



TECHNISCHE
UNIVERSITÄT
DARMSTADT

PROBING HALOS VIA NUCLEON REMOVAL

Halo Week, Gothenburg, Sweden

June 10-14, 2024

Alexandre Obertelli, TU Darmstadt

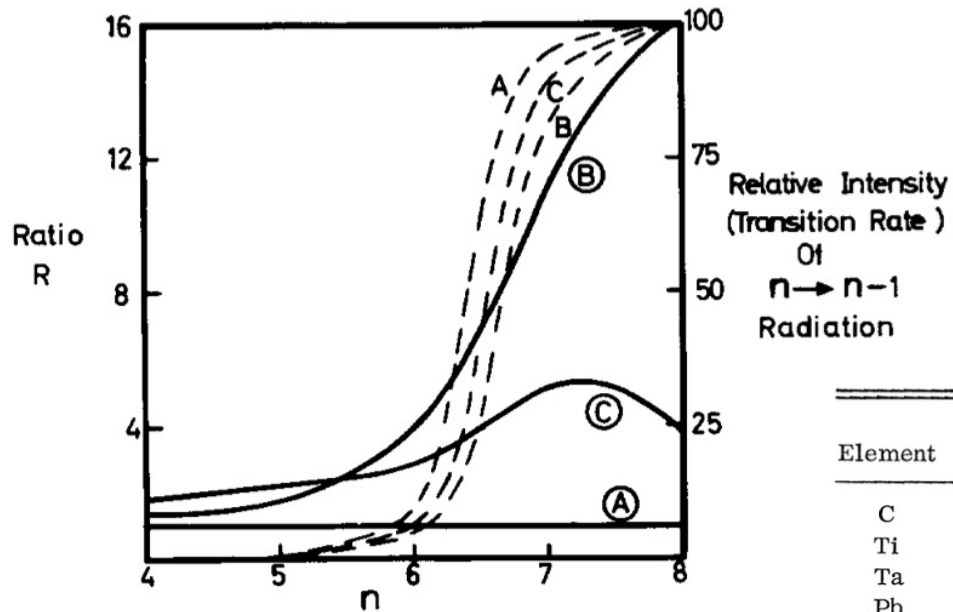


Alexander von
HUMBOLDT
STIFTUNG



HALOS AND SKINS: PRE-HISTORY

« Neutron atmosphere »
from kaon absorption



Burhop, NPB (1967)

« Neutron halo »
from antiproton annihilation

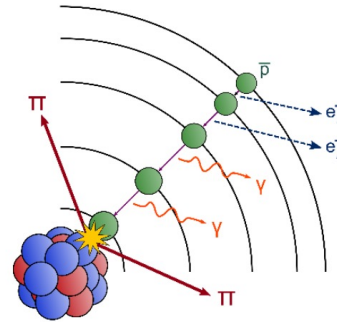
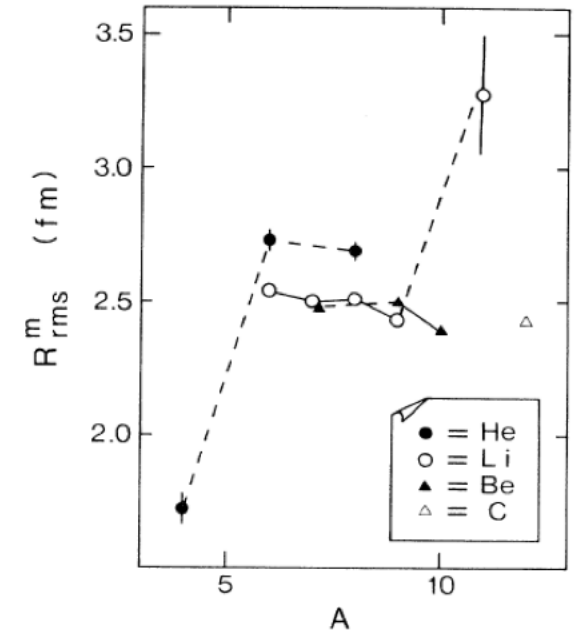


TABLE IV. "Halo factor" analysis.

Element	$N(\pi^-)$ $-N(\pi^+)$	$N(\bar{p}n)$	$N(\bar{p}p)$	$\frac{N(\bar{p}n)}{N(\bar{p}p)}$	$\frac{N(\bar{p}n)}{N(\bar{p}p)} \Big _c$	$\frac{N}{Z}$	Halo factor
C	2302	2586	4089	0.632	1.00	1.00	1.00
Ti	881	1067	1111	0.960	1.52	1.18	1.29 ± 0.21
Ta	1006	1276	931	1.371	2.17	1.48	1.46 ± 0.24
Pb	947	1216	534	2.270	3.59	1.54	2.34 ± 0.50

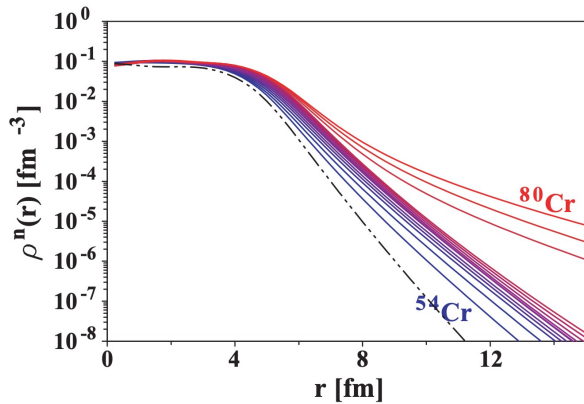
Buggs et al., PRL (1973)
followed up at LEAP, CERN, 90s

« Neutron halo »
from interaction cross section

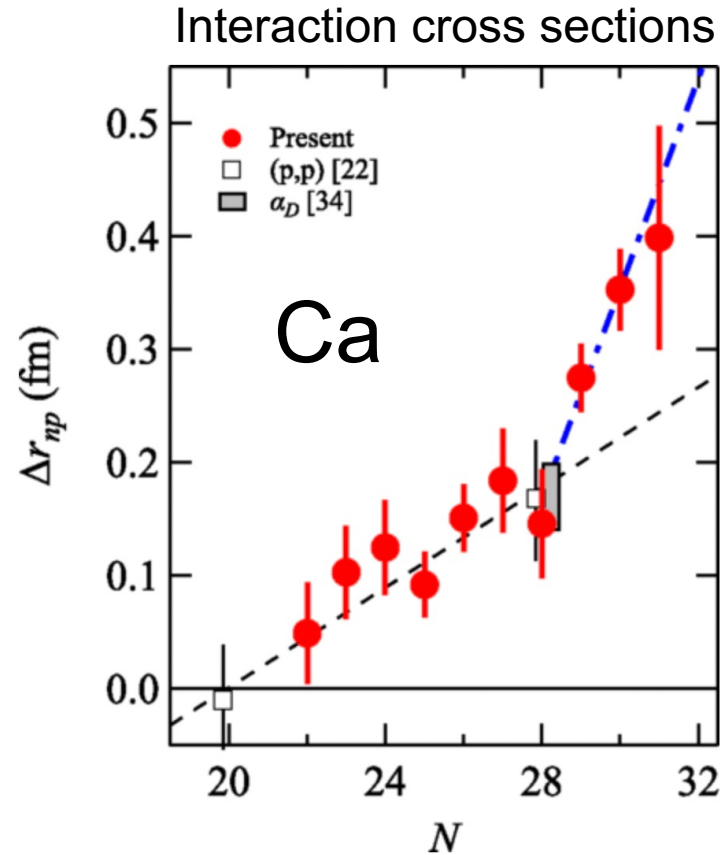


Tanihata et al., PRL (1985)

