

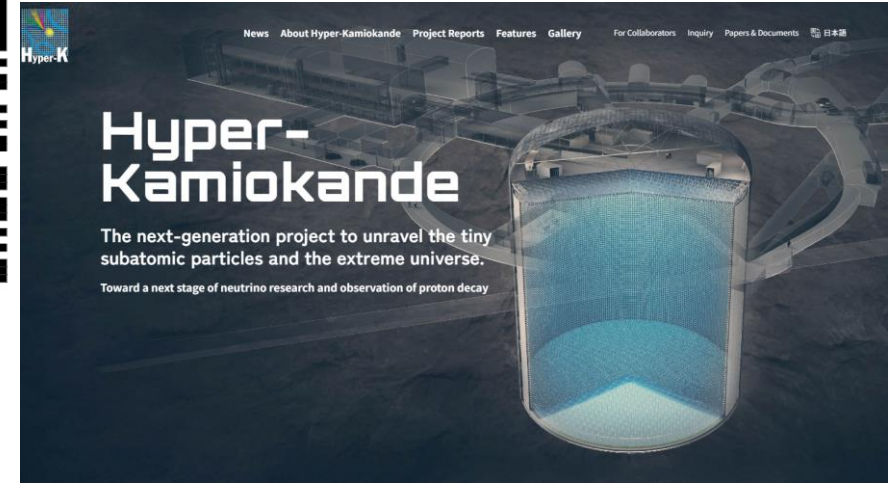
Status of the Hyper-K Experiment, and Mexican contributions so far

**Saul Cuen-Rochin (Tecnológico de Monterrey)
on behalf of the Hyper-K collaboration and Mexican group**

**The International Symposium on Very High Energy Cosmic Ray Interactions (ISVHECRI) 2024
Puerto Vallarta, Mexico
8 – 12 July 2024**

Agenda

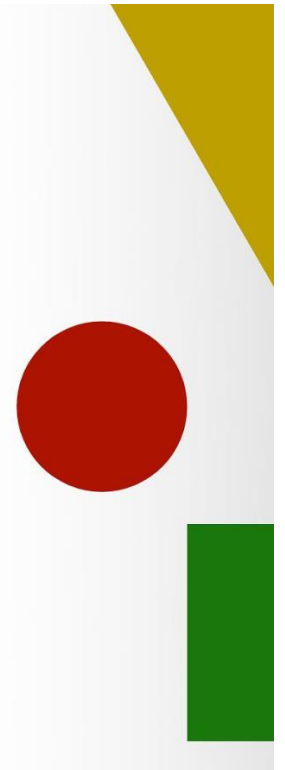
- Hyper-K project goals and status
- Work at local national institutions
 - Neutrino classification
 - Multi-PMT detector design and manufacturing



https://bio.site/Outreach_HK_MX

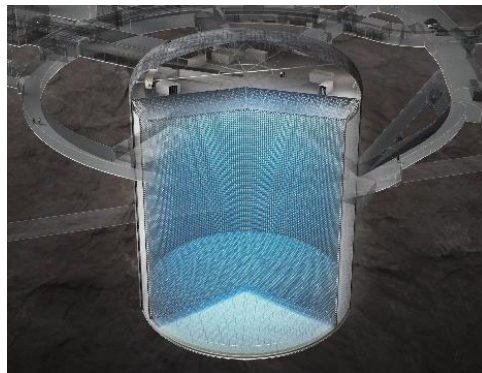
Hyper-Kamiokande Outreach MX Hyper-Kamiokande Outreach MX

SEÑOR		1	
NEUTRINO la partícula fantasma			
▶▶▶▶▶		ニュートリノ (neutrino)	
 tiene una velocidad cercana a la de la luz			
 tiene una masa muy pequeña			
 son muy difíciles de detectar			
electrón neutrino			
muón neutrino			
tau neutrino			
ハイパーカミオカンデ (Hyper-Kamiokande)			



Hyper-Kamiokande Project

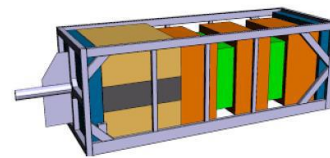
- The Hyper-Kamiokande project includes a far detector, a neutrino beam, and a neutrino near detector complex
 - Construct the Hyper-Kamiokande detector at Kamioka
 - Upgrade the J-PARC neutrino beam
 - Construct the Intermediate Water Cherenkov Detector (IWCD) at Tokai



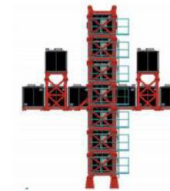
Hyper-Kamiokande detector
(Far detector)



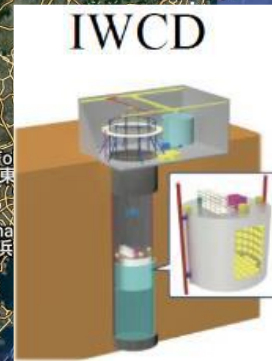
ND280



INGRID



IWCD



J-PARC



Three Generations of Water Cherenkov Detector in Kamioka

- **Kamiokande (1983 - 1996)**

- Atmospheric and solar neutrino “anomaly”
- Supernova 1987A

Birth of neutrino astrophysics

- **Super-Kamiokande (1996 - ongoing)**

- Proton decay: world best-limit
- Neutrino oscillation (atm/solar/LBL)
 - All mixing angles and Δm^2 s

Discovery of neutrino oscillations

- **Hyper-Kamiokande (2027 -)**

- Extended search for proton decay
- Precision measurement of neutrino oscillation including CPV and MO
- Neutrino astrophysics

Explore new physics



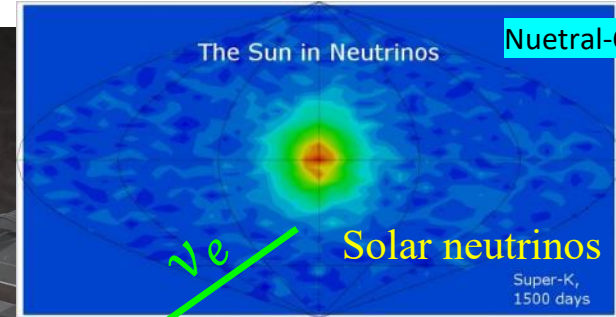
Hyper-K Observation Target



$$\nu_e + n \rightarrow e^- + p$$

$$\bar{\nu}_e + p \rightarrow e^+ + n$$

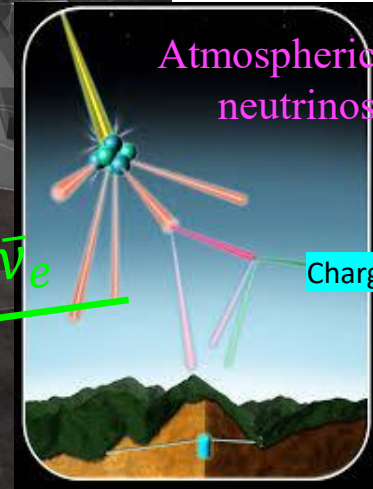
$\nu_e, \bar{\nu}_e, \nu_\mu, \bar{\nu}_\mu, \nu_\tau, \bar{\nu}_\tau$



Neutral-Current Interactions (NC)

$$\nu_e + e^- \rightarrow \nu_e + e^-$$

$$\nu + N \rightarrow \nu + N'$$



Charged-Current Interactions (CC)

$$\nu_e + n \rightarrow e^- + p$$

$$\bar{\nu}_e + p \rightarrow e^+ + n$$

$$\nu_\mu + n \rightarrow \mu^- + p$$

$$\nu_\tau + n \rightarrow \tau^- + p$$

$$\tau^- \rightarrow e^- + \bar{\nu}_e + \nu_\tau$$

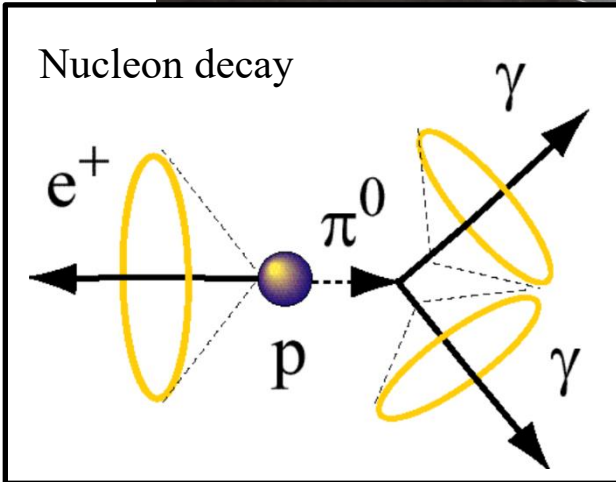
$$\tau^- \rightarrow \mu^- + \bar{\nu}_\mu + \nu_\tau$$

$$\tau^- \rightarrow \text{hadrons} + \nu_\tau$$

$$p \rightarrow e^+ + \pi^0$$

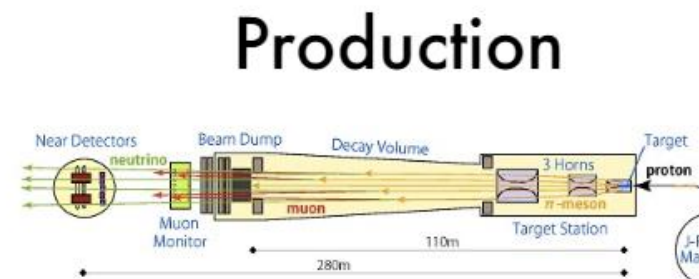
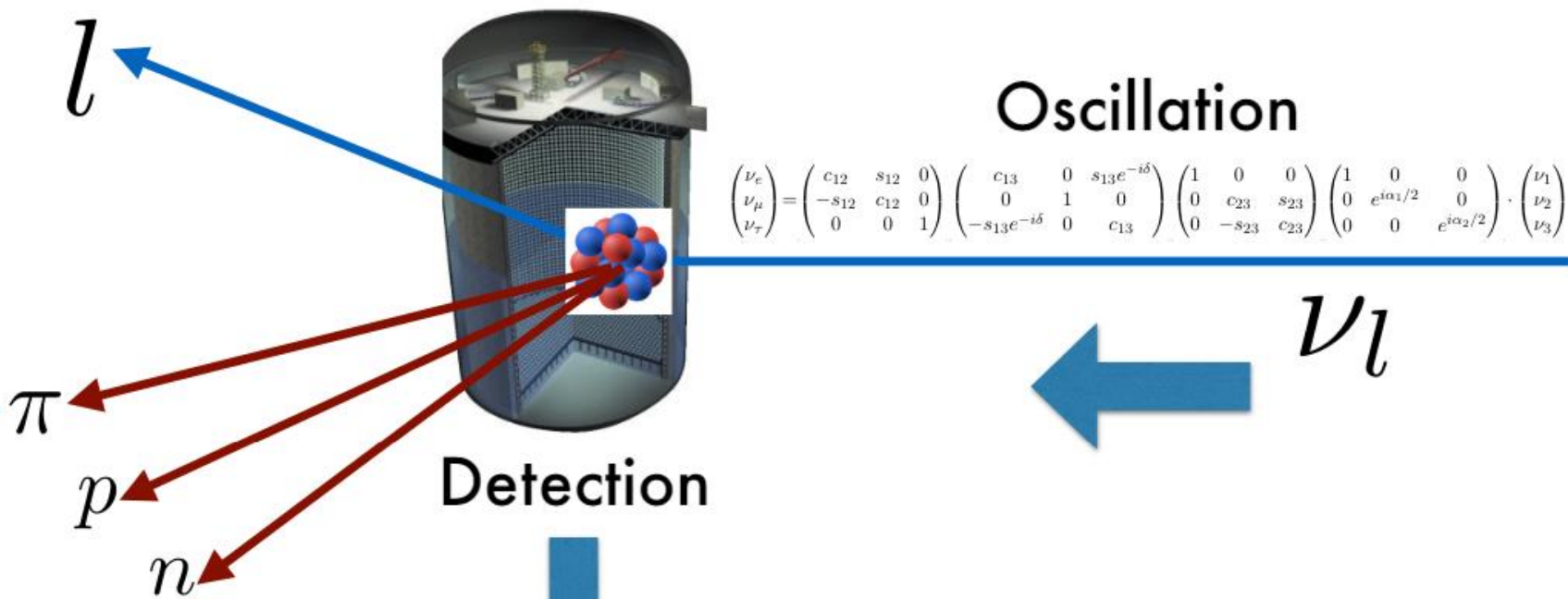
$$p \rightarrow \mu^+ + \pi^0$$

