

# Axions beyond the QCD axion

Stefania Gori  
UC Santa Cruz



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# The QCD axion. Present and future

Strong CP problem:

why is the QCD  $\bar{\theta}$  parameter so small?  $\mathcal{L}_{\text{QCD}} \supset \theta \frac{g^2}{32\pi^2} G_{\mu\nu} \tilde{G}^{\mu\nu}$

$$\bar{\theta} \equiv \theta + \arg(\det(Y_u Y_d))$$

QCD axion: elegant way to address this problem.

Dynamical solution to achieve:  $\bar{\theta} \lesssim 10^{-10}$  in agreement with neutron EDM constraints

The QCD axion mass is set by its decay constant,  $f_a$ :  $m_a f_a \sim f_\pi m_\pi$

The generic expectation is that it couples  $\sim 1/f_a$

Peccei-Quinn symmetry  
breaking scale

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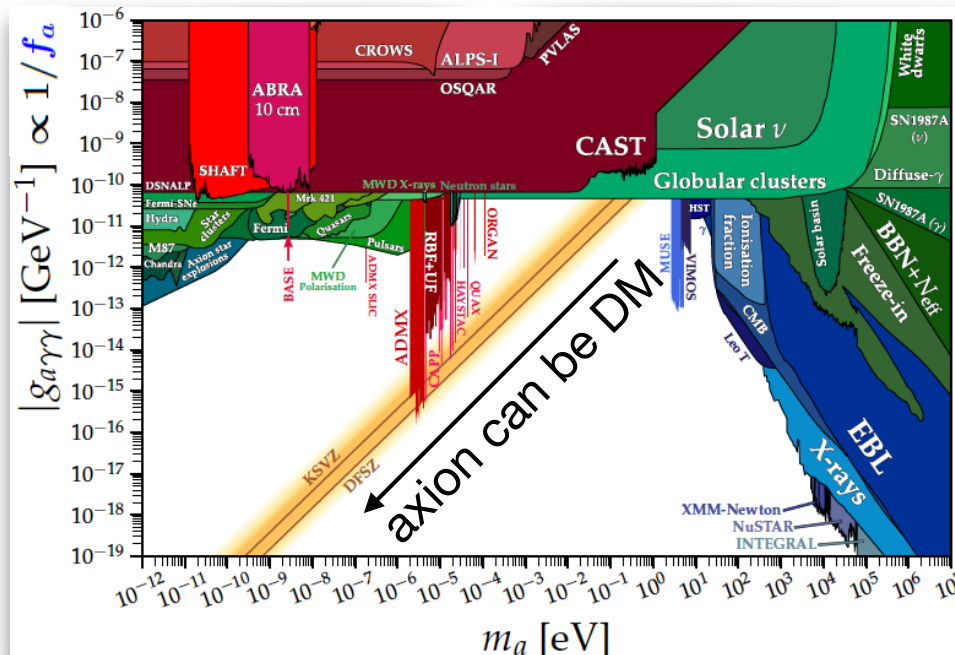
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Adams et al.,  
2203.14923



$$f_a \gtrsim \mathcal{O}(10^{11} \text{ GeV})$$

the QCD axion can easily be DM.  
It is more difficult if

$$\mathcal{O}(10^8 \text{ GeV}) \lesssim f_a \lesssim \mathcal{O}(10^{11} \text{ GeV})$$

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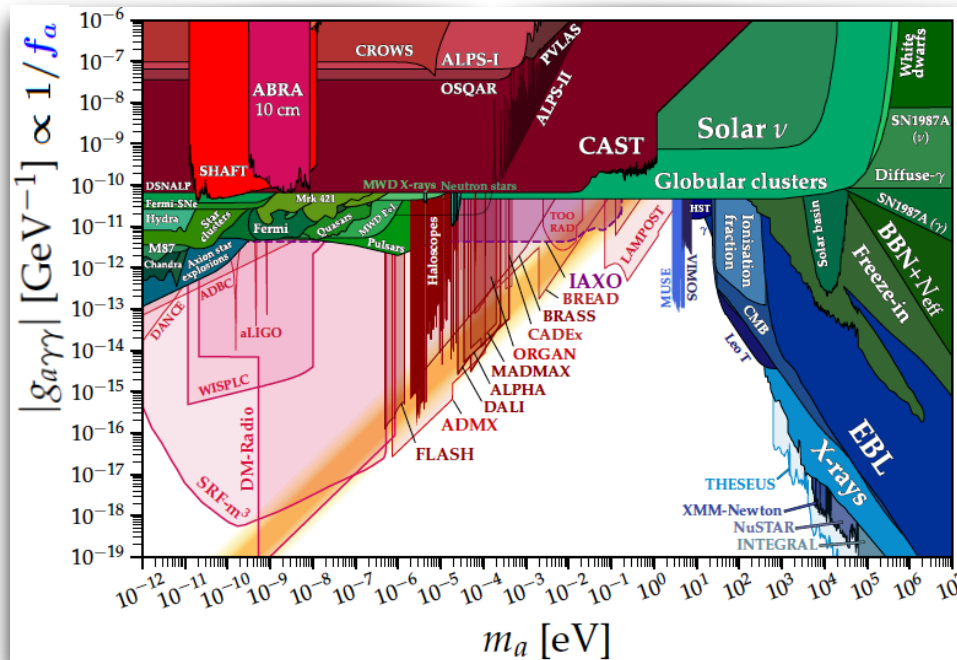
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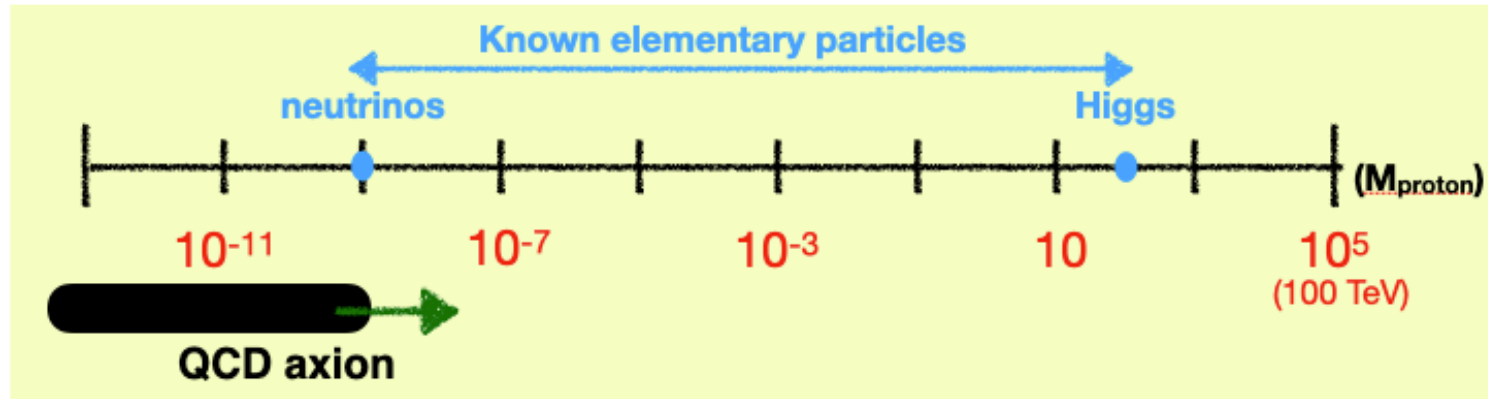
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# Heavier axions and the strong CP problem



Models where two or more axions naturally cooperate to address the strong CP problem.

(Recent references:

Agrawal, Howe, 1710.04213;

Foster, Kumar, Safdi, Soreq, 2208.10504, ...)

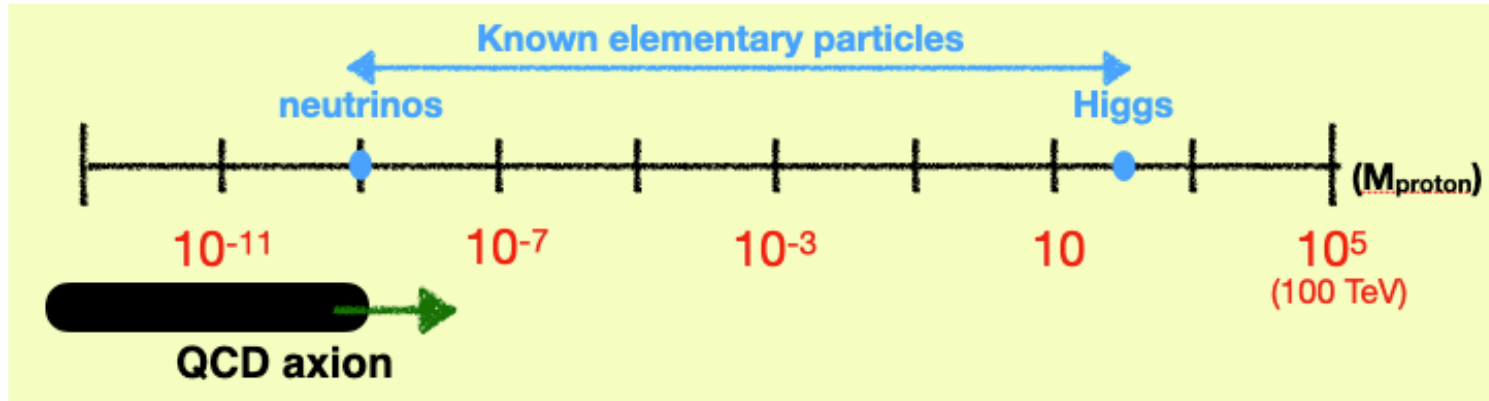
(2)

**Axion quality problem alleviated  
for heavy axions**

if PQ symmetry broken by  
dimension D operators at  $\Lambda_{\text{UV}}$ :

$$\delta\bar{\theta} \sim \frac{f_a^{D-2}}{m_a^2 \Lambda_{\text{UV}}^{D-4}}$$

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$$SU(3) \times SU(3) \rightarrow SU(3)_D = SU(3)_c$$

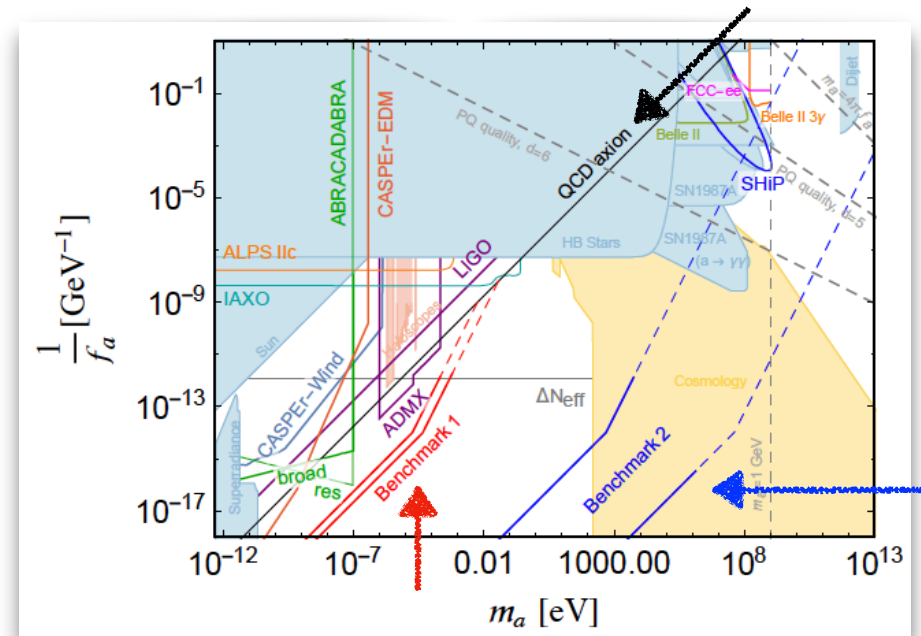
Lot of freedom in the  $(m_a - f_a)$  plane.

Also different hierarchies in axion-SM couplings.

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# For this talk

Variations of the minimal QCD axion:

**1)** The QCD axion as mediator between the SM and DM,  $\chi$ :  $\frac{c_\chi}{2f_a} \partial_\mu a \bar{\chi} \gamma^\mu \gamma_5 \chi$   
 $f_a \lesssim \mathcal{O}(10^{11} \text{GeV})$

Chapter 1.

**2)** Heavier axions and new phenomenology at meson factories

Chapter 2.

Main references:

1. Dror, SG, Munbodh, 2306.03145 (+ with Knapen, Lin, Munbodh, Suter, 2501.xxxxx)
2. Altmannshofer, Dror, SG, 2209.00665

# Chapter 1

## The QCD axion as mediator between the SM and DM

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# Motivations and model

- \* If  $f_a \lesssim \mathcal{O}(10^{11} \text{GeV})$ , typically axions do not constitute a sizable fraction of the DM energy density without additional dynamics beyond the misalignment mechanism ( $T_{\text{RH}} \lesssim f_a$ ) or the decay of cosmic defects ( $T_{\text{RH}} \gtrsim f_a$ )
- \* Experimental program to detect such low  $f_a$  axions: upcoming IAXO and ALPS II

Can we still have a connection to DM, even for such lower decay constants?

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**A minimal model:**

let's take the QCD axion (either KSVZ or DFSZ) and let's couple it to a singlet Dirac fermion DM candidate:

$$\mathcal{L} \supset \frac{c_\chi}{2f_a} \partial_\mu a \bar{\chi} \gamma^\mu \gamma_5 \chi$$

A small set of free parameters fixes the cosmology of the model:  $f_a, m_\chi, g_{a\chi} \equiv \frac{c_\chi m_\chi}{f_a}, T_{\text{RH}}$   
(eventually  $\tan\beta$  in DFSZ)

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Couplings	KSVZ	DFSZ
Gluons		$\frac{\alpha_s}{8\pi f_a}$
Photons	$-\frac{\alpha}{8\pi f_a} (1.924)$	$\frac{\alpha}{8\pi f_a} \left(\frac{8}{3} - 1.924\right)$
Quarks	Loop suppressed	up : $\frac{\cos^2 \beta}{6f_a}$ , down : $\frac{\sin^2 \beta}{6f_a}$
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Bounds from DM self-interaction:

$$c_\chi \frac{m_\chi}{f_a} \equiv g_{a\chi} \lesssim 0.21 \left( \frac{m_\chi}{1 \text{ MeV}} \right)^{3/4}$$

S.Gori  $f_a \gtrsim \begin{cases} 3.9 \times 10^8 \text{ GeV} & \text{(KSVZ)} \\ 1.2 \times 10^9 \text{ GeV} \sin^2 \beta & \text{(DFSZ-I)} \\ 1.2 \times 10^9 \text{ GeV} \cos^2 \beta & \text{(DFSZ-II)} \end{cases}$  supernova cooling bounds (axion-nucleon interaction)  
cooling bounds on red giant (axion-electron interaction)

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A small set of free

Couplings	KS
Gluons	
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Quarks	Loop su
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The well - motivated region of QCD axion parameter space is bounded:

$$\mathcal{O}(10^8 \text{ GeV}) \lesssim f_a \lesssim \mathcal{O}(10^{11} \text{ GeV})$$

astrophysical bounds

subdominant DM component

Type I :  $\frac{\sin^2 \beta}{6f_a}$ , Type II :  $-\frac{\cos^2 \beta}{6f_a}$

$$-\partial_\mu a \bar{\chi} \gamma^\mu \gamma_5 \chi$$

$$n_\chi, g_{a\chi} \equiv \frac{c_\chi m_\chi}{f_a}, T_{\text{RH}}$$

eventually  $\tan \beta$  in DFSZ)

DM self-interaction:

