



# Detecting Light DM via a Novel Torsion Balance

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Based on: *Detection of Dark Matter Coherent Scattering via Torsion Balance with Test Bodies of Different Sizes* [*arXiv: 2409.09950*] by Shigeki Matsumoto, J.S., Chuan-Yang Xing

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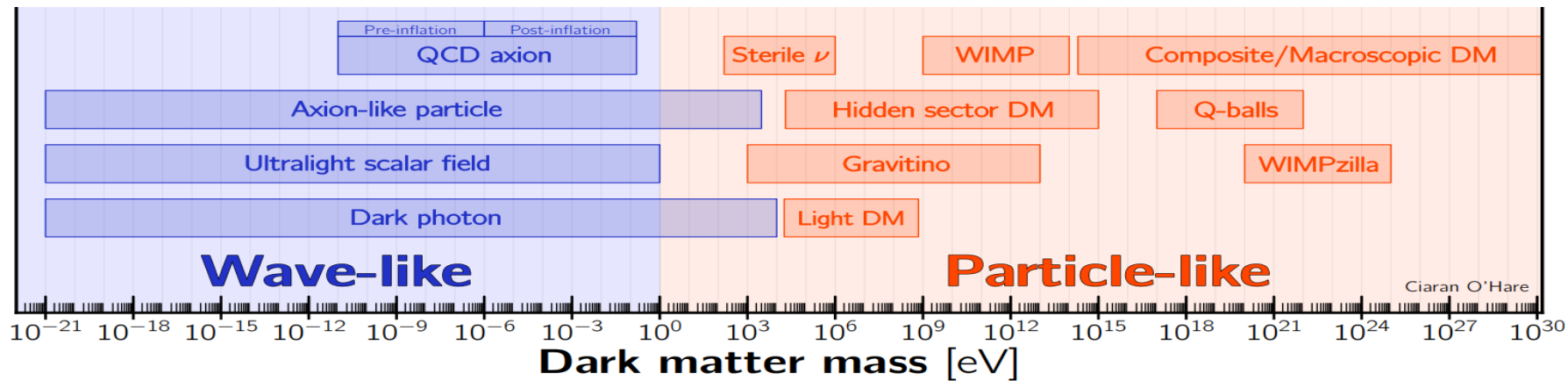
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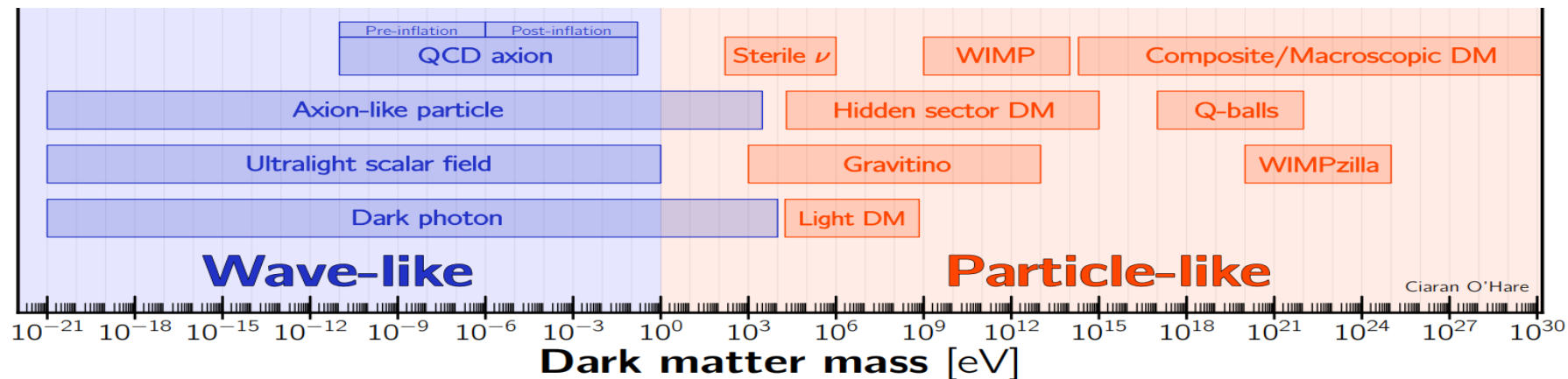
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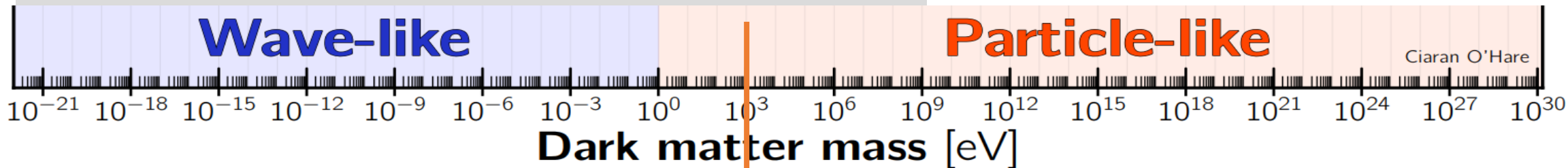
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- The detection of DM highly depends on the **DM mass** and its **interaction** with SM particles.
- In the same mass range, there can still be some common rules.

# 1.1 DM mass above keV



1) DM mass  $m_\chi > O(10\text{GeV})$ , (WIMP-like).

DM-nucleon elastic scattering

Kinetic energy  $T_\chi \sim m_\chi v^2 > \text{keV}$ .

2) DM mass  $\text{MeV} < m_\chi < O(10\text{GeV})$ ,

DM-electron elastic scattering

Kinetic energy  $T_\chi \sim m_\chi v^2 > 1\text{eV}$ .