$$\pi^0 \rightarrow \gamma + \gamma$$



Experiment

Analysis



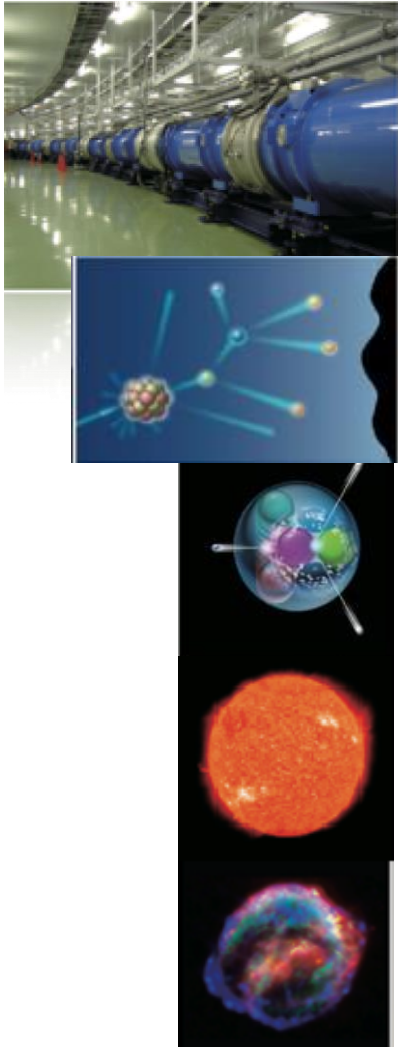
Cross section model
 $\sigma(E_\nu)$

Reconstruct

E_ν

Oscillation Parameters $\theta_e, \theta_\mu, \theta_\tau, \delta_{CP}$
 $\Delta m_e^2, \Delta m_\mu^2, \Delta m_\tau^2$

Hyper-K Target sensitivity

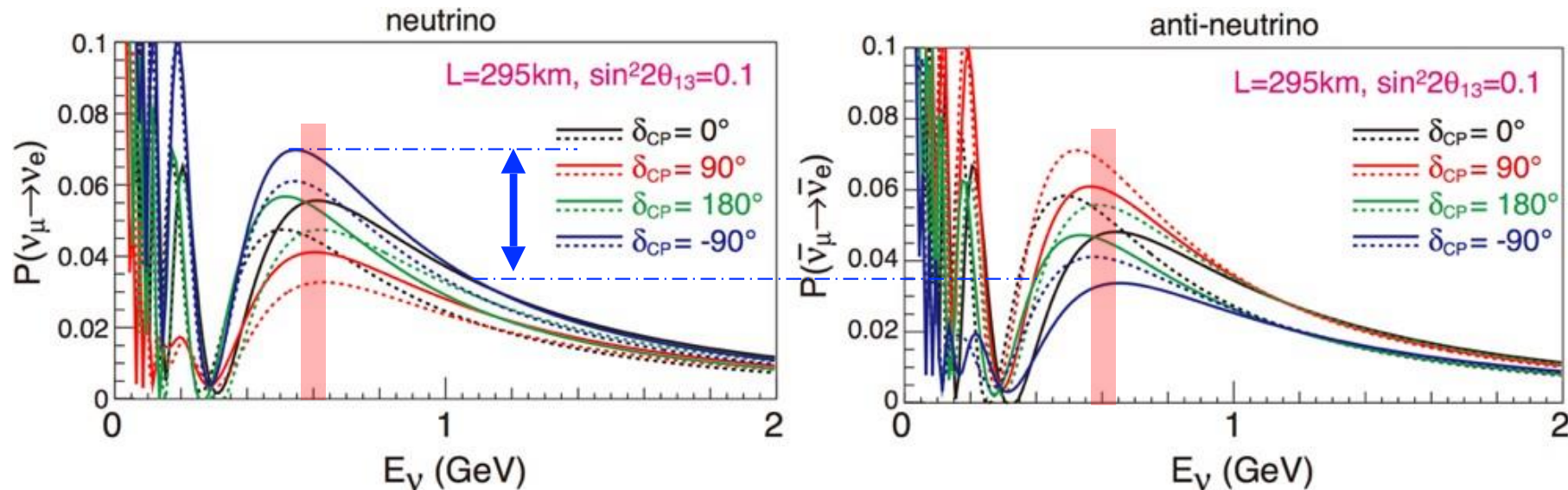


Physics category	Parameters	Sensitivity
LBL (1.3MW×10years)	δ precision	7° - 20°
	CPV coverage ($3/5\sigma$)	76%/58%
	$\sin^2\theta_{23}$ error (for 0.5)	± 0.017
ATM+LBL (10 years)	MO determination	$>3.8\sigma$
	Octant determination (3σ)	$ \theta_{23}-45^\circ >2^\circ$
Proton Decay (20 years)	τ for $e^+\pi^0$ (3σ)	1×10^{35} years
	τ for νK (3σ)	3×10^{34} years
Solar (10 years)	Day/Night (from 0/ from KL)	$8\sigma/4\sigma$
	Upturn	$>3\sigma$
Supernova	Burst (10kpc)	54k-90k
	Relic	70v's / 10 years

Long-baseline program with the J-PARC neutrino beam

Experimental setup

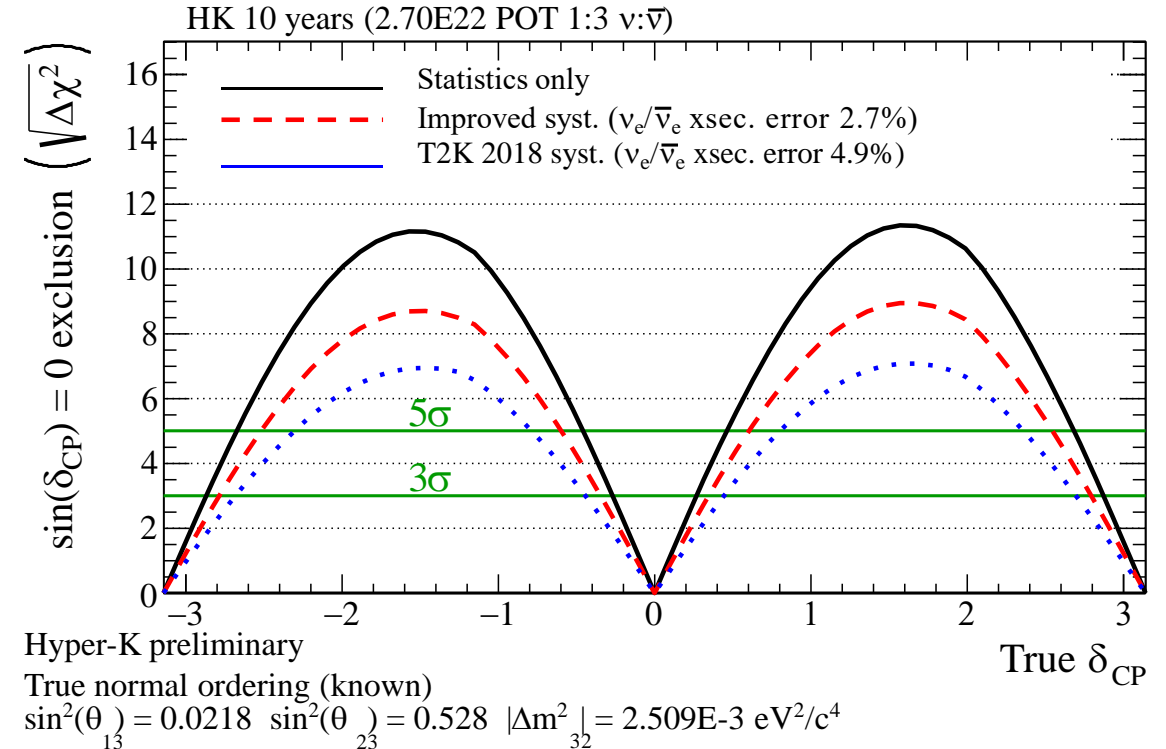
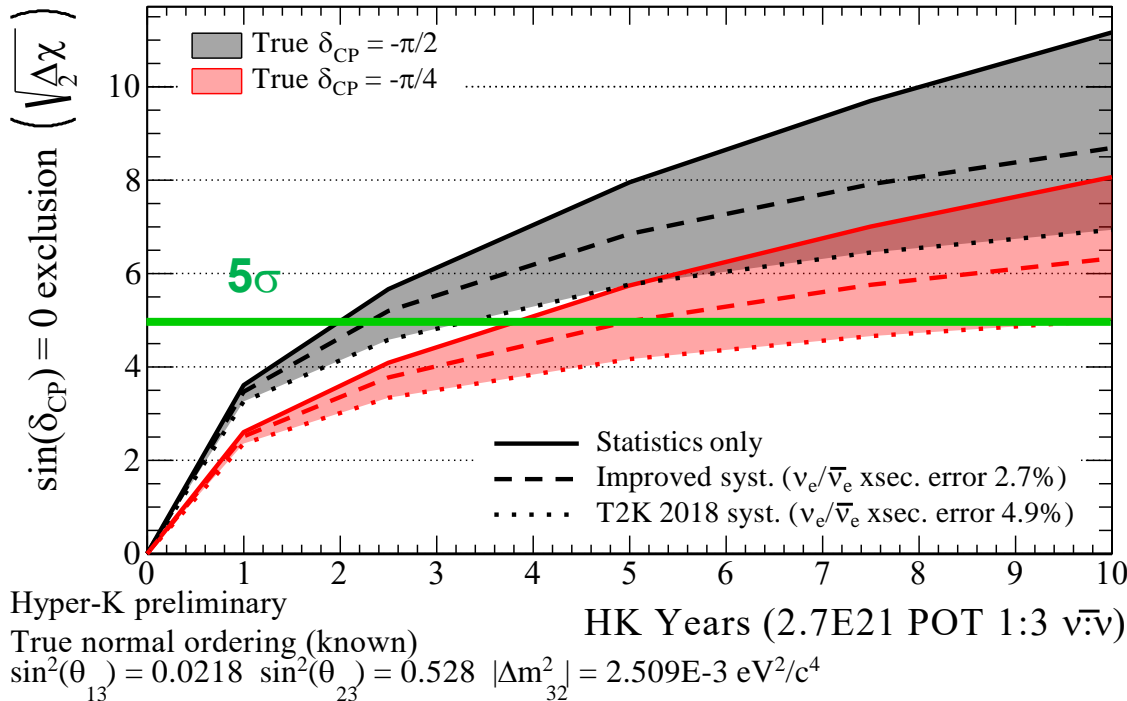
- 2.5° off-axis ν_μ and $\bar{\nu}_\mu$ beam peaked at 0.6 GeV (oscillation maximum at 295km)
 - Major interaction is QE: E_ν determined from (p, θ) of charged lepton
- Measures CP violation in neutrinos by comparing $P(\nu_\mu \rightarrow \nu_e)$ and $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$



