

# Charged Lepton Flavour Violation in the Inverse Seesaw

Based on the work “Charged lepton flavour violation from inverse seesaw with flavour and CP symmetries”, in collaboration with C. Hagedorn, Soon to be Published

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# Flash Introduction to Neutrino Physics

- 1930 : Pauli's hypothesis to explain  $\beta$ -decay
  - 1956: Discovered at Los Alamos (Reines and Cowan)
- Supposed massless and only LH



**Formulation of GSW theory  
of EW interactions (approx  
1960-1970)**

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- 1930 : Pauli's hypothesis to explain  $\beta$ -decay
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- Supposed massless and only LH
- 1968 : Detection of Solar Neutrinos (Homestake exp.)
  - Only 1/3 of electron neutrinos observed...
- 1985 : Detection of Atmospheric Neutrinos (Kamiokande and IMB)
  - Smaller ratio of muon neutrinos to electron neutrinos observed
- 1998 : First evidence of Atmospheric Neutrino Oscillations (Super-Kamiokande)
- 2000 : Discovery of Tau neutrino (DONUT collaboration)

**Formulation of GSW theory of EW interactions (approx 1960-1970)**

**Evidences of oscillation in the neutrino sector is incompatible with massless neutrinos!**

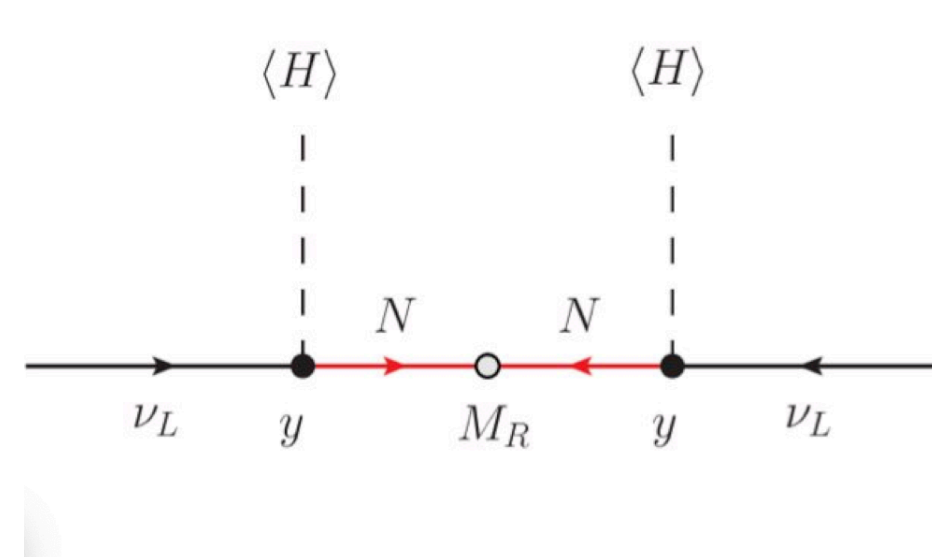
**Flavour dynamics needs explanations!!**

[<https://neutrinos.fnal.gov/history/>]

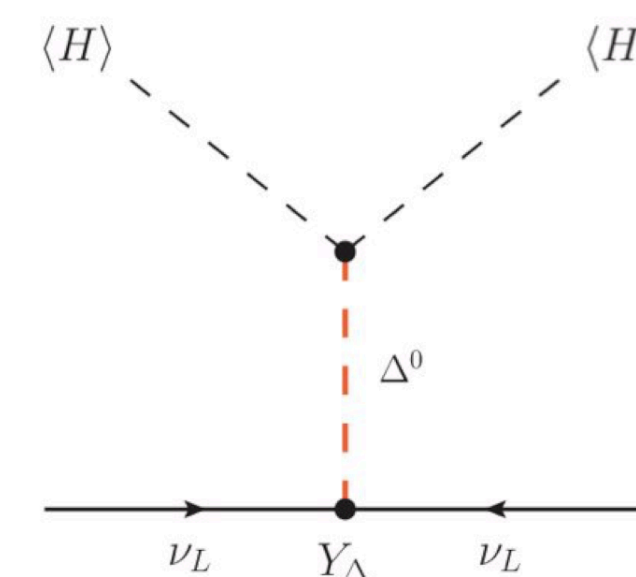
# Neutrino Mass Generation

$$\mathcal{O}_W^{(5)} = \frac{1}{\Lambda} \langle LLHH \rangle$$

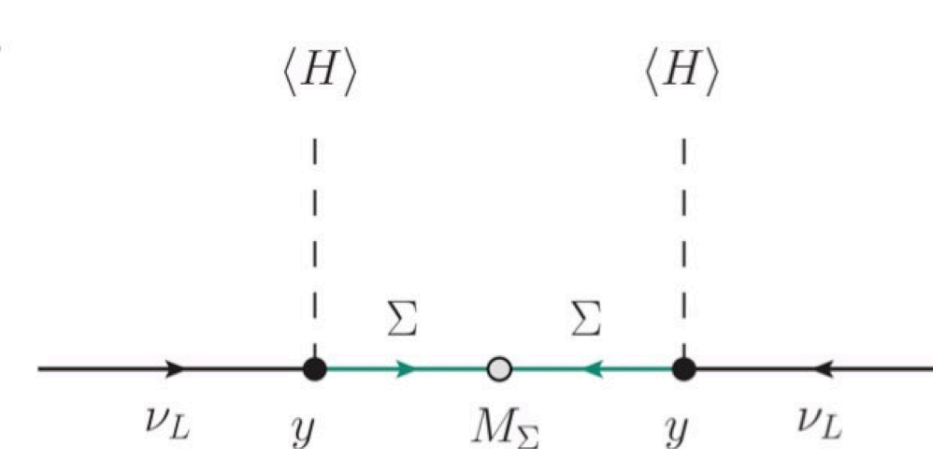
- Majorana Masses generated via Weinberg operators
- Common implementation: **Seesaw Mechanism**



$$m_\nu \sim y^2 \frac{v^2}{M_R}$$



$$m_\nu \sim \frac{v^2}{M_\Delta} Y_\Delta$$



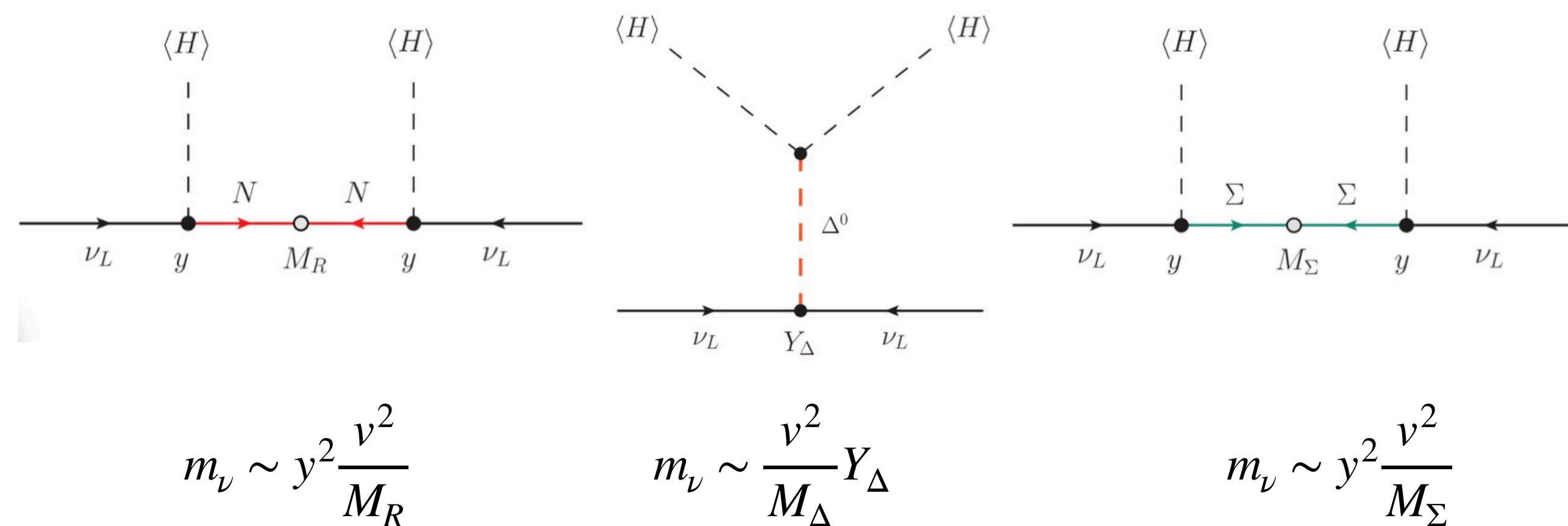
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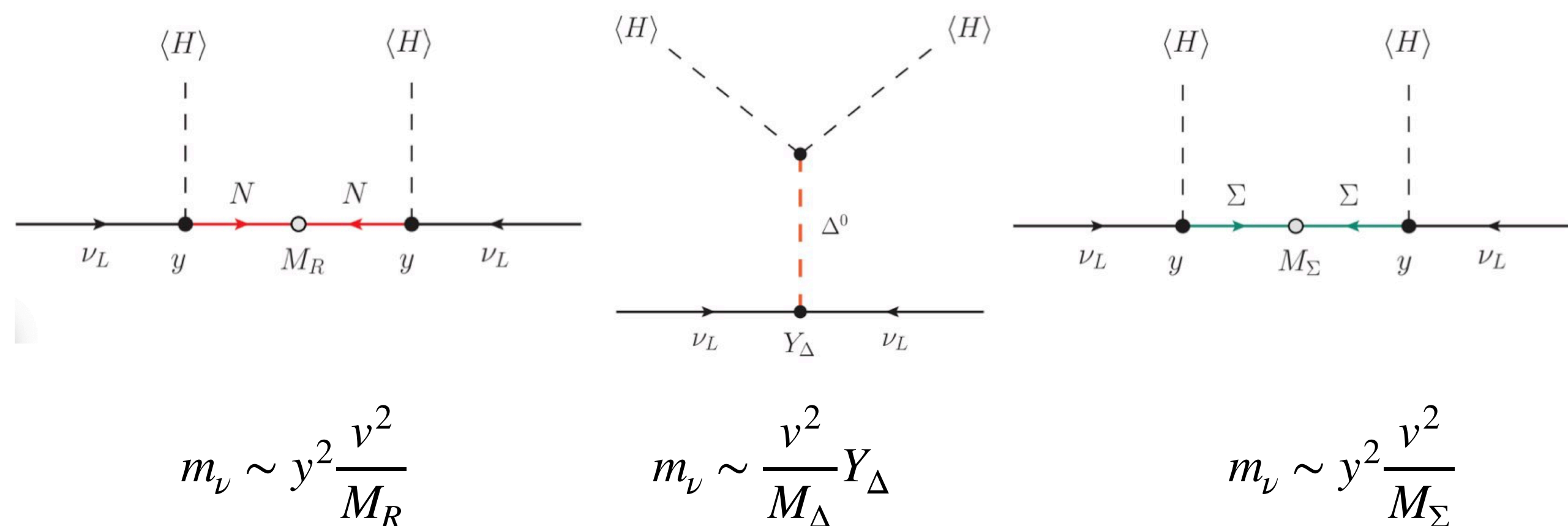
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**Seesaw Mechanism**

- Can we lower the mass scale?



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# Inverse SeeSaw (ISS) Framework

$$\mathcal{L}_m = \bar{L}^c Y_D H N^c + \bar{N} M_{NS} S + \frac{1}{2} \bar{S}^c \mu_S S + h.c. = (\bar{\nu}_L \quad \bar{N}^c \quad \bar{S})^c \mathcal{M} \begin{pmatrix} \nu_L \\ N^c \\ S \end{pmatrix}$$

$$\mathcal{M} = \begin{pmatrix} \emptyset & m_D & \emptyset \\ m_D^T & \emptyset & M_{NS} \\ \emptyset & M_{NS}^T & \mu_S \end{pmatrix}$$

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$$\mathcal{U}^T \mathcal{M} \mathcal{U} = \mathcal{M}^{(diag)}$$

$$\mathcal{U} = \begin{pmatrix} \tilde{U}_\nu & S \\ T & V \end{pmatrix}$$

$$m_\nu \approx m_D (M_{NS}^{-1})^T \mu_S M_{NS}^{-1} m_D^{-1} \sim y^2 \frac{\mu_0}{M_0^2}$$



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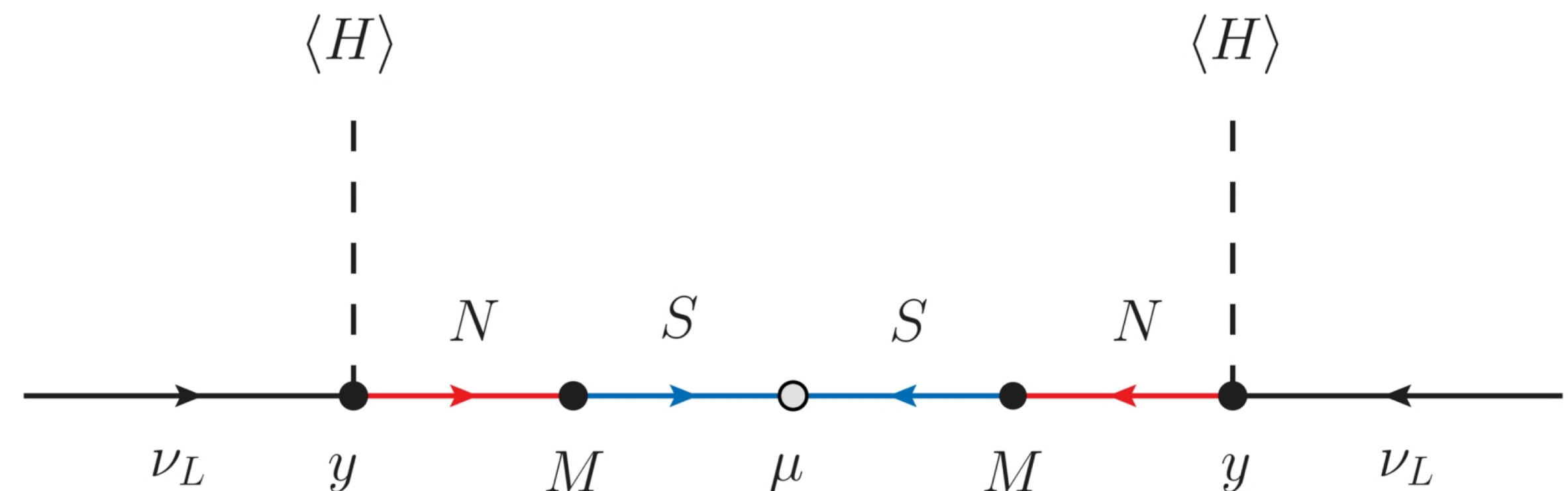
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- **Inverse Seesaw (ISS):**

- Light neutrino masses generated by the joint action of  $M_{NS}$  and small  $\mu_S$ .

- $\mu_S$  is a small source of Lepton Number violation



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In the basis in which charged lepton mass matrix is diagonal,  
 $\tilde{U}_\nu$  is the **(non-unitary)** leptonic mixing matrix

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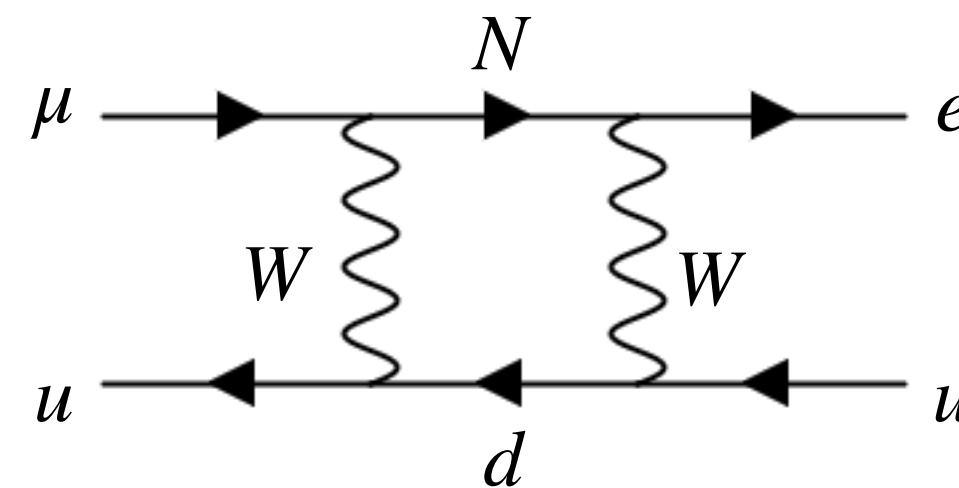
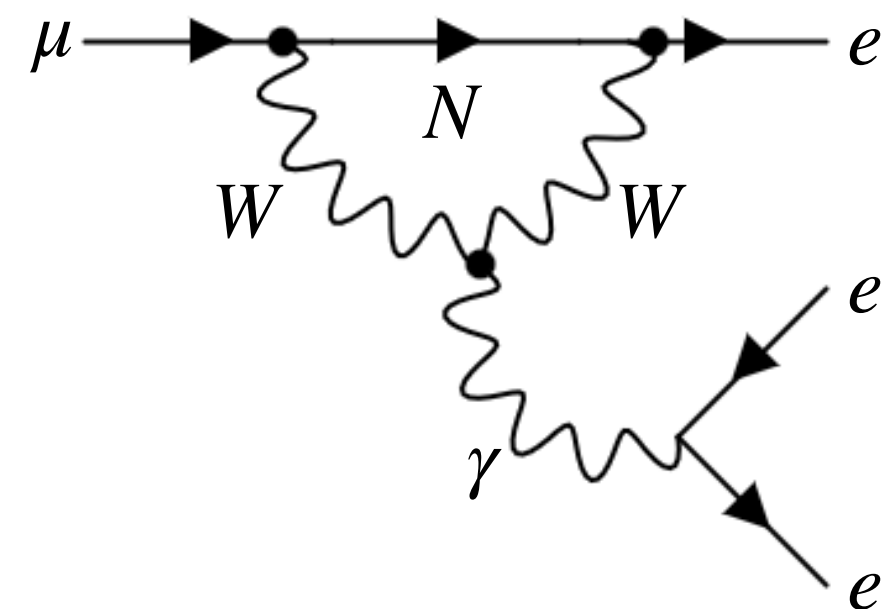
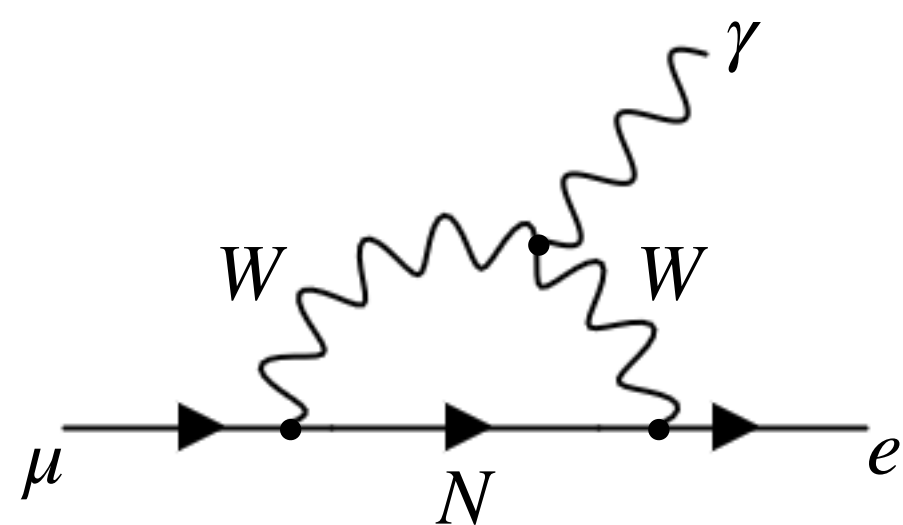
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$S, T$  describe the mixing of light and sterile neutrinos

