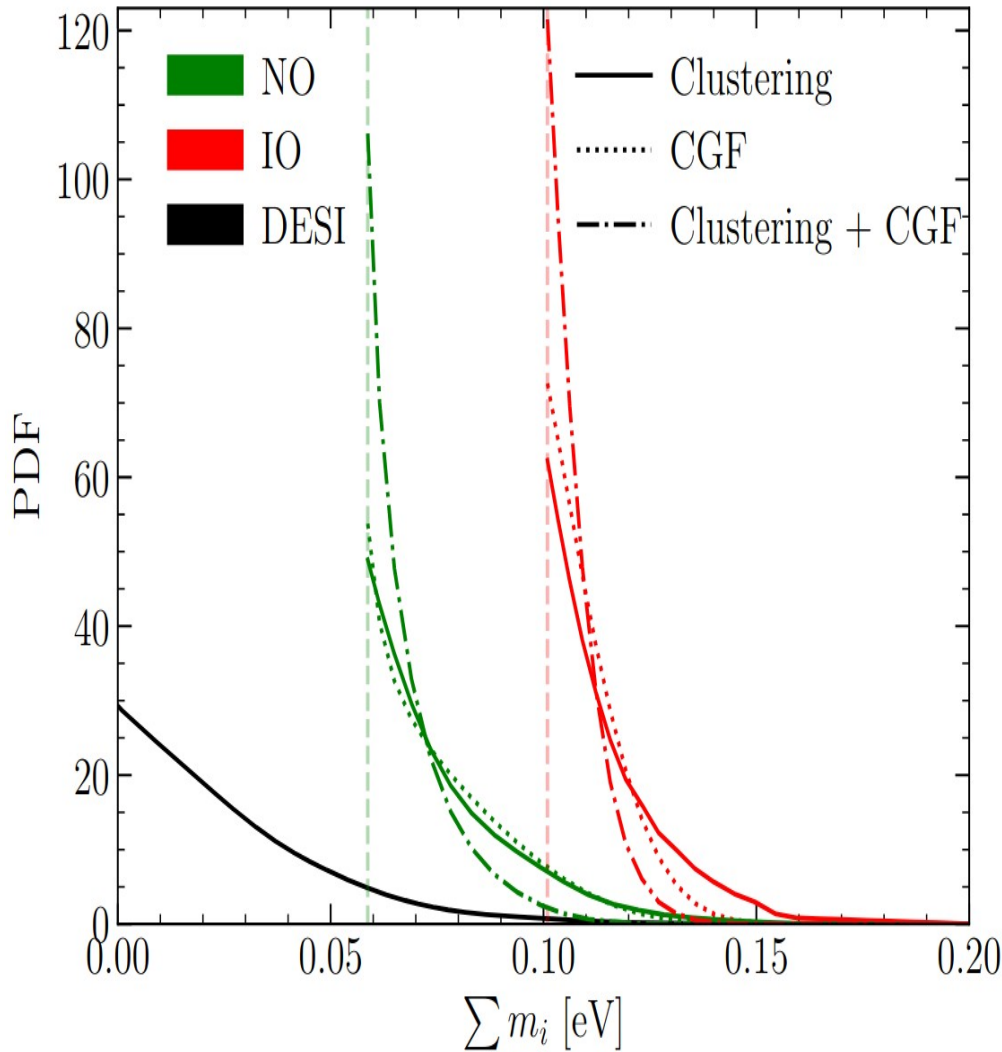
SFG, Pedro Pasquini, Liang Tan, **JCAP 05 (2024) 108** [arXiv:2312.16972]

Neutrino Mass Orderings

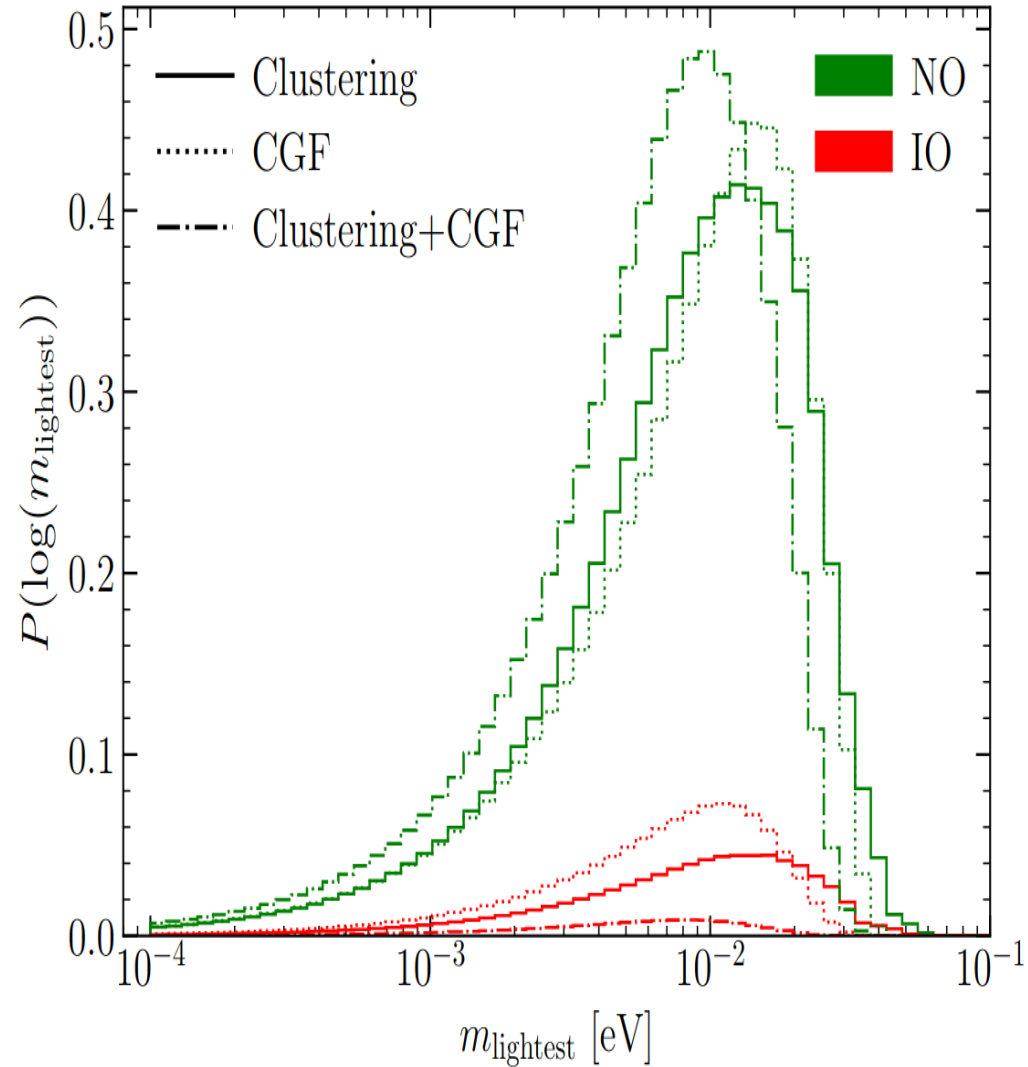


SFG & Liang Tan [arXiv:2409.11115]

Sensitivities on ν Mass Ordering



89.9% vs 10.1%



98.2% vs 1.8%

SFG & Liang Tan [arXiv:2409.11115]

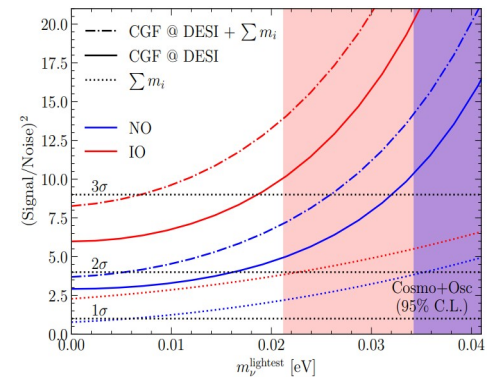
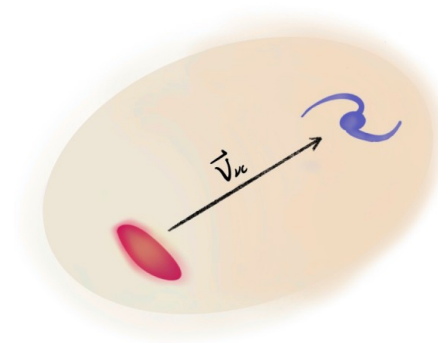
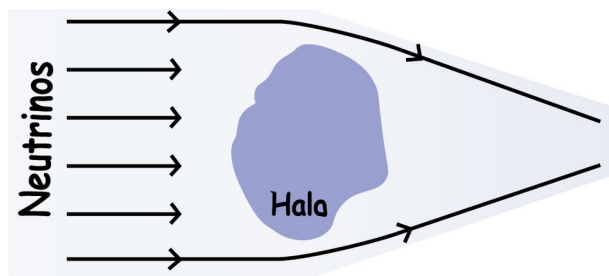
1) Overview of Neutrino Mass