- A few % statistical uncertainties after 10 years operation with $>1000 \nu_e$ and $\bar{\nu}_e$ signals

CP violation sensitivity

- Sensitivity CP violation with 1:3 ν : $\bar{\nu}$ beam



- With systematics and known mass ordering (MO): 2-3 years for 5σ sensitivity to exclude CP conservation for true $\delta_{CP} = -\pi/2$.
- After 10 years of operation, 60% of δ_{CP} values excluded at $> 5\sigma$

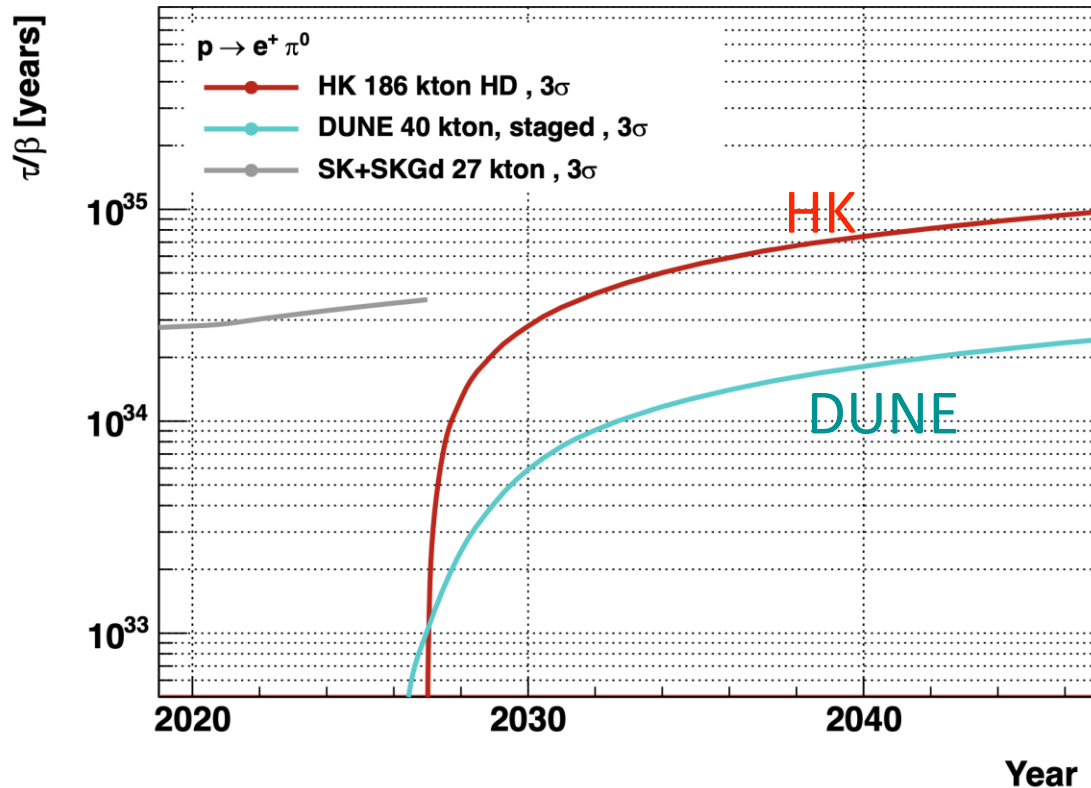
Nucleon decay search

- Nucleon decay is evidence of Beyond Standard Model (BSM) and Grand Unified Theories (GUT)
- Examples of proton decay sensitivity in two modes:

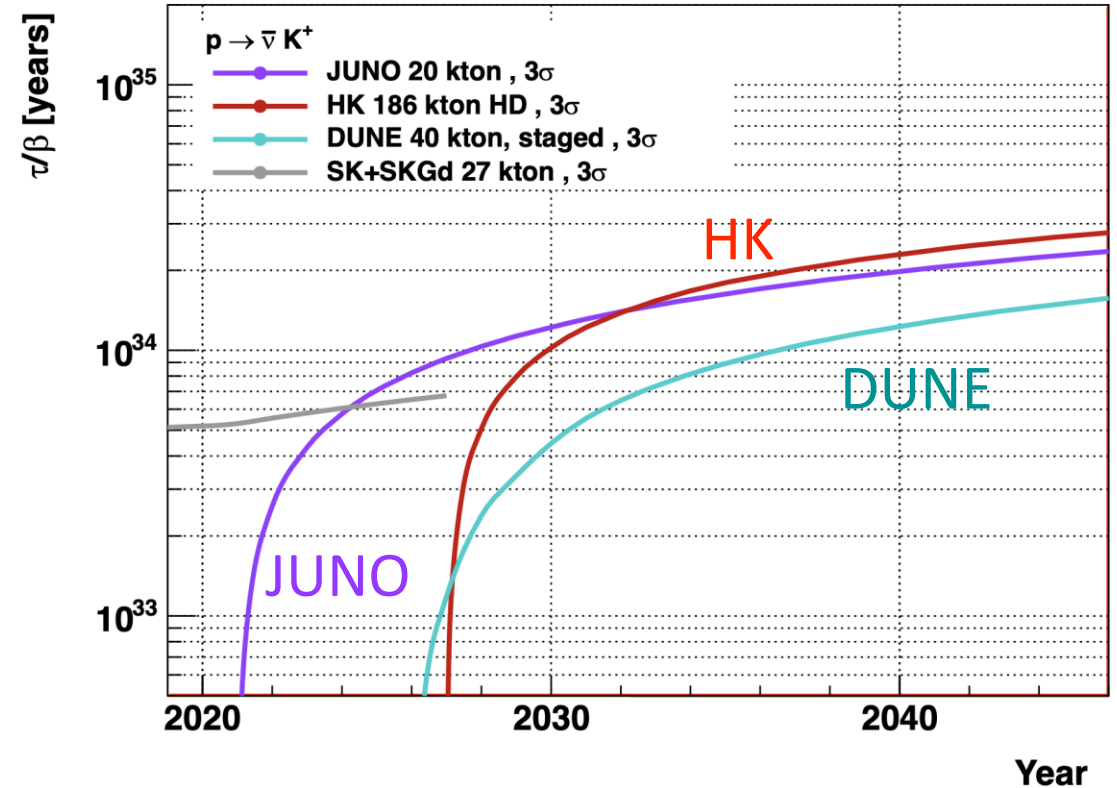
[HK] arXiv:1805.04163

[DUNE] arXiv:2002.03005

[JUNO] arXiv:1508.07166



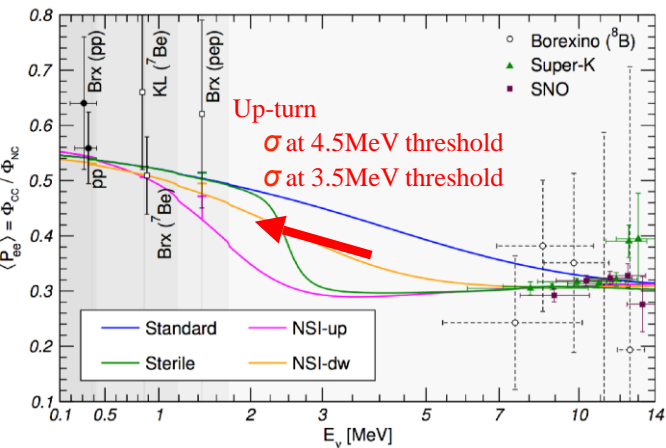
$\tau \sim 10^{35}$ years (3σ)



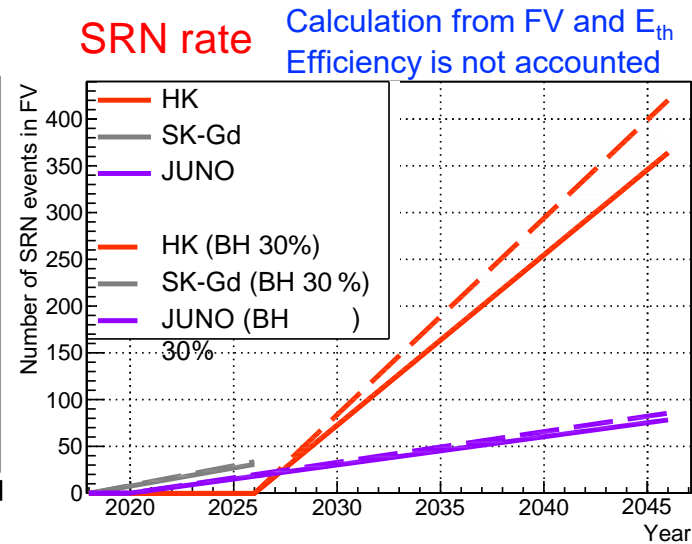
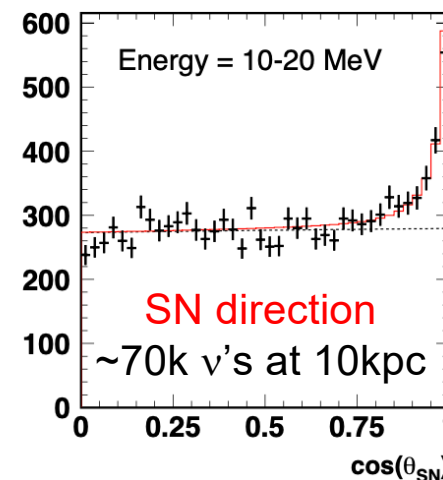
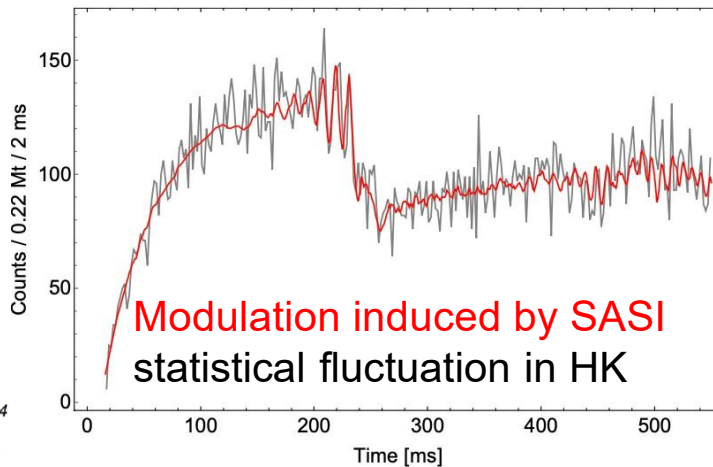
$\tau \sim 3 \times 10^{34}$ years (3σ)

Neutrino astrophysics

- Hyper-K is designed to be sensitive to neutrinos with energies starting from a few MeV, including time, energy and direction information. Unique role in multi-messenger observation
- **Solar neutrinos:** up-turn at vacuum-MSW transition, Day/Night asymmetry, hep neutrino observation
- **Supernova burst neutrinos:** explosion mechanism, BH/NS formation, alert with $\sim 1^\circ$ pointing
- **Supernova Relic Neutrinos (SRN):** stellar collapse, nucleosynthesis and history of the universe