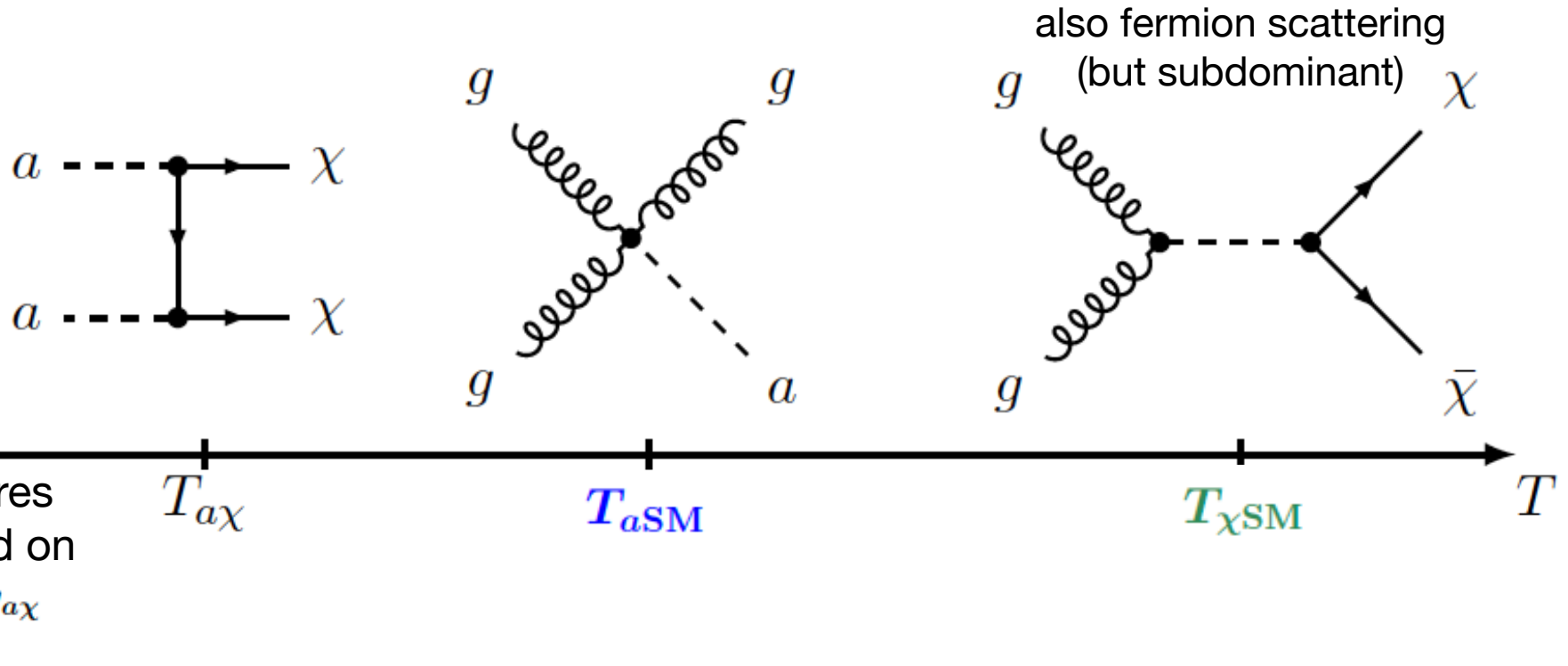
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supernova cooling bounds (axion-nucleon interaction)  
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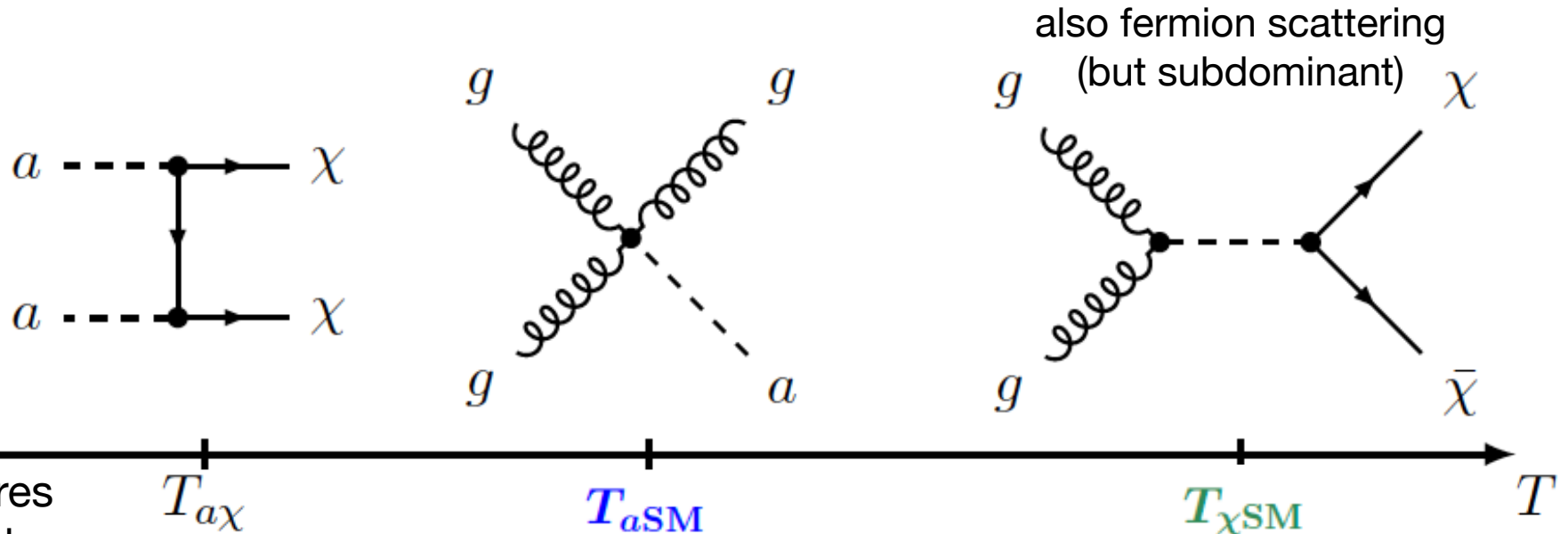
# Thermalizations

The thermal history depends on three classes of processes:



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temperatures only depend on

$f_a, m_\chi, g_{a\chi}$   
(mild)

$$T_{aSM} \simeq 2 \times 10^4 \text{ GeV} \left( \frac{f_a}{10^9 \text{ GeV}} \right)^2 \rightarrow T_{\chi SM} \gg T_{aSM}$$

$$T_{\chi SM} \simeq 4 \times 10^7 \text{ GeV} \left( \frac{f_a}{10^9 \text{ GeV}} \right)^2 \left( \frac{1}{g_{a\chi}} \right)^2$$

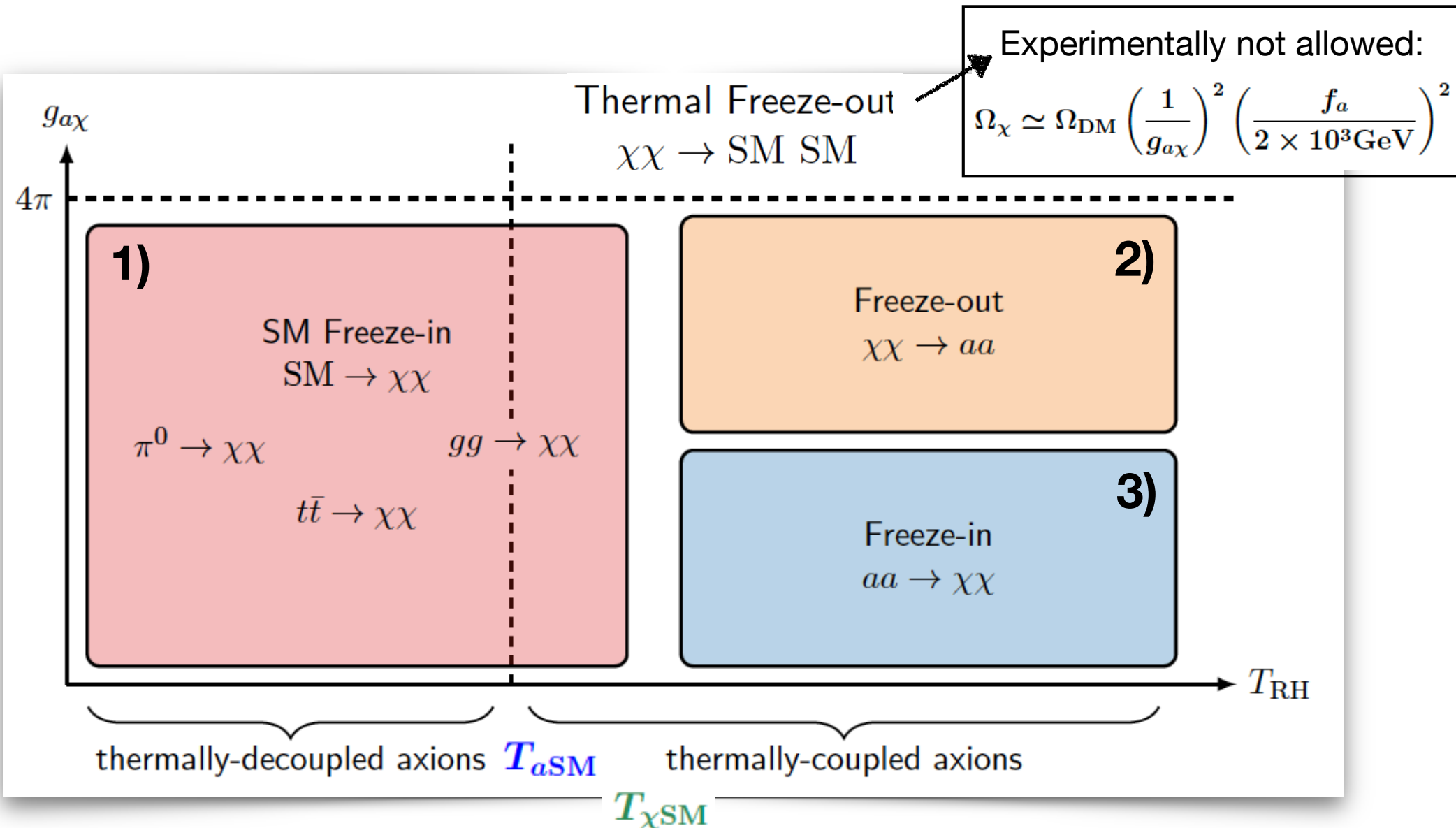
3 possible hierarchies:

(this is shown in the figure above)

- $T_{a\chi} \ll T_{aSM} \ll T_{\chi SM}$
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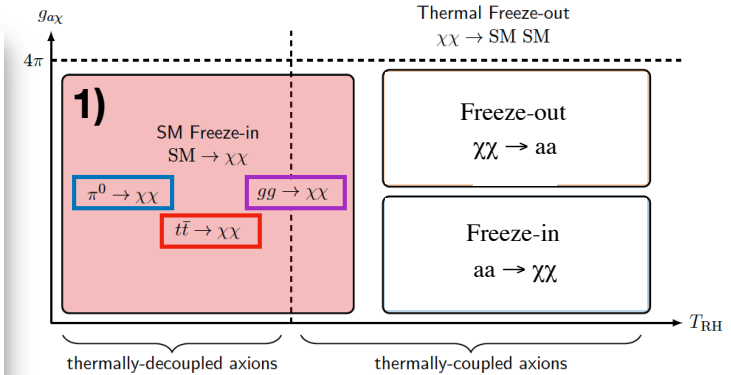
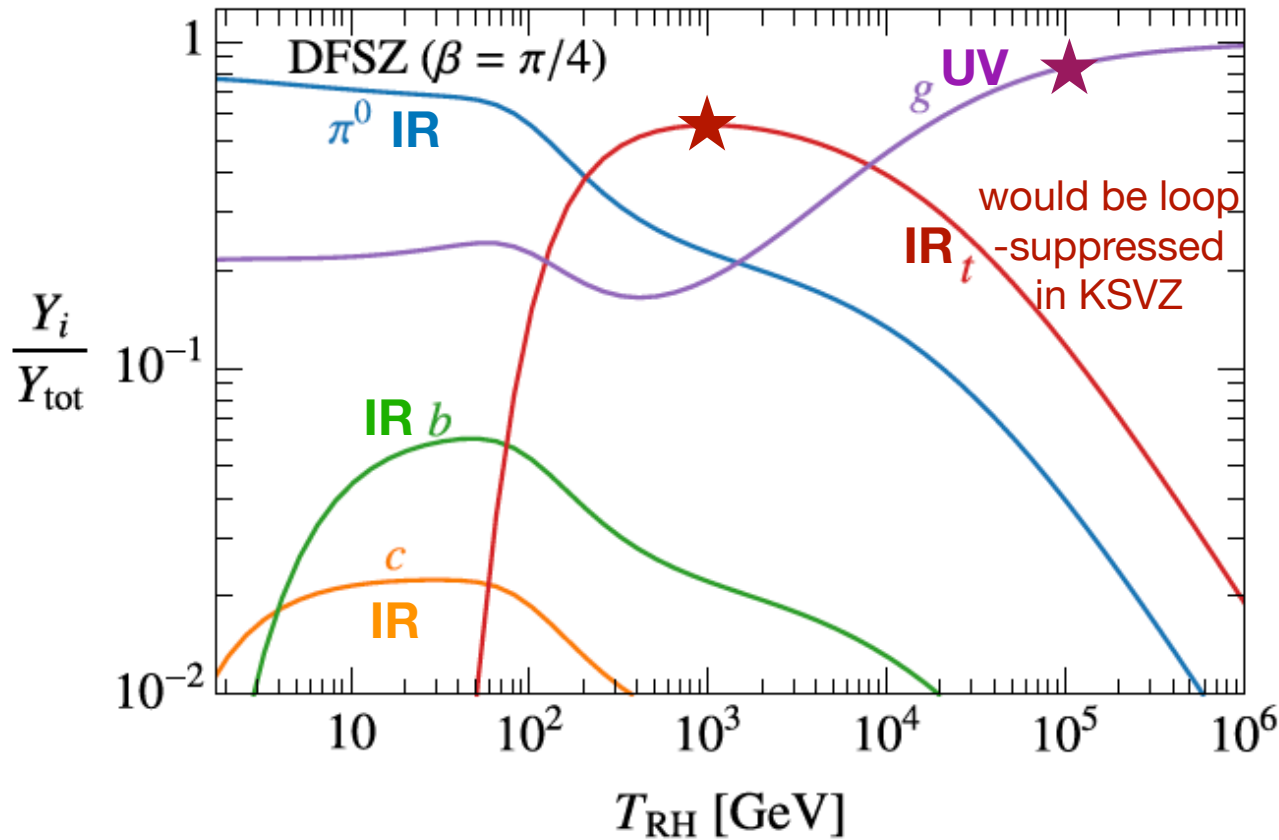
see also Salvio et al., 1310.6982

# A bird's-eye view





# 1) Dark Matter from SM freeze-in



Independent on  
 $f_a, g_{a\chi}$

To avoid DM-SM thermalization:

$$T_{RH} < T_{\chi SM} \simeq 4 \times 10^7 \text{ GeV} \left( \frac{f_a}{10^9 \text{ GeV}} \right)^2 \left( \frac{1}{g_{a\chi}} \right)^2$$

