



Impacts of ALP on dark photon searches

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Where is the Physics Beyond the SM?

ATLAS Heavy Particle Searches* 95% CL Upper Exclusion Limits

Status: March 2023

$$\int \mathcal{L} dt = (3.6 - 139) \text{ fb}^{-1}$$

ATLAS Preliminary

$$\sqrt{s} = 13 \text{ TeV}$$

Model	ℓ, γ	Jets [†]	$E_{\text{T}}^{\text{miss}}$	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Limit	Reference		
Extra dimen.	ADD $G_{KK} + g/q$	$0 e, \mu, \tau, \gamma$	$1-4 j$	Yes	139	M_D 11.2 TeV $n=2$	2102.10874	
	ADD non-resonant $\gamma\gamma$	2γ	-	-	36.7	M_S 8.6 TeV $n=3$ HLZ NLO	1707.04147	
	ADD QBH	-	$2 j$	-	139	M_{th} 9.4 TeV $n=6$	1910.08447	
	ADD BH multijet	-	$\geq 3 j$	-	3.6	M_{th} 9.55 TeV $n=6, M_D=3 \text{ TeV, rot BH}$	1512.02586	
	RS1 $G_{KK} \rightarrow \gamma\gamma$	2γ	-	-	139	G_{KK} mass 4.5 TeV $k/\overline{M}_{Pl}=0.1$	2102.13405	
	Bulk RS $G_{KK} \rightarrow WW/ZZ$	multi-channel	-	-	36.1	G_{KK} mass 2.3 TeV $k/\overline{M}_{Pl}=1.0$	1808.02380	
	Bulk RS $g_{KK} \rightarrow tt$	$1 e, \mu$	$\geq 1 b, \geq 1 J/2 j$	Yes	36.1	g_{KK} mass 3.8 TeV $\Gamma/m=15\%$	1804.10823	
2UED / RPP	$1 e, \mu$	$\geq 2 b, \geq 3 j$	Yes	36.1	KK mass 1.8 TeV Tier (1,1), $\mathcal{B}(A^{(1,1)} \rightarrow tt) = 1$	1803.09678		
Gauge bosons	SSM $Z' \rightarrow \ell\ell$	$2 e, \mu$	-	-	139	Z' mass 5.1 TeV	1903.06248	
	SSM $Z' \rightarrow \tau\tau$	2τ	-	-	36.1	Z' mass 2.42 TeV	1709.07242	
	Leptophobic $Z' \rightarrow bb$	-	$2 b$	-	36.1	Z' mass 2.1 TeV	1805.09299	
	Leptophobic $Z' \rightarrow tt$	$0 e, \mu$	$\geq 1 b, \geq 2 J$	Yes	139	Z' mass 4.1 TeV $\Gamma/m=1.2\%$	2005.05138	
	SSM $W' \rightarrow \ell\nu$	$1 e, \mu$	-	Yes	139	W' mass 6.0 TeV	1906.05609	
	SSM $W' \rightarrow \tau\nu$	1τ	-	Yes	139	W' mass 5.0 TeV	ATLAS-CONF-2021-025	
	SSM $W' \rightarrow tb$	-	$\geq 1 b, \geq 1 J$	-	139	W' mass 4.4 TeV	ATLAS-CONF-2021-043	
	HVT $W' \rightarrow WZ$ model B	$0-2 e, \mu$	$2 j / 1 J$	Yes	139	W' mass 4.3 TeV	2004.14636	
	HVT $W' \rightarrow WZ \rightarrow \ell\nu \ell' \ell'$ model C	$3 e, \mu$	$2 j$ (VBF)	Yes	139	W' mass 340 GeV	2207.03925	
	HVT $Z' \rightarrow WW$ model B	$1 e, \mu$	$2 j / 1 J$	Yes	139	Z' mass 3.9 TeV	2004.14636	
LRSM $W_R \rightarrow \mu N_R$	2μ	$1 J$	-	80	W_R mass 5.0 TeV $m(N_R)=0.5 \text{ TeV, } g_L = g_R$	1904.12679		
CI	CI $qqqq$	-	$2 j$	-	37.0	Λ 21.8 TeV η_{LL}	1703.09127	
	CI $\ell\ell qq$	$2 e, \mu$	-	-	139	Λ 35.8 TeV η_{LL}	2006.12946	
	CI $e e b s$	$2 e$	$1 b$	-	139	Λ 1.8 TeV $g_s=1$	2105.13847	
	CI $\mu\mu b s$	2μ	$1 b$	-	139	Λ 2.0 TeV $g_s=1$	2105.13847	
	CI $tttt$	$\geq 1 e, \mu$	$\geq 1 b, \geq 1 j$	Yes	36.1	Λ 2.57 TeV $ C_A =4\pi$	1811.02305	
	DM	Axial-vector med. (Dirac DM)	-	$2 j$	-	139	m_{med} 3.8 TeV	ATL-PHYS-PUB-2022-036
Pseudo-scalar med. (Dirac DM)		$0 e, \mu, \tau, \gamma$	$1-4 j$	Yes	139	m_{med} 376 GeV $g_q=0.25, g_\ell=1, m(\chi)=10 \text{ GeV}$	2102.10874	
Vector med. Z' -2HDM (Dirac DM)		$0 e, \mu$	$2 b$	Yes	139	$m_{Z'}$ 3.0 TeV $\tan\beta=1, g_\tau=1, m(\chi)=100 \text{ GeV}$	2108.13391	
Pseudo-scalar med. 2HDM+a		multi-channel	-	-	139	m_a 800 GeV $\tan\beta=1, g_\tau=1, m(\chi)=10 \text{ GeV}$	ATLAS-CONF-2021-036	
LQ	Scalar LQ 1 st gen	$2 e$	$\geq 2 j$	Yes	139	LQ mass 1.8 TeV $\beta=1$	2006.05872	
	Scalar LQ 2 nd gen	2μ	$\geq 2 j$	Yes	139	LQ mass 1.7 TeV $\beta=1$	2006.05872	
	Scalar LQ 3 rd gen	1τ	$2 b$	Yes	139	LQ_3^u mass 1.49 TeV $\mathcal{B}(LQ_3^u \rightarrow b\tau) = 1$	2303.01294	
	Scalar LQ 3 rd gen	$0 e, \mu$	$\geq 2 j, \geq 2 b$	Yes	139	LQ_3^d mass 1.24 TeV $\mathcal{B}(LQ_3^d \rightarrow \tau\nu) = 1$	2004.14060	
	Scalar LQ 3 rd gen	$\geq 2 e, \mu, \geq 1 \tau, \geq 1 j, \geq 1 b$	-	-	139	LQ_3^d mass 1.43 TeV $\mathcal{B}(LQ_3^d \rightarrow t\tau) = 1$	2101.11582	
	Scalar LQ 3 rd gen	$0 e, \mu, \geq 1 \tau$	$0-2 j, 2 b$	Yes	139	LQ_3^d mass 1.26 TeV $\mathcal{B}(LQ_3^d \rightarrow b\nu) = 1$	2101.12527	
	Vector LQ mix gen	multi-channel	$\geq 1 j, \geq 1 b$	Yes	139	LQ_3^v mass 2.0 TeV $\mathcal{B}(U_L \rightarrow t\mu) = 1, \text{Y-M coupl.}$	ATLAS-CONF-2022-052	
	Vector LQ 3 rd gen	$2 e, \mu, \tau$	$\geq 1 b$	Yes	139	LQ_3^v mass 1.96 TeV $\mathcal{B}(LQ_3^v \rightarrow b\tau) = 1, \text{Y-M coupl.}$	2303.01294	
Vector-like fermions	VLQ $TT \rightarrow Zt + X$	$2e/2\mu \geq 3e, \mu$	$\geq 1 b, \geq 1 j$	-	139	T mass 1.46 TeV	SU(2) doublet	2210.15413
	VLQ $BB \rightarrow Wt/Zb + X$	multi-channel	-	-	36.1	B mass 1.34 TeV	SU(2) doublet	1808.02343
	VLQ $T_{5/3} T_{5/3} T_{5/3} \rightarrow Wt + X$	$2(SS) \geq 3 e, \mu$	$\geq 1 b, \geq 1 j$	Yes	36.1	$T_{5/3}$ mass 1.64 TeV	$\mathcal{B}(T_{5/3} \rightarrow Wt) = 1, c(T_{5/3} Wt) = 1$	1807.11883
	VLQ $T \rightarrow Ht/Zt$	$1 e, \mu$	$\geq 1 b, \geq 3 j$	Yes	139	T mass 1.8 TeV	SU(2) singlet, $\kappa_T = 0.5$	ATLAS-CONF-2021-040
	VLQ $Y \rightarrow Wb$	$1 e, \mu$	$\geq 1 b, \geq 1 j$	Yes	36.1	Y mass 1.85 TeV	$\mathcal{B}(Y \rightarrow Wb) = 1, c_R(Wb) = 1$	1812.07343
	VLQ $B \rightarrow Hb$	$0 e, \mu$	$\geq 2b, \geq 1 j, \geq 1 j$	-	139	B mass 2.0 TeV	SU(2) doublet, $\kappa_B = 0.3$	ATLAS-CONF-2021-018
	VLL $\tau' \rightarrow Z\tau/H\tau$	multi-channel	$\geq 1 j$	Yes	139	τ' mass 898 GeV	SU(2) doublet	2303.05441
Exctd ferm.	Excited quark $q^* \rightarrow qg$	-	$2 j$	-	139	q^* mass 6.7 TeV	only u^* and d^* , $\Lambda = m(q^*)$	1910.08447
	Excited quark $q^* \rightarrow q\gamma$	1γ	$1 j$	-	36.7	q^* mass 5.3 TeV	only u^* and d^* , $\Lambda = m(q^*)$	1709.10440
	Excited quark $b^* \rightarrow bg$	-	$1 b, 1 j$	-	139	b^* mass 3.2 TeV	$\Lambda = 4.6 \text{ TeV}$	1910.08447
	Excited lepton τ^*	2τ	$\geq 2 j$	-	139	τ^* mass 4.6 TeV	$\Lambda = 4.6 \text{ TeV}$	2303.09444
Other	Type III Seesaw	$2,3,4 e, \mu$	$\geq 2 j$	Yes	139	N^0 mass 910 GeV	$m(W_R) = 4.1 \text{ TeV, } g_L = g_R$	2202.02039
	LRSM Majorana ν	2μ	$2 j$	-	36.1	N_R mass 3.2 TeV	1809.11105	
	Higgs triplet $H^{\pm\pm} \rightarrow W^{\pm} W^{\pm}$	$2,3,4 e, \mu$ (SS)	various	Yes	139	$H^{\pm\pm}$ mass 350 GeV	DY production	2101.11961
	Higgs triplet $H^{\pm\pm} \rightarrow \ell\ell$	$2,3,4 e, \mu$ (SS)	-	-	139	$H^{\pm\pm}$ mass 1.08 TeV	DY production	2211.07505
	Multi-charged particles	Multi-charged particles	-	-	139	multi-charged particle mass 1.50 TeV	DY production, $ \lambda = 5$	ATLAS-CONF-2022-034

new particles are too heavy?
too light? OR/AND couplings are too small?

*Only at
†Small



Tiny interactions?

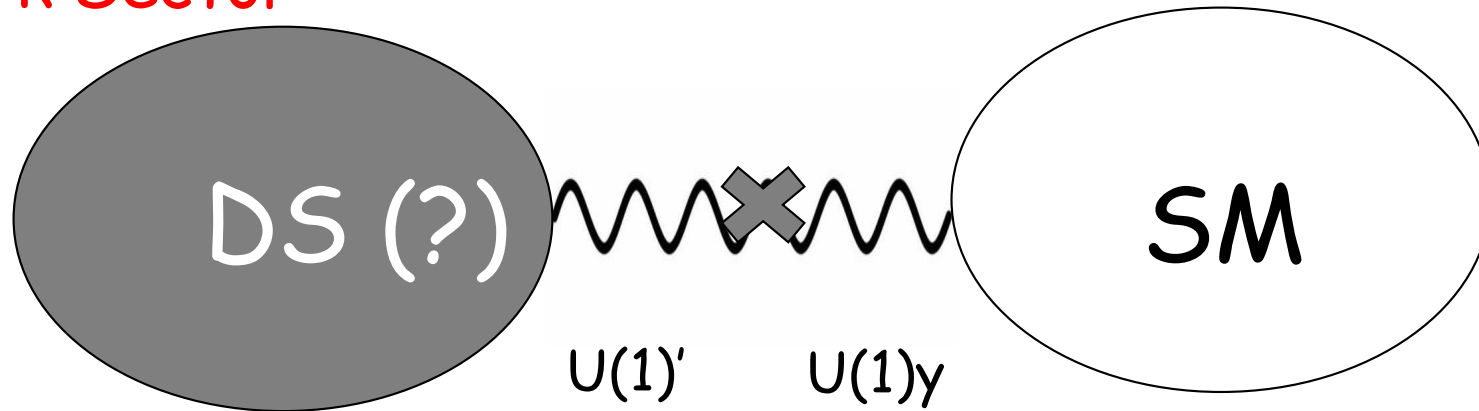
portals ?

Light particles?

dark photon ?

axion (axion-like) ?