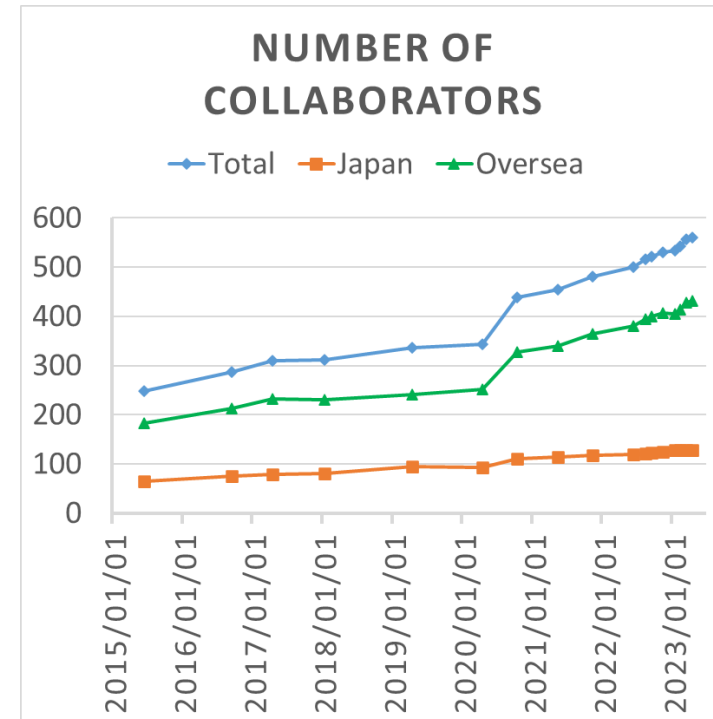


M. Maltoni et al., Phys. Eur. Phys. J. A52, 87 (2016)



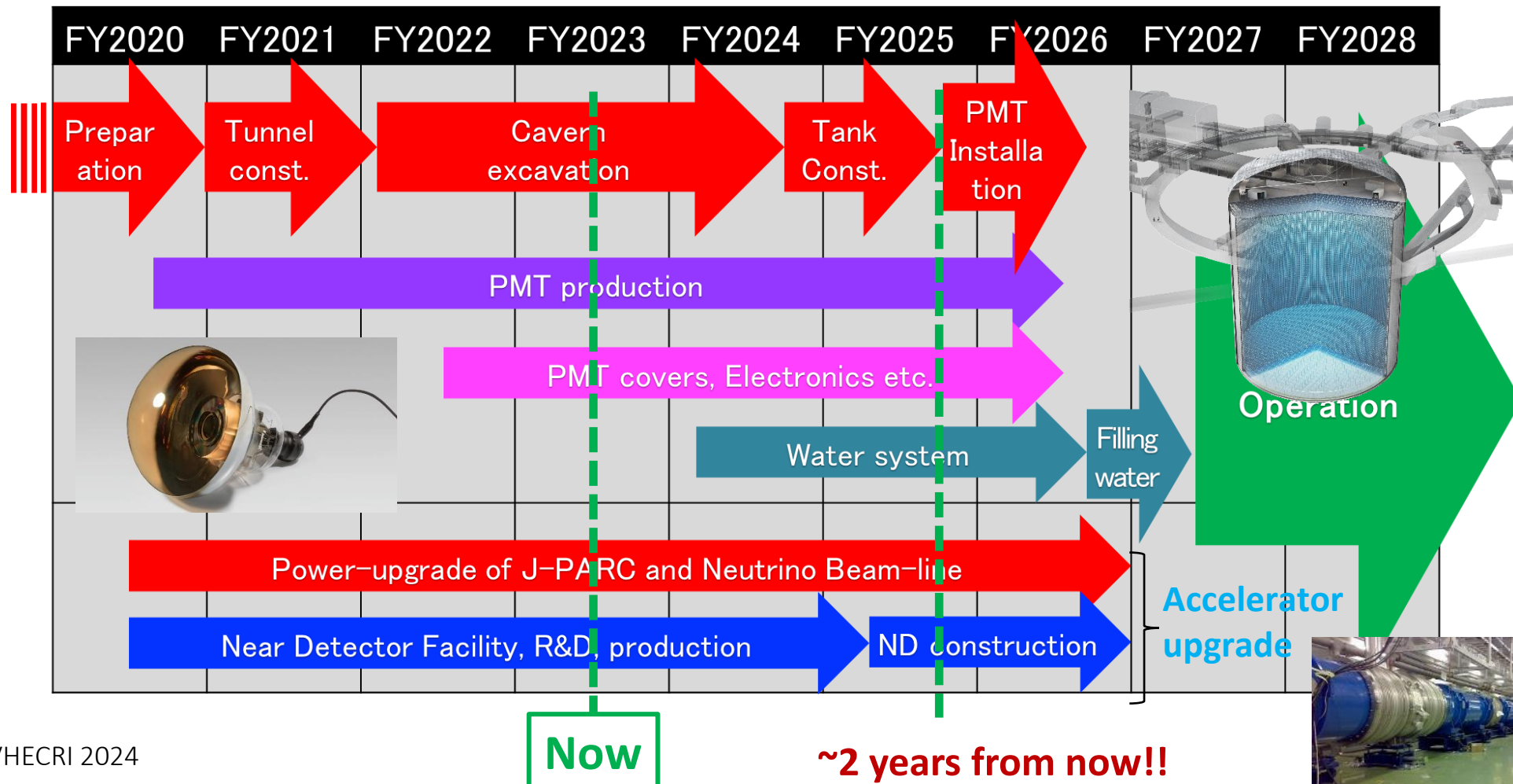
Hyper-Kamiokande Collaboration

- ~600 members located in 102 institutes from 22 countries
 - 25% Japanese / 75% non-Japanese
- Recently approved as a recognized experiment (RE45) at CERN
- March 2023:
our very 1st Collaboration meeting in person after COVID!

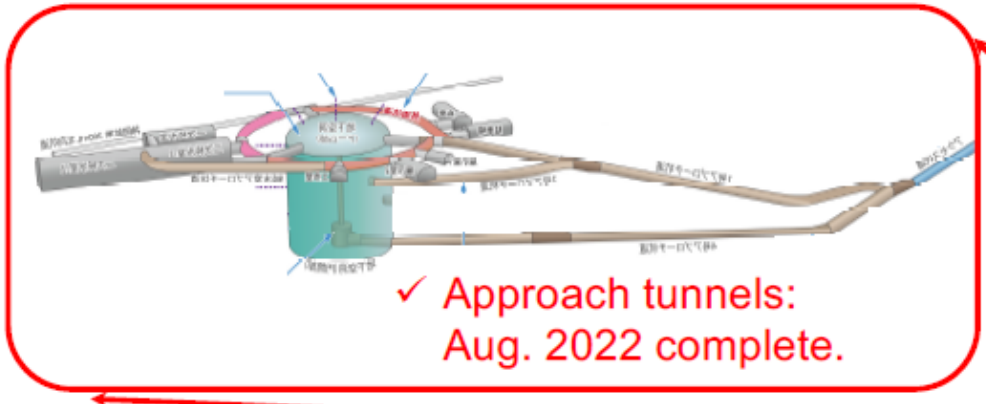


Hyper-K construction schedule

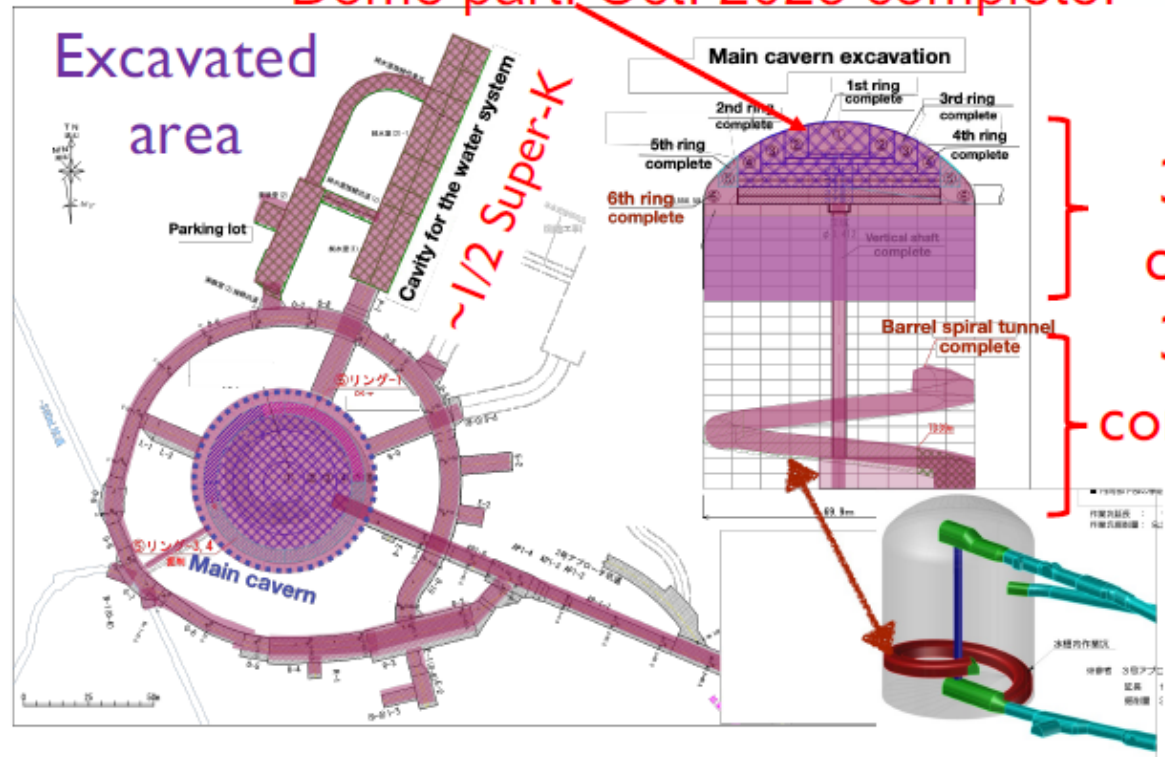
- The Hyper-K construction started in 2020 and will start operation in 2027.



Excavating the world's largest human-made cavern

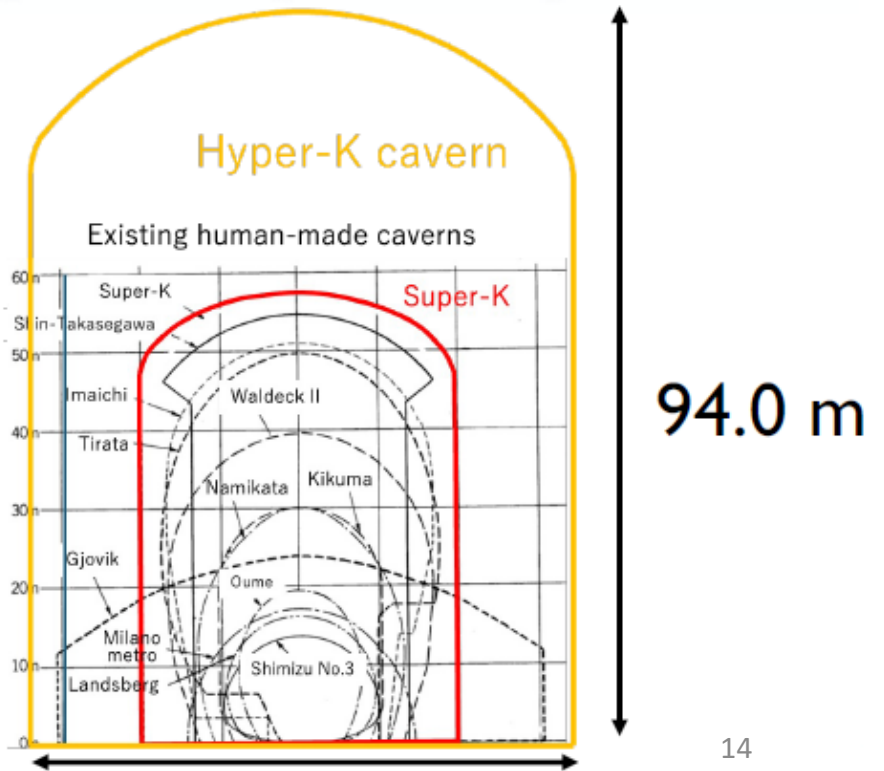


✓ Dome part: Oct. 2023 complete.