2) 3rd Cosmological Way

Cosmic Gravitational Focusing

Dipole Structure in Galaxy Correlation Function

Neutrino Mass Measurement @ DESI



3) Mass Ordering

4) Summary

Thank You

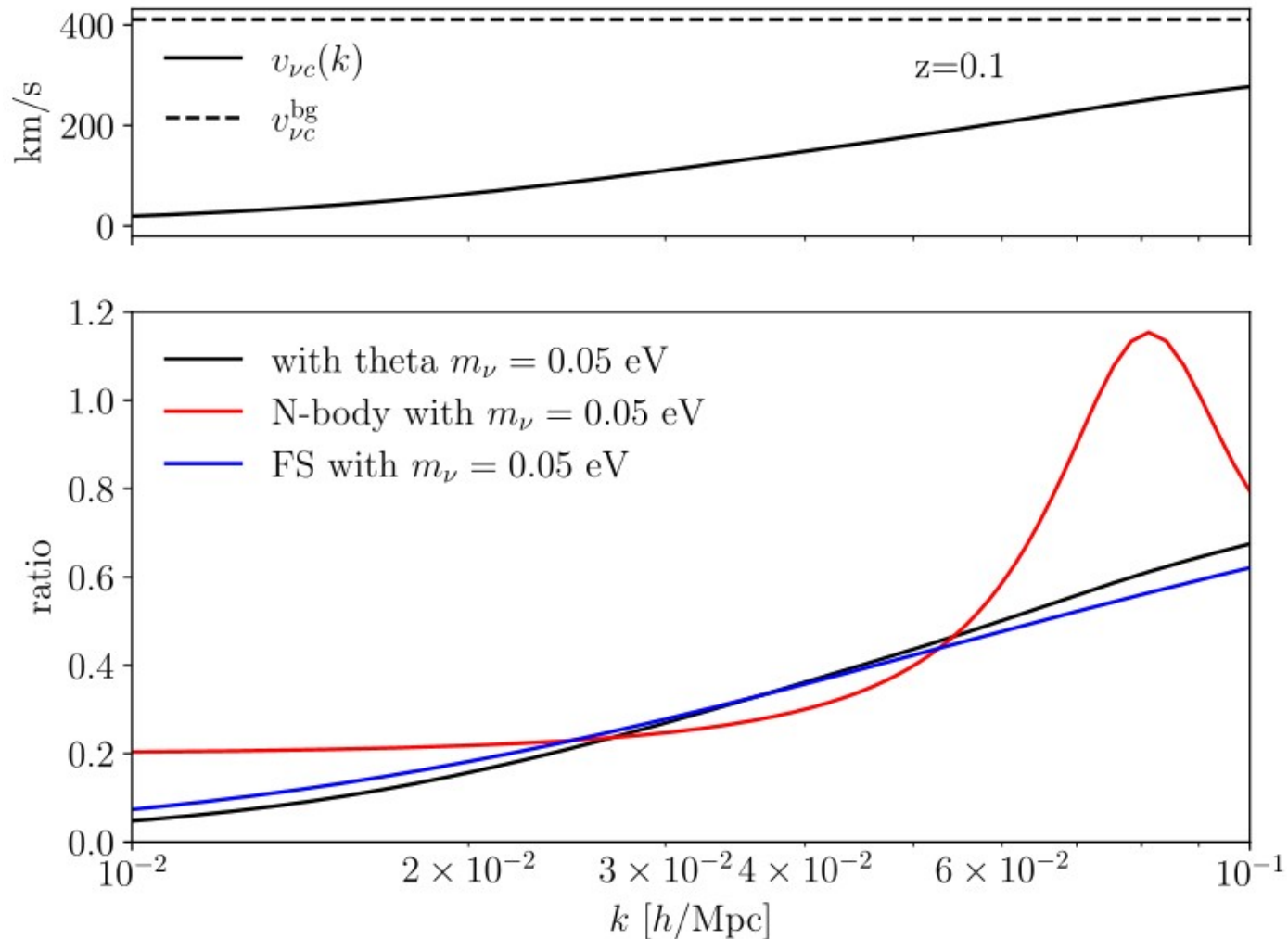
$$\text{Im}[\delta\rho_\nu] = \frac{\tilde{\Psi}}{4\pi} \sum_i (\mathbf{v}_{\nu ic} \cdot \hat{\mathbf{k}}) \left[m_i^4 \left(1 - \frac{m_i |\mathbf{v}_{\nu ic} \cdot \hat{\mathbf{k}}|}{2T_\nu} \right) + \pi^2 m_i^2 T_\nu^2 + \frac{14}{15} \pi^4 T_\nu^4 \right]$$

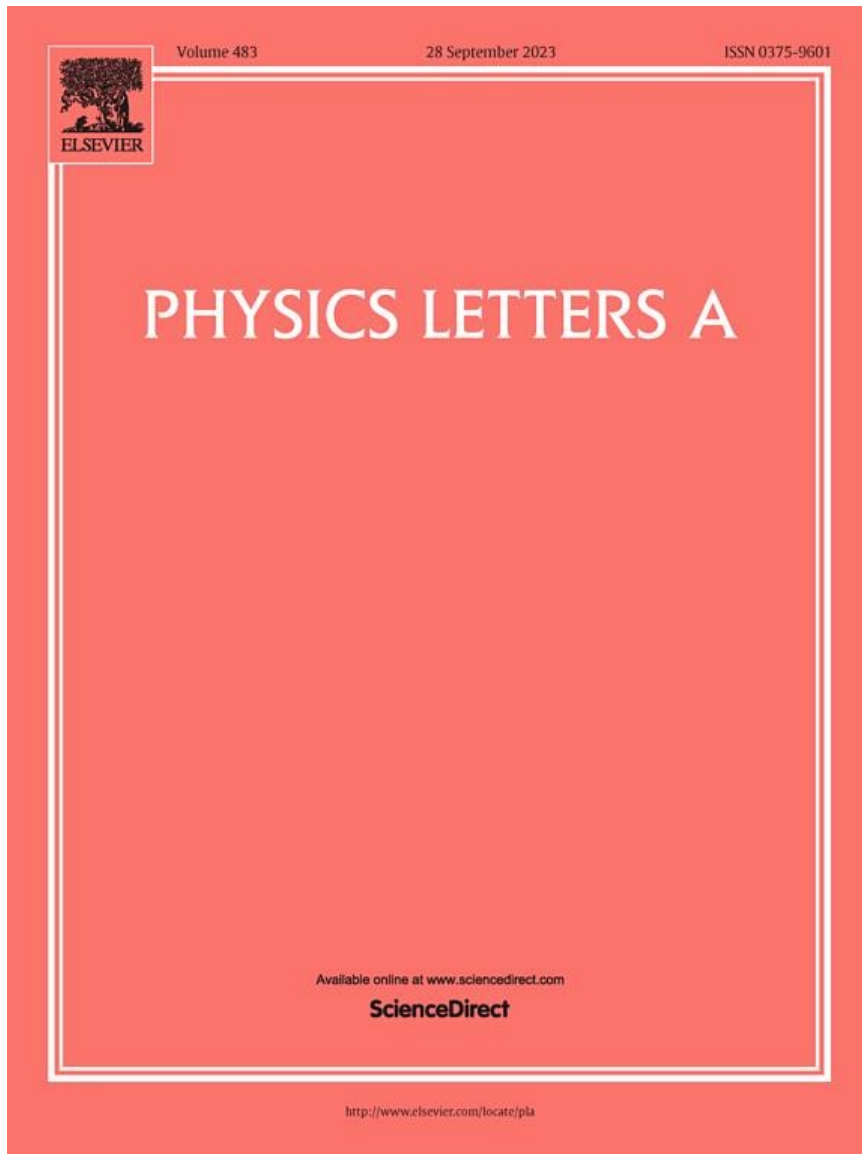
$$\langle (\mathbf{v}_{\nu ic} \cdot \hat{\mathbf{k}})^2 \rangle = \frac{1}{3} \int \frac{d|\mathbf{k}'|}{|\mathbf{k}'|} \Theta(|\mathbf{k}| - |\mathbf{k}'|) |\widetilde{W}(|\mathbf{k}'|R)|^2 \Delta_\zeta^2(\mathbf{k}') \left| \frac{T_{\theta_{\nu ic}}(\mathbf{k}', z)}{|\mathbf{k}'|} \right|^2,$$

