

# Disentangling new physics in rare meson decays

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Based on a work with  
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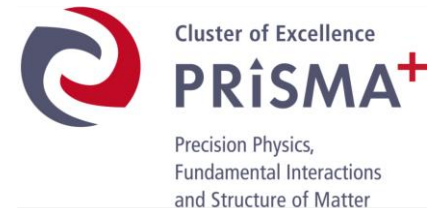


SFB 1258

Neutrinos  
Dark Matter  
Messengers



Martin A. Mojahed – Planck 2024



# Introduction and Motivation

The rare processes  $K \rightarrow \pi \nu \bar{\nu}$  and  $B \rightarrow K(K^*) \nu \bar{\nu}$  belong to theoretically cleanest (FCNC) processes.

Still room between SM prediction and experimental measurements (experimental bounds):

$$\mathcal{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu})_{\text{exp}} = (10.6_{-3.5}^{+4.0} \pm 0.9) \times 10^{-11},$$

$$\mathcal{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu})_{\text{SM}} = (8.60 \pm 0.42) \times 10^{-11},$$

$$\mathcal{B}(B^+ \rightarrow K^+ \nu \bar{\nu})_{\text{exp}} = (13 \pm 4) \times 10^{-6},$$

$$\mathcal{B}(B^+ \rightarrow K^+ \nu \bar{\nu})_{\text{SM}} = (4.92 \pm 0.30) \times 10^{-6},$$

Good place to **search for new physics**, for example hints of lepton-number violation (LNV).

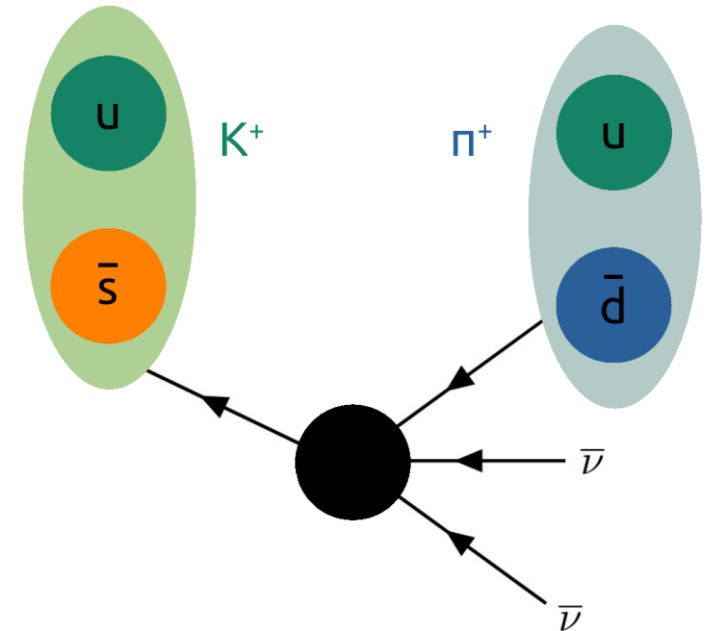
# Overview I: Formalism

Parametrize new physics (NP) with effective operators:

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \sum_{X=L,R} C_{\nu d}^{\text{VLX}} \mathcal{O}_{\nu d}^{\text{VLX}} + \left( \sum_{X=L,R} C_{\nu d}^{\text{SLX}} \mathcal{O}_{\nu d}^{\text{SLX}} + C_{\nu d}^{\text{TLL}} \mathcal{O}_{\nu d}^{\text{TLL}} + \text{h.c.} \right),$$

Unlike charged semileptonic decay, we cannot determine the nature of the final state neutrinos ( $\nu\nu$ ,  $\nu\bar{\nu}$ ,  $\bar{\nu}\bar{\nu}$ ).

