Summary

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Salvador Centelles Chuliá

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Neutrino masses
Observational footprints
High scale seesaws
Low scale seesaws

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Neutrino masses

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Neutrino masses

- Neutrino oscillations entering the precision (<1%) era. Simplest explanation for the L/E profile are massive neutrinos.
- For now, the neutrino mass mechanism is a mistery:
 - Tree level, radiative, extra dimensional origin...?
 - The scale(s) of relevant NP
 - Neutrino nature: Dirac or Majorana
 - Lepton number conservation/violation
- Neutrino mixing with new states could be the window to NP

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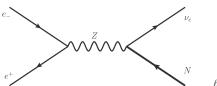
Observational footprints

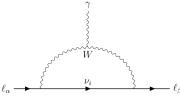
- Let us assume that neutrinos mix with some new heavy gauge singlet states
- Observational footprints are mainly
 - Collider direct production of m_{ν} mediators
 - cLFV (paradigmatically but not only $\mu^- \to e^- \gamma$)
 - Non-standard effects in neutrino propagation
- None of these observations would be a statement on neutrino. nature! See e.g. [1]

[1]Salvador Centelles Chuliá, Rahul Srivastava, and Avelino Vicente. "The inverse seesaw family: Dirac and Majorana". In: JHEP 03 (2021), p. 248. DOI: 10.1007/JHEP03(2021) 248. arXiv: 2011.06609到hep-p配. 夕久?

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• Typically suppressed by the 'active-heavy' mixing.

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Neutrino masses
Observational footprints

High scale seesaws

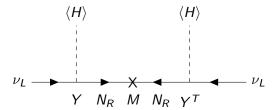
Low scale seesaws

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The type-I seesaw is a paradigmatic example of a high scale seesaw



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$M_{\nu} = \begin{pmatrix} 0 & Y \, \nu \\ Y^{\mathsf{T}} \, v & M \end{pmatrix}, \ U^{\mathsf{T}} M_{\nu} U = m_{\mathsf{d}}, \ U = \begin{pmatrix} \mathsf{N} & \mathsf{S} \\ \mathsf{X} & \mathsf{Y} \end{pmatrix}$

- *N* and *S* are the 'active-active' and 'active-sterile' mixings. By construction $N^{\dagger}N = 1 S^{\dagger}S$.
- In the seesaw expansion we loosely define the parameter $\varepsilon \sim O(Y \, v/M)$ and diagonalize perturbatively, finding

$$m_{\nu} = v^2 Y^T M^{-1} Y, \quad S^* = Y v M^{-1} V \sim \varepsilon, \quad \varepsilon^2 \sim O(m \nu / M)$$



Neutrino masses Observational footprints High scale seesaws

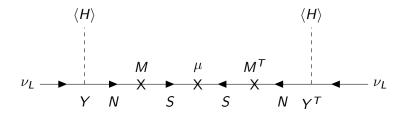
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 The inverse seesaw is a paradigmatic example of a low scale seesaw.



Inverse seesaw

• Introduces a new scale $\mu \ll \Lambda_{EW}$

$$M_{\nu} = \begin{pmatrix} 0 & Y v & 0 \\ Y^{T} v & \mu' & M \\ 0 & M^{T} & \mu \end{pmatrix}$$

In the seesaw expansion (one gen)

$$m_{\nu} = \frac{Y^2 v^2}{M^2} \mu, \quad S = \begin{pmatrix} \frac{Y v}{M^2} \mu & \frac{Y v}{M} \end{pmatrix} \sim \begin{pmatrix} \frac{m_{\nu}}{Y v} & \varepsilon \end{pmatrix}$$

ullet The second component can be % level, even if $m_
u o 0$

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Basic formulation
Neutrino scattering on a lepton target at zero distance

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Basic formulation

- 3 active neutrinos mix with *m* new (heavy) states
- The **unitary** mixing matrix is $(3 + m) \times (3 + m)$
- The upper 3 rows form a rectangular matrix K which characterizes the $\ell_{\alpha}-W-\nu_{i}$ interactions.