$$\langle (\mathbf{v}_{\nu ic} \cdot \hat{\mathbf{k}}) [\partial_z (\mathbf{v}_{\nu ic} \cdot \hat{\mathbf{k}})] \rangle = \frac{1}{3} \int \frac{d|\mathbf{k}'|}{|\mathbf{k}'|} \Theta(|\mathbf{k}| - |\mathbf{k}'|) |\widetilde{W}(|\mathbf{k}'|R)|^2 \Delta_\zeta^2(\mathbf{k}') \left[\frac{T_{\theta_{\nu ic}}}{|\mathbf{k}'|} \frac{\partial_z T_{\theta_{\nu ic}}}{|\mathbf{k}'|} \right],$$

$$\langle [\partial_z (\mathbf{v}_{\nu ic} \cdot \hat{\mathbf{k}})]^2 \rangle = \frac{1}{3} \int \frac{d|\mathbf{k}'|}{|\mathbf{k}'|} \Theta(|\mathbf{k}| - |\mathbf{k}'|) |\widetilde{W}(|\mathbf{k}'|R)|^2 \Delta_\zeta^2(\mathbf{k}') \left[\frac{\partial_z T_{\theta_{\nu ic}}}{|\mathbf{k}'|} \right]^2,$$

$$\langle (\mathbf{v}_{\nu c} \cdot \hat{\mathbf{k}})^2 \rangle = \frac{1}{3} \int \frac{d|\mathbf{k}'|}{|\mathbf{k}'|} \Theta(|\mathbf{k}| - |\mathbf{k}'|) \left| \widetilde{W}(|\mathbf{k}'|R) \right|^2 \Delta_{\zeta}^2(\mathbf{k}') \left| \frac{T_{\theta_{\nu ic}}(\mathbf{k}', z)}{|\mathbf{k}'|} \right|^2$$

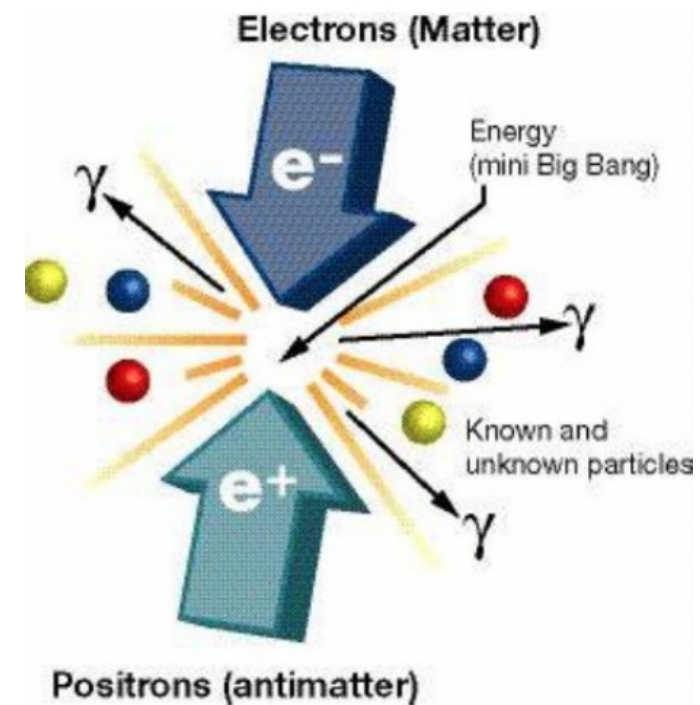
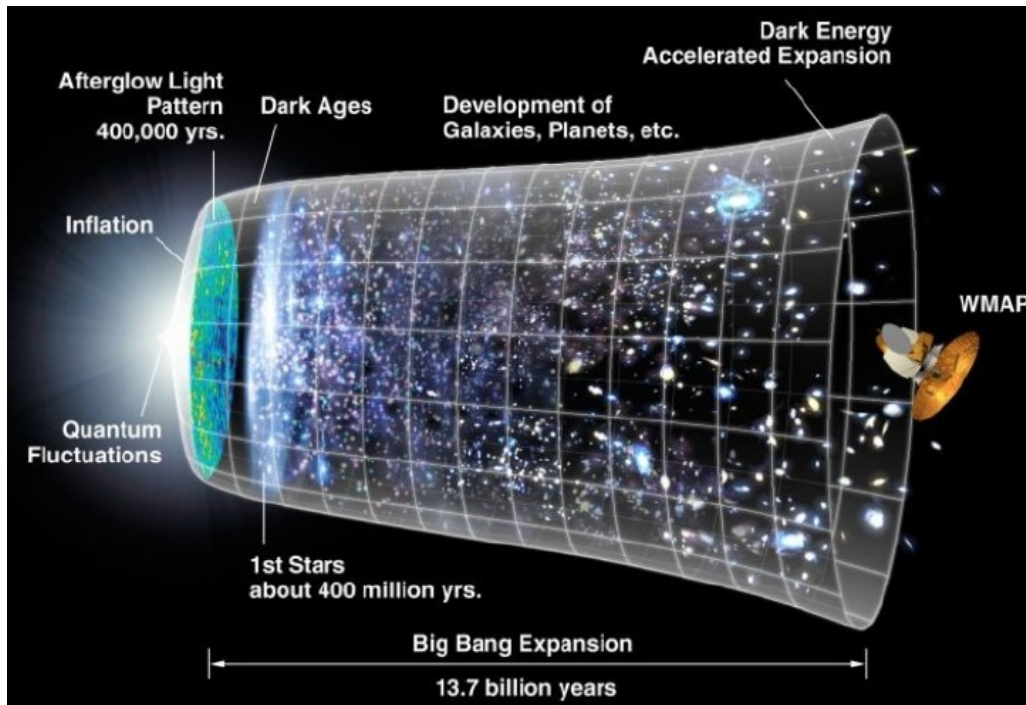




Aims & Scope

- Nonlinear science,
- Statistical physics,
- Mathematical and computational physics,
- AMO and physics of complex systems,
- Plasma and fluid physics,
- Optical physics,
- General and cross-disciplinary physics,
- Biological physics and nanoscience,
- Astrophysics, Particle physics and Cosmology.

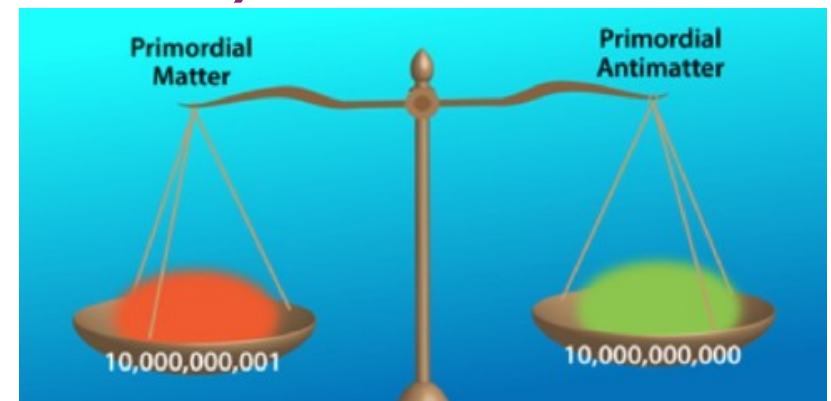
为什么存在物质世界?