3) DM mass  $\text{keV} < m_\chi < \text{MeV}$ ,

Further lower the threshold:

Scattering with optical trapped atoms, superconductors;

Boosted DM;

DM absorption:

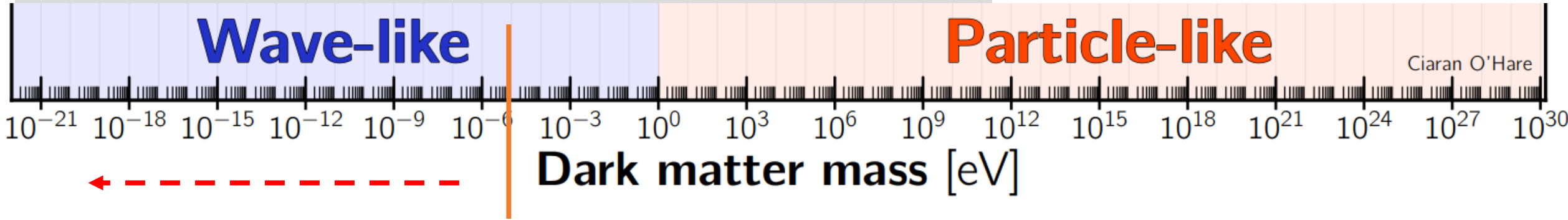
Hidden photon and Fermionic DM absorption

Kinetic energy  $T_\chi \sim m_\chi v^2 < 1\text{eV}$ .

[2203.08084, 2203.07492,  
2201.11497, 2111.03597v3...]

*Single Scattering Event Beyond the Detection Threshold*

# 1.2 DM mass below $10^{-5}$ eV



1) Pseudo-Scalar DM (axion, ALP).

[a-neutron: [CASPER, Comagnetometer...](#)]; [a-electron: [NV center...](#)]; [a-photon: ADMX (Haloscope)...]

2) Scalar DM.

[Atomic clock](#); [Atom interferometers...](#)

2) Vector DM (without Z2 symmetry).

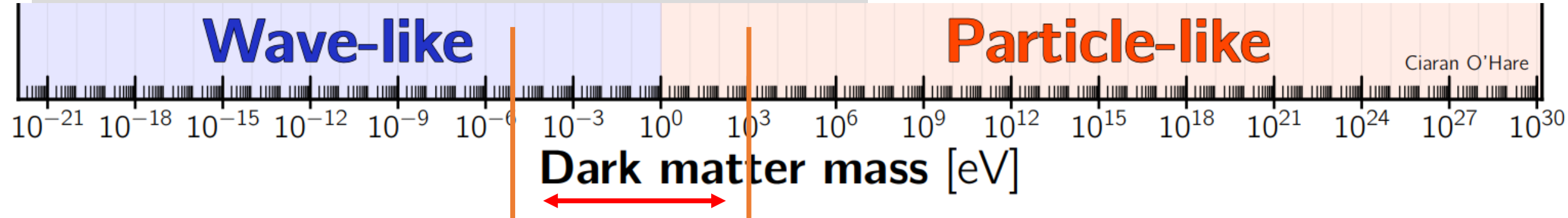
[B-L gauge boson: [Fifth force, Ligo...](#)]; [dark photon: [Haloscope...](#)]

[2203.14915v1]

*Coherent DM states as a background perturbation to affect the fundamental constants...*



# 1.3 DM mass in between



- States begin to overlap:  $n V_{dB} \geq 1$ . DM number density is  $n = \frac{\rho_\chi}{m_\chi}$  and the de Broglie volume of a particle is  $V_{dB} = (m_\chi v_\chi)^{-3}$ .
- Once  $m_\chi < 1$  eV, the quantum wave of DM particle starts to overlap. Thus, we call  $\sim 1$ eV a **crossover mass range between wave and particle DM**.
- The recoil energy from single DM scattering event is not enough. **Not particle-like enough.**
- The DM can not form coherent modes. **Not wave-like enough either.**
- **There are currently few sensitive detection methods particularly targeting it.**

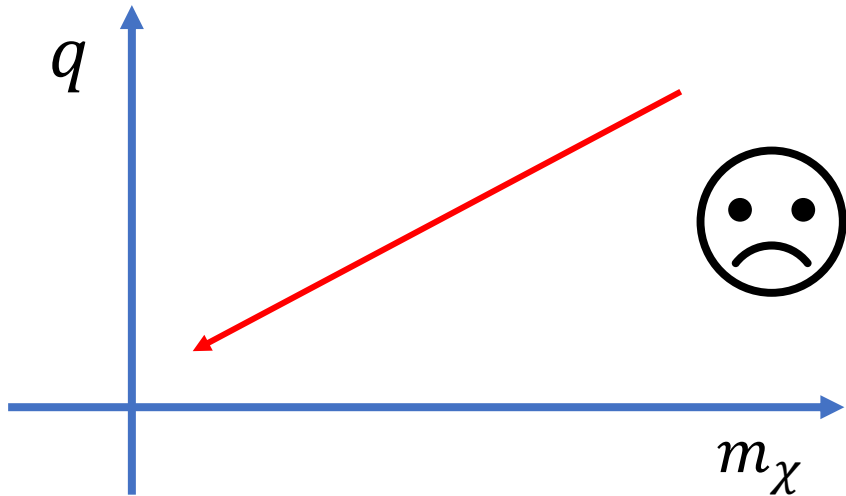
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Since all the particle targets we can find in SM is much heavier than  $1\text{eV}$ , the momentum transfer of a single DM scattering event is roughly  $q \sim p_\chi = m_\chi v_\chi = 10^{-3} m_\chi$ .