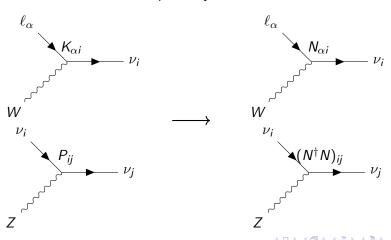
$$K = \begin{pmatrix} N & S \end{pmatrix}$$

$$K\,K^\dagger=1_{3 imes 3}$$

• The Z boson interaction is characterized by the $(3+m) \times (3+m)$ matrix $P = K^{\dagger} K \neq 1$



• If $E \ll M$ only the first 3×3 block of K and P can play a role, N and $N^{\dagger}N$, respectively.



Basic formulation

Important phenomenological consequences!

- CC can change flavour even at zero distance.
- NC is no longer diagonal.
- Observables at zero-distance depend on $(N^{\dagger}N)$. In the unitary limit this is the identity.
- Naive guess: Number of neutrino events in a given experiment is reduced compared to the unitary case. Not true!

- N plays a central role in this setup
- We parametrize N as [2]

$$N = \begin{pmatrix} \alpha_{11} & 0 & 0 \\ \alpha_{21} & \alpha_{22} & 0 \\ \alpha_{31} & \alpha_{31} & \alpha_{33} \end{pmatrix} \cdot U$$

- In the unitary limit, N = I and U is the unitary matrix responsible for the standard oscillations.
- Advantage of this parametrization: very clean theoretical interpretation in terms of mixing angles.
- The order of each parameter in the seesaw expansion is

$$\alpha_{ii}^2 \sim 1 - \varepsilon^2; \quad |\alpha_{ij}|^2 \sim \varepsilon^4$$

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^[2]F. J. Escrihuela et al. "On the description of nonunitary neutrino mixing". In: Phys. Rev. D 92.5 (2015). [Erratum: Phys.Rev.D 93, 119905 (2016)], p. 053009. DOI: 10.1103/PhysRevD.92.053009. arXiv: 1503.08879 [hep-ph].

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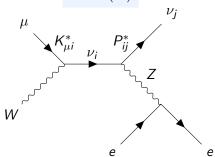
Neutrino scattering on a lepton target at zero distance

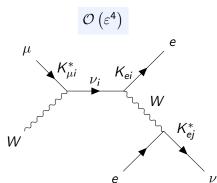
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 For concreteness, let's focus on the incoming muon neutrino case.

$$1-\mathcal{O}\left(arepsilon^2
ight)$$







Neutrino scattering on an electron target at zero distance

In the SM the cross section is given by

$$\left(\frac{d\sigma}{dT}\right)^{\rm SM} = \frac{2G_{\mu}^2 m_e}{\pi} \left(g_L^2 + g_R^2 \left(1 - \frac{T}{E\nu}\right)^2 - g_L g_R \frac{m_e T}{E_{\nu}^2}\right)$$

And in the presence of non-unitarity

$$\left(\frac{d\sigma}{dT}\right)^{NU} = \frac{\mathcal{P}_{\mu e}^{NC}}{(NN^{\dagger})_{ee}(NN^{\dagger})_{\mu\mu}} \left(\frac{d\sigma}{dT}\right)^{SM} + \frac{2m_{e}G_{\mu}^{2}}{\pi} \frac{\mathcal{R}e\left[\mathcal{P}_{\mu e}^{\text{int}}\right]}{(NN^{\dagger})_{ee}(NN^{\dagger})_{\mu\mu}} \left\{\frac{\mathcal{P}_{\mu e}^{CC}}{\mathcal{R}e\left[\mathcal{P}_{\mu e}^{\text{int}}\right]} + 2g_{L} - g_{R}\frac{m_{e}T}{E_{\nu}^{2}}\right\}$$

- This feature is not should by Covers (nurs NC)
- This feature is not shared by Cevens (pure NC) or inelastic scattering on nucleus (pure CC).
- However, this difference would be extremely hard to observe.
- It is also theoretically suppressed. Indeed, performing the the seesaw expansion and keeping only terms of $\mathcal{O}\left(\varepsilon\right)^2$ we find

$$\left(rac{d\sigma}{dT}
ight)^{\mathsf{NU}}pprox\left(2lpha_{\mathsf{22}}^{2}-lpha_{\mathsf{11}}^{2}
ight)\left(rac{d\sigma}{dT}
ight)^{\mathsf{SM}}+\mathcal{O}\left(arepsilon^{4}
ight)$$