- Primordial asymmetry diluted by **Inflation (暴涨理论)**

$$a(t) = e^{Ht}$$

- Starting from symmetric phase
- Why still some asymmetry?
- **Something is wrong!**



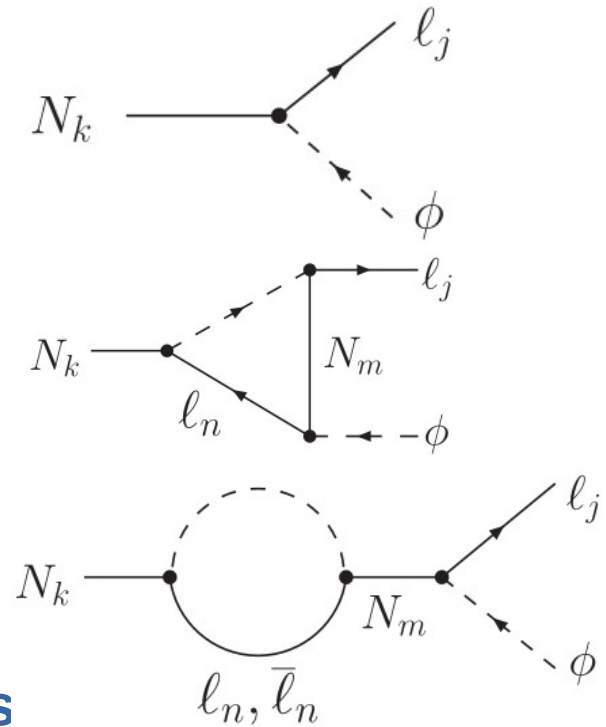
Leptogenesis (轻子创生机制)

宇宙早期有大量
右手中微子，随
着宇宙温度下
降，衰变到轻的
粒子

$$N_k \rightarrow \begin{cases} \ell_j + \bar{\phi} \\ \bar{\ell}_j + \phi \end{cases}$$



正反物质不对称



● Interference between tree & loop diagrams

$$\Gamma = \Gamma_{\text{tree}} + \Gamma_{\text{loop}}(+\delta_D, +\delta_M)$$

$$\bar{\Gamma} = \Gamma_{\text{tree}} + \Gamma_{\text{loop}}(-\delta_D, -\delta_M)$$

**The matter-antimatter asymmetry needs
Dirac/Majorana CP phases.**



Yanagida



Fukugita

- Three Neutrino Mixing

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \mathbf{U} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \begin{pmatrix} m_1 & & \\ & m_2 & \\ & & m_3 \end{pmatrix} = \mathbf{U}^\dagger \begin{pmatrix} A & B_1 & B_2 \\ & C_1 & D \\ & & C_2 \end{pmatrix} \mathbf{U}$$

- PMNS Matrix

$$\mathbf{U}_{\text{PMNS}} = \mathcal{P} \begin{pmatrix} c_s c_r & s_s c_r & s_r e^{-i\delta_D} \\ -s_s c_a - c_s s_a s_r e^{i\delta_D} & +c_s c_a - s_s s_a s_r e^{i\delta_D} & s_a c_r \\ +s_s s_a - c_s c_a s_r e^{i\delta_D} & -c_s s_a - s_s c_a s_r e^{i\delta_D} & c_a c_r \end{pmatrix} \mathcal{Q}$$

with $\mathcal{P} \equiv \text{diag}(e^{i\phi_1}, e^{i\phi_2}, e^{i\phi_3})$ & $\mathcal{Q} \equiv \text{diag}(e^{i\delta_{M1}}, 1, e^{i\delta_{M3}})$
[(s, a, r) \equiv (12, 23, 13) for (solar, atmospheric, reactor) angles]

如果中微子没有质量，就没有 CP 相角 δ_D !

Seesaw Mechanism (跷跷板机制)

- Heavy neutrinos (N)

$$\bar{\nu} M_D \mathcal{N} + h.c. + \bar{\mathcal{N}} M_N \mathcal{N} = \begin{pmatrix} \bar{\nu} & \bar{\mathcal{N}} \end{pmatrix} \begin{pmatrix} 0 & M_D \\ M_D^T & M_N \end{pmatrix} \begin{pmatrix} \nu \\ \mathcal{N} \end{pmatrix}$$

required by **Grand Unification Theory (大统一理论)**

- Seeaw Mechanism

The diagonalization of the full mass matrix

$$\begin{pmatrix} 0 & M_D \\ M_D^T & M_N \end{pmatrix} \Rightarrow M_\nu = -M_D \frac{1}{M_N} M_D^T$$

$$M_D \sim O(100) \text{ GeV}, \quad M_N \sim O(10^{15}) \text{ GeV} \quad \rightarrow \quad M_\nu \sim O(0.01) \text{ eV}$$

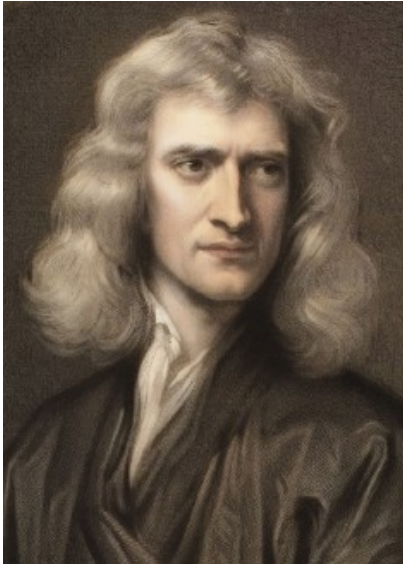
Light neutrino mass M_ν is suppressed by the heavy ones



柳田勉、何小刚：
小朋友，我们来玩
跷跷板吧

What is Mass? (质量是什么?)