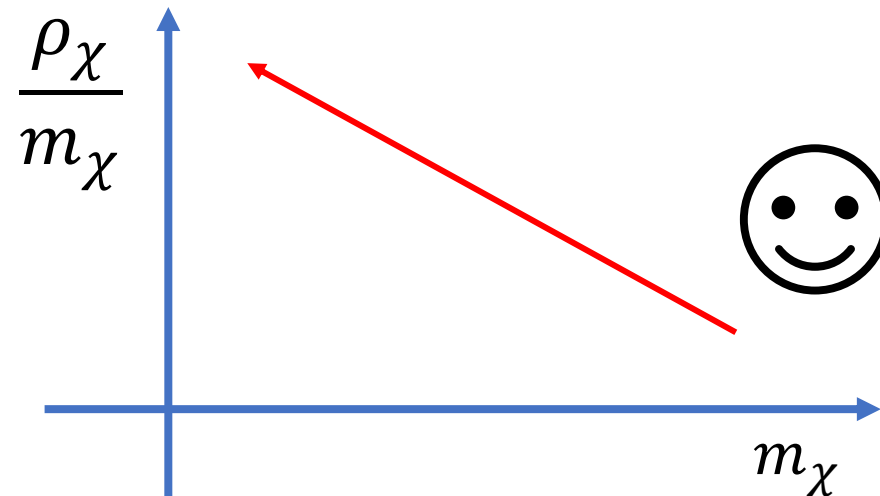
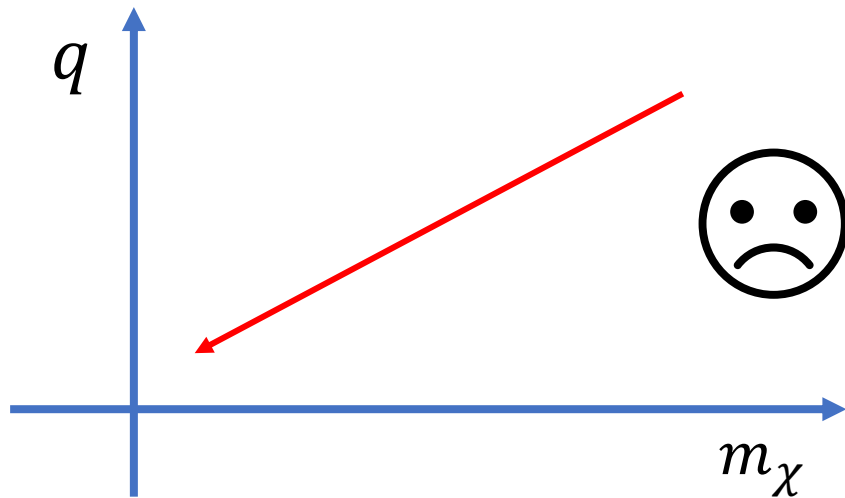
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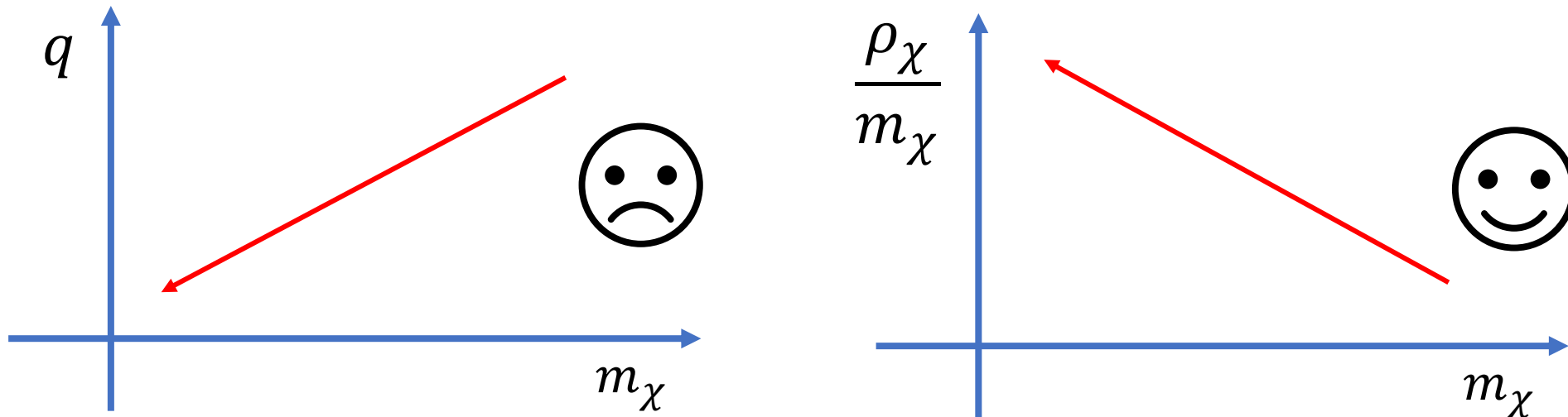
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*Since individual event cannot be detected, we need some collective effects.*

## 2.1 Acceleration (Force)

DM has a relative velocity with the earth. As a result the continuous scattering will induce some force and acceleration to the targets.

$$F = \frac{\rho_{\chi}}{m_{\chi}} \sigma_{tot} v q = \frac{\rho_{\chi}}{m_{\chi}} \sigma_{tot} v_{\chi} m_{\chi} v_{\chi} = \rho_{\chi} \sigma_{tot} v_{\chi}^2.$$