Neutrino scattering on an electron target

- If $E_{
 u} >$ 10 GeV we also have the purely CC process $u_{lpha} + e^-
 ightarrow
 u_j + \mu^-$
- ullet For incoming μ neutrinos, the probability factor is given by

$$P_{lpha\mu}=(\mathsf{N}^\dagger\mathsf{N})_{\mathsf{ee}}(\mathsf{N}^\dagger\mathsf{N})_{lpha\mu}(\mathsf{N}^\dagger\mathsf{N})_{\mulpha}$$

$$P_{\mu\mu} \approx 2\alpha_{22}^2 + \alpha_{11}^2 - 2 \sim 1 - \mathcal{O}\left(\varepsilon^2\right)$$
 $\sigma \approx P_{\mu\mu} \frac{G_F^2}{\pi} \left(2E_{\nu}m_e - m_{\mu}^2\right)$



- $u_{\mu} + e
 ightarrow
 u_{j} + e$, mainly NC (at first order in seesaw expansion)
- $\nu_{\mu} + e \rightarrow \nu_{j} + \mu$, purely CC
- Final number of events could be bigger or smaller than the expected in the SM (due to the redefinition of G_F):

$$e^-$$
 events, NC: $\frac{\#}{\#_{SM}} \approx 2\alpha_{22}^2 - \alpha_{11}^2 \sim 1 \pm \mathcal{O}(\varepsilon^2)$ μ^- events, CC: $\frac{\#}{\#_{SM}} \approx \alpha_{22}^2 \sim 1 - \mathcal{O}(\varepsilon^2)$

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We will now analyze how the Near Detector (ND) of DUNE can constraint the NU parameters. However:

- Neutrino scattering with leptons has lower statistics compared to inelastic scattering with nucleus (way lower cross section)
- The experiment is not optimized to search for scattering on leptons
- Big flux uncertainties
- Constraints on NU from EW precision measurements will be stronger than those from neutrino physics constraints.

- It is generally a good idea to find and explore complementary probes of a given phenomena
- Background under control (it is a cleaner process)
- Relatively less analyzed process compared to nucleus scattering

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- DUNE (Deep Underground Neutrino Experiment) is an ambitious neutrino experiment under construction. Will determine the mass ordering and improve precision on θ_{23} , δ_{CP} and θ_{13} [3]
- Two beam modes (neutrino/antineutrino), mainly u_{μ} or $ar{
 u}_{\mu}$

[3]B. Abi et al. "Long-baseline neutrino oscillation physics potential of the DUNE experiment". In: Eur. Phys. J. C 80.10 (2020), p. 978. DOI: 10.1140/epjc/s10052-020-08456-zz. arXim 2006 36043 hep-ex

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- Mainly for cross-checking the neutrino flux, but we can use it to do BSM analysis too
- Will be the first purely leptonic test of "zero distance neutral oscillations"

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The experiment

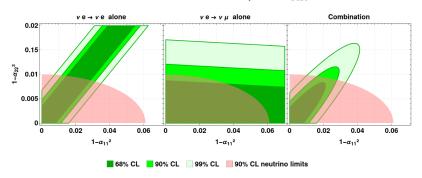
Analysis and sensitivity prospects

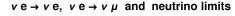
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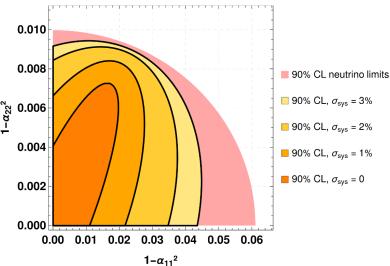


- We compute the expected number of events in the SM for each flavour component of the flux (3.5 years per mode).
- We compute the NU (global) factors at order ε^2 .
- We extract the sensitivity on NU parameters.
- We compare them with current neutrino limits.









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Summary

Take home ideas

- Low scale seesaws: A well motivated and broad class of models leading to rich phenomenology. Non-unitarity effects can be at the % level
- We have studied the effect of NU through the leptonic neutral current for the first time.
- The expected sensitivity will be competitive and complementary with other oscillation experiments (in particular on α_{11}^2 .

Thanks!