$(S, T \ll \tilde{U}_\nu) \Rightarrow$  **Can induce cLFV processes**





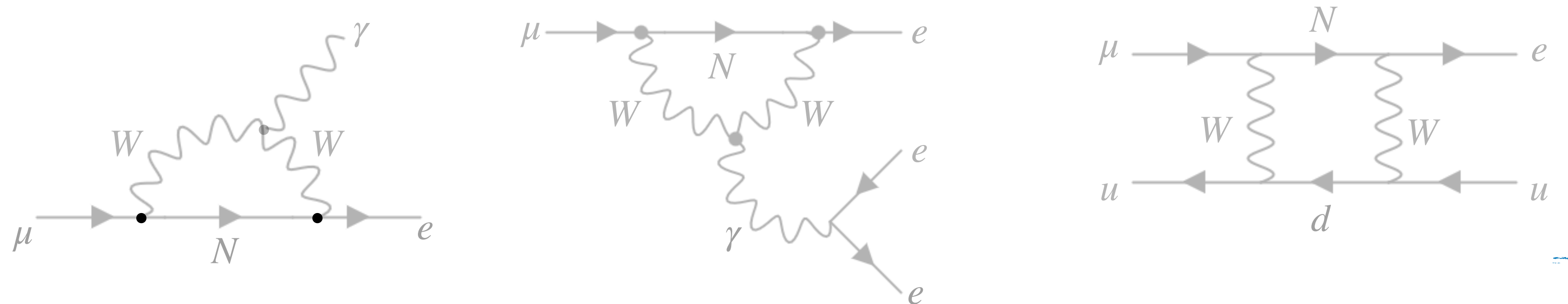
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**In this talk, we:**

- **Reproduce Neutrino Masses and mixing within an ISS framework equipped with some Flavour symmetry group**
- **Show that in such context cLFV processes are predicted, and how future experiment could probe our parameter space**

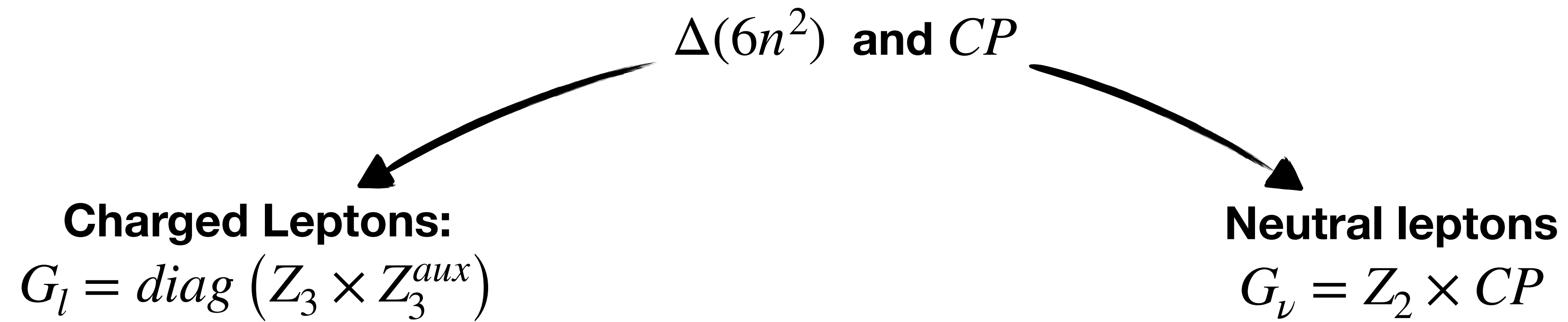
$(B, L \leq U_\nu) \rightarrow$  Can induce cLFV processes



# Flavour Symmetry Framework:

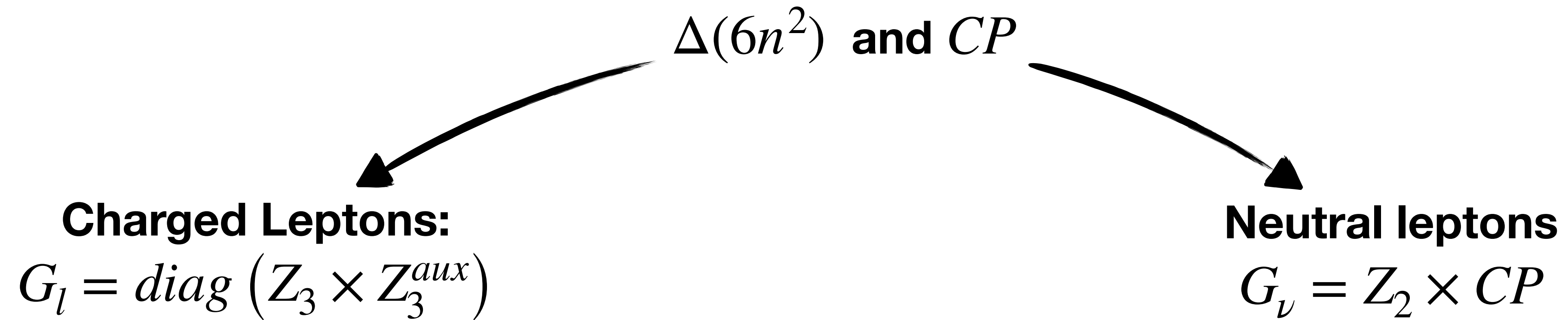
$\Delta(6n^2)$  and  $CP$

# Flavour Symmetry Framework:





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$L, N, S$  assigned to triplet representations of the flavour group;

Various ways of choosing the residual symmetry are possible

# Flavour Symmetry Framework:

$\Delta(6n^2)$  and  $CP$

**Charged Leptons:**

$$G_l = \text{diag} (Z_3 \times Z_3^{\text{aux}})$$

**Neutral leptons**

$$G_\nu = Z_2 \times CP$$

$$\mathcal{L}_m = \bar{L}^c Y_D H N^c + \bar{N} M_{NS} S + \frac{1}{2} \bar{S}^c \mu_S S + h.c.$$

C. Hagedorn, J. Kriewald, J. Orloff, A. M. Teixeira, 2107.07537

Different choices of  $\Delta(6n^2)$  representation result in different phenomenology:

- Option 1:  $L \sim 3$  ;  $N \sim 3$  ;  $S \sim 3$
- Option 2:  $L \sim 3$  ;  $N \sim 3'$  ;  $S \sim 3'$
- Option 3:  $L \sim 3$  ;  $N \sim 3$  ;  $S \sim 3'$

$$Y_D = y_0 \frac{\langle H \rangle}{M_0} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$M_{NS} = M_0 \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$\mu_S = U_S^*(\theta_S) \begin{pmatrix} \mu_1 & 0 & 0 \\ 0 & \mu_2 & 0 \\ 0 & 0 & \mu_3 \end{pmatrix} U_S^T(\theta_S)$$

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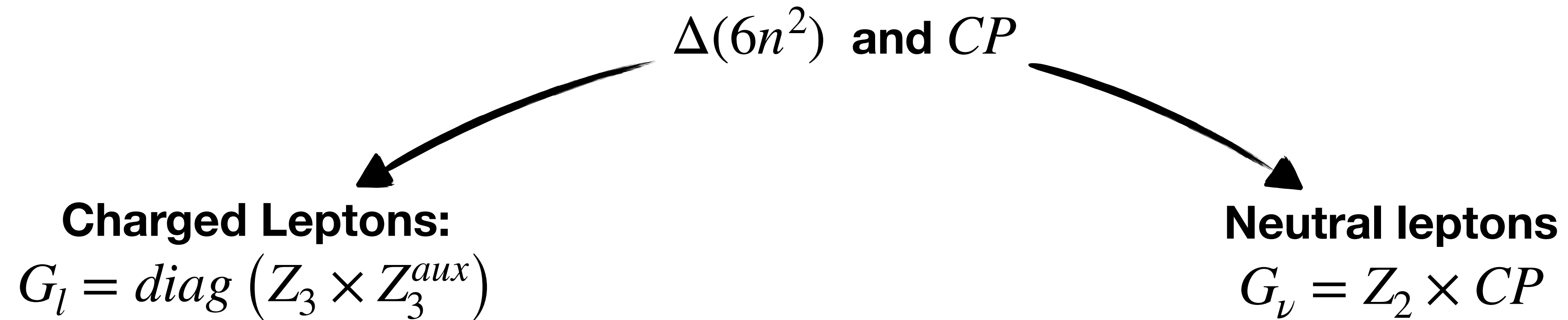
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$$M_{NS} = U_{NS}^N(\theta_N) \begin{pmatrix} M_1 & 0 & 0 \\ 0 & M_2 & 0 \\ 0 & 0 & M_3 \end{pmatrix} U_{NS}^{S\dagger}(\theta_S)$$

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In this discussion, we focus on **Option 2**

# Option 2

$$m_\nu \approx \frac{1}{M_0^2} m_D \mu_S m_D^T = \langle H \rangle \frac{\mu_0}{M_0^2} U_D^{(L)*}(\theta_L) \left[ \begin{pmatrix} y_1 & 0 & 0 \\ 0 & y_2 & 0 \\ 0 & 0 & y_3 \end{pmatrix} U_D^{(R)T}(\theta_R) \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} U_D^{(R)}(\theta_R) \begin{pmatrix} y_1 & 0 & 0 \\ 0 & y_2 & 0 \\ 0 & 0 & y_3 \end{pmatrix} \right] U_D^{(L)\dagger}(\theta_L)$$

Heavy neutral states mass matrix:

$$M_h \approx \begin{pmatrix} \emptyset & M_{NS} \\ M_{NS}^T & \mu_S \end{pmatrix}$$

Heavy neutral states spectrum: **Three approximately degenerate Pseudo-Dirac Pairs**

$$m_{(i=4,5,6)} \approx M_0 - \frac{1}{2} \mu_0 \quad m_{(i=7,8,9)} \approx M_0 + \frac{1}{2} \mu_0$$



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$$U_D^{(L)}(\theta_L) = \Omega(3) R_{ij}(\theta_L)$$

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Completely fixed by choice of residual symmetry

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Group theoretical parameters specify the residual symmetry :  $n, s, t$

**E.g.:**

$$\Omega(3)(u, v) = e^{i\frac{v\pi}{n}} U_{TB} R_{13} \begin{pmatrix} -\frac{u\pi}{2n} \\ \\ \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{-i\frac{v\pi}{n}} & 0 \\ 0 & 0 & -i \end{pmatrix}$$



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$R_{ij}(\theta_{L,R})$  is a rotation on the  $ij$  plane

Codifies residual freedom in the choice of  $\Omega$  matrices

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**Parameters of the theory:**

$$M_0, \mu_0, y_i, \theta_L, \theta_R$$

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$$M_0 \in [150 \text{ GeV}; 10 \text{ TeV}] \quad \mu_0 \in [0.1 \text{ keV}; 10^2 \text{ keV}] \quad \theta_R \in [0; 2\pi]$$

# Option 2

$$m_\nu \approx \frac{1}{M_0^2} m_D \mu_S m_D^T = \langle H \rangle \frac{\mu_0}{M_0^2} U_D^{(L)*}(\theta_L) \left[ \begin{pmatrix} y_1 & 0 & 0 \\ 0 & y_2 & 0 \\ 0 & 0 & y_3 \end{pmatrix} U_D^{(R)T}(\theta_R) \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} U_D^{(R)}(\theta_R) \begin{pmatrix} y_1 & 0 & 0 \\ 0 & y_2 & 0 \\ 0 & 0 & y_3 \end{pmatrix} \right] U_D^{(L)\dagger}(\theta_L)$$

- In our numerical analysis:

$$M_0 \in [150 \text{ GeV}; 10 \text{ TeV}] \quad \mu_0 \in [0.1 \text{ keV}; 10^2 \text{ keV}] \quad \theta_R \in [0; 2\pi]$$

- Fixed by fitting LO predictions of the masses to experimental values

# Option 2

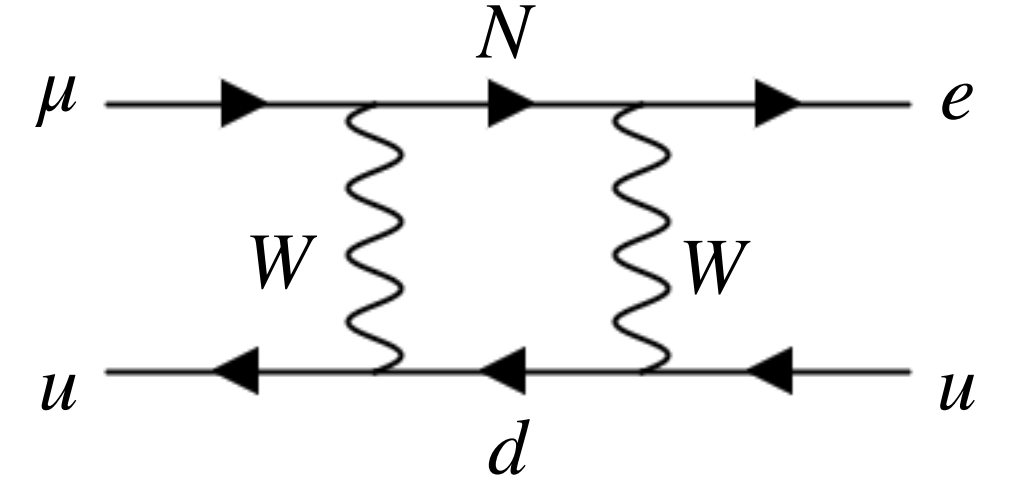
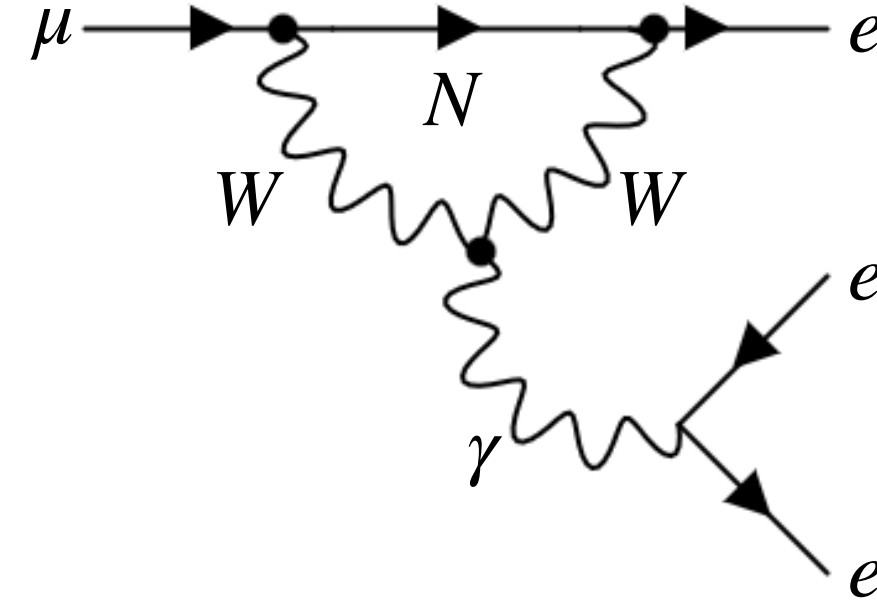
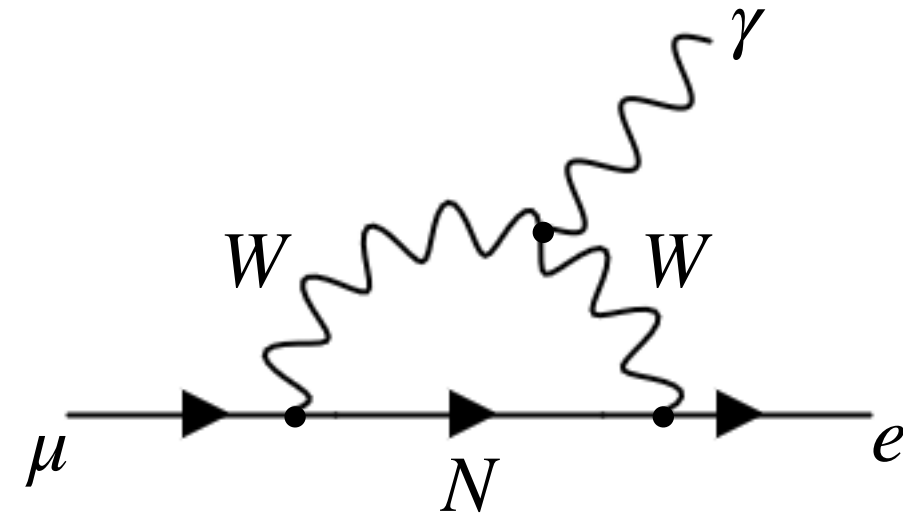
$$m_\nu \approx \frac{1}{M_0^2} m_D \mu_S m_D^T = \langle H \rangle \frac{\mu_0}{M_0^2} U_D^{(L)*}(\theta_L) \left[ \begin{pmatrix} y_1 & 0 & 0 \\ 0 & y_2 & 0 \\ 0 & 0 & y_3 \end{pmatrix} U_D^{(R)T}(\theta_R) \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} U_D^{(R)}(\theta_R) \begin{pmatrix} y_1 & 0 & 0 \\ 0 & y_2 & 0 \\ 0 & 0 & y_3 \end{pmatrix} \right] U_D^{(L)\dagger}(\theta_L)$$