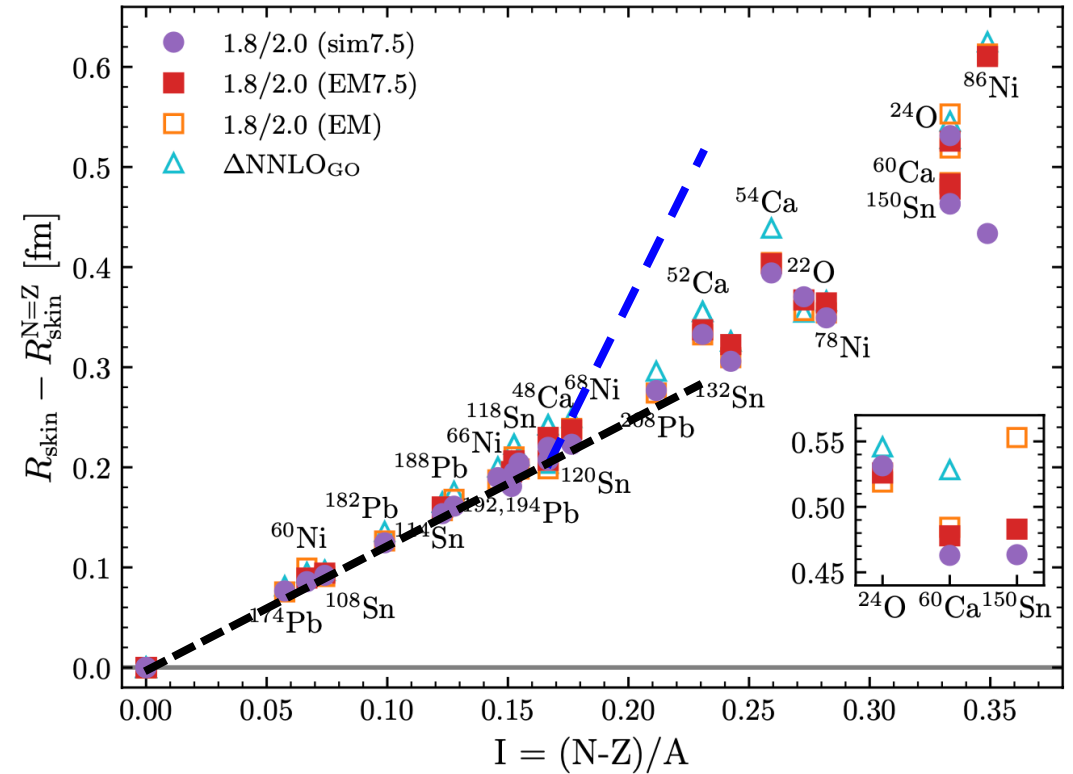
HALOS & SKINS IN MEDIUM-MASS NUCLEI



Rotival, Duguet, PRC (2009)



Tanaka et al., PRL (2020)



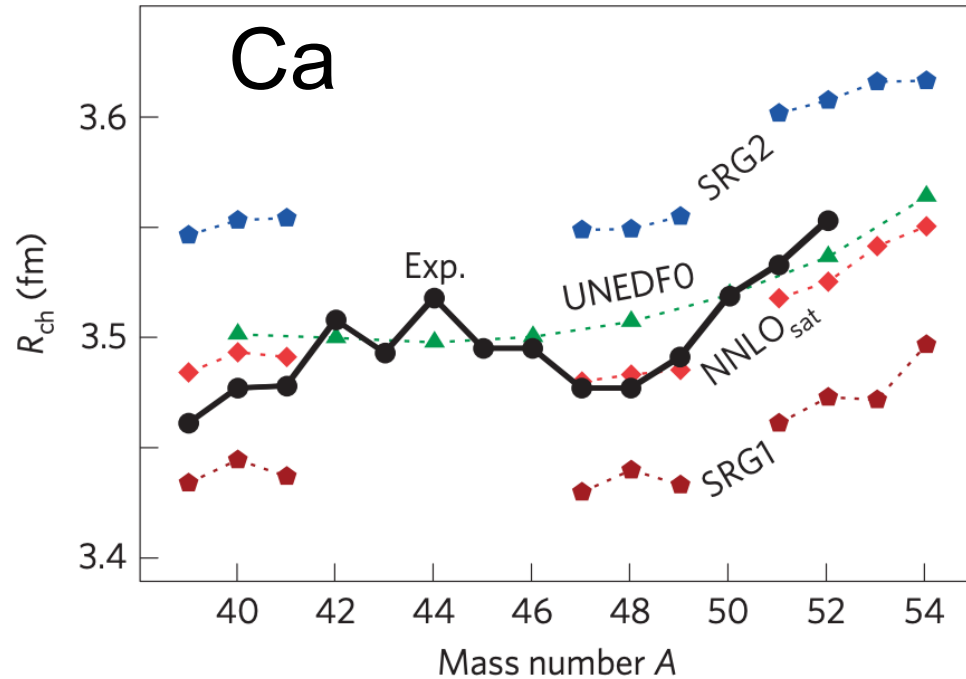
From Arthuis, Hebeler, Schwenk, arXiv (2024)

OUTLINE

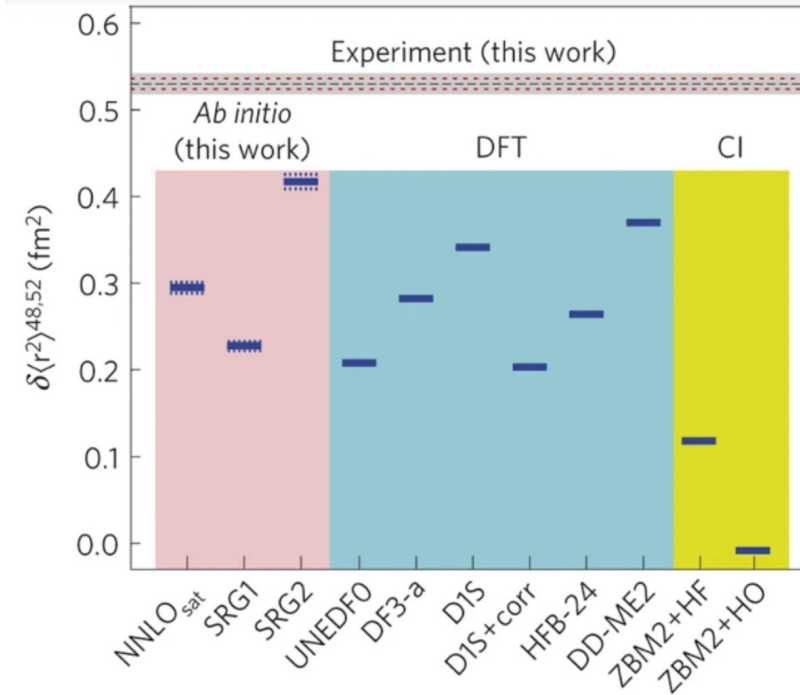
- **Sensitivity of quasifree scattering observables to the size of Ca isotopes**
Enciu et al., PRL (2022); Enciu et al., work in progress (2024)
- **PUMA: antimatter to probe the radial density tail**
Aumann et al., EPJA (2022); Fischer et al., NIMB (2024); Schlaich et al., IJMS (2024)