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- A unitary matrix of order n can be parametrized with $\frac{n(n-1)}{2}$ angles and n(n+1) phases.
- The complex rotation ω_{ij} has a mixing angle in the plane i-j and a phase.
- An (unphysical) diagonal matrix of phases times all the possible ω (in some order) parametrizes any unitary matrix

- We can choose the ordering to be (NS imes NS)(S imes NS)($\omega_{23}\omega_{13}\omega_{12}$)
- We identify the right hand side with the 'standard' mixing
- The first block cannot affect the 'active-active' 3 × 3 block
- The second block we subdivide in products of $\omega_{3j}\omega_{2j}\omega_{1j}$, which is lower tringular
- And a product of lower triangular matrices is also lower triangular
- diagonal entries are simply multiplications of cosines while off diagonal elements are proportional to sines (but are more complicated and include phases)



- As a simple example in the 3 + 1 scheme we get
- $\alpha_{ii} = c_{i4}$

 $\bullet \ \alpha_{ij} = s_{i4}s_{j4}e^{i(\phi_{i4}-\phi_{j4})}$



• $\nu_{\alpha} + N \rightarrow \nu_{i} + N$

- The COHERENT collaboration already detected the coherent scattering off nucleons in 2017 [4]
- Several experiments, including $Co\nu$ us in Heidelberg [5]
- For future sensitivity analysis see e.g. [6]

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^[4]D. Akimov et al. "Observation of Coherent Elastic Neutrino-Nucleus Scattering". In: Science 357.6356 (2017), pp. 1123–1126. DOI: 10.1126/science.aao0990. arXiv: 1708.01294 [nucl-ex].

^[5]H. Bonet et al. "Novel constraints on neutrino physics beyond the standard model from the CONUS experiment". In: JHEP 05 (2022), p. 085. DOI: 10.1007/JHEP05(2022)085. arXiv: 2110.02174 [hep-ph].

^[6]O. G. Miranda et al. "Future CEVNS experiments as probes of lepton unitarity and light-sterile neutrinos". In: Phys. Rev. D 102 (2020), p. 113014. DOI: 10.1103/PhysRevD.102.113014. arXiv: 2008.0275 [hep-ph].

Neutrino limits

Introduction

- See the analysis in[7]
- Combines data from long (NOvA, T2K, MINOS) and short baseline (NOMAD, NuTeV) experiments
- At 90% CL:

$$1 - \alpha_{11}^2 \le 6 \times 10^{-2}$$
$$1 - \alpha_{22}^2 \le 1 \times 10^{-2}.$$

[7]D. V. Forero et al. "Nonunitary neutrino mixing in short and long-baseline experiments". In: Phys. Rev. D 104.7 (2021), p. 075030. DOI: 10.1103/PhysRevD.104.075030. arXiv: 2163 01998 [hep-ph].

\mathcal{N}_{U}	ν mode		$\bar{\nu}$ mode		$\mathcal{N}_{ m NU}/\mathcal{N}_{ m U}$	Seesaw ord	er Main contribution
$\nu_a + e^- \rightarrow \nu_j + e^-$	events	σ	events	σ	$\mathcal{P} G_F^2/G_\mu^2$		
$ u_e$	2.800	80	1.530	50	$2\alpha_{11}^2 - \alpha_{21}^2$	$1 \pm \mathcal{O}\left(\varepsilon^2\right)$	NC + CC
$ u_{\mu}$	31.400	700	5.800	100	$2\alpha_{22}^2 - \alpha_{12}^2$	$1 \pm \mathcal{O}\left(\varepsilon^2\right)$) NC
$ar{ u}_e$	430	20	780	30	$2\alpha_{11}^2 - \alpha_{21}^2$	$1 \pm \mathcal{O}\left(\varepsilon^2\right)$	NC + CC
$ar{ u}_{\mu}$	3.200	80	20.000	400	$2\alpha_{22}^2 - \alpha_{12}^2$	$1 \pm \mathcal{O}\left(\varepsilon^2\right)$) NC
total	37.800	800	28.000	600			
\mathcal{N}_{U}	ν mode		$\bar{\nu} \text{ mode}$		$\mathcal{N}_{ m NU}/\mathcal{N}_{ m U}$	Seesaw order	Main contribution
$\nu_a + e^- \rightarrow \nu_j + \mu^-$	events	σ	events	σ	$\mathcal{P} G_F^2/G_\mu^2$		
$ u_e$	0	0	0	0	$ \alpha_{21} ^2$	$\mathcal{O}\left(arepsilon^4 ight)$	$\mathcal{O}\left(arepsilon^4 ight)$
$ u_{\mu}$	17.900	400	14.200	300	α_{22}^2	$1-\mathcal{O}\left(\varepsilon^2\right)$	CC
$ar{ u}_e$	380	20	230	20	α_{11}^2	$1 - \mathcal{O}\left(\varepsilon^2\right)$	CC
$ar{ u}_{\mu}$	0	0	0	0	$ \alpha_{21} ^{2}$	$\mathcal{O}\left(arepsilon^4 ight)$	$\mathcal{O}\left(arepsilon^4 ight)$
total	18.300	400	14.400	300			