- In our numerical analysis:

$$M_0 \in [150 \text{ GeV}; 10 \text{ TeV}] \quad \mu_0 \in [0.1 \text{ keV}; 10^2 \text{ keV}] \quad \theta_R \in [0; 2\pi]$$

- Fixed by fitting LO predictions of the masses to experimental values
- Fixed by fitting prediction of mixing to lepton mixing data

# cLFV in the ISS



Relevant (approximated) loop functions:

$$G_{\gamma}^{\beta\alpha} \approx \eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{\mu_0}{M_0}\right)$$

$$F_{\gamma}^{\beta\alpha} \approx -2 \left( \frac{7}{12} + \frac{1}{6} \log x_0 \right) \eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{m\mu_0}{M_0}\right)$$

$$F_Z^{\beta\alpha} \approx \eta_{\alpha\beta} (5 - 3 \log x_0) + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{\mu_0}{M_0}\right)$$

$$F_{box}^{\beta 3\alpha} \approx -2\eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y^2\mu_0}{M_0}, \frac{\mu_0^2}{M_0^2}\right)$$

$$F_{box}^{\mu e u u} = 2\eta_{e\mu} \left[ -4 - |V_{ub}|^2 (F_{box}(0, x_b) - F_{box}(x_0, x_b) - 4) \right]$$

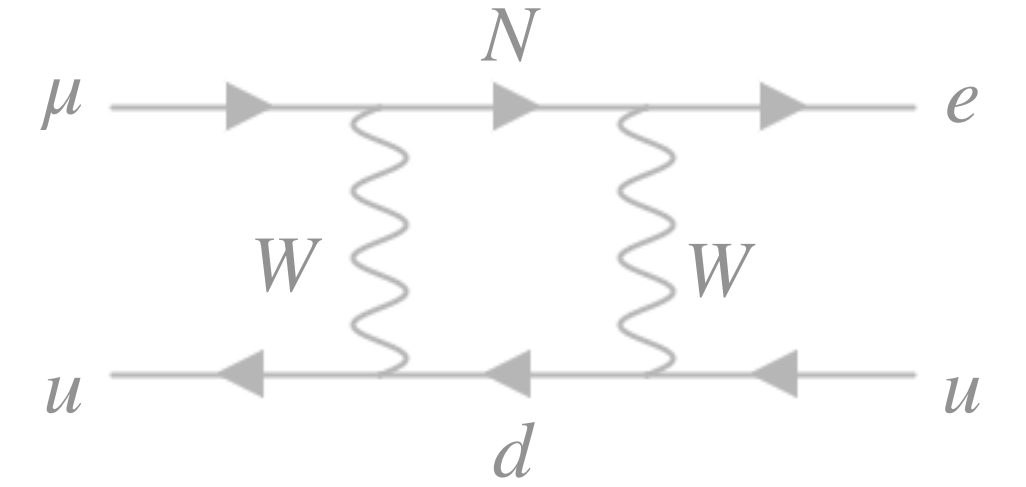
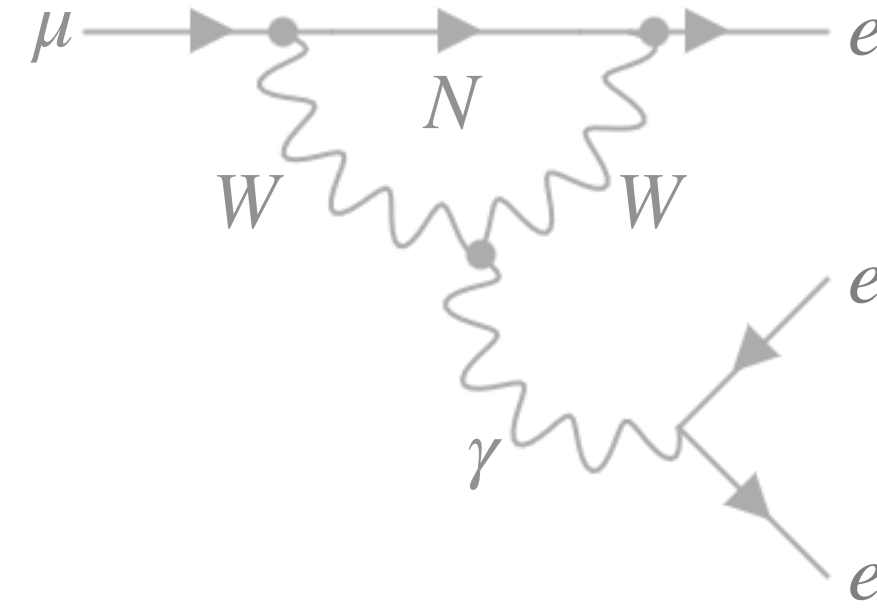
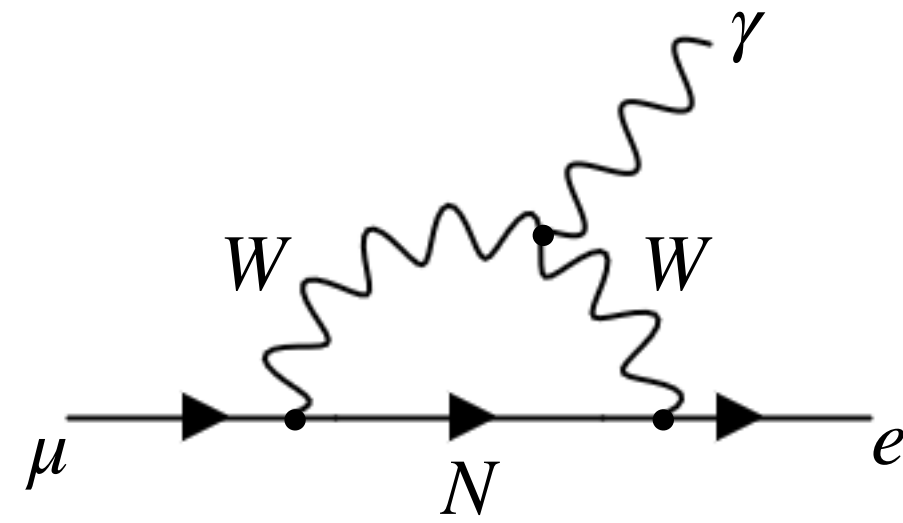
$$F_{box}^{\mu e d d} = 2\eta_{e\mu} \left[ 1 - |V_{td}|^2 (F_{Xbox}(0, x_t) - F_{Xbox}(x_0, x_t) + 1) \right]$$

Only LO in ISS framework

Considered degenerate heavy neutral states, **with mass  $M_0$**



# cLFV in the ISS



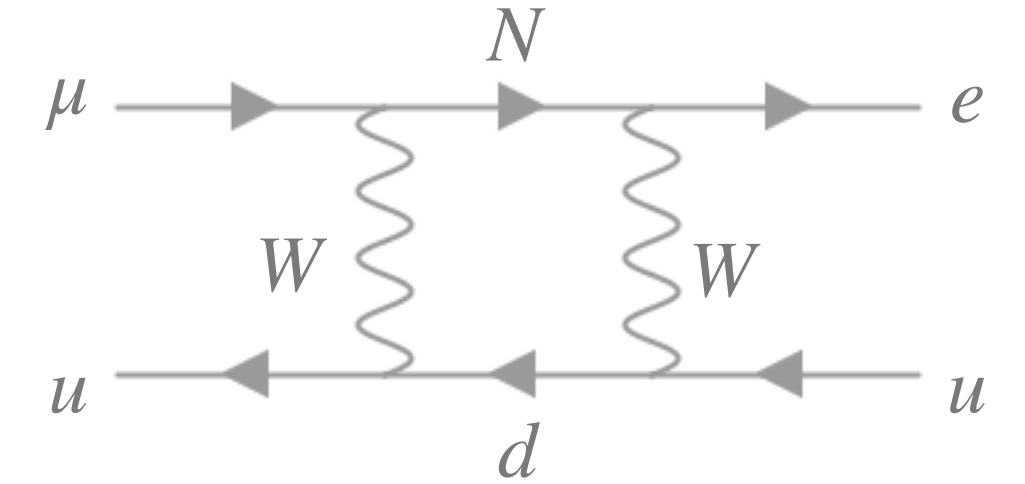
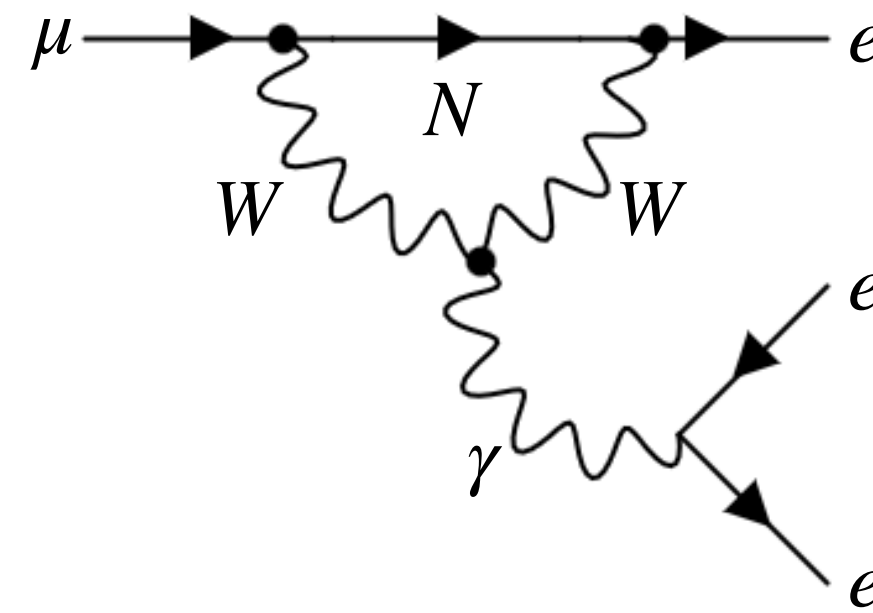
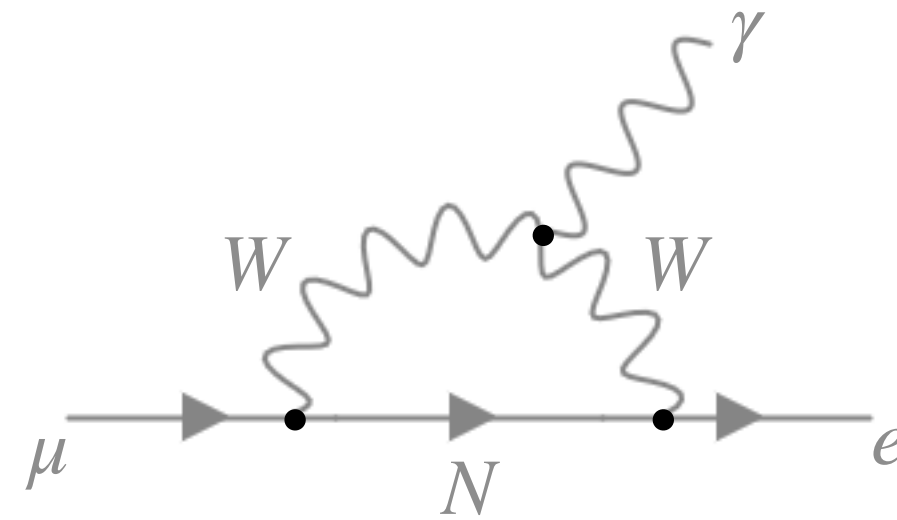
Relevant (approximated) loop functions:

$$\left. \begin{aligned}
 G_{\gamma}^{\beta\alpha} &\approx \eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{\mu_0}{M_0}\right) \\
 F_{\gamma}^{\beta\alpha} &\approx -2 \left( \frac{7}{12} + \frac{1}{6} \log x_0 \right) \eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{m\mu_0}{M_0}\right) \\
 F_Z^{\beta\alpha} &\approx \eta_{\alpha\beta} (5 - 3 \log x_0) + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{\mu_0}{M_0}\right) \\
 F_{box}^{\beta 3\alpha} &\approx -2\eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y^2\mu_0}{M_0}, \frac{\mu_0^2}{M_0^2}\right)
 \end{aligned} \right\} \text{Photon Penguins}$$

$$F_{box}^{\mu e u u} = 2\eta_{e\mu} \left[ -4 - |V_{ub}|^2 (F_{box}(0, x_b) - F_{box}(x_0, x_b) - 4) \right]$$

$$F_{box}^{\mu e d d} = 2\eta_{e\mu} \left[ 1 - |V_{td}|^2 (F_{Xbox}(0, x_t) - F_{Xbox}(x_0, x_t) + 1) \right]$$

# cLFV in the ISS



Relevant (approximated) loop functions:

$$G_{\gamma}^{\beta\alpha} \approx \eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{\mu_0}{M_0}\right)$$

$$F_{\gamma}^{\beta\alpha} \approx -2 \left( \frac{7}{12} + \frac{1}{6} \log x_0 \right) \eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{m\mu_0}{M_0}\right)$$

} Photon Penguins

$$F_Z^{\beta\alpha} \approx \eta_{\alpha\beta} (5 - 3 \log x_0) + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{\mu_0}{M_0}\right)$$

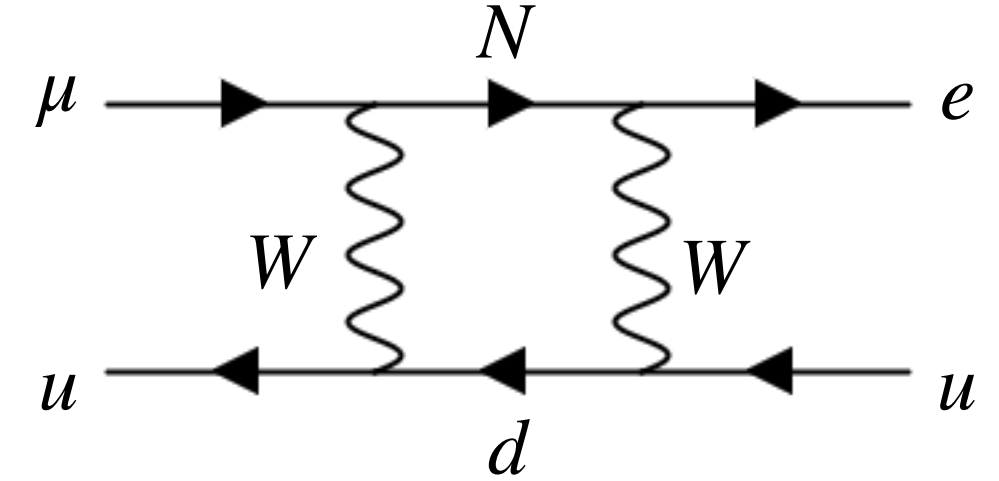
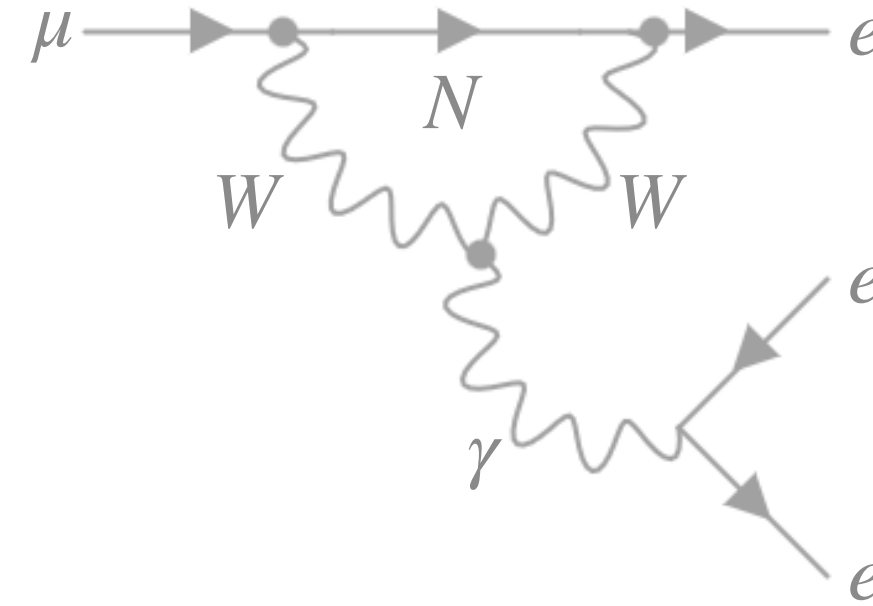
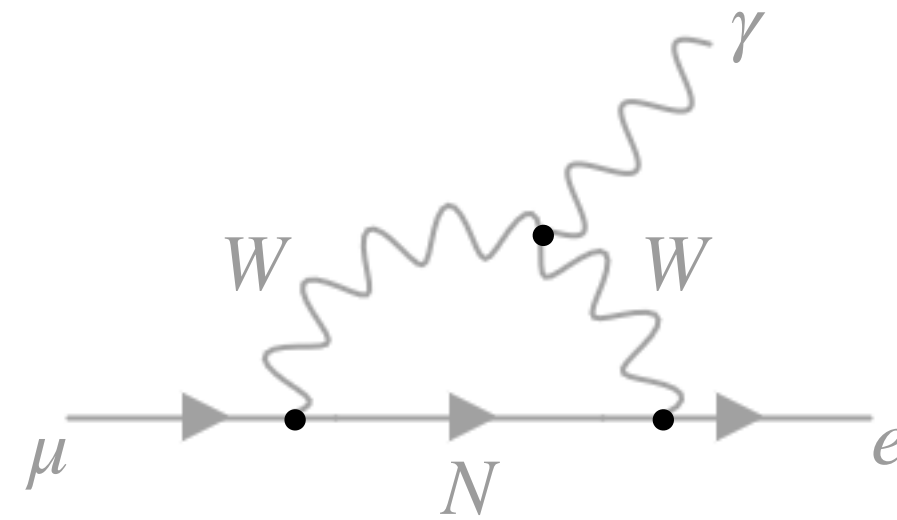
} Z Penguin

$$F_{box}^{\beta 3\alpha} \approx -2\eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y^2\mu_0}{M_0}, \frac{\mu_0^2}{M_0^2}\right)$$

$$F_{box}^{\mu e u u} = 2\eta_{e\mu} \left[ -4 - |V_{ub}|^2 (F_{box}(0, x_b) - F_{box}(x_0, x_b) - 4) \right]$$

$$F_{box}^{\mu e d d} = 2\eta_{e\mu} \left[ 1 - |V_{td}|^2 (F_{Xbox}(0, x_t) - F_{Xbox}(x_0, x_t) + 1) \right]$$