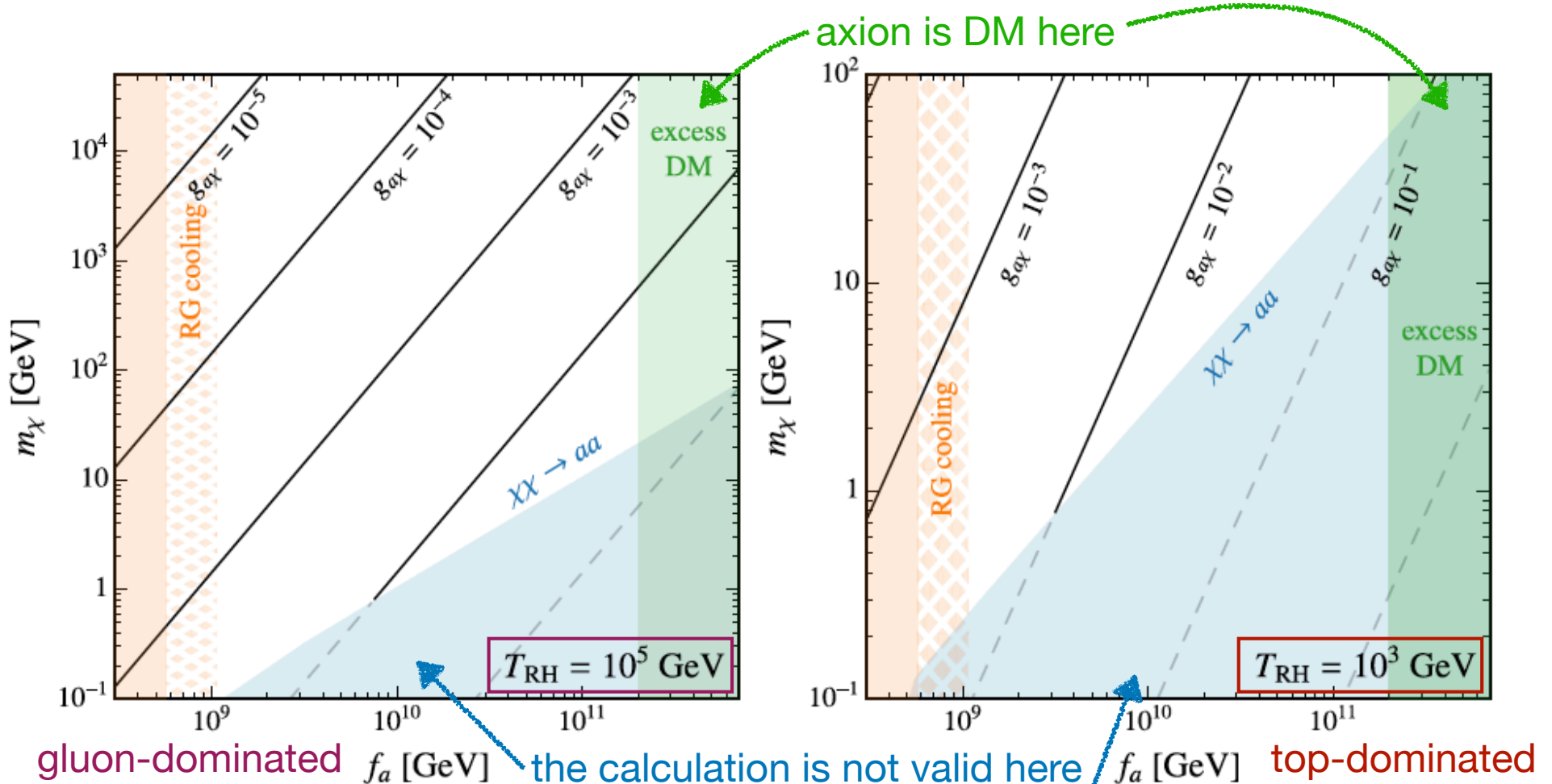
Simplifying assumption:

$$T_{RH} < T_{aSM} \simeq 2 \times 10^4 \text{ GeV} \left( \frac{f_a}{10^9 \text{ GeV}} \right)^2$$

such that the axion does not have  
a thermal abundance

# 1) SM freeze-in: the relic abundance

$g_{a\chi}$  needed to get the measured relic abundance



gluon-dominated

$f_a$  [GeV]

the calculation is not valid here since DM can freeze-out to axions

$f_a$  [GeV]

top-dominated

typically not natural regime since  $g_{a\chi} \equiv \frac{c_\chi m_\chi}{f_a}$

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$$n_\chi \sigma_{\chi\bar{\chi} \rightarrow aa} \gtrsim H$$

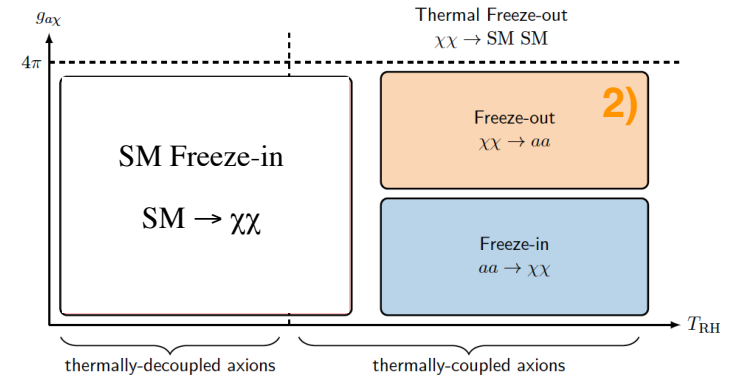
Decoupled freeze-out regime in Bharucha et al., 2209.03932

# 2), 3) Dark Matter from axion freeze-out or freeze-in

No coupling with the SM is necessary



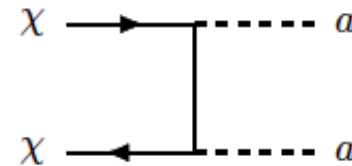
difficult to test experimentally



$$T_{RH} \gtrsim T_{\chi SM} \quad \text{secluded freeze-out}$$

2) DM in thermal contact with the SM in the early universe.

$$\frac{\Omega_\chi}{\Omega_{DM}} \sim \left( \frac{m_\chi}{1\text{GeV}} \right)^2 \left( \frac{4.4 \times 10^{-2}}{g_{a\chi}} \right)^4 \quad \text{unnatural regime}$$



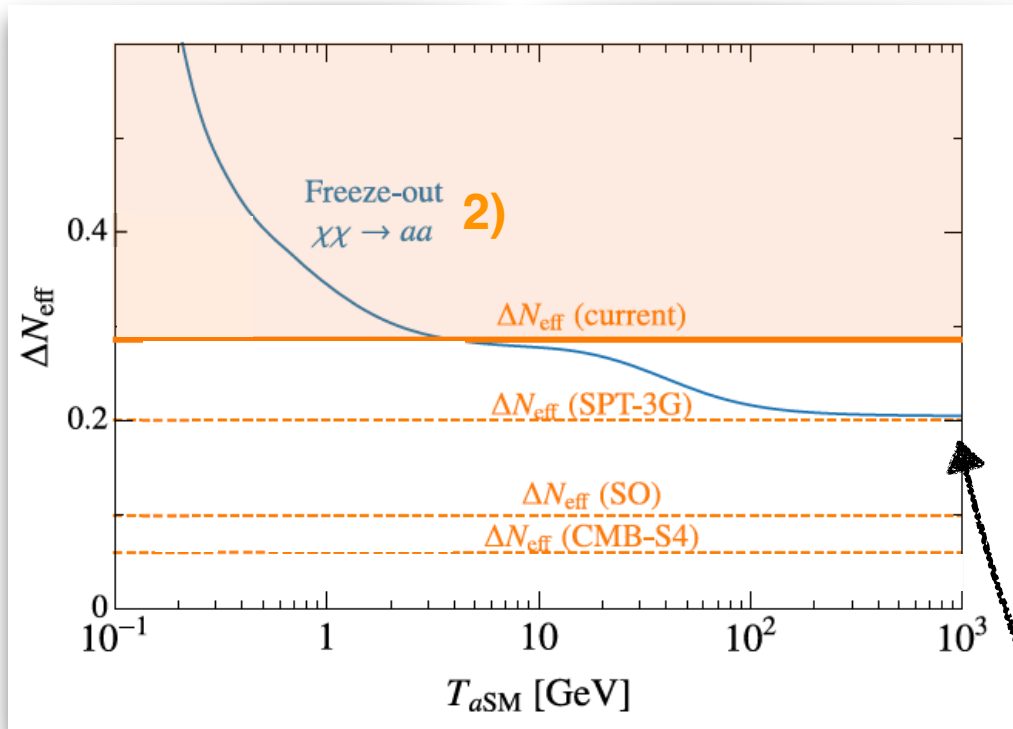
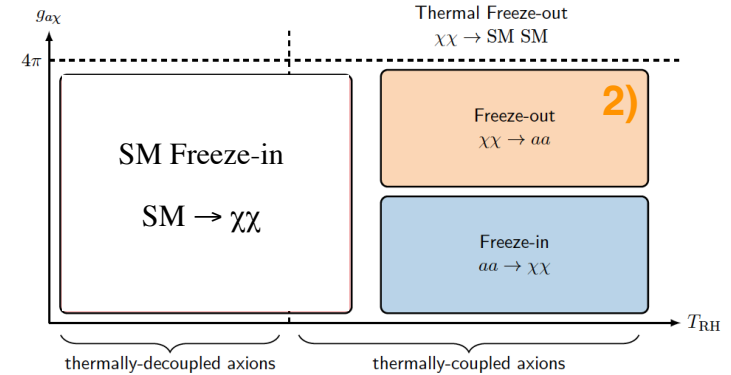
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DM annihilates into QCD axions, which remain relativistic today

➔ significant source of dark radiation.

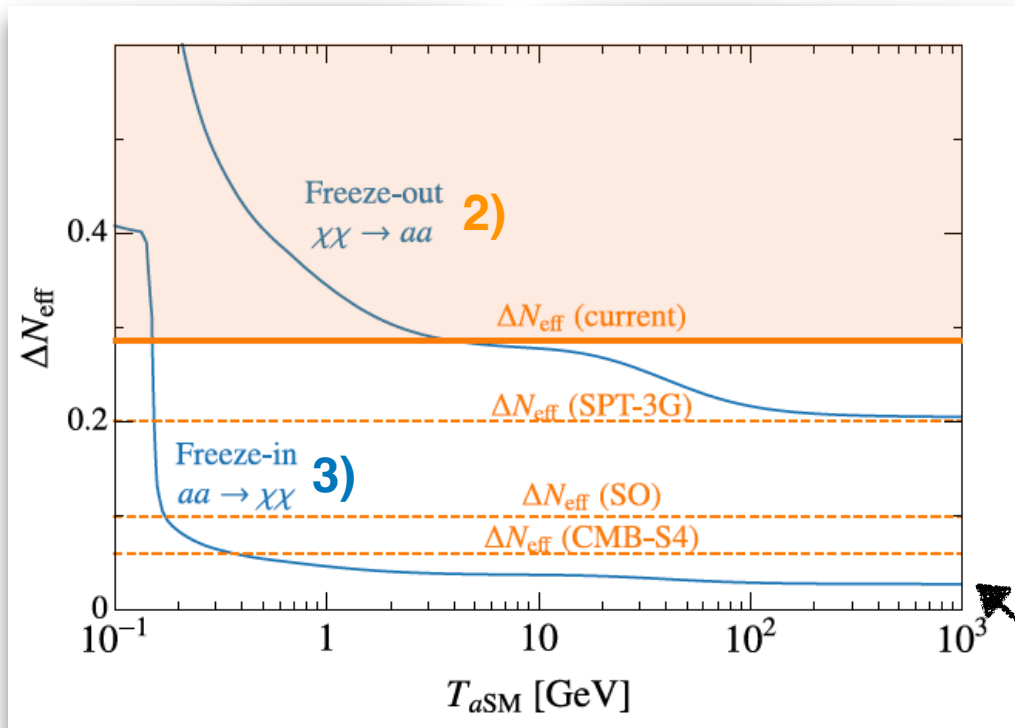
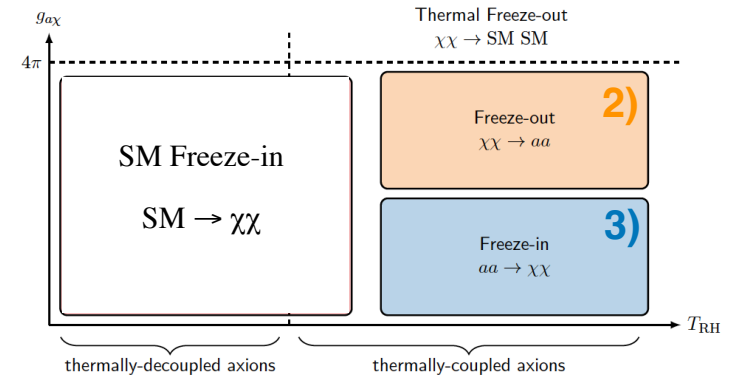
Future experiments will be able to completely probe the freeze-out scenario

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difficult to test experimentally



## secluded freeze-in

$$T_{aSM} \lesssim T_{RH} \lesssim T_{\chi SM} \text{ and } T_{RH} \gtrsim T_{ax}$$

3) axion is thermalized in the early universe

$$\frac{\Omega_\chi}{\Omega_{DM}} \sim \left( \frac{g_{a\chi}}{3 \times 10^{-6}} \right)^4 \quad \text{natural regime}$$

The QCD axions will contribute to  $N_{\text{eff}}$ , but now the DM does not have a sizable energy density  $\rightarrow$  the evolution of the dark sector temperature will be different than in the freeze-out case

The freeze-in scenario is pretty hidden, even to future experiments

# Axion mediators beyond the QCD axion

The model we discussed so far is compelling since it is highly predictive and some scenario can be probed in the future. However, **no laboratory-based probes**

Reason: large  $f_a$  that makes the coupling to the SM very small

**What if we extend this framework to much heavier axions/axion-like-particles?**

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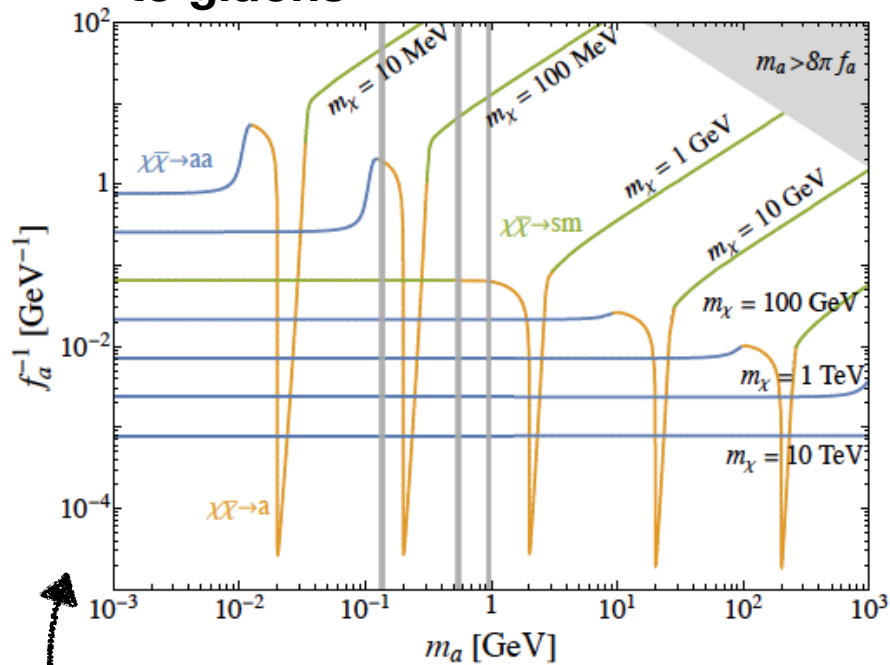
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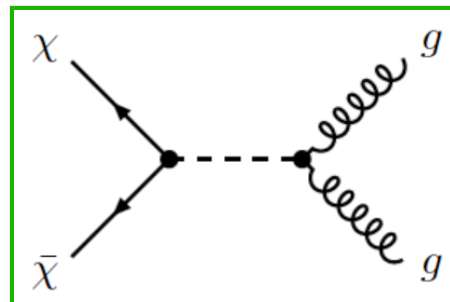
**ALP coupled to gluons**