3 Super-K completed.
3 Super-K coming ~1/2yr

69.0 m



Hyper-K main cavern excavation



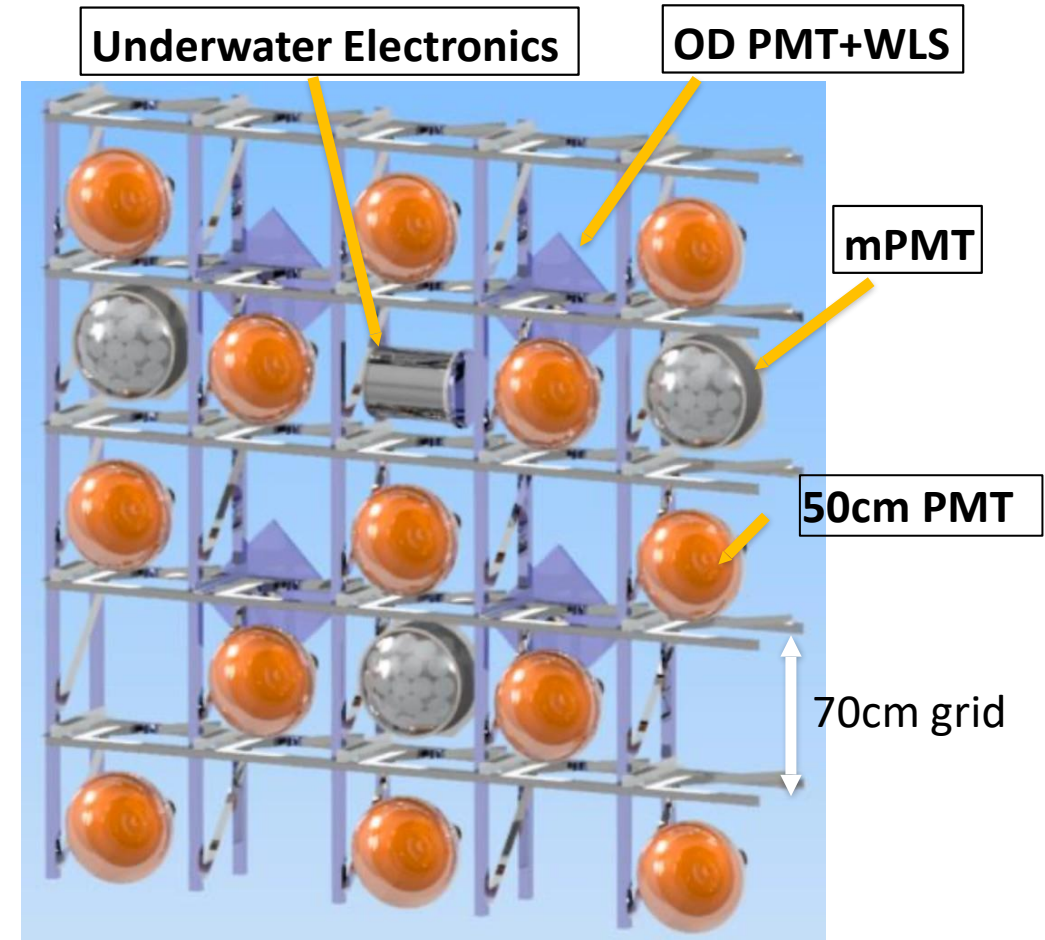
- **October 3, 2023:** Excavation of the dome section completed.
 - 69m diameter, 21m height
 - One of the largest human-made underground spaces.
- Now, the excavation of the barrel section is ongoing.



Excavation of the HK cavern will be completed by the end of this year!

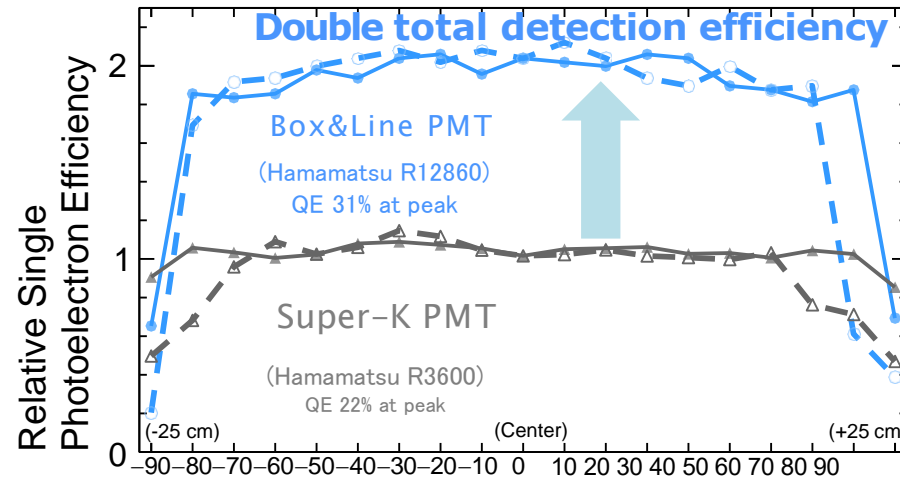
Hyper-K detector configuration

- **Inner Detector (ID)**
 - 20,000 – 20” PMT
 - 64.8m diameter, 65.8m height
 - 50cm PMTs will be installed
 - 800 multi-PMT modules (19 3” PMT each) will be integrated as hybrid configuration
- **Outer Detector (OD)**
 - 3,600 – 3” PMT
 - 1m (barrel) or 2m (top/bottom) thick
 - 3-inch PMT + WLS plate
 - Walls are covered with high-reflectivity Tyvek sheets
- **Under-water electronics**
 - Mitigate disadvantage of long cables



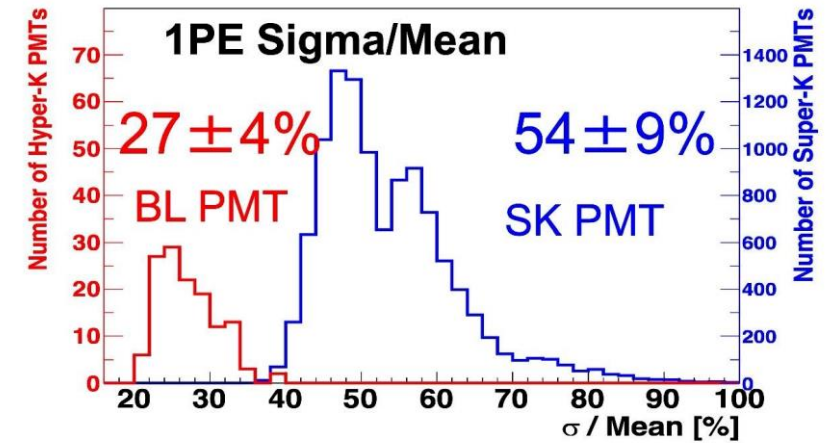
Hyper-K 50cm PMT performance

×2 better photodetection efficiency (QE×CE)

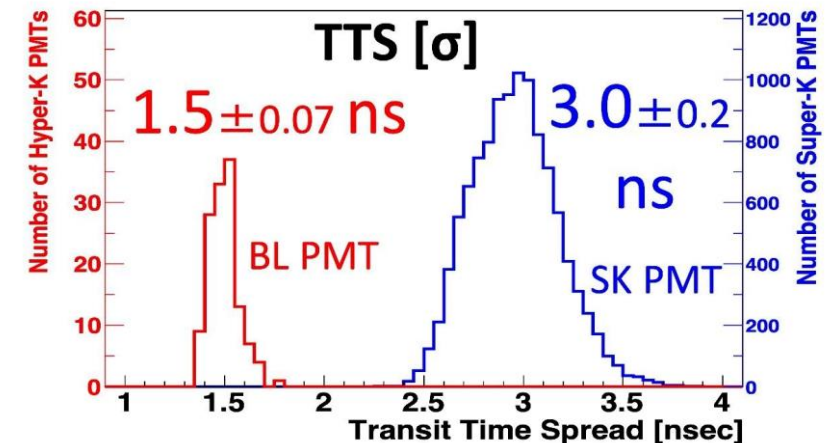


(Performance in SK tank, $1.7e7$ gain)

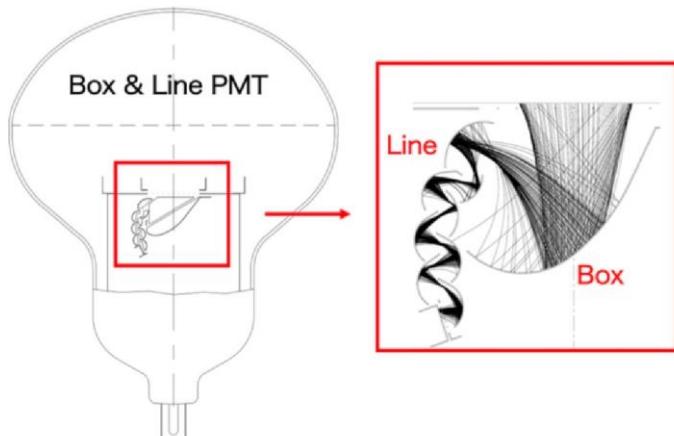
×2 better charge resolution



×2 better timing resolution



Box&Line dynode



×2 better pressure tolerance
→ enable deeper tank design,
project cost reduction

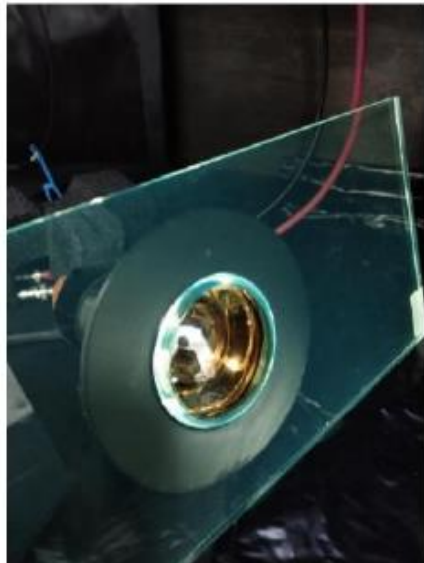
Low dark rate (4kHz) and RI



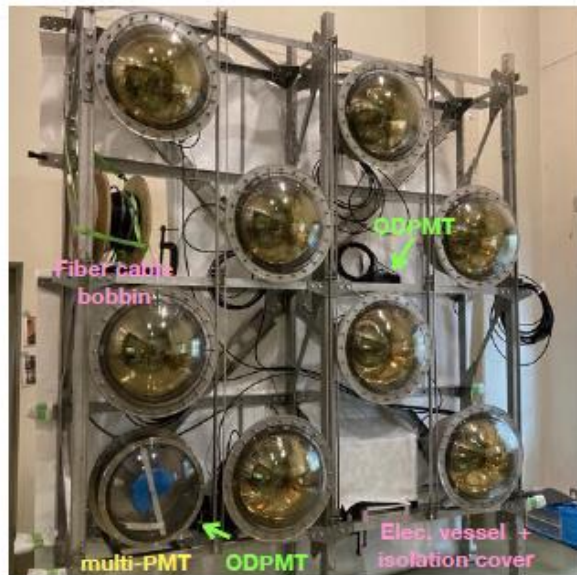
PMT production ongoing, > 10,000 delivered.
Screening both at Hamamatsu and Kamioka