CHARGE RADII OF CA ISOTOPES

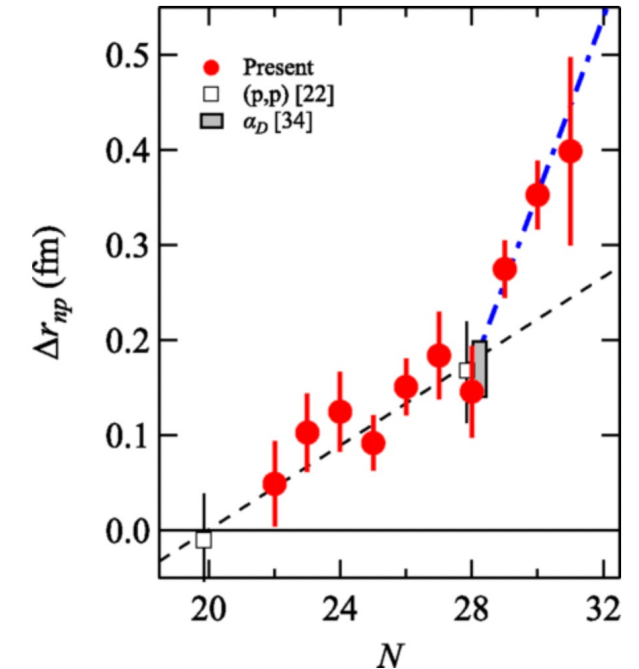
Ca, stable and short lived



Garcia Ruiz et al., Nat. Phys. (2016)



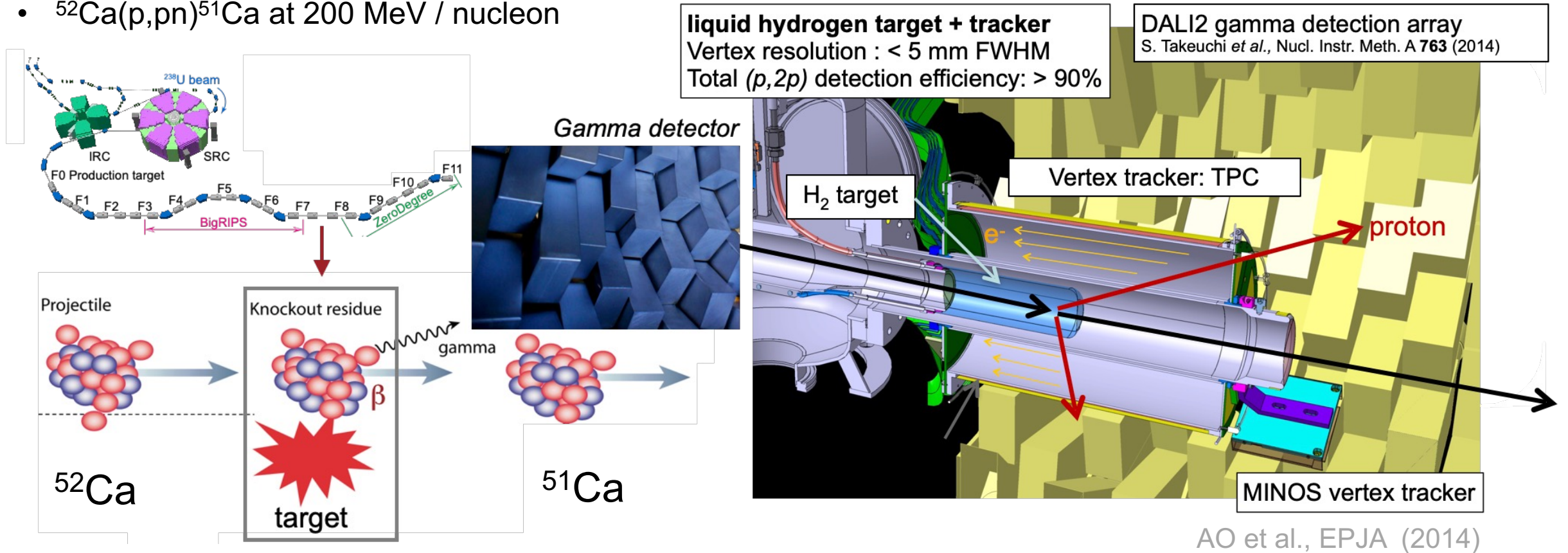
Tanaka et al., PRL (2020)



- electron scattering not yet possible for most RI (**see talk by Suda**)
- Isotopic shifts from laser spectroscopy give access to charge radii
- So far unexplained trend for neutron rich Ca and K isotopes
- Role of neutron distribution suggested to play a role [Bonnard et al., PRL \(2016\)](#)