Freeze Out

$$m_a f_a \neq m_\pi f_\pi$$



SM freeze-out is now a viable model



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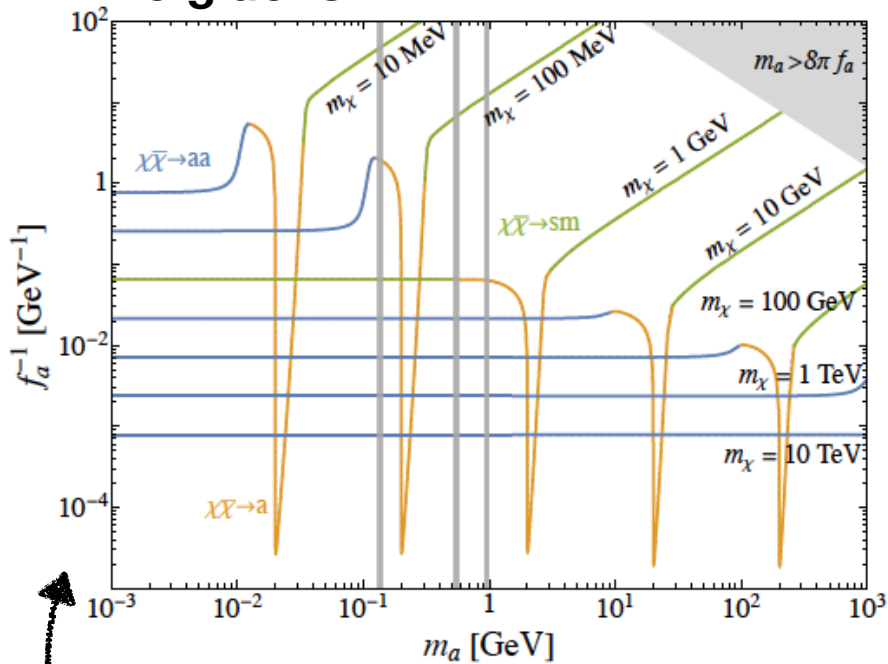
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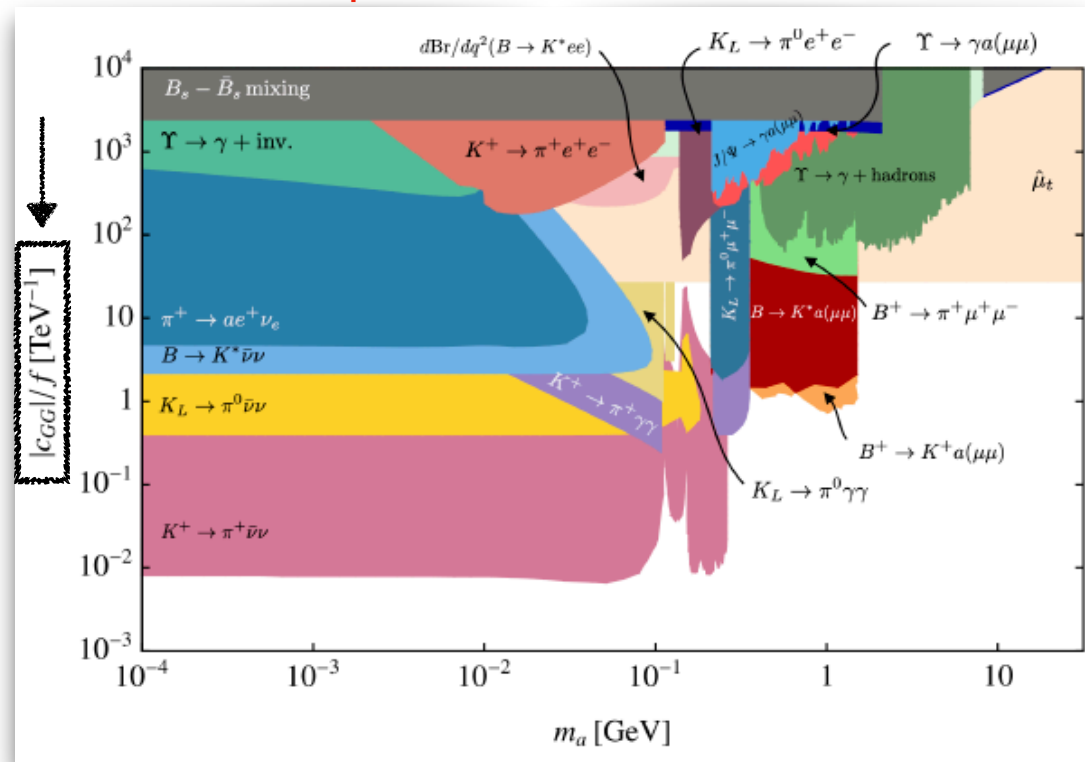
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## ALPs produced at accelerators



Bauer et al., 2110.10698



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Spin-dependent scattering with nuclei  
at direct detection experiments

with Knapen, Lin, Munbodh, Suter, 2501.xxxxx

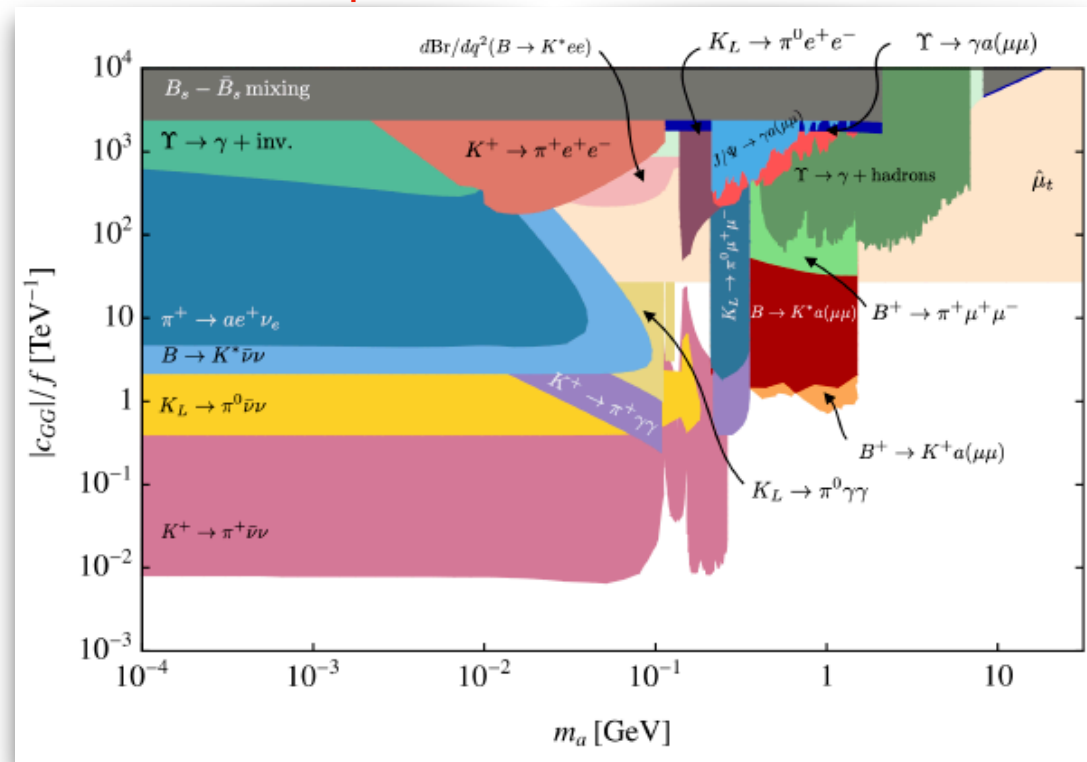
$$\mathcal{H} = -\frac{g_{a\chi}g_n}{m_n m_\chi} \frac{(\bar{q} \cdot \bar{S}_\chi)(\bar{q} \cdot \bar{S}_n)}{\bar{q}^2 + m_a^2} e^{i\bar{q}\bar{r}}$$

$g_n$  is the axion coupling to nuclei:

$$\mathcal{L}_{an} = g_n a \bar{n} \gamma^5 n$$

Computing rates for different materials as a function of the threshold energy, while being consistent with flavor constraints and DM self-interaction constraints

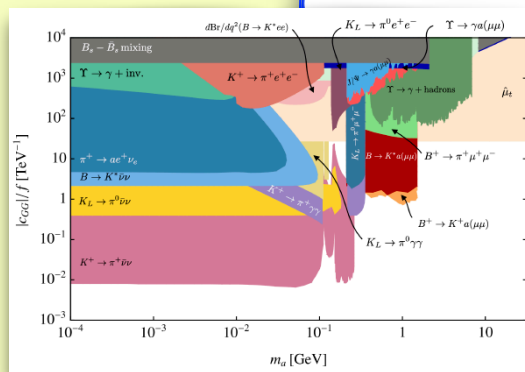
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Bauer et al., 2110.10698

# Chapter 2

## Heavier axions and new phenomenology at meson factories



any new search?

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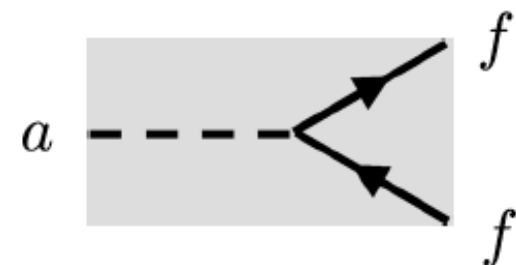
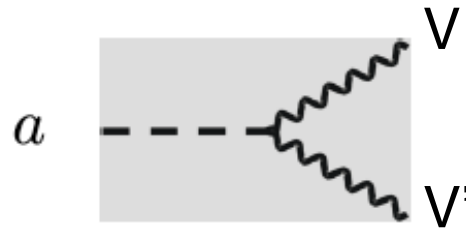
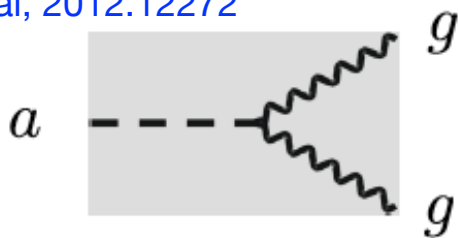
# Let's go back to the EFT for axions

At dimension 5, the most general Lagrangian for a spin 0, CP-odd particle with an approximate shift symmetry,  $a \rightarrow a+c$ :

Georgi, Kaplan, Randall 1986

$$\mathcal{L} \supset -\frac{g_{ag}}{4} a G_{\mu\nu}^a \tilde{G}^{a\mu\nu} - \frac{g_{aW}}{4} a W_{\mu\nu}^a \tilde{W}^{a\mu\nu} - \frac{g_{aB}}{4} a B_{\mu\nu} \tilde{B}^{\mu\nu} + ig_{af} (\partial_\mu a) (\bar{f} \gamma^\mu \gamma_5 f)$$

For the complete one-loop analysis, see  
 Bonilla et al, 2107.11392  
 Bauer et al, 2012.12272



Fast progressing number of theory studies +  
 experimental searches

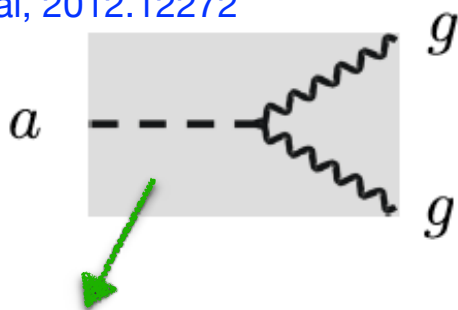
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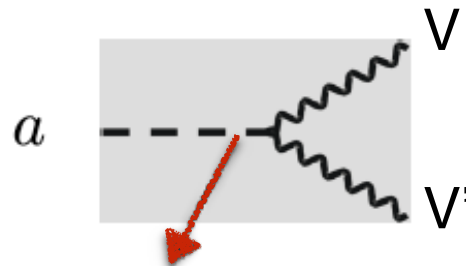
Georgi, Kaplan, Randall 1986

$$\mathcal{L} \supset -\frac{g_{ag}}{4} a G_{\mu\nu}^a \tilde{G}^{a\mu\nu} - \frac{g_{aW}}{4} a W_{\mu\nu}^a \tilde{W}^{a\mu\nu} - \frac{g_{aB}}{4} a B_{\mu\nu} \tilde{B}^{\mu\nu} + ig_{af} (\partial_\mu a) (\bar{f} \gamma^\mu \gamma_5 f)$$

For the complete one-loop analysis, see  
Bonilla et al, 2107.11392  
Bauer et al, 2012.12272



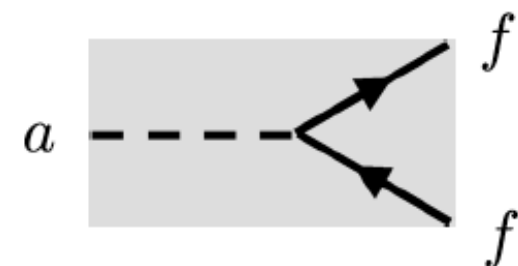
Minimal coupling expected if connection to the strong CP problem.



A axion-photon coupling is generated in the broken phase

$$g_{aB} \cos^2 \theta + g_{aW} \sin^2 \theta$$

This is the **main coupling** that has been considered for phenomenological studies of axions in the sub-GeV scale.



Fast progressing number of theory studies + experimental searches

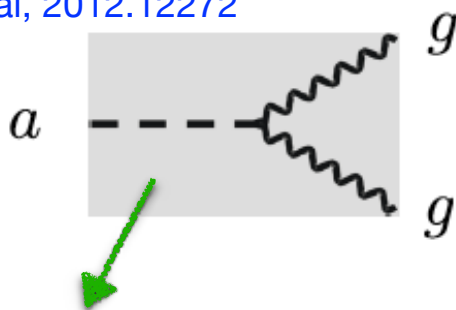
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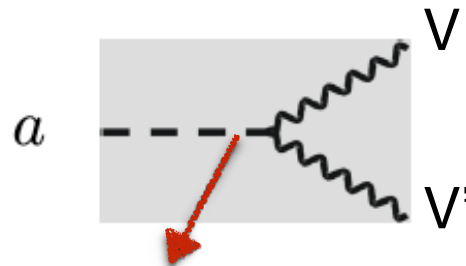
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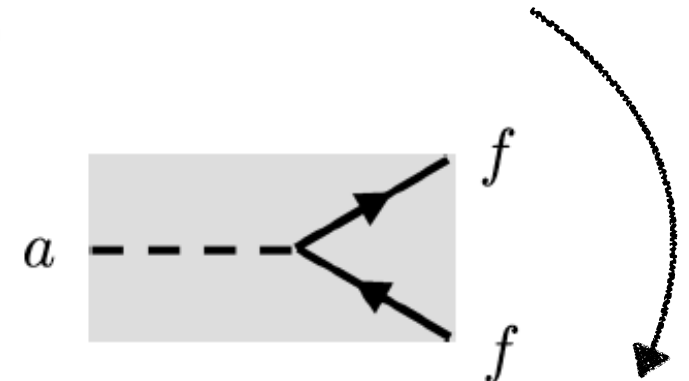
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These are **the least studied couplings**. Nevertheless they are present and sizable even in the minimal DFSZ QCD axion model.

(indeed the most stringent constraint on the DFSZ axion comes from its coupling to electrons (red giant))

**Fast progressing number of theory studies + experimental searches**

# Axion coupling to leptons

$$\frac{(\partial_\mu a)}{m_e} [\bar{e} \gamma^\mu (\bar{g}_{ee} + g_{ee} \gamma_5) e + g_\nu \bar{\nu} \gamma^\mu P_L \nu]$$

Altmannshofer, Dror,  
SG, 2209.00665

# Axion coupling to leptons

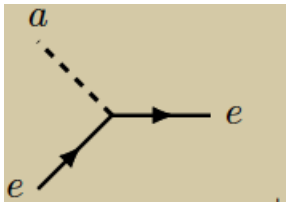
Altmannshofer, Dror,  
SG, 2209.00665

$$\frac{(\partial_\mu a)}{m_e} [\bar{e}\gamma^\mu (\bar{g}_{ee} + g_{ee}\gamma_5) e + g_\nu \bar{\nu}\gamma^\mu P_L \nu]$$

$$\mathcal{L} = -a \partial_\mu j_{PQ}^\mu$$

$$\partial_\mu j_{PQ}^\mu = g_{\ell\ell} (\bar{\ell} i \gamma_5 \ell)$$

“Standard”  
vertex



$$+ \frac{e^2}{16\pi^2 m_\ell} \frac{\bar{g}_{\ell\ell} - g_{\ell\ell} + g_{\nu\ell}}{4s_W^2} W_{\mu\nu}^+ \tilde{W}^{-,\mu\nu}$$

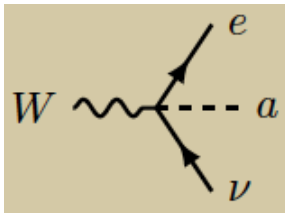
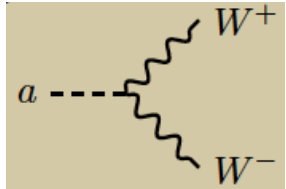
$$+ \frac{e^2}{16\pi^2 m_\ell} \frac{\bar{g}_{\ell\ell} - g_{\ell\ell}(1 - 4s_W^2)}{2c_W s_W} F_{\mu\nu} \tilde{Z}^{\mu\nu} - g_{\ell\ell} F_{\mu\nu} \tilde{F}^{\mu\nu} +$$

$$+ \frac{e^2}{16\pi^2 m_\ell} \frac{\bar{g}_{\ell\ell}(1 - 4s_W^2) - g_{\ell\ell}(1 - 4s_W^2 + 8s_W^4) + g_\nu}{8s_W^2 c_W^2} Z_{\mu\nu} \tilde{Z}^{\mu\nu}$$

$$+ \frac{ig}{2\sqrt{2}m_\ell} (g_{\ell\ell} - \bar{g}_{\ell\ell} + g_{\nu\ell}) (\bar{\ell} \gamma^\mu P_L \nu) W_\mu^-$$