Photosensors and underwater electronics

Outer detector: PMT+WLS plate

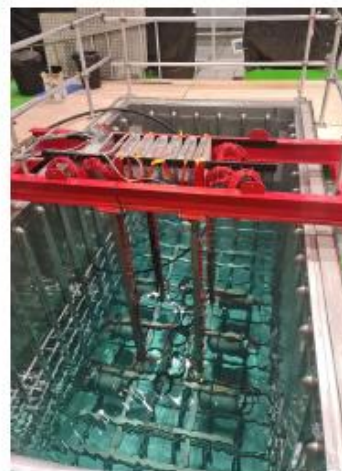


Photosensors/elec. mockup

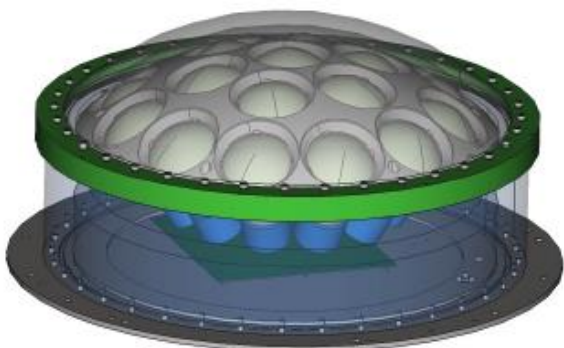


Underwater electronics:

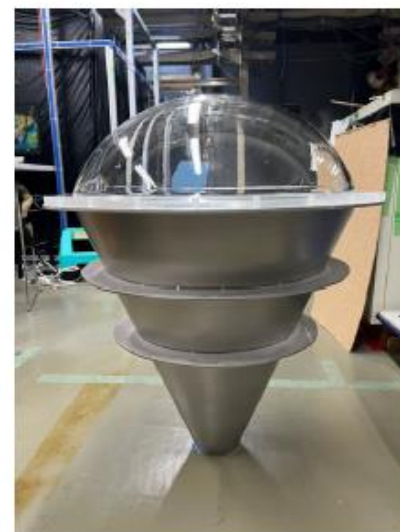
Case design and feedthrough



Multi-PMT module:



PMT cover



Design finalization ongoing

Hyper-K Calibration

- Various programs to determine detector parameters and measure systematics
- Pre-calibration of photosensors
- Photogrammetry
- Light Injection
 - Diffusers and collimators
 - mPMT system
 - OD injectors
- Electron LINAC
 - 3-24 MeV electrons
- Radioactive Sources
 - DT Source - 16N
 - AmBe + BGO – tagged neutrons
 - Ni/Cf - 9 MeV gamma cascade

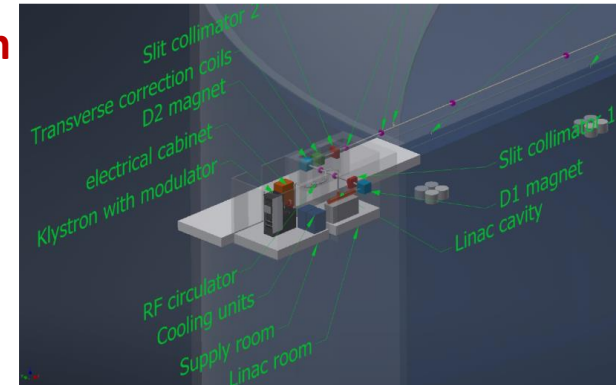
Photosensor Test Facility



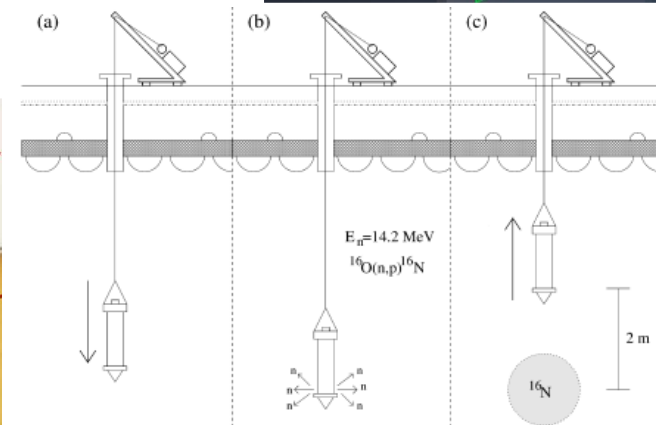
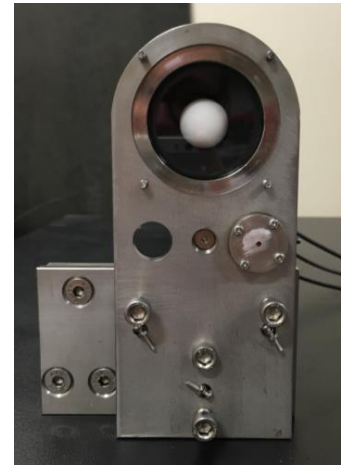
Photogrammetry testing



LINAC beam simulation



Light injectors

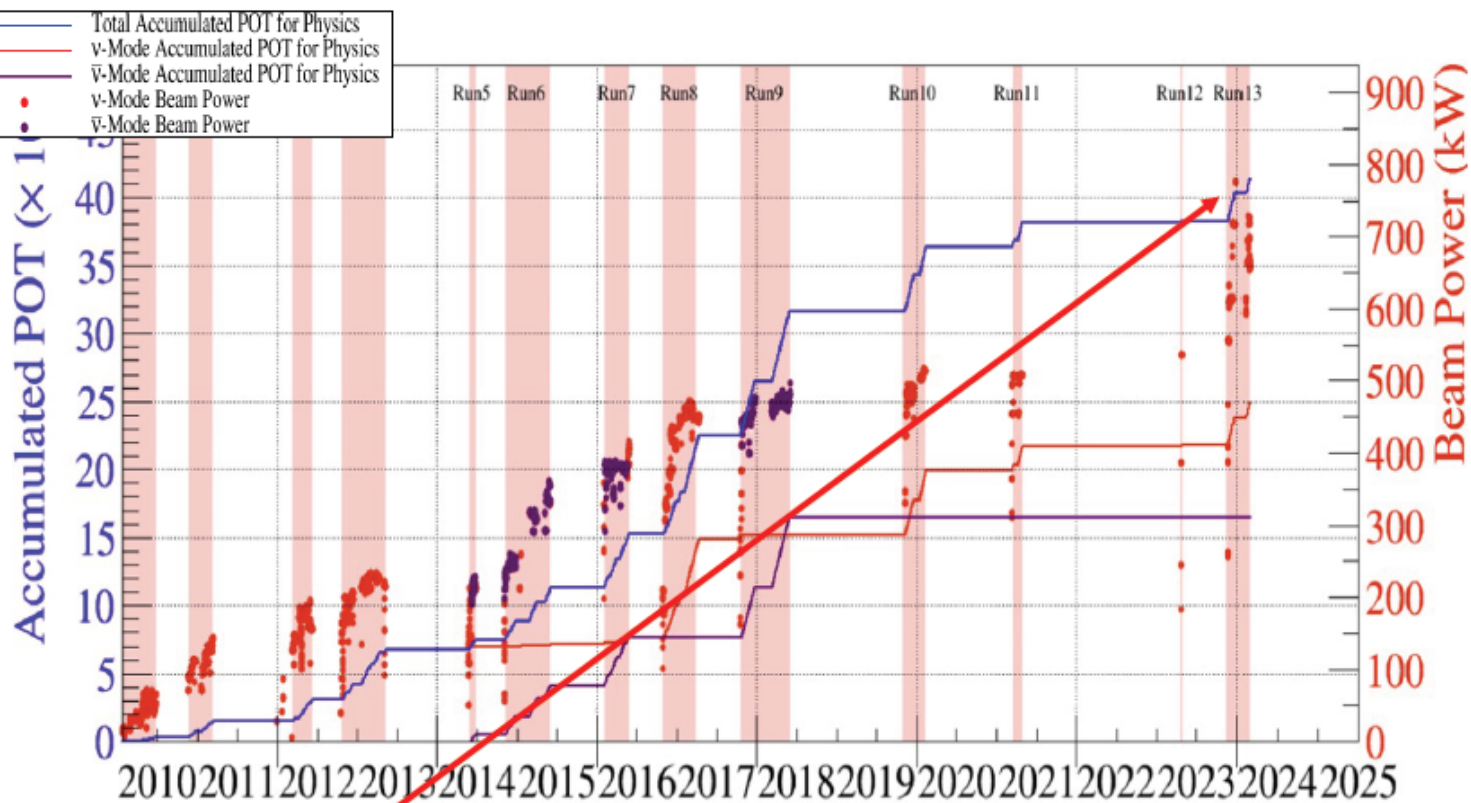


DT operation

Ni/Cf source



Beam: status and plan of power increase

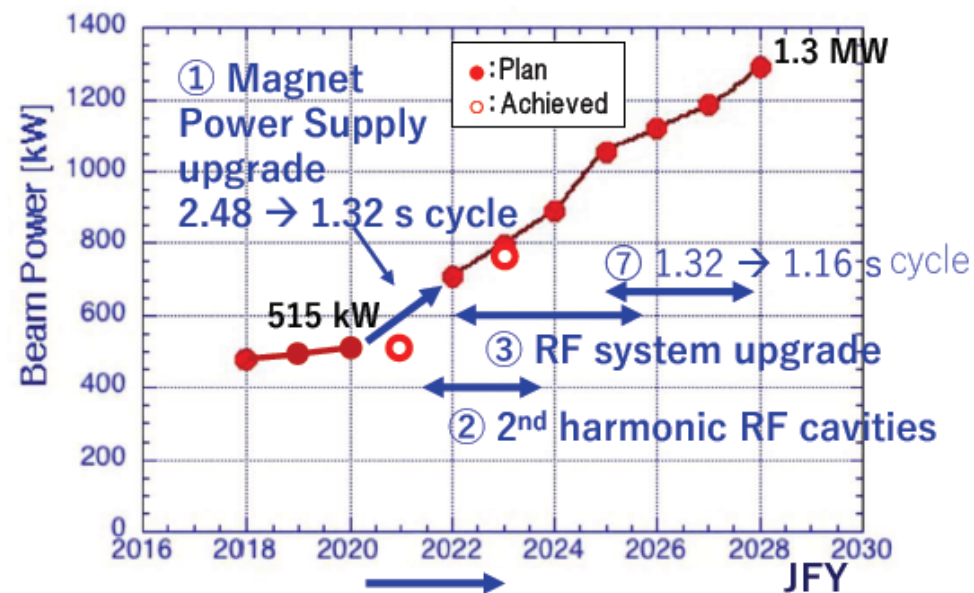


760 kW achieved already and **800 kW last week!**

Further beam power increase requires:

- Seeking beam loss with optics improvements
- More protons/pulse by upgrading RF system
- Further beam intensity increase will be done by **1.36 → 1.16 sec cycle**