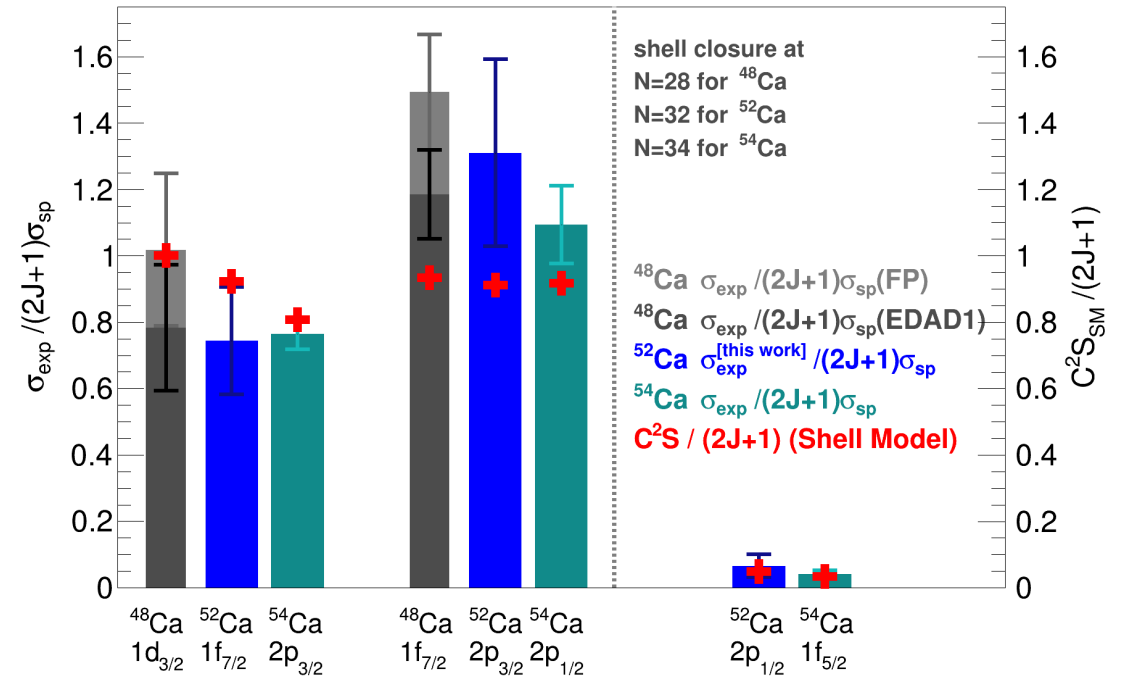
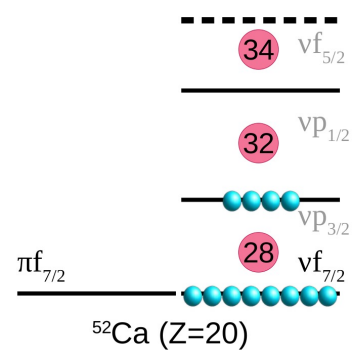
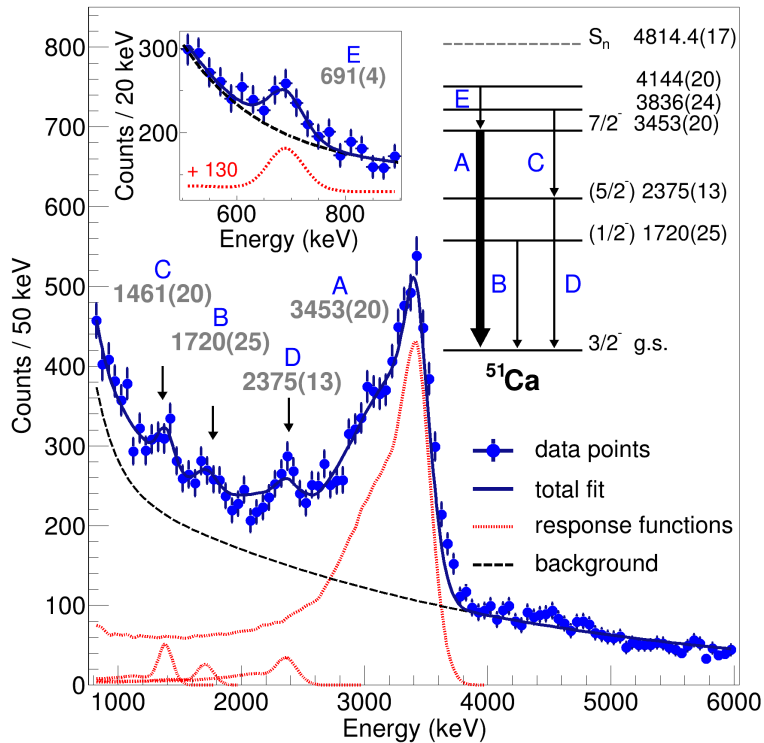
NEUTRON STRUCTURE OF ^{52}Ca

- RIBF, RIKEN
- ^{70}Zn primary beam at 345 MeV/nucleon, 200 pA
- $^{52}\text{Ca}(p,pn)^{51}\text{Ca}$ at 200 MeV / nucleon



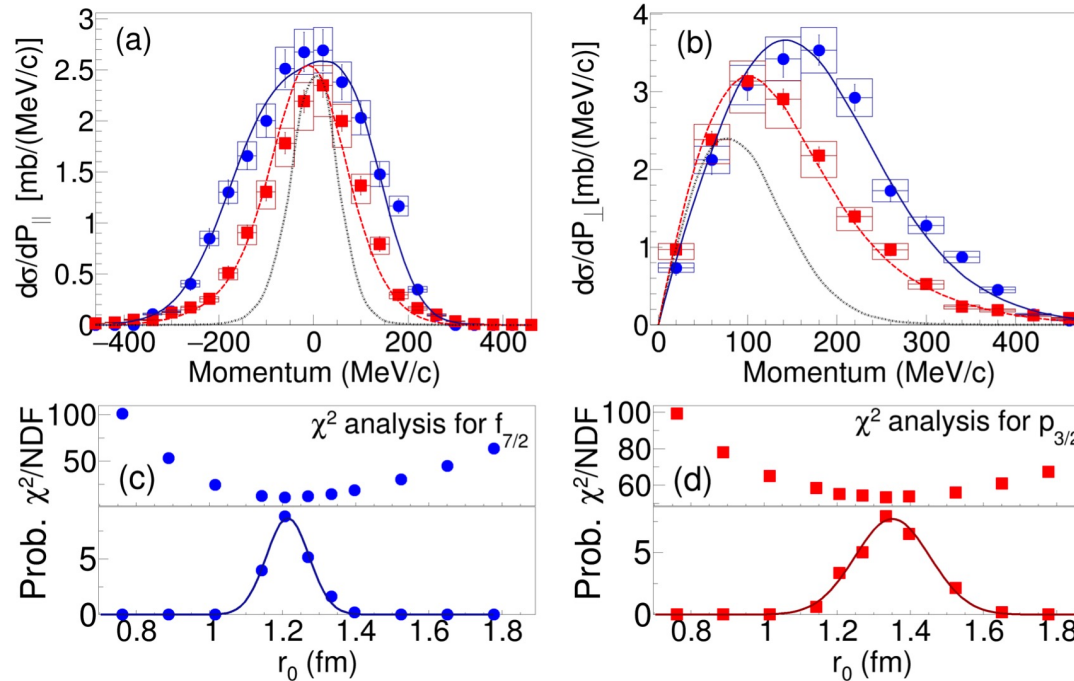
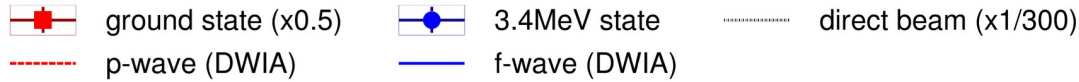
N=32 SHELL CLOSURE IN ^{52}Ca

- DWIA single-particle cross sections: K. Ogata, K. Yoshida
- ^{52}Ca presents the features of a neutron closed shell as ^{48}Ca , ^{54}Ca
- proton closed shell as well Sun et al., PLB (2020)



Enciu et al., PRL (2022)

EXCLUSIVE MOMENTUM DISTRIBUTIONS



- Momentum distributions of $\nu p_{3/2}$ and $\nu f_{7/2}$ single particle states, K. Yoshida (JAEA) and K. Ogata (Kyushu)

- DWIA transition amplitude:

$$T^{jlm} = \sqrt{S} \left\langle \chi_{p'}^{(-)} \chi_N^{(-)} \left| \tau_{pN} \right| \chi_p^{(+)} \psi_{jlm} \right\rangle$$