Weak  
vertex

Anomaly  
terms



only present if SU(2) (or weak) violation:

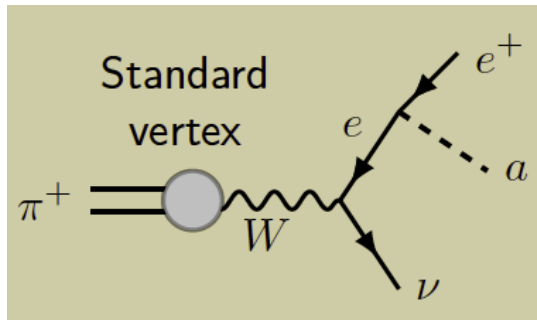
$$\bar{g}_{ee} - g_{ee} - g_\nu \neq 0$$

# Rate for charged current production

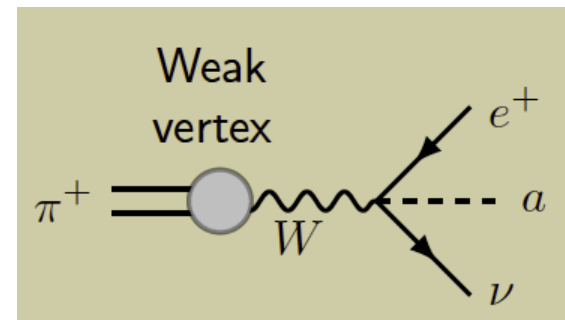
$$\text{BR}(\pi^+ \rightarrow e^+ a \nu) = \frac{1}{384\pi^2} \frac{m_\pi^4}{m_e^2 m_\mu^2} \left(1 - \frac{m_\mu^2}{m_\pi^2}\right)^{-2} \left[ \overbrace{(g_{ee} - \bar{g}_{ee} + g_\nu)^2 f_0 \left(\frac{m_a^2}{m_\pi^2}\right)}^{\neq 0 \text{ only if weak SU(2) violation}} \right. \\
 \left. + \frac{4m_e^2}{m_\pi^2} \left( 3(g_{ee})^2 f_3 \left(\frac{m_a^2}{m_\pi^2}\right) + 3(\bar{g}_{ee} - g_\nu)^2 f_4 \left(\frac{m_a^2}{m_\pi^2}\right) \right. \right. \\
 \left. \left. + 2g_{ee}(\bar{g}_{ee} - g_\nu) f_5 \left(\frac{m_a^2}{m_\pi^2}\right) \right) + \mathcal{O}\left(\frac{m_e^3}{m_\pi^3}\right) \right]$$

Helicity suppression is lifted only in the case of **weak SU(2) violation**

Weak preserving models



Weak violating models



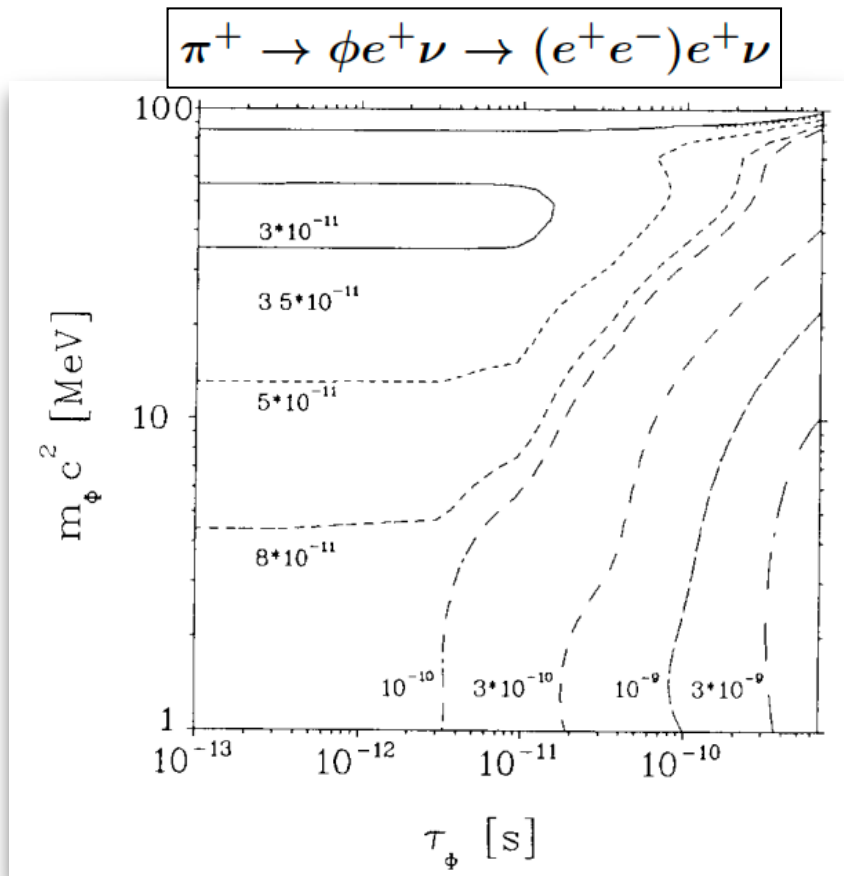
Similar results for  $K^+ \rightarrow ae^+\nu$ ,  $D_s^+ \rightarrow ae^+\nu$ ,  $B^+ \rightarrow ae^+\nu$



# The past search for $\pi^+ \rightarrow e^+ \nu (a \rightarrow e^+ e^-)$

In the late '80s,  
the SINDRUM experiment at PSI:

**Almost background free search**

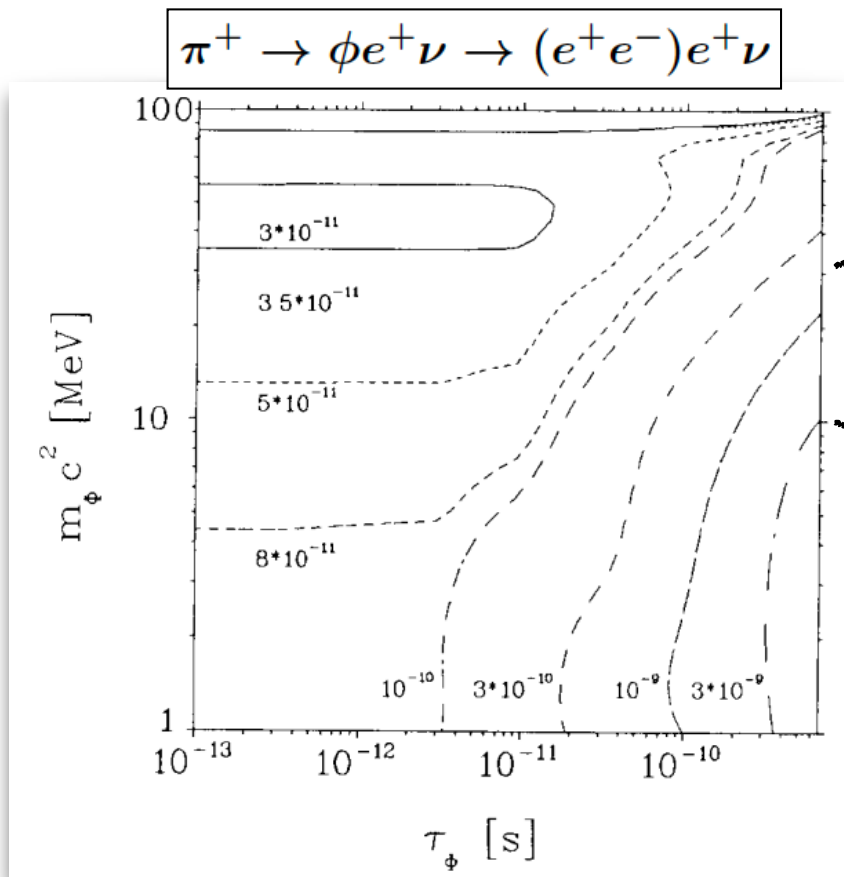


Eichler et al. Physics Letters B 175 (1986), no. 1 101–104

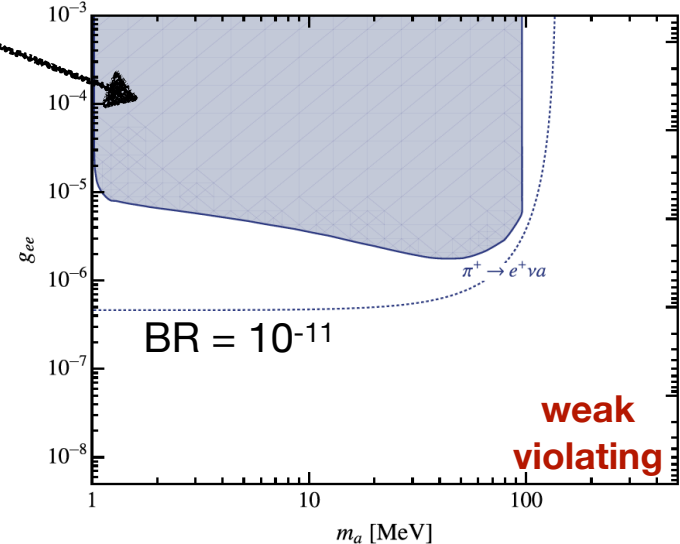
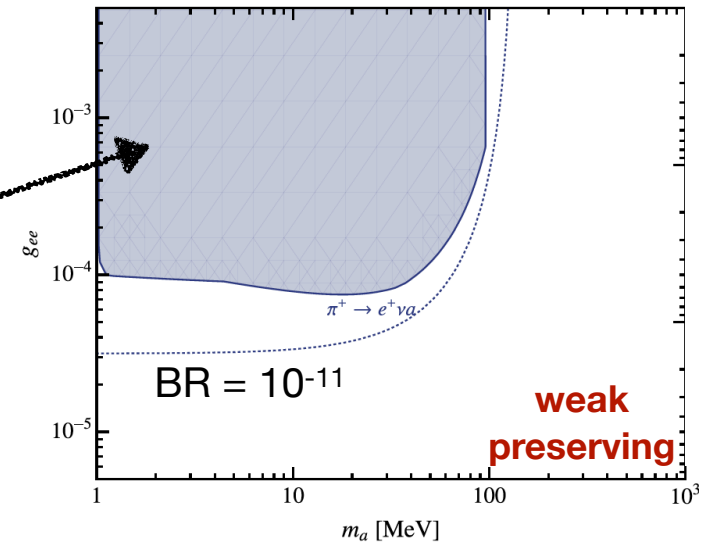
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$$f_a \sim m_e / g_{ee}$$

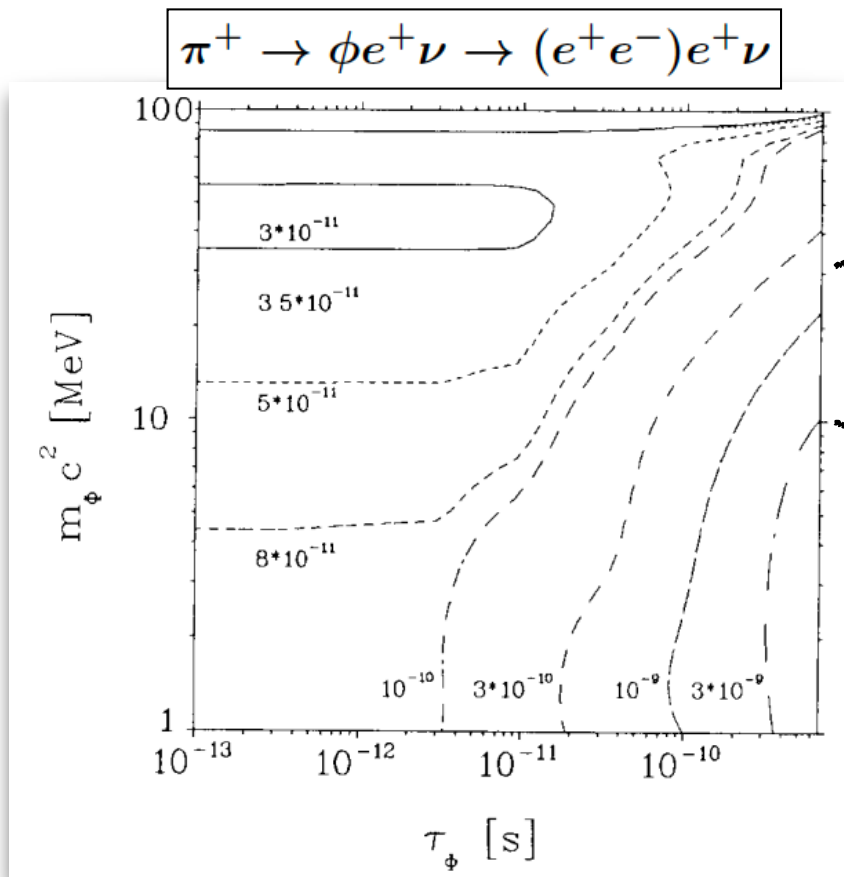


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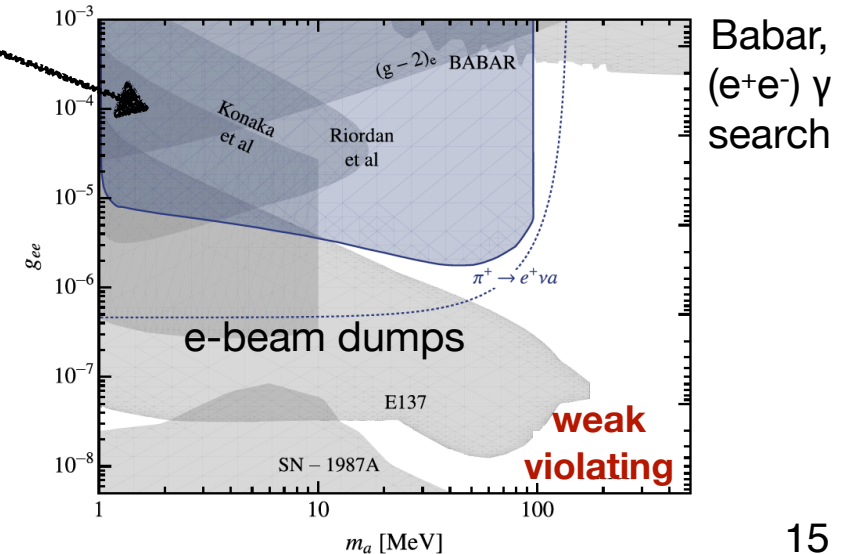
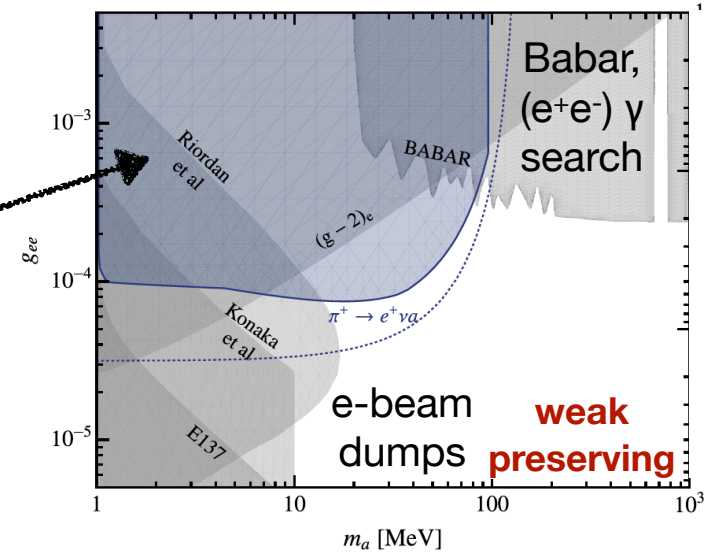
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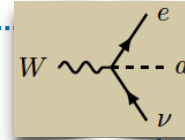
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$$f_a \sim m_e / g_{ee}$$