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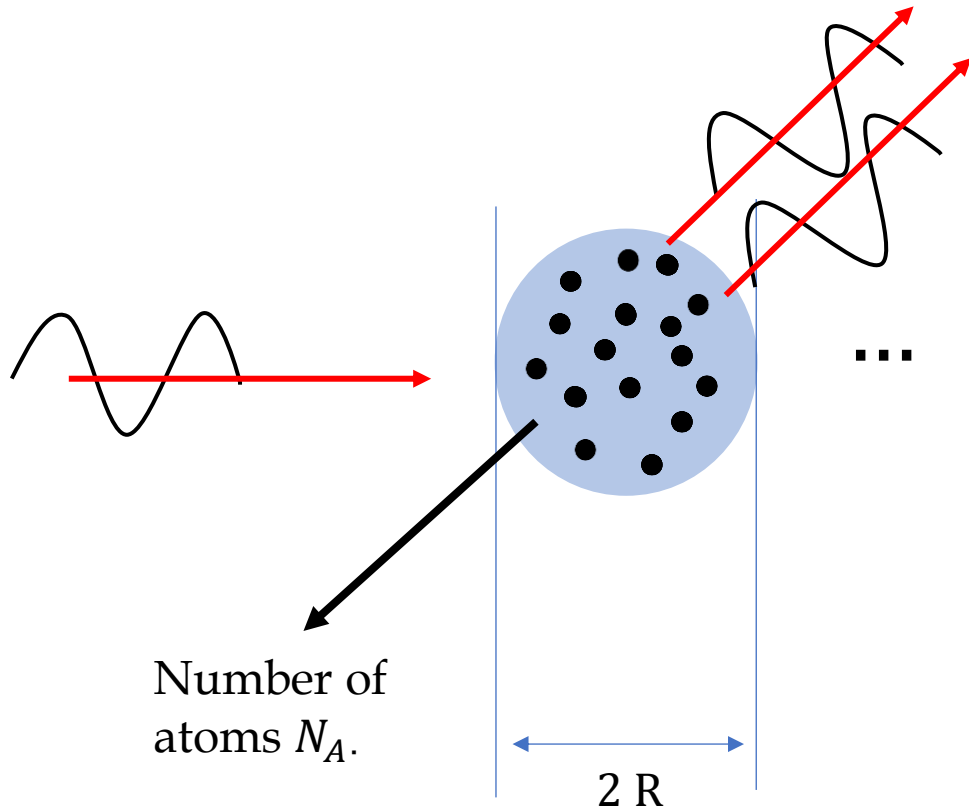
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- The mass of DM seems to be cancelled in the formalism.
- To enhance the signal, the only possibilities lies in the cross section.
- Fortunately, we **have coherent scattering effects** for light DM with a macroscopic wave length.

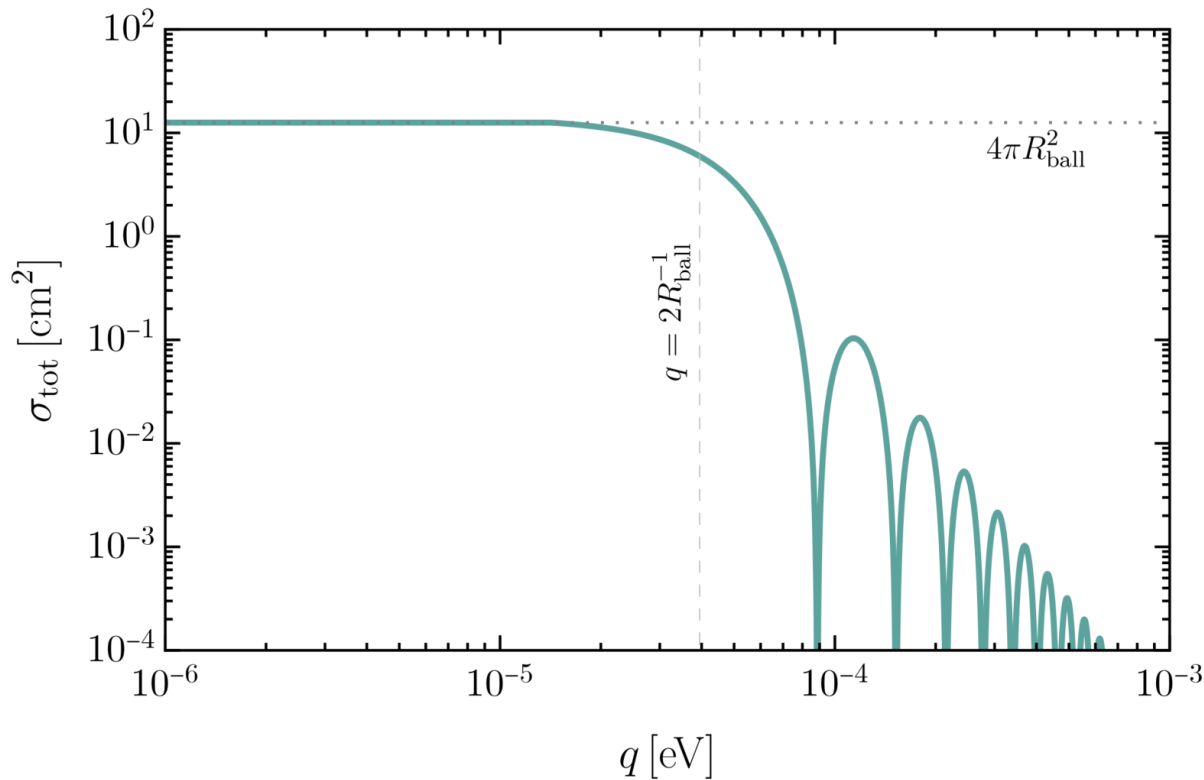


## 2.2 Coherent Scattering



- If we receive the outgoing wave at infinite far with initial and final momentum fixed, the only difference of these waves is the **phase factor**  $e^{iqr_j}$ .
- In the amplitude level,  $Amp \rightarrow Amp \sum e^{iqr_i}$ .
- In the cross section level,  $\sigma_{tot} = \sigma_{\chi A} \sum_{ij} e^{iq\Delta r_{ij}}$ .
- If  $q\Delta r_{ij} \ll 1$ ,  $e^{iq\Delta r_{ij}} = 1$ . The enhancement of total cross section is  $N_A^2$  as we expected. [This  $N$  square coherent factor is verified by neutrino scattering in CEvNS.]
- If  $q\Delta r_{ij} \gg 1$ , all the phases cancel each other except for the cases  $i = j, \Delta r_{ii} = 0$ . The enhancement factor is simply  $N_A$ .