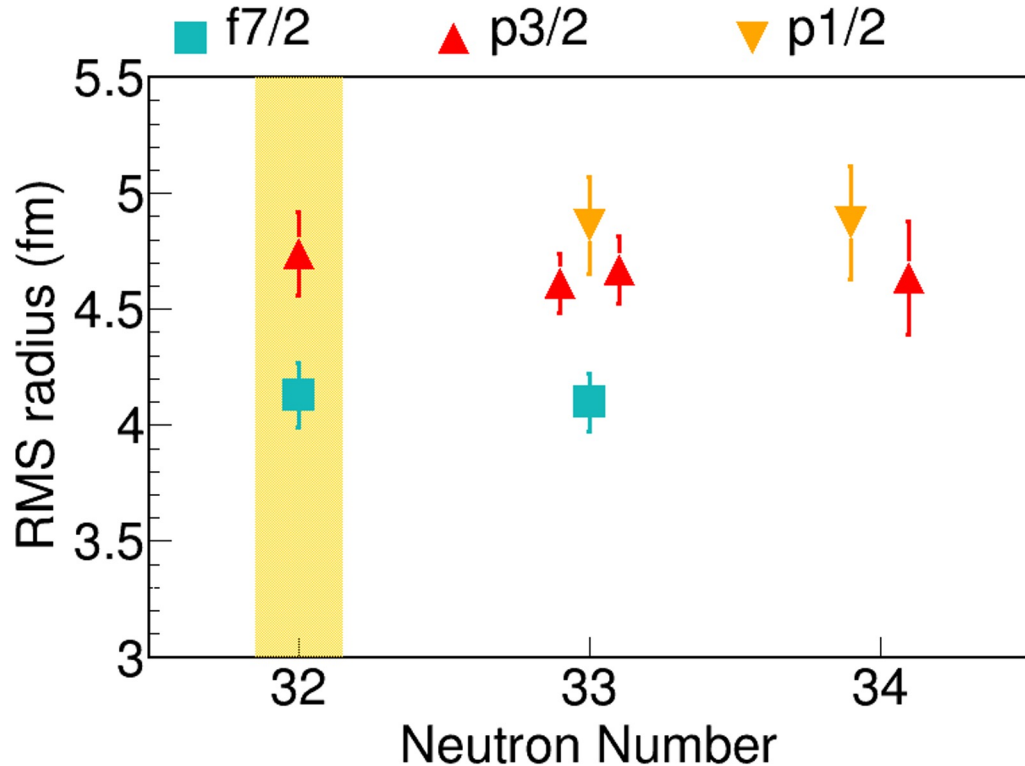
Distorted wavefunctions of the outgoing (-) and incoming (+) protons and knocked-out nucleon and transition form factor $\langle \psi_f | a_{nlj} | \psi_i \rangle \sim$ s.-p. wavefunction of the knocked-out nucleon.

Model inputs

- W.-S. wavefunctions, parameters with S_n , r_0 , a_0
OR microscopic transition amplitude
- Optical potential



NEUTRON TRANSITION AMPLITUDE RMS



- ^{52}Ca

Resulting rms radii:
 $\nu f_{7/2}$: 4.13(14) fm
 $\nu p_{3/2}$: 4.74(18) fm
 Difference: 0.61(23) fm

from HFB, SKM:

$\nu f_{7/2}$: 4.12 fm

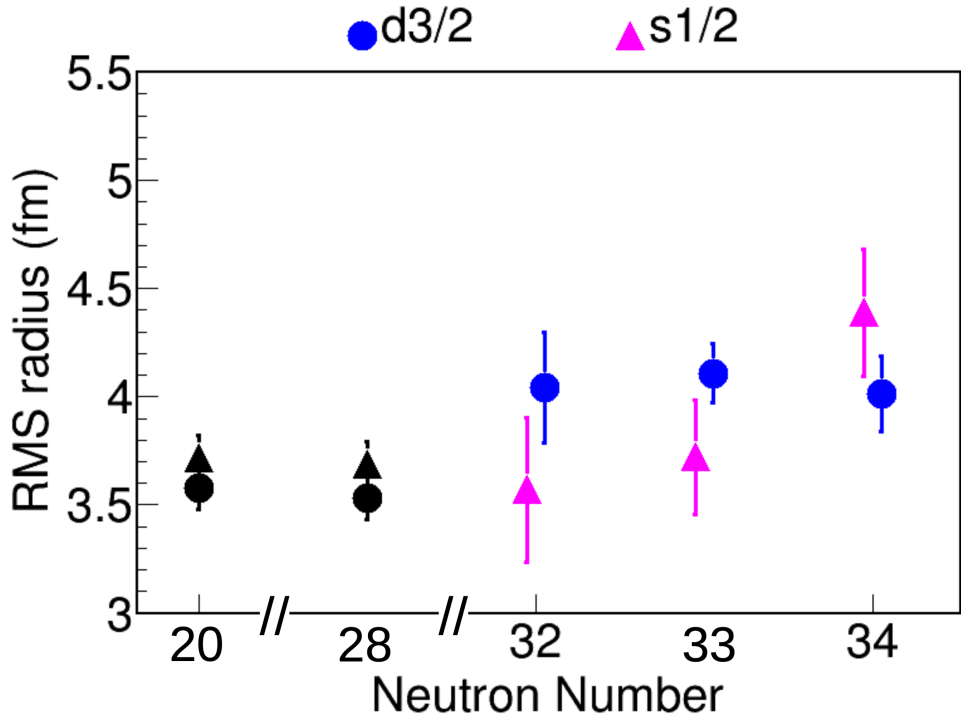
$\nu p_{3/2}$: 4.49 fm

Difference: 0.37 fm

- $^{53,54}\text{Ca}$: compatible results

- The approach gives comparable size for $\nu p_{3/2}$ and $\nu p_{1/2}$

PROTON TRANSITION AMPLITUDE RMS



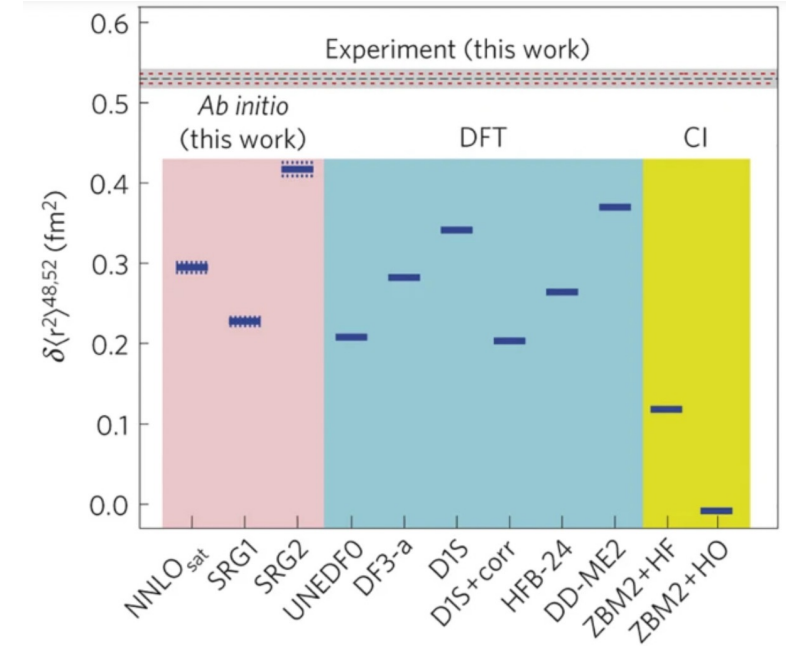
Assuming HF wavefunction and fixed proton orbitals up to $d_{5/2}$ (Si core), one gets:

$$\delta\langle r^2 \rangle^{48,52} = 0.76(63) \text{ fm}^2$$

Inconclusive due to

- systematics
- stat. uncertainties

New experiments accepted at RIBF for O, C, Ca isotopes (Nakamura, Obertelli)