Dark Sector



Fabbrichesi, Gabrielli, Lanfranchi, 2005.01515

messenger: vector-boson



"dark photon portal"

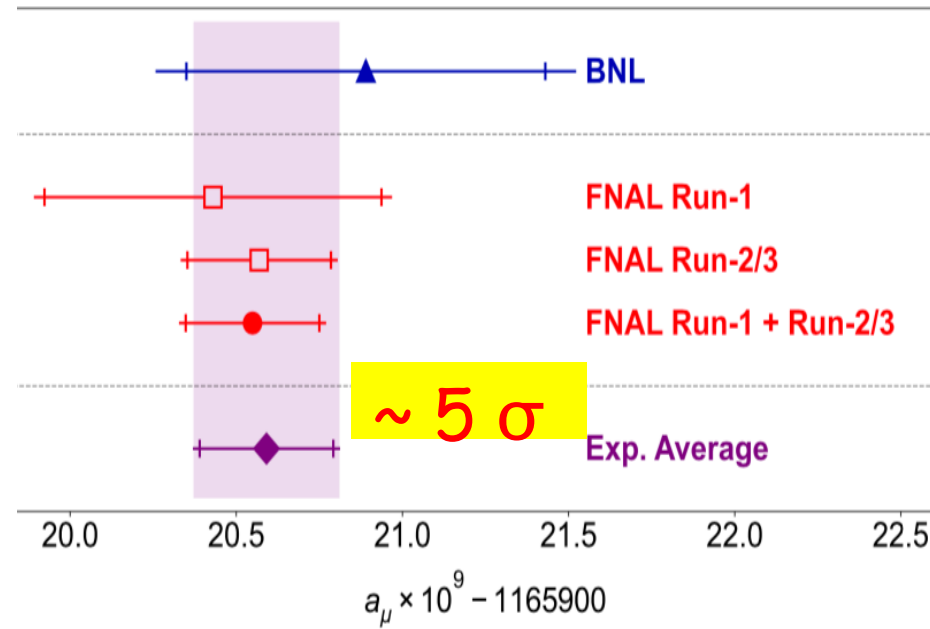
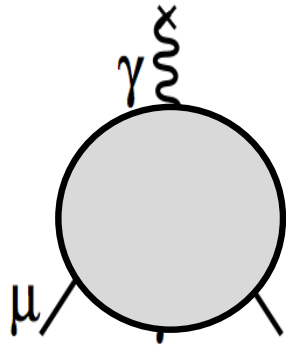
$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{DS} + \frac{\epsilon}{2c_W} B'_{\mu\nu} B^{\mu\nu}$$

$$B_\mu \rightarrow B_\mu - \epsilon/c_W B'_\mu$$

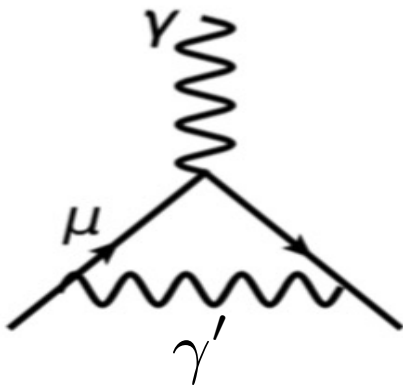
$$B'_\mu \rightarrow B'_\mu$$

$$\mathcal{L} \supset - \boxed{\epsilon e J_{EM}^\mu A'_\mu} - \epsilon \tan \theta_W \frac{m_{\gamma'}^2}{m_Z^2} g_Z J_{NC}^\mu A'_\mu$$

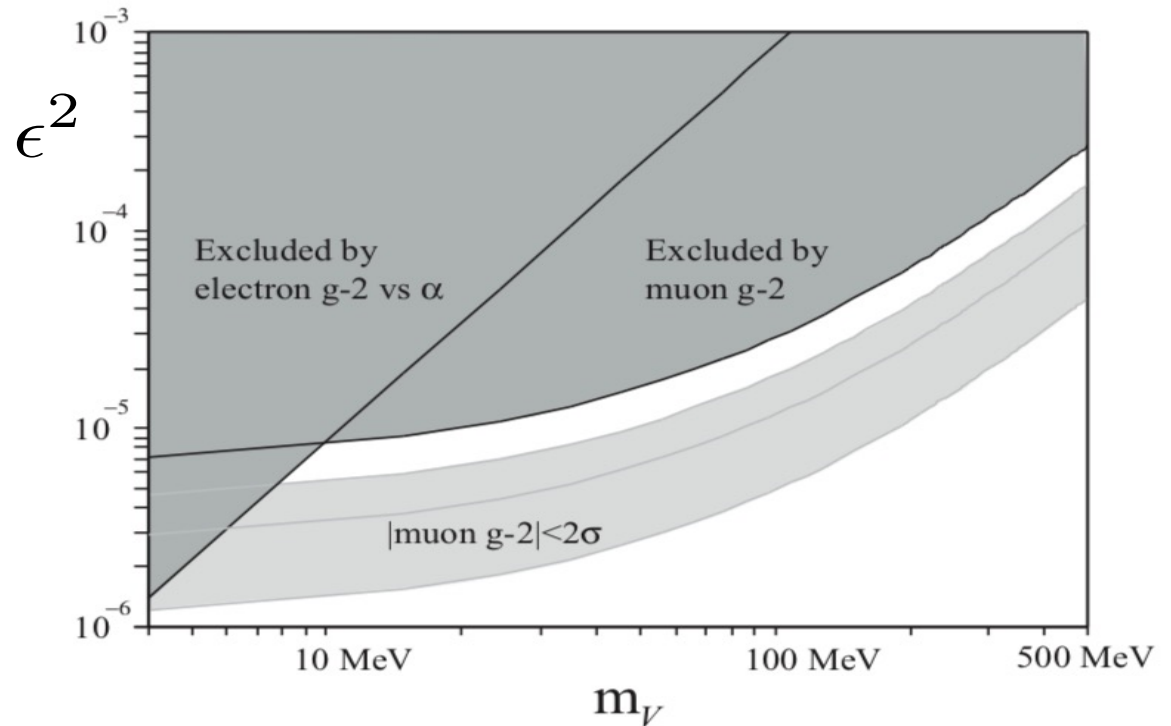
Anomalous muon g-2



$$\epsilon e J_\mu A'^\mu$$



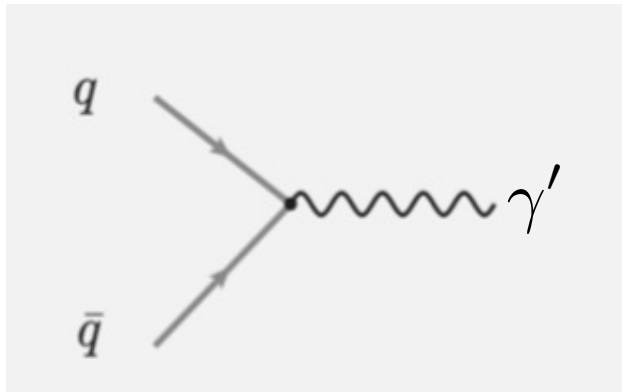
M. Pospelov, PRD 80 (2009) 095002



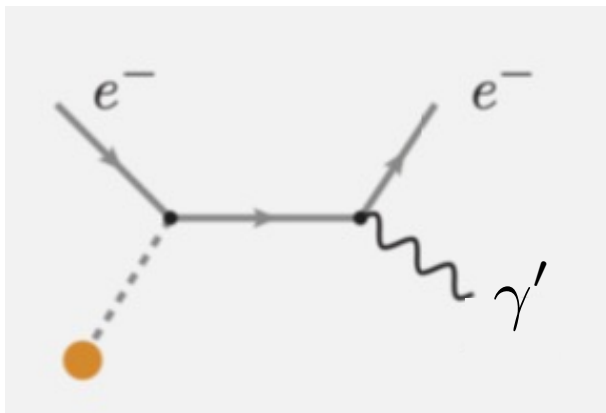
Dark photon search

productions: $\epsilon e J_\mu A'^\mu$

$$q\bar{q}(\text{or } e^+e^-) \rightarrow \gamma'$$

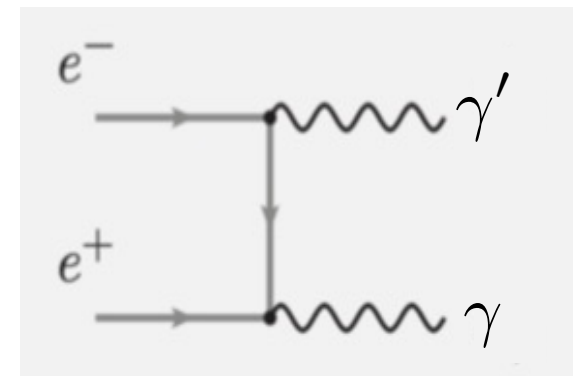


$$e^- Z \rightarrow e^- Z \gamma'$$



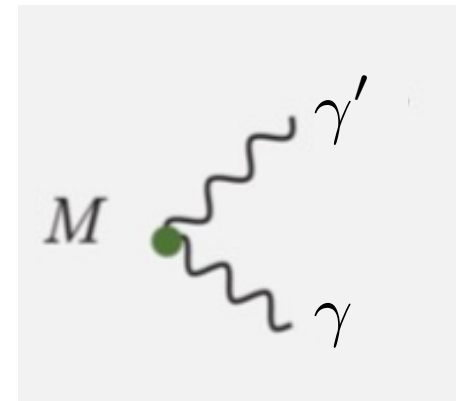
Fabbrichesi, Gabrielli, Lanfranchi, 2005.01515