# cLFV in the ISS



Relevant (approximated) loop functions:

$$G_{\gamma}^{\beta\alpha} \approx \eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{\mu_0}{M_0}\right)$$

$$F_{\gamma}^{\beta\alpha} \approx -2 \left( \frac{7}{12} + \frac{1}{6} \log x_0 \right) \eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{m u_0}{M_0}\right)$$

} Photon Penguins

$$F_Z^{\beta\alpha} \approx \eta_{\alpha\beta} (5 - 3 \log x_0) + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{\mu_0}{M_0}\right)$$

} Z Penguin

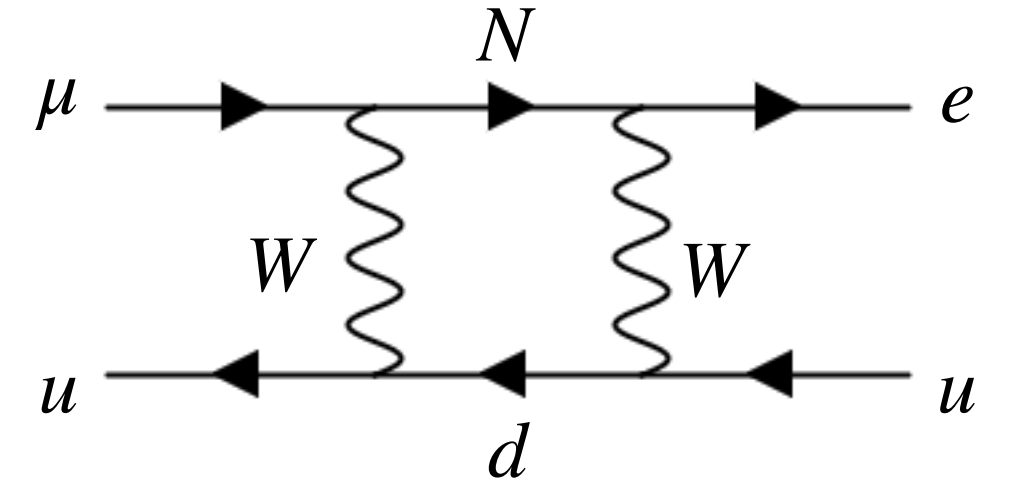
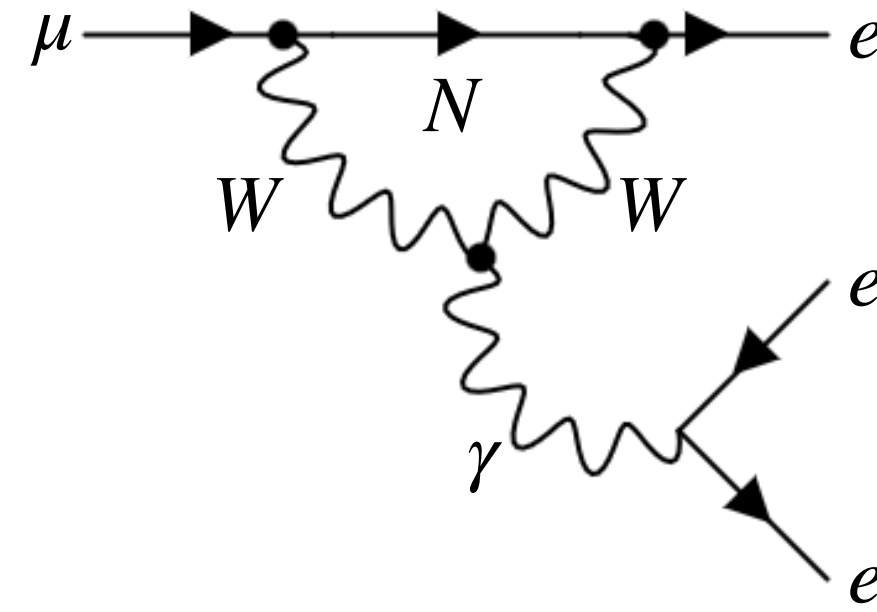
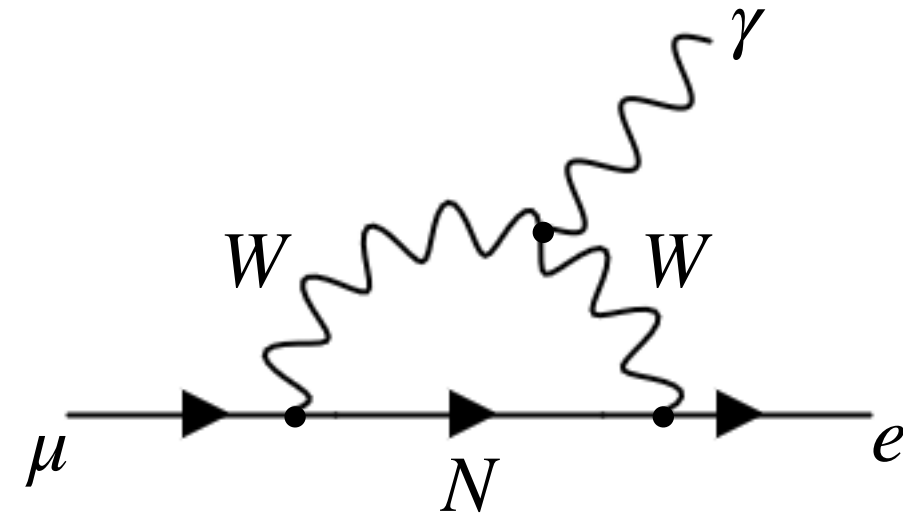
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} Box Diagrams

# cLFV in the ISS



Relevant (approximated) loop functions:

$$G_{\gamma}^{\beta\alpha} \approx \eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{\mu_0}{M_0}\right)$$

$$F_{\gamma}^{\beta\alpha} \approx -2 \left( \frac{7}{12} + \frac{1}{6} \log x_0 \right) \eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{m\mu_0}{M_0}\right)$$

$$F_Z^{\beta\alpha} \approx \eta_{\alpha\beta} (5 - 3 \log x_0) + \mathcal{O}\left(\frac{Y\mu_0}{M_0}, \frac{\mu_0}{M_0}\right)$$

$$F_{box}^{\beta3\alpha} \approx -2\eta_{\alpha\beta} + \mathcal{O}\left(\frac{Y^2\mu_0}{M_0}, \frac{\mu_0^2}{M_0^2}\right)$$

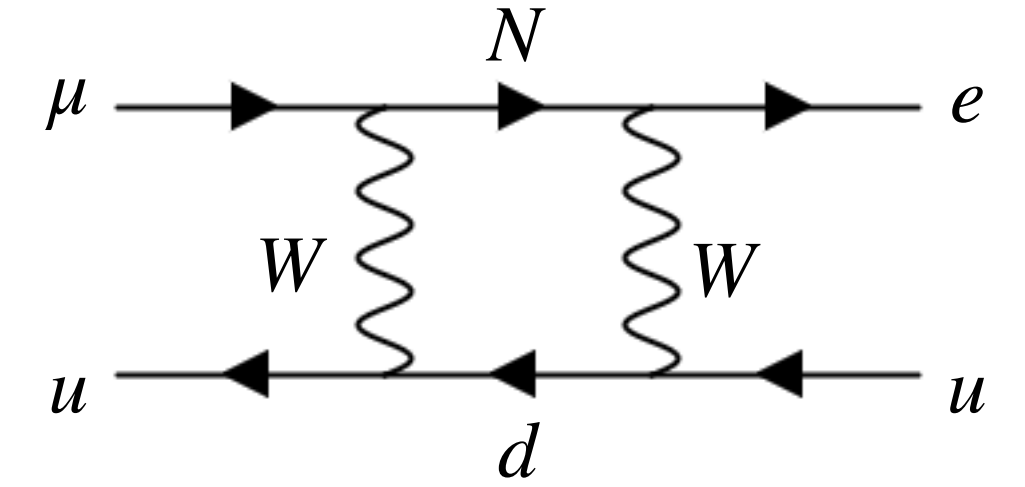
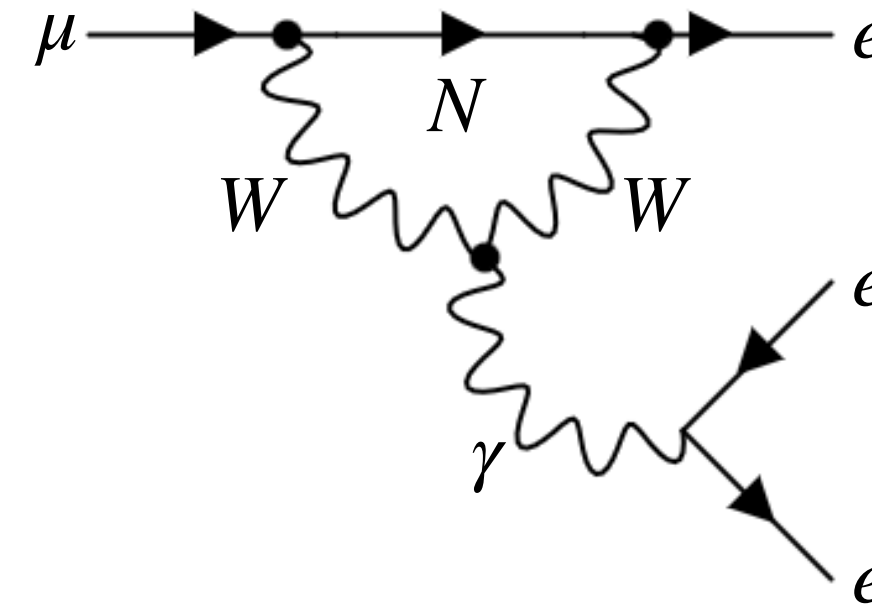
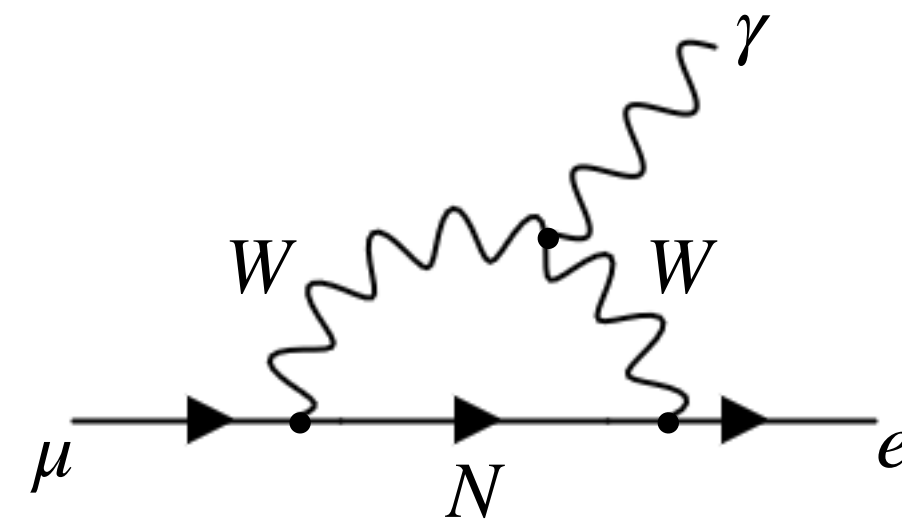
$$F_{box}^{\mu e u u} = 2\eta_{e\mu} \left[ -4 - |V_{ub}|^2 (F_{box}(0, x_b) - F_{box}(x_0, x_b) - 4) \right]$$

$$F_{box}^{\mu e d d} = 2\eta_{e\mu} \left[ 1 - |V_{td}|^2 (F_{Xbox}(0, x_t) - F_{Xbox}(x_0, x_t) + 1) \right]$$

Rates of the processes are proportional to  $\eta$  at LO in ISS framework



# cLFV in the ISS



Relevant (approximated) loop functions:

For a particular choice of residual symmetry (Case 2):

$F_{\gamma}^{\beta\alpha}$

$$\eta_{\mu e} = \frac{1}{6} \eta'_0 \left[ 2\Delta y_{21}^2 - \Delta y_{31}^2 \left( 1 - \cos(2\theta_L) \cos \phi_u - \sqrt{3} \left( \cos(2\theta_L) \sin \phi_u + i \sin(2\theta_L) \right) \right) \right]$$

$$F_{box}^{\mu e u u} = 2\eta_{e\mu} \left[ -4 - |V_{ub}|^2 \left( F_{box}(0, x_b) - F_{box}(x_0, x_b) - 4 \right) \right]$$

$$F_{box}^{\mu e d d} = 2\eta_{e\mu} \left[ 1 - |V_{td}|^2 \left( F_{Xbox}(0, x_t) - F_{Xbox}(x_0, x_t) + 1 \right) \right]$$