Oscillations

Introduction

- Long baseline experiments
- Short baseline (zero distance)
- Lepton flavour universality
 - π & K decays into μ^- and e^-
 - τ^- decays (hadrons or leptons) \leftarrow function of (α_{33})
 - ullet eta decays and CKM unitarity



- EW precision observables
 - W mass, s_W, Γ_z...
 - CDF-II W mass [8]

[8]Mattias Blennow et al. "Right-handed neutrinos and the CDF II anomaly". In: Phys. Rev. D 106.7 (2022), p. 073005. DOI: 10.1103/PhysRevD.106.073005. arXiv: 2204.04559 [hep∰h]. ∢ 🛢 ▶

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- See [9] for a nice update
- We can translate the decay rates of pions (or Kaons) to electrons and muons into couplings with the W. The experimental result

$$\left(\frac{g_e}{g_\mu}\right)^2 = 0.998 \pm 0.002$$

In the SM this ratio is 1. In the presence of non-unitarity

$$\left(\frac{g_e}{g_\mu}\right)^2 = 1 + \alpha_{11}^2 - \alpha_{22}^2$$

^[9]Douglas Bryman et al. "Testing Lepton Flavor Universality with Pion, Kaon, Tau, and Beta Decays". In: Ann. Rev. Nucl. Part. Sci. 72 (2022), pp. 69–91. DOI: 10.1146/annurev-nucl-110121-051223. arXiv: 2111.05338 [hep-ph].

• We can also compare the effective coupling of β and μ decays. In the SM

$$\left(rac{G_eta}{G_\mu}
ight)^2 = \sum_i |V_{ui}|^2 = 1$$

Different measurements of nuclear processes give

$$\sum_{i} |V_{ui}|^2 = 1 - (19.5 \pm 5.3) \times 10^{-4}$$

 Known as the Cabibbo anomaly. This anomaly only gets worse in the presence of (leptonic) non-unitarity

$$\left(\frac{G_{\beta}}{G_{\mu}}\right)^2 = 2 - \alpha_{22}^2 > 1$$



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 Including all the LFU and EW precision measurements gives a much stronger constraints than the ones obtained from DUNE-PRISM or oscillations [10]

95% CL:

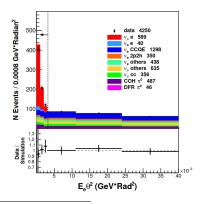
$$1 - \alpha_{11}^2 \le 2 \times 10^{-3}$$

 $1 - \alpha_{22}^2 \le 2 \times 10^{-4}$

 Caveat: The constraints are pushed towards zero due to the Cabibbo anomaly. Would be interested in seeing a similar analysis excluding the CKM unitarity test.

^{[10] (}attias Blennow et al. "Bounds on lepton non-unitarity and heavy neutrino mixing". In: JHEP 08 (2023), p. 030. DOI: 10.1007/JHEP08(2023)030. arXiv: 2306.01040 [hep-ph]. 4 () +

We can take advantage of the fact that the electron scattering will be mainly forward. See for example[11]



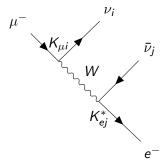
^{[11].} Zazueta et al. "Improved constraint on the MINER ν A medium energy neutrino flux using ν^- = $\rightarrow \nu^-$ e - data". In: Phys. Rev. D 107.1 (2023), p. 012001. DOI: 10.1103/PhysRevb.107.012001. arXiv: 2209.05540 [hep-ex].

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Redefinition of GF

- G_F is measured through the μ^- decay.
- Lepton non-unitarity modifies this decay. The effective muon decay G_{μ} is related to the 'real' G_F by



$$G_\mu^2 = (\mathit{N}^\dagger \mathit{N})_{\mu\mu} (\mathit{N}^\dagger \mathit{N})_{\mathsf{ee}} \, G_F^2
ightarrow rac{G_F^2}{G_\mu^2} = rac{1}{(\mathit{N}^\dagger \mathit{N})_{\mu\mu} (\mathit{N}^\dagger \mathit{N})_{\mathsf{ee}}} \geq 1$$