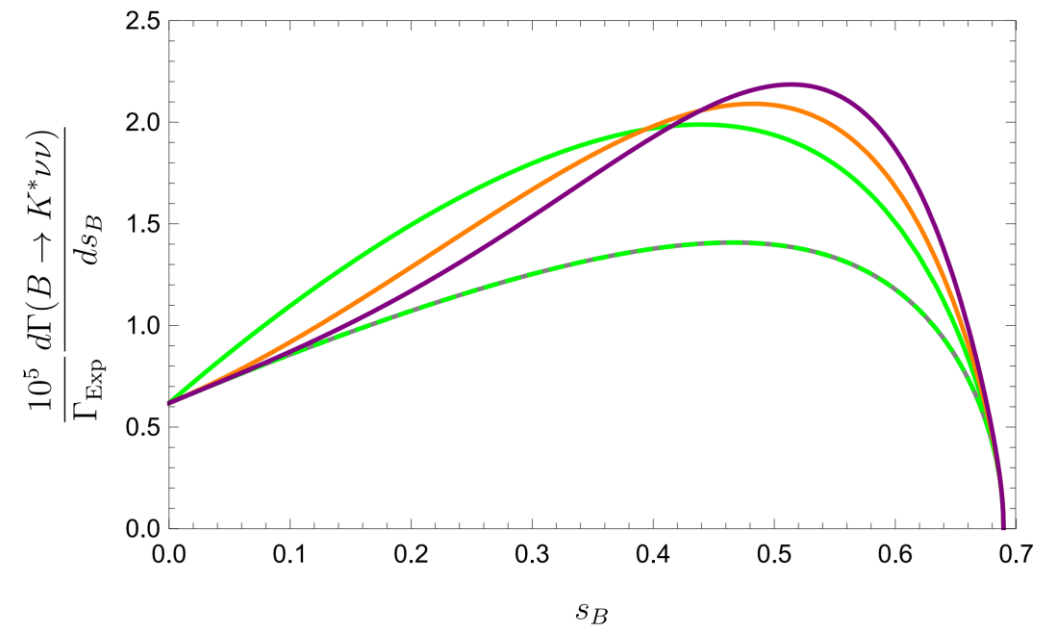
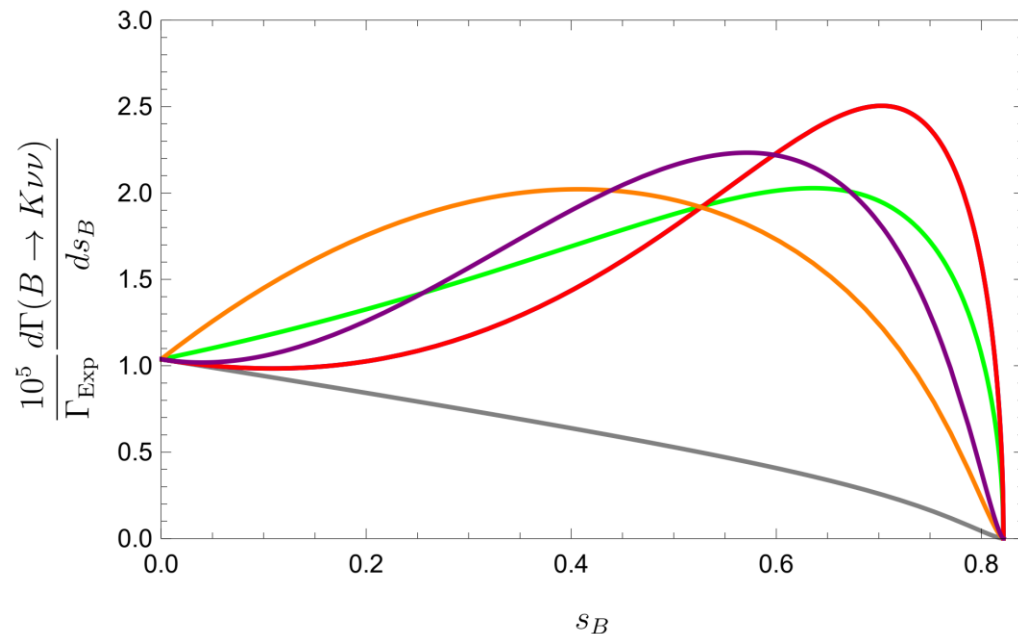
Measure missing invariant mass  $s$ , which leads to kinematic distributions  $\frac{d\Gamma}{ds}$ .



# Overview II: The main idea in a nutshell

Observation: Different operators give rise to different kinematic distributions.

Idea: Analyze kinematic distribution to disentangle contributing operators.



# LEFT setup:

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \sum_{X=L,R} C_{\nu d}^{\text{VLX}} \mathcal{O}_{\nu d}^{\text{VLX}} + \left( \sum_{X=L,R} C_{\nu d}^{\text{SLX}} \mathcal{O}_{\nu d}^{\text{SLX}} + C_{\nu d}^{\text{TLL}} \mathcal{O}_{\nu d}^{\text{TLL}} + \text{h.c.} \right),$$

$$\mathcal{O}_{\nu d}^{\text{VLL}} = (\bar{\nu}_L \gamma^\mu \nu_L) (\bar{d}_L \gamma_\mu d_L),$$

$$\mathcal{O}_{\nu d}^{\text{VLR}} = (\bar{\nu}_L \gamma^\mu \nu_L) (\bar{d}_R \gamma_\mu d_R),$$

$$\mathcal{O}_{\nu d}^{\text{SLL}} = (\bar{\nu}_L^c \nu_L) (\bar{d}_R d_L),$$

$$\mathcal{O}_{\nu d}^{\text{SLR}} = (\bar{\nu}_L^c \nu_L) (\bar{d}_L d_R),$$

$$\mathcal{O}_{\nu d}^{\text{TLL}} = (\bar{\nu}_L^c \sigma_{\mu\nu} \nu_L) (\bar{d}_R \sigma^{\mu\nu} d_L).$$

↑  
 LNV

# Example

$$\mathcal{D}_{BK}^{\text{exp}}(s) \equiv \frac{d\Gamma(B \rightarrow K\nu\hat{\nu})}{ds}, \quad \mathcal{D}_{BK}^{\text{exp}}(s) = C_S^{BK} f_S^{BK}(s) + C_T^{BK} f_T^{BK}(s) + C_V^{BK} f_V^{BK}(s),$$

$$C_S^{BK} = \sum_{\alpha \leq \beta} \left(1 - \frac{1}{2}\delta_{\alpha\beta}\right) \left( |C_{\nu d, \alpha\beta bs}^{\text{SLL}} + C_{\nu d, \alpha\beta bs}^{\text{SLR}}|^2 + |C_{\nu d, \alpha\beta sb}^{\text{SLL}} + C_{\nu d, \alpha\beta sb}^{\text{SLR}}|^2 \right),$$

$$C_T^{BK} = \sum_{\alpha < \beta} \left( |C_{\nu d, \alpha\beta bs}^{\text{TLL}}|^2 + |C_{\nu d, \alpha\beta sb}^{\text{TLL}}|^2 \right),$$

$$C_V^{BK} = \sum_{\alpha, \beta} \left(1 - \frac{1}{2}\delta_{\alpha\beta}\right) |C_{\nu d, \alpha\beta sb}^{\text{VLL}} + C_{\nu d, \alpha\beta sb}^{\text{VLR}}|^2,$$