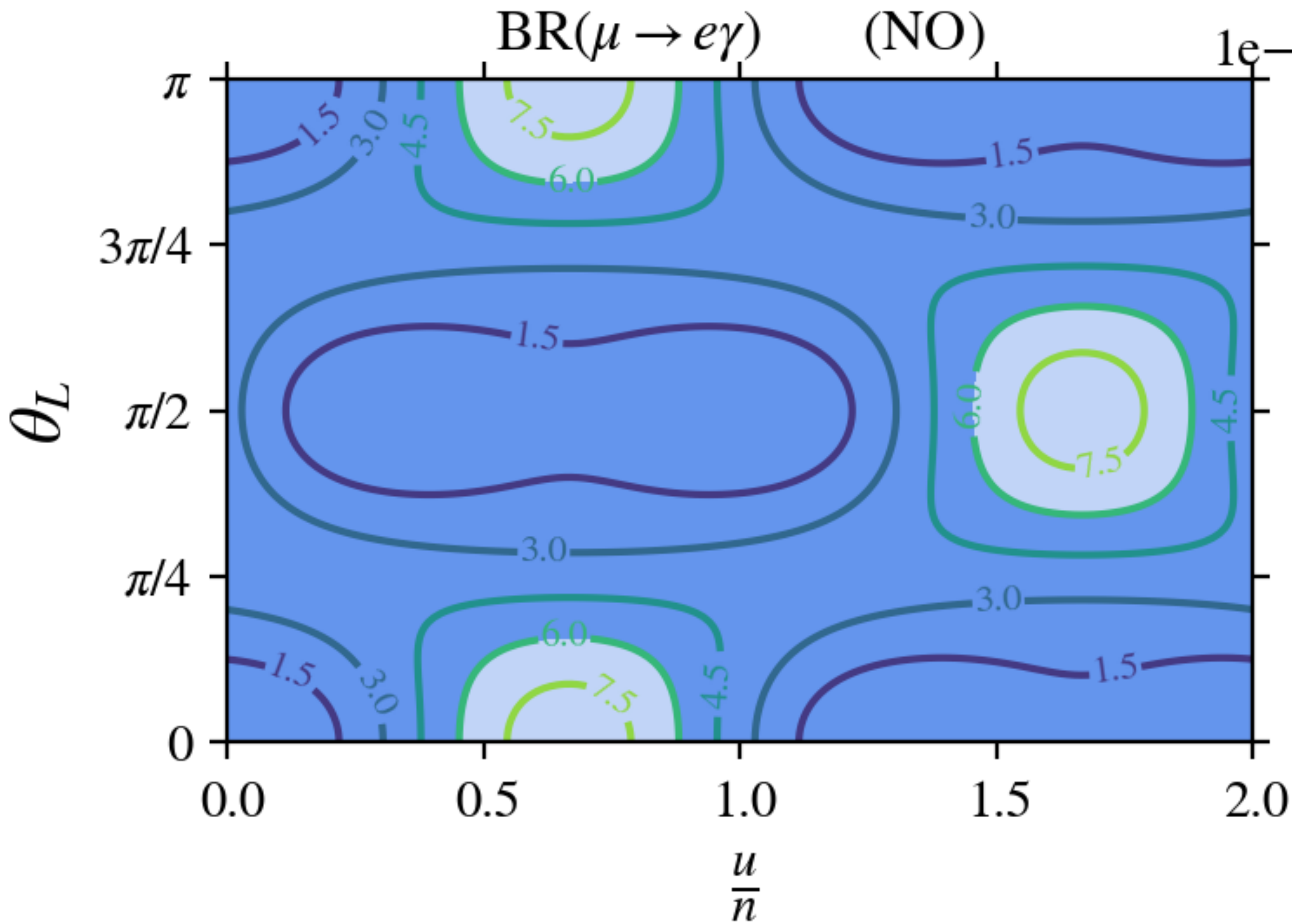
# Option 2: Case 2

$$M_0 = 3 \text{ TeV}$$

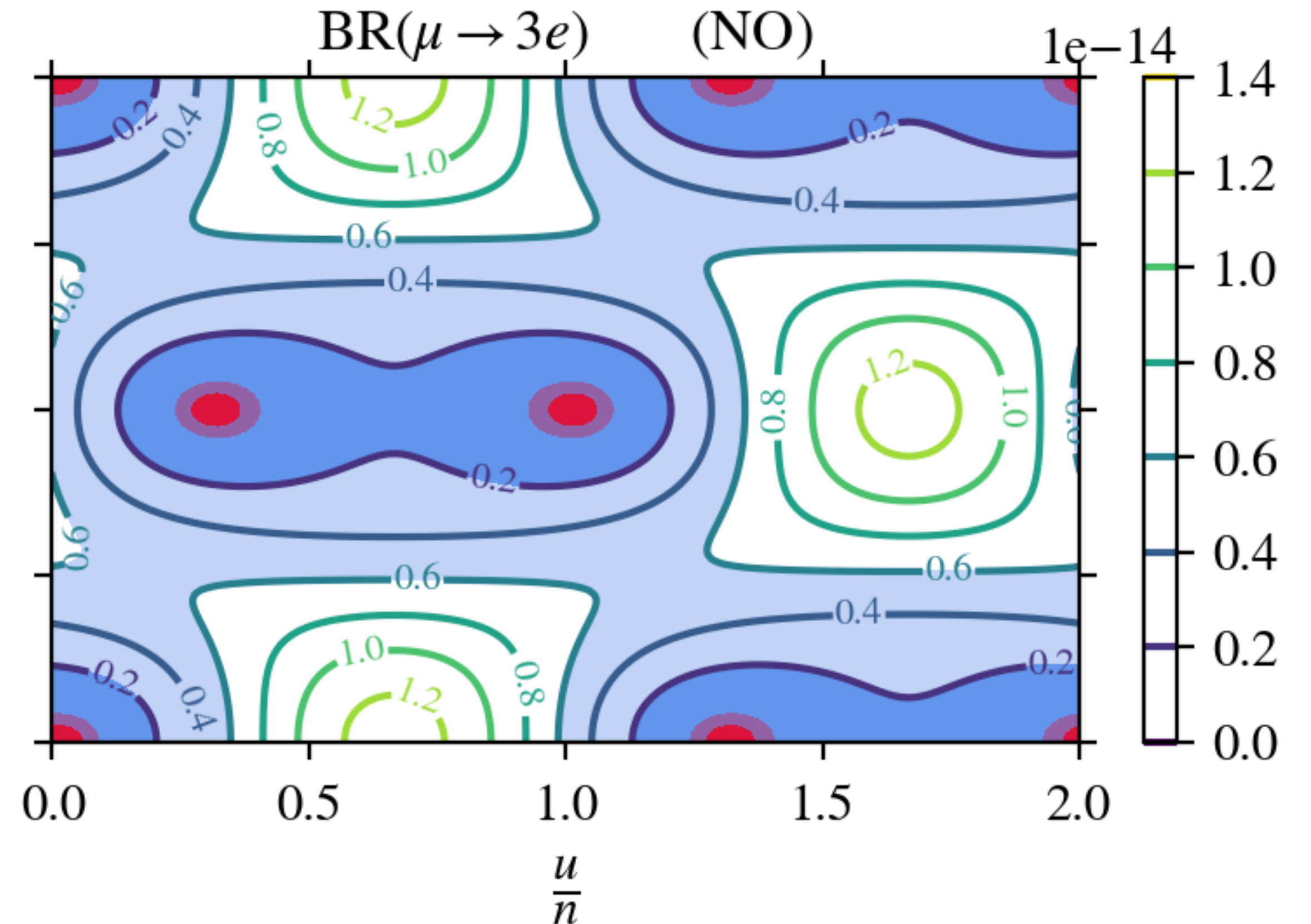
$$\mu_0 = 1 \text{ keV}$$

$$(u = 2s - t)$$

$t$  even



$$BR(\mu \rightarrow e\gamma) \lesssim 6 \times 10^{-14} \quad \text{Meg-II}$$



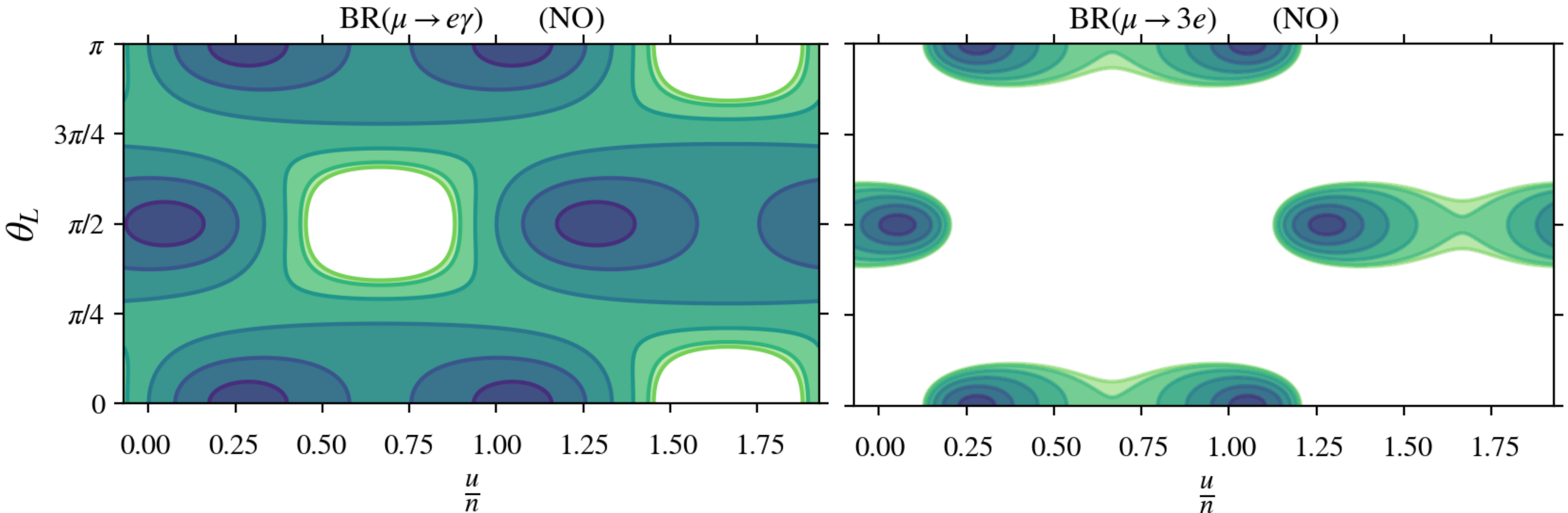
$$BR(\mu \rightarrow 3e) \lesssim 20(1) \times 10^{-16} \quad \text{Mu4E Phase-I (II)}$$

# Option 2: Case 2

- $\theta_R = 0.785$
- $\theta_R = 0.654$
- $\theta_R = 0.524$
- $\theta_R = 0.393$
- $\theta_R = 0.262$
- $\theta_R = 0.131$

$$M_0 = 3 \text{ TeV}$$
$$\mu_0 = 1 \text{ keV}$$

$t$  odd



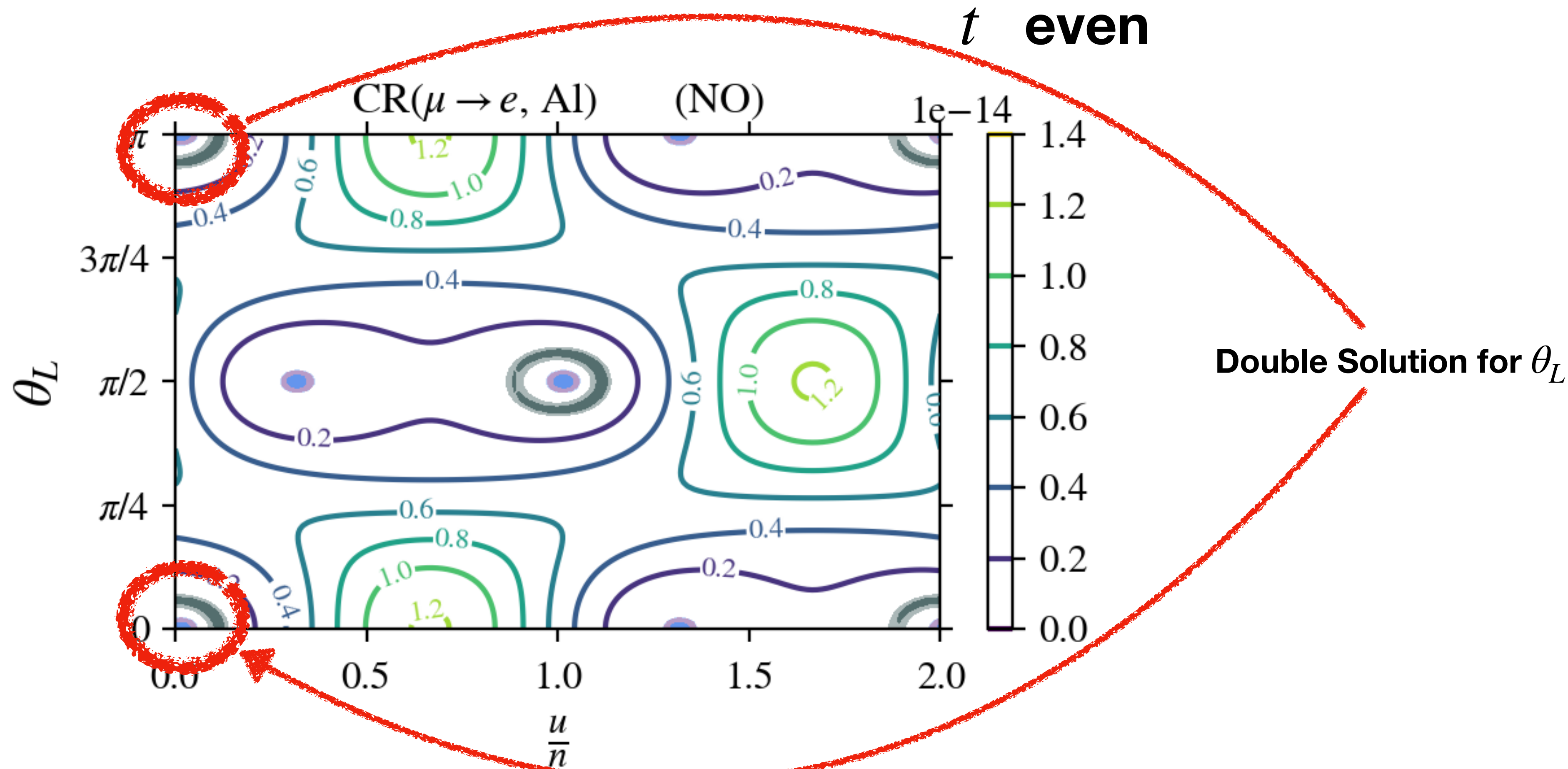
$BR(\mu \rightarrow e\gamma) \lesssim 6 \times 10^{-14}$  **Meg-II**

$BR(\mu \rightarrow 3e) \lesssim 20 \times 10^{-16}$  **Mu4E Phase-I (II)**



# Option 2: Case 2

$$M_0 = 3 \text{ TeV}$$
$$\mu_0 = 1 \text{ keV}$$



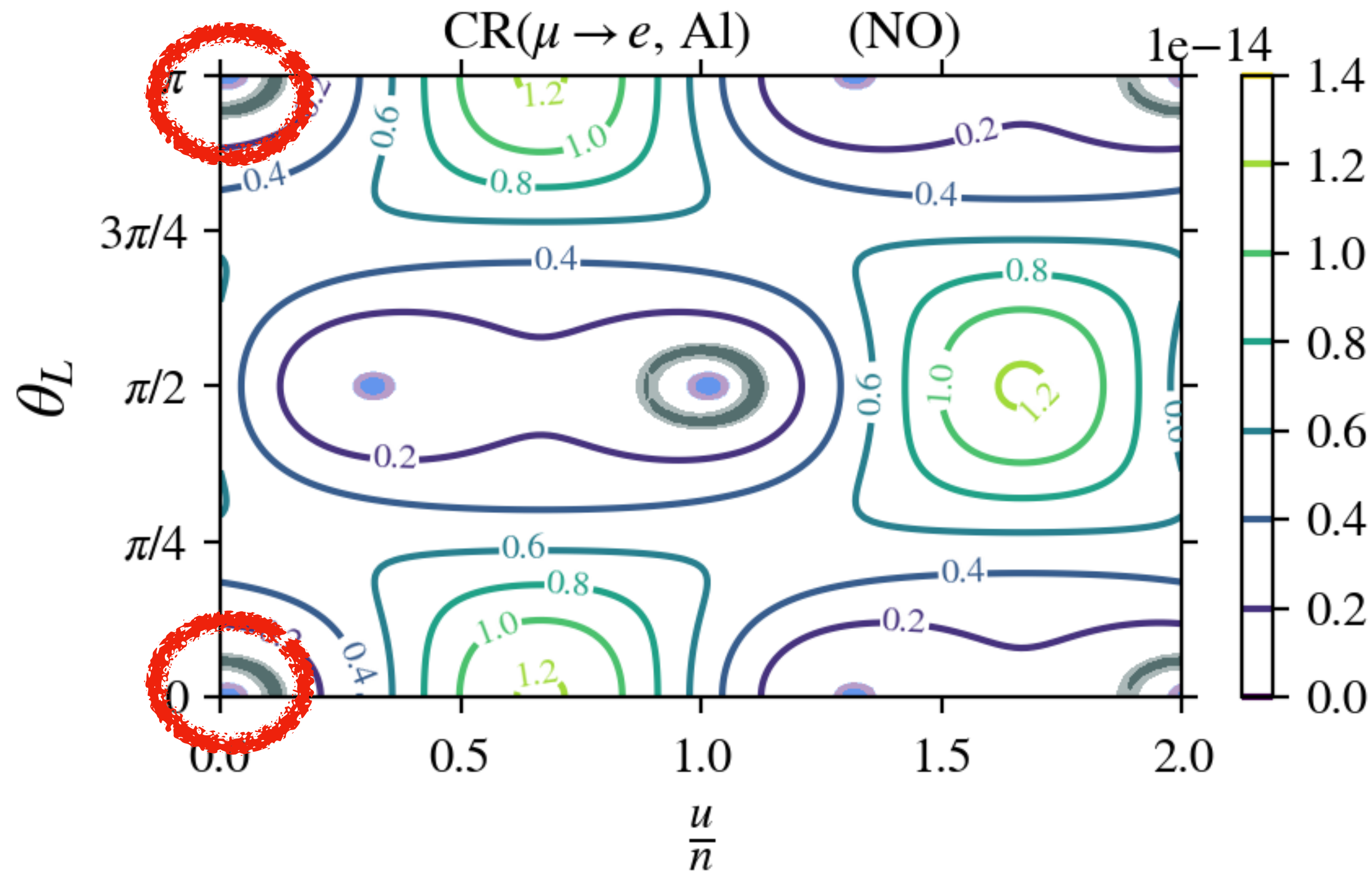
$$CR(\mu \rightarrow e\gamma) \lesssim 2.6(2.9) \times 10^{-17} \quad \text{COMET Phase-II (Mu2E)}$$

# Option 2: Case 2

$$M_0 = 3 \text{ TeV}$$

$$\mu_0 = 1 \text{ keV}$$

$t$  even



	$\sin^2 \theta_{12}$	$\sin^2 \theta_{13}$	$\sin^2 \theta_{23}$
$n = 14; s = t = 1$	0.34	0.021(0.022)	0.559(0.561)
$n = 14; s = t = 0$	0.341	0.022	0.5
$n = 14; s = 0; t = 1$	0.34	0.022	0.436
$n = 14; s = 1; t = 2$	0.341	0.022	0.5

$$\theta_L \approx 0.183 \text{ (2.959)}$$

Best fit results: ( $\chi_{mix}^2 \leq 12$ )

$$CR(\mu \rightarrow e\gamma) \lesssim 2.6(2.9) \times 10^{-17}$$

COMET Phase-II (Mu2E)

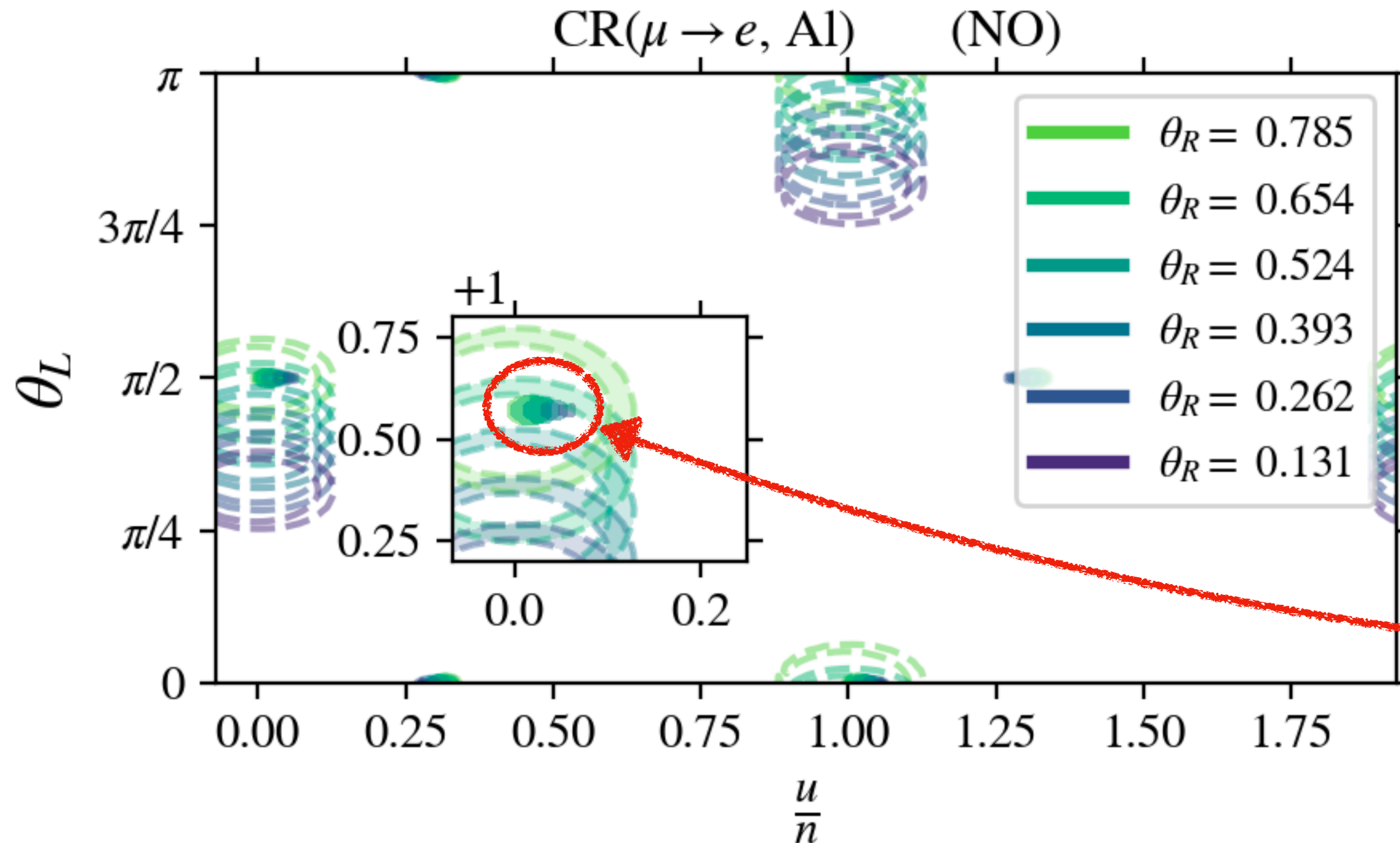


# Option 2: Case 2

$$M_0 = 3 \text{ TeV}$$

$$\mu_0 = 1 \text{ keV}$$

$t$  odd



$\theta_R$  Shifts best-fit regions upward:

$$\tan(\delta_\theta) = - \frac{y_i y_j \cot(2\theta_R)}{y_i^2 + y_j^2}$$

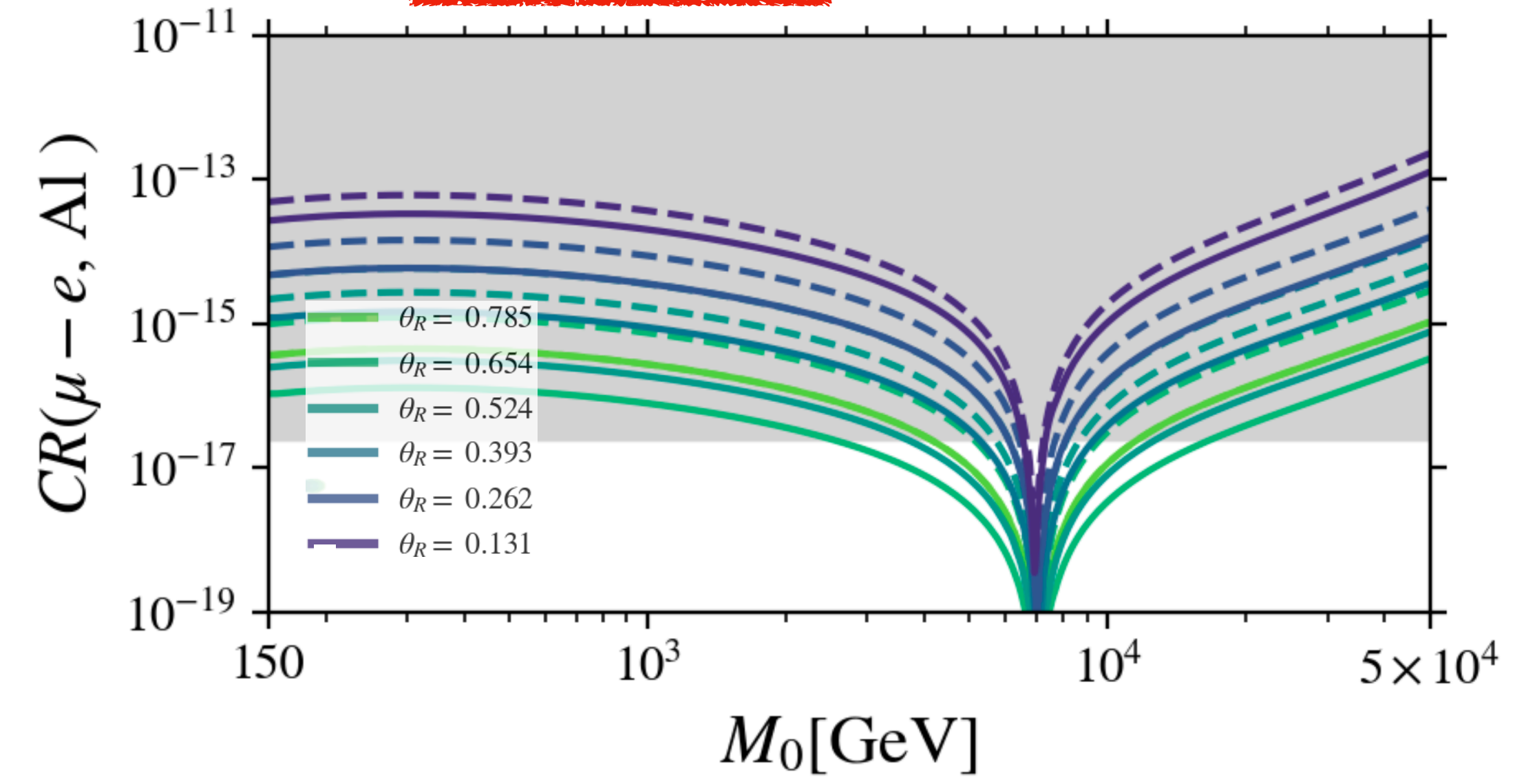
Different choices of  $\theta_R$  give different predictions for cLFV

$$CR(\mu \rightarrow e\gamma) \lesssim 2.6 \times 10^{-17}$$

**COMET Phase-II**

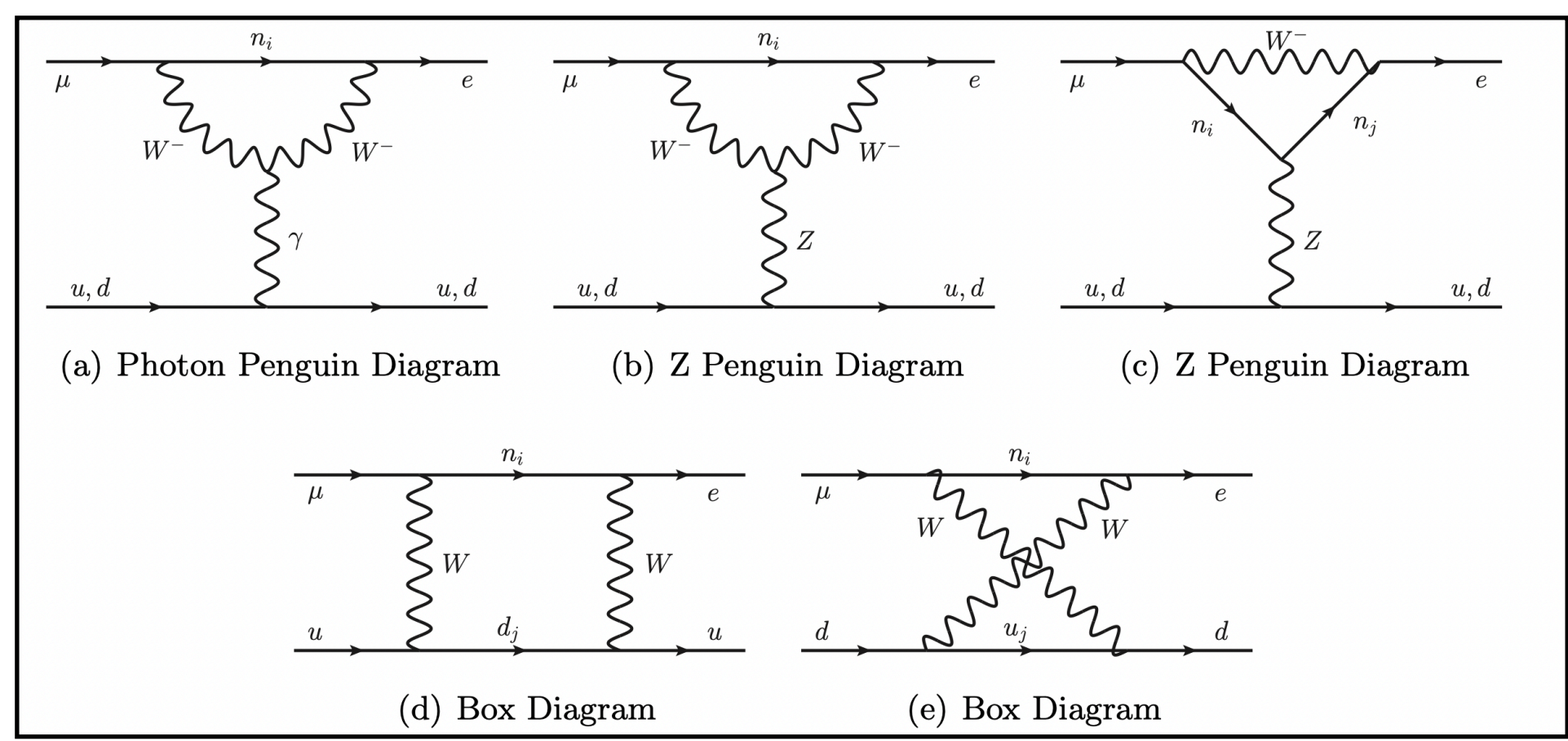
# Option 2: Case 2

$\mu_0 = 1 \text{ keV}$



$$CR(\mu - e, Al) = \frac{2G_F^2 \alpha_w^2 m_\mu^5}{(4\pi)^2 \Gamma_{capt}} \left| 4V^{(p)} \left( 2\tilde{F}_u^{\mu e} + \tilde{F}_d^{\mu e} \right) + 4V^{(n)} \left( \tilde{F}_u^{\mu e} + 2\tilde{F}_d^{\mu e} \right) + s_w^2 \frac{G_\gamma^{\mu e} D}{2e} \right|^2$$

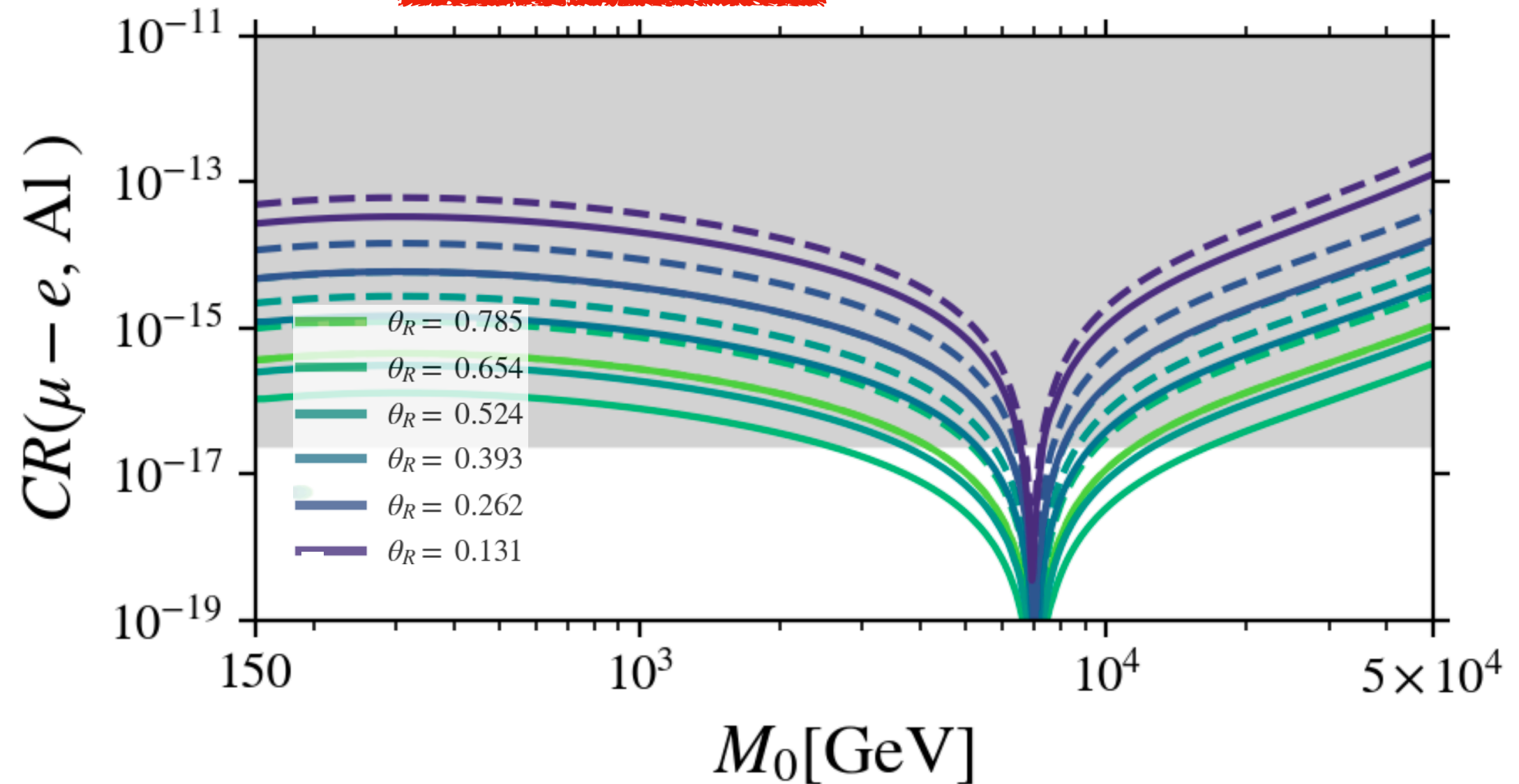
Different solutions for  $\theta_L$  (at fixed  $\theta_R$ ) also give different predictions for cLFV





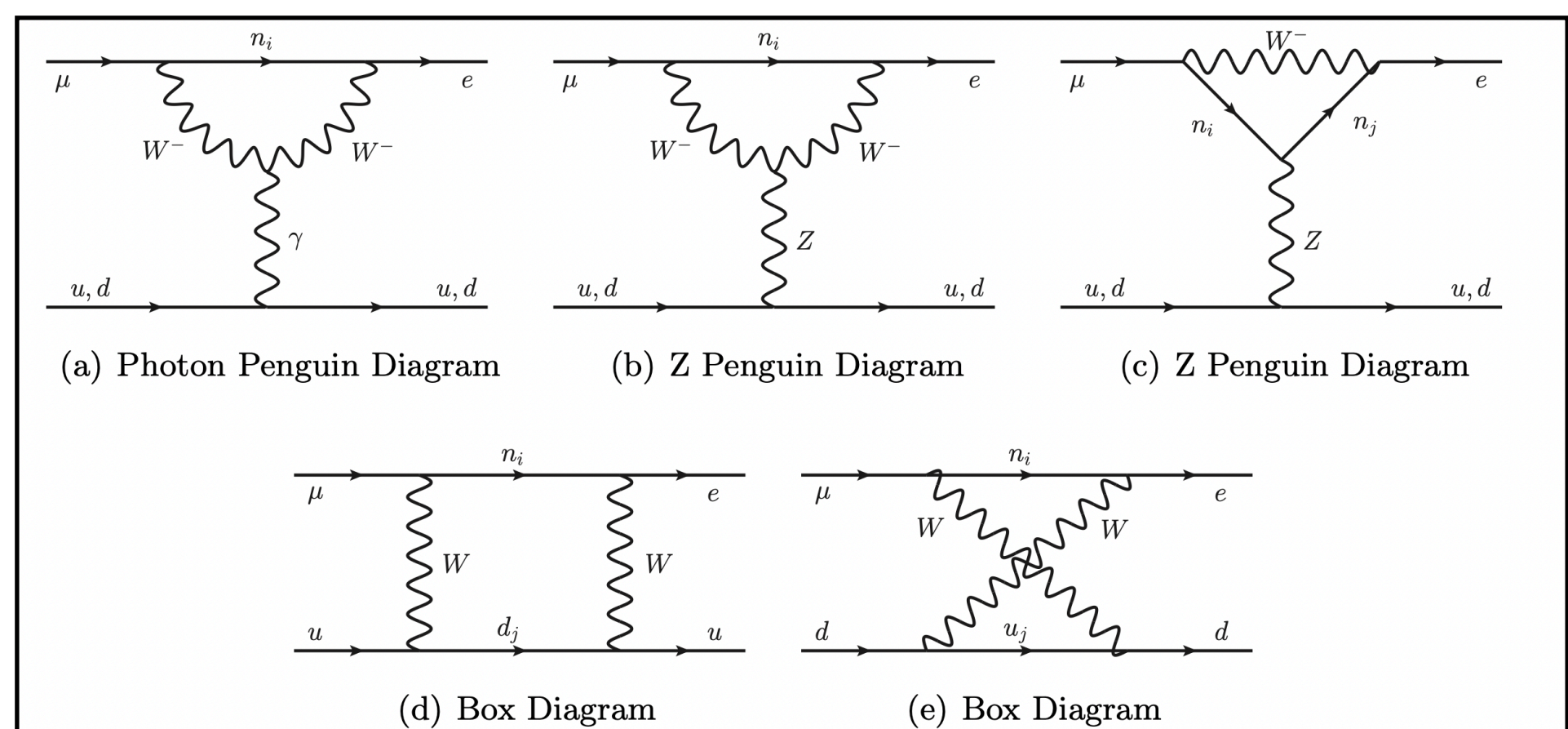
# Option 2: Case 2

$\mu_0 = 1 \text{ keV}$



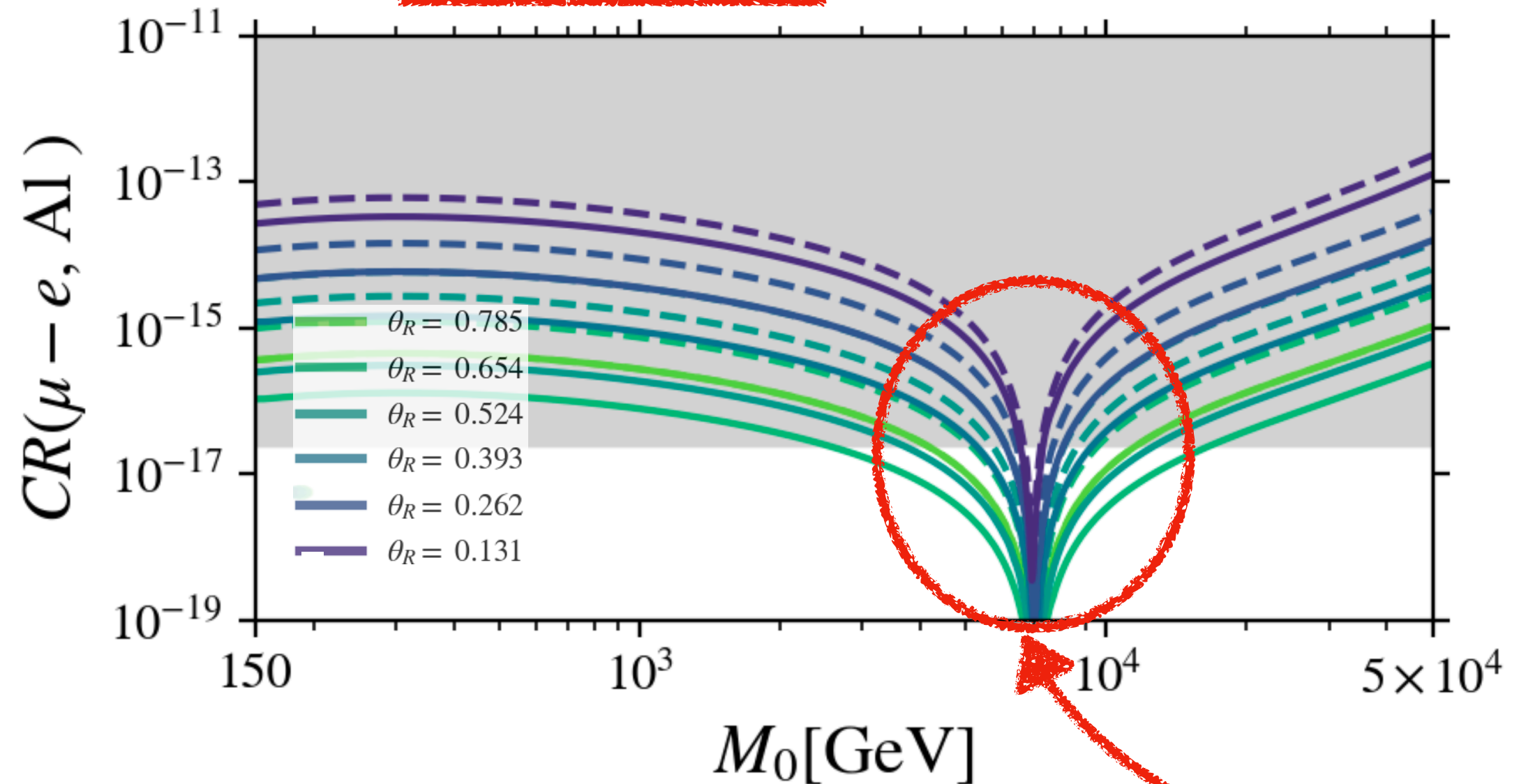
$$CR(\mu - e, Al) = \frac{2G_F^2 \alpha_w^2 m_\mu^5}{(4\pi)^2 \Gamma_{capt}} \left| 4V^{(p)} \left( 2\tilde{F}_u^{\mu e} + \tilde{F}_d^{\mu e} \right) + 4V^{(n)} \left( \tilde{F}_u^{\mu e} + 2\tilde{F}_d^{\mu e} \right) + s_w^2 \frac{G_\gamma^{\mu e} D}{2e} \right|^2$$