⁴⁰Ca and ⁴⁸Ca data from:

G. J. Kramer, "The proton spectral function of ⁴⁰Ca and ⁴⁸Ca studied with the (e,e'p) reaction", PhD, Amsterdam Univ., 1990

COMPARISON TO THEORY

[preliminary, removed from slides]

[1] HF calculations, HFBRAD code, SKM interaction

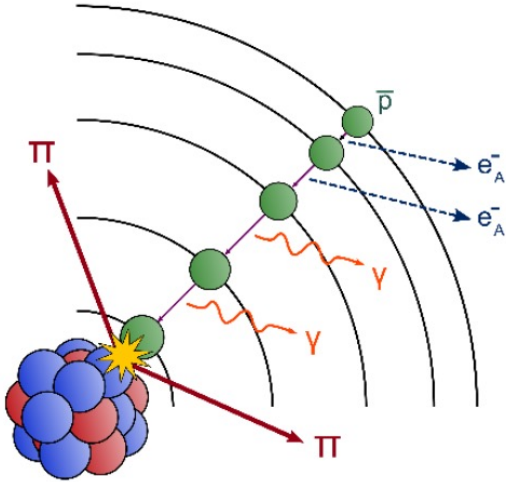
[2] W. Horiuchi, private communication
based on: W. Horiuchi et al., Phys. Rev. C (2020)

[3] J. Liu, private communication
based on: J. Liu et al., Phys. Lett. B (2020)
[Relativistic Hartree-Fock calculations]

IMSRG: **preliminary**, ongoing work
M. Heinz, T. Miyagi, A. Schwenk, A. Tichai

So far, unevolved one-body annihilation operator
Extension to transformed (many-body) operator ongoing

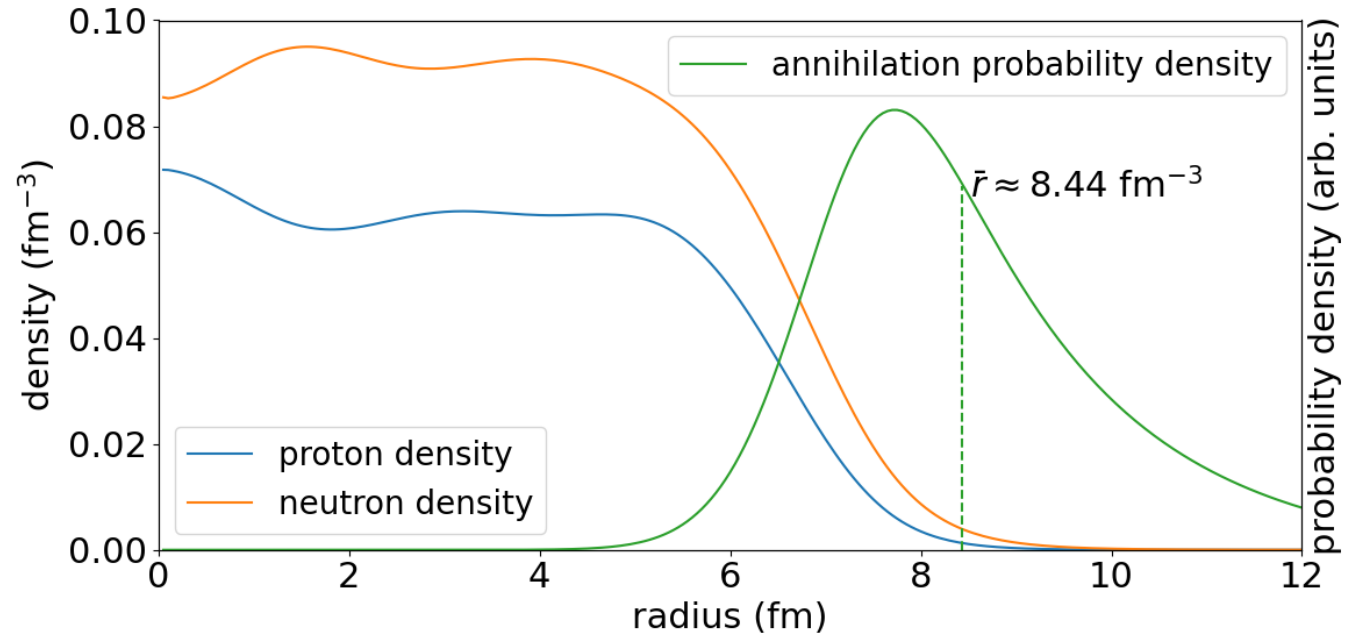
LOW ENERGY ANTIPROTONS AS PROBE



- \bar{p} captured in antiprotonic orbital (\sim QED)
- then annihilate in tail $\rho_{n/p}(r)$ (QCD)
- Conservation of total charge