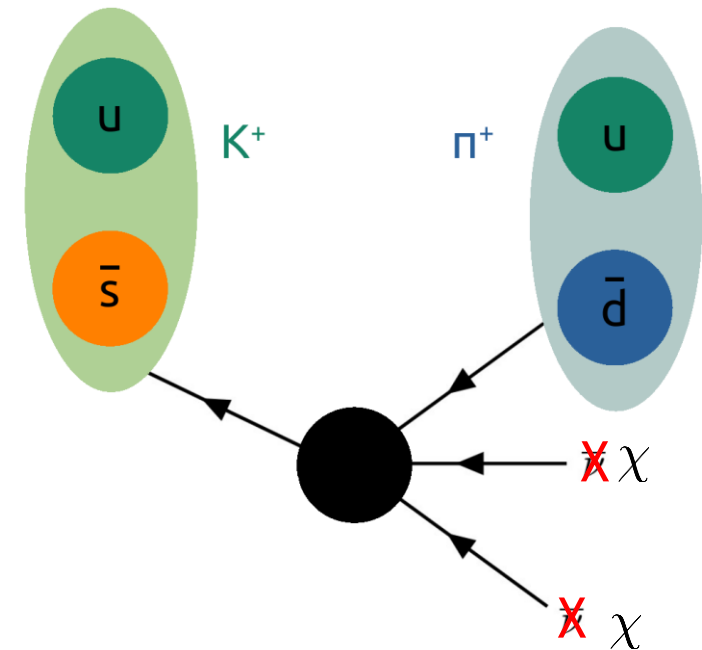
$$C_S^{BK} = \frac{D_{BK}^{\text{exp}}(s_1) [f_V^{BK}(s_2)f_T^{BK}(s_3) - f_V^{BK}(s_3)f_T^{BK}(s_2)] + \text{cyclic}}{f_S^{BK}(s_1) [f_V^{BK}(s_2)f_T^{BK}(s_3) - f_V^{BK}(s_3)f_T^{BK}(s_2)] + \text{cyclic}},$$

$$C_V^{BK} = \frac{D_{BK}^{\text{exp}}(s_1) [f_S^{BK}(s_2)f_T^{BK}(s_3) - f_S^{BK}(s_3)f_T^{BK}(s_2)] + \text{cyclic}}{f_V^{BK}(s_1) [f_S^{BK}(s_2)f_T^{BK}(s_3) - f_S^{BK}(s_3)f_T^{BK}(s_2)] + \text{cyclic}},$$

$$C_T^{BK} = \frac{D_{BK}^{\text{exp}}(s_1) [f_S^{BK}(s_2)f_V^{BK}(s_3) - f_S^{BK}(s_3)f_V^{BK}(s_2)] + \text{cyclic}}{f_T^{BK}(s_1) [f_S^{BK}(s_2)f_V^{BK}(s_3) - f_S^{BK}(s_3)f_V^{BK}(s_2)] + \text{cyclic}}.$$

# A Loophole

The final state neutrinos are not detected:  
What if the invisible part of the final state is  
not a pair of neutrinos?

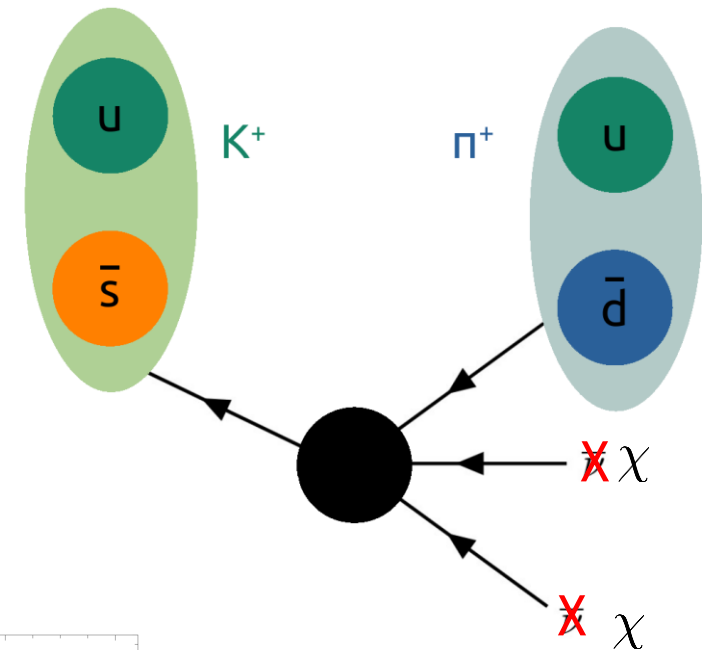
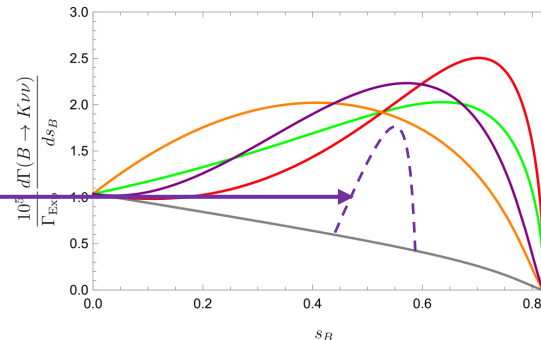


# Dark final state particles

Since neutrinos not detected, the invisible final state could e.g. be two dark sector particles  $\chi\chi$ .