Original power projection in MR Upgrade Plan

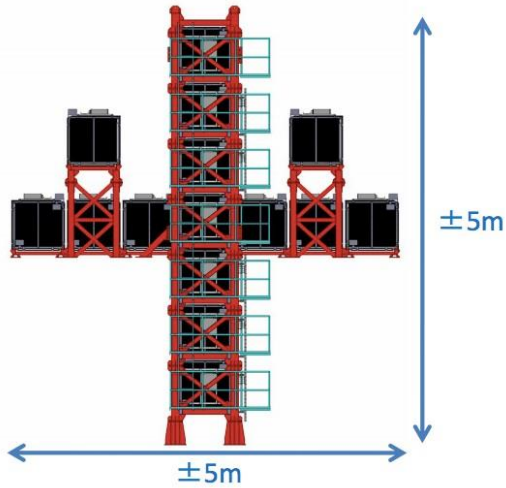


S. Igarashi, *et. al.*,
PTEP vol 2021,
Issue.3,p33

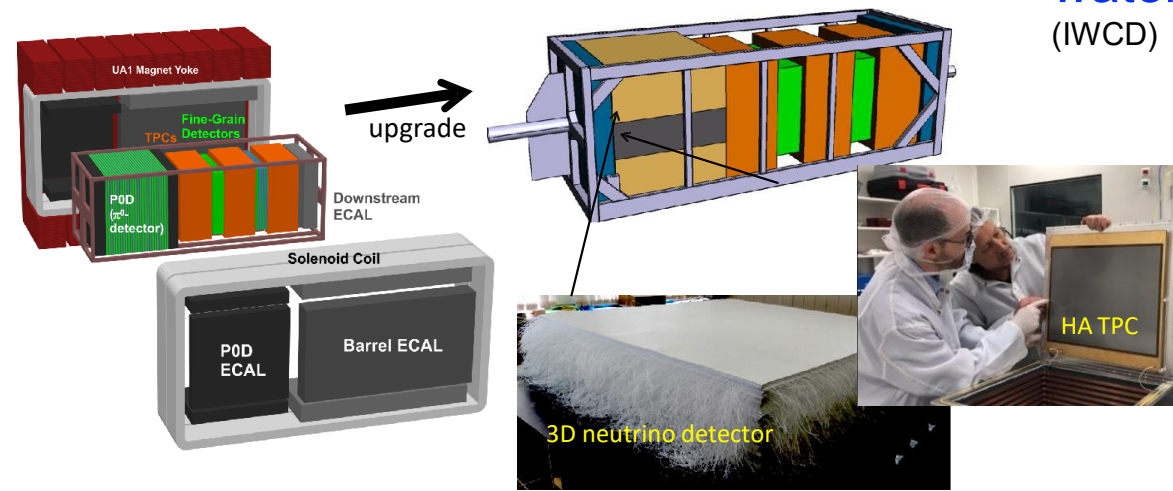
- ④ Collimator system
- ⑤ Injection/FX system
- ⑥ Beam Monitors (BPM circuits)

Neutrino detectors at J-PARC

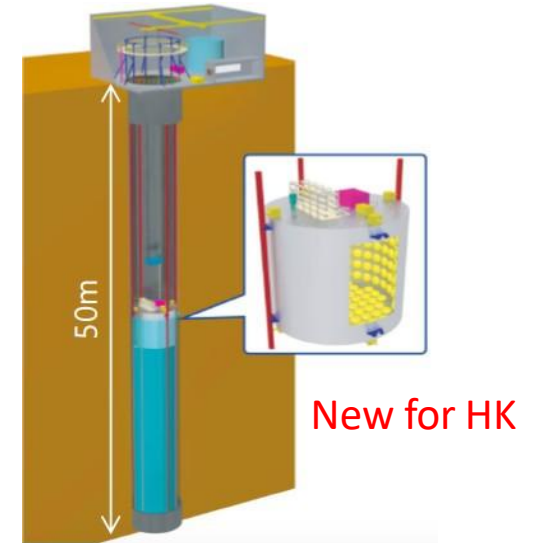
On-axis Detector
(INGRID)



Off-axis Magnetized Tracker
(ND280 → Upgrade for T2K → Upgrade for HK)



Off-axis spanning Intermediate water Cherenkov detector
(IWCD)



Critical components to precisely understand J-PARC beam and neutrino interactions:

- **On-axis detector:** Measure beam direction and event rate
- **Off-axis magnetized tracker:** Measure primary (anti)neutrino interaction rates, spectrum, and properties. Charge separation to measure wrong-sign background
→ Upgrade by T2K experiment and intensive discussion for further upgrade in HK-era is ongoing.
- **Intermediate WC detector:** H₂O target with off-axis angle spanning orientation.
→ Detector site investigation and conceptual facility design are ongoing.

Mexican funds awarded for Hyper-K

- CF-2023-G-643 "Construcción y comisionado de sensores de ciencia frontera para la detección de supernovas, materia oscura, y medición de la asimetría bariónica en el Universo, en experimentos de Neutrinos de nueva generación" (2023)
 - Grant holder: Eduardo de la Fuente Acosta (UdeG)
 - Institutions involved:
 - KAREN SALOME CABALLERO MORA (UNACH)
 - GIANNINA DALLE MESE ZAVALA (UAS)
 - ALEJANDRO KADSUMI TOMATANI SANCHEZ (TEC-GDL)
 - Saul Cuen Rochin (TEC-SIN)
- CBF2023-2024-427 "Deep Learning y Fabricación de Sensores de Ciencia de Frontera para Experimentos de Neutrinos de Próxima Generación" (2024)
 - Grant holder : Saul Cuen Rochin (TEC-SIN)
 - Institutions involved :
 - GIANNINA DALLE MESE ZAVALA (UAS)

MOU between Mexican Institutions and U.Tokyo/KEK should be ready by the end of 2024.



Mexican involvement in Hyper-K

Master thesis in progress (**TEC**):

- Neutrino Classification Through Deep Learning amid the Hyper-Kamiokande Project Development

Student: Maria Fernanda Romo Fuentes

Advisor: Luis Eduardo Falcon Morales

Doctoral thesis in progress (**UdeG**):

- Use of Machine Learning and Deep Learning in the reconstruction of high energy events for the Hyper Kamiokande

Student: Felipe Orozco Luna

Advisors: Eduardo de la Fuente, Luis Eduardo Falcon, Saul Cuen

Thesis open position (**UAS**):

- Analysis for supernova detection

Advisor: GIANNINA DALLE MESE ZAVALA

Thesis open position (**UNACH**):

- Analysis for supernova detection

Advisors: KAREN SALOME CABALLERO MORA

Thesis open position:

- Neutrino Classification with AI

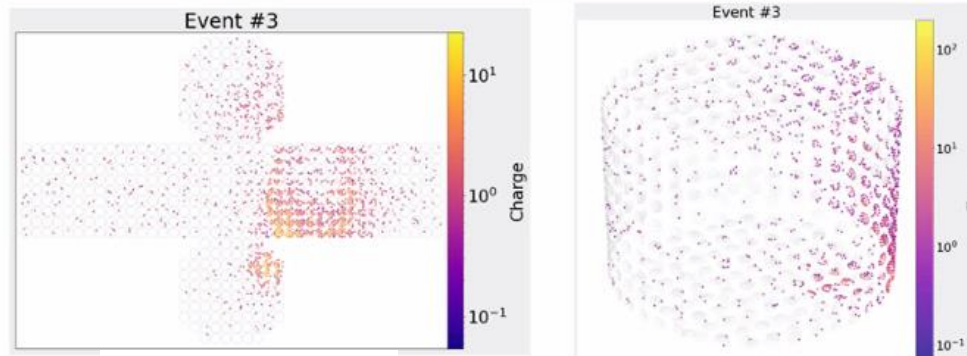
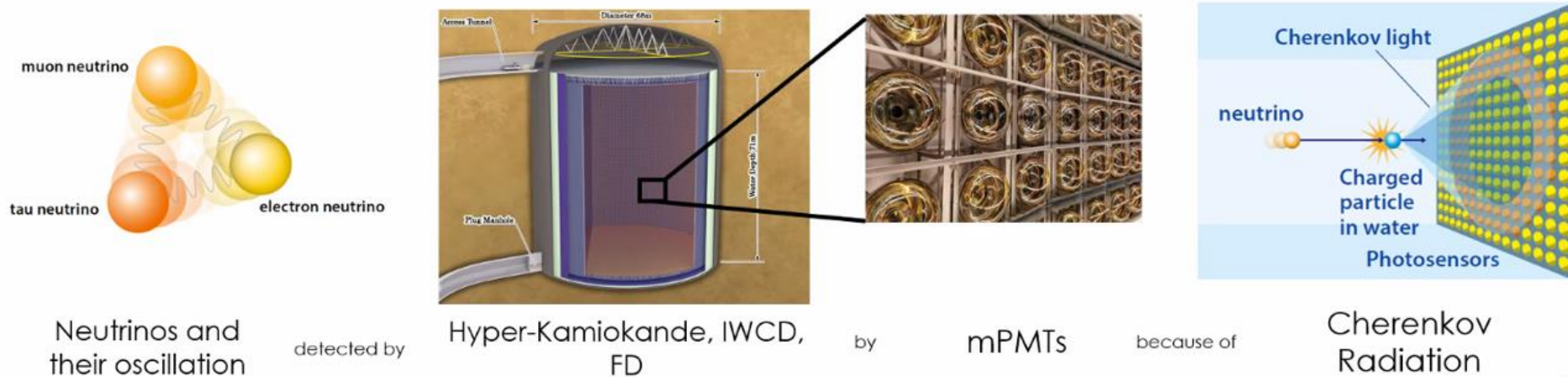
Advisors: Saul Cuen, Rajesh Biswal, Rodrigo Gamboa (volunteers?)

Thesis open position:

- mPMT design and manufacturing for Hyper-K

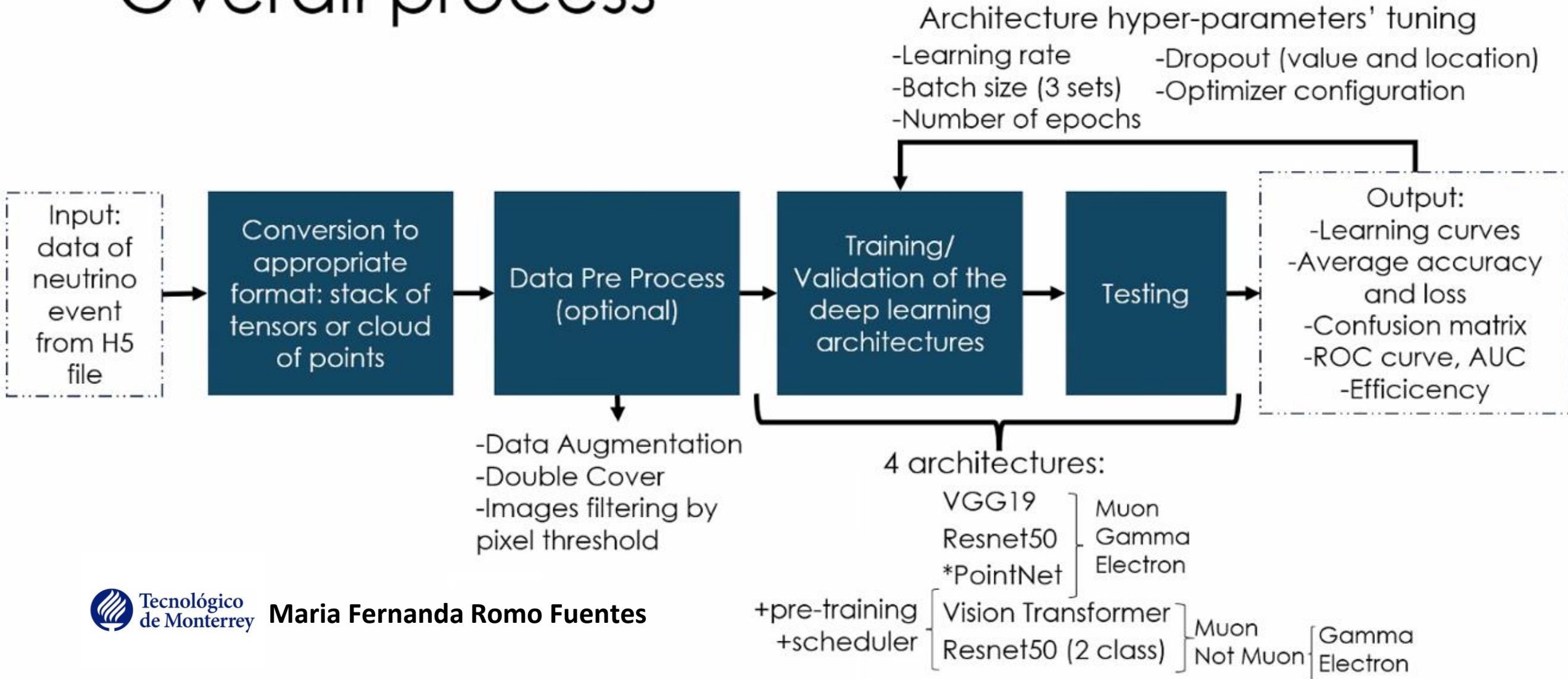
Advisors: Saul Cuen, Kadsumi Tomatani, (volunteers?)