# 2.2 Coherent Scattering



- Assuming  $N_A$  atoms in a **ball-like body** with radius  $R_{ball}$ , the cross section dependence on momentum transfer is shown as the green line.
- The coherent effects **start to disappear** when the momentum transfer roughly corresponds to the geometric size of the body  $\frac{1}{q} \sim R_{ball}$ .
- Naively, the largest cross section for the small  $q$  limit should be  $N_A^2 \sigma_{\chi A}$ . However, the cross section will **saturate to the geometric area** of the body  $4 \pi R_{ball}^2$  due to the unitarity of quantum mechanics.
- As a result,  $\sigma_{tot} = \text{Min} [N_A^2 \sigma_{\chi A}, 4 \pi R_{ball}^2]$ .

## 2.3 Enhancement of Acceleration

$$F = \frac{\rho_\chi}{m_\chi} \sigma v q = \frac{\rho_\chi}{m_\chi} \sigma_{tot} v_\chi m_\chi v_\chi = \rho_\chi \sigma v_\chi^2.$$

$$a = \frac{F}{m_{target}} = \frac{\rho_\chi \sigma_{tot} v_\chi^2}{m_{target}} = \frac{\rho_\chi N_A^2 \sigma_{\chi A} v_\chi^2}{N_A m_A} = \frac{\rho_\chi N_A \sigma_{\chi A} v_\chi^2}{m_A}.$$

The force or acceleration is now only proportional to the total cross section!

*This rough derivation shows that, with coherent scattering effect, the acceleration can be enhanced by a factor an atom number, which is of order  $10^{23}$  in macro-objects.*

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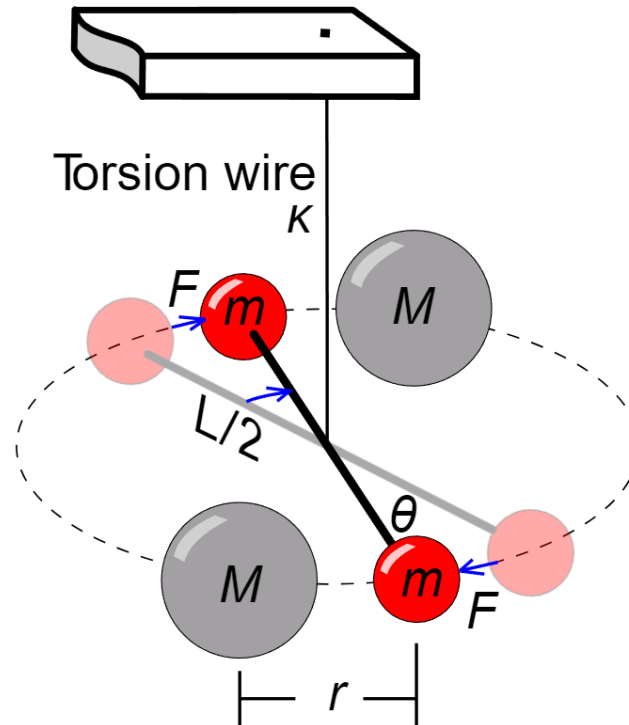
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Gravity is the weakest force.

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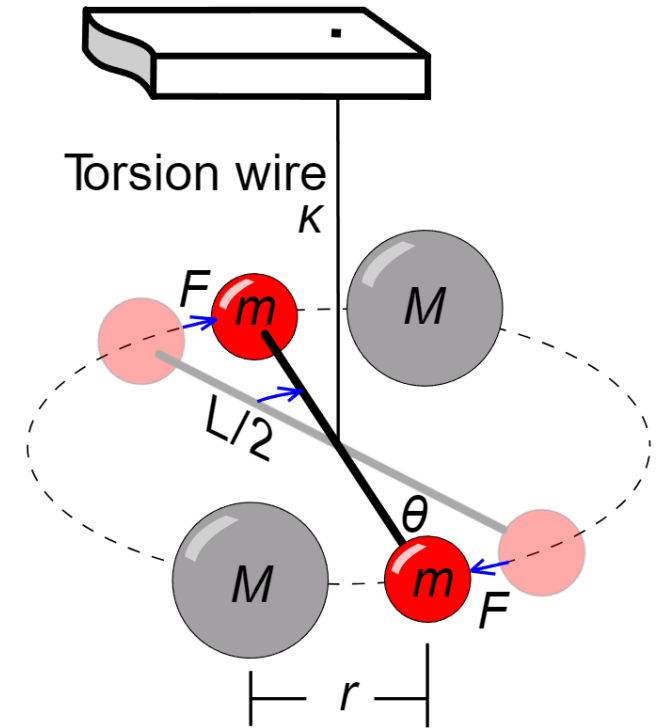
With a sensitivity:

$$\delta a = \frac{10^{-13} \text{ cm}}{\text{s}^2}.$$

[Cavendish Experiment (1797)]