We have taken into account the possibility of having two final state dark particles with spin 0, spin  $\frac{1}{2}$  and spin 1.

Note: A single invisible boson would lead to kinematic distributions with a single bump.





# Defining “kinematic structure”

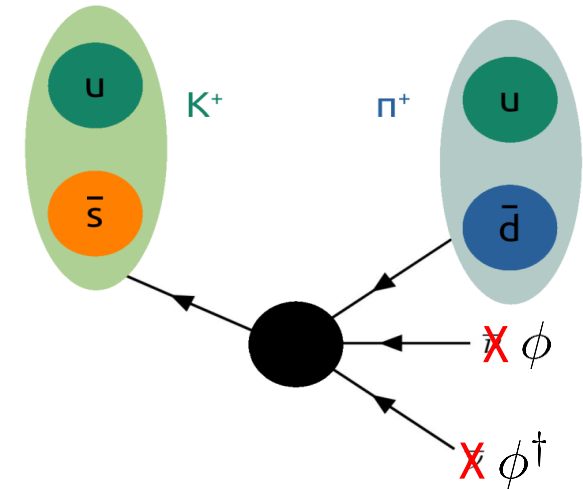
Operator:  $\mathcal{O}_{ab\phi}^S = (\bar{q}_a q_b)(\phi^\dagger \phi)$

Contribution to decay rate:  $\frac{d\Gamma(K^+ \rightarrow \pi^+ + \text{inv})}{ds} = \frac{B_+^2}{256\pi^3 m_{K^+}^3} \lambda^{1/2}(m_{K^+}^2, m_{\pi^+}^2, s) |f_0^{K^+}(s)|^2 |C_{ds\phi}^S|^2$

No dependence on  $s$ .

Kinematic structure.

Wilson coefficient.



# Going beyond LEFT: Dark LEFT

| Type         | Operator   | Kinematic structure in $K \rightarrow \pi + \cancel{E}$ |
|--------------|--|---|
| (SM)LEFT     | $\mathcal{O}_{\nu d}^{\text{SLL}}, \mathcal{O}_{\nu d}^{\text{SLR}}$ | $s \lambda^{1/2}  f_0^K ^2$                             |
|              | $\mathcal{O}_{\nu d}^{\text{VLL}}, \mathcal{O}_{\nu d}^{\text{VLR}}$ | $\lambda^{3/2}  f_+^K ^2$                               |
|              | $\mathcal{O}_{\nu d}^{\text{TLL}}$                                   | $s \lambda^{3/2}  f_T^K ^2$                             |
| Scalar DM    | $\mathcal{O}_{sd\phi}^S$   | $\lambda^{1/2}  f_0^K ^2$                               |
|              | $\mathcal{O}_{sd\phi}^V$   | $\lambda^{3/2}  f_+^K ^2$                               |
| Fermion DM   | $\mathcal{O}_{sd\chi_1}^S, \mathcal{O}_{sd\chi_2}^S$                 | $s \lambda^{1/2}  f_0^K ^2$                             |
|              | $\mathcal{O}_{sd\chi_1}^V, \mathcal{O}_{sd\chi_2}^V$                 | $\lambda^{3/2}  f_+^K ^2$                               |
|              | $\mathcal{O}_{sd\chi_1}^T, \mathcal{O}_{sd\chi_2}^T$                 | $s \lambda^{3/2}  f_T^K ^2$                             |
| Vector DM: A | $\mathcal{O}_{sdA}^S$  | $s^2 \lambda^{1/2}  f_0^K ^2$                           |
|              | $\mathcal{O}_{sdA2}^V$   | $s^2 \lambda^{1/2}  f_0^K ^2$                           |
|              | $\mathcal{O}_{sdA3}^V, \mathcal{O}_{sdA6}^V$                         | $s \lambda^{3/2}  f_+^K ^2$                             |
|              | $\mathcal{O}_{sdA4}^V, \mathcal{O}_{sdA5}^V$                         | $s^2 \lambda^{3/2}  f_+^K ^2$                           |
|              | $\mathcal{O}_{sdA1}^T$   | $s^2 \lambda^{3/2}  f_T^K ^2$                           |
| Vector DM: B | $\mathcal{O}_{sdB1}^S, \mathcal{O}_{sdB2}^S$                         | $s^2 \lambda^{1/2}  f_0^K ^2$                           |
|              | $\mathcal{O}_{sdB1}^T, \mathcal{O}_{sdB2}^T$                         | $s^2 \lambda^{3/2}  f_T^K ^2$                           |