- Mass is a long-standing issue. **Long history**



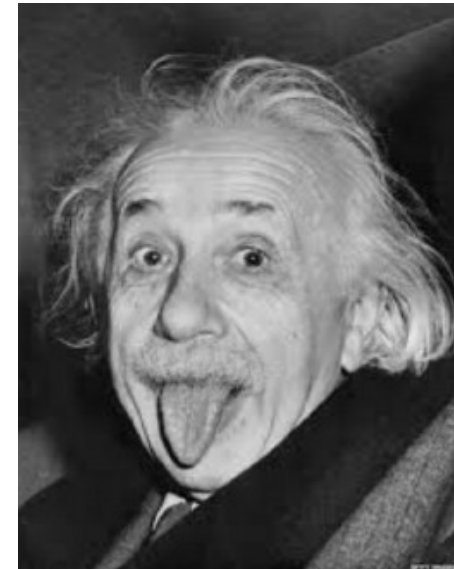
牛顿：如有雷同，纯属巧合

Inertial Mass

VS

Gravitational Mass

爱因斯坦：小样，你以为披个马甲我就不认识你了吗？



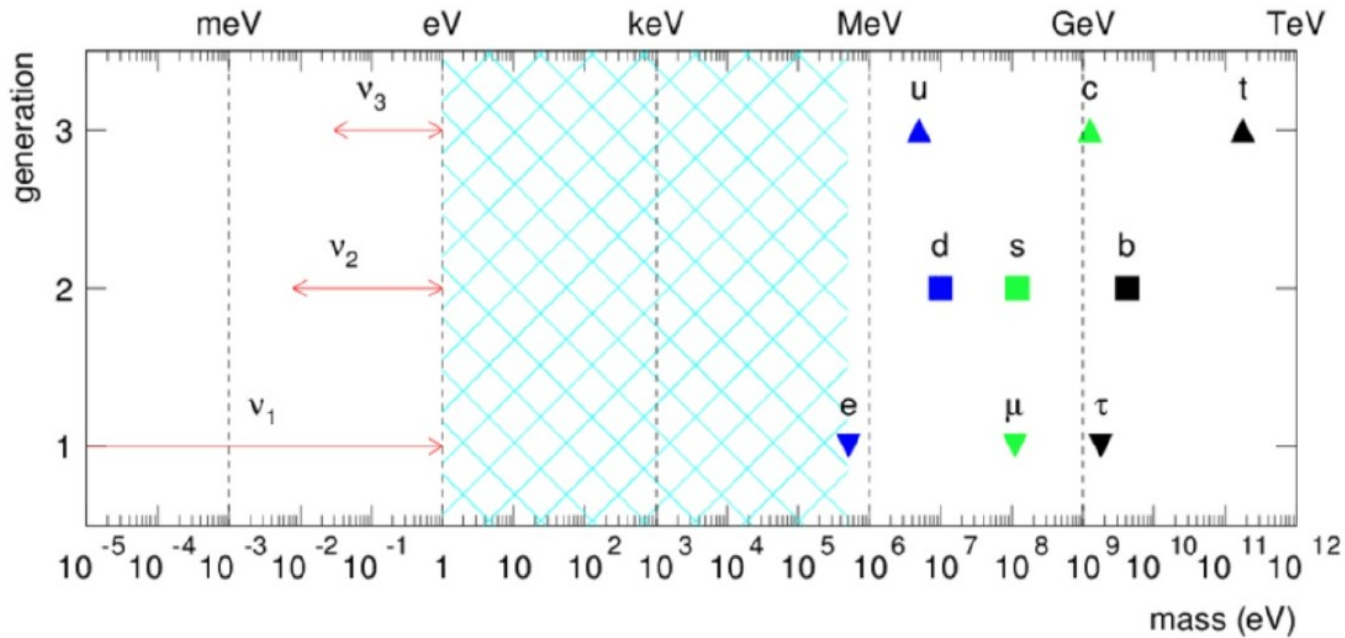
- Higgs Mechanism (希格斯机制)

$$y\bar{\psi}\phi\psi \quad \longrightarrow \quad \bar{\psi}M\psi = \bar{\psi}y\langle\phi\rangle\psi$$

希格斯：问题不在马甲，关键在于帽子！



Tiny Neutrino Mass



希格斯：大家都喜欢戴墨西哥草帽

$$m_t = 173 \text{ GeV} = 1.73 \times 10^{11} \text{ eV}$$

$$m_e = 0.51 \text{ MeV} = 5.1 \times 10^5 \text{ eV}$$

$$m_\nu \lesssim 10^{-2} \text{ eV}$$

How to explain the huge difference?

中微子：希格斯你这帽子太大了，我就一小孩，戴着不合适

ν Oscillation @ JUNO

SFG, Kaoru Hagiwara, Naotoshi Okamura, Yoshitaro Takaesu, **JHEP 05 (2013) 131** [arXiv:1210.8141]

Beta Decay

SFG, Manfred Lindner, **Phys.Rev.D 95 (2017) 3, 033003** [arXiv:1608.01618]

SFG & Werner Rodejohann, **Phys.Rev.D 96 (2017) 5, 055019** [arXiv:1707.07904]

SFG & Jing-Yu Zhu, **Chin.Phys.C 44 (2020) 8, 083103** [arXiv:1910.02666]

Radiative Emission of ν Pairs

SFG, Pedro Pasquini, **JHEP 12 (2023) 083** [arXiv:2306.12953]

Phys.Lett.B 841 (2023) 137911 [arXiv:2206.11717]

Eur.Phys.J.C 82 (2022) 3, 208 [arXiv:2110.03510]

Supernova ν Time Delay

SFG, Chui-Fan Kong, Alexei Smirnov, **PRL 133 (2024) 121802** [arXiv:2404.17352]

C ν B Detection

CMB & LSS

SFG, Pedro Pasquini, Liang Tan, **JCAP 05 (2024) 108** [arXiv:2312.16972]

SFG & Liang Tan [arXiv:2409.11115]

- Neutrino Kinetic & Mass Terms

$$\mathcal{L}_\nu = \bar{\nu}_\alpha (i\not{p}\delta_{\alpha\beta} + M_{\alpha\beta}) \nu_\beta$$

- Correlated Mass Matrix

$$\bar{\nu}_\alpha M_{\alpha\beta} \nu_\beta = -\frac{1}{2} \begin{pmatrix} \nu_e^T & \nu_\mu^T & \nu_\tau^T \end{pmatrix} \mathcal{C} \begin{pmatrix} A & B_1 & B_2 \\ & C_1 & D \\ & & C_2 \end{pmatrix} \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} + \dots$$

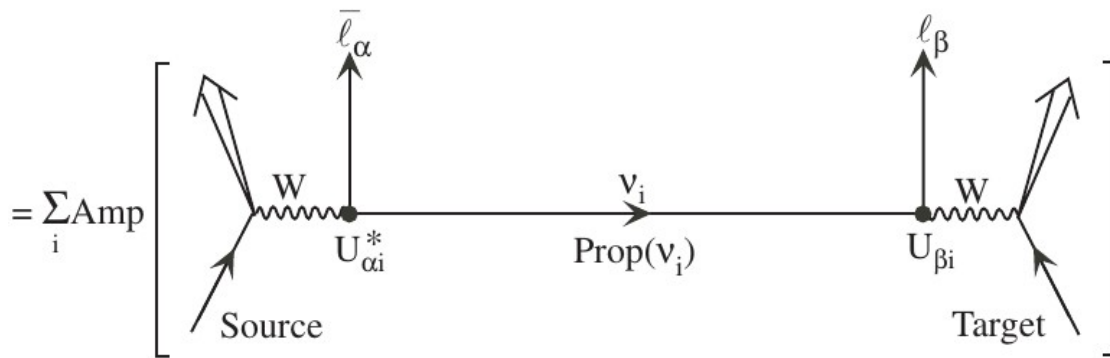
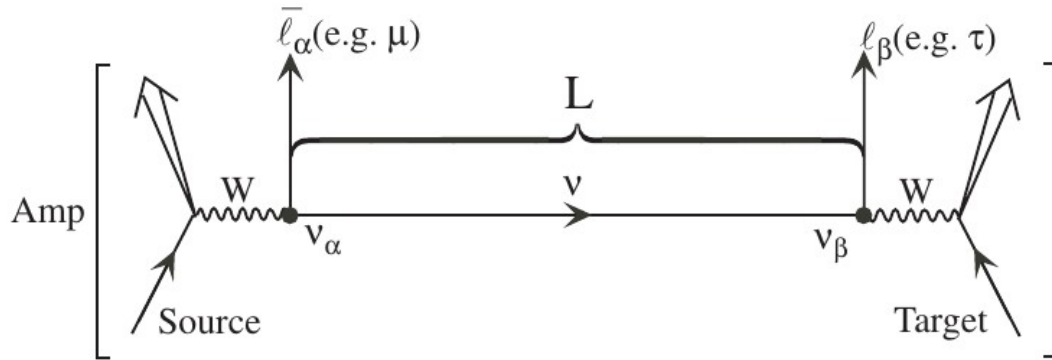
- Mass Diagonalization

$$\begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} = \mathbf{V}^\dagger \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} \quad \begin{pmatrix} m_1 & & \\ & m_2 & \\ & & m_3 \end{pmatrix} = \mathbf{V}^T \begin{pmatrix} A & B_1 & B_2 \\ & C_1 & D \\ & & C_2 \end{pmatrix} \mathbf{V}$$

Mass Eigenstate **Flavor Eigenstate**

Mass eigenstate is truly physical!

Neutrino Oscillation & Mass



$$\nu_\alpha = \sum_i U_{\alpha i} \nu_i$$

$$\rightarrow \sum_i U_{\alpha i} e^{i(E_i t - \vec{P}_i \cdot \vec{x})} \nu_i$$

$$= \sum_i U_{\alpha i} P_i U_{\beta i}^\dagger \nu_\beta$$

$$\equiv \sum_\beta A_{\alpha\beta} \nu_\beta$$

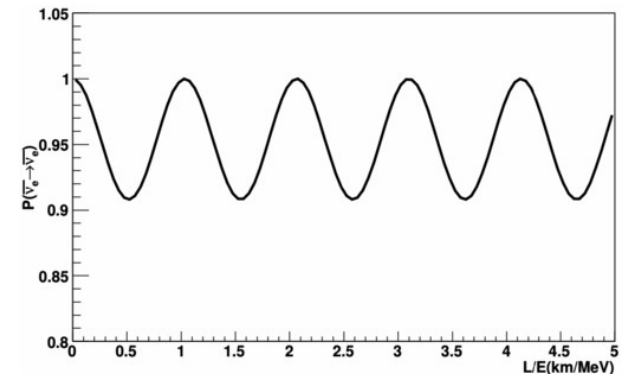
B. Kayser, [hep-ph/0506165]

$$P_{\alpha\beta} |_{\alpha \neq \beta} \equiv |A_{\alpha\beta}|^2 = \sin^2 2\theta \sin^2 \left(\delta m^2 \frac{L}{4E} \right)$$

↑
Mass

1st New Physics

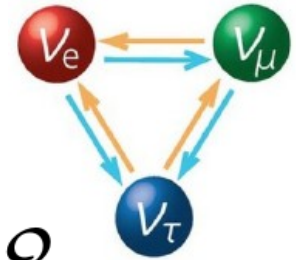
Quantum Interference over Macroscopic Distance!