$$q\bar{q}(\text{or } e^+e^-) \rightarrow \gamma'\gamma$$



meson decay

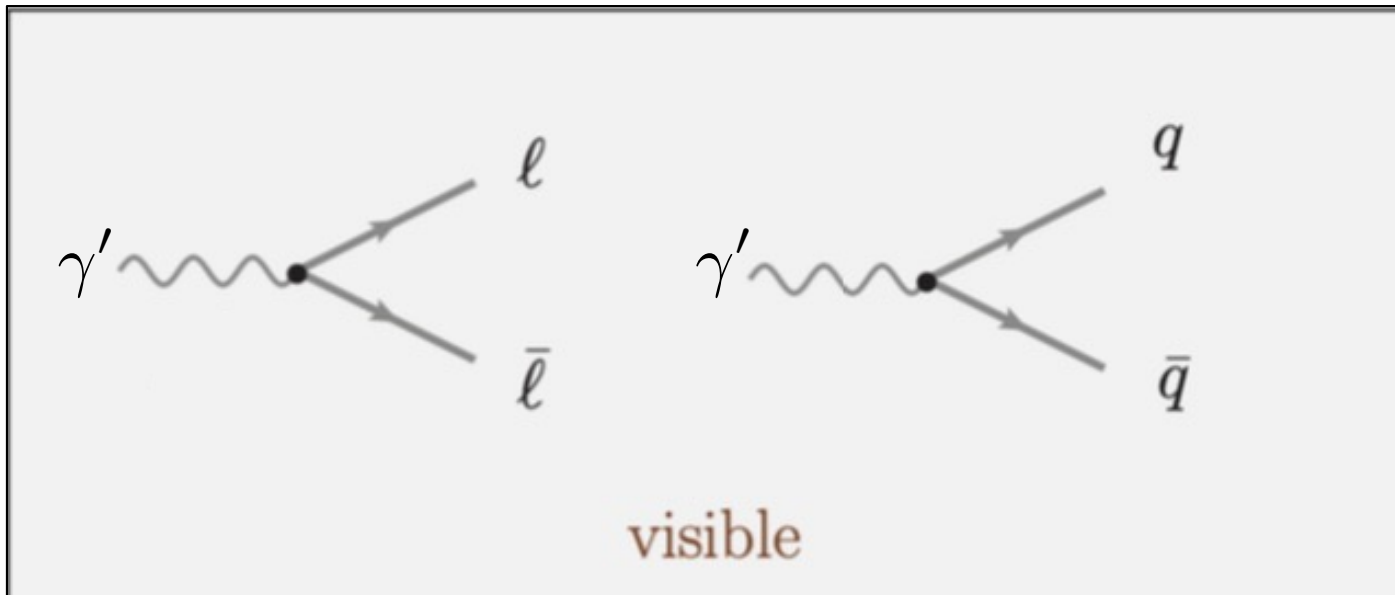
$$M \rightarrow \gamma'\gamma$$



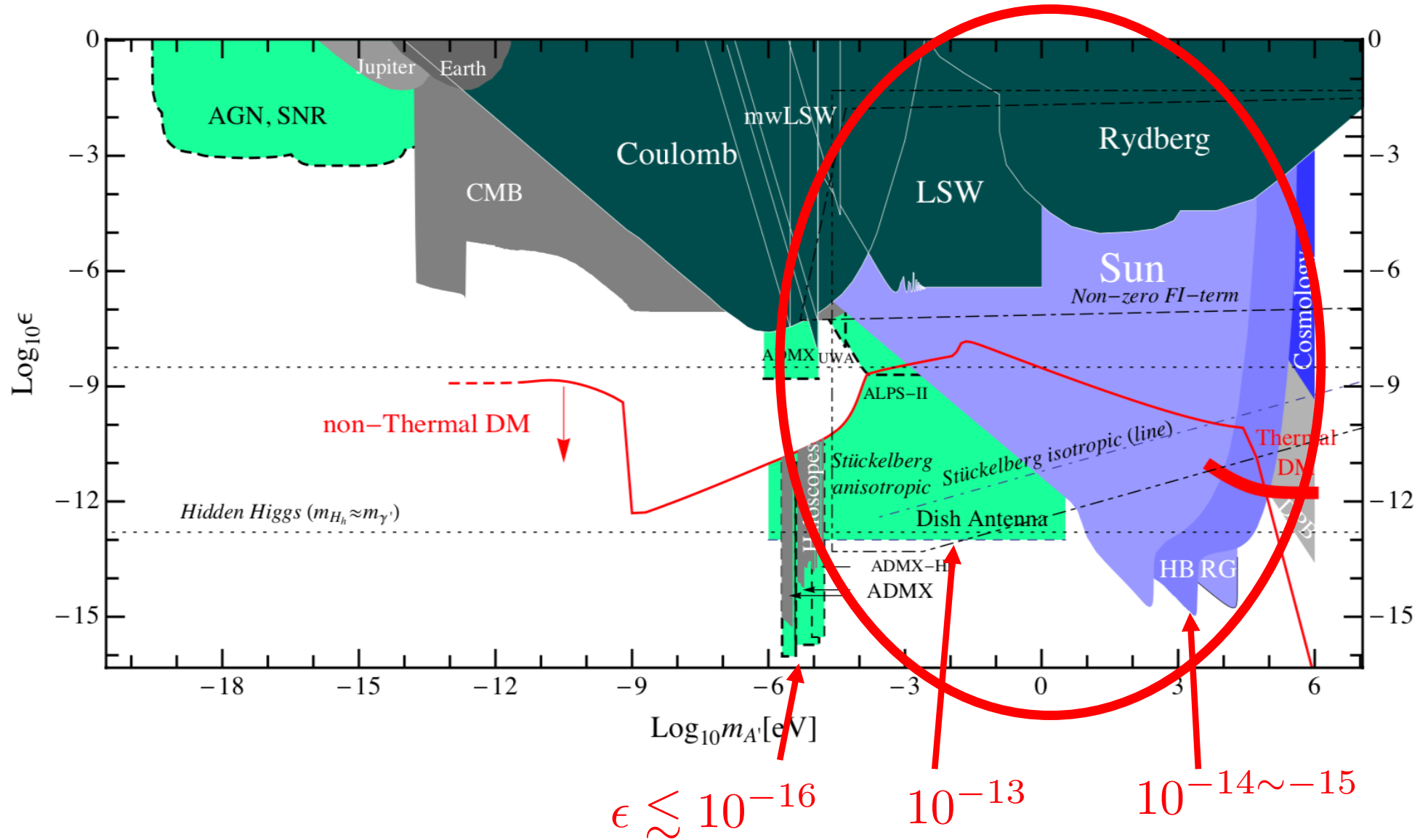
Dark photon search

Visible decays: $\epsilon e J_\mu A'^\mu$

Fabbrichesi, Gabrielli, Lanfranchi, 2005.01515



Dark photon constraints

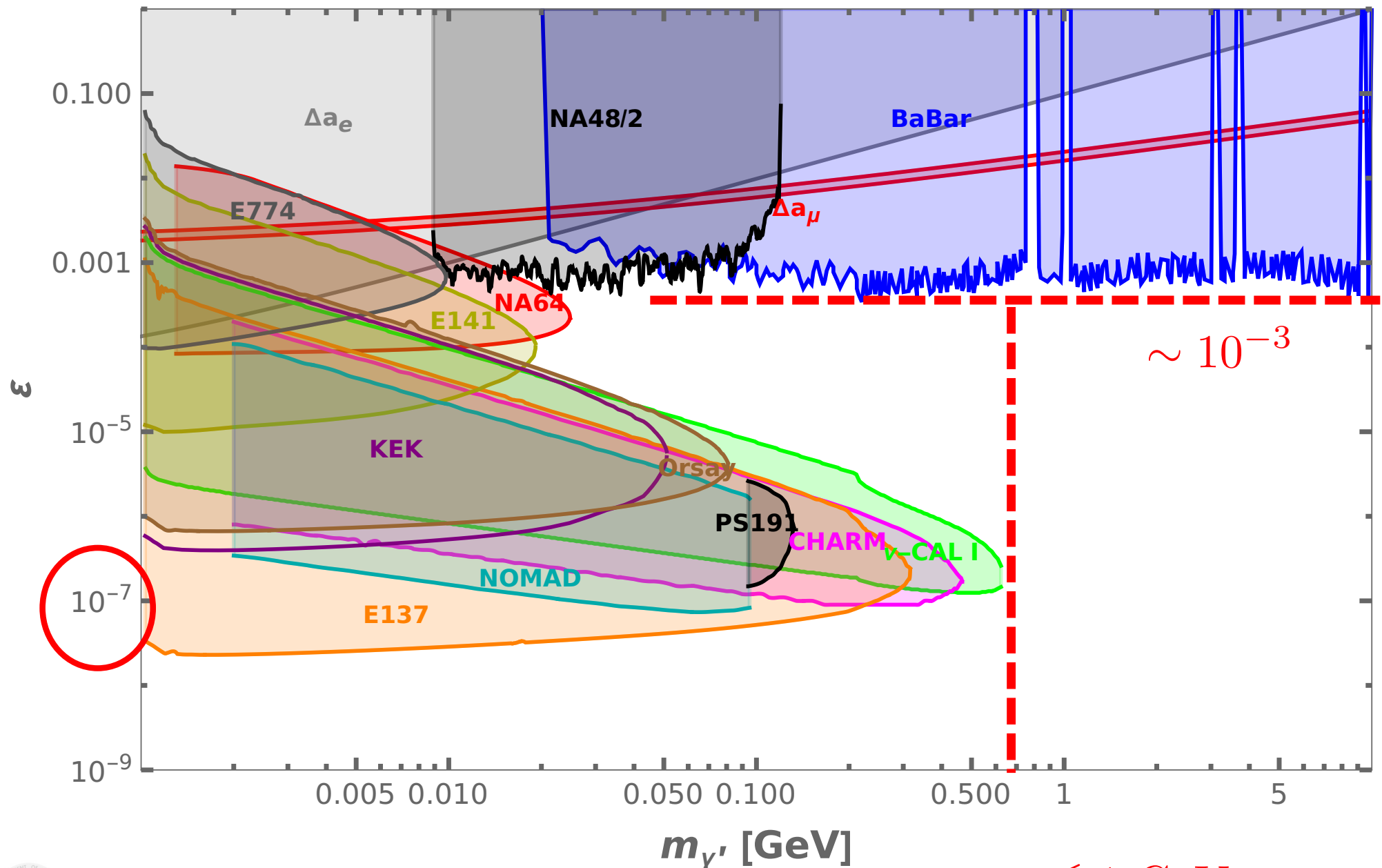


Stringent constraints for light dark photon with mass lighter than 1 MeV



Dark photon constraints

experiment Ilten, Soreq, Williams, Xue, 1801.04847



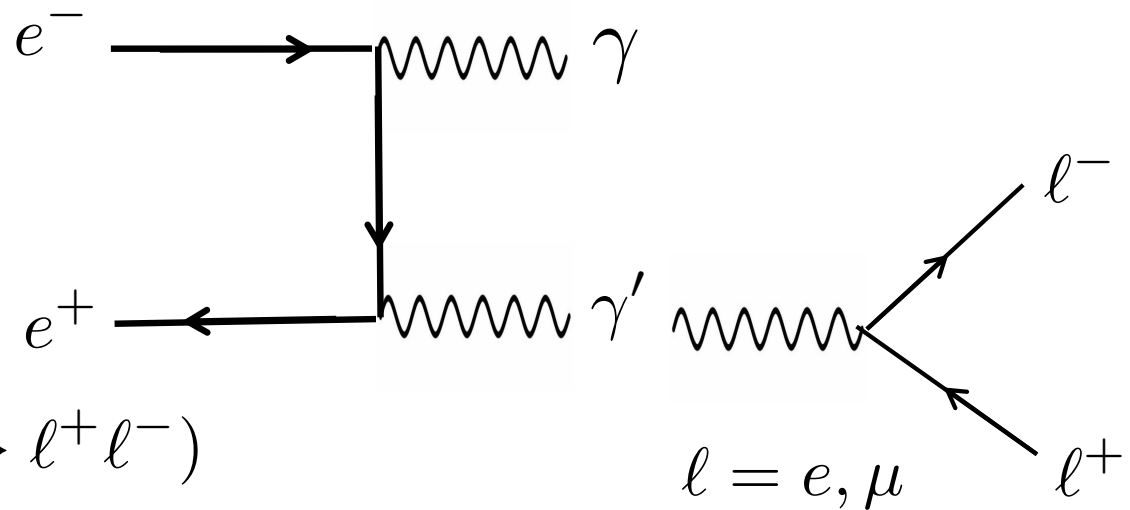
"heavy" dark photon: Stringent constraints for $m_{\gamma'} \lesssim 1$ GeV

Dark photon search

BaBar 514 fb^{-1}

$m: 0.02 \text{ GeV} \sim 10.2 \text{ GeV}$

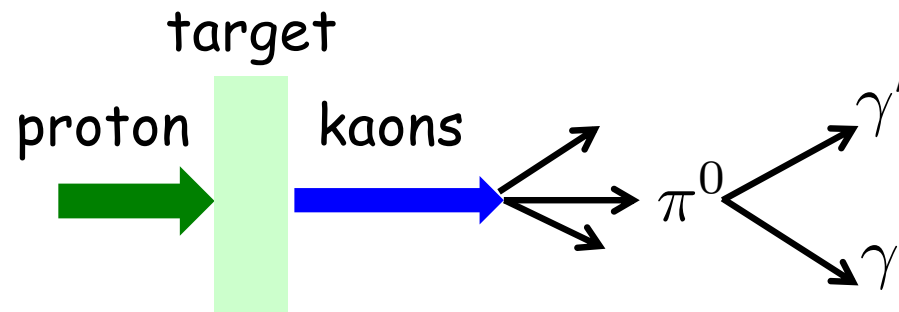
$$\sigma(e^+e^- \rightarrow \gamma\gamma') \times \text{Br}(\gamma' \rightarrow \ell^+\ell^-)$$



NA48/2

$$\sim 1.69 \times 10^7 \pi^0$$

$m: 9 \text{ MeV} \sim 70 \text{ MeV}$



$$\mathcal{B}(\pi^0 \rightarrow \gamma A') = 2\varepsilon^2 \left(1 - \frac{m_{A'}^2}{m_{\pi^0}^2}\right)^3 \mathcal{B}(\pi^0 \rightarrow \gamma\gamma)$$

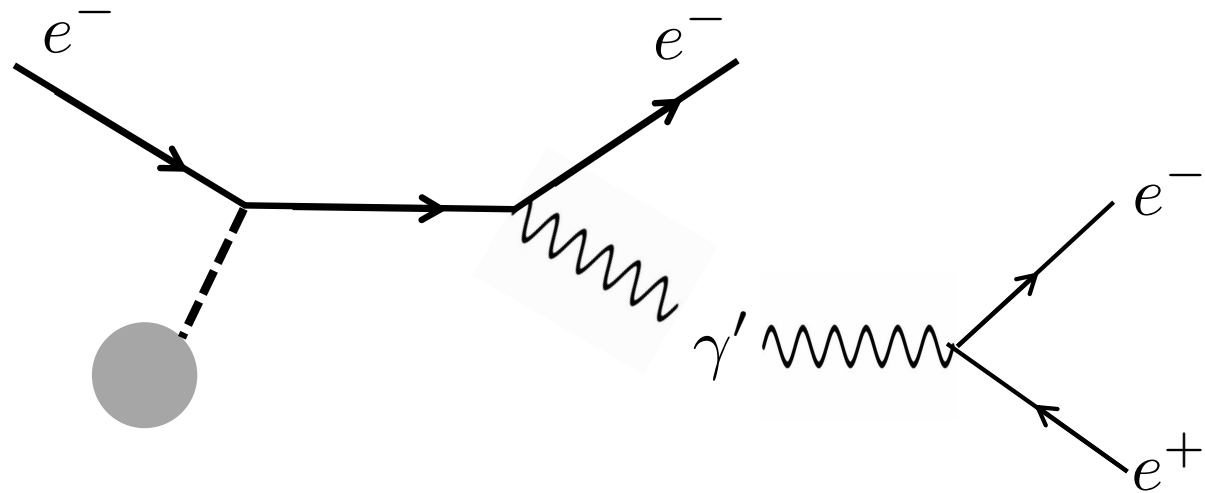
$$c\tau_{A'} = \hbar c / \Gamma_{A'} \approx 0.8 \mu\text{m} \times \left(\frac{10^{-6}}{\varepsilon^2}\right) \times \left(\frac{100 \text{ MeV}/c^2}{m_{A'}}\right)$$



Dark photon search

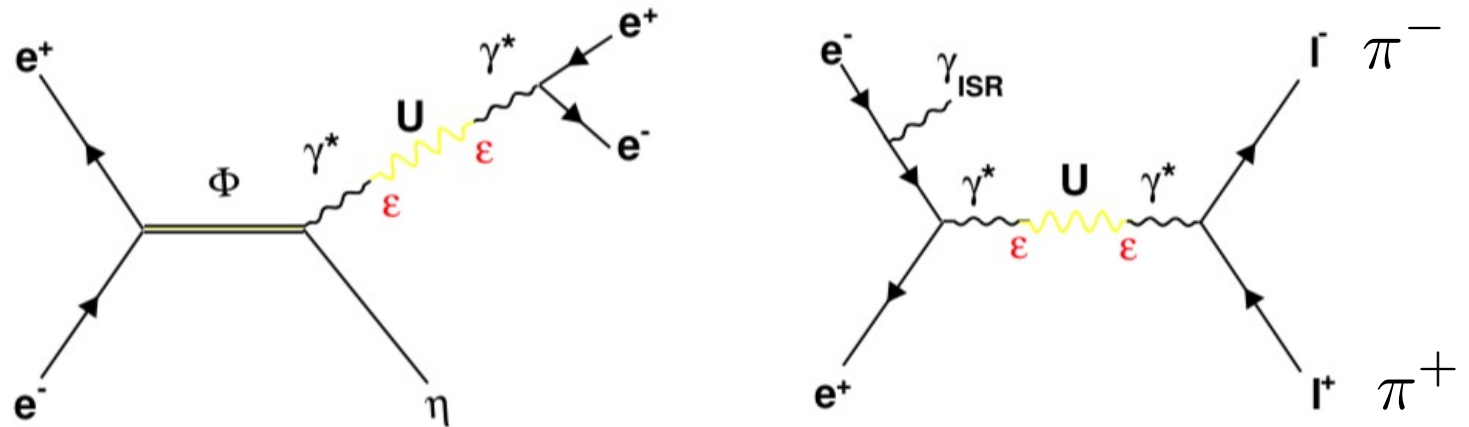
A1

40 MeV ~ 300 MeV



KLOE

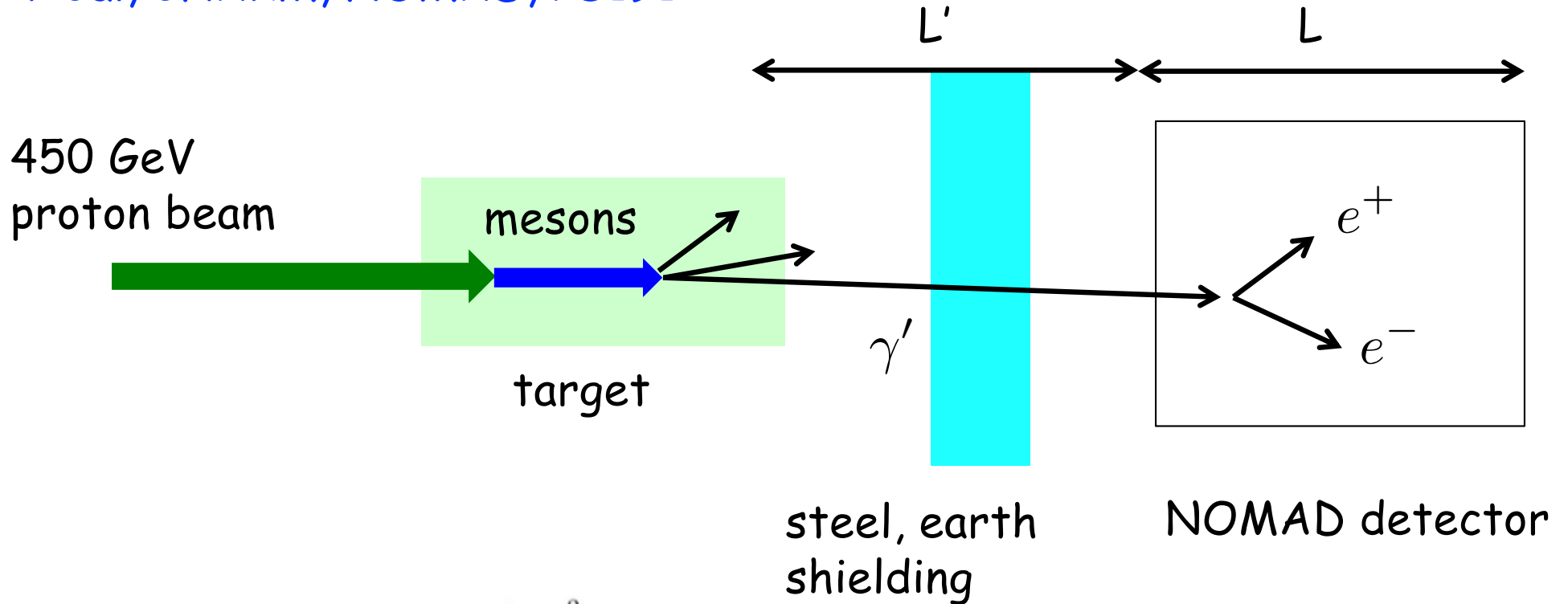
5 MeV ~ 987 MeV



$$\sigma(\phi \rightarrow \eta U) = \epsilon^2 |F_{\phi\eta}(m_U^2)|^2 \frac{\lambda^{3/2}(m_\phi^2, m_\eta^2, m_U^2)}{\lambda^{3/2}(m_\phi^2, m_\eta^2, 0)} \sigma(\phi \rightarrow \eta\gamma)$$