Up and down quark contributions have different sign due to different charge and weak isospin



# Option 2: Case 2

$\mu_0 = 1 \text{ keV}$

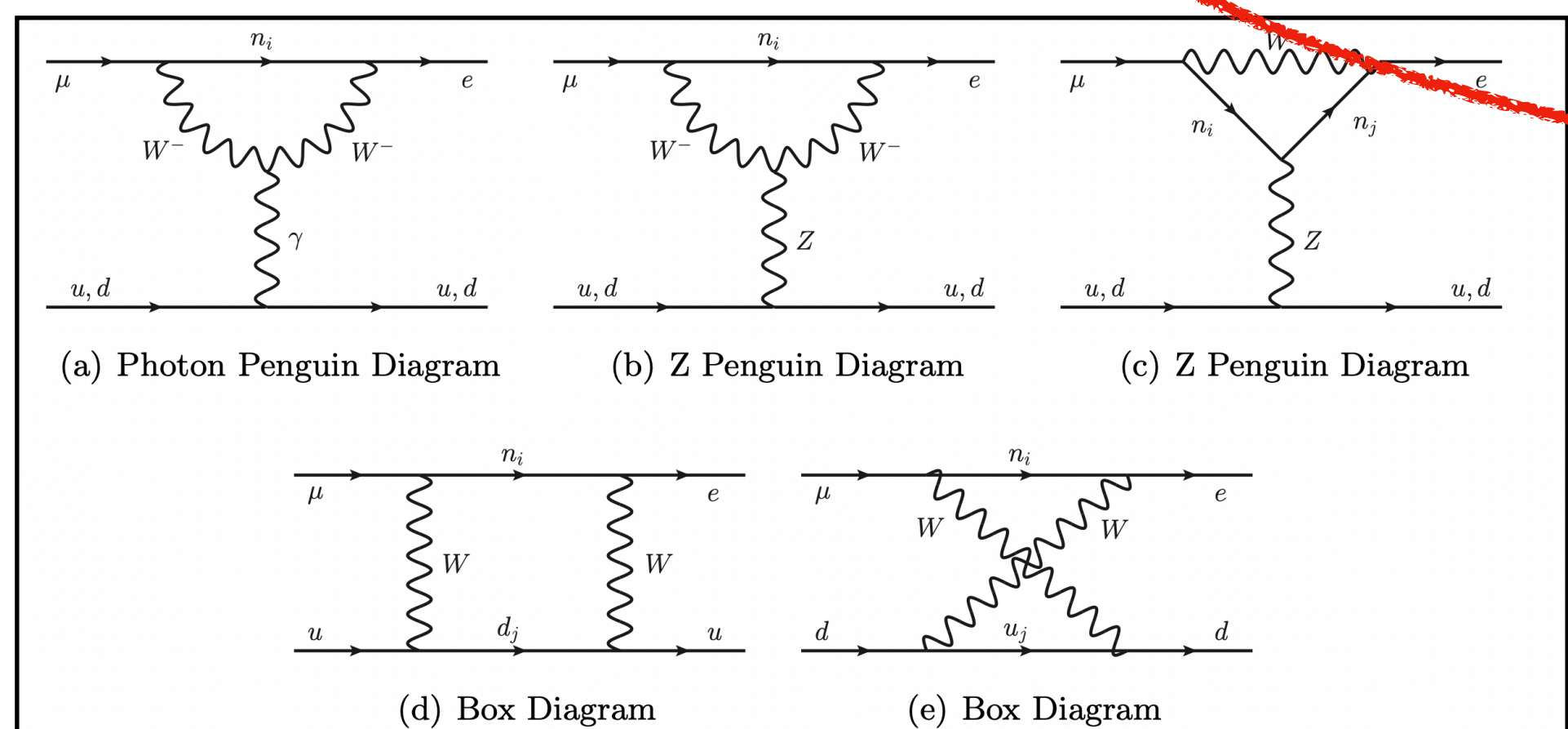


$$CR(\mu - e, Al) = \frac{2G_F^2 \alpha_w^2 m_\mu^5}{(4\pi)^2 \Gamma_{capt}} \left| 4V^{(p)} \left( 2\tilde{F}_u^{\mu e} + \tilde{F}_d^{\mu e} \right) + 4V^{(n)} \left( \tilde{F}_u^{\mu e} + 2\tilde{F}_d^{\mu e} \right) + s_w^2 \frac{G_\gamma^{\mu e} D}{2e} \right|^2$$

Up and down quark contributions have different sign due to different charge and weak isospin

Cancellation only depend on  $M_0$ :

$$M_0^2 = \exp \left( \frac{\frac{9}{8} V^{(n)} + \left( \frac{9}{8} + \frac{37}{12} s_w^2 \right) V^{(p)} - \frac{s_w^2}{16e} D}{\frac{3}{8} V^{(n)} + \left( \frac{4s_w^2}{3} - \frac{3}{8} \right) V^{(p)}} \right) M_W^2$$





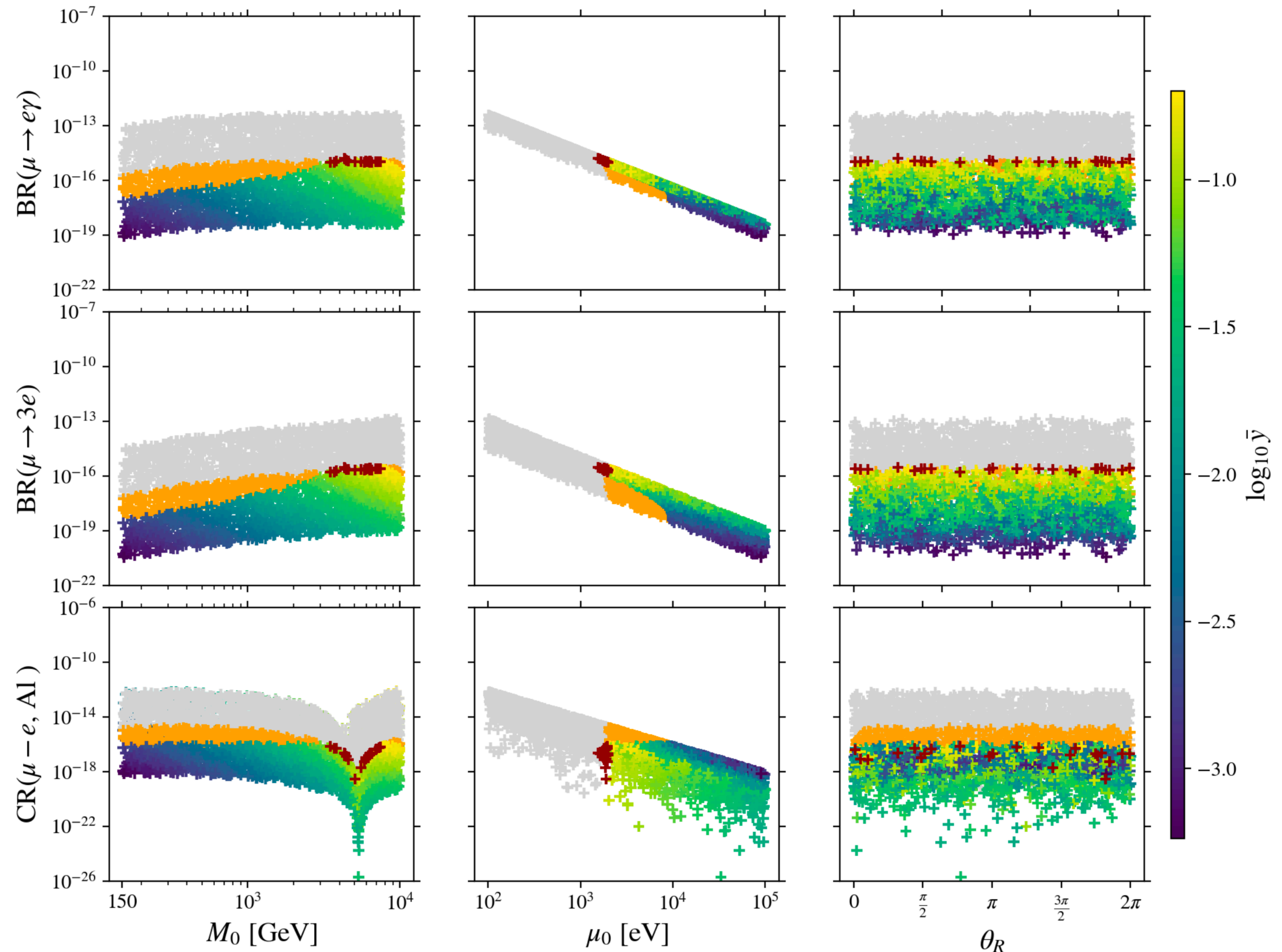
# Option 2: Case 2

Predictions are bounded by both experimental bounds on **cLFV processes** as by bounds on **unitarity violation**

$$\eta \lesssim \left( \begin{array}{ccc} 1.3 \times 10^{-3} & & \\ 1.2 \times 10^{-5} & 1.1 \times 10^{-5} & \\ 9.0 \times 10^{-4} & 5.7 \times 10^{-5} & 1.0 \times 10^{-3} \end{array} \right)$$