Neutrino classification



Detected event
has to be ↓
Reconstructed
starts by ↓
Identifying the
particles
involved in an
event

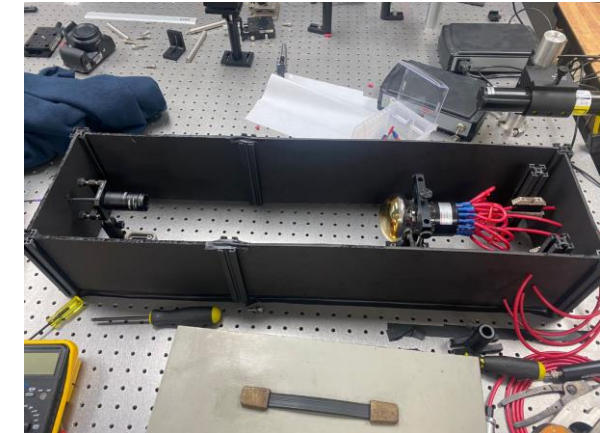
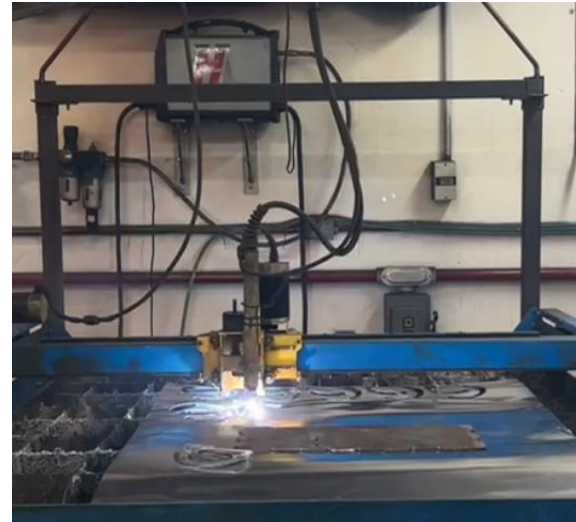
Overall process



mPMT prototype in Mexico

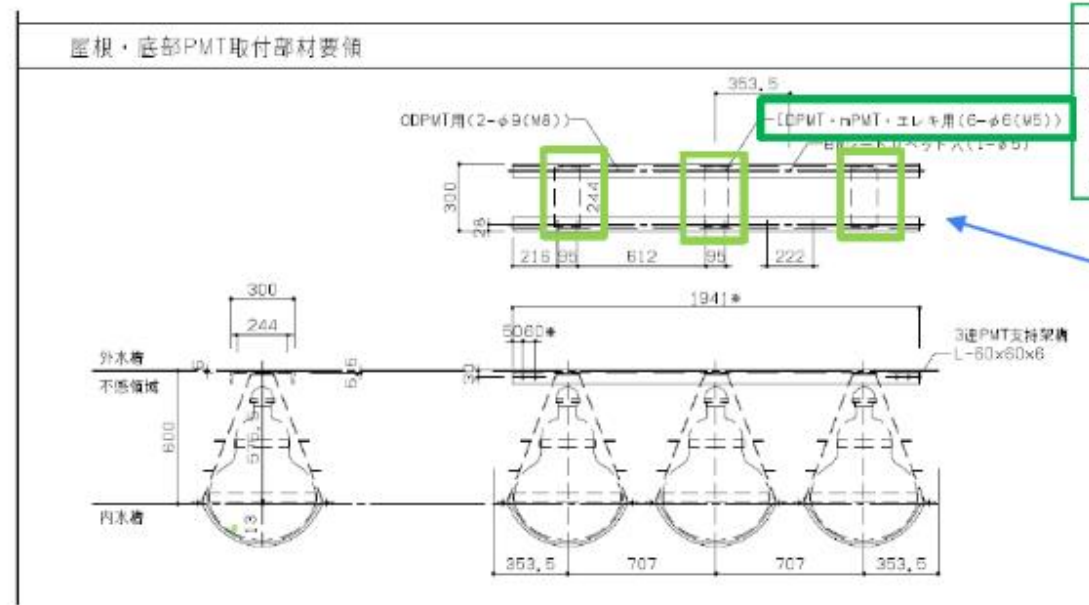
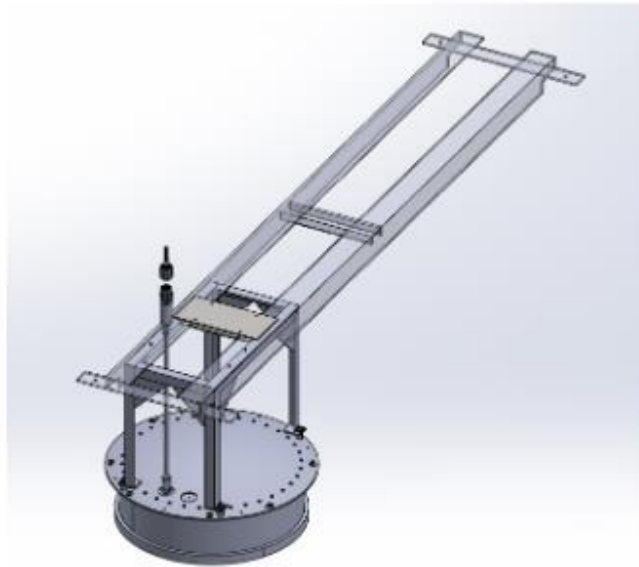
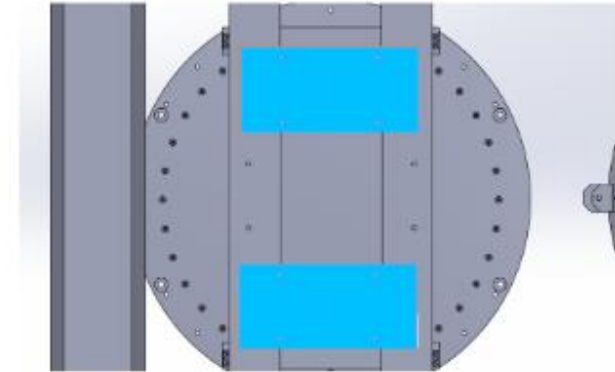
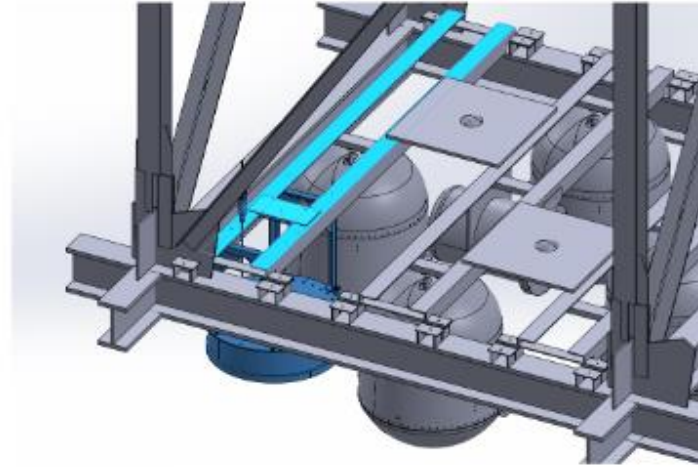
mPMT assembly and testing at TEC in collaboration with Professors Rodrigo Salmon, Kadsumi Tomatani, Raul Aranda, Christoper Falcon, Eduardo de la Fuente and Saul Cuen

- Mechanical metrology and assembly
- Setting un blackbox and optical testing for PMT check (student Roy Medina)



top & bottom mPMT support

Currently working on requirements from the integration group.



mPMT mechanical stress test

Top/bottom configuration

Barrel configuration

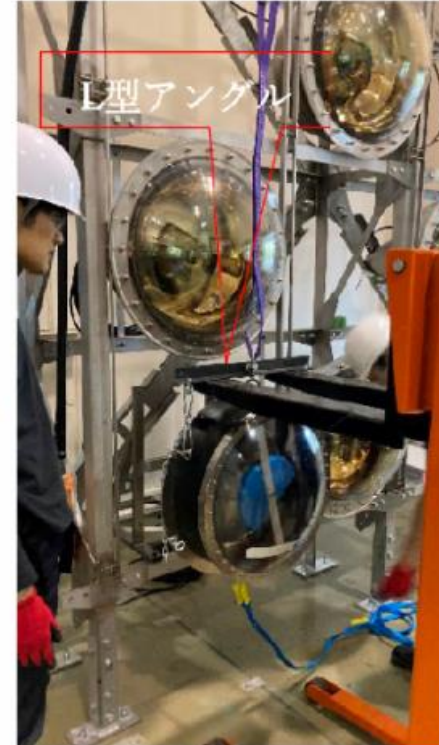
Transportation studies, and box design

- Compression
- Temperature
- Vibrations

(Kadsumi Tomatani, Christopher Falcon, Saul Cuen)



- mPMT installed successfully - procedure itself ok (possible change after talking to the inst. company)
- The main issue was the interference of the with the main frame due to
 - Enlarged gusset plate (cannot be modified)
 - Shifted front mounting holes (can/should be modified)



Top Installation Overview

- mPMT lifted by ceiling crane (not the original and final installation procedure) - successful
- Cause by the issues of lifting the 3-PMT module with middle space occupied



Conclusions

- **Hyper-Kamiokande is 3rd generation water Cherenkov detector in Kamioka**
- **Important physics targets**
 - Neutrino CP violation: Discovery with 5σ for $\sim 60\%$ parameter regions
 - Nucleon Decay Search for testing GUT: $\tau > 10^{35}$ years for $p \rightarrow e^+\pi^0$
 - Neutrino Astrophysics: Supernova neutrinos
- **Hyper-Kamiokande construction on schedule**
 - World's largest underground facility: 260 kton water Cherenkov detector
 - Access tunnel and cavern construction on track
 - 50cm PMT production underway
 - Other detector component designs being finalized
 - Neutrino beam upgrade to 1.3 MW
 - Near detector upgrade and design of intermediate detector being finalized
- **Hyper-Kamiokande will start operation in 2027.**