| Type   | Operator   | Kinematic structure in $B \rightarrow K^* + \cancel{E}$  |
|--|--|--|
| (SM)LEFT                                     | $\mathcal{O}_{\nu d}^{\text{SLL}}, \mathcal{O}_{\nu d}^{\text{SLR}}$                       | $s \lambda^{3/2}  A_0 ^2$  |
|  | $\mathcal{O}_{\nu d}^{\text{VLL}}, \mathcal{O}_{\nu d}^{\text{VLR}}$                       | $\lambda^{1/2} \left[ \frac{s\lambda}{(m_B+m_{K^*})^2}  V_0 ^2 + s(m_B+m_{K^*})^2  A_1 ^2 + 32m_B^2 m_{K^*}^2  A_{12} ^2 \right]$        |
|  | $\mathcal{O}_{\nu d}^{\text{TLL}}$   | $\lambda^{1/2} \left[ \lambda  T_1 ^2 + (m_B^2 - m_{K^*}^2)^2  T_2 ^2 + \frac{8m_B^2 m_{K^*}^2 s}{(m_B+m_{K^*})^2}  T_{23} ^2 \right]$   |
| Scalar DM                                    | $\mathcal{O}_{sb\phi}^P$   | $\lambda^{3/2}  A_0 ^2$  |
|  | $\mathcal{O}_{sb\phi}^A$   | $\lambda^{1/2} \left[ s(m_B+m_{K^*})^2  A_1 ^2 + 32m_B^2 m_{K^*}^2  A_{12} ^2 \right]$   |
|  | $\mathcal{O}_{sb\phi}^V$   | $s \lambda^{3/2}  V_0 ^2$  |
| Fermion DM                                   | $\mathcal{O}_{sb\chi_1}^P, \mathcal{O}_{sb\chi_2}^P$                                       | $s \lambda^{3/2}  A_0 ^2$  |
|  | $\mathcal{O}_{sb\chi_1}^A, \mathcal{O}_{sb\chi_2}^A$                                       | $\lambda^{1/2} \left[ s(m_B+m_{K^*})^2  A_1 ^2 + 32m_B^2 m_{K^*}^2  A_{12} ^2 \right]$   |
|  | $\mathcal{O}_{sb\chi_1}^V, \mathcal{O}_{sb\chi_2}^V$                                       | $s \lambda^{3/2}  V_0 ^2$  |
|  | $\mathcal{O}_{sb\chi_1}^T, \mathcal{O}_{sb\chi_2}^T$                                       | $\lambda^{1/2} \left[ \lambda  T_1 ^2 + (m_B^2 - m_{K^*}^2)^2  T_2 ^2 + \frac{8m_B^2 m_{K^*}^2 s}{(m_B+m_{K^*})^2}  T_{23} ^2 \right]$   |
| Vector DM: A                                 | $\mathcal{O}_{sbA}^P$  | $s^2 \lambda^{3/2}  A_0 ^2$  |
|  | $\mathcal{O}_{sbA1}^T$   | $s \lambda^{3/2}  T_1 ^2$  |
|  | $\mathcal{O}_{sbA2}^T$   | $s \lambda^{1/2} \left[ (m_B^2 - m_{K^*}^2)^2  T_2 ^2 + \frac{8m_B^2 m_{K^*}^2 s}{(m_B+m_{K^*})^2}  T_{23} ^2 \right]$                   |
|  | $\mathcal{O}_{sbA3}^V, \mathcal{O}_{sbA6}^V$   | $s^2 \lambda^{3/2}  V_0 ^2$  |
|  | $\mathcal{O}_{sbA4}^V, \mathcal{O}_{sbA5}^V$   | $s^3 \lambda^{3/2}  V_0 ^2$  |
|  | $\mathcal{O}_{sbA2}^A$   | $s^2 \lambda^{3/2}  A_0 ^2$  |
|  | $\mathcal{O}_{sbA3}^A, \mathcal{O}_{sbA6}^A$   | $s \lambda^{1/2} \left[ s(m_B+m_{K^*})^2  A_1 ^2 + 32m_B^2 m_{K^*}^2  A_{12} ^2 \right]$   |
| $\mathcal{O}_{sbA4}^A, \mathcal{O}_{sbA5}^A$ | $s^2 \lambda^{1/2} \left[ s(m_B+m_{K^*})^2  A_1 ^2 + 32m_B^2 m_{K^*}^2  A_{12} ^2 \right]$ |  |
| Vector DM: B                                 | $\mathcal{O}_{sbB1}^P, \mathcal{O}_{sbB2}^P$   | $s^2 \lambda^{3/2}  A_0 ^2$  |
|  | $\mathcal{O}_{sbB1}^T, \mathcal{O}_{sbB2}^T$   | $s \lambda^{1/2} \left[ \lambda  T_1 ^2 + (m_B^2 - m_{K^*}^2)^2  T_2 ^2 + \frac{8m_B^2 m_{K^*}^2 s}{(m_B+m_{K^*})^2}  T_{23} ^2 \right]$ |

# Implications of an observation of LNV in rare meson decays

What we could learn from observation of NP in rare meson decays.

Question:

Imagine one would find that a lepton-number violating (LNV) operator is contributing to rare meson decays.

What would be the ramifications for

- 1) UV physics?
- 2) Leptogenesis (LG)?

# LVN in rare meson decays

As an example, we consider the following SMEFT operator:

$$\mathcal{O}_{\bar{d}LQLH1} = \epsilon_{ij}\epsilon_{mn} (\bar{d}L^i) (\overline{Q^{Cj}} L^m) H^n.$$

This operator generates **scalar** and **tensor currents in LEFT**.

Lower bound on the associated new physics (NP) scale from Belle II:  $\Lambda_{\text{NP}} \approx 3.0 \text{ TeV}$ .