Dark photon search

ν -Cal, CHARM, NOMAD, PS191



$$\sigma(pp \rightarrow \gamma' X) = 2\epsilon^2 \left(1 - \frac{m_{\gamma'}^2}{m_{\pi^0}^2}\right)^3 \text{Br}(\pi^0 \rightarrow \gamma\gamma) \sigma(pp \rightarrow \pi^0 X)$$

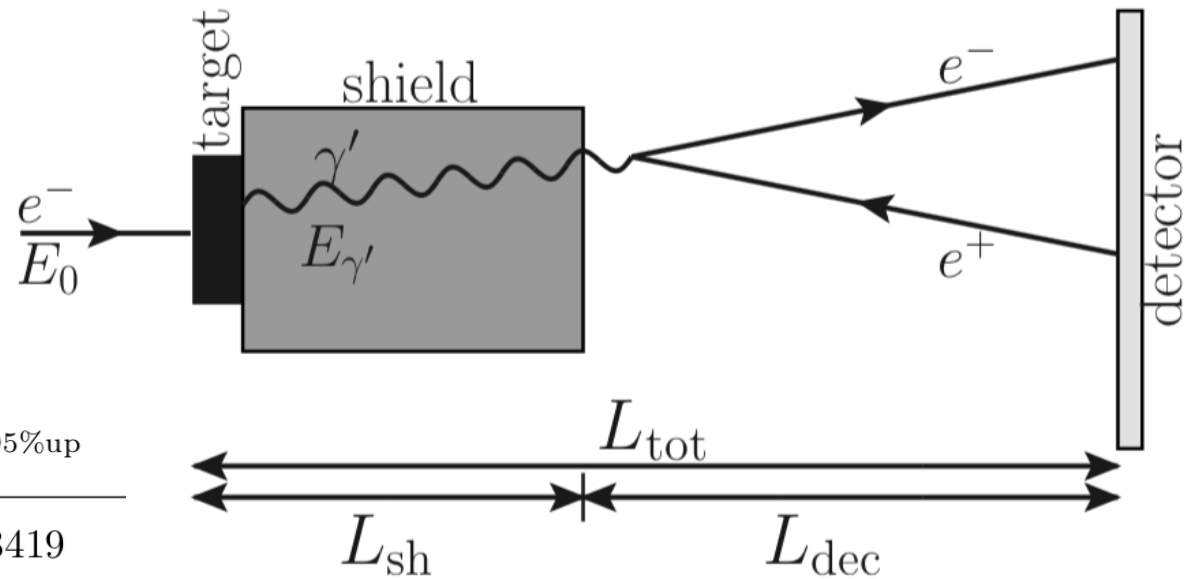
$$N \propto N_{POT} \times \sigma(\gamma' X) \times w$$

$$w = \exp^{-\frac{L' m_{\gamma'}}{\tau_{\gamma'} p_{\gamma'}} \left[1 - \exp^{-\frac{L m_{\gamma'}}{\tau_{\gamma'} p_{\gamma'}}\right]$$



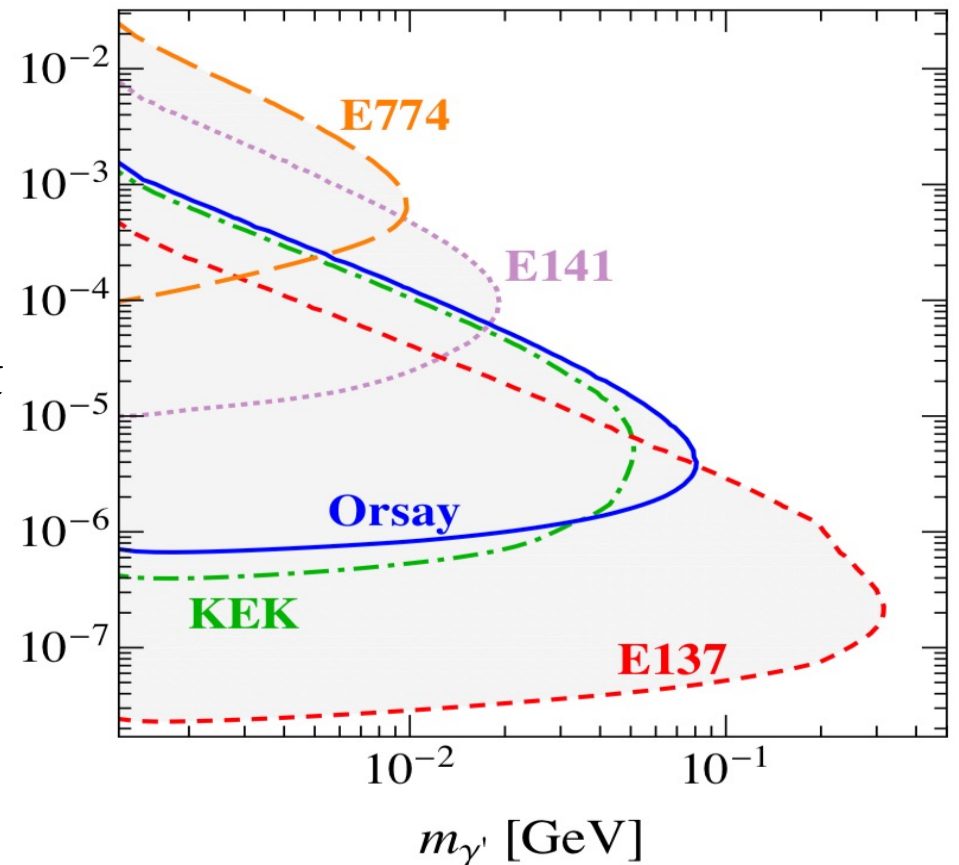
Dark photon search

E141, E137, E774,
KEK, Orsay



Experiment	L_{sh} [m]	L_{dec} [m]	N_{obs}	$N_{95\%up}$
E141 [47]	0.12	35	1126^{+1312}_{-1126}	3419
E137 [48]	179	204	0	3
E774 [49]	0.3	2	0_{-0}^{+9}	18
KEK [39]	2.4	2.2	0	3
Orsay [40]	1	2	0	3

$$N \simeq N_e \frac{N_0 X_0}{A} \int_{m_{\gamma'}}^{E_0 - m_e} dE_{\gamma'} \int_{E_{\gamma'} + m_e}^{E_0} dE_e \int_0^T dt \epsilon \left[I_e(E_0, E_e, t) \frac{1}{E_e} \frac{d\sigma}{dx_e} \Big|_{x_e = \frac{E_{\gamma'}}{E_e}} \right] e^{-L_{sh}/l_{\gamma'}} \left(1 - e^{-L_{dec}/l_{\gamma'}} \right) \text{BR}_{\text{detect}} \cdot$$



Axion-like Particle

$$\mathcal{L}_{\text{CP-viol.}} = \frac{\alpha_s}{8\pi} \theta G_{\mu\nu}^a \tilde{G}^{a,\mu\nu} \quad \bar{\theta} \equiv \theta + \arg \det M$$

$$|d_n| \sim \frac{e}{m_n} \left(\frac{m_q}{m_n} \right) |\bar{\theta}| \sim 10^{-16} |\bar{\theta}| \text{ e cm} < 2.9 \times 10^{-26} \text{ e cm} \longrightarrow |\bar{\theta}| \lesssim 10^{-10}$$

promote $\theta \rightarrow a/f_a$

$$U(1)_{\text{PQ}} \longrightarrow \text{axion} \quad m_a = \frac{m_\pi f_\pi}{f_a} \frac{\sqrt{m_u m_d}}{m_u + m_d} \simeq 0.6 \text{ meV} \times \left(\frac{10^{10} \text{ GeV}}{f_a} \right)$$

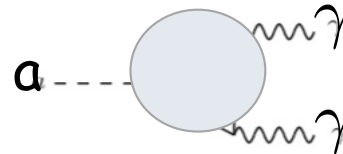
γ

Similar idea \rightarrow breaking of other global symmetry $U(1)_\phi$
 \rightarrow PNCB (ALP a)

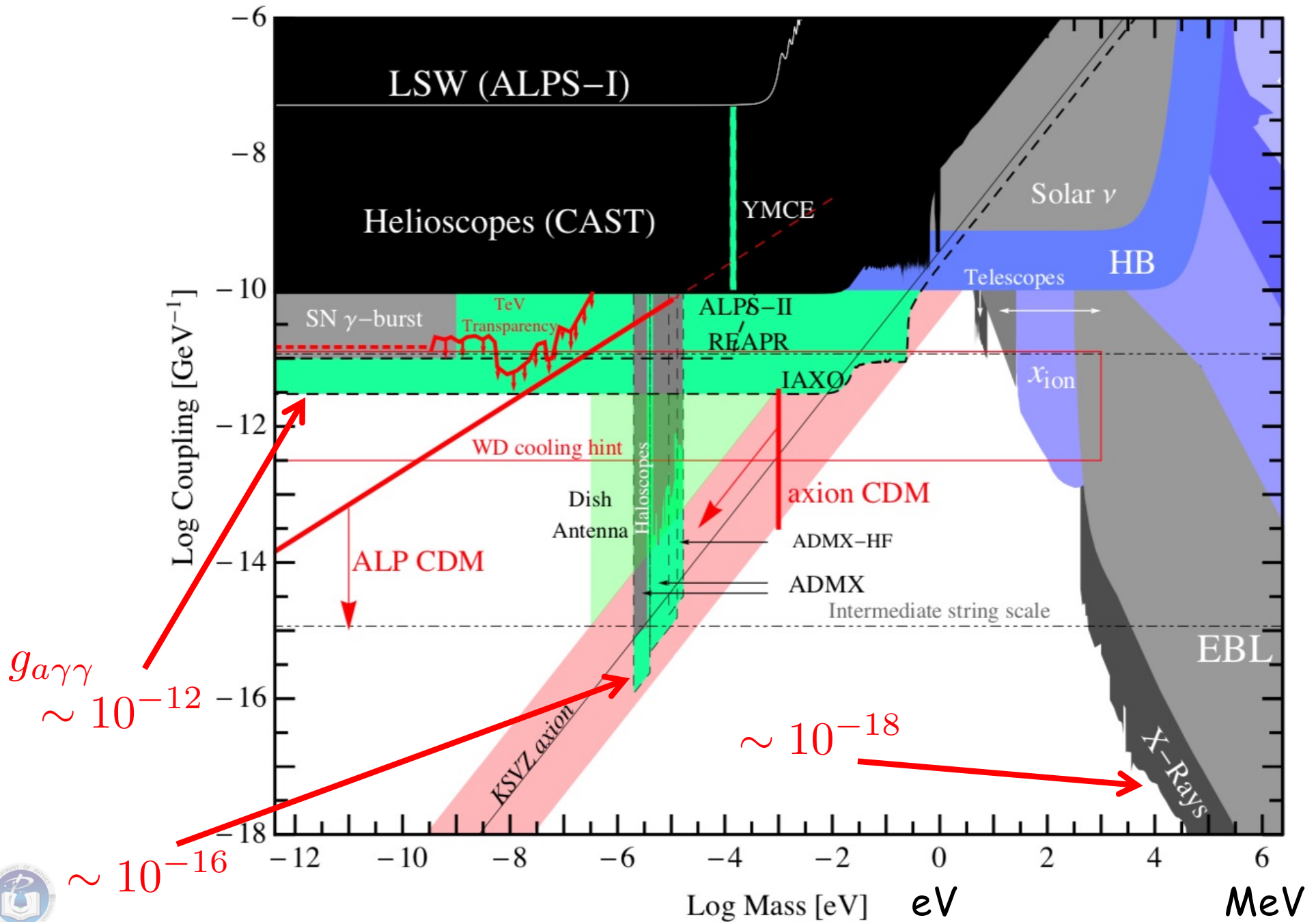
$$\phi = \frac{1}{\sqrt{2}} (v_\phi + s) e^{ia/v_\phi}$$