Three Neutrino Oscillation

- PMNS Matrix

$$U_{\text{PMNS}} = \mathcal{P} \begin{pmatrix} c_s c_r & s_s c_r & s_r e^{-i\delta_D} \\ -s_s c_a - c_s s_a s_r e^{i\delta_D} & +c_s c_a - s_s s_a s_r e^{i\delta_D} & s_a c_r \\ +s_s s_a - c_s c_a s_r e^{i\delta_D} & -c_s s_a - s_s c_a s_r e^{i\delta_D} & c_a c_r \end{pmatrix} \mathcal{Q}$$

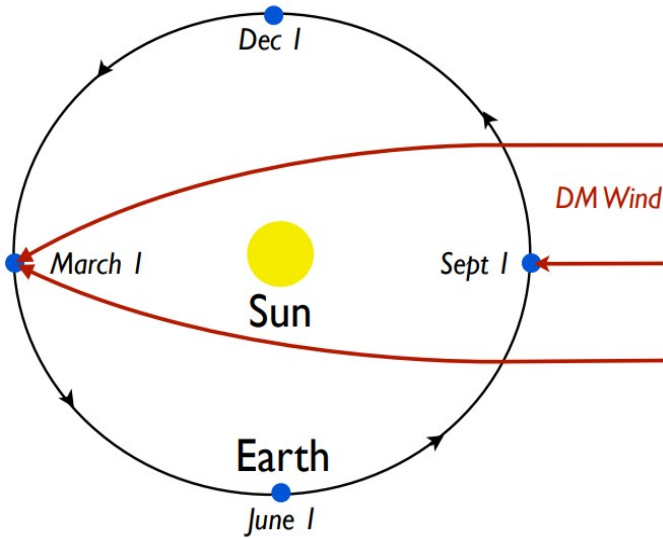


with $\mathcal{P} \equiv \text{diag}(e^{i\phi_1}, e^{i\phi_2}, e^{i\phi_3})$ & $\mathcal{Q} \equiv \text{diag}(e^{i\delta_{M1}}, 1, e^{i\delta_{M3}})$
 [(s, a, r) \equiv (12, 23, 13) for (solar, atmospheric, reactor) angles]

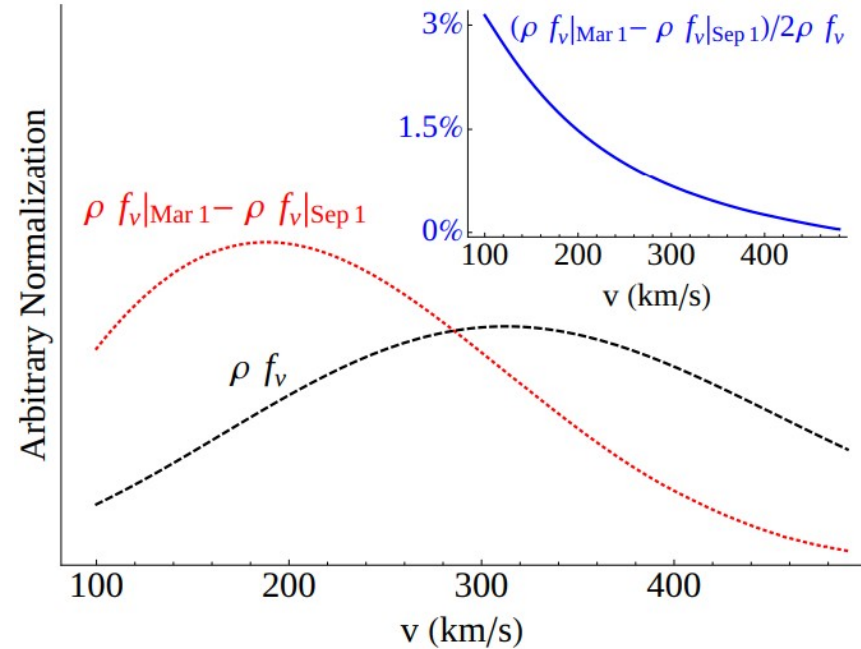
$$\begin{aligned} P(\nu_\alpha \rightarrow \nu_\beta) &= |\text{Amp}(\nu_\alpha \rightarrow \nu_\beta)|^2 \\ &= \delta_{\alpha\beta} - 4 \sum_{i>j} \Re(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2(\Delta m_{ij}^2 \frac{L}{4E}) \\ &\quad + 2 \sum_{i>j} \Im(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin(\Delta m_{ij}^2 \frac{L}{2E}) \end{aligned}$$

$$\Delta m_{ij}^2 \equiv m_i^2 - m_j^2 \qquad \Delta m_{ij}^2 \frac{L}{4E} = 1.27 \Delta m_{ij}^2 (\text{eV}^2) \frac{L (\text{km})}{E (\text{GeV})}$$