## 3.2 Asymmetry is needed

- The current torsion balance setup used to measure the weak gravity and test equivalence principle contains two balls with **same masses and same sizes**.
- The entire device has rotational symmetry around the central wire.
- As a result, the DM wind induces same forces on both sides. No signal can be detected.
- *Asymmetry is needed!*



# 4. Q: How to design the setup?

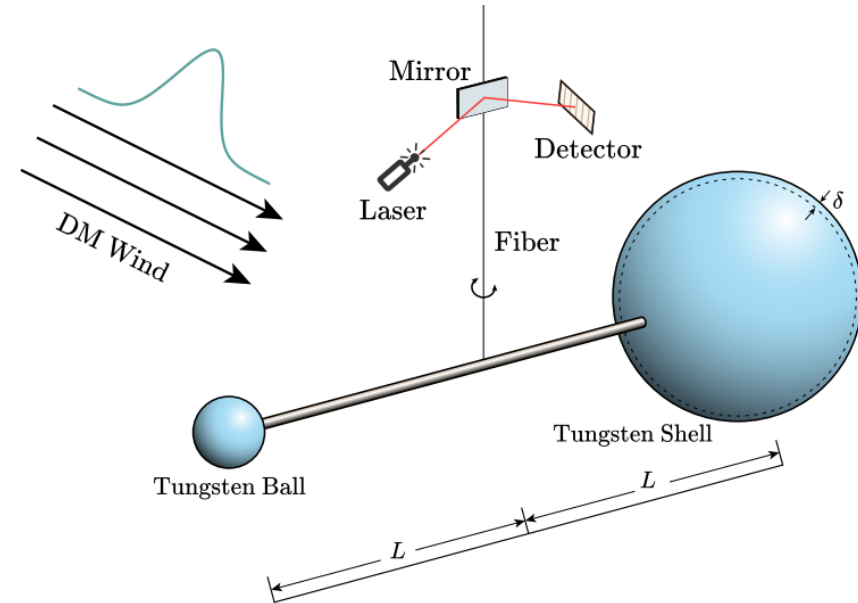


# 4. Q: How to design the setup?

- To make the torsion in balance, the two targets should have the same mass (If the arm lengths are the same).
- The masses means the same numbers of atoms.
- To have different coherent form factor, the two targets should have different geometric sizes.
- *Targets with equal masses but different sizes.*

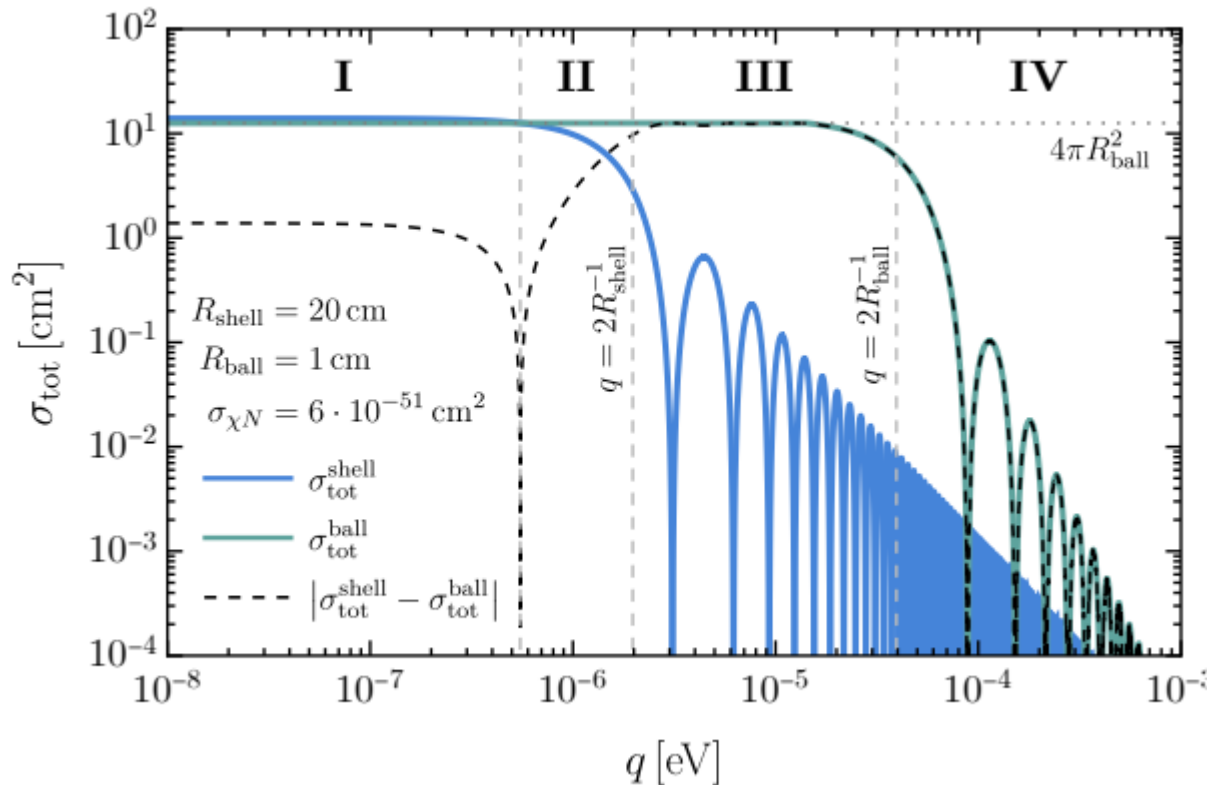
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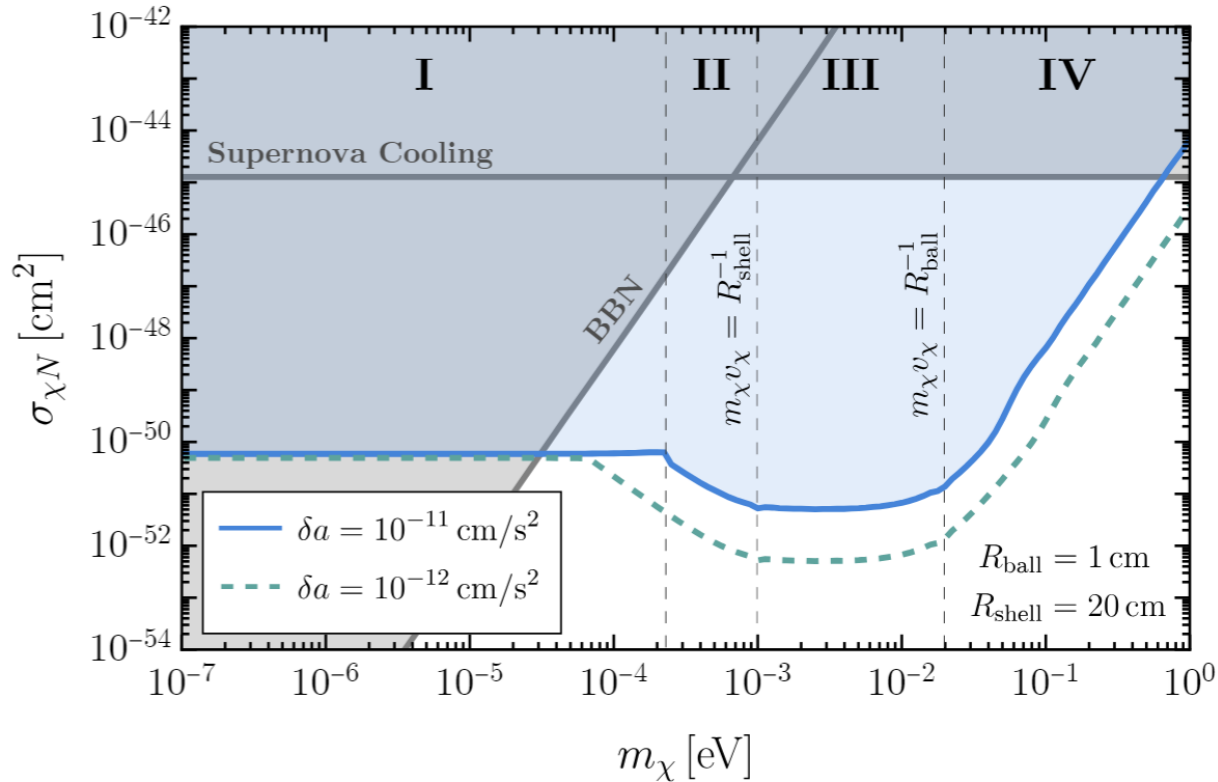
	Material	Density	Radius	Thickness	Mass
Ball	Tungsten	19.25 g/cm <sup>3</sup>	1 cm	-	80.6 g
Shell	Tungsten	19.25 g/cm <sup>3</sup>	20 cm	8.3 μm	80.6 g