# Other charged current decays

## Kaon decays

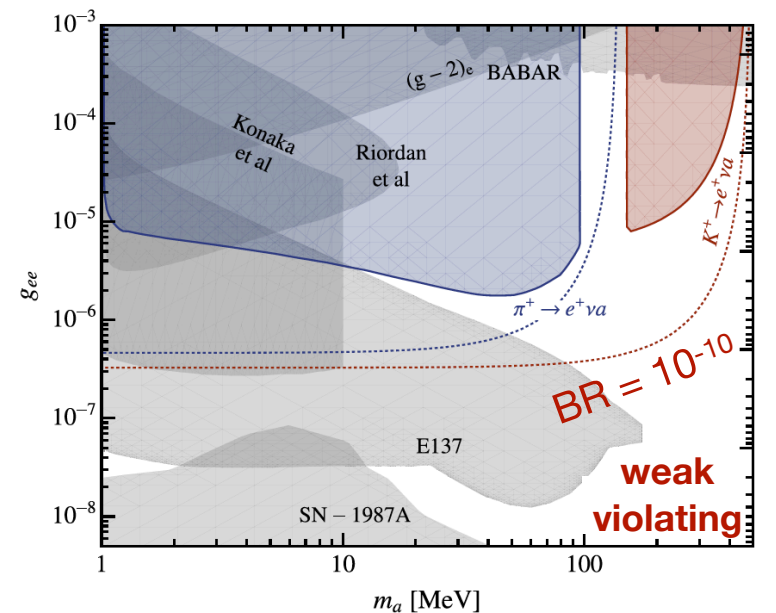
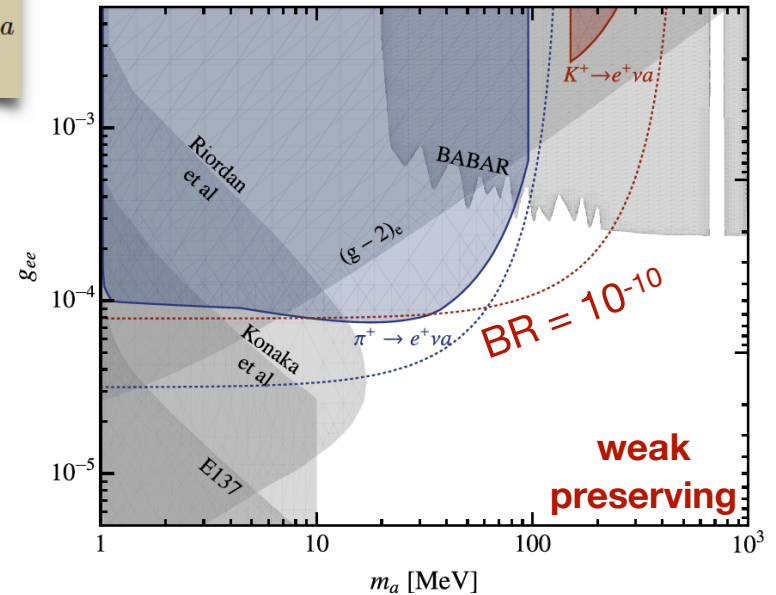


$$\frac{\text{BR}(K^+ \rightarrow e^+ \nu a)}{10^{-10}} \simeq \begin{cases} \left( \frac{g_{ee}}{3.3 \cdot 10^{-7}} \right)^2 & \text{weak violating} \\ \left( \frac{g_{ee}}{7.9 \cdot 10^{-5}} \right)^2 & \text{weak preserving} \end{cases}$$

hep-ex/0204006, E865 at BNL:

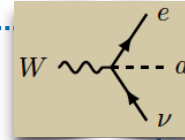
BR measured for  $m_{ee} > 150$  MeV

No dedicated search



# Other charged current decays

## Kaon decays

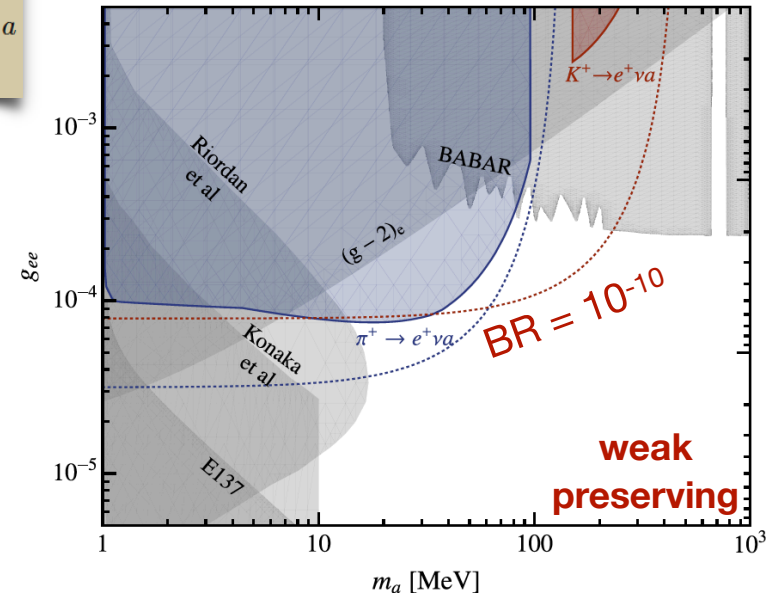


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## W boson decays

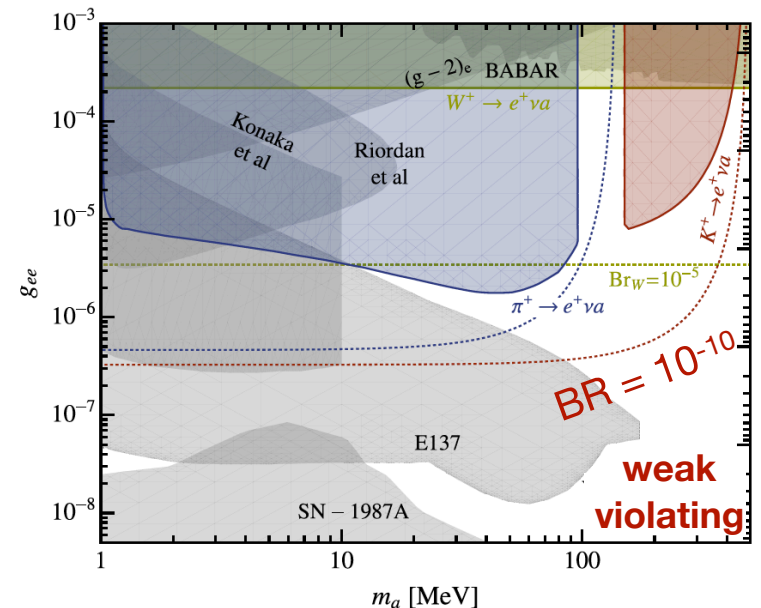
$$\frac{\text{BR}(W^+ \rightarrow \ell^+ \nu_\ell a)}{\text{BR}(W^+ \rightarrow e^+ \nu)} = \frac{3}{1024\pi^2} \frac{m_W^2}{m_\ell^2} (g_{\ell\ell} - \bar{g}_{\ell\ell} + g_{\nu\ell})^2$$

$\text{BR}(W^+ \rightarrow e^+ \nu_e a) \simeq \left( \frac{g_{ee}}{10^{-3}} \right)^2$  (only for weak violating)

No dedicated search

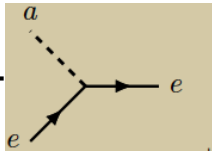
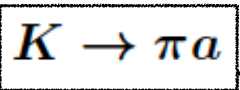
We impose:

the new width  $<$  exp uncertainty on W width

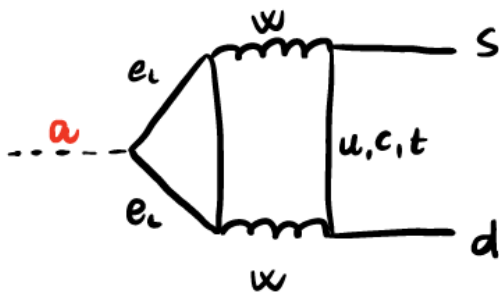


# Neutral current decays

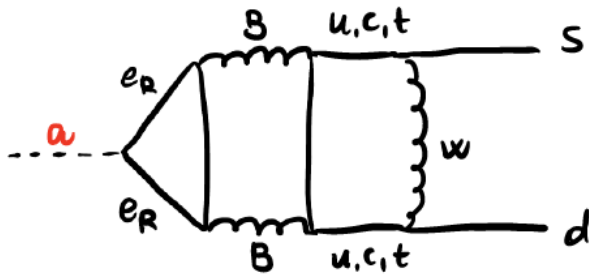
Neutral current meson decays are also generated at the 2 or 3-loop level (suppressed by CKM elements as well)



1. Axion with LH coupling



2. Axion with RH coupling



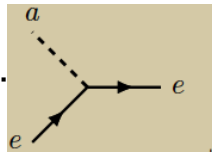
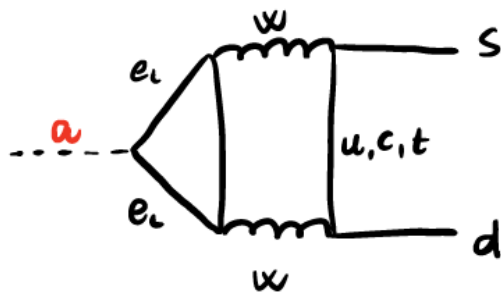
Similar diagrams for  $B \rightarrow K^{(*)} a$

# Neutral current decays

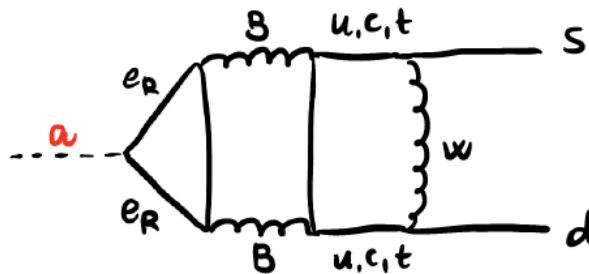
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$$K \rightarrow \pi a$$

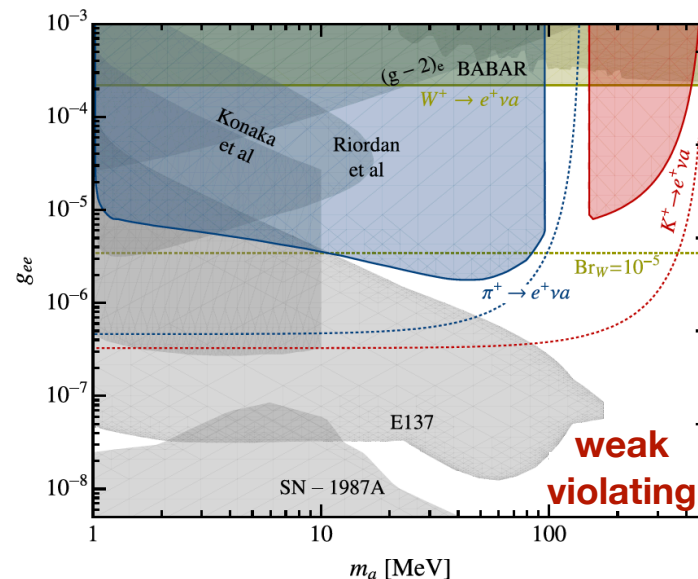
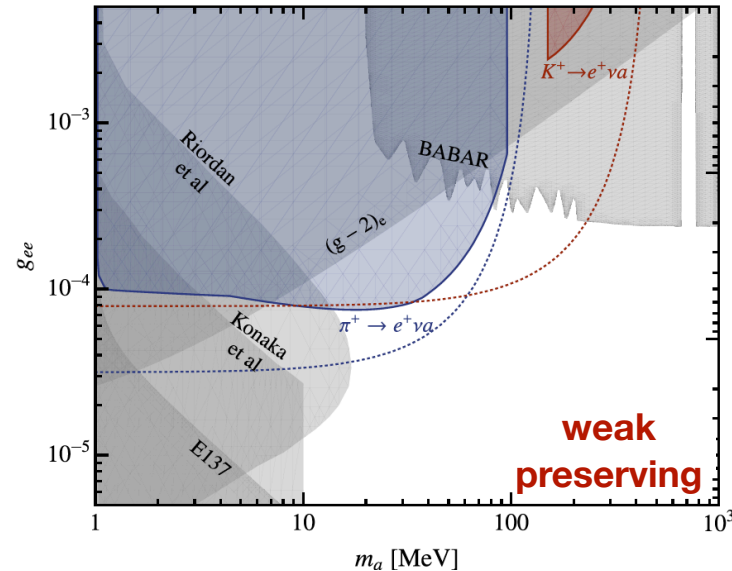
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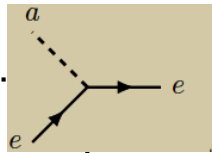
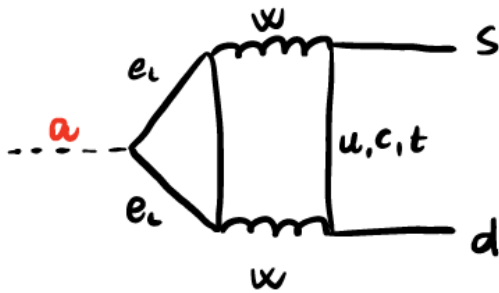


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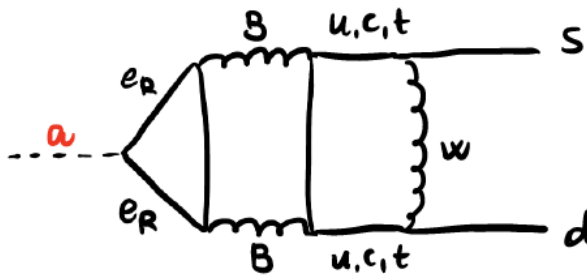
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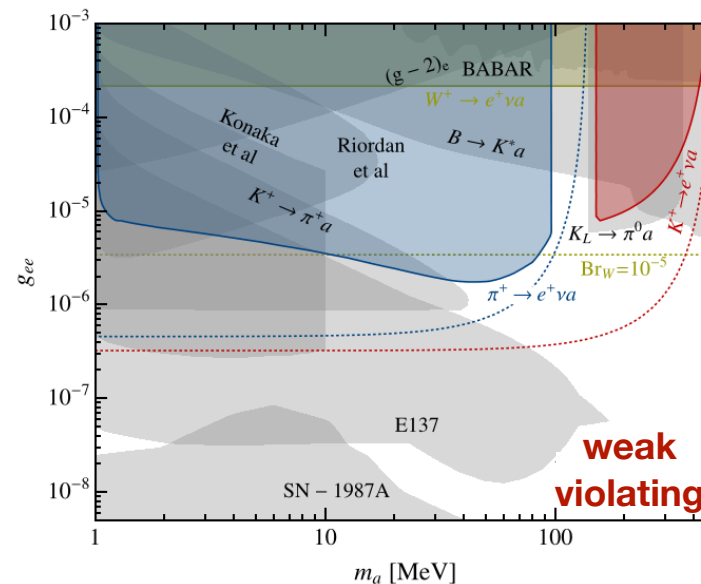
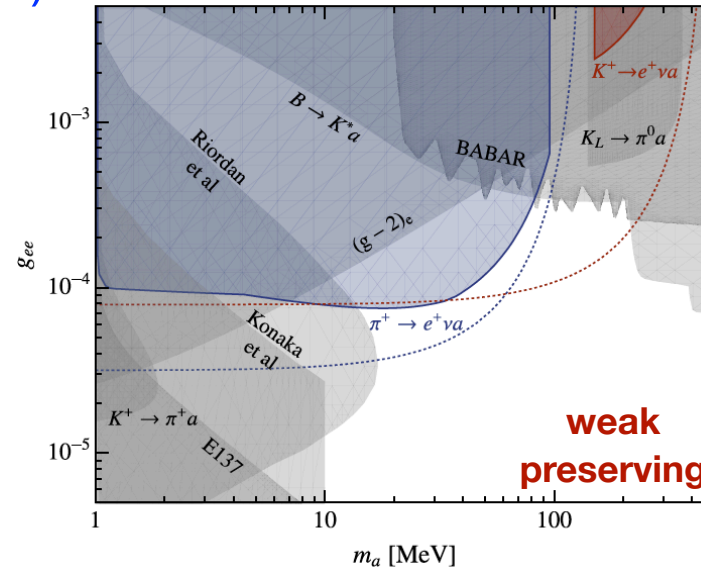
1. Axion with LH coupling



2. Axion with RH coupling



Similar diagrams for  $B \rightarrow K^{(*)} a$



reinterpreting past data (not dedicated searches)



# What about the future? Advances in meson factories

A big jump in luminosity is expected in the coming years

## Past/Present

## Future

Pion-  
factories

PIENU experiment at TRIUMF:  
 $\sim 10^{11}$   $\pi^+$

PIONEER experiment at PSI  
(phase 1 approved. Data in  $\sim 2028(?)$ ):  
 $\sim 10^{12}$   $\pi^+$

Kaon-  
factories

E949 at BNL:  $\sim 10^{12}$   $K^+$   
E391 at KEK:  $\sim 10^{12}$   $K_L$

NA62 at CERN:  $\sim 10^{13}$   $K^+$   
KOTO at JPARC:  $\sim 10^{14}$   $K_L$

B-  
factories

LHCb: more than  $\sim 5 \cdot 10^{12}$   $b$   
quarks produced so far;  
Belle (running until 2010):  
 $\sim 10^9$   $BB$ -pairs were produced.