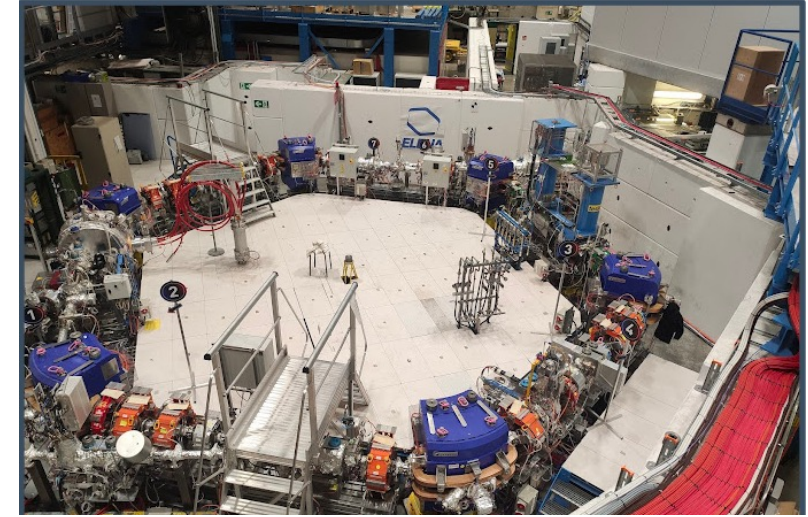
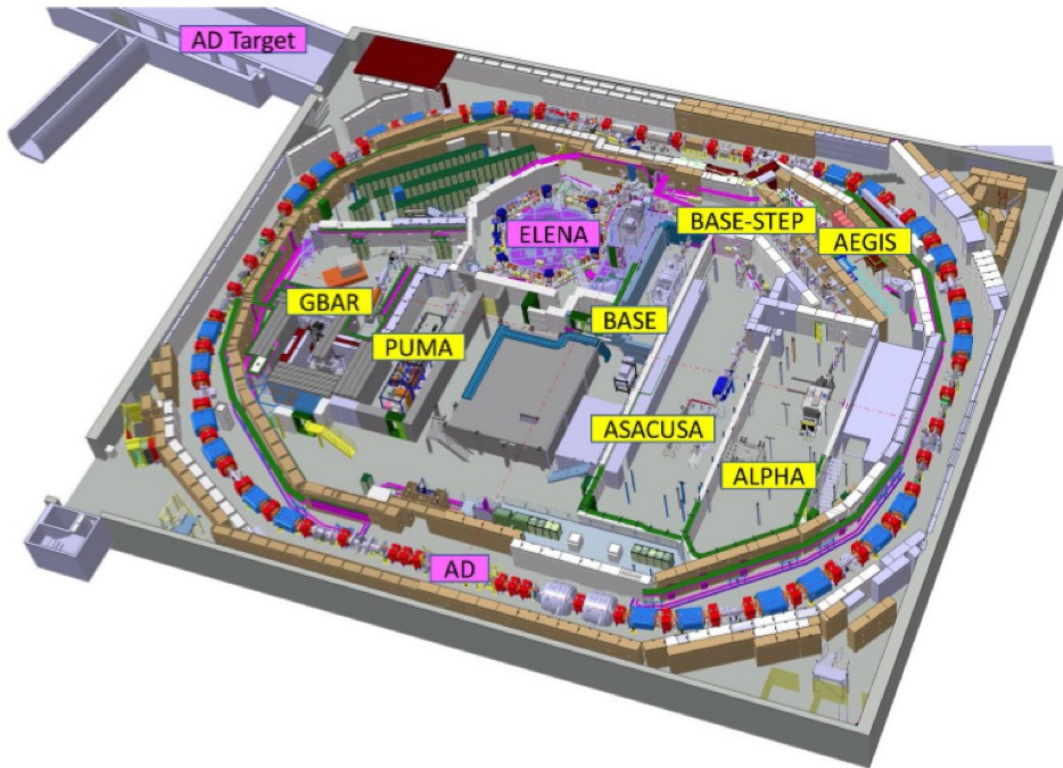
$$\sum_{\pi} q_{\pi} = \begin{cases} 0 & \text{for } \bar{p}p \\ -1 & \text{for } \bar{p}n \end{cases}$$

- Collaboration with Lazauskas, Hupin et al. (theory)

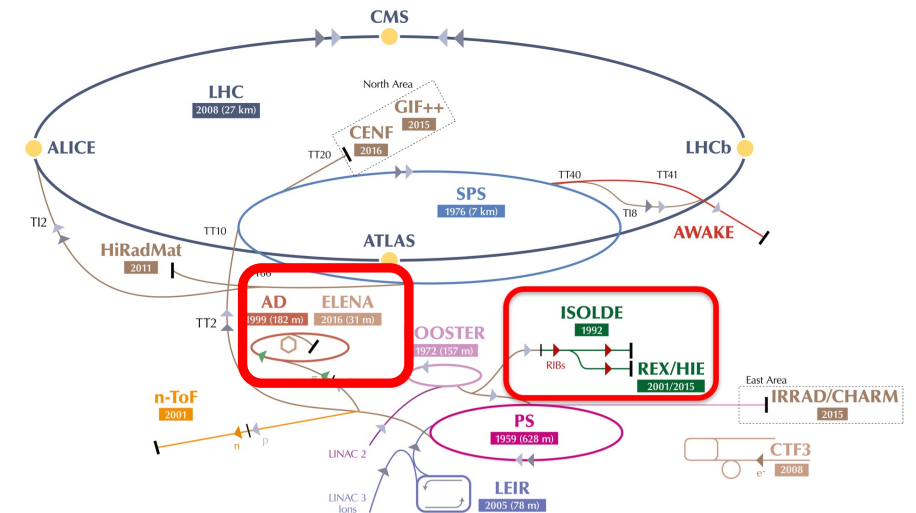


- First application of method: Bugg et al., PRL (1973)
- Application to rare isotopes first proposed by: Wada and Yamazaki, NIM B (2004)
- First experiments at PUMA, CERN, in 2025

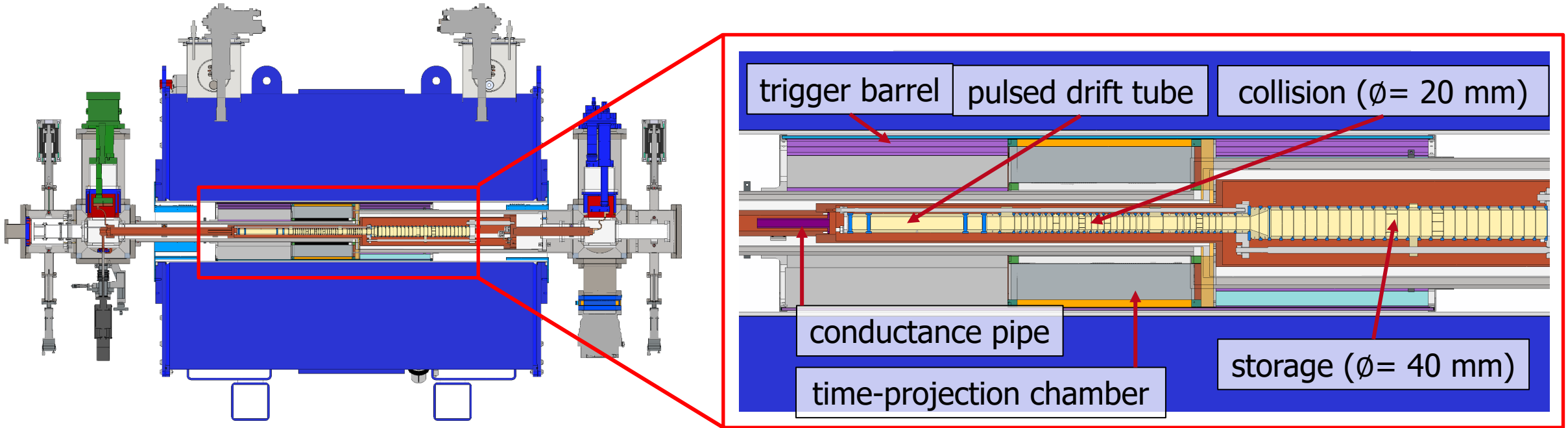
THE ANTIMATTER FACTORY AT CERN



- $1.5 \cdot 10^{13}$ protons at 26 GeV/c from PS
- $3 \cdot 10^7$ antiprotons at AD
- Deceleration: 5.3 MeV (AD), 100 keV (ELEN A)
- Experiments: $4 \cdot 10^6$ antiprotons every 110 s