# 5 Results



- Region I: DM wavelength much larger than both the radii of the ball and the shell.  
 $\sigma_{shell} = N_A^2 \sigma_{\chi A}$ ;  $\sigma_{ball} = 4\pi R_{ball}^2$ .
- Region II: DM wavelength  $\sim R_{shell}$ ;  
 Coherent effects of the shell starts to decrease.
- Region III:  $R_{ball} < \text{DM wavelength} < R_{shell}$ ;  
 Coherent scattering only with the ball.
- Region IV: DM wavelength  $< R_{ball}$ ;  
 Coherent effects gradually disappear.

# 5 Results



A reminder:  $q \sim p_\chi \sim m_\chi v_\chi \sim 10^{-3} m_\chi$

- The detection sensitivity of a stationary torsion with different test bodies is  $\delta a \sim 10^{-11}$  cm/s<sup>2</sup>. [Annals Phys. 26(1964) 442-517]
- Assume general models, such as  $\frac{\chi^2 G_{\mu\nu} G^{\mu\nu}}{\Lambda^2}$ ,  $\frac{m_i \chi^2 \bar{\psi}_i \psi_i}{\Lambda^2}$  with  $\psi_i$  the SM quarks.
- Constraint-1: supernova cooling due to the DM emission.
- Constraint-2: the DM will change the nucleon mass and affect the BBN.

# 5. Conclusion and Discussion

- How to detect DM?

The gap in the DM mass region in the crossover range of particle- and wave-DM

- How to detect DM in the above mass range?

Collective signals like force or acceleration.

- How to enhance the signal?