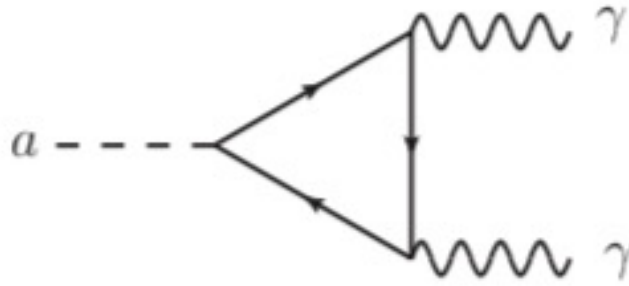
interactions:

$$\frac{1}{4} g_{a\gamma\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

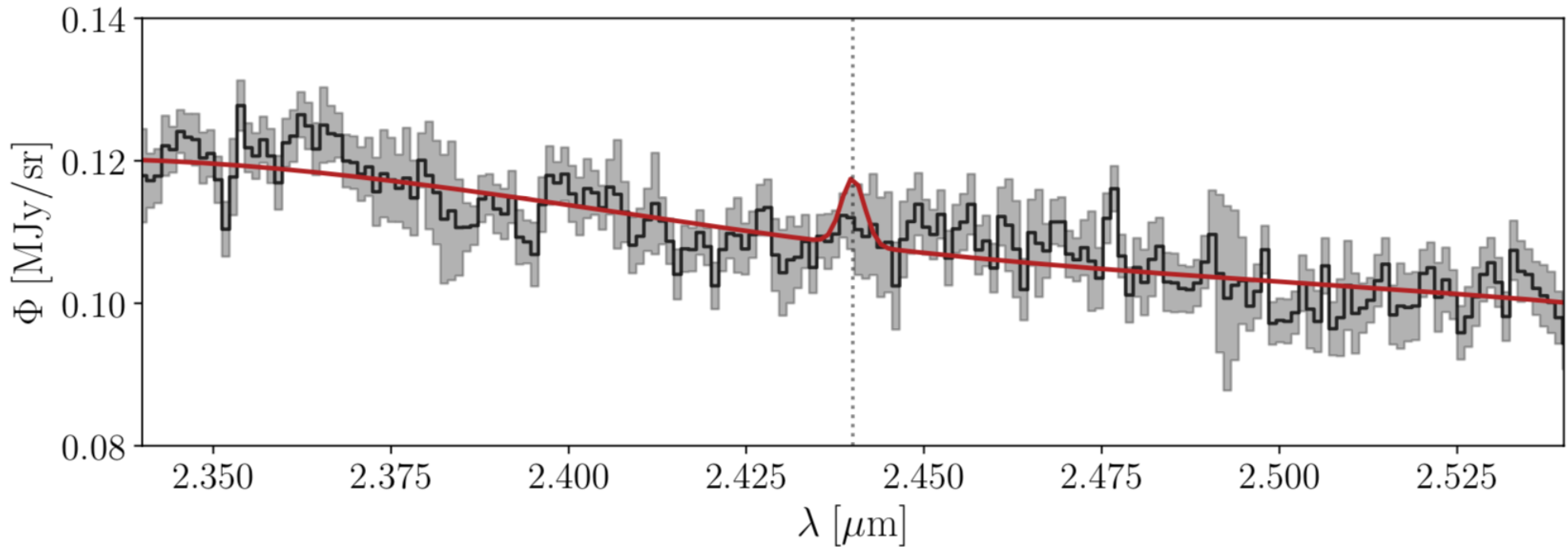


ALP constraints



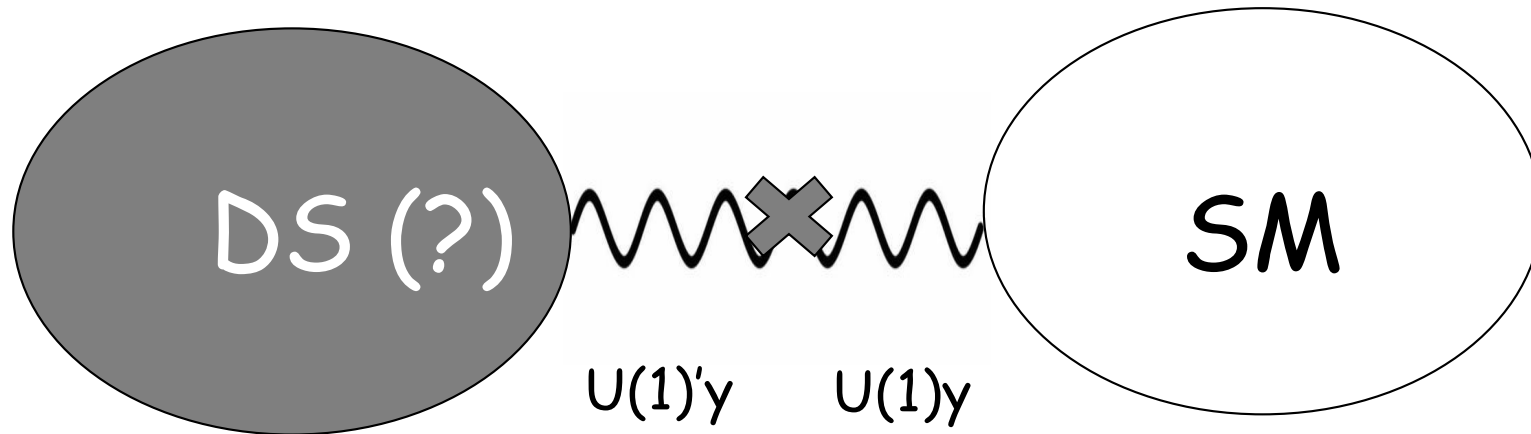


Janish, Pinetti, 2310.15396



JWST spectrum observation

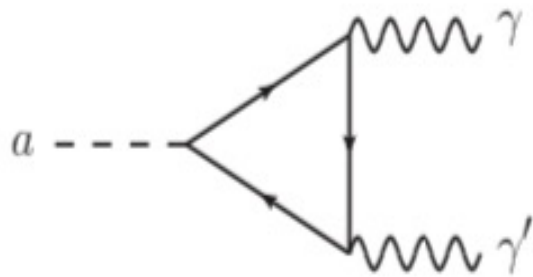
$$g_{a\gamma\gamma} < 2.4 \times 10^{-11} \text{ GeV}^{-1} \quad \text{for} \quad m_a = 1 \text{ eV}$$



if there exists axion-like particle in DS,

$$\mathcal{L} \supset \frac{1}{4} g_{a\gamma'\gamma'} a F'_{\mu\nu} \tilde{F}'^{\mu\nu} + \frac{1}{4} g_{a\gamma\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu} + \frac{1}{2} g_{a\gamma'\gamma} a F'_{\mu\nu} \tilde{F}^{\mu\nu}$$

Kaneta, Lee, Yun, PRL 118 (2017) 10, 101802



Induce a new decay channel
for dark photon

$$\gamma' \rightarrow a\gamma$$

photon + missing energy

An Example KSVZ-type

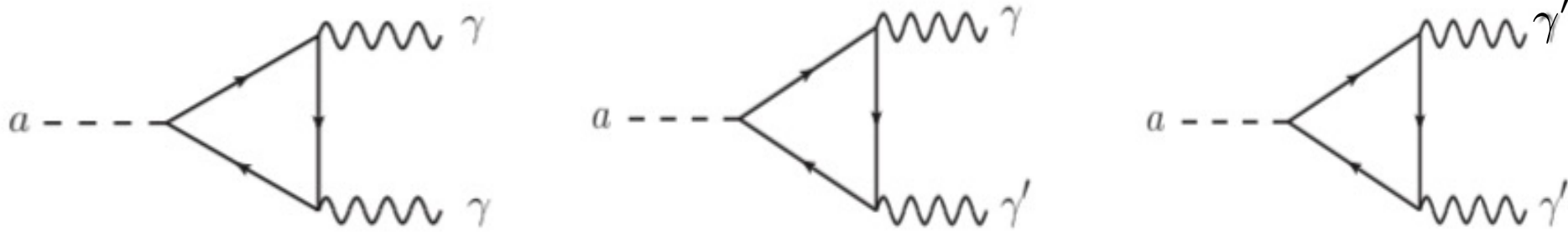


TABLE I. New fields and charge assignments in our model. Q_ψ is the electromagnetic charge of the exotic fermion ψ .

Field	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_{\text{Dark}}$	$U(1)_{\text{PQ}}$
ψ	3	1	Q_ψ	D_ψ	PQ_ψ
ψ^c	$\bar{3}$	1	$-Q_\psi$	$-D_\psi$	PQ_{ψ^c}
Φ_{PQ}	1	1	0	0	PQ_Φ
Φ_D	1	1	0	D_Φ	0

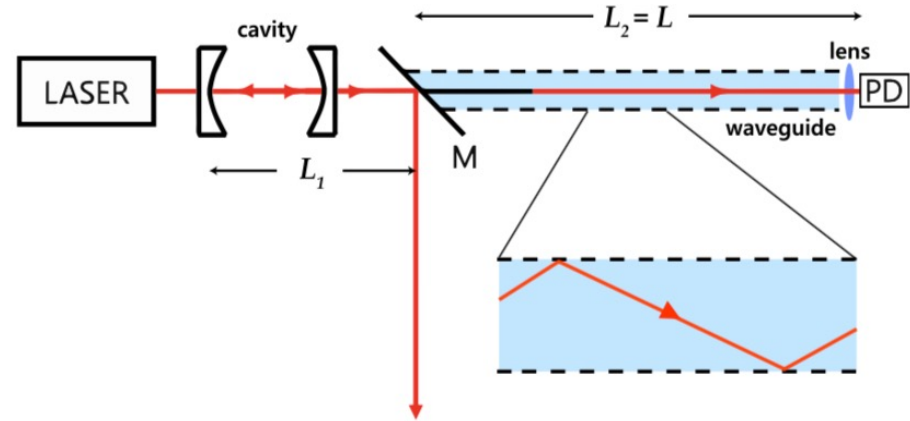
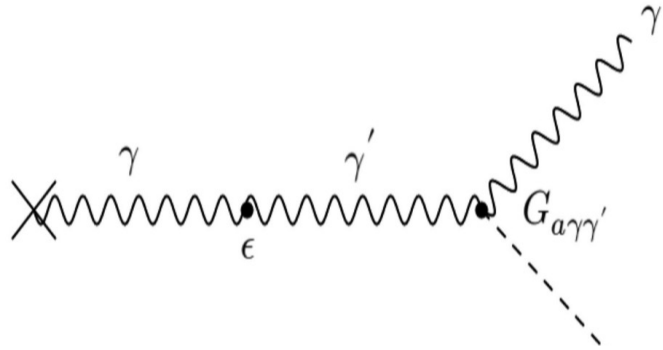
$$g_{a\gamma\gamma} = \frac{e^2}{8\pi^2} \frac{\text{PQ}_\Phi}{f_a} (2N_c Q_\psi^2)$$

Kaneta, Lee, Yun, PRL 118 (2017) 10, 101802

$$g_{a\gamma\gamma'} = \frac{ee'}{8\pi^2} \frac{\text{PQ}_\Phi}{f_a} (2N_c Q_\psi D_\psi) + \epsilon g_{a\gamma\gamma}$$

$$g_{a\gamma'\gamma'} = \frac{e'^2}{8\pi^2} \frac{\text{PQ}_\Phi}{f_a} (2N_c D_\psi^2) + 2\epsilon g_{a\gamma\gamma'}$$

Constraints of $g_{a\gamma\gamma'}$



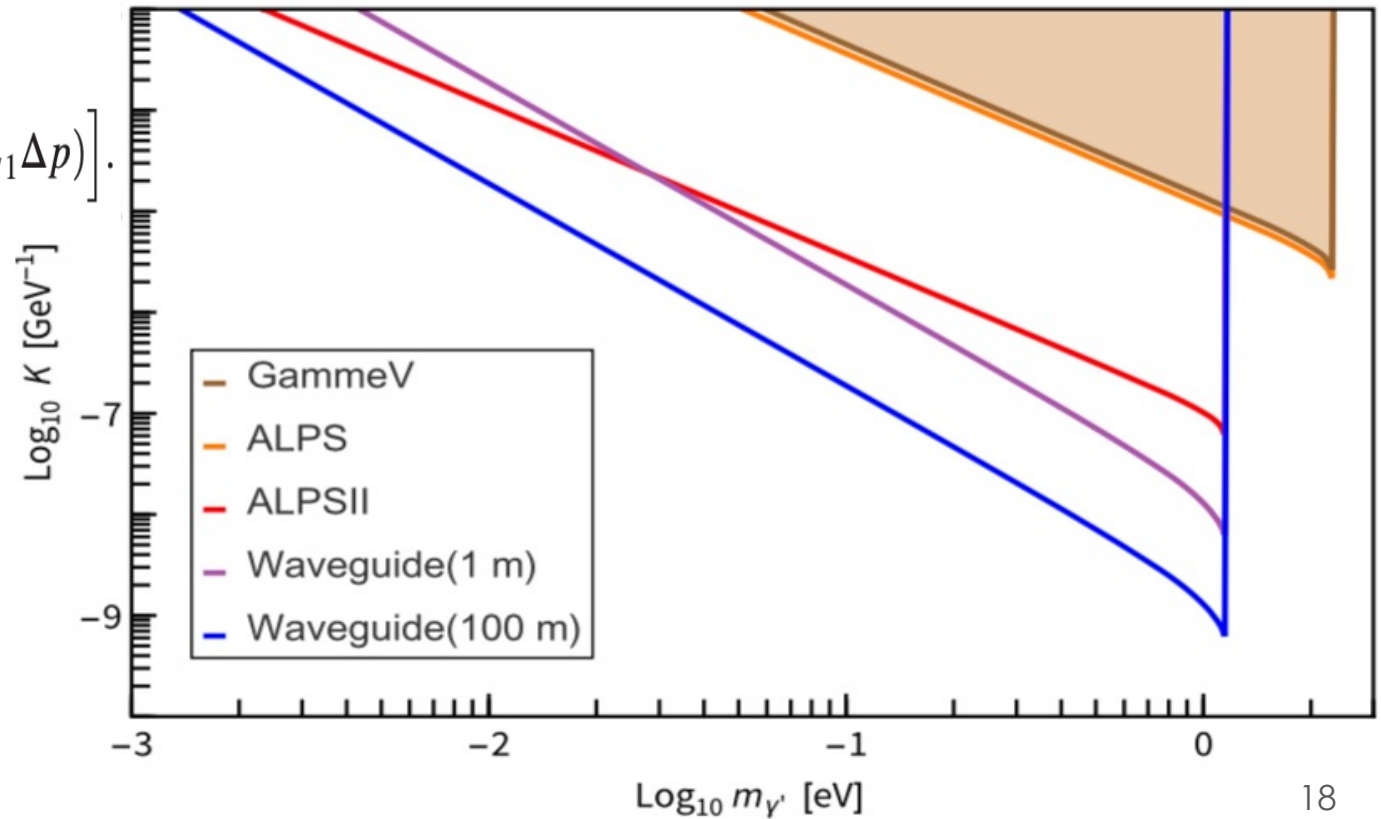
H.-S. Lee, J. Lee, J. Yi, PRD 106, 015011 (2022)