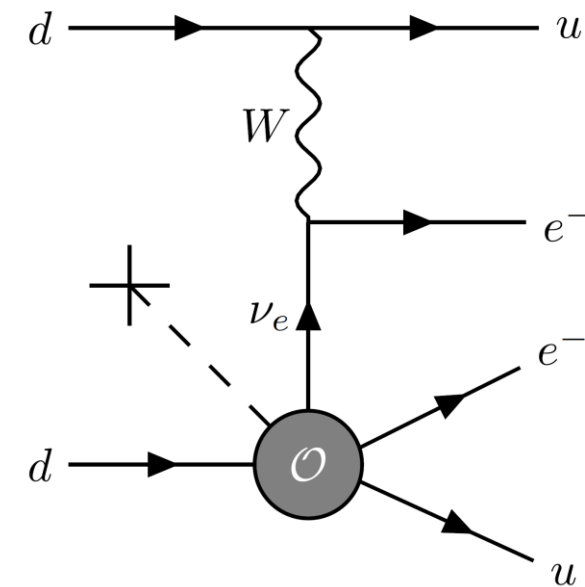
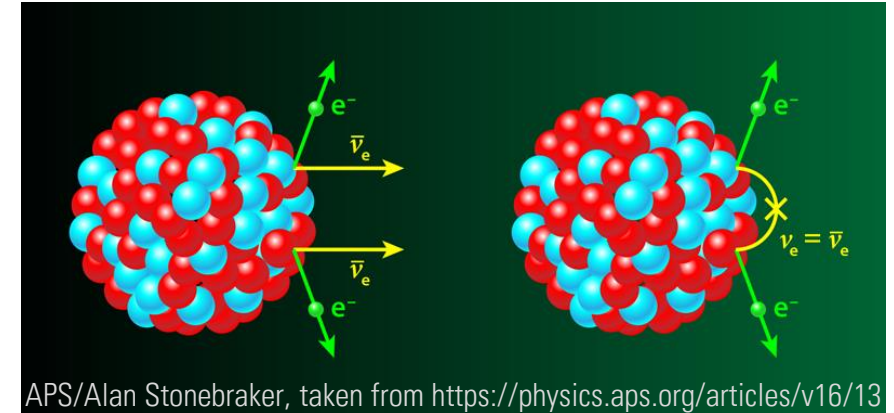
# 1) UV physics (I)

LN<sub>V</sub> operators are tightly constrained by non-observation of neutrinoless double beta decay.

Radiative corrections to neutrinoless double beta decay

leads to the following lower bound on the NP scale:

$$\Lambda_{\text{NP}} \approx 242 \text{ TeV}.$$



# 1) UV physics (II)

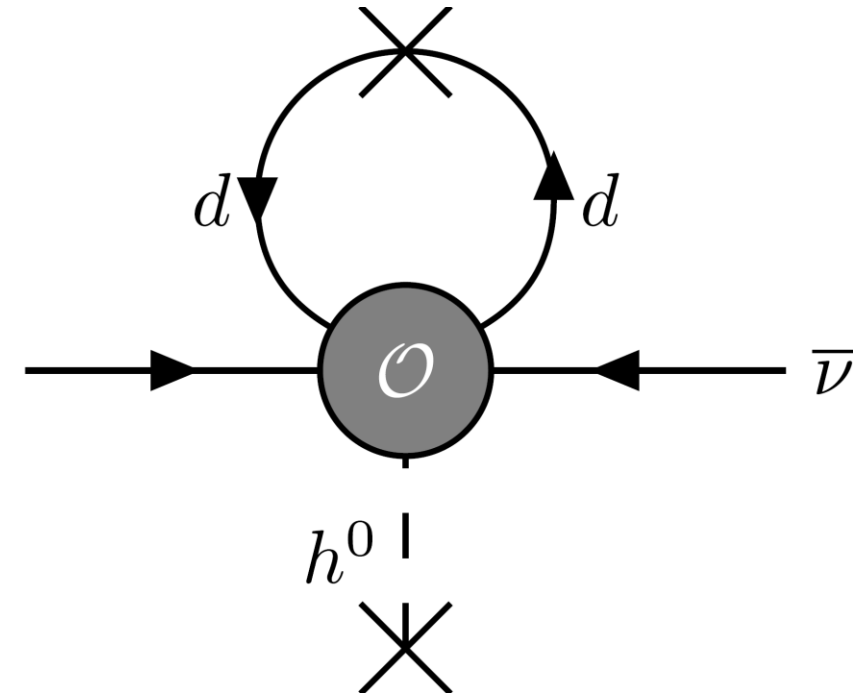
Higher-dimensional LNV operators give **radiative contributions to a Majorana neutrino mass,**