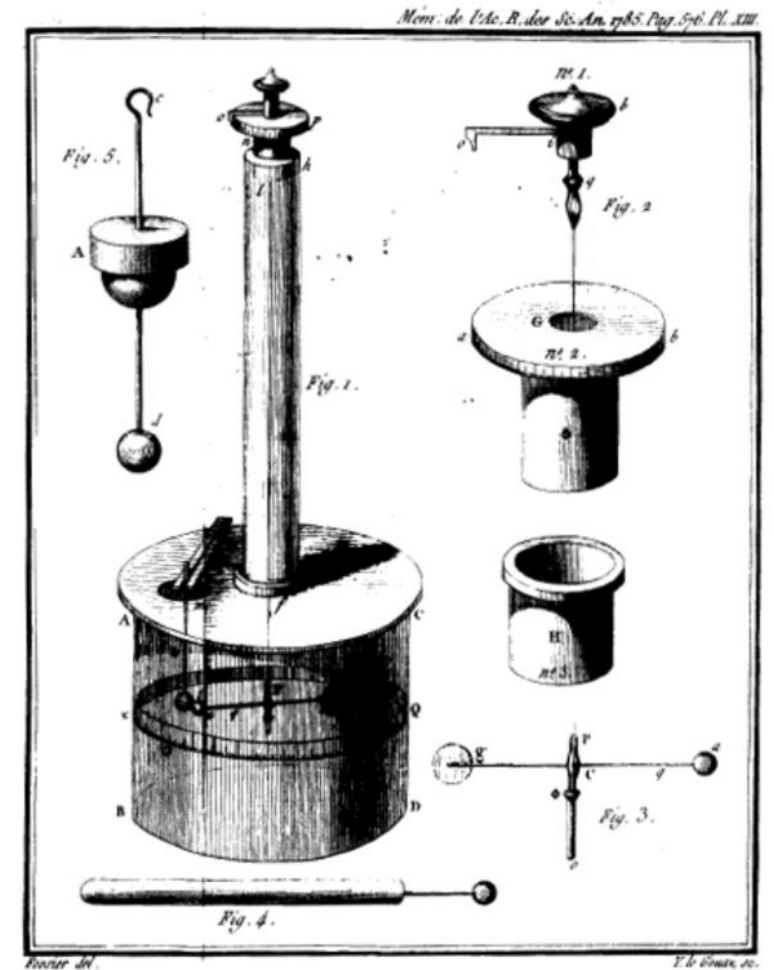
Coherent scattering of light DM.

- How to detect the DM-induced acceleration?

Torsion balance with test-bodies of equal mass but different sizes.

# An Interesting Discussion

- In 1785 (even earlier than Cavendish), Coulomb design **the exact same setup (a ball and a shell)** to measure the electric force.
- This is called as Coulomb's Torsion Balance[1785]!
- Since we are focus on the measurement of gravity over hundreds of years. Coulomb's torsion balance did not progress further.
- **Old devices, new purpose!**
- It is the time to develop Coulomb-type torsion balance again for DM detection!



# Backup: References

<sup>1</sup> In early studies, similar ideas were used for cosmic neutrino background ( $C\nu B$ ) detection [28, 32–35]. Unfortunately, due to the diffuse flux of cosmic neutrinos and their small scattering cross-section with nucleons, the resulting acceleration of the objects is far below the detectability threshold of current technologies.

Although the energy transfer of such a DM is insufficient to detect, their continuous momentum transfer can induce acceleration on macro-objects [28–31].<sup>1</sup>

# Backup: Unitarity

It is interesting that there is a maximum possible value for the cross section for each partial wave. From Eq. (20), the cross section for a given partial wave  $l$  is

$$\sigma_l = \frac{4\pi}{k^2}(2l + 1) \sin^2 \delta_l \leq \frac{4\pi}{k^2}(2l + 1). \quad (22)$$

This limit can be qualitatively understood in the following semi-classical argument. If you inject a particle with momentum  $p$  at the impact parameter  $b$ , it has angular momentum  $L = pb$ . On the other hand, the angular momentum is quantized in quantum mechanics,  $L = \hbar l$ . Let us say that the quantized angular momentum  $l$  corresponds roughly to the impact parameter  $b = L/p$  for  $\hbar l < L < \hbar(l + 1)$ , *i.e.*,

$$\frac{l}{k} \leq b \leq \frac{l + 1}{k}. \quad (23)$$

Assuming that the particle gets scattered with 100% probability when entering this ring, the classical cross section would be

$$\pi \left( \frac{l + 1}{k} \right)^2 - \pi \left( \frac{l}{k} \right)^2 = \pi \frac{2l + 1}{k^2}. \quad (24)$$

The unitarity limit is roughly the same as this semi-classical argument except for a factor of four.



# Backup: UV completion

As one would expect for a model of a light scalar with non-derivative couplings, the mass  $m_\phi$  is very fine-tuned. The dark matter relic abundance depends sensitively on  $m_\phi$  (for example through the misalignment mechanism [18-20]), so this tuning may have an anthropic origin [21]. On the other hand, a coupling of the form

$$\Delta\mathcal{L} = -\frac{\lambda_\phi}{4!}\phi^4 \quad (\text{B.1})$$

is allowed by all symmetries and is also UV sensitive.<sup>[6]</sup> This coupling is tightly constrained by constraints from structure formation [22]

$$\lambda_\phi \lesssim 3 \times 10^{-7} \left(\frac{m_\phi}{\text{eV}}\right)^4. \quad (\text{B.2})$$

$$\Delta\mathcal{L}_{\text{SM}} = -\frac{1}{2}m_\phi^2\phi^2 - \frac{\lambda_\phi}{4!}\phi^4 - \frac{1}{2}\epsilon\phi^2 H^\dagger H. \quad (\text{B.3})$$

This gives a  $\phi$  dependent shift in the Higgs VEV:

$$\frac{\delta v}{v} = \frac{\epsilon\phi^2}{2m_h^2}, \quad (\text{B.4})$$

where  $m_h = 125$  GeV is the physical Higgs mass. Below the electroweak symmetry breaking scale, this induces various couplings of  $\phi^2$  to standard model fields. For example, the coupling to SM fermions  $\psi$  is given by

$$\Delta\mathcal{L}_{\text{eff}} = -\sum_\psi \frac{1}{2f_\psi} \phi^2 \bar{\psi}\psi \quad \frac{1}{f_\psi} = \frac{\epsilon m_\psi}{m_h^2}. \quad (\text{B.5})$$

[arXiv: 2312.13345v1]