THE PUMA EXPERIMENT



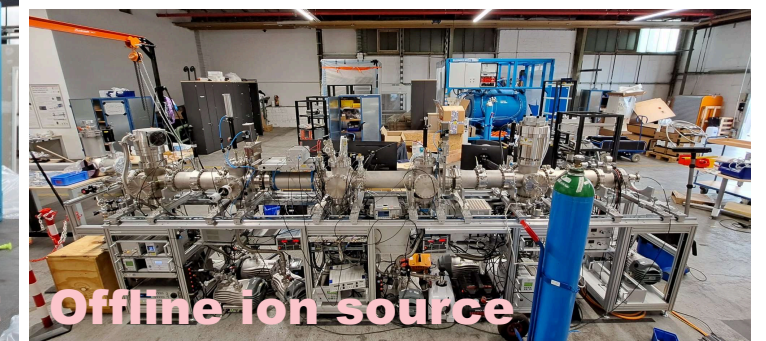
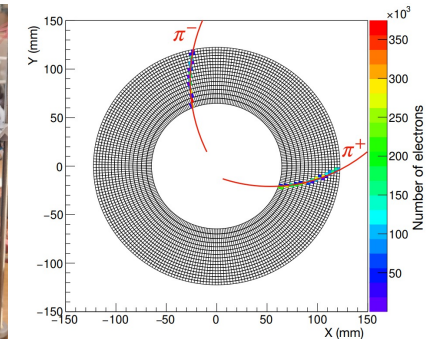
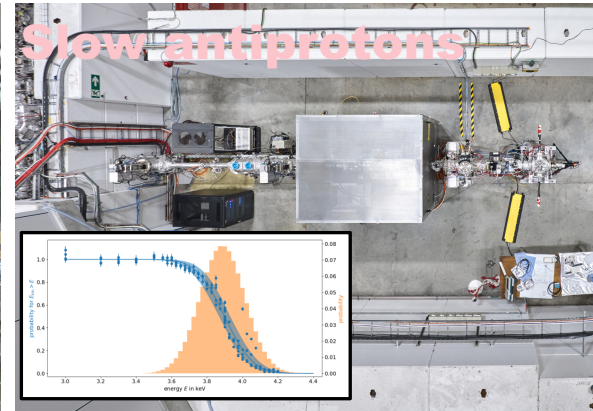
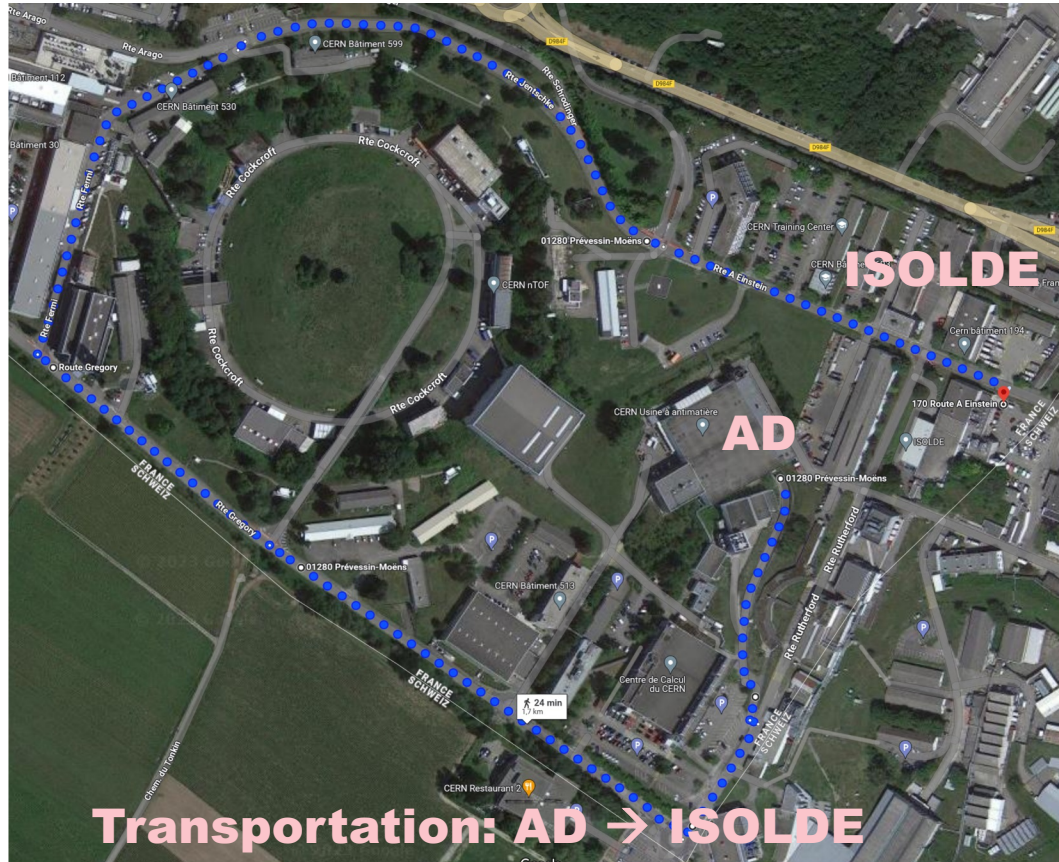
Objectives:

- store few 10^7 antiprotons (stage 1), 10^9 antiprotons (stage 2, $\equiv 10^7$ antiprotons / cm^3)
- Injection bunches of 10^{4-5} ions

Expected yields:

10^5 ions and 10^7 antiprotons in collision trap = 100 Hz of annihilations

THE PUMA EXPERIMENT



- Long antiproton trapping time already achieved. *Record*: BASE: > 50 years
- Transportation of antiprotons is also part of BASE-STEP (PI: C. Smorra)

FINAL STATE INTERACTIONS

M. Wada, Y. Yamazaki, NIMB (2004)

$M \setminus \Sigma_c$	-5	-4	-3	-2	-1	0	+1	+2	+3	+4	
0	0	0	0	0	0	0	0	0	0	0	(11386)*
1	0	0	0	0	17223	0	11233	0	0	0	28456
2	0	0	0	7530	0	21437	0	2844	0	0	31811
3	0	0	1029	0	11901	0	6591	0	179	0	19700
4	0	44	0	1904	0	4394	0	519	0	5	6866
5	1	0	99	0	979	0	451	0	13	0	1543
6	0	2	0	75	0	133	0	14	0	0	224
7	0	0	1	0	7	0	3	0	0	0	11
8	0	0	0	1	0	1	0	0	0	0	2
9	0	0	0	0	1	0	0	0	0	0	1
	1	46	1129	9510	30111	25965	18278	3377	192	5	88612

[preliminary, removed from slides]

* unobservable events

- 4-parameter model for FSIs S. Wychech
 ω^+, ω^- : absorption of $\pi^+, \pi^-, \lambda^+, \lambda^-$: charge exchange π^+, π^-, π^0
- Neural network Kubota et al., in preparation (2024)
Precision ~ 9%, Accuracy ~ 6 % benchmarked against INC simulations
- Room to improve performances

COLLABORATIONS

SAMURAI



PUMA



Theory (Ca isotopes) TUDa: M. Heinz, T. Miyagi, A. Schwenk, A. Tichai, JAEA: K. Yoshida, Kyushu: K. Ogata

SUMMARY AND PERSPECTIVES

- Investigation of **momentum distributions from quasifree-scattering** as observable sensitive to the spatial distribution of nucleons
Enciu et al., PRL (2022) & work in progress (2024)
- Low-energy antiprotons to measure the **ratio of neutron-to-proton annihilations**
PUMA collab., EPJA (2022); Kubota et al., in preparation (2024)
- Both observables require a reaction framework to extract nuclear structure theory
- Despite this limitation they may give access to information along isotopic chains not accesible otherwise