$$N_{\text{sub}} = \mathcal{P}_{A \rightarrow X}(L_1) \mathcal{P}_{X \rightarrow \text{decay}}(L_2) N_\gamma,$$

$$= \epsilon^2 N_\gamma \left[1 - e^{-\frac{m_{\gamma'} \Gamma L_2}{p}} \right]$$

$$\times \left[1 + e^{-\frac{m_{\gamma'} \Gamma L_1}{p}} - 2e^{-\frac{m_{\gamma'} \Gamma L_1}{2p}} \cos(L_1 \Delta p) \right].$$

$$K = \epsilon g_{a\gamma\gamma'}$$

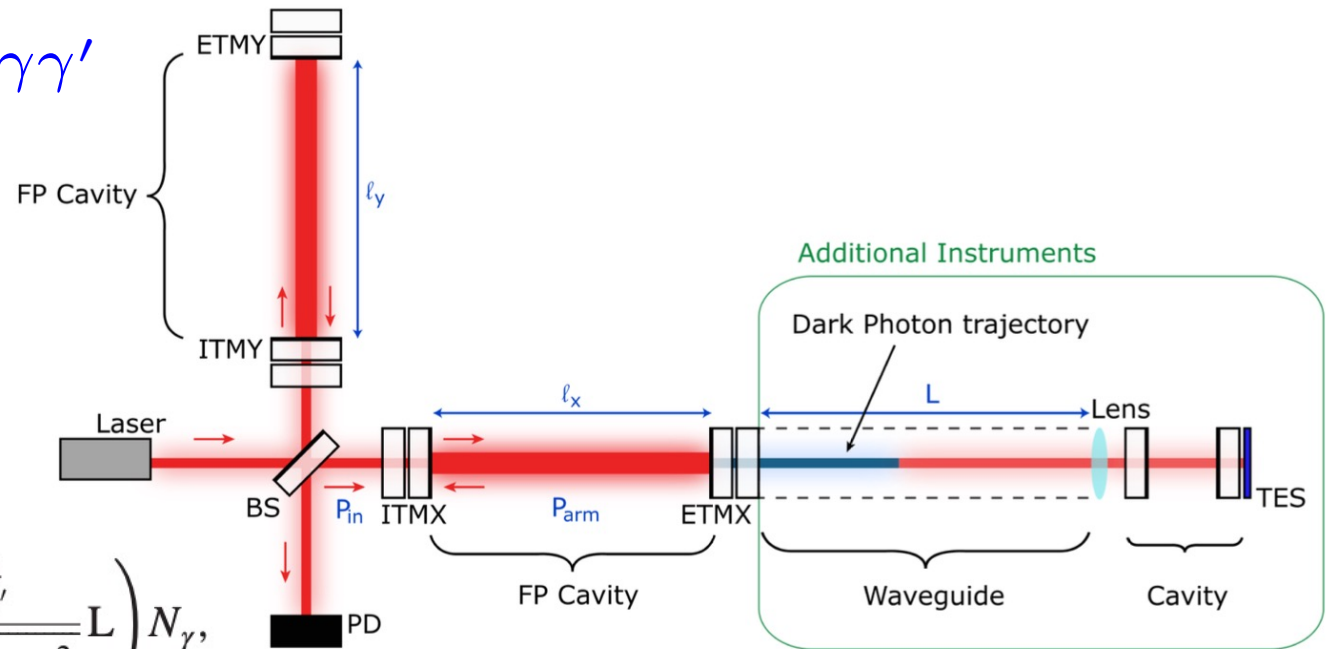


Constraints of $g_{a\gamma\gamma'}$

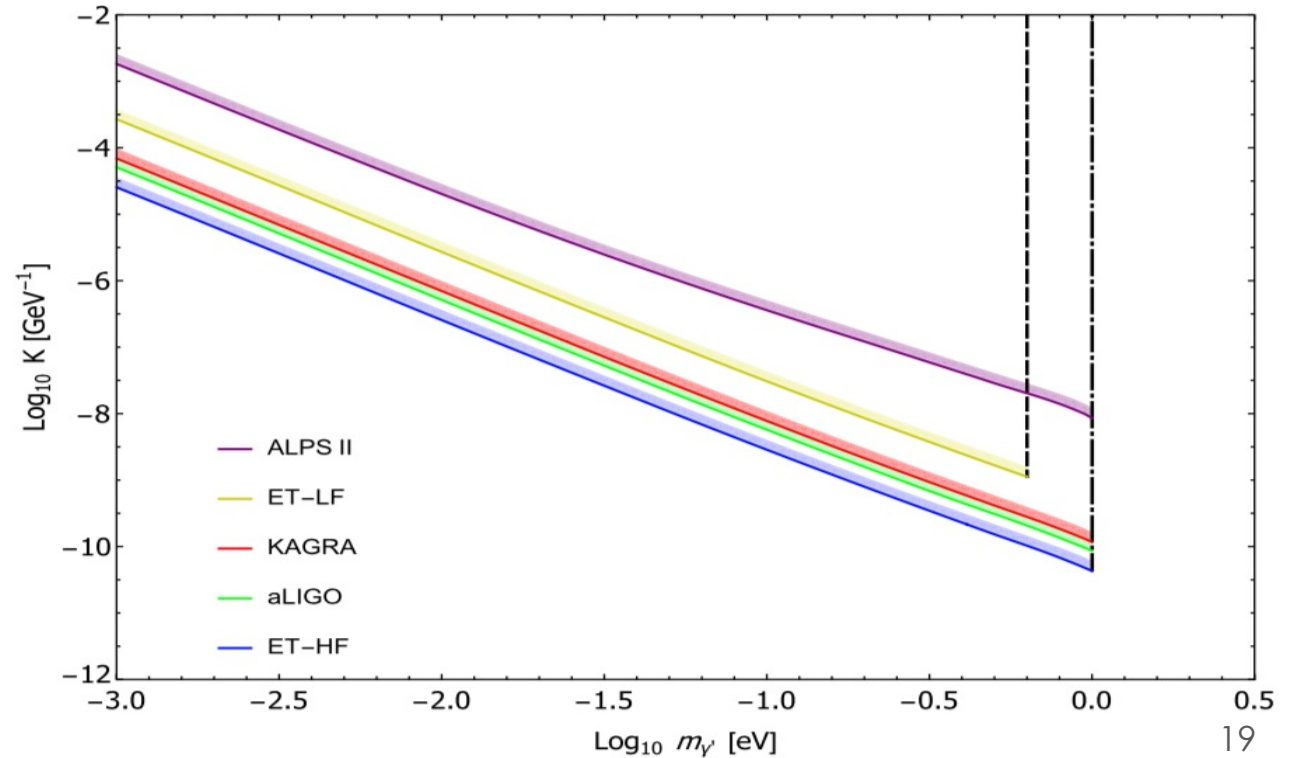
Ismail, Nugroho, Wong,
PRD 107, 082002 (2023)

$$\text{SNR} = \frac{N_s \sqrt{t_s}}{\sqrt{N_s + N_d}}$$

$$N_s^{\text{tot}} = \eta_{\text{kin}} \eta_{\text{cav}} \bar{n}_{\text{eff}} \eta_{\text{cav}}^{\text{WG}} \left(\frac{K^2}{96\pi} \frac{m_{\gamma'}^4}{\sqrt{\omega^2 - m_{\gamma'}^2}} L \right) N_{\gamma'}$$

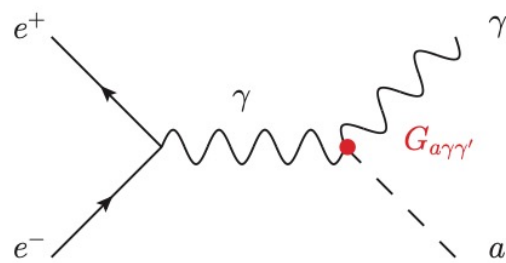


$$K = \epsilon g_{a\gamma\gamma'}$$



Constraints of $g_{a\gamma\gamma'}$

P. deNiverville, H.S. Lee, M.S. Seo, 1806.00757



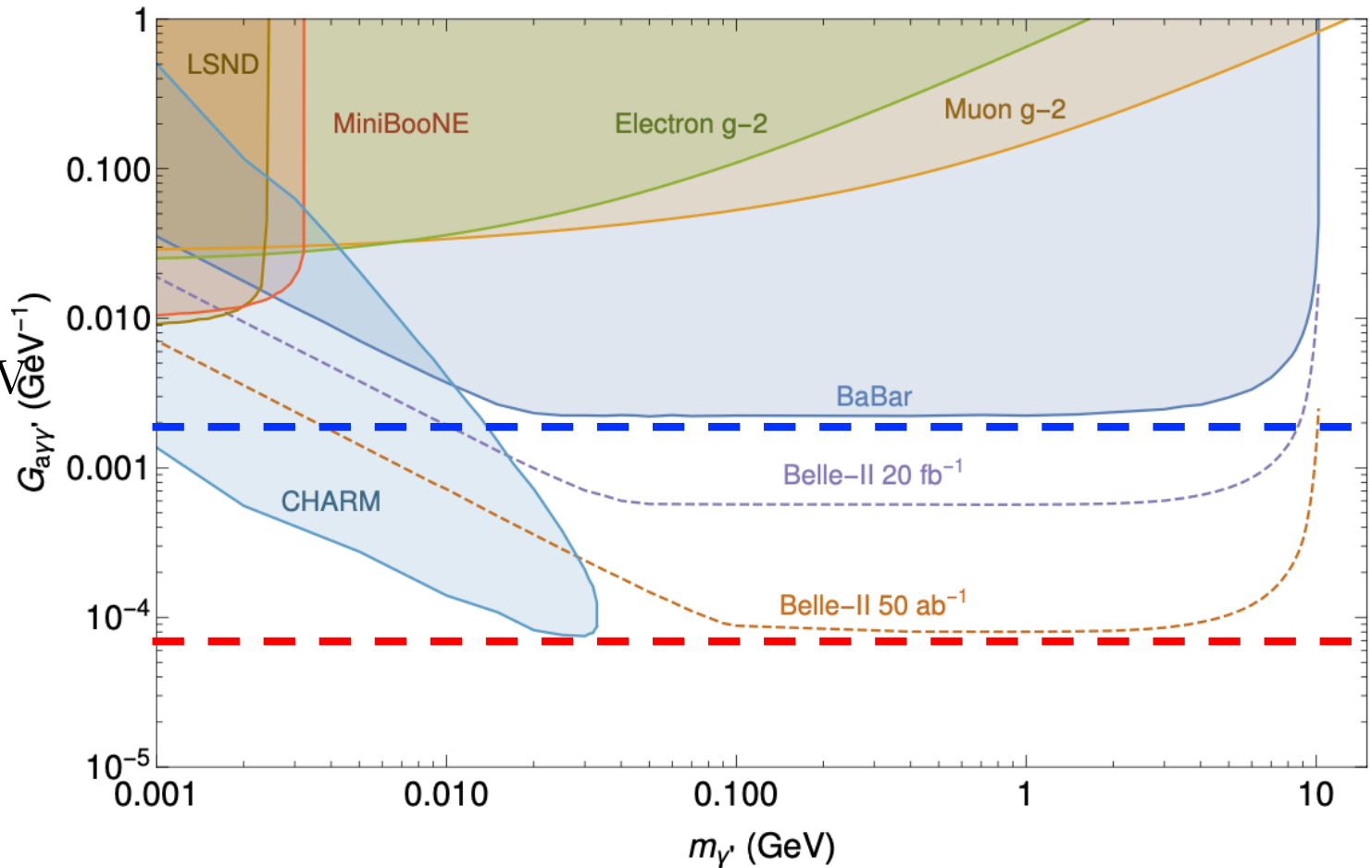
For $m_{\gamma'} \lesssim 50$ MeV

$$g_{a\gamma\gamma'} \lesssim 10^{-4}$$

For

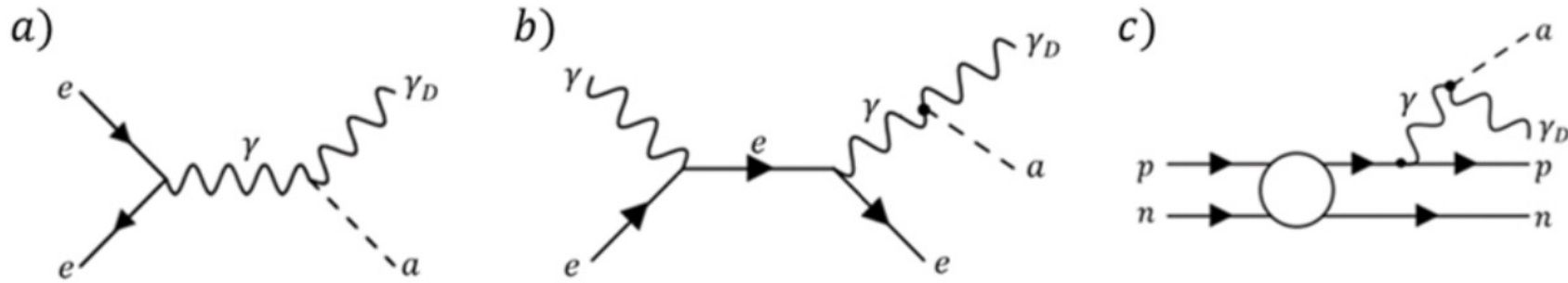
20 MeV $\lesssim m_{\gamma'} \lesssim 1$ GeV