LHCb:  $\sim 30$  times more  $b$  quarks  
will be produced by the end of the LHC;  
Belle-II:  $\sim 50$  times more  $BB$ -pairs  
will be produced.

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**Proton fixed target:**  $10^{18}$ ;  $10^{20}$   $\pi^+$

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**Proton fixed target:**  $10^{17}$ ;  $10^{20}$   $K^+$ ,  $K_L$

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Belle-II:  $\sim 50$  times more  $BB$ -pairs  
will be produced.

**Proton fixed target:**  $10^8$ ;  $10^{13}$   $B$

+ LHC,  
future  
colliders

**Proton fixed target** experiments also produce a huge statistics of mesons.

S.Gori E.g., **DarkQuest** (120 GeV protons,  $2 \cdot 10^{18}$  POT); **SHiP** (400 GeV protons,  $2 \cdot 10^{20}$  POT)

# Future axion searches

## PIONEER

is  $\text{BR}(\pi^+ \rightarrow e^+ \nu(a \rightarrow e^+ e^-)) = 10^{-11}$  reachable?  
 what about displaced axions?

## NA62

is  $\text{BR}(K^+ \rightarrow e^+ \nu(a \rightarrow e^+ e^-)) = 10^{-10}$  reachable?  
 what about displaced axions?

## Belle-II

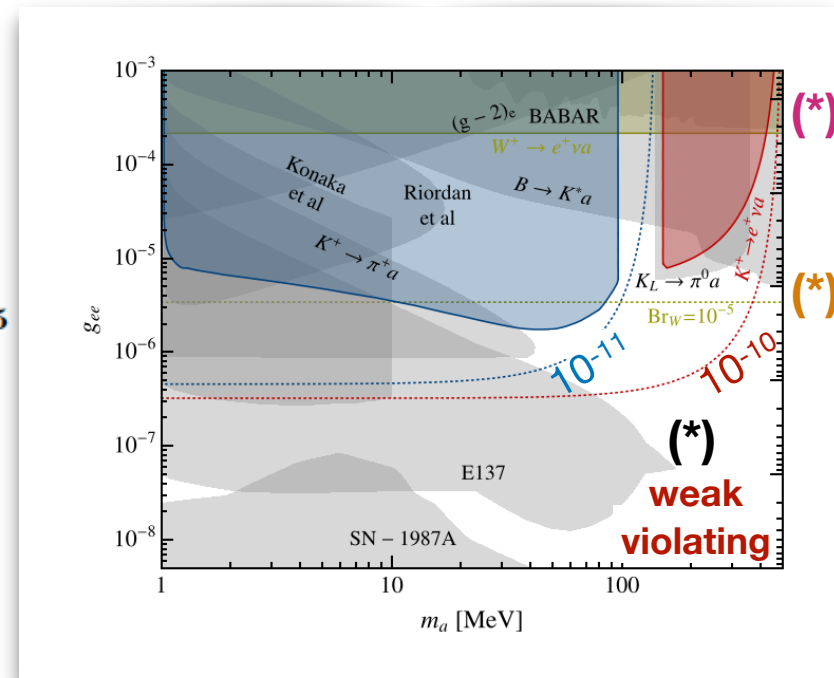
(\*) what about  $B^+ \rightarrow e^+ \nu(a \rightarrow e^+ e^-)$  ?

## LHC

is  $\text{BR}(W^+ \rightarrow e^+ \nu(a \rightarrow e^+ e^-)) = 10^{-5}$  reachable? (\*)

## DarkQuest & SHiP

Searches for a displaced axion decaying to  $e^+ e^-$  (\*)



# Conclusions & Outlook

Axions are very well motivated particles beyond the Standard Model.  
Large experimental effort to search for the minimal QCD axion.

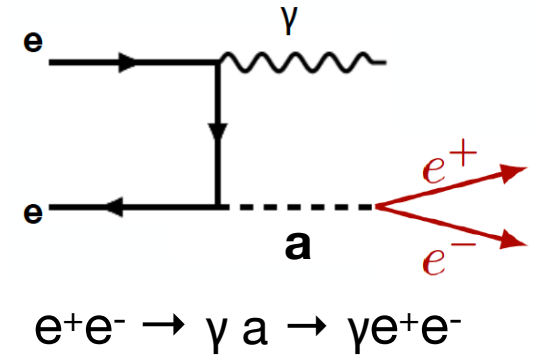
Axions can

- be a Dark Matter candidate or
- mediate interactions between the Dark Matter and the SM. Several cosmological models and tests with astrophysical and cosmological data

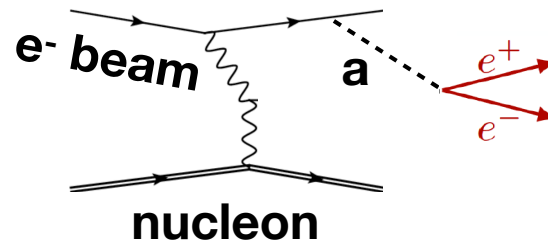
Heavier axions can still address the strong CP problem and lead to new interesting phenomenology for meson factories, LHC, and beam dump experiments.

# How to search for these axion interactions?

1. Reinterpretation: B- factories **direct production**  
(Babar. For prompt axions)

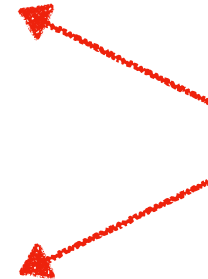
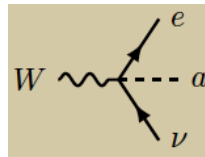


2. **Direct production** at electron beam dump experiments  
(For displaced axions)



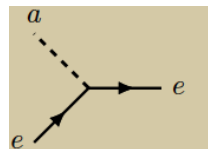
3. W and **meson decays** through **charged currents**  
(LHC, meson factories, proton beam dump experiments.  
For prompt and displaced axions).

**Example:**  $\pi^+ \rightarrow a e^+ \nu$   
**NEW**



Novel searches  
can be done in  
the coming years

4. **Meson decays** through **neutral currents**  
(meson factories, proton beam dump experiments.  
For prompt and displaced axions).

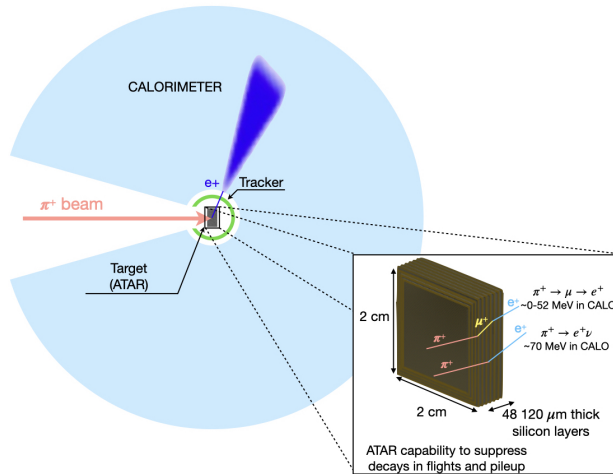


**Example:**  $B \rightarrow K a$  **Most studied**

# The most recent experiments

## PIONEER

**pion**-decay-at rest experiment

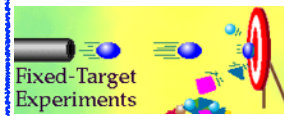
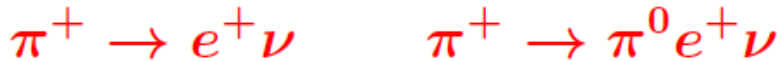


Experiment **approved** in summer '22 for **PSI** (Switzerland).

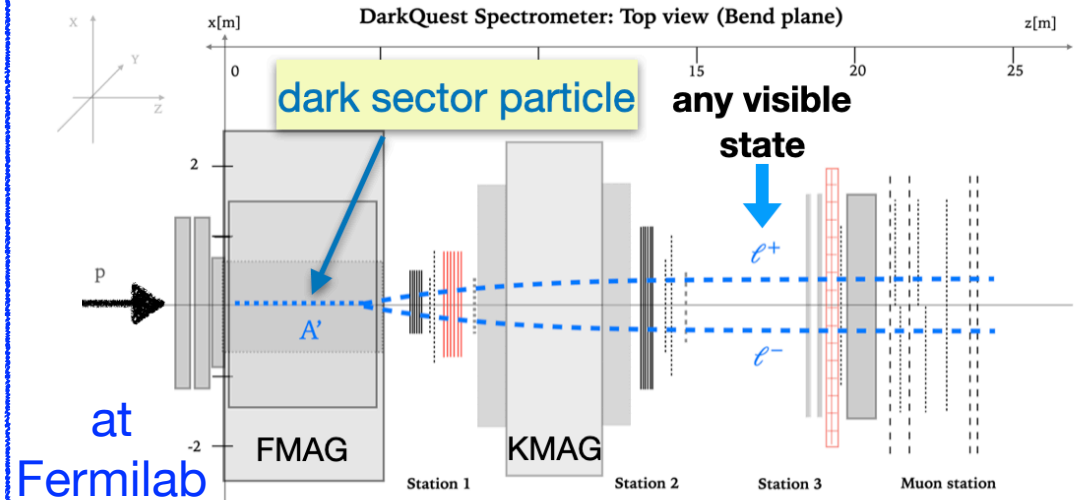
See the proposal:  
Altmannshofer et al., [2203.01981](#)

**Main physics goals:**

most precise measurement of  
**very rare pion decays**



## DarkQuest (upgrade of SpinQuest)



at  
Fermilab

**SpinQuest**  
Took data last  
**April-June**  
(only muon  
signatures)

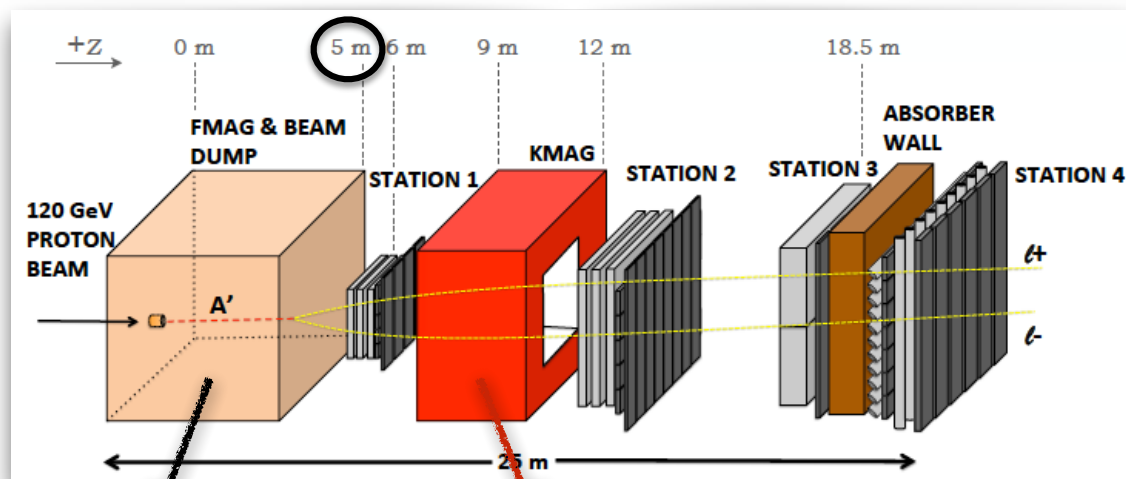
→ **DarkQuest**  
**proposed upgrade**  
(calorimeter +  
more tracking layers +  
hodoscope for triggering)

Initial proposal: Berlin, SG,  
Schuster, Toro, [1804.00661](#)

August '23:  
**We got funded!**

Similar in scope to the SHiP experiment at CERN

# Details on DarkQuest



## 1. Compact geometry



Sensitivity to (slightly) displaced dark particles with  $d > 5\text{ m}$

FMAG sweeps away soft SM radiation

2. KMAG separating even very forward charged particles ( $\Delta p_T \sim 0.4\text{ GeV}$ )



Identification of very light dark particles/squeezed spectra

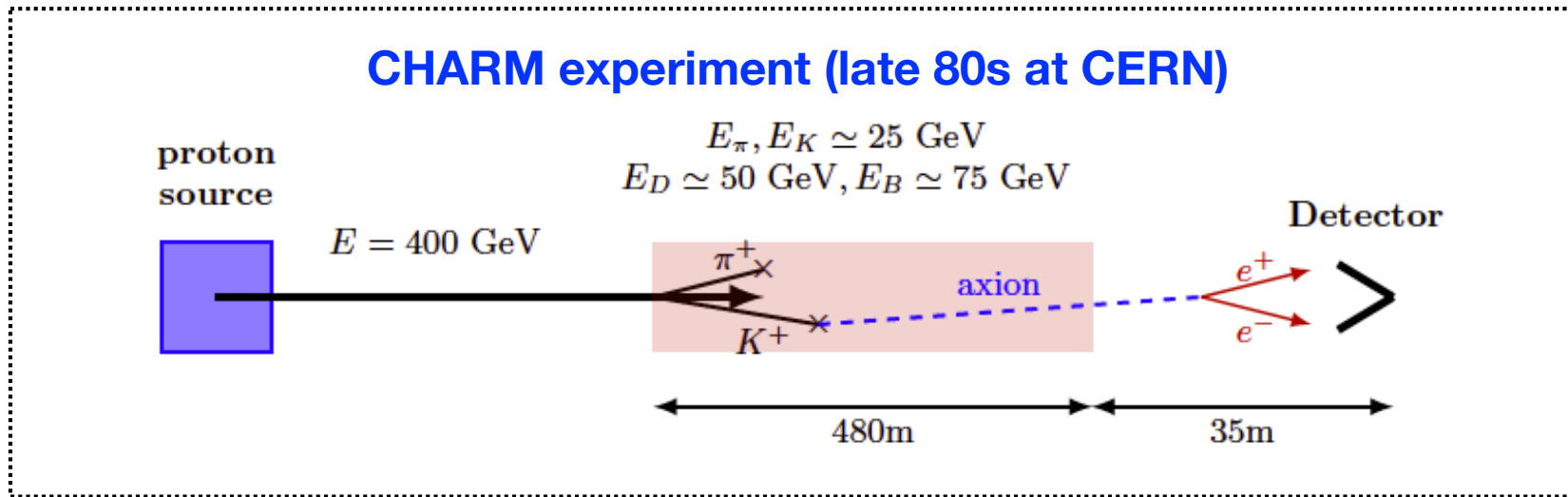
Experiment	Proton energy	POT	Dump	Decay volume
DarkQuest	120 GeV	$10^{18}$	5 m	10 m
CHARM	400 GeV	$2.4 \times 10^{18}$	480 m	35 m
LSND	800 MeV	$10^{22}$	30 m	10 m
NA62-dump	400 GeV	$5 \times 10^{19}$	100 m	250 m
SHiP	400 GeV	$2 \times 10^{20}$	30 m	50 m



Past

Proposed

# Proton beam dump experiments & axions



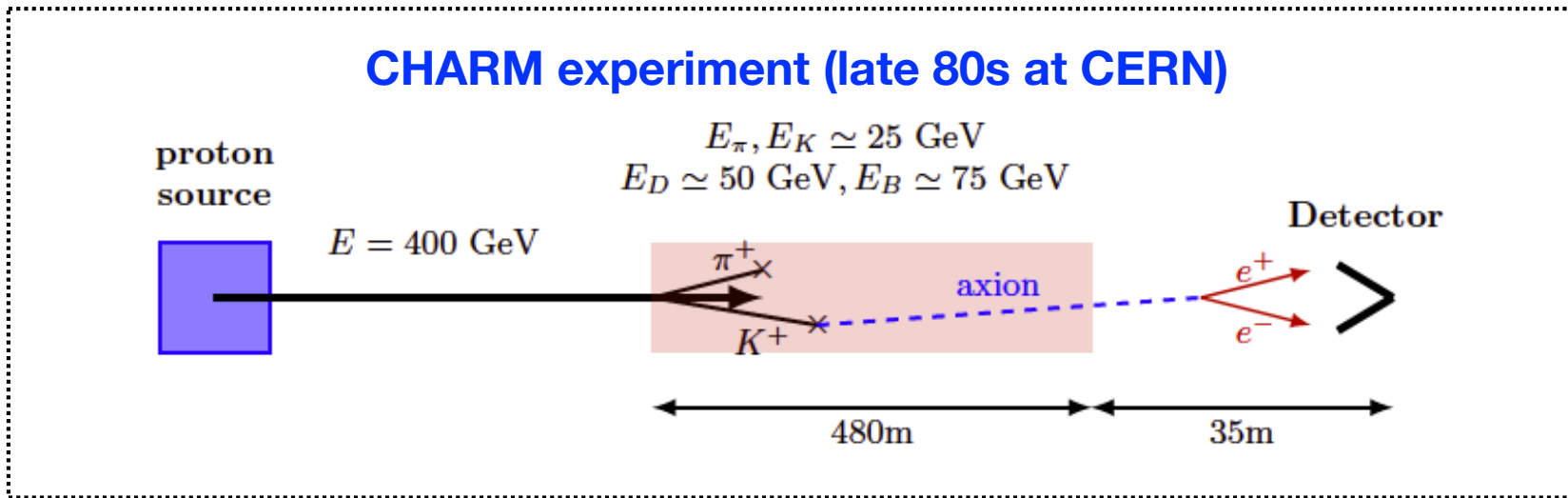
Past proton beam dump experiments produced unprecedented amounts of mesons. For example, the **CHARM** experiment collected  $\sim 10^{20}$  pions

Several studies for axions (or other dark sector particle) production at past proton beam dump experiments:  $K \rightarrow \pi a$  or  $B \rightarrow K a$

However, **charged current production has been neglected** ( $K^+ \rightarrow a e^+ \nu$ , ...)