$$\delta m_\nu \approx \frac{y_d v^2}{16\pi^2 \Lambda_{\text{NP}}}.$$

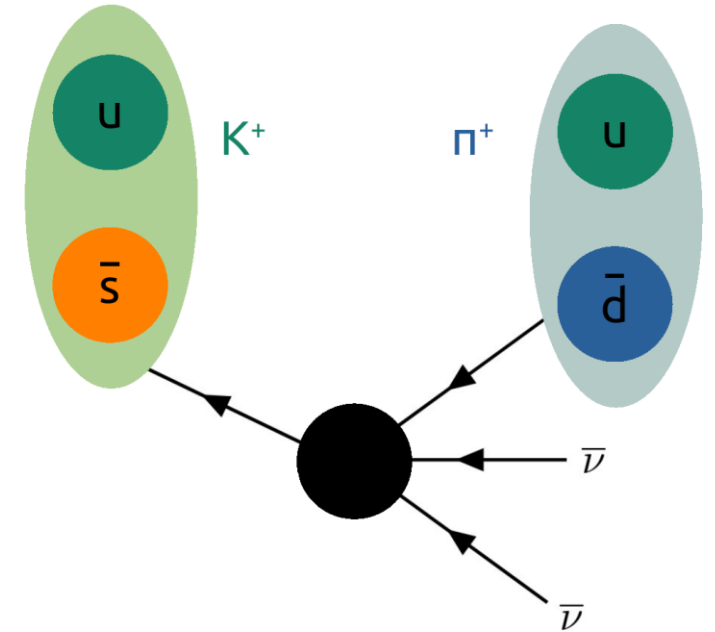
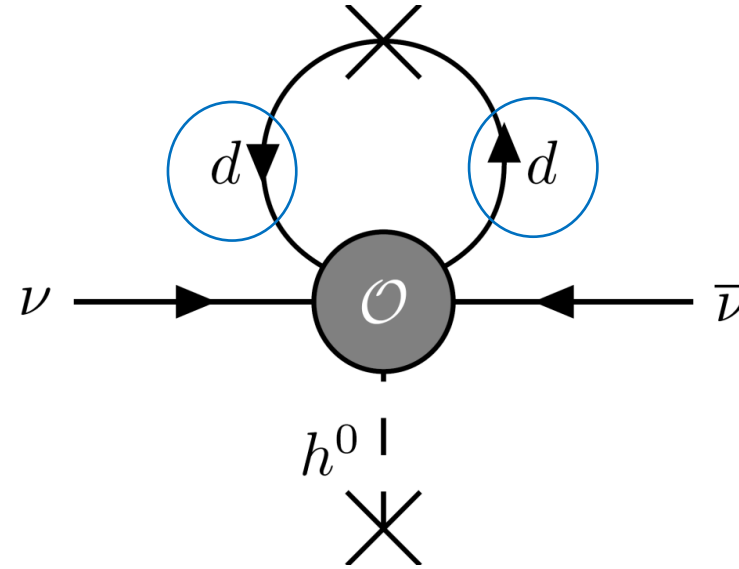
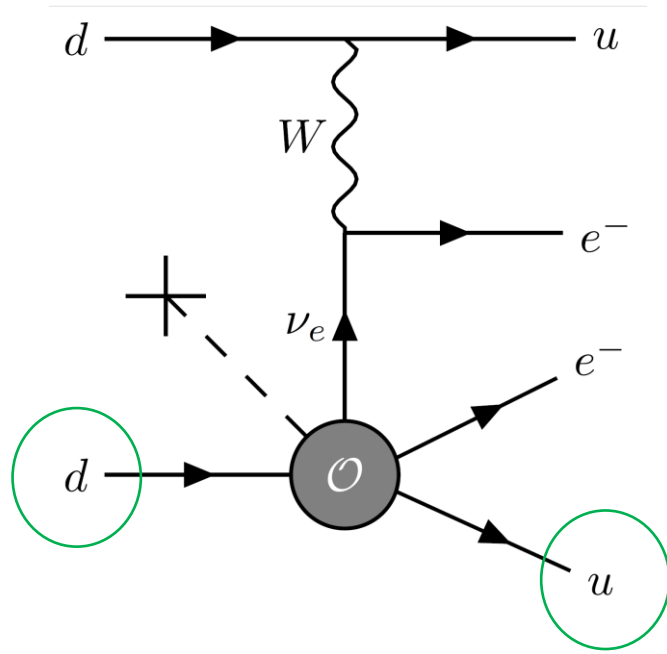
Requiring this contribution to not exceed 0.1 eV yields a lower limit on the NP scale

$$\Lambda_{\text{NP}} \approx 5 \cdot 10^4 \text{ TeV} \quad (10^9 \text{ GeV})$$

for the first (second) generation down-type quark Yukawa coupling.



# 1) UV physics (III)



Conclusion: Flavor non-democratic UV physics.

## 2) Leptogenesis

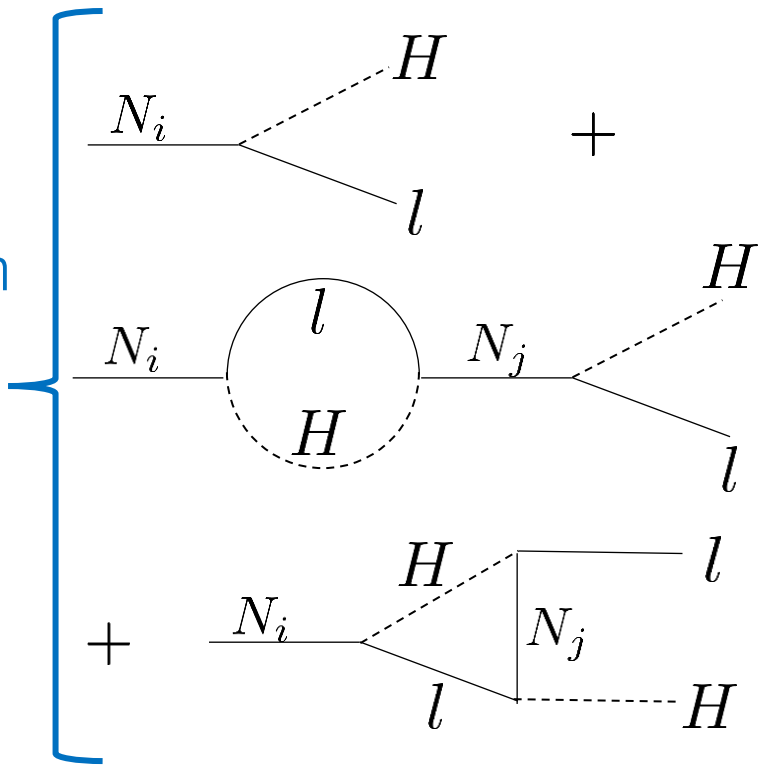
Leptogenesis denotes a class of scenarios for baryogenesis where a lepton asymmetry is generated via CP-violating decays of right-handed neutrinos.