M. Blennow, E. Fernández-Martínez, J. Hernández-García, J. López-Pavón, X., D. Naredo-Tuero; 2306.01040

Case 2),  $s=t=0$ , NO



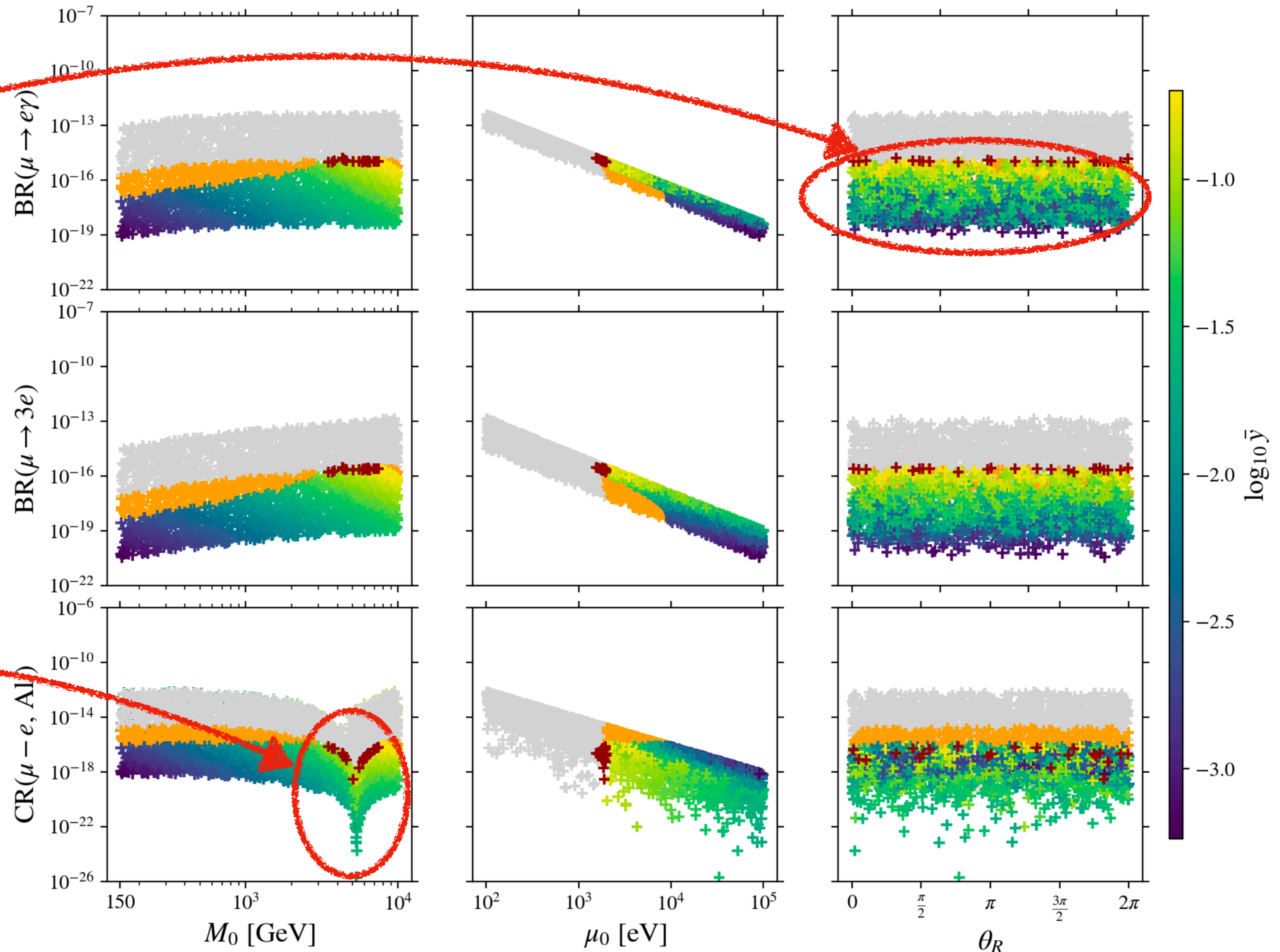


# Option 2: Case 2

Case 2),  $s=t=0$  , NO

No dependence from  $\theta_R$

Cancellation of contributions

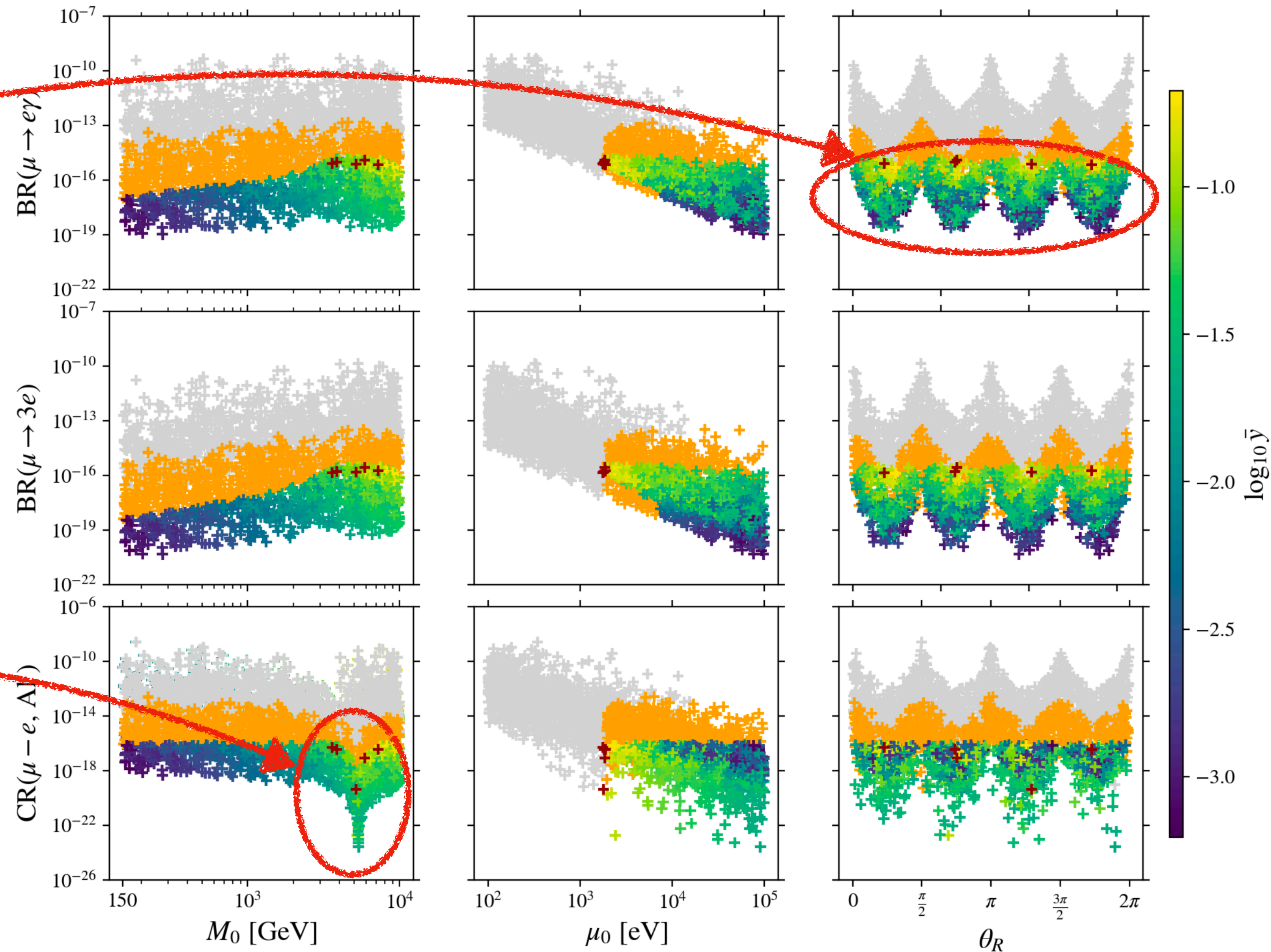


# Option 2: Case 2

Case 2),  $s=0, t=1$ , NO

Modulation in function of  $\theta_R$

Cancellation of contributions

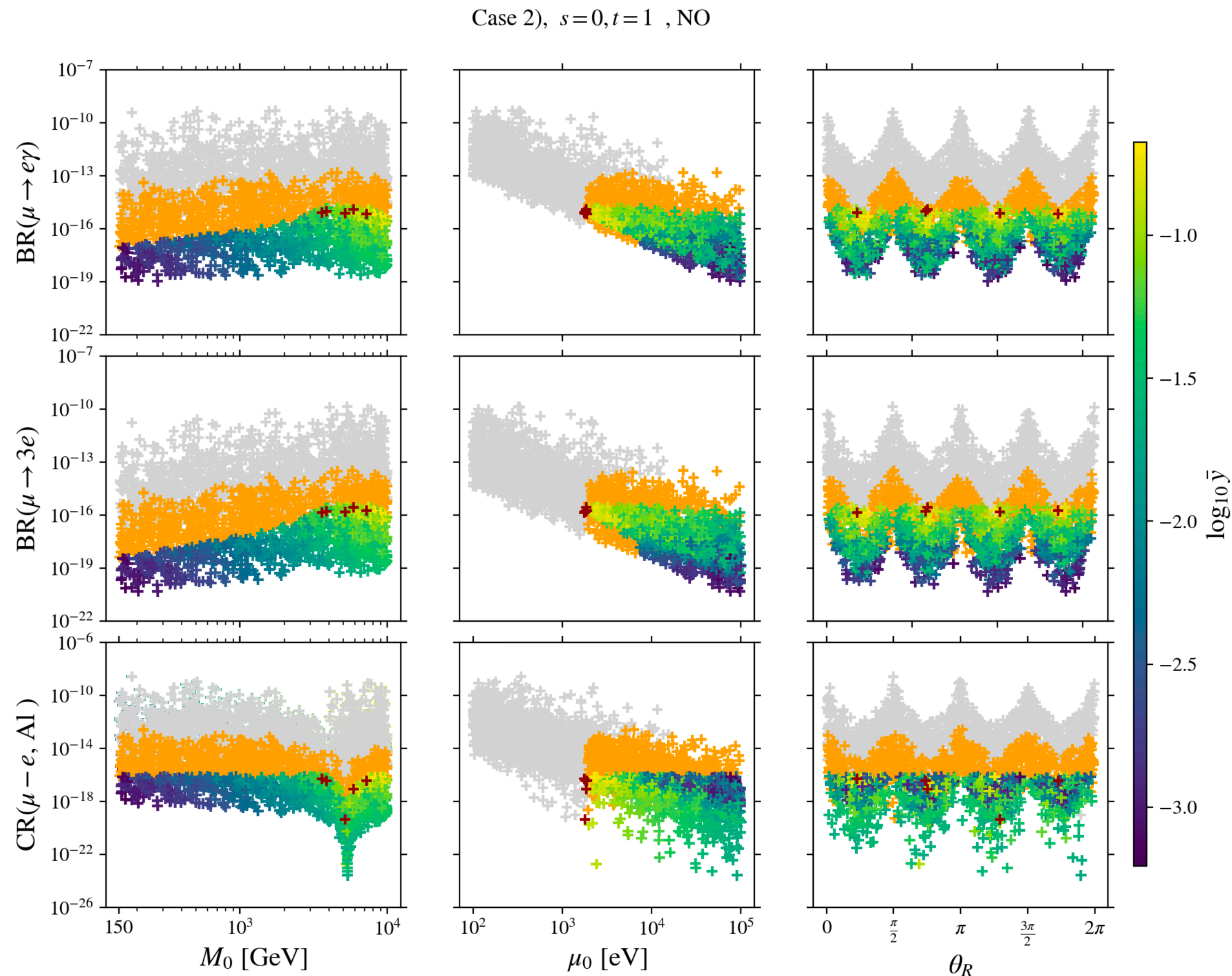




# Option 2: Case 2

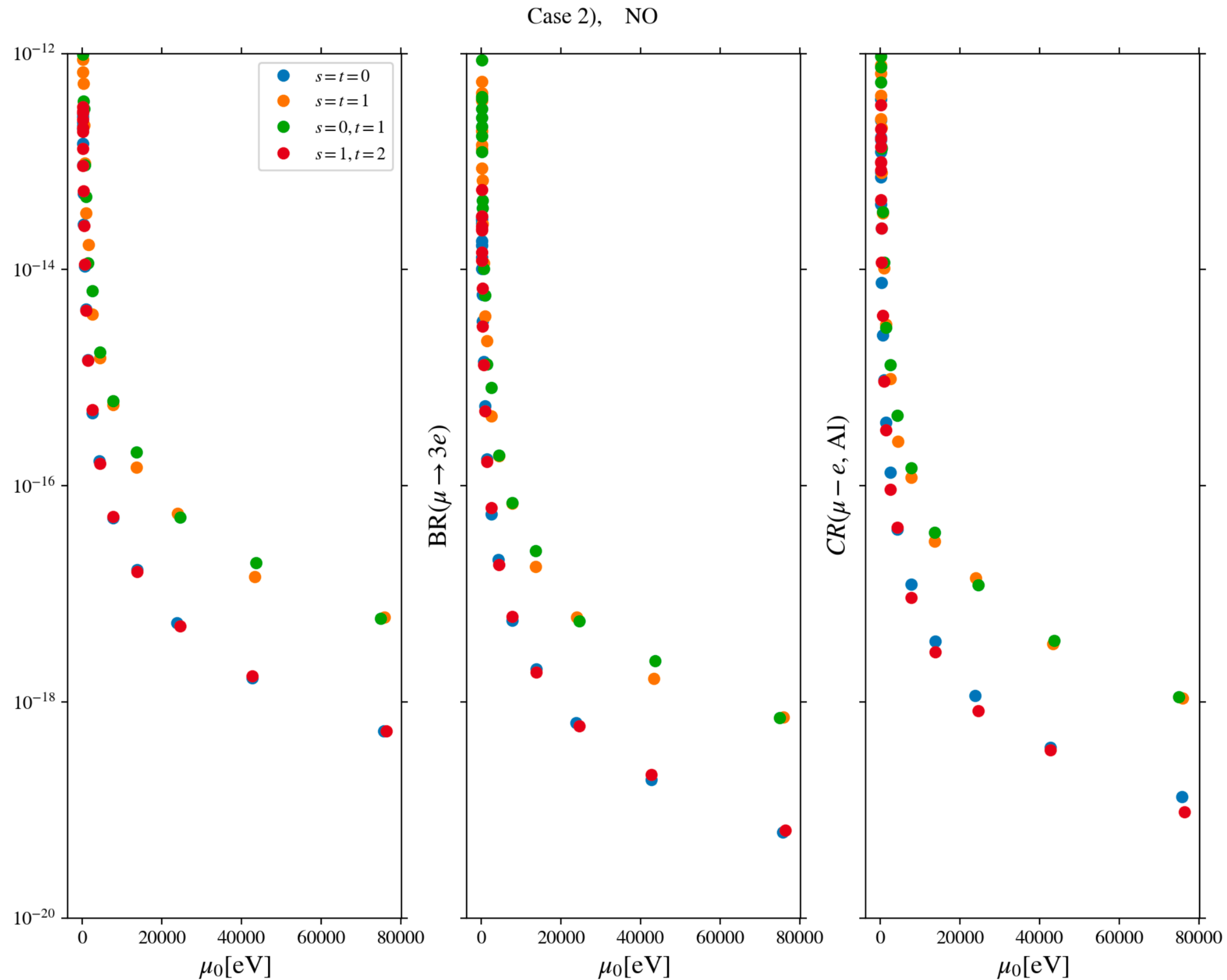
Predictions are compatible with future bounds on  $\mu - e$  transitions!

A lower limit for  $\mu_0$  can be extrapolated





# Option 2: Case 2



Binned averages of the scans:

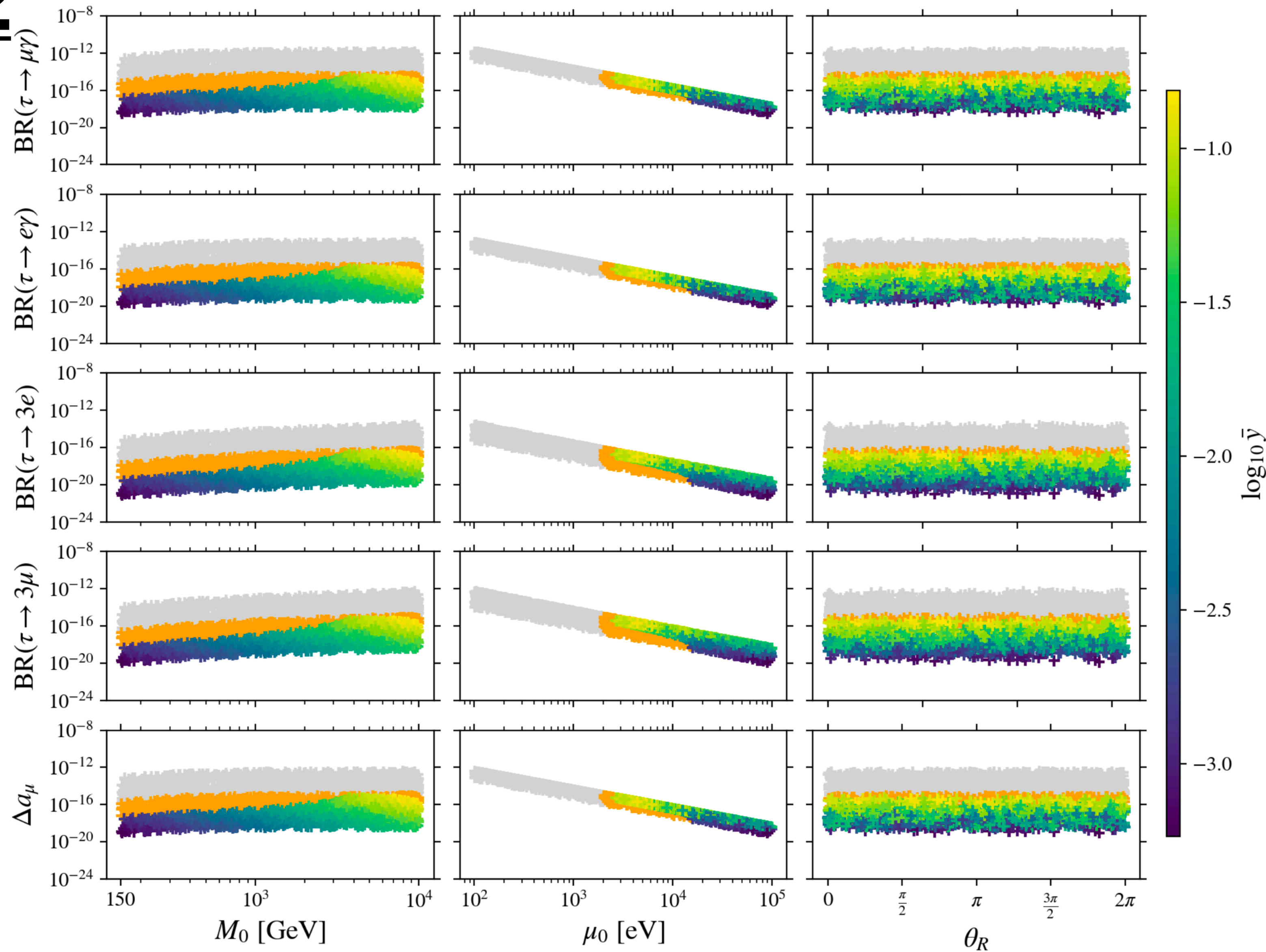
- $\theta_R$ -dependent case predicts slightly larger rates due to larger number of contributions to  $\eta_{e\mu}$
- Different choices of residual symmetry can lead to different predictions



# Option 2: Case 2

Case 2),  $n=14, t=0, s=0$ , NO

$\tau - \mu$  and  $\tau - e$  processes are non-constraining



# Conclusions:

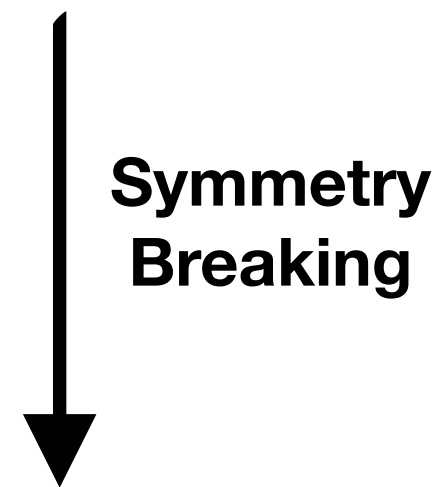
- ISS Framework is promising framework, due to the possibility of testing it at future experiments
- Flavour and  $CP$  symmetries are successful in reproducing lepton mixing data and light neutrino masses
- cLFV analysis in the ISS framework parameter space
- Other cLFV processes such as  $\tau - \mu$  and  $\tau - e$  transitions are still far from probing our parameter space

**Thank You For the Attention!**  
**Questions? Doubts?**

**Backup Slides**

In our theory:

- $e_R, \mu_R, \tau_R$  transform under an auxiliary  $Z_3^{aux}$
- $L, N, S$  transform as triplets of  $G_f = \Delta(6n^2)$
- $CP$  acts non trivially on flavour



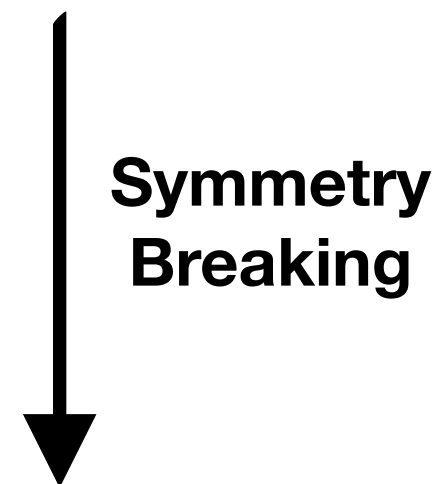
Breaking mechanism is irrelevant. After Breaking:

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$$\begin{cases} e_R \rightarrow e_R \\ \mu_R \rightarrow \omega \mu_R \\ \tau_R \rightarrow \omega^2 \tau_R \end{cases} \quad \text{with} \quad \omega = e^{\frac{2\pi i}{3}}$$

In our theory:

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$$\Delta(6n^2) \sim (\mathcal{I}_n \times \mathcal{I}_n) \rtimes \mathcal{S}_3$$

- Non-abelian subgroup of  $SU(3)$
- Spanned by four generators  $a, b, c, d$

$$a^n = b^n = (ab)^2 = c^n = d^n = 1$$

$$cd = dc$$

$$aca^{-1} = c^{-1}d^{-1}$$

$$ada^{-1} = d$$

$$bcb^{-1} = d^{-1}$$

$$bsdb^{-1} = c^{-1}$$

- General element of the group is written:

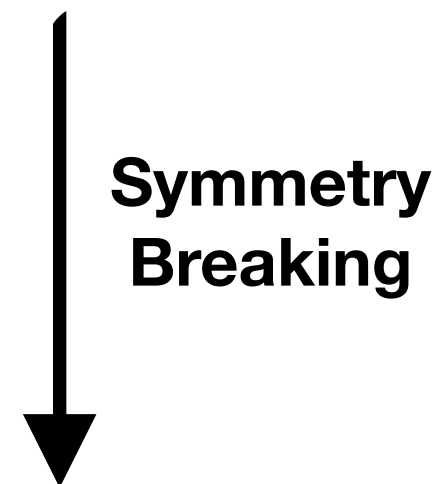
$$g = a^\alpha b^\beta c^\gamma d^\delta \quad \text{with} \quad \alpha, \beta, \gamma, \delta \in N$$

- Equipped with a variety of independent triplet representations



In our theory:

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- $CP$  represented by  $X(r)$  that acts on flavour
- Consistency requires :  $X(r)$  must be an automorphism of the flavour group:

$$\text{for } g(r) \in G_f \Rightarrow X(r)^* g(r) X(r) = g'(r) \in G_f$$

- We always choose:

$$X(r)^* X(r) = X(r) X(r)^* = 1$$

# About Option 2

$$m_\nu \approx \frac{1}{M_0^2} m_D \mu_S m_D^T =$$

$$= \langle H \rangle \frac{\mu_0}{M_0^2} U_D^{(L)*}(\theta_L) \left[ \begin{pmatrix} y_1 & 0 & 0 \\ 0 & y_2 & 0 \\ 0 & 0 & y_3 \end{pmatrix} U_D^{(R)T}(\theta_R) \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} U_D^{(R)}(\theta_R) \begin{pmatrix} y_1 & 0 & 0 \\ 0 & y_2 & 0 \\ 0 & 0 & y_3 \end{pmatrix} \right] U_D^{(L)\dagger}(\theta_L)$$

$$U_D^{(L)}(\theta_L) = \Omega(3) R_{ij}(\theta_L)$$

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The  $\Omega$  matrices:

- **Unitary**
- Depend on the choice of **residual symmetry** and *CP*

Parameters:

- Mass scales:  $M_0, \mu_0$
- Yukawas  $y_1, y_2, y_3$
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If it is diagonal :

- $U_D^{(L)}(\theta_L)$  diagonalises  $m_\nu$
- Mass eigenvalues are  $m_i = \frac{\langle H \rangle^2}{M_0^2} \mu_0$

If not diagonal:

- Redefinition  $\theta_L \rightarrow \tilde{\theta}_L \Rightarrow U_D^{(L)}(\tilde{\theta}_L)$  diagonalises  $m_\nu$

$$\tan(\tilde{\theta}_L - \theta_L) = - \frac{y_i y_j \tan(2\theta_R)}{y_i^2 + y_j^2}$$

- Mass eigenvalues are :

$$m_{i,j} = \frac{v^2}{4M_0^2} \mu_0 \left| \left( y_i^2 - y_j^2 \right) \cos(2\theta_R) \pm \sqrt{4y_i^2 y_j^2 + \left( y_i^2 - y_j^2 \right)^2 \cos^2(2\theta_R)} \right|$$

$$m_{k \neq i,j} = \frac{v^2}{2M_0^2} \mu_0 y_k^2$$

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## Outline of numerical study:

- Varied in ranges:

$$M_0 \in [150 \text{ GeV}; 10 \text{ TeV}] \quad \mu_0 \in [0.1 \text{ keV}; 10^2 \text{ keV}] \quad \theta_R \in [0; 2\pi]$$



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**Not diagonal for  $t$  odd!**



**Predictions are independent of  $\theta_R$  !**

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**Diagonal for  $t$  even!**



**Predictions depend on  $\theta_R$  !**

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