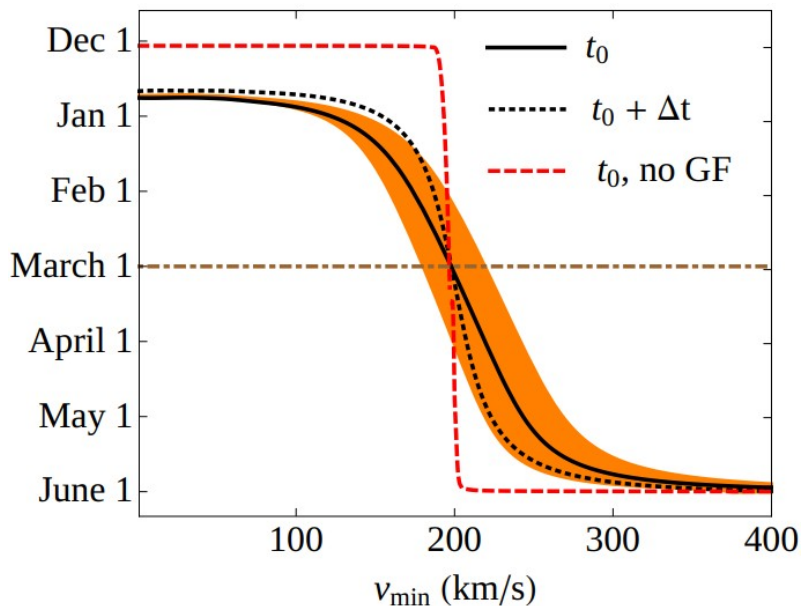
DM Gravitational Focusing



- Density enhancement



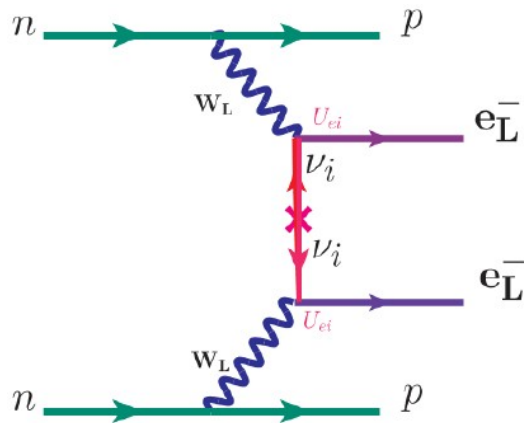
- Modulation phase shift



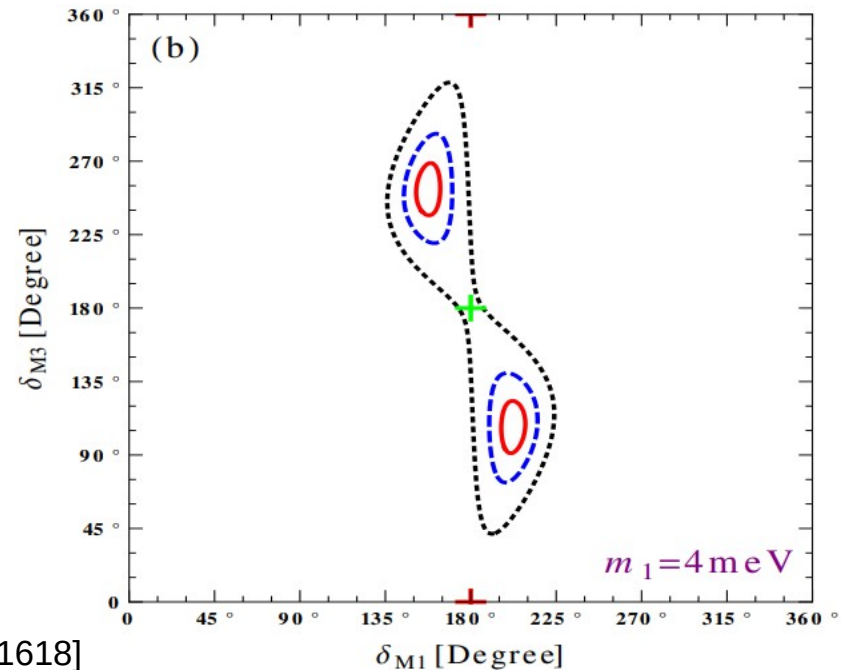
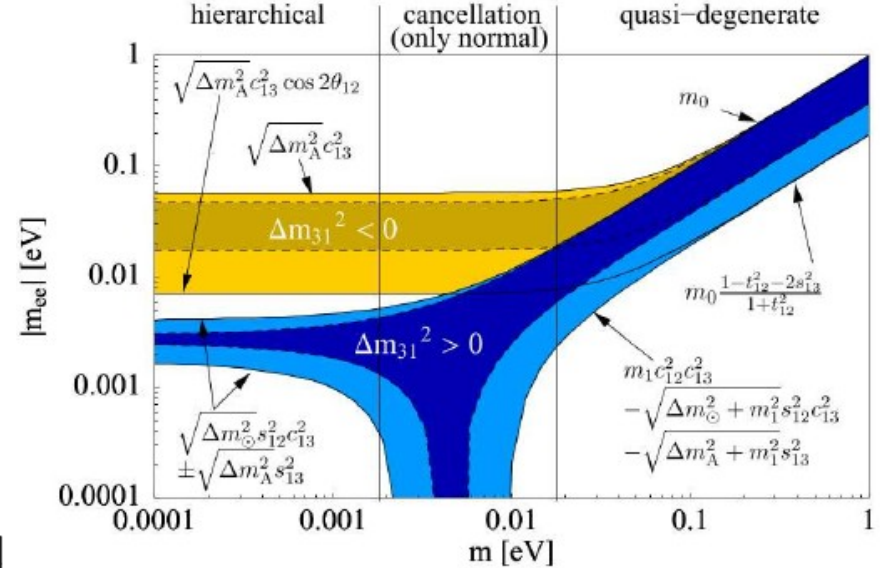
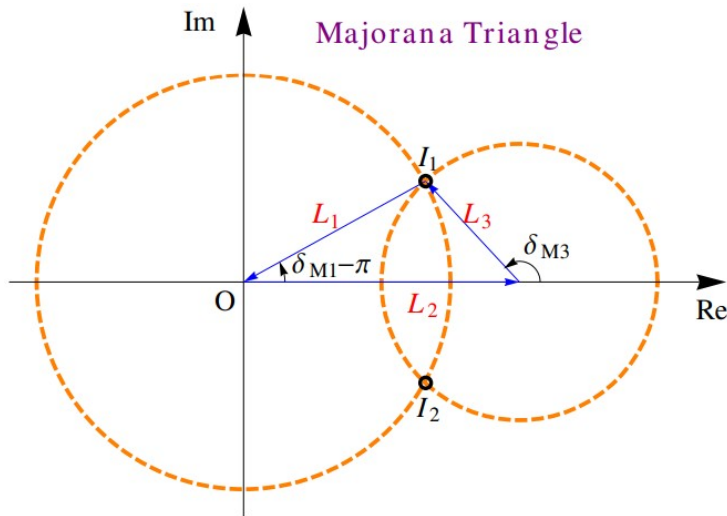
Lee, Lisanti, Peter, Safdi, Phys. Rev. Lett. 112, 011301 (2014) [arXiv:1308.1953]

Bozorgnia & Schwetz, JCAP 08 (2014) 013 [arXiv:1405.2340]

Neutrinoless Double Beta Decay

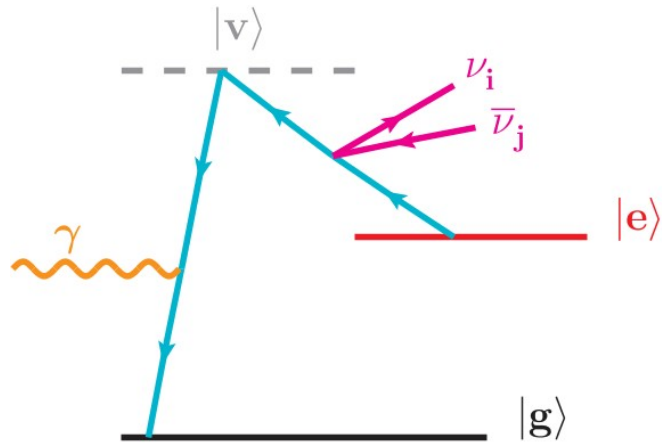


$$\langle m \rangle_{ee} \equiv \left| \sum_i m_i U_{ei}^2 \right| = \left| c_s^2 c_r^2 m_1 e^{i\delta_{M1}} + s_s^2 c_r^2 m_2 + s_r^2 m_3 e^{i\delta_{M3}} \right|$$



SFG & Manfred Lindner, PRD 95 (2017) No.3, 033003 [arXiv:1608.01618]