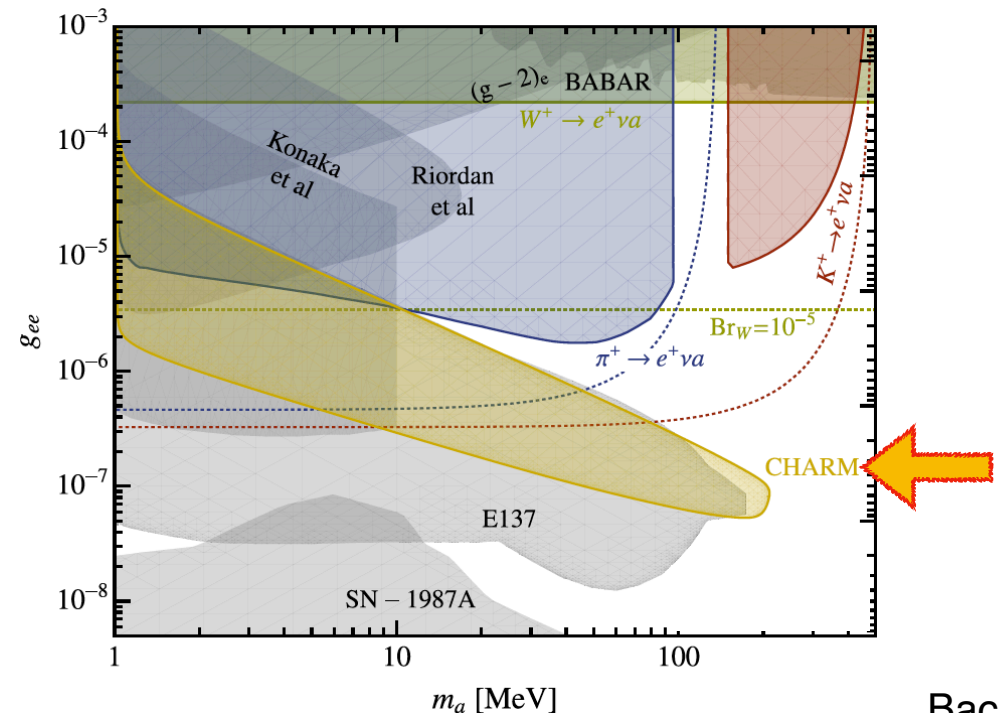


# Proton beam dump experiments & axions



New in our study:

- \* First use of  $D, D_s, B_c$  mesons
- \* Introduction of charged current meson decays

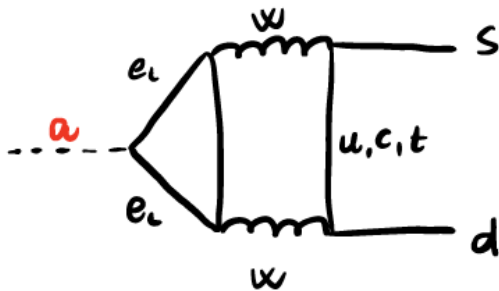


# Neutral current decays

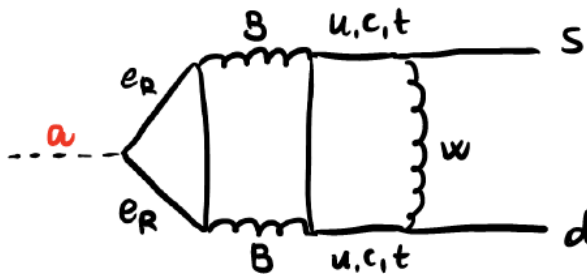
Neutral current meson decays are also generated at the 2 or 3-loop level (suppressed by CKM elements as well)

$$K \rightarrow \pi a$$

1. Axion with LH coupling



2. Axion with RH coupling



They arise from the anomaly terms induced at loop level

$$-\frac{g_{aW}}{4} a W_{\mu\nu}^a \tilde{W}^{a\mu\nu} - \frac{g_{aB}}{4} a B_{\mu\nu} \tilde{B}^{\mu\nu}$$

1. The sum of these diagrams is finite

$$\mathcal{L} \supset \frac{\partial_\mu a}{m_\ell} g_{d_i d_j} (\bar{d}_i \gamma^\mu P_L d_j) + \text{h.c.}$$

$$g_{d_i d_j} \simeq -\frac{g^2}{16\pi^2} V_{ti}^* V_{tj} \left[ \frac{g^2}{16\pi^2} \frac{3}{8} (g_{ee} - \bar{g}_{ee} - g_{\nu_e}) F(x_t) + \frac{g'^4}{(16\pi^2)^2} \frac{17}{96} (g_{ee} + \bar{g}_{ee}) \frac{m_t^2}{m_W^2} \log^2 \left( \frac{\Lambda^2}{m_t^2} \right) \right]$$

2. Log divergent

Similar diagrams for  $B \rightarrow K^{(*)} a$

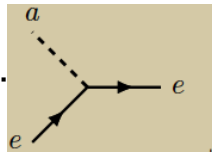
$$\Gamma(K^+ \rightarrow \pi^+ a) \propto |g_{asd}|^2$$

$$\Gamma(K_L \rightarrow \pi^0 a) \propto \text{Im}(g_{asd})^2$$

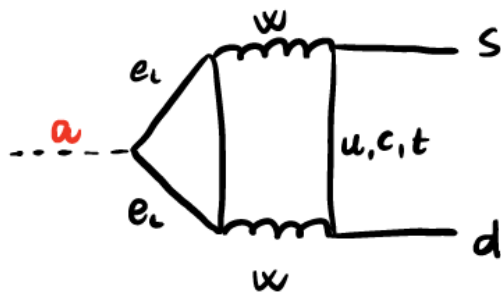
# Neutral current decays

Neutral current meson decays are also generated at the 2 or 3-loop level (suppressed by CKM elements as well)

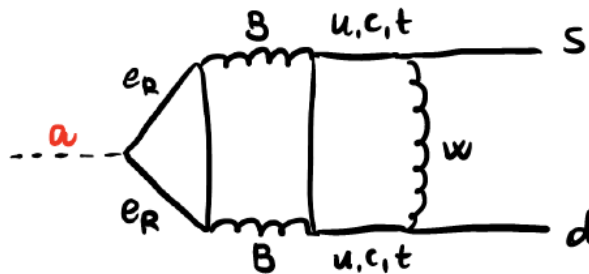
$$K \rightarrow \pi a$$



1. Axion with LH coupling



2. Axion with RH coupling



Similar diagrams for  $B \rightarrow K^{(*)} a$

**Main bounds: reinterpreting past data**

NA62, 2103.15389:  $K^+ \rightarrow \pi^+ + (a \rightarrow \text{invisible})$

KTeV, 0309072:  $K_L \rightarrow \pi^0 (a \rightarrow e^+ e^-)$   
 $m_{ee}$  in (140, 363) MeV

777 @ BNL, Phys. Rev. Lett. 59 (1987) 2832–2835:

$K^+ \rightarrow \pi^+ (a \rightarrow e^+ e^-)$   $m_a < 100$  MeV

see also Alves, Weiner, 1710.03764

LHCb, 1501.03038:  $B^0 \rightarrow K^{(*)0} (a \rightarrow e^+ e^-)$

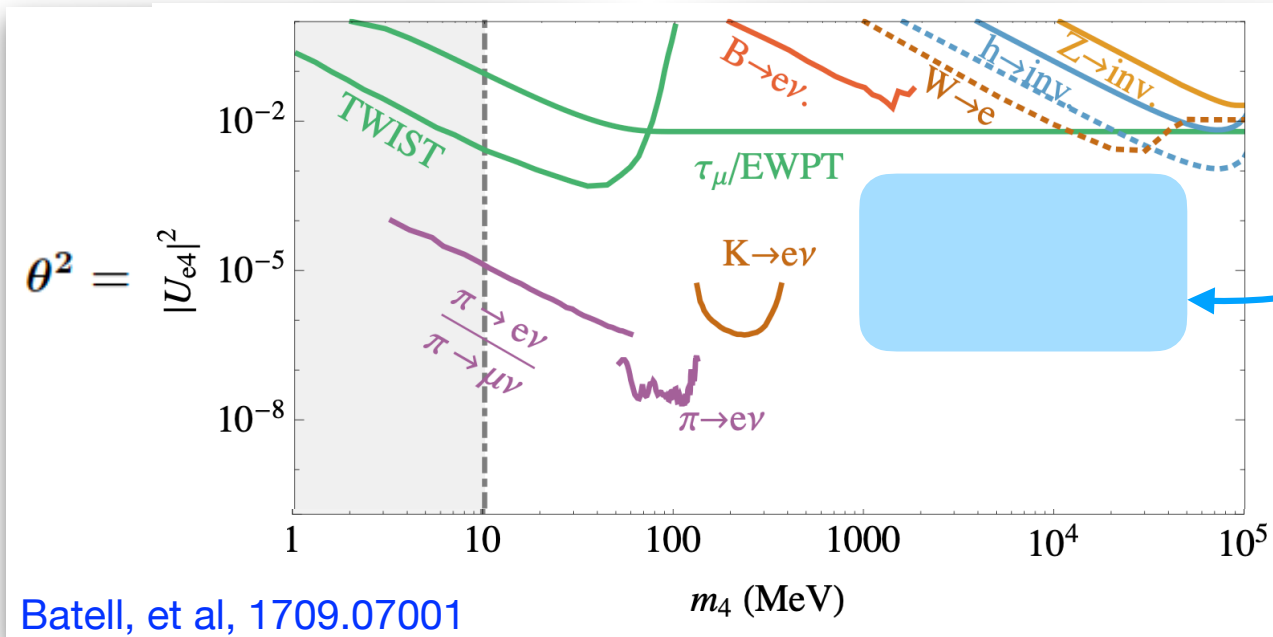
# Example of weak violating axions

$$\mathcal{L} \supset -yHLN^c - Me^{ia/f_a}NN^c + \text{h.c.}$$

N (with Dirac partner  $N^c$ )  
charged under PQ

After EWSB, sterile neutrinos and SM neutrinos mix and

$$\mathcal{L} \supset \frac{\theta^2}{f_a} \partial_\mu a (\bar{\nu}_e \gamma^\mu P_L \nu_e) \quad \rightarrow \quad \begin{cases} g_\nu = \frac{2\theta^2 m_e}{f_a} = 1.0 \times 10^{-5} \left(\frac{\theta}{0.1}\right)^2 \left(\frac{\text{GeV}}{f_a}\right) \\ g_{ee} = \bar{g}_{ee} = 0 \end{cases}$$



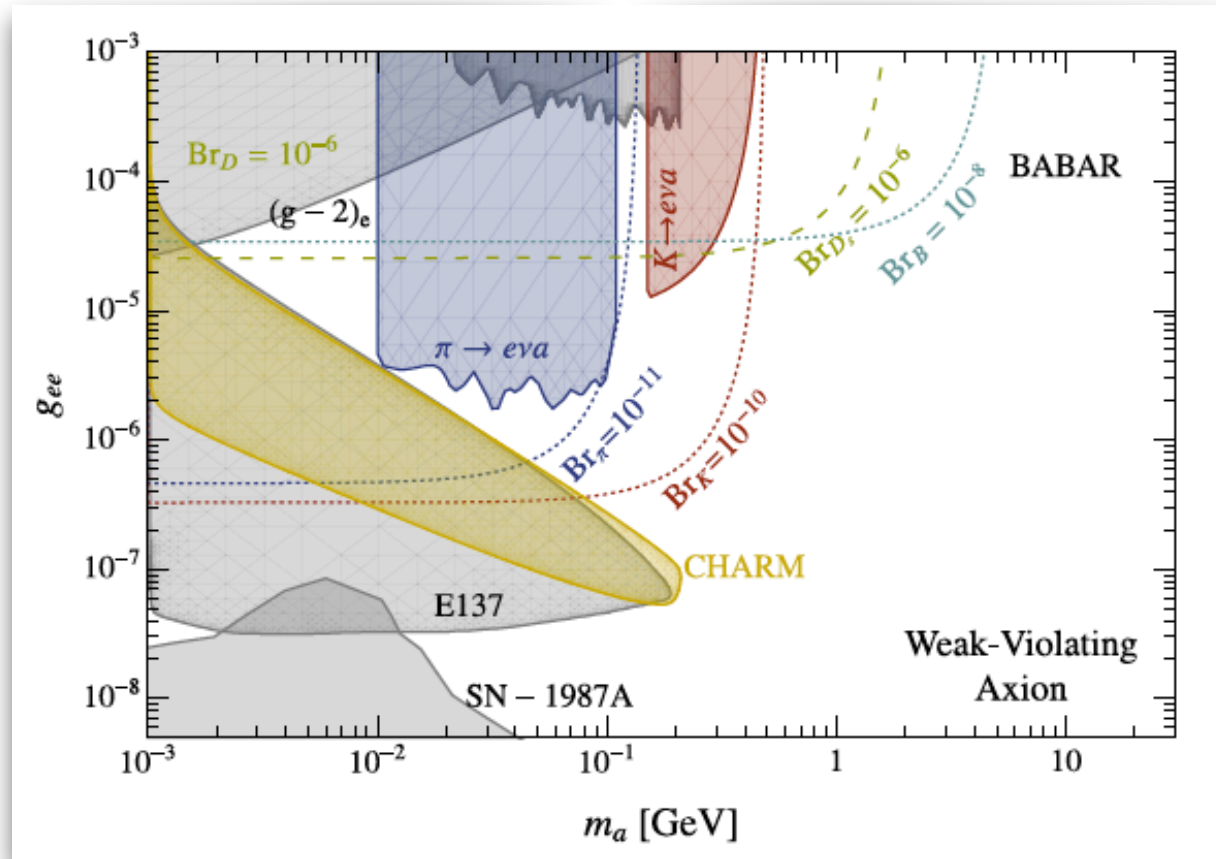
Batell, et al, 1709.07001

additional constraints  
from visibly decaying  
HNL (less robust)

Similar construction  
for vector-like leptons  
charged under PQ:

$$\bar{g}_{ee} - g_{ee} \neq 0, \quad g_{\nu_e} = 0$$

# Charged current heavy meson decays



# Lepton-coupled axion lifetime