$$g_{a\gamma\gamma'} \lesssim 10^{-3}$$



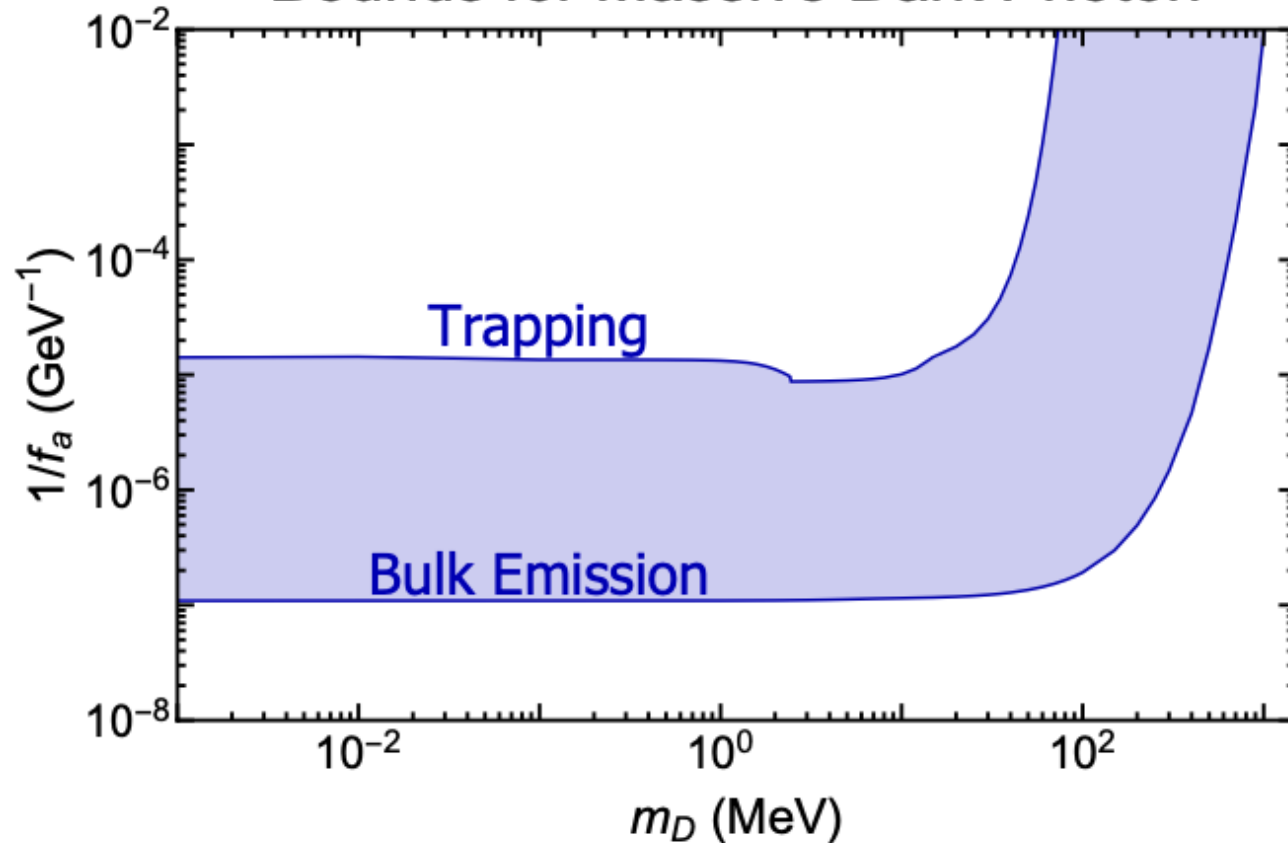
Constraints of $g_{a\gamma\gamma'}$

A. Hook, G. Marques-Tavares, C. Ristow, 2105.06476



Bounds for Massive Dark Photon

$$\mathcal{L} \supset \frac{1}{2} \frac{a}{f_a} F_{\mu\nu} \tilde{F}_D^{\mu\nu}$$



Decays

$$\Gamma(\gamma' \rightarrow \text{hadrons}) = \frac{\alpha\epsilon^2}{3} m_{\gamma'} \sqrt{1 - \frac{4m_\mu^2}{m_{\gamma'}^2}} \left(1 + \frac{2m_\mu^2}{m_{\gamma'}^2}\right) R$$

$$\Gamma(\gamma' \rightarrow l^+l^-) = \frac{\alpha\epsilon^2}{3} m_{\gamma'} \sqrt{1 - \frac{4m_l^2}{m_{\gamma'}^2}} \left(1 + \frac{2m_l^2}{m_{\gamma'}^2}\right)$$

$$\Gamma(\gamma' \rightarrow \gamma a) = \frac{g_{a\gamma\gamma'}^2}{96\pi} m_{\gamma'}^3 \left[1 - \frac{m_a^2}{m_{\gamma'}^2}\right]$$

$$\epsilon = 10^{-4}$$

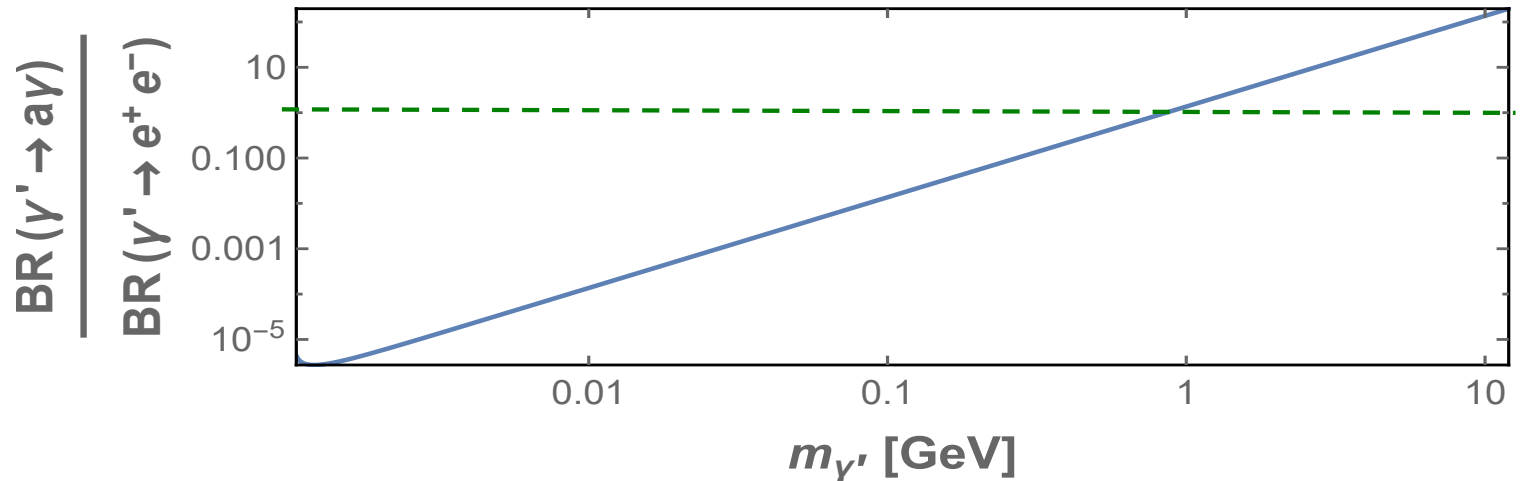
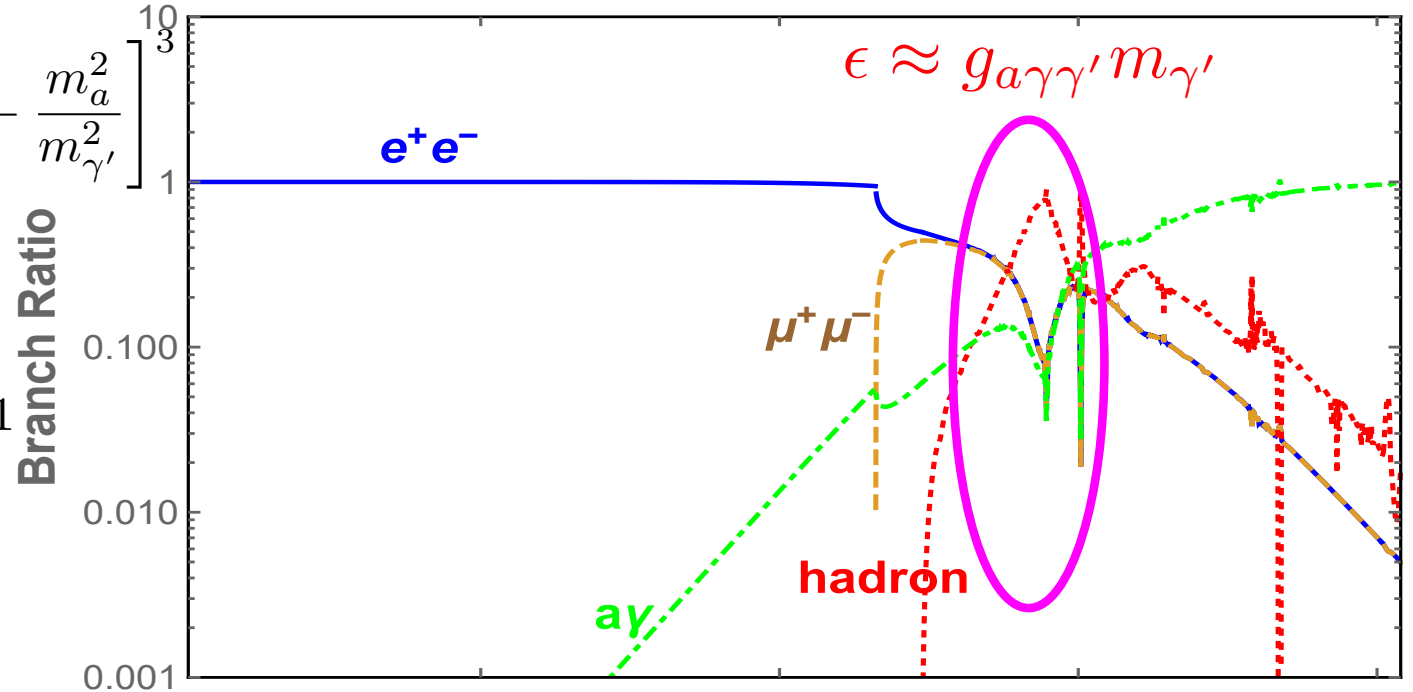
$$g_{a\gamma\gamma'} = 10^{-4} \text{ GeV}^{-1}$$

$$m_{\gamma'} \lesssim 300 \text{ MeV}$$

e^+e^- dominates

$$m_{\gamma'} \gtrsim 1 \text{ GeV}$$

$a\gamma$ is important



Decays

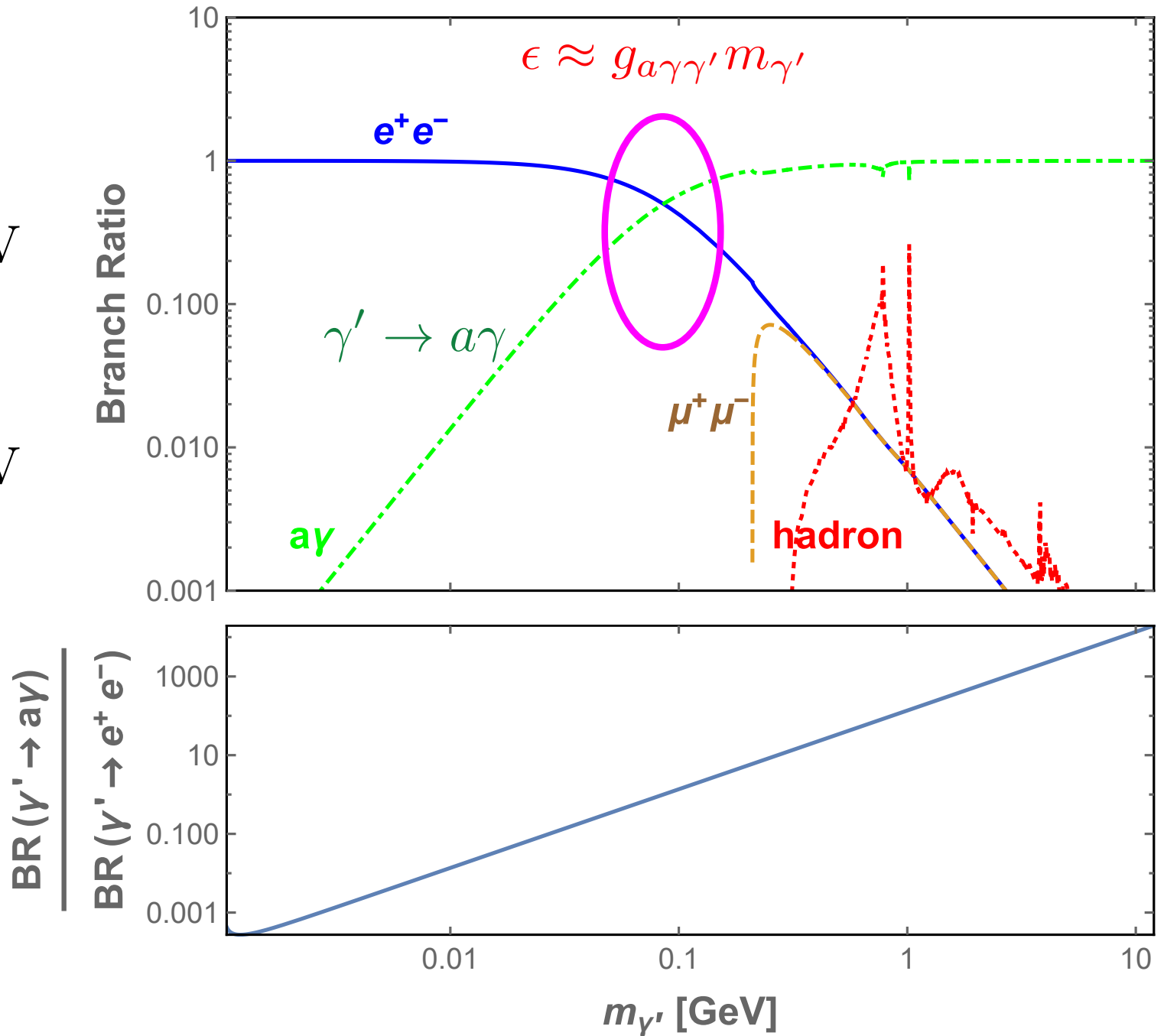
$$G_{a\gamma\gamma'} = 10^{-3} \text{ GeV}^{-1}, \epsilon = 10^{-4}$$

$m_{\gamma'} \lesssim 100 \text{ MeV}$

e^+e^- dominates

$m_{\gamma'} \gtrsim 100 \text{ MeV}$

$a\gamma$ is important



New Constraints

BaBar, KLOE, A1, NA48/2:

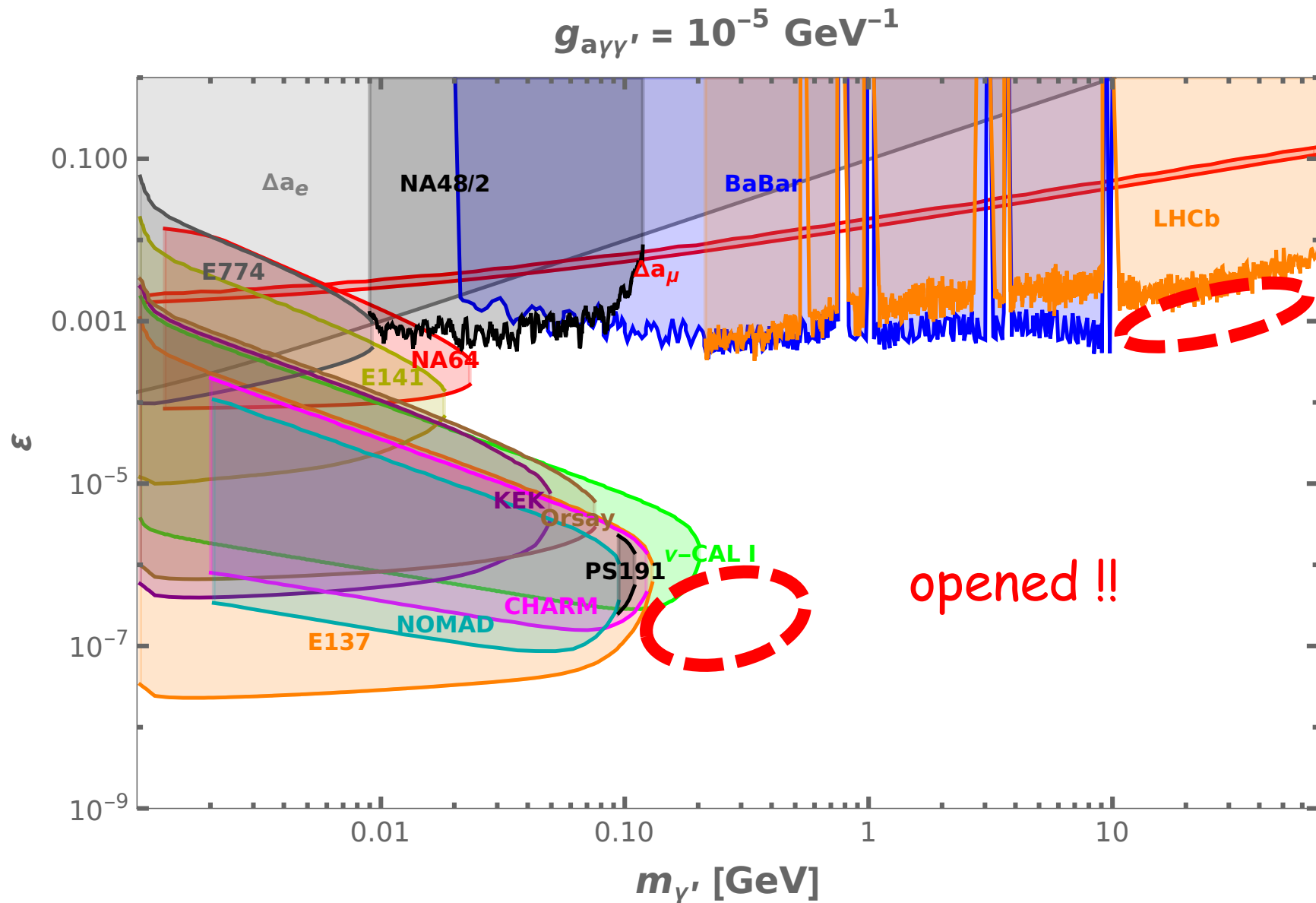
$$\begin{aligned} \ell^+ \ell^- \text{ events} &= \sigma \times \text{Br}(\gamma' \rightarrow \ell^+ \ell^-) \\ &\sim (\text{old}) \times \left(1 + \left(\frac{g_{a\gamma\gamma'} m_{\gamma'}}{e\epsilon}\right)^2\right)^{-1} \end{aligned}$$

beam dump

$$\begin{aligned} e^+ e^- \text{ events} \\ \sim (\text{old}) \times \exp\left(-\frac{L' m_{\gamma'}^4}{p_{\gamma'}}\right) \times \left(1 + \left(\frac{g_{a\gamma\gamma'} m_{\gamma'}}{e\epsilon}\right)^2\right)^{-1} \end{aligned}$$



New Constraints

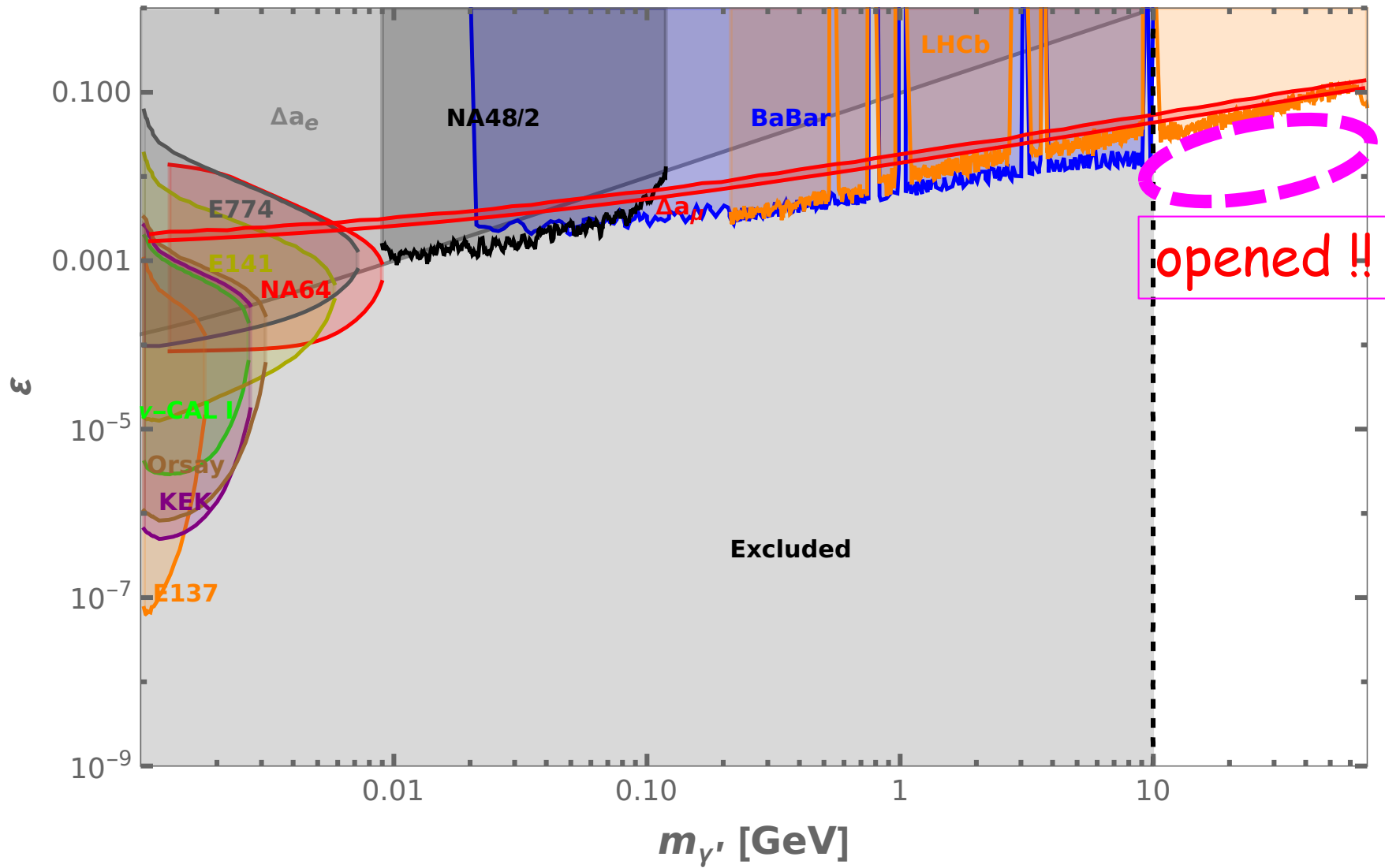


SN constraints are not included!



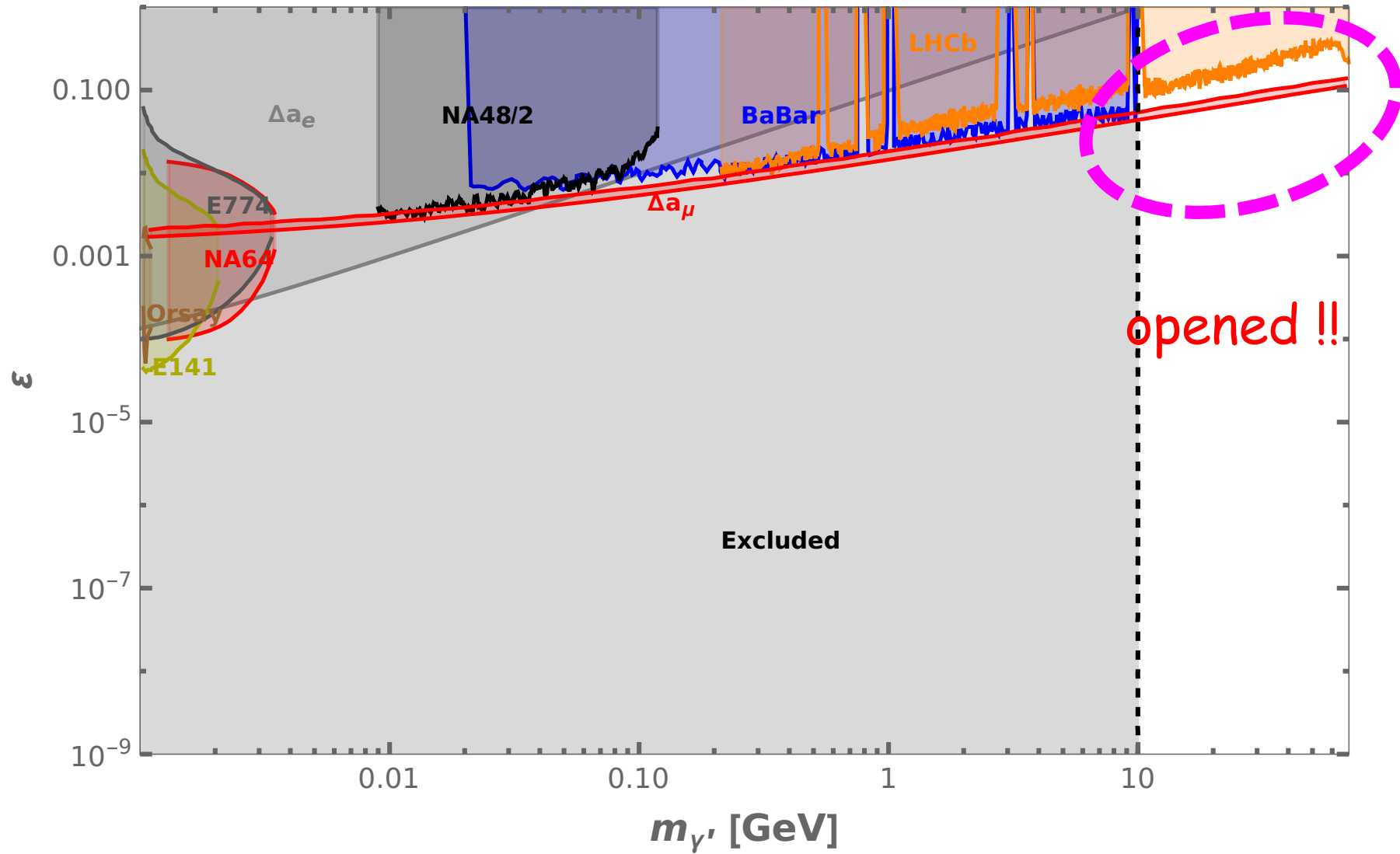
New Constraints

$$g_{a\gamma\gamma'} = 10^{-1} \text{ GeV}^{-1}$$



New Constraints

$$g_{\text{a}\gamma\gamma'} = 1 \text{ GeV}^{-1}$$



Summary

IF ALP lives in DS, it could couple to photon and dark photon !!

ALP could significantly change the constraints for dark photon (> 1 MeV)!

- ➔ may open or close parameter space
- ➔ especially for heavier dark photon
- ➔ may reopen the parameter to explanation muon $g-2$

Search strategy for dark photons should be revisited

- ➔ mono-photon with missing ET should be explored



Back UP



TABLE I. Specifications of the experimental setups we use in our analysis. The LSW experiments (GammeV, ALPS, ALPS II) are partly sensitive to the subfrequency new physics scenario. Our proposed experimental setup (waveguide) adopts a waveguide. The parameters for the LSW experiments including η_{eff} were taken from Refs. [7,30,31]. Some of the parameter values may not properly reflect the actual experiments. For instance, the η_{eff} of the LSW experiments would change if the frequency dependence is properly applied.

	a (mm)	L (m)	ω (eV)	N_γ (Hz)	N_{pass}	η_{eff}	N_d (Hz)	t_s (h)
Waveguide	...	1, 100	1.17	1.6×10^{20}	5000	0.54	10^{-6}	480
ALPS II	8.75 ^a	100	1.17	1.6×10^{20}	5000	0.95	10^{-6}	480
ALPS	7 ^b	7.6	2.33	2.6×10^{21}	1	0.9	0.0018	27
GammeV	25.5 ^c	7.2	2.33	6.6×10^{23}	1	0.25	130	24

^aThe ALPS II uses an optic suitable to collect a 17.5 mm diameter beam [32].

^bBecause we are unaware of the radius of the lens, we take the maximum of the vacuum tube size that can be inserted in the HERA dipole magnet [33].

^cWe take the size of the lens in front of the photomultiplier tube (PMT) [10].

TABLE I. Parameters of the experiments used in our calculation. The references for the primary FP cavity and laser setup of each experiment are presented in the table. In addition, we adopt the ALPS II cavity finesse to our proposed second cavity for GW experiments.

Parameters	ALPS II [14]	aLIGO [62]	KAGRA [63–65]	ET-HF [59–61]	ET-LF [59–61]
\mathcal{F}_{Cav}	7853	450	1550	880	880
P_{in} (W)	30	2600	412	5355	32
ω (eV)	1.165	1.165	1.165	1.165	0.799
L (m)	100	10	10	10	10
N_d (Hz)	10^{-6}	10^{-6}	10^{-6}	10^{-6}	10^{-6}
$\mathcal{F}_{\text{Cav}}^{\text{WG}}$...	7853	7853	7853	7853