Sphalerons: Lepton asymmetry  $\rightarrow$  baryon asymmetry.

LNV operator: Contributes to diminishing the generated lepton asymmetry.

Highly efficient in the low TEV range down to the electroweak scale.

Observation of LNV in rare meson decays  $\rightarrow$  High-scale LG under tension!

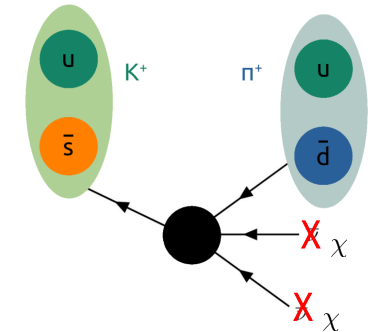
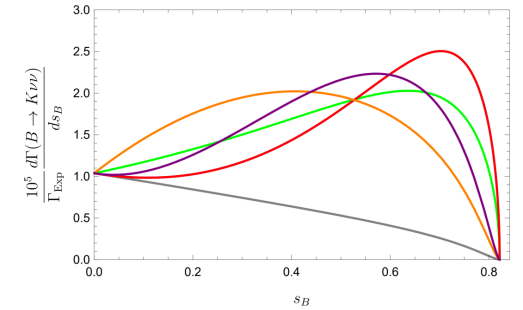




# Summary

- We considered strategies to disentangle NP in the processes  $K \rightarrow \pi \nu \bar{\nu}$  and  $B \rightarrow K(K^*) \nu \bar{\nu}$  using kinematic distributions (and beyond).
- It is very often possible to disentangle LEFT operators from two dark scalars or two dark vectors by analyzing kinematic distributions.
- We discussed implications of an observation of LNV NP in rare meson decays.

Take home message: By analyzing kinematic distributions of the rare processes  $K \rightarrow \pi \nu \bar{\nu}$  and  $B \rightarrow K(K^*) \nu \bar{\nu}$ : Disentangle effective operator origin of NP  $\rightarrow$  Search for e.g. LNV.



| Type         | Operator   | Kinematic structure in $K \rightarrow \pi + \mathcal{E}$ |
|--------------|--|--|
| (SM)LEFT     | $\mathcal{O}_{sd}^{\text{SLL}}, \mathcal{O}_{sd}^{\text{SLR}}$ | $s \lambda^{1/2}  f_0^K ^2$                              |
|              | $\mathcal{O}_{sd}^{\text{VLL}}, \mathcal{O}_{sd}^{\text{VLR}}$ | $\lambda^{3/2}  f_+^K ^2$                                |
|              | $\mathcal{O}_{sd}^{\text{TLL}}$                                | $s \lambda^{3/2}  f_T^K ^2$                              |
| Scalar DM    | $\mathcal{O}_{sd0}^S$  | $\lambda^{1/2}  f_0^K ^2$                                |
|              | $\mathcal{O}_{sd0}^V$  | $\lambda^{3/2}  f_+^K ^2$                                |
| Fermion DM   | $\mathcal{O}_{sd\chi 1}^S, \mathcal{O}_{sd\chi 2}^S$           | $s \lambda^{1/2}  f_0^K ^2$                              |
|              | $\mathcal{O}_{sd\chi 1}^V, \mathcal{O}_{sd\chi 2}^V$           | $\lambda^{3/2}  f_+^K ^2$                                |
|              | $\mathcal{O}_{sd\chi 1}^T, \mathcal{O}_{sd\chi 2}^T$           | $s \lambda^{3/2}  f_T^K ^2$                              |
| Vector DM: A | $\mathcal{O}_{sdA}^S$  | $s^2 \lambda^{1/2}  f_0^K ^2$                            |
|              | $\mathcal{O}_{sdA 2}^V$  | $s^2 \lambda^{1/2}  f_0^K ^2$                            |
|              | $\mathcal{O}_{sdA 3}^V, \mathcal{O}_{sdA 6}^V$                 | $s \lambda^{3/2}  f_+^K ^2$                              |
|              | $\mathcal{O}_{sdA 4}^V, \mathcal{O}_{sdA 5}^V$                 | $s^2 \lambda^{3/2}  f_+^K ^2$                            |
|              | $\mathcal{O}_{sdA 1}^T$  | $s^2 \lambda^{3/2}  f_T^K ^2$                            |
| Vector DM: B | $\mathcal{O}_{sdB 1}^S, \mathcal{O}_{sdB 2}^S$                 | $s^2 \lambda^{1/2}  f_0^K ^2$                            |
|              | $\mathcal{O}_{sdB 1}^T, \mathcal{O}_{sdB 2}^T$                 | $s^2 \lambda^{3/2}  f_T^K ^2$                            |



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Neutrinos  
Dark Matter  
Messengers



# Thank you for your attention!