Radiative Emission of ν Pairs

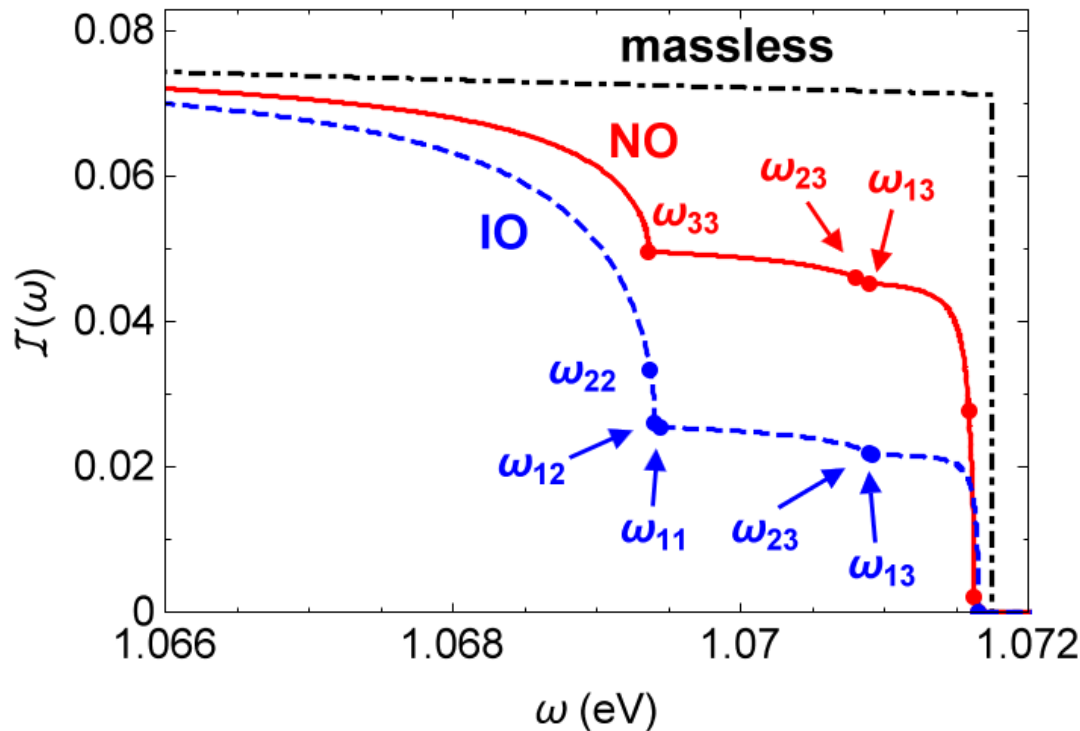


$$|e\rangle \rightarrow |v\rangle + \nu\bar{\nu}$$

$$E1 \times M1$$

M. Yoshimura, Phys. Rev. D 75, 113007 (2007)

Jue Zhang & Shun Zhou, Phys.Rev.D 93 (2016) 11, 113020

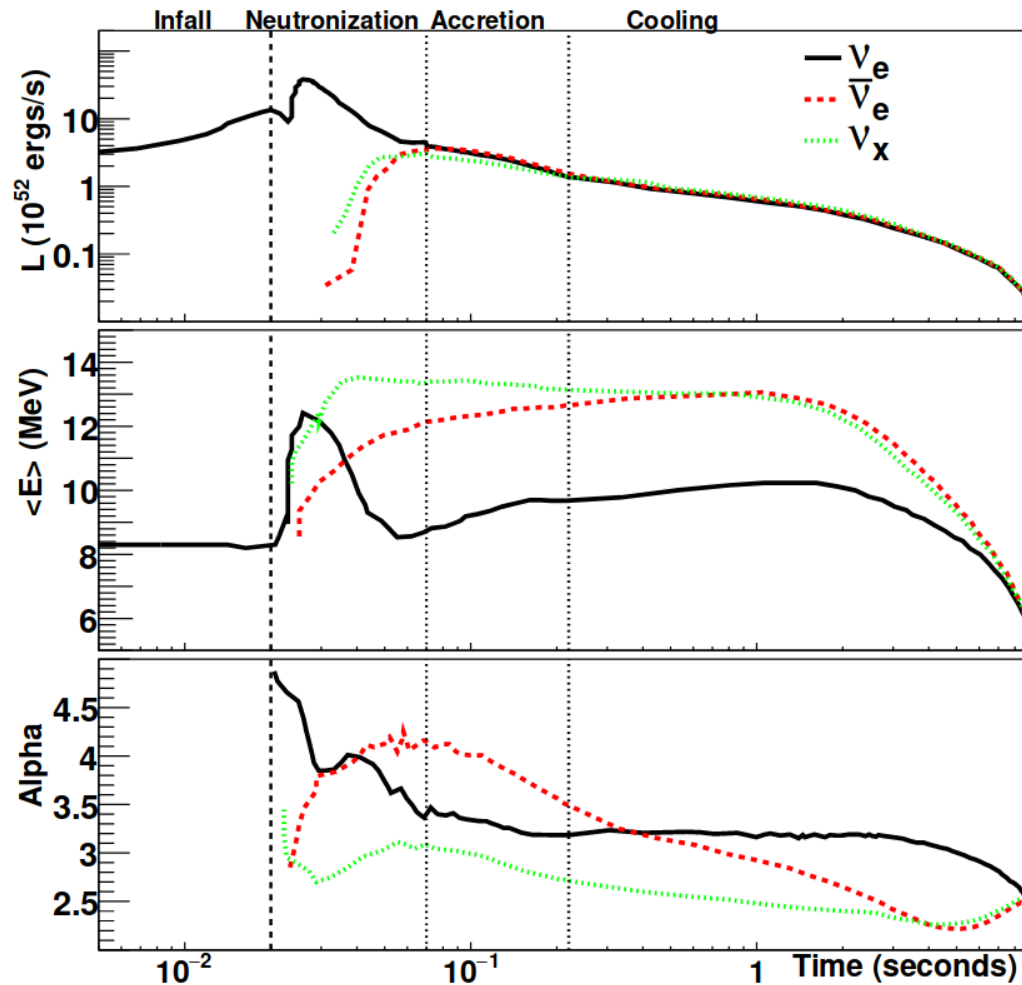


$$\omega_{ij}^{\max} \equiv \frac{E_e - E_g}{2} - \frac{1}{2} \frac{(m_i + m_j)^2}{(E_e - E_g)}$$

SFG & Pasquini

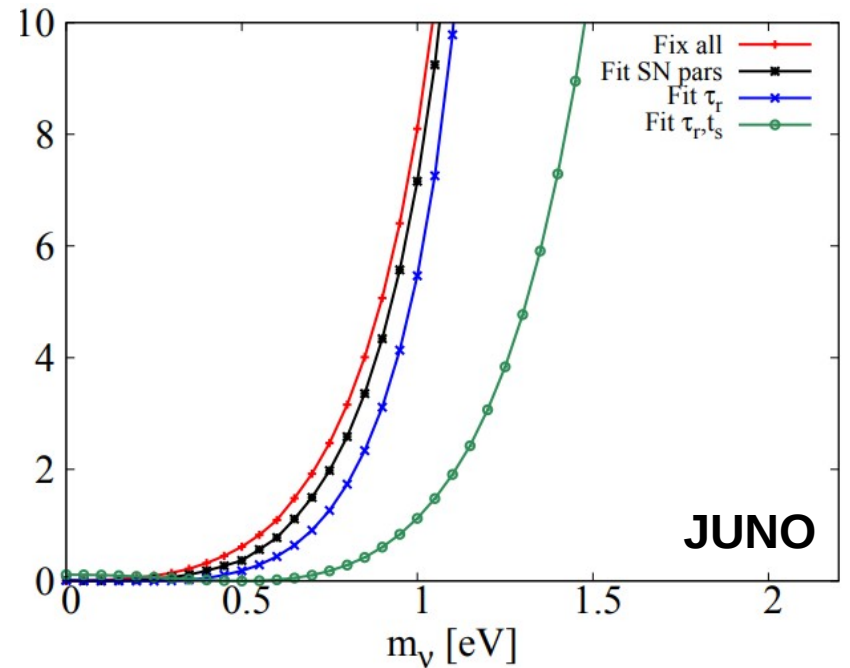
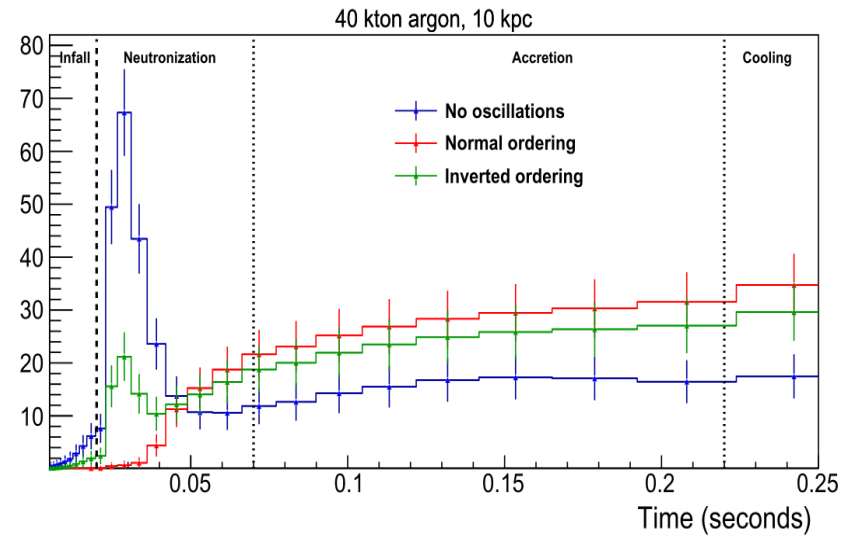
Eur.Phys.J.C 82 (2022) 3, 208;
Phys.Lett.B 841 (2023) 137911;
JHEP 12 (2023) 083

Time Delay in Supernova ν 's



DUNE, 2008.06647

SFG, Chui-Fan Kong, Alexei Smirnov,
PRL 133 (2024) 121802 [arXiv:2404.17352]



Lu, Cao, Li & Zhou [JCAP 05 (2015) 044]